#### ORIGINAL PAPER



# Study of sunspot cycles using fractal dimensions: wave-spectrum scaling

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#### Abstract

The fractal dimension analysis provides more appropriate evolution of several solar phenomena related to the sun and its environment. The novelty of this research is to use self-similar fractal dimension  $(FD<sub>S</sub>)$  and self-affine fractal dimension (FD<sub>A</sub>) to calculate fractal parameters including universal parameter such as the exponent scale  $\beta$ , spectral exponent ( $\alpha$ ), and fractal autocorrelation coefficient (C∇). First, the mean monthly data of each sunspot cycle from 1755 to 2008 (23 cycles) is analyzed separately. Then, the total data of 24 cycles is analyzed. The study focuses on finding an adequate value of the wavespectral exponent  $\alpha$  for which the cycles are more strongly correlated with each other. Self-similar fractal dimension is found to be more persistent and positively correlated as compared to self-affine fractal dimension. The fractal parameters are found to exist on a significant scale. The exponent scale  $\beta$  is calculated by both of the fractal dimensions FD<sub>S</sub> and FD<sub>A</sub>. Both the fractal dimensions are also related to the wave-spectral exponent  $\alpha$  which is calculated by the Hurst exponent (HE). The self-similar and self-affine spectral exponents  $\alpha_S$  and  $\alpha_A$  are used to determine whether the value of  $\alpha$  is greater than 2 or not. The spectrum for sunspot cycles is considered to be Gaussian if the value of  $\alpha$  is greater than 2. This demonstrates that the cycles are strongly correlated to other cycles. The self-similar fractal autocorrelation coefficient (C∇) is found to be more persistent and correlated as compared to the self-affine fractal dimension. It can be concluded that the fractal approach can study more rigorously the local and global aspects of the dynamical processes and activities associated with the sun and its climate.

Keywords Spectral analysis · Rescaled range analysis · Spectral exponent · Self-similar

## Introduction

A sunspot has an intense magnetic field and cool temperature, and the center temperature of a sunspot is 6400 °F. Galileo made drawings of sunspots in 1612 (TJO News [2006](#page-5-0)). Sunspots are associated with active regions, which are areas of locally increased magnetic flux of the sun (Kevin et al. [2014\)](#page-5-0). A maximum number of sunspots are known as solar maximum. Sunspots often appear near 30–35° north and south

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 $\boxtimes$  Asma Zaffar [asmazaffar105@gmail.com](mailto:asmazaffar105@gmail.com) of the sun's hemisphere with higher latitudes. The solar activity or sunspot is generated due to geomagnetic disturbances. Two types of magnetic lines are associated with sunspots. One is open whereas, the other is closed. Open lines extend out to long distances from the sun. The closed lines form loops and return back to the sun (Enfield et al. [1991\)](#page-5-0).

Sunspots consist of cycles, and each cycle has a different duration. The average duration of sunspot cycles is slightly greater than 11 years. In each cycle, the number of sunspots varies from a maximum to minimum and again back to maximum. Cycle 1 consists of 11.3 years, cycle 2 has 9 years, cycle 3 (9.3 years), cycle 4 (13.7 years), cycle 5 (12.6 years), cycle 6 (12.4 years), cycle 7 (10.5 years), cycle 8 (9.8 years), cycle 9 (12.4 years), cycle 10 (11.3 years), cycle 11 (11.8 years), cycle 12 (11.3 years), cycle 13 (11.9 years), cycle 14 (11.5 years), cycle 15 (10 years), cycle 16 (10.1 years), cycle 17 (10.4 years), cycle 18 (10.2 years), cycle 19 (10.5 years), cycle 20 (11.7 years), cycle 21 (10.3 years), cycle 22 (9.7 years), cycle 23 (11.7 years) and cycle 24 is proceeding which is started in January 2008 and will last until 2018.

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<span id="page-1-0"></span>As for the fractal dimensions are concerned, two basic types of fractals exist. One of the self-similar and the other is self-affine. In the self-similar type, the geometric object is composed into a union of rescaled copies of itself with uniform in all directions or rescaling isotropic. Whereas in the self-affine type, the geometric object is described as the union of rescaled copies of itself depending on the direction or rescaling anisotropic. The self-similarity and self-affinity both are linear transformations.

A commonly used method to calculate the fractal dimension is the Hausdorff-Besicovich method. Alternative methods such as a box counting method, rescaled range analysis, and Higuchi's method are also frequently used. Box dimension or box counting method is more appropriate than the methods mentioned above (Michael [1988\)](#page-5-0). Fractal dimension describes the roughness and smoothness of the data. In the case of the sun, it is used to determine the correlation between solar cycles (Gayathri and Selvaraj [2010](#page-5-0)). Fractal geometry plays an active role in the study of topography and spatial analysis. The fractal concept plays an essential role in the scaling symmetry. Scaling symmetry is defined as the geometric object size reduced or expanded, whereas the new object is the same as that of the original object characteristics. The fractal analysis is a common approach for self-affine structures. The fractal parameters are evaluated by using three methods that include scaling analysis, spatial correlation analysis, and spectral analysis (Chen [2010\)](#page-5-0). In spectral analysis, the energy spectrum and correlation function are able to convert into another Fourier transform (Chen [2009](#page-5-0)). The relations among different fractal parameters are calculated by using spectral analysis, which is based on correlation functions.

n this paper, we intend to explore the relationship between two methods of fractal dimension: one is the self-similar  $(FD<sub>S</sub>)$ which is calculated by the box counting method, and another is the self-affine  $(FD_A)$  which is calculated by rescaled range analysis of sunspot cycles. The second part consists of the wavespectrum scaling equations for calculating fractal dimensions of sunspot which are presented. The third section pertains to the discussion regarding the fractal geometry parameters of sunspot cycles and the fourth section is a conclusion.

# Mathematical models and fractal dimension relations

In this section, brief information about certain qualities of universal scaling laws is explored.

#### Spatial correlation dimensions

Fractal dimension can be measured with a characteristic scale. Three basic concepts regarding the fractal dimension of sunspots can be described as follows.

The plane of sunspots has Euclidean dimension  $d = 2$ . The smallest unit of the sunspot is considered as a point, so the topological dimension of sunspots is considered to be  $d_t = 0$ . So the fractal dimension of sunspots ranges from  $d_t$  $= 0$  to  $d = 2$ . Fractal analysis of sunspot time series presented here is based upon the existence of correlation among sunspot cycles. Therefore, this study is associates fractal analysis with correlation analysis. In this sense, the generalized fractal dimension is often called the correlation dimension (Chen and Jiang [2010;](#page-5-0) Grassberger and Procaccia, [1983\)](#page-5-0). Fractal dimension can be used to study time series data using the relationship between fractal dimension and Hurst exponent:

$$
FD_{S} = 2 - HE_{S}
$$
 (2.1)

FD equal to 1.5 indicates that the events are unpredictable and no correlation exists between two successive events. The process becomes more and more predictable as the value of the fractal dimension approaches to 1. Fractal dimension ranging from 1 to 1.5 indicates that the process is persistent. A rise in the value of the fractal dimension above 1.5 means that the process is anti-persistent. The value of  $FD = 1.5$  indicates that the data is random, whereas  $FD = 1$  reveals that the time series data is purely deterministic (Shaikh et al. [2008\)](#page-5-0).

The length of each sunspot cycle and  $FD<sub>S</sub>$  is related by the following equation:

$$
\delta(C) = \delta_1 C^{\text{FDs}-d} = \delta_1 C^{-\beta} \tag{2.2}
$$

where  $\delta_1$  is the proportionality coefficient and  $\beta = d - FD_S$ is the scaling exponent. Note that  $FD_s < d$  (Frankhauser [1998\)](#page-5-0). If the value of  $FD<sub>S</sub>$  lies between 1 and 2, then scaling exponents range forms 0 to 1. If  $FD<sub>S</sub> < 1$  or  $FD<sub>S</sub> > 2$ , then the value of  $\beta > 1$  or  $\beta < 0$ , respectively. For the solar cycle, the fractal dimension  $FD<sub>S</sub>$  can be considered as a one-point correlation dimension which indicates a zero-order correlation dimension. A correlation of zero order indicates that no relationship exists between cycles.

#### The wave-spectrum relation of sunspots

This section stresses upon the calculation of spectral exponent and spatial autocorrelation coefficient (spatial scaling). These two play an important role to study the spatial behavior of data. The data under consideration comprises 23 sunspot cycles. To perform spatial scaling, the correlation function associated with data is changed into an energy spectrum using Fourier transform (Chen [2009](#page-5-0)). In addition to other methods, Fourier transform can also be used to study similarity. In this method, relations of fractal parameters are determined by calculating spectral exponents. For this purpose, the following scaling law is used:

<span id="page-2-0"></span>
$$
f\left(\lambda \rho\right) = \lambda^{\beta} f\left(\lambda\right) \tag{2.3}
$$

Where  $\lambda$  denotes the scaling factor,  $\beta$  describes the scaling exponent ( $\beta = d - FD_s$ ), and  $\rho$  is called the length variable of each cycle.

Applying Fourier transform to 2.3, the following scaling relation is obtained:

$$
F(\lambda \gamma) = F[f(\lambda \rho)] = \lambda^{-(1-\beta)} F[f(\rho)]
$$
  
=  $\lambda^{-(1-\beta)} F(\gamma)$  (2.4)

where F is known as the Fourier operator and  $\gamma$  denotes the wave number, whereas  $F(\gamma)$  is the image function of  $f(\lambda)$ . Equation 2.4 finally gives the following wave-spectrum relation:

$$
S(\gamma) \propto \gamma^{-2(1-\beta)} \tag{2.5}
$$

Relation 2.5 provides a numerical relation between fractal dimension and the spectral exponents by taking  $\beta = d - FD_S$ (Eq. [2.2](#page-1-0)), thus:

$$
S(\gamma) \propto \gamma^{-2(1-d + FDs)} = \gamma^{-2(FDs - 1)} = \gamma^{-\alpha}
$$
 (2.6)

Here,

$$
\alpha = 2 \text{ (FDs-1)}\tag{2.7}
$$

where  $\alpha$  is the spectral exponent and considered to be a constant. When the range of fractal dimension is  $1 < FD < 2$ , then the spectral exponent correlation dimension is said to be a point-point correlation dimension. This means that there exists a spatial correlation between two adjacent points of each cycle. Spatial correlation describes the correlation between a spot's spatial direction and the average receiving spot again. Equation 2.7 gives the required relation between FD and  $\alpha$ . FD and  $\alpha$  possess one-point correlation dimension and the point-point correlation dimension, respectively. The onepoint correlation shows the spatial correlation between the given spot and other spots of the cycle. The relation between the fractal dimension (FD) and spectral exponent  $(\alpha)$  was introduced by Higuchi [\(1988\)](#page-5-0) and is described by the following relation:

$$
FD = (5-\alpha)/2 \tag{2.8}
$$

The relation among fractal dimension (FD), spectral exponent  $\alpha$ , and Hurst exponent HE is given by Burlaga (Burlaga and Klein [1986](#page-5-0); Turcotte [1992\)](#page-5-0) as given by Eq. 2.9.

$$
\alpha = 5 - 2FD = 2HE + 1\tag{2.9}
$$

 $\alpha = 0$  describes a white noise-like system. It means that the system is uncorrelated and the power spectrum is independent of the frequency.  $\alpha = 1$  is known as flicker or 1/f noise system which indicates a moderate correlation.  $\alpha = 2$  is called a Brownian noise-like system, which shows a strong correlation. In general, the fractal dimension FD lies in the range (0  $\langle$  FD  $\langle$  2), but the fractal dimension in the range (1  $\langle$  FD  $\langle$  2) indicates that the state under consideration is highly random and irregular. In such case, spectral exponents range as  $(1 < \alpha$ < 3) (Mandelbort and Van Ness [1968\)](#page-5-0).

The parameters to determine the fractal dimension can be calculated by using two methods: self-similarity and selfaffinity; consequently, two different fractal dimensions are obtained viz. Self-similar fractal dimension  $(FD<sub>S</sub>)$  and Selfaffine fractal dimension  $(FD_A)$ . Here, the wave-spectrum scaling is performed by calculating both the fractal dimensions, viz.  $FD<sub>S</sub>$  and  $FD<sub>A</sub>$ . Earlier, this sort of analysis was performed by Liu and Liu [\(1992\)](#page-5-0) and Mandelbrot [\(1999\)](#page-5-0) using  $FD_A$ . The Hurst exponent  $(HE_A)$  can be calculated by the method of rescaled range analysis (Hurst et al. [1965](#page-5-0)); H is described by the power function  $R(\tau)/S(\tau)=(\tau/2)^H$  (Feder [1988\)](#page-5-0). The Hurst exponent  $(HE<sub>S</sub>)$  can be calculated by the method of box counting technique. The relationship between  $FD_s$  and  $FD_A$ can be described from Eq, 2.7 and 2.9 giving 2.10:

$$
FD_{\rm S} = \frac{7}{2} - FD_{\rm A} \tag{2.10}
$$

The situation can be expressed logically by the following flowchart (Figure [1\)](#page-3-0).

The spatial activity of sunspots is supposed to be expressed as a fractal Brownian motion  $(fBm)$ ; thus, the value of FD of each cycle of sunspots lies between 1 and 2.

The relation between spatial autocorrelation coefficient ( $CV$ ) and HE based on  $fBm$  can be given as follows:

$$
CV = 2^{2HE-1} - 1
$$
\n(2.11)

where  $C∇$  represents the spatial autocorrelation coefficient, which is based on the multiple-lag 1-dimension spatial autocorrelation. Spatial autocorrelation is used to measure the correlation of spots with itself through the active region. It should be noted that the value of HE =  $\frac{1}{2}$  means that CV = 0 which is considered to be the Brownian motion. If HE  $> ½$ , then C $\nabla > 0$ representing that positive spatial autocorrelation exists. HE <  $\frac{1}{2}$  implies that CV < 0 which indicates that the spatial autocorrelation is negative.

The numerical relationship between different fractal parameters like FD<sub>S</sub>, FD<sub>A</sub>, HE<sub>S</sub>, HE<sub>A</sub>,  $\alpha$ <sub>S</sub>,  $\alpha$ <sub>A</sub>, C<sub>A</sub> $\nabla$ , and C<sub>S</sub> $\nabla$  can be computed by using Eqs. 2.9, 2.10, and 2.11. The results described in Table [1](#page-3-0) shows that all values are within a significant scale.  $FD_S$  ranging from 1 to 1.5 shows that the number of spots is correlated to each other and has linear behavior. The Hurst exponents (HE) ranging from 0.5 to 1 indicate the persistence of sunspot cycles. Relation 2.10 is theoretically valid when the fractal dimension  $FD<sub>S</sub>$  values range from 1.5 to 2.

<span id="page-3-0"></span>Fig. 1 A sketch map among the relationship of different fractal parameters



# Results and discussions

This study attempts to evaluate the fractal dimensions of sunspot data using scaling analysis and spectral analysis. Also, the cycle-wise spatial data correlation is determined. It is found that the correlation function is 1-dimensional (linear relation). Fractal parameters are calculated in a significant range using self-similar fractal dimension  $(FD<sub>S</sub>)$  and selfaffine fractal dimension (FDA). Table 1 indicates that the values of  $FD<sub>S</sub>$  and  $FD<sub>A</sub>$  in each cycle exist in a range from 1 to 1.5 which indicate that cycle is persistent, correlated, and predictable. The value of  $FD_A$  was found to be greater than FDs in all cycles. Similarly, both self-similar Hurst exponent  $HE<sub>S</sub>$  and self-affine Hurst exponent  $HE<sub>A</sub>$  values range from

Table 1 The numerical relationship between different fractal dimensions, Hurst exponent, autocorrelation coefficient and spectral exponents can be express as

Cycles	Duration	FD <sub>S</sub>	$FD_A$	HE <sub>S</sub>	$HE_{A}$	$\alpha_{\rm S}$	$\alpha_{A}$	$C_S \nabla$	$C_A \nabla$
1	Aug 1755 - Mar 1766	1.181	1.367	0.819	0.633	2.638	2.266	0.556	0.202
$\overline{c}$	Mar 1766 - Aug 1775	1.27	1.407	0.73	0.593	2.246	2.186	0.376	0.138
3	Aug 1775 - Jun 1784	1.327	1.587	0.672	0.413	2.344	1.826	0.269	$-0.114$
4	Jun 1784 - Jun 1798	1.318	1.446	0.682	0.554	2.364	2.108	0.287	0.078
5	Jun 1798 - Sep 1810	1.099	1.380	0.901	0.620	2.802	2.240	0.744	0.181
6	Sep 1810 - Dec 1823	1.136	1.363	0.864	0.637	2.728	2.274	0.656	0.209
7	Dec 1823 - Oct 1833	1.177	1.163	0.823	0.837	2.646	2.674	0.565	0.595
8	Oct 1833 - Sep 1843	1.296	1.223	0.704	0.772	2.408	2.554	0.327	0.458
9	Sep 1843 - Mar 1855	1.275	1.470	0.725	0.530	2.45	2.060	0.366	0.042
10	Mar 1855 - Feb 1867	1.267	1.313	0.733	0.687	2.466	2.374	0.381	0.296
11	Feb 1867 - Sep 1878	1.298	1.347	0.702	0.653	2.404	2.306	0.323	0.236
12	Sep 1878 - Jun 1890	1.165	1.463	0.835	0.537	2.67	2.074	0.591	0.053
13	Jun 1890 - Sep 1902	1.252	1.376	0.748	0.624	2.496	2.248	0.410	0.188
14	Sep 1902 - Dec 1913	1.141	1.400	0.859	0.600	2.718	2.200	0.645	0.149
15	Dec 1913 - May 1923	1.235	1.479	0.765	0.521	2.53	2.042	0.443	0.030
16	May 1923 - Sep 1933	1.149	1.391	0.851	0.604	2.702	2.218	0.627	0.155
17	Sep 1933 - Jan 1944	1.206	1.526	0.794	0.474	2.588	1.948	0.503	$-0.035$
18	Jan 1944 - Feb 1954	1.194	1.373	0.806	0.627	2.612	2.254	0.528	0.193
19	Feb 1954 - Oct 1964	1.266	1.410	0.734	0.590	2.468	2.180	0.383	0.133
20	Oct 1964 - May 1976	1.193	1.355	0.807	0.645	2.614	2.290	0.530	0.223
21	May 1976 - Mar 1986	1.198	1.187	0.802	0.813	2.604	2.626	0.520	0.543
22	Mar 1986 - Jun1996	1.225	1.436	0.775	0.563	2.55	2.128	0.464	0.091
23	Jun 1996 - Jan 2008	1.213	1.451	0.787	0.549	2.574	2.098	0.489	0.070
24	Aug 1755 - Jan 2008	1.002	1.354	0.998	0.649	2.996	2.110	0.994	0.229

Fig. 2 Comparative studies of wave-spectrum scaling parameter of self-similar and self-affine of each sunspot cycles



0.5 to 1 which reveals that each cycle is persistent.  $\text{HE}_\text{S}$  are found to be greater than  $HE_A$ . Since the fractal dimension lies between 1 and 2, the scaling exponent  $\beta$  ranges from 0 to 1. This range for  $\beta$  confirms its validity. The value of  $\beta$  shows that the data has a 1-point correlation dimension and the dynamical behavior of the data is linear. Table [1](#page-3-0) describes the numerical relation between the spectral exponent  $(\alpha)$  and autocorrelation coefficient (C∇) which are calculated by using both the self-similar fractal dimension and the self-affine fractal dimension. If  $1 <$  DF  $<$  2, then the spectral exponent  $\alpha$  has a range from 1 to 3.  $\alpha = 0$  indicates a white noise-like system which describes uncorrelated behavior.  $\alpha = 1$  is called a flicker which represents a moderately correlated behavior.  $\alpha = 2$  is called a Brownian noise-like system and shows a strong correlation. Self-similar spectral exponent  $(\alpha_s)$  and self-affine spectral exponent  $(\alpha_A)$  are calculated by using relation [2.8.](#page-2-0) Values of  $\alpha_s$  and  $\alpha_A$  reveal that the cycles behave like a Brownian noise. If  $1 <$  FD<sub>S</sub>  $<$  2, then the spectral exponent  $\alpha$  describes a point-point correlation. It is to mention that there exists a relationship between one-point correlation dimension  $FD<sub>S</sub>$  and the point-point correlation dimension ( $\alpha$ ). The autocorrelation coefficient (C∇) describes the multiple-lag 1-dimension spatial autocorrelation of sunspot cycles. The autocorrelation coefficient (C∇) is calculated by relation [2.11.](#page-2-0) C∇  $= 0$  if HE =  $\frac{1}{2}$  which indicates a Brownian motion. If HE >  $\frac{1}{2}$ , then C∇ > 0 which indicates a positive spatial autocorrelation. If HE <  $\frac{1}{2}$ , then CV < 0 which represents the negative spatial autocorrelation. The autocorrelation coefficients  $(C_S \nabla)$  show positive spatial autocorrelation except in case of  $C_A \nabla$  in cycle 3 and cycle 17 where the negative correlation appears. Figure [1](#page-3-0) indicates a sketch map of the relationship of different fractal parameters. Figure 2 represents the combined behavior of each solar cycle in one diagram. Figure 3 describes the compression between the self-similar fractal dimensions and self-affine fractal dimensions in each cycle.

Fig. 3 Compression of selfsimilar and self-affine fractal dimension of sunspots cycles



## <span id="page-5-0"></span>Conclusion

This study investigates the relationship between self-similar fractal dimensions  $(FD<sub>S</sub>)$  and self-affine fractal dimensions  $(FD_A)$ .  $FD_S$  is calculated by using a box counting method, and  $FD_A$  is calculated by rescale range analysis method. Hurst exponents are calculated by self-similar Hurts exponent  $(HE<sub>S</sub>)$  and self-affine Hurts exponent  $(HE<sub>A</sub>)$ . Spectral exponents are calculated using relation [2.8](#page-2-0). Respective values of  $\alpha_{\rm S}$  and  $\alpha_{\rm A}$  are compared using Eq. [2.9](#page-2-0). For the calculation of autocorrelation coefficient, Eq. [2.11](#page-2-0) is used. Table [1](#page-3-0) exhibits the numerical values of Hurst exponents, spectral exponent, and autocorrelation coefficient using self-similar and selfaffine techniques. Sunspot cycles are overall persistent and correlated.  $\text{HE}_\text{S}$  values are greater than  $\text{HE}_\text{A}$ . All values of the spectral exponent ( $\alpha_{\rm S}$  and  $\alpha_{\rm A}$ ) of each sunspot cycles behave like a Brownian noise. The autocorrelation coefficient in both cases lies in a valid range. Self-similar fractal dimensions  $(FD<sub>S</sub>)$  and self-affine fractal dimensions  $(FD<sub>A</sub>)$  obtained by using relation [2.10](#page-2-0) failed to give a well-defined relationship. This is because the method is useful only in case the fractal dimension ranges between 1.5 and 2. Here, the fractal dimensions of sunspot cycles are less than 1.5.

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