ORIGINAL PAPER



# Characterizations of complete spacelike submanifolds in the (n + p)-dimensional anti-de Sitter space of index q

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Received: 29 October 2015 / Accepted: 23 August 2016 / Published online: 1 September 2016 © Springer-Verlag Italia 2016

**Abstract** Our purpose in this paper is to study the geometry of *n*-dimensional complete spacelike submanifolds immersed in the (n + p)-dimensional anti-de Sitter space  $\mathbb{H}_q^{n+p}$  of index *q*, with  $1 \le q \le p$ . Under suitable constraints on the Ricci curvature and the second fundamental form, we show that a complete maximal spacelike submanifold of  $\mathbb{H}_q^{n+p}$  must be totally geodesic. Furthermore, we establish sufficient conditions to guarantee that a complete spacelike submanifold with nonzero parallel mean curvature vector in  $\mathbb{H}_p^{n+p}$  must be pseudo-umbilical, which means that its mean curvature vector is an umbilical direction.

**Keywords** Anti-de Sitter space · Complete spacelike submanifolds · Totally geodesic submanifolds · Parallel mean curvature vector · Pseudo-umbilical submanifolds

Mathematics Subject Classification Primary 53C42; Secondary 53A10 · 53C20 · 53C50

## **1** Introduction

Apart from their physical importance (see, for example, [25, 34]), the interest in the study of spacelike submanifolds immersed in a Lorentzian space is motivated by their nice Bernstein-type properties. For instance, it was proved by Calabi [10] (for  $n \le 4$ ) and by Cheng and Yau [14] (for all n) that the only complete maximal spacelike hypersurfaces of the Lorentz-Minkowski space  $\mathbb{L}^{n+1}$  are the spacelike hyperplanes. In [29], Nishikawa proved that a complete maximal spacelike hypersurface (that is, with mean curvature identically zero) in

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the de Sitter space  $\mathbb{S}_1^{n+1}$  must be totally geodesic. In [18], Goddard conjectured that the complete spacelike hypersurfaces of  $\mathbb{S}_1^{n+1}$  with constant mean curvature H must be totally umbilical. Ramanathan [32] proved Goddard's conjecture in  $\mathbb{S}_1^3$  for  $0 \le H \le 1$ . Moreover, for H > 1, he showed that the conjecture is false, as can be seen from an example due to Dajczer and Nomizu in [16]. Independently, Akutagawa [2] proved that Goddard's conjecture is true when either n = 2 and  $H^2 \le 1$  or  $n \ge 3$  and  $H^2 < \frac{4(n-1)}{n^2}$ . He also constructed complete spacelike rotation surfaces in  $\mathbb{S}_1^3$  having constant mean curvature H > 1 and which are not totally umbilical. Next, Montiel [26] showed that Goddard's conjecture is true provided that  $M^n$  is compact. Furthermore, he exhibited examples of complete spacelike hypersurfaces in  $\mathbb{S}_1^{n+1}$  with constant mean curvature H satisfying  $H^2 \ge \frac{4(n-1)}{n^2}$  and being non totally umbilical, the so-called hyperbolic cylinders.

In higher codimension, Cheng [12] extended Akutagawa's result for complete spacelike submanifolds with parallel mean curvature vector (that is, the mean curvature vector field is parallel as a section of the normal bundle) in the de Sitter space  $\mathbb{S}_p^{n+p}$  of index p. Afterwards, Aiyama [1] studied compact spacelike submanifolds in  $\mathbb{S}_p^{n+p}$  with parallel mean curvature vector and proved that if the normal connection of  $M^n$  is flat, then  $M^n$  is totally umbilical. Furthermore, she proved that a compact spacelike submanifold in  $\mathbb{S}_p^{n+p}$  with parallel mean curvature vector and nonnegative sectional curvature must be totally umbilical. Meanwhile, Alías and Romero [4] developed some integral formulas for compact spacelike submanifolds in  $\mathbb{S}_p^{n+p}$  which have a very clear geometric meaning and, as application, they obtained a Bernstein type result for complete maximal submanifolds in  $\mathbb{S}_q^{n+p}$ , extending a previous result due to Ishihara [19]. Moreover, they extended Ramanathan's result [32] showing that the only compact spacelike surfaces in  $\mathbb{S}_p^{2+p}$  with parallel mean curvature vector are the totally umbilical ones and, in particular, they also reproved Cheng's result [14] establishing that every complete spacelike surface in  $\mathbb{S}_p^{2+p}$  with parallel mean curvature vector such that  $H^2 < 1$  is totally umbilical. Next, Li [22] showed that Montiel's result [26] still holds for higher codimensional spacelike submanifolds in  $\mathbb{S}_p^{n+p}$ . More recently, Araújo and Barbosa [6], assuming appropriated controls on the second fundamental form and on the scalar curvature, extended the techniques developed in [23, 33, 35] and proved that a compact spacelike submanifold in  $\mathbb{S}_p^{n+p}$  with nonzero mean curvature and parallel mean curvature vector must be isometric to a sphere.

When the ambient spacetime is the anti-de Sitter space  $\mathbb{H}_1^{n+1}$ , Choi et al. [15] used the generalized maximum principle of Omori [30] and Yau [36] in order to obtain a Myers type theorem [28] concerning complete maximal spacelike hypersurfaces. More precisely, they showed that if the height function with respect to a timelike vector of such a hypersurface obeys a certain boundedness, then it must be totally geodesic. Extending a technique due to Yau [37], the first author jointly with Camargo [7] obtained another rigidity results to complete maximal spacelike hypersurfaces in  $\mathbb{H}_1^{n+1}$ , imposing suitable conditions on both the norm of the second fundamental form and a certain height function naturally attached to the hypersurface. Afterwards, working with a suitable warped product model for an open subset of  $\mathbb{H}_1^{n+1}$ , the same authors jointly with Caminha and Parente [8] extended the main result of [7] showing that if  $M^n$  is a complete spacelike hypersurface with constant mean curvature and bounded scalar curvature in  $\mathbb{H}_1^{n+1}$ , such that the gradient of its height function with respect to a timelike vector has integrable norm, then  $M^n$  must be totally umbilical. More recently, the first author jointly Aquino [5] obtained another characterizations theorems concerning complete constant mean curvature spacelike hypersurfaces of  $\mathbb{H}_1^{n+1}$ , under suitable constraints on the behavior of the Gauss mapping. In higher codimension, Ishihara [19] proved that a *n*dimensional complete maximal spacelike submanifold immersed in the anti-de Sitter space  $\mathbb{H}_p^{n+p}$  of index p must have the squared norm of the second fundamental form bounded from above by *np*. Moreover, the only ones that attain this estimate are the hyperbolic cylinders. Later on, Cheng [13] obtained a refinement of Ishihara's result [19] for the case of complete maximal spacelike surfaces immersed in  $\mathbb{H}_p^{2+p}$ .

Motivated by the works above described, our purpose in this paper is to study the geometry of complete spacelike submanifolds immersed in the anti-de Sitter space  $\mathbb{H}_{q}^{n+p}$  of index q. In this setting, we extend the technique due to Alías and Romero in [4] and, under appropriated constraints on the Ricci curvature and second fundamental form, we show that a complete maximal spacelike submanifold  $M^n$  of  $\mathbb{H}_q^{n+p}$  must be totally geodesic (see Theorem 1 and Corollaries 1 and 2). Furthermore, we establish sufficient conditions to guarantee that a complete spacelike submanifold with nonzero parallel mean curvature vector **H** in  $\mathbb{H}_{p}^{n+p}$ must be pseudo-umbilical, which means that H is an umbilical direction (see Theorem 2 and Corollary 3). Our approach is based on a generalized form of a maximum principle at the infinity of Yau [37] (see Lemma 1 and Remark 1).

#### 2 Preliminaries

Let  $\mathbb{R}_{q+1}^{n+p+1}$  be the (n+p+1)-dimensional semi-Euclidean space endowed with metric tensor  $\langle, \rangle$  of index q, with  $1 \le q \le p$ , given by

$$\langle v, w \rangle = \sum_{i=1}^{n+p-q} v_i w_i - \sum_{j=n+p-q+1}^{n+p+1} v_j w_j,$$

and let  $\mathbb{H}_{q}^{n+p}$  be the (n+p)-dimensional unitary anti-de Sitter space of index q, that is,

$$\mathbb{H}_{q}^{n+p} = \{ x \in \mathbb{R}_{q+1}^{n+p+1} ; \langle x, x \rangle = -1 \},\$$

which has constant sectional curvature equal to -1. Along this work, we will consider  $x: M^n \to \mathbb{H}_q^{n+p} \subset \mathbb{R}_{q+1}^{n+p+1}$  a spacelike submanifold isometrically immersed in  $\mathbb{H}_q^{n+p}$ . We recall that a submanifold immersed is said to be *spacelike* if its induced metric is positive definite. In this setting, we will denote by  $\nabla^{\circ}$ ,  $\overline{\nabla}$  and  $\nabla$  the Levi-Civita connections of  $\mathbb{R}^{n+p+1}_{q+1}$ ,  $\mathbb{H}^{n+p}_{q}$  and  $M^{n}$ , respectively, and  $\nabla^{\perp}$  will stand for the normal connection of  $M^n$  in  $\mathbb{H}_a^{n+p}$ .

We will denote by  $\alpha$  the second fundamental form of  $M^n$  in  $\mathbb{H}_q^{n+p}$  and by  $A_{\xi}$  the shape operator associated to a fixed vector field  $\xi$  normal to  $M^n$  in  $\mathbb{H}_q^{n+p}$ . We note that, for each  $\xi \in \mathfrak{X}^{\perp}(M), A_{\xi}$  is a symmetric endomorphism of the tangent space  $T_{x}M$  at  $x \in M^{n}$ . Moreover,  $A_{\xi}$  and  $\alpha$  are related by

$$\langle A_{\xi}X, Y \rangle = \langle \alpha(X, Y), \xi \rangle, \tag{2.1}$$

for all tangent vector fields  $X, Y \in \mathfrak{X}(M)$ .

We also recall that the Gauss and Weingarten formulas of  $M^n$  in  $\mathbb{H}^{n+p}_q$  are given by

$$\nabla_X^{\circ} Y = \overline{\nabla}_X Y + \langle X, Y \rangle_X = \nabla_X Y + \alpha(X, Y) + \langle X, Y \rangle_X, \tag{2.2}$$

and

$$\nabla_X^{\circ}\xi = \overline{\nabla}_X\xi = -A_{\xi}X + \nabla_X^{\perp}\xi, \qquad (2.3)$$

for all tangent vector fields  $X, Y \in \mathfrak{X}(M)$  and normal vector field  $\xi \in \mathfrak{X}^{\perp}(M)$ .

As in [31], the curvature tensor R of the spacelike submanifold  $M^n$  is given by

$$R(X, Y)Z = \nabla_{[X,Y]}Z - [\nabla_X, \nabla_Y]Z,$$

where [, ] denotes the Lie bracket and  $X, Y, Z \in \mathfrak{X}(M)$ .

A well known fact is that the curvature tensor R of  $M^n$  can be described in terms of its second fundamental form  $\alpha$  and the curvature tensor  $\overline{R}$  of the ambient spacetime  $\mathbb{H}_q^{n+p}$  by the so-called Gauss equation, which is given by

$$\langle R(X, Y)Z, W \rangle = \langle Y, Z \rangle \langle X, W \rangle - \langle X, Z \rangle \langle Y, W \rangle + \langle \alpha(X, W), \alpha(Y, Z) \rangle - \langle \alpha(X, Z), \alpha(Y, W) \rangle,$$
(2.4)

for all tangent vector fields  $X, Y, Z, W \in \mathfrak{X}(M)$ . Moreover, Codazzi equation asserts that

$$(\nabla_X A_{\xi})Y = (\nabla_Y A_{\xi})X, \tag{2.5}$$

for all  $X, Y \in \mathfrak{X}(M)$  and  $\xi \in \mathfrak{X}^{\perp}(M)$ .

We will define the mean curvature vector of  $M^n$  in  $\mathbb{H}_q^{n+p}$  by

$$\mathbf{H} = \frac{1}{n} \mathrm{tr}(\alpha).$$

We recall that  $M^n$  is called *maximal* when  $\mathbf{H} \equiv 0$  and we say that  $M^n$  has *parallel mean* curvature vector when  $\nabla_X^{\perp} \mathbf{H} \equiv 0$ , for every  $X \in \mathfrak{X}(M)$ . In this last case, when q = p and  $\mathbf{H} \neq 0$ , we have that  $\langle \mathbf{H}, \mathbf{H} \rangle$  is a negative constant along  $M^n$ . Moreover,  $M^n$  is called *totally* geodesic when its second fundamental form  $\alpha$  vanishes identically and it is called *totally* umbilical when

$$\alpha(X, Y) = \langle X, Y \rangle \mathbf{H}, \tag{2.6}$$

for all tangent vector fields  $X, Y \in \mathfrak{X}(M)$ .

We close this section describing the main analytical tool which is used along the proofs of our results in the next sections. In [37] Yau, generalizing a previous result due to Gaffney [17], established the following version of Stokes' Theorem on an *n*-dimensional, complete noncompact Riemannian manifold  $M^n$ : if  $\omega \in \Omega^{n-1}(M)$  is an integrable (n-1)-differential form on  $M^n$ , then there exists a sequence  $B_i$  of domains on  $M^n$  such that  $B_i \subset B_{i+1}$ ,  $M^n = \bigcup_{i>1} B_i$  and

$$\lim_{i \to +\infty} \int_{B_i} d\omega = 0.$$

Suppose that  $M^n$  is oriented by the volume element dM. If  $\omega = \iota_X dM$  is the contraction of dM in the direction of a smooth vector field X on  $M^n$ , then Caminha obtained a suitable consequence of Yau's result, which can be regarded as an extension of Hopf's maximum principle for complete Riemannian manifolds (cf. Proposition 2.1 of [9]). In what follows,  $\mathcal{L}^1(M)$  and div denote the space of Lebesgue integrable functions and the divergence on  $M^n$ , respectively.

**Lemma 1** Let X be a smooth vector field on the n-dimensional complete noncompact oriented Riemannian manifold  $M^n$ , such that divX does not change sign on  $M^n$ . If  $|X| \in \mathcal{L}^1(M)$ , then divX = 0.

*Remark 1* Lemma 1 can also be seen as a consequence of the version of Stokes' Theorem given by Karp in [20]. In fact, using Theorem in [20], condition  $|X| \in \mathcal{L}^1(M)$  can be weakened to the following technical condition:

$$\liminf_{r \to +\infty} \frac{1}{r} \int_{B(2r) \setminus B(r)} |X| dM = 0,$$

where B(r) denotes the geodesic ball of radius r center at some fixed origin  $o \in M^n$ . See also Corollary 1 and Remark in [20] for some another geometric conditions guaranteing this fact.

*Remark 2* Reasoning in a similar way of that in the beginning of Section 4 of [5] (see also Section 4 in [3]), it is not difficult to verify that there exist no *n*-dimensional compact (without boundary) spacelike submanifolds immersed in  $\mathbb{H}_p^{n+p}$ . Motivated by this fact, along this paper we will deal with complete spacelike submanifolds.

## **3** Complete maximal submanifolds immersed in $\mathbb{H}_{q}^{n+p}$

Let  $a \in \mathbb{R}_{q+1}^{n+p+1}$  be a fixed arbitrary vector and put

$$a = a^{\top} + a^N - \langle a, x \rangle x, \qquad (3.1)$$

where  $a^{\top} \in \mathfrak{X}(M)$  and  $a^N \in \mathfrak{X}^{\perp}(M)$  denote, respectively, the tangential and normal components of *a* with respect to  $M^n \hookrightarrow \mathbb{H}_q^{n+p}$ . By taking covariant derivative in (3.1) and using (2.2) and (2.3), we get for all tangent vector field  $X \in \mathfrak{X}(M)$  that

$$\nabla_X a^\top = A_{a^N} X + \langle a, x \rangle X \tag{3.2}$$

and

$$\nabla_X^{\perp} a^N = -\alpha(a^{\top}, X). \tag{3.3}$$

Hence, from (2.1) and (3.2) we obtain

$$\operatorname{div}(a^{\top}) = \operatorname{tr}(A_{a^{N}}) + n\langle a, x \rangle = n\langle a, \mathbf{H} \rangle + n\langle a, x \rangle.$$
(3.4)

Moreover, we also have that

$$\begin{aligned} \operatorname{tr}(\nabla_{a^{\top}} A_{\xi}) &= \sum_{i} \langle \nabla_{a^{\top}} A_{\xi} e_{i}, e_{i} \rangle - \sum_{i} \langle \nabla_{a^{\top}} e_{i}, A_{\xi} e_{i} \rangle \\ &+ n \langle \nabla_{a^{\top}}^{\perp} \mathbf{H}, \xi \rangle - \sum_{i} a^{\top} \langle A_{\xi} e_{i}, e_{i} \rangle. \end{aligned}$$

So, considering a local orthonormal frame  $\{e_1, \ldots, e_n\}$  on  $M^n$  such that  $A_{\xi}e_i = \lambda_i^{\xi}e_i$ , with a straightforward computation we can verify that

$$\operatorname{tr}(\nabla_{a^{\top}} A_{\xi}) = n \langle \nabla_{a^{\top}}^{\perp} \mathbf{H}, \xi \rangle.$$
(3.5)

From Codazzi Eq. (2.5) jointly with (3.2) and (3.5) we obtain, for all  $\xi \in \mathfrak{X}^{\perp}(M)$ ,

$$\operatorname{div}(A_{\xi}a^{\top}) = n \langle \nabla_{a^{\top}}^{\perp} \mathbf{H}, \xi \rangle + \operatorname{tr}(A_{a^{N}} \circ A_{\xi}) + \langle a, x \rangle \operatorname{tr}(A_{\xi}) + \sum_{i} \langle \alpha(a^{\top}, e_{i}), \nabla_{e_{i}}^{\perp} \xi \rangle.$$
(3.6)

On the other hand, taking the trace in Gauss Eq. (2.4), we have

$$\operatorname{Ric}(X,Y) = -(n-1)\langle X,Y\rangle + n\langle \alpha(X,Y),\mathbf{H}\rangle - \sum_{i} \langle \alpha(X,e_{i}),\alpha(Y,e_{i})\rangle, \quad (3.7)$$

where Ric denotes the Ricci curvature of  $M^n$ . Considering  $X = Y = a^{\top}$  in (3.7), we obtain

$$\operatorname{Ric}(a^{\top}, a^{\top}) = -(n-1)|a^{\top}|^{2} + n\langle \alpha(a^{\top}, a^{\top}), \mathbf{H} \rangle -\sum_{i} \langle \alpha(a^{\top}, e_{i}), \alpha(a^{\top}, e_{i}) \rangle.$$
(3.8)

Furthermore, from (3.3) and (3.6) we get

$$\operatorname{div}(A_{a^{N}}a^{\top}) = n \langle \nabla_{a^{\top}}^{\perp} \mathbf{H}, a^{N} \rangle + \operatorname{tr}(A_{a^{N}}^{2}) + \langle a, x \rangle \operatorname{tr}(A_{a^{N}}) - \sum_{i} \langle \alpha(a^{\top}, e_{i}), \alpha(a^{\top}, e_{i}) \rangle.$$
(3.9)

Hence, from (3.8) and (3.9) we conclude that

$$\operatorname{div}(A_{a^{N}}a^{\top}) = n \langle \nabla_{a^{\top}}^{\perp} \mathbf{H}, a^{N} \rangle + \operatorname{tr}(A_{a^{N}}^{2}) + \langle a, x \rangle \operatorname{tr}(A_{a^{N}}) + \operatorname{Ric}(a^{\top}, a^{\top}) + (n-1) |a^{\top}|^{2} - n \langle \alpha(a^{\top}, a^{\top}), \mathbf{H} \rangle.$$
(3.10)

Based on the previous computations, we obtain the following Bernstein type result concerning maximal submanifolds immersed in  $\mathbb{H}_{a}^{n+p}$ 

**Theorem 1** Let  $M^n$  be a complete maximal spacelike submanifold immersed in  $\mathbb{H}_q^{n+p}$ , with  $1 \leq q \leq p$ . Suppose that  $Ric \geq -(n-1)$  on  $M^n$ . If there exist p vectors  $a_1, \ldots, a_p \in \mathbb{R}_{q+1}^{n+p+1}$  such that  $a_1^N, \ldots, a_p^N$  are linearly independent, with  $A_{a_i^N}$  bounded on  $M^n$  and  $|a_i^\top| \in \mathcal{L}^1(M)$  for each  $1 \leq i \leq p$ , then  $M^n$  is totally geodesic.

*Proof* Let us consider  $a = a_i$  for some  $i \in \{1, ..., p\}$ . Provided that  $\mathbf{H} = 0$ , from (2.1) we see that Eq. (3.10) can be rewritten as follows

$$\operatorname{div}(A_{a^{N}}a^{\top}) = \operatorname{tr}(A_{a^{N}}^{2}) + \operatorname{Ric}(a^{\top}, a^{\top}) + (n-1)|a^{\top}|^{2}.$$
 (3.11)

Thus, since  $\operatorname{Ric}(a^{\top}, a^{\top}) \ge -(n-1)|a^{\top}|^2$ , from (3.11) we obtain that

$$\operatorname{div}(A_{a^N}a^\top) \ge 0. \tag{3.12}$$

Moreover, whereas  $A_{a^N}$  is bounded on  $M^n$ ,  $|A_{a^N}| \le C_1$ , for some constant  $C_1 > 0$ . Thus, as we are assuming that  $|a^\top| \in \mathcal{L}^1(M)$ , we have

$$|A_{a^{N}}a^{\top}| \le |A_{a^{N}}||a^{\top}| \le C_{1}|a^{\top}| \in \mathcal{L}^{1}(M).$$
(3.13)

Hence, taking into account (3.12) and (3.13), we can apply Lemma 1 to guarantee that  $\operatorname{div}(A_{a^N}a^{\top}) = 0$ . Consequently, returning to Eq. (3.11), we conclude that  $A_{a^N} \equiv 0$ . Therefore, since  $\alpha(X, Y) = \sum_{i=1}^{p} \langle A_{a_i^N} X, Y \rangle a_i^N$ , we have that  $M^n$  must be totally geodesic.  $\Box$ 

*Remark 3* Despite our assumption on the  $a_i^N$  in Theorem 1 to be a technical hypothesis, it is motivated by the fact that it occurs in a natural way in the context of spacelike hypersurfaces (see Corollary 2). In this sense, it is a mild hypothesis.

In the case that p = q, being  $M^n$  a maximal submanifold of  $\mathbb{H}_p^{n+p}$ , a classical result due to Ishihara [19] assures us that  $|A|^2 \leq np$  (see also Cheng [13] for the case n = 2). Moreover, each maximal submanifold in  $\mathbb{H}_p^{n+p}$  meets the condition Ric  $\geq -(n-1)$ . Thus, as a consequence of Theorem 1 we obtain

**Corollary 1** Let  $M^n$  be a complete maximal spacelike submanifold immersed in  $\mathbb{H}_p^{n+p}$ . If there exist p vectors  $a_1, \ldots, a_p \in \mathbb{R}_{p+1}^{n+p+1}$  such that  $a_1^N, \ldots, a_p^N$  are linearly independent, with  $|a_i^\top| \in \mathcal{L}^1(M)$  for each  $1 \le i \le p$ , then  $M^n$  is totally geodesic.

Taking into account that the warped product model  $-(-\frac{\pi}{2}, \frac{\pi}{2}) \times_{\cos t} \mathbb{H}^n$ , which is considered in [7] in order to prove their results, models just only an open subset of  $\mathbb{H}_1^{n+1}$  (cf. Example 4.3 of [27]), from Corollary 1 we obtain the following improvement of Theorem 1.2 of [7]

**Corollary 2** Let  $M^n$  be a complete maximal spacelike hypersurface immersed in  $\mathbb{H}_1^{n+1}$ . If there exists a vector  $a \in \mathbb{R}_2^{n+2}$  such that  $a^N$  does not vanish on  $M^n$  and  $|a^\top| \in \mathcal{L}^1(M)$ , then  $M^n$  is a totally geodesic hyperbolic space.

# 4 Submanifolds with parallel mean curvature vector in $\mathbb{H}_p^{n+p}$

In this section, we study the rigidity of a complete spacelike submanifold  $M^n$  of  $\mathbb{H}_p^{n+p}$  with nonzero parallel mean curvature vector **H**. For this, fixed a nonzero vector  $a \in \mathbb{R}_{p+1}^{n+p+1}$ , we observe that Eq. (3.4) gives us

$$\operatorname{div}\left(\langle a, \mathbf{H} \rangle a^{\top}\right) = \frac{1}{n} \operatorname{tr}(A_{a^{N}})^{2} + \langle a, x \rangle \operatorname{tr}(A_{a^{N}}) - \langle \alpha(a^{\top}, a^{\top}), \mathbf{H} \rangle + \langle a, \nabla_{a^{\top}}^{\perp} \mathbf{H} \rangle.$$
(4.1)

Thus, from (3.10) jointly with (4.1) we obtain

$$\operatorname{div}\left[ (A_{a^{N}} - \langle a, \mathbf{H} \rangle I) a^{\top} \right] = (n-1) \langle \nabla_{a^{\top}}^{\perp} \mathbf{H}, a^{N} \rangle + \operatorname{tr}(A_{a^{N}}^{2}) - \frac{1}{n} \operatorname{tr}(A_{a^{N}})^{2} + T(a^{\top}, a^{\top}),$$
(4.2)

where *I* denotes the identity operator in the algebra of smooth vector fields on  $M^n$  and, following the terminology established in [4], *T* stands for a covariant tensor on  $M^n$  which is given by

$$T(X, X) = \operatorname{Ric}(X, X) + (n-1) |X|^{2} - (n-1) \langle \alpha(X, X), \mathbf{H} \rangle.$$
(4.3)

According to [4,11], a spacelike submanifold  $M^n$  of  $\mathbb{H}_p^{n+p}$  with nonzero mean curvature vector **H** is said *pseudo-umbilical* if **H** is an umbilical direction. From (2.6) we see that a totally umbilical spacelike submanifold is always pseudo-umbilical. Conversely, we get

**Proposition 1** Let  $M^n$  be a complete pseudo-umbilical spacelike submanifold with nonzero parallel mean curvature vector  $\mathbf{H}$  in  $\mathbb{H}_p^{n+p}$ . If there exist p vectors  $a_1, \ldots, a_p \in \mathbb{R}_{p+1}^{n+p+1}$  such that  $a_1^N, \ldots, a_p^N$  are linearly independent, with  $\langle a_i, \mathbf{H} \rangle$  and  $A_{a_i^N}$  bounded on  $M^n$  and  $|a_i^\top| \in \mathcal{L}^1(M)$  for each  $1 \leq i \leq p$ , then  $M^n$  is totally umbilical.

*Proof* Let us consider  $a = a_i$  for some  $i \in \{1, ..., p\}$ . We have that

$$\left| (A_{a^N} - \langle a, \mathbf{H} \rangle I) a^\top \right| \le \left( |A_{a^N}| + |\langle a, \mathbf{H} \rangle| \right) |a^\top| \le C_2 |a^\top| \in \mathcal{L}^1(M).$$
(4.4)

Since we are assuming that  $M^n$  is pseudo-umbilical of  $\mathbb{H}_p^{n+p}$ , Lemma 4.1 of [4] assures that

$$\operatorname{Ric}(X, X) \ge -(n-1)|X|^2 + (n-1)\langle \alpha(X, X), \mathbf{H} \rangle, \tag{4.5}$$

for all  $X \in \mathfrak{X}(M)$ . Thus, from (4.3) and (4.5) we get that  $T(a^{\top}, a^{\top}) \ge 0$ . Moreover, we observe that the function  $u = \operatorname{tr}(A_{a^N}^2) - \frac{1}{n}\operatorname{tr}(A_{a^N})^2$  is always nonnegative with u = 0 if, and only if,  $a^N$  is a umbilical direction. From (4.2), we obtain

$$\operatorname{div}\left[ (A_{a^N} - \langle a, \mathbf{H} \rangle I) a^\top \right] \ge 0.$$
(4.6)

Thus, from (4.4) and (4.6), Lemma 1 assure us

$$\operatorname{tr}(A_{a^N}^2) - \frac{1}{n} \operatorname{tr}(A_{a^N})^2 + T(a^{\top}, a^{\top}) = 0.$$

Then,  $\operatorname{tr}(A_{a^N}^2) - \frac{1}{n}\operatorname{tr}(A_{a^N})^2 = 0$  and, hence,  $a^N$  is a umbilical direction of  $M^n$ . Therefore, since we are supposing the existence of such vectors  $a_1, \ldots, a_p \in \mathbb{R}_{p+1}^{n+p+1}$  whose normal projections  $a_1^N, \ldots, a_p^N$  with respect to  $M^n$  are linearly independent, we conclude that (2.6) holds, that is,  $M^n$  must be totally umbilical.

Proceeding, we establish sufficient conditions to guarantee that a spacelike submanifold immersed in  $\mathbb{H}_p^{n+p}$  with nonzero parallel mean curvature vector must be pseudo-umbilical.

**Theorem 2** Let  $M^n$  be a complete spacelike submanifold immersed in  $\mathbb{H}_p^{n+p}$  with nonzero parallel mean curvature vector  $\mathbf{H}$  and bounded normalized scalar curvature R. If there exists a nonzero vector  $a \in \mathbb{R}_{p+1}^{n+p+1}$  such that  $a^N$  is timelike, i.e.,  $\langle a^N, a^N \rangle < 0$ , collinear to  $\mathbf{H}$  and  $|a^{\top}| \in \mathcal{L}^1(M)$ , then  $M^n$  is pseudo-umbilical.

*Proof* Initially, taking a local orthonormal frame  $\{e_1, \ldots, e_n\}$  on  $M^n$ , from (3.7) we get that the squared norm of second form fundamental  $\alpha$  of  $M^n$  satisfies

$$|\alpha|^{2} = \sum_{i,j} |\alpha(e_{i}, e_{j})|^{2} = n^{2} \langle \mathbf{H}, \mathbf{H} \rangle + n(n-1)(R+1).$$
(4.7)

Now, let us consider a nonzero vector  $a \in \mathbb{R}_{p+1}^{n+p+1}$  such that  $a^N$  is timelike, collinear to **H** and with  $|a^\top| \in \mathcal{L}^1(M)$ . Since  $M^n$  has bounded normalized scalar curvature and nonzero parallel mean curvature vector **H**, from (4.7) we conclude that  $|\alpha|^2$  is bounded on  $M^n$ . So, taking  $\xi = \mathbf{H}$  in (3.6) we get

$$\operatorname{div}(A_{\mathbf{H}}a^{\top}) = \operatorname{tr}(A_{a^{N}} \circ A_{\mathbf{H}}) + \langle a, x \rangle \operatorname{tr}(A_{\mathbf{H}}), \tag{4.8}$$

where  $A_{\mathbf{H}}$  denotes the Weingarten operator associated to  $\mathbf{H}$ .

On the other hand, from (3.4) we have

$$\langle a, x \rangle = \frac{1}{n} \operatorname{div}(a^{\top}) - \langle a, \mathbf{H} \rangle.$$
 (4.9)

Consequently, from (4.8) and (4.9)

$$\operatorname{div}(A_{\mathbf{H}}a^{\top}) = \operatorname{tr}(A_{a^{N}} \circ A_{\mathbf{H}}) + \operatorname{tr}(A_{\mathbf{H}})\frac{1}{n}\operatorname{div}(a^{\top}) - \frac{1}{n}\operatorname{tr}(A_{a^{N}})\operatorname{tr}(A_{\mathbf{H}}).$$
(4.10)

Since

$$\operatorname{div}\left(\operatorname{tr}(A_{\mathbf{H}})a^{\top}\right) = \operatorname{tr}(A_{\mathbf{H}})\operatorname{div}(a^{\top}), \qquad (4.11)$$

from (4.10) and (4.11) we obtain

$$\operatorname{div} V = \operatorname{tr}(A_{a^N} \circ A_{\mathbf{H}}) - \frac{1}{n} \operatorname{tr}(A_{a^N}) \operatorname{tr}(A_{\mathbf{H}}), \qquad (4.12)$$

where V is a tangent vector field on  $M^n$  given by

$$V = \left(A_{\mathbf{H}} - \frac{1}{n} \operatorname{tr}(A_{\mathbf{H}})I\right) a^{\top}.$$

We note that, since we are supposing  $a^N$  timelike and collinear to **H**, there exists on  $M^n$  a smooth function  $\lambda$  having strict sign such that  $a^N = \lambda \mathbf{H}$ . Thus, from (2.3) and (4.12) we get

div 
$$V = \lambda \left( \operatorname{tr}(A_{\mathbf{H}}^2) - \frac{1}{n} \operatorname{tr}(A_{\mathbf{H}})^2 \right).$$
 (4.13)

Consequently, from (4.13) we conclude that div V does not change sign on  $M^n$ . Moreover, we also have that

$$|V| \le (|A_{\mathbf{H}}| + |\langle \mathbf{H}, \mathbf{H} \rangle|) |a^{\top}| \in \mathcal{L}^{1}(M).$$

Hence, we can apply once more Lemma 1 to assure that div V = 0 on  $M^n$ .

Therefore, returning to (4.13) we obtain that

$$\lambda \left( \operatorname{tr}(A_{\mathbf{H}}^2) - \frac{1}{n} \operatorname{tr}(A_{\mathbf{H}})^2 \right) = 0$$

which implies that H is an umbilical direction.

We observe that, in the case p = 1, the notion of pseudo-umbilical coincides with that of totally umbilical. Moreover, we note that the hypothesis that  $a^N$  is timelike amounts to the support function  $f_a = \langle a, v \rangle$  having strict sign on the spacelike hypersurface  $M^n \hookrightarrow \mathbb{H}_1^{n+1}$ , where v stands for the Gauss mapping of  $M^n$ . Consequently, taking into account the classification of the totally umbilical hypersurfaces of  $\mathbb{H}_1^{n+1}$  (see, for instance, Example 1 of [24]) and that Theorem 1 of [21] assures us that a complete constant mean curvature spacelike hypersurface of  $\mathbb{H}_1^{n+1}$  must have bounded second fundamental form (or, equivalently, bounded normalized scalar curvature), from Theorem 2 we obtain the following

**Corollary 3** Let  $M^n$  be a complete spacelike hypersurface immersed in  $\mathbb{H}_1^{n+1}$  with nonzero constant mean curvature. If there exists a nonzero vector  $a \in \mathbb{R}_2^{n+2}$  such that the support function  $f_a$  has strict sign on  $M^n$  and  $|a^\top| \in \mathcal{L}^1(M)$ , then  $M^n$  is a totally umbilical hyperbolic space.

Acknowledgements The first author is partially supported by CNPq, Brazil, Grant 303977/2015-9. The second author was partially supported by PNPD/UFCG/CAPES, Brazil. The third author is partially supported by CNPq, Brazil, grant 308757/2015-7. The authors would like to thank the referee for giving some valuable suggestions and comments which improved the paper.

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