A COMPARISON OF INSTRUMENTAL NEUTRON ACTIVATION ANALYSIS AND INDUCTIVELY COUPLED PLASMA-MASS SPECTROMETRY FOR TRACE ELEMENT DETERMINATION IN PETROLEUM GEOCHEMISTRY

R. H. FILBY,* S. D. OLSEN**

*Department of Chemistry and Nuclear Radiation Center, Washington State University, Pullman, WA 99164 (USA) **Rogaland Research Institute, P.O. Box 2503, 4004 Stavanger (Norway)

(Received January 4, 1994)

INAA and ICP-MS methods for determining trace elements in crude oils, oil fractions and sourcerock bitumens are compared. Results for a wide variety of crude oil types show good agreement for many elements for which the two techniques are suited and which are geochemically important (e.g., Ni, V, Fe, Se, As, etc.). For some other elements, the two techniques are complementary (e.g., B, Be, Cd, Br). The advantages and disadvantages of INAA and ICP-MS in geochemical exploration programs are reviewed and practical examples presented.

Trace elements are minor constituents of virtually all crude oils and their geochemical significance has been the subject of considerable study.^{1,2} For most crude oils, Ni and V are the most abundant elements and occur primarily as metalloporphyrins, predominantly etio or deoxophylloerythroetio (DPEP) porphyrins with small amounts of other structural types, e.g., benzo, tetrahydrobenzo.^{1,2} The Ni and V porphyrins have been used extensively as biomarkers in maturation, depositional environment, and correlation studies.^{1,3-6} In common with other biomarker methods, the use of metalloporphyrins for oil-oil and oil-source rock correlation studies is time-consuming and expensive. Several studies have reported the use of trace element abundances (or elemental ratios) to correlate oil families and oils to source rocks, even though, except for Ni and V, the molecular species of the trace elements are not known.^{2,7-10} The principle involved is that the bitumen generated from kerogen catagenesis in a source rock inherits a trace element distribution pattern from the parent kerogen or from mineral-kerogen interactions during catagenesis. Subsequent migration of the bitumen from the source rock to a trap or reservoir preserves this fingerprint and makes oil-oil and oil-source rock correlations feasible.

Instrumental neutron activation analysis has been used in several correlation studies using trace elements $^{7,8,ll-12}$ primarily because the method is non-destructive and has adequate sensitivity for a range of geochemically important elements, e.g., Ni, V, Se, Br and Fe.¹⁰ An alternative technique, ICP-MS,¹³ has the potential to determine more elements with lower detection limits than INAA. The work reported here compares INAA and ICP-MS for analysis of crude oils and related geochemical materials.

Experimental

Sample Selection: For the intercomparison study, six crude oils were chosen to represent a range of compositions from light oils with low trace element contents and high API gravities to heavy, low API gravity, oils with high Ni and V contents. These oils were supplied by Conoco, Inc., Ponca City, OK. The oils were analyzed by INAA and ICP-MS without further treatment. SRM 1634b Fuel Oil (NIST) was also analyzed by ICP-MS. A source rock bitumen from the New Albany shale (Indiana) was also analyzed.

<u>Neutron Activation Analysis</u>: The INAA method used was similar to that of JACOBS and FILBY.¹⁴ Samples (0.1-0.3 g) were weighed into 2/5 dram polyethylene irradiation vials which were sealed and then doubly encapsulated in 2 dram polyvials. Irradiation conditions and nuclides used are shown in Table 1.

	Irradiation time	Ge(Li) spectroscopy
Nuclide measured	at 5x10 ¹² n/cm ² /sec	aquisition time
28 _{Al,} 27 _{Mg,} 51 _{Ti} 52 _V	5 min	180s
38 _{Cl,} 56 _{Mn}	5 min	1000s
$^{76}_{As'}$ $^{82}_{Br}$, $^{72}_{Ga}$, $^{42}_{K}$ $^{140}_{La}$, 99mTc (for Mo), $^{24}_{Na}$, $^{153}_{Sm}$	8 hr	4000s
	8 hr	40,000s

Table l Irradiation conditions and nuclides measured by INAA



Figure 1. Schematic diagram of ICP-MS spectrometer.

Standards used were NIST SRM 1632a (Coal), SRM 1633a (Fly Ash), SRM 1571 (Orchard Leaves). Gamma-ray spectra were accumulated on a Gc(Li)-ND6700 spectrometer system, and data reduction was performed on a MicroVAX computer.

<u>ICP-MS Analysis</u>: Approximately 1.5g of oil was accurately weighed in a polyethylene container and 50 μ L of a 15 μ g/g In internal standard (in xylene) added. Ultrapure xylene (Merck AR) was added to give a total weight of 15.0g. The xylene solution was mixed thoroughly and stored in a closed polyethylene bottle. A VG Plasma Quad PQ2+ ICP-MS system was used and Ar (99.998%) was used as

Variable	Value	Variable	Value
Plasma RF power	1.8 kW	Nebulizer oxygen flow	0.11 L/min
Reflected power	25 watts	Spray chamber	double pass
ID of torch injector	1.5 mm	Spray chamber temperatur	re -15°C
Coolant argon flow	17 L/min	Sample uptake rate	0.5 ml/min
Auxiliary argon flow	1.2 L/min	Pt sampling cone orifice	0.75 mm
Nebulizer argon flow	0.85 L/min	Pt skimmer cone orifice	1.0 mm
Upgraded interface		Resolution	0.7 amu

Table 2 Instrument conditions for ICP-MS analysis



Figure 2. Mass spectrum of a North Sea oil (concentrations in ng/g listed above element symbol)

the plasma gas. Because the xylene solution was nebulized directly into the plasma, oxygen (99.998%) was added to reduce carbon build-up on the outer Pt cone inlet to the mass spectrometer. A schematic diagram of the ICP-MS system is shown in Figure 1 and instrument conditions are shown in Table 2. Standard curves for each element were established using CONOSTAN (Conoco Oil Co.) elemental standards (organic-solvent based). Figure 2 shows a mass spectrum of a North Sea oil.

Results and Discussion

Table 3 compares concentrations of Ni and V determined by ICP-MS and INAA in six crude oils and an asphaltene-rich bitumen extracted from the New Albany shale.⁵ Also shown in Table 3 are Ni and V data obtained by ICP-MS for the NIST SRM 1634b, Fuel Oil standard.

Agreement between the INAA and the ICP-MS data is generally good over a wide range of Ni and V concentrations and no systematic bias between the methods can be recognized. Data for a heavy and a light oil for elements other than Ni and V are shown in Table 4.

Samp	le <u>Nickel Co</u>	oncentration (L	<u>(g/g)</u>	Vanadium C	Concentration (µ	<u>g/g)¹</u>
	INAA	ICP-MS	Ref. Val	ue ² INAA	ICP-MS	Ref. Value ²
<u>Crude</u>	: Oils					
171-4	99.2±5.1	76±12	120	381±23	314±56	398
171-8	126±7	107±13	137	387±26	396±34	384
28-2	5.80±0.91	4.8±0.1	5.8	1.46±0.11	2.03±0.04	6.2
28-4	2.70±0.71	1.92±0.15	3.5	0.832±0.060	0.893±0.08	1.8
28-6	23.5±1.6	25.5±0.4	24.9	5.74±0.41	5.97±0.85	6.7
28-8	10.2±1.1	9.25±0.1	11.5	1.76±0.13	2.41±0.6	2.5
New A	Albany bitume	n			1 1	
	1900±110	2150±260		2030±150	1700±200	
NIST	SRM 1634b					
		29.7±2.7	28±2		58.1±6.2	55.4±1.1
			<u> </u>		<u> </u>	

Table 3 Ni and V concentrations ($\mu g/g$) in petroleum samples determined by INAA and ICP-MS.

1. Mean ± standard deviation.

2. Conoco, Inc. values for Conoco oils; precision \pm 5% RsD. Certified values for NIST SRM 1634b.

Table 4 indicates good agreement between ICP-MS and INAA, except for Cr and La in crude oil 171-4. The Cr value by ICP-MS may be high because of the difficulty of correcting the ${}^{53}Cr^+$ intensity for ArCH⁺ (53). The discrepancy in La values is currently unresolved. Table 4 also shows that several elements are determined by ICP-MS that cannot normally be determined by INAA in crude oils, e.g., B, Cd, Cu, Li, Pb, Sn, and Ti. Thus ICP-MS complements the suite of elements normally determined by INAA.

A comparison of detection limits for nebulization ICP-MS and INAA is shown in Table 5. For ICP-MS, detection limits are given for xylene which represent practical detection limits in mixed element standards. The determination limits in xylene represent five times the detection limit or a higher value based on apparent concentrations found when running different API gravity oils and where varying amounts of interferences from matrix ions such as ${}^{40}\text{Ar}{}^{12}\text{C}{}^{1}\text{H}^+$ occur. For INAA, detection limits vary with sample composition because radionuclides from elements other than that of interest constitute the Compton background on which the peak of interest is located. Values given are those for a typical intermediate API crude oil, uncontaminated by clay minerals and are calculated assuming a lg sample irradiated in a 5 x $10^{12} \text{ n/cm}^2\text{sec}$ neutron flux for 8.0 hr irradiation and 40,000 sec count on a 20% efficient Ge(Li) detector. For both ICP-MS and INAA the determination limits do not represent the ultimate detection limits which can often be approached by optimizing experimental conditions (i.e., irradiation time, count time) but do provide practical limits for routine analyses of crude oils. The use of single-ion monitoring using the graphite furnace (ETV-see Fig. I) attachement to ICP-MS increases the sensitivity 100 fold.

It can be seen from Table 5 that ICP-MS gives lower detection limits than INAA for many elements, e.g., Ag, Mo, Cd, Ba, Hg, In, Sn, Sr, and Zn and, more importantly, it is sensitive for a number of elements than cannot normally be determined by INAA, i.e., B, Be, Li, and Pb. For the geochemically

		Concentration	<u>(ng/g) or μg/g)^{l,2}</u>		
Element	Crude	Oil 171-4	Crude	Oil 28-2	
	ICP-MS	INAA	ICP-MS	INAA	
			,		
Ag	<13	ND	<13	ND	
As	17.7±2.6	19.5±3	<17	<12	
В	7900±660	ND	<395	ND	
Ва	44±18	ND	61±38	ND	
Be	<l< td=""><td>ND</td><td><l< td=""><td>ND</td><td></td></l<></td></l<>	ND	<l< td=""><td>ND</td><td></td></l<>	ND	
Br	<110	145±10	<110	164±14	
Cd	11±3.6	ND	<8	ND	
Со	168±4	162±10	64±1.7	6l±8.4	
Cr	3170±320	1120±70	<108	<100	
Cu	42±2.9	ND	9.2±2.5	ND	
Fe#	2.6±0.15	<5.3	0.69±0.04	<8.9	
Ga	1150±50	1250±60	<7	<18	
Hg	<47	<50	<27	22±26	
K#	ND	0.506±0.233	ND	2.62±0.46	
La	4.6±0.9	36.2±3.3	<3	<10	
Li	<l< td=""><td>ND</td><td>1.6±0.95</td><td>ND</td><td></td></l<>	ND	1.6±0.95	ND	
Mn	<2	ND	<2	ND	
Мо	2560±50	2340±170	<22	<440	
Na#	ND	3.02±0.15	ND	8.14±0.4	
Pb	<4	ND	<4	ND	
Sb	<12	<15	<10	<28	
Sc	ND	1.18±0.23	ND	< 0.87	
Se	415±52	276±38	<38	120±31	
Sm	<5.2	0.94±0.39	<5.2	<1.6	
Sn	123±8.5	ND	2540±29	ND	
Sr	<9	ND	<13	ND	
Ti#	7±4.7	<81	0.28±0.01	6.84±3.29	
U	<3.4	<22	<3.4	<28	
Zn	163±8.4	<94	97±4.5	<170	

 Table 4

 Comparison of INAA and ICP-MS data for trace elements in light (28-2) and heavy crude oils (171-4)

¹Concentration: Mean \pm standard deviation concentrations in ng/g except elements marked # in µg/g. ²< Values: Less than blank value for ICP-MS or 2 standard deviations of background of photo peak energy. ND--not determined/detected.

Detection Limit (ng/g)Determination Limit (ng/g)Determination Limit (ng/g)Ag 0.7 3.4 1000As 3.1 15 5 Ba 0.4 10 5000 Be 0.3 1.3 NDB10 25 NDNDBr10 550 10Cd 3.0 15 2000 Co 0.1 0.5 0.5 Cr 53 150 3000 20 Cu 63 9.2 46 500 Cu 65 8.6 43 $-$ Fe 57 113 570 50 In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 NDMn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 58 0.7 3.500 0.05 Se ND 3500 0.05 5 Se ND 3500 0.05 Se S2 18 42 10 Sn 0.9 4.3 $50,000$ Sr 0.3 1.5 5000 Fi 47 20 98 100 V 11 53 50 V 0.2 1.4 7.0 25	Element	ICP-M	INA A ²	
Limit (ng/g) Limit (ng/g) Limit (ng/g) Ag 0.7 3.4 1000 As 3.1 15 5 Ba 0.4 10 5000 Be 0.3 1.3 ND B10 25 ND ND Br 110 550 10 Cd 3.0 15 2000 Co 0.1 0.5 0.5 Cr 53 150 3000 20 Cu 63 9.2 46 500 Cu 63 9.2 46 50 Cu 65 8.6 43 - Fe 57 113 570 50 In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 10 Mo 1.4 7.0 100 NI 58 3.0 15 100		Detection	Determination	Determination
Ag 0.7 3.4 1000 As 3.1 15 5 Ba 0.4 10 5000 Be 0.3 1.3 NDB 10 25 NDNDBr 10 550 10 Cd 3.0 15 2000 Co 0.1 0.5 0.5 Cr 53 150 3000 20 Cu 65 8.6 43 -Fe 57 113 570 50 Hg 202 2.0 10 50 In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 NDMn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 58 3.0 15 100 Ni 60 10 50 -Pb 0.4 2.0 NDScND 3500 0.05 Se 822 18 42 10 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 -UNDND 100 V 11 53 500 Y 0.2 1.0 ND		Limit (ng/g)	Limit (ng/g)	Limit (ng/g)
As 3.1 15 5000 Ba 0.4 10 5000 Be 0.3 1.3 ND B 10 25 ND ND Br 100 550 10 Cd 3.0 15 2000 Co 0.1 0.5 0.5 Cr 53 150 3000 20 Cu 63 9.2 46 500 Cu 65 8.6 43 - Fe 57 113 570 50 Hg 202 2.0 10 50 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Si 60 14 7.0 100 Ni 60 10 50 - Pb 0.4 2.0 ND Si 60 1.5 <t< td=""><td>Ασ</td><td>07</td><td>34</td><td>1000</td></t<>	Ασ	07	34	1000
NSDADADASOUDBa 0.4 105000Be 0.3 1.3 NDB 10 25 NDNDBr10 550 10Cd 3.0 152000Co 0.1 0.5 0.5 Cr 53150300020Cu 63 9.2 46500Cu 64 8.6 43-Fe 5711357050Hg 2022.01050In 0.1 0.5 100La 0.2 0.8 5Li28140NDMn 7.9 401.0Ni 58 3.0 15100Ni 601050-Pb 0.4 2.0 NDSk 0.7 3.5 1.0Sk 0.3 1.55000Sk 0.3 1.55000Fi 47 20 98100Fi 48 5.6 28-UNDND100V11 53 50V 0.5 2.5 500Y 0.2 1.0 ND	As	31	15	5
Ba 0.4 10 10 500 Be 0.3 1.3 NDB 10 25 NDNDBr 110 550 10 Cd 3.0 15 2000 Co 0.1 0.5 0.5 Cr 53 150 3000 20 Cu 63 9.2 46 500 Cu 65 8.6 43 $-$ Fe 57 113 570 50 In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 NDMn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 60 10 50 $-$ Pb 0.4 2.0 NDScND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 $50,000$ Sr 0.3 1.5 5000 Fi 47 20 98 100 V 11 53 50 V 0.5 2.5 500 Y 0.2 1.0 ND V 14 7.0 100	Ra	0.4	10	5000
DefDefDefDefB1025NDNDBr1055010Cd3.0152000Co0.10.50.5Cr 53150300020Cu 639.246500Cu 658.643-Fe 5711357050Hg 2022.01050In0.10.5100La0.20.85Li28140NDMn7.9401.0Mo1.47.0100Ni 583.015100Sb0.73.51.0ScND35000.05Se 822184210Sn0.94.350,000Fi 472098100Fi 485.628-UNDND100V115350V0.52.5500Y0.21.0NDZn 661.47.025	Be	0.3	13	ND
Br 10 550 10 Br 10 550 10 Cd 3.0 15 2000 Co 0.1 0.5 0.5 $Cr 53$ 150 3000 20 $Cu 63$ 9.2 46 500 $Cu 65$ 8.6 43 $ Fe 57$ 113 570 50 $Ig 202$ 2.0 10 50 Li 28 140 ND La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 $Ni 58$ 3.0 15 100 Sb 0.7 3.5 1.0 Sc ND 3500 0.95 $Se 82$ 18 42 10 Sn 0.9 4.3 $50,000$ Sr 0.3 1.5 5000 Γ 43 5.6 28 $ U$ ND ND 100 V 11 53 50 V 11 53 50 V 144 7.0 25	B 10	25	ND	ND
LiINJSOISCd 3.0 152000Co 0.1 0.5 0.5 Cr 53 150 3000 20 Cu 63 9.2 46 500 Cu 65 8.6 43 -Fe 57 113 570 50 Hg 202 2.0 10 50 La 0.2 0.8 5 Li 28 140 NDMn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 ScND 3500 $-$ Pb 0.4 2.0 NDScND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 $50,000$ Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 $-$ UNDND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND	Br	110	550	10
Co D D D D D Co 0.1 0.5 0.5 Cr 53 150 3000 20 Cu 65 8.6 43 - Fe 57 113 570 50 Hg 202 2.0 10 50 In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 $50,000$ Sr 0.3 1.5 5000 Γi 47 20 98 100 Γi 48 5.6 28 - U ND ND 100 V 11 53 50 V 11 53 50 V 0.5 2.5 500 V 0.2 1.0 ND	Cd	3.0	15	2000
Cr5.15.05.05.0Cr53150300020Cu639.246500Cu658.643-Fe5711357050Hg2022.01050Li0.10.5100La0.20.85Li28140NDMn7.9401.0Mo1.47.0100Ni 583.015100Ni 601050-Pb0.42.0NDSb0.73.51.0ScND35000.05Se 82184210Sn0.94.350,000Sr0.31.55000Ti 472098100Ti 485.628-UNDND100V115350W0.52.5500Y0.21.0NDZn 661.47.025	Co	01	0.5	0.5
Cu 63 9.2 46 500 Cu 63 9.2 46 500 Cu 65 8.6 43 - Fe 57 II3 570 50 Hg 202 2.0 10 50 La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sc ND 3500 - Sk 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Sti 42 10 100 100 Fi 48 5.6 28 - <	Cr 53	150	3000	20
Cu 65 8.6 43 - Fe 57 113 570 50 Hg 202 2.0 10 50 In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Fi 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 V 11 53 50 V 0.2 1.0 ND	Cu 63	9.2	46	500
Fe 57 113 570 50 Hg 202 2.0 10 50 In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Fi 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND	Cu 65	86	43	-
Hg 202 2.0 10 50 In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND	Ee 57	113	570	50
In 0.1 0.5 100 La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND	Ησ 202	2.0	10	50
La 0.2 0.8 5 Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	In	01	0.5	100
Li 28 140 ND Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	 La	0.2	0.8	5
Mn 7.9 40 1.0 Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Li	28	140	ND
Mo 1.4 7.0 100 Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND	Mn	7.9	40	1.0
Ni 58 3.0 15 100 Ni 60 10 50 - Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Mo	1.4	7.0	100
Ni 60 10 50 - Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Ni 58	3.0	15	100
Pb 0.4 2.0 ND Sb 0.7 3.5 1.0 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Ni 60	10	50	-
Sb 0.7 3.5 1.0 Sb 0.7 3.50 0.05 Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Ph	04	2.0	ND
Sc ND 3500 0.05 Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Fi 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Sb	0.7	3.5	1.0
Se 82 18 42 10 Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND	Sc	ND	3500	0.05
Sn 0.9 4.3 50,000 Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Se 82	18	42	10
Sr 0.3 1.5 5000 Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Sn	0.9	4.3	50.000
Ti 47 20 98 100 Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Sr	0.3	1.5	5000
Ti 48 5.6 28 - U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	 Ti 47	20	98	100
U ND ND 100 V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	Ti 48	5.6	28	
V 11 53 50 W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	U	ND	ND	100
W 0.5 2.5 500 Y 0.2 1.0 ND Zn 66 1.4 7.0 25	v	11	53	50
Y 0.2 1.0 ND Zn 66 1.4 7.0 25	w	0.5	2.5	500
Zn 66 1.4 7.0 25	Y	0.2	1.0	ND
	Zn 66	1.4	7.0	25

Table 5 Detection Limits for ICP-MS and INAA for trace elements in oils.

Notes: ¹ICP-MS detection limits for mixtures of elemental standards; determination limits take account background ions from oil type matrix. Data corrected for solvent dilution (xylene). ND - not determined, ²INAA detection limits based on intermediate API oil; $\phi = 5 \times 10^{12} \text{ n/cm}^2 \text{sec}$; 8 hr (max) irradiation.

important trace elements in crude oils, Ni, V, and Fe, the two techniques are comparable in detection limits.

Both INAA ICP-MS are subject to interferences and other sources of systematic error. In the case of INAA the crude oil matrix, predominantly hydrocarbons and heteroatom-containing organic species, causes little attenuation of the neutron flux during irradiation. However, long irradiations (e.g., more than 30 mins) of light oils result in some sample loss by radiolysis/volatilization. Metal ions are associated with either the metalloporphyrins or with the high molecular weight involatile asphaltenes in the oil, hence negligible loss of metal species occurs as a result of hydrocarbon radiolysis. For such oils, irradiation in scaled polyethylene does not result in excessive pressure build-up because light gases resulting from radiolysis, CH_4 , C_2H_6 , H_2 , O_2 , etc. diffuse rapidly through the polyethylene wall. The potential sources of error in INAA are interfering reactions producing the nuclide of interest, or overlapping γ -rays. Table 6 lists potential interfering reactions for a crude oil matrix.

Interfering reactions in INAA			
Analytical reaction	Interfering reaction	Correction	
$5l_{V(n,v)}52_V$	$52_{Cr(n,p)}52_{V}$	None needed for typical oils where [V]>[Cr]	
⁵⁸ Ni(n,p) ⁵⁸ Co	⁵⁹ Co(n.2n) ⁵⁸ Co	Negligible cross section for thermalized neutron flux	
$58_{Fe(n,\gamma)}59_{Fe}$	$59_{Co(n,p)}59_{F}$	Negligible in crude oils, except for	
•	$^{62}Ni(n,\alpha)^{59}Fe$	very high Ni content	
$50_{\text{Ti}(n,\gamma)}51_{\text{Ti}}$	⁵¹ V(n,p) ⁵¹ Ti	Corrections needed for high V oils	
$75_{As(n,\gamma)}76_{As}$	76Sc(n,p) 76 As	Correction needed only for [Se] >> [As]	
		·	

Table 6

In ICP-MS the principal sources of interference are isobaric ions and plasma gas ions of coincident mass. High concentrations of low ionization energy elements (e.g., Na in sea water) in the sample may also depress plasma temperature but this is not a problem in petroleum analysis. Atomization of organic (xylene) solutions of crude oils into the plasma, however, requires the addition of O₂ to the gas stream to oxidize carbon to CO/CO2 to prevent soot build-up on the mass spectrometer inlet cone. This causes some plasma instability and also results in additional interfering gaseous ions, e.g., ArO+ (56), CO₂H+ (45), ArC⁺ (52). The major interfering gaseous ions are shown in Table 7. In general, isobaric ions of other elements (e.g., ⁸²Se - ⁸²Kr) are corrected for using known isotopic ratios of the analyte element and the interfering element.

Table 8 is an attempt to summarize the major advantages and disadvantages of NAA and ICP-MS as applied to petroleum geochemistry. The high sensitivity of ICP-MS for many elements is probably the most important advantage of ICP-MS for multi-element measurement in oil-oil and oil source rock correlations, and relative to other non-nuclear analytical methods (e.g., AAS, XRF), the method suffers from few sources of interference. For very light oils, and refined products which contain sub-ppm levels of most metals, ICP-MS is the method of choice. On the other hand, for kerogens, source rocks, and many bitumens and asphalts, the insolubility of these complex matrixes in organic solvents or typical aqueous acids and bases requires a difficult dissolution step and increases the blank value in ICP-MS.

Interest	Interfering Gas Ion	Correction
l0 _Р	12ct vary intense line overlan	ETV ICP MS: matrix removed by ashing
24 _{Ma}	$12C_{1}$ +	Lie 25Mg
- Mg 28s;	12_{C16} $14_{N_{c}}$ +	Use - Mg
40 _{Ca}	$40_{\Delta r}$ +	$U_{8} = 43_{C_2(0, 14\%)}$
44 _{Ca}	$12C^{16}O_{2}+$	Use 42 Ca
45 _{Sc}	$12C^{16}O_{2}H^{+}$	Small interference
⁴⁸ Ti	32 _S 16 _O +	Use ⁴⁷ Ti and ⁴⁸ Ti corrected for S
		content using ⁴⁰ Ar ³² S ⁺
51 _V	³⁵ Cl ¹⁶ O ⁺	Small interference in oils (monitor 51/53
		mass ratio)
⁵² Cr	$40_{\rm Ar}^{12}C^{+}$	Use ⁵³ Cr
⁵³ Cr	⁴⁰ Ar ¹³ C ⁺ ; ³⁷ Cl ¹⁶ O	Small interference in oils (monitor 51/53
		mass ratio)
56 _{Fe}	40Ar ¹⁶ O+	Use ⁵⁷ Fe
57 _{Fe}	$40_{\rm Ar}^{16}O^{1}H^{+}$	
⁷⁵ As	$40_{As}35Cl^{+}$	Small interference in oils (monitor 75/77
		mass ratio)

Table 7 Interfering gas ions in ICP-MS

For these materials the non-destructive INAA technique is superior. ICP-MS is capable of high sample throughput compared to INAA; however, the technique is more labor intensive which partially offsets this advantage.

ICP-MS is a complementary technique to INAA in a geochemical exploration program. Elements such as Pb, Tl, Bi, B, Be, Cd cannot be routinely determined by INAA and are best determined by ICP-MS. Elements of geochemical significance such as Ni, V, Fe, Co and other transition metals can be readily determined by both INAA and ICP-MS. In a typical geochemical correlation study in which oils, bitumens and kerogens, and source rocks are analyzed, the versatility and non-destructive nature of INAA is a major advantage--however, the combination of the two techniques provides the most powerful analytical tools for trace element determination in such a program.

Conclusions

ICP-MS and INAA are both extremely powerful methods for trace element determination in petroleum geochemistry. ICP-MS has the capacity to measure a large number of elements with high sensitivity and relatively few interferences. INAA lacks sensitivity for some elements but has sufficient sensitivity for the geochemically significant elements V, Ni, Fe, Se, Br, etc. Light oils are particularly suited to the ICP-MS technique whereas complex matrices such as kerogens, shales, asphalts are more suitably analyzed by INAA. Together, the two techniques offer the goechemist the most comprehensive suite of methods for trace element analysis available.

INAA	ICP-MS	
Adv	antages	
High sensitivity for many elements (sub ng to μ g)	High sensitivity for many elements - sensitivities similar over large mass range (sub ng to μg)	
Minimal sample preparation for a wide variety of matrixes a g karagone bitumone beau light aid	Many elements determined which are not routine by INAA, e.g., Cd, Tl, Bi, Be, I, Li, Pb, Pd, Pt, Sn	
Few interferences on elements of interest;	Isotope ratios of elements can be determined if significant variation present, e.g., Pb isotopes	
Good accuracy and precision demonstrated	High sample throughput (e.g., 20 samples per day)	
for many matrices Standards may be of different matrix to sample	On-line analyses of molecular sub-fractions separated by HPLC or SFC can be done	
Disad	Laser ablation can be used for solid samples	
<u></u>		
Several important elements cannot be routinely determined, e.g., Pb, Cd, B, Be	Light elements may have poor sensitivity because of interferences, e.g., ${}^{10}B$, $({}^{12}C) {}^{24}Mg({}^{12}C{}^{12}C)$, ${}^{28}Si({}^{12}C{}^{16}O)$	
Long analysis time because of decay periods		
Development and investigation from 10 and 10 and	Organic or aqueous solution required for nebulization ICP-MS	
Reactor and radioactivity handling required		
Very light oils undergo significant radiolysis	be in solution for nebulization ICP-MS	
	Pre-analysis treatment introduces blank values and sources of contamination/loss	

Table 8 Advantages and disadvantages of ICP-MS and INAA for geochemical analysis

The authors acknowledge the assistance of Eric Michael of Conoco, Inc., Ponca City, OK for supplying the crude oils and analytical data. The authors also are grateful to G. E. Mercer and Cathy Grimm of the Nuclear Radiation Center, Washington State University, for assistance in the INAA analyses.

*

References

- 1. R. H. FILBY, G. J. VAN BERKEL, in: Metal Complexes in Fossil Fuels, R. H. FILBY, J. F. BRANTHAVER, (Eds), ACS Symposium Ser. No. 344, Washington, DC, 1987, p. 2.
- 2. J. F. BRANTHAVER, R. H. FILBY, in: Metal Complexes in Fossil Fuels, R. H. FILBY, J. F. BRANTHAVER, (Eds), ACS Symposium Ser. No. 344, Washington, DC, 1987, p. 84.
- 3. E. W. BAKER, J. W. LOUDA, in: Biological Markets in the Sedimentary Record, R. B. JOHNS, (Ed.), Elsevier, Amsterdam, 1986, p. 125.
- 4. A. S. MACKENZIE, J. M. E. QUIRKE, J. R. MAXWELL, in: Advances in Organic Geochemistry, 1979, J. R. MAXWELL, A. G. DOUGLAS (Eds), Pergamon Press, Oxford, 1980, p. 239.
- 5. G. J. VAN BERKEL, J. M. E. QUIRKE, R. H. FILBY, Org. Geochem., 14 (1989) 119.
- 6. J. H. CHEN, R. P. PHILP, Chem. Geol., 91 (1991) 139.
- 7. B. HITCHON, R. H. FILBY, K. R. SHAH, Prepr. ACS Div. Petrol. Chem., 18 (1973) 623.
- A. V. HIRNER, in: Metal Complexes in Fossil Fuels, R. H. FILBY, J. F. BRANTHAVER, (Eds), ACS Symposium Ser. No. 344, Washington, DC, 1987, p. 146.
- 9. J. A. CURIALE, in: Metal Complexes in Fossil Fuels, R. H. FILBY, J. F. BRANTHAVER, (Eds), ACS Symposium Ser. No. 344, Washington, DC, 1987, p. 135.
- 10. J. R. ODERMATT, S. A. CURIALE, Chem., Geol., 91 (1991) 99.
- 11. B. HITCHON, R. H. FILBY, AAPG Bull., 68 (1984) 838.
- 12. J. ELLRICH, A. V. HIRNER, H. STARK, Chem. Geol., 48 (1985) 313.
- 13. M. I. CHICARELLI, C. B. ECKARDT, C. R. OWEN, J. R. MAXWELL, G. EGLINTON, R. C. HUTTON, A. N. EATON, Org. Geochem., 15 (1990) 267.
- 14. F. S. JACOBS, R. H. FILBY, Anal. Chem., 55 (1983) 74.