# **Characteristic shocks of crossing velocities in Magnetohydrodynamics**

### Guy BOILLAT<sup>\*</sup>, Augusto MURACCHINI<sup>†</sup> C.I.R.A.M., Università di Bologna Via Saragozza 8, 40123 Bologna (Italia) and Facoltà di Ingegneria dell'Università di Trento Mesiano di Trento, 38100 Trento (Italia)

#### **Abstract**

We illustrate, in the framework of magnetohydrodynamics, an application of crossing shock formulas to a characteristic shock moving perpendicularly to the magnetic field.

## **1 Introduction**

Consider the *conservative hyperbolic* system of N first order partial differential equations for the unknown *field*  $\mathbf{u}(t, x^i)$ 

$$
\partial_t \mathbf{u} + \partial_i \mathbf{f}^i(\mathbf{u}) = \mathbf{f}(\mathbf{u}), \qquad (i = 1, 2, 3)
$$
 (1)

with M *involutive constraints* 

$$
\partial_i \mathbf{c}^i(\mathbf{u}) = \mathbf{c}(\mathbf{u}) \tag{2}
$$

Let (1) supplemented by an additional conservation law (energy or entropy law)

$$
\partial_t h(\mathbf{u}) + \partial_i h^i(\mathbf{u}) = g(\mathbf{u})
$$
\n(3)

where  $h(\mathbf{u})$  is a *convex* function of **u**. The equation (3) can be obtained by multiplying (1) and (2) by suitable N and M vectors  $\mathbf{u}'$  and b respectively.

When a wave velocity  $\lambda$  of (1) is exceptional (see § 3) a characteristic shock may propagate with this velocity. Independently of the number of equations, the jump of the field u

$$
[\mathbf{u}]\stackrel{{\mathrm {\footnotesize def}}}{=} \mathbf{u} - \mathbf{u}_o
$$

<sup>\*</sup>Supported by C.N.R. (G.N.F.M.).

<sup>&</sup>lt;sup>†</sup>Supported by C.N.R. (G.N.F.M.) and MURST (Fondi Ricerche 40  $\%$  and 60  $\%$ ).

i.e. the difference of its values on the back and on the front of the shock surface, can be expressed as a combination of at most  $m + M + 2$  known vectors (*m* multiplicity of  $\lambda$ , M number of constraints).

The simplest case is that of linear equations where the  $f^{i}$ 's of (1) are linear functions of the field u

$$
\mathbf{f}^i = A^i \mathbf{u}
$$

so that the Rankine-Hugoniot equations (see  $(14)$  in § 4) are just

$$
(A_n - \lambda I)(\mathbf{u} - \mathbf{u}_o) = 0
$$

which shows that the jump

$$
[\mathbf{u}]=\alpha^I\mathbf{d}_I
$$

depends linearly on m parameters  $\alpha^{I}$  corresponding to the multiplicity of the eigenvalue.

When the conservative system  $(1)$  (with an additional law  $(3)$ ) is non linear it is still possible to solve the Rankine-Hugoniot equations and to give a simple expression for the jump of the *main field*  $\mathbf{u}'$  (§ 2)

$$
[\breve{\mathbf{u}}'] \stackrel{\text{def}}{=} \breve{\mathbf{u}}' - \breve{\mathbf{u}}'_o = \alpha^I \mathbf{1}_{I^o} + w \mathbf{g}_o
$$

where g and the left eigenvectors  $I_i$  are calculated for the value  $u_o$  of the field before the shock. The scalar w, a non linear function of the parameters  $\alpha^I$ , represents the non linear part of the shock. In particular this formula appears each time a characteristic velocity  $\lambda(\mathbf{u}, \vec{n})$  (then exceptional) with constant (i.e. independent of the direction  $\vec{n}$ ) multiplicity exists.

Instead when multiplicity is variable that is when it occurs only for some values of  $\vec{n}(\mathbf{u})$  (crossing velocities), a new relative vector,  $\mathbf{g}^r$ , has to be added and

$$
[\breve{\mathbf{u}}']=\alpha^I\mathbf{l}_{I^o}+w\mathbf{g}_o+\sigma\mathbf{g}_o^r
$$

Another adjunction may even be necessary when the involutive constraints (2) are associated with the field equations (1). In this case one has the most complicated expression

$$
[\breve{\mathbf{u}}'] = \alpha^I \mathbf{1}_{I^o} - (\mathbf{b} - \mathbf{b}_o)(M_{on} - \lambda_o I)^{-1} \nabla_o \mathbf{c}_{on} + w \mathbf{g}_o + \sigma \mathbf{g}_o^r, \quad I = 1, 2, ..., m \tag{4}
$$

where the second term is connected with the presence of constraints.

The aim of this paper is to illustrate this formula with an example taken from classical magnetohydrodynamics: a shock moving with the normal fluid velocity in a direction perpendicular to the magnetic field vector  $\vec{B}$ .

When the vector  $\vec{n}$  is orthogonal to the magnetic field no less than five velocities coincide (§ 8). Further a constraint (div  $\vec{B} = 0$ ) is present so that here eight vectors are necessary to describe the jump of the eight components of the main field. Surely it would have been better to find a physical example with less vectors than components. Nevertheless we show how to compute easily these various vectors without writing down matrices. Also we study in  $\S$  3, for a mathematical example, the behaviour of the radial velocities when the corresponding  $\lambda$ 's coincide.

### **2 Some remarks on field equations**

As it is well known the introduction of the main field

$$
\mathbf{u}' = \frac{\partial h}{\partial \mathbf{u}}\tag{5}
$$

allows the field equations (1) to be written in a Friedrichs-Lax-Godunov *symmetric form* [1], [2] by means of the Le Gendre transformation [3]-[8]

$$
h'(\mathbf{u}') = \breve{\mathbf{u}}' \cdot \mathbf{u} - h(\mathbf{u}) \tag{6}
$$

 $(\mathbf{\breve{u}})$  is the transpose of **u**) and the introduction of the quantities

$$
h'^{i}(\mathbf{u}') = \breve{\mathbf{u}}' \cdot \mathbf{f}^{i} + \mathbf{b} \cdot \mathbf{c}^{i} - h^{i}
$$
 (7)

In fact

$$
\mathbf{u} = \breve{\nabla}' h', \quad \mathbf{f}^i = \breve{\nabla}' h'^i - \breve{\mathbf{c}}^i \breve{\nabla}' \breve{\mathbf{b}}
$$

Further the involutive constraints must be such that [9]

$$
\nabla \mathbf{c}_n A_n = M_n \nabla \mathbf{c}_n \tag{8}
$$

with constant matrices  $\nabla \mathbf{c}^i$  and  $M^i$  and

$$
A^i \stackrel{\text{def}}{=} \nabla f^i, \quad A_n \stackrel{\text{def}}{=} \nabla f_n, \quad \mathbf{f}_n \stackrel{\text{def}}{=} \mathbf{f}^i n_i
$$

$$
\nabla \stackrel{\text{def}}{=} \left(\frac{\partial}{\partial \mathbf{u}}\right)_{\vec{n} = const.}, \quad \nabla' \stackrel{\text{def}}{=} \left(\frac{\partial}{\partial \mathbf{u}'}\right)_{\vec{n} = const.}
$$

## **3 Crossing velocities**

The wave velocities of propagation in the direction of the unit vector  $\vec{n}$  are given by the eigenvalues  $\lambda(\mathbf{u}, \vec{n})$  of the matrix  $A_n$ . Suppose that some eigenvalue  $\lambda$  has a constant (i.e. independent on  $\vec{n}$ ) multiplicity m and therefore m left and right eigenvectors

$$
l_I(\mathbf{u}, \vec{n}) (A_n - \lambda I) = 0
$$
,  $(A_n - \lambda I) \mathbf{d}_I(\mathbf{u}, \vec{n}) = 0$ ,  $I = 1, 2, ..., m$ 

When  $m > 1$  it is known that  $\lambda$  is *exceptional* that is

$$
\nabla \lambda \cdot \mathbf{d}_I \equiv 0 \tag{9}
$$

As a consequence a *characteristic shock* exists which propagates with velocity  $\lambda$ .

On the other hand if the number of equations (1) is  $N = \pm 2, \pm 3, \pm 4$  (mod 8) [10], [11] the eigenvalues of  $A_n$  cannot all be simple for every  $\vec{n}$ . Although these eigenvalues may still have constant multiplicity as for instance in the banal case of

$$
A_n = a_n I
$$

it may happen (in the so-called case of *variable multiplicity)* that two (or more) eigenvalues, say,  $\lambda^{(1)}(\mathbf{u}, \vec{n})$  and  $\lambda^{(2)}(\mathbf{u}, \vec{n})$  coincide only for some  $\vec{n} = \vec{n}_o(\mathbf{u})$ 

$$
\lambda^{(1)}(\mathbf{u},\vec{n}_o)=\lambda^{(2)}(\mathbf{u},\vec{n}_o)
$$

For the corresponding radial velocities (propagating the weak disturbances)

$$
\vec{\Lambda}^{(i)}(\mathbf{u},\vec{n}) \stackrel{\text{def}}{=} \lambda^{(i)}\vec{n} + \frac{\partial \lambda^{(i)}}{\partial \vec{n}} - \left(\vec{n} \cdot \frac{\partial \lambda^{(i)}}{\partial \vec{n}}\right)\vec{n} \tag{10}
$$

let us define

$$
\vec{\Omega}(\mathbf{u}, \vec{n}_o) = \lim_{\vec{n} \to \vec{n}_o} \left\{ \vec{\Lambda}^{(1)}(\mathbf{u}, \vec{n}) - \vec{\Lambda}^{(2)}(\mathbf{u}, \vec{n}) \right\}
$$

It follows that several cases are possible:

- 1)  $\vec{\Omega}$  does not exist
- 2)  $\vec{\Omega} = 0$
- 3)  $\vec{\Omega} \neq 0$

which are illustrated with the following mathematical example. Let

$$
A^i = \left[ \begin{array}{ccc} a^i & b^i & 0 \\ b^i & c^i & 0 \\ 0 & 0 & (a^i + c^i)/2 + b^i \end{array} \right]
$$

then

$$
\lambda^{(1)/(2)} = \frac{1}{2}(a_n + c_n \pm \sqrt{\Delta}), \quad \Delta = (a_n - c_n)^2 + 4b_n^2, \quad a_n = a^i n_i = \vec{a} \cdot \vec{n}
$$

$$
\lambda^{(3)} = \frac{a_n + c_n}{2} + b_n
$$

and

$$
\vec{\Lambda}^{(1)/(2)} = \frac{1}{2} \left( \vec{a} + \vec{c} \pm \frac{(a_n - c_n)(\vec{a} - \vec{c}) + 4b_n \vec{b}}{\sqrt{\Delta}} \right)
$$
(11)

$$
\vec{\Lambda}^{(3)} = \frac{\vec{a} + \vec{c}}{2} + \vec{b} \tag{12}
$$

Case 1) When  $\vec{n}$  tends to  $\vec{n}_o$  parallel to

 $(\vec{a}-\vec{c})\wedge \vec{b}$ 

the three velocities  $\lambda^{(1)}, \lambda^{(2)}, \lambda^{(3)}$  tend to  $a_n$  but  $\vec{\Lambda}^{(1)/(2)}$  have no limit and no  $\vec{\Omega}$ exists.

Case 2) If

$$
(\vec{a} - \vec{c}) \cdot \vec{n} = 0, \quad b_n > 0
$$

two velocities coincide

$$
\lambda^{(1)}=\lambda^{(3)}=a_n+b_n
$$

and since for such an  $\vec{n}$ , by (11), (12)

$$
\vec{\Lambda}^{(1)} = \vec{\Lambda}^{(3)} = \frac{\vec{a} + \vec{c}}{2} + \vec{b}
$$

$$
\vec{\Omega} = \vec{\Lambda}^{(1)} - \vec{\Lambda}^{(3)} = 0
$$

In these first two cases the differences of velocities  $\lambda^{(i)} - \lambda^{(j)}$  do not change sign when  $\vec{n}$  goes over  $\vec{n}_o$ .

Case 3) Suppose  $\vec{c} = \vec{a}$ .

$$
\lambda^{(1)} = \lambda^{(3)} = a_n + b_n, \qquad \lambda^{(2)} = a_n - b_n
$$

$$
\vec{\Lambda}^{(1)} = \vec{\Lambda}^{(3)} = \vec{a} + \vec{b}, \qquad \vec{\Lambda}^{(2)} = \vec{a} - \vec{b}
$$

so that when  $b_n$  tends to zero,  $\lambda^{(1)},\lambda^{(2)}, \lambda^{(3)}$  tend to  $a_n$  and  $\vec{\Omega}$  is different from zero

$$
\vec{\Lambda}^{(1)}-\vec{\Lambda}^{(2)}=2\vec{b}
$$

The velocities cross:  $\lambda^{(1)} - \lambda^{(2)} = 2b_n$  changes sign with  $b_n$ .

### **4 Explicit shock expression**

A natural question to ask is whether the exceptional property (9) still holds for variable multiplicity i.e. when  $\vec{n} = \vec{n}_o$ . It turns out [12] then that the important exceptionality is not so much that of  $\lambda$  but rather that of the difference of the velocities

$$
\nabla(\lambda^{(1)} - \lambda^{(2)}) \cdot \mathbf{d}_I = 0, \quad I = 1, 2, ..., m \tag{13}
$$

which implies the former one (9).

As already shown, a characteristic shock depending on  $m$  parameters propagating in a direction  $\vec{n}_o$  is possible when the conditions (13) are satisfied. This shock is obtained by solving the Rankine-Hugoniot equations involving the fields *Uo (unperturbed* field) and u *(perturbed* field) ahead and behind the shock front

$$
\mathbf{f}_n(\mathbf{u}) - \lambda_o \mathbf{u} = \mathbf{f}_n(\mathbf{u}_o) - \lambda_o \mathbf{u}_o, \qquad \mathbf{c}_n(\mathbf{u}) = \mathbf{c}_n(\mathbf{u}_o) \tag{14}
$$

together with

$$
\vec{n}(\mathbf{u}) = \vec{n}(\mathbf{u}_o) = \vec{n}_o, \qquad \lambda(\mathbf{u}, \vec{n}_o) = \lambda(\mathbf{u}_o, \vec{n}_o) = \lambda_o \tag{15}
$$

Quite generally, when  $\vec{\Omega} \neq 0$ , an explicit solution of these equations is given in terms of the *jump* of the main field [13]

$$
[\breve{\mathbf{u}}'] = \alpha^I \mathbf{I}_{I^o} - (\mathbf{b} - \mathbf{b}_o)(M_{on} - \lambda_o I)^{-1} \nabla_o \mathbf{c}_{on} + w \mathbf{g}_o + \sigma \mathbf{g}_o^r, \quad I = 1, 2, ..., m \tag{16}
$$

All quantities with the subscript zero are calculated for  $\mathbf{u} = \mathbf{u}_o$  and therefore depend only on the state before the shock.

In (16) the first term, where  $\alpha^I$  are m arbitrary parameters, represents a linear part of the shock while the second term is connected with the presence of constraints. The positive scalar

$$
w = [h'] - \breve{\mathbf{u}}_o[\mathbf{u}'] \tag{17}
$$

vanishes only when the shock does,  $g$  and  $g<sup>r</sup>$  are defined by the following formulas

$$
\mathbf{g}(A_n - \lambda I) = -\nabla \lambda, \quad \mathbf{g} \mathbf{d}_I = 0, \quad \mathbf{g}^r (A_n - \lambda I) = -\nabla (\lambda^1 - \lambda^2), \quad \mathbf{g}^r \mathbf{d}_I = 0.
$$
 (18)

Thus to the usual expression of a characteristic shock with constant multiplicity, an additional term, the last one of (16) is added which is specific of crossing eigenvalues.

The scalar  $\sigma$  comes from the general equality

$$
\phi^i \stackrel{\text{def}}{=} \check{\mathbf{u}}'[\mathbf{f}^i] - [h^i] + \mathbf{b}[\mathbf{c}^i] - w\Lambda_o^i \equiv [h'^i] - [\check{\mathbf{u}}']\mathbf{f}_o^i - [\mathbf{b}]\mathbf{c}_o^i - w\Lambda_o^i = \sigma\Omega_o^i \tag{19}
$$

and, as well as w, is a function of the m parameters  $\alpha^I$ .

In next sections we consider (16) in the case of magnetohydrodynamics by computing its termes.

## **5 The conservative form of the equations of magnetohydrodynamics**

The equations describing the magnetohydrodynamic motions are [14], [15]

$$
\partial_t \rho + \text{div}(\rho \vec{v}) = 0 \tag{20}
$$

$$
\partial_t \vec{B} - \text{rot}(\vec{v} \wedge \vec{B}) = 0 \tag{21}
$$

$$
\partial_t(\rho \vec{v}) + \text{div }\vec{\Pi} = 0 \tag{22}
$$

$$
\partial_t E + \operatorname{div} \vec{q} = 0 \tag{23}
$$

and express the conservation of the mass  $\rho$  (20), of the momentum  $\rho\vec{v}$  (22), of the total energy  $E(23)$  and the evolution of the magnetic field  $\vec{B}(21)$ . To the system (20)-(23) must be added the involutive constraint

$$
\operatorname{div} \vec{B} = 0 \tag{24}
$$

The quantities  $\vec{\Pi}$ , E,  $\vec{q}$ , appearing in (22),(23) are defined:

$$
\vec{\Pi} = (p + \frac{B^2}{2})I - \vec{B} \otimes \vec{B} + \rho \vec{v} \otimes \vec{v}, \quad \vec{q} = \rho \vec{v} \left( e + \frac{p}{\rho} + \frac{v^2}{2} \right) + \vec{B} \wedge (\vec{v} \wedge \vec{B})
$$

$$
E = \rho \frac{v^2}{2} + \rho e + \frac{B^2}{2}
$$

where  $\vec{v}$ , p are respectively the velocity and the pressure of the fluid; e is the density of internal energy satisfying

$$
de = TdS + pd\rho/\rho^2
$$

with entropy  $S$  and temperature  $T$ .

The system  $(20)-(23)$  with  $(24)$  is of the form  $(1)$  and  $(2)$  without the second members and

$$
\mathbf{u} = \begin{bmatrix} \rho \\ \vec{B} \\ \rho \vec{v} \\ E \end{bmatrix}, \qquad \mathbf{f}^i = \begin{bmatrix} \rho v^i \\ v^i \vec{B} - B^i \vec{v} \\ (p + \frac{B^2}{2}) e^i - B^i \vec{B} + \rho v^i \vec{v} \\ (E + p + \frac{B^2}{2}) v^i - (\vec{v} \cdot \vec{B}) B^i \end{bmatrix}, \qquad \mathbf{c}^i = B^i
$$

By taking as supplementary equation (3) the conservation of entropy

$$
\partial_t(\rho S) + \mathrm{div}(\rho S \vec{v}) = 0
$$

one has

$$
h(\mathbf{u}) = -\rho S, \qquad h^i(\mathbf{u}) = -\rho S v^i \qquad g(\mathbf{u}) = 0
$$

and therefore  $[16]$  by  $(5)-(7)$ 

$$
\breve{\mathbf{u}}' = \frac{1}{T} \left( G - \frac{v^2}{2}, \vec{B}, \vec{v}, -1 \right), \quad h' = \frac{1}{T} \left( p + \frac{B^2}{2} \right), \quad h'^i = \frac{v^i}{T} \left( p + \frac{B^2}{2} \right)
$$

where  $G = e + p/\rho - TS$  is the free enthalpy. The vector **b**, in (7), reduces here to a scalar factor of the single constraint div  $\vec{B} = 0$  and has already been calculated [16]. It is easy to see that

$$
b=\frac{\vec{B}\cdot\vec{v}}{T}
$$

#### **6 Wave velocities**

The formal substitution

$$
\partial_t \to -\lambda \delta, \qquad \partial_i \to n_i \delta
$$

permits us to write

$$
\delta \mathbf{f}_n - \lambda \delta \mathbf{u} = (A_n - \lambda I)\delta \mathbf{u} = 0 \tag{25}
$$

which shows immediately that  $\delta u$  is a linear combination of eigenvectors

$$
\delta \mathbf{u} = \beta^I \mathbf{d}_I \tag{26}
$$

Calculating the  $\lambda$  's one finds the Alfvén velocities [14], [15]

$$
\lambda^{(1)} = v_n + \frac{B_n}{\sqrt{\rho}}, \qquad \lambda^{(3)} = v_n - \frac{B_n}{\sqrt{\rho}}
$$

the contact velocity

$$
\lambda^{(2)} = v_n
$$

the slow velocities

$$
\lambda^{(4)/(5)} = v_n \pm \zeta, \qquad \zeta = \left\{ \frac{1}{2} (c^2 + \frac{B^2}{\rho} - \sqrt{\Delta}) \right\}^{\frac{1}{2}}
$$

the fast velocities for which a plus sign precedes the square root of

$$
\Delta = (c^2 + \frac{B^2}{\rho})^2 - 4\frac{B_n^2}{\rho}c^2, \qquad c = \sqrt{(\partial p/\partial \rho)s}
$$

and the null velocity.

By (10) the corresponding radial velocities are

$$
\vec{\Lambda}^{(1)/(3)} = \vec{v} \pm \frac{\vec{B}}{\sqrt{\rho}}, \quad \vec{\Lambda}^{(2)} = \vec{v}, \quad \vec{\Lambda}^{(4)/(5)}(\vec{n}) = \vec{v} \pm \zeta \vec{n} \pm \frac{c^2 B_n}{\rho \zeta \sqrt{\Delta}} (\vec{B} - B_n \vec{n}) \quad (27)
$$

# **7** Case  $\vec{\Omega} = 0$

For

$$
\vec{n}_o = \frac{\vec{B}}{\mid \vec{B} \mid}
$$

the Alfvén velocities become double since  $\cdot$ 

$$
\lambda^{(4)/(5)}(\vec{n}_o) = \lambda^{(1)/(3)}(\vec{n}_o) = \vec{v} \cdot \frac{\vec{B}}{|\vec{B}|} \pm \frac{|\vec{B}|}{\sqrt{\rho}}
$$

while (27) become

$$
\vec{\Lambda}^{(4)/(5)}(\vec{n}_o) = \vec{v} \pm \frac{\vec{B}}{\sqrt{\rho}} = \vec{\Lambda}^{(1)/(3)}
$$

and both  $\vec{\Omega}$ 's are zero

$$
\vec{\Lambda}^{(4)} - \vec{\Lambda}^{(1)} = 0, \quad \vec{\Lambda}^{(5)} - \vec{\Lambda}^{(3)} = 0
$$

## **8** Case of crossing velocities:  $\vec{\Omega} \neq 0$

From now on we consider this possibility. When  $\vec{\Omega}$  is not zero  $\vec{n}_o$  is not unique and the velocities cross, that is, their difference changes sign when  $\vec{n}$  goes over  $\vec{n}_o$ . This clearly occurs in magnetohydrodynamics when

$$
\vec{B} \cdot \vec{n}_o = 0 \tag{28}
$$

Then  $\lambda^{(1)}$  to  $\lambda^{(5)}$  are all equal to

$$
\lambda = v_n \tag{29}
$$

Observe that some of these velocities, for instance  $\lambda^{(2)}$ ,  $\lambda^{(4)}$ , still do not cross;  $\lambda^{(4)} - \lambda^{(2)} = \zeta \ge 0$  does not change sign in agreement with the fact that the corresponding  $\vec{\Omega}$  is zero:  $\vec{\Lambda}^{(4)}(\vec{n}_o) - \vec{\Lambda}^{(2)}(\vec{n}_o) = 0$ . However  $\lambda^{(1)}$ ,  $\lambda^{(2)}$  do

$$
\lambda^{(1)}(\vec{n}) - \lambda^{(2)}(\vec{n}) = \frac{B_n}{\sqrt{\rho}}, \qquad \vec{\Omega} = \vec{\Lambda}^{(1)}(\vec{n}_o) - \vec{\Lambda}^{(2)}(\vec{n}_o) = \frac{\vec{B}}{\sqrt{\rho}}
$$

The non vanishing  $\vec{\Omega}$ 's are proportional to this one as for instance  $\vec{\Lambda}^{(4)/(5)}(\vec{n}_o)$  –  $\vec{\Lambda}^{(1)/(3)}(\vec{n}_o).$ 

It is also easy to verify (13) in magnetohydrodynamics. Although one could check that

$$
\nabla(\lambda^{(1)}-\lambda^{(2)})(\vec{n}_o)=(\vec{0},\frac{\vec{n}_o}{\sqrt{\rho}},\vec{0},0)
$$

is orthogonal to the five right eigenvectors  $\mathbf{d}_I$  associated to (29) it is easier, by (26), to note that

$$
\delta(\lambda^{(1)} - \lambda^{(2)}) = \delta \frac{B_n}{\sqrt{\rho}} = \frac{\delta B_n}{\sqrt{\rho}} = 0
$$

according to the second equation (25) and (28)

$$
\delta(v_n \vec{B} - B_n \vec{v}) - \lambda \delta \vec{B} = 0
$$

multiplied by  $\vec{n}$ , for  $v_n \neq 0$ .

Equations (14), (15) read simply [5], [14], [15]

$$
p + \frac{B^2}{2} = p_o + \frac{B^2_o}{2}, \quad \vec{B} \cdot \vec{n} = \vec{B}_o \cdot \vec{n} = 0, \quad \vec{v} \cdot \vec{n} = \vec{v}_o \cdot \vec{n} = \lambda_o \tag{30}
$$

These are three conditions for eight field components leaving five arbitrary quantities.

We are now going to check  $(16)$  by computing its various components.

# **9** The left eigenvectors for  $B_n = 0, \lambda = v_n$

To avoid the writing of the matrix  $\mathcal{A}_n$  we set

$$
l \equiv (x, \vec{X}, \vec{Y}, y)
$$
 and form the product  $\Phi = l(f_n - \lambda u)$ .

By taking the variation, keeping 1 and  $\lambda$  constant, we get after setting  $\lambda = v_n, B_n = 0$ 

$$
(\delta \Phi)_{\lambda = v_n, B_n = 0} = \mathbf{1}(A_n - v_n I)\delta \mathbf{u} = 0
$$

for any  $\delta$ **u**. Explicitly

$$
\Phi = (v_n - \lambda)\{\rho x + \vec{B} \cdot \vec{X} + \rho \vec{v} \cdot \vec{Y} + yE\} - B_n\{\vec{v} \cdot \vec{X} + \vec{B} \cdot \vec{Y} + y\vec{B} \cdot \vec{v}\} + \left(p + \frac{B^2}{2}\right)(Y_n + yv_n)
$$

and so

$$
\delta \Phi = V \delta v_n + \mathcal{B} \delta B + \mathcal{P} \delta \left( p + \frac{B^2}{2} \right) = 0 \tag{31}
$$

with

$$
\mathcal{V} = \rho x + \vec{B} \cdot \vec{X} + \rho \vec{v} \cdot \vec{Y} + y(E + p + \frac{B^2}{2})
$$
  

$$
\mathcal{B} = -\vec{v} \cdot \vec{X} - \vec{B} \cdot (\vec{Y} + y\vec{v}) \qquad \mathcal{P} = Y_n + yv_n
$$

But  $V, B, P$  must be zero and therefore yield the solutions

$$
l_i = \left(\frac{v_i}{B^2} \vec{B} \cdot \vec{v} - \frac{1}{\rho} B_i, \quad \vec{e}_i, \quad -\frac{v_i}{B^2} \vec{B}, \quad 0\right) \quad i = 1, 2, 3
$$
\n
$$
l_4 = \left(\frac{p + B^2}{\rho} + e - \frac{v^2}{2}, \quad \vec{0}, \quad \vec{v}, \quad -1\right)
$$
\n
$$
l_5 = \left\{- (\vec{B} \wedge \vec{n}) \cdot \vec{v}, \quad \vec{0}, \quad \vec{B} \wedge \vec{n}, \quad 0\right\}
$$

## **10 Determination of the absolute**  and relative  $g's$

In the same manner by its very definition (18) the components of

$$
\mathbf{g} \equiv (x, \vec{X}, \vec{Y}, y)
$$

are determined by the conditions

$$
\delta\Phi + \delta v_n = 0 \tag{32}
$$

Vol. 3, 1996 Characteristic shocks of crossing velocities in magnetohydrodynamics 227

for all  $\delta$ **u** and

$$
\mathbf{g} \cdot \delta \mathbf{u} = 0 \tag{33}
$$

for all  $\delta$ **u** satisfying

$$
\delta v_n = 0, \quad \delta B_n = 0, \quad \delta \left( p + \frac{B^2}{2} \right) = 0 \tag{34}
$$

The conditions (32) are the same than (31) except for the first i.e.

$$
\mathcal{V}=-1,\quad \mathcal{B}=0,\quad \mathcal{P}=0
$$

By (33)

$$
(x + \vec{Y} \cdot \vec{v})\delta\rho + \vec{X} \cdot \delta\vec{B} + \rho\vec{Y} \cdot \delta\vec{v} + y\delta E = 0
$$
\n(35)

Since for a perfect polytropic fluid

$$
p = \mathcal{R}\rho T
$$
,  $p = a\rho^{\gamma} exp\left(\frac{\gamma - 1}{\mathcal{R}}S\right)$ ,  $E = p + \frac{B^2}{2} + \rho\frac{v^2}{2} + \frac{2 - \gamma}{\gamma - 1}p$ 

(35) becomes by (34), with  $\vec{v}_T = \vec{v} - v_n \vec{n}$ 

$$
\left(x + \vec{Y} \cdot \vec{v} + y\frac{v^2}{2}\right)\delta\rho + \left(\vec{X} - \frac{2-\gamma}{\gamma-1}y\vec{B}\right)\cdot \delta\vec{B}_T + \rho(\vec{Y} + y\vec{v})\cdot \delta\vec{v}_T = 0
$$

so that

$$
x + \vec{Y} \cdot \vec{v} + y\frac{v^2}{2} = 0\tag{36}
$$

$$
\vec{X}_T - \frac{2-\gamma}{\gamma - 1} y \vec{B}_T = 0 \tag{37}
$$

$$
\vec{Y}_T + y\vec{v}_T = 0\tag{38}
$$

which, together with  $(34)_1$ , give

$$
\mathbf{g} = \frac{1-\gamma}{\gamma p + B^2} \left\{ \frac{v^2}{2}, \quad \frac{2-\gamma}{\gamma-1} \left( \vec{B} - \frac{\vec{B} \cdot \vec{v}}{v_n} \vec{n} \right), \quad -\vec{v}, \quad 1 \right\}
$$

Now  $\mathbf{g}^r \equiv (x, \vec{X}, \vec{Y}, y)$  satisfies  $(36) - (38)$  and

$$
\mathcal{V}=0, \quad \mathcal{B}=-\frac{1}{\sqrt{\rho}}, \quad \mathcal{P}=0
$$

It appears immediately that

$$
\mathbf{g}^r = \begin{pmatrix} 0, & \frac{\vec{n}}{v_n\sqrt{\rho}}, & \vec{0}, & 0 \end{pmatrix}
$$

It is to be observed that  $g$  and  $g<sup>r</sup>$  by their definition (18) have the dimension of  $1/u$  or equivalently of  $u'/h'$ .

Further  $c_n = B_n$  and  $\nabla c_n = (0, \vec{n}, 0, 0)$  so that  $\nabla c_n f_n \equiv 0$   $\nabla c_n A_n \equiv 0$ implying, by (8),  $M_n = 0$ .

### **11 Jumps of the main field components**

Expliciting eqs. (16) we get, putting  $\tau = w/(\gamma p_o + B_o^2)$ ,

$$
\left[\frac{G-\frac{v^2}{2}}{T}\right] = \frac{\vec{\alpha}\cdot\vec{v}_o}{B_o^2}(\vec{B}_o\cdot\vec{v}_o) - \frac{1}{\rho_o}\vec{B}_o\cdot\vec{\alpha} + \frac{\alpha_4}{\rho_o}\left(B_o^2 - \frac{\rho_o v_o^2}{2} + \frac{\gamma}{\gamma - 1}p_o\right) + \\ -\alpha_5(\vec{B}_o\wedge\vec{n})\cdot\vec{v}_o + (1-\gamma)\frac{v_o^2}{2}\tau \tag{39}
$$

$$
\frac{B}{T} \Big|_{\alpha} = \vec{\alpha} - (\vec{\alpha} \cdot \vec{n})\vec{n} - (2 - \gamma)\tau \vec{B}_o \tag{40}
$$

$$
\begin{aligned}\n\left[\frac{D}{T}\right] &= \vec{\alpha} - (\vec{\alpha} \cdot \vec{n})\vec{n} - (2 - \gamma)\tau \vec{B}_o \\
\left[\frac{\vec{v}}{T}\right] &= -\frac{(\vec{\alpha} \cdot \vec{v}_o)}{B_o^2} \vec{B}_o + \alpha_5 (\vec{B}_o \wedge \vec{n}) + \left[\frac{1}{T}\right] \vec{v}_o\n\end{aligned} \tag{41}
$$

$$
\frac{1}{T} \bigg] = \alpha_4 + (\gamma - 1)\tau \tag{42}
$$

taking account of the condition

$$
v_{on}\vec{\alpha} \cdot \vec{n} + \left[\frac{\vec{B} \cdot \vec{v}}{T}\right] + (2 - \gamma)\tau \vec{B}_o \cdot \vec{v}_o + \frac{\sigma}{\sqrt{\rho}_o} = 0 \tag{43}
$$

which comes, from (30) i.e.  $\vec{n} \cdot \left[\frac{\vec{B}}{T}\right] = 0$ .

## **12** Determination of  $\sigma$  and  $w$

Since

$$
\left[\frac{\vec{B}\cdot\vec{v}}{T}\right] = \frac{[\vec{B}]\cdot[\vec{v}]}{T} + \vec{B}_o\left[\frac{\vec{v}}{T}\right] + v_o\left[\frac{\vec{B}}{T}\right] - \vec{B}_o \cdot \vec{v}_o\left[\frac{1}{T}\right]
$$

equation (43) can be rewritten

$$
\sigma = -\sqrt{\rho_o} \frac{[\vec{B}] \cdot [\vec{v}]}{T} \tag{44}
$$

It is a matter of a simple calculation to check (19) with  $\Lambda_o^i = v_o^i$ .

Substituting eqs. (39)–(42) in (44) one gets the following relation between  $\tau$ ,  $\sigma$ , and the five parameters  $\alpha$ 

$$
\frac{\sigma}{\sqrt{\rho}_o} \left\{ \frac{1}{T_o} + \alpha_4 + (\gamma - 1)\tau \right\} - \frac{\vec{\alpha} \cdot \vec{v}_o}{B_o^2} (\vec{\alpha} \cdot \vec{B}_o) + (\vec{\alpha} \cdot \vec{v}_o)(\alpha_4 + \tau) + \alpha_5 (\vec{B}_o \wedge \vec{n}) \cdot \vec{\alpha} = 0
$$

Another one is needed to determine w and  $\sigma$ . However if one inserts in the definition (17) of w the expressions (39)–(42) one merely obtains an identity.

In fact this new relation is obtained by writing

$$
p + \frac{B^2}{2} = p_o + \frac{B_o^2}{2}
$$
  
For a polytropic fluid  $\left[\frac{G}{T}\right] = -[S]$ 

and the relation determining  $\tau$  results by replacing the jumps by their expressions  $(39)–(42)$  in

$$
\ln \left(1-\frac{1}{2p_o}[B^2]\right)+\frac{\gamma}{\gamma-1}\ln \left(1+T_o\left[\frac{1}{T}\right]\right)+\frac{1}{\mathcal{R}}[S]=0
$$

Here this relation is rather complicated. It is useless to write it explicitly since the shock is not bounded [16]: the transversal part of  $\vec{v}$ , for instance, is not limited as it is easily seen by (30). A different situation may occur when the field equations derive from a variational principle. In that case  $w, \sigma$  lie on a quadric surface [13].

In relativistic magnetohydrodynamics an additional question arises: the subluminal character of wave and shock velocities [5], [17], [18]. It is also interesting to note that, in the relativistic case, the constraint can be eliminated as proved in [19], [20].

### **References**

- [1] S. K. GODUNOV, An interesting class of quasi-linear systems. *Soy. Math.*  2, 947 (1961); Symmetric form of the magnetohydrodynamic equations (in Russian). Preprint, *Akad. Nauk. SSSR Sib. Otd., Vychisl. Tsentr* 3, No.l, 26 (1972)
- [2] K. O. FRIEDRICHS, P. D. LAX, Systems of conservation equations with a convex extension, *Proc. Nat. Acad. Sci. U.S.A.* 68, 1686 (1971)
- [3] T. RUGGERI, A. STRUMI A, Main field and convex covariant density for quasi linear hyperbolic systems, *Ann. Inst. Henri Poincard* 34, 65 (1981)
- [4] A. M. ANILE, S. PENNISI, On the mathematical structure of the relativistic magneto-fluid dynamics. Ann. Inst. Henri Poincaré 46, 27 (1987)
- [5] A. M. ANILE, *Relativistic fluids and magneto-fluids.* Cambridge Monographs on Mathematical Physics, Cambridge University Press (1989)
- [6] I. MULLER, T. RUGGERI, *Extended Thermodynamics.* Springer-Verlag, New York, Berlin (1993)
- [7] G. BOILLAT, S. PLUCHINO, Sopra l'iperbolicità dei sistemi con vincoli e considerazioni sul superfluido e la magnetoidrodinamica, *ZAMP* 36, 893 (1985)
- [8] G. BOILLAT, Sur l'existence et la recherche d'équations de conservation supplémentaires pour les systèmes hyperboliques, *C. R. Acad. Sci. Paris* 278, série A, 909 (1974); Symétrisation des systèmes d'équations aux derivées partielles avec densité d'énergie convexe et contraintes. Ibid. 295, série I, 551 (1982)
- [9] G. BOILLAT, Involutions des syst6mes conservatifs, *C. R. Acad. Sci. Paris,*  307, série I, 891 (1988)
- [10] P. D. LAX, The multiplicity of eigenvalues, *Bull. Amer. Math. Soc.* 6, 213 (1982)
- [11] S. FRIEDLAND, J. W. ROBBIN, J. H. SYLVESTER, On the crossing rule, *Comm. Pure. Appl. Math.* 37, 19 (1984)
- [12] G. BOILLAT, A. MURACCHINI, Chocs caractéristiques de croisement,  $C$ . *R. Acad. Sci. Paris* 310, série I, 229 (1990)
- [13] G. BOILLAT, Expression explicite des chocs caractéristiques de croisement, *C.R. Acad. Sci. Paris* **312**, série I, 653 (1991)
- [14] A. JEFFREY, *Magnetohydrodynarnics.* Oliver and Boyd, London (1966)
- [15] H. CABANNES, *Theoretical Magnetofluiddynarnics.* Academic Press, New York (1970)
- [16] G. BOILLAT, A. MURACCHINI, The structure of the characteristic shocks in constrained symmetric systems with application to magnetohydrodynamics, *Wave Motion* 11,297 (1989)
- [17] A. LICHNEROWICZ, *Relativistic hydrodynamics and magnetohydrodynamics.* Benjamin, New York (1967); Ondes de chocs, ondes infinitésimales et rayons en hydrodynamique et magn~tohydrodynamique relativistes, in: *Relativistic Fluid Dynamics,* Ed. C. Cattaneo, Cremonese, Roma (1971)
- [18] G. BOILLAT, T. RUGGERI, Wave and shock velocities in relativistic magnetohydrodynamics compared with the speed of light, *Continuum Mech. Thatrnodyn.* 1, 47 (1989)
- [19] S. PENNISI, A covariant and extended model for relativistic magnetofluiddynamics, *Ann. Inst. Henri Poincard,* 58, 343 (1993)
- [20] G. BOILLAT, Sur l'dlimination des contraintes involutives, *C.R. Acad. Sci. Paris* 318, série I, 1053 (1994)

Received November 11, 1993 - Revised version received November 15, 1994