

# **Propagation of nonlinear dust‑acoustic solitary waves under the efect of non‑extensive electrons in inhomogeneous collisional magnetized dusty plasma**

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#### **Abstract**

In this paper, we have presented a study on the propagation of nonlinear Dust-Acoustic Solitary Waves (DASWs) in inhomogeneous magnetized dusty plasma. In this model, we have considered a system of collisional, magnetized dusty plasma, consisting of nonextensive electrons, Boltzmannean ions, and negatively charged dust grains in the plasma. We have also considered variable number density for the diferent dusty plasma components. Using the Reductive Perturbation Theory (RPT), the modifed Zakharov-Kuznetsov (m-ZK) equation is derived with the help of governing equations in the plasma. The solutions of m-ZK equation indicate the propagation of nonlinear DASWs. This study also shows how the inhomogeneity parameters and nonextensive electrons impact on the phase velocity, width and amplitude of the soliton propagating in inhomogeneous plasmas. In this investigation, we have also predicted some relations among the amplitude, width, and phase velocity of the DASWs which is relevant to Earth's magnetospheric plasma environment for the system.

**Keywords** Dusty plasma · RPT · ZK-equation · Non-extensive electrons · Inhomogeneous plasma

# **1 Introduction**

Solitons are those waves that maintain their structures by propagating at a particular speed. Solitons or solitary waves formed due to the cancellation of dispersive and nonlinear efects in the plasma. For the frst time, the Korteweg de-Vries (KdV) soliton has been studied to describe the ionacoustic solitons in an inhomogeneous plasma [[1](#page-6-0)]. The propagation of ion-acoustic solitary waves has also been studied in an inhomogeneous plasma system [[2\]](#page-6-1). Theoretically, Kuehl [[3](#page-6-2)] studied the ion-acoustic soliton refection in the inhomogeneous plasma. He observed variations in the amplitudes of the incident and refected solitons in the plasma. Later, considering fuid equations of plasma, Kuehl and Imen [\[4\]](#page-6-3) investigated the propagations of ion-acoustic soliton in inhomogeneous plasmas. Nejoh [[5](#page-6-4)] studied the

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waves in the presence of ions and κ-distributed electrons in the plasma.

Dusty plasma is one of the fundamental research topics in plasma physics. Dusty plasma supports a wide variety of interesting phenomena. In this system, plasma particles and dust grains together take part in the plasma system. Due to the addition of these dust grains in the plasma system, the plasma system becomes more complex in its behaviours. Therefore, they are termed complex systems. Goertz [[21\]](#page-6-18) and Northrop [[22](#page-6-19)] theoretically studied dusty plasmas for the inhomogeneous plasma system. Shukla and Silin [[23\]](#page-6-20) studied the Dust Ion-Acoustic Solitary Waves (DIAWs) for the isothermal plasma. They found that some novel low-frequency waves exist in the presence of charged dust particles in the plasma. They have also declared that there would have been some relation between those low-frequency waves and Saturn's F-ring. Many authors have studied Dust Acoustic Waves (DAWs) [[24\]](#page-6-21) and Dust Lattice Waves (DLW) [[25\]](#page-6-22) in diferent physical situations in the plasma. To understand the characteristics and properties of DLWs, DAWs, and DIAWs, researchers have tried several theoretical [[21](#page-6-18)[–23](#page-6-20)] and experimental studies [[26,](#page-6-23) [27](#page-6-24)]. Later, many studies have done on Dust-Acoustic Solitary Waves (DASWs) in inhomogeneous plasma under many physical situations. Dusty plasma has various connections with phenomena in astrophysical environments, space plasmas, fusion reactors, and as well as in laboratory experiments [\[28](#page-6-25), [29](#page-6-26)]. Gogoi and Deka [[30\]](#page-6-27) have investigated soliton propagation under the effect of dust charge and dust fuctuations in an inhomogeneous plasma. Theoretically, the DAWs have been studied in the presence of ions fuid, *K*-distributed superthermal electrons [\[31\]](#page-6-28), and negatively charged grains in the plasma. Akhtar et al. [[32\]](#page-6-29) studied the dust-cyclotron wave and modulational instability of DAWs with ions, electrons, and dust grains in magnetized plasma. Rehman et al. [[33\]](#page-6-30) investigated the linear and nonlinear magneto-acoustic wave propagation with pair-ion fullerene in an inhomogeneous plasma. Mushinzimana et al. [\[34\]](#page-6-31) studied the propagation of DIASWs and double layers under the efect of ions, Cairns-distributed electrons, and adiabatic positively charged dust grains in an inhomogeneous plasma. Dehingia and Deka [[35](#page-6-32)] have recently studied the impact of dust grains on DASWs propagating in an inhomogeneous dusty plasma under the efect of the magnetic feld. Dehingia and Deka [\[36](#page-6-33)] have studied the variations in the structures of DASWs propagating in an inhomogeneous unmagnetized plasma under the infuence of isothermal electrons. Thus, realizing the signifcance of studying DASWs in various astrophysical conditions, we have chosen our problem to investigate the propagation of nonlinear DASWs under the efect of non-extensive electrons in inhomogeneous collisional magnetized dusty plasma. In this problem, we have used RPT to derive the modifed Zakharov-Kuznetsov (m-ZK) equation. The solution of the m-ZK equation indicates the propagation of nonlinear DASWs in the plasma. In this study, we also discuss the variations of DA soliton amplitude and width depending on the variations of the inhomogeneity parameter and non-extensive electrons in the plasma. Results of this investigations concerning relations among the width, phase velocity, and amplitude of the DASWs are demonstrated in the earth's magnetospheric plasma environment in the model.

## **2 Governing equations**

We consider a three-dimensional, collisional, inhomogeneous warm dusty plasma with variation in number density for its components. The plasma model consists of Boltzmannean ions, non-extensive electrons, and massive dust grains with an external magnetic feld in the plasma. All the components of plasma and magnetic field *B*<sup>0</sup> = *B*<sup>0</sup> $\hat{x}$  are taken along *x*− direction. The dimensionless governing equations of continuity equation, momentum equation, Poisson's equation, Boltzmannean distribution of ions, nonextensive electron distribution, respectively, are given by

<span id="page-1-0"></span>
$$
\frac{\partial n_d}{\partial t} + \vec{\nabla} \cdot \left( n_d \vec{u}_d \right) = 0,\tag{1}
$$

<span id="page-1-1"></span>
$$
\frac{\partial \vec{u}_d}{\partial t} + \left(\vec{u}_d \cdot \vec{\nabla}\right) \vec{u}_d + \frac{5}{3} \sigma n_d^{-\frac{1}{3}} \vec{\nabla}\phi + \Omega(\vec{u}_d \times \hat{x}) + \Gamma \vec{u}_d = 0, \tag{2}
$$

<span id="page-1-2"></span>
$$
\vec{\nabla}^2 \phi = n_d + \frac{(n_e - \delta n_i)}{(\delta - 1)},
$$
\n(3)

<span id="page-1-3"></span>
$$
n_i = n_{i0}(x)e^{-s\phi},\tag{4}
$$

<span id="page-1-4"></span>
$$
n_e = n_{e0}(x)[1 + (q-1)\phi]^{\frac{(q+1)}{2(q-1)}}.
$$
\n(5)

Here, for equilibrium condition, the charge neutrality equation is given by  $\delta n_{i0}(x) = n_{e0}(x) + (\delta - 1)n_{d0}(x)$  where  $\delta = \frac{n_{i0}(x)}{n_{e0}(x)}$ . Also, we consider  $n_{e0}(x)$  and  $n_{i0}(x)$  are as unperturbed electron and ion number densities in the plasma. In our above equations,  $\Gamma$  is the collision frequency for neutral dust grains,  $\Omega$  is the dust cyclotron frequency,  $\phi$  is the electrostatic potential,  $n_d$  is the number density of dust grains, and  $\vec{u}_d$  is the dust fluid velocity in the plasma. We have also considered the normalized dimensionless parameters are  $\Omega = \frac{e_{0} Z_d}{m_d c}$ ,  $\sigma = \frac{T_d}{Z_d T_e}$ , and  $s = \frac{T_e}{T_i}$  where *c* is velocity of light,  $m_d$  is the mass of dust grains,  $Z_d$  is dust charge numbers,  $T_d$ is dust temperature,  $T_i$  is ion temperature, and  $T_e$  is electron temperatures in the plasma. Here,  $\alpha_e$  and  $\alpha_i$  are defined as the density gradient scale length for the electrons and ions, respectively, which can take either positive or negative

values for growing number densities. In our problem, we have used some physical parameters  $k$ ,  $\Omega$ ,  $\Gamma$ , and  $q$  to investigate the characteristics of DASWs depending on the typical nonthermal values [\[37](#page-6-34), [40\]](#page-6-35) of the dusty plasma in the system. In the governing Eqs.  $(1, 2, 3, 4, 5)$  $(1, 2, 3, 4, 5)$  $(1, 2, 3, 4, 5)$  $(1, 2, 3, 4, 5)$  $(1, 2, 3, 4, 5)$  $(1, 2, 3, 4, 5)$  $(1, 2, 3, 4, 5)$  $(1, 2, 3, 4, 5)$  $(1, 2, 3, 4, 5)$  we have considered the physical quantities are as follows:

$$
\nabla \lambda_{Dd} = \nabla \left( \sqrt{\frac{T_e}{4\pi e^2 Z_d}} \right), \Gamma = \left( \frac{1}{\omega_{pd}} \right) \lambda_{Dd},
$$

where  $\omega_{\rm nd}$  and  $\lambda_{\rm Dd}$  are the plasma frequency for dust grains and modifed Debye length, respectively. In this problem, for  $q \rightarrow 1$ , the electrons are distributed in Boltzmannean distribution, on the other hand for the real number  $q > -1$ , the electrons are considered to be in the form of unnormalized non-extensive electron distribution [\[39](#page-6-36)].

## **3 Derivation of m‑ZK equation**

To study the nonlinear propagation and characteristics of DAWs in inhomogeneous Magnetized Dusty Plasma (MDP) system, we use the RPT to derive the m-ZK equation. In order to use RPT [[16](#page-6-13), [36,](#page-6-33) [37\]](#page-6-34), we employ the standard stretched coordinates are as follows [[37\]](#page-6-34):

$$
\xi = \varepsilon^{\frac{1}{2}} \left( \int_{0}^{x} \frac{\partial x^{\prime}}{\lambda(x^{\prime})} - t \right), X = \varepsilon^{\frac{3}{2}} x, Y = \varepsilon^{\frac{1}{2}} y, Z = \varepsilon^{\frac{1}{2}} z, \Gamma = \varepsilon^{\frac{3}{2}} f_{1}
$$
\n(6)

Here, *M* is the phase velocity of DASWs and  $\epsilon$  is a smallness parameter which measures the size of the perturbation amplitude. To use RPT, we use some dependent variables are as follows:

$$
\begin{bmatrix} n_i \\ n_e \\ n_d \\ u_{dx} \\ u_{dy} \\ u_{dz} \\ \phi \end{bmatrix} = \begin{bmatrix} n_{i0}(X) \\ n_{e0}(X) \\ n_{d0}(X) \\ u_{d0}(X) \\ 0 \\ 0 \\ \phi_0(X) \end{bmatrix} + \varepsilon \begin{bmatrix} n_{i1} \\ n_{e1} \\ n_{d1} \\ u_{dx1} \\ 0 \\ 0 \\ 0 \\ 0 \\ \phi_1 \end{bmatrix} + \varepsilon^2 \begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \\ u_{dy1} \\ u_{dz1} \\ 0 \\ 0 \end{bmatrix} + \varepsilon^2 \begin{bmatrix} n_{i2} \\ n_{e2} \\ n_{d2} \\ u_{dz2} \\ u_{dz2} \\ u_{dz1} \\ u_{dz2} \\ \phi_2 \end{bmatrix} + \cdots
$$
\n(7)

Using RPT, we obtain the spatial gradient relations from the above Eqs. [1](#page-1-0), [2](#page-1-1), [3](#page-1-2), [4](#page-1-3), [5](#page-1-4), [6](#page-2-0), [7](#page-2-1) is as follows:

$$
\frac{\partial}{\partial \xi} \left( n_{i,e,d0} \right) = \frac{\partial \lambda}{\partial \xi} = \frac{\partial}{\partial \xi} \left( u_{dx0} \right) = \frac{\partial}{\partial \xi} \left( \phi_0(X) \right) = 0. \tag{8}
$$

At an equilibrium condition, we obtain

$$
5\sigma \frac{\partial}{\partial x} \left( n_{\partial}^{\frac{2}{3}} \right) + \frac{\partial}{\partial x} \left( n_{d0} u_{dx0} \right) = 0
$$
  
5\sigma \frac{\partial}{\partial x} \left( n\_{\partial}^{\frac{2}{3}} \right) + \frac{\partial}{\partial x} \left( u\_{dx0}^2 \right) - 2 \frac{\partial \phi\_0}{\partial x} + 2u\_{dx0} f\_1 = 0 (9)

Substituting the Eqs.  $(6)$  $(6)$  and  $(7)$  $(7)$  in the governing Eqs. [1,](#page-1-0) [2](#page-1-1), [3,](#page-1-2) [4](#page-1-3), [5,](#page-1-4) and comparing the coefficients of  $\epsilon$  for the smallest order, we have

$$
n_{d1} = -n_{d0} R \phi_1, \ u_{dx1} = -\lambda^{\prime} R \phi_1, \ u_{dz1} = \frac{\lambda^{\prime 2} R \partial \phi_1}{\Omega \partial Z}, \ u_{dz1} = \frac{\lambda^{\prime 2} R \partial \phi_1}{\Omega \partial Z} \}
$$
(10)

where 
$$
R = \frac{1}{(\lambda'^2 - \alpha n_{d0})}
$$
,  $\alpha = \frac{1}{\sqrt[3]{n_{d0}}} \frac{5}{3} \sigma$  and  $\lambda' = \lambda - u_{dx0}$ . The

value of phase velocity  $\lambda$  is found to have in the form as

$$
\lambda = u_{dx0} + \left( \alpha n_{d0} + \frac{n_{d0}(\delta - 1)}{\left(\frac{1+q}{2}\right)n_{e0}\left[1 - \left(\frac{q-3}{2}\right)\phi_0\right] + s\delta n_{i0}\left(1 - s\phi_0\right)} \right)^{\frac{1}{2}}.
$$
\n(11)

<span id="page-2-2"></span>Eq.  $(11)$  represents the phase velocity  $(\lambda)$  of the DAWs in the inhomogeneous plasma system.

Similarly, comparing the coefficients of  $\epsilon$  for the higher order perturbations, we have

<span id="page-2-6"></span>
$$
- \lambda \frac{\partial n_{d0}}{\partial \xi} + n_{d0} \frac{\partial n_{dx2}}{\partial \xi} + \frac{\partial}{\partial \xi} (n_{d1} u_{dx1})
$$
  
+ 
$$
\lambda \left[ \frac{\partial}{\partial X} (n_{d0} u_{dx1} + n_{d1} u_{dx0}) \right]
$$
  
+ 
$$
\lambda \left[ \frac{\partial}{\partial Y} (n_{d0} u_{dy2}) + \frac{\partial}{\partial Z} (n_{d0} u_{dz2}) \right] = 0
$$
 (12)

<span id="page-2-0"></span>
$$
- \lambda \frac{\partial u_{dx2}}{\partial \xi} - u_{dx1} \frac{\partial u_{dx1}}{\partial \xi} + \lambda \frac{\partial}{\partial X} (u_{dx0} u_{dx1}) + \alpha \frac{\partial n_{d2}}{\partial \xi} + \alpha \lambda \frac{\partial n_{d1}}{\partial X} - 1/3 \alpha \lambda (n_{d0})^{-1} n_{d1} \frac{\partial n_{d0}}{\partial X} - \frac{1}{3} \lambda \alpha (n_{d0})^{-1} n_{d1} \frac{\partial n_{d1}}{\partial \xi} - \lambda \frac{\partial \phi_1}{\partial X} - \lambda f_1 u_{dx1} = 0,
$$
\n(13)

<span id="page-2-3"></span><span id="page-2-1"></span>
$$
\frac{1}{\lambda^2} \frac{\partial^2 \phi_1}{\partial \xi^2} + \frac{\partial^2 \phi_1}{\partial Y^2} + \frac{\partial^2 \phi_1}{\partial Z^2} - n_{d2} - \frac{1}{(\delta - 1)} \left( n_{e2} - \delta n_{i2} \right) = 0,
$$
\n(14)

Where

$$
u_{d y 2} = -\frac{\lambda^{3} R}{\Omega^{2} \lambda} \left( \frac{\partial^{2} \phi_{1}}{\partial \xi \partial Z} \right), u_{d z 2} = -\frac{\lambda^{3} R}{\Omega^{2} \lambda} \left( \frac{\partial^{2} \phi_{1}}{\partial \xi \partial Y} \right).
$$
 (15)

Equation ([13\)](#page-2-3) shows the dependency of the phase velocity of soliton w.r.t. the variation in  $\phi_0$  and *x*/*L* where *L* refers as the density scale length taken to be an arbitrary value of 200. We have shown the graph for these variation in Figs. [1,](#page-3-0) [2](#page-3-1) and [3](#page-3-2), respectively. It also shows the decrease in the values of  $\lambda$  with the increasing values of *q* or *x*/*L* On

<span id="page-2-7"></span><span id="page-2-5"></span>2 Springer KCS 한국물리학회

<span id="page-2-4"></span> $\overline{a}$ 



<span id="page-3-0"></span>**Fig. 1** Plotting of  $\lambda$  vs  $x/L$  for  $\sigma = 10^{-2}$ ,  $s = 3$ ,  $\alpha_i = 0.9$ ,  $\alpha_e = 0.5$ ,  $u_{dx0} = 0.02$ 



<span id="page-3-1"></span>**Fig. 2** Plotting of  $\lambda$  vs  $x/L$  for  $\delta = 2$ ,  $s = 2$ ,  $x/L = 2$ ,  $\alpha_e = 0.5$ ,  $\alpha_i = 0.9, \sigma = 10^{-2}, u_{dx0} = 0.02$ 

the other hand, the phase velocity  $\lambda$  increases rapidly with the increase in the values of $\phi_0$ . However, the decrease in the values of  $\phi_0$  (negative values) and decreases the phase velocity of the DASW. Proceeding in this way, we observe an identical situation in the case of non-thermal ions present in the inhomogeneous dusty plasma system.

Using the Eqs. in Eqs.  $(10)$  $(10)$ ,  $(11)$  $(11)$  $(11)$  and  $(15)$  $(15)$  in Eqs.  $(12)$  $(12)$  $(12)$ , [\(13\)](#page-2-3), ([14](#page-2-7)) we eliminate second-order perturbed quantities and obtain the modifed ZK (m-ZK) equation is in the form as follows:

$$
\frac{\partial \phi_1}{\partial X} + A\phi_1 \frac{\partial \phi_1}{\partial \xi} + B\phi_1 \left( \frac{\partial^2 \phi_1}{\partial Y^2} + \frac{\partial^2 \phi_1}{\partial Z^2} \right) + C \frac{\partial^3 \phi_1}{\partial \xi^3} + D\phi_1 = 0,
$$
\n(16)

where,





<span id="page-3-2"></span>**Fig. 3** Plotting of *A* vs  $x/L$  for  $\sigma = 10^{-2}$ ,  $s = 3$ ,  $\alpha_i = 0.9$ ,  $\alpha_e = 0.5$ ,  $u_{dx0} = 0.02, \delta = 2, \Gamma = 0.001, \phi_0 = 0.2, \Omega = 0.05$ 

$$
A = E \left[ \frac{\left\{ 3R^2 n_{d0} \left( \lambda'^2 - \frac{1}{9} \alpha n_{d0} \right) \right\}}{-\frac{1}{R(\delta - 1)} \left\{ \frac{1}{4} n_{e0} (q - 3)(q - 1) + \delta s^2 n_{i0} \right\}} \right],
$$
  
\n
$$
B = -E \left( \frac{\lambda'^4 R n_{d0}}{\partial X} + \frac{1}{R} \right), C = -\frac{E}{\lambda^2 R},
$$
  
\n
$$
D = -E \lambda \left[ \lambda' \frac{\partial}{\partial X} (\lambda R n_{d0}) + \lambda' \frac{\partial}{\partial X} (R n_{d0} u_{dx0}) + n_{d0} \frac{\partial}{\partial X} (\lambda' u_{dx0} R) + \alpha n_{d0} \frac{\partial}{\partial X} (R n_{d0}) - \frac{1}{3} \alpha n_{d0} R \frac{\partial n_{d0}}{\partial X} + \lambda' n_{d0} R f_1 \right], E = \frac{1}{2\lambda^2 \lambda' R n_{d0}}.
$$

The above Eq.  $(16)$  is known as the modified ZK (m-ZK) Eq. [[37](#page-6-34)]. Equation [16](#page-3-3) indicates the propagation of nonlinear DASWs in the presence of non-extensive electron and inhomogeneity number density in an inhomogeneous dusty plasma. In the above m-ZK equation, *A* and *C* are the nonlinear and dispersive terms. However, the coefficient  $E$  occurs due to the inhomogeneity number density and the collision effect in the inhomogeneous dusty plasma system. In Fig. [4,](#page-4-0) we have seen that there is a variation in A, which is a function of *x*∕*L*. Figure [4](#page-4-0) also describes the dressed solitons for *A* < 0(0 <  $x/L$  < 1.2) and *A* > 0.

## **4 Solution of m‑ZK equation and result discussion**

To obtain the soliton solution of the derived m-ZK equation, we employ a standard transformation equation is as follows [[38](#page-6-37)]:

$$
\phi_1 = H(X, Y, Z, \xi)G(X),\tag{17}
$$

<span id="page-3-4"></span><span id="page-3-3"></span>where the amplitude factor  $G(X)$  is obtained as follows:



<span id="page-4-0"></span>**Fig. 4** Plotting of *G*(*X*) vs *x*/*L* for  $\delta = 2$ ,  $s = 4$ ,  $\alpha_e = 0.5$ ,  $\alpha_i = 0.9, \sigma = 10^{-2}, u_{dx0} = 0, \Gamma = 0.1, \Omega = 0.05, \phi_0 = -1$ 

$$
G(X) = \frac{1}{\left(R\lambda\sqrt{\lambda^{-}n_{d0}}\right)}e^{-\int_{0}^{x}\frac{f_1}{2\lambda}dx}
$$
\n(18)

Substituting Eq.  $(17)$  $(17)$  in  $(16)$ , we obtain

$$
\frac{\partial H}{\partial \tau} + A'H \frac{\partial H}{\partial \eta_1} + \frac{\partial}{\partial \eta_1} \left( \frac{\partial^2 H}{\partial \eta_2^2} + \frac{\partial^2 H}{\partial \eta_3^2} \right) + \frac{\partial^3 H}{\partial \eta_1^3} = 0, \quad (19)
$$

where  $A^{\prime}$ ,  $\tau$ ,  $\eta_1$ ,  $\eta_2$ , and  $\eta_3$  have their usual meaning which are defned as,

$$
\tau = \int_{-\infty}^{X} \sqrt{\frac{A}{C}} dX', A' = A(X)G(X),
$$

$$
\eta_1 = \sqrt{\frac{A}{C}} \xi, \eta_2 = \sqrt{\frac{A}{B}} Y, \text{ and } \eta_3 = \sqrt{\frac{A}{B}} Z
$$

Now, we define a new variable  $\eta = L_x \eta_1 + L_y \eta_2 + L_z \eta_3 - V\tau$ where *V* is the constant velocity. Here,  $L_x$ ,  $L_y$ ,  $L_z$  are the direction cosines of the wave vectors along the *𝜉*− axes, *Y*− axes, and *Z*− axes, respectively. Using the boundary conditions; for  $|\eta| \to \infty$ , we put  $H(\eta) \to 0$ ,  $\frac{\partial H}{\partial \eta_1} \to 0$ ,  $\frac{\partial^2 H}{\partial \eta_2^2} \to 0$ . Then, we obtain,

$$
H(\eta) = H'_m sech^2\left(\frac{\eta}{W}\right)
$$
\n(20)

Here,  $H'_m = \frac{6V}{L_x}$  is amplitude and  $W = 2\sqrt{\frac{L_x}{V}}$  $\frac{v_x}{V}$  is width of the DASWs in the plasma. Now, when we choose,  $\phi = \phi_1$ , and after performing some calculations, we obtain,

$$
\phi = A_m sech^2\left(\frac{\zeta}{\omega}\right),\tag{21}
$$



 $u_{dx0} = 0, \Gamma = 0.1, \Omega = 0.05, \phi_0 = -1$  **Fig.5** Plotting of  $\Delta$  vs  $x/L$  for  $s = 4, \sigma = 10^{-3}, \alpha_i = 0.6, \alpha_e = 0.5,$  $u_{dx0} = 0.2, \Gamma = 0.001, \delta = 0, \Omega = 0.5, \phi_0 = 0.2$ 

<span id="page-4-1"></span>where  $A_m$  and  $\Delta$  are the soliton amplitude and width of the DASWs in the plasma and are given by  $A_m = \left(\frac{6V}{L}\right)^n$ *Lx*  $\partial/\phi_0 \phi'$  and  $\Delta = 2 \sqrt{\frac{c}{\frac{V}{V} \Delta}}$  $\frac{C}{VL_xA\phi_{0\phi}}$ . Here, at *x* = 0, the amplitude factor of electrostatic potential is  $\phi_0$  and we consider  $\phi' = \frac{\phi'(x)}{\phi_0}$ . Also, the Mach number *M* for DASW is obtained as follows:

$$
M = \sqrt{\left(\frac{L_x \lambda}{L_x + V\lambda A \phi \phi'}\right)^2 + \frac{B}{C} \left\{\frac{L_x^2 (1 - L_x^2)}{L_y^2 + L_z^2}\right\}}.
$$
 (22)

In the above investigations, Fig. [4](#page-4-0) indicates the variations in  $G(X)$  versus  $x/L$  shown in Eq. (20). We have observed that  $G(X)$  attains its maximum value at the critical point  $(x/L) \approx 2.4$ . As the value of the non-extensive parameter (*q*) increases, decreases the value of critical point $(x/L)$ . Due to the dominance of the inhomogeneity parameter on the collision of the plasma particles, increases the value of *G*(*X*) w.r.t.*x*/*L*. Also, with the increasing value of  $x/L = (x/L)_{c}$ , a rapid decrease in the value of *G*(*X*) is seen accordingly. Similarly, when  $(x/L)$ <sub>*c*</sub>  $\lt x/L$ , the value of *G*(*X*) decreases due to the contraction of plasma inhomogeneities relative to the collisional efect in the plasma. Consequently, for the higher values of *x*∕*L*, *G*(*X*) obtains relatively smaller values in the above considered plasma system.

On the contrary, Fig. [5](#page-4-1), indicates the variations in the characteristics of DA solitons with *x*∕*L* and*q*. In Fig. [6,](#page-5-0) we have seen that there is a decrease in Mach number*M*, with the increase in  $\Omega$  or *x*/*L*. Consequently, the variations of three-dimensional electrostatics potential  $\phi$  and *x*/*L* are shown in Figs. [7](#page-5-1) and [8.](#page-5-2) From all the above observations, we have seen the width and amplitude of DASWs increase with increase in the value of *x*∕*L*. On the other hand, with the increase in the non-extensive effect  $(q)$  slightly



<span id="page-5-0"></span>**Fig.** 6 Plotting of *M* vs  $x/L$  for  $\sigma = 0, s = 3, \alpha_i = 0.9$ ,.  $\alpha_e = 0.5, u_{dx0} = 0, \phi_0 = -1, \Gamma = 0.01, q = 1, \delta = 4$ 



<span id="page-5-1"></span>**Fig.** 7 Plotting of  $\phi$  vs  $x/L$  and  $\xi$  for  $\sigma = 0, \alpha_i = 0.9, q = 2$ ,  $\alpha_e = 0.5, s = 3, u_{dx0} = 0, \phi_0 = -1, \Gamma = 0.01, q = 1, \delta = 4$ 

decrease the soliton amplitude but gets wider the width of the soliton in the plasma. Figures [7](#page-5-1) and [8](#page-5-2) also imply the movement of soliton towards  $\xi$ – direction with the increasing value of *x*∕*L* in the above considered dusty plasma model.

## **5 Conclusion**

In this paper, we have presented a study on the nonlinear characteristics of the DASWs propagating in an inhomogeneous collisional magnetized plasma. The plasma system consists of non-extensive electrons, positively and Boltzmannean distributed ions, and charged dust grains. Due to the efect of non-extensive electrons, collisional and



<span id="page-5-2"></span>**Fig. 8** Plotting of  $\phi$  vs  $\xi$  and *q* for  $\sigma = 0, \alpha_i = 0.9$ .  $x/L = 2, \alpha_e = 0.5, s = 3, u_{dr0} = 0, \phi_0 = -1, \Gamma = 0.01, q = 1, \delta = 4$ 

magnetic feld in the plasma, the variations of the soliton amplitude and widths are observed during the investigations of the plasma. All these things are studied in details in our present study. We have observed that the phase velocity of all the DASWs decreases with the increase in the gradient parameter connecting inhomogeneous factors *x*∕*L* and non-extensive electrons (*q*). However, using the RPT, we have derived a modifed ZK (m-ZK) equation with variable coefficients. During this investigation, we have observed a critical point for the nonlinear propagation of dressed (rarefactive/compressive) solitons in the plasma. After using a transformation equation, we obtained two type of soliton solution. One of them is amplitude factor, and the other is of *sech*<sup>2</sup> type soliton solution which is of bell shaped. The above results and discussions also indicate that with the increase in the *x*∕*L* or *q*, increases the width of DASW in the plasma. It is also observed that the variation of soliton amplitude remains same under the infuence of magnetic feld in the given considered plasma. But the magnetic feld infuences the speed of solitons or the phase velocity *M*. Our study may be helpful in understanding the various nonlinear wave phenomena of dusty plasma under the effect of nonextensive electrons and other physical situations for both theoretical and experimental studies.

**Data availability** There is no data associated with it.

### **Declarations**

**Conflict of interest** The authors declare no competing interests.

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