Noise-immune cavity-enhanced optical frequency comb spectroscopy: a sensitive technique for high-resolution broadband molecular detection

Amir Khodabakhsh · Alexandra C. Johansson · Aleksandra Foltynowicz

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Abstract Noise-immune cavity-enhanced optical frequency comb spectroscopy (NICE-OFCS) is a recently developed technique that utilizes phase modulation to obtain immunity to frequency-to-amplitude noise conversion by the cavity modes and yields high absorption sensitivity over a broad spectral range. We describe the principles of the technique and discuss possible comb-cavity matching solutions. We present a theoretical description of NICE-OFCS signals detected with a Fourier transform spectrometer (FTS) and validate the model by comparing it to experimental CO₂ spectra around 1,575 nm. Our system is based on an Er:fiber femtosecond laser locked to a cavity and phase-modulated at a frequency equal to a multiple of the cavity free spectral range (FSR). The NICE-OFCS signal is detected by a fast-scanning FTS equipped with a high-bandwidth commercial detector. We demonstrate a simple method of passive locking of the modulation frequency to the cavity FSR that significantly improves the long-term stability of the system, allowing averaging times on the order of minutes. Using a cavity with a finesse of \sim 9,000, we obtain absorption sensitivity of $6.4 \times 10^{-11} \text{ cm}^{-1} \text{ Hz}^{-1/2}$ per spectral element and concentration detection limit for CO_2 of 450 ppb Hz^{-1/2}, determined by multiline fitting.

1 Introduction

Cavity-enhanced optical frequency comb spectroscopy (CE-OFCS) [1, 2] has been gaining increased attention in the research community during the last decade because of

the unique combination of broad spectral bandwidth, high resolution, and absorption sensitivity that it provides. The technique allows detection of entire molecular absorption bands in acquisition times down to µs [3, 4] and enables simultaneous multispecies detection with high species selectivity [2, 5, 6]. Until now, the potential of CE-OFCS has been demonstrated for applications ranging from breath analysis [5, 6], time-resolved chemical sensing [4], flame characterization [7], and industrial process control [8] to environmental monitoring [9].

In order to take advantage of the benefits offered by optical frequency combs (OFC) for spectroscopy, various detection schemes have been developed capable of analyzing the broad spectral information carried by an OFC [10–13]. In general, they are based either on a dispersive element in combination with a detector, detector array or a camera [1, 5, 14–17], or on Fourier transform spectrometry (FTS) employing either a mechanical interferometer [18, 19] or the dual-comb approach [3, 20, 21]. Of these, a mechanical FTS provides the advantage of acquiring the entire spectral bandwidth of an OFC with high resolution using a single detector in acquisition times on the order of a second. This approach is easily implemented in both the near- [18, 19] and mid-infrared [6, 22] wavelength ranges, and it is compatible with commercial FTIR instruments [23, 24].

The broad equidistant spectrum of an optical frequency comb can be efficiently coupled to an external cavity by matching the repetition rate, f_{rep} , of the comb to the cavity free spectral range (FSR) and adjusting the carrierenvelope offset frequency, f_{ceo} , so that the comb lines come into resonance with the corresponding cavity modes. This matching has to be actively sustained in order to compensate for the drift of the cavity length and/or the comb parameters. One way of coupling the comb to the cavity is to dither the cavity modes or comb lines around the perfect

A. Khodabakhsh · A. C. Johansson · A. Foltynowicz (⊠) Department of Physics, Umeå University, 901 87 Umeå, Sweden e-mail: aleksandra.foltynowicz@physics.umu.se

matching condition, so that the comb comes periodically into resonance with the cavity. This approach is often used in combination with detection schemes based on dispersion elements, in which the cavity transmission is averaged over many dither cycles [5, 16]. However, it is not compatible with a fast-scanning FTS, in which an interferogram is acquired at a sampling rate on the order of Msamples/s, which is higher than the feasible dither frequency (usually in the kHz range, limited by the bandwidth of comb or cavity actuators). This precludes efficient averaging of the transmitted comb intensity and causes, in turn, distortion in the interferogram and the final absorption spectrum. Thus a tight lock of the comb to the cavity is required when a fastscanning FTS is used as a detection method [6, 19]. The tight comb-cavity lock introduces amplitude noise in the transmitted spectrum, originating from the residual fluctuations of the comb line frequencies relative to the center frequencies of the narrow cavity modes, a process called frequency-to-amplitude noise conversion. This noise can be reduced down to the shot noise level using an auto-balancing detector to subtract the signals from the two outputs of an interferometer [19]. However, this approach requires a custom-built detector, and the achieved noise reduction depends critically on the performance of the detector.

We have recently shown [25] that noise in the FTS-based CE-OFCS can instead be reduced in the optical domain by employing a concept known from continuous wave (cw) laser-based absorption spectroscopy and utilized in noiseimmune cavity-enhanced optical heterodyne molecular spectroscopy (NICE-OHMS) [26]. In NICE-OHMS, the laser light is phase-modulated at a frequency equal to the FSR of an external cavity, so that when the laser carrier frequency is locked to a cavity mode, its sidebands are transmitted through the adjacent cavity modes. As in ordinary frequency-modulation spectroscopy (FMS) [27], the transmitted intensity contains a beat signal at the modulation frequency whose amplitude is proportional to, depending on the detection phase, the difference between the phase shift of the carrier and the mean value of the phase shift of the sidebands, or the difference between the attenuations of the sidebands. When the modulation frequency is perfectly matched to the cavity FSR, the carrier and the sidebands are transmitted through their corresponding cavity modes in exactly the same way and any residual frequency noise on the laser affects them in an identical manner. Thus no signal is detected at the modulation frequency in the absence of a molecular or atomic absorber and the technique gains immunity to frequency-to-amplitude noise conversion by the cavity.

We implemented this concept in near-infrared CE-OFCS by locking an Er:fiber OFC to an external cavity, phasemodulating it at a frequency equal to the cavity FSR and detecting the transmitted intensity with a fast-scanning FTS equipped with a high-bandwidth detector. In the first demonstration of noise-immune cavity-enhanced optical frequency comb spectroscopy (NICE-OFCS) [25], we obtained close to shot-noise-limited absorption sensitivity of 4.3×10^{-10} cm⁻¹ Hz^{-1/2} per spectral element in a cavity with a finesse of ~2,000. Here we present the NICE-OFCS method in more detail; we discuss various comb-cavity matching solutions and describe the theoretical model of the signal. Compared to the first demonstration, we increase the detection sensitivity by using a cavity with a higher finesse, and we improve the long-term stability of the system by locking the modulation frequency to the cavity FSR. We show the validity of the theoretical model by comparing it to experimental NICE-OFCS signals from the overtone CO₂ band at 1,575 nm. We use multiline fitting and show that it improves the concentration detection limits compared to those obtained from the signal-to-noise ratio of a single absorption line.

2 Comb-cavity matching in NICE-OFCS

In NICE-OFCS an optical frequency comb is locked to a cavity and phase-modulated at a frequency equal to (a multiple of) the cavity FSR. Each comb line and sideband needs to be transmitted through a separate cavity mode in order to prevent overlap and unwanted interference between the various frequency components. To transmit all comb lines and their sidebands through the cavity, the ratio between the laser $f_{\rm rep}$ and the cavity FSR has to be at least 3, so that there are at least two 'empty' cavity modes in between those to which the comb lines are locked. The conceptually simplest way to achieve this is shown in Fig. 1a, where each triplet is indicated by a different color. Note that in this configuration the modulation frequency, f_m , can be equal to the FSR (as shown in the figure) as well as to $(3 m \pm 1)$ FSR, where m is an integer.

Since the typical repetition rate of commercially available visible and near-infrared comb sources is in the 80 to 250 MHz range, a linear cavity with FSR = $f_{rep}/3$ becomes impractically long. For example, the f_{rep} of our Er:fiber comb is equal to 250 MHz yielding a cavity FSR of 83.3 MHz, corresponding to a length of 1.8 m. To decrease the cavity length, other ratios of FSR and f_{rep} must be used, however, at the expense of losing some of the OFC lines in the transmission.

The comb-cavity matching solution that we used in the initial demonstration of NICE-OFCS [25] is shown in Fig. 1b. In this case every fourth OFC line is matched to every third cavity mode, and the modulation frequency can again be equal to the cavity FSR or to $(3 \ m \pm 1)$ FSR. Thus every fourth comb line and its sidebands are transmitted through the cavity, which for our comb had a compact



Fig. 1 Examples of comb-cavity matching schemes for implementing NICE-OFCS. The transmitted and reflected modes and sidebands are shown by *solid* and *dashed lines*, respectively. **a** Matching that allows transmitting all incident comb lines. **b** Matching used in the initial demonstration that yields a shorter cavity length. **c** Matching implemented in this work that allows transmitting more of the incident comb power than **b** and removes the beating of the sidebands of different comb lines at the modulation frequency

length of 45 cm. A disadvantage of this approach is that 75 % of the OFC power is reflected and lost for the spectroscopic measurement and that the sidebands of two different comb lines beat with each other at the modulation frequency, potentially introducing drift and distortion to the NICE-OFCS spectrum.

In order to increase the fraction of the transmitted power and remove the beating of the sidebands belonging to different comb lines while keeping the cavity length reasonable, we introduce in this work a different combcavity matching solution that is shown in Fig. 1c. In this solution, every third comb line is matched to every fourth cavity mode, so that 33 % of the incident comb lines are transmitted, and the modulation frequency can be equal to the cavity FSR or $(4 \ m \pm 1)$ FSR (the particular case of $f_{m2} = 3$ FSR₂ is shown in the figure). The cavity length is in our case increased to 80 cm, which also has the advantage of increasing the interaction length between the comb light and the sample compared to the previous cavity.

3 Theory of FTS-based NICE-OFCS signals

In NICE-OFCS, the phase-modulated OFC light that is transmitted through the cavity is analyzed by an FTS, where it is split into the two interferometer arms, reflected from moving mirrors (and thus Doppler-shifted), and finally recombined at the output of the FTS to yield an interferogram. When the OFC field is phase-modulated sinusoidally with a small modulation index, $\beta < 1$, one pair of sidebands will be created for each comb line, separated from it by the modulation frequency, $\omega_m = 2\pi f_m$ [27]. In the case when both mirrors in the FTS are moving (in opposite directions), the two phase-modulated and Doppler-shifted electric fields, E_+ , at the output of the FTS are given by [25]

$$E_{\pm} = \sum_{n} \sum_{k=-1,0,1} \frac{E_n}{4} J_k(\beta) T_{n,k} e^{i \left[(\omega_n + k\omega_m) \left(t \pm \frac{\Delta}{2c} \right) \right]} + c.c.,$$
(1)

where E_n is the field amplitude and $\omega_n = 2\pi (nf_{rep} + f_{ceo})$ is the angular frequency of the *n*th comb line, $J_k(\beta)$ is the Bessel function of order k, Δ is the optical path difference (OPD) between the arms in the FTS, c is the speed of light in vacuum, and $T_{n,k} = T(\omega_n + k\omega_m)$ is the complex transmission function of the cavity containing the analyte for the *n*th comb line (k = 0) or its sidebands ($k = \pm 1$), given by [6]

$$T_{n,k} = \frac{te^{-i\varphi_{n,k}/2 - \delta_{n,k} - i\phi_{n,k}}}{1 - re^{-i\varphi_{n,k} - 2\delta_{n,k} - 2i\phi_{n,k}}},$$
(2)

where *t* and *r* are the frequency-dependent intensity transmission and reflection coefficients of the cavity mirrors, respectively, and $\varphi_{n,k} = 2(\omega_n + k\omega_m)n_rL/c$ is the round-trip phase shift caused by the cavity, where in turn n_r is the index of refraction inside the cavity and *L* is the cavity length. The single-pass amplitude attenuation and phase shift of the light caused by the analyte are given by

$$\delta_{n,k} = \frac{1}{2} S c_{\rm rel} p L \operatorname{Re}_{\chi_{n,k}}$$
(3)

and

$$\phi_{n,k} = \frac{1}{2} Sc_{\rm rel} p L {\rm Im} \chi_{n,k}, \tag{4}$$

respectively, where *S* is the line strength (cm⁻²/atm), c_{rel} is the relative concentration of the analyte, *p* is the pressure (atm), and $\chi_{n,k}$ is the complex lineshape function (cm).

The intensity of the light at the output of the FTS contains a term oscillating at the modulation frequency that has two contributions. The first contribution is the beating of the comb lines of the field traveling in each arm of the FTS with their own sidebands, which after demodulation at f_m yields a DC offset. The second contribution is the beating of the comb lines of the electric field traveling in one arm with the sidebands of the field from the other arm (and vice versa), which constitutes the NICE-OFCS interferogram. The NICE-OFCS interferogram can be written as

Note that this expression reduces to the signal retrieved in cw NICE-OHMS when the OPD is equal to zero [28].

$$I_{\omega_{m}}^{\text{NICE-OFCS}} = J_{0}(\beta)J_{1}(\beta)\sin(\omega_{m}t) \times \sum_{n} I_{n} \left\{ \cos\left(\omega_{m}\frac{\Delta}{2c}\right)\cos\left(\omega_{n}\frac{\Delta}{c}\right) \text{Im}\left(T_{n,0}T_{n,-1}^{*} - T_{n,0}^{*}T_{n,+1}\right) \right. + \sin\left(\omega_{m}\frac{\Delta}{2c}\right)\sin\left(\omega_{n}\frac{\Delta}{c}\right) \text{Im}\left(T_{n,0}T_{n,-1}^{*} + T_{n,0}^{*}T_{n,+1}\right) \right\} - J_{0}(\beta)J_{1}(\beta)\cos(\omega_{m}t) \times \sum_{n} I_{n} \left\{ \cos\left(\omega_{m}\frac{\Delta}{2c}\right)\cos\left(\omega_{n}\frac{\Delta}{c}\right) \text{Re}\left(T_{n,0}T_{n,-1}^{*} - T_{n,0}^{*}T_{n,+1}\right) \right.$$
(5)

where the comb line intensity $I_n = c\varepsilon_0 E_n^2/2$ has been introduced. This signal contains one in-phase [sin $(\omega_m t)$], and outof-phase [cos $(\omega_m t)$] component. Each of these comprises, in turn, a sum of two terms consisting of a product of three factors. The first factor is the sine/cosine function of the downconverted modulation frequency, $\omega_m \Delta/2c$, which gives rise to a slowly varying envelope over the interferogram. The second factor is the sine/cosine function of the down-converted optical frequency, $\omega_n \Delta/c$, which, after summation over all frequencies, constitutes the quickly oscillating part of the interferogram. The last factor includes the contributions from the complex cavity transmission function for each comb line and its sidebands.

The full formula above must be used to model the signals measured for arbitrary absorption values. However, to show the functional form of the signal in a more intuitive way, it is useful to express it in the weakly absorbing sample approximation, where both $|\delta_{n,0} - \delta_{n,\pm 1}|$ and $|\phi_{n,0} - \phi_{n,\pm 1}|$ are $\ll 1$. For small molecular attenuation and phase shift, the cavity transmission function can be approximated by $T_{n,k} = 1 - 2F(\delta_{n,k} + i\phi_{n,k})/\pi$, where *F* is the frequency-dependent cavity finesse and where the cavity phase shift is assumed to be a multiple of 2π . Under these conditions, and neglecting the terms proportional to $\delta_{n,k}^2$ and $\phi_{n,k}^2$, Eq. (5) can be rewritten as The NICE-OFCS interferogram resides on a DC offset originating from the beating of the comb lines with their own sidebands, as described above. This offset, in the weakly absorbing sample approximation, can be expressed as

$$I_{\omega_m}^{\text{OFFSET}} = J_0(\beta) J_1(\beta) \sin(\omega_m t) \\ \times \sum_n I_n \cos\left(\omega_m \frac{\Delta}{2c}\right) \frac{2F}{\pi} \left[\phi_{n,-1} - 2\phi_{n,0} + \phi_{n,+1}\right] \\ + J_0(\beta) J_1(\beta) \cos(\omega_m t) \\ \times \sum_n I_n \cos\left(\omega_m \frac{\Delta}{2c}\right) \frac{2F}{\pi} \left[\delta_{n,-1} - \delta_{n,+1}\right].$$

$$(7)$$

It is important to note that this DC offset is zero in the absence of an absorber; moreover, it is noise-immune since the amplitude attenuations and phase shifts caused by the cavity cancel. This implies that the DC offset does not couple in any noise originating from the frequency-to-amplitude noise conversion, which is the key to the noise immunity of the FTS-based NICE-OFCS technique.

The intriguing feature of FTS-based NICE-OFCS, as compared to cw NICE-OHMS, is the existence of the

$$I_{\omega_m}^{\text{NICE-OFCS}} = J_0(\beta) J_1(\beta) \sin(\omega_m t) \times \sum_n I_n \left\{ \cos\left(\omega_m \frac{\Delta}{2c}\right) \cos\left(\omega_n \frac{\Delta}{c}\right) \frac{2F}{\pi} \left[\phi_{n,-1} - 2\phi_{n,0} + \phi_{n,+1}\right] \right. \\\left. + \sin\left(\omega_m \frac{\Delta}{2c}\right) \sin\left(\omega_n \frac{\Delta}{c}\right) \frac{2F}{\pi} \left[\phi_{n,-1} - \phi_{n,+1}\right] \right\} \\\left. + J_0(\beta) J_1(\beta) \cos(\omega_m t) \\\times \sum_n I_n \left\{ \cos\left(\omega_m \frac{\Delta}{2c}\right) \cos\left(\omega_n \frac{\Delta}{c}\right) \frac{2F}{\pi} \left[\delta_{n,-1} - \delta_{n,+1}\right] \\\left. - \sin\left(\omega_m \frac{\Delta}{2c}\right) \sin\left(\omega_n \frac{\Delta}{c}\right) \left[2 - \frac{2F}{\pi} \left(\delta_{n,-1} + 2\delta_{n,0} + \delta_{n,+1}\right)\right] \right\}.$$
(6)

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Fig. 2 Experimental setup: OFC, optical frequency comb; EOM, electro-optic modulator; FC, fiber collimator; $\lambda/2$, half-waveplate; $\lambda/4$, quarter-waveplate; PBS, polarizing beam splitter; FTS, Fourier transform spectrometer; BPF, band-pass filter; G, amplifier; Ph, phase

shifter; DBM, double-balanced mixer; LPF, low-pass filter; FFT, fast Fourier transform; DDS, direct digital synthesizer; f_m , NICE-OFCS modulation frequency; f_{PDH} , Pound-Drever-Hall modulation frequency

absorption term that violates the noise-immune condition [in the last line of Eqs. (5) and (6)]. This term is not zero in the absence of an absorber; it is also not noise-immune since the attenuations of a comb line and its sidebands combine with the same sign; hence, contributions from the frequency-to-amplitude noise conversion add up. However, after summation over all comb lines, the interferogram originating from this term has a form of a short burst. This implies that the noise couples in only during a small fraction (typically <1 %) of the OPD scan in the FTS, which makes this noise contribution negligible.

4 Experimental setup and procedures

The experimental setup is shown in Fig. 2. Our OFC source is an Er:fiber femtosecond laser (MenloSystems, FC1500-250-WG) with a repetition rate of 250 MHz that provides two low-power outputs (20 mW) directly from the oscillator and a high-power output (up to 500 mW) after a fiber amplifier, all in the 1.5- to 1.6-µm wavelength range. The output of the OFC is phase-modulated at two frequencies, $f_{\rm PDH}$ and f_m , using a single fiber-coupled lithium-niobate electro-optic modulator (EOM, Thorlabs, LN65S-FC). The lower frequency, $f_{\rm PDH}$, provided by a signal generator, is equal to 20 MHz and is used for generating the sidebands for the two-point Pound-Drever-Hall (PDH) locking of the OFC to the cavity [6]. The phase-modulated OFC light is coupled into free space and spatially mode-matched to the TEM_{00} mode of the cavity. The combination of the waveplates and the polarizing beam splitter picks up the light reflected from the cavity. The reflected beam is dispersed by a grating, and two different wavelength regions of the dispersed light (defined as locking points) are incident on two separate photodetectors (Electro-Optics Technology,

3000A). The outputs of the photodetectors are synchronously demodulated at f_{PDH} to yield two PDH error signals [29]. The first error signal is fed into a proportional-integral (PI) controller (New Focus, LB1005) that acts on the injection current of the OFC pump diode and controls the f_{ceo} of the OFC with a bandwidth of 150 kHz. The second error signal is split into two paths and fed to two separate PI controllers that work in parallel to control the f_{rep} of the OFC. The first controller (homebuilt) acts on the OFC resonator length via an intracavity piezoelectric transducer (PZT), controlling the f_{rep} with a low bandwidth (6 kHz) but a large range (5 kHz). The second one (New Focus, LB1005) is connected to an EOM inside the OFC resonator, which modulates the intracavity refractive index and thus changes the $f_{\rm rep}$ with a high bandwidth (500 kHz) but a small range (0.1 Hz) [30].

We use two cavities with different lengths and finesses: a 45-cm-long cavity (cavity 1, FSR₁ of 333.3 MHz) with a finesse of ~2,300 and a 80-cm-long cavity (cavity 2, FSR₂ of 187.5 MHz) with a finesse of ~9,000. Both cavities are made of pairs of identical mirrors (Layertec) with a 5-m radius of curvature. The comb-cavity matching for these cavities is shown in Fig. 1b and c, respectively. Cavity 1 yields a 1-GHz spacing between the transmitted OFC lines. For this cavity, we use a modulation frequency, f_{m1} , equal to the cavity FSR₁, i.e., 333.3 MHz, with a modulation index of 0.36. Cavity 2 yields a 750-MHz spacing between the transmitted OFC lines. For practical reason (availability of the RF components) for this cavity, we use a modulation frequency, f_{m2} , equal to three times the cavity FSR₂, i.e., 562.5 MHz, with a modulation index of 0.42. Because of the different losses in the cavity mirrors, cavity 1 transmits 66 % of the incident power per comb line, while cavity 2 transmits only 12 %. This implies that although cavity 2 transmits more comb lines than cavity 1,

2 Cavity 1 Units Length (L) 45 80 cm 187.5 Free spectral range (FSR) 333.3 MHz Finesse (F) @ 1,579 nm 9,400 2,150 _ Modulation frequency (f_m) 333.3 562.5 MHz Transmitted comb line 1 0.75 GHz spacing Transmitted comb lines 1/4 1/3 Transmission per comb line 66 12 % 70 80 % Transmitted bandwidth Total average transmitted 11.5 3.2 % power ${\rm cm}^{-1}\,{\rm Hz}^{-1/2}$ 4.2×10^{-9} Noise equivalent absorption 3.5×10^{-8} @ 1,579 nm 6.2×10^{-10} 6.4×10^{-11} $cm^{-1} Hz^{-1/2}$ Sensitivity per spectral element $ppm Hz^{-1/2}$ CO₂ concentration 3.3 0.45 detection limit

 Table 1
 The summary of the properties of the two cavities used in this work and the achieved performance

the average transmission per nanometer of spectral bandwidth is a quarter of that of cavity 1. (The different parameters of both cavities are summarized in Table 1.) In order to compensate for this difference in transmission, we use the oscillator output of the OFC with cavity 1 and the amplifier output with cavity 2. We adjust the amplifier power in front of the EOM using a variable fiber attenuator (OZ Optics) so that the total optical power transmitted through cavity 2 is the same as the power transmitted through cavity 1 using the oscillator output.

We target the $3v_1 + v_3$ overtone band of CO₂ around 1,575 nm, as it coincides with the maximum power of the oscillator output. When using the amplifier, we adjust its spectral shape so that it also has a maximum around this wavelength. We choose the locking points at 1,572 nm for the current lock and at 1,579 nm for the PZT + EOM lock in order to transmit only the fraction of the incident spectrum that overlaps with the CO₂ absorption band. Due to this choice of locking points, 70 % of the spectral range of the oscillator and 80 % of the spectral range of the amplifier output is transmitted through cavity 1 and 2, respectively. The resulting average power in cavity transmission is 1.4 mW for both cavities.

The cavity mirrors are mounted on stainless steel spacer tubes connected to a gas system in a flow configuration, which allows filling the cavities with gas samples at a desired pressure. The available gases are pure N₂, 1 % CO₂ in N₂, and 1,000 ppm CO₂ in N₂, the latter two specified with an accuracy of 1 %. The flows of the buffer gas (N₂) and the CO₂ standards are adjusted independently with programmable flow controllers (Bronkhorst, F-201CV). This allows mixing the CO_2 standards with pure N_2 to obtain lower CO_2 concentrations in the cavity. A vacuum pump is connected to the cavity outlet via a programmable pressure controller (Bronkhorst, P-702CV), which stabilizes the pressure inside the cavity.

In the first demonstration of NICE-OFCS [25], we used a free-running tuneable RF source to generate the modulation frequency and manually tuned it to the proper value by minimizing the noise in the NICE-OFCS signal. Since the RF source frequency was not locked to the cavity FSR, the perfect matching between the two frequencies was lost after some time due to the drift of the cavity length. In order to compensate for this drift, we now generate the modulation frequency with a direct digital synthesizer (DDS, Analog Devices, AD9915) using the fifth harmonic of the OFC repetition rate (i.e., 1.25 GHz) as a clock input. The clock signal is measured at the second oscillator output of the OFC with a fast photodetector (Electro-Optics Technology, 3000A), whose output is filtered with a band-pass filter to remove the other harmonics. In this configuration, the modulation frequency is locked to the repetition rate of the laser, which in turn is locked to the cavity FSR. This eliminates the need of manual tuning or long-term control of the modulation frequency to compensate the drifts in the cavity length and enhances the long-term stability of the system.

The light transmitted through the cavity is injected into a fast-scanning FTS, whose design is similar to that described in [19]. Two 2-inch-diameter gold-coated retro-reflectors (Edmund Optics) mounted back-to-back on a miniature linear translation stage (Parker, LX80L) constitute the ends of the moving arms of the interferometer. The scanning speed of the retro-reflectors is 0.2 m/s, which implies that the OPD between the interferometer arms is scanned at 0.8 m/s. The OPD is calibrated using a stabilized He-Ne laser (stability on the order of 10^{-7}) whose beam is propagating parallel to the comb beam in the FTS. The OFC output of the interferometer is detected by a 1-GHz bandwidth InGaAs detector (Electro-Optics Technology, 3000A), while the He-Ne output is detected by a 2-MHz bandwidth Si detector (Thorlabs, PDA36A-EC). The average comb power reaching the InGaAs detector is 0.37 mW for both cavities. The output of this detector is synchronously demodulated at the modulation frequency $(f_{m1} \text{ or } f_{m2})$ at a phase adjusted using a phase shifter between the DDS output and the local oscillator input of the double-balanced mixer (DBM). The He-Ne interferogram and the demodulated comb interferogram are recorded with a 2-channel data acquisition card (National Instruments, PCI-5922) at 5 Msample/s and 20-bit resolution. The comb interferogram is resampled at the zero-crossings and extrema of the He-Ne interferogram to yield an OPD-calibrated NICE-OFCS interferogram. Fast Fourier transform of the NICE-OFCS interferogram yields finally a NICE-OFCS signal.

5 Results

5.1 NICE-OFCS interferogram

The magnitude of the different absorption and dispersion factors in the FTS-based NICE-OFCS signal depends on the ratio of the modulation frequency to the linewidth of the probed transitions. Both dispersion factors present in the in-phase interferogram, as well as the first absorption factor present in the out-of-phase interferogram [see Eq. (6)], are maximized when the modulation frequency is roughly equal to the full width at half maximum (FWHM) linewidth of the probed absorption lines, as is the case in ordinary FMS [27] or NICE-OHMS [31]. For lower modulation frequencies (i.e., in the undermodulated case), these dispersion and absorption factors become a small fraction of ϕ_0 or δ_0 [27], respectively. The second absorption factor, however, has a different dependence on the ratio of f_m to the linewidth, since the attenuations of a comb line and its sidebands combine with the same sign. In contrast to the other three factors, this factor is on the order of δ_0 also in the undermodulated case. Moreover, it contains a background, which enables normalization and removes the need for calibration. These two features make detection of the out-ofphase interferogram the preferred mode of operation, since the noise increase caused by the partial loss of noise immunity is marginal (cf discussion at the end of Sect. 3).

Figure 3 shows the out-of-phase NICE-OFCS interferograms from both cavities recorded for an OPD scan from 0 to 120 cm, which lasts 1.5 s. The dominating features in the out-of-phase interferogram are the bursts originating from the background signal present in the last term of Eqs. (5) and (6). Due to the different effective comb line spacing in cavity transmission and different modulation frequencies, the positions of these bursts and their relative amplitudes are dissimilar for the two cavities. In general, the amplitudes of the bursts follow the sine envelope [the envelopes from Eqs. (5) and (6) are drawn by the dashed and dasheddotted lines in the figure as guidance to the eye], wherefore there are no bursts where this envelope is equal to zero. The small deviations from the amplitudes determined by the sine envelope are caused by the divergence of the OFC beam in the FTS, which decreases the interferogram amplitude for larger OPDs.

For detection of the NICE-OFCS signal, we limit the acquisition to a symmetric range around one burst (indicated in Fig. 3 for both cavities), which yields spectral resolution equal to the effective comb line spacing in cavity transmission. Such resolution is sufficient as long as the comb line spacing is lower than the linewidth of the absorption lines. As mentioned above, of the two molecular absorption terms in the out-of-phase interferogram, the one multiplied by the sine envelope [last line in Eqs. (5) and



Fig. 3 The out-of-phase NICE-OFCS interferogram as a function of optical path difference (*blue*) together with the sine (*dashed black*) and cosine (*dashed-dotted red*) envelope factors, drawn as a guidance to the eye, for **a** cavity 1 (L = 45 cm) and **b** cavity 2 (L = 80 cm)

(6)] dominates over the one multiplied by the cosine envelope [second-last line in Eqs. (5) and (6)] in the undermodulated case. In this case, detecting the burst at OPD = 30or 60 cm for cavity 1 yields the same NICE-OFCS signal, since both bursts are symmetrically placed with respect to the sine envelope and the contribution from the signal multiplied by the cosine envelope can be neglected. For cavity 2, we detect the burst at OPD = 80 cm, because it has the largest magnitude and is symmetric with respect to both envelopes. The scan over the indicated acquisition ranges takes 0.38 and 0.50 s for the two cavities, respectively. We acquire the signal only during one direction of the scan of the FTS retro-reflectors, which imposes dead time in repeated measurements that is used for processing and saving the data. As a result, for cavity 1 and 2 one interferogram is acquired every 1 s and every 1.6 s, respectively.

5.2 NICE-OFCS signal

Figure 4 shows CO_2 NICE-OFCS absorption signals from the two cavities normalized to corresponding background spectra taken with cavities filled with pure nitrogen (black curves, 20 averages). Cavity 1 was filled with 1 % of CO_2 in N₂ and cavity 2 with 200 ppm of CO_2 in N₂ (obtained by diluting the 1,000 ppm CO_2 mixture with pure N₂), both at a total pressure of 500 Torr (667 mbar). At this pressure, the average FWHM linewidth of the CO_2 absorption lines is ~3 GHz.

We compare the experimental absorption spectra to the theoretical model by performing a fit to the data. We calculate the theoretical spectrum using the last factor in the last line of Eq. (5), i.e., $\operatorname{Re}(T_{n,0}T_{n,-1}^* + T_{n,0}^*T_{n,+1})$, hence neglecting the term in the previous line, which is justified in the undermodulated case (the FWHM linewidth of the absorption lines is much broader than the modulation frequencies used). We also neglect the contribution from the slowly varying sine envelope of the interferogram. We calculate the attenuation and phase shift caused by the individual absorption lines using spectral line parameters from the HITRAN database [32] and a complex Voigt profile under the prevailing experimental conditions. We neglect the frequency offset between the comb lines and cavity modes, i.e., we put the cavity phase shift, φ_{nk} , to an integer multiple of 2π . We also take into account the wavelength dependence of the cavity finesse, which was measured by cavity ring-down. The measured values of the finesse are shown by black markers in Fig. 4a and d together with a third-order polynomial fit to the data (red) for the two cavities, respectively. The theoretical spectrum is fitted to the experimental data with CO_2 concentration as a fitting parameter, together with a third-order polynomial and a sum of low-frequency sinusoidal etalon fringes to correct the baseline. The experimental data (after baseline correction) and the fitted CO₂ spectra (inverted for clarity) are shown in black and red, respectively, in Fig. 4b and e, with the corresponding residuals in the lower panels, c and f.

As can be seen from the figure, the theoretical model and the experimental data show in general a good agreement over the entire spectral range. The main reason for the dissimilar structure in the residuals of the fits to the signals measured from the two cavities is that the two cavities are made with mirrors from different coating runs. Mirrors for cavity 1 were designed for minimum dispersion at 1,530 nm, while those for cavity 2 was designed for 1,570 nm. This implies that the spectra from cavity 1 are measured away from the design wavelength, which causes a slight walk-off between the comb lines and cavity modes and thus dispersion overshoots in the absorption lines [6]. This frequency offset between the comb lines and cavity modes was not included in the fitted model, which is the cause of the asymmetric structures visible in the residuum, e.g., around 1,570 and 1,578 nm. The spectra from cavity 2 are affected by the dispersion of the cavity mirrors to a much smaller extent, since they are measured close to the design wavelength of the mirrors. The structure visible in the residuum of the spectrum measured in cavity 2 is more symmetric around each absorption line and indicates that the fitted model needs to be expanded to include the sine envelope as well as the absorption factor multiplied by



Fig. 4 Experimental and theoretical NICE-OFCS signals from cavity 1 (*upper figure*) and 2 (*lower figure*). **a** and **d** Cavity finesse measured by cavity ring-down (*black markers*) along with a third-order polynomial fit (*red*). **b** and **e** Normalized NICE-OFCS absorption signals (*black*, 20 averages) from **b** 1 % CO₂ in N₂ in cavity 1 and **e** 200 ppm of CO₂ in N₂ in cavity 2, both at 500 Torr total pressure, along with fitted spectra calculated using the model described in the text (*red*, inverted for clarity). **c** and **f** Residuals of the fits

the cosine envelope factor. Once these factors are taken into account, the influence of the speed-dependent effects on molecular lineshapes might become visible. However, even using this simplified model, the CO₂ concentrations returned by the fits were 0.99 % for cavity 1 and 210 ppm for cavity 2. The first value is within the specified 1 % uncertainty of the CO₂ sample; the larger discrepancy of the second value (5 %) can partly be explained by inaccuracies in the gas mixing process.

5.3 Sensitivity

To estimate the absorption sensitivity, we take the ratio of two consecutive background spectra and fit and remove a baseline as in the treatment of the absorption signals described above for both cavities. We evaluate the noise at 1,579 nm, which corresponds to the center of the P branch of the CO₂ absorption band. The standard deviation of the noise is found to be $\sigma_1 = 2.5 \times 10^{-3}$ at 0.76 s for cavity 1, and $\sigma_2 = 2.0 \times 10^{-3}$ at 1 s for cavity 2. This corresponds to a noise equivalent absorption coefficient of 3.5×10^{-8} and 4.2×10^{-9} cm⁻¹ Hz^{-1/2} for the two cavities, respectively, calculated as $\sigma T^{1/2}/L_{\text{eff}}$, where *T* is the acquisition time of two interferograms and the effective length L_{eff} is defined as $2FL/\pi$, with finesse equal to 2,150 and 9,400 at 1,579 nm for the two cavities, respectively. The number of resolved elements, *M*, in the spectral range shown in Fig. 4 is 3,200 for cavity 1 and 4,300 for cavity 2. Thus, the absorption sensitivity per spectral element, calculated as $\sigma/L_{\text{eff}}(T/M)^{1/2}$, is equal to 6.2×10^{-10} cm⁻¹ Hz^{-1/2} for cavity 1 and 6.4×10^{-11} cm⁻¹ Hz^{-1/2} for cavity 2.

By calculating the CO₂ concentration that yields a NICE-OFCS absorption signal with amplitude equal to the standard deviation of the noise on the baseline, σ , for the strongest absorption line in the P branch, we find that the noise equivalent CO_2 concentration limit is equal to 22 and 2.5 ppm of CO₂ at 1-s integration time for cavity 1 and 2, respectively. A better way to estimate the concentration detection limit is to make use of the entire spectrum and the multiline fitting advantage [22]. To do this, we measure background spectra with cavities filled with pure nitrogen at 500 Torr for 1 h and normalize them to the first spectrum. Afterward we fit a sum of the simplified model of the NICE-OFCS CO₂ spectrum described above together with a baseline to these normalized spectra. The Allan-Werle plot [33] of the concentrations found from these fits is shown in Fig. 5 for cavity 1 (solid blue markers) and 2 (solid red markers). The two dashed black lines show the $\tau^{-1/2}$ dependence characteristic for the white-noise-dominated regime that is fitted to the corresponding measurement points. The CO₂ concentration detection limits, given by the slopes of the fitted lines, are 3.3 ppm $Hz^{-1/2}$ for cavity 1 and 0.45 ppm $Hz^{-1/2}$ for cavity 2. Note that these detection limits are a factor of ~6 lower than the noise equivalent CO₂ concentration calculated for a single absorption line in the center of the P branch. The fact that the detection limit for cavity 2 is \sim 7 times lower than for cavity 1 agrees well with the ratio of the products of the finesse and length of the two cavities, which determine the absorption path length enhancement $[(9,000 \times 80)/(2,300 \times 45) \approx 7]$. The Allan-Werle plot shows that the absolute minimum detectable concentration of CO₂ is 140 ppb after 520 s of integration in cavity 1, and 25 ppb after 330 s in cavity 2.

To demonstrate the long-term stability improvement when using the f_{rep} -referenced DDS instead of the fixedfrequency RF source for generating the modulation frequency, we show in Fig. 5 for comparison the concentrations retrieved from fits to normalized background spectra measured using the RF source and cavity 1 (open cyan markers) and 2 (open orange markers). For short integration times, using the DDS and the RF source yields almost the same noise levels since the RF source is initially tuned



Fig. 5 The Allan–Werle plot of the minimum detectable CO₂ concentration retrieved from fitting of NICE-OFCS CO₂ spectra to normalized background spectra measured in the two cavities using different modulation frequency sources (cavity 1 with DDS—*solid blue circular markers*, cavity 1 with RF source–*open cyan circular markers*, cavity 2 with DDS—*solid red square markers*, and cavity 2 with RF source—*open orange square markers*). The *dashed black lines* show the linear fits to the white-noise dominated regimes of each cavity measurement using the DDS

precisely to the correct frequency. At longer times, however, the cavity FSR drifts due to the thermal expansion of the cavity length causing a mismatch between the fixed RF frequency and the cavity FSR. This mismatch deteriorates the noise immunity and causes increased noise in the spectrum and thus worse detection limits. This is reflected in the linear drift visible in the Allan–Werle plot after 60 s for cavity 1 and after 90 s for cavity 2.

6 Conclusions

Fourier transform-based NICE-OFCS is a broadband, highly sensitive, and high-resolution spectroscopic technique that allows acquiring entire spectral bands in acquisition times on the order of a second. Thanks to the immunity of the technique to the frequency-to-amplitude noise conversion, the detection sensitivity is directly scalable with the cavity finesse. The NICE-OFCS detection scheme can be implemented in commercial FTIR instruments using a single high-speed detector at the interferometer output and standard RF components for phase-sensitive detection. The various possible comb-cavity matching solutions provide flexibility in the choice of cavity lengths and modulation frequencies for a given OFC with a fixed repetition rate. The long-term stability and low-noise performance of the system requires a perfect match of the modulation frequency to the cavity FSR. We have shown that this match

can be passively sustained by a simple approach in which the modulation frequency is generated by a DDS referenced to the repetition rate of the OFC (and thus to the cavity FSR due to the comb-cavity lock) instead of a fixedfrequency source. We achieved an absorption sensitivity of 6.4×10^{-11} cm⁻¹ Hz^{-1/2} per spectral element in a cavity with a finesse of ~9,000 and a minimum detectable concentration of 25 ppb of CO₂ in 330 s, measured at the $3\nu_1 + \nu_3$ overtone band around 1,575 nm.

The functional form of the FTS-based NICE-OFCS signal differs from that familiar from cw NICE-OHMS. In particular, there exists a term in which the molecular signal has a shape similar to direct cavity-enhanced absorption on top of a background signal, which makes the NICE-OFCS technique calibration-free (provided the cavity finesse is known). We have shown that by using a simplified form of the full model, given by this absorption term, one can extract the concentration from the measured spectra with <5 % error. Further improvements to the fitted model will lead to a better precision in concentration retrieval. Another step toward wider applicability of the NICE-OFCS technique is improving the spectral resolution by resolving the comb lines in order to allow for measurements of narrower absorption lines at lower pressures. The ability to fit a model to the experimental data provides also the possibility to utilize multiline fitting to obtain lower concentration detection limits compared to those calculated from the signal-to-noise ratio of a single absorption line in the spectrum.

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