= MECHANICS =

Trinomial Expansion of Kinetic-Energy Coefficients for Ideal Fluid at Motion of Two Spheres Near Their Contact

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Abstract—The trinomial asymptotic expansions of potential flow kinetic energy in an ideal fluid are constructed for the motion of two spheres of variable radii in the vicinity of their contact. Based on these expansions, it is possible to study the process of two pulsating gas bubbles approaching up to their contact.

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INTRODUCTION

It is convenient to solve the problem of the hydrodynamic interaction of two spheres moving in a potential flow of an ideal fluid using the method of generalized Lagrange coordinates. The coordinates of the centers of spheres and their radii are accepted as generalized coordinates, and their time derivatives are respectively taken as generalized velocities. The kinetic energy of the fluid is the Lagrange function. Such a method of solving the problem was first used by Kelvin and Tait [1], who obtained an expression for the kinetic energy of two spheres balls located far from each other. The exact expression for the kinetic energy of the fluid with solid spheres moving along the line of their centers was derived by Hicks [2] using the image method. Voinov [3–5] generalized the Hicks method for the motion of spheres of variable radii and obtained a general expression for the kinetic energy. In [6], a technique was developed for obtaining asymptotic expansions of the hydrodynamic interaction force with respect to a small gap between the spheres in the vicinity of their contact. With the help of this technique, the main logarithmic asymptotics for the small gap was calculated for solid spheres, and the next term of the expansion was found in [6]. In [7], it was shown how to find the main asymptotics of the forces of hydrodynamic interaction of bodies without using the exact

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solution. To do this, it suffices to single out the singular asymptotics of the velocity field in the vicinity of the contact.

It should be noted that the constant terms in the asymptotic expansion of hydrodynamic interaction forces are very important for studying the convergence of spheres of variable radius along with the logarithmic asymptotics. This study is devoted to constructing these terms. It opens the possibility of studying the process of approaching two pulsating gas bubbles up to their contact and to find the conditions under which no merging of bubbles occurs.

AN EXACT EXPRESSION FOR THE KINETIC ENERGY

The motion of two spheres of radii a_1 and a_2 along the line of their centers in a potential flow of an ideal fluid is considered. The velocities of the centers of spheres are u_1 and u_2 , respectively, and they are directed towards each other (Fig. 1). The kinetic energy T of the fluid is calculated using the potential of the velocity field. Its general expression first obtained by Voinov is given in [4, 5, 8]:

$$\frac{T}{2\pi\rho} = \sum_{i=1}^{2} (A_{i}u_{i}^{2} + Bu_{i}u_{k} + D_{i}\dot{a}_{i}^{2} + E\dot{a}_{i}\dot{a}_{k} + C_{ik}u_{i}\dot{a}_{k} + C_{ii}u_{i}\dot{a}_{i}),$$

 $i \neq k, \quad i, k = 1, 2,$
 $2A_{1} = \frac{a_{1}^{3}}{3} + \sum_{n=1}^{\infty} \left(\frac{a_{1}}{A_{n}^{1}}\right)^{3}, \quad 2B = \sum_{n=1}^{\infty} \left(\frac{a_{2}}{B_{n}^{1}}\right)^{3},$
 $D_{1} = a_{1}^{3} + \sum_{n=1}^{\infty} \frac{a_{1}^{3}}{A_{n}^{1}} \left[1 + ((B_{n}^{2})^{2} - 1)\ln\left(1 - \frac{1}{(B_{n}^{2})^{2}}\right)\right],$
(1)

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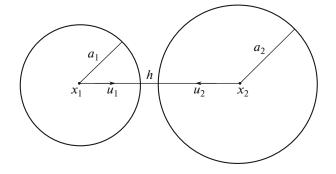


Fig. 1. Problem formulation.

$$E = \frac{(a_1 a_2)^2}{r} + \frac{(a_1 a_2)^2}{a_2} \sum_{n=1}^{\infty} \left(\frac{1}{B_{n+1}^2} - B_n^1 \ln\left(1 + \frac{1}{B_{n+1}^2}B_n^1\right) \right),$$

$$C_{11} = \sum_{n=1}^{\infty} \frac{a_1^3}{(A_n^1)^2 B_n^2}, \quad C_{12} = \sum_{n=1}^{\infty} \frac{a_1^3}{(B_n^2)^2 A_{n-1}^1}.$$

The series for A_2 , D_2 , C_{21} , C_{22} are obtained by permutation of 1 by 2 in the corresponding formulas. The coefficients A_n^i , B_n^i can be calculated from the recurrence relations for the distance between the centers of the spheres $r = x_2 - x_1$:

$$B_n^i = \frac{r}{a_i} A_{n-1}^k - \frac{a_k}{a_i} B_{n-1}^i, \quad A_n^i = \frac{r}{a_i} B_n^k - \frac{a_k}{a_i} A_{n-1}^i, \quad (2)$$

and the initial conditions $A_0^i = 1$, $B_0^i = 0$. It should be noted that B, E are independent of subscripts because the equality $a_1B_n^1 = a_2B_n^2$ holds.

The series of coefficients A_i , B, D_i , E, C_{ii} , C_{ik} for the kinetic energy converge as geometric progressions excluding the points of contact at which they converge as $1/n^3$.

EXPANSION IN THE VICINITY OF CONTACT

For the coefficients A_n^i, B_n^i , recurrence relations (2) can be solved as in [6, 9]:

$$A_n^i = \frac{\tau^{n+1} - \tau^{-n-1} + (\tau^n - \tau^{-n})a_i/a_k}{\tau - \tau^{-1}},$$
$$B_n^i = \frac{(\tau^n - \tau^{-n})r/a_i}{\tau - \tau^{-1}},$$

where τ is the root of the equation $r^2\tau = (a_1\tau + a_2)(a_1 + a_2\tau)$.

These expressions can be simplified by using the substitution $\tau = e^{\varepsilon}$ [10]:

$$A_n^i = \frac{r}{a_k} \frac{\sinh(n\varepsilon + \varepsilon_i)}{\sinh \varepsilon}, \quad B_n^i = \frac{r}{a_i} \frac{\sinh n\varepsilon}{\sinh \varepsilon}.$$

In this case, $\sinh \varepsilon_i = \frac{a_k}{r} \sinh \varepsilon$ and ε is found from the equation $r^2 = a_1^2 + a_2^2 + 2a_1a_2 \cosh \varepsilon$. Thus tending to the contact $r \to a_1 + a_2$, we obtain that $\varepsilon \to 0$ and the following limit expressions for A_n^i, B_n^i at $\varepsilon = 0$:

$$A_n^i = \frac{n + \alpha_k}{\alpha_k}, \quad B_n^i = \frac{n}{\alpha_i}, \quad \alpha_i = \frac{a_i}{a_1 + a_2}.$$
 (3)

Previously, Voinov proposed to determine the derivative $\frac{\partial T}{\partial r}$ in the vicinity of the contact using the following transformations [6]:

$$\sum_{n=1}^{\infty} f(n,\varepsilon) = \sum_{n=1}^{\lfloor 1/\varepsilon \rfloor} f(n,0)$$

$$-\int_{1}^{1/\varepsilon} f(x,0)dx + \int_{1}^{\infty} f(x,\varepsilon)dx + O(\varepsilon).$$
(4)

Applying transformation (4) to A_1 and B, Voinov found that [6]

$$\frac{1}{a'^2} \frac{\partial B}{\partial h} = \frac{1}{4} \ln \frac{h}{2a'} + \frac{1}{3} - \frac{\gamma}{2} + \frac{1}{2} (1 - 3\alpha_1 \alpha_2) \zeta(3) + O(h \ln h),$$
$$\frac{1}{a'^2} \frac{\partial A_1}{\partial h} = \frac{1}{4} \ln \frac{h}{2a'} + \frac{1}{3} - \frac{\gamma}{2}$$
$$+ \frac{1}{2} \sum_{n=1}^{\infty} \left[\frac{1}{n} - \frac{n(n+1)(n-1+3\alpha_2)}{(n+\alpha_2)^4} \right] + O(h \ln h),$$
$$a' = \frac{a_1 a_2}{a_1 + a_2}, \quad h = r - a_1 - a_2$$

and noticed that the series difference $\frac{1}{a'^2} \frac{\partial A_1}{\partial h} - \frac{1}{a'^2} \frac{\partial B}{\partial h}$ for the coefficient A_1 remains finite when the spheres approach their contact. It turns out that all other coefficients of kinetic energy $D_1, E, C_{11}/2, C_{12}/2$, have this property instead of only A_1 . This observation simplifies the deduction of the logarithmic and next terms of the asymptotic expansions for them.

Indeed, this property implies that all the coefficients of the kinetic energy have the same structure in the vicinity of the contact

$$\frac{1}{a'^{2}}\frac{\partial X}{\partial h} = \frac{1}{4}\ln\frac{h}{2a'} + X_{1} + O(h\ln h),$$
 (5)

where X is an arbitrary kinetic-energy coefficient and X_1 depends only on the radii of the spheres and does not depend on h.

The expansions for the coefficients A_1 , B, D_1 , E, $C_{11}/2, C_{12}/2$ are obtained by integrating using formula

DOKLADY PHYSICS Vol. 63 No. 12 2018

$$X(h) = X_0 + \int_0^h \frac{dX}{dh}$$
(6)
= $X_0 + a'^2 h \Big(\frac{1}{4} \Big(\ln \Big(\frac{h}{2a'} \Big) - 1 \Big) + X_1 \Big).$

This is the desired trinomial expansion of an arbitrary coefficient X(h) containing the constant term X_0 and terms like $h \ln h$ and h. We omit the remaining terms of the order of smallness of $h^2 \ln h$ and higher.

The expression for X_0 can be obtained from exact series (1) by substituting limit values (3) in them. We write out the dependencies of X_0 and X_1 for each kinetic-energy coefficient. For the first two coefficients A_1 and B, they have the following form:

$$X_{0} = a_{1}^{3} \left(\frac{1}{6} + \frac{1}{2} \sum_{n=1}^{\infty} \left(\frac{\alpha_{2}}{n + \alpha_{2}} \right)^{3} \right),$$

$$X_{1} = \frac{1}{3} - \frac{\gamma}{2} + \frac{1}{2} \sum_{n=1}^{\infty} \left[\frac{1}{n} - \frac{n(n+1)(n-1+3\alpha_{2})}{(n+\alpha_{2})^{4}} \right],$$

$$A_{1} = X_{0} + a'^{2}h \left(\frac{1}{4} \left(\ln \left(\frac{h}{2a'} \right) - 1 \right) + X_{1} \right),$$

$$X_{0} = \frac{1}{2} a'^{3} \zeta(3), \quad X_{1} = \frac{1}{3} - \frac{\gamma}{2} + \frac{1}{2} (1 - 3\alpha_{1}\alpha_{2}) \zeta(3),$$
(7)

$$2 = X_0 + a'^2 h \left(\frac{1}{4} \left(\ln \left(\frac{h}{2a'} \right) - 1 \right) + X_1 \right).$$

The remaining coefficients $D_1, E, C_{11}/2, C_{12}/2$, are calculated similarly by formula (6) in which only the expressions for X_0 and X_1 vary:

$$\begin{split} D_{1}: X_{0} &= a_{1}^{3} + a_{1}^{3} \sum_{n=1}^{\infty} \frac{\alpha_{2}}{n + \alpha_{2}} \bigg[1 + \bigg(\frac{n^{2}}{\alpha_{2}^{2}} - 1 \bigg) \ln \bigg(1 - \frac{\alpha_{2}^{2}}{n^{2}} \bigg) \bigg], \\ X_{1} &= \frac{1}{3} - \frac{\gamma}{2} + \sum_{n=1}^{\infty} \bigg[\frac{1}{6n} + \frac{(n - 1)(n + 1 - 3\alpha_{2})}{3n\alpha_{2}^{2}} \\ &\times \bigg(1 + \frac{n^{2}}{\alpha_{2}^{2}} \ln \bigg(1 - \frac{\alpha_{2}^{2}}{n^{2}} \bigg) \bigg) + \frac{\alpha_{1}^{2} - (n + 1)\alpha_{2}}{3n(n + \alpha_{2})^{2}} \bigg], \\ E: X_{0} &= \frac{(a_{1}a_{2})^{2}}{a_{1} + a_{2}} \bigg[1 + \sum_{n=1}^{\infty} \frac{1}{n + 1} \\ &\times \bigg(1 - \frac{n(n + 1)}{\alpha_{1}\alpha_{2}} \ln \bigg(1 + \frac{\alpha_{1}\alpha_{2}}{n(n + 1)} \bigg) \bigg) \bigg], \\ X_{1} &= \frac{1}{3} - \frac{\gamma}{2} - \frac{1}{2} + \sum_{n=1}^{\infty} \bigg[-\frac{1}{6(n + 1)} + \frac{n^{2} - 1 + 3\alpha_{1}\alpha_{2}}{3(n + 1)\alpha_{1}\alpha_{2}} \quad (8) \\ &\times \bigg(1 - \frac{n(n + 1)}{\alpha_{1}\alpha_{2}} \ln \bigg(1 + \frac{\alpha_{1}\alpha_{2}}{n(n + 1)} \bigg) \bigg) \bigg) \\ &+ \frac{(\alpha_{1} - \alpha_{2})^{2}}{3(n + 1)(n + \alpha_{1})(n + \alpha_{2})} \bigg], \end{split}$$

DOKLADY PHYSICS Vol. 63 No. 12 2018

$$C_{11}/2: X_0 = \frac{a_1^3}{2} \sum_{n=1}^{\infty} \frac{\alpha_2^3}{n(n+\alpha_2)^2},$$

$$X_1 = \frac{1}{3} - \frac{\gamma}{2}$$

$$+ \sum_{n=1}^{\infty} \left[\frac{2\alpha_2 n^2 + 3(\alpha_1^3 + \alpha_2^3 + \alpha_2^2)n + \alpha_2(\alpha_1^3 + \alpha_2^3 + 3\alpha_2^2)}{6n(n+\alpha_2)^3} \right],$$

$$C_{12}/2: X_0 = \frac{a_1^3}{2} \sum_{n=1}^{\infty} \frac{\alpha_2^3}{n^2(n-1+\alpha_2)},$$

$$X_1 = \frac{1}{3} - \frac{\gamma}{2} + \sum_{n=1}^{\infty} \left[\frac{-\alpha_1 n^2 + (3n-2\alpha_1)(\alpha_1^3 + \alpha_2^3)}{6n^2(n-\alpha_1)^2} \right]$$

where $\gamma \approx 0.577216$ is the Euler constant and the remaining coefficients of the kinetic energy can be obtained by permutation of 1 and 2.

All series for X_0 in formulas (7) and (8) are cited in [3–6]. The series for X_1 in formulas (8) were not calculated previously, but their contribution to the force is comparable with that from the series for X_0 .

PARTICULAR CASES

For spheres of identical radii $a_1 = a_2 = a$, the following substitution $\alpha_1 = \alpha_2 = 1/2, a' = a/2$ should be made in the kinetic energy coefficients formulas. Approximately calculating the numerical series in them, we obtain

$$A_{1} = A_{2} = a^{3} \times 0.192567$$

$$+ \frac{a^{2}}{4} h \left(\frac{1}{4} \left(\ln \left(\frac{h}{a} \right) - 1 \right) + 0.403378 \right),$$

$$B = a^{3} \times 0.0751286$$

$$+ \frac{a^{2}}{4} h \left(\frac{1}{4} \left(\ln \left(\frac{h}{a} \right) - 1 \right) + 0.194983 \right),$$

$$D_{1} = D_{2} = a^{3} \times 1.05634$$

$$+ \frac{a^{2}}{4} h \left(\frac{1}{4} \left(\ln \left(\frac{h}{a} \right) - 1 \right) + 0.148441 \right),$$

$$E = a^{3} \times 0.52088$$

$$+ \frac{a^{2}}{4} h \left(\frac{1}{4} \left(\ln \left(\frac{h}{a} \right) - 1 \right) - 0.581959 \right),$$

$$C_{11} = C_{22} = a^{3} \times 0.0731523$$

$$+ \frac{a^{2}}{4} h \left(\frac{1}{2} \left(\ln \left(\frac{h}{a} \right) - 1 \right) + 0.596124 \right),$$

$$C_{12} = C_{21} = a^{3} \times 0.281914$$

$$+ \frac{a^{2}}{4} h \left(\frac{1}{2} \left(\ln \left(\frac{h}{a} \right) - 1 \right) + 0.003238 \right).$$
(9)

At the contact of spheres h = 0 the numerical values of all coefficients were calculated in [3–5, 10]. No

numerical values of the coefficients of h for D_i , E, C_{ii} , C_{ik} were calculated previously.

Of interest is the case of the motion of spheres of identical radii and velocities $\dot{a}_1 = \dot{a}_2 = \dot{a}$, $u_1 = u_2 = u$, respectively. In this case, the kinetic energy determines the force acting on the sphere near the plane and has the form

$$\frac{T}{2\pi\rho} = Pu^2 + 2Qu\dot{a} + R\dot{a}^2.$$
(10)

The numerical values of the coefficients can be found using Eqs. (9)

$$P = 2A_{1} + 2B$$

= $a^{2}[0.53539a + 0.0491804h + 0.25h\ln(h/a)],$
 $2Q = 2C_{12} + 2C_{11}$
= $a^{2}[0.710132a - 0.200319h + 0.5h\ln(h/a)],$
 $R = 2D_{1} + 2E$
= $a^{2}[3.15443a - 0.466759h + 0.25h\ln(h/a)].$ (11)

HYDRODYNAMIC FORCE NEAR CONTACT

The hydrodynamic force acting on the first sphere can be found using the Lagrange formula:

$$F_{1} = -\frac{d}{dt}\frac{\partial T}{\partial \dot{x}_{1}} + \frac{\partial T}{\partial x_{1}} = -\frac{d}{dt}\frac{\partial T}{\partial u_{1}} - \frac{\partial T}{\partial h}$$

Using Eq. (1), we obtain that

$$\frac{F_1}{-2\pi\rho} = \frac{d}{dt} (2A_1u_1 + 2Bu_2 + C_{11}\dot{a}_1 + C_{12}\dot{a}_2) + \frac{\partial A_1}{\partial h}u_1^2 + 2\frac{\partial B}{\partial h}u_1u_2 + \frac{\partial A_2}{\partial h}u_2^2 + \frac{\partial D_1}{\partial h}\dot{a}_1^2 + 2\frac{\partial E}{\partial h}\dot{a}_1\dot{a}_2 + \frac{\partial D_2}{\partial h}\dot{a}_2^2 + \frac{\partial C_{11}}{\partial h}u_1\dot{a}_1 + \frac{\partial C_{12}}{\partial h}u_1\dot{a}_2 + \frac{\partial C_{21}}{\partial h}u_2\dot{a}_1 + \frac{\partial C_{22}}{\partial h}u_2\dot{a}_2.$$

Taking into account Eqs. (7) and (8), we can single out the main asymptotics for the force at $h \rightarrow 0$:

$$\frac{F_1}{2\pi\rho} = \frac{d}{dt} \left(\frac{1}{2} a'^2 h \ln\left(\frac{h}{2a'}\right) \dot{h} \right) - \frac{1}{4} a'^2 \ln\left(\frac{h}{2a'}\right) \dot{h}^2,$$

which agrees with [7]. If the distance between the spheres does not change, that is, $\dot{h} = -u_1 - u_2 - \dot{a}_1 - \dot{a}_2 = 0$, the logarithmic feature disappears as was indicated by Voinov in [4] without derivation.

We obtain an exact formula for the force acting on the sphere of variable radius a near the plane using Eq. (10) and its approximate expression using Eq. (11):

$$\frac{F_{1}}{-2\pi\rho} = \frac{d}{dt}(Pu + Q\dot{a}) + \frac{\partial P}{\partial h}u^{2} + 2\frac{\partial Q}{\partial h}u\dot{a} + \frac{\partial R}{\partial h}\dot{a}^{2}$$
$$= \frac{d}{dt}\left[a^{2}\left((0.355066a - 0.100159h)\dot{a} + (0.53539a + 0.0491804h)u - h\frac{1}{4}\ln\left(\frac{h}{a}\right)\frac{\dot{h}}{2}\right)\right]$$
$$+ a^{2}\left(0.29918u^{2} + 0.299681u\dot{a} - 0.216759\dot{a}^{2} + \frac{1}{4}\ln\left(\frac{h}{a}\right)\frac{\dot{h}^{2}}{4}\right),$$

where h/2 is the distance between the sphere and the plane, $u + \dot{a} = -\dot{h}/2$.

At $\dot{h} = 0$, we obtain $u = -\dot{a}$ and

$$\frac{F_1}{-2\pi\rho} = \frac{d}{dt} ((Q-P)\dot{a}) + \frac{\partial(P-2Q+R)}{\partial h} \dot{a}^2.$$

From this formula at h = 0, we find the force acting on the sphere with constant contact with the plane:

$$F_{1} = 2\pi\rho \left(0.180324 \frac{d}{dt} (a^{3}\dot{a}) + 0.217259a^{2}\dot{a}^{2} \right)$$
$$= 2\pi\rho a^{2} (0.758232\dot{a}^{2} + 0.180324a\ddot{a}).$$

In [4] the force is given in the following form:

$$F_1 = 2\pi\rho a^2 (0.75821\dot{a}^2 + 0.18032a\ddot{a}).$$

In case the bubble radius changes periodically, the

average force is $\langle F_1 \rangle = 2\pi\rho a^2 \times 0.217259 \langle \dot{a}^2 \rangle$, which corresponds to the attraction force.

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REFERENCES

- 1. W. T. B. Kelvin and P. G. Tait, *Treatise on Natural Philosophy* (Clarendon Press, London, 1867), Vol. 1.
- W. M. Hicks, Phil. Trans. Royal Soc. London 171, 455 (1880).
- O. V. Voinov. Vestn. MGU, Ser. Mat., Mekh., Astron., No. 5, 83 (1969).
- 4. O. V. Voinov, in *Sci. Conf., Theses Docl.* (In-t Mekh. MGU, M., 1970) [in Russian].
- O. V. Voinov and A. G. Petrov, Itogi Nauki i Tekhniki. Mekhanika Zhidkosti i Gaza 10, 86 (1976).
- 6. O. V. Voinov, Prikl. Mat. Mekh. 33 (4), 659 (1969).
- A. G. Petrov and A. A. Kharlamov, Fluid Dyn. 48 (5), 577 (2013).
- 8. A. G. Petrov, *Analytical Hydrodynamics* (Fizmatlit, Moscow, 2009) [in Russian].
- 9. C. Neumann, *Hydrodynamische Untersuchungen* (Teubner, Stuttgart, 1883).
- 10. A. G. Petrov, Fluid Dyn. 46 (4), 579 (2011).

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DOKLADY PHYSICS Vol. 63 No. 12 2018