#### **ORIGINAL ARTICLE**



# **Undiferenced ionospheric‑free ambiguity resolution using GLONASS data from inhomogeneous stations**

**Qile Zhao1 · Xiaotao Li1 · Yang Liu1 · Jianghui Geng1,2 · Jingnan Liu1**

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## **Abstract**

GLONASS frequency division multiple access signals render ambiguity resolution (AR) rather difcult because: (1) Different wavelengths are used by diferent satellites, and (2) pseudorange inter-frequency biases (IFBs) cannot be precisely modeled by means of a simple function. In this study, an AR approach based on the ionospheric-free combination with a wavelength of about 5.3 cm is assessed for GLONASS precise point positioning (PPP). This approach simplifes GLONASS AR because pseudorange IFBs do not matter, and PPP-AR can be enabled across inhomogeneous receivers. One month of GLONASS data from 165 European stations were processed for diferent network size and diferent durations of observation periods. We find that 89.9% of the fractional parts of ionospheric-free ambiguities agree well within  $\pm$  0.15 cycles for a small network (radius = 500 km), while 77.6% for a large network (radius = 2000 km). In case of the 3-hourly GLONASS-only static PPP solutions for the small network, reliable AR can be achieved where the number of fxed GLONASS ambiguities account for 97.6% within all candidate ambiguities. Meanwhile, the RMS of the east, north and up components with respect to daily solutions is improved from 1.0, 0.6, 1.2 cm to 0.4, 0.4, 1.1 cm, respectively. When GPS PPP-AR is carried out simultaneously, the positioning performance can be improved significantly such that the GLONASS ambiguity fixing rate rises from 74.4 to 95.4% in case of hourly solutions. Finally, we introduce ambiguity-fxed GLONASS orbits to re-attempt GLONASS PPP-AR in contrast to the above solutions with ambiguity-foat orbits. We fnd that ambiguity-fxed orbits lead to clearly better agreement among ionospheric-free ambiguity fractional parts in case of the large network, that is 80.5% of fractional parts fall in  $\pm$  0.15 cycles in contrast to 74.6% for the ambiguity-float orbits. We conclude that highly efficient GLONASS ionospheric-free PPP-AR is achievable in case of a few hours of data when GPS PPP-AR is also accomplished, and ambiguity-fxed GLONASS orbits will contribute signifcantly to PPP-AR over wide areas.

**Keywords** Precise point positioning · GLONASS · Ionospheric-free ambiguity resolution · Fractional-cycle bias · Inhomogeneous stations

## **Introduction**

In recent years, ambiguity resolution for precise pointing positioning (PPP-AR) has been of much interest since higher accuracy and faster initializations can be achieved in contrast to conventional foat solutions (Bertiger et al. [2010](#page-8-0); Collins et al. [2010;](#page-8-1) Laurichesse et al. [2009](#page-9-0)). Particularly, Ge et al. ([2008](#page-8-2)) developed a GPS PPP-AR approach where

 $\boxtimes$  Jianghui Geng jgeng@whu.edu.cn the fractional-cycle biases (FCBs) of uncalibrated phase delays (UPDs) are produced across a reference network and then delivered to PPP users to enable AR at a single station. While this approach applies to all GNSS emitting code division multiple access (CDMA) signals, it does not apply to the GLONASS constellation, which transmits frequency division multiple access (FDMA) signals at the moment.

One important reason is that inter-frequency biases (IFBs) are present within both GLONASS pseudorange and carrier phase observations, which have to be corrected before attempting GLONASS AR (Leick et al. [1995;](#page-9-1) Sleewaegen et al. [2012;](#page-9-2) Wang [2000](#page-9-3); Wang et al. [2001](#page-9-4)). Carrier phase IFBs are relatively stable over time and are usually presumed to precisely follow a linear function of the frequency channel number (Al-Shaery et al. [2013](#page-8-3); Geng et al. [2017](#page-8-4); Wanninger

<sup>&</sup>lt;sup>1</sup> GNSS Research Center, Wuhan University, Wuhan, China

<sup>&</sup>lt;sup>2</sup> Collaborative Innovation Center of Geospatial Technology, Wuhan, China

[2012](#page-9-5)). However, such linear function normally breaks down for pseudorange IFBs. It has been reported that pseudorange IFBs vary with receivers, antennas, domes and frmware and hence, cannot be modeled using a simple function but should be addressed for individual stations and satellites (Reussner and Wanninger [2011](#page-9-6); Shi et al. [2013;](#page-9-7) Yamada et al. [2010](#page-9-8)). It is the pseudorange IFBs that complicate GLONASS AR that relies on the Hatch–Melbourne–Wübbena combination observable (Hatch [1982\)](#page-9-9) for long baseline positioning or PPP across inhomogeneous stations.

At present, two approaches have been developed to avoid or diminish the adverse impact of pseudorange IFBs on GLONASS PPP-AR for inhomogeneous stations. One idea is that predetermined ionospheric products, e.g., global ionospheric maps, can be introduced into the wide-lane carrier phase combination observable to compute its ambiguity, rather than by means of the Hatch–Melbourne–Wübbena combination (Reussner and Wanninger [2011\)](#page-9-6). Based on raw observables, Geng and Bock ([2016](#page-8-5)) directly constrained the line of sight ionospheric parameters with external ionospheric products and presented the full schemes to accomplish GLONASS PPP-AR across inhomogeneous stations, with the extra beneft of speeding up PPP convergence to ambiguity-fxed solutions (Geng and Shi [2017](#page-8-6)). However, the performance of this strategy is subject to the accuracy of ionospheric products, and a regional ionospheric product is preferred (Geng and Shi [2017](#page-8-6)). On the other hand, Banville [\(2016\)](#page-8-7) proposed to fx ionospheric-free ambiguities with a wavelength of about 5.3 cm for GLONASS AR. Diferent from GPS ionospheric-free ambiguities, the GLONASS counterparts are integers in nature thanks to the ratio of 9/7 between the L1 and L2 frequencies. Banville [\(2016](#page-8-7)) calculated undiferenced ambiguities using PPP and formulated double-diference ambiguities over 12 baselines spanning 300–1400 km for integer cycle resolution. It is reported that ionospheric-free AR is indeed achievable, though shorter periods of data, e.g., 1 h, cannot efficiently guarantee a high fixing rate. Liu et al. ([2016](#page-9-10)) showed that GLONASS ionospheric-free AR over baselines of longer than 2000 km using daily data could be accomplished with a fixing rate of over 80% and the agreement of orbit overlaps can be improved by more than 20%.

This study is dedicated to GLONASS PPP-AR over inhomogeneous stations where we exploit the performance of resolving undifferenced ionospheric-free ambiguities, which are both immune to pseudorange IFBs and exemption of frst-order ionospheric delays. The only disadvantage of ionospheric-free ambiguities is their ultrashort wavelength of about 5.3 cm, which is exactly half of the narrow-lane wavelength. The ensuing question is what performance we can achieve when fxing them, especially in case of short durations of data, e.g., a few hours and networks of various scales, e.g., hundreds to thousands of kilometers, and in combination

with GPS PPP-AR. More interestingly, we will also investigate how the GLONASS orbits produced by fxing ionosphericfree ambiguities according to Liu et al. ([2016](#page-9-10)) beneft GLO-NASS PPP-AR. We frst present the mathematical model and the details of the ionospheric-free PPP-AR, then, describe the experiments and analyze the performance, and fnally draw the conclusions.

## **Methods**

In this section, we frst give a brief introduction about GLO-NASS ionospheric-free observable. In addition, based on the idea of fxing ionospheric-free ambiguities with a wavelength of about 5.3 cm for GLONASS proposed by Banville [\(2016\)](#page-8-7) and Liu et al. ([2016](#page-9-10)), the details of GLONASS ionosphericfree FCBs estimation are presented. The receiver carrier phase IFBs of GLONASS must be calibrated before the FCB estimation.

#### **GLONASS ionospheric‑free observable**

Undiferenced GLONASS observation equations of dualfrequency carrier phase and pseudorange measurements on frequency band  $(g = 1, 2)$  from receiver *a* to satellite *i* can be written as

$$
P_{g,a}^i = \rho_a^i + I_a^i / f_{g,k}^2 + B_{g,a}^i - B_g^i, \quad \varepsilon_{P_{g,a}^i} = \sigma_P
$$
  

$$
L_{g,a}^i = \rho_a^i - I_a^i / f_{g,k}^2 + b_{g,a}^i - b_g^i + \lambda_g^k N_{g,a}^i, \quad \varepsilon_{L_{g,a}^i} = \sigma_L
$$
 (1)

where  $P_{g,a}^i$  and  $L_{g,a}^i$  are pseudorange and carrier phase measurements in the unit of length, with precisions of  $\sigma_{\rm P}$  and  $\sigma_{\rm L}$ , respectively;  $\rho_a^i$  denotes the non-dispersive delay including the geometric distance, the tropospheric and relativistic effects; *k* is the frequency channel number of satellite  $i$ ;  $f_{g,k}$ is the frequency of channel *k* on frequency band  $g$  ( $g = 1, 2$ );  $I_a^i / f_{g,k}^2$  is the first-order ionospheric delay for  $f_{g,k} B_{g,a}^i$  and  $b_{g,a}^i$ are uncalibrated pseudorange and carrier phase biases at station a, respectively. Correspondingly,  $B^i_g$  and  $b^i_g$  are pseudorange and carrier phase biases at satellite *i*, respectively;  $N_{g,a}^i$ is the integer carrier phase ambiguity and is scaled by wavelength  $\lambda_g^k$ . Unmodeled errors such as multipath and higher order ionospheric efects are ignored for brevity.

For PPP, the ionospheric-free combination observable is usually used, that is

$$
P_{if,a}^i = \alpha P_{1,a}^i + \beta P_{2,a}^i = \rho_a^i + B_{if,a}^i - B_{if}^i
$$
  

$$
L_{if,a}^i = \alpha L_{1,a}^i + \beta L_{2,a}^i = \rho_a^i + \alpha \lambda_1^k N_{1,a}^i + \beta \lambda_2^k N_{2,a}^i + b_{if,a}^i - b_{if}^i
$$
  
(2)

where

$$
B_{\text{if},a}^{i} = \alpha B_{1,a}^{i} + \beta B_{2,a}^{i}
$$
  
\n
$$
B_{\text{if}}^{i} = \alpha B_{1}^{i} + \beta B_{2}^{i}
$$
  
\n
$$
b_{\text{if},a}^{i} = \alpha b_{1,a}^{i} + \beta b_{2,a}^{i}
$$
  
\n
$$
b_{\text{if}}^{i} = \alpha b_{1}^{i} + \beta b_{2}^{i}
$$

and

$$
\alpha = f_{1,k}^2 / (f_{1,k}^2 - f_{2,k}^2) \n\beta = -f_{2,k}^2 / (f_{1,k}^2 - f_{2,k}^2)
$$

The terms  $B_{if,a}^i$ ,  $B_{if}^i$ ,  $b_{if,a}^i$  and  $b_{if}^i$  cannot be directly estimated, which are, however, absorbed into either observation residuals or other parameters such as clocks and ambiguities (Geng et al. [2010a](#page-8-8)). For example,  $b_{if,a}^i$  and  $b_{if}^i$  will be combined with the ambiguity parameters, such that

$$
\lambda_{if}^{k} N_{if,a}^{i} = \alpha \lambda_{1}^{k} N_{1,a}^{i} + \beta \lambda_{1}^{k} N_{2,a}^{i} + b_{if,a}^{i} - b_{if}^{i}
$$
\n
$$
= \frac{9}{32} \lambda_{1}^{k} \left( 9N_{1,a}^{i} - 7N_{2,a}^{i} \right) + b_{if,a}^{i} - b_{if}^{i}
$$
\n
$$
= \frac{1}{2} \lambda_{n}^{k} \left( 9N_{1,a}^{i} - 7N_{2,a}^{i} \right) + b_{if,a}^{i} - b_{if}^{i}
$$
\n(3)

where  $\lambda_{if}^k$  and  $N_{if,a}^i$  are the ionospheric-free wavelength and ambiguity, respectively;  $\lambda_n^k$  is narrow-lane wavelength. Defining

$$
N_{N,a}^i = 9N_{1,a}^i - 7N_{2,a}^i
$$

equation  $(3)$  $(3)$  can then be rewritten as

$$
\lambda_{\text{if}}^k N_{\text{if},a}^i = \lambda_{\text{if}}^k N_{N,a}^i + b_{\text{if},a}^i - b_{\text{if}}^i \tag{4}
$$

 $N_{N,a}^i$  has integer nature and its wavelength,  $\lambda_{if}^k = \frac{1}{2} \lambda_n^k$ , is about 5.3 cm, which is exactly half of the narrow-lane wavelength (Dai [2000](#page-8-9); Roßbach [2000\)](#page-9-11).

## **GLONASS ionospheric‑free FCB estimation**

In the case of GPS data, single-diference ambiguities between satellite *i* and *j* are usually formed to eliminate the receiver dependent biases. However, considering that the wavelengths of GLONASS carrier phase data are not identical across diferent channels, and that carrier phase IFBs have a linear relationship with the channel number, we can obtain

$$
\lambda_{\text{if}}^k N_{\text{if},a}^i - \lambda_{\text{if}}^l N_{\text{if},a}^j = \Delta \lambda_{\text{if}}^{kl} N_{N,a}^i + \lambda_{\text{if}}^l \Delta N_{N,a}^{ij} + \tau \cdot \Delta b_{\text{if},a} - \Delta b_{\text{if}}^{ij}
$$
\n(5)

where the channel number of satellite *i* and *j* is *k* and *l*, respectively, and  $\tau = k - l$ ;  $\Delta b_{if,a}$  is the ionospheric-free IFBs defned between adjacent channel number; and

$$
\Delta N_{N,a}^{ij} = N_{N,a}^i - N_{N,a}^j
$$
  
\n
$$
\Delta \lambda_{if}^{kl} = \lambda_{if}^k - \lambda_{if}^l, \left| \Delta \lambda_{if}^{kl} \right| \le 0.245 \text{ mm}
$$
  
\n
$$
\Delta b_{if}^{ij} = b_{if}^i - b_{if}^j
$$

Note that  $N_{N,a}^i$  and  $\Delta N_{N,a}^{ij}$  cannot be estimated simultaneously due to rank deficiency. However,  $N_{N,a}^i$  can be approximated using its corresponding ionospheric-free ambiguity estimate  $N^i_{\text{if},a}$ . Since the wavelength of  $N^i_{N,a}$  is as small as 0.245 mm, the discrepancy between the true values of  $N_{if,a}^i$ and  $N_{N,a}^i$  will negligibly bias  $\Delta N_{N,a}^{ij}$ . Moreover,  $\Delta b_{if,a}$  can be calibrated using predetermined carrier phase IFBs (Wan-ninger [2012](#page-9-5); Geng et al. [2017](#page-8-4)). Assuming that  $\Delta b_{if,a}$  and  $N_{N,a}^i$  are known, we move them to the left side of ([5](#page-2-1)) and obtain

<span id="page-2-2"></span>
$$
l_{\text{if},a}^{kl} / \lambda_{\text{if}}^l = \Delta N_{N,a}^{ij} - \Delta b_{\text{if}}^{ij} / \lambda_{\text{if}}^l \tag{6}
$$

where

<span id="page-2-0"></span>
$$
l_{if,a}^{kl} = \lambda_{if}^k N_{if,a}^i - \lambda_{if}^l N_{if,a}^j - \Delta \lambda_{if}^{kl} N_{N,a}^i - \tau \cdot \Delta b_{if,a}
$$

Equation [\(6](#page-2-2)) is the basis for the estimation of GLONASS ionospheric-free FCBs, and the fractional parts of a satellite pair can be obtained.

Once we obtain the fractional parts of all relevant  $l_{if,a}^{kl}/\lambda_{if}^{l}$ for a satellite pair *i* and *j* using a reference network, we then average them to estimate single-diference FCBs (Ge et al. [2008;](#page-8-2) Cheng et al. [2017](#page-8-10)). Afterward, we select one satellite as the reference to convert all single-diference FCBs into undiferenced (pseudoabsolute) FCBs for PPP-AR. We note that rigorous FCB estimation should be achieved by frst fxing double-diference ambiguities of the reference network, which, however, is not implemented in this study for convenience (Geng et al. [2012](#page-8-11)).

#### **Data processing**

<span id="page-2-1"></span>We processed about 165 European stations from the EUREF (Europe Reference Frame) permanent network and IGS network, tracking both GPS and GLONASS, for the day of year (DOY) 300–330 in 2013 (Fig. [1](#page-3-0)). The number of receiver families involved in the network is seven (Table [1\)](#page-3-1). Again, we stress that ionospheric-free PPP-AR can be straightforwardly applied to GLONASS stations with diverse receiver types, antennas, domes and firmware. European Space Agency (ESA) fnal GPS/GLONASS orbits, 30-s satellite clock corrections and earth rotation parameters were used. A  $7^{\circ}$  cutoff angle was set for usable observations, but  $10^{\circ}$  for eligible ambiguities used for FCB estimation and ambiguity



<span id="page-3-0"></span>**Fig. 1** Diferent scale of GPS/GLONASS networks in Europe from the EUREF and IGS. Red solid circles denote stations used for the FCB estimation. Black open circles denote stations used to test PPP-

<span id="page-3-1"></span>**Table 1** Number of stations for each receiver family

Receiver family	Number of station
Javad	3
<b>JPS</b>	22.
Leica	81
<b>NovAtel</b>	4
Septentrio Polarx	5
<b>TPS</b>	13
Trimble	37

resolution. The precisions of pseudorange and carrier phase measurements were presumed to be 2 m and 2 cm, respectively. The phase center offset and variation corrections were also applied (Dach et al. [2011](#page-8-12); Schmid et al. [2007\)](#page-9-12). The phase wind-up efects were corrected (Wu et al. [1993](#page-9-13)). The mapping function of zenith troposphere delays was VMF1 (Vienna Mapping Function 1) (Boehm et al. [2006](#page-8-13)), and the power spectral density for zenith wet components was set as 2 cm/ $\sqrt{h}$ . The receiver position, the receiver clocks, the zenith tropospheric delays and horizontal gradients and the ambiguities were estimated. FCBs were re-estimated every 15 min, and no constraints on neighboring FCB estimates were applied. Rejection of outlier ambiguity fractional parts for the FCB estimation was carried out using a threshold of three times the standard deviation and the sign-constrained robust least squares method (Xu [2005\)](#page-9-14). Based on the method proposed by Liu et al. [\(2016](#page-9-10)), the receiver carrier phase IFBs were estimated with respect to the values listed by Wanninger [\(2012\)](#page-9-5). The LAMBDA (Teunissen [1995](#page-9-15); De Jonge and Tiberius [1996](#page-8-14)) method was used to search for integer ambiguities, and the ratio test, which is defned as the ratio between the weighted sum of the squared residuals of the second best solution to that of the best, was used to validate integer candidates. The threshold of the ratio test was 3.0 throughout in this study.

AR. Shown in panel left, middle and right are the small, medium and large networks, covering circular areas with radii of 500, 1000 and 2000 km centered at station WTZR, respectively

In order to investigate how the network extent afects the performance of ionospheric-free FCB estimation and PPP-AR, we tried three scales of networks which cover circular areas with radii of 500, 1000 and 2000 km centered at station WTZR. In the following, they are denoted as the small, medium and large networks including 52, 107 and 165 stations, respectively. Within each network, we chose a few tens of stations, i.e., 26, 37 and 55, respectively, to estimate FCBs whereas the remaining stations were used to test PPP-AR. For each network, we computed three types of solutions:

- "GLONASS-only and GLONASS fixed" solutions where only GLONASS data were processed, and only GLONASS ionospheric-free ambiguity fixing was attempted;
- "GLONASS + GPS and GLONASS fixed" solutions where both GLONASS and GPS data were processed, but only GLONASS ionospheric-free ambiguities were fixed;
- "GLONASS + GPS and both fixed" where both GLO-NASS and GPS data were processed, and all relevant ambiguities were fxed. Note that wide-lane and narrow-lane ambiguities were fxed for GPS.

In addition, for each type of solutions, we also tried diferent observation durations, i.e., 1, 2 and 3 h. Then the efficiency of GLONASS ionospheric-free ambiguity resolution and the positioning performance with respect to the observation durations are presented and discussed. Float solutions which are eligible for PPP-AR are defned as those that have fve and more undiferenced GLONASS ambiguities to be resolved. If five or more GLONASS ambiguities are resolved and the positioning diferences for the truth benchmark in the east and north components are both below 2 cm, this solution will be considered as a correctly fxed solution. In this study, daily solutions are taken as the truth benchmarks. The ambiguity fxing rate is

defned as the percentage of resolved ambiguity within all eligible ones. The correctly fxed ambiguity rate is defned as the percentage of correctly fxed ambiguities within all fxed ambiguities. Here, the daily ambiguity estimates were used as the reference values.

# **Results**

In this section, we will frst analyze the performance of GLONASS FCB estimation and then, inspect the efficiency of GLONASS ionospheric-free ambiguity resolution and the positioning performance, and fnally, the contribution of ambiguity-fxed GLONASS orbits to GLONASS PPP-AR is investigated.

## **GLONASS ionospheric‑free FCB estimation**

The ionospheric-free FCB estimates on day 300 for eight exemplary satellites and the distribution of relevant fractional parts after removal of FCBs for all GLONASS satellites over consecutive 31 days are shown in Fig. [2.](#page-4-0) We note that the FCBs are estimated using only the reference stations within each network, while the distribution of fractional parts in Fig. [2](#page-4-0) is based on the solutions from all stations. It can be seen that the ionospheric-free FCB time series change very smoothly and steadily over time. The FCBs for R12 from 11 to 14 h derived from the large network show drastic jumps of up to 0.5 cycles, which is marked with a dashed square. We contend that these jumps are due to the low elevation  $(< 20^{\circ})$  of R12 during the entire observation periods with respect to all stations, which leads to noisy ambiguity estimates. Overall, the ionospheric-free FCB estimates from all networks vary minimally within a continuous observation period.

The distribution of the fractional parts after removal of FCBs is an important diagnostics to assessing the quality of FCB estimations. The closer to zero such a fractional part is, the more accurate FCBs we have estimated and the more reliable PPP-AR can be achieved potentially. For the small network, 89.9% of all fractional parts fall in the range



<span id="page-4-0"></span>**Fig. 2** GLONASS ionospheric-free FCBs for eight satellites estimated every 15 min over reference stations on day 300 of 2013 (left column) and distribution of all fractional parts after removal of FCBs

over the 31 days (right column), for large (top), medium (middle) and small network (bottom).  $\sigma$  denotes the standard deviation of the fractional parts.  $\in$  means 'within' while  $\notin$  means 'outside'

of  $\pm$  0.15 cycles, 97.8% in  $\pm$  0.25 cycles and 0.6% outside  $\pm$  0.35 cycles with a standard deviation of 0.10 cycles. Comparatively, the medium network performs slightly worse, with 85.0% falling in  $\pm$  0.15 cycles and a standard deviation of 0.11 cycles. As the network size is increased, the performance of FCB estimation is worsened. In particular, for the large network, there are only 77.6% of fractional parts falling in  $\pm$  0.15 cycles. This phenomenon can be understood in terms of the fact that the larger the network is, the error correlation among stations gets weaker. As a result, the spatially correlated errors, such as orbits and atmosphere refractions, cannot be absorbed into the FCB estimates, resulting in poor agreement among the fractional parts of ionosphericfree ambiguities from all involved stations. Although Fig. [2](#page-4-0) shows that GLONASS ionospheric-free FCBs have excellent temporal stability and the fractional parts are overall in good agreement, their performance is still worse than those for the GLONASS narrow-lane FCBs in Geng and Bock ([2016](#page-8-5)). Even for the small network, for example, 89.9% of all fractional parts were located in the range of  $\pm$  0.15 cycles, which is below the 99% reported by them. This should be caused by the shorter wavelength of ionospheric-free ambiguities than that of narrow-lane ambiguities.

#### **GLONASS‑only PPP‑AR**

<span id="page-5-0"></span>**Table 2** Results of hourly static PPP with GLONASS-only data from rover stations over the

Stations involved in static PPP of this study are exclusive of reference stations, which are used to estimate FCBs above. Table [2](#page-5-0) shows GLONASS-only static PPP solutions. PPP-AR improves the positioning accuracy

signifcantly, since the RMS for the east, north and up components with respect to the daily position estimates for the small network is reduced from 1.5, 1.0, 2.3 cm for the float solutions to  $0.6$ ,  $0.5$ ,  $1.9$  cm for the fixed solutions, which equates an improvement of 60, 50 and 17%, respectively. Such improvement can also be clearly found for the medium and large networks, where the east component is improved most. When the observation period is increased from 1 to 2 and 3 h, the positioning accuracy of foat solutions steadily improves (Table [2\)](#page-5-0), and ionospheric-free ambiguity fxing still contributes signifcantly to the improvement in positioning accuracy, which agrees with Banville ([2016](#page-8-7)).

In the case of hourly data, the percentages of fxed solutions and fxed ambiguities are only 66.8 and 74.4%, respectively, for a small network, which is obviously worse than that of GPS (Geng et al. [2010b\)](#page-8-15). When the observation period is increased to 2 and 3 h, the percentage of correctly fxed solutions increases to 90.1 and 97.5%, respectively, while the percentage of fxed ambiguities increases to 91.7 and 97.6%, respectively. This is expected as longer observation periods should lead to higher accuracy of ambiguity estimates, which in turn improve the efficiency of ambiguity fxing. Furthermore, even for 1 h observation periods, the percentage of correctly fxed ambiguities can achieve 94.8 and 98.5% for 3 h, suggesting that this percentage is slightly afected by the length of observation periods. This indicates that once the GLONASS ionospheric-free ambiguities are presumably resolved after the LAMBDA search and the ratio test, they can normally be fxed to the correct integers, even for 1 h observation

Scale of network Fixed sol. [%] East [cm] fxed/foat North [cm] fxed/foat Up [cm] fxed/foat Fixed amb. [%] Correct amb. [%] *1 hourly* Small 66.8 0.6/1.5 0.5/1.0 1.9/2.3 74.4 94.8 Medium 63.4 0.6/1.5 0.6/1.0 1.9/2.3 71.8 94.0 Large 61.7 0.7/1.5 0.6/1.0 2.1/2.3 70.9 93.3 *2 hourly* Small 90.1 0.4/1.3 0.4/0.7 1.2/1.5 91.7 98.7 Medium 87.8 0.5/1.3 0.4/0.7 1.3/1.5 89.9 98.5 Large 84.4 0.5/1.2 0.5/0.7 1.5/1.6 86.8 98.1 *3 hourly* Small 97.5 0.4/1.0 0.4/0.6 1.1/1.2 97.6 98.5 Medium 97.5 0.4/1.0 0.4/0.6 1.2/1.3 97.1 98.5 Large 95.8 0.5/1.0 0.4/0.6 1.3/1.4 94.8 98.2

Column "Fixed sol." shows the percentage of correctly fxed solutions among all solutions to be fxed. In columns "East", "North", "Up" are the RMS (cm) of fxed and foat hourly solutions with respect to daily solutions, separated by "/". Column "Fixed amb." shows the percentage of resolved ambiguities within all eligible ones. Column "Correct amb." shows the percentage of correctly fxed ambiguities within all fxed ones, and the criterion of the correct ambiguities is the equivalence between the hourly and the daily ambiguity estimates

31 days

periods. Moreover, Table [2](#page-5-0) shows that when the observation periods reach 3 h, a fairly reliable AR can be achieved for GLONASS-only static PPP. For example, the percentages of the fxed solutions and the fxed ambiguities can reach 95.8 and 94.8% for the large network, respectively.

## **Integrated GLONASS and GPS PPP‑AR**

Table [3](#page-6-0) shows static PPP results for "GLONASS + GPS and GLONASS fxed" solutions and "GLONASS + GPS and both fxed" solutions. Taking the small network as an example, the percentage of correctly fxed solution increases from 66.8 to 80.6%, and the percentage of fxed GLONASS ambiguities also increases from 74.4 to 83.8% in case of 1 h observation periods, after introducing GPS data into GLONASS solutions while only GLONASS ambiguities are fxed. Meanwhile, the RMS for the east, north and up components with respect to daily solutions is decreased from 0.6, 0.5, 1.9 cm to 0.5, 0.5, 1.3 cm, respectively. If GPS PPP-AR is further performed, the percentage of correctly fxed solution can be improved from 80.6 to 95.9% and the percentage of fxed GLONASS ambiguities also improves from 83.8 to 95.4%. The RMS for the east, north and up components is reduced further to 0.4, 0.4, 1.1 cm, respectively.

When the length of observation period is increased from 1 to 2 and 3 h, the percentage of correctly fxed solutions increases from 80.6 to 95.6 and 99.4%, respectively, in case of the "GLONASS + GPS and GLONASS fxed" solutions. In contrast, the percentage of correctly fxed solutions increases from 95.9 to 99.0 and 99.4%, respectively, in case of the "GLONASS + GPS and both fixed" solutions. From Table [3,](#page-6-0) we can see that the performance of the "GLONASS + GPS and both fxed" solutions in case of 2 h observation periods is at the similar level to those of the "GLONASS-only and GLONASS fxed" and "GLO-NASS + GPS and GLONASS fixed" solutions in case of 3 h observation periods. In general, introducing GPS data and subsequently fxing GPS ambiguities can improve the positioning performance signifcantly, especially the success rate of GLONASS ambiguity resolution.

# **GLONASS PPP‑AR in case of ambiguity‑fxed GLONASS orbits**

GLONASS orbit errors are one of the major factors that afect the performance of PPP-AR, especially in case of a large network. The orbit accuracy can be signifcantly improved after successful ambiguity resolution over the reference network

<span id="page-6-0"></span>**Table 3** Results of hourly static PPP at rover stations over the 31 days



Column "Fixed sol." shows the percentage of correctly fxed solutions among all those to be fxed. In columns "East", "North", "Up" are the RMS (cm) of fixed and float solutions with respect to daily solutions, separated by "*i*". Column "Fixed amb." shows the percentage of resolved GLONASS ambiguities within all eligible GLONASS ones

used for orbit determination. In this section, we will use both ambiguity-fxed and ambiguity-foat orbits to investigate the performance of GLONASS PPP-AR. At present, the GLONASS products of IGS analysis centers are mainly based on the foat ambiguities, except for those from CODE (Centre for Orbit Determination in Europe). The ambiguity-fxed orbit product used in this section is derived using the method by Liu et al. [\(2016](#page-9-10)). We re-computed the "GLONASS-only and GLONASS fxed" solutions with the large network for day 300–330 in 2013.

The results are shown in Fig. [3](#page-7-0) and Table [4.](#page-7-1) Compared to the results from foat orbits, the results of fxed orbits have a better performance, and the percentages of the fractional parts within  $\pm$  0.15 cycles increase substantially from 74.6 to 80.5%, those within  $\pm$  0.25 cycles increase from 90.4 to 94.0% and those outside  $\pm$  0.35 cycles decrease from 3.6 to 2.0% on average. In particular, the percentages of the fractional parts within range  $\pm$  0.15 cycle vary from 72.2 to 85.4% for the fixed orbits, in contrast to 59.4–82.8% for the foat orbits, which shows that the performance with respect to the fxed orbits is better. The percentages of the correctly fxed solution with respect to the fxed orbits are 63.1%, but just 60.2% with respect to the foat orbits in case of 1 h observation periods. The percentages of fxed ambiguities within all candidate ambiguities are 71.4 and 68.6% for fxed and foat orbits, respectively. Compared to the foat orbits, the performance of static PPP-AR for the fxed orbits improves in terms of the fxing rate of ambiguities and the positioning accuracy. Therefore, for a large network spanning over a few 1000 km, ambiguity-fxed GLONASS orbit will clearly beneft GLONASS PPP-AR.

# **Conclusions**

This study investigates GLONASS PPP-AR based on the ionospheric-free ambiguities with a wavelength of about 5.3 cm, motivated by Banville [\(2016\)](#page-8-7) and Liu et al. ([2016](#page-9-10)).

<span id="page-7-0"></span>**Fig. 3** Distribution of all fractional parts after removal of FCBs over the 31 days. *σ* denotes the standard deviation of the fractional parts.  $\in$  means 'within' while ∉ means 'outside'. Red and black represent ambiguity-fxed and- foat orbits





Column "Fixed sol." shows the percentage of correctly fxed solutions within all those to be fxed. In columns "East", "North", "Up" are the RMS (cm) of fxed and foat hourly solutions with respect to daily solutions, separated by "/". Column "Fixed amb." shows the percentage of resolved GLONASS ambiguities within all candidate GLONASS ones

<span id="page-7-1"></span>**Table 4** Results of hourly static PPP of the ambiguity-fxed andfoat orbits from rover stations within large network over the 31 days

The data processing strategy is fairly straightforward as pseudorange IFBs are no longer a factor in deterring PPP-AR. With the data from day 300–331 in 2013 from about 165 stations in Europe, the performance of GLONASS PPP-AR is investigated with diferent network sizes, types of solution and lengths of the observation period.

The distribution of the fractional parts after removal of FCBs is an important diagnostics to assessing the quality of FCB estimations. For the small network, we fnd that 89.9% of all fractional parts agree well within  $\pm$  0.15 cycles with a standard deviation of 0.10 cycles. The performance slightly decreases when the size of the network increases, and 77.6% of fractional parts are within  $\pm$  0.15 cycle with a standard deviation of 0.14 cycle for the large network.

When the length of observation period reaches 3 h, a fairly reliable PPP-AR can be achieved for GLONASSonly static PPP. The percentages of fxed solution and fxed ambiguities can reach 97.5 and 97.6%, respectively, for the small network. The RMS of east, north and up directions decreases from 1.0, 0.6, 1.2 cm to 0.4, 0.4, 1.1 cm by 60.0, 33.3 and 8.3%, respectively.

Introducing GPS and simultaneously fxing GPS ambiguities, the performance of positioning is improved to a large extent, especially increasing the percentages of fxed GLO-NASS ambiguities. For the solution type "GLONASS + GPS and both fxed," the small network and the 1-h observation period, compared to solution type "GLONASS-only and GLONASS fxed" the percentage of fxed GLONASS ambiguity increases from 74.4 to 95.4%. For the 2-h observation period, the percentage of fxed ambiguity can reach 98.4% for the small network and also can reach 94.4% for the large network. The performance of static PPP-AR shows higher efficiency and reliability than the solution type "GLONASSonly and GLONASS fxed" of the 3-h observation period. Overall, we demonstrate that 2 h of data can usually ensure a reliable GLONASS PPP-AR which is comparable to what can be obtained with 3 h of data.

Ambiguity-fxed and- foat orbits are used to evaluate the efect of orbital errors on the performance of PPP-AR. The results show that for the large network, the fractional parts of ambiguity-fxed orbits have better consistency and show more stability from day to day than from ambiguityfoat orbits. The percentages of GLONASS fxed ambiguities are 71.4 and 68.6% for ambiguity-fxed and ambiguityfoat orbits for the 1-h observation period, respectively. A GLONASS ambiguity-fxed orbit is necessary, which makes AR more efective. Generally, a high precision orbit can efectively improve the performance of PPP-AR over a large network.

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**Qile Zhao** is a professor of GNSS Research Center of Wuhan Uni versity. He received his Ph.D. degree in Wuhan University in 2004. In 2006–2007, as a post doctoral fellow, he did his post doctoral program in DEOS, Delft University of Technology, the Netherlands. His current research interests are precise orbit determination of GNSS and low earth orbit satellites, and high-precision positioning using GNSS.

**Xiaotao Li** is currently a Ph.D. candidate at GNSS Research Center, Wuhan University. He received his bachelor degree in 2011 from Chang'an University. His current research mainly focuses on high-precision posi tioning using GNSS.



Yang Liu is currently a postdoctoral researcher at Wuhan University. He received his bachelor degree in geomatics engineering from School of Geodesy and Geomat ics, Wuhan University, in 2010. He received his Ph.D. degree in geodesy and surveying engineer ing from GNSS Research Center, Wuhan University, in 2016. His main research interests include real-time precise orbit determina tion of GNSS and LEO satellites and geophysical applications.



**Jianghui Geng** has been a professor in GNSS geodesy at Wuhan University since 2015. He graduated from the University of Nottingham in the UK in 2011. Afterward, he had an enterprise fellowship from the Nottingham Geospatial Institute in 2011 and a Green scholarship from Scripps Institution of Oceanography from 2012 to 2014. He was promoted as an assistant project scientist in 2015 (see homepage: pride.whu. edu.cn). His major research interest is high-precision GNSS.



**Jingnan Liu** graduated from the former Wuhan College of Surveying and Mapping in 1967 and received his master's degree in 1982. He was elected Academician of the Chinese Academy of Engineering in 1999. He has been a member of the Science and Technology Committee, Ministry of Education of China, in 1997–2009 and as an editorial board member of GPS Solutions in 1998–2000. He is currently an executive member of the council, Chinese Society for Geodesy Photogrammetry and Cartogra-

phy; the editorial board member of GPS World; and the coordinator of International GPS Geodynamics Services. His current research interests are satellite precise orbit determination.