Published for SISSA by 🖄 Springer

RECEIVED: April 25, 2017 REVISED: January 11, 2018 ACCEPTED: February 5, 2018 PUBLISHED: February 20, 2018

Runaway relaxion monodromy

Liam McAllister,^{*a*} Pedro Schwaller,^{*b,c*} Geraldine Servant,^{*b,d*} John Stout^{*a*} and Alexander Westphal^{*b*}

^a Department of Physics, Cornell University, Ithaca, New York, 14853, U.S.A.
^b DESY, Notkestrasse 85, D-22607 Hamburg, Germany
^c PRISMA Cluster of Excellence, Institut für Physik, Johannes Gutenberg-Universität, Mainz, Germany
^d II. Institute of Theoretical Physics, Univ. Hamburg, D-22761 Hamburg, Germany *E-mail:* mcallister@cornell.edu, pedro.schwaller@uni-mainz.de, geraldine.servant@desy.de, jes554@cornell.edu, alexander.westphal@desy.de

ABSTRACT: We examine the relaxion mechanism in string theory. An essential feature is that an axion winds over $N \gg 1$ fundamental periods. In string theory realizations via axion monodromy, this winding number corresponds to a physical charge carried by branes or fluxes. We show that — in the context of NS5-brane axion monodromy — this charge backreacts on the compact space, ruining the structure of the relaxion action. In particular, the barriers generated by strong gauge dynamics have height $\propto e^{-N}$, so the relaxion does not stop when the Higgs acquires a vev. Backreaction of monodromy charge can therefore spoil the relaxion mechanism. We comment on the limitations of technical naturalness arguments in this context.

KEYWORDS: Discrete Symmetries, Effective Field Theories, Flux compactifications

ARXIV EPRINT: 1610.05320





Contents

1	Intr	2			
2	Rel 2.1 2.2 2.3 2.4	Origin CHAI	rsymmetric models	12 12 13 13 15	
3	Relaxion monodromy 3.1 Axion monodromy in string theory			16	
	3.1	16			
	3.2	19			
	3.3	Fiveb	23		
4	Mic	26			
	4.1	Overv	view of microphysical constraints	27	
	4.2	Conse	equences of D3-brane backreaction	29	
		4.2.1	Universal effect on geometry	29	
		4.2.2	Tadpole constraint	30	
		4.2.3	Suppression from warping	31	
		4.2.4	Suppression from backreaction	32	
		4.2.5	Effects on the moduli potential	34	
		4.2.6	Effects on axion decay constants	36	
		4.2.7	Classical annihilation of the dipole	37	
		4.2.8	Antibrane tunneling and annihilation	38	
		4.2.9	Light NS5-brane modes	39	
5	Discussion and outlook				
	5.1	Exact	t discrete shift symmetries for relaxions	41	
	5.2	Const	traints from the Weak Gravity Conjecture	44	
6	Cor	Conclusions 4			
\mathbf{A}	Axions in string theory				
в	Necessity of the intersection				
\mathbf{C}	Type IIB supergravity with fivebranes				
	C.1	Conve	entions for type IIB supergravity	50	
	C.2	Einste	ein-frame potentials for fivebranes	51	
D	Backreaction on the internal space 54				

1 Introduction

Why is the Higgs mass so small? Graham, Kaplan, and Rajendran (GKR) have proposed a novel solution to the electroweak hierarchy problem, the *relaxion mechanism*, in which the evolution of an axion field ϕ drives the Higgs mass m_h to relax dynamically to a value much smaller than the cutoff, $|m_h^2| \ll M^2$ [1]. Achieving a large hierarchy in this way requires very small dimensionless couplings, as well as field excursions $\Delta \phi \gg M$, but GKR argued that the requisite couplings are technically natural.

In this work, we study the impact of ultraviolet completion on the relaxion mechanism. The large field excursions required by the mechanism, while technically natural in effective field theory, turn out to be *source terms* in string theory! Winding an axion ϕ over $N \gg 1$ fundamental periods leads to the accumulation of N units of monodromy charge, providing a large source term in ten dimensions. This changes the shape of the compactification and alters the couplings of the effective theory, eliminating the barrier that is needed to stop the relaxion once the Higgs acquires a vev.

The root of the problem is that new states linked to the monodromy charge, which are too massive in the initial configuration to be visible, are eventually drawn below the cutoff M. These new light states induce changes in the couplings of the effective theory. In particular, the gauge coupling $g_{\rm YM}$ of the gauge theory that generates the stopping potential receives a correction $\delta g_{\rm YM}^{-2} \sim N$. This leads to an exponential suppression of the stopping potential, with barrier heights $\sim e^{-N}$, and therefore to a runaway relaxion. This problem persists even in the limit in which the relaxion shift symmetry appears to be restored.

Although we work in string theory, and quantum gravity completion is the central question, our results do not hinge on super-Planckian displacements $\Delta \phi \gg M_{\rm pl}$, which are famously challenging in quantum gravity. The problems that we expose occur even for $\Delta \phi \ll M_{\rm pl}$. The core issue is indeed one of large displacements, but here large means compared to the natural scale (or periodicity) of the effective theory. When ϕ is an axion with decay constant f, the backreaction of monodromy charge is significant for $\Delta \phi \gg f$.

Our analysis does not amount to a complaint that the effective theories given in [1] contain small dimensionless parameters. Constructing a solution of string theory that yields an effective field theory containing small numbers plausibly requires fine-tuning, e.g. of the discrete data of a compactification. Quantifying this obvious issue is not our aim. The backreaction phenomenon that we identify is a much more severe problem: even granting fine-tuned data that gives rise to an apparently-suitable relaxion Lagrangian in the probe approximation that omits the monodromy charge as a source in ten dimensions, the full Lagrangian beyond the probe approximation is not of the form given in [1], and does not allow for relaxation of a hierarchy.

Our goal is to identify the challenges that confront the relaxion mechanism in string theory. Though we analyze a specific realization in type IIB string theory, we find a set of surprising, plausibly general, qualitative lessons about the nature of hierarchies and technical naturalness in low energy effective field theories descending from string theory.

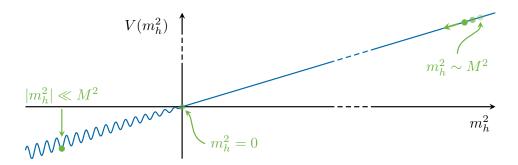


Figure 1. Schematic plot of the relaxion potential (1.1).

The remainder of section 1 is a microcosm of the paper. We begin with a review of the relaxion mechanism and then provide an overview of our results, leaving detailed analysis for the main text. The casual reader need only read section 1.

Overview of the relaxion. The simplest model of electroweak scale relaxation involves adding to the Standard Model a single axion ϕ , the relaxion, with the potential¹

$$V(\phi,h) = \left(M^2 - gM\left(\phi_{\text{init}} - \phi\right)\right)|h|^2 + gM^3\phi + V_{\text{stop}}(\phi,v).$$
(1.1)
(1.1)
(1.1)

Here h is the Higgs field and v is its vacuum expectation value, $v^2 \equiv \langle |h|^2 \rangle$, M is the cutoff of the effective field theory, and g is a dimensionless parameter that controls the explicit (albeit weak) complete breaking of the relaxion's perturbatively exact continuous shift symmetry $\phi \mapsto \phi + \text{const.}$ The coupling \triangle promotes the Higgs mass m_h^2 to a dynamical variable, so that evolution of ϕ scans over a range of Higgs masses, while \triangle is a potential that forces ϕ to smaller values, $\phi_{\text{final}} \ll \phi_{\text{init}}$. Finally, \bigcirc is a non-perturbatively generated, oscillatory "stopping potential" $V_{\text{stop}}(\phi, v) = V_{\text{stop}}(\phi + f, v)$, whose height grows with the Higgs vev v. For now, we take this potential to be

$$V_{\rm stop}(\phi, v) = \Lambda_c^3 v \cos\left(\frac{2\pi\phi}{f}\right) \tag{1.2}$$

with Λ_c the confinement scale of a gauge theory G to which ϕ has an axionic coupling, though we will consider more general potentials in section 2. This potential is generated by strong gauge dynamics and disappears when the theory is in a phase with unbroken chiral symmetry, i.e. in a phase with massless quarks. Thus, the stopping potential vanishes unless the Higgs has developed a vev.

The mechanism is illustrated in figure 1. The relaxion starts at a large value ϕ_{init} , where $m_{\text{h}}^2 \sim M^2$, and begins to slowly roll down the linear potential B. For generic initial conditions, the relaxion will roll a distance

$$\Delta \phi \sim M/g \tag{1.3}$$

¹We follow the same notation as [1], except that we take the coupling g to be dimensionless, $g_{\text{GKR}} = g \times M$, and shift the origin of the relaxion ϕ field space.

in field space before the Higgs becomes massless, $(A) = m_h^2 = 0$. The Higgs then develops a vev and the stopping potential is generated. The relaxion continues to roll, halting once the stopping potential grows strong enough to counterbalance the linear potential roughly when

$$\frac{v}{M} \sim g \left(\frac{M}{\Lambda_c}\right)^3 \frac{f}{M}.$$
 (1.4)

The hierarchy between the Higgs vev and the cutoff of the theory is thus controlled by the shift-symmetry breaking parameter g. In effective field theory, it is technically natural for g to be arbitrarily small. However, we will see that there are obstacles to such a structure in string theory.

Requirements for relaxation. We now summarize the necessary ingredients for a successful relaxation of the electroweak scale.

1. The Higgs mass must be made dynamical by introducing an axion² field ϕ with a coupling to the Higgs of the form

$$\mathcal{L}_h \supset \mathcal{G}(\phi) |h|^2 \tag{1.5}$$

where $\mathcal{G}(\phi)$ is some polynomial in ϕ . Evolution in ϕ scans over Higgs masses.

- 2. The dynamics of ϕ must be attractive, with the late-time (when $m_h^2 \sim 0$) behavior of ϕ being independent of the initial conditions.
- 3. ϕ must stop when the Higgs mass is approximately its observed, unnatural value.

For the evolution of the relaxion to be both attractive and dominated by classical dynamics, some friction is necessary. Therefore, the relaxion scenario has been assumed to take place during inflation (for an alternative source of friction from particle production see [2]). In this paper we will not discuss the underlying model of inflation (e.g. see [3-5]), nor its possible realization in string theory; these issues are the subject of an extensive literature (see for example [6]). We assume inflation to be operative, and concentrate instead on the relaxion potential and examine how it may arise in string theory constructions.

Typically, the stopping potential is generated by non-perturbative effects and is f-periodic. This ensures that only (A) and (B) explicitly break the discrete shift symmetry $\phi \mapsto \phi + f$, and protects against possibly disastrous corrections. The height of the stopping potential must depend on the Higgs vev.³ Furthermore, we require the minima of (1.1) to scan through Higgs masses finely enough so that a small overshoot does not dramatically increase the final electroweak scale; since the stopping potential minima are spaced roughly $\Delta \phi \sim f$ apart, this translates into the requirement that $\mathcal{G}'(\phi)f \ll v^2$.

²As is clear from the name, it is important that the relaxion ϕ be an axion: the axionic shift symmetry protects the potential against undesirable corrections. One could envision a more general relaxation scenario involving a field ϕ that is not an axion, but it would then be necessary to explain how the structures in (1.1) could be technically natural.

³This is not required when the stopping mechanism is realized by particle production, see e.g. [2].

While this appears to be a beautiful solution to the Higgs hierarchy problem, there is some cause for concern: g must be an exceptionally small number in order to generate a sizable hierarchy. The simplest model of [1] requires $g \sim 10^{-28}$; see section 2 for the requirements in variants of the model. It is reasonable to ask whether the associated large number 1/q infects any other terms in the effective action.

Note that although we used the same g in (A) and (B), these two terms could in principle be different. Let us temporarily distinguish them and denote the coupling in (A) as g_h . If $g \ll g_h$ at tree level, Higgs loops will drive the coupling in (B) to be of order g_h so that the two couplings in (A) and (B) are not very different. If, on the other hand, we take $g \gg g_h$ at tree level, this hierarchy is stable but the required field excursion in (1.3) increases to $M/g_h \gg M/g$. So, models with $g_h \sim g$ undergo the smallest field excursion, and for this reason we only consider one g coupling in eq. (1.1).

Although all the phenomena that we will uncover in this work can be encoded in an effective field theory, appropriately extended to include the effects of states that enter the spectrum as the relaxion makes its long excursion, these effects are not easily seen without the perspective of an ultraviolet theory. This is to say that the technical naturalness reasoning of [1] amounts to a set of premises about the field content and interactions of an effective theory, together with conclusions that indeed follow from those premises. In this work we question these premises, asking whether string theory imposes restrictions or refinements on the possible effective theories. We first critically examine technical naturalness arguments in this context and then turn to a string theory embedding of the relaxion.

Technical naturalness and large displacements. Technical naturalness is often used as a panacea in model building: one begins with a symmetry that protects against potentially disastrous quantum corrections and then weakly breaks it, confident that all corrections induced by this breaking are necessarily small. If g is a dimensionless parameter measuring the weak symmetry breaking, and the symmetry is restored for $g \to 0$, corrections in the effective theory are proportional to positive powers of g, and so are well-controlled for $g \ll 1$. However, this logic must be used with care in the presence of field excursions $\Delta \phi$ that are large compared to the effective theory's cutoff M. We now explain why technical naturalness arguments can fail in the context of effective field theories and how these failures can be exacerbated in quantum gravity — the essential problem being that corrections can depend both on g and on a potentially large dimensionless parameter, $\Delta \phi/M$.

As a toy example, consider a four-dimensional effective theory for a scalar field ϕ with Lagrangian

$$\mathcal{L} = -\frac{1}{2} (\partial \phi)^2 - M^4 \sum_{i=1}^{\infty} c_i \left(\frac{\phi}{M}\right)^{d_i} g^{e_i} , \qquad (1.6)$$

where M is a physical ultraviolet cutoff (the scale of some new physics), $g \ll 1$ is a dimensionless parameter, the c_i are dimensionless Wilson coefficients, and the d_i and e_i are non-negative numbers. As long as

$$e_i \neq 0 \qquad \forall i,$$
 (1.7)

all quantum corrections are proportional to powers of g, and the continuous shift symmetry $\phi \mapsto \phi + \text{const.}$ is restored in the limit $g \to 0$. However, we stress that (1.7) must be checked for every term in (1.6), as any $e_i = 0$ term, no matter how irrelevant, could potentially provide disastrous corrections.

Furthermore, at large displacements $\phi \gg M$, the condition (1.7) is far from sufficient to ensure that quantum corrections are under control at small but finite g. The theory contains a new large parameter, ϕ/M , and corrections proportional to $g^{e_i}(\phi/M)^{d_i}$ are not necessarily small for $g \ll 1$ and $\phi/M \gg 1$. Ensuring that the corrections to the classical equations of motion are small requires knowledge about the *entire sequence* $\{c_i, d_i, e_i\}$, and so the full Lagrangian (1.6).

In systems allowing axion monodromy, there is an additional subtlety: the limit $g \to 0$ is not smooth,⁴ because the field space *discontinuously* changes from a helix (for $g \neq 0$) to a circle (for g = 0). Standard technical naturalness arguments that rely on the $g \to 0$ limit can therefore become problematic.

Now suppose one obtains an effective theory from the top down, beginning in a vacuum of quantum gravity and integrating out Planck-scale degrees of freedom, for example by performing dimensional reduction in a string compactification with stabilized moduli. Then the low-energy theory in four dimensions could still take the form (1.6), but with two important caveats. First, the exponents d_i , e_i are dictated by the vacuum configuration of the underlying theory, and the condition (1.7) must be established rather than assumed. Second, in configurations with ≤ 4 supercharges in four dimensions, in practice one never obtains complete information about the infinite sum in (1.6): some terms can be computed in different approximations, but other terms remain incalculable, although they are in principle determined by the underlying vacuum.

Because we do not have the ability to compute every term in (1.6) in *any* halfwayrealistic solution of string theory, it is difficult to prove that (1.7) is possible in quantum gravity. As a result, there is a disjunction between bottom-up reasoning based on technical naturalness, and top-down reasoning based on obtaining effective theories from quantum gravity: the former strictly requires the condition (1.7), which appears not to be provable in quantum gravity.

In our view, the difficulty in establishing (1.7) in any particular solution of string theory is not just that the computation is challenging; it is that plausible general reasoning about black hole thermodynamics in quantum gravity suggests that (1.7) is in fact false. Exact continuous global internal symmetries are thought by many to be impossible in quantum gravity and have not appeared in string theory to date. We therefore expect quantum gravity to dramatically affect the $g \to 0$ limit. Although our results will turn out to be compatible with this general expectation, we do *not* rely on bottom-up reasoning about quantum gravity at any point in our analysis. In particular, we do not assume any form of the Weak Gravity Conjecture (WGC).⁵ Instead, as described above, we will directly analyze this limit in the context of NS5-brane axion monodromy.

⁴This observation led the authors of [7] to argue that the $g \to 0$ limit is not technically natural.

⁵For work applying the WGC to the relaxion, see e.g. [8, 9].

We will argue that axion monodromy in string theory is very generally characterized by the existence of one or more terms in the effective action (1.6) with $e_i = 0$, and the theory is poorly-controlled in the limit $g \to 0$, $\phi/M \to \infty$. The physical origin of these problematic terms is *backreaction by monodromy charge*, as we now explain.

New states from monodromy. In a viable relaxion theory, we must find that every shift symmetry breaking term in the relaxion Lagrangian is proportional to a power of g, the parameter that controls the weak breaking in (1.1). However, as we will explain qualitatively now and quantitatively in section 4, the monodromy charge

$$N \equiv \frac{\Delta \phi}{f} \,, \tag{1.8}$$

leads to corrections that are *not* dressed by powers of g, so that (1.7) does not hold. We will begin with an example and then draw more general lessons.

Suppose (cf. the detailed discussion in section 3.3) that the relaxion is associated to a two-cycle wrapped by an NS5-brane. Further, suppose that the stopping potential arises from the dynamics of a strongly-coupled non-Abelian gauge theory, with group G, living on a stack of D7-branes wrapping a four-cycle Σ_4 . The height of the stopping potential depends on the coupling $g_{\rm YM}$ of this D7-brane gauge theory:

$$|V_{\rm stop}| \propto \Lambda_c^3 \propto \exp\left(-\frac{8\pi^2}{g_{\rm YM}^2 c_G}\right)$$
 (1.9)

Here the constant c_G is determined by the type of non-perturbative effects that generate V_{stop} , and may be set to unity for our purposes. The gauge coupling function of G is proportional to the warped volume of Σ_4 , cf. (4.7):

$$\frac{1}{g_{\rm YM}^2} = \frac{\rm vol_W(\Sigma_4)}{2\pi\ell_s^4} \,. \tag{1.10}$$

When the system is wound up over N cycles, N units of monodromy charge — which in this scenario is D3-brane charge — accumulate on the NS5-brane. This charge is a source in the ten-dimensional Einstein equations, and so leads to changes in the metric of the internal space and the warp factor. The backreaction thus alters the warped volume $vol_W(\Sigma_4)$. Then, through (1.10), the gauge coupling function — and hence the height of the barriers — depend on N. In section 4.2.4, we will show that

$$\delta\left(\frac{8\pi^2}{g_{\rm YM}^2}\right) \sim N,$$
 (1.11)

without dependence on g.

The correction (1.11) can be understood in a dual description as resulting from new light states associated to the source of monodromy. The one-loop $\overline{\text{MS}} \beta$ -function in a Yang-Mills theory with n_{F} fermions, n_{S} complex scalars, and coupling constant g_{YM} can be written

$$\frac{\mathrm{d}}{\mathrm{d}\log\mu}\left(\frac{8\pi^2}{g_{\mathrm{YM}}^2}\right) = \frac{11}{3}T(\mathrm{Ad}) - \frac{2}{3}\sum_{i=1}^{n_{\mathrm{F}}}T(\mathrm{R}_i) - \frac{1}{3}\sum_{a=1}^{n_{\mathrm{S}}}T(\mathrm{R}_a),\tag{1.12}$$

where $T(\mathbf{R}_i)$ is the index of representation \mathbf{R}_i and Ad denotes the adjoint representation. The introduction of N light states will typically lead to a change

$$\delta\left(\frac{8\pi^2}{g_{\rm YM}^2}\right) = \gamma_{\rm br} N \,, \tag{1.13}$$

with $\gamma_{\rm br}$ a constant independent of N.

Where do these new light states come from? The N units of D3-brane charge in the NS5-brane can be viewed as resulting from N actual D3-branes (up to a binding energy that does not affect our argument). So there are N new states in the theory, corresponding to strings stretching from the D7-brane stack, where the gauge theory lives, to the D3-branes. These states transform in the fundamental of G, and so may be described as N species of quarks from the viewpoint of G. Including these species in loops leads to (1.13).

The lesson is that $\mathcal{O}(N)$ new states associated with the source of monodromy — in our examples, fundamental strings stretching from the source of monodromy to the gauge theory D-branes — can give large loop corrections. These states could easily be missed in field theory, but in a string theory configuration with two D-brane gauge theories G_1, G_2 , the presence of bifundamentals is hard to avoid. The only question is whether the bifundamentals are so massive that they are physically unimportant. In the simplest setting, we will find (cf. appendix B) that arbitrarily short — and hence, light — bifundamental strings are present. In more involved constructions the length of these strings may be increased, but we argue that a significant fine-tuning of $\mathcal{O}(N^{-1})$ remains necessary to control their effects.

The fact that for each unit of monodromy charge there is a new state coming down in mass that contributes to the gauge coupling of the effective theory — even though this state was far above the cutoff in the vacuum at zero winding — is a consequence of the structure of the ultraviolet completion. The new states described above arise from stretched strings, and so obviously have their origin in string theory *per se*, but there are also new states that arise simply from the presence of extra dimensions: these are Kaluza-Klein (KK) states made light by monodromy. Thus, our considerations can be extended to extra-dimensional "partial" ultraviolet completions of four-dimensional field theories, without invoking string theory.

Perhaps the simplest illustration of this phenomenon is the model of [10] (see also [11]), which describes axion monodromy arising from a Stueckelberg massive U(1) gauge field coupled to a massless charged scalar field in a five-dimensional spacetime with the extra dimension compactified on a circle,

$$S_{5D} = \int d^4x \int_{S^1} dy \sqrt{-g} \left(-\frac{1}{4} F_{MN} F^{MN} - \frac{1}{2} m^2 \mathcal{A}_M \mathcal{A}^M - (D_M \Phi)^{\dagger} (D^M \Phi) \right), \quad (1.14)$$

where $D_m = \partial_M - iqA_M$, $F_{MN} = \partial_{[M}A_{N]} = \partial_{[M}A_{N]}$, and $A_M = A_M - ie^{i\theta}\partial_M e^{-i\theta}$ denotes the Stueckelberg covariant U(1) gauge field. Now we perform a KK reduction on the circle, whose circumference we denote by $2\pi R$. We decompose the five-dimensional fields into an infinite series of discrete Fourier (KK) modes on the circle, and focus on the KK modes of the scalar Φ ,

$$\Phi(x^{\mu}, y) = \frac{1}{\sqrt{2\pi R}} \sum_{n \in \mathbb{Z}} \Phi_n(x^{\mu}) \exp\left(\frac{iny}{R}\right)$$
(1.15)

This yields the effective four-dimensional action

$$S_{5\mathrm{D}} \supset \int \mathrm{d}^4 x \, \left(\frac{1}{2} m^2 \phi^2 + \sum_{n \in \mathbb{Z}} \left(\frac{n}{R} - q\phi \right)^2 |\Phi_n|^2 \right). \tag{1.16}$$

Here, $\phi \sim A_5^{(0)}$ denotes the four-dimensional axion field corresponding to the fivedimensional gauge field Wilson line around the S¹. The axion ϕ evidently experiences monodromy, acquiring a quadratic potential.

The key observation is that the masses of the KK modes Φ_n ,

$$m_n^2 = \left(\frac{n}{R} - q\phi\right)^2 \tag{1.17}$$

depend on the vev of the axion ϕ . As ϕ scans across its field space, one KK mode after another falls below the cutoff R^{-1} in mass and thus enters the spectrum of the low-energy effective theory. In particular, as ϕ moves over N units of its fundamental domain, N KK modes fall below the cutoff R^{-1} , in analogy with the string theory effect discussed above.

We should clarify that in our examples, monodromy affects mass spectra in two very different ways. One effect is *shifting*, in which $\phi \mapsto \phi + f$ leaves the set of masses m in a sector invariant, but permutes the states associated with these masses. For example, in (1.17), changing $\phi \mapsto \phi + (qR)^{-1}$ increases by one unit the Kaluza-Klein charge of the state at each mass level.

The other effect is *compression*, in which a monodromy $\phi \mapsto \phi + f$ changes the mass spectrum. Typically, as the axion winds up and stores more energy, the masses in affected sectors are reduced. A shifting spectrum is compatible with an exact discrete shift symmetry of the theory; the number of states below a fixed cutoff does not change, but the labels of the states change. Compression violates even a discrete shift symmetry, as the number of states below a fixed cutoff depends on ϕ .

With this terminology, we remark that the five-dimensional example above displays only shifting, not compression. This is a consequence of the oversimplified nature of the model. We will show below that axion monodromy also causes compression of the mass spectrum of Kaluza-Klein excitations of an NS5-brane. Thus, stretched string states are not the only states that experience compression, and we expect that compressed spectra can arise in purely extra-dimensional scenarios without string theory.

Why do we not provide a purely four-dimensional field theory toy model showing the effects of shifting and compression, for instance in the case of axion monodromy from a four-form field strength [11-14]? The issue is that although the core mechanism of axion monodromy arising via the Stueckelberg mechanism can be described in four-dimensional field theory, the results of [11] make it clear that backreaction effects, including those of massive states entering the spectrum, are described by higher-derivative corrections arising from higher powers of the four-form field strength. These corrections must be determined

in the ultraviolet completion of gravity, as explicitly noted in [11] as well. That is, the twoderivative, four-dimensional field theory Kaloper-Sorbo model of axion monodromy [12] is not a magic wand that suppresses or controls backreaction effects.

Exponential suppression of the stopping potential. We have argued that backreaction by N units of monodromy charge leads to a large correction to the gauge coupling (1.13). Thus,

$$|V_{\rm stop}| \propto \exp\left(-\gamma_{\rm br}N\right),$$
 (1.18)

where $\gamma_{\rm br}$ is a number that has no parametric dependence on N or on the shift symmetry breaking parameter g. When $\gamma_{\rm br}$ is positive, the immediate and fatal consequence is the exponential suppression of the stopping potential.⁶ The stopping potential, including the backreaction effect encoded in (1.18), is far too small to halt the evolution when the Higgs is almost massless, $|m_{\rm h}|^2 \ll M^2$. The result is a runaway relaxion. If instead $\gamma_{\rm br}$ is negative, the story is more involved. But, as we will see in section 4.2.4, the result is still exponential suppression of the stopping potential.

Now to make things worse, there are two independent requirements that necessitate placing the source of monodromy in a region with a large background D3-brane charge N_{D3} that obeys $N_{D3} \gg N$. This background charge introduces an additional, larger exponential suppression of the stopping potential.

First, achieving a large hierarchy between the Higgs vev v and the cutoff M necessitates an extremely small g. This parameter controls how strongly the relaxion shift symmetry is broken and is determined by the amount of energy introduced into the configuration per unit of monodromy charge. Since the source of monodromy corresponds to a physical quantized object, the amount of energy introduced by an additional winding is, in a sense, irreducible. However, warping the source of monodromy reduces this quantum of energy compared to other scales in the problem. So, an extremely small g — and thus a large hierarchy — may be realized by placing the source of monodromy in a strongly warped region, as in figure 5 on page 25. As we will show in section 3.2, we may characterize this warping by the amount of effective D3-brane charge N_{D3} needed to create the warped throat, and

$$g \propto N_{\rm D3}^{-1}$$
. (1.19)

Moreover, we must require a large background D3-brane charge to retain computational control and to ensure stability of the ten-dimensional configuration. The D3-brane charge induced as the relaxion is wound must be a small correction to the charge of the ambient space for the background not to overwhelm the background configuration, and we therefore must require $N_{\text{D3}} \gg N$.

Now as in the preceding section, the D3-brane charge N_{D3} of the ambient space has an effect akin to that of N_{D3} actual D3-branes, which would give rise to N_{D3} species of

⁶Can one fine-tune $\gamma_{\rm br}$ to avoid the suppression in (1.18)? In section 4.2.3 we explain what such a tuning would correspond to in terms of compactification parameters, but it is already clear that this cannot be a satisfactory solution. Independent of where and how the stopping potential is generated, one would need to ensure that $\gamma_{\rm br} \leq \mathcal{O}(N^{-1})$, which by (1.3) reintroduces a tuning on the order of the hierarchy the mechanism was supposed to naturally explain.

quarks in the fundamental of G. Loops of these quarks yield

$$\delta\left(\frac{8\pi^2}{g_{\rm YM}^2}\right) = \gamma_{\rm bg} N_{\rm D3} \,, \tag{1.20}$$

where $\gamma_{\rm bg}$ is a positive constant.

We conclude that the stopping potential is *exponentially suppressed* by the warping required to achieve a weak shift symmetry breaking $g \ll 1$ and to maintain control over the model. Schematically, we have

$$|V_{\rm stop}| \propto \exp\left(-\gamma_{\rm bg} N_{\rm D3}\right) \,, \tag{1.21}$$

with $N_{\text{D3}} \gg N$. Here we remind the reader that $N \gg 1$ is the large number of windings required to substantially ameliorate the hierarchy problem. The suppression (1.21) renders the barriers utterly negligible.⁷ Using (1.19), (1.21) can be written as

$$|V_{\text{stop}}| \propto \exp\left(-\mathcal{O}\left(1/g\right)\right),$$
 (1.22)

so we see that the generated hierarchy is no longer proportional to g. Instead, suppressing the shift symmetry breaking scale simultaneously suppresses the stopping potential barriers, leading to a relaxion runaway.

Overview of the paper. We seek to understand to what extent the relaxion, as a mechanism to naturally generate large hierarchies, survives ultraviolet completion. We find that backreaction of monodromy charge — induced D3-brane charge in the NS5-brane model of axion monodromy we study in detail — dramatically alters the structure of the relaxion action. These corrections may be interpreted as the contribution of new light states which appear in the spectrum as one traverses the axion's fundamental domain, and are expected to be ubiquitous in models of axion monodromy. We thus find that a successful embedding of the relaxion mechanism requires a fine-tuning on the order of the hierarchy the mechanism sets out to generate, rendering the mechanism inoperable. Because we are interested in generic expectations and given (cf. table 1) the extreme dimensionless quantities involved, we work at the level of parametric estimates.

The organization of this paper is as follows. In section 2 we briefly survey relaxion models constructed in effective field theory, and identify the parameter ranges that allow relaxation of a hierarchy. In section 3 we introduce axion monodromy in string theory, emphasizing the fact that monodromy results from a physical, quantized source. We review the scenario of axion monodromy on NS5-branes, and then explain how the relaxion mechanism could be realized in this setting. In section 4 we determine the microphysical constraints that arise in such a realization. An executive summary appears in section 4.1. We discuss generalizations in section 5, and conclude in section 6. The appendices contain more technical material. Appendix A provides background on axions in string theory. In

⁷This effect holds regardless of the sign of $\gamma_{\rm br}$ in (1.18). But, for $\gamma_{\rm br} < 0$, (1.21) is the only relevant exponential suppression, while for $\gamma_{\rm br} > 0$ (1.18) amounts to an independent suppression $\sim e^{-N}$ which even on its own is sufficient to cause a runaway.

appendix B we prove that the D7-branes responsible for the stopping potential must intersect a holomorphic representative of the homology class $[\Sigma_2]$ wrapped by the NS5-brane. In appendix C we give the actions for D5-branes and NS5-branes in warped compactifications of type IIB string theory. In appendix D we analyze the backreaction of D3-brane and anti-D3-brane charge and tension on the metric of the internal space.

2 Relaxion zoology

We now briefly overview a selection of existing relaxion models in field theory. To present a unified synopsis of the genus *Relaxion* in its various speciations, we discuss these models in a consistent notation and, since the number of windings N is severely constrained in string theory, we pay special attention to the field excursions required to generate a large hierarchy.

For generic initial displacements, the relaxion must scan a field range $\Delta \phi \sim M/g$ to reach $m_h^2 = 0$. If we generalize (1.2) and (1.4) by including a more generic dependence on the Higgs vev v as in [15], schematically

$$V_{\text{stop}}(v,\phi) = \Lambda^4(v) \cos\left(\frac{2\pi\phi}{f}\right) = \epsilon \Lambda_c^4 \left(\frac{v}{\Lambda_c}\right)^r \cos\left(\frac{2\pi\phi}{f}\right)$$
(2.1)

and

$$\left(\frac{v}{M}\right)^r \sim \frac{g}{\epsilon} \frac{f}{M} \left(\frac{M}{\Lambda_c}\right)^{4-r} \tag{2.2}$$

with ϵ a constant coefficient, this excursion implies a winding charge of

$$N = \frac{\Delta\phi}{f} \sim \frac{1}{\epsilon} \left(\frac{M}{\Lambda_c}\right)^4 \left(\frac{\Lambda_c}{v}\right)^r.$$
 (2.3)

For $\Lambda_c \sim v$, N scales with the fourth power of the ratio of the cutoff scale M to the weak scale, further increasing if $\Lambda_c \ll v$.

2.1 Original models

Two explicit constructions were originally proposed in [1], and the dynamics of these models were explained in section 1. The relaxion in the first model (GKR1) is the QCD axion and the potential barriers are generated by strong chromodynamic forces. The potential barriers in (2.1) then scale as $\Lambda^4(v) \sim \Lambda^3_{\text{QCD}} m_u$, i.e. r = 1, $\Lambda_c \sim \Lambda_{\text{QCD}}$, and ϵ is the up-quark Yukawa coupling y_u . The main drawback of this model is that it destroys the solution to the strong CP problem. The PQ solution may be restored, as discussed in [1], by introducing additional dynamics at the end of inflation, which removes the slope of the relaxion potential at the end of inflation. However, in this case the cutoff scale cannot be pushed higher than $M \approx 30$ TeV. The hierarchy (2.3) is then multiplied by a factor of the QCD angle θ_{QCD} . In either case, the number of windings obeys $N \geq (M/\Lambda_{\text{QCD}})^4$.

Because the generated hierarchy (2.3) grows with the confinement scale Λ_c , the second model (GKR2) introduces a new strongly-interacting gauge sector G, whose axion is the relaxion. The PQ solution to the strong CP problem is then untouched and Λ_c can be

much larger than Λ_{QCD} . However, this does not allow one to make the barrier arbitrarily high. New electroweak scale fermions couple the Higgs sector to this new sector and the barrier height depends quadratically (r = 2) on the Higgs vev. But, a constant term (r = 0) will also be generated by quantum corrections so the barriers can schematically be written as [15]

$$\Lambda^4(v) \sim \epsilon \Lambda_c^4 \left(1 + \left(\frac{v}{\Lambda_c} \right)^2 \right).$$
(2.4)

So, the barrier will not depend strongly enough on the Higgs vev v for the relaxion mechanism to work unless $\Lambda_c \leq v$. Still, GKR2 can generate a much larger hierarchy $M \leq 10^8$ GeV than GKR1, with a similar parametric scaling of the number of windings $N \geq (M/v)^4$. Unfortunately, GKR2 requires that new electroweak scale fermions be put in by hand, and this coincidence of scales must be explained.

2.2 CHAIN

A solution to this coincidence problem was suggested in [15], and involves taking the barrier height to depend on an extra scalar field. The relaxion mechanism is then able to explain the near-criticality of the Higgs without a coincidence of scales. Instead, there is only one scale in the problem, the cutoff M, which also sets the barrier height $\Lambda_c \sim M$. The extra scalar σ , which need not be an axion, controls the height of the stopping potential,

$$\Lambda^4(h,\phi,\sigma) = \epsilon M^4 \left(\beta + c_\phi \frac{g\phi}{M} - c_\sigma \frac{g_\sigma \sigma}{M} + \frac{h^2}{M^2}\right).$$
(2.5)

The initial conditions are very different from both GKR1 and GKR2. At first, the barriers are large and the relaxion is stuck in one of its minima. As the second field σ evolves, its vev will eventually cancel this barrier and allow the relaxion to roll. In contrast with GKR2, there are no constraints on the decay constant f from reheating.

Given that now the barriers are allowed to be high, $\Lambda_c \gg v$, one might hope that the required number of windings for a given cutoff scale is substantially reduced. A more careful analysis, however, reveals that this is not the case. Instead, because classical evolution must dominate over quantum fluctuations, we require that $\epsilon \leq v^2/M^2$, while imposing that the Higgs barrier in (2.5) is solely responsible for stopping ϕ requires that $v^2 \sim gMf/\epsilon$. Together, these imply that

$$N \sim \frac{M}{gf} \gtrsim \left(\frac{M}{v}\right)^4,\tag{2.6}$$

and so the CHAIN model also requires a large number of windings to resolve the hierarchy problem.

A comparison of the three models is shown in figure 2 and table 1, while further phenomenological and cosmological constraints are discussed in [16].

2.3 Supersymmetric models

Inflation limits the achievable cutoff scale to $M \sim 10^9 \,\text{GeV}$. The energy stored in the relaxion must not dominate over the energy driving inflation,

$$M^4 < H^2 M_{\rm pl}^2$$
, (2.7)

	GKR 1	GKR 2	CHAIN with $f \sim M$
f	$f_{PQ} \sim 10^{10} - 10^{12} \text{ GeV}$	$\gtrsim M_{\rm GUT} \sim 10^{16} { m GeV}$	$\gtrsim M$
g	$\left(\frac{\Lambda_{\rm QCD}}{M}\right)^4 \frac{M}{f_{\rm PQ}} \theta_{\rm QCD} \lesssim 10^{-36}$	$\left(\frac{\Lambda_{\rm EW}}{M}\right)^4 \frac{M}{M_{\rm GUT}} \sim 10^{-30} - 10^{-20}$	$\lesssim v^4/M^4 \sim 10^{-26} - 10^{-6}$
M _{max}	$3\times 10^3~{\rm GeV}$	$10^8 { m ~GeV}$	$10^9 { m ~GeV}$
m_{ϕ}	$\frac{\Lambda_{\rm QCD}^2}{f_{\rm PQ}} \lesssim 10^{-11}~{\rm GeV}$	$\frac{\Lambda_{\rm EW}^2}{M_{ m GUT}} \lesssim 10^{-12} { m ~GeV}$	$\sqrt{gM^4/v^2} \lesssim v$
N	$\theta_{\rm QCD}^{-1} \left(\frac{M}{\Lambda_{\rm QCD}}\right)^4 \gtrsim 10^{30}$	$(M/\Lambda_{\rm EW})^4 \sim 10^8 - 10^{24}$	$g^{-1} \sim 10^6 - 10^{26}$

Table 1. Summary of parameter values in the three non-supersymmetric relaxion models discussed in section 2.

where H is the Hubble rate during inflation, and barriers cannot form unless $H < \Lambda_c$. This immediately implies a bound on the cutoff $M \lesssim \sqrt{vM_{\rm pl}} \sim 10^9$ GeV for GKR2. While this argument does not directly apply to the CHAIN model, there one finds the same bound $M \lesssim 10^9$ GeV. Since we must also explain the remaining hierarchy between the cutoff Mand the Planck scale $M_{\rm pl}$, a natural candidate solution is that supersymmetry is restored above M and the relaxion is embedded within a supersymmetric model.

A supersymmetric version of GKR1 was presented in [17], on which the following discussion is based. The relaxion becomes part of a chiral superfield S:

$$S = \frac{s+ia}{\sqrt{2}} + \sqrt{2}\theta\tilde{a} + \theta^2 F + \dots, \qquad (2.8)$$

which contains the (dimensionless) relaxion $a = \phi/f$, a srelaxion field s, and the relaxion \tilde{a} . The Peccei-Quinn symmetry acts as $S \mapsto S + i\alpha$. The linear term B in the relaxion potential (1.1) descends from the superpotential term

$$W \supset \frac{1}{2}mf^2S^2 \,. \tag{2.9}$$

Small $m \ll f$ is technically natural since m breaks the PQ symmetry, which is non-linearly realized via the term

$$W \supset \mu_0 \, e^{-qS} H_u H_d. \tag{2.10}$$

Apart from S, the model contains only SM particles and their superpartners (including the usual second Higgs doublet). The effective potential for s and a is then

$$V = \frac{1}{2}m^2 f^2 \left(s^2 + a^2\right) \kappa(s), \qquad (2.11)$$

with $\kappa(s)$ a function of s. As in all relaxion models, a starts out at a field value far away from its minimum at a = 0, and so breaks supersymmetry, with $F \propto ma$. As a evolves towards its minimum, it scans the SUSY breaking scale, and therefore the soft masses of the gauginos and the scalar superpartners. In particular, the determinant of the Higgs mass matrix was shown to scale as a^4 for $a \gg \mu_0/m$, far away from any electroweak symmetry breaking minima. As a approaches the critical value $a_* = \mu_0/m$, electroweak symmetry is broken, the Higgs(es) obtain a vacuum expectation value, and barriers appear that halt the evolution of a.

For a suitable choice of parameters, the model explains the hierarchy between the electroweak symmetry breaking scale v and the mass scale of the superpartners $\mu_0 \gg v$, thus solving the supersymmetric little hierarchy problem. According to [17], the number of windings scales as

$$N \sim \Delta a \sim \frac{f^2 \mu_0^2}{\Lambda_{\rm QCD}^4} \,, \tag{2.12}$$

where $f \sim 10^9 - 10^{12}$ GeV is the QCD axion decay constant and μ_0 plays the role of the UV cutoff. For $\mu_0 = 10^5$ GeV, a field excursion of $\Delta a \sim 10^{30}$ is required. Without further modifications, this model is phenomenologically unacceptable since it predicts $\theta_{\rm QCD} \sim \mathcal{O}(1)$. A variation with a non-QCD axion similar to GKR2 is briefly discussed in [17]. In this case, a larger range of decay constants is allowed — for $f = \mu_0$ and $\Lambda \sim v$ one obtains the same scaling as in GKR2, namely $N \sim \mu_0^4/v^4$.

A supersymmetrization of the CHAIN model was proposed in [18]. The philosophy is similar to the above discussion — now both the relaxion and the additional scalar σ are promoted to chiral superfields. The barriers for the relaxion are generated by a new SU(N_g) gauge theory, with confinement scale Λ_g , which communicates with the Higgs sector via a set of vector-like leptons. The required field excursion in this model is

$$N \equiv \frac{\Delta\phi}{f} \gtrsim \frac{m_{\rm SUSY}}{|m_{\rm S}|} \,, \tag{2.13}$$

where $m_{\rm SUSY} \sim \mu_0$ is the supersymmetry-breaking scale, and $m_{\rm S}$ is the relaxion mass coming from a term similar to (2.9). For $\Lambda_g \sim f \sim m_{\rm SUSY}$, a supersymmetry-breaking scale $\sim 10^9$ GeV may be generated through a field excursion of $N \sim 10^{27}$. So while the field excursion seems to grow more moderate as a function of the cutoff, it is still as large as in the earlier models for the largest possible cutoff (e.g. M^4/v^4 is of order 10^{26}). Instead, even the best case scenarios with $m_{\rm SUSY} \sim 10^4$ GeV require a large number of windings $N \gtrsim 10^8$.

2.4 Summary

The models presented above do not represent a complete classification of genus *Relaxion*. In particular, we are not considering models that rely on the alignment of multiple axions [19, 20] or use friction from particle production [2] to halt the evolution of ϕ . We discuss both of these further in section 5. However, the models that we examine represent a large cross-section of *Relaxion* and share a common trait: the required field excursion scales parametrically with the hierarchy generated, and so the associated number of windings around the relaxion field space $N \equiv \Delta \phi/f$ is enormous. In what follows, we will argue that N is a *physical charge* in string theory, which backreacts on the ten-dimensional configuration and tragically destroys the structures in (1.1), allowing for a runaway relaxion.

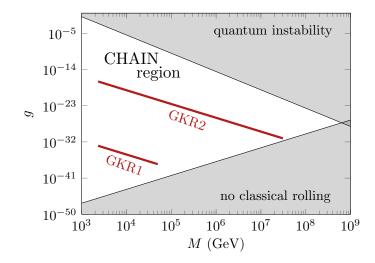


Figure 2. Schematic parameter space in the three main non-supersymmetric relaxion models. See [15] for the derivation of the constraints on the parameter space.

Relaxion Quantity	String Theory Origin
Axion ϕ	NS-NS or R-R $p\mbox{-}{\rm form}$ gauge field, dimensionally
	reduced along non-trivial p -cycle, §3.2
Discrete shift symmetry	Ten-dimensional NS-NS or R-R gauge symmetry,
$\phi \mapsto \phi + f$	exact in absence of brane or flux, $\S3.2$
Physical source of monodromy	Wrapped brane or flux along axion p -cycle, §3.2
explicitly breaks $\phi \mapsto \phi + f$	
Shift symmetry-breaking scale	Warped brane tension, $\S3.2$
gM^3f	
Winding number $N \equiv \Delta \phi / f$	Quantized monodromy charge, $\S3.2$
Axion decay constant f	Set by internal six-dimensional geometry, $\S A$
Stopping potential barrier	Set by warped volume of a four-cycle, $\S3.3$
height $\Lambda(v)$	

Table 2. A quick string theory-relaxion dictionary. This is a summary table, with more extended explanations given throughout the paper and in the appendices.

3 Relaxion monodromy

3.1 Axion monodromy in string theory

Axions are commonplace in string compactifications,⁸ and arise when a *p*-form gauge potential — either the NS-NS two-form B_2 or an R-R *p*-form C_p — is dimensionally reduced along a non-trivial cycle Σ_p in the compactification manifold X_6 . The ten-dimensional supergravity action is invariant under the gauge symmetry $B_2 \mapsto B_2 + d\Lambda_1$ and $C_p \mapsto C_p + d\Lambda_{p-1}$ which, upon reduction to four dimensions, ensures that the axion enjoys a perturbatively exact shift symmetry. For an axion *a* associated with a non-trivial cycle Σ_p , the shift symmetry $a \mapsto a + \text{const.}$ may be broken to a discrete shift symmetry by non-perturbative

 $^{^{8}}$ A detailed treatment of the material that follows can be found in [6], section 5.4.2. An overview is given in appendix A, and table 2 gives a simple dictionary.

effects, or completely broken by a brane wrapping Σ_p . In the latter case, the explicit breaking is proportional to the brane's tension. For example, if one wraps an NS5-brane along a two-cycle Σ_2 in the compactification manifold, the axion field c, defined by

$$c \equiv \frac{1}{\ell_s^2} \int_{\Sigma_2} C_2 \,, \tag{3.1}$$

experiences monodromy. The four-dimensional action for the dimensionless field c takes the form [21]

$$\mathcal{L} = -\frac{1}{2}f^2(\partial c)^2 - \varepsilon\mu_0^3 fc, \qquad (3.2)$$

where f is the axion decay constant,⁹ μ_0 is a parameter with dimensions of mass, determined by the geometry of X_6 , and ε parameterizes the warp factor at the location of the NS5-brane. (In terms of the warped line element (3.13), we have $\varepsilon = e^{4A_{\cup}}$.) We will explain the potential (3.2) in more detail in section 3.2.

Defining the canonically normalized axion $\phi \equiv fc$, we have

$$\mathcal{L} = -\frac{1}{2} (\partial \phi)^2 - \varepsilon \mu_0^3 \phi \,. \tag{3.3}$$

Comparing to the relaxion potential (1.1), we have the correspondence

$$gM^3 = \varepsilon \mu_0^3 \,. \tag{3.4}$$

So $\varepsilon \ll 1$ corresponds to $g \ll 1$ in the relaxion model. Since the breaking of the shift symmetry $\phi \mapsto \phi + \text{const.}$ is proportional to the warp factor at the location of the fivebrane, strong warping could lead to weak breaking of the symmetry, and hence to the small values of g required for a relaxion model.

The potential (3.3) has the desirable property that the entire potential is proportional to the warp factor, so it appears completely natural to make this potential small. However, a central observation of this paper is that achieving small g through warping, without unintended consequences elsewhere in the action, is challenging.

Requirements for axion monodromy. Let us first summarize the core ingredients mentioned above. For a model of axion monodromy in string theory, one requires:

- 1. An axion field descending from a *p*-form, and a *source of monodromy*: a brane, flux, or other physical ingredient that causes the configuration space to be a multi-cover of the axion circle, rather than just a single circle.
- 2. To have a plausible mechanism for making the breaking of the shift symmetry *weak*, the source of monodromy should be in a *warped region*.
- 3. Most of the issues that arise as possible obstacles become visible only in vacua with stabilized moduli: if one ignores the moduli sector, many problems disappear. But, of course, moduli stabilization is needed for a cosmological model. So the axion and the source of monodromy must be situated in a *vacuum with stabilized moduli*.

⁹The axion decay constant depends on the topology and geometry of the six-dimensional compact manifold X_6 : see appendix A.

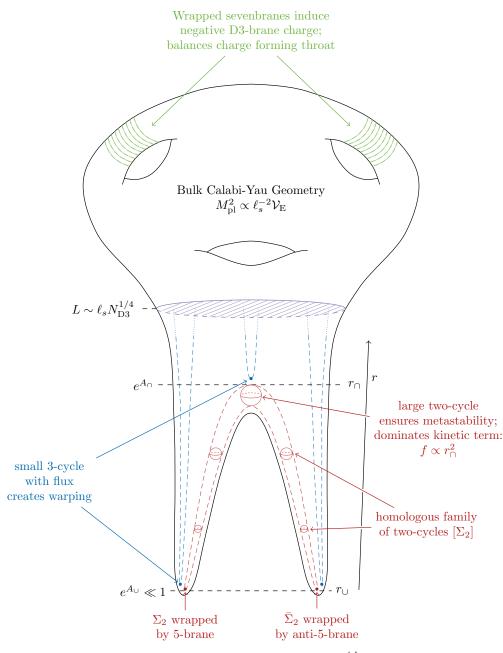




Figure 3. Minimal bifurcated warped throat setup for relaxion monodromy with 5-branes.

4. Since the compactification must have finite volume in order to lead to a finite fourdimensional Newton constant, Gauss's law imposes strict constraints on the charges in the compact space X_6 , and so we must *satisfy all tadpole conditions*.

There are many mechanisms in the literature that achieve (1), for instance [22]. But there is only one model currently available that achieves (1)-(3) [21, 23]: this is a model with an NS5-brane/anti-NS5-brane pair in a warped throat region of a type IIB flux compactification whose complex structure moduli are stabilized by fluxes, and whose Kähler moduli are stabilized by nonperturbative effects and possibly also by perturbative effects. We will call this model, whose detailed properties we will review in section 3.2, the NS5-brane model.

The central physics of the NS5-brane model is that transporting the dimensionless axion over a period induces one unit of D3-brane charge on the NS5-brane, and one unit of anti-D3-brane charge on the anti-NS5-brane. That is, "winding up" the axion by one cycle develops a D3-brane dipole in the compact space; the axis of the dipole is the line from the NS5-brane to the anti-NS5-brane. The entire dipole is in the infrared region of the warped throat where the fivebrane pair lives. See figure 3.

The key point is that the Lagrangian (3.3) arising from the NS5-brane DBI action, which is intended to be the relaxion Lagrangian, holds in the so-called *probe approximation*. That is, the potential in (3.3) follows from including the tension of the D3-branes and anti-D3-branes as a contribution to the four-dimensional vacuum energy, i.e. as a source in the four-dimensional Einstein equations, but *not* including this tension as a source in the ten-dimensional Einstein equations. The effects of a particular source on the tendimensional field configuration are termed the *backreaction* of that source, and so the probe approximation consists of neglecting the backreaction of D3-branes and anti-D3-branes.¹⁰

An immediate question is whether applying the probe approximation is consistent; in other words, can the backreaction of D3-branes be neglected? In the context of axion monodromy inflation in string theory, this question has been addressed, with the outcome that backreaction can be suppressed to some degree, by a variety of mechanisms, but nevertheless remains as a leading constraint on model-building [23]. However, the sources of backreaction are the D3-brane charge and tension, both proportional to the number of windings N of the axion. In the present context of relaxion monodromy, N needs to be extremely large, and so the problem of backreaction is much more severe than in the corresponding inflationary models. The constraints examined in [23] must therefore be revisited under this more severe test.

In this work, we will carefully examine the consequences of D3-brane backreaction for the NS5-brane model of relaxion monodromy in string theory. The first step is to explain how to compute the backreaction in this scenario.

3.2 Fivebrane axion monodromy

Our analysis will rely on detailed properties of the action for NS5-branes wrapping curves in a warped region of a type IIB flux compactification, so we now give some essential background. We will begin by discussing D5-branes, to facilitate comparison with the string theory literature, even though our eventual interest will be NS5-branes.

The action of a D5-brane is the sum of a Dirac-Born-Infeld term related to the world-volume \mathcal{W} of the brane,

$$S_{\text{DBI}} = -g_s T_5 \int_{\mathcal{W}} \mathrm{d}^6 \sigma \, e^{-\Phi} \sqrt{-\det(G_{ab} + \mathcal{F}_{ab})} \,, \tag{3.5}$$

¹⁰For brevity we will often speak of "D3-branes," "D3-brane backreaction," etc., with the understanding that both D3-branes and anti-D3-branes are included.

and a Chern-Simons term encoding the coupling of the D5-brane to the Ramond-Ramond p-form potentials C_0 , C_2 , C_4 , and C_6 ,

$$S_{\rm CS} = \mu_5 \int_{\mathcal{W}} \sum_p C_p \wedge e^{\mathcal{F}}, \qquad (3.6)$$

with $\mathcal{F} = B + 2\pi \alpha' F$. Here g_s is the string coupling, T_5 is the D5-brane tension, G_{ab} is the metric induced on the D5-brane, μ_5 is the D5-brane charge, and \mathcal{F} is the gauge-invariant two-form field strength on the D5-brane. The integral in (3.6) picks out the six-forms C_6 , $C_4 \wedge \mathcal{F}, C_2 \wedge \mathcal{F} \wedge \mathcal{F}$, and $C_0 \wedge \mathcal{F} \wedge \mathcal{F} \wedge \mathcal{F}$.

Now suppose that $\mathcal{W} = \mathcal{M}^{3,1} \times \Sigma_2$, with Σ_2 a two-cycle in the internal six-manifold X_6 . If the field strength \mathcal{F} obeys¹¹

$$\frac{1}{\ell_s^2} \int_{\Sigma_2} \mathcal{F} = N \in \mathbb{Z} \,, \tag{3.7}$$

then the Chern-Simons coupling becomes

$$\mu_5 \int_{\mathcal{W}} C_4 \wedge \mathcal{F} \to N \mu_3 \int_{\mathcal{M}^{3,1}} C_4.$$
(3.8)

The interaction (3.8) is precisely N times the Chern-Simons coupling of a single D3-brane to the Ramond-Ramond four-form potential C_4 , under which the D3-brane is electrically (and also magnetically) charged. The coupling (3.8) should be understood as a generalization of the worldline coupling

$$\mathcal{L}_{\rm int} = -\frac{e}{c} \int A_{\mu} \,\mathrm{d}x^{\mu} \tag{3.9}$$

in electromagnetism. In particular, (3.8) shows that a D5-brane wrapping Σ_2 , with N units of \mathcal{F} flux on Σ_2 , carries N units of D3-brane charge. Equivalently, the D5-brane can be said to contain N D3-branes dissolved in the D5-brane. This fact, while well-known, will be crucial for our considerations.

The Σ_2 -wrapping D5-brane can fluctuate in the space orthogonal to $\mathcal{M}^{3,1} \times \Sigma_2$. We denote these corresponding canonically-normalized fluctuations as X^i . Defining the dimensionless field

$$b \equiv \frac{1}{\ell_s^2} \int_{\Sigma_2} B_2 \,, \tag{3.10}$$

we may expand the DBI action (3.5) to second order in these fluctuations,¹²

$$S_{\text{DBI}} = -\frac{T_5}{2} \int d^4x \, d^2z \, \sqrt{-g_4} \Biggl(\sqrt{4\tilde{g}_2 + \ell_s^4 b^2} + \partial_\mu X^i \partial^\mu X^i + \frac{4\tilde{g}_2}{4\tilde{g}_2 + \ell_s^4 b^2} \partial_a X^i \partial^a X^i + \dots \Biggr),$$

$$(1) \qquad (2) \qquad (3.11)$$

¹¹We define the string length to be $\ell_s \equiv 2\pi \sqrt{\alpha'}$.

¹²As in appendix C, we denote $\mathcal{M}^{3,1}$ indices with μ, ν , etc.; Σ_2 indices with a, b, etc.; directions orthogonal to $\mathcal{M}^{3,1} \times \Sigma_2$ with indices i, j, etc.; and we parameterize Σ_2 using the coordinates y and z, with $dy \wedge dz = d^2 z$; see table 3.

where \tilde{g}_2 is the determinant of the metric on Σ_2 . Upon integrating over Σ_2 and denoting its volume as ℓ^2 , (1) yields a four-dimensional potential for b

$$V(b) = \frac{\varepsilon}{(2\pi)^3 \alpha'^2} \sqrt{\left(\frac{\ell}{\ell_s}\right)^4 + \frac{b^2}{4}},\tag{3.12}$$

In the absence of a wrapped D5-brane, b would enjoy an approximate continuous shift symmetry, $b \mapsto b + \text{const.}$, that is broken to a residual exact discrete shift symmetry, $b \mapsto b + 1$, by instanton effects.¹³ However, the potential (3.12) induced by the D5-brane completely breaks this symmetry. In fact, the D5-brane introduces a monodromy, in that upon traversing the axion circle, from $b \mapsto b + 1$, the potential energy is increased, rather than being periodic. For large b, the potential (3.12) becomes linear, as claimed in (3.3) for the related case of an NS5-brane.

The strength of this symmetry breaking is proportional to ε , the warp factor $\varepsilon = e^{4A_{\cup}}$ at the location of the fivebrane. In a warped compactification, the ten-dimensional metric takes the form

$$ds_{10}^2 = e^{2A(y)} g_{\mu\nu} dx^{\mu} dx^{\nu} + e^{-2A(y)} \tilde{g}_{mn} dy^m dy^n.$$
(3.13)

By placing the fivebranes in a warped throat, the energy of this shift symmetry breaking can be gravitationally redshifted to an energy much smaller than the natural scale of breaking due to unwarped fivebranes.

The monodromy is closely related to the induced D3-brane charge (3.8). Starting from an initial configuration with $b = b_0$ and moving to $b = b_0 + N$ (for N > 0) corresponds to shifting

$$\mathcal{F} \mapsto \mathcal{F} + N\omega_2 \,, \tag{3.14}$$

with ω_2 a two-form obeying $\ell_s^{-2} \int_{\Sigma_2} \omega_2 = 1$. This is an increase, of N units, of the gaugeinvariant field strength \mathcal{F} . This change is manifest in the potential (3.12), which increases linearly. The change is also visible in the D3-brane charge carried by the D5-brane, which increases by N units. We refer to this process as "winding up the axion N times."

A justifiable complaint at this stage is that in a compact space, the total D3-brane charge should be fixed: in fact it must vanish by Gauss's law. So winding up the axion would appear to be forbidden. However, to cancel the D5-brane tadpole, we may suppose that in addition to the D5-brane wrapping Σ_2 , there is an anti-D5-brane also wrapping Σ_2 . The anti-D5-brane Chern-Simons coupling differs from the D5-brane Chern-Simons coupling (3.6) by an overall minus sign. Thus, winding up the axion N times induces N units of D3-brane charge on the D5-brane, as well as -N units of D3-brane charge on the anti-D5-brane, so that no net D3-brane charge is produced, and if Gauss's law is obeyed in the initial configuration, it is also obeyed after winding.

A coincident D5-brane and anti-D5 brane will quickly annihilate. However, if a D5brane wraps Σ_2 , and an anti-D5-brane wraps a two-cycle $\overline{\Sigma}_2$ that is *homologous* to Σ_2 , but is not coincident with Σ_2 , then the D5-brane/anti-D5-brane configuration can be metastable and cosmologically long-lived [24]. Because the induced D3-brane charges are determined

 $^{^{13}}$ Moduli-stabilizing effects further break this symmetry, as explained in [21].

by the homology classes of Σ_2 and $\overline{\Sigma}_2$, if $[\Sigma_2] - [\overline{\Sigma}_2]$ is trivial in homology then no net D3brane charge is induced, just as in the case of a strictly coincident D5-brane/anti-D5-brane pair, and Gauss's law does not preclude winding up the axion.

Let us summarize the physics of B_2 monodromy from a wrapped D5-brane. The D5brane is a source of monodromy and gives rise to the non-periodic potential (3.12). The order parameter measuring the distance from the origin in the *b* field space is the number of windings, $N \in \mathbb{Z}$, which also counts the D3-brane charge induced on the D5-brane. This is the *monodromy charge* in the fivebrane model. Winding up corresponds to moving away from the origin in field space and storing energy in the form of the D3-branes dissolved in the D5-brane, and anti-D3-branes dissolved in the anti-D5-brane: that is, the energy is stored in the monodromy charge.

The potential (3.12) is that of a *probe* D5-brane, in the same sense that (3.9) includes the potential energy of an electron in a background electromagnetic field. However, just as (3.9) also encodes the fact that electrons source electromagnetic fields, the couplings (3.5)and (3.6) encode the effects that a D5-brane has on the background fields. To determine this *backreaction* of the D5-brane on the bulk field, including the metric and the *p*-form fields, we simply include the couplings (3.5) and (3.6) when varying the ten-dimensional action with respect to these fields φ . Schematically,

$$0 = \frac{\delta}{\delta\varphi} S_{10d,\text{bulk}} + \frac{\delta}{\delta\varphi} S_{\text{DBI}} + \frac{\delta}{\delta\varphi} S_{\text{CS}} \,. \tag{3.15}$$

Any D5-brane serves as a source for the ten-dimensional metric (it has tension), and as a source for C_6 . But a D5-brane with

$$\frac{1}{\ell_s^2} \int_{\Sigma_2} \mathcal{F} = N \neq 0 \tag{3.16}$$

also serves as a source for C_4 ; this is just to say that such a D5-brane carries D3-brane charge. The DBI action (3.5) may be interpreted as the product of the brane tension and its "effective volume," which grows with N. This growth has two principal effects. The mass of the five-brane is also, schematically, the product of its tension and this effective volume, and thus as N grows the charged D5-brane will more strongly source the ten-dimensional metric. Furthermore, there are Kaluza-Klein excitations arising from the dimensional reduction of (3.5) whose masses *decrease* as this effective volume grows; indeed, the dimensional reduction of (2) and (3) in (3.11)—which correspond to the transverse fluctuations of the five-brane — yield Kaluza-Klein modes with masses m_{bKK} that are smaller than the naive estimate $m_{KK} \propto \ell^{-1}$ by a factor of (see appendix C.2)

$$\frac{m_{\rm bKK}}{m_{\rm KK}} \sim \frac{\ell^2}{\sqrt{\ell^4 + \ell_s^4 b^2}} \,.$$
 (3.17)

Axions descending from B_2 generically suffer an η problem [21], meaning that in expansion around a vacuum with stabilized moduli, the actual potential for the axion, taking into account all couplings to moduli, is very different from the potential (3.12) that arises from the probe D5-brane action alone. This problem can be ameliorated by considering an

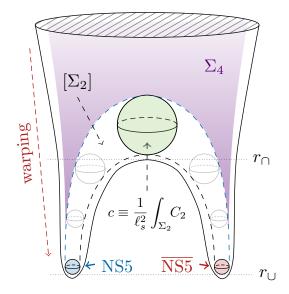


Figure 4. Ten-dimensional realization of B and C of (3.19). C is generated by strong gauge dynamics on seven-branes wrapping a divisor Σ_4 , which must necessarily intersect a holomorphic representative of the class $[\Sigma_2]$, such as those wrapped by the NS5-/anti-NS5-brane.

axion descending from the Ramond-Ramond two-form C_2 and exchanging the D5-branes in the above discussion for NS5-branes. The analogous potential is then given by

$$V(c) = \frac{\varepsilon}{(2\pi)^3 g_s \alpha'^2} \sqrt{\left(\frac{\ell}{\ell_s}\right)^4 + \frac{g_s^2 c^2}{4}}.$$
(3.18)

As we will argue in section 3.3, for a construction of a relaxion model via fivebrane axion monodromy in string theory one needs an extremely large winding $N \gg 1$. There is a correspondingly large induced D3-brane charge, the effect of which must be included in the ten-dimensional field equations. Backreaction of this charge and its effect on the five-brane cannot be neglected: the potential for the axion is no longer simply given by (3.12), and new light modes appear.

3.3 Fivebrane relaxion monodromy

To understand string theoretic constraints on the relaxion mechanism, we require an embedding of the four-dimensional potential

$$V(\phi, h) = \left(M^2 - g_h M (\phi_{\text{init}} - \phi) \right) |h|^2 + g M^3 \phi + V_{\text{stop}}(\phi, h) , \qquad (3.19)$$
(A) (B) (C)

or of something functionally equivalent, in a well-controlled compactification of string theory. As noted in the introduction, the ratio g_h/g need not be $\mathcal{O}(1)$, and so in (3.19) we distinguish between the two.

In section 3.2, we focused on realizing (B) as the potential energy of an NS5-/anti-NS5-brane pair wrapping the minimum volume representatives of the homology class [Σ_2] associated with the axion $c = \ell_s^{-2} \int_{\Sigma_2} C_2$, where $\phi \equiv fc$. (B) provides a potential that is self-similar (ignoring backreaction effects) over a very large distance $\Delta \phi \gg f$ in field space. Hubble friction eventually dominates and the late-time dynamics are independent of the initial conditions for ϕ .

Crucially, the small parameter g is controlled by the warp factor at the position r_{\cup} of the five-branes. Specifically,

$$gM^3 f \equiv \frac{2\pi}{\ell_s^4} e^{4A_{\cup}},$$
 (3.20)

where $e^{4A_{\cup}}$ is the warp factor at r_{\cup} , the location of the fivebranes and the bottom of the "tooth" in figure 4. We may think of the "roots" of the tooth as Klebanov-Strassler or similar warped throat geometries. Away from the tip, the warp factor is roughly $e^{4A} \sim r^4/L^4$, with L the characteristic size of the warped throat. A simple way to describe this warping is by the number of D3-branes it would take to form a similarly sized warped throat,

$$L^4 \sim g_s N_{\rm D3} \ell_s^4.$$
 (3.21)

As explained in section A, the axion decay constant f is determined by the radial position of the arch of the "tooth,"

$$f^2 \sim g_s \frac{r_{\cap}^2}{\ell_s^4},$$
 (3.22)

so the shift symmetry breaking scale is given by

$$gM^3 \sim \frac{2\pi}{g_s^{3/2} \ell_s^3 N_{\mathrm{D3}}} \left(\frac{r_{\mathrm{U}}}{r_{\mathrm{O}}}\right) \left(\frac{r_{\mathrm{U}}}{\ell_s}\right)^3. \tag{3.23}$$

The cutoff scale M depends on how the Higgs is realized and does not necessarily depend on the total D3-brane charge N_{D3} . However, regardless of where the Higgs is located in the internal space, the smallness of g is necessarily tied to a large N_{D3} . For example, if the Higgs sector is realized somewhere in the bulk geometry, then $M \propto N_{\text{D3}}^0$ and $g \propto N_{\text{D3}}^{-1}$, as in (3.23). If instead the Higgs sector is realized at the top of the warped throat at $r \sim L$ in figure 4, then we may take $M^3 \sim L^{-3}$ and so

$$g \sim \frac{1}{(g_s N_{\rm D3})^{1/4}} \left(\frac{r_{\rm \cup}}{r_{\rm \cap}}\right) \left(\frac{r_{\rm \cup}}{\ell_s}\right)^3. \tag{3.24}$$

We will not consider a Higgs realized deep within the warped throat, as this would lead to an exponential suppression of M, corresponding to a supersymmetric resolution of the hierarchy.

We will be agnostic about the detailed origin of the Higgs coupling (A). While its specific form would be relevant in a complete model, it is not needed to expose and quantify the issues that concern us here, which mainly deal with the interplay between the linear potential and the stopping potential. In the spirit of this agnosticism, we instead focus on the hierarchy generated between the string scale M_s and the electroweak scale v.

Even so, a concrete picture of one possibility may be helpful. The Higgs could arise from open strings stretching between stacks of D3-branes or D7-branes. The Higgs mass

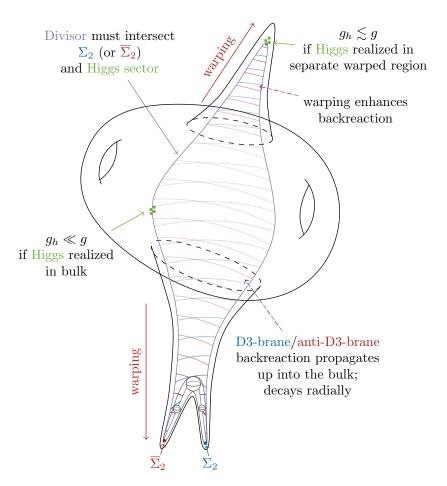


Figure 5. Schematic structure of the extra dimensions, showing a string theory setup that realizes the main relaxion features and couplings. The central region is the "bulk" of the extra dimensions, which does not experience position-dependent warping. The coupling g_h depends on where in the bulk Calabi-Yau the Higgs sector is realized.

is then proportional to the distance between the $U(1)_Y$ brane and the $SU(2)_W$ stack. The coupling \widehat{A} is generated by backreaction of the monodromy charge on the internal geometry, which changes distance between these branes and thus the Higgs mass, as in figure 5. Because this backreaction decays as it propagates throughout the six-dimensional space, there may be an appreciable hierarchy between g_h and g which depends on where the Higgs sector is realized in the internal geometry. We may mitigate this hierarchy somewhat by placing the Higgs in another warped throat — the backreaction will then be blue-shifted, leading to an increased coupling g_h — though, as explained above, placing the Higgs in a warped region will naturally suppress M.

Finally, for generic initial conditions, the relaxion traverses a distance $\Delta \phi \sim M/g_h$ in field space. This is associated with the dissipation of

$$N \sim \frac{\Delta\phi}{f} \sim \frac{M^4}{g_h M^3 f} \sim g_s N_{\rm D3} \left(\frac{g}{g_h}\right) \left(\frac{M}{M_s}\right)^4 \left(\frac{\ell_s}{r_{\cup}}\right)^4 \tag{3.25}$$

units of monodromy charge.

Barriers from D7-branes. We will be more specific about how (C) is realized. Perturbatively in g_s , the axion has a continuous shift symmetry $\phi \mapsto \phi + \text{const.}$ which is broken by non-perturbative effects (in g_s) to the discrete shift symmetry $\phi \mapsto \phi + f$. If (A) and (B) are to be the only terms that break this discrete shift symmetry, as is implicitly assumed in the relaxion construction, then V_{stop} must be generated non-perturbatively in g_s . As noted above, we take the Higgs and relaxion sectors to be separated in the internal geometry, and so in order for the stopping potential to depend on both of these sectors, it must be generated by physics on one or more extended objects — by either Euclidean D*p*-branes or strong gauge dynamics on a stack of D*p*-branes.

For simplicity, we will assume that V_{stop} is generated by the strong dynamics of a gauge theory, with group G, realized on a stack of D7-branes wrapping a holomorphic four-cycle Σ_4 , as illustrated in figure 4. The D7-branes couple to the C_2 axion through the Chern-Simons action

$$S_{\rm CS} \supset \mu_7 \int_{\mathcal{W}} \mathcal{F} \wedge C_2 \wedge \mathcal{F} \wedge \mathcal{F} \,. \tag{3.26}$$

A key observation is that the D7-branes must enter the warped throat region (see appendix B for a proof). The coupling (3.26) leads to a potential of the schematic form

$$V(\phi, v) = \Lambda_c^3 v \cos\left(\frac{2\pi\phi}{f}\right),\tag{3.27}$$

but can, in general, involve a more complicated polynomial of the Higgs vev v and a general f-periodic function in ϕ . The confinement scale Λ_c is naturally related to the string scale and the D7-brane gauge coupling $g_{\rm YM}$,

$$\Lambda_c^3 \propto \ell_s^{-3} \exp\left(-\frac{8\pi^2}{g_{\rm YM}^2 c_G}\right), \qquad (3.28)$$

where c_G is a constant determined by the particular effects that generate (3.27). For example, c_G is simply the dual Coxeter number of G if the stopping potential is realized through gaugino condensation. In known examples, c_G is at most $\mathcal{O}(10^2)$, and we will take $c_G = 1$ henceforth as it cannot be arbitrarily small. The generated hierarchy between the string and electroweak scales is then

$$\frac{M_s}{v} \propto g_s N_{\rm D3} \exp\left(-\frac{8\pi^2}{g_{\rm YM}^2}\right) \left(\frac{\ell_s}{r_{\rm \cup}}\right)^4. \tag{3.29}$$

Since $r_{\cup} \gtrsim \ell_s$, the hierarchy is controlled by the warp factor and at first sight appears to be proportional to N_{D3} . Thus, an arbitrarily large hierarchy could apparently be realized via substantially warping the source of monodromy. However, as we will discuss in section 4, this is too naive.

4 Microphysical constraints

Relaxation of a hierarchy by the relaxion mechanism occurs only in theories that meet several stringent requirements. Arguably the most challenging requirements from the viewpoint of ultraviolet completion in string theory are both the *large displacement* $\Delta \phi \sim M/g$, and the comparatively short stopping length. That is, the relaxion must evolve slowly over a large distance, gradually reducing the Higgs mass, but then rapidly come to rest after the Higgs acquires a vev. These disparate distance scales in field space correspond to very different energy scales in the potential: the final Higgs vev v is determined by the ratio of the shift symmetry breaking scale gM^3f to the stopping potential scale Λ_c^3 , cf. (1.4). In field theory, one can obtain a controllably large hierarchy by taking g to be extremely small while holding Λ_c fixed.

This limit is problematic in string theory. As we will show in section 4.2.3, the scale Λ_c depends on g, and is exponentially suppressed as $g \to 0$. This dramatically limits the hierarchy that can be generated.

At the same time, the large field excursion on its own implies that the initial configuration carries a very large monodromy charge. This gradually dissipating monodromy charge will serve as a changing source for the ten-dimensional equations of motion. For $N \gg 1$, this backreaction has profound effects on the compactification geometry and so on the four-dimensional relaxion potential (3.19).

It is tempting to argue that all of the corrections that result from backreaction must ultimately originate in the breaking of the axionic shift symmetry, and so must involve powers of the shift-symmetry breaking parameter g. This is *not* correct. In the NS5brane model, the breaking parameter g is small because the DBI action of an NS5-brane is proportional to the warp factor at the NS5-brane location, cf. (3.20). Backreaction effects sourced directly through the DBI action are indeed proportional to powers of g. However, the NS5-brane Chern-Simons action is not warped, and could not be: it is topological, and counts the (integer) D3-brane charge induced on the NS5-brane, i.e. the monodromy charge N.

Thus, backreaction effects sourced by the Chern-Simons action are proportional to N, without factors of g. For example, the integral of the R-R field strength F_5 over a Gaussian surface — say, an S⁵ — surrounding the NS5-brane is simply given by N, even as $g \to 0$. One consequence, as we shall see, is that the monodromy charge provides a large correction to the stopping potential.¹⁴

In this section we provide an array of calculations that reveal the concrete obstacles to achieving a large displacement and a short stopping length in the NS5-brane model.

4.1 Overview of microphysical constraints

We first preview a number of constraints on relaxion monodromy constructions, which originate from microphysical limitations on string compactifications that provide the desiderata listed in section 3.1. Each of these constraints will be detailed in turn in sections 4.2.2–4.2.6.

- §4.2.1 Universal effect on the geometry. The shape of the warped throat region is dramatically altered by backreaction, leading to large changes in the effective action.
- §4.2.2 Tadpole constraints. To accommodate $N \gg 1$ units of monodromy charge without the loss of perturbative control, we must construct a background throat with

¹⁴The backreaction sourced by this topological term does not need to propagate far to be "detected," i.e. to influence a significant term in the four-dimensional Lagrangian: see appendix B.

 $N_{\text{D3}} \gg N \gg 1$. Gauss's law — i.e., the D3-brane charge tadpole — then implies that there must be a source that is equivalent to $-N_{\text{D3}}$ D3-branes. To avoid the instabilities created by a large number of actual anti-D3-branes, this source must be supersymmetric, and arise from the topology of an elliptically-fibered fourfold: the D3-brane charge is then $-\chi/24$, where χ is the Euler number of the fourfold. The largest known Euler number of an elliptically-fibered fourfold is 1,820,448. So in this setting, N will have to be much smaller than 75,852.

- §4.2.3 Barrier suppression from warping. The D7-branes that generate the stopping potential must wrap a four-cycle Σ_4 that intersects the minimum volume two-cycle Σ_2 . The D7-brane gauge coupling function, which depends on the warped four-volume of Σ_4 , is then directly suppressed by the same warping responsible for the miniscule monodromy energy scale (3.20). A weakly broken shift symmetry therefore leads to extremely small barriers.
- §4.2.4 Barrier suppression from backreaction. The induced D3-brane charge and tension backreact on the warped four-volume of Σ_4 and therefore perturb the D7-brane gauge coupling function. This perturbation introduces an exponential dependence of the gauge coupling on the monodromy charge N, with no powers of $g \sim N_{\text{D3}}^{-1}$. This contradicts naive applications of technical naturalness: the dangerous term that arises is not negligible in the limit $g \to 0$ where the shift symmetry breaking is weak.
- §4.2.5 Effects on the moduli potential. The sources responsible for Kähler moduli stabilization are exponentially sensitive to perturbations of the warp factor. So the moduli potential depends on the relaxion field, i.e. there are new terms in the relaxion potential not captured by (3.19). This was extensively studied in [23].
- **§4.2.6** Effects on the axion decay constant. Large backreaction will also affect the axion decay constant, which depends on the volume of the cycle the axion threads as well as on the overall volume of the internal manifold.
- §4.2.7 Classical annihilation of the dipole. The compactification detailed in section 3.2 is metastable. The NS5-brane and anti-NS5-brane attract one another because of the induced D3-brane charge that each carries, but the fivebranes must stretch over a large-volume representative of $[\Sigma_2]$ in order to meet one another. This costs energy, because the fivebranes have tension. For modest windings N, the tension energy can be much larger than the Coulomb energy from the D3-branes and anti-D3-branes, and the system is controllably metastable. However, for $N \gg 1$, the Coulomb energy can overpower the tension energy, and the fivebrane/anti-fivebrane dipole can classically annihilate.
- §4.2.8 Constraints from anti-D3-brane annihilation. An anti-D3-brane at the tip of a large Klebanov-Strassler throat is a metastable and cosmologically long-lived configuration. However, the barrier that ensures metastability depends on the number of anti-D3-branes in the throat. For some number $N_{\rm KPV}$ of anti-D3-branes and thus for

windings $N \ge N_{\text{KPV}}$ — the barrier disappears and the anti-D3-branes can classically annihilate against the flux of the throat. Thus, the accumulation of anti-D3-branes on the anti-NS5-brane creates a risk of instability.

§4.2.9 Tunneling via light brane KK modes. The accumulation of D3-branes in the NS5brane pair leads to a reduction in the tension of the NS5-branes, and correspondingly a reduction in the mass of Kaluza-Klein excitations of the NS5-branes. This Kaluza-Klein spectrum has spacing proportional to m_0/N when the axion is wound up by N cycles, with m_0 associated to the IR scale of the warped throat. These light brane KK modes provide another pathway for classical annihilation of the dipole. If the throat is put at some temperature, say from a source of supersymmetry breaking elsewhere in the internal space, thermal fluctuations of the light brane KK modes could enable the NS5-branes to reach up towards one another, allowing for a quantum mechanical tunneling event.

4.2 Consequences of D3-brane backreaction

D3-branes and anti-D3-branes source warping, and so the D3-brane dipole that develops when the axion is wound up leads to a change in the local warp factor. The warped throat region is itself produced by some number N_{D3} of D3-branes that have dissolved into flux, and when the number of windings N becomes comparable to N_{D3} , the D3-brane dipole is a large correction to the background in which it is sitting. The probe approximation is not valid for such a configuration, and the backreaction of the D3-brane dipole affects many couplings in the four-dimensional theory.

4.2.1 Universal effect on geometry

The backreaction of the tension and charge of N induced D3-branes will be a small perturbation to the overall configuration as long as the ratio $g_s \ell_s^4 N/L^4$ is small, where L is the radius of the warped throat: see appendix D. Using N_{D3} to denote the effective D3-brane charge of the warped throat (3.21) we must require that

$$N \ll N_{\rm D3}.\tag{4.1}$$

To intuitively motivate (4.1), we may replace the N D3-branes with an AdS_5 warped throat with radius

$$R_N^4 \sim g_s \ell_s^4 N \tag{4.2}$$

via a geometric transition. The perturbed geometry will be drastically different unless size of this extra throat is much smaller than the original warped throat, $R_N^4 \ll L^4$. So, we require that $N \ll N_{\text{D3}}$ in order to maintain perturbative control.

The volume of the warped throat is necessarily bounded by the total volume of the internal space, $L^6 \leq \ell_s^6 \mathcal{V}_{\rm E}^{,15}$ which determines the four-dimensional Planck mass via $M_{\rm pl}^2 \ell_s^2 = 4\pi \mathcal{V}_{\rm E}$. From (4.1) we find the constraint

$$N \ll \frac{1}{g_s} \frac{M_{\rm pl}}{M_{\rm KK}} \,, \tag{4.3}$$

¹⁵We denote the total volume of the internal space X_6 , measured in Einstein frame, as $\ell_s^6 \mathcal{V}_{\rm E}$,

where $M_{\rm KK} = 1/(\ell_s \mathcal{V}_E^{1/6})$. This imposes a constraint on the number of windings for reasonable hierarchies between the compactification and Planck scales, and for reasonable values of g_s .

4.2.2 Tadpole constraint

The higher-dimensional equations of motion must be satisfied in a consistent string compactification. In particular, the higher-dimensional analog of Gauss's law for the five-form flux F_5 becomes a powerful constraint on the ten-dimensional configuration. In a noncompact manifold, flux lines are allowed to extend to infinity and Gauss's law places no constraint on the amount of charge allowed in a given configuration. However, in a compact manifold a flux line must end on a charge and Gauss's law provides a *tadpole constraint*: the total amount of D3-brane charge in the compactification must vanish. As discussed above, the warped throats pictured in figure 3 are supported by a total of N_{D3} units of D3-brane charge. The tadpole constraint requires that this charge be canceled elsewhere in the Calabi-Yau geometry.

This cancellation could occur by including anti-D3-branes elsewhere in the internal space, or by forming another, oppositely charged, warped throat elsewhere with a large amount of negative D3-brane charge. In both cases, the D3-branes supporting the relaxion's warped throat and these additional anti-D3-branes will attract and the entire model will generically be unstable.

Fortunately, there exist well-known sources of supersymmetric negative D3-brane charge, and thus one may satisfy the tadpole constraint while maintaining stability. Sevenbranes wrapping non-trivial cycles in the internal space provide curvature-induced negative D3-brane charge. F-theory compactified on elliptically-fibered Calabi-Yau fourfolds provides a framework for analyzing type IIB compactifications at arbitrary coupling, and the negative charge is related to the fourfold's Euler number $\chi(CY_4)$ via

$$N_{\rm D3}^{\rm CY_4} = -\frac{\chi(\rm CY_4)}{24}.$$
(4.4)

The largest known Euler number of an elliptic-fibered Calabi-Yau fourfold is $\chi(CY_4) = 1,820,448$ [25], which imposes the constraint

$$N_{\rm D3} \le 75,852.$$
 (4.5)

Requiring $N \ll N_{D3}$ to maintain control over the configuration, we then have the constraint

$$N \ll 75,852.$$
 (4.6)

The bound (4.5) on the Euler number of known fourfolds thus translates to a strong upper limit on the number of windings, and so constrains the maximum possible field excursion undergone by the relaxion.

The bound (4.6) applies only in the present case in which the monodromy charge is D3-brane charge. However, in alternative axion monodromy scenarios, it would still be necessary to arrange that the background solution at zero winding carries a large background monodromy charge analogous to N_{D3} . In such a setting we expect topological upper bounds analogous to (4.5) on the amount of monodromy charge that can be included without creating rapid instabilities.

4.2.3 Suppression from warping

Relaxation of a large hierarchy requires that the shift symmetry is very weakly broken, with $g \ll 1$. In the ten-dimensional model of section 3.3, the breaking is made small by placing the source of monodromy — NS5-branes wrapping the minimum-volume two-cycles Σ_2 and $\overline{\Sigma}_2$ — in a heavily warped region. However, we prove in appendix B that supersymmetric D7-branes can generate a relaxion stopping potential only if the four-cycle Σ_4 they wrap intersects a holomorphic representative of the class [Σ_2]. In the simplest constructions, the D7-brane stack responsible for the stopping potential descends into the warped region.¹⁶ As we will now see, elementary locality arguments show that the small parameter associated with this warping, g, in (A) and (B) of (3.19) then generically infects the stopping potential suppression of the stopping potential barriers.

The gauge coupling $g_{\rm YM}$ on a spacetime-filling D7-brane wrapping a four-cycle Σ_4 is proportional to the *warped* four-volume of Σ_4 in string units,

$$\frac{1}{g_{\rm YM}^2} = \frac{1}{2\pi\ell_s^4} \int_{\Sigma_4} d^4\xi \,\sqrt{\tilde{g}_4} \,e^{-4A},\tag{4.7}$$

with \tilde{g}_4 the induced, unwarped metric on Σ_4 . Defining a reference warp factor profile $\exp(4\bar{A}) = r^4/L^4$, cf. (3.21), we may express (4.7) as

$$g_{\rm YM}^{-2} = \alpha^{-1} g_s N_{\rm D3}, \tag{4.8}$$

where

$$\alpha^{-1} \propto \int_{\Sigma_4} \mathrm{d}^4 \xi \, \sqrt{\tilde{g}_4} \, r^{-4} e^{-4(A-\bar{A})}$$
(4.9)

is a dimensionless coefficient capturing the geometry of the embedding of Σ_4 in the warped throat.

We may estimate α as follows. We have shown that Σ_4 must reach down the warped throat to intersect Σ_2 at r_{\cup} . Assuming that Σ_4 roughly factorizes into a radial part and an angular part with volume \check{v} , that it extends up into the bulk geometry as in figure 5, and that the integral is dominated in the region where $A \sim \bar{A}$, we find

$$\alpha^{-1} \gtrsim \breve{v} \log\left(\frac{L}{r_{\cup}}\right) \sim \breve{v} \log\left(\frac{g_s N_{\mathrm{D3}} \ell_s^4}{r_{\cup}}\right).$$
 (4.10)

Importantly, α^{-1} is not naturally $\mathcal{O}(N_{\text{D3}}^{-1})$, and in fact grows with the size of the throat, $L^4 \propto N_{\text{D3}}$. So, $g_{\text{YM}}^{-2} \sim \mathcal{O}(N_{\text{D3}})$ unless the angular volume \breve{v} is finely tuned to be exceptionally small, to one part in g^{-1} , which is of order the desired hierarchy. In other words, fine-tuning the angular volume \breve{v} to eliminate the effects of this warping amounts to constructing the entire hierarchy by this fine-tuning. This suppression therefore renders the relaxation mechanism ineffectual.

From (3.28) and (4.8), the stopping potential is exponentially suppressed in N_{D3} ,

$$\Lambda_c^3 \propto \ell_s^{-3} \exp\left(-\gamma_{\rm bg} N_{\rm D3}\right),\tag{4.11}$$

¹⁶Even in more elaborate constructions, the backreaction problem persists, as we explain in section 4.2.5.

with $\gamma_{\rm bg} \sim 8\pi^2/(g_s \alpha c_G)$. The hierarchy generated including this suppression is then

$$\frac{M_s}{v} \sim g_s \gamma_{\rm bg}^{-1} \left(\gamma_{\rm bg} N_{\rm D3} e^{-\gamma_{\rm bg} N_{\rm D3}} \right) \left(\frac{\ell_s}{r_{\cup}} \right)^4, \tag{4.12}$$

and since $xe^{-x} \leq e^{-1}$, the maximum resolvable hierarchy is simply

$$\frac{M_s}{v} \sim g_s \gamma_{\rm bg}^{-1} \sim \alpha \, c_G \tag{4.13}$$

which is, crucially, not $\mathcal{O}(g^{-1})$ unless α is severely fine tuned.

Generically, the warping responsible for the suppression of the shift symmetry breaking energy scale also suppresses the scale of the stopping potential. This suppression drives a runaway relaxion, and precludes the dynamical generation of a large hierarchy in the absence of an acute fine tuning.

We expect this suppression to be very general. We argued in section 3.3 that the stopping potential must be generated by non-perturbative effects on a (p+1)-dimensional extended object, and Lorentz invariance requires this extended object to either fill space-time and wrap an internal cycle (p > 3) or else be instantonic. For a D*p*-brane wrapping a *p*-cycle Σ_p , the gauge coupling is given by

$$\frac{1}{g_{\rm YM,p}^2} = \frac{1}{2\pi\ell_s^{p+1}} \int_{\Sigma_p} \mathrm{d}^{p-3}\xi \,\sqrt{\tilde{g}_{p-3}} \,e^{(7-p)\Phi/4 - (p-3)A}.$$
(4.14)

Similarly, for a Euclidean Dp-brane wrapping the same cycle, the action is

$$S_{\rm EDp} = \frac{2\pi}{\ell_s^{p+1}} \int_{\Sigma_p} \mathrm{d}^{p+1} \xi \, \sqrt{\tilde{g}_{p+1}} \, e^{-(p-3)\Phi/4 - (p+1)A}. \tag{4.15}$$

Both depend on powers of e^{-A} and thus *positive* powers of N_{D3} . So, any potential barrier generated by these effects will suffer from the same exponential suppression, albeit with different powers of N_{D3} .

4.2.4 Suppression from backreaction

The backreaction of D3-brane charge is out of control unless the induced D3-brane charge N is a small fraction of the total D3-brane charge forming the throat, $N/N_{\rm D3} \ll 1$, so that we may perform a perturbative expansion of the ten-dimensional field configuration in this ratio. We should thus expect corrections to (3.19) to involve powers of $N/N_{\rm D3}$, which is consistent with the expectation that, because the monodromy charge is related to the shift symmetry breaking, any corrections due to backreaction will come dressed with powers of g. Crucially, however, it is *fractional* corrections to the field configurations — i.e. $\delta \varphi/\varphi$ for some field φ — that involve powers of $N/N_{\rm D3}$. If some quantity — say, a D7-brane gauge coupling function — also scaled with $N_{\rm D3} \propto g^{-1}$, the absolute (additive) correction to this quantity is *not* necessarily small when $N/N_{\rm D3} \ll 1$.

Indeed, the monodromy charge induces a perturbation to (4.7),

$$\delta\left(\frac{8\pi^2}{g_{\rm YM}^2}\right) \sim g_s N_{\rm D3} \int_{\Sigma_4} d^4 \xi \sqrt{\tilde{g}_4} \, e^{-4(A-A_0)} \, r^{-4} \underbrace{\left(\frac{1}{2} \frac{\delta \tilde{g}_4}{\tilde{g}_4} - \frac{\delta e^{4A}}{e^{4A}}\right)}_{\mathcal{O}(N/N_{\rm D3})} \propto g_s N \equiv \frac{\gamma_{\rm br} \phi}{f}. \tag{4.16}$$

As discussed in detail in appendix D, the fractional perturbations are $\mathcal{O}(N/N_{\text{D3}})$ and thus the entire perturbation to the gauge coupling is $\mathcal{O}(g_s N)$. We have again grouped specific geometric details into a coefficient γ_{br} .

In the introduction, we gave an interpretation of this backreaction in terms of new light states entering the spectrum of the theory upon a monodromy $\phi \mapsto \phi + f$. Open/closedstring duality dictates that the supergravity (closed-string channel) correction (4.16) must match the one-loop correction to the gauge coupling $g_{\rm YM}$ calculated in the open-string channel. In the open-string picture of the configuration pictured in figure 4, we are interested in the one-loop correction to the ${\rm SU}(N_c)$ gauge theory living on the D7-brane stack wrapping Σ_4 , in the presence of N D3-branes dissolved in the NS5-brane on Σ_2 and N anti-D3-branes dissolved in the anti-NS5-brane on $\bar{\Sigma}_2$. Crucially, the N D3-branes introduce N light 3-7 strings transforming in the fundamental of ${\rm SU}(N_c)$, which provide a contribution to the one-loop β -function (1.12).

Accounting for this backreaction changes the structure of the potential (3.19). In particular, from (4.16) the monodromy charge induces further relaxion-dependence of the height of the stopping potential barriers,

$$\Lambda^4(v) \to \Lambda^4(\phi) e^{-\gamma_{\rm br}\phi/f}.$$
(4.17)

A priori, it is not obvious that $\gamma_{\rm br}$ is either always positive or always negative, so we will consider $\gamma_{\rm br} > 0$ and $\gamma_{\rm br} < 0$ separately. Assuming that the Higgs quartic coupling takes the form

$$\mathcal{L}_h \supset -\frac{\lambda}{2} |h|^4, \tag{4.18}$$

v is given by

$$v(\phi) = \sqrt{\frac{gM}{\lambda} (\phi_{\rm h} - \phi)}, \qquad (4.19)$$

where $\phi_{\rm h} = \phi_{\rm init} - M/g$ is generically $\mathcal{O}(M/g)$, and thus the corresponding induced monodromy charge when the Higgs develops a vev is very large, $N_{\rm h} \equiv \phi_{\rm h}/f \gg 1$. Ignoring the backreaction effect (4.16) and assuming that $f \ll \lambda v^2/gM$, the relaxion will stop rolling when

$$\frac{\Lambda_c^3}{f} \sqrt{\frac{gM}{\lambda} \left(\phi_{\rm h} - \phi\right)} \sim gM^3 \tag{4.20}$$

and it will be stabilized at

$$\phi_h - \phi \sim \frac{\lambda}{gM} \left(\frac{gM^3f}{\Lambda_c^3}\right)^2.$$
 (4.21)

If we now include the backreaction (4.16), (4.20) becomes

$$\frac{2\gamma_{\rm br}}{f} \left(\phi_{\rm h} - \phi\right) e^{2\gamma_{\rm br}(\phi_{\rm h} - \phi)/f} \sim \frac{2\lambda}{gMf} \frac{\gamma_{\rm br}}{(1 + \gamma_{\rm br})^2} \left(\frac{gM^3f}{\Lambda_c^3}\right)^2 e^{2\gamma_{\rm br}\phi_{\rm h}/f}.$$
(4.22)

Because $\phi_{\rm h}/f \gg 1$, the asymptotic behavior of solutions to (4.22) is determined solely by the sign of $\gamma_{\rm br}$. For $\gamma_{\rm br} > 0$, the stopping potential barriers are exponentially suppressed by the backreaction and (4.22) predicts that the relaxion stops at

$$\phi \sim -\frac{f}{2\gamma_{\rm br}} \log\left(\frac{2\lambda\gamma_{\rm br}}{(1+\gamma_{\rm br})^2} \frac{gM^3f}{\Lambda_c^4} \frac{M^2}{\Lambda_c^2}\right) < 0.$$
(4.23)

However, the linear potential $gM^3\phi$ in (3.19) is only an approximation for a potential of the form (3.18) and cannot be used for arbitrarily small values of ϕ/f . From (4.23) we see that this approximation breaks down. We should therefore understand (4.23) as an indication that the relaxion stops roughly when it has dissipated all of its charge, near $\phi = 0$. The electroweak scale is then fixed at

$$v \sim \frac{gM}{\lambda} \phi_{\rm h} \sim \frac{M^2}{\lambda},$$
 (4.24)

leaving the hierarchy unresolved.

For $\beta < 0$, the barriers are exponentially *enhanced*, and the relaxion stops at

$$\phi \sim \phi_{\rm h} - \frac{\lambda}{gM} \left(\frac{gM^3f}{\Lambda_c^3}\right)^2 \frac{1}{(1+\gamma_{\rm br})^2} e^{2\gamma_{\rm br}\phi_{\rm h}/f},\tag{4.25}$$

and the electroweak scale

$$v \sim \frac{gM^3f}{\Lambda_c^3} \frac{1}{|1+\gamma_{\rm br}|} e^{-|\gamma_{\rm br}|\phi_{\rm h}/f}$$
 (4.26)

is suppressed by the backreaction.

Can one use this barrier enhancement to save the relaxion from the exponential suppression discussed in section 4.2.3? Unfortunately, this backreaction enhancement is not enough to overcome the suppression from warping. We may combine (4.26) with (4.11)to find

$$\frac{M_s}{v} \sim \frac{g_s |1+\beta|}{(\gamma_{\rm bg} - |\gamma_{\rm br}| N_{\rm h}/N_{\rm D3})} \left(\frac{\ell_s}{r_{\rm \cup}}\right)^4 \left[(\gamma_{\rm bg} N_{\rm D3} - |\gamma_{\rm br}| N_{\rm h}) e^{-\gamma_{\rm bg} N_{\rm D3} + |\gamma_{\rm br}| N_{\rm h}} \right].$$
(4.27)

Since we require that $N_{\rm h}/N_{\rm D3} \ll 1$ for control and we expect the geometric factors to be on the same order $\gamma_{\rm bg} \sim |\gamma_{\rm br}|$, (4.27) implies that the necessary fine-tuning is still of the same order as the hierarchy one wishes to generate.

4.2.5 Effects on the moduli potential

In section 4.2.4 we considered the backreaction of D3-brane charge on the gauge coupling of the D7-branes that generate the stopping potential. As shown in appendix B, this particular D7-brane stack must enter the strongly warped region, and so the backreaction does not need to propagate far to impact them. The result is a very large change in the gauge coupling of the D7-brane worldvolume theory, leading to exponential suppression of the stopping potential.

Let us now ask about the impact of backreaction on the moduli potential. In the NS5brane scenario, the Kähler moduli of the compactification are stabilized by nonperturbative effects on a collection of four-cycles, either Euclidean D3-branes or gaugino condensation on D7-branes. The moduli potential also involves exponentials of the warped volumes of these cycles. Backreaction of D3-brane charge will change the warped volumes of these cycles, and so the moduli potential will typically be a rapidly varying function of the relaxion ϕ .

The argument of appendix B does not imply that the four-cycles supporting the Kähler moduli potential enter the warped region, so in contrast to section 4.2.4, the backreaction

has to propagate across the internal geometry to influence the moduli potential. It is tempting to argue that backreaction has a negligible effect on a sufficiently distant fourcycle. This is *not* correct. We will give a heuristic explanation here, and refer the reader to [6] for a complete quantitative treatment.

To understand whether backreaction of D3-brane charge can decouple from D7-branes on a particular four-cycle Σ_4 , we work in the open string picture, where the effect of backreaction is translated into the open string one-loop threshold correction to the D7-brane gauge coupling. On very general grounds, this effect gives non-negligible contributions to the relaxion potential unless the masses M_{3-7} of the stretched open strings obey

$$M_{3-7} \gtrsim M_{\rm pl} \,, \tag{4.28}$$

for then the non-renormalizable operators coupling the relaxion to the moduli are suppressed by more than the Planck mass. In a compact space, the diameter of the space determines an upper bound on the mass M_{3-7} , and one finds that at weak coupling and large volume, $M_{3-7} \ll M_{\rm pl}$ [6] (cf. also [26]). This is easily checked in simple geometries, but holds more generally.¹⁷

The upshot is that the four-cycles supporting the moduli potential cannot be taken far enough away from the source of monodromy to avoid significant backreaction: the moduli potential depends strongly on ϕ . One consequence is that the relaxion potential is *not* simply given by the probe DBI action (3.18), but instead has important contributions from couplings to moduli. This is an incarnation of the eta problem, which hinders the construction of natural models of inflation.

If all the other obstacles enumerated here could be overcome in some manner, leaving only the problem of relaxion couplings to moduli induced by backreaction, then one could attempt to fine-tune the orientation of the source of monodromy with respect to the configuration of four-cycles in the bulk of the compactification. The idea is that if the leading multipoles of the backreaction can be made to vanish on the four-cycle "receiver" by finetuning the relative orientation, then the residual effect of the subleading multipoles might be negligible. This approach was proposed and analyzed in [23], where it was shown that the backreaction coefficient $\gamma_{\rm br}$ goes as $(r_{\odot}/r_{\rm bulk})^m$, with m an $\mathcal{O}(1)$ integer determined by the lowest unsuppressed multipole, and that one could realize $\gamma_{\rm br} \sim 10^{-2}$ even for relatively small hierarchies between $r_{\rm bulk}$ and r_{\odot} with moderate fine-tuning. So, for a modest winding number $N \sim 100$, this backreaction on the moduli potential can be ameliorated and the most dangerous couplings can be removed. It is not clear, however, that this method is applicable for the extremely large windings $N \gtrsim 10^6$ that arise in relaxion constructions.

Throughout this work we have taken the homology class $[\Sigma_2]$ wrapped by the NS5brane to be localized in the warped throat, as in figures 3 and 4. That is, we assumed that the harmonic two-form dual to $[\Sigma_2]$ is principally supported in the warped region, and every holomorphic representative of $[\Sigma_2]$ is in the warped region. This localization is automatic in the particular constructions given in [23, 24, 28]. The argument of appendix B

¹⁷This fact is responsible for the well-known problem that brane-antibrane potentials are generically too steep to support inflation [27].

only dictates that the object generating the relaxion stopping potential must intersect a holomorphic representative of $[\Sigma_2]$. If we place this holomorphic representative far from the warped throat, one might then hope to suppress this backreaction.

However, this placement has an unintended consequence. As discussed in appendix A, worldsheet instantons wrapping a holomorphic representative outside the warped throat have much smaller actions than those inside. In order to maintain control over the instanton expansion, the axion decay constant must be much larger than if all holomorphic representatives lived inside the warped throat, typically $f \sim 10^{-2} M_{\rm pl}$ [29]. The extremely large number of windings N then translates into a vastly super-Planckian field displacement. Quantum gravitational corrections, going in powers of $\phi/M_{\rm pl} \sim N$, will infect all sectors of the relaxion potential (1.1), including the stopping potential. So, even if one is able to realize such super-Planckian displacements in quantum gravity — a famously challenging problem — one must also ensure, by fine-tuning, that the stopping potential does not receive disastrous corrections. One may try to avoid such super-Planckian displacements by working with an anomalously small axion decay constant, but this again introduces a $\mathcal{O}(N^{-1})$ fine-tuning.¹⁸

4.2.6 Effects on axion decay constants

As discussed in appendix A, the relaxion decay constant f only depends on the sixdimensional metric \tilde{g}_{mn} , both through the explicit factors of \tilde{g}_{mn} in its definition (A.3) and implicitly via the defining equation of the harmonic form $\Delta \Omega = 0$. Additional D3brane charge will not perturb f: \tilde{g}_{mn} is Ricci-flat in supersymmetric compactifications and additional D3-brane charge will preserve the same supercharges as the three-branes forming the warped throat. However, anti-D3-branes break the remaining supersymmetry. The anti-D3-brane charge backreacts on the six-dimensional metric and perturbs the axion decay constant,

$$\frac{\delta f^2}{M_{\rm pl}^2} = \frac{g_s}{2\mathcal{V}_E\ell_s^6} \int \mathrm{d}^6 y \sqrt{\tilde{g}} \tilde{g}^{mp} \tilde{g}^{nq} \left(2\Omega_{mn}\delta\Omega_{pq} + \frac{1}{2} \tilde{g}^{rs}\delta\tilde{g}_{rs}\,\Omega_{mn}\Omega_{pq} - 2\Omega_{mn}\Omega_{ps}\delta\tilde{g}_{qr}\tilde{g}^{rs} \right). \tag{4.29}$$

The anti-D3-brane charge does not substantially perturb the four-dimensional Planck mass $M_{\rm pl}^2$, as $\mathcal{V}_{\rm E}$ is dominated by the volume of the bulk Calabi-Yau. In what follows, we estimate the size of each of these terms.

The first term in (4.29) vanishes at first order, as the perturbation $\delta\Omega$ is orthogonal to the unperturbed Ω . To analyze the contribution from the second and third terms in (4.29), we must backreact the anti-D3-brane charge on the metric \tilde{g}_{mn} . As detailed in appendix D,

¹⁸Note that our $\mathcal{O}(N^{-1})$ fine-tuning estimate is conservative even in the case where the D7-brane stack intersects a holomorphic representative of $[\Sigma_2]$ far from the NS5-/anti-NS5-brane. For example, in the NS5-brane model we study in this paper, it was explicitly shown in [23] that dodging the effects of this monodromy charge buildup requires arranging *tuned* brane configurations with $\mathcal{O}(N^{-1})$ precision. Similarly, the actual fine-tuning necessary to avoid this backreaction in the "sequestered" case may turn out to be stronger than the conservative estimate given here.

the dominant metric perturbations are

$$\tilde{g}_{mn} dy^m dy^n \sim \left(1 + \frac{\alpha N}{N_{\text{D3}}} \left(\frac{r'}{r}\right)^8\right) dr^2 + r^2 \left(1 + \frac{\beta N}{N_{\text{D3}}} \left(\frac{r'}{r}\right)^{19/2} \mathcal{Y}^{\frac{1}{2},\frac{1}{2},1}(\Psi)\right) \breve{g}_{\theta\phi} d\xi^\theta d\xi^\phi$$
(4.30)

where the coefficients α and β and the angular function $\mathcal{Y}^{\frac{1}{2},\frac{1}{2},1}(\Psi)$ depend on the details of the compactification. The perturbation to the decay constant (4.29) is then

$$\frac{\delta f^2}{M_{\rm pl}^2} \propto \frac{g_s}{2\mathcal{V}_E \ell_s^6} \frac{N}{N_{\rm D3}} \int_{r_{\rm U}}^{r_{\rm O}} \mathrm{d}r \, r \, \left(\frac{r_{\rm U}}{r}\right)^{19/2} \propto \frac{f^2}{M_{\rm pl}^2} \frac{N}{N_{\rm D3}} \left(\frac{r_{\rm U}}{r_{\rm O}}\right)^{19/2}.$$
(4.31)

Because of the large exponent this is a comparatively weak constraint.

4.2.7 Classical annihilation of the dipole

The fivebrane configuration detailed in section 3.2 is metastable. If the fivebranes were tensionless, the induced D3-brane charge on the NS5-brane would attract the induced anti-D3-brane charge on the anti-NS5-brane, and these branes would classically annihilate. However, this Coulomb attraction is balanced by the fivebrane tension — in order for the fivebranes to meet, they must stretch over the large two-cycle Σ_{\cap} at the junction of the two warped throats, r_{\cap} in figure 3, which costs an energy

$$V_t \sim \frac{2\pi}{\ell_s^4} \frac{e^{4A_{\cap}}}{\sqrt{g_s}} \sqrt{4\left(\frac{\operatorname{vol}\Sigma_{\cap}}{\ell_s^2}\right)^2 + N^2}.$$
(4.32)

This potential energy barrier ensures the configuration is metastable, and can be exponentially long-lived. However, for large enough winding, we expect the Coulomb force — which scales as N^2 — to overpower this "tension force," allowing the fivebranes to classically annihilate.

The potential energy density of a probe D3-brane is proportional to Φ_{-} ,

$$V = \frac{2\pi}{\ell_s^4} \left(e^{4A} - \alpha \right) = \frac{2\pi}{\ell_s^4} \Phi_-.$$
(4.33)

For N D3-branes and N anti-D3-branes, the Coulomb potential energy density is then

$$V_{\rm c} \sim \frac{2\pi N}{\ell_s^4} \delta \Phi_-, \tag{4.34}$$

where $\delta \Phi_{-}$ (cf. appendix D) is the perturbation to Φ_{-} due to N anti-D3-branes, measured at the location of the N D3-branes. Because the D3-branes live in a separate warped throat, $\delta \Phi_{-}$ must first propagate up the antibrane throat from the anti-NS5-brane location $r = r_{\cup}$ to the surface $r = r_{\cap}$,

$$\delta \Phi_{-,\overline{\text{D3}}}^{I_s} \sim -\frac{N}{N_{\text{D3}}} e^{4A_{\cup}} \left(\frac{r_{\cap}}{r_{\cup}}\right)^{2+\Delta_s},\tag{4.35}$$

which then propagates down the D3-brane warped throat via the homogeneous modes

$$\delta \Phi_{-,\mathrm{D3}}^{I_s} = c_1 \left(\frac{r_{\cap}}{r}\right)^{\Delta_s + 2} + c_2 \left(\frac{r}{r_{\cap}}\right)^{\Delta_s - 2}.$$
(4.36)

We may think of the perturbation (4.35) as specifying a boundary value for the perturbation (4.36). Generically, we have

$$c_1, c_2 \sim \frac{N}{N_{\rm D3}} e^{4A_{\cup}} \left(\frac{r_{\cap}}{r_{\cup}}\right)^{2+\Delta_s} \tag{4.37}$$

so that, at the position of the D3-brane charge,

$$\delta\Phi_{-,\mathrm{D3}}^{I_s} \sim -\frac{N}{N_{\mathrm{D3}}} e^{4A_{\cap}} + \frac{N}{N_{\mathrm{D3}}} e^{4A_{\cap}} \left(\frac{r_{\cap}}{r_{\cup}}\right)^{2\Delta_s}.$$
(4.38)

We then find that the Coulomb energy is roughly

$$V_c \sim -\frac{2\pi}{\ell_s^4} \frac{N^2}{N_{\rm D3}} e^{4A_{\rm fl}}.$$
(4.39)

Requiring that this be much less than the potential energy barrier (4.32) yields the constraint

$$N \ll \frac{N_{\rm D3}}{\sqrt{g_s}} + \frac{2\operatorname{vol}\Sigma_{\cap}}{\ell_s^2} + \mathcal{O}\left(\frac{1}{N_{\rm D3}^2}\left(\frac{\operatorname{vol}\Sigma_{\cap}}{\ell_s^2}\right)^4\right)$$
(4.40)

We should also account for the interaction energy between the pair of fivebranes. By performing an open string computation in an unwarped toroidal orbifold, [30] found a potential contribution that grows *logarithmically* with the fivebrane separation, and argued that this would also apply to warped geometries, with energy scale set by r_{\cap} . While it is not entirely clear that this logarithmic behavior arises in the actual NS5-brane configuration described in section 3, the corresponding potential energy contribution would take the schematic form

$$V_{55} \sim \frac{2\pi}{\ell_s^4} e^{4A_{\cap}} N_{\rm NS5}^2 \log\left(\frac{L}{r_{\cap}}\right) \sim \frac{2\pi}{\ell_s^4} e^{4A_{\cap}} \,. \tag{4.41}$$

We can ensure that this energy is much smaller than the uncharged tension energy (4.32) by imposing

$$\operatorname{vol}\Sigma_{\cap} \gg \sqrt{g_s}\ell_s^2$$
, (4.42)

which is necessary in any case to ensure the validity of the supergravity approximation.

4.2.8 Antibrane tunneling and annihilation

Consider a Klebanov-Strassler throat that arises from N_{D3} D3-branes probing a conifold with M_{KS} D5-branes wrapping the shrinking two-cycle. We take $N_{D3} = M_{KS}K_{KS}$; then K_{KS} is the number of units of H_3 flux on the *B*-cycle.

If N anti-D3-branes are placed at the tip of this throat, they create a metastable, exponentially long-lived state provided that $N \leq 0.08 M_{\rm KS}$ [31]. With more anti-D3-branes, $N \geq 0.08 M_{\rm KS}$, the anti-D3-branes rapidly annihilate [31] against the flux supporting the warped throat, decaying to the state with $K'_{\rm KS} = K_{\rm KS} - 1$. The D3-brane charge carried by flux is then $N'_{\rm D3} = N_{\rm D3} - M_{\rm KS}$, and $M_{\rm KS} - N$ D3-branes appear, but no anti-D3-branes remain.

While N_{D3} sets the overall scale of warping in the throat, M_{KS} sets the warp factor at the tip,

$$e^A \big|_{\cup} \sim \exp\left(-\frac{2\pi N_{\rm D3}}{3g_s M_{\rm KS}^2}\right) \tag{4.43}$$

and

$$g_s M_{\rm KS}^2 \lesssim N_{\rm D3} \tag{4.44}$$

if the warping is non-negligible. To avoid the KPV instability [31], the number of windings cannot exceed

$$N \ll 0.08 \, g_s^{-1/2} N_{\rm D3}^{1/2}. \tag{4.45}$$

4.2.9 Light NS5-brane modes

As discussed in section 3.2, dimensional reduction of the transverse fluctuations in the NS5-brane's position yields Kaluza-Klein excitations whose mass *decreases* as the relaxion is wound up. Intuitively, we may interpret the presence of two-form flux as increasing the effective volume of the NS5-brane. Since Kaluza-Klein masses will inversely scale with this effective volume, we should expect some modes to become light at large windings. As shown in appendix C.2, to second order the canonically normalized fluctuations are described by the action

$$S_{\rm NS5}^{(2)} = \int d^4x \sqrt{-g_4} \left(-V(c) - \frac{1}{2} g^{\mu\nu} \partial_\mu Y_I^{\hat{\imath}} \partial^\mu Y_{\hat{\imath}}^I - \frac{1}{2} m_I^2(c) Y_I^{\hat{\imath}} Y_{\hat{\imath}}^I \right. \\ \left. + g(c) \left(c \partial_\mu c \right) \left(Y_I^{\hat{\imath}} \partial^\mu Y_{\hat{\imath}}^I \right) - \frac{1}{2} g(c)^2 \left(\partial c \right)^2 Y_I^{\hat{\imath}} Y_{\hat{\imath}}^I \right)$$
(4.46)

with

$$m_I^2 \sim \frac{4}{g_s \ell_{\rm E}^2} \frac{e^{2A_{\cup}}}{N^2} \left(\frac{\ell_{\rm E}}{\ell_s}\right)^4 \quad \text{and} \quad g(N) \sim \frac{1}{2N^2}$$
(4.47)

at large winding $N^2 \gg 4(\ell_{\rm E}/\ell_s)^4/g_s$, where $\ell_{\rm E}^2$ is the Einstein-frame volume of the twocycle Σ_2 and the λ_I are eigenvalues of Σ_2 's Laplacian, labeled by the multi-index I. As discussed in the introduction, the appearance of light states is generic in realizations of monodromy in string theory, and one must ensure that these do not drastically affect the phenomenology.

The presence of $\mathcal{O}(N)$ light states in the spectrum, including 3-7 strings and KK excitations of the NS5-branes, can have a range of consequences. For example, modes with mass m < 3H/2 can fluctuate during inflation, storing energy and potentially impacting the late-time perturbations. Here we will examine just one effect of the KK modes, which is an enhanced probability of NS5-brane annihilation.

The masses of the canonically normalized fluctuations of the NS5-brane embedding (4.46) decrease with N, and one might worry that these light modes facilitate an additional instability. For example, if these modes are thermally excited by some source of supersymmetry breaking elsewhere in the compact space, then the NS5-branes can more readily reach each other and either classically or quantum-mechanically annihilate. A complete analysis of this process is beyond the scope of this work. We will instead use an approximate criterion for the onset of instability. The dominant instanton in the fourdimensional field theory responsible for the transition between the metastable and stable states (i.e., the states with and without the NS5/anti-NS5-brane dipole, respectively) will be SO(4)-symmetric, with radius determined by

$$R_* \sim \frac{T_{\rm D}}{\Delta V},\tag{4.48}$$

where, in the thin-wall approximation, $T_{\rm D}$ is the tension of the domain-wall interpolating between the two vacua and ΔV is their difference in energy. We then assume a loss of control when the typical thermal fluctuations of a spatial region of size R_* are comparable to the distance between the two fivebranes, $r_{\rm RMS} \sim r_{\rm O}$.

The difference in energies, in the probe approximation, is the potential energy contribution from the NS5-branes,

$$\Delta V = \frac{2\pi}{\ell_s^4} N \, e^{4A_{\cup}} \tag{4.49}$$

while the tension of the domain wall is determined by an NS5-brane winding N times around the minimum-volume three-cycle whose endpoints are Σ_2 and $\bar{\Sigma}_2$. As described in appendix C.2, the tension of the domain wall follows from the NS5-brane action and is roughly

$$T_{\rm D} \sim \frac{1}{\ell_s^3} N_{\rm D3}^{3/4} \left(e^{4A_{\rm \cap}} - e^{4A_{\rm \cup}} \right) \tag{4.50}$$

and so

$$R_* \sim \ell_s \frac{N_{\rm D3}^{3/4}}{N} \left(\frac{r_{\rm o}}{r_{\rm o}}\right)^4 \tag{4.51}$$

If the NS5-brane is in thermal equilibrium at temperature T, then a smooth excitation of size R_* in the canonically normalized fluctuations $Y_I^{\hat{i}}$ gains a thermal expectation value

$$\langle Y_I^{\hat{i}} Y_J^{\hat{j}} \rangle \sim \frac{T}{m_I^2} \frac{1}{R_*^3} \delta_{IJ} \delta^{\hat{i}\hat{j}}.$$
(4.52)

The thermal fluctuation in the radial direction, averaged over Σ_2 , is roughly

$$\langle \delta r^2 \rangle \sim g_s \ell_s^2 \left(\ell_s T\right) \frac{N^4}{N_{\rm D3}^{3/4}} \left(\frac{\ell_s}{\ell}\right)^2 \left(\frac{r_{\cup}}{r_{\cap}}\right)^{12} \tag{4.53}$$

where ℓ^2 is the unwarped volume of the two-cycle Σ_2 . The requirement that this is much smaller than the size of the dipole $\langle \delta r^2 \rangle \ll r_{\cap}^2$ imposes the weak constraint

$$N^{4} \ll e^{4A_{\cap}} \frac{N_{\rm D3}^{5/4}}{g_{s}^{1/2}(\ell_{s}T)} \left(\frac{\ell_{\rm E}}{\ell_{s}}\right)^{2} \left(\frac{r_{\cap}}{r_{\cup}}\right)^{10}.$$
(4.54)

5 Discussion and outlook

We have identified many obstacles to realizing the relaxion mechanism in string theory. Some of these obstacles are extremely general, while others apply only to NS5-brane monodromy, the particular example we studied in detail. We will now step back and give some perspective on our results, explaining their scope of validity.

Our first observation was that axion monodromy in string theory proceeds by the accumulation of monodromy charge, and the backreaction of this charge substantially changes the couplings of the axion. This applies to any realization of axion monodromy in string theory. Thus, any ultraviolet completion in string theory of a relaxation mechanism that involves axion displacements $\Delta \phi > f$ will be vulnerable to the backreaction of monodromy charge.

The effects of this backreaction will vary from one model to another. We focused on NS5-brane monodromy because this is, to our knowledge, the only scenario where the smallness of the shift symmetry breaking parameter g is natural — in this case, because of warping — while in alternative constructions in string theory, one must fine-tune discrete data to achieve small g. In the NS5-brane model, we found that the barriers in the stopping potential are exponentially small in the winding number $N \equiv \phi/f$, leading to a runaway relaxion. We expect this barrier suppression phenomenon to be rather general, but not universal. However, the particular effects of backreaction on the axion decay constants detailed in section 4.2.6, and the constraints from annihilation in section 4.2.7 and section 4.2.8, could be very different in other models.

Some of the challenges that we have identified might be milder in non-supersymmetric compactifications of string theory. In particular, in compactifications that break all supersymmetry at the Kaluza-Klein scale as in e.g. [32, 33], tadpole constraints on the total charge need not be a serious limitation. On the other hand, ensuring metastability of such a configuration can be very challenging. Moreover, for an embedding of the relaxion in a non-supersymmetric compactification, the absence of spacetime supersymmetry below the KK scale might require either the KK scale or even the string scale to arise as the regulator of the relaxion setup at the relaxion cutoff scale M.

We assumed that the periodic stopping potential arises from non-perturbative effects that couple locally to the axion. This local coupling then exposes the stopping potential to an exponential suppression from warping. However, the stopping mechanism could instead arise from other effects, for example from heavy states [34] coupled to the relaxion, or from the exponential production of massive particles [2], which are not necessarily susceptible to the same failure modes.¹⁹

5.1 Exact discrete shift symmetries for relaxions

Throughout this work we have considered axion monodromy, in which a source of monodromy completely breaks the shift symmetry of an axion. An important alternative is *alignment* of multiple periodic contributions to the axion potential, leaving an unbroken

 $^{^{19}\}mathrm{We}$ thank E. Silverstein for illuminating discussions of these points.

discrete shift symmetry. We now briefly outline this possibility and mention some of the obstacles to embedding this scenario in string theory.

The essential feature of the relaxion potential (1.1) is the combination of a slowlyvarying term (\mathbb{B}) , the linear $gM^3\phi$, and a quickly-varying term (\mathbb{C}) , the oscillatory $\Lambda_c^3 v \cos(2\pi\phi/f)$. As written, (\mathbb{B}) explicitly and completely breaks the discrete shift symmetry $\phi \mapsto \phi + f$. Alternatively, (\mathbb{B}) could represent the leading term in the expansion of a function that is invariant under a much larger discrete shift $\phi \mapsto \phi + kf$. For example, we could have

$$V_{\widehat{\mathbb{B}}} = gfM^3 \sin\left(\frac{2\pi\phi}{kf}\right),\tag{5.1}$$

where ϕ 's field space diameter is actually k times larger than would be naively inferred by only considering small displacements. We refer to these two cases as having an *explicitly* broken symmetry or an exact discrete symmetry, respectively. Thus far, we have only concentrated on the former. The explicit breaking is induced by a source of monodromy, an NS5-brane, and we have shown that the accumulation of monodromy charge leads to backreaction effects that spoil the relaxation mechanism. Given this difficulty, one might ask whether the relaxion mechanism could be more readily realized in a solution of string theory with an exact discrete shift symmetry.

As in the models with explicit breaking, the main difficulty in realizing a discrete shiftsymmetric relaxion lies in ensuring that the potential has structure over two — and only two — disparate scales. That is, the potential must roughly be the sum of two terms — a slowly varying term with periodicity f that apes the linear term $\widehat{\mathbb{B}}$ in (1.1), and a quickly varying stopping potential with periodicity $f_s \ll f$. One might take, as a toy model, a potential that is generated only by the instantons with winding number 1 and k, so that the potential takes the schematic form

$$V = M^2 e^{-S_1} \cos\left(\frac{2\pi\phi}{f}\right) |h|^2 + M^4 e^{-S_1} \cos\left(\frac{2\pi\phi}{f}\right) + M^3 e^{-S_k} v \cos\left(\frac{2\pi k\phi}{f}\right)$$
(5.2)

For $\phi \ll f$, the potential is approximately a "monomial with modulations" with $f_s = f/k$, and has the $\widehat{(A)} \otimes \widehat{(C)}$ structure of (1.1), with $g \propto e^{-S_1}$ and $\Lambda_c^3 \sim M^3 e^{-S_k}$. The analogue of (1.4) in this two period model is then

$$\frac{v}{M} \gtrsim \frac{1}{k} \left(\frac{M}{\Lambda_c}\right)^3 \sim \frac{1}{k} e^{S_k - S_1}.$$
(5.3)

Naively, the generated hierarchy grows with k.

However, there are many problems with this toy model. First and foremost, the action for a k-instanton is typically $S_k \ge kS_1$, and we require that $S_1 \gg 1$ in order to trust the instanton expansion. The stopping potential barriers will then shrink with k,

$$|V_{\rm stop}| \propto e^{-kS_1} \propto g^k \tag{5.4}$$

and, reminiscent of the suppression due to warping discussed in section 4.2.3, the maximum achievable hierarchy actually shrinks with k. The stopping potential is too small to stop the evolution near the point where the Higgs is massless. If some mechanism were able to enhance the k-instanton contribution, one must still explain the absence of j-instanton effects, with 1 < j < k, and we find it implausible that all such effects could be negligible for $k \gg 1.^{20}$ Furthermore, one would have to explain why the Higgs couples to instantons of winding 1 and k differently, and why the 1-instanton and k-instanton contributions do not both vanish when v = 0.

Many of these problems may be mitigated in models with multiple axions, as in the Kim-Nilles-Peloso mechanism [35] and kinetic alignment setups [36, 37] in the inflationary context. Scenarios for aligned relaxions have been presented in [19, 20, 38].²¹ A general multi-axion Lagrangian can be written (cf. e.g. [37]) as

$$\mathcal{L} = -K^{ij}\partial\phi_i\partial\phi_j - \sum_a \Lambda_a^4 \exp\left(-Q_a^{\ i}S_i\right) \left[1 - \cos\left(2\pi Q_a^{\ i}\phi_i\right)\right],\tag{5.5}$$

where K^{ij} is a positive definite kinetic matrix of real numbers, while $Q_a{}^i$ is a charge matrix containing integers. We imagine that there are "slow" and "fast" linear combinations of canonically normalized axions with effective decay constants f and $f_s \ll f$, respectively. The addition of another direction in field space solves several of the problems mentioned previously, at the cost of introducing much more complicated dynamics.

Foremost among the advantages is that the stopping potential is no longer necessarily suppressed. In the single axion model, the hierarchy between f_s and f — and so between v and M — was generated by a hierarchy in the charge matrix $Q_a^{\ i}$, and a high charge contribution is exponentially suppressed relative to a low charge contribution. In a multiaxion model, $f_s \ll f$ may instead be realized in the kinetic matrix K^{ij} , and this does not impose an exponential hierarchy in the associated barrier heights. Of course, the hierarchy in the kinetic matrix must then be explained, but it is much easier to realize a hierarchy in the eigenvalues of a matrix of real numbers than in a matrix of bounded integers, and one does not need to explain why instantons with winding j, 1 < j < k, do not contribute.

A very mild degree of alignment has been demonstrated in explicit examples [40], but whether alignment can yield large effective axion decay constants in string theory is an important open question, even for the $\mathcal{O}(100)$ enhancements that could suffice for inflation. It is not obvious to us that the vastly larger enhancements needed for a relaxion scenario are possible in known compactifications. For example, the "clockwork" mechanism [38, 41] requires a specific matrix of axion charges of instantons, and it remains to be seen whether this particular pattern of charges can arise in string theory. However, it is very plausible that linearly independent combinations of axions couple differently to the Higgs.

In summary, relaxion scenarios with exact discrete symmetries, built on the alignment of multiple instanton effects for one or more axions, are qualitatively different from the axion monodromy scenarios, with explicitly broken symmetry, considered in this work. However, *both* classes of models are vulnerable to ultraviolet physics. Axion monodromy scenarios suffer from the backreaction of monodromy charge, as we have explained. Aligned scenarios could avoid this problem, but require extremely special axion charges. These

²⁰Such a situation appears to conflict with the lattice form of the Weak Gravity Conjecture, but is already implausible regardless.

²¹See also [39] for a recent discussion of naturalness constraints on such scenarios.

charges are ultimately topological data dictated by the ultraviolet theory, and it is not clear that string theory allows strong enough alignment to permit relaxation of a large hierarchy. Furthermore, these multi-axion models have much more complicated dynamics, and it is not clear that the dynamical generation of a large hierarchy can proceed in a robust way.

5.2 Constraints from the Weak Gravity Conjecture

The Weak Gravity Conjecture (WGC), a class of conjectures asserting that gravity must be the weakest force [42–52], leads to (still conjectural) constraints on axion theories. One could therefore ask whether the WGC constraints relaxion monodromy scenarios. It does [8, 9], as we will briefly explain, but the known WGC constraints are far weaker than the limitations we have exposed in this work, which are independent of WGC considerations.

The WGC constrains monodromy scenarios by placing upper limits on the tension of domain walls. In the four-dimensional description of axion monodromy due to Kaloper, Lawrence, and Sorbo [11–13], Brown-Teitelboim domain walls connect different branches of the scalar potential. At the same time, when instanton effects lead to modulations of the axion potential, distinct critical points are connected across four-dimensional field theory domain walls, via Coleman-de Luccia tunneling. It turns out that the electric form of the WGC places constraints [8] on the domain walls of the Kaloper-Lawrence-Sorbo model, while the magnetic WGC places constraints on the field theory domain walls associated with instanton modulations [9]. In both cases one finds a bound on the domain wall tension [8, 9]

$$T < m f M_{\rm pl},\tag{5.6}$$

where m is the mass of the axion. For a relaxion model this implies a bound on the relaxion cutoff scale M of roughly the same order as the constraints already given in [1].

We conclude that the constraints arising from very general four-dimensional quantum gravity considerations, such as the WGC, do not automatically capture all of the effects of actual embeddings in quantum gravity. Examining such embeddings is therefore crucial for assessing the viability of the relaxion mechanism in string theory.²²

6 Conclusions

Could a portion of the observed hierarchy between the weak scale and the Planck scale be a consequence of dynamical relaxation of the Higgs mass during cosmological evolution? This striking idea is the core of the relaxion mechanism [1]. In this scenario, the relaxation of the Higgs mass is driven by the slow evolution of an axion field, the relaxion, whose shift symmetry is very weakly broken by a potential term that introduces monodromy. After relaxation over many cycles of monodromy, the Higgs mass passes through zero, causing

 $^{^{22}}$ We note that this view concurs with the results of [11], where some leading effects of backreaction on the axion Lagrangian are captured by a series of higher powers of gauge-invariant field strengths, whose coefficients must necessarily be determined in the UV theory, and in which strong backreaction effects can drive one far from the "natural" bottom-up estimates.

barriers to appear in the axion potential, and so halt the evolution. In effective field theory, the hierarchy that is generated is determined by the weak breaking parameter, and so is technically natural.

In this work we asked whether the relaxion mechanism survives ultraviolet completion in string theory. Do the essential components for the scenario exist in a well-controlled compactification, and do these components work in concert in string theory as they do in effective field theory?

We found that the key components of the scenario can indeed be realized in string theory. The mechanism of axion monodromy, first developed in the context of large-field inflation in string theory, can produce — in the probe approximation — the secular relaxion potential needed for slow relaxation over many fundamental axion periods. Moreover, the extremely low scale of the secular potential required for the relaxion mechanism can be explained by situating the source of monodromy in a strongly warped region. This is possible in one known scenario for axion monodromy in string theory, the NS5-brane model of [21], in which two-form axions acquire their potential from NS5-branes wrapping curves in a warped region.

However, our main result is that the structures required for monodromy in string theory present formidable and very general obstacles to a successful relaxion scenario in string theory. Monodromy proceeds by the accumulation of monodromy charge on a source of monodromy. As the relaxion rolls over N fundamental axion periods, it necessarily accumulates or discharges N units of monodromy charge. This large quantity of monodromy charge sources backreaction in the internal space, completely invalidating the probe approximation, and changing the couplings in the effective theory. The impact of monodromy charge is visible in a dual description as the appearance of N new light states.

We argued that the backreaction of monodromy charge can lead to disastrously large changes to the secular potential in *any* realization of the relaxion scenario via axion monodromy in string theory. In the specific case of the NS5-brane model, we computed the detailed form of these changes. The accumulation of monodromy charge suppresses the gauge coupling of the D7-brane gauge theory that generates the stopping potential. In the dual description, the N light states are charged under the D7-brane gauge group, and give a large threshold correction to the gauge coupling. The result is that the stopping potential is suppressed by a factor $\exp(-\gamma_{\rm br}N)$, where $\gamma_{\rm br}$ is a constant determined by the geometry. The stopping potential is therefore completely negligible, and cannot halt the evolution when the Higgs mass passes through zero. The Higgs mass indeed relaxes to smaller values, but this process continues far into the tachyonic regime.

While this detailed analysis was performed in the context of a specific model, we repeat that our findings are generic and are expected to apply to any model that relies upon a monodromy over many fundamental axion periods, regardless of the stopping mechanism. However, these constraints do not apply when the discrete shift symmetry remains unbroken, e.g. when the large field range is realized by the alignment of multiple axions.

In summary, we have shown that the physics of ultraviolet completion in string theory does not decouple from the dynamics of the relaxion mechanism. Our results do not exclude the dynamical relaxation of hierarchies in string theory, but in our view they do exclude technically natural dynamical relaxation driven by axion monodromy. It would be valuable to understand whether some of the difficulties we have uncovered result from limitations in existing constructions, or if instead they are consequences of general structures in quantum gravity.

Acknowledgments

We thank Nima Afkhami-Jeddi, Tom Hartman, Nemanja Kaloper, David E. Kaplan, Eric Kuflik, Cody Long, Miguel Montero, Surjeet Rajendran, Michael Stillman, Amir Tajdini, Irene Valenzuela, and Timo Weigand for valuable discussions. The work of L.M. is supported in part by an NSF CAREER Award. The work of G.S. is supported in part by the Helmholtz Association. The work of A.W. is supported by the ERC Consolidator Grant STRINGFLATION under the HORIZON 2020 contract no. 647995.

A Axions in string theory

There is an extensive literature on axions in string theory, but for the reader's convenience we now gather a few salient facts. We begin with the example of the Neveu-Schwarz two-form gauge potential B_2 .

Integrating a ten-dimensional *p*-form gauge potential C_p over a non-trivial *p*-cycle in the compactification manifold will give rise to an axion in four dimensions. The number of independent, non-trivial *p*-cycles then determines the maximal number of axions arising from C_p . The two-form B_2 has an associated field strength $H_3 \equiv dB_2$, and appears in the ten-dimensional type II and heterotic supergravity actions as²³

$$S_{\rm SUGRA} \supset -\frac{1}{4\kappa_{10}^2} \int d^{10} X \sqrt{-G^S} \, e^{-2\Phi} \, |H_3|^2 \tag{A.1}$$

Reducing this action along a six-dimensional compact space X_6 , each non-trivial two-cycle Σ_2^I with its associated harmonic form ω_2^I , $\ell_s^{-2} \int_{\Sigma_2^I} \omega_2^J = \delta_I^J$, gives rise to a four-dimensional axion $b_I(x)$,

$$b_I(x) \equiv \frac{1}{\ell_s^2} \int_{\Sigma_2^I} B_2, \tag{A.2}$$

with $B_2 = \sum_I b_I(x)\omega_2^I$. Upon dimensional reduction, the first term of (A.1) yields kinetic terms for the b_I axions,

$$S_{\rm kin} = -\frac{1}{2} \int \mathrm{d}^4 x \, \sqrt{-g} \, \gamma^{IJ} \partial_\mu b_I \partial^\mu b_J \qquad \qquad \frac{\gamma^{IJ}}{M_{\rm pl}^2} = \frac{g_s}{2\mathcal{V}_{\rm E}\ell_s^6} \int_{X_6} *_6 \omega_2^I \wedge \omega_2^J. \tag{A.3}$$

If a basis of harmonic forms ω_2^I is chosen such that γ^{IJ} is diagonal, then $\phi_I = f_I b_I$ (no sum) are the canonically normalized axion fields, whose decay constants are the eigenvalues of γ , $f_I = \operatorname{eig}_I \gamma^{JK}$. For example, if the compact space is a product of two-spheres, $X_6 = \mathrm{S}^2 \times \mathrm{S}^2 \times \mathrm{S}^2$, each with volume $L^2 \ell_s^2$, then we simply have $f_I^2 = M_{\rm pl}^2 L^{-4}/2$.

 $^{^{23}\}mathrm{Normalization}$ conventions appear in appendix C.

If X_6 is Calabi-Yau, then the axion decay constants for two-form axions b_I and c_I arising from the two-form potentials B_2 and C_2 , respectively, in type IIB string theory may be simply computed from the intersection numbers κ_{IJK} , the volumes $\ell_s^2 v^I$ of the twocycles Σ_2^I , and the overall total volume \mathcal{V}_E of X_6 . For example, for an axion $c_- = \ell_s^{-2} \int_{\Sigma_2^-} C_2$ associated to an orientifold-odd cycle Σ_2^- the axion decay constant is

$$\frac{f^2}{M_{\rm pl}^2} \sim g_s \frac{\kappa_{I--} v^I}{\mathcal{V}_{\rm E}}.\tag{A.4}$$

When fivebranes are introduced to create monodromy, the axion that experiences this monodromy will in general be a linear combination of the ω_2^I , which we call Ω . For example, in a variant of the axion monodromy construction detailed in section 3, the (rel)axion c(x)arises from a two-form Ω dual to the blowup cycle of an orbifold whose fixed point locus is Σ_0 , with dim_C $\Sigma_0 = 1$. As shown in [23], the support of $*\Omega \wedge \Omega$ is localized about Σ_0 . The six-dimensional metric is approximately a cone,

$$\tilde{g}_{mn} \mathrm{d} y^m \mathrm{d} y^n \approx \mathrm{d} r^2 + r^2 \check{g}_{\theta\phi} \mathrm{d} \Psi^\theta \mathrm{d} \Psi^\phi$$
 (A.5)

and $\Omega_{mn} \sim \Omega_{\theta\phi}$ to good approximation has its legs along the angular directions, so

$$\frac{f^2}{M_{\rm pl}^2} \sim \frac{g_s}{\mathcal{V}_{\rm E}} \frac{1}{\ell_s^6} \int_{r_{\cup}}^{r_{\cap}} \mathrm{d}r \, r \int \mathrm{d}^5 \Psi \sqrt{\check{g}} \, \check{g}^{\theta \phi} \check{g}^{\psi \chi} \Omega_{\theta \psi} \Omega_{\phi \chi} \approx \frac{g_s}{\mathcal{V}_{\rm E}} \frac{r_{\cap}^2}{\ell_s^2}. \tag{A.6}$$

Locally, we may think of the blow-up cycle as an Eguchi-Hanson space fibered over Σ_{0} . Since the integrand is highly localized about Σ_{0} , we have $\int *\Omega \wedge \Omega \approx \operatorname{vol}(\Sigma_{0})$, and because of the conical nature of the six-dimensional metric, $\operatorname{vol}(\Sigma_{0})$ is dominated by the contribution at r_{\cap} , so $\int *_{6} \Omega \wedge \Omega \propto r_{\cap}^{2}$.

The axion enjoys a continuous shift symmetry to all orders in perturbation theory in both g_s and α' . However, this continuous shift symmetry does not survive at the nonperturbative level, and is broken to a discrete shift symmetry by instantons carrying axion charge. In particular, fundamental strings are charged under B_2 , via the coupling

$$S(\mathcal{W}) = \ldots + \frac{i}{2\pi\alpha'} \int_{\mathcal{W}} d^2\sigma \sqrt{-h} \,\epsilon^{mn} B_{mn} + \ldots, \qquad (A.7)$$

where h is the metric on the string worldsheet \mathcal{W} , and m, n are two-dimensional indices tangent to \mathcal{W} . The string path integral receives a contribution from a Euclidean string whose worldsheet wraps Σ_2^I , termed a worldsheet instanton. This contribution will be proportional to e^{-S_I} , where $S_I = S(\Sigma_2^I)$. Because of the coupling (A.7),

$$S_I \supset 2\pi i b_I,$$
 (A.8)

and so the potential generated by these nonperturbative effects is still invariant under discrete shifts $b_I \mapsto b_I + N$, $N \in \mathbb{Z}$, as $e^{-S_I} \mapsto e^{-S_I + 2\pi i N} = e^{-S_I}$. Thus, worldsheet instantons break the perturbative, continuous shift symmetry of b_I to the discrete shift $b_I \mapsto b_I + 1$. The real part of the action (A.7) is proportional to the volume of Σ_2^I in string units, and worldsheet instanton contributions become more important as the volume shrinks. These contributions are difficult to compute, so requiring computational control of the effective action constrains the sizes of cycles, Σ_2^I , and thus the sizes of the axion decay constants. A standard requirement for control is that the sizes of all cycles are much larger than the string length, $v^{\alpha} \gg 1$.

However, the two-cycle Σ_2^I may sit in a warped region, with warp factor e^A . For twoform axions, (A.3) is unchanged — there is no explicit dependence on the warp factor. However, a ten-dimensional string will see a warped volume, and in particular the real part of the worldsheet instanton action is enhanced by factor e^{-2A} . This allows the two-cycle volumes v^{α} to be smaller by a factor of e^{-2A} without loss of control, and so the axion decay constant can be very small in a highly warped throat.

If we take v^{I} to measure the warped volume of Σ_{2}^{I} in string units, i.e. the volume a ten-dimensional string would measure, then we may write

$$\frac{f^2}{M_{\rm pl}^2} \sim g_s \frac{\kappa_{I--} v^I}{\mathcal{V}_{\rm E}} \left. e^{2A} \right|_{\Sigma_2^I} \ll g_s \frac{\kappa_{I--} v^I}{\mathcal{V}_{\rm E}},\tag{A.9}$$

keeping the constraint that $v^I \gg 1$.

B Necessity of the intersection

In NS5-brane axion monodromy, D3-brane charge accumulates on an NS5-brane that wraps a two-cycle $\Sigma_{\rm NS5}$ (denoted Σ_2 elsewhere in the text). Taking $c(x) \equiv \int_{\Sigma_{\rm NS5}} C_2$ to be the relaxion field, a stopping potential can be generated by strong gauge dynamics in a gauge theory G to which the relaxion has a nonvanishing axionic coupling $\lambda c(x)F \wedge F$, with λ a constant. We will take G to be realized on a stack of D7-branes wrapping a holomorphic four-cycle D (denoted Σ_4 elsewhere in the text). The backreaction of the D3-brane charge changes the supergravity background, with the strongest effects occurring near $\Sigma_{\rm NS5}$. In this appendix we show that any D for which $\lambda \neq 0$ necessarily intersects a holomorphic representative of the class [$\Sigma_{\rm NS5}$]. Thus, in the simplest setting with only two such representatives — each wrapped by a fivebrane — one cannot mitigate the backreaction by arranging that D is outside of the warped region. In models with more than two holomorphic representatives, avoiding this backreaction still entails an $\mathcal{O}(N^{-1})$ fine-tuning.

For our purposes, it suffices to show that D and $\Sigma_{\rm NS5}$ have at least one point in common, even though the intersection number $[D] \cap [\Sigma_{\rm NS5}]$ of the corresponding homology classes may vanish. We will use \cap_s to denote intersection as point sets, as distinct from the topological intersection $[\Sigma_1] \cap [\Sigma_2]$,²⁴ and we will show that $\lambda \neq 0$ implies that $D \cap_s \Sigma_{\rm NS5}$.

Consider a D7-brane that fills spacetime and wraps a smooth four-cycle $D \subset X$ in the internal space X. The D7-brane couples to C_2 axions via the Chern-Simons action

$$S_{\rm CS} = \mu_7 \int_{\mathcal{W}} \sum_p \iota^* C_p \wedge e^{\mathcal{F}} \supset \frac{\mu_7}{3!} \int_{\mathcal{W}} \iota^* C_2 \wedge \mathcal{F} \wedge \mathcal{F} \wedge \mathcal{F}, \tag{B.1}$$

²⁴Two submanifolds M, N, of X have $M \cap_s N \neq \emptyset$ if and only if M and N have at least one point in common, without regard to the orientation of M and N.

where $\mathcal{W} = \mathcal{M}^{3,1} \times D$ is the D7-brane worldvolume, $\iota : D \to X$ is the inclusion map of D into X, ι^* denotes the pullback onto D, F_2 is the field strength of the worldvolume gauge theory, and $\mathcal{F} = \iota^* B_2 + 2\pi \alpha' F_2$. The axionic coupling to the gauge theory G on a stack of D7-branes wrapping D is therefore

$$S_{\rm CS} \supset \frac{\mu_7}{2!} \left(\int_D \iota^* C_2 \wedge \mathcal{F} \right) \left(\int_{\mathcal{M}^{3,1}} \operatorname{tr} \mathcal{F} \wedge \mathcal{F} \right) \,. \tag{B.2}$$

Poincaré duality in D relates the flux \mathcal{F} to a two-cycle $S_{\mathcal{F}} \subset D$ which may further be viewed as a two-cycle $\iota_*S_{\mathcal{F}}$ in X, so that

$$\vartheta \equiv \int_D \iota^* C_2 \wedge \mathcal{F} = \int_{S_\mathcal{F}} \iota^* C_2 = \int_{\iota_* S_\mathcal{F}} C_2. \tag{B.3}$$

The holomorphic representative $\iota_*S_{\mathcal{F}}$ of the class $[\iota_*S_{\mathcal{F}}]$ is contained, as a point set, in D. So establishing the condition

$$\iota_* S_{\mathcal{F}} \cap_s \Sigma_{\mathrm{NS5}} \neq \emptyset \tag{B.4}$$

will imply our desired result $D \cap_s \Sigma_{NS5}$.

Now we choose a basis of nontrivial two-cycle classes, $\{[\Sigma_i]\}, I = 1, \ldots, p \equiv h^{1,1}$, to span $H_2(X, \mathbb{Z})$. Without loss of generality, we may take the NS5-brane class $[\Sigma_{NS5}]$ to be an element of this basis, say $[\Sigma_1] \equiv [\Sigma_{NS5}]$. There exists a dual basis of harmonic twoforms ω^J such that $\int_{\Sigma_I} \omega^J = \delta_I^J$. Expanding $\ell_s^{-2}C_2(x) = \sum_{i=1}^p c_I(x)\omega^I$, the relaxion field is $c_1(x) \equiv c(x)$. We may also expand

$$\iota_* S_{\mathcal{F}} = a_1[\Sigma_1] + \dots + a_p[\Sigma_p] + (\text{boundary}), \qquad (B.5)$$

for some integers a_I . Comparing to (B.3), we see that $\lambda \neq 0$ if and only if $a_1 \neq 0$.

The relation (B.5) with $a_1 \neq 0$ does not, on its own, imply (B.4). For example, consider a basis of homology $\{[\Sigma_1], [\Sigma_2]\}$, with minimum volume representatives $\{\Sigma_1, \Sigma_2\}$ that obey $\Sigma_I \cap_s \Sigma_J = \delta_{IJ}$. If $[S] = a_1[\Sigma_1] + a_2[\Sigma_2]$, then $S \cap_s \Sigma_1 \neq \emptyset \iff a_1 \neq 0$, regardless of the value of a_2 . But working in the basis $\{[\Sigma_1], [\Sigma'_2] = [\Sigma_2] - [\Sigma_1]\}$, for $a_1 \neq 0$ and $a_2 = 0$ we again have $S \cap_s \Sigma_1 \neq \emptyset$, while if $a_1 = a_2$ we have instead $S \cap_s \Sigma_1 = \emptyset$. Thus, the condition (B.4) depends on the relation between $[\Sigma_1]$ and $[\Sigma_2], \ldots [\Sigma_p]$, which we have not yet specified.

We may view this issue in a dual picture. The coupling (B.3) can be written as the triple intersection of three divisors in X,

$$\vartheta = [D] \cap [D_{\mathcal{F}}] \cap [D_{\text{NS5}}], \qquad (B.6)$$

where $D_{\mathcal{F}} = \text{PD}_X(\iota_*\mathcal{F})$ and $D_{\text{NS5}} = \text{PD}_X(\omega^1)$, with PD_X denoting the Poincaré dual in X. The divisor D_{NS5} is dual to the curve Σ_{NS5} , in that D_{NS5} is Poincaré dual to the two-form ω^1 that is the dual vector to Σ_{NS5} with respect to the pairing $\int_{\Sigma_I} \omega^J = \delta_I^J$. It follows that $[\Sigma_{\text{NS5}}] \cap [D_{\text{NS5}}] = 1$. Moreover, the requirement of a nonvanishing axionic coupling, $\vartheta \neq 0$, implies that $[D_{\text{NS5}}] \cap [D] \neq 0$. Now although $[\Sigma_{\text{NS5}}] \cap [D_{\text{NS5}}] \neq 0$ and $[D_{\text{NS5}}] \cap [D] \neq 0$, it appears that D_{NS5} could stretch between D and Σ_{NS5} , intersecting each, even though D and Σ_{NS5} remain widely separated. To exclude this possibility, we use further facts about $\Sigma_{\rm NS5}$ and $\iota_*S_{\mathcal{F}}$. Preserving supersymmetry in the D7-brane worldvolume requires that $\mathcal{F} \in H^2(D, \mathbb{Z})$ be of type (1, 1), and so its Poincaré dual $\iota_*S_{\mathcal{F}}$ is a holomorphic curve. Heuristically, $\iota_*S_{\mathcal{F}}$ can be viewed as the curve wrapped by a D5-brane dissolved in D: if the D7-brane were annihilated by introducing an anti-D7-brane, a D5-brane on $\iota_*S_{\mathcal{F}}$ would remain. Moreover, $\Sigma_{\rm NS5}$ is itself an irreducible holomorphic curve. (In a construction in which $\Sigma_{\rm NS5}$ is a sum of irreducible holomorphic curves, this argument can be applied to each component.) We can therefore express $\iota_*S_{\mathcal{F}}$ uniquely as a finite sum of distinct irreducible holomorphic curves $\{\sigma_A\}$, $A = 1, \ldots K$

$$\iota_* S_{\mathcal{F}} = a_1 \sigma_1 + a_2 \sigma_2 + \dots + a_K \sigma_K, \qquad a_A \in \mathbb{Z} \ge 0, \tag{B.7}$$

and we have shown above that a relaxionic coupling, $\lambda \neq 0$, requires $a_1 \neq 0$. Because the σ_A are distinct irreducible holomorphic curves, they intersect each other at most at points.²⁵ So ι_*S_F contains all but finitely many points of σ_1 , and thus the divisor D must intersect a holomorphic representative of $[\Sigma_{\rm NS5}]$. In the simplest model, this requires that D must descend into the strongly warped region to intersect (cf. figure 3) Σ_2 or $\bar{\Sigma}_2$.

It is possible for the curve class $[\Sigma_{\rm NS5}]$ to have more than two holomorphic representatives, which suggests a loophole that avoids the problematic backreaction discussed in the body of this work: design a curve class with a holomorphic representative far from the warped throat, so that the divisor D may lie outside the warped region yet still generate an axionic coupling. However, as we discuss at the end of section 4.2.5, a distant divisor is still vulnerable to this backreaction, which again can only be subdued by a significant $\mathcal{O}(N^{-1})$ fine-tuning.

C Type IIB supergravity with fivebranes

C.1 Conventions for type IIB supergravity

The bosonic part of the type IIB supergravity action in Einstein frame is

$$S_{\rm IIB} = \frac{1}{2\kappa_{10}^2} \int d^{10}X \,\sqrt{-G_{\rm E}} \left(R_{\rm E} - \frac{|\partial\tau|^2}{2({\rm Im}\,\tau)^2} - \frac{|G_3|^2}{2\,{\rm Im}\,\tau} - \frac{1}{4}|\tilde{F}_5|^2 \right) - \frac{i}{8\kappa_{10}^2} \int \frac{C_4 \wedge G_3 \wedge \bar{G}_3}{{\rm Im}\,\tau} \frac{1}{({\rm C}.1)} d^{10}X \,\sqrt{-G_{\rm E}} \left(\frac{1}{2} \left(\frac{1}{2} + \frac{1}{2} \right) - \frac{1}{2} \left(\frac{$$

with $2\kappa_{10}^2 = \ell_s^8/2\pi$, $G_3 \equiv F_3 - \tau H_3$, $\tau \equiv C_0 + ie^{-\Phi}$, $F_{p+1} = dC_p$, $H_3 = dB_2$, $\tilde{F}_5 = F_5 - \frac{1}{2}C_2 \wedge H_3 + \frac{1}{2}B_2 \wedge F_3$, and $\tilde{F}_5 = *_{10}\tilde{F}_5$ is imposed at the level of the equations of motion.

We define the string length

$$\ell_s^2 = (2\pi)^2 \alpha'.$$
 (C.2)

The actions for extremal Dp-branes and NS5-branes are given by

$$S_{\rm Dp} = -\mu_p \int d^{p+1} \xi \, e^{-\Phi} \sqrt{-\det \left(G_{ab} + B_{ab} + 2\pi \alpha' F_{ab}\right)} + S_{\rm CS} \tag{C.3}$$

²⁵Note that if the σ_A were simply a set of distinct, irreducible simplicial complexes, the relation (B.4) would not be automatic. If σ_1 intersected some of $\sigma_2, \ldots \sigma_K$ along suitable two-simplices, then $\sum_i a_A \sigma_A$ might have no points in common with σ_1 , because adding $a_2\sigma_2 + \ldots a_K\sigma_K$ could subtract all the points of σ_1 . For curves intersecting at most at points, this is not possible.

Directions	Indices
(3+1)-dim spacetime	μ, u, ho, \dots
6-dim internal space	m, n, \ldots
5-dim angular space	$ heta, \phi, \dots$
brane worldvolume	a, b, \ldots
transverse to worldvolume	i, j, k, \ldots
transverse vielbein	$\hat{\imath},\hat{\jmath},\hat{k},\ldots$
along cycle Σ_2	$lpha,eta,\ldots$

Table 3. Guide to index conventions.

and

$$S_{\rm NS5} = -\mu_5 \int d^6 \xi \, e^{-2\Phi} \sqrt{-\det\left(G_{ab} - e^{\Phi} \left(C_{ab} + 2\pi \alpha' F_{ab}\right)\right)} + S_{\rm CS},\tag{C.4}$$

respectively, where $\mu_p = 2\pi/\ell_s^{p+1}$, and F_{ab} is the gauge field strength on the brane worldvolume. The Chern-Simons term for a D*p*-brane reads

$$S_{\rm CS} = \mu_p \int \sum_n C_n \wedge e^{\mathcal{F}},\tag{C.5}$$

where we have introduced the notation $\mathcal{F} = B_2 + 2\pi \alpha' F$. The Chern-Simons piece sets the flux quantization condition

$$\frac{1}{\ell_s^{p+1}} \int_{\Sigma^p} F_{p+1} \in \mathbb{Z}.$$
(C.6)

We expand in a basis of $H^2(X_6)$, denoted $\omega^I(y)$, with normalization

$$\int_{\Sigma^J} \omega^I = \ell_s^2 \delta_J^I. \tag{C.7}$$

We list our index conventions in table 3.

C.2 Einstein-frame potentials for fivebranes

The DBI action for a D5-brane is

$$S_{\text{DBI}} = -\frac{2\pi}{\ell_s^6} \int d^6 \xi \, e^{-\Phi} \sqrt{-\det\left(G_{ab}^S + \mathcal{F}_{ab}\right)} \tag{C.8}$$

where G_{ab}^{S} and B_{ab} are the pull-backs onto the brane worldvolume of the ten-dimensional string-frame metric and NS-NS two-form B_2 , respectively.

The ten-dimensional string-frame metric is related to the Einstein-frame metric via

$$G_{MN}^{S} = e^{\Phi/2} G_{MN}^{E},$$
 (C.9)

which we assume takes a warped product form

$$G_{MN}^E \,\mathrm{d}X^M \,\mathrm{d}X^N = e^{2A(y)} g_{\mu\nu} \,\mathrm{d}x^\mu \,\mathrm{d}x^\nu + e^{-2A(y)} \tilde{g}_{mn} \,\mathrm{d}y^m \,\mathrm{d}y^n, \tag{C.10}$$

where $g_{\mu\nu}$ and \tilde{g}_{mn} are metrics on the four-dimensional spacetime and the six-dimensional internal space, respectively.

Defining

$$b(x) = \frac{1}{\ell_s^2} \int_{\Sigma_2} B_2 = \frac{1}{\ell_s^2} \int_{\Sigma_2} \ell_s^2 b(x) \, \mathrm{d}y \wedge \mathrm{d}z, \tag{C.11}$$

choosing coordinates on Σ_2 such that $dy \wedge dz$ is harmonic, and setting $F_{ab} = 0$, we may write

$$e^{\Phi/2}G_{ab}^E + B_{ab} = \begin{pmatrix} e^{\Phi/2}e^{2A(y)}g_{\mu\nu} & 0\\ 0 & \mathbf{m} \end{pmatrix}$$
(C.12)

with

$$\mathbf{m} = \begin{pmatrix} e^{\Phi/2} e^{-2A} \tilde{g}_{yy} & e^{\Phi/2} e^{-2A} \tilde{g}_{yz} + \ell_s^2 b/2 \\ e^{\Phi/2} e^{-2A} \tilde{g}_{yz} - \ell_s^2 b/2 & e^{\Phi/2} e^{-2A} \tilde{g}_{zz} \end{pmatrix}.$$
 (C.13)

With $\tilde{g}_2 \equiv \tilde{g}_{yy}\tilde{g}_{zz} - \tilde{g}_{yz}^2$, we have

$$\sqrt{-\det(G_{ab} + B_{ab})} = e^{\Phi + 4A} \sqrt{-\det g} \sqrt{e^{\Phi} e^{-4A} \tilde{g}_2 + \ell_s^4 b^2/4}.$$
 (C.14)

Upon integration over Σ_2 , we may take $e^{-4A}\tilde{g}_2 \rightarrow \ell_{\rm E}^4$, where $\ell_{\rm E}^2$ is the characteristic size of Σ_2 in the ten-dimensional *Einstein frame*. The DBI action, upon dimensional reduction, then yields a four-dimensional potential for the *b* axion,

$$S_{\text{DBI}} = \int \mathrm{d}^4 x \sqrt{-g} \left(-\frac{\pi e^{4A}}{\ell_s^4} \sqrt{\frac{g_s}{4} \left(\frac{\ell_{\text{E}}}{\ell_s}\right)^4 + b^2} \right). \tag{C.15}$$

The dimensional reduction of the NS5-brane action follows similarly.

We will also be interested in the spectrum of Kaluza-Klein excitations, which we now compute. Setting $F_{ab} = 0$, the NS5-brane action is

$$S_{\rm NS5} = -\frac{2\pi}{\ell_s^6} \int d^6 \xi \, e^{-2\Phi} \sqrt{-\det \mathbf{M}} \tag{C.16}$$

with $\mathbf{M}_{ab} = G_{ab} - e^{\Phi} C_{ab}$. Expanding in fluctuations $\delta \mathbf{M}_{ab}$, we have

$$\sqrt{-\det\left(\overline{\mathbf{M}} + \delta\mathbf{M}\right)} = \sqrt{-\det\overline{\mathbf{M}}} \left(1 + \frac{1}{2}\operatorname{tr}\overline{\mathbf{M}}^{-1}\delta\mathbf{M}\right), \qquad (C.17)$$

with

$$\overline{\mathbf{M}} = \begin{pmatrix} e^{\Phi/2} e^{2A} g_{\mu\nu} & 0\\ 0 & \mathbf{m} \end{pmatrix}$$
(C.18)

and

$$\mathbf{m} = \begin{pmatrix} e^{\Phi/2} e^{-2A} \tilde{g}_{yy} & e^{\Phi/2} e^{-2A} \tilde{g}_{yz} - e^{\Phi} \ell_s^2 c/2 \\ e^{\Phi/2} e^{-2A} \tilde{g}_{yz} + e^{\Phi} \ell_s^2 c/2 & e^{\Phi/2} e^{-2A} \tilde{g}_{zz} \end{pmatrix}.$$
 (C.19)

As above,

$$\sqrt{-\det \overline{\mathbf{M}}} = e^{\Phi + 4A} \sqrt{-g} \sqrt{e^{\Phi} e^{-4A} \tilde{g}_2 + e^{2\Phi} \ell_s^4 c^2/4}.$$
 (C.20)

The fluctuation $\delta \mathbf{M}$ arises from allowing the embedding of the NS5-brane to fluctuate. We may explicitly write the pull-back as

$$G_{ab} - e^{\Phi} C_{ab} = \Pi_{ab}^{MN} \left(G_{MN} - e^{\Phi} C_{MN} \right).$$
 (C.21)

If we take the embedding of the fivebrane to be specified by $X^M(\xi^a)$ and allow the brane to fluctuate in the transverse directions $X^M(\xi^a) = \delta^M_a \xi^a + \delta^M_j X^j(\xi^b)$, the projection operator is then

$$\Pi_{ab}^{MN} \equiv \frac{\partial X^M}{\partial \xi^a} \frac{\partial X^N}{\partial \xi^b} = \delta_a^M \delta_b^N + \delta_a^M \delta_j^N \frac{\partial X^j}{\partial \xi^b} + \delta_b^N \delta_i^M \frac{\partial X^i}{\partial \xi^b} + \delta_i^M \delta_j^N \frac{\partial X^i}{\partial \xi^a} \frac{\partial X^j}{\partial \xi^b}.$$
 (C.22)

Assuming a product metric $G_{ai}^E = 0$, we have

$$\delta \mathbf{M}_{ab} = e^{\Phi/2} e^{-2A} \tilde{g}_{ij} \frac{\partial X^i}{\partial \xi^a} \frac{\partial X^j}{\partial \xi^b}.$$
 (C.23)

and the NS5-brane action may be written

$$S_{\rm NS5} = -\frac{2\pi}{\ell_s^6} \int d^6 \xi \, e^{-\Phi} e^{4A} \sqrt{-g} \sqrt{e^{\Phi} e^{-4A} \tilde{g}_2 + e^{2\Phi} \ell_s^4 c^2/4} \left(1 + \frac{1}{2} e^{-4A} g^{\mu\nu} \tilde{g}_{ij} \partial_\mu X^i \partial_\nu X^j + \frac{1}{2} \frac{e^{\Phi} e^{-4A} \tilde{g}_2}{e^{\Phi} e^{-4A} \tilde{g}_2 + e^{2\Phi} \ell_s^4 c^2/4} \tilde{g}_{ij} \tilde{g}^{\alpha\beta} \tilde{\nabla}_\alpha X^i \tilde{\nabla}_\beta X^j \right). \quad (C.24)$$

We define the canonically normalized fields as

$$Y^{\hat{\imath}} = F E^{\hat{\imath}}_{\ j} X^{j} \tag{C.25}$$

with

$$F^{2}(x^{\mu}, y, z) = \frac{2\pi}{\ell_{s}^{6}} \frac{e^{-\Phi}}{\sqrt{\tilde{g}_{2}}} \sqrt{e^{\Phi} e^{-4A} \tilde{g}_{2} + e^{2\Phi} \ell_{s}^{4} c^{2}/4} \quad \text{and} \quad \tilde{g}_{ij} = \delta_{\hat{\imath}\hat{\jmath}} E^{\hat{\imath}}{}_{i} E^{\hat{\jmath}}{}_{j}. \quad (C.26)$$

We have assumed that $\tilde{g}_{i\alpha} = 0$, and thus $\tilde{\nabla}_{\alpha} E^{\hat{j}}_{\ j} = 0$. Decomposing in real \tilde{g}_2 harmonics gives

$$\tilde{\nabla}^2 \mathcal{Y}^I = -\frac{e^{-2A}}{\ell_{\rm E}^2} \lambda^I \mathcal{Y}^I \qquad Y^{\hat{\imath}} = \sum_I Y_I^{\hat{\imath}} \mathcal{Y}^I \qquad \int \mathrm{d}^2 \sigma \sqrt{\tilde{g}_2} \, \mathcal{Y}^I \, \mathcal{Y}^J = \delta^{IJ}. \tag{C.27}$$

The action is

$$S_{\rm NS5} = \int d^4x \, \sqrt{-g} \left(-V(c) - \frac{1}{2} g^{\mu\nu} \partial_\mu Y_I^{\hat{\imath}} \partial^\mu Y_{\hat{\imath}}^I - \frac{1}{2} m_I^2(c) Y_I^{\hat{\imath}} Y_{\hat{\imath}}^I \right. \\ \left. + g(c) \left(c \partial_\mu c \right) \left(Y_I^{\hat{\imath}} \partial^\mu Y_{\hat{\imath}}^I \right) - \frac{1}{2} g(c)^2 \left(\partial c \right)^2 Y_I^{\hat{\imath}} Y_{\hat{\imath}}^I \right)$$
(C.28)

with

$$V(c) \equiv \frac{\pi}{\ell_s^4} \frac{e^{4A}}{\sqrt{g_s}} \sqrt{4\left(\frac{\ell_{\rm E}}{\ell_s}\right)^4 + g_s c^2} \tag{C.29a}$$

$$g(c) \equiv \frac{g_s}{2} \left(4 \left(\frac{\ell_{\rm E}}{\ell_s} \right)^4 + g_s c^2 \right)^{-1} \tag{C.29b}$$

$$m_I^2(c) \equiv \frac{\mu_5^2 e^{-\Phi}}{F^4} \frac{e^{-2A}}{\ell_{\rm E}^2} \lambda_I = 4\lambda_I \frac{e^{2A}}{\ell_{\rm E}^2} \left(\frac{\ell_{\rm E}}{\ell_s}\right)^4 \left(4\left(\frac{\ell_{\rm E}}{\ell_s}\right)^4 + g_s c^2\right)^{-1}.$$
 (C.29c)

Finally, we will be interested in the tension of the domain wall interpolating between the metastable and stable states of the NS5-brane axion monodromy scenario. In the thin-wall approximation, the domain wall corresponds to an NS5-brane winding n times around the minimum volume three-cycle Σ_3 whose endpoints are Σ_2 and $\overline{\Sigma}_2$, the two-cycles wrapped by the NS5-brane and anti-NS5-brane, respectively. The tension can then be read off by reducing the action

$$S_{\rm NS5} = -\frac{2\pi}{\ell_s^6} \int_{\mathcal{D}} d^6 \xi \, e^{-2\Phi} \sqrt{-\det\left(G_{ab}^S - e^{\Phi}C_{ab}\right)} \to -T_D \int_{\mathcal{D}^{2,1}} d^3 x \, \sqrt{-\det\mathcal{P}(g_4)}. \quad (C.30)$$

where $\mathcal{P}(g_4)$ denotes the pullback of the spacetime metric $g_{\mu\nu}$ onto the world-volume of the domain wall $\mathcal{D}^{2,1}$.

We can gain intuition for this tension by modeling the three-cycle Σ_3 as

$$ds_{\Sigma_3}^2 = dr^2 + r^2 \left(dy^2 + dz^2 \right),$$
 (C.31)

where the two-torus volume form is $dy \wedge dz$, and $r \in [r_{\cup}, r_{\cap}]$. Then

$$G_{ab}^{S} - e^{\Phi}C_{ab} = \begin{pmatrix} e^{\Phi/2}e^{2A}\mathcal{P}(g_{4})_{\mu\nu} & 0 & 0\\ 0 & e^{-2A}e^{\Phi/2} & 0\\ 0 & 0 & \mathbf{m} \end{pmatrix}$$
(C.32)

where

$$\mathbf{m} = \begin{pmatrix} e^{-2A}e^{\Phi/2}r^2 & -e^{\Phi}\ell_s^2 c/2\\ -e^{\Phi}\ell_s^2 c/2 & e^{-2A}e^{\Phi/2}r^2 \end{pmatrix},$$
(C.33)

and so

$$S_{\rm NS5} = -\left(\frac{2\pi}{\ell_s^6} \int_{r_{\cup}}^{r_{\cap}} \mathrm{d}r \, e^{-3\Phi/4} e^{3A} \sqrt{e^{-4A}r^4 + e^{\Phi}\ell_s^4 c^2/4}\right) \int_{\mathcal{D}^{2,1}} \mathrm{d}^3 x \, \sqrt{-\det \mathcal{P}(g_4)}. \quad (C.34)$$

The tension then takes the form

$$T_{\rm D} = \frac{2\pi}{\ell_s^3} \frac{L^3}{g_s^{3/4} \ell_s^3} \frac{1}{4} \left(e^{4A_{\rm \cap}} - e^{4A_{\rm \cup}} \right) \sqrt{1 + \frac{g_s \ell_s^4 c^2}{4L^4}} \sim \frac{1}{\ell_s^3} N_{\rm D3}^{3/4} \left(e^{4A_{\rm \cap}} - e^{4A_{\rm \cup}} \right), \tag{C.35}$$

since we must have $g_s \ell_s^4 c^2 / 4L^4 \sim N^2 / N_{\rm D3} \ll 1$ to avoid the KPV instability.

D Backreaction on the internal space

When one introduces a source of monodromy in a compactification, and explicitly breaks supersymmetry, the corresponding stress-energy will backreact on the metric, affecting the parameters in the low-energy effective theory. In the NS5-brane model detailed in section 3.2, a key source of stress-energy is anti-D3-brane charge induced on the anti-NS5brane. Because D3-brane charge preserves the same supersymmetry as the background, it will not backreact on the internal metric at leading order.²⁶ However, the anti-D3brane charge will break the remaining supersymmetry of the background and perturb the

²⁶In the presence of anti-brane charge, the D3-brane charge will backreact on the internal metric at second order. Similarly, if the D3-brane charge is large enough a better description becomes available in which the D3-branes are dissolved into flux and a new warped throat is formed, corresponding to the analysis of section 4.2.1. In what follows, we will take $N \ll N_{D3}$ and assume that the induced anti-D3-brane charge may be thought of as a small perturbation to the geometry.

internal metric. Furthermore, both D3-brane and anti-D3-brane charge will perturb the warp factor e^{4A} . In this appendix we calculate the perturbations to the internal metric and the warp factor.

At the level of the supergravity equations of motion, we may approximate the N units of induced anti-D3-brane charge as N anti-D3-branes smeared about the anti-NS5-brane. These anti-D3-branes do not source the metric directly, but do so through a combination of the warp factor e^{4A} and the \tilde{F}_5 field which we denote Φ_- , where $\Phi_{\pm} \equiv e^{4A} \pm \alpha$, and

$$F_5 = (1 + *_{10}) \,\mathrm{d}\alpha(y) \wedge \mathrm{dvol}_{\mathbb{R}^{1,3}}.\tag{D.1}$$

The equations of motion for Φ_{\pm} and the internal Einstein equation read

$$\tilde{\nabla}^2 \Phi_+ = \frac{2}{\Phi_+ + \Phi_-} (\tilde{\nabla} \Phi_+)^2 + \frac{1}{2} g_s \ell_s^4 (\Phi_+ + \Phi_-)^2 \sum_i \delta(\mathrm{D3}_i)$$
(D.2a)

$$\tilde{\nabla}^2 \Phi_- = \frac{2}{\Phi_+ + \Phi_-} (\tilde{\nabla} \Phi_-)^2 + \frac{1}{2} g_s \ell_s^4 (\Phi_+ + \Phi_-)^2 \sum_i \delta(\overline{\text{D3}}_i)$$
(D.2b)

$$\tilde{R}_{mn} = \frac{2}{(\Phi_+ + \Phi_-)^2} \tilde{\nabla}_{(m} \Phi_+ \tilde{\nabla}_{n)} \Phi_-$$
(D.2c)

where we use $\tilde{\nabla}$, etc., to denote quantities related to the unwarped, internal metric \tilde{g}_{mn} . We treat the anti-D3-branes as a perturbation to an imaginary self-dual background, in which

$$\Phi_+ \approx \frac{2r^4}{L^4} \quad \text{and} \quad \Phi_- = 0,$$
(D.3)

 $L^4 \propto g_s N_{\rm D3} \ell_s^4$, and the internal metric is taken to be the conifold, a cone over T^{1,1},

$$\tilde{g}_{mn} \mathrm{d} y^m \mathrm{d} y^n = \mathrm{d} r^2 + r^2 \mathrm{d} s_{\mathrm{T}^{1,1}}^2 = \mathrm{d} r^2 + r^2 \breve{g}_{ij} \mathrm{d} \Psi^i \mathrm{d} \Psi^j.$$
 (D.4)

We linearize the system of equations (D.2) using the expansions²⁷

$$\tilde{g}_{mn} = \tilde{g}_{mn}^{(0)} + \delta \tilde{g}_{mn} \tag{D.5a}$$

$$\delta \tilde{g}_{rr} = \sum_{I_s} \tau^{I_s}(r) \mathcal{Y}^{I_s}(\Psi) \tag{D.5b}$$

$$\delta \tilde{g}_{r\theta} = \sum_{I_v} b^{I_v}(r) \mathcal{Y}_{\theta}^{I_v}(\Psi) \tag{D.5c}$$

$$\delta \tilde{g}_{\theta\phi} = \sum_{I_s} \frac{1}{5} \pi^{I_s}(r) \breve{g}_{\theta\phi} \mathcal{Y}^{I_s}(\Psi) + \sum_{I_t} \phi^{I_t}(r) \mathcal{Y}^{I_t}_{\theta\phi}(\Psi)$$
(D.5d)

$$\Phi_{-} = \Phi_{-}^{(0)} + \delta \Phi_{-}^{(1)}(r, \Psi) = \sum_{I_s} \delta \Phi_{-}^{I_s}(r) \mathcal{Y}^{I_s}(\Psi)$$
(D.5e)

where $\mathcal{Y}^{I_s}(\Psi)$, $\mathcal{Y}^{I_v}_{\theta}(\Psi)$ and $\mathcal{Y}^{I_t}_{\theta\phi}(\Psi)$ are the scalar, transverse vector, and transverse traceless tensor harmonics on $\mathbb{T}^{1,1}$, with appropriate Laplacian eigenvalues $\lambda(I_s)$, $\lambda(I_v)$, and $\lambda(I_t)$, respectively.

 $[\]overline{\tilde{g}_{ij}^{\text{here}}} = \tilde{g}_{\theta\phi}^{\text{there}}$.

The Einstein metric on $T^{1,1}$ is

$$ds_{T^{1,1}}^{2} = \frac{1}{6} \left(d\theta_{1}^{2} + \sin^{2}\theta_{1} d\varphi_{1}^{2} \right) + \frac{1}{6} \left(d\theta_{2}^{2} + \sin^{2}\theta_{2} d\varphi_{2}^{2} \right) + \frac{1}{9} \left(\cos\theta_{1} d\varphi_{1} + \cos\theta_{2} d\varphi_{2} + d\psi \right)^{2}, \qquad (D.6)$$

and in these coordinates a basis of scalar harmonics is given by

$$\mathcal{Y}^{I_s}(\varphi_1, \theta_1, \varphi_2, \theta_2, \psi) = e^{i\frac{R}{2}\psi} e^{im_1\varphi_1} e^{im_2\varphi_2} d_{m_1\frac{R}{2}}^{(j_1)}(\theta_1) d_{m_2\frac{R}{2}}^{(j_2)}(\theta_2)$$
(D.7)

where $I_s \equiv \{j_1, m_1, j_2, m_2, R\}$ is a multi-index, $d_{m_1m_2}^{(j)}(\theta)$ is the Wigner (small) d-matrix, and $\theta_i \in [0, \pi), \phi_i \in [0, 2\pi)$, and $\psi \in [0, 4\pi)$. We will only need the scalar harmonics in the following, with eigenvalues $\lambda_s(j_1, j_2, R) = 6(j_1(j_1 + 1) + j_2(j_2 + 1) - R^2/8)$.

We will only focus on the radial scaling of the dominant metric perturbation. The presence of the anti-D3-brane charge on the NS5-brane may be interpreted as N anti-D3-branes smeared over the two-cycle wrapped by the NS5-brane. The backreaction is heavily dependent on the geometric details of this smearing, so the reported radial scalings may be reduced by suitable geometric tuning. However, we expect a generic smearing to source all possible angular modes and any order-of-magnitude estimates to be set by the dominant mode.

When placed in the background (D.3), the anti-D3-branes will feel a force towards small r. In the actual configuration, interactions with the anti-NS5-brane provide a stabilizing force that keeps the three-brane charge localized around the two-cycle, but the system (D.2) does not account for this force. The effects of the stabilizing force could be included by sourcing appropriate perturbations in the warped throat, but doing so would leave the radial scaling of the dominant perturbation unchanged, and so our analysis applies in any case.

For an anti-D3-brane at (r', Ψ') , we find that

$$\delta\Phi_{-}^{I_{s}}(r;r',\Psi') = -\frac{1}{\Delta_{s}} \frac{g_{s}\ell_{s}^{4}}{L^{4}} \frac{r'^{4}}{L^{4}} \left(\left(\frac{r'}{r}\right)^{2+\Delta_{s}} \theta(r-r') + \left(\frac{r}{r'}\right)^{\Delta_{s}-2} \theta(r'-r) \right) \mathcal{Y}^{I_{s}}(\Psi'), \quad (D.8)$$

so in the area of interest,

$$\delta \Phi_{-}^{I_s}(r;r',\Psi') \propto -\frac{1}{\Delta_s} \frac{N}{N_{\rm D3}} \frac{r'^4}{L^4} \left(\frac{r'}{r}\right)^{2+\Delta_s},\tag{D.9}$$

where $\Delta_s \equiv \sqrt{4 + \lambda(I_s)}$. This Φ_- profile induces a metric perturbation

$$\pi^{0}(r) = 0$$
 $\pi^{I_{s}}(r) \propto r^{2} \frac{N}{N_{\text{D3}}} \left(\frac{r'}{r}\right)^{6+\Delta_{s}},$ (D.10a)

$$\tau^{0}(r) \propto \frac{N}{N_{\text{D3}}} \left(\frac{r'}{r}\right)^{8} \qquad \qquad \tau^{I_{s}}(r) \propto -\frac{N}{N_{\text{D3}}} \left(\frac{r'}{r}\right)^{6+\Delta_{s}}.$$
 (D.10b)

From the spectroscopy of T^{1,1}, the lowest scalar mode has quantum numbers $(\frac{1}{2}, \frac{1}{2}, \pm 1)$ and $\Delta_s = 7/2$, so the dominant metric perturbation is

$$\delta \tilde{g}_{mn} \mathrm{d} y^m \mathrm{d} y^n \propto \frac{N}{N_{\mathrm{D3}}} \left(\frac{r'}{r}\right)^{19/2} \mathcal{Y}^{\frac{1}{2},\frac{1}{2},1}(\Psi) r^2 \check{g}_{\theta\phi} \mathrm{d} \Psi^\theta \mathrm{d} \Psi^\phi \tag{D.11}$$

where $\mathcal{Y}^{\frac{1}{2},\frac{1}{2},1}(\Psi)$ is some real superposition of angular harmonics with $(j_1, j_2, R) = (\frac{1}{2}, \frac{1}{2}, \pm 1).$

The perturbations in both the warp factor and internal metric will alter the gauge coupling function,

$$\frac{8\pi^2}{g_{\rm YM}^2} = \frac{4\pi}{\ell_s^4} \int_{\Sigma_4} d^4 \xi \,\sqrt{\tilde{g}_4} \,e^{-4A},\tag{D.12}$$

on a stack of D7-branes wrapping a divisor Σ_4 , such that

$$\delta\left(\frac{8\pi^2}{g_{\rm YM}^2}\right) = \frac{4\pi}{\ell_s^4} \int_{\Sigma_4} d^4\xi \,\sqrt{\tilde{g}_4} \left(-2\Phi_+^{-2}\left(\delta\Phi_+^{(1)} + \delta\Phi_-^{(1)}\right) + \Phi_+^{-1}\tilde{g}^{ab}\delta\tilde{g}_{ab}\right). \tag{D.13}$$

Away from the NS5-brane, the supergravity approximation becomes accurate and we have shown above N D3-branes induce an $\mathcal{O}(N/N_{\text{D3}})$ fractional perturbation. This contribution is then

$$\delta\left(\frac{8\pi^2}{g_{\rm YM}^2}\right) \propto N_{\rm D3} \int_{\Sigma_4} \mathrm{d}^4 \xi \,\sqrt{\tilde{g}_4} \, r^{-4} \left(\underbrace{-2\Phi_+^{-1}\left(\delta\Phi_+^{(1)} + \delta\Phi_-^{(1)}\right) + \tilde{g}^{ab}\delta\tilde{g}_{ab}}_{\mathcal{O}(N/N_{\rm D3})}\right) \propto N. \tag{D.14}$$

Open Access. This article is distributed under the terms of the Creative Commons Attribution License (CC-BY 4.0), which permits any use, distribution and reproduction in any medium, provided the original author(s) and source are credited.

References

- P.W. Graham, D.E. Kaplan and S. Rajendran, Cosmological relaxation of the electroweak scale, Phys. Rev. Lett. 115 (2015) 221801 [arXiv:1504.07551] [INSPIRE].
- [2] A. Hook and G. Marques-Tavares, Relaxation from particle production, JHEP 12 (2016) 101 [arXiv:1607.01786] [INSPIRE].
- [3] S.P. Patil and P. Schwaller, Relaxing the electroweak scale: the role of broken dS symmetry, JHEP 02 (2016) 077 [arXiv:1507.08649] [INSPIRE].
- [4] S. Di Chiara et al., Relaxion cosmology and the price of fine-tuning, Phys. Rev. D 93 (2016) 103527 [arXiv:1511.02858] [INSPIRE].
- [5] T. Kobayashi, O. Seto, T. Shimomura and Y. Urakawa, *Relaxion window, Mod. Phys. Lett.* A 32 (2017) 1750142 [arXiv:1605.06908] [INSPIRE].
- [6] D. Baumann and L. McAllister, *Inflation and string theory*, Cambridge University Press, Cambridge U.K. (2015).
- [7] R.S. Gupta, Z. Komargodski, G. Perez and L. Ubaldi, Is the relaxion an axion?, JHEP 02 (2016) 166 [arXiv:1509.00047] [INSPIRE].
- [8] L.E. Ibáñez, M. Montero, A. Uranga and I. Valenzuela, *Relaxion monodromy and the weak gravity conjecture*, JHEP 04 (2016) 020 [arXiv:1512.00025] [INSPIRE].
- [9] A. Hebecker, F. Rompineve and A. Westphal, Axion monodromy and the weak gravity conjecture, JHEP 04 (2016) 157 [arXiv:1512.03768] [INSPIRE].
- [10] K. Furuuchi, Excursions through KK modes, JCAP 07 (2016) 008 [arXiv:1512.04684]
 [INSPIRE].

- [11] N. Kaloper and A. Lawrence, London equation for monodromy inflation, Phys. Rev. D 95 (2017) 063526 [arXiv:1607.06105] [INSPIRE].
- [12] N. Kaloper and L. Sorbo, A natural framework for chaotic inflation, Phys. Rev. Lett. 102 (2009) 121301 [arXiv:0811.1989] [INSPIRE].
- [13] N. Kaloper, A. Lawrence and L. Sorbo, An ignoble approach to large field inflation, JCAP
 03 (2011) 023 [arXiv:1101.0026] [INSPIRE].
- [14] N. Kaloper, M. Kleban, A. Lawrence and M.S. Sloth, Large field inflation and gravitational entropy, Phys. Rev. D 93 (2016) 043510 [arXiv:1511.05119] [INSPIRE].
- [15] J.R. Espinosa et al., Cosmological Higgs-axion interplay for a naturally small electroweak scale, Phys. Rev. Lett. 115 (2015) 251803 [arXiv:1506.09217] [INSPIRE].
- [16] K. Choi and S.H. Im, Constraints on relaxion windows, JHEP 12 (2016) 093
 [arXiv:1610.00680] [INSPIRE].
- B. Batell, G.F. Giudice and M. McCullough, Natural heavy supersymmetry, JHEP 12 (2015)
 162 [arXiv:1509.00834] [INSPIRE].
- [18] J.L. Evans, T. Gherghetta, N. Nagata and Z. Thomas, Naturalizing supersymmetry with a two-field relaxion mechanism, JHEP 09 (2016) 150 [arXiv:1602.04812] [INSPIRE].
- [19] K. Choi and S.H. Im, Realizing the relaxion from multiple axions and its UV completion with high scale supersymmetry, JHEP 01 (2016) 149 [arXiv:1511.00132] [INSPIRE].
- [20] N. Fonseca, L. de Lima, C.S. Machado and R.D. Matheus, Large field excursions from a few site relaxion model, Phys. Rev. D 94 (2016) 015010 [arXiv:1601.07183] [INSPIRE].
- [21] L. McAllister, E. Silverstein and A. Westphal, Gravity waves and linear inflation from axion monodromy, Phys. Rev. D 82 (2010) 046003 [arXiv:0808.0706] [INSPIRE].
- [22] E. Silverstein and A. Westphal, Monodromy in the CMB: gravity waves and string inflation, Phys. Rev. D 78 (2008) 106003 [arXiv:0803.3085] [INSPIRE].
- [23] R. Flauger, L. McAllister, E. Pajer, A. Westphal and G. Xu, Oscillations in the CMB from axion monodromy inflation, JCAP 06 (2010) 009 [arXiv:0907.2916] [INSPIRE].
- [24] M. Aganagic, C. Beem, J. Seo and C. Vafa, Geometrically induced metastability and Holography, Nucl. Phys. B 789 (2008) 382 [hep-th/0610249] [INSPIRE].
- [25] A. Klemm, B. Lian, S.S. Roan and S.-T. Yau, Calabi-Yau fourfolds for M-theory and F-theory compactifications, Nucl. Phys. B 518 (1998) 515 [hep-th/9701023] [INSPIRE].
- [26] M. Berg, D. Marsh, L. McAllister and E. Pajer, Sequestering in string compactifications, JHEP 06 (2011) 134 [arXiv:1012.1858] [INSPIRE].
- [27] C.P. Burgess et al., The inflationary brane anti-brane universe, JHEP 07 (2001) 047 [hep-th/0105204] [INSPIRE].
- [28] A. Retolaza, A.M. Uranga and A. Westphal, Bifid throats for axion monodromy inflation, JHEP 07 (2015) 099 [arXiv:1504.02103] [INSPIRE].
- [29] P. Svrček and E. Witten, Axions in string theory, JHEP 06 (2006) 051 [hep-th/0605206]
 [INSPIRE].
- [30] J.P. Conlon, Brane-antibrane backreaction in axion monodromy inflation, JCAP 01 (2012)
 033 [arXiv:1110.6454] [INSPIRE].

- [31] S. Kachru, J. Pearson and H.L. Verlinde, Brane/flux annihilation and the string dual of a nonsupersymmetric field theory, JHEP 06 (2002) 021 [hep-th/0112197] [INSPIRE].
- [32] A. Saltman and E. Silverstein, A new handle on de Sitter compactifications, JHEP 01 (2006)
 139 [hep-th/0411271] [INSPIRE].
- [33] L. McAllister, E. Silverstein, A. Westphal and T. Wrase, The powers of monodromy, JHEP 09 (2014) 123 [arXiv:1405.3652] [INSPIRE].
- [34] R. Flauger, M. Mirbabayi, L. Senatore and E. Silverstein, Productive interactions: heavy particles and non-Gaussianity, JCAP 10 (2017) 058 [arXiv:1606.00513] [INSPIRE].
- [35] J.E. Kim, H.P. Nilles and M. Peloso, Completing natural inflation, JCAP 01 (2005) 005 [hep-ph/0409138] [INSPIRE].
- [36] T.C. Bachlechner, M. Dias, J. Frazer and L. McAllister, Chaotic inflation with kinetic alignment of axion fields, Phys. Rev. D 91 (2015) 023520 [arXiv:1404.7496] [INSPIRE].
- [37] T.C. Bachlechner, C. Long and L. McAllister, *Planckian axions in string theory*, *JHEP* 12 (2015) 042 [arXiv:1412.1093] [INSPIRE].
- [38] D.E. Kaplan and R. Rattazzi, Large field excursions and approximate discrete symmetries from a clockwork axion, Phys. Rev. D 93 (2016) 085007 [arXiv:1511.01827] [INSPIRE].
- [39] T. Flacke, C. Frugiuele, E. Fuchs, R.S. Gupta and G. Perez, *Phenomenology of relaxion-Higgs mixing*, JHEP 06 (2017) 050 [arXiv:1610.02025] [INSPIRE].
- [40] C. Long, L. McAllister and J. Stout, Systematics of axion inflation in Calabi-Yau hypersurfaces, JHEP 02 (2017) 014 [arXiv:1603.01259] [INSPIRE].
- [41] K. Choi, H. Kim and S. Yun, Natural inflation with multiple sub-Planckian axions, Phys. Rev. D 90 (2014) 023545 [arXiv:1404.6209] [INSPIRE].
- [42] N. Arkani-Hamed, L. Motl, A. Nicolis and C. Vafa, The string landscape, black holes and gravity as the weakest force, JHEP 06 (2007) 060 [hep-th/0601001] [INSPIRE].
- [43] C. Cheung and G.N. Remmen, Naturalness and the weak gravity conjecture, Phys. Rev. Lett. 113 (2014) 051601 [arXiv:1402.2287] [INSPIRE].
- [44] A. de la Fuente, P. Saraswat and R. Sundrum, Natural inflation and quantum gravity, Phys. Rev. Lett. 114 (2015) 151303 [arXiv:1412.3457] [INSPIRE].
- [45] T. Rudelius, Constraints on axion inflation from the weak gravity conjecture, JCAP 09 (2015) 020 [arXiv:1503.00795] [INSPIRE].
- [46] M. Montero, A.M. Uranga and I. Valenzuela, *Transplanckian axions!?*, JHEP 08 (2015) 032 [arXiv:1503.03886] [INSPIRE].
- [47] J. Brown, W. Cottrell, G. Shiu and P. Soler, *Fencing in the swampland: quantum gravity constraints on large field inflation*, *JHEP* **10** (2015) 023 [arXiv:1503.04783] [INSPIRE].
- [48] T.C. Bachlechner, C. Long and L. McAllister, *Planckian axions and the weak gravity conjecture*, JHEP 01 (2016) 091 [arXiv:1503.07853] [INSPIRE].
- [49] J. Brown, W. Cottrell, G. Shiu and P. Soler, On axionic field ranges, loopholes and the weak gravity conjecture, JHEP 04 (2016) 017 [arXiv:1504.00659] [INSPIRE].
- [50] B. Heidenreich, M. Reece and T. Rudelius, Weak gravity strongly constrains large-field axion inflation, JHEP 12 (2015) 108 [arXiv:1506.03447] [INSPIRE].

- [51] B. Heidenreich, M. Reece and T. Rudelius, Sharpening the weak gravity conjecture with dimensional reduction, JHEP 02 (2016) 140 [arXiv:1509.06374] [INSPIRE].
- [52] P. Saraswat, Weak gravity conjecture and effective field theory, Phys. Rev. D 95 (2017) 025013 [arXiv:1608.06951] [INSPIRE].
- [53] S. Gandhi, L. McAllister and S. Sjors, A toolkit for perturbing flux compactifications, JHEP 12 (2011) 053 [arXiv:1106.0002] [INSPIRE].