Chapter 12 Inversion of the Heat Equation by a Block Based Algorithm Using Spline Wavelet Packets

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Abstract We present a robust algorithm starting from 1D or 2D discrete noised data to approximately invert the heat equation, which is an ill-conditioned problem. Relative contributions of the coherent structure and the noise in different frequency bands of the available data are different. We propose to solve the inversion problem separately in different frequency bands by methods similar to the Tikhonov regularization. This separation is achieved by using spline wavelet packets. The solutions are derived as linear combinations of those wavelet packets.

12.1 Introduction

The problems are formulated as follows: Let the functions $f(x) \in C^2(\mathbb{R}^1)$ and *f*(*x, y*) ∈ *C*²(\mathbb{R}^2) be compactly supported. Denote by U_t^1 and by U_t^2 the linear operators such that $\mathbf{U}_t^1 f(x) = g(x, t)$ and $\mathbf{U}_t^2 f(x, y) = g(x, y, t)$ where $g(x, t)$ and $g(x, y, t)$ are the solutions of the heat equations with the initial conditions $f(x)$ and $f(x, y)$, respectively:

$$
\frac{\partial g(x,t)}{\partial t} = g''_x(x,t), \qquad g(x,0) = f(x),
$$

\n
$$
\frac{\partial g(x,y,t)}{\partial t} = g''_x(x,y,t) + g''_y(x,y,t), \qquad g(x,y,0) = f(x,y).
$$
\n(12.1)

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S. Repin et al. (eds.), *Numerical Methods for Differential Equations, Optimization, and Technological Problems*, Computational Methods in Applied Sciences 27, DOI [10.1007/978-94-007-5288-7_12,](http://dx.doi.org/10.1007/978-94-007-5288-7_12) © Springer Science+Business Media Dordrecht 2013 219 **Problem 12.1** Let *t* be a fixed time parameter. Given $g(x, t) = U_t^1 f(x)$, find $f(x)$.

Problem 12.2 Let *t* be a fixed time parameter. Given $g(x, y, t) = \mathbf{U}_t^2 f(x, y)$, find $f(x, y)$.

For brevity, we concentrate on Problem [12.1.](#page-1-0) Extension to the 2D case is straightforward. The problem has explicit theoretical solutions [\[6](#page-17-0)]. We assume that the initial temperature distribution $f(x)$ is a *T*-periodic function. Consequently, $g(x, t) = U_t^1 f(x)$ is *T*-periodic as well. These functions can be expanded into the following Fourier series:

$$
f(x) = \frac{1}{T} \sum_{n \in \mathbb{Z}} f_n e^{2\pi i n x/T}, \qquad g(x, t) = \frac{1}{T} \sum_{n \in \mathbb{Z}} g_n(t) e^{2\pi i n x/T}, \qquad g_n(0) = f_n.
$$

If we know the function $g(x, t)$ at some fixed *t* then

$$
f_n = g_n(t)e^{t(2\pi n/T)^2} \Longrightarrow f(x) = \frac{1}{T} \sum_{n \in \mathbb{Z}} g_n(t)e^{t(2\pi n/T)^2} e^{2\pi i n x/T}.
$$
 (12.2)

In real life, the function $g(x, t)$ is known up to some errors, modeled as $\tilde{g}(x, t) =$ $g(x, t) + e(x)$. Generally, there is no reason to assume that the Fourier coefficients of the error tend to zero faster than *e*−*n*2*^t* (if they tend to zero at all). Therefore, accord-ing to [\(12.2](#page-1-1)), application of the direct inversion to the available data $\tilde{g}(x, t)$ results in an unstable solution. However, as the magnitude of the error function $|e(x)|$ becomes smaller, the function $\tilde{f}(x)$ can comprise strong high-frequency components, which do not exist in the original function $f(x)$. Therefore, a regularization, which provides a stability to the solution at the expense of deviation from the available data $\tilde{g}(x, t)$, is needed.

Typically, the data function $g(x, t)$ is sampled on a grid $\{x[k]\}$ and the samples are corrupted by a random noise, while the sought-after initial temperature distribution $f(x)$ is continuous. Therefore, it is reasonable to design approximated solutions as splines. Splines bridge the gap between the discrete input data and the continuous solution. To take into account different relative shares of the coherent signal and the noise in different frequency components of the available data, we propose to solve the inversion problem separately in different frequency bands. This approach significantly extends the adaptation abilities and the robustness of the method. Practically, this scheme is implemented via the application of the orthonormal spline wavelet packets, which are constructed by using the Spline Harmonic Analysis (SHA) framework. The wavelet packet transform splits the frequency domain of a signal into a set of bands whose overlap is minimal.

12.2 Elements of SHA

We briefly outline the basics of the SHA techniques. A detailed description is given, for example, in [\[2](#page-17-1), [3](#page-17-2)].

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We assume that $N = 2^{j}$, $j \in \mathbb{N}$ and $p = 2r > 0$ is an even integer. The space of *N*-periodic splines of even order $p = 2r$, which have nodes on the grid $\{k\}$, is denoted by $p\mathscr{S}$. A basis in $p\mathscr{S}$ is formed by translations of the centered periodic B-spline $B^p(x)$:

$$
S^{p}(x) = \sum_{k=0}^{N-1} q[k]B^{p}(x-k) \in {}^{p}\mathscr{S}.
$$

The B-spline $B^p(x)$ belongs to the space C^{p-2} . The circular convolution $B^p \star$ $B^r(x) = B^{p+r}(x) \Longrightarrow S_1^p \star S_2^r(x) = S_3^{p+r}(x) \in P^{+r} \mathscr{S}$. Thus, the circular convolution of two periodic splines is a spline. Therefore, splines are a proper tool for dealing with convolution-type problems where inversion of the heat equation belongs to.

There exist orthogonal bases in $P\mathscr{S}$ which resemble the Fourier basis. Denote $\omega \stackrel{\text{def}}{=} e^{2\pi i/N}$. The orthogonal basis of the space $P\mathscr{S}$ is constituted by *exponential splines*

$$
\beta^{p}[n](x) \stackrel{\text{def}}{=} \sum_{k=0}^{N} \omega^{-nk} B^{p}(x+k), \quad n = 0, ..., N-1.
$$

The following representation holds:

$$
S^{p}(x) = \sum_{k=0}^{N-1} q[k]B^{p}(x-k) = \frac{1}{N} \sum_{n=0}^{N-1} \widehat{q}[n] \beta^{p}[n](x).
$$

Here $\hat{q}[n] = \sum_{k=0}^{N} \omega^{-nk} q[k]$ is the discrete Fourier transform (DFT) of the coefficients $\{a[k]\}$. For further use, we single out the sequence cients $\{q[k]\}\$. For further use, we single out the sequence

$$
u^{p}[n] \stackrel{\text{def}}{=} \beta^{p}[n](0) = \sum_{k=0}^{N-1} \omega^{-nk} B^{p}(k), \qquad (12.3)
$$

which is the DFT of the sampled B-spline. The sequences $u^p[n]$ are N-periodic and strictly positive. The norms of the exponential splines are $\|\beta^p[n]\| = \sqrt{Nu^{2p}[n]}$. Thus, the splines,

$$
\gamma^{p}[n](x) \stackrel{\text{def}}{=} \frac{\beta^{p}[n](x)}{\|\beta^{p}[n]\|} = \frac{\beta^{p}[n](x)}{\sqrt{Nu^{2p}[n]}}, \quad n = 0, ..., N - 1,
$$

form an orthonormal basis of \mathcal{S}^p . The spline $S^p(x) \in {}^p \mathcal{S}$ is represented by

$$
S^{p}(x) = \sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} \sigma[n] \gamma^{p}[n](x), \quad \sigma[n] = \sqrt{u[n]^{2p}} \hat{q}[n]. \tag{12.4}
$$

This expansion imposes a specific form of the SHA methodology onto the spline space, where the splines $\left\{ \gamma^{p}[n](x) \right\}_{n=0}^{N-1}$ act as harmonics and the coordinates

 ${\sigma[n]}$, $n = 0, \ldots, N - 1$, which we refer to as to the SHA spectrum of the spline $S^p(x)$. They act as the Fourier coefficients. Many operations on splines are sig-nificantly simplified [\[2](#page-17-1), [3\]](#page-17-2). Denote by δ^2 the second central difference: $\delta^2 f(x) =$ $f(x+1) - 2f(x) + f(x-1)$ and

$$
w[n] \stackrel{\text{def}}{=} 4 \sin^2 \frac{\pi n}{M}, \qquad W[n] \stackrel{\text{def}}{=} \sqrt{\frac{u^{2(p-2)}[n]}{u^{2p}[n]}} w[n], \qquad V[n] \stackrel{\text{def}}{=} \frac{u^p[n]}{\sqrt{u^{2p}[n]}}.
$$
\n(12.5)

Then

$$
\delta^2 S^p(x) = -\sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} w[n] \sigma[n] \gamma^p[n](x), \qquad S^p(k) = \frac{1}{N} \sum_{n=0}^{N-1} \omega^{kn} V[n] \sigma[n], \tag{12.6}
$$

$$
S''(x) = -\sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} W[n] \sigma[n] \gamma^{p-2}[n](x), \qquad \left\| S'' \right\|^2 = \frac{1}{N} \sum_{n=0}^{N-1} \left| W[n] \sigma[n] \right|^2.
$$
\n(12.7)

It follows from [\(12.6\)](#page-3-0) that, if a spline $S^p(x)$ interpolates a sequence $y = \{y[k]\}$ at grid points $S^p(k) = y[k]$, then its SHA spectrum is

$$
\sigma[n] = \frac{\hat{y}[n]}{V[n]}, \quad \hat{y}[n] = \sum_{k=0}^{N-1} \omega^{-kn} y[k]. \tag{12.8}
$$

12.3 Global Regularized Spline Solution

We briefly outline the scheme for global solution, which is a realization of the Tikhonov regularization algorithm [[7\]](#page-17-3) in the space of periodic splines. A full presentation of the scheme is given in [\[1\]](#page-17-4).

To immerse Problem [12.1](#page-1-0) into the spline setting, V_t denotes the linear operator defined on the spline space $\mathbb{P}\mathscr{S}$ such that $\mathbf{V}_tS(x) = s(x, t)$, where $s(x, t)$ is the spline solution to the difference approximation of the heat equation

$$
\frac{\partial s(x,t)}{\partial t} = \delta_x^2 s(x,t), \qquad s(x,0) = S(x) \tag{12.9}
$$

from $\mathbb{P}\mathscr{S}$ (with respect to *x*).

Assume the spline $S^p(x)$ is represented by ([12.4](#page-2-0)). The spline $s(x, t) \in \mathcal{S}^p$ can be represented as

$$
s(x, t) = N^{-1/2} \sum_{n=0}^{N-1} \sigma[n](t) \gamma^{p}[n](x).
$$

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Using (12.6) , we get

$$
\sigma[n](t) = \eta[n](t)\sigma[n],\tag{12.10}
$$

where $\eta[n](t) = e^{-w[n]t}$. If we know the spline $s(x, t)$ by a fixed *t* then the spline *S(x)* becomes

$$
S(x) = \sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} e^{w[n]t} \sigma[n](t) \gamma^{p}[n](x).
$$

However, typically only the data vector $z = \{z[k]\}, k = 0, \ldots, N - 1$, is known, where $z[k] = g(k, t) + e_k$, $\mathbf{e} = \{e_k\}$ are the measurement errors, which we assume to be white noise.

The approximated solution to Problem [12.1](#page-1-0) is derived as a spline

$$
S_{\rho}(x) = \arg\min_{S \in \mathscr{S}^p} (\rho I(S) + E(S)),
$$

where

$$
I(S) \stackrel{\text{def}}{=} \|S''\|^2 = \frac{1}{N} \sum_{n=0}^{N-1} |W[n]\sigma[n]|^2,
$$

$$
E(S) \stackrel{\text{def}}{=} \sum_{k} (\mathbf{V}_t S(k) - z[k])^2 = \frac{1}{N} \sum_{n=0}^{N-1} |V[n]\sigma[n]\eta[n](t) - \hat{z}[n]|^2.
$$

A spline solution to the minimization problem is

$$
S_{\rho}(x) = \sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} \sigma[n](\rho) \gamma^{p}[n](x), \quad \sigma[n](\rho) = \frac{\overline{\eta}[n](t)\hat{z}[n]V[n]}{A[n](\rho)},
$$

where $\overline{\eta}[n](t)$ is the complex conjugate of $\eta[n](t)$, $A[n](\rho) \stackrel{\text{def}}{=} \rho W[n] +$ $(n[n](t)V[n])^2$.

Assume we are able to estimate the variance var $(\mathbf{e}) = \varepsilon^2$ of the error vector. The regularization parameter ρ is derived from the solution of the equation

$$
e(\rho) \stackrel{\text{def}}{=} E(S_{\rho})/N = \frac{1}{N^2} \sum_{n=0}^{N-1} \left(\frac{\rho W[n] |\hat{z}[n]|}{A[n](\rho)} \right)^2 = \varepsilon^2.
$$
 (12.11)

The function $e(\rho)$ grows strictly monotonically as $\rho \to \infty$ and $\lim_{\rho \to \infty} e(\rho) =$ N^{-2} $||z||^2$ **. If** N^{-1} $||z|| > \varepsilon$, then [\(12.11\)](#page-4-0) has a unique solution.

The parameter ρ , which provides a trade-off between approximation and regularization, depends on the relative shares of the coherent signal and the noise in the available data. These shares are different in different frequency components of the data. We propose to solve the problems separately in different frequency bands, while the regularization parameters are to be found according to the signal-to-noise

ratio in each band. It is achieved by the application of the orthonormal spline wavelet packet transform, which splits the frequency domain of a signal into a set of bands whose overlap is minimal. The SHA framework provides tools for the design of wavelet packets and for the efficient implementation of the algorithm.

12.4 Wavelet Packets

Denote by $^p\mathscr{S}_{r,0}$, $r \in \mathbb{N}$, the space of *N*-periodic splines of even order *p* on the grid $\{2^r k\}$. In the rest of the paper $N_r \stackrel{\text{def}}{=} N/2^r$, $n_r \stackrel{\text{def}}{=} n + N_r/2$. The space ${}^p \mathcal{S}_{r,0}$ is an N_r -dimensional space, where a basis consists of 2^r -sample shifts of the Bsplines B_r^p constructed on the grid $\{2^r k\}$. The inclusion relations between the spaces *p*S*r*₁,0</sub> ⊂ *p* $\mathscr{S}^{r-1,0}$ ⊂ ··· ⊂ *p* $\mathscr{S}^{0,0}$ ≡ \mathscr{S}^{p} hold. Similarly to the space \mathscr{S}^{p} , the orthogonal and orthonormal bases of $P\mathcal{S}_{r,0}$ are formed by the exponential splines

$$
\beta_{r,0}^p[n](x) \stackrel{\text{def}}{=} \sum_{k=0}^{N_r} \omega^{-2^r n k} B^p(x + 2^r k), \qquad u_r^p[n] \stackrel{\text{def}}{=} \beta_{r,0}^p[n](0),
$$

$$
\gamma_{r,0}^p[n](x) \stackrel{\text{def}}{=} \frac{\beta_{r,0}^p[n](x)}{\sqrt{N_r u_r^{2p}[n]}}.
$$

For the initial scale, we retain the notations $\gamma^{p}[n] \equiv \gamma^{p}_{0,0}[n](x)$, $u^{p}[n] \equiv u^{p}_{0}[n]$.

When it will not produce a confusion, we drop the order index P .

The two-scale relation between basis splines from adjacent spaces holds to be

$$
\gamma_{r,0}[n](x) = b_{r-1}[n]\gamma_{r-1,0}[n](x) + b_{r-1}[n_r]\gamma_{r-1,0}[n_r](x),\tag{12.12}
$$

where

$$
b_{r-1}[n] \stackrel{\text{def}}{=} \sqrt{\frac{u_{r-1}^{2p}[n]}{2u_r^{2p}[n]}} \cos^p\left(\frac{2^{r-1}\pi n}{N}\right), \quad n_r \stackrel{\text{def}}{=} n + \frac{N_r}{2}.
$$

Denote by ${}^p\mathscr{S}_{r,1}$ the orthogonal complement to ${}^p\mathscr{S}_{r,0}$ in the space ${}^p\mathscr{S}^{r-1,0}$. An orthonormal basis in $\mathcal{P}_{r,1}$ contains the splines

$$
\gamma_{r,1}[n](x) = \tilde{b}_{r-1}[n]\gamma_{r-1,0}[n](x) + \omega_{r-1}^{2^{r-1}n}[n_r]\gamma_{r-1,0}[n_r](x),\tag{12.13}
$$

where $\tilde{b}_{r-1}[n] \stackrel{\text{def}}{=} \omega^{2^{r-1}n} b_{r-1}[n_r]$. If $r > 1$, we can apply a similar procedure to the space ${}^p\mathscr{S}_{r-1,1}$. As a result, we get the decomposition ${}^p\mathscr{S}_{r-1,1} = {}^p\mathscr{S}_{r,2} \oplus {}^p\mathscr{S}_{r,3}$. By applying the same procedure to all the derived subspaces, we decompose the spline space $\mathbb{P}\mathscr{S}$ into a series of orthogonal sums

$$
{}^{p}\mathscr{S}={}^{p}\mathscr{S}_{1,0}\oplus{}^{p}\mathscr{S}_{1,1}={}^{p}\mathscr{S}_{2,0}\oplus{}^{p}\mathscr{S}_{2,1}\oplus{}^{p}\mathscr{S}_{2,2}\oplus{}^{p}\mathscr{S}_{2,3}=\cdots=\bigoplus_{l=0}^{2^{r}-1}{}^{p}\mathscr{S}_{r,l}.
$$

The orthonormal bases $\{\gamma_{r,l}[n](x)\}[n]$ of the subspaces $\gamma_{r,l}$ are derived iteratively by the two-scale relations using the coefficients $b_{r-1}[n]$ and $\tilde{b}_{r-1}[n]$.

Similarly to the Fourier exponentials, the exponential basis splines $\gamma_{r,l}[n](x)$ are complex-valued and are not localized in the space domain. However, their realvalued and well-localized counterparts satisfy

$$
\psi_{r,l}(x) \stackrel{\text{def}}{=} \sqrt{\frac{1}{N}} \sum_{n=0}^{N_r - 1} \gamma_{r,l}[n](x) \in {}^p \mathcal{S}_{r,l} \subset {}^p \mathcal{S}. \tag{12.14}
$$

These splines are called the spline wavelet packets. The shifts ${\psi_{r,l}(x - 2^r k)}$, $k =$ $0, \ldots, N_r - 1$, form an orthonormal basis for the space $\mathbb{P}\mathcal{S}_{r,l}$. Consequently, the union $\biguplus_{l=0}^{2^r-1} {\{\psi_{r,l}(x-2^r k)\}}$ forms an orthonormal basis for the entire space *pS*.

At the initial scale, the one-sample shifts of the splines

$$
\varphi^{p}(x) \stackrel{\text{def}}{=} \psi^{p}_{0,0}(x) = N^{-1/2} \sum_{n=0}^{N-1} \gamma^{p}[n](x)
$$

form an orthonormal basis.

All the spaces $^p\mathscr{S}_{r,l}$ belong to $^p\mathscr{S}$, thus, the wavelet packet $\psi_{r,l}(x)$ forms a subspace $^p\mathscr{S}_{r,l}$ and can be expanded over the orthonormal basis $\{\gamma^p[n](x)\}\$ of \mathscr{S}^p :

$$
\psi_{r,l}(x) = \sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} \nu_{r,l}[n] \gamma^{p}[n](x).
$$
 (12.15)

The SHA spectra $\{v_{r,l}[n]\}_{n=0}^{N-1}$ of the wavelet can be explicitly calculated using the two-scale relations.

Example 12.1 (The first decomposition scale $r = 1$) The SHA spectra are

$$
\nu_{1,0}[n] = \sqrt{2}b_0[n] = \sqrt{\frac{u^{2p}[n]}{u_1^{2p}[n]}} \cos^p \frac{\pi n}{N},
$$

$$
\nu_{1,1}[n] = \sqrt{2}\overline{\tilde{b}}_0[n] = \omega^{-n} \sqrt{\frac{u^{2p}[n_r]}{u_1^p[n]}} \sin^p \frac{\pi n}{N}.
$$

Example 12.2 (The second decomposition scale $r = 2$) The SHA spectra are

$$
\nu_{2,0}[n] = 2b_0[n]b_1[n], \qquad \nu_{2,1}[n] = 2b_0[n]\tilde{b}_1[n],
$$

$$
\nu_{2,2}[n] = 2\overline{\tilde{b}}_0[n]\overline{\tilde{b}}_1[n], \qquad \nu_{2,3}[n] = 2\overline{\tilde{b}}_0[n]b_1[n].
$$

Figure [12.1](#page-7-0) displays the wavelet packets from the first and the second decomposition scales with their SHA spectra.

Fig. 12.1 Wavelet packets of order 8 from the 1st (*top*) the 2nd decomposition scales with their SHA spectra (*bottom half-band*)

The wavelet packets are well localized in space. Their SHA spectra have a near rectangular shape (the higher the spline order is the closer the shape is to rectangular) and produce a sequence of partitions of the frequency band. The SHA spectrum of the wavelet packet $\psi_{r,l}(x)$ is effectively confined within the band

$$
A_{r,l} \stackrel{\text{def}}{=} \left[-\frac{(l+1)N}{2^{r+1}}, -\frac{(l)N}{2^{r+1}} \right] \cup \left[\frac{(l)N}{2^{r+1}}, \frac{(l+1)N}{2^{r+1}} \right], \quad l = 0, \dots, 2^r - 1. \tag{12.16}
$$

Since a spline from $\mathcal{P}_{r,l}$ is the linear combination of the wavelet packets

$$
S_{r,l}(x) = \sum_{k=0}^{N_r - 1} q_{r,l}[k] \psi_{r,l}^p(x - 2^r k), \qquad (12.17)
$$

then its SHA spectrum is effectively confined within the band *Λr,l*. This provides opportunities to approximate the heat inversion separately in different frequency bands.

12.5 Spline Wavelet Packet Transforms

Let a spline *S*(*x*) \in *PS* be represented by the orthonormal basis splines

$$
S(x) = \sqrt{\frac{1}{N}} \sum_{n=-N/2}^{N/2-1} \sigma[n] \gamma^{p}[n](x).
$$
 (12.18)

The sequence $\{\sigma[n]\}, n = -N/2, \ldots, N/2 - 1$, is the SHA spectrum of the spline *S(x)*. The space $P\mathscr{S}$ is the orthogonal sum of the subspaces $P\mathscr{S}_{1,0}$ and $P\mathscr{S}_{1,1}$ whose orthonormal bases are $\{\gamma_{1,0}^p[n](x)\}$ and $\{\gamma_{1,1}^p[n](x)\}$, respectively, where $n = 0, \ldots, N_1 - 1$ and $N_1 = N/2$. Thus, $S(x)$ can be represented as the sum of its orthogonal projections onto the subspaces ${}^p \mathcal{S}_{1,i}$, $i = 0, 1$: $S(x) = S_{1,0}(x) \oplus S_{1,1}(x)$, where

$$
S_{1,i}(x) \stackrel{\text{def}}{=} \sqrt{\frac{2}{N}} \sum_{n=0}^{N/2 - 1} \sigma_1[n] \gamma_{1,i}^p[n](x), \quad i = 0, 1. \tag{12.19}
$$

The orthonormality of the spline basis implies

$$
\sigma[n] = \sqrt{N} \langle S, \gamma^{p}[n] \rangle, \qquad \sigma_{1,i}[n] = \sqrt{N/2} \langle S, \gamma^{p}_{1,i}[n] \rangle, \quad i = 0, 1. \tag{12.20}
$$

By using the two-scale relations given by (12.12) and (12.13) , we derive for $n =$ $0, \ldots, N/2 - 1$

$$
\sigma_{1,0}[n] = \sqrt{N/2} \langle S, \gamma_{1,0}^p[n] \rangle = \sqrt{\frac{1}{2}} \big(b_0[n] \sigma[n] + b_0[n_1] \sigma[n_1] \big), \tag{12.21}
$$

$$
\sigma_{1,1}[n] = \sqrt{\frac{1}{2}(\tilde{b}_0[n]\sigma[n] + \tilde{b}_0[n_1]\sigma[n_1]), \quad n_1 = n + N/2.}
$$
 (12.22)

We can present (12.21) and (12.22) (12.22) (12.22) in a matrix form

$$
\begin{pmatrix}\n\sigma_{1,0}[n] \\
\sigma_{1,1}[n]\n\end{pmatrix} = \sqrt{\frac{1}{2}} \mathbf{A}_0[n] \cdot \begin{pmatrix}\n\sigma[n] \\
\sigma[n_1]\n\end{pmatrix}, \qquad \mathbf{A}_m[n] \stackrel{\text{def}}{=} \begin{pmatrix}\n\frac{b_m[n] b_m[n_{m+1}]}{b_m[n_{m+1}]}\n\end{pmatrix}.
$$
\n(12.23)

The coordinates of the projections of $S(x)$ onto the subspaces $P\mathscr{S}_{r,l}$ are derived iteratively:

$$
\begin{pmatrix}\n\sigma_{m,2l}[n] \\
\sigma_{m,2l+1}[n]\n\end{pmatrix} = \sqrt{\frac{1}{2}} \mathbf{A}_{m-1}[n] \cdot \begin{pmatrix}\n\sigma_{m-1,l}[n] \\
\sigma_{m-1,1}[n+N_r]\n\end{pmatrix} \quad \text{if } l \text{ is even,}
$$
\n
$$
\begin{pmatrix}\n\sigma_{m,2l+1}[n] \\
\sigma_{m,2l}[n]\n\end{pmatrix} = \sqrt{\frac{1}{2}} \mathbf{A}_{m-1}[n] \cdot \begin{pmatrix}\n\sigma_{m-1,l}[n] \\
\sigma_{m-1,1}[n+N_r]\n\end{pmatrix} \quad \text{if } l \text{ is odd.}
$$
\n(12.24)

12.6 Wavelet Packet Bases

Assume that the spline $S(x) \in P\mathscr{S}$ is expanded over the orthonormal bases

$$
S(x) = \frac{1}{\sqrt{N}} \sum_{n=0}^{N-1} \sigma[n] \gamma^{p}[n](x) = \sum_{k=0}^{N-1} q[k] \varphi^{p}(x - k),
$$

$$
\sigma[n] = \sum_{k=0}^{N-1} \omega^{-nk} q[k] = \hat{q}[n], \quad q[k] = \frac{1}{N} \sum_{n=0}^{N-1} \omega^{nk} \sigma[n].
$$

For example, if the samples $S(k) = y[k], k = 0, \ldots, N-1$, are available then [\(12.8\)](#page-3-1) claims that $\sigma[n]=\hat{y}[n]/V[n]$. Iterative application of the transform given by [\(12.24\)](#page-9-0) expands the projections splines $S_{r,l}(x) \in {}^p\mathscr{S}_{r,l}$ over the orthonormal bases $\{\gamma_{r,l}[n](x)\}\)$. Then, the coordinates $\{q_{r,l}[k]\}\$ of the alternative expan-sion ([12.17](#page-8-2)) over the orthonormal wavelet packet bases $\{\psi_{r,l}^p(x - 2^r k)\}\$ are derived by the application of the IDFT: $q_{r,l}[k] = N_r^{-1} \sum_{n=0}^{N_r-1} \omega^{2^r n k} \sigma_{r,l}[n]$. The subspace $\mathbb{P}\mathcal{S}_{r-1,l} = \mathbb{P}\mathcal{S}_{r,2l} \oplus \mathbb{P}\mathcal{S}_{r,2l+1}$. The spline $S_{r-1,l}(x)$ can be expanded either over the basis $\{\psi_{r-1,l}^p(x - 2^{r-1}k)\}_{k=0}^{N_{r-1}-1}$ or over the combined orthonormal basis $\{\psi_{r,2l}^p(x - 2^r k)\}\biguplus \{\psi_{r,2l}^p(x - 2^r k)\}_{k=0}^{N_r-1}$. The decision of which basis is preferable is made once a cost function is defined.

Consequently, once the wavelet packet transform of the spline $S(x)$ is implemented, a wide variety of orthonormal wavelet packet bases becomes available. A basis, which is optimal for a given spline with respect to a certain purpose, can be designed by the Best Basis algorithm [\[4](#page-17-5)], which compares the cost function of the "parent" spline *Sr*[−]1*,l(x)* with the cost of the "offsprings" *Sr,*2*l(x)* and *Sr,*2*l*+1*(x)*. Entropy us a typical cost function.

12.7 Parameterized Spline Solution in the Subspace $P\mathcal{S}_{r,l}$

12.7.1 Splines from the Subspaces $P\mathcal{S}_{r,l}$

The spline $S_{r,l}(x)$, which is the orthogonal projection of the spline $S(x)$ onto the subspace $P\mathcal{S}_{r,l}$, is expanded over the orthonormal wavelet packet basis as in [\(12.17\)](#page-8-2). On the other hand, the spline $S_{r,l}(x)$ belongs to the initial space $\mathbb{P}\mathscr{S}$ and can be expanded over the orthonormal basis of $P\mathscr{S}$

$$
S_{r,l}(x) = \sqrt{\frac{1}{N} \sum_{n=0}^{N-1} \zeta_{r,l}[n] \gamma^p[n](x)}, \quad \zeta_{r,l}[n] = \nu_{r,l}[n] \hat{q}_{r,l}[n]. \tag{12.25}
$$

The coefficients $v_{r,l}[n]$ are the SHA spectrum of the wavelet packet $\psi_{r,l}^p$. We emphasize that the DFT sequence $\hat{q}_{r,l}[n]$ is N_r -periodic, where $N_r = N/2^r$.

The projection coordinates ζ _{*r,l*}[*n*] can be expressed via the coordinates *σ*[*n*] of the spline $S(x)$.

Proposition 12.1 *The following representation of the projection coordinates holds*:

$$
\zeta_{r,l}[n] = \frac{\nu_{r,l}[n]}{2^r} \sum_{\lambda=-2^r/2}^{2^r/2-1} \sigma[n+\lambda N_r] \overline{\nu}_{r,l}[n+\lambda N_r] \approx \frac{1}{2^r} \sigma[n] |\nu_{r,l}[n]|^2. \quad (12.26)
$$

Remark 12.1 The higher the order *p* is, the closer $\zeta_{r,l}[n]$ is to $2^{-r}\sigma[n] |v_{r,l}[n]|^2$.

Equation ([12.9](#page-3-2)) implies that the application of the operator V_t to $S_{r,l}(x)$ results in

$$
S_{r,l}(x,t) = \mathbf{V}_t S_{r,l}(x) = \frac{1}{\sqrt{N}} \sum_{n=0}^{N-1} \eta[n](t) \zeta_{r,l}[n] \gamma^p[n](x).
$$
 (12.27)

From [\(12.6\)](#page-3-0), sampling of the spline $S(x, t) = V_t S(x)$ becomes

$$
y[k] \stackrel{\text{def}}{=} S(k,t) = \frac{1}{N} \sum_{n=0}^{N-1} \eta[n](t) \sigma[n] V[n] \omega^{kn}, \quad V[n] \stackrel{\text{def}}{=} \frac{u^p[n]}{\sqrt{u^{2p}[n]}},
$$

while sampling of $S_{r,l}(x, t) = V_t S_{r,l}(x)$

$$
y_{r,l}[k] \stackrel{\text{def}}{=} S_{r,l}(k,t) = \frac{1}{N} \sum_{n=0}^{N-1} \eta[n](t) \zeta_{r,l}[n] V[n] \omega^{kn}.
$$
 (12.28)

Equation ([12.26\)](#page-10-0) implies the approximated relation

$$
\hat{y}_{r,l}[n] \approx \frac{1}{2^r} \eta[n](t) \sigma[n] V[n] |v_{r,l}[n]|^2 = \frac{1}{2^r} \hat{y}[n] |v_{r,l}[n]|^2.
$$
 (12.29)

Remark 12.2 Equation ([12.29](#page-10-1)) can be interpreted in a sense that confinement of the operator's V_t domain from the whole spline space $P\mathscr{S}$ to the subspace $P\mathscr{S}_{r,l}$ effectively results in multiplication of the DFT $\hat{y}[n]$ of the sampled output $y[k]$ = $V_t S(k)$ with the factor $2^{-r} |v_r| [n] |^2$.

12.7.2 Parameterized Spline Solution

The scheme for a partial solution of Problem [12.1](#page-1-0) in the subspace $\mathbb{P}\mathcal{S}_{r,l}$ is very similar to the scheme of a global solution presented in Sect. [12.3](#page-3-3). By assumption, *t* is a known time parameter and the vector $\mathbf{z} = \{z[k] = g(k, t) + e_k\} = \mathbf{g} + \mathbf{e}$ is available, where $g(x, t) = U_t f(x)$. Then, a partial approximated inversion of the heat equation ([12.1](#page-0-0)) is derived as a spline

$$
S_{r,l}(x) = \sum_{k=0}^{N_r - 1} q_{r,l}[k] \psi_{r,l}^p(x - 2^r k) = \sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} \zeta_{r,l}[n] \gamma^p[n](x)
$$
(12.30)

such that the spline $S_{r,l}(x, t) = V_t S_{r,l}(x)$ approximates, in some sense, the available discrete data **z**. To be specific, Remark [12.2](#page-11-0) suggests that the sampled spline $S_{r,l}(k, t)$ should approximate the "filtered"

$$
\tilde{z}_{r,l}[k] \stackrel{\text{def}}{=} \frac{1}{N} \sum_{n=0}^{N-1} \omega^{kn} \hat{\tilde{z}}_{r,l}[n], \quad \text{where } \hat{\tilde{z}}_{r,l}[n] \stackrel{\text{def}}{=} \hat{z}[n] \frac{|\nu_{r,l}[n]|^2}{2^r}, \tag{12.31}
$$

rather than the entire data **z**. Similarly to Sect. [12.3](#page-3-3), we find a spline $S(\rho, x) \in \mathbb{P}\mathcal{S}_{r,l}$, which minimizes the functional $\rho I(S) + E_{r,l}(S)$, where

$$
I(S) \stackrel{\text{def}}{=} \|(S)''\|^2, \qquad E_{r,l}(S) \stackrel{\text{def}}{=} \sum_{k} (S(k,t) - \tilde{z}_{r,l}[k])^2, \qquad S(x,t) \stackrel{\text{def}}{=} \mathbf{V}_l S(x),
$$
\n(12.32)

and ρ is a numerical parameter.

Let a spline $s(x) \in \mathbb{P}\mathcal{S}_{r,l}$ be represented as in [\(12.30\)](#page-11-1)

$$
s(x) = \sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} \zeta_{r,l}[n] \gamma^{p}[n](x).
$$

Then,

$$
I(s) = \frac{1}{N} \sum_{n=0}^{N-1} W[n] |\zeta_{r,l}[n]|^2, \qquad E_{r,l}(s) = \frac{1}{N} \sum_{n=0}^{N-1} |\eta[n](t) \zeta_{r,l}[n] V[n] - \hat{\zeta}_{r,l}[n]|^2.
$$

The sequences $W[n]$ and $V[n]$ are defined in ([12.5](#page-3-4)).

A solution to the minimization problem is the spline from $P\mathcal{S}_{r,l}$

$$
S_{r,l}(\rho, x) = \sqrt{\frac{1}{N}} \sum_{n=0}^{N-1} \zeta_{r,l}(\rho) [n] \gamma^{p}[n](x), \qquad \zeta_{r,l}(\rho) [n] = \frac{\overline{\eta}[n](t) V[n] \hat{\zeta}_{r,l}(n)}{A[n](\rho)},
$$

\n
$$
A[n](\rho) \stackrel{\text{def}}{=} \rho W[n] + |\eta[n](t) V[n]|^{2}.
$$

Its samples on the grid points

$$
S_{r,l}(\rho,k) = \frac{1}{N} \sum_{n=0}^{N-1} \omega^{kn} \zeta_{r,l}(\rho) [n] V[n] = \frac{1}{N} \sum_{n=0}^{N-1} \omega^{kn} \frac{\overline{\eta}[n](t) V^2[n] \hat{\zeta}_{r,l}(n)}{A[n](\rho)}.
$$
 (12.33)

12.7.3 Selection of the Regularization Parameter

Assume that we are able to evaluate the errors vector $\mathbf{e} = \{e_k\}_{k=0}^{N-1}$, $e_k = N^{-1} \times$ $\sum_{n=0}^{N-1} \omega^{kn} \hat{e}[n]$, whose variance var $(\mathbf{e}) = \varepsilon^2 \approx N^{-1} \sum_{k=0}^{N-1} (e_k)^2$. Keeping ([12.31](#page-11-2)) in mind, denote

$$
e_{r,l}[k] \stackrel{\text{def}}{=} \frac{1}{N} \sum_{n=0}^{N-1} \omega^{kn} \hat{e}_{r,l}[n], \quad \text{where } \hat{e}_{r,l}[n] \stackrel{\text{def}}{=} \hat{e}[n] \frac{|\nu_{r,l}[n]|^2}{2^r},
$$

$$
(\varepsilon_{r,l})^2 \stackrel{\text{def}}{=} \sum_{k=0}^{N-1} (e_{r,l}[k])^2 = \frac{1}{N} \sum_{n=0}^{N-1} |\hat{e}_{r,l}[n]|^2.
$$

The function

$$
e_{r,l}(\rho) \stackrel{\text{def}}{=} E_{r,l}(S_{r,l}(\rho,\cdot)) = \frac{1}{N} \sum_{n=0}^{N-1} \left(\frac{\rho W[n] |\hat{\tilde{z}}_{r,l}[n]|}{A[n](\rho)} \right)^2
$$

grows monotonically from zero to $N^{-1} \sum_{n=0}^{N-1} |\hat{z}_{r,l}[n]|^2 = \sum_{k=0}^{N-1} (\tilde{z}_{r,l}[k])^2$ as ρ grows from zero to infinity. Therefore, we derive $\rho_{r,l}$ from the equation $e_{r,l}(\rho)$ = *(εr,l)*2.

12.7.4 Modeling the Noise

We assume that the error vector **e** is a zero mean Gaussian white noise. It is seen from [\(12.2\)](#page-1-1) that the Fourier coefficients of the function $g(x, t) = U_t f(x)$: $g_n(t) = f_n e^{-t(2\pi n/N)^2}$ are fast decaying when *n* is growing. Thus, the function $g(x, t)$ is efficiently bandlimited. Its significant Fourier coefficients $g_n(t)$ occupy

a relatively narrow band around zero, $-K(t) < n < K(t)$, $K(t) < N/2$, $K(t) \rightarrow 0$ as $t \to \infty$. Hence, the DFT coefficients of the data vector **z**: { $\hat{z}[n]$ } \approx { $\hat{e}[n]$ } as $n \in$ $[K(t), N/2 - 1] \cup [-N/2, -K(t)]$. By relying on the fact that the power spectrum $\{|\hat{e}[n]|^2\}$ of the white noise **e** is close to a constant for all $n = -N/2, \ldots, N/2 - 1$, it is possible to evaluate the variance

$$
\sigma^2 \approx \frac{1}{(N - K(t))^2} \sum_{n \in [K(t), N/2 - 1] \cup [-N/2, -K(t)]} |\hat{z}[n]|^2.
$$
 (12.34)

Then, the noise vector **e** is modeled as a zero mean Gaussian random process $\tilde{\mathbf{e}} =$ ${\{\tilde{e}_i\}}_{i=0}^{N-1}$, whose variance is σ^2 . Let ${\{\hat{e}[n]\}}_{n=-N/2}^{N/2-1}$ be the DFT spectrum of the model vector $\tilde{\mathbf{e}}$. Then, the values $(\varepsilon_{r,l})^2$, which are needed for the parameter ρ selection, are estimated as

$$
(\varepsilon_{r,l})^2 \approx \frac{1}{2^r N} \sum_{n=0}^{N-1} |(v_{r,l}[n])^2 \overline{\tilde{e}}[n]|^2.
$$
 (12.35)

Another option for the noise evaluation is to use the scheme in [\[5](#page-17-6)].

12.8 Spline Wavelet Packet Solution to Problem [12.1](#page-1-0)

The partial spline solution $S_{m,l}(\rho, x)$ of the inversion problem in the subspace $^p\mathscr{S}_{r,l}$ is derived from the filtered data such that the DFT is $\hat{\tilde{z}}_{m,l}[n] \stackrel{\text{def}}{=} \hat{z}[n] |v_{m,l}[n]|^2 2^{-m}$.

To determine an optimal set of the subspaces $P\mathscr{S}_{r,l}$, which reveal the internal structure of the data vector **z**, we construct the spline $Z(x) = \sum_{n=0}^{N-1} \xi[n]\gamma^{p}[n](x)$, $\xi[n]=\hat{\zeta}[n]/V[n]$, which interpolates the data **z**. Then, we apply the Best Basis algorithm to obtain the list $PL = \{(\bar{p}, \bar{l})\}$ such that the shifts of the wavelet packets $\psi_{\bar{m}\bar{j}}$ form an optimal basis for the spline $Z(x)$. The list *PL* determines the subspaces $P\mathcal{S}^{\tilde{m}, \tilde{l}}$, where the partial solutions for the inversion problem are to be derived. Due to the effective bandlimitedness of the function $g(x, t) = U_t^1 f(x)$, some subspaces $P \mathcal{S}^{\bar{m}, \bar{l}}$, which correspond to higher frequency bands are "empty" in a sense that they, actually, do not contain a contribution from the initial function *f (x)*. Such subspaces are discarded from the list *PL*.

A scheme for the approximated inversion of the heat equation

- 1. Calculate the coefficients $\eta[n](t)$ defined in ([12.10](#page-4-1)).
- 2. Construct the data interpolating spline $Z(x)$.
- 3. Implement the wavelet packet transform of order *p* of the spline *Z(x)*.
- 4. Apply the Best Basis algorithm to the transform coefficients to collect the list *PL* of relevant subspaces.
- 5. Reduce the list *PL* to *PL* by discarding the "empty" subspaces.
- 6. Evaluate the error vector to estimate the partial variances $(\varepsilon_{p,l})^2$, $(p,l) \in \overline{PL}$, of noise (see ([12.35](#page-13-0))).
- 7. Determine the optimal values $\rho_{\bar{m} \bar{l}}$ of the regularization parameter for each pair $(\bar{m}, \bar{l}) \in \overline{PL}$.
- 8. Find the partial solutions $S_{\overline{m}}$ $\overline{I}(\rho_{\overline{m}} l, x) \in {^p \mathscr{S}_{\overline{m}} l}$ for each pair $(\overline{m}, \overline{l}) \in \overline{PL}$ (see ([12.33\)](#page-12-0)).

The approximated solution to the inversion Problem [12.1](#page-1-0) is

$$
f(x) \approx S(x) = \sum_{(\overline{m}, \overline{l}) \in \overline{PL}} S_{\overline{m}, \overline{l}}(\rho_{\overline{m}, \overline{l}}, x) \in {^p \mathscr{\tilde{S}}}.
$$

Extension of the algorithm to the 2D case is straightforward once the tensor products of the basis splines are utilized:

$$
\gamma^{p}(x, y) \stackrel{\text{def}}{=} \gamma^{p}(x)\gamma^{p}(y), \qquad \varphi^{p}(x, y) \stackrel{\text{def}}{=} \varphi^{p}(x)\varphi^{p}(y),
$$

$$
\psi^{p}_{r,l,\tilde{l}}(x, y) \stackrel{\text{def}}{=} \psi^{p}_{r,l}(x)\psi^{p}_{r,\tilde{l}}(y).
$$

Figure [12.2](#page-15-0) displays the SHA spectra of two wavelet packets of order 10 from the second scale. We observe that the spectra have near-parallelepiped shape. The described algorithm can be utilized for signal and image denoising when the time parameter $t = 0$. In this case, the general scheme remains unchanged.

12.9 Numerical Examples

The following are examples, derived from three groups of experiments, on using the block-based methods for 2D images' restoration:

- Denoising: Restoration of objects corrupted by Gaussian noise (the time parameter $t = 0$).
- Pure blurred input: Restoration of blurred objects when the time parameter $t > 0$ and noise is not known. The advantage of the block based method over the global one materialized in the accurate tuning of the subspaces where the looked for solution is in the effective frequency domain of the blurred image.
- Noised blurred input: Restoration of objects from blurred inputs, which were corrupted by Gaussian noise.

These examples illustrate the difference between the performance of the global Tikhonov algorithm (GTA) presented in Sect. [12.3](#page-3-3) and of the Best Basis Algorithm (BBA). Visual perception is compared and the peak-signal-to-noise-ratio (PSNR). Three benchmark images each of which is presented by a 512×512 array of samples are used as the initial temperature distributions. The source images for the experiments are shown in Fig. [12.3.](#page-16-0)

Example 12.3 (Barbara Denoising) The "Barbara" image was corrupted by Gaussian zero-mean noise with standard deviations $STD = 25$. The time parameter is

Fig. 12.2 The SHA spectra of wavelet packets of order 10 from the second resolution scale. *Top*: $\psi_{2,2,3}^{10}(x, y)$. *Bottom*: $\psi_{2,3,1}^{10}(x, y)$

 $t = 0$. Figure [12.4](#page-16-1) displays fragments of the noised input image and of the image that was restored by the applications of GTA and BBA. We observe that BBA produces high PSNR values. The noise was suppressed almost completely. The GTA method did not succeed in noise suppression, although the texture is resolved a little bit better in comparison to BBA.

Fig. 12.3 *Left*: "Barbara". *Center*: "Lena". *Right*: Fingerprint

Fig. 12.4 "Barbara". *Left*: A noised image, STD = 25, PSNR = 20*.*17. *Center*: An image restored by GTA, PSNR = 24*.*12. *Right*: An image restored by BBA, PSNR = 25*.*77, spline wavelet packets of the fourth order from 4 levels were used

Fig. 12.5 "Fingerprint". *Left*: A blurred image, $t = 46$, PSNR = 15.71. *Center*: An image restored by GTA, PSNR = 17*.*57. *Right*: An image restored by BBA, PSNR = 20*.*22, spline wavelet packets of the fourth order from 3 levels were used

Example 12.4 (Restoration of a Strongly Blurred Fingerprint) In this example, the "Fingerprint" image was used as the initial temperature distribution. The input presents the temperature distribution when the time parameter was $t = 46$. The BBA restored the texture of the fingerprint, which was completely smeared in the input. The result produced by GTA was much worse. The results are illustrated in Fig. [12.5.](#page-16-2)

Example 12.5 (Restoration of Blurred and Noised "Lena") The "Lena" image was used as the initial temperature distribution. The input is the distribution when the time parameter was $t = 2.5$ corrupted by Gaussian noise whose $STD = 10$. The BBA-restored image is sharper compared to the GTA and its PSNR is higher. See Fig. [12.6](#page-17-7) for the results.

Fig. 12.6 "Lena". *Left*: A blurred noised image, $t = 2.5$, noise STD = 10, PSNR = 24.79. *Center*: An image restored by GTA, PSNR = 28*.*22. *Right*: An image restored by BBA, PSNR = 28*.*47, spline wavelet packets of the fourth order from 4 levels were used

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