## On the boundedness of Ricci curvature of an indefinite metric

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# 1. Introduction.

Among several conditions on boundedness of Ricci curvature that play important roles in general relativity we have [1, p. 95]

null convergence condition:  $R_{ab}w^aw^b \ge 0$  for all null vectors w:

time-like convergence condition:  $R_{ab}w^aw^b \ge 0$  for all time-like vectors.

In the present paper we consider the following conditions for a Lorentzian manifold M of dimension  $\geq 3$ :

(i)  $R_{ab}w^{a}w^{b} = 0$  for all null vectors w

(ii)  $|R_{ab}w^aw^b| \leq d$  for all time-like unit vectors w, where d is a certain positive number,

and prove that each of these conditions implies that M is an Einstein space, that is,  $R_{ab} = c g_{ab}$ . As a matter of fact, these results are valid for any metric of signature (-, ..., +, ...) and can be stated as therems in linear algebra:

**Theorem 1.** Let V be an n-dimensional real vector space with non-degenerate inner preduct  $\langle , \rangle$  of signature (-, ..., +, ...). If a bilinear symmetric function f on V satisfies the condition

(1) f(x, x) = 0 for all null vectors  $x \in V$ , then there is a constant c such that

$$f(x, y) = c\langle x, y \rangle$$
 for all  $x, y \in V$ .

**Theorem 2.** Let V be as in Theorem 1. If a bilinear symmetric function f on V satisfies the condition

(2)  $|f(x, y)| \leq d$  for all time-like unit vectors x, i.e.  $\langle x, x \rangle = -1$ , where d is a certain positive number, or

(2')  $|f(x, x)| \leq d$  for all space-like unit vectors x, i.e.  $\langle x, x \rangle = 1$ , where d is a certain positive number, then there is a constant c such that

$$f(x, y) = c\langle x, y \rangle$$
 for all  $x, y \in V$ .

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In section 2 we shall prove Theorem 1 as well as an equivalent result (Theorem 1a), which we use for the proof of Theorem 2. In section 3 we give a proof of Theorem 2 and add a remark on a result of Kulkarni [2] on sectional curvature of an indefinite metric.

## 2. Proof of Theorem 1.

One way of proving Theorem 1 is to use the same argument as that in [1, p. 61]. For the sake of completeness we provide the argument.

Let x be a time-like vector and y a space-like vector in V, and consider

$$p(t) = \langle x + ty, x + ty \rangle = \langle x, x \rangle + 2t \langle x, y \rangle + t^2 \langle y, y \rangle$$

and

$$q(t) = f(x + ty, x + ty) = f(x, x) + 2t f(x, y) + t^2 f(y, y),$$

which are polynomials of degree 2 in t. For t = 0, we have p(t) < 0 since x is time-like. For large enough |t|, we have p(t) > 0 since x is space-like. Thus there exist  $t_1 < 0 < t_2$  such that  $p(t_1) = p(t_2) = 0$  and

$$t_1 t_2 = \frac{\langle x, x \rangle}{\langle y, y \rangle} \cdot$$

By condition (1), we have  $q(t_1) = q(t_2) = 0$  and hence

$$t_1 t_2 = \frac{f(x, x)}{f(y, y)} \cdot$$

Therefore

$$\frac{\langle x, x \rangle}{\langle y, y \rangle} = \frac{f(x, x)}{f(y, y)} \quad \text{i.e.} \quad \frac{f(x, x)}{\langle x, x \rangle} = \frac{f(y, y)}{\langle y, y \rangle}, \text{ say, } c.$$

It follows that for any space-like or time-like vector z we get  $f(z,z) = c\langle z, z \rangle$ . This is valid for any null vector z as well. By polarization we easily get

$$f(z, w) = c\langle z, w \rangle$$
 for all  $z, w \in V$ .

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In order to state an equivalent result, we consider the following conditions for a bilinear symmetric function f:

(1a) If  $\langle x, x \rangle = -1$ ,  $\langle y, y \rangle = 1$  and  $\langle x, y \rangle = 0$ , then f(x, y) = 0;

(1b) If  $\langle x, x \rangle = -1$ ,  $\langle y, y \rangle = 1$  and  $\langle x, y \rangle = 0$ , then f(x, x) + f(y, y) = 0.

Lemma 1. (1) implies (1a) and (1b).

To prove this, let x, y be two vectors as in (1a). Then x + y and x - y are null vectors. By (1) we get

$$f(x + y, x + y) = f(x - y, x - y) = 0,$$

i.e.

$$f(x, x) + 2f(x, y) + f(y, y) = 0$$
  
$$f(x, x) - 2f(x, y) + f(y, y) = 0.$$

Hence f(x, y) = 0 and f(x, x) + f(y, y) = 0.

Lemma 2. (1a) implies (1b) and (1).

Let x, y be as in (1b). Then  $x_1 = \cos h tx + \sin h ty$  and  $y_1 = \sin h tx + \cos h ty$  form another orthonormal pair like  $\{x, y\}$ . Thus by (1a) we get

$$0 = f(x_1, y_1) =$$
  
= (cosh t sinh t) (f(x, x) + f(y, y)) + (sinh<sup>2</sup> t + cosh<sup>2</sup> t) f(x, y) =  
= (cosh t sinh t) (f(x, x) + f(y, y)),

since f(x, y) = 0. Thus for  $t \neq 0$  we get f(x, x) + f(y, y) = 0, proving (1b).

Now let u be a null vector. Then we can find x, y such that  $\langle x, x \rangle = -1$ ,  $\langle y, y \rangle = 1$ ,  $\langle x, y \rangle = 0$  and u = x + y. Then

$$f(u, u) = f(x + y, x + y) = f(x, x) + f(y, y) + 2f(x, y) = 0$$

by virtue of (1a) and (1b).

**Remark.** (1b) implies (1a), as can be proved in a similar way. Thus (1), (1a) and (1b) are equivalent.

We now give an alternate proof of the following

**Theorem 1a.** If a bilinear symmetric function f satisfies condition (1a), then there is a constant c such that

$$f(x, y) = c\langle x, y \rangle$$
 for all  $x, y \in V$ .

Let  $\{e_1, \ldots, e_r, e_{r+1}, \ldots, e_n\}$  be an orthonormal basis of V such that

$$\langle e_i, e_i \rangle = -1$$
 for  $1 \leq i \leq r$   
 $\langle e_i, e_j \rangle = 1$  for  $r+1 \leq j \leq n$ .

We shall prove that  $f(e_i, e_j) = 0$  for  $i \neq j$ . If  $1 \leq i \leq r$  and  $r + 1 \leq j \leq n$ , then this is satisfied by virtue of condition (1a). Now assume  $1 \leq i, j \leq r$  (the case where  $r + 1 \leq i, j \leq n$  is similar). Take any  $k \geq r + 1$  and set  $z = \sinh t e_i + \cosh t e_k$ . Then  $\langle z, z \rangle = 1$  and  $\langle z, e_j \rangle = 0$ . Thus condition (1a) implies

$$0 = f(z, e_j) = \sinh t \ f(e_i, e_j) + \cosh t \ f(e_k, e_j) = \sinh t \ f(e_i, e_j),$$

because  $f(e_k, e_j) = 0$ . For  $t \neq 0$ , we obtain  $f(e_i, e_j) = 0$ .

Now let  $c_i = f(e_i, e_i)$  for  $1 \le i \le n$ . By condition (1b) which follows from (1a), we have  $c_i + c_i = 0$  for  $1 \le i \le r$  and  $r + 1 \le j \le n$ . Thus

 $-c_1 = \dots = -c_r = c_{r+1} = \dots = c_n$ , say, c.

It follows that  $f(x, y) = c\langle x, y \rangle$  for all  $x, y \in V$ .

### 3. Proof of Theorem 2.

We shall prove Theorem 2 under assumption (2). Let x, y be two vectors such that  $\langle x, x \rangle = -1$ ,  $\langle y, y \rangle = 1$  and  $\langle x, y \rangle = 0$ . For |t| > 1, we have  $\langle tx + y, tx + y \rangle = 1 - t^2 < 0$ . Thus

$$u = \frac{tx + y}{(t^2 - 1)^{1/2}}$$

is a time-like unit vector. By assumption (2) we get

$$-d \leq \frac{f(tx + y, tx + y)}{t^2 - 1} \leq d$$

that is,

$$-d(t^{2}-1) \leq t^{2} f(x,x) + f(y,y) + 2t f(x,y) \leq d(t^{2}-1).$$

Let  $t \rightarrow 1$  from above. Then

$$f(x, x) + f(y, y) + 2f(x, y) = 0.$$

Let  $t \rightarrow -1$  from below. Then

$$f(x, x) + f(y, y) - 2f(x, y) = 0.$$

From these two equations, we get f(x, y) = 0 and f(x, x) + f(y, y) = 0. By Theorem 1a, we get the conclusion of Theorem 2.

The proof under assumption (2') is similar.

**Remark.** The above proof has been inspired by the work of Kulkarni [2]. He shows for an indefinite metric of signature (-, ..., +, ...) that if the sectional curvature function K is bounded from below (or from above) on the set of all nondegenerate 2-planes, then K is a constant function. For boundedness of K on all time-like (or space-like) 2-planes, we may establish the following.

**Proposition.** Let M be a manifold of dimension  $\geq 3$  with an indefinite metric of signature (-, ..., +, ...). If there is some d > 0 such that

$$|K(p)| \leq d$$
 for all time-like (or space-like) 2-planes p,

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then K is a constant function.

We note that one-sided boundedness on all time-like (or space-like) 2-planes:

$$K(p) \ge d$$
 or  $K(p) \le d$ 

does not imply that K is a constant function, as can be shown by using the spaces

$$S_1^2 \times \mathbb{R}$$
 or  $H_1^2 \times \mathbb{R}$ ,

where  $S_1^2$  (resp.  $H_1^2$ ) is the 2-dimensional Lorentz manifold of constant sectional curvature 1 (resp. -1). The same spaces serve as examples showing that one-sided boundedness of Ricci curvature on all time-like (or space-like) unit vectors:

$$R_{ab}w^a w^b \ge d$$
 or  $R_{ab}w^a w^b \le d$ 

does not imply that the space is Einstein.

#### References

- [1] Hawking, S. W., Ellis, G. F. R., The large-scale structure of space-time, Cambridge-London-New York-Melbourne, Cambridge University Press, 1973.
- [2] Kulkarni, R. S., The values of sectional curvature in indefinite metrics, Comm. Math. Helv. 54, 173-176 (1979).

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