



# The origin of low-MgO eclogite xenoliths from Obnazhennaya kimberlite, Siberian craton

Jing Sun<sup>1,2</sup> · Roberta L. Rudnick<sup>3</sup> · Sergey Kostrovitsky<sup>4,5</sup> · Tatiana Kalashnikova<sup>4</sup> · Kouki Kitajima<sup>6</sup> · Ranpeng Li<sup>3</sup> · Qiao Shu<sup>7,8</sup>

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## Abstract

The petrology, mineral major and trace-element concentrations, and garnet oxygen isotopic composition of low-MgO (11–16 wt%) eclogites from the Obnazhennaya kimberlite, Siberian craton, are used to infer their petrogenesis. These eclogites contain two types of compositionally distinct garnet: granular coarse garnet, and garnet exsolution (lamellae and fine-grained garnet) in clinopyroxene. The former record higher temperatures at lower pressures than the latter, which record the last stage of equilibrium at moderate pressure–temperature conditions 2.3–3.7 GPa and 855–1095 °C in the upper mantle at the time of entrainment. Although derived from the garnet stability field, these rocks have low-pressure cumulate protoliths containing plagioclase, olivine, and clinopyroxene as reflected by pronounced positive Eu and Sr anomalies in all eclogites, and low heavy rare earth element (HREE) contents in both minerals and reconstructed bulk rocks for a number of samples. Major elements, transition metals, and the HREE compositions of the reconstructed whole rocks are analogous to modern oceanic gabbro cumulates. Despite geochemical signatures supporting an oceanic crust origin, mantle-like  $\delta^{18}\text{O}$  of the garnets (5.07–5.62‰) for most samples indicates that the protoliths either did not interact with seawater or have coincidentally approximately normal igneous values. Some of the eclogite xenoliths have lower  $\text{SiO}_2$  contents and depleted light REE ( $(\text{Nd}/\text{Yb})_{\text{N}} < 1$ ) compared to modern oceanic gabbros, suggesting that they experienced partial melting. Positively inclined middle to heavy REE patterns ( $(\text{Dy}/\text{Yb})_{\text{N}} < 1$ ) of the reconstructed bulk rocks mostly result from repeated partial melting in the eclogite stability field, based on melting model calculations. We, therefore, suggest that the Obnazhennaya low-MgO eclogites may represent the gabbroic section of subducted or foundered basaltic crust that underwent continued partial melting processes at high pressures, where garnet was the main residual phase.

**Keywords** Obnazhennaya · Low-MgO eclogite xenolith · Low-pressure origin · Garnet exsolution · Cumulates

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✉ Jing Sun  
sunjingvv@163.com

✉ Qiao Shu  
shuqiao@mail.gyig.ac.cn

<sup>1</sup> State Key Laboratory of Petroleum Resources and Engineering, Beijing 102249, China

<sup>2</sup> College of Geosciences, China University of Petroleum (Beijing), Beijing 102249, China

<sup>3</sup> Department of Earth Science and Earth Research Institute, University of California, Santa Barbara, CA 93106, USA

<sup>4</sup> Institute of Geochemistry, Siberian Branch of the Russian Academy of Sciences, 664033 Irkutsk, Russia

<sup>5</sup> Institute of Earth Crust, Siberian Branch of the Russian Academy of Sciences, 664033 Irkutsk, Russia

<sup>6</sup> Department of Geoscience, University of Wisconsin-Madison, Madison, WI 53706, USA

<sup>7</sup> State Key Laboratory of Ore Deposit Geochemistry, Institute of Geochemistry, Chinese Academy of Sciences, Guiyang 550081, China

<sup>8</sup> CAS Center for Excellence in Comparative Planetology, Hefei, China

## Introduction

Eclogite is a volumetrically small (< 1%), but important and integral part of the continental lithosphere (Schulze 1989). Xenolithic eclogites exhibit a broad compositional spectrum (Aulbach and Jacob 2016), and their origins could vary accordingly. Several possible origins have been proposed for xenolithic eclogites: (1) cumulates of mantle melt formed at high pressure (e.g., O'Hara 1969); (2) residues of melted subducted oceanic crust (e.g., Ireland et al. 1994; Jacob and Foley 1999; Barth et al. 2001); (3) subducted low-pressure cumulates (e.g., Barth et al. 2002; Shu et al. 2016); and (4) products of melt-rock reaction (e.g., Smart et al. 2009; Wang et al. 2015). Due to their possible origin via ancient subduction, eclogites may hold clues to the past thermal, redox, and chemical state of the ambient convecting mantle (e.g., Aulbach and Stagno 2016; Aulbach and Arndt 2019). Hence, it is important to characterize eclogite xenoliths geochemically to place better constraints on their origins, and the implications for the operation of major earth processes in the past.

Eclogites with low bulk rock MgO contents are generally interpreted to originate as subducted oceanic crust. Examples include low-MgO eclogite xenoliths from South Africa (e.g., Gréau et al. 2011; Huang et al. 2012; Huang et al. 2014; Shu et al. 2013), Finland (Peltonen et al. 2002), West Africa (e.g., Barth et al. 2001; Aulbach et al. 2019), Slave craton, Canada (e.g., Aulbach et al. 2007; Smart et al. 2012) and from Udachnaya pipe in the middle of the Siberian craton (Agashev et al. 2018; and references therein). A further important locality for eclogite and pyroxenite xenoliths, the non-diamondiferous Obnazhennaya kimberlite, is located on the northeastern edge of the Siberian craton. These xenoliths can provide information on the shallower lithosphere evolution history of this region, and provide information about processes occurring at the edge of the platform.

Eclogite or pyroxenite xenoliths from the Obnazhennaya kimberlite pipe can be divided into two broad groups based on their petrological characteristics. One group of samples is biminerally, composed of garnet and clinopyroxene containing lamellae-shaped garnet exsolutions; they were named Group B/C eclogites in previous studies (Taylor et al. 2003) according to their garnet composition. The MgO content of reconstructed bulk rocks of this group is always lower than 16%, and we call this group "low-MgO eclogites". The other group of samples, previously named Group A eclogites (Qi et al. 1994; Taylor et al. 2003; Alifirova et al. 2012, 2015) or websterites (Spengler and Alifirova 2019), contain minor orthopyroxene and rutile in addition to garnet and clinopyroxene. This group of samples has higher MgO content in the reconstructed bulk rock, and is referred to as "high-MgO" here.

A large number of high-MgO eclogites and pyroxenites from Obnazhennaya have previously been studied (Qi et al. 1994; Taylor et al. 2003; Alifirova et al. 2012, 2015). However, far fewer studies have been carried out on the Obnazhennaya low-MgO eclogites.

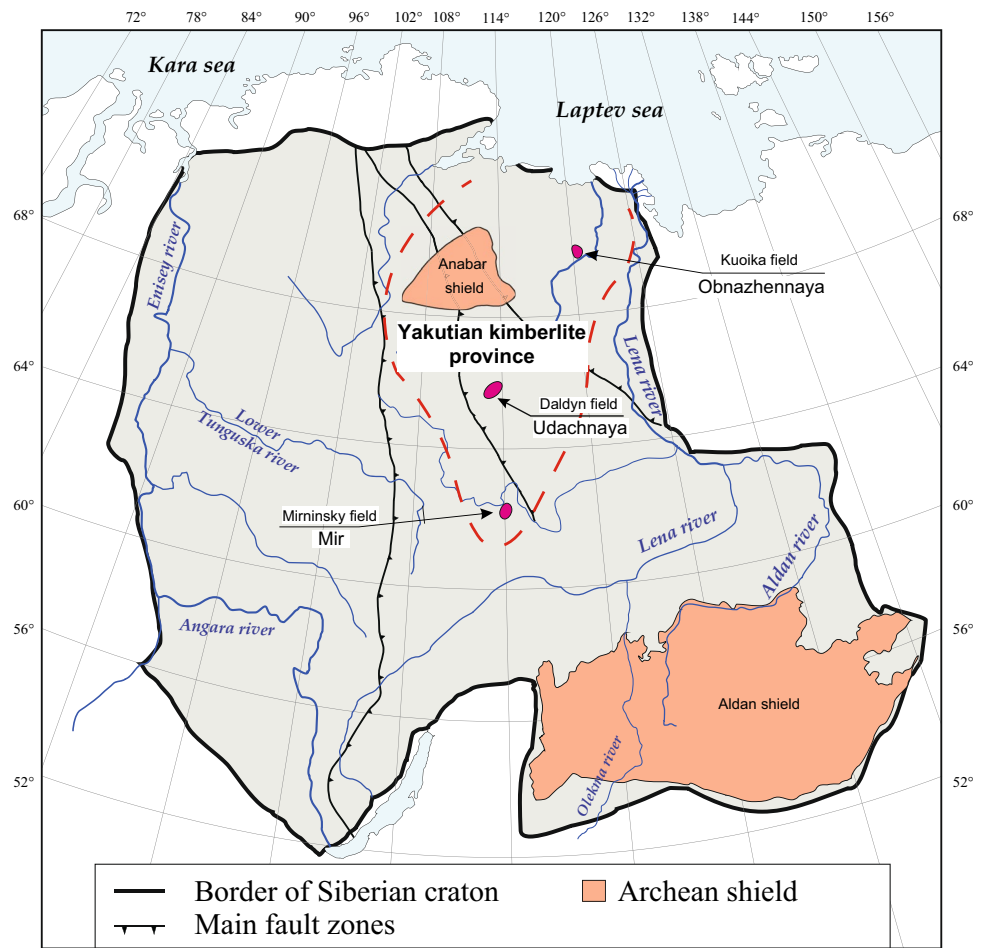
The eclogite xenoliths investigated here contain both granular garnet and garnet exsolutions in host clinopyroxenes, which allows us to place constraints not only on their origin and evolution but also the thermal state of the lithospheric mantle beneath the margin of the Siberian craton. In this paper, we present new major, trace element, and oxygen-isotope results for eight fresh low-MgO Obnazhennaya eclogites. We integrate our new data with the previously published data for three Group B eclogites from this pipe (Taylor et al. 2003) to decipher their origins and garnet exsolution history.

## Geological setting and prior work

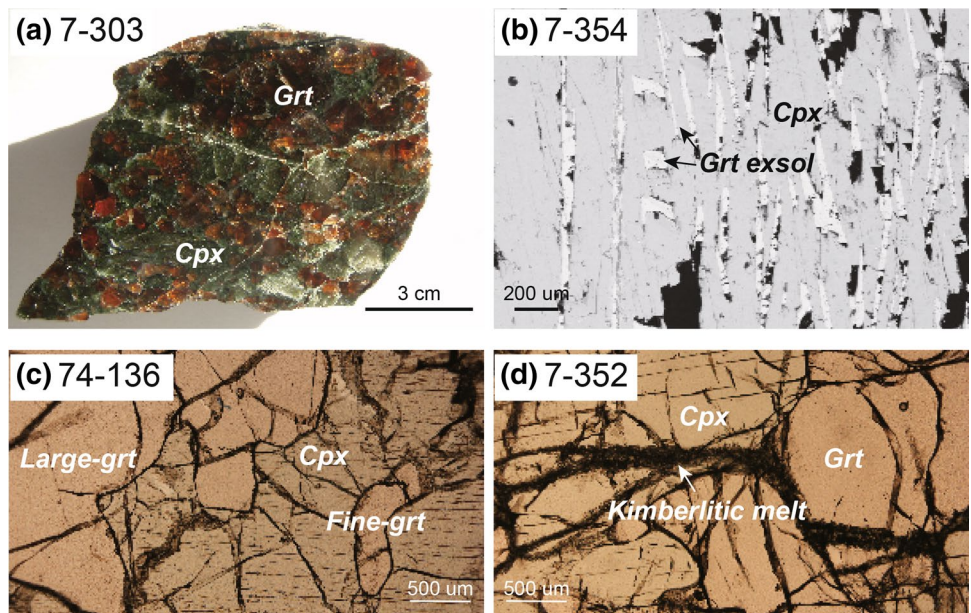
The Siberian platform extends more than 2500 km from Lake Baikal in the south to the Arctic Ocean in the north, and from the Yenisey River in the west to the Aldan shield in the east (Fig. 1). Moho depths beneath the platform range from 35 to 60 km (Rosen et al. 1994). Phanerozoic kimberlites with ages from 420 to 150 Ma (Sun et al. 2014) erupted within the region marked "Yakutian kimberlite Province" in Fig. 1. The Obnazhennaya kimberlite pipe lies within the Kuoika field, which occupies the extreme northeastern portion of the platform (Fig. 1). The age of the Obnazhennaya kimberlite is not well defined. Zircon and perovskite U–Pb dates from other pipes in the Kuoika field range from 156 to 167 Ma (Davis et al. 1980; Kinny et al. 1997; Blanco et al. 2013; Sun et al. 2014). Paleomagnetic studies of the Obnazhennaya pipe suggest an age of 168 Ma (Blanco et al. 2013). Therefore, ~ 160 Ma is assumed to be the eruption age of this pipe (Smelov and Zaitsev 2013).

We collected a suite of fresh samples and chose those without orthopyroxene and with low MgO contents in the reconstructed whole rock for this study. The size of most xenoliths was 4 cm × 6 cm × 4 cm (Fig. 2a). Eclogite and pyroxenite xenoliths from Obnazhennaya have previously been investigated for major, trace elements, oxygen isotope, and Sr–Nd isotopes by Qi et al. (1994); Taylor et al. (2003); Alifirova et al. (2012, 2015) and Spengler and Alifirova (2019). In these studies, nearly all of samples were of group A (or high-MgO eclogites and websterites). Only three samples from Taylor et al. (2003) are Group B eclogites with low MgO contents and without orthopyroxene and are, thus, similar to the samples in this study. No detailed petrographic descriptions for these three samples were reported, but their mineral major and trace-element compositions and garnet oxygen isotopes exist; these data are included and discussed in this study.

**Fig. 1** Simplified geological map of the Siberian platform, which is modified. The Obnazhennaya kimberlite pipe, host of the eclogite xenoliths in this study, is located at the north-eastern margin of the platform. The Udachnaya kimberlite pipe from the Daldyn filed and Mir kimberlite from Mirninsky field are located in the central part of the platform [Modified from Griffin et al. (1999)]



**Fig. 2** Typical textures in the Obnazhennaya eclogite suite. **a** Equigranular texture of garnet and clinopyroxene, sample 7-303; **b** garnet exsolution lamellae in clinopyroxene, sample 7-354; **c** coarse and fine garnets, sample 74-136; **d** a complex network of fractures filled with kimberlitic melt, sample 7-352. *Grt* garnet, *Cpx* clinopyroxene



## Analytical methods

### Modal mineralogy

Modal abundances were obtained by point counting of thin sections. The points were spaced 0.5 mm apart and between 200 and 600 points were counted, depending on the size of the thin section as well as the mineral size. Coarse-grained garnet, fine-grained garnet, garnet exsolution lamellae in clinopyroxene and clinopyroxene were distinguished in the point counting. Melt veins (described below) were omitted when calculating the modal mineralogy.

### Major and trace elements

Major-element compositions of minerals were analyzed with a JEOL JXA-8100 Electron Probe at the National Astronomical Observatories, Chinese Academic of Sciences, using a 15 keV accelerating voltage and a 10 nA beam current. Multiple grains of garnet and clinopyroxene were analyzed from each xenolith, with a minimum of five analyses per mineral grain. A variety of natural and synthetic standards were used to calibrate the major components, some of which were also measured as unknowns to monitor data quality (see Appendix 1). The ZAF model (Armstrong 1995) was applied for matrix correction. The two-sigma analytical precision of most elements (Si, Al, Ca, Fe, Mg, Na, Mn) analyzed by EPMA is ca. 2%.

Trace elements in clinopyroxenes and different types of garnets were analyzed using laser ablation inductively coupled plasma mass spectrometer (LA-ICP-MS) at China University of Geosciences, Beijing. The LA-ICP-MS system consists of a 193 nm pulsed ArF excimer laser coupled to an Agilent 7500 quadrupole ICP-MS. Isotopes were measured in peak-hopping mode. A spot size of 80  $\mu\text{m}$  and a repetition rate of 8 Hz were used. The NIST 612 glass was used as an external calibration standard and the isotope  $^{43}\text{Ca}$  was used as an internal standard. BHOV-2 glass was measured as an unknown to verify data accuracy, and their results are illustrated in Appendix 1. The analytical precision of most elements analyzed by the LA-ICP-MS method is ca. 5–10%.

### Oxygen isotopes by SIMS

Oxygen-isotope compositions of garnet were analyzed in situ using a CAMECA SIMS 1280 ion microprobe at the University of Wisconsin-Madison. Samples were lightly crushed and clean garnets (1–3 mm) were handpicked with a binocular microscope. Separated garnet grains were mounted within 7.5 mm of the center of an epoxy disk with the UWG-2 garnet standard at the center of the mount, and

then polished to a flat, smooth surface. The  $^{133}\text{Cs}^+$  primary ion beam (20 kV total accelerating voltage) was focused to a diameter of  $\sim 10 \mu\text{m}$  on the gold-coated sample surface. Secondary  $\text{O}^-$  ions were accelerated away from the sample by  $\sim 10 \text{ kV}$  and the analysis site was centered under a uniform electron field generated by a normal-incidence electron gun for charging compensation. Due to the different compositions between garnet standard UWG-2 and the Obnazhennaya garnets, a matrix effect is present. To account for the matrix effects, 16 in-house garnet standards from five garnet end members (pyrope, almandine, spessartine, grossular and andradite) were analyzed to calculate the working curve line and instrument bias relative to UWG-2 one day before running the unknowns. The Obnazhennaya garnets were then plotted on the working curve line based on their compositions, and the bias of each unknown relative to UWG-2 was calculated. Details are provided in Page et al. (2010) and Kitajima et al. (2016), and we provide only a short summary here. Four points on UWG-2 were analyzed before and after every 10–15 unknown sample points. The  $\delta^{18}\text{O}_{\text{VSMOW}}$  value of UWG-2 standard is +5.8‰ (Valley et al. 1995). Point-to-point precision for  $\delta^{18}\text{O}$  was typically better than 0.4‰ (2 SD).

## Results

### Petrography

The low-MgO Obnazhennaya eclogites are biminerally, with light to dark green clinopyroxene and orange to pale pink garnets. Accessory sulfides occur in most samples, whereas rutile, ilmenite and orthopyroxene were not observed. The modal mineralogy varies considerably, with garnet proportions of 27–76 vol% and clinopyroxene of 24–73 vol%. Complete petrographic descriptions are provided in Appendix 2.

The eclogites commonly have a planar fabric defined by subhedral to elongated garnets within a matrix of coarse and interlocking clinopyroxenes. Some eclogites have poikilitic microstructures with garnet enclosing clinopyroxene. All textural types of clinopyroxene typically contain needle and/or lamellae-shaped elongate garnets with a shape-preferred orientation (Fig. 2b), up to 0.7 mm across. By contrast, granular garnets contain no exsolution lamellae, and can be subdivided into two types based on their grain sizes: coarse and fine. The coarse-grained garnets (2–10 mm; Fig. 2c) represent 80 vol% of total garnets and have smooth rims, whereas fine-grained garnets (0.05–2 mm; Fig. 2c) have euhedral-to-subhedral texture and are embedded in clinopyroxenes.

All samples are fresh, as alteration products account for < 1 vol% of these samples, although fractures containing kimberlitic melt (Fig. 2d) cross-cut the samples and some



**Table 1** Major element compositions (wt%) of minerals from the Obnazhennaya eclogites

| Sample    | Mineral     | Modal | SiO <sub>2</sub> | TiO <sub>2</sub> | Al <sub>2</sub> O <sub>3</sub> | Cr <sub>2</sub> O <sub>3</sub> | FeO   | MnO  | MgO   | CaO   | Na <sub>2</sub> O | K <sub>2</sub> O | NiO  | Total  | Mg#   | Jd   |
|-----------|-------------|-------|------------------|------------------|--------------------------------|--------------------------------|-------|------|-------|-------|-------------------|------------------|------|--------|-------|------|
| Obn7-303  | Cpx         | 0.485 | 53.76            | 0.11             | 6.32                           | 0.05                           | 2.96  | 0.01 | 13.16 | 19.62 | 2.93              | 0.01             | 0.05 | 98.99  | 88.88 | 0.21 |
|           | Grt-coarse  | 0.439 | 40.43            | 0.04             | 22.73                          | 0.05                           | 13.53 | 0.38 | 13.33 | 8.80  | 0.02              | 0.01             | 0.01 | 99.33  | 63.94 |      |
|           | Grt-fine    | 0.032 | 40.10            | 0.02             | 22.58                          | 0.03                           | 13.26 | 0.38 | 11.96 | 10.82 | 0.01              | 0.00             | 0.01 | 99.17  | 61.87 |      |
|           | Grt-lamella | 0.044 | 40.00            | 0.01             | 22.21                          | 0.06                           | 12.57 | 0.34 | 10.30 | 13.37 | 0.01              | 0.00             | 0.01 | 98.87  | 59.61 |      |
| Obn7-350  | Cpx         | 0.451 | 53.30            | 0.14             | 6.17                           | 0.04                           | 2.61  | 0.04 | 13.44 | 17.62 | 3.25              | 0.01             | 0.12 | 96.72  | 90.27 | 0.25 |
|           | Grt-coarse  | 0.529 | 40.91            | 0.04             | 23.07                          | 0.04                           | 11.33 | 0.20 | 18.21 | 4.41  | 0.01              | 0.01             | 0.02 | 98.25  | 74.32 |      |
|           | Grt-fine    | 0.006 | 40.64            | 0.06             | 22.95                          | 0.03                           | 11.41 | 0.21 | 18.08 | 4.19  | 0.01              | 0.01             | 0.01 | 97.59  | 74.04 |      |
|           | Grt-lamella | 0.014 | 40.45            | 0.03             | 22.81                          | 0.05                           | 11.44 | 0.22 | 17.91 | 3.53  | 0.01              | 0.00             | 0.02 | 96.48  | 73.81 |      |
| Obn7-352  | Cpx         | 0.498 | 53.38            | 0.12             | 8.27                           | 0.10                           | 2.70  | 0.01 | 12.29 | 18.66 | 3.30              | 0.00             | 0.07 | 98.91  | 89.12 | 0.24 |
|           | Grt-coarse  | 0.488 | 41.05            | 0.03             | 22.91                          | 0.08                           | 12.79 | 0.31 | 14.61 | 8.12  | 0.01              | 0.01             | 0.01 | 99.92  | 67.27 |      |
|           | Grt-fine    | 0.014 | 40.71            | 0.04             | 22.60                          | 0.09                           | 12.32 | 0.30 | 13.19 | 9.81  | 0.01              | 0.01             | 0.01 | 99.07  | 65.84 |      |
| Obn7-354  | Cpx         | 0.727 | 53.14            | 0.08             | 6.98                           | 0.05                           | 2.93  | 0.02 | 12.73 | 19.14 | 3.02              | 0.01             | 0.07 | 98.18  | 88.65 | 0.22 |
|           | Grt-coarse  | 0.229 | 40.57            | 0.01             | 22.74                          | 0.03                           | 13.84 | 0.39 | 13.06 | 8.85  | 0.02              | 0.01             | 0.02 | 99.54  | 62.95 |      |
|           | Grt-fine    | 0.011 | 40.15            | 0.02             | 22.50                          | 0.03                           | 12.48 | 0.32 | 10.48 | 12.70 | 0.01              | 0.00             | 0.02 | 98.71  | 60.18 |      |
|           | Grt-lamella | 0.033 | 39.76            | 0.04             | 22.08                          | 0.03                           | 11.97 | 0.28 | 8.96  | 15.45 | 0.01              | 0.01             | 0.01 | 98.60  | 57.40 |      |
| Obn7-379  | Cpx         | 0.241 | 53.30            | 0.18             | 6.46                           | 0.08                           | 3.03  | 0.02 | 13.02 | 19.12 | 3.03              | 0.00             | 0.06 | 98.30  | 88.55 | 0.22 |
|           | Grt-coarse  | 0.728 | 40.45            | 0.05             | 22.66                          | 0.08                           | 13.31 | 0.37 | 13.96 | 8.25  | 0.02              | 0.01             | 0.02 | 99.15  | 65.38 |      |
|           | Grt-fine    | 0.011 | 40.55            | 0.03             | 22.63                          | 0.06                           | 13.18 | 0.35 | 13.00 | 9.68  | 0.01              | 0.01             | 0.00 | 99.51  | 63.96 |      |
|           | Grt-lamella | 0.020 | 39.92            | 0.02             | 22.45                          | 0.07                           | 12.99 | 0.35 | 11.94 | 10.64 | 0.02              | 0.01             | 0.01 | 98.41  | 62.33 |      |
| Obn7-394  | Cpx         | 0.569 | 51.53            | 0.05             | 15.39                          | 0.05                           | 1.12  | 0.00 | 8.90  | 17.66 | 4.44              | 0.01             | 0.13 | 99.28  | 93.46 | 0.31 |
|           | Grt-coarse  | 0.390 | 41.33            | 0.01             | 23.39                          | 0.28                           | 8.04  | 0.17 | 16.55 | 9.04  | 0.01              | 0.01             | 0.02 | 98.85  | 78.74 |      |
|           | Grt-fine    | 0.014 | 41.33            | 0.00             | 23.58                          | 0.05                           | 7.28  | 0.16 | 13.85 | 13.32 | 0.01              | 0.00             | 0.01 | 99.59  | 77.41 |      |
|           | Grt-lamella | 0.027 | 41.81            | 0.03             | 23.63                          | 0.07                           | 5.49  | 0.09 | 9.37  | 21.13 | 0.02              | 0.00             | 0.00 | 101.64 | 75.45 |      |
| Obn74-136 | Cpx         | 0.617 | 53.17            | 0.09             | 6.20                           | 0.04                           | 3.54  | 0.02 | 13.02 | 19.70 | 2.67              | 0.00             | 0.04 | 98.49  | 86.88 | 0.20 |
|           | Grt-coarse  | 0.330 | 40.05            | 0.02             | 22.45                          | 0.03                           | 15.15 | 0.40 | 12.44 | 8.35  | 0.00              | 0.00             | 0.00 | 98.91  | 59.64 |      |
|           | Grt-lamella | 0.053 | 39.80            | 0.03             | 22.32                          | 0.04                           | 14.22 | 0.32 | 9.18  | 13.61 | 0.01              | 0.00             | 0.01 | 99.53  | 53.75 |      |
| Obn74-639 | Cpx         | 0.488 | 53.76            | 0.04             | 6.15                           | 0.01                           | 3.21  | 0.03 | 13.39 | 19.85 | 2.67              | 0.01             | 0.05 | 99.17  | 88.26 | 0.20 |
|           | Grt-coarse  | 0.472 | 40.15            | 0.02             | 22.56                          | 0.01                           | 14.03 | 0.43 | 12.60 | 9.32  | 0.02              | 0.01             | 0.02 | 99.17  | 61.79 |      |
|           | Grt-fine    | 0.024 | 40.15            | 0.02             | 22.55                          | 0.01                           | 13.65 | 0.41 | 11.81 | 10.45 | 0.02              | 0.00             | 0.03 | 99.10  | 60.89 |      |
|           | Grt-lamella | 0.016 | 39.68            | 0.02             | 22.22                          | 0.01                           | 13.28 | 0.37 | 10.30 | 12.77 | 0.02              | 0.00             | 0.00 | 98.66  | 58.27 |      |

cpx clinopyroxene, grt garnet, grt-coarse garnet coarse grain with size larger than 2 mm, grt-fine garnet fine grain with size ranging from 0.05 to 2 mm, grt-lamella garnet lamellae in clinopyroxene, Jd Jadeite

mineral grains. Moreover, neither serpentine nor phlogopite was observed in any samples.

### Mineral major elements

Clinopyroxenes are homogeneous within the analytical uncertainty (Table 1). All eight samples have jadeite (Na/(Na+Ca) component in clinopyroxene > 0.2, indicating that all low-MgO samples are true eclogites (Clark et al. 1968), rather than garnet clinopyroxenites. Compared with the Obnazhennaya Group B eclogites previously studied (Taylor et al. 2003), clinopyroxenes in our samples have lower FeO contents (1.1–3.5 wt% vs. 4.3–5.1 wt%) and higher Mg# (Mg# = 100\*Mg/(Mg+Fe); 86.9–93.5 vs. 79.5–83.8). In addition, the clinopyroxenes in this study are slightly less jadeite-rich (0.20–0.31) compared to those of the

previous study (0.29–0.35; Taylor et al. 2003). Clinopyroxene in sample 7-394 contains the highest Al<sub>2</sub>O<sub>3</sub> (15.4 wt%) and Mg# (93.5), and the lowest MgO (8.90 wt%) and FeO (1.12 wt%) compared with the rest of the samples.

Garnets of different textural types are compositionally distinct, with the exception of those in sample 7-350. Compared to the coarse-grained garnets, fine-grained and lamellar garnets in clinopyroxene have higher CaO (9.7–21.1 wt% vs. 8.1–9.3 wt%) and lower MgO (9.0–13.9 wt% vs. 12.4–16.6 wt%). Lamellar garnets in each sample have the highest CaO (10.6–21.1 wt%) and lowest MgO (9.0–11.9 wt%) (Table 1; Appendix 3). Among these, garnet in sample 7-394 has the lowest FeO (5.5–8.0 wt%) and highest Mg# (75.4–78.7) of all garnet types. By contrast, garnets of different sizes are homogeneous in sample 7-350, and they have the highest

**Table 2** Trace-element composition of garnets, clinopyroxenes from Obnazhennaya eclogites in ppm

|     | 7-303-cpx | 7-303-grt-coarse | 7-303-grt-fine | 7-352-cpx | 7-352-grt-coarse | 7-352-grt-fine | 7-354-cpx | 7-354-grt-coarse | 7-354-grt-fine | 7-379-cpx | 7-379-grt-coarse | 7-379-grt-fine |
|-----|-----------|------------------|----------------|-----------|------------------|----------------|-----------|------------------|----------------|-----------|------------------|----------------|
| Vol | 0.529     | 0.439            | 0.032          | 0.498     | 0.488            | 0.014          | 0.727     | 0.262            | 0.011          | 0.261     | 0.728            | 0.011          |
| V   | 782       | 299              | 188            | 435       | 176              | 137            | 785       | 227              | 256            | 849       | 289              | 243            |
| Cr  | 319       | 287              | 151            | 626       | 683              | 569            | 192       | 188              | 200            | 427       | 451              | 434            |
| Th  | 0.070     | 0.001            | 0.001          | 0.011     | 0.001            | 0.001          | 0.070     | 0.002            | 0.002          | 0.049     | 0.001            | 0.17           |
| U   | 0.049     | 0.002            | 0.002          | 0.008     | 0.001            | 0.001          | 0.054     | 0.004            | 0.004          | 0.035     | 0.002            | 0.004          |
| Nb  | 0.66      | 0.010            | 0.009          | 0.020     | 0.003            | 0.002          | 0.82      | 0.025            | 0.031          | 0.35      | 0.011            | 0.015          |
| Ta  | 0.015     | 0.001            | 0.001          | 0.001     | 0.001            | 0.011          | 0.026     | 0.001            | 0.001          | 0.009     | 0.001            | 0.001          |
| La  | 0.653     | 0.005            | 0.002          | 0.084     | bdl              | 0.002          | 1.08      | 0.003            | 0.003          | 0.61      | 0.004            | 0.005          |
| Ce  | 1.21      | 0.009            | 0.007          | 0.14      | 0.002            | 0.003          | 2.15      | 0.013            | 0.019          | 1.05      | 0.008            | 0.011          |
| Pr  | 0.11      | 0.004            | 0.004          | 0.026     | 0.003            | 0.001          | 0.22      | 0.005            | 0.006          | 0.11      | 0.005            | 0.006          |
| Sr  | 29.5      | 0.025            | 0.07           | 9.78      | 0.014            | 0.011          | 40.2      | 0.020            | 0.028          | 30.5      | 0.020            | 0.037          |
| Nd  | 0.49      | 0.050            | 0.05           | 0.25      | 0.070            | 0.065          | 0.78      | 0.066            | 0.11           | 0.57      | 0.077            | 0.082          |
| Sm  | 0.10      | 0.14             | 0.09           | 0.19      | 0.23             | 0.18           | 0.078     | 0.095            | 0.14           | 0.13      | 0.15             | 0.15           |
| Zr  | 0.61      | 0.57             | 0.30           | 0.87      | 1.23             | 0.50           | 0.61      | 0.29             | 0.32           | 1.70      | 1.08             | 0.79           |
| Hf  | 0.033     | 0.023            | 0.01           | 0.054     | 0.051            | 0.018          | 0.029     | 0.008            | 0.010          | 0.11      | 0.030            | 0.012          |
| Eu  | 0.040     | 0.14             | 0.09           | 0.080     | 0.25             | 0.20           | 0.041     | 0.092            | 0.13           | 0.079     | 0.16             | 0.15           |
| Ti  | 662       | 272              | 157            | 948       | 393              | 252            | 513       | 170              | 210            | 875       | 317              | 254            |
| Gd  | 0.072     | 0.67             | 0.35           | 0.15      | 1.04             | 0.74           | 0.077     | 0.433            | 0.47           | 0.17      | 0.73             | 0.62           |
| Tb  | 0.008     | 0.21             | 0.14           | 0.018     | 0.29             | 0.19           | 0.006     | 0.154            | 0.12           | 0.021     | 0.23             | 0.19           |
| Dy  | 0.021     | 2.29             | 1.10           | 0.079     | 2.48             | 1.33           | 0.033     | 1.59             | 1.01           | 0.10      | 2.32             | 1.76           |
| Y   | 0.061     | 17.8             | 8.41           | 0.19      | 15.1             | 7.58           | 0.088     | 12.3             | 5.63           | 0.36      | 17.5             | 12.8           |
| Ho  | 0.003     | 0.67             | 0.31           | 0.007     | 0.62             | 0.30           | 0.003     | 0.45             | 0.22           | 0.015     | 0.64             | 0.48           |
| Er  | 0.005     | 2.14             | 0.97           | 0.018     | 1.81             | 0.77           | 0.006     | 1.53             | 0.55           | 0.030     | 2.18             | 1.58           |
| Tm  | 0.001     | 0.36             | 0.15           | 0.002     | 0.26             | 0.11           | 0.002     | 0.25             | 0.070          | 0.003     | 0.34             | 0.24           |
| Yb  | 0.003     | 2.62             | 1.07           | 0.007     | 2.03             | 0.70           | 0.004     | 1.90             | 0.36           | 0.015     | 2.77             | 1.69           |
| Lu  | 0.001     | 0.41             | 0.15           | 0.001     | 0.27             | 0.10           | 0.001     | 0.30             | 0.046          | 0.002     | 0.41             | 0.28           |

| Element | 7-394-cpx | 7-394-grt-coarse | 7-394-grt-fine | 74-136-cpx | 74-136-grt-coarse | 74-136-grt-lamellae | 74-639-cpx | 74-639-grt-coarse | 74-639-grt-fine | 7-350-cpx | 7-350-grt-coarse |
|---------|-----------|------------------|----------------|------------|-------------------|---------------------|------------|-------------------|-----------------|-----------|------------------|
| Vol     | 0.596     | 0.390            | 0.014          | 0.617      | 0.330             | 0.053               | 0.504      | 0.472             | 0.024           | 0.465     | 0.535            |
| V       | 183       | 51.0             | 55.0           | 509        | 197               | 143                 | 779        | 271               | 209             | 148       | 37.6             |
| Cr      | 407       | 255              | 299            | 247        | 231               | 198                 | 35.6       | 38.8              | 25.8            | 262       | 289              |
| Th      | 0.002     | 0.004            | 0.001          | 0.005      | bdl               | 0.001               | 0.003      | 0.001             | bdl             | 0.024     | 0.004            |
| U       | 0.004     | 0.11             | 0.001          | 0.003      | bdl               | 0.002               | 0.001      | 0.001             | bdl             | 0.007     | 0.003            |
| Nb      | 0.014     | 0.042            | 0.006          | 0.010      | 0.003             | 0.007               | 0.007      | 0.002             | bdl             | 0.036     | 0.005            |
| Ta      | 0.001     | 0.001            | 0.001          | 0.000      | bdl               | 0.0005              | 0.001      | 0.001             | 0.001           | 0.004     | 0.001            |
| La      | 0.004     | 0.003            | 0.024          | 0.052      | 0.002             | 0.002               | 0.15       | bdl               | 0.005           | 0.75      | 0.033            |

Table 2 (continued)

| Element | 7-394-cpx | 7-394-grt-coarse | 7-394-grt-fine | 74-136-cpx | 74-136-grt-coarse | 74-136-grt-lamellae | 74-639-cpx | 74-639-grt-coarse | 74-639-grt-fine | 7-350-cpx | 7-350-grt-coarse |
|---------|-----------|------------------|----------------|------------|-------------------|---------------------|------------|-------------------|-----------------|-----------|------------------|
| Ce      | 0.032     | 0.005            | 0.006          | 0.069      | 0.002             | 0.002               | 0.48       | 0.011             | 0.009           | 1.82      | 0.030            |
| Pr      | 0.021     | 0.004            | 0.010          | 0.008      | 0.003             | 0.001               | 0.070      | 0.008             | 0.004           | 0.27      | 0.008            |
| Sr      | 11.1      | 0.061            | 0.065          | 3.60       | bdl               | 0.013               | 22.3       | 0.032             | 0.025           | 168       | 0.047            |
| Nd      | 0.21      | 0.074            | 0.165          | 0.043      | 0.011             | 0.016               | 0.27       | 0.045             | 0.035           | 1.56      | 0.077            |
| Sm      | 0.068     | 0.17             | 0.19           | 0.051      | 0.046             | 0.052               | 0.040      | 0.056             | 0.061           | 0.57      | 0.16             |
| Zr      | 0.20      | 0.82             | 0.50           | 0.13       | 0.075             | 0.042               | 0.22       | 0.14              | 0.13            | 7.32      | 2.61             |
| Hf      | 0.015     | 0.032            | 0.020          | 0.014      | 0.006             | 0.003               | 0.018      | 0.012             | 0.003           | 0.24      | 0.040            |
| Eu      | 0.027     | 0.21             | 0.22           | 0.036      | 0.085             | 0.077               | 0.028      | 0.092             | 0.073           | 0.31      | 0.23             |
| Ti      | 364       | 107              | 106            | 510        | 197               | 141                 | 353        | 172               | 112             | 761       | 278              |
| Gd      | 0.023     | 0.41             | 0.38           | 0.088      | 0.42              | 0.33                | 0.070      | 0.47              | 0.35            | 0.40      | 0.58             |
| Tb      | 0.003     | 0.11             | 0.073          | 0.011      | 0.13              | 0.079               | 0.012      | 0.18              | 0.12            | 0.040     | 0.13             |
| Dy      | 0.014     | 0.83             | 0.46           | 0.056      | 1.12              | 0.62                | 0.046      | 1.73              | 1.00            | 0.13      | 0.91             |
| Y       | 0.017     | 5.30             | 2.34           | 0.16       | 7.88              | 3.67                | 0.16       | 13.4              | 6.63            | 0.36      | 4.71             |
| Ho      | 0.002     | 0.21             | 0.091          | 0.007      | 0.29              | 0.15                | 0.007      | 0.52              | 0.24            | 0.016     | 0.20             |
| Er      | 0.003     | 0.66             | 0.24           | 0.011      | 0.87              | 0.34                | 0.014      | 1.65              | 0.67            | 0.025     | 0.50             |
| Tm      | 0.002     | 0.11             | 0.036          | 0.002      | 0.13              | 0.043               | 0.001      | 0.26              | 0.085           | 0.003     | 0.063            |
| Yb      | 0.006     | 0.82             | 0.25           | 0.008      | 0.90              | 0.23                | 0.004      | 1.91              | 0.55            | 0.018     | 0.47             |
| Lu      | 0.002     | 0.13             | 0.033          | 0.002      | 0.13              | 0.030               | 0.001      | 0.28              | 0.073           | 0.001     | 0.078            |

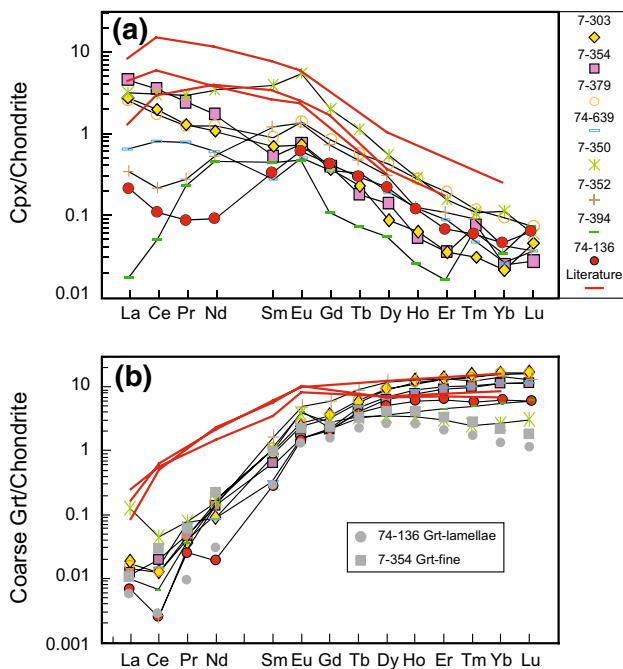
bd/ below detection limit

MgO (17.9–18.2 wt%), but the lowest CaO (3.5–4.4 wt%) contents.

## Trace elements

The trace-element compositions of minerals from Obnazhennaya eclogites are given in Table 2 and presented as chondrite-normalized (McDonough and Sun 1995) REE (rare earth element) plots in Fig. 3, along with previously published data (Taylor et al. 2003).

Clinopyroxenes exhibit variable REE concentrations and patterns. All are convex-upwards, but some are



**Fig. 3** **a** Chondrite-normalized REE patterns of clinopyroxenes from the Obnazhennaya low-MgO eclogites; **b** chondrite-normalized REE patterns of coarse garnets from the Obnazhennaya low-MgO eclogites. Grey cycles and squares are REE patterns of garnet lamellae in clinopyroxene from 74-136 and fine garnet from 7-354. Red lines in both **a** and **b** are Obnazhennaya Group B eclogites from Taylor et al. (2003). Chondrite values from McDonough and Sun (1995)

LREE-enriched, while others are LREE-depleted, and all exhibit positive Eu anomalies relative to their neighboring REE with  $\text{Eu}/\text{Eu}^*$  (chondrite-normalized  $\text{Eu}/(\text{Sm} \times \text{Gd})^{0.5}$ ) ranging from 1.4 to 2.4 (Fig. 3a). The REE patterns of clinopyroxenes in this study overlap with those in the previous study, but ours have overall lower REE concentrations and higher  $\text{Eu}/\text{Eu}^*$ .

Coarse-grained garnets show typical LREE-depleted chondrite-normalized REE patterns (Fig. 3b), with  $\text{La}_N$  of 0.1–2.3 and  $\text{Yb}_N$  of 1.6–9.8. These samples have similar HREE contents as garnets in the previous study (Taylor et al. 2003), but our samples have lower LREE contents. Coarse-grained garnets from both this and the previous study show variable, positive Eu anomalies (Fig. 3b), with  $\text{Eu}/\text{Eu}^*$  of 1.4–2.4. Fine-grained garnets from six samples were analyzed, but garnet lamellae in clinopyroxene could only be analyzed in one sample due to the typically small sizes. The fine-grained and lamellar garnet have similar REE patterns to each other, with overall lower HREE contents and negative sloped HREE patterns compared to the coarse-grained garnets (Fig. 3b).

## Reconstructed whole-rock compositions

Bulk rock chemical analyses of kimberlite-hosted xenolithic eclogites have been shown to represent mixtures of the original eclogite and host kimberlite (e.g., Barth et al. 2001; Jacob et al. 2005). We, therefore, calculated bulk rock compositions for major and trace elements from garnet and clinopyroxene using modal abundances (determined by point counting) and the results are listed in Tables 3 and 4, respectively. Garnets of all textural types are all accounted for in such calculations for both major and trace elements. Modal abundances of garnet lamellae in clinopyroxenes for some samples were considered part of the clinopyroxene mode while calculating their trace elements. Whole-rock reconstruction is not particularly sensitive to variations in modes as the calculated whole-rock trace-element patterns do not change significantly when the garnet (grt): clinopyroxene (cpx) ratios vary between 50:50 and 70:30 (Jacob and Foley 1999).

**Table 3** Estimated major element compositions (wt%) of reconstructed bulk rocks of Obnazhennaya eclogites

| Sample | SiO <sub>2</sub> | TiO <sub>2</sub> | Al <sub>2</sub> O <sub>3</sub> | Cr <sub>2</sub> O <sub>3</sub> | FeO   | MnO  | MgO   | CaO   | Na <sub>2</sub> O | K <sub>2</sub> O | NiO  | Total | Mg#  |
|--------|------------------|------------------|--------------------------------|--------------------------------|-------|------|-------|-------|-------------------|------------------|------|-------|------|
| 7-303  | 46.86            | 0.07             | 14.74                          | 0.05                           | 8.36  | 0.20 | 13.07 | 14.31 | 1.43              | 0.01             | 0.03 | 99.14 | 73.8 |
| 7-350  | 46.49            | 0.08             | 15.45                          | 0.04                           | 7.40  | 0.13 | 16.05 | 10.36 | 1.47              | 0.01             | 0.06 | 97.53 | 79.6 |
| 7-352  | 47.18            | 0.07             | 15.61                          | 0.09                           | 7.76  | 0.16 | 13.43 | 13.39 | 1.65              | 0.00             | 0.04 | 99.40 | 75.7 |
| 7-354  | 49.68            | 0.06             | 11.26                          | 0.04                           | 5.84  | 0.12 | 12.66 | 16.59 | 2.20              | 0.01             | 0.06 | 98.51 | 79.6 |
| 7-379  | 43.53            | 0.08             | 18.76                          | 0.08                           | 10.82 | 0.28 | 13.68 | 10.93 | 0.74              | 0.01             | 0.03 | 98.93 | 69.5 |
| 7-394  | 47.15            | 0.03             | 18.85                          | 0.14                           | 4.02  | 0.07 | 11.97 | 14.33 | 2.53              | 0.01             | 0.08 | 99.18 | 84.3 |
| 74-136 | 48.13            | 0.06             | 12.41                          | 0.04                           | 7.94  | 0.16 | 12.62 | 15.63 | 1.65              | 0.00             | 0.03 | 98.68 | 74.1 |
| 74-639 | 46.78            | 0.03             | 14.54                          | 0.01                           | 8.73  | 0.24 | 12.93 | 14.54 | 1.31              | 0.01             | 0.03 | 99.16 | 72.7 |



**Table 4** Estimated trace-element compositions (ppm) of reconstructed bulk rocks of Obnazhennaya eclogites

| Sample | 7-303 | 7-350 | 7-352 | 7-354 | 7-379 | 7-394 | 74-136 | 74-639 |
|--------|-------|-------|-------|-------|-------|-------|--------|--------|
| V      | 545   | 88.9  | 304   | 633   | 434   | 130   | 389    | 526    |
| Ni     | 240   | 478   | 311   | 443   | 208   | 641   | 220    | 271    |
| Cr     | 295   | 277   | 653   | 191   | 445   | 346   | 241    | 36.9   |
| Th     | 0.038 | 0.013 | 0.006 | 0.052 | 0.015 | 0.003 | 0.003  | 0.002  |
| U      | 0.027 | 0.005 | 0.004 | 0.040 | 0.010 | 0.045 | 0.002  | 0.001  |
| Nb     | 0.35  | 0.019 | 0.011 | 0.60  | 0.10  | 0.025 | 0.007  | 0.005  |
| Ta     | 0.008 | 0.002 | 0.001 | 0.019 | 0.003 | 0.001 | 0.0003 | 0.001  |
| La     | 0.35  | 0.37  | 0.042 | 0.78  | 0.16  | 0.004 | 0.033  | 0.078  |
| Ce     | 0.64  | 0.86  | 0.068 | 1.57  | 0.28  | 0.021 | 0.043  | 0.25   |
| Pr     | 0.062 | 0.13  | 0.014 | 0.16  | 0.033 | 0.014 | 0.006  | 0.039  |
| Sr     | 15.6  | 78.3  | 4.87  | 29.2  | 7.97  | 6.63  | 2.22   | 11.3   |
| Nd     | 0.28  | 0.76  | 0.16  | 0.59  | 0.20  | 0.16  | 0.031  | 0.16   |
| Sm     | 0.12  | 0.35  | 0.21  | 0.083 | 0.15  | 0.11  | 0.050  | 0.048  |
| Zr     | 0.57  | 4.80  | 1.04  | 0.53  | 1.24  | 0.44  | 0.11   | 0.18   |
| Hf     | 0.027 | 0.13  | 0.052 | 0.024 | 0.050 | 0.022 | 0.011  | 0.015  |
| Eu     | 0.083 | 0.27  | 0.16  | 0.056 | 0.14  | 0.10  | 0.055  | 0.059  |
| Ti     | 470   | 502   | 667   | 419   | 462   | 260   | 389    | 262    |
| Gd     | 0.33  | 0.50  | 0.59  | 0.17  | 0.58  | 0.18  | 0.21   | 0.26   |
| Tb     | 0.10  | 0.088 | 0.15  | 0.046 | 0.18  | 0.045 | 0.056  | 0.094  |
| Dy     | 1.02  | 0.54  | 1.27  | 0.45  | 1.74  | 0.34  | 0.45   | 0.87   |
| Y      | 7.83  | 2.69  | 7.56  | 3.35  | 13.0  | 2.11  | 2.97   | 6.56   |
| Ho     | 0.30  | 0.12  | 0.31  | 0.12  | 0.48  | 0.084 | 0.11   | 0.26   |
| Er     | 0.94  | 0.28  | 0.90  | 0.41  | 1.61  | 0.26  | 0.32   | 0.80   |
| Tm     | 0.16  | 0.035 | 0.13  | 0.068 | 0.25  | 0.044 | 0.046  | 0.12   |
| Yb     | 1.15  | 0.26  | 1.01  | 0.50  | 2.04  | 0.33  | 0.32   | 0.92   |
| Lu     | 0.18  | 0.042 | 0.13  | 0.079 | 0.30  | 0.052 | 0.047  | 0.14   |

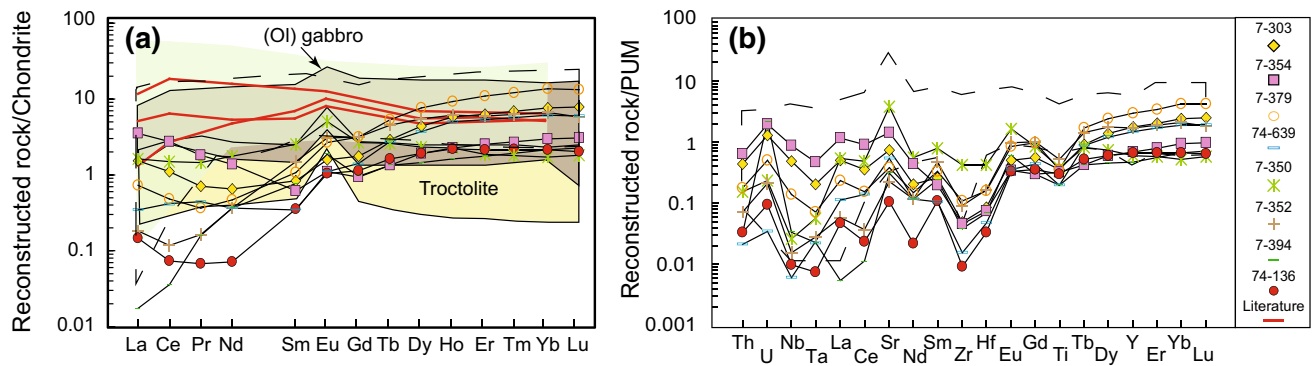
Reconstructed bulk major element compositions of the Obnazhennaya low-MgO eclogites have Mg-numbers ranging from 69.5 to 84.3, overlapping, but ranging to higher values than the three Group B samples from the previous study (60.8–74.4). Bulk CaO concentrations are high, from 10.9 to 16.6 wt%. Sample 7-394 has the highest Al<sub>2</sub>O<sub>3</sub> (18.9 wt%) and Na<sub>2</sub>O (2.5 wt%) contents and the highest Mg# (84.3). Sample 7-350 has the highest MgO (16.1 wt%) content, whereas most others have MgO contents ranging from 12.0 to 13.7 wt%.

All Obnazhennaya low-MgO eclogites except 7-350 have flat to positively sloped HREEs (heavy rare earth elements), MREE (medium rare earth elements)-depleted REE patterns ((Nd/Yb)<sub>N</sub> = 0.03–0.46) and positive Eu anomalies (Fig. 4a). Most of the samples are LREE-enriched, whereas others have flat to depleted LREE patterns. All samples are also enriched in Sr (positive Sr anomaly) relative to elements of similar incompatibility in basaltic systems (Fig. 4b). Sample 7-350, with the largest Eu and Sr anomaly, has a flat REE pattern with (Nd/Yb)<sub>N</sub> = 1.0. The MREE and HREE abundances of the eclogites are both lower and higher than those of the primitive mantle (Fig. 4b). Three Group B samples from the previous study show similar HREE contents, but

higher LREE concentrations compared to samples in this study (Qi et al. 1994; Taylor et al. 2003). Rutile dominates the budget of TiO<sub>2</sub>, Nb, Ta, Zr and Hf in eclogites (Rudnick et al. 2000) and negative anomalies for these elements in the reconstructed bulk rocks suggest that rutile may exist in Obnazhennaya low-MgO eclogites, but was not present in the thin sections we studied. Because of these uncertainties, the calculated TiO<sub>2</sub>, Nb and Ta in the reconstructed bulk rocks will not be used in further discussion.

### Oxygen isotopes

In eclogite xenoliths, garnet is considered to preserve oxygen isotopic ratios more reliably than clinopyroxene for two reasons (1) oxygen diffuses more slowly in garnet than clinopyroxene at mantle temperatures (Farver 2010; Zhang et al. 2016) and (2) garnet is more resistant to overprinting processes than clinopyroxene (Deines and Haggerty 2000). Deines et al. (1991) found that the  $\Delta^{18}\text{O}$  between garnet and clinopyroxene varies with temperature. Samples with equilibrium temperature of 800–1000 °C have 0.6–0.4‰  $\Delta^{18}\text{O}_{\text{grt-cpx}}$  (Fig. 8 in Radu et al. 2019), and the resultant difference between garnet and whole rock is less than 0.5‰



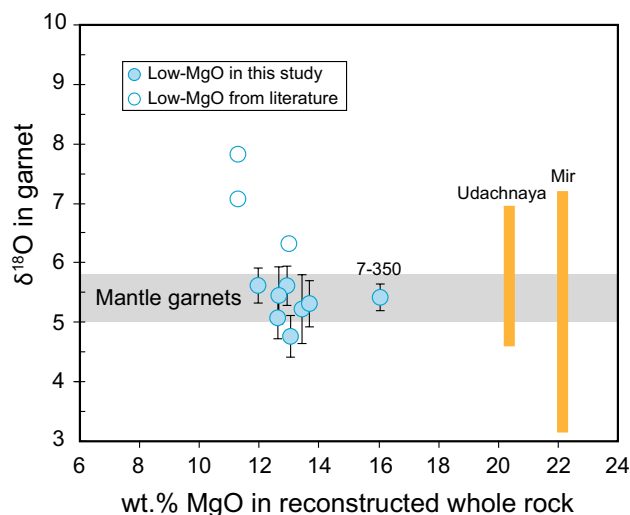
**Fig. 4** **a** Chondrite-normalized reconstructed bulk REE concentrations of low-MgO eclogites estimated from the REE concentrations in minerals and their modes from Table 2. The compositional field of olivine gabbros from the MARK region, mid-Atlantic ridge is shown as the brown field (Casey 1997) and the range of troctolite compositions from IODP Site U1309 in Mid-Atlantic Ridge (Godard et al. 2009) and the Pito Deep in the Pacific (Perk et al. 2007) is shown as a yellow field. Dashed line field and light green field mark the composition of Udachnaya low-MgO eclogites and Mir low-MgO eclogite

xenoliths, respectively. Data of low-MgO eclogites from Udachnaya and Mir kimberlite is from Agashev et al. (2018) and Beard et al. (1996), respectively. Red lines are Obnazhennaya Group B eclogites from Taylor et al. (2003). Data are normalized to chondrite values from McDonough and Sun (1995). **b** Primitive mantle-normalized reconstructed whole-rock trace-element diagram for Obnazhennaya low-MgO eclogites. Data are normalized to the primitive mantle of McDonough and Sun (1995). Dash line field represents the Udachnaya low-MgO eclogite xenoliths (Agashev et al. 2018)

(based on the function:  $\delta^{18}\text{O}_{\text{whole rock}} = \delta^{18}\text{O}_{\text{grt}} * \text{grt vol\%} + (\delta^{18}\text{O}_{\text{grt}} - \Delta^{18}\text{O}_{\text{grt-cpx}}) * \text{cpx vol\%}$ ). Therefore, the  $\delta^{18}\text{O}$  value in garnet approximates the  $\delta^{18}\text{O}$  value of the bulk rock (Korolev et al. 2018). Hence, garnet oxygen analyses

are generally used to represent eclogite whole-rock oxygen isotopic compositions.

Both coarse- and fine-grained garnets were analyzed, and in all cases, the oxygen-isotope compositions of different sized garnets are identical, within error. Garnets in this study have a restricted range of  $\delta^{18}\text{O}$  values (4.73–5.62‰; Table 5), most of which lie within the mantle range ( $\delta^{18}\text{O} = 5.5 \pm 0.4\text{‰}$ ; Matthey et al. 1994; Fig. 5). However, the three Group B eclogites from the previous study have heavier garnet  $\delta^{18}\text{O}$  values ranging from 6.3 to 7.8‰ (Taylor et al. 2003).



**Fig. 5** MgO in reconstructed whole-rock vs.  $\delta^{18}\text{O}$  in garnet for Obnazhennaya low-MgO eclogites. The grey filled depicts the range of mantle garnets that shifts towards lighter isotopic composition by 0.1‰ relative to the whole-rock mantle range due to the fractionation between garnet and whole rock at 1000–1200 °C (Deines et al. 1991). Vertical error bars are two sigma uncertainties for the oxygen isotopes. Open circles are Obnazhennaya Group B eclogites from Taylor et al. (2003).  $\delta^{18}\text{O}$  in garnet for Udachnaya low-MgO eclogites are from Jacob et al. (1994) and Shatsky et al. (2016) and data for Mir low-MgO eclogites are from Beard et al. (1996)

## Geothermobarometry

Temperatures were calculated using the garnet–clinopyroxene  $\text{Fe}^{2+}\text{-Mg}^{2+}$  exchange thermometer of Krogh (1988) and Krogh (2000) and pressures were calculated with the garnet–clinopyroxene geobarometer of Beyer et al. (2015). The Beyer et al. (2015) barometer formulation should only be applied to clinopyroxenes with low jadeite component and Si cations per formula unit below 1.985, as it is the case for all Obnazhennaya low-MgO eclogites. We applied these thermobarometers to the average EPMA analyses for clinopyroxenes and the garnet lamellae in clinopyroxenes, which should record the last stage of equilibration in the mantle.  $\text{Fe}^{3+}$  in garnet and clinopyroxene should be low to very low, as reflected by the sum of cations. Furthermore, Purwin et al. (2013) found, from high pressure experiments, that the partitioning of  $\text{Fe}^{3+}$  between the two phases is such that calculated temperatures are not significantly different

**Table 5** Oxygen-isotope composition ( $\delta^{18}\text{O}_{\text{VSMOW}}\%$ ) of garnets from Obnazhennaya eclogites

| Sample | $\delta^{18}\text{O}_{\text{VSMOW}}\%$ | 2SD  |
|--------|--|------|
| 7-303  | 4.73                                   | 0.37 |
| 7-352  | 5.22                                   | 0.58 |
| 74-639 | 5.61                                   | 0.33 |
| 7-379  | 5.31                                   | 0.39 |
| 7-354  | 5.45                                   | 0.48 |
| 7-394  | 5.62                                   | 0.29 |
| 74-136 | 5.07                                   | 0.35 |
| 7-350  | 5.42                                   | 0.22 |

$$\delta^{18}\text{O}_{\text{VSMOW}}\% = \left[ \frac{(^{18}\text{O}/^{16}\text{O})_{\text{sample}}}{(^{18}\text{O}/^{16}\text{O})_{\text{SMOW}}} - 1 \right] \times 1000$$

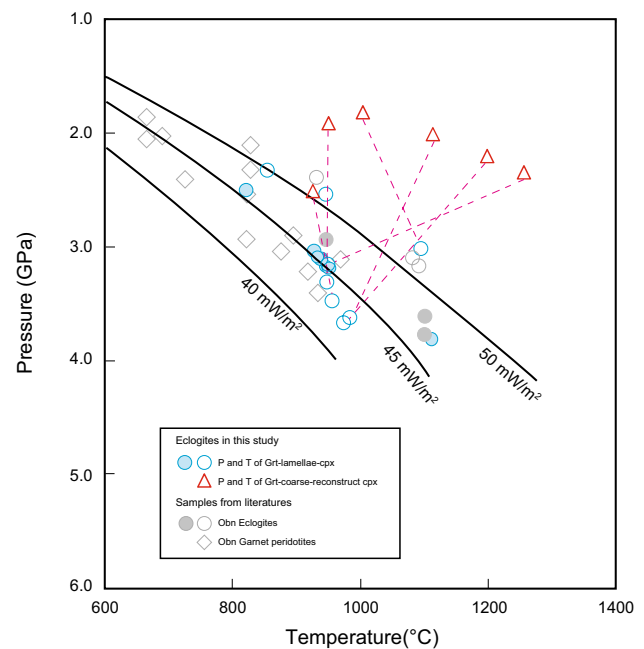
SMOW standard mean ocean water

if either  $\text{Fe}^{\text{total}}$  is used or if  $\text{Fe}^{3+}$  in both garnet and clinopyroxene are considered. Therefore,  $\text{Fe}^{\text{total}}$  was used in the calculations.

The calculated equilibrium temperatures and pressures are given in Table 6 and plotted in Fig. 6. Pressures and temperatures of the three Group B eclogites from Taylor et al. (2003) were calculated here as well using the above thermobarometers. Pressures and temperatures of garnet peridotites from Obnazhennaya pipe (Howarth et al. 2014) are also plotted in Fig. 6. They scatter around a conductive geotherm of about  $45 \text{ mW/m}^2$  (Chapman and Pollack 1977).

The temperatures and pressures of last equilibration, calculated from the garnet lamellae and their host clinopyroxenes range from 855 to 1095 °C and 2.3–3.7 GPa. The scatter for the eclogite PT conditions is much larger than for the garnet peridotites, which may be due to the larger uncertainty associated with the geothermobarometer combinations for eclogites relative to those for peridotites. Part of the eclogite PT conditions overlap with the geothermal gradient for peridotites, others are significantly higher. It appears that, overall, the Beyer et al. (2015) barometer tends to yield somewhat lower pressures and hence warmer geothermal gradient.

We further calculated PT conditions for the combination of reconstructed clinopyroxene (Table 7) + coarse-grained garnet, which may record the equilibrium conditions of the original eclogite-facies metamorphism. The results are also plotted in Fig. 6 (open red triangles) and connected by tie lines with their respective last equilibration, as recorded by the garnet lamellae PT conditions. The temperatures of the latter are lower by around 180 °C, on average, and the pressure is higher by ~1.5 GPa (Table 6).



**Fig. 6** Pressures and temperatures of the last equilibration of Obnazhennaya xenolithic low-MgO eclogites. Blue circles and red triangles represent Obnazhennaya eclogites in this study. Temperatures and pressures of blue circles are calculated from garnet-lamellae/garnet-fine (7-352)—host clinopyroxene, and the red triangles are calculated from garnet-coarse—reconstruct clinopyroxene. Grey circles are Obnazhennaya Group B eclogites calculated using the data from Taylor et al. (2003). Temperatures were calculated using the cpx-grt Fe–Mg exchange thermometer of Krogh (1988). Open symbols represent that pressures calculated using the barometer of Beyer et al. (2015), which is based on the incorporation of tetrahedrally coordinated aluminum in cpx coexisting with garnet, whereas filled symbols represent that pressures based on applying the temperature to the  $45 \text{ mW/m}^2$  geotherm. The temperatures and pressures of garnet-peridotites (grey diamonds) are from Howarth et al. (2014)

## Discussion

### Pressure–temperature history

The two generations of garnets preserved within the eclogites provide insights into the PT history experienced by the samples. Equilibrium pressure–temperature estimates from the last stage of equilibrium (garnet lamellae—host cpx) place these eclogite xenoliths at moderate pressure–temperature conditions 2.3–3.7 GPa and 855–1095 °C in the upper mantle, scattered around the  $45 \text{ mW/m}^2$  geotherm at the time of entrainment by the kimberlite. However, the combination of reconstructed initial clinopyroxene composition and garnet-coarse composition suggest that this suite of

**Table 6** Results of pressure and temperatures of low-MgO Obnazhennaya eclogites

| Sample         | T <sub>Cpx–Grt</sub> /°C<br>Grt lamella–cpx<br>Krogh (1988) | T <sub>Cpx–Grt</sub> /°C<br>Grt lamella–cpx<br>Krogh (2000) | P <sub>Cpx–Grt</sub> /GPa<br>Grt lamella–cpx<br>Beyer et al. (2015) | T <sub>Cpx–Grt</sub> /°C<br>Grt coarse–recon-<br>struct cpx<br>Krogh (1988) | P <sub>Cpx–Grt</sub> /GPa<br>Grt coarse–reconstruct cpx<br>Beyer et al. (2015) |
|----------------|---|---|---|---|--|
| This study     |   |   |   |   |  |
| 7-303          | 975   | 909   | 3.7   | 1198  | 2.2  |
| 7-350          | 855   | 840   | 2.3   | 1056  | 0.8  |
| 7-352          | 947   | 851   | 2.5   |   |  |
| 7-354          | 984   | 927   | 3.6   | 1114  | 2.0  |
| 7-379          | 950   | 870   | 3.2   | 1288  | 2.3  |
| 7-394          | 1095  | 1041  | 3.0   | 1003  | 1.8  |
| 74-136         | 949   | 888   | 3.3   | 950   | 1.9  |
| 74-639         | 957   | 887   | 3.5   | 927   | 2.5  |
| Previous study | Grt–cpx   |   | Grt–cpx   |   |  |
| O-501          | 1083  | 1038  | 3.1   |   |  |
| O-927          | 933   | 884   | 2.4   |   |  |
| O-82-91        | 1093  | 1045  | 3.2   |   |  |

xenoliths originally formed as higher temperatures eclogites 927–1288 °C, but at lower pressures (1.8–2.5 GPa; Fig. 6).

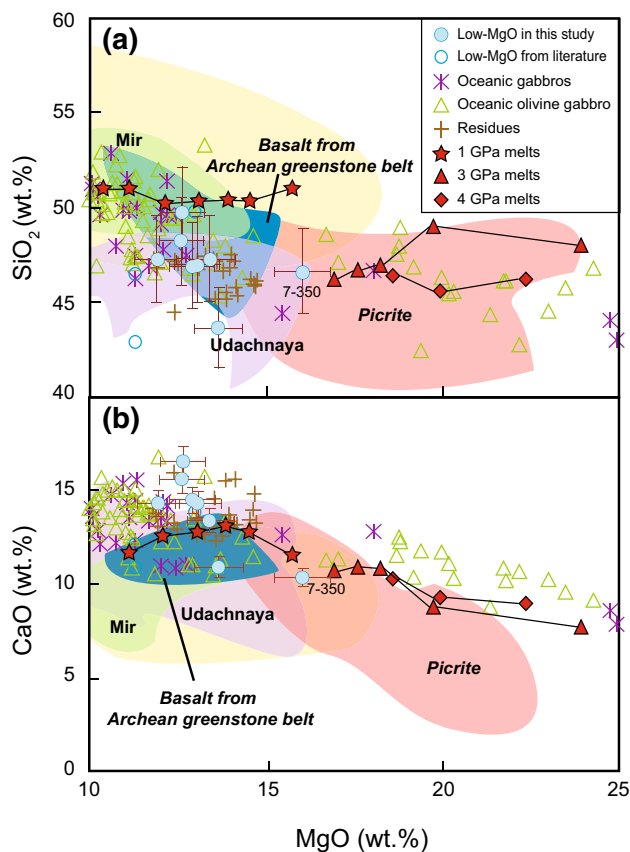
Garnet exsolution from clinopyroxene were previously described from Obnazhennaya by Jerde et al. (1993) from three corundum- and kyanite-bearing eclogites (only major elements) and in a conference abstract by Dobosi et al. (2007) from large clinopyroxene (3 cm; both major and trace elements). Similar exsolution features have also been found in xenolithic eclogites within kimberlite from other cratons (e.g., Beck 1907; Jerde et al. 1993), ultramafic rocks (Zhang and Liou 2003) and eclogites within serpentinites occurring in metamorphic complexes (e.g., Reiche and Bausch 1985). It was suggested in previous studies that any process involving a pressure increase and/or temperature decrease may produce garnet exsolution lamellae in host clinopyroxenes (e.g., Zhang and Liou 2003; Harte and Gurney 1975), because the solubility of Ca-(and Mg-) Tschermarks (CaTs and MgTs) is exceeded. The solubility of CaTs in clinopyroxene increases with temperature and decreases with pressure (e.g., Gasparik 1984).

Reconstructed clinopyroxenes in our samples have CaTs ranging from 12 to 16 mol% for most samples except for two with 4 and 6 mol%. This is much higher than in the measured clinopyroxenes (Table 7). Such precursor phases of garnet + clinopyroxene intergrowth were considered to be primary high-aluminum pyroxene formed under higher temperature conditions (Zhang and Liou 2003). The calculated temperatures from the reconstructed clinopyroxene and coarse garnet are indeed higher and support a higher temperature origin. The granular garnets are precursor phases and did not participate in the reaction during exsolution. This is supported by the different REE patterns of coarse

and lamellar garnets—the negative middle to heavy REE slopes of the lamellar garnets demonstrate equilibration with the host clinopyroxene but not with the granular garnets. Clinopyroxenes from experiments with a kyanite eclogite at 1100 °C and pressures between 2 and 3.5 GPa contain 15.7–19.3 wt% Al<sub>2</sub>O<sub>3</sub> both as jadeite and CaTs components (Morishita et al. 2004), higher than our reconstructed precursor clinopyroxenes (6.7–15.8 wt%). The experimental results at similar pressures as the Obnazhennaya eclogites are a guide for the possible PT conditions for the formation of the precursor high Al-pyroxenes. The experimental clinopyroxenes are higher in Al<sub>2</sub>O<sub>3</sub> because of the higher Na-content of the kyanite eclogite. The calculated PT conditions for the reconstructed cpx–granular grt pairs lie in the range of 2 GPa and 1000–1200 °C, far elevated above any continental conductive geothermal gradient, or modern subduction zone gradients. We explore, below, the implications of this PT history for the origin and devolution of the eclogites. First, though, we use their chemical characteristics to infer their origins.

### Evaluation of the effects of metasomatism

Before the chemical composition of eclogites can be used to infer their origin, one must first evaluate whether they have experienced any metasomatism, which may affect xenolithic eclogites both during their residence in the cratonic lithosphere and/or during entrainment in the host kimberlite (Barth et al. 2001; Huang et al. 2012, 2014; Shu et al. 2018). Metasomatism can alter the chemical and isotopic compositions of a rock at different scales, cryptically to modally, obscuring the primary composition to varying



**Fig. 7** Plot of MgO vs. SiO<sub>2</sub> (a) and CaO (b) for reconstructed Obnazhennaya eclogite whole-rock compositions. Blue circles: eclogites studied here. Blue open circles: Obnazhennaya Group B eclogites from Taylor et al. (2003). Pluses: Residues from eclogite melting experiments (Carroll and Wyllie 1990; Winther and Newton 1991; Sen and Dunn 1994; Rapp and Watson 1995), assuming equal proportions of residual garnet and cpx. Purple stars: oceanic gabbros (Casey 1997; Perk et al. 2007; Godard et al. 2009). Triangles: oceanic olivine gabbro (Casey 1997; Perk et al. 2007; Godard et al. 2009). Yellow field outlines data for Archean cratonic basalts and pink field outlines Archean picrite; dark blue field is for basalt from Archean greenstone belt (data from PetDB and geochemical rock database (Lehnert et al. 2000)). Green and purple fields are eclogites from Udachnaya (Agashev et al. 2018; Sobolev et al. 1994) and Mir (Beard et al. 1996; Snyder et al. 1997) kimberlites, respectively. Experimental pyrolite melts (red symbols) at 1, 3 and 4 GPa, respectively (Baker and Stolper 1994; Walter 1998)

degrees (Harte 1987). Therefore, it is important to evaluate the effects of metasomatic processes on the primary compositions of these xenoliths.

Several lines of evidence suggest that most eclogites in this study have not been strongly metasomatized: (1) phlogopite and amphibole, which are typically added to the cratonic mantle during modal metasomatism (e.g., eclogites from Kimberley, Kaapvaal craton; Jacob et al. 2009; Shu et al. 2018), are not observed in eclogites in this study. (2) Interaction with low-volume alkaline melts can lead to TiO<sub>2</sub> and high field-strength element (HFSE) enrichment in garnet

(e.g., Pearce 2008). Garnets from the Obnazhennaya eclogites do not show evidence for such melt-rock interaction. (3) Melt metasomatism is expected to increase the concentrations of most incompatible elements, like LREE, contents of the original eclogite (Ireland et al. 1994) and results in (La/Yb)<sub>N</sub> > 1. While many of the samples analyzed here show LREE-enrichment ((La/Yb)<sub>N</sub> > 1), which may be due to late-stage metasomatism, the enrichment appears to be confined to the lightest REE and does not extend into the MREE. Interpolating from the MREE–HREE, one can see that all samples in this study were originally LREE-depleted. The exception, sample O-927 from the previous study of Taylor et al. (2003), has a (La/Yb)<sub>N</sub> ratio of 1.8 (Fig. 4), which may reflect metasomatic overprinting. This sample has lower FeO and higher Mg#, which could be due to small degrees of metasomatic overprint after partial melting or metasomatism. However, partial melting will also result in strong depletion of SiO<sub>2</sub>, but the sample has the highest SiO<sub>2</sub> of all of the low-MgO Obnazhennaya eclogites. Combined with its enriched LREE pattern, we, therefore, favor metasomatism as an explanation for the low FeO and high SiO<sub>2</sub> content and Mg# for this sample. Hence, sample O-927 will be omitted from further discussion.

### Origin of the low-MgO eclogites

Low-MgO eclogite xenoliths have been reported from South Africa (e.g., Gréau et al. 2011; Huang et al. 2012, 2014; Shu et al. 2013), Finland (Peltonen et al. 2002; Smart et al. 2017a), West Africa (e.g., Hills and Haggerty 1989; Fung and Haggerty 1995; Barth et al. 2001; Aulbach and Arndt 2019), the Slave craton (e.g., Aulbach et al. 2007; Smart et al. 2012, 2014, 2017b) and Udachnaya (Agashev et al. 2018) and Mir (Snyder et al. 1997) in Siberia. Here, we evaluate models for the origin for the Obnazhennaya low-MgO eclogites: (1) as cumulates from mafic melts at higher pressures (“mantle” hypothesis) (Caporuscio and Smyth 1990; and references therein); (2) as residues from partial melting of subducted Archean basalts (Ireland et al. 1994; Barth et al. 2001); (3) as founded lower continental crust (Hills and Haggerty 1989; Pearson et al. 1991); (4) as founded or subducted low-pressure cumulates that may or may not have experienced partial melting (Barth et al. 2002; and references therein). We evaluate each of these hypotheses, in turn.

### High-pressure cumulates from mantle melts?

According to the “mantle” hypothesis, the eclogites should have oxygen isotopic compositions similar to those of primary mantle melts or their cumulates that formed at pressures exceeding 3 GPa (e.g., Hatton 1978). Most low-MgO eclogites from Obnazhennaya have equilibration pressures



**Table 7** Major element composition (wt%) of reconstruct clinopyroxene from the Obnazhennaya eclogites

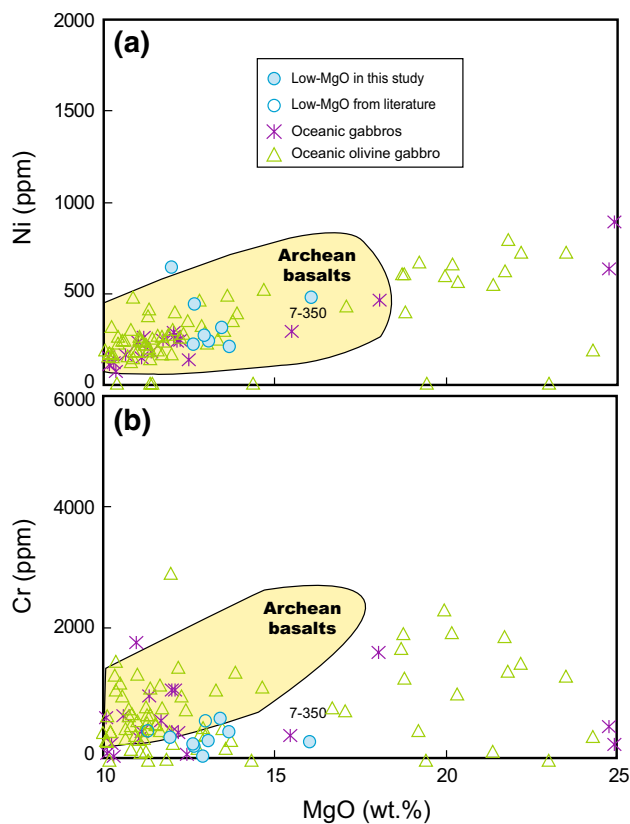
|                 | Vol  | SiO <sub>2</sub> | TiO <sub>2</sub> | Al <sub>2</sub> O <sub>3</sub> | Cr <sub>2</sub> O <sub>3</sub> | FeO   | MnO  | MgO   | CaO   | Na <sub>2</sub> O | K <sub>2</sub> O | NiO  | Total  | Jd   | Mg#  | CaTs |
|-----------------|------|------------------|------------------|--------------------------------|--------------------------------|-------|------|-------|-------|-------------------|------------------|------|--------|------|------|------|
| 7-303           |      |                  |                  |                                |                                |       |      |       |       |                   |                  |      |        |      |      |      |
| Host cpx        | 0.75 | 53.76            | 0.11             | 6.32                           | 0.05                           | 2.96  | 0.01 | 13.16 | 19.62 | 2.93              | 0.01             | 0.05 | 98.99  | 0.21 | 0.89 | 0.00 |
| Grt lamella     | 0.25 | 40.00            | 0.01             | 22.21                          | 0.06                           | 12.57 | 0.34 | 10.30 | 13.37 | 0.01              | 0.00             | 0.01 | 98.87  |      |      |      |
| Reconstruct cpx |      | 50.32            | 0.09             | 10.29                          | 0.05                           | 5.36  | 0.09 | 12.45 | 18.06 | 2.20              | 0.01             | 0.04 | 98.96  | 0.18 | 0.81 | 0.13 |
| 7-350           |      |                  |                  |                                |                                |       |      |       |       |                   |                  |      |        |      |      |      |
| Host cpx        | 0.73 | 53.30            | 0.14             | 6.17                           | 0.04                           | 2.61  | 0.04 | 13.44 | 17.62 | 3.25              | 0.01             | 0.12 | 96.72  | 0.25 | 0.90 | 0.01 |
| Grt lamella     | 0.27 | 40.45            | 0.03             | 22.81                          | 0.05                           | 11.44 | 0.22 | 17.91 | 3.53  | 0.01              | 0.00             | 0.02 | 96.48  |      |      |      |
| Reconstruct cpx |      | 49.83            | 0.11             | 10.66                          | 0.04                           | 4.99  | 0.09 | 14.65 | 13.82 | 2.37              | 0.01             | 0.09 | 96.66  | 0.24 | 0.84 | 0.14 |
| 7-354           |      |                  |                  |                                |                                |       |      |       |       |                   |                  |      |        |      |      |      |
| Host cpx        | 0.80 | 53.14            | 0.08             | 6.98                           | 0.05                           | 2.93  | 0.02 | 12.73 | 19.14 | 3.02              | 0.01             | 0.07 | 98.18  | 0.22 | 0.89 | 0.04 |
| Grt lamella     | 0.20 | 39.76            | 0.04             | 22.08                          | 0.03                           | 11.97 | 0.28 | 8.96  | 15.45 | 0.01              | 0.01             | 0.01 | 98.60  |      |      |      |
| Reconstruct cpx |      | 50.46            | 0.07             | 10.00                          | 0.05                           | 4.74  | 0.07 | 11.98 | 18.40 | 2.42              | 0.01             | 0.06 | 98.26  | 0.19 | 0.82 | 0.12 |
| 7-379           |      |                  |                  |                                |                                |       |      |       |       |                   |                  |      |        |      |      |      |
| Host cpx        | 0.70 | 53.30            | 0.18             | 6.46                           | 0.08                           | 3.03  | 0.02 | 13.02 | 19.12 | 3.03              | 0.00             | 0.06 | 98.30  | 0.22 | 0.89 | 0.02 |
| Grt lamella     | 0.30 | 39.92            | 0.02             | 22.45                          | 0.07                           | 12.99 | 0.35 | 11.94 | 10.64 | 0.02              | 0.01             | 0.01 | 98.41  |      |      |      |
| Reconstruct cpx |      | 49.29            | 0.13             | 11.26                          | 0.08                           | 6.02  | 0.12 | 12.70 | 16.57 | 2.12              | 0.01             | 0.05 | 98.34  | 0.19 | 0.79 | 0.16 |
| 7-394           |      |                  |                  |                                |                                |       |      |       |       |                   |                  |      |        |      |      |      |
| Host cpx        | 0.95 | 51.53            | 0.05             | 15.39                          | 0.05                           | 1.12  | 0.00 | 8.90  | 17.66 | 4.44              | 0.01             | 0.13 | 99.28  | 0.31 | 0.93 | 0.23 |
| Grt lamella     | 0.05 | 41.81            | 0.03             | 23.63                          | 0.07                           | 5.49  | 0.09 | 9.37  | 21.13 | 0.02              | 0.00             | 0.00 | 101.64 |      |      |      |
| Reconstruct cpx |      | 51.09            | 0.05             | 15.77                          | 0.05                           | 1.32  | 0.01 | 8.92  | 17.82 | 4.24              | 0.01             | 0.12 | 99.39  | 0.30 | 0.92 | 0.20 |
| 74-136          |      |                  |                  |                                |                                |       |      |       |       |                   |                  |      |        |      |      |      |
| Host cpx        | 0.92 | 53.17            | 0.09             | 6.20                           | 0.04                           | 3.54  | 0.02 | 13.02 | 19.70 | 2.67              | 0.00             | 0.04 | 98.49  | 0.20 | 0.87 | 0.03 |
| Grt lamella     | 0.08 | 39.80            | 0.03             | 22.32                          | 0.04                           | 14.22 | 0.32 | 9.18  | 13.61 | 0.01              | 0.00             | 0.01 | 99.53  |      |      |      |
| Reconstruct cpx |      | 52.11            | 0.08             | 7.47                           | 0.04                           | 4.38  | 0.04 | 12.71 | 19.22 | 2.46              | 0.00             | 0.04 | 98.57  | 0.19 | 0.84 | 0.06 |
| 74-639          |      |                  |                  |                                |                                |       |      |       |       |                   |                  |      |        |      |      |      |
| Host cpx        | 0.97 | 53.76            | 0.04             | 6.15                           | 0.01                           | 3.21  | 0.03 | 13.39 | 19.85 | 2.67              | 0.01             | 0.05 | 99.17  | 0.20 | 0.88 | 0.00 |
| Grt lamella     | 0.03 | 39.68            | 0.02             | 22.22                          | 0.01                           | 13.28 | 0.37 | 10.30 | 12.77 | 0.02              | 0.00             | 0.00 | 98.66  |      |      |      |
| Reconstruct cpx |      | 53.31            | 0.04             | 6.66                           | 0.01                           | 3.53  | 0.04 | 13.29 | 19.63 | 2.59              | 0.01             | 0.05 | 99.16  | 0.19 | 0.87 | 0.04 |

between 3.0 and 3.7 GPa (Fig. 6) and mantle-like oxygen isotopic compositions (Fig. 5), which seems to be consistent with the hypothesis. However, the major element compositions of the Obnazhennaya low-MgO eclogites differ significantly from primary and high-pressure mantle melts (Fig. 7). Picritic melts generated between 3 and 4 GPa have higher MgO and lower CaO contents than the Obnazhennaya low-MgO eclogites. Picritic melts derived from relative lower pressures, i.e., at 1 GPa, have moderate MgO content (10–16 wt%) that overlap well with the low-MgO eclogites. However, these melts show constant SiO<sub>2</sub> (~50 wt%) and CaO (11–13 wt%) contents at different MgO compositions, while the low-MgO Obnazhennaya eclogites display negative correlations between CaO vs. MgO and SiO<sub>2</sub> vs. MgO contents. Such negative correlations in the low-MgO eclogites argue against their origin as crystallized low-pressure picritic melts. In addition, five out of eight Obnazhennaya samples in this study have Ni contents lower than primary mantle melts (208–311 ppm vs. ≥320 ppm, Frey et al. 1978), though sulfide is not considered in the reconstruction of the

bulk rock, so the Ni contents discussed here are minimum values. Moreover, cumulates of garnet and cpx from primary mantle melts at pressures greater than 3 GPa (Walter 1998) cannot produce the positive Eu and Sr anomalies in our samples (Fig. 4a, b) and non-mantle-like oxygen isotopic compositions of sample 7-303 (4.73‰) in this study and samples in the previous study (7.1–7.8‰). Thus, the protoliths of the low-MgO Obnazhennaya eclogites are unlikely to be mantle melts crystallized at high pressures, nor high pressure cumulates.

#### Residues of altered and metamorphosed Archean basalt?

Most Obnazhennaya low-MgO eclogites have higher CaO, and Al<sub>2</sub>O<sub>3</sub> but lower FeO and SiO<sub>2</sub> contents than Archean cratonic basalts (yellow field in Fig. 7), but partially overlap with basalts from Archean greenstone belts and also modern oceanic gabbros, especially for SiO<sub>2</sub> and MgO (Figs. 7, 8; Appendix 4) (Winther and Newton 1991; Sen and Dunn 1994; Rapp and Watson 1995). Similar findings were



**Fig. 8** Whole-rock MgO (wt%) vs. Ni (a) and Cr (b). Symbols and data sources as in Fig. 7

reported by previous studies, e.g., the low-MgO eclogites from the Koidu kimberlite (Barth et al. 2001) have lower SiO<sub>2</sub> and higher CaO contents than Archean basalts, but overlap with the residues of partial melting experiments of Archean basalt. Therefore, Barth et al. (2001) interpreted the Koidu low-MgO eclogites as residues of silica-rich melt extraction from Archean basalts. Similarly, Taylor et al. (2003) interpreted Obnazhennaya Group B eclogites as residues after the extraction of a TTG melt. However, low silica contents may result not only from partial melting, but also from seafloor alteration (Ridley et al. 1994; Staudigel et al. 1995; Hart et al. 1999). Indeed, the elevated  $\delta^{18}\text{O}$  (7.1–7.8‰) of low-MgO Obnazhennaya eclogites from the previous study were interpreted to reflect low-T seafloor alteration. Hence, it is uncertain whether the samples depleted in SiO<sub>2</sub> in the previous study are the residues of Archean basalt partial melting. It seems that the protolith of some eclogites with the depleted SiO<sub>2</sub> can be residues from partial melting of Archean basalts. However, compared to the experimental melting residues of Archean basalt, most Obnazhennaya low-MgO eclogites from this study show higher MgO, CaO, and Al<sub>2</sub>O<sub>3</sub> but lower FeO contents, (Figs. 7, 8; Appendix 4) (Winther and Newton 1991; Sen

and Dunn 1994; Rapp and Watson 1995). Moreover, the positive slopes and high contents of the HREE in Obnazhennaya low-MgO eclogites (Fig. 4a) require the presence of large volumes of garnet in an open system—either during partial melting or during crystal accumulation. Therefore, the low-MgO Obnazhennaya eclogites are unlikely to be the residues of melting of Archean basalts.

#### Foundered lower continental crust?

Mafic granulites formed by crystallization of underplated basalts are typically encountered in the lower continental crust or near the crust–mantle boundary (Herzberg et al. 1983; O’Reilly and Griffin 2013). If such mafic rocks are tectonically buried, they could convert to eclogites, and founder to greater depths (Hills and Haggerty 1989; Pearson et al. 1991). Although P–T estimates indicate that Siberian eclogites and granulites do not exist at the same depths within the lithosphere (2.3–3.7 GPa vs. 0.8–1.2 GPa, respectively; Koreshkova et al. 2011; Moyen et al. 2017), it is conceivable that the Obnazhennaya low-MgO eclogites represent former lower continental crust that foundered to greater depths. However, this hypothesis cannot explain the mantle-like  $\delta^{18}\text{O}$  values observed in most Obnazhennaya low-MgO eclogites in this study, because the Oxygen-isotope compositions of most mafic lower crustal granulites are elevated well above the mantle value due to the assimilation of pre-existing lower crust (Fowler and Harmon 1990; Kempton and Harmon 1992).

#### Low-pressure cumulate protoliths?

As discussed above, the low-MgO Obnazhennaya eclogites do not have the compositions of Archean basalts (Figs. 7, 8), suggesting that they do not represent liquid compositions. By contrast, they have relatively high MgO contents and significant positive Eu anomalies in both calculated bulk rocks and minerals (Figs. 3, 4), which point toward a low-pressure cumulative origin, where plagioclase was a stable phase. Reconstructed bulk rocks show an anti-correlation between Eu/Eu\* and  $\Sigma\text{HREEs}$  (Appendix 5), which is consistent with plagioclase accumulation (Aulbach and Arndt 2019). Normative plagioclase (43%), diopside (27%) and olivine (23%) are similar to modes of (olivine) gabbroic cumulate rocks from modern MORB (Appendix 1). However, the reconstructed bulk rock compositions of the low-MgO eclogites have lower SiO<sub>2</sub> and higher CaO and MgO contents compared to oceanic gabbros. Such offset can be attributed to differences in the composition of their parental melts, or to later melt extraction (Fig. 7). When compared with pyrolite melts from different pressures (1 GPa, 3 GPa and 4 GPa; Walter 1998), most low-MgO eclogites (except 7-350 and two eclogites in the previous study) are depleted

in  $\text{SiO}_2$  and enriched in CaO and also  $\text{Al}_2\text{O}_3$  (Fig. 7), which could also reflect partial melting and extraction of silicic melts. The low-MgO eclogites are also depleted in LREE with  $(\text{Nd}/\text{Yb})_N$  ratios  $< 1$ , again consistent with partial melt extraction after protolith formation (Aulbach and Arndt 2019).

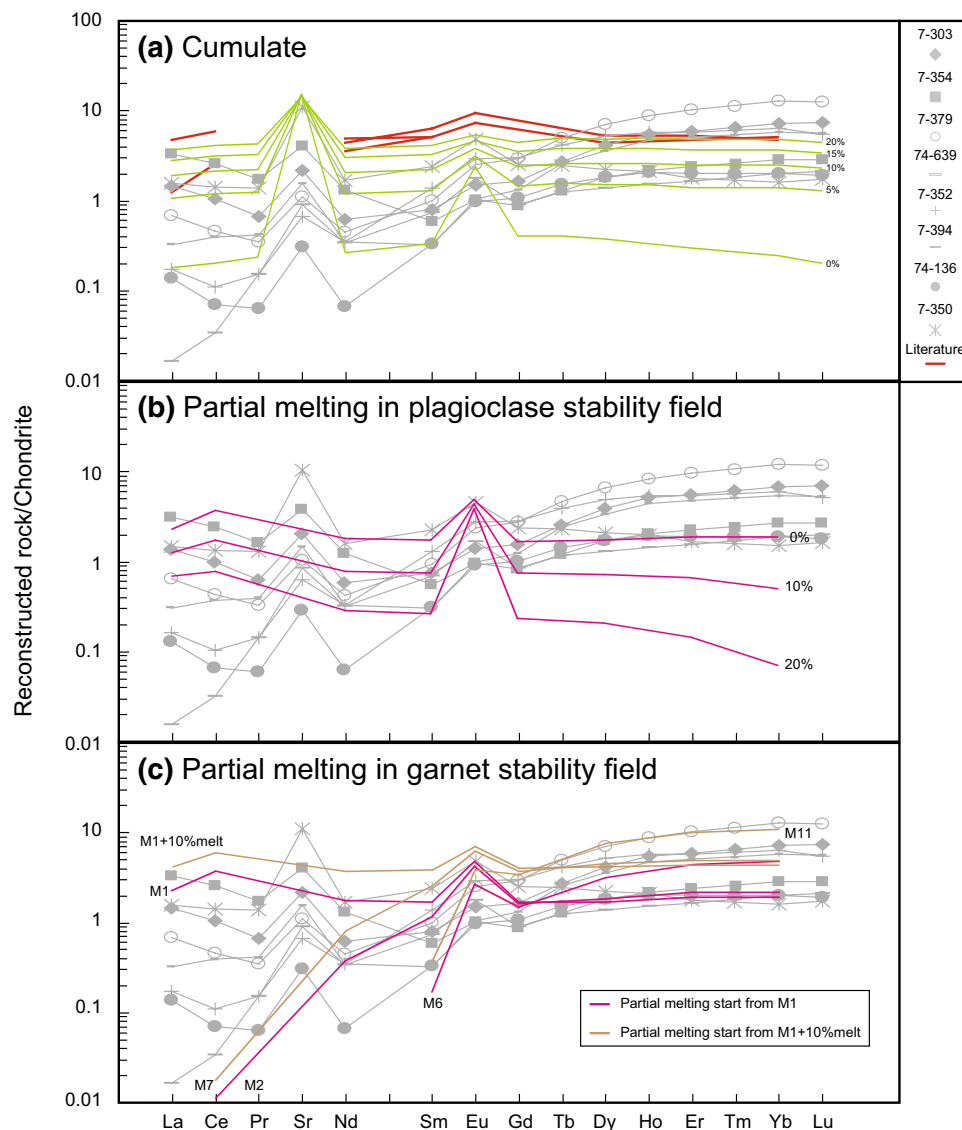
Therefore, we modeled the trace elements to test the hypothesis that low-MgO eclogites originated as melt residues of low-pressure cumulates. Although the age of Obnazhennaya eclogites is not known, two point Sm–Nd isochrons of garnet–clinopyroxene mineral pairs yield ages of 1.2 Ga, which suggest they are at least Mesoproterozoic in age (Taylor et al. 2003). The ambient mantle temperature earlier in Earth history is much hotter than in the modern time, which results in the difference in composition between Archean basalt and modern MORB (Herzberg et al. 2010). Archean basalt is more Mg-rich ( $\text{MgO} > 10\%$ ) than modern MORB (average  $\text{MgO} = 7.6\%$ ; Gale et al. 2013). Therefore, Archean basalts were used as the parental melt from which the low-pressure cumulates precipitate (with or without trapped melt). REE and Sr relative concentrations (Fig. 9) suggest the low-pressure ( $\leq 0.5$  GPa) accumulation of olivine, plagioclase, and clinopyroxene for all of the low-MgO Obnazhennaya samples (cf. Presnall et al. 1978). Model parameters are given in the caption of Fig. 9.

The composition of sample 7-350, with the lowest HREE contents ( $\text{Lu} = 0.6 \times \text{chondrite}$ ), the highest LREE content, and the largest positive Sr and Eu anomalies, is clearly indicative of a low-pressure, plagioclase-rich cumulate origin. The flat light to middle REE pattern suggests that this sample may not have experienced partial melting. Based on our calculations (Fig. 9a), a mixture of 50% plagioclase, 30% olivine, and 20% clinopyroxene cumulate can reproduce the eclogite composition. This is supported by the calculated CIPW norm which shows 48% plagioclase for this sample (Appendix 1). The trace-element composition of minerals for the model cumulate are taken from an olivine gabbro sample from IODP Site U1309 (Godard et al. 2009), which has a similar major element composition as 7-350 in this study. The  $(\text{Nd}/\text{Yb})_N$  ratio of such a cumulate is 0.92, matching the ratio of the sample 7-350 (1.0) and the two eclogites (O-82/91, O-501) in the previous study of Taylor et al. (2003) with  $(\text{Nd}/\text{Yb})_N$  ratios 0.91 and 0.96, respectively (Fig. 9a). These three samples have similar REE patterns as the low-pressure cumulate, but higher overall REE contents. We, therefore, mix variable amounts of Archean basalt with the cumulate melt to reproduce these three eclogite compositions. The compositions are successfully explained if the nonmodal batch melt is allowed to vary from 5 to 20% (Fig. 9a).

For samples with depleted features, e.g., low incompatible element and high and fractionated HREE, we used a fractional melting model of subducted basalt at different

pressures to determine how the partial melting process affects the chemical compositions of the low-MgO Obnazhennaya eclogites and to determine whether any constraints can be placed on the degrees and depths of partial melting. The average composition of modern-day oceanic gabbro (Casey 1997) is used as the starting composition (M1) for our model. The sample was melted in both the garnet and plagioclase stability field. Our calculations show that partial melting will not generate the observed positively sloped patterns for middle to heavy REE if the starting materials do not contain garnet. Instead it will generate negatively sloped HREE patterns and will significantly lower the HREE contents (Fig. 9b). However, repeated/continuous melting in the eclogite stability field with different mineral modal abundances will produce the higher and fractionated HREE pattern similar to our samples (Fig. 9c). We assume that different stages of partial melting affected these eclogites before they were brought back to the surface: Stage 1: 15% melting of eclogite with the composition of M1, mineral modal abundance ratios: grt/cpx = 60:40. The residue composition is M2; Stage 2: use M2 as a starting material, calculate 20% melt loss; grt/cpx ratio in the residue is 65:35. The residue composition is M3; Stage 3: use M3 as a starting material, remove another 20% of melt; grt/cpx ratio will be 70:30. The residue composition is then M4. Stage 4: starting with M4, remove another 20% of melt; grt/cpx ratio will be 75:25. The residue composition is M5. Stage 5: starting with M5, remove another 20% of melt; grt/cpx ratio will be 20:80, and M6 is the residue composition. The melting modal for our calculation has Grt/cpx ratio of 60/40 (details for modeling calculation are present in Appendix 1). The  $(\text{Dy}/\text{Yb})_N$  ratios of the residues, M1, M2, M3, M4, M5 and M6, are 0.91, 0.86, 0.80, 0.75, 0.70, 0.66, respectively (Fig. 9c; Appendix 1). They compare well with the low-MgO eclogites having fractionated HREE patterns (from 0.58 to 0.93). Continuous or repeated melting decreases cpx and increases the garnet modal abundance in the residues. This is consistent with the high modal garnet abundance in the sample suite.

Despite the successful match of the HREE fractionation between our model calculations and the low-MgO eclogites when using the average composition of MORB gabbro as starting material, a number of our samples have similarly fractionated patterns but higher HREE–MREE contents ( $(\text{Lu})_N = 4\text{--}10$  times Chondrite). When adding 10% Archean basaltic melt into the average composition of modern oceanic gabbro, the REE pattern and contents overlap well with these samples (Fig. 9c). Therefore, continuous partial melting of the low-MgO Obnazhennaya eclogites likely occurred in the eclogite stability field, after the cumulates (some with trapped melt) were brought to high pressures. Repeated melting of eclogite in the lithospheric mantle has also been explored by experimental studies (e.g., Rosenthal



**Fig. 9** **a** Low-MgO eclogites interpreted to be low-pressure cumulates are shown in grey. Green lines are a calculated ol-plag-cpx cumulate in equilibrium with an olivine gabbro from IODP Site U1309 (Godard et al. 2009) containing 0–20% Archean basalt. The cumulate consists of 50% plagioclase, 30% olivine, and 20% clinopyroxene, comparable to the CIPW norms of the sample 7–350. Red lines are low-MgO Obnazhennaya eclogites from Taylor et al. (2003). **b** Model calculations for various degrees of batch partial melting in plagioclase stability field. The starting composition is using the composition of gabbro from modern ocean (Casey 1997) and starting minerals are 50% plagioclase, 20% olivine and 30% clinopyroxene comparable to the average CIPW norms of the low-MgO Obnazhen-

naya eclogites. Pink lines are calculated compositions of variable degree partial melting. Numbers give the percentage of partial melting degree. **c** Model calculations for fractional partial melting in the eclogite stability field. Pink and light brown lines are calculated compositions of episode partial melting using M1 and M1 + 10% melt as starting material. M1 is the same gabbro composition as the starting material used in **b**. M1 + 10% melt is the M1 with 10% Archean basalts. M2 and M7 is 15% of melting with the starting composition of M1 and M1 + 10% melt, respectively. M6 and M11 is the residue of the fifth partial melting. Detail calculate description is provided in Appendix 1

et al. 2014), and was proposed as an important mechanism to cause refertilization and compositional heterogeneity of the upper mantle.

In summary, based on the major and trace-element compositions of low-MgO Obnazhennaya eclogites in this study,

their protoliths can be explained as plagioclase-dominated cumulates formed at low pressure that repeatedly partially melted to variable degrees after eclogite metamorphism.

## Oxygen isotopic features

The  $\delta^{18}\text{O}$  of modern ocean floor lithologies are significantly more variable than that of the mantle ( $5.5 \pm 0.4\text{‰}$ ; Matthey et al. 1994). Heavier  $\delta^{18}\text{O}$  values than that of the mantle are observed in sea floor basalts altered at relatively low temperature, while lighter  $\delta^{18}\text{O}$  values reflect high-temperature seafloor alteration in the lower parts of the oceanic crust (e.g., Gregory and Taylor 1981; McCulloch et al. 1981). Although Huang et al. (2014) attributed variable  $\delta^{18}\text{O}$  in eclogites from Roberts Victor, South Africa to melt metasomatism, we observe no correlation between  $\delta^{18}\text{O}$  of the low-MgO eclogites with incompatible trace elements, ruling out the metasomatism scenario. Therefore, we only consider seawater–rock exchange to explain the eclogite  $\delta^{18}\text{O}$  values.

The  $\delta^{18}\text{O}$  values of garnets from two Obnazhennaya eclogites in the previous study of Taylor et al. (2003) (7.1–7.8‰) are significantly heavier than the mantle value, which was attributed to low-temperature seafloor alteration. In addition, the garnet from sample 7-303 in our sample suite with lower  $\delta^{18}\text{O}$  values (4.73‰) also suggests a high-temperature seafloor alteration. Such alteration can also cause depletions of CaO and of  $\text{SiO}_2$ , which is consistent with the lower CaO and  $\text{SiO}_2$  content of these three samples relative to samples in this study.

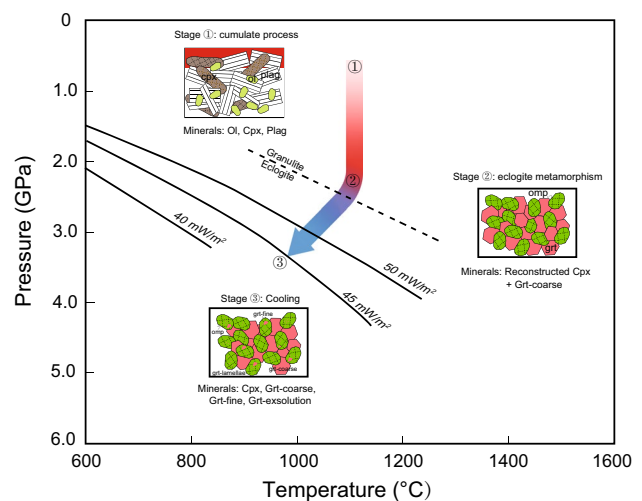
It remains unclear if the protoliths of most Obnazhennaya low-MgO eclogites with mantle-like  $\delta^{18}\text{O}$  were hydrothermally altered. Either the protoliths did not interact with seawater, or such interaction coincidentally produced normal igneous values. In the lowermost oceanic crust, hydrothermal alteration is localized along fractures and veins (Gregory and Taylor 1981; Dick et al. 2000). Thus, it is possible that most of the protoliths of the eclogites in this study did not experience hydrothermal alteration. Alternatively, seawater-derived hydrothermal fluids have variable  $\delta^{18}\text{O}$ , and a gabbro can be hydrothermally altered over a significant range of temperatures to produce a  $\delta^{18}\text{O}$  value that is coincidentally consistent to the mantle value (Barth et al. 2002). Moreover, if the cumulate formed deep within in a thicker oceanic crust, hydrothermal processes may not reach these depths, even at fractures generated by bending plates.

## Geological implications

If the Obnazhennaya eclogites are the products of low-pressure crystal accumulation and subduction of thermally equilibrated ancient oceanic crust, then a transient heating event must be postulated to produce the high temperatures at 2 GPa. This may occur during a rifting event at the edge of the craton, generating the high-temperature eclogites, followed by cooling and garnet exsolution.

Alternatively, the history of the low-MgO Obnazhennaya eclogite xenoliths may be as follows (Fig. 10): (1) accumulation of ol, cpx and plag at somewhat elevated pressure, i.e., ~0.5 GPa deep within a thick Archean oceanic crust, where high temperatures are preserved for a longer time; (2) subduction of the hot cumulate and metamorphism at high temperature to generates first a high-T mafic granulites and the high-T eclogites with Al-rich cpx; and (3) secular cooling to ambient mantle conditions or pressure increasing causing exsolution of garnets.

Eclogite xenoliths from several kimberlites (e.g., Udachnaya, Mir) located within the Siberian craton have been previously studied. Compared with low-MgO (Group B/C) eclogite xenoliths from Udachnaya and Mir kimberlite, samples from the Obnazhennaya kimberlite have comparable major, trace elements, (including positive Eu anomalies), and oxygen isotopic characteristics, which indicate their oceanic crust origin. However, most garnets from the Obnazhennaya eclogites show mantle-like oxygen isotopic ratios relative to those from the Udachnaya (4.9–7.0‰; Jacob et al. 1994; Shatsky et al. 2016) and Mir kimberlites (3.1–7.2‰; Beard et al. 1996). Additionally, Obnazhennaya low-MgO eclogite xenoliths have systematically lower LREE and incompatible elements than those from Udachnaya and Mir. These differences may reflect that the Udachnaya and Mir kimberlite pipes lie near the center of the Siberian craton, and thus could have tapped a source from deeper in the mantle keel. By contrast, Obnazhennaya kimberlite pipe is located on the edge of craton; therefore, their eclogite xenoliths remained at relative lower temperature within the mantle.



**Fig. 10** Schematic representation of the possible formation processes of the Obnazhennaya low-MgO eclogites suite



## Summary and conclusions

The low-MgO eclogites from the Obnazhennaya kimberlite pipe in Yakutia offer the rare occasion to deduce their protoliths from their geochemical properties but also to reconstruct the path that led to their transformation to eclogites using the garnet exsolution lamellae in their clinopyroxenes.

- (1) Petrographic and geochemical features of the Obnazhennaya low-MgO eclogite xenoliths in our study revealed a somewhat different petrogenesis than suggested in the previous study of Taylor et al. (2003). Our findings are consistent with an origin as subducted or foundered oceanic crust, as proposed by Taylor et al. (2003), but we suggest that the protoliths of low-MgO eclogite xenoliths were oceanic olivine gabbro cumulates with or without trapped melt that later partially melted, rather than basaltic oceanic crust.
- (2) The low-MgO eclogites from Obnazhennaya contain two types of garnet with different compositions: granular coarse garnet and garnet exsolution (lamellae and fine-grained) in clinopyroxene. Pressures and temperatures calculated from the last stage of equilibrium phases (garnet lamellae and their host clinopyroxene) suggest that the eclogites were derived from between 855 and 1095 °C and 2.3–3.7 GPa. By contrast, the coarse garnet and reconstructed clinopyroxene give an average about 2 GPa and a range of temperatures between 1000 and 1200 °C. Thus, the eclogites were brought to greater depths as they cooled.
- (3) Ratios of  $(\text{Nd}/\text{Yb})_N < 1$  and positively sloped HREE ( $(\text{Dy}/\text{Yb})_N < 1$ ) in most eclogites require repeated partial melting in the eclogite stability field.

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