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# Hunting for heavy majorana neutrinos with lepton number violating signatures at LHC

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ABSTRACT: The neutrinophilic two-Higgs-doublet model ( $\nu$ 2HDM) provides a natural way to generate tiny neutrino mass from interactions with the new doublet scalar  $\Phi_{\nu}$  $(H^{\pm}, H, A)$  and singlet neutrinos  $N_R$  of TeV scale. In this paper, we perform detailed simulations for the lepton number violating (LNV) signatures at LHC arising from cascade decays of the new scalars and neutrinos with the mass order  $m_{N_R} < m_{\Phi_{\nu}}$ . Under constraints from lepton flavor violating processes and direct collider searches, their decay properties are explored and lead to three types of LNV signatures:  $2\ell^{\pm}4j + \not{\!\!\!E}_T$ ,  $3\ell^{\pm}4j + \not{\!\!\!E}_T$ , and  $3\ell^{\pm}\ell^{\mp}4j$ . We find that the same-sign trilepton signature  $3\ell^{\pm}4j + \not{\!\!\!E}_T$  is quite unique and is the most promising discovery channel at the high-luminosity LHC. Our analysis also yields the 95% C.L. exclusion limits in the plane of the  $\Phi_{\nu}$  and  $N_R$  masses at 13 (14) TeV LHC with an integrated luminosity of 100 (3000) fb<sup>-1</sup>.

KEYWORDS: Beyond Standard Model, Neutrino Physics, Higgs Physics

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## 1 Introduction

Neutrino mass and mixing provides robust evidence for physics beyond the standard model (SM). Regarding SM as a low energy effective field theory, the tiny neutrino mass can be incorporated by the unique dimension-five Weinberg operator [1]. The high energy scale realization of this operator remains however to be a theoretical puzzle. If it is realized at the tree level, there are three possibilities to do so [2], which correspond exactly to the canonical type I [3–7], type II [8–12], and type III seesaw [13] respectively. The new particles introduced in these seesaws are typically so heavy that they are even beyond the reach of high energy colliders such as the Large Hadron Collider (LHC). To achieve tiny neutrino mass at relatively lower scales, more sophisticated mechanisms have been proposed, such as inverse [14–16] and linear seesaws [17–19], pushing the effective neutrino mass operator to even higher dimensions [20–31], or attributing it to a purely radiative effect [32–111]; see, e.g., refs. [112, 113] for reviews. In these generalized mechanisms, new particles could be at a TeV scale, making their detection in principle possible at LHC.

The demand for very heavy Majorana neutrinos in the type I seesaw arises partly from the difficulty to arrange for a naturally small Yukawa coupling between the light and heavy neutrinos which yields a Dirac mass for neutrinos. To relax the tension, a second Higgs doublet  $\Phi_{\nu}$  is introduced in the so-called neutrinophilic two-Higgs-doublet model ( $\nu$ 2HDM) [114–125]. By assigning a lepton number to  $\Phi_{\nu}$  but not to the Majorana neutrinos  $N_R$ , the Yukawa coupling between them and the SM lepton doublet  $L, \bar{L}\tilde{\Phi}_{\nu}N_R$ , is allowed by the lepton number  $U(1)_L$  symmetry, while the SM Higgs doublet  $\Phi$  is forbidden to couple to L and  $N_R$ . Assuming further  $U(1)_L$  is softly broken by a bilinear term  $\Phi^{\dagger}\Phi_{\nu}$  in the scalar potential,  $\Phi_{\nu}$  develops a naturally small vacuum expectation value (VEV) out of that of  $\Phi$ , so that a tiny neutrino mass becomes possible from a small Dirac mass without requiring a terribly heavy Majorana neutrino. Additional benefits such as a dark matter candidate [126–130] and leptogenesis [126, 131–135] have also been actively explored in the framework of  $\nu$ 2HDM.

The above discussion indicates that  $\nu$ 2HDM can be considered as a type-I-like seesaw, but with relative larger Yukawa couplings [136] among the SM leptons and the new scalar and heavy Majorana neutrinos. The new particles can now be naturally at a TeV scale and could be within the reach of LHC. The usual lepton number violating (LNV) signature for heavy neutrinos at LHC,  $pp \rightarrow \ell^{\pm} N_R \rightarrow \ell^{\pm} \ell^{\pm} jj$ , has been well studied [137–155], albeit with significant fine-tuning between tiny neutrino masses and detectable heavy-light neutrino mixing in type I seesaw (see, e.g., refs. [156, 157] for discussions on fine-tuning issues). In  $\nu$ 2HDM, the direct production of  $\ell^{\pm} N_R$  is still suppressed heavily by the small mixing, but now  $N_R$  could also arise as a decay product of other new particles that can be copiously produced at LHC [158–172]. This is indeed the case when  $N_R$  is lighter than  $\Phi_{\nu}$ which can be pair or associated produced via the Drell-Yan process followed by its decay to  $N_R$  via the Yukawa coupling. The further decay of  $N_R$ ,  $N_R \rightarrow \ell^{\pm} W^{\mp}$ , then triggers various LNV signatures at LHC [173]. The purpose of this work is to investigate these signatures by detailed simulations.

Our paper is organized as follows. In section 2, we introduce  $\nu$ 2HDM and consider constraints from lepton flavor violation (LFV) and direct collider searches. Then in section 3, we study the decay properties of neutrinophilic scalars and heavy neutrinos. The detailed simulation of various LNV signatures at LHC is performed in section 4. Finally, our conclusions are presented in the last section 5.

# 2 Model and constraints

#### 2.1 The model

The  $\nu$ 2HDM was previously suggested in ref. [114]. It introduces a new scalar doublet  $\Phi_{\nu}$  that has the same quantum numbers as the SM Higgs doublet  $\Phi$ , and three right-handed singlet neutrinos  $N_R$ . A global U(1)<sub>L</sub> symmetry is then imposed, under which  $\Phi_{\nu}$  has lepton number L = -1 and  $N_R$  has null lepton number. This then distinguishes between  $\Phi_{\nu}$  and  $\Phi$ : while  $\Phi_{\nu}$  couples to  $N_R$ ,  $\Phi$  couples only to SM fermions, thus avoiding flavor changing neutral currents at tree level in the Yukawa sector. When a soft U(1)<sub>L</sub> breaking term is introduced in the scalar potential, a small VEV of  $\Phi_{\nu}$  can naturally develop, making  $\nu$ 2HDM generically different from the conventional two-Higgs-doublet models [174].

Denoting the two scalar doublets as

$$\Phi = \begin{pmatrix} \phi^+ \\ (v + \phi^{0,r} + i\phi^{0,i})/\sqrt{2} \end{pmatrix}, \ \Phi_\nu = \begin{pmatrix} \phi_\nu^+ \\ (v_\nu + \phi_\nu^{0,r} + i\phi_\nu^{0,i})/\sqrt{2} \end{pmatrix},$$
(2.1)

the scalar potential is

$$V = -m_{\Phi}^{2} \Phi^{\dagger} \Phi + m_{\Phi_{\nu}}^{2} \Phi_{\nu}^{\dagger} \Phi_{\nu} + \frac{1}{2} \lambda_{1} (\Phi^{\dagger} \Phi)^{2} + \frac{1}{2} \lambda_{2} (\Phi_{\nu}^{\dagger} \Phi_{\nu})^{2} + \lambda_{3} (\Phi^{\dagger} \Phi) (\Phi_{\nu}^{\dagger} \Phi_{\nu}) + \lambda_{4} (\Phi^{\dagger} \Phi_{\nu}) (\Phi_{\nu}^{\dagger} \Phi) - (\mu^{2} \Phi^{\dagger} \Phi_{\nu} + \text{H.c.}).$$
(2.2)

Since vanishing of  $\mu^2$  would enhance symmetry, it can be considered as naturally small. Assuming  $m_{\Phi,\Phi_{\nu}}^2 > 0$  and  $\lambda$ s fulfil various bounded-from-below conditions such as  $\lambda_{1,2} > 0$ and  $\lambda_1 \lambda_2 \ge (\lambda_3 + \lambda_4)^2$  [174, 175],  $\Phi_{\nu}$  develops a small VEV out of that of  $\Phi$ :

$$v \simeq \sqrt{\frac{2m_{\Phi}^2}{\lambda_1}}, \ v_{\nu} \simeq \frac{\mu^2 v}{m_{\Phi_{\nu}}^2 + (\lambda_3 + \lambda_4)v^2/2}.$$
 (2.3)

To get some feel about the seesaw-like relation between the VEVs  $v_{\nu}$  and v, we may assume, for instance,  $m_{\Phi_{\nu}} \sim 500 \text{ GeV}$ ,  $\mu^2 \sim 10 \text{ GeV}^2$  to arrive at  $v_{\nu} \sim 10 \text{ MeV}$ . Since  $\mu^2$  is the only source of U(1)<sub>L</sub> breaking, its radiative correction is proportional to itself and only logarithmically sensitive to the cutoff [118]. The hierarchy  $v_{\nu} \ll v$  is thus stable against radiative corrections [132, 176, 177].

After electroweak symmetry breaking, the physical scalars are given by

$$H^{\pm} = \phi^{\pm} \sin\beta - \phi^{\pm}_{\nu} \cos\beta, \qquad \qquad A = \phi^{0,i} \sin\beta - \phi^{0,i}_{\nu} \cos\beta, \qquad (2.4)$$

$$H = \phi^{0,r} \sin \alpha - \phi^{0,r}_{\nu} \cos \alpha, \qquad \qquad h = -\phi^{0,r} \cos \alpha - \phi^{0,r}_{\nu} \sin \alpha, \qquad (2.5)$$

where the mixing angles  $\beta$  and  $\alpha$  are determined by

$$\tan \beta = \frac{v_{\nu}}{v}, \ \tan 2\alpha \simeq 2\frac{v_{\nu}}{v} \frac{-\mu^2 + (\lambda_3 + \lambda_4)vv_{\nu}}{-\mu^2 + \lambda_1 vv_{\nu}}, \tag{2.6}$$

and their masses are

$$m_{H^{\pm}}^2 \simeq m_{\Phi_{\nu}}^2 + \frac{1}{2}\lambda_3 v^2, \ m_A^2 \simeq m_H^2 \simeq m_{H^{\pm}}^2 + \frac{1}{2}\lambda_4 v^2, \ m_h^2 \simeq 2m_{\Phi}^2 = \lambda_1 v^2.$$
 (2.7)

Since  $v_{\nu} \ll v$  in our consideration here,  $\Phi$  is almost identical with the SM Higgs doublet, and h can be regarded as the boson of mass  $m_h = 125$  GeV discovered at LHC [178–180]. For simplicity, we will assume in our numerical analysis a degenerate mass spectrum of  $\Phi_{\nu}$ , i.e.,  $m_{H^{\pm}} = m_H = m_A = m_{\Phi_{\nu}}$ , which can be realized by taking  $\lambda_3 = \lambda_4 = 0$ . This simplification trivially saturates the bounded-from-below conditions mentioned above.

Under the  $U(1)_L$  symmetry, the Yukawa coupling and the Majorana mass terms for  $N_R$  are given by

$$-\mathcal{L}_N = y\overline{L}\widetilde{\Phi}_{\nu}N_R + \frac{1}{2}\overline{N_R^c}m_{N_R}N_R + \text{H.c.}, \qquad (2.8)$$

with  $\tilde{\Phi}_{\nu} = i\sigma_2 \Phi_{\nu}^*$ . The coupling y will be fixed by phenomenologically known neutrino parameters and theoretically expected values for  $v_{\nu}$  and  $m_{N_R}$ . Without loss of generality, we assume the charged leptons and  $N_R$  have been diagonalized. Note that the same Yukawa coupling can also be obtained by imposing a discrete  $\mathbb{Z}_2$  symmetry [121]. But the  $\mathbb{Z}_2$ scenario is found to be in severe tension with the electroweak precision tests, while the  $U(1)_L$  scenario is still viable [181]. Analogously to the type I seesaw, the mass matrix for light neutrinos can be derived approximately from eq. (2.8):

$$m_{\nu} = -\frac{1}{2} v_{\nu}^2 y \ m_{N_R}^{-1} y^T = U_{\text{PMNS}} \ \hat{m}_{\nu} U_{\text{PMNS}}^T, \qquad (2.9)$$

where  $\hat{m}_{\nu} = \text{diag}(m_1, m_2, m_3)$  is the diagonalized neutrino mass matrix and  $U_{\text{PMNS}}$  is the usual Pontecorvo-Maki-Nakagawa-Sakata (PMNS) matrix:

$$U_{\rm PMNS} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{-i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{-i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{-i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{-i\delta} & c_{23}c_{13} \end{pmatrix} \\ \times {\rm diag}(e^{i\varphi_1/2}, 1, e^{i\varphi_2/2}).$$

$$(2.10)$$

Here the shortcuts  $c_{ij} = \cos \theta_{ij}$  and  $s_{ij} = \sin \theta_{ij}$  are used,  $\delta$  is the Dirac phase and  $\varphi_{1,2}$  are the two Majorana phases. In our numerical sampling we will consider neutrino masses in either normal (NH) or inverted hierarchy (IH), and scan randomly the oscillation parameters in the  $2\sigma$  ranges of table I in ref. [182].

The Yukawa coupling matrix y in eq. (2.9) can be solved in terms of neutrino parameters,  $v_{\nu}$ ,  $m_{N_R}$ , and free parameters in the Casas-Ibarra parametrization [183, 184]:

$$y = \frac{\sqrt{2}}{v_{\nu}} U_{\text{PMNS}} \hat{m}_{\nu}^{1/2} R m_{N_R}^{1/2}, \qquad (2.11)$$

where R is a generalized orthogonal matrix:

$$R = \begin{pmatrix} u_{21} - \omega_{21} & 0\\ \omega_{21} & u_{21} & 0\\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} u_{31} & 0 - \omega_{31}\\ 0 & 1 & 0\\ \omega_{31} & 0 & u_{31} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0\\ 0 & u_{32} - \omega_{32}\\ 0 & \omega_{32} & u_{32} \end{pmatrix},$$
(2.12)

with  $u_{ij} = (1 - \omega_{ij}^2)^{1/2}$  and  $-1 \le \omega_{ij} \le 1$  when R is real. Then the mixing matrix between heavy and light neutrinos can be expressed as [185]

$$V_{\ell N} = \frac{y v_{\nu}}{\sqrt{2}} m_{N_R}^{-1} = U_{\text{PMNS}} \hat{m}_{\nu}^{1/2} R \ m_{N_R}^{-1/2}.$$
 (2.13)

Differently from the type I seesaw, the small neutrino mass may be attributed to a small  $v_{\nu}$  instead of a large  $m_{N_R}$  or a tiny y, making heavy Majorana neutrinos possibly within the reach of LHC. For instance,  $m_{\nu} \sim 0.01$  eV can be obtained with  $v_{\nu} \sim 10$  MeV,  $m_{N_R} \sim 200$  GeV, and  $y \sim 0.006$  whence we have  $V_{\ell N} \sim 10^{-7}$ . As will be clear in section 2.2, the choice of  $v_{\nu} \sim 10$  MeV mainly arises from tight constraints on lepton flavor violation. For such a small  $v_{\nu}$ , the modification to the SM Higgs properties due to the mixing between the doublet scalars is far below the current LHC sensitivity [186] and can be safely ignored. With our simplifying assumption of  $\lambda_{3,4} \approx 0$ , the charged scalars  $H^{\pm}$  do not contribute to the diphoton decay  $h \to \gamma \gamma$  either. In a word, no anomalous SM Higgs decays would be observable in this scenario.



**Figure 1.** BR( $\mu \to e\gamma$ ) as a function of the lightest neutrino mass for normal (left panel) and inverted (right) hierarchy at  $(m_{H^+}, m_{N_R}) = (300, 200)$  GeV.

#### 2.2 Constraints

Now we consider experimental constraints on the  $\nu$ 2HDM. Cosmological considerations have been used to set an upper bound on the sum of neutrino masses,  $\sum_i m_i < 0.23$  eV [187, 188]. The null result searching for neutrinoless double- $\beta$  decays requires the standard effective neutrino mass  $\langle m \rangle_{ee} = |\sum_i (V_{\text{PMNS}})_{ei}^2 m_i|$  to be less than 0.061 - 0.165 eV [189– 194], while the additional contribution from  $\Phi_{\nu}$  is negligible due to its neutrinophilic nature. It is thus safe to state that the lightest neutrino mass should be less than 0.1 eV.

The Yukawa coupling  $y\overline{L}\Phi_{\nu}N_R$  with a not too small y can mediate measurable LFV processes [195, 196]. The currently most stringent constraint comes from the MEG experiment on the decay  $\mu \to e\gamma$  with BR $(\mu \to e\gamma) < 4.2 \times 10^{-13}$  [197, 198], which is calculated as [195, 199]

$$BR(\mu \to e\gamma) = \frac{3\alpha}{64\pi G_F^2} \left| \sum_i \frac{y_{\mu i} y_{ei}^*}{m_{H^+}^2} F\left(\Delta_{H^+}^{N_{Ri}}\right) \right|^2,$$
(2.14)

where  $\Delta_B^A = m_A^2/m_B^2$  is the ratio of the squared masses and the loop function F(x) is

$$F(x) = \frac{1}{6(1-x)^4} \left( 1 - 6x + 3x^2 + 2x^3 - 6x^2 \ln x \right).$$
 (2.15)

While nondegenerate  $N_R$ s are required to generate nondegenerate light neutrinos, we assume approximately in our numerical estimates that they are nearly degenerate to reduce the number of free parameters. The branching ratio then simplifies to

$$BR(\mu \to e\gamma) \approx \frac{3\alpha}{16\pi G_F^2} \frac{m_{N_R}^2 |\tilde{m}_{\mu e}|^2}{m_{H^+}^4 v_{\nu}^4} \left| F\left(\Delta_{H^+}^{N_R}\right) \right|^2,$$
(2.16)

with  $\widetilde{m} = U_{\text{PMNS}} \hat{m}_{\nu} U_{\text{PMNS}}^{\dagger}$  and in particular

$$\tilde{m}_{\mu e} = c_{12}c_{13}c_{23}s_{12}(m_2 - m_1) + c_{13}s_{13}s_{23}e^{-i\delta}[(m_3 - m_2) + c_{12}^2(m_2 - m_1)].$$
(2.17)

Note that the dependence on the free Majorana phases and real matrix R drops out in this approximation, but the dependence on the Dirac phase  $\delta$  can be significant for LFV [200] as the mixing angle  $\theta_{13}$  is known to be not small.

We show in figure 1 BR( $\mu \to e\gamma$ ) as a function of the lightest neutrino mass in either NH or IH and for  $(m_{H^+}, m_{N_R}) = (300, 200)$  GeV. A too small value of  $v_{\nu} \sim 1$  MeV is obviously in conflict with the MEG bound. For nearly degenerate light neutrinos with the lightest mass  $m_{1/3} \gtrsim 0.01$  eV, we have  $m_j - m_i \approx \Delta m_{ji}^2/(2m_i)$ , which explains why  $|\tilde{m}_{\mu e}|$ and thus BR( $\mu \to e\gamma$ ) decrease as  $m_{1/3}$  increases. On the other hand, for the lightest mass  $m_{1/3} \lesssim 0.01$  eV, we have  $|m_j - m_i| \approx \sqrt{|\Delta m_{ji}^2|}$  with  $m_i$  being the lightest mass, so that  $\tilde{m}_{\mu e}$  and thus BR( $\mu \to e\gamma$ ) saturate to a constant. Using eq. (2.16), the upper bound on BR( $\mu \to e\gamma$ ) can be translated into a lower bound on  $m_{H^+}v_{\nu}$  for a given  $m_{N_R}$ :

$$m_{H^+} v_{\nu} \gtrsim \left[ \frac{1 \times 10^{-13}}{\text{BR}(\mu \to e\gamma)} \left( \frac{m_{N_R}}{100 \text{ GeV}} \right)^2 \right]^{1/4} \times 600 \text{ GeV} \cdot \text{MeV}, \quad (2.18)$$

with  $F(m_{N_R}^2/m_{H^+}^2) \sim 0.1$  when  $m_{N_R} \sim m_{\Phi_{\nu}}$ . Hence, for  $m_{N_R} \sim 200$  GeV for instance, the current MEG limit implies that  $m_{H^+}v_{\nu} \gtrsim 600$  GeV · MeV, which can also be seen in figure 1. Note that due to the existence of heavy Majorana neutrino  $N_R$ , the lower bound on  $v_{\nu}$  is about  $\sqrt{m_{N_R}/m_{\nu}} \sim 10^6$  times higher than those on the VEV of the Higgs triplet in type II seesaw [201–203] or of the neutrinophilic doublet in the Dirac scenario of  $\nu$ 2HDM [196].

### 3 Decay properties

The decay properties of the neutrinophilic scalars  $H^{\pm}$ , H, A and heavy Majorana neutrinos  $N_R$  have been discussed in refs. [114, 121]. For completeness, we first review briefly the scenario with  $m_{N_R} > m_{\Phi_{\nu}}$ . Then we concentrate on the opposite case with  $m_{N_R} < m_{\Phi_{\nu}}$ , where LNV signatures at LHC can arise.



Figure 2. Branching ratios of the charged scalar  $H^+$  as a function of the lightest neutrino mass for normal (left panel) and inverted (right) hierarchy.

$BR(H^+)$	$e^+N_{R1}$	$e^+N_{R2}$	$e^+N_{R3}$	$\mu^+ N_{R1}$	$\mu^+ N_{R2}$	$\mu^+ N_{R3}$	$\tau^+ N_{R1}$	$\tau^+ N_{R2}$	$\tau^+ N_{R3}$
NH	0.023	0.004	0.050	0.163	0.211	0.139	0.060	0.006	0.343
IH	0.358	0.015	0.109	0.005	0.192	0.027	0.009	0.197	0.088

**Table 1.** Branching ratios of the charged scalar  $H^+$  into  $\ell^+ N_{Ri}$  for NH and IH at  $m_{1/3} = 0.001$  eV and  $\omega_{ij} = 0.5$ .

Now we turn to the more interesting scenario  $m_{N_R} < m_{\Phi_{\nu}}$ . In this scenario, we can safely neglect the mixing between  $\Phi_{\nu}$  and  $\Phi$  with  $v_{\nu} \sim \mathcal{O}(10 \text{ MeV})$  [121], so that the dominant decays of the neutrinophilic scalars are  $H^+ \rightarrow \ell^+ N_R$  and  $H/A \rightarrow \nu N_R$ , with heavy Majorana neutrinos  $N_R$  decaying further into  $\ell^{\pm}W^{\mp}$ ,  $\nu Z$ , and  $\nu h$ . These decays are analyzed in the following subsections.

BR(H/A)	$\nu N_{R1}$	$\nu N_{R2}$	$\nu N_{R3}$
NH	0.246	0.221	0.533
IH	0.372	0.404	0.224

**Table 2.** Branching ratios of the neutral scalars H/A into  $\nu N_{Ri}$  for NH and IH at  $m_{1/3} = 0.001$  eV,  $\omega_{ij} = 0.5$  and upon summing over neutrinos  $\nu$ .

#### 3.1 Neutrinophilic scalars

In the scenario of  $m_{N_R} < m_{\Phi_{\nu}}$ , the neutrinophilic scalars decay into charged leptons or neutrinos and heavy Majorana neutrino  $N_R$  via the Yukawa coupling y. The partial decay widths are

$$\Gamma(H^+ \to \ell^+ N_{Ri}) = \frac{|y_{\ell i}|^2}{16\pi} m_{H^+} \left(1 - \Delta_{H^+}^{N_{Ri}}\right)^2, \qquad (3.1)$$

$$\Gamma(H/A \to \nu_{\ell} N_{Ri}) = \frac{|y_{\ell i}|^2}{16\pi} m_{H/A} \left(1 - \Delta_{H/A}^{N_{Ri}}\right)^2.$$
(3.2)

The branching ratios of the neutrinophilic scalars are only proportional to  $|y_{\ell i}|^2$  and are then determined by the neutrino parameters via eq. (2.11). As mentioned earlier, we randomly scan the oscillation parameters in their  $2\sigma$  ranges of ref. [182] when evaluating the branching ratios. In figure 2, we show the scanning results of BR $(H^+ \to \ell^+ N_{Ri})$  as a function of the lightest neutrino mass  $m_{1/3}$  in normal/inverted hierarchy by summing over the heavy Majorana neutrinos. We learn that in the nondegenerate neutrino mass region  $m_{1/3} \leq 0.1$  eV:

$$\sum_{i} \operatorname{BR}(H^+ \to e^+ N_{Ri}) < \sum_{i} \operatorname{BR}(H^+ \to \mu^+ N_{Ri}) \approx \sum_{i} \operatorname{BR}(H^+ \to \tau^+ N_{Ri}) \text{ for NH}, (3.3)$$

$$\sum_{i} \operatorname{BR}(H^+ \to e^+ N_{Ri}) > \sum_{i} \operatorname{BR}(H^+ \to \mu^+ N_{Ri}) \approx \sum_{i} \operatorname{BR}(H^+ \to \tau^+ N_{Ri}) \text{ for IH.} \quad (3.4)$$

So we expect that the neutrino mass hierarchy might be distinguishable at LHC by the decay products of the charged scalars  $H^{\pm}$ . To illustrate this, we use the best-fit values of the neutrino oscillation parameters in ref. [182] with the lightest neutrino mass  $m_{1/3} =$ 0.001 eV and  $\omega_{ij} = 0.5$  for the orthogonal R matrix to evaluate BR $(H^+ \rightarrow \ell^+ N_{Ri})$ . The results are shown in table 1 for both NH and IH, and will be employed for later signature simulations. From the table we are informed that a large hierarchy of individual branching ratios exists for both hierarchies with the largest being BR $(H^+ \rightarrow \tau^+ N_{R3}) = 0.343$  for NH and BR $(H^+ \rightarrow e^+ N_{R1}) = 0.358$  for IH.

Concerning the neutral scalars H and A, we sum over the light neutrinos when showing the scanning results, since they are invisible at colliders. In figure 3,  $BR(H/A \to \nu N_{Ri})$  is shown as a function of the lightest neutrino mass for NH and IH. No specific hierarchy in branching ratios is observed. A small difference between NH and IH is that the individual  $BR(H/A \to \nu N_{Ri})$  can exceed 0.5 in NH while it maximally reaches 0.5 in IH. The explicit values of  $BR(H/A \to \nu N_{Ri})$  are shown in table 2 using the same set of parameters as for  $BR(H^+ \to \ell^+ N_{Ri})$ . It is clear from the table that the largest individual branching



Figure 3. Branching ratios of the neutral scalars H/A as a function of the lightest neutrino mass for normal (left panel) and inverted (right) hierarchy.

ratio of H/A is BR $(H/A \rightarrow \nu N_{R3}) = 0.533$  for NH and BR $(H/A \rightarrow \nu N_{R2}) = 0.404$  for IH. And all branching ratios of  $H/A \rightarrow \nu N_{Ri}$  approach 1/3 when the light neutrinos are nearly degenerate.

## 3.2 Heavy Majorana neutrinos

When the heavy Majorana neutrinos are lighter than the neutrinophilic scalars, they decay via the small mixing with the light neutrinos. The partial decay widths are given by

$$\Gamma(N_{Ri} \to \ell^{\pm} W^{\mp}) = \frac{|V_{\ell i}|^2}{8\pi v^2} m_{N_{Ri}}^3 \left(1 - \Delta_{N_{Ri}}^W\right)^2 \left(1 + 2\Delta_{N_{Ri}}^W\right), \qquad (3.5)$$

$$\Gamma(N_{Ri} \to \nu_{\ell} Z) = \frac{|V_{\ell i}|^2}{16\pi v^2} m_{N_{Ri}}^3 \left(1 - \Delta_{N_{Ri}}^Z\right)^2 \left(1 + 2\Delta_{N_{Ri}}^Z\right), \qquad (3.6)$$

$$\Gamma(N_{Ri} \to \nu_{\ell} h) = \frac{|V_{\ell i}|^2}{16\pi v^2} m_{N_{Ri}}^3 \left(1 - \Delta_{N_{Ri}}^h\right)^2.$$
(3.7)

Due to the smallness of  $V_{\ell N}$ ,  $N_{Ri}$  are actually long-lived, which could lead to visible displaced vertices at LHC [185]. In collider phenomenology study, the decay  $N_{Ri} \rightarrow \ell^{\pm} W^{\mp}$ receives more attention, not only because it induces LNV signatures but also because it can be used to fully reconstruct the mass of  $N_{Ri}$ . In figure 4, we present the branching ratios of  $N_{Ri}$  as a function of  $m_{N_{Ri}}$  upon summing over the lepton flavors, which depend only on heavy neutrino masses  $m_{N_{Ri}}$ . As shown clearly, we have

$$BR(N_{Ri} \to \ell^{\pm} W^{\mp}) : BR(N_{Ri} \to \nu Z) : BR(N_{Ri} \to \nu h) \approx 2 : 1 : 1, \qquad (3.8)$$

in the large  $m_{N_{Ri}}$  limit, where the gauge bosons in the final state are mainly longitudinally polarized [142].

For the decays into charged leptons there are similar flavor relations among BR $(N_{Ri} \rightarrow \ell^{\pm}W^{\mp})$  as in the decays of  $H^+$ . In figure 5, the scanning results of BR $(N_{R1} \rightarrow \ell^{\pm}W^{\mp})$ 



Figure 4. Branching ratios of  $N_{Ri}$  as a function of  $m_{N_{Ri}}$  upon summing over the lepton flavors.



Figure 5. Branching ratios of  $N_{R1}$  into  $\ell^{\pm}W^{\mp}$  as a function of the lightest neutrino mass for normal (left panel) and inverted (right) hierarchy at  $m_{N_{R1}} = 200$  GeV.

are shown for  $\ell = e, \mu, \tau$  separately as a function of the lightest neutrino mass at  $m_{N_{R1}} = 200$  GeV, while the decays of  $N_{R2}$  and  $N_{R3}$  are similar. We observe the relations

$$BR(N_{R1} \to e^{\pm}W^{\mp}) < BR(N_{R1} \to \mu^{\pm}W^{\mp}) \approx BR(N_{R1} \to \tau^{\pm}W^{\mp}) \text{ for NH}, \quad (3.9)$$

$$BR(N_{R1} \to e^{\pm}W^{\mp}) > BR(N_{R1} \to \mu^{\pm}W^{\mp}) \approx BR(N_{R1} \to \tau^{\pm}W^{\mp}) \text{ for IH}, \quad (3.10)$$

when the lightest neutrino mass is less than 0.1 eV. These flavor relations can be better seen in figure 6 which shows the results of the averaged branching ratios  $\sum_i \text{BR}(N_{Ri} \rightarrow \ell^{\pm}W^{\mp})/3$ . Hence the decays of heavy Majorana neutrinos  $N_{Ri}$  into charged leptons could also be employed to distinguish between the neutrino mass hierarchies. Since  $\text{BR}(N_{Ri} \rightarrow$ 



Figure 6. The averaged branching ratios of  $N_{Ri}$  as a function of the lightest neutrino mass for normal (left panel) and inverted hierarchy (right) at  $m_{N_{Ri}} = 200$  GeV.

$\operatorname{FR}(N_{Ri})$	$e^{\pm}W^{\mp}$	$\mu^{\pm}W^{\mp}$	$\tau^{\pm}W^{\mp}$
$N_{R1}$	$0.093 \ (0.962)$	$0.663 \ (0.014)$	$0.244 \ (0.024)$
$N_{R2}$	$0.019 \ (0.037)$	$0.955\ (0.476)$	$0.026 \ (0.487)$
$N_{R3}$	$0.094\ (0.486)$	$0.261\ (0.122)$	$0.645\ (0.392)$

**Table 3.** Flavor ratios for heavy Majorana neutrinos  $N_{Ri}$  decaying into  $\ell^{\pm}W^{\mp}$  for NH (IH).

 $\ell^{\pm}W^{\mp}$ ) depends on  $m_{N_{Ri}}$ , we define the flavor ratio (FR)

$$\operatorname{FR}(N_{Ri} \to \ell^{\pm} W^{\mp}) = \frac{\operatorname{BR}(N_{Ri} \to \ell^{\pm} W^{\mp})}{\sum_{\ell} \operatorname{BR}(N_{Ri} \to \ell^{\pm} W^{\mp})} = \frac{|V_{\ell i}|^2}{\sum_{\ell} |V_{\ell i}|^2}, \quad (3.11)$$

which is independent of  $m_{N_{Ri}}$  but depends only on the neutrino oscillation parameters and the R matrix. The distributions of  $\operatorname{FR}(N_{Ri} \to \ell^{\pm} W^{\mp})$  and  $\sum_{i} \operatorname{FR}(N_{Ri} \to \ell^{\pm} W^{\mp})/3$ are similar to figures 5 and 6 respectively, up to a normalization factor of  $\sum_{\ell} \operatorname{BR}(N_{Ri} \to \ell^{\pm} W^{\mp})$ . In table 3, we show the values of  $\operatorname{FR}(N_{Ri} \to \ell^{\pm} W^{\mp})$  for the same set of parameters as for  $\operatorname{BR}(H^+ \to \ell^+ N_{Ri})$ . For NH,  $N_{R1,2} \to \mu^{\pm} W^{\mp}$  and  $N_{R3} \to \tau^{\pm} W^{\mp}$  are dominant while for IH  $N_{R1} \to e^{\pm} W^{\mp}$ ,  $N_{R2} \to \mu^{\pm} W^{\mp}/\tau^{\pm} W^{\mp}$ , and  $N_{R3} \to e^{\pm} W^{\mp}/\tau^{\pm} W^{\mp}$  take over. With the flavor ratios introduced above, we can easily acquire the branching ratios of  $N_{Ri} \to \ell^{\pm} W^{\mp}$  for any values of  $m_{N_{Ri}}$  by

$$BR(N_{Ri} \to \ell^{\pm} W^{\mp}) = FR(N_{Ri} \to \ell^{\pm} W^{\mp}) \times \sum_{\ell} BR(N_{Ri} \to \ell^{\pm} W^{\mp}), \qquad (3.12)$$

where  $\operatorname{FR}(N_{Ri} \to \ell^{\pm} W^{\mp})$  and  $\sum_{\ell} \operatorname{BR}(N_{Ri} \to \ell^{\pm} W^{\mp})$  are given in table 3 and figure 4, respectively.

Last but not least, we show in figure 7 the scatter plots for BR $(H^+ \rightarrow \ell^+ \ell^+ W^-)$ upon summing over the intermediate heavy Majorana neutrinos  $N_{Ri}$ . With  $W^-$  further



Figure 7. Branching ratios of  $H^+$  into  $\ell^+\ell^+W^-$  as a function of the lightest neutrino mass for normal (left panel) and inverted (right) hierarchy at  $(m_{H^+}, m_{N_R}) = (300, 200)$  GeV.



Figure 8. Cross sections for the pair and associated production of the neutrophilic scalars at 13 (14) TeV LHC as a function of their mass  $m_{H^+} = m_H = m_A = m_{\Phi_{\nu}}$ .

decaying hadronically, we have the LNV decay of the charged scalars  $H^+ \to \ell^+ \ell^+ j j$ , which contributes to several LNV signatures at LHC. Since  $\ell$  is identified as e and  $\mu$  at colliders, we have  $\mu^+\mu^+$  dominance for NH and  $e^+e^+$  dominance for IH when considering LNV signatures, which makes it possible to distinguish between the neutrino mass hierarchies.

# 4 Lepton number violating signatures

After the systematic study on the decay properties of the neutrinophilic scalars  $H^{\pm}$ , H, A and the heavy Majorana neutrinos  $N_R$  in section 3, we can now investigate the LNV signatures at LHC. Our simulation procedure is as follows. We first implement the  $\nu$ 2HDM into the package FeynRules [213–215] to generate the UFO [216] model file. The parton level

signal and corresponding background events are generated with MadGraph5\_aMC@NLO [217, 218] using the NNPDF2.3 [219, 220] LO parton distribution function set, and then pass through Pythia6 [221] to include showering and hadronization. Delphes3 [222, 223] is then employed for a fast detector simulation and MadAnalysis5 [224-226] for analysis. Finally, the 95% C.L. exclusion limits are acquired by employing CheckMATE [227, 228].

We first recall that the conventional well-studied LNV signature for heavy Majorana neutrinos,  $pp \to W^{\pm}/H^{\pm} \to \ell^{\pm}N_R \to \ell^{\pm}\ell^{\pm}jj$ , is also possible in  $\nu$ 2HDM, but its amplitude is suppressed by  $V_{\ell N} \sim 10^{-7}$  and  $v_{\nu}/v \sim 10^{-5}$  (for  $v_{\nu} \sim \mathcal{O}(10 \text{ MeV})$ ) respectively, making the signature practically unobservable at LHC. In contrast, due to the doublet nature of the neutrinophilic scalars, they can be pair and associated produced via the Drell-Yan processes

$$pp \to H^+H^-, \ H^\pm H, \ H^\pm A, \ HA.$$
 (4.1)

Their cross sections at LHC are presented in figure 8, which range from 400 to 0.01 fb in the mass interval 100 - 1000 GeV at 13 TeV, and become slightly enhanced at 14 TeV. There are many possible final states given by the decay channels of the scalars that we discussed in section 3 and the sequential decays of SM particles [173]. These channels lead to various signatures which are conventionally classified according to the multiplicities of charged leptons and jets. Among them the following three LNV signatures are most interesting and promising, and will be studied in section 4.1–4.3:

- $2\ell^{\pm}4j + \not\!\!\!E_T$  from  $H^{\pm}H$ ,  $H^{\pm}A$  and HA production,
- $3\ell^{\pm}4j + \not\!\!\!E_T$  from  $H^{\pm}H$ ,  $H^{\pm}A$  production,
- $3\ell^{\pm}\ell^{\mp}4j$  from  $H^+H^-$  production,

BP-A : 
$$m_{N_R} = 200 \text{GeV}, m_{\Phi_{\nu}} = 300 \text{GeV},$$
  
BP-B :  $m_{N_R} = 300 \text{GeV}, m_{\Phi_{\nu}} = 400 \text{GeV},$   
BP-C :  $m_{N_R} = 400 \text{GeV}, m_{\Phi_{\nu}} = 500 \text{GeV},$  (4.2)

which are still allowed by current constraints. The signals and backgrounds will be simulated at 13 (14) TeV with an integrated luminosity of 100 (3000) fb<sup>-1</sup>, or LHC13@100 (LHC14@3000) for short, and the corresponding exclusion limits will be derived as well.



Figure 9. Cross sections of LNV signatures at 13 (14) TeV LHC.

#### 4.1 Dilepton signature

The signature comes from pair and associated production of the doublet scalar  $\Phi_{\nu}$  and subsequent decays:

$$pp \to H^{\pm}H/A \to \ell^{\pm}N_R\nu N_R \to \ell^{\pm}\ell^{\pm}W^{\mp}\nu\nu Z/h \to \ell^{\pm}\ell^{\pm}jj\nu\nu jj,$$
 (4.3)

$$pp \to HA \to \nu N_R \nu N_R \to \nu \ell^{\pm} W^{\mp} \nu \ell^{\pm} W^{\mp} \to \nu \ell^{\pm} j j \nu \ell^{\pm} j j, \qquad (4.4)$$

where  $\ell = e, \mu$  for collider simulations. The major sources of background are  $t\bar{t}W, t\bar{t}Z$ and  $W^{\pm}W^{\pm}W^{\mp}jj$ . For the last one we use the MLM [229] matching scheme with xqcut = 25 (30) GeV for 13 (14) TeV LHC.

$$N(j) = 4, \ N(b) = 0, \ N(\ell^{\pm}) = 2.$$
 (4.6)



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Channels		Basic cuts in Eq: $(4.5)$	N(j) = 4	N(b) = 0	$N(\ell^{\pm}) = 2$	$S/\sqrt{S+B}$
	NH	51(1770)	16(569)	8.3 (319)	3.9(164)	0.65(4.64)
DI -A	IH	63~(2425)	20 (779)	10(436)	4.9(224)	0.79(6.20)
	NH	16(566)	5.0(179)	2.4 (90)	1.3(49)	0.22(1.44)
DI-D	IH	20(775)	6.1(245)	3.0 (124)	1.6(67)	0.27(1.96)
DD C	NH	5.8 (210)	1.8(65)	0.87(31)	0.47 (17)	0.08(0.51)
Dr-U	IH	7.2(288)	2.2(88)	1.1 (43)	0.58(23)	0.10 (0.70)
$t\bar{t}V$	V	1020 (35023)	325 (10917)	54 (1810)	18 (580)	
$t\bar{t}$	2	1043 (37909)	232 (8188)	39(1399)	2.9(172)	
WWW	N j j	155 (4623)	43(1213)	32 (901)	12(334)	

**Table 4**. Cut-flow for the SSD signature at three benchmark points in eq. (4.2) and dominant backgrounds at LHC13@100 (LHC14@3000) for both NH and IH.



**Figure 11**. Left (right): reconstruction of  $N_R$  ( $H^{\pm}$ ) via invariant mass  $M_{jj\ell^{\pm}}$  ( $M_{jj\ell^{\pm}\ell^{\pm}}$ ) at 13 TeV LHC for the SSD signature.

Here, the cut on the number of *b*-jet mainly aims to reduce the  $t\bar{t}W$  and  $t\bar{t}Z$  backgrounds. The identification of *b*-jets is performed with a tagging efficiency of 70%, a mis-tagging rate of 10% for *c*-jets and 1% for light-flavor jets, respectively [230, 231].

In table 4 we show the cut-flow for the SSD signature at the benchmark points and the dominant backgrounds, and list the statistical significance  $S/\sqrt{S+B}$  in the last column, where S(B) stands for the survival number of signal (background) events after applying all cuts. As shown in figure 10, the distributions of the background and signal events (especially for BP-A) are so similar for kinematic variables like  $p_T^j$  and  $\Delta R_{j\ell,\ell\ell}$  that it is hard to eliminate the background without sacrificing a big part of the signal events by the naive cuts in eqs. (4.5), (4.6). As a result, the significance can barely reach  $1\sigma$  at LHC13@100. However, at LHC14@3000, we may have a chance to probe the SSD signal for BP-A, even for which a more sophisticated and efficient cut strategy is highly desired to improve the testability of this SSD signature.

For the SSD signature the decay chains  $N_R \to \ell^{\pm} j j$  and  $H^{\pm} \to \ell^{\pm} \ell^{\pm} j j$  can be used to fully reconstruct the masses of heavy neutrinos  $N_{Ri}$  and charged scalars  $H^{\pm}$ . In figure 11, we depict the distributions for the reconstruction of  $N_R$  and  $H^{\pm}$  via the invariant mass



Figure 12. The expected 95% C.L. exclusion limits in the  $m_{N_R} - m_{\Phi_{\nu}}$  plane for the SSD signature at LHC13@100 and LHC14@3000.

 $M_{jj\ell^{\pm}}$  and  $M_{jj\ell^{\pm}\ell^{\pm}}$  respectively at 13 TeV LHC after applying all cuts in eqs. (4.5), (4.6). To render the resonance peaks more discernible, we have taken advantage of the expert mode of MadAnalysis5 [224–226] so as to pick up close  $jj\ell^{\pm}$  for  $N_R$  and  $jj\ell^{\pm}\ell^{\pm}$  for  $H^{\pm}$  in the final states. We see that  $N_R$  and  $H^{\pm}$  are apparently reconstructible, although the production rate for BP-B and BP-C is actually too small to be detected even at LHC14@3000.

Based on the above benchmark points study, it is worthwhile to figure out the exclusion limits of the SSD signature in the  $m_{N_R} - m_{\Phi_{\nu}}$  plane. In figure 12 we show the expected 95% C.L. exclusion limits at LHC13@100 and LHC14@3000 by employing CheckMATE [227, 228] with the cuts in eqs. (4.5), (4.6). We see that while LHC13@100 rules out a parameter region with  $m_{N_R} \leq 200$  GeV and  $m_{\Phi_{\nu}} \leq 250$  GeV, LHC14@3000 can exclude a larger region up to  $m_{N_R} \leq 350$  GeV and  $m_{\Phi_{\nu}} \leq 400$  GeV. Nevertheless, a compressed spectrum with  $m_{\Phi_{\nu}} \approx m_{N_R}$  is still allowed for both NH and IH, because due to the limited phase space for the  $\Phi_{\nu} \rightarrow N_R$  decay the momentum of the leptons or  $\not E_T$  is too small to pass the basic cuts in eq. (4.5). We also note that the limits are more stringent for IH than NH. This originates simply from the fact that the cross section for the SSD signature in the IH case is about 1.37 times as large as the one in the NH case as shown in figure 9. As also shown clearly in figure 12, although all three benchmark points are beyond the reach of LHC13@100, they are either within the reach (BP-A), on the edge (BP-B) or out (BP-C) of the exclusion capability of LHC14@3000.

#### 4.2 Trilepton signature

The SST signature, behaving as  $3\ell^{\pm}4j + \not\!\!\!E_T$ , is the most distinct one in this  $\nu$ 2HDM. It originates from the associated production  $H^{\pm}H$  and  $H^{\pm}A$ :

$$pp \to H^{\pm}H/A \to \ell^{\pm}N_R \nu N_R \to \ell^{\pm}\ell^{\pm}W^{\mp}\nu\ell^{\pm}W^{\mp} \to 3\ell^{\pm}4j + \not\!\!\!E_T.$$

$$(4.7)$$



**Figure 13**. Distributions of numbers of same-sign leptons  $N(\ell^{\pm})$  and jets N(j), transverse momenta  $p_T(j)$  and  $p_T(\ell)$ , missing transverse energy  $\not\!\!\!E_T$ , and relative distances  $\Delta R_{j\ell}$  for the SST signature and corresponding backgrounds at 13 TeV LHC.

Such a kind of signature is rather exotic in the SM as it violates the lepton number notably by three units in visible final states. The dominant backgrounds are generated from  $t\bar{t}W^{\pm}Z$ and  $t\bar{t}ZZ$ , where the  $t\bar{t}$  pair decays semi-leptonically and the bosons decay leptonically with one lepton missed from each Z. The cross sections of backgrounds are extremely small, making the SST signature easier to be tested on LHC.

Channels		Basic cuts in eq. $(4.5)$	$N(\ell^{\pm}) = 3$	$N(j) \ge 2$	$S/\sqrt{S+B}$
BD V	NH	11 (378)	3.7(127)	3.7(126)	1.85(10.8)
DI -A	IH	13 (458)	4.5(153)	4.5(152)	2.05(11.9)
PD P	NH	2.9(102)	1.1 (37)	1.0(36)	0.87(5.25)
DI -D	IH	3.5(124)	1.3(44)	1.3(44)	1.02(5.93)
BD C	NH	0.96 (36)	0.38(13)	0.37~(13)	0.45 (2.65)
DI-C	IH	1.2(43)	0.46(16)	0.44(16)	0.51 (3.08)
$t\bar{t}V$	V	1 (36)	0.25 (8.5)	0.24(8.3)	
VVV(V)		0.5~(16)	0.15(5)	0.08(2.7)	

Table 5. Same as table 4, but for the SST signature.



Figure 14. Same as figure 11, but for the SST signature.

background is rather clean while the signal itself has also a relatively small cross section, only few events could survive the exact selection of  $3\ell^{\pm}4j + \not\!\!\!E_T$ . We thus turn to a looser selection to pick up inclusive events,

$$N(\ell^{\pm}) = 3, \ N(j) \ge 2.$$
 (4.8)

According to figure 13 about one third of the signal events could pass this cut. Because of this looser selection, the backgrounds from VVV and VVVV (V = W/Z) should be taken into account as well.

The cut-flow for the signal at the three benchmark points in eq. (4.2) in both NH and IH cases and the dominant backgrounds is presented in table 5. The small backgrounds  $t\bar{t}W^{\pm}Z$  and  $t\bar{t}ZZ$  are summed to  $t\bar{t}VV$ , so are VVV and VVVV to VVV(V). At LHC13@100, only BP-A in the IH case could lead to a  $2\sigma$  excess. But at LHC14@3000, we will have a good chance to discover BP-A and BP-B, and even BP-C will result in about  $2.5\sigma$  ( $3\sigma$ ) excess for NH (IH). In figure 14, we plot the distributions in the invariant masses  $M_{jj\ell^{\pm}}$  and  $M_{jj\ell^{\pm}\ell^{\pm}}$  for the reconstruction of the  $N_R$  and  $H^{\pm}$  particles at 13 TeV LHC.

Based on the cuts in eqs. (4.5), (4.8) we now acquire the expected 95% C.L. exclusion limits at LHC13@100 and LHC14@300 by scanning in the  $m_{N_R} - m_{\Phi_{\nu}}$  plane. Our results are presented in figure 15. For instance, at LHC13@100,  $m_{\Phi_{\nu}}$  could be excluded up to



Figure 15. Same as figure 12, but for the SST signature.

about 350 GeV for NH or 380 GeV for IH at  $m_{N_R} \sim 150$  GeV, or conversely,  $m_{N_R}$  could be excluded up to about 280 GeV for NH or 305 GeV for IH at  $m_{\Phi_{\nu}} \sim 330$  GeV. With higher integrated luminosity at LHC14@3000, the exclusion limit would extend to  $m_{\Phi_{\nu}} \lesssim$ 600 GeV,  $m_{N_R} \lesssim 500$  GeV for NH, and even higher for IH, i.e.,  $m_{\Phi_{\nu}} \lesssim 640$  GeV,  $m_{N_R} \lesssim$ 530 GeV. It is evident that LHC13@100 (LHC14@3000) will be capable of excluding BP-A (BP-B and BP-C) through the SST signature.

## 4.3 Four-lepton signature

Finally we study the most exotic signature involving four leptons, as a result of the decay chains:

$$pp \to H^+H^- \to \ell^+ N_R \ell^- N_R \to \ell^+ W^{\mp} \ell^{\pm} \ell^- W^{\mp} \ell^{\pm} \to 3\ell^{\pm} \ell^{\mp} 4j.$$

$$\tag{4.9}$$

This is also the only case that involves two charged scalars in the intermediate state. The major sources of background are  $t\bar{t}h$ ,  $t\bar{t}V$ ,  $t\bar{t}t\bar{t}$ ,  $t\bar{t}VV$ , with  $V = W^{\pm}$ , Z. We found that the kinematical distributions for the four-lepton signature are similar to those for the SST signature. We impose the same basic cuts in eq. (4.5) but of course drop the cut on  $\not{E}_T$ , because no invisible particles are involved now. In addition, we apply the specific cuts for the four-lepton signature:

$$N(j) = 4, \ N(\ell^{\pm}) = 3, \ N(\ell^{\mp}) = 1.$$
 (4.10)

No further cuts are required, since we are practically background free at this stage as will be shown below.

In table 6 we show the cut-flow for the four-lepton signature at the three benchmark points in eq. (4.2) as well as for the dominant backgrounds. We notice that the backgrounds become zero after the selection cuts in eqs. (4.5), (4.10), and this increases the feasibility of the four-lepton signature remarkably. However, we should not forget that this signal is also the weakest among the three studied in this work. At LHC13@100, this signature is

Channels		Basic cuts	N(j) = 4	$N(\ell^{\pm}) = 3$ $N(\ell^{\mp}) = 1$	$S/\sqrt{S+B}$
BD V	NH	23 (805)	5.6(195)	0.74(29)	0.86(5.39)
DI -A	IH	40(1404)	9.8(340)	1.3(43)	$1.14\ (6.56)$
<b>BD B</b>	NH	7.2(242)	1.7(58)	0.27(8.7)	0.52(2.95)
DI -D	IH	13 (422)	3.0(102)	0.47(15)	0.69(3.87)
DD C	NH	2.4(89)	0.58(21)	0.09(3.2)	0.30(1.79)
Dr-U	IH	4.2(155)	1.0(37)	0.16(5.6)	0.40(2.37)
tīh	ı	403 (33431)	36~(2869)	0 (0)	
$t\bar{t}t$	$\overline{t}$	$125\ (11316)$	11 (971)	0 (0)	
$t\bar{t}V$		6043 (488674)	538 (41943)	0 (0)	

 Table 6. Same as table 4, but for the four-lepton signature.



Figure 16. Same as figure 11, but for the four-lepton signature.

undetectable even for BP-A. At LHC14@3000, BP-A will have a 5.39 (6.56) $\sigma$  significance in the case of NH (IH), while BP-B can only lead to a 2.95 (3.87) $\sigma$  significance. To speak roughly about the prospect to observe such signatures at LHC, the four-lepton signature lies in between the dilepton and trilepton signatures. Adopting the same method as for the other two signatures, we reconstruct in figure 16 the heavy Majorana neutrinos  $N_R$ and charged scalars  $H^{\pm}$  via the distributions in the invariant masses  $M_{jjl^{\pm}}$  and  $M_{jjl^{\pm}l^{\pm}}$ respectively. Although the peaks in the normalized distributions are most remarkable among the three, the number of signal events is also the least.

The expected 95% C.L. exclusion limits for the four-lepton signature are shown in figure 17. The region with  $m_{N_R} \leq 250$  GeV and  $m_{\Phi_{\nu}} \leq 300$  GeV could be excluded at LHC13@100, while LHC14@3000 could reach  $m_{N_R} \leq 450$  GeV and  $m_{\Phi_{\nu}} \leq 500$  GeV. Interestingly, we find that BP-A for IH but not for NH could be excluded by LHC13@100, while BP-B would be fully covered by LHC14@3000 for both NH and IH cases.

### 5 Conclusion

The neutrinophilic two-Higgs-doublet model provides a natural way to generate tiny neutrino mass with TeV scale Majorana neutrinos  $N_R$  and scalars  $\Phi_{\nu}$ , and results in interesting



Figure 17. Same as figure 12, but for the four-lepton signature.

phenomena such as LFV processes and LNV signatures at LHC. We found that the LFV processes of the charged leptons mediated by the Yukawa coupling  $\bar{L}\tilde{\Phi}_{\nu}N_R$  can be within the reach of current experiments. Using the most stringent constraint on the decay  $\mu \to e\gamma$ , we derived a combined tight bound on the  $\Phi_{\nu}$  mass and VEV,  $m_{H^+}v_{\nu} \gtrsim 600 \text{ GeV} \cdot \text{MeV}$ , for  $N_R$  of a few hundreds GeV.

Our work has been focused on the LNV signatures from heavy Majorana neutrinos at LHC. To achieve this, we systematically investigated the decay properties of the neutrinophilic scalars  $H^{\pm}$ , H, A and heavy Majorana neutrinos  $N_R$  for the mass order  $m_{N_R} < m_{\Phi_{\nu}}$ . Our results show that there is strong correlation between the neutrino mass hierarchy and the flavor fraction of charged leptons in the decays of  $H^{\pm}$  and  $N_R$ . In particular, we expect  $\mu^{\pm}\mu^{\pm}$  ( $e^{\pm}e^{\pm}$ ) dominance for NH (IH) in the LNV decay chain  $H^{\pm} \rightarrow \ell^{\pm}N_R \rightarrow \ell^{\pm}\ell^{\pm}jj$  ( $\ell = e, \mu$ ). In addition, the production rate for LNV signatures is larger in the case of IH than NH, making the physics signals in the former case more promising to be detected at LHC.

When the new scalars are heavier than the new neutrinos, they are first produced at LHC via the Drell-Yan processes and then cascade decay into the new neutrinos  $H^{\pm} \rightarrow \ell^{\pm} N_R$ ,  $H/A \rightarrow \nu N_R$  and the SM particles  $N_R \rightarrow \ell^{\pm} W^{\mp}$ ,  $\nu Z$ ,  $\nu h$ . This results in three kinds of LNV signatures:  $2\ell^{\pm}4j + \not{\!\!\!E}_T$ ,  $3\ell^{\pm}4j + \not{\!\!\!E}_T$ , and  $3\ell^{\pm}\ell^{\mp}4j$ . To illustrate the testability of such LNV signatures at LHC, we worked with three benchmark points (BP-A, -B, -C, for short),  $(m_{N_R}, m_{\Phi_{\nu}}) = (200, 300)$ , (300, 400), (400, 500) GeV, and performed detailed simulations. Our simulation results are recapitulated in table 7. We found that the SST signature is the most promising among the three. Although at LHC13@100 it is hard to observe excess events of the SST signature even for BP-A, we can discover BP-A and BP-B from the SST signature at LHC14@3000. Furthermore, BP-C with both scalars and neutrinos much heavier can also lead to about  $2.5\sigma$  ( $3\sigma$ ) excess for NH (IH) at LHC14@3000. With such high integrated luminosity there is also a good chance to probe the SSD and four-lepton signatures for BP-A. Conversely, if no excess is observed, the mass region of

LNV si	gnatures	$2\ell^{\pm}4jE_T$	$3\ell^{\pm}4jE_T$	$3\ell^{\pm}\ell^{\mp}4j$
RD Λ	NH	0.65(4.64)	1.85(10.8)	$0.86\ (5.39)$
DI -A	IH	0.79~(6.20)	2.05(11.9)	$1.14 \ (6.56)$
DD D	NH	0.22(1.44)	0.87 (5.25)	0.52(2.95)
DI -D	IH	0.27 (1.96)	1.02(5.93)	0.69(3.87)
BD C	NH	$0.08 \ (0.51)$	0.45 (2.65)	0.30(1.79)
DI -0	IH	$0.10 \ (0.70)$	$0.51 \ (3.08)$	0.40(2.37)
Ma	NH	200 (320)	280 (500)	230 (420)
$NI_{NR}$	IH	215 (340)	305~(530)	250 (460)
M	NH	230 (360)	350 (600)	270 (470)
$W \Phi_{\nu}$	IH	250 (390)	380~(640)	310~(530)

Table 7. Summary of statistical significance for three types of signatures at three benchmark points and of approximate lower bounds on new particle masses to be expected at LHC13@100 (LHC14@3000) for neutrino masses in either NH or IH.

 $m_{N_R} \lesssim 300 \text{ GeV}$  and  $m_{\Phi_{\nu}} \lesssim 350 \text{ GeV}$  will be excluded by LHC13@100, and the excluded region will be extended at LHC14@3000 to  $m_{N_R} \lesssim 500 \text{ GeV}$  and  $m_{\Phi_{\nu}} \lesssim 600 \text{ GeV}$  by the SST signature.

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