### **RESEARCH ARTICLE**



## Decay of the drag wake of a sphere at Reynolds number 10<sup>5</sup>

D. Curtis Saunders<sup>1</sup> · Justen A. Britt<sup>1</sup> · Scott Wunsch<sup>1</sup>

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#### Abstract

The drag wake of a dimpled sphere with  $Re = 10^5$  is studied experimentally using Stereo Particle Image Velocimetry to a downstream distance of ~ 90 diameters. The wake growth and velocity decay are analyzed and compared with previous dimpled sphere data for  $Re = 5 \cdot 10^4$ . Self-similar decay was observed with the ensemble mean axial velocity defect decaying as  $x^{-1}$  and the wake size growing as  $x^{1/2}$ . Due to the difference in Reynolds number, the two spheres have different drag coefficients (0.13 and 0.25, respectively), but these self-similar decay exponents were not observed to depend on drag coefficient or Reynolds number. The results suggest that the self-similar drag wake decay observed at laboratory scales may extrapolate to the larger Reynolds numbers typical of engineering and geophysical flows.

### **Graphical Abstract**





D. Curtis Saunders Curtis.Saunders@jhuapl.edu

## **1** Introduction

The spreading and decay of the drag wake of an object is relevant to many applications in geophysics and engineering. Examples include the wakes of mountains, seamounts, windmills, and buildings. Drag wakes are characterized by

<sup>&</sup>lt;sup>1</sup> The Johns Hopkins University Applied Physics Laboratory, Laurel, MD, USA

their mean velocity defect as a function of distance x downstream from the generating body. For axisymmetric bodies such as spheres, disks, and slender cylindrical bodies, selfsimilar decay of the wake is typically observed (Tennekes and Lumley 1972). In this case, the mean velocity defect u (the difference between the flow speed  $U_m$  and the mean velocity component in the direction of travel x) has the functional form

$$\langle u(x,r)\rangle \simeq U_o(x)f\left(\frac{r}{L}\right)$$
 (1)

$$L \sim \left(\frac{x}{D}\right)^{\alpha} \tag{2}$$

$$U_o \sim \left(\frac{x}{D}\right)^{-2\alpha} \tag{3}$$

where x/D is the downstream distance (measured in units of the body diameter D) and r is the radial distance from the center symmetry axis of the body. In this self-similar form, the radial profile of the wake velocity defect has the same shape f(r/L) at any downstream distance when scaled by the appropriate velocity and length scales. The appropriate length scale L grows with some power  $\alpha$  of x/D, while the velocity scale  $U_o$  decays twice as rapidly, as required by momentum conservation (Tennekes and Lumley 1972).

The classical self-similar scaling law (Swain 1929; Tennekes and Lumley 1972), in which the wake spreads with downstream distance x as  $x^{1/3}$  ( $\alpha = 1/3$  in Eqs. 1-3), has long been used to describe axisymmetric drag wakes. Although early experiments with spheres appeared to confirm this result (Bevilaqua and Lykoudis 1978), most recent experiments (Bonnier and Eiff 2002; Nedić et al. 2013; Obligado et al. 2016; Saunders et al. 2020) have demonstrated that axisymmetric drag wakes have more rapid wake spreading (as  $x^{1/2}$ , or  $\alpha \simeq 1/2$  in Eqs. 1-3) and decay than the classical result of  $\alpha = 1/3$ . Experiments with a sphere at a Reynolds number ( $U_m D/v$ ) of  $Re \simeq 10^4$  first demonstrated non-classical decay to a distance of  $\simeq 30$  body diameters. Non-classical scaling was extended to  $\simeq 50$  body diameters by Nedić et al. (2013). Although (Nedić et al. 2013) initially indicated that rough shapes (fractal plates) deviated from classical scaling while more smooth shapes (disks) did not, subsequent reanalysis in Obligado et al. (2016) found that all shapes studied had  $\alpha \simeq 1/2$ . However, other experiments with disks (Johanssan et al. 2003) have yielded classical scaling. Experiments with a dimpled sphere at  $Re \simeq 5 \cdot 10^4$ (Saunders et al. 2020) also found  $\alpha \simeq 1/2$  to a downstream distance of  $\sim 200$  body diameters. High resolution numerical simulations have also reproduced the more rapid  $x^{1/2}$ wake spreading (Dairay et al. 2015; Pal et al. 2017; Chongsiripinyo and Sarkar 2020; Nidhan et al. 2020), although (Nidhan et al. 2020) suggested a transition from  $x^{1/2}$  to the slower  $x^{1/3}$  growth at some downstream distance. While some questions remain, these results collectively appear to establish that non-classical ( $\alpha \simeq 1/2$ ) wake decay occurs for most axisymmetric bodies at laboratory Reynolds numbers  $(Re \sim 10^4)$ , at least over some initial distance downstream from the body.

Extrapolation of this laboratory scaling law to larger engineering applications with Reynolds numbers exceeding 10<sup>5</sup> remains an open question, as no data are available on wake decay at these Reynolds numbers. For spheres, the drag coefficient (and hence possibly the wake) undergoes a transition at Reynolds numbers of  $\sim 10^5$ , known as the 'drag crisis' (Achenbach 1972). As previous laboratory data were obtained at Reynolds numbers below this drag crisis, wake decay data in this transitional regime would provide insight into wake behavior at larger scales. To address this question, new Stereo Particle Image Velocimetry (SPIV) measurements of the decaying wake of a dimpled sphere with a Reynolds number of  $Re = 10^5$  are compared to previous results for  $Re = 5 \cdot 10^4$ . Laboratory data from Achenbach (1974) indicate that the drag coefficient  $C_D$  for a rough sphere decreases from 0.5 to 0.1-0.2 (the 'drag crisis') for Reynolds numbers in the range of ~  $5 \cdot 10^4$  to ~  $3 \cdot 10^5$ , with the transition point and minimum drag depending on surface roughness. As a consequence of the dimples on the sphere used here, the two Reynolds numbers compared fall within the drag crisis regime, and have different drag coefficients ( $C_D \simeq 0.13$  for the new  $Re = 10^5$  data,





(a) The dimpled sphere.

(b) A schematic of the experimental layout, illustrating the location of the towed sphere, SPIV cameras, and orientation of the laser illumination sheet.

compared to  $C_D \simeq 0.25$  for  $Re = 5 \cdot 10^4$  in Saunders et al. (2020)). These new data have been obtained near the drag minimum ( $C_D \simeq 0.1$ ), and hence approach the flow regime found in engineering applications. The goal of the present experiments is to determine if this change in drag results in a measurable difference in the wake decay. This will assist in determining the applicability of the laboratory wake decay data to larger scale engineering flows.

## 2 Experimental approach

Experiments were conducted in the same tow tank as Saunders et al. (2020) using a dimpled sphere with a diameter of D = 10 cm towed at a constant speed of  $U_m \simeq 1$  m/s. The sphere is a larger version of the one used in Saunders et al. (2020), and is shown in Fig. 1a. Figure 1b presents a schematic overhead view of the experiment. The tank width was 1.8 m and the water depth was 0.7 m. The blockage ratio (the square root of the ratio of sphere cross section to tank cross section) was  $\simeq 0.08$ , well below the threshold at which drag is impacted by the tank walls (Achenbach 1974). The sphere depth was approximately midway between the tank bottom and free surface. Surface disturbances were not observed during the experiments, indicating that interactions between the wake and free surface did not occur. A laser sheet originating beneath the tank floor and oriented perpendicular to the sphere direction of travel illuminated the particles within the wake. Two cameras in submerged boxes were located within the tank, one on each side of the wake at the same depth, to image the illuminated particles.

Stereo Particle Image Velocimetry (SPIV) data were acquired immediately following the passage of the sphere until wake structures reached the edge of the field of view of the SPIV system. The components of the SPIV system included the DaVis software (LaVision), two Imager sCMOS high speed cameras (LaVision), Camera Link HS Frame Grabber (LaVision), PTU-X Programmable Timing Unit (LaVision), and Litron Nano L 200-15 Pulsed Nd:YAG Laser. In addition, each camera was equipped with a Scheimpflug mount to correct for distortion due to the cameras not being perpendicular to the image plane. For these experiments the cameras were set to an angle of  $33^{\circ}$  from the *x*-axis (see Fig. 1b) and were fitted with 20 mm focal length lenses, resulting in an overlapping FOV of 73 cm in the *y* direction and 64 cm in the *z* direction. This angle is within the optimum range for SPIV systems, as determined by Lawson and Wu (1997).

The tank was seeded with 50  $\mu$ m neutrally buoyant particles with an average density of 1.03 g/cc. The Stokes drift time scale of the particles was  $\tau_p = 68 \ \mu s$  (Raffel et al. 2018). For these experiments the timescale of the flow  $\tau_f$  is given by the ratio of the length scale to the largest velocities in the wake. Therefore the Stokes number  $S_t = \frac{\tau_p}{\tau_c}$ , which is the ratio of these two scales, was approximately  $3 \cdot 10^{-7}$  between x/D = 15 and x/D = 90. This Stokes number satisfies the requirement  $(S_t \leq 10^{-1})$  for the seeding particles to follow the turbulent motions of the flow within the range of downstream distances studied. The 50  $\mu$ m particle size was smaller than the resolution of the PIV system, indicating that pixel locking effects would be present in the data. However, this pixel size was chosen to obtain the field of view necessary to image the entire wake cross section at  $x/D \sim 90$ . Pixel-locking effects were reduced by both defocusing the cameras and by the subpixel interpolation scheme (Michaelis et al. 2016). Any remaining bias errors were further reduced by azimuthally averaging the resulting velocity fields before quantifying the mean and fluctuating velocity fields. The tank was seeded such that each interrogation region contained on average 8-10 particles. A dual-head Nd-Yag laser (Litron, 200 mJ/pulse, 532 nm wavelength) was used to illuminate the particles. This system contains two pulsed lasers which were aligned such that their optical paths were nearly identical. The laser sheet thickness was approximately 3 cm, ensuring the out-of-plane displacement was less than one-quarter of the light sheet thickness. The

 Table 1
 SPIV parameters

Re	Laser pulse rate (Hz)	Image pair time spacing $\delta t$ (ms)	Min in plane velocity (cm/s)	Max in plane velocity (cm/s)	Uncer- tainty (cm/s)	No. of runs
$ \frac{1 \cdot 10^5}{\text{(Present data)}} $	15	10	1	16	0.6	65
5 · 10 <sup>4</sup> (Saun- ders et al. 2020)	10	10	0.8	13	0.2	42

 $Re = 1 \cdot 10^5$  corresponds to present study. Previous paremeters used in Saunders et al. (2020) are included as well corresponding to  $Re = 5 \cdot 10^4$ 



Fig. 2 Wake ensemble mean velocity components (u, v, and w) and fluctuations q at x/D = 25

lasers can be pulsed simultaneously or with a time delay between pulses. This time delay  $\delta t$ , combined with the desired spatial resolution and camera FOV, determines the range of flow velocity scales which can be measured.

Table 1 details the parameters used for the SPIV measurements. The SPIV processing was performed using an open-source MATLAB code called UVMAT (xxxx). The algorithm first performs PIV processing of the images from a single camera using a dual-pass method, where the first pass determines lower-resolution estimates of the velocity field within an interrogation region, while the second pass uses the estimates from the first pass to obtain more precise measurements of the velocity field at a higher resolution. The velocity fields from the two cameras are then combined using SPIV to determine all three velocity components within the plane of data. The first-pass PIV analysis was performed with  $128 \times 128$  pixel interrogation regions, a search-box size of  $147 \times 147$  pixels, a shift of (0,0) pixels, and a 50 % overlap of the interrogation regions. The secondpass analysis was performed with interrogation regions of 64  $\times$  64 pixels and a 50 % overlap, resulting in a velocity vector every 32 pixels in the y and z directions. These window sizes ensured that particle displacements were limited to 1/4 of the interrogation window, corresponding to a maximum in-plane velocity of 26 cm/s. Subpixel accuracy was obtained using the thin plate spline method. This is a multi-dimensional generalization of spline interpolation and is an optimum way

to interpolate data with minimal curvature of the interpolating function (UVMAT xxxx).

While the Stokes numbers indicated the particles would accurately follow turbulent motions within the fluid, the errors in detecting particle positions and ultimately fluid velocities were also quantified. Given the size of the particles ( $\leq 1$  pixel), we estimate a measurement uncertainty on the order of 0.3 pixels (Raffel et al. 2018). This corresponds to a velocity error of up to 0.6 cm/s. Typical particle displacements were from 3 to 6 pixels, corresponding to wake in-plane velocities from 7 to 11 cm/s. Additionally, the typical shear across the interrogation window was less than 0.02 pixels. Displacement estimates with correlations above 50% were kept and the remaining values were discarded.

Sixty-five runs were performed so that the results could be combined to produce estimates of the ensemble mean and fluctuating velocity fields. The SPIV settings for the present  $Re = 10^5$  experiments were sufficient to collect accurate velocity data in the downstream range of  $15 \le x/D \le 90$ . The prior experiments of Saunders et al. (2020) also yielded acceptable velocity data in this range, allowing direct comparison of the two Reynolds numbers. To obtain the mean axial velocity, the measured axial velocity fields from each experiment at a given x/D were averaged. The velocity fluctuation field is defined as the standard deviation of the velocity components over the ensemble of experiments at each x/D. The tow speed varied by less than 1 cm/s between runs so all SPIV velocities were normalized by the nominal tow





Fig. 3 Individual realizations of the axial velocity defect u at x/D = 25



Z/D

-3

-2

0

y/D

2

Fig. 4 Ensemble mean axial velocity defect  $\langle u \rangle / U_m$  at 4 down-track locations

0



**Fig. 5** Ensemble mean axial velocity defect  $U_m$  as a function of radius

speed of 1 m/s. There is also a variation in the depth and horizontal position of the wake center from run to run, due to slightly nonzero drift velocities which vary from run to

run. This effect was mitigated by averaging the runs.

# 3 Results

r/D at 5 down-track locations

Figure 2 presents the SPIV ensemble mean velocity field (averaged over 65 runs) measured at x/D = 25. The upper left panel shows the mean axial velocity defect *u* normalized by the sphere tow speed  $U_m$ . The wake appears as the yellow region in the center of the panel. This view demonstrates the approximately axisymmetric shape of the mean axial velocity defect. The upper right and lower left panels present the mean horizontal and vertical velocities, respectively. These components are an order of magnitude smaller than the mean axial velocities for the *v* and *w* components (~ 3 mm/s) at x/D = 25 are smaller than the SPIV accuracy for an individual run. Averaging the 65 runs reduces the effective error of the ensemble mean by a factor of  $1/\sqrt{65} \sim 1/8$ , for an error of ~ 0.8 mm/s in the ensemble mean.

The lower right panel of Fig. 2 presents the root-meansquare fluctuation velocity q for the ensemble of runs. Here q is defined as

$$q^{2} \equiv \frac{1}{2} \langle (u - \langle u \rangle)^{2} + (v - \langle v \rangle)^{2} + (w - \langle w \rangle)^{2} \rangle$$
(4)

At this x/D, the magnitude of these fluctuations is comparable to the mean axial velocity. The axial velocity defects of four individual runs at x/D = 25 are illustrated in Fig. 3.



**Fig. 6** Self-similar profiles of the mean axial velocity defect  $U_o$  as a function of radius  $r/L_u$  at the same locations as Fig. 5, using the similarity scales computed from Eqs. 7 and 8

There are large variations in axial velocity defect from run to run, and individual runs do not appear axisymmetric. As a result, no individual wake realization closely resembled the mean flow field. The variations from run to run are responsible for the magnitude of q seen in the lower right panel of Fig. 2. These variations also necessitate the large number of experimental runs (65) conducted to obtain quality ensemble-mean data.

The decay of the mean axial velocity defect is illustrated in Fig. 4, which presents the wake defect at four x/D locations from 15 to 90. The wake grows in size and its peak velocity decays with distance, as expected (note changing color scales). Observed deviations from axisymmetry tend to increase with downtrack distance, but remain modest even at x/D = 90. This asymmetry is attributed to a relative increase in wake fluctuations at greater x/D, requiring more runs to be averaged to obtain a quality estimate of the mean.

To explore self-similar decay, the wake size and mean axial velocity defect scales are characterized as a function of x/D using the same approach as Saunders et al. (2020). The SPIV ensemble mean axial velocity data in a small range of x/D are binned into radial segments (based on distance from the sphere center) and averaged over azimuth to produce profiles of  $\langle u(x, r) \rangle$ . This azimuthal averaging procedure exploits the expected axisymmetry of the wake. Here, radial bins of 0.1*D* and x/D bins of  $\pm 0.1x/D$  were used. Radial profiles of  $\langle u \rangle$  at 5 locations are presented in Fig. 5, demonstrating the spreading of the wake mean axial velocity defect as x/D increases.



**Fig. 7** Mean axial velocity decay (colored symbols) as a function of x/D, along with best-fit scaling exponents (black lines). Present  $Re = 10^5$  data are shown in red;  $Re = 5 \cdot 10^4$  data from Saunders et al. (2020) are shown in blue. Each panel presents a different method for

estimating the velocity scale: *Left:* Moments (Eqs. 5 and 6) *Center:* Gaussian fit *Right:* Maximum of the ensemble mean axial velocity defect

From the profiles in Fig. 5, the first two moments are computed at each x/D:

$$M_o(x) \equiv \int_0^\infty \langle u(x,r) \rangle r dr$$
(5)

$$M_2(x) \equiv \int_0^\infty \langle u(x,r) \rangle r^3 dr \tag{6}$$

In practice, the integration range is limited to r < 2.5D, to avoid contamination by measurement noise far from the wake center. Using these moments, the wake size  $L_u$  and mean axial velocity scale  $U_o$  are computed from

$$L_u^2 \equiv \frac{M_2}{M_o} \tag{7}$$

$$U_o \equiv \frac{M_o^2}{M_2} \tag{8}$$

These represent the typical length and velocity scales of the wake, based on the measured profiles in Fig. 5. This approach makes no assumption regarding the shape of the mean axial velocity defect profile. Figure 6 presents the mean axial velocity defect profiles from Fig. 5 normalized by the length scale  $L_u(x)$  and velocity scale  $U_o(x)$  computed using Eqs. 7 and 8. The results demonstrate that the mean axial velocity defect profile is approximately self-similar





when normalized by these scales. Also shown in Fig. 6 is a Gaussian fit to the profile (black line), which is found to be a reasonable approximation to the profile shape to  $r \simeq 2L_u$ .

The decay of the mean axial velocity defect scale  $U_{a}$ with x/D is presented in Fig. 7. The data were binned in distance x/D using a logarithmic scale with a bin width of  $\log_{10}(x/D) = 0.05$ . Logarithmic scales are used in Fig. 7 so the power law decay (Eq. 3) yields a straight line. Data from the present  $Re = 10^5$  experiments are shown in red. The panels present three independent measures of the velocity scale, all of which are expected to decay with the same exponent due to self-similarity. The left panel presents the result of Eq. 8, as in Saunders et al. (2020). This is the velocity scale that produced the self-similar collapse in Fig. 6. The center panel presents the velocity magnitude for a Gaussian fit to the profile at each x/D (as in Fig. 5), and the right panel shows the maximum of the ensemble-mean velocity defect at each x/D, as in Bonnier and Eiff (2002). The error bars in Fig. 7 represent one standard deviation of values within each logarithmic bin. The best-fit power law for each measure of the velocity scale is shown as a black dash-dot line. All three measures yield a best-fit decay exponent  $-2\alpha$  (Eq. 3) of approximately -1. Also shown in blue in each panel are the  $Re = 5 \cdot 10^4$  data from the previous dimpled sphere experiments of Saunders et al. (2020) for the same range of x/D. The wake ensemble mean axial velocity defect data for the two different values of Re closely align, with similar decay exponents of  $\sim -1$  for both the moments (left panel) and Gaussian fit (center panel) methods. The maximum of the mean velocity defect (right panel) for the previous  $Re = 5 \cdot 10^4$  data (blue) yielded a slower decay (exponent  $\sim -0.9$ ) but also exhibited greater scatter. However, the full



**Fig.9** Self-similar profiles of the mean axial velocity defect  $< EmphasisType = \varepsilon Underline\varepsilon > /U_o$  as a function of radius  $r/L_u$ 

data set in Saunders et al. (2020) for maximum mean velocity defect (covering a greater range of x/D) yielded a best-fit exponent of  $\approx -1$ , consistent with the other methods. The agreement between the two Reynolds numbers suggests that the observed wake mean axial velocity decay exponent of -1 may not change as *Re* increases from laboratory to engineering scales.

Spheres with different drag coefficients generate wakes with different total momentum fluxes, and this must be reflected in the mean wake velocity defect and/or length scales. The total momentum flux defect in the wake (found by integrating the mean velocity defect at any x/D and multiplying by  $U_m$ ) is proportional to  $U_m^2 M_o$ , and the drag coefficient can be estimated from the integrated mean axial velocity defect using

$$C_D = \frac{16M_o}{U_m D^2} = 16 \left(\frac{U_o}{U_m}\right) \left(\frac{L_u}{D}\right)^2 \tag{9}$$

The drag coefficient has been estimated here from the mean axial velocity data for  $15 \le x/D \le 30$ . For the present data with  $Re = 10^5$ , the result is  $C_D = 0.13 \pm 0.01$ , while the corresponding result for the  $Re = 5 \cdot 10^4$  dimpled sphere data of Saunders et al. (2020) is  $C_D = 0.25 \pm 0.01$ . This reduction in drag coefficient with Reynolds number is consistent with the 'drag crisis' for a rough sphere seen in the laboratory data of Achenbach (1974). The difference in drag coefficients between the two Reynolds numbers requires that the product  $U_{\rho}L_{\mu}^{2}$  for the two cases differ. The new  $Re = 10^{5}$  data in Fig. 7 have a velocity scale  $U_o/U_m$  that is 10–15 percent less than the previous  $Re = 5 \cdot 10^4$  values for a given x/D. Because of this small difference between the two values of  $U_a/U_m$ , it is expected that the length scale  $L_u/D$  must be larger for the larger  $C_D$  case, due to the greater total momentum of the wake.

Figure 8 presents the length scale  $L_{\mu}$  (Eq. 2) of the wake mean axial velocity defect as a function of x/D as computed from the mean axial velocity defect profiles in Fig. 5. The left panel presents the result from the moments (Eq. 7) of the profiles yielding the self-similar collapse in Fig. 6, while the right panel presents the result of a Gaussian fit to the profiles in Fig. 5. Present results for  $Re = 10^5$  are again shown in red, while prior results for  $Re = 5 \cdot 10^4$  are in blue (Saunders et al. 2020). The vertical axis is shifted by  $\log_{10}\sqrt{C_D}$  to approximately compensate for the difference in drag between the two cases, as suggested by Eq. 9 with  $U_o/U_m$  being approximately independent of  $C_D$ . The error bars represent one standard deviation of values within each logarithmic bin. The wake growth exponent is  $\alpha \simeq 1/2$ , as expected for consistency with the velocity decay exponent of  $-2\alpha \simeq -1$ in Fig. 7. The two Reynolds numbers have the same growth exponent, but the wake is smaller for the larger Re, lower



**Fig. 10** Velocity fluctuations  $q/U_m$  at 4 down-track locations



**Fig. 11** Self-similar profiles of the velocity fluctuation  $q/Q_o$  as a function of radius  $r/L_q$  at 5 down-track locations

drag case, with the reduction factor given approximately by the square root of the drag coefficient ratio.

Some previous studies (Johanssan et al. 2003; Nedić et al. 2013; Obligado et al. 2016) have utilized a virtual origin in x/D for the estimation of the wake growth and velocity decay exponents, while others (Bonnier and Eiff 2002; Saunders



et al. 2020) have not. The method of Obligado et al. (2016) has been applied to both data sets for  $U_o$  and  $L_u$  (not shown). In all cases, virtual origins were found to be less than 1.5D in magnitude, and the best-fit exponents within 0.05 of the values shown here. The modest impact of the virtual origin is attributed to the starting point being downstream (larger x/D) from the starting point of previous studies (Johanssan et al. 2003; Nedić et al. 2013; Obligado et al. 2016).

Figure 9 compares the mean axial velocity defect shape functions  $f(r/L_u)$  from Eq. 1 for the two Reynolds numbers. Here the velocity and radial distance are scaled by  $U_o$  and  $L_u$ , respectively. The shape function shown in red is an average of the shape functions at different x/D from Fig. 6 for  $Re = 10^5$ . Blue presents the corresponding average shape function from Saunders et al. (2020). The larger Reynolds number case appears to have a modestly larger shape function for  $r/L_u$  between ~ 0.5 and ~ 2, but the error bars overlap so this observation is not conclusive.

Other wake characteristics, such as velocity fluctuations and Reynolds stress, are also expected to demonstrate selfsimilar decay. The fluctuation q (Eq. 4 and Fig. 2, lower right panel) from the ensemble-mean wake velocity defect should follow a self-similar decay:

$$q(x,r) \simeq Q_o(x)g\left(\frac{r}{L_q}\right) \tag{10}$$



Fig. 12 Velocity fluctuation  $(Q_a)$  decay (colored symbols) as a function of x/D, along with best-fit decay exponents (black lines)





where  $Q_o$  scales with some power of x/D. The spreading and decay of the velocity fluctuations q at four x/D locations are presented in Fig. 10. Like the mean axial velocity defect (Fig. 4), the q field spreads and decays with downtrack distance (note changing color scales). Quality statistics for this quantity require more runs than are needed for the mean axial velocity defect, and the number of runs required increases with x/D. The present data yield approximately axisymmetric results for q to  $x/D \sim 60$ , and as can be seen in the lower panels of Fig. 10, symmetry is imperfect for the larger values of x/D shown. Self-similar decay of velocity fluctuations q is studied here only within this more limited range of x/D.

The characteristic scales of the velocity fluctuations q are computed as in Saunders et al. (2020) by replacing  $\langle u \rangle$  with q in the moment equations (Eqs. 5 and 6) to yield a second length scale  $L_q^2$  and velocity scale  $Q_o$ . Figure 11 presents the self-similar fluctuation profiles  $q(r/L_q)/Q_o$  for five values of x/D. Error bars, indicating the standard deviation in q within each data bin, are larger than in Fig. 6 due to the larger variability in this quantity than in the mean axial velocity defect. Although consistent with self-similarity, the larger error bars make it difficult to discern with confidence how well the profiles at different x/D collapse.

The decay of the fluctuation velocity scale  $Q_o$  is presented in the left panel of Fig. 12. In addition, the maximum fluctuation velocity in each x/D bin is included as  $Q_{max}$  in the right panel. For the present data (shown in red), the fluctuation velocity scale  $Q_o$  decays more slowly with x/D than  $U_o$ , with a decay exponent of  $\sim -0.4$ . Again, two differing approaches (the moment calculation and the maximum of q) are used to confirm the scaling, but the Gaussian fit is not used here as the shape profile is not Gaussian (see Fig. 11). Also shown in blue in Fig. 12 are previous results for  $Re = 5 \cdot 10^4$  from Saunders et al. (2020). The best-fit exponents for the two Reynolds numbers are both  $\simeq -0.4$ , and the difference in the magnitude of the velocity fluctuations between the two experiments is less than ten percent at each x/D.

A key observation regarding the fluctuation scale  $Q_o$  is that it decays much more slowly with x/D than the mean axial velocity defect scale  $U_o$ . This is illustrated in Fig. 13, which presents the ratio of the two velocity scales. As x/Dincreases, the wake is increasingly dominated by fluctuations. By  $x/D \sim 60$ , the fluctuations are nearly an order of magnitude larger than the mean axial velocity defect. This is true for both Reynolds numbers tested.



Fig. 14 Wake growth as a function of *x/D* based on velocity fluctuations, along with best-fit growth exponents

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Figure 14 presents the length scale  $L_a$ , along with its best-fit scaling exponent. The vertical axis shows  $L_a/\sqrt{C_D}$ for consistency with the presentation of  $L_{\mu}$  in Fig. 8. For self-similar wake decay it is expected that the length scales  $L_{\mu}$  and  $L_{a}$  for mean and fluctuating velocity are proportional. The velocity fluctuation scale  $L_a$  increases with x/Dwith an exponent of  $0.57 \pm 0.04$ , while  $L_{\mu}$  (Fig. 8) has a similar value of  $0.55 \pm 0.03$ . Self-similarity requires that these two exponents are equal, as there should be only one length scale (up to a constant of proportionality), and the data are consistent with this expectation. The fluctuation length scale  $L_a$  is typically about 60 percent larger than  $L_{\mu}$ , but the growth exponents are comparable. Figure 14 also shows the comparable result for  $Re = 5 \cdot 10^4$  in blue. These data have a similar growth exponent but a length scale (normalized by diameter) that is nearly twice as large as the  $Re = 10^5$  case. The smaller range of x/D shown for the  $Re = 5 \cdot 10^4$  case reflects the limitation of the smaller SPIV field of view in the earlier experiments.

The Reynolds stress is another wake characteristic expected to demonstrate self-similar decay. Data were analyzed for the radial component of the Reynolds stress  $R_{xr}$ , defined as

$$R_{xr} \equiv \langle uv_r \rangle - \langle u \rangle \langle v_r \rangle \tag{11}$$

where  $v_r$  is the radial velocity and  $\langle ... \rangle$  again denotes an average over the ensemble of experimental runs. For self-similar wake decay, this component of Reynolds stress can be expressed in terms of a magnitude  $R_o$  and a length scale  $L_R$  as

**Fig. 15** The radial component of the Reynolds stress. *Upper left*: Measured Reynolds stress at  $x/D \sim 36$ . *Upper right*: Selfsimilar profiles of Reynolds stress at four x/D. *Lower left*: The length scale  $L_R$  used to normalize the self-similar profiles, along with the length scale for the mean velocity defect  $L_u$ . *Lower right*: The magnitude  $R_o$ used to normalize the self-similar profiles, along with a line with exponent -1.42



 Table 2
 Exponent comparison

Quantity	Size $(L_u)$	Mean velocity $(U_o)$	Velocity fluctuations $(q)$
$Re = 10^5$ (present)	$0.55 \pm 0.07$	$-0.97 \pm 0.04$	$-0.38 \pm 0.03$
$Re = 5 \cdot 10^4$ (Saunders et al. 2020)	$0.57 \pm 0.07$	$-0.98 \pm 0.05$	$-0.43 \pm 0.05$

$$R_{xr}(x,r) \simeq R_o(x)h\left(\frac{r}{L_R}\right).$$
(12)

This is analogous to Eq. 1 for the mean velocity defect and Eq. 10 for the velocity fluctuations. This stress component has previously been shown to decay self-similarly to x/D = 50 in lower Reynolds number experimental (Nedić et al. 2013) and numerical (Obligado et al. 2016) studies of disks and fractal shapes, but the  $Re = 5 \cdot 10^4$  sphere experiments of Saunders et al. (2020) did not include a sufficient number of runs to obtain quality statistics on this quantity.

Figure 15 presents data from the present  $Re = 10^5$  experiments on the decay of the Reynolds stress component  $R_{xr}$ . The upper left panel presents  $R_{xr}$  at x/D = 36, averaged over a range of  $\pm 0.1x/D$ . The azimuthal symmetry is apparent, although imperfect as more data would likely be needed for better statistical convergence. Deviations from azimuthal symmetry were observed to increase with x/D. The stress is zero at the center of the wake, and has a maximum value at a radius of  $\simeq 1.5D$ . The maximum value is  $\simeq 10^{-4}U_m^2$ , approximately a quarter of the velocity fluctuations  $q^2$  at this x/D. The magnitude is smaller than  $q^2$  because the radial velocity  $v_r$  is smaller than the axial velocity (see Fig. 2), which also makes  $R_{xr}$  a more challenging measurement than  $q^2$ .

The upper right panel of Fig. 15 presents azimuthally averaged self-similar profiles for  $R_{yr}(r)$  at four x/D locations. As with the previously analyzed quantities, data are binned in x/D with a bin of  $\pm 0.1x/D$  to improve the statistics. The magnitude  $R_o$  and length scale  $L_R$  at each x/Dwere computed from the azimuthally averaged profiles using the moment method (Eqs. 5 and 6) analogous to the other quantities presented. The Reynolds stress from these locations approximately collapse, as expected. The shape is similar to that reported in Nedić et al. (2013); Obligado et al. (2016), with an off-center peak at  $r \sim 0.8L_{R}$ . The length scales  $L_R$  and magnitudes  $R_o$  computed for the self-similar profiles are shown in the lower panels of Fig. 15. The length scale  $L_R$  (lower left) is expected to be proportional to  $L_u$  and  $L_a$ . The dashed line, corresponding to  $2L_u$ , shows that the computed values of  $L_R$  are consistent with self-similarity. According to self-similar theory (Nedić et al. 2013; Obligado et al. 2016; Saunders et al. 2020), the magnitude  $R_{a}$ should obey  $R_o \sim U_o U_m dL_u/dx$ . For the measured best-fit scaling exponents of  $U_{\rho}$  (-0.97) and  $L_{\mu}$  (0.55), this yields a scaling exponent of -1.42. The black dashed line in the lower right panel of Fig. 15 shows that the computed values of  $R_o$  are consistent with this exponent. Best-fit values of the scaling exponents for  $L_R$  and  $R_o$  were not computed due to the relatively small range of x/D for which adequate statistics for  $R_{xr}$  were obtained. The data presented for  $R_{xr}$  are consistent with self-similar theory, but are not sufficient to show that the Reynolds stress obeys self-similar theory for the full range of data collected (to  $x/D \simeq 90$ ).

## **4** Discussion

This is the first experimental study of self-similar drag wake decay at a Reynolds number (based on diameter) of  $Re = 10^5$ extending to a downstream distance of x/D = 90. The mean axial velocity defect decayed with an exponent of  $\simeq -1$  while the wake size grew with exponent  $\simeq 0.5$ . The results are consistent with prior data for a sphere at  $Re = 5 \cdot 10^4$  (Saunders et al. 2020) as well as recent results from other axisymmetric shapes (Bonnier and Eiff 2002; Nedić et al. 2013; Obligado et al. 2016) at lower Re, but differ from the long-standing and previously accepted classical self-similar scaling exponents of -2/3 and 1/3, respectively (Swain 1929; Tennekes and Lumley 1972). Table 2 summarizes the wake scaling exponents found in the present study compared to previous experiments at a lower Reynolds number. Scaling exponents for the wake size, mean axial velocity magnitude, and velocity fluctuation magnitude are all identical to within error estimates for both Re. The similarity in wake decay between the two Reynolds numbers occurred despite a difference in drag coefficients of nearly a factor of two, owing to the flow transition associated with the 'drag crisis' at these Reynolds numbers (Achenbach 1974). This consistency in observed wake decay exponents despite the change in flow regime suggests that these laboratory results may pertain to larger Reynolds number engineering flows. (Note that the exponents for  $Re = 5 \cdot 10^4$  in Table 2 differ slightly from those reported in Saunders et al. (2020) because a larger range of x/D, obtained using multiple SPIV settings, was used. Here the fitting range of x/D was restricted to match the present  $Re = 10^5$  data, so that the best-fit exponents would reflect the same x/D range.)

Another observation common to both Reynolds numbers is the increasing importance of fluctuations relative to the ensemble mean as x/D increases. The fluctuation velocity scale decreases with an exponent of  $\simeq -0.4$ , less than half of the decay rate of the mean axial velocity defect. The ratio of fluctuations to the ensemble mean therefore grows with x/D. Individual realizations of the wake appear less and less like the ensemble mean as x/D increases.

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