

Chiral flavor violation from extended gauge mediation

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ABSTRACT: Models of extended gauge mediation, in which large A -terms arise through direct messenger-MSSM superpotential couplings, are well-motivated by the discovery of the 125 GeV Higgs. However, since these models are not necessarily MFV, the flavor constraints could be stringent. In this paper, we perform the first detailed and quantitative study of the flavor violation in these models. To facilitate our study, we introduce a new tool called `FormFlavor` for computing precision flavor observables in the general MSSM. We validate `FormFlavor` and our qualitative understanding of the flavor violation in these models by comparing against analytical expressions. Despite being non-MFV, we show that these models are protected against the strongest constraints by a special flavor texture, which we dub chiral flavor violation (χ FV). This results in only mild bounds from current experiments, and exciting prospects for experiments in the near future.

KEYWORDS: Supersymmetric Standard Model, Rare Decays, B-Physics, Kaon Physics

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1 Introduction

The SUSY flavor problem is a serious challenge for models of weak-scale supersymmetry. (For a review and original references, see [1, 2].) Soft SUSY-breaking introduces many sources of flavor violation beyond the Standard Model Yukawas. Generic points of the MSSM parameter space are ruled out by myriad precision flavor constraints, such as neutral

meson mixing and $b \rightarrow s\gamma$. Clearly, the underlying mechanism that generates these soft SUSY-breaking parameters from the hidden sector must be quite special. Historically, the lore has been that the mediation of SUSY-breaking must either be flavor-blind or obey the minimal flavor-violating (MFV) ansatz. Gauge mediated SUSY breaking (GMSB), which is manifestly flavor-blind, is one of the simplest solutions to the SUSY flavor problem (see [3] for a review and original references).

The 2012 discovery of a Standard Model-like Higgs with a mass near 125 GeV [4, 5] presents interesting challenges for models of GMSB, especially in the MSSM, where a 125 GeV Higgs implies either very heavy stops $\gtrsim 10$ TeV or large (multi-TeV) stop A -terms [6–10]. The heavy stop scenario is more fine-tuned and less interesting from both an experimental and theoretical point of view. The large A -term scenario allows for light stops, but a mechanism is required to generate these A -terms, which are absent at the messenger scale in GMSB.

Large A -terms can arise if the usual GMSB framework is extended to also include direct MSSM-messenger superpotential couplings. Using light (~ 1 TeV) stops and large A -terms, these models of extended GMSB (EGMSB) can give rise to the observed Higgs mass at fine-tuning levels close to the best achievable within the MSSM [11–28]. However, since these MSSM-messenger superpotentials are typically flavorful, they are in danger of reintroducing the SUSY flavor problem. Previous works on EGMSB have either assumed perfect alignment with the third-generation (to get a large stop A -term), or considered additional model building (such as Froggatt-Nielsen mechanisms) to ensure this alignment [18, 20, 21, 29–34]. In this paper, we will not presume any such alignment, and we will instead perform the first comprehensive study of the general flavor constraints on EGMSB models for the Higgs mass.

The precursor to this work was the complete classification of all renormalizable EGMSB couplings consistent with perturbative $SU(5)$ unification provided in [20] (see also [18]). By turning on one coupling at a time (perfect alignment with the third generation was assumed) and imposing the Higgs mass constraint, the landscape of EGMSB models was surveyed for their phenomenology and fine-tuning. It was shown that the EGMSB models that exhibit the smallest tuning are of the form,

$$W \ni \kappa_3 Q_3 \Phi \tilde{\Phi}, \quad W \ni \kappa_3 \bar{U}_3 \Phi \tilde{\Phi}, \quad W \ni \kappa_3 Q_3 H_u \Phi, \quad \text{or} \quad W \ni \kappa_3 \bar{U}_3 H_u \Phi. \quad (1.1)$$

Here $\Phi, \tilde{\Phi}$ are messenger fields transforming in appropriate representations of the SM gauge group. The first two (second two) models involve a single MSSM field (two MSSM fields) and so they were classified as “type I” (“type II”) models.

Notice that the least fine-tuned models in (1.1) are all flavorful. Type I Higgs models, i.e., $W \ni \kappa H_u \Phi \tilde{\Phi}$, while MFV, are more finely-tuned due to the “little A/m_H^2 problem” [13]. These models have an irreducible contribution to $m_{H_u}^2 = A_{\tilde{t}}^2 + \dots$ where $A_{\tilde{t}}$ is the stop A -term; thus a large $A_{\tilde{t}}$ is tied to a large $m_{H_u}^2$ and one does not improve tuning in these models by increasing the A -terms.

In this paper, we will study the effects that arise when the couplings of (1.1) are no longer required to align perfectly with the third-generation. For simplicity, we will

specialize to a pair of representative type I Q -class and U -class models, the models I.9 and I.13 from [20],

$$W \ni \kappa_i Q_i \Phi_D \Phi_L, \quad W \ni \kappa_i \bar{U}_i \Phi_{D_1} \Phi_{D_2} \quad (1.2)$$

where the messengers Φ_D and Φ_L have the SM gauge quantum numbers of the D and L matter fields. We have verified explicitly that the other type I squark models are very similar, both qualitatively and quantitatively. For reasons we will explain shortly, we expect the type II models are also qualitatively similar in their features. We will not consider the effect of turning on multiple EGMSB couplings, or couplings involving lepton flavor violation. Such couplings can be forbidden by appropriate choices of discrete symmetries. Finally, to focus exclusively on the SUSY flavor problem, we will not consider CP violation in this work, i.e., all the couplings κ_i are taken to be real in the mass basis of the standard model particles. The (possibly stringent) constraints from CP violating observables, such as ϵ_K and the neutron EDM, will be studied in a forthcoming publication [35].

The interactions in (1.2) result in flavor-violating contributions to the squark mass matrices

$$\delta\mathcal{M}_u^2 = \left(\begin{array}{c|c} \delta m_Q^2 & s_\beta v A_{\bar{u}}^\dagger \\ \hline s_\beta v A_{\bar{u}} & \delta m_U^2 \end{array} \right), \quad \delta\mathcal{M}_d^2 = \left(\begin{array}{c|c} \delta m_Q^2 & c_\beta v A_{\bar{d}}^\dagger \\ \hline c_\beta v A_{\bar{d}} & \delta m_D^2 \end{array} \right) \quad (1.3)$$

Here each block (clockwise from upper left: LL , LR , RR , RL) is a 3×3 matrix. These should be added to the flavor-conserving GMSB contributions and the supersymmetric contributions, and together they contribute to precision flavor observables through a variety of one-loop diagrams involving the squarks and other superpartners. The calculation of these one-loop diagrams for general MSSM spectra is a tedious and laborious task, ideally suited for a computer program. There are several such programs that are publicly available; unfortunately, we found that they were all unsuitable for our purpose. These programs either assumed MFV, did not have a sufficiently broad list of flavor observables, were numerically unstable, or were found to have bugs, likely introduced in transcribing formulas by hand from the literature.

As a result, we found it necessary to develop a new tool called **FormFlavor** for the study of flavor physics. **FormFlavor** takes general MSSM spectra and computes the contributions to the various flavor observables shown in table 1. The novel aspect of **FormFlavor** is that the computation of one-loop Wilson coefficients in the MSSM is done *completely from scratch*, using the general-purpose packages **FeynArts** [36] and **FormCalc** [37].¹ This avoids the problems associated with transcribing formulas from the literature, and it facilitates the inclusion of additional flavor observables in an automated and modular way. We intend to make **FormFlavor** publicly available; its usage and validation will be described in an upcoming publication [39].

Starting from the points of reduced tuning identified in [20], we will turn on $\kappa_{1,2}$ and use **FormFlavor** to investigate the constraints from precision flavor observables. Given that

¹While this work was in preparation, a new tool **FlavorKit** [38] was published with a very similar approach to calculating flavor observables from scratch. **FlavorKit** aims to be even more general, in that it can derive the one-loop Wilson coefficients for a general model, not just the MSSM. It would be interesting to compare the two codes in detail.

Observable	Experiment	SM prediction
Δm_K	$(3.484 \pm 0.006) \times 10^{-15}$ GeV	—
Δm_{B_d}	$(3.36 \pm 0.02) \times 10^{-13}$ GeV	$(3.56 \pm 0.60) \times 10^{-13}$ GeV [40]
Δm_{B_s}	$(1.169 \pm 0.0014) \times 10^{-11}$ GeV	$(1.13 \pm 0.17) \times 10^{-11}$ GeV [40]
Δm_D	$(6.2^{+2.7}_{-2.8}) \times 10^{-15}$ GeV	—
$Br(K^+ \rightarrow \pi^+ \nu \bar{\nu})$	$(1.7 \pm 1.1) \times 10^{-10}$	$(7.8 \pm 0.8) \times 10^{-11}$ [41]
$Br(B \rightarrow X_s \gamma)$	$(3.40 \pm 0.21) \times 10^{-4}$	$(3.15 \pm 0.23) \times 10^{-4}$ [42]
$Br(B \rightarrow X_d \gamma)$	$(1.41 \pm 0.57) \times 10^{-5}$ [43, 44]	$(1.54^{+0.26}_{-0.31}) \times 10^{-5}$ [44]
$Br(B_s \rightarrow \mu^+ \mu^-)$	$(2.9 \pm 0.7) \times 10^{-9}$ [45]	$(3.65 \pm 0.23) \times 10^{-9}$ [46]
$Br(B_d \rightarrow \mu^+ \mu^-)$	$(3.6^{+1.6}_{-1.4}) \times 10^{-10}$ [45]	$(1.06 \pm 0.09) \times 10^{-10}$ [46]

Table 1. Current experimental values based on PDG and HFAG fits [47, 48] except where noted. No reliable theoretical prediction for Δm_D currently exists. Although literature on the subject exists, we do not use a theoretical prediction for Δm_K (see the discussion of Δm_K in section 4.2 for more details).

they are not MFV, one might expect these constraints to be extremely stringent. Generic operator bounds put the scale of flavor violation at 20 PeV or higher from $K - \bar{K}$ and $D - \bar{D}$ mixing [2]. Even in SUSY, where one benefits from loop factors, etc., the bounds are close to 500 TeV (without using CP violation) [49, 50]. However, we will find that in these EGMSB models, the limits from flavor-violation are extremely mild — to the point that $\kappa_{1,2,3}$ can all be the same size and yet the model is not ruled out by flavor!

Much of this paper will be devoted to identifying the reasons for these surprisingly mild flavor constraints. Although the overall heavier mass scale required to raise the Higgs mass plays a role, the most important reason is the fact that the EGMSB models (1.1) only violate flavor through a spurion of either $SU(3)_Q$ or $SU(3)_U$, but not both. Because of this *chiral flavor violation* (χ FV), EGMSB models have a novel flavor texture — flavor violation primarily occurs only in either the left-chiral sector of the squark mass-squared matrix (for the Q -class models) or the right-chiral sector of the squark mass-squared matrix (for the U -class models). Only via communication through the SM Yukawas can the other chiral sector feel the flavor violation. As it turns out, the most stringent flavor constraints, which come from Δm_K and Δm_D , are vastly reduced when the flavor violation is restricted to only the left- or right-chiral sector. (It is also on these general grounds that we expect the type II models are similarly unconstrained by flavor, although it would be interesting to verify this in detail.) In a forthcoming work [51], we will study this new ansatz of χ FV in more generality, along the lines of what has been done for MFV. χ FV represents an interesting intermediate case between full flavor anarchy (which is known to be heavily constrained) and MFV (which is known to be basically unconstrained).

In order to validate and interpret our numerical findings, we will compare them against analytic expressions for the flavor observables. Historically, the mass insertion approximation (MIA) has been utilized to interpret the influence of flavor-violating squark masses on

precision flavor observables (see e.g., [2] for a review of the MIA and original references). However, the utility of the MIA is limited when one or more of these mass insertions are $\mathcal{O}(1)$. Since that is precisely the interesting region of parameter space for our EGMSB models, the traditional MIA cannot be used here. Fortunately, in these EGMSB models there is another handle we can use to obtain an analytic understanding. From two powers of our anti-fundamental spurion, κ_a , we can construct an adjoint+singlet of $SU(3)_Q$ or $SU(3)_U$,

$$K_{ab} \equiv \kappa_a^* \kappa_b \tag{1.4}$$

This matrix K governs all flavor violation from EGMSB. By exploiting the fact that it is only rank 1, together with the special properties of the soft masses for the type I models, we are able to obtain analytic formulas for the flavor observables that treat the flavor violation *exactly*. Then the only expansion that we do is in v/m_{SUSY} , which is still an excellent approximation in these models, even when the flavor violation is $\mathcal{O}(1)$. With this technique, we are able to make precise estimates of the supersymmetric contributions to flavor observables, validate our numerical results in detail, and understand their qualitative features.

Our paper is outlined as follows. In section 2, we describe the idea behind chiral flavor violation and how it applies to our models (1.1) in restricting the texture of the squark mass matrices. We then further restrict our attention to the type I squark models, and highlight some of their special features that will be useful in analyzing the flavor observables. In section 3, we turn to a detailed study of the flavor constraints on the EGMSB models (1.2). We illustrate these constraints in the κ_1 - κ_2 plane using a series of plots of flavor observables computed with `FormFlavor`. We also provide an analytic understanding of the features of these plots using the special properties of rank 1 χ FV and the type I squark models. Finally in section 4, we conclude with a brief summary of our results, and a discussion of the promising future prospects for precision flavor tests of these EGMSB models. In the appendices, we detail our parameter deformation and various subtleties that arise there, provide a brief description of `FormFlavor` (postponing a more detailed manual and validation for an upcoming work [39]), compile the necessary expressions for flavor observables using a uniform notation, and present some other formulas used in this work.

2 Chiral flavor violation and EGMSB

2.1 Chiral flavor violation

In the absence of Yukawa couplings, the Standard Model flavor symmetry group is

$$G_f = SU(3)_Q \times SU(3)_U \times SU(3)_D \times SU(3)_L \times SU(3)_E \tag{2.1}$$

The assumption of MFV is that this flavor symmetry is broken only by the Yukawa couplings y_u , y_d and y_ℓ , which transform as $(\mathbf{3}, \bar{\mathbf{3}})$ under $SU(3)_U \times SU(3)_Q$, $SU(3)_D \times SU(3)_Q$, and $SU(3)_E \times SU(3)_L$, respectively. As described in the introduction, the chiral flavor violation (χ FV) ansatz goes beyond MFV, and postulates that, in addition to the Yukawas, the

flavor symmetry is broken by additional spurions transforming only under a single $SU(3)$ of the full flavor group.

The χ FV ansatz results in a special texture of flavor violation in the MSSM, one which greatly suppresses the constraints from many flavor violating observables. In general, the 3×3 soft mass-squareds m_χ^2 transform as an adjoint+singlet of $SU(3)_X$, while the A -terms transform in the same way as the Yukawas. Using spurions from only a single $SU(3)_X$, one can obtain flavor violation only in m_χ^2 . To obtain flavor violation in any of the other soft terms, one must involve the Yukawa couplings. Thus these other soft masses inherit an MFV-like suppression.

The cases of interest for this paper are when $X = Q$ or U . (We reserve a more general treatment for an upcoming publication [51].) Let us now examine these in more detail. To leading order in the Yukawa couplings, the symmetries of Q -class χ FV imply that

$$\delta m_U^2 = y_u \Sigma y_u^\dagger, \quad \delta m_D^2 = y_d \Sigma y_d^\dagger, \quad A_{\bar{u}} = y_u \Gamma, \quad A_{\bar{d}} = y_d \Gamma \quad (2.2)$$

where Σ and Γ are built out of the $SU(3)_Q$ spurions. Substituting these into (1.3), we have under the third-generation dominant approximation (whereby y_u and y_d are nonzero only in the 33 component):

$$\delta \mathcal{M}_{\bar{u}}^2 \approx \left(\begin{array}{ccc|ccc} & & & 0 & 0 & m_t \Gamma_{31}^* \\ & & & 0 & 0 & m_t \Gamma_{32}^* \\ & & & 0 & 0 & m_t \Gamma_{33}^* \\ \hline 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 \\ m_t \Gamma_{31} & m_t \Gamma_{32} & m_t \Gamma_{33} & 0 & 0 & y_t^2 \Sigma_{33} \end{array} \right) \quad (2.3)$$

$$\delta \mathcal{M}_{\bar{d}}^2 \approx \left(\begin{array}{ccc|ccc} & & & 0 & 0 & m_b \Gamma_{31}^* \\ & & & 0 & 0 & m_b \Gamma_{32}^* \\ & & & 0 & 0 & m_b \Gamma_{33}^* \\ \hline 0 & 0 & 0 & 0 & 0 & 0 \\ 0 & 0 & 0 & 0 & 0 & 0 \\ m_b \Gamma_{31} & m_b \Gamma_{32} & m_b \Gamma_{33} & 0 & 0 & y_b^2 \Sigma_{33} \end{array} \right)$$

We see that within the third-generation dominant approximation, no flavor violation appears in the RR block for either the up or down squark mass matrices. Also, the only flavor violation in the LR block involves the 3rd generation, and is v/m_{SUSY} suppressed. These features greatly reduce the sensitivity of Q -class χ FV to precision flavor constraints. As we will explain in more detail in section 3.1.1, what typically provide the most stringent flavor bounds on new physics (i.e., Δm_K and Δm_D) involve the 1st and 2nd generations and involve simultaneous violation of flavor in both the left *and* right chiral sectors.

Meanwhile the U -class χ FV models are even more insulated from constraints. Here, on symmetry grounds, and again to leading order in the Yukawas, we must have

$$\delta m_Q^2 = y_u^\dagger \Sigma y_u, \quad A_{\bar{u}} = \Gamma y_u, \quad \delta m_D^2 = A_{\bar{d}} = 0 \quad (2.4)$$

where now Σ and Γ are built out of the $SU(3)_U$ spurions. Again substituting these into (1.3) in the third-generation dominant approximation, we find:

$$\begin{aligned}
 \delta\mathcal{M}_{\tilde{u}}^2 &\approx \left(\begin{array}{ccc|ccc}
 0 & 0 & 0 & 0 & 0 & 0 \\
 0 & 0 & 0 & 0 & 0 & 0 \\
 0 & 0 & y_t^2 \Sigma_{33} & m_t \Gamma_{13}^* & m_t \Gamma_{23}^* & m_t \Gamma_{33}^* \\
 \hline
 0 & 0 & m_t \Gamma_{13} & & & \\
 0 & 0 & m_t \Gamma_{23} & & \delta m_U^2 & \\
 0 & 0 & m_t \Gamma_{33} & & &
 \end{array} \right), \\
 \delta\mathcal{M}_{\tilde{d}}^2 &\approx \left(\begin{array}{ccc|c}
 0 & 0 & 0 & 0 \\
 0 & 0 & 0 & 0 \\
 0 & 0 & y_t^2 \Sigma_{33} & 0 \\
 \hline
 & & 0 & 0
 \end{array} \right) \tag{2.5}
 \end{aligned}$$

Thus there is no down-type flavor violation at all in the third-generation dominant approximation! As the majority of sensitive flavor probes involve the down sector, this eliminates almost all contributions to flavor observables. The only exception is Δm_D , which again has reduced sensitivity because there is no simultaneous left-right flavor violation in the 1st/2nd generations. (Down-type flavor observables can involve the up-squark RR block through chargino loops, but these must be mostly Higgsino-like, so they will be suppressed by Yukawa couplings.)

Finally, we should comment on the role of the RG. The RGEs from the messenger scale to the weak scale add more terms, but the symmetry-based arguments given above — which were truncated at leading order in the Yukawas — clearly continue to hold with the inclusion of higher orders in the Yukawas. In particular, terms that are zero in the third-generation dominant limit remain zero under RG evolution. For this reason, we do not need to concern ourselves with the details of RG running to gain a qualitative understanding of the EGMSB flavor-violating contributions to our flavor observables.

2.2 χ FV in type I squark models

The EGMSB models (1.1) are clearly examples of Q -class or U -class χ FV. In this subsection, we will discuss some further features that are specific to the type I squark models that will help us analyze the Wilson coefficients for flavor observables in the next section.

We begin by quoting the explicit formulas for the soft masses for general type I squark models (see appendix C). For concreteness, we focus on Q -class models, where we have (in addition to the GMSB contributions),

$$\begin{aligned}
 \delta m_Q^2 &= \frac{d_Q}{256\pi^4} \left((d_\phi + d_Q)\kappa^2 - 2C_r g_r^2 - \frac{16\pi^2}{3} h \left(\frac{\Lambda}{M} \right) \frac{\Lambda^2}{M^2} \right) K \Lambda^2 \\
 \delta m_U^2 &= -\frac{d_Q d_U^{QH}}{256\pi^4} y_u K y_u^\dagger \Lambda^2 \\
 \delta m_{H_u}^2 &= -\frac{3d_Q}{256\pi^4} \text{Tr}[y_u K y_u^\dagger] \Lambda^2 \\
 A_{\tilde{u}} &= -\frac{d_Q \Lambda}{16\pi^2} y_u K.
 \end{aligned} \tag{2.6}$$

Here we are neglecting the (numerically irrelevant) down-Yukawa contributions; K is the rank one matrix of couplings defined in (1.4); $d_U^{QH} = 2$ is a multiplicity factor; and $h(x)$ is an $\mathcal{O}(1)$ loop function (see (C.8) for the exact form). For our Q -class model of study (1.2), we have $d_Q = N = 6$, $d_\phi = 5$, where N is the number of messengers.²

We can see the χ FV texture of (2.3) quite clearly in the formulas for δm_U^2 and $A_{\bar{u}}$ in (2.6). There are also additional features of these explicit formulas that go beyond the χ FV ansatz. We notice that δm_Q^2 , together with the quantities Γ and Σ introduced in (2.2), are all proportional to K :

$$\delta m_Q^2 = \beta K, \quad \Gamma = \gamma K, \quad \Sigma = \sigma K \tag{2.7}$$

The simplicity of these relations is partly due to the rank 1 nature of the flavor violation. But in principle, on symmetry grounds alone, there could have been additional contributions to (2.7) proportional to the identity matrix and powers of the Yukawa couplings. These are absent due to the specific form of the type I squark couplings. This is the main reason we have focused on the type I models in this paper. The forms of the soft masses are more complicated in the type II models, and while we expect them to be similarly protected by their χ FV flavor texture, understanding them at the level of analytical detail that we apply to the type I models is more difficult.

Using the relations (2.7), we now discuss the diagonalization of the squark mass matrices. Because the LR blocks are suppressed by v/m_{SUSY} , the squark mass eigenvalues are given by those of the LL and RR blocks to a very good approximation. The RR blocks are already diagonal in the third-generation dominant approximation according to (2.3). The down-squark RR masses are just their minimal GMSB values m_0^2 , while the up-squark RR masses are given by m_0^2 , m_0^2 and $m_0^2 + \sigma y_t^2 \kappa_3^2 \equiv m_{RR}^2$. In practice the third eigenvalue is only a little offset from m_0^2 throughout the parameter space of our model, and this difference can be neglected.

The situation for the LL block is only a little more complicated. Since δm_Q^2 is just proportional to K , the LL block of the squark mass matrices is diagonalized by the unitary transformation

$$U_{LL} = \begin{pmatrix} \hat{\kappa}_1 & a_1 & b_1 \\ \hat{\kappa}_2 & a_2 & b_2 \\ \hat{\kappa}_3 & a_3 & b_3 \end{pmatrix} \tag{2.8}$$

where $\hat{\kappa}_a = \kappa_a/\kappa$ and \vec{a} and \vec{b} are any two orthonormal basis vectors for the orthogonal subspace $\vec{\kappa}^\perp$. The eigenvalues are m_0^2 , m_0^2 , and

$$m_S^2 \equiv m_0^2 + \beta \kappa^2 \tag{2.9}$$

(Here we are ignoring the small differences between the minimal GMSB contributions to the Q , U , D soft masses coming from $SU(2) \times U(1)$.) In our EGMSB models, m_S^2 corresponds roughly to the mass of the lightest squark, and moreover

$$m_S^2 \ll m_0^2 \tag{2.10}$$

²For the similar U -class model formulas, simply change $U \leftrightarrow Q$, $A_u \rightarrow A_u^\dagger$, $y_q \rightarrow y_q^\dagger$, and $K \rightarrow K^T$. The analogous multiplicity factor is $d_Q^{UH} = 1$, and for our U -class model, we have $d_U = 2N = 6$, $d_\phi = 4$.

Thus the Wilson coefficients will generally be dominated by this lightest squark running in the loop, and $m_{\tilde{g}}^2$ will play a central role in controlling the size of the flavor-violating effects in these models. In fact, because of (2.10), diagrams where the right-handed squarks propagate tend to be suppressed, and we will see that it suffices to focus on the LL flavor violation exclusively.

3 Flavor observables and constraints on EGMSB

In this section, we investigate the flavor constraints on our EGMSB models (1.2) in detail. For reference, the spectra corresponding to $\kappa_1 = \kappa_2 = 0$ are shown in figure 1. These are essentially the same points that were identified in [20] as being the least fine-tuned EGMSB models with $m_h = 125$ GeV. Starting from these flavor-aligned points, we will perform a numerical scan in the flavor-violating parameter space ($\kappa_1/\kappa_3, \kappa_2/\kappa_3$) of the models, using **FormFlavor** to compute the flavor observables and compare against their experimental values. Going through each observable X from table 1 in turn, we will exhibit plots of

$$\frac{|[X]_{\text{TH}} - [X]_{\text{EXP}}|}{[\sigma(X)]_{\text{TH+EXP}}} \tag{3.1}$$

Here $[X]_{\text{TH}}$ denotes the SM prediction together with the EGMSB contribution, and $[\sigma(X)]_{\text{TH+EXP}}$ is the theoretical and experimental errors added in quadrature. A contour of 2, which is roughly the 95% CL exclusion limit, will be taken to indicate the point at which the EGMSB model is excluded by the given flavor observable.³

$$\frac{|[\Delta m_X]_{\text{EGMSB}}|}{[\Delta m_X + 2\sigma(\Delta m_X)]_{\text{EXP}}} \tag{3.2}$$

Thus a contour of 2 does not represent a 95% confidence level exclusion for these observables. However, as it would require a substantial tuning for the standard model and the new physics contributions to cancel against one another, values larger than 1 are suspect.

As we deform away from $\kappa_1 = \kappa_2 = 0$, there are numerous subtleties that must be taken into account regarding how the other parameters of the model are varied. These subtleties and specifics of the procedure are described in appendix A. In short, κ_1 and κ_2 are introduced in such a way that the superpartner mass eigenvalues and “net” A -terms are essentially held fixed.

In order to validate the numerical results from **FormFlavor**, we will compare them against analytical formulas for the flavor observables. This will also shed further qualitative insights on the role of χFV in weakening the flavor constraints. As discussed in the Introduction, the usual mass insertion approximation fails due to $\mathcal{O}(1)$ entries. Instead, we will use flavor symmetries and the special features of the type I models discussed above to characterize their *exact* κ dependence, to leading order in an expansion in v/m_S . In general, the κ dependence arises through the lightest squark mass (2.9) and the unitary matrix U_{LL} given in (2.8). (The LR and RL blocks in (2.3) also depend on κ through (2.7),

³Two exceptions to this are Δm_K and Δm_D where the theory errors are uncontrolled. In these cases, we will plot instead:

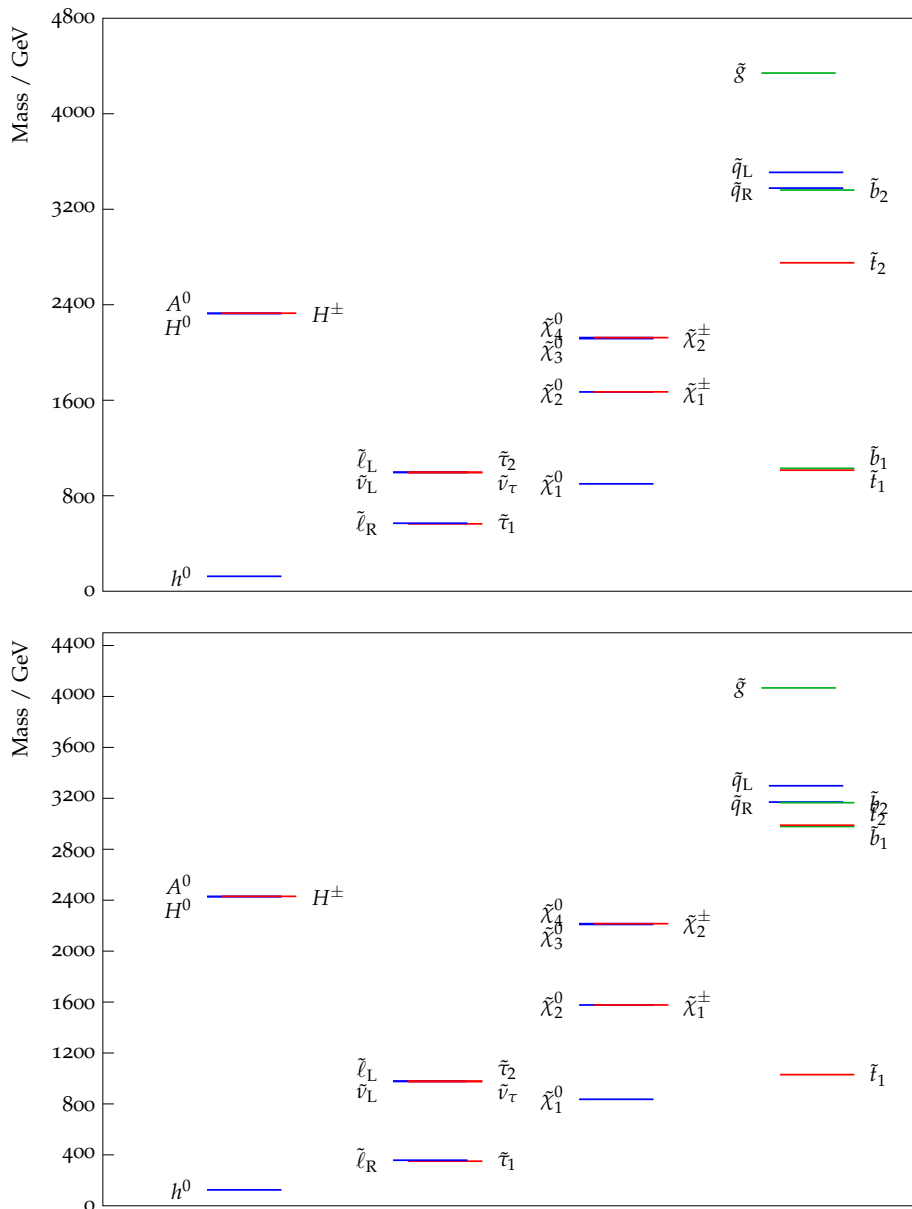


Figure 1. *Top:* mass spectrum of the $\kappa_i Q_i \Phi_D \Phi_L$ model at $\kappa_1 = \kappa_2 = 0$. *Bottom:* mass spectrum of the $\kappa_i U_i \Phi_{D1} \Phi_{D2}$ model at $\kappa_1 = \kappa_2 = 0$.

but as discussed in the previous section, this dependence can be generally be neglected due to the heaviness of the right-handed squarks and the extra v/m_S suppression.) The dependence on U_{LL} is constrained by the flavor symmetries, and as explained in the previous paragraph, the squark mass eigenvalues are mostly held fixed in our parameter space. Thus it will be possible to fully characterize the features of the **FormFlavor** plots in terms of very simple functions of κ .

3.1 Q-class models

3.1.1 Meson mixing

We begin with the $\Delta F = 2$ meson mixing observables. As we discuss in appendix D.1, for $X = K$ and D ,

$$\Delta m_X = 2 \operatorname{Re} \langle \bar{X} | H_{\text{eff}} | X \rangle \quad (3.3)$$

to a good approximation, while for $X = B_d$ and B_q ,

$$\Delta m_X = 2 |\langle \bar{X} | H_{\text{eff}} | X \rangle| \quad (3.4)$$

to a good approximation. Here H_{eff} is the $\Delta F = 2$ effective Hamiltonian; its local short-distance part (relevant for the MSSM contributions) is built out of the four-fermi operators,

$$\begin{aligned} (\mathcal{O}_S^{MN})_{ab} &= (\bar{q}_a P_M q_b)(\bar{q}_a P_N q_b) \\ (\mathcal{O}_V^{MN})_{ab} &= (\bar{q}_a \gamma^\mu P_M q_b)(\bar{q}_a \gamma_\mu P_N q_b) \\ (\mathcal{O}_T^{MN})_{ab} &= (\bar{q}_a \sigma^{\mu\nu} P_M q_b)(\bar{q}_a \sigma_{\mu\nu} P_N q_b) \end{aligned} \quad (3.5)$$

Here $M, N = L, R$ label the chirality of the incoming quarks; and $ab = 12, 13, 23$ for Δm_K , Δm_D , Δm_{B_d} and Δm_{B_s} , respectively, while q is up-type for Δm_D and down-type for the rest.

The full result of `FormFlavor` is shown in figure 2. In this subsection, we will endeavor to understand its features analytically using the χ FV ansatz and the special features of type I EGMSB identified in section 2.2.

The MSSM contributions to the $\Delta F = 2$ observables are due to box diagrams involving the squarks and the gauginos. Here the great simplification of χ FV is that any operator with an R index must transform non-trivially under $SU(3)_D$ or $SU(3)_U$; thus it is suppressed by χ FV in our Q -class models. Furthermore, the \mathcal{O}_S^{LL} and \mathcal{O}_T^{LL} operators all involve an $SU(2)_L$ -breaking chirality flip, so they are dropped in the $v = 0$ approximation. Therefore, the only unsuppressed Wilson coefficient is C_V^{LL} . This also happens to be the only contribution to the one-loop SM Wilson coefficients, which proceeds through W exchange.

Since C_V^{LL} transforms in the square of the adjoint+singlet representation of $SU(3)_Q$, with just left-handed squarks running in the loop, the only way it can depend on κ is,

$$(C_V^{LL})_{ab} = \frac{\hat{\kappa}_a^2 \hat{\kappa}_b^2}{m_S^2} f_1(m_S^2/m_0^2), \quad (3.6)$$

where f_1 is a dimensionless loop function depending on the LL squark mass eigenvalues.⁴ Dimensional analysis fixes the dependence on the masses, and the rest of the dependence must be from the unitary matrix (2.8). Under this simplification, the meson mixing contributions are of the form (see appendix D.1 for more details and specific values of the parameters),

$$[\langle \bar{X} | H_{\text{eff}} | X \rangle]_{\text{EGMSB}} \approx \frac{1}{3} m_X f_X^2 B_{V,X}^{LL} (C_V^{LL})_{ab}, \quad (3.7)$$

⁴All loop functions here and below will implicitly depend on the masses of the other superpartners running in the loop, e.g., the gaugino masses.

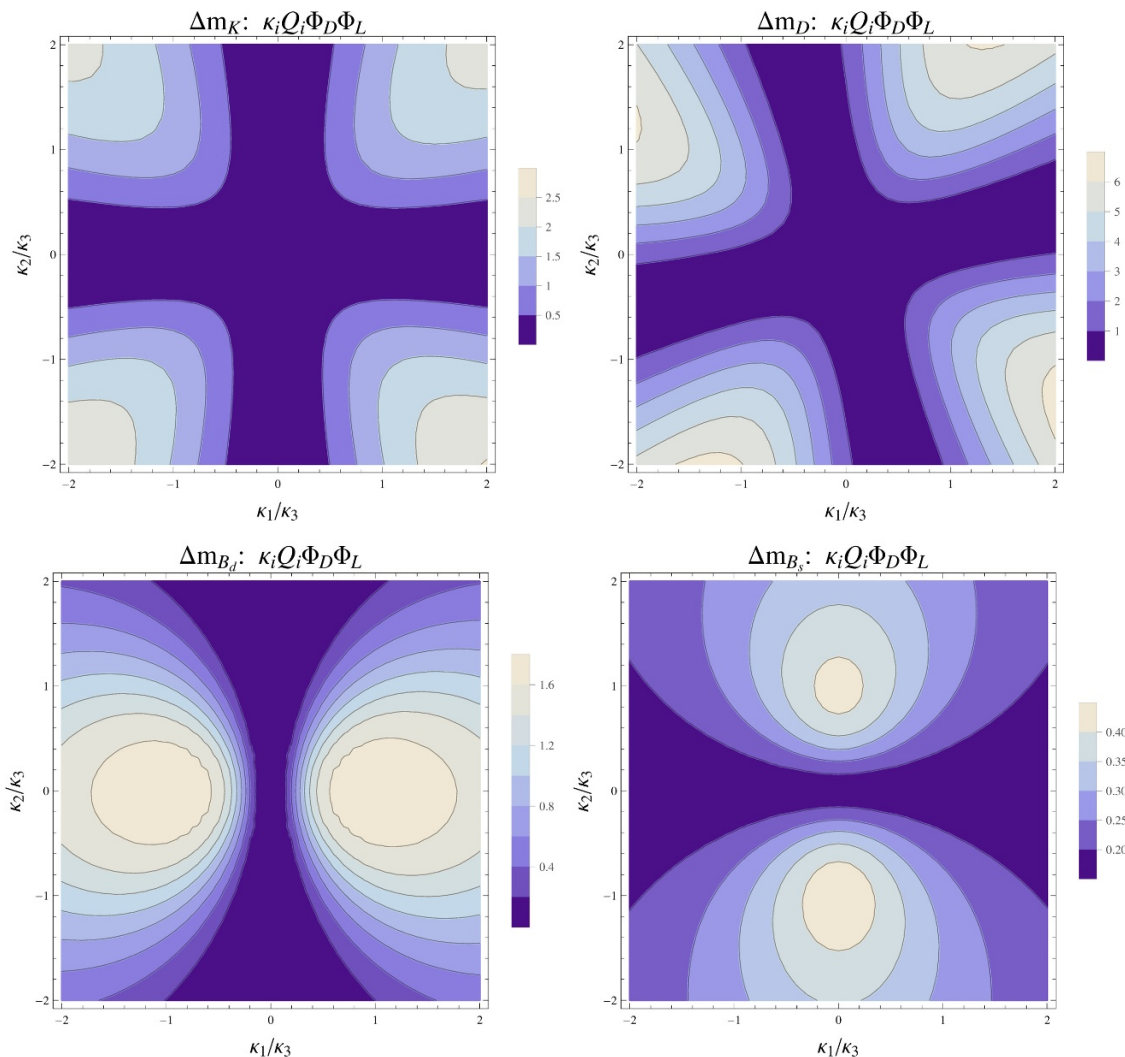


Figure 2. Plots of meson mixing observables the Q -class model, $\kappa_i Q_i \Phi_D \Phi_L$. Δm_K , Δm_D , Δm_{B_d} , Δm_{B_s} are presented in the upper left, upper right, lower left and lower right, respectively. Both Δm_K and Δm_D are presented as the `FormFlavor` output over the experimental value, while Δm_{B_d} and Δm_{B_s} are the difference between `FormFlavor` and the experimental value in units of the uncertainty.

where f_X is the decay constant for the meson; and $B_{V,X}^{LL}$ is an $\mathcal{O}(1)$ hadronic parameter. This simple formula suffices to accurately describe the $\Delta F = 2$ flavor observables.

Focusing on gluino boxes for simplicity,⁵ we find from explicit computation that the leading Wilson coefficient is given by

$$(C_V^{LL})_{ab} = -\hat{\kappa}_a^2 \hat{\kappa}_b^2 \frac{\alpha_s^2}{36m_S^2} f_{\tilde{g}}^{\Delta M, \text{box}}(x_g, x_q) \quad (3.8)$$

⁵Chargino and gluino-neutralino boxes are also typically comparable. However, they do not affect the qualitative discussion here. In fact, because the charginos enter with the opposite sign of the gluino and gluino+neutralino contributions, the gluinos alone provide a better estimate quantitatively than might be expected.

This clearly agrees with the general form (3.6). Here, $x_q = m_S^2/m_0^2$ and $x_g = m_{\tilde{g}}^2/m_0^2$, and $f_{\tilde{g}}^{\Delta M, \text{box}}(x_g, x_q) \approx 0.05$ is a loop function that we define in appendix E. As discussed earlier, since the relevant masses are roughly held fixed in our deformation, the loop function does not change across the parameter space. The QCD RG running from the SUSY scale to the meson scale is fairly mild for C_V^{LL} , yielding only a 20 – 30% suppression in the size of the Wilson coefficient [52].

The hadronic factors $f_X^2 B_{V,X}^{LL}$ are fairly similar across the mesons. We can define,

$$H_X \equiv \frac{f_X^2 B_{V,X}^{LL}}{f_K^2 B_{V,K}^{LL}}; \quad H_K = 1, \quad H_D \approx H_{B_d} \approx 2, \quad H_{B_s} \approx 3 \quad (3.9)$$

So we expect deviations of the form,

$$[\langle \bar{X} | H_{\text{eff}} | X \rangle]_{\text{EGMSB}} \sim -\hat{\kappa}_a^2 \hat{\kappa}_b^2 H_X m_X (10^{-13}), \quad (3.10)$$

Noting that $\hat{\kappa}_a^2 \hat{\kappa}_b^2$ is at most $\frac{1}{4}$, and comparing against table 1, we see that Δm_K and Δm_D should be most sensitive to EGMSB, while Δm_{B_d} should be barely sensitive, and Δm_{B_s} completely insensitive.⁶

These sensitivities are observed in the plots shown in figure 2. Moreover, the $\hat{\kappa}_a^2 \hat{\kappa}_b^2$ dependence is transparent in these plots. ΔB_d peaks at $(\kappa_1/\kappa_3, \kappa_2/\kappa_3) \sim (\pm 1, 0)$, while ΔB_s peaks at $\sim (0, \pm 1)$. For Δm_K and Δm_D , we can see that moving from $\kappa_1 = \kappa_2 = \kappa_3$ to the corners, constraints rise by the expected factor of $\sim 16/9$.⁷

These results from meson mixing may seem to conflict with the SUSY flavor problem, which suggests that, with $\mathcal{O}(1)$ flavor-violation, the SUSY scale needs to enter above ~ 500 TeV due to constraints from Δm_K and Δm_D [49, 50]. However, these constraints are driven by the C_S^{LR} Wilson coefficient, while χFV only generates the C_V^{LL} operator. In the MSSM, the contribution of the latter to $\Delta m_{K,D}$ is suppressed by $\sim 10^{-4} - 10^{-3}$ relative to the former. This is due to three separate effects that all work in the same direction. First, the hadronic matrix elements differ between these two operators,

$$\frac{\langle K | S_{LR} | \bar{K} \rangle}{\langle K | V_{LL} | \bar{K} \rangle} = \frac{3}{4} \frac{B_S^{LR}}{B_V^{LL}} R_K \sim 35, \quad \frac{\langle D | S_{LR} | \bar{D} \rangle}{\langle D | V_{LL} | \bar{D} \rangle} = \frac{3}{4} \frac{B_S^{LR}}{B_V^{LL}} R_D \sim 4. \quad (3.11)$$

Next, the SUSY contributions to the Wilson coefficients are also quite different in size. From the MIA with $\mathcal{O}(1)$ mass-insertions, it is easy to see that $C_S^{LR}/C_V^{LL} \sim 30$ [49]. Lastly, the QCD running from the SUSY scale to ~ 2 GeV suppresses C_V^{LL} by 20-30%, but *enhances* C_S^{LR} by a factor of ~ 3 [52]. All of these factors conspire to drop the scale of sensitivity to Δm_K and Δm_D to the TeV scale in χFV models.

⁶As discussed in the Introduction, we are not considering CP violation in this work, in particular ϵ_K . Although it depends on the precise value of the CP violating phase, the expectation is that this will place a meaningful, tighter constraint on the parameter space when the phase is large. This and other CPV observables will be studied in an upcoming paper [35].

⁷Note that Δm_D is rotated by the Cabbibo angle θ_c relative to the other observables. This is a consequence of the the fact that, as discussed in in appendix A, our κ_1 , κ_2 , and κ_3 directions are chosen to align with the low energy down, strange and bottom quark. For Δm_D , the dependence on κ_i proceeds through the LL sector of the up-squark mass matrix, where there is an additional rotation by V_{CKM} .

3.1.2 $K^\pm \rightarrow \pi^\pm \nu \bar{\nu}$

Unlike meson mixing in the previous subsection, the $K^\pm \rightarrow \pi^\pm \nu \bar{\nu}$ observable (along with all other $\Delta F = 1$ observables) enters as a matrix element squared, and thus interference with the standard model contribution can be important. The expression for this branching ratio is (see appendix D.2),

$$\text{BR}(K^\pm \rightarrow \pi^\pm \nu \bar{\nu}) = c_+ v^4 |C_{V,\text{SM}}^{LL} + C_V^{LL} + C_V^{RL}|^2 \quad (3.12)$$

where $c_+ v^4 = 4.9 \times 10^9 \text{ GeV}^4$, $C_{V,\text{SM}}^{LL} = (-1.21 + 0.39i) \times 10^{-10} \text{ GeV}^{-2}$, and all EGMSB effects are contained in C_V^{LL} and C_V^{RL} . These are the Wilson coefficients from the $\Delta F = 1$ effective Hamiltonian built out of the four-fermi operators

$$\begin{aligned} (\mathcal{O}_S^{MN})_{ab} &= (\bar{q}_a P_M q_b)(\bar{\ell} P_N \ell) \\ (\mathcal{O}_V^{MN})_{ab} &= (\bar{q}_a \gamma^\mu P_M q_b)(\bar{\ell} \gamma_\mu P_N \ell) \\ (\mathcal{O}_T^{MN})_{ab} &= (\bar{q}_a \sigma^{\mu\nu} P_M q_b)(\bar{\ell} \sigma_{\mu\nu} P_N \ell) \end{aligned} \quad (3.13)$$

Here $ab = 12$ for $K^\pm \rightarrow \pi^\pm \nu \bar{\nu}$. In the MSSM, these Wilson coefficients arise through one-loop box and Z -penguin diagrams.

The full FormFlavor result is shown in figure 3. The general trend of $K^\pm \rightarrow \pi^\pm \nu \bar{\nu}$ can again be understood through use of the features of rank 1 χ FV discussed in section 2.2. First of all, as for the $\Delta F = 2$ observables, $C_{V,ab}^{RL}$ must be zero in the third-generation dominant approximation — since it transforms in the adjoint+singlet of $\text{SU}(3)_D$, using the available spurions one can only obtain something that is nonzero in the 33 component. Thus, we can focus on $C_{V,ab}^{LL}$. This transforms in the adjoint+singlet of $\text{SU}(3)_Q$. Using flavor violation in the LL block only, the form of $C_{V,ab}^{LL}$ is constrained by the symmetries to be:

$$C_{V,ab}^{LL} = \frac{\hat{\kappa}_a \hat{\kappa}_b}{m_S^2} f_2 + \dots \quad (3.14)$$

where \dots contains higher orders in v/m_{SUSY} and other irrelevant terms, and f_2 is a dimensionless function of superpartner mass ratios. We note that (3.14) can come from box diagrams or Z -penguin diagrams, but in the latter case, the $1/m_Z^2$ from the Z propagator must be canceled out by two insertions of wino-Higgsino mixing. Insertions of LR mixing from the squark mass matrix would also cancel out the $1/m_Z^2$, but the heavy right-handed squark masses suppress these contributions enough that they may be ignored.

Now we will compare against explicit computations of the Wilson coefficients. We confirm that the C_V^{RL} coefficients are all negligible. For C_V^{LL} , we find that the chargino diagrams dominate, and they are given by

$$C_V^{LL} = \hat{\kappa}_1 \hat{\kappa}_2 \frac{\alpha_2^2}{12m_S^2} f_{\tilde{\chi}^\pm}^{K \rightarrow \pi \nu \nu}(x_\ell, x_2, x_\mu) \quad (3.15)$$

Again, $f_{\tilde{\chi}^\pm}^{K \rightarrow \pi \nu \nu}(x_\ell, x_2, x_\mu) \sim 1.5$ is a loop function of the sparticle mass-ratios $x_\ell = m_\ell^2/m_S^2$, $x_2 = M_2^2/m_S^2$, and $x_\mu = \mu^2/m_S^2$, defined in appendix E, and varies only mildly across the parameter space. This is fully consistent with the general form (3.14). This also explains

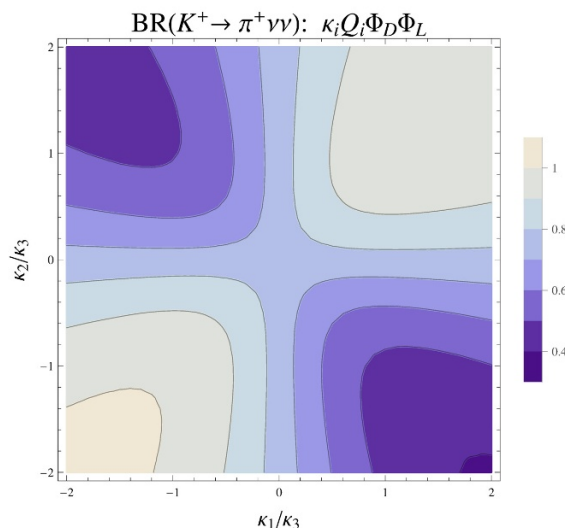


Figure 3. Contours of the difference in $\text{BR}(K^+ \rightarrow \pi^+ \nu \bar{\nu})$ between experiment and the SM + EGMSB predictions, given in units of the net uncertainty (added in quadrature). A contour of 2 corresponds roughly to a 95% exclusion.

why the charginos dominate over the gluinos, since the gluinos can only give rise to Z -penguin diagrams that are nonzero by virtue of down sector LR squark-mixing insertions.

Substituting in numerically for the loop function, m_S and α_2 , we find

$$C_V^{LL} \sim (8 \times 10^{-11} \text{ GeV}^{-2}) \hat{\kappa}_1 \hat{\kappa}_2 \quad (3.16)$$

This simple function of $\vec{\kappa}$, when added to the SM contribution and substituted into (3.12), reproduces well the features of figure 3. It grows in magnitude fastest along the lines $\kappa_1 = \pm \kappa_2$, asymptoting to the values $\pm 4 \times 10^{-11} \text{ GeV}^{-2}$. In the corners of the parameter space we obtain a deviation from the SM prediction of

$$\Delta \text{BR}(K^+ \rightarrow \pi^+ \nu \bar{\nu}) = c_+ v^4 \left(2 \text{Re} [C_{V,\text{SM}}^{LL} C_V^{LL*}] + |C_V^{LL}|^2 \right) \approx \mp 5 \times 10^{-11}, \quad (3.17)$$

This is smaller than the current experimental+theoretical uncertainty shown in table 1. However, because the SM prediction for the BR is a little lower than the experimentally observed value, moving along the $\kappa_1 = +\kappa_2$ line slightly exacerbates the difference, while moving along the $\kappa_1 = -\kappa_2$ line slightly lessens it.

3.1.3 $b \rightarrow s\gamma$ and $b \rightarrow d\gamma$

The expression for the $b \rightarrow s\gamma$ branching ratio (see appendix D.3) is,

$$\text{BR}(b \rightarrow s\gamma) = c_\gamma v^2 \left(|C_{A,\text{SM}}^L + C_A^L|^2 + |C_A^R|^2 \right), \quad (3.18)$$

where $c_\gamma v^2 = 1.97 \times 10^{12} \text{ GeV}^2$ and $C_{A,\text{SM}}^L \sim 1.3 \times 10^{-8} \text{ GeV}^{-1}$. The new physics is contained in C_A^L and C_A^R ; these come from the effective Hamiltonian built out of the dimension 5 operators

$$(\mathcal{O}_A^M)_{ab} = e(\bar{q}_a \sigma^{\mu\nu} P_M q_b) F_{\mu\nu} \quad (3.19)$$

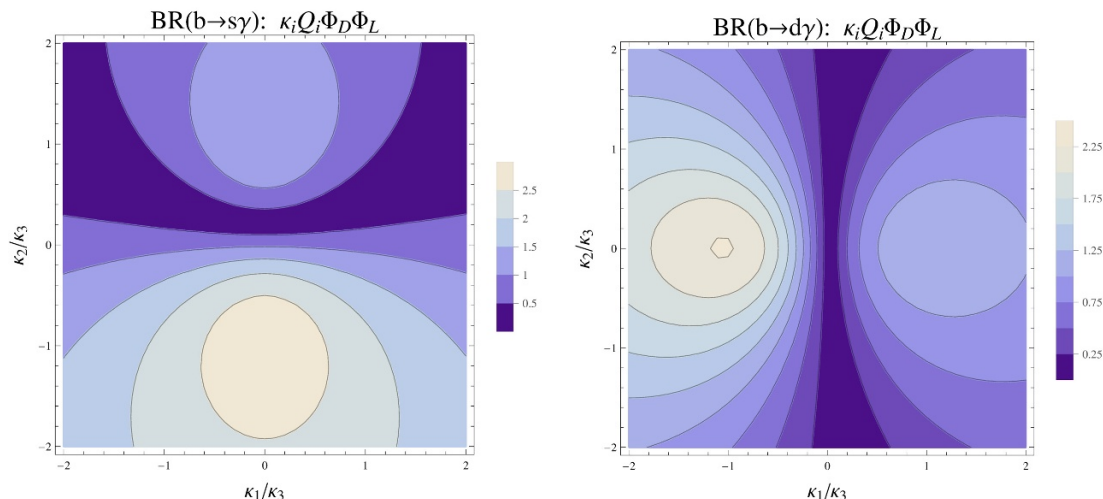


Figure 4. Sensitivity to $b \rightarrow s\gamma$ (left) and $b \rightarrow d\gamma$ (right), same conventions as figure 3.

where $ab = 32$ for $b \rightarrow s\gamma$. These operators require $SU(2)$ breaking, so their Wilson coefficients are zero in the $v \rightarrow 0$ limit. Naively, this would mean that the result is negligible in our zeroth-order expansion in v/m_{SUSY} . However, the SM Wilson coefficient is suppressed by m_b/v , so here we consider one higher order in the v/m_{SUSY} expansion in order to capture a numerically relevant result.

Since C_A^L (C_A^R) transforms in the $(\mathbf{3}, \bar{\mathbf{3}})$ of $SU(3)_D \times SU(3)_Q$ ($SU(3)_Q \times SU(3)_D$), it must be proportional to m_d acting on the left (right). The latter is zero in the third-generation dominant approximation, so it suffices to focus on C_A^L . With just left-handed squarks propagating in the loop, the Wilson coefficient must be given by:

$$C_{A,32}^L = \frac{(m_d K \tan \beta)_{32}}{\kappa^2} \frac{1}{m_S^2} f_6 = \hat{\kappa}_2 \hat{\kappa}_3 \frac{m_b \tan \beta}{m_S^2} f_6 \quad (3.20)$$

Here we assumed that m_d is accompanied by a $\tan \beta$ enhancement, otherwise the entire effect is numerically negligible.

Comparing with explicit calculation, we find again that charginos give the dominant contribution to the Wilson coefficient, through the quark-squark-Higgsino vertex. (Contributions of the form (3.20) can also arise through gluino and neutralino loops, but here the factor of $m_b \tan \beta$ arises through the LR block of the down-squark mass matrix, so the diagrams are suppressed by heavy right-handed squarks propagating in the loop.) The dominant chargino diagram gives,

$$C_A^L = \hat{\kappa}_2 \hat{\kappa}_3 \frac{m_b \tan \beta}{m_S^2} \frac{11\alpha_2}{288\pi} f_{\tilde{\chi}^\pm}^{b \rightarrow s/d\gamma, \text{peng}}(x_\mu, x_2), \quad (3.21)$$

where $x_\mu = \mu^2/m_S^2$, $x_2 = M_2^2/m_S^2$, and $f_{\tilde{\chi}^\pm}^{b \rightarrow s/d\gamma, \text{peng}}(x_\mu, x_2) \sim 0.5$ is a loop function defined in appendix E. This is fully in agreement with the general result (3.20). QCD running from the SUSY scale to the b pole [53] induces a mild 20% suppression to (3.21), and once again, we find this simple result is enough to account for the features of the `FormFlavor` plot in figure 4.

Substituting in numerically for our parameter space (we take $\tan \beta = 10$ as in [20]), we obtain:

$$C_A^L \sim (3 \times 10^{-9} \text{ GeV}^{-1}) \hat{\kappa}_2 \hat{\kappa}_3 \quad (3.22)$$

Clearly its effects are largest when $\kappa_1 = 0$, and along the κ_2 axis it has a maximum (minimum) at $\kappa_2/\kappa_3 = 1$ (-1), while asymptoting back to zero as $\kappa_2/\kappa_3 \rightarrow \pm\infty$. Substituting this into (3.18) together with the SM Wilson coefficient, we find at these extrema a deviation from the SM prediction of

$$[\Delta\text{BR}(b \rightarrow s\gamma)]_{\kappa_2=\pm\kappa_3} \sim c_\gamma v^2 (2 \text{Re} [C_{A,\text{SM}}^L C_A^{L*}]) \sim \pm 7 \times 10^{-5}, \quad (3.23)$$

Since the uncertainty on the measurement (combining the theoretical and experimental in quadrature) is 3.5×10^{-5} , regions of exclusion are to be expected for $b \rightarrow s\gamma$.

In practice, `FormFlavor` does find a region of exclusion in the bottom half of the parameter space, shown in figure 4. This exclusion is in part due to the current $\mathcal{O}(1\sigma)$ excess in the experimental measurement relative to the theoretical prediction,⁸ which allows the $\sim 2\sigma$ change to constrain a large region of negative κ_2 in the plot. Uncertainty on this observable is comparable in size between theory and experiment, so improvements on either side could make this observable more constraining and in a direction that other observables currently have no sensitivity.

The observable $b \rightarrow d\gamma$ could also potentially place constraints. The EGMSB contribution to the Wilson coefficient is the same as in (3.21)–(3.22), but with $\kappa_2 \leftrightarrow \kappa_1$. However, the SM contribution is of a different size, $C_{A,\text{SM}}^L \sim -(2.4 + 1.1i) \times 10^{-9} \text{ GeV}^{-1}$. Importantly, this is not much larger than the new physics contribution, so interference is very important. This yields an approximate deviation of,

$$\begin{aligned} [\Delta\text{BR}(b \rightarrow d\gamma)]_{\kappa_1=\kappa_3} &\sim c_\gamma v^2 \left(-2 |\text{Re} [C_{A,\text{SM}}^L C_A^{L*}]| + |C_A^L|^2 \right) \sim -0.8 \times 10^{-5} \\ [\Delta\text{BR}(b \rightarrow d\gamma)]_{\kappa_1=-\kappa_3} &\sim c_\gamma v^2 \left(2 |\text{Re} [C_{A,\text{SM}}^L C_A^{L*}]| + |C_A^L|^2 \right) \sim 1.6 \times 10^{-5}, \end{aligned} \quad (3.24)$$

while the net uncertainty is 6.3×10^{-6} . Again, a two sigma exclusion could be possible, but it would only be expected near $\frac{\kappa_1}{\kappa_3} = -1$ due to the interference acting constructively. This is borne out by figure 4, which shows a maximum deviation of ≈ 2.3 times the net uncertainty. Because the measurement of $b \rightarrow d\gamma$ is fairly recent, a future improvement that pushes the experimental uncertainty to the 10-15% level would allow for $b \rightarrow d\gamma$ to place much tighter constraints on the region of constructive interference, i.e., $\kappa_1 < 0$.

3.1.4 $B_q \rightarrow \mu^+ \mu^-$

As shown in figure 5, neither $B_s \rightarrow \mu^+ \mu^-$ nor $B_d \rightarrow \mu^+ \mu^-$ is at all close to constraining our EGMSB models. We include a brief discussion of these observables just for completeness sake.

⁸Recently, an improved theoretical prediction of $b \rightarrow s\gamma$ was released [54]. This work predicts that $\text{BR}(b \rightarrow s\gamma) = (3.36 \pm 0.23) \times 10^{-4}$, which is significantly more in line with the experimental value. Unfortunately, insufficient details are provided in that work for us to modify `FormFlavor` to account for this prediction.

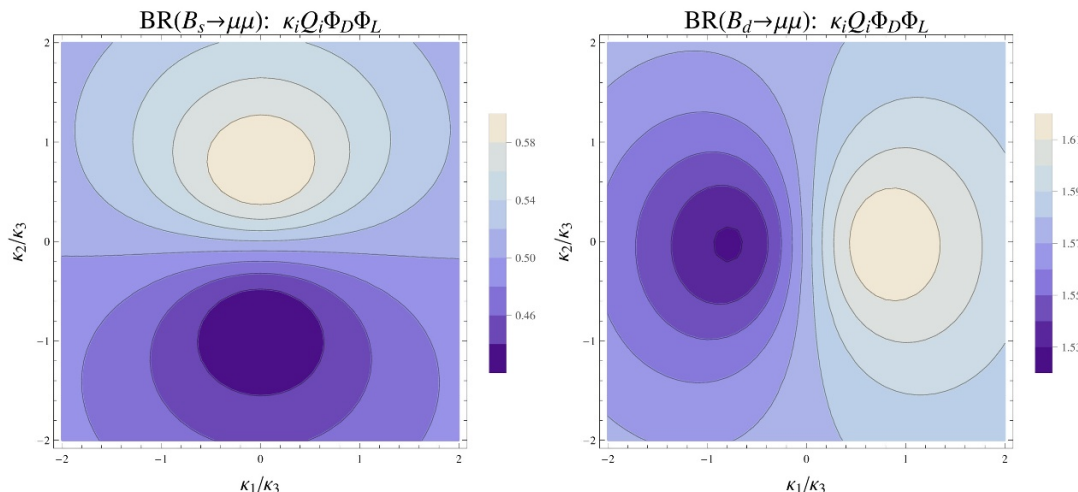


Figure 5. Sensitivity to $B_s \rightarrow \mu\mu$ and $B_d \rightarrow \mu\mu$, same conventions as figure 3. Neither is constraining over the parameter space.

The branching ratio for $B_q \rightarrow \mu^+\mu^-$ is [55, 56] (see appendix D.4 for more detailed formulas):

$$\text{BR}(B_q \rightarrow \mu^+\mu^-) = X_q \left\{ \left(1 - \frac{4m_\mu^2}{m_{B_q}^2} \right) |F_S^{(q)}|^2 + |F_P^{(q)} + F_A^{(q)}|^2 \right\} \quad (3.25)$$

where $X_s = 5.36 \times 10^7$ and $X_d = 3.97 \times 10^7$, and

$$\begin{aligned} F_S^{(q)} &= \frac{m_{B_q}^3}{m_b + m_q} (C_S^{LL} + C_S^{LR} - C_S^{RR} - C_S^{RL}), \\ F_P^{(q)} &= \frac{m_{B_q}^3}{m_b + m_q} (-C_S^{LL} + C_S^{LR} - C_S^{RR} + C_S^{RL}), \\ F_A^{(q)} &= 2m_{B_q}m_\mu (C_V^{LL} - C_V^{LR} + C_V^{RR} - C_V^{RL}), \end{aligned} \quad (3.26)$$

Here the C_Y^{MN} coefficients are those for the $\Delta F = 1$ effective Hamiltonian introduced in (3.13). The three-loop Standard Model contribution [57] is $F_{A,\text{SM}}^{(d)} = (1.5 - 0.6i) \times 10^{-9}$ and $F_{A,\text{SM}}^{(s)} = (-7.9 - 0.1i) \times 10^{-9}$.

In the MSSM, the dominant contributions to $B_q \rightarrow \mu^+\mu^-$ are the $\tan^3 \beta$ enhanced wave function correction diagrams with a heavy higgs propagator, see figure 6. Here gluinos, charginos and neutralinos can run in the loop, with both CP even and odd higgs states along the penguin line. For simplicity, we will quote the result only for gluinos; the answer for the others is very similar. As our higgs states are heavy, we use the relation $m_A^2 \approx m_H^2$ to simplify expressions.

These diagrams contribute to C_S^{MN} , and the discussion is similar to that of $b \rightarrow q\gamma$. C_S^{LN} transforms in the $(\mathbf{3}, \bar{\mathbf{3}})$ of $\text{SU}(3)_D \times \text{SU}(3)_Q$ while C_S^{RN} transforms in the $(\mathbf{3}, \bar{\mathbf{3}})$ of $\text{SU}(3)_Q \times \text{SU}(3)_D$. Thus as before, the latter is zero thanks to χFV and we can focus on the former. Here we can afford to work at $v = 0$ since we are not concerned with Z -penguins.

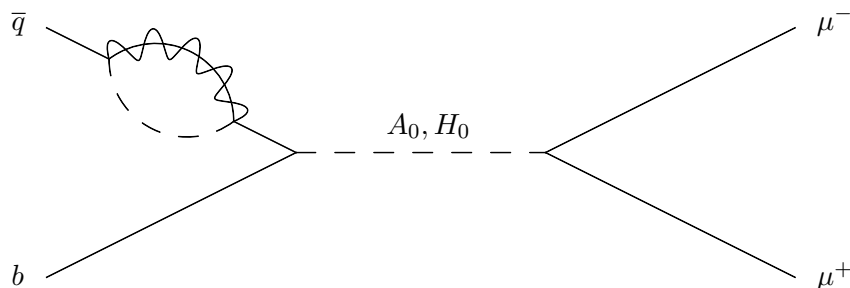


Figure 6. Wave function correction diagrams that provide the leading, $\tan^3 \beta$ contributions to $B_q \rightarrow \mu^+ \mu^-$. Loop can be a gluino, chargino, or neutralino.

With only left-handed squarks propagating in the loops, by symmetries the answer must be of the form:

$$C_{S,3a}^{LM} = \frac{(Ky_d)_{3a}}{\kappa^2 m_S^2} f_5 = \hat{\kappa}_3 \hat{\kappa}_a \frac{y_b}{m_S^2} f_5 \quad (3.27)$$

where $a = 2$ for $B_s \rightarrow \mu\mu$ and $a = 1$ for $B_d \rightarrow \mu\mu$. Now the Yukawa coupling needed on symmetry grounds arises from the Higgs-quark-quark coupling. Explicit computation gives

$$C_{S,\hat{g}}^{LR} = \hat{\kappa}_3 \hat{\kappa}_a \frac{m_b \tan \beta}{m_W} \frac{m_\mu \tan^2 \beta}{m_W} \frac{4\alpha_2 \alpha_s}{3m_A^2} \frac{\mu}{m_S} f_{\hat{g}}^{B_q \rightarrow \mu^+ \mu^-, \text{h-peng}}(x_q, x_g) \quad (3.28)$$

where $x_g = m_{\hat{g}}^2/m_0^2$ and $x_q = m_S^2/m_0^2$ and the loop function (see appendix E) is $f_{\hat{g}}^{B_q \rightarrow \mu^+ \mu^-, \text{h-peng}}(x_q, x_g) \sim 0.1$. Since $m_A \sim m_{\text{SUSY}}$, this leading contribution is of the form (3.27) as expected.

From the above expressions, we can translate to the phenomenologically useful parameters

$$F_P^{(s)} = \frac{m_{B_s}^3}{m_b + m_s} C_S^{LR} \sim (2 \times 10^{-10}) \hat{\kappa}_3 \hat{\kappa}_2 \quad (3.29)$$

$$F_P^{(d)} = \frac{m_{B_d}^3}{m_b + m_d} C_S^{LR} \sim (2 \times 10^{-10}) \hat{\kappa}_3 \hat{\kappa}_1.$$

These translate into maximum deviations from the SM branching fractions by

$$\begin{aligned} [\Delta\text{BR}(B_s \rightarrow \mu^+ \mu^-)]_{\text{MAX}} &\simeq X_s \left\{ 2 \text{Re} \left[F_{A,\text{SM}}^{(s)*} F_P^{(s)} \right] \right\} \sim 9 \times 10^{-11} \\ [\Delta\text{BR}(B_d \rightarrow \mu^+ \mu^-)]_{\text{MAX}} &\simeq X_d \left\{ 2 \text{Re} \left[F_{A,\text{SM}}^{(d)*} F_P^{(d)} \right] \right\} \sim 1.2 \times 10^{-11} \end{aligned} \quad (3.30)$$

These deviations are an order of magnitude smaller than the uncertainties on their respective measurements, and therefore neither $B_s \rightarrow \mu^+ \mu^-$ nor $B_d \rightarrow \mu^+ \mu^-$ place any meaningful constraints on our models.⁹

⁹In order for $B_s \rightarrow \mu^+ \mu^-$ to place any meaningful constraint, one would have to take $\tan \beta$ much larger than the $\tan \beta \sim 10$ that we have assumed in this paper, e.g., $\tan \beta \gtrsim 30$.

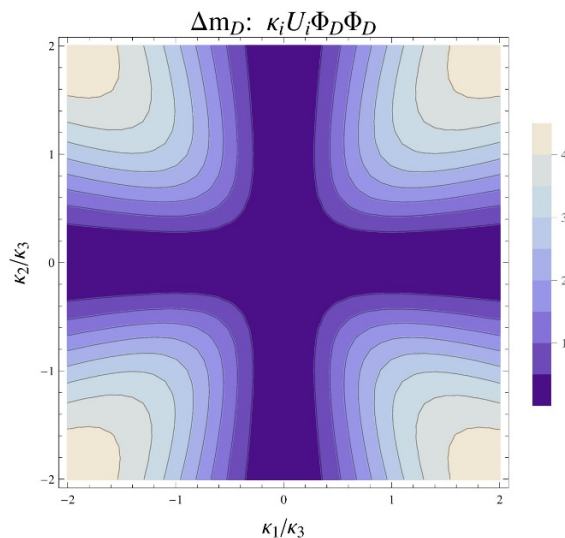


Figure 7. Plot of Δm_D meson mixing observable for a U -class model.

3.2 U -class models

As discussed in section 2.1, all U -class χ FV models receive very few constraints from flavor observables. First, according to the χ FV texture of U -class models (2.5), all flavor-violation is restricted to the up-squark sector only. As most potentially constraining flavor observables have external down-type quarks, chargino diagrams are required for sensitivity to flavor-violation. However, as is further shown in (2.5), there is no flavor-violation in the LL block; thus pure wino diagrams cannot contribute. Higgsino diagrams introduce Yukawa couplings which suppress contributions to down sector observables enough that none of these would be remotely constraining in the foreseeable future. We have verified all of these general results in the context of our type I EGMSB models.

There can be constraints from the D -meson system. Contributions to Δm_D are as in the Q -class models, only now with the C_V^{RR} contribution dominating, so we expect very similar constraints from Δm_D . This is indeed shown in figure 7. Again, although the contour of 2 does not represent a 95% confidence level exclusion, values larger than 1 necessitate a cancellation between the standard model and the new physics contributions. It should be noted that a viable possibility is that the standard model contribution is in fact much smaller than the observed value, and the contour of one is actually where EGMSB entirely accounts for the Δm_D measurement. As with the Q -class models, the sensitivity vanishes as either $\kappa_1 \rightarrow 0$ or $\kappa_2 \rightarrow 0$, and increases most rapidly along the diagonals.

4 Conclusions

4.1 Summary and discussion

In this work, we performed a detailed investigation into the precision flavor constraints on extended GMSB models. These models, where the usual GMSB contributions are augmented by direct matter-messenger couplings, are well motivated in light of the recent

discovery of a Higgs boson near 125 GeV. However, since these models are not necessarily MFV, in their full, three-family generalizations they could be potentially dangerous from the point of view of flavor.

Our work required a computer program that could turn general MSSM spectra into precision flavor observables. We found that existing programs had various limitations — they either assumed MFV, were numerically unstable, or had incorrectly transcribed formulas from the literature. This motivated us to develop `FormFlavor`, a comprehensive package that computes flavor observables *ab initio* starting from the Feynman rules, and uses a modular framework that enables us to add new observables in a uniform and straightforward way.

Using `FormFlavor`, we studied the flavor constraints on the three-family generalizations of the EGMSB models of [20]. The results we encountered from this systematic study were interesting and unexpected. We found that despite the introduction of $\mathcal{O}(1)$ flavor-violating couplings, there are currently very few constraints on EGMSB models. To validate the numerical results of `FormFlavor`, we compared them in detail with analytic formulas for the Wilson coefficients derived using a combination of flavor symmetry arguments and direct calculation, and we found excellent agreement.

The mild flavor constraints in these models are illustrated in the summary plot of figure 8. *U*-class models only receive constraints from Δm_D . *Q*-class models are constrained by Δm_K and Δm_D in the corners of the plot, while the radiative $b \rightarrow s\gamma$ and $b \rightarrow d\gamma$ each exclude a single bubble near $\kappa_2 = -\kappa_3$ and $\kappa_1 = -\kappa_3$, respectively. We note that the excluded region from $b \rightarrow s\gamma$ is largely due to the current $\sim 1\sigma$ discrepancy between the theoretical prediction and the measurement, which new theoretical work suggests will disappear [54].

The results from these models may seem at odds with the SUSY flavor problem. We have argued that the mildness of the flavor violation in these models originates from the fact that they obey the “chiral flavor violation” ansatz, whereby flavor is violated only by the Yukawas and spurions of a single $SU(3)$ of the full $SU(3)^5$ SM flavor symmetry. We showed that χ FV prevents many of the most problematic contributions to flavor observables from arising in the MSSM, and allows for $\mathcal{O}(1)$ flavor-violation in EGMSB models.

4.2 Future constraints

Although constraints are currently very mild, there is immense potential to further probe these models in the near future.

- For Δm_K , short-distance predictions exist with moderate uncertainty, i.e., $(3.1 \pm 1.2) \times 10^{-15}$ GeV [58], but the long-distance contributions are currently unknown. However, an accurate, full calculation on the lattice, including both long- and short-distance contributions, may be coming in the near future, as promising preliminary work on the subject shows [59, 60].
- Expected incremental improvements to the theoretical uncertainty of the standard model prediction could make ΔB_d into a constraining observable soon. Estimates

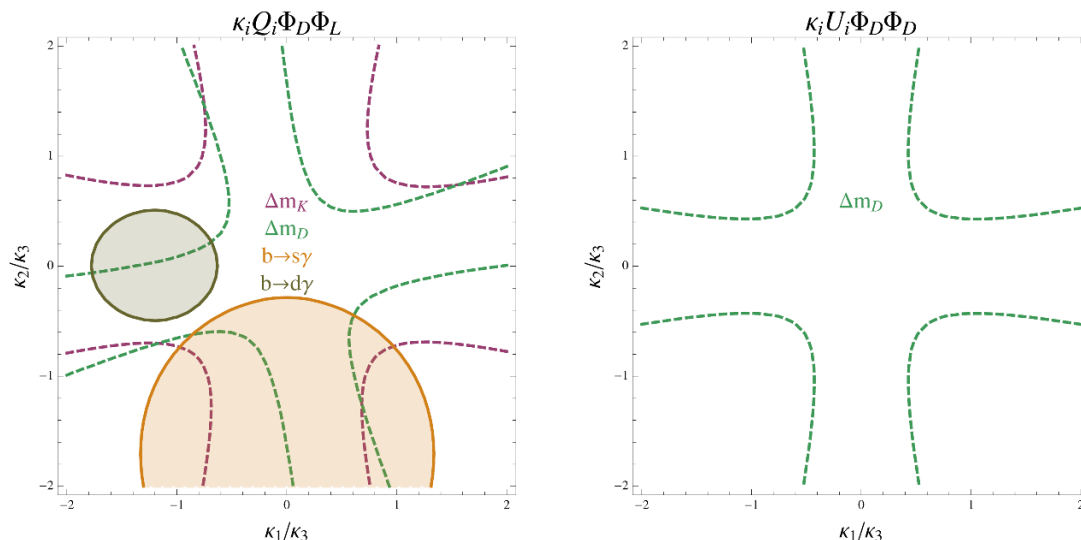


Figure 8. Flavor sensitivity for a Q -class (left) and U -class (right) model. Shaded regions of this EGMSB parameter space are excluded to 95% confidence level by current measurements of flavor observables as described in the text. Dashed lines indicate the locations where the EGMSB contribution to Δm_K and Δm_D are equal to the measured value (as these currently cannot provide a genuine exclusion due to poorly controlled theoretical uncertainties). In both cases, observables not shown are not constraining.

suggest that both the bag parameter, B_{V,B_d}^{LL} , and the relevant CKM elements could be calculated on the lattice to significantly improved levels by 2018 [61].

- Belle II [62] is expected to make significantly improved measurements to both $b \rightarrow s\gamma$ and $b \rightarrow d\gamma$ [63], which would allow for both of these observables to constrain more of the parameter space.
- NA62 [64] at CERN is projected to be able to measure $\text{BR}(K^\pm \rightarrow \pi^\pm \nu \bar{\nu})$ to a precision of about 10% [65] of the SM value. This is an improvement of more than an order of magnitude relative to the current measurement. In a few years, $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ will be one of the the most sensitive flavor observable to Q -class models. Perhaps more importantly, if NA62 were to measure a deviation from the standard model prediction, this model would provide a significant motivation to invest in an experiment like ORKA [66], that would be able to hone in on the parameter space.

Due to these potentially significant theoretical and experimental improvements, much of the parameter space in Q -class models could be probed in just a few years. The projected sensitivities are shown in figure 9. Although Q -class models have an exciting future in flavor, U -class models remain completely unconstrained. Short of lattice predictions for Δm_D , no observables in the current program are sensitive to these these models. Charm factories and precision top studies could someday explore this space in, for instance, $c \rightarrow u\gamma$ or $t \rightarrow c/u\gamma$. However, for the moment, U -class models are resilient against flavor constraints.

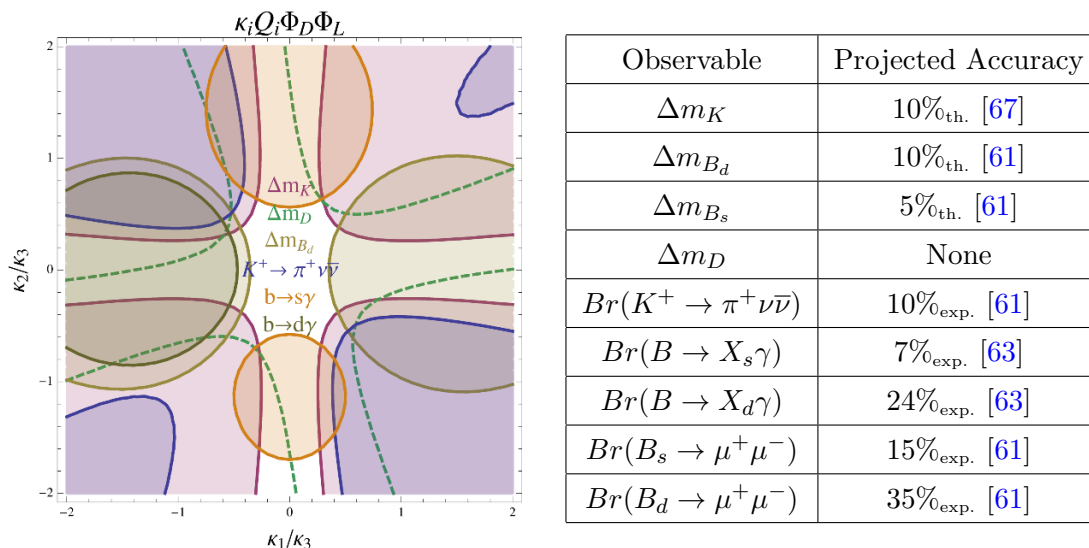


Figure 9. Future projections for flavor constraints on EGMSB in the ~ 3 –5 year range. In the figure, the errors have been updated to the fractional values in the accompanying table, and the central values have set equal between theory and experiment. Depicted is the same Q -class model from figure 8. The dashed green line, which is unchanged from figure 8, indicates where the EGMSB contribution to Δm_D is equal to the measured value. Observables not shown are not constraining.

4.3 Future directions

There are many avenues for future investigations. Here we list a few.

- As we alluded to earlier, the χ FV ansatz is a general paradigm, and it provides a novel and realistic solution to the SUSY flavor problem. This texture, its possible origins, and consequences are worthy of further study [51].
- Another interesting question is that of CP violation in these models. The goal of this work was to focus on the SUSY flavor problem, and so the EGMSB couplings were intentionally assumed to be real. Allowing for $\arg \kappa_i \neq 0$ would give nontrivial contributions to CP-violating observables, some of which (such as ϵ_K and the neutron EDM) are typically extremely constraining. An interesting question is to what extent the χ FV texture protects EGMSB models from the SUSY CP problem before there is any conflict with data. We plan to study this in detail in an upcoming work [35].
- The possibility of heavily mixed squarks allows for very interesting collider signatures [68–71]. These EGMSB models provide a flavor-safe proof-of-concept motivation for experimental searches at ATLAS and CMS.
- While we made some effort to ensure that the flavor-violating A -terms do not destabilize the vacuum (see the discussion in appendix A), it would be interesting to study in more detail the vacuum stability of these flavor-violating EGMSB models, along the lines of [72].

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A Details of the deformation and parameter scan

In this appendix, we detail our parameter scan. In particular, we specify the deformation about the best points in [20] that we use. Our type I models are characterized by the following parameter space:

$$(\kappa_1, \kappa_2, \kappa_3, \Lambda/M, \Lambda) \tag{A.1}$$

In [20], the models were studied at the $\kappa_1 = \kappa_2 = 0$ point. For each κ_3 and Λ/M , Λ (which sets the overall scale of the superpartner spectrum) was increased until $m_h = 125$ GeV was achieved. Then the fine-tuning of the point was estimated and the region of least fine-tuning in $(\kappa_3, \Lambda/M)$ space was identified. Our aim is to investigate these regions in the presence of nontrivial flavor violation, parametrized by κ_1 and κ_2 . A full optimization of the fine-tuning would involve a five-dimensional scan in the $(\kappa_1, \kappa_2, \kappa_3, \Lambda/M, M)$ parameter space. Such an endeavor is not computationally feasible and moreover is unnecessary. The main question we intend to explore is how qualitatively dangerous the flavor-physics contributions are in these EGMSB models. In principle, this question can be answered using a simpler scan where we essentially fix $(\kappa_3, \Lambda/M, M)$ as in [20] and then vary κ_1, κ_2 . However, there are a number of subtleties to take into account when doing so, most of which stem from the fact that the points from [20] tended to have light stops as a result of a cancellation (this is unsurprising due to the tuning involved).

- We must be careful to choose $(\kappa_3, \Lambda/M)$ such that the point at $\kappa_1 = \kappa_2 = 0$ is unconstrained by Run I searches. In practice, to get a model that is not constrained by the LHC and has a good tuning value, we will lower $\frac{\Lambda}{M}$ away from the least-tuned point identified in [20]. This increases the fine-tuning required in the models only by about 10%. The parameters we choose (at the origin) are $\kappa_3 = 0.858$, $N = 6$, $\frac{\Lambda}{M} = 0.347$, and $M = 312$ TeV for I.9, and $\kappa_3 = 0.908$, $N = 3$, $\frac{\Lambda}{M} = 0.290$, and $M = 350$ TeV for I.13. The spectra are shown in figure 1.
- Despite our focus on CP conserving observables in this work, there could, in principle, be significant constraints from ϵ_K introduced solely through the CP violating phase of the CKM.¹⁰ In order to avoid this constraint, in our Q-class models, we

¹⁰We thank W. Altmannshofer for bringing this to our attention.

choose κ_1 , κ_2 and κ_3 to align with the low-energy down, strange and bottom quark, respectively. Thus in our choice of interaction basis, the down Yukawa is diagonal, but the up Yukawa is multiplied by V_{CKM} . We ignore the small differences with our previous work where the alignment was with the top quark. The U -class models retain alignment with the up-type quarks.

- Turning on κ_1 and κ_2 while holding fixed the other parameters can significantly modify the spectrum, as shown in (2.6). At the least finely-tuned points, where a cancellation results in a stop lighter than the other squarks, this can either lead to stop tachyons, or it can lead to the stops being so heavy that the hypercharge tadpole contribution to the RG running quickly drives the sleptons tachyonic. To avoid these undesirable features, as we turn on $\kappa_{1,2}$, we fix Λ , but vary M , so that the one-loop contribution adapts in such a way that the lightest squark eigenvalue of the LL block in Q -class or RR block in U -class models is held fixed. This deformation enables us to prevent the interesting flavor-violation from vanishing and to maintain the squark masses at sensible values.
- Additionally, as can be seen from (2.6), if we were to fix κ_3 and turn on $\kappa_{1,2}$, then the “net” A -term would increase. The “net” A -term in the type I models is aligned with the lightest eigenvalue direction of the rank 1 block, i.e., in Q -class models,

$$\mathcal{L} \supset -\frac{\kappa\kappa_3 y_t \Lambda}{16\pi^2} H_u \tilde{Q}_S \tilde{t}_R. \tag{A.2}$$

The A -terms in all orthogonal squark directions vanish in the third-generation dominant limit. Increasing A/m_S is potentially dangerous for vacuum stability [72]. In order to avoid these issues, we require that this “net” A -term remains constant. To achieve this we fix $\kappa_3\kappa$ everywhere in the parameter space so that as we increase κ_1 and κ_2 , κ_3 decreases to compensate.

- Finally, this deformation of $\kappa_{1,2,3}$ will also affect the Higgs mass. The one-loop corrections to the Higgs mass in the presence of general flavor violation have been computed in [73, 74], and our preliminary studies of these suggest that m_h may drift down by several GeV as we move out in $\kappa_{1,2}$. However, the two-loop corrections are not yet known, and if the usual non-flavor-violating MSSM is any guide, these are likely to be important for an accurate determination of the Higgs mass. While it would be interesting to study this further, it is beyond the scope of the work. At the very least, one could imagine increasing the overall scale of the superpartners in order to compensate for any decrease in the Higgs mass. This would only serve to further alleviate the flavor constraints, so the qualitative value of χ_{FV} to these EGMSB models is unaffected.

Across a grid in (κ_1, κ_2) constructed via this deformation of the benchmark point, we generate the soft spectrum at the messenger scale. All of the couplings and soft masses are then evolved down from the messenger scale to the SUSY scale using the fully general

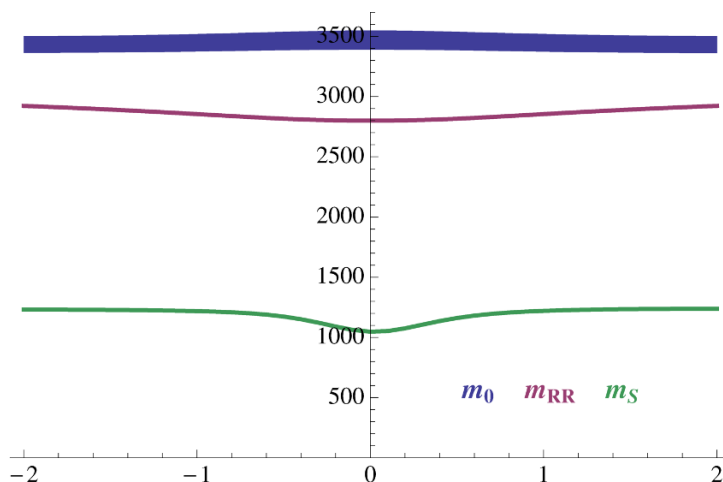


Figure 10. The spectra of squark masses (in GeV) for model I.9 as we move in $\kappa_1 = \kappa_2$. Aside from the three lightest eigenvalues, $m_{S,up} = m_{S,down}$, and m_{RR} , the other nine squark eigenvalues all fall in the range of the blue band near 3.5 TeV. The width of the band is largely due to the SU(2) GMSB contributions to the left-handed squarks.

3×3 MSSM β -functions.¹¹ At the SUSY scale, we apply the BMPZ QCD threshold corrections to the squarks and gluinos [75]. Finally, we use our new `Mathematica` package `FormFlavor` to compute the Wilson coefficients, RG evolve these coefficients down to the scale of interest (e.g., m_b) for each flavor observable, and compute the contributions to each flavor observable there. (The details of `FormFlavor` are briefly explained in appendix B and will be further fleshed out when the package is made publicly available [39].)

B Details of FormFlavor

`FormFlavor` is a general, modular flavor package written in `Mathematica` that takes general MSSM spectra as inputs and calculates a variety of flavor observables *in situ* starting from the Feynman rules. The package is flexible — new flavor observables can be straightforwardly added to it and, although currently implemented only for the MSSM, it can be extended to other models readily. Also, since processes are calculated in a uniform manner from first principles (rather than hardcoding formulas from the literature), the reliability of the code is greatly enhanced.

In `FormFlavor`, the process under consideration is specified in `FeynArts` [36] and all the relevant topologies for the various sparticle mediators are generated there. These topologies are then converted into their respective amplitudes via `FormCalc` [37]. The `FormCalc` output is analytically transformed into a particular Wilson operator basis for each process of interest, and the Wilson coefficients are extracted. Having performed this once for an observable the analytic expressions may be written to a process file and used in subsequent runs. The Wilson coefficients may then be piped into functions that calculate the flavor observables.

¹¹For simplicity, we do not run back and forth between the IR and UV to better specify the scale.

The main components of `FormFlavor` are as follows:

- Automated generation of one-loop diagrams for flavor processes (using `FeynArts`)
- Automated calculation of Wilson coefficients from Feynman diagrams (using `FormCalc`)
- Library of analytic one-loop integral functions
- Routines for general MSSM spectrum input (SLHA2 [76] compatible)
- Routines for converting Wilson coefficients to flavor observables. (This step contains hardcoded formulas from the literature and is not MSSM specific.) The following processes are currently implemented:

CP-conserving observables

- * Meson Mixings [52, 77]: ΔM_K [78], ΔM_D [79], ΔM_{B_d} and ΔM_{B_s} [40, 80]
- * Leptonic decays: $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ [41, 81, 82], $B_s \rightarrow \mu^+ \mu^-$ and $B_d \rightarrow \mu^+ \mu^-$ [46, 55–57]
- * Radiative Decays: $b \rightarrow s \gamma$ [42, 83], $b \rightarrow d \gamma$ [44, 84]

CP-violating observables

- * Meson Mixings: ϵ_K
 - * Leptonic decays: $K^0 \rightarrow \pi^0 \nu \bar{\nu}$
 - * Electric dipole moments: neutron-EDM
- MSSM RGE — Full three-family MSSM renormalization group evolution, including arbitrary CP-phases, implemented with the aid of `SARAH` [85]

This powerful flavor package will be released to the public and described in more detail in an upcoming work [39].

C Soft terms for type I squark models

In this appendix, we will specialize the fully general EGMSB formulas from [20] to the case of the type I squark models with general flavor-violating couplings κ_i . These were quoted in (2.6) for the Q -class models,

$$W_{\text{EGMSB}} = \kappa_i Q_i \sum_{A=1}^N \Phi_A \tilde{\Phi}_A \tag{C.1}$$

and we again focus on this case here. To translate to U -class models, one must take $U \leftrightarrow Q$, $y_u \rightarrow y_u^\dagger$, $A_{\tilde{u}} \rightarrow A_{\tilde{u}}^\dagger$ and $K \rightarrow K^T$ in all expressions.

Before we begin, we must address a convention difference between this paper and that of [20] concerning the treatment of the right-handed quark chiral superfields \bar{U} , \bar{D} . In this paper, the Yukawas, squark mass matrices and A -terms are defined as:

$$\begin{aligned} W &\supset H_u \bar{U}_i y_{u,ij} Q_j + H_d \bar{D}_i y_{d,ij} Q_j \\ \mathcal{L} &\supset \tilde{Q}_i^* \delta m_{Q,ij}^2 \tilde{Q}_j + \tilde{U}_i^* \delta m_{U,ij}^2 \tilde{U}_j + \tilde{D}_i^* \delta m_{D,ij}^2 \tilde{D}_j \\ &\quad + \left(H_u \tilde{U}_i^* A_{\tilde{u},ij} \tilde{Q}_j + H_d \tilde{D}_i^* A_{\tilde{d},ij} \tilde{Q}_j + \text{h.c.} \right) \end{aligned} \quad (\text{C.2})$$

Thus, the Yukawas, A -terms, $m_{\tilde{U}}^2$ and $m_{\tilde{D}}^2$ are all transposed relative to those of [20]. Note that in these conventions, \tilde{U}^* , \tilde{D}^* are the scalar components of the *chiral* supermultiplets \bar{U} , \bar{D} .

The starting point of our derivation of (2.6) is eq. (2.19) in [20], which we repeat here for convenience:

$$\begin{aligned} A_{ab} &= -\frac{1}{32\pi^2} d_a^{BC} \lambda_{aBC}^* \lambda_{bBC} \Lambda \\ \delta m_{ab}^2 &= \frac{1}{256\pi^4} \left(d_a^{BC} d_b^{CD} \lambda_{aBC}^* \lambda_{bCE} \lambda_{cBD} \lambda_{cDE}^* + \frac{1}{4} d_a^{BC} d_b^{DE} \lambda_{aBC}^* \lambda_{cBC} \lambda_{cDE}^* \lambda_{bDE} \right. \\ &\quad \left. - \frac{1}{2} d_a^{cd} d_c^{BC} y_{acd} y_{bde} \lambda_{cBC} \lambda_{eBC}^* - d_a^{BC} C_r^{aBC} g_r^2 \lambda_{aBC}^* \lambda_{bBC} \right) \Lambda^2 \end{aligned} \quad (\text{C.3})$$

Here a, b, \dots run over all MSSM fields (including all gauge and flavor degrees of freedom). Meanwhile A, B, \dots run over all messenger fields similarly. In this appendix, we will use i, j, \dots to denote MSSM flavor indices.

From (C.3), we immediately obtain the bilinear A -terms after summing over the N messenger multiplets and substituting $\lambda_{aBC} \rightarrow \kappa_i$ and $d_a^{BC} \rightarrow d_Q \propto N$

$$A_{Q,ij} = -\frac{1}{16\pi^2} d_Q \kappa_i^* \kappa_j \Lambda \quad (\text{C.4})$$

This becomes the trilinear A -term $A_{\tilde{u}}$ used in the paper via:

$$A_{Q,ij} F_{Q_i}^\dagger \tilde{Q}_j \rightarrow y_{u,ik} A_{Q,kj} H_u \tilde{U}_i^* \tilde{Q}_j \equiv A_{\tilde{u},ij} H_u \tilde{U}_i^* \tilde{Q}_j \quad (\text{C.5})$$

Next let's consider the EGMSB contributions to the soft mass-squareds, starting with $m_{\tilde{Q}}^2$. From (C.3), we obtain:

$$\begin{aligned} \delta m_{Q,ij}^2 &= \frac{1}{256\pi^4} \left(d_Q d_\phi \kappa_i^* \kappa_j \kappa_k \kappa_k^* + d_Q^2 \kappa_i^* \kappa_k^* \kappa_k \kappa_j - 2d_Q C_r g_r^2 \kappa_i^* \kappa_j \right) \Lambda^2 \\ &= \frac{d_Q}{256\pi^4} \left((d_\phi + d_Q) \kappa^2 - 2C_r g_r^2 \right) K_{ij} \Lambda^2 \end{aligned} \quad (\text{C.6})$$

In the first line, we have introduced $d_B^{cD} \rightarrow \frac{1}{2} d_\phi$, and we have used the fact that Q is the only MSSM field coupling to the messengers to set $d_a^{cd} d_c^{BC}$ to zero. Additionally, as Q couples directly to the messengers, there is a one-loop term suppressed by $\frac{\Lambda^2}{M^2}$ [13],

$$\delta \left(m_{Q,ij}^{1\text{-loop}} \right)^2 = -\frac{16\pi^2}{3} h \left(\frac{\Lambda}{M} \right) \frac{\Lambda^2}{M^2} \frac{d_Q K_{ij} \Lambda^2}{256\pi^4} \quad (\text{C.7})$$

where $h(x)$ is a loop function given by

$$h(x) = \frac{3}{x^4} \left((x-2) \ln(1-x) - (x+2) \ln(1+x) \right) = 1 + \frac{4}{5}x^2 + \mathcal{O}(x^4). \quad (\text{C.8})$$

The EGMSB contribution to $m_{H_u}^2$ is much simpler. Here only the third term of δm_{ab}^2 in (C.3) contributes:

$$\delta m_{H_u}^2 = -\frac{3d_Q}{256\pi^4} y_{u,ij}^* y_{u,ik} \kappa_j \kappa_k^* \Lambda^2 = -\frac{3d_Q}{256\pi^4} \text{Tr} \left[y_u K y_u^\dagger \right] \Lambda^2 \quad (\text{C.9})$$

where we have used $d_H^{QU} = 3$.

Lastly, the EGMSB contribution to m_U^2 also comes from just the third term of δm_{ab}^2 :

$$\delta m_{U,ij}^2 = -\frac{d_U^{QH} d_Q}{256\pi^4} y_{u,ik}^* y_{j\ell} \kappa_k \kappa_\ell^* \Lambda^2 = -\frac{d_U^{QH} d_Q}{256\pi^4} \left(y_u K y_u^\dagger \right)_{ji} \Lambda^2 \quad (\text{C.10})$$

where $d_U^{QH} = 2$. Taking into account the need to transpose m_U^2 to translate between the conventions of [20] and those of this paper, we obtain the correct result quoted in (2.6).

D Formulas for flavor observables

In this appendix, we collect formulas from the literature for the various flavor observables considered in this work. Along the way, we will streamline the different notations scattered throughout the literature into a uniform convention.

The uniform operator basis we will use was introduced in the text; we repeat it here for convenience. For dimension 5, we have:

$$\mathcal{O}_A^M(f_1, f_2) = e \bar{f}_1 \sigma^{\mu\nu} P_M f_2 F_{\mu\nu} \quad (\text{D.1})$$

$$\mathcal{O}_G^M(f_1, f_2) = g \bar{f}_1 \sigma^{\mu\nu} P_M f_2 G_{\mu\nu} \quad (\text{D.2})$$

For dimension 6, we have:

$$\mathcal{O}_S^{MN}(f_1, f_2, f_3, f_4) = (\bar{f}_1 P_M f_2) (\bar{f}_3 P_N f_4) \quad (\text{D.3})$$

$$\mathcal{O}_V^{MN}(f_1, f_2, f_3, f_4) = (\bar{f}_1 \gamma^\mu P_M f_2) (\bar{f}_3 \gamma_\mu P_N f_4) \quad (\text{D.4})$$

$$\mathcal{O}_T^{MN}(f_1, f_2, f_3, f_4) = (\bar{f}_1 \sigma^{\mu\nu} P_M f_2) (\bar{f}_3 \sigma_{\mu\nu} P_N f_4) \quad (\text{D.5})$$

where $M, N = L, R$. $P_R = \frac{1}{2}(1 + \gamma_5)$, $P_L = \frac{1}{2}(1 - \gamma_5)$ are projection operators, $\sigma^{\mu\nu} = \frac{1}{2}[\gamma^\mu, \gamma^\nu]$ and if f_i carry color indices, they are contracted within a bilinear factor.

The general effective Hamiltonian is then:

$$H_{\text{eff}}(f_i) = \sum C_X^M(f_i) \mathcal{O}_X^M(f_i) + \sum C_X^{MN}(f_i) \mathcal{O}_X^{MN}(f_i) \quad (\text{D.6})$$

where the sums runs over a complete basis of independent operators.

Meson	m_X (GeV)	f_X (GeV)	R_X	B_V^{LL}	B_V^{LR}	B_S^{LR}	B_S^{LL}	B_T^{LL}	$C_{V,SM}^{LL} _{\mu=m_b}$ (GeV ⁻²)
Δm_K	0.4976	0.160	24.3	0.56	0.85	1.08	0.62	0.43	—
Δm_D	1.8645	0.209	3.20	0.76	0.97	0.95	0.64	0.39	—
Δm_{B_d}	5.2796	0.191	1.65	0.84	1.47	0.95	0.72	0.61	$(2.34 - 2.20i) \times 10^{-12}$
Δm_{B_s}	5.3668	0.228	1.65	0.88	1.57	0.93	0.73	0.62	$(6.96 - 0.26i) \times 10^{-11}$

Table 2. Properties of the meson mixing used in this work. f_X and $B_{V,X}^{LL}$ come from the FLAG review [77]. The other four non-perturbative B -parameters are taken from several sources: for Δm_K from [78] (at $\mu = 2$ GeV), Δm_D from [79] (at $\mu = 3$ GeV, rescaled for a common R_D , and converted to our basis using $B_T^{LL} = \frac{5}{3}B_2 - \frac{2}{3}B_3$), and for B_s and B_d from [80] (at $\mu = m_b = 4.2$ GeV and converted to our basis). $C_{V,SM}^{LL}|_{\mu=m_b}$ are the FormFlavor values in the CKM basis of the PDG [47].

D.1 Meson Mixing Δm_X

This corresponds to a $\Delta F = 2$ effective Hamiltonian, with $f_1 = f_3 = q_1$, and $f_2 = f_4 = q_2$, with $(q_1, q_2) = (s, d), (c, u), (b, d), (b, s)$ for K, D, B_d and B_s respectively. The quantities relevant for mixing are derived from the effective Hamiltonian as

$$\begin{aligned} \langle \bar{X} | H_{\text{eff}} | X \rangle &\equiv M_{X,12} - \frac{i}{2} \Gamma_{X,12} \\ \langle X | H_{\text{eff}} | \bar{X} \rangle &= M_{X,12}^* - \frac{i}{2} \Gamma_{X,12}^* \end{aligned} \quad (\text{D.7})$$

$M_{X,12}$ and $\Gamma_{X,12}$ are, respectively, the dispersive and absorptive parts of $\langle \bar{X} | H_{\text{eff}} | X \rangle$. In terms of these quantities, the mass splitting is given by:

$$\Delta m_X = 2 \text{Re} \left[\left(M_{X,12} - \frac{i}{2} \Gamma_{X,12} \right) \sqrt{\frac{M_{X,12}^* - \frac{i}{2} \Gamma_{X,12}^*}{M_{X,12} - \frac{i}{2} \Gamma_{X,12}}} \right] \quad (\text{D.8})$$

For $X = B_d$ and B_s where $\Gamma_{X,12} \ll M_{X,12}$, this is well approximated by

$$\Delta m_X \approx 2|M_{X,12}| \approx 2|\langle \bar{X} | H_{\text{eff}} | X \rangle| \quad (\text{D.9})$$

For $X = K$ and D , where experimentally one finds that $M_{X,12}/\Gamma_{X,12}$ is approximately real, and $M_{X,12}$ and $\Gamma_{X,12}$ are both predicted to be approximately real in the standard CKM convention (where the CPV phase is primarily in V_{td} and V_{ub}), one has to a good approximation

$$\Delta m_X \approx 2 \text{Re} M_{X,12} \approx 2 \text{Re} \langle \bar{X} | H_{\text{eff}} | X \rangle \quad (\text{D.10})$$

Finally, the short-distance part of the matrix elements in (D.9) and (D.10) (which is all that is relevant for the B_q systems and for new physics) are given by

$$\begin{aligned} \langle \bar{X} | H_{\text{eff}} | X \rangle &= \frac{m_X f_X^2}{24} \left(8B_V^{LL} (C_V^{LL} + C_V^{RR}) - R_X \left[4B_V^{LR} C_V^{LR} - 6B_S^{LR} C_S^{LR} \right. \right. \\ &\quad \left. \left. + 5B_S^{LL} (C_S^{LL} + C_S^{RR}) + 12B_T^{LL} (C_T^{LL} + C_T^{RR}) \right] \right) \end{aligned} \quad (\text{D.11})$$

where,

$$R_X = \left(\frac{m_X}{m_{q_1} + m_{q_2}} \right)^2, \quad (\text{D.12})$$

and the B -parameters are non-perturbative corrections that have been computed on the lattice. In the Standard Model, only $C_V^{LL} \neq 0$. These values and the other parameters relevant for meson mixing are shown in table 2.

D.2 $K^\pm \rightarrow \pi^\pm \nu \bar{\nu}$

This corresponds to a dimension 6 effective Hamiltonian with $f_1 = s$, $f_2 = d$, $f_3 = f_4 = \nu_\ell$. The branching ratios for $K^\pm \rightarrow \pi^\pm \nu \bar{\nu}$ are given by [81],

$$\text{BR}(K^\pm \rightarrow \pi^\pm \nu \bar{\nu}) = \frac{c_+}{3} v^4 \sum_{\ell=e,\mu,\tau} |C_{V,\ell}^{LL} + C_{V,\ell}^{RL}|^2 \quad (\text{D.13})$$

where $v = 246$ GeV and

$$c_+ = \frac{3 r_{K^+}}{2 |V_{us}|^2} \text{BR}(K^\pm \rightarrow \pi^0 e^\pm \nu) = 1.35. \quad (\text{D.14})$$

Here, the branching ratio has been included to remove dependence on the hadronic matrix element, and $r_{K^+} = 0.901$ contains isospin violating quark mass effects and electroweak corrections computed in [86].

In the standard model, the contributions from top loops are have no sensitivity to different generations, but the charm loop contributions do. In particular, the charm contributions are the same for e and μ , but differ for τ . Thus, the SM contribution can be expressed as,

$$C_{V,SM,\ell}^{LL} = \frac{\alpha_2}{\pi v^2} \left(\lambda_c X_c^\ell + \lambda_t X_t \right) \quad (\text{D.15})$$

where $\lambda_i = V_{is}^* V_{id}$, $X_t = 1.469$ [41], $X_c^e = X_c^\mu = 1.055 \times 10^{-3}$, and $X_c^\tau = 7.01 \times 10^{-4}$ [81].

In the absence of lepton flavor-violating new physics effects, the contributions from new physics are the same across lepton generations, i.e., $C_{V,NP,\ell}^{XY} = C_{V,NP,\ell'}^{XY} \equiv C_{V,NP}^{XY}$. We can then give a simpler form to equation (D.13)

$$\text{BR}(K^\pm \rightarrow \pi^\pm \nu \bar{\nu}) = c_+ v^4 |C_V^{LL} + C_V^{RL}|^2 \quad (\text{D.16})$$

where $C_V^{XY} = C_{V,NP}^{XY} + C_{V,SM}^{XY}$, with

$$\begin{aligned} C_{V,SM}^{RL} &= 0 \\ C_{V,SM}^{LL} &= \frac{\alpha_2}{\pi v^2} (\lambda_c P_c + \lambda_t X_t) = (-12.2 + 3.6i) \times 10^{-11} \text{ GeV}^{-2} \\ P_c &= \left(\frac{2}{3} X_{NL}^e + \frac{1}{3} X_{NL}^\tau \right) \sim 9.37 \times 10^{-4}. \end{aligned} \quad (\text{D.17})$$

Here, P_c is the charm contribution averaged over the different neutrino flavors. Using P_c simplifies the expression, but reduces the standard model charm contribution by about 3%. This represents only a 0.3% decrease in the overall SM contribution, which is completely negligible when compared to the theoretical uncertainty. Most importantly, the interference effects with new physics are properly captured under this simplification.

D.3 $b \rightarrow s\gamma$ and $b \rightarrow d\gamma$

These observables correspond to dimension 5 effective Hamiltonians with $f_1 = b$ and $f_2 = s$ or d . The branching ratio is given by,

$$\text{BR}(b \rightarrow q\gamma) = c_\gamma v^2 \left(|C_A^L|^2 + |C_A^R|^2 \right) \quad (\text{D.18})$$

where [87]

$$\begin{aligned} c_\gamma &= (8\pi^2)^2 \frac{6}{\pi} \frac{\text{BR}(b \rightarrow X_c e \nu)_{\text{EXP}}}{\Phi |V_{cb}|^2} \frac{v^2}{m_b^2} \alpha_{\text{EM}} = 3.3 \times 10^7 \\ \Phi &= \left| \frac{V_{ub}}{V_{cb}} \right|^2 \frac{\text{BR}(b \rightarrow X_c e \nu)}{\text{BR}(b \rightarrow X_u e \nu)} = 0.58 \end{aligned} \quad (\text{D.19})$$

where the branching ratio to charm decays is used to remove sensitivity to the hadronic matrix element, and the Φ factor is introduced to account for the nontrivial phase space factor in the compared branching ratio (due mostly to the charm quark mass).

Only C_A^L contains a standard model contribution, which is,

$$[C_{A,\text{SM}}^L]_{b \rightarrow s\gamma} = \frac{V_{ts}^* V_{tb} m_b}{8\pi^2 v^2} X_{\text{SM}} \sim (1.3 + 0.03i) \times 10^{-8} \text{ GeV}^{-1} \quad (\text{D.20})$$

where $X_{\text{SM}} = -0.3736$.

The observable $b \rightarrow d\gamma$ is defined completely analogously. The only difference is that the the standard model contribution differs. In particular,

$$[C_{A,\text{SM}}^L]_{b \rightarrow d\gamma} = \frac{V_{td}^* V_{tb} m_b}{8\pi^2 v^2} X_{\text{SM}} \sim -(2.4 + 1.1i) \times 10^{-9} \text{ GeV}^{-1} \quad (\text{D.21})$$

with the same $X_{\text{SM}} = -0.3736$ as above.

D.4 $B_s \rightarrow \mu^+ \mu^-$ and $B_d \rightarrow \mu^+ \mu^-$

This corresponds to a dimension 6 effective Hamiltonian with $f_1 = b$, $f_2 = s, d$ and $f_3 = f_4 = \mu$. The branching ratio for $B_{s,d} \rightarrow \mu^+ \mu^-$ is [55, 56]:

$$\text{BR}(B_i \rightarrow \mu^+ \mu^-) = X_i \left\{ \left(1 - \frac{4m_\mu^2}{m_{B_i}^2} \right) |F_S^{(i)}|^2 + |F_P^{(i)} + F_A^{(i)}|^2 \right\} \quad (\text{D.22})$$

where

$$\begin{aligned} X_i &= \frac{f_{B_i}^2}{128\pi m_{B_i} \Gamma_{B_i}} \sqrt{1 - \frac{4m_\mu^2}{m_{B_i}^2}} \implies X_s = 5.36 \times 10^7 \text{ and } X_d = 3.97 \times 10^7, \\ F_S^{(i)} &= \frac{m_{B_i}^3}{m_b + m_i} (C_S^{LL} + C_S^{LR} - C_S^{RR} - C_S^{RL}), \\ F_P^{(i)} &= \frac{m_{B_i}^3}{m_b + m_i} (-C_S^{LL} + C_S^{LR} - C_S^{RR} + C_S^{RL}), \\ F_A^{(i)} &= 2m_{B_i} m_\mu (C_V^{LL} - C_V^{LR} + C_V^{RR} - C_V^{RL}), \end{aligned} \quad (\text{D.23})$$

and $\Gamma_{B_d} \equiv \tau_{B_d}^{-1} = 4.33 \times 10^{-13} \text{ GeV}$ and $\Gamma_{B_s} \equiv \tau_{B_s}^{-1} = 4.49 \times 10^{-13} \text{ GeV}$. The three-loop Standard Model contribution to $B_i \rightarrow \mu^+ \mu^-$ can be expressed as [57]:

$$F_{A,\text{SM}}^{(i)} = 0.4802 \times \frac{4\alpha_2 V_{tb}^* V_{ti} m_\mu m_{B_i}}{\pi v^2}; \quad F_{S,\text{SM}}^{(i)} = F_{P,\text{SM}}^{(i)} = 0 \quad (\text{D.24})$$

so that $F_{A,\text{SM}}^{(d)} = (1.5 - 0.6i) \times 10^{-9}$ and $F_{A,\text{SM}}^{(s)} = (-7.9 - 0.1i) \times 10^{-9}$

E Rank 1 χ FV loop functions

We compile some of the loop functions for our rank 1 χ FV approximation below.

Starting with meson-mixing, we have,

$$\begin{aligned} f_{\tilde{g}}^{\Delta M, \text{box}}(x_g, x_q) &= \frac{2x_q \log x_q (11x_q^3 + 6x_g x_q^2 - 2x_g^2 x_q - 13x_g x_q - 2x_g^2)}{(x_q - 1)(x_q - x_g)^3} \\ &+ \frac{2x_q x_g \log x_g (x_q - 1)^2 (2x_q x_g + 13x_q - 17x_g^2 + 2x_g)}{(x_g - 1)^3 (x_q - x_g)^3} \\ &- \frac{x_q (19x_q^2 x_g + 11x_q^2 + 3x_q x_g^2 - 74x_q x_g + 11x_q + 8x_g^3 + 3x_g^2 + 19x_g)}{(x_g - 1)^2 (x_q - x_g)^2} \end{aligned} \quad (\text{E.1})$$

Here, $x_q = m_S^2/m_0^2$ and $x_g = m_S^2/m_{\tilde{g}}^2$; where m_0 and m_S are the heaviest and lightest squark mass eigenvalues.

For $K^\pm \rightarrow \pi^\pm \nu \bar{\nu}$, there are two contributions,

$$f_{\tilde{\chi}^\pm}^{K \rightarrow \pi \nu \nu} (x_\ell, x_2, x_\mu) = f_{\tilde{\chi}^\pm}^{K \rightarrow \pi \nu \nu, \text{box}} (x_\ell, x_2) + f_{\tilde{\chi}^\pm}^{K \rightarrow \pi \nu \nu, \text{peng}} (x_2, x_\mu). \quad (\text{E.2})$$

In practice, $f_{\tilde{\chi}^\pm}^{K \rightarrow \pi \nu \nu, \text{peng}} (x_\ell, x_2)$ is numerically only $\mathcal{O}(10\%)$ of $f_{\tilde{\chi}^\pm}^{K \rightarrow \pi \nu \nu, \text{box}} (x_\ell, x_2)$, so, out of simplicity, we only present the box contribution for interpreting our results,

$$\begin{aligned} f_{\tilde{\chi}^\pm}^{K \rightarrow \pi \nu \nu, \text{box}} (x_\ell, x_2) &= 6 \left[\frac{x_2 (x_2^2 - x_\ell) \log x_2}{(x_2 - 1)^2 (x_2 - x_\ell)^2} - \frac{x_2}{(x_2 - 1)(x_2 - x_\ell)} \right. \\ &\quad \left. - \frac{x_2 x_\ell \log x_\ell}{(x_2 - x_\ell)^2 (x_\ell - 1)} \right] \end{aligned} \quad (\text{E.3})$$

Here $x_\ell = m_{\tilde{l}}^2/m_S^2$, $x_2 = M_2^2/m_S^2$ and $x_\mu = \mu^2/m_S^2$. We have also dropped the neutralino boxes and Z -penguins that contribute at the $\mathcal{O}(10\%)$ level or below.

For $b \rightarrow q \gamma$

$$\begin{aligned} f_{\tilde{\chi}^\pm}^{b \rightarrow s/d \gamma, \text{peng}} (x_\mu, x_2) &= \frac{6\sqrt{x_\mu x_2}}{11} \left[\frac{5x_2 x_\mu - 7x_\mu - 7x_2 + 9}{(x_2 - 1)^2 (x_\mu - 1)^2} + \frac{2(2x_2 - 3) \log x_2}{(x_2 - 1)^3 (x_2 - x_\mu)} \right. \\ &\quad \left. - \frac{2(2x_\mu - 3) \log x_\mu}{(x_2 - x_\mu)(x_\mu - 1)^3} \right] \end{aligned} \quad (\text{E.4})$$

Here $x_\mu = \mu^2/m_S^2$, $x_2 = M_2^2/m_S^2$ and m_S is the lightest up squark mass eigenvalue.

For $B_q \rightarrow \mu^- \mu^+$

$$f_{\tilde{g}}^{B_q \rightarrow \mu^+ \mu^-, \text{h-peng}}(x_q, x_g) = \sqrt{x_g x_q} \left[\frac{(x_q - 1) x_g \log x_g}{(x_g - 1)^2 (x_g - x_q)} + \frac{1}{(x_g - 1)} - \frac{x_q \log x_q}{(x_q - 1)(x_g - x_q)} \right] \quad (\text{E.5})$$

$$f_{\tilde{\chi}}^{B_q \rightarrow \mu^+ \mu^-, \text{h-peng}}(x_q, x_\mu, x_2) = \sqrt{x_2} \left[\frac{x_2 (x_q - 1) \log x_2}{(x_2 - 1)(x_2 - x_\mu)(x_2 - x_q)} \right. \\ \left. + \frac{x_\mu (x_q - 1) \log x_\mu}{(x_\mu - x_2)(x_\mu - 1)(x_\mu - x_q)} + \frac{x_q \log x_q}{(x_\mu - x_q)(x_2 - x_q)} \right] \quad (\text{E.6})$$

Here $x_g = m_g^2/m_0^2$, $x_\mu = \mu^2/m_0^2$, $x_2 = M_2^2/m_0^2$ and $x_q = m_S^2/m_0^2$; m_S and m_0 in this context are the lightest and heaviest squark mass eigenvalues.

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