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The Spin ± 1 Teukolsky Equations and the Maxwell System on Schwarzschild

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Abstract. In this note, we prove decay for the spin ± 1 Teukolsky equations on the Schwarzschild spacetime. These equations are those satisfied by the extreme components (α and α) of the Maxwell field, when expressed with respect to a null frame. The subject has already been addressed in the literature, and the interest in the present approach lies in the connection with the recent work by Dafermos, Holzegel and Rodnianski on linearized gravity (Dafermos et al. in The linear stability of the Schwarzschild solution to gravitational perturbations, 2016. arXiv:1601.06467). In analogy with the spin ± 2 case, it seems difficult to directly prove Morawetz estimates for solutions to the spin ± 1 Teukolsky equations. By performing a differential transformation on the extreme components α and $\underline{\alpha}$, we obtain quantities which satisfy a Fackerell-Ipser Equation, which does admit a straightforward Morawetz estimate and is the key to the decay estimates. This approach is exactly analogous to the strategy appearing in the aforementioned work on linearized gravity. We achieve inverse polynomial decay estimates by a streamlined version of the physical space r^p method of Dafermos and Rodnianski. Furthermore, we are also able to prove decay for all the components of the Maxwell system. The transformation that we use is a physical space version of a fixed-frequency transformation which appeared in the work of Chandrasekhar (Proc R Soc Lond Ser A 348(1652):39-55, 1976). The present note is a version of the author's master thesis and also serves the "pedagogical" purpose to be as complete as possible in the presentation.

1. Introduction

The subject of black-hole stability has received a good amount of attention lately, as research efforts are focused on proving the full nonlinear stability of the Kerr family of black holes. See the lecture notes [10] for a comprehensive introduction on the topic. The interest in the aforementioned problem stems from the fundamental question of whether the Kerr solution indeed provides an appropriate description of physical reality.

In an attempt to address the fully nonlinear problem, researchers have been following a natural path: first, one studies the covariant scalar wave equation (spin 0). Then, one studies the Maxwell equations (in which the extreme components satisfy spin ± 1 Teukolsky equations). Finally, one seeks to study the linearized Einstein equations (in which the extreme curvature components satisfy spin ± 2 Teukolsky equations). Eventually, this process leads to a deeper understanding of the linear structure, which is a crucial step needed to address the nonlinear stability of the Kerr family.

Hence, the subject of decay of linear waves on a black-hole background has been recently studied, with many contributions by different research groups. For the first step of the "linear program," these efforts culminated in the proof of decay of scalar waves on a Kerr background for |a| < M, by Dafermos et al. [12].

To proceed in the outlined program, the Maxwell equations have also been studied. Boundedness and decay for solutions to the Maxwell equations have been first proved by Blue [4] on the Schwarzschild background. There were advances in extending these results to the Kerr setting in the slowly rotating case ($a \ll M$) by Andersson and Blue [2]. Furthermore, Sterbenz and Tataru proved local energy decay for the Maxwell field on a large class of spherically symmetric spacetimes in [17]. In addition, Metcalfe, Tataru and Tohaneanu proved pointwise decay for the Maxwell field under the assumption of local energy decay, for a fairly general class of asymptotically flat spacetimes, in the paper [14]. See also [13].

Finally, Andersson, Bäckdahl and Blue found a new way of producing robust energy estimates on the Schwarzschild background, exploiting a superenergy tensor. The relevant paper is [1].

Recently, there has also been a substantial advance in the last step of this "linear program," i.e., the proof of linear stability of the Schwarzschild metric under gravitational perturbations. Indeed, a recent result of Dafermos et al. [11] shows that the Schwarzschild metric is stable under linearized gravitational perturbations.

1.1. Maxwell: Why Another Proof?

In this paper, we return to the topic of decay of the Maxwell field on a curved background. The aim of the note is threefold: first, we provide a simple proof of pointwise decay of the Maxwell field on the Schwarzschild background. Second, we adopt the further "didactic" aim to be as detailed as possible in our exposition. The third and main motivation for this work, though, lies in the connection with the aforementioned problem of linear stability of the Schwarzschild metric. It has become evident that a very similar approach to the one in [11] can be adopted to address the decay properties of the spin ± 1 Teukolsky equations and of the Maxwell system on Schwarzschild.

We briefly recall the strategy followed by the authors in [11]. Given a solution to the equation for the extreme curvature components (i.e., the spin ± 2 Teukolsky equation), the authors find a second-order differential transformation that performs the following: if we apply the transformation on the

extreme curvature component, the resulting expression satisfies a "good" equation (i.e., an equation for which Morawetz and energy estimates can be proved). The resulting equation is called the Regge–Wheeler equation. The relevant quantity enables us to estimate all the components of the field, just by using transport equations.

We wish to follow the same path in the context of the Maxwell system, which we study on the Schwarzschild background. We start from the spin ± 1 Teukolsky equations, which are satisfied by the extreme components α and $\underline{\alpha}$. We exploit an elementary transformation (which can be found, in its fixedfrequency form, in the work of Chandrasekhar [6]). The transformation takes the spin ± 1 Teukolsky equations into a "good" equation, called the Fackerell– Ipser equation. The transformed quantity has the required property of vanishing on the zeroth mode and of satisfying good integrated decay estimates. We finally use the transformed quantity to estimate, via transport equations, first the extreme components α and $\underline{\alpha}$, and subsequently the remaining components of the Maxwell field.

The approach outlined above differs from previous work ([1,4]) in two main aspects. The first main difference is the analysis of the "good" quantity arising from the Chandrasekhar transformation. Secondly, the way we perform estimates (specifically, Morawetz estimates and decay estimates) is closer in spirit to the works [9,11] and uses the "new physical space method" by Dafermos and Rodnianski. We will touch upon these points in Sect. D in "Appendix," where we sketch a comparison with the aforementioned works [1,4].

Let me finally note that earlier versions of the results contained in this note were originally obtained in my master thesis at ETH Zürich [15].

1.2. Outline of the Note

We will first motivate our analysis in Sect. 2, giving an outline of the work [11], as well as a sketch of the argument in the present note.

We subsequently introduce some necessary notation and the null decomposition of the Maxwell system in Sect. 3, as well as the crucial transformation which takes the spin ± 1 Teukolsky equations into the Fackerell–Ipser equation.

We then proceed to state the main results of this note in Sect. 4:

- decay for solutions to the spin ± 1 Teukolsky equations (Theorem 4.1),
- decay for solutions to the Maxwell system (Theorem 4.2).

Subsequently, we prove integrated decay estimates for solutions to the Fackerell–Ipser equation, in Sect. 5. We follow the r^p -method approach by Dafermos–Rodnianski, as in [9].

From these integrated estimates, via a combination of Sobolev embedding and the Grönwall inequality, we obtain decay for solutions to the spin ± 1 Teukolsky equations in Sect. 6.

We improve the decay for the α component in Sect. 7.

We then extend the decay to the full Maxwell system in Sect. 8.

We conclude the note with the appendices, in which we collect important lemmas and calculations. In "Appendix D," we compare our result with those achieved in previous work by Blue [4] and Andersson, Bäckdahl, Blue [1].

2. Motivation and Main Idea of the Work

2.1. Introducing the Work on Linearized Gravity

It is sensible to recall here the strategy the authors follow in the paper [11]. We warn the reader that this brief subsection does not have any claim of completeness, and we refer to the paper [11] for the full details.

In said work, the authors consider the vacuum Einstein equations:

$$\mathbf{Ric}_{\mu\nu} = 0, \tag{2.1}$$

with respect to the unknown metric g, which is considered to be "near" the Schwarzschild metric. In the following, bold typeface will always denote quantities associated with the metric g, whereas quantities in non-bolded typeface will always be associated with the Schwarzschild metric g.

The authors perform a suitable linearization as follows. Let $(\hat{L}, \hat{L}, e_1, e_2)$ be a normalized null frame with respect to the metric g. Also, \star denotes the Hodge dual with respect to two indices:

$${}^{\star} \boldsymbol{R}_{\mu\nu\kappa\lambda} := \frac{1}{2} \boldsymbol{\varepsilon}_{\mu\nu\alpha\beta} \boldsymbol{R}_{\kappa\lambda}^{\ \alpha\beta}.$$
(2.2)

Note that, here, \star is in bold typeface, hence the Hodge dual is calculated using the (bold) volume form ε , which is the natural volume form induced by the metric g.

In order to write the linearized Einstein equations, the authors decompose the field into null frame. They consider the Ricci coefficients:

$$\begin{split} \chi_{AB} &:= g(\nabla_A L, e_B), \qquad \underline{\chi}_{AB} := g(\nabla_A \underline{L}, e_B), \\ \eta_A &:= -\frac{1}{2}g(\nabla_{\underline{\hat{L}}} e_A, \hat{L}), \qquad \underline{\eta}_A := -\frac{1}{2}g(\nabla_{\underline{\hat{L}}} e_A, \underline{\hat{L}}) \\ \hat{\omega} &:= \frac{1}{2}g(\nabla_{\underline{\hat{L}}} \underline{\hat{L}}, \hat{L}), \qquad \underline{\hat{\omega}} := \frac{1}{2}g(\nabla_{\underline{\hat{L}}} \hat{L}, \underline{\hat{L}}), \\ \zeta &:= \frac{1}{2}g(\nabla_A \hat{L}, \underline{\hat{L}}). \end{split}$$
(2.3)

The authors consider the null components of the Riemann tensor as well, i.e., the set of components

$$\begin{aligned} \boldsymbol{\alpha}_{AB} &:= \boldsymbol{R}(\boldsymbol{e}_{A}, \hat{\boldsymbol{L}}, \boldsymbol{e}_{B}, \hat{\boldsymbol{L}}), & \underline{\boldsymbol{\alpha}}_{AB} &:= \boldsymbol{R}(\boldsymbol{e}_{A}, \underline{\hat{\boldsymbol{L}}}, \boldsymbol{e}_{B}, \underline{\hat{\boldsymbol{L}}}), \\ \boldsymbol{\beta}_{A} &:= \frac{1}{2} \boldsymbol{R}(\boldsymbol{e}_{A}, \hat{\boldsymbol{L}}, \underline{\hat{\boldsymbol{L}}}, \hat{\boldsymbol{L}}), & \underline{\boldsymbol{\beta}}_{A} &:= \boldsymbol{R}(\boldsymbol{e}_{A}, \underline{\hat{\boldsymbol{L}}}, \underline{\hat{\boldsymbol{L}}}, \hat{\boldsymbol{L}}), \\ \boldsymbol{\rho} &:= \frac{1}{4} \boldsymbol{R}(\hat{\boldsymbol{L}}, \underline{\hat{\boldsymbol{L}}}, \hat{\boldsymbol{L}}, \underline{\hat{\boldsymbol{L}}}), & \boldsymbol{\sigma} &:= \frac{1}{4} {}^{*} \boldsymbol{R}(\hat{\boldsymbol{L}}, \underline{\hat{\boldsymbol{L}}}, \hat{\boldsymbol{L}}, \underline{\hat{\boldsymbol{L}}}). \end{aligned}$$
(2.4)

They proceed to linearize the Einstein vacuum equations (2.1) around the Schwarzschild metric using this null framework. In other words, they write the unknown metric

$$\boldsymbol{g} = g + \overset{(1)}{g},$$

where g is the Schwarzschild metric and $\overset{(1)}{g}$ is the variation of the metric. They plug this expression for g into Eq. (2.1), decompose into null components and

eliminate the nonlinear terms in $\overset{(1)}{g}$. They therefore obtain a suitable linearization of the Einstein equations around Schwarzschild.

The resulting equations are *coupled*, meaning that in every equation we have more than one component of the field. In the following, we adopt the notation from [11]. In particular, quantities with the superscript ⁽¹⁾ correspond to the variation of the metric, whereas quantities without subscript or superscript correspond to their original Schwarzschild values.

Remarkably, the perturbed extreme components $\stackrel{(1)}{\alpha}$ and $\stackrel{(1)}{\alpha}$ were shown by Teukolsky in [18] to satisfy *decoupled* equations. In the notation of [11], the equation for $\stackrel{(1)}{\alpha}$ is

On functions, we have the definition $\Omega \nabla_3 = \partial_u$, $\Omega \nabla_4 = \partial_v$. Here, $\Omega = \sqrt{1-2M/r}$. ∇ indicates the induced connection on the spheres of constant (u, v) Schwarzschild coordinates, and Δ indicates the corresponding covariant Laplacian. Also, the definition of non-bolded Ricci coefficients and non-bolded null components is exactly as in (2.3) and (2.4), replacing all the boldface quantities by the non-bolded ones. Furthermore, $e_3 = \underline{\hat{L}} = \Omega^{-1}\partial_u$, $e_4 = \hat{L} = \Omega^{-1}\partial_v$.

The core of the proof is the following: starting from Eq. (2.5), the authors find a quantity $\stackrel{(1)}{\Psi}$ which satisfies a "good" equation. Such quantity $\stackrel{(1)}{\Psi}$ (and analogously its companion $\stackrel{(1)}{\Psi}$) can be defined in terms of the sole extreme component $\stackrel{(1)}{\alpha}$ (resp. $\stackrel{(1)}{\alpha}$), which satisfies a spin ± 2 Teukolsky equation. Here are some definitions

As usual, the superscript ⁽¹⁾ indicates that the quantity is the one relative to the perturbed metric. As before, on functions, $\Omega \nabla_3 = \partial_u$, $\Omega \nabla_4 = \partial_v$. We finally define the rescaled versions of $\stackrel{(1)}{P}$ and $\stackrel{(1)}{\underline{P}}$:

$$\stackrel{\scriptscriptstyle(1)}{\Psi}:=r^5\stackrel{\scriptscriptstyle(1)}{P},\quad \stackrel{\scriptscriptstyle(1)}{\underline{\Psi}}:=r^5\stackrel{\scriptscriptstyle(1)}{\underline{P}}.$$

In this setting, $\stackrel{(1)}{\Psi}$ satisfies the **Regge–Wheeler equation**, where we define $\mu = 2M/r$:

$$\Omega \nabla_{3} (\Omega \nabla_{4} \overset{(1)}{\Psi}) - (1-\mu) \Delta \overset{(1)}{\Psi} + \left(\frac{4}{r^{2}} - \frac{6M}{r^{3}}\right) (1-\mu) \overset{(1)}{\Psi} = 0.$$
(2.6)

Here, as before, $\Omega = \sqrt{1 - 2M/r}$. Also, on functions, $\Omega \nabla_3 = \partial_u$, $\Omega \nabla_4 = \partial_v$. The same equation is satisfied by $\underline{\Psi}^{(1)}$. *Remark* 2.1. Most importantly, solutions to Eq. (2.6) satisfy an energy conservation inequality and a Morawetz estimate.

Remark 2.2. Equation (2.6) is a *tensorial* equation. This equation is typically stated by other authors in the corresponding scalar form. See Remark 7.1 in [11] for the form of the corresponding scalar equation.

Remark 2.3. We further notice that $\stackrel{(1)}{\Psi}$ can also be defined solely in terms of the middle components $\stackrel{(1)}{\sigma}$ and $\stackrel{(1)}{\rho}$ via angular derivation:

$$\Psi_{AB}^{(1)} := r^5 \left(\mathcal{D}_2^{\star} \mathcal{D}_1^{\star} (-\stackrel{(1)}{\rho}, \stackrel{(1)}{\sigma}) + \frac{3}{4} \rho \operatorname{tr} \chi (\stackrel{(1)}{\hat{\chi}} - \stackrel{(1)}{\hat{\chi}}) \right).$$
(2.7)

Here, the definition of the spherical operators \mathcal{P}_1^{\star} and \mathcal{P}_2^{\star} is as follows:

Remark 2.4. An additional difficulty, in the linearized gravity case, is the existence of *pure gauge solutions*. These are solutions of the linearized Einstein equations arising from the diffeomorphism invariance of the Einstein equations. For the full formulation, we refer to the paper [11], especially Sections 2.1.4 and 6.1. The Maxwell case considered here does not suffer from such difficulties, as there is a predetermined gauge (given by our choice of L and \underline{L}) in which the field satisfies good decay properties.

Remark 2.5. There is a connection between the kernel of $\mathcal{P}_2^* \mathcal{P}_1^*$ and solutions of the linearized gravity equations corresponding to "infinitesimal perturbations toward Kerr" of the Schwarzschild solution. As we shall see, this has an analogy in the Maxwell case.

2.2. Non-radiating Modes

Let us now turn our attention to the Maxwell system on the Schwarzschild spacetime. We remark that the Maxwell equations possess non-trivial stationary solutions, whose null components decay at spacelike infinity.

It is an easy computation to show that the following expression gives stationary solutions to the Maxwell system on Schwarzschild:

Here, q_E and q_B are two real parameters, respectively, the "electric charge" and the "magnetic charge."

Excluding such "stationary modes" is a crucial element of every proof of decay of the Maxwell field on the Schwarzschild manifold.

2.3. Key to the Proof for the Spin $\pm\,1$ Teukolsky Equations and the Maxwell System

We consider the Maxwell equations for the unknown two-form F on a fixed Schwarzschild metric as a background.

Let us first anticipate some useful notation from Sect. 3. Let us consider the Regge–Wheeler coordinates $(t, r_*, \theta, \varphi)$, with $r_* := r + 2M \log(r - 2M) - 3M - 2M \log M$. We define $L := \partial_t + \partial_{r_*}$, and $\underline{L} := \partial_t - \partial_{r_*}$. We then let

$$\alpha(V) := F(L, V), \quad \underline{\alpha}(V) := F(\underline{L}, V),$$

for any vector field V tangent to the spheres $S_{t,r}$ of constant t and r coordinates. We subsequently regard α and $\underline{\alpha}$ as one-forms on the vector bundle \mathcal{B} composed of vectors everywhere tangent to the spheres $S_{t,r}$. Since we denote (θ^A, θ^B) a generic local coordinate system for the sphere \mathbb{S}^2 , we will write α_A and $\underline{\alpha}_A$ (using capitalized roman indices) to denote the one-forms obtained by the above procedure. For the complete definitions, we refer the reader to Sects. 3 and 3.2.

As was proved by Bardeen and Press in [3] (in its scalar, fixed-frequency version), the extreme components (α and $\underline{\alpha}$) satisfy the so-called **spin** $\pm \mathbf{1}$ **Teukolsky equations**:

$$\nabla \underline{V}_{\underline{L}} \nabla L(r\alpha_A) + \frac{2}{r} \left(1 - \frac{3M}{r} \right) \nabla \underline{L}(r\alpha_A) - (1 - \mu) \Delta (r\alpha_A) + \frac{1 - \mu}{r^2} r\alpha_A = 0,$$
(2.10)

$$\nabla \underline{V}_{\underline{L}} \nabla L(r\underline{\alpha}_A) - \frac{2}{r} \left(1 - \frac{3M}{r} \right) \nabla L(r\underline{\alpha}_A) - (1 - \mu) \Delta (r\underline{\alpha}_A) + \frac{1 - \mu}{r^2} r\underline{\alpha}_A = 0.$$
(2.11)

For the ∇ notation, and for the definition of Δ , we refer to Sect. 3. Here, as usual, $\mu = 2M/r$.

Remark 2.6. Notice that these are tensorial equations, cf. Remark 2.2.

Remark 2.7. The stationary solutions considered in Sect. 2.2 have vanishing α and $\underline{\alpha}$ components. In other words, the extreme components α and $\underline{\alpha}$ do not "see" the stationary modes. Hence, we seek to define a quantity starting from α and $\underline{\alpha}$.

These equations are poorly behaved from the point of view of energy estimates, due to the presence of bad first-order terms. We now consider the quantities

$$\phi_A := \frac{r^2}{1-\mu} \nabla \underline{\nabla}_{\underline{L}}(r\alpha_A),$$

$$\underline{\phi}_A := \frac{r^2}{1-\mu} \nabla \nabla_{\underline{L}}(r\underline{\alpha}_A).$$
(2.12)

Remark 2.8. A fixed-frequency version of this transformation appeared in its scalar form in the work by Chandrasekhar [6], in the case of linear electromagnetism on a fixed Kerr background. Similarly, a fixed-frequency version of the transformation relating Eqs. (2.5) and (2.6) appeared in the work of Chandrasekhar [5] on the linearized gravitational perturbations of a Schwarzschild black hole.

Remark 2.9. Notice that, in view of the Maxwell equations, in the notation of the previous subsection,

$$\phi_A = r^3 \mathcal{D}_1^\star(\rho, \sigma).$$

It can be shown by direct calculation from (2.10) and (2.11) that ϕ_A and ϕ_A both satisfy the following tensorial **Fackerell–Ipser equation**:

$$\nabla_{\underline{L}} \nabla_{\underline{L}} \phi - (1-\mu) \Delta \phi + V \phi = 0, \qquad (2.13)$$

with $V = \frac{1-\mu}{r^2}$.

Remark 2.10. Notice that this equation is analogous to (2.6). As in the case of linearized gravity, this equation is also usually stated in the literature as a scalar equation. The scalar form of the Fackerell–Ipser equation for the unknown u is the following:

$$(1-\mu)^{-1}L\underline{L}u - \not\Delta u = 0.$$
(2.14)

In particular, this is the wave equation satisfied by $r^2 \rho$ and by $r^2 \sigma$. Commuting the equation with the angular operator $r \nabla$ leads to the appearance of the additional zeroth order term. This is closely related to the case of linearized gravity, see Remark 7.1 in [11], and is consistent with Remark 2.9.

Equation (2.13) is now the key to the argument: it admits robust energy estimates, and furthermore the quantity ϕ enables us to estimate the quantities α and $\underline{\alpha}$.

Remark 2.11. Let us notice here that the argument contained in this note (in particular, the proof of Theorem 4.1) does not suffer from the difficulty arising from the "pure gauge solutions," which are present in the linearized gravity case [11].

3. Preliminaries and Notation

Having settled the heuristics, we proceed to the actual setup.

- Let $g_{\mathbb{S}^2}$ be the standard metric on the sphere \mathbb{S}^2 .
- Let S_e be the following smooth Lorentzian manifold without boundary: $S_e := (t, r, \omega) \in \mathbb{R} \times (2M, \infty) \times \mathbb{S}^2$. The metric tensor g_e on S_e is defined as follows:

$$g_e := -(1-\mu)\mathrm{d}t \otimes \mathrm{d}t + (1-\mu)^{-1}\mathrm{d}r \otimes \mathrm{d}r + r^2 g_{\mathbb{S}^2}.$$

Here, $\mu := \frac{2M}{r}$. We call S_e the **open exterior Schwarzschild spacetime**, or simply **Schwarzschild exterior**. Note that this set does not have a boundary.

• When denoting subsets of S_e determined by some property, we shorten the notation in the following way:

{property} := { $(u, v, \omega) \in S_e$: property}.

Hence, for instance, the set $\{(u, v, \omega) \in S_e : u \ge u_0\}$ is denoted by $\{u \ge u_0\}$.

- Let ∇ the Levi-Civita connection associated with g_e .
- Other coordinates:
 - (t, r_*, ω) , with $r_* := r + 2M \log(r 2M) 3M 2M \log M$,
 - (t^*, r_1, ω) with $t^* := t + 2M \log(r 2M), r_1 := r$,
 - (u, v, ω) with $u := t r_*, v := t + r_*.$
- We will adopt the following convention throughout the paper. If y > 2M is a real number, which we think of as a radius r, we will define y_* to be the corresponding r_* coordinate, that is

$$y_* := y + 2M \log(y - 2M) - 3M - 2M \log M.$$

This will allow us to make sense of expression such as, for instance, $(y_1 + y_2)_*$, if y_1, y_2 are reals, both greater than 2M.

• Define a local frame field: $(L, \underline{L}, \partial_{\theta^A}, \partial_{\theta^B})$ such that

$$L := \partial_t + \partial_{r_*},$$

$$\underline{L} := \partial_t - \partial_{r_*},$$

and such that ∂_{θ^A} and ∂_{θ^B} are local vector fields induced by a system of local coordinates (θ^A, θ^B) for \mathbb{S}^2 .

- S_e embeds isometrically in \mathcal{M} , the maximally extended Schwarzschild spacetime. Let us call such isometric embedding $i: S_e \to \mathcal{M}$. For details about the precise definition of \mathcal{M} and the form of i, see [10], Section 2.3.
- Consider the Kruskal coordinates (T, R, θ, φ) on \mathcal{M} , as in [10], Section 2.3. We denote by \mathfrak{V} the set of vector fields on $T\mathcal{M}$:

$$\mathfrak{V} := \left\{ \frac{\partial}{\partial T}, \frac{\partial}{\partial R}, \Omega_1, \Omega_2, \Omega_3 \right\},\,$$

where the $\{\Omega_i\}$ are rotation Killing fields, such that $\operatorname{span}(\mathfrak{V}) = T\mathcal{M}$.

- We will adopt the following convention throughout the paper: a function f is smooth on an open set $U \subset S_e$ (denoted by $f \in C^{\infty}(U)$) if there exists an open set $O \subset \mathcal{M}$, such that $O \supset \overline{i(U)}$ and there exists a smooth function $\tilde{f} \in C^{\infty}(O)$ (seen as a subset of \mathcal{M}) which restricts to f on i(U).
- Let again be $U \subset S_e$ an open set, and let \mathcal{V} be either TU or a derived bundle of it (i.e., a tensor product of some copies of TU with some copies of its dual). We say that a section V of \mathcal{V} is smooth if the following holds. We push V forward via i to obtain a section V' of \mathcal{V}' , a derived bundle of $T\mathcal{M}$. We then express the components of V' in the frame \mathfrak{V} (or the corresponding derived frame), obtaining a collection of functions $(f_i)_{i=1,...,5}: i(U) \to \mathbb{R}, n \in \mathbb{N}$. For V to be smooth, we require that the f_i 's all be extendible to smooth functions $\overline{f_i}$ on an open set $O \supset \overline{i(U)}$, as in the previous bullet. We denote by $\Gamma(\mathcal{V})$ the vectorspace of all smooth sections of such bundle.

Remark 3.1. Notice that this definition encodes the notion of "smoothness up to the event horizon," for instance, if $U = S_e$, or if $U = \{t^* > a\}$, with a > 0.

Remark 3.2. \underline{L} vanishes as $r \to 2M$, $\left(1 - \frac{2M}{r}\right)^{-1} \underline{L}$ is a smooth vector field on the set $U := \{t^* > a\}$, with $a \in \mathbb{R}$, according to our definition (it is a smooth section of TU).

• Let $k \in \mathbb{N}$. We denote by $\Lambda^k(\mathcal{V})$ the vectorspace of smooth antisymmetric k-forms, which is the space of smooth sections of the bundle $\mathcal{V}^* \otimes \cdots \otimes \mathcal{V}^*$, which are antisymmetric with respect to the permutation

k times of any two arguments.

- We introduce tensor fields tangent to the spheres of constant r.
 - Let $S_{\tilde{t},\tilde{r}} \subset \mathcal{S}_e$ be the set

$$S_{\tilde{t},\tilde{r}} := \{(t,r,\omega) \in \mathcal{S}_e, t = \tilde{t}, r = \tilde{r}\}.$$

- Consider $TS_{\tilde{t},\tilde{r}} \subset TS_e$, and let

$$\mathcal{B} := \bigcup_{\tilde{t} \in (-\infty,\infty), \tilde{r} \in (2M,\infty)} TS_{\tilde{t},\tilde{r}} \subset TS_e.$$

Notice that \mathcal{B} is the bundle tangent to each sphere of constant t, r.

- Sections of \mathcal{B} can be seen as sections of $T\mathcal{S}_e$, due to the fact that $\mathcal{B} \subset T\mathcal{S}_e$. Hence, we say that a section W of \mathcal{B} (or a derived bundle thereof) is smooth if the corresponding section of $T\mathcal{S}_e$ (or a derived bundle thereof) is.
- If $\widetilde{\mathcal{B}}$ is \mathcal{B} or a derived bundle thereof, we denote by $\Gamma(\widetilde{\mathcal{B}})$ the vector space of all smooth sections of $\widetilde{\mathcal{B}}$. Similarly, $\Lambda^k(\widetilde{\mathcal{B}})$ is the space of alternating k-forms on $\widetilde{\mathcal{B}}$.
- Indices for tensors in TS_e and derived bundles will be indicated by Greek letters $\mu, \nu, \kappa \dots$ Indices for tensors in \mathcal{B} will be indicated by uppercase Latin letters: A, B, C, \dots
- Let $\not g$ be the induced metric on spheres of constant r. Technically, this is a smooth section $\not g \in \Gamma(\mathcal{B}^* \otimes \mathcal{B}^*)$. On each sphere $S_{t,r}, \not g$ is the round metric.
- Let $\not \in_{AB} \in \Lambda^2(\mathcal{B})$ be the induced volume form on the spheres $S_{\tilde{t},\tilde{r}}$.
- Let $(\cdot)^{\perp}$: $T\mathcal{S}_e \to \mathcal{B}$ be the orthogonal projection on the spheres $S_{\tilde{t},\tilde{r}}$.
- Let $V, W \in \Gamma(\mathcal{B})$. We define a connection on \mathcal{B} by

$$\nabla W := (\nabla_V W)^{\perp}. \tag{3.1}$$

This connection coincides, on the spheres, with the Levi-Civita connection induced by the induced metric g.

– We define two other differential operators on $\Gamma(\mathcal{B})$ in the following way:

$$\label{eq:V_LV} \ensuremath{\overline{\nabla}}_L V := (\nabla_L V)^\perp, \quad \ensuremath{\overline{\nabla}}_{\underline{L}} V := (\nabla_{\underline{L}} V)^\perp.$$

– The previous differential operators can be extended to derived bundles from \mathcal{B} in the usual way, asking that they satisfy the Leibnitz rule.

- We define the induced covariant curl and divergence in the following way. Let $\omega \in \Gamma(\mathcal{B}^*)$:

- We introduce the foliation needed in the note, and relevant Sobolev norms on it.
 - Let $\Omega := \{\Omega_1, \Omega_2, \Omega_3\}$ be a set of angular Killing fields of S_e whose elements, at each point of S_e , span all directions in \mathcal{B} . Let $\tilde{\Omega}$ be the renormalized version

$$\tilde{\Omega} := \{\Omega_1/r, \Omega_2/r, \Omega_3/r\}.$$

- Let $k \geq 0$, let $\iota_k^{\tilde{\Omega}}$ (resp. $\iota_{\leq k}^{\tilde{\Omega}}$) be the set of all ordered lists of length k (resp. $\leq k$) composed of elements of $\tilde{\Omega}$, and analogously let ι_k^{Ω} (resp. $\iota_{\leq k}^{\tilde{\Omega}}$) be the set of all ordered lists of length k (resp. $\leq k$) composed of elements of Ω . Elements in $\iota_k^{\tilde{\Omega}}$ and ι_k^{Ω} will be referred to as **multi-indices**.
- Let η be a covariant tensor field on \mathcal{B} . If $J = (V_1, \ldots, V_k)$ is a multiindex, and let $\not L$ be the Lie derivative induced by the connection ∇ . Let $X \in \Gamma(\mathcal{B})$. We let

$$\nabla^{J}\eta := \nabla_{V_{1}} \cdots \nabla_{V_{k}}\eta, \quad \pounds^{J}\eta := \pounds_{V_{1}} \cdots \pounds_{V_{k}}\eta, \quad (\nabla_{X})^{k}\eta := \underbrace{\nabla_{X} \cdots \nabla_{X}}_{k-\text{times}}\eta.$$
(3.2)

– Let $n \ge 0$ be an integer, $\eta \in \Gamma((\mathcal{B}^*)^n)$. We define the angular norm of η as

$$|\eta|(t,r,\omega) := \sum_{J \in \iota_n^{\bar{\Omega}}} |\eta(J)|,$$

where we assumed, if $J = (V_1, \ldots, V_n)$,

$$\eta(J) = \eta(V_1, \ldots, V_n).$$

– Given $\tilde{u}, \tilde{v} \in \mathbb{R}$, we define

$$\overline{C}_{\tilde{u},\tilde{v}} := \{(u,v,\omega) \in \mathcal{S}_e, u = \tilde{u}, v \ge \tilde{v}\}, \\
\underline{C}_{\tilde{u},\tilde{v}} := \{(u,v,\omega) \in \mathcal{S}_e, u \ge \tilde{u}, v = \tilde{v}\}, \\
C_{\tilde{u},\tilde{v}} := \overline{C}_{\tilde{u},\tilde{v}} \cup \underline{C}_{\tilde{u},\tilde{v}}, \\
\overline{C}_{\tilde{u}} := \{(u,v,\omega) \in \mathcal{S}_e, u = \tilde{u}\}, \\
\underline{C}_{\tilde{v}} := \{(u,v,\omega) \in \mathcal{S}_e, v = \tilde{v}\}.$$

- Let *i* be the inclusion of C_{u_0,v_0} into S_e . Let η be a covariant section of $i^*\mathcal{B}$ or one of its derived bundles. We define the following fluxes

$$\mathcal{F}_{u}^{T}[\eta](v_{1}, v_{2})$$

$$:= \int_{v_{1}}^{v_{2}} \int_{\mathbb{S}^{2}} [|\nabla_{L}\eta|^{2} + (1-\mu)|\nabla\eta|^{2} + V|\eta|^{2}](u, v, \omega) \mathrm{d}v \mathrm{d}\mathbb{S}^{2}(\omega),$$
(3.3)

$$\mathcal{F}_{v}^{T}[\eta](u_{1}, u_{2})$$

$$:= \int_{u_{1}}^{u_{2}} \int_{\mathbb{S}^{2}} [|\nabla \underline{L}\eta|^{2} + (1-\mu)|\nabla \eta|^{2} + V|\eta|^{2}](u, v, \omega) \mathrm{d}u \mathrm{d}\mathbb{S}^{2}(\omega),$$
(3.4)

$$\begin{aligned} \mathcal{F}_{v}^{N}[\eta](u_{1}, u_{2}) \\ &:= \int_{u_{1}}^{u_{2}} \int_{\mathbb{S}^{2}} [(1-\mu)^{-1} |\nabla_{\underline{L}} \eta|^{2} + (1-\mu) |\nabla \eta|^{2} \\ &+ V |\eta|^{2}](u, v, \omega) \mathrm{d}u \mathrm{d}\mathbb{S}^{2}(\omega), \end{aligned}$$
(3.5)

$$\mathcal{F}^{\infty}[\eta](u_1, v_1) := \mathcal{F}^T_{u_1}[\eta](v_1, \infty) + \mathcal{F}^N_{v_1}[\eta](u_1, \infty).$$
(3.6)

– Let $q, x, s \in \mathbb{N}_{\geq 0}$. Let η as above. We define the weighted Sobolev norms

$$\begin{aligned} \|\eta\|_{C_{u_{0},v_{0}};q;x,s} &:= \sum_{i=0}^{x} \sum_{J \in \iota_{\leq s}^{\Omega}} \left\{ \mathcal{F}^{\infty}[(\nabla_{T})^{i} \mathcal{L}^{J} \eta](u_{0},v_{0}) \right. \\ &+ \int_{\overline{C}_{u_{0}} \cap \{r \geq R\}} r^{q} |\nabla_{L}(\nabla_{T})^{\min\{i,(x-1)^{+}\}} \mathcal{L}^{J} \eta|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \right\}. \end{aligned} (3.7)$$

3.1. The Maxwell System

Let F be an antisymmetric 2-form on S_e . Let us introduce the Maxwell Equations:

$$dF = 0, \quad d \star F = 0.$$
 (3.8)

Here, \star denotes the Hodge dual operator. More explicitly, if G is a two-form,

$$(\star G)_{\mu\nu} = \frac{1}{2} \varepsilon_{\alpha\beta\gamma\delta} G^{\gamma\delta}.$$
(3.9)

Equivalently, the system can be written as

$$\nabla_{[\mu}F_{\kappa\lambda]} = 0, \quad \nabla^{\mu}F_{\mu\nu} = 0.$$

Here, square brackets denote antisymmetrization of indices.

3.2. The Null Decomposition of the Maxwell System

Definition 3.3. Let $F \in \Lambda^2(\mathcal{S}_e)$. We define $\alpha, \underline{\alpha} \in \Gamma(\mathcal{B}^*)$, and $\rho, \sigma \in \mathcal{C}^{\infty}(\mathcal{S}_e)$ by the following relations:

$$\begin{aligned}
\alpha(V) &:= F(V, L), \\
\underline{\alpha}(V) &:= F(V, \underline{L}), \\
\rho &:= \frac{1}{2} \left(1 - \frac{2M}{r} \right)^{-1} F(\underline{L}, L), \\
\sigma &:= \frac{1}{2} \notin^{CD} F_{CD}.
\end{aligned}$$
(3.10)

for all $V \in \Gamma(\mathcal{B})$.

Remark 3.4. α and $\underline{\alpha}$ can also be viewed as one-forms in $\Gamma(T^*S_e)$, by requiring that they vanish on L and \underline{L} . Furthermore, in the definition of σ , we consider F as an element of $\Lambda^2(\mathcal{B})$, by restriction.

Remark 3.5. Note that all the previously defined quantities are smooth on S_e (up to the horizon), in the sense of our definition. This is due to the fact that the vector field $(1 - \mu)^{-1}\underline{L}$ is smooth on S_e , according to our definition.

Having introduced these quantities, we write the Maxwell system with respect to them. We have the following proposition.

Proposition 3.6. Let $F \in \Lambda^2(S_e)$, and let F satisfy the Maxwell system (3.8) on S_e . Then, defining the objects $\alpha, \underline{\alpha}, \rho, \sigma$ as in (3.10) we have that

$$\frac{1}{r} \nabla \underline{\mathcal{I}}(r\alpha_A) - (1-\mu)(\nabla_A \rho + \not \epsilon_{AB} \nabla^B \sigma) = 0, \qquad (3.12)$$

$$c \psi r l \,\underline{\alpha} - 2 \frac{1-\mu}{r} \sigma + \nabla \underline{L} \sigma = 0, \qquad (3.13)$$

$$- d\not v \,\underline{\alpha} + 2\frac{1-\mu}{r}\rho - \nabla \underline{L}\rho = 0, \qquad (3.14)$$

$$c \psi r l \ \alpha + 2 \frac{1-\mu}{r} \sigma + \nabla_L \sigma = 0, \tag{3.15}$$

$$d\not v \alpha - 2\frac{1-\mu}{r}\rho - \nabla _L \rho = 0.$$
(3.16)

Furthermore, the extreme components α and $\underline{\alpha}$ satisfy the spin ± 1 Teukolsky equations:

$$\nabla \underline{L} \nabla \underline{L} (r\alpha_A) + \frac{2}{r} \left(1 - \frac{3M}{r} \right) \nabla \underline{L} (r\alpha_A) - (1 - \mu) \Delta (r\alpha_A) + \frac{1 - \mu}{r^2} r\alpha_A = 0,$$
(3.17)

$$\nabla \underline{L} \nabla L(r\underline{\alpha}_A) - \frac{2}{r} \left(1 - \frac{3M}{r} \right) \nabla L(r\underline{\alpha}_A) - (1 - \mu) \Delta (r\underline{\alpha}_A) + \frac{1 - \mu}{r^2} r\underline{\alpha}_A = 0.$$
(3.18)

Proof of Proposition 3.6. We postpone the relevant calculations to "Appendix," Sect. A. \Box

3.3. Derivation of the Fackerell–Ipser Equation

We now proceed to introduce the crucial quantities ϕ and ϕ , and we prove that, if we only require the spin ± 1 Teukolsky equations to hold for α and α , then ϕ and ϕ satisfy the so-called Fackerell–Ipser equation.

Proposition 3.7. Let α satisfy the spin +1 Teukolsky equation (3.17) and let $\underline{\alpha}$ satisfy the -1 Teukolsky equation (3.18) on S_e . Then, we define

$$\phi_A := \frac{r^2}{1-\mu} \nabla \underline{L}(r\alpha_A),$$

$$\underline{\phi}_A := \frac{r^2}{1-\mu} \nabla \underline{L}(r\underline{\alpha}_A).$$
(3.19)

Under these hypotheses, ϕ and ϕ satisfy the Fackerell–Ipser Equation:

$$\nabla \underline{L} \nabla_L \phi_A - (1-\mu) \Delta(\phi_A) + \frac{1-\mu}{r^2} \phi_A = 0, \qquad (3.20)$$

$$\nabla \underline{L} \nabla \underline{L} \phi_A - (1-\mu) \Delta (\underline{\phi}_A) + \frac{1-\mu}{r^2} \underline{\phi}_A = 0.$$
(3.21)

Remark 3.8. We remark that, if we further assume that α and $\underline{\alpha}$ are part of a solution $(\rho, \sigma, \alpha, \underline{\alpha})$ of the Maxwell equations (3.11)–(3.16) on S_e , the following relations hold true:

We also remark that, in this case, the tensorial Fackerell–Ipser Equation can be obtained from the wave equation (scalar Fackerell–Ipser) satisfied by the middle components, commuting with the projected covariant angular derivative ∇_A .

Proof of Proposition 3.7. It is a straightforward calculation from the Teukolsky equation. We restrict to $\underline{\phi}$, the reasoning for ϕ being analogous. First of all, we notice that the Teukolsky equation for $\underline{\alpha}$ is equivalent to

For,

$$\underline{L}\left(\frac{r^2}{1-\mu}\right) = -\frac{r^2}{1-\mu}\frac{2}{r}\left(1-\frac{3M}{r}\right).$$

Multiply Eq. (3.22) by $\frac{r^2}{1-\mu}$ and subsequently take the ∇_L derivative of both sides. We obtain, since $[\nabla_L, r^2 \Delta] = 0$,

$$\nabla_L \nabla_L \left(\frac{r^2}{1-\mu} \nabla_L(r\underline{\alpha}_A) \right) - r^2 \Delta \nabla_L(r\underline{\alpha}_A) + \nabla_L(r\underline{\alpha}_A) = 0.$$

This implies the claim.

4. Statements of the Main Results

In this section, we state the main results of the present note. The first result, Theorem 4.1, only deals with solutions to the spin ± 1 Teukolsky equations on a fixed Schwazrzschild background, and provides decay rates for them. The second result, Theorem 4.2, concerns a solution F to the full Maxwell system on a fixed Schwazrschild background, and derives decay estimates for the relevant quantities using Theorem 4.1.

Theorem 4.1 (Decay of solutions to the Teukolsky equations). There exist a positive real number $R_* > 0$ and a positive constant C depending only on M and R_* such that letting (u_0, v_0) be real numbers such that $v_0 - u_0 = 2R_*$,

we have the following. Let $\alpha, \underline{\alpha} \in \Gamma(\mathcal{B}^*)$ be solutions to the spin ± 1 Teukolsky equations on $[u_0, \infty) \times [v_0, \infty) \times \mathbb{S}^2 \subset \mathcal{S}_e$:

$$\nabla \underline{L} \nabla \underline{L} (r \alpha_A) + \frac{2}{r} \left(1 - \frac{3M}{r} \right) \nabla \underline{L} (r \alpha_A) - (1 - \mu) \Delta (r \alpha_A) + \frac{1 - \mu}{r^2} r \alpha_A = 0,$$

$$\nabla \underline{L} \nabla \underline{L} (r \underline{\alpha}_A) - \frac{2}{r} \left(1 - \frac{3M}{r} \right) \nabla \underline{L} (r \underline{\alpha}_A) - (1 - \mu) \Delta (r \underline{\alpha}_A) + \frac{1 - \mu}{r^2} r \underline{\alpha}_A = 0.$$

Let ϕ , ϕ be the related quantities as in (3.19). Under these assumptions, ϕ and ϕ satisfy the Morawetz estimate (5.8) of Lemma 5.2, as well as the hierarchy of integrated estimates (5.23) and (5.24).

Furthermore, let χ be a smooth cutoff function such that $\chi(r) = 1$ for $r \geq 3M$, and $\chi(r) = 0$ for $r \in [2M, 3/2M]$. Let $\underline{\widetilde{\alpha}} = (1 - \mu)^{-1} \underline{\alpha}$ and $\overline{\Psi} := \chi(r)(1 - \mu)^{-1}r^3 \alpha$.

Under these conditions, we have the pointwise estimates:

$$|\alpha| \le C \frac{\left\|\overline{\Psi}\right\|_{C_{u_0,v_0};0;0,2} + \|\phi\|_{C_{u_0,v_0};2;2,1}}{v} \quad and \tag{4.1}$$

$$|\underline{\widetilde{\alpha}}| \le C \frac{\|\underline{\widetilde{\alpha}}\|_{C_{u_0,v_0};0;0,0} + \|\underline{\phi}\|_{C_{u_0,v_0};2;2,1}}{v}$$
(4.2)

on $\{r_* \leq R_*\} \cap \{u \geq u_0\} \cap \{v \geq v_0\}$. Also, we have

$$\alpha| \le C \frac{\left\|\overline{\Psi}\right\|_{C_{u_0,v_0};0;0,2} + \|\phi\|_{C_{u_0,v_0};2;q,1}}{v^{q/2}r^{3/2}} \quad for \ q \in \{0,1,2\}, \ and$$
(4.3)

$$|\alpha| \le C \frac{\|\Psi\|_{C_{u_0,v_0};2;0,2} + \|\phi\|_{C_{u_0,v_0};2;2,2}}{(|u|+1)^{\frac{1}{2}}r^3}, \quad and$$

$$(4.4)$$

$$|\underline{\alpha}| \le C \frac{\|\underline{\widetilde{\alpha}}\|_{C_{u_0,v_0};0;0,0} + \|\underline{\phi}\|_{C_{u_0,v_0};2;2,1}}{(|u|+1)r}$$
(4.5)

on $\{r_* \geq R_*\} \cap \{u \geq u_0\} \cap \{v \geq v_0\}$. Here, we used the definition of norm in (3.7).

Theorem 4.2 (Decay of solutions to the Maxwell system). There exist a positive real number $R_* > 0$ and a positive constant C depending only on M and R_* such that, letting (u_0, v_0) be real numbers such that $v_0 - u_0 = 2R_*$, we have the following. Let $F \in \Lambda^2(TS_e)$ be a solution to the Maxwell system on $\{u \ge u_0\} \cap \{v \ge v_0\}$:

$$\mathrm{d}F = 0, \quad \mathrm{d} \star F = 0.$$

Recall the definition of the null components $\alpha, \underline{\alpha}, \rho, \sigma$ (3.10). We let

$$\rho_s := \frac{R^2}{4\pi} \int_{\omega \in \mathbb{S}^2} \rho(u_0, v_0, \omega) \mathrm{d}\mathbb{S}^2(\omega), \quad \sigma_s := \frac{R^2}{4\pi} \int_{\omega \in \mathbb{S}^2} \sigma(u_0, v_0, \omega) \mathrm{d}\mathbb{S}^2(\omega).$$

$$(4.6)$$

Recall the definition of ϕ and $\underline{\phi}$ from (3.19). Let χ be a smooth cutoff function such that $\chi(r) = 1$ for $r \geq 3M$, and $\chi(r) = 0$ for $r \in [2M, 3/2M]$. Let us furthermore set $\underline{\widetilde{\alpha}} = (1-\mu)^{-1}\underline{\alpha}, \overline{\Psi} := \chi(r)(1-\mu)^{-1}r^3\alpha$, and

$$M_{\alpha} := \left\| \overline{\Psi} \right\|_{C_{u_0, v_0}; 2; 0, 2} + \left\| \phi \right\|_{C_{u_0, v_0}; 2; 2, 1}, \tag{4.7}$$

$$M_{\rho,\sigma} := \|\phi\|_{C_{u_0,v_0};2;2,1} + \|\underline{\phi}\|_{C_{u_0,v_0};2;2,1}, \qquad (4.8)$$

$$M_{\underline{\alpha}} := \|\underline{\widetilde{\alpha}}\|_{C_{u_0,v_0};0;0,0} + \|\underline{\phi}\|_{C_{u_0,v_0};2;2,1}.$$
(4.9)

Then, we have

$$|\alpha|, (1-\mu)^{-1}|\underline{\alpha}| \le Cv^{-1}(M_{\alpha} + M_{\underline{\alpha}}), \tag{4.10}$$

$$\left|\rho - \frac{\rho_s}{r^2}\right|, \left|\sigma - \frac{\sigma_s}{r^2}\right| \le Cv^{-1}M_{\rho,\sigma} \tag{4.11}$$

 $on \ \{r_* \leq R_*\} \cap \{u \geq u_0\} \cap \{v \geq v_0\}.$ Furthermore,

$$|\alpha| \le C(|u|+1)^{-\frac{1}{2}} r^{-3} M_{\alpha}, \qquad (4.12)$$

$$|\underline{\alpha}| \le C(|u|+1)^{-1} r^{-1} M_{\underline{\alpha}}, \qquad (4.13)$$

$$\left|\rho - \frac{\rho_s}{r^2}\right|, \left|\sigma - \frac{\sigma_s}{r^2}\right| \le C(|u|+1)^{-\frac{1}{2}}r^{-2}M_{\rho,\sigma},$$
(4.14)

$$\left| \rho - \frac{\rho_s}{r^2} \right|, \left| \sigma - \frac{\sigma_s}{r^2} \right| \le C(|u| + 1)^{-1} r^{-3/2} M_{\rho,\sigma}$$
(4.15)

on $\{r_* \ge R_*\} \cap \{u \ge u_0\} \cap \{v \ge v_0\}.$

A few remarks are in order.

Remark 4.3. We will not delve into the issue of optimal well-posedness statements for the Maxwell system or for the spin ± 1 Teukolsky equations here. Let us just remark that, in the smooth category, well-posedness for the characteristic initial value problem follows, in both situations, from ideas contained in the work by Rendall [16].

Remark 4.4. We notice that Theorem 4.2 gives the decay rate v^{-1} for all components of the field in the region $\{r \leq R\}$. Furthermore, for all components of the field, we have the **uniform peeling estimates**, on a fixed outgoing null cone $\overline{C}_{\tilde{u}} \cap \{r \geq R\}$:

$$\begin{aligned} |\alpha| &\lesssim r^{-3} (|\tilde{u}|+1)^{-\frac{1}{2}}, \\ |\rho|, |\sigma| &\lesssim r^{-2} (|\tilde{u}|+1)^{-\frac{1}{2}}, \\ |\underline{\alpha}| &\lesssim r^{-1} (|\tilde{u}|+1)^{-1}. \end{aligned}$$
(4.16)

Here, we supposed for simplicity that $M_{\rho,\sigma}$, M_{α} , $M_{\underline{\alpha}}$ all be finite, and that $\rho_s = \sigma_s = 0$. The bound for α is the strongest one, corresponding to inequality (4.4).

We underline the difference between estimates (4.3) and (4.4). In the former, we require less of the initial data to obtain a lower decay rate. In the latter, we have a larger weight on the *L*-derivative of Ψ , and we obtain a uniform peeling estimate for α . Weaker requirements on initial data, though implying weaker decay, may be useful for applications to nonlinear problems, in view of a bootstrap argument. An example is the original proof of the nonlinear stability of the Minkowski spacetime by Christodoulou and Klainerman [8], in which the authors do not need optimal decay rates in order to close the argument. In fact, the failure of peeling to hold has a physical interpretation [7].

Remark 4.5 (On initial data). We remark that, in order to solve the full Maxwell system (3.11)–(3.16), it is enough to impose initial data for α and $\underline{\alpha}$ on the set \mathcal{C}_{u_0,v_0} . For, then, all first derivatives of α and $\underline{\alpha}$ along \mathcal{C}_{u_0,v_0} can be recovered via the spin ± 1 Teukolsky equations. Then, we can solve for α and $\underline{\alpha}$ in $\{u \geq u_0\} \cap \{v \geq v_0\}$, again from the Teukolsky equations. Finally, we can use relations (3.11) and (3.12) to recover all angular derivatives of ρ and σ . This defines uniquely a solution up to the stationary solutions (2.9). The resulting quantities $\alpha, \underline{\alpha}, \sigma, \rho$ then satisfy the full Maxwell system (3.11)–(3.16).

Remark 4.6. For notational convenience, the norms we defined in Eq. (3.7) are not intrinsic to the surface C_{u_0,v_0} . Nevertheless, using the Fackerell–Ipser equation for ϕ and ϕ , it can be shown that

$$\begin{split} \|\phi\|_{C_{u_0,v_0};2;2,0}^2 &\leq \int_{\overline{C}_{u_0}\cap\{r\geq R\}} \sum_{I\in\iota_{\leq 3}^+} (|\nabla_L \nabla^I \phi|^2 + |\nabla \nabla^I \phi|^2 + r^{-2} |\nabla^I \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2 \\ &+ \int_{\overline{C}_{u_0}\cap\{r\geq R\}} (r^2 |\nabla_L \nabla_L \phi|^2 + r^2 |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2 \\ &+ \int_{\underline{C}_{u_0}\cap\{r\leq R\}} \sum_{J\in\iota_{\leq 4}^-} |\nabla^J \phi| (1-\mu) \mathrm{d}u \mathrm{d}\mathbb{S}^2. \end{split}$$

$$(4.17)$$

Here, $\iota_{\leq 2}^+$ (resp. $\iota_{\leq 2}^-$) is the set of all ordered lists of length ≤ 2 composed of elements of $\Omega \cup \{L\}$ (resp. $\Omega \cup \{(1-\mu)^{-1}\underline{L}\}$).

Estimate (4.17) implies in particular that the norm $\|\phi\|_{C_{u_0,v_0};2;2,0}^2$ can be controlled in terms of a norm intrinsic to the surface C_{u_0,v_0} . Similar expressions hold for α and $\underline{\alpha}$.

Furthermore notice that estimate (4.17) "loses derivatives." On the lefthand side, the norm $\|\phi\|_{C_{u_0,v_0};2;2,0}^2$ depends on 3 "unweighted" derivatives and on 2 "weighted" derivatives. The norm on the right-hand side of (4.17), on the other hand, depends on 4 "unweighted" derivatives and 2 "weighted" derivatives.

Remark 4.7 (On the propagation of decay from initial data). Let ω be a nontrivial one-form on \mathbb{S}^2 . Suppose for ease of exposition that $\mathbb{A}_{\mathbb{S}^2}\omega = 2\omega$. Let $f_1(r)$ be a smooth function of r. Following Remark 4.5, let us set initial data for α on \mathcal{C}_{u_0,v_0} in the following way:

$$\alpha = f_1(r)\omega \quad \text{on } \mathcal{C}_{u_0,v_0}. \tag{4.18}$$

We then use the relation (spin +1 Teukolsky equation) to induce data for ϕ :

$$\nabla \!\!\!\!/_L \phi_A = r^2 \Delta (r \alpha_A) - r \alpha_A.$$

From the latter, it follows that initial data for ϕ satisfies

$$\phi(u_0, v) = \left(\int_{v_0}^v r(u_0, \tilde{v}) f_1(r(u_0, \tilde{v})) \mathrm{d}\tilde{v}\right) \omega + \eta, \tag{4.19}$$

where η is a fixed one-form on \mathbb{S}^2 . Here, $r(u_0, \tilde{v})$ denotes the *r*-coordinate of the point (u_0, \tilde{v}) in (u, v)-coordinates. Let us denote

$$f_2(\bar{r}) := \int_{v_0}^v r(u_0, \tilde{v}) f_1(r(u_0, \tilde{v})) \mathrm{d}\tilde{v},$$

whenever the *r*-coordinate of the point (u_0, v) is \bar{r} .

Now, let $s \in (\frac{3}{2}, 1)$, and let us suppose the following on the function f_1 :

$$|f_1(r)| \sim r^{-1-s}, \quad |f_1'(r)| \sim r^{-2-s}, \quad |f_1''(r)| \sim r^{-3-s},$$
 (4.20)

as $r \to \infty$.

It then follows that

$$|f_2(r)| \sim 1, \quad |f'_2(r)| \sim r^{-s}, \quad |f''_2(r)| \sim r^{-s-1}, \quad |f''_2(r)| \sim r^{-s-2}, \quad (4.21)$$

It follows now, from (4.17) and the form of ϕ (4.19), that

$$\|\phi\|_{C_{u_0,v_0};2;2,1} < \infty.$$

Under conditions (4.20), we furthermore have that, recalling $\overline{\Psi}_A = \chi(r)r^3(1-\mu)^{-1}\alpha_A$, with $\chi(r)$ smooth supported away from r = 2M, such that $\chi(r) = 1$ for $r \geq 3M$,

$$\left\|\overline{\Psi}\right\|_{C_{u_0,v_0};0;0,2}<\infty.$$

In this case, the assumptions of estimate (4.3) are satisfied with q = 2, and we obtain the bound $|\alpha| \leq Cr^{-\frac{5}{2}}$ for α along any fixed outgoing cone, whereas we supposed that $|\alpha|$ is asymptotic to r^{-1-s} on \mathcal{C}_{u_0,v_0} , with $s+1 > \frac{5}{2}$. In this case, we do not recover the initial decay.

On the other hand, if we require $s \ge 1$, we obtain

$$\left\|\overline{\Psi}\right\|_{C_{u_0,v_0};2;0,2}<\infty,$$

and the assumptions of estimate (4.4) are satisfied. We then have $|\alpha| \leq Cr^{-3}$ for some constant C, along a fixed outgoing null cone.

In particular, for s = 1, we are able to propagate the r^{-3} initial decay. Similar statements hold for $\underline{\alpha}$, ρ , σ .

We finally remark that, if we were to a sharper decay than r^{-3} for α on the initial cone \mathcal{C}_{u_0,v_0} , generically, it would not propagate.

5. Estimates on the Fackerell–Ipser Equation

In this section, we prove integrated decay estimates for solutions to the Fackerell–Ipser equation. The estimates and the methods to obtain them are very similar to those in [11]. The results contained in this section are of independent interest, and the section can be read independently from the rest of the paper, starting from the assumption that ϕ only satisfies the Fackerell–Ipser equation.

We do not prove pointwise decay for ϕ , as it clearly follows from the ideas in the proof of Theorem 4.2, cf. Remark 3.8.

Let us now proceed to the setup. Let $v_2 \ge v_1 \ge v_0$ and $u_2 \ge u_1 \ge u_0$. Let $V = (1 - \mu)/r^2$. Recall the definition of the null fluxes and of the Sobolev norms:

$$\mathcal{F}_{u}^{T}[\phi](v_{1}, v_{2}) \\ := \int_{v_{1}}^{v_{2}} \int_{\mathbb{S}^{2}} [|\nabla_{L}\phi|^{2} + (1-\mu)|\nabla\phi|^{2} + V|\phi|^{2}](u, v, \omega) \mathrm{d}v \mathrm{d}\mathbb{S}^{2}(\omega),$$

$$\mathcal{F}_{v}^{T}[\phi](u_{1}, u_{2})$$
(5.1)

$$:= \int_{u_1}^{u_2} \int_{\mathbb{S}^2} [|\nabla \!\!\!/_{\underline{L}} \phi|^2 + (1-\mu) |\nabla \!\!\!/ \phi|^2 + V |\phi|^2](u,v,\omega) \mathrm{d}u \mathrm{d}\mathbb{S}^2(\omega), \tag{5.2}$$

$$\mathcal{F}_{v}^{N}[\phi](u_{1}, u_{2})$$

$$:= \int_{u_{1}}^{u_{2}} \int_{\mathbb{S}^{2}} [(1-\mu)^{-1} |\nabla \underline{L} \phi|^{2} + (1-\mu) |\nabla \phi|^{2} + V |\phi|^{2}](u, v, \omega) \mathrm{d}u \mathrm{d}\mathbb{S}^{2}(\omega),$$
(5.3)

$$\mathcal{F}^{\infty}[\phi](u_{1}, v_{1}) := \mathcal{F}^{T}_{u_{1}}[\phi](v_{1}, \infty) + \mathcal{F}^{N}_{v_{1}}[\phi](u_{1}, \infty),$$

$$:= \|\phi\|_{C_{u_{0}, v_{0}}; q; x, s}$$

$$\sum_{i=0}^{x} \sum_{J \in \iota^{\Omega}_{\leq s}} \left\{ \mathcal{F}^{\infty}[(\nabla_{T})^{i} \mathcal{L}^{J} \phi](u_{0}, v_{0}) + \int r^{q} |\nabla_{\tau}(\nabla_{T})^{\min\{i, (x-1)^{+}\}} \mathcal{L}^{J} \phi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \right\}$$
(5.4)
$$(5.4)$$

$$+ \int_{\overline{C}_{u_0} \cap \{r \ge R\}} r^q |\nabla_L (\nabla_T)^{\min\{i, (x-1)^+\}} \mathcal{L}^J \phi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 \bigg\}.$$

$$(5.5)$$

Let $R_* > 0$. We define the spacetime regions (Fig. 1):

$$\mathfrak{D}_{u_1}^{u_2} := \{r \ge R, u \in [u_1, u_2]\}, \\
\mathfrak{E}_{v_1}^{v_2} := \{r \le R, v \in [v_1, v_2]\}, \\
\mathfrak{F}_{u_1}^{u_2} := \mathfrak{D}_{u_1}^{u_2} \cup \mathfrak{E}_{v_1}^{v_2}, \text{ such that } v_1 - u_1 = 2R_* \text{ and } v_2 - u_2 = 2R_*.$$
(5.6)

5.1. Energy Conservation

Lemma 5.1. Let ϕ be a smooth solution to the Fackerell–Ipser Equation (3.20) on $\{u \ge u_0\} \cap \{v \ge v_0\}$. Let $v_2 \ge v_1 \ge v_0$ and $u_2 \ge u_1 \ge u_0$. Defining the fluxes as in (5.1) and (5.2), we have that

$$\mathcal{F}_{u_2}^T[\phi](v_1, v_2) + \mathcal{F}_{v_2}^T[\phi](u_1, u_2) = \mathcal{F}_{v_1}^T[\phi](u_1, u_2) + \mathcal{F}_{u_1}^T[\phi](v_1, v_2).$$
(5.7)

Proof. First, we notice that the Fackerell–Ipser Equation implies:

$$\begin{split} (\boldsymbol{\nabla}_{\underline{L}} + \boldsymbol{\nabla}_{L}) \int_{\mathbb{S}^{2}} \left\{ | \boldsymbol{\nabla}_{L} \phi |^{2} + | \boldsymbol{\nabla}_{\underline{L}} \phi |^{2} + 2 \frac{1-\mu}{r^{2}} | r \boldsymbol{\nabla} \phi |^{2} + 2 V | \phi |^{2} \right\} \mathrm{d} \mathbb{S}^{2} \\ &+ (\boldsymbol{\nabla}_{\underline{L}} - \boldsymbol{\nabla}_{L}) \int_{\mathbb{S}^{2}} \left\{ | \boldsymbol{\nabla}_{L} \phi |^{2} - | \boldsymbol{\nabla}_{\underline{L}} \phi |^{2} \right\} \mathrm{d} \mathbb{S}^{2} = 0. \end{split}$$

Integrating with respect to dudv yields the claim.



FIGURE 1. Penrose diagram depicting of the spacetime regions

5.2. Morawetz Estimate

Lemma 5.2. There exists a positive constant C such that the following holds. Let ϕ be a smooth solution to the Fackerell–Ipser Equation (3.20) on $\{u \ge u_0\} \cap \{v \ge v_0\}$. Defining the fluxes as in (5.1), (5.2), we have that

$$\begin{split} &\int_{u_1}^{u_2} \int_{v_1}^{v_2} \int_{\mathbb{S}^2} \left\{ \frac{1}{r^2} |\nabla_L \phi - \nabla_{\underline{L}} \phi|^2 + \frac{(r - 3M)^2}{r^3} \left(|\nabla \phi|^2 + \frac{1}{r} |\nabla_L \phi + \nabla_{\underline{L}} \phi|^2 \right) \\ &+ \frac{1}{r^3} |\phi|^2 \right\} (1 - \mu) \mathrm{d} u \mathrm{d} v \mathrm{d} \mathbb{S}^2 \\ &\leq C(\mathcal{F}_{v_1}^T[\phi](u_1, u_2) + \mathcal{F}_{u_1}^T[\phi](v_1, v_2)). \end{split}$$
(5.8)

Proof of Lemma 5.2. We consider the following identities, which follow from the Fackerell–Ipser equation. Let $f: S_e \to \mathbb{R}$ be a smooth radial function. Let $(\cdot)'$ denote differentiation with respect to ∂_{r_*} . Recall that $V = \frac{1-\mu}{r^2}$:

$$\begin{aligned} (\nabla \underline{L} + \nabla L) &\int_{\mathbb{S}^2} f\left\{ |\nabla_L \phi|^2 - |\nabla \underline{L} \phi|^2 \right\} d\mathbb{S}^2 \\ &+ (\nabla \underline{L} - \nabla L) \int_{\mathbb{S}^2} f\left\{ |\nabla_L \phi|^2 + |\nabla \underline{L} \phi|^2 - 2\frac{1-\mu}{r^2} |r \nabla \phi|^2 - 2V |\phi|^2 \right\} d\mathbb{S}^2 \\ &+ \int_{\mathbb{S}^2} \left\{ 2f'(|\nabla_L \phi|^2 + |\nabla \underline{L} \phi|^2) - 4\partial_{r^\star} \left(f \frac{1-\mu}{r^2} \right) |r \nabla \phi|^2 \\ &- 4\partial_{r^\star} (fV) |\phi|^2 \right\} d\mathbb{S}^2 = 0. \end{aligned}$$

$$(5.9)$$

We also have

$$(\nabla \underline{L} + \nabla \underline{L}) \int_{\mathbb{S}^2} \left(f' \phi \cdot (\nabla \underline{L} + \nabla \underline{L}) \phi \right) d\mathbb{S}^2$$

$$- (\nabla \underline{L} - \nabla \underline{L}) \int_{\mathbb{S}^2} \left(f' \phi \cdot (\nabla \underline{L} - \nabla \underline{L}) \phi + f'' |\phi|^2 \right) d\mathbb{S}^2$$

$$+ \int_{\mathbb{S}^2} \left\{ -2f''' |\phi|^2 - 4f' \nabla \underline{L} \phi \cdot \nabla \underline{L} \phi$$

$$+ 4f' \left(\frac{1-\mu}{r^2} |r \nabla \phi|^2 + V |\phi|^2 \right) \right\} d\mathbb{S}^2 = 0.$$

$$(5.10)$$

In the previous equation, the dot \cdot indicated that we are contracting with $\not g$. Let us now add the previous Eqs. (5.9) and (5.10), to get

$$\begin{aligned} \left(\nabla \underline{L} + \nabla L \right) & \int_{\mathbb{S}^2} \left(f \left\{ |\nabla_L \phi|^2 - |\nabla_{\underline{L}} \phi|^2 \right\} + f' \phi \cdot \left(\nabla_{\underline{L}} + \nabla_L \right) \phi \right) d\mathbb{S}^2 \\ &+ \left(\nabla_{\underline{L}} - \nabla_L \right) \int_{\mathbb{S}^2} \left(f \left\{ |\nabla_L \phi|^2 + |\nabla_{\underline{L}} \phi|^2 - 2\frac{1-\mu}{r^2} |r \nabla \phi|^2 - 2V |\phi|^2 \right\} \\ &- f' \phi \cdot \left(\nabla_{\underline{L}} - \nabla_L \right) \phi - f'' |\phi|^2 \right) d\mathbb{S}^2 \\ &+ \int_{\mathbb{S}^2} \left\{ 2f' (|\nabla_L \phi - \nabla_{\underline{L}} \phi|^2) + |r \nabla \phi|^2 \left[-4f \left(\frac{1-\mu}{r^2} \right)' \right] \\ &+ |\phi|^2 (-4fV' - 2f''') \right\} d\mathbb{S}^2 = 0. \end{aligned}$$

$$(5.11)$$

We now proceed to integrate Eq. (5.11) on spacetime against the form dudv. By Lemma B.1, we note that for the bulk term to be positive, it suffices that there exists a c > 0 such that the following three inequalities hold true:

$$-2\frac{(V+\frac{1}{r^2}(1-\mu))'}{1-\mu}f - \frac{f'''}{1-\mu} \ge \frac{c}{r^3},$$
(5.12)

$$-4f\left(\frac{1-\mu}{r^2}\right)' > 0,\tag{5.13}$$

$$f' > 0.$$
 (5.14)

By choosing $f(r) := \left(1 + \frac{M}{r}\right) \left(1 - \frac{3M}{r}\right)$, let us calculate, as in [11],

$$f' = (1 - \mu) \left(\frac{2M}{r^2} + \frac{6M^2}{r^3}\right),$$

$$\begin{split} f'' &= (1-\mu)\partial_r(f') = \frac{2M\left(-48M^3 + 30M^2r + Mr^2 - 2r^3\right)}{r^6},\\ f''' &= (1-\mu)\partial_r(f'') = \frac{4M(1-\mu)\left(144M^3 - 75M^2r - 2Mr^2 + 3r^3\right)}{r^7},\\ \left(V + \frac{1-\mu}{r^2}\right)' &= 2\left(\frac{1-\mu}{r^2}\right)' = -4\frac{(1-\mu)}{r^3}\left(1 - \frac{3M}{r}\right),\\ &-4f\left(\frac{1-\mu}{r^2}\right)' = 8\frac{(1-\mu)}{r^3}\left(1 - \frac{3M}{r}\right)^2\left(1 + \frac{M}{r}\right). \end{split}$$

These calculations immediately show that conditions (5.13) and (5.14) are satisfied. Plugging these calculations in condition (5.12), multiplying by $\frac{1}{4}r^3$, and renaming $x := \frac{M}{r}$, we obtain that it suffices, for (5.12) to hold, that there exists c > 0 such that, for all $x \in [0, 1/2]$,

$$-144x^4 + 93x^3 + 8x^2 - 13x + 2 \ge c.$$

An analysis of the polynomial reveals that this inequality is true for all $x \in [0, 1/2]$ upon choosing, for instance, c = 1/8 (in fact, the minimum is at x = 1/2, where the polynomial takes the value 1/8).

We therefore obtain the following estimate, making use of the positivity of the angular terms:

$$\int_{u_{1}}^{u_{2}} \int_{v_{1}}^{v_{2}} \int_{\mathbb{S}^{2}} \left\{ \frac{1}{r^{2}} |\nabla_{L}\phi - \nabla_{\underline{L}}\phi|^{2} + \frac{(r-3M)^{2}}{r^{3}} |\nabla\phi|^{2} + \frac{1}{r^{3}} |\phi|^{2} \right\} (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \qquad (5.15)$$

$$\leq C(F_{u_{1}}^{T}[\phi](v_{1}, v_{2}) + F_{v_{1}}^{T}[\phi](v_{1}, v_{2})).$$

We can recover the missing derivative by integrating Eq. (5.9) with a monotonically increasing f, which vanishes of third order at r = 3M, and get:

$$\int_{u_{1}}^{u_{2}} \int_{v_{1}}^{v_{2}} \int_{\mathbb{S}^{2}} \left\{ \frac{1}{r^{2}} |\nabla_{L}\phi - \nabla_{\underline{L}}\phi|^{2} + \frac{(r-3M)^{2}}{r^{3}} \left(|\nabla\phi|^{2} + \frac{1}{r} |\nabla_{L}\phi + \nabla_{\underline{L}}\phi|^{2} \right) \\
+ \frac{1}{r^{3}} |\phi|^{2} \right\} (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\
\leq C(\mathcal{F}_{v_{1}}^{T}[\phi](u_{1}, u_{2}) + \mathcal{F}_{u_{1}}^{T}[\phi](v_{1}, v_{2})).$$
(5.16)

This is the claim.

5.3. The Redshift Estimate

Lemma 5.3 (The redshift estimate). There exists a positive constant C such that the following holds. Let ϕ be a smooth solution to the Fackerell–Ipser

Equation (3.20) on $\{u \geq u_0\} \cap \{v \geq v_0\}$. Let $u_2 \geq u_1 \geq u_0$ and $v_2 \geq v_1 \geq v_0$. Let $\chi(r)$ be a smooth cutoff function such that $\chi_{\mathcal{H}^+}(r) = 1$ for $r \in (2M, 5/2M)$ and $\chi_{\mathcal{H}^+}(r) = 0$ for $r \geq 3M$. Defining the fluxes as in (5.1)–(5.4), we have that

$$\begin{aligned} \mathcal{F}^{\infty}(u_{2},v_{2}) &\leq C\mathcal{F}^{\infty}(u_{1},v_{1}), \end{aligned} \tag{5.17} \\ \int_{u_{1}}^{u_{2}} \int_{v_{1}}^{v_{2}} \int_{\mathbb{S}^{2}} \left\{ \frac{1}{r^{2}} | \nabla_{L}\phi - \nabla_{\underline{L}}\phi |^{2} + \frac{(r-3M)^{2}}{r^{3}} \left(| \nabla\phi|^{2} + \frac{1}{r} | \nabla_{L}\phi + \nabla_{\underline{L}}\phi |^{2} \right) + \frac{1}{r^{3}} |\phi|^{2} \right\} (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &+ \int_{u_{1}}^{u_{2}} \int_{v_{1}}^{v_{2}} \int_{\mathbb{S}^{2}} \chi_{\mathcal{H}^{+}}(r) | \nabla_{(1-\mu)^{-1}\underline{L}}\phi |^{2} (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &\leq C(\mathcal{F}_{v_{1}}^{N}[\phi](u_{1},u_{2}) + \mathcal{F}_{u_{1}}^{T}[\phi](v_{1},v_{2})). \end{aligned} \tag{5.18}$$

Proof. Let $2M < r_c < 3M$. Let h(r) be a smooth radial function such that $h(r) = (1 - \mu)^{-1}$ for $r \in (2M, r_c)$, h(r) = 0 for $r \in [3M, +\infty)$. The following relation can be deduced from the Fackerell–Ipser equation:

$$L\{h(r)|\nabla \underline{L}\phi|^{2}\} + \underline{L}\{(1-\mu)h(r)|\nabla \phi|^{2}\} + \underline{L}\{\frac{(1-\mu)}{r^{2}}h(r)|\phi|^{2}\} - L\{h(r)\}|\nabla \underline{L}\phi|^{2} - \underline{L}\{\frac{(1-\mu)}{r^{2}}h(r)\}|\phi|^{2} - r^{2}\underline{L}\{\frac{(1-\mu)}{r^{2}}h(r)\}|\nabla \phi|^{2} \stackrel{\text{s}^{2}}{=} 0.$$
(5.19)

Here, $A(u, v, \omega) \stackrel{\mathbb{S}^2}{=} 0$ denotes equality after integration on \mathbb{S}^2 against the form $d\mathbb{S}^2$, i.e., $\int_{\mathbb{S}^2} A(u, v, \omega) d\mathbb{S}^2(\omega) = 0$. We proceed to integrate (5.19) against the form dudv in the region $\{u_1 \leq u \leq u_2\} \cap \{v_1 \leq v \leq v_2\}$. The second claim is achieved absorbing the error terms arising from h with estimate (5.8) (note that we only require one derivative on the initial slice).

Analyzing the boundary terms arising from integration of (5.19) and adding a multiple of (5.7), we obtain the first claim.

5.4. The Estimates Required for the r^p Method

Definition 5.4. Let $\chi_{\mathcal{H}^+}$ be the smooth cutoff function of Lemma 5.3. We define $m[\phi]$ to be the integrand in the left-hand side of Eq. (5.18):

$$m[\phi] := \frac{1}{r^2} |\nabla_L \phi - \nabla_{\underline{L}} \phi|^2 + \frac{(r-3M)^2}{r^3} \left(|\nabla \phi|^2 + \frac{1}{r} |\nabla_L \phi + \nabla_{\underline{L}} \phi|^2 \right) + \frac{1}{r^3} |\phi|^2 + \chi_{\mathcal{H}^+}(r) |\nabla_{(1-\mu)^{-1}\underline{L}} \phi|^2.$$
(5.20)

Furthermore, we let

$$m_{\rm all}[\phi] := m[\phi] + m[\nabla_T \phi]. \tag{5.21}$$

Remark 5.5. Let us notice that the Fackerell–Ipser Equation implies that there exists a constant C > 0, depending only on M, such that

$$\begin{split} &\int_{u_1}^{\infty} \int_{v_1}^{\infty} \int_{\mathbb{S}^2} m_{\text{all}}[\phi](1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 \\ &\geq C \int_{u_1}^{\infty} \int_{v_1}^{\infty} \int_{\mathbb{S}^2} \left(\frac{1}{r^2} (|\nabla_L \phi|^2 + |\nabla_{(1-\mu)^{-1}\underline{L}} \phi|^2) \right. \\ &\left. + \frac{1}{r^3} |\phi|^2 + \frac{1}{r} |\nabla \phi|^2 \right) (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2. \end{split}$$

This is seen by the following observation. It is evident that $m_{\rm all}[\phi]$ controls in a pointwise fashion all non-degenerate derivatives at r = 2M and at r = 3M, with the exception of the angular derivatives at r = 3M. To bound the angular derivatives at the photon sphere, we contract the Fackerell–Ipser equation with $g(r)\phi^A$, where g(r) is a radial cutoff function which is 1 in a neighborhood of r = 3M, and vanishes outside of a neighborhood of r = 3M. We then integrate by parts (in the \underline{L} direction for the term $\nabla_L \nabla_{\underline{L}}$, and in the angular direction for the angular term) and see that all the error terms can be bounded in terms of $m_{\rm all}[\phi]$. We conclude that $m_{\rm all}[\phi]$ controls all non-degenerate derivatives of ϕ everywhere on S_e .

Remark 5.6. From the commutation relations: $[\nabla_T, \nabla_L] = [\nabla_T, r \nabla_{\underline{L}}] = [\nabla_T, \nabla_A] = 0$, we obtain that $\nabla_T \phi$ satisfies the Fackerell–Ipser equation, if ϕ does. The reasoning which led to the proof of Lemma 5.3 then yields

$$\int_{u_1}^{\infty} \int_{v_1}^{\infty} \int_{\mathbb{S}^2} m_{\text{all}}[\phi](1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C(\mathcal{F}^{\infty}[\phi](u_1,v_1) + \mathcal{F}^{\infty}[\nabla_T \phi](u_1,v_1)).$$
(5.22)

Lemma 5.7 (p-hierarchy). There exists a positive number R_* and a positive constant C such that the following holds. Let ϕ be a smooth solution to the Fackerell–Ipser Equation (3.20) on $\{u \ge u_0\} \cap \{v \ge v_0\}$, let $u_2 \ge u_1 \ge u_0$, $v_2 \ge v_1 \ge v_0$. Let P(u) be $(u, 2R_* + u)$. Defining fluxes as in (5.4), we have that

$$\int_{\overline{C}_{u_2} \cap \{r \ge R\}} (r^2 |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2
+ \int_{\mathfrak{D}_{u_1}^{u_2}} \left[r |\nabla_L \phi|^2 + |\nabla\phi|^2 + r^{-2} |\phi|^2 \right] (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C \left(\mathcal{F}^{\infty}[\phi](P(u_1)) \right) + \int_{\overline{C}_{u_1} \cap \{r \ge R\}} (r^2 |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2.$$
(5.23)

Also,

$$\int_{\overline{C}_{u_2} \cap \{r \ge R\}} (r |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2
+ \int_{\mathfrak{D}_{u_1}^{u_2}} \left[|\nabla_L \phi|^2 + |\nabla \phi|^2 + r^{-2} |\phi|^2 \right] (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C \left(\mathcal{F}^{\infty}[\phi](P(u_1))\right) + \int_{\overline{C}_{u_1} \cap \{r \ge R\}} (r |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2.$$
(5.24)

Proof. Let $p, k \in \mathbb{R}$. Let us consider the identity, which follows from the Fackerell–Ipser Equation (3.20):

$$\underline{L}\left\{\frac{r^{p}}{(1-\mu)^{k}}|\nabla_{L}\phi|^{2}\right\} + L\left\{\frac{r^{p}}{(1-\mu)^{k-1}}|\nabla\phi|^{2}\right\} + L\left\{V\frac{r^{p}}{(1-\mu)^{k}}|\phi|^{2}\right\} - \underline{L}\left\{\frac{r^{p}}{(1-\mu)^{k}}\right\}|\nabla_{L}\phi|^{2} - L\left\{V\frac{r^{p}}{(1-\mu)^{k}}\right\}|\phi|^{2} + \left[(2-p)r^{p-1}(1-\mu)^{2-k} + r^{p}(k-1)(1-\mu)^{-k+1}\frac{2M}{r^{2}}\right]|\nabla\phi|^{2} \stackrel{\mathbb{S}^{2}}{=} 0. \quad (5.25)$$

Here, $\stackrel{\mathbb{S}^2}{=}$ denotes equality after integration on \mathbb{S}^2 with respect to the volume form $d\mathbb{S}^2$. We fix $R_* > 0$ big enough so that the following holds in the region $r_* \geq R_*$:

$$-\underline{L}\left(\frac{r^p}{(1-\mu)^k}\right) \ge \frac{1}{2}r^{p-1},\tag{5.26}$$

when p is either 1 or 2 and k is either 1 or 2.

Let us furthermore calculate

$$-L\left(\frac{Vr^{p}}{(1-\mu)^{k}}\right) = (1-\mu)^{2-k}(2-p)r^{p-3} - 2Mr^{p-4}(1-k)(1-\mu)^{-k+1}.$$
 (5.27)

Recall the definition of \mathfrak{D} :

$$\mathfrak{D}_{u_1}^{u_2} := \{ r \ge R, u \in [u_1, u_2] \}.$$

By possibly choosing R bigger, we assume that R - M > 3M. Let f be a smooth radial function such that

$$f(r) := \begin{cases} 1 \text{ for } r \ge R, \\ 0 \text{ for } r \le R - M. \end{cases}$$

We then integrate the expression

$$\underline{L}\left\{\frac{r^{p}}{(1-\mu)^{k}}f|\nabla_{L}\phi|^{2}\right\} + L\left\{\frac{r^{p}}{(1-\mu)^{k-1}}f|\nabla\phi|^{2}\right\} + L\left\{V\frac{r^{p}}{(1-\mu)^{k}}f|\phi|^{2}\right\} - \underline{L}\left\{\frac{r^{p}}{(1-\mu)^{k}}f\right\}|\nabla_{L}\phi|^{2} - L\left\{V\frac{r^{p}}{(1-\mu)^{k}}f\right\}|\phi|^{2}$$
(5.28)

on the region $J^+(S_{u_1,v_1})$, with $v_1 = 2R_* + u_1$.

We choose p = 2, k = 2, and we use the Morawetz estimate (5.8) in order to bound the spacetime error term in the strip $r \in [R - M, R]$. We obtain

$$\begin{split} &\int_{\overline{C}_{u_2}\cap\{r\geq R\}} (r^2 |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2 \\ &\quad + \int_{\mathfrak{D}_{u_1}^{u_2}} \left[r |\nabla_L \phi|^2 + |\nabla\phi|^2 + r^{-2} |\phi|^2 \right] (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 \\ &\leq C \mathcal{F}^{\infty}[\phi](P(u_1)) + \int_{\overline{C}_{u_1}\cap\{r\geq R\}} (r^2 |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2. \end{split}$$
(5.29)

Choosing p = 1, k = 1, we obtain, instead

$$\int_{\overline{C}_{u_2} \cap \{r \ge R\}} (r |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2
+ \int_{\mathfrak{D}_{u_1}^{u_2}} \left[|\nabla_L \phi|^2 + |\nabla\phi|^2 + r^{-2} |\phi|^2 \right] (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C \mathcal{F}^{\infty}[\phi](P(u_1)) + \int_{\overline{C}_{u_1} \cap \{r \ge R\}} (r |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2.$$
(5.30)

This proves the lemma.

c

5.5. Application of the r^p Method: Decay of Null Fluxes

We now apply the r^p -method of Dafermos and Rodnianski to prove integrated decay for ϕ . We prove the following lemma.

Lemma 5.8. There exists a positive number R_* and a positive constant C such that the following holds. Let ϕ be a smooth solution to the Fackerell–Ipser Equation (3.20) on $\{u \ge u_0\} \cap \{v \ge v_0\}$, let $u \ge u_0$. Let P(u) be $(u, 2R_* + u)$. We have the decay of the flux:

$$\mathcal{F}^{\infty}[\phi](P(u)) \le Cu^{-2} \|\phi\|_{C_{u_0,v_0};2;2,0}.$$
(5.31)

Here, we used the definition of norm in (3.7).

Proof. We define the sequence $u_n := 2^n(|u_0| + M)$. Inequality (5.23) now yields:

$$\int_{\overline{C}_{u_n} \cap \{r \ge R\}} r^2 |\nabla_L \phi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C \left(\mathcal{F}^{\infty}[\phi](P(u_0))\right) + \int_{\overline{C}_{u_0} \cap \{r \ge R\}} (r^2 |\nabla_L \phi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2.$$
(5.32)

Also,

$$(u_{n+1} - u_n) \int_{\overline{C}_{\bar{u}_n} \cap \{r \ge R\}} r |\nabla_L \phi|^2 dv d\mathbb{S}^2$$

$$\leq \int_{\mathfrak{D}_{u_n}^{u_{n+1}}} r |\nabla_L \phi|^2 (1 - \mu) du dv d\mathbb{S}^2$$

$$\leq C \left(\mathcal{F}^{\infty}[\phi](P(u_0)) \right) + \int_{\overline{C}_{u_n} \cap \{r \ge R\}} (r^2 |\nabla_L \phi|^2) dv d\mathbb{S}^2.$$
(5.33)

This holds for some $\tilde{u}_n \in (u_n, u_{n+1}).$ All in all, we have, for this new sequence $\tilde{u}_n,$

$$\int_{\overline{C}_{\tilde{u}_n} \cap \{r \ge R\}} r |\nabla_L \phi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C \tilde{u}_n^{-1} \mathcal{F}^{\infty}[\phi](P(u_0)) + \tilde{u}_n^{-1} \int_{\overline{C}_{u_0} \cap \{r \ge R\}} r^2 |\nabla_L \phi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2.$$
(5.34)

Recall the definition of the spacetime regions:

$$\begin{aligned} \mathfrak{E}_{v_1}^{v_2} &:= \{ r \le R, v \in [v_1, v_2] \} \,, \\ \mathfrak{F}_{u_1}^{u_2} &:= \mathfrak{D}_{u_1}^{u_2} \cup \mathfrak{E}_{v_1}^{v_2} \,, \text{ such that } v_1 - u_1 = 2R_* \text{ and } v_2 - u_2 = 2R_*. \end{aligned}$$

Recall furthermore the definition of $m[\phi]$, as in (5.20). We employ Eq. (5.24) and the Morawetz estimate (recall: Eq. (5.18) and Definition 5.4), to obtain the following:

$$\int_{\mathfrak{E}_{v_1}^{v_2}} m[\phi](1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 + \int_{\overline{C}_{u_2} \cap \{r \ge R\}} r |\nabla_L \phi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 + \int_{\mathfrak{D}_{u_1}^{u_2}} \left[|\nabla_L \phi|^2 + |\nabla\phi|^2 + r^{-2} |\phi|^2 \right] (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 \leq C \mathcal{F}^{\infty}[\phi](P(u_1)) + \int_{\overline{C}_{u_1} \cap \{r \ge R\}} r |\nabla_L \phi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2.$$
(5.35)

Recall now that $m_{\rm all}[\phi] := m[\phi] + m[\nabla_T \phi]$. Plugging now the previous sequence \tilde{u}_n into the formula (5.35), summing it with estimate (5.22) and using Fubini's theorem, we obtain a second sequence $\tilde{\tilde{u}}_n$ (with $\tilde{\tilde{u}}_n + \tilde{\tilde{v}}_n = 2R_*$) such that

$$\begin{split} \tilde{\tilde{u}}_{n} \int_{\mathbb{S}^{2}} \int_{\underline{C}_{\tilde{\tilde{v}}_{n}} \cap \{r \leq R\}} m_{\mathrm{all}}[\phi] \mathrm{d}u \mathrm{d}\mathbb{S}^{2} \\ &+ \tilde{\tilde{u}}_{n} \int_{\mathbb{S}^{2}} \int_{\overline{C}_{\tilde{\tilde{u}}_{n}} \cap \{r \geq R\}} \left[|\nabla_{L}\phi|^{2} + |\nabla\phi|^{2} + r^{-2}|\phi|^{2} \right] \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &\leq C \int_{\mathfrak{E}_{\tilde{\tilde{v}}_{1}}^{\tilde{v}_{2}}} m_{\mathrm{all}}[\phi] \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &+ C \int_{\mathfrak{D}_{\tilde{u}_{1}}^{\tilde{u}_{2}}} \left[|\nabla_{L}\phi|^{2} + |\nabla\phi|^{2} + r^{-2}|\phi|^{2} \right] (1-\mu) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &\leq C \left(\mathcal{F}^{\infty}[\phi](P(\tilde{u}_{n})) + \mathcal{F}^{\infty}[\nabla_{T}\phi](P(\tilde{u}_{n})) \right) \\ &+ \int_{\overline{C}_{\tilde{u}_{n}} \cap \{r \geq R\}} r |\nabla_{L}\phi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &\leq C \left(\mathcal{F}^{\infty}[\phi](P(\tilde{u}_{n})) + \mathcal{F}^{\infty}[\nabla_{T}\phi](P(\tilde{u}_{n})) \right) + C\tilde{u}_{n}^{-1} \mathcal{F}^{\infty}[\phi](P(u_{0})) \\ &+ \tilde{u}_{n}^{-1} \int_{\overline{C}_{u_{0}}} r^{2} |\nabla_{L}\phi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2}. \end{split}$$

$$(5.36)$$

We commute the Fackerell–Ipser Equation (3.20) with ∇_T , and we see that

$$\mathcal{F}^{\infty}[\nabla_T \phi](P(u))$$

decays, by the same reasoning for the decay of $\mathcal{F}^{\infty}[\phi](P(u))$. We use the monotonicity of energy given in (5.17) to eliminate the restriction to the dyadic sequence. We finally obtain the bound

$$\begin{aligned} \mathcal{F}^{\infty}[\phi](P(u)) &\leq Cu^{-2} \sum_{i=0}^{2} \mathcal{F}^{\infty}[(\nabla_{T})^{i}\phi](P(u_{0})) \\ &+ \sum_{i=0}^{2} \int_{\overline{C}_{u_{0}} \cap \{r \geq R\}} r^{2} |\nabla_{L}(\nabla_{T})^{\min\{i,1\}}\phi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2}. \end{aligned}$$
the claim.

This is the claim.

Corollary 5.9. Assume the hypotheses of Lemma 5.8. Consider the Lie derivative \mathcal{L} induced from the connection ∇ . Recall the definition of the set $\Omega :=$ $\{\Omega_1, \Omega_2, \Omega_3\}$. The following estimate holds, for $j \in \{1, 2, 3\}$:

$$\mathcal{F}^{\infty}[\mathcal{L}_{\Omega_{j}}\phi](P(u)) \le Cu^{-2} \|\phi\|_{C_{u_{0},v_{0}};2;2,1}.$$
(5.37)

Here, we used the definition of norm in (3.7).

Proof. Let $\eta \in \Gamma(\mathcal{B}^*)$. An easy calculation implies that we have the commutation relations:

$$[\mathcal{L}_{\Omega_i}, \nabla_L]\eta = 0, \quad [\mathcal{L}_{\Omega_i}, \nabla_{\underline{L}}]\eta = 0.$$

Furthermore, since the Ω_i 's are Killing vectors for the induced metric $\not q$ on the spheres, by Equation 3.25 in [8], we have that

$$[\mathcal{L}_{\Omega_i}, \nabla] = 0.$$

We therefore obtain that $\mathcal{L}_{\Omega_i}\phi$ also satisfies the Fackerell–Ipser equation (3.20). The proof then proceeds as in Lemma 5.8.

5.6. The r^p -Method Revisited

We will now derive decay for a null flux, crucial for the decay of the extreme components.

Lemma 5.10. There exists a positive number R_* and a positive constant C such that the following holds. Let ϕ be a smooth solution to the Fackerell–Ipser Equation (3.20) on $\{u \ge u_0\} \cap \{v \ge v_0\}$. We define the following flux:

$$\mathcal{F}_{p}^{\sim}[\eta](\tilde{u},\tilde{v}) := \int_{\underline{C}_{\tilde{v}} \cap \{r \ge R\} \cap \{u \ge \tilde{u}\}} |\eta|^{2} (u,\tilde{v},\omega) (r(u,\tilde{v}))^{p} \mathrm{d}u \mathrm{d}\mathbb{S}^{2}.$$
(5.38)

Let $u \ge u_0$, $v \ge v_0$. We have the decay estimates:

$$\mathcal{F}_{0}^{\sim}[\phi](u_{0},v) \leq C \|\phi\|_{C_{u_{0},v_{0}};2;0,0}^{2}, \\
\mathcal{F}_{2}^{\sim}[\nabla\phi](u_{0},v) \leq C \|\phi\|_{C_{u_{0},v_{0}};2;0,0}^{2}, \\
\mathcal{F}_{-1}^{\sim}[\phi]\left(\frac{v}{2},v\right) \leq C \frac{\|\phi\|_{C_{u_{0},v_{0}};2;1,0}^{2}}{v}, \\
\mathcal{F}_{1}^{\sim}[\nabla\phi]\left(\frac{v}{2},v\right) \leq C \frac{\|\phi\|_{C_{u_{0},v_{0}};2;1,0}^{2}}{v}, \\$$
(5.39)

$${}^{\sim}_{1}[\nabla \phi]\left(\frac{v}{2}, v\right) \le C \frac{\|\phi\|^{2}_{C_{u_{0}, v_{0}}; 2; 1, 0}}{v},$$
(5.40)



FIGURE 2. Penrose diagram of the spacetime regions

$$\mathcal{F}_{-2}^{\sim}[\phi]\left(\frac{v}{2},v\right) \le C \frac{\|\phi\|_{C_{u_0,v_0};2;2,0}^2}{v^2},$$

$$\mathcal{F}_{0}^{\sim}[\nabla\phi]\left(\frac{v}{2},v\right) \le C \frac{\|\phi\|_{C_{u_0,v_0};2;2,0}^2}{v^2}.$$
 (5.41)

Proof. We choose R such that R - M > 3M. We choose a smooth radial function f with the following requirements:

$$f(r) = \begin{cases} 1 & \text{for } r \ge R, \\ 0 & \text{for } r \le R - M. \end{cases}$$

Let $v_1 \ge v_0$, $u_1 \ge u_0$. We define the spacetime region (Fig. 2):

$$\mathfrak{W}(u_1, R, R_1, v_1) := \{ u \ge u_1 \} \cap \{ r \ge R \} \cap \{ 2(R_1)_* + u_1 \le v \le v_1 \}.$$

To shorten notation, we let $\mathfrak{W}_1 := \mathfrak{W}(u_1, R, R, v_1)$, and $\mathfrak{W}_2 := \mathfrak{W}(u_1, R - M, R, v_1)$. We integrate the expression (5.28) on the region \mathfrak{W}_2 against the form $dudvd\mathbb{S}^2$. We obtain, using (5.25),

$$\begin{split} &\int_{\{v=v_1\}\cap\mathfrak{W}_1} \left\{ \frac{r^p}{(1-\mu)^{k-1}} |\nabla\!\!\!/\phi|^2 + V \frac{r^p}{(1-\mu)^k} |\phi|^2 \right\} \mathrm{d}u\mathrm{d}\mathbb{S}^2 \\ &+ \int_{\mathfrak{W}_1} \left\{ -\underline{L} \left\{ \frac{r^p}{(1-\mu)^k} \right\} |\nabla\!\!\!/_L \phi|^2 - L \left\{ V \frac{r^p}{(1-\mu)^k} \right\} |\phi|^2 \right\} \mathrm{d}u\mathrm{d}v\mathrm{d}\mathbb{S}^2 \\ &+ \int_{\mathfrak{W}_1} \left\{ (2-p)r^{p-1}(1-\mu)^{2-k} \\ &+ r^p(k-1)(1-\mu)^{-k+1} \frac{2M}{r^2} \right\} |\nabla\!\!\!/\phi|^2 \mathrm{d}u\mathrm{d}v\mathrm{d}\mathbb{S}^2 \\ &\lesssim \int_{\{u=u_1\}\cap\mathfrak{W}_1} \left\{ \frac{r^p}{(1-\mu)^k} |\nabla\!\!\!/_L \phi|^2 \right\} \mathrm{d}v\mathrm{d}\mathbb{S}^2 + \mathcal{F}^\infty[\phi](u_1,u_1+2R_*). \quad (5.42) \end{split}$$

Here, we used the Morawetz estimate (5.18) as well as the conservation of energy (5.17) to bound the errors (both boundary and spacetime) arising from the cutoff f (here is where we use that R > 4M not to lose derivatives).

First, recall: $V = \frac{1-\mu}{r^2}$. Also, by possibly increasing R,

$$-\underline{L}\left\{\frac{r^{p}}{(1-\mu)^{k}}\right\} \ge \frac{1}{2}r^{p-1}$$
(5.43)

on \mathfrak{W}_1 , for p = 0, 1, 2.

We now plug p = 2, k = 1 in Eq. (5.42). We obtain that there exists a positive constant K > 0 such that the following holds:

$$\int_{\{v=v_1\}\cap\mathfrak{W}_1} \left\{ r^2 |\nabla\phi|^2 + |\phi|^2 \right\} \mathrm{d}u\mathrm{d}\mathbb{S}^2 + \int_{\mathfrak{W}_1} Kr |\nabla_L\phi|^2 \mathrm{d}u\mathrm{d}v\mathrm{d}\mathbb{S}^2$$
$$\lesssim \int_{\{u=u_1\}\cap\mathfrak{W}_1} \frac{r^2}{(1-\mu)} |\nabla_L\phi|^2 \mathrm{d}v\mathrm{d}\mathbb{S}^2 + \mathcal{F}^{\infty}[\phi](u_1, u_1 + 2R_*). \quad (5.44)$$

In particular, this means that the flux

$$\int_{\{v=v_1\}\cap\mathfrak{W}_1} \left\{ r^2 |\nabla\!\!\!/ \phi|^2 + |\phi|^2 \right\} \mathrm{d} u \mathrm{d} \mathbb{S}^2 \le C \, \|\phi\|_{C_{u_0,v_0};2;0,0}^2$$

Furthermore, using again Eq. (5.42), with p = 1, k = 1, we obtain:

$$\int_{\{v=v_1\}\cap\mathfrak{W}_1} \left\{ r |\nabla\!\!\!/\phi|^2 + r^{-1} |\phi|^2 \right\} \mathrm{d}u \mathrm{d}\mathbb{S}^2$$

$$\lesssim \int_{\{u=u_1\}\cap\mathfrak{W}_1} r |\nabla\!\!\!/_L \phi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 + \mathcal{F}^\infty[\phi](u_1, u_1 + 2R_*). \tag{5.45}$$

Using inequality (5.45) with the choice $u_1 = v/2$, $v_1 = v$, together with (5.36), (5.34), (5.24) and the monotonicity of $\mathcal{F}^{\infty}[\phi]$ (5.17) to eliminate the restriction to the dyadic sequence, we obtain

$$\mathcal{F}_{-1}^{\sim}[\phi](v/2,v) + \mathcal{F}_{1}^{\sim}[\nabla\phi](v/2,v) \le Cv^{-1} \|\phi\|_{C_{u_{0},v_{0}};2;1,0}^{2}.$$
(5.46)

Furthermore, using again Eq. (5.42), with p = 1, k = 1, together with inequality (5.36) commuted once with ∇_T , we obtain similarly

$$\mathcal{F}_{-2}^{\sim}[\phi](v/2,v) + \mathcal{F}_{0}^{\sim}[\nabla\phi](v/2,v) \le Cv^{-2} \|\phi\|_{C_{u_{0},v_{0}};2;2,0}^{2}.$$
(5.47)

This finishes the proof of the lemma.

6. Decay for Solutions to the Spin ± 1 Teukolsky Equations

In this section, we assume that α and $\underline{\alpha}$ are solutions to the spin ± 1 Teukolsky equations, and we prove items (4.1)–(4.3), (4.5) of Theorem 4.1. We postpone the proof of item 4.4 to the next section, which is self-contained and in particular does not depend on the estimates for the incoming fluxes in Lemma 5.10.

Proof of estimates (4.1)-(4.3), (4.5) in Theorem 4.1. We divide the proof into four steps.

Step 1: Estimates on α , Region of Unbounded r

We recall the definition of ϕ :

$$\phi_A = \frac{r^2}{1-\mu} \nabla \underline{V}_{\underline{L}}(r\alpha_A). \tag{6.1}$$

Let $p \in \mathbb{R}$. It follows that

$$\nabla \underline{L}(r^{1+p}\alpha_A) = -p(1-\mu)r^p\alpha_A + (1-\mu)r^{p-2}\phi_A.$$
 (6.2)

Let $2p > \varepsilon > 0$, an application of the Cauchy–Schwarz inequality implies

$$\nabla \underline{L}(r^{2+2p}|\alpha|^2) = 2r^{1+p}\alpha^A(-p(1-\mu)r^p\alpha_A + (1-\mu)r^{p-2}\phi_A)
\leq -(2p-\varepsilon)(1-\mu)r^{1+2p}|\alpha|^2 + \varepsilon^{-1}(1-\mu)r^{2p-3}|\phi|^2.$$
(6.3)

A straightforward calculation implies that $[r\nabla, \nabla_{\underline{L}}] = 0$. Commuting Eq. (6.2) twice with $r\nabla$, we obtain

$$\nabla_{\underline{L}}(r^{3+p}\nabla_{C}\nabla_{B}\alpha_{A}) = -p(1-\mu)r^{p+2}\nabla_{C}\nabla_{B}\alpha_{A} + (1-\mu)r^{p}\nabla_{C}\nabla_{B}\phi_{A}.$$

This implies:

$$\underbrace{\nabla_{\underline{L}}(r^{2+2p}|r^2 \nabla \nabla \alpha|^2)}_{\leq -(2p-\varepsilon)(1-\mu)r^{1+2p}|r^2 \nabla \nabla \alpha|^2 + \varepsilon^{-1}(1-\mu)r^{2p-3}|r^2 \nabla \nabla \phi|^2.$$
(6.4)

We let $\tilde{v} \ge v_0$, $\tilde{u} \ge u_0$, and $\tilde{v} - \tilde{u} = 2(\tilde{r})_*$. Let r_0 such that $\tilde{v} - u_0 = 2(r_0)_*$. We now integrate Eq. (6.3) and (6.4) on $\underline{C}_{\tilde{v}} \cap \{r \ge \tilde{r}\} \cap \{u \ge u_0\}$. We obtain:

$$\int_{\mathbb{S}^{2}} \tilde{r}^{2+2p} (|\tilde{r}^{2} \nabla \nabla \alpha|^{2} (\tilde{u}, \tilde{v}, \omega) + |\alpha|^{2} (\tilde{u}, \tilde{v}, \omega)) d\mathbb{S}^{2} (\omega)
\lesssim r_{0}^{2+2p} \int_{\mathbb{S}^{2}} (|r_{0}^{2} \nabla \nabla \alpha|^{2} (u_{0}, \tilde{v}, \omega) + |\alpha|^{2} (\tilde{u}, \tilde{v}, \omega)) d\mathbb{S}^{2} (\omega)
+ \varepsilon^{-1} \int_{\underline{C}_{\tilde{v}} \cap \{r \geq \tilde{r}\} \cap \{u \geq u_{0}\}} r^{2p-3} (|\phi|^{2} (u, \tilde{v}) + |r^{2} \nabla \nabla \phi|^{2} (u, \tilde{v})) d\mathbb{S}^{2} du. \quad (6.5)$$

Recall the definition of the cutoff χ . It is a smooth function such that $\chi(r) = 1$ for $r \geq 3M$, and $\chi(r) = 0$ for $r \in [2M, 5/2M]$.

Letting $\overline{\Psi} := \chi(r)(1-\mu)^{-1}r^3\alpha$, we trivially bound the initial boundary term by data:

$$|\alpha| \le Cr^{-\frac{5}{2}} \left\| \overline{\Psi} \right\|_{C_{u_0, v_0}; 0; 0, 2}.$$
(6.6)

Furthermore, we bound the integrals of ϕ appearing in the RHS of (6.5) by Lemma 5.10. We first write, using the definition of Lie derivative:

$$\begin{split} &\int_{\underline{C}_{\bar{v}}\cap\{r\geq\tilde{r}\}\cap\{u\geq u_0\}} r^{2p-3}(|\phi|^2(u,\tilde{v})+|r^2\nabla\nabla\phi|^2(u,\tilde{v}))\mathrm{d}\mathbb{S}^2\mathrm{d}u\\ &\leq C\int_{\underline{C}_{\bar{v}}\cap\{r\geq\tilde{r}\}\cap\{u\geq u_0\}} r^{2p-3}(|\phi|^2(u,\tilde{v})+r^2|\nabla\phi|^2(u,\tilde{v}))\mathrm{d}\mathbb{S}^2\mathrm{d}u\\ &+C\int_{\underline{C}_{\bar{v}}\cap\{r\geq\tilde{r}\}\cap\{u\geq u_0\}} r^{2p-1}\sum_{j=1}^3|\nabla\mathcal{L}_{\Omega_j}\phi|^2(u,\tilde{v}))\mathrm{d}\mathbb{S}^2\mathrm{d}u. \end{split}$$

From here, we consider two cases:

• If $\tilde{u} \leq \tilde{v}/2$, estimate (5.39) is sufficient to conclude. Let \tilde{r} such that $\tilde{r}_* = \frac{1}{2}(\tilde{v} - \tilde{u})$. Then,

$$\begin{split} &\int_{\mathbb{S}^{2}} \tilde{r}^{5}(|\tilde{r}^{2}\nabla \nabla \alpha|^{2}(\tilde{u},\tilde{v},\omega) + |\alpha|^{2}(\tilde{u},\tilde{v},\omega))d\mathbb{S}^{2}(\omega) \\ &\lesssim r_{0}^{5}\int_{\mathbb{S}^{2}}(|r_{0}^{2}\nabla \nabla \alpha|^{2}(u_{0},\tilde{v},\omega) + |\alpha|^{2}(\tilde{u},\tilde{v},\omega))d\mathbb{S}^{2}(\omega) \\ &+ \varepsilon^{-1}\int_{\underline{C}_{\tilde{v}}\cap\{r\geq\tilde{r}\}\cap\{u\geq u_{0}\}}(|\phi|^{2}(u,\tilde{v}) + r^{2}|\nabla \phi|^{2}(u,\tilde{v}))d\mathbb{S}^{2}du \\ &+ \varepsilon^{-1}\int_{\underline{C}_{\tilde{v}}\cap\{r\geq\tilde{r}\}\cap\{u\geq u_{0}\}}r^{2}\sum_{j=1}^{3}|\nabla \pounds_{\Omega_{j}}\phi|^{2}(u,\tilde{v})d\mathbb{S}^{2}du \\ &\lesssim \left\|\overline{\Psi}\right\|_{C_{u_{0},v_{0}};0;0,2}^{2} + \mathcal{F}_{0}^{\sim}[\phi](u_{0},\tilde{v}) + \mathcal{F}_{2}^{\sim}[\nabla \phi](u_{0},\tilde{v}) \\ &+ \sum_{i=1}^{3}\mathcal{F}_{2}^{\sim}[\nabla \pounds_{\Omega_{i}}\phi](u_{0},\tilde{v}) \\ &\lesssim \left\|\overline{\Psi}\right\|_{C_{u_{0},v_{0}};0;0,2}^{2} + \left\|\phi\right\|_{C_{u_{0},v_{0}};2;0,0}^{2} + \sum_{i=1}^{3}\left\|\pounds_{\Omega_{i}}\phi\right\|_{C_{u_{0},v_{0}};2;0,0}^{2} \\ &\lesssim \left\|\overline{\Psi}\right\|_{C_{u_{0},v_{0}};0;0,2}^{2} + \left\|\phi\right\|_{C_{u_{0},v_{0}};2;0,1}^{2}. \end{split}$$

Note that, if $\tilde{u} \leq \tilde{v}/2$, then $\tilde{r}_* = \frac{1}{2}(\tilde{v} - \tilde{u}) \geq \frac{1}{4}v$. The Sobolev embedding for one-forms (see Lemma C.2 in "Appendix") then implies the bound

$$|\alpha(u,v,\omega)| \le C(\left\|\overline{\Psi}\right\|_{C_{u_0,v_0};0;0,2} + \left\|\phi\right\|_{C_{u_0,v_0};2;0,1})v^{-\frac{5}{2}},$$

in the region $\{r \ge R_*\} \cap \{u \le v/2\}.$

• If $\tilde{u} \geq \tilde{v}/2$, we consider again Eq. (6.5). Notice that $r_0 \gtrsim (r_0)_* = \frac{1}{2}(\tilde{v} - u_0) \gtrsim \tilde{v}/2$. We estimate

$$\int_{\mathbb{S}^2} \tilde{r}^{2+2p} (|\tilde{r}^2 \nabla \nabla \alpha|^2 (\tilde{u}, \tilde{v}, \omega) + |\alpha|^2 (\tilde{u}, \tilde{v}, \omega)) d\mathbb{S}^2(\omega)$$

$$\lesssim r_0^{2p+2-5} \underbrace{\int_{\mathbb{S}^2} r_0^5 (|r_0^2 \nabla \nabla \alpha|^2 (u_0, \tilde{v}, \omega) + |\alpha|^2 (u_0, \tilde{v}, \omega)) d\mathbb{S}^2(\omega)}_{(i)}$$

$$\underbrace{+ \underbrace{\mathcal{F}_{2p-3}^{\sim}[\phi](\tilde{v}/2,\tilde{v}) + \underbrace{\mathcal{F}_{2p-1}^{\sim}[\nabla \phi](\tilde{v}/2,\tilde{v}) + \sum_{i=1}^{3} \mathcal{F}_{2p-1}^{\sim}[\nabla \pounds_{\Omega_{i}}\phi](\tilde{v}/2,\tilde{v})}_{(ii)}}_{(iii)} \\ + \underbrace{\int_{\underline{C}_{\tilde{v}} \cap \{\tilde{v}/2 \ge u \ge u_{0}\}} r^{2p-3} |\phi|^{2}(u,\tilde{v}) + r^{2p-1} |\nabla \phi|^{2}(u,\tilde{v}) \mathrm{d}\mathbb{S}^{2} \mathrm{d}u}_{(iii)}}_{(iii)} \\ + \underbrace{\int_{\underline{C}_{\tilde{v}} \cap \{\tilde{v}/2 \ge u \ge u_{0}\}} r^{2p-1} \sum_{j=1}^{3} |\nabla \pounds_{\Omega_{j}}\phi|^{2}(u,\tilde{v}) \mathrm{d}\mathbb{S}^{2} \mathrm{d}u}_{(iv)}}_{(iv)}.$$

Now, by inequality (6.6),

$$(i) \lesssim \left\|\overline{\Psi}\right\|_{C_{u_0,v_0};0;0,2}^2.$$

Concerning (ii), we have, by Lemma 5.10, that

$$\begin{aligned} \mathcal{F}_{2p-3}^{\sim}[\phi](\tilde{v}/2,\tilde{v}) + \mathcal{F}_{2p-1}^{\sim}[\nabla\phi](\tilde{v}/2,\tilde{v}) &\lesssim v^{2p-3} \|\phi\|_{C_{u_0,v_0};2;3-2p,0}^2 \,, \\ &\sum_{i=1}^3 \mathcal{F}_{2p-1}^{\sim}[\nabla \not\!\!{L}_{\Omega_i}\phi](\tilde{v}/2,\tilde{v}) \lesssim v^{2p-3} \left\| \not\!\!{L}_{\Omega_i}\phi \right\|_{C_{u_0,v_0};2;3-2p,1}^2 \end{aligned}$$

Concerning (*iii*), we have that the *r*-coordinate of the point of coordinates $(\tilde{v}, \tilde{v}/2)$ is such that $r(\tilde{v}, \tilde{v}/2) \gtrsim \tilde{v}, \tilde{v}/2 = \tilde{v}/2$. Then, if $p \in \{3/2, 1, 1/2\}, 2p-3 \leq 0$, hence

$$\begin{aligned} (iii) &\lesssim (r(\tilde{v}, \tilde{v}/2))^{2p-3} \int_{\underline{C}_{\tilde{v}} \cap \{\tilde{v}/2 \ge u \ge u_0\}} (|\phi|^2(u, \tilde{v}) + r^2 |\nabla \phi|^2(u, \tilde{v})) \mathrm{d}\mathbb{S}^2 \mathrm{d}u \\ &\lesssim (\tilde{v})^{2p-3} \|\phi\|_{C_{u_0, v_0}; 2; 0, 0} \,. \end{aligned}$$

Similarly,

$$(iv) \lesssim (\tilde{v})^{2p-3} \|\phi\|_{C_{u_0,v_0};2;0,1}.$$

Using the Sobolev embedding, Lemma C.2 in "Appendix," setting q = -2p + 3, we finally have

$$|\alpha(u,v,\omega)| \le C(\left\|\overline{\Psi}\right\|_{C_{u_0,v_0};0;0,2} + \left\|\phi\right\|_{C_{u_0,v_0};2;q,1}) v^{-q/2} r^{-3/2}$$

in the region $\{r \ge R_*\} \cap \{u \ge v/2\}$, for $q \in \{0, 1, 2\}$.

Summarizing, we have the claim (4.3). This concludes Step 1.

Step 2: Estimates on α , Region of Bounded r

We let $2M < r_c < R$, $v \ge v_0$ $u \ge u_0$, and we integrate Eqs. (6.3) and (6.4) on

$$\underline{C}_v \cap \{r_c \le r \le R\}.$$

We already know by estimate (6.7) in Step 1 that there exists a constant C depending only on R such that

$$\int_{\mathbb{S}^2} (|\nabla \nabla \alpha|^2(u, v, \omega) + |\alpha|^2(u, v, \omega)) d\mathbb{S}^2$$

$$\leq C_R v^{-2} (\|\overline{\Psi}\|_{C_{u_0, v_0}; 0; 0, 2}^2 + \|\phi\|_{C_{u_0, v_0}; 2; 2, 1}^2).$$

Here, u and v are such that $v - u = 2R_*$. From the estimates on the flux of ϕ (5.31), (5.37), and the Sobolev embedding C.2, we finally obtain

$$|\alpha|(u,v,\omega) \le Cv^{-1}(\left\|\overline{\Psi}\right\|_{C_{u_0,v_0};0;0,2} + \|\phi\|_{C_{u_0,v_0};2;2,1}),\tag{6.8}$$

where $(u, v, \omega) \in \{u \ge u_0\} \cap \{v \ge v_0\} \cap \{r \le R\}.$

Step 3: Estimates on $\underline{\alpha}$, Region of Bounded r

We set

$$\underline{\widetilde{\alpha}} := (1-\mu)^{-1}\underline{\alpha}$$

It follows that

$$\begin{aligned} & \nabla_L ((1-\mu)^{-2} r^2 |\underline{\alpha}|^2) \\ &= L((1-\mu)^{-2}) r^2 |\underline{\alpha}|^2 + 2(1-\mu)^{-2} r \underline{\alpha}^A \nabla_L (r \underline{\alpha}_A) \\ &= -2(1-\mu)^{-2} \frac{2M}{r^2} r^2 |\underline{\alpha}|^2 + 2(1-\mu)^{-2} r \underline{\alpha}^A \nabla_L (r \underline{\alpha}_A). \end{aligned}$$
(6.9)

Recall the definition of ϕ :

This implies:

$$\nabla \nabla_L (r^2 |\underline{\widetilde{\alpha}}|^2) = -4M |\underline{\widetilde{\alpha}}|^2 + 2r^{-1} \underline{\widetilde{\alpha}}^A \underline{\phi}_A.$$
(6.10)

From this, it follows that

Defining now $A^2(u,v) := \int_{\mathbb{S}^2} |\underline{\widetilde{\alpha}}|^2(u,v,\omega) d\mathbb{S}^2(\omega)$, we have the same equation for A^2 :

We integrate this inequality on the interval $[v_0, v]$. We obtain

$$L\left(\exp\left(\int_{v_0}^{v} \frac{4M(1-\varepsilon)}{r(u,v')^2} \mathrm{d}v'\right) r(u,v)^2 A(u,v)^2\right)$$

$$\leq \frac{1}{4Mr(u,v)^2\varepsilon} \exp\left(\int_{v_0}^{v} \frac{4M(1-\varepsilon)}{r(u,v')^2} \mathrm{d}v'\right) \int_{\mathbb{S}^2} |\underline{\phi}|^2 \mathrm{d}\mathbb{S}^2.$$
(6.13)

Consider u as fixed. Since we restrict to the region $\{r \leq R\}$, we have that the function

$$F(u,v) := \int_{v_0}^{v} \frac{4M(1-\varepsilon)}{r(u,v')^2} \mathrm{d}v'$$
(6.14)

is monotonically increasing and satisfies the inequalities

$$\frac{4M(1-\varepsilon)}{R^2}(v-v_0) \le F(u,v) \le M^{-1}(1-\varepsilon)(v-v_0).$$
(6.15)

By integrating inequality (6.13), we obtain that there exists a constant C_R such that

$$\exp(F(v))r(u,v)^{2}A(u,v)^{2} - r(u,v_{0})^{2}A(u,v_{0})^{2}$$

$$\leq C_{R} \int_{\mathbb{S}^{2}} \int_{v_{0}}^{v} (r(u,v'))^{-2} \exp(F(u,v')) |\underline{\phi}|^{2} \mathrm{d}v' \mathrm{d}\mathbb{S}^{2}.$$
(6.16)

We split the integral on the right-hand side into two:

$$\int_{\mathbb{S}^{2}} \int_{v_{0}}^{v} \exp(F(u, v'))(r(u, v'))^{-2} |\underline{\phi}|^{2} dv' d\mathbb{S}^{2}
= \int_{\mathbb{S}^{2}} \int_{v_{0}}^{v/A_{0}} \exp(F(u, v'))(r(u, v'))^{-2} |\underline{\phi}|^{2} dv' d\mathbb{S}^{2}
+ \int_{\mathbb{S}^{2}} \int_{v/A_{0}}^{v} \exp(F(u, v'))(r(u, v'))^{-2} |\underline{\phi}|^{2} dv' d\mathbb{S}^{2}, \quad (6.17)$$

with $A_0 > 1$. We subsequently claim that:

$$\exp(-F(u,v)) \int_{\mathbb{S}^2} \int_{v_0}^{v/A_0} \exp(F(u,v'))(r(u,v'))^{-2} |\underline{\phi}|^2 dv' d\mathbb{S}^2$$

$$\leq C \left\|\underline{\phi}\right\|_{C_{u_0,v_0};0;0,0}^2 v^{-2},$$
(6.18)

and

$$\exp(-F(u,v)) \int_{\mathbb{S}^2} \int_{v/A_0}^{v} \exp(F(u,v')) (r(u,v'))^{-2} |\underline{\phi}|^2 dv' d\mathbb{S}^2$$

$$\leq C \left\| \underline{\phi} \right\|_{C_{u_0,v_0};2;2,1} v^{-2}.$$
 (6.19)

• In order to prove (6.18), we recall the inequalities (6.15), which we write in the following way:

$$A_1v + A_2 \le F(u, v) \le A_3v + A_4, \tag{6.20}$$

with $A_1 < A_3$ positive constants. Choose $A_0 > 0$ so that

$$-A_1 + A_3 / A_0 < 0.$$

Note that A_0 depends only on the value of R. Then,

$$\exp(-F(u,v)) \int_{\mathbb{S}^2} \int_{v_0}^{v/A_0} \exp(F(u,v'))(r(u,v'))^{-2} |\underline{\phi}(u,v')|^2 dv' d\mathbb{S}^2$$

$$\leq \exp(-F(u,v) + F(u,v/A_0)) \int_{\mathbb{S}^2} \int_{v_0}^{\infty} (r(u,v'))^{-2} |\underline{\phi}(u,v')|^2 dv' d\mathbb{S}^2$$

$$\lesssim \exp((-A_1 + A_3/A_0)v) \int_{\mathbb{S}^2} \int_{v_0}^{\infty} (r(u,v'))^{-2} |\underline{\phi}(u,v')|^2 dv' d\mathbb{S}^2. \quad (6.21)$$

We now use the fact that ϕ satisfies the Fackerell–Ipser equation: we have the energy conservation statement (5.17). We obtain the first claim (6.18).

• For the claim (6.19), we first estimate, given that, for fixed u, F(u, v) is non-decreasing in v,

$$\exp(-(F(v))) \int_{\mathbb{S}^2} \int_{v/A_0}^v \exp(F(v'))(r(u,v'))^{-2} |\underline{\phi}|^2 \mathrm{d}v' \mathrm{d}\mathbb{S}^2$$
$$\leq \int_{\mathbb{S}^2} \int_{v/A_0}^v (r(u,v'))^{-2} |\underline{\phi}|^2 \mathrm{d}v' \mathrm{d}\mathbb{S}^2.$$

We consequently notice that, by the energy conservation statement (5.7), the resulting flux satisfies

$$\int_{\mathbb{S}^2} \int_{v/A_0}^v (r(u,v'))^{-2} |\underline{\phi}|^2 \mathrm{d}v' \mathrm{d}\mathbb{S}^2 \le C\mathcal{F}^\infty[\underline{\phi}](Q(v)).$$
(6.22)

where the point $Q(v) = ((A_0)^{-1}v - 2R_*, (A_0)^{-1}v)$. We now use the bound in Lemma 5.8 in order to obtain

$$\int_{\mathbb{S}^2} \int_{v/A_0}^v (r(u,v'))^{-2} |\underline{\phi}|^2 \mathrm{d}v' \mathrm{d}\mathbb{S}^2 \le Cv^{-2} \left\|\underline{\phi}\right\|_{C_{u_0,v_0};2;2,1}^2.$$
(6.23)

This is claim (6.19).

We now use Eq. (6.16), together with the fact that

 $r(u, v_0)^2 A(u, v_0)^2 \le C_R \|\underline{\widetilde{\alpha}}\|_{C_{u_0, v_0}; 0; 0, 0}$

on the region $\{v \ge v_0\} \cap \{r \le R\}$. We obtain finally that A satisfies, in the region $\mathcal{J}^+(C_{u_0,v_0}) \cap \{r \le R\}$,

$$A \le Cv^{-1}(\|\underline{\widetilde{\alpha}}\|_{C_{u_0,v_0};0;0,0} + \|\underline{\phi}\|_{C_{u_0,v_0};2;2,0}).$$

We notice that Eq. (6.11) holds with $\underline{\widetilde{\alpha}}$ replaced by $r^2 \nabla \nabla \underline{\widetilde{\alpha}}$ (this follows by taking the $r \nabla$ derivative of the defining relation of $\underline{\phi}$). Using the Sobolev embedding (Lemma C.2), we obtain the decay for α in the region $\{r \leq R\}$:

$$|\underline{\alpha}| \le C(1-\mu)v^{-1}(\left\|\underline{\phi}\right\|_{C_{u_0,v_0};2;2,1} + \left\|(1-\mu)^{-1}\underline{\alpha}\right\|_{C_{u_0,v_0};0;0,2}).$$
 (6.24)

Step 4: Estimates on $\underline{\alpha}$, Region of Unbounded r

Setting p = 0 in Eq. (6.9), we obtain, using Cauchy–Schwarz on the right:

We integrate this equation on cones of constant u coordinate, starting from $\{r = R\}$. We use the Sobolev embedding C.2 and the estimates (5.31). We finally obtain:

$$|\underline{\alpha}| \le \frac{C}{(|u|+1)r} (\left\|\underline{\phi}\right\|_{C_{u_0,v_0};2;2,1} + \left\|(1-\mu)^{-1}\underline{\alpha}\right\|_{C_{u_0,v_0};0;0,2}),$$
(6.26)

on the region $\{r \ge R\} \cap \mathcal{J}^+(u_0, v_0)$. This concludes the proof of the proposition.

7. Improved Decay for α

In this section, we prove estimate (4.4) of Theorem 4.1. We state again the result, as it is of independent interest. It is essentially an application of the r^p method to the spin +1 Teukolsky equation.

Proposition 7.1. There exists $R_* > 0$ and a constant C > 0 such that the following holds. Let α satisfy the spin +1 Teukolsky equation on $\{u \ge u_0\} \cap \{v \ge v_0\}$, with $v_0 - u_0 = 2R_*$. Let χ be a smooth cutoff function such that $\chi(r) = 1$ for $r \ge 3M$, and $\chi(r) = 0$ for $r \in [2M, 3/2M]$.

We let $\overline{\Psi} := \chi(r)(1-\mu)^{-1}r^3\alpha$. In these conditions, we have the following bound, valid in the region $\{r \geq R\}$:

$$|\alpha| \le C(|u|+1)^{-\frac{1}{2}}r^{-3}(\left\|\overline{\Psi}\right\|_{C_{u_0,v_0};2;0,2} + \|\phi\|_{C_{u_0,v_0};2;2,2}).$$

In order to prove the proposition, we make use of the following lemma. The lemma gives decay estimates on the boundary terms at $\{r = R\}$.

Lemma 7.2. Under the assumptions of Proposition 7.1, we have the following inequality, valid for $u_1 \ge u_0$:

$$\begin{split} &\int_{\overline{C}_{u_{1}}\cap\{r\geq R\}} (r^{2}|\alpha|^{2}+r^{2}|r\nabla \alpha|^{2}) \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &+\int_{\{r=R\}\cap\{u\geq u_{1}\}} (|\nabla \alpha|^{2}+|\alpha|^{2}) \mathrm{d}\mathcal{T} + \int_{\mathfrak{D}_{u_{1}}^{\infty}} (|r\nabla \alpha|^{2}+|\alpha|^{2}) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &\leq (1+|u_{1}|)^{-2} \sum_{i=0}^{2} \mathcal{F}^{\infty}[(\nabla_{T})^{i}\phi](P(u_{0})) \\ &+ (1+|u_{1}|)^{-2} \sum_{i=0}^{2} \int_{\overline{C}_{u_{0}}\cap\{r\geq R\}} r^{2} |\nabla_{L}(\nabla_{T})^{\min\{i,1\}}\phi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &+ (1+|u_{1}|)^{-2} \int_{\overline{C}_{u_{0}}\cap\{r\geq R\}} (r^{4}|\alpha|^{2}+r^{4}|r\nabla \alpha|^{2}) \mathrm{d}v \mathrm{d}\mathbb{S}^{2}. \end{split}$$

$$(7.1)$$

Here, as usual, $d\mathcal{T}$ is the induced volume form on the hypersurface $\{r = R\}$.

Recall now the definition of angular multi-indices and repeated Lie derivative of (3.2). Letting $I \in \iota_{\leq 2}^{\Omega}$, the estimate (7.1) holds as well when all the occurrences of the symbol α in (7.1) are replaced with $\mathcal{L}^{I}\alpha$, and all the occurrences of the symbol ϕ in (7.1) are replaced with $\mathcal{L}^{I}\phi$.

Proof of Lemma 7.2. Let us notice that the quantity ϕ satisfies the following equation, upon substitution in the Teukolsky equation for α (3.17):

$$\nabla _{L}\phi_{A} - r^{2} \Delta (r\alpha_{A}) + r\alpha_{A} = 0.$$
(7.2)

Let f(r) be a smooth radial function. We now multiply Eq. (7.2) by $(1 - \mu)f(r)\phi^A$, and we obtain the identity, valid upon integration on \mathbb{S}^2 :

$$\frac{1}{2}L\left\{(1-\mu)f(r)|\phi|^{2}\right\} - \frac{1}{2}L\left\{(1-\mu)f(r)\right\}|\phi|^{2} + \frac{1}{2}\underline{L}\left\{f(r)r^{2}|r^{2}\nabla\!\!\!/\alpha|^{2}\right\} - \frac{1}{2}\underline{L}\left(f(r)r^{2}\right)|r^{2}\nabla\!\!\!/\alpha|^{2} + \frac{1}{2}\underline{L}\left\{f(r)r^{2}|r\alpha|^{2}\right\} - \frac{1}{2}\underline{L}\left(f(r)r^{2}\right)|r\alpha|^{2} \stackrel{\mathbb{S}^{2}}{=} 0.$$
(7.3)

Let us then integrate with respect to the form dudv in the spacetime region $\{r \geq R\} \cap \mathfrak{D}_{u_1}^{u_2} \cap \{v \leq v_{\max}\}$. We obtain, taking $v_{\max} \to \infty$, and averaging in ϕ ,

$$\int_{\overline{C}_{u_{2}} \cap \{r \ge R\}} f(r)r^{2}(|r\alpha|^{2} + |r^{2}\nabla \alpha|^{2}) dv dS^{2}
- \int_{\overline{C}_{u_{1}} \cap \{r \ge R\}} f(r)r^{2}(|r\alpha|^{2} + |r^{2}\nabla \alpha|^{2}) dv dS^{2}
+ \int_{\{r=R\} \cap \{u_{1} \le u \le u_{2}\}} f(R)R^{2}(|R^{2}\nabla \alpha|^{2} + |R\alpha|^{2}) d\mathcal{T}
- \int_{\mathfrak{D}_{u_{1}}^{u_{2}}} \underline{L}(r^{2}f(r))(|r^{2}\nabla \alpha|^{2} + |r\alpha|^{2}) du dv dS^{2}
\leq C(f, R) \int_{\{R \le r \le R+M\} \cap \{u \ge u_{1}\}} (|\phi|^{2} + |\nabla_{L}\phi|^{2}) du dv dS^{2}
+ \int_{\mathfrak{D}_{u_{1}}^{u_{2}}} L((1-\mu)f(r))|\phi|^{2} du dv dS^{2}.$$
(7.4)

Here, we supposed f to be positive and smooth. $d\mathcal{T}$ is, as before, the induced volume form on the hypersurface $\{r = R\}$.

We choose now $f(r) = (1 - \mu)^{-1}$ in inequality (7.4), and we obtain

$$\int_{\overline{C}_{u_2} \cap \{r \ge R\}} (r^4 |\alpha|^2 + r^4 |r \nabla \alpha|^2) dv d\mathbb{S}^2
+ \int_{\{r=R\} \cap \{u_1 \le u \le u_2\}} (|\nabla \alpha|^2 + |\alpha|^2) d\mathcal{T}
+ \int_{\mathfrak{D}_{u_1}^{u_2}} r(|r^2 \nabla \alpha|^2 + |r\alpha|^2) du dv d\mathbb{S}^2
\leq C \int_{\overline{C}_{u_1} \cap \{r \ge R\}} (r^4 |\alpha|^2 + r^4 |r \nabla \alpha|^2) dv d\mathbb{S}^2 + C\mathcal{F}^{\infty}[\phi](P(u_1)). \quad (7.5)$$

Recall that, here, $P(u) = (u, u + 2R_*)$. This implies that, along a dyadic sequence u_n , we have

$$\int_{\overline{C}_{u_n} \cap \{r \ge R\}} r(|r\alpha|^2 + |r^2 \nabla \alpha|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C u_n^{-1} \int_{\overline{C}_{u_0} \cap \{r \ge R\}} (r^4 |\alpha|^2 + r^4 |r \nabla \alpha|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2 + u_n^{-1} \mathcal{F}^{\infty}[\phi](P(u_0)).$$
(7.6)

We choose now $f(r) = r^{-1}(1-\mu)^{-1}$ in inequality (7.4), and we obtain, discarding the last term in (7.4) (indeed, $L((1-\mu)f) < 0$):

$$\int_{\overline{C}_{u_2} \cap \{r \ge R\}} (r^3 |\alpha|^2 + r^3 |r \nabla \alpha|^2) dv d\mathbb{S}^2
+ \int_{\{r=R\} \cap \{u_1 \le u \le u_2\}} (|\nabla \alpha|^2 + |\alpha|^2) d\mathcal{T}
+ \int_{\mathfrak{D}_{u_1}^{u_2}} (|r^2 \nabla \alpha|^2 + |r\alpha|^2) du dv d\mathbb{S}^2
\leq C \int_{\overline{C}_{u_1} \cap \{r \ge R\}} (r^3 |\alpha|^2 + r^3 |r \nabla \alpha|^2) dv d\mathbb{S}^2
+ C \int_{\{R \le r \le R+M\} \cap \{u \ge u_1\}} (|\phi|^2 + |\nabla_L \phi|^2) du dv d\mathbb{S}^2$$
(7.7)

From inequality (5.36), we now have, along a dyadic sequence \tilde{u}_n , that

$$\tilde{u}_n \mathcal{F}^{\infty}[\phi](P(\tilde{u}_n)) \leq C \underbrace{\left(\mathcal{F}^{\infty}[\phi](P(u_0)) + \mathcal{F}^{\infty}[\nabla_T \phi](P(u_0)) + \int_{\overline{C}_{u_0} \cap \{r \ge R\}} |\nabla_L \phi|^2 r \mathrm{d}v \mathrm{d}\mathbb{S}^2 \right)}_{:=W[\phi](P(u_0))}.$$

By the Morawetz estimate for ϕ , then, we have (without loss of generality, we can assume r > 3M):

$$\int_{\{R \le r \le R+M\} \cap \{u \ge u_1\}} (|\phi|^2 + |\nabla_L \phi|^2) \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 \\
\le (1 + |u_1|)^{-1} W[\phi](P(u_0)).$$
(7.8)

We therefore have the inequality, valid for $u_2 \ge u_n$:

$$\int_{\overline{C}_{u_{2}}\cap\{r\geq R\}} (r^{3}|\alpha|^{2} + r^{3}|r\nabla\alpha|^{2}) dv d\mathbb{S}^{2}
+ \int_{\{r=R\}\cap\{u_{n}\leq u\leq u_{2}\}} (|\nabla\alpha|^{2} + |\alpha|^{2}) d\mathcal{T}
+ \int_{\mathfrak{D}_{u_{n}}^{u_{2}}} (|r^{2}\nabla\alpha|^{2} + |r\alpha|^{2}) du dv d\mathbb{S}^{2}
\leq u_{n}^{-1}W[\phi](P(u_{0})) + Cu_{n}^{-1} \int_{\overline{C}_{u_{0}}\cap\{r\geq R\}} (r^{4}|\alpha|^{2} + r^{4}|r\nabla\alpha|^{2}) dv d\mathbb{S}^{2}.$$
(7.9)

From the last display, in particular, it follows that, for $u_1 \ge u_0$,

$$\begin{split} &\int_{\overline{C}_{u_1}\cap\{r\geq R\}} (r^3|\alpha|^2 + r^3|r\nabla \alpha|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2 + \int_{\{r=R\}\cap\{u\geq u_1\}} (|\nabla \alpha|^2 + |\alpha|^2) \mathrm{d}\mathcal{T} \\ &\leq C u_1^{-1} W[\phi](P(u_0)) + C u_1^{-1} \int_{\overline{C}_{u_0}\cap\{r\geq R\}} (r^4|\alpha|^2 + r^4|r\nabla \alpha|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2. \end{split}$$

From (7.9), it furthermore follows, along a dyadic sequence \bar{u}_n ,

$$\int_{\overline{C}_{\bar{u}_n} \cap \{r \ge R\}} (|r\alpha|^2 + |r^2 \nabla \alpha|^2) dv d\mathbb{S}^2
\leq C \bar{u}_n^{-2} \int_{\overline{C}_{u_0} \cap \{r \ge R\}} (r^4 |\alpha|^2 + r^4 |r \nabla \alpha|^2) dv d\mathbb{S}^2 + \bar{u}_n^{-2} W[\phi](P(u_0)).$$
(7.10)

We choose now $f(r) = r^{-2}(1-\mu)^{-1}$ in inequality (7.4), and we obtain finally

$$\int_{\overline{C}_{u_{2}}\cap\{r\geq R\}} (r^{2}|\alpha|^{2} + r^{2}|r\nabla\alpha|^{2}) dv d\mathbb{S}^{2}
+ \int_{\{r=R\}\cap\{u_{1}\leq u\leq u_{2}\}} (|\nabla\alpha|^{2} + |\alpha|^{2}) d\mathcal{T}
+ \int_{\mathfrak{D}_{u_{1}}^{u_{2}}} (|r\nabla\alpha|^{2} + |\alpha|^{2}) du dv d\mathbb{S}^{2}
\leq \int_{\overline{C}_{u_{1}}\cap\{r\geq R\}} (r^{2}|\alpha|^{2} + r^{2}|r\nabla\alpha|^{2}) dv d\mathbb{S}^{2}
+ \int_{\{R\leq r\leq R+M\}\cap\{u\geq u_{1}\}} (|\phi|^{2} + |\nabla_{L}\phi|^{2}) du dv d\mathbb{S}^{2}.$$
(7.11)

We now use the Morawetz estimate for ϕ , as well as the flux decay in Lemma 5.8, to obtain

It is trivial to remove the restriction to the dyadic sequence, due to the monotonicity of the fluxes considered on the left-hand side of (7.12).

Similarly, it is straightforward to deduce the decay estimate, valid for all $u_1 \ge u_0$:

$$\int_{\overline{C}_{u_1} \cap \{r \ge R\}} (r^2 |\alpha|^2 + r^2 |r \nabla \alpha|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2$$
$$\leq (1 + |u_1|)^{-2} \sum_{i=0}^2 \mathcal{F}^{\infty}[(\nabla_T)^i \phi](P(u_0))$$

$$+ (1+|u_{1}|)^{-2} \sum_{i=0}^{2} \int_{\overline{C}_{u_{0}} \cap \{r \ge R\}} r^{2} |\nabla_{L}(\nabla_{T})^{\min\{i,1\}} \phi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} + (1+|u_{1}|)^{-2} \int_{\overline{C}_{u_{0}} \cap \{r \ge R\}} (r^{4}|\alpha|^{2} + r^{4}|r\nabla\alpha|^{2}) \mathrm{d}v \mathrm{d}\mathbb{S}^{2}.$$
(7.13)

The lemma is thus proved by combining the $\not L^I$ -commuted versions of displays (7.12) and (7.13).

Proof of Proposition 7.1 and of (4.4). Let us consider the Teukolsky equation for α (3.17), and write it in the following way:

$$\begin{aligned} \nabla_{L} \left(\frac{r^{2}}{1-\mu} \nabla_{\underline{L}}(r\alpha_{A}) \right) &- r^{2} \Delta(r\alpha_{A}) + r\alpha_{A} = 0 \\ \iff \nabla_{L} \nabla_{\underline{L}} \left(\frac{r^{2}}{1-\mu} r\alpha_{A} \right) - \nabla_{L} \left(\underline{L} \left(\frac{r^{2}}{1-\mu} \right) r\alpha_{A} \right) \\ &- r^{2} \Delta(r\alpha_{A}) + r\alpha_{A} = 0 \end{aligned}$$

$$(7.14)$$

$$\iff \nabla_{L} \nabla_{\underline{L}} \left(\frac{r^{2}}{1-\mu} r\alpha_{A} \right) - \nabla_{L} \left(\frac{1-\mu}{r^{2}} \underline{L} \left(\frac{r^{2}}{1-\mu} \right) \frac{r^{2}}{1-\mu} r\alpha_{A} \right) \\ &- r^{2} \Delta(r\alpha_{A}) + r\alpha_{A} = 0. \end{aligned}$$

Letting $\Psi_A := (1-\mu)^{-1} r^3 \alpha_A$, we have the following equation for Ψ_A :

which implies

$$\nabla \underline{L} \nabla \underline{V}_L \Psi_A - L \left(\frac{1-\mu}{r^2} \underline{L} \left(\frac{r^2}{1-\mu} \right) \right) \Psi_A - \left(\frac{1-\mu}{r^2} \underline{L} \left(\frac{r^2}{1-\mu} \right) \right) \nabla \underline{V}_L \Psi_A - \left((1-\mu) \underline{A} \Psi_A + \frac{1-\mu}{r^2} \Psi_A = 0. \right)$$
(7.16)

Let f=f(u,v) be a smooth function. Multiply the last display through by $f(u,v) \not \! \nabla_L \Psi^A.$ We obtain

$$\begin{split} &-f(u,v)\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right)\nabla\!\!\!/_L\Psi_A\nabla\!\!\!/_L\Psi^A\\ &=-f(u,v)\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right)|\nabla\!\!\!/_L\Psi|^2,\\ &-(1-\mu)f(u,v)\Delta\!\!\!/\Psi^A\nabla\!\!\!/_L\Psi^A=-f(u,v)\frac{1-\mu}{r^2}r^2\Delta\!\!\!/\Psi^A\nabla\!\!\!/_L\Psi^A\\ &\stackrel{\mathbb{S}^2}{=}f(u,v)\frac{1-\mu}{r^2}\nabla\!\!\!/_L(r\nabla\!\!\!/_B\Psi_A)(r\nabla\!\!\!/^B\Psi^A)\\ &=\frac{1}{2}L\left\{f(u,v)(1-\mu)|\nabla\!\!\!/\Psi|^2\right\}-\frac{1}{2}r^2L\left(f(u,v)\frac{1-\mu}{r^2}\right)|\nabla\!\!\!/\Psi|^2,\\ f(u,v)\frac{1-\mu}{r^2}\Psi^A\nabla\!\!\!/_L\Psi_A=\frac{1}{2}f(u,v)\frac{1-\mu}{r^2}L|\Psi|^2\\ &=\frac{1}{2}L\left(f(u,v)\frac{1-\mu}{r^2}|\Psi|^2\right)-\frac{1}{2}L\left(f(u,v)\frac{1-\mu}{r^2}\right)|\Psi|^2. \end{split}$$

We now proceed to integrate the resulting identity on the region $\mathfrak{D}_{u_1}^{u_2} \cap \{v \leq v_{\max}\}$. We notice the following, from the Poincaré inequality for one-forms (Lemma B.1):

$$\begin{split} &-f(u,v)L\left\{\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right\}|\Psi|^2 + f(u,v)\frac{1-\mu}{r^2}|\Psi|^2\\ &+f(u,v)(1-\mu)|\nabla\Psi|^2\\ &\geq f(u,v)\left\{-L\left\{\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right\} + 2\frac{1-\mu}{r^2}\right\}|\Psi|^2\\ &= 12Mr^{-3}f(u,v)(1-\mu)|\Psi|^2 \ge 0. \end{split}$$

Hence, if f is a positive function, we can discard the corresponding incoming null flux on v_{max} .

We now choose $f(u, v) = (1 - \mu)^{-1}v^2$. It is easy to verify that there exists a value u_{in} such that the following holds, for $u \ge u_{\text{in}}$ and $u + v \ge 2R_*$:

$$-L\left(\frac{v^2}{r^2}\right) \ge 0, \quad -L\left(\frac{v^2}{r^3}\right) \ge 0$$

We now use the previous display, along with the Poincaré estimate for oneforms (Lemma B.1) to obtain positivity of the bulk terms in Ψ , for $u \ge u_{in}$:

$$\begin{split} &\frac{1}{2}L\left\{f(u,v)L\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right)\right\}|\Psi|^2 - \frac{1}{2}L\left(\frac{v^2}{r^2}\right)|\Psi|^2\\ &\quad -\frac{1}{2}L\left(\frac{v^2}{r^2}\right)|\nabla\!\!\!\!|\Psi|^2 \ge \frac{1}{2}L\left\{-2\frac{v^2}{r^2} + f(u,v)L\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right)\right\}|\Psi|^2\\ &\quad = \frac{1}{2}L\left\{-12M\frac{v^2}{r^3}\right\} \ge 0. \end{split}$$

We therefore obtain the following estimate, valid for $u_1, u_2 \ge u_{in}$:

$$\int_{\overline{C}_{u_2} \cap \{r \ge R\}} v^2 |\nabla_L \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 - \int_{\overline{C}_{u_1} \cap \{r \ge R\}} v^2 |\nabla_L \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C \int_{\{r=R\} \cap \{u_1 \le u \le u_2\}} (|\Psi|^2 + |\nabla\Psi|^2) \mathrm{d}\mathcal{T}.$$
(7.17)

Similarly, we have the commuted version of the previous bound:

$$\sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_2} \cap \{r \geq R\}} v^2 |\nabla_L \mathcal{L}^I \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2$$

$$- \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_1} \cap \{r \geq R\}} v^2 |\nabla_L \mathcal{L}^I \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2$$

$$\leq C \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\{r = R\} \cap \{u_1 \leq u \leq u_2\}} (|\mathcal{L}^I \Psi|^2 + |\nabla \mathcal{L}^I \Psi|^2) \mathrm{d}\mathcal{T}.$$
(7.18)

We now choose $f(u,v) = (1-\mu)^{-1}r^2$. We notice that the only spacetime term remaining in either $|\Psi|^2$ or $|\nabla \Psi|^2$ is

$$\int_{\mathfrak{D}_{u_1}^{u_2}} L\left\{f(u,v)L\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right)\right\} |\Psi|^2 \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2.$$
(7.19)

With our choice of f, we have

$$L\left\{f(u,v)L\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right)\right\} = 12M\frac{1-\mu}{r^2}.$$
(7.20)

Hence, we obtain the following estimate, possibly restricting to ${\cal R}$ large enough,

$$\int_{\overline{C}_{u_2} \cap \{r \ge R\}} r^2 |\nabla_L \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 - \int_{\overline{C}_{u_1} \cap \{r \ge R\}} r^2 |\nabla_L \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2
+ \int_{\mathfrak{D}_{u_1}^{u_2}} r |\nabla_L \Psi|^2 \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 + \int_{\mathfrak{D}_{u_1}^{u_2}} r^{-2} |\Psi|^2 \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C_R \int_{\{r=R\} \cap \{u_1 \le u \le u_2\}} (|\Psi|^2 + |\nabla\Psi|^2) \mathrm{d}\mathcal{T}.$$
(7.21)

Similarly, we obtain the commuted estimate

$$\sum_{I \in \iota_{\leq 2}^{\Omega}} \left\{ \int_{\overline{C}_{u_{2}} \cap \{r \geq R\}} r^{2} |\nabla_{L} \mathcal{L}^{I} \Psi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} - \int_{\overline{C}_{u_{1}} \cap \{r \geq R\}} r^{2} |\nabla_{L} \mathcal{L}^{I} \Psi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} + \int_{\mathfrak{D}_{u_{1}}^{u_{2}}} r |\nabla_{L} \mathcal{L}^{I} \Psi|^{2} \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} + \int_{\mathfrak{D}_{u_{1}}^{u_{2}}} r^{-2} |\mathcal{L}^{I} \Psi|^{2} \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \right\}$$

$$\leq C_{R} \sum_{I \in \iota_{\leq 2}} \int_{\{r=R\} \cap \{u_{1} \leq u \leq u_{2}\}} (|\mathcal{L}^{I} \Psi|^{2} + |\nabla \mathcal{L}^{I} \Psi|^{2}) \mathrm{d}\mathcal{T}.$$

$$(7.22)$$

By an analogous reasoning, we obtain the following estimate, choosing $f(r) = (1-\mu)^{-1}r$:

$$\int_{\overline{C}_{u_2}\cap\{r\geq R\}} r|\nabla_L\Psi|^2 \mathrm{d}v\mathrm{d}\mathbb{S}^2 - \int_{\overline{C}_{u_1}\cap\{r\geq R\}} r|\nabla_L\Psi|^2 \mathrm{d}v\mathrm{d}\mathbb{S}^2
+ \int_{\mathfrak{D}_{u_1}^{u_2}} |\nabla_L\Psi|^2 \mathrm{d}u\mathrm{d}v\mathrm{d}\mathbb{S}^2 + \int_{\mathfrak{D}_{u_1}^{u_2}} r^{-3}|\Psi|^2 \mathrm{d}u\mathrm{d}v\mathrm{d}\mathbb{S}^2
\leq C_R \int_{\{r=R\}\cap\{u_1\leq u\leq u_2\}} (|\Psi|^2 + |\nabla\Psi|^2)\mathrm{d}\mathcal{T}.$$
(7.23)

We also obtain the corresponding commuted estimate:

$$\sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_2} \cap \{r \geq R\}} r |\nabla_L \mathcal{L}^I \Psi|^2 dv dS^2$$

$$= \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_1} \cap \{r \geq R\}} r |\nabla_L \mathcal{L}^I \Psi|^2 dv dS^2$$

$$+ \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\mathfrak{D}_{u_1}^{u_2}} |\nabla_L \mathcal{L}^I \Psi|^2 du dv dS^2$$

$$+ \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\mathfrak{D}_{u_1}^{u_2}} r^{-3} |\mathcal{L}^I \Psi|^2 du dv dS^2$$

$$\leq C_R \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\{r=R\} \cap \{u_1 \leq u \leq u_2\}} (|\mathcal{L}^I \Psi|^2 + |\nabla \mathcal{L}^I \Psi|^2) d\mathcal{T}.$$
(7.24)

We now choose $f(u, v) = (1 - \mu)^{-1}$. We look again at the resulting combination of spacetime terms in either $|\Psi|^2$ or $|\nabla \Psi|^2$ and use the Poincaré inequality for one-forms:

$$\begin{split} &\frac{1}{2}L\left\{(1-\mu)^{-1}L\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right) - r^{-2}\right\}|\Psi|^2 - \frac{1}{2}r^2L(r^{-2})|\nabla\!\!\!/\Psi|^2\\ &\geq \frac{1}{2}L\left\{(1-\mu)^{-1}L\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right) - 2r^{-2}\right\}|\Psi|^2. \end{split}$$

We now notice:

$$L\left\{ (1-\mu)^{-1}L\left(\frac{1-\mu}{r^2}\underline{L}\left(\frac{r^2}{1-\mu}\right)\right) - 2r^{-2} \right\} = (1-\mu)\partial_r \left\{ -2r^{-2} - \partial_r \left(\frac{(1-\mu)^2}{r^2}\partial_r \left(\frac{r^2}{1-\mu}\right)\right) \right\} = (1-\mu)\frac{36M}{r^4}.$$

Hence, we obtain the following estimate:

$$\int_{\overline{C}_{u_2}\cap\{r\geq R\}} |\nabla_L \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 - \int_{\overline{C}_{u_1}\cap\{r\geq R\}} |\nabla_L \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2
+ \int_{\mathfrak{D}_{u_1}^{u_2}} r^{-1} |\nabla_L \Psi|^2 \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 + \int_{\mathfrak{D}_{u_1}^{u_2}} r^{-4} |\Psi|^2 \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2
\leq C_R \int_{\{r=R\}\cap\{u_1\leq u\leq u_2\}} (|\Psi|^2 + |\nabla\Psi|^2) \mathrm{d}\mathcal{T}.$$
(7.25)

Also, we obtain the following commuted version of the previous estimate:

$$\sum_{I \in \iota_{\leq 2}^{\Omega}} \left\{ \int_{\overline{C}_{u_2} \cap \{r \geq R\}} |\nabla_L \mathcal{L}^I \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 - \int_{\overline{C}_{u_1} \cap \{r \geq R\}} |\nabla_L \mathcal{L}^I \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 \right. \\ \left. + \int_{\mathfrak{D}_{u_1}^{u_2}} r^{-1} |\nabla_L \mathcal{L}^I \Psi|^2 \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 + \int_{\mathfrak{D}_{u_1}^{u_2}} r^{-4} |\mathcal{L}^I \Psi|^2 \mathrm{d}u \mathrm{d}v \mathrm{d}\mathbb{S}^2 \right\}$$

$$\leq C_R \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\{r=R\} \cap \{u_1 \leq u \leq u_2\}} (|\mathcal{L}^I \Psi|^2 + |\nabla \mathcal{L}^I \Psi|^2) \mathrm{d}\mathcal{T}.$$

$$(7.26)$$

Notice that, in particular, from Lemma 7.2, we have the following estimate:

$$\sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\{u \geq u_0\} \cap \{r = R\}} (|\nabla \mathcal{L}^I \alpha|^2 + |\mathcal{L}^I \alpha|^2) dv d\mathbb{S}^2$$

$$\leq \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_0} \cap \{r \geq R\}} (r^{-4} |\mathcal{L}^I \Psi|^2 + r^{-2} |\nabla \mathcal{L}^I \Psi|^2) dv d\mathbb{S}^2$$

$$+ \sum_{I \in \iota_{\leq 2}^{\Omega}} \mathcal{F}^{\infty} [\mathcal{L}^I \phi] (P(u_0)).$$
(7.27)

Now, using inequalities (7.18), (7.22) and (7.27), we have the following uniform bound for the flux in $r^2 |\nabla_L \Psi|$:

$$\sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u} \cap \{r \geq R\}} v^{2} |\nabla_{L} \not\!\!{L}^{I} \Psi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \leq C(\left\|\overline{\Psi}\right\|_{C_{u_{0},v_{0}};2;0,2}^{2} + \left\|\phi\right\|_{C_{u_{0},v_{0}};0;0,2}^{2}).$$

$$(7.28)$$

Here, we used the definition $\overline{\Psi} := \chi(r)\Psi$, where χ is a smooth cutoff function as in the statement of the proposition, i.e., such that $\chi(r) = 0$ for $r \in [2M, 5/2M]$, and $\chi(r) = 1$ for $r \in [3M, \infty)$.

Now, from (7.22) it follows that there exists a dyadic sequence u_n such that

$$\begin{split} &\sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_n} \cap \{r \geq R\}} \left(r | \nabla_L \mathcal{L}^I \Psi|^2 + r^{-2} | \mathcal{L}^I \Psi|^2 \right) \mathrm{d}v \mathrm{d}\mathbb{S}^2 \\ &\leq C u_n^{-1} \left\{ \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_0} \cap \{r \geq R\}} (r^{-4} | \mathcal{L}^I \Psi|^2 + r^{-2} | \nabla \mathcal{L}^I \Psi|^2) \mathrm{d}v \mathrm{d}\mathbb{S}^2 \right. \\ &+ \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_0} \cap \{r \geq R\}} r^2 | \nabla_L \mathcal{L}^I \Psi|^2 \mathrm{d}v \mathrm{d}\mathbb{S}^2 + \sum_{I \in \iota_{\leq 2}^{\Omega}} \mathcal{F}^{\infty}[\mathcal{L}^I \phi](P(u_0)) \right\}. \end{split}$$
(7.29)

Recall that, from inequality (7.5), the following bound holds for all $u \ge u_0$:

$$\sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u} \cap \{r \geq R\}} (|\nabla \not{\mathcal{L}}^{I} \Psi|^{2} + r^{-2} | \not{\mathcal{L}}^{I} \Psi|^{2}) \mathrm{d}v \mathrm{d}\mathbb{S}^{2}$$

$$\leq C(1 + |u|)^{-1} \sum_{I \in \iota_{\leq 2}^{\Omega}} \left(W[\not{\mathcal{L}}^{I} \phi](P(u_{0})) + \int_{\overline{C}_{u_{0}} \cap \{r \geq R\}} (r^{-2} | \not{\mathcal{L}}^{I} \Psi|^{2} + |\nabla \not{\mathcal{L}}^{I} \Psi|^{2} + r^{2} |\nabla_{L} \not{\mathcal{L}}^{I} \Psi|^{2}) \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \right).$$

$$(7.30)$$

We plug the sequence $\{u_n\}$ in estimate (7.24) and use (7.30) to bound the terms on the right-hand side, and we obtain that there exists a second dyadic sequence $\{\bar{u}_n\}$ such that there holds

$$\sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{\bar{u}_n} \cap \{r \geq R\}} \left(|\nabla_L \mathcal{L}^I \Psi|^2 + r^{-3} |\mathcal{L}^I \Psi|^2 \right) dv dS^2$$

$$\leq C(|\bar{u}_n| + 1)^{-2} \left\{ \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_0} \cap \{r \geq R\}} (r^{-4} |\mathcal{L}^I \Psi|^2 + r^{-2} |\nabla \mathcal{L}^I \Psi|^2) dv dS^2$$

$$+ \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{u_0} \cap \{r \geq R\}} r^2 |\nabla_L \mathcal{L}^I \Psi|^2 dv dS^2 + \sum_{I \in \iota_{\leq 2}^{\Omega}} \mathcal{F}^{\infty}[\mathcal{L}^I \phi](P(u_0))$$

$$+ \sum_{I \in \iota_{\leq 2}^{\Omega}} W[\mathcal{L}^I \phi](P(u_0)) \right\}.$$
(7.31)

We now wish to remove the restriction to the dyadic sequence on the integral

$$\int_{\overline{C}_{\bar{u}_n} \cap \{r \ge R\}} \left(|\nabla_L \mathcal{L}^I \Psi|^2 + r^{-3} |\mathcal{L}^I \Psi|^2 \right) \mathrm{d}v \mathrm{d}\mathbb{S}^2.$$

Concerning the term in $\nabla_L \Psi$, we have that, from inequality (7.31), (7.25) and (7.1), the following bound holds for all $u \ge u_0$:

$$\begin{split} &\int_{\overline{C}_{u}\cap\{r\geq R\}} |\nabla_{L}\Psi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &\leq (1+|u|)^{-2} \sum_{i=0}^{2} \mathcal{F}^{\infty}[(\nabla_{T})^{i}\phi](P(u_{0})) \\ &+ (1+|u|)^{-2} \sum_{i=0}^{2} \int_{\overline{C}_{u_{0}}\cap\{r\geq R\}} r^{2} |\nabla_{L}(\nabla_{T})^{\min\{i,1\}}\phi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &+ C(1+|u|)^{-2} \int_{\overline{C}_{u_{0}}\cap\{r\geq R\}} (r^{-2}|\Psi|^{2} + |\nabla\Psi|^{2} + r^{2}|\nabla_{L}\Psi|^{2}) \mathrm{d}v \mathrm{d}\mathbb{S}^{2}. \end{split}$$

$$(7.32)$$

Similarly, considering the corresponding commuted estimates, if $I \in \iota_{\leq 2}^{\Omega}$, we have, for $u \geq u_0$,

$$\int_{\overline{C}_{u} \cap \{r \geq R\}} |\nabla_{L} \mathcal{L}^{I} \Psi|^{2} dv d\mathbb{S}^{2}
\leq C(|u|+1)^{-2} (||\Psi||^{2}_{C_{u_{0},v_{0}};2;0,2} + ||\phi||^{2}_{C_{u_{0},v_{0}};2;2,2}).$$
(7.33)

Now, from the commuted version of (7.1), we obtain, if $I \in \iota_{\leq 2}^{\Omega}$ and $u \geq u_0$:

$$\int_{\overline{C}_{u} \cap \{r \geq R\}} (r^{-4} | \mathcal{L}^{I} \Psi |^{2} + r^{-2} | \nabla \mathcal{L}^{I} \Psi |^{2}) dv dS^{2}
\leq C(|u|+1)^{-2} (\|\Psi\|^{2}_{C_{u_{0},v_{0}};2;0,2} + \|\phi\|^{2}_{C_{u_{0},v_{0}};2;2,2}).$$
(7.34)

We now have, using Lemma C.5 in "Appendix," letting (u, v, ω) a point in (u, v)-coordinates:

$$|\Psi(u,v,\omega)|^2 \le C(|u|+1)^{-1}(||\Psi||^2_{C_{u_0,v_0};2;0,2} + ||\phi||^2_{C_{u_0,v_0};2;2,2}),$$

if $v - u \ge 2R_*$. This implies the claim.

8. Decay Estimates for σ and ρ

In this section, we suppose that $F \in \Lambda^2(\mathcal{S}_e)$ is a solution to the full Maxwell system, and we prove Theorem 4.2.

Proof of Theorem 4.2. Let $F \in \Lambda^2(\mathcal{S}_e)$ satisfy the Maxwell Equations (3.11) to (3.16). Then, the extreme components α and $\underline{\alpha}$ satisfy the spin ± 1 Teukolsky Equations [(3.17) and (3.18)], and hence we have the required decay rates for α and $\underline{\alpha}$ from Theorem 4.1.

Hence, the proof reduces to proving decay for the middle components σ and ρ . As noticed in Remark 3.8, we have

Hence, it is clear that estimates (5.31) and (5.37) still hold with ϕ replaced by either $r^3 \nabla \phi$ or $r^3 \nabla \sigma$.

Let us now define $\rho_s(u, v)$ and $\sigma_s(u, v)$:

$$\rho_s(u,v) := \frac{R^2}{4\pi} \int_{\omega \in \mathbb{S}^2} \rho(u,v,\omega) \mathrm{d}\mathbb{S}^2(\omega),$$

$$\sigma_s(u,v) := \frac{R^2}{4\pi} \int_{\omega \in \mathbb{S}^2} \sigma(u,v,\omega) \mathrm{d}\mathbb{S}^2(\omega).$$
(8.1)

We notice that, by integrating each of the Maxwell Equations (3.13)–(3.16) on \mathbb{S}^2 , we have, for all $(u, v) \in \{u \ge u_0\} \cap \{v \ge v_0\}$,

$$\sigma_s(u, v) = \sigma_s(u_0, v_0), \quad \rho_s(u, v) = \rho_s(u_0, v_0).$$
(8.2)

Let us restrict our attention to the estimates for ρ . The estimates for σ can be obtained in a very analogous manner.

Step 1: Region of Bounded r

We first consider the region $\{r \leq R\} \cap \mathcal{J}^+(C_{u_0,v_0})$. Estimates (5.31) and (5.37) imply:

$$\mathcal{F}^{\infty}[\nabla\!\!\!/\rho](P(u)) + \sum_{i=1}^{3} \mathcal{F}^{\infty}[\nabla\!\!\!/\mathcal{L}_{\Omega_{i}}\rho](P(u)) \le Cu^{-2}(M_{\rho,\sigma})^{2}.$$
(8.3)

Here, $M_{\rho,\sigma}$ is as in Eq. (4.8).

Let $2M < r_c < R$. We now use the Sobolev lemma in "Appendix C.4" to obtain, if $\tilde{v} \ge v_0$,

$$\sup_{\substack{(u,\tilde{v},\omega)\in \underline{C}_{\tilde{v}}\cap\{r_c\leq r\leq R\}}} |\rho(u,\tilde{v},\omega) - (r(u,\tilde{v}))^{-2}\rho_s(u,\tilde{v})|^2$$
$$\leq C \int_{\underline{C}_{\tilde{v}}\cap\{r_c\leq r\leq R\}} (|\nabla\!\!\!\nabla\!\!\!\nabla\rho|^2 + |(1-\mu)^{-1}\!\!\!\nabla\!\!\!\underline{\nabla}\!\!\!\nabla\rho|^2)(1-\mu)\mathrm{d}\mathbb{S}^2\mathrm{d}u. \quad (8.4)$$

By writing the expression of the Lie derivative, we obtain that there exist positive constants C_0 , C_1 , C_2 such that, in the region $\{r \leq R\}$,

$$\sum_{i=1}^{3} |(1-\mu)^{-1} \nabla \underline{\mathcal{L}} \mathcal{L}_{\Omega_i} \nabla \rho|^2 \ge |(1-\mu)^{-1} \nabla \underline{\mathcal{L}} \nabla \nabla \rho| - C_1 |\nabla \rho|^2 - C_2 |\nabla \nabla \rho|^2 \quad (8.5)$$

This implies that, possibly renaming C_1 and C_2 , the following inequality holds (recall: $[\mathcal{L}_{\Omega_i}, \nabla] = 0$):

$$\int_{\underline{C}_{\tilde{v}} \cap \{r_c \le r \le R\}} (1-\mu) (|\nabla \nabla \rho|^2 + |(1-\mu)^{-1} \nabla \underline{L} \nabla \nabla \rho|^2) d\mathbb{S}^2 du$$
$$\leq C_1 \mathcal{F}^{\infty} [\nabla \rho] (P(\tilde{u})) + C_2 \sum_{i=1}^3 \mathcal{F}^{\infty} [\nabla \mathcal{L}_{\Omega_i} \rho] P(\tilde{u}),$$

with $\tilde{v} - \tilde{u} = 2R_*$. (4.11) now easily follows from the previous display, along with (8.3), in the region $\{u \ge u_0\} \cap \{v \ge v_0\}$.

Step 2: Region of Unbounded r

Let $v \ge v_1 \ge v_0$ and $u = u_1 \ge u_0$, and let $v_1 - u_1 = 2R_*$ and $v - u = 2r_*$. We begin by noticing, by Lemma C.1 and the definition of ϕ :

$$|\rho - r^{-2}\rho_s|^2 \lesssim \int_{\mathbb{S}^2} r^4 |\nabla \nabla \rho|^2 \mathrm{d}\mathbb{S}^2 = \int_{\mathbb{S}^2} r^{-2} |\nabla \phi|^2 \mathrm{d}\mathbb{S}^2.$$
(8.6)

Now, the definition of Lie derivative yields:

$$\mathcal{L}_{\Omega_i}\phi_A = \nabla_{\Omega_i}\phi_A + \phi(\nabla_A\Omega_i).$$

This implies the pointwise bound:

$$|\nabla_{L}\nabla_{\Omega_{i}}\phi|^{2} = \nabla_{L}\nabla_{\Omega_{i}}\phi_{A}\nabla_{L}\nabla_{\Omega_{i}}\phi^{A}$$
$$= \nabla_{L}(\mathcal{L}_{\Omega_{i}}\phi_{A} - \phi(\nabla_{A}\Omega_{i}))\nabla_{L}(\mathcal{L}_{\Omega_{i}}\phi^{A} - \phi(\nabla^{A}\Omega_{i})) \lesssim |\nabla_{L}\mathcal{L}_{\Omega_{i}}\phi|^{2} + |\nabla_{L}\phi|^{2}.$$
(8.7)

Furthermore, we have

$$|\nabla_L \nabla \phi|^2 \lesssim \sum_{i=1}^3 \frac{1}{r^2} |\nabla_L \nabla_{\Omega_i} \phi^2| + \frac{1}{r^2} |\nabla \phi|^2.$$
(8.8)

We now use the Cauchy–Schwarz inequality, as well as the computation in Eq. (8.7) to obtain the following chain of estimates.

$$\begin{split} r \int_{\mathbb{S}^{2}} | \nabla \phi |^{2}(u, v, \omega) \mathrm{d}\mathbb{S}^{2}(\omega) \\ \lesssim \int_{\overline{C}_{u} \cap \{r \geq R\}} | \nabla \phi |^{2}(u, \tilde{v}, \omega) \mathrm{d}\mathbb{S}^{2} \mathrm{d}\tilde{v} \\ &+ \int_{\overline{C}_{u} \cap \{r \geq R\}} r(u, \tilde{v}) \left(| \nabla \phi | | \nabla_{L} \nabla \phi | \right) (u, \tilde{v}, \omega) \mathrm{d}\mathbb{S}^{2}(\omega) \mathrm{d}\tilde{v} \\ \stackrel{(8.8)}{\lesssim} \int_{\overline{C}_{u} \cap \{r \geq R\}} \sum_{i=1}^{3} \left(| \nabla \phi | | \nabla_{L} \nabla_{\Omega_{i}} \phi | \right) \mathrm{d}\mathbb{S}^{2}(\omega) \mathrm{d}\tilde{v} \\ &+ \int_{\overline{C}_{u} \cap \{r \geq R\}} | \nabla \phi |^{2} \mathrm{d}\mathbb{S}^{2}(\omega) \mathrm{d}\tilde{v} \\ \stackrel{(8.7)}{\lesssim} (M_{\rho, \sigma})^{2} (|u| + 1)^{-2} \\ &+ \left(\int_{\overline{C}_{u} \cap \{r \geq R\}} | \nabla \phi |^{2} \mathrm{d}\mathbb{S}^{2} \mathrm{d}\tilde{v} \right)^{\frac{1}{2}} \left(\int_{\overline{C}_{u} \cap \{r \geq R\}} \sum_{i=1}^{3} | \nabla_{L} \mathcal{L}_{\Omega_{i}} \phi |^{2} \mathrm{d}\mathbb{S}^{2} \mathrm{d}\tilde{v} \right)^{\frac{1}{2}} \\ &+ \left(\int_{\overline{C}_{u} \cap \{r \geq R\}} | \nabla \phi |^{2} \mathrm{d}\mathbb{S}^{2} \mathrm{d}\tilde{v} \right)^{\frac{1}{2}} \left(\int_{\overline{C}_{u} \cap \{r \geq R\}} | \nabla_{L} \phi |^{2} \mathrm{d}\mathbb{S}^{2} \mathrm{d}\tilde{v} \right)^{\frac{1}{2}} \\ &\lesssim (M_{\rho, \sigma})^{2} (|u| + 1)^{-2}. \end{split}$$

Remark 8.1. Notice that the flux \mathcal{F}^{∞} allows us to estimate not only the $\nabla _{L} \phi$ term, but also the $\nabla \phi$ term.

This implies finally that

$$|\rho(u, v, \omega) - r^{-2}\rho_s(u_0, v_0)| \le Cr^{-\frac{3}{2}}(|u|+1)^{-1}M_{\rho,\sigma}.$$
(8.9)

Similarly, we compute

$$r^{2} \int_{\mathbb{S}^{2}} |\nabla \phi|^{2}(u, v, \omega) d\mathbb{S}^{2}(\omega) \lesssim \int_{\mathbb{S}^{2}} \sum_{i=1}^{3} \mathscr{G} \left(\nabla_{\Omega_{i}} \phi, \nabla_{\Omega_{i}} \phi \right) (u, v, \omega) d\mathbb{S}^{2}(\omega)$$

$$\lesssim \int_{\overline{C}_{u} \cap \{r \ge R\}} |\nabla \phi|^{2}(u, \tilde{v}, \omega) d\tilde{v} d\mathbb{S}^{2}(\omega) \qquad (8.10)$$

$$+ \int_{\overline{C}_{u} \cap \{r \ge R\}} \nabla_{L} \left(\sum_{i=1}^{3} \mathscr{G} \left(\nabla_{\Omega_{i}} \phi, \nabla_{\Omega_{i}} \phi \right) \right) (u, \tilde{v}, \omega) d\tilde{v} d\mathbb{S}^{2}(\omega).$$

The first term in the right-hand side of the last display is estimated by (5.31). We again use Eq. (8.7) to obtain, for the second term in (8.10), the following chain of estimates.

The last inequality follows by the decay estimates (5.37), (5.31), and the uniform boundedness estimate (5.23) applied to the flux containing the *L*-derivative.

This implies the decay rate for ρ :

$$|\rho(u, v, \omega) - r^{-2}\rho_s(u_0, v_0)| \le C(|u| + 1)^{-\frac{1}{2}}r^{-2}M_{\rho,\sigma}.$$
(8.11)

This concludes the proof of the theorem.

Appendix A. Derivation of the Null Decomposition of the Maxwell System and of the Spin ± 1 Teukolsky Equations

Proof of Proposition 3.6. The proof is by calculation in the null frame and Hodge dualization.

Step 1: The Full Maxwell System

Recall: $L := \partial_{r_{\star}} + \partial_t = \partial_u$, $\underline{L} := \partial_t - \partial_{r_{\star}} = \partial_u$. In the following calculations, when an uppercase letter appears, it signifies contraction with one of the basis elements ∂_{θ^A} , ∂_{θ^B} , where θ^A and θ^B are local coordinates for \mathbb{S}^2 .

Now, set $e_A := \partial_{\theta^A}$ and $e_B := \partial_{\theta^B}$. Let us now calculate

$$\begin{split} \nabla_{e_{A}}e_{B} &= \bigvee_{e_{A}}e_{B} - \frac{1}{2}(1-\mu)^{-1}\left(g(\nabla_{e_{A}}e_{B},\underline{L})L + g(\nabla_{e_{A}}e_{B},L)\underline{L}\right) \\ &= \bigvee_{e_{A}}e_{B} + \frac{1}{2}(1-\mu)^{-1}\left(g(e_{B},\nabla_{e_{A}}\underline{L})L + g(e_{B},\nabla_{e_{A}}L)\underline{L}\right) \\ &= \bigvee_{e_{A}}e_{B} + \frac{1}{2}(1-\mu)^{-1}\left(g(e_{B},-\frac{1-\mu}{r}e_{A})L + g(e_{B},\frac{1-\mu}{r}e_{A})\underline{L}\right) \\ &= \bigvee_{e_{A}}e_{B} + \frac{1}{2r}(\underline{L}-L)\not_{AB}. \end{split}$$
(A.1)

We begin by calculating:

$$\nabla_A F_{BL} = e_A F(e_B, L) - F(\nabla_{e_A} e_B, L) - F(e_B, \nabla_{e_A} L)$$

$$= e_A F(e_B, L) - F(\nabla_{e_A} e_B, L)$$

$$- \frac{1}{2r} F((\underline{L} - L) \not g_{AB}, L) - F(e_B, \frac{1 - \mu}{r} e_A)$$

$$= \nabla_A \alpha_B - \frac{1 - \mu}{r} \rho \not g_{AB} + \frac{1 - \mu}{r} \sigma \not \epsilon_{AB}.$$
(A.2)

Furthermore, we have:

$$\nabla_{A}F_{B\underline{L}} = e_{A}F(e_{B},\underline{L}) - F(\nabla_{e_{A}}e_{B},\underline{L}) - F(e_{B},\nabla_{e_{A}}\underline{L})$$

$$= e_{A}F(e_{B},\underline{L}) - F(\nabla_{e_{A}}e_{B},\underline{L})$$

$$-\frac{1}{2r}F((\underline{L}-L)\not_{AB},\underline{L}) - F(e_{B},-\frac{1-\mu}{r}e_{A})$$

$$= \nabla_{A}\underline{\alpha}_{B} - \frac{1-\mu}{r}\rho\not_{AB} - \frac{1-\mu}{r}\sigma\not_{AB}.$$
(A.3)

By the null decomposition of the Hodge dual of F, it follows that

We also have:

$$\frac{1}{2}\nabla_A F_{\underline{L}L} = \frac{1}{2}\nabla_A (F(\underline{L}, L)) - \frac{1}{2}F(\nabla_A \underline{L}, L) - \frac{1}{2}F(\underline{L}, \nabla_A L)
= \nabla_A \rho + \frac{1}{2}\frac{1-\mu}{r}(\underline{\alpha}_A + \alpha_A).$$
(A.4)

Again, by taking the Hodge dual,

$$\frac{1}{2} \nabla_A {}^*F_{\underline{L}L} = \nabla\!\!\!/_A \rho + \frac{1}{2} \frac{1-\mu}{r} (-\not\!\!/_{CA} \underline{\alpha}^C + \not\!\!/_{CA} \alpha^C)$$

Now, use Eqs. (A.1) and (A.4) to get:

$$\begin{aligned} \nabla_L F_{\underline{L}A} + \nabla_A F_{L\underline{L}} + \nabla_{\underline{L}} F_{AL} &= -\nabla_L F_{A\underline{L}} + \nabla_A F_{L\underline{L}} + \nabla_{\underline{L}} F_{AL} \\ &= -\nabla_L \underline{\alpha}_A + \nabla_{\underline{L}} \alpha_A - 2(1-\mu) \nabla_A \rho - F(\nabla_A L, \underline{L}) - F(L, \nabla_A \underline{L}) \\ &= -\nabla_L \underline{\alpha}_A + \nabla_{\underline{L}} \alpha_A - 2(1-\mu) \nabla_A \rho - \frac{1-\mu}{r} \alpha_A - \frac{1-\mu}{r} \underline{\alpha}_A \\ &= -\frac{1}{r} \nabla_L (r\underline{\alpha}_A) + \frac{1}{r} \nabla_{\underline{L}} (r\alpha_A) - 2(1-\mu) \nabla_A \rho = 0. \end{aligned}$$
(A.5)

By taking the dual of the last equation, we obtain

$$\nabla_L^* F_{\underline{L}A} + \nabla_A^* F_{L\underline{L}} + \nabla_{\underline{L}}^* F_{AL} = \frac{1}{r} \nabla_L (r \phi_{BA} \underline{\alpha}^B) + \frac{1}{r} \nabla_{\underline{L}} (r \phi_{BA} \alpha^B) - 2(1-\mu) \nabla_A \sigma = 0$$
(A.6)

The last display is equivalent to:

We therefore obtain Eqs. (3.11) and (3.12):

$$\frac{1}{r} \nabla _{L}(r \underline{\alpha}_{A}) + (1 - \mu) (\nabla _{A} \rho - \not \epsilon_{AB} \nabla^{B} \sigma) = 0,$$

$$\frac{1}{r} \nabla _{\underline{L}}(r \alpha_{A}) - (1 - \mu) (\nabla _{A} \rho + \not \epsilon_{AB} \nabla^{B} \sigma) = 0.$$
(A.8)

Now, let us calculate, with the aid of (A.3), the following expression:

$$\nabla_{A}F_{B\underline{L}} + \nabla_{\underline{L}}F_{AB} + \nabla_{B}F_{\underline{L}A}$$

$$= \nabla_{A}\underline{\alpha}_{B} - \frac{1-\mu}{r}\rho \phi_{AB} - \frac{1-\mu}{r}\sigma \phi_{AB} +$$

$$- \nabla_{B}\underline{\alpha}_{A} + \frac{1-\mu}{r}\rho \phi_{AB} - \frac{1-\mu}{r}\sigma \phi_{AB} + \phi_{AB}\nabla_{\underline{L}}\sigma = 0.$$
(A.9)

Contracting the last display with \notin^{AB} , we obtain the following equation:

By taking the dual of the last equation, we obtain furthermore

$$-\mathrm{d} \dot{\psi} \ \underline{\alpha} + 2 \frac{1-\mu}{r} \rho - \nabla \underline{\mu} \rho = 0.$$
(A.11)

We finally compute, with the aid of (A.2), the following expression:

$$\nabla_{A}F_{BL} + \nabla_{L}F_{AB} + \nabla_{B}F_{LA}$$

$$= \nabla_{A}\alpha_{B} - \frac{1-\mu}{r}\rho \phi_{AB} + \frac{1-\mu}{r}\sigma \phi_{AB} + \frac{1-\mu}{r}\sigma \phi_{AB} + \frac{1-\mu}{r}\sigma \phi_{AB} + \frac{1-\mu}{r}\sigma \phi_{AB} + \nabla_{L}\sigma$$

$$= \operatorname{cufrl} \alpha + 2\frac{1-\mu}{r}\sigma + \nabla_{L}\sigma = 0.$$
(A.12)

We obtain:

$$\operatorname{curl} \alpha + 2\frac{1-\mu}{r}\sigma + \nabla_L \sigma = 0.$$

Taking the dual of the last expression yields

This concludes the derivation of the null decomposition of the Maxwell system.

Step 2: Spin ± 1 Teukolsky Equations

We now turn to the derivation of the spin ± 1 Teukolsky equations. Recall the following facts, which can be checked by explicit calculation:

Operate now on Eq. (3.11) with \underline{L} , in order to obtain

$$\begin{split} & \nabla_{\underline{L}} \nabla_L (r \underline{\alpha}_A) + \underline{L} (1-\mu) r (\nabla_A \rho - \boldsymbol{\xi}_{AB} \nabla^B \sigma) \\ & + \underbrace{(1-\mu) r (\nabla_A \nabla_{\underline{L}} \rho - \boldsymbol{\xi}_{AB} \nabla^B \nabla_{\underline{L}} \sigma)}_{(*)} = 0. \end{split}$$

Using the expression for $\underline{L}\rho$ (3.14) and $\underline{L}\sigma$ (3.13), we obtain

Hence,

$$\begin{split} & \nabla \underline{k} \nabla_L (r \underline{\alpha}_A) - (1-\mu) \frac{2M}{r^2} r (\nabla_A \rho - \not \epsilon_{AB} \nabla^B \sigma) \\ & + 2(1-\mu)^2 (\nabla_A \rho - \not \epsilon_{AB} \nabla^B \sigma) + (1-\mu) r \cdot \text{ANG} = 0. \end{split}$$

This implies, upon substitution using (3.11) again,

$$\nabla \underline{k} \nabla \underline{k} \nabla L(r \underline{\alpha}_A) - \frac{2}{r} \left(1 - \frac{3M}{r} \right) \nabla L(r \underline{\alpha}_A) - (1 - \mu)r \cdot \text{ANG} = 0.$$

Using now Lemma A.1, we obtain the claim for $\underline{\alpha}$. The reasoning for α is analogous.

Lemma A.1. Let \mathbb{S}^2 be endowed with the standard metric $g_{\mathbb{S}^2}$ and denote by D the Levi-Civita connection associated with such metric on \mathbb{S}^2 . Let ε_{AB} be the standard volume form. Let ω be a smooth one-form on \mathbb{S}^2 , let div and curl the associated covariant divergence and curl:

$$div\omega := g^{AB} D_A \omega_B \quad curl\omega := \varepsilon^{AB} D_A \omega_B.$$

Let $\Delta := g_{\mathbb{S}^2}^{AB} D_A D_B$ be the covariant Laplacian. Then, we have:

$$D_A div\omega - \varepsilon_{AB} D^B curl\omega = \Delta \omega_A - \omega_A.$$
(A.14)

Proof. Let $p \in \mathbb{S}^2$. Let us fix a vector $V \in T_p \mathbb{S}^2$, and let us set up coordinates (θ, φ) such that the coordinates of p are $(\pi/2, 0)$, $\partial_{\theta}|_p = V$, and finally the metric in these local coordinates is represented by the two-form $d\theta \otimes d\theta + \sin^2 \theta d\varphi \otimes d\varphi$. Let

$$\mathcal{T}_A := D_A \operatorname{div} \omega - \varepsilon_{AB} D^B \operatorname{curl} \omega. \tag{A.15}$$

Then, since the left-hand side of (A.14) is a tensor, we have

$$\begin{aligned} \mathcal{T}(\partial_{\theta})|_{p} &= \mathcal{T}(V) = \underbrace{D_{\partial_{\theta}} \operatorname{div} \omega}_{(1)} - \underbrace{\varepsilon_{\theta B} D^{B} \operatorname{curl} \omega}_{(2)}. \\ (1) &= \partial_{\theta} \left(\partial_{\theta} \omega(\partial_{\theta}) - \omega(D_{\partial_{\theta}} \partial_{\theta}) + \frac{\partial_{\varphi}}{\sin \theta} \omega\left(\frac{\partial_{\varphi}}{\sin \theta}\right) - \omega\left(D_{\frac{\partial_{\varphi}}{\sin \theta}} \frac{\partial_{\varphi}}{\sin \theta}\right) \right) \Big|_{(\pi/2,0)} \\ &= \left(\partial_{\theta} \partial_{\theta} \omega(\partial_{\theta}) + \partial_{\theta} \left(\frac{\partial_{\varphi}}{\sin \theta} \omega\left(\frac{\partial_{\varphi}}{\sin \theta}\right)\right) - \omega(\partial_{\theta}) \right) \Big|_{(\pi/2,0)} \\ &- (2) &= -\frac{\partial_{\varphi}}{\sin \theta} \left((D_{\partial_{\theta}} \omega) \left(\frac{\partial_{\varphi}}{\sin \theta}\right) - (D_{\frac{\partial_{\varphi}}{\sin \theta}} \omega) (\partial_{\theta}) \right) \Big|_{(\pi/2,0)} \\ &= -\frac{\partial_{\varphi}}{\sin \theta} \left(\partial_{\theta} \omega\left(\frac{\partial_{\varphi}}{\sin \theta}\right) + \frac{\cos \theta}{\sin^{2} \theta} \omega(\partial_{\varphi}) \right) \Big|_{(\pi/2,0)} + \frac{\partial_{\varphi}}{\sin \theta} \frac{\partial_{\varphi}}{\sin \theta} \omega(\partial_{\theta}) \Big|_{(\pi/2,0)}. \end{aligned}$$

Upon summation,

$$\mathcal{T}(\partial_{\theta})|_{p} = \left. \left(\partial_{\theta} \partial_{\theta} \omega(\partial_{\theta}) - \omega(\partial_{\theta}) \right) \right|_{(\pi/2,0)} + \left. \frac{\partial_{\varphi}}{\sin \theta} \frac{\partial_{\varphi}}{\sin \theta} \omega(\partial_{\theta}) \right|_{(\pi/2,0)}$$

A calculation of the covariant Laplacian in coordinates (θ, φ) at $(\pi/2, 0)$ yields the claim.

Appendix B. A Poincaré Lemma

Let us first set some notation for this section.

- Consider the sphere \mathbb{S}^2 , with the standard metric $g_{\mathbb{S}^2}$ and the standard volume form ε .
- Denote by D the Levi-Civita connection on the sphere \mathbb{S}^2 related to the standard metric $g_{\mathbb{S}^2}$.
- Denote by Δ the covariant Laplacian associated with $D, \Delta := g_{\mathbb{S}^2}^{AB} D_A D_B$.

• Denote by Δ_H the Hodge-de-Rham Laplacian:

$$\Delta_H := d\delta + \delta d,$$

where $\delta := -\star d\star$ is the codifferential. Here, if ω is a one-form, the Hodge dual $(\star \omega)$ is still a one-form, defined to satisfy

$$(\star\omega)_A := \varepsilon_{AB}\omega^B$$

• Let our convention on the Riemann tensor be

$$R(X,Y)Z := (D_X D_Y - D_Y D_X - D_{[X,Y]})Z,$$

where X, Y, Z are smooth vector fields. Consequently, we denote

$$\operatorname{Rm}_{ABCD} := g_{\mathbb{S}^2}(R(\partial_{\theta^A}, \partial_{\theta^B})\partial_{\theta^C}, \partial_{\theta^D}).$$

Here, (θ^1, θ^2) is a local coordinate system on \mathbb{S}^2 , and the previous equation defines uniquely the 4-covariant tensor Rm.

- We define the Ricci tensor so that $\operatorname{Ric}_{AB} := \operatorname{Rm}_{AC}{}^{C}{}_{B}$.
- Under these conventions, if ω_A is a one-form, the commutation relation holds:

$$D_A D_B \omega_C - D_B D_A \omega_C = -\operatorname{Rm}_{ABC}{}^D \omega_D.$$

The main purpose of this section is to give a proof of the following elementary inequality.

Lemma B.1 (Poincaré inequality for one-forms on \mathbb{S}^2). Let ω be a smooth one-form on \mathbb{S}^2 . We have the inequality

$$\int_{\mathbb{S}^2} |D\omega|^2 \mathrm{d}\mathbb{S}^2 \ge \int_{\mathbb{S}^2} |\omega|^2 \mathrm{d}\mathbb{S}^2.$$
(B.1)

We first need a classical result.

Lemma B.2 (Hodge). Let ω be a smooth one-form on \mathbb{S}^2 , then there exist smooth functions f and $g \in \mathcal{C}^{\infty}(\mathbb{S}^2)$ such that

 $\omega = df + \star dg.$

We deduce a simple case of the Bochner–Weitzenböck identity.

Lemma B.3. For any one-form ω on the sphere \mathbb{S}^2 , we have

$$\int_{\mathbb{S}^2} |D\omega|^2 d\mathbb{S}^2 = \int_{\mathbb{S}^2} (\Delta_H \omega)_A \omega^A d\mathbb{S}^2 - \int_{\mathbb{S}^2} |\omega|^2 d\mathbb{S}^2.$$
(B.2)

Proof. We compute

$$g^{AB}D_AD_BD_Cf = g^{AB}D_AD_CD_Bf = g^{AB}D_CD_AD_Bf - g^{AB}\operatorname{Rm}_{ACB}{}^ED_Ef.$$

Now, on functions, $\Delta_H = -\Delta$, and $\operatorname{Rm}_{AC}^{AE} = -\operatorname{Ric}_C^E$. Therefore, we have, considering Ric as a map from one-forms to one-forms:

$$\Delta(df) = -\Delta_H(df) + \operatorname{Ric}(df),$$

for $f \in \mathcal{C}^{\infty}(\mathbb{S}^2)$. Since $D_A \varepsilon_{BC} = 0$, the same holds for the dual:

$$\Delta(\star dg) = -\Delta_H(\star dg) + \operatorname{Ric}(\star dg),$$

for $g \in \mathcal{C}^{\infty}(\mathbb{S}^2)$.

Consider now a one-form ω , by the Hodge decomposition (Lemma B.2), we then have

$$(\Delta\omega)_A = -(\Delta_H \omega)_A + \operatorname{Ric}_A{}^B \omega_B. \tag{B.3}$$

Now, $\operatorname{Ric} = g_{\mathbb{S}^2}$. We now contract the previous display with ω^A . We integrate by parts on \mathbb{S}^2 in order to obtain the claim.

We then characterize the spectrum of Δ_H on one-forms.

Lemma B.4. The smooth one-form ω is an eigenvector for Δ_H if and only if it is of the form $\omega = df + \star dg$, with f and g smooth eigenfunctions of Δ_H with the same eigenvalue.

Proof. If $\Delta_H f = \lambda f$ and $\Delta_H g = \lambda g$, then

$$\Delta_H(df) = d\delta df + 0 = \lambda df,$$

also

$$\Delta_H(\star dg) = d\delta d(\star dg) + 0 = \lambda \star dg.$$

On the other hand, if ω is a one-form on the sphere, we have, by the Hodge theorem, that there exist f and g such that $\omega = df + \star dg$. Then, imposing the eigenvalue condition, we have

$$\Delta_H(df + \star dg) = \lambda(df + \star dg) \implies d(\Delta_H f - \lambda f) = \star d(\Delta_H g - \lambda g).$$
(B.4)

This implies, since the only harmonic functions on the sphere are the constant functions, $\Delta_H f - \lambda f = c_1$, $\Delta_H g - \lambda g = c_2$. By considering $f_1 := f + c_1/\lambda$ and $g_1 := g + c_2/\lambda$, we obtain two functions f_1 and g_1 in the conditions.

We are now ready to prove Lemma B.1.

Proof of Lemma B.1. By Lemma B.4, the spectrum of Δ_H on \mathbb{S}^2 on one-forms is the same as the spectrum on functions.

Let us now take the Hodge decomposition of ω , $\omega = df + \star dg$, and write f, g in the spherical harmonic decomposition. We obtain sequences f_k, g_k such that

$$f_k \stackrel{L^2(\mathbb{S}^2)}{\to} f, \quad g_k \stackrel{L^2(\mathbb{S}^2)}{\to} g.$$

Since f, g are smooth, and hence in particular belong to $H^2(\mathbb{S}^2)$, the approximation by spherical harmonics is in $H^2(\mathbb{S}^2)$:

$$f_k \stackrel{H^2(\mathbb{S}^2)}{\to} f, \quad g_k \stackrel{H^2(\mathbb{S}^2)}{\to} g.$$

Letting now $\omega_A^{(k)} := df_k + \star dg_k$, we have, from Eq. (B.2), and from the fact that the smallest eigenvalue of Δ_H is 2,

$$\int_{\mathbb{S}^2} |D\omega^{(k)}|^2 \mathrm{d}\mathbb{S}^2 \ge \int_{\mathbb{S}^2} (2|\omega^{(k)}|^2 - |\omega^{(k)}|^2) \mathrm{d}\mathbb{S}^2 = \int_{\mathbb{S}^2} |\omega^{(k)}|^2 \mathrm{d}\mathbb{S}^2$$

Taking the limit $k \to \infty$, and using the continuity of the norm, we obtain the claim.

Appendix C. Sobolev Lemmas

These Sobolev lemmas are used throughout the note. The proofs being standard, we do not provide them.

Lemma C.1 (Sobolev estimate for scalar functions on the sphere). Consider $(\mathbb{S}^2, g_{\mathbb{S}^2})$, the two-sphere with the standard metric, let D be the associated Levi-Civita connection, and let f be a smooth function $f : \mathbb{S}^2 \to \mathbb{R}$. Let $\overline{f} := \frac{1}{4\pi} \int_{\mathbb{S}^2} f d\mathbb{S}^2$ be the spherical average of f. There exists a universal constant C such that

$$\sup_{\mathbb{S}^2} |f - \bar{f}|^2 \le C \int_{\mathbb{S}^2} |DDf|^2 \mathrm{d}\mathbb{S}^2.$$
(C.1)

Lemma C.2. There is a universal constant C such that, for any one-form η on \mathbb{S}^2 , the following inequality holds:

$$\sup_{\mathbb{S}^2} |\eta|^2 \le C \int_{\mathbb{S}^2} (|DD\eta|^2 + |\eta|^2) d\mathbb{S}^2.$$
 (C.2)

Lemma C.3 (1-d Sobolev embedding). Let $(a, b) \subset \mathbb{R}$, with $-\infty < a < b < \infty$. Let $f: (a, b) \to \mathbb{R}$, $f \in W^{1,1}(\mathbb{R})$. Then, there holds:

$$\left\| f(x) - (b-a)^{-1} \int_{a}^{b} f(y) \mathrm{d}y \right\|_{L^{\infty}(\mathbb{R})} \le \|f'\|_{L^{1}(\mathbb{R})}.$$
 (C.3)

A straightforward application of Lemma C.3 yields the following lemma.

Lemma C.4 (Sobolev inequality involving only certain derivatives). Let R > 2M. Let $f : S_e \to \mathbb{R}$ be a smooth function. Let $v_0 \ge 0$. Let \overline{f} be the mean of f over the spheres:

$$\bar{f}(u,v) := \frac{1}{4\pi} \int_{\mathbb{S}^2} f(u,v,\omega) \mathrm{d}\mathbb{S}^2(\omega).$$
 (C.4)

Then, there exists a constant C = C(R) such that, for any $\tilde{v} \ge v_0$,

$$\sup_{\substack{(u,\tilde{v},\omega)\in\underline{C}_{\tilde{v}}\cap\{2M\leq r\leq R\}\\\leq C\int_{\underline{C}_{\tilde{v}}\cap\{2M\leq r\leq R\}}} |f(u,\tilde{v},\omega)-\bar{f}(u,\tilde{v})|^{2}$$

$$\leq C\int_{\underline{C}_{\tilde{v}}\cap\{2M\leq r\leq R\}} (|\nabla\nabla f|^{2}+|(1-\mu)^{-1}\nabla \underline{L}\nabla\nabla f|^{2})(1-\mu)\mathrm{d}\mathbb{S}^{2}\mathrm{d}u$$
(C.5)

Finally, we state the following lemma, which we need in the proof of Proposition 7.1.

Lemma C.5. For any R > 0, there exists a constant C > 0 such that the following holds. Let $\Psi \in \Lambda^1(\mathcal{B})$ be a one-form tangent to the spheres of constant *r*-coordinate. Then, we have

$$\begin{split} \Psi|^{2}(\bar{u},\bar{v},\omega) &\leq C \sum_{I \in \iota_{\leq 2}^{\Omega}} \int_{\overline{C}_{\bar{u}} \cap \{R \leq r \leq 2R\}} |\mathcal{L}^{I}\Psi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \\ &+ \sum_{I \in \iota_{\leq 2}^{\Omega}} \left(\int_{\overline{C}_{\bar{u}} \cap \{R \leq r \leq 2R\}} |\mathcal{L}^{I}\Psi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} + \int_{\overline{C}_{\bar{u}} \cap \{r \geq R\}} |\nabla_{L}\mathcal{L}^{I}\Psi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \right)^{\frac{1}{2}} \\ &\times \left(\int_{\overline{C}_{\bar{u}} \cap \{r \geq R\}} v^{2} |\nabla_{L}\mathcal{L}^{I}\Psi|^{2} \mathrm{d}v \mathrm{d}\mathbb{S}^{2} \right)^{\frac{1}{2}} \end{split}$$

$$(C.6)$$

Here, \bar{u} and \bar{v} are such that

 $\bar{v} - \bar{u} \ge 2R_* = 2(R + 2M\log(R - 2M) - 3M - 2M\log(M)).$

Sketch of proof. The lemma is a straightforward consequence of the Sobolev inequality on spheres (Lemma C.2), the Cauchy–Schwarz inequality, as well as the one-dimensional Hardy inequality:

$$\int_0^\infty x^{-2} (F(x))^2 \mathrm{d}x \le C \int_0^\infty \left(\frac{df}{dx}\right)^2 \mathrm{d}x.$$
 (C.7)

Here, f is any smooth real-valued function on \mathbb{R} , and F is its primitive (in x) which vanishes at 0.

Appendix D. Comparison with Previous Work

In this section, we compare our approach with the work of Blue [4] and with the work of Andersson, Bäckdahl and Blue [1].

We briefly recall the main strategy in [4]. There, the author considers the multiplier estimate arising from the "K-vector field":

$$K := u^2 \partial_u + v^2 \partial_v,$$

on the full Maxwell system (here, as usual, $u = t - r_*$, and $v = t + r_*$). The ensuing estimate displays a negative bulk term close to the photon sphere at r = 3M, comprised solely of angular derivatives of the ρ, σ components. In order to estimate these bad bulk terms, he then considers the *scalar* Fackerell– Ipser equation [Eq. (2.14)] satisfied by the middle components ρ and σ , and proves a Morawetz estimate. The aforementioned K-multiplier estimate, along with Sobolev embeddings, finally implies pointwise decay.

Let us note here that the Morawetz estimate in [4] is obtained directly on the scalar Fackerell–Ipser equation, which displays static solutions. There, the Morawetz bulk is controlled by the initial *conformal Morawetz energy* (the energy arising from the K-multiplier), which is infinite for static solutions. In our case, since we have effectively "commuted with an angular derivative," we are able to prove a Morawetz estimate (5.8) which is controlled by initial data terms at the level of energy with no r-weights, and which furthermore vanish on static solutions. This allows us to include the case of static solutions in a natural way. The main two differences between our approach and that of [4] are the following. Firstly, we consider the quantities ϕ_A and ϕ_A , which do not appear in [4]. Furthermore, the methods to prove decay are different: in our result, we rely on the *new physical space method* of Dafermos–Rodnianski [9] (which is also in the spirit of the work on linearized gravity [11]), whereas in [4] the author makes essential use of the K-multiplier.

The initial assumptions differ between the two approaches: our technique allows us to work separating the Teukolsky equations from the rest of the system. Finally, the decay rates we obtain are analogous to those in [4], with minor differences.

To conclude this section, we compare our findings with those of [1]. In this work, whose sole purpose is to prove integrated local energy decay, the authors obtain the following estimate:

$$\begin{split} &\int_{J^{+}(\{t=t_{1}\})} \frac{(r-3M)^{2}}{r^{3}} (|\beta_{\hat{X}}|^{2} + |\beta_{\hat{Y}}|^{2}) \\ &+ M \frac{M(r-2M)}{r^{3}} |\beta_{\hat{Z}}|^{2} + \frac{M(r-3M)^{2}(r-2M)}{r^{5}} |\beta_{\hat{T}}|^{2} \mathrm{dVol} \\ &\lesssim E(\{t=t_{1}\}) \end{split} \tag{D.1}$$

Here,

$$\begin{split} \beta_{\hat{T}} &= (1-\mu)^{-\frac{1}{2}} \nabla_{\partial_t} (r\rho + ir\sigma), \\ \beta_{\hat{X}} &= \nabla_{\partial_\theta} (\rho + i\sigma), \\ \beta_{\hat{Y}} &= \csc \theta \nabla_{\partial_\phi} (\rho + i\sigma), \\ \beta_{\hat{Z}} &= \frac{1}{r} \partial_r (r^2 (\rho + i\sigma)), \end{split}$$

and

$$E(\{t=t_1\}) := \int_{\{t=t_1\}} (|\beta_{\hat{X}}|^2 + |\beta_{\hat{Y}}|^2 + |\beta_{\hat{Z}}|^2 + |\beta_{\hat{T}}|^2) r^2 \mathrm{d}r \sin\theta \mathrm{d}\theta \mathrm{d}\phi.$$

In this case, a first-order Morawetz estimate is achieved which does not "see" the stationary Coulomb solutions (indeed, the quantity β vanishes for stationary solutions of Maxwell on Schwarzschild).

If we compare the estimate proved in [1] [Eq. (D.1)] with our Morawetz estimate [Eq. (5.8)], it is evident that our estimate lies at the level of two derivatives of the field F, whereas estimate (D.1) contains only one derivative of the field on both sides. Hence, in our approach, we require control on more derivatives in the initial data, to obtain a Morawetz estimate which in turn controls more derivatives of the solution.

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References

- Andersson, L., Bäckdahl, T., Blue, P.: Decay of solutions to the Maxwell equation on the Schwarzschild background. Class. Quantum Gravity 33(8), 085010 (2016)
- [2] Andersson, L., Bäckdahl, T., Blue, P.: Decay of solutions to the Maxwell equation on the Schwarzschild background. Class. Quantum Gravity 33(8), 085010, 20 (2016)
- [3] Bardeen, J.M., Press, W.H.: Radiation fields in the Schwarzschild background. J. Math. Phys. 14, 7–19 (1973)
- [4] Blue, P.: Decay of the Maxwell field on the Schwarzschild manifold. J. Hyperbolic Differ. Equ. 5(4), 807–856 (2008)
- [5] Chandrasekhar, S.: On the equations governing the perturbations of the Schwarzschild black hole. Proc. R. Soc. Lond. Ser. A 343, 289–298 (1975)
- [6] Chandrasekhar, S.: On a transformation of Teukolsky's equation and the electromagnetic perturbations of the Kerr black hole. Proc. R. Soc. Lond. Ser. A 348(1652), 39–55 (1976)
- [7] Christodoulou, D.: On the global initial value problem and the issue of singularities. Class. Quantum Gravity 16(12A), A23–A35 (1999). https://doi.org/10. 1088/0264-9381/16/12A/302
- [8] Christodoulou, D., Klainerman, S.: The Global Nonlinear Stability of the Minkowski Space. Princeton Mathematical Series, vol. 41. Princeton University Press, Princeton (1993)
- [9] Dafermos, M., Rodnianski, I.: A new physical-space approach to decay for the wave equation with applications to black hole spacetimes. In: XVIth International Congress on Mathematical Physics, pp. 421–432. World Scientific Publishing, Hackensack, NJ (2010)
- [10] Dafermos, M., Rodnianski, I.: Lectures on black holes and linear waves. In: Evolution Equations, vol. 17 of Clay Mathematics Proceedings, pp. 97–207. American Mathematical Society, Providence, RI (2013)
- [11] Dafermos, M., Holzegel, G., Rodnianski, I.: The linear stability of the Schwarzschild solution to gravitational perturbations. Preprint arXiv:1601.06467 (2016)
- [12] Dafermos, M., Rodnianski, I., Shlapentokh-Rothman, Y.: Decay for solutions of the wave equation on Kerr exterior spacetimes III: the full subextremal case |a| < M. Ann. Math. **183**(3), 787–913 (2016)
- Ghanem, S.: On uniform decay of the Maxwell fields on black hole space-times. Preprint arXiv:1409.8040 (2014)

- [14] Metcalfe, J., Tataru, D., Tohaneanu, M.: Pointwise decay for the Maxwell field on black hole space-times. Preprint arXiv:1411.3693 (2015)
- [15] Pasqualotto, F.: Decay of the Maxwell field on the Schwarzschild manifold. Master thesis. ETH Zürich (2014)
- [16] Rendall, A.D.: Reduction of the characteristic initial value problem to the Cauchy problem and its applications to the Einstein equations. Proc. R. Soc. Lond. Ser. A 427(1872), 221–239 (1990)
- [17] Sterbenz, J., Tataru, D.: Local energy decay for maxwell fields part i: spherically symmetric black-hole backgrounds. Int. Math. Res. Not. 2015(11), 3298–3342 (2015)
- [18] Teukolsky, S.A.: Perturbations of a rotating black hole. I. Fundamental equations for gravitational, electromagnetic, and neutrino-field perturbations. Astrophys. J. 185, 635–648 (1973)

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