# **Regularity criteria for the 3D magneto-micropolar fluid equations via the direction of the velocity**

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**Abstract.** We consider sufficient conditions to ensure the smoothness of solutions to 3D magneto-micropolar fluid equations. It involves only the direction of the velocity and the magnetic field. Our result extends to the cases of Navier–Stokes and MHD equations.

**Keywords.** Magneto-micropolar fluid equations; regularity criteria; direction of velocity.

**2010 Mathematics Subject Classification.** 35Q35, 76W05, 35B65.

#### **1. Introduction and the main result**

In this paper, we consider the 3D magneto-micropolar fluid equations studied by Galdi and Rionero [\[5\]](#page-5-0):

<span id="page-0-0"></span>
$$
\begin{cases}\n\partial_t u + u \cdot \nabla u - (\mu + \chi) \Delta u - b \cdot \nabla b + \nabla (p + b^2) - \chi \nabla \times \omega = 0, \\
\partial_t \omega - \gamma \Delta \omega - \kappa \nabla \text{div } \omega + 2\chi \omega + u \cdot \nabla \omega - \chi \nabla \times u = 0, \\
\partial_t b - \nu \Delta b + u \cdot \nabla b - b \cdot \nabla u = 0, \\
\nabla \cdot u = \nabla \cdot b = 0, \\
u(x, 0) = u_0(x), \omega(0, x) = \omega_0(x), b(0, x) = b_0(x).\n\end{cases} (1.1)
$$

Here  $u = u(x, t)$ ,  $b = b(x, t)$ ,  $\omega = \omega(x, t)$  represent the velocity field, the magnetic field and the micro-rotational velocity respectively; *p* denotes the hydrodynamic pressure;  $\mu > 0$  is the kinematic viscosity,  $\chi > 0$  is the vortex viscosity,  $\kappa > 0$  and  $\gamma > 0$  are the spin viscosities,  $1/\nu$  (with  $\nu > 0$ ) is the magnetic Reynold; while  $u_0, b_0, \omega_0$  are the corresponding initial data with  $\nabla \cdot u_0 = \nabla \cdot b_0 = 0$ .

The global weak solution to system [\(1.1\)](#page-0-0) is established by Rojas-Medar and Boldrini [\[10\]](#page-5-1), while the local strong solutions are given by Rojas-Medar [\[9\]](#page-5-2). However, whether or not the local strong solutions can exist globally is still an open problem. Thus regularity criteria appears. Notice that:

(1) Yuan [\[14\]](#page-6-0) first established the following fundamental regularity criterion in terms of the velocity or its gradient

<span id="page-0-1"></span>
$$
u \in L^{p}(0, T; L^{q}(\mathbb{R}^{3})), \quad \frac{2}{p} + \frac{3}{q} = 1, \quad 3 < q \leq \infty
$$
 (1.2)

and

$$
\nabla u \in L^{p}(0, T; L^{q}(\mathbb{R}^{3})), \quad \frac{2}{p} + \frac{3}{q} = 2, \quad \frac{3}{2} < q \leq \infty.
$$
 (1.3)

Then Gala [\[4\]](#page-5-3) extended it to the Morrey–Campanato spaces, Zhang *et al* [\[16\]](#page-6-1) improved it to some more general Triebel–Lizorkin spaces.

- (2) When  $\omega = b = 0$ , system [\(1.1\)](#page-0-0) is just the classical Naiver–Stokes equations. Serrin [\[11\]](#page-6-2), Prodi [\[8\]](#page-5-4) and Beirão da Veiga [\[1\]](#page-5-5) proved regularity if some scaling-invariant norm of *u* or  $\nabla u$  is bounded.
- (3) When  $\omega = 0$ , system (1,1) is then the 3D MHD equations. He and Xin [\[6\]](#page-5-6), and Zhou [\[19\]](#page-6-3) gave criteria similar to the case for Navier–Stokes equations.
- (4) When  $b = 0$ , system [\(1.1\)](#page-0-0) is reduced to the micropolar fluid equations, Yuan [\[13\]](#page-6-4) gave some criteria in Lorentz spaces.

For later developments, see [\[2,](#page-5-7) [3,](#page-5-8) [15,](#page-6-5) [18\]](#page-6-6) and references cited therein. Recently, Vasseur [\[12\]](#page-6-7) proved that if

<span id="page-1-0"></span>
$$
\operatorname{div} \frac{u}{|u|} \in L^p(0, T; L^q(\mathbb{R}^3)), \quad \frac{2}{p} + \frac{3}{q} \le \frac{1}{2}, \quad p \ge 4, \quad q \ge 6,\tag{1.4}
$$

then the solutions to the Navier–Stokes equations are smooth. Later, Luo [\[7\]](#page-5-9) extended [\(1.4\)](#page-1-0) to the MHD equations, but involves the magnetic field also. We now extend the result of Vasseur  $[12]$  and Luo  $[7]$  to system  $(1.1)$ . The main result is the following:

**Theorem 1.1.** *Let*  $u_0, \omega_0, b_0 \in H^1(\mathbb{R}^3)$  *with*  $\nabla \cdot u_0 = \nabla \cdot b_0 = 0$  *in the sense of distributions. Suppose that*  $(u, \omega, b)$  *is a strong solution to*  $(1.1)$  *in*  $(0, T)$  *such that* 

$$
u, \, \omega, \, b \in C((0, T); H^1(\mathbb{R}^3)) \cap C((0, T); H^2(\mathbb{R}^3))
$$

*and*  $\nabla \cdot u = \nabla \cdot b = 0$ . If additionally,

<span id="page-1-2"></span>
$$
\text{div}\frac{u}{|u|} \in L^p(0, T; L^q(\mathbb{R}^3)), \ \frac{2}{p} + \frac{3}{q} \leq \frac{1}{2}, \ 4 \leq p < \infty, \ \ 6 \leq q \leq \infty \tag{1.5}
$$

*and*

<span id="page-1-1"></span>
$$
b \in L^{r}(0, T; L^{s}(\mathbb{R}^{3})), \quad \frac{2}{r} + \frac{3}{s} \leq 1, \quad 2 \leq r < \infty, \quad 3 \leq s \leq \infty,\tag{1.6}
$$

*then the solution can be extended smoothly beyond*  $t = T$ .

*Remark* 1.1*.* Theorem 1.1 shows that it is enough to control the rate of change in the direction of the velocity and the norm of *b* to get full regularity of the solutions. Notice that we add no conditions on the micro-rotational velocity *ω*.

*Remark* 1.2*.* Our theorem covers the results of Vasseur [\[12\]](#page-6-7) and Luo [\[7\]](#page-5-9) for Navier– Stokes and MHD equations, respectively. Observe that the condition [\(1.6\)](#page-1-1) is a scaling-invariant, but [\(1.5\)](#page-1-2) is not. Whether or not the 1*/*2 in [\(1.5\)](#page-1-2) can increase to 1 is our future work.

Before giving a proof, let us first recall the definition of weak solutions to system  $(1.1)$ .

#### DEFINITION 1.1

Let  $u_0, \omega_0, b_0 \in L^2(\mathbb{R}^3)$  with  $\nabla \cdot u_0 = \nabla \cdot b_0 = 0$ . A triple  $(u, \omega, b)$  of measurable functions on  $\mathbb{R}^3 \times (0, T)$  is said to be a weak solution of system [\(1.1\)](#page-0-0) if

- (1)  $u, \omega, b \in L^{\infty}(0, T; L^2) \cap L^2(0, T; H^1)$  with  $\nabla \cdot u = \nabla \cdot b = 0;$
- (2) System [\(1.1\)](#page-0-0) holds in the sense of distributions.

*Remark* 1.3*.* Testing  $(1.1)_1$  $(1.1)_1$ ,  $(1.1)_2$ ,  $(1.1)_3$  by *u*,  $\omega$ , *b* respectively, after suitable integration by parts, one has the energy inequality:

$$
\| (u(t), \omega(t), b(t)) \|_{L^2}^2 + 2(\mu + \chi) \int_0^t \|\nabla u(s)\|_{L^2}^2 ds + 2\gamma \int_0^t \|\nabla \omega(s)\|_{L^2}^2 ds + 2\nu \int_0^t \|\nabla b(s)\|_{L^2}^2 ds + 2\chi \int_0^t \|\omega(s)\|_{L^2}^2 ds \le \| (u_0, \omega_0, b_0) \|_{L^2}^2.
$$
 (1.7)

Throughout the proof in the next section, we shall frequently use the following interpolation inequality (see [\[17\]](#page-6-8)):

<span id="page-2-0"></span>
$$
\|u\|_{p,q} \leqslant C \|u\|_{\infty,2}^{\frac{3}{q}-\frac{1}{2}} \|\nabla u\|_{2,2}^{\frac{3}{2}-\frac{3}{q}} \leqslant C \left( \|u\|_{\infty,2} + \|\nabla u\|_{2,2} \right),\tag{1.8}
$$

for *(p, q)* satisfying

$$
\frac{2}{p} + \frac{3}{q} \geqslant \frac{3}{2}, \quad 2 \leqslant q \leqslant 6.
$$

In this paper, we shall use standard notations for Lebesgue space  $L^q(\mathbb{R}^3)$  endowed with the norm  $\|\cdot\|_q$ , and anisotropic Lebesgue space  $L^p(I; L^q(\mathbb{R}^3))$  endowed with the norm *|∙|<i>n*<sub>*n*</sub></sub> *<i>l*  $\subset \mathbb{R}^+$  is an interval. A constant  $C$  ( $C = C(*, *, ...$ ) which depends on the parameters) may differ from line to line.

### **2. Proof of Theorem 1.1**

By decreasing *p* or *r* if necessary, we may assume that

$$
\frac{2}{p} + \frac{3}{q} = \frac{1}{2}, \quad \frac{2}{r} + \frac{3}{s} = 1.
$$

For an  $\varepsilon > 0$  to be chosen sufficiently small (see [\(2.12\)](#page-5-10)), choose  $t_1 \in (0, T)$  such that

$$
\left\| \operatorname{div} \frac{u}{|u|} \right\|_{p,q} < \varepsilon \tag{2.1}
$$

and

$$
\|b\|_{r,s} < \varepsilon. \tag{2.2}
$$

Hereafter, the integrals are over  $\mathbb{R}^3 \times (t_1, T)$ .

Utilizing the regularity criteria [\(1.2\)](#page-0-1), we complete the proof of Theorem 1.1 provided

$$
u \in L^{8}(t_{1}, T; L^{4}(\mathbb{R}^{3})). \tag{2.3}
$$

To this end, denote by

<span id="page-3-3"></span>
$$
I = ||u||^{2}||_{\infty,2} + ||\nabla|u|^{2}||_{2,2} + ||b||^{2}||_{\infty,2} + ||\nabla|b|^{2}||_{2,2}.
$$
 (2.4)

Multiplying  $(1.1)_1$  $(1.1)_1$ ,  $(1.1)_3$  by  $|u|^2 u$ ,  $|b|^2 b$  respectively and integrating over  $\mathbb{R}^3 \times$  $(t_1, T)$ , we find that

<span id="page-3-0"></span>
$$
\frac{1}{4}||u||^{2}||_{\infty,2}^{2} + \frac{\mu + \chi}{2}||\nabla|u|^{2}||_{2,2}^{2} + (\mu + \chi)||u||\nabla u||_{2,2}^{2}
$$
\n
$$
= \int_{t_{1}}^{T} \int_{\mathbb{R}^{3}} -b \cdot \nabla(|u|^{2}u) \cdot b + (p + |b|^{2})u \cdot \nabla |u|^{2}
$$
\n
$$
+ \chi \omega \cdot \nabla \times (|u|^{2}u) dx dt, \qquad (2.5)
$$

as well as

<span id="page-3-1"></span>
$$
\frac{1}{4}|||b|^2||_{\infty,2}^2 + \frac{\nu}{2}||\nabla|b|^2||_{2,2}^2 + \nu |||b||\nabla b||_{2,2}^2
$$
\n
$$
= \int_{t_1}^T \int_{\mathbb{R}^3} -b \cdot \nabla(|b|^2 b) \cdot u \, dx \, dt. \tag{2.6}
$$

Using  $(2.5)$  and  $(2.6)$ , notice that (see [\[12\]](#page-6-7))

$$
\operatorname{div} \frac{u}{|u|} = -\frac{u}{|u|^2} \nabla u
$$

and (see  $[13]$ )

$$
|\nabla| u \|\leqslant |\nabla u|.
$$

Thus we have

<span id="page-3-2"></span>
$$
I^{2} + || |u|| |\nabla u|||_{2,2}^{2} + || |b|| |\nabla b|||_{2,2}^{2}
$$
  
\n
$$
\leq C \int_{t_{1}}^{T} \int_{\mathbb{R}^{3}} |b|^{2} |u| (|u|| |\nabla u| + |b|| |\nabla b|) \, dx dt
$$
  
\n
$$
+ C \int_{t_{1}}^{T} \int_{\mathbb{R}^{3}} (p + |b|^{2}) |u|^{3} | \, du \frac{u}{|u|} | \, dx dt
$$
  
\n
$$
+ C \int_{t_{1}}^{T} \int_{\mathbb{R}^{3}} | \omega | |u|^{2} | \nabla u | \, dx dt
$$
\n(2.7)  
\n
$$
\equiv I_{1} + I_{2} + I_{3}.
$$

Here *C* is a constant depending only on  $\mu$ ,  $\chi$ ,  $\nu$ .

Using Cauchy–Schwartz inequality,  $I_1$  can be bounded as

$$
I_1 \leq \frac{1}{4} ||u|| |\nabla u||_{2,2}^2 + \frac{1}{2} ||b|| |\nabla b||_{2,2}^2 + C ||b|^2 |u||_{2,2}^2.
$$

By generalized Hölder inequality and [\(1.8\)](#page-2-0), it follows that

$$
\begin{aligned} |||b|^2 |u||_{2,2}^2 &\leq ||b||_{r,s}^2 ||b||_{a,b}^2 ||u||_{c,d}^2 \\ &= ||b||_{r,s}^2 |||b|^2 ||\frac{a}{2}, \frac{b}{2} |||u|^2 ||\frac{c}{2}, \frac{d}{2} \\ &\leq C\varepsilon^2 I^2, \end{aligned}
$$

*,*

where

$$
\begin{cases} \frac{1}{r} + \frac{1}{a} + \frac{1}{c} = \frac{1}{2} = \frac{1}{s} + \frac{1}{b} + \frac{1}{d} \\ \frac{2}{a/2} + \frac{3}{b/2} = \frac{3}{2}, \\ \frac{2}{c/2} + \frac{3}{d/2} = \frac{3}{2}, \end{cases}
$$

and we have used [\(1.8\)](#page-2-0). In fact, we can choose

$$
a = c = \frac{4r}{r-2}
$$
,  $b = d = \frac{4s}{s-2}$ ,

where  $r$ ,  $s$  are as in  $(1.6)$ . Thus

<span id="page-4-0"></span>
$$
I_1 \leqslant \frac{1}{4} \left\| |u| \left| \nabla u \right| \right\|_{2,2}^2 + \frac{1}{2} \left\| |b| \left| \nabla b \right| \right\|_{2,2}^2 + C\varepsilon I^2. \tag{2.9}
$$

For  $I_2$ , let us first take divergence of  $(1.1)<sub>1</sub>$  $(1.1)<sub>1</sub>$  to see

$$
-\Delta p = \sum_{i,j=1}^{3} \partial_{ij} (u_i u_j - b_i b_j + \delta_{ij} |b|^2),
$$

thus classical Calderón-Zygmund estimates imply

$$
||p||_{a,b} \leqslant C(||u|^2||_{a,b} + ||b|^2||_{a,b}),
$$

invoking again the generalized Hölder inequality and [\(1.8\)](#page-2-0),

<span id="page-4-1"></span>
$$
I_2 \leq C \|p + |b|^2 \|a, b\| \|u\|^3 \|c, d\| \text{div} \frac{u}{|u|} \|_{p,q}
$$
  
\n
$$
\leq C \varepsilon \left( \|u\|_{a_1, b_1} \|u\|_{3c, 3d} + \|b\|_{a_1, b_1} \|b\|_{3c, 3d} \right) \|u\|_{3c, 3d}^3
$$
  
\n
$$
= C \varepsilon \|u\|_{a_1, b_1} \|u\|^2 \|_{\frac{3c}{2}, \frac{3d}{2}}^2 + C \varepsilon \|b\|_{a_1, b_1} \|b\|^2 \|_{\frac{3c}{2}, \frac{3d}{2}}^{\frac{3}{2}} \|u\|^2 \|_{\frac{3c}{2}, \frac{3d}{2}}^{\frac{3}{2}}
$$
  
\n
$$
\leq C \varepsilon I^2,
$$
\n(2.10)

where

$$
\begin{cases}\n\frac{1}{a} + \frac{1}{c} + \frac{1}{p} = 1 = \frac{1}{b} + \frac{1}{d} + \frac{1}{q}, \n\frac{1}{a} = \frac{1}{a_1} + \frac{1}{3c}, \quad \frac{1}{b} = \frac{1}{b_1} + \frac{1}{3d}, \n\frac{2}{a_1} + \frac{3}{b_1} = \frac{3}{2}, \quad \frac{2}{3c/2} + \frac{3}{3d/2} = \frac{3}{2}.\n\end{cases}
$$

In fact, we can choose

$$
a = c = \frac{2}{3}a_1 = \frac{2p}{p-1}
$$
,  $b = d = \frac{2}{3}b_1 = \frac{2q}{q-1}$ ,

where  $p, q$  are as in  $(1.5)$ .

Finally, using  $(1.8)$ ,  $I_3$  is treated as

<span id="page-5-11"></span>
$$
I_3 \leq \frac{1}{4} |||u|| |\nabla u|||^2_{2,2} + C |||w|| ||u|||^2_{2,2}
$$
  
\n
$$
\leq \frac{1}{4} |||u|| |\nabla u|||^2_{2,2} + C ||w||^2_{2,4} ||u||^2_{\infty,4}
$$
  
\n
$$
\leq \frac{1}{4} |||u|| |\nabla u|||^2_{2,2} + C I
$$
  
\n
$$
\leq \frac{1}{4} |||u|| |\nabla u|||^2_{2,2} + C \varepsilon I^2 + C.
$$
 (2.11)

Combining the estimates for  $I_1$ ,  $I_2$ ,  $I_3$ , i.e.  $(2.9)$ ,  $(2.10)$ ,  $(2.11)$ , and substituting into [\(2.7\)](#page-3-2), we find

$$
I^{2} + \frac{1}{2} |||u|| |\nabla u|||_{2,2}^{2} + \frac{1}{2} |||b|| |\nabla b|||_{2,2}^{2} \leq 3C\varepsilon I^{2} + C,
$$

where *C* is the generic constant appearing in  $(2.9)$ ,  $(2.10)$  and  $(2.11)$ . Thus, we see that

$$
I\leqslant \sqrt{2C}<\infty,
$$

provided

<span id="page-5-10"></span>
$$
\varepsilon = \frac{1}{6C}.\tag{2.12}
$$

Consequently, by  $(2.4)$ , we have

$$
u \in L^{\infty}(t_1, T, L^{4}(\mathbb{R}^3)) \subset L^{8}(t_1, T, L^{4}(\mathbb{R}^3)).
$$

The proof is completed.

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