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# Variability of the microbial community in the western Antarctic Peninsula from late fall to spring during a low ice cover year

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Abstract Although winter conditions play a major role in determining the productivity of the western Antarctic Peninsula (WAP) waters for the following spring and summer, a few studies have dealt with the seasonal variability of microorganisms in the WAP in winter. Moreover, because of regional warming, sea-ice retreat is happening earlier in spring, at the onset of the production season. In this context, this study describes the dynamics of the marine microbial community in the Melchior Archipelago (WAP) from fall to spring 2006. Samples were collected monthly to biweekly at four depths from the surface to the aphotic layer. The abundance and carbon content of bacteria, phytoplankton and microzooplankton were analyzed using flow cytometry and inverted microscopy, and bacterial richness was examined by PCR–DGGE. As expected, due to the extreme environmental conditions, the microbial community abundance and biomass were low in fall and

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winter. Bacterial abundance ranged from 1.2 to 2.8  $\times$  $10<sup>5</sup>$  cells ml<sup>-1</sup> showing a slight increase in spring. Phytoplankton biomass was low and dominated by small cells ( $\langle 2 \mu m \rangle$ ) in fall and winter (average chlorophyll a concentration, Chl-a, of, respectively, 0.3 and 0.13  $\mu$ g l<sup>-1</sup>). Phytoplankton biomass increased in spring (Chl-a up to 1.13  $\mu$ g l<sup>-1</sup>), and, despite potentially adequate growth conditions, this rise was small and phytoplankton was still dominated by small cells  $(2-20 \mu m)$ . In addition, the early disappearing of sea-ice in spring 2006 let the surface water exposed to ultraviolet B radiations (UVBR, 280–320 nm), which seemed to have a negative impact on the microbial community in surface waters.

Keywords Microbial food web · Antarctica · Sea-ice · Ozone hole - UV radiations

## Introduction

The Southern Ocean, and hence the WAP waters, undergo extreme seasonal fluctuations in terms of light regime, seaice concentration and productivity (Delille [2004\)](#page-13-0). However, most oceanographic studies in the WAP have taken place during the more productive periods (i.e. late spring– early summer) when strong phytoplankton blooms occur (e.g. Smith et al. [2008](#page-15-0); Vernet et al. [2008;](#page-15-0) Reiss et al. [2009](#page-14-0)). Indeed, a few studies have dealt with the seasonal variability of microorganisms in the WAP through winter, and these studies focused on the distribution and diversity of bacterioplankton (Murray et al. [1998;](#page-14-0) Church et al. [2003](#page-13-0)), microzooplankton grazing and phytoplankton growth (Brightman and Smith [1989](#page-13-0); Pearce et al. [2008](#page-14-0)) and the distribution of krill (Lawson et al. [2004](#page-14-0)). This is attributable to very difficult sampling conditions during

winter and to a specific interest in highly productive periods. Nevertheless, because winter conditions will in part determine the production for the following spring and summer, winter studies are necessary. Indeed, winter seaice conditions are a major determinant of the productivity of WAP waters during the subsequent spring and summer (Smith et al. [1998;](#page-14-0) Smith et al. [2001](#page-15-0)). Poor winter sea-ice conditions may also favor the presence of salps, which have a low nutritive value for higher predators, over krill, which represent a major link to higher trophic levels in the WAP (Siegel and Loeb [1995;](#page-14-0) Loeb et al. [1997\)](#page-14-0).

Because of global warming and seasonal stratospheric ozone layer breakdown, the WAP is subjected to changing environmental conditions, and more particularly, during winter and spring. The ''ozone hole'', which refers to a strong reduction in the concentration of stratospheric ozone, has been observed every spring over Antarctica for the last 20 years (McKenzie et al. [2007](#page-14-0)). In consequence, the intensities of UVBR reaching the surface of the Southern Ocean have increased (Arrigo [1994](#page-13-0); Frederick and Lubin [1994\)](#page-13-0). On the other hand, the WAP has experienced a significant rise in air temperatures during the last 50 years  $(+0.56^{\circ}C$  per decade; Marshall et al. [2002](#page-14-0)). This rise was important mostly in winter  $(+1.09^{\circ}C$  per decade; Turner et al. [2005](#page-15-0)) and coincided with a decrease in the duration of sea-ice cover and more particularly an earlier retreat of sea-ice (Stammerjohn et al. [2008](#page-15-0)). As a consequence, sea-ice retreat may now coincide with the seasonally occurring ozone hole and lead to potentially harmful conditions where no or low ice cover can shield the water column from harmful UVBR (Cockell and Córdoba-Jabonero [2004](#page-14-0); Lesser et al. 2004).

UVBR and ultraviolet A radiation (UVAR, 320– 400 nm) are recognized to have significant impacts on marine organisms, including DNA damage (Häder and Sinha [2005\)](#page-13-0) and photoinhibition (e.g. Cullen et al. [1992](#page-13-0); Neale et al. [1998\)](#page-14-0). These radiations may also affect indirectly the microbial community through trophic interactions (Mostajir et al. [1999](#page-14-0); Davidson and Belbin [2002](#page-13-0); Ferreyra et al. [2006](#page-13-0)). On the other hand, global warming may have both positive and negative effects on the microbial community of the WAP. Higher temperatures should promote higher growth and photosynthesis rates and should favor enzyme-based photorepair mechanisms (Bouchard et al. [2006\)](#page-13-0). Moreover, the predicted earlier retreat of sea-ice in the WAP may lead to a longer growth season, as long as nutrients are available. Under this assumption, it has been hypothesized that overall plankton biomass will be higher in the future for the Southern Ocean (Arrigo and Thomas [2004;](#page-13-0) Sarmiento et al. [2004\)](#page-14-0) as it has already been modeled for the Arctic Ocean (Arrigo et al. [2008\)](#page-13-0). However, by changing the dynamics of sea-ice and water column stratification, global warming might also have negative effects on the microbial community of the WAP (Sarmiento et al. [2004;](#page-14-0) Häder et al. [2007\)](#page-13-0). Finally, earlier ice melting due to global warming may expose the growing phytoplankton populations to high UVBR during the ozone hole period (Häder et al.  $2007$ ).

The present study was conducted in the WAP over the 2006 austral late-fall, winter and early-spring seasons in the Scholaert Channel, in the vicinity of the Melchior Archipelago. The aim of this research was to study the dynamics of the microbial community during this time of the year when productivity is minimum (late fall and winter) and at the onset of the production season (early spring) when the ozone hole may threaten the WAP microbial community.

## Materials and methods

### Sampling

This project was undertaken in a shallow bay (maximum depth of 150 m) of the Melchior Archipelago (Scholaert Channel) in the WAP (Fig. [1\)](#page-2-0), where the Sedna IV sailboat wintered from June 3rd to November 10th 2006. A GUV-510 surface spectral radiometer (Biospherical Instruments Inc., USA; later on abbreviated as GUV) was fixed on the Sedna IV boat at a shadow-free site to monitor incident UVBR, UVAR and photosynthetically available radiation (PAR, 400–700 nm) every 15 min during the whole study period. Sampling was performed aboard rubber boats at a fixed station (64°19′31″S, 62°53′42″W) located ~1 km from the coast. Vertical profiles of water column characteristics were obtained using a Conductivity, Temperature and Depth data logger (CTD; model SBE 19 Plus, Sea-Bird Electronics, Inc.) equipped with a fluorescence sensor (Seapoint Chlorophyll Fluorometer; Seapoint Sensors, Inc.). Light penetration within the water column was calculated from vertical profiles using a PUV-542T spectral radiometer (Biospherical Instruments Inc., USA; later on abbreviated as PUV). The GUV and the PUV instruments were set to measure the irradiances of wavelengths both in the visible (PAR, in  $\mu$ E cm<sup>-2</sup> s<sup>-1</sup>) and in the ultraviolet range (305 and 313 nm for UVBR, and 320, 340 and 380 nm for UVAR, in  $\mu$ W cm<sup>-2</sup>). The PUV profiler was cast in the water column before each sampling to determine sampling depths. The four sampling depths consisted of three depths in the euphotic zone (100, 50 and 0.1% of incident PAR) and one depth within the aphotic zone  $(Z_{\text{aph}}$ . 50 m below 1% of incident irradiance). During the last four sampling dates in spring, the second depth (50% of incident irradiance) was modified to sample the maximum of Chl-a. Moreover, when the difference between two preset depths was less than 10 m, only one sample was taken at an average in-between depth (see Fig. [2](#page-2-0) for details of all

<span id="page-2-0"></span>Fig. 1 The western Antarctic Peninsula (WAP) and the position of the sampling site north of Anvers Island (cross)





Fig. 2 Sampling depths for each day of sampling (given in Julian days). Depths corresponded to the penetration of light within the water column:  $100\%$ ,  $50\%$  (or maximum of Chl-a) and  $0.1\%$  of incident irradiance and  $Z_{\text{aph}}$ 

sampling depths). Water was then sampled using 5 l Niskin bottles.

Samples for the analysis of nutrients (nitrate, nitrite, phosphate and silicate) were filtered through precombusted Whatman GF/F filters, and filtrates were kept frozen  $(-20^{\circ}C)$  until analysis. For phytoplankton pigment analysis, 950–1,000 ml of seawater was filtered on a 25-mmdiameter Whatman GF/F glass fiber filter. Filters were then wrapped in aluminum foil, immediately frozen in liquid nitrogen and stored at  $-80^{\circ}$ C until analysis. Water aliquots for flow cytometric analyses of bacteria and picophytoplankton and nanophytoplankton were fixed with glutaraldehyde (final concentration 0.1%) into 4.5-ml cryovials in triplicate and stored at  $-80^{\circ}$ C. Samples for phytoplankton (nanophytoplankton and microphytoplankton) and heterotrophic protist identification and enumeration were fixed in acid iodine Lugol solution at 2% final v/v concentrations and kept at  $4^{\circ}$ C in 300-ml glass amber bottles. Bacterial richness was estimated from two 500-ml aliquots of seawater filtered on white polycarbonate membrane (PC) of  $25 \text{ mm}$  diameter and  $0.2 \mu \text{m}$  pore size  $(ISOPORE<sup>TM</sup>)$ , and filters were subsequently stored at  $-80^{\circ}$ C in sterile Petri dishes. Bacterial richness was only analyzed at the two extreme depths (the surface and  $Z_{\text{aph}}$ ).

## Analysis of physical conditions

Daily stratospheric ozone concentration over the study site was obtained from [http://toms.gsfc.nasa.gov/.](http://toms.gsfc.nasa.gov/) The ozone hole was considered to be present over the study site when stratospheric ozone thickness was below 220 Dobson Units (DU) according to NASA's definition ([http://ozonewatch.](http://ozonewatch.gsfc.nasa.gov/) [gsfc.nasa.gov/](http://ozonewatch.gsfc.nasa.gov/)). Sea-ice concentration data were obtained from NASA's Nimbus-7 Scanning Multichannel Microwave Radiometer (SMMR) and Defense Meteorological Satellite Program (DMSP) -F13 Special Sensor Microwave/Imager (SSM/I). Daily sea-ice percent concentration for the year 2006 was obtained from the EOS Distributed Active Archive Center (DAAC) at the National Snow and Ice Data Center, University of Colorado in Boulder, Colorado [\(http://nsidc.org](http://nsidc.org)) (Cavalieri et al. [2008](#page-13-0)). Sea-ice concentration was defined as the percent cover of sea-ice on a  $25 \times 25$  km square around the sampling site. For the purpose of this study, sea-ice retreat was defined as the last day of the year that sea-ice concentration was above 15% (Stammerjohn et al. [2008](#page-15-0)).

The fluorescence profiles were calibrated with extracted values of Chl-a, determined by HPLC (see below). For each sampling day, Chl-a concentrations were depth-integrated from the surface to the bottom of the euphotic layer,

<span id="page-3-0"></span>and mean values were calculated by dividing the depthintegrated value by the integration depth (Heywood et al. [2006\)](#page-13-0). The stability of the water column was estimated as in Sabatini et al. [\(2004](#page-14-0)) from the Simpson stability parameter (Simpson [1981](#page-14-0)), which is a measure of the mechanical work required to vertically mix the water column. A value  $>40$  J m<sup>-3</sup> for the Simpson parameter was considered to define a well-stratified water column (Sabatini et al. [2004](#page-14-0)). The depth of the upper mixed layer  $(Z<sub>UMI</sub>)$  was measured using the threshold gradient method described in Thomson and Fine [\(2003](#page-15-0)).

The attenuation coefficients,  $K<sub>d</sub>$ , were calculated from the Beer–Lambert's law (Kirk [1983](#page-14-0)). The 10 and 1% penetration depths of incident light corresponding to each wavelength were calculated as:  $Z_{10\%} = 2.3/K_d$  and  $Z_{1\%} = 4.6051/K_d$  (Kirk [1983\)](#page-14-0). The depth of the euphotic zone was defined as  $Z_{1\%}$  (PAR) ( $Z_{eu}$ ), and the photoactive depth of each UV radiation (UVR) wavelength  $(Z_{ph}(\lambda))$ ; i.e. the depth at which the effects of UVBR and UVAR cease) as  $Z_{10\%}$  (Neale et al. [2003](#page-14-0)). Because the water column was very clear during fall, winter and early spring (i.e. the sampling period), the attenuation of light within the water column did not show important variation. Therefore, the average  $K_d(\lambda)$  is presented in Table 1 along with the depth penetration of each wavelength.

The penetration of UVR within the water column was difficult to determine during austral winter due to the very low light reaching the surface of the ocean and the PUV detection limits. Because the GUV recorded the incident irradiance above the water surface every day, it was possible to estimate the irradiance just below the water surface for every day of the sampling period from GUV records. To do so, the average  $K_d$  of each wavelength was considered and the transmission of light through the water surface was estimated. The penetration of light depends upon the illumination conditions (i.e. direct versus diffuse sky light) and the sea roughness (Arrigo [1994](#page-13-0)). While the transmittance of diffuse light is nearly constant, with an approximate value of 93.4% (Morel [1991\)](#page-14-0), the transmittance of direct light strongly depends upon the solar zenith angle and the sea surface roughness (Haltrin et al. [2000](#page-13-0)). However, the GUV recorded the total incident irradiance and did not differentiate between direct and diffuse light. The large specular reflexion values at a high sun zenith angle (e.g.  $60^{\circ}$ ), however, are compensated by the increasing proportion of the diffuse component of the total incident irradiance. Calculations show that the transmittance coefficients of total incident irradiance for a sun zenith angle  $>60^\circ$  under a clear sky are, on average, about 84% in the visible (500 nm) and 93% in the UVBR (305 nm; S. Bélanger, personal communication). To simplify, a transmittance value of 90% was chosen to estimate the irradiance just below the water surface from GUV records for the whole sampling period. This method was tested by comparing the results with the actual measures of each wavelength obtained with the PUV when the water was sampled. The method proved to be a good approximation for each wavelength with all correlation coefficients between 0.8 and 0.85 ( $P < 0.01$ ).

Because sea-ice cover is a major physical barrier, the penetration of light within the water column will be considered insignificant when sea-ice cover was above 15%. To determine the effect of UVBR on surface waters, each UVR wavelength was weighted for its potential impact on microorganisms DNA (Setlow [1974\)](#page-14-0). Moreover,  $E_{\text{ph}}$ , the total irradiance integrated from the surface to the photoactive depth of each UVR wavelength  $(Z_{\text{ph}})$ , was calculated following the equation of Neale et al. [\(1991](#page-14-0)) that was slightly modified:

$$
\int\limits_0^{z_{ph}} E(\lambda).dz = \frac{E_{0-}(\lambda)}{K_d(\lambda)}\Big(1-e^{-K_d(\lambda).z_{ph}(\lambda)}\Big)
$$

with  $E_0(\lambda)$  the incident irradiance just below the water surface and  $E_z(\lambda)$  the irradiance at depth z. Therefore, the potential impact of UVR on microorganisms' DNA could be estimated for the surface waters (i.e. to the depth of the photoactive zone). The results obtained were correlated with ozone concentration to determine the effect of the ozone hole in increasing UVBR transmission through the atmosphere. As expected,  $E_{\rm ph}$  only correlated significantly with ozone thickness for UVBR wavelengths (i.e. with 313 nm,  $r = -0.44$ ,  $P < 0.01$ ). The penetration of 305 nm within the atmosphere is more sensitive to the concentration of stratospheric ozone than 313 nm. Therefore, one would expect  $E_{\rm ph}$  (305 nm) to correlate with ozone concentration. However, because 305 nm is attenuated in the water column much faster than 313 nm and because the photoactive depth of 305 nm (7 m) is shallower than that of 313 nm (10 m), 313 nm was a better representative of the actual level of UVBR in the surface waters.

**Table 1** Average  $K_d(\lambda)$  ( $\pm$ SD) for UVBR, UVAR and PAR and corresponding depth ( $m \pm$  SD) for the penetration of 10% of UVBR and UVAR  $(Z_{\text{ph}})$  and of 1% of PAR  $(Z_{\text{eu}})$  in the water column

	PAR	$305 \text{ nm}$	313 nm	$320 \text{ nm}$	340 nm	380 nm
$K_{d}(\lambda)$		$0.06 \pm 0.02$ $0.33 \pm 0.03$ $0.24 \pm 0.04$ $0.19 \pm 0.03$ $0.15 \pm 0.02$ $0.09 \pm 0.01$				
$Z_{\rm ph}$ (UVBR & UVAR) and $Z_{\rm eu}$ (PAR) in meters	$72 \pm 24$	$7 \pm 0.56$			$9.6 \pm 1.46$ $12.4 \pm 2.27$ $15.7 \pm 1.74$ $25.4 \pm 3.39$	

### Sample analysis

Nutrient concentrations were determined using a Bran Luebbe AutoAnalyzer 3 system following Grasshof et al. [\(1983](#page-13-0)). Phytoplankton pigments were measured by high performance liquid chromatography (HPLC) using a Thermo Fisher system according to the method of Zapata et al. [\(2000](#page-15-0)). Pigments were detected using a photodiodearray spectrophotometric detector (UV-6000) in series with a FL 3000 fluorescence detector. Pigments were identified using their retention time, visible spectrum and comparison with standards from DHI Water and Environment (Hørsholm, Denmark).

Bacteria, picophytoplankton and nanophytoplankton were enumerated using an EPICS<sup>®</sup> ALTRA<sup>TM</sup> cell sorting flow cytometer (Beckman Coulter<sup>®</sup> Inc.) equipped with a laser emitting at 488 nm. Fluorescent beads (Fluoresbrite YG microspheres 1 and 1.9  $\mu$ m for bacteria and phytoplankton, respectively, Polysciences<sup>TM</sup>) were systematically added to each sample as an internal standard to normalize cell fluorescence emission and light scatter values. For the analysis of bacterial abundance, frozen samples were thawed, and two subsamples of 1 ml were half-diluted in TE  $10\times$  buffer (100 mM Tris–HCl, 10 mM EDTA, pH 8.0). One ml of the resulting diluted sample was stained with 0.3  $\mu$ l of SYBR® Green I nucleic acid gel stain  $(Ci = 10,000 \times$ , Invitrogen, Inc., incubated for 10 min at room temperature in the dark (Lebaron et al. [2001\)](#page-14-0) and analyzed for 240 s. For the analysis of phytoplankton abundance, frozen samples were thawed, and 1 ml of the sample was analyzed for 300 s. To calculate bacterial and phytoplankton cell abundances, the volume analyzed was calculated by weighing samples before and after each run. Total free bacteria (TB) were detected in a plot of green fluorescence recorded at  $530 \pm 30$  nm (FL1) versus side angle light scatter (SSC). Bacteria with high and low nucleic acid content (HNA and LNA subgroups, respectively) were discriminated by gating the FL1-versus-SSC plot, and respective abundances of both subgroups were determined (Lebaron et al. [2001\)](#page-14-0). For the purpose of this study, TB abundance was used to describe the bacterial community distribution, and %HNA (the ratio of HNA cells on TB) was used to describe the physiological structure of the bacterial community as suggested by different studies (Gasol et al. [1999;](#page-13-0) Gasol and Giorgio [2000](#page-13-0); Lebaron et al. [2001](#page-14-0); Vaqué et al. [2002](#page-15-0)). Picophytoplankton and nanophytoplankton were detected in a plot of red fluorescence recorded at  $675 \pm 5$  nm (FL4) versus forward side scatter (FSC). Picophytoplankton was defined as phytoplankton smaller than 2 µm and nanophytoplankton as phytoplankton comprised between  $2$  and  $20 \mu m$ . All cytometric analyses were performed under the Expo 32 v1.2b software (Beckman Coulter<sup>®</sup> Inc.). The size of organisms was determined through the construction of a calibration curve using 1, 2, 4, 5.9, 9.7 and 15.5  $\mu$ m diameter large polystyrene microspheres (Molecular Probes Inc.). The carbon content of microorganisms, based on their abundance and size measured by flow cytometry, was calculated using different conversion factors for the different size classes: 220 fg C  $\mu$ m<sup>-3</sup> for nanophyto-plankton (Tarran et al. [2006\)](#page-15-0), 1.5 pg C cell<sup>-1</sup> for picophytoplankton and 12 fg C cell<sup>-1</sup> for bacteria (Zubkov et al. [2000a](#page-15-0), [b](#page-15-0)).

Furthermore, nanophytoplankton, microphytoplankton and microzooplankton were identified and enumerated on Lugol-fixed samples using a Leitz Diavert inverted, phasecontrast microscope. Due to time constraints, only samples from the second depth, corresponding to the maximum of Chl-a, were analyzed. The biovolume (V) of cells was calculated using the geometric shapes proposed by Hillebrand et al. [\(1999](#page-13-0)) and corrected to account for cell shrinkage caused by fixation of samples (Montagnes et al. [1994](#page-14-0)). The carbon content of cells was calculated with two different carbon-to-volume ratios: pg C cell<sup>-1</sup> = 0.288  $V^{0.811}$  for diatoms and pg C cell<sup>-1</sup> = 0.216 V<sup>0.939</sup> for all other groups (Menden-Deuer and Lessard [2000\)](#page-14-0).

Bacterial richness was analyzed at the two extreme depths (100% of incident irradiance and  $Z_{\text{anh}}$ ). Total DNA extraction was performed using a classic Phenol–chloroform–isoamyl alcohol  $(25/24/1; 900 \mu l)$  protocol. PCR amplification of the 16S rDNA gene was then performed using a Mastercycler epS (Eppendorf) thermal cycler following the method proposed by Schäfer and Muyzer [\(2001](#page-14-0)). Three PCR amplifications were performed on each DNA sample to overcome the effect of PCR biases (Perreault et al. [2007\)](#page-14-0). Amplicons were then purified with the MinElute (QIAGEN) columns according to the manufacturer's instructions and stored at  $-20^{\circ}$ C until denaturing gradient gel electrophoresis (DGGE) analysis. DGGE was performed using a DGGE-4001-Rev-B (C.B.S. Scientific Company, CA, USA) system according to Schäfer and Muyzer ([2001\)](#page-14-0). Gels were then stained with a half-diluted solution of SYBR Green I (10,000 $\times$ , Molecular Probes, Oregon) for 1 h according to the manufacturer's instructions. Gels were photographed under UV light, and DGGE profiles were analyzed using an AlphaImager<sup>®</sup> HP (Alpha-Innotech). The number of bands, corresponding to different operational taxonomic units (OTU), was determined visually for each sample. A similarity matrix using Jaccard's distance index was used to compare the fingerprints.

## Statistical analyses

Seasons were defined according to the southern hemisphere winter solstice (June 21) as well as the spring equinox

(September 22). Statistical analyses were performed using the Statistica software (StatSoft<sup>®</sup> Inc). Normality was tested using Kolmogorov–Smirnov's and Shapiro–Wilk's tests, and homoscedasticity was verified using Levine's test. When normality and/or homoscedasticity were not met, data were transformed. Pearson (r) or Spearman rank  $(r<sub>s</sub>)$  correlation coefficients were calculated to study the interdependence between abiotic (PAR, UVBR, UVAR, nutrient concentration, temperature, salinity, dissolved organic carbon concentration (DOC, data not shown) and particulate organic carbon and nitrogen concentrations (POC and PON, data not shown)) and biotic variables (TB, %HNA, pico-, nano-, microphytoplankton and microzooplankton abundances). The temperature and salinity of each sampling depth were obtained from the CTD profiles. The concentrations of DOC, POC and PON were obtained from Wang et al. [\(2009](#page-15-0)). Differences between depths were analyzed by t tests, whereas seasonal differences for biological and physical parameters were tested with ANOVAs.

To study the effects of stratospheric ozone loss and related UVBR increase on microorganisms in the field, the ratio between the abundance of organisms at the surface and at a sampling depth immediately below the UML was correlated with  $E_{\text{ph}}$  (313 nm), which best reflects the loss of stratospheric ozone and the increase in UVBR penetrating through the water column. These calculations were performed for bacterioplankton, %HNA, bacterial richness and for picophytoplankton and nanophytoplankton only because microphytoplankton and microzooplankton data only existed at the depth of 50% of incident irradiance. The rationale for the use of such biological ratios is that the changes in the relative abundance of microorganisms above and below the upper mixed layer would reflect significant effects of UVBR in surface waters (Helbling et al. [2005;](#page-13-0) Hernando and Ferreyra [2005\)](#page-13-0). The ratios for picophytoplankton and nanophytoplankton were  $log_{10}$ -transformed to meet the assumption of normality and homoscedasticity, while  $E_{ph}$  (313 nm) was arcsine (square root) transformed.

## Results

#### Environmental conditions

In Fig. [3a](#page-6-0), the irradiances at 305, 313 and 380 nm, estimated just under the water surface, are presented for the sampling period. An important daily variation in irradiance, related to cloud cover and albedo, can be observed for all wavelengths. As expected, an increase in UVAR (380 nm) was observed from winter to spring (i.e. a slope of 0.21  $\mu$ W cm<sup>-2</sup> day<sup>-1</sup>). Although not presented in Fig. [3](#page-6-0)a,

the same pattern was observed for both 320 and 340 nm. Shorter wavelengths such as 305 and 313 nm also increased regularly during spring time, but their rise was less pronounced and punctuated by episodes of significant elevation. Because 305 and 313 nm correlated negatively with ozone thickness  $(r = -0.66, P < 0.01$  and  $r = -0.47$ ,  $P < 0.01$  for 305 and 313 nm, respectively), we can assess that these episodes of increased UVBR corresponded to both ozone loss periods and the onset of the spring season. Indeed, a significant decrease in ozone concentration (i.e. an ozone hole with ozone thickness \220 DU) was observed over the study site from the 31st of August to the 9th of October 2006 (day 243 to day 282 included; Fig. [3](#page-6-0)b). During this 40-day period, the minimum ozone thickness observed was 132 DU, and ozone thickness was below 220 DU on 33 occasions (averaging  $192.9 \pm 37$  DU).

Sea-ice covered the sampling area during winter, except from the 10th to the 16th of September (day 253 to day 259 included) when it suddenly retreated (Fig. [3c](#page-6-0)). Ozone thickness was below 220 DU on 6 occasions during this transient retreat. Sea-ice then covered again the sampling area from the 17th of September to the 12th of October (day 260 to day 285 included), and therefore, sea-ice still covered the sampling area 4 days after the ozone hole disappeared. The water column was thus protected from increased UVBR by sea-ice cover during most of the ozone hole period. However, as seen in Fig. [3](#page-6-0)b, because of the onset of the spring season and because ozone recovery was not immediate after the ozone hole period, high intensities of UVBR penetrated the water column for the remaining of the sampling period. Indeed, high values of  $E_{\rm ph}$  (313 nm) were encountered between the 13th of October and the 10th of November (day 286 to day 314, Fig. [3](#page-6-0)b). For example, during this period, the intensity at 320 nm just below the water surface varied from 4 to 17.8  $\mu$ W cm<sup>-2</sup> (with an average of 9.7  $\mu$ W cm<sup>-2</sup>). As a comparison, Helbling et al. [\(1995](#page-13-0)) found that only 3 and 0.1% of two bacterial strains (Acinetobacter sp. and Bacillus sp. respectively) could survive a radiation intensity of 9.6  $\mu$ W cm<sup>-2</sup> at 320 nm. In addition, Helbling et al. ([1994\)](#page-13-0) found that an intensity of 6  $\mu$ W cm<sup>-2</sup> at 320 nm at the surface could reduce primary production by 85%.

The Simpson parameter ranged between 21 and [3](#page-6-0)2 J.m<sup>-3</sup> during fall (average  $Z_{UML}$  of 35 m; Fig. 3d) and averaged 12.4 J m<sup>-3</sup> during winter (average  $Z_{UML}$  of 48 m). During spring, the Simpson parameter increased regularly and was above 40 J  $\text{m}^{-3}$  after the 23rd of October (day 296). As a consequence, the UML became shallower in spring and was 22.6 m deep on average and as shallow as 16 m on day 296 (23rd of October). In addition, the photoactive depth,  $Z_{ph}$ , for UVBR laid from 10 to 12 m deep in spring, and the UML was therefore highly exposed to UVBR at this time of the year (Table [1](#page-3-0)).

<span id="page-6-0"></span>Fig. 3 a Incident irradiance at 305, 313 and 380 nm estimated below the water surface at local noon time ( $\mu$ W cm<sup>-2</sup>); b Stratospheric ozone thickness (DU) and  $E_{\rm ph}$  (313 nm); c Seaice concentration (%) and d Depth of the UML (m, black dashed line) and contour plot of  $\sigma_t$  for the sampling period



Surface water temperature was always below zero during the study. It decreased from fall (average of  $-0.78 \pm 0.2$ °C) to winter (average of  $-1.62 \pm 0.13$ °C), before increasing in spring (average of  $-1.04 \pm 0.47$ °C). Surface water salinity followed an opposite pattern, increased from fall (average of  $33.91 \pm 0.05$  PSU) to winter (average of  $34.04 \pm 0.04$  PSU), before decreasing in spring (average of  $33.94 \pm 0.11$  PSU). During this study, average concentrations of phosphate, nitrate and nitrite, and silicate through the water column were  $2.3 \pm 0.3$ ,  $27.7 \pm 1.8$  and  $37 \pm 8.5$  µM, respectively.

Across all depths, nutrient concentrations did not vary significantly ( $P > 0.05$ , Table [2\)](#page-7-0) except at the surface in spring where nitrite and nitrate concentration decreased with a minimum observed concentration of  $23.1 \mu M$ .

Bacterioplankton seasonal and vertical distributions and richness

During our survey, TB abundance ranged from  $1.2 \times 10^5$ to 2.8  $\times$  10<sup>5</sup> cells ml<sup>-1</sup> (Fig. [4](#page-8-0)a; Table [3\)](#page-9-0), decreased with depth ( $r_s = -0.47$ ,  $P \lt 0.01$ ), was significantly lower in

<span id="page-7-0"></span>Table 2 Average nutrient concentration ( $\mu$ M  $\pm$  SD) for each season (fall, winter and spring) and depth (100, 50 or 0.1% of incident irradiance and  $Z_{\text{sub}}$ )

		$Phosphate$ Nitrite + Nitrate Silicate	
100%			
Fall		$1.98 \pm 0.04$ 25.83 $\pm$ 1.24	$36.83 \pm 15.45$
Winter		$2.02 \pm 0.18$ 27.87 $\pm$ 0.58	$34.97 \pm 9.09$
Spring		$2.01 \pm 0.08$ 25.04 $\pm$ 1.09	$46.58 \pm 11.99$
$50\%$ or max. Chl-a			
Fall		$2.05 \pm 0.07$ 25.45 $\pm$ 0.35	$47.93 \pm 1.8$
Winter		$2.02 \pm 0.79$ 27.84 $\pm$ 1.71	$37.17 \pm 4.31$
Spring		$2.39 \pm 0.08$ 28.58 $\pm$ 0.72	$34.50 \pm 4.8$
$0.1\%$			
Fall		$2.15 \pm 0.28$ 28.05 $\pm$ 0.99	$33.55 \pm 1.84$
Winter		$2.11 \pm 0.14$ 27.64 $\pm$ 0.34	$30.3 \pm 1.1$
Spring		$2.42 \pm 0.14$ 29.73 $\pm$ 0.94	$37.84 \pm 6.83$
$Z_{\rm aph}$			
Fall	$2.28 \pm 0.11$	$28.58 \pm 0.25$	$33.88 \pm 1.38$
Winter		$1.94 \pm 0.49$ 28.28 $\pm$ 0.88	$32.79 \pm 3$
Spring	$2.55 \pm 0.2$	$27.98 \pm 1.58$	$31.82 \pm 3.17$

winter ( $P < 0.01$ ), more specifically for 100 and 50% of incident irradiance ( $P \lt 0.01$  and  $P \lt 0.05$ , respectively), and increased from winter to spring. At the surface and 50% of incident irradiance, TB correlated positively with total phytoplankton abundance  $(r = 0.64$  and  $r = 0.73$ ,  $P < 0.01$ ), nanophytoplankton abundance ( $r = 0.55$ ,  $P < 0.05$  and  $r = 0.65$ ,  $P < 0.01$ ) and water temperature  $(r = 0.66, P < 0.05 \text{ and } r = 0.64, P < 0.01)$ . In contrast, no significant seasonal variations in TB were observed at 0.1% of incident irradiance and  $Z_{\text{aph}}$ . In addition, the biomass of bacterioplankton at 50% of incident irradiance increased significantly  $(P < 0.01)$  from winter to spring (average of 2.8  $\pm$  0.44 µg l<sup>-1</sup>; Fig. [5](#page-10-0)a).

The bacterial community was constantly dominated by HNA (Table [3\)](#page-9-0), and %HNA increased significantly with depth  $(r<sub>s</sub> = 0.32, P < 0.05)$ . No seasonal variations in %HNA were observed, and %HNA was not related to any biotic or abiotic variables. On the other hand, bacterial richness averaged  $52.4 \pm 6.9$  OTUs per sample during the whole sampling period (Table [3\)](#page-9-0). The Jaccard index showed a gradual decrease in bacterial richness in surface samples through the whole survey, while bacterial richness increased at  $Z_{\text{aph}}$  (Table [3\)](#page-9-0). The decrease in bacterial richness at the surface was mainly attributable to the disappearance of non-dominant OTUs (represented by low intensity bands), while the contrary was true for  $Z_{\text{anh}}$ . Moreover, bacterial richness at the surface in spring was negatively correlated with  $E_{\text{ph}}$  (313 nm) ( $r = -0.77$ ,  $P < 0.01$ ), suggesting a possible cause for the decrease in bacterial richness at the surface in the spring.

Phytoplankton seasonal and vertical distribution and composition

Mean Chl-a integrated over the water column remained low during the whole period of study and never exceeded 0.3  $\mu$ g l<sup>-1</sup>, averaging 0.13  $\mu$ g l<sup>-1</sup> during fall and winter and 0.22  $\mu$ g l<sup>-1</sup> in spring (Fig. [4b](#page-8-0)). As shown in Fig. 4b, a sub-surface Chl-a maximum developed during spring time between  $\sim$  30 and  $\sim$  50 m deep (43 m on average), and the maximum estimated Chl-a concentration was 1.13  $\mu$ g l<sup>-1</sup>.

In terms of size classes, microphytoplankton was rare, and phytoplankton was always dominated numerically by picophytoplankton and nanophytoplankton from fall to spring. Microphytoplankton abundance at the depth of 50% irradiance averaged 1.38  $\pm$  1.4 cells ml<sup>-1</sup> throughout the sampling period (Table [3](#page-9-0)) and was higher in spring (average of 2.06  $\pm$  1.69 cells ml<sup>-1</sup>) than during either fall or winter (average of  $0.45 \pm 0.11$  and  $0.73 \pm 0.47$ cells  $ml^{-1}$ , respectively). Nanophytoplankton abundance ranged from 0.13 to 4.56  $\times$  10<sup>3</sup> cells ml<sup>-1</sup> and was significantly higher in spring than during either fall or winter  $(P<0.01)$ . More specifically, nanophytoplankton abundance increased significantly across seasons at the surface and 50% of incident irradiance  $(P < 0.01;$  Fig. [4c](#page-8-0); Table [3](#page-9-0)). In addition, nanophytoplankton abundance showed a positive correlation with PAR  $(r = 0.52)$ ,  $P < 0.05$ ) for the surface, suggesting that it responded to increasing light availability in spring. Picophytoplankton abundance ranged from 0.09 to 2.24  $\times$  10<sup>3</sup> cells ml<sup>-1</sup> and was more abundant at the surface and 50% of incident irradiance (ranging from 0.41 to 2.24  $\times$  10<sup>3</sup> cells ml<sup>-1</sup>) than at 0.1% of incident irradiance and  $Z_{\text{aph}}$  (ranging from 0.09 to 1.[4](#page-8-0)4  $\times$  10<sup>[3](#page-9-0)</sup> cells ml<sup>-1</sup>, Fig. 4d; Table 3). In contrast to nanophytoplankton, picophytoplankton average abundance  $(0.81 \pm 0.44 \times 10^3 \text{ cells m}^{-1})$  did not vary significantly across seasons ( $P > 0.05$ ), except at the surface where it decreased in spring  $(P < 0.01)$ . It also appears that the ratios of nanophytoplankton and picophytoplankton abundance at the surface and below the pycnocline (Nano $_{\text{surf}}$ /Nano $_{\text{depth}}$  and Pico $_{\text{surf}}$ /Pico $_{\text{depth}}$ ) both showed negative correlation with  $E_{ph}$  (313 nm), although it was only significant for picophytoplankton ( $r = -0.51$ ,  $P > 0.05$  and  $r = -0.77, P < 0.01$ , for nanophytoplankton and picophytoplankton, respectively). This may suggest that UVBR had a negative impact on phytoplankton and more particularly on picophytoplankton.

In terms of community composition, dominant groups included pico-sized and nano-sized prymnesiophytes and autotrophic flagellates (respectively, averaging  $39 \pm 8.5\%$ and  $42.7 \pm 7.8\%$  of the phytoplankton abundance). Phytoplankton was further composed of nano-sized dinoflagellates such as Gymnodinium spp. and Gyrodinium spp. during winter (11.3  $\pm$  4% of phytoplankton abundance)

<span id="page-8-0"></span>Fig. 4 Contour plots of a bacterial abundance  $(10^5 \text{ cells ml}^{-1})$ ; **b** Chl-a concentration ( $\mu$ g l<sup>-1</sup>); c nanophytoplankton abundance (cells  $ml^{-1}$ ) and d picophytoplankton abundance  $(cells \text{ ml}^{-1})$  for the sampling period. On each graph, the depth of the UML is indicated by a black line



and of nano-sized cryptophytes and pennate diatoms (respectively,  $7.6 \pm 7.9\%$  and  $3.3 \pm 1.5\%$  of phytoplankton abundance) and micro-sized centric diatoms such as Chaetoceros aff. tenuissimus and Corethron hystrix  $(1.4 \pm 1\%$  of phytoplankton abundance) during spring.

The carbon biomass of phytoplankton increased from fall and winter  $(7.2 \pm 2.3 \text{ and } 8.8 \pm 4.1 \text{ µg C } 1^{-1}$ , respectively) to spring  $(25.5 \pm 9.4 \,\mu g \, C \, 1^{-1}$  $(25.5 \pm 9.4 \,\mu g \, C \, 1^{-1}$  $(25.5 \pm 9.4 \,\mu g \, C \, 1^{-1}$ ; Fig. 5b). At the same time, the mean Chl- $a$  at the depth of 50% of incident irradiance decreased from fall (0.21  $\mu$ g l<sup>-1</sup> in average) to winter (0.12  $\mu$ g l<sup>-1</sup> in average) before increasing again in spring (0.41  $\mu$ g l<sup>-1</sup> in average). Although prymnesiophytes and autotrophic flagellates dominated the phytoplankton in terms of abundance, they never contributed significantly to the total autotrophic carbon biomass but on one occasion during spring time (26th of October or day 299) when prymnesiophytes accounted for 49% of the phytoplankton carbon biomass. In agreement with microscopic data, an increase in pigments found in prymnesiophytes, such as Chl  $c_3$  and 19'-hexanoyloxyfucoxanthin, was noticed on day 299 (data from M. Lionard and S. Roy, not shown). During fall and winter, nano-sized dinoflagellates accounted for most of the autotrophic carbon biomass (averaging  $51.3 \pm 20.4\%$ ). The typical dinoflagellate marker peridinin was not detected, but other pigments (Chl  $c_2$ , Chl  $c_3$ , 19'-hexanoyloxyfucoxanthin, fucoxanthin and diadinoxanthin) were present and could indicate the presence of dinoflagellates with acyl-fucoxanthin instead of peridinin. During spring time, the phytoplankton carbon biomass was dominated by micro-sized centric and pennate diatoms  $(29.6 \pm 15.2\%$  and  $11.7 \pm 10.3\%$ , respectively) and by nano-sized dinoflagellates (25.6  $\pm$  15% in average).

<span id="page-9-0"></span>Table 3 Seasonal average (fall, winter and spring) of the abundance of total bacteria (TB,  $10^5$  cells ml<sup>-1</sup>), %HNA, bacterial richness (OTUs), picophytoplankton (Pico, cells  $ml^{-1}$ ), nanophytoplankton (Nano, cells ml<sup>-1</sup>), microphytoplankton (Microphyto, cells ml<sup>-1</sup>) and microzooplankton (Microzoo, cells  $ml^{-1}$ ) at the depth of 100, 50

and 0.1% of incident irradiance and at  $Z_{\text{aph}}$ . Bacterial richness was only determined at 100% and  $Z_{\text{aph}}$ . Microphytoplankton and microzooplankton were only enumerated at the depth of 50% of incident irradiance

Depth	Season	TB	%HNA	Richness	Pico	Nano	Microphyto	Microzoo
100%	Fall	$2.43 \pm 0.22$	$58.2 \pm 1.8$	$52.67 \pm 1.16$	$1,373 \pm 305$	$1,063 \pm 120$	ND.	ND
	Winter	$1.66 \pm 0.33$	$66 \pm 4$	$55.75 \pm 1.5$	$876 \pm 264$	$1,194 \pm 498$	ND.	ND
	Spring	$2.35 \pm 0.30$	$64.4 \pm 4$	$50.11 \pm 6.11$	$557 \pm 113$	$2,456 \pm 716$	ND.	ND.
50\% or max Chl- $a$	Fall	$2.24 \pm 0.17$	$58.6 \pm 0.6$	ND.	$1,354 \pm 311$	$1,071 \pm 102$	$0.45 \pm 0.11$	$50.3 \pm 15.4$
	Winter	$1.64 \pm 0.37$	$65.6 \pm 5.8$	ND.	$905 \pm 298$	$1.316 \pm 654$	$0.73 \pm 0.47$	$51.5 \pm 18.4$
	Spring	$2.29 \pm 0.36$	$65.4 \pm 3.1$	ND	$1,056 \pm 638$	$2.920 \pm 920$	$2.06 \pm 1.69$	$108 \pm 52.6$
$0.1\%$	Fall	$2.14 \pm 0.62$	$70 \pm 7.1$	ND.	$492 \pm 236$	$317 \pm 115$	ND.	ND.
	Winter	$1.69 \pm 0.32$	$69.6 \pm 3.6$	ND.	$464 \pm 427$	$697 \pm 756$	ND.	ND.
	Spring	$1.96 \pm 0.17$	$66.5 \pm 2.8$	ND	$920 \pm 369$	$1,480 \pm 696$	ND.	ND.
$Z_{\rm aph}$	Fall	$1.83 \pm 0.30$	$59 \pm 5.2$	$44.33 \pm 4.16$	$339 \pm 86$	$240 \pm 55$	ND.	ND.
	Winter	$1.71 \pm 0.24$	$70.5 \pm 3.7$	$48 \pm 8.37$	$481 \pm 486$	$754 \pm 924$	ND.	ND.
	Spring	$1.77 \pm 0.19$	$68.1 \pm 3$	$57.78 \pm 5.95$	$617 \pm 239$	$898 \pm 305$	ND.	ND.

ND not determined

## Microzooplankton seasonal distribution and composition

Microzooplankton was composed of small heterotrophic flagellates (diameter range  $3-10 \mu m$ ), choanoflagellates (diameter range  $10-20 \mu m$ ), a  $10-\mu m$ -diameter phagotrophic flagellate, Telonema sp. and ciliates dominated by Strombidium aff. striatum (30 µm diameter) and oligotrichous ciliates (diameter range  $10-50 \mu m$ ). Average microzooplankton abundance increased from fall and winter  $(50.3 \pm 15.4 \text{ and } 51.5 \pm 18.4 \text{ cells } \text{ml}^{-1}$ , respectively) to spring  $(108 \pm 52.6 \text{ cells m}^{-1}, \text{Table 3}).$ 

The abundance of small heterotrophic flagellates averaged 37.9, 38.4 and 71 cells  $ml^{-1}$  for fall, winter and spring, respectively. This represented  $68 \pm 14.4\%$  of all microzooplankton for the whole sampling period, and small heterotrophic flagellates always dominated microzooplankton in terms of abundance. In contrast, they only accounted for a small part of the microzooplankton carbon biomass  $(0.82 \pm 0.49 \text{ µg C1}^{-1} \text{ or } 17.1 \pm 7.4\% \text{ of the}$ microzooplankton biomass). Ciliates' abundance and biomass were always low throughout the sampling period  $(0.5 \pm 0.4 \text{ cells m}^{-1} \text{ and } 0.25 \pm 0.2 \text{ µg C}^{-1}),$  accounting for only  $0.9 \pm 0.9\%$  and  $6.4 \pm 6.3\%$  of the microzooplankton abundance and biomass, respectively. During fall and winter, choanoflagellates composed the remaining of the microzooplankton averaging  $7.1 \pm 4.2$  cells ml<sup>-1</sup> and  $2.5 \pm 1.5 \,\text{µg C} \, \text{I}^{-1}$ . This represented  $13.8 \pm 7.4\%$ and  $63 \pm 25.4\%$  of the microzooplankton abundance and biomass, respectively. A shift in the composition of microzooplankton occurred at spring time when Telonema sp. abundance and biomass increased (35  $\pm$  14 cells ml<sup>-1</sup> and

 $3.8 \pm 1.6$  µg C l<sup>-1</sup>). This represented  $35.3 \pm 15.5\%$  and  $65 \pm 17.3\%$  of the microzooplankton abundance and biomass, respectively.

As shown in Fig. [5a](#page-10-0), microzooplankton biomass at 50% of incident irradiance did not change from fall (average of  $3.6 \pm 0.9$  µg l<sup>-1</sup>) to winter (average of  $3.7 \pm 0.16$  µg l<sup>-1</sup>) but increased in spring (average of  $5.7 \pm 1.3 \,\mathrm{\mu g\ l}^{-1}$ ), although this difference was not significant ( $P > 0.05$ ). In addition, microzooplankton biomass was positively correlated with phytoplankton and bacterioplankton biomass  $(r = 0.73$  and  $r = 0.71$  for phytoplankton and bacterioplankton, respectively,  $P < 0.01$ ).

## Discussion

#### Fall and winter conditions

As expected at this high latitude (Pearce et al. [2007\)](#page-14-0), light was nearly absent in the study area during fall and winter while water temperature was below zero, creating inhospitable conditions for marine organisms. Despite nutrient availability, mean integrated Chl-a was low in fall and winter (averaging 0.3 and 0.13  $\mu$ g l<sup>-1</sup>, respectively) due to the absence of light and very low water temperature (below  $0^{\circ}$ C). During this low production period, bacterioplankton and microzooplankton biomass was also low compared to more productive times of the year. Finally, small cells (i.e. pico-sized and nano-sized) composed most of the autotrophic and heterotrophic biomass, suggesting the predominance of a microbial food web ecosystem (Legendre and Rassoulzadegan [1996\)](#page-14-0).

<span id="page-10-0"></span>Fig. 5 Carbon content at the depth of 50% of incident irradiance of a microzooplankton and bacterioplankton ( $\mu$ g l<sup>-1</sup>) and b picophytoplankton, nanophytoplankton and microphytoplankton ( $\mu$ g l<sup>-1</sup>)



The Southern Ocean, and hence the WAP waters, have been hypothesized to be rectified (one-way) annual  $CO<sub>2</sub>$ sinks, mainly because sea-ice is present during the less productive period of the year, while in the summer, when sea-ice is absent, production is maximum (Takahashi et al. [2002;](#page-15-0) Ducklow et al. [2007](#page-13-0)). Because of regional warming, Stammerjohn et al. ([2008\)](#page-15-0) have observed a decreasing winter sea-ice concentration in the WAP along with a later advance and an earlier retreat over the last decade. Therefore, there is now some doubt regarding the rectified sink hypothesis for WAP waters (Metzl [2009](#page-14-0); Takahashi et al. [2009](#page-15-0)). In fact, sea-ice concentration was particularly low during the 2006 winter compared to the monthly average sea-ice concentrations for the last decade (Moreau et al. unpublished data). Indeed, in 2006, July and August average sea-ice concentrations were, respectively, 8 and 17% below the monthly average sea-ice concentration from

1998 to 2007. If, in the future, sea-ice is absent during the less productive seasons, the WAP waters may no longer be a rectified  $CO<sub>2</sub>$  sink.

Not surprisingly, Wang et al. [\(2009](#page-15-0)) determined that our study site acted as a  $CO<sub>2</sub>$  source during winter. Because water temperature ranges are very small in Antarctic waters,  $CO<sub>2</sub>$  fluxes in Antarctica are mainly driven by biological uptake and respiration (Takahashi et al. [2002](#page-15-0)). Moreover, the presence of a microbial food web leads the ocean to act as a source of  $CO<sub>2</sub>$  to the atmosphere (Schloss et al. [2007](#page-14-0)). We found that a microbial food web dominated at our study site during fall and winter. Therefore, the potential absence of a seaice cover during winter in the WAP may be a concern. Future information on microorganisms' composition and dynamics during winter will be important to determine carbon fluxes in Antarctica.

### Spring: the onset of the production season

The average day of sea-ice retreat in this region for the last decade (1998–2007) was day 318 (Moreau et al., unpublished data). In 2006, sea-ice retreated on day 285 and therefore 33 days before the average retreat of sea-ice for the last decade. The disappearing of sea-ice corresponded to an increase in incident irradiance and water temperature along with high nutrient concentrations, creating suitable growth conditions for phytoplankton and the microbial community. An increase in phytoplankton abundance and biomass was indeed observed in spring, immediately following the retreat of sea-ice. This rise was slight, reaching a maximum Chl-a concentration of 1.13  $\mu$ g l<sup>-1</sup>, representative of what Morán et al.  $(2006)$  $(2006)$  reported as the initial phase of the bloom period in the WAP. Morán et al. ([2006\)](#page-14-0) also associated temporary increase in particulate extracellular release (PER) with changes in the community composition toward the dominance of larger cells. This is in agreement with Wang et al. ([2009\)](#page-15-0) who observed a clear DOC increase in spring during our study, and with our observation that, during this phase, the community composition changed and most of the autotrophic biomass consisted of nano-sized cells. Therefore, the community started to resemble the richer herbivorous food web usually found in WAP waters in the late spring–summer season (e.g. Villafañe et al. [1993](#page-15-0); Moline and Prézelin [1996](#page-14-0); Garibotti et al. [2003](#page-13-0)).

Even though we observed some degree of phytoplankton biomass accumulation in spring following the immediate retreat of sea-ice, this accumulation was not as important as sometimes witnessed in the coastal WAP (e.g. in November 2002; Smith et al. [2008\)](#page-15-0). In fact, from data collected near Palmer station, less than 80 km away from where this study was performed, this initial phase of the bloom lasted until mid to late-December when a significant spring bloom truly started (reaching up to 20.5  $\mu$ g l<sup>-1</sup> on the 26th of December 2006; M. Vernet, personal communication). Several factors may have prevented or delayed the phytoplankton bloom in November 2006. Of these factors, iron is not a limiting factor in Antarctic coastal waters (Martin et al. [1990](#page-14-0)) and probably did not limit phytoplankton biomass accumulation during our study. Compared to other studies (e.g. Holm-Hansen et al. [1994\)](#page-13-0), phosphate, nitrate  $+$ nitrite, and silicate were in excess to support high levels of primary productivity and never became limiting throughout the study either. According to Mitchell and Holm-Hansen ([1991](#page-14-0)), the depth of the UML and of the euphotic zone is critical in determining phytoplankton growth in the WAP waters. During this study, the UML was shallow in spring after sea-ice had retreated (22.6 m deep on average and as shallow as 16 m on day 296, 23rd of October), and  $Z_{eu}$  was deep (average of 72 m). Therefore, oceanographic conditions seemed to be adequate for growth.

Microzooplankton grazing may have delayed or prevented the expected phytoplankton bloom since it can be important in Antarctic waters (Pearce et al. [2008\)](#page-14-0). Indeed, microzooplankton biomass was positively correlated with phytoplankton biomass, and, as suggested by Buck and Garrison ([1988\)](#page-13-0), this may imply that grazing controlled the phytoplankton community. Choanoflagellates largely dominated the microzooplankton biomass during fall and winter. Choanoflagellates are commonly found in Antarctica where they prey on bacteria and picophytoplankton (Buck and Garrison [1988\)](#page-13-0), both of which were numerically dominant during fall and winter. At the beginning of spring, the composition of microzooplankton shifted, and Telonema sp. dominated the microzooplankton biomass. Phagotrophic flagellates such as Telonema sp. are known to prey on larger cells such as nanophytoplankton (Brandt and Sleigh [2000\)](#page-13-0), which dominated the phytoplankton community in spring. Therefore, microzooplankton may have controlled phytoplankton to some extent. Moreover, such a switch in the type and size of microzooplankton from winter to spring has already been observed in Antarctica by Pearce et al. [\(2008](#page-14-0)) who found that microzooplankton was mainly composed of bacterivores during winter and herbivores during more productive periods (i.e. spring and summer).

Despite these apparently optimal growth conditions in spring, we observed only a small increase in bacterial abundance and no significant shift in the bacterial physiological structure (based on the %HNA). The factors that usually control bacterioplankton distribution, and that could have prevented a bacterioplankton bloom during this study, include resource limitation (Morán et al. [2002](#page-14-0)), grazing (Buck and Garrison [1988\)](#page-13-0), viral lysis (Pearce et al. [2007](#page-14-0)) or other deleterious factors such as UVBR penetrating the water column (Buma et al. [2001\)](#page-13-0). Phytoplankton is an important source of nutrients for bacterioplankton (Morán et al.  $2006$ ). In our study, we observed that bacterial abundance in the euphotic zone was positively correlated with phytoplankton abundance and more particularly with nanophytoplankton. Moreover, because of the low bacterial biomass encountered in this study, we believe that phytoplankton production met bacterial carbon demand in the investigated area and that bacterioplankton was not resource limited. In fact, a lag of approximately 1 month between the accumulation of new production in spring and the increase in bacterial abundance has previously been observed in Antarctic waters (Pearce et al. [2007](#page-14-0), [2008\)](#page-14-0). It is possible that the same lag was present here, but sampling stopped too early to detect the subsequent bacterial increase. On the other hand, microzooplankton biomass was positively correlated with bacterial biomass, and, as suggested by Buck and Garrison [\(1988](#page-13-0)), this may imply that grazing had a major influence on the bacterial community.

Spring-time ozone hole and UVBR effects on surface waters

Ozone concentration in August 2006 was 9 DU lower than the average August ozone concentration from 1998 to 2007. The average ozone concentrations in September and October 2006 were very low (202 and 244 DU, respectively) and were consistent with the average ozone concentration observed from 1998 to 2007 (Moreau et al., unpublished data). Finally, 2006 is the year when the ozone hole attained its maximum area coverage [\(http://](http://ozonewatch.gsfc.nasa.gov/) [ozonewatch.gsfc.nasa.gov/\)](http://ozonewatch.gsfc.nasa.gov/). During our survey, sea-ice cover still protected the water column from increased UVBR 4 days after the ozone hole period (13th of October or day 285, Fig. [3](#page-6-0)c). However, ozone recovery was not immediate after day 285, and ''low'' ozone thickness episodes (average of 293.8  $\pm$  51.3 DU) were observed during ozone recovery (mid-October to December, Fig. [3](#page-6-0)b). As a consequence, high-intensity UVBR reached the surface of the water column after the ozone hole period and for the remaining of the sampling period (from the 13th of October to the 10th of November, Fig. [3b](#page-6-0)). Indeed, the intensities of UVBR measured under the water surface during this study are comparable to other studies that reported strong inhibition of primary production (Helbling et al. [1994](#page-13-0)) or mass mortality of bacterioplankton (Helbling et al. [1995\)](#page-13-0).

Bacterial richness was negatively correlated with  $E_{\text{ph}}$ (313 nm) during our study, suggesting negative effects of UVBR. In fact, bacterial richness decreased at the surface in spring, which is in accordance with the findings of Murray et al. ([1998\)](#page-14-0). These authors believed that other factors such as water column stability, depth or season could be responsible for the decrease in bacterial richness observed in spring. Considering our data, it seems that UVBR might also influence bacterial richness in surface waters in spring, possibly by selecting for UVBR-resistant species. However, this result needs further studies to be confirmed. Moreover, the ratios  $Nano<sub>surf</sub>/Nano<sub>depth</sub>$  and Pico<sub>surf</sub>/Pico<sub>depth</sub> were both inversely correlated with  $E_{\text{ph}}$ (313 nm) (although it was only significant for picophytoplankton). This suggests that UVBR had negative effects on phytoplankton in surface waters (for a review of potential effects see Häder and Sinha [2005](#page-13-0)).

UVBR are deleterious for marine organisms, and photorepair processes are more efficient in deeper waters where low values of the UVBR:UVAR ratio are encountered (Kaiser and Herndl [1997](#page-14-0)). Moreover, the degree and extent of exposure to UVBR are important in determining the impact these radiations can have (Helbling et al. [1996](#page-13-0); Davidson and Belbin [2002\)](#page-13-0). Therefore, water column stratification will be a major factor determining how these radiations affect the microbial community (Xenopoulos et al. [2000;](#page-15-0) Hernando and Ferreyra [2005\)](#page-13-0). When increased UVBR reached the water column of the WAP in spring, the upper mixed layer was as shallow as 16 m and the photoactive depth for UVBR was relatively deep (10–12 m). Therefore, the upper mixed layer was highly exposed to UVBR. Under these conditions, microorganisms were confined within a shallow UML and were probably undergoing significant stress from UVBR (Häder and Sinha [2005](#page-13-0); Hernando et al. [2005](#page-13-0)). DNA photorepair was probably less efficient, and microorganisms would have accumulated DNA damage (i.e. cyclobutane pyrimidine dimers or CPDs; one of the major DNA lesion caused by UVR) (Häder and Sinha  $2005$ ). Moreover, as described by Helbling et al. ([2005\)](#page-13-0), photoinhibition was probably strong in surface waters. Therefore, we hypothesize that UVBR had an impact on surface water microorganisms early in spring, when no ice cover could shield the water column.

## Conclusions

In conclusion, the abundance and biomass of the microbial community (bacterioplankton, phytoplankton and microzooplankton) were low in fall and winter 2006 at our study site, and the microbial community was dominated by small cells (pico-sized). Following the retreat of sea-ice, an increase in bacterial and phytoplankton abundance and biomass was observed in spring. However, despite seemingly optimal growth conditions, this increase was small. Microzooplankton grazing may have prevented or delayed the typical WAP bloom. In addition, the early disappearing of sea-ice in spring 2006 let the water column exposed to increased UVBR, which seemed to have had an impact on bacterial richness and on phytoplankton abundance and composition in surface waters. This study suggests that the dynamics of the microbial community in the WAP at this time of the year can be controlled by the interaction of several biological and physical factors. In the context of global warming, decreasing sea-ice season and delayed stratospheric ozone recovery for the next decades, more studies of the WAP microbial food web functioning during winter and early spring should be encouraged.

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