

Ion Acoustic Shock Wave Excitations Around the Critical Values in an Unmagnetized Pair–Ion Plasma

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Abstract

This work deals with the ion-acoustic shock waves (IASWs) around the critical values in an unmagnetized pair–ions plasma with (α, q) -distributed electrons by formulating the correct stationary solution of Burgers-type equations with higher-order corrections. By considering higher-order correction of the reductive perturbation technique, the modifed Burgers(mB)-, and mixed modified Burgers(mmB)-type equations are derived. With the changes of viscosity coefficients of positive and negative ions, the electrostatic IASWs and normalized electric felds are investigated around the critical values. The current studies might be very useful to understand the behavior of shocks around the critical values in the F- and D-regions of Earth's ionosphere, and the later experimental verifcation in plasma laboratory.

Keywords Unmagnetized pair–ions plasma · Shock · mB-type equation · mmB-type equation · Stationary shock wave solution

1 Introduction

It is well confrmed that plasmas with negative ions (NIs) have attracted a lot of interest from the investigators because of extensive range of technological tools such as a neutral metal source [[1](#page-7-0)], semiconductor and material processing [[2](#page-7-1)], Q-machine plasmas [\[3](#page-7-2)], etc. Production of NIs by the inclusion of a small amount of SF_6 gas to low levels ($\approx 0.2eV$) of potassium plasma in a Q-machine is already confrmed in Ref. [\[4](#page-7-3)]. Goeler and Ohe [\[5](#page-7-4)] generated a *CsCl* beam into a hot tungsten plate of a Q-plasma machine that produces plasma containing *Cs*⁺, *Cl*[−] and electrons. In addition, NIs are found in the astrophysical environments, particularly in the D-, and F-region of the Earth's ionosphere [[6–](#page-7-5)[10\]](#page-7-6). It is a convincing evidence that the Cassini spacecraft has confrmed the exitance of heavy NIs in the upper region of Titan's atmosphere [\[6](#page-7-5)]. Due to the existence of pair ions not only in laboratory but also in many space and astronomical environment, many researchers [[10](#page-7-6)[–15](#page-7-7)] investigated the propagation of IASWs in plasmas composing of various types of positive ions (PIs), (e.g., K^+ , Cs^+ , C_{60}^+ , etc.) and NIs (e.g., SF_6^- , Cs^- , C_{60}^- , etc.).

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 1 Department of Mathematics, Chittagong University of Engineering and Technology, Chittagong, Bangladesh However, the most reliable method is the formation of shocks for a reason that it spreads further. Actually, the IASWs are produced in plasmas because the relative velocity between the rarefaction wave and the plasma overtakes its IA speed, where the frequency of the rarefaction wave and the surrounding medium is comparatively the same. Adak et al. [[14\]](#page-7-8) reported IASWs in the (C_{60}^+, C_{60}^-) plasma by obtaining the KdV Burgers equation. Hussain et al. [[15\]](#page-7-7) theoretically described IASWs in negative ions plasma having nonextensive electrons. Recently, Alam and Talukder [\[11\]](#page-7-9) investigated collisions between two IASWs by suggesting that the plasma system consists of pair ions and isothermal electrons. Furthermore, the energy of electrons may be nonthermal and less (subthermal) or greater (superthermal) than its isothermal energy in many environments. When the energy of electrons becomes subthermal or superthermal, one can use the nonextensive velocity distribution function to determine the density function of electron as $N_e = N_{e0} \left[1 + (q-1)(e\Phi/k_B T_e) \right]^{\frac{7(q+1)}{2(q-1)}},$ where Φ , N_{e0} , q , T_e , e and k_B are the electrostatic potential, unperturbed electron density, strength of nonextensivity, electron temperature, electron charge and Boltzmann constant, respectively $[16]$ $[16]$. The effects of electron nonextensivity on IASWs have already studied by many researchers [\[16](#page-7-10)[–20](#page-8-0)]. For instance, Ferdousi et al. [\[20](#page-8-0)] have reported the effects of electron nonextensivity on the properties of IASWs in an unmagnetized three-component plasma. They have shown that

SWEs are signifcantly modifed and both compressive and rarefective shock waves are supported by the infuence of the strength of nonextensivity. But, the nonextensive distribution function fails to address the physical issues for the nonthermal electron populations. As a result, one can consider model an electron distribution with a population of fast particle by taking the Cairns velocity distribution function $[21-23]$ $[21-23]$ leading to $N_e = N_{e0} \left[1 - (4\alpha/1 + 3\alpha)(e\Phi/k_B T_e) + (4\alpha/1$ $(+3\alpha)(e\Phi/k_BT_e)^2$ exp $(e\Phi/k_BT_e)$, where α is a parameter determining the fast particles present in the model considered. Later, Tribeche et al. [\[24\]](#page-8-3) have proposed a unique distribution function, the so-called the (α, q) -velocity distribution function by generalizing the work of Cairns et al. [\[21\]](#page-8-1), which provides a better ft of the space observations due to the fexibility provided by the nonextensivity parameter. They have obtained electron density by integrating (α, q) -velocity distribution function over all velocity space as

$$
N_e = N_{e0} \left[1 + (q - 1) \frac{e\Phi}{k_B T_e} \right]_{\frac{q(q+1)}{2(q-1)}}^{q(p+1)} \left[1 - \frac{16qa}{3 - 14q + 15q^2 + 12\alpha} \left(\frac{e\Phi}{k_B T_e} \right) + \frac{16q(2q - 1)\alpha}{3 - 14q + 15q^2 + 12\alpha} \left(\frac{e\Phi}{k_B T_e} \right)^2 \right].
$$

One can easily reduce the above electron density to the nonextensive and nonthermal electron density by setting $\alpha = 0$ and $q \rightarrow 1$, respectively. Thus, a unique distribution function, like (α, q) -velocity distribution function [\[24\]](#page-8-3) can be assumed for examining all the cases of electrons energies. Very recently, Hafez et al. [[10\]](#page-7-6) proposed an NI plasma for understanding the nature of shocks and overtaking collisions of multishocks by deriving a Burgers-like equation. They reported that the compressive and rarefactive electrostatic shocks are supported in the aforementioned plasma system. But, they ignored the features of electrostatic shocks not only around the critical values (CVs) but also at the CVs. Also, the shock waves phenomena around CVs are reported by incorrectly defned solution of mB and mmB equations in most of the previous studies [\[25](#page-8-4)[–27\]](#page-8-5), which is not useful for further verifcation in laboratory plasmas. It is therefore essential to study shock wave propagation in the plasmas around CVs by formulating the appropriate solutions of mB and mmB equations.

Thus, this work explores the electrostatic nonlinear propagation of IASWs, not only around CVs but also at CVs in a pair ions plasma system, by deriving mB- and mmB-type equations with their useful solutions. The efect of some parameters on the shocks around CVs and almost at CVs are investigated.

2 Governing Equations

To achieve our goal, the following normalized model equations are considered, by assuming PI (M_{+i} and temperature T_{+i}), NI (like SF_6^- with mass M_{-i} and temperature T_{-i}) and (α, q) - distributed electrons, along with $1 = N_{r1} + N_{r2}$, where $N_{r1} = N_{-i0}/N_{+i0}$, $N_{r2} = N_{e0}/N_{+i0}$, $N_{+i0}(N_{-i0})$ and N_{e0} are respectively the PIs (NIs) and electrons unperturbed densities [\[10](#page-7-6)]:

$$
\frac{\partial N_{ji}}{\partial t} + \frac{\partial (N_{ji} U_{ji})}{\partial z} = 0, j = +, -,
$$
\n(1)

$$
\frac{\partial U_{+i}}{\partial t} + U_{+i} \frac{\partial U_{+i}}{\partial z} + \frac{\partial \Phi}{\partial z} + \mu_{+i} \frac{\partial^2 U_{+i}}{\partial z^2} = 0,\tag{2}
$$

$$
\frac{\partial U_{-i}}{\partial t} + U_{-i} \frac{\partial U_{-i}}{\partial z} - M_{\pm} \frac{\partial \Phi}{\partial z} + \frac{\delta_{ei}}{N_{-i}} \frac{\partial N_{-i}}{\partial z} + \mu_{-i} \frac{\partial^2 U_{-i}}{\partial z^2} = 0,
$$
\n(3)

$$
\frac{\partial^2 \Phi}{\partial z^2} = N_{pi} - N_{ni} - N_{r2} \left\{ \left[1 + (q - 1)\Phi \right]^{\frac{(q+1)}{2(q-1)}} \times \left[1 - \frac{16qa}{3 - 14q + 15q^2 + 12\alpha} \Phi \right] + \frac{16q(2q - 1)\alpha}{3 - 14q + 15q^2 + 12\alpha} \Phi^2 \right] \right\},
$$
\n(4)

where,

$$
M_{\pm} = \frac{M_{+i}}{M_{-i}}, \delta_{ei} = \frac{T_e}{(1 - N_{r1})T_{-i}}.
$$
\n(5)

Here, N_{+i} (N_{-i}) and N_e are respectively the normalized PIs (NIs) and electrons densities normalized by N_{+i0} , U_{+i} (U_{-i}) is the normalized PIs (NIs) fuid velocity normalized by the PI speed $C_{+is} = \sqrt{k_B T_e / M_{\pm}} \sqrt{(N_{r1} + N_{r2}) / N_{r2} (1 - N_{r1})}$, Φ is the normalized electrostatic potential, $\Phi \rightarrow e\Phi/k_BT_e$, *t* is the time variable normalized by $\omega_{+i}^{-1} = \lambda_{De} / C_{+iS}$, *z* is the space variable normalized by $\lambda_{De} = \sqrt[k]{k_B T_e / 4 \pi N_{e0} e^2}$, and μ_{+i} (μ_{-i}) is the normalized PIs (NIs) viscosity coefficient normalized by $\omega_{+i}^{-1}/M_{+}N_{+i0}C_{+i}^{2}$ ($\omega_{+i}^{-1}/M_{-i}N_{-i0}C_{+i}^{2}$). Additionally, one can use (i) $q \rightarrow 1$ and $\alpha = 0$ for the Maxwell-Boltzmann distributed electrons, (ii) $q \to 1$ and $\alpha \neq 0$ for the Cairns distributed electrons and (iii) $\alpha = 0$ for superthermal $(0 < q < 1)$ and subthermal $(q > 1)$ electrons, respectively, where *q* is the strength of nonextensivity and α is measuring the population of nonthermal electrons.

3 Mathematical Analysis

3.1 Formation of mBE with Stationary Wave Solution

In order to study the shock wave phenomena around the critical values, one can consider the stretching coordinates by taking the higher-order correction of reductive perturbative method [\[25\]](#page-8-4) as

$$
X = \epsilon^2 (z - V_p t), T = \epsilon^4 t,
$$
\n⁽⁶⁾

and the expansions for physical quantities as

$$
\begin{bmatrix} N_{+i} \\ N_{-i} \\ U_{+i} \\ U_{+i} \\ \Phi \end{bmatrix} = \begin{bmatrix} 1 \\ N_{r1} \\ 0 \\ 0 \\ 0 \end{bmatrix} + \sum_{i}^{\infty} \epsilon^{i} \begin{bmatrix} N_{+i}^{(i)} \\ N_{-i}^{(i)} \\ U_{-i}^{(i)} \\ U_{-i}^{(i)} \\ U_{-i}^{(i)} \\ \Phi^{(i)} \end{bmatrix},
$$
(7)

where V_p and ϵ are the IA linear phase speed and a small quantity measuring the weakness of dissipation. By system-atically inserting Eqs. [\(6](#page-2-0)) and [\(7\)](#page-2-1) into Eqs. ([1\)](#page-1-0)–[\(4](#page-1-1)), the different order of ϵ equations are easily determined. The $O(\epsilon^3)$ -order equations gives the following expressions for $N_{pi}^{(1)}$, $N_{ni}^{(1)}$, $U_{ni}^{(1)}$ and V_p :

$$
C_f = \left[\frac{3}{V_p^4} - \frac{3N_{r1}V_p^2 M_{\pm}^2}{\left(V_p^2 - \delta_{ei}\right)^3} + \frac{\delta_{ei}N_{r1}M_{\pm}^2}{\left(V_p^2 - \delta_{ei}\right)^3} - 2N_{r2}\Omega_2 \right], \quad (12)
$$

and

 \mathbf{r}

$$
\Omega_2 = \left[\frac{(q+1)(3-q)}{8} - \frac{q+1}{2} \times \frac{16q\alpha}{3 - 14q + 15q^2 + 12\alpha} + \frac{16q(2q-1)\alpha}{3} - 14q + 15q^2 + 12\alpha \right].
$$
\n(13)

It is noted that one can determine the critical values (CVs) by setting $C_f = 0$. But, it is very difficult to formulate the mathematical expression for CV. Figure [1](#page-2-3) shows the variation of C_f with regard to q and α by considering the other

$$
N_{+i}^{(1)} = \frac{1}{V_p^2} \Phi^{(1)}, U_{+i}^{(1)} = \frac{1}{V_p} \Phi^{(1)}, N_{-i}^{(1)} = -\frac{N_{r1} M_{\pm}}{V_p^2 - \delta_{ei}} \Phi^{(1)}, U_{-i}^{(1)} = -\frac{V_p M_{\pm}}{V_p^2 - \delta_{ei}} \Phi^{(1)},
$$
\n(8)

$$
V_p = \sqrt{\frac{\left(1 + N_{r1}M_{\pm} + \delta_{ei}N_{r2}\Omega_1\right) \pm \sqrt{\left(1 + N_{r1}M_{\pm} + \delta_{ei}N_{r2}\Omega_1\right)^2 - 4N_{r2}\Omega_1\delta_{ei}}}{2N_{r2}\delta_{ei}},\tag{9}
$$

where

$$
\Omega_1 = \left[\frac{q+1}{2} - \frac{16q\alpha}{3 - 14q + 15q^2 + 12\alpha} \right].
$$
 (10)

It is obviously found from Eq. [\(9\)](#page-2-2) that V_p is strongly dependent on N_{r1} , M_{\pm} , δ_{ei} , N_{r2} , α and q , but not on μ_{+i} and μ_{-i} . It is validated for $(1 + N_{r1}M_{\pm} + \delta_{ei}N_{r2}\Omega_1)^2 - 4N_{r2}\Omega_1\delta_{ei} \ge 0$. Again, the $O(\epsilon^4)$ -order equations (ignored for convenience) yields the following relations:

$$
N_{+i}^{(2)} = \frac{1}{V_p^2} \left[\frac{3}{2V_p^2} {\Phi^{(1)}}^2 + {\Phi^{(2)}} \right],
$$

\n
$$
U_{+i}^{(2)} = \frac{1}{V_p^2} \left[\frac{1}{2V_p^2} {\Phi^{(1)}}^2 + {\Phi^{(2)}} \right],
$$

\n
$$
N_{-i}^{(2)} = \frac{N_{r1}M_{\pm}}{V_p^2 - \delta_{ei}} \left[\frac{M_{\pm}(3V_p^2 - \delta_{ei})}{2(V_p^2 - \delta_{ei})^2} {\Phi^{(1)}}^2 - {\Phi^{(2)}} \right],
$$

\n
$$
U_{-i}^{(1)} = \frac{V_pM_{\pm}}{V_p^2 - \delta_{ei}} \left[\frac{M_{\pm}(V_p^2 + \delta_{ei})}{2(V_p^2 - \delta_{ei})^2} {\Phi^{(1)}}^2 - {\Phi^{(2)}} \right],
$$

\n
$$
- C_f {\Phi^{(1)}}^2 = 0,
$$

\n(11)

where

parameters constant. It is obviously found from Fig. [1](#page-2-3) that the CVs only occur when the electron energy becomes less than its thermal energy, that is, for the subthermal

Fig. 1 Variation of C_f with regard to q and α . The other parameters are considered as $\dot{M}_r = 3.75$, $T_{ni} = 0.05eV$, $T_e = 0.2eV$, $N_{r1} = 0.5$ and $N_{r2} = 0.01$

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electrons based on the considered parametric values as in Ref. [[10](#page-7-6), [11\]](#page-7-9). For instance, the CV q_c for q is determined as $q_c = 3.098638852$ by considering $\alpha = 0$, $M_+ = 3.75$, $T_{-i} = 0.05eV$, $T_e = 0.19eV$, $N_{r1} = 0.5$ and $N_{r2} = 0.01$. One can also find the other CV as $q_c = 6.176859718$ by considering α = 0, M_{\pm} = 3.75, T_{-i} = 0.05*eV*, T_e = 0.1*eV*, N_{r1} = 0.5 and $N_{r2} = 0.1$. It is also provided that the CVs are only supported for subthermal electrons.

Finally, the $O(\epsilon^5)$ -order provides the following equations:

$$
\frac{1}{V_p^2} \frac{\partial \Phi^{(1)}}{\partial T} - V_p \frac{\partial N_{+i}^{(3)}}{\partial X} + \frac{\partial U_{+i}^{(3)}}{\partial X} + \frac{6}{V_p^5} {\{\Phi^{(1)}\}}^2
$$
\n
$$
\frac{\partial \Phi^{(1)}}{\partial X} + \frac{2}{V_p^3} \frac{\partial}{\partial X} [\Phi^{(1)} \Phi^{(2)}] = 0,
$$
\n(14)

$$
\frac{1}{V_p} \frac{\partial \Phi^{(1)}}{\partial T} - V_p \frac{\partial U_{+i}^{(3)}}{\partial X} + \frac{\partial \Phi^{(3)}}{\partial X} + \frac{3}{2V_p^4} {\{\Phi^{(1)}\}}^2
$$
\n
$$
\frac{\partial \Phi^{(1)}}{\partial X} + \frac{1}{V_p^2} \frac{\partial}{\partial X} [\Phi^{(1)} \Phi^{(2)}] + \frac{\mu_{+i}}{V_p} \frac{\partial^2 \Phi^{(1)}}{\partial X^2} = 0,
$$
\n(15)

$$
-\frac{N_{r1}M_{\pm}}{V_{p}^{2} - \delta_{ei}}\frac{\partial \Phi^{(1)}}{\partial T} - V_{p}\frac{\partial N_{-i}^{(3)}}{\partial X} + N_{r1}\frac{\partial U_{-i}^{(3)}}{\partial X} - \frac{6N_{r1}V_{p}^{3}M_{\pm}^{3}}{\left(V_{p}^{2} - \delta_{ei}\right)^{4}} + \frac{2N_{r1}V_{p}M_{\pm}^{2}}{\left(V_{p}^{2} - \delta_{ei}\right)^{2}}\frac{\partial}{\partial X}\left[\Phi^{(1)}\Phi^{(2)}\right] = 0,
$$
\n(16)

$$
-\frac{V_p M_{\pm}}{V_p^2 - \delta_{ei}} \frac{\partial \Phi^{(1)}}{\partial T} - V_p \frac{\partial U_{-i}^{(3)}}{\partial X} + \frac{\delta_{ei}}{N_{r1}} \frac{\partial N_{-i}^{(3)}}{\partial X} - M_{\pm} \frac{\partial \Phi^{(3)}}{\partial X} -\frac{\mu_{-i} V_p M_{\pm}}{V_p^2 - \delta_{ei}} \frac{\partial^2 \Phi^{(1)}}{\partial X^2} + \frac{M_{\pm}^2}{(V_p^2 - \delta_{ei})} \frac{\partial}{\partial X} [\Phi^{(1)} \Phi^{(2)}] + \left[\frac{3V_p^2 M_{\pm}^3 (V_p^2 + \delta_{ei})}{2(V_p^2 - \delta_{ei})} + \frac{3\delta_{ei} M_{\pm}^3 (3V_p^2 - \delta_{ei})}{2(V_p^2 - \delta_{ei})^4} - \frac{\delta_{ei} M_{\pm}^3}{(V_p^2 - \delta_{ei})^3} \right] {\Phi^{(1)}}^2 \frac{\partial \Phi^{(1)}}{\partial X} = 0,
$$
\n(17)

$$
0 = -N_{r2}\Omega_1 \frac{\partial \Phi^{(3)}}{\partial X} - 3N_{r2}\Omega_3 [\Phi^{(1)}]^{2} \frac{\partial \Phi^{(1)}}{\partial X}
$$

$$
-2N_{r2}\Omega_2 \frac{\partial}{\partial X} [\Phi^{(1)}\Phi^{(2)}] + \frac{\partial N_{+i}^{(3)}}{\partial X} - \frac{\partial N_{-i}^{(3)}}{\partial X}, \qquad (18)
$$

where,

$$
\Omega_3 = \left[\frac{(q+1)(3-q)(5-q)}{48} - \frac{2q\alpha(q+1)(3-q)}{3 - 14q + 15q^2 + 12} + \frac{8q(2q-1)\alpha(q+1)}{3 - 14q + 15q^2 + 12} \right].
$$
\n(19)

By simplifying the above equations, the following equation is derived:

$$
A\frac{\partial \Phi^{(1)}}{\partial T} + B\{\Phi^{(1)}\}^2 \frac{\partial \Phi^{(1)}}{\partial X} + C\frac{\partial^2 \Phi^{(1)}}{\partial X^2} = 0,
$$
 (20)

which is the mB-type equation. The coefficients of Eq. (20) (20) are determined as

$$
A = 2\left[\frac{V_p^2 - \delta_{ei}}{V_p} + \frac{N_{r1}V_p^3 M_{\pm}}{(V_p^2 - \delta_{ei})}\right],
$$

\n
$$
B = \frac{15(V_p^2 - \delta_{ei})}{2V_p^4} - 3N_{r2}\Omega_3 V_p^2 (V_p^2 - \delta_{ei})
$$

\n
$$
+ \frac{3N_{r1}V_p^2 M_{\pm}^3 \left(5V_p^4 - 2V_p^2 \delta_{ei} + \delta_{ei}^2\right)}{2\left(V_p^2 - \delta_{ei}\right)^4}
$$

\n
$$
+ \frac{\delta_{ei}N_{r1}M_{\pm}^3 V_p^2}{\left(V_p^2 - \delta_{ei}\right)^3},
$$

\n
$$
C = \frac{\mu_{+i}\left(V_p^2 - \delta_{ei}\right)}{V_p} + \frac{N_{r1}\mu_{-i}V_p^3 M_{\pm}}{\left(V_p^2 - \delta_{ei}\right)}.
$$
 (21)

In the previous literature $[25-27]$ $[25-27]$ $[25-27]$, the solution of mB equation (see [Appendix](#page-7-11)) is incorrectly defned, which is not useful for further verifcation in laboratory plasmas. To determine the correct stationary shock wave solution of Eq. (20) , one can convert Eq. (20) (20) (20) by considering $\Phi^{(1)}(X,T) = \Psi(\zeta)$ with $\zeta = X - V_rT$ (V_r is the constant speed of reference frame) to the following form:

$$
C\frac{d\Psi}{d\zeta} = V_r A \Psi - \frac{B}{3} \Psi^3. \tag{22}
$$

This implies that

$$
\int \frac{C}{V_r A \Psi - \frac{B}{3} \Psi^3} d\Psi = \int d\zeta.
$$
\n(23)

By simplifying Eq. (23) (23) (23) , the stationary shock wave solution of Eq. ([20\)](#page-3-0) is obtained as

$$
\Phi^{(1)} = \sqrt{\Phi_A \left\{ 1 + \tan h \left(\frac{\zeta}{\Phi_W} \right) \right\}},\tag{24}
$$

where $\Phi_A = (3V_rA/2B)$ and $\Phi_W = (C/V_rB)$ are respectively the amplitude and thickness of IASWs around the critical values. The verifcation of the obtained solution is given in [Appendix.](#page-7-11)

3.2 Formation of mmB‑Type Equation with Wave Solution

One can easily find that Eq. (20) is not useful to study the shock wave phenomena not only at CVs but also around CVs. To study the electrostatic shocks not only around CVs but also at CVs, one needs to derive another evolution equation. To do so, one can take $C_f = C_f^0$ for *q* (say) around its q_c as

$$
C_f^0 = h \left(\frac{\partial C_f}{\partial q} \right)_{q = q_c} |q - q_c| = S G \epsilon,
$$
\n(25)

where $|q - q_v| \equiv \epsilon$ (because $|q - q_c|$ is small quantity), $S = 1(-1)$ for $q > q_c (q < q_c)$ and $G = \left(\frac{\partial C_f}{\partial q_c}\right)$ *𝜕q* λ . As a result, $q=q_c$ one can re-evaluate from Eq. ([4\)](#page-1-1) by adding $\rho^{(2)} \equiv -\varepsilon^3 \frac{1}{2} S G \Phi^2$ with the $O(\epsilon^5)$ equation and yields

$$
0 = -N_{r2}\Omega_1 \frac{\partial \Phi^{(3)}}{\partial X} + SG\Phi^{(1)}\frac{\partial \Phi^{(1)}}{\partial X} - 3N_{r2}\Omega_3 [\Phi^{(1)}]^2 \frac{\partial \Phi^{(1)}}{\partial X}
$$

$$
-3N_{r2}\Omega_2 \frac{\partial}{\partial X} [\Phi^{(1)}\Phi^{(2)}] + \frac{\partial N_{+i}^{(3)}}{\partial X} - \frac{\partial N_{-i}^{(3)}}{\partial X},
$$
(26)

Now, simplifying Eqs. (14) (14) – (17) (17) and (26) (26) , one obtains

$$
A\frac{\partial\Phi^{(1)}}{\partial T} + SD\Phi^{(1)}\frac{\partial\Phi^{(1)}}{\partial X} + B\{\Phi^{(1)}\}^2\frac{\partial\Phi^{(1)}}{\partial X} + C\frac{\partial^2\Phi^{(1)}}{\partial x^2} = 0,
$$
\n(27)

where

$$
D = V_p^2 \left(V_p^2 - \delta_{ei} \right) \left(\frac{\partial C_f}{\partial q} \right)_{q = q_c}.
$$
 (28)

Equation ([27](#page-4-1)) is so-called the mmB-type equation because the additional nonlinear term is occurred with the mBE. One can easily convert mmB-type equation not only to mB-type equation but also to B-type equation. It is noted that Eq. ([27](#page-4-1)) does not support the IASWs around CVs but also at CVs, which is the main advantage of this equation.

The appropriate stationary shock wave solution of mmBtype equation can be defned as

$$
\Phi^{(1)} = \sqrt{\Phi_{mA} \left\{ 1 + \tan h \left(\frac{\zeta}{\Phi_{mW}} \right) \right\}} - \frac{SD}{2B},\tag{29}
$$

where

$$
\Phi_{mA} = (3Ck/4B), \Phi_W = (2/k),
$$
\n
$$
k = \left[S^2 \left(\frac{D}{A} \right)^2 \pm \sqrt{S^4 \left(\frac{D}{A} \right)^4 + 32CV_r \left(\frac{B}{A} \right)^2} \right] \div \left[4C \left(\frac{B}{A} \right) \right].
$$
\n(30)

4 Discussion

This section explores the effect of μ_{+i} , μ_{-i} and V_r with physical explanation on the electrostatic shocks around CVs and at the neighborhood of CVs.

Figure [2a](#page-5-0), b display the effect of μ_{+i} and μ_{-i} on the electrostatic IASWs around q_c . It is observed that μ_{+i} and μ_{-i} strongly plays an important role to the formation of monotonically shock waves around q_c . In addition, the thickness of shocks are remarkably increased and slightly increased with the increases of μ_{+i} and μ_{-i} , respectively. It is noted that the effect of the viscosity coefficient of NIs and PIs on shock wave excitation can be determined on the basis of collective friction between the layers of the plasma concentration system. In fact, viscosity is the force of collective friction between layers of fluid in the aforementioned plasmas. With the decrease of μ_{+i} and μ_{-i} , the collective friction force is decreased. As a result, the thickness of shocks is decreased. The effect of V_r on electrostatic IASWs around q_c is displayed in Fig. [3](#page-5-1). Figure 3 obviously shows that the amplitude and thickness of IASWs are significantly increased with the increase of V_r . Finally, the variation of normalized electric field $(E = -\text{grad}\Phi)$ around q_c for different values of μ_{+i} and μ_{-i} is presented in Fig. [4.](#page-6-0) It is observed that the electric field becomes monotonically hump-shaped with the increase of μ_{+i} and μ_{-i} . Consequently, the electric field is propagating narrowly with the increase of μ_{+i} , but smoothly with the increase of μ_{-i} (See Fig. [4](#page-6-0)).

Figure [5](#page-6-1)a, b shows the shape of electrostatic ion acoustic mmB shocks for $q = 6 < q_c = 6.176859718$ and $q = 6.5 > q_c = 6.176859718$ with $M_r = 3.75$, $T_{ni} = 0.05$, $T_e = 0.1$, $N_{r1} = 0.5$, $N_{r2} = 0.1$, $V_r = 0.1$, $\mu_{-i} = 0.01$ and $\mu_{+i} = 0.5$. Whereas Fig. [6a](#page-7-12), b display the effect of μ_{+i} and μ_{-i} on the electrostatic IASWs at the neighboring of q_c , that is at $q = 6.5$. It is found that the monotonic correct solution of the shocks are only generated for the viscosity coeffcient of NIs in which both amplitude and thickness are increased with the increase of the viscosity coefficient of

Fig. 2 Electrostatic mB shocks around the critical value $q_c = 3.098638852$, that is $q = 3.5$ for different values of (**a**) μ_{-i} and (**b**) μ_{+i} , with $M_r = 3.75$, $T_{ni} = 0.05$, $T_e = 0.19$, $N_{r1} = 0.5$, $N_{r2} = 0.01$ and $V_r = 0.01$

NIs at the neighboring CVs. However, the amplitude and thickness of IASWs are decreased with the increase of the viscosity coefficient of PIs at the neighboring CVs. It might be predicted from this work that one can study the real

Fig. 3 Effect of the reference speed V_r on electrostatic mB shocks around the critical value $q_c = 3.098638852$, that is $q = 3.5$, with $M_r = 3.75$, $T_{ni} = 0.05$, $T_e = 0.19$, $N_{r1} = 0.5$, $N_{r2} = 0.01$, $\mu_{-i} = 0.05$, and $\mu_{+i} = 0.3$

features of shock wave excitations around CVs by formulating the correct solution of the mB and mmB equations.

To conclude, the plasma system involving pair-ions and generalized distributed electrons has been considered for reporting the features of electrostatic IASWs around CVs. To do so, the mB- and mmB-type equations have been obtained by employing the reductive perturbation method, which reveals the shocks only around the critical values in the plasmas. The correct stationary shock wave solutions of these equations are first time reported. The effect of μ_{+i} , μ_{-i} and V_r on the electrostatic mB and mmB shocks around and at the neighboring of CVs are discussed. It is found that the mB-type equation is supporting monotonic shocks around CVs in which the amplitude is unchanged but the thickness is increased with the increase of μ_{+i} and μ_{-i} . Subsequently, the monotonically hump-shaped electric field is propagating narrowly and smoothly with the increase of μ_{+i} and μ_{-i} , respectively, due to the correct solution of the mB-type equation. However, the mmB-type equation supports monotonic shocks around CVs with the influence of μ_{-i} only, in which both the amplitudes and the thickness are increased with the increase of μ_{-i} but decreased with the increase of μ_{+i} . It may be concluded that the results presented in this work are not only very helpful to understand the broadband shocks noise in the D- and F-regions of the Earth's ionosphere around the critical values, but also in plasma laboratory experiments.

Fig. 4 Variation of the normalized electric field around the critical value $q_c = 3.098638852$, that is $q = 3.5$, for different values of (**a**) μ_{-i} and (**b**) μ_{+i} , with the same typical values of Fig. [1](#page-2-3)

Fig. 5 Electrostatic mmB shocks (a) $q = 6 < q_c = 6.176859718$ and (b) $q = 6.5 > q_c = 6.176859718$, with $M_r = 3.75$, $T_{ni} = 0.05$, $T_e = 0.1$, $N_{r1} = 0.5$, $N_{r2} = 0.1$, $V_r = 0.1$, $\mu_{-i} = 0.01$ and $\mu_{+i} = 0.5$

Fig. 6 Effect of (**a**) μ_{-i} and (**b**) μ_{+i} on the mmB shocks at $q = 6.5$

Appendix

Verification of solution as in Eq. [\(23](#page-3-1)) with the aid of Maple 18:

$$
\begin{aligned}\n>\zeta := X - V_r T \\
>\Phi_A := (3V_r A/2B) \\
>\Phi_W := (C/V_r B) \\
>\Phi^{(1)} = \sqrt{\Phi_A \left\{ 1 + \tan h \left(\frac{\zeta}{\Phi_W} \right) \right\}} \\
> \text{diff}(\Phi^{(1)}, T) \\
> \text{diff}(\Phi^{(1)}, X) \\
> \text{diff}(\Phi^{(1)}, X, X) \\
> A * \text{diff}(\Phi^{(1)}, T) + B * (\Phi^{(1)})^2 * \text{diff}(\Phi^{(1)}, X) + C * \text{diff}(\Phi^{(1)}, X, X) \\
> 0\n\end{aligned}
$$

It is noted that the stationary solution

$$
\gg \phi^{(1)} = \sqrt{\left(\frac{3u_0}{2B}\right)\left\{1 - \tan h \left(\frac{\zeta - u_0 \tau}{\sqrt{\frac{C}{u_0}}}\right)\right\}}
$$

provided in the previous literature [[25–](#page-8-4)[27](#page-8-5)] does not satisfy the following modifed Burgers equation:

$$
\frac{\partial \phi^{(1)}}{\partial \tau} + B\left\{\phi^{(1)}\right\}^2 \frac{\partial \Phi^{(1)}}{\partial \zeta} = C \frac{\partial^2 \phi^{(1)}}{\partial \zeta^2},
$$

which clearly indicates that the above solution of mB equation is not useful for later experimental verifcation.

Data Availability Data sharing is not applicable to this article as no new data were created or analyzed in this study.

Declarations

Conflicts of Interest The authors declare that they have no confict of interest concerning the publication of this manuscript.

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