

# **OH radical measurements in combustion environments using wavelength modulation spectroscopy and dual‑frequency comb spectroscopy near 1491 nm**

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#### **Abstract**

Hydroxyl radical (OH) is a key intermediate reactive species during combustion processes relevant to power production, transportation, and manufacturing. We demonstrate an OH sensor based on in situ laser absorption spectroscopy for deployment in industrial conditions. The sensor relies on telecommunications-fber-coupled, tunable-diode-laser absorption spectroscopy of an OH transition near 1491 nm. By employing wavelength modulation spectroscopy, the sensor is capable of in situ, quantitative detection of OH down to mole fraction values of  $10^{-5}$  over a 75-cm pathlength. To increase the accuracy of the OH sensor, we perform the frst dual-comb spectroscopy measurement above a fame and use the results to create an absorption database of water vapor transitions from 1489.2 to 1492.5 nm at temperatures up to 2165 K. The database is included in the analysis procedure for the tunable diode laser sensor to account for the water vapor absorption that overlaps with the OH absorption. The utility of the laser sensor is demonstrated by characterizing the concentration of OH radical above a catalytic combustor under diferent operating conditions.

# **1 Introduction**

Hydroxyl radical (OH) is of diagnostic interest because of its key role in combustion kinetics and surface modifcation manufacturing processes. In combustion, it is a prevalent intermediate that can map the fame structure and indicate combustion completeness. In manufacturing, OH stimulates surface treatment of polypropylene [\[1](#page-12-0), [2\]](#page-12-1) and also appears in dielectric discharge plasmas used for material and gas stream processing [\[3](#page-12-2)[–6\]](#page-12-3). Thus, diagnostics that provide temporal and spatial measurements of OH can aid in optimizing the operation of industrial fame systems. However, OH is diffcult to measure in industrial settings because extractive samples from techniques such as gas chromatography are not representative of fame conditions (OH is highly reactive

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 $\boxtimes$  Torrey R. S. Hayden torrey.hayden@colorado.edu and temperature sensitive), and in situ measurements are complicated by high temperatures, limited optical access, and generally harsh deployment conditions.

Laser-induced fluorescence (LIF) and planar laserinduced fluorescence (PLIF) are the most widespread approaches to measuring OH and have been used with great success to image OH concentrations in many combustion applications [[7–](#page-12-4)[15\]](#page-12-5). These methods have provided measurements of small quantities of OH radical in a variety of combustion settings; however, they require some combination of high-power laser equipment, signifcant optical access, and/or sensitive optical sub-systems that are challenging for industrial application and real-time monitoring.

Laser absorption spectroscopy is a preferred option for probing many industrial combustion and plasma systems because it is quantitative, relatively simple, can probe large systems, and requires only one or two small optical ports. In laser absorption spectroscopy, the thermodynamic properties (temperature, pressure, etc.) and concentration of a species of interest can be measured by their relationship to the absorption of laser light with a photon energy matching the energy of a quantum mechanical transition of the target species. Various research groups have used absorption spectroscopy to probe the electronic transitions of OH

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in the ultraviolet–visible (UV–Vis) wavelength range for a variety of applications including shock-tubes, engines, and the atmosphere [\[4](#page-12-6), [16–](#page-12-7)[20](#page-12-8)]. Although the electronic transitions of OH in the UV–Vis range are very strong, it can also be desirable to probe weaker ro-vibrational transitions at near-infrared (NIR) wavelengths. NIR sensors are desirable because of the low-cost and fber-coupled nature of the available lasers. Despite these advantages, there has been limited work on OH absorption spectroscopy in this region. Both Peeters et al. [[21\]](#page-12-9) and Rutkowski et al. [[22\]](#page-12-10) used cavityenhanced absorption techniques near 1500 nm to detect OH concentrations in fame environments. For large, long-term industrial deployment, it is desirable to pursue open-path, non-cavity enhanced approaches such as direct absorption or wavelength modulation spectroscopy with tunable diode lasers.

Wavelength modulation spectroscopy (WMS) is a form of absorption spectroscopy that achieves higher sensitivity than traditional direct absorption spectroscopy by reducing the efects of system noise and intensity fuctuations. Aizawa [\[23](#page-12-11), [24\]](#page-12-12) previously demonstrated the use of WMS to detect OH radical near 1550 nm. These papers highlight the challenges of making accurate OH measurements, namely low OH absorption signal levels and interference from hightemperature  $H_2O$  absorption.

It is not possible to fnd OH transitions in the infrared spectral regions that are free from significant (10–80%) interference from nearby  $H_2O$  transitions. The overlap between the absorption of OH and  $H<sub>2</sub>O$  in the 1490–1492 nm region is shown in the simulated absorbance in Fig. [1](#page-1-0). This overlap is signifcant because a WMS measurement detects the combined absorption of any overlapping species. In order to calculate OH mole fraction from the measured WMS signal, a model of the underlying  $H_2O$  absorption must be included in the ft of the measured signal to account for the added signal. Although spectral databases such as HITRAN or HITEMP contain absorption transition information for  $H_2O$ near 1491 nm, the high-temperature  $H_2O$  parameters in this wavelength region are extrapolated from ambient conditions and have never been experimentally investigated or validated [\[25,](#page-12-13) [26\]](#page-12-14). It has been shown by several past researchers that these databases are increasingly uncertain as temperatures approach combustion conditions [[27](#page-13-0)–[29\]](#page-13-1). Hence, careful characterization of the  $H_2O$  transitions neighboring the strong OH transitions of interest in the 1491-nm range are required in order to develop an accurate OH sensor.

Here, we improve the accuracy of an OH sensor operating near 1491 nm by developing an updated high-temperature water vapor absorption database in this region using dual-frequency comb spectroscopy and multispectral ftting. A premixed ribbon burner system is used as the stable high-temperature environment in which to measure



<span id="page-1-0"></span>Fig. 1 Simulated absorbance of  $15\%$  H<sub>2</sub>O and  $0.3\%$  OH mole fraction from 200.9 to 201.4 THz (1489–1493 nm) using HITEMP 2010 for OH and the spectral linelist developed in this paper for  $H_2O$ . Both species were simulated at 2200 K, 820 mbar, and a pathlength of 30.7 cm. Shaded region indicates the tuning range of DFB diode laser used for WMS measurements. The range of the dual –frequency comb spectrometer extends beyond the plot from 196.2 to 217.2 THz (1380–1528 nm)

these spectral parameters. The light from a dual-frequency comb spectrometer is passed through the ribbon burner system operating at several diferent equivalence ratios in order to measure absorption parameters at diferent thermodynamic conditions. We then extract linestrength and linecenter for the strongest transitions using the multispectral ftting code described in [[30](#page-13-2), [31](#page-13-3)]. The dual-frequency comb spectrometer can measure all of the  $H_2O$  and OH transitions simultaneously with broad bandwidth, high resolution, and a well-known wavelength axis to minimize the uncertainty of the extracted spectral parameters [[29](#page-13-1)]. These are the frst dual-comb spectroscopy measurements in a fame environment.

Finally, we build on the work of Aizawa et al. by incorporating the  $H<sub>2</sub>O$  absorption database into a WMS-based sensor for measuring OH at 1491 nm. We utilize 1*f* normalization of the 2*f* signal to mitigate beam steering in the WMS sensor and use an improved WMS ftting routine to reduce uncertainties in extracted thermodynamic values. We demonstrate the sensor by probing the product gases above the surface of a catalytic combustor (which is nominally expected to be a radical-free heat source) for OH presence at diferent combustion conditions. This demonstration shows the beneft of the WMS OH sensor, which in this case is able to characterize conditions under which the combustor produces OH radical above the surface.

#### **2 Measurement approaches**

## **2.1 Absorption spectroscopy**

Wavelength modulation spectroscopy and dual-frequency comb spectroscopy are both forms of absorption spectroscopy. Absorption spectroscopy is the quantifcation of the light absorbed by a gas sample as a function of the incident wavelength. The light is absorbed if the wavelength matches the energy of a quantum transition between energy states in an atom or molecule. The amount of light absorbed, which is quantified as absorbance,  $\alpha_{\lambda}$ , is directly related to the bulk thermodynamic properties of the probed gas sample through Beer's law, Eq  $(1)$  $(1)$ :

$$
\alpha_{\lambda} = -\ln\left(\frac{I_{\rm t}}{I_0}\right)_{\lambda} = S(T)P\chi L\phi_{\lambda} \tag{1}
$$

where  $I_t$  is the transmitted intensity of the light after it has passed through the sample,  $I_0$  is the initial (un-attenuated) intensity, and  $\lambda$  is the wavelength of the laser light that passes through the gas sample.

As shown in Eq. ([1](#page-2-0)), the absorbance can be represented by the product of the linestrength of the absorbing transition at a given temperature  $S(T)$ , the total pressure  $P$ , the species mole fraction of the absorbing gas  $X_{\text{abs}}$ , the pathlength of the laser across the gas *L*, and the absorption lineshape  $\phi_{\lambda}$ . Thus, the absorbance is explicitly related to the thermodynamic properties *T*, *X*, *P*, and implicitly related to the quantum properties of the target gas through the linestrength and lineshape terms.

In many atmospheric-pressure systems, it is sufficient to model the lineshape as a Voigt profle—a convolution of the Gaussian and Lorentz profles that captures the collisional and Doppler broadening of the absorption transition. The four components of the Voigt profle are the center wavelength, integrated area, Lorentz width, and Doppler width. The Doppler width is not transition-specifc and is calculated only from the wavelength of the transition and the temperature of the gas. The other quantities are transition-specifc and are calculated from linecenter, pressure-shift coefficient, linestrength, foreign and self-broadening coefficients, and their temperature dependence, which are all catalogued in spectral databases such as HITRAN and HITEMP or measured line lists [\[25](#page-12-13), [26](#page-12-14), [28](#page-13-4), [29](#page-13-1)].

## <span id="page-2-2"></span>**2.2 Scanned wavelength modulation spectroscopy (WMS) for industrial OH sensing**

For the industrial OH sensor, we use 1*f*-normalized WMS, which is a sensitive form of absorption spectroscopy. WMS is an excellent choice for industrial applications because of its kilohertz time resolution, compact size, relatively low

cost, and compatibility with telecom fber optic cables of the semiconductor diode lasers [\[32](#page-13-5)[–46](#page-13-6)]. In WMS, the wavelength and intensity of the laser are modulated at a fast frequency, *f*, (typically 10–100's of kHz), which results in a series of harmonics in the detector signal. The harmonics in the detector signal can be isolated using a lock-in amplifer, which acts as a bandpass flter centered at the frequency of interest. We are particularly interested in the second harmonic signal, 2*f*, because its magnitude is related to the absorbance and can be used to extract thermodynamic properties. Normalization of the 2*f* by the frst harmonic, 2*f*/1*f*, enables absolute measurements and removes the efects of intensity fuctuations due to beam steering, window fouling, and vibrating optics [\[36](#page-13-7)].

<span id="page-2-0"></span>The experimental equipment used for WMS measurements is shown in the top-right box in Fig. [2](#page-3-0). Laser current and temperature are controlled by an ILX laser diode controller. Two National Instruments PCI-6110 data acquisition boards supply the modulation voltage to the ILX laser diode controller and acquire the output voltages from the photodetectors. Light from the three lasers is combined and passed to two fbers, one that sends the light to the gas sample via a transmit collimator and another that passes the light directly to a photodetector for intensity monitoring or alternatively to a fiber-based etalon ( $FSR = 636$  MHz) to characterize the wavelength modulation.

By ftting the WMS signal to obtain the integrated absorbance according to the procedures in [[47](#page-13-8)], we eliminate the dependence on the lineshape,  $\phi_{\lambda}$ , as the integral of  $\phi_{\lambda}$  is unity by definition. The result is shown in Eq.  $(2)$  $(2)$  below:

<span id="page-2-1"></span>
$$
A = \int \alpha d\lambda = S(T)PX_{\text{abs}}L
$$
 (2)

The integrated area is thus only a function of the linestrength, pressure, absorbing-species mole fraction, and pathlength. Temperature can be calculated using the ratio,  $A_1/A_2$ , from two transitions of the same species [[48\]](#page-13-9), removing the dependence of the temperature measurement on pressure and mole fraction. Absorbing-species mole fraction can then be calculated by solving Eq. [\(2](#page-2-1)) for  $X_{\text{abs}}$ , with measured pressure, pathlength, and temperature.

As discussed in the previous section and shown in Fig. [1](#page-1-0), an absorption measurement of OH at 1491 nm will also include overlapping  $H<sub>2</sub>O$  absorption. Therefore, analysis of OH must remove the  $H_2O$  absorption contribution in order to isolate the OH integrated absorbance. Therefore, we measure the temperature and  $H_2O$  mole fraction using two diode lasers which probe two interference-free  $H<sub>2</sub>O$  transitions at 1392 nm and 1469 nm [\[47\]](#page-13-8). The integrated areas extracted from these two transitions are used to calculate temperature and  $H_2O$  mole fraction following the technique described above. A third diode laser probes



<span id="page-3-0"></span>**Fig. 2** Schematic of the optical setup used for all ribbon burner experiments. (left) Light from two frequency combs is combined to pass across the fame 7 mm above the edge of the burner for extraction of high-temperature spectral parameters. (top-right) Experimen-

tal setup for WMS OH sensor in the premixed fame of the ribbon burner. (bottom-right) A thermocouple is mounted on a translation stage coincident with the laser beam in the fame to test for fame uniformity

the OH transition near 1491 nm. We use the measured temperature and  $H_2O$  mole fraction to simulate  $H_2O$  absorbance as a function of wavelength in the 1491-nm region and incorporate this simulated absorbance into the ft of the OH transition. Specifcally, the ft treats the OH transition as a floating Voigt lineshape added to the  $H_2O$  absorbance. The summed absorbance of  $H_2O$  and OH is then incorporated into Eq. ([1\)](#page-2-0) with the measured intensity and wavelength variations of the diode laser to simulate the transmitted intensity,  $I_t$ . The simulated transmitted intensity passes through the lock-in amplifer to isolate the 2*f*/1*f* signal, shown in Fig. [3](#page-4-0). The simulated 2*f*/1*f* signal is then ft to the measured data by varying the parameters of the Voigt lineshape of the single OH transition until residuals between the model and the measured WMS signal are minimized. An example of a measured 2*f*/1*f* signal and the corresponding model ft using the database derived here in this paper is shown at the bottom of Fig. [3](#page-4-0). The integrated absorbance for the single OH transition that minimizes residuals is used to calculate OH mole fraction.

The absolute accuracy of the OH measurement is highly dependent on the quality of the underlying  $H_2O$  database used to simulate  $H_2O$  absorbance as a function of wavelength in the analysis. There have been no known database measurements of  $H<sub>2</sub>O$  in this 1490–1493 nm region, which means that the best available data are from HITEMP 2010. The line parameters of HITEMP in this region have 10–20% uncertainty as they are calculated from quantum models. This uncertainty correlates to signifcant uncertainty in the OH measurement. Hence, we measure the spectral line parameters of the  $H<sub>2</sub>O$  transitions in this region using dualfrequency comb spectroscopy to improve the accuracy of our OH measurement.

## **2.3 Dual‑frequency comb spectroscopy (DCS) for H2O database development**

Dual-frequency comb spectrometers are an emerging class of laser spectrometers that provide simultaneously broad spectral bandwidth, high spectral resolution, and excellent wavelength accuracy [\[49](#page-13-10)]. With these characteristics, the dual-frequency comb spectrometer is well-suited for the development of spectral databases, particularly for hightemperature environments [\[28](#page-13-4), [29](#page-13-1), [50,](#page-13-11) [51\]](#page-13-12). Dual-frequency comb spectrometers are recently also capable of feld deployment to industrial environments [\[52](#page-13-13), [53\]](#page-13-14), but currently at a higher cost and slower measurement rate than tunable diode laser systems. Thus, we use frequency comb spectroscopy here to develop the high-temperature database for interfering water vapor absorption in the region near the OH transition of interest, but still apply tunable diode laser WMS for the primary industrial OH sensor.

The dual-comb spectrometer employed here consists of two linear-cavity, Er-doped fber frequency comb lasers with a 200-MHz pulse repetition rate. Each comb is individually amplifed using erbium-doped fber amplifers and then spectrally broadened with a section of highly non-linear fber (HNLF). In the optical frequency domain, each comb

<span id="page-4-0"></span>



produces hundreds of thousands of modes, which appear as narrow linewidth "comb teeth", spaced by the pulse repetition rate of the laser. The broadened spectrum is then spatially fltered to the wavelength range of interest using a grating-based flter. This fltered light then passes through the gas sample of interest to probe the absorption in the NIR and infer the thermodynamic properties.

Once the light from both lasers passes through the gas sample, they interfere on a single photodetector. In order to read the signals from each comb tooth, the two combs are stabilized to slightly diferent pulse repetition rates, which results in a diferent frequency spacing of the comb teeth for each laser. Therefore, when the lasers are combined, the frequency diference between corresponding "comb teeth" results in a unique RF heterodyne frequency at an integer multiple of the diference between repetition rates,  $\Delta f_{\text{ren}}$ . The resulting time-domain signal on the detector is a series of interferograms which repeat at the diference frequency between the two comb repetition rates. A Fourier transform of these interferograms reveals the transmission spectrum including attenuation due to absorption from the measured gas. Further description of

the specifc DCS method and spectrometer applied here can be found in Refs. [\[52–](#page-13-13)[54](#page-13-15)].

# **3 High-temperature H<sub>2</sub>O absorption database development using dual‑comb spectroscopy in a premixed fame**

We utilize dual-frequency comb spectroscopy (DCS) measurements in a well-characterized fame to develop a high-temperature  $H<sub>2</sub>O$  absorption database near 1.5 um. The absorption line parameters are extracted from DCS data using the multispectral ftting code described in Refs [\[30,](#page-13-2) [31\]](#page-13-3). This section will provide an overview of the database extraction process, as highlighted in Fig. [4](#page-5-0).

To extract a new database from measurements in the fame, we provide the multispectral ftting code with three inputs: (1) an estimated linelist to provide seed input values for the location, strength, and shape of transitions; (2) measured transmission data from the fame; and (3) the known thermodynamic properties (temperature, pressure, etc.) along the laser line-of-sight during the measurement. We utilize the  $H<sub>2</sub>O$  line parameters from HITEMP 2010 for the seed absorption database [[25](#page-12-13)]. DCS measurements near 1491 nm at three diferent fame conditions provide the input transmission data. Fits of the DCS-measured transmission from a diferent wavelength region where a published high-temperature experimental database exists [\[29\]](#page-13-1) provide the line of sight thermodynamic properties of the fame. The multispectral ftting routine uses the three inputs to ft all three data sets simultaneously to determine the optimal set of absorption line parameters for the new database. We then compare a simulation created using the new high-temperature  $H_2O$  database to measured DCS data at a fourth premixed fame condition to test the new parameters.

#### <span id="page-5-1"></span>**3.1 Transmission measurements**

The multispectral ftting routine is able to generate the best possible list of spectral absorption parameters if it is provided measurements at diferent temperatures, pressures, and species mole fractions. Many of the  $H_2O$  transitions measured in combustion conditions near 1491 nm are not detectable (without cavity enhancement) except at high temperatures; thus, we require a gas sample above 1700 K containing sufficient  $H_2O$  to generate detectable absorption for stable fts. By probing a premixed fame, we can achieve high temperatures in a uniform manner. The burner system used in this work for extraction of line parameters is a ribbon burner, which is a burner consisting of a series of small flame jets [\[55](#page-13-16)]. These burners have been shown to produce stable fames that result in uniform temperature felds that can be used for industrial processing [\[56\]](#page-13-17). Since the ribbon burner is a premixed fame, OH will be produced in the reacting zone of the fame and then decrease toward equilibrium values for the temperatures present above the fame. We present a characterization of the OH within and above the reacting zone of the ribbon burner in Sect. [4.1](#page-10-0) to demonstrate that the WMS sensor indeed measures OH instead of additional H2O absorption.



<span id="page-5-0"></span>Fig. 4 Flow chart of high-temperature,  $H_2O$  absorption experimental linelist development. The purple trace is the transmission measured by dual-frequency comb spectroscopy above a premixed fame operating at 90 W/cm<sup>2</sup> and  $\phi$  = 0.9. Shading highlights wavelength regions used to provide the transmission data for the linelist (orange) and temperature and  $H_2O$  mole fraction (blue) inputs to multispectral

<span id="page-6-0"></span>**Table 1** Premixed ribbon burner conditions for the four measurements used for spectral line parameter measurements

Condition #	Fuel (SLPM) Air (SLPM)		Equiva- lence ratio	Power flux (W/cm <sup>2</sup> )
1	6.28	76.1	0.8	79
$\overline{c}$	7.00	75.3	0.9	88
3	7.36	75.3	0.95	93
4	7.70	74.6		97

The ribbon burner combusts premixed methane and air at controlled conditions, as described in Ref. [[45\]](#page-13-18). Low-power fluxes, approximately 90  $W/cm^2$ , were selected to keep the visible fame short and gas velocity low. By keeping the fame short, the laser can probe the product gases above the visible fame at a distance close to the burner surface, which helps to minimize non-uniformity across the fame due to buoyancy-driven necking and entrainment of surrounding air (which occurs at larger heights above the burner). Air and fuel fow rates were adjusted to achieve varying fuel–air equivalence ratios while maintaining a consistent fame height. The laser light from the frequency combs passes along the length of the burner at approximately 7 mm above the surface, as shown in Fig. [2](#page-3-0).

The fame conditions and reactant fow rates for the four measurement conditions are given in Table [1.](#page-6-0) Data using the dual comb spectrometer are collected at each of the four conditions for 30 min and a background data set is collected for 20 min while the air is fowing at 100 SLPM. A sample of the transmission spectrum is shown in Fig. [4,](#page-5-0) with the gold region highlighting the wavelength range of interest for OH detection. The 30-min averaged transmission data for three out of the four conditions are used with the background data as an input for the multispectral ftting. Condition #3 is retained as a test case to validate the spectral parameters measured from the other three cases.

## **3.2 Line‑of‑sight thermodynamic properties of the premixed fame**

In order to minimize uncertainty in the extracted line parameters, it is important to create a known, spatially uniform thermodynamic profle along the laser beam. We characterize the spatial uniformity of the ribbon burner test bed by probing the temperature profle along the beam path with a type-K thermocouple. The thermocouple bead is 1.65 mm, approximately the size of the laser beam diameter. The thermocouple is placed in the beam path on a horizontal translation stage, as shown in Fig. [2](#page-3-0). Figure [5](#page-6-1) shows the thermocouple-measured temperature normalized by the maximum temperature. There is 1.4% maximum variation measured between the bounds of the burner (denoted by



<span id="page-6-1"></span>**Fig. 5** Normalized temperature measured with a type-K thermocouple in the beam path

gold-dash lines) and a steep gradient on each side. Similar uniformity is achieved for all four trials. The 297-mm width of the uniform region is taken to be the pathlength for the multispectral ftting.

The thermocouple characterizes the uniformity along the beam path, but does not accurately measure the absolute temperature due to radiative, convective, and conductive losses around the thermocouple tip. Instead, we utilize the dual-comb spectrometer measurements from 1455.3 to 1469.5 nm (204–206 THz) where a high-temperature database has already been created using a tube furnace [\[29](#page-13-1)] to calculate the thermodynamic properties for each condition. The measured transmission within this wavelength range (highlighted in the blue box in Fig. [4\)](#page-5-0) is ft with a model to extract the temperature,  $H_2O$ , and OH mole fractions. The ftting routine (described in Ref. [[57\]](#page-13-19)) uses line parameters provided by Schroeder et al. [[29\]](#page-13-1) for H<sub>2</sub>O and HITEMP 2010 [\[25](#page-12-13)] for OH. The resulting measured thermodynamic properties for all four trials are shown in Table [2](#page-7-0). The equilibrium  $H<sub>2</sub>O$  and OH mole fractions are calculated using NASA CEA [\[58](#page-13-20)] at the measured temperatures. The measured OH mole fraction is on average 20% higher than the equilibrium calculations. This trend is consistent with the expectation that OH concentrations will decrease toward the equilibrium value as the reaction comes to completion above the visible region of the fame.

### **3.3 High‑temperature absorption linelist**

We now use three of the four measured conditions to create the experimental linelist, leaving the data condition #3 to test the experimental linelist. To begin, we evaluate the linestrengths at both 296 K and 2200 K for the 83,000 lines in

Condition #	ratio	Equivalence Measured tem- perature(K)	Adiabatic tem- perature(K)	Measured $H2O$ mole fraction	Equilibrium $H_2O$ mole fraction	Measured OH mole fraction	Equilibrium OH mole fraction
	0.8	2019	1994	0.162	0.153	0.00275	0.00212
$\overline{2}$	0.9	2119	2130	0.181	0.169	0.00354	0.00316
	0.95	2154	2184	0.187	0.177	0.00389	0.00341
4		2165	2220	0.191	0.184	0.00385	0.00316

<span id="page-7-0"></span>**Table 2** Flame conditions measured by DCS and used for the multispectral fitting routine to extract H2O absorption transition parameters near 1491 nm

the 1489–1492.5 nm (200.8–201.4 THz) spectral region of the HITEMP database, and apply a  $3 \times 10^{-26}$  cm<sup>-1</sup>/atm linestrength flter. The threshold reduces the database to 1983 H2O lines and 4 OH lines. These lines produce absorption features above the noise foor at the measured conditions. The multispectral ftting routine uses the line parameters from the seed list as initial guess values to simulate absorbance as a function of wavelength as a series of Voigt profles. This simulated line-by-line absorbance from  $H_2O$  and OH is combined with a 9th order Chebyshev polynomial baseline to produce a transmission model that is compared to the measured transmission for a given condition. Residuals between this model and the measured transmission spectra are calculated for each condition.

The multispectral program iteratively fts the Chebyshev baseline and  $H<sub>2</sub>O$  spectral parameters to the measured transmission spectra. OH parameters are not varied because it is difficult to achieve controlled quantities of OH and the database parameters are typically assumed to be sufficient, such as in Ref [[22](#page-12-10)]. The linecenters are scaled to burner conditions using the HITEMP air-shift parameter with a temperature-dependence exponent of unity [\[50](#page-13-11)]. The selfshift was fxed to zero because HITEMP does not provide ab initio coefficients and Schroeder et al. [[29\]](#page-13-1) was not able



<span id="page-7-1"></span>Fig. 6 (Top) Relative transmission measured during trial 2 with baseline fit from multispectral analysis. (Middle) Baseline-corrected transmission (black), multispectral ft result (green), and ft residual (blue)

to observe the  $H_2O$  self-shift in a multispectral fit to a neighboring spectral region. An example of this ftting is shown in the top panel of Fig. [6,](#page-7-1) which shows the measured relative transmission for condition #2 and the Chebyshev baseline. First, the Chebyshev polynomial coefficients are fit to each transmission spectrum using the seed database absorption model. Then, the reference linestrengths  $S_{296}$  are floated to minimize ft residuals. If the foated linestrength of a particular transition is not stable within the reported statistical uncertainty of the multispectral ft after ten iterations, the linestrength is fxed back to its HITEMP value. Of the 1983 HITEMP transitions above the measurement threshold in this spectral region, 82 were found to produce stable linestrength fts that reduced the multispectral ft residual. All others were fxed at the HITEMP value because the signalto-noise ratio and range of thermodynamic conditions were insufficient to produce stable fits. Next, self-broadening coefficients are systematically floated in the multispectral ft. Self-broadening was chosen over foreign broadening as the measured gases are  $18-20\%$  H<sub>2</sub>O and H<sub>2</sub>O induces signifcant self-broadening. Of the 82 transitions with foated

linestrength, only 14 produced stable self-broadening fts that reduced the ft residual. Finally, a single linecenter was foated for a strong transition that showed clear evidence of an incorrect linecenter. The spectral parameters of the 28 H<sub>2</sub>O doublets in the spectral range are constrained to each other according to the procedure in [\[29](#page-13-1), [51\]](#page-13-12). The baselinecorrected transmission spectra for condition #2, normalized to 100, is shown in black in the middle panel of Fig. [6](#page-7-1) with a simulation of the fnal linelist and the corresponding residuals on the bottom panel.

Having extracted absorption line parameters for  $H_2O$  in the 1489–1492.5 nm (200.8–201.3 THz) region, we test the new experimental linelist using the measured spectrum of condition  $\#3$  ( $\phi$  = 0.95, 93 W/cm<sup>2</sup>). The measured absorbance is shown in black in Fig. [7.](#page-8-0) We compare the measured absorbance to simulated absorbance created from HITEMP 2010 (green) and the experimental line list (pink). Each simulation is compared to the measured data to produce the residuals shown in the top axis of Fig. [7](#page-8-0). It is clear from Fig. [7](#page-8-0) that the residuals for the simulation using the experimental linelist are smaller across almost all frequencies, suggesting the experimental linelist



<span id="page-8-0"></span>**Fig. 7** Measured absorbance as a function of optical frequency as compared to absorbance simulated with HITEMP 2010 (green) and the experimental database (pink) at the measured thermodynamic properties. The resulting residuals between model and ft are shown as the top traces

improves on HITEMP 2010 under these conditions. The inset of Fig. [7](#page-8-0) illustrates the improved agreement near the OH transition that will be used by the WMS sensor. For linestrength, 36 of the new linestrengths deviate less than 50% and 65 deviated less than 200% from HITEMP 2010.

Uncertainties of the extracted line parameters in the new linelist fall into two categories: transition-specifc and transition-independent. The transition-independent uncertainties are based on the systematic uncertainties of the experiment and will be the same for all extracted parameters. By simplifying the equation for temperature-dependent linestrength:

$$
S(T) = S_0 \frac{Q(T_0)}{Q(T)} \left(\frac{T_0}{T}\right) \exp\left(-\frac{hcE''}{k} \left(\frac{1}{T} - \frac{1}{T_0}\right)\right)
$$

$$
\times \left(1 - \exp\left(\frac{-hc\nu_0}{kT}\right)\right) \left(1 - \exp\left(\frac{-hc\nu_0}{kT_0}\right)\right)^{-1}
$$
(3)

where *T* is the temperature of interest,  $S_0$  or  $S_{296 \text{ K}}$  is the strength of the absorption transition at the reference temperature  $T_0$  (296 K),  $Q(T)$  is the partition function at a given temperature, *E*″ is the lower state energy state of a transition,  $v_0$  is the center frequency of the transition, and *hc* and *k* are Planck's constant and the Boltzmann constant, respectively. Because  $v_0$  will be approximately the same for all transitions measured in this work we can simplify this expression to:

$$
S(T) = S_{296K}\beta(T) \tag{4}
$$

where  $\beta(T)$  contains the temperature dependence of the linestrength, simplifed to:

$$
\beta(T) = \left(\frac{296}{T}\right)^{2.77} \exp\left(-\frac{1.435E''}{T}\right) \tag{5}
$$

We can then solve Eq. [\(2](#page-2-1)) for  $S_{296K}$  to give:

$$
S_{296K} = \frac{A}{\beta(T)PXL}
$$
\n(6)

The uncertainty on each of the independent variables on the right-hand side of Eq. [\(6](#page-9-0)) contributes to the overall systematic linestrength uncertainty. Temperature and mole fraction uncertainties from the ft to the 1455.3–1469.5 nm (204–206 THz) spectral region propagate to  $S<sub>296K</sub>$  uncertainties according to Eqs. ([5\)](#page-9-1) and [\(6](#page-9-0)). In particular, temperature uncertainty will propagate as an uncertainty of *β*(*T*). Pressure and path length uncertainty are captured in the uncertainty of the mole fraction ft, and thus a single value of uncertainty for the product of *PXL* is reported based on the mole fraction fit uncertainty. The integrated area (*A*) uncertainty is attributed to any nonlinearity in DCS measurements.

We can, therefore, estimate the transition-independent uncertainty of the extracted reference linestrength to be the sum of the uncertainties of each of these quantities in quadrature,

$$
\frac{\delta S_{296K}}{S_{296K}} = \sqrt{\left(\frac{\delta A}{A}\right)^2 + \left(\frac{\delta \beta}{\beta}\right)^2 + \left(\frac{\delta PXL}{PXL}\right)^2}
$$
(7)

which is approximately 7%. This analysis assumes 0.5% uncertainty in *A* based on a conservative estimate of the frequency comb linearity in these experiments [[59\]](#page-13-21), 5.5% uncertainty in  $\beta$  from temperature uncertainty, and an average *E*″ and 3.8% uncertainty in *PXL* from the uncertainty in DCS ftting for *X*.

We estimate the transition-specifc uncertainty as a product of statistical uncertainty of the multispectral ftting routine and the stability of the ft. The statistical uncertainty for each fit parameter is extracted with the new line parameter and is tabulated in the Appendix A of ESM. We determined the fit stability by perturbing the stable  $S_{296K}$  and self-broadening values by a normal distribution of uncertainties with a standard deviation of 10% and iterating the multispectral ftting with these new parameters. Select parameters were unstable in this perturbation study due to the narrow range of conditions covered in the data. This perturbation uncertainty is also tabulated with the new linelist parameters in the Appendix A of ESM. These transition-specifc linestrength uncertainties should be included together with the transitionindependent linestrength uncertainty (7%) to determine the overall linestrength uncertainty for any particular transition.

<span id="page-9-1"></span>It is important to note that the linestrength and self-broadening values measured in this work are primarily for use under similar temperature and pressure conditions as those demonstrated here. Because of the small temperature and pressure range over which these parameters were obtained, it is not suggested that these parameters be extrapolated over large ranges, as one might do with lists obtained from data over larger ranges.

# <span id="page-9-0"></span>**4 Demonstration of WMS OH radical sensor**

While the broadband, high-resolution nature of DCS provides greater accuracy and fexibility, WMS is a robust, high time-resolution, lower-cost option for a targeted measurement of a particular species. As described in Sect. [2.2,](#page-2-2) OH and  $H_2O$  significantly overlap and interfere with one another in the near-infrared spectral region. Therefore, the OH mole fraction must be extracted from measured WMS signals which include absorption by both  $H<sub>2</sub>O$  and OH. We use the measured high-temperature  $H<sub>2</sub>O$  linelist as the basis for data analysis in the WMS sensor. The new WMS sensor is used to characterize OH in both a premixed fame system known to contain OH, and a catalytic combustor system with unknown OH content.

#### <span id="page-10-0"></span>**4.1 OH measurements in a premixed fame**

We first demonstrate our WMS sensor in a system that is known to produce OH. The premixed ribbon burner described in Sect.  $3.1$  and used to produce the H<sub>2</sub>O linelist, was shown via modeling by Sullivan et al. [\[60](#page-13-22)] to have OH mole fractions up to  $7 \times 10^{-3}$ . We use higher methane and air fow rates than those used to produce absorption spectra for the linelist retrieval (equivalence ratio of 0.95 and power fux of approximately 500 W/cm<sup>2</sup>). This results in visible flames that are several centimeters tall. Measurements are made across the fame at 2-mm vertical intervals for the frst 1 cm and 5-mm vertical intervals up to 5 cm. Each measurement is averaged for 30–60 s to account for turbulent fuctuations in the fame.

Figure [8a](#page-10-1)–c shows the measured 2*f*/1*f* signals at three of the heights above the burner surface to illustrate how the H2O and OH absorption signals change in the premixed fame. The signals at 1392 nm and 1469 nm, shown in Fig. [8](#page-10-1)a, b, represent  $H_2O$  absorption with two different lower state energies (*E*″), which give the features diferent temperature dependencies. The consistent magnitude of the signals in both (a) and (b) at all three heights suggests that  $H<sub>2</sub>O$  and temperature are relatively constant at these three heights. On the other hand, the measured 2f/1f signal at 1491 nm in Fig. [8](#page-10-1)c drops by 50% between 7 and 38 mm, which follows the expected trend for OH. If the signal arose from H2O absorption from any of a variety of lower state energy levels, we would not expect such a change. This is further highlighted in Fig. [8](#page-10-1)d, which shows the vertical profiles for temperature,  $H_2O$  and OH mole fraction in the premixed flame. Both temperature and  $H<sub>2</sub>O$  increase initially and then become stable as the combustion process is complete. The green (triangle) trace in 8(d) shows that the



<span id="page-10-1"></span>**Fig.** 8 Measured 2 $f/1f$  at three heights above burner exit using lasers at **a** 1391.7 nm to measure H<sub>2</sub>O, **b** 1468.9 nm to measure H<sub>2</sub>O and **c** 1490.8 nm to measure OH. **d** Measured average temperature, H2O, and OH mole fractions as the beam is scanned vertically in the fame

measured OH mole fraction in the fame using the 1491 nm WMS measurements initially increases, followed by a rapid decline. We measure a peak mole fraction of OH of 0.006 with an uncertainty of  $8 \times 10^{-4}$ , which is very close to the maximum of approximately 0.007 predicted by the simulations in Ref. [\[60](#page-13-22)].

#### **4.2 OH measurement in a catalytic combustor**

In industry, catalytic combustors can be used to reduce emissions from processes requiring combustion, and as a nominally non-reacting heat source. Introducing a catalyst lowers the energy required for the dissociation of combustion reactants, with the goal of obtaining more complete combustion with lower emissions [[61,](#page-13-23) [62\]](#page-13-24). Because the catalyst reduces combustion temperatures to 1400–1600 K, we do not expect to fnd equilibrium OH mole fraction levels above 10−4 if combustion is complete. Therefore, sensing OH at the exit of a catalytic combustor can provide insight into the efectiveness of the catalyst at achieving complete combustion, or the health of a catalyst in long-term monitoring applications.

We send the combined laser light of the WMS sensor in a triple pass arrangement across the surface of the combustor to probe the product gases exiting the combustor. Four combustor power fluxes were selected from 5 to 20 W/cm<sup>2</sup>. At 5 W/cm<sup>2</sup> the catalytic bed is barely glowing, which suggests low temperatures that are not conducive to the existence of OH. On the other hand, when the burner operates at 20 W/cm<sup>2</sup> there is a visible blue flame at the surface of the catalytic bed that indicates a high likelihood of continued combustion and the possible presence of detectable OH concentrations. At each power fux the equivalence ratio is varied by fixing the methane flow rate and varying the air flow rate to change the combustion chemistry and conditions at

the surface. Thus, the  $\phi$  = 0.6 case corresponds to the highest air fow rate and, therefore, the highest exit velocity for a given power fux.

Figure [9](#page-11-0)a, b shows the measured temperature and  $H_2O$ mole fraction as a function of equivalence ratio for the four power fux conditions selected. Vertical uncertainty bars are 4% and 7%, respectively, which is the sensor uncertainty derived from validations under known conditions [\[47](#page-13-8)] and the uncertainty of experimental conditions. The increase in both properties with equivalence ratio matches expectations that each will reach a maximum near or slightly above the stoichiometric  $(\phi = 1)$  condition. Additionally, the traces in Fig. [9](#page-11-0)a confrm that higher temperatures occur at higher fuel flow rates.

The OH mole fraction trends are shown in Fig. [9c](#page-11-0). Timeresolved OH mole fraction results indicate that the precision of the measurement is about  $1.6 \times 10^{-5}$  at 1700 K for the data presented here with a pathlength of 75 cm and  $2.9 \times 10^{-4}$  at 2000 K with a pathlength of 90 cm. The precision is used here as an estimate of the lowest detectable level of OH. The uncertainty of OH mole fraction values is approximately 9%, which is calculated by adding in quadrature the uncertainties of temperature,  $H<sub>2</sub>O$  mole fraction, pressure, pathlength and *S*(*T*) as per Eq. ([2\)](#page-2-1).

Combustion should be complete, and the equilibrium level of OH at the gas temperatures above the catalyst should be  $\sim 10^{-4}$ ; thus, it is surprising to see indications of OH radical present above the catalytic combustor surface. For the lower power fux conditions, OH concentration decreases with equivalence ratio. This trend opposes both the expected trend of increasing equilibrium levels of OH with increasing temperature and increasing OH concentrations in premixed combustion as equivalence ratio increases [[60](#page-13-22)]. In the case of the lowest power flux  $(5 \text{ W/cm}^2)$ , the OH mole fraction



<span id="page-11-0"></span>**Fig. 9 a** Temperature, **b**  $H_2O$  mole fraction and **c** OH mole fraction as a function of equivalence ratio for a range of power fuxes measured at the surface of a catalytic combustor using WMS. Uncertainty bars for **a** and **b** are 4% and 7%, respectively, from measurements in

a fxed-temperature furnace in Ref. [\[36\]](#page-13-7). The uncertainty bars for **c** are 9%, which are estimated from the combined  $H_2O$  and temperature measurement uncertainties and uncertainty in  $S_{296 \text{ K}}$  values extracted in this work

measured at  $\phi$  = 0.6 is significantly larger than the other conditions (where OH levels are very close to the detection limit of the sensor). The OH, therefore, appears to trend with air fow rate, instead of temperature or equivalence ratio, indicating that the catalyst may not have sufficient residence time or high enough temperatures in the catalyst-bed to react all of the fuel and oxidizer at higher fow rate cases. Whether this trend is due to a saturation of the available active sites on the catalyst or residence time issues is not clear, but this condition is not ideal for certain industrial applications where reactive radicals in the product gases of the catalytic combustor are not desired. Overall, these measurements demonstrate the utility of the WMS OH sensor for studying the efectiveness of the catalyst under diferent conditions or for long-term industrial monitoring of catalyst health.

# **5 Conclusion**

In this work, we employ both wavelength modulation spectroscopy and dual-comb spectroscopy to develop an accurate and sensitive OH radical sensor near 1491 nm. The frst dual-comb spectroscopy measurements above a controlled premixed fame are performed in order to develop a high-temperature absorption model for  $H<sub>2</sub>O$  near 1491 nm. The air and fuel fow rates of a premixed ribbon burner are adjusted to achieve four diferent fame conditions for use in multispectral ftting to extract an empirical absorption linelist valid for similar combustion temperatures. Simulations of absorbance using the new experimental linelist are in better agreement with measured absorbance data at our fame conditions, as compared to simulations derived from HITEMP 2010.

The new experimental linelist is incorporated as part of an industrial OH sensor using wavelength modulation spectroscopy at 1491 nm, in order to properly account for unavoidable  $H_2O$  interference. The linelist serves as the basis for simulating absorbance as a function of wavelength for  $H_2O$ from 1489 to 1493 nm so that  $H<sub>2</sub>O$  is accounted for when measuring the OH contribution of the WMS signal. Using this linelist results in a signifcant increase in measured OH, matching well with previous work done in premixed fames. Our sensor has an OH mole fraction detection limit of  $1.6 \times 10^{-5}$  at 1700 K with 75 cm pathlength.

We apply the sensor for characterization of a premixed fame from a ribbon burner and the gases above a catalytic combustor. Measurements of OH at 1491 nm in the premixed fame match the expected trend for OH. Probing the gases at the exit of the catalytic combustor at a variety of power fux and equivalence ratio conditions indicate the existence of additional combustion above the surface of the catalytic bed at high velocity and/or higher fuel fow rate conditions. The combined sensitivity of WMS, together with the accuracy provided by the high-temperature spectral empirical linelist derived from DCS, make the sensor ideal for application to catalytic combustion, plasmas or premixed fame processes in industry. The time-resolved and accurate nature of the sensor is also ideal for long-term monitoring and control.

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