Insufficiency of Karmarkar's Condition

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Received June 3, 1980

Abstract

It is shown that Karmarkar's condition which has been claimed to be a necessary and sufficient condition for the spherically symmetric (s.s.) space-time to be of class 1 ceases to be sufficient in case $R_{2323} = 0$. An s.s. line element which satisfies Karmarkar's condition but ceases to be of class 1 is given. This line element describes a conformally flat perfect fluid distribution with zero density.

The problem of imbedding of an s.s. space-time into a five-dimensional flat space, has been investigated in great detail by Eiesland [1], Karmarkar [2], and Takeno [3]. Karmarkar considered an s.s. line element in the form

$$ds^{2} = -A \, d\rho^{2} - B(d\theta^{2} + \sin^{2} \theta \, d\phi^{2}) + C \, d\tau^{2} + 2D \, d\rho \, d\tau \tag{1}$$

where A, B, C, and D are functions of ρ and τ only. Karmarkar has claimed that the necessary and sufficient condition for (1) to be of class 1 is

$$3F^{2} + 8\pi F (4T_{2}^{2} - T_{1}^{1} - T_{4}^{4}) - 64\pi^{2} (T_{1}^{1}T_{4}^{4} - T_{1}^{4}T_{4}^{1}) = 0$$
(2)

where

$$F \equiv -\frac{R_{2323}}{B^2 \sin^2 \theta} = \frac{1}{B} - \frac{1}{4B^2} (CB'^2 - 2D\dot{B}B' - A\dot{B}^2)(AC + D^2)^{-1}$$
(3)

and T_i^j are the components of the energy-momentum tensor defined through Einstein field equations. The overhead dot and prime denote a partial differentiation with respect to ρ and τ , respectively. The condition (2) was first obtained by Eiesland [1] in the form

$$R_{1414}R_{2323} - R_{1212}R_{3434} - R_{1224}R_{1334} = 0, \tag{4}$$

113 0001-7701/82/0200-0113\$03.00/0 © 1982 Plenum Publishing Corporation in terms of curvature components arising out of (1). The object of this note is to show that the condition (2) is only a necessary condition for the space-time (1) to be of class 1.

By the use of the transformation

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$$r = B^{1/2}$$
$$dt = \frac{E(\rho, \tau)}{B^{\prime 2}} \left[2B^{1/2} (A\dot{B} + DB^{\prime}) dr + (CB^{\prime 2} - 2D\dot{B}B^{\prime} - A\dot{B}^{2}) d\tau \right]$$
(5)

where $E(\rho, \tau)$ is an integrating factor, the line element (1) reduces to the form

$$ds^{2} = (CB'^{2} - 2DBB' - AB^{2})^{-1} \left[-4B(AC + D^{2}) dr^{2} + \frac{B'^{2}}{E^{2}} dt^{2} \right]$$
$$-r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2})$$
(6)

When $R_{2323} = 0$, the line element (6) reduces to

$$ds^{2} = -dr^{2} - r^{2}(d\theta^{2} + \sin^{2}\theta \, d\phi^{2}) + \frac{B'^{2}}{E^{2}} \left(CB'^{2} - 2D\dot{B}B' - A\dot{B}^{2}\right)^{-1} dt^{2} \quad (7)$$

as a consequence of (3). Thus it is possible to reduce (1) to the form

$$ds^{2} = -dr^{2} - r^{2}(d\theta^{2} + \sin^{2}\theta \, d\phi^{2}) + e^{\nu} \, dt^{2}$$
(8)

where $v \equiv v(r, t)$ whenever $R_{2323} = 0$ and $B' \neq 0$. However, when $\dot{B} \neq 0$, a similar reduction is possible. In case \dot{B} and B' both vanish simultaneously the spacetime (1) ceases to be of class 1.

It is well known [4] that for a Riemannian space to be of class 1, a necessary and sufficient condition is that a second-order symmetric tensor satisfying the following conditions

$$R_{ijkl} = e(b_{ik}b_{jl} - b_{il}b_{jk}), \quad e = \pm 1$$
(9)

$$b_{ij;k} - b_{ik;j} = 0 \tag{10}$$

should exist. In view of the fact that the only nonzero components of the curvature tensor for (8) are R_{1414} , R_{2424} , and R_{3434} , it is found that the equations (9) imply the following relations:

$$b_{22}b_{33} = 0$$
, $b_{24}b_{33} = 0$, $b_{22}b_{44} - b_{24}^2 \neq 0$, $b_{33}b_{44} \neq 0$

which are clearly inconsistent. Hence the space-time (8) is not of class 1. Thus it has been established that the s.s. space-time ceases to be of class 1 when $R_{2323} = 0$.

Calculations show that the only nonzero components of T_i^j for the line element (8) are T_1^1 , $T_2^2 = T_3^3$, and $R_{2323} = 0$. Therefore, the condition (2) is identically satisfied. Thus it is shown that Karmarkar's condition (2) holds good and yet the space-time (8) is not of class 1. This establishes the insufficiency of Kar-

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markar's condition. In light of this result the claim made by Tikekar [5] and Krishna Rao [6] that every conformally flat perfect fluid was of class 1 could not be upheld. In fact the s.s. line element (8) with $e^{\nu} = (H + \frac{1}{4}Gr^2)^2$, where H and G are functions of t alone, describes a perfect fluid with zero density, it is conformal to a flat space, yet it is not of class 1. A more general metric is furnished by Barnes [7].

The authors are grateful to the referee for his valuable suggestions.

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