Regularity criteria for the 3D magneto-micropolar fluid equations via the direction of the velocity

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Abstract. We consider sufficient conditions to ensure the smoothness of solutions to 3D magneto-micropolar fluid equations. It involves only the direction of the velocity and the magnetic field. Our result extends to the cases of Navier–Stokes and MHD equations.

Keywords. Magneto-micropolar fluid equations; regularity criteria; direction of velocity.

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1. Introduction and the main result

In this paper, we consider the 3D magneto-micropolar fluid equations studied by Galdi and Rionero [5]:

$$\begin{cases} \partial_t u + u \cdot \nabla u - (\mu + \chi) \Delta u - b \cdot \nabla b + \nabla (p + b^2) - \chi \nabla \times \omega = 0, \\ \partial_t \omega - \gamma \Delta \omega - \kappa \nabla \operatorname{div} \omega + 2\chi \omega + u \cdot \nabla \omega - \chi \nabla \times u = 0, \\ \partial_t b - \nu \Delta b + u \cdot \nabla b - b \cdot \nabla u = 0, \\ \nabla \cdot u = \nabla \cdot b = 0, \\ u(x, 0) = u_0(x), \, \omega(0, x) = \omega_0(x), \, b(0, x) = b_0(x). \end{cases}$$
(1.1)

Here u = u(x, t), b = b(x, t), $\omega = \omega(x, t)$ represent the velocity field, the magnetic field and the micro-rotational velocity respectively; p denotes the hydrodynamic pressure; $\mu > 0$ is the kinematic viscosity, $\chi > 0$ is the vortex viscosity, $\kappa > 0$ and $\gamma > 0$ are the spin viscosities, $1/\nu$ (with $\nu > 0$) is the magnetic Reynold; while u_0 , b_0 , ω_0 are the corresponding initial data with $\nabla \cdot u_0 = \nabla \cdot b_0 = 0$.

The global weak solution to system (1.1) is established by Rojas-Medar and Boldrini [10], while the local strong solutions are given by Rojas-Medar [9]. However, whether or not the local strong solutions can exist globally is still an open problem. Thus regularity criteria appears. Notice that:

(1) Yuan [14] first established the following fundamental regularity criterion in terms of the velocity or its gradient

$$u \in L^{p}(0, T; L^{q}(\mathbb{R}^{3})), \quad \frac{2}{p} + \frac{3}{q} = 1, \quad 3 < q \leq \infty$$
 (1.2)

and

$$\nabla u \in L^p(0,T; L^q(\mathbb{R}^3)), \quad \frac{2}{p} + \frac{3}{q} = 2, \quad \frac{3}{2} < q \leq \infty.$$
 (1.3)

Then Gala [4] extended it to the Morrey–Campanato spaces, Zhang *et al* [16] improved it to some more general Triebel–Lizorkin spaces.

- (2) When ω = b = 0, system (1.1) is just the classical Naiver–Stokes equations. Serrin [11], Prodi [8] and Beirão da Veiga [1] proved regularity if some scaling-invariant norm of u or ∇u is bounded.
- (3) When $\omega = 0$, system (1.1) is then the 3D MHD equations, He and Xin [6], and Zhou [19] gave criteria similar to the case for Navier–Stokes equations.
- (4) When b = 0, system (1.1) is reduced to the micropolar fluid equations, Yuan [13] gave some criteria in Lorentz spaces.

For later developments, see [2, 3, 15, 18] and references cited therein. Recently, Vasseur [12] proved that if

$$\operatorname{div}\frac{u}{|u|} \in L^{p}(0, T; L^{q}(\mathbb{R}^{3})), \quad \frac{2}{p} + \frac{3}{q} \leqslant \frac{1}{2}, \quad p \ge 4, \quad q \ge 6,$$
(1.4)

then the solutions to the Navier–Stokes equations are smooth. Later, Luo [7] extended (1.4) to the MHD equations, but involves the magnetic field also. We now extend the result of Vasseur [12] and Luo [7] to system (1.1). The main result is the following:

Theorem 1.1. Let $u_0, \omega_0, b_0 \in H^1(\mathbb{R}^3)$ with $\nabla \cdot u_0 = \nabla \cdot b_0 = 0$ in the sense of distributions. Suppose that (u, ω, b) is a strong solution to (1.1) in (0, T) such that

$$u, \omega, b \in C((0, T); H^1(\mathbb{R}^3)) \cap C((0, T); H^2(\mathbb{R}^3))$$

and $\nabla \cdot u = \nabla \cdot b = 0$. If additionally,

$$\operatorname{div}\frac{u}{|u|} \in L^{p}(0,T; L^{q}(\mathbb{R}^{3})), \ \frac{2}{p} + \frac{3}{q} \leqslant \frac{1}{2}, \ 4 \leqslant p < \infty, \ 6 \leqslant q \leqslant \infty$$

$$(1.5)$$

and

$$b \in L^r(0, T; L^s(\mathbb{R}^3)), \quad \frac{2}{r} + \frac{3}{s} \leqslant 1, \quad 2 \leqslant r < \infty, \quad 3 \leqslant s \leqslant \infty,$$
 (1.6)

then the solution can be extended smoothly beyond t = T.

Remark 1.1. Theorem 1.1 shows that it is enough to control the rate of change in the direction of the velocity and the norm of *b* to get full regularity of the solutions. Notice that we add no conditions on the micro-rotational velocity ω .

Remark 1.2. Our theorem covers the results of Vasseur [12] and Luo [7] for Navier–Stokes and MHD equations, respectively. Observe that the condition (1.6) is a scaling-invariant, but (1.5) is not. Whether or not the 1/2 in (1.5) can increase to 1 is our future work.

Before giving a proof, let us first recall the definition of weak solutions to system (1.1).

DEFINITION 1.1

Let $u_0, \omega_0, b_0 \in L^2(\mathbb{R}^3)$ with $\nabla \cdot u_0 = \nabla \cdot b_0 = 0$. A triple (u, ω, b) of measurable functions on $\mathbb{R}^3 \times (0, T)$ is said to be a weak solution of system (1.1) if

- (1) $u, \omega, b \in L^{\infty}(0, T; L^2) \cap L^2(0, T; H^1)$ with $\nabla \cdot u = \nabla \cdot b = 0$;
- (2) System (1.1) holds in the sense of distributions.

Remark 1.3. Testing $(1.1)_1$, $(1.1)_2$, $(1.1)_3$ by u, ω, b respectively, after suitable integration by parts, one has the energy inequality:

$$\|(u(t), \omega(t), b(t))\|_{L^{2}}^{2} + 2(\mu + \chi) \int_{0}^{t} \|\nabla u(s)\|_{L^{2}}^{2} ds + 2\gamma \int_{0}^{t} \|\nabla \omega(s)\|_{L^{2}}^{2} ds + 2\nu \int_{0}^{t} \|\nabla b(s)\|_{L^{2}}^{2} ds + 2\chi \int_{0}^{t} \|\omega(s)\|_{L^{2}}^{2} ds \leq \|(u_{0}, \omega_{0}, b_{0})\|_{L^{2}}^{2}.$$
(1.7)

Throughout the proof in the next section, we shall frequently use the following interpolation inequality (see [17]):

$$\|u\|_{p,q} \leq C \|u\|_{\infty,2}^{\frac{3}{q}-\frac{1}{2}} \|\nabla u\|_{2,2}^{\frac{3}{2}-\frac{3}{q}} \leq C \left(\|u\|_{\infty,2} + \|\nabla u\|_{2,2}\right),$$
(1.8)

for (p, q) satisfying

$$\frac{2}{p} + \frac{3}{q} \geqslant \frac{3}{2}, \quad 2 \leqslant q \leqslant 6.$$

In this paper, we shall use standard notations for Lebesgue space $L^q(\mathbb{R}^3)$ endowed with the norm $\|\cdot\|_q$, and anisotropic Lebesgue space $L^p(I; L^q(\mathbb{R}^3))$ endowed with the norm $\|\cdot\|_{p,q}$. Here $I \subset \mathbb{R}^+$ is an interval. A constant C (C = C(*, *, ...) which depends on the parameters) may differ from line to line.

2. Proof of Theorem 1.1

By decreasing p or r if necessary, we may assume that

$$\frac{2}{p} + \frac{3}{q} = \frac{1}{2}, \quad \frac{2}{r} + \frac{3}{s} = 1.$$

For an $\varepsilon > 0$ to be chosen sufficiently small (see (2.12)), choose $t_1 \in (0, T)$ such that

$$\left\| \operatorname{div} \frac{u}{|u|} \right\|_{p,q} < \varepsilon \tag{2.1}$$

and

$$\|b\|_{r,s} < \varepsilon. \tag{2.2}$$

Hereafter, the integrals are over $\mathbb{R}^3 \times (t_1, T)$.

Utilizing the regularity criteria (1.2), we complete the proof of Theorem 1.1 provided

$$u \in L^{8}(t_{1}, T; L^{4}(\mathbb{R}^{3})).$$
 (2.3)

To this end, denote by

$$I = ||u|^{2}||_{\infty,2} + ||\nabla|u|^{2}||_{2,2} + ||b|^{2}||_{\infty,2} + ||\nabla|b|^{2}||_{2,2}.$$
(2.4)

Multiplying $(1.1)_1, (1.1)_3$ by $|u|^2 u, |b|^2 b$ respectively and integrating over $\mathbb{R}^3 \times (t_1, T)$, we find that

$$\frac{1}{4} \| |u|^2 \|_{\infty,2}^2 + \frac{\mu + \chi}{2} \| \nabla |u|^2 \|_{2,2}^2 + (\mu + \chi) \| |u| \| \nabla u \| \|_{2,2}^2
= \int_{t_1}^T \int_{\mathbb{R}^3} -b \cdot \nabla (|u|^2 u) \cdot b + (p + |b|^2) u \cdot \nabla |u|^2
+ \chi \omega \cdot \nabla \times (|u|^2 u) dx dt,$$
(2.5)

as well as

$$\frac{1}{4} ||b|^2||_{\infty,2}^2 + \frac{\nu}{2} ||\nabla|b|^2||_{2,2}^2 + \nu ||b||\nabla b|||_{2,2}^2$$
$$= \int_{t_1}^T \int_{\mathbb{R}^3} -b \cdot \nabla(|b|^2 b) \cdot u \, dx dt.$$
(2.6)

Using (2.5) and (2.6), notice that (see [12])

$$\operatorname{div}\frac{u}{|u|} = -\frac{u}{|u|^2}\nabla u$$

and (see [13])

$$|\nabla|\,u\|\leqslant |\nabla u|\,.$$

Thus we have

$$I^{2} + |||u| |\nabla u|||_{2,2}^{2} + |||b| |\nabla b|||_{2,2}^{2}$$

$$\leq C \int_{t_{1}}^{T} \int_{\mathbb{R}^{3}} |b|^{2} |u| (|u| |\nabla u| + |b| |\nabla b|) \, dx dt$$

$$+ C \int_{t_{1}}^{T} \int_{\mathbb{R}^{3}} (p + |b|^{2}) |u|^{3} \left| \operatorname{div} \frac{u}{|u|} \right| \, dx dt$$

$$+ C \int_{t_{1}}^{T} \int_{\mathbb{R}^{3}} |\omega| |u|^{2} |\nabla u| \, dx dt \qquad (2.7)$$

$$\equiv I_{1} + I_{2} + I_{3}. \qquad (2.8)$$

Here *C* is a constant depending only on μ , χ , ν .

Using Cauchy–Schwartz inequality, I_1 can be bounded as

$$I_1 \leqslant \frac{1}{4} |||u|| |\nabla u|||_{2,2}^2 + \frac{1}{2} |||b|| |\nabla b|||_{2,2}^2 + C |||b|^2|u||_{2,2}^2.$$

By generalized Hölder inequality and (1.8), it follows that

$$\begin{aligned} \||b|^{2}|u|\|_{2,2}^{2} &\leq \|b\|_{r,s}^{2} \|b\|_{a,b}^{2} \|u\|_{c,d}^{2} \\ &= \|b\|_{r,s}^{2} \||b|^{2}\|_{\frac{a}{2},\frac{b}{2}} \||u|^{2}\|_{\frac{c}{2},\frac{d}{2}} \\ &\leq C\varepsilon^{2}I^{2}, \end{aligned}$$

where

$$\frac{1}{r} + \frac{1}{a} + \frac{1}{c} = \frac{1}{2} = \frac{1}{s} + \frac{1}{b} + \frac{1}{d}$$
$$\frac{2}{a/2} + \frac{3}{b/2} = \frac{3}{2},$$
$$\frac{2}{c/2} + \frac{3}{d/2} = \frac{3}{2},$$

and we have used (1.8). In fact, we can choose

$$a = c = \frac{4r}{r-2}, \quad b = d = \frac{4s}{s-2},$$

where r, s are as in (1.6). Thus

$$I_{1} \leq \frac{1}{4} \||u| \|\nabla u\|_{2,2}^{2} + \frac{1}{2} \||b| \|\nabla b\|_{2,2}^{2} + C\varepsilon I^{2}.$$

$$(2.9)$$

For I_2 , let us first take divergence of $(1.1)_1$ to see

$$-\Delta p = \sum_{i,j=1}^{3} \partial_{ij} (u_i u_j - b_i b_j + \delta_{ij} |b|^2),$$

thus classical Calderón-Zygmund estimates imply

$$\|p\|_{a,b} \leq C(\||u|^2\|_{a,b} + \||b|^2\|_{a,b}),$$

invoking again the generalized Hölder inequality and (1.8),

$$I_{2} \leqslant C \|p + |b|^{2} \|_{a,b} \||u|^{3} \|_{c,d} \|\operatorname{div} \frac{u}{|u|} \|_{p,q}$$

$$\leqslant C \varepsilon \left(\|u\|_{a_{1},b_{1}} \|u\|_{3c,3d} + \|b\|_{a_{1},b_{1}} \|b\|_{3c,3d} \right) \|u\|_{3c,3d}^{3}$$

$$= C \varepsilon \|u\|_{a_{1},b_{1}} \||u|^{2} \|_{\frac{3c}{2},\frac{3d}{2}}^{2} + C \varepsilon \|b\|_{a_{1},b_{1}} \||b|^{2} \|_{\frac{3c}{2},\frac{3d}{2}}^{\frac{1}{2}} \||u|^{2} \|_{\frac{3c}{2},\frac{3d}{2}}^{\frac{3}{2}}$$

$$\leqslant C \varepsilon I^{2}, \qquad (2.10)$$

where

$$\frac{1}{a} + \frac{1}{c} + \frac{1}{p} = 1 = \frac{1}{b} + \frac{1}{d} + \frac{1}{q},$$

$$\frac{1}{a} = \frac{1}{a_1} + \frac{1}{3c}, \quad \frac{1}{b} = \frac{1}{b_1} + \frac{1}{3d},$$

$$\frac{2}{a_1} + \frac{3}{b_1} = \frac{3}{2}, \quad \frac{2}{3c/2} + \frac{3}{3d/2} = \frac{3}{2}$$

In fact, we can choose

$$a = c = \frac{2}{3}a_1 = \frac{2p}{p-1}, \quad b = d = \frac{2}{3}b_1 = \frac{2q}{q-1},$$

where p, q are as in (1.5).

Finally, using (1.8), I_3 is treated as

$$I_{3} \leqslant \frac{1}{4} |||u|| \nabla u||_{2,2}^{2} + C |||\omega||u|||_{2,2}^{2}$$

$$\leqslant \frac{1}{4} |||u|| \nabla u||_{2,2}^{2} + C ||\omega||_{2,4}^{2} ||u||_{\infty,4}^{2}$$

$$\leqslant \frac{1}{4} |||u|| \nabla u||_{2,2}^{2} + CI$$

$$\leqslant \frac{1}{4} |||u|| \nabla u||_{2,2}^{2} + C\varepsilon I^{2} + C.$$
(2.11)

Combining the estimates for I_1 , I_2 , I_3 , i.e. (2.9), (2.10), (2.11), and substituting into (2.7), we find

$$I^{2} + \frac{1}{2} |||u|| |\nabla u|||_{2,2}^{2} + \frac{1}{2} |||b|| |\nabla b|||_{2,2}^{2} \leq 3C\varepsilon I^{2} + C,$$

where C is the generic constant appearing in (2.9), (2.10) and (2.11). Thus, we see that

$$I\leqslant \sqrt{2C}<\infty,$$

provided

$$\varepsilon = \frac{1}{6C}.\tag{2.12}$$

Consequently, by (2.4), we have

$$u \in L^{\infty}(t_1, T, L^4(\mathbb{R}^3)) \subset L^8(t_1, T, L^4(\mathbb{R}^3)).$$

The proof is completed.

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