

The origin of coarse garnet peridotites in cratonic lithosphere: new data on xenoliths from the Udachnaya kimberlite, central Siberia

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Abstract We report new textural and chemical data for 10 garnet peridotite xenoliths from the Udachnaya kimberlite and examine them together with recent data on another 21 xenoliths from the 80–220 km depth range. The samples are very fresh (LOI near zero), modally homogeneous and large (>100 g). Some coarse-grained peridotites show incipient stages of deformation with <10 % neoblasts at grain boundaries of coarse olivine. Such microstructures can only be recognized in very fresh rocks, because fine-grained interstitial olivine is strongly affected by alteration, and may have been overlooked in previous studies of altered peridotite xenoliths in the Siberian and other cratons. Some of the garnet peridotites are similar in composition to low-opx Udachnaya spinel harzburgites (previously interpreted as pristine melt extraction residues), but the majority show post-melting enrichments in Fe and Ti. The least metasomatized coarse peridotites were formed by 30–38 % of polybaric fractional melting

between 7 and 4 GPa and $\leq 1\text{--}3$ GPa. Our data together with experimental results suggest that garnet in these rocks, as well as in some other cratonic peridotites elsewhere, may be a residual mineral, which has survived partial melting together with olivine and opx. Many coarse and all deformed garnet peridotites from Udachnaya underwent modal metasomatism through interaction of the melting residues with Fe-, Al-, Si-, Ti-, REE-rich melts, which precipitated cpx, less commonly additional garnet. The xenoliths define a complex geotherm probably affected by thermal perturbations shortly before the intrusion of the host kimberlite magmas. The deformation in the lower lithosphere may be linked to metasomatism.

Keywords Garnet peridotite · Mantle xenolith · Craton · Melting residue · Metasomatism · Deformation

Introduction

Garnet peridotites are by far the most common rock type in cratonic mantle, particularly in the depth range 80–250 km (Boyd 1989; Lee et al. 2011; Pearson et al. 2003). Understanding their origin is essential to get a better insight into the origin of cratonic lithosphere. Garnet peridotites are usually considered to be residues of high degrees of melt extraction that experienced a range of post-melting chemical and modal enrichments commonly referred to as metasomatism and/or “re-fertilization” (Carlson et al. 2005; Ionov et al. 2010; Pearson et al. 1995; Shimizu et al. 1997; Simon et al. 2003), in many cases accompanied by deformation (Boyd and Mertzman 1987). This intrinsic complexity of the origin of garnet peridotites makes it difficult, on the one hand, to define their melting conditions (because of variable metasomatic imprints) and, on the

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other hand, to trace a sequence of various enrichment events, which may have taken place in several stages. As a result, a large number of processes, tectonic settings and evolution paths have been proposed for the formation of garnet-facies cratonic mantle involving various residual peridotites and their interaction with intra-plate alkali-rich silicate or carbonatitic melts (Rudnick et al. 1993; Simon et al. 2003) or subduction-related melts/fluids (Kelemen et al. 1998; Pearson and Wittig 2008).

Tracing the melt extraction history of garnet-facies peridotites may be easier for xenolith suites that contain both garnet and spinel peridotites because the latter are usually less metasomatized and considered as being closer in modal and chemical composition to initial melting residues (Bernstein et al. 1998, 2007; Boyd 1989; Wittig et al. 2008). Some major insights into melting conditions and tectonic settings during the formation of cratonic mantle have been obtained through studies of spinel peridotites. A very promising approach to shed more light on the origin of garnet peridotites is to examine a large suite of cratonic peridotites including both garnet- and spinel-facies rocks suitable to obtain high-quality petro-geochemical data.

In this paper, we explore the origin of a unique suite of large, fresh and homogeneous garnet peridotite xenoliths from the Udachnaya kimberlite in the central Siberian craton (Agashev et al. 2013; Ashchepkov et al. 2010; Boyd et al. 1997; Ionov et al. 2010) in relation to a recent study that addressed the formation of spinel harzburgites from the same locality (Doucet et al. 2012). We report new modal, major and trace element data for 10 xenoliths and examine them together with recently published data on another 21 garnet peridotites (Ionov et al. 2010), which collectively represent an almost complete lithospheric mantle section beneath Udachnaya. We focus on the origin of “coarse” peridotites that experienced no or very little deformation. Our main objectives are to better constrain the formation conditions of initial “protoliths” of the garnet peridotites by partial melting (depths, melting degrees, tectonic settings) and the origin of garnet (residual vs. metasomatic), which have not been addressed in detail by earlier work on Udachnaya (Boyd et al. 1997; Ionov et al. 2010). We also examine the nature and sequence of processes that transformed the melting residues to the garnet-facies rocks sampled by kimberlite magmatism as well as links between their modal and chemical composition, microstructures and locations in the lithospheric profile.

Geological setting, sample selection and preparation

Xenoliths in this study are from the Udachnaya-East kimberlite (66°26'N, 112°19'E) erupted 360 Myr ago (Kinny et al. 1997) through the PR₂-AR₁ Daldyn block of the

central Siberian craton (Rosen et al. 1994) (Fig. 1). Early studies of mantle materials at Udachnaya (Boyd et al. 1997; Spetsius and Serenko 1990; Zinchuk et al. 1993) were based on samples from shallow levels in the open-pit diamond mine, which are strongly serpentinized (LOI 3–15 % for peridotite xenoliths). Samples reported here were collected in 2003–2009 in the 420–640 m depth range near the centre of the pipe. They are hosted by remarkably well-preserved type-I kimberlites, which contain fresh, unserpentinized olivine and groundmass (Kamenetsky et al. 2008), and are basically unaffected by syn- and post-magmatic alteration (Kamenetsky et al. 2009a, b, 2012).

Large, homogeneous and fresh coarse-grained garnet peridotites, which contain unaltered olivine, garnet and pyroxenes (Fig. 2a), were selected among several hundred xenoliths. They are 15–30 cm in size and have ellipsoidal shapes. We rejected heterogeneous xenoliths, for example, those with cm-scale garnet-rich segregations (Fig. 2b), and rare rock types, like wehrlites, which may be of magmatic origin. The new samples are coarse or only moderately deformed; sheared peridotites examined in this study are from previous work (Table 1). A sufficiently large amount of fresh material (≥ 100 g taking into account grain size) was taken from xenolith cores to provide representative whole-rock samples (Boyd 1989) and was crushed to <5–10 mm in a steel jaw crusher. Splits of crushed samples (~50 g) were ground in agate to fine powder. Olivine, pyroxenes and garnet were handpicked from sieved 0.5–1.0 mm size fractions of the remaining material to produce grain mounts for micro-beam analyses.

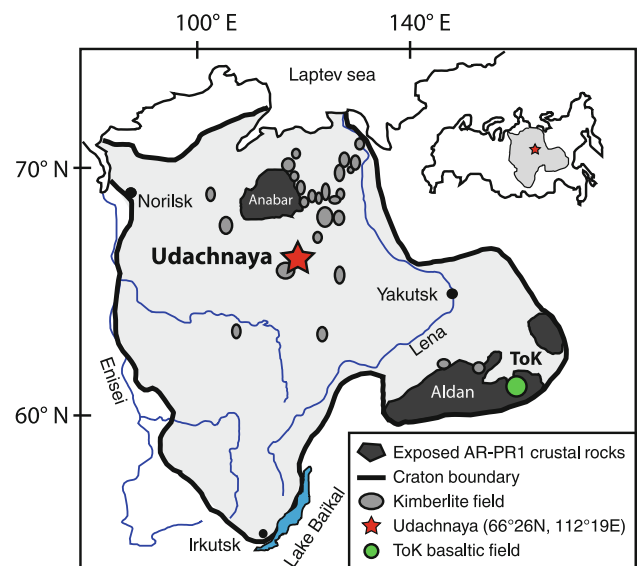
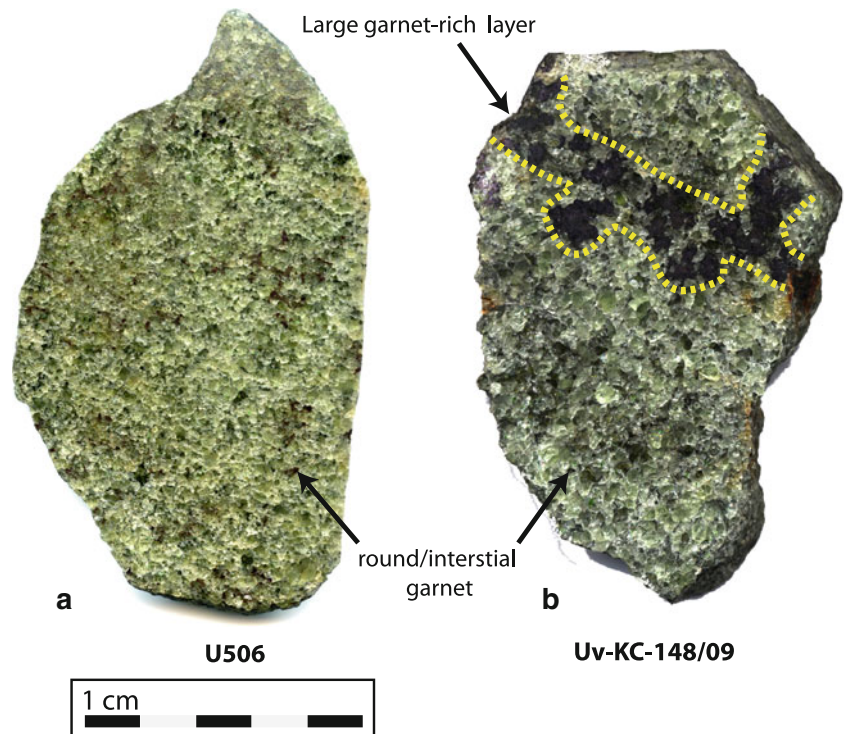


Fig. 1 Location map of Udachnaya, kimberlite fields and areas of exposed early Precambrian crust (Aldan and Anabar shield) in the Siberian craton. Adapted from Ashchepkov et al. (2010)

Fig. 2 Photographs of sawn hand specimens from fresh granular peridotite xenoliths collected at the Udachnaya kimberlite pipe. **a** Garnet harzburgite U506.

b Heterogeneous garnet peridotite Uv-KC-148/09 collected together with peridotite xenoliths used in this study



Analytical methods

The abundances of major and minor elements in bulk-rocks were determined by wavelength-dispersive (WD) X-ray fluorescence (XRF) spectrometry at J. Gutenberg University, Mainz. The rock powders were first ignited for 3 h at 1,000 °C to convert all FeO into Fe₂O₃ and expel water and CO₂. Glass beads, produced by fusing 0.8 g of the ignited powders with 4.8 g of dried LiB₄O₇ (1:7 dilution), were analysed on a Philips PW1404 instrument using ultramafic and mafic reference samples as external standards. Reference sample JP-1 was run as an unknown to control precision and accuracy with results close to recommended values; full duplicates of Uv-KC-03/08 reproduced within 0.05 wt% for LOI, SiO₂, MgO and FeO, and ≤0.01 wt% for other elements (Online Resource 1).

Major element compositions of minerals were obtained by WD electron probe micro-analysis (EPMA) at Laboratoire Magmas et Volcans (LMV, Clermont-Ferrand) on a CAMECA SX-100 using 15 keV voltage, 15 nA sample current and counting times of 20–30 s for peaks and 10 s for background; standards were natural and synthetic minerals; ZAF correction was applied. The minerals were normally analysed in grain mounts, usually 4 different grains for clinopyroxene (cpx) and garnet, and 6 grains for olivine and orthopyroxene (opx), less commonly in thin sections to explore heterogeneities and zoning of mineral grains and their relation to textural position.

Whole-rock trace element compositions were determined by inductively coupled plasma mass-spectrometry (ICPMS) at Université Montpellier II following a modified method of Ionov et al. (1992). The powders (100 mg) were dissolved in HF-HClO₄. Dried samples were dissolved in HNO₃ and diluted in 2 % HNO₃ to 1:2,000 shortly before analysis. The solutions were analysed on an Element XR instrument together with four blanks, three duplicates (separate dissolutions) and two duplicates of JP-1. Chemical blanks were 0.5 ppm for Ba, 0.002–0.1 ppm for Li, Sr, Zr and Rb, 0.01–0.001 ppm for Y, Nb, La, Ce, Hf and Th, and <0.001 ppm for other rare earth elements (REE), Ta and U.

Minerals were analysed for trace elements by laser-ablation (LA) ICPMS in grain mounts at the LMV. The AGILENT 7500 ICPMS instrument is coupled with an Excimer 193 nm Resonetics M-50E ATL operated at 7 Hz, 6 mJ cm⁻² pulse energy and 70–100 μm beam size. Helium was used as carrier gas. Acquisition time was 90 s for background and 60 s for each signal. Reference sample SRN NIST 612 was used as an external standard (Gagnon et al. 2008). Data were reduced with GLITTER software.

Petrography

Photomicrographs of representative xenoliths are shown in Fig. 3; those for all xenoliths in this study (full-size images

Table 1 Summary of petrological data for Udachnaya xenoliths in this study

Sample no	WR (g)	Rock type	Mg# (ol)	Cr# (gar)	P (GPa)	T, °C (Ca-opx)	P (GPa)	T, °C (Taylor)	Calculated modal abundances, wt%			
									ol	opx	cpx	gar
<i>Coarse garnet peridotites</i>												
Uv-02/03	90	Hzb	0.925	0.243	4.5	848	–	–	80.2	15.9	0.0	3.9
Uv-604/09	188	Hzb	0.924	0.150	3.5	939	–	–	84.6	10.7	0.0	4.7
Uv-419/09	380	Hzb	0.922	0.164	5.0	931	5.1	952	69.4	21.5	4.5	4.6
U29	390	Lh	0.922	0.183	5.3	940	5.1	886	68.6	20.2	5.6	2.1
U64	150	Hzb	0.909	0.336	6.0	1,175	6.2	1,219	77.1	16.6	3.8	2.5
U260	350	Hzb	0.917	0.223	3.5	846	–	–	90.0	4.5	4.5	1.0
U280	510	Hzb	0.919	0.127	3.6	909	2.7	720	85.1	9.3	1.3	4.3
U283	143	Hzb	0.920	0.149	3.6	883	3.2	796	84.5	10.2	1.4	3.7
U501	220	Hzb	0.917	0.149	4.8	873	4.9	882	82.4	6.6	4.3	6.7
U506	190	Hzb	0.925	0.187	6.1	1,031	5.5	943	77.2	16.0	2.6	4.2
U508	110	Hzb	0.923	0.159	3.3	844			85.4	7.9	0.0	6.7
U1147	140	Hzb	0.925	0.139	5.1	991	4.7	915	64.5	26.4	1.0	8.1
U1188	200	Hzb	0.916	0.270	6.6	1,318	6.6	1,322	69.9	23.1	2.0	5.0
<i>Transitional garnet peridotites</i>												
Uv-01/03	160	Hzb	0.917	0.365	6.2	1,251	–	–	62.6	33.0	0.0	4.4
Uv-408/09	350	Hzb	0.908	0.170	6.3	1,300	6.4	1,301	67.4	22.0	4.2	6.4
Uv-103/03	133	Hzb	0.914	0.413	6.5	1,300	6.5	1,305	81.1	13.8	3.5	1.6
Uv-421/09	510	Hzb	0.908	0.210	5.6	1,256	5.7	1,278	70.9	20.8	3.5	4.8
Uv-573/09	270	Lh	0.911	0.144	6.0	1,283	6.1	1,289	69.6	17.4	5.6	7.4
Uv-KC-03/08	277	Hzb	0.910	0.181	6.3	1,294	6.3	1,305	83.2	9.4	4.1	3.3
Uv-KC-67/08	280	Hzb	0.912	0.243	6.2	1,286	6.3	1,300	68.7	23.3	3.9	4.1
U4	204	Lh	0.912	0.214	5.8	1,283	5.8	1,301	70.8	19.3	5.5	4.4
U10	200	Hzb	0.911	0.302	6.2	1,273	6.3	1,298	81.4	12.1	4.0	2.5
U50	220	Lh	0.910	0.140	6.1	1,297	6.1	1,294	65.2	22.0	5.9	6.9
U71	310	Lh	0.912	0.176	6.4	1,306	6.3	1,296	76.5	11.2	4.7	7.6
<i>Sheared garnet peridotites</i>												
U9	220	Hzb	0.905	0.294	5.8	1,252	–	–	70.3	24.1	0.0	5.6
U57	550	Hzb	0.911	0.341	6.5	1,291	6.5	1,299	85.2	8.5	4.0	2.3
U70	410	Lh	0.901	0.154	5.2	1,224	5.2	1,235	79.0	9.5	6.0	5.5
U148	91	Lh	0.906	0.139	5.9	1,258	6.0	1,271	69.7	16.4	6.2	7.7
U183	190	Hzb	0.915	0.285	6.3	1,198	6.5	1,244	80.8	13.5	1.9	3.8
U503	140	Hzb	0.912	0.313	6.4	1,283	6.6	1,320	75.2	14.0	3.7	7.1
U532	204	Lh	0.910	0.311	5.8	1,277	5.8	1,289	86.4	5.7	4.9	3.0

Samples with prefix “Uv” are from this study, those with prefix “U” are from Ionov et al. (2010). Hzb, harzburgite; Lh, lherzolite. Transitional peridotites contain <10–20 % of anhedral neoblasts; the neoblasts in sheared peridotites are smaller and may make up 60–90 % of total olivine. WR, weight of crushed whole-rock samples. Mg#, Mg/(Mg + Fe)_{at}; Cr#, Cr/(Al + Cr)_{at}. Ol, olivine; opx, orthopyroxene; cpx, clinopyroxene; gar, garnet. Equilibration pressures (*P*) and temperatures (*T*) are from Nickel and Green (1985) opx-gar barometer, Taylor (1998) cpx-opx thermometer and Brey and Köhler (1990) Ca-in-opx thermometer corrected as in Nimis and Grütter (2010). Modal estimates were obtained by least-squares from whole-rock and mineral major oxide compositions

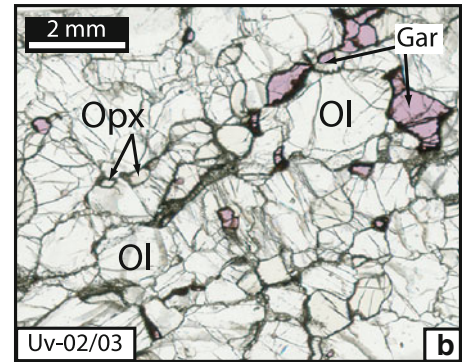
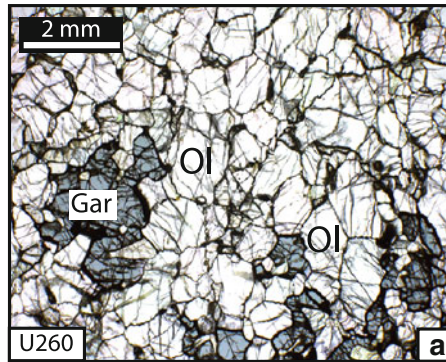
of 4.5 cm × 2.5 cm sections in transmitted plane-polarized and cross-polarized light) are given in Online Resource 2. Based on microstructures, they are grouped into coarse, transitional and sheared.

Only three out of ten new medium- to coarse-grained xenoliths selected for this study can be considered as classical “coarse” peridotites with protogranular to mosaic

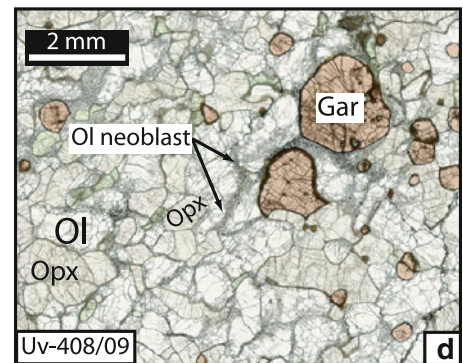
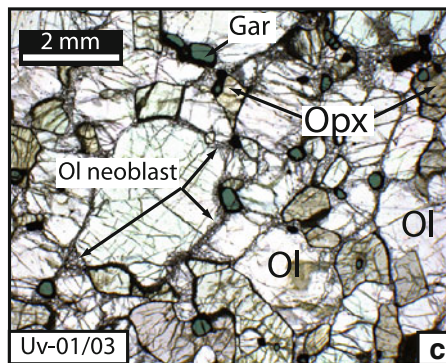
equigranular microstructures (Fig. 3a, b). Another seven contain small amounts (usually <10 %) of anhedral olivine neoblasts and sub-grains at rims of coarse, strained olivine (Fig. 3c, d) and thus must be, strictly speaking, classified as porphyroclastic even though the contents of neoblasts are very low. Following Ionov et al. (2010), we classify such xenoliths as “transitional” (between coarse and sheared),

Fig. 3 Photomicrographs of garnet peridotites from Udachnaya in transmitted plane-polarized light showing representative microstructures; scale bars are at upper left, sample numbers are at bottom left. *Ol* Olivine, *Opx* orthopyroxene, *Cpx* clinopyroxene, *Gar* garnet. Two samples are shown for each microstructural type: **a–b** coarse, **c–d** transitional (incipient deformation, rare sub-grains and neoblasts at rims of coarse olivine), **e–f** sheared (olivine mainly as fine-grained neoblasts)

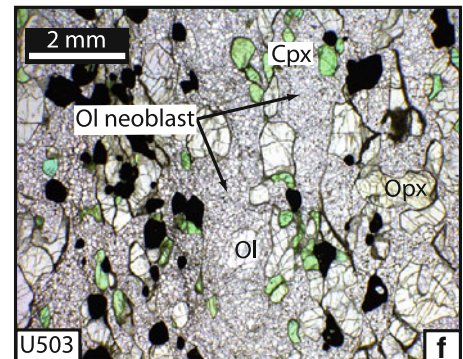
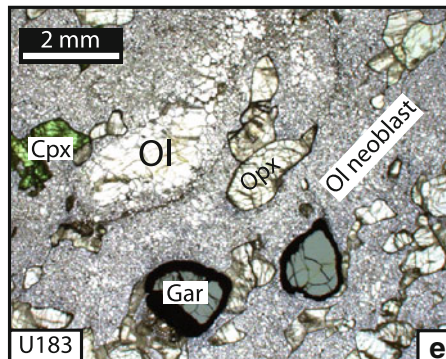
Coarse garnet peridotites



Transitional garnet peridotites

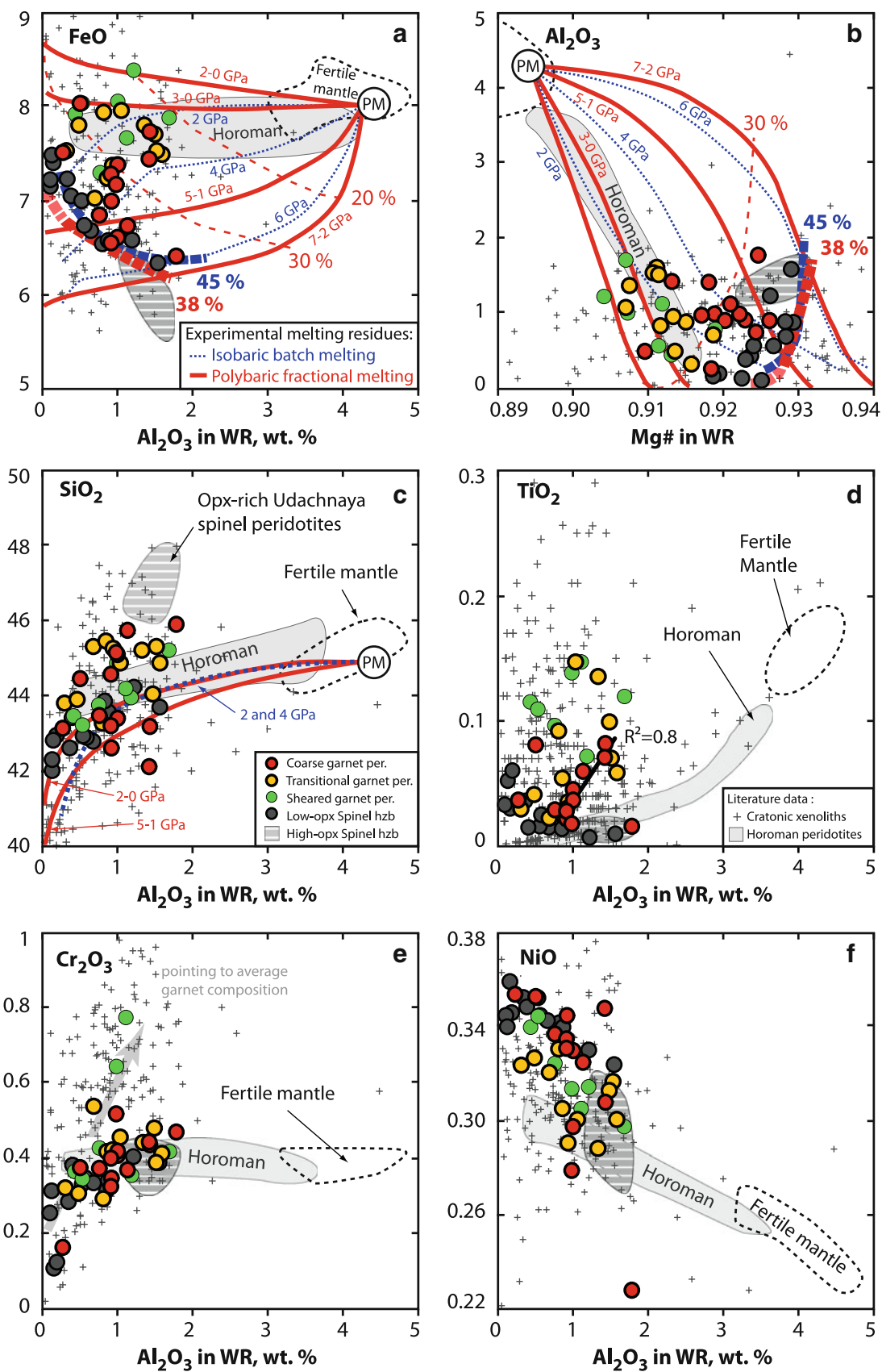


Sheared garnet peridotites



that is, showing incipient stages of deformation. The transitional microstructures with rare neoblasts essentially restricted to grain boundaries of coarse olivine (Fig. 3c, d) can only be recognized in very fresh rocks, like those in this work because fine-grained interstitial olivine is the first to be affected by alteration. We argue that such microstructures may have been overlooked in previous studies of altered peridotite xenoliths in the Siberian and other cratons. If this is correct, some data on “coarse” cratonic peridotites in the literature may actually have been obtained on transitional rocks with low deformation degrees.

Together with xenoliths initially reported by Ionov et al. (2010), 13 samples in this study are coarse and 11 transitional. Nearly all (19 out of 24) are harzburgites ($\leq 5\%$ cpx; four are cpx-free), five are low-cpx lherzolites (5–6 % cpx). Olivine usually has equant shapes and larger grains than opx (Fig. 3a–d); cpx forms small grains in the vicinity of opx. Garnet (1–8 %) ranges from rare interstitial grains to large irregular aggregates or equant grains with smooth boundaries (Fig. 3a). Seven sheared peridotites are harzburgites or low-cpx lherzolites (5–6 % cpx) and contain 2–7 % garnet. They are porphyroclastic, usually with $>50\%$ of olivine as fine-grained neoblasts (Fig. 3e, f) and



◀ **Fig. 4** Co-variation plots of major and minor oxides (wt%) and Mg# [$\text{Mg}/(\text{Mg} + \text{Fe})_{\text{at}}$] for coarse, transitional and sheared whole-rock garnet peridotites from Udachnaya in this study [see (a), (c) and (d) for symbols]. Also shown are primitive mantle (PM) after McDonough and Sun (1995); fertile off-craton garnet and spinel peridotite xenoliths from Vitim and Tariat in central Asia (*fine black dashed lines*; Ionov et al. 2005a; Ionov and Hofmann 2007); low-opx (*dark grey circles*) and high-opx (*grey dashed field*) spinel peridotites from Udachnaya after Doucet et al. (2012); Horoman massif peridotites (*grey contours*; Takazawa et al. 2000); cratonic peridotite xenoliths (*thin crosses*) from the Kaapvaal (Simon et al. 2007; Pearson et al. 2004), Tanzanian (Lee and Rudnick 1999), North Atlantic (Bernstein et al. 1998, 2006; Wittig et al. 2008) and Slave (Irvine et al. 2003; Kopylova and Russell 2000) cratons. *Thin dashed blue lines* show isobaric batch melting residues formed at 2, 4 and 6 GPa; continuous red lines are residues of polybaric fractional melting at 2–0, 3–0, 5–1 and 7–2 GPa (Herzberg 2004). *Thick dashed blue lines* show residues of 45 % batch melting between 2 and 6 GPa; thick dashed red lines show 38 % of polybaric fractional melting

relics of strained olivines with abundant sub-grains. Proportions of coarse and neoblastic olivine, as well as the size of the neoblasts, may vary locally. Pyroxenes are elongated and commonly show alteration at rims and cracks (Fig. 3e). Garnets are often round; they range in size from tiny grains to >2 mm and usually have thick kelyphitic rims (Fig. 3e, f). The recrystallization shows little, if any, effect on garnet.

Even though we took care to avoid xenoliths with kimberlite veins, some interstitial kimberlite-related materials, sometimes with late-stage phlogopite, were found in thin sections in some samples (Online Resource 3). They may affect whole-rock abundances of alkalis, but not of major oxides (Si, Mg, Al, Fe), whose contents in the kimberlites are not very different from those in the peridotites (Kamenetsky et al. 2012).

Major element and modal compositions

Whole-rock (WR) major oxide and LOI data are given in Online Resource 1 and are shown as co-variation plots in Fig. 4. The LOI values range from 1.5 % to –0.5 % (average 0.3 %); they indicate that secondary alteration is limited, that is, nearly absent in many samples and low in others, even in samples with fine-grained olivine sensitive to alteration. The negative LOI, that is, gain of mass on ignition (up to 0.5 %) means that oxidation of FeO to Fe₂O₃ is more significant than the loss of volatiles (usually H₂O and CO₂ from alteration products) if the latter are present.

Coarse and deformed peridotites have similar ranges of SiO₂ (42–46 %), Al₂O₃ (0.3–1.8 %) and CaO (0.6–1.7 %), but coarse peridotites tend to have higher Mg#_{WR} [$\text{Mg}/(\text{Mg} + \text{Fe})_{\text{at}}$] (average: ~0.920 vs. 0.912) and lower FeO (average: 7.1 vs. 7.7 %); sheared peridotites have the highest TiO₂ (Fig. 4).

On diagrams of Al₂O₃ versus FeO and Mg# (Fig. 4a, b), some coarse garnet peridotites plot close to low-opx spinel peridotites from Udachnaya, which were previously interpreted as pristine melting residues based on Al–Fe–Mg# relations in comparison with experimental data (Doucet et al. 2012). The majority of coarse garnet peridotites, however, have higher FeO at given Al₂O₃ than the spinel peridotites, that is, show Fe-enrichments relative to inferred melting residues. The garnet peridotites also tend to have higher TiO₂ than the spinel peridotites and residual peridotite suites, for example, the Horoman massif (Fig. 4d). Some garnet peridotites have slightly higher SiO₂ than the low-opx spinel peridotites (average: 44.5 vs. 43 %) (Fig. 4c), but none contains as much silica as opx-rich spinel peridotites from Udachnaya (average 46.8 %). The range of Cr₂O₃ in garnet peridotites (0.12–0.78 %) extends to higher values than in the spinel peridotites (0.1–0.4 %), but very high Cr₂O₃ (>0.6 %), common in published data on cratonic peridotites, was found only in two sheared, garnet-rich samples in our suite (Fig. 4e). The contents of NiO in many coarse peridotites are close to those in spinel peridotites (0.33–0.35 %), but the deformed rocks (and four coarse peridotites with >20 % opx) have lower NiO (0.23–0.33 %; Fig. 4f).

EPMA obtained in this study is given in Online Resource 1. Mineral grains are normally not zoned. Mg# of olivine (Mg#_{OI}) ranges from 0.900 to 0.925 and is generally lower in deformed (0.900–0.917) than in coarse garnet peridotites, the latter commonly have Mg#_{OI} ≥ 0.92, which overlap the lowest Mg#_{OI} in spinel peridotites (Fig. 5a–c). Mg#_{OI} and Mg#_{Opx} define a positive linear co-variation (Fig. 5a), indicative of equilibrium Mg–Fe distribution between olivine and opx. By contrast, Mg#_{OI} shows no regular correlation with Mg# in coexisting cpx (Fig. 5a). This may either indicate that cpx in many garnet peridotites is not chemically equilibrated with coexisting olivine and thus may be of late-stage origin or be due to the effects of pressure and temperature on cpx–garnet equilibria (Brey and Köhler 1990). The contents of Al₂O₃ in opx from garnet peridotites are lower than in opx from nearly all spinel peridotites (except four ultra-depleted rocks with <0.2 wt% Al₂O₃) because Al strongly partitions to garnet; Al₂O₃ in opx is negatively correlated with Mg#_{OI} in deformed peridotites (Fig. 5b).

Classification schemes for mantle-derived garnets are mainly based on their Ca–Cr relations (Grütter et al. 2004; Sobolev 1974). CaO and Cr₂O₃ contents of garnets in this study range broadly and are positively correlated (Fig. 5d) as is common for garnets in cratonic peridotite xenoliths (Grütter et al. 2004; Pearson et al. 2003). Only one garnet (from coarse cpx-free sample Uv-02/03) can be classified as “harzburgitic” G10 (low-Ca), the most abundant type recovered in diamond exploration the world over, which is

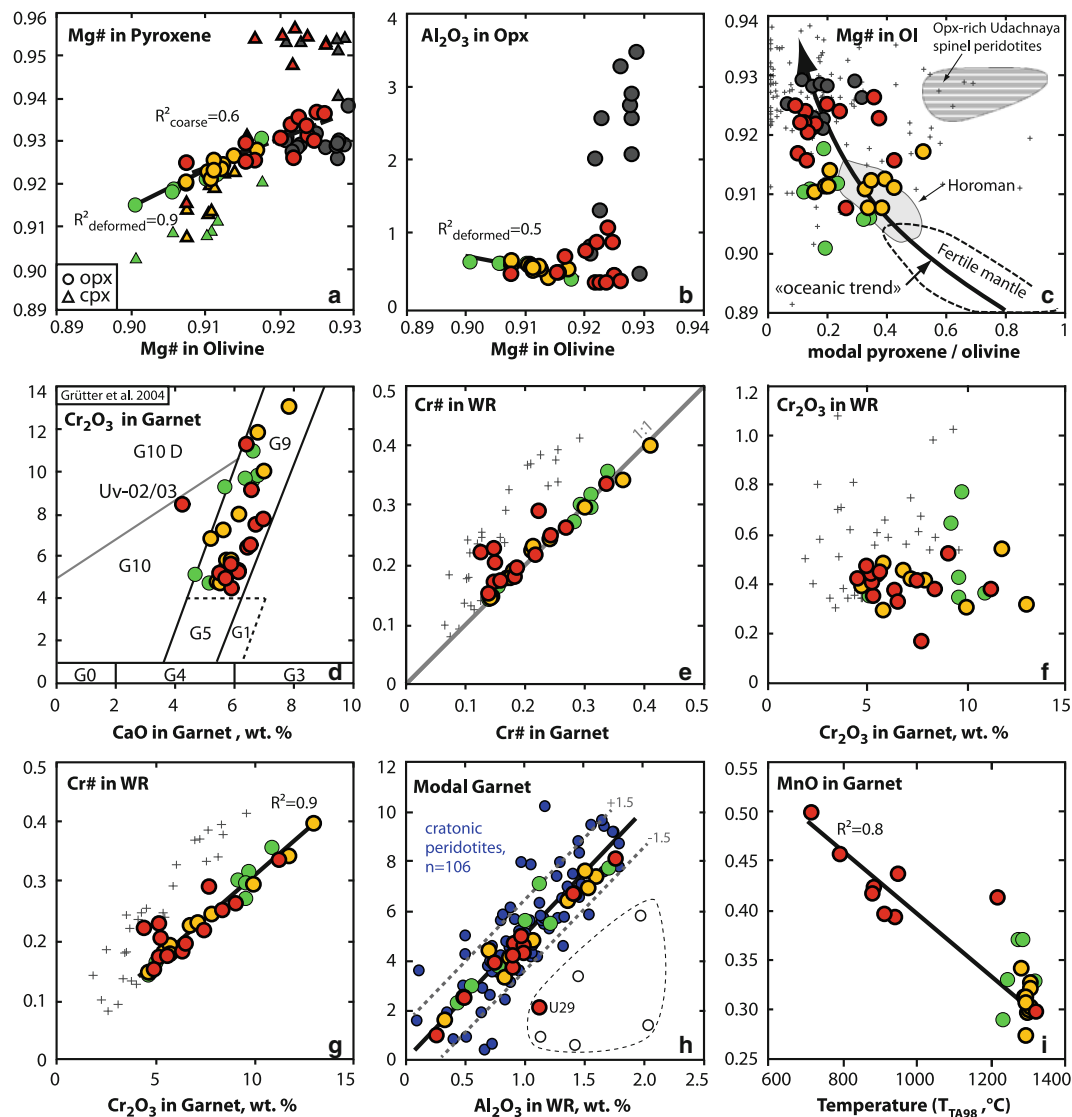


Fig. 5 Co-variation plots of major element and modal abundances (wt%) and ratios for olivine, pyroxenes and garnet from Udachnaya peridotites. **a** $Mg\#_{Oliv}$ [$Mg/(Mg + Fe)_{at}$ in olivine] versus $Mg\#_{opx}$ and $Mg\#_{cpx}$; **b** $Mg\#_{Oliv}$ versus Al_2O_3 in opx; **c** modal pyroxene/olivine versus $Mg\#_{Oliv}$; **d** classification scheme for mantle-derived garnet from Grütter et al. (2004); **e** $Cr\#$ [$Cr/(Cr + Al)_{at}$] in garnet versus $Cr\#$ in WR; **f** Cr_2O_3 in garnets versus Cr_2O_3 in WR and **g** $Cr\#$ in WR; **h** Al_2O_3 in WR versus modal garnet for Udachnaya garnet peridotites

(this study) and for worldwide cratonic peridotites (blue circles Boyd et al. 1997; Lee and Rudnick 1999; Kopylova and Russell 2000; Irvine et al. 2003; Simon et al. 2007; Pearson et al. 2004; Wittig et al. 2008; fine dashed lines, cpx- and spl-rich peridotites); **i** Equilibration temperature after Taylor (1998) (T_{T98} , °C) versus MnO in garnet. Also shown are Horoman peridotites, fertile off-craton mantle and worldwide cratonic peridotite xenoliths [see Fig. 4 for symbols and references]

believed to originate from cpx-free rocks (Grütter et al. 2004). Garnets from all other samples in this study, including four cpx-free harzburgites, fall into the “Iherzolitic” G9 field (Fig. 5d). $Cr\#_{Gar}$ [$Cr/(Cr + Al)_{at}$ in garnet] are identical to $Cr\#_{WR}$ in the deformed, and all but four, coarse peridotites, the latter showing slightly higher $Cr\#_{WR}$ like many samples from literature (Fig. 5e). $Cr\#_{Gar}$ show a positive correlation with CaO and a negative correlation with MgO. The contents of Cr_2O_3 in garnets are not correlated with whole-rock Cr_2O_3 (Fig. 5f), but show an excellent ($r^2 = 0.9$) positive linear correlation with

$Cr\#_{WR}$ (Fig. 5g). $Cr\#$ appears to be a very useful chemical parameter of mantle garnets because it offers direct links to whole-rock chemistry.

Modal abundances were calculated from major element compositions of bulk-rocks and minerals. On average, coarse peridotites have slightly higher modal olivine (78 vs. 75 %) , lower opx (14.5 vs. 16.4 %) and cpx (2.4 vs. 4 %) than deformed peridotites, but similar garnet (~5 %; Fig. 6a). Coarse garnet peridotites plot close to spinel peridotites on a diagram of $Mg\#_{Oliv}$ versus pyroxene/olivine ratios, but generally apart from deformed peridotites

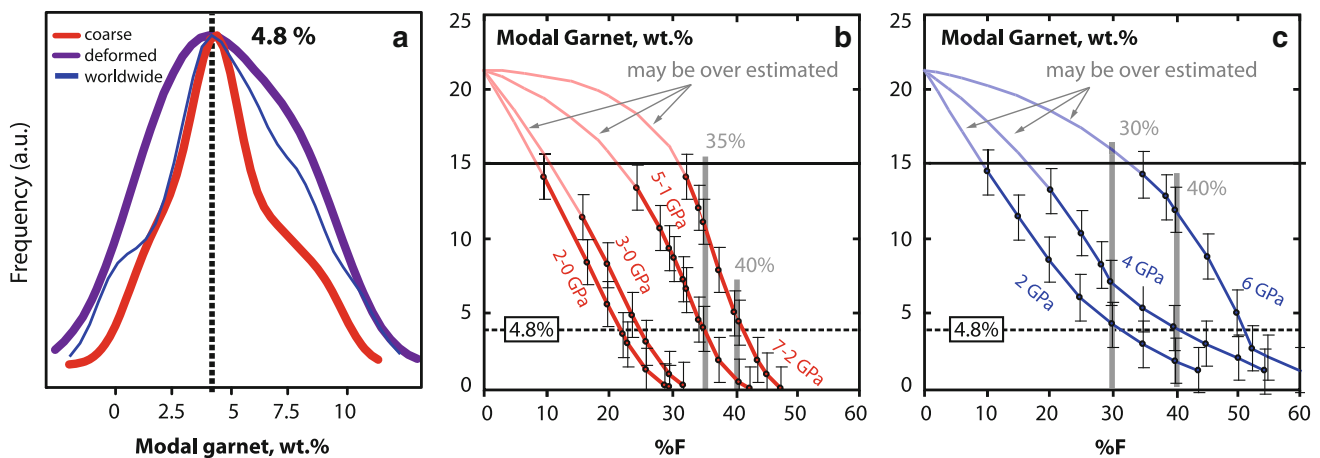


Fig. 6 **a** Frequency distribution diagram for modal garnet in coarse (red line) and deformed (purple line) peridotite xenoliths from Udachnaya and worldwide cratonic peridotite xenoliths (blue line) [see Fig. 5 for references]. **b–c** Calculated modal garnet in melt residues versus degree of partial melting (%F) for polybaric fractional melting (**b**) and isobaric batch melting (**c**). Modal garnet in melting residues in experiments of Herzberg (2004) was estimated as follows: Garnet (%) = Al_2O_3 (wt% in WR) \times 5–0.15 defined by a linear

correlation between Al_2O_3 in WR and modal garnet for cratonic peridotites (Fig. 5h). The equation is well constrained (error 1.5 %) for $\text{Al}_2\text{O}_3 \leq 2\%$ (12 % garnet), that is, modal garnet may be overestimated for $\text{Al}_2\text{O}_3 > 2\%$. Horizontal dashed lines in **b–c** show the most common modal garnet values (4.8 %) in peridotites from Udachnaya and worldwide cratons; thick grey vertical lines show melting degree estimates for different pressures at 4.8 % residual garnet

(Fig. 5c), which is similar to cratonic peridotites elsewhere (Pearson and Wittig 2008; Simon et al. 2007). Modal garnet is positively correlated with whole-rock Al_2O_3 ($r^2 = 0.9$) (Fig. 5h). A comparison with literature data shows that such linear correlation may be a general feature of cratonic garnet peridotites. Thus, garnet abundances are controlled by whole-rock Al_2O_3 while garnet major oxide composition mainly depends on $\text{Cr}\#\text{WR}$ or $\text{Cr}/\text{Al}\text{WR}$.

Temperature and pressure estimates

Equilibration pressures (P) and temperatures (T) were calculated from average core compositions of minerals using several poly- and mono-mineral thermobarometers commonly applied to cratonic peridotites: the cpx-opx thermometer of Taylor (1998) (T_{TA98}), the Ca-in-opx thermometer of Brey and Köhler (1990) ($T_{\text{Ca-in-opx}}$) corrected by Nimis and Grütter (2010), the opx-gar barometer of Nickel and Green (1985) (P_{NG85}) as well as the single-cpx thermo-barometer of Nimis and Taylor (2000) (P_{NT00}) and the single-garnet thermo-barometer of Ryan et al. (1996) ($P_{\text{Cr-TNi}}$). The latter is based on the abundances of Cr (for P) and Ni (for T) and is mainly used to outline cratonic geotherms from data on garnet xenocrysts if no garnet peridotite xenoliths are available (Griffin et al. 1999). A recent re-assessment of thermo-barometry methods (Nimis and Grütter 2010) considers the $P_{\text{NG85}}-T_{\text{TA98}}$ combination the most accurate for mantle peridotites.

The $P_{\text{NG85}}-T_{\text{TA98}}$ estimates (Table 1; Fig. 7a) define a broad P – T range for coarse peridotites (2.7–6.6 GPa,

720–1,322 °C) and a narrow P – T range for deformed peridotites (5.2–6.6 GPa, 1,235–1,320 °C). They suggest a lithosphere-asthenosphere boundary beneath Udachnaya at ~ 7 GPa (~ 220 km) and show that both deformed and coarse peridotites occurred near the base of the lithosphere at the time of kimberlite eruption (~ 360 Ma). P – T values from most other methods yield somewhat different ranges (Table 1; Fig. 7b, c) but are generally consistent with the $P_{\text{NG85}}-T_{\text{TA98}}$ estimates, that is, are scattered between the 35 and 45 mW m^{-2} model conductive geotherms (Pollack and Chapman 1977) and define no single P – T gradient, similar to published data for Udachnaya xenoliths obtained using different samples and methods (Boyd et al. 1997; Goncharov et al. 2012; Ionov et al. 2010).

The single-garnet thermo-barometer of Ryan et al. (1996) yields coherent P – T estimates for low- T ($\leq 1,050$ °C) peridotites, which plot between the 35 and 40 mW m^{-2} conductive geotherms, but unrealistic values (low P , high T) for high- T peridotites, which plot above the graphite/diamond boundary and partly on the mantle adiabat (Fig. 7d). This method, calibrated assuming Cr equilibrium between garnet and Cr-rich spinel, underestimates pressure if spinel is absent. The agreement between the single-garnet and poly-mineral thermo-barometers for the low- T Udachnaya peridotites may imply that their garnets are in equilibrium with Cr-spinel, that is, that accessory spinel may be present in some of the rocks, or that garnet peridotites are intercalated with spinel peridotites on a metre scale. The latter is consistent with overlapping T estimates for the Udachnaya spinel and low- T garnet peridotites (Doucet et al. 2012; Goncharov et al.

Fig. 7 Plots of equilibration temperature (T) versus pressure (P) for Udachnaya peridotites in this study. **a** $P_{NG85}-T_{TA98}$, **b** $P_{NG85}-T_{Ca-in-opx}$, **c** $P_{NT00}-T_{NT00}$ and **d** $P_{Cr}-T_{Ni}$ [see Fig. 4 for symbols and text for references and details]. $P-T$ fields for spinel peridotites were calculated based on T estimates assuming that the samples plot between the 35 and 45 $mW m^{-2}$ conductive geotherms of Pollack and Chapman (1977) (grey dashed lines). Also shown are mantle adiabat and the graphite-diamond transition boundary (Rudnick and Nyblade 1999) as well as $P-T$ estimates for garnet peridotite xenoliths from South Africa (small black crosses Simon et al. 2007; Boyd et al. 2004) and North America (triangles Kopylova and Caro 2004; Sand et al. 2009; Wittig et al. 2008; Schmidberger and Francis 1999) re-calculated with the same thermo-barometers as for samples in this study

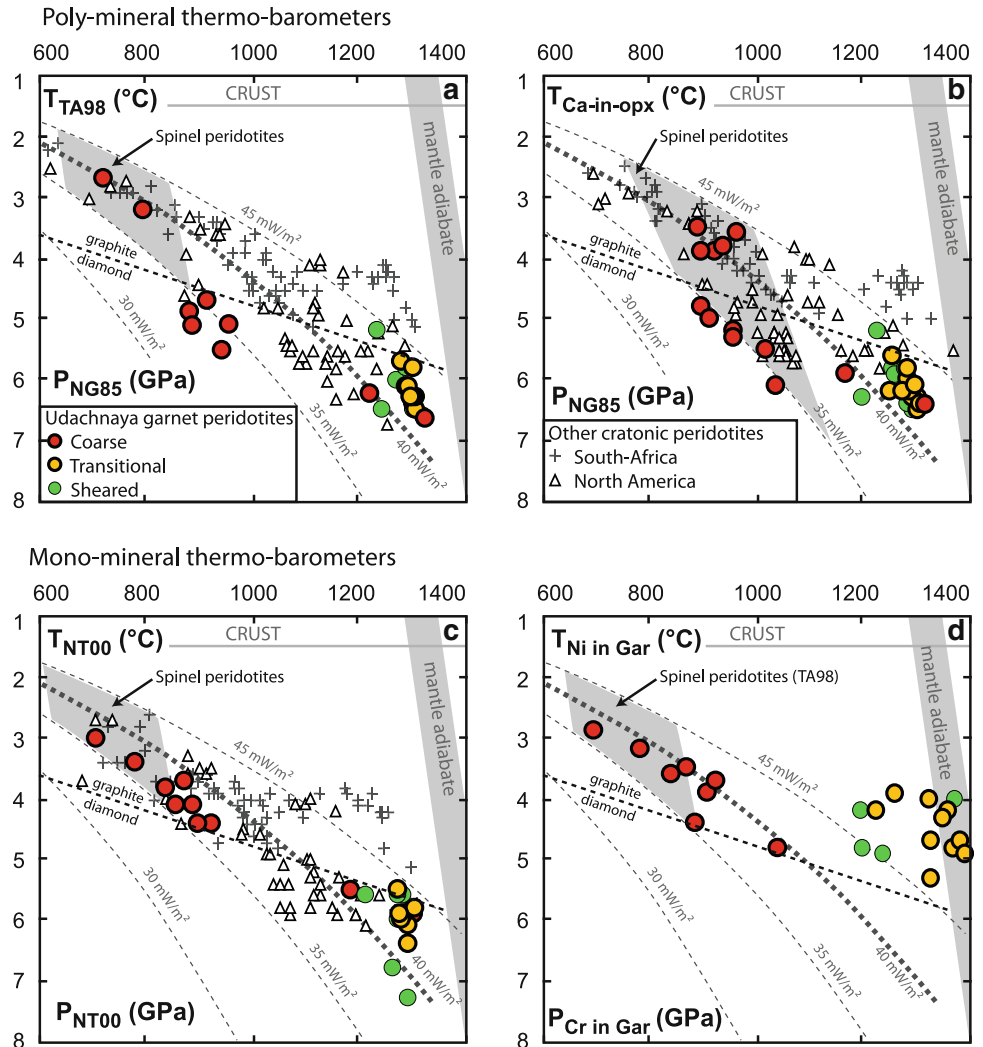


Fig. 8 Primitive mantle-normalized (McDonough and Sun 1995) rare earth element (REE) patterns for garnet in coarse (a), transitional (b) and sheared (c) Udachnaya peridotites, and d REE partitioning between garnet and cpx ($D^{gar/cpx}$) for Udachnaya peridotites in this study, off-craton peridotite xenoliths from Vitim (black line (Ionov 2004) and experimental data compiled for pMELTS software (dashed black line (Ghiorso et al. 2002; Smith and Asimow 2005)

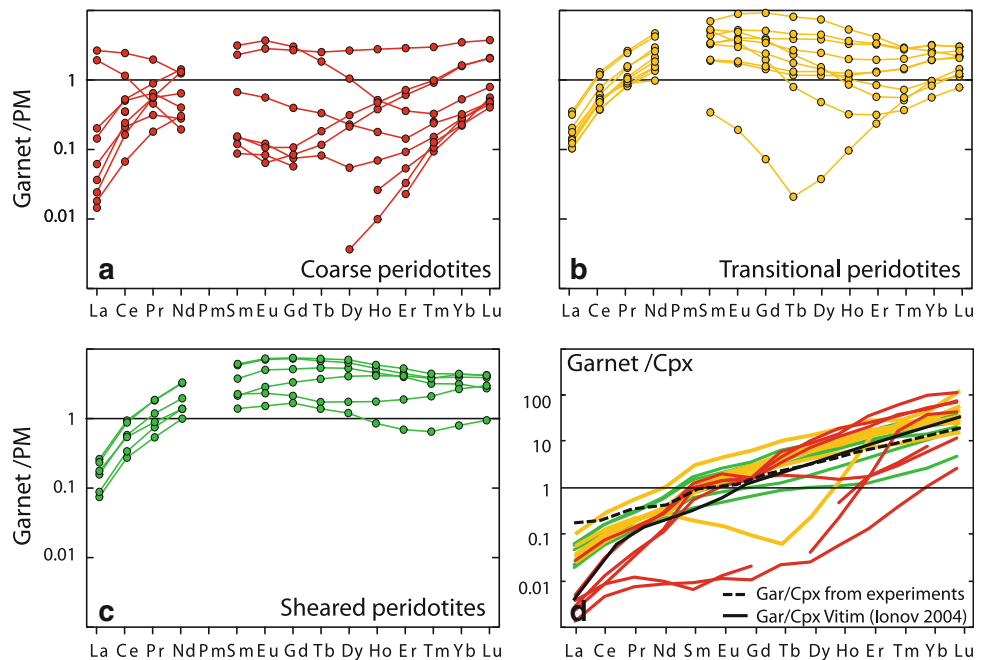
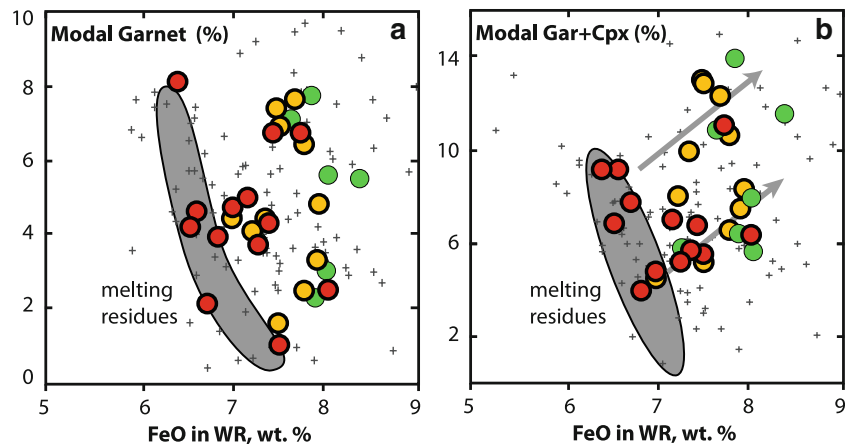


Fig. 9 Co-variation plots of FeO (wt%) in WR garnet peridotites from this study versus modal garnet (a) and modal garnet + cpx (b) in comparison with the field (grey) for melting residues estimated from Udachnaya spinel peridotites (Doucet et al. 2012) and worldwide cratonic garnet peridotites (crosses, see Fig. 4 for references)



2012; Ionov et al. 2010). To sum up, the method of Ryan et al. (1996) applied to our samples may only be valid at $\leq 4\text{--}5$ GPa (130–150 km) and ignores the complexity of the perturbed local geotherm. The application of this technique to assess variations of mantle compositions with depth for the Siberian craton (Griffin et al. 1999), in particular at >130 km, does not appear to yield valid results and is not consistent with the vertical distribution of rock types established by xenolith studies (Ionov et al. 2010).

The contents of MnO in garnets from this study are negatively correlated with equilibration T (Fig. 5i) and show no correlation with other oxides in garnet or with MnO in olivine. Mn-in-garnet thermometers for cratonic mantle were earlier proposed by Delaney et al. (1979) and Creighton (2009). The broad T range for coarse Udachnaya peridotites also affects the T -dependant Fe–Mg distribution between olivine and opx (Mori and Green 1978), which shows more scatter for coarse than for deformed peridotites relative to $\text{Mg}\#_{\text{Ol}}$ versus $\text{Mg}\#_{\text{Opx}}$ co-variation trends (Fig. 5a) and explains the lack of correlation of $\text{Mg}\#$ with major oxides in garnet.

Trace element compositions

LA-ICPMS analyses of garnet, cpx and opx are given in Online Resource 1 and are shown as trace element patterns normalized to primitive mantle (PM; McDonough and Sun (1995) in Fig. 8 and Online Resource 4. Garnets show a broad range of REE patterns, from LREE–MREE-depleted to “sinusoidal” and “humped” (high-MREE). In the majority of coarse peridotites, garnets show depleted to sinusoidal REE, except U501, which has nearly flat MREE–HREE with LREE-depletion. Garnets in transitional and sheared rocks tend to have higher MREE with mildly sinusoidal to humped patterns (Fig. 8a–c). Clinopyroxenes generally show continuous enrichments from HREE to

MREE and LREE, sometimes with nearly flat LREE, and only minor differences between coarse and deformed peridotites. Gar/cpx ratios for REE ($D^{\text{gar/cpx}}$) in several coarse peridotites are much lower, in particular for LREE, than those for mantle xenoliths from Vitim (Ionov 2004) and experimental data compiled for pMELTS software (Ghiorso et al. 2002; Jenner and O’Neill 2012), which may be attributed to incomplete equilibration between garnet and cpx (Simon et al. 2003). Data for opx are less precise because of abundances 1–3 orders of magnitude lower than for garnet and cpx. In general, their REE patterns resemble those in coexisting cpx both in coarse and deformed peridotites (Online Resource 4).

Trace element compositions of WR peridotites are listed in Online Resource 1 and illustrated in Online Resource 4. Their patterns for highly to moderately incompatible elements (MREE to Th) are similar to those of host kimberlites (Kamenetsky et al. 2012) and appear to be mainly controlled by contamination with host magmas, consistent with evidence for kimberlite-related interstitial material. By contrast, the HREE patterns (a decrease from Lu to Tm) are opposite to those of the kimberlites and may be similar to those in the mantle source for peridotites. Similar observations were reported from other cratonic garnet peridotite suites (Wittig et al. 2008).

To estimate REE abundances in peridotites before contamination by the host magma, we calculate them from modal compositions and mineral LA-ICPMS data (Online Resources 1 and 4). The contamination-free compositions of coarse peridotites show continuous HREE-depletions from Lu to Ho or spoon-shaped HREE–MREE (Online Resource 4), except for U501, which has a flat pattern. Deformed peridotites have sinusoidal, flat and humped patterns, except for U4, which shows depletion from Lu to Tb and enrichment from Tb to La. These estimates may be a better approximation to “true” HREE–MREE abundances in peridotites than measured ones because these elements in mantle peridotites mainly reside in major minerals

(rather than in accessories and interstitially) (Bedini and Bodinier 1999; Eggins et al. 1998; Sharygin et al. 2007).

Pristine melting residues among garnet peridotites and their formation conditions

To identify “residual” garnet peridotites, whose modal and major element compositions have largely remained intact after partial melting, we compare them with experimental results on melting of fertile mantle (Herzberg 2004; Herzberg and O’Hara 2002; Walter 1998, 2003) and with low-opx spinel harzburgite xenoliths from Udachnaya, which were previously interpreted as melting residues (Doucet et al. 2012). Seven out of 13 coarse garnet peridotites in this study (Uv02/03, Uv419/09, Uv604/09, U29, U260, U506 and U1147) overlap the trend defined on the diagram of bulk-rock Al_2O_3 versus FeO (Fig. 4a) by the spinel harzburgites and also are close to them on plots versus SiO_2 , TiO_2 and NiO (Fig. 4c, d, f). Their modal averages (76 % olivine, 16 % opx, <3 % cpx, 4 % gar) and $\text{Mg}\#_{\text{Ol}}$ (0.923) are similar to those for the spinel harzburgites as well (Figs. 5c and 9); these values may represent the compositional range for the initial residual protolith during craton formation.

Comparisons with experimental data indicate that degrees and depth of partial melting for these rocks appear to be similar to those for the Udachnaya spinel harzburgites. All these rocks overlap lines of equal melt extraction (~ 45 % anhydrous batch melting or ~ 38 % fractional melting) as a function of pressure (Fig. 4a). It is also

possible that some garnet peridotites were formed by somewhat lower melting degrees, down to 30 %, if their iron contents have not been affected by metasomatism (Fig. 9). To sum up, the protolith of garnet peridotites was formed by 30–38 % of fractional polybaric melt extraction, which began at 7–4 GPa and ended at ≤ 1 –3 GPa (Doucet et al. 2012).

Depth range of melting may have important consequences on Cr contents and Cr/Al ratios in the residues because Cr is much more compatible in spinel than in garnet (Canil 2004). Extended melting in the spinel stability field produces Cr-rich residues, whereas melting in the garnet stability field yields Cr-poor residues. The contents of Cr_2O_3 in the majority of cratonic peridotites, reported in the literature, are much higher (>0.45 %) than in reliable data on fertile to refractory massif peridotites and off-craton peridotite xenoliths (Fig. 4e). This has been interpreted as a result of high degrees of melt extraction in shallow, spinel-facies mantle (Canil 2004; Lee et al. 2011), even though well-documented harzburgite series formed by melt extraction in shallow mantle do not normally show $\text{Cr}_2\text{O}_3 > 0.45$ % (Ionov 2010; Takazawa et al. 2000). By contrast, Cr_2O_3 in the majority of Udachnaya garnet peridotites are similar to those in off-craton peridotites and are positively correlated with Al_2O_3 (Fig. 4e). Four Udachnaya garnet peridotites have $\text{Cr}_2\text{O}_3 \geq 0.5$ %, but three of them are deformed, and their high Cr could be due to mechanical sorting of garnet (or its crystallization) in shear zones, complementary to low Cr_2O_3 (≤ 0.3 %) in four other deformed xenoliths.

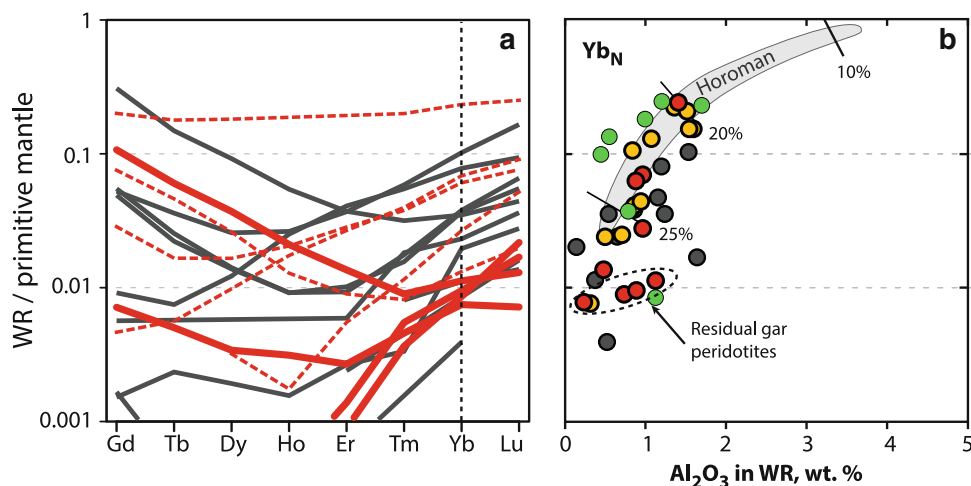


Fig. 10 **a** Primitive mantle-normalized (McDonough and Sun 1995) whole-rock MREE-HREE patterns for coarse garnet peridotites in this study (red) and residual spinel peridotites from Udachnaya (Doucet et al. 2012); continuous thick red lines show xenoliths from this study with HREE patterns of melting residues unaffected by post-melting processes. **b** Whole-rock abundances of Yb normalized to primitive mantle (Yb_N) for Udachnaya garnet (this study) and spinel peridotite

xenoliths (Doucet et al. 2012) in comparison with those in whole-rock Horoman peridotites (Takazawa et al. 2000) and results of trace element modelling for non-modal fractional melting of primitive mantle (McDonough and Sun 1995) modified after Johnson (1998) and Simon et al. (2007) taking into account modal variation of cpx and garnet during decompression melting at 5–1 GPa

The Cr_2O_3 variation range in all but three deformed Udachnaya garnet peridotites is not likely to be due to “nugget effect”, advocated by Canil (2004) because our samples are large (Table 1) and because of the excellent positive correlation between Cr_2O_3 or Cr# in garnet with Cr# in bulk-rocks ($r^2 = 0.8\text{--}0.9$; Fig. 5e, g), which shows that $\text{Cr}\#_{\text{gar}}$ and $\text{Cr}\#_{\text{WR}}$ are equilibrated. The low Cr_2O_3 in both garnet and spinel peridotites suggest that partial melting mainly took place in the garnet stability field, leading to Cr-depletion in the residues, whereas residual garnet was absent or rapidly exhausted for the rocks with $\geq 0.4\%$ Cr_2O_3 . By contrast to our Udachnaya suite, the majority of cratonic peridotite xenoliths reported in the literature shows much higher $\text{Cr}\#_{\text{WR}}$ than $\text{Cr}\#_{\text{gar}}$ (Fig. 5e). For some rocks, this might be due to local chemical disequilibrium, for example, for recently formed garnets. We argue, however, that in many cases, this could be related to analytical problems, in particular when Cr, unlike in this study, is not determined by XRF in fused discs, but analysed separately as a “trace element”.

Heavy REE are sensitive to the presence of garnet during mantle melting due to high garnet/melt partition coefficients and can be used to constrain depth and degrees of melting (Canil 2004; Simon et al. 2007; Wittig et al. 2008). The broad HREE range in calculated whole-rocks (Fig. 10a) may be mainly controlled by modal garnet in the residues. Figure 10b shows Yb abundances in melting residues of primitive mantle (McDonough and Sun 1995) estimated after Johnson (1998) and Simon et al. (2007) from 5 to 1 GPa (Fig. 10b), which takes into account modal garnet variations in experimental residues (Fig. 6b, c). This modelling shows that melting in the garnet stability field may produce HREE-depleted melting residues. However, our modelling does not consider the change of melting parameters (i.e. P , melting reaction, trace element partition coefficients, etc.) during decompression melting, by contrast with thermodynamic modelling (e.g. pMELTS, calibrated for melting from 1 to 3 GPa [Ghiorso et al. 2002]). This can produce uncertainties in trace element behaviour during melting. Contamination-free whole-rock HREE of residual coarse garnet peridotites are too low for residues of 35–45 % of melting in the garnet stability field starting at 3 and 5 GPa and may require at least 20–25 % melting of spinel facies. This is generally in agreement with the estimates of melting degrees and pressures from major oxide data ($\sim 30\text{--}40\%$ between 7 and 4 and $\leq 1\text{--}3$ GPa) and Cr_2O_3 contents.

One coarse, and nearly all deformed garnet peridotites, has nearly flat or Gd–Tm-enriched ($\text{Lu} < \text{Gd}$) patterns, which cannot be produced by melt extraction, and are therefore the result of metasomatism.

The origin of garnet and cpx and their modal variations

Garnet and cpx in cratonic peridotites may have different origins: (1) residual phases after melt extraction, (2) exsolution from high- T , Ca–Al-rich opx on cooling and (3) post-melting enrichments. Experiments give useful but limited hints on garnet and cpx abundances in melting residues. The majority of experimental studies on fertile mantle melting address major element compositions of the residues (Herzberg 2004; Herzberg and O'Hara 2002; Walter 1998), but give limited results on their modal abundances, for example, Walter (1998) provides modal compositions of melting residues only for batch melting at 3 and 7 GPa. Here, we use the well-defined correlation between whole-rock Al_2O_3 and modal garnet in cratonic peridotites (Fig. 5h) expressed as: Modal garnet (%) = Al_2O_3 (wt% in WR) $\times 5 - 0.15$ ($\pm 1.5\%$) to estimate modal garnet in melting residues from data sets of Herzberg (2004) for polybaric fractional melting (starting at 2, 3, 5 and 7 GPa; Fig. 6b) and isobaric batch melting (at 2, 4 and 6 GPa; Fig. 6c). The estimates show that melting residues (re-equilibrated at lower T 's in the lithospheric mantle after the melting) may contain up to 5 % of garnet after 35–40 % of polybaric fractional melting starting at high P (5 and 7 GPa), but that garnet is exhausted after 30 % of melt extraction starting at lower P (Fig. 6b). Garnet is more stable during isobaric batch melting [Fig. 6c and Walter (1998)]. About 2/3 of Udachnaya garnet peridotites (i.e. 2/3 coarse and 1/2 deformed) have $\leq 5\%$ garnet (Fig. 5h). The modal garnet maximum for the whole suite is 4.8 % (Fig. 6a). Hence, garnet in many coarse and transitional peridotites from Udachnaya may be a residual phase after partial melting, like olivine and opx, which is also consistent with depletions from Lu to Gd in such garnets (Fig. 8a, b). By comparison, garnet in xenoliths with $>5\%$ garnet, as well as LREE–MREE enrichments in all garnets, must be a result of post-melting processes (Shimizu 1999; Shimizu et al. 1997; Simon et al. 2003).

Melting experiments and studies of natural residual peridotites show that melting residues may contain small amounts of cpx (up to 2.5 %), for example, due to exsolution from residual opx (Canil and Wei 1992; Saltzer et al. 2001; Takazawa et al. 2000; Walter 1998; Wood and Banno 1973). The majority of Udachnaya garnet peridotites have modal cpx $>2.5\%$. Moreover, poor correlation of Mg# in cpx and in coexisting olivine (Fig. 5a), together with REE disequilibrium between cpx and garnet (Fig. 8d), indicate that cpx may not be chemically equilibrated with olivine in some rocks. This cpx may be post-melting, late-stage in origin (Canil and Wei 1992; Saltzer et al. 2001; Simon et al. 2003).

The variation ranges of FeO (6.4–8.4 %), TiO₂ (0.02–0.15 %) and Mg#_{WR} (0.90–0.92) in the Udachnaya garnet peridotites can be viewed as trends of Fe–Ti enrichments of initial melting residues represented by spinel harzburgites (Fig. 4), similar to those described in many cratonic xenolith suites (Ionov et al. 2005b; Kopylova and Russell 2000; Lee and Rudnick 1999; Rudnick et al. 1993; Wittig et al. 2008). Garnet and cpx abundances of Udachnaya peridotites correlate positively with FeO (Fig. 9) and negatively with Mg#_{WR}. These trends, for cratonic peridotites from Udachnaya and elsewhere, are interpreted here as a result of interaction between residual protoliths of garnet peridotites, produced in a range of pressures, with liquids, which precipitated cpx and garnet either during or after craton formation. These processes generally affect deformed peridotites more than coarse peridotites.

Mantle metasomatism in time and space

Peridotite xenoliths from Udachnaya, which equilibrated at <5.5 GPa (i.e. coarse garnet and spinel peridotites), tend to have higher Mg#_{OI}, Mg#_{WR} and lower HREE than predominantly deformed peridotites equilibrated at >5.5 GPa (Figs. 4b and 5a and Online Resource 4). The upper part of the lithospheric profile has better preserved the chemical compositions of its residual protoliths than the lower part as earlier inferred for the Siberian (Boyd et al. 1997; Ionov et al. 2010), Kaapvaal (Boyd and Mertzman 1987; Simon et al. 2007) and Slave cratons (Kopylova and Russell 2000; Kopylova et al. 1999). However, we find no clear correlation of melt depletion and metasomatism indices with equilibration temperature or depth for Udachnaya garnet peridotites. This contrasts with the increase in “fertility” (lower Mg#_{WR}, higher cpx) with depth documented in the Slave craton (Kopylova and Russell 2000). Despite the absence of clear relationships between the *P–T* values, deformation degrees and the magnitude of metasomatism, the majority of sheared garnet peridotites have the lowest Mg#_{WR}, highest FeO, and greatest trace element enrichments in whole-rock, cpx and garnet. This strongly suggests that the deformation may be contemporary to metasomatism.

Peridotites in this study have complex contamination-free whole-rock trace element compositions, which cannot be produced by a single process. HREE in residual garnet peridotites can be explained by high degrees of melt extraction, whereas LREE–MREE may reflect post-melting enrichments (Fig. 10). We have shown that the cpx compositions in the Udachnaya garnet peridotites are most likely linked to metasomatism. To constrain the nature of metasomatic agents that produced incompatible element enrichment of the cpx (and possibly of some garnets), we

calculated the REE contents of hypothetical melts in equilibrium with cpx in the xenoliths (Fig. 11). The calculated REE ranges in the melts are similar to those in the Udachnaya kimberlites (Kamenetsky et al. 2012).

Some previous studies (Simon et al. 2003; Stachel and Harris 1997) linked kimberlite circulation in the lithospheric mantle to crystallization of phlogopite and cpx in peridotite xenoliths following the reaction: gar + opx + fluid → phlogopite + diopside (Boyd et al. 1997; Grégoire et al. 2002; van Achterbergh et al. 2001). The majority of garnet and spinel peridotites from Udachnaya in this and related studies (Doucet et al. 2012; Ionov et al. 2010) are mica-free. Phlogopite, when present, only occurs as a late-stage phase related to kimberlite veins (Online

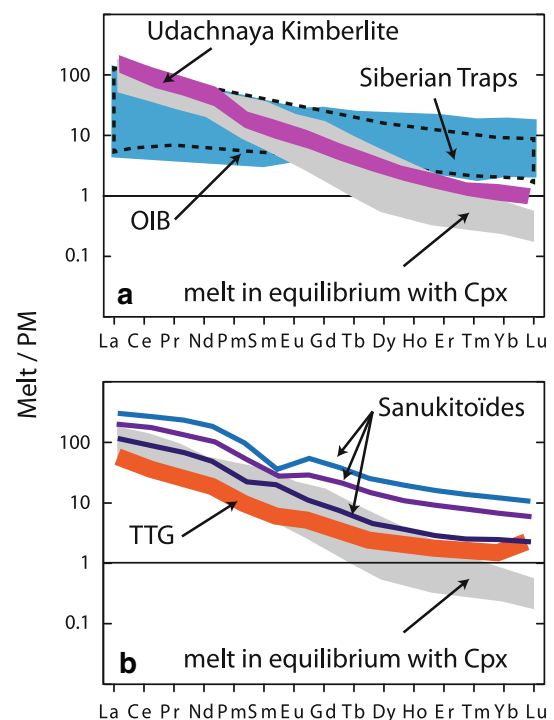


Fig. 11 Primitive mantle-normalized (McDonough and Sun 1995) REE field for hypothetical melts in equilibrium with cpx (grey field) calculated using mineral/melt partition coefficients compiled for pMELTS software (Ghiorso et al. 2002; Smith and Asimow 2005). Also shown are Udachnaya kimberlite compositions (pink field; Kamenetsky et al. 2012) (a), Siberian trap basalts (blue field; Hawkesworth and Gallaher 1993; Hawkesworth et al. 1995; Kamo et al. 2003; Lightfoot et al. 1990; Reichow et al. 2005) (a), Hawaiian ocean island basalts (fine dashed line; Clague and Moore 2002; Feigenson et al. 2003; Fekiacova et al. 2007; Huang and Frey 2003; Kimura et al. 2006; Lipman et al. 2006; Morgan et al. 2007; Pearce et al. 1999; Ren et al. 2009; Sims et al. 1999; Van Der Zander et al. 2010; Wanless et al. 2006; Xu et al. 2005) (a), South African Tonalite-Trondhjemite-Granodiorite (TTG, orange field; Martin and Moyen 2002) (b) and South African sanukitoides (MME (microgranular mafic enclaves), clear blue field; enderbite, dark purple field; mafic dykes, dark blue field; Laurent et al. 2011) (b). Udachnaya kimberlite melts appear to be the best potential metasomatic agents to account for the cpx compositions

Resource 3). The common chemical homogeneity of minerals and local disequilibrium between cpx and garnet indicate that the kimberlite veins must have been emplaced shortly before kimberlite eruption that brought up the xenoliths to the surface.

Conclusions

1. Coarse and deformed garnet peridotites in this study span a P – T range from 2.5 to 6.6 GPa and 720–1,320 °C. They range from near-pristine melting residues to Fe–Ti-enriched rocks. About ½ of coarse peridotites are similar in modal, major and trace element compositions to low-opx Udachnaya spinel harzburgites previously interpreted as melting residues. The protoliths of the garnet peridotites may have been formed by 30–38 % of anhydrous polybaric partial melting between 7 and 4 GPa and ≤ 1 –3 GPa, that is, at similar or slightly lower melting degrees than the spinel harzburgites.
2. Garnets in many coarse and transitional Udachnaya xenoliths with ≤ 5 % of modal garnet, as well as in other cratonic peridotites, may be residual phases that survived partial melting.
3. The majority of Udachnaya garnet peridotites experienced modal metasomatism, which precipitated cpx and HREE-MREE-rich garnets and produced enrichments in Fe, Ti, Al and Ca. The metasomatism can be linked to a range of silicate- and carbonate-rich liquids. The latest event took place shortly before the kimberlite eruption and was related to localized deformation.

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