

# Solubility of Nitrogen, Carbon, and Hydrogen in FeO–Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> Melt and Liquid Iron Alloy: Influence of Oxygen Fugacity

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**Abstract**—Reactions of nitrogen, carbon, and hydrogen with FeO–Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> melts, liquid iron alloys, and graphite were investigated at 4 GPa, 1550°C, and  $fO_2$  values 1.5–3.0 orders of magnitude below  $fO_2(IW)$ . A number of features important for the understanding of the formation conditions of volatile nitrogen compounds during melting of the Earth's early reduced mantle were revealed. The nitrogen content of melt increases with decreasing  $fO_2$  from 0.96 wt % at  $\Delta\log fO_2(IW) = -1.4$  to 4.1 wt % at  $\Delta\log fO_2(IW) = -3.0$ , whereas the hydrogen content of melt is weakly dependent on  $fO_2$  and lies within 0.40–0.47 wt %. The carbon content is approximately 0.3–0.5 wt %. The IR and Raman spectroscopy of the glasses indicated that the dissolution of nitrogen, carbon, and hydrogen in silicate liquids is accompanied by the formation of NH<sub>3</sub>, N<sub>2</sub>, and CH<sub>4</sub> molecules, as well as NH<sub>2</sub><sup>-</sup>, NH<sub>2</sub><sup>+</sup>, NH<sub>4</sub><sup>+</sup>, and CH<sub>3</sub><sup>-</sup> complexes. Hydrogen is dissolved in melts as OH<sup>-</sup>, H<sub>2</sub>O, and H<sub>2</sub>. The experiments also demonstrated the presence of species with C=O double bonds in the melts. It was found that the solubility of nitrogen in FeO–Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> melts increases in the presence of carbon owing to the formation of species with C–N bonds in the silicate liquid. One of the most remarkable features of nitrogen, carbon, and hydrogen interaction with FeO–Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> melts is a significant change in the proportions of N–C–H–O species at  $fO_2$  2–3 orders of magnitude below  $fO_2(IW)$ . Under these conditions, a sharp decrease in the contents of NH<sub>4</sub><sup>+</sup>, NH<sub>2</sub><sup>+</sup> (O–NH<sub>2</sub>), OH, H<sub>2</sub>O, and CH<sub>4</sub> is accompanied by enrichment in NH<sub>2</sub><sup>-</sup> (≡Si–NH<sub>2</sub>) and NH<sub>3</sub>. As a result, NH<sub>3</sub> becomes the dominant nitrogen species in the melt. The investigation revealed high nitrogen solubility in iron alloys at  $fO_2 < fO_2(IW)$ . The nitrogen content increases from 2.47 wt % at  $\Delta\log fO_2(IW) = -1.4$  to 3.63 wt % at  $\Delta\log fO_2(IW) = -3.0$ . The carbon content of N–C–Fe alloys ranges from 2.3 to 3.8 wt % and decreases with decreasing  $fO_2$ . The siderophile behavior of nitrogen at  $fO_2 < fO_2(IW)$  suggests that part of nitrogen could be dissolved in iron alloys during large-scale melting of the early reduced mantle with subsequent nitrogen burial in the Earth's metallic core. It was suggested that the self-oxidation of magmas in the Earth's early mantle with the release of reduced N–C–H–O volatiles could be one of the reasons of extensive nitrogen degassing.

**Keywords:** nitrogen, carbon, hydrogen, oxygen fugacity, Earth's early mantle, magma ocean, iron alloy, degassing

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## PROBLEM FORMULATION

The behavior of nitrogen during the chemical differentiation of terrestrial materials is still a topic of debate (e.g., Javoy, 1997; Cartigny et al., 2001; Marty and Dauphas, 2003; Marty, 2012). Nitrogen is the main component of the Earth's present-day atmosphere, but the early atmosphere could be dramatically different (Catling and Claire, 2005; Shaw, 2008; Zahnle et al., 2010). Its gas components were probably NH<sub>3</sub>, CH<sub>4</sub>,

H<sub>2</sub>, and H<sub>2</sub>O. It was suggested that the reduced atmosphere could be a source of prebiotic molecules, whose formation had led to the origin of life on Earth (e.g., Galimov, 2004).

The evaluation of the geochemical history of nitrogen and formation of its possible mantle reservoirs is strongly hindered by the poor knowledge of the influence of temperature, pressure, and redox conditions on the distribution of nitrogen among the mantle, core, products of planetary melting, and a deep fluid phase.

Important characteristics of nitrogen behavior can be gained from the experimental investigations of nitrogen dissolution in minerals, silicate and metallic melts, and fluids under mantle pressures. In particular, it was shown that nitrogen could be incorporated under mantle conditions in some silicates and spinels as the osbornite component or in aluminum-bearing silicates (Watenphul et al., 2009). The stability of iron nitrides at high temperatures and pressures (Adler and Williams, 2005; Hasegawa and Yagi, 2005) makes them viable candidates for the formation of nitrogen reservoirs in the mantle and core. The experiments of Watenphul et al. (2010) and Li et al. (2013) suggested that nitrogen could be dissolved in the products of magma ocean crystallization and be retained, thus, in the minerals of the Earth's early mantle. Li and Keppler (2014) concluded that  $\text{NH}_3$  is the dominant nitrogen compound in the deep parts of the reduced mantle.

One of the endogenic sources of nitrogen, carbon, and hydrogen in the early atmosphere is the gas exchange between the zones of extensive mantle melting (magma ocean) and the surface of the early Earth. The magma ocean was connected with the forming planetary core through the gravitational migration of metallic iron in the molten silicate mantle (Li and Agee, 1996; Richter and Drake, 2000; Wood et al., 2006; Frost et al., 2008). The segregation of the core had to be accompanied by the formation of volatile compounds in the zones of large-scale melting of the early Earth; their composition was controlled by the interaction of C, N, H, and O with silicate and metallic melts. Therefore, the gas regime of the early Earth had to be related to the character of dissolution of volatile components in silicate and metallic melts at  $f\text{O}_2$  values 4–8 orders of magnitude lower than those characteristic of present-day mantle magmas. The speciation of N, C, H, and O in silicate and metallic phases can be experimentally determined at  $f\text{O}_2$  values prevailing during the formation and evolution of the magma ocean.

The influence of pressure, oxygen fugacity ( $f\text{O}_2$ ), and hydrogen fugacity ( $f\text{H}_2$ ) on the solubility of N–H–O volatile compounds in silicate melts was studied by Fogel (1994), Libourel et al. (2003), Miyazaki et al. (2004), Roskosz et al. (2006), Mysen et al. (2008, 2014); Mysen and Fogel (2010), Mysen (2013), and Kadik et al. (2011, 2013). Under reducing conditions, the chemical interaction of nitrogen with hydrogen and silicate liquid components produces N–H species ( $\text{NH}_3$ ,  $\text{NH}_2^-$ , and  $\text{NH}_2^+$ ), among which the  $\text{NH}_3$  molecule can be dominant at  $f\text{O}_2$  below the Fe–FeO (IW) buffer equilibrium. Hence, it can be suggested that reduced magmas were the source of  $\text{NH}_3$  for the reduced atmosphere (e.g., Roskosz et al., 2006; Kadik et al., 2011, 2013; Mysen, 2012; Li and Keppler, 2014). The experiments of Kadik et al. (2011, 2013) and Roskosz et al. (2013) established for the first time a significant decrease in nitrogen partition coefficient between metal and silicate liquid at high pressures. This effect

could be the reason for the depletion of nitrogen relative to other volatiles in the bulk Earth composition, which was detected by Marty (2012). Experimental investigations of the solubility of C–O–H compounds in silicate melts at  $f\text{O}_2 < f\text{O}_2$  (IW) established the formation of C–H species in them ( $\text{CH}_4$ ,  $\text{CH}_3^-$ , and complexes with C=O double bonds) (Kadik et al., 2004, 2011, 2014; Mysen et al., 2009, 2011; Hirschmann, 2012; Ardia et al., 2013; Dasgupta et al., 2013; Wetzel et al., 2013; Stanley et al., 2014). These experiments were used to develop models describing the dependence of the contents of C–H–O species in planetary magmas on the redox conditions of their formation.

The separate investigation of the solubility of N–H–O and C–H–O volatiles in melts leaves unanswered the question on possible coupling of nitrogen and carbon effects on the formation of N–C–H–O molecules and complexes, in particular, on the proportions of  $\text{NH}_3$ ,  $\text{CH}_4$ , and  $\text{H}_2\text{O}$  in the melts. By analogy with metallurgical systems, the solubilities of carbon and nitrogen in iron alloys are expected to be interrelated (e.g., Bouchard and Bale, 1995; Ma, 2001). In particular, the solubility of  $\text{N}_2$  in iron alloy decreases significantly at increasing C content. Another poorly understood factor that may have a significant influence on the behavior of nitrogen and carbon in reduced magmas is the incorporation of Si in the structure of iron alloy. It results in a decrease in the solubility of carbon and nitrogen in it (e.g., Bouchard and Bale, 1995; Ma, 2001). This effect is especially important at  $\Delta\log f\text{O}_2(\text{IW}) < -2$ , because the Si content of iron alloy in equilibrium with silicate melt may be as high as several weight percent at such  $f\text{O}_2$  values (e.g., Gessman et al., 2001; Ricolleau et al., 2011).

Our experimental investigations aimed at elucidating the character of simultaneous dissolution of nitrogen, carbon, and hydrogen in iron-bearing magmatic melts and equilibrium liquid iron alloy at  $f\text{O}_2$  values two–five orders of magnitude below  $f\text{O}_2$  (IW). Such extremely low  $f\text{O}_2$  values are thought to be characteristic of the early chemical differentiation of the Earth during its melting and segregation of iron phase (e.g., Wood et al., 2006; Javoy et al., 2010; Frost et al., 2008). Despite the short duration of this period in Earth history, it probably had a profound influence on the contents of volatile components in the mantle after the completion of planetary core formation.

## EXPERIMENTAL AND ANALYTICAL METHODS

### *Starting Materials*

In order to evaluate the character of dissolution of nitrogen and its hydrogen- and oxygen-bearing compounds in the reduced mantle of the early Earth, melts of mafic–ultramafic composition should have been used; however, we selected a model FeO– $\text{Na}_2\text{O}$ –

**Table 1.** Compositions of starting NaAlSi<sub>3</sub>O<sub>8</sub> + FeO + Si<sub>3</sub>N<sub>4</sub> mixtures used in the experiments, wt %

Mixture no.	Starting mixture	SiO <sub>2</sub>	Al <sub>2</sub> O <sub>3</sub>	FeO	Na <sub>2</sub> O	Si	N
L5	99%AbFeO + 1%Si <sub>3</sub> N <sub>4</sub>	54.4	15.4	19.8	9.4	0.6	0.4
L6	97%AbFeO + 3%Si <sub>3</sub> N <sub>4</sub>	53.3	15.1	19.4	9.2	1.8	1.2
L7	95%AbFeO + 5%Si <sub>3</sub> N <sub>4</sub>	52.2	14.8	19.0	9.0	3.0	2.0
L8	93%AbFeO + 7%Si <sub>3</sub> N <sub>4</sub>	51.1	14.5	18.6	8.8	4.2	2.8

AbFeO is a mixture of albite glass (80 wt %) and FeO (20 wt %).

SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melt for our study. As was previously shown (Kadik et al., 2011, 2013), silicate melts of such compositions are readily quenched to glasses. This enables the application of spectral methods for the determination of the mechanisms of dissolution of volatile components in melts at high temperatures and pressures.

The starting material for experiments was a finely dispersed mixture of albite glass NaAlSi<sub>3</sub>O<sub>8</sub> (80 wt %) and FeO (20 wt %), which was blended with 1, 3, 5, or 7 wt % of powdered silicon nitride (Si<sub>3</sub>N<sub>4</sub>). The latter imposed low *f*O<sub>2</sub> values during experiments owing to the oxidation of Si<sub>3</sub>N<sub>4</sub> interacting with iron-bearing silicate melt and H<sub>2</sub>. Melting reactions at high pressures were investigated in the presence of a graphite disk, which provided conditions of system saturation with respect to free carbon. The albite glass was prepared from high-purity SiO<sub>2</sub> and Al<sub>2</sub>O<sub>3</sub> reagents annealed before weighing at 1100°C and Na<sub>2</sub>CO<sub>3</sub> dried at 200°C. An oxide mixture of the stoichiometric albite composition was ground in an agate mortar under ethanol for more than 3 h. The oxide–carbonate mixture was heated up to the temperature of Na<sub>2</sub>CO<sub>3</sub> decarbonation (~900°C) at a rate of 10°C/min. Then, the temperature increased up to 1400°C, and the sample was held at this temperature for 1 h and quenched in air. The quenching rate was 100–200°C/s. At least three separate fragments of the glass were analyzed to check for its homogeneity. The glass was ground to 20 μm and stored in a desiccator until use in the experiments. The FeO powder was dried at 200°C for 24 h and added in a desired proportion to the powdered glass. The mixture was placed into an agate mortar and ground under ethanol for 3 h, after which the glass + FeO + Si<sub>3</sub>N<sub>4</sub> mixture was heated to 300°C, kept at this temperature for 6 h, and stored subsequently in a desiccator. The chemical compositions of the starting mixtures are given in Table 1.

### Experiments

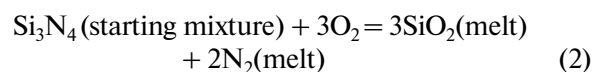
Experiments were conducted using an anvil with hole apparatus with a working volume of 6 cm<sup>3</sup> at 4 GPa and 1550°C (Litvin, 1979, 1991; Kadik et al., 2013) under controlled hydrogen fugacity (Kadik et al., 2004, 2013). The experimental setup is a 2000-t uniaxial press equipped with a high-pressure system consisting of two tungsten carbide anvils with truncated cone-shaped

holes. The anvils are axially aligned and separated by a pyrophyllite gasket. The resulting cavity comprises a solid-phase cell with a sample assembly. Pressure was measured at ambient temperature with an accuracy of ±0.02 GPa using Bi and Tl sensors (Kadik et al., 2013), and temperature was measured using a Pt30%Rh/Pt6%Rh thermocouple inserted radially into the center of the cell between two capsules with an accuracy of ±5°C at 1500°C and ±10°C at 1600°C. During the experiments, *f*H<sub>2</sub> was buffered through the dissociation of H<sub>2</sub>O absorbed by fine-grained limestone or pyrophyllite outside the graphite heater up to the attainment of equal chemical potentials of H<sub>2</sub> in the sample and in the external solid oxygen buffer (e.g., Eugster and Wones, 1962; Kadik et al., 2004). It is assumed that *f*O<sub>2</sub> is buffered by the tungsten carbide anvils of the high-pressure system in accordance to the reaction WC + O<sub>2</sub> = WO<sub>2</sub> + C (WCWO buffer) (Taylor and Foley, 1989). The presence of traces of H<sub>2</sub>O and the graphite heater allows us to expect the formation of a C–O–H fluid phase saturated in elemental carbon. The H<sub>2</sub>O content of C–O–H fluid at 4 GPa, 1550°C, and *f*O<sub>2</sub> of the WCWO buffer is probably close to the maximum. The mole fraction of H<sub>2</sub> in the C–O–H fluid is up to 0.015, which corresponds to log*f*H<sub>2</sub> = 2.78 bar.

For each experiment, 200–300 mg of the prepared mixture was loaded into a Pt capsule. To prevent interaction between Fe-bearing melt and the walls of the Pt capsule, the sample was isolated by a 0.5-mm thick W foil. The duration of experiments was 30–60 min. Quenching was performed by shutting off the power to the heater at an initial rate of 200°C/s. Oxygen fugacity in the experimental sample was controlled by a redox reaction between externally buffered hydrogen and components of Fe-bearing melt, which was reduced with the release of oxygen and metallic Fe:



The initial silicon nitride (Si<sub>3</sub>N<sub>4</sub>) is unstable under the experimental conditions and completely consumed via the oxidation reaction



and the released nitrogen reacts subsequently with the components of silicate melt and hydrogen. The redox reaction results in the attainment of the three-phase

equilibrium iron alloy + silicate melt + graphite (graphite disk in the Pt capsule). Owing to the interaction of  $\text{Si}_3\text{N}_4$  with silicate melt and  $\text{O}_2$  produced by FeO reduction, the  $f\text{O}_2$  value within the Pt capsule appears to be significantly below  $f\text{O}_2(IW)$ . The decrease of  $f\text{O}_2$  is proportional to the amount of  $\text{Si}_3\text{N}_4$  in the initial mixture.

#### *Electron Microprobe Analysis*

Electron microprobe analysis was conducted at the Vernadsky Institute of Geochemistry and Analytical Chemistry, Russian Academy of Sciences using a CAMECA Camebax SX-100 instrument with four vertical spectrometers in two steps. The measurement of major elements was followed by the analysis of carbon, which required a special procedure. Carbon was measured using a PC2 pseudocrystal with  $2d = 97.46 \text{ \AA}$ . In our case, the spectral peaks of carbon did not overlap with those of other measured elements. The standards for carbon were pure carbon and WC for the analysis of the metallic phase and dolomite,  $\text{CaMgCO}_3$ , for the silicate phase. In order to improve the statistical parameters of carbon analysis, the counting time on the peak was increased up to 30 s. The detection limit for carbon was  $\sim 0.2 \text{ wt } \%$ .

#### *Ion Microprobe Analysis*

Hydrogen was analyzed in glasses by secondary ion mass spectrometry using a CAMECA IMS-4F ion microprobe at the Yaroslavl' Filial of the Physical Technological Institute, Russian Academy of Sciences. Secondary ion beams of  $^1\text{H}^+$  and  $^{30}\text{Si}^+$  were used as analytical and standard signals (Smirnov et al., 1995; Sobolev and Chaussidon, 1996). To maintain charge neutrality, the sample was coated with a 20-nm Au film. The H/Si ratio of glasses was estimated using the curve of Sobolev and Chaussidon (1996), which was calibrated on the basis of the  $\text{H}_2\text{O}/\text{SiO}_2$  (wt %) ratios of glasses containing 0.09–8.00 wt %  $\text{H}_2\text{O}$ . The processing of analytical results included the subtraction of the effect of the silicate matrix at  $\text{SiO}_2$  contents within 49–71 wt %.

#### *Calculation of Oxygen Fugacity*

Oxygen fugacity in experiments was calculated relative to the iron–wüstite (IW) buffer equilibrium:  $\Delta\log f\text{O}_2(IW) = 2\log(a_{\text{FeO}}/a_{\text{Fe}})$ , where  $a_{\text{FeO}}$  is the activity of FeO in silicate melt, and  $a_{\text{Fe}}$  is the activity of Fe in liquid metallic alloy. The activities  $a_{\text{FeO}}$  and  $a_{\text{Fe}}$  were estimated from the activity coefficients  $\gamma$  and mole fractions  $x$  of the components,  $a = x\gamma$ , which yields  $\Delta\log f\text{O}_2(IW) = 2\log(x_{\text{FeO}}/x_{\text{Fe}}) + 2\log(\gamma_{\text{FeO}}/\gamma_{\text{Fe}})$ . The activity coefficient of Fe ( $\gamma_{\text{Fe}}$ ) was taken to be one, because at Fe contents higher than  $\sim 80 \text{ wt } \%$ , its behavior in liquids is governed mainly by Raoult's law (Hult-

gren, 1973). Similar to the study of Kadik et al. (2014),  $\gamma_{\text{FeO}}$  was taken to be 1.3.

## EXPERIMENTAL RESULTS

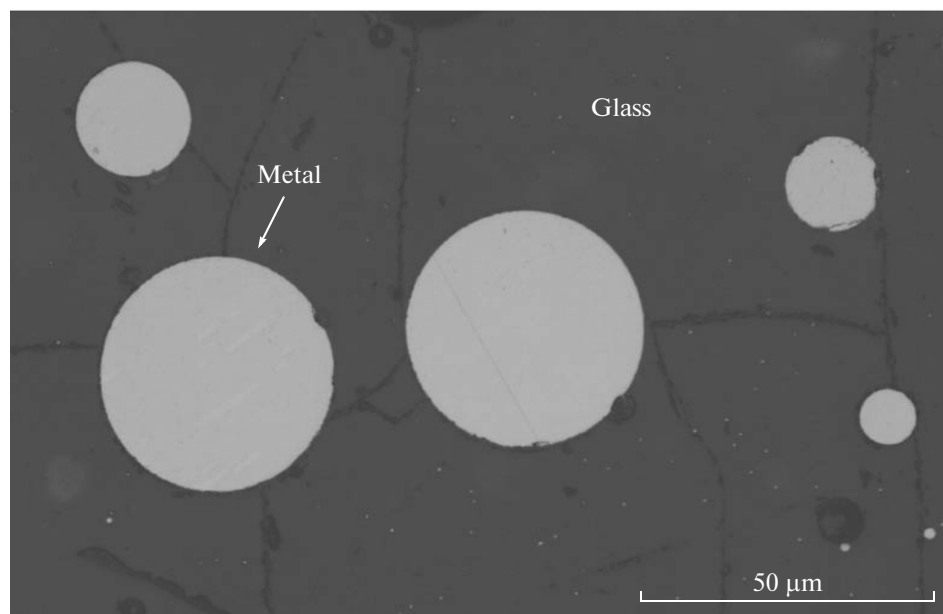
The experimental products are transparent glasses with embedded iron alloy globules 30–50  $\mu\text{m}$  in size, and smaller iron alloy globules, no larger than 1–2  $\mu\text{m}$ . Graphite crystals  $\sim 2$ –5  $\mu\text{m}$  in size were occasionally observed on the surface of Fe globules (Fig. 1). The graphite disk that was used in the experiments as a source of carbon is preserved and shows sharp contacts with glass. The color of the glass varies from greenish in experiments at  $\Delta\log f\text{O}_2(IW)$  from  $-1.4$  to  $-2.3$  to colorless in experiment at  $\Delta\log f\text{O}_2(IW) = -3.0$ . The spherical shape of the Fe metal phase indicates that it was liquid during the experiment. The microscopic texture of the globules resulted from quench crystallization of liquid metal during cooling (Fig. 1).

Gas inclusions were never observed in the glasses under a microscope. This indicates that melt saturation with respect to volatile compounds of nitrogen, hydrogen, and carbon with the formation of a fluid phase was not reached in the experiments.

The chemical compositions of experimental glasses show a considerable decrease in FeO content with decreasing  $f\text{O}_2$  (Table 2). This results from the reduction of FeO in the melt with the formation of Fe alloy in accordance with reaction (1). An increase in  $\text{SiO}_2$  content in the melt with decreasing  $f\text{O}_2$  is caused by the simultaneous occurrence of reactions (1) and (2), which lead to the removal of FeO from melt owing to its reduction and the addition of a certain amount  $\text{SiO}_2$  to the melt owing to  $\text{Si}_3\text{N}_4$  oxidation in the starting mixture. The chemical composition of the glasses and optical examination indicate complete  $\text{Si}_3\text{N}_4$  oxidation via reaction (2). The distribution of the components of glass and metal is relatively uniform, which indicates the attainment of chemical equilibrium during the experiments.

The content of nitrogen in melt increases with decreasing  $f\text{O}_2$  from 0.96 wt % N at  $\Delta\log f\text{O}_2(IW) = -1.4$  to 4.1 wt % N at  $\Delta\log f\text{O}_2(IW) = 3.0$ , whereas the hydrogen content of glass is insensitive to  $f\text{O}_2$  and lies within 0.40–0.47 wt %. The carbon content is approximately 0.3–0.5 wt %. The nitrogen content of iron alloy increases from 0.96 wt % N at  $\Delta\log f\text{O}_2(IW) = -1.4$  to 6.1 wt % N at  $\Delta\log f\text{O}_2(IW) = -3.0$ . The carbon content decreases with decreasing  $f\text{O}_2$  and ranges from 2.3 to 3.8 wt %.

The difference between FeO contents in FeO– $\text{Na}_2\text{O}$ – $\text{SiO}_2$ – $\text{Al}_2\text{O}_3$  melts and in the starting silicate mixture (20 wt %) can be used to estimate the amounts of Fe alloy and oxygen formed through melt reduction (reaction 1) and the amount of oxygen consumed for  $\text{Si}_3\text{N}_4$  oxidation (reaction 2) (Fig. 2). These estimated indicated that the amount of oxygen produced by FeO



**Fig. 1.** Reflected-light photomicrograph of glass with metallic iron droplets and graphite crystals in the quench products of an experiment at 4 GPa, 1550°C, and  $\Delta\log fO_2(IW) = -2.3$ .

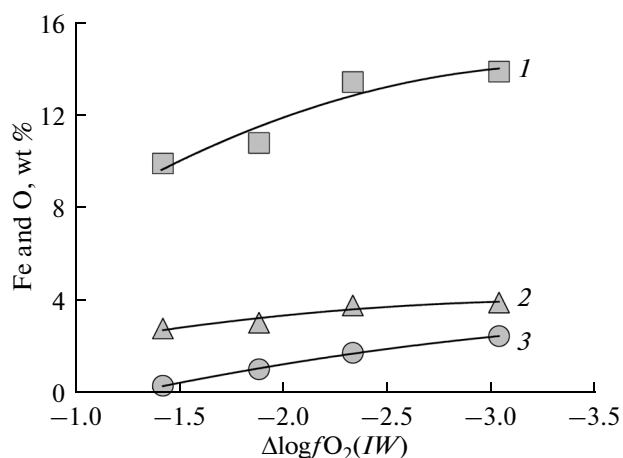
reduction is higher than the amount of oxygen necessary for Si oxidation with  $SiO_2$  formation. It can be suggested that the excess oxygen corresponds to its amount consumed by the formation of nitrogen, carbon, and hydrogen species in the silicate liquid. The most important component is  $H_2O$ . Part of oxygen could be dissolved in the Fe alloy. Under the conditions of our experiments, the oxygen content of iron alloy could be as high as ~0.5 wt % (e.g., Ricolleau et al., 2011).

#### SOLUBILITY OF REDUCED N–H–O AND C–H–O VOLATILES IN SILICATE MELTS

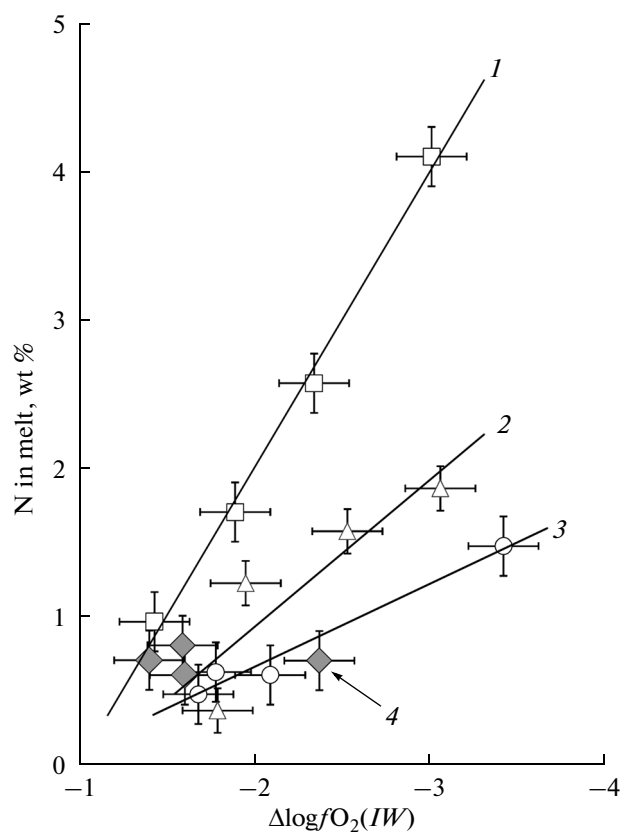
One of the most remarkable features of nitrogen interaction with silicate melts is significant dependence of its solubility on  $fO_2$  and  $fH_2$ . At  $fO_2$  values of the modern mantle approaching the level of the fayalite–magnetite–quartz buffer,  $fO_2$  (FMQ), nitrogen solubility in basaltic melts is rather low (less than 1 ppm) because of the molecular mechanism of  $N_2$  dissolution in the silicate melt structure (e.g., Libourel et al., 2003). However, the nitrogen solubility increases considerably at  $fO_2$  values below  $fO_2(IW)$ , and the total nitrogen content of melt may be as high as hundreds of ppm at atmospheric pressure (Libourel et al., 2003; Miyazaki et al., 2004). At high pressures and  $fO_2 < fO_2(IW)$ , the bulk nitrogen solubility in melt in the presence of hydrogen is up to several weight percent (Roskosz et al., 2006, 2013; Mysen et al., 2008, 2013, 2014; Mysen and Fogel, 2010; Kadik et al., 2011, 2013). The dramatic increase in nitrogen solubility is related to a change in the mechanism of its dissolution in silicate melts, which is controlled by the chemical interaction of nitrogen with

hydrogen and Si–O groups with the formation of the molecular and ionic species  $NH_3$ ,  $NH_2^-$ ,  $NH_2^+$ , and  $NH_4^+$ .

Similar to nitrogen, carbon solubility in silicate melts depends on  $fO_2$  and  $fH_2$ . At  $fO_2$  values near  $fO_2(FMQ)$ , the dominant C–H–O species in magmas



**Fig. 2.** Amounts of Fe and O in the FeO–Na<sub>2</sub>O–SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melt during experiments in accordance with reaction (1). (1) Amount of Fe, (2) amount of O, and (3) amount of O consumed for  $Si_3N_4$  oxidation in the starting mixture and  $SiO_2$  formation via reaction (2). The difference between variants 1 and 2 corresponds to the amount of oxygen consumed by the formation of H–O complexes in melt and oxygen dissolution in liquid iron alloy.



**Fig. 3.** Influence of  $fO_2$  on nitrogen content in FeO–Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> melts at the dissolution of N–C–H–O and N–H–O volatile compounds: (1) 4 GPa, 1550°C, melt + N–C–H–O volatile components (this study); (2) 4 GPa, 1550°C, melt + N–H–O volatile components (Kadik et al., 2013); (3) 1.5 GPa, 1400°C, melt + N–C–H–O volatile components (Kadik et al., 2011); and (4) model chondritic melt + N<sub>2</sub> at 1.8–8.2 GPa, 2500–2650 (Roskosz et al., 2013).

are CO<sub>2</sub> and CO<sub>3</sub><sup>2-</sup> (e.g., Holloway, 1981; Morizet et al., 2010). At  $fO_2 < fO_2(IW)$ , the mechanism of carbon dissolution changes fundamentally, and the speciation of carbon in silicate melts is dominated by CH<sub>4</sub>, CH<sub>3</sub>, and carbonyl groups, whereas CO<sub>2</sub> and CO<sub>3</sub><sup>2-</sup> occur in minor amounts (e.g., Kadik et al., 2004, 2014; Mysen et al., 2011; Mysen, 2012; Ardia et al., 2013; Dasgupta et al., 2013; Wetzel et al., 2013; Stanley et al., 2014). Compared with nitrogen, the solubility of carbon in silicate liquids at  $fO_2 < fO_2(IW)$  is less significant. According to Ardia et al. (2013), Dasgupta et al. (2013), Wetzel et al. (2013), and Stanley et al. (2014), the total carbon content of basic and ultrabasic melts saturated with respect to graphite at high temperatures and pressures is tens to hundreds of ppm. The experiments of Mysen et al. (2009, 2011) and Kadik et al. (2004, 2014) at similar conditions showed somewhat higher carbon contents in model silicate melts, 0.1–0.5 wt %. This could be related to the presence of H<sub>2</sub>O in the melt, which is

believed to increase CH<sub>4</sub> solubility in silicate liquids (Ardia et al., 2013; Kadik et al., 2014).

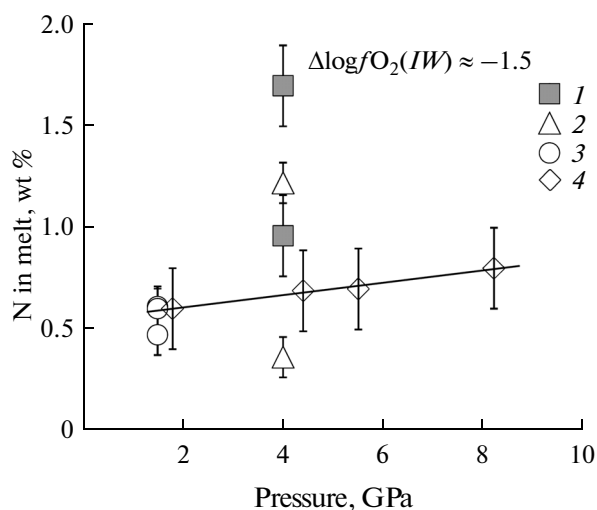
The goal of our experiments was to elucidate the character of simultaneous dissolution of nitrogen and carbon in silicate liquids at  $fO_2 < fO_2(IW)$ .

A comparison of experimental data on the solubility of nitrogen, carbon, and hydrogen in FeO–Na<sub>2</sub>O–SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melts at 4 GPa, 1550°C, and  $fO_2 < fO_2(IW)$  with nitrogen and hydrogen solubility in the same silicate liquids under the same  $T$ – $P$ – $fO_2$  conditions without carbon (Kadik et al., 2014) suggests that carbon affects nitrogen solubility in melts (Fig. 3). The total nitrogen content of melt is higher in the carbon-bearing systems. The influence of carbon on nitrogen solubility increases with decreasing  $fO_2$ . As will be shown below, the most probable reason for this is the formation of complexes in the melts with C–N bonds, the fraction of which increases with decreasing  $fO_2$ . A comparison of experimental data obtained at 4 GPa with the solubility of nitrogen and carbon in FeO–Na<sub>2</sub>O–SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melts at a lower pressure of 1.5 GPa and 1400°C indicates the influence of pressure on reactions in the melt. In the latter case, the melts show lower nitrogen contents at the same  $fO_2$  values (Fig. 3).

Experimental investigations of nitrogen interaction with melts of chondritic composition in the absence of hydrogen at pressures from 1 to 18 GPa and  $fO_2$  values 1.5–2.0 orders of magnitude below  $fO_2(IW)$  demonstrated that nitrogen solubility increases initially with increasing pressure but remains constant at pressures higher than ~4 GPa (Roskosz et al., 2013). At a constant  $fO_2$  value, the solubility of nitrogen in silicate melt is pressure-independent between 4 and 15 GPa (Fig. 4). It should be noted that, at  $fO_2$  values 1.5–2.0 orders of magnitude below  $fO_2(IW)$ , the solubility of nitrogen in chondritic melt is similar to that in FeO–Na<sub>2</sub>O–SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melts (Fig. 4). However, the nitrogen solubility in FeO–Na<sub>2</sub>O–SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melts is higher at lower  $fO_2$  values, which is probably related to significant changes in the speciation of N–C–H–O fluids in silicate liquids.

#### SOLUBILITY OF N AND C IN LIQUID IRON ALLOY

Experimental studies (Kadik et al., 2011, 2013) revealed that nitrogen solubility in liquid iron alloys at high pressures and low  $fO_2$  is 2–3 orders of magnitude higher than that in metals under ambient pressure. The solubility of nitrogen in liquid iron alloys in equilibrium with FeO–Na<sub>2</sub>O–SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melts is 2–4 wt % at 1.5 GPa, 1400°C or 4 GPa, 1550°C, and  $fO_2$  2–4 log units below  $fO_2(IW)$  and no higher than tens of ppm at atmospheric pressure (e.g., Bouchard and Bale, 1995). The investigations of Roskosz et al. (2013) over a wide range of pressure (1–18 GPa) and  $fO_2$  values 1.5–



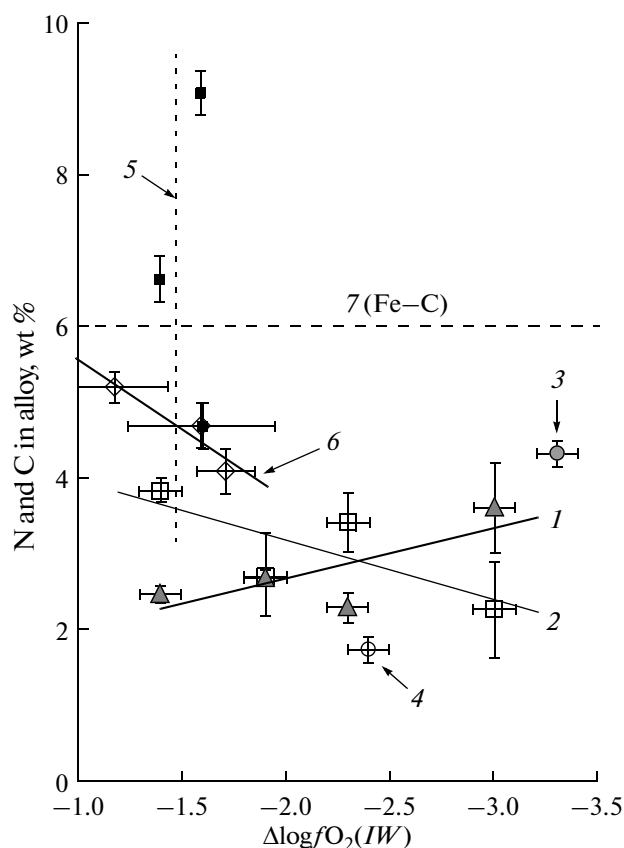
**Fig. 4.** Influence of pressure on the nitrogen content in FeO–Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> melts and model CI chondrite melt at  $f_{\text{O}_2}$  1.5 orders of magnitude below  $f_{\text{O}_2}(IW)$  at the dissolution of N<sub>2</sub> and N–C–H–O and N–H–O volatile compounds: (1) 4 GPa, 1550°C, melt + N–C–H–O volatile compounds (this study); (2) 4 GPa, 1550°C, melt + N–H–O volatile compounds (Kadik et al., 2013); (3) 1.5 GPa, 1400°C, melt + N–C–H–O volatile compounds (Kadik et al., 2011); and (4) model chondrite melt + N<sub>2</sub> at 1.8–8.2 GPa, 2500–2650 K (Roskosz et al., 2013).

2.0 orders of magnitude below  $f_{\text{O}_2}(IW)$  also indicated high nitrogen solubility in liquid iron alloys reaching 10–12% at 10–18 GPa. These authors observed a significant decrease in nitrogen solubility in FeNi alloy with increasing Ni content.

Similar to our previous results (Kadik et al., 2011, 2013), the present experiments indicate that nitrogen solubility in iron alloys at high pressures and  $f_{\text{O}_2} < f_{\text{O}_2}(IW)$  is high and  $f_{\text{O}_2}$ -dependent. The nitrogen content of liquid iron alloy increases with decreasing  $f_{\text{O}_2}$  (Fig. 5) from 2.5% at  $\Delta\log f_{\text{O}_2}(IW) = -1.4$  to 3.6% at  $\Delta\log f_{\text{O}_2}(IW) = -3.0$ . These nitrogen contents are somewhat higher than those documented in liquid iron alloys at 1.5 GPa (Kadik et al., 2011) and 4.0 GPa (Kadik et al., 2013) in equilibrium with FeO–Na<sub>2</sub>O–SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melts (Fig. 4).

In general, at the same pressures and  $f_{\text{O}_2}$  values two orders of magnitude below  $f_{\text{O}_2}(IW)$ , the nitrogen contents of liquid iron alloys observed in the present experiments and by Kadik et al. (2011, 2013) are in adequate agreement with nitrogen solubility in liquid iron alloys with low nickel contents (Roskosz et al., 2013) (Fig. 4).

The observed carbon contents in N–C–Fe alloys (2.3–3.8 wt %) are lower than the equilibrium carbon solubility in metal in the Fe–C system under the  $T$ – $P$  conditions of our experiments (~6 wt %) (e.g., Wood, 1993). A decrease in carbon solubility in liquid iron alloys was observed in the experiments of Takahashi



**Fig. 5.** Influence of  $f_{\text{O}_2}$  on the contents of nitrogen and carbon in liquid Fe alloy in equilibrium with FeO–Na<sub>2</sub>O–Al<sub>2</sub>O<sub>3</sub>–SiO<sub>2</sub> melt: (1) nitrogen in N–C–Fe alloy at 4 GPa and 1550°C (this study); (2) carbon in N–C–Fe alloy at 4 GPa and 1550°C (this study); (3) nitrogen in N–Fe alloy at 4 GPa and 1550°C (Kadik et al., 2013); (4) nitrogen in N–C–Fe alloy at 1.5 GPa and 1400°C (Kadik et al., 2011); (5) nitrogen in N–Ni–Fe alloy in equilibrium with chondrite melt at 1.8, 4.4, and 8.5 GPa and 2227–2327°C (Roskosz et al., 2013); (6) carbon in Fe alloy at low Si content, C–Mg–Fe–Si–O system at 4 GPa and 1600°C (Takahashi et al., 2013); and (7) (Fe–C) carbon solubility in Fe alloy in the Fe–C system at 4 GPa and 1600°C (Wood, 1993).

et al. (2013), Chi et al. (2014), and Kadik et al. (2014). It is related to the incorporation of Si in the structure of the metallic phase (e.g., Bouchard and Bale, 1995). Iron alloys in equilibrium with silicate liquids at  $f_{\text{O}_2}$  3–4 orders of magnitude below  $f_{\text{O}_2}(IW)$  are considerably enriched in Si and depleted in C (up to negligible contents) at decreasing  $f_{\text{O}_2}$  (Takahashi et al., 2013; Kadik et al., 2014). Since the Si content of N–C–Fe alloys was no higher than 0.1–0.2 wt % in our experiments, it can be suggested that the decrease of carbon content was caused by the influence of nitrogen. The investigation of metallurgical systems (e.g., Bouchard and Bale, 1995) showed that nitrogen dissolution in iron alloys is accompanied by their significant depletion in carbon. Another possible factor contributing to the depression of carbon solubility is the dissolution of oxygen in N–

**Table 2.** Chemical compositions of glasses and iron alloy globules, wt %

Run							
glass	$\Delta \log fO_2(IW)$	SiO <sub>2</sub>	Al <sub>2</sub> O <sub>3</sub>	FeO	Na <sub>2</sub> O	N	total
854	-1.4	62.42(0.14)	12.73(0.05)	13.45(0.19)	9.94(0.18)	0.96(0.12)	99.50
855	-1.9	67.23(0.26)	13.27(0.20)	7.97(0.21)	9.93(0.19)	1.69(0.21)	99.20
856	-2.3	70.11(0.47)	13.53(0.20)	4.22(0.61)	9.23(0.80)	2.57(0.31)	99.48
857	-3.0	74.52(0.24)	13.60(0.06)	2.17(0.07)	7.56(0.10)	4.19(0.12)	102.05
Iron alloy globules							
glass	$\Delta \log fO_2(IW)$	Fe	Si	C	N	W	total
854	-1.4	92.46(0.68)	0.05(0.01)	3.83(0.16)	2.52(0.07)	0.30(0.06)	99.16
855	-1.9	93.15(0.15)	0.12(0.05)	2.71(0.56)	2.7(0.08)	0.14(0.04)	98.82
856	-2.3	92.92(0.35)	0.05(0.02)	3.43(0.38)	2.19(0.19)	0.21(0.15)	98.80
857	-3.0	93.06(0.19)	0.038(0.01)	2.29(0.63)	3.63(0.16)	0.10(0.07)	99.12

C–Fe alloys, the fraction of which increases in the system with increasing degree of reduction of FeO melt (Fig. 2).

#### INFRARED AND RAMAN SPECTROSCOPY OF N–C–H–O-BEARING GLASSES

The IR and Raman spectroscopy of experimental products (quenched glasses) was used to evaluate the mechanisms of nitrogen, carbon, and hydrogen dissolution in FeO–Na<sub>2</sub>O–SiO<sub>2</sub>–Al<sub>2</sub>O<sub>3</sub> melts and the influence of oxygen fugacity, pressure, and temperature. The results of spectroscopy in the frequency range 1300–4500 cm<sup>-1</sup> indicate the presence of N–H–O and C–H–O species in melts (Figs. 6a, 7a).

##### IR Spectroscopy

The broad asymmetric absorption feature with peaks at 3608, 3523, and 3436 cm<sup>-1</sup> (Fig. 6a) is related to the stretching vibrations of hydroxyl groups (OH<sup>-</sup>) and H<sub>2</sub>O molecules (e.g., Stolper, 1982; Nowak and Behrens, 1995; Dianov et al., 2000; Mandeville et al., 2002). The sharp peaks at 1627–1612 cm<sup>-1</sup> correspond to the bending vibrations of H<sub>2</sub>O molecules. The weak peaks at 4543 and 4972 cm<sup>-1</sup> are attributed to the stretching vibrations of OH<sup>-</sup> and H<sub>2</sub>O in the former case and H<sub>2</sub>O in the latter case. The 4130 cm<sup>-1</sup> weak absorption band is assigned to molecular H<sub>2</sub> dissolved in glass (Shelby, 1994; Schmidt et al., 1998; Plotnichenko et al., 2005).

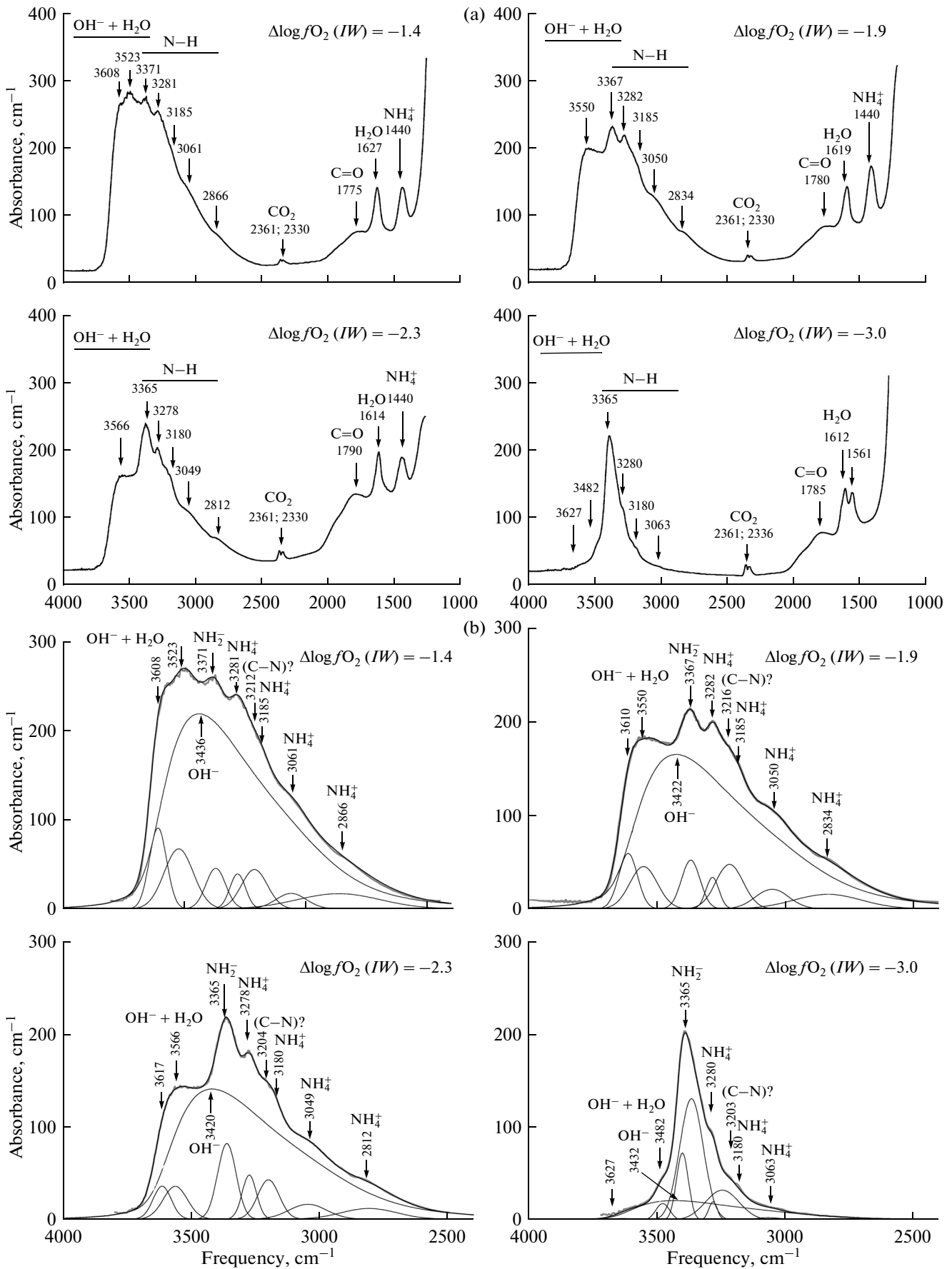
The double absorption feature with maxima at 2330 and 2361 cm<sup>-1</sup> (Fig. 6a) is interpreted as due to the stretching vibrations of C–O bonds in a CO<sub>2</sub> molecule

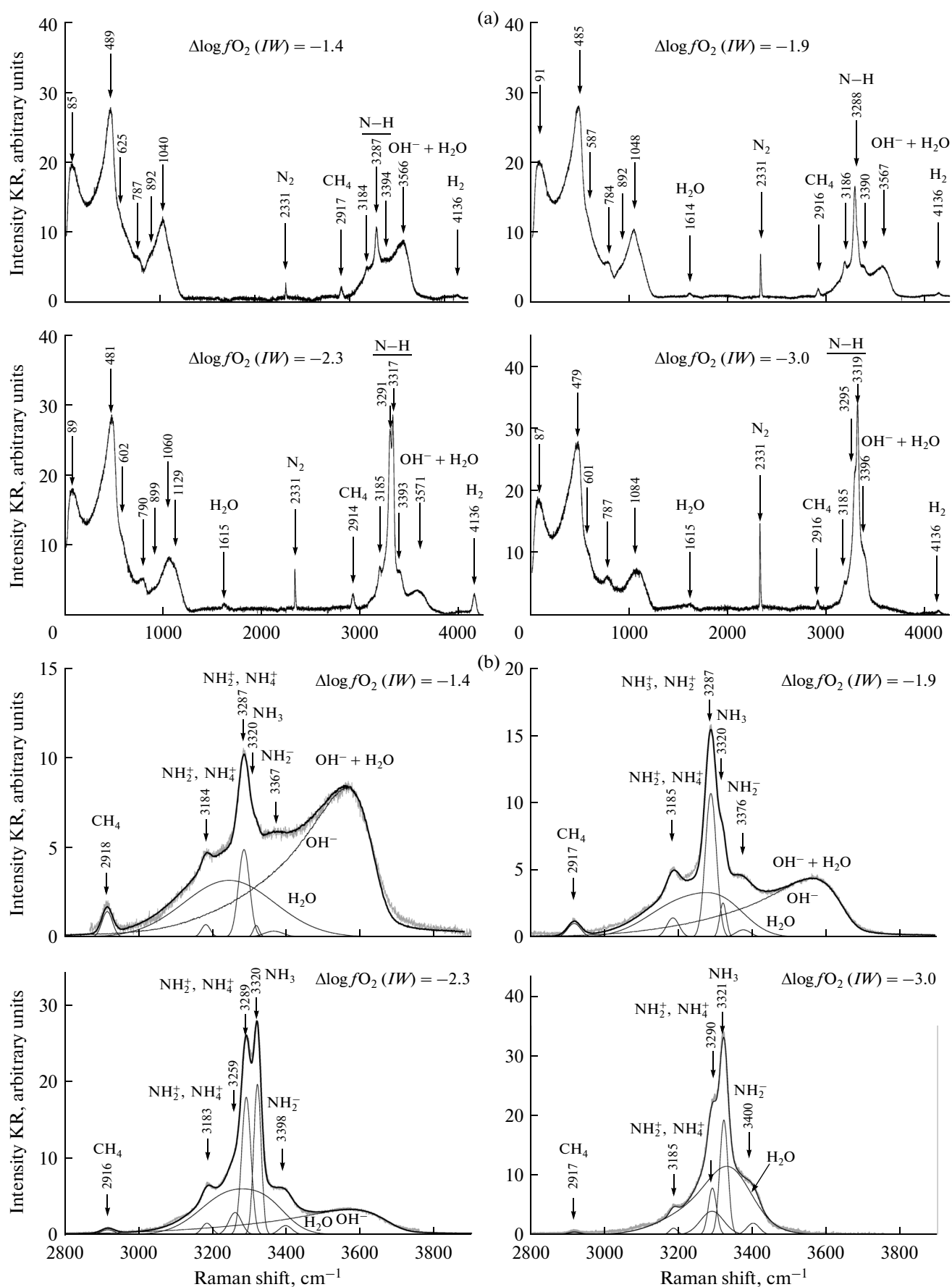
dissolved in glass (e.g., Fine and Stolper, 1985; Kohn et al., 1991; Morizet et al., 2010; Mysen, 2012). The broad absorption band at ~1775–1790 cm<sup>-1</sup> could associate with the stretching vibrations of the C=O double bond (Cataliotti and Jones, 1971). It should be noted that such peaks were observed in the IR spectra of C–H–O-bearing glasses obtained in experiments at  $fO_2 < fO_2(IW)$  (Kadik et al., 2011, 2013, 2014). Additional investigations are necessary to identify the species with C=O bonds in the glasses.

The absorption IR bands at 1440 cm<sup>-1</sup> and in the range 2800–3390 cm<sup>-1</sup> are due to vibrations of bonds in N–C–H–O molecules and complexes occurring in the glass structure (Fig. 6b). It is suggested that the intense absorption peaks at 1440 and 3280 cm<sup>-1</sup> and less intense peaks at 2866, 3061, and 3185 cm<sup>-1</sup> are related to ammonium groups, NH<sub>4</sub><sup>+</sup>, in melt. It should be noted that peaks at these frequencies were practically absent in the IR spectra of glasses produced at  $\Delta \log fO_2(IW) = -3.0$ , i.e., at the lowest  $fO_2$  conditions of our experiments (Fig. 6b). The assignment of absorption at these frequencies to NH<sub>4</sub><sup>+</sup> groups is based on a comparison of the IR spectra of N–C–H–O-bearing glasses with the IR spectra of NH<sub>4</sub><sup>+</sup>-bearing silicates, NH<sub>4</sub>AlSi<sub>3</sub>O<sub>8</sub> (hollandite), (NH<sub>4</sub>)<sub>2</sub>Si<sub>4</sub>O<sub>9</sub> (wadeite), NH<sub>4</sub>AlSi<sub>3</sub>O<sub>8</sub>H<sub>2</sub>O (cymrite) (Watenphul et al., 2009), and free NH<sub>3</sub> molecule (Herzberg, 1966) (Table 3). This inference is in agreement with the character of the IR spectra of other silicates containing NH<sub>4</sub><sup>+</sup> (micas) (Busigny et al., 2004; Harlov et al., 2001a, 2001b). The IR spectra of diopside synthesized at high pressures in the presence of nitrogen and hydrogen exhibit a strong absorption band at

**Fig. 6.** Infrared spectra of glasses containing N–C–O–H compounds in the ranges (a) 5000–1000 cm<sup>-1</sup> and (b) 2800–3800 cm<sup>-1</sup> and deconvolution of the IR spectra into separate bands. The recorded spectra were deconvoluted to Gaussian components taking into account the position and width at half-height of spectral components.







←  
**Fig. 7.** Raman spectra of glasses containing N–C–O–H compounds in the ranges (a) 5000–1000  $\text{cm}^{-1}$  and (b) 2800–3800  $\text{cm}^{-1}$  and deconvolution of the spectra into separate bands. The recorded spectra were deconvoluted to Gaussian components taking into account the position and width at half-height of spectral components. The spectra were normalized to the 490  $\text{cm}^{-1}$  band.

1414  $\text{cm}^{-1}$  (Watenphul et al., 2010). It is suggested that it corresponds to the  $\nu_4$  deformation vibrations of  $\text{NH}_4^+$  in the diopside structure.

The IR absorption peaks at 3371–3365  $\text{cm}^{-1}$  are believed to correspond to the stretching vibrations of N–H bonds in  $\equiv\text{Si}-\text{NH}_2$  complexes, which are formed owing to nitrogen and hydrogen interaction with silicate liquid (Mysen et al., 2008).

The IR spectra of the experimental samples exhibit a weak absorption band at 3203–3216  $\text{cm}^{-1}$  (Fig. 6b), which was not observed in our previous experiments. In this case, it can be assigned to a molecule or a complex containing both nitrogen and carbon, for instance,  $\equiv\text{C}-\text{NH}_2$  (Schrader, 1989). This problem requires further study.

### Raman Spectroscopy

The Raman spectra of glasses in the range 3300–3700  $\text{cm}^{-1}$  comprise a broad asymmetric band with maxima at 3566–3571  $\text{cm}^{-1}$ , which are similar to those observed in the Raman spectra of hydrous glasses (e.g., Mysen and Virgo, 1986), hydrogen-bearing glasses (Luth et al., 1987), and glasses containing carbon, hydrogen, and nitrogen (Kadik et al., 2004, 2013, 2014; Mysen et al., 2008, 2009). They correspond to vibrations of O–H bonds in  $\text{OH}^-$  groups and  $\text{H}_2\text{O}$  molecules in the silicate melt structure. The weak peak at

1615  $\text{cm}^{-1}$  is associated with bending vibrations of O–H bonds in  $\text{H}_2\text{O}$  molecules dissolved in melt. The minor band at 4136  $\text{cm}^{-1}$  corresponds to the vibration of H–H bonds of molecular  $\text{H}_2$  in glass (Luth et al., 1987; Plotnichenko et al., 2005).

The band centered at 2916–2918  $\text{cm}^{-1}$  is attributed to the methane molecule,  $\text{CH}_4$  (Socrates, 2001). A similar band was observed in the spectra of fluid inclusions in  $\text{CH}_4$ -bearing minerals (Pasteris et al., 1990; Seitz et al., 1987; Dubessy et al., 1999, 2001) and silicate melts containing carbon and hydrogen at low  $f\text{O}_2$  (Kadik et al., 2004, 2011; Mysen et al., 2009; Mysen and Yamashita, 2010; Mysen, 2012). The broadening and displacement of this band compared to the Raman spectra of gaseous methane are due to the interaction of  $\text{CH}_4$  molecules with the glass matrix.

According to Mysen et al. (2009, 2011) and Mysen and Yamashita (2010), the broadening of this band is related to the reaction of  $\text{CH}_4$  with oxygen of the silicate melt network with the formation of  $\text{CH}_3^-(\equiv\text{Si}-\text{O}-\text{CH}_3)$  complexes, the vibrations of which are observed at the same or a similar frequency as those in methane (Socrates, 2001). However, it should be noted that in the Raman spectra of organic compounds containing  $\text{CH}_3$  end-groups (for instance, ethane,  $\text{CH}_3-\text{CH}_3$ ), two bands are observed in this range.

**Table 3.** Frequencies of IR absorption bands of  $\text{NH}_4^+$ -bearing silicates assigned to different vibration modes ( $\nu$ ,  $\text{cm}^{-1}$ ) according to Watenphul et al. (2009) compared with the IR spectral characteristics of nitrogen-bearing glasses (this study)

$\nu$ ( $\text{cm}^{-1}$ )	$(\text{NH}_4)^+$	$\text{NH}_4\text{-Hld}$	$\text{NH}_4\text{-SWd}$	$\text{NH}_4\text{ Cym}$	NCHO-melt
$\nu_4$	1397	1402 sh (m)	1397 sh (m)	1404 sh (w)	1438 (s)
		1436 (s)	1422 (s)	1423 (s)	
		1459 sh (m)	1438 sh (m)	1447 sh (m)	
$\nu_2$	1685	1669 (w)	1671 (w)		2866 (m)
		2878 (m)	2842 (m)	2852 (w)	
$2\nu_4$		3043 (m)	2967 sh (m)	3091 sh (m)	3061 (m)
$\nu_2 + \nu_4$			3025 (m)		
$2\nu_2$		3140 (s)	3140 (s)	3179 sh (m)	3185 (m)
$\nu_3$	3134	3223 (m)	3320 (m)	3233 sh (m)	3285 (s)
		3281 sh (w)		3294 (s)	
		3333 sh (w)			

$(\text{NH}_4)^+$  denotes the values for free  $\text{NH}_4^+$  according to Herzberg (1966); values for hollandite (*Hld*,  $\text{NH}_4\text{AlSi}_3\text{O}_8$ ), Si-wadeite (*SWd*,  $(\text{NH}_4)_2\text{Si}_4\text{O}_9$ ), and cymrite (*Cym*,  $\text{NH}_4\text{AlSi}_3\text{O}_8\text{H}_2\text{O}$ ) are after Watenphul et al. (2009); NCHO-melt is the results of this study; intensities of bands: *sh*, band shoulder in the spectrum; *w*, weak; *m*, medium; and *s*, strong.

The strong narrow peak at  $2331\text{ cm}^{-1}$  corresponds to molecular nitrogen dissolved in glass. At normal temperature and pressure, the vibration of  $\text{N}_2$  molecules in gas is observed at  $2335\text{ cm}^{-1}$  (e.g., Lofithus and Krupnie, 1977). A similar peak of  $\text{N}_2$  in glass was documented by Roskosz et al. (2006), Mysen et al. (2008), Mysen and Fogel (2010), and Kadik et al. (2011, 2013), who studied the solubility of nitrogen and its compounds in silicate melts at high pressures.

The Raman spectra of glasses display intense peaks at  $3287\text{--}3295$  and  $3320\text{ cm}^{-1}$ , as well as peaks at  $3185$  and  $3387\text{--}3390\text{ cm}^{-1}$  (Fig. 7b) occurring in the frequency region corresponding to bond vibrations in N–H–O species (e.g., Dickinson et al., 1929; Yeo and Ford, 1994; Kowal, 2002). Their possible nature in N–H-bearing silicate melts under reduced conditions were investigated by Roskosz et al. (2006), Mysen et al. (2008, 2010), and Kadik et al. (2013). The peak at  $3320\text{ cm}^{-1}$  is similar to that observed in the Raman spectra of gaseous  $\text{NH}_3$  molecules (Dickinson et al., 1929; Wang et al., 1973; Durig et al., 1989). The Raman band at  $3287\text{--}3295\text{ cm}^{-1}$  corresponds to N–H vibrations of the amine group  $\text{NH}_2^-$  ( $\equiv\text{Si}\text{--}\text{NH}_2$ ) (e.g., Socrates, 2001), which is produced by the chemical interaction of nitrogen and hydrogen with silicate melts (Mysen et al., 2008). The band around  $3288\text{ cm}^{-1}$  probably associates with the N–H vibrations of a modified hydrogen bond in the  $\text{NH}_3$  molecule (Yeo and Ford, 1994), and the bands at  $3185\text{ cm}^{-1}$  are related to the vibrations of  $\text{NH}_2^+$  ( $\equiv\text{Si}\text{--}\text{O}\text{--}\text{NH}_2$ ) (Mysen et al., 2008). It is thought that the  $\equiv\text{Si}\text{--}\text{O}\text{--}\text{NH}_2$  complexes are also products of chemical interaction of nitrogen and hydrogen with silicate melt. However, the Raman peaks at  $3288$  and  $3185\text{ cm}^{-1}$  can be interpreted as resulting from the vibration of bonds in  $\text{NH}_4^+$ , which was suggested based on the IR spectra of the same glasses. We do not exclude the possibility of a different interpretation of bands at  $3185$ ,  $3290$ , and  $3390\text{ cm}^{-1}$  in the Raman spectra of glasses. They could associate with the  $2\nu_4$ ,  $\nu_1$ , and  $\nu_3$  vibrations in the  $\text{NH}_3$  molecule, which can be inferred from the comparison with the spectra of ammonium (Nakamoto, 1978).

#### COMPARISON WITH PREVIOUS WORK

The IR and Raman spectroscopy of the glasses showed that the dissolution of N–C–O–H volatiles in  $\text{FeO}\text{--}\text{Na}_2\text{O}\text{--}\text{SiO}_2\text{--}\text{Al}_2\text{O}_3$  melts in equilibrium with liquid iron alloy and graphite at 4 GPa,  $1550^\circ\text{C}$ , and  $f\text{O}_2$  values 2–4 orders of magnitude below  $f\text{O}_2(IW)$  is accompanied by the formation of  $\text{NH}_3$ ,  $\text{N}_2$ , and  $\text{CH}_4$  molecules and  $\equiv\text{Si}\text{--}\text{NH}_2$ ,  $\equiv\text{Si}\text{--}\text{O}\text{--}\text{NH}_2$ , and  $\text{NH}_4^+$  complexes. The melts contain very minor amounts of  $\text{CO}_3^{2-}$ . This is consistent with the estimates of the influence of  $f\text{O}_2$  on the stability of  $\text{CO}_3^{2-}$  in magmatic melts (Hirschmann and Withers, 2008; Stanley et al., 2014),

which indicated very low  $\text{CO}_3^{2-}$  contents in silicate liquids at  $f\text{O}_2$  values two orders of magnitude below  $f\text{O}_2(IW)$ . In addition to compounds with nitrogen and carbon, dissolved hydrogen exists in melt as  $\text{H}_2$  and  $\text{H}_2\text{O}$  molecules and hydroxyl groups,  $\text{OH}^-$ . The volatile N–C–O–H species in  $\text{FeO}\text{--}\text{Na}_2\text{O}\text{--}\text{SiO}_2\text{--}\text{Al}_2\text{O}_3$  melts in equilibrium with liquid iron alloy and graphite are in general identical to those observed during nitrogen and carbon dissolution in iron-free binary silicate melts of alkaline and calc-alkaline composition (Mysen et al., 2008; Mysen and Yamashita, 2010; Mysen and Fogel, 2010), Fe-bearing aluminosilicate melt (Kadik et al., 2010, 2013, 2014), and basic melts (Kadik et al., 2004; Dasgupta et al., 2013; Ardia et al., 2013; Wetzel et al., 2013). In contrast to the results for basaltic melts (Wetzel et al., 2013; Stanley et al., 2014), the  $\text{FeO}\text{--}\text{Na}_2\text{O}\text{--}\text{SiO}_2\text{--}\text{Al}_2\text{O}_3$  glasses do not show a band at  $2110\text{ cm}^{-1}$  in Raman spectra (Wetzel et al., 2013) and a band at  $2205\text{ cm}^{-1}$  in IR spectra (Stanley et al., 2014); these bands associate with the formation of complexes with C=O bonds in basaltic liquids. The vibrations at  $2110\text{ cm}^{-1}$  in the Raman spectra and at  $2205\text{ cm}^{-1}$  in IR spectra were assigned to the  $\text{Fe}(\text{CO})_5$  and  $\text{Fe}(\text{CO})_6^{2+}$ , respectively. It is conceivable that the absence of these species in the  $\text{FeO}\text{--}\text{Na}_2\text{O}\text{--}\text{SiO}_2\text{--}\text{Al}_2\text{O}_3$  melts was due to the lower  $f\text{O}_2$  values in our experiments compared with those of Wetzel et al. (2013) and Stanley et al. (2014). In the former case,  $f\text{O}_2$  was 2–4 orders of magnitude below  $f\text{O}_2(IW)$ , and, in the later case,  $f\text{O}_2$  values are close to or 1–2 orders of magnitude higher than  $f\text{O}_2(IW)$ . On the other hand, it is possible that complexes with C=O double bonds are formed in the  $\text{FeO}\text{--}\text{Na}_2\text{O}\text{--}\text{SiO}_2\text{--}\text{Al}_2\text{O}_3$  melts at  $f\text{O}_2 < f\text{O}_2(IW)$ , but their nature is probably different and poorly known. Their formation in melts is supported by IR absorption bands at  $1775\text{--}1785\text{ cm}^{-1}$  in the  $\text{FeO}\text{--}\text{Na}_2\text{O}\text{--}\text{SiO}_2\text{--}\text{Al}_2\text{O}_3$  glasses (Fig. 6a). The formation of complexes with C=O double bonds in silicate liquids at  $f\text{O}_2 < f\text{O}_2(IW)$  was established by Kadik et al. (2011, 2013).

The determination of nitrogen and hydrogen solubility in  $\text{FeO}\text{--}\text{Na}_2\text{O}\text{--}\text{SiO}_2\text{--}\text{Al}_2\text{O}_3$  melts both in the absence (Kadik et al., 2013) and in the presence of free carbon (graphite) in the system (this study) implies the presence of generally similar N–H–O molecules and complexes at 4 GPa,  $1550^\circ\text{C}$ , and  $f\text{O}_2 < f\text{O}_2(IW)$ . An exception is an absorption band in the IR spectra of the glasses at  $3203\text{--}3216\text{ cm}^{-1}$ , which appears in silicate liquids saturated in carbon and nitrogen.

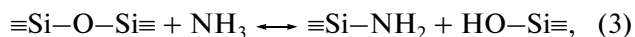
#### INFLUENCE OF $f\text{O}_2$ ON THE FORMATION OF N–C–O–H SPECIES IN MELT

Figures 8 and 9 show the normalized spectral characteristics of glasses,  $Abs/Abs_0$  and  $I/I_0$ , where  $Abs_0$  is the absorption coefficient for a band of the IR spectra and  $I_0$  is the coefficient of the integral intensity of

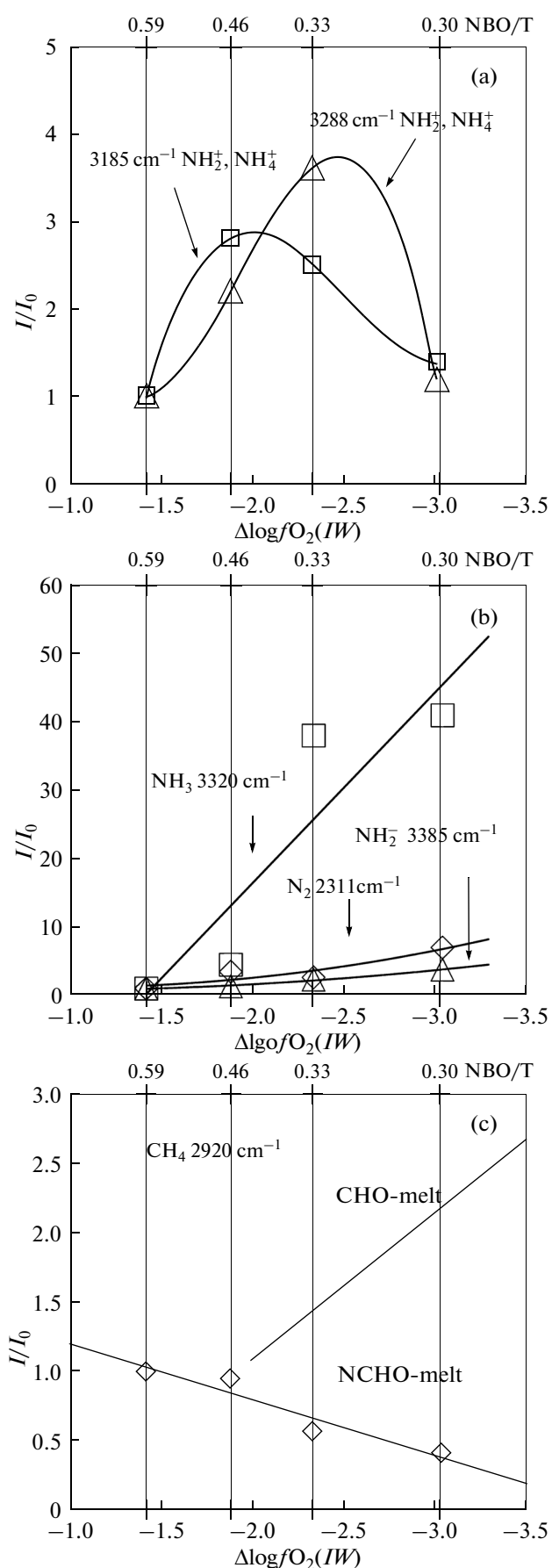
Raman bands at  $\Delta \log fO_2(IW) = -1.4$ . The degree of silicate melt polymerization is characterized in these diagrams by the NBO/T ratio (Mysen, 2012), where NBO is the number of nonbridging oxygens, and T is the number of tetrahedral cations. A decrease in NBO/T corresponds to an increase in the degree of silicate melt polymerization.

Variations in  $I/I_0$  and  $Abs/Abs_0$  of the Raman and IR spectra of the glasses indicate an increase in the content of the reduced nitrogen species  $\equiv Si-NH_2$  and  $\equiv Si-O-NH_2$  at decreasing  $fO_2$  (Figs. 8b, 9b). The proportion of the integral intensities of the 3320 and 3385  $cm^{-1}$  bands shows that the content of molecular  $NH_3$  increases sharply relative to that of  $\equiv Si-NH_2$ . At  $fO_2$  values 2–3 orders of magnitude below  $fO_2(IW)$ , the reduction of nitrogen is accompanied by a considerable decrease in the contents of  $\equiv Si-O-NH_2$  and  $NH_4^+$  species (Figs. 8a, 9a). Hence, it can be suggested that  $NH_3$  molecules become the dominant nitrogen species in melts at low  $fO_2$ . Despite the low  $fO_2$  values, molecular  $N_2$  is present in the melt, and its amount increases with decreasing  $fO_2$  and NBO/T (Fig. 8b). The intensity of the 2331  $cm^{-1}$  band in the Raman spectra indicates that the total content of molecular nitrogen in melts is not high.

According to the interpretation of Mysen and Fogel (2010) and Mysen (2013), the formation of  $\equiv Si-NH_2$  complexes is controlled by the chemical interaction of  $NH_3$  with the silicate melt matrix, which results in the appearance of  $Si-NH_2$  and  $Si-OH$  bonds through the reaction:

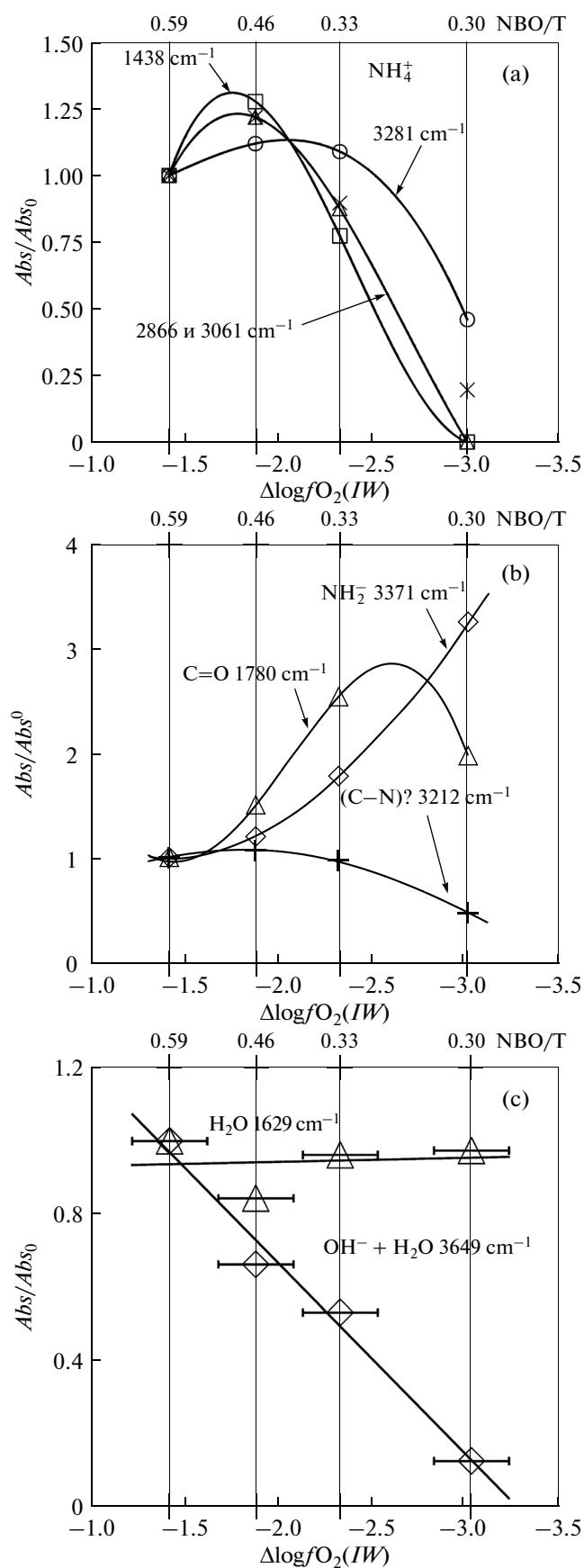


where  $\equiv Si-O-Si\equiv$  denotes a bridging oxygen in the glass matrix, and  $Si-NH_2$  and  $Si-OH$  are the formed end-groups. The interaction of  $NH_3$  with the silicate melt matrix is similar to  $H_2O$  dissolution with the for-



**Fig. 8.** Variations in the  $I/I_0$  ratio of the Raman spectra of glasses characterizing the influence of  $fO_2$  on the proportions of N–C–O–H species in  $FeO-Na_2O-Al_2O_3-SiO_2$  melts at 4 GPa and 1550°C.  $I_0$  is the integral intensity of Raman bands at  $\Delta \log fO_2(IW) = -1.4$ .

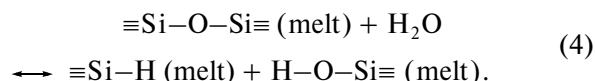
$I/I_0$  values are given for the Raman bands of (a)  $NH_2^+$ ,  $NH_4^+$  (3185  $cm^{-1}$ ) and  $NH_2^+$ ,  $NH_4^+$  (3288  $cm^{-1}$ ); (b)  $NH_3$  (3320  $cm^{-1}$ ),  $NH_2$  (3385  $cm^{-1}$ ), and  $N_2$  (2331  $cm^{-1}$ ); and (c)  $CH_4$  (2917  $cm^{-1}$ ) according to this study (NCHO-melt) and Kadik et al. (2014) (CHO-melt) and  $H_2$  (4136  $cm^{-1}$ ). The NBO/T values of melts are given accounting for the total content of  $H_2O$  in 14 glasses determined by IR spectroscopy: 0.59 (0.18), 0.46 (0.14), 0.33 (0.06), and 0.30 (0.04) corresponding to experimental  $fO_2$  values. Values in parentheses are NBO/T values ignoring the  $H_2O$  content of glasses.



**Fig. 9.** Variations in the  $Abs/Abs_0$  ratio of the IR spectra of glasses characterizing the influence of  $fO_2$  on the proportions of N–C–O–H species in  $FeO-Na_2O-Al_2O_3-SiO_2$  melts at 4 GPa and 1550°C.  $Abs_0$  is the absorption coefficients of IR bands at  $\Delta\log fO_2(IW) = -1.4$ .

$Abs/Abs_0$  values are given for the IR bands of (a)  $NH_4^+$  (1438, 2688, 30161, and 3281  $cm^{-1}$ ); (b)  $NH_2^-$  (3371  $cm^{-1}$ ), (C–N)? (3212  $cm^{-1}$ ), and C=O (1780  $cm^{-1}$ ); and (c)  $OH^-$  (3549  $cm^{-1}$ ) and  $H_2O$  (1629  $cm^{-1}$ ). See Fig. 8 for the determination of NBO/T values.

mation of hydroxyl groups,  $OH^-$  (e.g., Mysen and Virgo, 1986; Kohn et al., 1991):



The stronger changes in the integral intensity of the Raman band at 3320  $cm^{-1}$  ( $NH_3$ ) compared with the 3385  $cm^{-1}$  band ( $\equiv Si-NH_2$ ) indicate an increase in  $NH_3/\equiv Si-NH_2$  content with decreasing  $fO_2$  and NBO/T (Fig. 8b). However, according to reaction (3), a decrease in melt NBO/T, i.e., the formation of silicate melts with higher bridging oxygen contents, should have decreased the  $NH_3/NH_2^-$  ratio, which was not the case in our experiments. Therefore, it can be suggested that  $fO_2$  plays a dominant role in the formation of  $NH_3$  and  $\equiv Si-NH_2$  in the melt through redox reactions involving simultaneously nitrogen, carbon, hydrogen, and oxygen.

It is believed that, similar to noble gases, the  $N_2$ ,  $NH_3$ , and  $CH_4$  molecules reside in structural cavities of silicate liquids (Carroll and Webster, 1994; Mysen and Richet, 2005, ch. 16). This is indicated by the increasing solubility of  $N_2$ ,  $NH_3$ , and  $CH_4$  with increasing degree of silicate melt polymerization and, correspondingly, increasing number of sites suitable for these molecules in the silicate liquid matrix (Mysen et al., 2011; Dasgupta et al., 2013). Within this model, the increase in  $N_2$  and  $NH_3$  contents in  $FeO-Na_2O-SiO_2-Al_2O_3$  melts with decreasing  $fO_2$  can be attributed to an increase in the degree of silicate melt polymerization induced by FeO reduction. On the other hand, the  $CH_4$  content decreases with decreasing NBO/T. Similar to the case of the formation of N–H bonds, this may indicate that a change in the melt structure accompanying its polymerization is not the only factor of  $CH_4$  dissolution.

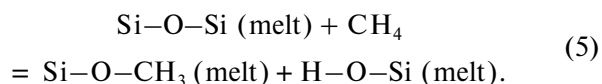
The  $Abs/Abs_0$  values that are assigned to complexes with C=O double bonds increase sharply with decreasing  $fO_2$  and reach maxima at around  $\Delta\log fO_2(IW) = -3$  (Fig. 8c). A further  $fO_2$  decrease is accompanied by a decrease in  $Abs/Abs_0$ . The existence of a maximum in the  $fO_2$  dependence of the content of complexes with C=O double bonds suggests that these complexes react

with other C–O–H species in melt at  $fO_2$  values in the system below  $\Delta\log fO_2(IW) = -1.4$ .

The observed  $Abs/Abs_0$  values indicate that a decrease in  $fO_2$  is accompanied by the occurrence of redox reactions resulting in a significant decrease in the content of  $OH^-$  groups in melt at minor variations in molecular  $H_2O$  content (Fig. 9c) and, correspondingly, a decrease in  $OH^-/H_2O$ . The  $fO_2$  dependence of  $I/I_0$  for  $H_2$  displays a sharp maximum and a significant decrease below  $\Delta\log fO_2 \approx -2.5$ . This tendency for  $H_2$  deserves further experimental investigations. Owing to the high diffusion mobility of hydrogen molecules in glasses, part of hydrogen could be lost from the glasses after the completion of melting experiments at high pressures and temperatures. However, there is evidence that the diffusive loss of hydrogen from glasses is not significant, if at all. This is supported by observations of hydrogen contents in glasses obtained by quenching carbon- and water-bearing melts (Hirschmann et al., 2012), which did not change during the prolonged storage of the glasses after the experiments.

## $H_2O$ CONTENT IN MELTS

Total water contents ( $OH^- + H_2O$ ) in glasses were determined on the basis of the Bouguer–Lambert–Beer law using  $Abs_{3548}$  values from the IR spectra in the region of ( $OH^- + H_2O$ ) vibrations at  $3548\text{ cm}^{-1}$  (e.g., Mercier et al., 2010). The molar extinction coefficient,  $\epsilon_{3548}$ , was estimated from its dependence on glass NBO/T calibrated by Mercier et al. (2010). The obtained ( $OH^- + H_2O$ ) contents recalculated to  $H_2O$  are 2.93, 1.98, 1.65, and 0.41 wt % at  $\Delta\log fO_2(IW)$  values of  $-1.4$ ,  $-1.8$ ,  $-2.3$ , and  $-3.0$ , respectively. The high  $H_2O$  contents result from the character of redox reactions in our experiments. They are related to FeO reduction in the melts via reaction (1) with release of considerable amounts of oxygen (Fig. 3), which contributes to the formation of oxidized carbon and hydrogen species at the given  $fH_2$  value. These include complexes with C=O double bonds,  $OH^-$  groups, and  $H_2O$  molecules. The formation of  $OH^-$  groups could be additionally related to the interaction of molecular  $NH_3$  with the silicate melt matrix via reaction (4), which was discussed by Mysen et al. (2013, 2014). Similarly,  $OH^-$  groups may be produced by the interaction of molecular  $CH_4$  with silicate melts (Kadik et al., 2014):



Thus, the experiments suggest that  $OH^-$  groups and molecular  $H_2O$  can be stable in silicate liquids at certain  $fO_2$  values below  $fO_2(IW)$ .

## NITROGEN DEGASSING DURING MELTING OF THE EARTH'S EARLY REDUCED MANTLE

The investigation of the nature of redox reactions of nitrogen, carbon, and hydrogen with FeO– $Na_2O$ – $Al_2O_3$ – $SiO_2$  melts and liquid iron alloys at 4 GPa,  $1550^\circ\text{C}$ , and  $fO_2$  values 1.5–3 orders of magnitude below  $fO_2(IW)$  has a number of important implications for the estimation of conditions under which volatile nitrogen compounds formed during the melting of the Earth's early reduced mantle.

According to estimates (Marty, 1995; Javoy, 1997; Cartigny et al., 2001), the Earth's mantle contains between 2 and 40 ppm nitrogen. The solubility of nitrogen in silicate melts at high pressure and  $fO_2 < fO_2(IW)$  is much higher. Therefore, it is reasonable to assume that most of terrestrial nitrogen was initially dissolved in the magma ocean (e.g., Li et al., 2013). Its further behavior during the evolution of the early reduced mantle had to be controlled by nitrogen partitioning between silicate melt, metallic phase, and products of magma ocean crystallization and degassing. Experimental studies (Mysen et al., 2008, 2013, 2014; Mysen and Fogel, 2010; Kadik et al., 2011, 2013; Roskosz et al., 2013; Li et al., 2013, 2015; Li and Keppler, 2014) have demonstrated the possible occurrence of such partitioning during the early geochemical history of the Earth.

If the nitrogen content of the material from which the Earth was formed was similar to the nitrogen abundance in chondrites,  $\sim 1100$  ppm (Keridge, 1985; Javoy, 1997; Marty, 2012), either more than 95% of primordial nitrogen had to be lost from the Earth to space during accretion or a significant amount of nitrogen should still be retained in the Earth's interiors (e.g., Marty, 2012). The relative abundance of nitrogen compared with other gases in the present-day atmosphere (Marty, 2012) suggests that the nitrogen loss to space, if occurred, could not be considerable. This motivated searches for silicate and metallic phases, either liquid or crystalline, in the mantle that could account for the burial of nitrogen in the planetary interiors.

This study confirmed the conclusions of Kadik et al. (2011, 2013) and Roskosz et al. (2013) on high nitrogen solubility in iron alloys at high pressures and  $fO_2 < fO_2(IW)$ . It increases with increasing pressure and hydrogen content and decreasing  $fO_2$ . The nitrogen content of iron alloy may be as high as 4–6 wt % at 1.5–5 GPa and up to  $\sim 10$  wt % at 9–12 GPa (Roskosz et al., 2013). Thus, nitrogen exhibits siderophile behavior under the conditions of magma ocean formation at  $fO_2 < fO_2(IW)$ . This implies that part of nitrogen could be dissolved in iron alloys segregating from the melting products of the early reduced mantle and buried subsequently in the Earth's metallic core.

The experimental results of Watenphul et al. (2010) and Li et al. (2013) allow us to suggest that nitrogen

could be dissolved in the crystallization products of the magma ocean and, thus, be retained in the minerals of the Earth's early mantle. Nitrogen solubility in forsterite, pyroxenes, and pyrope at 15–35 kbar is between 3 and 100 ppm; it increases with decreasing  $fO_2$  and is controlled by the formation of ammonium cation ( $NH_4^+$ ) in the crystals. It should be noted that, according to our experiments, a similar mechanism of nitrogen dissolution is characteristic of  $FeO-Na_2O-Al_2O_3-SiO_2$  melts; however, it appeared that the stability of  $NH_4^+$  is significantly dependent on  $fO_2$ . At 4 GPa and  $fO_2$  values 2.5–3.0 orders of magnitude lower than  $fO_2(IW)$ , nitrogen dissolution in aluminosilicate liquids does not result in the formation of  $NH_4^+$ . If this tendency will also prove to be valid for crystals, nitrogen burial in the products of magma ocean crystallization will be restricted to  $fO_2$  values 2.5–3.0 orders of magnitude below  $fO_2(IW)$ .

The nitrogen solubility in silicate melts and mantle crystalline phases decreases with increasing  $fO_2$  (Li et al., 2013). Therefore, it can be supposed that an increase in  $fO_2$  during the chemical differentiation of the Earth's early reduced mantle could cause short-term intense nitrogen release from deep magmas and crystalline phases. The investigations of Galimov (2005), Wood et al. (2006), Corgne et al. (2008), and Javoy et al. (2010) suggested that an increase in  $fO_2$  in the deep parts of the early mantle and magma ocean could be related to the self-oxidation of planetary material owing to the redistribution of Fe, Si, and O between silicate and metallic phases. According to Galimov (2005), reduced conditions were characteristic of the mantle during the initial stage of the formation of the Earth's core. The next evolutionary stage involved interaction between the silicate mantle and the metallic core:  $3Fe^{2+}O$  (mantle)  $\rightarrow$   $Fe^0$  (core) +  $Fe_2^{3+}O_3$  (mantle), which was accompanied by an increase in  $fO_2$  in the planetary interiors, probably up to the high  $fO_2$  values that are observed in the upper mantle. According to estimates, this stage could last 150–300 Myr. The model of heterogeneous accretion (Rubie et al., 2011) postulates an increase in  $fO_2$  with increasing mass of the growing planet within the  $\Delta\log fO_2(IW)$  range from  $-5$  to  $-2$ . The increase in  $fO_2$  in the deep parts of the Earth's early mantle could be caused by Si transfer owing to interaction between the silicate mantle and metallic core:  $SiO_2$  (silicate) +  $2Fe$  (metal) =  $Si$  (metal) +  $2FeO$  (silicate) (Frost et al., 2008; Javoy et al., 2010). The incorporation of Si in the core increased the abundance of FeO in the mantle, which became progressively oxidized as  $\Delta\log fO_2(IW)$  increased supposedly from  $-4.5$  to  $-1.5$ . It is assumed that  $fO_2$  increased in the magma ocean owing to FeO disproportionation during (Mg,Fe,Al)(Al,Si) $O_3$  perovskite crystallization via the reaction  $3Fe^{2+}O$  (melt) +  $Al_2O_3$  (melt) =  $2Fe^{3+}$

$AlO_3$  (perovskite) +  $Fe^0$  (metal) (Frost et al., 2008; Wood et al., 2006). The precipitation of a metallic phase was accompanied by the gradual oxidation of the magma ocean and a  $\Delta\log fO_2(IW)$  increase from  $-5$  to  $-2$ . According to the experiments reported here and recent data of Kadik et al. (2011, 2013) and Roskosz et al. (2013), this region of  $fO_2$  values corresponds to conditions under which the supposed initial nitrogen content in the Earth's early mantle could be completely dissolved in the equilibrium melt and metallic phases. The degassing of such melts with dissolved nitrogen would require a more significant increase in  $fO_2$  in the zones of early mantle melting up to the condition  $fO_2 > fO_2(IW)$ .

Such conditions were probably reached in the upper parts of the magma ocean owing to  $fO_2$  stratification with depth (Hirschmann et al., 2012). The distribution of  $fO_2$  with depth is controlled by the influence of pressure, temperature, and convective mixing of melts in the ocean on the  $Fe_3^+/Fe_2^+$  ratio and  $fO_2$  in silicate liquid and, correspondingly, on the behavior of volatile components. It is also possible that the stratification resulted from the self-oxidation of melts at their low-pressure degassing owing to the release of reduced N–C–H–O species from the melts. This process had to result in an increase in  $fO_2$  owing to changes in the proportions of oxidized and reduced compounds of nitrogen, carbon, hydrogen, and oxygen in the melt. This is supported by the experimental results of Holloway and Jakobsson (1986), who studied equilibria between silicate melts, graphite, and C–O–H volatile compounds at high temperatures and pressures and  $fO_2 = fO_2(IW)$ . They showed that the formation of fluids enriched in  $CH_4$ ,  $CO$ , and  $H_2$  results in a significant increase in the fraction of oxygen-rich C–O–H species in the melt, primarily,  $H_2O$ . The fractional degassing of reduced gaseous components will enhance this effect resulting in an additional increase in  $fO_2$  (Kadik and Lukanin, 1986).

Shallow local magmatic degassing at the contact with the early atmosphere could be sustained by the convective transport of dissolved volatile components from the deep levels of the magma ocean. It could also be enhanced by the directional crystallization of the magma ocean, which is suggested by the model of Elkins-Tanton (2008).

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