High-resolution microwave frequency dissemination on an 86-km urban optical link

O. Lopez · A. Amy-Klein · M. Lours · C. Chardonnet · G. Santarelli

Received: 19 July 2009 / Revised version: 9 November 2009 / Published online: 21 November 2009 © Springer-Verlag 2009

Abstract We report the first demonstration of a longdistance ultra-stable frequency dissemination in the microwave range. A 9.15-GHz signal is transferred through an 86-km urban optical link with a fractional frequency instability of 1.3×10^{-15} at 1-s integration time and below 10^{-18} at one day. The optical link phase noise compensation is performed with a round-trip method. To achieve such a result we implement light polarisation scrambling and dispersion compensation. This link outperforms all the previous radio-frequency links and compares well with recently demonstrated full optical links.

1 Introduction

Ultra-stable frequency transfer between remotely located laboratories is required in time and frequency metrology, fundamental physics, particle accelerators and astronomy. Distant clock comparisons are currently performed using satellites, by two-way satellite time and frequency transfer, or through the global positioning system. Both methods are limited to a 10^{-15} fractional frequency instability for one day of averaging time [1]. This is insufficient to transfer the properties of modern cold atom microwave frequency standards, which have demonstrated frequency instability of a few 10^{-16} at one day [2, 3]. Beyond metrology, high-resolution clock comparison is essential for advanced tests

O. Lopez · A. Amy-Klein · C. Chardonnet Laboratoire de Physique des Lasers, Université Paris 13, CNRS, 99 Av. J.B. Clément, 93430 Villetaneuse, France

M. Lours · G. Santarelli (⊠) LNE-SYRTE, Observatoire de Paris, CNRS, UPMC, 61 Av. de l'Observatoire, Paris, France e-mail: giorgio.santarelli@obspm.fr in fundamental physics, such as tests of the stability of fundamental constants [4–6].

To overcome current free space link limitations, the transmission of standard frequencies over optical fibers has been investigated for several years [7-10]. This technique takes advantage of the low attenuation, high reliability and continuous availability of fibers.

Radio-frequency (RF) transmission using amplitude modulation of an optical carrier at 1 GHz has demonstrated a frequency instability as low as 5×10^{-15} at 1 s and 2×10^{-18} at one day over 86 km [11]. Direct optical frequency transfer [12–17] can provide even better stability and can be extended to greater distances. For both methods a phase noise correction is needed to compensate for the fluctuation of the propagation delay due to mechanical perturbation and temperature variation along the fiber. For this purpose the so-called round-trip method is used.

Radio-frequency transmission over an optical link has already been demonstrated at 100 MHz, 1 GHz and a few tens and hundreds of GHz [8, 9, 11, 18-20]. In this paper, we report on the transmission of a microwave frequency reference signal at 9.15 GHz over an 86-km urban fiber link connecting our two laboratories, LPL and LNE-SYRTE. This work pursues the development of stable frequency distribution of a reference signal already demonstrated between our two laboratories in the RF domain at 100 MHz [8] and 1 GHz [11]. The fiber optical length fluctuations induce phase fluctuations proportional to the modulation frequency; thus, moving to higher frequency potentially leads to an increase of the signal-to-noise ratio of the detected fiber phase fluctuations. Moreover, a microwave frequency of about 10 GHz is well suited to applications concerning particle accelerators [21] and astronomy and is close to the 9.192-GHz caesium transition frequency used for the definition of the SI second. In the following, we describe the new

set-up. Then we present the resulting performance, discuss the limitations and conclude.

2 Experimental set-up

The schematic of the optical link and compensation system is displayed in Fig. 1. The link is obtained by cascading two 43-km twin fibers connecting the laboratories LPL and LNE-SYRTE. Both ends are collocated at LPL, one labelled as local and the other as remote.

The performance of our previous optical link was limited by the signal-to-noise ratio at detection and the phase noise and long-term instability of the laser source, which induced a parasitic noise due to the fiber dispersion [11]. Thus, two main changes have been implemented. The reference carrier frequency has been increased from 1 GHz to 9.15 GHz inducing a complete change of the compensation system, laser sources and detection set-up. Moreover, a section of negative-dispersion fiber has been added to the link to compensate for its dispersion.

At each end two very low phase noise microwave synthesis systems generate all the signals necessary for the compensation system. At the local end, a 9.15-GHz yttrium iron garnet microwave oscillator (YIG) is phase locked to a low phase noise 100-MHz voltage-controlled quartz oscillator (VCXO) with a double-stage RF down-conversion approach using a 200-MHz sampling mixer [22, 23]. At the remote end a similar technique is used to phase lock a second microwave oscillator operating at 9.25 GHz by harmonic sampling to a low phase noise 1-GHz surface acoustic wave oscillator (SAW). Each microwave signal directly modulates

the beam intensity of a 6-mW distributed feedback laser diode (DFB LD) at 1.55 µm using an electro-absorption modulator (EAM), the DFB LD and the EAM being integrated on the same chip (Mitsubishi FU-653SEA). The modulation power of about 5 dBm is injected in the EAM using a bias tee to optimize the EAM operation point with a DC bias of a few volts. The modulation is detected with a fast photodiode (10-GHz bandwidth, Discovery Semiconductors DSC50S) at the other end of the fiber. We measured the phase noise of the emitter/receiver part (DFB LD, photodiode and microwave amplifier) to be $-105 \text{ dB } \text{rad}^2/\text{Hz}$ at 1-Hz offset, rolling down with a close to 1/f slope up to 1 kHz. This residual noise is compatible with a frequency instability better than 10^{-15} at 1 s. At the remote end the 9.25-GHz microwave signal is mixed with the incoming 9.15-GHz signal detected by the photodiode and the resulting beat-note signal at 100 MHz is mixed to DC with the SAW output digitally divided by 10. This error signal is used to phase lock the SAW to the incoming signal and consequently the 9.25 GHz signal with a bandwidth of a few kHz. In this way the 9.25-GHz YIG signal reproduces the oneway phase perturbations of the link. This signal modulates a second laser diode whose output is injected back into the link using an optical circulator. At the local end the 9.25-GHz backward signal is detected with a second fast photodiode (DSC50S). This signal carries the round-trip phase perturbations accumulated over the link. In order to compensate for these perturbations, the phase is revealed by successive mixing processes with the input 9.15-GHz signal and the 100 MHz from the VCXO. This generates an error signal, which is processed by a simple low-frequency loop filter



Fig. 1 Full scheme of the microwave frequency transfer. SAW: surface acoustic wave oscillator; *PLL*: phase-locked loop; *OC*: optical circulator; *EDFA*: erbium-doped fiber amplifier; *EAM-DFB LD*: electro-

absorption modulator distributed feedback laser diode, YIG: yttrium iron garnet microwave oscillator

and applied to two variable delay lines at the local input in order to cancel out the link phase perturbations.

One variable delay line is home made with a 15-m-long optical fiber wrapped around a 20-mm-diameter cylindrical piezoelectric transducer (PZT) to correct the fast and small phase perturbations [20]. Voltage applied to the PZT stretches the fiber with a dynamic range of 15 ps and a bandwidth of about 1 kHz. The second variable delay line consists of a 2.5-km-long fiber spool in an enclosure which is temperature controlled by thermoelectric modules. This device corrects slow and large phase perturbations with a dynamic range of 50°C (88 ps/°C). It enables the correction of a global temperature change of 1.5° C of the optical link, whereas the temperature change is typically less than 0.5° C per week.

The use of two different microwave frequencies for the forward and backward signals (9.15 GHz and 9.25 GHz) prevents interferences between the main signal, Brillouin backscattering and parasitic reflections from connectors and splices along the link.

The polarisation mode dispersion (PMD) and the chromatic dispersion are two detrimental fiber propagation effects that limit the performance of the correction system as discussed in [11, 24]. Both phenomena induce an asymmetry of the phase perturbation in forward and backward propagation directions along the link, with the result that the round-trip phase fluctuations do not exactly correspond to twice the forward phase fluctuations. To minimize PMD, the laser beam polarisation is directly scrambled at each emitted source by a three-axis polarisation scrambler at frequencies higher than the inverse of the light propagation delay in the overall round trip (0.5 ms). Each scrambler acts as three cascaded variable retardation waveplates excited at different resonant frequencies (approximately 60 kHz, 100 kHz and 130 kHz). This enables the exploration of all polarisation states. The chromatic dispersion of the fiber ($D \sim$ -17 ps/km/nm) converts the laser diodes' frequency noise into an excess phase noise of the optical link [11]. To reduce this effect, 11 km of highly negative dispersion fiber is inserted at the local end. Thus, the total dispersion is reduced to less than 5% of the original 86-km fiber dispersion. This also prevents periodic signal fading and extinction along the link. Without this compensation the dispersion creates a differential phase shift between the microwave sidebands of the optical signal. It amounts to $\Delta \phi = 2\pi cDL(\Omega/\omega)^2$ with D the fiber dispersion, L the fiber length and Ω and ω the microwave and optical angular frequencies, respectively. This leads to a modulation of the detected microwave signal amplitude with the fiber length with a first complete extinction at around 22 km [11].

Two optical amplifiers (EDFAs) are used to optimize the signal-to-noise ratio at detection and to compensate for the

11 dB additional optical losses introduced by the negativedispersion fiber. Despite these EDFAs and additional attenuation, the signal-to-noise ratio has not been degraded thanks to the higher modulation frequency compared to a 1-GHz system [11].

3 Results and discussion

We measured the fractional stability of the compensated link by analysing the phase variation between the local end 9.15 GHz and the remote end 9.25 GHz (Fig. 1). The beat signal at 9.25 GHz is down converted to DC by mixing with the 9.15-GHz input and the 100-MHz local oscillator. Figure 2 shows the residual phase noise spectral density of the compensated 86-km optical link (red trace) measured at 9.25 GHz, the system noise floor obtained after replacing the fiber link with an equivalent optical attenuator (black dotted trace) and the noise of the laser diode, photodiode and amplifier (blue dashed trace). The latter noise does not limit the compensation system. The red trace shows a rather dense distribution of spurious bright lines due to the polarisation scrambling process and their multi-frequency mixing products converted into phase variations. These unwanted lines are not a limiting factor for distant clocks' comparisons. However, in order to transfer an ultra-stable oscillator, it is necessary to remove the spurious lines beyond 50–100 Hz. This could be done by locking a commercially available low-phase-noise oscillator to the transmitted signal with a bandwidth around 50 Hz and reducing its phase noise close to the carrier.

Figure 3 shows the temporal behaviour of the link propagation delay in open and closed loop over four days. The open-loop signal was obtained simultaneously with



Fig. 2 Phase noise spectral density measured at 9.25 GHz of the compensated link (*red trace*), the system noise with an optical attenuator to shorten the link (*black dotted trace*) and the laser diode and photodiode noise (*blue dashed trace*)



Fig. 3 Temporal behaviour of (**a**) the propagation delay of the free-running optical link (*black trace*, one data point every 100 s) and (**b**) the end-to-end phase of the compensated link (sampled every 100 s, *blue trace*)



Fig. 4 Fractional frequency stability of (a) the free-running 86-km link (*black diamonds*), (b) the 1-GHz compensation system [11] (*blue squares*) and (c) the compensated link at 9.15 GHz (*red circles*, 15-Hz measurement frequency bandwidth)

the closed-loop phase by measuring the temperature of the 2.5-km heated spool used for the correction and converting it in phase. The free-running fiber propagation delay spans over about 2 ns while the closed-loop signal is confined well below 200 fs peak-to-peak. This demonstrates that the correction systems have a rejection factor close to 10^4 . Figure 4 shows the Allan deviation calculated from the propagation delay data measured on the 86-km compensated link. A 15-Hz low-pass filter is used to remove the spurious lines and the high-frequency phase noise of the microwave signal before phase sampling. We obtain a frequency instability of the link of 1.3×10^{-15} at 1-s integration time and much better than 10^{-18} at one-day integration time (red circles). This result is compared with previous results at 1 GHz (blue squares) [11] together with the free-running frequency sta-



Fig. 5 Fractional frequency instability of the 86-km link, in various configurations to highlight the different limitations of the system: without dispersion-compensation fiber (*blue triangles*), forward-polarisation scrambling only (*black squares*), compensated 86-km link (*red circles*), system noise floor with an optical attenuation to shorten the link (*purple diamonds*)

bility of the link (black diamonds). This demonstrates that we have significantly improved the frequency stability of the link by using amplitude modulation at higher microwave frequencies and minimising the unwanted fiber propagation effects. A linear fit calculated over the entire dataset shows a frequency bias of about 2×10^{-19} , compatible with zero within the error bars.

The ultimate link stability is still an open question and we need to analyse the different sources limiting the stability. To demonstrate the PMD deleterious effects on the fiber propagation, we have removed the polarisation scrambler at the remote end and performed a phase measurement. In this case the long-term stability of the link is severely degraded (black squares in Fig. 5) and reaches a flicker floor at around 10^{-17} . We have also replaced the negative-dispersion fiber with an equivalent optical attenuator. In this case both short-term and long-term stability are affected (blue triangles in Fig. 5). The compensation system noise floor is shown in Fig. 5 (purple diamonds), where the 86-km link has been replaced by an equivalent optical attenuator. Below 2000 s, there is a good agreement between the stability of the 86-km link (red circles) and the compensation system floor. The signal-tonoise ratio at the optical detection is limiting this noise floor at short term. The frequency instability scales with approximately a $\tau^{-2/3}$ slope. This noise level has been obtained with a careful control of the thermal environment of the electronics package and the uncommon sections of the optical paths. Over the long term, the optical link stability is no longer limited by the compensator floor. This long-term instability can be attributed to residual PMD effects or amplitude modulation to phase modulation (AM/PM) conversion in photodiodes and mixers [25-27].

4 Conclusion

We have demonstrated an 86-km compensated optical link using an urban telecom network with a frequency instability of 1.3×10^{-15} at 1 s and below 10^{-18} after one day of integration time. These results were obtained due to the use of two slightly different modulation frequencies for the two different propagation directions, the scrambling of the light polarisation state and the compensation of the fiber chromatic dispersion. The ultimate limitation of the compensator itself is reached and cannot be overcome easily. The stability of the present system enables comparison of the best frequency standards, in both the microwave and optical regions, over distances of up to 100 km. For longer distances, attenuation along the fiber is a crucial limitation. Optical amplifiers, which limit the frequency instability at a level slightly below 10^{-14} at 1 GHz, have a negligible effect at 10 GHz and can be used to recover a sufficient signalto-noise ratio for distances up to about 200 km. For longer spans, an all-optical compensation system is more promising. Modern dispersion-shifted fibers with lower PMD values allow even better performance and can potentially extend the link beyond 200 km.

Acknowledgements We acknowledge funding support from the Ministère de la Recherche and European Space Agency/ESOC.

References

- A. Bauch, J. Achkar, S. Bize, D. Calonico, R. Dach, R. Hlavac, L. Lorini, T. Parker, G. Petit, D. Piester, K. Szymaniec, P. Uhrich, Metrologia 43, 109 (2006)
- C. Vian, P. Rosenbusch, H. Marion, S. Bize, L. Cacciapuoti, S. Zhang, M. Abgrall, D. Chambon, I. Maksimovic, P. Laurent, G. Santarelli, A. Clairon, A. Luiten, M. Tobar, C. Salomon, IEEE Trans. Instrum. Meas. 54, 833 (2005)
- S. Weyers, B. Lipphardt, H. Schnatz, Phys. Rev. A 79, 031803R (2009)
- 4. J.P. Uzan, Rev. Mod. Phys. 75, 403 (2003)

- 5. S.G. Karshenboim, Can. J. Phys. 83, 767 (2005)
- 6. V.V. Flambaum, Phys. Rev. D 69, 115006 (2004)
- A. Amy-Klein, A. Goncharov, C. Daussy, C. Grain, O. Lopez, G. Santarelli, C. Chardonnet, Appl. Phys. B 78, 25 (2004)
- F. Narbonneau, M. Lours, S. Bize, A. Clairon, G. Santarelli, O. Lopez, C. Daussy, A. Amy-Klein, C. Chardonnet, Rev. Sci. Instrum. 77, 064701 (2006)
- M. Calhoun, S. Huang, R.L. Tjoelker, Proc. IEEE Spec. Issue Tech. Adv. Deep Space Commun. Track 95, 1931 (2007)
- S.M. Foreman, K.W. Holman, D.D. Hudson, D.J. Jones, J. Ye, Rev. Sci. Instrum. 78, 021101 (2007)
- O. Lopez, A. Amy-Klein, C. Daussy, Ch. Chardonnet, F. Narbonneau, M. Lours, G. Santarelli, Eur. Phys. J. D 48, 35 (2008)
- 12. N.R. Newbury, P.A. Williams, W.C. Swann, Opt. Lett. **32**, 3056 (2007)
- P.A. Williams, W.C. Swann, N.R. Newbury, J. Opt. Soc. Am. B 25, 1284 (2008)
- M. Musha, F. Hong, K. Nakagawa, K. Ueda, Opt. Express 16, 16459 (2008)
- H. Jiang, F. Kéfélian, S. Crane, O. Lopez, M. Lours, J. Millo, D. Holleville, P. Lemonde, Ch. Chardonnet, A. Amy-Klein, G. Santarelli, J. Opt. Soc. Am. B 25, 2029 (2008)
- F. Kéfélian, O. Lopez, H. Jiang, Ch. Chardonnet, A. Amy-Klein, G. Santarelli, Opt. Lett. 34, 1573 (2009)
- G. Grosche, O. Terra, K. Predehl, R. Holzwarth, B. Lipphardt, F. Vogt, U. Sterr, H. Schnatz, Opt. Lett. 34, 2270 (2009)
- M. Fujieda, M. Kumagai, T. Gotoh, M. Hosokawa, IEEE Trans. Instrum. Meas. 58, 1223 (2009)
- M. Kumagai, M. Fujieda, S. Nagano, M. Hosokawa, Opt. Lett. 34, 2949 (2009)
- B. Shillue, S. AlBanna, L. D'Addario, in Proc. IEEE Int. Top. Meet. Microwave Photonics (MWP 2004), 2004, pp. 201–204
- R. Wilcox, J.M. Byrd, L. Doolittle, G. Huang, J.W. Staples, Opt. Lett. 34, 3050 (2009)
- G.D. Rovera, G. Santarelli, A. Clairon, IEEE Trans. Ultrason. Ferroelectr. Freq. Control 43, 354 (1996)
- D. Chambon, M. Lours, F. Narbonneau, M.E. Tobar, A. Clairon, G. Santarelli, Rev. Sci. Instrum. 76, 094704 (2005)
- P. Shen, N.J. Gomes, W.P. Shillue, S. AlBanna, J. Lightwave Technol. 26, 2754 (2008)
- D. Eliyahu, D. Seidel, L. Maleki, IEEE Trans. Microwave Theory Tech. 56, 449 (2008)
- G. Cibiel, M. Régis, E. Tournier, O. Llopis, IEEE Trans. Ultrason., Ferroelectr. Freq. Control 49, 784 (2002)
- L.M. Nelson, F.L. Walls, in *Proc. IEEE Frequency Control Symp.*, 1992, pp. 831–837. DOI:10.1109/FREQ.1992.269954