

Analysis of Global and Local Hydrodynamic Instabilities on a High‑Speed Jet Difusion Flame via Time‑Resolved 3D Measurements

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Abstract

Measurements of 3D fame edge dynamics were made on a high-speed jet difusion fame to assess the global/local hydrodynamic instability. The fame was generated by issuing highspeed ethylene (U_i =170 m/s) into a low-speed vitiated hot coflow (U_c =1.5 m/s), resulting in a hydrodynamic shear layer instability at the interface between combustion products and ambient fow. The measurements used a high-speed camera combined with nine-headed fber endoscopes to simultaneously collect both the soot radiation and chemiluminescence projections of the fame from nine views, based on which 3D fame edges were obtained via computed tomography at 15 kHz. The measurements clearly capture the time-varying, 3D instantaneous fame edge structures with fne-scale corrugations, enabling the observation of small-scale vortices' evolution. The fame edge deformations induced by those vortices were calculated globally and locally to infer the relationship between global and local fame edge oscillations. Results show that various local oscillation frequencies exist at different locations along the fow direction for such a highly sheared fame. They are dominated by the periodical formation and motion of various localized, small-scale vortices. The local oscillations are much severer than the global oscillation, indicating a self-suppression of the instabilities between these local oscillations. The suppression mechanism is attributed to the constructive and destructive interference behavior of the local disturbances. The global oscillation of the fame edge turns out to be the linear superposition of local oscillations at diferent locations.

Keywords 3D measurements · Flame edge dynamics · Global and local hydrodynamic instabilities · Non-premixed · Computed tomography

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1 Introduction

Combustion instabilities remain a critical issue that limits the development of a wide operating range, low-emission combustion systems. One of the essential mechanisms that induce and sustain unstable combustion is the coupling of acoustic instability with self-excited hydrodynamic instability. Hydrodynamic instability is associated with various factors, including the buoyancy, shear layer, density and fraction gradients, etc., which poses a challenging feld of research.

Experimental measurements to characterize the hydrodynamic instability are desired. Of particular interests are the measurements of fame edge dynamics since the timevarying fame edge behaviors are strongly related to the fame-vortex interaction and thus can provide insights regarding the phenomenology of the hydrodynamic instability of the fame (Lee and Santavicca [2003;](#page-20-0) Kolhe and Agrawal [2007;](#page-20-1) Schlimpert et al. [2016](#page-20-2)). The major challenges of the experimental investigations are thus attributed to the time-varying nature of these fows. Past eforts for the measurements of fame edge dynamics have been primarily based on the 2D techniques, such as the luminescence/ chemiluminescence of self-emissions of combustion products (Shin et al. [2011](#page-20-3); Allison et al. [2017\)](#page-19-0), planar Mie-scattering of seeded tracer particles (Alabdeli et al. [2006;](#page-19-1) Shanbhogue et al. [2009a;](#page-20-4) Long and Hargrave [2011](#page-20-5)), Schlieren imaging (Kolhe and Agrawal [2007](#page-20-1)) and planar laser-induced fuorescence (PLIF) (Steinberg et al. [2010;](#page-21-0) Allison et al. [2015](#page-19-2); Doleiden et al. [2019\)](#page-19-3). In these techniques, the 2D fame edge contours are frst extracted based on the measured signal distributions, and their motions are then tracked shot-to-shot to infer the fame fuctuation behaviors.

However, the major disadvantage of these techniques is that they only provide incomplete information on the fame edge motion. For example, the line-of-sight nature of the luminescence/chemiluminescence technique prevents it from resolving localized phenomena or spatially resolved fame edge structure. Planar Mie scattering and PLIF can overcome the line-of-sight limitation, but they cannot provide the fame oscillation information out of the measurement plane. This disadvantage becomes severer for the conditions where various local fame oscillations exist in the fame at diferent locations. For instance, Kolhe and Agrawal ([2007](#page-20-1)) found that in a jet difusion fame, the instabilities developed in the shear layer between the fame edge and outer vortices vary along the axial direction resulting in diferent local dominant frequencies. Subsequent numerical studies (Nichols and Schmid [2008;](#page-20-6) Qadri et al. [2015](#page-20-7)) further identified multiple types of instability modes that originated from diferent locations of the fame that govern the global instability dynamics. This requires time-resolved 3D measurements further to quantify the relationship between the local and global instabilities and provide validation data to numerical results.

Fortunately, extending the capability of the above-mentioned 2D techniques to 3D measurements becomes possible with the recent advancement in lasers, cameras, and computing techniques. The strategies include laser sheet scanning (Cho et al. [2014;](#page-19-4) Weinkauff et al. [2015\)](#page-21-1) and the computed tomography (CT). Notably, recently, a significant amount of work has demonstrated the potential of the CT strategy combined with luminescence/chemiluminescence (Floyd et al. [2011;](#page-19-5) Worth and Dawson [2012;](#page-21-2) Moeck et al. [2013;](#page-20-8) Li and Ma [2015;](#page-20-9) Huang et al. [2019\)](#page-20-10), Mie scattering (Upton et al. [2011](#page-21-3); Lei et al. [2014](#page-20-11)), Schlieren (Ishino et al. [2015;](#page-20-12) Nicolas et al. [2015](#page-20-13)), and LIF (Ma et al. [2017;](#page-20-14) Li et al. [2017](#page-20-15); Halls et al. [2017](#page-20-16); Pareja et al. [2019\)](#page-20-17) to measure the 3D fame surface properties with sufficient temporal resolution (i.e., up to 20 kHz Ma et al. [2016\)](#page-20-18).

However, to date, most demonstrations of CT fame measurements are dedicated to improving the CT technique itself. The post-processing and analysis of the retrieved timeseries of 3D fame data to understand combustion phenomena are rare. Wiseman et al. ([2017\)](#page-21-4) have developed a series of post-processing algorithms for measuring fame surface area, fame curvature, fame thickness, and the normal component based on the reconstructed 3D data of a weakly turbulent premixed fame. In this work, we aim to show how the measured time-varying 3D data can be utilized to obtain useful information such as fame edge deformations to study local/global hydrodynamic instability on a high-speed jet difusion fame. Another aim is to identify how the local oscillations at diferent locations afect the global oscillation of the fame and to complement the theory of fame transfer functions (FTF) or fame describing functions (FDF) (Candel et al. [2014](#page-19-6); Gatti et al. [2019;](#page-19-7) Gaudron et al. [2019](#page-19-8)).

With the above understandings, this work reports an application of CT technique combined with luminescence to resolve the 3D fame edge dynamics of a high-speed jet diffusion fame at 15 kHz, based on which the relationship of the local and global hydrodynamic instabilities is identifed. The measurements were obtained using a single camera with a nine-headed fber bundle, an approach recently developed to signifcantly facilitate the implementation of the CT technique (Anikin et al. [2010\)](#page-19-9). This work aims to address two crucial aspects in the tomographic measurements of turbulent fames: (a) the validity and capability using a single camera combined with fber endoscopes to resolve the 3D structures and dynamics of high-speed, highly turbulent fames beyond tens of kilohertz; (b) using the 3D data (i.e., tracking the 3D fame edge) to interpret the hydrodynamic instabilities in such highly sheared fames.

In the rest of this paper, Sect. [2](#page-2-0) describes the experimental setup, including details of the burner used in this work, testing conditions, and the hardware used in the 3D diagnostics. Section [3](#page-4-0) describes the 3D tomographic reconstruction methods and validation results. Section [4](#page-8-0) discusses the 3D measurement of the fame edge dynamics and the analysis of the local and global hydrodynamic instabilities. Section [5](#page-18-0) then summarizes and concludes the paper.

2 Experimental Setup

2.1 Burner and Operating Conditions

Figure [1](#page-3-0)a shows the schematic of the burner tested in this study. The burner was designed to generate a high-speed, highly sheared jet difusion fame. The specifcs of this burner have been described in our previous work (Liu et al. [2019a\)](#page-20-19), and only a summary is provided here. The burner consisted of a primary central tube (1 mm ID), from which issued pure C_2H_4 . The central tube was surrounded by an 80 mm ID hot coflow. The hot coflow was generated by a lean premixed C_2H_4/a ir flame stabilized on an annular stainless-steel honeycomb. The equivalence ratio of the coflow was 0.48, and the velocity of the coflow was 1.5 m/s. The tip of the central nozzle was 20 mm above the honeycomb. To generate the turbulent flames, pure C_2H_4 was continuously issued from the central tube at 1.2 atm into the hot coflow, resulting in a jet exit velocity of $U=170$ m/s and jet Reynolds number of Re=22,516. At this specific operating condition, the overall flame length was \sim 27 cm, as shown in the time-averaged fame luminescence image in Fig. [1](#page-3-0)b. The white box in Fig. [1b](#page-3-0) centered at $Z=16$ cm represents the field-of-view in the subsequent CT 3D measurements.

2.2 3D Measurement System

Figure [2](#page-3-1) shows the schematic of the experimental system to obtain instantaneous 3D fame measurements. As mentioned in the operating condition of the burner, the target fames were highly turbulent and high-speed. Therefore, the projection measurements from various orientations were captured simultaneously and instantaneously. As shown in Fig. [2,](#page-3-1) a customized fber bundle and one CMOS camera (iX-speed 720) were applied to accomplish such projection measurements. More specifcally, the customized fber bundle consisted of nine inputs, which were then assembled into one output. Each input was an assembly of 90,000 single-mode fibers in an array of 300×300 , so the output end transmitted a total of $810,000$ ($9 \times 90,000$) image elements. The fiber endoscopes and the camera were aligned so that one image element transmitted by one fber approximately corresponded to one pixel on the camera. In the experiments, each fber endoscope input was equipped with an AF Nikon lens $(50 \text{ mm } f/1.4D)$. The nine fiber bundles were installed on an annular rail so that they were all at the same co-plane and with the same objective distance to the target fame. The signal from all fber endoscopes at the output end was then recorded on the same CMOS array of the camera. The camera was equipped with another AF Nikon lens (50 mm f/1.4D) and a broadband filter (i.e., 400–700 nm). The cameras were operated at a frame rate of 15 kHz and an exposure time of 66.12 μs.

To better illustrate the geometry parameters of the setup, the following Cartesian coordinate system was established as follows. The origin was defned at the center of the burner, the X-axis was defned along the optical axis of fber endoscope 1, the Y-axis was defned to be perpendicular to the X-axis, and the Z-axis was defned to be along the direction of the fow. Based on this defnition, the angles formed between the optical axis of each fiber endoscope and the positive X-axis were 0° , 16.6°, 35.0°, 52.4°, 74.6°, 93.0°, 110.4°, 124.2°, 142.0°, respectively, determined by an improved camera calibration method (Xia et al. [2019\)](#page-21-5). The objective distance of all fber endoscopes to the center of the burner was 39.3 cm. These measured projections were then used as the inputs to an inversion algorithm as described in Sect. [3](#page-4-0) to reconstruct the 3D luminescence feld of the target fame.

It is worth noting that with the above-mentioned measurement scheme and the difusion nature of the ethylene fame, the signal measured contains both the thermal radiation of soot and the chemiluminescence in the visible range. However, as the goal of this work is to investigate the hydrodynamic shear layer instability, the fame edge dynamics extracted based on the luminescence signals (soot thermal radiation) can also represent the vortexinduced hydrodynamic instability of the fame for the following reasons. First, even though there are deviations in the position of chemical radicals (i.e., OH/OH*) and soot particles, their motion/fuctuation behaviors should be similar since they both fuctuate as a response to the local vortex (i.e., the vortex formation, rolling and propagation). The time scales for both soot formation and chemical reaction are signifcantly faster compared to the vortex propagation time scales. It is demonstrated in Katta et al. ([2009\)](#page-20-20) that in an ethylene jet diffusion flame, the magnitude of flame oscillations decreases with the amount of soot generated in the fame; however, the frequency of the oscillations does not change. The same oscillation frequency of soot particles and other combustion radicals indicates that they respond to the same vortex of the turbulent fow feld. The suppression of the oscillation magnitude by the increased soot concentration is, however, due to the cooling efects of thermal radiation. In the present work, with the injecting velocity of 170 m/s, the Froude number is calculated to be larger than 10. When the Froude number is larger than 5, the jet fame oscillation is suggested to be dominated by the momentum (Delichatsios [1993](#page-19-10)). Thus, a stronger response of the combustion products (whether they are intermediate radicals or soot particles) is expected to the local vortex. Second, we have performed a test experiment to examine the oscillation behaviors of the soot particles and the OH* radicals. In the experiment, two high-speed cameras (equipped with appropriate flters) were used to measure the simultaneous soot radiation signals and OH* chemiluminescence, respectively. Results show that although the soot radiation signals and OH* chemiluminescence are with deviation in the position, the oscillating frequencies of fame edge obtained based on two signals at diferent downstream locations are almost identical. We cannot perform the 3D OH* chemiluminescence measurement since the current fber endoscopes do not transmit the ultraviolet signals.

3 3D Tomographic Reconstruction and Validation

3.1 Mathematic Formulation

Figure [3](#page-5-0) shows the mathematical formulation of a typical 3D tomographic problem of turbulent fames. In the problem, the 3D distribution of the sought fame property is discretized into voxels, denoted as $f(x, y, z)$. An imaging lens collects the 2D line-of-sight

Fig. 3 Mathematical formulation of a reacting fow tomography problem

integrated emission signals of the fame from diferent orientations. Those 2D images are named as projections and denoted as $P_{\theta}(s)$. The measured $P_{\theta}(s)$ depends on the viewing angle *θ*, which represents the orientation of *f* with respect to the acquisition plane; *s* denotes the projection coordinate. In experimental measurement setups that the dimension of the flame is much smaller than the objective distance, the measured $P_\theta(s)$ corresponds to parallel lines perpendicular to the Z-axis. Thus, $P_{\theta}(s)$ is known as the Radon transform of $f(x, y, z)$ and can be formulated as:

$$
P_{\theta}(s) = R\{f(x, y, z)\} = \int_{\psi(\theta, s)} f(x, y, z)dl
$$
 (1)

where *R* denotes the Radon transform, and $\psi(\theta, s)$ is a parameter representing the line sight with a given viewing angle θ and the distance of the line to the origin. The imaging model used to compute the sensitivity matrix in this work is the parallel beam imaging model. Generally, the use of parallel beam projection theory is an approximation since the camera collects light in a fan-shaped manner. When the object distance relative to the object diameter is large, the rays can be assumed to be parallel (Floyd et al. [2009](#page-19-11)). In the present case, the object distance (39.3 cm) is sufficiently larger than the object diameter (6.4 cm) . Thus, it can be assumed that the error resulted from the parallel beam assumption is negligible.

The goal of the above tomography problem is to solve $f(x, y, z)$ with measured $P_{\theta}(s)$. Suppose the projections are sufficient (i.e., hundreds) and are sampled continuously around the flame, as in the medical tomography problem. In that case, the $f(x, y, z)$ may be reconstructed analytically using the Fourier slice theorem, which states that the Fourier transform of $P_{\theta}(s)$ with respect to *s* at a constant z_0 equals to the 2D Fourier transform of $f(x,$ *y*, z_0) along s with constant viewing angle θ (Kak et al. [2002\)](#page-20-21). Unfortunately, in turbulent fame tomography, it is impractical to obtain many such projections. In turbulent fame tomography, the projections are sampled discretely from a few viewing orientations, and the *f* can be reconstructed using an algebraic iteration technique in this case. An algebraic algorithm, named the multiplicative algebraic reconstruction technique (MART), is used to perform the tomography reconstruction in the current work. Mohri et al. ([2017\)](#page-20-22) have conducted a thorough study to evaluate a similar algebraic iterative algorithm (e.g., ART) on the tomographic reconstruction of highly turbulent swirl fames. They concluded that in the most sparsely sampled cases, e.g., when N_a (# of views) = 3 and N_a = 4, decreasing the total number of unknowns, or voxels, actually aids the system to generate better results. They thus predicted that even a small number of views (e.g., 4) appears to be sufficient to observe larger asymmetries in the flame, such as those caused by coherent vortex cores or maybe even thermoacoustic fuctuations with a suitably reduced number of voxels. In the present work, the ratio of the total number of voxels (e.g., $150 \times 150 \times 150$) to the total number of pixels (e.g., $9 \times 300 \times 300$) is ~4, which is in a safe range for 9 views to yield a good reconstruction according to Mohri et al. ([2017\)](#page-20-22). The domain dimensions are $6.4 \times 6.4 \times 6.4$ cm³ with a grid spacing of ~0.43 mm. The termination criterion, as described in Mishra et al. [\(2004](#page-20-23)), is employed in this work to control the termination of convergence in MART algorithm.

3.2 Validation of the Method

This work presents two kinds of validation for the tomography algorithm to reconstruct the 3D structure of the highly turbulent fame. The frst validation involves a numerical phantom test. The second validation involves the reconstruction of 3D fame using partial experimental measured projections, the reproduction of projection which is not used in the foregoing reconstruction, and the comparison between the reproduced and measured projections. This section describes the frst validation. The second validation will be illustrated with the experimental results in Sect. [4.1](#page-8-1).

In the phantom test, the 3D OH radical distribution of the same jet difusion fame computed by the large eddy simulation (LES) of our previous work (Liu et al. [2019a](#page-20-19)) was used as the phantom. Projections of this phantom were calculated according to Eq. ([1](#page-5-1)) with the same nine viewing orientations as in the experimental setup of Fig. [2.](#page-3-1) The chain branching reaction $H + O_2 \rightarrow OH + O$ is used for the LES-calculations of OH. Though the OH distribution used in the phantom study does not match with the reconstructed feld in the experiments (i.e., soot radiation), it is assumed they exhibit the fame structure at the same turbulence level. And the purpose of the LES phantom study is to assess and validate the capability of the CT technique to resolve the fame structure at a similar turbulence level to the experiment.

A total of 0.6° of the view registration (VR) error, a major source of noise for combus-tion tomography (Liu et al. [2019b\)](#page-20-24), was randomly added to the projections. The reconstruction is shown and compared with the phantom in the second row of Fig. [4.](#page-7-0) In the second set of validation, both 0.6° of view registration (VR) error and 2% of Gaussian noise (estimated according to the signal-to-noise ratio) were artifcially added to the projections to further quantify the efect of the VR error and imaging noise on the reconstruction. The reconstruction and comparison are shown in the third row of Fig. [4](#page-7-0). In Fig. [4](#page-7-0), the comparison from three sets of horizontal slices of the original phantom and reconstructions is shown. The comparison at those horizontal planes allows a critical assessment of the reconstruction quality, as suggested in Mohri et al. [\(2017](#page-20-22)). To quantify the reconstruction quality, the overall reconstruction error (e) and correlation coefficient (r) were calculated as:

$$
e = \frac{\sum_{n=1}^{N} |f_n^{rec} - f_n|}{\sum_{n=1}^{N} f_n} \text{ and } r = \frac{\text{cov}(f_f^{rec})}{\sigma_f \sigma_{f^{rec}}} \tag{2}
$$

Fig. 4 Comparison of the phantom and reconstruction at three sets of horizontal slices

where *n* is the voxel index, *N* the total number of voxels; *f* represents the known phantom, *f rec* the reconstruction; cov and *σ* represent the convolution and standard derivation, respectively. Good agreement can be observed between the original and reconstructed fame structures with both VR error and imaging noise considered. In addition, the reconstruction error resulting from Gaussian noise is below half of that resulted from view registration uncertainty, indicating that the angle calibration uncertainty brings primary reconstruction error. The Gaussian noise induces little degradation in the correlation coefficient.

Figure [5](#page-8-2) further shows the overall reconstruction error and correlation coefficient along the axial direction with both the VR error and image noises considered. As shown, the reconstruction fdelity obviously decreases (*e* increases and *r* decreases) with height above the burner, indicating larger deviations from the perfect reconstruction. This is due to the increased corrugation and wrinkling of the downstream fame structure. The overall reconstruction accuracy is within acceptance (with *e* below 10% and *r* larger than 97%). Note that, in the phantom validation, the fame extends all the way to the last voxel row on the top of the computational domain. Thus, a dip in error observed at the largest height is due to the stochastic nature of turbulent fows. The error of the reconstruction in the black background region is below 1% by calculating the error for vertical slices. Results in Figs. [4](#page-7-0) and [5](#page-8-2) demonstrate the capability and robustness of the CT system in the present work to resolve both the large-scale structures and details of the target fame with acceptable accuracy when noisy projections are involved.

Based on the pure tomography algorithm validation, the sensitivity of the discretization scheme on the fame edge deformation is also evaluated. More specifcally, a

high-resolution LES phantom (with a dimension of $256³$ determined by the maximum usage of 128G RAM of the workstation) is used to generate the nine projections with the same orientations in the experiments. Those projections are used to reconstruct three sets of fame distribution: high-resolution reconstruction (with the same dimension to phantom), low-resolution reconstruction (with the dimension of $150³$), and low-resolution reconstruction (i.e., $150³$) with noise (including both VR error and image noise). The fame edge deformation between two consecutive time instances is calculated and compared for phantom and reconstructions. Comparison shows good agreement (i.e., $e=1.74\%$ for high-resolution reconstruction, $e=5.19\%$ for low-resolution reconstruction, and *e*=5.65% for low-resolution reconstruction with noises).

Further validation of the CT technique is illustrated in Sect. [4](#page-8-0) by comparing the particular features of the fame between the reconstruction and the raw projection measurement.

4 Experimental Results and Discussion

4.1 Tomographic Reconstructions

After the description and numerical validation of the tomography reconstruction, this section shows the experimental demonstration and validation of tomographic reconstruction on the jet difusion fame. First, Fig. [6](#page-9-0) shows a set of measured raw projection examples containing the soot radiation and visible chemiluminescence captured by all nine fber endoscopes and the camera. The camera captured nine projections transmitted by fber endoscope inputs 1 through 9 simultaneously, with the projection transmitted by each fber endoscope occupying approximately one-ninth of the camera chip (i.e., $\sim 300 \times 300$ pixels). The only pre-processing step performed on these raw images before reconstruction is the signal normalization of nine projections based on the linearity calibration of signal transmission of nine fber endoscopes. As shown in Fig. [6,](#page-9-0) the fame is highly turbulent and has a distinctly diferent structure in each view, illustrating the 3D nature of the fame structure and the necessity to resolve it.

Fig. 6 A sample of instantaneous projection measurement containing the soot radiation and visible chemiluminescence captured by nine-headed fber endoscopes from nine diferent perspectives

The projections shown in Fig. [6](#page-9-0) were fed into the subsequent tomography reconstruction program using the MART algorithm to retrieve the 3D instantaneous luminescence distribution of the fame. The reconstruction results are shown in Fig. [7](#page-10-0) from four diferent viewing angles. As seen, the reconstruction successfully resolves both the overall fame structure and also the detailed turbulent wrinkles (as highlighted in the ovals). For example, the measured projection on the seventh fiber endoscopes shown in Fig. 6 (F-7) indicates that the fame has an approximate ''7" shape when viewed from this particular orientation. The 3D reconstruction shown in Fig. [7c](#page-10-0) captured the overall "7" shape. Furthermore, the measured projection reveals fame holes and the fame tongue at various locations, as highlighted in the ovals. These holes and the fame tongue are successfully captured by the 3D reconstruction. Besides, several vertical and horizontal slices of reconstruction are shown in Fig. [7e](#page-10-0) to further exhibit the spatially resolved distribution of the reconstructed feld.

Figure [8](#page-10-1) shows a more detailed comparison of the specifc feature of the fame tongue, as seen in Fig. [6](#page-9-0) highlighted with ovals to further validate the 3D reconstruction. This flame tongue can be seen in the 1st and 2nd projections in Fig. [6](#page-9-0) of the luminescence from the fame. Figure [8](#page-10-1)a shows a zoom-in raw image of this feature from the 1st projection. The reconstruction of this feature is shown in Fig. [8b](#page-10-1), c from the side and the top view, respectively. It can be seen from Fig. [8b](#page-10-1) that the thickness of the reconstructed feature in the axial direction (1.29 mm) is a good representation of its thickness in the raw data (0.85 mm).

Fig. 7 First row: 3D reconstructions of fame from four diferent views; second row: several vertical and horizontal slices of the reconstruction

Fig. 8 The fame tongue as seen **a** in the original data from fber endoscope 1, **b** in the reconstruction viewing along the *X* axis, and **c** in the reconstruction from the top

Though the spatial resolution of 3D reconstruction cannot compete with the raw 2D data, it provides 3D information in all three directions. For instance, Fig. [8](#page-10-1)c shows the structure and size of the fame tongue from the top view. However, such information cannot be revealed from the original 2D projection data.

Though comparisons in Figs. 7 and 8 are not sufficient to prove the correctness of the reconstruction, they provide a necessary condition. In order to further quantitatively

Fig. 10 Flame edge contours in planes perpendicular to the burner centerline. Heights above the burner exit are indicated. The edge contour of the fame tongue is observed at *Z*=18.2 cm

validate the method, the tomographic reconstruction was performed again using the frst eight (i.e., F-1 to F-8) projections in Fig. [6](#page-9-0). The 9th projection was reproduced based on Eq. ([1](#page-5-1)) and then was compared with the 9th measured projection. The comparison was shown in Fig. [9.](#page-11-0) Good agreement can be seen from the measured and reproduced projections. The correlation coefficient was calculated to be larger than 98.5%, further validating the high fdelity of the reconstruction technique.

Based on the 3D reconstruction in Fig. [7,](#page-10-0) Fig. [10](#page-11-1) shows the horizontal fame edge contours extracted from the 3D reconstruction shown in Fig. [7](#page-10-0). Note that it is usually more diffcult to capture such multiple horizontal slices of fame simultaneously at diferent heights with conventional 2D measurements. In this work, the fame edge contour was determined based on a thresholding method (Ma et al. [2016](#page-20-18)). In this method, the average background level is determined by comparing the raw projection measurements with/without fames (i.e., 1000 frames for each case). And the background level is calculated to be 0.1% of the maximum luminescence intensity. This background level is then used as the threshold value to the 3D reconstructions to determine the fame edge. Other threshold values, e.g., 0.2% and 0.5%, are tested to extract the fame edges, and similar contours are obtained. It should be noted that after the threshold value is determined, both inner boundaries and outer boundaries can be extracted. The outer boundaries represent the fame edge, and the inner boundaries indicate the holes inside the fame. We exclude the inner boundaries by judging if the boundary is enclosed by another outer boundary, if yes, we call it the

'child feature' while the outer boundary is called the 'parent feature' accordingly. Only the parent boundaries are extracted in this work. In this way, the fame edge contours shown in Fig. [10](#page-11-1) represent the outermost layer of the raw voxels. As discussed previously (see Sect. [2](#page-2-0)), the tomographic reconstruction procedure produces 150 planes spaced 0.43 mm vertically from 130 to 194 mm above the fuel jet exit. In Fig. [10](#page-11-1), only ten of these slices are shown at a vertical spacing of \sim 4.3 mm. Those horizontal flame edge contours show a narrower, more corrugated structure at the lower height and a broader, smoother structure at the higher height. The analysis of fuctuation behaviors of those fame edge contours will be described in the following section (see Sect. [4.3\)](#page-13-0).

4.2 Time‑Varying Flame Dynamics

To display the time-series of the 3D reconstructions, two movie fles are frst provided: the first video shows the evolution of the 3D flame structures in a time period of 6 ms, containing 130 consecutive frames of 3D instantaneous tomography reconstructions (Visualization 1); the second video shows a 360° rotating animation of one 3D reconstructed fame (Visualization 2).

Four sample frames in the time-series movie with a relatively large time interval (Δt) of 1.6 ms are shown in Fig. [11,](#page-12-0) which reveals several critical physical phenomena of the flame. For example, at $t=0$ ms, an isolated soot pocket rises from the left bottom of the FOV. In the subsequent images at $t = 1.6$ ms and 3.2 ms, the pocket increases in volume and merges into the main fame, respectively. Also, note that during the time from 3.2 to 4.8 ms, large-scale fame wrinkling, a result of vortex formation, roll-up, and convection, is clearly observed, indicative of the existence of local vortex-induced instability. The diameter of the formed vortex at $t = 4.8$ ms is \sim 7 mm, as shown in Fig. [11](#page-12-0)d. The downstream movement speed of the vortex is \sim 15 m/s calculated by tracking its centroid in the 15 consecutive reconstructions from 3.2 to 4.8 ms. During this time period, the vortex induces an average air-penetration volume of 4300 mm^3 to the flame (by counting the zero-value voxel number in the fame region). By screening the reconstructed data, a vortex with a similar scale and convective speed as in Fig. [11](#page-12-0)d occurs periodically almost every 3 ms around this location. This implies the local fame edge disturbance induced by the vortex oscillates at a frequency of \sim 333 Hz. And the mean magnitude of the flame edge deformation due to the vortex stretching at each height is estimated to be 96 mm² (i.e., 4300 mm³/15 m/s/3 ms). The above analysis shows the usefulness of time-varying 3D fame data to estimate the vortex-induced fame edge oscillations. The next section will further quantify the fame

Fig. 11 Evolution of overall 3D flame edge structures at four time instances with $\Delta t = 1.6$ ms

dynamics by calculating the fame sheet density (*FSD*), the oscillation amplitude (*A*′), frequency (f) , and phase (φ) of both the global and local flame edge deformation.

4.3 Quantifcation of Flame Dynamics

First, the 3D *FSD* is calculated based on the fame edge surface area per unit volume (using a sample sized of 400 frames of reconstruction):

$$
FSD = \lim_{\Delta x \to 0} \frac{\overline{A}}{\Delta x^3}
$$
 (3)

where \overline{A} is the time-averaged surface area in the interrogation box. At each point of the three-dimensional flame edge brush \overline{A} is measured by counting the flame edge voxels within an interrogation box of $\Delta x^3 = 1.28$ mm \times 1.28 mm \times 1.28 mm. Distributions of the 3D *FSD* at several diferent downstream distances are shown in Fig. [12](#page-13-1). As shown, the high value of the *FSD* distributes in an irregular region close to the central axis at the lower height (i.e., $Z = 134$ mm), which indicates a severer fluctuation of the flame edge at this location and a possible global extinction occurring upstream of this height since the *FSD* distribution appears to be an accumulation of a series of small and unsteady fame edge contours. The global extinction at the lower heights can also be inferred from the time-varying movie (Visualization 1). Note that the observation of global extinction is consistent with the previous LES results where the blow-of appears with high-frequency and high-magnitude fuctuations of the fame surfaces (Zhang and Mastorakos [2015](#page-21-6)). The fame tends to be relatively stable with the height increase as the *FSDs* are observed to distribute in a relatively stationary toroidal region (i.e., *Z*=158 mm to 182 mm), indicating the re-ignition and re-establishment of the fame. Such a stable region of the fame spans~30 mm, and the fame becomes unstable again by observing the gradual shrink of the inner core of the toroid (i.e., *Z*=190 mm).

The *FSD* shown in Fig. [12](#page-13-1) indicates the significant variance of local flame fluctuations along downstream distances. To further investigate the local fame fuctuation behaviors and how they afect the global fuctuation of the fame. The global and local fame edge deformation are calculated. Figure [13](#page-14-0) shows the global and local fame edge deformation at two consecutive time instants and the calculation strategy. For example, the blue and red

Fig. 12 Flame sheet density at diferent heights along downstream distances

Fig. 13 Global (**a**) and local (**b**) fame edge deformation during two consecutive time instants with the time interval of 66.12 μs; the blue surface (or solid black line) represents the fame edge at the former time instant, the red surface (or red dashed line) at the latter time instant

surfaces shown in Fig. [13a](#page-14-0) are the global fame edge surfaces extracted at two consecutive time instants. The time interval is $66.12 \mu s$. A localized small flame pocket of the flame is zoomed-in in Fig. [13a](#page-14-0) to show the discrepancy of the two fame edge surfaces. As can be seen, the fame edge slightly changes, indicating the time resolution of 15 kHz is suf-ficient to capture the flame edge change in the present flame. Figure [13](#page-14-0)b shows two examples of the local fame edge deformations at two diferent heights, wherein the solid black line represents the fame edge at the former time instant, the red dashed line at the latter time instant. It is clearly observed that the deformation magnitude varies for diferent parts of the fame contour. Some parts of the contour are compressed by the surrounding fow, some are stretched, and some are almost unchanged. And the fame edge was observed to respond primarily to large-scale fuctuations in the fow at the shear layer. Those observations frst demonstrate the utility of the 3D technique to analyze the fame edge deformation. In contrast, 2D techniques only capture the fame deformation from one specifc orientation. Second, the results indicate a signifcant interaction of the local fow turbulence with the fame edge. For the local fame edge deformation, the deformation magnitude (*A*′) is defned and calculated as the summation of the area formed by the two intersected fame edge contours at each height (i.e., $A' = \sum A'_i$, a sample of A'_i is marked as the dashed region in Fig. [13b](#page-14-0)). Similarly, the global flame deformation magnitude (V') is defined and calculated as the summation of the volume formed by the two intersected fame edge surfaces.

The calculated global fame edge deformation magnitude during a time period of 26 ms (corresponding to 400 frames) is shown in Fig. $14a$. In this time period, V' varies in a range of 5000–8000 mm³, and the mean flame edge deformation is 6175 mm^3 . The RMS variation of V' is 595 mm³, approximately 9.6% of the mean flame edge deformation. Fig. [14b](#page-15-0) shows the power spectral density (PSD) and the frequency response of the V' in Fig. [14](#page-15-0)a. The dominant frequency response of the V' occurs at 113 Hz and 301 Hz, respectively.

Similarly, the local fame edge deformation magnitude and PSD at two local heights of the flame in Fig. [13](#page-14-0)b are shown in Fig. [15](#page-15-1). The A' varies by as much as 200 mm² at both heights. The mean values of A' are 109 mm² and 101 mm², close to the estimated mean value in Fig. [11](#page-12-0) (i.e., 96 mm²). The calculated RMS fluctuations of the A' are 26.9% of the mean flame edge deformation at $Z=134$ mm and 32.8% of the mean flame edge deformation at $Z=190$ mm, respectively. These values are notably larger than that observed in the global deformation (9.6%), indicating that the local fuctuations are much severer than the global fuctuation, and it appears to exist a self-suppression of the instabilities between these local oscillations. The largest frequency response of the

Fig. 15 Local fame edge deformation fuctuation magnitude and power spectrum density of frequency response at two downstream distances: **a** $Z = 134$ mm and **b** $Z = 190$ mm. The red dashed line shows the mean value of *A'*

flame edge deformation at $Z=134$ mm occurs at 301 Hz, while there are four dominant frequencies (113 Hz, 263 Hz, 376 Hz, and 489 Hz) at $Z = 190$ mm. Two of them (113 Hz and 301 Hz) are matched with the dominant frequencies in the global fame edge deformation. Another dominant frequency of 376 Hz at $Z=193$ mm is close to the estimated frequency of 333 Hz based on the movement of the local vortex shown in Fig. [11](#page-12-0). Those observations imply a complicated coupling of local oscillations to govern the global oscillation.

It is thus needed to evaluate the local fame edge fuctuations along the whole height of the fame in the feld of view (FOV). Figure [16](#page-16-0) frst shows the mean deformation magnitude (A'_{mean}) of the local flame edge oscillations at different downstream locations.

As shown, the *A*′ mean plot shows two distinct regions: a rapid descending region and a re-ascending region. This observation agrees well with the *FSD* plots in Fig. [12](#page-13-1), in which the fame is undergoing an "unstable, stable, and unstable" cycle. The *FSD*s in Fig. [12](#page-13-1) suggest that the unstable region is caused by the local extinction of the fame, which results in the large magnitude of the flame edge deformation. The A'_{mean} plot also suggests that the stable region maintains in a range of 35 mm (from $Z=148$ mm to 183 mm), in which the *A*′ mean is below the averaged value (dashed line).

Figure [17](#page-16-1) shows the largest frequency response of the local fame edge oscillations at diferent downstream locations (solid black line). Based on the similarity of the frequency responses at diferent locations, the whole fame height can be divided into six subregions. The circled number above the line in Fig. [17](#page-16-1) represents the subregion index, and the star next to the number indicates there exist other high-power–frequency responses in this subregion. In the first subregion $(Z=130 \text{ mm}$ to 146 mm), the single, largest frequency response occurs at 301 Hz (referring to Fig. [15](#page-15-1)a). In the second subregion ($Z=146$ mm to 153 mm), the largest frequency response occurs at 902 Hz, and other high-power

Fig. 17 Distribution of the local frequency response and evolution of the global frequency as a function of downstream distance

frequencies include 113 Hz and 301 Hz. In the third subregion $(Z=153 \text{ mm})$ to 158 mm), the largest frequency response occurs at 113 Hz, and other high-power frequencies include 301 Hz and 902 Hz. The same three dominant frequencies exist in the second and third subregions, although the largest frequency shifts. The fourth subregion $(Z=158 \text{ mm})$ to 172 mm) is analogous to the frst regime, and wherein the single, largest frequency response occurs at 301 Hz. In the fifth subregion $(Z=172 \text{ mm})$ to 188 mm), the largest frequency remains 301 Hz, and other high-power frequencies include 113 Hz and 225 Hz. In the last subregion, the largest frequency shifts between the four dominant frequencies (i.e., 113 Hz, 263 Hz, 376 Hz, and 489 Hz (referring to Fig. [15b](#page-15-1)).

From the above observations, the frequency of 301 Hz dominates in most regions of the fame in the FOV (i.e., occupying 92% of the total height). The second most infuential frequency is 113 Hz, dominating a region occupying 52% of the total height. Thus, a natural expectation of the global frequency in the FOV should be 301 Hz. However, the fact is the power density of 113 Hz is markedly larger than that of 301 Hz in the *PSD* plot of global instability shown in Fig. [14](#page-15-0). To examine how the local frequencies result in the global instability, the red dashed line shows the evolution of the global frequency as a function of downstream distance. The global frequency along the axial direction shown in Fig. [17](#page-16-1) (red dashed line) is calculated by considering a local–global region from Z_0 to Z . As shown, the global frequency keeps at 301 Hz after the frst two subregions. It transitions into 113 Hz in the third subregion. Strangely, the global frequency maintains to be 113 Hz after the third subregion, although there still exists the strongest frequency of 301 Hz in the fourth and ffth subregions.

To explain the shift behavior of the global frequency, Fig. [18](#page-17-0)a shows the phase of local fame edge oscillations along the Z axis. As shown, the phases of oscillations with various frequencies all display continuity at subsequent downstream distances, suggesting the different vortex disturbances moving downstream. The continuous variation of the phase also provides evidence for the soot particles oscillating with the local vortex. As the slope of the phase is indicative of the vortex convective speed along the fame edge, it can be seen that the same vortex-induced disturbance would experience a velocity change when encountering other disturbances. The typical vortex convection speeds in the 6 subregions are

20 m/s, 18 m/s, 15.8 m/s, 15.7 m/s, 13.1 m/s and 11.9 m/s, respectively. The abrupt change of the phase of 301 Hz mode is due to the occurrence of a new high-frequency disturbance of 902 Hz, which is dominating in subregion 2.

The phase offset of different frequency oscillations varies markedly, resulting in the constructive and destructive interference of these oscillations. It exhibits constructive edge oscillation interference when $\Delta \varphi$ is between 0 and 0.5π radians, and the oscillations neutralize when $\Delta\varphi$ is between 0.5π and 1.5π radians. For example, the maximum phase offset of f_z =301 Hz is 2.7 π radian throughout the height in Fig. [18a](#page-17-0), suggestive of the alternation occurrence of the constructive and destructive interference. For the $f_z = 113$ Hz, the phase offset is a range of 0.9π radian, indicative of the major constructive interference. Figure [18](#page-17-0)b further quantifes the constructive and destructive interferences by showing the superposed power spectral density of these two oscillations (with $f_z = 301$ Hz and $f_z = 113$ Hz, respectively) along the *Z* axis. The linear superposition principle is used. As shown, for $f_z = 301$ Hz, the superposed *PSD* increases in the first and second subregions. It begins to decrease in the third subregion because the phase ofset between this subregion and the frst subregion is approachingπ rad. A similar decrease in supposed *PSD* due to the π rad phase offset is also observed in the sixth subregion. For $f_z = 113$ Hz, the superposed *PSD* keeps increasing as the phase offset is always smaller than π rad. It is close to π rad in the sixth subregion, leading to the weak growth of superposed *PSD* in this subregion. The final result of the superposition shows the stronger *PSD* of $f_z = 113$ Hz compared to $f_z = 301$ Hz. The result agrees with that shown in Fig. [14](#page-15-0)b, indicating that the global oscillation of the fame results from the linear superposition of local oscillations at diferent locations.

5 Summary and Conclusions

Time-resolved 3D measurements of fame edge dynamics were performed on a highspeed jet difusion fame at 15 kHz. The measurements were performed based on emission tomography implemented with optical endoscopes. Soot radiation and visible chemiluminescence signals are captured to perform the 3D reconstructions of the fame. The 3D reconstructions can resolve the 3D instantaneous fame edge structures with fne-scale corrugations. Instantaneous fame topologies were successfully captured, and highly transient fame dynamics, such as the formation and convection of vortical structures, were also observed by the 3D measurements. The analysis of fame edge deformations shows the vortex-fame interaction dominates the hydrodynamic instability of the fame. Various local frequencies of instability existed along the downstream distance induced by the localized, periodical vortex formations and convections. Those local oscillations exhibited diferent constructive and destructive interference behaviors. The global oscillation of the fame results from the linear superposition of local oscillations at diferent locations.

The contributions of the present work are twofold. First, it emphasizes the importance of establishing the relationship between the global and local instabilities, as it is always the case in the practical industry to estimate the global instability of combustor from a localized measurement point/region. The presented method can be readily extended for the conditions where multiple sources of instabilities (i.e., thermo-acoustic instabilities) are involved. Second, the measurement results of the spatially resolved local frequencies distribution along the fow direction are a direct representation of the local vorticity

distribution (Shanbhogue et al. [2009b\)](#page-20-25), but is obtained without introducing further velocity feld measurements.

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Compliance with Ethical Standards

Confict of interest The authors declare that they have no confict of interest.

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