

# **Analysis of laser absorption gas sensors employing scanned‑wavelength modulation spectroscopy with 1***f***‑phase detection**

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## **Abstract**

The recently introduced wavelength-modulation spectroscopy with 1*f*-phase detection (WMS- $\theta_{1f}$ ) technique showed promising results with potentially improved measurement precision over the popular 1*f*-normalized WMS-*nf* (WMS-*nf*/1*f*) technique. Like WMS- $nfl$ <sup>1</sup>*f*, WMS- $\theta_{1f}$  enjoys the typical benefits of WMS methods, including low-frequency noise rejection, correction for non-absorbing losses, and insensitivity to the broadband absorption spectra of interfering species. In this work, we performed a detailed analysis of the spectrally resolved scanned-wavelength WMS- $\theta_{1f}$  measurement technique and its direct comparison against the expected performance of scanned-wavelength WMS-*nf*/1*f* and scanned-wavelength direct-absorption spectroscopy (SDAS) measurements. This simulation-based analysis identifed specifc operating regimes in which the performance of WMS- $\theta_{1f}$  measurements in terms of accuracy is expected to be greater than the performance of WMS-*nf*/1*f* or SDAS. Additionally, improved guidelines for the optimal selection of laser-tuning parameters, including an explicit optimization of the optical scan depth parameter, were developed. Experiments with a  $CO<sub>2</sub>$  static cell perturbed by a high-speed air jet corroborated the simulation-based findings. Finally, a practical demonstration of a WMS- $\theta_{1f}$  sensor for measuring temperature and  $H_2O$  mole fraction in the exhaust of a CH<sub>4</sub>/air flat-flame burner was presented, with the results confirming model predictions of the superior precision of WMS- $\theta_{1f}$  relative to WMS- $nfl$  *f* and SDAS.

# **1 Introduction**

Tunable diode-laser absorption spectroscopy (TDLAS) is a proven technology widely utilized in combustion, hypersonics, industrial process control, and environmental sensing for quantitative in situ measurements of multiple gas properties such as species concentration, temperature, pressure, and velocity  $[1-5]$  $[1-5]$ . Scanned-wavelength direct-absorption spectroscopy (SDAS), in which gas parameters are inferred from the measured transmission fraction of a laser source as it tunes across absorption features of a target species, is the most widely used TDLAS technique due to its simplicity and robustness in relatively quiescent environments. For harsh environments where signal distortions such as optical emission, beam-steering, étalons, and general engineering

 $\boxtimes$  Wen Yu Peng wypeng@stanford.edu noise are substantial, wavelength modulation spectroscopy (WMS) is the preferred measurement technique due to its ability to suppress the effects of many forms of signal distortions. This is accomplished by modulating the wavelength of the laser at a high frequency, thereby encoding information related to the absorption spectrum at high frequencies and rejecting low-frequency distortions with lock-in flters. Further, wavelength-independent multiplicative signal distortions (e.g. beam-steering) can be corrected with the 1*f*-normalized WMS-*nf* detection scheme (for *n >* 2 and referred to as WMS-*nf*/1*f*), thereby enabling a calibrationand background-free gas absorption measurement technique. In exchange for added sensor design and signal-processing complexity, WMS-*nf*/1*f* is demonstratively a better measurement technique than SDAS for noisy conditions and has been successfully deployed in a variety of harsh environ-ments that hinder SDAS sensing [[6–](#page-21-2)[13\]](#page-22-0).

Until recently, researchers have focused on optimizing, analyzing, and deploying WMS-*nf*/1*f* sensors for diverse environments. A new WMS-based technique, here referred to as WMS-1*f*-phase detection (shortened to WMS- $\theta_{1f}$  or  $\theta_{1f}$ ), was only recently introduced and explored in several

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publications [[14,](#page-22-1) [15](#page-22-2)]. In WMS- $\theta_{1f}$ , the phase angle of the 1st harmonic is used for detecting gas absorption. The resulting WMS- $\theta_{1f}$  signal as the laser is tuned across an absorption feature approximately tracks the frst derivative of the absorption spectrum and, therefore, is sensitive to both absorption linestrength and lineshape. Additionally, WMS- $\theta_{1f}$  signals enjoy all of the standard benefits of WMS-*nf*/1*f* such as beam-steering-distortion- and 1/*f*-additive-noiserejection. The measured  $\theta_{1f}$  signal can then be fitted with a lineshape model to infer gas properties using a ftting routine similar to the algorithm presented in Goldenstein et al.  $[16]$  $[16]$ . Also, since the technique utilizes only the 1st harmonic of the transmitted intensity waveform, it is easier to select appropriate modulation frequencies for frequency-domain multiplexing of multiple lasers [[17\]](#page-22-4) onto the same detector without encountering cross-talk issues between the *nf* harmonics of the modulation signal.

Yang et al. [\[14](#page-22-1)] presented a theoretical description and experimental demonstration of the WMS- $\theta_{1f}$  method for CO<sub>2</sub> detection in a static gas cell. Upadhyay et al. [[15\]](#page-22-2) presented a theoretical description and experimental demonstration of the WMS- $R_{1f}/Y_{1f}$  method (which carries the same information as  $\theta_{1f}$  and is, therefore, indistinct. See Sect. [2.2](#page-2-0) for clarification) also for  $CO<sub>2</sub>$  detection in a static gas cell and the exhaust of a jet engine combustor. Both research teams directly compared the WMS- $\theta_{1f}$  or - $R_{1f}/Y_{1f}$  results against WMS-*2f*/1*f* results and demonstrated that the new techniques performed better than WMS-*2f*/1*f* in terms of Allan variance and accuracy. However, these comparisons were performed using identical laser-tuning parameters for both the  $\theta_{1f}$  and WMS-*2f*/1*f* measurements rather than at optimized laser-tuning parameters for each respective measurement technique. To appropriately demonstrate the superior performance of WMS- $\theta_{1f}$  relative to WMS-2*f*/1*f*, the laser-tuning parameters should be set at optimal conditions for each method independently.

In this work, we extended the framework developed in Yang et al. and Upadhyay et al. and performed an analysis of the WMS- $\theta_{1f}$  measurement technique and its expected performance in harsh environments. This analysis included a theoretical description of WMS- $\theta_{1f}$  (Sect. [2.2](#page-2-0)), a discussion on the WMS- $\theta_{1f}$  spectral fitting routine (Sect. [2.3\)](#page-4-0), and a numerical assessment of the technique's sensitivity to environmental distortions commonly encountered in harsh measurement environments (Sects. [3.2](#page-6-0) and [3.3\)](#page-12-0). The new distortion analysis techniques also enabled the direct comparison of the expected performance of WMS- $\theta_{1f}$ , WMS-*nf*/1*f*, and SDAS in arbitrary measurement environments, leading to new guidelines for selecting the optimal measurement strategy and laser-tuning parameters. These guidelines are distinct from procedures previously reported in the literature

[[8,](#page-21-3) [17,](#page-22-4) [18\]](#page-22-5). They represent the frst-reported analysis and optimization scheme that (1) quantitatively identifes the measurement technique (out of WMS- $\theta_{1f}$ , *nf*/1*f*, and SDAS) that is expected to perform the best for a given TDLAS application and (2) identifes laser operating parameters that explicitly minimize the impact of environmental signal distortions on the accuracy of WMS- or SDAS-based spectroscopic measurements. Several important results were achieved from this analysis.

- 1. Assuming broadband 1/*f* stochastic additive and multiplicative environmental signal distortions (e.g. optical emission and beam-steering), WMS- $\theta_{1f}$  measurements are generally more accurate than all WMS-*nf*/1*f* and SDAS techniques regardless of the selected laser-tuning parameters.
- 2. The optical scan depth parameter, often neglected in laser-tuning parameter optimization procedures for spectrally resolved WMS-*nf*/1*f* techniques, significantly afects the accuracy of all WMS-based measurement techniques. Simulations and experiments (Sects. [3.2.2](#page-9-0) and [4.1](#page-14-0)) suggest selecting modulation and scan depths that obey certain linear relationships to minimize the expected error.
- 3. The frequency-domain characteristics of the stochastic signal distortions dictate the relative performance of the various WMS measurement techniques. Assuming the spectral amplitude of the distortions decays as a function of 1/*f*<sup>*p*</sup> for some roll-off exponent, *p*, it was found that, above a certain value of *p*, WMS-*nf*/1*f* measurements became more accurate than WMS- $\theta_{1f}$  measurements. This provides useful guidelines for selecting the best WMS technique for a given measurement environment.
- 4. Conversely, for signal distortions with *p* near zero (i.e. white noise), the performance of SDAS-based sensors was found to exceed WMS-*nf*/1*f*-based sensors while rivaling the performance of WMS- $\theta_{1f}$ . This demonstrates that there are certain environments in which the use of WMS instead of SDAS is actually detrimental to the performance of a TDLAS sensor.
- 5. WMS- $\theta_{1f}$  is qualitatively as sensitive to wavelengthdependent distortions (e.g. étalons) as WMS-*2f*/1*f*. Similar to WMS-*2f*/1*f*, the measurement accuracy is less adversely impacted by étalon distortions with free spectral ranges that are sufficiently small or large relative to the absorption linewidth.
- 6. Short-period étalon distortions were found to impact the accuracy of SDAS measurements more strongly than WMS- $\theta_{1f}$ . For applications using long optical cavities (e.g. cavity-enhanced techniques) where étalon interfer-

ence fringes generally exhibit low free-spectral range, it may be beneficial to use WMS- $\theta_{1f}$  instead of SDAS.

The fndings of the simulation-based analysis were corroborated with experimental measurements of  $CO<sub>2</sub>$  in a static cell with the laser beam perturbed by a high-speed air jet in free space to simulate broadband environmental distortions. Results from these experiments demonstrate the improved accuracy and precision of the WMS- $\theta_{1f}$  results against all WMS-*nf*/1*f* and SDAS strategies in distortion environments with specific frequency characteristics. The WMS- $\theta_{1f}$  technique was then practically demonstrated by measuring the temperature and mole fraction of  $H_2O$  within the combustion product stream of a  $CH<sub>4</sub>/air$  flat-flame burner. Allan deviations of the WMS- $\theta_{1f}$  results were found to be lower than the Allan deviations of the various WMS-*nf*/1*f* techniques, further demonstrating the utility of the WMS- $\theta_{1f}$  technique for highprecision measurements in harsh environments.

# **2 Scanned‑WMS‑1***f***‑phase detection principles**

The scanned-WMS-1*f*-phase detection technique is grounded in WMS theory, which is well documented in the literature [\[17–](#page-22-4)[21\]](#page-22-6). The fundamentals of LAS and WMS are briefy reviewed in this section to clarify notation. These are followed by a theoretical description of WMS- $\theta_{1f}$  signals and a discussion on the spectral ftting routine used to infer lineshape parameters from a  $\theta_{1f}$  measurement. Additional nuances related to the  $\theta_{1f}$  fitting routine not seen in typical WMS-*nfl* 1*f* ftting routines are highlighted and discussed.

## **2.1 Laser absorption spectroscopy**

The spectral transmissivity,  $\tau_{\nu}$ , defined as the ratio between the transmitted  $(I_t)$  and incident  $(I_0)$  intensities, of monochromatic radiation with frequency  $v$  (cm<sup>-1</sup>) through a uniform ideal gas is described by the Beer–Lambert relation:

$$
\tau_{v} = \left(\frac{I_{t}}{I_{0}}\right)_{v} = \exp(-\alpha_{v}) = \exp\left\{-\sum_{k} S_{k} P \chi L \phi_{v,k}\right\},\quad(1)
$$

where  $\alpha_{\nu}$  is the spectral absorbance,  $S_k$  (cm<sup>-2</sup> atm<sup>-1</sup>) is the linestrength of the *k*th absorbing transition, *P* (atm) is the absolute gas pressure,  $\chi$  is the mole fraction of the absorbing species, *L* (cm) is the optical path length, and  $\phi_{\nu k}$  (cm) is the frequency-dependent normalized lineshape of the *k*th absorbing transition.  $S_k$  is purely a function of temperature, physical constants, and known molecular parameters [[22\]](#page-22-7) and can be calculated as follows:

$$
S_{k}(T) = S_{k}(T_{0}) \frac{T_{0}}{T} \frac{Q(T_{0})}{Q(T)} \exp\left[-\frac{hcE_{j}''}{k_{B}} \left(\frac{1}{T} - \frac{1}{T_{0}}\right)\right]
$$

$$
\times \left[1 - \exp\left(-\frac{hc\nu_{0,k}}{k_{B}T}\right)\right] \left[1 - \exp\left(-\frac{hc\nu_{0,k}}{k_{B}T_{0}}\right)\right]^{-1}, \tag{2}
$$

where  $S_k(T_0)$  is the linestrength at some reference temperature  $T_0$  (296 K throughout this work), *Q* is the total internal partition function of the absorbing species,  $E_k''(\text{cm}^{-1})$  is the lower state energy of the transition, and  $v_{0,k}$  (cm<sup>-1</sup>) is the transition vacuum linecenter. The constants *h* (erg s), *c* (cm s<sup>-1</sup>), and  $k_B$  (erg K<sup>-1</sup>) are the Planck constant, speed of light, and Boltzmann constant, respectively.

The normalized Voigt lineshape function, which accounts for Doppler and collisional broadening efects, is widely used to model  $\phi_{v,k}$ . These effects are characterized by the Doppler broadening full-width half-maximum (FWHM),  $\Delta v_D$  (cm<sup>-1</sup>), and the collisional broadening FWHM,  $\Delta v_C$  (cm<sup>-1</sup>), given by

<span id="page-2-1"></span>
$$
\Delta v_{\rm D} = v_0 \sqrt{\frac{8 \ln(2) k_{\rm B} T}{mc^2}}\tag{3}
$$

$$
\Delta v_{\rm C} = 2P \sum_{k} \chi_{k} \gamma_{k,0} \left(\frac{T_0}{T}\right)^{n_k},\tag{4}
$$

where *m* (g) is the absorbing species molecular mass,  $\gamma_{k,0}$  (cm<sup>-1</sup> atm<sup>-1</sup>) is the collisional broadening coefficient of the *k*th colliding species at the reference temperature, and  $n_k$  is the temperature-dependent exponent of the broadening coefficients.

For measurements of *T*, *P*, or  $\chi$  in spectrally resolved TDLAS, it is often convenient to defne the integrated absorbance,  $A_k$  (cm<sup>-1</sup>), as the integral of the absorbance for a single absorption transition:

<span id="page-2-3"></span>
$$
A_k = \int_{-\infty}^{\infty} S_k P \chi L \phi_k(v) \mathrm{d}v = S_k P \chi L. \tag{5}
$$

## <span id="page-2-0"></span>**2.2 Scanned-WMS-** $\theta_{1f}$

<span id="page-2-2"></span>Extensive information describing scanned-WMS can be found in the literature  $[16, 21, 23]$  $[16, 21, 23]$  $[16, 21, 23]$  $[16, 21, 23]$  $[16, 21, 23]$ ; therefore, only a brief overview of its theoretical framework and its relation to WMS- $\theta_{1f}$  is provided here. WMS fundamentally involves modulating the laser injection current with a sinusoidal waveform of frequency  $f_m$  to tune the laser wavelength around the linecenter of an absorption feature. In scanned-WMS, an additional lower frequency sinusoidal scanning waveform of frequency  $f_s$  is added to the high-frequency modulation to produce spectrally resolved WMS signals

that can be used for lineshape fitting. The intensity  $(I_0(t))$ and optical frequency  $(v(t))$  response of a laser to sinusoidal injection-current tuning can be described as general Fourier series:

$$
I_0(t) = \overline{I_0} \left[ 1 + \sum_{k=1}^{\infty} i_{s,k} \cos(2\pi k f_s t + \psi_{s,k}) + \sum_{k=1}^{\infty} i_{m,k} \cos(2\pi k f_m t + \psi_{m,k}) \right],
$$
 (6)

$$
v(t) = \overline{v} + \sum_{k=1}^{\infty} a_{s,k} \cos(2\pi k f_s t + \phi_{s,k})
$$
  
+ 
$$
\sum_{k=1}^{\infty} a_{m,k} \cos(2\pi k f_m t + \phi_{m,k}),
$$
 (7)

where the subscripts s and m denote parameters associated with the low-frequency scan and high-frequency modulation, respectively.  $\overline{I_0}$  is the mean intensity,  $i_k$  are the normalized intensity scan or modulation amplitudes,  $\psi_k$  are the intensity scan or modulation phases,  $\bar{v}$  (cm<sup>-1</sup>) is the mean optical frequency,  $a_k$  (cm<sup>-1</sup>) are the optical scan or modulation depths, and  $\phi_k$  are the wavelength scan or modulation phases. The index terms  $k = 1$  and  $k > 1$  represent the linear and nonlinear modulation components, respectively, with  $i_1 \gg i_{k>1}$  and  $a_1 \gg a_{k>1}$  in general. All laser response parameters can be characterized in the laboratory and do not change signifcantly over time for a given sensor system and laser settings.

It is useful to define the modulation and scan indices ( $\mu$ and  $\sigma$ , respectively) as the linear optical modulation and scan depths  $(a_{m,1}$  and  $a_{s,1}$ , respectively) non-dimensionalized by the lineshape half-width half-maximum of the targeted absorption transition. Diferent absorption spectra and lasertuning parameters yield similar results when compared using these non-dimensional parameters. For a single absorption transition described by a Voigt profile,  $\mu$  and  $\sigma$  can be approximated as follows:

$$
\mu = 2a_{m,1}/(\Delta v_C + \Delta v_D),\tag{8}
$$

$$
\sigma = 2a_{s,1}/(\Delta v_C + \Delta v_D). \tag{9}
$$

The interaction between  $I_0(t)$ ,  $v(t)$ , and the spectral transmissivity  $(\tau_{\nu})$  through the probed gaseous medium introduces new frequency content at the harmonics of  $f<sub>m</sub>$  and beat frequencies of  $f_s$  and  $f_m$  in the transmitted intensity,  $I_t(t)$ . The resulting analytical expression for  $I_t(t)$  is highly complex (see Appendix A in [[21\]](#page-22-6), for example) and most practical implementations of WMS simulate rather than analytically compute the harmonic signals [\[23](#page-22-8)]. To understand the qualitative behavior of scanned-WMS signals, it is

more instructive to restrict the analysis to assume (1) no lowfrequency scan (i.e.  $i_{s,k} = a_{s,k} = 0$ ), (2) nonlinear terms are neglected (i.e.  $i_{m,k>1} = a_{m,k>1} = 0$ ), and (3)  $\phi_{m,1} = 0$ , which is achieved in practice by setting an appropriate temporal offset in the data time series. With this simplification,  $I_t(t)$ can be written as follows:

<span id="page-3-4"></span>
$$
I_{\rm t}(t) = G_{\rm opt} I_0(t) \cdot \tau(v(t))
$$
  
=  $G_{\rm opt} \overline{I_0} [1 + i_{\rm m,1} \cos(2\pi f_{\rm m} t + \psi_{\rm m,1})] \cdot \sum_{k=0}^{\infty} H_k \cos(2\pi k f_{\rm m} t),$  (10)

<span id="page-3-5"></span><span id="page-3-0"></span>where  $G_{opt}$  is some constant experimental opto-electronic detection gain and  $H_k$  are the cosine-series coefficients of the time-domain spectral transmissivity function:

<span id="page-3-1"></span>
$$
H_k = \begin{cases} \frac{1}{2\pi} \int_{-\pi}^{\pi} \tau(\overline{v} + a_{m,1} \cos \theta) d\theta & \text{if } k = 0, \\ \frac{1}{\pi} \int_{-\pi}^{\pi} \tau(\overline{v} + a_{m,1} \cos \theta) \cos k\theta d\theta & \text{if } k \neq 0. \end{cases}
$$
(11)

We can now define the *n*th harmonic WMS signal,  $R_{nf}$ , to be the complex Fourier coefficient of the transmitted intensity corresponding to frequency *nf*m, which is extracted from practical measurements using a lock-in/low-pass flter operation:

<span id="page-3-2"></span>
$$
R_{nf} = \text{LPF}(t) * [I_{t}(t) \cdot \exp(j2\pi nf_{\text{m}}t)], \qquad (12)
$$

where  $LPF(t)$  is the impulse response function of the lowpass flter, ∗ is the convolution operator, and *j* is the imaginary unit. Combining Eqs. [10](#page-3-0), [11,](#page-3-1) and [12](#page-3-2), the expression for the complex phase angle of  $R_{1f}$ ,  $\theta_{1f}$ , can be given as follows:

<span id="page-3-3"></span>
$$
\theta_{1f} = \angle (R_{1f}),
$$
  
= tan<sup>-1</sup>  $\left\{ \frac{i_{m,1} \sin \psi_{m,1} \left( H_0 - \frac{H_2}{2} \right)}{H_1 + i_{m,1} \cos \psi_{m,1} \left( H_0 + \frac{H_2}{2} \right)} \right\}.$  (13)

<span id="page-3-6"></span>This expression can be viewed as the WMS-1*f* phase for a single center optical frequency,  $\overline{v}$ . By varying the center optical frequency with a low-frequency scan, a spectrally resolved  $\theta_{1f}$  signal can be measured. An example of such a signal is shown in the bottom panel of Fig. [1](#page-4-1) for single-Voigt profles of various integrated absorbances and constant  $\Delta v_C = 0.1$  cm<sup>-1</sup> and  $\Delta v_D = 0.01$  cm<sup>-1</sup>. Note that the arctangent function used to calculate  $\angle(R_{1f})$  throughout this work is the two-parameter arctangent function with output space  $\in [-\pi, \pi]$ . This choice was arbitrary and has no effect on results when compared to the standard single-parameter arctangent function with output space  $\in [-\pi/2, \pi/2]$ . Several important characteristics related to  $\theta_{1f}$  can be observed from Eq. [13](#page-3-3) and Fig. [1](#page-4-1).



<span id="page-4-1"></span>**Fig. 1** (Top panel) Spectral transmissivity (black) and its derivative (blue) of an example absorption transition described by a single Voigt profile with  $A = 0.15$  cm<sup>-1</sup>,  $\Delta v_C = 0.1$  cm<sup>-1</sup>, and  $\Delta v_D = 0.01$  cm<sup>-1</sup>. (Middle panel) Simulated incident (blue) and transmitted laser intensity (red) for a scanned-WMS signal with  $a_{m,1} = 0.1$  cm<sup>-1</sup> and  $a_{s,1} = 0.5$  cm<sup>-1</sup>. (Bottom panel) The resulting WMS- $\theta_{1f}$  signals for various integrated absorbances. Note how the  $\theta_{1f}$  amplitude scales with *A* and qualitatively tracks the profile of  $d\tau/d\nu$ 

- 1. Similar to WMS-*nf*/1*f* signals,  $\theta_{1f}$  signals are independent of  $G_{opt}I_0$  (as shown in Eq. [13](#page-3-3)) and, therefore, do not require *in situ* intensity calibration while naturally correcting for additive and multiplicative signal distortions such as optical emission, beam-steering, vibrations, and particulate matter occlusion with frequency-domain bandwidth below  $f_m/2$ .
- 2. In the absence of absorption,  $\theta_{1f} = \psi_{m,1}$ , which is simply the phase of the intensity modulation waveform. Additional temporal offsets in the data acquisition system will similarly manifest as constant phase offsets in the  $\theta_{1f}$  signal that can be fitted as part of the spectral fitting routine. This is unlike the phase-sensitive WMS- $R_{1f}/Y_{1f}$ (equivalent to  $|R_{1f}|/Im(R_{1f})$  in the current nomenclature) detection strategy used in Upadhyay et al. [[15\]](#page-22-2).
- 3. *H*<sub>1</sub> is the dominant contributor to changes in  $\theta_{1f}$  relative to the non-absorbing background. As noted in Reid and Labrie [\[18](#page-22-5)],  $H_1 \propto -d\tau/d\nu$  in the  $a_{m,1} \to 0$  limit. There-



<span id="page-4-2"></span>Fig. 2 Phase offset between the intensity- and optical frequencymodulation waveforms as a function of modulation frequency for an injection-current-tuned distributed-feedback tunable diode laser operating near 1.47 μm

fore, the scanned- $\theta_{1f}$  signal approximately tracks the frst derivative of the transmission spectrum as shown in the top panel of Fig. [1.](#page-4-1) Additionally, since  $d\tau/d\nu$  is a function of the absorption spectrum,  $\theta_{1f}$  signals respond directly to the thermodynamic conditions of the absorbing species and, therefore, can be used for spectroscopic measurements.

4.  $\theta_{1f}$  does not respond meaningfully to absorption if the phase offset between the intensity modulation waveform and the wavelength modulation waveform is close to a multiple of  $\pi$  (i.e.  $\sin(\psi_{m,1}) \approx 0$ ). In practice, the thermal mass of typical semiconductor laser optical cavities acts as a low-pass flter with some ohmic heating-induced phase lag between the optical and intensity tuning waveforms, resulting in  $\pi < \psi_{m,1} < 3\pi/2$  [[24](#page-22-9)]. This lag is further pronounced and asymptotes towards  $3\pi/2$  at high modulation frequencies as evidenced in Fig. [2](#page-4-2) for a distributed-feedback tunable diode laser (DFB-TDL) operating near 1.47 μm (NEL Photonics). As will be shown in Sect. [3.2,](#page-6-0)  $\theta_{1f}$  signals are least sensitive to environmental distortions when  $\psi_{m,1} = 3\pi/2$ .

# <span id="page-4-0"></span>**2.3 Fitting of WMS-** $\theta_{1f}$  **signals**

The fitting routine used to fit a simulated scanned- $\theta_{1f}$  signal,  $\theta_{1f, \text{sim}}(t)$ , to a measured  $\theta_{1f}$  signal,  $\theta_{1f, \text{meas}}(t)$ , is similar to the least-squares WMS spectral ftting routine described in Goldenstein et al. [[16](#page-22-3)] but with several important differences. A fowchart describing the algorithm is shown in Fig. [3](#page-5-0). First, the user initializes guess values for the spectroscopic parameters associated with the Voigt lineshape profile  $(A, v_0, \Delta v_C)$  and a background phase,  $\delta$  [rad]. Note that  $\Delta v_{\rm D}$  can be calculated from temperature using Eq. [3,](#page-2-1) which in many cases is either well known or can be inferred by probing multiple absorption transitions of the same species. Hence,  $\Delta v_{\rm D}$  is generally a fixed value within the fitting routine. Next, these Voigt profle parameters are combined



<span id="page-5-0"></span>**Fig. 3** Flowchart illustrating the WMS- $\theta_{1f}$  spectral fitting routine

with the simulated or previously measured  $I_0(t)$  and  $v(t)$  to calculate the simulated transmitted intensity  $I_t(t)$  using Eq. [1.](#page-2-2) The resulting simulated  $I_t(t)$  is then lock-in filtered to determine the simulated  $\theta_{1f}$  signal ( $\theta_{1f, \text{sim}}(t)$ ) using Eqs. [12](#page-3-2) and [13](#page-3-3). The background phase,  $\delta$ , is a constant offset added to the simulated  $\theta_{1f}$  to account for variations in the timing of  $I_0(t)$  relative to the measured  $I_t(t)$  caused by variations in the trigger point of a data acquisition system.

The sum-of-squared errors (SSE) between the simulated and measured  $\theta_{1f}(t)$  signals is then computed. However, phase angles are modular (e.g. 0 and  $2\pi$  are identical), which can lead to errors in calculating the true SSE without appropriate corrections. For example, −3*𝜋*∕4 and 3*𝜋*∕4 are separated in phase space by only  $\pi/2$ , yet the standard definition of SSE would assume a separation of  $3\pi/2$ . Therefore, the typical defnition of SSE needs to be modifed to account for the modularity:

$$
SSE = \sum_{t} \{ \min[|\theta_{1f, \text{meas}}(t) - \theta_{1f, \text{sim}}(t) + (2\pi, 0, -2\pi)| \}]^{2}.
$$
\n(14)

By taking the minimum of  $|\theta_{1f,meas}(t) - \theta_{1f,sim}(t)|$  offset by  $2\pi$ , 0, or  $-2\pi$ , this expression ensures that only the true angular separation between  $\theta_{1f \text{.meas}}(t)$  and  $\theta_{1f \text{.sim}}(t)$  is used to compute SSE. This efect is illustrated in Fig. [4](#page-5-1) for an optically thick  $\theta_{1f}$  signal that exceeds  $\theta_{1f} = \pi$  and wraps into  $\theta_{1f} = -\pi$  near  $t = 4.75$  ms. Without the modular SSE, the ftting routine would not be able to account for the apparent discontinuity in phase angles and converge onto correct values for the spectroscopic parameters.

The parameter vector,  $\beta$ , composed of the Voigt profile spectroscopic parameters and  $\delta$ , is then numerically perturbed to compute (via one-sided fnite diferencing) the



<span id="page-5-1"></span>**Fig. 4** Example WMS- $\theta_{1f}$  signal (black) for an optically thick absorbing gas medium. Without appropriately accounting for the phase discontinuity near  $t = 4.75$  ms with a modified definition for sum-ofsquared-errors, the WMS- $\theta_{1f}$  spectral fitting routine would not be able to converge onto the best-ft signal shown in red

approximate gradient of the SSE with respect to  $\beta$ ,  $\nabla_{\beta}$ SSE. Stochastic Polak–Ribiere conjugate gradient-descent (SCGD)  $[25]$  $[25]$  $[25]$  is then used to iteratively update  $\beta$  until SSE reaches a minimum. Note that this optimization scheme is diferent than the Levenberg–Marquardt algorithm [\[26,](#page-22-11) [27](#page-22-12)] commonly used in the literature to search for the bestfit  $\beta$ . Although the convergence rate for Levenberg–Marquardt is generally faster than SCGD, in practice, SCGD was observed to be more numerically stable at converging onto optimal values, especially for heavily distorted signals. Additionally, SCGD allows the user to arbitrarily specify the cost function for minimization rather than being restricted to SSE.

# <span id="page-5-2"></span>**3** WMS- $\theta_{1f}$  signal distortion analysis **and laser‑tuning parameter optimization**

With the theoretical basis and practical ftting guidelines for WMS- $\theta_{1f}$  signals established in the preceding section, we now analyze and optimize WMS- $\theta_{1f}$ -based sensors for measurements in harsh environments. Integral to this analysis are direct comparisons against the well-established WMS-*nf*/1*f* and SDAS techniques used throughout the literature with the explicit goal of understanding the expected performance of WMS- $\theta_{1f}$  sensors for practical sensing applications.

It should be noted that the vast majority of TDLAS use cases involve the accurate determination of the mole fraction, temperature, and/or pressure of the absorbing species. All of these parameters directly factor into the integrated absorbance, *A*, as shown in Eq. [5.](#page-2-3) Hence, the current analysis focuses specifcally on quantifying the impact of environmental distortions on the resulting uncertainty in the best-ft *A*. Though uncertainties in best-fit  $\Delta v_{\rm C}$ ,  $v_{\rm 0}$ , and  $\delta$  were not evaluated, the methods used in this work can be generalized for such purposes.

# **3.1 Difculties with the standard WMS optimization approach**

The standard approach for optimizing the laser-tuning parameters of WMS-based sensors (referred to as the "standard optimization approach" or SOA) maximizes the expected peak WMS signal strength near the transition linecenter for a given measurement environment [\[8](#page-21-3), [17](#page-22-4), [18](#page-22-5)]. For example, the peak WMS-2*f*, -3*f*, and -4*f* signal responses occur when the modulation index,  $\mu$ , is set to 2.2, 3.35, and 4.25, respectively, for a single absorption transition described by a Voigt profle. This scheme is intuitive because it theoretically maximizes the signal-to-noise ratio (SNR) of the WMS signals relative to background distortions in  $I_t$  from the measurement environment. Since the amplitude of the WMS harmonics responds in proportion to the integrated absorbance, measurement precision would be maximized at these optimal setpoints. For spectrally resolved scanned-WMS signals, however, additional optimization factors must be considered as illustrated below.

Figure [5](#page-6-1) plots the peak-normalized linecenter signal responses for WMS- $\theta_{1f}$ , -2*f*/1*f*, -3*f*/1*f*, and -4*f*/1*f* as a function of  $\mu$  for an example absorption transition described by a Voigt profile with peak absorbance  $= 0.063$  and  $\Delta v_C/\Delta v_D = 10$ . Here, the response of WMS- $\theta_{1f}$  is defined as the phase diference between the maximum and minimum phases in a  $\theta_{1f}$  signal. Like what was observed in Goldenstein et al. [\[16\]](#page-22-3), the maximum signal responses for WMS-2*f*/1*f*,  $-3f/1f$ , and  $-4f/1f$  signals occur near  $\mu = 1.1, 1.6,$  and 1.9, respectively. In contrast, the maximum response for



<span id="page-6-1"></span>**Fig. 5** Peak-normalized signal responses for WMS- $\theta_{1f}$ , -2*f*/1*f*, -3*f*/1*f*, and  $-4f/1f$  as a function of modulation depth,  $\mu$  for a single absorption transition with peak absorbance of 0.063 and  $\Delta v_C / \Delta v_D = 10$ . The simulation assumes  $i_{m,1} \propto a_{m,1}$  similar to the behavior of real semiconductor lasers. Note that the optimal  $\mu$  for maximum  $\theta_{1f}$  signal response is 0

WMS- $\theta_{1f}$  occurs when  $\mu \to 0$ , a behavior first observed in Yang et al. [[14\]](#page-22-1). This presents several inconsistencies and weaknesses in the SOA:

- 1. The modulation index,  $\mu$ , can never be truly zero because  $H_{k>0} = 0$  when  $\mu = 0$  (see Eq. [11](#page-3-1)). As a result, based on Eq. [13](#page-3-3),  $\theta_{1f}$  = const. and does not respond to gas absorption.
- 2. Although the SOA maximizes the expected signal responses, it assumes that the noise amplitude is independent of laser-tuning parameters. It is not guaranteed that WMS signals are minimally sensitive to signal distortions within the measurement environment at the optimal values of  $\mu$  in Fig. [5.](#page-6-1)
- 3. The Allan variance of measured  $CO<sub>2</sub>$  mole fractions in the static cell experiments using the WMS- $\theta_{1f}$  method presented in Yang et al. [\[14\]](#page-22-1) decreased for all integration times with increasing  $\mu$ , which is opposite to the expected trend based on Fig. [5.](#page-6-1)

These difficulties suggest that the SOA fails to account for additional complexities of spectrally resolved scanned-WMS signals and its interaction with environmental distortions. Thus, to address these complexities, a new analysis and optimization scheme that directly considers the combined efects of spectral ftting and environmental distortions is needed.

## <span id="page-6-0"></span>**3.2 Additive and multiplicative signal distortions**

We now consider the effects of additive (e.g. optical emission and offsets in the detection and digitization electronics chain) and multiplicative (e.g. non-absorption transmission losses from beam-steering or particulate matter occlusion) signal distortions on the accuracy of WMS- $\theta_{1f}$  and -*nf*/1*f* measurements. The current analysis specifically evaluates the effects of broadband distortion signatures with bandwidth on the order of the modulation frequency,  $f<sub>m</sub>$ . This was done because it is well established both theoretically and experimentally that additive and multiplicative distortions with bandwidth less than  $f_{\rm m}/2$  can be corrected by WMS-*nf*/1*f* and WMS- $\theta_{1f}$ [\[18](#page-22-5), [28\]](#page-22-13). These broadband distortions are often the dominant source of error for WMS-based measurements in harsh environments (see [[12](#page-22-14), [29](#page-22-15)], for example) and, therefore, deserve signifcant attention. It should be noted that a similar analysis was performed for SDAS and its comparison against the various WMS techniques is included in "Appendix" and summarized in Sect. [3.4.](#page-14-1)

Rather than mathematically modeling how these distortions affect measurement uncertainty, it is sufficient and significantly more efficient and illustrative to numerically compute measurement uncertainty by perturbing the ideal transmitted intensity,  $I_t$ , with simulated distortions from a target measurement environment. The frequency content and temporal behavior of these signal distortions can vary widely depending on the measurement environment and experimental hardware. Thus, to maximize the generality of the current analysis, 1/*f* random pink noise was used to simulate the additive  $(\kappa(t))$  and multiplicative  $(\rho(t))$  distortion time series, which may be constructed as follows:

$$
\kappa(t) = \xi_{\kappa} \cdot \left[ (1/f\text{-filter}) \ast \mathcal{N}_t(0, 1) \right],\tag{15}
$$

$$
\rho(t) = \exp\left[\xi_{\rho} \cdot (1/f\text{-filter}) * \mathcal{N}_t(0, 1)\right],\tag{16}
$$

where  $\xi_{\kappa}$  and  $\xi_{\theta}$  are user-defined additive and multiplicative distortion amplitudes, respectively, and  $\mathcal{N}_t(0, 1)$  denotes a Gaussian white noise time series constructed by randomly sampling a normal distribution with zero mean and standard deviation = 1. The 1/*f*-filter transfer function,  $G_{1/f}(f)$ , has zero phase with frequency-dependent gain defned as follows:

$$
\left| G_{1/f}(f) \right| = \frac{1}{1 + \left( \frac{f}{f_{1/2}} \right)^p},\tag{17}
$$

where  $f_{1/2}$  is the frequency "elbow" corresponding to  $G_{1/f} = 1/2$  and kept at  $f_{1/2} = 0.1 f_m$  throughout this work to ensure that  $G_{1/f}$  asymptotically decays as a function of  $1/f<sup>p</sup>$  at the harmonics of the modulation frequency. *p* is the frequency roll-off exponent, which sets the relative impact of the distortion signals on the modulation harmonics.  $G_{1/f}$ is plotted as a function of  $f/f<sub>m</sub>$  in Fig. [6](#page-7-0) (log–log space) for various values of *p* with important frequencies labeled. Note that the choice of  $\xi_{\kappa}$  and  $\xi_{\theta}$  are inconsequential to the conclusions drawn from this work as long as they are small relative to  $I_0$  (to prevent exceedingly poor spectral fits) and kept constant when comparing the results of the various



WMS techniques against each other. The roll-off exponent, *p*, however, is an important parameter and is the subject of Sect. [3.2.3](#page-11-0).  $\kappa(t)$  and  $\rho(t)$  were then used to distort *I<sub>t</sub>*(*t*) to yield the distorted transmitted intensity,  $\widetilde{I_t(t)}$ , as follows:

$$
\widetilde{I_t(t)} = \rho(t) \cdot I_t(t) + \kappa(t). \tag{18}
$$

<span id="page-7-3"></span><span id="page-7-2"></span> $\widetilde{I_t(t)}$  was then passed through the lock-in filter and angle calculation operations (Eqs. [12](#page-3-2) and [13\)](#page-3-3) to determine the distorted  $\theta_{1f}$  signal,  $\widetilde{\theta_{1f}}$ . An example of the effects of such a distortion on a TDLAS experiment for an absorbing gas described by a single Voigt profle is shown in Fig. [7](#page-7-1) for the transmitted intensity (top panel) and best-fit  $\theta_{1f}$ signal and residual (middle and bottom panel) assuming  $\xi_{\kappa} = 0.1I_0$ ,  $\xi_{\theta} = 0.1$ , and  $p = 1$ . Note that the distortion introduced a 4.96% error in the best-ft integrated absorbance relative to the known value.

<span id="page-7-4"></span>The WMS- $\theta_{1f}$  and WMS-*nf*/1*f* spectral fitting routines were then used to compute best-ft values of *A* for the distorted signal, denoted as *Ã*, for a large collection (size *N*) of randomly distorted  $\widetilde{I_t(t)}$ . A sample-averaged error in the best-ft integrated absorbance can then computed as follows:



<span id="page-7-0"></span>**Fig. 6** Gain as a function of frequency normalized by the modulation frequency ((log-log space) for various values of the roll-off exponent, *p*, for the 1/*f* flter used to generate the stochastic additive and multiplicative distortion time series in Eqs. [15](#page-7-2) and [16](#page-7-3)

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<span id="page-7-1"></span>**Fig. 7** Example effect of  $1/f$  additive and multiplicative distortions on the transmitted intensity (top panel) and the resulting best-ft WMS- $\theta_{1f}$  signal and residuals (middle and bottom panels) for an absorbing gas described by a single Voigt profle. The distortion introduced a 4.96% error in the best-ft integrated absorbance relative to the known value

$$
\epsilon_A = \frac{1}{N} \sum_{k=1}^{N} \left| \frac{\widetilde{A}_k - A_{\text{known}}}{A_{\text{known}}} \right|,
$$
\n(19)

where *A*known is the user-defned value for *A* used in the simulated absorption spectrum. *N* was set to be 200 for the results presented in this article, which was observed to be sufficient to converge onto a statistically stable value for  $\epsilon_A$ .  $\epsilon_A$  can be interpreted as the true figure of merit for optimizing the laser-tuning parameters of any scanned-WMS technique with respect to additive and multiplicative distortions in the environment. Rather than optimizing the expected WMS response just near linecenter,  $\epsilon_A$  explicitly evaluates the combined efects of environmental distortions, laser-tuning parameters, and the spectral ftting routine on the results of a scanned-WMS measurement. Laser-tuning parameters should thus be selected such that  $\epsilon_A$  is at a global minimum.

#### **3.2.1 IM–FM phase delay optimization**

As noted in Sect. [2.2,](#page-2-0) WMS- $\theta_{1f}$  signals are sensitive to the linear phase delay between the intensity modulation (IM, described by the m-component of Eq. [6\)](#page-3-4) and the frequency modulation (FM, described by the m-component of Eq. [7\)](#page-3-5) waveforms, defined here as  $\psi_{m,1} - \phi_{m,1}$ . In particular, for operating conditions where mod $(\psi_{m,1} - \phi_{m,1}, \pi) = 0$ , the WMS- $\theta_{1f}$  signal is insensitive to gas absorption and cannot be used for spectroscopic measurement. Consequently,  $\epsilon_A$ for WMS- $\theta_{1f}$  signals is expected to depend strongly on the IM–FM phase delay.

Figure [8](#page-8-0) plots  $\epsilon_A$  for WMS- $\theta_{1f}$  as a function of the IM–FM phase delay for various modulation indices and at a fxed scan index,  $\sigma$ , equal to 9. All simulations assume fixed laser tuning, distortion, and absorption spectrum parameters as listed in Table [1](#page-8-1).  $f_{\text{cutoff}}$  here represents the cutoff frequency of the brick-wall low-pass flter used to compute the WMS harmonics. The range of IM–FM phase delays considered here was  $\pi \rightarrow 3\pi/2$  corresponding to the physical limits of injection current-tuned lasers (based on thermal lag argu-ments discussed in Sect. [2.2](#page-2-0)), with  $\pi$  corresponding to  $f_m \to 0$  and  $3\pi/2$  corresponding to  $f_m \to \infty$ .

As can be seen from Fig.  $8$ , the minimum sampleaveraged error in the best-ft integrated absorbance corresponds to an IM–FM phase delay of  $3\pi/2$  and its optimal



<span id="page-8-0"></span>**Fig. 8** Sample-averaged error in the best-ft integrated absorbance as measured with a least-squares best-fit of distorted WMS- $\theta_{1f}$  signals plotted as a function of the intensity-optical frequency modulation phase delay for various modulation indices (red, blue, and green). Note that for all modulation indices, the minimum error occurs when  $\psi_{m,1} - \phi_{m,1} = 3\pi/2$ , which corresponds to the IM–FM phase delay at infnite modulation frequency for injection current-tuned lasers. Mean error rate of distorted WMS-2 $f/1f$  signals is shown in black for  $\mu = 4$ to demonstrate the near independence of WMS-*nf*/1*f* measurement accuracy on the IM–FM phase delay

positioning is independent of the modulation index. More generally, the  $3\pi/2$  optimality criterion was verified to be independent of the characteristics of the distortion environment and laser-tuning parameters (data not shown for brevity). This demonstrates that it is beneficial to modulate the laser at as high of a frequency as possible to maximize the IM–FM phase delay and thus desensitize the sensor to additive and multiplicative distortions. However, above a phase delay of approximately  $5\pi/4$ ,  $\epsilon_A$  quickly asymptotes towards the minimum value and becomes nearly independent of the IM–FM phase delay; hence, there are diminishing returns for modulating at increasingly high frequencies. For the example laser used to generate Fig. [2,](#page-4-2) modulation frequencies above 200 kHz correspond to IM–FM phase delays that are in the asymptotic region of Fig. [8.](#page-8-0)

For WMS- $n f/1 f$  measurements,  $\epsilon_A$  is expected to be nearly independent of the IM–FM phase delay because the analytical expressions for  $|R_{nf}/R_{1f}|$  are weakly dependent on  $\psi_{m,1} - \phi_{m,1}$  [\[17\]](#page-22-4). This is demonstrated in the black curve in Fig. [8](#page-8-0), which plots  $\epsilon_A$  for WMS-2*f*/1*f* measurements as a function of  $\psi_{m,1} - \phi_{m,1}$  for  $\mu = 4$  and  $\sigma = 9$ . As can be seen,

<span id="page-8-1"></span>**Table 1** Laser tuning, distortion, and absorption spectrum parameters used to numerically compute  $\epsilon_A$  in Figs. [8](#page-8-0), [9,](#page-9-1) and [10](#page-10-0)

$\mathcal{L}$					$\sigma$
$J_{\rm m}$		∍∩	אכ	$J_1/2$	
$100$ kHz	1 kHz	0.01	$0.01I_0$	$0.1 f_{\rm m}$	Q
$f_{\rm cutoff}$			$\Delta v_C$	$\Delta v_{\rm D}$	$V_0$
$10f_{\rm s}$		$0.05$ cm <sup>-1</sup>	$0.1 \text{ cm}^{-1}$	$0.01 \text{ cm}^{-1}$	$\mathcal V$

 $\epsilon_A$  remains nearly constant as a function of  $\psi_{m,1} - \phi_{m,1}$ . In practice, the variation of  $\psi_{m,1} - \phi_{m,1}$  with modulation frequency for specifc lasers is strongly dependent on the laser architecture and intrinsic properties. For applications where the IM–FM phase delay cannot reach the asymptotic region of Fig. [8](#page-8-0) with phase delay greater than  $5\pi/4$  (e.g. for systems that are unable to modulate the lasers at sufficiently high frequencies or for lasers with exceptionally low thermal mass), WMS-*nf*/1*f* may be the preferred measurement technique relative to WMS- $\theta_{1f}$  since its accuracy is nearly independent of  $\psi_{m,1} - \phi_{m,1}$ .

## <span id="page-9-0"></span>**3.2.2 Scan and modulation index optimization for**  $p = 1$

We now optimize the scan and modulation indices of the various WMS methods under the infuence of additive and multiplicative distortions for the specifc case where the frequency roll-off exponent  $(p \text{ in Eq. 17})$  equals unity. This value of *p* was chosen because such distortion signatures are commonly found in nature as the limiting statistical distribution of diverse stochastic processes [\[30](#page-22-16)]. Despite this choice, the methods used in this work can be used to assess the performance of WMS sensors for arbitrary values of *p*.

Similar to the IM–FM phase delay optimization scheme,  $\epsilon_4$  can be evaluated for a variety of scan and modulation indices ( $\sigma$  and  $\mu$ , respectively) for a given absorption spectrum and distortion environment to determine optimal values for  $\mu$  and  $\sigma$ , denoted as  $\mu^*$  and  $\sigma^*$ , respectively. We first restrict the analysis specifically on the impact of  $\mu$  while keeping  $\sigma$ constant, similar to what is considered in the standard optimization approach. The results are shown in Fig. [9](#page-9-1), which plots  $\epsilon_A$  as a function of  $\mu$  for WMS- $\theta_{1f}$  and WMS-*nf*/1*f* with  $n \in [2, 3, 4]$ . All simulations assume the same low-frequency scan, pink-noise, and gas absorption parameters listed in Table [1.](#page-8-1) Additionally, laser-tuning parameters in Eqs. [6](#page-3-4) and



<span id="page-9-1"></span>**Fig. 9** Sample-averaged error in the best-ft integrated absorbance as a function of modulation index for WMS- $\theta_{1f}$  and various WMS-*nf*/1*f* signals distorted by random 1/*f* additive and multiplicative noise

[7](#page-3-5) used to simulate the results were experimentally determined values from a DFB-TDL operating near 2.017 μm (Nanoplus) with  $f_m = 100$  kHz and  $f_s = 1$  kHz. For the  $\theta_{1f}$ results,  $\psi_{m,1} - \phi_{m,1}$  was measured to be 1.36 $\pi$  at  $f_m = 100$ kHz, which is within the asymptotic region of Fig. [8.](#page-8-0)

Several trends can be observed and conclusions drawn from Fig. [9:](#page-9-1)

- 1. Regardless of  $\mu$ , WMS- $\theta_{1f}$  exhibits improved measurement accuracy over the various WMS-*nf*/1*f* methods. In general, the accuracy of WMS-*nf*/1*f* worsens as *n* increases. This trend agrees with the superior Allan variance of the WMS- $\theta_{1f}$  or WMS- $R_{1f}/Y_{1f}$  measurements shown in Yang et al. [[14\]](#page-22-1) and Upadhyay et al. [\[15](#page-22-2)] when compared to concurrent WMS-2*f*/1*f* measurements.
- 2.  $\epsilon_A$  for  $\theta_{1f}$  measurements decreases as a function of  $\mu$ . This behavior contrasts with the SOA of WMS- $\theta_{1f}$ shown in black in Fig. [5,](#page-6-1) where  $\epsilon_A$  for WMS- $\theta_{1f}$  signals are expected to increase as a function of  $\mu$  due to the lower  $\theta_{1f}$  signals response at higher values of  $\mu$ . Likewise for the various WMS-*nf*/1*f* methods, besides local minima near  $\mu = 3, 4.5,$  and 5.3 for WMS-2*f*/1*f*, -3*f*/1*f*, and -4 $f/1f$ , respectively,  $\epsilon_A$  decreases as a function of  $\mu$ in general.

Given the unexpected behavior of  $\epsilon_A$  versus  $\mu$  compared to the expectation from the standard optimization approach, it is natural to question whether the choice of  $\sigma$  has an influence on  $\epsilon_A$ . Past researchers provided heuristic  $\sigma$ -selection guidelines  $[16]$  that only suggest selecting a sufficiently large  $\sigma$  to spectrally resolve the WMS- $n f/1 f$  signal lobes but do not assess how  $\sigma$  impacts measurement accuracy. Fig-ure [10](#page-10-0)a shows the numerically calculated  $\epsilon_A$  for WMS- $\theta_{1f}$ color-mapped onto a matrix of  $\mu$ – $\sigma$  pairs, with black and dark blue colors representing better measurement accuracy. Note the existence of multiple valleys of optimal  $\mu$  and  $\sigma$ with troughs best described by the lines  $\sigma = 1.05\mu + 1.17$ (labeled as "Valley 1") and  $\sigma = 0.93\mu - 0.73$  ("Valley 2").  $\mu$ <sup>– $\sigma$ </sup> pairs along these troughs may be considered as "good" choices" for minimizing  $\theta_{1f}$  sensitivity to additive and multiplicative signal distortions with roll-off exponent, *p*, equal to 1. The global minimum for  $\epsilon_A$  exists within Valley 2 at  $(\mu^*, \sigma^*)$  = (6.4, 5.2). However, since these  $\mu$  and  $\sigma$ values are relatively large and difficult to achieve except for absorption spectra with small linewidths (e.g. gases at low pressures) or exceptionally wide-tuning lasers, it is more practically relevant to consider the behavior of  $\epsilon_A$  at lower  $\mu$  and  $\sigma$  values. Figure [10](#page-10-0)b shows a cutaway of the region between  $\mu = 2 \rightarrow 5$  and  $\sigma = 3 \rightarrow 7$  bisected by Valley 1. It can be seen that there exists a  $(\mu^*, \sigma^*)$  = (4.5, 5.6) pair that locally minimizes  $\epsilon_A$  with more reasonably achievable scan and modulation depths. For practical environments where laser-tuning range relative to absorption linewidth is limited,



<span id="page-10-0"></span>**Fig. 10** Sample-averaged error in the best-ft integrated absorbance color-mapped onto a matrix of scan and modulation index pairs for **a**, **b** WMS- $\theta_{1f}$ , **c** 2*f*/1*f*, and **d** 3*f*/1*f* signals distorted by random 1/*f* addi-

tive and multiplicative noise with  $p = 1$ . "Valleys" of optimal  $\mu-\sigma$ pairs are highlighted in dotted white lines. Global optima are marked with red diamonds

it is possible to select slightly suboptimal  $\mu$ – $\sigma$  pairs along Valley 1 such as (2.5, 3.8) without signifcant loss in accuracy. Repeated simulations indicated that the  $\mu-\sigma$  trough equation and the optimal  $(\mu^*, \sigma^*)$  pair were insensitive to IM–FM phase delays,  $\xi_{\rho}$ ,  $\xi_{\kappa}$ , and peak absorbances at a constant  $p = 1$  (results not shown for brevity).

Similar analyses can be performed for distorted WMS-2*f*/1*f* and -3*f*/1*f* signals, whose results are shown in Fig. [10](#page-10-0)c, d, respectively. Note that the  $\epsilon_A$  color-scales for these two fgures were kept consistent with the color scale used in Fig. [10](#page-10-0)a; hence,  $\epsilon_A$  values exceeding 0.3% were displayed as uniformly dark red. Several conclusions can be drawn from comparing  $\epsilon_A$  among the different WMS techniques.

1. Like in Fig. [10](#page-10-0)a, there exist multiple valleys of optimal  $\mu$  and  $\sigma$ . For WMS-2*f*/1*f*, the troughs of these valleys are described by the equations  $\sigma = 1.04\mu + 1.08$ and  $\sigma = 0.2\mu + 0.5$  for Valleys 1 and 2, respectively. The global optimum lies within Valley 2 at  $(\mu^*, \sigma^*)$  = (6.4, 1.7). For WMS-3*f*/1*f*, the troughs are described by  $\sigma = 1.24\mu - 0.16$  and  $\sigma = 0.52\mu - 0.75$  for Valleys 1 and 2, respectively, with the global optimum lying in Valley 2 at  $(\mu^*, \sigma^*)$  = (7.8, 3.5).

2. Although the global minima in  $\epsilon_A$  for WMS-2*f*/1*f* and  $-3f/1f$  are comparable to the minimum for  $\theta_{1f}$ , the WMS*nf*/1*f* strategies generally do not perform as well as  $\theta_{1f}$ at low modulation and scan indices. In practical harsh environments with pronounced lineshape broadening effects, lasers are often incapable of reaching the  $\mu$  and *𝜎* needed to reach the optimal valleys of the WMS-*nf*/1*f* techniques, suggesting that WMS- $\theta_{1f}$  would be the preferred sensing technique for distortion-heavy environments or where maximum measurement rate is important.

It is difficult to identify precise reasons for the trends observed in Fig. [10](#page-10-0) in the absence of a mathematical model. It is, however, possible to elucidate qualitative intuitions that can explain key features. For example, to explain the existence of the valleys of optimal  $\mu$ – $\sigma$  pairs, we propose the following line of reasoning.



<span id="page-11-1"></span>**Fig. 11** Frequency spectrum of the transmitted intensity of a scannedwavelength WMS experiment near the frst harmonic with (left panel)  $(\mu, \sigma) = (2.5, 3.8)$  and (right panel)  $\sigma$  doubled to 7.6 while keeping  $\mu$ constant. Note the weakening of the harmonic sidebands relative to the noise floor with increasing  $\sigma$ 

- 1. Assuming a constant  $\mu$ , for low values of  $\sigma$ , the laser does not sample over a sufficient fraction of the absorption spectrum during each scan cycle, leading to an insufficiently constrained fitting problem that leads to large errors. Specifically, at low values of  $\sigma$ , *A* and  $\Delta v_C$ become less distinct in their efects on simulated WMS signals, resulting in best-ft *A* that are susceptible to signal distortions. This explains the generally high  $\epsilon_A$ values at low  $\sigma$  in Fig. [10.](#page-10-0)
- 2. Conversely, for exceedingly large values of  $\sigma$ , the laser tunes across the absorption spectrum very quickly during each scan cycle, thus requiring larger amplitudes at high sideband frequencies to represent the signals. Figure [11](#page-11-1) illustrates this effect for the sidebands surrounding the frst harmonic of the transmitted intensity of a WMS experiment. As can be seen, when  $\sigma$  is doubled from 3.8 (left) to 7.6 (right) at constant  $\mu = 2.5$ , the harmonic sidebands become weaker on average and more widely distributed in frequency (not visible here but demonstrated in Fig. 7 of [[17](#page-22-4)]), resulting in lower average SNR in the sidebands. Since these sidebands encode information needed for WMS spectral ftting, the lower SNR is expected to contribute to greater  $\epsilon_A$ .
- 3. As a result of the previous two arguments, for each  $\mu$ , a certain optimal  $\sigma$  exists that is both large enough to suffciently constrain the WMS spectral ftting routine and small enough to maintain good SNR in the harmonic sidebands relative to the noise floor.
- 4. In general, as  $\mu$  increases at a constant  $\sigma$ , the width of the WMS signal lobes in the time-domain increases. This requires a corresponding increase in the minimum  $\sigma$  needed to fully resolve the WMS signal lobes and sufficiently constrain the WMS spectral fitting routine. Together with the preceding observations, the net result



<span id="page-11-2"></span>**Fig. 12** Simulated WMS-2*f*/1*f* signals over a single up-scan for a range of  $\mu$ – $\sigma$  pairs obeying the two  $\mu$ – $\sigma$  valley trough equations observed in Fig. [10c](#page-10-0)

is the existence of valleys of optimal  $\mu$ – $\sigma$  pairs exhibiting positive slope as evidenced in Fig. [10](#page-10-0).

It is also possible to qualitatively explain the existence of multiple optimal  $\mu$ – $\sigma$  valleys in Fig. [10](#page-10-0). Figure [12](#page-11-2) plots simulated WMS-2*f*/1*f* signals over a single up-scan for a range of  $\mu$ – $\sigma$  pairs evaluated along the optimal trough equations describing Valley 1 (left) and Valley 2 (right) observed in Fig. [10](#page-10-0)c. As can be seen, WMS-2*f*/1*f* signals along Valley 1 show the three canonical WMS-2*f*/1*f* lobes while signals along Valley 2 show only the central lobe near transition linecenter. This observation applies equally to WMS- $\theta_{1f}$  and WMS-3 $f/1f$  signals along their respective  $\mu-\sigma$  valleys (not shown). It is, therefore, reasonable to conclude that the various  $\mu$ – $\sigma$  trough equations optimize the laser tuning parameters for diferent numbers of WMS lobes observed in the measured signal.

## <span id="page-11-0"></span>**3.2.3** Dependence on the frequency roll-off exponent, *p*

Section [3.2.2](#page-9-0) demonstrated that WMS- $\theta_{1f}$  measurements are generally less sensitive to additive and multiplicative distortions assuming that the frequency content of the distortions rolls off as a function of  $1/f$ . However, as the frequency roll-off exponent,  $p$ , increases, the impact of the distortions on the  $n > 1$  harmonics of  $f_m$  become weaker relative to the impact on the frst harmonic. Therefore, it is expected that there exists some value of *p* at which WMS-*nf*/1*f* measurements are more accurate than WMS- $\theta_{1f}$ . Similarly, as  $p \to 0$ , the distortion signal approaches white noise and equally impacts the signals of all WMS harmonics. Since most of the absorption-induced frequency content in  $I_t(t)$  is concentrated near the first harmonic, it is expected that WMS- $\theta_{1f}$ will perform better than *nf*/1*f* at low values of *p*.

Figure [13](#page-12-1) confrms the preceding arguments. Here, the accuracy of the various WMS techniques were evaluated



<span id="page-12-1"></span>**Fig. 13** Sample-averaged error in the best-ft integrated absorbance,  $\epsilon_A$ , as a function of the frequency roll-off exponent, *p*, for WMS- $\theta_{1f}$ and various WMS-*nf*/1*f* signals distorted by random 1/*f*<sup>*p*</sup> additive and multiplicative noise. Results were normalized by the error in WMS- $\theta_{1f}, \epsilon_A |_{\theta_{1f}}$ , at each value of *p* 

as a function of *p* assuming the same simulated absorption spectrum and distortion amplitudes listed in Table [1.](#page-8-1)  $\sigma$  and  $\mu$  for each WMS technique were set to the optimal values as calculated for  $p = 1$  using the methods in Sect. [3.2.2.](#page-9-0) The resulting  $\epsilon_A$  values were then normalized by the calculated  $\epsilon_A$  for WMS- $\theta_{1f}$  and plotted in Fig. [13](#page-12-1) as a function of *p*. As can be seen, below  $p = 1.25$ , WMS- $\theta_{1f}$  performs better than WMS- $n f/1 f$ , consistent with the  $p = 1$  results in Sect. [3.2.2.](#page-9-0) Above  $p = 1.25$ , the greater roll-off exponent desensitizes WMS-2*f*/1*f* measurements to the frequency content of the distortions, allowing second harmonic detection to perform better than WMS- $\theta_{1f}$ . WMS-3*f*/1*f* and -4*f*/1*f* never perform better than  $-2f/1f$  and do not exceed  $\theta_{1f}$  performance until  $p > 2.1$ . It should be noted that the specific intersection points in the roll-off exponent shown in Fig. [13](#page-12-1) are specifc to the selected laser-tuning parameters and may vary as a function of  $\mu$ ,  $\sigma$ , and distortion amplitudes. However, the general trends shown in Fig. [13](#page-12-1) are representative of all conditions.

In practice, the true value of *p* depends strongly on the frequency-domain characteristics of the distortion sources within the measurement environment. For example, experiments in non-isothermal jet flows  $[31]$  $[31]$  show that the density fuctuation spectrum (and by extension, the beam-steering spectrum) within the inertial subrange of the potential core decays with  $p = 3.4$ , which is well outside the region in Fig. [13](#page-12-1) where WMS- $\theta_{1f}$  would be the preferred measurement technique. Conversely, TDLAS distortions within the combustion annulus of a rotating detonation engine—a representative environment where additive and multiplicative signal distortions are strong and diverse in source—were found to exhibit a  $p = 1.1$  roll-off exponent [\[32](#page-22-18)], indicating that WMS- $\theta_{1f}$  may be the preferred measurement technique in such an environment. As a result, it is important for sensor designers to assess the expected frequency-dependence of the distortion characteristics within a measurement environment to select an appropriate WMS detection scheme. Practically, one can characterize the additive and multiplicative distortion signatures within a measurement environment by measuring the transmitted intensity of a laser operating at constant intensity and at a non-absorbing wavelength. The information gathered from these initial experiments can guide the selection of the optimal WMS technique and lasertuning parameters.

#### <span id="page-12-0"></span>**3.3 Wavelength‑dependent signal distortions**

In some environments, wavelength-dependent distortions resulting from étalon cavities can be difficult to avoid and are often major sources of uncertainty in practical TDLAS experiments. Additionally, these distortions often vary with time due to vibration- or temperature-induced changes in alignment and can signifcantly impact the long-term stability of laser-based sensors. Therefore, it is important to assess the impact of these distortions on the accuracy of WMSbased measurements. A comparison with SDAS-based sensors was also performed and is included in "Appendix" and summarized in Sect. [3.4](#page-14-1).

Étalon-induced wavelength-dependent distortions generally manifest as periodic offsets in  $\alpha_{\nu}$ . To first order, these offsets may be approximated as sinusoidal distortions to the spectral absorbance with some free spectral range,  $T_D$  (cm<sup>-1</sup>), and peak offset,  $\delta_{\nu}$  (cm<sup>-1</sup>), relative to the transition linecenter:

$$
\widetilde{\alpha}_{v} = \alpha_{v} + \xi_{\text{etln}} \cos \left( \frac{2\pi}{T_{\text{D}}} \left( v - \left( v_{0} + \delta_{v} \right) \right) \right),\tag{20}
$$

where  $\tilde{\alpha}_{\nu}$  is the distorted spectral absorbance and  $\xi_{\text{eth}}$  is the étalon distortion amplitude.  $\tilde{\alpha}$  was then used to simulate a distorted  $I_t$ , which was spectrally fitted to determine the best-ft integrated absorbance, *Ã*. An example of this distortion and its effect on the best-fit WMS- $\theta_{1f}$  signal is shown in Fig. [14](#page-13-0) for an absorption spectrum described by  $(A, \Delta v_C, \Delta v_D, \tilde{v}) = (0.05, 0.1, 0.01, \overline{v})$  distorted with  $\xi_{\text{etln}} = 0.1 \alpha_{\text{max}}$ ,  $\delta_v = 0$ , and  $T_D = 0.4 \text{ cm}^{-1}$ .  $\mu$  and  $\sigma$  for this example were set at 3.64 and 5.45, respectively. As can be seen, the wavelength-dependent distortion to the absorption spectrum introduced a 4.26% error in the best-ft integrated absorbance relative to the true value.

The error of the distorted best-fit integrated absorbance depends, in general, on  $T_D$ ,  $a_{m,1}$ ,  $a_{s,1}$ , FWHM of the absorption transition,  $\xi_{\text{eth}}$ , and  $\delta_{\nu}$ . To simplify the analysis,  $a_{s,1}$  and  $\xi_{\text{eth}}$  were kept constant and  $T_D$  was nondimensionalized with  $a_{m,1}$  to form  $\eta = T_D/(2a_{m,1})$ , identical to the scheme in [[16\]](#page-22-3). Like the definition for  $\epsilon_A$ , we similarly defne an average error in the integrated absorbance due to



<span id="page-13-0"></span>**Fig. 14** Example efect of wavelength-dependent distortion on the spectral absorbance (top panel) and the resulting best-fit  $WMS- $\theta$ <sub>1f</sub>$ signal and residuals (middle and bottom panel) for an absorbing gas described by a single Voigt profle. The distortion introduced a 4.26% error in the best-ft integrated absorbance relative to the true value

wavelength-dependent distortions,  $\epsilon_{A,\text{etln}}$ , for a given  $\eta$  and  $\mu$  pair as follows:

$$
\epsilon_{A,\text{etln}} = \frac{1}{N} \sum_{k=1}^{N} \left| \frac{\widetilde{A}(\delta_{\nu} = kT_D/N) - A_{\text{known}}}{A_{\text{known}}} \right|,
$$
\n(21)

where *N* covers a range of individual  $\delta_{\nu}$  used for computing the average (100 was used in this work, sufficient to converge

onto stable values of  $\epsilon_{A,\text{eth}}$ ). This analysis extends on the methods shown in Goldenstein et al. [[16\]](#page-22-3). Specifcally, this analysis computes the étalon-induced error in the best-ft integrated absorbance rather than just the sum-of-squarederrors in the WMS-*nf*/1*f* signal. In addition, this analysis extends the assessment to the full range of  $\delta_{\nu}$ , supplementing the work of  $[16]$  $[16]$  at just  $\delta_{\nu} = 0$ .

Figure [15](#page-13-1) shows how  $\epsilon_{A,\text{eth}}$  varies as a function of  $\eta$  for various  $\mu$  for (a) WMS- $\theta_{1f}$  and (b) -2*f*/1*f*. Results are shown for a single Voigt profile described by  $(A, \Delta v_C, \Delta v_D, v_0) = (0.05, 0.1, 0.01, \overline{v})$ .  $\sigma$  and  $\xi_{\text{eth}}$  were set at 3.8 and 1% of the peak absorbance, respectively. Again, several trends can be observed from these results.

- 1. Similar to the observations in [\[16\]](#page-22-3),  $\epsilon_{A,\text{eth}}$  approaches zero when  $\eta$  approaches zero or infinity for all of the WMS techniques considered here. Also,  $\epsilon_{A,\text{eth}}$  decreases in general with increasing  $\mu$ .
- 2. The primary error band, representing the qualitative range of  $\eta$  in which  $\epsilon_{A,\text{eth}}$  is greatest, is highlighted with black arrows in Fig. [15](#page-13-1). As can be seen for WMS- $\theta_{1f}$ (Fig.  $15a$ ), the primary error band is wider in  $\eta$  than for WMS-2*f*/1*f* (Fig. [15](#page-13-1)b). However, the maximum errors for  $\theta_{1f}$  within this band are less than the maximum errors for WMS-2 $f/1f$ . Additionally,  $\theta_{1f}$  is more forgiving for distortions with small  $\eta$  (e.g. low free-spectral range étalons in the measurement environment) than WMS-2*f*/1*f*.
- 3. Results for higher harmonic WMS-*nf*/1*f* signals (e.g. WMS-3*f*/1*f* or -4*f*/1*f*, not shown here for brevity) indicate that the width of the primary error bands decreases with increasing *n*. However,  $\epsilon_{A,\text{eth}}$  within the primary error band worsens with increasing *n*, indicating less forgiveness at higher detection harmonics for wavelengthdependent distortions with sub-optimal *η*.



<span id="page-13-1"></span>**Fig. 15** Wavelength-dependent-distortion-induced errors in the best-fit integrated absorbance for **a** WMS- $\theta_{1f}$  and **b** -2*f*/1*f*. Distortion amplitude was 1% of the peak absorbance. Each curve represents a diferent value of modulation index

From these observations, we can conclude that WMS- $\theta_{1f}$ qualitatively performs just as well as typical WMS-*nf*/1*f* methods under wavelength-dependent distortions. For some low- $\eta$  measurement environments, WMS- $\theta_{1f}$  may even be the preferred measurement technique for maximum accuracy in the best-ft integrated absorbance. It should be noted that the results for optimizing WMS-based sensors against étalon noise may confict with the optimization criteria for additive and multiplicative distortions discussed in Sect. [3.2.](#page-6-0) The authors again emphasize the need for sensor designers to appropriately characterize and assess the relative impact of the various sources of signal distortion in the measurement environment to select optimal laser tuning parameters.

# <span id="page-14-1"></span>**3.4 Comparisons with scanned‑wavelength direct absorption spectroscopy**

Scanned-wavelength direct-absorption spectroscopy (SDAS) is a TDLAS technique where the laser source is periodically wavelength-tuned across an absorption transition. The measured transmitted intensity is used to directly calculate the absorption spectrum via Eq. [1](#page-2-2), which can be ftted with absorption lineshapes to infer spectroscopic parameters not unlike the spectral ftting strategy in Sect. [2.3](#page-4-0). Due to its simplicity in implementation and ease of real-time data visualization and troubleshooting as compared to WMSbased techniques, SDAS is, in general, the most widely used TDLAS sensing technique. As a result, there is considerable need to understand and quantify the relative performance of SDAS and WMS to justify the added complexity and overhead needed to implement WMS-based strategies. Fortunately, the analysis techniques described in the preceding sections can be generalized to SDAS, thereby allowing direct comparisons between SDAS and WMS.

Details related to the SDAS signal distortion analysis may be found in "Appendix". Key fndings are summarized below.

- 1. There exists an optimal modulation index,  $\mu^* = 5.6$ , that minimizes  $\epsilon_A$  for SDAS measurements when subjected to stochastic additive and multiplicative noise (see Fig. [22](#page-19-0)). To be consistent with the conventions used for WMS, modulation index for SDAS is defned as the ratio between the optical frequency scanning range to the half-width half-maximum of the absorption feature.
- 2. When subjected to the identical  $p = 1$  additive and multiplicative distortions as used in Sect. [3.2.2,](#page-9-0) SDAS was found to perform signifcantly worse than the various WMS-based sensors (see Fig. [22](#page-19-0)a). This agrees with intuition because the frequency content of the distortions are concentrated at low frequencies near the harmonics of the SDAS measurements. Since SDAS does not flter

or correct for these distortions, it is expected that the performance of SDAS will be more heavily impacted.

- 3. However, for  $p = 0$  additive and multiplicative distortions (i.e. white noise) where the distortions equally impact the SDAS and WMS harmonics, SDAS was found to perform as well as WMS- $\theta_{1f}$  and significantly better than either WMS-2*f*/1*f* or -3*f*/1*f* (see Fig. [22](#page-19-0)b). This suggests that, for certain environments with distortion characteristics that exhibit a roll-off exponent near zero (e.g. for quiescent environments where electronics noise is dominant), it may be detrimental to use WMS instead of SDAS in terms of sensor accuracy.
- 4. It is important to note that the intensity normalization and lock-in fltering operations inherent in WMS-based strategies are able to correct additive and multiplicative distortions with bandwidth below *f*m∕2. Since SDAS cannot intrinsically correct for these low-frequency distortions, for applications where strong, pure-tone oscillations are expected (e.g. for in situ monitoring of turbomachinery), SDAS measurement accuracy can be signifcantly worse than WMS measurement accuracy (see Fig. [23](#page-20-0)).
- 5. The accuracy of SDAS sensors is more strongly afected by short-period étalons (see Fig. [24](#page-20-1)), which is opposite the behavior of WMS- $\theta_{1f}$  sensors as demonstrated in Fig. [15](#page-13-1)a. This suggests that, for environments where wavelength-dependent distortions are the dominant source of measurement uncertainty, there are distinct ranges of  $\eta$  at which either WMS- $\theta_{1f}$  or SDAS is the preferred measurement technique.

These fndings show that SDAS is not always the inferior measurement technique when compared to WMS. This again reinforces the need to characterize the expected distortion environment for a target application and to use the analysis techniques introduced in this work in order to select the optimal sensing strategy and laser tuning parameters.

# **4 Experimental validation and demonstration**

Two sets of experiments were performed both to validate the simulation-based results shown in Sect. [3](#page-5-2) (using a static gas cell) and to provide a demonstration of the utility of WMS- $\theta_{1f}$  for practical diagnostics (using a methane-air flame).

## <span id="page-14-0"></span>**4.1 CO<sub>2</sub> static cell with high-speed jet distortions**

Section [3.2](#page-6-0) numerically demonstrated the superior performance of scanned-WMS- $\theta_{1f}$  over WMS- $n f / 1 f$  with respect to best-ft integrated absorbance accuracy under additive and multiplicative signal distortions below certain values



<span id="page-15-0"></span>Fig. 16 Illustration of the experimental setup for the  $CO<sub>2</sub>$  detection experiments. The static  $CO<sub>2</sub>$  gas cell provided a constant gas absorption environment while the high-speed air jet and polarizer perturbed the transmitted intensity to simulate stochastic additive and multiplicative signal distortions

of  $p$ . To validate these findings, a series of  $CO<sub>2</sub>$  detection experiments were performed using a static optical gas cell and high-speed jet-induced optical perturbations. Figure [16](#page-15-0) shows the experimental setup used for these experiments with major elements labeled. A fiber-coupled DFB-TDL (Nanoplus) operating near 2.017 μm was used to target the  $P(24)$  absorption transition of the  $CO<sub>2</sub>$  $v = (2, 0, 1) \leftarrow (0, 0, 0)$  combination absorption band with vacuum linecenter near 4957.08 cm<sup>−</sup><sup>1</sup> . Light from the laser was collimated (Thorlabs F028APC-2000) and transmitted through a 21.44-cm-long optical gas cell capped with wedged sapphire windows. The laser was controlled and modulated with an Arroyo Instruments 6305 controller and a National Instruments USB-6353 waveform generator, respectively. Photodetector voltage and waveform generator timing pulses were collected and digitized by a Pico Technologies PicoScope 4824 at a rate of 10 MS s<sup>-1</sup>. Mixtures of  $CO<sub>2</sub>$  diluted in N<sub>2</sub> were prepared manometrically within the gas cell using high-purity  $CO<sub>2</sub>$  and  $N<sub>2</sub>$  gas bottles (Praxair) and a barometer (Setra).

A high-speed air jet fed with a 75 psig reservoir was mounted immediately downstream of the gas cell to generate turbulence-induced density gradients that perturb the path of the transmitted light. Beam-steering experiments with the laser beam placed near the jet exit suggested a preferred Kelvin–Helmholtz instability mode near 52 kHz. This preferred mode quickly transitioned into fully developed turbulence beyond two jet diameters away from the exit plane (see [\[31](#page-22-17)]), thereby approximating stochastic broadband multiplicative distortions in the transmitted intensity. The transmitted laser intensity was collected by an anti-reflection-coated  $f = 30$ mm  $CaF<sub>2</sub>$  lens and measured by an InGaAs photodetector (Thorlabs PDA10D). Because the distortion amplitude increases as the detector was moved away from the focus of the collection lens, the detector chip was not placed at the focus to control the amplitude of the air-jet distortions. Finally, a rotatable linear polarizer (Thorlabs LPMIR050) was placed between the laser and the static cell to variably



<span id="page-15-1"></span>**Fig. 17** Measured **a** time series and **b** frequency spectra of the transmitted intensity of a constant-intensity laser source with (black) and without (blue) distortions generated within the turbulent wake of a high-speed air jet. Power-law best-fts to the frequency spectra are shown in green and red

attenuate the incident laser intensity. This has the efect of boosting both the detection and electronics noise relative to the amplitude of the laser signal, thereby simulating an environment with stochastic additive signal distortions.

Figure [17](#page-15-1) shows a sample photodetector (a) time series and (b) frequency spectrum (in log–log space) when the laser is operated at a constant intensity without (black) and with (blue) the air jet distortions. Both the time series and frequency spectra were vertically shifted in the fgure for clarity. Without the jet distortions, the detection and digitization electronics noise exhibits a  $p = 0.22$  decay exponent based on power-law ftting (shown in green in Fig. [17b](#page-15-1)) of the spectrum. Similarly, the best-ft to the jet distortions (shown in red in Fig. [17](#page-15-1)b) shows stochastic frequency content ranging from  $f = 0-1000$  kHz with  $f_{1/2} = 278$  kHz and  $p = 3.3$ . Note that it is possible to subject the various detection harmonics to diferent efective values of *p* by judiciously selecting appropriate modulation frequencies. For example, by choosing  $f_m = 50$  kHz, the detection harmonics

up to  $n = 4$  will experience essentially white noise ( $p = 0$ ) distortions from the jet. Additionally, Fig. [17](#page-15-1)a essentially provides experimentally measured additive  $(\kappa(t))$  in Eq. [15\)](#page-7-2) and multiplicative  $(\rho(t)$  in Eq. [16\)](#page-7-3) distortion time series. With this information, identical  $\kappa(t)$  and  $\rho(t)$  signals can be used to distort multiple ideal  $I_t(t)$  signals. This way, it is possible to compare the results of data collected from distinct  $\mu$  $-\sigma$  pairs across multiple hours without having to adjust for the temporal variability of the jet or electronics distortion spectra.

Figure [18](#page-16-0) shows  $\epsilon_A$  color-mapped onto a range of  $\mu-\sigma$ pairs for (a) WMS- $\theta_{1f}$ , (b) -2*f*/1*f*, (c) -3*f*/1*f*, and (d) -4*f*/1*f*. These results were generated for  $f_m = 100$  kHz,  $f_s = 1$ kHz, and a static cell containing  $9.59\%$  CO<sub>2</sub> by mole in N<sub>2</sub> balance at a constant room temperature of 297 K and pressure of 1 atm. The IM–FM phase delay for this modulation frequency was measured to be  $1.36\pi$ , well within the asymptotic region of Fig. [13](#page-12-1). Brick-wall flters with cutoff frequency at  $10f<sub>s</sub>$  were used for the low-pass operation for computing the WMS harmonics. Concurrent scanned-wavelength direct absorption spectroscopy (SDAS) measurements of the absorption spectrum indicated an  $A_{\text{known}} = 0.05 \text{ cm}^{-1}$  and  $\Delta v_C = 0.144 \text{ cm}^{-1}$ . The ideal transmitted intensity  $I_t(t)$  for each  $\mu-\sigma$  pair as measured by the photodetector was perturbed by rescaled versions of the  $\rho(t)$  and  $\kappa(t)$  distortion time series measured in Fig. [17](#page-15-1). The rescaling factor was chosen such that the spectral power of  $\rho(t)$  and  $\kappa(t)$  at zero frequency were equal to 5% of the average transmitted intensity, which has the effect of setting  $\xi_{\kappa} = 0.05\overline{I_0}$  and  $\xi_{\rho} = 0.05$ . At  $f_m = 100$ kHz, the effective frequency roll-off exponent of the resulting distortion signal was equal to 0.92 based on power-law ftting of the frequency spectrum within the 100–400 kHz range.

Note that Fig. [18](#page-16-0) is heavily undersampled in the  $\mu-\sigma$ domain relative to Fig. [10](#page-10-0) due to the extensive amount of time needed to collect data for the 252 distinct  $(\mu, \sigma)$  pairs shown here. Nevertheless, it is possible to observe several trends from these results.



<span id="page-16-0"></span>**Fig. 18** Sample-averaged error the best-fit integrated absorbance,  $\epsilon_A$ , color-mapped onto a matrix of scan and modulation index pairs for experimentally measured **a** WMS- $\theta_{1f}$ , **b** 2*f*/1*f*, **c** 3*f*/1*f* and **d** 4*f*/1*f* signals distorted by experimentally measured jet-induced beam-steering

and electronics noise with an effective  $p = 0.92$ . Identical optimal trough equations from Fig. [10](#page-10-0) were superposed in dotted white lines to illustrate the close agreement between the simulated and measured optimal valleys

- 1. The data exhibit valleys of optimal  $\mu$ – $\sigma$  pairs identical to what was observed in Fig. [10.](#page-10-0) The trough equations in Fig. [10](#page-10-0) were superimposed onto Fig. [18](#page-16-0) with white dotted lines to illustrate the near perfect alignment between the simulated and measured optimal valleys.
- 2. Along the valleys,  $\epsilon_A$  for WMS- $\theta_{1f}$  is, in general, lower than  $\epsilon_A$  for the various WMS- $n f/1 f$  methods, with WMS-2*f* / 1*f* , 3*f* / 1*f*, and 4*f* / 1*f* performing on average 22%, 31%, and 20% worse than WMS- $\theta_{1f}$ , respectively. This is consistent with the results in [3.2.3](#page-11-0), where it is expected that WMS- $\theta_{1f}$  performs better than -*nf*/1*f* for  $p = 0.92$ .

Concurrent SDAS measurements using identical measurement rates,  $\rho(t)$  and  $\kappa(t)$  distortion time series, and the optimal SDAS modulation index,  $\mu^*$ , of 5.6 were found exhibit an  $\epsilon_A$  of 2.91%, a value greater than the minimum  $\epsilon_A$  for all of the WMS methods shown in Fig. [18.](#page-16-0) This is consistent with the fndings in Sect. [3.4](#page-14-1) and "Appendix", which predicted that the accuracy of SDAS-based sensors will be worse than WMS-based sensors for stochastic distortion environments characterized by values of *p* near 1.

# **4.2 Flat-flame burner H<sub>2</sub>O speciation and thermometry demonstration**

The experimental setup shown in Fig. [19](#page-17-0) was used as a practical demonstration of the WMS- $\theta_{1f}$  technique. Two fber-coupled DFB-TDLs (NEL Photonics) were used to probe H<sub>2</sub>O transitions near 7185.59 and 6806.06 cm<sup>-1</sup> (1391.7 and 1469.3 nm, respectively) to measure the temperature and mole fraction of  $H_2O$  within the exhaust stream of a premixed  $CH<sub>4</sub>/air$  flat-flame burner. Light from the two lasers was fber-combined, collimated, and transmitted through the 25.4-cm-long burner apparatus and collected with an InGaAs photodetector. To maximize fame-induced distortions on the measured signals, a focusing lens was not installed in front of the photodetector. Transition linestrengths measured in Goldenstein et al. [\[33\]](#page-22-19) were used to calculate temperature based on the bestft integrated absorbance of the two transitions. The laser



<span id="page-17-0"></span>**Fig. 19** Illustration of the experimental setup for  $H_2O$  concentration exhibits several important characteristics. and temperature detection within the exhaust of a flat  $CH<sub>4</sub>/air$  burner

targeting the transition near 7185.59 cm<sup>−</sup><sup>1</sup> was modulated and scanned at  $(f_m, f_s) = (200, 1)$  kHz with optical tuning depths of  $(a_{m,1}, a_{s,1}) = (0.138, 0.209)$  cm<sup>-1</sup>, roughly corresponding to  $(\mu, \sigma) = (3, 4.3)$  for typical combustion product stream composition and temperatures. Similarly, the laser targeting the transition near 6806.06 cm<sup>−</sup><sup>1</sup> was modulated and scanned at  $(f_m, f_s) = (250, 1)$  kHz with optical tuning depths of  $(a_{m,1}, a_{s,1}) = (0.097, 0.140)$  cm<sup>-1</sup>, which also roughly corresponded to  $(\mu, \sigma) = (3, 4.3)$  at typical burner conditions. This  $\mu-\sigma$  pair is located within the trough labeled as "Valley 1" in Fig. [10a](#page-10-0). IM–FM phase offsets were measured to be  $1.21\pi$  and  $1.28\pi$  for the 7185.59 and 6806.06 cm<sup>−</sup><sup>1</sup> lasers, respectively. The scanning frequency of 1 kHz provided a measurement rate of 2 kHz (one measurement per up- or down-scan). Brick-wall flters with cutoff frequency at  $10f_s$  were used to compute the WMS harmonics.

Figure [20a](#page-18-0) shows examples of the measured and best-ft WMS- $\theta_{1f}$  for the absorption transition near 7185.59 cm<sup>-1</sup> probing a CH<sub>4</sub>/air mixture with equivalence ratio,  $\Phi$ , of one. Data were collected over a period of 1 s and processed using WMS- $\theta_{1f}$ , 2*f*/1*f*, and 3*f*/1*f* spectral fitting, yielding 2000 measurement samples across the collection period. Figure [20b](#page-18-0) shows the measured temperature and mole fraction time series processed using WMS- $\theta_{1f}$ . Long timescale fluctuations in both measurements refect physical unsteadiness in the fame whereas short timescale fuctuations represent non-physical measurement uncertainties resulting from environmental signal distortions. Figure [20c](#page-18-0), d shows the Allan deviation of the measured  $H_2O$  mole fraction and temperature, respectively, for WMS- $\theta_{1f}$ , -2*f*/1*f*, and -3*f*/1*f*. For both measurements, the WMS- $\theta_{1f}$  technique shows lower Allan deviations than either WMS-2*f*/1*f* or WMS-3*f*/1*f* regardless of the integration time, a result consistent with experimental observations in Yang et al. [\[14\]](#page-22-1) and which demonstrates the benefits of using WMS- $\theta_{1f}$  for practical sensing.

# **5 Conclusion**

A theoretical description of the scanned-wavelength WMS- $\theta_{1f}$  gas sensing technique and its associated spectral ftting algorithm was presented. The expected performance of the  $\theta_{1f}$  strategy was directly compared against the popular WMS-*nf*/1*f* and SDAS measurement techniques using simulations-based analyses of common signal distortions experienced in harsh measurement environments. The simulations were then experimentally validated in a staticgas-cell experiment subjected to broadband additive and multiplicative signal distortions. The WMS- $\theta_{1f}$  strategy



<span id="page-18-0"></span>**Fig. 20** a Sample spectral fit of the WMS- $\theta_{1f}$  signal for the H<sub>2</sub>O absorption transition near 7185.59 cm−<sup>1</sup> . **b** Measured 1-s temperature (red) and H<sub>2</sub>O mole fraction (blue) time-histories using the WMS- $\theta_{1f}$ technique within the exhaust of a  $CH<sub>4</sub>/air$  flat-flame burner operating

- 1. Like WMS- $n f / 1 f$ , WMS- $\theta_{1f}$  signals are immune to nonabsorbing transmission losses and interfering emission with frequencies less than *f*<sub>m</sub>/2. Additionally, the sensitivity of  $\theta_{1f}$  signal to absorption lineshape enables the simultaneous ftting of lineshape parameters without prior knowledge or models of their behavior.
- 2. Both the optical scan and modulation depth parameters are important drivers of measurement accuracy for WMS-based sensors subjected to broadband environmental distortion. In contrast with the standard optimization strategy of selecting a modulation depth that maximizes the peak WMS response, the proposed method used in this work simultaneously optimizes the two parameters for a given measurement environment.
- 3. There exist certain ranges of frequency roll-off exponent, *p*, in the distortion spectrum at which WMS- $\theta_{1f}$  is expected to perform better than the various WMS-*nf*/1*f* strategies. For the simulations, lasers, and detection systems used in this work,  $\theta_{1f}$  was shown to be the preferred measurement technique for distortion signals that roll of

at equivalence ratio of 1. Allan deviations of **c** H<sub>2</sub>O mole fraction and **d** temperature of the various WMS techniques for the same data set shown in **b**

slower than  $1/f^{1.25}$  in the region of frequencies containing the detection harmonics.

- 4. For stochastic distortion environments with values of *p* near zero, SDAS is expected to perform better than the various WMS-*nf*/1*f* methods regardless of the selected laser tuning parameters. However, WMS- $\theta_{1f}$ performance is generally on par with SDAS for  $p = 0$ , demonstrating that WMS- $\theta_{1f}$  is a suitable substitute for SDAS in applications with distortions exhibiting weak frequency dependence in the spectral amplitude.
- 5. Like WMS- $n f / 1 f$ , the modulation depth of WMS- $\theta_{1f}$  signals can be chosen such that the results are insensitive to wavelength-dependent optical distortions. WMS- $\theta_{1f}$ measurements were also found to be insensitive to étalon interference patterns with low free spectral range, opposite the behavior of SDAS measurements.

It is clear from these results that WMS- $\theta_{1f}$  offers several potential benefts over traditional WMS-*nf*/1*f* and SDAS strategies for high-precision measurements of gas parameters

in harsh measurement environments. Future research seeks to deploy  $\theta_{1f}$ -based sensors for practical applications as well as to extend the analysis techniques developed in this work for selecting optimal laser-tuning parameters that maximize the accuracy of the best-fit  $\Delta v_{\rm C}$  and  $v_{\rm 0}$  parameters for pressure and velocimetry applications, respectively.

Additionally, the measurement technique-agnostic analysis procedures introduced in this work provide, to the authors' knowledge, the frst set of tools for sensor designers to directly compare the relative performance of diferent TDLAS techniques for a given measurement application. These tools are expected to aid sensor designers in determining the true optimal sensing strategy and laser tuning parameters for any target environment.

# **6 Appendix: SDAS signal distortion analysis**

The intensity and optical frequency tuning waveforms of a typical SDAS sensor can be described with sawtooth functions as follows:

$$
I_0(t) = 2\overline{I_0} \left( 2f_s t - \lfloor 2f_s t \rfloor \right) \tag{22}
$$

$$
v(t) = -2a_{m,1}\left(2f_s t - \lfloor 2f_s t \rfloor\right) + a_{m,1}.\tag{23}
$$

Here,  $\vert \vert$  is the floor function used to compute the sawtooth function. The definitions of  $I_0$ ,  $a_{m,1}$ , and  $f_s$  are consistent with the conventions used in Eqs. [6](#page-3-4) and [7](#page-3-5) for scanned-WMS signals. These defnitions are also consistent with the conventions used to define modulation index,  $\mu$  (Eq. [8](#page-3-6)), and non-dimensionalized wavelength-dependent distortion parameter,  $\eta$  (Sect. [3.3\)](#page-12-0). Note that the frequency of the sawtooth was set at  $2f<sub>s</sub>$  to yield a SDAS measurement rate equal to the  $2f_s$  measurement rate of scanned-WMS sensors.

## **6.1 Additive and multiplicative distortions**

To directly compare the WMS-specifc results in Sect. [3.2](#page-6-0) against SDAS,  $I_0(t)$  from Eq. [22](#page-19-1) was attenuated and distorted by the same absorption spectrum and  $\kappa(t)$  and  $\rho(t)$  distortion time series as defned by the parameters listed in Table [1](#page-8-1) to yield  $\widetilde{I_1(t)}$ . The distorted transmitted intensity was then used to compute the distorted absorption spectrum,  $\tilde{\alpha}_{\nu}$ . Each measurement period was then ftted with Voigt profles [[2\]](#page-21-4) to yield the best-ft spectroscopic parameters. An example of this distortion and its efect on the best-ft integrated absorbance is shown in Fig. [21](#page-19-2), with the distortions introducing a 0.262% error in the best-ft integrated absorbance relative to the known value.

Using the analysis methods described in Sect. [3,](#page-5-2) it is possible to compute the sample-averaged error in the best-ft



<span id="page-19-2"></span>Fig. 21 Example effect of stochastic  $1/f$  additive and multiplicative signal distortions on the measured spectral absorbance (black, top panel) and the resulting best-ft absorbance (red, top panel) and residual (bottom panel). Parameters used to simulate these results are listed in Table [1](#page-8-1). The distortions introduced a 0.262% error in the best-ft integrated absorbance relative to the known value

<span id="page-19-1"></span>

<span id="page-19-0"></span>**Fig. 22** Sample-averaged error in the best-ft integrated absorbance for SDAS (green), WMS- $\theta_{1f}$  (black), -2*f*/1*f* (blue), and -3*f*/1*f* (red) as a function of modulation index,  $\mu$ , when perturbed by stochastic additive and multiplicative noise with roll-off exponent,  $p$ , equal to  $a$  1, and **b** 0.  $\mu$ - $\sigma$  pairs for the WMS methods follow the linear relationship labeled as "Valley 2" in Fig. [10.](#page-10-0) Parameters used to simulate these results are listed in Table [1](#page-8-1)



<span id="page-20-0"></span>Fig. 23 Example effect of pure-tone additive and multiplicative distortions on measured **a** SDAS spectra and **b** WMS- $\theta_{1f}$  signals. In both cases, the ideal transmitted intensity was perturbed by pure-tone sinusoidal distortion time series with frequency  $f_s/10$  and  $f_m/10$  for SDAS and WMS- $\theta_{1f}$ , respectively. Note that the distortions negligibly affect the WMS- $\theta_{1f}$  spectrum due to the lock-in filtering and intensity normalization operations



<span id="page-20-1"></span>**Fig. 24** Étalon-induced errors in the best-ft integrated absorbance of SDAS measurements as a function of  $\eta$ . The distortion amplitude was set at 1% of the peak absorbance. Each curve represents a diferent value of modulation index

integrated absorbance,  $\epsilon_A$ , for SDAS measurements as a function of  $\mu$ , which is shown in green in Fig. [22a](#page-19-0). Super-posed on Fig. [22](#page-19-0)a are the  $\epsilon_A$  for WMS- $\theta_{1f}$  (black), -2*f*/1*f* (blue), and -3*f*/1*f* (red) evaluated along the valleys of optimal  $\mu$ <sup>– $\sigma$ </sup> shown in Fig. [10](#page-10-0) ("Valley 2" for each WMS technique was used since they contain the global optima). Several trends can be observed from these results.

- 1.  $\epsilon_A$  for SDAS is consistently higher than the various WMS measurement techniques assuming the roll-off exponent, *p*, is equal to 1. This trend agrees with intuition because the frequency content of the distortion time series are primarily concentrated at low frequencies with  $f < f_{1/2}$ , which in these simulations correspond to  $f < 10f_s$ . Since SDAS does not flter or correct for low-frequency noise signatures, it is expected that the performance of SDAS will be more heavily impacted by the distortions than WMS.
- 2. There exists an optimal  $\mu^* = 5.6$  that minimizes  $\epsilon_A$  for SDAS sensors. To the authors' knowledge, this represents the frst reported observation of optimal tuning parameters for SDAS sensors in the presence of environmental distortions. However, the  $\epsilon_A - \mu$  curve flattens and becomes largely insensitive to  $\mu$  at values of  $\mu$  larger than 5.6, indicating that it is not severely detrimental to SDAS sensor performance if the laser is over-modulated beyond  $\mu = 5.6$ . This characteristic is critical for applications where multiple absorption transitions are simultaneously measured within each laser scan.

We also evaluated the case where the additive and multiplicative distortion signals exhibit uniform power in the frequency domain (i.e.  $p = 0$  or white noise). This is an important test case for practical applications because the primary sources of stochastic signal distortions in quiescent environments (e.g. static cells or environmental monitoring devices in general) are noise in the light detection and electronic digitization hardware, which often exhibit roll-off exponents close to zero (as evidenced in Fig. [17b](#page-15-1)). The results of this exercise are shown in Fig. [22](#page-19-0)b. As can be seen, when subjected to  $p = 0$  distortions, SDAS performs better than the various WMS-*nf*/1*f* techniques while rivaling the performance of WMS- $\theta_{1f}$  sensors. These observations suggest that, for certain environments with stochastic distortions exhibiting sufficiently low  $p$ , implementing WMS instead of SDAS may actually be detrimental to sensor performance.

Although the preceding results demonstrate that SDAS sensors perform comparably to WMS-based sensors for certain broadband, stochastic distortion environments, it is important to note that SDAS does not enjoy the benefts of intensity normalization and optical emission rejection that is inherent in the WMS- $\theta_{1f}$  and -*nf*/1*f* techniques. When SDAS signals are perturbed by strong pure-tone additive and

multiplicative distortions (e.g. from natural vibrations in the measurement environment), it becomes difficult to infer an appropriate  $I_0(t)$  from the measurement, resulting in significant errors. WMS- $\theta_{1f}$  and -*nf*/1*f* do not suffer from these diffculties because all additive and multiplicative distortions with frequency below *f*<sub>m</sub>/2 are removed by the lock-in filter and intensity normalization operations. These distinctions are illustrated in Fig. [23,](#page-20-0) where pure-tone sinusoidal  $\rho(t) = 1 + 0.1 \sin\left(\frac{2\pi f_s}{10}t\right)$  and  $\kappa(t) = 0.1\overline{I_0} \sin\left(\frac{2\pi f_s}{10}t\right)$  distortion time series were used to perturb the SDAS absorption profle in Fig. [23](#page-20-0)a, resulting in a 2.87% error in the best-ft integrated absorbance. These results are compared against Fig. [23](#page-20-0)b, where pure-tone sinusoidal  $\rho(t) = 1 + 0.1 \sin\left(\frac{2\pi f_m}{10}t\right)$ and  $\kappa(t) = 0.1\overline{I_0} \sin\left(\frac{2\pi f_m}{10}t\right)$  distortion time series were used to perturb the WMS- $\theta_{1f}$  signal. As can be seen, the resulting WMS- $\theta_{1f}$  signal was negligibly affected by the pure-tone distortions, resulting in an error in the best-ft integrated absorbance of only 0.0635%.

It is worth noting that there are well-established postprocessing techniques such as polynomial baseline ftting that can correct for signal distortions and somewhat improve the accuracy of SDAS measurements. For the example shown in Fig. [23](#page-20-0)a, a second-order polynomial ftting of the non-absorbing wings of the absorption transition yielded a best-ft integrated absorbance error of 0.59%, a signifcant improvement relative to the uncorrected signal. However, these techniques are generally artifcial methods that add additional free parameters to the post-processing work flow that may not be grounded in reality. Additionally, polynomial baseline ftting cannot account for intra-scan additive offsets in  $I_t(t)$  (e.g. due to optical emission), which would result in incorrect measurements of the absorbance. As a result, correction schemes for SDAS measurements are, at best, unreliable in highly dynamic environments. This is in contrast with WMS- $\theta_{1f}$  and -*nf*/1*f*, both of which correct for additive and multiplicative distortions by directly decoupling these distortions from the spectroscopic measurement a priori rather than through artifcial corrections in post-processing.

#### **6.2 A.2 Wavelength‑dependent distortions**

The procedures described in Sect. [3.3](#page-12-0) were used to compute the average expected error in the best-ft integrated absorbance,  $\epsilon_{A,\text{eth}}$ , for an SDAS measurement distorted by étalon-like interference fringes. Figure [24](#page-20-1) shows how the  $\epsilon_{A,\text{eth}}$  of SDAS varies as a function of  $\eta$  for various values of  $\mu$ . Like in Fig. [15,](#page-13-1)  $\xi_{\text{eth}}$  was set to be 1% of the peak absorbance.

Several observations can be made from comparing Fig. [24](#page-20-1) against Fig. [15](#page-13-1).

- 1. Like the various WMS methods considered in Fig. [15,](#page-13-1)  $\epsilon_{A,\text{eth}}$  vanishes when  $\eta$  approaches zero or infinity and decreases in general with increasing  $\mu$ .
- 2. The primary error band for SDAS is narrower than the bands for WMS- $\theta_{1f}$  and 2*f*/1*f* and is confined to lower values of  $\eta$ , indicating that short-period étalons most adversely impact the accuracy of SDAS sensors. This is the opposite of the behavior observed for WMS- $\theta_{1f}$ (Fig. [15](#page-13-1)a). In certain environments were cavity lengths are long (e.g. cavity-enhanced techniques) and, therefore, étalon free-spectral ranges are short, it may be advisable to use WMS- $\theta_{1f}$  to reduce sensitivity to wavelength-dependent distortions.
- 3.  $\epsilon_{A, \text{eth}}$  within the primary error band for SDAS are, in general, less than either WMS- $\theta_{1f}$  or WMS-2*f*/1*f*, indicating that, at suboptimal values of  $\eta$ , SDAS may be more forgiving than WMS-based methods for wavelength-dependent distortions.

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