

Improved laser absorption spectroscopy measurements of fame temperature via a collisional line-mixing model for CO₂ spectra **near 4.17 µm**

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Abstract

An experimental and theoretical study of CO_2 absorption spectroscopy near 4.172 μ m was conducted to acquire measurements of gas temperature and $CO₂$ concentration in flames at near-atmospheric pressure. Scanned-wavelength laser absorption spectroscopy measurements were obtained in laminar non-premixed fames produced using either a Hencken burner or inside a fxed-volume combustion vessel. For the Hencken burner fame, the non-uniform distribution of gas conditions along the line-of-sight was estimated from computational fuid dynamics (CFD) simulations and accounted for in simulations of path-integrated absorbance spectra. Gas properties were inferred from measured absorbance spectra using two models: (1) an isolated line model employing the Voigt Profle and (2) a collisional line-mixing model. Regarding the latter, the relaxation matrix for $CO₂$ in the line-mixing model was modified and scaled based on an empirically derived collisional-relaxation matrix for $CO₂$ –Ar. The collisional line-mixing model reduced the residuals of the best-fit spectra by approximately a factor of 4. In addition, the fame temperature obtained from the line-mixing model agreed well with that predicted by adiabatic equilibrium calculations and was typically 10% more accurate than those produced by the isolated line model, thereby illustrating the importance of accounting for collisional line mixing at the conditions and wavelengths studied despite the modest gas pressure and high temperature.

1 Introduction

Carbon dioxide $(CO₂)$ is a major product of hydrocarbon fuels and is, therefore, an attractive candidate for providing temperature measurements in combustion environments using laser absorption spectroscopy. Previously, $CO₂$ absorption measurements were primarily conducted using transitions in the near-infrared. This includes measurements in the $(0,0^0,0 \rightarrow 2,2^0,1)$ vibrational band near 1.5 μ m [\[1](#page-9-0)], as well as in the $(0,0^0,0 \rightarrow 1,2^0,1)$ vibrational band near 2.0 μ m [\[2–](#page-9-1)[4\]](#page-9-2). The advent of tunable diode lasers in the extended near-infrared (2.7 μm) has enabled absorption measurements with signifcantly larger signal levels, for example,

measurements in the $(0,0^0,0 \rightarrow 1,0^0,1)$ and $(0,0^0,0 \rightarrow 0,2^0,1)$ combination bands have demonstrated much better signalto-noise ratio, especially for high-temperature combustion applications $[5-7]$ $[5-7]$ $[5-7]$. Perhaps more significantly, the development of interband cascade lasers (ICL) and quantum cascade lasers (QCL) has provided access to even stronger absorption transitions in the mid-infrared. For example, the R(76) transition in the $(0,0^0,0 \rightarrow 0,0^0,1)$ fundamental band transition near 4.3 μ m was used for sensitive CO₂ concentration measurements during thermal decomposition of C3–C5 ethyl esters by Ren et al. [\[8](#page-9-5)] using an external cavity QCL. In addition, Spearrin et al. measured the concentration of $CO₂$ in a scramjet combustor at the same wavelength [\[9](#page-9-6)].

More recently, absorption measurements near the $(0,0^0,0 \rightarrow 0,0^0,1)$ bandhead at 4.172 µm have become popular due to the good signal-to-noise ratio, high temperature sensitivity and isolation from interfering absorption lines of, for example, $H₂O$ [[10](#page-9-7)]. Temperature measurements in flames at the $(0,0^0,0 \rightarrow 0,0^0,1)$ bandhead were first demonstrated by Villareal et al. [\[11\]](#page-10-0) on a McKenna burner and was latter demonstrated in several diferent combustion

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facilities $[12-14]$ $[12-14]$ $[12-14]$. Accurate flame thermometry was also obtained by scanning across multiple absorption transitions within the $(0,0^0,0 \rightarrow 0,0^0,1)$ band [[10](#page-9-7), [14](#page-10-2)[–16\]](#page-10-3) or by combining simultaneous measurements in the $(0,0^0,0 \rightarrow 0,0^0,1)$ band and the $(0,0^0,0 \rightarrow 0,2^0,1)$ combination band [\[17](#page-10-4)]. In many of the aforementioned studies [[12](#page-10-1), [13\]](#page-10-5), however, simulated spectra performed using an isolated line model and the HITEMP [[18\]](#page-10-6) database did not produce satisfactory agreement with the experimentally measured spectra, and gas-cell experiments were needed to correct the line positions and line intensities. Some authors pointed out that the possible cause for poor agreement could be collisional line-mixing effects $[13]$ $[13]$.

Collisional line-mixing (LM) results from collisioninduced transfer of rotational state populations and it can cause absorption to shift from the wings to the center of a vibrational band. Collisional line mixing is most signifcant when the collisional linewidth is comparable to the spacing of the rotational lines. This occurs when the gas number density or collision frequency is high, and/or when the rotational lines are tightly spaced, for example, near bandheads. Conventionally, collisional line-mixing efects are often only considered important at high pressures. In this work, we show that collisional line mixing has a signifcant impact on the absorbance spectrum of CO_2 's v_3 fundamental band near the bandhead in atmospheric pressure fames.

For $CO₂$, many studies have been devoted to developing databases for the collisional-broadening coefficients among diferent collisional partners while accounting for collisional line mixing. For example, those gaseous species include, but are not limited to O_2 , CO_2 , N_2 [[19–](#page-10-7)[21\]](#page-10-8) as well as by noble gases like Ar $[21-25]$ $[21-25]$. Among those studies, Brownsword et al. [[23](#page-10-10)] concluded that the *J*-dependent collisional linewidths for CO_2 –Ar and CO_2 –N₂ differed only by a scalar and the dependence of the collisional linewidth on the vibrational band was negligible. This indicates that the same set of collisional-broadening coefficients can be applied to both Raman and absorption spectroscopy calculations for diferent bands. On the other hand, many studies have been devoted to developing semi-empirical scaling laws $[24, 26-30]$ $[24, 26-30]$ $[24, 26-30]$ $[24, 26-30]$ or by performing ab initio calculations $[31-35]$ $[31-35]$ to compute the state-to-state collision rate and, thus, the collisional-broadening coefficient for $CO₂-CO₂$ collisions or for $CO₂$ –Ar collisions. There are also some databases that can provide necessary parameters for $CO₂$ absorbance calculations that account for collisional line mixing [[36](#page-10-16)].

However, direct experimental measurements and validation of $CO₂$ line-mixing parameters at combustion relevant conditions are rare. Recently, Lee et al. [[37,](#page-10-17) [38](#page-10-18)] frst reported measurements of the $CO₂–Ar$ collisional linewidth for absorption transitions near 4.172 μm which correspond to rotational levels $J=99-145$. The authors acquired such measurements in a shock tube in $CO₂-Ar$ mixtures at

temperatures up to 3000 K. They also empirically derived the scaling coefficients for the Modified exponential Energy Gap (MEG) law to calculate the *J*-dependent CO_2 -Ar stateto-state collision rate. Their model covered all the major absorption lines at and near the bandhead and showed excellent agreement between measured and modeled spectra at pressures up to 58 atm. In addition, they modifed the collisional line-mixing model extracted from the $CO₂–Ar$ measurements by a scaling constant to apply it to more practical situations such as inside a rotating detonation engine [\[39](#page-10-19)]. The spectral-ftting results were satisfactory and were greatly improved over those produced using an isolated line model with the Voigt profle. However, the gaseous compositions in practical combustion applications are usually more complex than in a shock tube. In addition, it is difficult to rigorously evaluate the accuracy of temperature measurements obtained in detonation engines.

In this work, we demonstrate that accounting for collisional line mixing not only improves residuals between measured and modeled $CO₂$ absorbance spectra near 4.172 μm but also gives more accurate thermometry measurements in several near-atmospheric pressure fames. The experimental measurements were performed in a laminar non-premixed fame established over a Hencken burner as well as a spherically expanding fame established inside a fxed-volume combustion vessel. The measured fame temperatures in both of those two test facilities were close to their adiabatic values. For Hencken burner fames, the efects of non-uniform distribution of fame properties were accounted through CFD simulations. A Boltzmann profle $[16, 40]$ $[16, 40]$ $[16, 40]$ $[16, 40]$ $[16, 40]$ was assumed for flame temperature and $CO₂$ concentration distributions along the line-of-sight in calculating path-integrated absorbance spectra. A collisional linemixing model was implemented using the G-matrix method. The collisional-relaxation matrix was adjusted by scaling the empirically derived collision rate for $CO₂$ –Ar model [[37\]](#page-10-17) with a scalar that was inferred from the measured spectra using a similar, but diferent approach to Nair et al. [[39](#page-10-19)]. The results illustrate that this approach provides improved thermometry measurements even when the gaseous mixtures are not known a priori. The validity of this approach is discussed in detail later.

2 Experimental setup

2.1 Optical equipment

Figure [1](#page-2-0) shows a schematic of the experimental setup. A continuous-wave distributed-feedback (DFB) interbandcascade laser (ICL) with a center wavelength of 4.172 μm (Nanoplus) was used for all measurements. The function generator (Tektronix AFG3022C) was used to generate

Fig. 1 Schematic of the experimental setup used to characterize burner flames with laser-absorption measurements of CO₂ near 4.172 μm

10 kHz triangle waveforms that were input into the laser diode controller (Stanford Research System, LDTC0520) to scan the injection current of the ICL. A $CaF₂$ beamsplitter was used to create a second beam which was passed through a Germanium etalon to characterize how the relative frequency of the laser varied throughout each scan, the absolute frequencies, on the other hand, could be determined by comparing the absorption features to numerical simulations. The free spectral range (FSR) of the etalon used in this work is 0.012 cm^{-1} . The laser intensity was measured by two thermoelectrically cooled mercury cadmium telluride (MCT) photovoltaic detectors (PVI-3TE, Vigo System).

3 Combustion equipment

Flame temperature measurements were acquired in both laminar non-premixed flames and laminar premixed fames. The laminar non-premixed fame was produced using a Hencken burner, and previous studies have shown that with sufficiently high flow rate, the heat loss to the burner is insignifcant and the fame temperature at 0.8–3 cm above the burner surface is approximately the adiabatic flame temperature for H_2 -Air and CH₄-Air flames [[41](#page-10-21), [42](#page-10-22)]. Figure [2](#page-2-1) shows a schematic of the homemade

Fig. 2 Schematic of the Hencken burner used to produce laminar non-premixed fames

Hencken burner. The laser beam was passed 2 cm above the burner surface where the fames were stable and the fame temperature was near equal to the adiabatic fame temperature.

A CFD simulation was performed with the kinetic model developed by Qi et al. [\[43\]](#page-10-23) for DME–air fames to estimate and account for the impact of the non-uniform distributions of flame temperature and $CO₂$ concentration on measurement accuracy. The air flow rate was set as 15 standard liter per minute (SLM), while the co-flow N_2 flow rate was set as 10 SLM to be consistent with the experimental condition. The fuel and the oxidizer are sufficiently mixed within less than one millimeter above the burner surface, and the regime of interest in this work is at 2 cm above the burner. Thus, we simplifed the dense distributions of the fuel tubes and air tubes and treated the burner geometry as a 25.4 mm diameter tube where the reactants are fully premixed. This treatment was also used in some previous numerical studies of Hencken burner fame structure, and good agreement between the simulation and the experimental measurements were reported [[44,](#page-10-24) [45](#page-11-0)]. An example of two-dimensional temperature and $CO₂$ mole fraction distributions provided by CFD for the stoichiometric DME–air fame are shown in Fig. [3](#page-3-0)

As is shown in Fig. [3](#page-3-0), the temperature and $CO₂$ concentration profles are uniform along the optical path in the core of the Hencken fame. The boundary layers can be modeled using the simplifed 2-T/trapezoid profle [\[46\]](#page-11-1) or more realistic Boltzmann profile [[16](#page-10-3), [40](#page-10-20)]. In this work, the Boltzmann profile in the form of Eq. (1) (1) (1) was used to model the flame temperature and $CO₂$ concentration distribution in path-integrated absorbance calculations. In Eq. ([1](#page-3-1)), *y* denotes the physical property (temperature or $CO₂$ concentration), *x* represents the spatial location along the line-of-sight, $C_{1,y}$ and $C_{2,y}$ are the values of the physical properties in the fame core (determined from spectral ftting to absorbance spectra) and the ambient condition $(C_{2,T} = 300 \text{ K}, C_{2,x} = 700 \text{ ppm})$, respectively, C_3 is the spatial location where the fame temperature or concentration is equal to $\frac{C_{1,y} + C_{2,y}}{2}$ and $\frac{1}{C_4}$ is to characterize the gradient of the flame periphery. Both C_3 and C_4 can be inferred from the CFD simulation. For example, for the fame condition shown in Fig. [3,](#page-3-0) $C_{3,T} = 1.4$ cm, $C_{4,T} = 0.23$ cm and $C_{3,X} = 1.35$ cm, $C_{4,X} = 0.2$ cm.

$$
y = C_{2,y} + \frac{C_{1,y} - C_{2,y}}{1 + \exp\left(\frac{x - C_3}{C_4}\right)}.
$$
\n(1)

The laminar premixed fame studied was a spherically expanding fame which was generated inside a fxed-volume combustion vessel [[47](#page-11-2), [48\]](#page-11-3) since this vessel has been widely used to study laminar burning velocities. In this work, ethylene–air mixtures were ignited at the center of the combustion vessel at atmospheric pressure. The combustion process and the subsequent expansion of the fame causes the pressure to rise which was characterized using a pressure transducer. In addition, a high-speed camera was used to monitor the fame expansion to provide timeresolved measurements of the fame diameter (i.e., the absorbing path length).

100 80 Height (mm) 60 40

Fig. 3 CFD calculation of the **a** $CO₂$ concentration and **b** temperature distribution for DME– Air Hencken burner fame for stoichiometric conditions. Only one half of the fame is shown because it is center-symmetric

4 Theory of absorption spectroscopy

4.1 Isolated line model

At low to moderate densities, the isolated line approximation is typically extremely accurate and the absorbance spectrum of a given molecule can be calculated for a non-uniform lineof-sight using Eq. ([2\)](#page-4-0).

$$
\alpha(\nu - \nu_0) = \int_0^L n X_{\text{abs}}(x) \cdot \sum_i S_i(T(x)) \phi_i(\nu - \nu_0, \Delta \nu_{\text{D}}, \Delta \nu_{\text{C}}) dx.
$$
\n(2)

Here, the subscript *i* denotes a given absorption transition, L (cm) is the entire absorbing path length and x is the position variable representing the location along the LOS, n (mol·cm⁻³) is the total gas number density and $X_{\text{abs}}(x)$ and $T(x)$ are the spatially dependent mole fraction and temperature of the absorbing species and are given in the form of Eq. (1) (1) , $S_i(T)$ (cm⁻¹/mol·cm⁻²) is the transition line strength, and ϕ_i (cm) is the transition lineshape function. The Voigt profile is very commonly used to model the transition lineshape and it is given by the convolution of the Gaussian profle $G(\nu', \Delta \nu_{\rm D})$ and Lorentzian profile $L(\nu', \Delta \nu_{\rm C})$:

$$
\phi(\nu', \Delta v_{\rm D}, \Delta v_{\rm C}) = \int_{-\infty}^{+\infty} G(\nu'', \Delta v_{\rm D}) L(\nu' - \nu'', \Delta v_{\rm C}) \, \mathrm{d}\nu'', \tag{3}
$$

where v' is the relative wavenumber with regard to the transition linecenter v_0 (cm⁻¹), Δv _D (cm⁻¹) and Δv _C (cm⁻¹) are the Doppler full-width at half-maximum (FWHM) and collisional FWHM, respectively.

4.2 Line‑mixing model

At high density or for densely packed transitions, collisional coupling (i.e., line mixing) can have a large impact on a molecule's absorbance spectrum and, thus, the isolated line approximation is no longer accurate. In this case, the socalled G-matrix formula can be used to calculate the absorbance spectrum using Eq. ([4\)](#page-4-1) for a non-uniform line-of-sight:

$$
\alpha(\nu) = \int_{0}^{L} \frac{n X_{\text{abs}}(x)}{\pi} \text{Im}(\mathbf{d} \cdot \tilde{\mathbf{G}}^{-1}(\nu) \cdot \tilde{\mathbf{p}} \cdot \mathbf{d}^{T}) dx,
$$
 (4)

where $\tilde{\rho}$ is a $N \times N$ diagonal matrix where *N* is the number of absorption transitions. The diagonal elements of $\tilde{\rho}$, ρ_i , are given by the Boltzmann population fraction of the lower state of the transition and its off-diagonal elements are set to zero. The row vector **d** and its transpose \mathbf{d}^T represent the transition amplitude, with its element given by Eq. (5).

$$
d_i = \sqrt{\frac{S_i(T(x))}{\rho_i}}.\tag{5}
$$

The matrix \tilde{G} accounts for the collisional broadening and line mixing and can be given as:

$$
G(v) = v\tilde{\mathbf{I}} - \tilde{\mathbf{v}}_0 + iP\tilde{\mathbf{W}},
$$
\n(6)

where $\tilde{\mathbf{I}}$ is the *N*×*N* identity matrix, $\tilde{\mathbf{v}}_0$ is a diagonal matrix with its nonzero element $v_{0,i}$ (cm⁻¹) corresponding to the transition frequencies, P (atm) is the pressure, and \tilde{W} is the collisional-relaxation matrix where its diagonal and offdiagonal elements are given by Eq. [\(7](#page-4-2)).

$$
W_{ij} = \begin{cases} \gamma_i - i\delta_i, & i = j \\ R_{ij}, & i \neq j. \end{cases}
$$
 (7)

The subscripts *i* and *j* denote the rotational quantum num-bers. The diagonal elements of Eq. ([7\)](#page-4-2) $\gamma_i - i\delta_i$ were taken from the HITEMP database $[18]$ $[18]$. The off-diagonal elements of *Wij* correspond to the state-to-state collision rate. The collision rate is provided by the modifed exponential energy gap (MEG) law as shown in Eq. ([8\)](#page-4-3).

$$
R_{ij} = a_1(T) \left[\frac{1 + a_4 \left(\frac{E_i''}{a_2 k_B T} \right)}{1 + a_4 \left(\frac{E_i''}{k_B T} \right)} \right] \times \exp \left[\frac{-a_3 \left(E_j'' - E_i'' \right)}{k_B T} \right].
$$
 (8)

The collisional-broadening coefficients and the MEG law coefficients used here were taken from Lee et al.[[37\]](#page-10-17) which are strictly valid for the $CO₂–Ar$ collision system. In fames, the compositions of gaseous species are much more complex. As a result, in this work the collision rate obtained from the $CO₂$ –Ar system was scaled by an empirically derived scaling constant, *k*, according to Eq. ([9\)](#page-4-4).

$$
\tilde{\mathbf{W}}_{\text{CO}_2-\text{mix}}(T) = \tilde{\mathbf{W}}_{\text{CO}_2-Ar}(T) \cdot k = \tilde{\mathbf{W}}_{\text{CO}_2-Ar}(T) \left(\frac{\gamma_{\text{CO}_2-\text{mix}}(T)}{\gamma_{\text{CO}_2-Ar}(T)} \right),\tag{9}
$$

$$
\gamma_{\text{CO}_2\text{-mix}} = \sum_B X_B \gamma_{\text{CO}_2\text{-B}}.\tag{10}
$$

The subscript $CO₂–B$ denotes the collision partner between CO_2 and an arbitrary gaseous species B, X_B is its corresponding mole fraction. It should be noted that in Eqs. (9) and (10) (10) , the species-specific collisionalbroadening coefficient is also rotational-level (i.e., state) dependent. Ideally, Eqs. ([9\)](#page-4-4) and ([10\)](#page-4-5) should be evaluated for all the relevant rotational levels. Unfortunately, a thorough and experimentally verifed database of collisionalbroadening parameters between $CO₂$ and major combustion species like N_2 , O_2 , H_2O , etc. is currently unavailable especially for high rotational quantum numbers. In the

work by Nair et al. [[39\]](#page-10-19), *γ*_{CO₂−*B*} in Eq. [\(10\)](#page-4-5) was referenced to the $J'' = 100$ transition and the collisional-broadening coefficients for $CO₂$ colliding with several major combustion species were provided by Rosenmann et al. [[49](#page-11-4)]. This can be justifed by the fact that at high temperatures, the broadening coefficients between $CO₂$ and other gaseous species are weakly dependent on *J* [[50,](#page-11-5) [51](#page-11-6)]. In this work, we further took advantage of this approximation and chose using the *J*-averaged collisional-broadening coefficient to calculate the scaling parameter *k* as the initial guess for the relaxation matrix calculation for several fame gases and conditions [\[49\]](#page-11-4).

$$
\overline{\gamma}(T) = \sum_{i=0}^{J_{\text{max}}} \rho_i \gamma_i(T), \qquad (11)
$$

where ρ_i is the fractional population in the rotational level *i* according to the Boltzmann distribution, and γ_i is the corresponding broadening coefficient.

Fig. 4 Value of *k* for various collision partners or a mixture of collision partners corresponding to DME–air fame as a function of equivalence ratio. Note, the value of *k* varies with equivalence ratio for a single collision partner due to the temperature and concentration varying

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Figure [4](#page-5-0) shows how $k = X_B \cdot \gamma_{\text{CO}_2 - B} / \gamma_{\text{CO}_2 - \text{Ar}}$ varies with equivalence ratio for several collision partners and a mixture of collision partners corresponding to premixed DME–air fame gas at equilibrium. The gaseous compositions and the temperatures were calculated assuming at constant enthalpy and pressure condition. It can be concluded from Fig. [4](#page-5-0) that *k* does not vary signifcantly with the equivalence ratios for the mixture, this is largely due to the fact that the collisional rates for $CO₂–Ar$ and $CO₂–N₂$ systems are highly alike [[28](#page-10-25)]. To account for the uncertainties from inexact gas compositions and from the calculated collisional-broadening coefficients etc., in this work, we set k as a free parameter in the spectral-fitting routine used for data processing and the calculations based on Eqs. ([9\)](#page-4-4)–([11](#page-4-5)) were only used as the initial guess value for *k* in the spectral-ftting routine. We will also show how diferent values of *k* affect the best-fit temperatures in Sect. 4.3.

5 Results and discussion

5.1 Preprocessing

Figure [5](#page-5-1) shows example single-scan measurements of *I*o, I_t , and the corresponding CO_2 absorbance spectrum. I_0 was recorded before each measurement. To mitigate the impact of beam steering induced by the flame, I_0 was recorded in the open air as well as in H_2 –Air flames at diferent equivalence ratios. An example measurement of I_0 is shown in Fig. [5b](#page-5-1) which illustrates that I_0 was not signifcantly impacted by the fame. In addition, Fig. [5a](#page-5-1) and c shows that CO_2 absorption near 4.172 μ m is very weak at room temperature, indicating the ambient absorption has negligible effects on the baseline and flame measurements.

Fig. 5 a Example of the measured transmitted light intensity I_t and the unabsorbed background I_0 . **b** Measured unabsorbed background signal in the open air (solid line) and in the stoichiometric H_2 –Air fame (dash line). **c** Simulated absorbance spectrum based on HITEMP database and Voigt profle

5.2 Spectral‑ftting routine

The spectral-ftting routine is described as follows. The experimentally measured spectra were preprocessed as described in Sect. 4.1. We used the profle-ftting strategy as introduced in Eqs. ([2\)](#page-4-0) and ([4](#page-4-1)) to account for the spatial inhomogeneity in Hencken burner fames. The temperature and species concentration profles were given in the form of Eq. [\(1](#page-3-1)) and $C_{1,T}$ and $C_{1,X}$ (the values of temperature and CO_2 mole fraction in the fame core) were determined from the spectral-fitting routine. Parameters like $C_{2,T}$ and $C_{2,X}$ were pre-set ($C_{2,T}$ = 300 K, $C_{2,x}$ = 700 ppm), while $C_{3,T}$, $C_{3,x}$ and $C_{4,T}$, $C_{4,X}$ were inferred from the CFD simulation. For the isolated line model, the theoretical spectra were calculated according to Eqs. ([2](#page-4-0)) and ([3\)](#page-4-6). For the line-mixing model, the theoretical spectra were generated according to Eqs. [\(4\)](#page-4-1)–[\(10\)](#page-4-7). The transition linestrength, $S_i(T)$, and the unperturbed transition frequency $v_{0,i}$ were taken from the HITEMP database [\[18](#page-10-6)]. The collisional linewidth γ_i for CO₂–Ar and the off-diagonal collisional rates R_{ij} were taken from Lee et al. [[37\]](#page-10-17). The temperature (C_{1T}) , C O₂ concentration (C_{1X}) , the scaling parameter k of the relaxation matrix were set as free parameters in the spectral-ftting routine. In addition, a wavelength shift parameter and a baseline shift parameter for correcting I_o were included in the line-mixing model as well

Table 1 Summary of free parameters used in the spectral-ftting routine

Fitting parameter	Floating range
Concentration $(C_{I,X})$	5 to 20\%
Temperature (C_{LT})	1200 to 2800 K
Pressure	0.95 to 1.05 atm
Wavelength shift parameter	-0.02 to 0.02 nm
Baseline shift parameter	-0.05 to 0.05
k (line-mixing calculation only)	$1.5 \text{ to } 2.5$

 0.5

Fig. 6 Comparisons of best-ft absorbance spectra for the same experimental data obtained using **a** the isolated line model and **b** the line-mixing model, both of which account for line-of-sight non-uniformities. LM2 denotes the best-ft from assuming uniform distribution of flame temperature and $CO₂$ concentration along the line-ofsight with an optical path length of 2.54 cm. The measurement was performed in a stoichiometric DME–air fame and the experimental spectra was averaged over 100 consecutive scans

as the isolated line model. A summary of the free parameters used in the spectral-ftting routine is shown in Table [1](#page-6-0).

For each dataset, the best-ft spectrum for the frst scan/ measurement was ftted using the brute-force search combined with the Broyden–Fletcher–Goldfarb–Shanno (BFGS) method [[52\]](#page-11-7) to ensure that the ftting routine converged on the global minimum. After that, the best-ft values for each free parameter was used as the initial guess values for each consecutive measurement and the least-squares method was used to fnd the best-ft parameters. This is to reduce the load of numerical calculations as well as to fnd the global optimum solution.

5.3 Hencken burner fame

Laser-absorption measurements were frst acquired in a laminar non-premixed DME–air fame. The volumetric fow rate of the air was set at 15 standard liter per minute (SLM) so that the overall speed of the fuel–air mixture was slightly higher than the laminar fame speed of DME at the stoichiometric combustion condition. An example of the measured $CO₂$ absorbance spectrum compared to best-fit spectra obtained using the conventional isolated line model (i.e., Voigt model) and the line-mixing (LM) model are shown in Fig. [6.](#page-6-1) The agreement between the measured and best-ft spectra is much better for the line-mixing model than the isolated line model, especially at the bandhead where the transitions overlap most signifcantly, the overall relative ftting residual from the isolated line model (15%) is about 3 times larger than from the line-mixing model (4%). In addition, we also compared the spectral-ftting results obtained with LM model and by assuming a uniform distribution of flame temperature and $CO₂$ concentration (denoted as LM2) in Fig. [6b](#page-6-1). The best-ft spectra were not signifcantly diferent between two assumptions; however, the ftting accuracy was slightly improved near the absorption bandhead using the non-uniform temperature and $CO₂$ concentration profile

Fig. 7 Summary of **a** measured core flame temperature $(C_{1,T})$ and \bf{b} core $CO₂$ mole fractions $(C₁ x)$ for DME–air flames with various equivalence ratio acquired using the line-mixing model or the Voigt model

as inferred from CFD calculations. Further, based on the uniform fame property assumption, the best-ft temperature T_{LM2} = 2217 K was slightly lower than T_{LM} = 2285 K and the best-fit concentration $X_{LM2} = 13.8\%$ was higher than $X_{LM} = 12.9\%$.

In addition, Fig. [7](#page-7-0) shows a comparison of the measured core fame temperature obtained using the two models at various equivalence ratios. The adiabatic fame temperature was calculated assuming chemical equilibrium using Cantera [\[53\]](#page-11-8) with the kinetic model developed by Qi et al. [[43\]](#page-10-23). For the stoichiometric combustion condition, the fame temperature and $CO₂$ mole fraction in the flame core obtained from CFD simulations are also shown.

In Fig. [7](#page-7-0), the error bar was evaluated based on the standard deviation from the 100 measured spectra acquired at 10 kHz rate as well as the uncertainty from the spectralftting routine, the magnitude of the error bar was 40–120 K for flame temperature and $0.5-0.8\%$ for $CO₂$ mole fraction measurement and for equivalence ratios from 0.7 to 1.6. The absolute uncertainty of fame temperature, on the other hand, was about 30–80 K as estimated from previous femtosecond CARS measurement [[54,](#page-11-9) [55](#page-11-10)]. For all the fame conditions studied, the temperature measured using the Voigt model were about 100–200 K higher than those obtained using the line-mixing model. This is caused by the fact that Voigt model underestimates the absorbance at the bandhead. For stoichiometric combustion condition, the fame core

temperature from CFD calculation was 140 K lower than adiabatic fame temperature, this is caused by the fact that heat loss due to radiation and convection were considered in the CFD simulation. The agreement between the temperature obtained using the line-mixing model and the adiabatic temperature were in general satisfactory except at high equivalence ratios. For the fuel rich conditions, the measured temperatures obtained using the line-mixing model are 50–125 K lower than the adiabatic fame temperatures. The behavior was also found in other studies investigating fuel rich fames with carbonaceous fuels [[56](#page-11-11), [57\]](#page-11-12) and is attributed to the radiative heat loss from soot. The best-fit $CO₂$ mole fractions from line-mixing calculations are 1–3% higher than the adiabatic equilibrium. On the other hand, at stoichiometric combustion condition, the best-fit $CO₂$ mole fraction is 2% lower than it was obtained from the CFD calculations. Considering the fact that many previous studies have shown that the experimentally measured fame temperatures were usually only 30–80 K lower than their adiabatic values [[54,](#page-11-9) [55](#page-11-10)], it is likely that our current CFD calculation could overestimate the heat loss due to radiation and convection. While the adiabatic equilibrium calculations do not consider heat loss, it is reasonable to infer that the real fame temperatures and $CO₂$ concentrations were in between the values obtained from CFD simulation and adiabatic calculations as indicated by the laser absorption measurements.

Fig. 8 a The experimentally ftted scaling parameter *k* for the collisional coefficient as compared to the scaling parameters calculated from *J*-averaged collisional-broadening coefficient and from the collisional-broadening coefficient referenced from $J = 100$. **b** The best-fit temperatures corresponding to diferent values of *k*. The results shown are for DME–air fames

To better understand the accuracy of the line-mixing model, we compared the best-ft value of *k* with that predicted by two models. Figure [8](#page-7-1)a shows how the measured and predicted values of *k* difer as a function of equivalence ratio. The diference between the best-ft (i.e., ftted) *k* and the theoretical predictions is not surprising, considering the fact that the *J*-dependencies cannot be totally neglected. In addition, there are some studies showing discrepancies between the experimentally measured linewidth and those predicted by the calculations proposed by Rosenman et al. [\[58\]](#page-11-13).

Figure [8](#page-7-1)b illustrates how the measured temperature varies with equivalence ratio and with the method for determining *k* (inferring it or fxing it to diferent predicted values). Increasing or decreasing *k* leads to the same trend in temperature. While there is signifcant uncertainty in the most accurate value for *k* and how to infer or predict it, it is clear that direct determination of the collisional-relaxation matrix for each mixture can help increase the accuracy of the temperature measured using the line-mixing model.

5.4 Spherical fame inside constant‑volume combustion vessel

Inside the fxed-volume combustion vessel, the pressure was monitored using a pressure transducer. In the experiment, an ethylene–air mixture with an equivalence ratio of 2.2 was frst ignited in the center of the combustion vessel with the gas initially at 1 atm. The combustion process then caused the pressure to increase throughout the test. The pressure measured by the transducer was used as an input to the spectral-ftting routine. The optical path length through the spherical fame can be precisely determined through highspeed imaging as shown in Fig. [9](#page-8-0). Measured time histories of temperature, $CO₂$ concentration and pressure are shown in Fig. [10.](#page-8-1)

The temperature rise is obvious and expected during ignition; however, the drop in $CO₂$ concentration might be

Fig. 10 Measured time histories of **a** the gas temperature, **b** the $CO₂$ concentration, and **c** the gas pressure for the ethylene–air fame inside the constant-volume combustion vessel. The temperature and concentration of $CO₂$ were determined using the line-mixing model. The fuel–air equivalence ratio was 2.2 and the shaded area denotes the uncertainty due to spectral-ftting routine

related to uncertainty in the fame diameter when it was small, as well as the low signal-to-noise ratio at early times. In general, the measured temperature and $CO₂$ concentration agree well with the adiabatic fame calculations after the fame entered a quasi-steady propagation stage.

Examples of the measured absorbance spectra and the corresponding best-ft spectra acquired during diferent periods of time during the spherical fame expansion are presented in Fig. [11.](#page-9-8) In general, the agreement between the measured $CO₂$ absorbance spectra and the best-fit spectra produced using the line-mixing model are satisfactory, particularly when the pressure is near 1 atm. The

Fig. 9 Example images of fame front and schematic of the fxedvolume combustion vessel used to generate laminar, spherically expanding, premixed fames

Fig. 11 Spectral-ftting results of $CO₂$ absorption spectra obtained using the line mixing model for spectra that were measured in the spherically expanding fame. The solid black curve denotes the experimentally measured spectra and the red dash curve represents the best-ft spectra

magnitude of the measured absorbance increases in time due to the growth of the flame. At $t = 80$ ms, differences between the experimentally measured spectra and the bestft spectrum become more obvious than at earlier times (lower pressures). While the pressure as well as the scaling parameter increase overtime, this might indicate that the current treatment of using one single scaling parameter for the G-matrix calculation correction could fail for some intermediate pressures.

6 Conclusions

Laser absorption spectroscopy measurements of gas temperature and $CO₂$ concentration were acquired in laminar non-premixed fames and spherical premixed fames with varying equivalence ratio. The measurements were acquired by scanning the wavelength of a distributed-feedback interband-cascade laser across transitions near the bandhead of the fundamental v_3 absorption band of CO_2 near 4.17 μ m. A collisional line mixing model based on previous studies of high pressure CO_2 –Ar mixtures [[37\]](#page-10-17) was modified to provide high fidelity measurements in flames. The simulated spectra produced using the collisional line mixing model showed signifcantly better agreement with the experimentally measured spectra compared to simulated spectra produced using the conventional isolated line approximation and the Voigt profle. The temperature corresponding to the best-ft Voigt model signifcantly exceeded the values obtained using the line-mixing model, the latter of which agreed well with adiabatic equilibrium fame temperatures. As a result, this work illustrated the importance of accounting for collisional line mixing for the $CO₂$ transitions studied despite the modest pressure and high temperature of the fame gas. Future work will include a more accurate characterization of the collisional-broadening coefficients for different collisional species and to develop a more versatile line mixing model for low to moderately high pressures.

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