

RECEIVED: April 4, 2012 ACCEPTED: April 10, 2012 PUBLISHED: April 30, 2012

# Meson Thermalization in Various Dimensions

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ABSTRACT: In gauge/gravity duality framework the thermalization of mesons in strongly coupled (p+1)-dimensional gauge theories is studied for a general Dp-Dq system,  $q \geq p$ , using the flavour Dq-brane as a probe. Thermalization corresponds to the horizon formation on the flavour Dq-brane. We calculate the thermalization time-scale due to a time-dependent change in the baryon number chemical potential, baryon injection in the field theory. We observe that for such a general system it has a universal behaviour depending only on the t'Hooft coupling constant and the two parameters which describe how we inject baryons into the system. We show that this universal behaviour is independent of the details of the theory whether it is conformal and/or supersymmetric.

KEYWORDS: Gauge-gravity correspondence, D-branes, Holography and quark-gluon plasmas

ARXIV EPRINT: 1203.3425

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## 1 Introduction and results

One of the interesting experiments in particle physics is the heavy ion collision done at RHIC (Relativistic Heavy Ion Collider) and LHC (Large Hadron Collider). In these experimental set-ups two pancakes of heavy nuclei such as Gold(Au) or Lead(Pb) are collided at a relativistic speed. Very soon after the collision a new phase of matter called QGP (Quark-Gluon Plasma) is produced. Understanding the properties of the QGP has attracted a lot of attention. It's been realized through hydrodynamic simulations and experimental observations that the plasma is strongly coupled with very low viscosity over entropy density,  $(\frac{\eta}{s})$  [1, 2].

A very challenging observation is the very rapid thermalization of plasma. At almost 1fm/c after the collision the plasma reaches local equilibrium (thermalizes) and its dynamics is approximately described by ideal fluid hydrodynamics [3, 4]. Since the system is strongly coupled and perturbative calculations can not be applied, it is difficult to study an out-of-equilibrium process such as thermalization. Models of QCD such as lattice QCD have not been able to achieve much to explain this phenomenon.

Recently AdS/CFT correspondence [5–7], or more broadly, the gauge/gravity duality has gained some success in making toy models to describe this process. Generally speaking, the AdS/CFT duality has provided a powerful framework to study strongly coupled gauge theories. According to the original statement of the correspondence, type IIB string theory on  $AdS_5 \times S^5$  is dual to D=4,  $\mathcal{N}=4$ , SU(N) super Yang-Milles (SYM) theory. This idea was then generalized to a broader class of gauge/gravity dualities in various dimensions [8]. It was shown that in the large N limit, the strongly coupled SU(N) gauge theory living on the worldvolume of N coincident Dp-branes is dual to the supergravity on the near horizon geometry of the Dp-branes. Although for p=3 the gauge theory enjoys conformal symmetry, the gauge coupling is dimensionful for the other values of p.

The absence of conformal invariance in the gauge theory leads to the radial variation of both the string coupling (dilaton) and the spacetime curvature in the dual description. This is due to the fact that duality relates the radial coordinate transverse to the Dp-brane to the energy scale in the gauge theory. The supergravity background is only reliable for weak string coupling and small curvature. This is provided in a regime of energy where [8]

$$1 \ll g_{\text{eff}} \ll N^{\frac{4}{7-p}} \ . \tag{1.1}$$

 $g_{\text{eff}}$  is the dimensionless effective coupling,  $g_{\text{eff}} = g_{YM}^2 N u^{p-3}$ , where u is the radial coordinate. In this intermediate regime of energy the dual gauge theory is always strongly coupled.

AdS/CFT techniques have been applied to study rapid thermalization observed at RHIC [9–20]. The thermalization in field theory which is an out-of-equilibrium process happens after an injection of energy into the system. In the gauge/gravity context this injection can be done in two different ways: by directly adding a time-dependent source to the boundary field theory which corresponds to turning on a non-normalizable mode [9–11] or by introducing a time-dependent coupling in the field theory which is represented by nontrivial time-dependent classical solutions of some kinds of probe branes [18–20]. In the first method the source is assumed to be nonzero only for a limited time interval. In the bulk gravity dual this corresponds to the collapse of matter and formation of the horizon; in other words black hole formation. In [9, 10] the authors have been able to quantitatively give an approximate value for the isotropization (thermalization) time of the plasma which is close to the experimental result. In the other method the horizon formation happens on the probe brane and then the energy can flow from the probe brane to the bulk gravitational degrees of freedom. This corresponds to the dissipation of energy into the field theory.

Note that in the first method the field theory is pure YM and conformal. While in the other one the field theory contains fundamental matters and mesons in addition to the pure YM. Studying the thermalization in the meson sector is the subject of an interesting work [20] where the authors have modeled the thermalization by baryon injection into the system. This idea comes from the observed sudden change in the baryon chemical potential at QGP production.

The addition of fundamental matter to the SYM theory which lives on the boundary of AdS space is done by adding the flavour branes to the dual gravity system. This has been proposed for the first time in [21] where  $N_f$  D7-branes are added to the AdS<sub>5</sub> ×

 $S^5$  background in the probe limit where  $\frac{N_f}{N}$  is small. This will reduce the number of supersymmetries of the original theory from 16 to 8 and will add a matter hypermultiplet in the fundamental representation to the SYM theory. The hypermultiplet describes the dynamical quarks living in four dimensions and the fluctuations of the flavour brane explain the meson spectrum of the field theory. In another interesting work this system has been extended to more general settings of Dp-Dq brane configurations [22, 23].

In the D3-D7 brane set-up the authors of [20] have studied thermalization in the meson sector of the SYM theory. The thermalization is modeled by a sudden change in the baryon number. In the gravity dual description this change is realized by throwing the baryonically-charged fundamental strings from the boundary to the bulk. The end points of the strings stretched between D3 and D7-branes act as the source for the gauge field on the D7-brane. Therefore this set-up provides a time-dependent gauge field configuration. Using DBI action this induces a time-dependent metric and eventually emergence of an apparent horizon on the D7-brane. Hence the baryon injection leads to the horizon formation on the D7-brane which is the signal of thermalization in the meson sector of the field theory.

An intriguing observation of [20] is that the thermalization time in the field theory for the D3-D7 system can be written only in terms of a few parameters which are: the gauge theory t'Hooft coupling ( $\lambda = g_{YM}^2 N$ ), maximum baryon number density( $n_B$ ) and the inverse of the variation time-scale of the baryon number ( $\omega$ ). Note that  $n_B$  and  $\omega$  describe how the baryon charge density changes in the system. Therefore the thermalization time is independent of the details of the theory by which we mean the form of the Lagrangian. The authors generally argue that this possible universal behaviour might be generalized to other gauge theories. In this paper we would like to examine this idea and generalize their computation to the Dp-Dq system.

Interestingly for such a general system where the gauge theory is not conformal, the background metric is not AdS and as mentioned above gauge/gravity duality can be trusted only in an intermediate range of energy [8], we observe similar universal scaling behaviour. To be more explicit, by universality we mean the thermalization time of the form

$$t_{th} \sim \left(\frac{\lambda^{\alpha}}{n_B^2 \ \omega^2}\right)^{\beta},$$
 (1.2)

where as we will see in the following sections  $\alpha$  and  $\beta$  are fixed in terms of p using the equation of the apparent horizon.

The thermalization time is obtained by studying the dynamics of the scalar mesons in Dp-Dq system after the time-dependent baryon injection. This can be done by investigating the fluctuations of the transverse and parallel directions of the Dq-brane to the Dp-brane. We will show that although the equations of the apparent horizon are different in these two cases, the thermalization time-scales are yet the same.

We will also observe that even when the background is not supersymmetric, which means  $q \neq p$ , p+2, p+4, this behaviour still persists. The only change in (1.2) occurs in  $\alpha$  and  $\beta$ . These results confirm, to some extent, the claim of [20] that the thermalization time-scale shows a universal behaviour.

### 2 Review on Dp-Dq system

In the introduction we mentioned that we are interested in studying the thermalization of mesons in strongly coupled field theories in the context of gauge/gravity duality. Therefore in this section we will give a brief review on Dp-Dq brane configurations that are used to explain the meson spectrum in the gauge theory.

The supergravity solution corresponding to near horizon limit of N coincident Dp-branes is [8]

$$ds^{2} = H^{-\frac{1}{2}}(-dt^{2} + dx_{p}^{2}) + H^{\frac{1}{2}}(du^{2} + u^{2}d\Omega_{8-p}^{2}),$$
  

$$d\Omega_{8-p}^{2} = d\theta^{2} + \sin^{2}\theta d\Omega_{k}^{2} + \cos^{2}\theta d\Omega_{7-p-k}^{2},$$
(2.1)

where

$$H(u) = \left(\frac{R}{u}\right)^{7-p}, \quad e^{\phi} = H^{\frac{3-p}{4}}, \quad C_{01...p} = H^{-1},$$
 (2.2)

written in the string frame. The Dilaton field is represented by  $\phi$  and  $C_{01...p}$  is a (p+1)form field coupled to Dp-branes. The length scale R is defined in terms of the string length
scale  $l_s = \sqrt{\alpha'}$  and the string coupling  $g_s = e^{\phi_{\infty}}$ 

$$R^{7-p} = (4\pi)^{\frac{5-p}{2}} \Gamma\left(\frac{7-p}{2}\right) g_s N l_s^{7-p}.$$
 (2.3)

It is easy to see that for the special case p=3, (2.1) reduces to  $AdS_5 \times S^5$ . We are considering  $p \leq 4$  cases because there is no decoupling limit for  $p \geq 5$  and hence no dual gauge theory [8].

According to gauge/gravity duality, a strongly coupled SU(N) SYM theory living on the (p+1)-dimensional worldvolume of N coincident Dp-branes is dual to the supergravity on the above background (2.1) in the large N limit [8]. The isometry group of this background geometry is  $SO(1, p) \times SO(9 - p)$  which in the dual gauge theory corresponds to space-time Lorentz symmetry and R-symmetry, respectively. The gauge theory coupling constant is related to the string coupling and the string length through

$$g_{YM}^2 = 2\pi g_s (2\pi l_s)^{p-3}. (2.4)$$

One can add a stack of  $N_f$  Dq-branes to the above background and study it in the probe limit where  $\frac{N_f}{N}$  is small. The low energy effective action for a Dq-brane in an arbitrary background is [24]

$$S = S_{\text{DBI}} + S_{\text{CS}},$$

$$S_{\text{DBI}} = -\tau_q \int d^{q+1} \xi \ e^{-\phi} \sqrt{-\det(g_{ab} + B_{ab} + 2\pi\alpha' F_{ab})},$$

$$S_{\text{CS}} = \tau_q \int P[\Sigma C^{(n)} e^B] e^{2\pi\alpha' F},$$
(2.5)

where induced metric  $g_{ab}$  and Kalb-Ramond field  $B_{ab}$  are given by

$$g_{ab} = G_{MN} \partial_a X^M \partial_b X^N,$$
  

$$B_{ab} = B_{MN} \partial_a X^M \partial_b X^N.$$
(2.6)

 $a, b, \ldots$  indices are used to describe worldvolume coordinate running over  $0, 1, \ldots, q$ . The capital indices  $M, N, \ldots$  are used to denote space-time coordinates.  $G_{MN}$  is the background metric that is introduced in (2.1) and  $F_{ab}$  is the field strength of the gauge field living on the Dq-brane. Note that in the background we consider here the B-field is zero. In the Chern-Simons part,  $C^{(n)}$  denotes the (n+1)-form Ramond-Ramond potential and  $P[\ldots]$  is the pull-back of the bulk space-time tensors to the Dq-brane worldvolume. The Dq-brane tension is

$$\tau_q = \frac{1}{(2\pi)^q \ l_s^{q+1} g_s} \ . \tag{2.7}$$

In order to embed a probe Dq-brane in the background (2.1) one can introduce the following change of coordinates

$$\rho = u \sin \theta ,$$

$$\sigma = u \cos \theta .$$
(2.8)

Hence the background metric becomes

$$ds^{2} = H^{-\frac{1}{2}}(-dt^{2} + dx_{p}^{2}) + H^{\frac{1}{2}}(d\rho^{2} + \rho^{2}d\Omega_{k}^{2} + d\sigma^{2} + \sigma^{2}d\Omega_{7-p-k}^{2}).$$
 (2.9)

This configuration of the Dp-Dq branes can be schematically shown as

$$t x_1 \dots x_d x_{d+1} \dots x_p \rho \Omega_k \sigma \Omega_{7-p-k}$$

$$Dp \times \times \times \times \times \times \times \times \times \times$$

$$Da \times \times$$

$$(2.10)$$

where

$$q = k + d + 1. (2.11)$$

In order to get a stable background one has to consider the brane configurations which preserve supersymmetry. This condition dictates that q should be p, p + 2, p + 4 and correspondingly k is 1, 2 and 3. Note that this embedding breaks the initial isometry group of the Dp-brane to

$$SO(1,d) \times SO(p-d) \times SO(k+1) \times SO(8-p-k). \tag{2.12}$$

Supersymmetrically embedding  $N_f$  flavour Dq-branes into the background obtained from the near horizon limit of N Dp-branes corresponds to coupling  $N_f$  flavour dynamical matter fields in the fundamental representation (dynamical quarks) to the SYM theory which lives on the Dp-brane. This reduces the number of supersymmetries of the original theory from 16 to 8. The fundamental matter fields which form a  $\mathcal{N}=2$  hypermultiplet arise as the lightest modes of the strings stretched between Dp and Dq-branes.

The spectrum of the fluctuations of a Dq-brane  $(N_f = 1)$  corresponds to the mesonic spectrum of a SU(N) SYM theory in p+1 dimensions which is coupled to a hypermultiplet in the fundamental representation [22, 23]. Note that in the case where q = p + 4, the fundamental fields live on the full p+1 dimensions of the gauge theory. But for q = p+2 and q = p they are confined to a (d+1)-dimensional defect where the Dp and Dq-branes coincide.

#### 3 Thermalization

The thermalization in field theory is the result of the injection of energy into the system. One way to realize it in the context of gauge/gravity duality is to inject baryon charge to the strongly coupled field theory. This has been done for the first time in [20] where the dynamics of mesons or in other words the thermalization of mesons in D3-D7 system has been studied. In this section we will elaborate more on this process by generalizing it to Dp-Dq system.

#### 3.1 Baryon injection

The injection of baryons into the strongly coupled field theory is presented in the gauge/gravity context by throwing byronically charged F-strings on the flavour Dq-brane. The end points of the strings stretched between Dq and Dp-branes resemble quarks in the gauge theory. The quarks are assumed to be massless. They act as a source for the background gauge field on the flavour brane. The injection of quarks into the system results in a sudden change in the baryon number chemical potential.

In the AdS/CFT correspondence, static baryon chemical potential is identified with the time component of the gauge field in the following way

$$\mu = \int_0^\infty du \, \partial_u A_t(u) ,$$
  
=  $A_t(\infty) - A_t(0) ,$  (3.1)

where u represents the radial direction in the AdS space.<sup>1</sup> Therefore in order to explain a time-dependent chemical potential which results from the injection of quarks to the system we must add the source term

$$S_{\text{current}} = \int d^{q+1}\xi \sqrt{-g} A_a J^a, \qquad (3.2)$$

to the action (2.5). Note that the currents,  $J^a$ s, are now time-dependent and will produce time-dependent gauge fields.

Since the end points of the open strings move on the light geodesics, without loss of generality, the current components can be fixed to be  $J^{\rho}$  and  $J^{t}$ . We assume that the massless quarks are moving along the null line  $x^{-}$  defined as

$$x^{\pm} = t \pm z = t \mp \int H^{\frac{1}{2}} d\rho .$$
 (3.3)

Hence the currents are only functions of  $x^-$ .

With the above assumptions the Dq-brane DBI action for the background (2.9) in the presence of the source terms reduces to<sup>2</sup>

$$S = -\tau_q V_d \text{Vol}(S^k) \int dt d\rho \left[ H^{\frac{p-q+2(k-1)}{4}} \rho^k \sqrt{1 - (2\pi\alpha')^2 F_{t\rho}^2} - A_t j^t - A_\rho j^\rho \right]. \tag{3.4}$$

<sup>&</sup>lt;sup>1</sup>Note that this identification is done in the gauge where  $A_u$  is zero.

<sup>&</sup>lt;sup>2</sup>Note that  $\tau_q j^a = \frac{R^8}{r^5} J^a$ .

To solve the equation of motion for the gauge field we can consistently choose

$$j^{\rho} = -H^{\frac{-1}{2}}j^{t} = g'(x^{-}), \tag{3.5}$$

which satisfies the current conservation equation

$$\nabla_a J^a = 0. ag{3.6}$$

Therefore the classical time-dependent solution for the gauge field is

$$(2\pi\alpha')F_{t\rho} = \frac{g}{\sqrt{g^2 + (2\pi\alpha')^2 \rho^{2k} H^{\frac{p-q+2(k-1)}{2}}}},$$
(3.7)

where the following identity has been used

$$\partial_{+} \left( \rho^{k} H^{\frac{p-q+2(k-1)}{4}} \frac{(2\pi\alpha')^{2} F_{t\rho}}{\sqrt{1 - (2\pi\alpha')^{2} F_{t\rho}^{2}}} \right) = 0.$$
 (3.8)

It's been already explained that the sudden change in the baryon number density can locally model the collision of the heavy nuclei at RHIC. The baryon number density is defined as  $n_B = \frac{n}{N}$  where n is the quark density number. In our set-up the quarks live on the (d+1)-dimensional defect field theory. Therefore the quark density number is  $n = \frac{\tilde{n}}{Vol_d}$  where  $\tilde{n}$  is the quark number (number of open string end points). The time-dependent quark number can be calculated by taking the integral of the time component of the external source over volume space of the Dq-brane

$$\tilde{n} = \int d^q \xi \sqrt{-g} J^t \ . \tag{3.9}$$

Therefore replacing (3.5) in the above equation leads to

$$g(x^{-}) = \frac{\gamma}{\text{Vol}(S^{k})} (2\pi)^{\frac{q-p}{2}} (2\pi\alpha')^{\frac{q-p+4}{2}} \lambda \ n_{B}(x^{-}), \tag{3.10}$$

where  $\lambda = g_{YM}^2 N$  and  $\gamma$  is a parameter that should be fixed according to the convention. For instance in the case of D3-D7 system [20],  $\gamma$  is  $\frac{1}{4\pi}$ .

So far all the calculations have been done for general values of p and q. In order to have a stable background in the limit of zero temperature, we consider supersymmetric systems where q = p + 4, p + 2, p and k = 3, 2, 1, respectively. For such configurations we have

$$p - q + 2(k - 1) = 0. (3.11)$$

Therefore one can see that the p and q dependence in  $F_{t\rho}$  and  $g(x^{-})$  disappears and they reduce to

$$(2\pi\alpha')F_{t\rho} = \frac{g}{\sqrt{g^2 + (2\pi\alpha')^2 \rho^{2k}}},$$
(3.12)

and

$$g(x^{-}) = \begin{cases} \frac{\gamma}{\text{Vol}(S^{3})} (2\pi)^{2} (2\pi\alpha')^{4} \lambda \ n_{B}(x^{-}), & q = p + 4, \\ \frac{\gamma}{\text{Vol}(S^{2})} (2\pi) (2\pi\alpha')^{3} \lambda \ n_{B}(x^{-}), & q = p + 2, \\ \frac{\gamma}{\text{Vol}(S^{1})} (2\pi\alpha')^{2} \lambda \ n_{B}(x^{-}), & q = p. \end{cases}$$
(3.13)

Time-dependent gauge field on the Dq-brane modifies the metric observed by the open strings. This can be explicitly seen by using the DBI action. We will show in the following section that this field configuration can create an apparent horizon on the Dq-brane which signals the thermalization in the dual strongly coupled field theory.

#### 3.2 Apparent horizon formation

Thermalization usually refers to an increase in the temperature of a thermodynamic system from zero to a nonzero value. We know that in the context of gauge/gravity duality the vacuum state in field theory corresponds to pure AdS space in the bulk and the thermal mixed state is dual to the AdS-black hole. So the horizon formation in the bulk can be considered as a signal of thermalization in the dual gauge theory. In order to study the process of thermalization in a strongly coupled field theory the gauge/gravity techniques can be applied [9–11]. In this framework the horizon formation usually refers to the apparent horizon formation since in contrast to the event horizon it can be obtained locally and the global knowledge of the bulk solution is not needed [9, 10].

In the Dp-Dq set-up the meson dynamics can be used to study thermalization in the field theory. The mesons are the open strings (scalar or gauge fields) living on the Dq-brane [25]. The Dq-brane can fluctuate in the directions which are transverse or parallel to the background Dp-branes. In the following subsections we will study the dynamics of these two kinds of scalar fluctuations in the presence of a time-dependent gauge field on the Dq-brane. It leads to calculating the modified metric on the Dq-brane which is called the open string metric [26]. Thus we can obtain the equation of the apparent horizon and compute the thermalization time-scale.

#### 3.2.1 Transverse fluctuation

Consider Dq-brane embedding (2.9). The transverse fluctuations are along  $\sigma$  and  $\Omega_{7-p-k}$  directions. Due to the isometry SO(8-p-k) we assume only  $\delta\sigma$  to be nonzero. This plays the role of the scalar meson whose dynamics we would like to study. We expand the DBI action (2.5) to quadratic order in  $\delta\sigma$  in the background solution  $F_{t\rho}$  (3.7). Note that since the RR field does not couple to the end point of the string on the brane it will not enter in the calculations for the open string metric. Therefore from now on in our calculations we neglect the CS action. Hence the action for the scalar meson is

$$S = -\frac{1}{2} \int dt d\rho d^d x^i d^k \theta^\alpha \sqrt{-\tilde{g}} \ \tilde{g}^{ab} \partial_a \delta \sigma \partial_b \delta \sigma, \tag{3.14}$$

where

$$-\tilde{g}_{tt} = \tau_{q}^{\frac{2}{q-1}} [1 - (2\pi\alpha')^{2} F_{t\rho}^{2}]^{\frac{2-q}{1-q}} H^{\frac{p-q}{2(q-1)}},$$

$$\tilde{g}_{\rho\rho} = \tau_{q}^{\frac{2}{q-1}} [1 - (2\pi\alpha')^{2} F_{t\rho}^{2}]^{\frac{2-q}{1-q}} H^{\frac{p+q-2}{2(q-1)}},$$

$$\tilde{g}_{ij} = \tau_{q}^{\frac{2}{q-1}} [1 - (2\pi\alpha')^{2} F_{t\rho}^{2}]^{\frac{1}{1-q}} H^{\frac{p-q}{2(q-1)}} \delta_{ij},$$

$$\tilde{g}_{\alpha\beta} = \tau_{q}^{\frac{2}{q-1}} [1 - (2\pi\alpha')^{2} F_{t\rho}^{2}]^{\frac{1}{1-q}} \rho^{2} H^{\frac{p+q-2}{2(q-1)}} \hat{g}_{\alpha\beta}.$$
(3.15)

 $i=1,\ldots,d$  is the spatial directions of the (d+1)-dimensional defect.  $\theta^{\alpha}$   $(\alpha=1,\ldots,k)$  is the angular variable on  $S^k$  and  $\hat{g}_{\alpha\beta}$  is the metric on the unit k-sphere. The surface of the presumably apparent horizon, which is defined locally as a surface whose area variation vanishes along the null ray which is normal to the surface, is given by<sup>3</sup>

$$V_{\text{surface}} = \int d^{d}x^{i} d^{k} \theta^{\alpha} \left( \prod_{i=1}^{d} \tilde{g}_{ii} \prod_{\alpha=1}^{k} \tilde{g}_{\alpha\alpha} \right)^{\frac{1}{2}}$$

$$= V_{d} \text{Vol}(S^{k}) \frac{\tau_{q}}{2\pi\alpha'} H^{\frac{1}{2}} \sqrt{g^{2} + (2\pi\alpha')^{2} \rho^{2k} H^{\frac{p-q+2(k-1)}{2}}},$$
(3.16)

where this should satisfy

$$dV|_{dt=-dz} = 0. (3.17)$$

Therefore the equation for the apparent horizon reduces to

$$4gg'H^{\frac{1}{2}}\rho + (p-7)g^2 + \nu(2\pi\alpha')^2H^{\frac{p-q+2(k-1)}{2}}\rho^{2k} = 0,$$
(3.18)

where  $\nu = \frac{1}{2} [(p-7)(p-q+2k)+4k]$ . We call this the master equation. For supersymmetric cases, the master equation can be simplified further

$$4gg'H^{\frac{1}{2}}\rho + (p-7)g^2 + (p-7+2k)(2\pi\alpha')^2\rho^{2k} = 0.$$
(3.19)

Since working with z coordinate is more appropriate we replace  $\rho$  with z in (3.19) and get

$$2(5-p)gg'z + (p-7)g^2 + (2\pi\alpha')^2(p-7+2k)\left(\frac{5-p}{2}\right)^{\frac{4k}{p-5}}z^{\frac{4k}{p-5}}R^{\frac{2k(p-7)}{p-5}} = 0, \quad (3.20)$$

where the length scale R can be written in terms of the t'Hooft coupling constant  $\lambda$  as

$$R^{\frac{p-7}{p-5}} = \sqrt{2}(2\pi\alpha')(2\pi)^{\frac{p+1}{2(p-5)}} \Gamma\left(\frac{7-p}{2}\right)^{\frac{1}{5-p}} \lambda^{\frac{1}{5-p}}.$$
 (3.21)

The formula (3.20) is the main equation that will be used to calculate the thermalization time for the transverse fluctuations.

$$V_{\text{surface}} = V_d \text{Vol}(S^k) \frac{\tau_q}{2\pi\alpha'} H^{\frac{1}{2}} \sqrt{g^2 + (2\pi\alpha')^2 \rho^{2k}}.$$

<sup>&</sup>lt;sup>3</sup>One can see that for the supersymmetric cases the apparent horizon surface area simplifies to

#### 3.2.2 Parallel fluctuation

As it was mentioned earlier in addition to the transverse fluctuations of the Dq-brane one can also study the ones which are parallel to the background Dp-branes. These parallel fluctuations exist only for q = p + 2 and p. This can be easily seen from the configuration (2.10). Note that for q = p + 2 we have d = p - 1 and the only parallel fluctuation is along p direction. For q = p there are two directions for parallel fluctuations, p - 1 and p, and the fundamental matter lives on the spatial (d = p - 2)-dimensional defect.

Similar to the transverse fluctuations the DBI action reduces to

$$S = -\frac{1}{2} \int dt d\rho d^d x^i d^k \theta^\alpha \sqrt{-\tilde{g}} \ \tilde{g}^{ab} \partial_a \delta x^n \partial_b \delta x^n \,, \tag{3.22}$$

where  $\delta x^n$  represents small fluctuation parallel to Dp-branes. The components of the metric are

$$-\tilde{g}_{tt} = \tau_q^{\frac{2}{q-1}} [1 - (2\pi\alpha')^2 F_{t\rho}^2]^{\frac{2-q}{1-q}} H^{\frac{p-q-4}{2(q-1)}},$$

$$\tilde{g}_{\rho\rho} = \tau_q^{\frac{2}{q-1}} [1 - (2\pi\alpha')^2 F_{t\rho}^2]^{\frac{2-q}{1-q}} H^{\frac{p+q-6}{2(q-1)}},$$

$$\tilde{g}_{ij} = \tau_q^{\frac{2}{q-1}} [1 - (2\pi\alpha')^2 F_{t\rho}^2]^{\frac{1}{1-q}} H^{\frac{p-q-4}{2(q-1)}} \delta_{ij},$$

$$\tilde{g}_{\alpha\beta} = \tau_q^{\frac{2}{q-1}} [1 - (2\pi\alpha')^2 F_{t\rho}^2]^{\frac{1}{1-q}} \rho^2 H^{\frac{p+q-6}{2(q-1)}} \hat{g}_{\alpha\beta}.$$
(3.23)

Therefore we can now calculate the equation for the apparent horizon area<sup>4</sup>

$$V_{\text{surface}} = V_d \text{Vol}(S^k) \frac{\tau_q}{2\pi\alpha'} H^{-\frac{1}{2}} \sqrt{g^2 + (2\pi\alpha')^2 \rho^{2k} H^{\frac{p-q+2(k-1)}{2}}}.$$
 (3.24)

Similar to the previous subsection we apply the condition (3.17) and the master equation for the parallel fluctuations reduces to

$$2(5-p)gg'z - (p-7)g^2 - (2\pi\alpha')^2(p-7-2k)\left(\frac{5-p}{2}\right)^{\frac{4k}{p-5}}z^{\frac{4k}{p-5}}R^{\frac{2k(p-7)}{p-5}} = 0, \quad (3.25)$$

where the length scale R was introduced in (3.21). The master equation (3.25) is very similar to the one obtained for transverse fluctuations (3.20) except for a couple of sign differences. Note that we have calculated the master equation for the supersymmetric cases.

Before closing this section we would like to emphasize on the observation that for the general values of q, not necessarily supersymmetric ones, the only change in the equations (3.20) and (3.25) appears in the power of z, in their last term. Since p-q+2(k-1) is not zero any more in the non-supersymmetric brane configurations the power of z,  $\frac{4k}{p-5}$ , is replaced by  $\frac{p-q+2(3k-1)}{p-5}$ . Note that this is a rough argument and we have ignored discussing the stability of the backgrounds. We will see from our results in the following section that this indicates that the thermalization time scale behaviour does not rely on supersymmetry.

$$V_{\text{surface}} = V_d \text{Vol}(S^k) \frac{\tau_q}{2\pi\alpha'} H^{-\frac{1}{2}} \sqrt{g^2 + (2\pi\alpha')^2 \rho^{2k}}.$$

<sup>&</sup>lt;sup>4</sup>In supersymmetric cases we have

### 4 Baryon injection model

To be able to determine the time-scale of thermalization, the source function  $g(x^{-})$  should be specified. Similar to the function considered in [20], we assume

$$g(x^{-}) = \begin{cases} 0 & x^{-} < 0 \\ g_{\text{max}} & \omega x^{-} & 0 < x^{-} < \frac{1}{\omega} \\ g_{\text{max}} & \frac{1}{\omega} < x^{-} \end{cases}$$
(4.1)

Using the equation (3.10) it shows that the baryon density  $n_B$  is zero in the beginning which resembles the situation before the collision of heavy ions. At  $x^- = 0$  it starts changing to reach a maximum value at  $x^- = \frac{1}{\omega}$  and remains constant since then. The maximum value of source function  $g_{\text{max}}$  is given by (3.13) where  $n_B(x^-)$  is replaced by its constant maximum value,  $n_B$ 

$$g_{\text{max}} = \frac{\gamma}{\text{Vol}(S^k)} (2\pi)^{\frac{q-p}{2}} (2\pi\alpha')^{\frac{q-p+4}{2}} \lambda \ n_B.$$
 (4.2)

This is similar to the situation where the baryonically-charged heavy ions approach each other and the baryon number chemical potential changes locally.

The time-dependence of the baryon number can be chosen to have a general form,  $(\omega x^-)^n$ , where  $n \ge 0.5$  This is a dimensionless combination and guarantees that  $g(x^-)$  gets its maximal value at  $x^- = \frac{1}{\omega}$ .

We consider the supersymmetric configurations which fall into these categories:

#### 4.1 (p, p + 4) system

In this set-up of brane configuration we consider q to be p+4 where  $0 \le p < 5$ . As it's been shown in (2.10) there is no parallel fluctuations in this case and one needs to consider only the transverse fluctuations.

#### 4.1.1 Transverse fluctuation

If we define  $g(x^-) = g_{\text{max}} y(x^-)$  where  $g_{\text{max}}$  is given by (4.2) and set k to be 3, the master equation (3.20) becomes

$$(p-7)y^{2} + 2(5-p)yy'z + \frac{1}{8\gamma^{2}}(p-1)\left(\frac{5-p}{2}\right)^{\frac{12}{p-5}}(2\pi)^{\frac{3(p+1)}{p-5}} \Gamma\left(\frac{7-p}{2}\right)^{\frac{6}{5-p}} \frac{\lambda^{\frac{2(p-2)}{5-p}}}{n_{B}^{2}} z^{\frac{12}{p-5}} = 0.$$

$$(4.3)$$

To warm up, we start with the D4-D8 brane configuration.

$$t_{th} \sim \left(\frac{\lambda^{\alpha}}{n_B^2 \omega^{2n}}\right)^{\beta}.$$

Note that, in contrast to transverse fluctuations, n = 0 does not give a real solution for parallel ones. We will see this in subsection 4.2.2.

<sup>&</sup>lt;sup>5</sup>If we assume  $g(x^-)$  to scale as  $g_{\text{max}}(\omega x^-)^n$  the change in the thermalization time appears in the overall power  $\beta$  and the power of  $\omega$  as

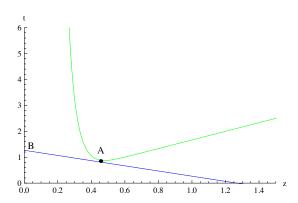


Figure 1. The location of the apparent horizon for  $t < z + \frac{1}{\omega}$  in the t-z plane. The green curve represents (4.5a) and the blue line shows the tangent null ray. The parameters are chosen as  $\omega = 0.5$  and  $\frac{\lambda^4}{\gamma^2 n_B^2} = \frac{1}{\omega^{12}}$  for D4-D8 system.

#### **D4-D8**

The master equation is then

$$\frac{3\lambda^4}{4096\pi^{12}\gamma^2 n_B^2 z^{12}} - 3y^2 + 2zyy' = 0. \tag{4.4}$$

The location of the apparent horizon on the D8-brane is obtained by solving the above equation and determine t in terms of z. We get

$$t = \frac{4z}{3} + \frac{1}{192} \sqrt{4096z^2 + \frac{9\lambda^4}{z^{12}\pi^{12}\gamma^2 n_B^2 \omega^2}}, \qquad t < z + \frac{1}{\omega}, \qquad (4.5a)$$

$$z = 0.16 \left(\frac{\lambda^2}{\gamma n_B}\right)^{\frac{1}{6}}, \qquad t > z + \frac{1}{\omega}. \tag{4.5b}$$

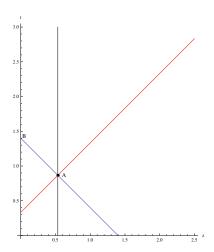
The curve (4.5a) and the line (4.5b) are plotted in figures 1 and 2 in green and black colours, respectively. The goal of these calculations is to compute the thermalization time in the boundary field theory. This in fact is the time that a boundary observer starts to see the apparent horizon formation on the D8-brane. For  $t < z + \frac{1}{\omega}$  the null ray which is tangent to (4.5a) is the earliest one that conveys the information to the boundary. It has been shown by the blue line in figure 1. The other null rays which cross (4.5a) reach the boundary at later times. The tangent point A coordinates are

$$z_A = 0.46 \left(\frac{\lambda^2}{\gamma n_B \omega}\right)^{\frac{1}{7}}, \quad t_A = 0.81 \left(\frac{\lambda^2}{\gamma n_B \omega}\right)^{\frac{1}{7}}.$$
 (4.6)

So the thermalization time is

$$t_{th} = t_A + z_A \sim \left(\frac{\lambda^2}{n_B \omega}\right)^{\frac{1}{7}}. (4.7)$$

An interesting observation is that the thermalization time depends only on  $\lambda$  and the baryon injection parameters  $n_B$  and  $\omega$ .



**Figure 2**. The location of the apparent horizon for  $t>z+\frac{1}{\omega}$  in the t-z plane. The black, red and blue lines represent (4.5b),  $t=z+\frac{1}{\omega}$  and the earliest null ray which reaches the boundary. The parameters are chosen as  $\omega=3$  and  $(\frac{\lambda^2}{\gamma n_B})^{\frac{1}{6}}=\frac{10}{\omega}$  for D4-D8 system.

Similar argument can be made for (4.5b). In that case the thermalization time is determined by the point A coordinates in figure 2. Point A is where the line (4.5b) crosses the ingoing null ray  $t = z + \frac{1}{\omega}$ , red line in that figure. So the null ray which passes through this point and reaches the boundary gives the earliest time that a boundary observer sees the horizon formation. This happens especially if

$$\left(\frac{\lambda^2}{n_B\omega}\right)^{\frac{1}{7}} \gg \frac{1}{\omega}\,,\tag{4.8}$$

and thermalization time becomes

$$t_{th} = 2z_A + \frac{1}{\omega} \,, \tag{4.9}$$

where

$$t_A = z_A + \frac{1}{\omega} \,, \tag{4.10a}$$

$$z_A = 0.16 \left(\frac{\lambda^2}{\gamma n_B}\right)^{\frac{1}{6}}.\tag{4.10b}$$

Inequality (4.8) then indicates

$$t_{th} \sim \left(\frac{\lambda^2}{n_B}\right)^{\frac{1}{6}}.\tag{4.11}$$

#### Dp-D(p+4)

The above calculation can be generalized to other values of p. We conclude that

•  $0 \le p (\ne 1) < 5$ 

$$t_{th} \sim \left(\frac{\lambda^{\frac{2(p-2)}{5-p}}}{n_B^2 \omega^2}\right)^{\frac{5-p}{2(11-p)}},$$
 (4.12)

and if  $\left(\frac{\lambda^{\frac{2(p-2)}{5-p}}}{n_B^2\omega^2}\right)^{\frac{5-p}{2(11-p)}} \gg \frac{1}{\omega}$  we have

$$t_{th} \sim \left(\frac{\lambda^{\frac{2(p-2)}{5-p}}}{n_B^2}\right)^{\frac{5-p}{12}}.$$
 (4.13)

• p = 1

Setting p = 1 the system behaves differently from the other values of p. This is due to the fact that at p = 1 one of the terms in the master equation vanishes and this trivializes the calculations.

## 4.2 (p, p + 2) system

Consider a D(p+2)-brane supersymmetrically embedded in the background obtained from the near horizon geometry of N coincident Dp-branes where  $1 \le p < 5$ . k equals 2 and the dual gauge theory is (p+1)-dimensional, but the fundamental hypermultiplet has been introduced on a (d+1=p)-dimensional defect. There are three and one set of fluctuations along transverse and parallel directions, respectively. In the following subsections we will discuss them individually.

### 4.2.1 Transverse fluctuations

For this configuration the master equation (3.20) reduces to

$$2(5-p)yy'z + (p-7)y^{2} + \frac{16}{9}(p-3)(2\pi)^{\frac{2(p+1)}{p-5}} \left(\frac{5-p}{2}\right)^{\frac{8}{p-5}} \Gamma\left(\frac{7-p}{2}\right)^{\frac{4}{5-p}} \frac{\lambda^{\frac{2(p-3)}{5-p}} z^{\frac{8}{p-5}}}{\gamma^{2} n_{B}^{2}} = 0.$$

$$(4.14)$$

If we repeat the same calculations done in the previous section, the thermalization time for arbitrary values of p is

• 
$$1 \le p \ne 3 < 5$$

$$t_{th} \sim \left(\frac{\lambda^{\frac{2(p-3)}{5-p}}}{n_B^2 \omega^2}\right)^{\frac{5-p}{2(9-p)}},$$
 (4.15)

and if the time scale lies in the limit where  $\left(\frac{\lambda^{\frac{2(p-3)}{5-p}}}{n_B^2\omega^2}\right)^{\frac{5-p}{2(9-p)}} \gg \frac{1}{\omega}$ , we obtain

$$t_{th} \sim \left(\frac{\lambda^{\frac{2(p-3)}{5-p}}}{n_B^2}\right)^{\frac{p-5}{8}}.$$
 (4.16)

• p = 3

Note that similar to p = 1 case in the Dp-D(p + 4) system, choosing p = 3 sets one of the terms in the master equation to zero and the thermalization time can not be obtained. Interestingly we will see in the following that the parallel fluctuation produces the result we expect for p = 3.

#### 4.2.2 Parallel fluctuations

As it was mentioned earlier there is only one parallel fluctuation for Dp-D(p+2) system which is along the  $x^p$  direction. The master equation (3.25) for k=2 reduces to

$$2(5-p)gg'z - (p-7)g^2 - (2\pi\alpha')^2(p-11)\left(\frac{5-p}{2}\right)^{\frac{8}{p-5}}z^{\frac{8}{p-5}}R^{\frac{4(p-7)}{p-5}} = 0.$$
 (4.17)

Let us start with an explicit example.

#### **D4-D6**

Setting p = 4 the apparent horizon equation becomes

$$3y^2 + 2zyy' + \frac{7\lambda^2}{36\pi^8\gamma^2 n_B^2 z^8} = 0. {(4.18)}$$

This looks similar to the master equation for D4-D8 brane system (4.4). We follow the same analogy here. The solution to the above equation which specifies the location of the apparent horizon reads

$$t = \frac{2z}{3} + \frac{1}{3}\sqrt{z^2 - \frac{7\lambda^2}{12z^8\pi^8\gamma^2n_B^2\omega^2}}, \quad t < z + \frac{1}{\omega}.$$
 (4.19)

Note that there is no real solution for z when  $t > z + \frac{1}{\omega}$ . Thus the thermalization time is only fixed by the solution (4.19). For  $t < z + \frac{1}{\omega}$  the expression under the square root must be positive. Fortunately we can always find a solution for z that respects this condition. The coordinates of the tangent point A to the curve (4.19) will determine the thermalization time. This has been fully discussed in D4-D8 case so we state the result. The coordinates of the tangent point A are

$$z_A = \frac{0.39}{\gamma^{1/5}} \left(\frac{\lambda}{n_B \omega}\right)^{1/5},$$

$$t_A = \frac{0.45}{\gamma^{1/5}} \left(\frac{\lambda}{n_B \omega}\right)^{1/5}.$$

$$(4.20)$$

Hence the thermalization time on the boundary reads

$$t_{th} = t_A + z_A \sim \left(\frac{\lambda}{n_B \omega}\right)^{1/5}.$$
 (4.21)

If we set p=4 in the result for the transverse fluctuation (4.15) we get the same thermalization time scale, up to a numeric constant. It is interesting that although the apparent horizon equations are different for transverse and parallel fluctuations their thermalization times observed on the boundary exactly resemble each other.

### Dp-D(p+2)

The above result can be generalized to other values of p and we get

$$t_{th} \sim \left(\frac{\lambda^{\frac{2(p-3)}{5-p}}}{n_B^2 \omega^2}\right)^{\frac{5-p}{2(9-p)}}$$
 (4.22)

One can conclude that the thermalization time for parallel fluctuations and transverse ones (4.15) look completely alike for various values of p. An interesting observation is that the time scale (4.22) works fine for p=3 case in contrast to the transverse one. We can see that for p=3 there is no dependence on the t'Hooft coupling  $\lambda$  in the thermalization time scale.

#### 4.3 (p,p) system

Consider the brane configuration where q=p. In order to supersymmetrically embed one Dp-brane in the background geometry of N Dp-branes we choose  $2 \le p < 5$ . As it was mentioned earlier the fundamental hypermultiplet is confined to a (d+1=p-1)-dimensional surface. k=1 for this set-up. Therefore there are two sets of both, parallel and transverse fluctuations which we will elaborate more on them in the following.

#### 4.3.1 Transverse fluctuations

The master equation (3.20) in this case reduces to

$$(p-7)y^{2} + 2(5-p)yy'z + \frac{1}{2\gamma^{2}}(p-5)\left(\frac{5-p}{2}\right)^{\frac{4}{p-5}}(2\pi)^{\frac{3(p-3)}{p-5}} \Gamma\left(\frac{7-p}{2}\right)^{\frac{2}{5-p}} \frac{\lambda^{\frac{2(p-4)}{5-p}}}{n_{P}^{2}} z^{\frac{4}{p-5}} = 0.$$

$$(4.23)$$

The thermalization time-scale for general values of p becomes

•  $2 \le p < 5$ 

$$t_{th} \sim \left(\frac{\lambda^{\frac{2(p-4)}{5-p}}}{n_B^2 \omega^2}\right)^{\frac{5-p}{2(7-p)}},$$
 (4.24)

and in the limit where  $\left(\frac{\lambda^{\frac{2(p-4)}{5-p}}}{n_B^2\omega^2}\right)^{\frac{5-p}{2(7-p)}} \gg \frac{1}{\omega}$ , we have

$$t_{th} \sim \left(\frac{\lambda^{\frac{2(p-4)}{5-p}}}{n_B^2}\right)^{\frac{p-5}{4}}.$$
 (4.25)

#### 4.3.2 Parallel fluctuation

For Dp-Dp system there are two directions for parallel fluctuations. The master equation (3.25) reads

$$-(p-7)y^{2} + 2(5-p)yy'z - \frac{1}{2\gamma^{2}}(p-9)\left(\frac{5-p}{2}\right)^{\frac{4}{p-5}}(2\pi)^{\frac{3(p-3)}{p-5}} \Gamma\left(\frac{7-p}{2}\right)^{\frac{2}{5-p}} \frac{\lambda^{\frac{2(p-4)}{5-p}}}{n_{R}^{2}} z^{\frac{4}{p-5}} = 0.$$

$$(4.26)$$

Interestingly we observe that similar to Dp-D(p+2) system, the thermalization time for parallel fluctuations has the same scaling behaviour as the transverse fluctuations. Therefore for general p we obtain

• 
$$2 \le p < 5$$

$$t_{th} \sim \left(\frac{\lambda^{\frac{2(p-4)}{5-p}}}{n_B^2 \omega^2}\right)^{\frac{5-p}{2(7-p)}}.$$
(4.27)

#### 4.4 Concluding remarks

We conclude this section mentioning some interesting observations. It's been already mentioned that the thermalization time for parallel and transverse fluctuations are the same, up to a numeric coefficient, even though the apparent horizon equations are different. The theories we discussed here are not necessarily conformal and their dual background geometries are not AdS. But the time-scale behaviour still gets the same form as an AdS background with a conformal field theory dual such as D3-D7 system. Even for general values of q for which the solution is not supersymmetric the general form of the thermalization time (1.2) persists. This seems to approve the claim that the thermalization time-scale behaves universally. But one may ask to what extent this discussion works. We leave it as an open question.

Let us consider the case where p=3. As it is well known the dual gauge theory is  $\mathcal{N}=4$  SYM theory which lives on the 4-dimensional boundary of AdS space. If we require to preserve supersymmetry we can add three types of probe flavour branes to this background. They are D3, D5, and D7-branes. Adding these branes will modify the dual gauge theory to  $\mathcal{N}=4$  coupled to  $\mathcal{N}=2$  fundamental hypermultiplet. The dynamical quarks live on (1+1), (2+1) and (3+1)-dimensional defects for q=3, 5 and 7, respectively. Interestingly we see that although the 4-dimensional gauge theory is the same for all of these brane set-ups these theories produce different time scales for thermalization which are

$$t_{th} \sim \begin{cases} \left(\frac{\lambda}{n_B^2 \omega^2}\right)^{\frac{1}{8}}, & D3 - D7, \\ \left(\frac{1}{n_B^2 \omega^2}\right)^{\frac{1}{6}}, & D3 - D5, \\ \left(\frac{1}{n_B^2 \omega^2 \lambda}\right)^{\frac{1}{4}}, & D3 - D3. \end{cases}$$
(4.28)

Therefore we can conclude that the thermalization time scales differently with the t'Hooft coupling,  $\lambda$ . The main difference between these theories come from the fact that the dynamical quarks are confined to defects with different dimensions.

Another intriguing observation is when we consider the dimension of the defect to be (2+1). The brane configurations to get it are D2-D6, D3-D5 and D4-D4. One can see that in all of these cases the dependence on  $\lambda$  in the thermalization time disappears. It is very interesting that for this specific choice all the brane configurations give the same value for the thermalization time which is  $\left(\frac{1}{n_B^2\omega^2}\right)^{\frac{1}{6}}$ . Note that this doesn't happen in the other dimensions for the defect such as (1+1) and (0+1) dimensions.

In this paper we have studied thermalization for the scalar mesons. One can generalize this calculation and study the dynamics of the vector mesons. Moreover the universal behaviour of the thermalization time-scale can be investigated in other gravity backgrounds such as anisotropic Lifshitz ones. It is also interesting to study the time-scale of the change in the temperature if one starts from thermal field theory, dual to black hole background, instead of zero-temperature one and then injects energy into it. We postpone these to the future works.

## Acknowledgments

We would like to thank N. Abbasi, M. Alishahiha, A. Davody and M. M. Sheikh-Jabbari for fruitful comments and discussions.

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