# **Locally Divergence-Free Discontinuous Galerkin Methods for MHD Equations**

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In this paper, we continue our investigation of the locally divergence-free discontinuous Galerkin method, originally developed for the linear Maxwell equations (*J. Comput. Phys*. **194** 588–610 (2004)), to solve the nonlinear ideal magnetohydrodynamics (MHD) equations. The distinctive feature of such method is the use of approximate solutions that are exactly divergence-free inside each element for the magnetic field. As a consequence, this method has a smaller computational cost than the traditional discontinuous Galerkin method with standard piecewise polynomial spaces. We formulate the locally divergence-free discontinuous Galerkin method for the MHD equations and perform extensive one and two-dimensional numerical experiments for both smooth solutions and solutions with discontinuities. Our computational results demonstrate that the locally divergence-free discontinuous Galerkin method, with a reduced cost comparing to the traditional discontinuous Galerkin method, can maintain the same accuracy for smooth solutions and can enhance the numerical stability of the scheme and reduce certain nonphysical features in some of the test cases.

**KEY WORDS:** Discontinuous Galerkin method; divergence-free solutions; magnetohydrodynamics (MHD) equations.

# **1. INTRODUCTION**

In this paper we continue our development of the locally divergence-free discontinuous Galerkin method, studied first in [11] jointly with Cockburn for solving the linear Maxwell equations, to solve the two dimensional nonlinear ideal magnetohydrodynamics (MHD) equations. The method

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can be applied to three dimensional cases with no essential difficulty, but we will not consider it in this paper.

The locally divergence-free discontinuous Galerkin method is based on the traditional discontinuous Galerkin methods equipped with TVD Runge–Kutta time discretization (RKDG) [9, 12, 13]. The distinctive feature of the locally divergence-free discontinuous Galerkin method is the use of approximate solutions that are exactly divergence-free inside each element for certain components which should be divergence-free according to the PDEs, for example the magnetic field in the MHD equations.

There are many partial differential equations with solutions which are divergence-free. Examples include the incompressible Euler and Navier–Stokes equations, the MHD equations, and the Maxwell equations. For some of the problems, such as the MHD equations and the Maxwell equations, the solutions of the PDE should automatically satisfy the divergence-free condition if the initial data is divergence-free, but it is widely known that negligence in dealing with the divergence-free condition numerically sometimes can lead to serious defects, see, e.g. [17, 20, 6].

Attempts have been made in the literature to enforce explicitly the divergence-free condition. In the following we will review some of the divergence "cleaning" techniques in the context of MHD equations.

One way for the divergence correction is by projection, which, based on "Hodge decomposition", was first suggested by Brackbill and Barnes [6] in the context of MHD. The numerical magnetic field **B** is projected into a zero divergence vector space and the projected **B** is used in the next time step. To implement this, usually a Poisson equation needs to be solved. Another method is Powell's source term formulation [22], which is derived from the physical laws if  $\nabla \cdot \mathbf{B} = 0$  is not used. Source terms proportional to ∇ ·**B** are added to the equations, which makes the system well behaved but nonconservative. For MHD equations, there is also the generalized Lagrange multiplier method, which was proposed by Dedner *et al*. [15]. With this method, the divergence errors could be transported to the domain boundaries with certain speed and damped at the same time.

Another approach, often referred to as "constrained transport methods", was first brought up by Yee [32] in electromagnetics, and then adapted to MHD equations by Brecht *et al*. [7], Evans and Hawley [16], Stone and Norman [29], etc. In this approach, a staggered mesh is used, and a suitably defined discrete approximation to the divergence of the solution can be maintained exactly zero. This method has been further developed recently by combining with the higher order Godunov type schemes by Dai and Woodward [14], Ryu *et al*. [25], Balsara and Spicer [5], etc. In [3, 4], Balsara developed such divergence-free reconstruction strategy in an adaptive mesh setting. In [30], Toth compared some of the

methods mentioned above and introduced the field/flux-interpolated central difference (CD) approaches, in which no staggered mesh is needed.

In the context of Stokes equations and the stationary Navier–Stokes equations, Baker *et al*. [2, 19] introduced a discontinuous Galerkin method by using an interior penalty method with locally divergence-free approximate solutions. Optimal error estimates were proven.

Similar to [11] for the two-dimensional Maxwell equations, we follow the approach of Baker *et al*. [2, 19] and use the locally divergence-free piecewise polynomials as the solution space in the discontinuous Galerkin method to solve ideal MHD equations. Because the space is smaller, we can save in computational cost when using the locally divergence-free piecewise polynomial space compared with the standard piecewise polynomial space for the RKDG method. Theoretical study in [11], and numerical experiments in [11] and in this paper, indicate that this saving is not at the price of the degeneration of the solutions in stability and accuracy. On the contrary, both accuracy and stability can be enhanced in many cases. Compared with the Maxwell equations, the nonlinearity gives the solutions of MHD equations much more complicated features. The RKDG method with locally divergence-free piecewise polynomial space can enhance numerical stability in several test cases, such as the Orszag– Tang vortex example and the shock reflection example. Also, nonphysical features are reduced in several test cases, such as the rotor example.

The paper is organized as follows. In Sec. 2, we describe the equations and introduce the locally divergence-free space as well as the numerical formulation of the algorithm. The numerical results are presented in Sec. 3. Concluding remarks are made in Sec. 4.

# **2. EQUATIONS AND LOCALLY DIVERGENCE-FREE DISCONTINUOUS GALERKIN METHOD**

Electrically conducting fluid flow in which the electromagnetic forces can be of the same order or even greater than the hydrodynamic ones is often modeled by MHD equations. The ideal MHD equations consist of a set of nonlinear hyperbolic equations,

$$
\partial_t \rho + \nabla \cdot (\rho \mathbf{u}) = 0,
$$
  

$$
\partial_t (\rho \mathbf{u}) + \nabla \cdot \left[ \rho \mathbf{u} \mathbf{u}^T + \left( p + \frac{1}{2} |\mathbf{B}|^2 \right) \mathcal{I} - \mathbf{B} \mathbf{B}^T \right] = 0,
$$
  

$$
\partial_t \mathbf{B} + \nabla \cdot (\mathbf{u} \mathbf{B}^T - \mathbf{B} \mathbf{u}^T) = 0,
$$
  

$$
\partial_t E + \nabla \cdot \left[ \left( E + p + \frac{1}{2} |\mathbf{B}|^2 \right) \mathbf{u} - \mathbf{B} (\mathbf{u} \cdot \mathbf{B}) \right] = 0
$$
 (2.1)

with the additional divergence constraint

$$
\nabla \cdot \mathbf{B} = 0. \tag{2.2}
$$

Here  $\rho$ ,  $p$ ,  $\mathbf{u} = (u_x, u_y, u_z)$ ,  $\mathbf{B} = (B_x, B_y, B_z)$  and *E* denote the mass density, the hydrodynamic pressure, the velocity field, the magnetic field, and the total energy, respectively. The ratio of the specific heats is given by *γ* and

$$
E = \frac{1}{2}\rho |\mathbf{u}|^2 + \frac{1}{2}|\mathbf{B}|^2 + \frac{p}{\gamma - 1}.
$$

This system combines the equations of gas dynamics with Maxwell equations for problems in which relativistic, viscous, and resistive effects can be neglected; the permeability is set to be unity. If the initial magnetic field satisfies the divergence-free condition (2.2), the exact solution will automatically satisfy the constraint (2.2) for all time.

We rewrite Eq.  $(2.1)$  in the conservative form

$$
\mathbf{U}_t + \nabla \cdot \mathbf{F}(\mathbf{U}) = 0,\tag{2.3}
$$

where

$$
\mathbf{U} = (\rho, \rho u_x, \rho u_y, \rho u_z, B_x, B_y, B_z, E)^T.
$$

$$
F_1(\mathbf{U}) = (\rho u_x, \rho u_x^2 + p + \frac{1}{2} |\mathbf{B}|^2 - B_x^2, \rho u_x u_y
$$
  
-B\_x B\_y, \rho u\_x u\_z - B\_x B\_z, 0, u\_x B\_y - u\_y B\_x, u\_x B\_z  
-u\_z B\_x, u\_x (E + p + \frac{1}{2} |\mathbf{B}|^2) - B\_x (u\_x B\_x + u\_y B\_y + u\_z B\_z))^T.

$$
F_2(\mathbf{U}) = (\rho u_y, \rho u_y u_x - B_y B_x, \rho u_y^2 + p + \frac{1}{2} |\mathbf{B}|^2 - B_y^2,
$$
  
\n
$$
\rho u_y u_z - B_y B_z, u_y B_x - u_x B_y, 0, u_y B_z - u_z B_y,
$$
  
\n
$$
u_y (E + p + \frac{1}{2} |\mathbf{B}|^2) - B_y (u_x B_x + u_y B_y + u_z B_z))^T.
$$

For notational convenience, we will also use  $\mathbf{U}^p = (\rho, u_x, u_y, u_z, B_x, B_y,$  $B_z$ ,  $p$ <sup>T</sup> to denote the primitive variables and use  $\widetilde{U}$  (or  $\widetilde{U}^p$ ,  $\widetilde{F}_1$ ) to denote a vector which contains all the components of **U** (or  $\mathbf{U}^p$ ,  $F_1$ ) except the fifth one.

Starting with a triangulation  $\mathcal{T}_h$  of the domain  $\Omega$ , with the element being denoted by K, the edge by e, and the outward unit normal by  $\mathbf{n} =$ **, following the usual definition of discontinuous Galerkin** 

methods for conservation laws, e.g. [9, 12], we obtain the RKDG formulation for (2.3): find  $U_h \in V_h$ , such that for all  $v \in V_h$ ,

$$
\int_{K} \mathbf{U}_{\mathbf{h}t} \cdot \mathbf{v} d\mathbf{x} + \sum_{e \in \partial K} \int_{e} \mathbf{h}_{e,K}(\mathbf{U}_{\mathbf{h}}^{\text{int}(K)}, \mathbf{U}_{\mathbf{h}}^{\text{ext}(K)}, \mathbf{n}_{e,K}) \cdot \mathbf{v} ds \qquad (2.4)
$$

$$
-\int_{K} \mathbf{F}(\mathbf{U}_{\mathbf{h}}) \cdot \nabla \mathbf{v} d\mathbf{x} = 0, \ \forall K
$$

holds, where  $V_h$  is the solution space, which is the same as the test space and given by

$$
\mathbf{V}_h = \mathbf{V}_h^k = \left\{ \mathbf{v} : \mathbf{v}|_K \in \mathbf{P}^k(K), \left( \frac{\partial v_5}{\partial x} + \frac{\partial v_6}{\partial y} \right) |_K = 0, \forall K \right\}
$$
 (2.5)

with  $\mathbf{P}^{k}(K) = (P^{k}(K))^{8}$ , and  $P^{k}(K)$  denotes the space of polynomials in K of degree at most *k*. By this definition, we will have solution spaces which contain magnetic fields with zero divergence within each element. We call these spaces the locally divergence-free spaces, and call the discontinuous Galerkin methods with these solution spaces the locally divergence-free discontinuous Galerkin methods. Since the solution space is a subspace of the standard piecewise polynomial space, we expect a saving in computational cost by using the locally divergence-free spaces. More precisely, the dimension of  $V_h^k|_K$  is  $((k+1)(k+4)/2) + 3(k+1)(k+2)$ , which asymptotically is 7/8 of  $4(k+1)(k+2)$ , the dimension of  $(P^k(K))^8$ . It is easy to obtain local bases for  $\{(v_5, v_6) : \mathbf{v} = \{v_i\}_{i=1}^8, \mathbf{v} \in \mathbf{V}_h^k\}$  within *K*. For example, if *K* is a rectangle, with center  $(x_i, y_j)$ , and width  $\Delta x_i, \Delta y_j$ , if we denote  $\bar{X} = (x - x_i)/\Delta x_i$ ,  $\bar{Y} = (y - y_i)/\Delta y_i$ , one set of orthogonal bases would be, when  $k = 1$ ,

$$
\begin{pmatrix} 1 \\ 0 \end{pmatrix}
$$
,  $\begin{pmatrix} \Delta x_i \bar{X} \\ -\Delta y_j \bar{Y} \end{pmatrix}$ ,  $\begin{pmatrix} \bar{Y} \\ 0 \end{pmatrix}$ ,  $\begin{pmatrix} 0 \\ 1 \end{pmatrix}$ ,  $\begin{pmatrix} 0 \\ \bar{X} \end{pmatrix}$ .

For  $k = 2$ , we need to add

$$
\begin{pmatrix}\n\Delta x_i (12\overline{X}^2 - 1) \\
-24\Delta y_j \overline{X} \overline{Y}\n\end{pmatrix}, \quad\n\begin{pmatrix}\n-24\Delta x_i \overline{X} \overline{Y} \\
\Delta y_j (12\overline{Y}^2 - 1)\n\end{pmatrix}, \quad\n\begin{pmatrix}\n12\overline{Y}^2 - 1 \\
0\n\end{pmatrix}, \quad\n\begin{pmatrix}\n0 \\
12\overline{X}^2 - 1\n\end{pmatrix}.
$$

In general, the local bases  $\{(v_5, v_6) : \mathbf{v} = \{v_i\}_{i=1}^8, \mathbf{v} \in \mathbf{V}_h^k\}$  in *K* could be generated by taking the *curl* of bases of  $P^{k+1}(\overline{K})$ .

In (2.4),  $\mathbf{h}_{e,K}(\mathbf{v}^{\text{int}(K)}, \mathbf{v}^{\text{ext}(K)}, \mathbf{n}_{e,K})$  is the numerical flux, which is an exact or approximate Riemann solver, consistent with  $\mathbf{F}(\mathbf{U}) \cdot \mathbf{n}_{e,K}$  and conservative,

$$
\mathbf{h}_{e,K}(\mathbf{v}^{\text{int}(K)},\mathbf{v}^{\text{ext}(K)},\mathbf{n}_{e,K})+\mathbf{h}_{e,K'}(\mathbf{v}^{\text{int}(K')},\mathbf{v}^{\text{ext}(K')},\mathbf{n}_{e,K'})=0, K\cap K'=e,
$$

here  $\mathbf{v}^{\text{int}(K)}$ ,  $\mathbf{v}^{\text{ext}(K)}$  are the limits of **v** at the interface *e* from the interior and exterior of *K*, respectively. The one we will use in our numerical examples is the local Lax–Friedrichs flux,

$$
\mathbf{h}_{e,K}(\mathbf{a},\mathbf{b},\mathbf{n}_{e,K}) = \frac{1}{2} [F(\mathbf{a}) \cdot \mathbf{n}_{e,K} + F(\mathbf{b}) \cdot \mathbf{n}_{e,K} - \alpha_{e,K}(\mathbf{b}-\mathbf{a})],\tag{2.6}
$$

where  $\alpha_{e,K}$  is an estimate of the largest absolute value of eigenvalues of the Jacobi  $(\partial/\partial v)$ **F**(**v**) · **n**<sub>*e,K*</sub> in the neighborhood of the edge *e*.

For discontinuous Galerkin methods applied to nonlinear systems such as (2.1), nonlinear limiters are often needed. In this paper we use the minmod TVB slope limiter by Shu [26] and Cockburn *et al*. [10, 12], which has a parameter *M* related to the magnitude of the second derivatives of the solution at smooth extrema. The limiter is implemented in local characteristic fields and we use the eigenvector system from [18]. We will not give the details of the limiter but mention that after the limiting process, the magnetic fields should still be locally divergence-free. It is a difficult task and an active research area to find good limiters for the discontinuous Galerkin methods in the presence of strong shocks, to obtain nonoscillatory results which are high order accurate in smooth regions. The emphasis of this paper is not on limiters, thus we will not discuss the optimal choice of limiters, rather we will use the minmod TVB slope limiter with suitable *M* (not necessarily optimal *M*) in the numerical tests. We do point out that sometimes the numerical results are sensitive to the choice of limiters. We refer the readers to [23, 24] for a recent effort in designing a new class of limiters for the discontinuous Galerkin method based on weighted essentially nonoscillatory (WENO) schemes.

#### **3. NUMERICAL EXAMPLES**

In this section, we present the numerical results of the discontinuous Galerkin methods for one dimensional test problems using the standard piecewise *P<sup>k</sup>* elements, and two dimensional test problems using both the standard piecewise  $P^k$  elements and locally divergence-free piecewise  $P^k$  elements, or just one of them if the results are similar, with  $k = 1, 2$ , namely the second and third-order methods. We include the one dimensional examples here for the sake of completeness. By the notations introduced in Sec. 2, the one-dimensional system could be written as follows

 $\widetilde{\mathbf{U}}_t + (\widetilde{F}_1(\widetilde{\mathbf{U}}))_x = 0.$ 

The divergence-free condition on the magnetic field is reduced to  $B_x = \text{con}$ stant. In all tests, the third-order TVD Runge–Kutta time discretization [28] will be used. The time step will be dynamically determined by

$$
\Delta t = \frac{C_{cfl} \min(\Delta x)}{\max(|u_x| + c_f)}
$$

in the one dimensional case and by

$$
\Delta t = C_{cfl} / \left\{ \frac{\max(|u_x| + c_f^x)}{\min(\Delta x)} + \frac{\max(|u_y| + c_f^y)}{\min(\Delta y)} \right\}
$$

in the two-dimensional case, where  $c_f$  ( $c_f^x$ ,  $c_f^y$ ) is the fast speed (in *x*, *y* direction, see [18] for the definition), the maxima (minima) are taken over all the computational cells, and the CFL number  $C_{cfl}$  is taken as 0.3 for the  $P<sup>1</sup>$  case and 0.18 for the  $P<sup>2</sup>$  case. In order to compute the integrals in formulation (2.4) numerically, the proper quadrature rules need to be used [12]. Uniform rectangular meshes are used in the computation unless otherwise indicated. The method, however, can be applied on arbitrary triangulation.

#### **3.1. One-Dimensional Riemann Problems**

In this Section, we solve standard one-dimensional Riemann problems, see e.g. [18]. The first Riemann problem is given by

$$
\widetilde{\mathbf{U}}^p = \begin{cases} (1.000, 0, 0, 0, +1, 0, 1.0) & \text{for } x < 0, \\ (0.125, 0, 0, 0, -1, 0, 0.1) & \text{for } x > 0 \end{cases}
$$

with  $B_x = 0.75$  and  $\gamma = 2$  on the computational domain [−1, 1]. Note that the hydrodynamic data is the same as the one for Sod's Riemann problem. This is the example used by Brio and Wu in [8] to show the formation of the compound wave in MHD.

The solution at  $t = 0.2$  is shown in Fig. 1, which includes the left moving waves: the fast rarefaction wave, the intermediate shock attached by a slow rarefaction wave; and the right moving waves: the contact discontinuity, a slow shock, and a fast rarefaction wave. The results obtained with 5000 cells serve as a reference in Fig. 1. We can see that all the waves are resolved very well.

The second Riemann problem is

$$
\widetilde{\mathbf{U}}^p = \begin{cases} (1.000, 0, 0, 0, +1, 0, 1000) & \text{for } x < 0, \\ (0.125, 0, 0, 0, -1, 0, 0.10) & \text{for } x > 0 \end{cases}
$$



**Fig. 1.** The first one-dimensional Riemann example. *P*<sup>2</sup> with 800 cells (circles) on a background solid line computed with 5000 cells.  $t = 0.2$  and  $M = 1$ . The symbol *FR* denotes a fast rarefaction wave; *SM* is a compound wave (an intermediate shock followed by a slow rarefaction wave); *C* is a contact discontinuity; *SS* is a slow shock. Top:  $\rho$ ; middle left:  $u_x$ ; middle right:  $u_y$ ; bottom left:  $B_y$ ; bottom right: p.



**Fig. 2.** The second one-dimensional Riemann example.  $P^2$  with 600 cells (circles) on a background solid line of the exact solution.  $t = 0.012$  and  $M = 1$ . Top left:  $\rho$ ; top right:  $u_x$ ; bottom left:  $B_y$ ; bottom right: *p*.

with  $B_x = 0$  and  $\gamma = 2$ . This problem is used to evaluate the code for high Mach number flow. If one regards the term  $p + (1/2)|\mathbf{B}|^2$  as the "hydrodynamic pressure", the system becomes a standard hydrodynamical Riemann problem, thus the exact solution can be easily obtained. The computational domain is taken to be  $[-1, 1]$ . The solution at  $t = 0.012$  is shown in Fig. 2, with the solid line as the exact solution.

In both problems we have applied the TVB minmod limiter with the TVB constant  $M = 1$  in the local characteristic fields computed by the eigenvectors evaluated with the cell averages, see [9, 12] for the details of the implementation of such limiters. Our experience obtained from these one-dimensional experiments is that the limiter should be used in the local characteristic fields rather than in the physical fields, in order to effectively control the spurious oscillations. This is consistent with the experience of hydrodynamic computations using the discontinuous Galerkin methods [10, 12].

Mesh	$L^2$ error	Order	$L^\infty$ error	Order
		P <sup>1</sup>		
$16 \times 16$	$1.11E - 01$		$1.77E - 01$	
$32 \times 32$	$1.84E - 02$	2.59	$3.19E - 02$	2.47
$64 \times 64$	$3.36E - 03$	2.46	$5.66E - 03$	2.50
$128 \times 128$	$7.29E - 04$	2.20	$1.39E - 03$	2.02
		P <sup>2</sup>		
$16 \times 16$	$1.14E - 03$		$8.34E - 03$	
$32 \times 32$	$1.44E - 04$	2.98	$1.21E - 03$	2.78
$64 \times 64$	$1.81E - 0.5$	2.99	$1.58E - 04$	2.94
$128 \times 128$	$2.27E - 06$	3.00	$1.99E - 0.5$	2.99

**Table I.**  $L^2$  and  $L^\infty$  Errors and Orders for  $\rho$  in the First Two-Dimensional Smooth Example with Initial Condition  $\rho(x, y) = 2 + \sin(x + y)$ ,  $u_x = u_y = 1$ ,  $p = 5$  at  $t = 7$ 

### **3.2. Two-Dimensional Examples**

## *3.2.1. Examples with Smooth Solutions*

The first example with smooth solutions is essentially a scalar problem, since there is only one nontrivial component  $\rho$  in the solution. The domain is taken to be  $[0, 2\pi] \times [0, 2\pi]$  and  $\gamma = 2$ . The initial data is given by  $U^p = (\rho_0(x, y), 1, 1, 0, 0, 0, 0, 5)$  with  $\rho_0(x, y) = 2 + \sin(x + y)$ . Periodic boundary conditions are used. We compute until  $t = 7$ , after approximately two time periods. The exact solution for this problem is simply  $\rho(x, y, t)$  =  $\rho_0(x - u_x t, y - u_y t)$  with other components remaining constants.

From Table I, we can see that we obtain second order accuracy for the  $P<sup>1</sup>$  elements, and third order accuracy for the  $P<sup>2</sup>$  elements, both in  $L<sup>2</sup>$ and  $L^{\infty}$  norms, which are the optimal orders we expect. The drawback of this example is that only the hydrodynamical part of the system is tested.

Next we look at a genuinely two dimensional vortex problem, which was originally suggested by Shu [27] in the hydrodynamical system, and was adapted to the MHD equations by Balsara [4]. The solution is a smooth vortex stably convected with the velocity field and the magnetic field. The unperturbed magnetohydrodynamic flow with  $(\rho, u_x, u_y, B_x, B_y, p) = (1, 1, 1, 0, 0, 1)$  is initialized on the computational domain  $[-10, 10] \times [-10, 10]$  with  $\gamma = 5/3$ . The vortex is introduced through the fluctuation in the velocity and magnetic fields given by

$$
(\delta u_x, \delta u_y) = \frac{\eta}{2\pi} \nabla \times \exp\{0.5(1 - r^2)\},\
$$

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$$
(\delta B_x, \delta B_y) = \frac{\xi}{2\pi} \nabla \times \exp\{0.5(1 - r^2)\},
$$

where  $r^2 = x^2 + y^2$ , and the dynamical balance is obtained through the perturbation on pressure by

$$
\delta p = (\xi^2 (1 - r^2) - \eta^2) \frac{1}{8\pi^2} \exp(1 - r^2).
$$

We use  $\eta = 1$ ,  $\xi = 1$  in our computation. Periodic boundary conditions are used. The exact solution is just the one obtained from the initial configuration propagating with speed *(*1*,* 1*)*, or mathematically given by  $U^p(x, y, t) = U_0^p(x - t, y - t).$ 

Table II shows the  $L^2$  errors and orders of accuracy for some representative components at  $t = 20$ . The errors are computed within [−5*,* 5]×[−5*,* 5] which contains the vortex. We can see that for both locally divergence-free piecewise  $P<sup>1</sup>$  elements and standard piecewise  $P<sup>1</sup>$  elements, second order accuracy is obtained. When  $k = 2$ , the locally divergence-free piecewise polynomial spaces give us better results for magnetic fields, compared with the standard  $P^2$  spaces, though both of the solution spaces give us comparable results of accuracy order close to three for all components.

Similar to the Maxwell equations [11], we compute the global divergence (see [11] for the definition) of the magnetic fields of the RKDG solutions using both the standard polynomial spaces and the locally divergence-free spaces in Fig. 3. It shows that in general, the RKDG methods with the locally divergence-free spaces render smaller global divergence.

#### *3.2.2. Two-Dimensional Kelvin–Helmholtz Instability*

The Kelvin–Helmholtz instability can arise when two superposed fluids flow one over the other in a relative speed. Here we will use the discontinuous Galerkin methods to solve the two dimensional periodic and convective models with transverse magnetic field configuration (see [18] and [31] for more details).

The initial stationary configuration is given by  $U^p = (1, u_{x0}, 0, 0, 0, 0, ...)$ 1, 0.5), where  $u_{x0}(x, y) = \tanh(y)$ . At  $t = 0$ , a small perturbation is introduced to  $u_x$  by

$$
\delta u_x = \begin{cases}\n-0.008 \sin(\frac{2\pi x}{\lambda})/(1+y^2) & \text{if } -\frac{\lambda}{2} < x < \frac{\lambda}{2}, \\
0, & \text{otherwise.}\n\end{cases}
$$

The computational domain is  $[-\frac{L}{2}, \frac{L}{2}] \times [0, H]$ .  $\gamma = 2$ . In the periodic model,  $L = \lambda = 5\pi$ ,  $H = 10$ . The periodic boundary condition is used in the

	$\rho$		$u_x$		$B_x$		$\boldsymbol{p}$				
Mesh	Error	Order	Error	Order	Error	Order	Error	Order			
P <sup>1</sup>											
$32 \times 32$	$3.89E - 04$		$2.12E - 02$		$2.11E - 02$		$2.57E - 03$				
$64 \times 64$	$2.96E - 04$	0.39	$8.56E - 03$	1.30	$8.33E - 03$	1.34	$1.22E - 03$	1.07			
$128 \times 128$	$1.13E - 04$	1.39	$2.08E - 03$	2.04	$2.03E - 03$	2.04	$2.92E - 04$	2.06			
$256 \times 256$	$2.80E - 05$	2.01	$4.65E - 04$	2.16	$4.60E - 04$	2.14	$6.04E - 05$	2.28			
$P1-LDF$											
$32 \times 32$	$3.35E - 04$		$2.06E - 02$		$2.05E - 02$		$2.63E - 03$				
$64 \times 64$	$2.47E - 04$	0.44	$8.43E - 03$	1.29	$8.06E - 03$	1.35	$1.28E - 03$	1.04			
$128 \times 128$	$9.99E - 05$	1.31	$2.05E - 03$	2.04	$1.94E - 03$	2.06	$3.04E - 04$	2.07			
$256 \times 256$	$2.63E - 05$	1.93	$4.61E - 04$	2.16	$4.51E - 04$	2.10	$6.18E - 05$	2.30			
$\mathbf{p}^2$											
$32 \times 32$	$2.06E - 04$		$1.04E - 03$		$1.07E - 03$		$3.63E - 04$				
$64 \times 64$	$2.40E - 05$	3.10	$6.30E - 05$	4.05	$9.36E - 05$	3.52	$4.02E - 05$	3.17			
$128 \times 128$	$3.08E - 06$	2.96	$5.80E - 06$	3.44	$1.53E - 05$	2.61	$4.99E - 06$	3.01			
$256 \times 256$	$6.07E - 07$	2.34	$7.78E - 07$	2.90	$2.65E - 06$	2.53	$9.63E - 07$	2.37			
$512 \times 512$	$1.06E - 07$	2.52	$1.16E - 07$	2.75	$4.28E - 07$	2.63	$1.68E - 07$	2.52			
$P^2$ -LDF											
$32 \times 32$	$2.05E - 04$		$1.12E - 03$		$1.14E - 03$		$3.56E - 04$				
$64 \times 64$	$2.59E - 05$	2.98	$6.44E - 05$	4.12	$6.01E - 0.5$	4.24	$4.27E - 0.5$	3.06			
$128 \times 128$	$4.48E - 06$	2.53	$6.48E - 06$	3.31	$5.61E - 06$	3.42	$7.35E - 06$	2.54			
$256 \times 256$	$7.33E - 07$	2.61	$8.87E - 07$	2.87	$6.96E - 07$	3.01	$1.20E - 06$	2.61			
$512 \times 512$	$1.05E - 07$	2.80	$1.19E - 07$	2.89	$8.74E - 08$	2.99	$1.72E - 07$	2.80			

**Table II.** *L*<sup>2</sup> Errors and Orders of Accuracy in the Two-Dimensional Smooth Vortex Example Computed Within  $[-5, 5] \times [-5, 5]$  at  $t = 20$ 

*x* direction, and the characteristic boundary condition is used for the top boundary. At  $y = 0$ , the boundary condition is obtained from the following symmetry: under the transformation  $(x, y) \rightarrow -(x, y)$ ,  $\rho$ ,  $p$ , and  $B_z$  are symmetric and  $u_x$ ,  $u_y$  are antisymmetric.

Since there are more features developed around  $y = 0$ , the Robert transform

$$
y \leftarrow \frac{H \sinh(3y/H)}{\sinh(3)}
$$

is used on the meshes in the *y* direction, which renders the mesh denser near  $y=0$  and coarser near  $y=H$ .

In the following computation,  $B_x$ ,  $B_y$  are always zero and there is not much difference between the results using piecewise  $P^k$  and locally divergence-free piecewise  $P^k$ . In Fig. 4, we show the results at  $t = 144$  with



**Fig. 3.** The divergence of **B** against time t for the two-dimensional smooth vortex example.



**Fig. 4.** Kelvin–Helmholtz instability in the periodic model. 20 equally spaced contours of  $\rho \in [0.82, 1.17]$ . *P*<sup>2</sup> on a 48 × 30 mesh with Robert transform in the *y* direction. *M* = 1.

the TVB parameter  $M = 1$ . One can see the vortex flow generated from  $y = 0$ , off which there is some weak shock developed. Since there is no strong structure in this model, the computation is not very sensitive to the limiter parameter *M*.

Next, we check the numerical convergence of this method by looking at the cuts for  $\rho$  at  $y = (3/10)H$  on uniform meshes:  $48 \times 30$ ,  $80 \times 100$  and  $100 \times 150$ . *M* = 1 and piecewise *P*<sup>2</sup> are used. In Fig. 5, the convergence is



**Fig. 5.** Kelvin–Helmholtz instability in the periodic model. Cuts of  $\rho$  at  $y = (3/10)H$  to see convergence.  $P^2$  on uniform meshes with  $M=1$ . Dotted line:  $48 \times 30$ ; dashed line:  $80 \times 100$ ; solid line:  $100 \times 150$ . Zoom-in plot is on the right.



**Fig. 6.** Kelvin–Helmholtz instability in the periodic model. Time evolution of the transverse kinetic energy to see convergence.  $P^2$  on uniform meshes with  $M = 1$ . Dotted line:  $48 \times 30$ ; dashed line:  $80 \times 100$ ; solid line:  $100 \times 150$ .

seen clearly, even around the rough area near the left boundary. Another way to see the convergence is to look at the time evolution curves of the total transverse kinetic energy over [−*L/*2*, L/*2]×[−*H,H*], which, for this example, is defined as  $\int_{-\frac{L}{2}}^{\frac{L}{2}} \int_{-H}^{H} (\frac{1}{2}\rho u_y^2) dx dy$  (see Fig. 6). All the curves on the given meshes coincide with each other very well.

For the convective model, the same initial profile is used as in the periodic case. In contrast to the periodic model, the domain is enlarged,



**Fig. 7.** Kelvin–Helmholtz instability in the convective model at  $t = 120$ . 15 equally spaced contours for  $\rho \in [0.67, 1.24]$  (top) and for  $p \in [0.24, 0.79]$  (bottom).  $P^2$  with  $M = 0$  on a 528  $\times$ 96 mesh. Robert transform is used in the *y* direction.

and the features developed as time evolves are allowed to convect away freely in the *x* direction. Using the same notation as in the periodic case,  $L = 11\lambda = 55\pi$ ,  $H = 20$ . A 528 × 48 mesh is used with the same Robert transform in the *y* direction as in the periodic model. We observe the formation of larger vortices, and the shocks formed off these vortices are much stronger than the one found in the periodic system. Contour plots for density and pressure at  $t = 120$  and  $t = 145$  are shown in Figs 7 and 8, respectively. The overall features in this calculation are similar to those obtained in [18] and [31].

#### *3.2.3. Two-Dimensional Riemann Problem*

The initial data for this two-dimensional Riemann problem is chosen so that the solutions of the three of the four one-dimensional Riemann problems are simple waves: if one denotes the quadrants by Roman numbers counter-clockwisely starting from  $\{(x, y) : x > 0, y > 0\}$ , we have shocks for the problem II↔III and III↔IV and a rarefaction wave for I↔II (see [15]).



**Fig. 8.** Kelvin–Helmholtz instability in the convective model at  $t = 145$ . 15 equally spaced contours for  $\rho \in [0.44, 1.19]$  (top) and for  $p \in [0.11, 0.73]$  (bottom).  $P^2$  with  $M = 0$  on a  $528 \times 96$  mesh. Robert transform is used in the *y* direction.

The initial condition is given as follows:

$$
\mathbf{U} = \begin{cases}\n(0.9308, 1.4557, -0.4633, & 0.0575, 0.3501, 0.9830, 0.3050, 5.0838) & \text{for } x > 0, \ y > 0 \\
(1.0304, 1.5774, -1.0455, & -0.1016, 0.3501, 0.5078, 0.1576, 5.7813) & \text{for } x < 0, \ y > 0 \\
(1.0000, 1.7500, -1.0000, & 0.0000, 0.5642, 0.5078, 0.2539, 6.0000) & \text{for } x < 0, \ y < 0 \\
(1.8887, 0.2334, -1.7422, & 0.0733, 0.5642, 0.9830, 0.4915, 12.999) & \text{for } x > 0, \ y < 0\n\end{cases}
$$

and the computational domain is taken to be  $[-1, 1] \times [-1, 1]$ .  $\gamma = 5/3$ . For the boundary conditions, instead of using the Dirichlet boundary conditions from numerical solutions of the one dimensional Riemann problems as in Dedner *et al*. [15], we simply use the Neumann boundary conditions.

Contours of  $B_x$  and  $B_y$  obtained with the standard piecewise  $P^2$ elements and locally divergence-free piecewise *P*<sup>2</sup> elements are shown in



**Fig. 9.** Two-dimensional Riemann problem. 15 equally spaced contours of  $B_x \in [0.40, 1.16]$ (top) and of  $B_y \in [0.54, 1.31]$  (bottom) at  $t = 0.2$ . 100 × 100 mesh size with  $M = 1$ . Left column:  $P^2$ ; right column:  $P^2$ -LDF.

Fig. 9. We do not seem to have any problem in keeping  $B_y$  constant across the shock in the II $\leftrightarrow$ III Riemann problem, or causing any strong distortions in the magnetic field components behind the rarefaction wave of the I $\leftrightarrow$ II Riemann problem, as in some methods described in [15]. We also look at the similar cuts as in Fig. 7 from [15], shown in our Fig. 10, and we observe much smaller oscillations in our results.

#### *3.2.4. Orszag–Tang Vortex Example*

In this section, we will consider the development of the Orszag–Tang vortex example [21], which is a widely used test example in the literature because of the complex interaction between several shocks generated as the whole system evolves. Starting from a smooth state, after the transition period, the system will go to turbulence. The initial setup is the same as the one used in [18]



**Fig. 10.** Two-dimensional Riemann problem. Cuts at  $x = 0.93$  and  $t = 0.2$  on  $100 \times 100$  mesh with  $M = 1$ . Dotted line:  $P^2$ -LDF; solid line: one-dimensional reference results. Top left:  $\rho$ ; top right:  $u_x$ ; bottom left:  $B_x$ ; bottom right:  $B_y$ .

$$
\mathbf{U}^{p} = (\gamma^{2}, -\sin y, \sin x, 0, -\sin y, \sin 2x, 0, \gamma)
$$

with  $\gamma = 5/3$ , and the computational domain is  $[0, 2\pi] \times [0, 2\pi]$  with periodic boundary conditions.

The time development of density is shown in Fig. 11. One can see that in the early stage, the solution is quite smooth. At  $t = 2$ , shocks have already appeared. At later times, for example, at  $t = 3, 4$ , the shocks interact with each other and the structure gets quite complicated.

During the computation, we notice that schemes with different solution spaces behave quite differently, in the sense that they are able to keep the computation from blowing up or delay the blowing up. For example, by using the  $192 \times 192$  mesh with  $M = 1$ , computation with piecewise  $P<sup>1</sup>$ elements could reach  $t = 3.45$ , yet using locally divergence-free piecewise  $P<sup>1</sup>$  elements, we could do the computation at least until  $t = 30$  (and it seems that the computation could go on forever). This seems to indicate that the elimination of numerical divergence inside each element by the locally divergence-free discontinuous Galerkin method has enhanced the



**Fig. 11.** Orszag–Tang vortex example. The development of  $\rho$  with time: Top left:  $t = 0.5$ ; top right:  $t = 2$ ; bottom left:  $t = 3$ ; bottom right:  $t = 4$ . 15 equally spaced contours with ranges  $[2.33, 5.58]$ ,  $[0.96, 5.87]$ ,  $[1.56, 5.74]$  and  $[1.33, 5.45]$ , respectively.  $P^2$ -LDF on  $192 \times 192$  mesh with  $M = 1$ .

numerical stability. When  $k = 2$ , we do not have the same luck, but still by using locally divergence-free piecewise  $P^2$  elements, the computation could go on until  $t = 4.14$ , instead of  $t = 3.60$  when the piecewise  $P^2$  elements are used. The instability is probably still due to numerical divergence, as even the locally divergence-free discontinuous Galerkin solution has numerical divergence manifested by the jumps of the normal component of the magnetic field across cell interfaces. Notice that for this example, computation with WENO methods also blows up without divergence correction [18]. This stability problem might also be related to the choice of the limiters (see, e.g. [30]). Further investigations are needed to clarify these issues and these will be left for future work.

For this example, we also compare the results using locally divergence-free and globally divergence-free [11] piecewise polynomial solution

spaces, see Fig. 12, as well as the results using locally divergence-free and standard piecewise polynomial solution spaces, see Fig. 13, by looking at the cuts of pressure at different times when  $y = 0.8\pi$ . One can see that when the solutions are smooth, these solution spaces give almost the same results. Although there are slight differences in the solutions as shocks are developed and interact with each other at later time, the main structures captured by these three solution spaces are still basically the same.

We again, as in the smooth vortex example, monitor the size of the global divergence of the magnetic fields in the RKDG solutions using both the standard polynomial spaces and the locally divergence-free spaces in Fig. 14. Though the actual curves depend on the limiter parameter *M*, in general, the RKDG methods with the locally divergence-free spaces give smaller global divergence.

#### *3.2.5. Shock Reflection*

In this section, the shock reflection example suggested in [1] is considered. This problem was originally constructed such that a 29◦ reflected shock is the equilibrium solution across a Cartesian tube. The setup here is the same as in [15]. We take  $[-1, 1] \times [-0.5, 0.5]$  as the computational domain, and  $\gamma = 1.4$ . Setting

$$
\mathbf{U}_l^p = (1.0, 2.9, 0.0, 0.0, 0.5, 0.0, 0.0, 5/7),
$$

$$
\mathbf{U}_t^p = (1.4598, 2.7170, -0.4049, 0.0, 0.6838, -0.1019, 0.0, 1.2229),
$$

we use  $\mathbf{U}^p = \mathbf{U}_l^p$  as the initial data,  $\mathbf{U}^p = \mathbf{U}_l^p$  and  $\mathbf{U}^p = \mathbf{U}_l^p$  as Dirichlet boundary conditions on top and left, and reflective and outgoing boundary conditions for the lower and right boundaries, respectively.

The results for density at  $t = 2$  are shown in Fig. 15 on a  $100 \times 100$ mesh with  $M = 1$ . The result on the left is from the computation using standard piecewise  $P^2$  elements, which is very oscillatory compared with the one using locally divergence-free piecewise  $P^2$  elements. Similar oscillatory features were also observed in [1] and were attributed to the lack of the divergence cleaning (see also [15]). The oscillatory features using the standard piecewise  $P^k$  elements finally make the computation blow up. For example, for  $k = 1$ , the blowing up happens at t=3.64. Computation using locally divergence-free  $P^k$  could however continue stably.

To see more details, we plot the contours for  $\rho$  at different times in Fig. 16. One could see that at early time, for example at  $t = 0.2$ , when the shock is not fully developed, the two solutions behave quite similarly. There are some small numerical structures around the top left corner,



**Fig. 12.** Orszag–Tang vortex example. The pressure distribution along  $y = 0.8\pi$  on 100  $\times$ 100 mesh with  $M = 1$ . From top to bottom:  $t = 0.5$ ,  $t = 2.0$ ,  $t = 3.0$ ,  $t = 4.0$ . Solid line: globally divergence-free *P*2; dash-dotted line: locally divergence-free *P*2.



**Fig. 13.** Orszag–Tang vortex example. The pressure distribution along  $y = 0.8\pi$  on 100  $\times$ 100 mesh with  $M = 1$ . From top to bottom:  $t = 0.5$ ,  $t = 2.0$ ,  $t = 3.0$ ,  $t = 4.0$ . Solid line: locally divergence-free *P*2; dash-dotted line: standard solution space *P*2.



**Fig. 14.** The divergence of **B** against time t for the Orszag–Tang vortex example on  $100 \times 100$  mesh with  $M = 0$  when  $k = 2$ .



**Fig. 15.** Shock reflection. 15 equally spaced contours of  $\rho \in [0.86, 2.18]$  at  $t = 2$  on  $100 \times 100$ mesh with  $M = 1$ . Left:  $P^2$ ; right:  $P^2$ -LDF.

which is quite common in the computation of such problems. When time goes on, the situation is changed: in the case with the locally divergencefree solution space, those small numerical structures stemming from the corner are slowly smoothed out, and convected away through the right boundary, so the whole computation moves on stably. For the standard piecewise  $P^k$  computation, however, there are small scale structures generated to the right of the shocks, for example at  $t = 0.6$ , which accumulate, form oscillations and make the computation unstable eventually. What we see here is quite similar to what we have seen in the Orszag–Tang vortex example: the locally divergence-free space has some "smoothing" effect in the computation, which could enhance the stability of the scheme. Notice



**Fig. 16.** Shock reflection. Thirty equally spaced contours of  $\rho$  at different times on 100  $\times$ 100 mesh with *M* = 1. From top to bottom:  $ρ ∈ [1.03, 1.50]$  at  $t = 0.2$ ,  $ρ ∈ [1.02, 1.52]$  at  $t = 0.6$ ,  $\rho \in [1.05, 1.92]$  at  $t = 1.0$ , and  $\rho \in [1.05, 2.10]$  at  $t = 1.4$ . Left:  $P^2$ ; right:  $P^2$ -LDF.

there is some boundary-layer-like structure near the top boundary of the domain (see the right column in Fig. 16) which is some intermediate structure in the transition period and will disappear as time goes on. This intermediate feature is related to the actual boundary treatment as well as the solution spaces we are using.

Moreover, the computation using the standard piecewise  $P^k$  is very sensitive to the limiter parameter  $M$ ; the larger  $M$  is, the earlier the computation blows up. As for the computation with locally divergence-free  $P^k$ , the parameter *M* has no apparent influence to the stability.

#### *3.2.6. Rotor*

In this section, we study the rotor example from [30].

The setup of this problem is as follows. There is a dense rotating disk of fluid located in the central area, while the ambient fluid is at rest. A "taper" function is used between these two areas, which helps to reduce initial transient.

The computational domain is taken as  $[0, 1] \times [0, 1]$ .  $\gamma = 5/3$ . The initial condition is given by

$$
\mathbf{U}^{p} = (\rho_0, u_{x0}, u_{y0}, 0, 2.5/\sqrt{4\pi}, 0, 0, 0.5)
$$

with

$$
(\rho_0(\mathbf{x}), u_{x0}(\mathbf{x}), u_{y0}(\mathbf{x})) = \begin{cases} (10, -(y - 0.5)/r_0, (x - 0.5)/r_0) & \text{if } r < r_0, \\ (1 + 9f, -f \cdot (y - 0.5)/r, f \cdot (x - 0.5)/r) & \text{if } r_0 < r < r_1, \\ (1, 0, 0) & \text{if } r > r_1, \end{cases}
$$

where  $r_0 = 0.1$ ,  $r_1 = 0.115$ ,  $f = (r_1 - r)/(r_1 - r_0)$ ,  $r = [(x - 0.5)^2 + (y (0.5)^2$ <sup>1/2</sup>. Note that the rotor is not in equilibrium, since the centrifugal forces are not balanced. The magnetic field will confine the rotating dense fluid into an oblate shape. Periodic boundary conditions are used.

The results at  $t = 0.295$  on a  $100 \times 100$  mesh with  $M = 10$  are shown in Fig. 17, which are the density *ρ*, pressure *p*, hydrodynamic Mach number  $|\mathbf{u}|/c$  with sound speed  $c = \sqrt{\gamma p/\rho}$ , and also the magnetic pressure  $\mathbf{B}^2/2$ . The solutions are very well resolved (see the results in [30] for a reference).

Similarly to what the authors of [5] and [30] have observed, when we zoom in the central part of Mach number obtained by using the standard piecewise  $P^k$  (left column in Fig. 18), we see some "distortion" around the central rotating area, for example near the points (0.40, 0.45) and (0.60, 0.55). Besides, we also find some extra features around the points (0.50, 0.35) and (0.50, 0.65). We say these features are "extra" since they are not observed in [5] and [30], nor in our results using the locally divergence-free  $P<sup>k</sup>$  (right column in Fig. 18). There is no distortion in our locally divergence-free solutions, which are quite comparable with the "good" results in [5] and [30]. This indicates that using the locally divergence-free space can reduce nonphysical features in the computation.

We also plot cuts from Mach number to see the convergence of the methods and the different behavior of the two solution spaces, in Fig. 19. These cuts are at  $x = 0.413$  and  $x = 0.545$ . It can be seen that the results using locally divergence-free solution space behave better in convergence, and they give relatively smoother solutions.



**Fig. 17.** Rotor. From top to bottom: 15 equally spaced contours of density *ρ* ∈ [0*.*730*,* 7*.*330]; pressure *p* ∈ [0*.*058*,* 0*.*733]; Mach number |**u**|*/c* ∈ [0*.*147*,* 2*.*270]; magnetic pressure  $\mathbf{B}^2/2 \in [0.059, 0.655]$ . At  $t = 0.295$  on  $100 \times 100$  mesh with  $M = 10$ . Left:  $P^2$ ; right:  $\overline{P^2$ -LDF.



**Fig. 18.** Rotor. Zoom-in central part for Mach number. 30 equally spaced contours with range [0.18, 3.12] at  $t = 0.295$  with  $M = 10$ . From top to bottom:  $100 \times 100$  mesh,  $200 \times 200$ mesh, 400×400 mesh, 800×800 mesh. Left: *P*2; right: *P*2-LDF.



**Fig. 19.** Rotors. Cuts of Mach number. Left column: at  $x = 0.413$ ; right column: at  $x =$ 0.545. The first two rows are for convergence: dotted lines are for  $400 \times 400$  mesh, solid lines are for  $800 \times 800$  mesh; the top row is for  $P^2$ , the middle row is for  $P^2$ -LDF. The last row is for comparison of  $P^2$  and  $P^2$ -LDF: dotted lines are for  $P^2$ , solid lines are for  $P^2$ -LDF.  $M = 10$ .

#### **4. CONCLUDING REMARKS**

Discontinuous Galerkin method using locally divergence-free piecewise polynomial solution spaces seems to be very effective for solving the ideal MHD equations. It produces results no worse than the ones obtained from the discontinuous Galerkin methods using the standard piecewise polynomial spaces with the same polynomial degree, yet has a smaller computational cost. More importantly, this locally divergence-free discontinuous Galerkin method could enhance the numerical stability and reduce some nonphysical features in the solutions.

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## **REFERENCES**

- 1. Aslan, N., and Kammash, T. (1997). Developing numerical fluxes with new sonic fix for MHD equations. *J. Comput. Phys.* **133**, 43–55.
- 2. Baker, G. A., Jureidini, W. N., and Karakashian, O. A. (1990). Piecewise solenoidal vector fields and the Stokes problem. *SIAM J. Numer. Anal.* **27**, 1466–1485.
- 3. Balsara, D. S. (2001). Divergence-free adaptive mesh refinement for magnetohydrodynamics. *J. Comput. Phys.* **174**, 614–648.
- 4. Balsara, D. S. (2004). Second order accurate schemes for magnetohydrodynamics with divergence-free reconstruction. *Astrophys. J. Suppl. Ser.* **151**, 149–184.
- 5. Balsara, D. S., and Spicer, D. S. (1999). A staggered mesh algorithm using high order Godunov fluxes to ensure solenoidal magnetic fields in magnetohydrodynamic simulations. *J. Comput. Phys.* **149**, 270–292.
- 6. Brackbill, J. U., and Barnes, D. C. (1980). The effect of nonzero ∇ · **B** on the numerical solution of the magnetohydrodynamic equations. *J. Comput. Phys.* **35**, 426–430.
- 7. Brecht, S. H., Lyon, J. G., Fedder, J.A., and Hain, K. (1981). A simulation study of eastwest IMF effects on the magnetosphere. *Geophys. Res. Lett.* **8**, 397–400.
- 8. Brio, M., and Wu, C. C. (1988). An upwind differencing scheme for the equations of ideal magnetohydrodynamics. *J. Comput. Phys.* **75**, 400–422.
- 9. Cockburn, B., Hou, S., and Shu, C.-W. (1990). The Runge–Kutta local projection discontinuous Galerkin finite element method for conservation laws IV: The multidimensional case. *Math. Comput.* **54**, 545–581.
- 10. Cockburn, B., Lin, S.-Y., and Shu, C.-W. (1989). TVB Runge–Kutta local projection discontinuous Galerkin finite element method for conservation laws III: One dimensional systems. *J. Comput. Phys.* **84**, 90–113.
- 11. Cockburn, B., Li, F., and Shu, C.-W. (2004). Locally divergence-free discontinuous Galerkin methods for the Maxwell equations. *J. Comput. Phys.* **194**, 588–610.
- 12. Cockburn, B., and Shu, C.-W. (1998). The Runge–Kutta discontinuous Galerkin method for conservation laws V: Multidimensional systems. *J. Comput. Phys.* **141**, 199–224.
- 13. Cockburn, B., and Shu, C.-W. (2001). Runge–Kutta Discontinuous Galerkin methods for convection-dominated problems. *J. Sci. Comput.* **16**, 173–261.
- 14. Dai, W., and Woodward, P. R. (1998). A simple finite difference scheme for multidimensional magnetohydrodynamic equations. *J. Comput. Phys.* **142**, 331–369.
- 15. Dedner, A., Kemm, F., Kröner, D., Munz, C.-D., Schnitner, T., and Wesenberg, M., (2002). Hyperbolic divergence cleaning for the MHD equations. *J. Comput. Phys.* **175**, 645–673.
- 16. Evans, C. R., and Hawley, J. F. (1988). Simulation of magnetohydrodynamic flows: A constrained transport method. *Astrophys. J.* **332**, 659–677.
- 17. Jiang, B.-N., Wu, J., and Povinelli, L. A. (1996). The origin of spurious solutions in computational electromagnetics. *J. Comput. Phys.* **125**, 104–123.
- 18. Jiang, G.-S., and Wu, C.-C. (1999). A high-order WENO finite difference scheme for the equations of ideal magnetohydrodynamics. *J. Comput. Phys.* **150**, 561–594.
- 19. Karakashian, O. A., and Jureidini, W. N. (1998). A nonconforming finite element method for the stationary Navier–Stokes equations. *SIAM J. Numer. Anal.* **35**, 93–120.
- 20. Munz, C.-D., Omnes, P., Schneider, R., Sonnendrücker, E., and Voβ, U. (2000). Divergence correction techniques for Maxwell solvers based on a hyperbolic model. *J. Comput. Phys.* **161**, 484–511.
- 21. Orszag, S. A., and Tang, C.-M. (1979). Small-scale structure of two-dimensional magnetohydrodynamic turbulence. *J. Fluid. Mech.* **90**, 129–143.
- 22. Powell K. G. (1994). *An Approximate Riemann solver for Magnetohydrodynamics (that works in more than one dimension)*, ICASE report No. 94–24, Langley, VA.
- 23. Qiu, J., and Shu, C.-W. Runge–Kutta discontinuous Galerkin method using WENO limiters. *SIAM J. Sci. Comput.*, to appear.
- 24. Qiu, J., and Shu, C.-W. (2003). Hermite WENO schemes and their application as limiters for Runge–Kutta discontinuous Galerkin method: One dimensional case. *J. Comput. Phys.* **193**, 115–135.
- 25. Ryu, D., Miniati, F., Jones, T. W., and Frank, A. (1998). A divergence-free upwind code for multi-dimensional magnetohydrodynamic flows. *Astrophys. J.* **509**, 244–255.
- 26. Shu, C.-W. (1987). TVB uniformly high-order schemes for conservation laws. *Math. Comput.* **49**, 105–121.
- 27. Shu, C.-W. (1998). Essentially non-oscillatory and weighted essentially non-oscillatory schemes for hyperbolic conservation laws. In Cockburn, B., Johnson, C., Shu, C.-W., and Tadmor, E., Quarteroni, A. (eds.), Advanced Numerical Approximation of Nonlinear Hyperbolic Equations. Lecture Notes in Mathematics, Springer, Berlin, Vol. **1697**, 325–432.
- 28. Shu, C.-W. and Osher, S. (1988). Efficient implementation of essentially non-oscillatory shock-capturing schemes. *J. Comput. Phys.* **77**, 439–471.
- 29. Stone, J. M., and Norman, M. L. (1992). ZEUS-2D: A radiation magnetohydrodynamics code for astrophysical flows in two space dimensions. II The magnetohydrodynamic algorithms and tests. *Astrophys. J. Suppl. Ser.* **80**, 791–818.
- 30. Toth, G. (2000). The  $\nabla \cdot \mathbf{B} = 0$  constraint in shock-capturing magnetohydrodynamics codes, *J. Comput. Phys.* **161**, 605–652.
- 31. Wu, C. C. (1986). An Kelvin–Helmholtz instability at the magnetopause boundary. *J. Geophys. Res.* **91**, 3042–3060.
- 32. Yee, K. S. (1966). Numerical solution of initial boundary value problems involving Maxwell's equations in isotropic media. *IEEE Trans. on Antenna Propagation* **AP-14**, 302–307.