

Nikolay Ivanov Kolev

# Multiphase Flow Dynamics

**1** FUNDAMENTALS

Fourth Edition



Springer

# Multiphase Flow Dynamics 1

Nikolay Ivanov Kolev

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Fundamentals

 Springer

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*To Iva, Rali and Sonja with love!*



*Rügen, July, 2004, Nikolay Ivanov Kolev, 48×36 cm, oil on linen*



*Nikolay Ivanov Kolev, PhD, DrSc  
Born 1.8.1951, Gabrowo, Bulgaria*

# **A Few Words about the New Editions of Volumes 1 through 5**

The present content and format of the fourth, improved, and extended edition of Volumes 1, 2, and 3, and the second improved and extended edition of Volumes 4 and 5 were achieved after I received many communications from all over the world from colleagues and friends commenting on different aspects or requesting additional information. Of course, misprints and some layout deficiencies in the previous editions, for which I apologize very much, have also been removed, as is usual for subsequent editions of such voluminous 3000-page monographs. I thank everyone who contributed in this way to improving the five volumes!

The new editions contain my experiences in different subjects, collected during my daily work in this field since 1975. They include my own new results and the new information collected by colleagues since the previous editions. The overwhelming literature in multiphase fluid dynamics that has appeared in the last 40 years practically prohibits a complete overview by a single person. This is the reason why, inevitably, one or other colleague may feel that his personal scientific achievements are not reflected in this book, for which I apologize very much. However, it is the responsibility of transferring knowledge to the next generation that drove me to write these, definitely not perfect, books. I hope that they will help young scientists and engineers to design better facilities than those created by my generation.

29.12.2010  
Herzogenaurach

Nikolay Ivanov Kolev

# Introduction

Multiphase flows, such as rainy or snowy winds, tornadoes, typhoons, air and water pollution, volcanic activities, etc., see Fig.1, are not only part of our natural environment but also are working processes in a variety of conventional and nuclear power plants, combustion engines, propulsion systems, flows inside the human body, oil and gas production, and transport, chemical industry, biological industry, process technology in the metallurgical industry or in food production, etc.



**Fig. 1** The fascinating picture of the start of a discovery, a piece of universe, a tornado, a volcano, flows in the human heart, or even the “pure” water or the sky in Van Gogh’s painting are, in fact, different forms of multiphase flows

The list is by far not exhaustive. For instance, everything to do with phase changes is associated with multiphase flows. The industrial use of multiphase systems requires methods for predicting their behavior. This explains the “explosion” of scientific publications in this field in the last 50 years. Some countries, such as Japan, have declared this field to be of strategic importance for future technological development.

Probably the first known systematic study on two-phase flow was done during the Second World War by the Soviet scientist *Teletov* and published later in 1958 with the title “On the problem of fluid dynamics of two-phase mixtures”. Two books that appeared in Russia and the USA in 1969 by *Mamaev* et al. and by *Wallis* played an important role in educating a generation of scientists in this discipline, including me. Both books contain valuable information, mainly on steady state flows in pipes. In 1974 *Hewitt* and *Hall-Taylor* published the book “Annular two-phase flow”, which also considers steady state pipe flows. The usefulness of the idea of a three-fluid description of two-phase flows was clearly demonstrated

on annular flows with entrainment and deposition. In 1975 *Ishii* published the book “Thermo-fluid dynamic theory of two-phase flow”, which contained a rigorous derivation of *time*-averaged conservation equations for the so called two-fluid separated and diffusion momentum equations models. This book founded the basics for new measurement methods appearing on the market later. The book was updated in 2006 by *Ishii* and *Hibiki* who included new information about the interfacial area density modeling in one-dimensional flows, which had been developed by the authors for several years. *R. Nigmatulin* published “Fundamentals of mechanics of heterogeneous media” in Russian in 1978. The book mainly considers one-dimensional two-phase flows. Interesting particular wave dynamics solutions are obtained for specific sets of assumptions for dispersed systems. The book was extended mainly with mechanical interaction constitutive relations and translated into English in 1991. The next important book for two-phase steam-water flow in turbines was published by *Deich* and *Philipoff* in 1981, in Russian. Again, mainly steady state, one-dimensional flows are considered. In the same year *Delhaye* et al. published “Thermohydraulics of two-phase systems for industrial design and nuclear engineering”. The book contains the main ideas of local volume averaging and considers mainly many steady state one-dimensional flows. One year later, in 1982, *Hetsroni* edited the “Handbook of multiphase systems”, which contained the state of the art of constitutive interfacial relationships for practical use. The book is still a valuable source of empirical information for different disciplines dealing with multiphase flows. In 2006 *Crowe* 2006 edited the “Multiphase flow handbook”, which contained an updated state of the art of constitutive interfacial relationships for practical use. In the monograph “Interfacial transport phenomena” published by *Slattery* in 1990 complete, rigorous derivations of the local volume-averaged two-fluid conservation equations are presented together with a variety of aspects of the fundamentals of the interfacial processes based on his long years of work. *Slattery*’s first edition appeared in 1978. Some aspects of the heat and mass transfer theory of two-phase flow are now included in modern textbooks such as “Thermodynamics” by *Baer* (1996) and “Technical thermodynamics” by *Stephan* and *Mayinger* (1998).

It is noticeable that none of the above mentioned books is devoted in particular to *numerical methods* of solution of the fundamental systems of partial differential equations describing multiphase flows. Analytical methods still do not exist. In 1986 I published the book “Transient two-phase flows” with Springer-Verlag, in German, and discussed several engineering methods and practical examples for integrating systems of partial differential equations describing two- and three-fluid flows in pipes.

Since 1984 I have worked intensively on creating numerical algorithms for describing complicated multiphase multicomponent flows in pipe networks and complex *three-dimensional* geometries mainly for nuclear safety applications. Note that the mathematical description of multidimensional two-phase and multiphase flows is a scientific discipline that has seen considerable activity in the last 30 years. In addition, for yeas thousands of scientists have collected experimental information in this field. However, there is still a lack of systematic presentation of the theory and practice of *numerical multiphase fluid dynamics*. This book is intended to fill this gap.

Numerical multiphase fluid dynamics is the science of the derivation and the numerical integration of the conservation equations reflecting the mass momentum and energy conservation for multiphase processes in nature and technology at different scales in time and space. The emphasis of this book is on the generic links within computational predictive models between

- fundamentals,
- numerical methods,
- empirical information about constitutive interfacial phenomena, and
- a comparison with experimental data at different levels of complexity.

The reader will realize how strong the mutual influence of the four model constituencies is. There are still many attempts to attack these problems using single-phase fluid mechanics by simply extending existing single-phase computer codes with additional fields and linking with differential terms outside the code without increasing the strength of the feedback in the numerical integration methods. The success of this approach in describing low concentration suspensions and dispersed systems without strong thermal interactions should not confuse the engineer about the real limitations of this method.

This monograph can also be considered as a handbook on the numerical modeling of three strongly interacting fluids with dynamic fragmentation and coalescence representing multiphase multicomponent systems. Some aspects of the author's ideas, such as the three-fluid entropy concept with dynamic fragmentation and coalescence for describing multiphase, multicomponent flows by local volume-averaged and time-averaged conservation equations, were published previously in separate papers but are collected here in a single context for the first time. An important contribution of this book to the state of the art is also the rigorous thermodynamic treatment of multiphase systems consisting of different mixtures. It is also the first time that the basics of the boundary fitted description of multiphase flows and an appropriate numerical method for integrating them with proven convergence has been published. It is well known in engineering practice that "the devil is hidden in the details". This book gives many hints and details on how to design computational methods for multiphase flow analysis and demonstrates the power of the method in the attached compact disc and in the last chapter in Volume 2 by presenting successful comparisons between predictions and experimental data or analytical benchmarks for a class of problems with a complexity not known in the multiphase literature until now. It starts with the single-phase U-tube problem and ends with explosive interaction between molten melt and cold water in complicated 3D geometry and condensation shocks in complicated pipe networks containing acoustically interacting valves and other components.

Volume 3 is devoted to selected subjects in multiphase fluid dynamics that are very important for practical applications but could not find place in the first two volumes of this work.

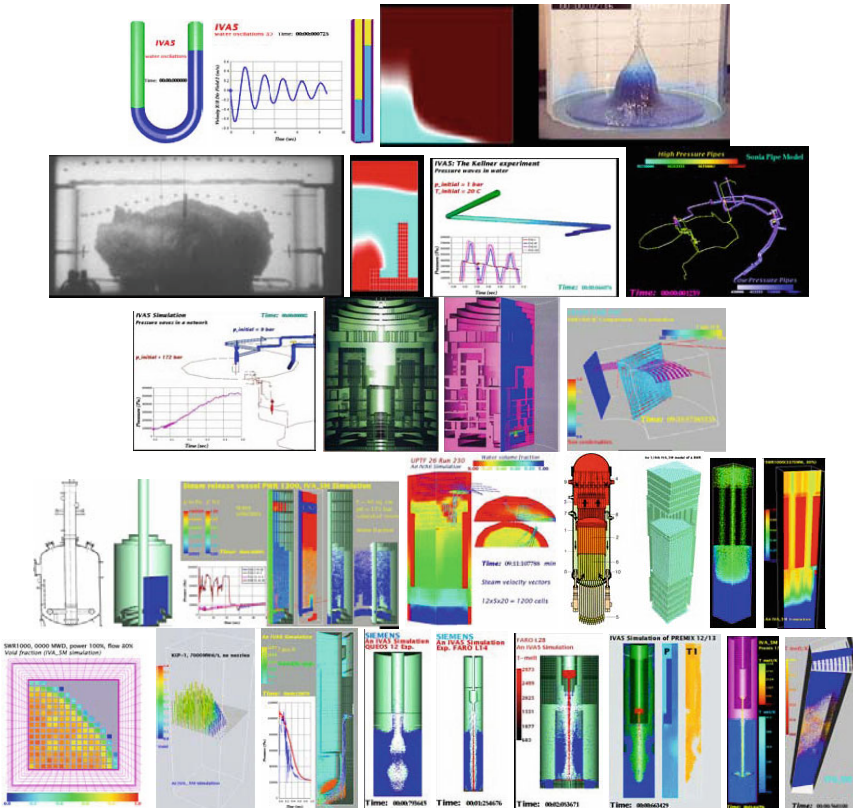
The state of the art of turbulence modeling in multiphase flows is also presented. As an introduction, some basics of the single-phase boundary layer theory, including some important scales and flow oscillation characteristics in pipes and rod bundles are presented. Then the scales characterizing the dispersed flow systems are presented. The description of the turbulence is provided at different levels of complexity: simple algebraic models for eddy viscosity, algebraic models based on the *Boussinesq* hypothesis, modification of the boundary layer share due to modification of the bulk turbulence, and modification of the boundary layer share due to nucleate boiling. Then the role of the following forces on the mathematical description of turbulent flows is discussed: the lift force, the lubrication force in the wall boundary layer, and the dispersion force. A pragmatic generalization of the *k-eps* models for continuous velocity fields is proposed, which contains flows in large volumes and flows in porous structures. Its large eddy simulation variant is also presented. A method to derive source and sinks terms for multiphase *k-eps* models is presented. A set of 13 single- and two-phase benchmarks for verification of *k-eps* models in system computer codes are provided and reproduced with the IVA computer code as an example of the application of the theory. This methodology is intended to help engineers and scientists to introduce this technology step by step in their own engineering practice.

In many practical applications gases are dissolved in liquids under given conditions, released under other conditions, and therefore affect technical processes for good or for bad. There is almost no systematic description of this subject in the literature. That is why I decided to collect in Volume 3 useful information on the solubility of oxygen, nitrogen, hydrogen, and carbon dioxide in water, valid within large intervals of pressures and temperatures, provide appropriate mathematical approximation functions, and validate them. In addition, methods for computation of the diffusion coefficients are described. With this information solution and dissolution dynamics in multiphase fluid flows can be analyzed. For this purpose, the nonequilibrium absorption and release on bubble, droplet, and film surfaces under different conditions are mathematically described.

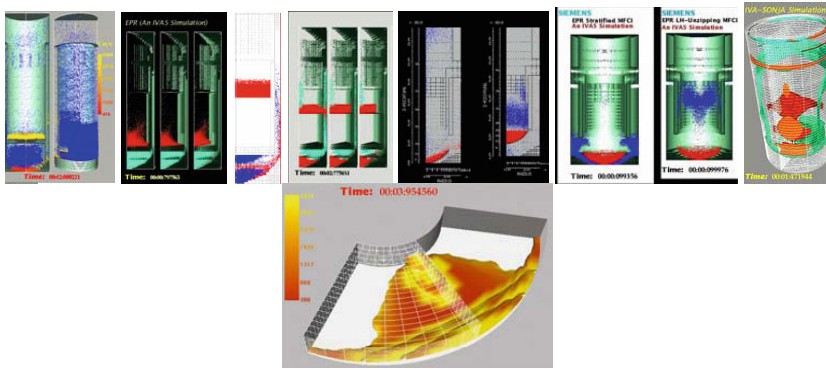
Volume 4 is devoted to nuclear thermal hydraulics, which is a substantial part of nuclear reactor safety. It provides knowledge and mathematical tools for the adequate description of the process of transferring the fission heat released in materials due to nuclear reactions into its environment. The heat release inside the fuel, the temperature fields in the fuels, and the “simple” boiling flow in a pipe, are introduced step by step, using ideas of different complexity like equilibrium, nonequilibrium, homogeneity, and nonhomogeneity. Then the “simple” three-fluid boiling flow in a pipe is described by gradually involving mechanisms like entrainment and deposition, dynamic fragmentation, collisions, and coalescence, turbulence. All heat transfer mechanisms are introduced gradually and their uncertainty is discussed. Different techniques like boundary layer treatments or integral methods are introduced. Comparisons with experimental data at each step demonstrate the success of the different ideas and models. After an introduction into the design of reactor pressure vessels for pressurized and boiling water reactors, the accuracy of modern methods is demonstrated using a large number of experimental data sets for steady and transient flows in heated bundles. Starting with single



pipe boiling going through boiling in a rod bundles the analysis of the complete vessel, including the reactor, is finally demonstrated. Then a powerful method for nonlinear stability analysis of flow boiling and condensation is introduced. Models are presented and their accuracies in describing critical multiphase flow at different level of complexity are investigated. The basics of the design of steam generators, moisture separators, and emergency condensers are presented. Methods for analyzing a complex pipe network flows with components like pumps, valves, etc., are also presented. Methods for the analysis of important aspects of severe accidents like melt-water interactions and external cooling and cooling of layers of molten nuclear reactor material are presented. Valuable sets of thermophysical and transport properties for severe accident analysis are presented for the following materials: uranium dioxide, zirconium dioxide, stainless steel, zirconium, aluminum, aluminum oxide, silicon dioxide, iron oxide, molybdenum, boron oxide, reactor corium, sodium, lead, bismuth, and lead-bismuth eutectic alloy. The emphasis is on the complete and consistent thermodynamical sets of analytical approximations appropriate for computational analysis. Thus the book presents a complete coverage of modern nuclear thermal hydrodynamics.







**Fig. 2** Examples of multiphase flows in nuclear technology. See <http://www.herzovision.de/kolev-nikolay/>

Herzogenaurach, Winter 2010

Nikolay Ivanov Kolev

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# Summary

This monograph contains theory, methods, and practical experience for describing complex transient multiphase processes in arbitrary geometrical configurations. It is intended to help applied scientists and practicing engineers to better understand natural and industrial processes containing dynamic evolutions of complex multiphase flows. The book is also intended to be a useful source of information for students in higher semesters and in PhD programs.

This monograph consists of five volumes:

Vol. 1 Fundamentals, 4<sup>th</sup> ed. (14 chapters and 2 appendixes), 782 pages.

Vol. 2 Mechanical interactions, 4<sup>th</sup> ed. (11 chapters), 364 pages,

Vol. 3 Thermal interactions, 4<sup>th</sup> ed. (16 chapters), 678 pages.

Vol. 4 Turbulence, gas absorption and release by liquid, diesel fuel properties, 2<sup>nd</sup> ed. (13 chapters), 328 pages.

Vol. 5 Nuclear thermal hydraulics, 2<sup>nd</sup> ed. (17 chapters), 848 pages.

In Volume 1 the concept of three-fluid modeling is presented in detail “from the origin to the applications”. This includes the derivation of local volume- and time-averaged equations and their working forms, the development of methods for their numerical integration, and finally a variety of solutions for different problems of practical interest.

Special attention is paid in Volume 1 to the link between partial differential equations and constitutive relations, called closure laws, without providing any information on the closure laws.

Volumes 2 and 3 are devoted to these important constitutive relations for the mathematical description of the mechanical and thermal interactions. The structure of the volumes is, in fact, a state of the art review and a selection of the best available approaches for describing interfacial transfer processes. In many cases, the original contribution of the author is incorporated in the overall presentation. The most important aspects of the presentation are that they stem from the author’s long years of experience in developing computer codes. The emphasis is on the practical use of these relationships: either as stand-alone estimation methods or within a framework of computer codes.

Volume 4 is devoted to turbulence in multiphase flows.

Nuclear thermal hydraulics is the science providing knowledge about the physical processes occurring during the transfer of the fission heat released in structural materials due to nuclear reactions into its environment. Along its way to the environment thermal energy is organized to provide useful mechanical work or useful heat, or both. Volume 5 is devoted to the nuclear thermal hydraulics. In a way this

is the most essential application of multiphase fluid dynamics in analyzing steady and transient processes in nuclear power plants.

In particular in Volume 1, Chapters 1, 2, 3, and 5, the concept of three-fluid modeling is introduced. Each field consists of multicomponents grouped into an inert and a noninert components group. Each field has its own velocity in space and its own temperature, allowing mechanical and thermodynamic nonequilibrium among the fields. The idea of dynamic fragmentation and coalescence is introduced. Using the *Slattery–Whitaker* local spatial averaging theorem and the *Leibnitz* rule, the local volume-averaged mass, momentum and energy conservation equations are rigorously derived for heterogeneous porous structures. Successively time averaging is performed. A discussion is provided on particle size spectra and averaging, cutting off the lower part of the spectrum due to mass transfer, the effect of the averaging on the effective velocity difference, etc. Chapter 1 also contains brief remarks on the kinematic velocity of density wave propagation in porous structures and on the diffusion term of void propagation in the case of pooling all the mechanical interactions in this kind of formalism. In the derivation of the momentum equations special attention is paid to rearranging the pressure surface integrals in order to demonstrate the physical meaning of the originating source terms in the averaged systems and their link to hyperbolicity. The Reynolds stress concept is introduced for multiphase flows. Chapter 2 also contains a collection of constitutive relations for lift- and virtual mass forces, for wall boundary layer forces, for forces causing turbulent diffusion, and for forces forcing the rejection of droplet deposition on a wall with evaporation.

Before deriving the energy conservation in Chapter 5, I provide Chapter 3 in which it is shown how to generate thermodynamic properties and the substantial derivatives for different kinds of mixtures by knowing the properties of the particular constituents. It contains the generalization of the theory of the equations of states for arbitrary real mixtures. With one and the same formalism a mixture of miscible and immiscible components in arbitrary solid, liquid, or gaseous states mixed and/or dissolved can be treated. This is a powerful method for creating a universal flow analyzer. Chapter 3 contains additional information on the construction of the saturation line by knowing pressure or temperature. An application of the material given in Chapter 3 is given in the new Volume 3 of this work to diesel fuel, where an inherently consistent set of equations of state for both gas and liquid is formulated. In addition, a section defining the equilibrium of gases dissolved in liquids is provided. These basics are then used in Volume 3 to construct approximations for the equilibrium solution concentrations of  $H_2$ ,  $O_2$ ,  $N_2$  and  $CO_2$  in water and to describe the nonequilibrium solution and dissolution at bubble, droplet, and film interfaces, which extend the applicability of the methods of multiphase fluid dynamics to flows with nonequilibrium solution and dissolution of gases. The generalizations of Chapter 3 are also used in Chapter 17 of Volume 4 to represent a variety of thermal properties including sodium vapor properties. An additional appendix to Chapter 3 shows a table where the partial derivatives of different forms of the equation of state is provided. This chapter provides the information necessary to understand the entropy concept, which is presented in Chapter 5.

In the author's experience understanding the complex energy conservation for multiphase systems and especially the entropy concept is very difficult for most students and practicing engineers. This is why Chapter 4 is provided as an introduction, showing the variety of notations of the energy conservation principle for single-phase multicomponent flows. Chapter 4 further contains a careful state of the art review for the application of the method of characteristics for modeling 1D and 2D flows in engineering practice.

The local volume-averaged and time-averaged energy conservation equation is derived in Chapter 5 in different notational forms in terms of specific internal energy, specific enthalpy, specific entropy, and temperatures. The introduction of the entropy principle for such complex systems is given in detail in order to enable the practical use of the entropy concept. The useful "conservation of volume" equation is also derived. Chapter 5 contains an additional example of the computation of irreversible viscous dissipation in the boundary layer. For easy application additional sections have been added to Chapter 5, which contain the different notations of energy conservation for lumped parameter volumes and steady state flows. The limiting case with gas flow in a pipe is considered in order to show the important difference to the existing gas dynamics solution where the irreversible heat dissipation due to friction is correctly taken into account.

Examples for a better understanding are given for the simple cases of lumped parameters – Chapter 6, infinite heat exchange without interfacial mass transfer, discharge of gas from a volume, injection of inert gas in a closed volume initially filled with inert gas, heat input in a gas in a closed volume, steam injection in a steam–air mixture, chemical reaction in a gas mixture in a closed volume, and hydrogen combustion in an inert atmosphere. Chapter 6 has been extended with cases including details of the modeling of combustion and detonation of hydrogen by taking into account the equilibrium dissociation.

The exergy for a multiphase, multicomponent system is introduced in Chapter 7 and discussed for the example of judging the efficiency of a heat pump.

Simplification of the resulting system of PDEs to the case of one-dimensional flow is presented in Chapter 8. Some interesting aspects of fluid structure coupling, such as pipe deformation due to temporal pressure change in the flow and forces acting on the internal pipe walls are discussed. The idea of algebraic slip is presented. From the system thus obtained the next step of simplification leads to the system of ordinary differential equations describing the critical multiphase, multicomponent flow by means of three velocity fields. Modeling of valves and pumps is discussed in the context of the modeling of networks consisting of pipes, valves, pumps, and other different components.

Another case of simplification of the theory of multiphase flows is presented in Chapter 9, where the theory of continuum sound waves and discontinuous shock waves for melt-water interaction is presented. In order to easily understand it, the corresponding theory for single- and two-phase flows is reviewed as an introduction. Finally, an interesting application for the interaction of molten uranium and aluminum oxides with water, as well of the interaction of molten iron with water is presented. Chapter 9 also deals with detonation during melt-water interaction. To better put this information into the context of the detonation theory, additional

introductory information is given for the detonation of hydrogen in closed pipes, taking into account the dissociation of the generated steam.

Chapter 10 is devoted to the derivation of the conservation equations for multiphase, multicomponent, multivelocity field flow in general curvilinear coordinate systems. For a better understanding of the mathematical basics used in this chapter two appendixes are provided: Appendix 1 in which a brief introduction to vector analysis is given and Appendix 2 in which the basics of the coordinate transformation theory are summarized.

A new Chapter 11 gives the mathematical tools for computing eigenvalues and eigenvectors and for determination of the type of systems of partial differential equations. The procedure for the transformation of a hyperbolic system into canonical form is also provided. Then the relations between eigenvalues and critical flow and between eigenvalues and the propagation velocity of small perturbations are briefly defined. This is, in fact, a translation of one chapter of my first book published in German by Springer in 1986. This completes the basics of the multiphase, multicomponent flow dynamics.

Chapter 12 describes numerical solution methods for different multiphase flow problems. The first-order donor-cell method is presented in detail by discretizing the governing equations, creating a strong interfacial velocity coupling, and strong pressure-velocity coupling. Different approximations for the pressure equations are derived and three different solution methods are discussed in detail. One of them is based on the Newton iterations for minimizing the residuals by using the conjugate gradients. A method for temperature inversion is presented. Several details are given, which enables scientists and engineers to use this chapter for their own computer code development, such as the integration procedure (implicit method), the time step, and accuracy control. Finally, some high-order discretization schemes for convection-diffusion terms such as space exponential scheme and other high-order up-winding schemas are presented. Different analytical derivations are provided in Appendixes 12.1–12.8, including the analytical derivatives of the residual error of each equation with respect to the dependent variables. Some important basic definitions that are required for describing pipe networks are introduced. In addition, the variation of volume-porosity with time is systematically incorporated into the numerical formalism.

Chapter 13 presents a numerical solution method for multiphase flow problems in multiple blocks of curvilinear coordinate systems, generalizing, in fact, the experience gained in Chapter 12. Several important details of how to derive explicit pressure equations are provided. The advantage of using orthogonal grids also is easily derived from this chapter. Appendixes 1 and 2 of Volume I contain some additional information about orthogonal grid generation.

Chapter 14 provides several numerical simulations as illustrations of the power of the methods presented in this monograph. A compact disc that contains films corresponding to particular cases discussed in this chapter is attached. The films can be played with any tool capable of accepting *avi*- or animated *gif*-files.

As has already been mentioned, Volumes 2 and 3 are devoted to the so called closure laws: the important constitutive relations for mechanical and thermal interactions. The structure of the volume has the character of a state of the art review and a selection of the best available approaches for describing interfacial

processes. In many cases, the original contribution of the author is incorporated into the overall presentation. The most important aspects of the presentation are that they stem from the author's long years of experience in developing computer codes. The emphasis is on the practical use of these relationships: either as stand alone estimation methods or within a framework of computer codes.

Volume 4 is devoted to selected chapters of the multiphase fluid dynamics that are important for practical applications: The state of the art of the turbulence modeling in multiphase flows is presented. As an introduction, some basics of single-phase boundary layer theory, including some important scales and flow oscillation characteristics in pipes and rod bundles are presented. Then the scales characterizing dispersed flow systems are presented. The description of turbulence is provided at different level of complexity: simple algebraic models for eddy viscosity, algebraic models based on the *Boussinesq* hypothesis, modification of the boundary layer share due to modification of the bulk turbulence, and modification of the boundary layer share due to nucleate boiling. Then the role of the following forces on the mathematical description of turbulent flows is discussed: the lift force, the lubrication force in the wall boundary layer, and the dispersion force. A pragmatic generalization of the k-eps models for continuous velocity fields, which contains flows in large volumes and flows in porous structures, is proposed. A method of how to derive source and sink terms for multiphase k-eps models is presented. A set of 13 single- and two phase benchmarks for the verification of k-eps models in system computer codes is provided and reproduced with the IVA computer code as an example of the application of the theory. This methodology is intended to help other engineers and scientists to introduce this technology step by step in their own engineering practice.

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In order to allow the application of the theory from the first three volumes also to processes in combustion engines, a systematic set of internally consistent state equations for diesel fuel gas and liquid valid in a broad range of changing pressures and temperatures is provided.

Volume 5 is devoted to nuclear thermal hydraulics, which is a substantial part of nuclear reactor safety. It provides knowledge and mathematical tools for an adequate description of the process of the transfer of the fission heat released in materials due to nuclear reactions into its environment. It step by step introduces the reader into the understanding of the "simple" boiling flow in a pipe described mathematically using ideas of different complexity like equilibrium, nonequilibrium, homogeneity, and nonhomogeneity. Then the mathematical description of

the heat release inside the fuel, the resulting temperature distribution inside the fuels, and the interaction of the fuel with the cooling fluid are introduced. Next, the “simple” three-fluid boiling flow in a pipe is described by gradually involving the mechanisms like entrainment and deposition, dynamic fragmentation, collisions, coalescence, and turbulence. All heat transfer mechanisms are introduced gradually by discussing their uncertainty. Different techniques are introduced, like boundary layer treatments or integral methods. Comparisons with experimental data at each step demonstrate the success of the different ideas and models. After an introduction into the design of the reactor pressure vessels for pressurized and boiling water reactors the accuracy of modern methods is demonstrated using a large number of experimental data sets for steady and transient flows in heated bundles. Starting with single pipe boiling going through to boiling in rod bundles the analysis of the complete vessel including the reactor is finally demonstrated. Then a powerful method for nonlinear stability analysis of flow boiling and condensation is introduced. Models are presented and their accuracies for describing critical multiphase flow at different level of complexity are investigated. The basics of the design of steam generators, moisture separators, and emergency condensers are presented. Methods for analyzing complex pipe network flows with components like pumps, valves, etc., are also presented. Methods for the analysis of important aspects of severe accidents like melt-water interactions, external cooling, and cooling of layers of molten nuclear reactor material are presented. Valuable sets of thermophysical and transport properties for severe accident analysis are presented for the following materials: uranium dioxide, zirconium dioxide, stainless steel, zirconium, aluminum, aluminum oxide, silicon dioxide, iron oxide, molybdenum, boron oxide, reactor corium, sodium, lead, bismuth, and lead-bismuth eutectic alloy. The emphasis is on the complete and consistent thermodynamical sets of analytical approximations appropriate for computational analysis. Thus, the book presents a complete coverage of modern nuclear thermal hydrodynamics.

29.12.2010  
Herzogenaurach

# Nomenclature

## Latin

$A$	cross-section, $m^2$
$\mathbf{A}$	surface vector
$a$	speed of sound, $m/s$
$a_{lw}$	surface of the field $l$ wetting the wall $w$ per unit flow volume $\sum_{l=1}^{l_{\max}} Vol_l$ belonging to control volume $Vol$ (local volume interface area density of the structure $w$ ), $m^{-1}$
$a_{l\sigma}$	surface of the velocity field $l$ contacting the neighboring fields per unit flow volume $\sum_{l=1}^{l_{\max}} Vol_l$ belonging to control volume $Vol$ (local volume interface area density of the velocity field $l$ ), $m^{-1}$
$a_l$	total surface of the velocity field $l$ per unit flow volume $\sum_{l=1}^{l_{\max}} Vol_l$ belonging to control volume $Vol$ (local volume interface area density of the velocity field $l$ ), $m^{-1}$
$Cu_i$	Courant criterion corresponding to each eigenvalue, <i>dimensionless</i>
$C_{il}$	mass concentration of the inert component $i$ in the velocity field $l$
$c$	coefficients, <i>dimensionless</i>
$C_m$	mass concentration of the component $m$ in the velocity field, <i>dimensionless</i>
$C_i$	mass concentration of the component $i$ in the velocity field, <i>dimensionless</i>
$c_p$	specific heat at constant pressure, $J/(kgK)$
$c^{vm}$	virtual mass force coefficient, <i>dimensionless</i>
$c^d$	drag force coefficient, <i>dimensionless</i>
$c^L$	lift force coefficient, <i>dimensionless</i>
$D_{hy}$	hydraulic diameter (4 times cross-sectional area / perimeter), $m$
$D_{3E}$	diameter of the entrained droplets, $m$
$D_{ld}$	size of the bubbles produced after one nucleation cycle on the solid structure, bubble departure diameter, $m$



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$D_{ldm}$	size of bubbles produced after one nucleation cycle on the inert solid particles of field $m = 2, 3$
$D_{lch}$	critical size for homogeneous nucleation, $m$
$D_{lcd}$	critical size in presence of dissolved gases, $m$
$D'_l$	most probable particle size, $m$
$D_l$	characteristic length of the velocity field $l$ , particle size in case of fragmented field, $m$
$D_{il}^l$	coefficient of molecular diffusion for species $i$ into the field $l$ , $m^2 / s$
$D_{il}^t$	coefficient of turbulent diffusion, $m^2 / s$
$D_{il}^*$	total diffusion coefficient, $m^2 / s$
$DC_{il}$	right-hand side of the nonconservative conservation equation for the inert component, $kg / (sm^3)$
$D$	diffusivity, $m^2 / s$
$d$	total differential
$E$	total energy, $J$
$e$	specific internal energy, $J/kg$
$F(\xi)$	function introduced first in Eq. (42), Chapter 2
$F, f(...)$	function of (...)
$f$	force per unit flow volume, $N / m^3$
$f$	fraction of entrained melt or water in the detonation theory
$F_{lv}$	surfaces separating the velocity field $l$ from the neighboring structure within $Vol$ , $m^2$
$F_{l\sigma}$	surfaces separating the velocity field $l$ from the neighboring velocity field within $Vol$ , $m^2$
$F$	surface defining the control volume $Vol$ , $m^2$
$f_{im}$	frequency of the nuclei generated from one activated seed on the particle belonging to the donor velocity field $m$ , $s^{-1}$
$f_{lv}$	frequency of the bubble generation from one activated seed on the channel wall, $s^{-1}$
$f_{l,coal}$	coalescence frequency, $s^{-1}$
$g$	acceleration due to gravity, $m / s^2$
$H$	height, $m$
$h$	specific enthalpy, $J/kg$
$h_i$	eigenvectors corresponding to each eigenvalue
<b>I</b>	unit matrix, <i>dimensionless</i>
<b>i</b>	unit vector along the $x$ -axis
<b>J</b>	matrix, <i>Jacobian</i>
<b>j</b>	unit vector along the $y$ -axis

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<b>k</b>	unit vector along the $k$ -axis
$k$	cell number
$k$	kinetic energy of turbulent pulsation, $m^2 / s^2$
$k_{il}^T$	coefficient of thermodiffusion, <i>dimensionless</i>
$k_{il}^P$	coefficient of barodiffusion, <i>dimensionless</i>
$L$	length, $m$
$M_i$	kg-mole mass of the species $i$ , $kg/mole$
$m$	total mass, $kg$
$\mathbf{n}_{\Delta V}$	unit vector pointing along $\Delta \mathbf{V}_{ml}$ , <i>dimensionless</i>
$\mathbf{n}$	unit vector pointing outwards from the control volume $Vol$ , <i>dimensionless</i>
$\mathbf{n}_{le}$	unit surface vector pointing outwards from the control volume $Vol$
$\mathbf{n}_{l\sigma}$	unit interface vector pointing outwards from the velocity field $l$
$n_{il}$	number of the particle from species $i$ per unit flow volume, $m^{-3}$
$n_l$	number of particles of field $l$ per unit flow volume, particle number density of the velocity field $l$ , $m^{-3}$
$\dot{n}_{coal}$	number of particles disappearing due to coalescence per unit time and unit volume, $m^{-3}$
$\dot{n}_{l,kin}$	particle production rate due to nucleation during evaporation or condensation, $1/(m^3 s)$
$n_{lw}''$	number of the activated seeds on unit area of the wall, $m^{-2}$
$\dot{n}_{lh}$	number of the nuclei generated by homogeneous nucleation in the donor velocity field per unit time and unit volume of the flow, $1/(m^3 s)$
$\dot{n}_{l,dis}$	number of the nuclei generated from dissolved gases in the donor velocity field per unit time and unit volume of the flow, $1/(m^3 s)$
$\dot{n}_{l,sp}$	number of particles of the velocity field $l$ arising due to hydrodynamic disintegration per unit time and unit volume of the flow, $1/(m^3 s)$
$P$	probability
$P$	irreversibly dissipated power from the viscous forces due to deformation of the local volume and time average velocities in the space, $W/kg$
$Per$	perimeter, $m$
$p_{li}$	$l = 1$ : partial pressure inside the velocity field $l$ $l = 2,3$ : pressure of the velocity field $l$
$p$	pressure, $Pa$
$\dot{q}'''$	thermal power per unit flow volume introduced into the fluid, $W/m^3$
$\dot{q}_{\sigma l}'''$	$l = 1,2,3$ . Thermal power per unit flow volume introduced from the interface into the velocity field $l$ , $W/m^3$

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$\dot{q}_{w\sigma l}''$	thermal power per unit flow volume introduced from the structure interface into the velocity field $l$ , $W / m^3$
$R$	mean radius of the interface curvature, $m$
$\mathbf{r}(x,y,z)$	position vector, $m$
$R$	(with indexes) gas constant, $J/(kgK)$
$\mathbf{s}$	arc length vector, $m$
$S$	total entropy, $J/K$
$s$	specific entropy, $J/(kgK)$
$Sc^t$	turbulent <i>Schmidt</i> number, <i>dimensionless</i>
$Sc^m$	is the turbulent <i>Schmidt</i> number for particle diffusion, <i>dimensionless</i>
$T$	temperature, $K$
$T_l$	temperature of the velocity field $l$ , $K$
$\mathbf{T}$	shear stress tensor, $N / m^2$
$\mathbf{t}$	unit tangent vector
$U$	dependent variables vector
$Vol$	control volume, $m^3$
$Vol^{1/3}$	size of the control volume, $m$
$Vol_l$	volume available for the field $l$ inside the control volume, $m^3$
$\sum_{l=1}^{l_{\max}} Vol_l$	volume available for the flow inside the control volume, $m^3$
$\mathbf{V}$	instantaneous fluid velocity with components, $u, v, w$ in $r, \theta$ , and $z$ direction, $m/s$
$\mathbf{V}_l^\tau$	instantaneous field velocity with components, $u_l^\theta, v_l^\tau, w_l^\tau$ in $r, \theta$ , and $z$ direction, $m/s$
$\mathbf{V}_l$	time-averaged velocity, $m/s$
$\mathbf{V}_l'$	pulsation component of the instantaneous velocity field, $m/s$
$\Delta \mathbf{V}_{lm}$	$\mathbf{V}_l - \mathbf{V}_m$ , velocity difference, disperse phase $l$ , continuous phase $m$ carrying $l$ , $m/s$
$\delta_i V_l^\tau$	diffusion velocity, $m/s$
$\mathbf{V}_{l\sigma}^\tau$	interface velocity vector, $m/s$
$\mathbf{V}_l^\tau \gamma$	instantaneous vector with components, $u_l^\theta \gamma_r, v_l^\tau \gamma_\theta, w_l^\tau \gamma_z$ in $r, \theta$ , and $z$ directions, $m/s$
$v$	specific volume, $m^3 / kg$
$x$	mass fraction, <i>dimensionless</i>
$y$	distance between the bottom of the pipe and the center of mass of the liquid, $m$
$\times$	vector product

**Greek**

$\alpha_l$	part of $\gamma_v Vol$ available to the velocity field $l$ , local instantaneous volume fraction of the velocity field $l$ , <i>dimensionless</i>
$\alpha_{il}$	the same as $\alpha_l$ in the case of gas mixtures; in the case of mixtures consisting of liquid and macroscopic solid particles, the part of $\gamma_v Vol$ available to the inert component $i$ of the velocity field $l$ , local instantaneous volume fraction of the inert component $i$ of the velocity field $l$ , <i>dimensionless</i>
$\alpha_{l,max}$	$\approx 0.62$ , limit for the closest possible packing of particles, <i>dimensionless</i>
$\gamma_v$	the part of $dVol$ available for the flow, volumetric porosity, <i>dimensionless</i>
$\gamma$	surface permeability, <i>dimensionless</i>
$\bar{\gamma}$	directional surface permeability with components $\gamma_r, \gamma_\theta, \gamma_z$ , <i>dimensionless</i>
$\Delta$	finite difference
$\delta$	small deviation with respect to a given value
$\delta_l$	$= 1$ for continuous field; $= 0$ for disperse field, <i>dimensionless</i>
$\partial$	partial differential
$\varepsilon$	dissipation rate for kinetic energy from turbulent fluctuation, power irreversibly dissipated by the viscous forces due to turbulent fluctuations, $W/kg$
$\eta$	dynamic viscosity, $kg/(ms)$
$\theta$	$\theta$ -coordinate in the cylindrical or spherical coordinate systems, <i>rad</i>
$\kappa$	$= 0$ for Cartesian coordinates, $= 1$ for cylindrical coordinates
$\kappa$	isentropic exponent
$\kappa_l$	curvature of the surface of the velocity field $l$ , <i>m</i>
$\lambda$	thermal conductivity, $W/(mK)$
$\lambda$	eigenvalue
$\mu_l^{\tau}$	local volume-averaged mass transferred into the velocity field $l$ per unit time and unit mixture flow volume, local volume-averaged instantaneous mass source density of the velocity field $l$ , $kg/(m^3s)$
$\mu_l$	time average of $\mu_l^{\tau}$ , $kg/(m^3s)$
$\mu_{wl}$	mass transport from exterior source into the velocity field $l$ , $kg/(m^3s)$
$\mu_{il}^{\tau}$	local volume-averaged inert mass from species $i$ transferred into the velocity field $l$ per unit time and unit mixture flow volume, local volume-averaged

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	instantaneous mass source density of the inert component $i$ of the velocity field $l$ , $kg/(m^3 s)$
$\mu_{il}$	time average of $\mu_{il}^\tau$ , $kg/(m^3 s)$
$\mu_{iml}^\tau$	local volume-averaged instantaneous mass source density of the inert component $i$ of the velocity field $l$ due to mass transfer from field $m$ , $kg/(m^3 s)$
$\mu_{iml}$	time average of $\mu_{iml}^\tau$ , $kg/(m^3 s)$
$\mu_{ilm}^\tau$	local volume-averaged instantaneous mass source density of the inert component $i$ of the velocity field $l$ due to mass transfer from field $l$ into velocity field $m$ , $kg/(m^3 s)$
$\mu_{ilm}$	time average of $\mu_{ilm}^\tau$ , $kg/(m^3 s)$
$\nu$	cinematic viscosity, $m^2/s$
$\nu_l^t$	coefficient of turbulent cinematic viscosity, $m^2/s$
$\nu_l^m$	coefficient of turbulent particles diffusion, $m^2/s$
$\xi$	angle between $\mathbf{n}_{l\sigma}$ and $\Delta\mathbf{V}_{lm}$ , $rad$
$\rho$	density, $kg/m^3$
$\rho$	instantaneous density, density; without indexes, mixture density, $kg/m^3$
$\rho_l$	instantaneous field density, $kg/m^3$
$\rho_{il}$	instantaneous inert component density of the velocity field $l$ , $kg/m^3$
$\langle\rho_l\rangle^l$	intrinsic local volume-averaged phase density, $kg/m^3$
$(\rho w)_{23}$	entrainment mass flow rate, $kg/(m^2 s)$
$(\rho w)_{32}$	deposition mass flow rate, $kg/(m^2 s)$
$(\rho_l \mathbf{V}_l^\tau)^{le}$	local intrinsic surface mass flow rate, $kg/(m^2 s)$
$\sigma$ , $\sigma_{12}$	surface tension between phases 1 and 2, $N/m$
$\tau$	time, $s$
$\varphi$	angle giving the projection of the position of the surface point in the plane normal to $\Delta\mathbf{V}_{lm}$ , $rad$
$\chi_l^{m\sigma}$	the product of the effective heat transfer coefficient and the interfacial area density, $W/(m^3 K)$ . The subscript $l$ denotes inside the velocity field $l$ . The superscript $m\sigma$ denotes location at the interface $\sigma$ dividing field $m$ from field $l$ . The superscript is only used if the interfacial heat transfer is associated with mass transfer. If there is heat transfer only, the

linearized interaction coefficient is assigned the subscript  $ml$  only, indicating the interface at which the heat transfer takes place.

### ***Subscripts***

$c$	continuous
$d$	disperse
$lm$	from $l$ to $m$ or $l$ acting on $m$
$w$	region "outside of the flow"
$e$	entrances and exits for control volume $Vol$
$l$	velocity field $l$ , intrinsic field average
$i$	inert components inside the field $l$ , noncondensable gases in the gas field $l = 1$ , or microscopic particles in water in field 2 or 3
$i$	corresponding to the eigenvalue $\lambda_i$ in Chapter 4
$M$	noninert component
$m$	mixture of entrained coolant and entrained melt debris that is in thermal and mechanical equilibrium behind the shock front
$ml$	from $m$ into $l$
$iml$	from $im$ into $il$
max	maximum number of points
$n$	inert component
0	at the beginning of the time step
$E$	entrainment
$coal$	coalescence
$sp$	splitting, fragmentation
$\sigma$	interface
$\tau$	old time level
$\tau + \Delta\tau$	new time level
*	initial
0	reference conditions
$p, v, s$	at constant $p, v, s$ , respectively
L	left
R	right
1	vapor or in front of the shock wave
2	water or behind the shock wave
3	melt
4	entrained coolant behind the front – entrained coolant
5	microparticles after the thermal interaction – entrained melt

### ***Superscripts***

$\hat{\phantom{x}}$	time fluctuation
'	saturated steam
"	saturated liquid

"	saturated solid phase
$A$	air
$d$	drag
$e$	heterogeneous
$i$	component (either gas or solid particles) of the velocity field
$i_{\max}$	maximum for the number of the components inside the velocity field
$L$	lift
$l$	intrinsic field average
$le$	intrinsic surface average
$l\sigma$	averaged over the surface of the sphere
$m$	component
$n$	normal
$n$	old iteration
$n+1$	new iteration
$t$	turbulent, tangential
$vm$	virtual mass
$\tau$	temporal, instantaneous
–	averaging sign

### **Operators**

$\nabla \cdot$	divergence
$\nabla$	gradient
$\nabla_n$	normal component of the gradient
$\nabla_t$	tangential component of the gradient
$\nabla_l$	surface gradient operator, $1/m$
$\nabla^2$	<i>Laplacian</i>
$\langle \rangle$	local volume average
$\langle \rangle^l$	local intrinsic volume average
$\langle \rangle^{le}$	local intrinsic surface average

### **Nomenclature required for coordinate transformations**

- $(x, y, z)$  coordinates of a Cartesian, left oriented coordinate system (*Euclidean* space). Another notation which is simultaneously used is  $x_i$  ( $i = 1, 2, 3$ ):  
 $x_1, x_2, x_3$
- $(\xi, \eta, \zeta)$  coordinates of the curvilinear coordinate system called transformed coordinate system. Another notation which is simultaneously used is  $\xi^i$  ( $i = 1, 2, 3$ ):  $\xi^1, \xi^2, \xi^3$

- $\mathbf{V}_{cs}$  the velocity of the curvilinear coordinate system
- $\sqrt{g}$  *Jacobian* determinant or *Jacobian* of the coordinate transformation  
 $x = f(\xi, \eta, \zeta)$ ,  $y = g(\xi, \eta, \zeta)$ ,  $z = h(\xi, \eta, \zeta)$
- $a_{ij}$  elements of the *Jacobian* determinant
- $a^{ij}$  elements of the determinant transferring the partial derivatives with respect to the transformed coordinates into partial derivatives with respect to the physical coordinates. The second superscript indicates the Cartesian components of the contravariant vectors
- $(\mathbf{a}_1, \mathbf{a}_2, \mathbf{a}_3)$  covariant base vectors of the curvilinear coordinate system tangent vectors to the three curvilinear coordinate lines represented by  $(\xi, \eta, \zeta)$
- $(\mathbf{a}^1, \mathbf{a}^2, \mathbf{a}^3)$  contravariant base vectors, normal to a coordinate surface on which the coordinates  $\xi$ ,  $\eta$  and  $\zeta$  are constant, respectively
- $g_{ij}$  covariant metric tensor (symmetric)
- $g^{ij}$  contravariant metric tensor (symmetric)
- $(\mathbf{e}^1, \mathbf{e}^2, \mathbf{e}^3)$  unit vectors normal to a coordinate surface on which the coordinates  $\xi$ ,  $\eta$  and  $\zeta$  are constant, respectively
- $\mathbf{V}^i$  =  $\mathbf{a}^i \cdot \mathbf{V}$ , contravariant components of the vector  $\mathbf{V}$
- $\mathbf{V}_i$  =  $\mathbf{a}_i \cdot \mathbf{V}$ , covariant components of the vector  $\mathbf{V}$
- $(\gamma_\xi, \gamma_\eta, \gamma_\zeta)$  permeabilities of coordinate surfaces on which the coordinates  $\xi$ ,  $\eta$  and  $\zeta$  are constant, respectively

**Greek**

A, $\alpha$ Alpha	I, $\iota$ Iota	$\Sigma$ , $\sigma$ Sigma
B, $\beta$ Beta	K, $\kappa$ Kappa	T, $\tau$ Tau
$\Gamma$ , $\gamma$ Gamma	$\Lambda$ , $\lambda$ Lambda	$\Phi$ , $\phi$ Phi
$\Delta$ , $\delta$ Delta	M, $\mu$ Mu	X, $\chi$ Chi
E, $\varepsilon$ Epsilon	N, $\nu$ Nu	Y, $\upsilon$ Ypsilon
Z, $\zeta$ Zeta	$\Xi$ , $\xi$ Xi	$\Psi$ , $\psi$ Psi
H, $\eta$ Eta	O, $o$ Omikron	$\Omega$ , $\omega$ Omega
$\Theta$ , $\vartheta$ Theta	$\Pi$ , $\pi$ Pi	
	P, $\rho$ Rho	



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# 1 Mass conservation

“...all changes in the nature happen so that the mass lost by one body is added to other...”.

1748, *M. Lomonosov's* letter to *L. Euler*

## 1.1 Introduction

The creation of computer codes for modeling multiphase flows in industrial facilities is very complicated, time-consuming, and expensive. This is why the fundamentals on which such codes are based are subject to continuous review in order to incorporate the state of the art of knowledge into the current version of the code in question. An important element of the codes is the system of partial differential equations governing the flow. The understanding of each particular term in these equations is very important for the application.

From the large number of formulations of the conservation equations for multiphase flows, see *Slattery* (1990), *Hetstrony* (1982), *Delhaye et al.* (1981), *Ishii* (1975) and the references given therein, the local volume averaging as founded by *Anderson and Jackson* (1967), *Slattery* (1967), and *Whitaker* (1967) is selected because it is rigorous and mathematically elegant. The heterogeneous porous media formulation introduced by *Gentry et al.* (1966), commented on by *Hirt* (1993), and used by *Sha et al.* (1984), is then implanted into the formalism as a geometrical skeleton because of its practical usefulness. For technical structures, the introduction of the local volume porosity and directional permeability is a convenient formalism to describe the real distances between the flow volumes. Simplifications of geometry that lead to distortion of the modeled acoustic process characteristics of the systems are not necessary. Beyond these concepts, I include inert components in each velocity field, and introduce the concept of dynamic particle fragmentation and coalescence. Then I perform subsequent time averaging. The link between kinetic generation of particles and the mass source terms is specified. The link between mass source terms and the change in bubble/droplet size due to evaporation or condensation is presented for local volume-averaged and time-averaged source terms. The concept of monodispersity is discussed, and a method is proposed for computation of the disappearance of particles due to evaporation and condensation.

Having started this work in 1983 to create the foundations of the IVA computer code, I believed then and now after 23 years, I am convinced that this concept is

unique and that it is the most effective concept for deriving multiphase equations that are applicable to flows in complicated technical facilities.

Here I will present the derivation of the mass conservation equation for a single component inside the field, for the mixture constituting the field. Then I will present the derivation of the particle number density equations. This chapter is an improved version of the work published in *Kolev (1994b)*.

## 1.2 Basic definitions

Consider the idealized flow patterns of the three-phase flows shown in Fig. 1.1. What do all these flow patterns have in common? How can they be described by a single mathematical formalism that allows transition from any flow pattern to any other flow pattern? How can the complicated interactions between the participating components be described as simply as possible? In the discussion that follows a number of basic definitions are introduced or reiterated so as to permit successful treatment of such flows using a unified algorithm.

*Zemansky (1968)*, p. 566, defines “a phase as a system or a portion of the system composed of any number of chemical constituents satisfying the requirements (a) that it is homogeneous and (b) that it has a definite boundary”. In this sense, consideration of a multiphase flow system assumes the consistency described below:

1) A homogeneous gas phase having a definite boundary. The gas phase can be continuous (nonstructured) or dispersed (structured). The gas phase in the gas/droplet flow is an example of a nonstructured gas phase. The bubbles in the bubble/liquid flow are an example of a structured gas phase. Besides the steam, the gas phase contains  $i_{1\max} - 1$  groups of species of *inert* components. Inert means that these components do not experience changes in the aggregate state during the process of consideration. The gas components occupy the entire gas volume in accordance with *Dalton's* law, and therefore possess the definite boundary of the gas phase themselves. The gas phase velocity field will be denoted as no.1 and  $l = 1$  assigned to this. Velocity field no.1 has the temperature  $T_1$ .

2) A mixture of liquid water and a number of species  $i_{2\max} - 1$  of microscopic solid particles. The particles are homogeneous and have a definite boundary surrounded by liquid only. The water has its own definite outer boundary. An example of such a mixture is a liquid film carrying  $i_{2\max} - 1$  groups of radioactive solid species. This means that the mixture of water with  $i_{2\max} - 1$  groups of species is a mixture of  $i_{2\max}$  phases. The mixture velocity field defined here will be denoted as no. 2. Velocity field no. 2 can likewise be nonstructured or structured. Velocity field no. 2 has the temperature  $T_2$ .

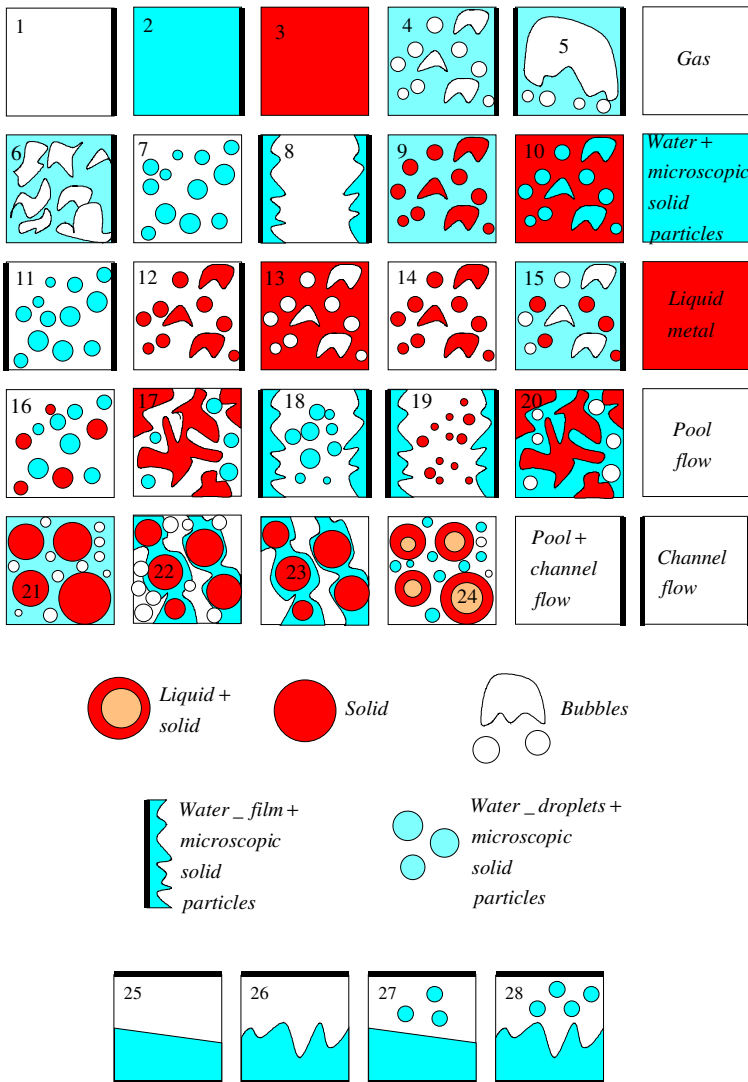


Fig. 1.1 Multiphase flow patterns

3) A mixture of the type of velocity field no. 2 of liquid with  $i_{3\max} - 1$  groups of species of microscopic solid particles. The mixture can be nonstructured or structured; this is denoted as velocity field no. 3. Velocity field no. 3 has the temperature  $T_3$ . In the event that only one inert component occupies velocity field no. 3, this component will be allowed to be either a liquid, or a homogeneous liquid–solid mixture being in thermodynamic equilibrium, or solid particles only. This allows versatile application of the concept.

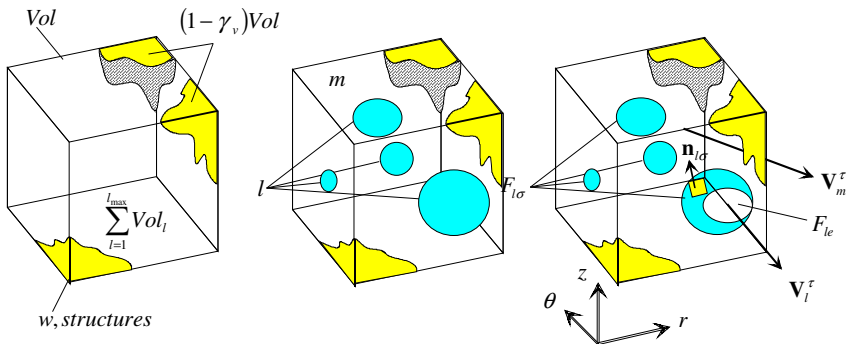
The entire flow under consideration consists of  $i_{2\max} + i_{3\max} + 1$  phases and is conditionally divided into three velocity fields. In order to avoid the need to write indices to the indices,  $il$  will be written in place of  $i_i$ .

An example demonstrating the necessity for the use of three velocity fields is the modeling of a mixture consisting of gas, water, and a liquid metal, whose densities have the approximate ratios  $1 : 10^3 : 10^4$ . In a transient, these three fields will have velocities and temperatures that differ considerably.

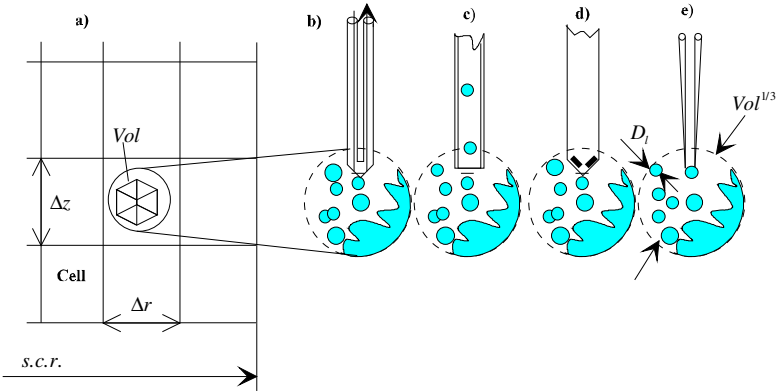
Turning our attention again to the flow patterns depicted in Fig. 1.1 and keeping in mind that in reality these change their characteristic sizes chaotically, it is obvious that it is not possible to determine the details of the thermodynamic and flow parameters for each component of each velocity field. Consequently, some type of averaging must be implemented. Following *Anderson and Jackson (1967)*, *Slattery (1967)*, and *Whitaker (1967)*, a control volume is allocated to every point in space, and the thermodynamic and flow properties averaged for each velocity field are assigned to the center of the control volume. Since there is a control volume associated with every point in space, a field of average values can be generated for all thermodynamic and flow properties. The field of average properties is, therefore, smooth, and space derivatives of these averaged properties exist.

Consider the control volume  $Vol$  occupied by a nonmovable structure in addition to the three velocity fields, see Figs. 1.2 and 1.3. While the individual  $l$ -field volumes  $Vol_l$  may be functions of time and space, the control volume  $Vol$  is not.

Velocity field  $l$  is taken to have the characteristic length  $D_l$  (e.g., bubble size, droplet size), which is much larger than the molecular free path (Fig. 1.3). The size of the computational region of interest here is much larger than the size of the local structure  $D_l$  and larger than the spatial changes in the flow parameters of interest. The choice of the size of the control volume, which is of the order of



**Fig. 1.2** Control volume for definition of the mass conservation equation, partially occupied by structure and two different velocity fields



**Fig. 1.3** Comparison of possible scales of local volume averaging, scale of the measuring devices, scale of the computational region (s.c.r.) and the global flow dimensions

$Vol^{1/3}$  has a major impact on the meaning of the averaged values. From the various approaches possible, two are meaningful:

(i) The size of the control volume is larger than the characteristic configuration length  $D_l$  of the field  $l$ , i.e.,

$$\text{molecular free path} \ll D_l \ll Vol^{1/3} \ll \text{size of the computational region} \quad (1.1)$$

This approach is useful for fine dispersed flows.

(ii) The size of the control volume is comparable to the characteristic configuration length  $D_l$  of the field  $l$ , i.e.,

$$\text{molecular free path} \ll D_l \approx Vol^{1/3} \ll \text{size of the computational region} \quad (1.2)$$

This approach is meaningful for direct simulation or for simulation of flow patterns with stratification, for example.

The geometry of the nonmovable structures inside  $Vol$  is characterized by the volumetric porosity  $\gamma_v$ , which is defined as the ratio of the volume occupied by the flow mixture inside the control volume,  $\sum_{l=1}^3 Vol_l$ , and the control volume  $Vol$ :

$$\gamma_v = \frac{\sum_{l=1}^3 Vol_l}{Vol} . \quad (1.3)$$

Consequently, the part of  $Vol$  occupied by structures is  $1 - \gamma_v$ . Inside the volume available for the flow,  $\sum_{l=1}^3 Vol_l$ , it is assumed that three velocity fields coexist. The instantaneous geometry of the velocity field is defined in a similar way to the

nonmovable structure. The local instantaneous volume fraction of the velocity field  $l$ ,  $\alpha_l$ , is defined as the part of  $\gamma_v Vol$  occupied by the velocity field  $l$ :

$$\alpha_l = \frac{Vol_l}{\sum_{l=1}^3 Vol_l} = \frac{Vol_l}{\gamma_v Vol}. \quad (1.4)$$

In general, this varies with time and location. By definition, we have

$$\sum_{l=1}^3 \alpha_l = 1. \quad (1.4a)$$

On the basis of the definitions introduced above, the part of the control volume  $Vol$  occupied by the velocity field  $l$  is

$$Vol_l = \alpha_l \gamma_v Vol. \quad (1.5)$$

The discussion here is restricted to the right-handed Cartesian and cylindrical coordinates designated by  $r$ ,  $\theta$ , and  $z$ . The instantaneous field velocity  $\mathbf{V}_l^r(r, \theta, z, \tau)$  is defined through the entire region occupied by the velocity field  $l$ . The components of the velocity vectors along the axes are

$$\mathbf{V}_l^r = (u_l^r, v_l^r, w_l^r). \quad (1.6)$$

The control volume  $Vol$  is bounded by a surface  $F$ . The velocity field within the volume  $Vol_l$  is bounded by a surface  $F_l$ . The orientation of the smooth  $l$ -surfaces in space with respect to the field  $l$  is given with the unit vector  $\mathbf{n}_l$  attached to the  $l$ -surface and pointing outwards of the field  $l$ . The interface belongs to both neighboring phases. The surfaces  $F$ ,  $F_l$  are defined as scalars. The surface  $F_l$  has the following three constituents:

- (1)  $F_{le}$ , the control volume enters and exits crossing the field  $l$ , being also part of the control volume surface  $F$ .
- (2)  $F_{l\sigma}$ , the interface between the field  $l$  and the surrounding field or fields  $m$ .
- (3)  $F_{lw}$ , the interface between the solid structure, wall, and the field.

Consider the small part of the surface  $F$ ,  $\Delta F$ . The part of this surface occupied by the flow is  $\sum_{l=1}^3 \Delta F_l$ . The ratio

$$\gamma = \frac{1}{\Delta F} \sum_{l=1}^3 \Delta F_{le} \quad (1.7)$$

is also known as *surface permeability*. We consider the surface permeability as the ratio of scalars. Here

$$dF_{le} < \Delta F_{le} < F_{le}, \quad (1.8)$$

and

$$dF < \Delta F < F. \quad (1.9)$$

Note the difference between surface permeability and the “permeability coefficients” in standard use for the description of the pressure drop in porous media. Also note that the values of  $\gamma_v$  at the center and at the surface of the control volume may be different, because they belong to different averaging volumes. This definition of  $\gamma_v$  is sufficient to describe the isotropic porosity that is found in homogeneous porous bodies. For most of the technical structures forming channels for multiphase flow it is extremely convenient to introduce the surface permeability  $\gamma$  in addition to the volumetric porosity  $\gamma_v$  in order to describe obstacles to the flow inside the space of interest, i.e., nonisotropic porosity (heterogeneous porous body). As far I know, *Gentry et al.* (1966) introduced this concept for the first time in 1966, using the terms volume and surface flow fractions for volumetric porosity and surface permeability. Valuable comments for the practical use of the concept are given by *Hirt* (1993). Note that inclined surfaces in orthogonal structured grids can also be modeled by this method by specifying correctly the part of the boundary cells that is occupied by flow and the part of the surfaces that is open for the flow.  $\gamma$ -values are likewise permitted to be prescribed functions of time in addition to functions of space. Additionally, the vector

$$\vec{\gamma} = \mathbf{n}_{l,x} \gamma + \mathbf{n}_{l,y} \gamma + \mathbf{n}_{l,z} \gamma = \vec{\gamma}_x + \vec{\gamma}_y + \vec{\gamma}_z \quad (1.10)$$

is referred to as the directional surface permeability.

The counterpart of the surface permeability  $\gamma$  for each velocity field is then introduced as field surface fraction  $\alpha_l^e$  (sometimes called the heterogeneous volume fraction), which is the part of the surface  $\gamma\Delta F$  occupied or passed by the field  $l$ :

$$\alpha_l^e = \Delta F_{l,e} / (\gamma\Delta F). \quad (1.11)$$

This means that

$$\sum_{l=1}^3 \alpha_l^e = 1. \quad (1.11a)$$

The impact of the choice of the size for the control volume can now be clearly understood by examining the relationship between the heterogeneous volume fraction and the local volume fraction. A large size for the control volume compared to the characteristic configuration length, assumption (i), leads to

$$\alpha_l^e = \alpha_l. \quad (1.12)$$



This assumption is reasonable for dispersed flows. If the size of the control volume is comparable with the characteristic configuration length of the field  $l$ , assumption (ii), the following results:

$$\alpha_l^e \neq \alpha_l. \quad (1.13)$$

In both cases local volume averaging is mathematically permissible.

Consider any scalar property of field  $l$  designated with  $\varphi_l$ . Inside the space occupied by the field  $l$  this property is considered to be smooth. In general, the averages of the property  $\varphi_l$  given below are required for the derivation of local volume average equations.

(1) The *local volume average* is defined by

$$\langle \varphi_l \rangle := \frac{1}{Vol} \int_{Vol_l} \varphi_l dVol. \quad (1.14)$$

Note that when the property  $\varphi_l$  is constant, the local volume average given by Eq. (1.14) does not equal this constant.

(2) The *intrinsic field average* is defined by

$$\langle \varphi_l \rangle^l := \frac{1}{Vol_l} \int_{Vol_l} \varphi_l dVol. \quad (1.15)$$

These two average properties are obviously related by

$$\langle \varphi_l \rangle = \alpha_l \gamma_v \langle \varphi_l \rangle^l. \quad (1.16)$$

The averages defined by Eqs. (1.14) and (1.15) are applicable not only for scalars but also for vectors and tensors.

The macroscopic density  $\langle \rho_l \rangle^l$  is the intrinsic field average over the space occupied by velocity field  $l$ . Outside the velocity field  $l$  the density  $\langle \rho_l \rangle^l$  is not defined. The macroscopic density, simply referred to as density in the following, obeys the law expressed by the macroscopic equation of state. The equation of state describes the interdependence between density, the intrinsic field-averaged pressure, and the intrinsic field-averaged temperature,  $\langle \rho_l \rangle^l = f(\langle p_l \rangle^l, \langle T_l \rangle^l)$  frequently denoted in short as  $\rho_l = \rho_l(p_l, T_l)$  in the sense of intrinsic volume average.

3) The *intrinsic surface average* is defined by

$$\langle \varphi_l \rangle^{le} := \frac{1}{\Delta F_{le}} \int_{\Delta F_{le}} \varphi_l \mathbf{n}_l dF. \quad (1.17)$$

The result of the surface averaging of a scalar is a vector. Therefore  $\langle \varphi_i \rangle^{le}$  is a vector. The product of the scalar property  $\varphi_i$  and  $\mathbf{V}_i^\tau$ ,  $\varphi_i \mathbf{V}_i^\tau$  is referred to as the flux of  $\varphi_i$  along the flow direction. This is a vector. The scalar product of the three-dimensional vector field  $\varphi_i \mathbf{V}_i^\tau$  across the oriented surface  $dF$  in the direction  $\mathbf{n}_i$ ,  $\varphi_i \mathbf{V}_i^\tau \cdot \mathbf{n}_i dF$ , is the flow of property  $\varphi_i$  normal through the surface  $dF$ . The vector form of Eq. (1.17) is

$$\langle \varphi_i \mathbf{V}_i^\tau \rangle^{le} := \frac{1}{\Delta F_{le}} \int_{\Delta F_{le}} \varphi_i \mathbf{V}_i^\tau \cdot \mathbf{n}_i dF \quad (1.18)$$

Note that the intrinsic surface average of a scalar is a vector and the intrinsic surface average of a vector is a scalar. The notation  $\langle \rangle$ ,  $\langle \rangle^l$ , and  $\langle \rangle^{le}$  should be considered as operators defined by Eqs. (1.14), (1.15), and (1.17). Here, I adapted the notation used by *Whitaker* (1985).

For practical applications the following definition of the weighted average velocity is used:

$$\langle \mathbf{V}_i^\tau \rangle^{le,\rho} = \langle \rho_i \mathbf{V}_i^\tau \rangle^{le} / \langle \rho_i \rangle^l. \quad (1.19)$$

### 1.3 Nonstructured and structured fields

As has already been mentioned, within the averaging volume the fields are essentially allowed to be (a) nonstructured or (b) structured. For the nonstructured fields the spatial variation of the properties within the continuous volume is smooth, and, therefore, space differentiating of the properties is allowed inside the space occupied by the field. Consequently the *Gauss–Ostrogradskii* theorem for transformation of surface integrals into volume integrals is applicable. This is not the case for a structured fluid, where the properties within the continuous fragment are smooth, but the field volume inside the control volume consists of several fragments. In this case, a simple extension of the *Gauss–Ostrogradskii* theorem for transfer of surface integral into volume integral over the fragmented volume occupied by the field is not possible unless one extends the definition of the properties in the sense of distributions, see the discussion by *Gray and Lee* in (1977), or uses some type of special treatment as described in the next section.

## 1.4 The *Slattery* and *Whitaker* local spatial averaging theorem

In the following text the operator  $\nabla \langle \cdot \rangle$  is used for the *gradient* of a scalar that is a vector, and  $\nabla \cdot \langle \cdot \rangle$  for the *divergence* of a vector that is a scalar. The following mathematical tools are used to derive local volume-averaged field conservation equations for the property  $\varphi_l$ .

(a) The *spatial averaging theorem*, Eq. (22) in *Whitaker* (1985),

$$\nabla \langle \varphi_l \rangle = \nabla \left( \alpha_l \gamma_v \langle \varphi_l \rangle^l \right) = \frac{1}{Vol} \int_{F_{ie}} \varphi_l \mathbf{n}_l dF. \quad (1.20)$$

Here  $\nabla \langle \varphi_l \rangle$  is a vector. The vector form of the spatial averaging theorem is given by Eq. (23) in *Whitaker* (1985) as

$$\nabla \cdot \langle \varphi_l \mathbf{V}_l^\tau \rangle = \nabla \cdot \left( \alpha_l \gamma_v \langle \varphi_l \mathbf{V}_l^\tau \rangle^l \right) = \frac{1}{Vol} \int_{F_{ie}} \varphi_l \mathbf{V}_l^\tau \cdot \mathbf{n}_l dF, \quad (1.21)$$

Here  $\nabla \cdot \langle \varphi_l \mathbf{V}_l^\tau \rangle$  is a scalar. For the validation of Eqs. (1.20) and (1.21) see *Ander-son* and *Jackson* (1967), *Slattery* (1967), and *Whitaker* (1967, 1985).

Useful consequences of Eq. (1.20) are obtained for  $\varphi_l = 1$ :

$$\nabla \langle \alpha_l \gamma_v \rangle = \frac{1}{Vol} \int_{F_{ie}} \mathbf{n}_l dF = \frac{1}{Vol} \int_F \alpha_l^e \gamma \mathbf{n}_l dF' = \nabla \langle \alpha_l^e \gamma \rangle. \quad (1.22)$$

Taking into account Eqs. (1.4a) and (1.11a), the sum of the above equation for all three fields yields

$$\nabla \langle \gamma_v \rangle = \frac{1}{Vol} \int_F \gamma \mathbf{n}_l dF' = \nabla \gamma. \quad (1.23)$$

It follows from Eq. (1.21) that

$$\nabla \cdot \langle \varphi_l \mathbf{V}_l^\tau \rangle = \frac{1}{Vol} \int_{F_{ie}} \varphi_l \mathbf{V}_l^\tau \cdot \mathbf{n}_l dF = \frac{1}{Vol} \int_F \alpha_l^e \gamma \langle \varphi_l \mathbf{V}_l^\tau \rangle^{el} \cdot \mathbf{n}_l dF'. \quad (1.24)$$

By introducing local surface averaging, Eq. (1.18), the last surface integral can be transformed into the divergence for a smooth vector field

$$\nabla \cdot \langle \varphi_l \mathbf{V}_l^\tau \rangle = \nabla \cdot \left( \alpha_l^e \gamma \langle \varphi_l \mathbf{V}_l^\tau \rangle^{el} \right). \quad (1.25)$$

This equation is similar to Eq. (2.26) obtained by *Sha et al.* (1984), where, however, the surface averaging was performed over the entire surface  $F$  of the control volume rather than over local  $\Delta F$ , which is a part of  $F$ .

**Note:** Comparing Eq. (1.25) with Eq. (1.21) we realize something very interesting

$$\nabla \cdot \langle \varphi_l \mathbf{V}_l^\tau \rangle = \nabla \cdot \left( \alpha_l \gamma_v \langle \varphi_l \mathbf{V}_l^\tau \rangle^l \right) = \nabla \cdot \left( \alpha_l^e \gamma \langle \varphi_l \mathbf{V}_l^\tau \rangle^{el} \right). \quad (1.25a)$$

This means that one can work either with local volume-averaged fluxes  $\langle \varphi_l \mathbf{V}_l^\tau \rangle^l$  or with local surface-averaged fluxes  $\langle \varphi_l \mathbf{V}_l^\tau \rangle^{el}$ . In the first case, one has to use in the divergence expression the local volume fractions  $\alpha_l$  and  $\gamma_v$ , and in the second case the local surface fractions  $\alpha_l^e$  and  $\gamma$ .

There are many literature sources where this difference is not clearly made.

(b) The *Gauss–Ostrogradskii* theorem is applied to the volume  $Vol_l$ . The resulting expression is divided by  $Vol$ . The result is

$$\frac{1}{Vol} \int_{Vol_l} \nabla \varphi_l dVol = \frac{1}{Vol} \int_{F_{le}} \varphi_l \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \varphi_l \cdot \mathbf{n}_l dF. \quad (1.26)$$

Replacing the first integral of the RHS of Eq. (1.26) with Eq. (1.20) one obtains

$$\langle \nabla \varphi_l \rangle = \nabla \langle \varphi_l \rangle + \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \varphi_l \cdot \mathbf{n}_l dF. \quad (1.27)$$

This equation expresses the volume average of derivatives in the form of derivatives of volume average and a surface integral. While the derivatives of the non-average property  $\varphi_l$  may be a nonsmooth function of location, the local volume-averaged property  $\langle \varphi_l \rangle$  is a smooth function of location and can be differentiated.

This is an extremely interesting consequence of the *Slattery–Whitaker spatial averaging theorem*, which allows one to write local volume average differential conservation equations for nonstructured as well for structured fluids. For one-dimensional flow Eq. (1.27) reduces to Eq. (3) by *Delhaye*, p. 160, published in 1981.

The analogous form for a vector is

$$\langle \nabla \cdot (\varphi_l \mathbf{V}_l^\tau) \rangle = \nabla \cdot \langle \varphi_l \mathbf{V}_l^\tau \rangle + \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \varphi_l \mathbf{V}_l^\tau \cdot \mathbf{n}_l dF. \quad (1.28)$$

A different route to this result was given by *Gray and Lee* (1977). These authors extended the definition of the field properties in the sense of distributions.

Useful consequences of Eq. (1.27) are obtained by setting  $\varphi_l = 1$

$$\nabla (\alpha_l \gamma_v) = -\frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \mathbf{n}_l dF = \nabla (\alpha_l^e \gamma) \quad (1.29)$$

compare this with Eq. (1.22). Summing over all the fields and keeping in mind (Eq. (1.4a)) that  $\sum_{l=1}^3 \alpha_l = 1$  and  $\mathbf{n}_w = -\mathbf{n}_l$ , the following is obtained:

$$\nabla \gamma_v = -\frac{1}{Vol} \sum_{l=1}^3 \int_{F_{l\sigma}} \mathbf{n}_l dF - \frac{1}{Vol} \sum_{l=1}^3 \int_{F_{lw}} \mathbf{n}_l dF = \frac{1}{Vol} \int_{F_w} \mathbf{n}_w dF . \quad (1.30)$$

While the first integral is a sum of repeating terms with alternating signs and, therefore, is equal to zero, the second integral will be zero only in the case of immersed structure inclusions. Otherwise the second integral is not equal to zero.

### 1.5 General transport equation (*Leibnitz rule*)

The general transport equation applied to the field volume  $Vol_l$  is used in the form

$$\frac{d}{d\tau} \int_{Vol_l} \varphi_l dVol = \int_{Vol_l} \frac{\partial \varphi_l}{\partial \tau} dVol + \int_{F_{l\sigma} + F_{lw} + F_{le}} \varphi_l \mathbf{V}_{l\sigma}^\tau \cdot \mathbf{n}_l dF \quad (1.31)$$

using Eq. (1.14), divided by the constant  $Vol$  and rearranged,

$$\left\langle \frac{\partial \varphi_l}{\partial \tau} \right\rangle = \frac{d}{d\tau} \langle \varphi_l \rangle - \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw} + F_{le}} \varphi_l \mathbf{V}_{l\sigma}^\tau \cdot \mathbf{n}_l dF . \quad (1.32)$$

$\mathbf{V}_{l\sigma}^\tau$  is the instantaneous interface velocity. Replacing the total differential

$$\frac{d}{d\tau} \langle \varphi_l \rangle = \frac{\partial}{\partial \tau} \langle \varphi_l \rangle + \nabla \cdot \langle \varphi_l \mathbf{V}_l^\tau \rangle ,$$

and using Eq. (1.21) we can see that the components at the control volume surfaces cancel and, therefore,

$$\left\langle \frac{\partial \varphi_l}{\partial \tau} \right\rangle = \frac{\partial}{\partial \tau} \langle \varphi_l \rangle - \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \varphi_l \mathbf{V}_{l\sigma}^\tau \cdot \mathbf{n}_l dF . \quad (1.32a)$$

Since  $\langle \varphi_l \rangle = \alpha_l \gamma_v \langle \varphi_l \rangle^l$  (Eq. (1.16)), we finally obtain

$$\left\langle \frac{\partial \varphi_l}{\partial \tau} \right\rangle = \frac{\partial}{\partial \tau} (\alpha_l \gamma_v \langle \varphi_l \rangle^l) - \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \varphi_l \mathbf{V}_{l\sigma}^\tau \cdot \mathbf{n}_l dF . \quad (1.32b)$$

A useful consequence is obtained by setting  $\varphi_l = 1$ :

$$\frac{\partial}{\partial \tau} (\alpha_l \gamma_v) = \frac{1}{Vol} \int_{F_{l\sigma} + F_{lv}} \mathbf{V}_{l\sigma}^\tau \cdot \mathbf{n}_l dF. \quad (1.33)$$

## 1.6 Local volume-averaged mass conservation equation

The principle of the conservation of mass for the control volume  $Vol$  can be expressed verbally as follows:

The change in the mass of the velocity field  $l$  within  $Vol$  with time equals the net mass flow of the field  $l$  through the surface  $F$  and through the interface of the velocity field  $F_{l\sigma}$ .

The conservation equation for the property  $\varphi_l$  valid inside  $Vol_l$  is the classical one:

$$\frac{\partial \varphi_l}{\partial \tau} + \nabla \cdot (\varphi_l \mathbf{V}_l^\tau) = 0. \quad (1.34)$$

Performing volume averaging,

$$\left\langle \frac{\partial \varphi_l}{\partial \tau} \right\rangle + \left\langle \nabla \cdot (\varphi_l \mathbf{V}_l^\tau) \right\rangle = 0, \quad (1.35)$$

and using Eqs. (1.32a) and (1.28), the local average of the derivatives is replaced by derivatives of the local average and additional terms. The result is

$$\frac{\partial}{\partial \tau} \langle \varphi_l \rangle + \nabla \cdot \langle \varphi_l \mathbf{V}_l^\tau \rangle = -\frac{1}{Vol} \int_{F_{l\sigma} + F_{lv}} \varphi_l (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot \mathbf{n}_l dF, \quad (1.36)$$

or, replacing the volume average with its equivalents,

$$\frac{\partial}{\partial \tau} (\alpha_l \gamma_v \langle \varphi_l \rangle^l) + \nabla \cdot (\alpha_l \gamma \langle \varphi_l \mathbf{V}_l^\tau \rangle^{le}) = -\frac{1}{Vol} \int_{F_{l\sigma} + F_{lv}} \varphi_l (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot \mathbf{n}_l dF, \quad (1.37)$$

which is valid for nonstructured as well for structured velocity fields. Finally, the mass conservation equation is easily obtained by setting  $\varphi_l = \rho_l$  and using the definition of the weighted average velocity (Eq. (1.19)). The result is

$$\frac{\partial}{\partial \tau} (\alpha_l \langle \rho_l \rangle^l \gamma_v) + \nabla \cdot (\alpha_l \langle \rho_l \rangle^l \langle \mathbf{V}_l^\tau \rangle^{le} \gamma) = \gamma_v \mu_l^\tau, \quad (1.38)$$

which is Eq. (3.9) or (5.7) in *Sha et al. (1984)* with  $\gamma_v$  incorporated into the time derivatives. Leaving  $\gamma_v$  under the time derivatives allows modeling of structure with varying  $\gamma_v$  over time, e.g., deformable structures that depend on local pressure differences or structures deformed by actions independent of the flow parameter but governing the flow. Here

$$\begin{aligned} \mu_i^\tau &= -\frac{1}{\gamma_v \text{Vol}} \int_{F_{l\sigma} + F_{lw}} \rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot \mathbf{n}_l dF \\ &= -\frac{F_{l\sigma}}{\sum_{l=1}^3 \text{Vol}_l} \frac{1}{F_{l\sigma}} \int_{F_{l\sigma}} \rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot \mathbf{n}_{l\sigma} dF - \frac{F_{lw}}{\sum_{l=1}^3 \text{Vol}_l} \frac{1}{F_{lw}} \int_{F_{lw}} \rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{lw}^\tau) \cdot \mathbf{n}_{lw} dF \end{aligned} \quad (1.39)$$

is the local volume-averaged mass transferred into the velocity field  $l$  per unit time and unit mixture flow volume  $\sum_{l=1}^3 \text{Vol}_l$ . This term is referred to as the local volume-averaged instantaneous mass source density in  $kg/(m^3s)$  for the velocity field  $l$ . Note the difference between the mass source per unit volume of flow,  $\mu_i^\tau$ , and the mass source per unit of the control volume  $\text{Vol}$ ,  $\gamma_v \mu_i^\tau$ . The ratio of the interface surface to the mixture volume  $F_{l\sigma} / \sum_{l=1}^3 \text{Vol}_l$  is referred to in the literature as *interfacial area density*. It is customary in two-phase flow theory to incorporate local volume interface density of the velocity field  $l$  in the form of the interface per unit flow volume  $\sum_{l=1}^3 \text{Vol}_l$  belonging to the control volume  $\text{Vol}$ :

$$a_{l\sigma} = F_{l\sigma} / \sum_{l=1}^3 \text{Vol}_l. \quad (1.40)$$

This is an important dependent variable. In general, it varies with time and location. The means for describing this variable are discussed in detail in Sections 1.10 and 1.11. In addition, the local volume interface density of the structure  $w$  is defined as follows:

$$a_{lw} = F_{lw} / \sum_{l=1}^3 \text{Vol}_l = 4 / D_{hy}, \quad (1.41)$$

where  $D_{hy}$  is the hydraulic diameter (4 times cross-sectional area divided by perimeter).

For numerical integration it should not be forgotten that the volume fraction at the surfaces of the discretization volume may differ from the local volume fraction of the cells, as has already been mentioned, simply because these are associated with

different points in space. Setting these volume fractions to be equal, as is usually done in the widely used donor-cell method, introduces nonphysical diffusion. This kind of diffusion is in no way associated with numerical diffusion. It influences the modeling of the disperse flow pattern less than free surface modeling does. The performance of the numerical model can be improved even for the case of free surfaces in the control volume by extension of the numerical technique developed by *Hirt* and *Nichols* (1981) to cover multiphase flows.

If the interface is immaterial and consequently does not accumulate mass, we have

$$\rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot \mathbf{n}_l + \rho_m (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) \cdot \mathbf{n}_m = 0. \quad (1.42)$$

This is the instantaneous mass balance of the interface. Note that  $\mathbf{n}_l = -\mathbf{n}_m$ , and that the contact discontinuity velocity is common for both sides of the interface,

$$\mathbf{V}_{l\sigma}^\tau = \mathbf{V}_{m\sigma}^\tau = \mathbf{V}_{lm}^\tau. \quad (1.43)$$

Interfaces at which we have

$$\mathbf{V}_l^\tau \cdot \mathbf{n}_l = \mathbf{V}_m^\tau \cdot \mathbf{n}_m \quad (1.44)$$

are called *impermeable*. Such a surface may be a solid–fluid or fluid–fluid interface (*Whitaker* 1977, p. 307). Surfaces at which

$$\mathbf{V}_l^\tau = \mathbf{V}_m^\tau = \mathbf{0} \quad (1.45)$$

are called *impermeable fixed* surfaces, (*Whitaker* 1977, p. 307). It is convenient to write the mass flow rates perpendicular to the interface in the following form:

$$(\rho w)_{lm} = \rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_l \quad (1.46)$$

$$(\rho w)_{ml} = \rho_m (\mathbf{V}_m^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_m. \quad (1.47)$$

Then Eq. (1.42) reads as

$$(\rho w)_{lm} = -(\rho w)_{ml}. \quad (1.48)$$

Consequently the relations between the velocity components normal to the interface and the density are

$$(\mathbf{V}_l^\tau - \mathbf{V}_m^\tau) \cdot \mathbf{n}_l = \left( \frac{1}{\rho_l} - \frac{1}{\rho_m} \right) (\rho w)_{lm}, \quad (1.49)$$

for the case of mass transfer across the interface and

$$V_{lm}^{n,\tau} = \frac{\rho_l V_l^{n,\tau} - \rho_m V_m^{n,\tau}}{\rho_l - \rho_m}, \quad (1.50)$$



$$V_l^{n,\tau} - V_m^{n,\tau} = \frac{\rho_m}{\rho_l - \rho_m} (V_m^{n,\tau} - V_l^{n,\tau}), \quad (1.51)$$

$$V_m^{n,\tau} - V_l^{n,\tau} = \frac{\rho_l}{\rho_l - \rho_m} (V_m^{n,\tau} - V_l^{n,\tau}), \quad (1.52)$$

for “shock wave discontinuity“ without mass transfer. The volume average mass balance of the surface  $F_{l\sigma}$  common to the fields  $m$  and  $l$  is

$$\frac{1}{Vol} \int_{F_{lm}} [\rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) - \rho_m (\mathbf{V}_m^\tau - \mathbf{V}_{lm}^\tau)] \cdot \mathbf{n}_m dF = 0. \quad (1.53)$$

## 1.7 Time average

Splitting time dependent variables, e.g.,  $\Phi^\tau$  and  $\Psi^\tau$ , into their mean  $\Phi$ ,  $\Psi$  and fluctuation parts  $\Phi'$ ,  $\Psi'$ ,

$$\Phi^\tau = \Phi + \Phi', \quad \Psi^\tau = \Psi + \Psi', \quad (1.54)$$

and time averaging them in the following way:

$$\Phi = \frac{1}{2\Delta\tau} \int_{-\Delta\tau}^{\Delta\tau} \Phi^\tau d\tau, \quad (1.55)$$

is known in the literature as *Reynolds* averaging (*Reynolds*. 1894). Here  $\Delta\tau$  is a time scale, large relative to the time scale of turbulent fluctuations, and small relative to the time scale that we wish to resolve. This averaging process has the following properties:

$$\overline{\Phi^\tau} = \Phi, \quad \overline{\Psi^\tau} = \Psi, \quad (1.56)$$

$$\overline{a\Phi^\tau + b\Psi^\tau} = a\Phi + b\Psi, \quad (1.57)$$

$$\overline{\Phi^\tau\Psi^\tau} = \Phi\Psi + \overline{\Phi'\Psi'}, \quad (1.58)$$

$$\frac{\partial\overline{\Phi^\tau}}{\partial\tau} = \frac{\partial\Phi}{\partial\tau}, \quad \nabla\overline{\Phi^\tau} = \nabla\Phi. \quad (1.59)$$

The instantaneous surface-averaged velocity of the field  $l$ ,  $\langle \mathbf{V}_l^\tau \rangle^{le}$ , can be expressed as the sum of the surface-averaged velocity, which is subsequently time averaged,

$$V_l := \overline{\langle \mathbf{V}_l^\tau \rangle^{le}} \quad (1.60)$$

and a pulsation component  $V_l'$ ,

$$\langle \mathbf{V}_l^r \rangle^{le} = V_l + V_l' \quad (1.61)$$

as proposed by *Reynolds*. Substituting for the appropriate terms in the mass conservation equation, performing time averaging, and dropping any averaging signs for the sake of simple notation leads to

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l \gamma_v) + \nabla \cdot (\alpha_l \rho_l \mathbf{V}_l \gamma) = \gamma_v \mu_l. \quad (1.62)$$

For convenience of numerical integration, the source term  $\mu_l$  is split up into a sum of pairs of nonnegative terms:

$$\mu_l = \sum_{m=1}^{l_{\max}, w} (\mu_{ml} - \mu_{lm}). \quad (1.63)$$

Thus, source terms with two subscripts are nonnegative. Two successive subscripts denote the direction of the mass transfer. For instance,  $ml$  denotes the transferred mass per unit time and unit volume of the flow from velocity field  $m$  into  $l$ . As a consequence of this definition, source terms with two identical subscripts are equal to zero  $\mu_{mm} = 0$ .  $w$  is used to denote the region outside of the flow.  $\mu_{wl}$  denotes mass transport from an exterior source into the velocity field  $l$ .

If an interface is immaterial and consequently does not accumulate mass, the following equation holds at this interface:

$$\sum_{l=1}^3 \mu_l = 0. \quad (1.64)$$

For  $\gamma_v = \gamma = 1$ , Eq. (1.62) has been successfully used in thousands of publications in two-phase flow literature, tacitly in the majority of cases in the sense of a local volume average and successively time-averaged equation. *There is no doubt as to its validity* for velocities far below the velocity of light. If used incorrectly, however, this can give rise to nonsolvable numerical problems. An example for usage that can lead to such problems is the use of the left-hand side in the sense of local volume and time average and the right-hand side in the sense of instantaneous sources. For processes with intense mass transfer (condensation shocks, flushing, etc.) this is incorrect. The right-hand side must be the local volume average, with this successively time averaged over the current time step.

After some rearrangements, Eq. (1.62) can be written in the following form:

$$\frac{\partial \alpha_l}{\partial \tau} + \nabla \cdot \left( \frac{\gamma \mathbf{V}_l}{\gamma_v} \alpha_l \right) = \frac{\mu_l}{\rho_l} - \alpha_l \left( \frac{\partial \ln \gamma_v \rho_l}{\partial \tau} + \frac{\gamma \mathbf{V}_l}{\gamma_v} \cdot \nabla \ln \gamma_v \rho_l \right).$$

We can see that the effective convection velocity of the volumetric fraction inside a complex structure is not the velocity itself but the expression  $\gamma \mathbf{V}_i / \gamma_v$ .

To describe multiphase flows by using *gross* spatial discretization the effective flux of the dispersed component caused by turbulent fluctuations has to be taken into account. If no turbulence pulsation forces are taken into account in the momentum equation, the term  $\overline{\rho_i \alpha_i' \mathbf{V}'_i}$  has to be modeled. Idealizing the transversal bubble movement as a diffusion process

$$\left( \overline{\rho \alpha' \mathbf{V}'} \right)_i = -\rho_i D'_{\alpha,12} \nabla \alpha_i,$$

from pipe experimental data within gas mass flow fractions of  $0 < X_1 < 0.04$  *Serizawa et al. (1975)* deduced a diffusion coefficient of the following order of magnitude:

$$D'_{\alpha,12} \approx (1 \div 2.5) \times 10^{-4} \text{ m}^2/\text{s}.$$

Some authors describe diffusion of bubbles using the mixture length approach  $D'_{\alpha,12} \approx v'_2 \ell_{2,mix}$ , where  $v'_2$  is the liquid pulsation velocity. In any case, the process of turbulent bubble diffusion is still not well understood.

## 1.8 Local volume-averaged component conservation equations

The definition given previously should now be recalled, i.e., that each velocity field consists of one noninert component and several inert components. Each component is designated with  $i_l$ . As mentioned in Section 1.2, in order to avoid complicated indices, the designation  $il$  will be used in the following.

The local volume-averaged instantaneous mass conservation equation for the microscopic component  $il$  in the velocity field  $l$  can be expressed as follows:

Inside the part of the control volume filled with the component  $il$ , the net mass flow of the component  $il$  must equal the rate of increase in mass for the component  $il$ , or mathematically:

$$\begin{aligned} \frac{\partial}{\partial \tau} \left( \alpha_{il} \langle \rho_{il} \rangle^l \gamma_v \right) + \nabla \cdot \left( \alpha_{il} \langle \rho_{il} \rangle^l \langle \mathbf{V}_{il}^\tau \rangle^{le} \gamma \right) \\ = - \frac{1}{Vol} \int_{F_{i\sigma} + F_{iv}} \alpha_{il} \rho_{il} (\mathbf{V}_{il}^\tau - \mathbf{V}_{i\sigma}^\tau) \cdot \mathbf{n}_i dF = \gamma_v \mu_{il}^\tau. \end{aligned} \quad (1.65)$$

For a gas mixture it follows from *Dalton's* law that  $\alpha_{il} = \alpha_l$ , whereas for mixtures consisting of liquid and macroscopic solid particles  $\alpha_{il} \neq \alpha_l$ . The instantaneous mass concentration of the component  $i$  in  $l$  is now defined by:

$$C_{il}^\tau = \alpha_{il} \langle \rho_{il} \rangle^l / \left( \alpha_l \langle \rho_l \rangle^l \right). \quad (1.66)$$

The center of mass (c.m.) velocity is given by intrinsic surface-averaged field velocity  $\langle \mathbf{V}_l^\tau \rangle^{le}$  introduced previously:

$$\alpha_l \langle \rho_l \rangle^l \langle \mathbf{V}_l^\tau \rangle^{le} = \sum_{i=1}^{i_{\max}} \alpha_{il} \langle \rho_{il} \rangle^l \langle \mathbf{V}_{il}^\tau \rangle^{le} = \alpha_l \langle \rho_l \rangle^l \sum_{i=1}^{i_{\max}} C_{il}^\tau \langle \mathbf{V}_{il}^\tau \rangle^{le}. \quad (1.67)$$

Consequently

$$\langle \mathbf{V}_l^\tau \rangle^{le} = \sum_{i=1}^{i_{\max}} C_{il}^\tau \langle \mathbf{V}_{il}^\tau \rangle^{le}. \quad (1.68)$$

For the description of transport of the microscopic component  $il$  in the velocity field  $l$  it is convenient to replace the velocity component  $\langle \mathbf{V}_{il}^\tau \rangle^{le}$  by the sum of the center of mass velocity for the particular field  $\langle \mathbf{V}_i^\tau \rangle^{le}$  and the deviation from the c.m. velocity or the so-called *diffusion velocity* of the inert component  $\delta_i \langle \mathbf{V}_i^\tau \rangle^{le}$ , which yields the following:

$$\langle \mathbf{V}_{il}^\tau \rangle^{le} = \langle \mathbf{V}_i^\tau \rangle^{le} + \left( \delta \langle \mathbf{V}_i^\tau \rangle^{le} \right)_i. \quad (1.69)$$

*Fick* (1855) noticed that the mass flow rate of the inert component with respect to the total mass flow rate of the continuous mixture including the inert component is proportional to the gradient of the concentration of this inert component

$$\alpha_{il} \langle \rho_{il} \rangle^l \left( \delta \langle \mathbf{V}_i^\tau \rangle^{le} \right)_i = -\alpha_l \langle \rho_l \rangle^l \delta_l D_{il}^l \nabla C_{il}^\tau. \quad (1.70)$$

The coefficient of proportionality,  $D_{il}^l$ , is known as the coefficient of molecular diffusion. The diffusion mass flow rate is directed from regions with higher concentration to regions with lower concentration, with this reflected by the minus sign in the assumption made by *Fick* (which has subsequently come to be known as *Fick's* law), because many processes in nature and industrial equipment are successfully described mathematically by the above approach – so called diffusion processes. Here  $\delta_l = 1$  for a continuous field and  $\delta_l = 0$  for a disperse field. Substitution in Eq. (1.65) yields

$$\frac{\partial}{\partial \tau} \left( \alpha_l \langle \rho_l \rangle^l C_{il}^\tau \gamma_v \right) + \nabla \cdot \left( \alpha_l \langle \rho_l \rangle^l \langle \mathbf{V}_i^\tau \rangle^{le} C_{il}^\tau \gamma \right) - \nabla \cdot \left( \alpha_l \langle \rho_l \rangle^l \delta_l D_{il}^l \nabla C_{il}^\tau \gamma \right) = \gamma_v \mu_{il}^\tau. \quad (1.71)$$

Molecular diffusion has a microscopic character, as it is caused by molecular interactions. The general expression (1.70) is actually

$$\alpha_{il} \langle \rho_{il} \rangle^l \left( \delta \langle \mathbf{V}_l^\tau \rangle^{le} \right)_i = -\alpha_l \langle \rho_{il} \rangle^l \delta_l D_{il}^l \left( \nabla C_{il}^\tau + \frac{k_{il}^T}{T_l} \nabla T_l + \frac{k_{il}^P}{p} \nabla p \right). \quad (1.72)$$

$k_{il}^T$  and  $k_{il}^P$  are the dimensionless coefficients of thermodiffusion and barodiffusion, which are functions of the component concentrations, *Grigorieva and Zorina* (1988). They tend to zero in the limiting cases of pure substances. Note that

$$\sum_{i=1}^{i_{\max}} D_{il}^l \nabla C_{il}^\tau = 0, \quad (1.73)$$

$$\sum_{i=1}^{i_{\max}} k_{il}^T = 0, \quad (1.74)$$

$$\sum_{i=1}^{i_{\max}} k_{il}^P = 0. \quad (1.75)$$

The effect of the barodiffusion is usually negligible. Thermodiffusion may be substantial if the components have a quite different molecular mass (e.g., a mixture of hydrogen and freon), large temperature gradients, and concentrations about  $1/i_{\max}$ . Usually, thermodiffusion is also neglected in practical computations, *Grigorieva and Zorina* (1988). The special theoretical treatment and the experimental experience of how to determine the molecular diffusion coefficients in multicomponent mixtures, is a science in its own right. This topic is beyond the scope of this book. In this context, it should merely be noted that in line with the thermodynamics of irreversible processes, the thermal diffusivity and the diffusion coefficients influence each other. The interested reader can find useful information in *Reid et al.* (1982).

## 1.9 Local volume- and time-averaged conservation equations

As has already been mentioned, the c.m. velocity of the velocity field can be expressed as time-averaged c.m. velocity and a pulsation component as proposed by *Reynolds*, see Eq. (1.61). The same can be performed for the concentrations,

$$C_{il}^\tau = C_{il} + C_{il}', \quad (1.76)$$

where

$$C_{il} := \overline{C_{il}^\tau}. \quad (1.77)$$

Substituting appropriately in Eq. (1.71), performing time averaging, and omitting averaging signs

$$\overline{\alpha_l \langle \rho_l \rangle^l C_{il}^\tau} = \overline{\alpha_l \langle \rho_l \rangle^l} \overline{C_{il}^\tau} \equiv \alpha_l \rho_l C_{il}, \quad (1.78)$$

$$\overline{\alpha_l \langle \rho_l \rangle^l C_{il}'} = 0, \quad (1.79)$$

$$\overline{\alpha_l \langle \rho_l \rangle^l C_{il} \mathbf{V}_l} = \overline{\alpha_l \langle \rho_l \rangle^l} \overline{C_{il} \mathbf{V}_l} \equiv \alpha_l \rho_l C_{il} \mathbf{V}_l, \quad (1.80)$$

$$\overline{\alpha_l \langle \rho_l \rangle^l C_{il} \mathbf{V}_l'} = 0, \quad (1.81)$$

$$\overline{\alpha_l \langle \rho_l \rangle^l C_{il}' \mathbf{V}_l} = 0, \quad (1.82)$$

$$\overline{\alpha_l \langle \rho_l \rangle^l \delta_l D_{il}' \nabla C_{il}} = \overline{\alpha_l \langle \rho_l \rangle^l} \overline{\delta_l D_{il}' \nabla C_{il}} \equiv \alpha_l \rho_l \delta_l D_{il}' \nabla C_{il}, \quad (1.83)$$

$$\overline{\alpha_l \langle \rho_l \rangle^l \delta_l D_{il}' \nabla C_{il}'} = 0, \quad (1.84)$$

$$\overline{\mu_{il}^\tau} = \mu_{il}, \quad (1.85)$$

the following is obtained:

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l C_{il} \gamma_v) + \nabla \cdot (\alpha_l \rho_l C_{il} \mathbf{V}_l \gamma) + \nabla \cdot (\alpha_l \rho_l \overline{C_{il}' \mathbf{V}_l' \gamma}) - \nabla \cdot (\alpha_l \rho_l \delta_l D_{il}' \nabla C_{il} \gamma) = \gamma_v \mu_{il}. \quad (1.86)$$

For the sake of simplicity, averaging signs are omitted except in the turbulent diffusion term, which will be discussed next. The diffusion can also have macroscopic character, being caused by the macroscopic strokes between large eddies with dimensions considerably larger than the molecular dimensions – turbulent diffusion. In a mixture at rest the molecular strokes are the only mechanism driving diffusion. In real flows both mechanisms are observed. The higher the velocity of the flow, the higher the effect of the turbulent diffusion. *O. Reynolds* assumed for isotropic diffusion

$$\overline{C_{il}' \mathbf{V}_l'} = -D_{il}' \nabla C_{il}, \quad (1.87)$$

where the coefficient of turbulent isotropic diffusion  $D_{il}'$  is proportional to the coefficient of turbulent kinematic viscosity (this is not valid for turbulence of electroconductive liquids in a strong magnetic field):

$$D_{il}' = \nu_l' / Sc', \quad (1.88)$$

where the proportionality is determined by the turbulent *Schmidt* number (e.g.,  $Sc' = 0.77$  if no other information is available). This coefficient is not a thermo-

dynamic property of the material as the molecular coefficient is, but forms a property of the flow. Thus, having for the total diffusion coefficient

$$D_{il}^* = \delta_l D_{il}^l + D_{il}^l, \quad (1.89)$$

in a final step, the local volume and time-averaged mass conservation equation valid for each species  $il$  inside each velocity field  $l$  is obtained:

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l C_{il} \gamma_v) + \nabla \cdot [\alpha_l \rho_l \gamma (\mathbf{V}_l C_{il} - D_{il}^* \nabla C_{il})] \\ & = \gamma_v \sum_{m=1}^{3,w} (\mu_{iml} - \mu_{ilm}) \equiv \gamma_v \mu_{il}. \end{aligned} \quad (1.90)$$

For  $\alpha_l = 1$ ,  $\gamma_v = \gamma = 1$  this is the well known concentration equation from single-phase flow dynamics. There is no doubt as to its general validity as previously mentioned for velocities much smaller than the velocity of light. Equation (1.90) has been successfully used in IVA codes with diffusion terms neglected, see *Kolev* (1985a, 1985b, 1986a, 1986b, 1987, 1990, 1991a, 1991b, 1993a, 1993b, 1993c) and with diffusion terms *Kolev* (1999).

If the surface is immaterial and does not accumulate mass from the inert component, then the mass jump condition is  $\sum_{m=1}^3 \mu_{il}$ .

Note that there is no net mass diffusively transported across a cross-section perpendicular to the strongest concentration gradient. Mathematically it is expressed as follows:

$$\sum_{i=1}^{i_{\max}} (D_{il}^* \nabla C_{il}) = 0, \quad (1.91)$$

or

$$D_{1l}^* \nabla C_{1l} = - \sum_{i=2}^{i_{\max}} (D_{il}^* \nabla C_{il}). \quad (1.92)$$

For monodisperse particles, i.e., particles with constant particle size  $D_{il}$  inside the control volume, Eq. (1.90) can be divided by the mass of the single particle in order to obtain an alternative form

$$\frac{\partial}{\partial \tau} (n_{il} \gamma_v) + \nabla \cdot [\gamma (\mathbf{V}_l n_{il} - D_{il}^* \nabla n_{il})] = \gamma_v \mu_{il} / \left( \rho_{il} \frac{1}{6} \pi D_{il}^3 \right). \quad (1.93)$$

Here

$$n_{il} = \alpha_l \rho_l C_{il} / \left( \rho_{il} \frac{1}{6} \pi D_{il}^3 \right) \quad (1.94)$$

is the number of particles of species  $i$  per unit flow volume.

After expansion of the first two terms using the chain rule and a comparison with the local volume and time-averaged mass conservation equation of the velocity field, the nonconservative form is obtained as follows:

$$\alpha_l \rho_l \left( \gamma_v \frac{\partial C_{il}}{\partial \tau} + \mathbf{V}_l \nabla C_{il} \right) - \nabla \left( \alpha_l \rho_l D_{il}^* \nabla C_{il} \right) = \gamma_v (\mu_{il} - C_{il} \mu_l). \quad (1.95)$$

For numerical integration it is more convenient to use the so called semiconservative form, Eq. (1.64), obtained by splitting the right-hand side into nonnegative source and sink terms and then placing the sources at the left-hand side:

$$\alpha_l \rho_l \left( \gamma_v \frac{\partial C_{il}}{\partial \tau} + \mathbf{V}_l \nabla C_{il} \right) - \nabla \left( \alpha_l \rho_l D_{il}^* \nabla C_{il} \right) + \gamma_v \mu_l^+ C_{il} = \gamma_v DC_{il}. \quad (1.96)$$

Here

$$\mu_{il}^+ = \sum_{\substack{m=1 \\ m \neq l}}^3 \mu_{iml}, \quad (1.97)$$

$$DC_{il} = \mu_{il} - C_{il} (\mu_l - \mu_l^+). \quad (1.98)$$

For mass transport among the velocity fields and between the exterior sources and the velocity fields it is assumed that the mass convectively leaving the velocity field has the concentrations of the inert components and the mass entering the field has the concentrations of the donor field.

For the particular case of flow consisting of air-steam and two additional liquid fields with one species of solid particles, as postulated for the IVA codes, *Kolev* (1985a, 1986a, 1986b, 1987, 1991a, 1991b, 1993b, 1993c, 1993d, 1993e, 1993f, 1994a, 1996, 1998, 1999), the following are obtained

$$DC_{i1} = C_{i1} (\mu_{12} + \mu_{13}), \quad (1.99)$$

$$DC_{i2} = C_{i2} \mu_{21} + C_{i3} \mu_{32}, \quad (1.100)$$

$$DC_{i3} = C_{i3} \mu_{31} + C_{i2} \mu_{23}, \quad (1.101)$$

and

$$\mu_1^+ = \mu_{21} + \mu_{31}, \quad (1.102)$$

$$\mu_2^+ = \mu_{12} + \mu_{32}, \quad (1.103)$$



$$\mu_{31}^+ = \mu_{13} + \mu_{23}, \quad (1.104)$$

or, in abbreviated form

$$\mu_{il}^+ = \mu_i^+. \quad (1.105)$$

The simplicity of the concentration equation is the reason for the choice of  $C_{il}$  as elements of the dependent variables' vector describing the flow. As a consequence of this choice, the equation of state for the multicomponent mixture,  $\rho_l = \rho_l(p, T_l, C_{il}$  for all  $i$ ), has to be derived from the equation of state for the elementary components of the mixture,  $\rho_{il} = \rho_{il}(p_{il}, T_{il})$ . This has already been performed by the author and published in *Kolev* (1990). An extended variant of this work is given Chapter 3 of this monograph.

Finally, let us write the mass jump condition at the interface.

Instant:

$$\alpha_{il}^{\tau\sigma} \rho_{il}^\sigma (\mathbf{V}_{il}^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_l + \alpha_{im}^{\tau\sigma} \rho_{im}^\sigma (\mathbf{V}_{im}^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_m = 0. \quad (1.106)$$

In terms of local volume-averaged parameter:

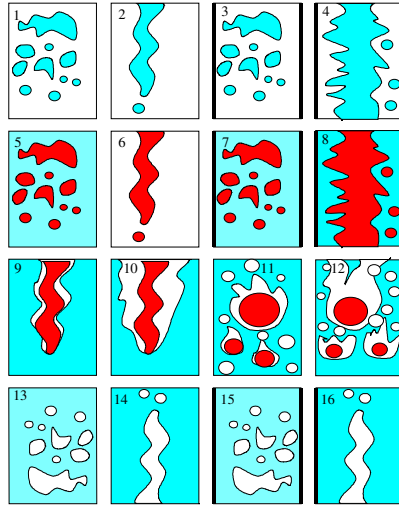
$$\begin{aligned} & C_{il}^{\tau\sigma} \rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_l - \rho_l D_{il}^l \nabla C_{il}^{\tau\sigma} \cdot \mathbf{n}_l \\ & + C_{im}^{\tau\sigma} \rho_m (\mathbf{V}_m^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_m - \rho_m D_{im}^l \nabla C_{im}^{\tau\sigma} \cdot \mathbf{n}_m = 0. \end{aligned} \quad (1.107)$$

Local volume and time average:

$$\begin{aligned} & C_{il}^{m\sigma} \rho_l (\mathbf{V}_l - \mathbf{V}_{lm}) \cdot \mathbf{n}_l - \rho_l D_{il}^* \nabla C_{il}^{m\sigma} \cdot \mathbf{n}_l \\ & + C_{im}^{l\sigma} \rho_m (\mathbf{V}_m - \mathbf{V}_{lm}) \cdot \mathbf{n}_m - \rho_m D_{im}^* \nabla C_{im}^{l\sigma} \cdot \mathbf{n}_m = 0. \end{aligned} \quad (1.108)$$

## 1.10 Conservation equations for the number density of particles

Modeling of gases and liquids that disintegrate in a finite time from a continuum to a spectrum of particles and vice versa, see Fig. 1.4, requires, in addition to the mass conservation equation, equations balancing the number of particles per



**Fig. 1.4** Fragmentation mechanism. (1) Acceleration and turbulence-induced droplet fragmentation in pool flow; (2) jet disintegration in pool; (3) acceleration and turbulence-induced droplet fragmentation in channels; (4) jet fragmentation in channels; (5) droplet in pool; (6) jet in pool; (7) droplet in channel; (8) jet in channel; Liquid metal jet disintegration in liquid: (9) with film boiling – thin film; (10) with film boiling and strong radiation (thick film); liquid metal droplet disintegration in liquid: (11) with film boiling – thin film; (12) with film boiling and strong radiation (thick film); (13) bubble fragmentation in pool; (14) gas jet disintegration in pool; (15) bubble fragmentation in channels; and (16) gas jet disintegration in channels

unit volume of the mixture. Such particles may originate from nucleation or may be the result of convection from the neighboring cells.

If the flow is dispersed there are several possible reasons for the presence of a spectrum of particle sizes in the local control volume. A practicable simplification is the used concept of monodispersity in the control volume. This concept is associated with the assumption that the particles in each control volume are represented by a single local volume and time-average particle size. As the volume fraction  $\alpha_l$  and the local volume and time-average particle number density  $n_l$  are known, it is then possible to compute the representative local volume and time-average particle diameter

$$D_l = \left( \frac{6 \alpha_l}{\pi n_l} \right)^{1/3} \quad (1.109)$$

for any control volume. The approach used to obtain Eq. (1.93) illustrates the philosophy behind the method for obtaining the equation for the conservation of the number of macroscopic particles (droplets or bubbles) per unit flow volume:

$$\frac{\partial}{\partial \tau} (n_l \gamma_v) + \nabla \cdot \left[ \left( \mathbf{V}_l n_l - \frac{\mathbf{v}_l^t}{Sc^m} \nabla n_l \right) \gamma \right] = \gamma_v (\dot{n}_{l,kin} - \dot{n}_{l,coal} + \dot{n}_{l,sp}), \quad (1.110)$$

for  $\alpha_l > 0$ . For simplicity, any averaging signs are omitted.  $n_l$  takes values between zero and approximately  $10^{12}$  per cubic meter ( $10^{12} \text{ m}^{-3}$  is presumably the upper limit for nucleation during evaporation or condensation).  $n_l = 0$  for an existing velocity field here will be associated with a continuous fluid.  $0 < n_l Vol_{cel} < 1$ , where  $Vol_{cel}$  is the volume under consideration, can be interpreted as a continuous fluid (jet or free surface) with an excited interface.  $n_l Vol_{cel} \rightarrow 0$  means a stable surface and  $n_l Vol_{cel} \rightarrow 1$  means an unstable interface.  $n_l Vol_{cel} > 1$  means that there is no longer a continuous velocity field  $l$ . Here  $-\mathbf{v}_l^m / Sc^m$  is the diffusive flux for particle number density resulting from the fact that the particles are in random motion in the presence of a particle density gradient. In accordance with *Batchelor* (1970),

$$-\frac{\mathbf{v}_l^m}{Sc^m} \approx const D_l \overline{(\mathbf{V}_l^t \cdot \mathbf{V}_l^t)^{1/2}} = const D_l (\Delta \mathbf{V}_{lm} \cdot \Delta \mathbf{V}_{lm})^{1/2} [H(\alpha_l)]^{1/2}, \quad (1.111)$$

where the constant is approximately unity,

$$H(\alpha_l) \approx \frac{\alpha_l}{\alpha_{l,max}} \left( 1 - \frac{\alpha_l}{\alpha_{l,max}} \right), \quad (1.112)$$

and  $\alpha_{l,max} \approx 0.62$  is the limit for the closest possible packing of particles.  $\overline{(\mathbf{V}_l^t \cdot \mathbf{V}_l^t)}$  is the mean square of the velocity fluctuations and  $(\Delta \mathbf{V}_{lm} \cdot \Delta \mathbf{V}_{lm})^{1/2}$  is the magnitude of the relative velocity. This choice meets the requirements for disappearance of velocity fluctuations at the limit  $\alpha_l \rightarrow 0$  and at the limit for the closest possible packing  $\alpha_l \rightarrow \alpha_{l,max}$ .

For the locally monodisperse system there is a unique relationship between volume, concentration, particle number density, and the local particle length scale or the interfacial area density  $a_d$  (interfacial area  $F_d$  divided by the mixture flow

volume  $\sum_{l=1}^{l_{max}} Vol_l$ ),

$$\begin{aligned} a_d &= F_d / \sum_{l=1}^{l_{max}} Vol_l = n_d F_{d, \text{single particle}} = \alpha_d F_{d, \text{single particle}} / Vol_{d, \text{single particle}} \\ &= \alpha_d \frac{6}{D_d} = 6 \left( \frac{\pi}{6} n_d \right)^{1/3} \alpha_d^{2/3}. \end{aligned} \quad (1.113)$$

Therefore the use of one of the variables  $n_d$  or  $D_d$  or  $a_d$  is equivalent and only the simplicity of the conservation equation dictates preference for  $n_d$ .

The production terms on the right-hand side of Eq. (1.109) can be classified as kinetic and nonkinetic terms. The kinetic terms are

$$\dot{n}_{l,kin} = f_{lw} \frac{4}{D_{hy}} n_{lw}'' + \alpha_m f_{im} \sum_{i=1}^{i_{max}} n_{im} + \alpha_m (\dot{n}_{l,h} + \dot{n}_{l,dis}). \quad (1.114)$$

In the case of bubble flow,  $l = 1$ ,  $m = 2$ . For this case, the following then apply:

$f_{lw} \frac{4}{D_{hy}} n_{lw}''$  is the number of bubbles generated at the wall per unit time and unit volume of the flow (wall cavity nucleation rate).  $f_{lw}$  is the frequency of bubble generation at one activated seed on the channel wall.  $n_{lw}''$  is the number of activated seeds per unit area of the wall.

The term  $\alpha_m f_{im} \sum_{i=1}^{i_{max}} n_{im}$  gives the number of bubbles generated from the solid particles homogeneously mixed with the second velocity field per unit time and unit volume of the flow.  $f_{im}$  is the frequency of bubbles generated from one activated seed on the particle belonging to the second velocity field.  $\dot{n}_{l,h} + \dot{n}_{l,dis}$  is the bulk liquid nucleation rate consisting of the number of bubbles generated by homogeneous nucleation in the second velocity field per unit time and unit volume of the flow,  $\dot{n}_{l,h}$ , and the number of bubbles generated from the dissolved gases in the second velocity field per unit time and unit volume of the flow,  $\dot{n}_{l,dis}$ .

The nonkinetic *mechanical* terms are

$$\dot{n}_{l,coal} = n_l f_{l,coal} / 2, \quad (1.115)$$

which denotes the number of bubbles that disappear due to *coalescence* per unit time and unit volume of the flow, and

$$\dot{n}_{l,sp} = n_l f_{l,sp}, \quad (1.116)$$

which is the number of the bubbles arising due to hydrodynamic bubble disintegration per unit time and unit volume of the flow.  $f_{l,coal}$  is the *coalescence frequency of two colliding particles*.  $f_{l,sp}$  is the *fragmentation frequency of single particle*. The coalescence frequency of a single particle is defined as a product of the *collision frequency per single particle*  $f_{d,col}$  and the *coalescence probability* of two colliding particles  $P_{d,coal}$ . Different causes for the collisions are associated with different collision frequencies  $f_{d,col}^s$ ,  $f_{d,col}^{no}$  and  $f_{d,col}^o$ . Similarly, different causes for collisions are associated with different

coalescence probabilities  $P_{d,coal}^{no}$ ,  $P_{d,coal}^o$ .  $P_{d,coal}^o$  is the probability of oscillatory coalescence (due to turbulent oscillations).  $P_{d,coal}^{no}$  is the probability of nonoscillatory coalescence (due to nonuniform velocity field in space). Thus the final expression takes the form

$$f_{d,coal} = (f_{d,coal}^s + f_{d,coal}^{no}) P_{d,coal}^{no} + f_{d,coal}^o P_{d,coal}^o. \quad (1.117)$$

$f_{d,coal}^o$  is the frequency of collision due to turbulence fluctuation of the particles. It depends on the fluctuation component of the velocity of the particles,  $V_d'$ ,

$$f_{d,coal}^o = f(\Delta V_d', \alpha_d, \dots, \dots). \quad (1.118)$$

The superscript  $o$  is used to remind us that this is an oscillatory frequency.  $f_{d,coal}^s$  and  $f_{d,coal}^{no}$  are the frequencies of collision due to the convective motion. The splitting into two components is due to the use of the concept of monodispersity. The computational collisions,  $f_{d,coal}^{no}$ , are caused only by the mean relative velocity between the field with averaged particle size and the surrounding continuum  $\Delta V_{dd}^{no}$ . Correct mathematical averaging of the velocity difference gives an additional component  $\Delta V_{dd}^s$  associated with  $f_{d,coal}^s$ .  $f_{d,coal}^s$  is zero for real monodisperse systems. Thus

$$f_{d,coal}^{no} = f(\Delta V_{dd}^{no}, \alpha_d, \dots, \dots), \quad (1.119)$$

and

$$f_{d,coal}^s = f(\Delta V_{dd}^s, \alpha_d, \dots, \dots). \quad (1.120)$$

The superscript  $no$  is used to remind us that this is a nonoscillatory frequency. The superscript  $s$  is used to remind us that this frequency is associated with a spectrum of particle sizes. The source terms for the conservation equation for droplet number density,  $l = 3$  are defined in a similar way.

For channel flow it is assumed here that the second velocity field is continuous and that the third is disperse. In this case, besides the fragmentation and coalescence that also exist in pool flow, the entrainment and deposition will influence not only the mass but also the particle number density balance

$$\mu_{23} - \mu_{32} = \frac{4}{D_h} (1 - \alpha_2)^{1/2} [(\rho w)_{23} - (\rho w)_{32}], \quad (1.121)$$

$$\dot{n}_{23} - \dot{n}_{32} = \frac{6}{\pi} \frac{4}{D_h} (1 - \alpha_2)^{1/2} [(\rho w)_{23} / (\rho_2 D_{3E}^3) - (\rho w)_{32} / (\rho_3 D_3^3)], \quad (1.122)$$

where  $(\rho w)_{23}$  and  $(\rho w)_{32}$  are entrainment and deposition mass flow rate, and  $D_{3E}$  is the diameter of the entrained droplets.

The large variety of phenomena leading to fragmentation and coalescence are discussed by the author in *Kolev* (1993a) and more recently in *Kolev* (1998). In Volume II of this monograph the reader will find the current status of empirical knowledge in this field and the way in which this can be used to compute the fragmentation and coalescence production rates.

It is now appropriate to highlight an interesting potential of this concept. During numerical integration the multiplying of  $n_l$  with  $\sum_{l=1}^3 Vol_l$  gives the number of particles in a single control volume or in a single discretization volume. If there is more than one particle in the cell,

$$n_l \sum_{l=1}^3 Vol_l > 1, \quad (1.123)$$

the field is dispersed. Consequently, the mechanisms governing fragmentation and coalescence are the mechanisms for dispersed field  $l$  surrounded by  $m$ , e.g., droplets or bubble fragmentation and coalescence. If

$$n_l \sum_{l=1}^3 Vol_l \leq 1, \quad (1.124)$$

the field is continuous. This means that the mechanisms controlling potential fragmentation are the mechanisms known for continuous field  $l$  surrounded by the field  $m$ . An example is jet fragmentation, where both participating fields at the jet region are continuous. In this case there is no coalescence. Thus, natural transition from continuum to disperse and vice versa can be modeled. In addition, a very important memory effect of the multiphase structure is taken into account in this modeling approach. An example of the use of this approach was given in *Kolev* (1993b) for interaction and fragmentation of gas, molten metal, and water during a transient.

Equation (1.109) without diffusion and nonkinetic production terms was successfully used by *Kocamustafaogulari* and *Ishii* (1983) for the modeling of single-component boiling systems. A comparison of nonequilibrium model predictions with experimental data for flashing in *Laval* nozzles performed by the author (1985b) showed that an additional differential equation for the description of the particle number density is necessary to obtain a more accurate prediction than the widespread approach of assuming an almost arbitrary number density within the range of  $10^9$  to  $10^{13}$  cm<sup>-3</sup>. *Riznic* and *Ishii* (1989) applied the *Kocamustafaogulari* and *Ishii* approach with success to the modeling of single-component flashing systems. *Deich* and *Philipoff* (1981) analyzed the pressure and temperature distribution inside the eddies of subcooled steam and came to the conclusion that eddies with appropriate dimensions serve as nuclei for condensation, which provides further substantiation for the use of the method discussed above.

The derived particle number density conservation equation can be used to compute the surface energy associated with the interface. Multiplying the bubble number density,  $n_l$ , by the surface energy of a single bubble,  $\pi D_1^2 \sigma_{12}$ , the surface energy per unit mixture volume is then obtained. The conservation equation for this energy is similar to Eq. (1.109). It is interesting to note that the surface energy is transported by the convection and diffusion of the *discrete* velocity fields but that the terms supplying the change in this energy (kinetic origination, collision, splitting) result from the energy of the surrounding *continuous* liquid.

## 1.11 Implication of the assumption of monodispersity in a cell

### 1.11.1 Particle size spectrum and averaging

The result of the initial, boundary conditions, convection, local collision and coalescence, fragmentation entrainment, and deposition is a spectrum of particle size. In this section we will demonstrate what influence the replacement of the spectrum of particle sizes with a single particle size may have on heat and mass transfer processes and how it can be incorporated approximately in the analysis.

To illustrate this we use the *Nukama–Tanasawa* distribution observed in 1931

$$\frac{P(D_{di})}{D'_d} = 4 \left( \frac{D_{di}}{D'_d} \right)^2 e^{-2(D_{di}/D'_d)}. \quad (1.125)$$

Here  $P(D_{di})$  is the probability that a particle size is between  $D_{di}$  and  $D_{di} + \delta D_{di}$ .  $D'_d$  is the most probable particle size, i.e., the size where the probability distribution function has its maximum value. The particle sizes may take values between zero and a maximum value

$$0 < D_{di} < D_{d,\max}. \quad (1.126)$$

Thus, if we know  $D'_d$  and  $D_{d,\max}$ , the particle distribution is uniquely characterized. *MacVean* (see *Wallis* (1969)) found that a great deal of data could be correlated by assuming that

$$D'_d = D_d / 2, \quad (1.127)$$

where  $D_d$  is the volume-averaged particle size. The relationship between  $D_d$  and the maximum particle size,  $D_{d,\max}$ , is reported as

$$D_{d,\max} \approx (2.04 \text{ to } 3.13) D_d \quad (1.128)$$

(see *Pilch* et al. (1981) and *Kataoka* et al. (1983), among others). Other distributions give slightly different results, e.g., *Kolomentzev* and *Dushkin* (1985)

$$\frac{P(D_{di})}{D'_d} = 8 \left( \frac{D_{di}}{D'_d} \right)^2 e^{-4(D_{di}/D'_d)^2} \quad (1.129)$$

came to  $D_{d,\max} \approx 1.33 D_d$ . Other distributions are reported by *Rosin and Rammler* (1933), *Griffith* (1943), and *Mugele and Evans* (1951).

In contrast to Eq. (1.113), for monodisperse spherical particles the interfacial area density in the general case is

$$\begin{aligned} a_d &= F_d \left/ \sum_{l=1}^{l_{\max}} Vol_l \right. = n_d F_{d,\text{single particle}} = \alpha_d F_{d,\text{single particle}} / Vol_{d,\text{single particle}} = \\ \frac{6\alpha_d}{D_d^{sm}} &= \frac{6\alpha_d}{D_d} \frac{D_d}{D_d^{sm}}. \end{aligned} \quad (1.130)$$

Here

$$D_d^{sm} = \left[ \int_0^{D_{d,\max}} D_d^3 P(D_d) dD_d \right/ \int_0^{D_{d,\max}} D_d^2 P(D_d) dD_d \right] \quad (1.131)$$

is the so-called *Sauter* mean diameter, see *Sauter* (1929). A droplet with a diameter equal to the *Sauter* mean diameter has the same surface to volume ratio as the entire spray. Using the *Nukama–Tanasawa* distribution, Eq. (1.125), and Eqs. (1.127) and (1.128), after evaluating Eq. (1.131) analytically we obtain for the *Sauter* mean diameter

$$D_d^{sm} \approx (1.154 \text{ to } 1.238) D_d \quad (D_d^{sm} < 1.25 D_d),$$

or

$$D_d / D_d^{sm} \approx 0.867 \text{ to } 0.807 > 0.8, \quad (1.132)$$

which is perfectly confirmed by the measurements summarized by *Faeth* (1995) –  $D_d^{sm} / D_d \approx 1.2$ . This is a very useful result, allowing use of the volume fraction  $\alpha_d$  and the particle number density  $n_d$  in cases where the assumption of monodispersity does not hold.

### 1.11.2 Cutting of the lower part of the spectrum due to mass transfer

The use of the particle number density conservation equations for each of the three velocity fields as already implemented in the IVA computer code series since 1985 is an extremely practical method of modeling the scale of the field. A number of important features of the concept of monodispersity will now be discussed.

The time-averaged source terms associated with the change in the aggregate state for the mass conservation equations are



$$\begin{aligned}
\mu_i &= \frac{1}{\Delta\tau} \int_0^{\Delta\tau} \frac{d}{d\tau} \left( n_i \rho_i \frac{\pi}{6} D_i^3 \right) d\tau = \left[ \Delta \left( n_i \rho_i \frac{\pi}{6} D_i^3 \right) \right] / \Delta\tau \\
&= \rho_{i0} \frac{\pi}{6} \left[ D_{id}^3 f_{iw} \frac{4}{D_{iy}} n_{iw}'' + D_{idm}^3 \alpha_m f_{im} \sum_{i=1}^{i_{\max}} n_{im} + \alpha_m (D_{i,h}^3 \dot{n}_{i,h} + D_{i,dis}^3 \dot{n}_{i,dis}) \right] \\
&+ \alpha_{i0} \rho_{i0} \left[ \rho_i D_i^3 / (\rho_{i0} D_{i0}^3) - 1 \right] / \Delta\tau. \tag{1.133}
\end{aligned}$$

Here  $\rho_{i0}, n_{i0}$  and  $D_{i0}$  are density, particle number density, and particle size at the beginning of the time step  $\Delta\tau$ .  $D_{id}$  is the size with which the bubbles are produced after one nucleation cycle on the solid structure, i.e., the particle departure diameter.  $D_{idm}$  is the size with which the bubbles are produced after one nucleation cycle on the inert solid particles of the field  $m = 2$ . If the nucleation takes place in the bulk of the donor field, the size is equal to the smallest stable size,  $D_{i,h}$ , or in case of nucleation on dissolved gases,  $D_{i,dis}$ . More information on the modeling of particulate processes is provided in Volume II of this monograph.

The first term in the above equation gives the time and local volume-averaged mass production due to nucleation; the second term gives the time and local volume average of mass production due to mass change for existing particles.

It is important to note that the second term describes changes in the particle size due to evaporation and condensation. In accordance with the concept of monodispersity, the mass changes leading to a decrease in the local volume and time-average particle size also cause the disappearance of those particles in the spectrum with sizes smaller than the size change. This statement, together with the assumption made for the form of the spectrum, provides the basics for computation of the averaged number of disappearing particles. Suppose the particle size distribution obeys the *Nukiyama-Tanasawa* (1931) law

$$P(D_i) = 4(D_i / D_i')^2 \exp[-2(D_i / D_i')]. \tag{1.134}$$

Suppose the form of the distribution remains unchanged during the time step. Having in mind that  $\mu_{im} = -\alpha_{i0} \rho_{i0} \left[ \rho_i D_i^3 / (\rho_{i0} D_{i0}^3) - 1 \right] / \Delta\tau$  the volume-averaged diameter changes within this time by

$$\begin{aligned}
\Delta D_i &= D_{i0} - D_i = D_{i0} (1 - D_i / D_{i0}) = D_{i0} \left\{ 1 - \left[ \left( 1 - \frac{\mu_{im} \Delta\tau}{\alpha_{i0} \rho_{i0}} \right) \frac{\rho_{i0}}{\rho_i} \right]^{1/3} \right\} \\
&\approx D_{i0} \left\{ 1 - \left[ \left( 1 - \frac{\mu_{im} \Delta\tau}{\alpha_{i0} \rho_{i0}} \right) \right]^{1/3} \right\}. \tag{1.135}
\end{aligned}$$

and all particles having sizes  $0 \leq D_i \leq 2\Delta D_i$ , namely,

$$n_i \int_0^{2\Delta D_i} P(D_i) dD_i = 2n_i \left\{ 1 - \exp\left(-4 \frac{\Delta D_i}{D_i'}\right) \left[ 1 + 4 \frac{\Delta D_i}{D_i'} \left( 1 + 2 \frac{\Delta D_i}{D_i'} \right) \right] \right\}, \quad (1.136)$$

disappear. The averaged particle sink per unit time and unit mixture volume is consequently

$$\dot{n}_{i,\text{spectrum\_cut}} = \left( \frac{n_i}{\Delta \tau} \right) 2 \left\{ 1 - \exp\left(-4 \frac{\Delta D_i}{D_i'}\right) \left[ 1 + 4 \frac{\Delta D_i}{D_i'} \left( 1 + 2 \frac{\Delta D_i}{D_i'} \right) \right] \right\}, \quad (1.137)$$

for  $\Delta D_i \geq D_i'$ . The latter condition means that the averaged particle size is  $\Delta D_i$ , with particle sizes ranging approximately within the interval  $2\Delta D_i$ . It was found that for shock condensation of bubbles the inclusion of Eq. (1.137) for prediction of a reduction in the number density for disappearance of steam volume fractions is very important, *Kolev* (1993d, 1993e, 1993f, 1996, 1999). The same is true for the disappearance of water droplets in highly superheated gases due to evaporation. The latter can be the case in steam explosion analysis where gas is superheated due to previous contact with melt followed by water droplet entry into the superheated gas regions. Combustion processes and spray cooling of hot gas jets are further examples.

Equation (1.137) can be written in the following form:

$$\dot{n}_{i,\text{spectrum\_cut}} = n_i f_{i,\text{spectrum\_cut}}, \quad (1.138)$$

where

$$f_{i,\text{spectrum\_cut}} = \frac{2}{\Delta \tau} \left\{ 1 - \exp\left(-4 \frac{\Delta D_i}{D_i'}\right) \left[ 1 + 4 \frac{\Delta D_i}{D_i'} \left( 1 + 2 \frac{\Delta D_i}{D_i'} \right) \right] \right\} \quad (1.139)$$

is the frequency of particle disappearance due to spectrum cutting.

### 1.11.3 The effect of averaging on the effective velocity difference

Consider  $i$  groups of particles with different diameters having a concentration per unit volume called particle number density

$$n_{di} = n_d P(D_{di}) \quad i = 1, I, \quad (1.140)$$

which is dependent on the diameter. Here  $P(D_{di})$  is the probability of a particle having a size between  $D_{di}$  and  $D_{di} + \delta D_{di}$ . We assume that the continuous velocity field possesses a constant velocity  $w_c$ . The difference of the continuum and the particle velocity is approximately described by the following equations:

$$\left(1 + b_d c_{cd}^{vm}\right) \frac{\partial \Delta w_{cd}}{\partial \tau} + \left(\frac{1}{\rho_c} - \frac{1}{\rho_d}\right) \nabla p + b_d \frac{3}{4} \frac{c_{cd}^d}{D_d} |\Delta w_{cd}| \Delta w_{cd} = 0, \quad (1.141)$$

$$\left(1 + b_{di} c_{cdi}^{vm}\right) \frac{\partial \Delta w_{cdi}}{\partial \tau} + \left(\frac{1}{\rho_c} - \frac{1}{\rho_d}\right) \nabla p + b_{di} \frac{3}{4} \frac{c_{cdi}^d}{D_{di}} |\Delta w_{cdi}| \Delta w_{cdi} = 0, \quad (1.142)$$

where  $b_d = (\alpha_c \rho_c + \alpha_d \rho_d) / (\rho_c \rho_d)$ , and  $b_{di} = (\alpha_c \rho_c + \alpha_{di} \rho_d) / (\rho_c \rho_d)$ . In this case we see that each particular group possesses a different relative velocity with respect to the continuum and consequently the groups move relatively to each other. Therefore, *collision caused by particle-particle relative velocity in real nature is an inevitable phenomenon*. For steady flow from Eqs. (1.141) and (1.142) we have

$$\Delta w_{cdi} = \Delta w_{cd} \left( \frac{b_d}{b_{di}} \frac{c_d^d}{c_{cdi}^d} \frac{D_{di}}{D_d} \right)^{1/2} = \Delta w_{cd} f_i. \quad (1.143)$$

Obviously there is a resulting average velocity

$$\Delta V_{dd} = \frac{1}{n_d} \sum_i n_{di} |\Delta w_{cd} - \Delta w_{cdi}| = |\Delta w_{cd}| \frac{1}{n_d} \sum_i n_{di} |1 - f_i| = |\Delta w_{cd}| \left| \sum_i P(D_{di}) |1 - f_i| \right|, \quad (1.144)$$

which is the

$$\frac{1}{n_d} \sum_i n_{di} |1 - f_i| \quad (1.145)$$

part of the average relative velocity of the droplets with respect to the continuum. For monodispersed particles

$$f_1 = f_2 = \dots = f_i = 1 \quad (1.146)$$

and

$$\Delta V_{dd} = 0. \quad (1.147)$$

The expression

$$\frac{1}{n_d} \sum_i n_{di} |1 - f_i| = \int_0^{D_{d,\max}/D_d} P(D_{di}) |1 - f_i| d(D_{di}/D_d), \quad (1.148)$$

can be numerically estimated using known distribution functions, e.g., the *Nukia-ma-Tanasava* distribution (1939) as given by Eq. (1.125). Furthermore,

$$f_i = \left( \frac{b^* + 1}{b^* + \alpha_{di} / \alpha_d} \frac{c_d^d}{c_{cdi}^d} \frac{D_{di}}{D_d} \right)^{1/2}, \quad (1.149)$$

where  $b^* = \alpha_c \rho_c / \alpha_d \rho_d$  can be expressed as function of  $D_{di} / D_d$ , having in mind that  $\alpha_{di} = (\pi D_{di}^3 / 6) n_{di}$  and  $\alpha_d = (\pi D_d^3 / 6) n_d$  and using Eq. (1.140),

$$\alpha_{di} / \alpha_d = (D_{di} / D_d)^3 P(D_{di}). \quad (1.150)$$

## 1.12 Stratified structure

If the selected scale of discretization  $\delta$  is smaller than the characteristic length of the fields,

$$\delta |\kappa_l| < 1, \quad (1.151)$$

the local control volume contains a surface dividing for instance field  $l$  from field  $m$ . The unit vector  $\mathbf{n}_l$  pointing outwards from the field  $l$  is an important local characteristic of this surface. It is defined from the spatial distribution of the volume fraction of the field  $l$  as follows:

$$\mathbf{n}_l = - \frac{\nabla(\alpha_l \gamma)}{|\nabla(\alpha_l \gamma)|}. \quad (1.152)$$

The curvature of the surface is conveniently computed following the derivation of *Brackbill et al.* (1992),

$$\kappa_l = \nabla \cdot \mathbf{n}_l, \quad (1.153)$$

and is defined as positive if the center of the curvature is in field  $l$ .

## 1.13 Final remarks and conclusions

Local volume averaging as founded by *Anderson and Jackson, Slattery, Whitaker*, is applied to derive mass conservation equations for multiphase multicomponent flows conditionally divided into three velocity fields. The heterogeneous porous media formulation introduced by *Gentry, Martin and Daly*, commented on by *Hirt*, and used by *Sha, Chao and Soo*, was then implanted into the formalism as a geometrical skeleton because of its practical usefulness. Beyond these concepts, inert components in each velocity field are included, and the concept of dynamic particle fragmentation and coalescence is introduced. Then subsequent time averaging was performed.

The result of this derivation yields the following three local volume and time-average equations applicable for each velocity field. Here, the equations are written in the scalar form for the most frequently used Cartesian and cylindrical coordinate systems to simplify their direct use by the reader for his particular application.

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (\alpha_l \rho_l \gamma_v) + \frac{1}{r^\kappa} \frac{\partial}{\partial r} (r^\kappa \alpha_l \rho_l u_l \gamma_r) + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} (\alpha_l \rho_l v_l \gamma_\theta) + \frac{\partial}{\partial z} (\alpha_l \rho_l w_l \gamma_z) \\
& = \gamma_v \sum_{m=1}^{l_{\max}, w} (\mu_{ml} - \mu_{lm}), \tag{1.154}
\end{aligned}$$

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (\alpha_l \rho_l C_{il} \gamma_v) + \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left[ r^\kappa \alpha_l \rho_l \left( u_l C_{il} - D_{il}^* \frac{\partial C_{il}}{\partial r} \right) \gamma_r \right] \\
& + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left[ \alpha_l \rho_l \left( v_l C_{il} - D_{il}^* \frac{1}{r^\kappa} \frac{\partial C_{il}}{\partial \theta} \right) \gamma_\theta \right] + \frac{\partial}{\partial z} \left[ \alpha_l \rho_l \left( w_l C_{il} - D_{il}^* \frac{\partial C_{il}}{\partial z} \right) \gamma_z \right] \\
& = \gamma_v \sum_{m=1}^{l_{\max}, w} (\mu_{iml} - \mu_{ilm}) \quad \text{for } \alpha_l \geq 0, \tag{1.155}
\end{aligned}$$

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (n_l \gamma_v) + \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left[ r^\kappa \left( u_l n_l - \frac{v_l^m}{Sc_l^m} \frac{\partial n_l}{\partial r} \right) \gamma_r \right] + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left[ \left( v_l n_l - \frac{v_l^t}{Sc_l^t} \frac{1}{r^\kappa} \frac{\partial n_l}{\partial \theta} \right) \gamma_\theta \right] \\
& + \frac{\partial}{\partial z} \left[ \left( w_l n_l - \frac{v_l^t}{Sc_l^t} \frac{\partial n_l}{\partial z} \right) \gamma_z \right] = \gamma_v (\dot{n}_{l,kin} - \dot{n}_{l,coal} + \dot{n}_{l,sp}) \\
& \equiv \gamma_v n_l \left\{ \frac{1}{n_l} \dot{n}_{l,kin} + f_{l,sp} - \frac{1}{2} \left[ (f_{d,col}^s + f_{d,col}^{no}) P_{d,coal}^{no} + f_{d,col}^o P_{d,coal}^o \right] - f_{l,spectrum\_cut} \right\} \\
& \text{for } \alpha_l \geq 0. \tag{1.156}
\end{aligned}$$

For flow in a pipe the entrainment and deposition terms have to be taken into account. The link between kinetic generation of particles and the mass source terms is specified. The link between mass sources and the change in bubble/droplet size due to evaporation or condensation has been presented for local volume and time-averaged source terms. The concept of monodispersity has been discussed and a method proposed for computation of the disappearance of particles due to evaporation and condensation.

The following conclusions can be drawn from this derivation:

1. For numerical integration, the size of the discretization for the above local volume and time-average equations is allowed to be smaller than, equal to, or larger than the characteristic length scale  $D_l$  of the velocity field, unless other mathematical constraints associated with the numerical method used or concept used for the computation of the fragmentation and coalescence sources are imposed. The characteristic length scale  $D_l$  is understood to be:  $D_l = 2/|\kappa_l|$ , for a continuum, and  $D_l =$  particle size, for a disperse field.
2. The size of the spatial discretization for the resulting equations should not be an order of magnitude larger than the characteristic length scale of the fragmentation. This ensures that important local events will not be obscured as a result of averaging of the properties in the finite volume. For example, insufficiently fine spatial discretization on modeling of the melt–water interaction may lead to

nonprediction of steam explosions for situations where explosions have been observed in experiments, because averaging over large volume has regions with intensive fragmentation and others without such fragmentation.

3. The use of the concept of monodispersity is associated with the effect of disappearance of part of the particles during a given time step in the case of coincident mass transfer that reduces the local volume fraction of the phase.
4. Kinetic particle source terms are directly related to mass source terms in the mass conservation equation.
5. Nonkinetic particle source terms such as fragmentation or coalescence influence the average particle size and, with the size, the amount of evaporation or condensation. This means that the nonkinetic particle source terms indirectly influence the mass of the field. These terms introduce important *time delays* compared to the approach where the quasistatic particle size under local conditions is used. This is of major significance in the analysis of severe accidents such as steam explosions or processes associated with condensation shocks.

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# 2 Conservation of Momentum

*The time rate for change in momentum of a body equals the net force exerted on it.*

*Isaac Newton, Philosophiae Naturalis Principia, 1687*

## 2.1 Introduction

As in Chapter 1, from the large number of formulations of the conservation equations for multiphase flows, local volume averaging as founded by *Anderson* and *Jackson, Slattery, and Whitaker* was selected to derive rigorously the momentum equations for multiphase flows conditionally and divided into three velocity fields. The heterogeneous porous-media formulation introduced by *Gentry* et al., commented on by *Hirt*, and used by *Sha, Chao, and Soo*, is then implanted into the formalism as a geometrical skeleton. Beyond these concepts, I perform subsequent time averaging. This yields a working form that is applicable to a large variety of problems. All interfacial integrals are suitably transformed in order to enable practical application. Some minor simplifications are introduced in the finally obtained general equation and working equations for each of the three velocity fields are recommended for general use in multiphase fluid dynamic analysis.

This chapter is an improved and extended version of the work published in *Kolev* (1994b). The strategy followed is: We first apply the momentum equations for each of the velocity fields, excluding the interfaces by replacing their actions by forces. Then, we write a force balance at the interfaces, considering them as immaterial and therefore inertialess. This interfacial force balance links the momentum equations that are valid for the both sides of the interface.

## 2.2 Local volume-averaged momentum equations

### 2.2.1 Single-phase momentum equations

The time rate for change in momentum of a body relative to an inertial frame of reference equals the net force exerted on it (*Newton*).

Applied to a single continuum, this principle results in *Euler's* first law of continuum mechanics (*Truesdell*, 1968). When applied to each velocity field within the control volume, except the interface, this principle yields the well known local instantaneous momentum equation

$$\frac{\partial}{\partial \tau} (\rho_l \mathbf{V}_l^\tau) + \nabla \cdot (\rho_l \mathbf{V}_l^\tau \mathbf{V}_l^\tau - \mathbf{T}_l^\tau) + \nabla p_l^\tau + \rho_l \mathbf{g} = 0, \quad (2.1)$$

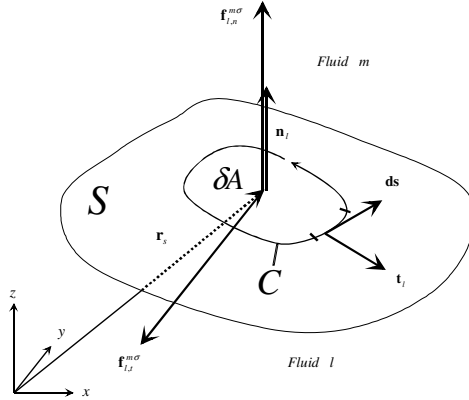
which is valid only inside the velocity field  $l$  excluding the interface. Here, the positive velocity direction gives the negative force direction – a commonly used definition. The total stress tensor is split into  $p_l^\tau \mathbf{I}$  and  $\mathbf{T}_l^\tau$ .  $p_l^\tau$  is the static pressure inside field  $l$ ,  $\mathbf{I}$  is the unit matrix, and  $\mathbf{T}_l^\tau$  is the shear stress tensor.  $\mathbf{V}_l^\tau \mathbf{V}_l^\tau$  is the dyadic product of two vectors  $\mathbf{V}_l^\tau$  and  $\mathbf{V}_l^\tau$  – see Appendix 1. It is a second-order tensor.  $\mathbf{g}$  is the vector of the gravitational acceleration. Equation (2.1) is the generally accepted balance of momentum for the single phase for velocities that are much smaller than the velocity of light.

### 2.2.2 Interface force balance (momentum jump condition)

Next, we abstract a volume around the interface having thickness  $\varepsilon$  converging to zero. Mechanical decoupling of the control volume from the adjacent fields requires replacing the action of the forces on the volume by equivalent forces. In this case, *Cauchy's* lemma holds:

The stress vectors acting upon opposite sides of the same surface at a given point are equal in magnitude and opposite in direction, *Truesdell* (1968, p.32).

*Characterization of the interface surface tension:* Consider two fluids with different densities, fluid  $m$ , and fluid  $l$  as presented in Fig. 2.1.



**Fig. 2.1** Surface force – geometry definitions

Fluid  $m$  is a gas and fluid  $l$  is a liquid,  $\rho_l > \rho_m$ . A three-dimensional surface  $S$ , described by the position vector  $\mathbf{r}_s(x, y, z)$ , separates both fluids. We call such fluids *immiscible*. The unit normal vector of the liquid interface  $\mathbf{n}_l$  points outside the liquid  $l$  and is an important local characteristic of this surface. It is defined from the spatial distribution of the volume fraction of the field  $l$  as follows:

$$\mathbf{n}_l = -\frac{\nabla(\alpha_l \gamma)}{|\nabla(\alpha_l \gamma)|}. \quad (2.2)$$

Due to different molecular attraction forces at the two sides of the surface a resulting attraction force with special properties arises at the surface. The force exists only at the surface and acts at the denser fluid  $l$ . This force is called *surface force*. The surface force per unit mixture volume is denoted by  $\mathbf{f}_l^{m\sigma}$ . The subscript  $l$  indicates that the force acts at the field  $l$ , and the subscript  $m\sigma$  indicates that the surface is an interface with field  $m$ . From the surface  $S$  we extract an infinitesimal part  $\delta A$  around the point  $\mathbf{r}_s(x, y, z)$  so that

$$\delta \mathbf{A} = \mathbf{n}_l \delta A. \quad (2.3)$$

The closed curve  $C$  contains  $\delta A$ . The closed curve  $C$  is oriented counterclockwise. Consider the infinitesimal directed line element  $d\mathbf{s}$  called the *arc length vector*. The unit tangent vector to the surface  $S$  at  $C$  that is perpendicular to  $d\mathbf{s}$  is  $\mathbf{t}$ . In this case, the following relation holds:

$$\mathbf{t} ds = d\mathbf{s} \times \mathbf{n}_l. \quad (2.4)$$

The surface force exerted on the surface  $\delta A$  by the surface outside of  $\delta A$  across the directed line element  $d\mathbf{s}$  is equal to  $\sigma_{lm} \mathbf{t} ds$ . Here  $\sigma_{lm}$  is a material property being a force tangential to  $S$  per unit length called *surface tension*. It may vary

with surface properties such as temperature, concentration of the impurities of the liquid, etc. The net surface force on the element  $\delta A$  is then obtained by summing all forces  $\sigma_{lm} \mathbf{t} ds$  exerted on each element of arc length  $ds$ ,  $\int_C \sigma_{lm} \mathbf{t} ds$ . Using

Eq. (2.4) gives

$$\int_C \sigma_{lm} \mathbf{t} ds = \int_C \sigma_{lm} \mathbf{ds} \times \mathbf{n}_l . \quad (2.5)$$

The *Stokes* theorem, see *Thomas et al. (1998)* for derivation, allows one to transfer the integral over a closed curve to an integral over the surface closed by this curve

$$\int_C \mathbf{ds} \times \mathbf{n}_l \sigma_{lm} = \iint_S (\mathbf{n}_l \times \nabla) \times \mathbf{n}_l \sigma_{lm} dA . \quad (2.6)$$

We see that for the *infinitesimal* surface  $\delta A$  the *surface force per unit interface* is

$$(\mathbf{n}_l \times \nabla) \times \mathbf{n}_l \sigma_{lm} = \sigma_{lm} [(\mathbf{n}_l \times \nabla) \times \mathbf{n}_l] + (\mathbf{n}_l \times \nabla \sigma_{lm}) \times \mathbf{n}_l . \quad (2.7)$$

This force can be split into a normal and a tangential component by splitting the gradient into a sum of normal and tangential components

$$\nabla = \nabla_n + \nabla_t , \quad (2.8)$$

where

$$\nabla_n = (\nabla \cdot \mathbf{n}_l) \mathbf{n}_l \quad (2.9)$$

and

$$\nabla_t = \nabla - \nabla_n . \quad (2.10)$$

Using this splitting and after some mathematical manipulation, *Brackbill et al. (1992)* simplified Eq. (2.7) and finally obtained the very important result,

$$(\mathbf{n}_l \times \nabla) \times \mathbf{n}_l \sigma = -\mathbf{n}_l \sigma_{lm} (\nabla \cdot \mathbf{n}_l) + \nabla_t \sigma_{lm} = \mathbf{n}_l \sigma_{lm} \kappa_l + \nabla_t \sigma_{lm} , \quad (2.11)$$

where

$$\kappa_l = -(\nabla \cdot \mathbf{n}_l) = \nabla \cdot \left[ \frac{\nabla(\alpha_l \gamma)}{|\nabla(\alpha_l \gamma)|} \right] \quad (2.12)$$

is the curvature of the interface defined only by the gradient of the interface unit vector. Equation (2.12) is used in Eq. (5.4) in *Drazin and Reid (1981, p. 23)*. Note that the mathematical definition of curvature is the sum of the two principal curvatures which are magnitudes of two vectors. This sum is always positive. The expression resulting from Eq. (2.12) defines curvature with sign, which means with its orientation. The curvature is positive if the center of the curvature is in the fluid *m*. In other words, the positive curvature  $\kappa_l$  is oriented along the normal vector  $\mathbf{n}_l$ .

As an example, let us estimate the curvature of a liquid layer in stratified flow between two horizontal planes with a gap equal to  $H$ . The interface is described by the curve  $z^* = \frac{z}{H} = \alpha_2(x)$  being in the plane  $y = \text{const}$ . As coordinates we use  $(x, y, z^*)$ . The gradient of the liquid volume fraction is then

$$\nabla \alpha_2(x) = \frac{\partial \alpha_2(x)}{\partial x} \mathbf{i} + \frac{\partial \alpha_2(x)}{\partial z^*} \mathbf{k} = \frac{\partial \alpha_2(x)}{\partial x} \mathbf{i} + \frac{\partial z^*}{\partial z^*} \mathbf{k} = \frac{\partial \alpha_2(x)}{\partial x} \mathbf{i} + \mathbf{k}.$$

The magnitude of the gradient is then

$$|\nabla \alpha_2(x)| = \left\{ 1 + \left[ \frac{\partial \alpha_2(x)}{\partial x} \right]^2 \right\}^{1/2}.$$

The normal vector is

$$\mathbf{n}_2 = -\frac{\nabla \alpha_2(x)}{|\nabla \alpha_2(x)|} = -\left( \frac{\frac{\partial \alpha_2(x)}{\partial x}}{\left\{ 1 + \left[ \frac{\partial \alpha_2(x)}{\partial x} \right]^2 \right\}^{1/2}} \mathbf{i} + \frac{1}{\left\{ 1 + \left[ \frac{\partial \alpha_2(x)}{\partial x} \right]^2 \right\}^{1/2}} \mathbf{k} \right).$$

The curvature in accordance with Eq. (2.12) is then

$$\begin{aligned} \kappa_l &= -(\nabla \cdot \mathbf{n}_l) = \nabla \cdot \left[ \frac{\nabla(\alpha_l \gamma)}{|\nabla(\alpha_l \gamma)|} \right] \\ &= \nabla \cdot \left( \frac{\frac{\partial \alpha_2(x)}{\partial x}}{\left\{ 1 + \left[ \frac{\partial \alpha_2(x)}{\partial x} \right]^2 \right\}^{1/2}} \mathbf{i} + \frac{1}{\left\{ 1 + \left[ \frac{\partial \alpha_2(x)}{\partial x} \right]^2 \right\}^{1/2}} \mathbf{k} \right) \\ &= \frac{\partial}{\partial x} \frac{\frac{\partial \alpha_2(x)}{\partial x}}{\left\{ 1 + \left[ \frac{\partial \alpha_2(x)}{\partial x} \right]^2 \right\}^{1/2}} + \frac{\partial}{\partial z^*} \frac{1}{\left\{ 1 + \left[ \frac{\partial \alpha_2(x)}{\partial x} \right]^2 \right\}^{1/2}}. \end{aligned}$$

Bearing in mind that

$$\frac{\partial}{\partial z^*} \frac{1}{\left\{1 + \left[\frac{\partial \alpha_2(x)}{\partial x}\right]^2\right\}^{1/2}} = - \frac{\frac{\partial \alpha_2(x)}{\partial x} \frac{\partial^2 \alpha_2(x)}{\partial x \partial z^*}}{\left\{1 + \left[\frac{\partial \alpha_2(x)}{\partial x}\right]^2\right\}^{3/2}} = 0,$$

one finally obtains the well known expression

$$\kappa_2 = \frac{\partial^2 \alpha_2(x)}{\partial x^2} \bigg/ \left\{1 + \left[\frac{\partial \alpha_2(x)}{\partial x}\right]^2\right\}^{3/2}.$$

The higher pressure is in the fluid medium on the concave side of the interface, since surface force in case of constant surface tension is a net normal force directed toward the center of curvature of the interface.

Note that the term  $\sigma_{lm} \kappa_l$  becomes important for

$$1/|\kappa_l| < 0.001m. \quad (2.13)$$

The term  $\nabla_t \sigma_{lm}$  can be expressed as

$$\nabla_t \sigma_{lm} = (\nabla_t T_l) \frac{\partial \sigma_{lm}}{\partial T_l} + \sum_{i=1}^{i_{\max}} (\nabla_t C_{il}) \frac{\partial \sigma_{lm}}{\partial C_{il}}. \quad (2.14)$$

These terms describe the well known *Marangoni* effect.

The grid density for computational analysis can be judged by comparing the size of the control volume  $\Delta x$  with the curvature. Obviously, if and only if the volume is small enough,

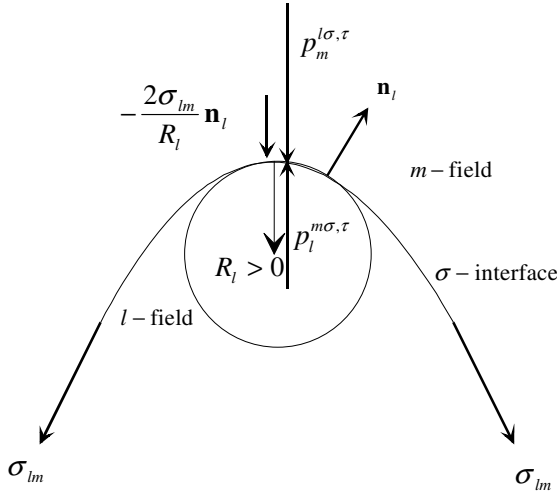
$$\Delta x |\kappa_l| < 1, \quad (2.15)$$

the curvature can be resolved by the computational analysis.

*No velocity variation across the interface:* Consider the case for which there is no velocity variation across the interface (e.g., stagnant fluids) – Fig. 2.2.

The interface pressure  $p_l^{m\sigma,\tau}$  is the normal force per unit surface acting on field  $l$ . This force acts *inside* the field  $l$  in the immediate vicinity of the interface and in the opposite direction on the interface control volume. It is *different* from the surface force exerted by the pressure of the neighboring velocity field,  $p_m^{l\sigma,\tau}$ . If there are no other forces except pressure and surface tension, we have

$$p_l^{m\sigma,\tau} \mathbf{n}_l + p_m^{l\sigma,\tau} \mathbf{n}_m + \nabla_t \sigma_{lm} + \sigma_{lm} \kappa_l \mathbf{n}_l = 0. \quad (2.16)$$



**Fig. 2.2** Interface force equilibrium without mass transfer and Marangoni effect (Laplace)

This equation is known as the *Laplace* equation. If, in addition, we have viscous forces acting on the two sides the momentum balance is

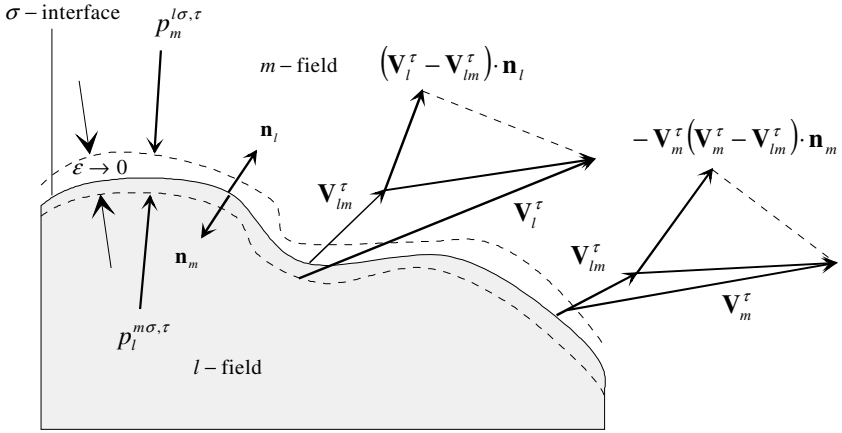
$$\left(\mathbf{T}_l^{m\sigma,\tau} - p_l^{m\sigma,\tau}\mathbf{I}\right) \cdot \mathbf{n}_l + \left(\mathbf{T}_m^{l\sigma,\tau} - p_m^{l\sigma,\tau}\mathbf{I}\right) \cdot \mathbf{n}_m - \nabla_l \sigma_{lm} - \sigma_{lm} \kappa_l \mathbf{n}_l = 0. \quad (2.17)$$

*Velocity variation across the interface:* If mass is transferred from one to the other fluid for whatever reason, the interface moves in space not only convectively but is also controlled by the amount of mass transferred between the fields – Fig. 2.3. In this case, the interface velocity  $\mathbf{V}_{lm}^\tau$  is not equal to the neighboring field velocities. The mass flow rate

$$(\rho w)_{lm} = \rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_l$$

enters the interface control volume and exerts the force  $\rho_l \mathbf{V}_l^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_l$  per unit surface on it. Note that this force has the same direction as the pressure force inside the field  $l$ . Similarly, we have a reactive force  $\rho_m \mathbf{V}_m^\tau (\mathbf{V}_m^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_m$  exerted per unit surface on the control volume by the leaving mass flow rate. Assuming that the control volume moves with the normal component of the interface velocity,  $\mathbf{V}_{lm}^\tau \cdot \mathbf{n}_l$ , we obtain the following force balance:

$$\begin{aligned} & -\rho_l \mathbf{V}_l^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_l + \left(\mathbf{T}_l^{m\sigma,\tau} - p_l^{m\sigma,\tau}\mathbf{I}\right) \cdot \mathbf{n}_l \\ & -\rho_m \mathbf{V}_m^\tau (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) \cdot \mathbf{n}_m + \left(-p_{m\sigma}^\tau \mathbf{I} + \mathbf{T}_{m\sigma}^\tau\right) \cdot \mathbf{n}_m - \nabla_l \sigma_{lm} - \sigma_{lm} \kappa_l \mathbf{n}_l = 0. \end{aligned} \quad (2.18)$$



**Fig. 2.3** Definition of the interface characteristics

This is the general form of the interfacial momentum jump condition. It is convenient to rewrite the above equation by using the mass jump condition at the interface

$$\rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_l + \rho_m (\mathbf{V}_m^\tau - \mathbf{V}_{lm}^\tau) \cdot \mathbf{n}_m = 0, \quad (2.19)$$

which is Eq. (1.42). The result is

$$\left[ \rho_l (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) (\mathbf{V}_m^\tau - \mathbf{V}_l^\tau) + (p_m^{l\sigma,\tau} - p_l^{m\sigma,\tau}) \mathbf{I} + \mathbf{T}_l^{m\sigma,\tau} - \mathbf{T}_m^{l\sigma,\tau} - \sigma_{lm} \kappa_l \right] \cdot \mathbf{n}_l - \nabla_l \sigma_{lm} = 0 \quad (2.20)$$

or

$$(\rho w)_{lm} (\mathbf{V}_m^\tau - \mathbf{V}_l^\tau) + (p_m^{l\sigma,\tau} - p_l^{m\sigma,\tau}) \mathbf{n}_l + \left[ \mathbf{T}_l^{m\sigma,\tau} - \mathbf{T}_m^{l\sigma,\tau} - \sigma_{lm} \kappa_l \right] \cdot \mathbf{n}_l - \nabla_l \sigma_{lm} = 0. \quad (2.21)$$

The projection of this force to the normal direction is obtained by scalar multiplication of the above equation with the unit vector  $\mathbf{n}_l$ . The result is

$$\begin{aligned} & (\rho w)_{lm} (\mathbf{V}_m^\tau - \mathbf{V}_l^\tau) \cdot \mathbf{n}_l + p_m^{l\sigma,\tau} - p_l^{m\sigma,\tau} - \sigma_{lm} \kappa_l + \left[ (\mathbf{T}_l^{m\sigma,\tau} - \mathbf{T}_m^{l\sigma,\tau}) \cdot \mathbf{n}_l \right] \cdot \mathbf{n}_l \\ & - (\nabla_l \sigma_{lm}) \cdot \mathbf{n}_l = 0. \end{aligned} \quad (2.22)$$

Using the mass conservation at the interface we finally have an important force balance normal to the interface

$$\begin{aligned} & (\rho w)_{lm}^2 \left( \frac{1}{\rho_m} - \frac{1}{\rho_l} \right) + p_m^{l\sigma,\tau} - p_l^{m\sigma,\tau} - \sigma_{lm} \kappa_l + \left[ (\mathbf{T}_l^{m\sigma,\tau} - \mathbf{T}_m^{l\sigma,\tau}) \cdot \mathbf{n}_l \right] \cdot \mathbf{n}_l \\ & - (\nabla_l \sigma_{lm}) \cdot \mathbf{n}_l = 0. \end{aligned} \quad (2.23)$$



Neglecting all forces except those caused by pressure and interfacial mass transfer results in the surprising conclusion that during the mass transfer the pressure in the denser fluid is always larger than the pressure in the lighter fluid independently of the direction of the mass transfer – *Delhay* (1981, p. 52, Eq. (2.64)).

For the limiting case of no interfacial mass transfer and dominance of the pressure difference, the velocity of the interface can be expressed as a function of the pressure difference and the velocities in the bulk of the fields,

$$V_{lm}^{n,\tau} = V_l^{n,\tau} - \frac{p_m^{l\sigma,\tau} - p_l^{m\sigma,\tau}}{\rho_l (V_m^{n,\tau} - V_l^{n,\tau})}. \quad (2.24)$$

This velocity is called contact discontinuity velocity. Replacing the discontinuity velocity with Eq. (1.42), we obtain

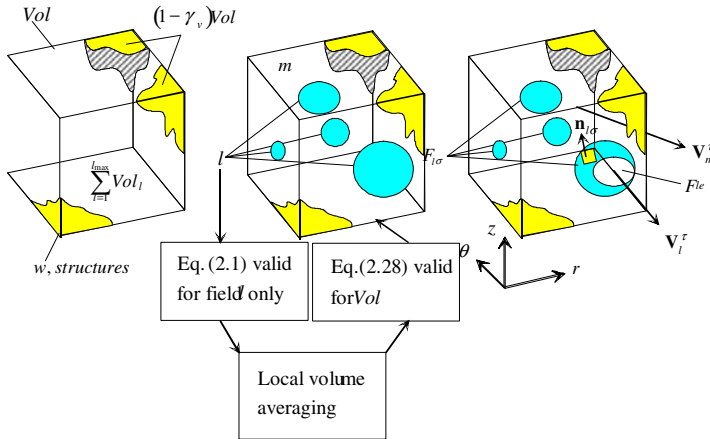
$$(V_l^{n,\tau} - V_m^{n,\tau})^2 = \frac{\rho_l - \rho_m}{\rho_l \rho_m} (p_l^{m\sigma,\tau} - p_m^{l\sigma,\tau}). \quad (2.25)$$

For the case  $\rho_l \gg \rho_m$ , we have the expected result that the pressure difference equals the stagnation pressure at the side of the lighter medium

$$p_l^{m\sigma,\tau} - p_m^{l\sigma,\tau} = \rho_m (V_l^{n,\tau} - V_m^{n,\tau})^2. \quad (2.26)$$

### 2.2.3 Local volume averaging of the single-phase momentum equation

The aim here is to average Eq. (2.1) over the total control volume – see Fig. 2.4.



**Fig. 2.4** Definition regions for single-phase instantaneous momentum balance and the local volume average balance

The mathematical tools used to derive local volume-averaged field conservation equations for the property being any scalar, vectorial, or tensorial function of time and location are once again *Slattery–Whitaker’s* spatial averaging theorem, together with the *Gauss–Ostrogradskii* theorem and the general transport equation (*Leibnitz* rule), see *Anderson and Jackson* (1967), *Slattery* (1967), *Whitaker* (1967, 1985, 1969), and *Gray and Lee* (1977). Applying the local volume average to Eq. (2.1), the following is obtained:

$$\left\langle \frac{\partial}{\partial \tau} (\rho_l \mathbf{V}_l^\tau) \right\rangle + \left\langle \nabla \cdot (\rho_l \mathbf{V}_l^\tau \mathbf{V}_l^\tau) \right\rangle - \left\langle \nabla \cdot (\mathbf{T}_l^\tau) \right\rangle + \left\langle \nabla p_l^\tau \right\rangle + \left\langle \rho_l \mathbf{g} \right\rangle = 0, \quad (2.27)$$

or using Eqs. (1.28), (1.32), and (1.28), (*Kolev*, 1994),

$$\begin{aligned} & \frac{\partial}{\partial \tau} \left\langle \rho_l \mathbf{V}_l^\tau \right\rangle + \nabla \cdot \left\langle \rho_l \mathbf{V}_l^\tau \mathbf{V}_l^\tau \right\rangle + \nabla \cdot \left\langle p_l^\tau \mathbf{I} - \mathbf{T}_l^\tau \right\rangle + \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \rho_l \mathbf{V}_l^\tau (\mathbf{V}_l^\tau - \mathbf{V}_l^{\sigma,\tau}) \cdot \mathbf{n}_i dF \\ & - \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} (-p_l^\tau \mathbf{I} + \mathbf{T}_l^\tau) \cdot \mathbf{n}_i dF + \mathbf{g} \rho_l = 0. \end{aligned} \quad (2.28)$$

The volume average momentum equation can be rewritten using the *weighted average*, see Eq. (1.19), *Kolev* (1994),

$$\left\langle \mathbf{V}_l^\tau \right\rangle^{le,\rho} = \left\langle \rho_l \mathbf{V}_l^\tau \right\rangle^{le} / \left\langle \rho_l \right\rangle^l, \quad (2.29)$$

and

$$\left\langle \mathbf{V}_l^\tau \right\rangle^{le} \left\langle \mathbf{V}_l^\tau \right\rangle^{le} = \left\langle \rho_l \mathbf{V}_l^\tau \mathbf{V}_l^\tau \right\rangle^{le} / \left\langle \rho_l \right\rangle^l. \quad (2.30)$$

The result is

$$\begin{aligned} & \frac{\partial}{\partial \tau} \left( \alpha_l \gamma_v \left\langle \rho_l \right\rangle^l \left\langle \mathbf{V}_l^\tau \right\rangle^{le} \right) + \nabla \cdot \left( \alpha_l^e \gamma \left\langle \rho_l \right\rangle^l \left\langle \mathbf{V}_l^\tau \right\rangle^{le} \left\langle \mathbf{V}_l^\tau \right\rangle^{le} \right) - \nabla \cdot \left( \alpha_l^e \gamma \left\langle \mathbf{T}_l^\tau \right\rangle^{le} \right) \\ & + \nabla \left( \alpha_l^e \gamma \left\langle p_l^\tau \right\rangle^{le} \right) + \alpha_l \gamma_v \left\langle \rho_l \right\rangle^l \mathbf{g} - \frac{1}{Vol} \int_{F_{l\sigma}} \left[ -p_l^{m\sigma,\tau} \mathbf{I} + \mathbf{T}_l^{m\sigma,\tau} - \rho_l \mathbf{V}_l^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) \right] \cdot \mathbf{n}_i dF \\ & - \frac{1}{Vol} \int_{F_{lw}} \left[ -p_l^{w\sigma,\tau} \mathbf{I} + \mathbf{T}_l^{w\sigma,\tau} - \rho_l \mathbf{V}_l^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lw}^\tau) \right] \cdot \mathbf{n}_i dF = 0. \end{aligned} \quad (2.31)$$

We assume that the weighted average of products can be replaced by the products of the average. This should be borne in mind when constructing a numerical algorithm for solving the final system and selecting the size of the finite volume so as to be not so large as to violate the validity of this assumption.

Note the differences between Eq. (2.31) and the final result obtained by *Ishii* (1975), Eq. (3.16):

- (a) the directional permeability is used here instead of the volumetric porosity in the pressure gradient term, and

- (b) as for the mass conservation equation in *Kolev* (1994a), the volumetric porosity is kept below the time differential since it can be a function of time in a number of interesting applications.

For the case  $\gamma_v = \gamma = 1$  and one-dimensional flow, Eq. (2.31) reduces to Eq. (15) derived by *Delhay* in *Hetstrony* (1982, p. 163).

Equation (2.31), the rigorously derived local volume average momentum equation, is not amenable to direct use in computational models without further transformation. In order to facilitate its practical use

- (a) the integral expression must be evaluated, and  
 (b) the time averaging must be performed subsequently.

These steps, which go beyond *Sha* et al. (1984), are performed in Sects. 2.3 and 2.4.

## 2.3 Rearrangement of the surface integrals

The expression under the surface integral is replaced by its equivalent from the momentum jump condition, Eq. (2.18):

$$\begin{aligned} & -\frac{1}{Vol} \int_{F_{l\sigma}} \left[ -\rho_l \mathbf{V}_l^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) + \mathbf{T}_l^{m\sigma,\tau} - p_l^{m\sigma,\tau} \mathbf{I} \right] \cdot \mathbf{n}_l dF \\ & = -\frac{1}{Vol} \int_{F_{l\sigma}} \left\{ \left[ -\rho_m \mathbf{V}_m^\tau (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) + \mathbf{T}_m^{l\sigma,\tau} - p_m^{l\sigma,\tau} \mathbf{I} + \sigma_{lm} \kappa_l \right] \cdot \mathbf{n}_l + \nabla_l \sigma_{lm} \right\} dF. \end{aligned} \quad (2.32)$$

Note that there are no surface force terms in Eq. (2.31). Equation (2.32) reflects the action of the surface forces and the action of the stresses caused by the surrounding field  $m$  on  $l$ . Note also that if the momentum equations are written for two neighboring fields the interface forces will appear in both equations with opposite sign only if the equations are applied to a common dividing surface. Otherwise, the exchange terms in question will be nonsymmetric.

The *intrinsic surface-averaged* field pressure at the entrances and exits of the control volume crossing the field  $m$  is  $\langle p_m^\tau \rangle^{me}$ . We call it bulk pressure inside the velocity field  $m$  at this particular surface. The interfacial  $m$ -side pressure  $p_m^{l\sigma,\tau}$  can be expressed as the sum of the intrinsic averaged pressure,  $\langle p_m^\tau \rangle^{me}$ , which is not a function of the position at the interface inside the control volume and can be taken outside of the integral sign, and a pressure difference  $\Delta p_m^{l\sigma,\tau}$ , which is a function of the position at the interface in the control volume

$$p_m^{l\sigma,\tau} = \langle p_m^\tau \rangle^{me} + \Delta p_m^{l\sigma,\tau}. \quad (2.33)$$

The same is performed for all other fields. Similarly, the surface pressure of the field structure interface is expressed as

$$p_i^{w\sigma,\tau} = \langle p_i^\tau \rangle^{le} + \Delta p_i^{w\sigma,\tau}. \quad (2.34)$$

The interfacial pressure differs from the bulk pressures of the corresponding fields  $m\sigma$ , and  $lw$ ,  $\Delta p_m^{l\sigma,\tau} \neq 0$  and  $\Delta p_l^{w\sigma,\tau} \neq 0$ . This occurs because obstacles to the continuous phase can cause local velocity decreases or increases at the interface velocity boundary layer, resulting in increased or decreased pressure relative to bulk pressure. In order to estimate the surface integrals the exact dependence of the pressure as a function of the position at the interface inside the control volume must be elaborated for each idealized flow pattern and form of the structure. The same is valid for the viscous shear stresses  $\mathbf{T}_m^{l\sigma,\tau}$  and  $\mathbf{T}_l^{w\sigma,\tau}$ . Note that in order to estimate the interfacial pressure integrals for practical use of the momentum equations the time averaging must be performed first to make it admissible to use the real pressure distributions measured experimentally on bodies in turbulent flow.

Substituting Eqs. (2.32)–(2.34) into Eq. (2.31), we obtain

$$\begin{aligned} & \frac{\partial}{\partial \tau} \left( \alpha_i \gamma_v \langle \rho_i \rangle^l \langle \mathbf{V}_i^\tau \rangle^{le} \right) + \nabla \cdot \left( \alpha_i^e \gamma \langle \rho_i \rangle^l \langle \mathbf{V}_i^\tau \rangle^{le} \langle \mathbf{V}_i^\tau \rangle^{le} \right) - \nabla \cdot \left( \alpha_i^e \gamma \langle \mathbf{T}_i^\tau \rangle^{le} \right) \\ & + \nabla \left( \alpha_i^e \gamma \langle p_i^\tau \rangle^{le} \right) + \alpha_i \gamma_v \langle \rho_i \rangle^l \mathbf{g} + \frac{1}{Vol} \int_{F_{l\sigma}} \langle p_m^\tau \rangle^{me} \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{lw}} \langle p_l^\tau \rangle^{le} \mathbf{n}_l dF \\ & - \frac{1}{Vol} \int_{F_{l\sigma}} \left\{ \left[ -\Delta p_m^{l\sigma,\tau} \mathbf{I} + \mathbf{T}_m^{l\sigma,\tau} - \rho_m \mathbf{V}_m^\tau (\mathbf{V}_m^\tau - \mathbf{V}_{lm}^\tau) + \sigma_{lm} \boldsymbol{\kappa}_l \right] \cdot \mathbf{n}_l + \nabla_t \sigma_{lm} \right\} dF \\ & - \frac{1}{Vol} \int_{F_{lw}} \left[ -\Delta p_l^{w\sigma,\tau} \mathbf{I} + \mathbf{T}_l^{w\sigma,\tau} - \rho_l \mathbf{V}_l^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lw}^\tau) \right] \cdot \mathbf{n}_l dF = 0. \end{aligned} \quad (2.35)$$

Keeping in mind that  $\langle p_m^\tau \rangle^{me}$  and  $\langle p_l^\tau \rangle^{le}$  are not functions of the interface position inside the control volume, and making use of Eq. (1.32) (see also *Kolev (1994)*), the bulk pressure integrals can then be rewritten as follows:

$$\begin{aligned} & \langle p_m^\tau \rangle^{me} \frac{1}{Vol} \int_{F_{l\sigma}} \mathbf{n}_l dF + \langle p_l^\tau \rangle^{le} \frac{1}{Vol} \int_{F_{lw}} \mathbf{n}_l dF \\ & = -\langle p_m^\tau \rangle^{me} \nabla \left( \alpha_i^e \gamma \right) + \left( \langle p_l^\tau \rangle^{le} - \langle p_m^\tau \rangle^{me} \right) \frac{1}{Vol} \int_{F_{lw}} \mathbf{n}_l dF. \end{aligned} \quad (2.36)$$

*Rearranging the surface tension integrals:* The surface force per unit volume of the mixture is, in fact, the local volume average of the surface tension force

$$\mathbf{f}_l^{m\sigma} = -\frac{1}{Vol} \iint_{F_{l\sigma}} (\mathbf{n}_l \sigma_{lm} \boldsymbol{\kappa} + \nabla_t \sigma_{lm}) dF. \quad (2.37)$$

Note that the orientation of this force is defined with respect the coordinate system given in Fig. 2.1. Using Eq. (1.29) we have

$$\begin{aligned}
\mathbf{f}_i^{m\sigma} &= -\sigma_{lm}\kappa_l \frac{1}{Vol} \iint_{F_{i\sigma}} \mathbf{n}_l \sigma_{lm} \kappa_l dF - \frac{1}{Vol} \iint_{F_{i\sigma}} (\nabla_i \sigma_{lm}) dF \\
&= \sigma_{lm}\kappa_l \left[ \nabla(\alpha_i^e \gamma) + \frac{1}{Vol} \int_{F_{iw}} \mathbf{n}_l dF \right] - \frac{1}{Vol} \iint_{F_{i\sigma}} (\nabla_i \sigma_{lm}) dF, \tag{2.38}
\end{aligned}$$

with normal

$$\begin{aligned}
\mathbf{f}_{i,n}^{m\sigma} &= \sigma_{lm}\kappa_l \left[ \nabla(\alpha_i^e \gamma) + \frac{1}{Vol} \int_{F_{iw}} \mathbf{n}_l dF \right] = \sigma_{lm} \nabla \cdot \left[ \frac{\nabla(\alpha_i \gamma)}{|\nabla(\alpha_i \gamma)|} \right] \left[ \nabla(\alpha_i^e \gamma) + \frac{1}{Vol} \int_{F_{iw}} \mathbf{n}_l dF \right] \\
\text{for } \alpha_i > 0 & \tag{2.39}
\end{aligned}$$

and tangential

$$\mathbf{f}_{i,t}^{m\sigma} = -\frac{1}{Vol} \iint_{F_{i\sigma}} (\nabla_i \sigma_{lm}) dF \approx -a_{lm} \nabla_i \sigma_{lm} \tag{2.40}$$

surface force components per unit volume of the mixture, respectively. Here  $a_{lm}$  is the interfacial area density. In the literature, the local volume averaged surface force is sometimes called the *continuum surface force* or abbreviated as CSF, see *Brackbill et al. (1992)*.

Note that if the surface tension is a constant in space there is no resulting tangential force component. At plane surfaces the curvature is zero and therefore there is no normal force acting at the liquid. If the liquid consists of clouds of spheres the local surface force creates only a difference in pressures inside and outside the sphere, but there is no net force influencing the total movement either of a single droplet or of the cloud of the droplets in the space due to this force. We express this fact by multiplying the surface force by the function  $\delta_l$  being 1 for the continuum and 0 for the disperse field.

Using Eqs. (2.36) and (2.38) the pressure and the surface tension terms can be rearranged as follows:

$$\begin{aligned}
&\nabla \cdot \left( \alpha_i^e \gamma \langle p_l^\tau \rangle^{le} \right) + \frac{1}{Vol} \int_{F_{i\sigma}} \langle p_m^\tau \rangle^{me} \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{iw}} \langle p_l^\tau \rangle^{le} \mathbf{n}_l dF \\
&- \frac{1}{Vol} \iint_{F_{i\sigma}} (\mathbf{n}_l \sigma_{lm} \kappa + \nabla_i \sigma_{lm}) dF \\
&= \nabla \cdot \left( \alpha_i^e \gamma \langle p_l^\tau \rangle^{le} \right) - \langle p_m^\tau \rangle^{me} \nabla(\alpha_i^e \gamma) + \left( \langle p_l^\tau \rangle^{le} - \langle p_m^\tau \rangle^{me} \right) \frac{1}{Vol} \int_{F_{iw}} \mathbf{n}_l dF \\
&+ \delta_l \sigma_{lm} \kappa_l \left[ \nabla(\alpha_i^e \gamma) + \frac{1}{Vol} \int_{F_{iw}} \mathbf{n}_l dF \right] - \delta_l \frac{1}{Vol} \iint_{F_{i\sigma}} (\nabla_i \sigma_{lm}) dF
\end{aligned}$$

$$\begin{aligned}
&= \alpha_i^e \gamma \mathcal{N} \langle p_i^\tau \rangle^{le} + \left( \langle p_i^\tau \rangle^{le} - \langle p_m^\tau \rangle^{me} + \delta_i \sigma_{lm} \kappa_l \right) \left( \nabla (\alpha_i^e \gamma) + \frac{1}{Vol} \int_{F_w} \mathbf{n}_l dF \right) \\
&- \delta_i \frac{1}{Vol} \iint_{F_\sigma} (\nabla_i \sigma_{lm}) dF. \tag{2.41}
\end{aligned}$$

Note that in Eq. (2.41)

$$\alpha_i^e \gamma \mathcal{N} \cdot \langle p_i^\tau \rangle^{le}$$

stands for

$$\nabla \cdot \left( \alpha_i^e \gamma \langle p_i^\tau \rangle^{le} \right) - \langle p_i^\tau \rangle^{le} \nabla \cdot (\alpha_i^e \gamma).$$

*Rearranging the integrals defining interfacial momentum transfer due to mass transfer:* Again using the mass jump condition at the interface, which is Eq. (1.42),

$$\rho_m \mathbf{V}_m^\tau (\mathbf{V}_m^\tau - \mathbf{V}_{lm}^\tau) = \rho_l \mathbf{V}_m^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau), \tag{2.42}$$

the surface integral is rearranged as follows:

$$-\frac{1}{Vol} \int_{F_\sigma} \left[ -\rho_m \mathbf{V}_m^\tau (\mathbf{V}_m^\tau - \mathbf{V}_{lm}^\tau) \right] \cdot \mathbf{n}_l dF = \frac{1}{Vol} \int_{F_\sigma} \left[ \rho_l \mathbf{V}_m^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lm}^\tau) \right] \cdot \mathbf{n}_l dF. \tag{2.43}$$

For practical applications, the mass source term is split into nonnegative components, as has already been explained in Chap. 1 (see also *Kolev (1994a)*). In addition, it is assumed that the mass emitted from a field has the velocity of the donor field. As a result, the local volume-averaged interfacial forces related to mass transfer through the interfaces can be rewritten as follows:

(a) The components related to mass injection into, or suction from, the field through the field structure interface are replaced by the “donor” hypothesis

$$-\frac{1}{Vol} \int_{F_w} \left[ \rho_l \mathbf{V}_m^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{lw}^\tau) \right] \cdot \mathbf{n}_l dF = \gamma_v \left( \mu_{wl}^\tau \langle \mathbf{V}_w^\tau \rangle^{we} - \mu_{lw}^\tau \langle \mathbf{V}_l^\tau \rangle^{le} \right). \tag{2.44}$$

(b) The components due to evaporation, condensation, entrainment, and deposition are replaced by the “donor” hypothesis

$$-\frac{1}{Vol} \int_{F_\sigma} \left[ \rho_l \mathbf{V}_i^\tau (\mathbf{V}_l^\tau - \mathbf{V}_{i\sigma}^\tau) \right] \cdot \mathbf{n}_l dF = \gamma_v \sum_{m=1}^3 \left( \mu_{ml}^\tau \langle \mathbf{V}_m^\tau \rangle^{me} - \mu_{lm}^\tau \langle \mathbf{V}_l^\tau \rangle^{le} \right). \tag{2.45}$$

The sum of all interface mass transfer components is then

$$\gamma_v \sum_{m=1}^{3,w} \left( \mu_{ml}^\tau \langle \mathbf{V}_m^\tau \rangle^{me} - \mu_{lm}^\tau \langle \mathbf{V}_l^\tau \rangle^{le} \right). \tag{2.46}$$

Thus, the final form obtained for the local volume average momentum equation is as follows:

$$\begin{aligned}
& \frac{\partial}{\partial \tau} \left( \alpha_l \gamma_v \langle \rho_l \rangle^l \langle \mathbf{V}_l^\tau \rangle^{le} \right) + \nabla \cdot \left( \alpha_l^e \gamma \langle \rho_l \rangle^l \langle \mathbf{V}_l^\tau \rangle^{le} \langle \mathbf{V}_l^\tau \rangle^{le} \right) - \nabla \cdot \left( \alpha_l^e \gamma \langle \mathbf{T}_l^\tau \rangle^{le} \right) + \alpha_l \gamma_v \langle \rho_l \rangle^l \mathbf{g} \\
& + \alpha_l^e \gamma \bar{\nabla} \langle p_l^\tau \rangle^{le} + \left( \langle p_l^\tau \rangle^{le} - \langle p_m^\tau \rangle^{me} + \delta_l \sigma_{lm} \kappa_l \right) \left( \nabla (\alpha_l^e \gamma) + \frac{1}{Vol} \int_{F_{hw}} \mathbf{n}_l dF \right) \\
& - \delta_l \frac{1}{Vol} \int_{F_{\sigma}} (\nabla_l \sigma_{lm}) dF + \frac{1}{Vol} \int_{F_{\sigma}} (\Delta p_m^{l\sigma, \tau} \mathbf{I} - \mathbf{T}_m^{l\sigma, \tau}) \cdot \mathbf{n}_l dF \\
& + \frac{1}{Vol} \int_{F_{hw}} (\Delta p_l^{w\sigma, \tau} \mathbf{I} - \mathbf{T}_l^{w\sigma, \tau}) \cdot \mathbf{n}_l dF = \gamma_v \sum_{m=1}^{3,w} \left( \mu_{ml}^\tau \langle \mathbf{V}_m^\tau \rangle^{me} - \mu_{lm}^\tau \langle \mathbf{V}_l^\tau \rangle^{le} \right), \quad (2.47)
\end{aligned}$$

and is independent of whether the field is structured or nonstructured. Before we continue with the estimation of the remaining integrals, we will first perform a time averaging of Eq. (2.47).

## 2.4 Local volume average and time average

The instantaneous surface-averaged velocity of the field  $l$ ,  $\langle \mathbf{V}_l^\tau \rangle^{le}$ , can be expressed as the sum of the surface-averaged velocity, which is subsequently time-averaged,

$$V_l = \overline{\langle \mathbf{V}_l^\tau \rangle^{le}} \quad (2.48)$$

and a pulsation component  $V_l'$ ,

$$\langle \mathbf{V}_l^\tau \rangle^{le} = V_l + V_l', \quad (2.49)$$

as proposed by *Reynolds*. The fluctuation of the velocity is the predominant fluctuation component relative to, say, the fluctuation of  $\alpha_l$  or  $\rho_l$ . Introduction of Eq. (2.49) into the momentum conservation equation and time averaging yields

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (\alpha_l \rho_l \mathbf{V}_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l \mathbf{V}_l \mathbf{V}_l \gamma) + \nabla \cdot \left[ \alpha_l^e \gamma (\rho_l \overline{\mathbf{V}_l \mathbf{V}_l'} - \delta_l \mathbf{T}_l) \right] + \alpha_l^e \gamma \bar{\nabla} p_l + \alpha_l \gamma_v \rho_l \mathbf{g} \\
& + (p_l - p_m + \delta_l \sigma_{lm} \kappa_l) \left( \nabla (\alpha_l^e \gamma) + \frac{1}{Vol} \int_{F_{hw}} \mathbf{n}_l dF \right) - \delta_l \frac{1}{Vol} \int_{F_{\sigma}} (\nabla_l \sigma_{lm}) dF \\
& + \frac{1}{Vol} \int_{F_{\sigma}} (\Delta p_m^{l\sigma} \mathbf{I} - \mathbf{T}_m^{l\sigma}) \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{hw}} (\Delta p_l^{w\sigma} \mathbf{I} - \mathbf{T}_l^{w\sigma}) \cdot \mathbf{n}_l dF
\end{aligned}$$

$$= \gamma_v \sum_{\substack{k=1 \\ k \neq l}}^{3,w} (\mu_{kl} \mathbf{V}_k - \mu_{lk} \mathbf{V}_l), \quad (2.50a)$$

see also Appendix 2.1. It is evident from Eq. (2.50a) that the products of the pulsation velocity components, called *Reynolds stresses*, act on the flow, introducing additional macroscopic cohesion inside the velocity field. Equation (2.50a) is applied on the field  $l$  including the surface up to the  $m$ -side interface.

For dispersed flows it is convenient to have also the momentum equation of the continuum in a primitive form without using the momentum jump condition. In this case, the time average of Eq. (2.31) for field  $m$  after introducing Eq. (2.33) is

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_m \rho_m \mathbf{V}_m \gamma_v) + \nabla \cdot (\alpha_m^e \rho_m \mathbf{V}_m \mathbf{V}_m \gamma) + \nabla \cdot \left[ \alpha_m^e \gamma (\rho_m \overline{\mathbf{V}'_m \mathbf{V}'_m} - \delta_m \mathbf{T}_m) \right] + \alpha_m \gamma_v \rho_m \mathbf{g} \\ & + \alpha_m^e \mathcal{N} p_m - \frac{1}{Vol} \int_{F_{m\sigma}} (\Delta p_m^{l\sigma} \mathbf{I} - \mathbf{T}_m^{l\sigma}) \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{mv}} (\Delta p_m^{w\sigma} \mathbf{I} - \mathbf{T}_m^{w\sigma}) \cdot \mathbf{n}_m dF \\ & = \gamma_v \sum_{\substack{k=1 \\ k \neq m}}^{3,w} (\mu_{km} \mathbf{V}_k - \mu_{mk} \mathbf{V}_m). \end{aligned} \quad (2.50b)$$

Comparing Eq. (2.50a) with Eq. (2.50b) we realize that the term

$$- \frac{1}{Vol} \int_{F_{m\sigma}} (\Delta p_m^{l\sigma} \mathbf{I} - \mathbf{T}_m^{l\sigma}) \cdot \mathbf{n}_l dF$$

appears in both equations with opposite sign. For practical computation we recommend the use of a couple of equations having common interface in order to easily control the momentum conservation at the selected common interface.

## 2.5 Dispersed phase in a laminar continuum – pseudo turbulence

It is known that even low-velocity potential flow over a family of spheres is associated with natural fluctuations of the continuum. The produced oscillations of the laminar continuum are called *pseudo turbulence* by some authors. The averaged pressure over the dispersed particles surface is smaller than the volume averaged pressure. Therefore, in flows with spatially changing concentration of the disperse



phase, an additional force acts towards the concentration gradients. For bubbly flow, *Nigmatulin* (1979) obtained the analytical expression

$$-\overline{\mathbf{V}'_c \mathbf{V}'_c} = |\Delta \mathbf{V}_{cd}|^2 \begin{vmatrix} \frac{4}{20} & 0 & 0 \\ 0 & \frac{3}{20} & 0 \\ 0 & 0 & \frac{3}{20} \end{vmatrix}.$$

*Van Wijngaarden* (1982) used this expression multiplied by  $\alpha_d$ .

## 2.6 Viscous and *Reynolds* stresses

The solid body rotation and translation of the fluid element does not cause any deformation and, therefore, no internal viscous stresses in the fluid. Only the deformation of the fluid element causes viscous stress resisting this deformation. For estimation of the relationship between deformation and viscous stress, the heuristic approach proposed by *Helmholtz* and *Stokes* (1845a and b) can be used for the continuous, intrinsic isotropic, nonstructured field, see *Schlichting* (1959, p. 58). The background conditions behind this approach will now be recalled: (a) the field is a continuum, (b) small velocity changes are considered, (c) only the linear part of the *Taylor* series is taken into account, and (d) linear dependence between stresses and velocity deformations (*Newtonian* continuum). The mathematical notation for this hypothesis is

$$\begin{aligned} \mathbf{T}_\eta &= \eta \left[ \nabla \mathbf{V} + (\nabla \mathbf{V})^T - \frac{2}{3} (\nabla \cdot \mathbf{V}) \mathbf{I} \right] = \eta \left[ 2\mathbf{D} - \frac{2}{3} (\nabla \cdot \mathbf{V}) \mathbf{I} \right] \\ &= \eta \begin{pmatrix} 2 \left( \frac{\partial u}{\partial x} - \frac{1}{3} \nabla \cdot \mathbf{V} \right) & \frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} & \frac{\partial w}{\partial x} + \frac{\partial u}{\partial z} \\ \frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} & 2 \left( \frac{\partial v}{\partial y} - \frac{1}{3} \nabla \cdot \mathbf{V} \right) & \frac{\partial w}{\partial y} + \frac{\partial v}{\partial z} \\ \frac{\partial u}{\partial z} + \frac{\partial w}{\partial x} & \frac{\partial v}{\partial z} + \frac{\partial w}{\partial y} & 2 \left( \frac{\partial w}{\partial z} - \frac{1}{3} \nabla \cdot \mathbf{V} \right) \end{pmatrix}, \end{aligned} \quad (2.51)$$

where  $\mathbf{T}$  is the second-order tensor for the viscous momentum flux,  $\nabla \mathbf{V}$  is the dyadic product of the *nabla* operator and the velocity vector (a second-order tensor), and  $^T$  designates the transposed tensor. Note that the *nabla* operator of the velocity vector,

$$\nabla \mathbf{V} = \mathbf{D} + \mathbf{W}$$

consists of a symmetric part

$$\mathbf{D} = \frac{1}{2} [\nabla \mathbf{V} + (\nabla \mathbf{V})^T],$$

called *deformation rate*, and a skew part

$$\mathbf{W} = \frac{1}{2} [\nabla \mathbf{V} - (\nabla \mathbf{V})^T]$$

called *spin* or *vortices tensor*.

$$\nabla \cdot \mathbf{V} = \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \frac{\partial w}{\partial z}$$

is the divergence of the velocity vector. *Stokes* called the term containing the divergence of the velocity vector the *rate of cubic dilatation*. The hypothesis says that the relation between viscous stresses and the deformation rate of a control volume is linear and the proportionality factor is the dynamic viscosity  $\eta$ , that solid body translations and rotations do not contribute to the viscous forces, that the shear stresses are symmetric, and that the relation between volumetric and the shear viscosity is such that the pressure always equals one third of the sum of the normal stresses. *Stokes* ingeniously argues each of these points in his paper. In the multiphase continuous field models, as long as they are resolved with very fine grids this stress tensor reflects the real one.

I recommend to anyone having the serious intention to understand flows to study this paper. Alternatively, see *Schlichting* (1959, p.60), where it is explained that Eq. (2.51) contains the *Stokes* result for the relation of the bulk viscosity equal to  $-2/3$  dynamic viscosity. From the mechanical equilibrium condition for all angular momenta around an axis for vanishing dimensions of the control volume, one obtains the symmetry of the components of the viscous stress tensor, see *Schlichting* (1959, p. 50) for the *Cartesian* coordinates.

For practical use it is convenient to write the viscous stress tensor for Cartesian and cylindrical coordinates as follows:

$$\mathbf{T}_\eta = \eta \begin{pmatrix} 2 \left[ \frac{\partial u}{\partial r} - \frac{1}{3} (\nabla \cdot \mathbf{V}) \right] & \frac{1}{r^\kappa} \frac{\partial u}{\partial \theta} - \kappa \frac{v}{r^\kappa} + \frac{\partial v}{\partial r} & \frac{\partial u}{\partial z} + \frac{\partial w}{\partial r} \\ \frac{1}{r^\kappa} \frac{\partial u}{\partial \theta} - \kappa \frac{v}{r^\kappa} + \frac{\partial v}{\partial r} & 2 \left( \frac{1}{r^\kappa} \frac{\partial v}{\partial \theta} + \kappa \frac{v}{r^\kappa} - \frac{1}{3} (\nabla \cdot \mathbf{V}) \right) & \frac{\partial v}{\partial z} + \frac{1}{r^\kappa} \frac{\partial w}{\partial \theta} \\ \frac{\partial u}{\partial z} + \frac{\partial w}{\partial r} & \frac{\partial v}{\partial z} + \frac{1}{r^\kappa} \frac{\partial w}{\partial \theta} & 2 \left[ \frac{\partial w}{\partial z} - \frac{1}{3} (\nabla \cdot \mathbf{V}) \right] \end{pmatrix},$$

where

$$\nabla \cdot \mathbf{V} = \frac{1}{r^\kappa} \frac{\partial}{\partial r} (r^\kappa u) + \frac{1}{r^\kappa} \frac{\partial v}{\partial \theta} + \frac{\partial w}{\partial z}.$$

For cylindrical coordinates  $\kappa = 1$ . For Cartesian coordinates set  $\kappa = 0$  and replace  $r, \theta, z$  with  $x, y, z$ , respectively.

Now let us step to the turbulence stress tensor  $-\rho_l \overline{\mathbf{V}'_l \mathbf{V}'_l}$  called the *Reynolds* stress tensor. The search for a quantitative estimation of the *Reynolds* stresses for multiphase flows is in its initial stage. A possible step in the right direction, in analogy to single-phase turbulence, is the use of the *Boussinesq* hypothesis (1877) for the viscosity of turbulent eddies inside the velocity field. *Boussinesq* introduced the idea of turbulent eddy viscosity inside the velocity field,

$$\begin{aligned}
 -\rho_l \overline{\mathbf{V}'_l \mathbf{V}'_l} &= \eta_l^t \left[ 2\mathbf{D}_l - \frac{2}{3}(\nabla \cdot \mathbf{V}_l) \mathbf{I} \right] \\
 &= \eta_l^t \begin{pmatrix} 2\left(\frac{\partial u}{\partial x} - \frac{1}{3}\nabla \cdot \mathbf{V}\right) & \frac{\partial v}{\partial x} + \frac{\partial u}{\partial y} & \frac{\partial w}{\partial x} + \frac{\partial u}{\partial z} \\ \frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} & 2\left(\frac{\partial v}{\partial y} - \frac{1}{3}\nabla \cdot \mathbf{V}\right) & \frac{\partial w}{\partial y} + \frac{\partial v}{\partial z} \\ \frac{\partial u}{\partial z} + \frac{\partial w}{\partial x} & \frac{\partial v}{\partial z} + \frac{\partial w}{\partial y} & 2\left(\frac{\partial w}{\partial z} - \frac{1}{3}\nabla \cdot \mathbf{V}\right) \end{pmatrix}_l, \quad (2.52)
 \end{aligned}$$

so that it has the same structure as the *Stokes* hypothesis. The corresponding *Reynolds* stresses are

$$\begin{aligned}
 -\rho_l u'_l u'_l &= \tau'_{l,xx} = \eta_l^t 2 \left( \frac{\partial u_l}{\partial x} - \frac{1}{3} \nabla \cdot \mathbf{V}_l \right), \\
 -\rho_l v'_l v'_l &= \tau'_{l,yy} = \eta_l^t 2 \left( \frac{\partial v_l}{\partial y} - \frac{1}{3} \nabla \cdot \mathbf{V}_l \right), \\
 -\rho_l w'_l w'_l &= \tau'_{l,zz} = \eta_l^t 2 \left( \frac{\partial w_l}{\partial z} - \frac{1}{3} \nabla \cdot \mathbf{V}_l \right), \\
 -\rho_l u'_l v'_l &= \tau'_{l,xy} = \eta_l^t \left( \frac{\partial v_l}{\partial x} + \frac{\partial u_l}{\partial y} \right), \\
 -\rho_l u'_l w'_l &= \tau'_{l,xz} = \eta_l^t \left( \frac{\partial w_l}{\partial x} + \frac{\partial u_l}{\partial z} \right), \\
 -\rho_l v'_l w'_l &= \tau'_{l,yz} = \eta_l^t \left( \frac{\partial w_l}{\partial y} + \frac{\partial v_l}{\partial z} \right).
 \end{aligned}$$

The dynamic turbulent viscosity now is a flow property and remains to be estimated. Note that at a given point this is a single value for all directions. Strictly speaking, this approach is valid for isotropic turbulence because there is a single eddy viscosity assumed to be valid for all directions.

An alternative notation of the term  $\nabla \cdot [\alpha_i^e (\rho_l \overline{\mathbf{V}_l \mathbf{V}_l'}) \gamma]$  is given here for isotropic turbulence, for which

$$u'_i u'_i = v'_i v'_i = w'_i w'_i = \frac{2}{3} k_i : \quad (2.53)$$

$$\begin{aligned} \nabla \cdot [(\alpha_i^e \rho_l \mathbf{V}_l \mathbf{V}_l') \gamma] &= \begin{pmatrix} \frac{\partial \gamma_x \alpha_i^e \rho_l u'_i u'_i}{\partial x} + \frac{\partial \gamma_y \alpha_i^e \rho_l u'_i v'_i}{\partial y} + \frac{\partial \gamma_z \alpha_i^e \rho_l u'_i w'_i}{\partial z} \\ \frac{\partial \gamma_x \alpha_i^e \rho_l v'_i u'_i}{\partial x} + \frac{\partial \gamma_y \alpha_i^e \rho_l v'_i v'_i}{\partial y} + \frac{\partial \gamma_z \alpha_i^e \rho_l v'_i w'_i}{\partial z} \\ \frac{\partial \gamma_x \alpha_i^e \rho_l w'_i u'_i}{\partial x} + \frac{\partial \gamma_y \alpha_i^e \rho_l w'_i v'_i}{\partial y} + \frac{\partial \gamma_z \alpha_i^e \rho_l w'_i w'_i}{\partial z} \end{pmatrix} \\ &= \begin{pmatrix} \frac{\partial \gamma_y \alpha_i^e \rho_l u'_i v'_i}{\partial y} + \frac{\partial \gamma_z \alpha_i^e \rho_l u'_i w'_i}{\partial z} \\ \frac{\partial \gamma_x \alpha_i^e \rho_l v'_i u'_i}{\partial x} + \frac{\partial \gamma_z \alpha_i^e \rho_l v'_i w'_i}{\partial z} \\ \frac{\partial \gamma_x \alpha_i^e \rho_l w'_i u'_i}{\partial x} + \frac{\partial \gamma_y \alpha_i^e \rho_l w'_i v'_i}{\partial y} \end{pmatrix} + \frac{2}{3} \nabla (\gamma \alpha_i^e \rho_l k_i) \\ &= -\tilde{\mathbf{S}}_i + \frac{2}{3} \nabla (\gamma \alpha_i^e \rho_l k_i), \end{aligned} \quad (2.54)$$

where

$$\tilde{\mathbf{S}}_i = \begin{pmatrix} \frac{\partial}{\partial y} \left[ \gamma_y \alpha_i^e \rho_l v'_i \left( \frac{\partial v'_i}{\partial x} + \frac{\partial u'_i}{\partial y} \right) \right] + \frac{\partial}{\partial z} \left[ \gamma_z \alpha_i^e \rho_l v'_i \left( \frac{\partial w'_i}{\partial x} + \frac{\partial u'_i}{\partial z} \right) \right] \\ \frac{\partial}{\partial x} \left[ \gamma_x \alpha_i^e \rho_l v'_i \left( \frac{\partial v'_i}{\partial x} + \frac{\partial u'_i}{\partial y} \right) \right] + \frac{\partial}{\partial z} \left[ \gamma_z \alpha_i^e \rho_l v'_i \left( \frac{\partial w'_i}{\partial y} + \frac{\partial v'_i}{\partial z} \right) \right] \\ \frac{\partial}{\partial x} \left[ \gamma_x \alpha_i^e \rho_l v'_i \left( \frac{\partial w'_i}{\partial x} + \frac{\partial u'_i}{\partial z} \right) \right] + \frac{\partial}{\partial y} \left[ \gamma_y \alpha_i^e \rho_l v'_i \left( \frac{\partial w'_i}{\partial y} + \frac{\partial v'_i}{\partial z} \right) \right] \end{pmatrix}. \quad (2.55)$$

Here the diagonal symmetric term  $\frac{2}{3} \nabla (\gamma \alpha_i^e \rho_l k_i)$  is considered as a dispersion force and is directly computed from the turbulent kinetic energy delivered by the turbulence model.

It is important to emphasize that, in spite of the fact that several processes in single-phase fluid dynamics can be successfully described by the *Helmholtz–Stokes* and by the *Boussinesq* hypotheses; these hypotheses have never been derived from experiments or proven by abstract arguments. This limitation of the hypotheses should be borne in mind when they are applied.

While the heuristic approach proposed by *Helmholtz* and *Stokes*, Eq. (2.51), is valid only for the continuous part of each velocity field, the *Boussinesq* hypothesis is useful for continuous and disperse velocity fields. This behavior is again described here by introducing for each velocity field the multiplier  $\delta_i$  in Eq. (2.50), where  $\delta_i = 0$  for dispersed field and  $\delta_i = 1$  for continuous nonstructured fields.

For a single field,  $\delta_i = 1$ , the description of the viscous and *Reynolds* stresses reduces to the widely accepted expression.

It is plausible to define the turbulent pressure as

$$p'_i = \rho_i \frac{2}{3} k_i = \rho_i \frac{1}{3} (u'u' + v'v' + w'w'), \quad (2.56)$$

and to consider the term  $-\alpha_i^e \gamma \nabla p'_i$  as absorbed from the pressure term  $\alpha_i^e \gamma \nabla p_i$  and the term

$$-p'_i \nabla (\alpha_i^e \gamma) = -\frac{1}{Vol} \int_{F_{i\sigma}} p'_i \mathbf{n}_i dF + \frac{1}{Vol} \int_{F_{iw}} p'_i \mathbf{n}_i dF \quad (2.57)$$

as included in the pressure differences between the bulk pressure and boundary layer pressure. This could mean that  $p'_i$  no longer needs to appear in the notation. Until the correctness of this agglomeration of the terms is not strictly proven I do not recommend it.

It is interesting to note that from the kinetic theory for two colliding particles plus their added mass the following is obtained:

$$\nabla (\alpha_i^e \gamma p'_i) \approx \nabla \left[ \alpha_i^e \gamma (\rho_i + c_i^{vm} \rho_m) \frac{1}{3} \mathbf{V}_i'^2 \right]. \quad (2.58)$$

## 2.7 Nonequal bulk and boundary layer pressures

### 2.7.1 Continuous interface

#### 2.7.1.1 3D flows

Examples for the existence of continuous interfaces are the stratified pool, see Fig. 2.5, and annular pipe flow.

The treatment of the interface depends very much on the numerical method used. If the numerical method is able to resolve the interface itself and the two attached boundary layers (see, for instance, *Hirt* and *Nichols* (1981) for more information), the interface momentum jump condition is the only information needed to close the mathematical description of the interface. If this is not the case, special treatment of the processes at the interface is necessary. In this section we discuss some possibilities.

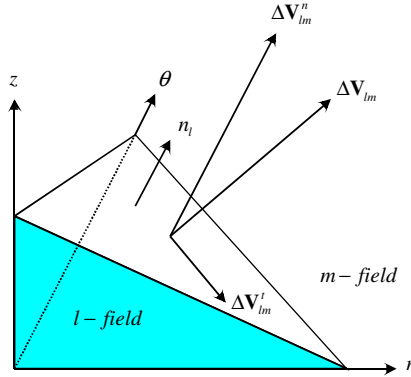


Fig. 2.5 Continuous interface

Consider a liquid ( $l$ )-gas ( $m$ ) flow without mass transfer. The computational cells are so large that the surface is at best represented by piecewise planes at which the surface tension is neglected. The compressibility of the gas is much larger than the compressibility of the liquid. In this case, we can assume that there is almost no difference between the bulk and liquid side interface pressure

$$\Delta p_l^{m\sigma} = 0. \quad (2.59)$$

The pressure change across the gas side boundary layer is then approximated by the stagnation pressure

$$\Delta p_m^{l\sigma} = \rho_m (V_l^n - V_m^n)^2 = \rho_m |\Delta \mathbf{V}_{lm}^n|^2 \text{sign}(V_l^n - V_m^n), \quad (2.60)$$

see Fig. 2.6. The normal velocity difference required in the above expression can be obtained by splitting the relative velocity vector at the interface  $\Delta \mathbf{V}_{lm}^n$  into a component that is parallel to  $\mathbf{n}_l$

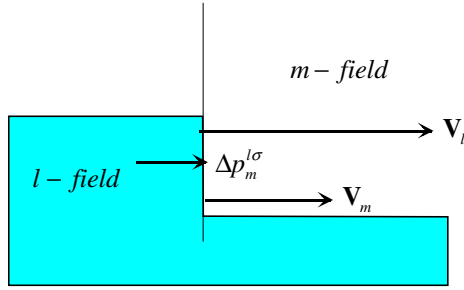
$$\Delta \mathbf{V}_{lm}^n = \text{proj}_{\mathbf{n}_l} \Delta \mathbf{V}_{lm} = \left( \frac{\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l}{\mathbf{n}_l \cdot \mathbf{n}_l} \right) \mathbf{n}_l = (\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l) \mathbf{n}_l, \quad (2.61)$$

with a magnitude

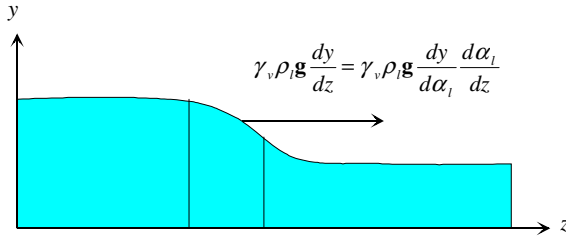
$$|\Delta \mathbf{V}_{lm}^n| = |\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l| = \sqrt{[n_{lx}(u_l - u_m)]^2 + [n_{ly}(v_l - v_m)]^2 + [n_{lz}(w_l - w_m)]^2} \quad (2.62)$$

and a component orthogonal to  $\mathbf{n}_l$ ,

$$\begin{aligned} \Delta \mathbf{V}_{lm}^t &= \Delta \mathbf{V}_{lm} - \text{proj}_{\mathbf{n}_l} \Delta \mathbf{V}_{lm} = \Delta \mathbf{V}_{lm} - (\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l) \mathbf{n}_l \\ &= [\Delta u_{lm} - n_{lx}(\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l)] \mathbf{i} + [\Delta v_{lm} - n_{ly}(\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l)] \mathbf{j} + [\Delta w_{lm} - n_{lz}(\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l)] \mathbf{k}. \end{aligned} \quad (2.63)$$



**Fig. 2.6** Stagnation pressure in stratified flow



**Fig. 2.7** Geodesic pressure force

In a similar way, the stagnation pressure difference at the field-structure interface can be estimated,

$$\Delta p_l^{w\sigma} = \rho_l (V_l^n - V_w^n)^2 \text{sign}(V_l^n - V_w^n). \quad (2.64)$$

For the case of  $\Delta p_m^{l\sigma} \approx \text{const}$  within  $Vol$  the following can be written:

$$\frac{1}{Vol} \int_{F_{l\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l dF = \Delta p_m^{l\sigma} \frac{1}{Vol} \int_{F_{l\sigma}} \mathbf{n}_l dF = -\Delta p_m^{l\sigma} \left[ \nabla(\alpha_l^e \gamma) + \frac{1}{Vol} \int_{F_w} \mathbf{n}_l dF \right]. \quad (2.65)$$

For  $\mathbf{V}_l^n > \mathbf{V}_m^n$  and decreasing  $\alpha_l^e \gamma$  in space, this force resists the field  $l$ . If there is no difference in the average normal velocities at the interface the above term is zero.

If the tangential average velocity difference differs from zero, there is a tangential viscous shear force

$$-\frac{1}{Vol} \int_{F_{l\sigma}} \mathbf{T}_{m\sigma} \cdot \mathbf{n}_{l\sigma} dF = c_{ml} |\Delta \mathbf{V}_{lm}^t| |\Delta \mathbf{V}_{lm}^t|. \quad (2.66)$$

Here  $c_{ml}$  has to be computed using empirical correlation in the case of the large scale of the cells not resolving the details of the boundary layer. For a wavy sur-

face, the interface share coefficient should be increased by a component for form drag caused by the nonuniform pressure distribution, which results in an additional tangential force. Similarly, the viscous shear stress of the wall is

$$-\frac{1}{Vol} \int_{F_w} \mathbf{T}_{lw} \cdot \mathbf{n}_{lw} dF = c_{wl} |\mathbf{V}_l^t| \mathbf{V}_l^t, \quad (2.67)$$

where, like in the previous case,  $c_{wl}$  has to be computed using empirical correlation. Note that in the case of stratified rectangular duct flow in the  $z$  direction, see Fig. 2.7, the change in the liquid thickness causes a lateral geodesic pressure force

$$\gamma_v \rho_l \mathbf{g} \frac{dy}{dz} = \gamma_v \rho_l \mathbf{g} \frac{dy}{d\alpha_1} \frac{d\alpha_1}{dz}. \quad (2.68)$$

Here  $y$  is the distance between the bottom of the duct and the center of mass of the liquid. In three-dimensional models this force automatically arises due to differences in the local bulk pressure having the geodesic pressure as a component. This force should be taken into account in one-dimensional models. If this force is neglected, the one-dimensional model will not be able to predict water flow in a horizontal pipe with negligible gas-induced shear. In the next section we consider this problem in more detail.

### 2.7.1.2 Stratified flow in horizontal or inclined rectangular channels

*Geometrical characteristics:* Stratified flow may exist in regions with such relative velocities between the liquid and the gas which does not cause instabilities leading to slugging.

Some important geometrical characteristics are specified here – compare with Fig. 2.8. The perimeter of the pipe is then

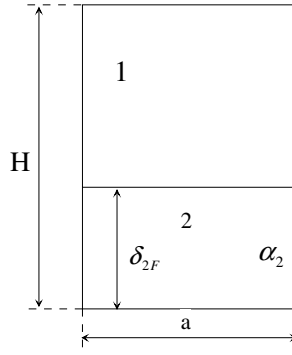
$$Per_{1w} = 2(a + H), \quad (2.69)$$

and the wetted perimeters for the gas and the liquid parts are

$$Per_{1w} = a + 2(H - \delta_{2F}) = a + 2\alpha_1 H, \quad (2.70)$$

$$Per_{2w} = a + 2\delta_{1F} = a + 2\alpha_2 H. \quad (2.71)$$





**Fig. 2.8** Definition of the geometrical characteristics of the stratified flow

The gas–liquid interface median is then  $a$ , and the liquid level

$$\delta_{2F} = \alpha_2 H. \quad (2.72)$$

The hydraulic diameters for the gas and the liquid for computation of the pressure drop due to friction with the wall are, therefore,

$$D_{h1} = 4\alpha_1 F / Per_{1w} = \frac{4\alpha_1 aH}{a + 2H\alpha_1}, \quad (2.73)$$

$$D_{h2} = 4\alpha_2 F / Per_{2w} = \frac{4\alpha_2 aH}{a + 2H\alpha_2}, \quad (2.74)$$

and the corresponding *Reynolds* numbers

$$Re_1 = \frac{\alpha_1 \rho_1 w_1}{\eta_1} \frac{4\alpha_1 aH}{a + 2H\alpha_1}, \quad (2.75)$$

$$Re_2 = \frac{\alpha_2 \rho_2 w_2}{\eta_2} \frac{4\alpha_2 aH}{a + 2H\alpha_2}. \quad (2.76)$$

Here  $F$  is the channel cross-section and  $Per_{1w}$  and  $Per_{2w}$  are the perimeters wet by gas and film, respectively. If one considers the gas-core of the flow, the hydraulic diameter for computation of the pressure loss component due to the gas-liquid friction is then

$$D_{h12} = 4\alpha_1 F / (Per_{1w} + a) = \frac{2\alpha_1 aH}{a + H\alpha_1} \quad (2.77)$$

and the corresponding *Reynolds* number

$$Re_1 = \frac{\alpha_1 \rho_1 |w_1 - w_2|}{\eta_1} \frac{2\alpha_1 aH}{a + H\alpha_1}. \quad (2.78)$$

The gas-wall, liquid-wall, and gas-liquid interfacial area densities are

$$a_{1w} = \frac{Per_{1w}}{F} = \frac{a + 2\alpha_1 H}{aH}, \quad (2.79)$$

$$a_{2w} = \frac{Per_{2w}}{F} = \frac{a + 2\alpha_2 H}{aH}, \quad (2.80)$$

$$a_{12} = \frac{a}{F} = \frac{1}{H}. \quad (2.81)$$

For the estimation of the flow pattern transition criterion, the following expression is sometimes required:

$$\frac{d\alpha_2}{d\delta_{2F}} = \frac{1}{H}. \quad (2.82)$$

Using the geometric characteristics and the Reynolds numbers the interfacial interaction coefficients can be computed by means of empirical correlations as discussed in Volume II of this monograph.

*Gravitational (hydrostatic) pressure variation across the flow cross-section of horizontal pipe:* In stratified flow the gravitation is a dominant force. The cross-section averaged gas gravity pressure difference with respect to the interface is

$$\Delta p_1^{2\sigma} = \rho_1 g \alpha_1 \frac{H}{2}. \quad (2.83)$$

The cross-section averaged liquid gravity pressure difference with respect to the interface is

$$\Delta p_2^{1\sigma} = -\rho_2 g \alpha_2 \frac{H}{2}. \quad (2.84)$$

The cross-section averaged field pressures in terms of the interfacial pressure are then

$$p_1 = p_1^{2\sigma} - \Delta p_1^{2\sigma} = p_1^{2\sigma} - \rho_1 g \alpha_1 \frac{H}{2}, \quad (2.85)$$

$$p_2 = p_2^{1\sigma} - \Delta p_2^{1\sigma} = p_2^{1\sigma} + \rho_2 g \alpha_2 \frac{H}{2}, \quad (2.86)$$

recalling the definition in Eqs. (2.33) and (2.34). The averaged pressure in the cross-section can be expressed as the cross-section weighted averaged pressures inside the fields

$$p = \alpha_1 p_1 + \alpha_2 p_2 = \alpha_1 p_1^{2\sigma} + \alpha_2 p_2^{1\sigma} + g \frac{H}{2} (\alpha_2^2 \rho_2 - \alpha_1^2 \rho_1). \quad (2.87)$$

Neglecting the surface tension, we obtain

$$p \approx p_\sigma + g \frac{H}{2} (\alpha_2^2 \rho_2 - \alpha_1^2 \rho_1), \quad (2.88)$$

or

$$p_\sigma \approx p - g \frac{H}{2} (\alpha_2^2 \rho_2 - \alpha_1^2 \rho_1). \quad (2.89)$$

The averaged pressure in the gas and in the liquid phase can be then expressed as a function of the system averaged pressure and geometrical characteristics by replacing  $p_\sigma$  in Eqs. (2.85) and (2.86)

$$p_1 = p - \frac{H}{2} g \alpha_2 \rho \quad \text{for } \alpha_1 > 0, \quad (2.90)$$

$$p_2 = p + \frac{H}{2} g \alpha_1 \rho \quad \text{for } \alpha_2 > 0, \quad (2.91)$$

where  $\rho = \alpha_1 \rho_1 + \alpha_2 \rho_2$  is the homogeneous mixture density. The check  $\alpha_1 p_1 + \alpha_2 p_2 = p$  proves the correctness of the computation. The difference between both averaged pressures is then

$$\Delta p_{21} = p_2 - p_1 = \frac{H}{2} g \rho \quad \text{for } \alpha_1 > 0 \text{ and } \alpha_2 > 0, \quad (2.92)$$

*Delhaye* (1981, p.89). Therefore

$$p_1 = p - \alpha_2 \Delta p_{21}, \quad (2.93)$$

$$p_2 = p + \alpha_1 \Delta p_{21}, \quad (2.94)$$

and consequently

$$\frac{\partial p_1}{\partial z} = \frac{\partial p}{\partial z} - \Delta p_{21} \frac{\partial \alpha_2}{\partial z} - \alpha_2 \frac{\partial \Delta p_{21}}{\partial z}, \quad (2.95)$$

$$\frac{\partial p_2}{\partial z} = \frac{\partial p}{\partial z} + \Delta p_{21} \frac{\partial \alpha_1}{\partial z} + \alpha_1 \frac{\partial \Delta p_{21}}{\partial z}. \quad (2.96)$$

Now we can write the specific form of the following general terms of the momentum equation:

$$\begin{aligned} & \dots + \alpha_i^e \gamma_z \frac{\partial p_i}{\partial z} + (p_l - p_m - \delta_i \sigma_{lm} \kappa_l - \Delta p_m^{l\sigma}) \left[ \frac{\partial}{\partial z} (\alpha_i^e \gamma_z) - \delta_i \frac{\partial \gamma_z}{\partial z} \right] \\ & - \Delta p_i^{w\sigma} \delta_i \frac{\partial \gamma_z}{\partial z} + \gamma_v \alpha_i \rho_l g \cos \varphi \dots \end{aligned} \quad (2.97)$$

Note that both fields are continuous and, therefore,  $\delta_l = 1$  and the effect of the surface tension is neglected

$$\dots + \alpha_1 \gamma_z \frac{\partial p_1}{\partial z} - \gamma_z (\Delta p_{21} + \Delta p_2^{1\sigma}) \frac{\partial \alpha_1}{\partial z} - \Delta p_1^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (2.98)$$

$$\dots + \alpha_2 \gamma_z \frac{\partial p_2}{\partial z} + \gamma_z (\Delta p_{21} - \Delta p_1^{2\sigma}) \frac{\partial \alpha_2}{\partial z} - \Delta p_2^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (2.99)$$

Substituting the field pressures, we finally obtain

$$\dots + \gamma_z \left[ \alpha_1 \frac{\partial p}{\partial z} - \alpha_1 \alpha_2 \frac{\partial \Delta p_{21}}{\partial z} - (\alpha_2 \Delta p_{21} + \Delta p_2^{1\sigma}) \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_1^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (2.100)$$

$$\dots + \gamma_z \left[ \alpha_2 \frac{\partial p}{\partial z} + \alpha_1 \alpha_2 \frac{\partial \Delta p_{21}}{\partial z} - (\alpha_1 \Delta p_{21} - \Delta p_1^{2\sigma}) \frac{\partial \alpha_2}{\partial z} \right] - \Delta p_2^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (2.101)$$

Assuming that the change of the densities contributes much less to the change of

$$\frac{\partial \Delta p_{21}}{\partial z} \approx \frac{\partial \Delta p_{21}}{\partial \alpha_1} \frac{\partial \alpha_1}{\partial z} = g \frac{H}{2} (\rho_1 - \rho_2) \frac{\partial \alpha_1}{\partial z} \quad (2.102)$$

then the change of the local volume fraction becomes

$$\dots + \gamma_z \left[ \alpha_1 \frac{\partial p}{\partial z} - g H \alpha_1 \alpha_2 (\rho_1 - \rho_2) \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_1^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (2.103)$$

$$\dots + \gamma_z \left[ \alpha_2 \frac{\partial p}{\partial z} + g H \alpha_1 \alpha_2 (\rho_1 - \rho_2) \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_2^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (2.104)$$

As a plausibility check note that for  $\alpha_1 \rightarrow 0$  or  $\alpha_2 \rightarrow 0$  the term containing the derivative of the volume fraction of velocity field 1 converges to zero and the momentum equations take the expected form. The sum of the two momentum equations gives

$$\dots + \gamma_z \frac{\partial p}{\partial z} - (\Delta p_1^{w\sigma} + \Delta p_2^{w\sigma}) \frac{\partial \gamma_z}{\partial z} \dots \quad (2.105)$$

Note that the gravitational force is already taken into account in Eqs. (2.103) and (2.104) and there is no need for an additional term  $\gamma_z \alpha_1 \rho_1 g \cos \varphi$ . Stability criteria for the stratified flow can be obtained from the eigenvalue analysis. For simplicity, assuming incompressible flow the mass and momentum equations of stratified flow in a straight pipe with constant cross-section section are

$$\alpha_1 \frac{\partial w_1}{\partial z} + \alpha_2 \frac{\partial w_2}{\partial z} + (w_1 - w_2) \frac{\partial \alpha_1}{\partial z} = 0, \quad (2.106)$$

$$\frac{\partial \alpha_1}{\partial \tau} + \alpha_1 \frac{\partial w_1}{\partial z} + w_1 \frac{\partial \alpha_1}{\partial z} = 0, \quad (2.107)$$

$$\frac{\partial w_1}{\partial \tau} + w_1 \frac{\partial w_1}{\partial z} + \frac{1}{\rho_1} \frac{\partial p}{\partial z} + gH \frac{\alpha_2 (\rho_2 - \rho_1)}{\rho_1} \frac{\partial \alpha_1}{\partial z} = 0, \quad (2.108)$$

$$\frac{\partial w_2}{\partial \tau} + w_2 \frac{\partial w_2}{\partial z} + \frac{1}{\rho_2} \frac{\partial p}{\partial z} - gH \frac{\alpha_1 (\rho_2 - \rho_1)}{\rho_2} \frac{\partial \alpha_1}{\partial z} = 0, \quad (2.109)$$

or in matrix notation

$$\begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \frac{\partial}{\partial \tau} \begin{pmatrix} p \\ \alpha_1 \\ w_1 \\ w_2 \end{pmatrix} + \begin{pmatrix} 0 & (w_1 - w_2) & \alpha_1 & \alpha_2 \\ 0 & w_1 & \alpha_1 & 0 \\ \frac{1}{\rho_1} & gH \frac{\alpha_2 (\rho_2 - \rho_1)}{\rho_1} & w_1 & 0 \\ \frac{1}{\rho_2} & -gH \frac{\alpha_1 (\rho_2 - \rho_1)}{\rho_2} & 0 & w_2 \end{pmatrix} \frac{\partial}{\partial z} \begin{pmatrix} p \\ \alpha_1 \\ w_1 \\ w_2 \end{pmatrix} = 0. \quad (2.110)$$

For the reader who is not familiar with the analysis of the type of a system of partial differential equations by first computing the eigenvalues, eigenvectors, and canonical forms it is recommended to first read Section 11 before continuing here.

The eigenvalues are defined by the characteristics equations

$$\begin{pmatrix} 0 & (w_1 - w_2) & \alpha_1 & \alpha_2 \\ 0 & w_1 - \lambda & \alpha_1 & 0 \\ \frac{1}{\rho_1} & gH \frac{\alpha_2 (\rho_2 - \rho_1)}{\rho_1} & w_1 - \lambda & 0 \\ \frac{1}{\rho_2} & -gH \frac{\alpha_1 (\rho_2 - \rho_1)}{\rho_2} & 0 & w_2 - \lambda \end{pmatrix} = 0, \quad (2.111)$$

or

$$\begin{aligned} & \frac{\alpha_1}{\rho_1} (w_1 - w_2)(w_2 - \lambda) - \frac{\alpha_1}{\rho_1} (w_1 - \lambda)(w_2 - \lambda) - \frac{\alpha_2}{\rho_2} (w_1 - \lambda)^2 \\ & + gH \frac{\alpha_1 \alpha_2 (\rho_2 - \rho_1)}{\rho_1 \rho_2} = 0, \end{aligned} \quad (2.112)$$

or

$$\left( \frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2} \right) \lambda^2 - 2 \left( \frac{\alpha_1}{\rho_1} w_2 + \frac{\alpha_2}{\rho_2} w_1 \right) \lambda + \frac{\alpha_1}{\rho_1} w_2^2 + \frac{\alpha_2}{\rho_2} w_1^2 - gH \frac{\alpha_1 \alpha_2 (\rho_2 - \rho_1)}{\rho_1 \rho_2} = 0. \quad (2.113)$$

This equation is, in fact, consistent with the long wave gravity theory by *Milne-Thomson* (1968). There are two eigenvalues

$$\lambda_{1,2} = \frac{\frac{\alpha_1}{\rho_1} w_2 + \frac{\alpha_2}{\rho_2} w_1 \pm \sqrt{\left(\frac{\alpha_1}{\rho_1} w_2 + \frac{\alpha_2}{\rho_2} w_1\right)^2 - \left(\frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2}\right) \left(\frac{\alpha_1}{\rho_1} w_2^2 + \frac{\alpha_2}{\rho_2} w_1^2 - gH \frac{\alpha_1 \alpha_2 (\rho_2 - \rho_1)}{\rho_1 \rho_2}\right)}}{\left(\frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2}\right)}, \quad (2.114)$$

or after rearranging

$$\lambda_{1,2} = \frac{\frac{\alpha_1}{\rho_1} w_2 + \frac{\alpha_2}{\rho_2} w_1 \pm \sqrt{\left[ gH (\rho_2 - \rho_1) \left(\frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2}\right) - (w_1 - w_2)^2 \right] \frac{\alpha_1 \alpha_2}{\rho_1 \rho_2}}}{\left(\frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2}\right)}, \quad (2.115)$$

which are real and different from each other if

$$(w_1 - w_2)^2 < gH (\rho_2 - \rho_1) \left(\frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2}\right). \quad (2.116)$$

In fact this is the *Kelvin–Helmholtz* stability criterion. If the above condition is satisfied the system describing the flow is hyperbolic. In nature, violation of the above condition results in flow patterns that are different from the stratified one. Condition (2.116) is equivalent to Eq. (2.216) derived by *Delhaye* (1981, p. 90). In 1992 *Brauner* and *Maron* included the surface tension effect in their stability analysis and obtained

$$(w_1 - w_2)^2 < H \left(\frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2}\right) \left[ g (\rho_2 - \rho_1) + \sigma_{12} k^2 \right],$$

see Eq. (28a) in *Brauner* and *Maron* (1992), which for neglected surface tension results in Eq. (2.116). Here  $k$  is the real wave number.

### 2.7.1.3 Stratified flow in horizontal or inclined pipes

*Geometrical characteristics:* The geometric flow characteristics for round pipes are nonlinearly dependent on the liquid level, which makes the computation somewhat more complicated.

Some important geometrical characteristics are specified here – compare with Fig. 2.9. The angle with the origin of the pipe axis defined between the upwards oriented vertical and the liquid-gas-wall triple point is defined as a function of the liquid volume fraction by the equation

$$f(\theta) = -(1 - \alpha_2)\pi + \theta - \sin\theta \cos\theta = 0. \quad (2.117)$$

The derivative

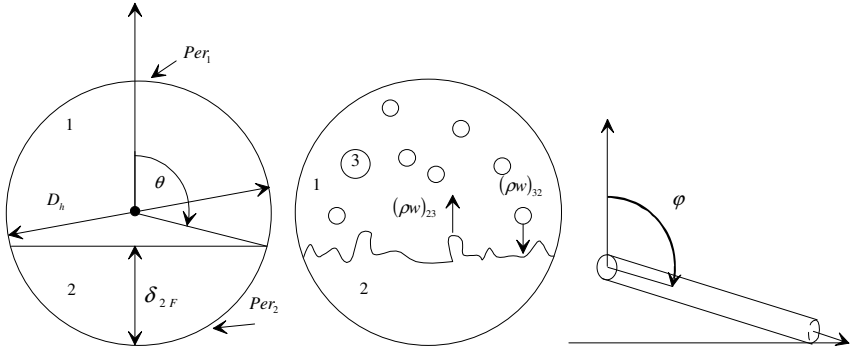
$$\frac{d\theta}{d\alpha_1} = \frac{\pi}{2\sin^2\theta} \quad (2.118)$$

will be used later. Bearing in mind that

$$\frac{df}{d\theta} = 2\sin^2\theta, \quad (2.119)$$

the solution with respect to the angle can be obtained by using the *Newton* iteration method as follows:

$$\theta = \theta_0 - \frac{f_0}{df/d\theta} = \theta_0 + \frac{(1 - \alpha_2)\pi - \theta_0 + \sin\theta_0 \cos\theta_0}{2\sin^2\theta_0}, \quad (2.120)$$



**Fig. 2.9** Definition of the geometrical characteristics of the stratified flow

where subscript  $\theta$  indicates the previous guess in the iteration. The iteration starts with an initial value of  $\pi/2$  (Kolev 1977). In 1999 *Biberg* proposed an accurate direct approximation

$$\theta = \pi\alpha_2 + \left(\frac{3\pi}{2}\right)^{1/3} \left[1 - 2\alpha_2 + \alpha_2^{1/3} - (1 - \alpha_2)^{1/3}\right] \quad (2.120b)$$

with an error of less than  $\pm 0.002\text{rad}$  or

$$\theta = \pi\alpha_2 + \left(\frac{3\pi}{2}\right)^{1/3} \left[1 - 2\alpha_2 + \alpha_2^{1/3} - (1 - \alpha_2)^{1/3}\right] - \frac{1}{200}\alpha_2(1 - \alpha_2)(1 - 2\alpha_2)\left\{1 + 4\left[\alpha_2^2 + (1 - \alpha_2)^2\right]\right\}, \quad (2.120c)$$

with an error less than  $\pm 0.00005 \text{ rad}$ . The perimeter of the pipe is then

$$Per_{1w} = \pi D_h, \quad (2.121)$$

and the wetted perimeters for the gas and the liquid parts are

$$Per_{1w} = \theta D_h, \quad (2.122)$$

$$Per_{2w} = (\pi - \theta) D_h. \quad (2.123)$$

The gas-liquid interface median is then

$$b = D_h \sin(\pi - \theta) = D_h \sin \theta, \quad (2.124)$$

and the liquid level is

$$\delta_{2F} = \frac{1}{2} D_h (1 + \cos \theta). \quad (2.125)$$

The hydraulic diameters for the gas and the liquid for computation of the pressure drop due to friction with the wall are, therefore,

$$D_{h1} = 4\alpha_1 F / Per_{1w} = \frac{\pi}{\theta} \alpha_1 D_h, \quad (2.126)$$

$$D_{h2} = 4\alpha_2 F / Per_{2w} = \frac{\pi}{\pi - \theta} \alpha_2 D_h, \quad (2.127)$$

and the corresponding *Reynolds* numbers are

$$Re_1 = \frac{\alpha_1 \rho_1 w_1 D_h}{\eta_1} \frac{\pi}{\theta}, \quad (2.128)$$

$$Re_2 = \frac{\alpha_2 \rho_2 w_2 D_h}{\eta_2} \frac{\pi}{\pi - \theta}. \quad (2.129)$$

Here  $F$  is the channel cross-section section and  $Per_1$  and  $Per_2$  are the wet perimeters of gas and film, respectively.

If one considers the core of the flow the hydraulic diameter for computation of the friction pressure loss component at the gas-liquid interface is then

$$D_{h12} = 4\alpha_1 F / (Per_{1w} + b) = \frac{\pi}{\theta + \sin \theta} \alpha_1 D_h, \quad (2.130)$$

and the corresponding *Reynolds* number is

$$Re_1 = \frac{\alpha_1 \rho_1 |w_1 - w_2| D_h}{\eta_1} \frac{\pi}{\theta + \sin \theta}. \quad (2.131)$$

The gas-wall, liquid-wall, and gas-liquid interfacial area densities are



$$a_{1w} = \frac{Per_{1w}}{F} = \frac{4\alpha_1}{D_{h1}} = \frac{\theta}{\pi} \frac{4}{D_h}, \quad (2.132)$$

$$a_{2w} = \frac{Per_{2w}}{F} = \frac{4\alpha_2}{D_{h2}} = \frac{\pi - \theta}{\pi} \frac{4}{D_h}, \quad (2.133)$$

$$a_{12} = \frac{b}{F} = \frac{\sin \theta}{\pi} \frac{4}{D_h}. \quad (2.134)$$

Some authors approximated this relation for a smooth interface with

$$a_{12} \cong \frac{8}{\pi D_h} \sqrt{\alpha_2 (1 - \alpha_2)}, \quad (2.135)$$

which in view of the accurate computation presented above is no longer necessary. For the estimation of the flow pattern transition criterion, the following expression is sometimes required:

$$\frac{d\alpha_2}{d\delta_{2F}} = \frac{4}{D_h} \frac{\sin \theta}{\pi}. \quad (2.136)$$

*Gravitational (hydrostatic) pressure variation across the flow cross-section section of a horizontal pipe:* In stratified flow the gravitation is a dominant force. The cross-section section averaged gas gravity pressure difference with respect to the interface is

$$\Delta p_1^{2\sigma} = \rho_1 g D_h \left( \frac{\sin^3 \theta}{3\pi\alpha_1} - \frac{1}{2} \cos \theta \right). \quad (2.137)$$

The cross-section averaged liquid gravity pressure difference with respect to the bottom of the pipe is

$$\Delta p_2^{1\sigma} = -\rho_2 g D_h \left( \frac{\sin^3 \theta}{3\pi\alpha_2} + \frac{1}{2} \cos \theta \right). \quad (2.138)$$

With respect to the interfacial pressure we have

$$p_1 = p_1^{2\sigma} - \Delta p_1^{2\sigma} = p_1^{2\sigma} - \rho_1 g D_h \left( \frac{\sin^3 \theta}{3\pi\alpha_1} - \frac{1}{2} \cos \theta \right), \quad (2.139)$$

$$p_2 = p_2^{1\sigma} - \Delta p_2^{1\sigma} = p_2^{1\sigma} + \rho_2 g D_h \left( \frac{\sin^3 \theta}{3\pi\alpha_2} + \frac{1}{2} \cos \theta \right), \quad (2.140)$$

which are, in fact, Eqs. (55) and (60) in *Ransom et al. (1987, p. 30)*. The averaged pressure in the cross-section section can be expressed as the cross-section section weighted averaged pressures inside the fields

$$p = \alpha_1 p_1 + \alpha_2 p_2 = \alpha_1 p_1^{2\sigma} + \alpha_2 p_2^{1\sigma} + gD_h \left[ \frac{\sin^3 \theta}{3\pi} (\rho_2 - \rho_1) + (\alpha_2 \rho_2 + \alpha_1 \rho_1) \frac{1}{2} \cos \theta \right]. \quad (2.141)$$

Neglecting the surface tension we obtain

$$p \approx p_\sigma + gD_h \left[ \frac{\sin^3 \theta}{3\pi} (\rho_2 - \rho_1) + (\alpha_2 \rho_2 + \alpha_1 \rho_1) \frac{1}{2} \cos \theta \right], \quad (2.142)$$

or

$$p_\sigma \approx p - gD_h \left[ \frac{\sin^3 \theta}{3\pi} (\rho_2 - \rho_1) + (\alpha_2 \rho_2 + \alpha_1 \rho_1) \frac{1}{2} \cos \theta \right]. \quad (2.143)$$

The averaged pressure in the gas phase and in the liquid phase can then be expressed as a function of the system's averaged pressure and geometrical characteristics:

$$p_1 = p - gD_h \left[ \frac{\sin^3 \theta}{3\pi\alpha_1} (\alpha_1 \rho_2 + \alpha_2 \rho_1) + \alpha_2 (\rho_2 - \rho_1) \frac{1}{2} \cos \theta \right] \quad \text{for } \alpha_1 > 0, \quad (2.144)$$

$$p_2 = p + gD_h \left[ \frac{\sin^3 \theta}{3\pi\alpha_2} (\alpha_1 \rho_2 + \alpha_2 \rho_1) + \alpha_1 (\rho_2 - \rho_1) \frac{1}{2} \cos \theta \right] \quad \text{for } \alpha_2 > 0. \quad (2.145)$$

The check  $\alpha_1 p_1 + \alpha_2 p_2 = p$  proves the correctness of the computation. The difference between both averaged pressures is then

$$\Delta p_{21} = p_2 - p_1 = gD_h \left[ \frac{\sin^3 \theta}{3\pi} \left( \frac{\rho_2}{\alpha_2} + \frac{\rho_1}{\alpha_1} \right) + (\rho_2 - \rho_1) \frac{1}{2} \cos \theta \right] \quad \text{for } \alpha_1 > 0 \text{ and } \alpha_2 > 0, \quad (2.146)$$

and, therefore, Eqs. (2.93)–(2.96) are valid also for a circular pipe. Assuming that the change of the densities contributes much less to the change of  $\Delta p_{21}$ , the change of the local volume fraction results in

$$\begin{aligned} \frac{\partial \Delta p_{21}}{\partial z} &\approx \frac{\partial \Delta p_{21}}{\partial \alpha_1} \frac{\partial \alpha_1}{\partial z} \\ &= gD_h \left[ \frac{\sin^3 \theta}{3\pi} \left( \frac{\rho_2}{\alpha_2^2} - \frac{\rho_1}{\alpha_1^2} \right) + \left( \frac{\rho_2}{\alpha_2} + \frac{\rho_1}{\alpha_1} \right) \frac{1}{2} \cos \theta - (\rho_2 - \rho_1) \frac{\pi}{4 \sin \theta} \right] \frac{\partial \alpha_1}{\partial z}. \end{aligned} \quad (2.147)$$

This equation was obtained by taking into account that  $\theta$  is also an implicit function of  $\alpha_1$  through Eq. (2.117) and using Eq. (2.118). Using Eqs. (2.156), (2.146), (2.137), and (2.138) and substituting into the momentum equations (2.100) and (2.101), we finally obtain

$$\dots + \gamma_z \left[ \alpha_1 \frac{\partial p}{\partial z} + \alpha_1 \alpha_2 (\rho_2 - \rho_1) \frac{\pi g D_h}{4 \sin \theta} \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_1^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (2.148)$$

$$\dots + \gamma_z \left[ \alpha_2 \frac{\partial p}{\partial z} - \alpha_1 \alpha_2 (\rho_2 - \rho_1) \frac{\pi g D_h}{4 \sin \theta} \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_2^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (2.149)$$

As a plausibility check note that for  $\alpha_1 \rightarrow 0$  the term containing the derivative of the volume fraction of velocity field 1 converges to zero and the momentum equation for field 2 takes the expected form. The sum of the two equations is then

$$\dots + \gamma_z \frac{\partial p}{\partial z} - (\Delta p_1^{w\sigma} + \Delta p_2^{w\sigma}) \frac{\partial \gamma_z}{\partial z} \dots \quad (2.150)$$

Comparing with the momentum equations for the rectangular channel we find instead of  $H$  the length  $\frac{\pi D_h}{4 \sin \theta} = \frac{\pi D_h^2}{4b}$ , which is the height of the rectangular channel having the same cross-section section as the pipe and a base equal to the gas-liquid median from Eq. (2.124). Therefore, the stability condition for stratified flow in a pipe is

$$(w_1 - w_2)^2 < g \frac{\pi D_h}{4 \sin \theta} (\rho_2 - \rho_1) \left( \frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2} \right), \quad (2.151a)$$

or in an alternative form

$$(w_1 - w_2)^2 < g (\rho_2 - \rho_1) \left( \frac{\alpha_1}{\rho_1} + \frac{\alpha_2}{\rho_2} \right) \frac{d\delta_{2F}}{d\alpha_2}, \quad (2.151b)$$

which is a generalized *Kelvin–Helmholtz* stability criterion valid for pipes with arbitrary cross-section section. Substituting Eq. (2.82) in the above equation we obtain Eq. (2.116) for rectangular channels. This result is, in fact, Eq. (26) in *Barnea and Taitel* (1994) for the inviscid case. It is identical with Eq. (6.9), p. 313, obtained by *de Crecy* in 1986.

Dividing the momentum equations by  $\alpha_1 \rho_1$  and  $\alpha_2 \rho_2$ , respectively, and subtracting the second from the first, we obtain

$$\dots + \gamma_z \left[ \left( \frac{1}{\rho_1} - \frac{1}{\rho_2} \right) \frac{\partial p}{\partial z} + \frac{\rho}{\rho_1 \rho_2} (\rho_2 - \rho_1) \frac{\pi g D_h}{4 \sin \theta} \frac{\partial \alpha_1}{\partial z} \right] - \left( \frac{\Delta p_1^{w\sigma}}{\alpha_1 \rho_1} - \frac{\Delta p_2^{w\sigma}}{\alpha_2 \rho_2} \right) \frac{\partial \gamma_z}{\partial z} \dots \quad (2.152)$$

The second term in the first brackets is exactly Eq. (58) obtained by *Ransom et al.* (1987, p. 30).

Note the general notation of the coefficients of  $\partial\alpha_1/\partial z$  in Eqs. (2.103) and (2.104) and Eqs. (2.148) and (2.149)

$$\alpha_1 \alpha_2 (\rho_2 - \rho_1) g \frac{F}{b} \frac{\partial\alpha_1}{\partial z},$$

where  $F$  is the channel cross-section section and  $b$  is the gas-liquid interface median. This result was obtained by *de Crecy* (1986, p. 312) in his Eq. (6.3).

*Teletov and Mamaev et al.*, see *Mamaev et al.* (1969), presented analytical solutions for stratified flow between two parallel plates and stratified flow in circular tubes, respectively. The gas and liquid phases are considered incompressible. No heat and mass transfer is considered. The pressure gradient is assumed constant and the flow is considered to be stationary and fully developed. The velocities at the wall are assumed to be zero and the velocity at the interface is assumed to be equal for both phases. The solution for the circular tube is found after introducing a bipolar coordinate transformation and integration, and is expressed as cross-section averaged velocities as a function of the flow parameter in the form

$$w_1 = \frac{\pi R^4}{8\eta_1} \left[ \frac{dp}{dz} - \rho_1 g \cos(\mathbf{g}, z) \right] \varphi_1(\alpha_1, \eta_2/\eta_1),$$

$$w_2 = \frac{\pi R^4}{8\eta_2} \left[ \frac{dp}{dz} - \rho_2 g \cos(\mathbf{g}, z) \right] \varphi_2(\alpha_1, \eta_2/\eta_1).$$

$\varphi_1(\alpha_1, \eta_2/\eta_1)$  and  $\varphi_2(\alpha_1, \eta_2/\eta_1)$  are complicated integral functions. For practical use they are presented in graphical form.

## 2.7.2 Dispersed interface

### 2.7.2.1 General

In this section, we provide a guide for derivation of a constitutive relation for mechanical interaction between a dispersed field  $l$  and the surrounding continuum  $m$ . An example for such flow is bubbly flow. In other words, we discuss a possible simplification of the surface integrals in Eq. (2.50). For a dispersed phase  $l$ , the viscous shear at the interface is negligible for non-*Stokes* flows:

$$-\frac{1}{Vol} \int_{F_{l\sigma}} \mathbf{T}_l^{m\sigma} \cdot \mathbf{n}_l dF \approx 0. \quad (2.153)$$

The viscous effects in the continuum at the interface are also neglected,

$$-\frac{1}{Vol} \int_{F_{m\sigma}} \mathbf{T}_m^{l\sigma} \cdot \mathbf{n}_m dF \approx 0. \quad (2.154)$$

Note that a 1 cm bubble in water having a relative velocity of 10 cm/s has a *Reynolds* number of about 100. For Reynolds numbers less than 24 the viscous effect in the continuum is important. For larger Reynolds numbers, which is often the case in nature, the viscous effects can be neglected. The force of the *Marangoni* effect,

$$\delta_l \frac{1}{Vol} \int_{F_{i\sigma}} (\nabla_i \sigma_{lm}) dF = 0, \quad (2.155)$$

can be neglected for the majority of macroscopic processes. If the dispersed phase is assumed to have no wall contact,  $F_{lw} = 0$ , the following results:

$$\frac{1}{Vol} \int_{F_{lw}} (\Delta p_l^{w\sigma} \mathbf{I} - \mathbf{T}_l^{w\sigma}) \cdot \mathbf{n}_l dF = 0. \quad (2.156)$$

The difference between the bubble bulk pressure and the bubble interface pressure is also negligible,

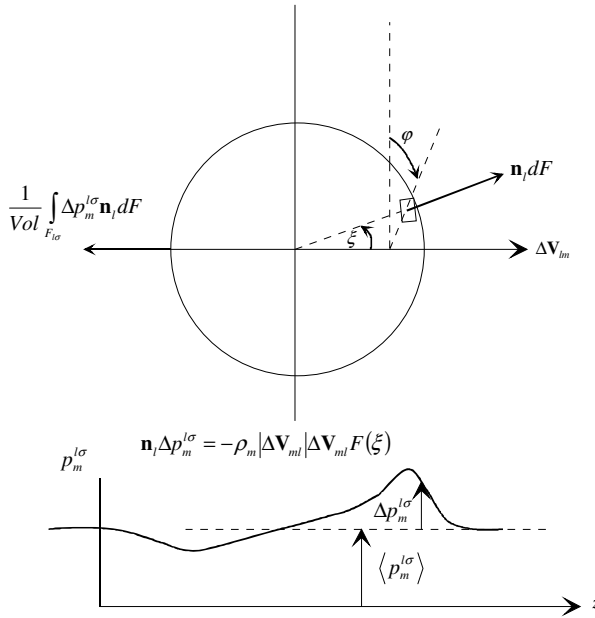
$$+ \frac{1}{Vol} \int_{F_{m\sigma}} \Delta p_l^{m\sigma} \mathbf{n}_m dF \approx 0. \quad (2.157)$$

The momentum equation for the dispersed field is

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l \mathbf{V}_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l \mathbf{V}_l \mathbf{V}_l \gamma) + \nabla \cdot [\alpha_l^e (\rho_l \overline{\mathbf{V}_l \mathbf{V}_l'}) \gamma] + \alpha_l^e \gamma \mathcal{N} p_l + \alpha_l \gamma_v \rho_l \mathbf{g} \\ & + (p_l - p_m + \delta_l \sigma_{lm} \kappa_l) \nabla (\alpha_l^e \gamma) + \frac{1}{Vol} \int_{F_{l\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l dF = \gamma_v \sum_{\substack{k=1 \\ k \neq l}}^{3,w} (\mu_{kl} \mathbf{V}_k - \mu_{lk} \mathbf{V}_l). \end{aligned} \quad (2.158)$$

The dispersed equation for the continuum (2.50b) is

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_m \rho_m \mathbf{V}_m \gamma_v) + \nabla \cdot (\alpha_m^e \rho_m \mathbf{V}_m \mathbf{V}_m \gamma) + \nabla \cdot [\alpha_m^e \gamma (\rho_m \overline{\mathbf{V}_m \mathbf{V}_m'} - \delta_m \mathbf{T}_m)] + \alpha_m \gamma_v \rho_m \mathbf{g} \\ & + \alpha_m^e \gamma \mathcal{N} p_m - \frac{1}{Vol} \int_{F_{m\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{mv}} (\Delta p_m^{w\sigma} \mathbf{I} - \mathbf{T}_m^{w\sigma}) \cdot \mathbf{n}_m dF \\ & = \gamma_v \sum_{\substack{k=1 \\ k \neq m}}^{3,w} (\mu_{km} \mathbf{V}_k - \mu_{mk} \mathbf{V}_m) \end{aligned} \quad (2.159)$$



**Fig. 2.10** Difference between the bulk pressure and interfacial pressure inside the velocity field  $m$  for steady state flow

The term in the momentum equation for the dispersed field that remains to be estimated is

$$\frac{1}{Vol} \int_{F_{l\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l dF .$$

For this, information is needed about the pressure distribution over the surface of a single particle, see Fig. 2.10. For illustration of the estimation of this integral we assume a family of monodisperse spheres. For dispersed flow with a small concentration of the dispersed phase, the flow about each sphere can be considered as unaffected by its neighbors. The interface pressure distribution along the surface can be generally represented by the following expression:

$$\Delta p_m^{l\sigma} \mathbf{n}_l = -\frac{1}{2} \rho_m R_l \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \cos \xi + \rho_m |\Delta \mathbf{V}_{ml}|^2 F(\xi), \quad (2.160)$$

see, for example, *Stuhmiller (1977)*. Here a spherical coordinate system is used with the main axis along  $\Delta \mathbf{V}_{lm}$ . The polar angle  $\xi$  is measured with respect to the direction of  $\Delta \mathbf{V}_{lm}$ . The azimuthal angle is  $\phi$ . The force per unit surface is split into a component parallel to  $\Delta \mathbf{V}_{lm}$  and a component perpendicular to  $\Delta \mathbf{V}_{lm}$ . The integration is then performed. Note that Eq. (2.160) does not depend on  $\phi$  and,

therefore, the perpendicular component for the symmetric body is  $\mathbf{0}$ . Note also that the pressure distribution in Eq. (2.160) does not take into account the spatial variation of the continuum velocity. The latter will give rise to a force component perpendicular to  $\Delta\mathbf{V}_{lm}$  even for a symmetric body. For the integration we need the following relations. The differential surface element of the sphere (rotational body) is  $dF = R_l^2 \sin \xi \, d\xi \, d\varphi$ . The projection of the interfacial pressure force on direction  $\Delta\mathbf{V}_{lm}$  for a single particle is, therefore,

$$\frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l \cos \xi \, dF = R_l^2 \int_0^{2\pi} \int_0^{\pi} \Delta p_m^{l\sigma} \mathbf{n}_l \cos \xi \sin \xi \, d\xi \, d\varphi, \quad (2.161)$$

and on the plane normal to  $\Delta\mathbf{V}_{lm}$

$$\frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l \sin \xi \, dF = R_l^2 \int_0^{2\pi} \int_0^{\pi} \Delta p_m^{l\sigma} \mathbf{n}_l \sin^2 \xi \, d\xi \, d\varphi. \quad (2.162)$$

For our case of a rotational body, the above integral gives  $\mathbf{0}$ . The collective force acting on the cloud of  $n_l \sum_{l=1}^3 Vol_l$  spheres in the control volume  $Vol$  per unit control volume in the axial direction is, therefore,

$$\frac{\sum_{l=1}^3 Vol_l}{Vol} n_l R_l^2 \int_0^{2\pi} \int_0^{\pi} \Delta p_m^{l\sigma} \mathbf{n}_l \cos \xi \sin \xi \, d\xi \, d\varphi = \gamma_v n_l R_l^2 \int_0^{2\pi} \int_0^{\pi} \Delta p_m^{l\sigma} \mathbf{n}_l \cos \xi \sin \xi \, d\xi \, d\varphi, \quad (2.163)$$

and in the plane normal to  $\Delta\mathbf{V}_{ml}$

$$\frac{\sum_{l=1}^3 Vol_l}{Vol} n_l R_l^2 \int_0^{2\pi} \int_0^{\pi} \Delta p_m^{l\sigma} \mathbf{n}_l \sin^2 \xi \, d\xi \, d\varphi = \gamma_v n_l R_l^2 \int_0^{2\pi} \int_0^{\pi} \Delta p_m^{l\sigma} \mathbf{n}_l \sin^2 \xi \, d\xi \, d\varphi. \quad (2.164)$$

Estimation of the integrals (2.163) and (2.164) provides a practical approach for computing the interfacial forces. What remains after the integration and some rearrangements given in the next section is

$$\begin{aligned} \frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l \, dF &= -\gamma_v \alpha_l \frac{1}{2} \rho_m \left[ \frac{\partial}{\partial \tau} \Delta\mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta\mathbf{V}_{ml} \right] \\ &+ 0.37 c_{ml}^d \rho_m |\Delta\mathbf{V}_{ml}|^2 \nabla (\alpha_l^e \gamma) - \gamma_v \alpha_l \rho_m \frac{1}{D_l} \frac{3}{4} c_{ml}^d |\Delta\mathbf{V}_{ml}| \Delta\mathbf{V}_{ml} \\ &= \gamma_v (\mathbf{f}_l^{vm} + \mathbf{f}_l^d) + \Delta p_m^{l\sigma*} \frac{1}{Vol} \int_{F_{i\sigma}} \mathbf{n}_l \, dF. \end{aligned} \quad (2.165)$$

This result for one-dimensional flow was obtained by *Stuhmiller* (1977). If we take the nonisotropy of the continuum velocity field into account in Eq. (2.160), we obtain the general form

$$\frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l dF = \gamma_v (\mathbf{f}_l^{vm} + \mathbf{f}_l^d + \mathbf{f}_l^L) + \Delta p_m^{l\sigma*} \frac{1}{Vol} \int_{F_i^{m\sigma}} \mathbf{n}_l dF. \quad (2.166)$$

The force components  $\mathbf{f}_l^{vm}$ ,  $\mathbf{f}_l^d$ ,  $\mathbf{f}_l^L$ , and  $\Delta p_m^{l\sigma*} \frac{1}{Vol} \int_{F_i^{m\sigma}} \mathbf{n}_l dF$  are called virtual mass

force, drag force, lift force, and stagnation pressure force, respectively. A detailed discussion of these is given below. Empirical information on how to compute these forces is given in Volume II.

The pressure distribution around a particle may influence the local mass transfer. In the case of strong thermodynamic nonequilibrium, the larger pressure difference across the interface at the stagnation point may lead to lower evaporation compared to the rear point. It may lead to a reactive resulting force at the droplet, which manifests itself as an effective drag reduction. A strong condensation may lead to the opposite effect. Although such arguments may sound reasonable, one should be careful because there is no accurate theoretical or experimental treatment of this problem.

### 2.7.2.2 Virtual mass force

Consider the integral defined by Eq. (2.161) taking over the first term of Eq. (2.160)

$$\begin{aligned} \gamma_v \mathbf{f}_l^{vm} &= -\gamma_v n_l \frac{1}{2} \rho_m \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right] R_l^3 \int_0^{2\pi} d\varphi \int_0^\pi \cos^2 \xi \sin \xi d\xi \\ &= -\gamma_v n_l \frac{4}{3} \pi R_l^3 \frac{1}{2} \rho_m \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \\ &= -\gamma_v \alpha_l \frac{1}{2} \rho_m \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right]. \end{aligned} \quad (2.167)$$

The force  $\mathbf{f}_l^{vm}$  is the *virtual mass force* per unit mixture volume. Here, the virtual mass coefficient is  $c_{ml}^{vm} = 1/2$ . The general form of the virtual mass force with the accuracy of an empirical coefficient was first proposed first *Prandtl* (1952), *Lamb* (1945), and *Milne-Thomson* (1968) in the same form

$$\mathbf{f}_d^{vm} = -\alpha_d \rho_c c_d^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{cd} + (\mathbf{V}_d \cdot \nabla) \Delta \mathbf{V}_{cd} \right], \quad (2.168)$$

where the subscripts *c* and *d* mean continuous and disperse, respectively. The scalar force components in *Cartesian* and in *cylindrical* coordinates are



$$f_{d,r}^{vm} = -\alpha_d \rho_c c_d^{vm} \left\{ \begin{array}{l} \frac{\partial}{\partial \tau} (u_c - u_d) + u_d \frac{\partial}{\partial r} (u_c - u_d) \\ + v_d \frac{1}{r^x} \frac{\partial}{\partial \theta} (u_c - u_d) + w_d \frac{\partial}{\partial z} (u_c - u_d) \end{array} \right\}, \quad (2.169)$$

$$f_{d,\theta}^{vm} = -\alpha_d \rho_c c_d^{vm} \left\{ \begin{array}{l} \frac{\partial}{\partial \tau} (v_c - v_d) + u_d \frac{\partial}{\partial r} (v_c - v_d) \\ + v_d \frac{1}{r^x} \frac{\partial}{\partial \theta} (v_c - v_d) + w_d \frac{\partial}{\partial z} (v_c - v_d) \end{array} \right\}, \quad (2.170)$$

$$f_{d,z}^{vm} = -\alpha_d \rho_c c_d^{vm} \left\{ \begin{array}{l} \frac{\partial}{\partial \tau} (w_c - w_d) + u_d \frac{\partial}{\partial r} (w_c - w_d) \\ + v_d \frac{1}{r^x} \frac{\partial}{\partial \theta} (w_c - w_d) + w_d \frac{\partial}{\partial z} (w_c - w_d) \end{array} \right\}. \quad (2.171)$$

The virtual mass force is experienced by the body as if it were to have an additional mass during its translation relative to the continuum. This explains the other name used for this force, *added mass force*. For larger particle concentrations,  $c_d^{vm}$  is a function of  $\alpha_r$ . Expressions for practical computation of the virtual mass coefficient for dispersed fields with larger concentration can be found in *Biesheuvel and van Wijngaarden (1984)*, *Biesheuvel and Spollstra (1989)*, *Cook and Harlow (1983, 1984)*, *Ishii and Michima (1984)*, *Lahey (1991)*, *Lamb (1945)*, *Milne-Thomson (1968)*, *Mokeyev (1977)*, *No and Kazimi (1985)*, *Prandtl (1952)*, *Ruggles et al. (1988)*, *van Wijngaarden (1976)*, *Winatabe et al. (1990)*, *Wallis (1969)*, and *Zuber (1964)*. These references represent the state of the art in this field. *Winatabe et al. (1990)* proposed the use of only the transient part of the virtual mass force for the case of strong transients.

*Lamb (1945)* computed the virtual mass coefficient for partials in potential flow with ellipsoidal shape defined by

$$\frac{x^2}{R_x} + \frac{y^2}{R_y} + \frac{z^2}{R_z} = 1,$$

where the lengths of the principal axis are  $R_x$ ,  $R_y$  and  $R_z$ , and the relative velocity is parallel to the x-axis as follows:

$$c_{cd}^{vm} = \frac{a_0}{2 - a_0},$$

$$a_0 = R_x R_y R_z \int_0^{\infty} \frac{d\lambda}{(R_x^2 + \lambda) \sqrt{(R_x^2 + \lambda)(R_y^2 + \lambda)(R_z^2 + \lambda)}}.$$

*Bournaski* (1992) evaluated some values as given in Table 2.1.

**Table 2.1.** Virtual mass coefficients for an ellipsoid

Shape of particles	Translation parallel to axis		
	$x$ $c_{cd,x}^{vm}$	$y$ $c_{cd,y}^{vm}$	$z$ $c_{cd,z}^{vm}$
$R_x = R_y = R_z$ , sphere	1/2	1/2	1/2
$R_x = R_y = R_z/2$ , rotary ellipsoid	0.704	0.704	0.210
$R_x = R_y = R_z/3$ , rotary ellipsoid	0.803	0.803	0.122
$R_x = R_y = R_z/4$ , rotary ellipsoid	0.859	0.859	0.081
$R_x = (2/3)R_y = R_z/2$ , unrotary ellipsoid	0.936	0.439	0.268
$R_x = R_y/2 = R_z/4$ , unrotary ellipsoid	1.516	0.398	0.126

For a single ellipsoid bubble with axis aspect ratio  $\chi$

$$c_{21}^{vm} = \frac{(\chi^2 - 1)^{1/2} - \cos^{-1} \chi^{-1}}{\cos^{-1} \chi^{-1} - (\chi^2 - 1)^{1/2} / \chi^2},$$

*van Wijngaarden* (1998).

*Lance* and *Bataille* (1991) reported experiments showing that for a 5mm deforming bubble the virtual mass coefficient is in the region:  $1.2 < c_{21}^{vm} < 3.4$ . For a family of spherical bubbles:

$$c_{21}^{vm} = \frac{1}{2}(1 + 2.78\alpha_1),$$

dilute bubble dispersion, interaction between two equally sized bubbles, *van Wijngaarden* (1976);

$$c_{21}^{vm} = \frac{1}{2} \frac{1 + 2\alpha_1}{1 - \alpha_1},$$

no interaction with the neighboring bubbles, *Zuber* (1964);

$$c_{21}^{vm} = \frac{1}{2}(1 + 3\alpha_1), \quad \alpha \rightarrow 0,$$

*Zuber* (1964);

$$c_{21}^{vm} = \frac{1}{2}(1 + 3.32\alpha_1),$$

analogously to thermal conductivity in composite material, *Jeffrey* (1973);

$$c_{21}^{vm} = \frac{1}{2} \left[ \frac{1}{2} - 1.98 \ln(0.62 - \alpha_1) \right],$$

the approximation for  $\alpha_1 \leq 0.35$ , *Biesheuvel* and *Spoelstra* (1989). *Laurien* and *Niemann* (2004) used direct numerical simulation and come to

$$c_{21}^{vm} = 0.5 + 1.63\alpha_1 + 3.85\alpha_1^2.$$

*Kendoush* (2006) considered the separation of the velocity profile around a single sphere and obtained a virtual mass coefficient depending on the separation angle. *Pougatch*, *Salcudean*, *Chan*, and *Knapper* (2008) concluded that the virtual mass force cannot be larger than the inertia force for accelerating the remaining liquid and, therefore, the coefficient is naturally limited by

$$c_{21}^{vm} \leq \alpha_2 / \alpha_1.$$

### 2.7.2.3 Form drag and stagnation pressure force

For inviscid (ideal) potential flow we have in Eq. (2.160)

$$F(\xi) = \frac{1}{8}(9 \cos^2 \xi - 5), \quad (2.172)$$

see *Lamb* (1945). This profile does not give any resulting force component (*d'Alembert's paradox*).

In nature  $F(\xi)$  gives a nonsymmetric profile, as indicated in Fig. 2.10. For an idealized nonsymmetric fore-aft profile, *Nigmatulin* (1979) estimated the integrals of Eqs. (2.160) and (2.161) for the second term of the right-hand side of Eq. (2.159) for bubbles in bubble-liquid flows. *Biesheuvel* and *van Wijngaarden* (1984) analytically computed the coefficients for spherical bubbles for *Nigmatulin's* derivation.

A more general approach was proposed by *Hwang* and *Schen* (1992). For the general case,  $F(\xi)$  is determined experimentally, see *Schlichting* (1959, p.21, Fig. 1.11). *Hwang* and *Schen* (1992) provided a method for computing the pressure distribution around a sphere for Reynolds numbers greater than 3000, where

$$F(\xi) = \frac{1}{4} \left\{ \frac{2\lambda^2 - 1}{\lambda^2 + 4} + \frac{9}{2} \frac{e^{-\lambda\xi}}{\lambda^2 + 4} \left[ 2\lambda \sin(2\xi) + \cos(2\xi) \right] \right\}, \quad (2.173)$$

see Eq. (12) in *Hwang and Schen (1992)*. When the drag coefficient  $c_{ml}^d$  is known and the equation

$$c_{ml}^d = 45 \frac{1 - e^{-\pi\lambda}}{(\lambda^2 + 4)(\lambda^2 + 16)} \quad (2.174)$$

is solved for the smaller real root of  $\lambda$ , there is a method for estimating the integrals analytically. The reader can find the final result of this derivation, characterized by anisotropic forces in *Hwang and Schen (1992)*.

Before *Hwang and Schen* published their work, *Stuhmiller (1977)* had already rewritten the term  $F(\xi)$  as follows:

$$F(\xi) = \langle F(\xi) \rangle^{l\sigma} + F(\xi) - \langle F(\xi) \rangle^{l\sigma}, \quad (2.175)$$

where  $\langle F(\xi) \rangle^{l\sigma}$  represents the surface average over a single sphere. For a *turbulent pressure distribution* around a sphere as given by *Schlichting (1959)*, the interface average of the function  $F(\xi)$  is

$$\langle F(\xi) \rangle^{l\sigma} = -0.37c_{ml}^d, \quad (2.176)$$

where  $c_{ml}^d$  is the form drag coefficient for single particles. In the literature sometimes  $\langle F(\xi) \rangle^{l\sigma}$  is set to the constant value  $1/4$  – see, for instance, *Lamb (1945)*. This means that the integral over the second term of Eq. (2.160) can be split into two parts. The first part can be estimated directly exploiting the fact that  $\langle F(\xi) \rangle^{l\sigma}$  does not depend on the position at the interface. The result is

$$\begin{aligned} \rho_m |\Delta \mathbf{V}_{ml}|^2 \int_{F_{l\sigma}} \langle F(\xi) \rangle^{l\sigma} \mathbf{n}_l dF &= -\rho_m |\Delta \mathbf{V}_{ml}|^2 \langle F(\xi) \rangle^{l\sigma} \left[ \nabla(\alpha_l^e \gamma) + \frac{1}{Vol} \int_{F_{lw}} \mathbf{n}_l dF \right] \\ &= 0.37c_{ml}^d \rho_m |\Delta \mathbf{V}_{ml}|^2 \left[ \nabla(\alpha_l^e \gamma) + \frac{1}{Vol} \int_{F_{lw}} \mathbf{n}_l dF \right] = 0.37c_{ml}^d \rho_m |\Delta \mathbf{V}_{ml}|^2 \nabla(\alpha_l^e \gamma), \end{aligned} \quad (2.177)$$

which results in an effective stagnation pressure difference

$$\Delta p_m^{l\sigma^*} = -0.37c_{ml}^d \rho_m |\Delta \mathbf{V}_{ml}|^2 \quad (2.178)$$

similar to that discussed for stratified flow (Eq. (2.64)). To derive Eq. (2.177), Eq. (29) from *Kolev (1994b)* is also used together with the fact that there is no contact between the dispersed field and the wall,  $F_{lw} = 0$ .

The second part is the net force experienced by the particle due to nonuniform pressure distribution around the particle, the so-called form drag force:

$$\begin{aligned}
\gamma_v \mathbf{f}_l^d &= -\gamma_v n_l \rho_m |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml} R_l^2 \int_0^{2\pi} d\varphi \int_0^\pi \left[ F(\xi) - \langle F(\xi) \rangle^{l\sigma} \right] \cos \xi \sin \xi d\xi \\
&= -\gamma_v n_l \rho_m |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml} c_{ml}^d \frac{1}{2} \pi R_l^2 = -\gamma_v \left[ \alpha_l / \left( \frac{4}{3} \pi R_l^3 \right) \right] \rho_m |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml} c_{ml}^d \frac{1}{2} \pi R_l^2 \\
&= -\gamma_v \alpha_l \rho_m \frac{1}{D_l} \frac{3}{4} c_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml}. \tag{2.179}
\end{aligned}$$

The form drag force per unit mixture volume is, therefore,

$$\mathbf{f}_l^d = -\alpha_l \rho_m \frac{1}{D_l} \frac{3}{4} c_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml}. \tag{2.180}$$

For larger volume fractions  $\alpha_l$  one should take into account the dependence of the drag force on the volume fraction – also see *Ishii and Mishima* (1984) and *Zuber* (1964), for example. For drag forces in two-phase flows, the study by *Ishii and Mishima* (1984) is recommended.

#### 2.7.2.4 Lift force

**Note on particle rotation:** A rotating sphere obeys the law of conservation of momentum

$$I_d \frac{d\omega_d}{d\tau} = -C_d^\omega \left( \frac{D_d}{2} \right)^5 \frac{1}{2} \rho_d |\omega_d| \omega_d.$$

Here, the particle rotation velocity is  $\omega_d$  and  $I_d$  is the particle's moment of inertia.

$$C_d^\omega = \frac{c_1}{(\text{Re}_{cd}^\omega)^{1/2}} + \frac{c_2}{\text{Re}_{cd}^\omega} + c_3 \text{Re}_{cd}^\omega$$

is a coefficient depending on the rotational *Reynolds* number

$$\text{Re}_{cd}^\omega = \left( \frac{D_d}{2} \right) |\omega_d| / \nu_c.$$

The  $c$  coefficients are given by *Yamamoto et al.* (2001) in the following table:

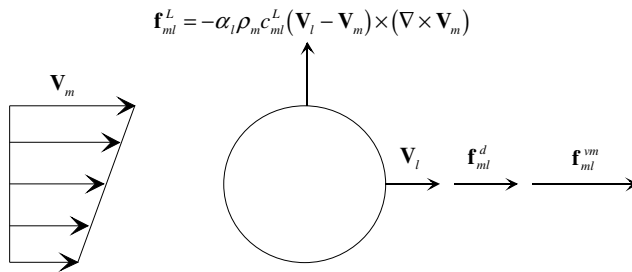
$\text{Re}_{cd}^\omega$	0 to 1	1 to 10	10 to 20	20 to 50	>50
$c_1$	0	0	5.32	6.44	6.45
$c_2$	50.27	50.27	37.2	32.2	32.1
$c_3$	0	0.0418	5.32	6.44	6.45

We learn from this dependence that small and light particles can be more easily put in rotation compared to heavy and large particles. The following three main idealizations give an idea of the origin of the so-called lift force:

- a) A rotating symmetric particle in symmetric flow of continuum experiences a lift force called the *Magnus* force (named after the Berlin physicist *Gustav Magnus*, 1802–1870). The curiosity of Lord *Rayleigh* in regard to the trajectory of the tennis ball led him in 1877 to the corresponding explanation. The force was analytically estimated by *Jukowski* and independently by *Kutta*, see *Albring* (1970, p. 75).
- b) A nonrotating symmetric particle in nonsymmetric continuum flow as presented in Fig. 2.11 experiences lift force, *Jukowski*.
- c) A nonrotating asymmetric particle in symmetric continuum flow experiences lift force, *Jukowski*.

The force component perpendicular to the relative velocity direction is called the lateral or *lift force*. The lift force is zero for symmetric bodies exposed to symmetrical flow

$$\mathbf{f}_l^L = 0. \tag{2.181}$$



**Fig. 2.11** Drag, virtual mass, and lift forces acting simultaneously on the field *l*

A symmetric body exposed to asymmetrical flow experiences a lateral force – see Fig. 2.11. The lift force is similar in nature to the aerodynamic lift of an airfoil, but differs in that it is a result of the gradient in the continuum velocity field over a symmetric body rather than a uniform flow over an asymmetric airfoil. The general form of the lateral lift force for inviscid flows is given by *Drew* and *Lahey* (1987)

$$\mathbf{f}_{cd}^L = -\alpha_d \rho_c c_{cd}^L (\mathbf{V}_d - \mathbf{V}_c) \times (\nabla \times \mathbf{V}_c). \tag{2.182}$$

The scalar components for *Cartesian* and *cylindrical* coordinates are

$$f_{cd,r}^L = -\alpha_d \rho_c c_{cd}^L \left\{ (v_d - v_c) \left[ \frac{1}{r^\kappa} \frac{\partial}{\partial r} (r^\kappa v_c) - \frac{1}{r^\kappa} \frac{\partial u_c}{\partial \theta} \right] - (w_d - w_c) \left( \frac{\partial u_c}{\partial z} - \frac{\partial w_c}{\partial r} \right) \right\}, \tag{2.183}$$

$$f_{cd,\theta}^L = -\alpha_d \rho_c c_{cd}^L \left\{ (w_d - w_c) \left[ \frac{1}{r^\kappa} \frac{\partial w_c}{\partial \theta} - \frac{\partial v_c}{\partial z} \right] - (u_d - u_c) \left( \frac{1}{r^\kappa} \frac{\partial}{\partial r} (r^\kappa v_c) - \frac{1}{r^\kappa} \frac{\partial u_c}{\partial \theta} \right) \right\}, \quad (2.184)$$

$$f_{cd,z}^L = -\alpha_d \rho_c c_{cd}^L \left\{ (u_d - u_c) \left[ \frac{\partial u_c}{\partial z} - \frac{\partial w_c}{\partial r} \right] - (v_d - v_c) \left( \frac{1}{r^\kappa} \frac{\partial w_c}{\partial \theta} - \frac{\partial v_c}{\partial z} \right) \right\}. \quad (2.185)$$

The lift coefficient must be derived experimentally. The reader will find information on modeling of the lift force in *Deich and Philipoff* (1981), *Staffman* (1965), *Bernemann et al.* (1991), *Soo and Tung* (1972), *Ho and Leal* (1976), *Vasseur and Cox* (1976), *Drew and Lahey* (1987, 1990), *Erichhorn and Small* (1969), and *Bataille et al.* (1990).

For negligible particle rotation, *Staffman* (1965, 1968) derived a negligible particle *Reynolds* number and small gradients of the continuum velocity the analytical expression for the shear lift force

$$c_{21}^L = 3.084 v_2^{1/2} \left/ \left( D_1 \left| \frac{dw_2}{dr} \right|^{1/2} \right) \right.$$

Inside the boundary layer of bubbly flow having  $w_1 > w_2$  and  $\partial w_2 / \partial r < 0$ , the lift force pushes the bubbles towards the wall. Note that the spatial resolution in discrete analyses must be fine enough in order to accurately compute the rotation of the continuous velocity field. Bad resolution, such in so-called subchannel analyses, produces only useless noise that makes the use of this force meaningless.

*Mei* (1992) proposed an expression that can be used for larger particle *Reynolds* numbers

$$c_{21}^L = Mei \ 3.084 v_2^{1/2} \left/ \left( D_1 \left| \frac{dw_2}{dr} \right|^{1/2} \right) \right.,$$

where

$$Mei = (1 - 0.3314 \omega_2^{1/2}) \exp(-0.1 Re_{12}) + 0.3314 \omega_2^{1/2}, \quad Re_{12} \leq 40,$$

$$Mei = 0.0524 (\omega_2 Re_{12})^{1/2}, \quad Re_{12} > 40,$$

and  $Re_{12} = \Delta w_{12} D_1 / \nu_2$ ,  $\omega_2 = \frac{D_1/2}{w_2 - w_1} \left| \frac{dw_2}{dr} \right|$ . In a later work, *Klausner et al.* (1993)

found that the lift force on a bubble attached to a wall can be computed using

$$c_{21}^L = \frac{16}{3} 3.877 \omega_2^{3/2} (0.014 \beta^2 + Re_{12}^{-2})^{1/4},$$

which is valid for larger Reynolds numbers than the previous relation. In a later work *Mei and Klausner* (1995) proposed to use interpolation between the *Stafman's* results for small Reynolds numbers and *Auton's* (1987) results for large Reynolds numbers:

$$c_{21}^L = \frac{3}{8\omega_2^{1/2}} \left\{ \frac{16}{9} \omega_2 + \left[ \frac{1.72J\sqrt{2\omega_2/\text{Re}_{12}}}{\text{Re}_{12}^{1/2}} \right]^2 \right\}^{1/2},$$

$$J = 0.6765 \left\{ 1 + \tanh \left[ 2.5 \left( \log_{10} \sqrt{2\beta/\text{Re}_{12}} + 0.191 \right) \right] \right\}$$

$$\times \left\{ 0.667 + \tanh \left[ 6 \left( \log_{10} \sqrt{2\beta/\text{Re}_{12}} - 0.32 \right) \right] \right\}$$

There are other expressions for the lift force on a single bubble. *Tomiya et al.* (2002) measured trajectories of single bubbles in simple shear flows of glycerol–water solution. They obtained the following empirical correlation:

$$c_{21}^L = \min \left[ 0.288 \tanh (0.121 \text{Re}_{12}), f(E\ddot{\omega}_{1m}) \right] \quad \text{for } E\ddot{\omega}_{1m} < 4,$$

$$c_{21}^L = f(E\ddot{\omega}_{1m}) = 0.00105E\ddot{\omega}_{1m}^3 - 0.0159E\ddot{\omega}_{1m}^2 - 0.0204E\ddot{\omega}_{1m} + 0.474$$

$$\text{for } 4 \leq E\ddot{\omega}_{1m} \leq 10.7,$$

$$c_{21}^L = -0.29 \quad \text{for } 10.7 < E\ddot{\omega}_{1m},$$

based on experiments within the region of parameters defined by  $1.39 \leq E\ddot{\omega}_{1m} \leq 5.74$ ,  $-5.5 \leq \log_{10} Mo_{12} \leq -2.8$ , and  $0 < |\nabla \times \mathbf{V}_2| \leq 8.3s^{-1}$ . The lift coefficient varied in this region between about 0.3 and  $-0.3$ . Here a modified *Eötvös* and *Morton* numbers are computed using the horizontal bubble size

$$E\ddot{\omega}_{1m} = g(\rho_2 - \rho_1)D_{1,\max}^2 / \sigma_{12},$$

$$Mo_{12} = g(\rho_2 - \rho_1)\eta_2^4 / (\rho_2^2 \sigma_{12}).$$

The aspect ratio of the bubble is computed by using the *Wellek et al.* (1966) correlation

$$D_{1,\max} / D_{1,\min} = 1 + 0.163E\ddot{\omega}_{1m}^{0.757}.$$

In accordance with the *Tomiya et al.* correlation, the lift coefficient for a bubble with a diameter of 3 mm in an air–water system is equal to 0.288. *Zun* (1980) performed measurements and estimated a value for small bubbles of about 0.3. *Naciri et al.* (1992) experimentally measured the lift coefficient of a bubble in a vortex to be 0.25.



It should be emphasized that the above reviewed considerations are for a single object in shear flow. The presence of multiple objects in share flow is found to influence this force too.

The importance of the findings by *Tomiyama et al. (2002)* is in the observation that for large bubbles the lift force changes sign. *Krepper et al. (2005)* observed experimentally that in vertical bubbly flow the void profile depends on the bubble size spectrum. For a spectrum with predominantly small sized bubbles a wall void peaking is observed. The level of the wall peaking depends on the turbulence in the liquid and on the stagnation pressure force. For spectra having predominantly large bubbles the central void, peaking is observed. This effect was reproduced by *Krepper et al. (2005)* by using lift force applied on multiple groups with  $c_{21}^L = 0.05$  for  $D_1 < 0.006m$  and  $c_{21}^L = -0.05$  for  $D_1 \geq 0.006m$ . The improvement going from 1 to 2 groups was considerable. No substantial change was reported if more then 8 size groups were used.

Using a combination of the radial liquid and gas momentum equations

$$p(r) = p(R) - (1 - \alpha_1) \rho_2 \overline{u^{r^2}} - \int_R^r \frac{(1 - \alpha_1) \rho_2 (\overline{u^{r^2}} - \overline{v^{r^2}})}{r^*} dr^* .$$

and the measured fluctuation velocities in the radial and in the azimuthal direction, *Wang et al. (1987)* explained why the bubble peaking for upwards flows is observed close to the wall. Later it was found that this is valid for bubbles with small sizes. Close to the wall the authors observed that a) the velocity gradient has a maximal, b) the velocity fluctuations have maximum, and c) the static pressure has a minimum. Using the radial momentum equations for gas and liquids

$$\begin{aligned} -\alpha_1 \frac{dp}{dr} + \frac{1}{r} \frac{d}{dr} (\alpha_1 r \tau_{1,rr}) - \frac{1}{r} \alpha_1 \tau_{1,\theta\theta} + f_{21}^L &= 0 \\ -(1 - \alpha_1) \frac{dp}{dr} + \frac{1}{r} \frac{d}{dr} [r(1 - \alpha_1) \tau_{2,rr}] - \frac{1}{r} (1 - \alpha_1) \tau_{2,\theta\theta} - f_{21}^L &= 0, \end{aligned}$$

it is possible to estimate the radial pressure distribution and the lift force, knowing from measurements the void and the velocity profiles with their fluctuations  $\tau_{1,rr} \approx 0$ ,  $\tau_{1,\theta\theta} \approx 0$ ,  $\tau_{2,rr} = -\rho_2 \overline{u_2^{r^2}}$ ,  $\tau_{2,\theta\theta} = -\rho_2 \overline{v_2^{r^2}}$ . This is the approach used by *Wang et al.* to gain expression for the lift force in bubbly flow based on groups of variables that come from the theory of the lift force on a single object. *Wang et al.* introduced the influence of the local volume fraction into the lift coefficient

$$c_{21}^L(\xi) = 0.01 + \frac{0.49}{\pi} \cot^{-1} \frac{\log \xi + 9.3168}{0.1963} ,$$

as a function of

$$\xi = \exp(-\alpha_1) 2\omega_2 \left( \frac{D_1}{D_{hyd}} \frac{1}{Re_{12}} \right)^2 \left( \frac{w_1}{\Delta w_{12\infty}} \right)^2,$$

where  $\Delta w_{12\infty} = 1.18(g\sigma/\rho_2)^{1/4}$ . This coefficient varies within 0.01 and 0.1 in accordance to Wang's et al. data. The disadvantage of this approach is that due to the dependence  $\xi = \xi(D_{hyd})$  the correlation depends on one global geometry characteristic and cannot be applied locally.

For a down flow of buoyant bubbles the lift-force is directed toward the center of the pipe. As a result, no wall peaking of the void fraction is experimentally observed in turbulent bubbly flow. The level of the wall peaking depends of the turbulence in the liquid and on the stagnation pressure force.

**Conclusions:** (a) The spatial resolution in finite volume analyses must be fine enough in order to accurately compute the rotation of the continuous velocity field. Bad resolution such as in the so-called subchannel analyses produces only useless noise that makes the use of this force meaningless. (b) There is no method known to me that is based on local conditions and that allows taking into account the effect of multiple objects on the lift force. (c) The other problem is that small bubbles will probably rotate and the application of lift force derived for nonrotating objects in shear flows is questionable. (d) Heavy solid particles carried by gas are rather subject to lift force than to Magnus force because due to their inertia they will hardly take the rotation of the surrounding continuum.

### 2.7.2.5 Interfacial structure forces

The continuous field interacts with the wall structures. Careful estimation of the surface integral

$$\frac{1}{Vol} \int_{F_{mw}} \Delta p_m^{w\sigma} \mathbf{n}_m dF$$

is required, especially in the case of variable geometry of the structure in space. The discussion of flow on immersed bodies given in Sections 2.6.2.2 through 2.6.2.4 is also valid for the case of a porous solid structure with a characteristic size of  $D_w$ . This means that the stagnation pressure force, form drag, virtual mass force, and lift force must also be incorporated,

$$\frac{1}{Vol} \int_{F_{mw}} (\Delta p_m^{w\sigma} \mathbf{I} - \mathbf{T}_m^{w\sigma}) \cdot \mathbf{n}_m dF = \gamma_v (f_{vm}^d + f_{vm}^{vm} + f_{vm}^L) + \Delta p_m^{w\sigma*} \frac{1}{Vol} \int_{F_{mw}} \mathbf{n}_m dF. \quad (2.186)$$

For a continuous velocity field wetting the total structure,  $F_{mw} = F_w$  and, therefore,

$$\frac{1}{Vol} \int_{F_{mv}} \mathbf{n}_m dF = \frac{1}{Vol} \int_{F_w} \mathbf{n}_m dF = \nabla(1-\gamma) = -\nabla\gamma. \quad (2.187)$$

which is in fact Eq. (1.29) with  $\mathbf{n}_l = -\mathbf{n}_w$  and  $\alpha_l^e = 1$ . Consequently

$$\frac{1}{Vol} \int_{F_{mv}} (\Delta p_m^{w\sigma} \mathbf{I} - \mathbf{T}_m^{w\sigma}) \cdot \mathbf{n}_m dF = \gamma_v (f_{wm}^d + f_{wm}^{vm} + f_{wm}^L) - \Delta p_m^{w\sigma} \nabla\gamma. \quad (2.188)$$

Here the following applies for a disperse structure (flow through porous media):

$$\mathbf{f}_{wm}^{vm} = \rho_m c_w^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{mw} + (\mathbf{V}_w \cdot \nabla) \Delta \mathbf{V}_{mw} \right] = \overline{c_w^{vm}} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{mw} + (\mathbf{V}_w \cdot \nabla) \Delta \mathbf{V}_{mw} \right], \quad (2.189)$$

$$\mathbf{f}_{wm}^d = \rho_m \frac{1}{D_w} \frac{3}{4} c_{mw}^d |\Delta \mathbf{V}_{mw}| \Delta \mathbf{V}_{mw} = \overline{c_{mw}^d} |\Delta \mathbf{V}_{mw}| \Delta \mathbf{V}_{mw}, \quad (2.190)$$

$$\mathbf{f}_{wm}^L = \rho_m c_{mw}^L (\mathbf{V}_w - \mathbf{V}_m) \times (\nabla \times \mathbf{V}_m) = \overline{c_{mw}^L} (\mathbf{V}_w - \mathbf{V}_m) \times (\nabla \times \mathbf{V}_m). \quad (2.191)$$

For the case of a wall at rest we have

$$\mathbf{f}_{wm}^{vm} = \rho_m c_w^{vm} \frac{\partial \mathbf{V}_m}{\partial \tau} = \overline{c_w^{vm}} \frac{\partial \mathbf{V}_m}{\partial \tau}, \quad (2.192)$$

$$\mathbf{f}_{wm}^d = \rho_m \frac{1}{D_w} \frac{3}{4} c_{mw}^d |\mathbf{V}_m| \mathbf{V}_m = \overline{c_{mw}^d} |\mathbf{V}_m| \mathbf{V}_m, \quad (2.193)$$

$$\mathbf{f}_{wm}^L = -\rho_m c_{mw}^L \mathbf{V}_m \times (\nabla \times \mathbf{V}_m) = -\overline{c_{mw}^L} \mathbf{V}_m \times (\nabla \times \mathbf{V}_m). \quad (2.194)$$

The shear (friction) force for channels is usually incorporated into  $\overline{c_{mw}^d}$ . The same is performed for the drag resulting from local changes in the flow cross-section for the specific flow direction.

### 2.7.2.6 Force in the wall boundary layer

Note that no bubbles are observed at the wall for adiabatic flows. This led *Antal et al.* (1991) to the conclusion that there is a special force at the wall similar to the lubrication force that pushes the bubbles away from the surface,

$$\mathbf{f}_{cd}^{Lw} = \frac{\alpha_d \rho_c |\hat{\mathbf{V}}|^2}{R_d} \left( -0.104 - 0.06 |\Delta V_{cd}| + 0.147 \frac{R_d}{y_0} \right) \mathbf{n}_w,$$

where  $y_0$  is the distance between the bubble and the wall,  $\mathbf{n}_w$  is the unit outward normal vector on the surface of the wall, and

$$\hat{\mathbf{V}} = \mathbf{V}_d - \mathbf{V}_c - [\mathbf{n}_w \cdot (\mathbf{V}_d - \mathbf{V}_c)] \mathbf{n}_w.$$

### 2.7.2.7 Force causing turbulent diffusion

It is experimentally observed that turbulence in the continuous phase tends to smooth the volumetric concentrations of the dispersed phase. In other words, the pulsation in the continuum producing the force  $\nabla \cdot [\alpha_d^e (\rho_d \overline{\mathbf{V}'_d \mathbf{V}'_d}) \gamma]$  forces the particles to move from the places with higher concentration to the places with lower concentration. For homogeneous turbulence

$$\nabla \cdot [\alpha_d^e (\rho_d \overline{\mathbf{V}'_d \mathbf{V}'_d}) \gamma] = \nabla [\alpha_d^e (\rho_d 2k_d) \gamma].$$

For bubbly flow  $\nabla [\alpha_d^e (\rho_d 2k_d) \gamma] \approx \nabla [\alpha_d^e (\rho_d 2k_c) \gamma]$ . Here,  $k_c$  is the specific turbulent kinetic energy of the continuous phase. For bubbly flow, *Lopez de Bertodano* (1992) proposed the following form of this force:

$$\mathbf{f}_{cd}^t = \nabla \cdot [\alpha_d^e (\rho_d \overline{\mathbf{V}'_d \mathbf{V}'_d}) \gamma] \approx -c_{cd}^t \rho_c k_c \nabla \alpha_d,$$

with  $c_{cd}^t = 0.1$  proposed by *Lahey et al.* (1993). *Shi et al.* (2005) performed a *Favre* (mass-weighted) averaging of the classical form of the drag form. His final expression for the dispersed form is

$$\mathbf{f}_{cd}^t = -\frac{3}{4} \frac{c_{cd}^d}{D_d} \frac{v_c'}{Sh_c^t} \rho_c |\mathbf{V}_c - \mathbf{V}_d| \nabla \alpha_d,$$

where the turbulent *Schmidt* number for the continuous field,  $Sh_c^t$ , is set to 1.

### 2.7.2.8 Force causing rejection of droplet deposition at the wall

Consider very strong evaporation of a film. The deposition mass flow rate of droplets is  $(\rho w)_{32}$  with a velocity  $w_{32} = (\rho w)_{32} / \rho_3$ , which is perpendicular to the wall. The film evaporation emits a vapor with velocity that has a component opposing the droplet deposition velocity  $w_{1, \text{evaporation}} = \dot{q}'' / \rho_1 \Delta h$ . Therefore, the droplet in the proximity to the wall experiences an additional drag force that opposes its movement towards the wall

$$\mathbf{f}_{cd}^{dw} = \rho_c \frac{1}{D_d} \frac{3}{4} c_{cd}^{dw} |w_{1, \text{evaporation}} - w_{32}| (w_{1, \text{evaporation}} - w_{32}) \mathbf{n}_w.$$

As we can see this force is important a) for high pressure because the continuum density is high, b) for high heat fluxes, and c) for low turbulence in the vapor phase.

## 2.8 Working form for the dispersed and continuous phase

Thus the momentum equation of a dispersed velocity field takes the form

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_l \rho_l \mathbf{V}_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l \mathbf{V}_l \mathbf{V}_l \gamma) + \nabla \cdot \left[ \alpha_l^e (\rho_l \overline{\mathbf{V}_l \mathbf{V}_l'}) \gamma \right] + \alpha_l^e \gamma \overline{\mathcal{N}} p_l + \alpha_l \gamma_v \rho_l \mathbf{g} \\
 & + (p_l - p_m + \delta_l \sigma_{lm} \kappa_l) \nabla (\alpha_l^e \gamma) - \Delta p_m^{l\sigma^*} \nabla (\alpha_l^e \gamma) \\
 & - \gamma_v \alpha_l \rho_m \left\{ \begin{array}{l} c_{ml}^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \\ -c_{ml}^L \Delta \mathbf{V}_{ml} \times (\nabla \times \mathbf{V}_m) + \frac{1}{D_l} \frac{3}{4} c_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml} \end{array} \right\} \\
 & = \gamma_v \sum_{\substack{k=1 \\ k \neq l}}^{3,w} (\mu_{kl} \mathbf{V}_k - \mu_{lk} \mathbf{V}_l). \tag{2.195}
 \end{aligned}$$

In the case of isotropic turbulence we have

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_l \rho_l \mathbf{V}_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l \mathbf{V}_l \mathbf{V}_l \gamma) - \tilde{\mathcal{S}}_l + \frac{2}{3} \nabla (\gamma \alpha_l^e \rho_l k_l) + \alpha_l^e \gamma \overline{\mathcal{N}} p_l + \alpha_l \gamma_v \rho_l \mathbf{g} \\
 & + (p_l - p_m + \delta_l \sigma_{lm} \kappa_l - \Delta p_m^{l\sigma^*}) \nabla (\alpha_l^e \gamma) \\
 & - \gamma_v \alpha_l \rho_m \left\{ \begin{array}{l} c_{ml}^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \\ -c_{ml}^L \Delta \mathbf{V}_{ml} \times (\nabla \times \mathbf{V}_m) + \frac{1}{D_l} \frac{3}{4} c_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml} \end{array} \right\} \\
 & = \gamma_v \sum_{\substack{k=1 \\ k \neq l}}^{3,w} (\mu_{kl} \mathbf{V}_k - \mu_{lk} \mathbf{V}_l). \tag{2.195b}
 \end{aligned}$$

The momentum equation of the continuous phase so far takes the following form:

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_m \rho_m \mathbf{V}_m \gamma_v) + \nabla \cdot (\alpha_m^e \rho_m \mathbf{V}_m \mathbf{V}_m \gamma) \\
 & - \tilde{\mathcal{S}}_m + \frac{2}{3} \nabla (\gamma \alpha_m^e \rho_m k_m) - \nabla \cdot \left( \alpha_m^e \gamma_m \left[ 2\mathbf{D}_m - \frac{2}{3} (\nabla \cdot \mathbf{V}_m) \mathbf{I} \right] \right) + \alpha_m^e \gamma \overline{\mathcal{N}} p_m \\
 & + \alpha_m \gamma_v \rho_m \mathbf{g} + \Delta p_m^{l\sigma^*} \nabla (\alpha_l^e \gamma) - \Delta p_m^{w\sigma^*} \nabla \gamma
 \end{aligned}$$

$$\begin{aligned}
& +\gamma_v \alpha_l \rho_m \left\{ \begin{array}{l} c_{ml}^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \\ -c_{ml}^L \Delta \mathbf{V}_{ml} \times (\nabla \times \mathbf{V}_m) + \frac{1}{D_l} \frac{3}{4} c_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml} \end{array} \right\} \\
& +\gamma_v \rho_m \left[ c_w^{vm} \frac{1}{2} \frac{\partial \mathbf{V}_m}{\partial \tau} - c_{mw}^L \mathbf{V}_m \times (\nabla \times \mathbf{V}_m) + \frac{1}{D_w} \frac{3}{4} c_{mw}^d |\mathbf{V}_m| \mathbf{V}_m \right] \\
& = \gamma_v \sum_{\substack{k=1 \\ k \neq m}}^{3,w} (\mu_{km} \mathbf{V}_k - \mu_{mk} \mathbf{V}_m). \tag{2.196}
\end{aligned}$$

In the case of isotropic turbulence, we have

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (\alpha_m \rho_m \mathbf{V}_m \gamma_v) + \nabla \cdot (\alpha_m^e \rho_m \mathbf{V}_m \mathbf{V}_m \gamma) + \nabla \cdot [\alpha_m^e \gamma (\rho_m \overline{\mathbf{V}_m \mathbf{V}_m} - \mathbf{T}_m)] + \alpha_m^e \gamma \nabla p_m \\
& + \alpha_m \gamma_v \rho_m \mathbf{g} + \Delta p_m^{l\sigma^*} \nabla (\alpha_l^e \gamma) - \Delta p_m^{w\sigma^*} \nabla \gamma \\
& +\gamma_v \alpha_l \rho_m \left\{ \begin{array}{l} c_{ml}^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \\ -c_{ml}^L \Delta \mathbf{V}_{ml} \times (\nabla \times \mathbf{V}_m) + \frac{1}{D_l} \frac{3}{4} c_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml} \end{array} \right\} \\
& +\gamma_v \rho_m \left[ c_w^{vm} \frac{1}{2} \frac{\partial \mathbf{V}_m}{\partial \tau} - c_{mw}^L \mathbf{V}_m \times (\nabla \times \mathbf{V}_m) + \frac{1}{D_w} \frac{3}{4} c_{mw}^d |\mathbf{V}_m| \mathbf{V}_m \right] \\
& = \gamma_v \sum_{\substack{k=1 \\ k \neq m}}^{3,w} (\mu_{km} \mathbf{V}_k - \mu_{mk} \mathbf{V}_m). \tag{2.196b}
\end{aligned}$$

The relation between the two bulk pressures is given by the momentum jump condition. With the assumptions made in Section 2.6.2.1, the momentum jump condition, Eq. (2.23), reduces to

$$(\rho w)_{ml}^2 \left( \frac{1}{\rho_l} - \frac{1}{\rho_m} \right) + p_l^{m\sigma, \tau} - p_m^{l\sigma, \tau} - \sigma_{ml} \kappa_m = 0. \tag{2.197}$$

We exchange the subscripts  $l$  and  $m$  because the surface tension is assumed to belong to the liquid phase. After time averaging we obtain

$$(\rho w)_{ml}^2 \left( \frac{1}{\rho_l} - \frac{1}{\rho_m} \right) + p_l^{m\sigma} - p_m^{l\sigma} + \sigma_{ml} \kappa_l = 0, \quad (2.198)$$

or in terms of bulk pressure

$$p_l - p_m + \sigma_{ml} \kappa_l = \Delta p_m^{l\sigma} + (\rho w)_{ml}^2 \left( \frac{1}{\rho_m} - \frac{1}{\rho_l} \right). \quad (2.199)$$

Actually, the pressure difference  $\Delta p_m^{l\sigma}$  varies over the surface and some surface-averaged value

$$\frac{1}{F_{l\sigma}} \left| \int_{F_{l\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l dF \right| = \frac{Vol}{F_{l\sigma}} \left| \frac{1}{Vol} \int_{F_{l\sigma}} \Delta p_m^{l\sigma} \mathbf{n}_l dF \right| \quad (2.200)$$

may be used. There is no experience in this field and future investigations are necessary. Approximations are thinkable for predominant surface tension and low mass transfer

$$p_l - p_m + \sigma_{ml} \kappa_l \approx 0, \quad (2.201)$$

or for predominant mass transfer

$$p_l - p_m + \sigma_{ml} \kappa_l \approx (\rho w)_{ml}^2 \left( \frac{1}{\rho_m} - \frac{1}{\rho_l} \right). \quad (2.202)$$

Note that for spheres  $\kappa_l = 1/R_l + 1/R_l = 2/R_l$ . Remember that if the radius is inside the field the curvature is negative.

Now let us as a practical illustration of the application of the theory analyze the eigenvalues for bubble flow without mass transfer in a one-dimensional horizontal channel with constant cross-section section, assuming noncompressible phases and neglecting the diffusion terms. The governing system then simplifies to

$$p = p_2, \quad (2.203)$$

$$p_1 = p + \sigma_{12} \kappa_2, \quad (2.204)$$

$$\alpha_1 \frac{\partial w_1}{\partial z} + \alpha_2 \frac{\partial w_2}{\partial z} + (w_1 - w_2) \frac{\partial \alpha_1}{\partial z} = 0 \quad (2.205)$$

$$\frac{\partial \alpha_1}{\partial \tau} + \alpha_1 \frac{\partial w_1}{\partial z} + w_1 \frac{\partial \alpha_1}{\partial z} = 0 \quad (2.206)$$

$$\begin{aligned} & \alpha_1 \left( \rho_1 + \rho_2 c_{21}^{vm} \right) \left( \frac{\partial w_1}{\partial \tau} + w_1 \frac{\partial w_1}{\partial z} \right) - \alpha_1 \rho_2 c_{21}^{vm} \left( \frac{\partial w_2}{\partial \tau} + w_1 \frac{\partial w_2}{\partial z} \right) + \alpha_1 \frac{\partial p}{\partial z} - \Delta p_2^{1\sigma} \frac{\partial \alpha_1}{\partial z} \\ & = \alpha_1 \rho_2 \frac{1}{D_1} \frac{3}{4} c_{21}^d |\Delta w_{21}| \Delta w_{21}, \end{aligned} \quad (2.207)$$

$$\begin{aligned}
& -\alpha_1 \rho_2 c_{21}^{vm} \left( \frac{\partial w_1}{\partial \tau} + w_1 \frac{\partial w_1}{\partial z} \right) + (\alpha_2 + \alpha_1 c_{21}^{vm} + c_w^{vm}) \rho_2 \frac{\partial w_2}{\partial \tau} + (\alpha_2 w_2 + \alpha_1 c_{21}^{vm} w_1) \rho_2 \frac{\partial w_2}{\partial z} \\
& + \alpha_2 \frac{\partial p}{\partial z} + \Delta p_2^{1\sigma*} \frac{\partial \alpha_1}{\partial z} = -\frac{3}{4} \rho_2 \left( \alpha_1 \frac{1}{D_1} c_{21}^d |\Delta w_{21}| \Delta w_{21} + \frac{1}{D_w} c_{2w}^d |w_2| w_2 \right)
\end{aligned} \tag{2.208}$$

In matrix notation we have

$$\begin{aligned}
& \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & \alpha_1 (\rho_1 + \rho_2 c_{21}^{vm}) & -\alpha_1 \rho_2 c_{21}^{vm} \\ 0 & 0 & -\alpha_1 \rho_2 c_{21}^{vm} & (\alpha_2 + \alpha_1 c_{21}^{vm} + c_w^{vm}) \rho_2 \end{pmatrix} \frac{\partial}{\partial \tau} \begin{pmatrix} p \\ \alpha_1 \\ w_1 \\ w_2 \end{pmatrix} \\
& + \begin{pmatrix} 0 & w_1 - w_2 & \alpha_1 & \alpha_2 \\ 0 & w_1 & \alpha_1 & 0 \\ \alpha_1 & -\Delta p_2^{1\sigma*} & \alpha_1 (\rho_1 + \rho_2 c_{21}^{vm}) w_1 & -\alpha_1 \rho_2 c_{21}^{vm} w_1 \\ \alpha_2 & \Delta p_2^{1\sigma*} & -\alpha_1 \rho_2 c_{21}^{vm} w_1 & (\alpha_2 w_2 + \alpha_1 c_{21}^{vm} w_1) \rho_2 \end{pmatrix} \frac{\partial}{\partial z} \begin{pmatrix} p \\ \alpha_1 \\ w_1 \\ w_2 \end{pmatrix} \\
& = \begin{pmatrix} 0 \\ 0 \\ \alpha_1 \rho_2 \frac{1}{D_1} \frac{3}{4} c_{21}^d |\Delta w_{21}| \Delta w_{21} \\ -\frac{3}{4} \rho_2 \left( \alpha_1 \frac{1}{D_1} c_{21}^d |\Delta w_{21}| \Delta w_{21} + \frac{1}{D_w} c_{2w}^d |w_2| w_2 \right) \end{pmatrix}.
\end{aligned} \tag{2.209}$$

The reader unfamiliar with the analysis of the type of a system of partial differential equations by first computing the eigenvalues, eigenvectors, and canonical forms, is recommended to first read Section 11 before continuing here.

The characteristic equation for determining the eigenvalues is then

$$\begin{pmatrix} 0 & w_1 - w_2 & \alpha_1 & \alpha_2 \\ 0 & w_1 - \lambda & \alpha_1 & 0 \\ \alpha_1 & -\Delta p_2^{1\sigma*} & \alpha_1 (\rho_1 + \rho_2 c_{21}^{vm}) (w_1 - \lambda) & -\alpha_1 \rho_2 c_{21}^{vm} (w_1 - \lambda) \\ \alpha_2 & \Delta p_2^{1\sigma*} & -\alpha_1 \rho_2 c_{21}^{vm} (w_1 - \lambda) & \begin{bmatrix} \alpha_2 w_2 + \alpha_1 c_{21}^{vm} w_1 \\ -\lambda (\alpha_2 + \alpha_1 c_{21}^{vm} + c_w^{vm}) \end{bmatrix} \rho_2 \end{pmatrix} = 0.$$

$$(2.210)$$



or

$$\alpha_2 \left( c_{21}^{vm} + \alpha_2 \frac{\rho_1}{\rho_2} \right) (w_1 - \lambda)^2 + \alpha_1 (w_2 - \lambda) \left[ \alpha_2 w_2 + \alpha_1 c_{21}^{vm} w_1 - \lambda (\alpha_2 + \alpha_1 c_{21}^{vm} + c_w^{vm}) \right] - \alpha_1 \alpha_2 (w_1 - w_2) c_{21}^{vm} (w_1 - \lambda) + \alpha_2 \frac{\Delta p_2^{1\sigma^*}}{\rho_2} = 0 \quad (2.211)$$

or

$$a\lambda^2 - 2b\lambda + c = 0 \quad (2.212)$$

where

$$a = \alpha_2 \left( \alpha_1 - \alpha_2 \frac{\rho_1}{\rho_2} \right) + (\alpha_1^2 - \alpha_2) c_{21}^{vm} + \alpha_1 c_w^{vm} > 0, \quad (2.213)$$

$$b = \frac{1}{2} \left\{ \left[ \alpha_1 w_2 + (\alpha_1^2 - 2\alpha_2 - \alpha_1 \alpha_2) w_1 \right] c_{21}^{vm} + \alpha_1 (2\alpha_2 + c_w^{vm}) w_2 - 2w_1 \alpha_2^2 \frac{\rho_1}{\rho_2} \right\}, \quad (2.214)$$

$$c = (\alpha_1 w_1 w_2 + \alpha_2^2 w_1^2) c_{21}^{vm} + \alpha_1 \alpha_2 w_2^2 + \alpha_2^2 \frac{\rho_1}{\rho_2} w_1^2 + \alpha_2 \frac{\Delta p_2^{1\sigma^*}}{\rho_2}, \quad (2.215)$$

with two real solutions

$$\lambda_{1,2} = \frac{b \pm \sqrt{b^2 - ac}}{a} \quad (2.216)$$

for

$$b^2 > ac \quad (2.217)$$

which is, in fact, the stability criterion for bubbly flow.

## 2.9 General working form for dispersed and continuous phases

In this section, we write a single equation valid for both disperse and continuous phases. First we compare the terms in the two equations

$$\mathbf{f}_{d\alpha} = (p_d - p_c + \delta_d \sigma_{dc} \kappa_d - \Delta p_c^{d\sigma^*}) \nabla (\alpha_d^e \gamma), \quad (2.218)$$

and

$$\mathbf{f}_{c\alpha} = \Delta p_c^{d\sigma^*} \nabla (\alpha_d^e \gamma), \quad (2.219)$$

and realize that in both cases the force is a function of the gradient of the disperse volume fraction multiplied by different multipliers. With this notation, we have

$$\begin{aligned}
 & \left. \frac{\partial}{\partial \tau} (\alpha_i \rho_l \mathbf{V}_i \gamma_v) + \nabla \cdot \left( \alpha_i^e \rho_l \gamma \left\{ \mathbf{V}_i \mathbf{V}_l - v_l^* \left[ \nabla \mathbf{V}_l + (\nabla \mathbf{V}_l)^T - \frac{2}{3} (\nabla \cdot \mathbf{V}_l) \mathbf{I} \right] \right\} \right) \right. \\
 & + \alpha_i^e \gamma \nabla p_l + \alpha_i \rho_l \mathbf{g} \gamma_v - \Delta p_l^{w\sigma^*} \nabla \gamma + \mathbf{f}_{l\alpha} \\
 & - \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^3 \left\{ \begin{aligned} & \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta \mathbf{V}_{ml} \\ & + \bar{c}_{ml}^L (\mathbf{V}_l - \mathbf{V}_m) \times (\nabla \times \mathbf{V}_m) + \bar{c}_{ml}^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \end{aligned} \right\} \\
 & + \gamma_v \left[ \bar{c}_{lw}^d |\mathbf{V}_l| \mathbf{V}_l + \rho_l c_{lw}^{vm} \frac{\partial \mathbf{V}_l}{\partial \tau} - \rho_l c_{lw}^L \mathbf{V}_l \times (\nabla \times \mathbf{V}_l) \right] = \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} \mathbf{V}_m - \mu_{lm} \mathbf{V}_l).
 \end{aligned} \tag{2.220}$$

Note that

$$\Delta p_d^{w\sigma^*} = 0, \tag{2.221}$$

$$\Delta p_c^{w\sigma^*} \neq 0. \tag{2.222}$$

Similarly

$$c_{dw}^d, c_{dw}^{vm}, c_{dw}^L = 0, \tag{2.223}$$

and

$$c_{cw}^d, c_{cw}^{vm}, c_{cw}^L \neq 0, \tag{2.224}$$

if the continuum is in a contact with the wall. For easy programming, a couple of simple drag, lift, and virtual mass coefficients combined as follows are introduced for each field

$$\bar{c}_{ml}^d = \bar{c}_{lm}^d = \frac{3}{4} (\alpha_m \rho_l c_{lm}^d / D_m + \alpha_l \rho_m c_{ml}^d / D_l), \tag{2.225}$$

$$\bar{c}_{ml}^L = \bar{c}_{lm}^L = \alpha_m \rho_l c_{lm}^L + \alpha_l \rho_m c_{ml}^L, \tag{2.226}$$

$$\bar{c}_{ml}^{vm} = \bar{c}_{lm}^{vm} = \alpha_m \rho_l c_{lm}^{vm} + \alpha_l \rho_m c_{ml}^{vm}. \tag{2.227}$$

If field  $l$  is disperse and surrounded by the continuous  $m$ -field, the coefficient  $c_{ml}$  is not equal to zero and  $c_{lm}$  is equal to zero, and vice versa. In other words if the second subscript refers to a disperse field, the local size of dispersion is positive and the coefficients are not equal to zero. For application in computer codes the following general notation is recommended:

$$\mathbf{f}_l = \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \bar{c}_{ml}^d \mathbf{f}_{ml} \quad (2.228)$$

to take into account the fact that a control volume may contain two dispersed fields carried by one continuous field. This approach together with implicit discretization of the momentum equations, their strong coupling through a special numerical procedure, and the comparison with data for three-phase bubble flow was presented by *Kolev et al. (1991)*. In Volume II the reader will find additional information on practical computation of drag forces in multiphase flows.

Equation (2.220) is the rigorously derived local volume and time average momentum balance for multiphase flows in heterogeneous porous structures conditionally divided into three velocity fields.

The nonconservative form of the momentum conservation equation in component notation is given in Appendix 2.3. In the same appendix some interesting single-phase analytical solutions are given. They can be used as benchmarks for testing the accuracy of the numerical solution methods.

## 2.10 Some practical simplifications

Equation (2.220) has been used since 1984 in the IVA1 to IVA6 computer codes *Kolev (1985, 1986a, 1986b, 1987, 1991a, 1991b, 1993a, 1993b, 1993c, 1993d, 1994, 1996, 1999)* with the following simplifications:

$$p_l - p_m + \frac{2\sigma_{lm}}{R_l} - \Delta p_m^{l\sigma^*} \approx 0, \quad (2.229)$$

$$p_l \approx p_m = p, \quad (2.230)$$

$$v_l' \approx 0. \quad (2.231)$$

Assumption (2.229) is quite close to the local volume and time average interfacial jump condition at the interface and, therefore, does not lead to any problems for slow interfacial mass transfer.

Assumption (2.230) leads to the so-called *single-pressure model*. It should be emphasized that the most important interfacial pressure forces, which are considerably larger in magnitude than the error introduced by the single-pressure assumption, have already been taken into account. This assumption likewise does not lead to any problems. In this type of single-pressure model the hyperbolicity is

preserved due to the stabilizing viscous, drag, and virtual mass terms. Neglect of the viscous, drag and virtual mass terms leads to unphysical models.

Assumption (2.231) was dictated by a lack of knowledge. When information for  $v_i'$  becomes available, this can easily be included, as the viscous terms have already been taken into account.

The resulting simplified form of Eq. (2.220) is, therefore,

$$\begin{aligned}
& \frac{\partial}{\partial \tau} \left( \alpha_i \rho_i \mathbf{V}_i \gamma_v \right) + \nabla \cdot \left\{ \alpha_i^e \rho_i \gamma_v \left[ \mathbf{V}_i \mathbf{V}_i - v_i^* \left[ \nabla \mathbf{V}_i + (\nabla \mathbf{V}_i)^T - \frac{2}{3} (\nabla \cdot \mathbf{V}_i) \mathbf{I} \right] \right] \right\} \\
& + \alpha_i^e \gamma_v \mathcal{N} p + \alpha_i \rho_i \mathbf{g} \gamma_v \\
& - \gamma_v \sum_{\substack{m=1 \\ m \neq i}}^3 \left\{ \begin{aligned} & \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \cdot \Delta \mathbf{V}_{ml} \\ & + \bar{c}_{ml}^L (\mathbf{V}_i - \mathbf{V}_m) \times (\nabla \times \mathbf{V}_m) + \bar{c}_{ml}^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_i \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \end{aligned} \right\} \\
& + \gamma_v \left[ \bar{c}_{lw}^d |\mathbf{V}_l| \cdot \mathbf{V}_l + \rho_l c_{lw}^{vm} \frac{\partial \mathbf{V}_l}{\partial \tau} - \rho_l c_{lw}^L \mathbf{V}_l \times (\nabla \times \mathbf{V}_l) \right] \\
& = \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} \mathbf{V}_m - \mu_{lm} \mathbf{V}_l). \tag{2.232}
\end{aligned}$$

The form of Eq. (2.232) is sometimes called *conservative* in order to distinguish it from the *nonconservative* form. The nonconservative form is derived by applying the chain rule to the first two terms and inserting the mass-conservation equation (1.45). The resulting equation,

$$\begin{aligned}
& \alpha_i \rho_i \left[ \gamma_v \frac{\partial \mathbf{V}_i}{\partial \tau} + (\mathbf{V}_i \mathcal{N}) \mathbf{V}_i \right] - \nabla \cdot \left\{ \alpha_i^e \rho_i \gamma_v^* \left[ \nabla \mathbf{V}_i + (\nabla \mathbf{V}_i)^T - \frac{2}{3} (\nabla \cdot \mathbf{V}_i) \mathbf{I} \right] \right\} \\
& + \alpha_i^e \gamma_v \mathcal{N} p + \alpha_i \rho_i \mathbf{g} \gamma_v \\
& - \gamma_v \sum_{\substack{m=1 \\ m \neq i}}^3 \left\{ \begin{aligned} & \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \cdot \Delta \mathbf{V}_{ml} \\ & + \bar{c}_{ml}^L (\mathbf{V}_i - \mathbf{V}_m) \times (\nabla \times \mathbf{V}_m) + \bar{c}_{ml}^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_{ml} + (\mathbf{V}_i \cdot \nabla) \Delta \mathbf{V}_{ml} \right] \end{aligned} \right\} \\
& + \gamma_v \left[ \bar{c}_{lw}^d |\mathbf{V}_l| \cdot \mathbf{V}_l + \rho_l c_{lw}^{vm} \frac{\partial \mathbf{V}_l}{\partial \tau} - \rho_l c_{lw}^L \mathbf{V}_l \times (\nabla \times \mathbf{V}_l) \right]
\end{aligned}$$

$$= \gamma_v \left\{ \sum_{m=1}^3 [\mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)] + \mu_{wl} (\mathbf{V}_{wl} - \mathbf{V}_l) + \mu_{lw} (\mathbf{V}_{lw} - \mathbf{V}_l) \right\}, \quad (2.233)$$

contains some extremely interesting information, namely:

- (a) mass sinks of the velocity field  $l$  have no influence on the velocity change (an exception is the controlled flow suction from the structure through the structure interface), and
- (b) mass sources from a donor field whose velocities differ from the velocity of the receiving field influence the velocity change.

To facilitate the direct use of the vector equation (2.333), we give its scalar components for the most frequently used cylindrical,  $\kappa = 1$ , and Cartesian,  $\kappa = 0$ , coordinate systems.

$r$  direction

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l u_l \gamma_v) + \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left[ r^\kappa \alpha_l \rho_l \left( u_l u_l - v_l^* \frac{\partial u_l}{\partial r} \right) \gamma_r \right] \\ & + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left[ \alpha_l \rho_l \left( v_l u_l - v_l^* \frac{1}{r^\kappa} \frac{\partial u_l}{\partial \theta} \right) \gamma_\theta \right] \\ & - \frac{\kappa}{r^\kappa} \alpha_l \rho_l \left[ v_l v_l - \frac{2}{r^\kappa} v_l^* \left( \frac{\partial v_l}{\partial \theta} + u_l \right) \right] \gamma_\theta + \frac{\partial}{\partial z} \left[ \alpha_l \rho_l \left( w_l u_l - v_l^* \frac{\partial u_l}{\partial z} \right) \gamma_z \right] \\ & + \alpha_l \gamma_r \frac{\partial p}{\partial r} + (\alpha_l \rho_l g_r + f_{lu}) \gamma_v = \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} u_{ml} - \mu_{lm} u_l) + f_{vlu}, \end{aligned} \quad (2.234)$$

where

$$\begin{aligned} f_{vlu} &= \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left[ r^\kappa \alpha_l \rho_l v_l^* \frac{\partial u_l}{\partial r} \gamma_r \right] + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left[ \alpha_l \rho_l v_l^* r^\kappa \frac{\partial}{\partial r} \left( \frac{v_l}{r^\kappa} \right) \gamma_\theta \right] \\ & + \frac{\partial}{\partial z} \left( \alpha_l \rho_l v_l^* \frac{\partial w_l}{\partial z} \gamma_z \right) - \frac{2}{3} \frac{1}{r^\kappa} \left\{ \frac{\partial}{\partial r} \left[ r^\kappa (\alpha_l \rho_l v_l^* \nabla \cdot \mathbf{V}_l) \gamma_r \right] - (\alpha_l \rho_l v_l^* \nabla \cdot \mathbf{V}_l) \gamma_\theta \right\}; \end{aligned} \quad (2.235)$$

$\theta$  direction

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l v_l \gamma_v) + \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left[ r^\kappa \alpha_l \rho_l \left( u_l v_l - v_l^* \frac{\partial v_l}{\partial r} \right) \gamma_r \right] \\ & + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left[ \alpha_l \rho_l \left( v_l v_l - v_l^* \frac{1}{r^\kappa} \frac{\partial v_l}{\partial \theta} \right) \gamma_\theta \right] \end{aligned}$$

$$\begin{aligned}
& + \frac{\kappa}{r^\kappa} \alpha_i \rho_l \left[ v_l u_l - r^\kappa v_l^* \frac{\partial}{\partial r} \left( \frac{v_l}{r^\kappa} \right) - v_l^* \frac{1}{r^\kappa} \frac{\partial u_l}{\partial \theta} \right] \gamma_\theta + \frac{\partial}{\partial z} \left[ \alpha_i \rho_l \left( w_l v_l - v_l^* \frac{\partial v_l}{\partial z} \right) \gamma_z \right] \\
& + \alpha_i \gamma_\theta \frac{1}{r^\kappa} \frac{\partial p}{\partial \theta} + (\alpha_i \rho_l g_\theta + f_{lv}) \gamma_v = \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} v_{ml} - \mu_{lm} v_l) + f_{vlv}, \quad (2.236)
\end{aligned}$$

where

$$\begin{aligned}
f_{vlv} &= \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left[ \alpha_i \rho_l v_l^* \left( \frac{\partial u_l}{\partial \theta} - v_l \right) \gamma_r \right] + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left[ \alpha_i \rho_l v_l^* \frac{1}{r^\kappa} \left( \frac{\partial v_l}{\partial \theta} + 2u_l \right) \gamma_\theta \right] \\
& + \frac{\partial}{\partial z} \left( \alpha_i \rho_l v_l^* \frac{1}{r^\kappa} \frac{\partial w_l}{\partial \theta} \gamma_z \right) - \frac{2}{3} \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left[ (\alpha_i \rho_l v_l^* \nabla \cdot \mathbf{V}_l) \gamma_\theta \right]; \quad (2.237)
\end{aligned}$$

$z$  direction

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (\alpha_i \rho_l w_l \gamma_v) + \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left[ r^\kappa \alpha_i \rho_l \left( u_l w_l - v_l^* \frac{\partial w_l}{\partial r} \right) \gamma_r \right] \\
& + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left[ \alpha_i \rho_l \left( v_l w_l - v_l^* \frac{1}{r^\kappa} \frac{\partial w_l}{\partial \theta} \right) \gamma_\theta \right] + \frac{\partial}{\partial z} \left[ \alpha_i \rho_l \left( w_l w_l - v_l^* \frac{\partial w_l}{\partial z} \right) \gamma_z \right] \\
& + \alpha_i \gamma_z \frac{\partial p}{\partial z} + (\alpha_i \rho_l g_z + f_{lv}) \gamma_v = \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} w_{ml} - \mu_{lm} w_l) + f_{vlv}, \quad (2.238)
\end{aligned}$$

where

$$\begin{aligned}
f_{vlv} &= \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left( r^\kappa \alpha_i \rho_l v_l^* \frac{\partial u_l}{\partial z} \gamma_r \right) + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left( \alpha_i \rho_l v_l^* \frac{\partial v_l}{\partial z} \gamma_\theta \right) + \frac{\partial}{\partial z} \left( \alpha_i \rho_l v_l^* \frac{\partial w_l}{\partial z} \gamma_z \right) \\
& - \frac{2}{3} \frac{\partial}{\partial z} \left[ (\alpha_i \rho_l v_l^* \nabla \cdot \mathbf{V}_l) \gamma_z \right]. \quad (2.239)
\end{aligned}$$

Here

$$\nabla \cdot \mathbf{V}_l = \left[ \frac{1}{r^\kappa} \frac{\partial}{\partial r} (r^\kappa u_l) + \frac{1}{r^\kappa} \frac{\partial v_l}{\partial \theta} + \frac{\partial w_l}{\partial z} \right]. \quad (2.240)$$

Note that all interfacial forces are designated with  $\mathbf{f}_l$  with components  $f_{lu}$ ,  $f_{lv}$ , and  $f_{lw}$ .

The viscous terms have been rearranged in order to obtain a *convection-diffusion form* for the left-hand side of the momentum equations. The residual terms are pooled into the momentum source terms  $\mathbf{f}_v$ . This notation is justified for two reasons. First, for a single-phase flow in a pool (unrestricted flows) and a constant density, the source terms  $\mathbf{f}_v$  are equal to zero, which *intuitively leads to the idea that the main viscous influence is outside the  $\mathbf{f}_v$  source terms*. This argument led some authors to derive an explicit discretization for the  $\mathbf{f}_v$  source terms for

single-phase flow applications, see *Trent and Eyster* (1983), or even to neglect these source terms. Second, methods with known mathematical properties for the discretization of convection-diffusion equations have already been developed, and these can be applied directly.

Note that for Cartesian coordinates the convective components contain spatial derivatives. In the case of cylindrical coordinates the convective part contains in addition two components not containing spatial derivatives. The component

$$\text{centrifugal force} = -\frac{\kappa}{r^\kappa} \alpha_1 \rho_1 v_1 v_1 \gamma_\theta \quad (2.241)$$

is known in the literature as centrifugal force. It gives the effective force component in the  $r$  direction resulting from fluid motion in the  $\theta$  direction. The component

$$\text{Coriolis force} = \frac{\kappa}{r^\kappa} \alpha_1 \rho_1 v_1 u_1 \gamma_\theta \quad (2.242)$$

is known in the literature as the *Coriolis* force. It is an effective force component in the  $\theta$  direction when there is flow in both the  $r$  and  $\theta$  directions. This results in the components of the viscous stress tensor,  $\tau_{1,\theta\theta}$  and  $\tau_{1,r\theta}$  corresponding to these forces and acting in the opposite directions.

## 2.11 Conclusion

The positive experience with Eqs. (2.334)–(2.340) in the development of the IVA code *Kolev* (1985, 1993e, 1996, 1999) allows one to recommend these equations for general use. One should keep in mind for application purposes that both sides of the equations are local volume and time averages.

Understanding of the local volume and time average momentum equations is a prerequisite for understanding the second law of thermodynamics and its extremely interesting application to yield a simple description of this highly complicated system. As a next step in this direction, a rigorous formulation of the equations reflecting the second law of thermodynamics for this multiphase, multicomponent system has been successfully derived. The result of this derivation has formed the subject of a Chapter 5, see also *Kolev* (1995), and a comment to this publication (*Kolev* 1997).

## Appendix 2.1

Substituting in the momentum conservation equation and performing time averaging,

$$\overline{\langle p_i^\tau \rangle}^{lc} = p_i,$$

$$\overline{\langle p_m^\tau \rangle^{me}} = p_m,$$

$$\overline{\alpha_l \gamma_v \langle \rho_l \rangle^l \mathbf{V}_l} = \alpha_l \gamma_v \rho_l \mathbf{V}_l,$$

$$\overline{\alpha_l \gamma_v \langle \rho_l \rangle^l \mathbf{V}_l'} = 0,$$

$$\overline{\alpha_l^e \gamma \langle \rho_l \rangle^l \mathbf{V}_l \mathbf{V}_l} = \alpha_l^e \gamma \rho_l \mathbf{V}_l \mathbf{V}_l,$$

$$\overline{\alpha_l^e \gamma \langle \rho_l \rangle^l \mathbf{V}_l \mathbf{V}_l'} = 0,$$

$$\overline{\alpha_l^e \gamma \langle \rho_l \rangle^l \mathbf{V}_l \mathbf{V}_l'} = \alpha_l^e \gamma \rho_l \overline{\mathbf{V}_l \mathbf{V}_l'},$$

$$\overline{\nabla \cdot (\alpha_l^e \gamma \langle \mathbf{T}_l^\tau \rangle^{le})} = \nabla \cdot (\alpha_l^e \gamma \mathbf{T}_l),$$

$$\overline{\alpha_l^e \mathcal{N} \cdot \langle p_l^\tau \rangle^{le}} = \alpha_l^e \mathcal{N} \cdot p_l,$$

$$\overline{\alpha_l \gamma_v \langle \rho_l \rangle^l \mathbf{g}} = \alpha_l \gamma_v \rho_l \mathbf{g},$$

$$\sum_{m=1}^3 \left( \overline{\mu_{ml}^\tau \langle \mathbf{V}_m^\tau \rangle^{me}} - \overline{\mu_{lm}^\tau \langle \mathbf{V}_l^\tau \rangle^{le}} \right) \approx \sum_{m=1}^3 (\mu_{ml} \mathbf{V}_m - \mu_{lm} \mathbf{V}_l),$$

and

$$+ (p_l - p_m + \delta_l \sigma_{lm} \kappa_l) \left( \nabla (\alpha_l^e \gamma) + \frac{1}{Vol} \int_{F_{lv}} \mathbf{n}_l dF \right) - \delta_l \frac{1}{Vol} \int_{F_{l\sigma}} (\nabla_l \sigma_{lm}) dF$$

$$+ \frac{1}{Vol} \int_{F_{l\sigma}} (\Delta p_m^{l\sigma} \mathbf{I} - \mathbf{T}_m^{l\sigma}) \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{lv}} (\Delta p_l^{w\sigma} \mathbf{I} - \mathbf{T}_l^{w\sigma}) \cdot \mathbf{n}_l dF$$

$$= \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} \mathbf{V}_m - \mu_{lm} \mathbf{V}_l)$$

$$\int_{F_{l\sigma}} \left\{ (-\Delta p_m^{l\sigma} \mathbf{I} + \mathbf{T}_m^{l\sigma}) \cdot \mathbf{n}_l - \nabla_l \sigma_{lm} \right\} dF + \int_{F_{lv}} (-\Delta p_l^{w\sigma} \mathbf{I} + \mathbf{T}_l^{w\sigma}) \cdot \mathbf{n}_l dF$$

$$= \int_{F_{l\sigma}} \left\{ (-\Delta p_m^{l\sigma} \mathbf{I} + \mathbf{T}_m^{l\sigma}) \cdot \mathbf{n}_l - \nabla_l \sigma_{lm} \right\} dF + \int_{F_{lv}} (-\Delta p_l^{w\sigma} \mathbf{I} + \mathbf{T}_l^{w\sigma}) dF$$

one obtains Eq. (2.50a).



## Appendix 2.2

The normal velocity difference can be obtained by splitting the relative velocity vector at the interface  $\Delta \mathbf{V}_{lm}$  into a component that is parallel to  $\mathbf{n}_l$

$$\Delta \mathbf{V}_{lm}^n = \text{proj}_{\mathbf{n}_l} \Delta \mathbf{V}_{lm} = \left( \frac{\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l}{\mathbf{n}_l \cdot \mathbf{n}_l} \right) \mathbf{n}_l = (\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l) \mathbf{n}_l,$$

with a magnitude

$$|\Delta \mathbf{V}_{lm}^n| = |\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l| = \sqrt{[n_{lx}(u_l - u_m)]^2 + [n_{ly}(v_l - v_m)]^2 + [n_{lz}(w_l - w_m)]^2}$$

and a component orthogonal to  $\mathbf{n}_l$ ,

$$\begin{aligned} \Delta \mathbf{V}_{lm}^t &= \Delta \mathbf{V}_{lm} - \text{proj}_{\mathbf{n}_l} \Delta \mathbf{V}_{lm} = \Delta \mathbf{V}_{lm} - (\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l) \mathbf{n}_l \\ &= [\Delta u_{lm} - n_{lx}(\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l)] \mathbf{i} + [\Delta v_{lm} - n_{ly}(\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l)] \mathbf{j} + [\Delta w_{lm} - n_{lz}(\Delta \mathbf{V}_{lm} \cdot \mathbf{n}_l)] \mathbf{k} \end{aligned}$$

## Appendix 2.3

The nonconservative form of Eqs. (2.195) is

$r$  direction:

$$\begin{aligned} & \alpha_l \rho_l \left( \gamma_v \frac{\partial u_l}{\partial \tau} + \gamma_r \frac{1}{2} \frac{\partial u_l^2}{\partial r} + v_l \gamma_\theta \frac{1}{r^\kappa} \frac{\partial u_l}{\partial \theta} + w_l \gamma_z \frac{\partial u_l}{\partial z} \right) \\ & - \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left( r^\kappa \alpha_l \rho_l v_l^* \frac{\partial u_l}{\partial r} \gamma_r \right) - \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left( \alpha_l \rho_l v_l^* \frac{1}{r^\kappa} \frac{\partial u_l}{\partial \theta} \gamma_\theta \right) - \frac{\partial}{\partial z} \left( \alpha_l \rho_l v_l^* \frac{\partial u_l}{\partial z} \gamma_z \right) \\ & - \frac{\kappa}{r^\kappa} \alpha_l \rho_l \left[ v_l v_l - \frac{2}{r^\kappa} v_l^* \left( \frac{\partial v_l}{\partial \theta} + u_l \right) \right] \gamma_\theta + \alpha_l \gamma_r \frac{\partial p}{\partial r} + (\alpha_l \rho_l g_r + f_{lu}) \gamma_v \\ & = \gamma_v \left\{ \sum_{m=1}^3 [\mu_{ml}(u_m - u_l)] + \mu_{wl}(u_{wl} - u_l) - \mu_{lw}(u_{lw} - u_l) \right\} + f_{vlu}; \end{aligned}$$

$\theta$  direction:

$$\begin{aligned} & \alpha_l \rho_l \left( \gamma_v \frac{\partial v_l}{\partial \tau} + u_l \gamma_r \frac{\partial v_l}{\partial r} + \gamma_\theta \frac{1}{2} \frac{1}{r^\kappa} \frac{\partial v_l^2}{\partial \theta} + w_l \gamma_z \frac{\partial v_l}{\partial z} \right) \\ & - \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left( r^\kappa \alpha_l \rho_l v_l^* \frac{\partial v_l}{\partial r} \gamma_r \right) - \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left( \alpha_l \rho_l v_l^* \frac{1}{r^\kappa} \frac{\partial v_l}{\partial \theta} \gamma_\theta \right) - \frac{\partial}{\partial z} \left( \alpha_l \rho_l v_l^* \frac{\partial v_l}{\partial z} \gamma_z \right) \end{aligned}$$

$$\begin{aligned}
& + \frac{\kappa}{r^\kappa} \alpha_l \rho_l \left[ v_l u_l - r^\kappa v_l^* \frac{\partial}{\partial r} \left( \frac{v_l}{r^\kappa} \right) - v_l^* \frac{1}{r^\kappa} \frac{\partial u_l}{\partial \theta} \right] \gamma_\theta + \alpha_l \gamma_\theta \frac{1}{r^\kappa} \frac{\partial p}{\partial \theta} + (\alpha_l \rho_l g_\theta + f_{l\nu}) \gamma_\nu \cdot \\
& = \gamma_\nu \left\{ \sum_{m=1}^3 [\mu_{ml} (v_m - v_l)] + \mu_{wl} (v_{wl} - v_l) - \mu_{lw} (v_{lw} - v_l) \right\} + f_{l\nu};
\end{aligned}$$

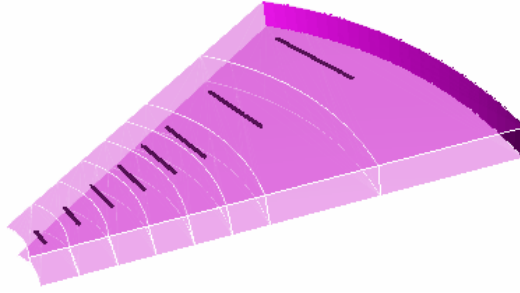
$z$  direction:

$$\begin{aligned}
& \alpha_l \rho_l \left( \gamma_\nu \frac{\partial w_l}{\partial \tau} + u_l \gamma_r \frac{\partial w_l}{\partial r} + v_l \gamma_\theta \frac{1}{r^\kappa} \frac{\partial w_l}{\partial \theta} + \gamma_z \frac{1}{2} \frac{\partial w_l^2}{\partial z} \right) \\
& - \frac{1}{r^\kappa} \frac{\partial}{\partial r} \left( r^\kappa \alpha_l \rho_l v_l^* \frac{\partial w_l}{\partial r} \gamma_r \right) - \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} \left( \alpha_l \rho_l v_l^* \frac{1}{r^\kappa} \frac{\partial w_l}{\partial \theta} \gamma_\theta \right) - \frac{\partial}{\partial z} \left( \alpha_l \rho_l v_l^* \frac{\partial w_l}{\partial z} \gamma_z \right) \\
& + \alpha_l \gamma_z \frac{\partial p}{\partial z} + (\alpha_l \rho_l g_z + f_{lw}) \gamma_\nu \\
& = \gamma_\nu \left\{ \sum_{m=1}^3 [\mu_{ml} (w_m - w_l)] + \mu_{wl} (w_{wl} - w_l) - \mu_{lw} (w_{lw} - w_l) \right\} + f_{lw}.
\end{aligned}$$

**Some simple single-phase test cases:** For testing numerical solutions it is important to provide a set of simple benchmarks having analytical solutions. Some of them are presented here.

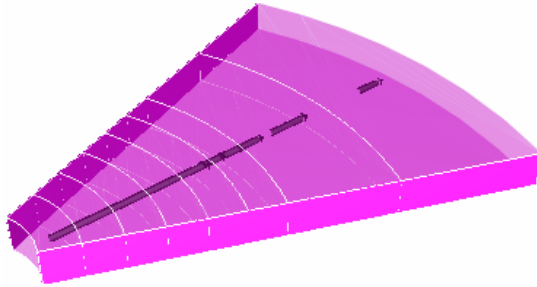
*Rigid body steady rotation problem:* This test problem presents a hollow cylinder with symmetric flow in the azimuth direction, see Fig. A2.3-1. No axial and radial flow exists. The mass conservation equation gives  $\frac{\partial v}{\partial \theta} = 0$ . The  $r$  direction momentum equation simplifies to  $\rho \frac{v^2}{r} = \frac{\partial p}{\partial r}$ , and the  $\theta$  direction momentum equation gives  $\frac{\partial p}{\partial \theta} = 0$ . For constant rotational frequency  $\omega$ , ( $v(r) = r\omega$ ), the analytical solution of the radial momentum equation is  $p - p_0 = \frac{1}{2} \rho \omega^2 (r^2 - r_0^2)$  or

$$p - p_0 = \frac{1}{2} \rho [v(r)]^2 \left[ 1 - \left( \frac{r_0}{r} \right)^2 \right].$$



**Fig. A2.3-1** Geometry of the test problem *rigid body steady rotation*

*Pure radial symmetric flow:* This test problem presents a hollow cylinder with symmetric flow in the radial direction, see Fig. A2.3-2.



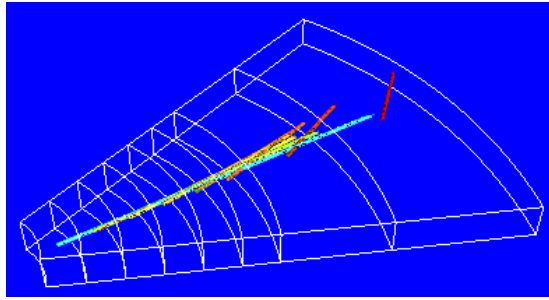
**Fig. A2.3-2** Geometry of the test problem *pure radial symmetric flow*

No axial and azimuthal flow exists. The mass conservation equation gives  $\frac{\partial}{\partial r}(ru) = 0$ . The  $r$  direction momentum equation simplifies to  $\rho \frac{1}{2} \frac{\partial u^2}{\partial r} = -\frac{\partial p}{\partial r}$ , and the  $\theta$  direction momentum equation gives  $\frac{\partial p}{\partial \theta} = 0$ . From the mass conservation we have  $u = u_0 \frac{r_0}{r}$ . The analytical solution of the radial momentum equation is the well-known *Bernoulli* equation  $p - p_0 = -\frac{1}{2} \rho (u^2 - u_0^2)$  or

$$p - p_0 = \frac{1}{2} \rho u_0^2 \left[ 1 - \left( \frac{r_0}{r} \right)^2 \right].$$

*Radial-azimuthal symmetric flow:* This test problem presents a hollow cylinder with symmetric flow in the radial and azimuthal directions – it is, in fact, a superposition of the previous two cases, rigid body steady rotation and pure radial symmetric flow, see Fig. A2.3-3. No axial flow exists. The mass conservation eq-

uation gives  $\frac{\partial}{\partial r}(ru) = 0$ . The  $r$  direction momentum equation simplifies to  $\rho \left( \frac{1}{2} \frac{\partial u^2}{\partial r} - \frac{v^2}{r} \right) = -\frac{\partial p}{\partial r}$ , and the  $\theta$ -direction momentum equation gives  $\rho u \left( \frac{\partial v}{\partial r} + \frac{v}{r} \right) = -\frac{1}{r} \frac{\partial p}{\partial \theta}$ . From the mass conservation we have  $u = u_0 \frac{r_0}{r}$ . From the azimuthal symmetry,  $\frac{\partial p}{\partial \theta} = 0$ , the  $\theta$  direction momentum equation gives  $\frac{\partial v}{\partial r} + \frac{v}{r} = 0$  or  $v = v_0 \frac{r_0}{r}$ . Taking into account the both solutions of the mass and of the  $\theta$  momentum equation the radial momentum equation gives  $\rho \frac{1}{r} (u^2 + v^2) = \frac{\partial p}{\partial r}$  or  $\rho (u_0^2 + v_0^2) r_0^2 \frac{1}{r^3} = \frac{\partial p}{\partial r}$  or

$$p - p_0 = \rho (u_0^2 + v_0^2) r_0^2 \frac{1}{2} \left( \frac{1}{r_0^2} - \frac{1}{r^2} \right).$$


**Fig. A2.3-3** Geometry of the test problem *radial-azimuthal symmetric flow*

*Trajectory of particles in a known gas field:* Consider flow of particles with very small concentrations in a known gas velocity field. Compute the trajectory of a particle with mass  $m_i$  taking into account only the drag force and assuming the validity of *Stokes' law*. Such a task was usually solved in the past for computing trajectories of particles in cyclone separators. If in such computations, the trajectory ends in the particle capturing devices, this class characterized by particle size and starting coordinate is considered as removed from the gas flow. The three simplified momentum equations are then

$$\rho_i \frac{\pi D_{3,i}^3}{6} \left( \frac{du_{3,i}}{d\tau} - \frac{v_{3,i}^2}{r} \right) = 3\pi\eta_1 D_{3,i} (u_1 - u_{3,i}), \quad \frac{du_{3,i}}{d\tau} = \frac{v_{3,i}^2}{r} + \frac{18\eta_1}{\rho_i D_{3,i}^2} (u_1 - u_{3,i}),$$

$$\rho_i \frac{\pi D_{3,i}^3}{6} \left( \frac{dv_{3,i}}{d\tau} + \frac{v_{3,i} u_{3,i}}{r} \right) = 3\pi\eta_1 D_{3,i} (v_1 - v_{3,i}), \quad \frac{dv_{3,i}}{d\tau} = -\frac{v_{3,i} u_{3,i}}{r} + \frac{18\eta_1}{\rho_i D_{3,i}^2} (v_1 - v_{3,i}),$$

$$\rho_i \frac{\pi D_{3,i}^3}{6} \frac{dw_{3,i}}{d\tau} = 3\pi\eta_1 D_{3,i} (w_1 - w_{3,i}), \quad \frac{dw_{3,i}}{d\tau} = \frac{18\eta_1}{\rho_i D_{3,i}^2} (w_1 - w_{3,i}).$$

For constant gas velocity and particle size the analytical solution is provided by *Crowe and Patt (1974)*

$$\begin{aligned} w_{3,i} &= w_1 - (w_1 - w_{3,i,a}) \exp(-\Delta\tau/\Delta\tau_{13}), \\ u_{3,i} &= u_1 - (u_1 - u_{3,i,a}) \exp(-\Delta\tau/\Delta\tau_{13}) + [1 - \exp(-\Delta\tau/\Delta\tau_{13})] \Delta\tau_{13} v_{3,i}^2 / r, \\ v_{3,i} &= v_1 - (v_1 - v_{3,i,a}) \exp(-\Delta\tau/\Delta\tau_{13}) - [1 - \exp(-\Delta\tau/\Delta\tau_{13})] \Delta\tau_{13} v_{3,i} u_{3,i} / r, \end{aligned}$$

where

$$\Delta\tau_{13} = (\rho_3 D_{3,i}^2) / (18\eta_1)$$

is the *Stokes* relaxation time constant. Knowing the initial position and the velocity, the position after the time interval  $\Delta\tau$  can be computed by using the *Euler* method.

*The Kreith and Sonju solution for the decay of turbulent swirl in a pipe: Kreith and Sonju (1965)* analyzed steady turbulent swirl in a pipe. After making several reasonable simplifying assumptions the authors arrived at the following form of the tangential momentum equation:

$$w \frac{\partial v}{\partial z} = (v + v') \left( \frac{\partial^2 v}{\partial r^2} + \frac{1}{r} \frac{\partial v}{\partial r} - \frac{v}{r^2} \right),$$

or in nondimensional form

$$\frac{\partial \bar{v}}{\partial \bar{z}} = \frac{\bar{v}}{\text{Re}} \left( \frac{\partial^2 \bar{v}}{\partial \bar{r}^2} + \frac{1}{\bar{r}} \frac{\partial \bar{v}}{\partial \bar{r}} - \frac{\bar{v}}{\bar{r}^2} \right),$$

where  $\bar{v} = v/w_{\max}$ ,  $\bar{r} = r/R$ ,  $\bar{v} = (v + v')/\nu$ ,  $\bar{z} = z/R$ , and  $\text{Re} = wR/\nu$ . The authors solved this equation by separation of the variables for the following boundary conditions:  $v = 0$  at  $r = 0$  and  $r = R$ ,  $v(r, 0) = f(r)$  at  $z = 0$ . The initial condition was gained from experimental data for the initial distribution of the tangential velocity behind a tape swirler

$$\bar{v}(\bar{r}, 0) = \left[ 6.3\bar{r} - 0.013(1.1 - \bar{r})^{-2.68} \right] / \bar{\Delta z}_{ts},$$

where  $\bar{\Delta z}_{ts} = \Delta z_{ts}/R$ ,  $\Delta z_{ts}$  is the length of the tape swirler making a complete 360° rotation.

$$\begin{aligned} \bar{v}(\bar{r}, \bar{z}) &= \frac{7.78}{H} J_1(3.832\bar{r}) \exp\left(-16.7 \frac{\bar{v}\bar{z}}{\text{Re}}\right) - \frac{5.26}{H} J_1(7.016\bar{r}) \exp\left(-55.7 \frac{\bar{v}\bar{z}}{\text{Re}}\right) \\ &+ \frac{3.93}{H} J_1(10.174\bar{r}) \exp\left(-117.9 \frac{\bar{v}\bar{z}}{\text{Re}}\right) - \frac{3.16}{H} J_1(13.324\bar{r}) \exp\left(-203.7 \frac{\bar{v}\bar{z}}{\text{Re}}\right) + \dots \end{aligned}$$

$J_1$  is the *Bessel's* function of the first kind of order one. From experimental data the relation  $\bar{v} = 1 + 2.03 \times 10^{-3} \text{Re}^{0.86}$  was recommended for  $4 \times 10^4 < \text{Re} < 1 \times 10^6$ . Experimental data for  $\text{Re} = 18\,000$  and  $61\,000$  validate the approximate solution.

The data indicate the initial swirl decay to be about 20% at  $\bar{z} = 100$ . There are authors trying to represent the decay by a single exponential function. From the data collected by *Steenberger* (1995) it is visible that the decay coefficient is a decreasing function with increasing Reynolds number as manifested by the above solution. Note the practical importance of this solution. Having the rotation introduced by twisted tapes in the cylinder particle trajectories can be computed and, therefore, the efficiency of separation devices can be judged.

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## 3 Derivatives for the equations of state

*Derivation of partial derivatives for isothermal multi-component mixtures needed for development of universal models for multi-phase multi-component flows is presented. The equations of state and the derivative approximations as functions of temperature and pressure for the elementary mixture constituents are assumed to be known. The so called universal fluid is introduced consisting of an arbitrary number of miscible and non-miscible components. This fluid model describes in its limiting cases gas, or gas mixture, or liquid, or solution of liquids, or gas-liquid solutions with an arbitrary number of gaseous and liquid components, or gas-liquid solutions containing immiscible liquid or solid particles. In addition, one component liquid-gas and solid-liquid equilibrium mixtures are considered.*

### 3.1 Introduction

Numerical modeling of complicated physical phenomena such as multi-component multi-phase flows is a powerful tool supplementing experiments and enabling optimum design of complicated technical facilities. The wide range of computer codes developed over the past 30 years for the description of multi-dimensional single-, two- and multiphase flows inevitably leads to the step of developing a universal flow analyzer. Such a computer code should model transient and steady-state three-dimensional flows in a complicated geometry with arbitrary internals. The flow should be described by multi-velocity fields, each of them consisting of an arbitrary number of chemical components.

The local volume- and time-averaged fundamental equations are derived by applying the so-called instantaneous equations inside the velocity field and averaging these over space and time. But even the instantaneous conservation equations for mass, momentum and energy are averages too in the sense that the motions of the individual molecules are averaged. This leads to loss of information on the thermodynamic behavior of the system. The information lost must be provided by the state and transport equations which already incorporate the averaging procedure by virtue of their derivation.

For a mathematical description of the flow, time and three space coordinates are generally used as independent variables, with a set of time- and space-dependent variables, e.g.  $p$  and  $T$  as dependent variables. Besides the velocities that describe the flow, there are dependent variables that describe the thermodynamic state of the particular velocity field. This group of variables has to consist of mutually independent variables, e.g.  $(p, T)$ , or  $(p, h)$ , or  $(p, s)$ , or  $(p, \rho)$ , etc., see for example, *Kestin* (1979). Some valuable references are also given in Appendix 3.2.

All other properties of the velocity field are a unique function of these mutually independent variables. Analytical or tabular approximations of the equation of state and of the transport properties as a function of pressure ( $p$ ) and temperature ( $T$ ) are in general use. As already mentioned, the set of dependent variables can be transformed from  $(p, T)$  to another set, e.g.  $(p, h)$  or  $(p, s)$ , and the system of PDEs can be rearranged in terms of the chosen set. From the large variety of systems that result, only one has the remarkable quality of being the simplest. Interestingly enough, it is just this system which is obtained if the specific entropy of the velocity fields is used as an element of the dependent variable vector. This will be demonstrated in Chapters 4 and 5. An interesting dilemma thus arises: on the one hand, the desire to integrate the simplest possible system of PDEs in order to save computer time, to reduce errors during code development, and to concentrate on other related models incorporated into the code; on the other, the desire to use an existing library of analytical approximations for the equation of state and for the transport equation as a function of pressure and temperature. These two tendencies are not contradictory. A solution is easily obtained if the equation of state is used in the form

$$T = T(p, s), \quad dT = \left( \frac{\partial T}{\partial p} \right)_s dp + \left( \frac{\partial T}{\partial s} \right)_p ds \quad (3.1)$$

and

$$\rho = \rho(p, s), \quad d\rho = \left( \frac{\partial \rho}{\partial p} \right)_s dp + \left( \frac{\partial \rho}{\partial s} \right)_p ds \quad (3.2)$$

where the derivatives  $(\partial T / \partial p)_s$ ,  $(\partial T / \partial s)_p$ ,  $(\partial \rho / \partial p)_s$ , and  $(\partial \rho / \partial s)_p$  are functions of  $(p, T)$ . The derivatives  $(\partial \rho / \partial p)_s$  and  $(\partial \rho / \partial s)_p$  are needed for development of the very important link between the density change and the change in the other dependent variables describing the flow. This link is, at the same time, the link between the mass, momentum, and energy-conservation equations. This construct for the solution methods is called the *entropy concept*.

The results presented can be used in the framework of the entropy concept, or in any other concepts for development of solution methods in the field of multiphase fluid dynamics. The expressions used in the IVA computer code family *Ko-lev* (1994a, 1995, 1998) are the simple cases for two components of the general expressions presented in this chapter for the  $n$ -component. Complete implementation of the theory for computational analyses was performed by this author in *Ko-lev* (1996, 1999, 2002).

### 3.2 Multi-component mixtures of miscible and non-miscible components

Every fluid that seems to be pure in nature is in fact a mixture of substances. The idealization pure applied to a fluid is very helpful in science. In many cases replacing mixtures of one predominant fluid with some traces of other substances is a very good approximation to work with. But there are a lot of other applications where even these small “impurities” substantially control processes. Examples are the initiation of phase transitions in meta-stable liquids and gases – the so called heterogeneous nucleation process starting at proffered “impurities” such as fine gas bubbles or solid particles. Another example is the solution and dissolution process of gases. In some technological applications such processes may lead to explosive gas mixtures which under some circumstances may ignite and destroy the facility.

The results obtained in *Kolev* (1986a, b) for mixtures consisting of one real and one ideal gas were generalized in *Kolev* (1991) for mixtures with an unlimited number of chemical gas components. Chapter 3.2 in *Kolev* (2002) was an extended version of *Kolev* (1986a, b). Here we extend the theory once again for mixtures consisting simultaneously of miscible and non-miscible chemical constituents.

Consider a mixture of  $i_{\max} = m_{\max} + n_{\max}$  components, where  $m_{\max}$  is the number of the miscible components and  $n_{\max}$  is the number of the non-miscible components.

Examples of miscible components are:

- a) Gases inside a gas mixture;
- b) Gases dissolved in liquid;
- c) Liquids dissolved in liquid.

In this case each component occupies the total volume occupied by the mixture itself. The volumetric fraction of each miscible component  $\alpha_i$  in the multiphase flow is therefore equal to the volume fraction of the corresponding velocity field,

$\alpha$ , and the definition of the density of the mixture  $\alpha\rho = \sum_{i=1}^{i_{\max}} \alpha_i \rho_i$  simplifies to

$\rho = \sum_{i=1}^{i_{\max}} \rho_i$ . As a result, the definition of the mass concentrations simplifies to

$C_i = \alpha_i \rho_i / (\alpha\rho) = \rho_i / \rho$ . The total pressure  $p$  is then the sum of the partial pressures of the miscible components,  $p = \sum_{i=1}^{m_{\max}} p_i$ . This relation expresses the law of

*Dalton*.

Examples of non-miscible components are:

- a) Mixtures of non-miscible liquids;
- b) Mixture of solid particles and liquid.

As already said, in a mixture consisting of miscible components, each component occupies the total volume occupied by the mixture itself. This is not the case for

mixtures consisting of non-miscible components. The volume fraction occupied by each species differs from the volume fraction occupied by the mixture itself.

The definition of the mixture density is then  $\left(\sum_{i=1}^{i_{\max}} \alpha_i\right) \rho = \sum_{i=1}^{i_{\max}} \alpha_i \rho_i$ , and the definition of the mass concentrations therefore

$$C_i = \alpha_i \rho_i / \left(\sum_{i=1}^{i_{\max}} \alpha_i \rho_i\right) = \alpha_i \rho_i / \left[\left(\sum_{i=1}^{i_{\max}} \alpha_i\right) \rho\right]$$

does not simplify, as does that for gas mixtures.

The non-miscible components experience the same total pressure which is equal to the total mixture pressure  $p_n = p$ , where  $n = m_{\max} + 1, m_{\max} + n_{\max}$ .

Examples of mixture of miscible and non-miscible fluids are in fact all fluids in the nature such us:

- a) Liquid water containing dissolved gases and impurities such as solid particles of different chemical components;
- b) Lava consisting of several molten components containing dissolved gases and solid particles from other than the molten species;
- c) Air containing microscopic impurities such as dust.

There are many examples of dramatic events happening in nature and in technology due to the release of the dissolved components inside mixtures:

- a) Volcanic explosions initiated first by pressure release and followed by a dramatic release of the dissolved gases;
- b) Choked flow in propulsion machines caused by gas release due to a pressure drop;
- c) Hydrogen and oxygen release and accumulation in nuclear power plant pipelines that could be ignited and cause explosive damage etc.

### 3.2.1 Computation of partial pressures for known mass concentrations, system pressure and temperature

*Task definition.* The mass concentration of each constituent, designated  $C_i$  is

$$C_i := \rho_i / \sum_{m=1}^{i_{\max}} \rho_m = \rho_i / \rho, \quad (3.3)$$

where

$$\frac{1}{\rho} = \frac{1 - \sum_{n=m_{\max}+1}^{i_{\max}} C_n}{\sum_{m=1}^{m_{\max}} \rho_m(p_m, T)} + \sum_{n=m_{\max}+1}^{i_{\max}} \frac{C_n}{\rho_n(p, T)}, \quad (3.4)$$

is the mixture density. Equations (3.3) and (3.4) are valid for mixture consisting simultaneously of both miscible and non-miscible components. The system pressure, temperature of the mixture and mass concentration of each constituent,

$$p, T, C_{1,2,\dots,i_{\max}-1}, C_{i_{\max}} = 1 - \sum_{k=1}^{i_{\max}-1} C_k$$

are, by definition, known. The partial pressure of each particular miscible component is sought.

**Solution.** We start with the  $m_{\max} - 1$  definition equations for the mass concentrations,

$$\rho_i(p_i, T) = C_i \rho, \quad i = 1, m_{\max} - 1, \tag{3.5}$$

and with Dalton's law valid for the miscible components (i.e., the sum of the partial pressures = system pressure)

$$p = \sum_{m=1}^{m_{\max}} p_m. \tag{3.6}$$

In the general case of real fluids, Eqs. (3.5) and (3.6) are a system of non-linear algebraic equations for the unknowns  $p_m, m = 1, m_{\max}$ , which can be solved numerically by iterations. Before showing this solution we will first give the well-known solution for which the miscible components are perfect fluids. It can be used as a first approximation for the numerical solution.

*The solution if the miscible components are perfect fluids.* For miscible fluids assumed to be perfect we have

$$\rho_i = \frac{p_i}{R_i T}. \tag{3.7}$$

This simplifies the system (3.5), (3.6) considerably. We first substitute the mixture density in Eq. (3.5) using (3.4). Then we substitute each of the  $m$  densities of the miscible fluids using (3.7) in the resulting equation. After multiplying both sides of the resulting equation by  $T$  we finally obtain

$$\frac{p_i}{R_i} = C_i \left[ \frac{1 - \sum_{n=m_{\max}+1}^{i_{\max}} C_n}{\sum_{i=m}^{m_{\max}} \frac{p_m}{R_m}} + \frac{1}{T} \sum_{n=m_{\max}+1}^{i_{\max}} \frac{C_n}{\rho_n(p, T)} \right]^{-1}, \quad i = 1, m_{\max} - 1. \tag{3.8}$$

Now the system (3.6) and (3.8) can be solved analytically for the unknown partial pressures in the following way. Select the component,  $m_{\max}$ , whose pressure  $p_{m_{\max}}$  must be calculated first. Add the  $m_{\max} - 1$  equations (3.8)



$$\sum_{k=1}^{m_{\max}-1} p_k = \sum_{k=1}^{m_{\max}-1} R_k C_k \left[ \frac{1 - \sum_{n=m_{\max}+1}^{i_{\max}} C_n}{\sum_{i=m}^{m_{\max}} \frac{p_m}{R_m}} + \frac{1}{T} \sum_{n=m_{\max}+1}^{i_{\max}} \frac{C_n}{\rho_n(p, T)} \right]^{-1}, \quad (3.9)$$

rewrite Eq. (3.6) in the following form

$$p - p_{m_{\max}} = \sum_{k=1}^{m_{\max}-1} p_k, \quad (3.10)$$

and replace the sum of the left-hand side of Eq. (3.9) by Eq. (3.10) to obtain

$$p - p_{m_{\max}} = \sum_{k=1}^{m_{\max}-1} R_k C_k \left[ \frac{1 - \sum_{n=m_{\max}+1}^{i_{\max}} C_n}{\sum_{i=m}^{m_{\max}} \frac{p_m}{R_m}} + \frac{1}{T} \sum_{n=m_{\max}+1}^{i_{\max}} \frac{C_n}{\rho_n(p, T)} \right]^{-1}. \quad (3.11)$$

Replace the sum  $\sum_{m=1}^{m_{\max}} \frac{p_m}{R_m}$  in this equation by the definition equation (3.3) for  $C_{m_{\max}}$

$$\sum_{m=1}^{m_{\max}} \frac{p_m}{R_m} = \frac{p_{m_{\max}}}{C_{m_{\max}} R_{m_{\max}}}, \quad (3.12)$$

to obtain the quadratic equation

$$\begin{aligned} & \frac{p_{m_{\max}}^2}{T} \sum_{n=m_{\max}+1}^{i_{\max}} \frac{C_n}{\rho_n(p, T)} \\ & + \left[ \sum_{k=1}^{m_{\max}} C_k R_k - C_{m_{\max}} R_{m_{\max}} \sum_{n=m_{\max}+1}^{i_{\max}} C_n - \frac{p}{T} \sum_{n=m_{\max}+1}^{i_{\max}} \frac{C_n}{\rho_n(p, T)} \right] p_{m_{\max}} \\ & - \left( 1 - \sum_{n=m_{\max}+1}^{i_{\max}} C_n \right) C_{m_{\max}} R_{m_{\max}} p = 0. \end{aligned} \quad (3.13)$$

The solution of this equation gives the result we are looking for. Analogously, we compute each of the other partial pressures of the miscible components.

*Limiting case for no non-miscible components.* For the limiting case of no non-miscible components in the mixture the solution is

$$p_{m_{\max}} = \frac{R_{m_{\max}} C_{m_{\max}}}{\sum_{k=1}^{m_{\max}} R_k C_k} p = Y_{m_{\max}} p, \quad (3.14)$$

see *Elsner* (1974) for comparison. Analogously, we compute each of the other partial pressures

$$p_i = Y_i p, \quad (3.15)$$

where

$$Y_i = \frac{R_i C_i}{\sum_{k=1}^{i_{\max}} R_k C_k} = \frac{C_i / M_i}{\sum_{k=1}^{i_{\max}} C_k / M_k} \quad (3.16)$$

is the *molar concentration* of the  $i$ -th component of the gas mixture. Here  $M_i$  is *kg-mole* of the constituent  $i$ . Note that for a description of transport processes without chemical reactions the use of the mass concentrations is much more convenient than the use of the molar concentrations. For perfect fluids the reverse computation of the mass concentrations if the *kg-mole* concentrations are known is useful

$$C_i = \frac{Y_i M_i}{\sum_{k=1}^{i_{\max}} Y_k M_k}. \quad (3.17)$$

*The general solution.* In reality, the mixture does not consist of perfect fluids but of real fluids. As a result, the solution already obtained can be used as a first approximation for the exact solution of the system (3.5) and (3.6), which then must be obtained by iteration. For this purpose, we use the standard method of *Newton-Raphson*, seeking the zeros for the following functions

$$\varepsilon_i = \rho_i(p_i, T) - C_i \rho, \quad i = 1, m_{\max} - 1, \quad (3.18)$$

$$\varepsilon_{m_{\max}} = p - \sum_{m=1}^{m_{\max}} p_m. \quad (3.19)$$

Small deviations in  $p_m$ ,  $\delta p_m$ ,  $m = 1, m_{\max} - 1$  lead to a corresponding deviation in  $\varepsilon_m$

$$\delta \varepsilon_i = \left( \frac{\partial \rho_i}{\partial p_i} \right)_T \delta p_i - C_i \delta \rho, \quad i = 1, m_{\max} - 1, \quad (3.20)$$

and

$$\delta \varepsilon_{m_{\max}} = - \sum_{m=1}^{m_{\max}} \delta p_m. \quad (3.21)$$

Further to this, the task is reduced to finding such a  $\delta p_m$ , which for a given initial approximation  $p_m$ , and

$$\left( \frac{\partial p_m}{\partial p_m} \right)_T = f(p_m, T) \quad (3.22)$$

minimizes the residuals of the functions, namely,

$$\delta \varepsilon_m = -\varepsilon_m(p_m, T) \quad (3.23)$$

or

$$\mathbf{J} \overline{\delta p} = -\mathbf{e} \quad (3.24)$$

or

$$\overline{\delta p} = -\mathbf{J}^{-1} \mathbf{e}, \quad (3.25)$$

where the algebraic vector of the pressure increments is

$$\overline{\delta p} = \begin{pmatrix} p_1^{n+1} - p_1^n \\ p_2^{n+1} - p_2^n \\ \dots \\ p_{m_{\max}}^{n+1} - p_{m_{\max}}^n \end{pmatrix}, \quad (3.26)$$

and the residuals

$$\mathbf{e} = \begin{pmatrix} \rho_1(p_1, T) - C_1 \rho \\ \rho_2(p_2, T) - C_2 \rho \\ \dots \\ p - \sum_{m=1}^{m_{\max}} p_m \end{pmatrix}. \quad (3.27)$$

Knowing the residuals and the *Jacobian* in the previous iteration step  $n$ , the solution improves as follows

$$\overline{p}^{n+1} = \overline{p}^n - (\mathbf{J}^n)^{-1} \mathbf{e}. \quad (3.28)$$

Because  $p$  and  $T$  are constant the term  $\sum_{i=m_{\max}+1}^{i_{\max}} \frac{C_i}{\rho_i(p, T)}$  in the density expression is also a constant. Therefore

$$d\rho = \frac{\left(1 - \sum_{i=m_{\max}+1}^{i_{\max}} C_i\right) \rho^2}{\left[\sum_{i=1}^{m_{\max}} \rho_i(p_i, T)\right]^2} \sum_{i=1}^{m_{\max}} d\rho_i(p_i, T), \quad (3.29)$$

and the *Jacobian* takes the form

$$J_{kj} = \partial \varepsilon_k / \partial p_j = \left\{ \delta_{kj} - C_k \frac{\left(1 - \sum_{i=m_{\max}+1}^{i_{\max}} C_i\right) \rho^2}{\left[\sum_{i=1}^{m_{\max}} \rho_i(p_i, T)\right]^2} \right\} \left( \frac{\partial \rho_k}{\partial p_k} \right)_T \quad \text{for } k=1, \dots, m_{\max} - 1, \quad (3.30)$$

$$J_{m_{\max}j} = -1. \quad (3.31)$$

$\delta_{kj}$  in Eq. (3.30) is the *Kroneker* delta function ( $\delta_{kj} = 0$  for  $j \neq k$ , and  $\delta_{kj} = 1$  otherwise). For non-existing non-miscible components Eq. (3.30) reduces to Eq. (3.30) in *Kolev (2002)*. If one part of the components is taken to be a perfect gas, the corresponding derivatives in the *Jacobian* are simply

$$\left( \frac{\partial \rho_i}{\partial p_i} \right)_T = \frac{1}{R_i T}. \quad (3.32)$$

But in the general case

$$\left( \frac{\partial \rho_i}{\partial p_i} \right)_T = \frac{1}{R_i(p_i, T) T}, \quad (3.33)$$

where

$$R_i(p_i, T) = 1 / \left[ T \left( \frac{\partial \rho_i}{\partial p_i} \right)_T \right] \quad (3.34)$$

is not a constant, but a function of  $p_i$  and  $T$  because

$$T \left( \frac{\partial \rho_i}{\partial p_i} \right)_T = f(p_i, T). \quad (3.35)$$

Using  $(p, T, \rho_i)$  instead of  $(p, T, C_i)$  as dependent variables we eliminate the need for computing the partial pressure by iterations because we have directly

$p_i = p_i(\rho_i, T)$ . This approach has its limitation if used with digital computers due to the so called truncation error. Thus below a given *Mach* number

$$Ma_1 < 10^{(n1+n2)/2} \quad (3.36)$$

the density change is then below the truncation error for computers

$$\frac{\Delta\rho_1}{\rho_{10}} \approx 10^{n1} \quad (3.37)$$

and cannot influence the pressure field any more below the value,

$$\frac{\Delta p}{\rho_{10} w_{10}^2} \approx 10^{n2} \quad (3.38)$$

due to  $\Delta p \approx a^2 \Delta \rho$ , (3.39)

$$\frac{\Delta p}{\rho_{10} w_{10}^2} \approx \frac{\Delta \rho_1}{\rho_{10}} / Ma_1^2, \quad (3.40)$$

see *Issa* (1983). More recent discussion of this topic is given in *Sesternhenn* et al. (1999).

Thus, after  $p$ ,  $C_i$  and  $T$  are known, the partial pressure and therefore the densities can be calculated as

$$\rho_i = \rho_i(p_i, T) \quad (3.41)$$

as can the partial derivatives for each of the components  $\left(\frac{\partial \rho_i}{\partial p_i}\right)_T$  and  $\left(\frac{\partial \rho_i}{\partial T}\right)_{p_i}$ .

For a perfect gas the densities and their derivatives can easily be computed from the state equation  $\rho_i = \rho_i(p_i, T)$ . For real gases, it is assumed that the analytical expression for  $\rho_i = \rho_i(p_i, T)$  is known. It is therefore easy to derive analytical expressions to compute the corresponding derivatives. The differential form of the equation of state

$$\rho_i = \rho_i(p_i, T), \quad (3.42)$$

namely,

$$d\rho_i = \left(\frac{\partial \rho_i}{\partial p_i}\right)_T dp_i + \left(\frac{\partial \rho_i}{\partial T}\right)_{p_i} dT, \quad (3.43)$$

for each of the components is uniquely defined.

### 3.2.2 Partial derivatives of the equation of state $\rho = \rho(p, T, C_{2, \dots, i_{\max}})$

The mixture density

$$\rho = \left[ \frac{1 - \sum_{n=m_{\max}+1}^{i_{\max}} C_n}{\sum_{m=1}^{m_{\max}} \rho_m(p_m, T)} + \sum_{n=m_{\max}+1}^{i_{\max}} \frac{C_n}{\rho_n(p, T)} \right]^{-1} = \rho(C_{1,2, \dots, i_{\max}}, p, T) \quad (3.44)$$

is obviously a function of  $C_{1,2, \dots, i_{\max}}$ ,  $p$  and  $T$ . Note that only  $i_{\max} - 1$  concentrations are mutually independent due to the fact that

$$\sum_{i=1}^{i_{\max}} C_i = 1. \quad (3.45)$$

Consequently,  $\rho$  is simply a function of only  $i_{\max} - 1$  mutually independent concentrations,

$$\rho = \rho(C_{2, \dots, i_{\max}}, p, T). \quad (3.46)$$

Further we solve the following task: Let us assume that  $C_{2, \dots, i_{\max}}$ ,  $p$  and  $T$  are given, from which, as already shown, all  $p_i, \rho_i, \left(\frac{\partial \rho_i}{\partial p_i}\right)_T, \left(\frac{\partial \rho_i}{\partial T}\right)_{p_i}$  and  $\rho$ , respectively can be calculated. The expressions defining the partial derivatives for the mixture in the differential form of the equation of state (3.46) are sought:

$$d\rho = \left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C\text{'s}} dp + \left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C\text{'s}} dT + \sum_{i=2}^{i_{\max}} \left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all } C\text{'s}_{\text{except } C_i}} dC_i \quad (3.47)$$

The partial derivatives  $\left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C\text{'s}}$ ,  $\left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C\text{'s}}$  and  $\left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all } C\text{'s}_{\text{except } C_i}}$  fol-

low immediately from the differential form of the equation  $p = \sum_{m=1}^{m_{\max}} p_m$ ,

$$dp = \sum_{m=1}^{m_{\max}} d[p_m(\rho_m, T)] = \sum_{m=1}^{m_{\max}} d[p_m(C_m, \rho, T)], \quad (3.48)$$

where

$$dp_m = \left(\frac{\partial p_m}{\partial \rho_m}\right)_T C_m d\rho + \rho \left(\frac{\partial p_m}{\partial \rho_m}\right)_T dC_m + \left(\frac{\partial p_m}{\partial T}\right)_{\rho_m} dT, \quad (3.49)$$

or

$$\begin{aligned}
 dp &= \sum_{m=1}^{m_{\max}} d[p_m(\rho_m, T)] = \sum_{m=1}^{m_{\max}} d[p_m(C_m \rho, T)] \\
 &= \sum_{m=1}^{m_{\max}} \left[ \left( \frac{\partial p_m}{\partial \rho_m} \right)_T (C_m d\rho + \rho dC_m) + \left( \frac{\partial p_m}{\partial T} \right)_{\rho_m} dT \right] \\
 &= \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial \rho_i} \right)_T C_m \right] d\rho + \rho \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial \rho_m} \right)_T dC_m + \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial T} \right)_{\rho_m} \right] dT \quad (3.50)
 \end{aligned}$$

or after substituting

$$dC_1 = -\sum_{i=2}^{i_{\max}} dC_i = -\sum_{i=2}^{m_{\max}} dC_i - \sum_{i=m_{\max}+1}^{i_{\max}} dC_i \quad (3.51)$$

$$\begin{aligned}
 dp &= \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial \rho_m} \right)_T C_m \right] d\rho + \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial T} \right)_{\rho_m} \right] dT \\
 &+ \rho \sum_{m=2}^{m_{\max}} \left[ \left( \frac{\partial p_m}{\partial \rho_m} \right)_T - \left( \frac{\partial p_1}{\partial \rho_1} \right)_T \right] dC_m - \rho \left( \frac{\partial p_1}{\partial \rho_1} \right)_T \sum_{n=m_{\max}+1}^{i_{\max}} dC_n \quad (3.52)
 \end{aligned}$$

namely,

$$\left( \frac{\partial p}{\partial \rho} \right)_{T, \text{all } C\text{'s}} = 1 / \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial \rho_m} \right)_T C_m \right], \quad (3.53)$$

$$\left( \frac{\partial p}{\partial T} \right)_{\rho, \text{all } C\text{'s}} = - \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial T} \right)_{\rho_m} \right] / \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial \rho_m} \right)_T C_m \right], \quad (3.54)$$

$$\left( \frac{\partial p}{\partial C_i} \right)_{\rho, T, \text{all } C\text{'s except } C_i} = -\rho \left[ \left( \frac{\partial p_i}{\partial \rho_i} \right)_T - \left( \frac{\partial p_1}{\partial \rho_1} \right)_T \right] / \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial \rho_m} \right)_T C_m \right], \quad (3.55)$$

for  $i = 2, m_{\max}$ , and

$$\left( \frac{\partial p}{\partial C_i} \right)_{\rho, T, \text{all } C\text{'s except } C_i} = \rho \left( \frac{\partial p_1}{\partial \rho_1} \right)_T / \left[ \sum_{m=1}^{m_{\max}} \left( \frac{\partial p_m}{\partial \rho_m} \right)_T C_m \right], \quad (3.56)$$

for  $i = m_{\max} + 1, i_{\max}$ . Here an arbitrary existing component is denoted with 1,  $C_1 > 0$ . The equations above contain derivatives  $\left(\frac{\partial p_i}{\partial p_i}\right)_T$ ,  $\left(\frac{\partial p_i}{\partial T}\right)_{p_i}$ . Only the derivatives  $(\partial p_i / \partial p_i)_T$  and  $(\partial p_i / \partial T)_{p_i}$  of the simple components are known, so that the expressions for the mixture derivatives (3.53) through (3.56) in which the known component derivatives explicitly occur are easily obtained by replacing

$$\left(\frac{\partial p_i}{\partial p_i}\right)_T = 1 / \left(\frac{\partial p_i}{\partial p_i}\right)_T, \quad (3.57)$$

$$\left(\frac{\partial p_i}{\partial T}\right)_{p_i} = -\left(\frac{\partial p_i}{\partial T}\right)_{p_i} / \left(\frac{\partial p_i}{\partial p_i}\right)_T. \quad (3.58)$$

The final result is

$$\left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C\text{'s}} = 1 / \left[ \sum_{m=1}^{m_{\max}} C_m / \left(\frac{\partial \rho_m}{\partial p_m}\right)_T \right], \quad (3.59)$$

$$\left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C\text{'s}} = \left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C\text{'s}} \sum_{m=1}^{m_{\max}} \left(\frac{\partial \rho_m}{\partial T}\right)_{p_m} / \left(\frac{\partial \rho_m}{\partial p_m}\right)_T, \quad (3.60)$$

$$\left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all } C\text{'s except } C_i} = -\rho \left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C\text{'s}} \left[ 1 / \left(\frac{\partial p_i}{\partial p_i}\right)_T - 1 / \left(\frac{\partial p_i}{\partial p_i}\right)_T \right], \quad (3.61)$$

for  $i = 2, m_{\max}$ , and

$$\left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all } C\text{'s except } C_i} = \rho \left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C\text{'s}} / \left(\frac{\partial p_i}{\partial p_i}\right)_T, \quad (3.62)$$

for  $i = m_{\max} + 1, i_{\max}$ . With these results Eq. (3.49) can be rewritten as a function of the pressure, temperature and concentrations



$$\begin{aligned}
 dp_m = & \frac{C_m \left( \frac{\partial p_m}{\partial \rho_m} \right)_T}{\sum_{i=1}^{m_{\max}} C_i \left( \frac{\partial p_i}{\partial \rho_i} \right)_T} dp + \frac{1}{\left( \frac{\partial p_m}{\partial \rho_m} \right)_T} \left[ C_m \frac{\sum_{i=1}^{m_{\max}} \left( \frac{\partial \rho_i}{\partial T} \right)_{p_i} / \left( \frac{\partial \rho_i}{\partial p_i} \right)_T - \left( \frac{\partial \rho_m}{\partial T} \right)_{p_m}}{\sum_{i=1}^{m_{\max}} C_i \left( \frac{\partial p_i}{\partial \rho_i} \right)_T} \right] dT \\
 & + \rho \left( \frac{\partial p_m}{\partial \rho_m} \right)_T \left[ dC_m - \frac{C_m}{\sum_{i=1}^{m_{\max}} \left( \frac{\partial p_i}{\partial \rho_i} \right)_T} \sum_{i=1}^{m_{\max}} \left( \frac{\partial p_i}{\partial \rho_i} \right)_T dC_i \right], \quad (3.63)
 \end{aligned}$$

or

$$dp_m = \left( \frac{\partial p_m}{\partial p} \right)_{T, \text{all\_}C\text{'s}} dp + \left( \frac{\partial p_m}{\partial T} \right)_{p, \text{all\_}C\text{'s}} dT + \sum_{i=2}^{m_{\max}} \left( \frac{\partial p_m}{\partial C_i} \right)_{p, T, \text{all\_}C\text{'s\_except\_}C_i} dC_i. \quad (3.64)$$

Therefore

$$\left( \frac{\partial p_m}{\partial p} \right)_{T, \text{all\_}C\text{'s}} = \frac{C_m / \left( \frac{\partial \rho_m}{\partial p_m} \right)_T}{\sum_{i=1}^{m_{\max}} C_i / \left( \frac{\partial \rho_i}{\partial p_i} \right)_T}, \quad (3.65)$$

$$\left( \frac{\partial p_m}{\partial T} \right)_{p, \text{all\_}C\text{'s}} = \frac{1}{\left( \frac{\partial p_m}{\partial \rho_m} \right)_T} \left[ C_m \frac{\sum_{i=1}^{m_{\max}} \left( \frac{\partial \rho_i}{\partial T} \right)_{p_i} / \left( \frac{\partial \rho_i}{\partial p_i} \right)_T - \left( \frac{\partial \rho_m}{\partial T} \right)_{p_m}}{\sum_{i=1}^{m_{\max}} C_i / \left( \frac{\partial \rho_i}{\partial p_i} \right)_T} \right], \quad (3.66)$$

$$\left( \frac{\partial p_m}{\partial C_i} \right)_{p, T, \text{all\_}C\text{'s\_except\_}C_i} = \rho \left( \frac{\partial p_m}{\partial \rho_m} \right)_T \left[ \delta_{mi} - C_m \frac{\left( \frac{\partial p_i}{\partial \rho_i} \right)_T - \left( \frac{\partial p_1}{\partial \rho_1} \right)_T}{\sum_{i=1}^{m_{\max}} C_i \left( \frac{\partial p_i}{\partial \rho_i} \right)_T} \right]. \quad (3.67)$$

For the limiting case, where all of the components are taken to be perfect gases, we have

$$\left(\frac{\partial \rho_i}{\partial p_i}\right)_T = \frac{1}{R_i T}, \quad (3.68)$$

$$\left(\frac{\partial \rho_i}{\partial T}\right)_{p_i} = -\frac{\rho_i}{T}, \quad (3.69)$$

and

$$\left(\frac{\partial \rho_1}{\partial p_1}\right)_T = \frac{1}{R_1 T}, \quad (3.70)$$

$$\left(\frac{\partial \rho_1}{\partial T}\right)_{p_1} = -\frac{\rho_1}{T}. \quad (3.71)$$

Substituting in the equations defining the derivatives this yields

$$\left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C's} = 1 / \left(T \sum_{i=1}^{i_{\max}} C_i R_i\right) = \frac{1}{RT}, \quad (3.72)$$

$$\left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C's} = -\left(\sum_{i=1}^{i_{\max}} \rho_i R_i\right) / \left(T \sum_{i=1}^{i_{\max}} C_i R_i\right) = -\rho / T, \quad (3.73)$$

$$\left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all } C's \text{ except } C_i} = -\rho (R_i - R_1) / R, \quad (3.74)$$

where  $R = \sum_{i=1}^{i_{\max}} C_i R_i$ .

Now assume that the gas consists only of one  $i$ -th component,  $C_1 = 0$ ,  $\rho_1 = 0$ , and  $C_i = 1$ . Using Eq. (3.50), this trivial case yields

$$\left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C's} = \frac{1}{R_i T}, \quad (3.75)$$

$$\left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C's} = -\rho_i / T, \quad (3.76)$$

and

$$\left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all } C's \text{ except } C_i} = 0. \quad (3.77)$$

The derivatives are not defined in the case of missing miscible components because in this case Eq. (3.6) from which we started our derivation is no longer valid. In this case the Eq. (3.44) simplifies fortunately to

$$\frac{1}{\rho} = \sum_{i=1}^{i_{\max}} \frac{C_i}{\rho_i(p, T)} = 1/\rho(C_{1,2,\dots,i_{\max}}, p, T). \quad (3.78)$$

The derivatives are then obtained from the differential form

$$\begin{aligned} d\rho &= \sum_{i=1}^{i_{\max}} C_i \frac{\rho^2}{\rho_i^2} d\rho_i - \rho^2 \sum_{i=1}^{i_{\max}} \frac{1}{\rho_i} dC_i \\ &= \left[ \sum_{i=1}^{i_{\max}} C_i \frac{\rho^2}{\rho_i^2} \left( \frac{\partial \rho_i}{\partial p} \right)_T \right] dp + \left[ \sum_{i=1}^{i_{\max}} C_i \frac{\rho^2}{\rho_i^2} \left( \frac{\partial \rho_i}{\partial T} \right)_p \right] dT - \rho^2 \sum_{i=2}^{i_{\max}} \left( \frac{1}{\rho_i} - \frac{1}{\rho_1} \right) dC_i. \end{aligned} \quad (3.79)$$

The final result is

$$\left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all } C\text{'s}} = \sum_{i=1}^{i_{\max}} C_i \frac{\rho^2}{\rho_i^2} \left( \frac{\partial \rho_i}{\partial p} \right)_T, \quad (3.80)$$

$$\left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all } C\text{'s}} = \sum_{i=1}^{i_{\max}} C_i \frac{\rho^2}{\rho_i^2} \left( \frac{\partial \rho_i}{\partial T} \right)_p, \quad (3.81)$$

$$\left( \frac{\partial \rho}{\partial C_i} \right)_{p, T, \text{all } C\text{'s}_{\text{except } C_i}} = -\rho^2 \left( \frac{1}{\rho_i} - \frac{1}{\rho_1} \right), \quad (3.82)$$

If some of the non-miscible components are considered as non-compressible, their density derivatives are simply set to zero.

### 3.2.3 Partial derivatives in the equation of state $T = T(\varphi, p, C_{2,\dots,i_{\max}})$ , where $\varphi = s, h, e$

The next step is to define the partial derivatives in the linearized equation of state

$$T = T(\varphi, p, C_{2,\dots,i_{\max}}) \quad (3.83)$$

namely

$$dT = \left( \frac{\partial T}{\partial p} \right)_{\varphi, \text{all } C_i} dp + \left( \frac{\partial T}{\partial \varphi} \right)_{p, \text{all } C_i} d\varphi + \sum_{i=2}^{i_{\max}} \left( \frac{\partial T}{\partial C_i} \right)_{p, \varphi, \text{all } C_i \text{ except } C_i} dC_i, \quad (3.84)$$

where  $\varphi$  may be one of the following variables: specific entropy, specific enthalpy or specific internal energy,

$$\varphi = s, h, e.$$

This is very important for the construction of a numerical algorithm within the framework of the *entropy, or enthalpy or energy concepts*. Remember that this kind of algorithm describes the behavior of the flow using the specific mixture property  $\varphi$  as elements of the dependent variable vector. As the increments  $\delta p, \delta \varphi$  and  $\delta C_i$  are known from the numerical integration of the system of PDEs governing the flow, it is then possible to compute the corresponding increment of the temperature  $T$ , and therefore, the particular thermo-physical and transport properties of the velocity field, which depend on  $T, p$  and  $C_i$ .

Begin with the definition equation for the specific mixture property

$$\rho\varphi = \sum_{i=1}^{i_{\max}} \rho_i \varphi_i \quad (3.85)$$

or

$$\varphi = \sum_{i=1}^{i_{\max}} C_i \varphi_i. \quad (3.86)$$

After differentiating the definition equation for the specific mixture entropy (3.86), the following is obtained

$$d\varphi = \sum_{i=1}^{i_{\max}} C_i d\varphi_i + \sum_{i=1}^{i_{\max}} \varphi_i dC_i. \quad (3.87)$$

An arbitrary, but existing component is again denoted with 1, where the mass concentration of this component is uniquely defined by knowing the other concentrations

$$C_1 = 1 - \sum_{i=2}^{i_{\max}} C_i \quad (3.88)$$

and therefore

$$dC_1 = - \sum_{i=2}^{i_{\max}} dC_i. \quad (3.89)$$

Note that this is the only criterion for selecting the above component denoted with 1. Its concentration is not calculated using the differential conservation equation, but merely using Eq. (3.88).

*Enthalpy concept*  $\varphi = h$ . In this case Eq. (3.87) is

$$dh = \sum_{i=1}^{i_{\max}} C_i dh_i + \sum_{i=1}^{i_{\max}} h_i dC_i. \quad (3.90)$$

Replace in the above equation the differentials of the specific enthalpies using the caloric equations

$$dh_i = c_{pi} dT + \left( \frac{\partial h_i}{\partial p_i} \right)_T dp_i, \quad (3.91)$$

to obtain

$$dh = \left( \sum_{i=1}^{i_{\max}} C_i c_{pi} \right) dT + \sum_{i=1}^{m_{\max}} C_i \left( \frac{\partial h_i}{\partial p_i} \right)_T dp_i + \left[ \sum_{i=m_{\max}+1}^{i_{\max}} C_i \left( \frac{\partial h_i}{\partial p} \right)_T \right] dp + \sum_{i=1}^{i_{\max}} h_i dC_i. \quad (3.92)$$

Replace  $dp_i$  in the above equation by means of the differentiated equation of state  $p_i = p_i(\rho_i, T) = p_i(C_i, \rho, T)$ , namely

$$dp_i = \left( \frac{\partial p_i}{\partial \rho_i} \right)_T C_i d\rho + \rho \left( \frac{\partial p_i}{\partial \rho_i} \right)_T dC_i + \left( \frac{\partial p_i}{\partial T} \right)_{\rho_i} dT, \quad (3.93)$$

to obtain

$$\begin{aligned} dh = & \left[ \left( \sum_{i=1}^{i_{\max}} C_i c_{pi} \right) + \sum_{i=1}^{m_{\max}} C_i \left( \frac{\partial h_i}{\partial p_i} \right)_T \left( \frac{\partial p_i}{\partial T} \right)_{\rho_i} \right] dT + \left[ \sum_{i=1}^{m_{\max}} C_i^2 \left( \frac{\partial h_i}{\partial p_i} \right)_T \left( \frac{\partial p_i}{\partial \rho_i} \right)_T \right] d\rho \\ & + \left[ \sum_{i=m_{\max}+1}^{i_{\max}} C_i \left( \frac{\partial h_i}{\partial p} \right)_T \right] dp + \sum_{i=1}^{i_{\max}} h_i dC_i + \sum_{i=1}^{m_{\max}} \rho C_i \left( \frac{\partial h_i}{\partial p_i} \right)_T \left( \frac{\partial p_i}{\partial \rho_i} \right)_T dC_i. \end{aligned} \quad (3.94)$$

Substituting  $dC_1$  from Eq. (3.89) and  $d\rho$  from Eq. (3.47) into Eq. (3.94) yields

$$\begin{aligned} dh = & \left[ \sum_{i=1}^{m_{\max}} C_i \left\{ c_{pi} + \left( \frac{\partial h_i}{\partial p_i} \right)_T \left[ \left( \frac{\partial p_i}{\partial T} \right)_{\rho_i} + C_i \left( \frac{\partial p_i}{\partial \rho_i} \right)_T \left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all } C's} \right] \right\} + \sum_{i=m_{\max}+1}^{i_{\max}} C_i c_{pi} \right] dT \\ & + \left\{ \left[ \sum_{i=1}^{m_{\max}} C_i^2 \left( \frac{\partial h_i}{\partial p_i} \right)_T \left( \frac{\partial p_i}{\partial \rho_i} \right)_T \right] \left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all } C's} + \sum_{i=m_{\max}+1}^{i_{\max}} C_i \left( \frac{\partial h_i}{\partial p} \right)_T \right\} dp \end{aligned}$$

$$\begin{aligned}
 & + \sum_{i=2}^{i_{\max}} \left\{ h_i - h_1 + \left[ \sum_{i=1}^{m_{\max}} C_i^2 \left( \frac{\partial h_i}{\partial p_i} \right)_T \left( \frac{\partial p_i}{\partial \rho_i} \right)_T \right] \left( \frac{\partial \rho}{\partial C_i} \right)_{p,T,\text{all } C's \text{ except } C_i} \right\} dC_i \\
 & + \sum_{i=2}^{m_{\max}} \rho \left[ C_i \left( \frac{\partial h_i}{\partial p_i} \right)_T \left( \frac{\partial p_i}{\partial \rho_i} \right)_T - C_1 \left( \frac{\partial h_1}{\partial p_1} \right)_T \left( \frac{\partial p_1}{\partial \rho_1} \right)_T \right] dC_i, \quad (3.95)
 \end{aligned}$$

or

$$dh = c_p dT + \left( \frac{\partial h}{\partial p} \right)_{T,\text{all } C's} dp + \sum_{i=2}^{i_{\max}} \left( \frac{\partial h}{\partial C_i} \right)_{p,T,\text{all } C's \text{ except } C_i} dC_i \quad (3.96)$$

or

$$dT = \frac{dh}{c_p} - \frac{1}{c_p} \left( \frac{\partial h}{\partial p} \right)_{T,\text{all } C's} dp - \frac{1}{c_p} \sum_{i=2}^{i_{\max}} \left( \frac{\partial h}{\partial C_i} \right)_{p,T,\text{all } C's \text{ except } C_i} dC_i \quad (3.97)$$

where

$$\begin{aligned}
 c_p & = \sum_{i=m_{\max}+1}^{i_{\max}} C_i c_{pi} + \sum_{i=1}^{m_{\max}} C_i \left\{ c_{pi} + \left( \frac{\partial h_i}{\partial p_i} \right)_T \left[ \left( \frac{\partial p_i}{\partial T} \right)_{\rho_i} + C_i \left( \frac{\partial p_i}{\partial \rho_i} \right)_T \left( \frac{\partial \rho}{\partial T} \right)_{p,\text{all } C's} \right] \right\} \\
 & = \sum_{i=m_{\max}+1}^{i_{\max}} C_i c_{pi} + \sum_{i=1}^{m_{\max}} C_i \left\{ c_{pi} + \frac{(\partial h_i / \partial p_i)_T}{(\partial \rho_i / \partial p_i)_T} \left[ C_i \left( \frac{\partial \rho}{\partial T} \right)_{p,\text{all } C's} - \left( \frac{\partial \rho_i}{\partial T} \right)_{\rho_i} \right] \right\}, \quad (3.98)
 \end{aligned}$$

$$\left( \frac{\partial h}{\partial p} \right)_{T,\text{all } C's} = \left( \frac{\partial \rho}{\partial p} \right)_{T,\text{all } C's} \left[ \sum_{i=1}^{m_{\max}} C_i^2 \left( \frac{\partial h_i}{\partial p_i} \right)_T / \left( \frac{\partial \rho_i}{\partial p_i} \right)_T \right] + \sum_{i=m_{\max}+1}^{i_{\max}} C_i \left( \frac{\partial h_i}{\partial p} \right)_T, \quad (3.99)$$

$$\left( \frac{\partial h}{\partial C_i} \right)_{p,T,\text{all } C's \text{ except } C_i} = h_i - h_1 + \Delta h_i^{np} \quad (3.100)$$

and

$$\Delta h_i^{np} = \rho \left[ C_i \left( \frac{\partial h_i}{\partial p_i} \right)_T \middle/ \left( \frac{\partial p_i}{\partial p_i} \right)_T - C_1 \left( \frac{\partial h_1}{\partial p_1} \right)_T \middle/ \left( \frac{\partial p_1}{\partial p_1} \right)_T \right] + \left[ \sum_{i=1}^{m_{\max}} C_i^2 \left( \frac{\partial h_i}{\partial p_i} \right)_T \middle/ \left( \frac{\partial p_i}{\partial p_i} \right)_T \right] \left( \frac{\partial \rho}{\partial C_i} \right)_{p,T,all\_C's\_except\_C_i}, \quad (3.101)$$

for  $i = 2, m_{\max}$  and

$$\Delta h_i^{np} = \left[ \sum_{i=1}^{m_{\max}} C_i^2 \left( \frac{\partial h_i}{\partial p_i} \right)_T \middle/ \left( \frac{\partial p_i}{\partial p_i} \right)_T \right] \left( \frac{\partial \rho}{\partial C_i} \right)_{p,T,all\_C's\_except\_C_i} \quad (3.102)$$

for  $i = m_{\max} + 1, i_{\max}$ .

Important: note that for non-existing miscible components the concentration denoted with “1” in Eq. (3.100) is the first non-miscible component.

Equation (3.100) consists of two parts. In the case of a mixture consisting of perfect fluids, the second part is equal to zero,  $\Delta h_i^{np} = 0$ , because the enthalpies do not depend on the corresponding partial pressures. This also illustrates the meaning of the superscript  $np$  which stands for non-perfect fluid. From Eq. (3.97) we obtain the analytical expressions for the following derivatives

$$\left( \frac{\partial T}{\partial p} \right)_{h,all\_C's} = -\frac{1}{c_p} \left( \frac{\partial h}{\partial p} \right)_{T,all\_C's}, \quad (3.103)$$

$$\left( \frac{\partial T}{\partial h} \right)_{p,all\_C's} = \frac{1}{c_p}, \quad (3.104)$$

$$\left( \frac{\partial T}{\partial C_i} \right)_{p,h,all\_C's\_except\_C_i} = -\frac{1}{c_p} \left( \frac{\partial h}{\partial C_i} \right)_{p,T,all\_C's\_except\_C_i}. \quad (3.105)$$

Checking the above derivation: for a perfect gas mixture, and taking into account that

$$\left( \frac{\partial h_i}{\partial p_i} \right)_T = 0, \quad (3.106)$$

$$\left(\frac{\partial h_1}{\partial p_1}\right)_T = 0, \quad (3.107)$$

we obtain

$$c_p = \sum_{i=1}^{i_{\max}} C_i c_{pi}, \quad (3.108)$$

$$\left(\frac{\partial h}{\partial p}\right)_{T, \text{all } C_i} = 0, \quad (3.109)$$

and

$$\left(\frac{\partial h}{\partial C_i}\right)_{p, T, \text{all } C_i \text{ except } C_i} = h_i - h_1, \quad (3.110)$$

which is the result expected.

*Energy concept*  $\phi = e$ . The derivation already presented for the specific enthalpy is formally identical with the derivation for the specific internal energy. One has to replace formally  $h$  with  $e$  and write instead  $c_{pi}$ ,  $\left(\frac{\partial e_i}{\partial T}\right)_{p_i, T}$ .

*Entropy concept*  $\phi = s$ . Then use the definition equation for the particular specific entropies

$$T \rho_i ds_i = \rho_i dh_i - dp_i \quad (3.111)$$

or divided by  $\rho$

$$TC_i ds_i = C_i dh_i - \frac{dp_i}{\rho}. \quad (3.112)$$

Substituting in this equation  $dh_i$  using Eq. (3.91), summing the resulting  $i_{\max}$  equations and comparing the result with Eq. (3.87) we finally obtain

$$Tds = \left(\sum_{i=1}^{i_{\max}} C_i dh_i\right) - \frac{dp}{\rho} + T \sum_{i=1}^{i_{\max}} s_i dC_i. \quad (3.113)$$

Note that this is equivalent to the *Gibbs* equation for mixtures

$$\begin{aligned} Tds &= \sum_{i=1}^{i_{\max}} d(C_i h_i) - \frac{dp}{\rho} + T \sum_{i=1}^{i_{\max}} s_i dC_i - \left(\sum_{i=1}^{i_{\max}} h_i dC_i\right) = dh - \frac{dp}{\rho} - \sum_{i=1}^{i_{\max}} (h_i - Ts_i) dC_i \\ &= de + pd\left(\frac{1}{\rho}\right) - \sum_{i=1}^{i_{\max}} (h_i - Ts_i) dC_i, \end{aligned} \quad (3.114)$$



in which the  $h_i - Ts_i$  is the so called *Gibbs potential* for the single component. The entropy change due to the mixing process is

$$ds_{\text{mixing}} = -\frac{1}{T} \sum_{i=1}^{i_{\text{max}}} (h_i - Ts_i) dC_i . \quad (3.115)$$

Substituting  $dC_1$  from Eq. (3.89) into Eq. (3.115) results in

$$ds_{\text{mixing}} = -\frac{1}{T} \sum_{i=2}^{i_{\text{max}}} [h_i - h_1 - T(s_i - s_1)] dC_i . \quad (3.116a)$$

This result follows not from Eq. (3.114) directly but from the comparison of Eq. (3.114) with the energy conservation equation for the mixture. For a mixing process in which there is no total mass change in the system,  $\sum_{k=1}^{i_{\text{max}}} Y_k M_k = \text{const}$ ,

Eq. (3.116a) can be rearranged using Eq. (3.17)

$$ds_{\text{mixing}} = -\frac{1}{T} \sum_{i=2}^{i_{\text{max}}} C_i [h_i - h_1 - T(s_i - s_1)] d \ln Y_i . \quad (3.116b)$$

Remember, that in case of mixing of perfect fluids the molar concentration is equivalent to the ratio of the partial pressure to the total pressure.

Replace the differentials in Eq. (3.114) of the mixture specific enthalpy using the caloric equations (3.96) to obtain

$$ds = \frac{c_p}{T} dT + \frac{1}{\rho T} \left[ \rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all } C's} - 1 \right] dp + \sum_{i=2}^{i_{\text{max}}} \left( s_i - s_1 + \frac{\Delta h_i^{np}}{T} \right) dC_i , \quad (3.117)$$

or

$$dT = \frac{T}{c_p} ds - \frac{\rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all } C's} - 1}{\rho c_p} dp - \frac{T}{c_p} \sum_{i=2}^{i_{\text{max}}} \left( \frac{\partial s}{\partial C_i} \right)_{p, T, \text{all } C's \text{ except } C_i} dC_i \quad (3.118)$$

where

$$\left( \frac{\partial s}{\partial C_i} \right)_{p, T, \text{all } C's \text{ except } C_i} = s_i - s_1 + \Delta s_i^{np} \quad (3.119)$$

where

$$\Delta s_i^{np} = \frac{\Delta h_i^{np}}{T}. \quad (3.120)$$

Comparing Eq. (3.100) with Eq. (3.119) we realize that

$$\left( \frac{\partial s}{\partial C_i} \right)_{p,T,\text{all}_{C's}_{\text{except}_{C_i}}} = \frac{1}{T} \left( \frac{\partial h}{\partial C_i} \right)_{p,T,\text{all}_{C's}_{\text{except}_{C_i}}}. \quad (3.121)$$

Like Eq. (3.100) Eq. (3.119) also consists of two parts. In the case of a mixture consisting of perfect fluids, the second part is equal to zero,  $\Delta s_{il}^{np} = 0$ , because the enthalpies do not depend on the corresponding partial pressures. This again illustrates the meaning of the superscript *np* which stands for non-perfect fluid. From Eq. (3.118) analytical expressions for the following derivatives are obtained

$$\left( \frac{\partial T}{\partial p} \right)_{s,\text{all}_{C's}} = \frac{1}{\rho c_p} \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T,\text{all}_{C's}} \right], \quad (3.122)$$

$$\left( \frac{\partial T}{\partial s} \right)_{p,\text{all}_{C's}} = \frac{T}{c_p}, \quad (3.123)$$

$$\left( \frac{\partial T}{\partial C_i} \right)_{p,s,\text{all}_{C's}_{\text{except}_{C_i}}} = - \frac{T}{c_p} \left( \frac{\partial s}{\partial C_i} \right)_{p,T,\text{all}_{C's}_{\text{except}_{C_i}}}. \quad (3.124)$$

Checking the above derivation: for a perfect gas mixture, and taking into account that

$$\left( \frac{\partial h_i}{\partial p_i} \right)_T = 0, \quad (3.125)$$

$$\left( \frac{\partial h_1}{\partial p_1} \right)_T = 0, \quad (3.126)$$

we obtain

$$c_p = \sum_{i=1}^{i_{\max}} C_i c_{pi}, \quad (3.127)$$

$$\left(\frac{\partial h}{\partial p}\right)_{T,all\_C's} = 0, \quad (3.128)$$

and

$$\left(\frac{\partial s}{\partial C_i}\right)_{p,T,all\_C's\_except\_C_i} = s_i - s_1, \quad (3.129)$$

which is the result expected.

Thus, as already mentioned, after the required derivatives have been computed in Eq. (3.118) in the manner shown, the temperature difference can be easily computed:

$$\begin{aligned} T - T_a &= \left(\frac{\partial T}{\partial p}\right)_{s,all\_C's} (p - p_a) + \left(\frac{\partial T}{\partial s}\right)_{p,all\_C's} (s - s_a) \\ &+ \sum_{i=2}^{i_{\max}} \left(\frac{\partial T}{\partial C_i}\right)_{p,s,all\_C's\_except\_C_i} (C_i - C_{ia}), \end{aligned} \quad (3.130)$$

which corresponds to the increments for the pressure, entropy and concentrations. When the increments  $\delta p$ ,  $\delta s$  and  $\delta C_i$  are of considerable magnitude, it is better to take into account their non-linear dependencies in Eq. (3.118),

$$\frac{dT}{T} = \frac{1}{c_p} \left[ ds - \sum_{i=2}^{i_{\max}} \left(\frac{\partial s}{\partial C_i}\right)_{p,T,all\_C's\_except\_C_i} dC_i \right] + \frac{p}{\rho T c_p} \left[ 1 - \rho \left(\frac{\partial h}{\partial p}\right)_{T,all\_C's} \right] \frac{dp}{p} \quad (3.131)$$

in the following way: substitute

$$\Delta s^* = \frac{1}{c_p} \left[ s - s_a - \sum_{i=2}^{i_{\max}} \left(\frac{\partial s}{\partial C_i}\right)_{p,T,all\_C's\_except\_C_i} (C - C_{ia}) \right] \quad (3.132)$$

$$\bar{R} = \frac{p}{\rho T} \left[ 1 - \rho \left(\frac{\partial h}{\partial p}\right)_{T,all\_C's} \right] \quad (3.133)$$

and integrate to obtain

$$\ln(T/T_a) = \Delta s^* + (\bar{R}/c_p) \ln(p/p_a) \quad (3.134)$$

or

$$T = T_a e^{\Delta s^*} (p/p_a)^{\bar{R}/c_p}, \quad (3.135)$$

or for a quasi-linear relationship between temperature  $T$  and specific enthalpy  $h$

$$h = h_a e^{\Delta s^*} (p/p_a)^{\bar{R}/c_p} . \quad (3.136)$$

This is a canonical equation, as all thermodynamic properties of the gas mixture can be computed as a function of  $(s, p, C_i)$  by differentiating. For  $\Delta s^* = 0$ ,

$$T = T_a (p/p_a)^{\bar{R}/c_p} . \quad (3.137)$$

Equation (3.130) is also useful for exact solution of the task  $T = ?$  if  $s$ ,  $p$  and  $C_{2,3,\dots,i_{\max}}$  are known. This can be performed by iterations (*Newton* method) as follows: the error for the previous step designated with superscript  $n$  is

$$\Delta s^n = s - s^n (T^n, p, C_{2,3,\dots,i_{\max}}) . \quad (3.138)$$

A new temperature  $T^{n+1}$  can be computed so as to obtain  $\Delta s^{n+1} = 0$ ,

$$\Delta s^{n+1} - \Delta s^n = - \left( \frac{\partial s}{\partial T} \right)_{p, \text{all } C_s} (T^{n+1} - T^n) \quad (3.139)$$

or solving with respect to  $T^{n+1}$  and using Eq. (3.123)

$$T^{n+1} = T^n \left\{ 1 + \left[ s - s^n (T^n, p, C_{2,3,\dots,i_{\max}}) \right] / c_p \right\} \quad (3.140)$$

or, in more precise form

$$T^{n+1} = T^n \exp \left\{ \left[ s - s^n (T^n, p, C_{2,3,\dots,i_{\max}}) \right] / c_p \right\} \quad (3.141)$$

### 3.2.4 Chemical potential

#### 3.2.4.1 Gibbs function

Neither the specific entropy, nor the specific enthalpy, nor the specific internal energy can be measured directly. The quantities that can be measured directly are pressure, temperature and concentrations. It is interesting to know whether Eq. (3.114) can be rewritten as a function of measurable variables  $f = f(p, T, C's)$ . The answer is yes and the result is

$$d(h - Ts) = \frac{dp}{\rho} - s dT + \sum_{i=1}^{i_{\max}} (h_i - Ts_i) dC_i . \quad (3.142)$$

Introducing the so called *free enthalpy* or *Gibbs function*,

$$g = h - Ts \quad (3.143)$$

Eq. (3.142) takes the remarkable form

$$dg = \frac{dp}{\rho} - sdT + \sum_{i=2}^{i_{\max}} [h_i - h_1 - T(s_i - s_1)] dC_i, \quad (3.144)$$

with

$$\left( \frac{\partial g}{\partial p} \right)_{T, \text{all } C_i} = \frac{1}{\rho}, \quad (3.145)$$

$$\left( \frac{\partial g}{\partial T} \right)_{p, \text{all } C_i} = -s, \quad (3.146)$$

$$\left( \frac{\partial g}{\partial C_i} \right)_{p, s, \text{all } C_j \text{ except } C_i} = h_i - h_1 - T(s_i - s_1). \quad (3.147)$$

### 3.2.4.2 Definition of the chemical equilibrium

We identify a mixture that is at constant mixture pressure and temperature and that does not change the concentrations of its constituent as a mixture in *chemical equilibrium*.

The remarkable property of Eq. (3.144) is that it provides a quantitative definition of the chemical equilibrium, namely:

$$dg = 0 \quad \text{or} \quad g = \text{const} \quad (3.148)$$

Now consider the following chemical reaction

$$\sum_{i=1}^{i_{\max}} n_i \text{Symb}_i = 0. \quad (3.149)$$

Here  $\text{Symb}_i$  is the chemical identification symbol of the  $i$ -th species and  $n_i$  is the stoichiometric coefficient equal to the number of *kg-moles* of each species that participates in the reaction. Usually  $n_i < 0$  for *reactants* that are reducing their mass and  $n_i > 0$  for *products* that are increasing their mass in the mixture. An example is  $-2H_2 - O_2 + 2H_2O = 0$ . During the progress of a *stoichiometric reac-*

tion all concentrations change not arbitrarily but so that the mass change of a single species is always proportional to  $n_i M_i$ . This is demonstrated as follows. The mass balance of the stoichiometric reaction (3.149) gives

$$\sum_{i=1}^{i_{\max}} n_i M_i = 0. \quad (3.150)$$

Applied to our last example:  $-2 \times 2kg - 1 \times 32kg + 2 \times 18kg = 0$ . Note that  $n_i M_i$  is a natural constant belonging to the specific chemical reaction. On other hand, the sum of all mass sources per unit volume of the mixture is equal to zero,

$$\sum_{i=1}^{i_{\max}} \mu_i = 0. \quad (3.151)$$

Selecting one arbitrary, but existing, mass source term designated with subscript 1 (usually having the minimal initial concentration if it is going to be consumed) and rearranging we have

$$\sum_{i=1}^{i_{\max}} \mu_1 \frac{\mu_i}{\mu_1} = 0. \quad (3.152)$$

Equations (3.150) and (3.151) can then, and only then, be satisfied simultaneously if

$$\mu_i = \frac{n_i M_i}{n_1 M_1} \mu_1. \quad (3.153)$$

Applying this to our example we obtain

$$\mu_{H_2} = \mu_{H_2}, \quad \mu_{O_2} = (32/4) \mu_{H_2} = 8 \mu_{H_2}, \quad \mu_{H_2O} = -(36/4) \mu_{H_2} = -9 \mu_{H_2}.$$

For the closed volume in which the chemical reaction happens the mass conservation for each species gives

$$dC_i / d\tau = \mu_i / \rho, \quad (3.154)$$

and therefore

$$\frac{dC_i / d\tau}{dC_1 / d\tau} = \frac{\mu_i / \rho}{\mu_1 / \rho}, \quad (3.155)$$

or

$$dC_i = n_i M_i d \frac{C_1}{n_1 M_1} = n_i M_i d \xi, \quad (3.156)$$

where  $\xi$  is sometimes called in the literature the *reaction progress variable*. Substituting in Eq. (3.142) we obtain

$$d(h - Ts) = \frac{dp}{\rho} - s dT + \left[ \sum_{i=1}^{i_{\max}} (h_i - Ts_i) n_i M_i \right] d\xi. \quad (3.157)$$

For the case of chemical equilibrium,  $dg = 0$ , and at constant mixture pressure and temperature,  $dp = 0$  and  $dT = 0$ , Eq. (3.157) reads

$$\sum_{i=1}^{i_{\max}} (h_i - Ts_i) n_i M_i = 0. \quad (3.158)$$

This means that the function  $g = g(\xi)$  possesses an extreme  $dg = 0$  if Eq. (3.158) is fulfilled for all components. For a single component Eq. (3.142) results in

$$d(h_i - Ts_i) = \frac{dp_i}{\rho_i} - s_i dT. \quad (3.159)$$

For the case of a constant mixture temperature valid for each species,  $dT = 0$ , we have

$$d(h_i - Ts_i) = \frac{dp_i}{\rho_i}, \quad (3.160)$$

or after integration

$$h_i - Ts_i = h_{i0} - Ts_{i0} + \int_{p_0}^{p_i} \frac{dp'_i}{\rho_i}. \quad (3.161)$$

For perfect gases Eq. (3.161) results in

$$h_i - Ts_i = h_{i0} - Ts_{i0} + TR_i \int_{p_0}^{p_i} \frac{dp'_i}{P_i} = h_{i0} - Ts_{i0} + TR_i \ln \left( \frac{P_i}{P_0} \right). \quad (3.162)$$

Substituting (3.162) in Eq. (3.158) we obtain

$$\sum_{i=1}^{i_{\max}} (h_{i0} - Ts_{i0}) M_i n_i + T \sum_{i=1}^{i_{\max}} n_i M_i R_i \ln \left( \frac{P_i}{P_0} \right) = 0, \quad (3.163)$$

or bearing in mind that

$$R = M_i R_i, \quad (3.164)$$

with  $R = 8314 \text{ J/(kg-mol K)}$  being the universal gas constant,

$$\sum_{i=1}^{i_{\max}} (h_{i0} - Ts_{i0}) M_i n_i + TR \ln \prod_{i=1}^{i_{\max}} \left( \frac{p_i}{p_0} \right)^{n_i} = 0. \quad (3.165)$$

Introducing the so called *pressure equilibrium factor*

$$K_p = \prod_{i=1}^{i_{\max}} \left( \frac{p_i}{p_0} \right)^{n_i} = \frac{\prod_{i=1}^{i_{\max}} p_i^{n_i}}{p_0^n} = \left( \frac{p}{p_0} \right)^n \prod_{i=1}^{i_{\max}} Y_i^{n_i} = \left( \frac{p}{p_0} \right)^n K_y, \quad (3.166)$$

where

$$n = \sum_{i=1}^{i_{\max}} n_i, \quad (3.167)$$

$$K_y = \prod_{i=1}^{i_{\max}} Y_i^{n_i}, \quad (3.168)$$

and solving Eq. (3.165) with respect to  $K_p$  results in the expression well known in chemical thermodynamics defining mathematically the chemical equilibrium

$$K_p = \exp \left[ -\frac{1}{TR} \sum_{i=1}^{i_{\max}} (h_{i0} - Ts_{i0}) n_i M_i \right], \quad (3.169a)$$

see for comparison in Warnatz et al. (2001) p. 42, or

$$K_y = \left( \frac{p_0}{p} \right)^n \exp \left[ -\frac{1}{TR} \sum_{i=1}^{i_{\max}} (h_{i0} - Ts_{i0}) n_i M_i \right]. \quad (3.169b)$$

Therefore the pressure equilibrium factor *for the particular reaction* (3.149) can be analytically computed from the properties of the constituents. Another reaction generates another expression for the pressure equilibrium constant.

The expression

$$\Delta G_{ch,0} = \sum_{i=1}^{i_{\max}} (h_{i0} - Ts_{i0}) n_i M_i = \sum_{i=1}^{i_{\max}} h_{i0} n_i M_i - T \sum_{i=1}^{i_{\max}} s_{i0} n_i M_i, \quad (3.170)$$

is called the *molar Gibbs energy of the chemical reaction*, Warnatz et al. (2001). The expressions



$$\Delta H_{ch,0} = \sum_{i=1}^{i_{\max}} h_{i0} n_i M_i, \quad (3.171)$$

$$\Delta S_{ch,0} = \sum_{i=1}^{i_{\max}} s_{i0} n_i M_i, \quad (3.172)$$

are called the enthalpy and entropy changes of the mixture due to the chemical reaction, respectively. With this Eq. (3.169b) can also be used in the form

$$K_y = \left( \frac{P_0}{p} \right)^n \exp \left( -\frac{\Delta H_{ch,0}}{TR} + \frac{\Delta S_{ch,0}}{R} \right) = \left( \frac{P_0}{p} \right)^n \exp \left( -\frac{\Delta H_{ch,0}/R}{T} + \sum_{i=1}^{i_{\max}} n_i \frac{s_{i0}}{R_i} \right), \quad (3.173)$$

The choice of the reference pressure varies in the literature. Some authors set it to unity. In this case  $K_p = p^n K_y$ . If the mixture pressure is selected as the reference pressure we have  $K_p = K_y$ . In this case  $\Delta G_{ch,0} = G(p, T)$ . In this case Eq. (3.162) in its integral form  $\Delta g_i = TR_i d \ln(p_i/p)$  define a dimensionless property  $f_i = p_i/p$  called the *fugacity* of component  $i$  in the solution of perfect fluids. The general definition of the fugacity of component  $i$  is  $d \ln f_i = dg_i/(TR_i)$ .

### 3.2.4.3 Partial pressures of perfect fluid compounds in chemical equilibrium

Given the temperature  $T$  and the total pressure  $p$  of a mixture of  $j_{\max}$  compounds that may react in a number  $i_{\max}$  of chemical reactions, then

$$\sum_{j=1}^{j_{\max}} n_{ij} \text{Symb}_j = 0, \quad \text{for } i = 1, i_{\max}. \quad (3.174)$$

Here  $n_{ij}$  is the stoichiometric coefficient,  $< 0$  for reactants and  $> 0$  for products. We look for a solution of  $j_{\max}$  molar concentrations

$$\mathbf{Y}^T = (Y_1, Y_2, \dots, Y_{j_{\max}}), \quad (3.175)$$

for which the system is in chemical equilibrium. We know from *Dalton's* law that the system pressure is the sum of the partial pressures

$$p = \sum_{j=1}^{j_{\max}} p_j, \quad Y_i = \frac{p_i}{p}, \quad \sum_{j=1}^{j_{\max}} Y_j = 1. \quad (3.176)$$

For each chemical reaction  $i$  we have the condition enforcing chemical equilibrium (3.174),

$$K_{y,i} = \prod_{j=1}^{j_{\max}} Y_j^{n_{ij}} = \left( \frac{p_0}{p} \right)^{n_i} \exp \left( -\frac{\Delta H_{ch,0,i}/R}{T} + \sum_{j=1}^{j_{\max}} n_{ij} \frac{s_{j0}}{R_j} \right), \text{ for } i = 1, i_{\max}. \quad (3.177)$$

where

$$n_i = \sum_{j=1}^{j_{\max}} n_{ij}. \quad (3.178)$$

Here  $R$  is the universal gas constant,  $R_j$  is the gas constant of the  $j$ -th component,

$$\Delta H_{ch,0,i} = \sum_{j=1}^{j_{\max}} h_{0j} n_{ij} M_j \quad (3.179)$$

and

$$\bar{s}_{0j} = s_{0j}/R_j \quad (3.180)$$

are the dimensionless reference entropies of the species  $j$ . Therefore we have  $j_{\max}$  unknowns and  $1+i_{\max}$  equations (3.176), (3.177). The missing  $j_{\max} - i_{\max} - 1$  are obtained from the *proportions* defined by the mass conservation equations

$$\sum_{j=1}^{j_{\max}} n_{ij} M_j = 0, \text{ for } i = 1, i_{\max}, \quad (3.181)$$

resulting in

$$\frac{dC_j}{n_{ij} M_{ij}} = \frac{dC_{ij^*}}{n_{ij^*} M_{ij^*}}, \text{ for } i = 1, i_{\max}, j = 1, j_{\max}, \quad (3.182)$$

where  $ij^*$  refers for each  $i$ -reaction to the minimum but existing compound  $j$ , or

$$\frac{dY_j}{n_{ij}} = \frac{dY_{ij^*}}{n_{ij^*}}, \text{ for } i = 1, i_{\max}, j = 1, j_{\max}. \quad (3.183)$$

Observe that in a chemical reaction for reactants that completely disappear we have  $dY_j = -Y_j$ , and for products that do not exist before the reaction but just originate we have  $dY_j = Y_j$ . Comparing this with the definition of the stoichiometric coefficient for which  $n_{ij} < 0$  for reactants and  $> 0$  for products, we have

$$\frac{Y_j}{|n_{ij}|} = \frac{Y_{ij^*}}{|n_{ij^*}|}, \text{ for } i = 1, i_{\max}, j = 1, j_{\max}. \quad (3.184)$$

**Significance.** The theory of the chemical equilibrium has very wide-ranging applications that are particularly important to physical chemists, metallurgists, solid-state physicists, and many engineers. It permits us to predict the equilibrium

composition of a mixture of chemicals in an isolated vessel; it tells us how many phases of an alloy can exist together at any particular temperature and pressure etc. Examples of the application of these theoretical results are given in Chapter 6.

### 3.2.4.4 Phase equilibrium

The idea of chemical equilibrium is extendable to the so called phase equilibrium. If we apply the definitions of the *Gibbs* function, Eq. (3.144), to a mixture of two phases designated with ' for the liquid and '' for its vapor we have

$$dg = \frac{dp}{\rho} - sdT + [h'' - h' - T(s'' - s')]dC' = \frac{dp}{\rho} - sdT - [h'' - h' - T(s'' - s')]dC'' \quad (3.185)$$

Because evaporation at a given constant pressure happens for pure liquids at a constant temperature the relation between the entropy change and the enthalpy change is

$$s'' - s' = \frac{h'' - h'}{T} \quad (3.186)$$

This relation makes the *Gibbs* function of the mixture independent of the vapor mass concentration change at constant pressure and temperature. Therefore the equilibrium between the liquid and its vapor for any pressure and temperature is defined by

$$dg = 0 \quad (3.187)$$

Equation (3.186) can be rewritten in the form  $h' - Ts' = h'' - Ts''$ .

In other words, the equilibrium between the liquid and its vapor is defined by the equality of the specific *Gibbs* functions for each phase

$$g' = g'' \quad (3.188)$$

If the liquid and the vapor initially in equilibrium are disturbed by  $dp$  resulting in  $dg'$  and  $dg''$ , respectively, but in such a way that the resultant mixture is again in equilibrium we have

$$g' + dg' = g'' + dg'' \quad (3.189)$$

or after using Eq. (3.188)

$$dg' = dg'' \quad (3.190)$$

resulting in

$$v' dp - s' dT = v'' dp - s'' dT \quad (3.191)$$

and finally

$$\frac{dT}{dp} = \frac{v'' - v'}{s'' - s'} \quad (3.192)$$

derived for the first time by *Benoit Pierre Emile Clapeyron* in 1834. The significance of *Clapeyron's* equation is in the description of the  $p$ - $T$  line dividing the stable single phase state from the metastable state. This line is called the *saturation line*, *Clausius*, see in *Elsner* (1974) p. 327. Another derivation of Eq. (3.192) is obtained if one visualizes in the  $T$ - $s$  and  $p$ - $v$  diagrams the evaporation process as a *Carnot cycle* for infinitesimal values of  $dT$  and  $dp$ . Equalizing the work computed from both diagrams for the cycle results in Eq. (3.193). Using Eq. (3.186) the above relation reads

$$\frac{dT}{dp} = T \frac{v'' - v'}{h'' - h'} \quad (3.193)$$

*Clapeyron's* equations also define the equilibrium state for the case of an equilibrium mixture of liquid and solid. In this case the corresponding properties along the solidification line have to be used.

At low pressure for which the density difference between liquid and vapor is very large Eq. (3.193) simplifies to  $dT/dp = T v''/\Delta h$ . This equation has been known since 1828 in France as the *August equation*. Assuming the vapor behaves as a perfect gas results in  $dp/p = (\Delta h/R) dT/T^2$ , where  $R$  is the vapor gas constant for the specific substance, (*R. Clausius* 1850). Assuming that the latent heat of evaporation is a constant, integrating between an initial state 0 and a final state and rearranging results in a useful expression for extrapolation along the saturation line around the known state 0,

$$\frac{p - p_0}{p_0} = \exp\left(\frac{\Delta h}{RT} \frac{T - T_0}{T_0}\right) - 1. \quad (3.194)$$

A next step in the improvements of Eq. (3.194) is to be assumed that  $\Delta h \approx \Delta h_0 + \alpha T$ , resulting in  $dp/p = (\Delta h_0/RT^2) dT + (\alpha/RT) dT$ . After integrating we obtain

$$\ln p = -\frac{\Delta h_0}{RT} + \frac{\alpha}{R} \ln T + \frac{\Delta h_0}{RT_0} + \ln p_0 - \frac{\alpha}{R} \ln T_0, \quad (3.195)$$

or

$$\ln p = A/T + B \ln T + C, \quad (3.196)$$

which is a form empirically proposed by *G. R. Kirchhoff* in 1858. Sometimes only the boiling properties at atmospheric pressure are known. Equation (3.194) allows us to compute the saturation temperature at a pressure different from the atmospheric pressure as long as the pressure remains far below the critical pressure. Equation (3.196) allows the three unknown constants to be fitted on three measured points. In fact the non-linear dependence of the group  $(v'' - v')/(h'' - h')$  on the temperature requires a non-linear approximation of the saturation line of the type  $\ln p = f(T)$  e.g.  $\Delta h \approx a_1 + a_2 T + a_3 T^2$ . In this case

$$\frac{dp}{p} = \frac{1}{R} \left( a_1 \frac{dT}{T^2} + a_2 \frac{dT}{T} + a_3 dT \right).$$

After integration between two pressure-temperature points we obtain

$$R \ln \frac{p}{p_0} = a_1 \left( \frac{1}{T_0} - \frac{1}{T} \right) + a_2 \ln \frac{T}{T_0} + a_3 (T - T_0).$$

The reference point can be also the critical point. In this case we have an accurate representation of the saturation line that satisfy the *Clausius* equation, is consistent with definition of the latent heat of vaporization and with the critical point of the fluid. This equation for is very useful in the practice. If the pressure is the known variable it is useful to approximate the above equation with  $T'(p) = 1/\left[ a + b(\log_{10} p) \right]$  as a first approximation and the solve it by iterations

$$T^{n+1} = a_1 \left/ \left[ \frac{1}{T_c} - R \ln \frac{p}{p_c} - a_2 \ln \frac{T}{T_c} - a_3 (T - T_c) \right] \right.$$

### 3.2.4.5 Equilibrium of gas solution in water

Consider absorption of a gas  $i$  by water having temperature  $T_2$ . The pressure in the gas space is  $p_{i,gas}$ . The partial pressure of the specie  $i$  in the water is  $p_{i,aq}$ . The stoichiometric reaction is then  $-1 \times I_{gas} + 1 \times I_{aq} = 0$ . The solution is reaching equilibrium. The equilibrium is described by Eq. (3.169a)

$$\left( \frac{p_{i,gas}}{p_0} \right)^{-1} \left( \frac{p_{i,aq}}{p_0} \right)^1 = \exp \left\{ -\frac{M_i}{RT_2} \left[ h_{i0,aq} - h_{i0,gas} - T_2 (s_{i0,aq} - s_{i0,gas}) \right] \right\}.$$

The equation can be rewritten as,

$$\frac{p_{i,aq}/p}{p_{i,gas}/p} = \exp \left\{ -\frac{M_i}{RT_2} \left[ \begin{array}{l} h_{i0,aq}(T_2, p_0) - h_{i0,gas}(T_2, p_0) \\ -T_2 [s_{i0,aq}(T_2, p_0) - s_{i0,gas}(T_2, p_0)] \end{array} \right] \right\}.$$

Considering the gas-water solution as a perfect fluid solution we obtain for the saturated molar concentration of specie  $i$  in the water the final expression

$$Y_{i,aq} = Y_{i,gas} \exp \left\{ -\frac{M_i}{RT_2} \left[ \begin{array}{l} h_{i0,aq}(T_2, p_0) - h_{i0,gas}(T_2, p_0) \\ -T_2 [s_{i0,aq}(T_2, p_0) - s_{i0,gas}(T_2, p_0)] \end{array} \right] \right\}.$$

Note that the specific enthalpies and entropies at the reference state for gas and gas dissolved in water have to differ in order to obtain a solution. Several values for aqueous gas solutions are tabulated in *Oelkers et al. (1995)*. For known relations  $c_{pi}(T)$  and  $c_{pi,aq}(T)$  the above expression may be explicitly expressed as a function of temperature.

If the exponential expression is only a temperature function the results reflects the empirically found *Henry's law*. The *Henry's law* says

“The mass of non condensable gas dissolved in a liquid is proportional to the partial pressure of the gas above the liquid with which it is in equilibrium”

Frequently in the chemical thermodynamic literature the solubility data are approximated by the *Henry's law* in the following form

$$p_{li} = k_{H,2i}(p, T_2) Y_{2i,\infty}.$$

Here the partial pressure of specie  $i$  in the gas phase is  $p_{li}$  and  $Y_{2i,\infty}$  is the saturation molar concentration of the same specie in the liquid.

$$k_{H,2i}(p, T_2) = p \left/ \exp \left\{ -\frac{M_i}{RT_2} \left[ \begin{array}{l} h_{i0,aq}(T_2, p_0) - h_{i0,gas}(T_2, p_0) \\ -T_2 [s_{i0,aq}(T_2, p_0) - s_{i0,gas}(T_2, p_0)] \end{array} \right] \right\} \right.$$

is called the *Henry's coefficient*. Note, that this idealization does not hold for many cases. There is a non linear pressure dependence of the *Henry's coefficient*

too especially for very high pressures. Knowing the *Henry's* coefficient the molar concentration of the saturated solution is then

$$Y_{2i,\infty} = p_{li}/k_{H,2i}(T_2),$$

and the corresponding mass concentration is

$$C_{2i,\infty} = \frac{Y_{2i,\infty} M_{2i,\infty}}{Y_{2i,\infty} M_{2i,\infty} + (1 - Y_{2i,\infty}) M_{H_2O}}.$$

If the mass concentration of the saturated solution is known the molar concentration is easily computed by

$$Y_{2i,\infty} = \frac{C_{2i,\infty}/M_{2i,\infty}}{C_{2i,\infty}/M_{2i,\infty} + (1 - C_{2i,\infty})/M_{H_2O}}.$$

### 3.2.5 Partial derivatives in the equation of state $\rho = \rho(p, \varphi, C_{2,\dots,i_{\max}})$ , where $\varphi = s, h, e$

*Enthalpy concept*  $\varphi = h$ . Finally, we replace  $dT$  in Eq. (3.96) using Eq. (3.47) and substitute

$$\left(\frac{\partial \rho}{\partial p}\right)_{h, \text{all } C's} = \left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C's} - \left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C's} \frac{1}{c_p} \left(\frac{\partial h}{\partial p}\right)_{T, \text{all } C's}, \quad (3.197)$$

$$\left(\frac{\partial \rho}{\partial h}\right)_{p, \text{all } C's} = \frac{1}{c_p} \left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C's}, \quad (3.198)$$

$$\begin{aligned} & \left(\frac{\partial \rho}{\partial C_i}\right)_{p, h, \text{all } C's \text{ except } C_i} \\ &= \left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all } C's \text{ except } C_i} - \left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C's} \frac{1}{c_p} \left(\frac{\partial h}{\partial C_i}\right)_{p, T, \text{all } C's \text{ except } C_i}, \end{aligned} \quad (3.199)$$

to obtain

$$d\rho = \left(\frac{\partial \rho}{\partial p}\right)_{h, \text{all } C's} dp + \left(\frac{\partial \rho}{\partial h}\right)_{p, \text{all } C's} dh + \sum_{i=2}^{i_{\max}} \left(\frac{\partial \rho}{\partial C_i}\right)_{p, h, \text{all } C's \text{ except } C_i} dC_i \quad (3.200)$$

which is the differential form for the equation of state

$$\rho = \rho(p, h, C_{2, \dots, n_{\max}}). \quad (3.201)$$

*Energy concept*  $\varphi = e$ . Again, the derivation already presented for the specific enthalpy is formally identical with the derivation for the specific internal energy.

One has to replace formally  $h$  with  $e$  and write instead  $c_{pi}$ ,  $\left(\frac{\partial e_i}{\partial T}\right)_{p_i, T}$ .

*Entropy concept*  $\varphi = s$ : We replace  $dT$  in Eq. (3.118) using Eq. (3.47) and substitute

$$\left(\frac{\partial \rho}{\partial p}\right)_{s, \text{all}_{-}C's} = \left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all}_{-}C's} - \left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all}_{-}C's} \frac{\rho \left(\frac{\partial h}{\partial p}\right)_{T, \text{all}_{-}C's}^{-1}}{\rho c_p} = \frac{1}{a^2} = \frac{\rho}{\kappa p}, \quad (3.202)$$

$$\left(\frac{\partial \rho}{\partial s}\right)_{p, \text{all}_{-}C's} = \left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all}_{-}C's} \frac{T}{c_p}, \quad (3.203)$$

$$\begin{aligned} & \left(\frac{\partial \rho}{\partial C_i}\right)_{p, s, \text{all}_{-}C's_{-}\text{except}_{-}C_i} \\ &= \left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all}_{-}C's_{-}\text{except}_{-}C_i} - \left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all}_{-}C's} \frac{T}{c_p} \left(\frac{\partial s}{\partial C_i}\right)_{p, T, \text{all}_{-}C's_{-}\text{except}_{-}C_i}, \end{aligned} \quad (3.204)$$

to obtain

$$d\rho = \frac{d\rho}{a^2} + \left(\frac{\partial \rho}{\partial s}\right)_{p, \text{all}_{-}C's} ds + \sum_{i=2}^{i_{\max}} \left(\frac{\partial \rho}{\partial C_i}\right)_{p, s, \text{all}_{-}C's_{-}\text{except}_{-}C_i} dC_i \quad (3.205)$$

or

$$d\rho = \frac{\rho}{\kappa} \frac{dp}{p} + \left(\frac{\partial \rho}{\partial s}\right)_{p, \text{all}_{-}C's} ds + \sum_{i=2}^{i_{\max}} \left(\frac{\partial \rho}{\partial C_i}\right)_{p, s, \text{all}_{-}C's_{-}\text{except}_{-}C_i} dC_i, \quad (3.206)$$

which is the differential form for the equation of state

$$\rho = \rho(p, s, C_{2, \dots, n_{\max}}). \quad (3.207)$$

The value  $\kappa$  in Eq. (3.206) is defined as



$$\begin{aligned} \kappa &= \frac{\rho}{p} \left/ \left[ \left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all\_}C's} - \left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all\_}C's} \frac{\rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all\_}C's} - 1}{\rho c_p} \right] \right. \\ &= c_p \left/ \left[ \frac{p}{\rho} \left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all\_}C's} c_p + \frac{T}{\rho} \left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all\_}C's} \bar{R} \right] = \frac{1}{\frac{p}{\rho} \left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all\_}C's}} \frac{c_p}{c_v}, \end{aligned} \quad (3.208)$$

where

$$\bar{R} = \frac{p}{\rho T} \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all\_}C's} \right], \quad (3.209)$$

$$c_v = c_p + \frac{\frac{T}{\rho} \left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all\_}C's}}{\frac{p}{\rho} \left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all\_}C's}} \bar{R} = c_p + \frac{\left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all\_}C's}}{\left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all\_}C's}} \frac{1}{\rho} \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all\_}C's} \right], \quad (3.210)$$

is generally valid for real gases as well for perfect gases.  $\bar{R}$  is used to denote the *pseudo gas constant*, which is of course not in fact a constant for real gases.  $\kappa$  is the isentropic exponent for real gas mixtures. Note, that  $\kappa$  is not identical to the isentropic exponent for perfect gases in all cases. Only for mixtures of perfect gases does the *pseudo gas constant*  $\bar{R}$  reduce to the gas constant for perfect mixtures because  $(\partial h / \partial p)_{T, \text{all\_}C's} = 0$ , and the expression above for the  $\kappa$  coefficient for real gases then reduces to the usual expression defining the isentropic exponent for perfect gas. For proof, we insert Eqs. (3.75), (3.76) into Eq. (3.208). The result is

$$\kappa = \frac{c_p}{c_p - R}. \quad (3.211)$$

An example for application of the general theory to an air-steam mixture is given in Appendix 3.1.

### 3.3 Mixture of liquid and microscopic solid particles of different chemical substances

This is a limiting case of the theory presented in the previous Section 3.2.

### 3.3.1 Partial derivatives in the equation of state $\rho = \rho(p, T, C_2, \dots, C_{i_{\max}})$

For this particular case the derivatives in Eq. (3.47) given by Eqs. (3.80) through (3.82) reduce to

$$\left(\frac{\partial \rho}{\partial p}\right)_{T, \text{all } C's} = C_1 \frac{\rho^2}{\rho_1^2} \left(\frac{\partial \rho_1}{\partial p}\right)_T, \quad (3.212)$$

$$\left(\frac{\partial \rho}{\partial T}\right)_{p, \text{all } C's} = C_1 \frac{\rho^2}{\rho_1^2} \left(\frac{\partial \rho_1}{\partial T}\right)_p, \quad (3.213)$$

$$\left(\frac{\partial \rho}{\partial C_i}\right)_{p, T, \text{all } C's \text{ except } C_i} = \rho^2 \left(\frac{1}{\rho_1} - \frac{1}{\rho_i}\right). \quad (3.214)$$

For missing inert components in the liquid,  $\sum_{i=2}^{i_{\max}} C_i = 0$ , and  $\rho = \rho_1$ , Eqs. (3.213) and (3.214) take values characteristic of the pure liquid. If the mixture consists solely of inert components,  $\sum_{i=2}^{i_{\max}} C_i = 1$ , Eqs. (3.212) and (3.213) are not defined. In this case  $\delta\rho = 0$  holds by definition.

Another consequence of Eq. (3.214) is that the concentration change for a given component causes a density change if and only if the density differs from the density of the liquid,  $\rho_i \neq \rho_1$ . In all other cases,  $\rho_i = \rho_1$ , the concentration change does not lead to a change in mixture density.

### 3.3.2 Partial derivatives in the equation of state $T = T(p, \varphi, C_2, \dots, C_{i_{\max}})$ where $\varphi = h, e, s$

For this simple case we obtain the following.

*Enthalpy concept*  $\varphi = h$  :

$$c_p = \sum_{i=1}^{i_{\max}} C_i c_{pi}, \quad (3.215)$$

$$\left(\frac{\partial h}{\partial p}\right)_{T, \text{all } C's} = C_1 \left(\frac{\partial h_1}{\partial p}\right)_T. \quad (3.216)$$

Energy concept  $\varphi = e$  :

$$\left(\frac{\partial e}{\partial T}\right)_p = \sum_{i=1}^{i_{\max}} C_i \left(\frac{\partial e_i}{\partial T}\right)_p, \quad (3.217)$$

$$\left(\frac{\partial e}{\partial p}\right)_{T, \text{all } C_i} = C_1 \left(\frac{\partial e_1}{\partial p}\right)_T. \quad (3.218)$$

Entropy concept  $\varphi = s$  :

$$\left(\frac{\partial s}{\partial C_i}\right)_{p, T, \text{all } C_j \text{ except } C_i} = s_i - s_1. \quad (3.219)$$

If the increments  $\delta p$ ,  $\delta s$  and  $\delta C_i$  are known, the temperature increments  $\delta T$  can be computed from Eq. (3.118) in a manner analogous to that used in the preceding section. In the event that the increments  $\delta p$ ,  $\delta s$  and  $\delta C_i$  are large in magnitude, it is better to take into account the non-linear character of Eq. (3.118), as already demonstrated using Eqs. (3.131) through (3.141). If  $\Delta s^* = 0$ , Eq. (3.137) will once again result. For  $p = \text{const}$  Eq. (3.136) yields

$$T = T_a e^{\Delta s^*}. \quad (3.220)$$

Usually  $\bar{R}$  is significantly smaller for liquids or mixtures of liquids than for gases because of the marked differences in the densities. The temperature change caused by compression or decompression is thus observable after significant changes in pressure. Usually, the entropy change due to the heat and mass transfer causes a change in the liquid temperature  $T \approx T_a e^{\Delta s^*}$ . The partial derivatives in the equation of state  $\rho = \rho(p, s, C_{2, \dots, n_{\max}})$  are formally identical to those shown in Section 2.4, and are therefore not presented here.

### 3.4 Single-component equilibrium fluid

The restriction applied that one velocity field is in thermodynamic equilibrium is very strong and is, in fact, not necessary for an adequate description of multi-phase flow behavior. This assumption is, however, used in a number of applications to yield an approximate estimate of the order of magnitude of the processes. For this reason, this section briefly summarizes the derivatives calculated previously (see *Kolev* (1986b), for example) for equilibrium mixtures.

Consider the properties of a fluid, assuming that it is in a thermodynamic equilibrium. Additional use will be made here of the following superscripts:

' , saturated liquid,

" , saturated steam,  
 "' , saturated solid phase.

The properties on the saturation lines are functions of pressure only

$$T' = T'(p), T'' = T''(p), \rho', \rho'', \rho''', s', s'', s''' = f(p).$$

The same is valid for the properties

$$d\rho' / dp, d\rho'' / dp, d\rho''' / dp, ds' / dp, ds'' / dp, ds''' / dp = f(p).$$

As a result, four regions are distinguished.

### 3.4.1 Superheated vapor

For  $s > s''$  superheated steam exists with properties that are functions of the pressure and of the temperature

$$\rho = \rho(p, T), \quad (3.221)$$

$$d\rho = \left( \frac{\partial \rho}{\partial p} \right)_T dp + \left( \frac{\partial \rho}{\partial T} \right)_p dT, \quad (3.222)$$

$$\left( \frac{\partial \rho}{\partial p} \right)_T, \left( \frac{\partial \rho}{\partial T} \right)_p = f(p, T), \quad (3.223)$$

$$T = T(p, s), \quad (3.224)$$

$$\frac{dT}{T} = \frac{ds}{c_p} + \frac{\bar{R}}{c_p} \frac{dp}{p}, \quad (3.225)$$

$$c_p = c_p(p, T), \quad (3.226)$$

$$\bar{R} = \frac{p}{\rho T} \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_T \right] = \bar{R}(p, T), \quad (3.227)$$

$$\rho = \rho(p, s), \quad (3.228)$$

$$d\rho = \frac{dp}{a^2} + \left( \frac{\partial \rho}{\partial s} \right)_p ds = \frac{\rho}{\kappa} \frac{dp}{p} + \left( \frac{\partial \rho}{\partial s} \right)_p ds, \quad (3.229)$$

$$\frac{1}{a^2} = \frac{\rho}{\kappa p} = \left( \frac{\partial \rho}{\partial p} \right)_T - \left( \frac{\partial \rho}{\partial T} \right)_p \frac{\rho \left( \frac{\partial h}{\partial p} \right)_T - 1}{\rho c_p}, \quad (3.230)$$

$$\left(\frac{\partial p}{\partial s}\right)_p = \left(\frac{\partial p}{\partial T}\right)_p \frac{T}{c_p}, \quad (3.231)$$

where

$$c_p = c_p(p, T), \quad (3.232)$$

$$\left(\frac{\partial h}{\partial p}\right)_T = f(p, T). \quad (3.233)$$

A very useful retaliation for gaining this derivative from experimental data is obtained by using the so called *Schwarz* theorem:

$$\left[\frac{\partial}{\partial T}\left(\frac{\partial s}{\partial p}\right)_T\right]_p = \left[\frac{\partial}{\partial p}\left(\frac{\partial s}{\partial T}\right)_p\right]_T, \quad \left\{\frac{\partial}{\partial T}\left[\frac{1}{T}\left(\frac{\partial h}{\partial p}\right)_T - \frac{v}{T}\right]\right\}_p = \left\{\frac{\partial}{\partial p}\left[\frac{1}{T}\left(\frac{\partial h}{\partial T}\right)_p\right]\right\}_T, \\ -\frac{1}{T^2}\left(\frac{\partial h}{\partial p}\right)_T + \frac{1}{T}\left(\frac{\partial^2 h}{\partial p \partial T}\right) + \frac{1}{T^2}v - \frac{1}{T}\left(\frac{\partial v}{\partial T}\right)_p = \frac{1}{T}\left(\frac{\partial^2 h}{\partial T \partial p}\right),$$

resulting in

$$\left(\frac{\partial h}{\partial p}\right)_T = v - T\left(\frac{\partial v}{\partial T}\right)_p.$$

For a perfect gas, the state equations and the derivatives are easily computed. For real gases, it is assumed that analytical expressions for the state equation are known, and therefore analytical expressions for the derivatives required can easily be derived.

### 3.4.2 Reconstruction of equation of state by using a limited amount of data available

For many new application fields the information regarding thermodynamic data is very limited but order of magnitude analyses are required for the practice. Therefore we present next a few brief examples of how to construct an approximate closed description of the thermodynamic state using only a few data points.

#### 3.4.2.1 Constant thermal expansion coefficient and isothermal compressibility coefficient

The volumetric thermal expansion coefficient defined as

$$\beta = \frac{1}{v} \left( \frac{dv}{dT} \right)_p = - \frac{1}{\rho} \left( \frac{d\rho}{dT} \right)_p, \quad (3.234)$$

and the isothermal coefficient of compressibility defined as

$$k = - \frac{1}{v} \left( \frac{dv}{dp} \right)_T = \frac{1}{\rho} \left( \frac{d\rho}{dp} \right)_T, \quad (3.235)$$

are measured experimentally. As an example solid  $Al_2O_3$  has the properties  $\beta = 5.0 \times 10^{-5} K^{-1}$ ,  $k = 2.8 \times 10^{-12} Pa^{-1}$  *McCahan and Shepherd (1993)*. The reciprocal value of the isothermal compressibility is known as elasticity modulus,  $E = k^{-1}$ , having dimensions of pressure,  $Pa$ . Thus the equation of state

$$\rho = \rho(T, p) \quad (3.236)$$

can be easily constructed starting with

$$\frac{d\rho}{\rho} = \frac{1}{\rho} \left( \frac{\partial \rho}{\partial T} \right)_p dT + \frac{1}{\rho} \left( \frac{\partial \rho}{\partial p} \right)_T dp = -\beta dT + k dp, \quad (3.237)$$

and integrating with respect to some reference temperature and pressure

$$\rho = \rho_0 \exp[-\beta(T - T_0) + k(p - p_0)]. \quad (3.238)$$

For construction of the equation of state for specific enthalpy, entropy and internal energy the measured dependence on the specific heat as a function of temperature

$$c_p = c_p(T) \quad (3.239)$$

is necessary. Using the differential relationships

$$\begin{aligned} dh &= c_p dT + \frac{1}{\rho} \left[ 1 + T \frac{1}{\rho} \left( \frac{\partial \rho}{\partial T} \right)_p \right] dp \\ &= c_p dT + \frac{(1 - \beta T)}{\rho_0} \exp[\beta(T - T_0) - k(p - p_0)] dp, \end{aligned} \quad (3.240)$$

$$ds = \frac{c_p}{T} dT + \frac{1}{\rho^2} \left( \frac{\partial \rho}{\partial T} \right)_p dp = \frac{c_p}{T} dT - \frac{\beta}{\rho_0} \exp[\beta(T - T_0) - k(p - p_0)] dp, \quad (3.241)$$

$$\begin{aligned} de &= dh - \frac{1}{\rho} dp + \frac{p}{\rho^2} d\rho = \left[ c_p + \frac{p}{\rho^2} \left( \frac{\partial \rho}{\partial T} \right)_p \right] dT + \frac{1}{\rho} \left[ \frac{T}{\rho} \left( \frac{\partial \rho}{\partial T} \right)_p + \frac{p}{\rho} \left( \frac{\partial \rho}{\partial p} \right)_T \right] dp \\ &= \left( c_p - \frac{p}{\rho} \beta \right) dT + \frac{1}{\rho} (-\beta T + kp) dp \end{aligned}$$

$$\begin{aligned}
&= c_p dT - \frac{p \exp[-k(p-p_0)]}{\rho_0} \beta \exp[\beta(T-T_0)] dT \\
&- \frac{\beta T}{\rho_0} \exp[\beta(T-T_0)] \exp[-k(p-p_0)] dp \\
&+ \frac{1}{\rho_0} \exp[\beta(T-T_0)] \exp[-k(p-p_0)] k p dp, \quad (3.242)
\end{aligned}$$

and integrating from some reference pressure and temperature we obtain

$$h = h_0 + \int_{T_0}^T c_p(T) dT - \frac{(1-\beta T)}{k\rho_0} \exp[\beta(T-T_0)] \{ \exp[-k(p-p_0)] - 1 \}, \quad (3.243)$$

$$s = s_0 + \int_{T_0}^T \frac{c_p(T)}{T} dT + \frac{1}{k\rho_0} \beta \exp[\beta(T-T_0)] \{ \exp[-k(p-p_0)] - 1 \}, \quad (3.244)$$

$$\begin{aligned}
e &= e_0 + \int_{T_0}^T c_p(T) dT - \frac{p}{\rho_0} \exp[-k(p-p_0)] \{ \exp[\beta(T-T_0)] - 1 \} \\
&+ \frac{\beta T}{k\rho_0} \exp[\beta(T-T_0)] \{ \exp[-k(p-p_0)] - 1 \} \\
&+ \frac{1}{\rho_0} \exp[\beta(T-T_0)] \int_{p_0}^p \exp[-k(p-p_0)] k p dp. \quad (3.245)
\end{aligned}$$

Many authors enforce  $s = 0$  for  $T = 0$ . This is called in some text books the third law of the thermodynamics *Cordfunke* and *Konings* (1990) as proposed by *Max Plank* in 1912. Note that the solution of the last integral is quite complicated

$$\int_{p_0}^p \exp[-k(p-p_0)] k p dp = -\frac{\sqrt{\pi}}{2k} \exp\left(\frac{1}{4} k^2 p_0^2\right) \left[ \operatorname{erf}\left(-kp + \frac{1}{2} k p_0\right) + \operatorname{erf}\left(\frac{1}{2} k p_0\right) \right]. \quad (3.246)$$

In addition from the equation of state in the form

$$d\rho = -\beta\rho \frac{T}{c_p} ds + \left\{ \rho k - \frac{T}{c_p} \beta^2 \right\} dp = \left( \frac{\partial \rho}{\partial s} \right)_p ds + \left( \frac{\partial \rho}{\partial p} \right)_s dp, \quad (3.247)$$

we obtain the derivatives

$$\left( \frac{\partial \rho}{\partial s} \right)_p = -T \beta \rho / c_p, \quad (3.248)$$

$$\left(\frac{\partial \rho}{\partial p}\right)_s = \rho k - T \beta^2 / c_p. \quad (3.249)$$

The velocity of sound is therefore

$$a = \frac{1}{\sqrt{\rho k - T \beta^2 / c_p}}. \quad (3.250)$$

The derivatives of the enthalpy, entropy and specific internal energy with respect to pressure and temperature follow immediately from Eqs. (3.228) through (3.242).

$$\left(\frac{\partial h}{\partial T}\right)_p = c_p, \quad (3.251)$$

$$\left(\frac{\partial h}{\partial p}\right)_T = \frac{(1 - \beta T)}{\rho_0} \exp[\beta(T - T_0) - k(p - p_0)], \quad (3.252)$$

$$\left(\frac{\partial s}{\partial T}\right)_p = \frac{c_p}{T}, \quad (3.253)$$

$$\left(\frac{\partial s}{\partial p}\right)_T = -\frac{\beta}{\rho_0} \exp[\beta(T - T_0) - k(p - p_0)], \quad (3.254)$$

$$\left(\frac{\partial e}{\partial T}\right)_p = c_p - \frac{\beta p}{\rho_0} \exp[\beta(T - T_0) - k(p - p_0)], \quad (3.255)$$

$$\left(\frac{\partial e}{\partial p}\right)_T = \frac{kp - \beta T}{\rho_0} \exp[\beta(T - T_0) - k(p - p_0)]. \quad (3.256)$$

### 3.4.2.2 Properties known only at given a pressure as temperature functions

Frequently in the literature there are measured properties at atmospheric pressure  $p_0$  given as temperature functions only

$$\rho_{p_0} = \rho_{p_0}(T), \quad (3.257)$$

or

$$\left(\frac{\partial \rho}{\partial T}\right)_p = f(T), \quad (3.258)$$

$$c_p = c_p(T). \quad (3.259)$$



In addition the velocity of sound is measured

$$a = a(T). \quad (3.260)$$

For practical application of this information the reconstruction of the properties in consistent data functions is required. Next we construct the analytical form of the density as a function of temperature and pressure. The density derivative with respect to pressure at constant temperature

$$\left(\frac{d\rho}{dp}\right)_T = \frac{1}{a^2} + T\beta^2 / c_p = \frac{1}{a^2} + \frac{T}{\rho^2 c_p} \left(\frac{d\rho}{dT}\right)_p^2 = f(T) \quad (3.261)$$

is obviously a function of the temperature only. Therefore the density can be computed by integrating analytically the equation of state

$$d\rho = \left(\frac{\partial\rho}{\partial T}\right)_p dT + \left(\frac{\partial\rho}{\partial p}\right)_T dp. \quad (3.262)$$

The result is

$$\rho = \rho_{p_0} + \left(\frac{\partial\rho}{\partial p}\right)_T (p - p_0). \quad (3.263)$$

Using the differential equations

$$dh = c_p dT + \frac{1}{\rho} \left[ 1 + T \frac{1}{\rho} \left(\frac{\partial\rho}{\partial T}\right)_p \right] dp = c_p dT + \frac{dp}{\rho} + T \frac{1}{\rho^2} \left(\frac{\partial\rho}{\partial T}\right)_p dp, \quad (3.264)$$

$$ds = \frac{c_p}{T} dT + \frac{1}{\rho^2} \left(\frac{\partial\rho}{\partial T}\right)_p dp, \quad (3.265)$$

$$\begin{aligned} de &= dh - \frac{1}{\rho} dp + \frac{p}{\rho^2} d\rho = \left[ c_p + \frac{p}{\rho^2} \left(\frac{\partial\rho}{\partial T}\right)_p \right] dT + \frac{1}{\rho} \left[ \frac{T}{\rho} \left(\frac{\partial\rho}{\partial T}\right)_p + \frac{p}{\rho} \left(\frac{\partial\rho}{\partial p}\right)_T \right] dp \\ &= c_p dT + \frac{p}{\rho^2} \left(\frac{\partial\rho}{\partial T}\right)_p dT + T \left(\frac{\partial\rho}{\partial T}\right)_p \frac{dp}{\left[ \rho_{p_0} + \left(\frac{\partial\rho}{\partial p}\right)_T (p - p_0) \right]^2} \\ &\quad + \left(\frac{\partial\rho}{\partial p}\right)_T \frac{p dp}{\left[ \rho_{p_0} + \left(\frac{\partial\rho}{\partial p}\right)_T (p - p_0) \right]^2}, \end{aligned} \quad (3.266)$$

keeping in mind that

$$\int_{p_0}^p \frac{dp}{\rho_{p_0} + \left(\frac{\partial \rho}{\partial p}\right)_T (p - p_0)} = \frac{1}{\left(\frac{\partial \rho}{\partial p}\right)_T} \ln \frac{\rho}{\rho_{p_0}}, \quad (3.267)$$

$$\int_{p_0}^p \frac{dp}{\left[\rho_{p_0} + \left(\frac{\partial \rho}{\partial p}\right)_T (p - p_0)\right]^2} = \frac{p - p_0}{\rho_{p_0} \rho}, \quad (3.268)$$

$$\begin{aligned} & \left(\frac{\partial \rho}{\partial p}\right)_T \int_{p_0}^p \frac{p dp}{\left[\rho_{p_0} + \left(\frac{\partial \rho}{\partial p}\right)_T (p - p_0)\right]^2} \\ &= \frac{1}{\left(\frac{\partial \rho}{\partial p}\right)_T} \left\{ \left(\frac{1}{\rho} - \frac{1}{\rho_{p_0}}\right) \left[\rho_{p_0} - \left(\frac{\partial \rho}{\partial p}\right)_T p_0\right] + \ln \frac{\rho}{\rho_{p_0}} \right\}, \end{aligned} \quad (3.269)$$

and integrating from some reference pressure and temperature we obtain

$$h = h_0 + \int_{T_0}^T c_p(T) dT + \frac{1}{\left(\frac{\partial \rho}{\partial p}\right)_T} \ln \frac{\rho}{\rho_{p_0}} + T \left(\frac{\partial \rho}{\partial T}\right)_p \frac{p - p_0}{\rho_{p_0} \rho}, \quad (3.270)$$

$$\begin{aligned} s &= s_0 + \int_{T_0}^T \frac{c_p}{T} dT + \left(\frac{\partial \rho}{\partial T}\right)_p \frac{dp}{\left[\rho_{p_0} + \left(\frac{\partial \rho}{\partial p}\right)_T (p - p_0)\right]^2} \\ &= s_0 + \int_{T_0}^T \frac{c_p}{T} dT + \left(\frac{\partial \rho}{\partial T}\right)_p \frac{p - p_0}{\rho_{p_0} \rho}, \end{aligned} \quad (3.271)$$

$$\begin{aligned} e &= e_0 + \int_{T_0}^T c_p dT + p \int_{T_0}^T \frac{1}{\rho^2} \left(\frac{\partial \rho}{\partial T}\right)_p dT \\ &+ T \left(\frac{\partial \rho}{\partial T}\right)_p \frac{p - p_0}{\rho_{p_0} \rho} + \frac{1}{\left(\frac{\partial \rho}{\partial p}\right)_T} \left\{ \left(\frac{1}{\rho} - \frac{1}{\rho_{p_0}}\right) \left[\rho_{p_0} - \left(\frac{\partial \rho}{\partial p}\right)_T p_0\right] + \ln \frac{\rho}{\rho_{p_0}} \right\}. \end{aligned} \quad (3.272)$$

Again the derivatives are

$$\left(\frac{\partial h}{\partial T}\right)_p = c_p, \quad (3.273)$$

$$\left(\frac{\partial h}{\partial p}\right)_T = \frac{1}{\rho} \left[ 1 + T \frac{1}{\rho} \left(\frac{\partial \rho}{\partial T}\right)_p \right], \quad (3.274)$$

$$\left(\frac{\partial s}{\partial T}\right)_p = \frac{c_p}{T}, \quad (3.275)$$

$$\left(\frac{\partial s}{\partial p}\right)_T = \frac{1}{\rho^2} \left(\frac{\partial \rho}{\partial T}\right)_p, \quad (3.276)$$

$$\left(\frac{\partial e}{\partial T}\right)_p = \left[ c_p + \frac{p}{\rho^2} \left(\frac{\partial \rho}{\partial T}\right)_p \right], \quad (3.277)$$

$$\left(\frac{\partial e}{\partial p}\right)_T = \frac{1}{\rho^2} \left[ T \left(\frac{\partial \rho}{\partial T}\right)_p + p \left(\frac{\partial \rho}{\partial p}\right)_T \right]. \quad (3.278)$$

### 3.4.2.3 Constant density approximations

For an idealized case of constant density we obtain

$$\rho = \rho_0, \quad (3.279)$$

$$h = h_0 + \int_{T_0}^T c_p(T) dT, \quad (3.280)$$

$$s = s_0 + \int_{T_0}^T \frac{c_p}{T} dT, \quad (3.281)$$

$$e = e_0 + \int_{T_0}^T c_p dT. \quad (3.282)$$

For zero initial values at the reference point we have

$$e \equiv h. \quad (3.283)$$

Within this concept internal pressure for the incompressible field is not defined. Submerged in fluid the interface averaged stress is the pressure of the system.

### 3.4.2.4 Perfect gas approximations

For a perfect gas defined by

$$p = RT\rho \quad (3.284)$$

with constant specific heat at constant pressure the differential Eqs. (3.241, 3.242, 3.243) result in

$$dh = c_p dT + \frac{1}{\rho} \left[ 1 + T \frac{1}{\rho} \left( \frac{\partial \rho}{\partial T} \right)_p \right] dp = c_p dT, \quad (3.285)$$

$$ds = c_p \frac{dT}{T} + \frac{1}{\rho^2} \left( \frac{\partial \rho}{\partial T} \right)_p dp = c_p \frac{dT}{T} - R \frac{dp}{p}, \quad (3.286)$$

$$\begin{aligned} de &= dh - \frac{1}{\rho} dp + \frac{p}{\rho^2} d\rho = \left[ c_p + \frac{p}{\rho^2} \left( \frac{\partial \rho}{\partial T} \right)_p \right] dT + \frac{1}{\rho} \left[ \frac{T}{\rho} \left( \frac{\partial \rho}{\partial T} \right)_p + \frac{p}{\rho} \left( \frac{\partial \rho}{\partial p} \right)_T \right] dp \\ &= (c_p - R) dT = c_v dT, \end{aligned} \quad (3.287)$$

or

$$h = c_p (T - T_0), \quad (3.288)$$

$$e = c_v (T - T_0), \quad (3.289)$$

$$s = c_p \ln \frac{T}{T_0} - R \ln \frac{p}{p_0}. \quad (3.290)$$

Obviously for  $p \leq 0 \ln p/p_0$  and consequently  $s$  is not defined. With other words: Gases does not exist at zero pressure and gases do not possess tension state. Modeling multi-phase flows using multi-velocity field models offers the possibility of postulating one of the mixtures described above for each of the fields. The combination of velocity fields and an appropriate number of components in each field make it possible to describe a large variety of flows observed in nature and in the field of engineering.

### 3.4.3 Vapor-liquid mixture in thermodynamic equilibrium

For  $s' < s < s''$  a saturated liquid coexists with its saturated vapor in thermodynamic equilibrium after having enough time for relaxation. The mass concentration of the vapor in the equilibrium mixture is denoted with  $C''$ . The mass concentration is used to generate the following definitions for the specific entropy and the density of the mixtures

$$s = C'' s'' + (1 - C'') s', \quad (3.291)$$

$$\frac{1}{\rho} = \frac{C''}{\rho''} + \frac{1 - C''}{\rho'}. \quad (3.292)$$

As  $p$  (and therefore  $s''$ ,  $s'$ ) and  $s$  are known, this permits easy computation of  $C''$  from Eq. (3.291), i.e.,

$$C'' = \frac{s - s'}{s'' - s'}. \quad (3.293)$$

As  $C''$  and the pressure are known, the density is uniquely defined by Eq. (3.292), i.e.,

$$\rho = \rho(p, s), \quad (3.294)$$

or

$$d\rho = \frac{dp}{a^2} + \left( \frac{\partial \rho}{\partial s} \right)_p ds. \quad (3.295)$$

The derivatives can be easily obtained after differentiating Eqs. (3.291) and (3.292) and after eliminating  $dC''$ . The result is

$$\left( \frac{\partial \rho}{\partial s} \right)_p = - \frac{\rho^2}{\rho' \rho''} \frac{\rho' - \rho''}{s'' - s'}, \quad (3.296)$$

$$\frac{1}{a^2} = \frac{\rho}{\kappa p} = \rho^2 \left\{ \frac{C''}{\rho'^2} \frac{d\rho''}{dp} + \frac{1 - C''}{\rho'^2} \frac{d\rho'}{dp} - \frac{1}{\rho^2} \left( \frac{\partial \rho}{\partial s} \right)_p \left[ C'' \frac{ds''}{dp} + (1 - C'') \frac{ds'}{dp} \right] \right\}, \quad (3.297)$$

*Kolev (1986b)*. As Eq. (3.296) shows, the density change with the entropy for constant pressure depends strongly on  $\rho$ , and therefore on the concentration  $C''$ .

### 3.4.4 Liquid-solid mixture in thermodynamic equilibrium

For  $s''' < s < s'$ , an equilibrium mixture of liquid and solid phases (designated with superscripts  $'$  and  $'''$  respectively) exists. Replacing the superscript  $''$  with  $'''$  in the formalism of the previous section the required equation of state and the corresponding derivatives are then obtained.

### 3.4.5 Solid phase

For  $s < s'''$ , a solid phase exists. It is assumed here that analytical expressions for the corresponding derivatives are available. Solid phases can be approximately treated as shown in Section 3.4.2.1.

### 3.5 Extension state of liquids

Superheated liquids are typical examples for liquids in tension state. Their internal pressure is higher than the surrounding pressure and the molecules starts to build clusters of vapor nucleolus. The liquid properties are then computed extrapolating the approximations into the region between the saturation- and the spinoidal line.

Liquids at low pressure can be brought in extension state for a very short time too. This state may lead to negative pressures. Many times in the past, researchers have observed in clever experiments such states. Such states are usual if at low pressure strong pressure waves are reflected. Similarly, fast closing valves interrupting fast liquid flows, results also in tension state of the liquid behind the valves. The question of practical interests is how to compute the properties at negative pressure. The reader will find seldom information to this subject in the literature. *Worthington* (1892) reported "...In the neighborhood of the zero pressure the absolute coefficient of volume elasticity of alcohol is the same for extension as for compression, and so far as the observation shows is constant between pressure of -17 and atmospheres 12." It is confirmed also by *Meyer* (1911), reporting that

"...the extension coefficient of water  $\beta_T = \frac{1}{v} \left( \frac{\partial v}{\partial p} \right)_T$  was measured depending on

the temperature from 0 to 31°C and pressures between -30 to 7 atmospheres and found in agreement with the compression coefficient at this temperatures." *Skripov* et al. (1980) reported properties for heavy water and recommended use of linear change of the density with the pressure in the meta-stable region. Therefore, a useful strategy as a first approximation is to fix the derivatives at the corresponding temperature and minimum pressure and using them to linearly extrapolate densities, enthalpies, entropies etc. Regarding the transport properties, again they can be fixed at the given temperature and minimum pressure for which they are valid and then used in the region of negative pressure until better approach is found.

### Appendix 3.1 Application of the theory to steam-air mixtures

Consider a mixture consisting of steam and air. This mixture is common in nature. To facilitate easy application, the definitions obtained for the partial derivatives are reduced to those for a two-component mixture consisting of one inert and one non-inert component. Here air is designated with *A* and steam with *S*, with air taken as a single ideal gas. The result is

$$\left( \frac{\partial \rho}{\partial p} \right)_{T, all\ C's} = 1 / \left[ C_A R_A T + (1 - C_A) / \left( \frac{\partial \rho_S}{\partial p_S} \right)_T \right], \quad (\text{A.1})$$

$$\left(\frac{\partial \rho}{\partial T}\right)_{p,all\_C's} = \left(\frac{\partial \rho}{\partial p}\right)_{T,all\_C's} \left[ -\rho_A R_A + \left(\frac{\partial \rho_s}{\partial T}\right)_{ps} / \left(\frac{\partial \rho_s}{\partial p_s}\right)_T \right], \quad (A.2)$$

$$\left(\frac{\partial \rho}{\partial C_A}\right)_{p,T,all\_C's\_except\_C_A} = -\rho \left(\frac{\partial \rho}{\partial p}\right)_{T,all\_C's} \left[ R_A T - \frac{1}{(\partial \rho_s / \partial p_s)_T} \right], \quad (A.3)$$

$$\begin{aligned} c_p &= C_A c_{pA} + (1 - C_A) \left\{ c_{pS} + \frac{(\partial h_s / \partial p_s)_T}{(\partial \rho_s / \partial p_s)_T} \left[ (1 - C_A) \left(\frac{\partial \rho}{\partial T}\right)_{p,all\_C's} - \left(\frac{\partial \rho_s}{\partial T}\right)_{ps} \right] \right\} \\ &= C_A c_{pA} + (1 - C_A) c_{pS} \\ &\quad - C_A (1 - C_A) R_A \left[ (1 - C_A) \rho + T \left(\frac{\partial \rho_s}{\partial T}\right)_{ps} \right] \frac{(\partial h_s / \partial p_s)_T}{1 - C_A \left[ 1 - R_A T (\partial \rho_s / \partial p_s)_T \right]}, \end{aligned} \quad (A.4)$$

$$\begin{aligned} \left(\frac{\partial h}{\partial p}\right)_{T,all\_C's} &= \left(\frac{\partial \rho}{\partial p}\right)_{T,all\_C's} (1 - C_A)^2 \left(\frac{\partial h_s}{\partial p_s}\right)_T / \left(\frac{\partial \rho_s}{\partial p_s}\right)_T \\ &= (1 - C_A)^2 \left(\frac{\partial h_s}{\partial p_s}\right)_T / \left\{ 1 - C_A \left[ 1 - R_A T \left(\frac{\partial \rho_s}{\partial p_s}\right)_T \right] \right\}, \end{aligned} \quad (A.5)$$

$$\begin{aligned} &\left(\frac{\partial s}{\partial C_A}\right)_{p,T,all\_C's\_except\_C_i} \\ &= s_A - s_S - \frac{\rho}{T} (1 - C_A) \left[ \left(\frac{\partial h_s}{\partial p_s}\right)_T / \left(\frac{\partial \rho_s}{\partial p_s}\right)_T \right] \left[ 1 - (1 - C_A) \frac{1}{\rho} \left(\frac{\partial \rho}{\partial C_A}\right)_{p,T,all\_C's\_except\_C_i} \right] \\ &= s_A - s_S - (1 - C_A) \left(\frac{\partial h_s}{\partial p_s}\right)_T \rho R_A / \left\{ 1 - C_A \left[ 1 - R_A T \left(\frac{\partial \rho_s}{\partial p_s}\right)_T \right] \right\}, \end{aligned} \quad (A.6)$$

and

$$\left(\frac{\partial T}{\partial p}\right)_{s,all\_C's} = \frac{1}{\rho c_p} \left[ 1 - \rho \left(\frac{\partial h}{\partial p}\right)_{T,all\_C's} \right], \quad (A.7)$$

$$\left(\frac{\partial T}{\partial s}\right)_{p,all\_C's} = \frac{T}{c_p}, \quad (A.8)$$

$$\left(\frac{\partial T}{\partial C_A}\right)_{p,s,\text{all } C'_s \text{ except } C_i} = -\frac{T}{c_p} \left(\frac{\partial s}{\partial C_A}\right)_{p,T,\text{all } C'_s \text{ except } C_A} \quad (\text{A.9})$$

## Appendix 3.2 Useful references for computing properties of single constituents

In this appendix a number of references that provide a useful set of approximations for thermodynamic and transport properties of simple constituents are summarized. These approximations are an example of a simple state equation library.

The library discussed below as an example is used in the computer code IVA3 *Kolev* (1991a, b, c). This library consists of a set of analytical approximations for the following *simple* substances: air, water, steam, uranium dioxide in solid, liquid and equilibrium solid/liquid state. Alternatively, the analytical approximations for stainless steel or corium can be used instead of the approximation for uranium dioxide properties.

For air, the *Irvine* and *Liley* (1984) approximations of  $\rho, c_p, h = f(T)$  and  $s = s(T, p = \text{const})$  are used where the influence of the pressure on the entropy,  $-R \ln(p/p_0)$  is added, where  $p_0$  is a reference pressure (e.g.  $10^5$ ) and  $R$  is the gas constant of air. For steam, the *Irvine* and *Liley* (1984) approximations of  $\rho, c_p, s, h = f(T, p)$  are used. For water the *Rivkin* and *Kremnevskaya* (1977) approximations of  $\rho, c_p, h = f(T, p)$  are used. For metastable water, the above analytical properties from *Rivkin* and *Kremnevskaya* (1977) are extrapolated, taking into account the discussion by *Skripov* et al. (1980). The water entropy as a function of temperature and pressure is computed as follows. First the saturation entropy as a function of liquid temperature is computed,  $s' = s'[p'(T)]$ , using analytical approximations proposed by *Irvine* and *Liley* (1984), with the pressure correction  $s = s[s', p'(T) - p]$  then introduced as proposed by *Garland* and *Hand* (1989).

The analytical approximations by *Irvine* and *Liley* (1984) for the steam/water saturation line,  $p' = p'(T)$  and  $T' = T'(p)$  are used in IVA3. The *Clausius-Clapeyron* equation for  $dp'/dT$  was obtained by taking the first derivative of the analytical approximation with respect to temperature.

The steam/water saturation properties are computed as a function of  $p$  and  $T'(p)$  using the above approximations.

Analytical approximations for the properties  $\rho, c_p, s, h = f(T)$  of solid and liquid uranium dioxide as proposed by *Fischer* (1990), *Chawla* et al. (1981) and *Fink* et al. (1982) are used.

For the solid stainless steel properties  $\rho, c_p, s, h = f(T)$  the analytical approximations proposed by *Chawla* et al. (1981) are recommended. For liquid



stainless steel properties  $\rho, c_p, s, h = f(T)$  the approximations proposed by *Chawla et al. (1981)* and *Touloukian and Makita (1970)* are recommended.

For the solid/liquid two-phase region for *liquid metal*, the assumptions of thermodynamic equilibrium within the velocity field are used, and the properties then computed as explained in *Kolev (1991d)*.

The derivatives  $(\partial h / \partial p)_T, (\partial \rho / \partial T)_p, (\partial \rho / \partial p)_T$  are easily obtained by differentiating the corresponding analytical approximations.

The transport properties of the *simple* substances are computed as follows. Thermal conductivity and dynamic viscosity of air, and steam  $\lambda = \lambda(T)$  - after *Irvine and Lilly (1984)*. The water thermal conductivity  $\lambda = \lambda(T, p)$  is computed using the *Rivkin and Alexandrov (1975)* approximation and the water dynamic viscosity  $\eta = \eta(T, p)$  and surface tension  $\sigma = \sigma(T)$  using the TRAC approximation, *Liles et al. (1981)*.

The thermal conductivity and dynamic viscosity of the air-steam mixture are computed using the mole weight method of *Wilke (1950)*.

The thermal conductivity of solid and liquid uranium dioxide and stainless steel, as well as the dynamic viscosity and surface tension of liquid uranium dioxide and steel, are computed using the *Chawla et al. (1981)* approximations.

*Hill and Miyagawa* proposed in 1997 to use The IAPS Industrial Formulation 1997 for the Thermodynamic Properties of Water and Steam, see in *Wagner et al. (2000)*, in tabulated form. The use of an equidistant grid on the two independent axes is proposed and storing not only the properties but also their derivatives at a center of the computational cell. Marking this point with the integer coordinates  $(i, j)$  the interpolation is made by *Taylor* series expansion

$$f(x, y) = f_{i,j} + (x - x_i) \left( \frac{\partial f}{\partial x} \right)_{y,i,j} + \frac{1}{2} (x - x_i)^2 \left( \frac{\partial^2 f}{\partial x^2} \right)_{y,i,j} + (y - y_j) \left( \frac{\partial f}{\partial y} \right)_{x,i,j} + \frac{1}{2} (y - y_j)^2 \left( \frac{\partial^2 f}{\partial y^2} \right)_{x,i,j} + (x - x_i)(y - y_j) \left( \frac{\partial^2 f}{\partial x \partial y} \right)_{i,j}$$

In addition to first derivatives, which are usually available as analytical derivatives from the basic functions, the second derivatives have to be computed numerically, e.g. with steps of 0.001 K and 1000 Pa. Six constants per property and cell center have to be stored. The search of the integer address is quick because of the selected equidistant grid. The reversed functions are also easily obtained by solving analytically the above quadratic equation. Appropriate selected grids contain the critical point at the boundary of a cell (not at the center). Thus, for two such neighboring cells sub- and supercritical approximations are used, respectively.  $x_i$  and  $y_j$  may be slightly removed from the center if the saturation line crosses the cell. In this case the properties at the stable site of the saturation line are used. The authors also recommended that instead of the functions  $v, h, s = f(p, T)$ , the more appropriate  $p v, h, p s = f(p, T)$  be used, which give finite values for pressures tending to zero.

Reliable source of water and steam approximations is *Wagner and Kruse (1998)* providing also some inverted approximations for the sub-critical region. Some inverted approximations for water and steam are also available in *Meyer-Pittroff et al. (1969)*. General approaches for constructing properties for liquids and gases are given in *Reid et al. (1982)*. Large data base for caloric and transport properties for pure substances and mixtures is given in *Vargaftik et al. (1996)*. The NIST-JANAF thermo-chemical tables edited by *Chase (1998)* are inevitable for computational analysis of reactive flows.

Useful differential thermodynamic relationships are given in Appendix 3.3. The reader can find in *Elizer et al. (2002)* the relativistic modification to this relation.

### Appendix 3.3 Useful definitions and relations between thermodynamic quantities

		$T$	$s$	$p$	$v$
$\frac{\partial}{\partial p}$	$ _v$	$\frac{k}{\alpha}$	$\frac{kc_v}{\alpha T}$	1	0
	$ _T$	0	$-\alpha v$	1	$-kv$
	$ _s$	$\frac{\alpha T v}{c_p}$	0	1	$-\frac{c_v}{c_p} vk$
$\frac{\partial}{\partial v}$	$ _p$	$\frac{1}{\alpha v}$	$\frac{c_p}{\alpha T v}$	0	1
	$ _T$	0	$\frac{\alpha}{k}$	$-\frac{1}{k\alpha}$	1
	$ _s$	$-\frac{\alpha T}{kc_v}$	0	$-\frac{c_p}{c_v vk}$	1
$\frac{\partial}{\partial T}$	$ _p$	1	$\frac{c_p}{T}$	0	$\alpha v$
	$ _v$	1	$\frac{c_v}{T}$	$\frac{\alpha}{k}$	0
	$ _s$	1	0	$\frac{c_p}{\alpha T v}$	$\frac{c_v k}{\alpha T}$
$\frac{\partial}{\partial s}$	$ _p$	$\frac{T}{c_p}$	1	0	$\frac{\alpha T v}{c_p}$
	$ _v$	$\frac{T}{c_v}$	1	$\frac{\alpha T}{kc_v}$	0
	$ _T$	0	1	$-\frac{1}{\alpha v}$	$\frac{k}{\alpha}$

Specific capacity at constant volume	$c_v = \left( \frac{\partial u}{\partial T} \right)_v$
Specific capacity at constant pressure	$c_p = \left( \frac{\partial u}{\partial T} \right)_p + p \left( \frac{\partial v}{\partial T} \right)_p$
Thermal expansion coefficient	$\alpha = \frac{1}{v} \left( \frac{\partial v}{\partial T} \right)_p$
Compressibility	$k = -\frac{1}{v} \left( \frac{\partial v}{\partial p} \right)_T$
Coefficient of thermal strain	$\sigma = \frac{1}{p} \left( \frac{\partial p}{\partial T} \right)_v$

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# 4 On the variety of notations of the energy conservation for single-phase flow

The Greek word *τροπή* stems from the verb *εντροπείη* – conversion, transformation, and was used by *R. Clausius* in a sense of “value representing the sum of all transformations necessary to bring each body or system of bodies to their present state” (1854 *Ann. Phys.*, 125, p. 390; 1868 *Phil. Mag.*, 35, p. 419; 1875 “*Die mechanische Wärmetheorie*”, vol. 1).

*This chapter recalls the achievements of the classical thermodynamics for describing the thermodynamic behavior of flows. At the end of the 19th century this knowledge had already formed the classical technical thermodynamics.*

*The classical thermodynamics makes use of different mathematical notation of the first principles. The purpose of this chapter is to remember that all mathematically correct transformations from one notation into the other are simply logical and consistent reflections of the same physics. The use of the specific entropy is not an exception. The importance of the entropy is that it allows one to reflect nature in the simplest mathematical way.*

*Even being known since 100 years the basic system of partial differential equations is still not analytically solved for the general case. The requirements to solve this system numerically and the virtual possibility to do this by using modern computers is the charming characteristic of the science today. The obtained numerical results have always to be critically examined. As we demonstrate here for pressure wave analysis lower order numerical methods predict acceptably only the first few cycles but then the solution degrades destroyed by numerical diffusion. The inability of the lower order numerical methods to solve fluid mechanics equations for initial and boundary conditions defining strong gradients is not evidence that the equations are wrong. That high accuracy solution methods are necessary to satisfy future needs of the industry is beyond question.*

## 4.1 Introduction

Chapter 4 is intended to serve as an introduction to Chapter 5.

The computational fluid mechanics produces a large number of publications in which the mathematical notation of the basic principles and of the thermodynamic relationships is often taken by the authors for self understanding and is rarely

explained in detail. One of the prominent examples is the different notation of the energy conservation for flows in the literature. The vector of the dependent variables frequently used may contain either specific energy, or specific enthalpy, or specific entropy, or temperature etc. among other variables. If doing the transformation from one vector into the other correctly the obtained systems of partial differential equation must be completely equivalent to each other. Nevertheless, sometimes the message of the different notations of the first principle is misunderstood by interpreting one of them as “wrong” and other as right. That is why recalling the basics once more seems to be of practical use in order to avoid misunderstandings.

## 4.2 Mass and momentum conservation, energy conservation

The fluid mechanics emerges as a scientific discipline by first introducing the idea of *control volume*. The flowing continuum can be described mathematically by abstracting a control volume inside the flow and describing what happens in this control volume. If the control volume is stationary to the frame of reference the resulting description of the flow is frequently called *Eulerian* description. If the control volume is stationary to the coordinate system following the trajectory of fluid particles the description is called *Lagrangian*. *Oswatitsch's* “Gasdynamik” (1952) is my favorite recommendation to start with this topic because of its uncomplicated introduction into the “language of the gas dynamics”. Next we write a system of quasi linear partial differential equations describing the behavior of a single-phase flow in an *Eulerian* control volume without derivation.

The idea of conservation of matter was already expressed in ancient philosophy. Much later, in 1748, *M. Lomonosov* wrote in a letter to *L. Euler* “... all changes in the nature happens so that the mass lost by one body is added to other ...”. Equation (4.1) reflects the principle of conservation of mass per unit volume of the flow probably first mathematically expressed in its present form by *D’Alembert* (1743). The second equation is nothing else than the first principle of *Newton* (1872) applied on a continuum passing through the control volume probably for the first time by *Euler* (1737-1755), see *Euler* (1757). Again the equation is written per unit flow volume. The third equation reflects the principle of the conservation of energy, *Mayer* (1841-1851), see *Mayer* (1757), a medical doctor finding that the humans loose more energy in colder regions in the globe then in hotter, which is reflected in the color of the blood due to the different degree of blood oxidation.

$$\frac{\partial \rho}{\partial \tau} + \nabla \cdot (\rho \mathbf{V}) = 0 \quad (4.1)$$

$$\rho \left( \frac{\partial \mathbf{V}}{\partial \tau} + \mathbf{V} \nabla \cdot \mathbf{V} \right) + \nabla \cdot p + \rho \mathbf{g} = 0 \quad (4.2)$$

$$\frac{\partial}{\partial \tau} \left[ \rho \left( e + \frac{1}{2} V^2 \right) \right] + \nabla \cdot \left[ \rho \left( e + p / \rho + \frac{1}{2} V^2 \right) \mathbf{V} \right] + \rho \mathbf{g} \cdot \mathbf{V} = 0 \quad (4.3)$$

Here the usual notations of the thermodynamics are used.

Note that these equations are

- a) local volume averaged which means that the behavior of a single molecule interacting with its neighbors is not described in detail.

Secondly, they are

- b) not time-averaged conservation equation for
- c) flow without internal or external heat sources.

That this equation does not reflect all of the complexity of the flows in nature is evident from the fact that

- d) the forces resisting the flow due to viscosity are neglected, and consequently
- e) the power component of these forces into the energy conservation equation is also neglected.

Nevertheless the system (4.1) through (4.3) reflects several important characteristics observed in real flows and is frequently used in science and practical computations. In other words this system describes well wave dynamics but does not describe any diffusion processes in the flow, like viscous energy dissipation, heat conduction etc.

### 4.3 Simple notation of the energy conservation equation

The energy conservation equation can be substantially simplified as it will be demonstrated here without any loss of generality. To achieve this first the scalar product of Eq. (4.2) is constructed with the field velocity  $\mathbf{V}$ . The result is a scalar expressing the mechanical energy balance. Subtracting this result from the energy equation results in

$$\frac{\partial}{\partial \tau} (\rho e) + \nabla \cdot (\rho e \mathbf{V}) + p \nabla \cdot \mathbf{V} = 0. \quad (4.4)$$

Obviously, this equation is much more compact than the primitive form of the energy conservation equation (4.3). Additional simplification is reached if differentiating the first two terms and comparing with the mass conservation equation. Dividing by the density results in

$$\frac{\partial e}{\partial \tau} + (\mathbf{V} \cdot \nabla) e + \frac{p}{\rho} \nabla \cdot \mathbf{V} = 0. \quad (4.4)$$



Note, that the system consisting of Eqs. (4.1, 4.2, 4.3) is completely equivalent to the system consisting of Eqs. (4.1, 4.2, 4.4 or 4.4a).

## 4.4 The entropy

Equation (4.4) is called frequently in the literature also energy equation. It can be additionally simplified by introducing a new variable. In a closed system the introduced external heat,  $dq$ , may be transformed into internal specific energy  $de$ , or into expansion work  $p dv$ , or into both simultaneously, in a way that no energy is lost  $dq = de + p dv$  – energy conservation principle. A unique functional dependence of the type  $q = q(e, v)$  cannot be derived. In the mathematics the linear form  $dq = de + p dv$  in which  $dq$  is not a total differential is called *Pfaff's form*. The *Gibbs relation*,

$$T ds = de + p dv = de - \frac{p}{\rho^2} d\rho, \quad (4.5)$$

*Gibbs* (1878), simply says that per unit mass the change of the internal energy  $de$  and the expansion work  $p dv$  in the control volume can be expressed as a change of an single variable denoted with  $s$ , provided  $T$  is the so-called integrating denominator. That  $T$  is the integrating denominator was proven 30 years earlier. The integrating denominator itself was found to be the absolute temperature  $T$  (*W. Thomson* 1848). *M. Plank* showed that among the many choices for the integrating denominator only the absolute temperature allows to integrate independently from the trajectory of the transformation *Plank* (1964). The new remarkable variable was called by *R. Clausius* specific *entropy*, see *Clausius* (1854). Specific – because related to unit mass of the fluid. The Greek word for entropy, *τροπή*, stems from the verb *εντροπειη* which means conversion or transformation, and is used by *Clausius* in a sense of “...value representing the sum of all transformations necessary to bring each body or system of bodies to their present state”. The introduction of the specific entropy allows simple notation of the second principle of the thermodynamics formulated also by *Clausius*. *Gibbs* used Eq. (4.5) as a mathematical expression of the so called equilibrium principle “... after the entropy of the system reaches the maximum the system is in a state of equilibrium ...”. Equation (4.5) has another remarkable property. It is the differential expression of a unique equation of state  $s = s(e, v)$  in which the partial derivatives  $(\partial e / \partial s)_v = T$  and  $(\partial e / \partial v)_s = -p$  are technically measurable properties.

Note that the introduction of the new variable entropy plays also an important role in the statistical mechanics and in the informatics which will not be discussed here.

Thus, making use of the *Gibbs* relation, of the mass conservation equation, and of Eq. (4.5), Eq. (4.4a) is easily transformed into the simplest notation expressing the energy conservation law

$$\frac{\partial s}{\partial \tau} + (\mathbf{V} \cdot \nabla) \cdot s = 0. \quad (4.6)$$

Again note that the system consisting of Eqs. (4.1, 4.2, 4.6) is completely equivalent to the system consisting of Eqs. (4.1, 4.2, 4.4), and consequently to the original system containing Eqs. (4.1, 4.2, 4.3).

## 4.5 Equation of state

The macroscopic behavior of fluids can be described by describing the behavior of the simple molecule in interaction with others in a control volume *Boltzman* (1909). The technical thermodynamics makes use of the so called equation of state in which the local volume- and time-averaged behavior of the fluid is reflected rather than describing the behavior of the single molecule (see the *Riemann*'s work from 1859 in which he combines the experimental findings by *Gay-Lussac* (1806), *Boyle-Mariotte* to one equation of state, or any thermodynamic text book). One example for the equation of state is the dependence of the specific internal energy on the volume-averaged density and on the volume-averaged temperatures, the so called caloric equation of state,

$$e = e(\rho, T). \quad (4.7)$$

Note the useful definition of the specific capacity at constant volume

$$\left( \frac{\partial e}{\partial T} \right)_{\rho} = \left( \frac{\partial e}{\partial T} \right)_{v} = c_v \quad (4.8)$$

Perfect fluids are defined by the dependence of the specific internal energy on the temperature only, but not on the density (remember the famous expansion experiments by *Gay-Lussac*). Variety of forms of the equation of state is possible which may lead to a variety of notations of the energy conservation principle.

## 4.6 Variety of notation of the energy conservation principle

### 4.6.1 Temperature

Equation (4.4) can be rewritten in terms of temperature using the equation of state (4.7) and the mass conservation equation (4.1). The result is

$$\frac{\partial T}{\partial \tau} + (\mathbf{V} \cdot \nabla) T + \frac{p}{\rho c_v} \left[ 1 - \frac{\rho^2}{p} \left( \frac{\partial e}{\partial \rho} \right)_T \right] \nabla \cdot \mathbf{V} = 0, \quad (4.9)$$

an equation containing only one time derivative. Note that

$$\frac{p}{T} \left[ 1 - \frac{\rho^2}{p} \left( \frac{\partial e}{\partial \rho} \right)_T \right] \equiv \left( \frac{\partial p}{\partial T} \right)_v, \quad (4.10)$$

which is an important thermodynamic relationship, see *Reynolds and Perkins* (1977) p. 266 Eq. 8.54. Again note that the system consisting of Eqs. (4.1, 4.2, 4.9) is completely equivalent to the system consisting of Eqs. (4.1, 4.2, 4.4a), and consequently to the original system containing Eqs. (4.1, 4.2, 4.3).

#### 4.6.2 Specific enthalpy

Another specific variable frequently used in the thermodynamic is the specific enthalpy defined as follows

$$h = e + p / \rho. \quad (4.11)$$

Using the differential form of this definition

$$d(\rho e) + dp = d(\rho h) \quad (4.12)$$

Eq. (4.4) can be written in terms of the specific enthalpy and pressure

$$\frac{\partial}{\partial \tau}(\rho h) + \nabla \cdot (\rho h \mathbf{V}) - \left( \frac{\partial p}{\partial \tau} + \mathbf{V} \cdot \nabla p \right) = 0. \quad (4.13)$$

Again note that the system consisting of Eqs. (4.1, 4.2, 4.13) is completely equivalent to the system consisting of Eqs. (4.1, 4.2, 4.4), and consequently to the original system containing Eqs. (4.1, 4.2, 4.3). Moreover, it is easily shown that using the definition equation for the specific enthalpy (4.11) the *Gibbs* relation (4.5) takes the form

$$T \rho ds = \rho dh - dp \quad (4.14)$$

which immediately transforms Eq. (4.13) into the entropy equation (4.6) again. Equation (4.13) is easily transferred in terms of temperature and pressure by using the following equation of state

$$h = h(T, p) \quad (4.15)$$

or in differential form

$$dh = c_p dT + \left( \frac{\partial h}{\partial p} \right)_T dp. \quad (4.16)$$

The resulting equation is

$$\frac{\partial T}{\partial \tau} + (\mathbf{V} \cdot \nabla) T - \frac{1}{\rho c_p} \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_T \right] \left[ \frac{\partial p}{\partial \tau} + (\mathbf{V} \cdot \nabla) p \right] = 0. \quad (4.17)$$

Again note that the system consisting of Eqs. (4.1, 4.2, 4.17) is completely equivalent to the system consisting of Eqs. (4.1, 4.2, 4.13), and consequently to the original system containing Eqs. (4.1, 4.2, 4.3).

## 4.7 Summary of different notations

It follows from Sections 4.3 through 4.6 that the system of equations (4.1, 4.2, 4.3) is completely equivalent to one of the systems containing equations (4.1) and (4.2), and one of the following expressions of the first principle

$$\frac{\partial e}{\partial \tau} + (\mathbf{V} \cdot \nabla) e + \frac{p}{\rho} \nabla \cdot \mathbf{V} = 0, \quad (4.4a)$$

$$\frac{\partial s}{\partial \tau} + (\mathbf{V} \cdot \nabla) s = 0, \quad (4.6)$$

$$\frac{\partial T}{\partial \tau} + (\mathbf{V} \cdot \nabla) T + \frac{p}{\rho c_v} \left[ 1 - \frac{\rho^2}{p} \left( \frac{\partial e}{\partial \rho} \right)_T \right] \nabla \cdot \mathbf{V} = 0, \quad (4.9)$$

$$\frac{\partial h}{\partial \tau} + (\mathbf{V} \cdot \nabla) h - \frac{1}{\rho} \left[ \frac{\partial p}{\partial \tau} + (\mathbf{V} \cdot \nabla) p \right] = 0, \quad (4.13)$$

$$\frac{\partial T}{\partial \tau} + (\mathbf{V} \cdot \nabla) T - \frac{1}{\rho c_p} \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_T \right] \left[ \frac{\partial p}{\partial \tau} + (\mathbf{V} \cdot \nabla) p \right] = 0. \quad (4.17)$$

These equations are derived without specifying the type of the fluid as perfect or real gas. They are valid for both cases. The reader will find in *McDonald et al.* (1978) notation using as dependent variables:  $(w, h, p)$ ,  $(G, h, p)$ ,  $(w, \rho, p)$ ,  $(G, \rho, p)$ ,  $(w, \rho, h)$ ,  $(G, \rho, h)$ ,  $(w, \rho, e)$ ,  $(G, \rho, e)$  and in *Thorley and Tiley* (1987)  $(\rho, h, w)$  and  $(p, T, w)$ .

## 4.8 The equivalence of the canonical forms

The system of partial differential equations can be further considerably simplified if they are from a special type called hyperbolic. *Riemann* found that such a system can be transferred to the so called canonical form which consists of ordinary differential equations. Next we will start with two different forms of the system, one using specific entropy and the other using specific internal energy. We then give the eigen values, the eigen vectors and the canonical form of both systems. It will be shown again that the so obtained forms are mathematically identical.

Before performing the analysis of the type of the system we simplify the mass conservation equation as follows. The density in the mass conservation equation is expressed as function of pressure and specific entropy

$$\rho = \rho(p, s) \quad (4.18)$$

or in differential form

$$d\rho = \frac{d\rho}{dp} + \left( \frac{\partial \rho}{\partial s} \right)_p ds. \quad (4.19)$$

Substituting Eq. (4.19) into Eq. (4.1) and taking into account the entropy Eq. (4.6) the mass conservation equation takes the form of Eq. (4.21). Thus using the following vector of dependent variables

$$U = (p, w, s) \quad (4.20)$$

the flow is completely described by the system of quasi linear partial differential equations

$$\frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} + \rho a^2 \frac{\partial w}{\partial z} = 0, \quad (4.21)$$

$$\frac{\partial w}{\partial \tau} + w \frac{\partial w}{\partial z} + \frac{1}{\rho} \frac{\partial p}{\partial z} + g_z = 0, \quad (4.2)$$

$$\frac{\partial s}{\partial \tau} + w \frac{\partial s}{\partial z} = 0. \quad (4.6)$$

It is well known that the eigen values of this system defined by

$$\begin{vmatrix} w - \lambda & \rho a^2 & 0 \\ 1/\rho & w - \lambda & 0 \\ 0 & 0 & w - \lambda \end{vmatrix} = 0, \quad (4.22)$$

are real and different from each other

$$\lambda_{1,2} = w \pm a, \quad \lambda_3 = w. \quad (4.23, 4.24, 4.25)$$

The eigen vectors of the transposed characteristic matrix

$$\begin{pmatrix} w - \lambda_i & 1/\rho & 0 \\ \rho a^2 & w - \lambda_i & 0 \\ 0 & 0 & w - \lambda_i \end{pmatrix} \begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = 0, \quad (4.26)$$

are linearly independent

$$h_1 = \begin{pmatrix} 1/(\rho a) \\ 1 \\ 0 \end{pmatrix}, h_2 = \begin{pmatrix} -1/(\rho a) \\ 1 \\ 0 \end{pmatrix}, \text{ and } h_3 = \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}. \quad (4.27, 4.28, 4.29)$$

Therefore the quasi linear system of partial differential equations is hyperbolic and can be transformed into the very simple and useful canonical form

$$\frac{dz}{d\tau} = w + a, \quad \frac{1}{\rho a} \frac{dp}{d\tau} + \frac{dw}{d\tau} = -g, \quad (4.30, 4.31)$$

$$\frac{dz}{d\tau} = w - a, \quad -\frac{1}{\rho a} \frac{dp}{d\tau} + \frac{dw}{d\tau} = -g, \quad (4.32, 4.33)$$

$$\frac{dz}{d\tau} = w, \quad \frac{ds}{d\tau} = 0. \quad (4.34, 4.35)$$

As far as I know Eqs. (4.30 to 4.35) were obtained for the first time by *Riemann* in 1859. Remember that each of the ordinary differential equations (4.31, 4.33, 4.35) is valid in its own coordinate system attached along one of the three characteristic lines defined by the equations (4.30, 4.32, 4.34), respectively. Equations (4.34, 4.35) contain a very interesting analytical solution  $s = \text{const}$ . It means that an infinitesimal control volume moving with the fluid velocity along the line defined by Eq. (4.34) in the  $(\tau, z)$  plane does not change a specific property defined as specific entropy.

This change of state is called isentropic. Comparing with Eq. (4.5) it simply means that within this control volume the energy required for volume expansion per unit flow volume reduces the specific internal energy per unit time by the same amount.

Now let us perform the same analysis starting from the system containing the first principle in terms of specific internal energy. We use this time as dependent variables vector,

$$U = (p, w, e), \quad (4.36)$$

and the system describing the flow

$$\frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} + \rho a^2 \frac{\partial w}{\partial z} = 0 \quad (4.21)$$

$$\frac{\partial w}{\partial \tau} + w \frac{\partial w}{\partial z} + \frac{1}{\rho} \frac{\partial p}{\partial z} + g_z = 0 \quad (4.2)$$

$$\frac{\partial e}{\partial \tau} + w \frac{\partial e}{\partial z} + \frac{p}{\rho} \frac{\partial w}{\partial z} = 0. \quad (4.4a)$$

If you are not familiar with the analysis of the type of a system of partial differential equations by first computing the eigen values, eigenvectors and canonical forms it is recommended first to read Section 11 before continuing here.

From the characteristic equation

$$\begin{vmatrix} w - \lambda & \rho a^2 & 0 \\ 1/\rho & w - \lambda & 0 \\ 0 & \frac{p}{\rho} & w - \lambda \end{vmatrix} = 0, \quad (4.37)$$

we obtain the same eigen values, Eqs. (4.23, 4.24, 4.25), as in the previous case, which is not surprising for equivalent systems. Using the eigen vectors of the transposed characteristic matrix

$$\begin{pmatrix} w - \lambda & 1/\rho & 0 \\ \rho a^2 & w - \lambda & \frac{p}{\rho} \\ 0 & 0 & w - \lambda \end{pmatrix} \begin{pmatrix} h_1 \\ h_2 \\ h_3 \end{pmatrix} = 0, \quad (4.38)$$

we can transform the system of partial differential equations again into a system of ordinary equations along the characteristic lines. Along the first two characteristic lines both canonical equations are identical with the system using the specific entropy as a dependent variable. Along the third characteristic line we have a canonical equation which is in terms of specific internal energy and pressure,

$$\frac{dz}{d\tau} = w, \quad \frac{de}{d\tau} - \left( \frac{p}{\rho a} \right)^2 \frac{1}{p} \frac{dp}{d\tau} = 0. \quad (4.39, 4.40)$$

For the general case it can be shown that the following relation holds

$$\left[ T + \frac{p}{\rho^2} \left( \frac{\partial \rho}{\partial s} \right)_p \right] ds = de - \left( \frac{p}{\rho a} \right)^2 \frac{1}{p} dp. \quad (4.41)$$

Applying this relation to the canonical equation (4.40) results exactly in Eq. (4.35). This demonstrates that also the canonical forms of the system using different sets as depending variables are mathematically equivalent.

## 4.9 Equivalence of the analytical solutions

That the analytical solutions of the equivalent systems of equations written in terms of different dependent variables are equivalent to each other is obvious.

## 4.10 Equivalence of the numerical solutions?

### 4.10.1 Explicit first order method of characteristics

Next let us apply the simple explicit first order method of characteristics for solving the system for a horizontal equidistant discretized pipe,

$$\Delta z_1 = (w + a) \Delta \tau, \quad \frac{1}{\rho a} (p^{\tau+\Delta\tau} - p_1) + w^{\tau+\Delta\tau} - w_1 = 0, \quad (4.42, 4.43)$$

$$\Delta z_2 = (w - a) \Delta \tau, \quad -\frac{1}{\rho a} (p^{\tau+\Delta\tau} - p_2) + w^{\tau+\Delta\tau} - w_2 = 0, \quad (4.44, 4.45)$$

$$\Delta z_3 = w \Delta \tau, \quad s^{\tau+\Delta\tau} - s_3 = 0, \quad (4.46, 4.47)$$

where indices  $i$  refers to values of the dependent variables at  $(\tau, z - \lambda_i \Delta \tau)$ . There are a variety of methods to compute  $U_i$  as a function of the neighboring values. One of the simplest is the linear interpolation

$$U_i = \frac{1}{2} [1 + \text{sign}(Cu_i)] [U_{k-1} + (U - U_{k-1})(1 - Cu_i)] \\ + \frac{1}{2} [1 - \text{sign}(Cu_i)] [U - (U_{k+1} - U)Cu_i], \quad (4.48)$$

where

$$Cu_i = \frac{\lambda_i \Delta \tau}{\Delta z}. \quad (4.49)$$

The result is

$$w^{\tau+\Delta\tau} = \frac{1}{2} \left[ -\frac{1}{\rho a} (p_2 - p_1) + w_1 + w_2 \right] \quad (4.50)$$

$$p^{\tau+\Delta\tau} = \frac{1}{2} [p_2 + p_1 - (w_2 - w_1) \rho a] \quad (4.51)$$

$$s^{\tau+\Delta\tau} = s_3. \quad (4.52)$$



Obviously, the pressure and the velocity in the point  $(\tau + \Delta\tau, z)$  depend on the pressure and the velocities at points 1 and 2. The old time value of the specific entropy is simply shifted along the third characteristic from the point  $(\tau, z - w\Delta\tau)$  to the point  $(\tau + \Delta\tau, z)$ . This result is derived without any assumption of the type of the fluid.

In case of closed ends of the pipe we have the following results. For the left end  $w_L = 0$ ,  $s_3 = s_L$ , and therefore  $s^{\tau+\Delta\tau} = s_L$ , and from Eq. (4.45)  $p^{\tau+\Delta\tau} = p_2 - w_2\rho a$ . Similarly for the right end  $w_R = 0$ ,  $s_3 = s_R$ , and therefore  $s^{\tau+\Delta\tau} = s_R$ , and from Eq. (4.43)  $p^{\tau+\Delta\tau} = p_1 + w_1\rho a$ . For open ends only the canonical equations whose characteristics are inside the pipe are used. For the missing equations boundary conditions have to be provided.

*Joukowski* (1898) derived the closed end relation already in 1898. He explained for the first time the physics behind the maximum pressure spikes for water pipe lines in which valves are closed quickly.

Therefore, fast closing valves produce stagnation pressure spikes that are proportional (a) to the undisturbed initial velocity and (b) to the sound velocity,

$$p^{\tau+\Delta\tau} - p_1 = \rho w_1 a, \text{ Eq. (16), } \textit{Joukowski} \text{ (1898).}$$

This is important knowledge for designing the pipe supports in the engineering practice. Either the designer of the valves have to implement solution not allowing fast closing, or if fast closing is desired for some reasons, the pipes have to sustain the forces originating after the closing. Farther more *Joukowski* uses the *Riemann's* method to transfer the mass and momentum equations into canonical forms and proposed for the first time numerical solution method which is known now as the "method of characteristics". Not having computers, *Joukowski* proposed graphical method of characteristics. In his important for the modern hydrodynamics work, *Joukowski* also considered the influence of the elasticity of the pipe walls on the velocity of sound, Eq. (15) in *Joukowski* (1898).

The reader will find examples for numerical solutions based on the single-phase method of characteristics as follows:

**1D-flows:** For diabatic three-equations model see *Namatame* and *Kobayashi* (1974), for two-equations model for pipe-networks see *Katkovskiy* and *Poletaev* (1975), *Henshaw* (1987), *Weisman* and *Tenter* (1981), for two-equations model for pipe-networks with pumps see *Capozza* (1986), *Chen* et al. (1974), *Shin* and *Chen* (1975).

**2D-flows:** For single component flow see *George and Doshi (1994)*, *Henshaw (1987)*, *Weisman and Tenter (1981)*, for multi-component steady reactive flows see *Zucrow and Hoffman (1977)*, for isentropic flow see *Shin and Valentin (1974)*.

**1D homogeneous two-phase flows:** The extension of the single phase flow models to two-phase equilibrium models is very easy. The mixture specific entropy is defined as follows

$$s = X_{1,eq} s'' + (1 - X_{1,eq}) s',$$

resulting in equilibrium vapor mass flow fraction

$$X_{1,eq} = \frac{s - s'}{s'' - s'}.$$

The specific mixture volume is then

$$\begin{aligned} v &= X_{1,eq} v'' + (1 - X_{1,eq}) v' \\ &= s \frac{v'' - v'}{s'' - s'} + \frac{s'' v' - s' v''}{s'' - s'} = s \frac{dT'}{dp} + \frac{s'' v' - s' v''}{s'' - s'} = f(s, p) = sf_1 + f_2 = v(p, s). \end{aligned}$$

Therefore a convenient equation of state is  $v = v(p, s)$  with differential form

$$dv = \left( \frac{\partial v}{\partial p} \right)_s dp + \left( \frac{\partial v}{\partial s} \right)_p ds,$$

where

$$\left( \frac{\partial v}{\partial p} \right)_s = s \frac{df_1}{dp} + \frac{df_2}{dp},$$

$$\left( \frac{\partial v}{\partial s} \right)_p = f_1.$$

We realize that the partial derivative of the mixture specific volume with respect to the specific mixture entropy is equivalent to the *Clapeyron's* expression

$$\frac{dT'}{dp} = \frac{v'' - v'}{s'' - s'}.$$

Approximating the function  $f_1$  and  $f_2$  with smooth pressure functions is very convenient for using them in fast computational algorithms. The velocity of sound  $a$  is easily derived if we use the above expressions to obtain the equation of state  $\rho = 1/v = \rho(p, s)$  or in differential form

$$d\rho = \frac{dp}{a^2} + \left( \frac{\partial \rho}{\partial s} \right)_p ds,$$

where

$$\frac{1}{a^2} = \left( \frac{\partial \rho}{\partial p} \right)_s = -\frac{1}{v^2} \left( s \frac{df_1}{dp} + \frac{df_2}{dp} \right)$$

$$\left( \frac{\partial \rho}{\partial s} \right)_p = -f_1/v^2.$$

Alternative expression for the velocity of sound,

$$\frac{1}{(\rho a)^2} = \frac{v'' - v'}{s'' - s'} \left[ X_{1,eq} \frac{ds''}{dp} + (1 - X_{1,eq}) \frac{ds'}{dp} \right] - X_{1,eq} \frac{dv''}{dp} + (1 - X_{1,eq}) \frac{dv'}{dp},$$

is frequently used in the literature but it is obviously computationally more expensive compared to the previous expression.

Homogeneous non equilibrium 4-equation model is solved by the method of characteristics on structured grid using second order method by *Köberlein* (1972), *Grillenberger* (1981). Homogeneous equilibrium 3-equations models with variety of notations for the energy conservation are reported by *McDonald* et al. (1978). Different variants of the methods of characteristics are proposed in this work with the most interesting one based on two grids. On the even grids a conservative form of the conservation equations is used guaranteeing strict conservation, and on the odd grids – the canonical form using the method of characteristics. Homogeneous equilibrium 3-equations and homogeneous non-equilibrium 5-equations model are solved by the method of characteristics on non structured grid by *Hancox* and *Banerjee* (1977), *Hancox* et al. (1978). Homogeneous non-equilibrium 5-equations model is solved by the method of characteristics on non structured grid by *Ferch* (1979); For two fluid model without thermal interaction see *Weisman* and *Tenter* (1981).

**1D non-homogeneous two-phase flows:** For isentropic drift flux model see *Kaizerman* et al. (1983). Analytical solution for concentration waves in boiling channel at constant pressure and known mass flow is provided by *Santalo* and *Lahey* (1972).

If the spatial derivatives in the canonical form are computed by finite differences for the old time plane by using weighing depending on the direction of the velocity the method is called method of lines or pseudo-characteristics method. Examples are available by *Wang* and *Johnson* (1981), *Kolev* and *Carver* (1981). Positioning adaptively more grid points along the pipe axis where the gradients of the variables are stronger is a efficient technique to increase the accuracy of the method of lines by low computational costs *Hu* and *Schiesser* (1981).

To get an impression about the physical meaning of Eq. (4.52) we rewrite it for a perfect gas

$$\frac{T^{\tau+\Delta\tau}}{T_3} = \left( \frac{p^{\tau+\Delta\tau}}{p_3} \right)^{R/c_p}. \quad (4.53)$$

This is usually used as a benchmark for any other numerical solution using first order explicit methods.

Applying again the explicit method of characteristics to the system in terms of the specific internal energy we obtain the same results for the pressure and velocities. The relation along the third characteristic line should be considered very carefully. If the differential form

$$de = \left( \frac{p}{\rho a} \right)^2 \frac{dp}{p} \quad (4.54)$$

is analytically integrated for a perfect gas it results exactly in Eq. (4.53). This is not the case if the differential form is written in finite difference form,

$$e^{\tau+\Delta\tau} = e_3 + \frac{p}{(\rho a)^2} (p^{\tau+\Delta\tau} - p_3), \quad (4.55)$$

because for the computation of the coefficient  $\frac{p}{(\rho a)^2}$  an assumption has to be made of how it has to be computed. For this reason the result,

$$T^{\tau+\Delta\tau} - T_3 = \frac{T}{p} \frac{R}{c_p} (p^{\tau+\Delta\tau} - p_3), \quad (4.56)$$

does not fit to the exact benchmark solution. Over many time steps it may accumulate to considerable error.

#### 4.10.2 The perfect gas shock tube: benchmark for numerical methods

What is a perfect gas? The experiment that demonstrates the properties of a fluid classified as a perfect fluid was performed in 1806 by *Gay-Lussac* and repeated with a more sophisticated apparatus in 1845 by *Joule* see *Gay-Lussac* (1807) and *Joule* (1884). In this remarkable experiment a closed space with rigid and thermally insulated walls is separated by a membrane into two parts. The left part contained gas with temperature  $T$  and pressure  $p$ . The right part was evacuated. The separating membrane was broken and the gas expands decreasing first temperature and pressure in the left side and increasing temperature and pressure in the right side. After enough time something remarkable happens. The temperature everywhere approaches the initial temperature. For the same mass and internal

energy, that is for the same specific internal energy, the temperature remains constant in spite of the change of the specific volume. Mathematically this means

$$\left(\frac{\partial e}{\partial v}\right)_\tau = 0, \quad (4.57)$$

and therefore

$$e = e(T). \quad (4.58)$$

This reflects the condition which has to be satisfied by the caloric equation of state of the fluid to classify it as a perfect fluid. Real fluids do not satisfy this condition *Elsner* (1974). That is why after such experiments the temperature will not be the same as the initial temperature even if the experimental apparatus could provide perfect energy conservation inside the control volume.

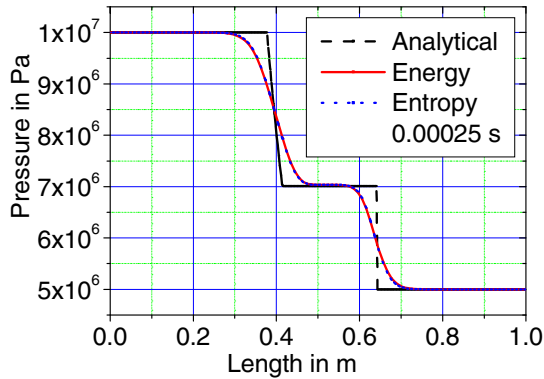
The modern expression for this apparatus is a shock tube. Shock tubes are important experimental facilities for studying dynamic processes in fluids. In the numerical fluid dynamics they serve as a benchmark problem on which the accuracy of the numerical methods can be examined. Text book analytical solution for a perfect gas is available; see for instance *Hoffmann* and *Chang* (2000). Procedure to its use is given in Appendix A4.1 and comparison of this solution with some numerical solutions is given in Appendix A4.2.

Remember that there are two important error sources in the numerical solution: discretization errors and truncation errors. The discretization errors are associated with the approximations used to represent derivatives at given points. Usually series expansion is used and the terms not taken into account are computed to estimate the error. To make the error plausible it can be expressed in changes of measurable variables. The truncation errors themselves are associated with inability of the digital computers to work with as much digits as necessary. In the course of the computation errors generated from both sources accumulate. *Under given circumstances the accumulated error can be so big that the obtained solutions are no longer consistent with the initial conditions.*

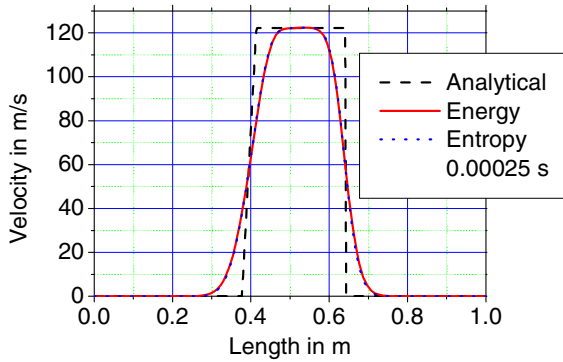
Let us perform such a test in order to quantify the error of the numerical solution of the system of partial equations. Consider a pipe with constant cross section  $A$  and length  $L$ . Discretize it equidistantly by using  $k_{\max}$  points.  $k_{\max}$  is an even number. Assign for time  $\tau$  initial values for  $T$  and  $p$  as follows:  $T = T^*$ ,  $w = 0$  for  $k = 1, k_{\max}$ ,  $p = p_L$  for  $k = 1, k_{\max}/2$  and  $p = p_R$  for  $k = k_{\max}/2, k_{\max}$ . Having the temperature and pressure in each point all other variables required for the analysis can be computed as follows  $\rho = p/(RT)$ ,  $a = \sqrt{\kappa RT}$ , where  $\kappa = c_p/(c_p - R)$ ,  $(\partial\rho/\partial s)_p = -p/(RTc_p)$ ,  $s = c_p \ln(T/T_0) - R \ln(p/p_0)$ ,  $e = c_v(T - T_0)$ ,  $h = c_p(T - T_0)$ . Select the time step  $\Delta\tau$  satisfying the criterion

$$\max(Cu_i) < 1. \quad (4.59)$$

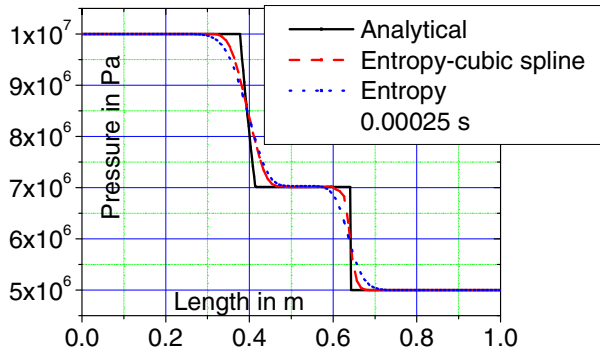
Compute all dependent variables in the new time plane  $\tau + \Delta\tau$  taking into account that the velocities at the both ends are zero. If the entropy or the specific internal energy is one of the dependent variables compute the temperature as either using  $T = T_0 (p/p_0)^{R/c_p} \exp(s/c_p)$  or using  $T = T_0 + e/c_v$ , respectively. Then all properties can be computed as function of pressure and temperatures. Compare the differences between the initial total energy of the system at the beginning of the process to the actual energy and estimate the error made. Advance the computation up to obtaining the steady state solution and compare the temperature with the initial temperature.



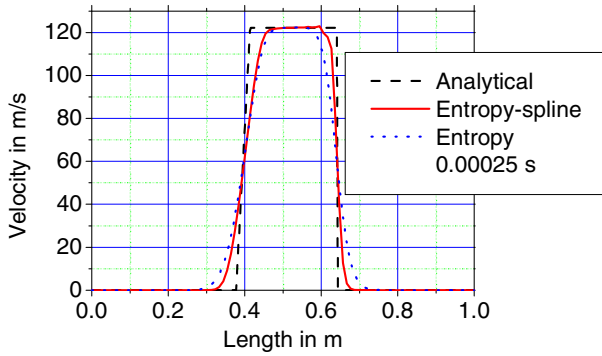
**Fig. 4.1** Pressure as function of the distance from the left closed end of the pipe at  $\tau = 250\mu s$ . Comparison between the analytical and numerical solutions. The systems of PDE's contain the energy conservation principle in terms of specific internal energy or specific entropy. Both numerical solutions are indistinguishable



**Fig. 4.2** Velocity as function of the distance from the left closed end of the pipe at  $\tau = 250\mu s$ . Comparison between the analytical and numerical solutions. The systems of PDE's contain the energy conservation principle in terms of specific internal energy or specific entropy. Both numerical solutions are indistinguishable

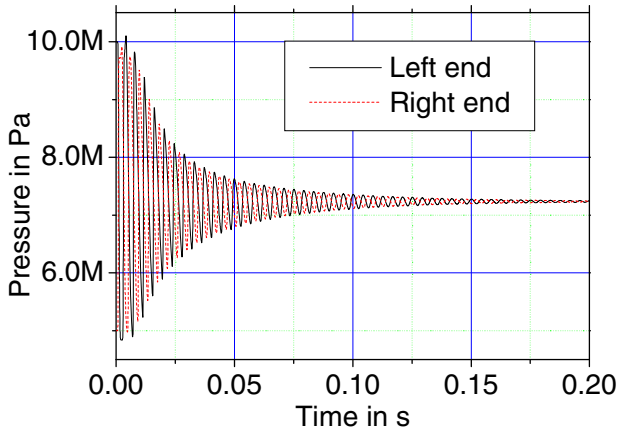


**Fig. 4.3** Pressure as function of the distance from the left closed end of the pipe at  $\tau = 250\mu s$ . Comparison between the analytical and two numerical solutions obtained by using the explicit method of characteristics: one with piecewise linear interpolation, the second with cubic spline interpolation. The systems of PDE's contain the energy conservation principle in terms of specific internal entropy. As expected increasing the order of the spatial discretization increases the accuracy

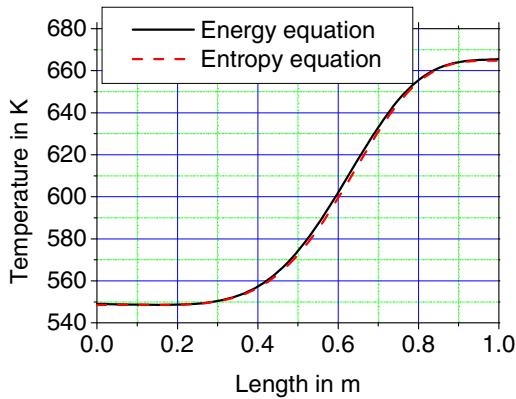


**Fig. 4.4** Velocity as function of the distance from the left closed end of the pipe at  $\tau = 250\mu s$ . Comparison between the analytical and two numerical solutions obtained by using the explicit method of characteristics: one with piecewise linear interpolation, the second with cubic spline interpolation. The systems of PDE's contain the energy conservation principle in terms of specific internal entropy. As expected increasing the order of the spatial discretization increases the accuracy





**Fig. 4.5** Pressures at the both ends of the pipe as functions of time. The energy conservation is used in terms of specific internal energy. The result using the energy conservation in terms of specific entropy is almost indistinguishable



**Fig. 4.6** Spatial distribution of the gas temperature at the 0.4th second. Comparison between the predictions using the energy conservation in terms of specific internal energy or entropy

Taking into account the equidistant discretization and the piecewise linear dependence of the variables between two adjacent points, the initial mass, the total energy and the total entropy in the pipe are

$$\frac{m_0}{\Delta z A} = \frac{1}{2} \sum_{k=1}^{k_{\max}-1} (\rho_k + \rho_{k+1}), \quad (4.60)$$

$$\frac{E_0}{\Delta z A} = \frac{1}{2} \sum_{k=1}^{k_{\max}-1} \left[ \rho_k \left( e_k + \frac{1}{2} w_k^2 \right) + \rho_{k+1} \left( e_{k+1} + \frac{1}{2} w_{k+1}^2 \right) \right], \quad (4.61)$$

$$\frac{S_0}{\Delta z A} = \frac{1}{2} \sum_{k=1}^{k_{\max}-1} (\rho_k s_k + \rho_{k+1} s_{k+1}), \quad (4.62)$$

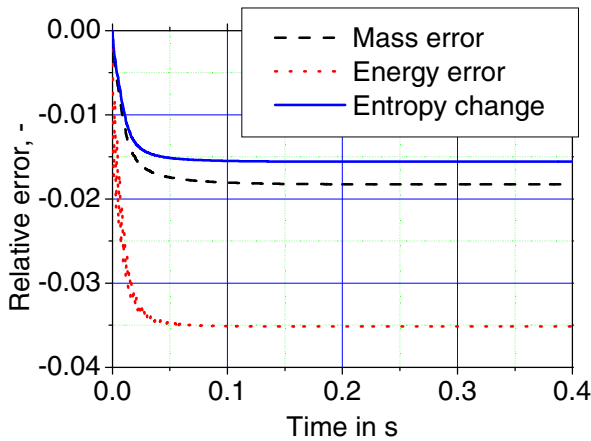
respectively. After each time step we compute the mass, the total energy and the total entropy of the system. Thus, the relative mass and total energy errors are

$$\varepsilon_m = \frac{m}{m_0} - 1, \quad (4.63)$$

$$\varepsilon_E = \frac{E}{E_0} - 1, \quad (4.64)$$

respectively. Next we show a numerical computation for initial conditions of 100 and 50 *bar* respectively and initial temperature 600 *K*. The material properties used are  $c_p = 1034.09$ ,  $R = 287.04$ ,  $T_0 = 1$  and  $p_0 = 1$ . The small computer code used here was written in FORTRAN 90 and runs using 32 or 16 bits floating point operations on real numbers. The computation was performed using a SGI Octane work station with a R10000 processor.

Figures 4.1 and 4.2 show comparisons between an analytical solution *Oertel* (1966) for the problem defined at  $\tau = 250 \mu s$ . The time step used here was 0.1  $\mu s$ . We see that both systems deliver indistinguishable numerical solutions. Both numerical solutions suffer from considerable numerical diffusion. Repeating the computation with single precision does not change the numerical diffusion. This is evidence that the numerical solution suffers from a considerable *truncation* error introduced by the lower order approximation of the PDE's with finite difference algebraic equations in time and space. Improvement is, as expected, achievable by increasing the order of the discretization as demonstrated in Figs. 4.3 and 4.4. The latter results are obtained by using instead of piecewise linear interpolation piecewise cubic spline interpolation.



**Fig. 4.7** Accumulated relative error as a function of time. The conservation of energy is used in terms of specific entropy

Next we perform a computation for a much longer time. The time step was forced to satisfy the condition  $Cu_i \equiv \lambda_i \Delta \tau / \Delta z = 0.1$ . Figure 4.5 shows the pressure at the both ends of the pipe as a function of time. The computation was performed with a system containing the energy conservation in terms of specific energy. We see pressure waves are traveling through the pipe up to 0.2 s. The computation was stopped at 0.4th second when the “mechanical equilibrium” inside the pipe is almost reached. As it will be explained later the damping of the oscillations is an artificial product of the numerical method with a large truncation error acting as dissipative force. Now we repeat the same computation with a system containing the energy conservation in terms of specific entropy. The comparison of the two results shows that they are hardly distinguishable. The spatial distribution of the temperature at the 0.4th second is presented in Fig. 4.6. We found that the difference between the two results is very small.

Figure 4.7 presents the accumulation of the numerical errors as a function of time. 3.5% energy error means that the averaged temperature is predicted with an error  $\Delta T = \varepsilon_E (T^* - T_0) \approx 21K$ . Again the accuracy of both numerical solutions is hardly distinguishable.

As noted at the beginning of our discussion, there are no dissipative diffusion terms in the solved equations. A prominent consequence of this model simplification is manifested by the temperature distribution after the “end” of the dynamic oscillations, Fig. 4.6. Actually, the nature will enforce spatial thermal equilibrium after some time due to molecular diffusion.

Remarkable is the fact that the largest error originates in the initial phase of the oscillation process where the replacement of complicated non-linear spatial

distribution with piecewise linear distribution is obviously not appropriate. Lower order numerical methods predict acceptably only the first few cycles but then the solution degrades destroyed by numerical diffusion. The inability of the lower order numerical methods to solve fluid mechanics equations for initial and boundary conditions defining strong gradients is not evidence that the equations are wrong. It is beyond question that high accuracy solution methods are necessary to satisfy future needs of the industry. Such a method is discussed in Section 12.15.3.

This computational analysis shows a behavior of the numerical solutions which is indeed expected. The solutions varies slightly only due to the additional error made by discretization of two terms in the energy conservation equation (4.4a) instead of one in the entropy equation (4.6).

Secondly, the viscosity of the fluid resists the relative movement of the fluid particles. To illustrate this effect let us introduce the viscous force in accordance with the *Stokes* (1845) hypothesis for a compressible fluid into the momentum equation. For simplicity, we use the one-dimensional notation. The corresponding energy change per unit time and unit flow volume is introduced into the energy conservation equation. For comparison we also introduce a term  $\dot{q}'''$  representing the heat introduction into the flow per unit time and unit flow volume. The modified systems is then

$$\frac{\partial \rho}{\partial \tau} + \frac{\partial}{\partial z}(\rho w) = 0 \quad (4.65)$$

$$\frac{\partial}{\partial \tau}(\rho w) + \frac{\partial}{\partial z}(\rho w w) + \frac{\partial p}{\partial z} + \rho g_z = \frac{4}{3} \frac{\partial}{\partial z} \left( \rho \nu \frac{\partial w}{\partial z} \right) \quad (4.66)$$

$$\frac{\partial}{\partial \tau} \left[ \rho \left( e + \frac{1}{2} w^2 \right) \right] + \frac{\partial}{\partial z} \left[ \rho \left( e + p / \rho + \frac{1}{2} w^2 \right) w \right] + \rho g_z w = \frac{4}{3} \frac{\partial}{\partial z} \left( \rho \nu w \frac{\partial w}{\partial z} \right) + \dot{q}''' \quad (4.67)$$

Performing the same transformation as described in the previous chapter we obtain the corresponding equations

$$\frac{\partial \rho}{\partial \tau} + w \frac{\partial \rho}{\partial z} + \rho a^2 \frac{\partial w}{\partial z} = - \left( \frac{\partial \rho}{\partial s} \right)_s \frac{a^2}{\rho T} \left[ \dot{q}''' + \frac{4}{3} \rho \nu \left( \frac{\partial w}{\partial z} \right)^2 \right] \quad (4.68)$$

$$\frac{\partial w}{\partial \tau} + w \frac{\partial w}{\partial z} + \frac{1}{\rho} \frac{\partial p}{\partial z} = -g_z + \frac{1}{\rho} \frac{4}{3} \frac{\partial}{\partial z} \left( \rho \nu \frac{\partial w}{\partial z} \right) \quad (4.69)$$

$$\frac{\partial e}{\partial \tau} + w \frac{\partial e}{\partial z} + \frac{p}{\rho} \frac{\partial w}{\partial z} = \frac{1}{\rho} \left[ \dot{q}''' + \frac{4}{3} \rho \nu \left( \frac{\partial w}{\partial z} \right)^2 \right] \quad (4.70)$$

or

$$\frac{\partial s}{\partial \tau} + w \frac{\partial s}{\partial z} = \frac{1}{\rho T} \left[ \dot{q}''' + \frac{4}{3} \rho \nu \left( \frac{\partial w}{\partial z} \right)^2 \right] \quad (4.71)$$

Again the systems containing the energy conservation in terms of specific internal energy and those in terms of specific entropy are identical. We see that the viscous dissipation per unit flow volume  $w \frac{4}{3} \frac{\partial}{\partial z} \left( \rho \nu \frac{\partial w}{\partial z} \right)$  cancels. This is the reversible component of the viscous dissipation of mechanical energy. The parts which remains  $\frac{4}{3} \rho \nu \left( \frac{\partial w}{\partial z} \right)^2$  is called irreversible dissipation. Looking at Eq. (4.70) we immediately recognize that it acts in a way transferring mechanical energy in thermal energy. The molecular cinematic viscosity has typically very small values, e.g. 40/100000  $m^2/s$  for air at atmospheric conditions. Numerical solutions with large truncation errors such as the solutions discussed above introduce much more numerical diffusion so that for such methods the effect of this term can not be actually seen. In nature this effect is also small except for singularities in the flow. Such a singularity is a shock front. In a shock front the dissociation processes can make this value effectively considerable higher *Oertel* (1966).

The pseudo-canonical forms of the two systems are

$$\frac{dz}{d\tau} = w + a, \quad (4.72)$$

$$\frac{1}{\rho a} \frac{dp}{d\tau} + \frac{dw}{d\tau} = -\frac{1}{\rho} \left( \frac{\partial \rho}{\partial s} \right)_p \frac{a}{\rho T} \left[ \dot{q}''' + \frac{4}{3} \rho \nu \left( \frac{\partial w}{\partial z} \right)^2 \right] - g_z + \frac{1}{\rho} \frac{4}{3} \frac{\partial}{\partial z} \left( \rho \nu \frac{\partial w}{\partial z} \right) \quad (4.73)$$

$$\frac{dz}{d\tau} = w - a, \quad (4.74)$$

$$-\frac{1}{\rho a} \frac{dp}{d\tau} + \frac{dw}{d\tau} = \frac{1}{\rho} \left( \frac{\partial \rho}{\partial s} \right)_p \frac{a}{\rho T} \left[ \dot{q}''' + \frac{4}{3} \rho \nu \left( \frac{\partial w}{\partial z} \right)^2 \right] - g_z + \frac{1}{\rho} \frac{4}{3} \frac{\partial}{\partial z} \left( \rho \nu \frac{\partial w}{\partial z} \right) \quad (4.75)$$

$$\frac{dz}{d\tau} = w, \quad \frac{ds}{d\tau} = \frac{1}{\rho T} \left[ \dot{q}''' + \frac{4}{3} \rho \nu \left( \frac{\partial w}{\partial z} \right)^2 \right]. \quad (4.76, 4.77)$$

or

$$\frac{dz}{d\tau} = w, \quad \frac{de}{d\tau} - \frac{p}{(\rho a)^2} \frac{dp}{d\tau} = \left[ 1 + \frac{p}{\rho T} \frac{1}{\rho} \left( \frac{\partial p}{\partial s} \right)_p \right] \frac{1}{\rho} \left[ \dot{q}'' + \frac{4}{3} \rho v \left( \frac{\partial w}{\partial z} \right)^2 \right]. \quad (4.78, 4.79)$$

Pseudo-canonical because the dissipative terms are considered as a sources for the hyperbolic part.

## 4.11 Interpenetrating fluids

Mixtures are an example of interpenetrating fluids. For an *infinite time* gas species penetrate the space occupied by the total mixture in a way that in each point the same concentration is observed. This is known as a *Dalton's law*. What does *infinite time* mean? The time scale for complete mixing depends on the size of the considered volume, of the specific properties of the constituents to each other and of course on the initial state of the system. The conservation equations for mass, momentum and energy can be written in the same way as for the single component by considering that the single component occupies all the volume, possesses locally the same temperature as the mixture temperature and exerts on walls partial pressure corresponding to each component so that the sum of the partial pressures is equal to the total pressure. It is common practice to introduce some averaged properties and to sum the component equations. The equations are first local volume averaged and then time averaged. For simplicity we do not consider any dissipative effects like diffusion, viscous effects etc. except fluid mixing due to convection and local mass sources. We give here the final results for one-dimensional flow in a thermally isolated non-deformable pipe with constant cross section without derivation.

$$\frac{\partial \rho}{\partial \tau} + \frac{\partial}{\partial z} (\rho w) = \mu \quad (4.80)$$

$$\frac{\partial C_i}{\partial \tau} + w \frac{\partial C_i}{\partial z} = \frac{1}{\rho} (\mu_i - C_i \mu) \quad i = 2, i_{\max} \quad (4.81)$$

$$\frac{\partial w}{\partial \tau} + w \frac{\partial w}{\partial z} + \frac{1}{\rho} \frac{\partial p}{\partial z} + g_z = 0, \quad (4.82)$$

$$\rho \left( \frac{\partial h}{\partial \tau} + w \frac{\partial h}{\partial z} \right) - \left( \frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} \right) = \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) \quad (4.83)$$

where

$$h = \sum_{i=1}^{i_{\max}} C_i h_i. \quad (4.84)$$

The first equation reflects the mass conservation for the mixture. The second group of  $i_{\max} - 1$  equations reflects the mass conservation for each velocity field. Note that the sum of all mass concentrations is equal to one,

$$\sum_{i=1}^{i_{\max}} C_i = 1, \quad (4.85)$$

or

$$C_1 = 1 - \sum_{i=2}^{i_{\max}} C_i \quad (4.86)$$

Equation (4.83) is remarkable. It says that at constant pressure the injected mass sources contribute to the total enthalpy change with their deviation of the external component enthalpies from the corresponding component enthalpy inside the flow. More information about the rigorous derivation of the above set of equations for the general case of multi-phase multi-component flows is presented in the next Chapter (first published in *Kolev* (1990, 1994a, b, 1995, 1997a)). Next we will write the energy conservation equation in terms of entropy. For this purpose define the mixture entropy in a similar way as the mixture enthalpy

$$s = \sum_{i=1}^{i_{\max}} C_i s_i \quad (4.87)$$

Obviously

$$Tds = T \sum_{i=1}^{i_{\max}} C_i ds_i + T \sum_{i=2}^{i_{\max}} (s_i - s_1) dC_i. \quad (4.88)$$

Having in mind that

$$C_i = \frac{\rho_i}{\rho} \quad (4.89)$$

and replacing with the *Gibbs* equation (4.14) for each component  $T\rho_i ds_i = \rho_i dh_i - dp_i$  or

$$TC_i ds_i = C_i dh_i - \frac{dp_i}{\rho} \quad (4.90)$$

and rearranging results in the *Gibbs* mixture equation (compare with Eq. (12) in *ibbs* (1992))

$$Tds = \sum_{i=1}^{i_{\max}} C_i dh_i - \frac{dp}{\rho} + T \sum_{i=2}^{i_{\max}} (s_i - s_1) dC_i = dh - \frac{dp}{\rho} - \sum_{i=2}^{i_{\max}} [h_i - Ts_i - (h_1 - Ts_1)] dC_i . \quad (4.91)$$

Now, Eq. (4.83) is easily transformed into the entropy equation

$$\rho T \left( \frac{\partial s}{\partial \tau} + w \frac{\partial s}{\partial z} \right) = \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) + T \sum_{i=1}^{i_{\max}} (\mu_i - \mu C_i) s_i \quad (4.92)$$

using Eqs. (4.91) and (4.81). Again we see that the entropy equation is simpler than Eq. (4.83). The temperature equation can be derived from Eq. (4.92) by using the equation of state, Eq. (3.98),

$$\rho T ds = \rho c_p dT - \left( 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all } C_i} \right) dp + \rho T \sum_{i=2}^{i_{\max}} \left( \frac{\partial s}{\partial C_i} \right)_{p, T, \text{all } C_i \text{ except } C_i} dC_i \quad (4.93)$$

derived by this author in *Kolev (1990)*. Here

$$\left( \frac{\partial s}{\partial C_i} \right)_{p, T, \text{all } C_i \text{ except } C_i} = s_i - s_1 + \Delta s_i^{np} \quad (4.94)$$

In the intermediate result

$$\begin{aligned} & \rho c_p \left( \frac{\partial T}{\partial \tau} + w \frac{\partial T}{\partial z} \right) - \left( 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all } C_i} \right) \left( \frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} \right) \\ & + \rho T \sum_{i=2}^{i_{\max}} \left( \frac{\partial s}{\partial C_i} \right)_{p, T, \text{all } C_i \text{ except } C_i} \left( \frac{\partial C_i}{\partial \tau} + w \frac{\partial C_i}{\partial z} \right) = \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) + T \sum_{i=1}^{i_{\max}} (\mu_i - \mu C_i) s_i \end{aligned} \quad (4.95)$$

we replace the RHS of the concentration equations instead of the total variation of the concentrations. Keeping in mind that

$$\sum_{i=1}^{i_{\max}} (\mu_i - \mu C_i) s_i - \sum_{i=2}^{i_{\max}} (s_i - s_1) (\mu_i - C_i \mu) = 0 \quad (4.96)$$

we obtain a very informative form of the temperature equation

$$\rho c_p \left( \frac{\partial T}{\partial \tau} + w \frac{\partial T}{\partial z} \right) - \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all } C_i} \right] \left( \frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} \right)$$



$$= \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) - T \sum_{i=2}^{i_{\max}} \Delta s_i^{np} (\mu_i - C_i \mu). \quad (4.97)$$

The mixture mass conservation equation can be rearranged similarly as in the single component case by using the equation of state

$$d\rho = \frac{d\rho}{a^2} + \left( \frac{\partial \rho}{\partial s} \right)_{p, \text{all\_}C's} ds + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho}{\partial C_i} \right)_{p, s, \text{all\_}C's\_ \text{except\_}C_i} dC_i \quad (4.98)$$

namely

$$\begin{aligned} & \frac{1}{a^2} \left( \frac{\partial \rho}{\partial \tau} + w \frac{\partial \rho}{\partial z} \right) + \left( \frac{\partial \rho}{\partial s} \right)_{p, \text{all\_}C's} \left( \frac{\partial s}{\partial \tau} + w \frac{\partial s}{\partial z} \right) + \\ & \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho}{\partial C_i} \right)_{p, s, \text{all\_}C's\_ \text{except\_}C_i} \left( \frac{\partial C_i}{\partial \tau} + w \frac{\partial C_i}{\partial z} \right) + \rho \frac{\partial w}{\partial z} = \mu \end{aligned} \quad (4.99)$$

or

$$\frac{\partial \rho}{\partial \tau} + w \frac{\partial \rho}{\partial z} + \rho a^2 \frac{\partial w}{\partial z} = Dp \quad (4.100)$$

where

$$Dp = a^2 \left\{ \begin{array}{l} \mu - \frac{1}{\rho T} \left( \frac{\partial \rho}{\partial s} \right)_{p, \text{all\_}C's} \left( \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) + T \sum_{i=1}^{i_{\max}} (\mu_i - \mu C_i) s_i \right) \\ - \frac{1}{\rho} \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho}{\partial C_i} \right)_{p, s, \text{all\_}C's\_ \text{except\_}C_i} (\mu_i - C_i \mu) \end{array} \right\}. \quad (4.101)$$

The source term can be further simplified by using the following expressions for the derivatives

$$\left( \frac{\partial \rho}{\partial s} \right)_{p, \text{all\_}C's} = \left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all\_}C's} \frac{T}{c_p}, \quad (4.102)$$

$$\begin{aligned} & \left( \frac{\partial \rho}{\partial C_i} \right)_{p, s, \text{all\_}C's\_ \text{except\_}C_i} = \left( \frac{\partial \rho}{\partial C_i} \right)_{p, T, \text{all\_}C's\_ \text{except\_}C_i} \\ & - \left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all\_}C's} \frac{T}{c_p} \left( \frac{\partial s}{\partial C_i} \right)_{p, T, \text{all\_}C's\_ \text{except\_}C_i}, \end{aligned} \quad (4.103)$$

and Eq. (4.96). The result is

$$Dp = a^2 \left\{ \begin{array}{l} \mu - \frac{1}{\rho c_p} \left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all } C_i} \left[ \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) - T \sum_{i=2}^{i_{\max}} \Delta S_i^{np} (\mu_i - C_i \mu) \right] \\ - \frac{1}{\rho} \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho}{\partial C_i} \right)_{p, T, \text{all } C_i \text{ except } C_i} (\mu_i - C_i \mu) \end{array} \right\}. \quad (4.104)$$

For mixture of ideal gases the specific component enthalpies is not a function of pressure, therefore  $\left( \frac{\partial h}{\partial p} \right)_{T, \text{all } C_i} = 0$ , and  $\Delta S_i^{np} = 0$ ,  $\left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all } C_i} = \frac{1}{RT}$ ,

$$\left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all } C_i} = -\frac{\rho}{T}, \quad \left( \frac{\partial \rho}{\partial C_i} \right)_{p, T, \text{all } C_i \text{ except } C_i} = -\frac{\rho}{R} (R_i - R_1). \text{ Consequently}$$

$$\rho c_p \left( \frac{\partial T}{\partial \tau} + w \frac{\partial T}{\partial z} \right) - \left( \frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} \right) = \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i), \quad (4.105)$$

and

$$Dp = a^2 \left[ \mu + \frac{1}{c_p T} \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) + \frac{1}{R} \sum_{i=2}^{i_{\max}} (R_i - R_1) (\mu_i - C_i \mu) \right] \quad (4.106)$$

The equation of state for a mixture of  $i_{\max}$  ideal gases is  $\rho = p/(RT)$ , where

$$R = \sum_{i=1}^{i_{\max}} C_i R_i, \quad a = \sqrt{\kappa RT}, \quad \text{where } \kappa = c_p / (c_p - R), \quad c_p = \sum_{i=1}^{i_{\max}} C_i c_{pi}, \quad c_v = \sum_{i=1}^{i_{\max}} C_i c_{vi},$$

$$p_i = Y_i p, \quad \text{where } Y_i = \frac{R_i}{R} C_i, \quad s = \sum_{i=1}^{i_{\max}} C_i s_i = \sum_{i=1}^{i_{\max}} C_i c_{pi} \ln(T/T_{0i}),$$

$$- \sum_{i=1}^{i_{\max}} C_i R_i \ln(p_i/p_{0i}) \quad \text{where } s_i = c_{pi} \ln(T/T_{0i}) - R_i \ln(p_i/p_{0i}),$$

$$e = \sum_{i=1}^{i_{\max}} C_i e_i = \sum_{i=1}^{i_{\max}} C_i c_{vi} (T - T_{0i}), \quad \text{where } e_i = c_{vi} (T - T_{0i}),$$

$$h = \sum_{i=1}^{i_{\max}} C_i h_i = \sum_{i=1}^{i_{\max}} C_i c_{pi} (T - T_{0i}) \quad \text{where } h_i = c_{pi} (T - T_{0i}).$$

The caloric equation of state can be considerably simplified if one uses the same reference point for integration constants set to zero, that is,  $T_{0i} = T_0$  and  $p_{0i} = p_0$ . The result is

$$\begin{aligned}
 s &= \sum_{i=1}^{i_{\max}} C_i s_i = c_p \ln(T/T_0) - \sum_{i=1}^{i_{\max}} C_i R_i \ln(p_i/p_0) \\
 &= c_p \ln(T/T_0) - R \ln\left(\frac{p}{p_0}\right) - \sum_{i=1}^{i_{\max}} C_i R_i \ln\left(\frac{C_i R_i}{R}\right) \\
 &= c_p \ln(T/T_0) - R \left( \ln \frac{p}{p_0} + \sum_{i=1}^{i_{\max}} Y_i \ln Y_i \right), \\
 T &= T_0 \exp \left[ \frac{s}{c_p} + \frac{R}{c_p} \ln\left(\frac{p}{p_0}\right) + \frac{1}{c_p} \sum_{i=1}^{i_{\max}} C_i R_i \ln\left(\frac{C_i R_i}{R}\right) \right] \\
 &= T_0 \left( \frac{p}{p_0} \right)^{\frac{R}{c_p}} \exp\left(\frac{s}{c_p}\right) \prod_{i=1}^{i_{\max}} \left( \frac{C_i R_i}{R} \right)^{\frac{C_i R_i}{c_p}}, \\
 e &= c_v (T - T_0), \\
 h &= c_p (T - T_0).
 \end{aligned}$$

For two components  $M$  and  $n$  we have

$$T = T_0 \left( \frac{p}{p_0} \right)^{\frac{R}{c_p}} \exp\left(\frac{s}{c_p}\right) \left( \frac{C_n R_n}{R} \right)^{\frac{C_n R_n}{c_p}} \left[ \frac{(1 - C_n) R_M}{R} \right]^{\frac{(1 - C_n) R_M}{c_p}}.$$

## 4.12 Summary of different notations for interpenetrating fluids

For convenience we summarize the results of Section 4.11. The system of partial differential equations describing the multi-component flow is

$$\frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} + \rho a^2 \frac{\partial w}{\partial z} = Dp \quad (4.107)$$

$$\frac{\partial C_i}{\partial \tau} + w \frac{\partial C_i}{\partial z} = \frac{1}{\rho} (\mu_i - C_i \mu) \quad i = 2, i_{\max} \quad (4.108)$$

$$\frac{\partial w}{\partial \tau} + w \frac{\partial w}{\partial z} + \frac{1}{\rho} \frac{\partial p}{\partial z} + g_z = 0, \quad (4.109)$$

$$\rho T \left( \frac{\partial s}{\partial \tau} + w \frac{\partial s}{\partial z} \right) = \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) + T \sum_{i=1}^{i_{\max}} (\mu_i - \mu C_i) s_i \quad (4.110)$$

or

$$\rho \left( \frac{\partial h}{\partial \tau} + w \frac{\partial h}{\partial z} \right) - \left( \frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} \right) = \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) \quad (4.111)$$

or

$$\begin{aligned} \rho c_p \left( \frac{\partial T}{\partial \tau} + w \frac{\partial T}{\partial z} \right) - \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T, all\_C's} \right] \left( \frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} \right) \\ = \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) - T \sum_{i=2}^{i_{\max}} \Delta s_i^{np} (\mu_i - C_i \mu). \end{aligned} \quad (4.112)$$

As for the single-component flows, the energy conservation can be noted in terms of different variables. Equations (4.110), (4.111) and (4.112) which are completely equivalent are few of the possible examples. Preferring one of them for practical analysis is only a matter of convenience. Obviously the entropy concept offers again the simplest equation. Analysis of the type of the system is again greatly simplified by using the entropy equation. Thus, using the following vector of dependent variables

$$U = (p, w, s, C_{i=2, i_{\max}}) \quad (4.113)$$

the flow is completely described by the system of quasi linear partial differential equations

$$\frac{\partial p}{\partial \tau} + w \frac{\partial p}{\partial z} + \rho a^2 \frac{\partial w}{\partial z} = Dp \quad (4.114)$$

$$\frac{\partial w}{\partial \tau} + w \frac{\partial w}{\partial z} + \frac{1}{\rho} \frac{\partial p}{\partial z} + g_z = 0, \quad (4.115)$$

$$\frac{\partial s}{\partial \tau} + w \frac{\partial s}{\partial z} = Ds \quad (4.116)$$

$$\frac{\partial C_i}{\partial \tau} + w \frac{\partial C_i}{\partial z} = DC_i \quad i = 2, i_{\max} \quad (4.117)$$

where

$$Ds = \frac{1}{\rho T} \left[ \sum_{i=1}^{i_{\max}} \mu_i (h_i^\sigma - h_i) + T \sum_{i=1}^{i_{\max}} (\mu_i - \mu C_i) s_i \right] \quad (4.118)$$

$$DC_i = \frac{1}{\rho} (\mu_i - C_i \mu) \quad (4.119)$$

The eigen values of this system are real and different from each other

$$\lambda_{1,2} = w \pm a, \quad \lambda_{3,i=2,j_{\max}} = w. \quad (4.120, 4.121, 4.122)$$

The eigen vectors of the transposed characteristic matrix are linearly independent. Therefore the quasi linear system of partial differential equations is hyperbolic and can be transformed into the very simple and useful canonical form

$$\frac{dz}{d\tau} = w + a, \quad \frac{1}{\rho a} \frac{dp}{d\tau} + \frac{dw}{d\tau} = \frac{1}{\rho a} Dp - g \quad (4.123, 4.124)$$

$$\frac{dz}{d\tau} = w - a, \quad -\frac{1}{\rho a} \frac{dp}{d\tau} + \frac{dw}{d\tau} = -\frac{1}{\rho a} Dp - g \quad (4.125, 4.126)$$

$$\frac{dz}{d\tau} = w, \quad \frac{ds}{d\tau} = Ds, \quad (4.127, 4.128)$$

$$\frac{dz}{d\tau} = w, \quad \frac{dC_i}{d\tau} = DC_i. \quad (4.129, 4.130)$$

We see the remarkable behavior that concentrations and entropy change along the characteristic line defined with the mixture velocity. For the here considered simplified case if there are no mass sources in the pipe the concentrations and the entropy do not change along the characteristic line for a given time step - convective transport only. In case of mass sources, the mixing process changes the concentration and increases the entropy along the characteristic line for a given time step. One should again not forget that at the beginning of Section 4.11 we have neglected all other dissipative effects which cause entropy increase in the real nature and that our final system is only an idealization of the nature.

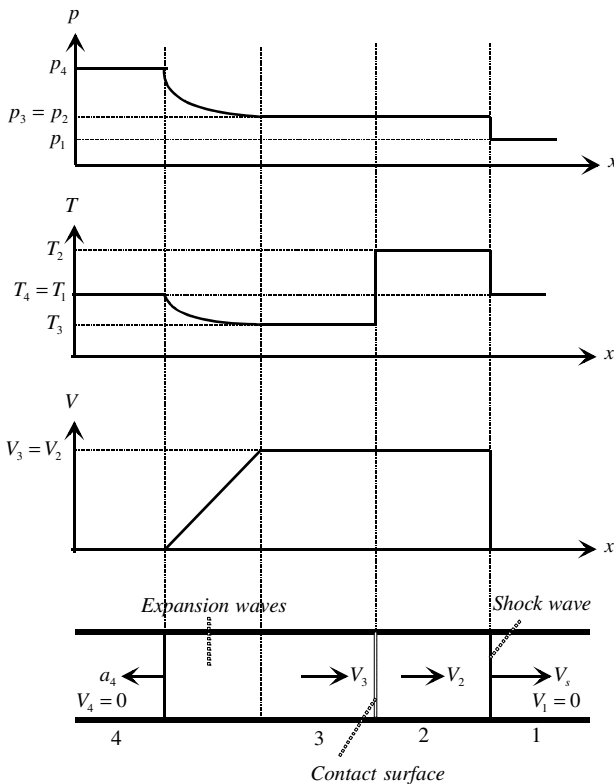
## Appendix 4.1 Analytical solution of the shock tube problem

Consider a pipe with a constant cross section filled with perfect gas. The left halve of the pipe is separated from the right with a membrane that separates high pressure region 4, from low pressure region 1 – see Fig- 4.1-1. The membrane is removed and a shock wave propagates from the left to the right and a refraction

wave from right to the left. There is a text book solution for this case that is widely used as a benchmark for numerical methods. This solution is given below.

1) First the pressure ratio  $p_2/p_1$  is found by solving the following equations by iterations

$$\frac{p_4}{p_1} = \frac{p_2}{p_1} \left\{ 1 - \frac{(\kappa_4 - 1) \frac{a_1}{a_4} \left( \frac{p_2}{p_1} - 1 \right)}{\left[ 4\kappa_1^2 + 2\kappa_1 (\kappa_1 + 1) \left( \frac{p_2}{p_1} - 1 \right) \right]^{1/2}} \right\}^{\frac{2\kappa_4}{\kappa_4 - 1}} \quad (1)$$



**Fig. A4.1-1** Analytical solution to the perfect gas shock wave problem in a pipe with constant cross section

2) The shock Mach number is found by solving the following equation by iterations

$$\frac{p_4}{p_1} = \frac{\kappa_1 - 1}{\kappa_1 + 1} \left( \frac{2\kappa_1}{\kappa_1 - 1} Ma_s^2 - 1 \right) \left[ 1 - \frac{\frac{\kappa_4 - 1}{\kappa_4 + 1} a_1 (Ma_s^2 - 1)}{Ma_s} \right]^{\frac{2\kappa_4}{\kappa_4 - 1}}. \quad (2)$$

3) Knowing the pressure ratio  $p_2/p_1$  all parameter behind the shock are computed as follows

$$\frac{T_2}{T_1} = \frac{p_2}{p_1} \frac{\frac{\kappa_1 + 1}{\kappa_1 - 1} + \frac{p_2}{p_1}}{1 + \frac{\kappa_1 + 1}{\kappa_1 - 1} \frac{p_2}{p_1}}, \quad (3)$$

$$\frac{\rho_2}{\rho_1} = \frac{1 + \frac{\kappa_1 + 1}{\kappa_1 - 1} \frac{p_2}{p_1}}{\frac{\kappa_1 + 1}{\kappa_1 - 1} + \frac{p_2}{p_1}}. \quad (4)$$

4) The shock velocity and the shock Mach number are computed alternatively by

$$V_s = a_1 \left[ 1 + \frac{\kappa_1 + 1}{2\kappa_1} \left( \frac{p_2}{p_1} - 1 \right) \right]^{1/2}, \quad (5)$$

$$Ma_s = V_s / a_1. \quad (6)$$

The velocity of the gas behind the shock is then

$$V_2 = \frac{a_1}{\kappa_1} \left( \frac{p_2}{p_1} - 1 \right) \left( \frac{\frac{2\kappa_1}{\kappa_1 + 1}}{\frac{\kappa_1 - 1}{\kappa_1 + 1} + \frac{p_2}{p_1}} \right)^{1/2}. \quad (7)$$

The parameter in the region 3 are computed as follows

$$\frac{p_3}{p_4} = \frac{p_3}{p_1} \frac{p_1}{p_4} = \frac{p_2}{p_1} \frac{p_1}{p_4}, \quad (8)$$

$$\frac{T_3}{T_4} = \left( \frac{p_3}{p_4} \right)^{\frac{\kappa_4 - 1}{\kappa_4}}, \quad (9)$$

$$\frac{\rho_3}{\rho_4} = \left( \frac{p_3}{p_4} \right)^{\frac{1}{\kappa_4}}, \quad (10)$$

$$\frac{a_4}{a_3} = \left( \frac{p_4}{p_3} \right)^{\frac{\kappa_4-1}{2\kappa_4}}, \quad (11)$$

$$Ma_3 = \frac{2}{\kappa_4-1} \left[ \left( \frac{p_4}{p_3} \right)^{\frac{\kappa_4-1}{2\kappa_4}} - 1 \right], \quad (11)$$

$$V_3 = Ma_3 a_3. \quad (12)$$

The flow parameters in the refraction wave are functions of time and pace. The relation between the position and time along the backwards characteristic

$$\frac{dx}{d\tau} = V - a, \quad (13)$$

is

$$x = (V - a)\tau. \quad (14)$$

From

$$V + \frac{2a}{\kappa-1} = const, \quad (15)$$

The following relations results

$$\frac{a}{a_4} = 1 - \frac{\kappa-1}{2} \frac{V}{a_4}, \quad (16)$$

$$V = \frac{2}{\kappa+1} \left( a_4 + \frac{x}{\tau} \right), \quad (17)$$

The other parameters in the refraction region are then

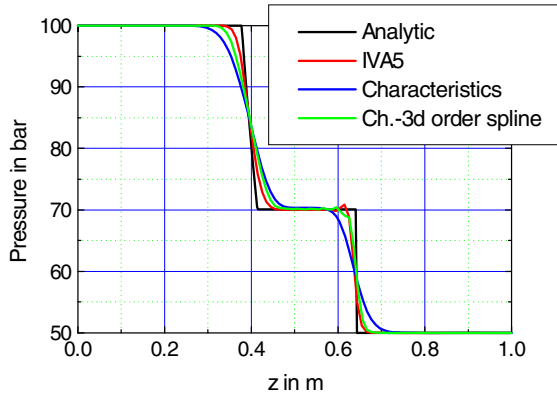
$$a = (\kappa RT)^{1/2}, \quad (18)$$

$$\frac{\rho}{\rho_4} = \left( 1 - \frac{\kappa-1}{2} \frac{V}{a_4} \right)^{\frac{2}{\kappa-1}}, \quad (19)$$

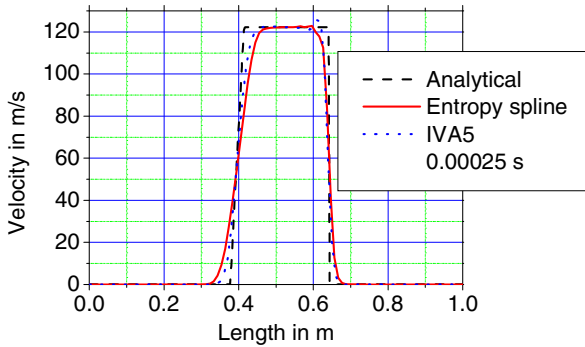
$$\frac{p}{p_4} = \left( 1 - \frac{\kappa-1}{2} \frac{V}{a_4} \right)^{\frac{2\kappa}{\kappa-1}}. \quad (20)$$



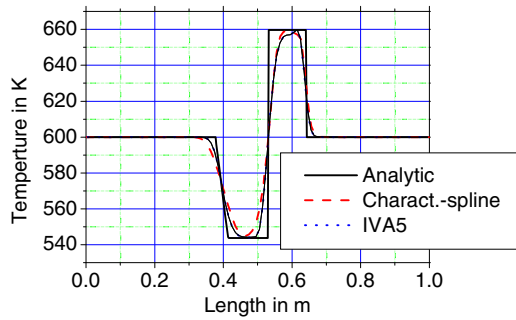
## Appendix 4.2 Achievable accuracy of the donor-cell method for single-phase flows



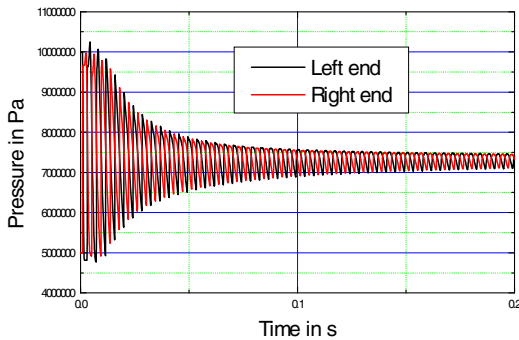
**Fig. A4.2-1** Pressure as function of the distance from the left closed end of the pipe at  $\tau = 250\mu s$ . Comparison between the analytical and four numerical solutions obtained by using the explicit method of characteristics: piecewise linear interpolation using entropy equation, the second solution with cubic spline interpolation, and the third solution is the IVA5 solution



**Fig. A4.2-2** Velocity as function of the distance from the left closed end of the pipe at  $\tau = 250\mu s$ . Comparison between the analytical and numerical solutions. The systems of PDE's contain energy conservation principle in terms of specific internal energy in the first case or specific entropy as used in IVA5 computer code in the second case



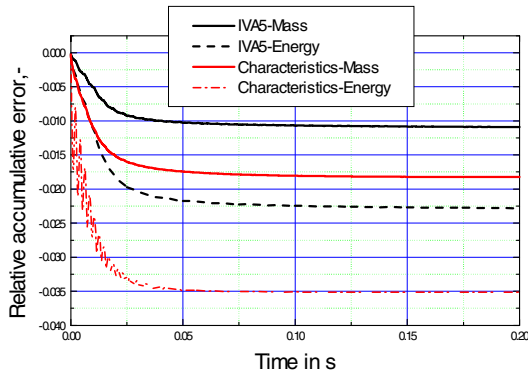
**Fig. A4.2-3** Temperature as function of the distance from the left closed end of the pipe at  $\tau = 250\mu s$ . Comparison between analytical and two numerical solutions. The systems of PDE's contain the energy conservation principle in terms of specific internal energy in the first case or specific entropy as used in IVA5 computer code in the second case



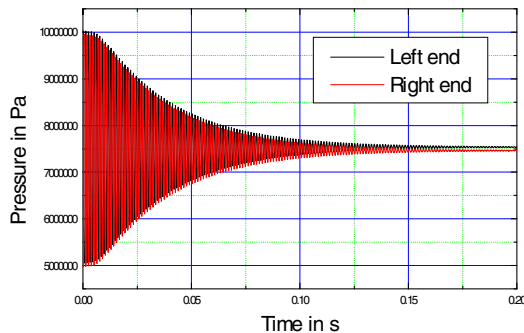
**Fig. A4.2-4** Pressures at the both ends of the pipe as a functions of time (air). The energy conservation is used in terms of specific entropy in IVA5

For many years the first order donor-cell method has been widely used in the multi-phase fluid mechanics because of its simplicity and stability. It is interesting to know how this method compares with the other already discussed methods based on the shock tube problem. Figure A4.2-1 shows the pressure as a function of time computed with the computer code IVA5 which exploits the first order donor-cell method as discussed in *Kolev (1997a, 1999)*. The initial conditions are the same as discussed before. The code uses equation of state for air as real gas. As we see from Figs. A4.2-1 through A4.2-3 the IVA5 donor-cell method is more accurate than the first order method of characteristics and less accurate than the method of characteristic using third order spatial interpolation. The pressure at both ends of the pipe as a function of time is presented in Fig. A4.2-4. Comparing Fig. A4.2-4 with Fig. 4.5 we see almost indistinguishable solutions. The total mass and energy

conservation error is shown in Fig. A4.2-5. Again we observe from Fig. A4.2-5 that the IVA5 numerical method is better than the first order method of characteristics.

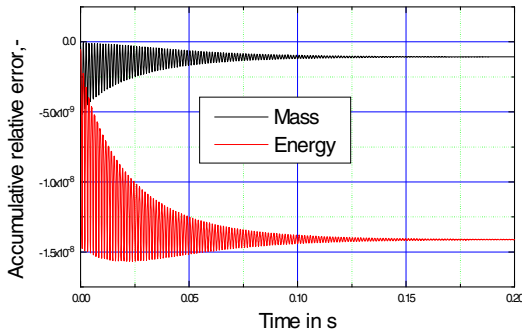


**Fig. A4.2-5** Comparison between the overall mass and energy conservation error of the IVA5 numerical method and the first order method of characteristics



**Fig. A4.2-6** Pressure at the ends of the pipe as a function of time (corium melt 3300 K). The energy conservation is used in terms of specific entropy in IVA5

Now we perform a similar test with IVA5 replacing the air with a molten oxide mixture called corium at 3300 K initial temperature. The initial pressure distribution is the same as in the previous examples. The velocity of sound of the system is around 1500 m/s. For this test we developed a special set of equation of state and their derivatives which are strictly consistent to each other. The pressure at both ends as function of time is shown in Fig. A4.2-6 and the corresponding accumulative relative mass and energy conservation error is given in Fig. A4.2-7. We see the order of magnitude is  $10^{-7}$  which is much better than for the case of the strongly compressible gas. The slight difference in steady state pressures is due to the geodetic pressure difference because the pipe is considered to be vertical.



**Fig. A4.2-7** Relative overall mass and energy conservation error as function of time (corium melt 3300 K). The energy conservation is used in terms of specific entropy in IVA5

For practical applications in large scale facilities spatial resolution of one centimeter is hardly achievable and therefore the expected errors for single phase-flows will be higher corresponding, among others, to the size of the cells.

For multi-phase flow in *Euler* representation there are additional sources of conservation errors, with one of them being associated with the limited minimum of the used time steps. An example for three-phase flow case is given in *Kolev* (1996, 1997b) where the accumulative relative mass error for two seconds transient was below 0.1% and the accumulative energy conservation error was below 1%.

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# 5 First and second laws of the thermodynamics

*Local volume and time averaging is used to derive rigorous energy equations for multi-phase flows in heterogeneous porous media. The flow is conditionally divided into three velocity fields. Each of the fields consists of several chemical components. Using the conservation equations for mass and momentum and the Gibbs equation, entropy equations are rigorously derived. It is shown that the use of the specific entropy as one of the dependent variables results in the simplest method for describing and modeling such a complicated thermodynamic system. A working form of the final entropy equation is recommended for general use in multi-phase flow dynamics.*

## 5.1 Introduction

As in Chapters 1 and 2, from the large number of formulations of the conservation equations for multi-phase flows the local volume averaging as founded by *Anderson* and *Jackson*, *Slattery*, and *Whitaker* was selected to derive rigorously the energy conservation equations for multi-phase flows conditionally divided into three velocity fields. The heterogeneous porous-media formulation introduced by *Gentry* et al., commented by *Hirt*, and used by *Sha*, *Chao* and *Soo*, is then implanted into the formalism as a geometrical skeleton.

Beyond these concepts: I introduce in each of the velocity fields several constituents; The energy equations obtained in this way are then rearranged into the entropy equation using the mass and momentum equations, thereby reflecting the second law of thermodynamics; Then I perform subsequent time averaging; All interfacial integrals are suitably transformed in order to enable practical application.

This yields working equations for each of the three velocity fields that are recommended for general use in multi-phase fluid dynamic analysis.

As far as the author is aware, it was the first time in *Kolev* (1995) that a formulation of the second law of thermodynamics has been presented for such a complicated thermodynamic system as the multi-phase flows consisting of three velocity fields in porous structure, with each of these consisting of several chemical components. The most interesting result of this work is the simplicity of the local volume and time-averaged entropy equation (5.125) finally obtained

$$\rho_l \left[ \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) s_l \right] - \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \mathcal{W} T_l) \\ - \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \nabla C_{il} \right] \right\} = \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right] \quad \text{for } \alpha_l \geq 0,$$

where

$$DT_l^N = \alpha_l \rho_l \left[ \delta_l (P_{k,l} + \varepsilon_{\eta,l}) + \varepsilon_l' \right] + \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_{iwl} (h_{iwl} - h_l) + \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}} \left[ \mu_{Mml} (h_{Ml}^\sigma - h_{Ml}) \right] \\ + \sum_{n=1}^{n_{\max}} \mu_{nml} (h_{nm} - h_{nl}) \\ + \frac{1}{2} \left[ \mu_{wl} (\mathbf{V}_{wl} - \mathbf{V}_l)^2 - \mu_{lw} (\mathbf{V}_{lw} - \mathbf{V}_l)^2 + \sum_{m=1}^3 \mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)^2 \right. \\ \left. + \mu_{wl} \overline{(\mathbf{V}'_{wl} - \mathbf{V}'_l)^2} - \mu_{lw} \overline{(\mathbf{V}'_{lw} - \mathbf{V}'_l)^2} + \sum_{m=1}^3 \mu_{ml} \overline{(\mathbf{V}'_m - \mathbf{V}'_l)^2} \right].$$

This equation is more suitable for general use than the various forms of the energy equation e.g. Eq. (5.109) written in terms of the specific internal energy,

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l e_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l e_l \mathbf{V}_l \gamma) + p_l \left[ \frac{\partial}{\partial \tau} (\alpha_l \gamma_v) + \nabla \cdot (\alpha_l^e \mathbf{V}_l \gamma) \right] \\ - \nabla \cdot (\alpha_l^e \rho_l D_l^e \gamma \cdot \nabla e_l) + \overline{p_l' \nabla \cdot (\alpha_l^e \mathbf{V}_l' \gamma)} = \gamma_v D e_l^*,$$

where

$$D e_l^* = \alpha_l \rho_l \left[ \delta_l (P_{k,l} + \varepsilon_{\eta,l}) + \varepsilon_l' \right] + \dot{q}''' + \frac{1}{2} \left[ \mu_{wl} (\mathbf{V}_{wl} - \mathbf{V}_l)^2 - \mu_{lw} (\mathbf{V}_{lw} - \mathbf{V}_l)^2 \right. \\ \left. + \sum_{m=1}^3 \mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)^2 \right] \\ + \frac{1}{2} \left[ \mu_{wl} \overline{(\mathbf{V}'_{wl} - \mathbf{V}'_l)^2} - \mu_{lw} \overline{(\mathbf{V}'_{lw} - \mathbf{V}'_l)^2} + \sum_{m=1}^3 \mu_{ml} \overline{(\mathbf{V}'_m - \mathbf{V}'_l)^2} \right] \\ + \sum_{i=1}^{i_{\max}} (\mu_{iwl} h_{iwl} - \mu_{ilw} h_{il}) + \sum_{m=1}^3 \left[ \mu_{Mml} h_{Ml}^\sigma + \sum_{n=1}^{n_{\max}} (\mu_{nml} h_{nm} - \mu_{nlm} h_{nl}) \right]$$



or Eq. (5.176) written in terms of temperature

$$\rho_l c_{pl} \left[ \alpha_l \gamma_v \frac{\partial T_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) T_l \right] - \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } C'_s} \right] \left[ \alpha_l \gamma_v \frac{\partial p}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) p \right] \\ - \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \bar{\mathcal{N}} T) + T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{pp} \nabla (\alpha_l^e \rho_l D_{il}^* \gamma \bar{\mathcal{N}} C_{il}) = \gamma_v \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{pp} (\mu_{il} - C_{il} \mu_l) \right],$$

or Eq. (5.115) written in terms of specific enthalpy

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l h_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l \mathbf{V}_l h_l \gamma) - \left( \alpha_l \gamma_v \frac{\partial p_l}{\partial \tau} + \alpha_l^e \mathbf{V}_l \gamma \cdot \nabla p_l \right) \\ - \nabla \cdot \left\{ \alpha_l^e \left[ \lambda_l^* \nabla T_l + \rho_l \sum_{i=2}^{i_{\max}} (h_{il} - h_{l1}) D_{il}^* \nabla C_{il} \right] \gamma \right\} - \alpha_l^e \overline{\mathbf{V}_l' \gamma \cdot \nabla p_l'} \\ + \delta_l (\alpha_l^e \gamma \bar{\mathcal{N}} p_l) \cdot \sum_{i=1}^{i_{\max}} D_{il}^l \nabla \ln C_{il} = \gamma_v DT_l^N + \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^3 \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) h_{il}.$$

Another important result is the so called volume conservation equation

$$\frac{\gamma_v}{\rho a^2} \frac{\partial p}{\partial \tau} + \sum_{i=1}^{i_{\max}} \frac{\alpha_i}{\rho_i \alpha_i^2} (\mathbf{V}_i \gamma \cdot \nabla) p + \nabla \cdot \sum_{i=1}^{i_{\max}} (\alpha_i \mathbf{V}_i \gamma) = \sum_{i=1}^{i_{\max}} D \alpha_i - \frac{\partial \gamma_v}{\partial \tau},$$

where

$$D \alpha_l = \frac{1}{\rho_l} \left\{ \gamma_v \mu_l - \frac{1}{\rho_l} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, \text{all } C'_{ij} s} D s_l^N + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{li}} \right)_{p, s, \text{all } C'_{ij} s \text{ except } C_{l1}} D C_{il}^N \right] \right\} \\ D C_{il}^N = \nabla (\alpha_l \rho_l D_{il}^* \gamma \bar{\mathcal{N}} C_{il}) + \gamma_v (\mu_{il} - C_{il} \mu_l), \\ D s_l^N = \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \bar{\mathcal{N}} T_l) + \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \sum_{i=1}^{i_{\max}} s_{il} D_{il}^* \nabla C_{il} \right] \right\} \\ + \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right].$$

This equation can be used instead of the one of the field mass conservation equations.

This Chapter is an extended version of the work published in *Kolev (1995, 1997a)*.

## 5.2 Instantaneous local volume average energy equations

The energy principle formulated for a volume occupied by the velocity field  $l$  only is as follows:

The sum of the rates of energy added to the velocity field from the surroundings due to conduction and convection plus the rate of work done on the velocity field is equal to the rate of change in the energy of the velocity field as it flows through a volume occupied by this velocity field.

The instantaneous energy conservation equation written per unit field volume is thus

$$\sum_{i=1}^{i_{\max}} \left\{ \frac{\partial}{\partial \tau} \left[ \rho_{il} \left( e_{il}^{\tau} + \frac{1}{2} V_{il}^{\tau 2} \right) \right] + \nabla \cdot \left[ \rho_{il} \left( e_{il}^{\tau} + p_{il}^{\tau} / \rho_{il} + \frac{1}{2} V_{il}^{\tau 2} \right) \mathbf{V}_{il}^{\tau} \right] \right\} - \nabla \cdot (\mathbf{T}_{\eta,l}^{\tau} \cdot \mathbf{V}_l^{\tau}) - \nabla \cdot (\lambda_l^{\tau} \nabla T_l) + \rho_l \mathbf{g} \cdot \mathbf{V}_l^{\tau} = 0, \quad (5.1)$$

where  $V_{il}^{\tau 2} = \mathbf{V}_{il}^{\tau} \cdot \mathbf{V}_{il}^{\tau}$  is a scalar. The scalar  $\lambda_l^{\tau}$  in the *Fourier's* law of heat conduction  $\dot{q}_l^{\tau} = -\lambda_l^{\tau} \nabla T_l$  is called isotropic thermal conductivity. There is no doubt about the validity of equation (5.1) for velocities much less than the velocity of light. In contrast to the *Sha et al.* (1984) derivation, a multi-component velocity field is considered here instead of a single component. This will allow for derivation of mixture properties for the field that are strictly consistent with the first law of thermodynamics.

The *specific enthalpy* of each component

$$h_{il}^{\tau} = e_{il}^{\tau} + p_{il}^{\tau} / \rho_{il} \quad (5.2)$$

naturally arises in the second differential term. Historically, the specific enthalpy was introduced as a very convenient variable to describe steady-state processes. We replace the specific internal energy also in the first differential term by using

$$e_{il}^{\tau} = h_{il}^{\tau} - p_{il}^{\tau} / \rho_{il}. \quad (5.3)$$

Performing local volume averaging on Eq. (5.1) as already described in Chapter 1, one obtains

$$\sum_{i=1}^{i_{\max}} \left\{ \left\langle \frac{\partial}{\partial \tau} \left[ \rho_{il} \left( h_{il}^{\tau} + \frac{1}{2} V_{il}^{\tau 2} \right) \right] \right\rangle + \left\langle \nabla \cdot \left[ \rho_{il} \left( h_{il}^{\tau} + \frac{1}{2} V_{il}^{\tau 2} \right) \mathbf{V}_{il}^{\tau} \right] \right\rangle - \left\langle \frac{\partial p_{il}^{\tau}}{\partial \tau} \right\rangle \right\} - \left\langle \nabla \cdot (\mathbf{T}_{\eta,l}^{\tau} \cdot \mathbf{V}_l^{\tau}) + \nabla \cdot (\lambda_l^{\tau} \nabla T_l) \right\rangle + \left\langle \rho_l \mathbf{g} \cdot \mathbf{V}_l^{\tau} \right\rangle = 0. \quad (5.4)$$

The time derivative of the pressure is averaged using Eq. (1.32) (*Leibnitz* rule) as follows:

$$-\left\langle \frac{\partial p_{il}^\tau}{\partial \tau} \right\rangle = -\frac{\partial}{\partial \tau} \left( \gamma_v \alpha_{il} \langle p_{il}^\tau \rangle^{il} \right) + \frac{1}{Vol} \int_{F_{i\sigma} + F_{iw}} p_{il}^\tau \mathbf{V}_{il\sigma}^\tau \cdot \mathbf{n}_l dF. \quad (5.5)$$

As in Chapter 2, the pressure at the interface  $F_{i\sigma}$  is expressed in the form of the sum of bulk averaged pressure, which is independent on the location of the interface, and of the deviation from the bulk pressure, which depends on the location of the surface:

$$p_{il}^\tau = \langle p_{il}^\tau \rangle^{il} + \Delta p_{il\sigma}^\tau \quad (5.6)$$

An analogous separation can be performed at the structure interface  $F_{iw}$

$$p_{il}^\tau = \langle p_{il}^\tau \rangle^{il} + \Delta p_{ilw}^\tau. \quad (5.7)$$

Substituting Eqs. (5.6) and (5.7) into Eq. (5.5), and using Eq. (1.33), the following is obtained

$$\begin{aligned} & \frac{1}{Vol} \int_{F_{i\sigma} + F_{iw}} p_{il}^\tau \mathbf{V}_{il\sigma}^\tau \cdot \mathbf{n}_l dF \\ &= \langle p_{il}^\tau \rangle^{il} \frac{1}{Vol} \int_{F_{i\sigma} + F_{iw}} \mathbf{V}_{il\sigma}^\tau \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_{il\sigma}^\tau \mathbf{V}_{il\sigma}^\tau \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{iw}} \Delta p_{ilw}^\tau \mathbf{V}_{ilw}^\tau \cdot \mathbf{n}_l dF \\ &= \langle p_{il}^\tau \rangle^{il} \frac{\partial}{\partial \tau} (\alpha_{il} \gamma_v) + \frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_{il\sigma}^\tau \mathbf{V}_{il\sigma}^\tau \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{iw}} \Delta p_{ilw}^\tau \mathbf{V}_{ilw}^\tau \cdot \mathbf{n}_l dF. \quad (5.8) \end{aligned}$$

The first term is the power introduced into the velocity field due to a change of the field component volume per unit time and unit control volume (expansion or compression power). Thus, the final result for the averaging of the time derivative of the pressure is

$$\begin{aligned} & -\left\langle \frac{\partial p_{il}^\tau}{\partial \tau} \right\rangle = -\frac{\partial}{\partial \tau} \left( \gamma_v \alpha_{il} \langle p_{il}^\tau \rangle^{il} \right) + \langle p_{il}^\tau \rangle^{il} \frac{\partial}{\partial \tau} (\alpha_{il} \gamma_v) \\ & + \frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_{il\sigma}^\tau \mathbf{V}_{il\sigma}^\tau \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{iw}} \Delta p_{ilw}^\tau \mathbf{V}_{ilw}^\tau \cdot \mathbf{n}_l dF \\ & = -\alpha_{il} \gamma_v \frac{\partial}{\partial \tau} \langle p_{il}^\tau \rangle^{il} + \frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_{il\sigma}^\tau \mathbf{V}_{il\sigma}^\tau \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{iw}} \Delta p_{ilw}^\tau \mathbf{V}_{ilw}^\tau \cdot \mathbf{n}_l dF. \quad (5.9) \end{aligned}$$

Applying Eqs. (1.28), (1.25) and (1.32) to the other terms of Eq. (5.4), and taking into account Eq. (5.9), one obtains

$$\begin{aligned}
 & \left. \sum_{i=1}^{i_{\max}} \left\{ \frac{\partial}{\partial \tau} \left[ \gamma_v \alpha_{il} \left\langle \rho_{il} \left( h_{il}^\tau + \frac{1}{2} V_{il}^{\tau 2} \right) \right\rangle^{il} \right] \right. \right. \\
 & \left. \left. + \nabla \cdot \left[ \gamma \alpha_{il}^e \left\langle \rho_{il} \left( h_{il}^\tau + \frac{1}{2} V_{il}^{\tau 2} \right) \mathbf{V}_{il}^\tau \right\rangle^{ile} \right] - \gamma_v \alpha_{il} \frac{\partial}{\partial \tau} \langle p_{il}^\tau \rangle^{il} \right\} \right. \\
 & - \nabla \cdot \left( \alpha_i^e \gamma \langle \mathbf{T}_{\eta,l}^\tau \cdot \mathbf{V}_l^\tau \rangle^{le} \right) - \nabla \cdot \left( \alpha_i^e \gamma \langle \lambda_l^l \nabla T_l \rangle^{le} \right) + \alpha_i \gamma_v \langle \rho_l (\mathbf{g} \cdot \mathbf{V}_l^\tau) \rangle^l \\
 & + \frac{1}{Vol} \int_{F_{\sigma} + F_{lw}} \left[ \sum_{i=1}^{i_{\max}} \rho_{il} \left( h_{il}^\tau + \frac{1}{2} V_{il}^{\tau 2} \right) (\mathbf{V}_{il}^\tau - \mathbf{V}_{il\sigma}^\tau) \cdot \mathbf{n}_i dF \right] \\
 & + \frac{1}{Vol} \int_{F_{\sigma}} \Delta p_{l\sigma}^\tau \mathbf{V}_{l\sigma}^\tau \cdot \mathbf{n}_i dF + \frac{1}{Vol} \int_{F_{lw}} \Delta p_{lw}^\tau \mathbf{V}_{lw}^\tau \cdot \mathbf{n}_i dF \\
 & - \frac{1}{Vol} \int_{F_{\sigma} + F_{lw}} \mathbf{T}_{\eta,l}^\tau \cdot \mathbf{V}_l^\tau \cdot \mathbf{n}_i dF - \frac{1}{Vol} \int_{F_{\sigma} + F_{lw}} \lambda_l^l \nabla T_l \cdot \mathbf{n}_i dF = 0. \tag{5.10}
 \end{aligned}$$

The  $\delta_l$ -identifier for the dispersed field is now introduced into the heat conduction term, which thus becomes

$$\nabla \cdot \left( \alpha_i^e \gamma \delta_l \langle \lambda_l^l \nabla T_l \rangle^{le} \right). \tag{5.11}$$

The above simply means that there is no heat transfer through heat conduction for dispersed field,  $\delta_l = 0$ , and that heat conduction is taken into account for a continuous field,  $\delta_l = 1$ .

The heat source term

$$\frac{1}{Vol} \int_{F_{\sigma} + F_{lw}} \lambda_l^l \nabla T_l \cdot \mathbf{n}_i dF = \gamma_v \left( \dot{q}_{wl}''' + \sum_{\substack{m=1 \\ m \neq l}}^3 \dot{q}_{m\sigma l}''' \right) = \gamma_v \dot{q}_l''' \tag{5.12}$$

takes into account the sum of the heat power per unit control volume introduced into the velocity field through the interface,  $\gamma_v \dot{q}_{m\sigma l}'''$ , and through the wall,  $\gamma_v \dot{q}_{wl}'''$ .

The integral terms containing the velocity difference  $\mathbf{V}_{il}^\tau - \mathbf{V}_{il\sigma}^\tau$  represent the energy transfer due to (a) evaporation, or (b) condensation, or (c) entrainment, or (d) deposition at the interface  $F_{\sigma}$ , or (e) due to injection or (f) suction through the solid interface  $F_{lw}$ . It is very important to note that  $h_{il}^\tau$  under the interface integral is taken *inside the field l* at the immediate interface neighborhood denoted with  $\sigma$ . The subscript  $M$  is introduced here to designate the non-inert component

within the field.  $n$  used to designate the inert component. The total number of inert components in each field is  $i_{\max} - 1$ .

Consequently

$$h_{Ml}^{\tau\sigma} = h''(T_{12}^{\sigma}) \quad \text{or} \quad h_{Ml}^{\tau\sigma} = h''(T_{13}^{\sigma}) \quad \text{for } l=1, \quad (5.13)$$

$$h_{M2}^{\tau\sigma} = h'(p) \quad \text{for } l=2, \quad (5.14)$$

$$h_{M3}^{\tau\sigma} = h'(p) \quad \text{for } l=3, C_{M3} > 0, \quad (5.15)$$

for all mass flows entering field  $l$ . Mass flows leaving field  $l$  possess the  $l$ -properties. One now assumes that  $\mathbf{V}_{il}^{\tau} = \mathbf{V}_l^{\tau}$  at the  $l$  side of the interface, with these terms then split into non-negative components. It is necessary to distinguish between mass transfer due to change in the state of aggregate, and mass transfer such as entrainment, deposition, etc. resulting from mechanical macroscopic forces.

(a) The mass transfer term at the field-solid interface is decomposed as follows:

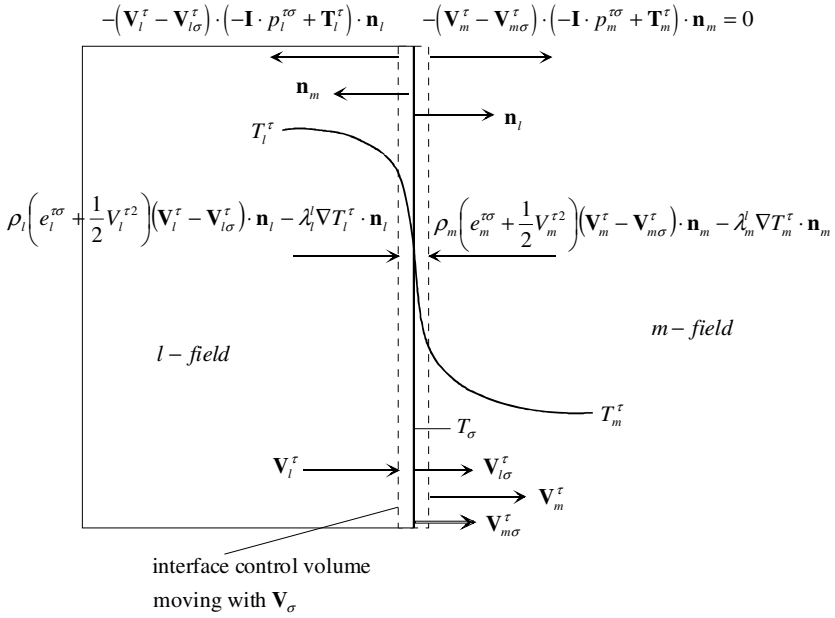
$$\begin{aligned} & -\frac{1}{Vol} \int_{F_w} \left[ \sum_{i=1}^{i_{\max}} \rho_{il} \left( h_{il}^{\tau} + \frac{1}{2} V_{il}^{\tau 2} \right) (\mathbf{V}_{il}^{\tau} - \mathbf{V}_{i\sigma}^{\tau}) \cdot \mathbf{n}_l dF \right] \\ & = \gamma_v \left\{ \sum_{i=1}^{i_{\max}} \left[ \left( \mu_{iwl}^{\tau} \langle h_{iw}^{\tau} \rangle^{iw} - \mu_{ilw}^{\tau} \langle h_{il}^{\tau} \rangle^{il} \right) \right] + \mu_{wl}^{\tau} \frac{1}{2} \langle V_{wl}^{\tau 2} \rangle^{we} - \mu_{lw}^{\tau} \frac{1}{2} \langle V_{lw}^{\tau 2} \rangle^{le} \right\}; \quad (5.16) \end{aligned}$$

(b) The mass transfer term at the field interface is decomposed as follows

$$\begin{aligned} & -\frac{1}{Vol} \int_{F_{\sigma}} \left[ \sum_{i=1}^{i_{\max}} \rho_{il} \left( h_{il}^{\tau} + \frac{1}{2} V_{il}^{\tau 2} \right) (\mathbf{V}_{il}^{\tau} - \mathbf{V}_{i\sigma}^{\tau}) \cdot \mathbf{n}_l dF \right] \\ & = \gamma_v \left\{ \sum_{m=1}^3 \left\{ \left( \mu_{Mml}^{\tau} h_{Ml}^{\tau\sigma} - \mu_{Mlm}^{\tau} \langle h_{Ml}^{\tau} \rangle^{Ml} \right) + \sum_{n=1}^{n_{\max}} \left( \mu_{nml}^{\tau} \langle h_{nm}^{\tau} \rangle^{nm} - \mu_{nlm}^{\tau} \langle h_{nl}^{\tau} \rangle^{nl} \right) \right\} \right. \\ & \quad \left. + \mu_{ml}^{\tau} \frac{1}{2} \langle V_m^{\tau 2} \rangle^{me} - \mu_{lm}^{\tau} \frac{1}{2} \langle V_l^{\tau 2} \rangle^{le} \right\}; \quad (5.17) \end{aligned}$$

**Attention:** The interfacial mass transfer is either convective or is a change of the state of aggregate (evaporation or condensation). Eq. (5.17) is strictly valid for convective interfacial mass transfer. In case of evaporation or condensation the specific enthalpy of the leaving mass flow rate may be different than the volume

averaged field enthalpy resulting in:  $\sum_{m=1}^3 \left[ \left( \mu_{Mml}^{\tau} h_{Ml}^{\tau\sigma} - \mu_{Mlm}^{\tau} h_{Ml}^{\tau\sigma} \right) + \dots \right] \dots$


**Fig. 5.1** Interfacial energy transfer

The sum of Eqs. (5.12), (5.16) and (5.17) gives

$$\begin{aligned}
 q_l^* &= \dot{q}_l^m + \sum_{i=1}^{i_{\max}} \left[ \left( \mu_{iwl}^\tau \langle h_{iw}^\tau \rangle^{iw} - \mu_{ilw}^\tau \langle h_{il}^\tau \rangle^{il} \right) \right] \\
 &+ \sum_{m=1}^3 \left\{ \left( \mu_{Mml}^\tau h_{Ml}^{\tau\sigma} - \mu_{Mlm}^\tau \langle h_{Ml}^\tau \rangle^{Ml} \right) + \sum_{n=1}^{n_{\max}} \left( \mu_{nml}^\tau \langle h_{nm}^\tau \rangle^{nm} - \mu_{nlm}^\tau \langle h_{nl}^\tau \rangle^{nl} \right) \right\} \\
 &+ \mu_{wl}^\tau \frac{1}{2} \langle V_{wl}^{\tau 2} \rangle^{we} - \mu_{lw}^\tau \frac{1}{2} \langle V_{lw}^{\tau 2} \rangle^{le} + \frac{1}{2} \sum_{m=1}^3 \left( \mu_{ml}^\tau \langle V_m^{\tau 2} \rangle^{me} - \mu_{lm}^\tau \langle V_l^{\tau 2} \rangle^{le} \right). \quad (5.18)
 \end{aligned}$$

The *interfacial energy jump condition* is introduced at this point, with interface  $lm$  considered as *non-material*, see Fig. 5.1,

$$\begin{aligned}
 &\rho_l \left( e_l^{\tau\sigma} + \frac{1}{2} V_l^{\tau 2} \right) (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot \mathbf{n}_l - \lambda_l^l \nabla T_l^\tau \cdot \mathbf{n}_l \\
 &+ \rho_m \left( e_m^{\tau\sigma} + \frac{1}{2} V_m^{\tau 2} \right) (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) \cdot \mathbf{n}_m - \lambda_m^l \nabla T_m^\tau \cdot \mathbf{n}_m \\
 &- (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot (-\mathbf{I} \cdot p_l^{\tau\sigma} + \mathbf{T}_l^\tau) \cdot \mathbf{n}_l - (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) \cdot (-\mathbf{I} \cdot p_m^{\tau\sigma} + \mathbf{T}_m^\tau) \cdot \mathbf{n}_m = 0 \quad (5.19)
 \end{aligned}$$

or

$$\left[ \begin{array}{l} \rho_l \left( h_l^{\tau\sigma} + \frac{1}{2} V_l^{\tau 2} \right) (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) - \rho_m \left( h_m^{\tau\sigma} + \frac{1}{2} V_m^{\tau 2} \right) (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) \\ -\lambda_l^l \nabla T_l^\tau + \lambda_m^l \nabla T_m^\tau - (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot \mathbf{T}_l^\tau + (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) \cdot \mathbf{T}_m^\tau \end{array} \right] \cdot \mathbf{n}_l = 0. \quad (5.20)$$

Using Eqs. (1.42b) and (1.42c), Eq. (5.20) can be simplified as follows

$$\left[ \begin{array}{l} \left[ h_l^{\tau,m\sigma} - h_m^{\tau,l\sigma} + \frac{1}{2} (V_l^{\tau 2} - V_m^{\tau 2}) \right] \frac{\rho_m \rho_l}{\rho_l - \rho_m} (\mathbf{V}_m^\tau - \mathbf{V}_l^\tau) \\ -\lambda_l^l \nabla T_l^\tau + \lambda_m^l \nabla T_m^\tau + \frac{\rho_l \rho_m}{\rho_l - \rho_m} (\mathbf{V}_m^\tau - \mathbf{V}_l^\tau) \cdot \left( \frac{\mathbf{T}_m^\tau}{\rho_m} - \frac{\mathbf{T}_l^\tau}{\rho_l} \right) \end{array} \right] \cdot \mathbf{n}_l = 0, \quad (5.21)$$

or

$$\left[ h_l^{\tau,m\sigma} - h_m^{\tau,l\sigma} + \frac{1}{2} (V_l^{\tau 2} - V_m^{\tau 2}) + \frac{\mathbf{T}_m^\tau}{\rho_m} - \frac{\mathbf{T}_l^\tau}{\rho_l} \right] (\rho w)_{lm}^\tau + (\lambda_m^l \nabla T_m^\tau - \lambda_l^l \nabla T_l^\tau) \cdot \mathbf{n}_l = 0.$$

or

$$(\rho w)_{lm}^\tau = - \frac{(\lambda_m^l \nabla T_m^\tau - \lambda_l^l \nabla T_l^\tau) \cdot \mathbf{n}_l}{h_l^{\tau,m\sigma} - h_m^{\tau,l\sigma} + \frac{1}{2} (V_l^{\tau 2} - V_m^{\tau 2}) + \frac{\mathbf{T}_m^\tau}{\rho_m} - \frac{\mathbf{T}_l^\tau}{\rho_l}}.$$

We realize that if there is no mass transfer across the interfacial contact discontinuity heat conduction is the only mechanism transferring energy across. In the simple case of no heat conduction at both sides of the interface and zero stress tensors the energy jump condition simplifies to

$$h_l^{\tau\sigma} - h_m^{\tau\sigma} + \frac{1}{2} (V_l^{\tau 2} - V_m^{\tau 2}) = 0. \quad (5.22)$$

Integrating Eq. (5.20) over the interface inside the control volume and dividing by the control volume we obtain

$$\frac{1}{Vol} \int_{F_{l\sigma} + F_w} \left[ \begin{array}{l} \rho_l \left( h_l^{\tau\sigma} + \frac{1}{2} V_l^{\tau 2} \right) (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) - \rho_m \left( h_m^{\tau\sigma} + \frac{1}{2} V_m^{\tau 2} \right) (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) \\ -\lambda_l^l \nabla T_l^\tau + \lambda_m^l \nabla T_m^\tau - (\mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau) \cdot \mathbf{T}_l^\tau + (\mathbf{V}_m^\tau - \mathbf{V}_{m\sigma}^\tau) \cdot \mathbf{T}_m^\tau \end{array} \right] \cdot \mathbf{n}_l dF = 0. \quad (5.23)$$

In the event of evaporation or condensation this yields

$$\begin{aligned} & \mu_{Ml}^{\tau} \left( h_{Ml}^{\tau\sigma} + \frac{1}{2} V_l^{\tau 2} \right) - \dot{q}_{m\sigma l}^{\prime\prime} - \mu_{Mm}^{\tau} \left( h_{Mm}^{\tau\sigma} + \frac{1}{2} V_m^{\tau 2} \right) - \dot{q}_{l\sigma m}^{\prime\prime} \\ & + \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \left[ -(\mathbf{V}_l^{\tau} - \mathbf{V}_{l\sigma}^{\tau}) \cdot \mathbf{T}_l^{\tau} + (\mathbf{V}_m^{\tau} - \mathbf{V}_{m\sigma}^{\tau}) \cdot \mathbf{T}_m^{\tau} \right] \cdot \mathbf{n}_l dF = 0. \end{aligned} \quad (5.24)$$

Postulating the weighted average

$$\langle \rho_{il} h_{il}^{\tau} \rangle^{il} / \langle \rho_{il} \rangle^{il} = \langle h_{il}^{\tau} \rangle^{il}, \quad (5.25)$$

$$\langle \rho_{il} V_{il}^{\tau 2} \rangle^{ile} / \langle \rho_{il} \rangle^{ile} = \langle V_{il}^{\tau 2} \rangle^{ile}, \quad (5.26)$$

$$\langle \rho_{il} V_{il}^{\tau 2} \mathbf{V}_{il}^{\tau} \rangle^{ile} / \langle \rho_{il} \rangle^{ile} = \langle V_{il}^{\tau 2} \rangle^{ile} \langle \mathbf{V}_{il}^{\tau} \rangle^{ile}, \quad (5.27)$$

one can then write the *conservative form* of the energy conservation equation in the following form

$$\begin{aligned} & \left. \sum_{i=1}^{i_{\max}} \left\{ \frac{\partial}{\partial \tau} \left[ \gamma_v \alpha_{il} \langle \rho_{il} \rangle^{il} \left( \langle h_{il}^{\tau} \rangle^{il} + \frac{1}{2} \langle V_{il}^{\tau 2} \rangle^{ile} \right) \right] \right. \right. \\ & \left. \left. + \nabla \cdot \left[ \gamma \alpha_{il}^e \langle \rho_{il} \rangle^{il} \langle \mathbf{V}_{il}^{\tau} \rangle^{ile} \left( \langle h_{il}^{\tau} \rangle^{il} + \frac{1}{2} \langle V_{il}^{\tau 2} \rangle^{ile} \right) \right] - \gamma_v \alpha_{il} \frac{\partial}{\partial \tau} \langle p_{il}^{\tau} \rangle^{il} \right\} \right. \\ & - \nabla \cdot \left( \alpha_i^e \gamma \langle \mathbf{T}_{\eta, l}^{\tau} \cdot \mathbf{V}_l^{\tau} \rangle^{le} \right) - \nabla \cdot \left( \alpha_i^e \gamma \delta_l \langle \lambda_l' \nabla T_l \rangle^{le} \right) + \alpha_i \gamma_v \langle \rho_l (\mathbf{g} \cdot \mathbf{V}_l^{\tau}) \rangle^l \\ & + \frac{1}{Vol} \int_{F_{l\sigma}} \Delta p_{l\sigma}^{\tau} \mathbf{V}_{l\sigma}^{\tau} \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{lw}} \Delta p_{lw}^{\tau} \mathbf{V}_{lw}^{\tau} \cdot \mathbf{n}_l dF - \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \mathbf{T}_l^{\tau} \cdot \mathbf{V}_l^{\tau} \cdot \mathbf{n}_l dF = \gamma_v q_l^* . \end{aligned} \quad (5.28)$$

Using the chain rule, differentiation can now be performed on the first two terms. Comparing them with the mass conservation equations (1.38) one then obtains the *non-conservative form* of the energy equation:

$$\begin{aligned} & \left. \sum_{i=1}^{i_{\max}} \left\{ \langle \rho_{il} \rangle^{il} \left[ \alpha_{il} \gamma_v \frac{\partial}{\partial \tau} \left( \langle h_{il}^{\tau} \rangle^{il} + \frac{1}{2} \langle V_{il}^{\tau 2} \rangle^{ile} \right) \right. \right. \right. \\ & \left. \left. \left. + \alpha_{il}^e \langle \mathbf{V}_{il}^{\tau} \rangle^{ile} \gamma \cdot \nabla \left( \langle h_{il}^{\tau} \rangle^{il} + \frac{1}{2} \langle V_{il}^{\tau 2} \rangle^{ile} \right) \right] - \gamma_v \alpha_{il} \frac{\partial}{\partial \tau} \langle p_{il}^{\tau} \rangle^{il} \right\} \right. \\ & - \nabla \cdot \left( \alpha_i^e \gamma \langle \mathbf{T}_{\eta, l}^{\tau} \cdot \mathbf{V}_l^{\tau} \rangle^{le} \right) - \nabla \cdot \left( \alpha_i^e \gamma \delta_l \langle \lambda_l' \nabla T_l \rangle^{le} \right) + \alpha_i \gamma_v \langle \rho_l (\mathbf{g} \cdot \mathbf{V}_l^{\tau}) \rangle^l \end{aligned}$$



$$+ \frac{1}{Vol} \int_{F_{i\sigma}} \Delta p_{i\sigma}^{\tau} \mathbf{V}_{i\sigma}^{\tau} \cdot \mathbf{n}_i dF + \frac{1}{Vol} \int_{F_{iw}} \Delta p_{iw}^{\tau} \mathbf{V}_{iw}^{\tau} \cdot \mathbf{n}_i dF - \frac{1}{Vol} \int_{F_{i\sigma} + F_{iw}} \mathbf{T}_i^{\tau} \cdot \mathbf{V}_i^{\tau} \cdot \mathbf{n}_i dF = \gamma_v q_i^{*N} \quad (5.29)$$

where

$$\begin{aligned}
 q_i^{*N} = & \dot{q}_i''' + \sum_{i=1}^{i_{\max}} \mu_{iw}^{\tau} \left( \langle h_{iw}^{\tau} \rangle^{iw} - \langle h_{il}^{\tau} \rangle^{il} \right) \\
 & + \sum_{m=1}^3 \left[ \mu_{Mml}^{\tau} \left( h_{Ml}^{\tau\sigma} - \langle h_{Ml}^{\tau} \rangle^{Ml} \right) + \sum_{n=1}^{n_{\max}} \mu_{nml}^{\tau} \left( \langle h_{nm} \rangle^{nm} - \langle h_{nl}^{\tau} \rangle^{nl} \right) \right] \\
 & + \mu_{wl}^{\tau} \frac{1}{2} \left( \langle V_{wl}^{\tau 2} \rangle^{we} - \langle V_l^{\tau 2} \rangle^{le} \right) - \mu_{lw}^{\tau} \frac{1}{2} \left( \langle V_{lw}^{\tau 2} \rangle^{le} - \langle V_l^{\tau 2} \rangle^{le} \right) \\
 & + \frac{1}{2} \sum_{m=1}^3 \mu_{ml}^{\tau} \left( \langle V_m^{\tau 2} \rangle^{me} - \langle V_l^{\tau 2} \rangle^{le} \right). \quad (5.30)
 \end{aligned}$$

The superscript  $N$  stands here to remember that this RHS belongs to the non-conservative notation of the energy conservation equation.

### 5.3 Dalton and Fick's laws, center of mass mixture velocity, caloric mixture properties

As mentioned in Chapter 1, for a gas mixture it follows from the Dalton's law that  $\alpha_{il} = \alpha_i$ , whereas for mixtures consisting of liquid and macroscopic solid particles  $\alpha_{il} \neq \alpha_i$ . The instantaneous mass concentration of the component  $i$  in  $l$  is defined by Eq. (1.49),

$$C_{il}^{\tau} = \alpha_{il} \langle \rho_{il} \rangle^l / \left( \alpha_l \langle \rho_l \rangle^l \right). \quad (5.31)$$

The center of mass (c. m.) velocity is given by intrinsic surface-averaged field velocity  $\langle \mathbf{V}_l^{\tau} \rangle^{le}$ . Equation (1.50) can be rewritten as

$$\alpha_l \langle \rho_l \rangle^l \langle \mathbf{V}_l^{\tau} \rangle^{le} = \sum_{i=1}^{i_{\max}} \alpha_{il} \langle \rho_{il} \rangle^l \langle \mathbf{V}_{il}^{\tau} \rangle^{le} = \alpha_l \langle \rho_l \rangle^l \sum_{i=1}^{i_{\max}} C_{il}^{\tau} \langle \mathbf{V}_{il}^{\tau} \rangle^{le}. \quad (5.32)$$

Consequently

$$\langle \mathbf{V}_l^{\tau} \rangle^{le} = \sum_{i=1}^{i_{\max}} C_{il}^{\tau} \langle \mathbf{V}_{il}^{\tau} \rangle^{le}. \quad (5.33)$$

As for the derivation of the mass conservation equation, it is convenient for description of transport of the microscopic component  $il$  in the velocity field  $l$  to replace the velocity component  $\langle \mathbf{V}_{il}^\tau \rangle^{le}$  by the sum of the center of mass velocity for the particular field  $\langle \mathbf{V}_l^\tau \rangle^{le}$  and the deviation from the c. m. velocity or the so called *diffusion velocity* of the inert component  $\delta_i \langle \mathbf{V}_l^\tau \rangle^{le}$ , this yielding the following

$$\langle \mathbf{V}_{il}^\tau \rangle^{le} = \langle \mathbf{V}_l^\tau \rangle^{le} + \left( \delta \langle \mathbf{V}_l^\tau \rangle^{le} \right)_i. \quad (5.34)$$

*Fick* (1855) noticed that the mass flow rate of the inert component with respect to the total mass flow rate of the continuous mixture including the inert component is proportional to the gradient of the concentration of this inert component

$$\alpha_{il} \langle \rho_{il} \rangle^l \left( \delta \langle \mathbf{V}_l^\tau \rangle^{le} \right)_i = -\alpha_i^e \delta_i D_{il}^l \nabla \langle \rho_{il} \rangle^{il} = -\alpha_i \langle \rho_l \rangle^l \delta_i D_{il}^l \nabla \langle C_{il}^\tau \rangle^{il}, \quad (5.35)$$

or divided by the component density

$$\alpha_{il}^e \delta \mathbf{V}_{il}^\tau = -\alpha_i^e \delta_i D_{il}^l \frac{1}{\langle C_{il}^\tau \rangle^{il}} \nabla \langle C_{il}^\tau \rangle^{il} = -\alpha_i^e \delta_i D_{il}^l \nabla \ln \langle C_{il}^\tau \rangle^{il}. \quad (5.36)$$

The coefficient of proportionality,  $D_{il}^l$ , is known as the *isotropic coefficient of molecular diffusion*. The diffusion mass flow rate is directed from regions with higher concentration to regions with lower concentration, with this reflected by the minus sign in the assumption made by *Fick* (which has subsequently come to be known as *Fick's law*), because many processes in nature and industrial equipment are successfully described mathematically by the above approach – so called diffusion processes. Here  $\delta_i = 1$  for a continuous field and  $\delta_i = 0$  for a disperse field. Molecular diffusion has microscopic character, as it is caused by molecular interactions. The special theoretical treatment and the experimental experience of how to determine the molecular diffusion constant in multi-component mixtures, is a science in its own right. This topic is beyond the scope of this chapter. In this context, it should merely be only noted that in line with the thermodynamics of irreversible processes, the thermal diffusivity and the diffusion coefficients influence each other. The interested reader can find useful information in *Reid et al.* (1982).

One should keep in mind that there is no molecular net mass diffusively transported across a cross section perpendicular to the strongest concentration gradient. This is mathematically expressed as follows

$$\sum_{i=1}^{i_{\max}} \left( D_{il}^l \nabla \langle C_{il}^\tau \rangle^{il} \right) = 0, \quad (5.37)$$

or

$$D_{il}' \nabla \langle C_{il}^\tau \rangle^{il} = - \sum_{i=2}^{i_{\max}} \left( D_{il}' \nabla \langle C_{il}^\tau \rangle^{il} \right). \quad (5.38)$$

Therefore, the diffusion can be considered as volumetric replacement of groups of molecules of one specie with group of molecules of other specie. The compensating volumetric flow is called sometimes *Stefan's flow*, *Stefan* (1874).

The caloric mixture properties naturally arise after summing all the energy conservation equations in conservative form. These are defined as follows

$$\langle \varphi_l^\tau \rangle^l = \frac{\sum_{i=1}^{i_{\max}} \alpha_{il} \langle \rho_{il} \rangle^{il} \langle \varphi_{il}^\tau \rangle^{il}}{\alpha_l \langle \rho_l \rangle^l} = \sum_{i=1}^{i_{\max}} \left( \langle C_{il} \rangle^{il} \langle \varphi_{il}^\tau \rangle^{il} \right), \quad (5.39)$$

where

$$\varphi = h, e, s. \quad (5.40)$$

## 5.4 Enthalpy equation

For practical applications it is extremely convenient to simplify the energy equation in a way that the mechanical energy terms disappear. The resulting equation is called the *enthalpy equation*. The enthalpy equation will now be derived.

The non-conservative form of the momentum equation (2.31) using Eq. (1.38) is

$$\begin{aligned} & \langle \rho_l \rangle^l \left[ \alpha_l \gamma_v \frac{\partial}{\partial \tau} \langle \mathbf{V}_l^\tau \rangle^{le} + \alpha_l^e \langle \mathbf{V}_l^\tau \rangle^{le} \gamma \mathcal{N} \cdot \langle \mathbf{V}_l^\tau \rangle^{le} \right] - \nabla \cdot \left( \alpha_l^e \gamma \langle \mathbf{T}_{\eta,l}^\tau \rangle^{le} \right) + \alpha_l^e \gamma \mathcal{N} \cdot \langle p_l^\tau \rangle^{le} \\ & + \frac{1}{Vol} \int_{F_{l\sigma}} \Delta p_{l\sigma}^\tau \cdot \mathbf{n}_{l\sigma} dF + \alpha_l \gamma_v \langle \rho_l \rangle^l \mathbf{g} + \frac{1}{Vol} \int_{F_{lw}} \Delta p_{lw}^\tau \cdot \mathbf{n}_{l\sigma} dF - \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \mathbf{T}_l^\tau \cdot \mathbf{n}_l dF \\ & = \gamma_v \left\{ \mu_{wl}^\tau \left( \langle \mathbf{V}_{wl}^\tau \rangle^{we} - \langle \mathbf{V}_l^\tau \rangle^{le} \right) - \mu_{lw}^\tau \left( \langle \mathbf{V}_{lw}^\tau \rangle^{we} - \langle \mathbf{V}_l^\tau \rangle^{le} \right) + \sum_{m=1}^3 \mu_{ml}^\tau \left( \langle \mathbf{V}_m^\tau \rangle^{me} - \langle \mathbf{V}_l^\tau \rangle^{le} \right) \right\}. \end{aligned} \quad (5.41)$$

Here, Eqs. (5.6) and (5.7) are used, and the integral containing  $\langle p_l^\tau \rangle^{le}$  is rearranged using Eq. (1.29). The scalar product of Eq. (5.41) is constructed with the field velocity. The result is a scalar expressing the mechanical energy balance. Subtracting this result from the energy equation and bearing in mind that

$$\frac{1}{2} \left( \langle V_m^{\tau 2} \rangle^{me} - \langle V_l^{\tau 2} \rangle^{le} \right) - \langle \mathbf{V}_l^\tau \rangle^{le} \cdot \left( \langle \mathbf{V}_m^\tau \rangle^{me} - \langle \mathbf{V}_l^\tau \rangle^{le} \right) = \frac{1}{2} \left( \langle V_m^\tau \rangle^{me} - \langle \mathbf{V}_l^\tau \rangle^{le} \right)^2, \quad (5.42)$$

and

$$\begin{aligned}
& \mu_{wl}^{\tau} \frac{1}{2} \left( \langle V_{wl}^{\tau 2} \rangle^{we} - \langle V_l^{\tau 2} \rangle^{le} \right) - \mu_{lw}^{\tau} \frac{1}{2} \left( \langle V_{lw}^{\tau 2} \rangle^{we} - \langle V_l^{\tau 2} \rangle^{le} \right) \\
& - \langle V_l^{\tau} \rangle^{le} \left[ \mu_{wl}^{\tau} \left( \langle V_{wl}^{\tau} \rangle^{we} - \langle V_l^{\tau} \rangle^{le} \right) - \mu_{lw}^{\tau} \left( \langle V_{lw}^{\tau} \rangle^{we} - \langle V_l^{\tau} \rangle^{le} \right) \right] \\
& = \mu_{wl}^{\tau} \frac{1}{2} \left( \langle V_{wl}^{\tau} \rangle^{we} - \langle V_l^{\tau} \rangle^{le} \right)^2 - \mu_{lw}^{\tau} \frac{1}{2} \left( \langle V_{lw}^{\tau} \rangle^{we} - \langle V_l^{\tau} \rangle^{le} \right)^2, \tag{5.43}
\end{aligned}$$

one then obtains the *non-conservative form of the enthalpy equation*:

$$\begin{aligned}
& \left. \sum_{i=1}^{i_{\max}} \left\{ \begin{aligned} & \langle \rho_{il} \rangle^{il} \left[ \alpha_{il} \gamma_v \frac{\partial}{\partial \tau} \langle h_{il}^{\tau} \rangle^{il} + \alpha_{il}^e \langle V_{il}^{\tau} \rangle^{ile} \gamma \cdot \nabla \langle h_{il}^{\tau} \rangle^{il} \right] \\ & - \left[ \alpha_{il} \gamma_v \frac{\partial}{\partial \tau} \langle p_{il}^{\tau} \rangle^{il} + \alpha_{il}^e \langle V_{il}^{\tau} \rangle^{ile} \gamma \cdot \nabla \langle p_{il}^{\tau} \rangle^{il} \right] \end{aligned} \right\} \right. \\
& \left. - \nabla \cdot \left( \alpha_l^e \gamma \langle \lambda_l^{\tau} \nabla T_l \rangle^{le} \right) = \gamma_v \left( \alpha_l \rho_l P_{\eta,l}^{\tau} + q_l^{\tau N} \right) + E_l^{*\tau}. \tag{5.44}
\end{aligned}$$

where

$$\begin{aligned}
q_l^{\tau N} & = \dot{q}_l^{\tau} + \sum_{i=1}^{i_{\max}} \mu_{iwl}^{\tau} \left( \langle h_{iw}^{\tau} \rangle^{iw} - \langle h_{il}^{\tau} \rangle^{il} \right) \\
& + \sum_{m=1}^3 \left[ \mu_{Mml}^{\tau} \left( h_{Ml}^{\tau\sigma} - \langle h_{Ml}^{\tau} \rangle^{Ml} \right) + \sum_{n=1}^{n_{\max}} \mu_{nml}^{\tau} \left( \langle h_{nm} \rangle^{nm} - \langle h_{nl}^{\tau} \rangle^{nl} \right) \right] \\
& - \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \mu_{Mlm}^{\tau} \left( h_{Ml}^{\tau\sigma} - \langle h_{Ml}^{\tau} \rangle^{Ml} \right) + \sum_{n=1}^{n_{\max}} \mu_{nlm}^{\tau} \left( \langle h_{nl} \rangle^{nl} - \langle h_{nl}^{\tau} \rangle^{nl} \right) \right] \\
& + \mu_{wl}^{\tau} \frac{1}{2} \left( \langle V_{wl}^{\tau} \rangle^{we} - \langle V_l^{\tau} \rangle^{le} \right)^2 - \mu_{lw}^{\tau} \frac{1}{2} \left( \langle V_{lw}^{\tau} \rangle^{le} - \langle V_l^{\tau} \rangle^{le} \right)^2 \\
& + \frac{1}{2} \sum_{m=1}^3 \mu_{ml}^{\tau} \left( \langle V_m^{\tau} \rangle^{me} - \langle V_l^{\tau} \rangle^{le} \right)^2. \tag{5.45}
\end{aligned}$$

The corresponding conservative form is

$$\left. \begin{aligned} & \sum_{i=1}^{i_{\max}} \left\{ \frac{\partial}{\partial \tau} \left( \alpha_{il} \langle \rho_{il} \rangle^{il} \langle h_{il}^\tau \rangle^{il} \gamma_v \right) + \nabla \cdot \left( \alpha_{il}^e \langle \rho_{il} \rangle^{il} \langle \mathbf{V}_{il}^\tau \rangle^{ile} \langle h_{il}^\tau \rangle^{il} \gamma \right) \right\} \\ & - \left[ \alpha_{il} \gamma_v \frac{\partial}{\partial \tau} \langle p_{il}^\tau \rangle^{il} + \alpha_{il}^e \langle \mathbf{V}_{il}^\tau \rangle^{ile} \gamma \cdot \nabla \langle p_{il}^\tau \rangle^{il} \right] \end{aligned} \right\}$$

$$-\nabla \cdot \left( \alpha_l^e \gamma \delta_l \langle \lambda_l' \nabla T_l \rangle^{le} \right) = \gamma_v \left( \alpha_l \rho_l P_{\eta,l}^\tau + q_l \right) + E_l^{*\tau}, \quad (5.46)$$

where

$$\begin{aligned} q_l &= \dot{q}_l^m + \sum_{i=1}^{i_{\max}} \left( \mu_{iwl}^\tau \langle h_{iw}^\tau \rangle^{iw} - \mu_{dlw}^\tau \langle h_{il}^\tau \rangle^{il} \right) \\ &+ \sum_{m=1}^3 \left[ \left( \mu_{Mml}^\tau h_{Ml}^{\tau\sigma} - \mu_{Mlm}^\tau \langle h_{Ml}^\tau \rangle^{Ml} \right) + \sum_{n=1}^{n_{\max}} \left( \mu_{nml}^\tau \langle h_{nm}^\tau \rangle^{nm} - \mu_{nlm}^\tau \langle h_{il}^\tau \rangle^{nl} \right) \right] \\ &+ \mu_{wl}^\tau \frac{1}{2} \left( \langle \mathbf{V}_{wl}^\tau \rangle^{we} - \langle \mathbf{V}_l^\tau \rangle^{le} \right)^2 - \mu_{lw}^\tau \frac{1}{2} \left( \langle \mathbf{V}_{lw}^\tau \rangle^{le} - \langle \mathbf{V}_l^\tau \rangle^{le} \right)^2 \\ &+ \frac{1}{2} \sum_{m=1}^3 \mu_{ml}^\tau \left( \langle \mathbf{V}_m^\tau \rangle^{me} - \langle \mathbf{V}_l^\tau \rangle^{le} \right)^2 \equiv q_l^{\tau N} + \sum_{i=1}^{i_{\max}} \mu_{il}^\tau \langle h_{il} \rangle^{il}. \end{aligned} \quad (5.47)$$

Here

$$\begin{aligned} \gamma_v \alpha_l \rho_l P_{\eta,l}^\tau &= \nabla \cdot \left( \alpha_l^e \gamma \langle \mathbf{T}_{\eta,l}^\tau \cdot \mathbf{V}_l^\tau \rangle^{le} \right) - \langle \mathbf{V}_l^\tau \rangle^{le} \cdot \nabla \cdot \left( \alpha_l^e \gamma \langle \mathbf{T}_{\eta,l}^\tau \rangle^{le} \right) \\ &= \alpha_l^e \gamma \cdot \langle \mathbf{T}_{\eta,l}^\tau : \nabla \cdot \mathbf{V}_l^\tau \rangle^{le}, \end{aligned} \quad (5.48)$$

is the *irreversible bulk viscous dissipation*. The term,

$$\begin{aligned} E_l^{*\tau} &= \frac{1}{Vol} \int_{F_{l\sigma}} \Delta p_{l\sigma}^\tau \left( \langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_{l\sigma}^\tau \right) \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{lw}} \Delta p_{lw}^\tau \left( \langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_{lw}^\tau \right) \cdot \mathbf{n}_l dF \\ &- \frac{1}{Vol} \int_{F_{l\sigma} + F_{lw}} \mathbf{T}_l^\tau \cdot \left( \langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_l^\tau \right) \cdot \mathbf{n}_l dF = \frac{1}{Vol} \int_{F_{l\sigma}} \frac{\Delta p_{l\sigma}^\tau}{\rho_l} \rho_l \left( \mathbf{V}_l^\tau - \mathbf{V}_{l\sigma}^\tau \right) \cdot \mathbf{n}_l dF \\ &+ \frac{1}{Vol} \int_{F_{lw}} \frac{\Delta p_{lw}^\tau}{\rho_l} \rho_l \left( \mathbf{V}_l^\tau - \mathbf{V}_{lw}^\tau \right) \cdot \mathbf{n}_l dF - \frac{1}{Vol} \int_{F_{l\sigma}} \left( -\Delta p_{l\sigma}^\tau + \mathbf{T}_l^\tau \right) \cdot \left( \langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_l^\tau \right) \cdot \mathbf{n}_l dF \\ &- \frac{1}{Vol} \int_{F_{lw}} \left( -\Delta p_{lw}^\tau + \mathbf{T}_l^\tau \right) \cdot \left( \langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_l^\tau \right) \cdot \mathbf{n}_l dF \end{aligned} \quad (5.49)$$

is the irreversible power dissipation caused by the interface mass transfer between two regions with different velocities. A good approximation of the last two terms is obtained if one assumes

$$\mathbf{V}_l^\tau \approx \langle \mathbf{V}_m^\tau \rangle^{me} \quad \text{at } F_{l\sigma} \quad (5.50)$$

and

$$\mathbf{V}_l^\tau = 0 \quad \text{at } F_{lw} \quad (5.51)$$

(non-slip condition), namely

$$\begin{aligned} & -\frac{1}{Vol} \int_{F_{l\sigma}} (-\Delta p_{l\sigma}^\tau + \mathbf{T}_l^\tau) \cdot (\langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_l^\tau) \cdot \mathbf{n}_l dF \\ & \approx -(\langle \mathbf{V}_l^\tau \rangle^{le} - \langle \mathbf{V}_m^\tau \rangle^{me}) \frac{1}{Vol} \int_{F_{l\sigma}} (-\Delta p_{l\sigma}^\tau + \mathbf{T}_l^\tau) \cdot \mathbf{n}_l dF \end{aligned} \quad (5.52)$$

and

$$-\frac{1}{Vol} \int_{F_{lw}} (-\Delta p_{lw}^\tau + \mathbf{T}_l^\tau) \cdot (\langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_l^\tau) \cdot \mathbf{n}_l dF \approx -\langle \mathbf{V}_l^\tau \rangle^{le} \frac{1}{Vol} \int_{F_{lw}} (-\Delta p_{lw}^\tau + \mathbf{T}_l^\tau) \cdot \mathbf{n}_l dF. \quad (5.53)$$

The order of magnitude of the first two terms in Eq. (5.49) is

$$\begin{aligned} & \frac{1}{Vol} \int_{F_{l\sigma}} \Delta p_{l\sigma}^\tau (\langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_l^\tau) \cdot \mathbf{n}_l dF + \frac{1}{Vol} \int_{F_{lw}} \Delta p_{lw}^\tau (\langle \mathbf{V}_l^\tau \rangle^{le} - \mathbf{V}_l^\tau) \cdot \mathbf{n}_l dF \\ & \approx \sum_{m=1}^{3,w} \frac{\overline{\Delta p_{lm\sigma}^\tau}}{\rho_l} (\mu_{ml} - \mu_{lm}), \end{aligned} \quad (5.54)$$

where  $\overline{\Delta p_{lm\sigma}^\tau} / \rho_l$  is an averaged pressure difference between the bulk pressure and the boundary layer pressure inside the velocity field  $l$ .

Performing the summation in Eq. (5.46), using the *Dalton's* law, substituting the *Fick's* laws in the thus obtained equation, and applying the definitions given in Chapter 3 yields

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \langle \rho_l^l \rangle \langle h_l^\tau \rangle^l \gamma_v) + \nabla \cdot (\alpha_l^e \langle \rho_l^l \rangle \langle \mathbf{V}_l^\tau \rangle^{le} \langle h_l^\tau \rangle^l \gamma) \\ & - \left[ \alpha_l \gamma_v \frac{\partial}{\partial \tau} \langle p_l^\tau \rangle^l + \alpha_l^e \langle \mathbf{V}_l^\tau \rangle^{le} \gamma \cdot \nabla \langle p_l^\tau \rangle^l \right] \end{aligned}$$

$$\begin{aligned}
& -\nabla \cdot \left\{ \alpha_i^e \gamma \delta_i \left[ \langle \lambda_i^l \nabla T_l \rangle^{le} + \langle \rho_l \rangle^l \sum_{i=1}^{i_{\max}} \left( \langle h_{il}^\tau \rangle^{il} D_{il}^l \nabla \langle C_{il}^\tau \rangle^{il} \right) \right] \right\} \\
& + \delta_i \left( \alpha_i^e \bar{\gamma} \langle p_i^\tau \rangle^l \right) \cdot \sum_{i=1}^{i_{\max}} D_{il}^l \nabla \ln \langle C_{il}^\tau \rangle^{il} = \gamma_v \left( \alpha_i \rho_l \varepsilon_{\eta,l}^\tau + q_l \right) + E_i^{*\tau} . \quad (5.55)
\end{aligned}$$

The non-conservative form of the equation (5.55) is readily obtained by differentiating the first two terms and comparing them with the field mass conservation equation. The result is

$$\begin{aligned}
& \alpha_i \langle \rho_l \rangle^l \gamma_v \frac{\partial \langle h_i^\tau \rangle^l}{\partial \tau} + \alpha_i^e \langle \rho_l \rangle^l \langle \mathbf{V}_i^\tau \rangle^{le} \bar{\gamma} \langle h_i^\tau \rangle^l \\
& - \left[ \alpha_i \gamma_v \frac{\partial \langle p_i^\tau \rangle^l}{\partial \tau} + \alpha_i^e \langle \mathbf{V}_i^\tau \rangle^{le} \gamma \cdot \nabla \langle p_i^\tau \rangle^l \right] \\
& - \nabla \cdot \left\{ \alpha_i^e \gamma \delta_i \left[ \langle \lambda_i^l \nabla T_l \rangle^{le} + \sum_{i=1}^{i_{\max}} \left( \langle \rho_{il} \rangle^{il} \langle h_{il}^\tau \rangle^{il} D_{il}^l \nabla \ln \langle C_{il}^\tau \rangle^{il} \right) \right] \right\} \\
& + \delta_i \left( \alpha_i^e \bar{\gamma} \langle p_i^\tau \rangle^l \right) \cdot \sum_{i=1}^{i_{\max}} D_{il}^l \nabla \ln \langle C_{il}^\tau \rangle^{il} = \gamma_v \left( \alpha_i \rho_l \varepsilon_{\eta,l}^\tau + q_l - \mu_i^\tau \langle h_i^\tau \rangle^l \right) + E_i^{*\tau} . \quad (5.56)
\end{aligned}$$

Remember that the sum of all  $C$ 's inside the velocity field is equal to one so that one of the concentrations depends on all the others.

## 5.5 Internal energy equation

Engineers sometimes have in their personal library approximations for the state variables and transport properties in terms of the specific internal energy. For practical use in this case, Eq. (5.46) can be rewritten in terms of the specific internal energy:

$$\sum_{i=1}^{i_{\max}} \left\{ \begin{aligned} & \frac{\partial}{\partial \tau} \left( \alpha_{il} \langle \rho_{il} \rangle^{il} \langle e_{il}^\tau \rangle^{il} \gamma_v \right) + \nabla \cdot \left( \alpha_{il}^e \langle \rho_{il} \rangle^{il} \langle \mathbf{V}_{il}^\tau \rangle^{ile} \langle e_{il}^\tau \rangle^{il} \gamma \right) \\ & - \langle p_{il}^\tau \rangle^{il} \left[ \frac{\partial}{\partial \tau} (\alpha_{il} \gamma_v) + \nabla \cdot \left( \alpha_{il}^e \langle \mathbf{V}_{il}^\tau \rangle^{ile} \gamma \right) \right] \end{aligned} \right\}$$

$$-\nabla \cdot \left( \alpha_i^e \gamma \delta_i \langle \lambda_i^l \nabla T_i \rangle^{le} \right) = \gamma_v \left( \alpha_i \rho_i P_{\eta,l}^\tau + q_l \right) + E_i^{*\tau}, \quad (5.57)$$

or using the definitions introduced in Chapter 3

$$\begin{aligned} & \frac{\partial}{\partial \tau} \left( \alpha_i \langle \rho_i \rangle^l \langle e_i^\tau \rangle \gamma_v \right) + \nabla \cdot \left( \alpha_i^e \langle \rho_i \rangle^l \langle \mathbf{V}_i^\tau \rangle^{le} \langle e_i^\tau \rangle \gamma \right) \\ & - \langle p_i^\tau \rangle^l \left[ \frac{\partial}{\partial \tau} (\alpha_i \gamma_v) + \nabla \cdot \left( \alpha_i^e \langle \mathbf{V}_i^\tau \rangle^{le} \gamma \right) \right] \\ & - \nabla \cdot \left\{ \alpha_i^e \gamma \delta_i \left[ \langle \lambda_i^l \nabla T_i \rangle^{le} + \langle \rho_i \rangle^l \sum_{i=1}^{i_{\max}} \left( \langle e_{ii}^\tau \rangle^{il} D_{ii}^l \nabla \langle C_{ii}^\tau \rangle^{il} \right) \right] \right\} \\ & + \delta_i \sum_{i=1}^{i_{\max}} \left[ \langle p_{ii}^\tau \rangle^{il} \nabla \left( \alpha_i^e D_{ii}^l \gamma \mathcal{N} \ln \langle C_{ii}^\tau \rangle^{il} \right) \right] = \gamma_v \left( \alpha_i \rho_i \varepsilon_{\eta,l}^\tau + q_l \right) + E_i^{*\tau}. \quad (5.58) \end{aligned}$$

## 5.6 Entropy equation

The basic idea of the *Legendre* transformation is briefly introduced here: If  $f$  is a function of  $n$  variables,  $x_1, x_2, \dots, x_n$ ,  $f = f(x_1, x_2, \dots, x_n)$ , and for  $j$  of them the partial derivatives  $\partial f / \partial x_i$  are known, the *Legendre* transformation  $z$  is defined by

$$z = f - \sum_{i=1}^j \frac{\partial f}{\partial x_i} x_i = z \left( \frac{\partial f}{\partial x_1}, \frac{\partial f}{\partial x_2}, \dots, \frac{\partial f}{\partial x_j}, x_{j+1}, \dots, x_n \right),$$

for which the differential is  $dz$  is defined as follows

$$dz = \sum_{i=j+1}^n \frac{\partial f}{\partial x_i} dx_i - \sum_{i=1}^j x_i d \frac{\partial f}{\partial x_i}.$$

So if the specific volume is a function of the specific entropy and internal energy  $v = v(s, e)$  the specific internal energy  $e = e(s, v)$  is its *Legendre* transformation.

The derivatives of  $e$  are measurable variables  $(\partial e / \partial s)_v = T$  and  $(\partial e / \partial v)_s = -p$ . The differential  $de$  is then  $de = Tds - pdv$  or

$$Tds = de + pdv. \quad (5.59)$$



This equation is called *Gibbs* equation (1878). The thermodynamic definitions of temperature  $T$  (absolute temperature - *Kelvin* 1848) and of the thermodynamic pressure  $p$  are used here. This is a differential equation of state which is extremely important in the thermodynamic theory of compressible substances. It relates the difference in entropy between any two infinitesimally separated states to the infinitesimal differences in internal energy and volume between those states. Note that the *Gibbs* equation is valid only if  $s$  is a *smooth* function of  $e$  and  $v$ , i.e. where the differentials  $de$  and  $dv$  are uniquely defined (smooth equation of state), see in *Gibbs* (1892). Using the definition of enthalpy (a mixture of thermodynamic and mechanical properties)

$$h = e + pv \quad (5.60)$$

the *Gibbs* equation can be written as

$$T \rho ds = \rho dh - dp. \quad (5.61)$$

After substituting for  $\langle \rho_{il} \rangle^{il} d \langle h_{il}^\tau \rangle^{il} - d \langle p_{il}^\tau \rangle^{il}$  in Eq. (5.44) with the *Gibbs* definitions of the *specific entropy* of the corresponding components,

$$\langle \rho_{il} \rangle^{il} d \langle h_{il}^\tau \rangle^{il} - d \langle p_{il}^\tau \rangle^{il} = T_l \langle \rho_{il} \rangle^{il} d \langle s_{il}^\tau \rangle^{il}, \quad (5.61)$$

one obtains

$$T_l \sum_{i=1}^{i_{\max}} \left\{ \langle \rho_{il} \rangle^{il} \left[ \alpha_{il} \gamma_v \frac{\partial}{\partial \tau} \langle s_{il}^\tau \rangle^{il} + \alpha_{il}^e \langle \mathbf{V}_{il}^\tau \rangle^{ile} \gamma \cdot \nabla \langle s_{il}^\tau \rangle^{il} \right] \right\} \\ - \nabla \cdot \left( \alpha_l^e \gamma \langle \lambda_l^l \nabla T_l \rangle^{le} \right) = \alpha_l^e \gamma \cdot \langle \mathbf{T}_{\eta,l}^\tau : \nabla \cdot \mathbf{V}_l^\tau \rangle^{le} + E_l^{*\tau} + \gamma_v q_l^N. \quad (5.62)$$

The following rearrangements are performed to give the conservative form of Eq. (5.62). Each of the mass conservation equations is multiplied by the corresponding component-specific entropy and field temperature. All the mass conservation equations are then added to Eq. (5.62). The differential terms are lumped together using the reverse chain rule of differentiation. The resulting *conservative form* is:

$$T_l \left\{ \frac{\partial}{\partial \tau} \left[ \gamma_v \sum_{i=1}^{i_{\max}} \left( \alpha_{il} \langle \rho_{il} \rangle^{il} \langle s_{il}^\tau \rangle^{il} \right) \right] + \nabla \cdot \left[ \gamma \sum_{i=1}^{i_{\max}} \left( \alpha_{il}^e \langle \rho_{il} \rangle^{il} \langle \mathbf{V}_{il}^\tau \rangle^{ile} \langle s_{il}^\tau \rangle^{il} \right) \right] \right\} \\ - \nabla \cdot \left( \alpha_l^e \gamma \delta_l \langle \lambda_l^l \nabla T_l \rangle^{le} \right) = \gamma_v \left( \alpha_l \rho_l P_{\eta,l}^\tau + q_l^N + T_l \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml}^\tau - \mu_{ilm}^\tau) \langle s_{il}^\tau \rangle^{il} \right) + E_l^{*\tau}. \quad (5.63)$$

After replacing the instantaneous values of the component velocities with the sums of the c. m. field velocities plus the deviations from the c. m. velocities,

Eq. (5.34), and applying *Fick's law*, Eq. (5.36) and keeping in mind that there is no molecular net mass diffusively transported across a cross section perpendicular to the strongest concentration gradient, Eqs. (5.37, 5.38), defining the *specific mixture entropy of the velocity field*,  $\langle s_i^\tau \rangle^l$ , by Eq. (5.39), introducing the *Prandtl number*

$$\text{Pr}_i^l = \rho_l c_{pl} v_i^l / \lambda_i^l, \quad (5.64)$$

and dividing by  $T_l$ , the following form for the local instantaneous entropy equations is obtained:

$$\begin{aligned} & \frac{\partial}{\partial \tau} \left( \alpha_l \langle \rho_l \rangle^l \langle s_i^\tau \rangle^l \gamma_v \right) + \nabla \cdot \left( \alpha_l^e \langle \rho_l \rangle^l \langle s_i^\tau \rangle^l \langle \mathbf{V}_l^\tau \rangle^{le} \gamma \right) \\ & - \nabla \cdot \left[ \alpha_l^e \delta_l \langle \rho_l \rangle^l \gamma \left( \sum_{i=2}^{i_{\max}} \left( \langle s_{il}^\tau \rangle^{il} - \langle s_{il}^\tau \rangle^{ll} \right) D_{il}^l \nabla \langle C_{il}^\tau \rangle^{il} \right) \right] \\ & - \frac{1}{T_l} \nabla \cdot \left( \alpha_l^e \gamma \delta_l \left\langle \rho_l c_{pl} \frac{v_i^l}{\text{Pr}_i^l} \nabla T_l \right\rangle^{le} \right) \\ & = \gamma_v \left[ \left( \alpha_l \rho_l \varepsilon_{\eta,l}^\tau + q_l^N \right) / T_l + \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} \left( \mu_{iml}^\tau - \mu_{ilm}^\tau \right) \langle s_{il}^\tau \rangle^{il} \right] + E_l^{*\tau}. \end{aligned} \quad (5.65)$$

Observe the form of the RHS of Eq. (5.65).

$$\begin{aligned} & \left( \alpha_l \rho_l \varepsilon_{\eta,l}^\tau + q_l^N \right) / T_l + \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} \left( \mu_{iml}^\tau - \mu_{ilm}^\tau \right) \langle s_{il}^\tau \rangle^{il} \\ & = \frac{\dot{q}_l^m}{T_l} + \sum_{i=1}^{i_{\max}} \left[ \mu_{iwl}^\tau \left( \frac{\langle h_{iw}^\tau \rangle^{iw} - \langle h_{il}^\tau \rangle^{il}}{T_l} + \langle s_{il}^\tau \rangle^{il} \right) + \mu_{ilw}^\tau \left( \frac{\langle h_{il}^\tau \rangle^{il} - \langle h_{il}^\tau \rangle^{il}}{T_l} + \langle s_{il}^\tau \rangle^{il} \right) \right] \\ & + \sum_{m=1}^3 \left[ \mu_{Mml}^\tau \left( \frac{h_{Ml}^{\tau\sigma} - \langle h_{Ml}^\tau \rangle^{Ml}}{T_l} + \langle s_{Ml}^\tau \rangle^{Ml} \right) + \sum_{n=1}^{n_{\max}} \mu_{nml}^\tau \left( \frac{\langle h_{nm} \rangle^{nm} - \langle h_{nl}^\tau \rangle^{nl}}{T_l} + \langle s_{nl}^\tau \rangle^{nl} \right) \right] \end{aligned}$$

$$\begin{aligned}
& - \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \mu_{Mlm}^{\tau} \left( \frac{h_{Ml}^{\tau\sigma} - \langle h_{Ml}^{\tau} \rangle^{Ml}}{T_l} + \langle s_{Ml}^{\tau} \rangle^{Ml} \right) + \sum_{n=1}^{n_{\max}} \mu_{nlm}^{\tau} \left( \frac{\langle h_{nl} \rangle^{nl} - \langle h_{nl}^{\tau} \rangle^{nl}}{T_l} + \langle s_{nl}^{\tau} \rangle^{nl} \right) \right] \\
& + \frac{1}{T_l} \left[ \alpha_l \rho_l \varepsilon_{\eta,l}^{\tau} + \mu_{wl}^{\tau} \frac{1}{2} \left( \langle \mathbf{V}_{wl}^{\tau} \rangle^{we} - \langle \mathbf{V}_l^{\tau} \rangle^{le} \right)^2 - \mu_{lw}^{\tau} \frac{1}{2} \left( \langle \mathbf{V}_{lw}^{\tau} \rangle^{le} - \langle \mathbf{V}_l^{\tau} \rangle^{le} \right)^2 \right].
\end{aligned}$$

Obviously *if and only if* the pressure at the boundary inside the field is equal to the field averaged pressure the sources can be rewritten as follows

$$\begin{aligned}
& \left( \alpha_l \rho_l \varepsilon_{\eta,l}^{\tau} + q_l^N \right) / T_l + \sum_{\substack{m=1 \\ m \neq l}}^3 \sum_{i=1}^{i_{\max}} \left( \mu_{iml}^{\tau} - \mu_{ilm}^{\tau} \right) \langle s_{il}^{\tau} \rangle^{il} = \frac{\dot{q}_l^m}{T_l} + \sum_{i=1}^{i_{\max}} \left[ \mu_{iwl}^{\tau} \langle s_{iw}^{\tau} \rangle^{iw} - \mu_{ilw}^{\tau} \langle s_{il}^{\tau} \rangle^{il} \right] \\
& + \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \mu_{Mml}^{\tau} s_{Ml}^{\tau\sigma} - \mu_{Mlm}^{\tau} s_{Ml}^{\tau\sigma} + \sum_{n=1}^{n_{\max}} \left( \mu_{nml}^{\tau} \langle s_{nm} \rangle^{nm} - \mu_{nlm}^{\tau} \langle s_{nl} \rangle^{nl} \right) \right] \\
& + \frac{1}{T_l} \left[ \alpha_l \rho_l \varepsilon_{\eta,l}^{\tau} + \mu_{wl}^{\tau} \frac{1}{2} \left( \langle \mathbf{V}_{wl}^{\tau} \rangle^{we} - \langle \mathbf{V}_l^{\tau} \rangle^{le} \right)^2 - \mu_{lw}^{\tau} \frac{1}{2} \left( \langle \mathbf{V}_{lw}^{\tau} \rangle^{le} - \langle \mathbf{V}_l^{\tau} \rangle^{le} \right)^2 \right].
\end{aligned}$$

Otherwise, the primary form has to be used.

We discussed in Chapter 4 a variety of notations of the energy conservation principle for single-phase flow. All of them are more complicated than the entropy notation. As for the single-phase flow comparing the multi-phase entropy equation (5.65) with the energy conservation equations in terms of specific enthalpy and specific internal energy (5.56, 5.58), it is evident that the computational effort required to approximate

- 1 time derivative,
- 1 divergence term,
- 2 diffusion terms, and
- 1 tensor product

is much less than the computational effort required to discretize

- 2 time derivatives,
- 2 divergence terms,
- 2 diffusion terms, and
- 1 tensor product.

Bearing in mind that each divergence term contains at least 3 differential terms, it is obvious that *for computational analysis the use of the entropy equation instead of energy equation in any of its variants is the most cost effective way for the numerical modeling of flows.*

**CONCLUSION:** The use of the specific entropies as components of the dependent variable vector, *the entropy concept, gives the simplest form for the mathematical description of the flow.* The use of any other state variables instead of the entropies makes the description more complicated.

## 5.7 Local volume- and time-averaged entropy equation

The procedure we use to obtain a time-averaged entropy equation will be used also for deriving the time-averaged enthalpy, or specific internal energy equations. Here we introduce the time averaging rules. The instantaneous surface-averaged velocity of the field  $l$ ,  $\langle \mathbf{V}_l^\tau \rangle^{le}$ , can be expressed as the sum of the surface-averaged velocity which is subsequently time averaged,

$$V_l := \overline{\langle \mathbf{V}_l^\tau \rangle^{le}} \quad (5.66)$$

and a fluctuation component  $V_l'$ ,

$$\langle \mathbf{V}_l^\tau \rangle^{le} = V_l + V_l' \quad (5.67)$$

as proposed by *Reynolds*. The same is performed for the pressure

$$\langle p_l^\tau \rangle^l = p_l + p_l', \quad (5.68)$$

and for any specific property

$$\langle \varphi_{il}^\tau \rangle^l = \varphi_{il} + \varphi_{il}', \quad (5.69)$$

where  $\varphi = s, h, e, T, C$ . Here the fluctuations of densities, volume fractions are neglected as they are only of small magnitude relative to the velocity and entropy fluctuations.

The time averaging rules we use are:

$$\overline{\alpha_l \langle \rho_l \rangle^l \varphi_{il}} = \alpha_l \langle \rho_l \rangle^l \varphi_{il}, \quad (5.70)$$

$$\overline{\alpha_l \langle \rho_l \rangle^l \varphi_{il}'} = 0, \quad (5.71)$$

$$\overline{\alpha_l^e \langle \rho_l \rangle^l \varphi_{il} \mathbf{V}_l} = \alpha_l^e \langle \rho_l \rangle^l \varphi_{il} \mathbf{V}_l, \quad (5.72)$$

$$\overline{\alpha_l^e \langle \rho_l \rangle^l \varphi_{il} \mathbf{V}_l'} = 0, \quad (5.73)$$

$$\overline{\alpha_l^e \langle \rho_l \rangle^l \varphi_{il}' \mathbf{V}_l} = 0, \quad (5.74)$$

$$\overline{\varphi'_{il} D'_{il} \nabla C_{il}} = 0, \quad (5.75)$$

$$\overline{\varphi_{il} D'_{il} \nabla C'_{il}} = 0, \quad (5.76)$$

$$\overline{\varphi_{il} D'_{il} \nabla C_{il}} = \overline{\varphi_{il} D'_{il} \nabla C_{il}}, \quad (5.77)$$

$$\overline{(\mu_{iml}^\tau - \mu_{ilm}^\tau) \langle \varphi_{il}^\tau \rangle^{il}} = (\mu_{iml} - \mu_{ilm}) \varphi_{il} \quad (5.78)$$

$$\overline{\mu_{il}^\tau} = \mu_{il}, \quad (5.79)$$

$$\overline{\mu_{ml}^\tau (\mathbf{V}'_m + \mathbf{V}'_m - \mathbf{V}_l - \mathbf{V}'_l)^2} = \overline{\mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)^2} + \overline{\mu_{ml} (\mathbf{V}'_m - \mathbf{V}'_l)^2}, \quad (5.80)$$

$$\overline{\mathbf{T}'_l : \nabla \cdot \mathbf{V}_l} = \overline{\mathbf{T}'_l : \nabla \cdot \mathbf{V}_l}, \quad (5.81)$$

$$\overline{\mathbf{T}_l : \nabla \cdot \mathbf{V}'_l} = 0, \quad (5.82)$$

$$\overline{\mathbf{T}_l : \nabla \cdot \mathbf{V}_l} = \overline{\mathbf{T}_l : \nabla \cdot \mathbf{V}_l}, \quad (5.83)$$

$$\overline{\mu_{iwl}^\tau \left( \langle h_{iwl}^\tau \rangle^{iw} - \langle h_{il}^\tau \rangle^{il} \right)} = \mu_{iwl} (h_{iw} - h_{il}) \quad (5.84)$$

$$\overline{\mu_{Mml}^\tau \left( h_{Ml}^{\tau\sigma} - \langle h_{Ml}^\tau \rangle^{Ml} \right)} = \mu_{Mml} (h_{Ml}^\sigma - h_{Ml}) \quad (5.85)$$

$$\overline{\mu_{nml}^\tau \left( \langle h_{nm}^\tau \rangle^{nm} - \langle h_{nl}^\tau \rangle^{nl} \right)} = \mu_{nml} (h_{nm} - h_{nl}) \quad (5.86)$$

$$\overline{q_l^{\tau N}} = q_l^N \quad (5.87)$$

$$\begin{aligned} q_l^N &= \dot{q}_l^N + \sum_{i=1}^{i_{\max}} \mu_{iwl} (h_{iw} - h_{il}) + \sum_{m=1}^3 \left[ \mu_{Mml} (h_{Ml}^\sigma - h_{Ml}) + \sum_{n=1}^{n_{\max}} \mu_{nml} (h_{nm} - h_{nl}) \right] \\ &+ \frac{1}{2} \left[ \mu_{wl} (\mathbf{V}_{wl} - \mathbf{V}_l)^2 - \mu_{lw} (\mathbf{V}_{lw} - \mathbf{V}_l)^2 + \sum_{m=1}^3 \mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)^2 \right] \\ &+ \frac{1}{2} \left[ \mu_{wl} \overline{(\mathbf{V}'_{wl} - \mathbf{V}'_l)^2} - \mu_{lw} \overline{(\mathbf{V}'_{lw} - \mathbf{V}'_l)^2} + \sum_{m=1}^3 \mu_{ml} \overline{(\mathbf{V}'_m - \mathbf{V}'_l)^2} \right] \end{aligned} \quad (5.88)$$

Substituting in Eq. (5.65) the above mentioned variables as consisting of mean values and fluctuations and performing time averaging following the above rules we obtain

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l s_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l s_l \mathbf{V}_l \gamma)$$

$$\begin{aligned}
 & -\frac{1}{T_l} \nabla \cdot \left( \alpha_i^e \gamma \delta_l \rho_l c_{pl} \frac{V_l^l}{Pr_l^l} \nabla T_l \right) - \nabla \cdot \left[ \alpha_i^e \delta_l \rho_l \gamma \left( \sum_{i=1}^{i_{\max}} s_{il} D_{il}^l \nabla C_{il} \right) \right] \\
 & + \nabla \cdot \left[ \alpha_i^e \rho_l \left( \overline{s_l' \mathbf{V}_l'} - \sum_{i=1}^{i_{\max}} \overline{s_{il}' D_{il}' \nabla C_{il}'} \right) \gamma \right] = \gamma_v \frac{1}{T_l} DT_l^N + \gamma_v \sum_{m=1}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} \cdot
 \end{aligned} \tag{5.89}$$

Here

$$\begin{aligned}
 DT_l^N &= \alpha_l \rho_l (P_{k,l} + \delta_l \varepsilon_{\eta,l} + \varepsilon_l) + E_l^* + E_l^{r*} + \dot{q}_l''' + \sum_{i=1}^{i_{\max}} \mu_{iwl} (h_{iwl} - h_{il}) \\
 & + \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}} \left[ \mu_{Mml} (h_{Ml}^\sigma - h_{Ml}) - \mu_{Mlm} (h_{Ml}^\sigma - h_{Ml}) + \sum_{n=1}^{n_{\max}} \mu_{nml} (h_{nm} - h_{nl}) \right] \\
 & + \frac{1}{2} \left[ \begin{aligned} & \mu_{wl} (\mathbf{V}_{wl} - \mathbf{V}_l)^2 - \mu_{lw} (\mathbf{V}_{lw} - \mathbf{V}_l)^2 + \sum_{m=1}^3 \mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)^2 \\ & + \mu_{wl} (\overline{\mathbf{V}'_{wl}} - \mathbf{V}_l')^2 - \mu_{lw} (\overline{\mathbf{V}'_{lw}} - \mathbf{V}_l')^2 + \sum_{m=1}^3 \mu_{ml} (\overline{\mathbf{V}'_m} - \mathbf{V}_l')^2 \end{aligned} \right], \tag{5.90}
 \end{aligned}$$

$$\gamma_v \alpha_l \rho_l \varepsilon_{\eta,l} = \alpha_l^e \gamma \cdot (\mathbf{T}_{\eta,l} : \nabla \cdot \mathbf{V}_l) \tag{5.91}$$

is the *irreversible dissipated power* caused by the viscous forces *due to deformation of the mean values of the velocities* in the space, and

$$\gamma_v \alpha_l \rho_l \varepsilon_l' = \alpha_l^e \gamma \cdot (\mathbf{T}_l' : \nabla \cdot \mathbf{V}_l') \tag{5.92}$$

is the *irreversibly dissipated power* in the viscous fluid due to *turbulent pulsations*, and

$$\gamma_v \alpha_l \rho_l P_{k,l} = \alpha_l^e \gamma \cdot (\mathbf{T}_l' : \nabla \cdot \mathbf{V}_l)$$

is the power needed for production of turbulence. By modeling of the turbulence usually the last term is removed from the energy conservation equation and introduced as a generation term for the turbulent kinetic energy.

It is evident that the mass transfer between the velocity fields and between the fields and external sources causes additional entropy transport as a result of the different pulsation characteristics of the donor and receiver fields.

For simplicity's sake, averaging signs are omitted except in the turbulent diffusion term, which will be discussed next. The diffusion can also have *macroscopic* character, being caused by the macroscopic strokes between eddies with dimensions considerably larger than the molecular dimensions: turbulent diffusion. In a mixture at rest, *molecular* strokes represent the only mechanism

driving diffusion. In real flows both mechanisms are observed. The higher the velocity of the flow, the higher the effect of turbulent diffusion.

To permit practical application of the entropy equation, it is necessary to define more accurately the term

$$\nabla \cdot \left[ \alpha_i^e \rho_l \left( s_l' \mathbf{V}_l' - \sum_{i=1}^{i_{\max}} s_{il}' D_{il}' \nabla C_{il}' \right) \gamma \right].$$

A possible assumption for this case is that the mechanism of entropy transport caused by fluctuations is a diffusion-like mechanism, which means that

$$\begin{aligned} & \nabla \cdot \left[ \alpha_i^e \rho_l \left( s_l' \mathbf{V}_l' - \sum_{i=1}^{i_{\max}} s_{il}' D_{il}' \nabla C_{il}' \right) \gamma \right] \\ &= -\frac{1}{T_l} \nabla \cdot \left[ \alpha_i^e \rho_l \left( c_{pl} \frac{v_l'}{\text{Pr}_l'} \nabla T_l \right) \gamma \right] - \nabla \cdot \left[ \alpha_i^e \rho_l \left( \sum_{i=1}^{i_{\max}} s_{il}' \frac{v_l'}{Sc_l'} \nabla C_{il}' \right) \gamma \right] \end{aligned} \quad (5.93)$$

where

$$\text{Pr}_l' = \rho_l c_{pl} v_l' / \lambda_l' \quad (5.94)$$

is the turbulent *Prandtl* number and  $\lambda_l'$  is the *turbulent coefficient of thermal conductivity* or *eddy conductivity*. Note that the thermal diffusion

$$\lambda_l = \rho_l c_{pl} v_l / \text{Pr}_l \quad (5.95)$$

is a thermodynamic property of the continuum  $l$ , and

$$\lambda_l' = \rho_l c_{pl} v_l' / \text{Pr}_l' \quad (5.96)$$

is a mechanical property of the flowing field  $l$ . In channels,

$$\text{Pr}_l' \approx 0.7 \dots 0.9, \quad (5.97)$$

whereas for flow in jets, i.e., in *free turbulence*, the value is closer to 0.5 -see *Bird et al.* (1960),

$$\text{Pr}_l' \approx 0.25 \dots 2.5 \quad (5.98)$$

for water, air and steam - see *Hammond* (1985)]. The turbulent *Schmidt* number is defined as

$$Sc_l' = v_l' / D_l' . \quad (5.99)$$

Here again the turbulent diffusion coefficient

$$D_l' = v_l' / Sc_l' \quad (5.100)$$

is a mechanical property of the flowing field  $l$ . As a result, the final form of the entropy equation is

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l s_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l s_l \mathbf{V}_l \gamma) - \frac{1}{T_l} \nabla \cdot \left[ \alpha_l^e \rho_l c_{pl} \left( \delta_l \frac{V_l}{Pr_l} + \frac{V_l'}{Pr_l'} \right) \gamma \mathcal{N} T_l \right] \\ & - \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \sum_{i=1}^{i_{\max}} s_{il} \left( \delta_l \frac{V_l}{Sc_l} + \frac{V_l'}{Sc_l'} \right) \nabla C_{il} \right] \right\} = \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{m=1}^{3,w} \sum_{\substack{i=1 \\ m \neq i}}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} \right], \end{aligned} \quad (5.101)$$

with this called the *conservative form*.

Keeping in mind that there is no net mass diffusively transported across a cross section perpendicular to the strongest concentration gradient

$$\sum_{i=1}^{i_{\max}} (D_{il}^* \nabla C_{il}) = 0, \quad (5.102)$$

results in

$$D_{il}^* \nabla C_{il} = - \sum_{i=2}^{i_{\max}} (D_{il}^* \nabla C_{il}). \quad (5.103)$$

where the effective conductivity, and the diffusivity are

$$\lambda_l^* = \delta_l \lambda_l + \rho_l c_{pl} \frac{V_l'}{Pr_l'} = \rho_l c_{pl} \left( \delta_l \frac{V_l}{Pr_l} + \frac{V_l'}{Pr_l'} \right), \quad (5.104)$$

$$D_{il}^* = \delta_l D_{il} + \frac{V_l'}{Sc_l'} = \delta_l \frac{V_l}{Sc_l} + \frac{V_l'}{Sc_l'}, \quad (5.105)$$

we obtain finally

$$\boxed{\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l s_l \gamma_v) + \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ s_l \mathbf{V}_l - \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \nabla C_{il} \right] \right\} \\ & - \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \mathcal{N} T_l) = \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{m=1}^{3,w} \sum_{\substack{i=1 \\ m \neq i}}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} \right]. \end{aligned}} \quad (5.101b)$$

For a mixture consisting of several inert components and one no-inert component it is advisable to select subscript  $i = 1 = M$  for the non-inert component and consider the mixture of the inert components as an ideal gas. In this case we have



$$\dots = \nabla \cdot \left[ \alpha_i^e \rho_l \gamma (s_{nl} - s_{Ml}) D_{nl}^* \nabla C_{nl} \right] \dots \quad (5.101c)$$

This method is used in IVA6 computer code *Kolev* (1997b) developed by this author.

## 5.8 Local volume- and time-averaged internal energy equation

Following the same procedure as for the entropy equation we perform time averaging of Eq. (5.58). The result is

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l e_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l e_l \mathbf{V}_l \gamma) + p_l \left[ \frac{\partial}{\partial \tau} (\alpha_l \gamma_v) + \nabla \cdot (\alpha_l^e \mathbf{V}_l \gamma) \right] \\ & - \nabla \cdot \left[ \alpha_l^e \rho_l \left( \delta_l c_{pl} \frac{V_l'}{Pr_l'} \nabla T_l + e_l' \overline{\mathbf{V}_l'} \right) \gamma \right] \\ & - \nabla \cdot \left[ \alpha_l^e \rho_l \left( \delta_l \sum_{i=1}^{l_{\max}} e_{il} D_{il}' \nabla C_{il} + \sum_{i=1}^{l_{\max}} e_{il}' D_{il}' \nabla C_{il}' \right) \gamma \right] + \overline{p_l' \nabla \cdot (\alpha_l^e \mathbf{V}_l' \gamma)} = \gamma_v D e_l^* \end{aligned} \quad (5.106)$$

where

$$\begin{aligned} D e_l^* &= \alpha_l \rho_l \left( P_{k,l} + \delta_l \varepsilon_{\eta,l} + \varepsilon_l' \right) + \dot{q}''' \\ & + \frac{1}{2} \left[ \begin{aligned} & \mu_{wl} (\mathbf{V}_{wl} - \mathbf{V}_l)^2 - \mu_{lw} (\mathbf{V}_{lw} - \mathbf{V}_l)^2 + \sum_{\substack{m=1 \\ m \neq l}}^3 \mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)^2 \\ & + \mu_{wl} (\overline{\mathbf{V}_{wl}' - \mathbf{V}_l'})^2 - \mu_{lw} (\overline{\mathbf{V}_{lw}' - \mathbf{V}_l'})^2 + \sum_{\substack{m=1 \\ m \neq l}}^3 \mu_{ml} (\overline{\mathbf{V}_m' - \mathbf{V}_l'})^2 \end{aligned} \right] \\ & + \sum_{i=1}^{l_{\max}} (\mu_{iwl} h_{iwl} - \mu_{ilw} h_{il}) + \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}} \left[ \mu_{Mml} h_{Ml}^\sigma + \sum_{n=1}^{n_{\max}} (\mu_{nml} h_{nm} - \mu_{nlm} h_{nl}) \right] \end{aligned} \quad (5.107)$$

or introducing

$$\nabla \cdot \left[ \alpha_l^e \rho_l \left( e_l' \overline{\mathbf{V}_l'} - \delta_l \sum_{i=1}^{l_{\max}} e_{il}' D_{il}' \nabla C_{il}' \right) \gamma \right]$$

$$= -\nabla \cdot \left[ \alpha_i^e \rho_l \left( c_{pl} \frac{v_l^t}{Pr_l^t} \nabla T_l + \sum_{i=1}^{i_{\max}} e_{il} \frac{v_l^t}{Sc_{il}^t} \nabla C_{il} \right) \gamma \right] \quad (5.108)$$

we obtain finally

$$\boxed{\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_i \rho_l e_l \gamma_v) + \nabla \cdot (\alpha_i^e \rho_l e_l \mathbf{V}_l \gamma) + p_l \left[ \frac{\partial}{\partial \tau} (\alpha_i \gamma_v) + \nabla \cdot (\alpha_i^e \mathbf{V}_l \gamma) \right] \\ & - \nabla \cdot \left\{ \alpha_i^e \gamma \left[ \lambda_i^* \nabla T_l + \rho_l \sum_{i=2}^{i_{\max}} (e_{il} - e_{il}) D_{il}^* \nabla C_{il} \right] \right\} + \overline{p_l \nabla \cdot (\alpha_i^e \mathbf{V}_l \gamma)} = \gamma_v D e_l^* \end{aligned}} \quad (5.109)$$

The expression defining the diffusion coefficient for the specific internal mixture energy is then

$$\begin{aligned} & \nabla \cdot (\alpha_i^e \rho_l D_l^e \gamma \cdot \nabla e_l) \\ & = \nabla \cdot \left\{ \alpha_i^e \rho_l \gamma \left[ c_{pl} \left( \delta_l \frac{v_l}{Pr_l} + \frac{v_l^t}{Pr_l^t} \right) \nabla T_l + \sum_{i=2}^{i_{\max}} (e_{il} - e_{il}) D_{il}^* \nabla C_{il} \right] \right\} \end{aligned} \quad (5.110a)$$

which for ideal gas mixtures where the specific internal energy is not a function of pressure reduces to

$$\nabla \cdot (\alpha_i^e \rho_l D_l^e \gamma \cdot \nabla e_l) = \nabla \cdot \left\{ \alpha_i^e \rho_l \left[ \frac{\lambda_l^*}{\rho_l c_{pl}} + \sum_{i=1}^{i_{\max}} e_{il} D_{il}^* / \left( \frac{\partial e_l}{\partial C_{il}} \right)_{p, T_l} \right] \gamma \nabla e_l \right\} \quad (5.110b)$$

where

$$D_l^e = \frac{\lambda_l^*}{\rho_l c_{pl}} + \sum_{i=1}^{i_{\max}} e_{il} D_{il}^* / \left( \frac{\partial e_l}{\partial C_{il}} \right)_{p, T_l} \quad (5.111)$$

This equation can be used to derive an alternative form for computation of the entropy diffusion coefficient

$$\nabla \cdot (\alpha_i^e \rho_l D_l^s \gamma \cdot \nabla s_l) = \left[ \nabla \cdot (\alpha_i^e \rho_l D_l^e \gamma \cdot \nabla e_l) - \overline{p_l \nabla \cdot (\alpha_i^e \mathbf{V}_l \gamma)} \right] / T_l \quad (5.112)$$

## 5.9 Local volume- and time-averaged specific enthalpy equation

Following the same procedure as for the entropy equation we perform time averaging of Eq. (5.55). The result is

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_l \rho_l h_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l \mathbf{V}_l h_l \gamma) - \left( \alpha_l \gamma_v \frac{\partial p_l}{\partial \tau} + \alpha_l^e \mathbf{V}_l \gamma \cdot \nabla p_l \right) \\
 & - \nabla \cdot \left\{ \alpha_l^e \gamma \left[ \delta_l \lambda_l \nabla T_l - \rho_l \overline{\mathbf{V}_l h_l'} + \delta_l \rho_l \sum_{i=1}^{i_{\max}} (h_{il} D_{il} \nabla C_{il}) + \rho_l \sum_{i=1}^{i_{\max}} (h_{il}' D_{il} \nabla C_{il}') \right] \right\} \\
 & - \alpha_l^e \overline{\mathbf{V}_l' \gamma \cdot \nabla p_l'} + \delta_l (\alpha_l^e \gamma \overline{\mathcal{N} p_l}) \cdot \sum_{i=1}^{i_{\max}} D_{il}' \nabla \ln C_{il} \\
 & = \gamma_v D T_l^N + \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) h_{il}
 \end{aligned} \tag{5.113}$$

Introducing

$$\begin{aligned}
 & \nabla \cdot \left\{ \alpha_l^e \rho_l \left[ \overline{\mathbf{V}_l h_l'} - \sum_{i=1}^{i_{\max}} (h_{il}' D_{il} \nabla C_{il}') \right] \gamma \right\} \\
 & = -\nabla \cdot \left[ \alpha_l^e \rho_l \left( \delta_l c_{pl} \frac{\mathbf{v}_l'}{P_l^t} \nabla T_l + \sum_{i=1}^{i_{\max}} h_{il} \frac{\mathbf{v}_l'}{S C_l^t} \nabla C_{il} \right) \gamma \right]
 \end{aligned} \tag{5.114}$$

we obtain

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_l \rho_l h_l \gamma_v) + \nabla \cdot (\alpha_l^e \rho_l \mathbf{V}_l h_l \gamma) - \left( \alpha_l \gamma_v \frac{\partial p_l}{\partial \tau} + \alpha_l^e \mathbf{V}_l \gamma \cdot \nabla p_l \right) \\
 & - \nabla \cdot \left\{ \alpha_l^e \left[ \lambda_l^* \nabla T_l + \rho_l \sum_{i=2}^{i_{\max}} (h_{il} - h_{li}) D_{il}^* \nabla C_{il} \right] \gamma \right\} - \alpha_l^e \overline{\mathbf{V}_l' \gamma \cdot \nabla p_l'} \\
 & + \delta_l (\alpha_l^e \gamma \overline{\mathcal{N} p_l}) \cdot \sum_{i=1}^{i_{\max}} D_{il}' \nabla \ln C_{il} = \gamma_v D T_l^N + \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) h_{il}.
 \end{aligned} \tag{5.115}$$

The non-conservative form is then

$$\begin{aligned}
 & \alpha_l \rho_l \gamma_v \frac{\partial h_l}{\partial \tau} + \alpha_l^e \rho_l (\mathbf{V}_l \gamma \cdot \nabla) h_l - \left( \alpha_l \gamma_v \frac{\partial p_l}{\partial \tau} + \alpha_l^e \mathbf{V}_l \gamma \cdot \nabla p_l \right) \\
 & - \nabla \cdot \left\{ \alpha_l^e \left[ \lambda_l^* \nabla T_l + \rho_l \sum_{i=2}^{i_{\max}} (h_{il} - h_{li}) D_{il}^* \nabla C_{il} \right] \gamma \right\} - \alpha_l^e \overline{\mathbf{V}_l' \gamma \cdot \nabla p_l'}
 \end{aligned}$$

$$+\delta_l (\alpha_l^e \gamma \bar{V} p_l) \cdot \sum_{i=1}^{i_{\max}} D_{il}^i \nabla \ln C_{il} = \gamma_v D T_l^N. \quad (5.115)$$

The definition equation for an enthalpy diffusion coefficient is then

$$\nabla \cdot (\alpha_l^e \rho_l D_l^h \gamma \bar{V} h_l) = \nabla \cdot \left\{ \alpha_l^e \rho_l \left[ \frac{\lambda_l^*}{\rho_l} \nabla T_l + \sum_{i=2}^{i_{\max}} (h_{il} - h_{l1}) D_{il}^* \nabla C_{il} \right] \gamma \right\}. \quad (5.116)$$

For mixtures of two ideal gases  $n$  and  $M$  we have

$$\nabla \cdot (\alpha_l^e \rho_l D_l^h \gamma \cdot \nabla h_l) = \nabla \cdot \left\{ \alpha_l^e \rho_l \left[ \frac{\lambda_l^*}{\rho_l} \nabla T_l + (h_{nl} - h_{Ml}) D_{nl}^* \nabla C_{nl} \right] \gamma \right\} \quad (5.117a)$$

and an equation of state in differential form

$$\nabla h_l = c_{pl} \nabla T_l + (h_{nl} - h_{Ml}) \nabla C_{nl},$$

which finally gives

$$\nabla \cdot (\alpha_l^e \rho_l D_l^h \gamma \bar{V} h_l) = \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \lambda_l^* (1 - Le_{il}) \nabla T + \rho_l D_{nl}^* \nabla h_l \right] \right\} \quad (5.117b)$$

where

$$Le_{il} = \frac{\rho_l c_{pl} D_{nl}^*}{\lambda_l^*}$$

is the *Lewis-Semenov* number. For a lot of binary gas mixtures  $Le_{il} \approx 1$  and

$$\nabla \cdot (\alpha_l^e \rho_l D_l^h \gamma \bar{V} h_l) = \nabla \cdot (\alpha_l^e \rho_l D_{nl}^* \gamma \bar{V} h_l), \quad (5.117c)$$

which results in

$$D_l^h = D_{nl}^*,$$

*Grigorieva and Zorina* (1988) p. 267. This relation is widely used for simulation of diffusion in binary gas mixtures. If both components possess equal specific capacities at constant pressures, and therefore  $h_{nl} \approx h_{Ml}$ , we have

$$\nabla \cdot (\alpha_l^e \rho_l D_l^h \gamma \cdot \nabla h_l) = \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \bar{V} T_l). \quad (5.117d)$$

This condition is fulfilled for many liquid mixtures and solutions. In this case the diffusion enthalpy transport looks like the *Fick's* law and is totally controlled by the temperature gradient.

## 5.10 Non-conservative and semi-conservative forms of the entropy equation

The *non-conservative* form of the entropy equation is obtained by differentiating the first and second terms of the conservative form appropriately and comparing with the mass conservation equation. The result is

$$\begin{aligned}
 & \rho_l \left[ \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) s_l \right] \\
 & - \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \mathcal{N} T) - \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \nabla C_{il} \right] \right\} \\
 & = \gamma_v \left( \frac{1}{T_l} D T_l^N + \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}, w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - \mu_l s_l \right) \equiv \gamma_v D s_l^N. \quad (5.118)
 \end{aligned}$$

The superscript  $N$  stands to remember that the RHS is for the non-conservative form of the energy conservation equation written in entropy form. *Note* that in addition to the enthalpy source terms divided by the field temperature, new terms arise. In fact, these denote the difference between the sum of the mass source per unit time and mixture volume multiplied by the corresponding specific component entropies and the product of the specific field entropy and the field mass source density.

$$\gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}, w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - s_l \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}, w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) \right]. \quad (5.119)$$

Splitting the mass source term into *sources* and *sinks*

$$\mu_l = \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}, w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) = \mu_l^+ - \mu_l^- \quad (5.120)$$

where

$$\mu_l^+ = \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}, w} \sum_{i=1}^{i_{\max}} \mu_{iml} \geq 0, \quad (5.121)$$

$$\mu_l^- = \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}, w} \sum_{i=1}^{i_{\max}} \mu_{ilm} \geq 0, \quad (5.122)$$

$$\mu_l^+ s_l \geq 0, \quad (5.123)$$

$$\mu_l^- s_l \geq 0, \quad (5.124)$$

one then obtains the *final semi-conservative* form for the entropy equation

$$\begin{aligned} & \rho_l \left[ \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) s_l \right] - \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \mathcal{N} T) \\ & - \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \nabla C_{il} \right] \right\} + \gamma_v \mu_l^+ s_l = \gamma_v D s_l, \end{aligned} \quad (5.125)$$

in which

$$\begin{aligned} D s_l &= \frac{1}{T_l} D T_l^N + \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - \mu_l s_l + \mu_l^+ s_l \\ &= \frac{1}{T_l} D T_l^N + \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} + \mu_l^- s_l \\ &= \frac{1}{T_l} D T_l^N + \gamma_v \left\{ \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left[ \mu_{lm} s_l + \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} \right] \right\} \end{aligned} \quad (5.126)$$

The three forms of the entropy equation, conservative (5.110b), non-conservative (5.118), and semi-conservative (5.125) are mathematically identical.

The introduction of the semi-conservative form is perfectly suited to numerical integration because it ensures proper initialization of the value for the entropy in a computational cell in which a previously non-existent field is just in the course of origination.

## 5.11 Comments on the source terms in the mixture entropy equation

The following terms in the semi-conservative entropy equation (5.125),

$$D s_l = \dots + \gamma_v \left\{ \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left[ \mu_{lm} s_l + \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} \right] \right\}$$

$$+ \frac{\gamma_v}{T_l} \left\{ \sum_{i=1}^{i_{\max}} \mu_{iwl} (h_{iwl} - h_{il}) + \sum_{m=1}^3 \left[ \mu_{Mml} (h_{Ml}^\sigma - h_{Ml}) + \sum_{n=1}^{n_{\max}} \mu_{nml} (h_{nm} - h_{nl}) \right] \right\}, \quad (5.127)$$

are discussed in some detail in this section in order to facilitate their practical application. For the wall source terms, that is, for injection from the wall into the flow and suction from the flow into the wall one can write

$$\mu_{iwl} = C_{iwl} \mu_{wl}, \quad (5.128)$$

$$\mu_{ilw} = C_{ilw} \mu_{lw}, \quad (5.129)$$

and therefore

$$Ds_l = \dots + \gamma_v \mu_{wl} \sum_{i=1}^{i_{\max}} C_{iwl} \left[ s_{il} + (h_{iwl} - h_{il}) / T_l \right] + \gamma_v \sum_{m=1}^{i_{\max}} \left\{ \mu_{lm} s_l + \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} \right\} \\ + \frac{\gamma_v}{T_l} \sum_{m=1}^3 \left[ \mu_{Mml} (h_{Ml}^\sigma - h_{Ml}) + \sum_{n=1}^{n_{\max}} \mu_{nml} (h_{nm} - h_{nl}) \right]. \quad (5.130)$$

Similar relationships as Eqs. (5.128) and (5.129) are valid for entrainment from field 2 to 3 and deposition from field 3 to 2, that is for  $lm = 23, 32$ . Note that, say, for steam condensation from multi-component gas mixtures,  $lm = 12, 13$ , and for evaporation,  $lm = 21, 31$ , the following then applies:  $\mu_{ilm} \neq \mu_{lm} C_{il}$ . For practical applications it is convenient to group all inert and non-inert components in pseudo-two-component mixtures inside the field  $l$  designated simply with  $n$  and  $M$ , respectively.

The most important mass transfer mechanisms in the three velocity fields flow will now be taken into account in accordance with the following assumptions:

- No solution of gas into field 2 and/or 3;
- No gas dissolution from field 2 and/or 3;
- Evaporation from field 2 and/or 3 into the gas field 1 is allowed;
- Condensation of the component  $M$  from the gas field onto the interface of fields 2 and/or 3 is allowed;
- Injection from the wall into all of the fields is allowed;
- Suction from all of the fields through the wall is allowed.

The mass transfer terms are written in the following form:

$$\mu_{nlm} = \begin{vmatrix} 0 & 0 & 0 & \mu_{n1w} = C_{n1} \mu_{1w} \\ 0 & 0 & \mu_{n23} = C_{n2} \mu_{23} & \mu_{n2w} = C_{n2} \mu_{2w} \\ 0 & \mu_{n32} = C_{n3} \mu_{32} & 0 & \mu_{n3w} = C_{n3} \mu_{3w} \\ \mu_{nw1} = C_{nw1} \mu_{w1} & \mu_{nw2} = C_{nw2} \mu_{w2} & \mu_{nw3} = C_{nw3} \mu_{w3} & 0 \end{vmatrix}, \quad (5.131)$$

$$\mu_{Mlm} = \begin{vmatrix} 0 & \mu_{M12} = \mu_{12} & \mu_{M13} = \mu_{13} & \mu_{M1w} = C_{M1}\mu_{1w} \\ \mu_{M21} = \mu_{21} & 0 & \mu_{M23} = C_{M2}\mu_{23} & \mu_{M2w} = C_{M2}\mu_{2w} \\ \mu_{M31} = \mu_{31} & \mu_{M32} = C_{M3}\mu_{32} & 0 & \mu_{M3w} = C_{M3}\mu_{3w} \\ \mu_{Mw1} = C_{Mw1}\mu_{w1} & \mu_{Mw2} = C_{Mw2}\mu_{w2} & \mu_{Mw3} = C_{Mw3}\mu_{w3} & 0 \end{vmatrix}. \quad (5.132)$$

Except for the diagonal elements, the zeros in Eq. (5.132) result from assumptions a) and b). Assumptions c) and d) give rise to the terms

$$\mu_{Mlm} = \begin{vmatrix} 0 & \mu_{M12} = \mu_{12} & \mu_{M13} = \mu_{13} & \dots \\ \mu_{M21} = \mu_{21} & 0 & \dots & \dots \\ \mu_{M31} = \mu_{31} & \dots & 0 & \dots \\ \dots & \dots & \dots & 0 \end{vmatrix} \quad (5.133)$$

in the above equation.

Figures 5.2 and 5.3 illustrate the *l*-field side interface properties for the three fields under consideration.

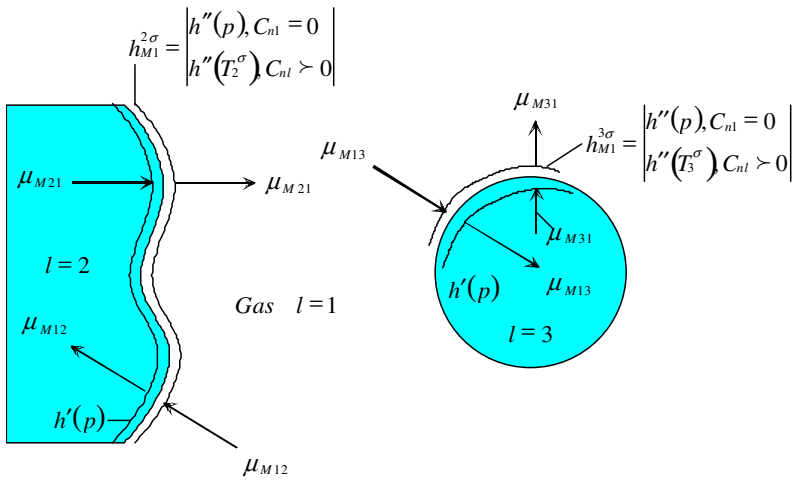
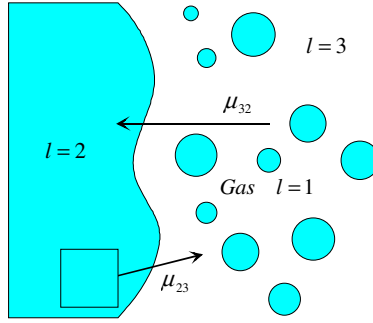


Fig. 5.2 Evaporation and condensation mass transfer. Interface properties





**Fig. 5.3** Entrainment and deposition mass transfer

Taking into account Eqs. (5.131) and (5.132), the specific form for the source terms is:

$$\begin{aligned}
 Ds_1 &= \dot{q}_1^m / T_1 \\
 &+ \gamma_v \mu_{w1} \left\{ (1 - C_{nw1}) [s_{M1} + (h_{Mw1} - h_{M1}) / T_1] + C_{nw1} [s_{n1} + (h_{nw1} - h_{n1}) / T_1] \right\} \\
 &+ \gamma_v \left[ \begin{array}{l} \mu_{21} [s_{M1} + (h_{M1}^{2\sigma} - h_{M1}) / T_1] - \mu_{12} C_{n1} (s_{M1} - s_{n1}) \\ \mu_{31} [s_{M1} + (h_{M1}^{3\sigma} - h_{M1}) / T_1] - \mu_{13} C_{n1} (s_{M1} - s_{n1}) \end{array} \right] + \alpha_1 \rho_1 (\delta_1 P_{k1}^e + \varepsilon_1) / T_1 \\
 &+ \frac{1}{2} \left[ \mu_{w1} (\mathbf{V}_{w1} - \mathbf{V}_1)^2 - \mu_{1w} (\mathbf{V}_{1w} - \mathbf{V}_1)^2 + \sum_{m=2,3} \mu_{m1} (\mathbf{V}_m - \mathbf{V}_1)^2 \right] / T_1 \\
 &+ \frac{1}{2} \left[ \overline{\mu_{w1} (\mathbf{V}'_{w1} - \mathbf{V}'_1)^2} - \overline{\mu_{1w} (\mathbf{V}'_{1w} - \mathbf{V}'_1)^2} + \sum_{m=2,3} \overline{\mu_{m1} (\mathbf{V}'_m - \mathbf{V}'_1)^2} \right] / T_1, \quad (5.134)
 \end{aligned}$$

$$\begin{aligned}
 Ds_2 &= \dot{q}_2^m / T_2 \\
 &+ \gamma_v \mu_{w2} \left\{ (1 - C_{nw2}) [s_{M2} + (h_{Mw2} - h_{M2}) / T_2] + C_{nw2} [s_{n2} + (h_{nw2} - h_{n2}) / T_2] \right\} \\
 &+ \gamma_v \mu_{32} \left\{ (1 - C_{n3}) [s_{M2} + (h_{M3} - h_{M2}) / T_2] + C_{n3} [s_{n2} + (h_{n3} - h_{n2}) / T_2] \right\} \\
 &+ \gamma_v \left\{ \mu_{12} [s_{M2} + (h_{M2}^\sigma - h_{M2}) / T_2] - \mu_{21} C_{n2} (s_{M2} - s_{n2}) \right\} + \alpha_2 \rho_2 (\delta_2 P_{k2}^e + \varepsilon_2) / T_2 \\
 &+ \frac{1}{2} \left[ \mu_{w2} (\mathbf{V}_{w2} - \mathbf{V}_2)^2 - \mu_{2w} (\mathbf{V}_{2w} - \mathbf{V}_2)^2 + \sum_{m=1,3} \mu_{m2} (\mathbf{V}_m - \mathbf{V}_2)^2 \right] / T_2
 \end{aligned}$$

$$+ \frac{1}{2} \left[ \overline{\mu_{w2} (\mathbf{V}'_{w2} - \mathbf{V}'_2)^2} - \overline{\mu_{2w} (\mathbf{V}'_{2w} - \mathbf{V}'_2)^2} + \sum_{m=1,3} \overline{\mu_{m2} (\mathbf{V}'_m - \mathbf{V}'_2)^2} \right] / T_2, \quad (5.135)$$

$$Ds_3 = \dot{q}_3''' / T_3$$

$$\begin{aligned} & + \gamma_v \mu_{w3} \left\{ (1 - C_{nw3}) \left[ s_{M3} + (h_{Mw3} - h_{M3}) / T_3 \right] + C_{nw3} \left[ s_{n3} + (h_{nw3} - h_{n3}) / T_3 \right] \right\} \\ & + \gamma_v \mu_{23} \left\{ (1 - C_{n2}) \left[ s_{M3} + (h_{M2} - h_{M3}) / T_3 \right] + C_{n2} \left[ s_{n3} + (h_{n2} - h_{n3}) / T_3 \right] \right\} \\ & + \gamma_v \left\{ \mu_{13} \left[ s_{M3} + (h_{M3}^\sigma - h_{M3}) / T_3 \right] - \mu_{31} C_{n3} (s_{M3} - s_{n3}) \right\} + \alpha_3 \rho_3 (\delta_3 P_{k3}^r + \varepsilon_3) / T_3 \\ & + \frac{1}{2} \left[ \mu_{w3} (\mathbf{V}_{w3} - \mathbf{V}_3)^2 - \mu_{3w} (\mathbf{V}_{3w} - \mathbf{V}_3)^2 + \sum_{m=1,2} \mu_{m3} (\mathbf{V}_m - \mathbf{V}_3)^2 \right] / T_3 \\ & + \frac{1}{2} \left[ \overline{\mu_{w3} (\mathbf{V}'_{w3} - \mathbf{V}'_3)^2} - \overline{\mu_{3w} (\mathbf{V}'_{3w} - \mathbf{V}'_3)^2} + \sum_{m=1,2} \overline{\mu_{m3} (\mathbf{V}'_m - \mathbf{V}'_3)^2} \right] / T_3. \quad (5.136) \end{aligned}$$

The above equations can be simplified by taking into account

$$s_{iwl} = s_{il} + (h_{iwl} - h_{il}) / T_l, \quad (5.137)$$

$$s_{wl} = \sum_{i=1}^{i_{\max}} C_{iwl} s_{iwl}, \quad (5.138)$$

$$s_{il} = s_{im} + (h_{il} - h_{im}) / T_m, \quad (5.139)$$

$$s_l = \sum_{i=1}^{i_{\max}} C_{il} s_{il}, \quad (5.140)$$

$$s_{Ml}^\sigma = s_{Ml} + (h_{Ml}^\sigma - h_{Ml}) / T_l, \quad (5.141)$$

for  $l = 2, 3$  and  $m = 3, 2$  respectively.

$$\begin{aligned} Ds_1 = & \dots + \gamma_v \mu_{w1} \left\{ (1 - C_{nw1}) \left[ s_{M1} + (h_{Mw1} - h_{M1}) / T_1 \right] + C_{nw1} \left[ s_{n1} + (h_{nw1} - h_{n1}) / T_1 \right] \right\} \\ & + \gamma_v \left[ \begin{aligned} & (\mu_{21} + \mu_{31}) s_{M1} + \left[ \mu_{21} (h_{M1}^{2\sigma} - h_{M1}) + \mu_{31} (h_{M1}^{3\sigma} - h_{M1}) \right] / T_1 \\ & - (\mu_{12} + \mu_{13}) C_{n1} (s_{M1} - s_{n1}) \end{aligned} \right] + \dots, \quad (5.142) \end{aligned}$$

$$Ds_2 = \dots + \gamma_v \left[ \mu_{w2} s_{w2} + \mu_{32} s_3 + \mu_{12} s'(p) \right] + \dots, \quad (5.143)$$

$$Ds_3 = \dots + \gamma_v \left[ \mu_{w3} s_{w3} + \mu_{23} s_2 + \mu_{13} s'(p) \right] + \dots \quad (5.144)$$

Note that the mass transfer terms can give rise to entropy change even if mass leaves the field with entropy at the interface not equal to the intrinsic average field entropy. There is an additional source of entropy change if the non-inert component leaves the field with entropy at the interface not equal to the intrinsic average field entropy as a result of evaporation or condensation.

For evaporation only within the closed control volume one obtains

$$(\mu_{12} + \mu_{13}) s_1 = \mu_{21} \left[ s_{M1} + \left( h''(T_2^\sigma) - h_{M1} \right) / T_1 \right] + \mu_{31} \left[ s_{M1} + \left( h''(T_3^\sigma) - h_{M1} \right) / T_1 \right] \quad (5.145)$$

or

$$s_1 = \left\{ \mu_{21} \left[ s_{M1} + \left( h''(T_2^\sigma) - h_{M1} \right) / T_1 \right] + \mu_{31} \left[ s_{M1} + \left( h''(T_3^\sigma) - h_{M1} \right) / T_1 \right] \right\} / (\mu_{12} + \mu_{13}) . \quad (5.146)$$

For condensation and deposition the following is obtained for the specific entropy of the second velocity field

$$s_2 = \left[ \mu_{32} s_3 + \mu_{12} s'(p) \right] / (\mu_{12} + \mu_{32}) . \quad (5.147)$$

For condensation and entrainment the following is obtained for the specific entropy of the third velocity field

$$s_3 = \left[ \mu_{23} s_2 + \mu_{13} s'(p) \right] / (\mu_{13} + \mu_{23}) . \quad (5.148)$$

A frequently used simplification of the energy jump condition for computing the resulting evaporation or condensation is

$$\dot{q}_{m\sigma l}''' + \dot{q}_{l\sigma m}''' + \mu_{Ml}^\tau (h_{Mm}^\sigma - h_{Ml}^\sigma) = 0, \quad (5.149)$$

or if  $-(\dot{q}_{m\sigma l}''' + \dot{q}_{l\sigma m}''') / (h_{Mm}^\sigma - h_{Ml}^\sigma) > 0$ , then

$$\mu_{Mml} = -(\dot{q}_{m\sigma l}''' + \dot{q}_{l\sigma m}''') / (h_{Mm}^\sigma - h_{Ml}^\sigma) \quad \text{and} \quad \mu_{Mlm} = 0, \quad (5.150)$$

else

$$\mu_{Mlm} = (\dot{q}_{m\sigma l}''' + \dot{q}_{l\sigma m}''') / (h_{Ml}^\sigma - h_{Mm}^\sigma) \quad \text{and} \quad \mu_{Mml} = 0. \quad (5.151)$$

## 5.12 Viscous dissipation

Now we will discuss three important terms reflecting the irreversible part of the dissipation of the mechanical energy. The *irreversibly dissipated power* in the

viscous fluid due to *turbulent pulsations* and due to *change of the mean velocity* in space is

$$\gamma_v \alpha_l \rho_l \varepsilon_l = \gamma_v \alpha_l \rho_l (\varepsilon'_l + \varepsilon_{\eta,l}) = \alpha_l^e \gamma \cdot [(\mathbf{T}'_l : \nabla \cdot \mathbf{V}'_l) + (\mathbf{T}_{\eta,l} : \nabla \cdot \mathbf{V}_l)]. \quad (5.152)$$

These components can not be returned back as a mechanical energy of the flow. They express quantitatively the transfer of mechanical in thermal energy in the field  $l$ . In a notation common for Cartesian and cylindrical coordinates the irreversibly dissipated power in the viscous fluid due to turbulent pulsations is expressed as follows

$$\begin{aligned} \frac{\varepsilon'_l}{v_l} = & 2 \left\{ \gamma_r \overline{\left( \frac{\partial u'_l}{\partial r} \right)^2} + \gamma_\theta \overline{\left[ \frac{1}{r^\kappa} \left( \frac{\partial v'_l}{\partial \theta} + \kappa u'_l \right) \right]^2} + \gamma_z \overline{\left( \frac{\partial w'_l}{\partial z} \right)^2} \right\} \\ & + \overline{\left[ \frac{\partial v'_l}{\partial r} + \frac{1}{r^\kappa} \left( \frac{\partial u'_l}{\partial \theta} - \kappa v'_l \right) \right]} \overline{\left[ \gamma_r \frac{\partial v'_l}{\partial r} + \gamma_\theta \frac{1}{r^\kappa} \left( \frac{\partial u'_l}{\partial \theta} - \kappa v'_l \right) \right]} \\ & + \overline{\left[ \frac{\partial w'_l}{\partial r} + \frac{\partial u'_l}{\partial z} \right]} \overline{\left[ \gamma_r \frac{\partial w'_l}{\partial r} + \gamma_z \frac{\partial u'_l}{\partial z} \right]} + \overline{\left[ \frac{1}{r^\kappa} \frac{\partial w'_l}{\partial \theta} + \frac{\partial v'_l}{\partial z} \right]} \overline{\left[ \gamma_\theta \frac{1}{r^\kappa} \frac{\partial w'_l}{\partial \theta} + \gamma_z \frac{\partial v'_l}{\partial z} \right]} \\ & - \frac{2}{3} \overline{\left[ \frac{\partial u'_l}{\partial r} + \frac{1}{r^\kappa} \left( \frac{\partial v'_l}{\partial \theta} + \kappa u'_l \right) + \frac{\partial w'_l}{\partial z} \right]} \overline{\left[ \gamma_r \frac{\partial u'_l}{\partial r} + \gamma_\theta \frac{1}{r^\kappa} \left( \frac{\partial v'_l}{\partial \theta} + \kappa u'_l \right) + \gamma_z \frac{\partial w'_l}{\partial z} \right]} \geq 0. \end{aligned} \quad (5.153)$$

All terms in the right-hand side of this equation are time averages. Similarly, the irreversibly dissipated power in the viscous fluid due to deformation of the mean velocity field in space is expressed as follows

$$\begin{aligned} \gamma_v \alpha_l \rho_l \varepsilon_{\eta,l} &= \alpha_l^e \gamma \cdot (\mathbf{T}_{\eta,l} : \nabla \cdot \mathbf{V}_l) \\ &= \alpha_l^e \rho_l v_l \left( \begin{aligned} & 2 \left\{ \gamma_r \overline{\left( \frac{\partial u_l}{\partial r} \right)^2} + \gamma_\theta \overline{\left[ \frac{1}{r^\kappa} \left( \frac{\partial v_l}{\partial \theta} + \kappa u_l \right) \right]^2} + \gamma_z \overline{\left( \frac{\partial w_l}{\partial z} \right)^2} \right\} \\ & - \frac{2}{3} (\nabla \cdot \mathbf{V}_l) (\gamma \nabla \cdot \mathbf{V}_l) + \tilde{S}_{k,l}^2 \end{aligned} \right), \end{aligned} \quad (5.154)$$

where

$$\gamma \nabla \cdot \mathbf{V}_l = \gamma_r \frac{\partial u_l}{\partial r} + \gamma_\theta \frac{1}{r^\kappa} \left( \frac{\partial v_l}{\partial \theta} + \kappa u_l \right) + \gamma_z \frac{\partial w_l}{\partial z}, \quad (5.155)$$

$$\nabla \cdot \mathbf{V}_l = \frac{\partial u_l}{\partial r} + \frac{1}{r^\kappa} \left( \frac{\partial v_l}{\partial \theta} + \kappa u_l \right) + \frac{\partial w_l}{\partial z}, \quad (5.156)$$

$$\begin{aligned} \tilde{S}_{k,l}^2 = & \left[ \frac{\partial v_l}{\partial r} + \frac{1}{r^\kappa} \left( \frac{\partial u_l}{\partial \theta} - \kappa v_l \right) \right] \left[ \gamma_r \frac{\partial v_l}{\partial r} + \gamma_\theta \frac{1}{r^\kappa} \left( \frac{\partial u_l}{\partial \theta} - \kappa v_l \right) \right] \\ & + \left( \frac{\partial w_l}{\partial r} + \frac{\partial u_l}{\partial z} \right) \left( \gamma_r \frac{\partial w_l}{\partial r} + \gamma_z \frac{\partial u_l}{\partial z} \right) \\ & + \left( \frac{1}{r^\kappa} \frac{\partial w_l}{\partial \theta} + \frac{\partial v_l}{\partial z} \right) \left( \gamma_\theta \frac{1}{r^\kappa} \frac{\partial w_l}{\partial \theta} + \gamma_z \frac{\partial v_l}{\partial z} \right). \end{aligned} \quad (5.157)$$

Here the *Stokes* hypothesis is used. For single-phase flow,  $\alpha_l = 1$ , in free three-dimensional space,  $\gamma = 1$ , the above equation then reducing to the form obtained for the first time by *Rayleigh*. Note that in a turbulent pipe flow in the viscous boundary layer  $\varepsilon'_l = 0$  and  $\varepsilon_{\eta,l} > 0$ . Outside the boundary layer for relatively flat velocity profiles  $\varepsilon'_l > 0$  and  $\varepsilon_{\eta,l} \rightarrow 0$

The specific irreversibly dissipated power per unit viscous fluid mass due to turbulent pulsations

$$\varepsilon_l = \varepsilon_{\eta,l} + \varepsilon'_l \quad (5.158)$$

is used as important dependent variable characterizing the turbulence in the field. It is subject of model description. This power is considered to be constantly removed from the specific turbulent kinetic energy per unit mass of the flow field defined as follows

$$k_l = \frac{1}{2} (u_l'^2 + v_l'^2 + w_l'^2). \quad (5.159)$$

This is the second dependent variable for the velocity field which is also a subject of modeling.

In fact, Eq. (5.152) is the definition equation for the viscous dissipation rate,  $\varepsilon_l$  of the turbulent kinetic energy  $k_l$ . Here it is evident that  $\varepsilon_l$  is

- (a) a non-negative quadratic form,  $\varepsilon_l \geq 0$ ,
- (b) its mathematical description does not depend on the rotation of the coordinate system, and
- (c) it contains no derivatives of the viscosity, compare with *Zierep* (1983) for single phase flow.

The term

$$\gamma_v \alpha_l \rho_l P_{k,l} = \alpha_l^e \gamma \cdot (\mathbf{T}' : \nabla \cdot \mathbf{V}_l) = \alpha_l^e \left[ \nabla \gamma \cdot (\mathbf{T}' \cdot \mathbf{V}_l) - \mathbf{V}_l \cdot (\nabla \gamma \cdot \mathbf{T}') \right]$$

$$= \alpha_i^e \left[ \begin{array}{l} \tau'_{xx} \gamma_x \frac{\partial u}{\partial x} + \tau'_{yy} \gamma_y \frac{\partial v}{\partial y} + \tau'_{zz} \gamma_z \frac{\partial w}{\partial z} \\ + \tau'_{xy} \left( \gamma_y \frac{\partial u}{\partial y} + \gamma_x \frac{\partial v}{\partial x} \right) + \tau'_{zx} \left( \gamma_z \frac{\partial u}{\partial z} + \gamma_x \frac{\partial w}{\partial x} \right) + \tau'_{zy} \left( \gamma_z \frac{\partial v}{\partial z} + \gamma_y \frac{\partial w}{\partial y} \right) \end{array} \right]. \quad (5.160)$$

is considered to be a generation of turbulent kinetic energy, a turbulence source term. It is removed from the energy conservation and introduced as a source term in the balance equation for the turbulent kinetic energy. Inserting the *Reynolds* stresses by using the *Boussinesq* (1877) hypothesis results in common notation for Cartesian and cylindrical coordinates

$$\begin{aligned} \gamma_v \overline{P_{k,l}} := \gamma_v \frac{\alpha_l P_{k,l}}{\alpha_l^e v_l^e} = 2 \left\{ \gamma_r \left( \frac{\partial u_l}{\partial r} \right)^2 + \gamma_\theta \left[ \frac{1}{r^k} \left( \frac{\partial v_l}{\partial \theta} + \kappa u_l \right) \right]^2 + \gamma_z \left( \frac{\partial w_l}{\partial z} \right)^2 \right\} \\ - \frac{2}{3} (\nabla \cdot \mathbf{V}_l) (\gamma \nabla \cdot \mathbf{V}_l) + \tilde{S}_{k,l}^2. \end{aligned} \quad (5.161)$$

Compare this expression with Eqs. (5.153) and (5.154) and recognize the difference.

An alternative notation of the Eq. (5.161) is given for isotropic turbulence

$$\alpha_i^e \gamma (\mathbf{T}' : \nabla \mathbf{V}_l) = -\alpha_i^e \rho_l \frac{2}{3} k_l \gamma \nabla \cdot \mathbf{V}_l + \alpha_i^e \rho_l v_l^e \tilde{S}_{k,l}^2. \quad (5.162)$$

Nothing that the pressure pulsation caused the eddies is

$$p' = \rho_l v_l'^2 = \rho_l \frac{2}{3} k_l \quad (5.163)$$

the term

$$\alpha_i \rho_l \frac{2}{3} k_l \gamma \nabla \cdot \mathbf{V}_l \equiv p dVol\text{-work} \quad (5.164)$$

is immediately recognized as the mechanical expansion or compression *p dVol*-work.

For the case of steady-state single-phase,  $\alpha_i = 1$ , incompressible ( $\nabla \cdot \mathbf{V}_l = 0$ ), isentropic, and developed flow,  $P_{k,l} + \mathcal{E}_i' = 0$ , with equal velocity gradients in all directions, *Taylor* (1935) noticed that

$$\varepsilon_i \approx 12 v_l \left( \frac{\partial w_l}{\partial z} \right)^2. \quad (5.165)$$

The modeling of turbulence in multi-phase flows constitutes an exciting challenge for theoreticians and experimental scientists. We collect useful information to this subject in Volume 3 of this monograph.

**Problem 1:** Compute the power dissipation in *laminar* flow inside a pipe.

**Solution to problem 1:** For axially symmetric flow in a pipe the Eqs. (5.163) and (5.152) reduce to

$$\begin{aligned} -\alpha_i^e \rho_i \varepsilon_{\eta,l} &= \frac{1}{Vol} \int_{Vol} \alpha_i^e (\mathbf{T}_{\eta,l} : \nabla \mathbf{V}_l) dVol = \frac{4}{\pi D_h^2 \Delta z} \int_0^{\Delta z} \int_0^{R_h} \alpha_i^e (\mathbf{T}_{\eta,l} : \nabla \mathbf{V}_l) 2\pi r dr dz \\ &= \int_0^1 \int_0^1 \alpha_i^e (\mathbf{T}_{\eta,l} : \nabla \mathbf{V}_l) d(r/R_h)^2 d(z/\Delta z) \end{aligned} \quad (5.166)$$

$$-\alpha_i^e \rho_i \varepsilon_i' = \frac{1}{Vol} \int_{Vol} \alpha_i^e (\mathbf{T}_l' : \nabla \mathbf{V}_l') dVol = \int_0^1 \int_0^1 \alpha_i^e (\mathbf{T}_l' : \nabla \mathbf{V}_l') d(r/R_h)^2 d(z/\Delta z). \quad (5.167)$$

For a single-phase flow Eq. (5.166) reduces to

$$\rho_l \varepsilon_{\eta,l} = \int_0^1 \left( \frac{dw}{dr} \right)^2 d(r/R_h)^2. \quad (5.168)$$

As an illustration of how to estimate the above integral, laminar flow in a pipe in line with the *Hagen-Poiseuille* law will now be considered, with the velocity distribution

$$w(r) = -\frac{dp}{dz} \frac{1}{4\eta} (R_h^2 - r^2), \quad (5.169)$$

$$\rho \varepsilon_k = \left( \frac{dp}{dz} \right)^2 \frac{D_h^2}{32\eta}. \quad (5.170)$$

Bearing in mind that

$$\frac{dp}{dz} = 32\eta \bar{w} / D_h^2, \quad (5.171)$$

one then obtains

$$\rho \varepsilon_k = \bar{w} \frac{-dp}{dz}, \quad (5.172)$$

i.e., the *frictional pressure loss per unit mixture volume multiplied by the averaged flow velocity gives the irreversible part of the dissipated energy due to friction per unit time and unit flow volume*. There are a number of industrial processes where this component is important, e.g., heating of liquid in a circuit due to heat dissipation from pumping, and gas flow in very long pipes.

**Problem 2:** Compute the power dissipation in *turbulent* flow inside a pipe.

**Solution to problem 2:** The predominant cross section averaged velocity is  $\bar{w}$ . The *irreversible dissipated power* per unit flow volume caused by the viscous forces is  $\rho(\varepsilon_\eta + \varepsilon')$ . The *irreversible dissipated power* per unit flow volume caused by the viscous forces *due to deformation of the mean values of the velocities* in the space is  $\rho\varepsilon_\eta = (\mathbf{T}_\eta : \nabla \cdot \mathbf{V})$ . In case of turbulence flow the second component, the *irreversibly dissipated power* per unit flow volume in the viscous fluid due to *turbulent pulsations* is  $\rho\varepsilon' = (\mathbf{T}' : \nabla \cdot \mathbf{V}')$ . The friction force per unit flow volume can be computed in this case using correlations based on the macroscopic flow,

$$f_w = \frac{\lambda_{fr}}{D_{hyd}} \frac{1}{2} \rho \bar{w}^2. \quad (5.173)$$

The power per unit flow volume needed to overcome this force is  $f_w \bar{w}$ . Therefore

$$f_w w = \rho(P_k + \varepsilon_\eta + \varepsilon') = (\mathbf{T}' : \nabla \cdot \mathbf{V}) + (\mathbf{T}_\eta : \nabla \cdot \mathbf{V}) + (\mathbf{T}' : \nabla \cdot \mathbf{V}'). \quad (5.174)$$

With other words, in turbulent flow there are two components of the mechanical energy dissipation: the viscous dissipation mainly in the viscous sub-layer where the velocity gradient possesses a maximum and the dissipation due to disappearing of microscopic eddies. In the flow considered here the following simplification can be used

$$\mathbf{T}_\eta : \nabla \cdot \mathbf{V} \approx \rho \nu \left( \frac{d\bar{w}}{dy} \right)^2, \quad (5.175)$$

$$(\mathbf{T}' : \nabla \cdot \mathbf{V}') \approx 0. \quad (5.176)$$

Within the laminar sub-layer,  $y^+ \leq 5$ , there is no turbulence energy dissipation,  $\mathbf{T}' : \nabla \cdot \mathbf{V}' = 0$ . In this region the velocity is linear function of the wall distance

$$w^+ = y^+ \quad \text{or} \quad w = y \frac{\bar{w}^2}{\nu} \frac{\lambda_{fr}}{8}, \quad \frac{dw}{dy} = \frac{\bar{w}^2}{\nu} \frac{\lambda_{fr}}{8} \quad \text{and consequently}$$

$$\mathbf{T}_\eta : \nabla \cdot \mathbf{V} \approx \rho \frac{\bar{w}^4}{\nu} \left( \frac{\lambda_{fr}}{8} \right)^2. \quad (5.177)$$

The ratio of the viscous boundary layer volume to the total flow volume is

$$\frac{\Pi \Delta z y_{lim}}{F_{flow} \Delta z} = \frac{4 y_{lim}}{D_h}. \quad (5.178)$$



Here  $\Pi$  is the wetted perimeter and  $F_{flow}$  is the flow cross section. The *irreversible dissipated power* per unit flow volume caused by the viscous forces *due to deformation of the mean values of the velocities* in the space is then approximately

$$\rho \varepsilon_\eta = \frac{\Pi \Delta z}{F_{flow} \Delta z} \int_0^{y_{lim}} (\mathbf{T} : \nabla \cdot \mathbf{V}) dy = \frac{4 y_{lim}}{D_h} \rho \frac{\bar{w}^4}{\nu} \left( \frac{\lambda_{fr}}{8} \right)^2. \quad (5.179)$$

From  $f_w w = \rho (P_k + \varepsilon_\eta + \varepsilon)$  we compute *irreversibly dissipated power* per unit flow volume in the viscous fluid due to *turbulent pulsations*

$$\rho P_k = f_w w - \rho \varepsilon_\eta - \rho \varepsilon' \quad (5.180)$$

or

$$\rho P_k = \frac{\lambda_{fr}}{D_h} \frac{1}{2} \rho \bar{w}^3 \left( 1 - y_{lim}^+ \sqrt{\frac{\lambda_{fr}}{8}} \right), \quad (5.181)$$

compare with *Chandesris et al. (2005)*. In Volume 3 I will extend this method to multiphase flows.

## 5.13 Temperature equation

The purpose of this section is to rewrite the entropy equation in terms of the field temperature and system pressure. It will be shown that the temperature and pressure changes do not depend on the absolute values of the specific component entropies, and therefore on the selection of the reference temperature and pressure to define the zero specific entropies. Important differences between mixtures of perfect and non-perfect gases will also be demonstrated.

In Chapter 3 the relationship between the field temperature,  $T_l$ , and the field properties  $(s_l, C_{il}, p)$ , Eq. (3.106), is found to be

$$c_{pl} \frac{dT_l}{T_l} - \bar{R}_l \frac{dp_l}{p_l} = ds_l - \sum_{i=2}^{i_{max}} \left( \frac{\partial s_l}{\partial C_{il}} \right)_{p, T_l, all\_C's\_except\_C_{il}} dC_{il}, \quad (5.182)$$

where

$$\left( \frac{\partial s}{\partial C_i} \right)_{p, T, all\_C's\_except\_C_i} = s_{il} - s_{l1} + \Delta s_{il}^{np}. \quad (5.183)$$

One of the mass concentrations, arbitrarily numbered with subscript 1,  $C_{1l}$ , depends on all others and is computed as all others are known,

$$C_{1l} = 1 - \sum_{i=2}^{i_{\max}} C_{il} . \quad (5.184)$$

Equation (3.86) consists of two parts. For the case of a mixture consisting of ideal gases the second part is equal to zero,

$$\Delta s_{il}^{np} = 0 , \quad (5.185)$$

this also demonstrating the meaning of the subscript  $np$ , which stands for non-perfect fluid.

The non-conservative form of the entropy equation, Eq. (5.118), is

$$\begin{aligned} & \rho_l \left[ \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) s_l \right] - \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \mathcal{N} T_l) - \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \sum_{i=1}^{i_{\max}} s_{il} D_{il}^* \nabla C_{il} \right] \right\} \\ & = \gamma_v \left[ \frac{1}{T_l} D T_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right] , \end{aligned} \quad (5.186)$$

The non-conservative form of the mass conservation for each component inside the velocity field, Eq. (1.95), is

$$\alpha_l \rho_l \left( \gamma_v \frac{\partial C_{il}}{\partial \tau} + \mathbf{V}_l \gamma \mathcal{N} C_{il} \right) - \nabla \cdot (\alpha_l \rho_l D_{il}^* \gamma \mathcal{N} C_{il}) = \gamma_v (\mu_{il} - C_{il} \mu_l) . \quad (5.187)$$

The non-conservative form of the entropy equation in terms of temperature and pressure is obtained by multiplying the  $i_{\max} - 1$  mass conservation equations (5.187) by  $s_{il} - s_{il} + \Delta s_{il}^{np}$  and subtracting them from the equation (5.186). The result is simplified by using

$$\sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) - \sum_{i=2}^{i_{\max}} (s_{il} - s_{il} + \Delta s_{il}^{np}) (\mu_{il} - C_{il} \mu_l) = - \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) , \quad (5.188)$$

namely

$$\begin{aligned} & \rho_l \left[ \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) s_l \right] - \rho_l \sum_{i=2}^{i_{\max}} (s_{il} - s_{il} + \Delta s_{il}^{np}) \left( \alpha_l \gamma_v \frac{\partial C_{il}}{\partial \tau} + \alpha_l^e \mathbf{V}_l \gamma \mathcal{N} C_{il} \right) \\ & - \frac{1}{T_l} \nabla \cdot \left[ \alpha_l^e \lambda_l^* \gamma \mathcal{N} T_l \right] - \nabla \cdot \left[ \alpha_l^e \gamma \left( \sum_{i=1}^{i_{\max}} s_{il} \rho_l D_{il}^* \nabla C_{il} \right) \right] \\ & + \sum_{i=2}^{i_{\max}} (s_{il} - s_{il} + \Delta s_{il}^{np}) \nabla \cdot (\alpha_l^e \rho_l D_{il}^* \gamma \mathcal{N} C_{il}) = \gamma_v \left[ \frac{1}{T_l} D T_l^N - \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right] . \end{aligned} \quad (5.189)$$

Further simplification is obtained by using

$$\begin{aligned}
 & \rho_l \left[ \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) s_l \right] \\
 & - \rho_l \sum_{i=2}^{i_{\max}} \left( \frac{\partial s_l}{\partial C_{il}} \right)_{p, T_l, \text{all\_}C^s \text{\_except\_}C_{il}} \left( \alpha_l \gamma_v \frac{\partial C_{il}}{\partial \tau} + \alpha_l^e \mathbf{V}_l \gamma \mathcal{N} C_{il} \right) \\
 & = \left\{ \rho_l c_{pl} \left[ \alpha_l \gamma_v \frac{\partial T_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) T_l \right] - \frac{\rho_l \bar{R}_l T_l}{p} \left[ \alpha_l \gamma_v \frac{\partial p}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) p \right] \right\} / T_l,
 \end{aligned} \tag{5.190}$$

where Eq. (3.121) is

$$\frac{\rho_l \bar{R}_l T_l}{p} = \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all\_}C^s} \right]. \tag{5.191}$$

The last two terms of the left-hand side are simplified by using Eq. (5.103)

$$\begin{aligned}
 & \nabla \cdot \left[ \alpha_l^e \gamma \left( \sum_{i=1}^{i_{\max}} s_{il} \rho_l D_{il}^* \nabla C_{il} \right) \right] - \sum_{i=2}^{i_{\max}} (s_{il} - s_{ll} + \Delta s_{il}^{np}) \nabla (\alpha_l^e \rho_l D_{il}^* \mathcal{N} C_{il}) \\
 & = \nabla \cdot \left[ \alpha_l^e \gamma \left( \sum_{i=2}^{i_{\max}} (s_{il} - s_{ll}) \rho_l D_{il}^* \nabla C_{il} \right) \right] - \sum_{i=2}^{i_{\max}} (s_{il} - s_{ll} + \Delta s_{il}^{np}) \nabla (\alpha_l^e \rho_l D_{il}^* \mathcal{N} C_{il}) \\
 & = \alpha_l^e \rho_l \gamma \sum_{i=2}^{i_{\max}} (D_{il}^* \nabla C_{il}) \nabla (s_{il} - s_{ll}) - \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \nabla (\alpha_l^e \rho_l D_{il}^* \mathcal{N} C_{il})
 \end{aligned} \tag{5.192}$$

Neglecting the second order terms

$$\alpha_l^e \rho_l \gamma \sum_{i=2}^{i_{\max}} (D_{il}^* \nabla C_{il}) \nabla (s_{il} - s_{ll}), \tag{5.193}$$

we obtain the form below that is very convenient for practical applications

$$\boxed{
 \begin{aligned}
 & \rho_l c_{pl} \left[ \alpha_l \gamma_v \frac{\partial T_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) T_l \right] \\
 & - \nabla \cdot (\alpha_l^e \lambda_l^* \mathcal{N} T) - \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all\_}C^s} \right] \left[ \alpha_l \gamma_v \frac{\partial p}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) p \right]
 \end{aligned}
 }$$

$$+T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \nabla \left( \alpha_i^e \rho_l D_{il}^* \mathcal{N} C_{il} \right) = \gamma_v \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \left( \mu_{il} - C_{il} \mu_l \right) \right]. \quad (5.194)$$

This is the entropy equation rewritten in terms of temperature and pressure. For perfect gas mixtures, the following is obtained

$$\begin{aligned} \rho_l c_{pl} \left[ \alpha_i \gamma_v \frac{\partial T_l}{\partial \tau} + \left( \alpha_i^e \mathbf{V}_l \gamma \cdot \nabla \right) T_l \right] - \nabla \cdot \left( \alpha_i^e \lambda_i^* \mathcal{N} T \right) \\ - \left[ \alpha_i \gamma_v \frac{\partial p}{\partial \tau} + \left( \alpha_i^e \mathbf{V}_l \gamma \cdot \nabla \right) p \right] = \gamma_v DT_l^N. \end{aligned} \quad (5.195)$$

### Conclusions:

- The temperature change caused by the injection into the velocity field depends on the difference between the specific enthalpy of the donor and the specific enthalpy of the velocity field. It is important to note that this is an *enthalpy difference* but not, say, differences in the specific internal energies, entropies, etc.
- Equation (5.194) does not contain any specific entropies. As a result, the temperature change described by Eq. (5.194) does not depend on the reference temperature and pressure for computation of the specific entropies for the specific components.
- Consider diffusion in a gas mixture consisting of non-perfect gases in adiabatic closed space. The initial state is characterized by spatially uniform temperature. At nearly constant system pressure Eq. (5.194) reduces to

$$\gamma_v \rho_l c_{pl} \frac{\partial T_l}{\partial \tau} \approx -T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \nabla \left( \rho_l D_{il}^* \mathcal{N} C_{il} \right), \quad (5.196)$$

which demonstrates that a temporal temperature change takes place at the locations of considerable diffusion. Bad numerical resolution can amplify this effect, leading to considerable local error in the numerical analysis.

- Injection of non-perfect gas components into a volume initially filled with perfect gas can also give rise to behavior different from that of injection of a perfect gas component, due to the differences in the temperature and pressure equations describing both cases. This will be demonstrated in a simple case in the next section.

## 5.14 Second law of the thermodynamics

The entropy equation reflects very interesting physical phenomena. It is evident that velocity gradients in continuum of viscous fluid cause energy dissipation,  $\varepsilon_{\eta,l} + \varepsilon'_l \geq 0$ , and may generate turbulent kinetic energy  $P_{k,l} \geq 0$ . The turbulent kinetic energy increases the turbulent viscosity according to the *Prandtl-Kolmogorov* law, and helps to reduce the velocity gradients. The irreversible

dissipation of kinetic energy caused by the turbulent pulsation increases the specific internal energy of the continuum field,  $\varepsilon'_i \geq 0$ . This dissipation decreases the specific turbulent kinetic energy directly. The same action is caused by the direct viscous dissipation  $\varepsilon_{\eta,i}$ .

If the equation is applied to a single velocity field in a closed system without interaction with external mass, momentum or energy sources, the change in the specific entropy of the system will be non-negative, as the sum of the dissipation terms,  $\varepsilon_{\eta,i} + \varepsilon'_i$ , is non-negative. This expresses the *second law of thermodynamics*. The second law tells us in what direction a process will develop in nature for closed and isolated systems.

The process will proceed in a direction such that the entropy of the system always increases, or at best stays the same,  $\varepsilon_{\eta,i} + \varepsilon'_i = 0$ , - entropy principle.

This information is not contained in the first law of thermodynamics. It results only after combining the three conservation principles (mass, momentum and energy) and introducing a *Legendre* transformation in the form of a *Gibbs* equation. In a way, it is a general expression of these conservation principles.

The entropy equation is not only very informative, but as already mentioned, it is very convenient for numerical integration because of its *simplicity* compared to the primitive form of the energy principle. This is the reason why the *specific entropies* of the velocity fields together with the concentrations of the inert component  $C_{ni}$  were chosen for use as *components of the dependent variables* vector as already mentioned in *Kolev (1994a)*. This unique combination of the dependent variables *simply minimizes the computational effort* associated with numerical integration and therefore makes the computer code *faster* and the analysis *cheaper*. It also makes the computer code architecture simple and allows the inclusion of more physical phenomena within a general flow model.

I call the flow modeling concept which makes use of the specific entropies of the velocity fields as components of the dependent variables vector the *entropy concept*.

## 5.15 Mixture volume conservation equation

Any numerical method in fluid mechanics should provide correct coupling between pressure changes and velocity changes. There are different ways of achieving this. One of these is faster than the others. Probably the fastest of these methods uses a specially derived equation, referred to as here the *mixture volume conservation equation*, MVCE. The MVCE is a linear combination of the mass

conservation equations. The purpose of this section is to derive the analytical form of the MVCE, to discuss the physical meaning of each separate term and finally to show the reasons that make this equation so appropriate for use in constructing numerical schemes for complicated multi-phase flows.

The MVCE was obtained as follows:

- (a) The mass conservation equations (1.62) were differentiated using the chain rule.
- (b) Each equation was divided by the associated density. The resulting equations are dimensionless,  $m^3/m^3$ , and reflect the volume change balance for each velocity field per unit time and per unit mixture volume.

$$\gamma_v \frac{\partial \alpha_l}{\partial \tau} + \alpha_l \frac{\partial \gamma_v}{\partial \tau} + \frac{1}{\rho_l} \left[ \alpha_l \gamma_v \frac{\partial \rho_l}{\partial \tau} + (\alpha_l \mathbf{V}_l \gamma \cdot \nabla) \rho_l \right] + \nabla \cdot (\alpha_l \mathbf{V}_l \gamma) = \gamma_v \frac{\mu_l}{\rho_l} \quad (5.197)$$

- (c) The volume conservation equations obtained in this way were added and the fact that

$$\sum_{l=1}^3 \alpha_l = 1 \quad (5.198)$$

and

$$\sum_{l=1}^3 d\alpha_l = 0 \quad (5.199)$$

was used to cancel the sum of the time derivatives of the volume concentrations.

$$\sum_{l=1}^{l_{\max}} \frac{1}{\rho_l} \left[ \alpha_l \gamma_v \frac{\partial \rho_l}{\partial \tau} + (\alpha_l \mathbf{V}_l \gamma \cdot \nabla) \rho_l \right] + \nabla \cdot \sum_{l=1}^{l_{\max}} (\alpha_l \mathbf{V}_l \gamma) = \gamma_v \sum_{l=1}^{l_{\max}} \frac{\mu_l}{\rho_l} - \frac{\partial \gamma_v}{\partial \tau} \quad (5.200)$$

- (d) The density derivatives were substituted using the differential form of the equation of state for each velocity field (3.137)

$$d\rho_l = \frac{dp_l}{a_l^2} + \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, \text{all } C_i^s} ds_l + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{li}} \right)_{p, s, \text{all } C_i^s \text{ except } C_{li}} dC_{li} \quad (5.201)$$

The result is

$$\gamma_v \left( \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} \frac{\partial p_l}{\partial \tau} \right) + \sum_{l=1}^{l_{\max}} \left[ \frac{\alpha_l}{\rho_l a_l^2} (\mathbf{V}_l \gamma \cdot \nabla) p_l \right] + \nabla \cdot \sum_{l=1}^{l_{\max}} (\alpha_l \mathbf{V}_l \gamma)$$

$$\begin{aligned}
&= \sum_{l=1}^{i_{\max}} \frac{1}{\rho_l} \left\{ \gamma_v \mu_l - \frac{1}{\rho_l} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, \text{all\_}C'_l s} \rho_l \left[ \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + (\alpha_l \mathbf{V}_l \gamma \cdot \nabla) s_l \right] \right. \right. \\
&\quad \left. \left. + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{li}} \right)_{p, s, \text{all\_}C'_s \text{ \_except\_} C_i} \rho_l \left[ \alpha_l \gamma_v \frac{\partial C_{li}}{\partial \tau} + (\alpha_l \mathbf{V}_l \gamma \cdot \nabla) C_{li} \right] \right] \right\} \\
&\quad - \frac{\partial \gamma_v}{\partial \tau} \tag{5.202}
\end{aligned}$$

(e) The concentration equations (1.96) in Chapter 1 and the non-conservative form of the entropy equation (5.118) in this Chapter are compared with the LHS of Eq. (5.202). Making the substitution

$$\rho_l \left[ \alpha_l \gamma_v \frac{\partial C_{il}}{\partial \tau} + (\alpha_l \mathbf{V}_l \gamma \cdot \nabla) C_{il} \right] = DC_{il}^N, \tag{5.203}$$

$$\rho_l \left[ \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + (\alpha_l \mathbf{V}_l \gamma \cdot \nabla) s_l \right] = Ds_l^N, \tag{5.204}$$

where

$$\overline{DC_{il}^N} = \nabla (\alpha_l \rho_l D_{il}^* \mathcal{N} C_{il}) + \gamma_v (\mu_{il} - C_{il} \mu_l), \tag{5.205}$$

$$\begin{aligned}
\overline{Ds_l^N} &= \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \mathcal{N} T_l) + \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \sum_{i=1}^{i_{\max}} s_{il} D_{il}^* \nabla C_{il} \right] \right\} \\
&+ \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right], \tag{5.206}
\end{aligned}$$

we obtain the final form for the MVCE

$$\gamma_v \left( \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} \frac{\partial p_l}{\partial \tau} \right) + \sum_{l=1}^{l_{\max}} \left[ \frac{\alpha_l}{\rho_l a_l^2} (\mathbf{V}_l \gamma \cdot \nabla) p_l \right] + \nabla \cdot \sum_{l=1}^{l_{\max}} (\alpha_l \mathbf{V}_l \gamma) = \sum_{l=1}^{l_{\max}} D\alpha_l - \frac{\partial \gamma_v}{\partial \tau}$$

which for negligible differences in the field pressures is

$$\boxed{ \frac{\gamma_v}{\rho a^2} \frac{\partial p}{\partial \tau} + \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} (\mathbf{V}_l \gamma \cdot \nabla) p + \nabla \cdot \sum_{l=1}^{l_{\max}} (\alpha_l \mathbf{V}_l \gamma) = \sum_{l=1}^{l_{\max}} D\alpha_l - \frac{\partial \gamma_v}{\partial \tau}, \tag{5.207} }$$

or in scalar form for practical application in Cartesian and cylindrical coordinates

$$\frac{\gamma_v}{\rho a^2} \frac{\partial p}{\partial \tau} + \sum_{l=1}^3 \frac{\alpha_l}{\rho_l a_l^2} \left( u_l \gamma_r \frac{\partial p}{\partial r} + v_l \gamma_\theta \frac{\partial p}{\partial \theta} + w_l \gamma_z \frac{\partial p}{\partial z} \right)$$

$$+ \frac{\partial}{\partial r} (\alpha_l \mu_l \gamma_r) + \frac{1}{r^\kappa} \frac{\partial}{\partial \theta} (r^\kappa \alpha_l \nu_l \gamma_\theta) + \frac{\partial}{\partial z} (\alpha_l w_l \gamma_z) = \sum_{l=1}^3 D\alpha_l - \frac{\partial \gamma_v}{\partial \tau} \quad (5.208)$$

where

$$D\alpha_l = \frac{1}{\rho_l} \left\{ \gamma_v \mu_l - \frac{1}{\rho_l} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, \text{all}_{-C'_i s}} \overline{Ds_l^N} + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{il}} \right)_{p, s, \text{all}_{-C'_i s_{\text{except}}_{-C_{il}}}} \overline{DC_{il}^N} \right] \right\} \quad (5.209)$$

Replacing the derivatives of the mixture density with the Eqs. (110) and (111) obtained in *Kolev* (1991), one obtains

$$D\alpha_l = \gamma_v \frac{\mu_l}{\rho_l} - \frac{1}{\rho_l^2} \left[ \left( \frac{\partial \rho_l}{\partial T_l} \right)_{p, \text{all}_{-C's}} \frac{T_l}{c_{pl}} \left[ \overline{Ds_l^N} - \sum_{i=2}^{i_{\max}} \left( \frac{\partial s_l}{\partial C_{il}} \right)_{p, T_l, \text{all}_{-C's_{\text{except}}_{-C_{il}}}} \overline{DC_{il}^N} \right] + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{il}} \right)_{p, T_l, \text{all}_{-C's_{\text{except}}_{-C_{il}}}} \overline{DC_{il}^N} \right]. \quad (5.210)$$

Further simplification is obtained by using Eqs. (5.205) and (5.206), subsequently the chain rules and neglecting the second order term

$$\alpha_l \rho_l \gamma T_l \sum_{i=2}^{i_{\max}} \nabla (D_{il}^* \nabla C_{il}) \nabla (s_{il} - s_{il}). \quad (5.211)$$

The result is

$$\begin{aligned} & T_l \left[ \overline{Ds_l^N} - \sum_{i=2}^{i_{\max}} \left( \frac{\partial s_l}{\partial C_{il}} \right)_{p, T_l, \text{all}_{-C's_{\text{except}}_{-C_{il}}}} \overline{DC_{il}^N} \right] \\ & \approx \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \mathcal{N} T) + \gamma_v D T_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \left[ \nabla (\alpha_l^e \rho_l D_{il}^* \gamma \mathcal{N} C_{il}) + \gamma_v (\mu_{il} - C_{il} \mu_l) \right], \end{aligned} \quad (5.212)$$

For perfect gas mixtures the following alone is obtained

$$T_l \left[ \overline{Ds_l^N} - \sum_{i=2}^{i_{\max}} \left( \frac{\partial s_l}{\partial C_{il}} \right)_{p, T_l, \text{all}_{-C's_{\text{except}}_{-C_{il}}}} \overline{DC_{il}^N} \right] \approx \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \mathcal{N} T) + \gamma_v \overline{DT_l^N} \quad (5.213)$$

$a$  is the *sonic velocity* in a homogeneous multi-phase mixture and is defined as follows



$$\frac{1}{\rho a^2} = \sum_{l=1}^3 \frac{\alpha_l}{\rho_l a_l^2} = \frac{1}{p} \sum_{l=1}^3 \frac{\alpha_l}{\kappa_l} = \frac{1}{\kappa p}, \quad (5.214)$$

and

$$\rho = \sum_{l=1}^3 \alpha_l \rho_l \quad (5.215)$$

is the mixture density. Equation (5.207) was already derived in *Kolev* (1986) and published in *Kolev* (1987) p.100.

The MVCE equation can be directly discretized and incorporated into the numerical scheme. Another possibility is to follow the same scheme as for deriving the MVCE analytically but starting with already discretized mass conservation equations. The coupling finally obtained is then *strictly consistent* with the discretized form of the mass conservation equations.

Alternative forms of the MVCE can also be used, e.g.

$$\frac{\gamma_v}{\kappa p} \frac{\partial p}{\partial \tau} + \sum_{l=1}^3 \frac{\alpha_l}{\kappa_l p} \mathbf{V}_l \gamma \cdot \nabla p + \nabla \cdot \left( \sum_{l=1}^3 \alpha_l \mathbf{V}_l \gamma \right) = \sum_{l=1}^3 D \alpha_l - \frac{\partial \gamma_v}{\partial \tau}, \quad (5.216)$$

or certain integrated forms

$$\frac{\gamma_v}{\kappa} \frac{\partial}{\partial \tau} \ln p + \sum_{l=1}^3 \frac{\alpha_l}{\kappa_l} \mathbf{V}_l \gamma \cdot \nabla \ln p + \nabla \cdot \left( \sum_{l=1}^3 \alpha_l \mathbf{V}_l \gamma \right) = \sum_{l=1}^3 D \alpha_l - \frac{\partial \gamma_v}{\partial \tau}, \quad (5.217)$$

where

$$\kappa = \rho a^2 / p, \quad (5.218)$$

$$\kappa_l = \rho_l a_l^2 / p \quad (5.219)$$

are the mixture isentropic exponent and the isentropic exponent of each particular velocity field, respectively.

For completeness of the theory the MVCE equation valid in the case of steady-state non-compressible flow will also be given

$$\nabla \cdot \left( \sum_{l=1}^3 \alpha_l \mathbf{V}_l \gamma \right) = \nabla \cdot (\mathbf{J} \gamma) = \sum_{l=1}^3 D \alpha_l, \quad (5.220)$$

where

$$\mathbf{J} = \sum_{l=1}^3 \alpha_l \mathbf{V}_l = \sum_{l=1}^3 j_l \quad (5.221)$$

is the *volumetric mixture flow rate* in  $m^3/(m^2s)$ . If diffusion is neglected and no mass exchange takes place between the velocity fields or between the flow and external sources, this gives

$$\nabla \cdot (\mathbf{J} \gamma) = 0 \quad (5.222)$$

or

$$\mathbf{J}\gamma = \text{const.} \quad (5.223)$$

The MVCE has the remarkable feature that it couples the temporal pressure change through the compressibility  $1/(\rho a^2)$  with

- a) the convective specific volume change  $\nabla \cdot \left( \sum_{i=1}^3 \alpha_i \mathbf{V}_i \gamma \right)$  for the control volume;
- b) the change in the specific volume of the mixture associated with the net specific volume change for the mixture due to the mass sources  $\gamma_v \sum_{i=1}^3 \mu_i / \rho_i$ ;
- c) the density change due to the spatial pressure, entropy and concentration changes for the associated velocity field.

Another specific property of Eq. (5.207) should be emphasized. The second term on the left-hand side represents the dimensionless changes in the density. This fact allows one to use *up-wind* discretization even for the pressure terms (*donor-cell* concept), because one in fact discretizes the dimensionless density change within the interval  $\Delta r, \Delta \theta, \Delta z$ .

The above features of this equation make it very suitable for coupling with the momentum equations in order to derive an equation for the mixtures which is similar to the *Poisson* equation for single-phase flow.

For the case of negligible diffusion the right-hand side of the MVCE contains no differential terms:

$$D\alpha_i = \frac{\gamma_v}{\rho_i} \left\{ \mu_i - \left[ \frac{\partial \rho_i}{\partial s_i} (Ds_i - \mu_i^+ s_i) + \sum_{i=1}^{i_{\max}} \frac{\partial \rho_i}{\partial C_{ii}} (\mu_{ii} - \mu_{ii}^+ C_{ii}) \right] / \rho_i \right\}. \quad (5.224)$$

This means that during numerical integration the influence of the changes of entropies and concentrations on the pressure field can be taken into account in a single step only (without outer iterations). The computer code architecture is thus extremely simplified, speeding up the numerical integration and therefore making it cheaper. This is an important feature of the entropy concept presented here and used in the IVA computer code development compared to the concepts of other computer codes, e.g. TRAC *Liles et al.* (1978, 1981), COBRA-TF *Kelly and Kohrt* (1983), COBRA/TRAC *Thurgood et al.* (1983) and AFDM *Bohl et al.* (1988).

The right-hand side of the pressure equation, Eq. (5.207), does not contain any specific entropy. As a result, the pressure change described by Eq. (5.207) is not dependent on the reference temperature and pressure for computation of the specific entropies for the specific components.

Comparing Eqs. (5.212) and (5.213) it is evident that if systems in which at least one of the gas components deviates from the perfect gas behavior are approximated as consisting of perfect gasses, an entropy imbalance of about

$$-\sum_{i=2}^{i_{\max}} \Delta s_{ii}^{np} \left[ \nabla \cdot (\alpha_i^e \rho_i D_{ii}^* \gamma \mathcal{N} C_{ii}) + \gamma_v (\mu_{ii} - C_{ii} \mu_i) \right] \quad (5.225)$$

is introduced. This is a very surprising result. I call this result the *non-perfect gas paradox*.

## 5.16 Linearized form of the source term for the temperature equation

The source terms with fluctuation components neglected

$$\begin{aligned}
 DT_l^N = & \alpha_l \rho_l (\delta_l P_{\eta,l}^r + \varepsilon_l) + \dot{q}'' + \sum_{i=1}^{i_{\max}} \mu_{iwl} (h_{iwl} - h_{il}) \\
 & + \sum_{m=1}^3 \left[ \mu_{Mml} (h_{Ml}^\sigma - h_{Ml}) + \sum_{n=1}^{n_{\max}} \mu_{nml} (h_{nm} - h_{nl}) \right] \quad (5.226)
 \end{aligned}$$

will be next rewritten for each velocity field in a form that allows the use of implicit integration schemes. The conditions governing heat and mass transfer can change during the time step, thereby influencing the averaged rate for the transported quantity. This feedback effect during a single time step may be crucial in the case of a number of applications associated with strong heat and mass transfer processes. The first work known to this author formalizing source terms is those by *Solbrig et al.* published in 1977. *Solbrig et al.* considered mass and energy equations in a closed system for two single component phases. The meta-stable state of the phases was not allowed in their work. Instead the pressure depending mass sources are defined adjusting the state of each phase to the saturation with the changing pressure. But even the meta-stable state is the driving force for spontaneous evaporation and condensation, and therefore has to be taken into account as we do in our analysis.

The following definitions and assumptions are used here:

$$c_{pw2} = \sum_{i=1}^{i_{\max}} C_{iw2} c_{piw2} \quad (5.227)$$

$$c_{pw3} = \sum_{i=1}^{i_{\max}} C_{iw3} c_{piw3} \quad (5.228)$$

$$c_{p2} = \sum_{i=1}^{i_{\max}} C_{i2} c_{pi2} \quad (5.229)$$

$$c_{p3} = \sum_{i=1}^{i_{\max}} C_{i3} c_{pi3} \quad (5.230)$$

$$h_{M1}^{2\sigma} - h_{M1} \approx c_{pM1} (T_{M1}^{2\sigma} - T_1) \quad \text{and} \quad p_{M1}^{2\sigma} \approx p_{M1} \quad (5.231)$$

$$h_{M1}^{3\sigma} - h_{M1} \approx c_{pM1} (T_{M1}^{3\sigma} - T_1) \quad \text{and} \quad p_{M1}^{3\sigma} \approx p_{M1}, \quad (5.232)$$

$$h_{M2}^{1\sigma} - h_{M2} \approx c_{pM2} (T_2^{1\sigma} - T_2), \quad (5.233)$$

$$h_{M3}^{1\sigma} - h_{M3} \approx c_{pM3} (T_3^{1\sigma} - T_3), \quad (5.234)$$

$$C_{M2} (h_{M2} - h_{M3}) + \sum_{n=1}^{n_{\max}} C_{n2} (h_{n2} - h_{n3}) \approx c_{p2} (T_2 - T_3), \quad (5.235)$$

$$C_{M3} (h_{M3} - h_{M2}) + \sum_{n=1}^{n_{\max}} C_{n3} (h_{n3} - h_{n2}) \approx c_{p3} (T_3 - T_2), \quad (5.236)$$

$$\sum_{i=1}^{i_{\max}} C_{i2} (h_{iw2} - h_{i2}) \approx c_{pw2} (T_w - T_2), \quad (5.237)$$

$$\sum_{i=1}^{i_{\max}} C_{i3} (h_{iw3} - h_{i3}) \approx c_{pw3} (T_w - T_3). \quad (5.238)$$

It is assumed that the properties of the mass leaving the velocity field are equal to the properties of this donor field. Here I introduce the product of the effective heat transfer coefficient and the interfacial area density and designate this by  $\chi_l^{m\sigma}$ . The subscript  $l$  designates the location inside the velocity field  $l$  and the superscript  $m\sigma$  designates the location at the interface  $\sigma$  dividing field  $m$  from field  $l$ . Superscripts are only used if the interfacial heat transfer is associated with mass transfer. If there is heat transfer only, the linearized interaction coefficient is assigned the subscript  $ml$  only, this indicating the interface at which the heat transfer takes place. These rules are valid also for the wall. Actually, the wall is treated as an additional field. The result is

$$\begin{aligned} DT_1^N &= \alpha_1 \rho_1 (\delta_1 P_{k1}^r + \varepsilon_1) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1) + (\mu_{21}^w + \mu_{31}^w) c_{pM1} (T_{M1}^{w\sigma} - T_1) \\ &+ \chi_{w1} (T_w - T_1) + \chi_{21} (T_2 - T_1) + \chi_{31} (T_3 - T_1) + \chi_1^{2\sigma} (T_1^{2\sigma} - T_1) + \chi_1^{3\sigma} (T_1^{3\sigma} - T_1) \\ &+ \sum_{i=1}^{i_{\max}} \mu_{iw1} (h_{iw1} - h_{i1}) + c_{pM1} [\mu_{21} (T_1^{2\sigma} - T_1) + \mu_{31} (T_1^{3\sigma} - T_1)], \end{aligned} \quad (5.239)$$

$$\begin{aligned} DT_2^N &= \alpha_2 \rho_2 (\delta_2 P_{k2}^r + \varepsilon_2) + \chi_2^{w\sigma} (T_2^{w\sigma} - T_2) + \mu_{12}^w c_{pM2} (T_2^{w\sigma} - T_2) \\ &+ (\chi_{w2} + \mu_{w2} c_{pw2}) (T_w - T_2) + (\chi_{32} + \mu_{32} c_{p3}) (T_3 - T_2) - \chi_{21} (T_2 - T_1) \\ &+ \chi_2^{1\sigma} (T_2^{1\sigma} - T_2) + \mu_{12} c_{pM2} (T_2^{1\sigma} - T_2), \end{aligned} \quad (5.240)$$

$$\begin{aligned}
DT_3^N &= \alpha_3 \rho_3 (\delta_3 P_{k3}^\tau + \varepsilon_3) + \chi_3^{w\sigma} (T_3^{w\sigma} - T_3) + \mu_{13}^w c_{pM3} (T_3^{w\sigma} - T_3) \\
&+ (\chi_{w3} + \mu_{w3} c_{pw3}) (T_w - T_3) - (\chi_{32} + \mu_{23} c_{p2}) (T_3 - T_2) - \chi_{31} (T_3 - T_1) \\
&+ \chi_3^{1\sigma} (T_3^{1\sigma} - T_3) + \mu_{13} c_{pM3} (T_3^{1\sigma} - T_3). \tag{5.241}
\end{aligned}$$

The energy jump condition at the interfaces yields the following for the *condensation* sources

$$\mu_{12} = \psi_{12} \frac{\chi_2^{1\sigma} (T_2^{1\sigma} - T_2) + \chi_1^{2\sigma} (T_1^{2\sigma} - T_1)}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}} \geq 0, \tag{5.242}$$

$$\mu_{13} = \psi_{13} \frac{\chi_3^{1\sigma} (T_3^{1\sigma} - T_3) + \chi_1^{3\sigma} (T_1^{3\sigma} - T_1)}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}} \geq 0, \tag{5.243}$$

for  $C_{M1} > 0$ . Here the integer switches

$$\psi_{12} = \frac{1}{2} \left\{ 1 + \text{sign} \left[ \chi_2^{1\sigma} (T_2^{1\sigma} - T_2) + \chi_1^{2\sigma} (T_1^{2\sigma} - T_1) \right] \right\}, \tag{5.244}$$

$$\psi_{13} = \frac{1}{2} \left\{ 1 + \text{sign} \left[ \chi_3^{1\sigma} (T_3^{1\sigma} - T_3) + \chi_1^{3\sigma} (T_1^{3\sigma} - T_1) \right] \right\} \tag{5.245}$$

guarantee that the mass sources are non-negative. For  $C_{M1} = 0$  we have  $\mu_{12} = 0$ , and  $\mu_{13} = 0$  because there is nothing to condense. The energy jump condition at the interfaces yields the following for the *evaporation* sources:

$$\mu_{21} = -(1 - \psi_{12}) \frac{\chi_2^{1\sigma} (T_2^{1\sigma} - T_2) + \chi_1^{2\sigma} (T_1^{2\sigma} - T_1)}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}} \geq 0, \text{ for } C_{M2} > 0, \tag{5.246}$$

$$\mu_{31} = -(1 - \psi_{13}) \frac{\chi_3^{1\sigma} (T_3^{1\sigma} - T_3) + \chi_1^{3\sigma} (T_1^{3\sigma} - T_1)}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}} \geq 0, \text{ for } C_{M3} > 0. \tag{5.247}$$

For  $C_{M2} = 0$  we have  $\mu_{21} = 0$  because there is nothing to evaporate. Similarly for  $C_{M3} = 0$  we have  $\mu_{31} = 0$ .

The mass transfer at heated or cooled walls is defined similarly:

*Condensation:*

$$\mu_{12}^w = \psi_{12}^w f_{\text{cond} \rightarrow \text{film}} \frac{\chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1)}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}} \geq 0, \tag{5.248}$$

$$\mu_{13}^w = \psi_{13}^w \left(1 - f_{cond \rightarrow film}\right) \frac{\chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1)}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}} \geq 0. \quad (5.249)$$

Here we introduce the factor  $f_{cond \rightarrow film}$  which determines how much of the condensed vapor is going into the film. For  $f_{cond \rightarrow film} = 1$  we have film condensation at the wall. For  $f_{cond \rightarrow film} = 0$  we have droplet condensation at the wall. The decision which of the mechanisms is active depends on the wettability of the surface and its orientation in space.

*Evaporation:*

$$\begin{aligned} \mu_{21}^w &= -\psi_{21}^w \frac{\chi_w^{2\sigma} (T_w^{2\sigma} - T_w) + \chi_2^{w\sigma} (T_2^{w\sigma} - T_2)}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}} \\ &= -\psi_{21}^w \frac{\chi_w^{2\sigma} [T'(p) - T_w] + \chi_2^{w\sigma} [T'(p) - T_2]}{h''(p) - h'(p)} \geq 0, \text{ for } C_{M2} > 0, \end{aligned} \quad (5.250)$$

$$\begin{aligned} \mu_{31}^w &= -\psi_{31}^w \frac{\chi_w^{3\sigma} (T_w^{3\sigma} - T_w) + \chi_3^{w\sigma} (T_3^{w\sigma} - T_3)}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}} \\ &= -\psi_{31}^w \frac{\chi_w^{3\sigma} [T'(p) - T_w] + \chi_3^{w\sigma} [T'(p) - T_3]}{h''(p) - h'(p)} \geq 0, \text{ for } C_{M3} > 0. \end{aligned} \quad (5.251)$$

In deriving Eqs. (5.231) and (5.232) the following assumptions are made

$$T_w^{2\sigma}, T_2^{w\sigma}, T_w^{3\sigma}, T_3^{w\sigma} = T'(p), \quad (5.252)$$

$$h_{M1}^{2\sigma}, h_{M1}^{3\sigma} = h''(p), \quad (5.253)$$

and

$$h_{M2}^{1\sigma}, h_{M3}^{1\sigma} = h'(p). \quad (5.254)$$

Here again the integer switches

$$\psi_{21}^w = \frac{1}{2} \left\{ 1 - \text{sign} \left( \chi_w^{2\sigma} [T'(p) - T_w] + \chi_2^{w\sigma} [T'(p) - T_2] \right) \right\}, \quad (5.255)$$

$$\psi_{31}^w = \frac{1}{2} \left\{ 1 - \text{sign} \left( \chi_w^{3\sigma} [T'(p) - T_w] + \chi_3^{w\sigma} [T'(p) - T_3] \right) \right\}, \quad (5.256)$$

$$\psi_{12}^w = \frac{1}{2} \left\{ 1 + \text{sign} \left[ \chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1) \right] \right\}, \quad (5.257)$$

and conditions for film condensation otherwise  $\psi_{12}^w = 0$ ,

$$\psi_{13}^w = \frac{1}{2} \left\{ 1 + \text{sign} \left[ \chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1) \right] \right\} \quad (5.258)$$

and conditions for droplet condensation otherwise  $\psi_{13}^w = 0$ , guarantee that the mass sources are non-negative.

Substituting for the evaporation and condensation mass sources one obtains

$$\begin{aligned} DT_1^N &= \alpha_1 \rho_1 (\delta_1 P_{k1}^\tau + \varepsilon_1) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1) \\ &- \left( \begin{aligned} &\psi_{21}^w \left\{ \chi_w^{2\sigma} [T'(p) - T_w] + \chi_2^{w\sigma} [T'(p) - T_2] \right\} \\ &+ \psi_{31}^w \left\{ \chi_w^{3\sigma} [T'(p) - T_w] + \chi_3^{w\sigma} [T'(p) - T_3] \right\} \end{aligned} \right) \frac{c_{pM1}}{h''(p) - h'(p)} (T_1^{w\sigma} - T_1) \\ &+ \chi_{w1} (T_w - T_1) + \chi_{21} (T_2 - T_1) + \chi_{31} (T_3 - T_1) + \sum_{i=1}^{i_{\max}} \mu_{iw1} (h_{iw1} - h_{i1}) \\ &+ \chi_1^{2\sigma} (T_1^{2\sigma} - T_1) - (1 - \psi_{12}) \left[ \chi_2^{1\sigma} (T_2^{1\sigma} - T_2) + \chi_1^{2\sigma} (T_1^{2\sigma} - T_1) \right] \frac{c_{pM1} (T_1^{2\sigma} - T_1)}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}} \\ &+ \chi_1^{3\sigma} (T_1^{3\sigma} - T_1) - (1 - \psi_{13}) \left[ \chi_3^{1\sigma} (T_3^{1\sigma} - T_3) + \chi_1^{3\sigma} (T_1^{3\sigma} - T_1) \right] \frac{c_{pM1} (T_1^{3\sigma} - T_1)}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}}, \end{aligned} \quad (5.259)$$

$$\begin{aligned} DT_2^N &= \alpha_2 \rho_2 (\delta_2 P_{k2}^\tau + \varepsilon_2) + \chi_2^{w\sigma} (T_2^{w\sigma} - T_2) \\ &+ \psi_{12}^w f_{\text{cond} \rightarrow \text{film}} \left[ \chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1) \right] \frac{c_{pM2}}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}} (T_2^{w\sigma} - T_2) \\ &+ (\chi_{w2} + \mu_{w2} c_{pw2}) (T_w - T_2) + (\chi_{32} + \mu_{32} c_{p3}) (T_3 - T_2) - \chi_{21} (T_2 - T_1) \\ &+ \chi_2^{1\sigma} (T_2^{1\sigma} - T_2) + \psi_{12} \left[ \chi_2^{1\sigma} (T_2^{1\sigma} - T_2) + \chi_1^{2\sigma} (T_1^{2\sigma} - T_1) \right] \frac{c_{pM2} (T_2^{1\sigma} - T_2)}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}}, \end{aligned} \quad (5.260)$$

$$\begin{aligned}
DT_3^N &= \alpha_3 \rho_3 (\delta_3 P_{k3}^\tau + \varepsilon_3) + \chi_3^{w\sigma} (T_3^{w\sigma} - T_3) \\
&+ \psi_{13}^w (1 - f_{cond \rightarrow film}) \left[ \chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1) \right] \frac{c_{pM3}}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}} (T_3^{w\sigma} - T_3) \\
&+ (\chi_{w3} + \mu_{w3} c_{pw3}) (T_w - T_3) - (\chi_{32} + \mu_{23} c_{p2}) (T_3 - T_2) - \chi_{31} (T_3 - T_1) \\
&+ \chi_3^{1\sigma} (T_3^{1\sigma} - T_3) + \psi_{13} \left[ \chi_3^{1\sigma} (T_3^{1\sigma} - T_3) + \chi_1^{3\sigma} (T_1^{3\sigma} - T_1) \right] \frac{c_{pM3} (T_3^{1\sigma} - T_3)}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}}.
\end{aligned} \tag{5.261}$$

The above expressions can be rewritten in the compact form

$$DT_i^N = c_i^T - \sum_{k=1}^3 a_{ik}^{*T} T_k \tag{5.262}$$

where

$$\begin{aligned}
c_1^T &= \alpha_1 \rho_1 (\delta_1 P_{k1}^\tau + \varepsilon_1) + \sum_{i=1}^{i_{\max}} \mu_{iw1} (h_{iw1} - h_{i1}) + \chi_{w1} T_w \\
&+ \chi_1^{w\sigma} T_1^{w\sigma} - \left( \begin{aligned} &\left[ \psi_{21}^w (\chi_w^{2\sigma} + \chi_2^{w\sigma}) + \psi_{31}^w (\chi_w^{3\sigma} + \chi_3^{w\sigma}) \right] T'(p) \\ &- (\psi_{21}^w \chi_w^{2\sigma} + \psi_{31}^w \chi_w^{3\sigma}) T_w \end{aligned} \right) \frac{c_{pM1}}{h''(p) - h'(p)} T_1^{w\sigma} \\
&+ \left[ \chi_1^{2\sigma} - \frac{(1 - \psi_{12}) c_{pM1}}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}} (\chi_2^{1\sigma} T_2^{1\sigma} + \chi_1^{2\sigma} T_1^{2\sigma}) \right] T_1^{2\sigma} \\
&+ \left[ \chi_1^{3\sigma} - \frac{(1 - \psi_{13}) c_{pM1}}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}} (\chi_3^{1\sigma} T_3^{1\sigma} + \chi_1^{3\sigma} T_1^{3\sigma}) \right] T_1^{3\sigma}, \tag{5.263} \\
a_{11}^{*T} &= +\chi_1^{w\sigma} + \chi_{21} + \chi_{31} + \chi_1^{2\sigma} + \chi_1^{3\sigma} + \chi_{w1} \\
&- \left( \begin{aligned} &\left\{ \psi_{21}^w \left[ \chi_w^{2\sigma} [T'(p) - T_w] + \chi_2^{w\sigma} [T'(p) - T_2] \right] \right\} \\ &+ \left\{ \psi_{31}^w \left[ \chi_w^{3\sigma} [T'(p) - T_w] + \chi_3^{w\sigma} [T'(p) - T_3] \right] \right\} \end{aligned} \right) \frac{c_{pM1}}{h''(p) - h'(p)}
\end{aligned}$$



$$\begin{aligned}
& -\frac{\chi_2^{1\sigma}(1-\psi_{12})c_{pM1}}{h_{M1}^{2\sigma}-h_{M2}^{1\sigma}}(T_2^{1\sigma}-T_2)+\frac{\chi_1^{2\sigma}(1-\psi_{12})c_{pM1}}{h_{M1}^{2\sigma}-h_{M2}^{1\sigma}}(T_1-2T_1^{2\sigma}) \\
& -\frac{\chi_3^{1\sigma}(1-\psi_{13})c_{pM1}}{h_{M1}^{3\sigma}-h_{M3}^{1\sigma}}(T_3^{1\sigma}-T_3)+\frac{\chi_1^{3\sigma}(1-\psi_{13})c_{pM1}}{h_{M1}^{3\sigma}-h_{M3}^{1\sigma}}(T_1-2T_1^{3\sigma}), \quad (5.264)
\end{aligned}$$

$$a_{12}^{*T} = -\chi_{21} - \frac{\chi_2^{1\sigma}(1-\psi_{12})c_{pM1}}{h_{M1}^{2\sigma}-h_{M2}^{1\sigma}}T_1^{2\sigma} - \frac{\psi_{21}\chi_2^{w\sigma}c_{pM1}}{h''(p)-h'(p)}T_1^{w\sigma}, \quad (5.265)$$

$$a_{13}^{*T} = -\chi_{31} - \frac{\chi_3^{1\sigma}(1-\psi_{13})c_{pM1}}{h_{M1}^{3\sigma}-h_{M3}^{1\sigma}}T_1^{3\sigma} - \frac{\psi_{31}\chi_3^{w\sigma}c_{pM1}}{h''(p)-h'(p)}T_1^{w\sigma}, \quad (5.266)$$

$$\begin{aligned}
c_2^T &= \alpha_2\rho_2(\delta_2P_{k2}^\tau + \varepsilon_2) + \chi_2^{w\sigma}T_2^{w\sigma} + \chi_2^{1\sigma}T_2^{1\sigma} + (\chi_{w2} + \mu_{w2}c_{pw2})T_w \\
& + \left\{ \begin{array}{l} \psi_{12}^w f_{cond \rightarrow film} \left[ \chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} T_1^{w\sigma} \right] T_2^{w\sigma} \\ + \psi_{12} \left[ \chi_2^{1\sigma} T_2^{1\sigma} + \chi_1^{2\sigma} T_1^{2\sigma} \right] T_2^{1\sigma} \end{array} \right\} \frac{c_{pM2}}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}}, \quad (5.267)
\end{aligned}$$

$$a_{21}^{*T} = (\psi_{12}^w f_{cond \rightarrow film} \chi_1^{w\sigma} T_2^{w\sigma} + \psi_{12} \chi_1^{2\sigma} T_2^{1\sigma}) \frac{c_{pM2}}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}} - \chi_{21}, \quad (5.268)$$

$$\begin{aligned}
a_{22}^{*T} &= \chi_2^{w\sigma} + \chi_2^{1\sigma} + \chi_{w2} + \mu_{w2}c_{pw2} + \chi_{32} + \mu_{32}c_{p3} + \chi_{21} \\
& + \left\{ \begin{array}{l} \psi_{12}^w f_{cond \rightarrow film} \left[ \chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1) \right] \\ + \psi_{12} \left[ \chi_2^{1\sigma} (2T_2^{1\sigma} - T_2) + \chi_1^{2\sigma} (T_1^{2\sigma} - T_1) \right] \end{array} \right\} \frac{c_{pM2}}{h_{M1}^{2\sigma} - h_{M2}^{1\sigma}}, \quad (5.269)
\end{aligned}$$

$$a_{23}^{*T} = -\chi_{32} - \mu_{32}c_{p3}, \quad (5.270)$$

$$\begin{aligned}
c_3^T &= \alpha_3\rho_3(\delta_3P_{k3}^\tau + \varepsilon_3) + \chi_3^{w\sigma}T_3^{w\sigma} + \chi_3^{1\sigma}T_3^{1\sigma} + (\chi_{w3} + \mu_{w3}c_{pw3})T_w \\
& + \left\{ \begin{array}{l} \psi_{13}^w (1 - f_{cond \rightarrow film}) \left[ \chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} T_1^{w\sigma} \right] T_3^{w\sigma} \\ + \psi_{13} (\chi_3^{1\sigma} T_3^{1\sigma} + \chi_1^{3\sigma} T_1^{3\sigma}) T_3^{1\sigma} \end{array} \right\} \frac{c_{pM3}}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}}, \quad (5.271)
\end{aligned}$$

$$a_{31}^{*T} = \left[ \psi_{13}^w (1 - f_{cond \rightarrow film}) \chi_1^{w\sigma} T_3^{w\sigma} + \psi_{13} \chi_1^{3\sigma} T_3^{1\sigma} \right] \frac{c_{pM3}}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}} - \chi_{31}, \quad (5.272)$$

$$a_{32}^{*T} = -\chi_{32} - \mu_{23}c_{p2}, \quad (5.273)$$

$$a_{33}^{*T} = \chi_3^{w\sigma} + \chi_3^{1\sigma} + \chi_{w3} + \mu_{w3}c_{pw3} + \chi_{32} + \mu_{23}c_{p2} + \chi_{31} \\ + \left\{ \begin{array}{l} \psi_{13}^w (1 - f_{cond \rightarrow film}) \left[ \chi_w^{1\sigma} (T_w^{1\sigma} - T_w) + \chi_1^{w\sigma} (T_1^{w\sigma} - T_1) \right] \\ + \psi_{13} \left[ \chi_3^{1\sigma} (2T_3^{1\sigma} - T_3) + \chi_1^{3\sigma} (T_1^{3\sigma} - T_1) \right] \end{array} \right\} \frac{c_{pM3}}{h_{M1}^{3\sigma} - h_{M3}^{1\sigma}}, \quad (5.274)$$

## 5.17 Interface conditions

*Steam only:* For the case of steam only in the gas field, i.e.  $C_{M1} = 1$ , one has

$$T_3^{1\sigma} = T_2^{1\sigma} = T_1^{2\sigma} = T_1^{3\sigma} = T'(p) \approx T'(p_a) + \frac{dT'}{dp}(p - p_a) \\ \approx T'(p_a) + \frac{dT'}{dp} \frac{dp}{d\tau} \Delta\tau, \quad (5.275)$$

$$h_{M3}^{1\sigma} = h_{M2}^{1\sigma} = h'(p), \quad (5.276)$$

$$h_{M1}^{2\sigma} = h_{M1}^{3\sigma} = h''(p). \quad (5.277)$$

Here  $p_a$  is the reference pressure for the previous time step. Note that the pressure  $p$  can change during the time interval considered, which can reduce or even stop the condensation and promote evaporation, for example.

*Flashing:* For the case of spontaneous evaporation of the water, e.g. from the second velocity field,

$$T_2 > \left( 1 + \frac{\chi_1^{2\sigma}}{\chi_2^{1\sigma}} \right) T'(p) - \frac{\chi_1^{2\sigma}}{\chi_2^{1\sigma}} T_1 \quad (5.278)$$

one has

$$T_2^{1\sigma} = T_1^{2\sigma} = T'(p) \approx T'(p_a) + \frac{dT'}{dp}(p - p_a), \quad (5.279)$$

$$h_{M1}^{2\sigma} = h''(p). \quad (5.280)$$

Similarly, for the case of spontaneous evaporation of the water from the third velocity field,

$$T_3 > \left(1 + \frac{\chi_1^{3\sigma}}{\chi_3^{1\sigma}}\right) T'(p) - \frac{\chi_1^{3\sigma}}{\chi_3^{1\sigma}} T_1 \quad (5.281)$$

one has

$$T_3^{1\sigma} = T_1^{3\sigma} = T'(p) \approx T'(p_a) + \frac{dT'}{dp}(p - p_a), \quad (5.282)$$

$$h_{M1}^{3\sigma} = h''(p). \quad (5.283)$$

*Boiling at the wall:* If boiling conditions are in force at the wall one has

$$T_{M1}^{w\sigma} = T'(p) \approx T'(p_a) + \frac{dT'}{dp}(p - p_a). \quad (5.284)$$

*Non-condensibles present in the gas field:* This case is much more complicated than the single component case, especially for diffusion-controlled interfacial mass transfer.

The liquid side interface temperature during condensation processes is the saturation temperature at the local partial steam pressure

$$T_3^{1\sigma} = T_2^{1\sigma} = T'(p_{M1}) = T'(p_{M1a}) + \frac{dT'}{dp}(p_{M1} - p_{M1a}). \quad (5.285)$$

## 5.18 Lumped parameter volumes

If in a practical application a good mixing at any time within a control volume can be considered as a good approximation, the energy conservation equations in any of their variants simplify considerable. All convection and diffusion terms disappear. What remains are the time derivatives and the source terms. We summarize the results of this section for this particular case in order to make easy practical applications. Thus, we have for the entropy, specific internal energy, the temperature and the specific enthalpy equations the following result:

$$\rho_l \alpha_l \frac{ds_l}{d\tau} = \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \quad \text{for } \alpha_l \geq 0, \quad (5.286)$$

$$\frac{d}{d\tau} (\alpha_l \rho_l e_l \gamma_v) + p_l \frac{d}{d\tau} (\alpha_l \gamma_v) = \gamma_v De_l^*, \quad (5.287)$$

$$\rho_l c_{p,l} \alpha_l \frac{dT_l}{d\tau} - \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } C's} \right] \alpha_l \frac{dp}{d\tau}$$

$$= DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l), \quad (5.288)$$

$$\frac{d}{d\tau} (\alpha_l \rho_l h_l \gamma_v) - \alpha_l \gamma_v \frac{d\rho_l}{d\tau} = \gamma_v DT_l^N + \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) h_{il}. \quad (5.289)$$

The volume conservation equation divided by the volumetric porosity for this case reduces to

$$\frac{1}{\rho a^2} \frac{dp}{d\tau} = \sum_{l=1}^{l_{\max}} D\alpha_l - \frac{d}{d\tau} \ln \gamma_v, \quad (5.290)$$

where

$$D\alpha_l = \frac{\mu_l}{\rho_l} - \frac{1}{\rho_l^2} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, \text{all } C_{il}' s} \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right] + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{li}} \right)_{p, s, \text{all } C_{il}' s \text{ except } C_{li}} (\mu_{il} - C_{il} \mu_l) \right]. \quad (5.291)$$

## 5.19 Steady state

Now we consider a steady state flows. Neglecting the time derivatives we obtain the following different forms of the energy conservation.

$$\begin{aligned} & \rho_l (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) s_l - \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \mathcal{N} T) - \nabla \cdot \left\{ \alpha_l^e \rho_l \gamma \left[ \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \nabla C_{il} \right] \right\} \\ &= \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - \mu_l s_l \right], \end{aligned} \quad (5.292)$$

$$\begin{aligned} & \nabla \cdot (\alpha_l^e \rho_l e_l \mathbf{V}_l \gamma) + p_l \nabla \cdot (\alpha_l^e \mathbf{V}_l \gamma) - \nabla \cdot \left\{ \alpha_l^e \gamma \left[ \lambda_l^* \nabla T_l + \rho_l \sum_{i=2}^{i_{\max}} (e_{il} - e_{il}) D_{il}^* \nabla C_{il} \right] \right\} \\ &+ \overline{p_l \nabla \cdot (\alpha_l^e \mathbf{V}_l \gamma)} = \gamma_v De_l^*, \end{aligned} \quad (5.293)$$

$$\rho_l c_{pl} \left( \alpha_l^e \mathbf{V}_l \gamma \cdot \nabla \right) T_l - \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } C's} \right] \left( \alpha_l^e \mathbf{V}_l \gamma \cdot \nabla \right) p - \nabla \cdot \left( \alpha_l^e \lambda_l^* \gamma \nabla T \right) + T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \nabla \left( \alpha_l^e \rho_l D_{il}^* \gamma \nabla C_{il} \right) = \gamma_v \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \left( \mu_{il} - C_{il} \mu_l \right) \right], \quad (5.294)$$

$$\alpha_l^e \rho_l \left( \mathbf{V}_l \gamma \cdot \nabla \right) h_l - \alpha_l^e \mathbf{V}_l \gamma \cdot \nabla p_l - \nabla \cdot \left[ \alpha_l^e \gamma \lambda_l^* \nabla T_l \right] - \nabla \cdot \left[ \alpha_l^e \gamma \rho_l \sum_{i=2}^{i_{\max}} \left( h_{il} - h_{il} \right) D_{il}^* \nabla C_{il} \right] - \alpha_l^e \overline{\mathbf{V}'_l \gamma \cdot \nabla p'_l} + \delta_l \left( \alpha_l^e \gamma \nabla p_l \right) \cdot \sum_{i=1}^{i_{\max}} D_{il}^l \nabla \ln C_{il} = \gamma_v DT_l^N. \quad (5.295)$$

For single phase single component flows without mass transfer through the external sources we obtain and neglecting second order terms including conduction we obtain the well known forms:

$$\rho \left( \mathbf{V} \gamma \cdot \nabla \right) s = \gamma_v \frac{1}{T} DT^N, \quad (5.296)$$

$$\nabla \cdot \left( \rho e \mathbf{V} \gamma \right) + p \nabla \cdot \left( \mathbf{V} \gamma \right) = \gamma_v De^*, \quad (5.297)$$

$$\rho c_p \left( \mathbf{V} \gamma \cdot \nabla \right) T - \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_T \right] \left( \mathbf{V} \gamma \cdot \nabla \right) p = \gamma_v DT^N, \quad (5.298)$$

$$\rho \left( \mathbf{V} \gamma \cdot \nabla \right) h - \mathbf{V} \gamma \cdot \nabla p = \gamma_v DT^N, \quad (5.299)$$

where

$$DT^N = De^* = \rho \left( P_k + \varepsilon \right) + \dot{q}''' . \quad (5.300)$$

Again note the simplicity of the entropy notation of the conservation of energy

$$\left( \mathbf{V} \gamma \cdot \nabla \right) s = \gamma_v P_\eta / T + \dot{q}''' / \rho. \quad (5.301)$$

Remember that in a adiabatic pipe flow in which  $\dot{q}''' = 0$ , the friction dissipation causes entropy production due to

$$P_\eta = \frac{\lambda_{fr}}{D_h} \frac{1}{2} \bar{w}^3 y_{\text{lim}}^+ \sqrt{\frac{\lambda_{fr}}{8}} > 0. \quad (5.302)$$

This is in the technology very important if one has to do with expansion of gases in long pipes. Without taking this term into account the gas may “freeze in the computation” which does not happen in the reality.

**Problem:** Write the conservation equation for a pipe single-phase single-component flow without mass transfer.

**Solution:** The mass momentum and energy conservation equations are:

$$\frac{d}{dz}(\rho w \gamma_z) = 0 \text{ or } \frac{1}{w} \frac{dw}{dz} + \frac{1}{\rho} \frac{d\rho}{dz} = -\frac{1}{\gamma_z} \frac{d\gamma_z}{dz}, \quad (5.303)$$

$$\frac{1}{2} \frac{dw^2}{dz} + \frac{1}{\rho} \frac{dp}{dz} + \left( g_z + \frac{\lambda_{fr}}{D_h} \frac{1}{2} w^2 \right) \frac{\gamma_v}{\gamma_z} = 0, \quad (5.304)$$

$$\frac{ds}{dz} = \frac{\gamma_v}{\gamma_z} \frac{\lambda_{fr}}{D_h} \frac{1}{2} w^2 y_{\text{lim}}^+ \sqrt{\frac{\lambda_{fr}}{8}} / T + \dot{q}''' / (\rho w \gamma_z) = s_z. \quad (5.305)$$

**Problem:** Assume the flow is a perfect gas and rewrite the above system in terms of temperature and pressure.

**Solution:** Having in mind  $\rho = \frac{p}{RT}$ ,  $\kappa = c_p / (c_p - R)$ ,  $a^2 = \kappa RT$  and

$s = c_p \ln \frac{T}{T_0} - R \ln \frac{p}{p_0}$  we obtain

$$\frac{1}{w} \frac{dw}{dz} + \frac{1}{p} \frac{dp}{dz} - \frac{1}{T} \frac{dT}{dz} = -\frac{1}{\gamma_z} \frac{d\gamma_z}{dz}, \quad (5.306)$$

$$\frac{1}{2} \frac{dw^2}{dz} + \frac{RT}{p} \frac{dp}{dz} = -\left( g_z + \frac{\lambda_{fr}}{D_h} \frac{1}{2} w^2 \right) \frac{\gamma_v}{\gamma_z}, \quad (5.307)$$

$$\frac{1}{T} \frac{dT}{dz} - \frac{R}{c_p p} \frac{dp}{dz} = \frac{s_z}{c_p}. \quad (5.308)$$

Solving with respect to the derivatives we finally obtain:

$$\frac{dp}{dz} = -\frac{\rho w^2 \left( \frac{s_z}{c_p} - \frac{1}{\gamma_z} \frac{d\gamma_z}{dz} \right) + \rho \left( g_z + \frac{\lambda_{fr}}{D_h} \frac{1}{2} w^2 \right) \frac{\gamma_v}{\gamma_z}}{1 - \frac{w^2}{a^2}}, \quad (5.309)$$

$$\frac{1}{w} \frac{dw}{dz} = \frac{s_z}{c_p} - \frac{1}{\gamma_z} \frac{d\gamma_z}{dz} - \frac{1}{\kappa p} \frac{dp}{dz} = \frac{\frac{s_z}{c_p} - \frac{1}{\gamma_z} \frac{d\gamma_z}{dz} + \frac{1}{a^2} \left( g_z + \frac{\lambda_{fr}}{D_h} \frac{1}{2} w^2 \right) \frac{\gamma_v}{\gamma_z}}{1 - \frac{w^2}{a^2}}, \quad (5.310)$$

$$\frac{dT}{dz} = \frac{T}{c_p} \left( s_z + \frac{R}{p} \frac{dp}{dz} \right). \quad (5.311)$$

The obtained system of non-homogeneous non-linear ordinary differential equation indicates important behavior immediately. For constant pipe cross section  $\gamma_z = const$ , the nominator of the pressure gradient is positive. For sub-critical flow,  $w < a$ , the pressure gradient is therefore negative. Consequently, the velocity gradient in the second equation is positive. If the pipe is long enough for a given pressure difference acting at the both ends and the velocity approach the sound velocity at the exit of the pipe,  $w \rightarrow a$ , the pressure gradient tends to minus infinity. We call such flow *critical single phase flow*. Because the multi-phase flows are compressible flows they obey also such behavior in pipes. Finding the conditions for the critical flow will be one of the tasks solved in the next chapters. We learn on this example how to proceed in multiphase flows to.

Note the difference to the solution repeated in many gas-dynamics text books, *Albring* (1970), *Oswatitsch* (1952), *Shapiro* (1953), by using the energy conservation for adiabatic flow neglecting the gravitation and the friction component in the energy equation

$$\rho w \gamma_z \frac{d}{dz} \left( h + \frac{1}{2} w^2 \right) + \cancel{\gamma_v \rho w g_z} = \cancel{\gamma_v \dot{q}_l} + \gamma_v \frac{\lambda_{fr}}{D_h} \frac{1}{2} w^2 \gamma_{lim}^+ \sqrt{\frac{\lambda_{fr}}{8}}. \quad (5.312)$$

In this case the energy conservation for perfect fluid simplifies to

$$c_p \frac{dT}{dz} = -\frac{1}{2} \frac{dw^2}{dz} \quad \text{or} \quad \frac{1}{T} \frac{dT}{dz} = -(\kappa - 1) M^2 \frac{1}{w} \frac{dw}{dz}, \quad (5.313)$$

and allows to write the definition of the mach number,  $M^2 = \frac{w^2}{\kappa RT}$ , and the mass conservation equations in the following forms

$$\frac{1}{w} \frac{dw}{dz} = \frac{1}{\left[ 2 + (\kappa - 1) M^2 \right] M^2} \frac{dM^2}{dz}, \quad (5.314)$$

$$\frac{1}{p} \frac{dp}{dz} = -\frac{1 + (\kappa - 1) M^2}{\left[ 2 + (\kappa - 1) M^2 \right] M^2} \frac{dM^2}{dz}, \quad (5.315)$$

respectively. This allows in the rearranged momentum equation for separation of the variables

$$\frac{1-M^2}{\left(1+\frac{\kappa-1}{2}M^2\right)M^4}dM^2 = -\frac{\lambda_{fr}}{D_h}\kappa dz, \quad (5.316)$$

*Albring* (1970) p. 315, and for integration it analytically for constant friction coefficient

$$\frac{(\kappa+1)}{2\kappa} \ln \frac{\left(\frac{2}{M_1^2} + \kappa - 1\right)}{\left(\frac{2}{M_0^2} + \kappa - 1\right)} + \frac{1}{\kappa} \left(\frac{1}{M_0^2} - \frac{1}{M_1^2}\right) = -\frac{\lambda_{fr}}{D_h} \Delta z. \quad (5.317)$$

For critical flow at the exit  $M_1 = 1$  the results simplifies to

$$\frac{(\kappa+1)}{2\kappa} \ln \left[ 1 - \frac{2}{1+\kappa} \left( 1 - \frac{1}{M_0^2} \right) \right] + \frac{1}{\kappa} \left( 1 - \frac{1}{M_0^2} \right) = \frac{\lambda_{fr}}{D_h} \Delta z, \quad (5.318)$$

which is Eq. 21.48 in *Albring* (1970) p. 315. For frictionless flow the above equation is satisfied only for  $M_0 = 1$ .

## 5.20 Final remarks

It should be emphasized that in the temperature equation, just as in the energy equation, in the enthalpy equation and in the internal energy equations, the specific enthalpy occurs in the mass transfer terms. In the entropy equations the interfacial mass source terms are associated with specific entropies.

The most important result of this chapter is the rigorous derivation of the entropy equation, which reflects the second law of thermodynamics for multi-phase systems consisting of several chemical components which are conditionally divided into three velocity fields. It was shown that the use of specific entropy, introduced by *Clausius* and recognized by *Gibbs* as a very important quantity more than one hundred years ago, provides the simplest method for modern description and modeling of complicated flow systems. The experience gained by the author of this work in the development of the various versions of the IVA computer code, *Kolev* (1985-1995), which are based on the entropy concept allows for the recommendation of the local volume and time average entropy equation for general use in multi-phase flow dynamics.

## References

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# 6 Some simple applications of mass and energy conservation

*This Chapter contains some simple cases illustrating the use of the mass and the energy conservation for describing the thermodynamic state of multi-component single-phase systems. The results are useful themselves or for use as a benchmarks for testing the performance of computer codes.*

## 6.1 Infinite heat exchange without interfacial mass transfer

Consider the simple case of a three-field mixture having constant component mass concentration without wall interaction, without any mass transfer, and with instantaneous heat exchange that equalizes the field temperatures at any moment. Equation (5.176) reduces to

$$\alpha_l \rho_l c_{pl} \frac{dT}{T} - \alpha_l \rho_l \bar{R}_l \frac{dp}{p} = \frac{\dot{q}_l'''}{T} . \quad (6.1)$$

After dividing by the mixture density

$$\rho = \sum_{l=1}^3 \alpha_l \rho_l , \quad (6.2)$$

and summing all three equations one obtains

$$\frac{dT}{T} = \frac{\bar{R}}{c_p} \frac{dp}{p} = \frac{n-1}{n} \frac{dp}{p} , \quad (6.3)$$

where the mixture specific heat is

$$c_p = \sum_{l=1}^3 \frac{\alpha_l \rho_l}{\rho} c_{pl} , \quad (6.4)$$

and the effective gas constant is

$$R = \sum_{l=1}^3 \frac{\alpha_l \rho_l}{\rho} \bar{R}_l . \quad (6.5)$$

Note that for the second and the third field the pseudo-gas constant is negligibly small. The polytropic exponent is

$$n = \frac{c_p}{c_p - R}. \quad (6.6)$$

Integration of Eq. (6.3) yields

$$\frac{T}{T_0} = \left( \frac{p}{p_0} \right)^{\frac{n-1}{n}}. \quad (6.7)$$

This change of state is associated with considerable entropy change for the gas phase

$$ds_1 = -c_{p1} \left( \frac{1}{n} - \frac{1}{\kappa_1} \right) \frac{dp}{p}, \quad (6.8)$$

and for the other fields

$$ds_l = c_{pl} \frac{n-1}{n} \frac{dp}{p}, \quad \text{for } l = 2, 3. \quad (6.9)$$

There are a number of cases where such simplification forms a useful approximation instead of modeling all details of the heat transfer for inert velocity fields. This result was obtained for two velocity field by *Tangren et al.* in 1949. In expanding steady state flow the steady state energy conservation for each field gives

$$\frac{ds_1}{dp} \frac{dp}{dz} = \frac{1}{X_1 G T_1} (\dot{q}_{21}'' + \dot{q}_{31}''), \quad (6.10)$$

$$\frac{ds_2}{dp} \frac{dp}{dz} = -\frac{1}{X_2 G T_2} \dot{q}_{21}'', \quad (6.11)$$

$$\frac{ds_3}{dp} \frac{dp}{dz} = -\frac{1}{X_3 G T_3} \dot{q}_{31}''. \quad (6.12)$$

Therefore the transferred thermal power per unit mixture volume from the liquid to the gas necessary to equalize the temperatures is a function of the pressure gradient

$$\dot{q}_{21}'' + \dot{q}_{31}'' = -c_{p1} \left( \frac{1}{n} - \frac{1}{\kappa_1} \right) \frac{X_1 G T_1}{p} \frac{dp}{dz}, \quad (6.13)$$

$$\dot{q}_{21}'' = -X_2 G T_2 \frac{ds_2}{dp} \frac{dp}{dz} = -X_2 G T_2 c_{p2} \frac{n-1}{n} \frac{1}{p} \frac{dp}{dz}, \quad (6.14)$$

$$\dot{q}_{31}'' = -X_3 G T_3 \frac{ds_3}{dp} \frac{dp}{dz} = -X_3 G T c_{p3} \frac{n-1}{n} \frac{1}{p} \frac{dp}{dz}. \quad (6.15)$$

Useful relations are

$$\frac{d\rho_1}{dp} = \frac{\rho_1}{np}, \quad (6.16)$$

the integrated form

$$\frac{\rho_1}{\rho_{10}} = \left( \frac{p}{p_0} \right)^{\frac{1}{n}} \quad (6.17)$$

and

$$\frac{d\rho_1}{ds_1} = - \frac{\rho_1}{nc_{p1} \left( \frac{1}{n} - \frac{1}{\kappa_1} \right)}. \quad (6.18)$$

## 6.2 Discharge of gas from a volume

Consider a volume  $V$  with initial pressure and temperature  $p_0$  and  $T_0$ , respectively. The discharged mass per unit time and volume is

$$\mu_w = \rho_{out} w \frac{F}{V}. \quad (6.19)$$

Here  $F$  is the flow cross section. The transient behavior of the pressure and temperature is then described by

$$\frac{1}{a^2} \frac{dp}{d\tau} = -\rho_{out} w \frac{F}{V} \quad (6.20)$$

$$\rho c_p \frac{dT}{d\tau} - \frac{dp}{d\tau} = 0 \quad \text{i.e.} \quad p = p_0 \left( \frac{T}{T_0} \right)^{\frac{\kappa}{\kappa-1}}. \quad (6.21)$$

Note that for perfect gases the last equation results in the isentropic relation. The discharge velocity and the corresponding densities are

$$w = \sqrt{\kappa R T} \sqrt{\frac{2}{\kappa+1}}, \quad (6.22)$$

$$\rho_{out} = \rho \left( \frac{2}{\kappa+1} \right)^{\frac{1}{\kappa-1}}, \quad (6.23)$$

for  $\varepsilon \geq \varepsilon^*$ , and

$$w^2 = 2 \frac{p}{\rho} \frac{\kappa}{\kappa-1} \left( 1 - \varepsilon^{\frac{\kappa-1}{\kappa}} \right) = \kappa RT \frac{2}{\kappa-1} \left( 1 - \varepsilon^{\frac{\kappa-1}{\kappa}} \right), \quad (6.24)$$

$$\rho_{out} = \rho \left[ 1 - \frac{1}{2} (\kappa-1) \frac{w^2}{\kappa RT} \right]^{\frac{1}{\kappa-1}} = \rho \varepsilon^{\frac{1}{\kappa}} \quad (6.25)$$

for  $\varepsilon < \varepsilon^*$ , which means for critical and sub-critical flow, respectively. This result was obtained by *de Saint Venant* and *Wantzel* in 1838. Here

$$\varepsilon = \frac{p_{out}}{p} \quad (6.26)$$

is the pressure ratio and

$$\varepsilon^* = \left( \frac{2}{\kappa+1} \right)^{\frac{\kappa}{\kappa-1}} \quad (6.27)$$

is the critical pressure ratio – see *Oswatitsch* (1952), *Landau and Lifshitz* (1987), p. 319. For the case of critical outflow we have

$$\frac{dp}{d\tau} = -\frac{F}{V} \kappa p \left( \frac{2}{\kappa+1} \right)^{\frac{1}{2} \frac{\kappa+1}{\kappa-1}} \sqrt{\kappa RT}, \quad (6.28)$$

$$\rho c_p \frac{dT}{d\tau} - \frac{dp}{d\tau} = 0. \quad (6.29)$$

Replacing the pressure derivative in the second equation we obtain

$$\rho c_p \frac{dT}{d\tau} = -\frac{F}{V} \kappa p \left( \frac{2}{\kappa+1} \right)^{\frac{1}{2} \frac{\kappa+1}{\kappa-1}} \sqrt{\kappa RT}, \quad (6.30)$$

or

$$\frac{1}{T^{3/2}} \frac{dT}{d\tau} = -\frac{F}{V} (\kappa-1) \left( \frac{2}{\kappa+1} \right)^{\frac{1}{2} \frac{\kappa+1}{\kappa-1}} \sqrt{\kappa R}, \quad (6.31)$$

or after integrating

$$T = \left[ \frac{1}{T_0^{1/2}} + \frac{1}{2} \frac{F}{V} (\kappa-1) \left( \frac{2}{\kappa+1} \right)^{\frac{1}{2} \frac{\kappa+1}{\kappa-1}} \sqrt{\kappa R} \tau \right]^{-2}. \quad (6.32)$$

The pressure as a function of time is computed from the isentropic equation as mentioned before.

For the sub-critical case we have

$$\frac{1}{p} \frac{dp}{d\tau} = -\kappa \sqrt{\kappa RT} \left( \frac{2}{\kappa-1} \right)^{1/2} \varepsilon^{\frac{1}{\kappa}} \left( 1 - \varepsilon^{\frac{\kappa-1}{\kappa}} \right)^{1/2} \frac{F}{V}. \quad (6.33)$$

Replacing the temperature using the isentropic relation we obtain

$$\frac{dp}{d\tau} = -\kappa p_0 \sqrt{\kappa RT_0} \left( \frac{p_{out}}{p_0} \right)^{\frac{3\kappa-1}{2\kappa}} \left( \frac{2}{\kappa-1} \right)^{1/2} \varepsilon^{\frac{1-3\kappa}{2\kappa}} \left( 1 - \varepsilon^{\frac{\kappa-1}{\kappa}} \right)^{1/2} \frac{F}{V}, \quad (6.34)$$

or

$$\frac{d \left( 1 - \varepsilon^{\frac{\kappa-1}{\kappa}} \right)}{\varepsilon \left( 1 - \varepsilon^{\frac{\kappa-1}{\kappa}} \right)^{1/2}} = \frac{\kappa^2}{\kappa-1} \left( \frac{2}{\kappa-1} \right)^{1/2} p_0 \sqrt{\kappa RT_0} \left( \frac{p_{out}}{p_0} \right)^{\frac{\kappa-1}{2\kappa}} \frac{F}{V} d\tau. \quad (6.35)$$

Here the subscript 0 means an arbitrary state, e.g. the initial state. The equation has to be integrated numerically.

### 6.3 Injection of inert gas in a closed volume initially filled with inert gas

For the case of inert gas injection with constant mass source density  $\mu_{nw}$  and constant temperature  $T_w$  in a closed volume initially filled with inert gas at temperature and pressure  $T_0$  and  $p_0$ , respectively, we have

$$\frac{1}{a^2} \frac{dp}{d\tau} = \mu_{nw} + \frac{1}{Tc_p} \left[ \mu_{nw} (h_{nw} - h_n) \right] \quad (6.36)$$

$$\rho c_p \frac{dT}{d\tau} - \frac{dp}{d\tau} = \mu_{nw} (h_{nw} - h_n). \quad (6.37)$$

Rearranging

$$\frac{dp}{d\tau} = \mu_{nw} a^2 \frac{T_w}{T} \quad \text{or} \quad \frac{dp}{d\tau} = \mu_{nw} \kappa RT_w \quad \text{or} \quad p = p_0 + \mu_{nw} \kappa RT_w \tau, \quad (6.38)$$

$$\rho c_p \frac{dT}{d\tau} - \frac{dp}{d\tau} = \mu_{nw} c_p (T_w - T) \quad \text{or} \quad \frac{dT}{\kappa T_w T - T^2} = \frac{d\tau}{\frac{p_0}{\mu_{nw} R} + \kappa T_w \tau}, \quad (6.39)$$

and integrating we obtain

$$p = p_0 + \mu_{nw} \kappa R T_w \tau \quad (6.40)$$

$$T = T_0 \frac{1 + \kappa \frac{\mu_{nw} R T_w \tau}{p_0}}{1 + \frac{\mu_{nw} R T_0 \tau}{p_0}}, \quad \frac{T - T_0}{T_0} = \frac{\kappa T_w - T_0}{\frac{p_0}{\mu_{nw} R \tau} + T_0}. \quad (6.41)$$

We see that in this case the pressure increases linearly with time and the temperature tends to an asymptotic for infinite time

$$T(\tau \rightarrow \infty) = \kappa T_w. \quad (6.42)$$

## 6.4 Heat input in a gas in a closed volume

For the case of no chemical reaction but introduction of heat only into the gas mixture we have

$$\frac{dp}{d\tau} = \frac{a^2}{T c_p} \dot{q}''', \quad (6.43)$$

$$\rho c_p \frac{dT}{d\tau} - \frac{dp}{d\tau} = \dot{q}'''. \quad (6.44)$$

Bearing in mind that

$$a^2 = \kappa R T, \quad (6.45)$$

$$\frac{R}{c_p} = \frac{\kappa - 1}{\kappa}, \quad (6.46)$$

and after rearrangement we obtain

$$\frac{dp}{d\tau} = (\kappa - 1) \dot{q}''', \quad (6.47)$$

$$\frac{d}{d\tau} \ln T = (\kappa - 1) \frac{\dot{q}'''}{p}. \quad (6.48)$$

We recognize that

$$\frac{d}{d\tau} \ln T = \frac{d}{d\tau} \ln p \quad (6.49)$$

results in a polytropic change of state with polytropic exponent equal to one,



$$T = \frac{T_0}{p_0} p. \quad (6.50)$$

Integrating over the time  $\Delta\tau$  for initial conditions  $p_0$  and  $T_0$  and assuming that the isentropic exponent does not depend on temperature and pressure we finally obtain

$$p = p_0 + (\kappa - 1) \int_0^{\Delta\tau} \dot{q}'''' d\tau, \quad (6.51)$$

$$T = \frac{T_0}{p_0} p = T_0 \left[ 1 + \frac{\kappa - 1}{p_0} \int_0^{\Delta\tau} \dot{q}'''' d\tau \right]. \quad (6.52)$$

If the heat source term is not a time function we have

$$p = p_0 + (\kappa - 1) \dot{q}'''' \Delta\tau, \quad (6.53)$$

$$T = \frac{T_0}{p_0} p = T_0 \left[ 1 + \frac{\kappa - 1}{p_0} \dot{q}'''' \Delta\tau \right]. \quad (6.54)$$

## 6.5 Steam injection in a steam-air mixture

Consider a closed volume filled with air,  $C_n = 1$ , at initial temperature  $T = 273.15 + 20K$  and pressure  $p = 1 \times 10^5 Pa$ . The volume is adiabatic,  $\dot{q}'''' = 0$ . Inject into this volume steam at a temperature  $T_w = 273.15 + 100K$  and intensity  $\mu = \mu_M = \mu^+ = const$ . There is no mass generation for the inert component (air), that is  $\mu_{n1} = 0$ . Compute the change in the concentration, pressure and temperature with time. The air concentration in the volume is governed by

$$\frac{dC_n}{d\tau} = -C_n \mu / \rho, \quad (6.55)$$

or integrating within a time step  $\Delta\tau$ ,

$$C_{n,\tau+\Delta\tau} = C_n / \exp \left( \int_{\tau}^{\tau+\Delta\tau} \frac{\mu}{\rho} d\tau \right). \quad (6.56)$$

For a closed adiabatic volume occupied by a single-phase mixture, one has the simple form of the entropy equation in terms of temperature and pressure

$$\rho c_p \frac{dT}{d\tau} - \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T,all\_Cs} \right] \frac{dp}{d\tau} = \mu \left[ c_{pM} (T_w - T) + C_n T \Delta S_n^{np} \right]. \quad (6.57)$$

Taking into account that

$$\Delta s_n^{np} = -(1 - C_n) \left( \frac{\partial h_M}{\partial p_M} \right)_T \rho R_n / \left\{ 1 - C_n \left[ 1 - R_n T \left( \frac{\partial \rho_M}{\partial p_M} \right)_T \right] \right\} \geq 0 \quad (6.58)$$

one obtains

$$\begin{aligned} & \rho c_p \frac{dT}{d\tau} - \left[ 1 - \rho \left( \frac{\partial h}{\partial p} \right)_{T, \text{all } C's} \right] \frac{dp}{d\tau} \\ &= \mu \left[ c_{pM} (T_w - T) - C_n (1 - C_n) \rho \frac{R_n T \left( \frac{\partial h_M}{\partial p_M} \right)_T}{1 - C_n \left[ 1 - R_n T \left( \frac{\partial \rho_M}{\partial p_M} \right)_T \right]} \right]. \end{aligned} \quad (6.59)$$

For steam

$$\left( \frac{\partial h_M}{\partial p_M} \right)_T < 0, \quad (6.60)$$

$$R_n T \left( \frac{\partial \rho_M}{\partial p_M} \right)_T \approx \frac{R_n}{R_M} = \frac{287.22}{461.631} < 1, \quad (6.61)$$

and therefore

$$\frac{R_n T \left( \frac{\partial h_M}{\partial p_M} \right)_T}{1 - C_n \left[ 1 - R_n T \left( \frac{\partial \rho_M}{\partial p_M} \right)_T \right]} < 0. \quad (6.62)$$

Consider the case of an open volume in which in addition to the inlet flow there is an outlet flow that guarantees nearly constant pressure. We assume an ideal, immediate intermixing of the introduced steam with the air in the volume. For this particular case and ideal gas mixtures one has

$$\rho c_p \frac{dT}{d\tau} = \mu c_{pM} (T_w - T). \quad (6.63)$$

It is evident that for  $T_w > T$  the volume temperature should increase monotonically until reaching  $T_w$ . For steam considered as a real gas one obtains

$$\rho c_p \frac{dT}{d\tau} = \mu \left[ c_{pM} (T_w - T) - C_n (1 - C_n) \rho \frac{R_n T \left( \frac{\partial h_M}{\partial p_M} \right)_T}{1 - C_n \left[ 1 - R_n T \left( \frac{\partial \rho_M}{\partial p_M} \right)_T \right]} \right]. \quad (6.64)$$

On the basis of Eq. (6.62), the temperature will increase somewhat more strongly and monotonically if the mixture is considered to be a non-perfect gas mixture. For single component flow the non-perfect gas term is equal to zero because of the multiplier  $C_n(1 - C_n)$ .

For a closed volume, the pressure is described by

$$\frac{1}{a^2} \frac{\partial p}{\partial \tau} = \mu \left\{ \begin{array}{l} 1 + C_n \frac{1}{\rho} \left( \frac{\partial \rho}{\partial C_n} \right)_{p,T, \text{all } C_s \text{ except } C_i} \\ - \frac{1}{\rho} \left( \frac{\partial \rho}{\partial T} \right)_{p, \text{all } C_s} \frac{1}{c_p} \left[ c_{pM} (T_w - T) + T \Delta s_n^{np} C_n \right] \end{array} \right\}. \quad (6.65)$$

After substituting the derivatives (A-1), (A-2) and (A-3) from Chapter 3 one obtains

$$\begin{aligned} \frac{1}{\mu a^2} \frac{\partial p}{\partial \tau} = & 1 + \left( \frac{\partial \rho}{\partial p} \right)_{T, \text{all } C_s} \left\{ C_n \left[ \frac{1}{(\partial \rho_M / \partial p_M)_T} - R_n T \right] \right. \\ & \left. + \frac{1}{\rho c_p} \left[ \rho_n R_n - \frac{\left( \frac{\partial \rho_M}{\partial T} \right)_{p_M}}{\left( \frac{\partial \rho_M}{\partial p_{M1}} \right)_T} \right] \left[ c_{pM} (T_w - T) \right] \right\} + T \Delta s_n^{np} C_n \end{aligned} \quad (6.66)$$

For a mixture of perfect gases this gives

$$\frac{1}{a_1^2} \frac{\partial p}{\partial \tau} = \mu_1 \left( 1 + C_{n1} \frac{R_{M1} - R_{n1}}{R_1} + \frac{c_{pM1}}{c_{p1}} \frac{T_{w1} - T_1}{T_1} \right). \quad (6.67)$$

In a similar manner to the temperature changes, the pressure change will be somewhat more pronounced in a non-perfect gas than in a perfect gas mixture.

For practical application the perfect gas mixture approximation can be used:

$$\frac{d}{d\tau} \ln C_n = - \frac{\mu_{Mw} RT}{p}, \quad (6.68)$$

$$\frac{dp}{d\tau} = \mu_{Mw} c_{pM} \left[ \left( 1 - \frac{\kappa}{\kappa_M} \right) T + (\kappa - 1) T_w \right], \quad (6.69)$$

$$\frac{dT}{d\tau} = \frac{\mu_{Mw} R c_{pM}}{p c_p} \left( \kappa T_w T - \frac{\kappa}{\kappa_M} T^2 \right) \quad \text{or} \quad \frac{d}{d\tau} \ln \frac{T}{\kappa T_w - \frac{\kappa}{\kappa_M} T} = \frac{\mu_{Mw} R c_{pM}}{p c_p} \kappa T_w, \quad (6.70)$$

$$R = C_n R_n + (1 - C_n) R_M, \quad (6.71)$$

$$c_p = C_n c_{pn} + (1 - C_n) c_{pM}, \quad (6.72)$$

$$\kappa = \frac{c_p}{c_p - R}, \quad (6.73)$$

$$\kappa_M = \frac{c_{pM}}{c_{pM} - R_M}. \quad (6.74)$$

A simple numerical method can be easily constructed to integrate this system of three non-linear ordinary differential equations.

## 6.6 Heat removal from a closed volume containing equilibrium two-phase mixture

**Problem:** Given a vessel with volume  $Vol$  filled with *saturated* water at steam. The steam volume is *stratified* above the water volume. The initial steam volume fraction is  $\alpha_{10}$ . The initial pressure is  $p_0$ . The power  $\dot{Q}_w^{1\sigma}$  is removed from the steam volume leading to condensation of the steam. Assume that the process always occurs at saturation and compute the pressure development as a function of time. Assume  $Vol = 140 \text{ m}^3$ ,  $\alpha_{10} = 0.75, 0.5, 0.25$ , and  $\dot{Q}_w^{1\sigma} = 60$  and  $30 \text{ MW}$ .

**Solution to the problem:** First I take the volume concentration equation (5.207) together with (5.90) and apply it the total volume:

$$\frac{1}{\rho a^2} \frac{dp}{d\tau} = \sum_{l=1}^{l_{\max}} \left[ \frac{\mu_l}{\rho_l} - \frac{1}{\rho_l^2} \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, \text{all } C_{i,s}} \frac{\dot{q}_l'''}{T_l} \right] = \sum_{l=1}^{l_{\max}} \left[ \frac{\mu_l}{\rho_l} - \frac{1}{\rho_l^2} \frac{T}{c_{pl}} \left( \frac{\partial \rho_l}{\partial T} \right)_p \frac{\dot{q}_l'''}{T_l} \right].$$

Using the assumption of thermal equilibrium  $\frac{1}{\rho a^2} = \frac{\alpha_1}{\rho'' a''^2} + \frac{1-\alpha_1}{\rho' a'^2}$ ,  $\mu_1 = -\mu_{12}$ ,  $\mu_2 = \mu_{12}$ ,  $\dot{q}_1'' = -\dot{q}_w^{m1\sigma} = -\dot{Q}_w^{l\sigma}/Vol$ ,  $\dot{q}_2'' = 0$ ,  $\mu_{12} = \frac{\dot{q}_w^{m1\sigma}}{h'' - h'}$ ,  $\frac{v'' - v'}{h'' - h'} = \frac{1}{T} \frac{dT'}{dp}$ ,

Clapeyron (1834),  $\left(\frac{\partial \rho''}{\partial T}\right)_p = 0$ , the volume conservation equation reduces to

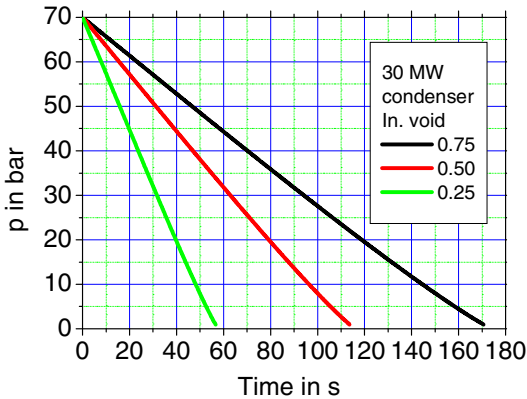
$$\frac{dp}{d\tau} = -\frac{1}{T'} \frac{dT'}{dp} \frac{\dot{Q}_w^{l\sigma}}{Vol} \bigg/ \frac{1}{\rho a^2}.$$

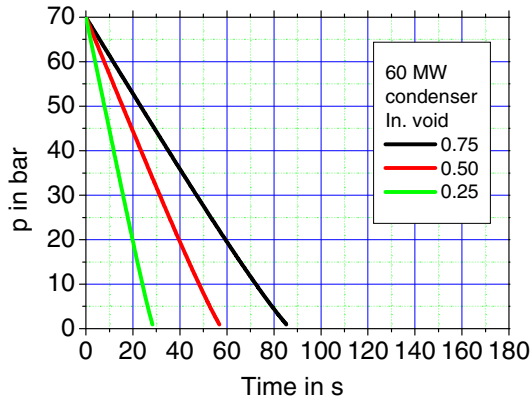
Integrating over a time step assuming that the expression under the integral is an averaged value over the time step results in

$$p = p_0 - \left( \frac{1}{T'} \frac{dT'}{dp} \frac{\dot{Q}_w^{l\sigma}}{Vol} \bigg/ \frac{1}{\rho a^2} \right) \Delta\tau \approx p_0 - \left( \frac{1}{T'} \frac{dT'}{dp} \frac{\dot{Q}_w^{l\sigma}}{Vol} \frac{\rho'' a''^2}{\alpha_1} \right) \Delta\tau.$$

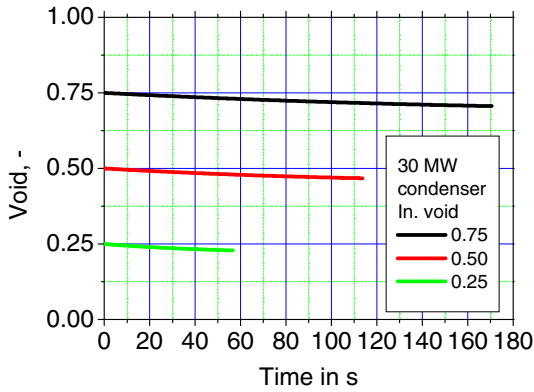
Note that the mass conservation  $\frac{d}{d\tau} [\alpha_1 \rho'' + (1-\alpha_1) \rho'] = 0$  gives

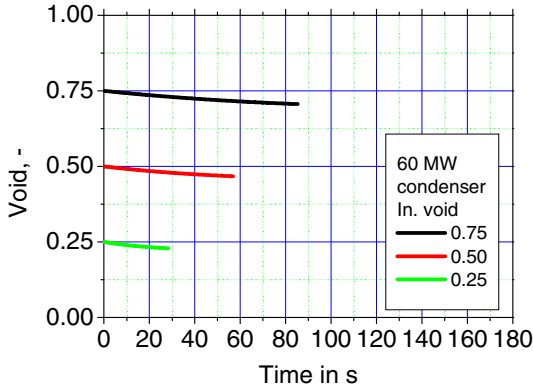
$$\alpha_1 = \alpha_{10} + \frac{p - p_0}{\rho' - \rho''} \left[ \alpha_1 \frac{d\rho''}{dp} + (1-\alpha_1) \frac{d\rho'}{dp} \right].$$





**Fig. 6.1** Pressure as a function of time for the three different initial void fractions and for 60 and 30 MW cooling power





**Fig. 6.2** Void fraction as a function of time for the three different initial void fractions and for 60 and 30 MW cooling power

Figures 6.1 a) and b) show the pressure dependences as a function of time for the three different initial void fractions and for 60 and 30 MW cooling power. Figures 6.2 a) and b) show the corresponding void fraction dependences as a function of time for the three different initial void fractions and for 60 and 30 MW cooling power. The simple form of the volume conservation equation possesses remarkable messages. We realize that the smaller the void fraction, the stronger the influence of the heat removal on the decompression. Therefore, the dynamics of such a system depends on the void volume  $\alpha_1 Vol$ . This is clearly visible from Figs. 6.1 a) and b). Contrary to the strong pressure reduction, we see a relatively small variation of the void fraction from Figs. 6.2 a) and b).

## 6.7 Chemical reaction in a gas mixture in a closed volume

For the case of a closed control volume filled with a multi-component gas mixture we have

$$\rho c_p \frac{dT}{d\tau} - \frac{dp}{d\tau} = DT^N, \quad (6.75)$$

$$\frac{1}{\rho a^2} \frac{dp}{d\tau} = D, \quad (6.76)$$

where for a mixture consisting of perfect gases

$$DT^N = \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_i (h_i^* - h_i), \quad (6.77)$$

$$\begin{aligned} \rho D &= \frac{1}{c_p T} DT^N + \frac{1}{R} \sum_{i=2}^{i_{\max}} \mu_i (R_i - R_1) \\ &= \frac{1}{c_p T} \left[ \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_i (h_i^* - h_i) \right] + \frac{1}{R} \sum_{i=2}^{i_{\max}} \mu_i (R_i - R_1). \end{aligned} \quad (6.78)$$

For  $\mu_i < 0$ ,  $h_i^* = h_i$ . For the case  $\mu_i > 0$  the assumption that the origination enthalpy is equal to the component specific enthalpy at some specified reference temperature and pressure,

$$h_i^* = h_{i,\text{ref}}(T_{\text{ref}}, p_{\text{ref}}), \quad (6.79)$$

simplifies the computation but should be considered in computing a proper energy source due to chemical reaction. The final form of the temperature and the pressure change equations is

$$\frac{dT}{d\tau} = \frac{1}{\rho c_p} \left( 1 + \frac{a^2}{c_p T} \right) \left[ \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_i (h_i^* - h_i) \right] + \frac{a^2}{\rho c_p R} \sum_{i=2}^{i_{\max}} \mu_i (R_i - R_1), \quad (6.80)$$

$$\frac{dp}{d\tau} = \frac{a^2}{c_p T} \left[ \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_i (h_i^* - h_i) \right] + \frac{a^2}{R} \sum_{i=2}^{i_{\max}} \mu_i (R_i - R_1). \quad (6.81)$$

Bearing in mind that

$$a^2 = \kappa RT, \quad (6.82)$$

$$\frac{R}{c_p} = \frac{\kappa - 1}{\kappa}, \quad (6.83)$$

$$\begin{aligned} \frac{d}{d\tau} \ln T &= (\kappa - 1) \frac{1}{p} \left[ \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_i (h_i^* - h_i) + T \sum_{i=2}^{i_{\max}} \mu_i (R_i - R_1) \right] \\ &= (\kappa - 1) \frac{1}{p} \left[ \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_i (h_i^* - h_i + TR_i) \right], \end{aligned} \quad (6.84)$$

$$\frac{dp}{d\tau} = (\kappa - 1) \left[ \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_i (h_i^* - h_i) \right] + \kappa T \sum_{i=2}^{i_{\max}} \mu_i (R_i - R_1). \quad (6.85)$$

These are remarkable equations. From Eq. (6.84) we realize that in the case of a chemical reaction in order to have a constant temperature the following condition has to hold

$$\dot{q}''' = - \sum_{i=1}^{i_{\max}} \mu_i (h_i^* - h_i + TR_i). \quad (6.86)$$



## 6.8 Hydrogen combustion in an inert atmosphere

Mixtures of hydrogen and oxygen react at atmospheric conditions over millions of years to produce water *Kolarov* (1970). At an elevated temperature of 673.15K this process happens within 80 days and at 773.15K within 2h *Crussard* (1907). A real combustion is possible if some specific thermodynamic conditions are satisfied as will be discussed in a moment.

### 6.8.1 Simple introduction to combustion kinetics

A single chemical combustion reaction containing  $N$  initial components and final products can be mathematically described by the so called stoichiometric equation



see in *Kuo* (1986). Here  $M_i$  is the chemical identification symbol of the  $i$ -th species before and after the reaction,  $\nu_i'$  is the stoichiometric coefficient of the *initial* substances, and  $\nu_i''$  is the stoichiometric coefficient of the *final* substances.  $N$  is the total number of the components participating in the reaction. The *order of the chemical reaction* is defined by

$$n = \sum_{i=1}^N \nu_i'. \quad (6.88)$$

The reaction velocity as found by *Beketov* in 1865, and by *Guldberg* and *Baare* in 1867, has the general form

$$\frac{dY_i^*}{d\tau} = (\nu_i'' - \nu_i') k(T) \prod_{i=1}^N Y_i^{*\nu_i'}, \quad (6.89)$$

where  $k(T)$  is the velocity coefficient of the reaction and

$$Y_{i,1}^* = \frac{\rho_1 C_{i,1}}{M_{i,1}} \quad (6.90)$$

is the molar density in  $kg\text{-mole}/m^3$ . The dependence of the reaction velocity on the temperature is described by the *Arrhenius* law,

$$k(T) = A \exp\left[-E_a / (RT)\right], \quad (6.91)$$

see in *Bartlmä* (1975), where  $A = 5 \times 10^6 [m^3/(kg\text{-mole})]^2/s$ ,  $R = 8\,314 \text{ J}/(kg\text{-mole K})$  is the general gas constant, and  $E = 78 \times 10^6 \text{ J}/(kg\text{-mole})$  is the activation energy. Instead of complete modeling of the complicated  $H_2\text{-}O_2$  kinetics, as is usually done in rocket propulsion systems design, there is a simplified possibility to describe the global chemical combustion that is frequently used in the literature:



Therefore for a closed control volume we have

$$\mu_{H_2,1} = \alpha_1 M_{H_2,1} \frac{dY_{H_2,1}^*}{d\tau} = -\alpha_1 M_{H_2,1} 2k(T) Y_{H_2,1}^{*2} Y_{O_2,1}^*, \quad (6.93)$$

$$\mu_{O_2,1} = \alpha_1 M_{O_2,1} \frac{dY_{O_2,1}^*}{d\tau} = -\alpha_1 M_{O_2,1} k(T) Y_{H_2,1}^{*2} Y_{O_2,1}^*. \quad (6.94)$$

One can easily see in the kinetic model the expected relations: a) the mass of hydrogen and oxygen is consumed to produce water-steam

$$\mu_{H_2,1} + \mu_{O_2,1} + \mu_{H_2O,1} = 0, \quad (6.95)$$

and b) the ratio between the reacted mass source terms is related to the mole masses as follows

$$\frac{\mu_{O_2,1}}{\mu_{H_2,1}} = \frac{1}{2} \frac{M_{O_2}}{M_{H_2}} = 8, \quad (6.96)$$

$$\frac{\mu_{H_2O,1}}{\mu_{H_2,1}} = -\left(1 + \frac{1}{2} \frac{M_{O_2}}{M_{H_2}}\right) = -9. \quad (6.97)$$

Therefore, once the mass source term of the hydrogen mass conservation equation is computed all other mass sources are easily derived from Eqs. (6.96) and (6.97).

Note that usually the *fuel equivalence ratio*

$$\Phi = \frac{C_{H_2}}{1 - C_{H_2}} \frac{1 - C_{st,H_2}}{C_{st,H_2}} = \frac{Y_{H_2}}{1 - Y_{H_2}} \frac{1 - Y_{st,H_2}}{Y_{st,H_2}}, \quad (6.98)$$

with a stoichiometric mass concentration  $C_{st,H_2}$ , is used to classify the premixed combustion process into three groups,

$$\text{rich combustion:} \quad \Phi > 1, \quad (6.99)$$

$$\text{stoichiometric combustion:} \quad \Phi = 1, \quad (6.100)$$

$$\text{lean combustion:} \quad \Phi < 1. \quad (6.101)$$

In the limiting case of complete consumption of the fuel or of the oxidizer over the time step  $\Delta\tau$  we have

$$\mu_{H_2,1} = -\min\left(\frac{\rho_1 C_{H_2,1}}{\Delta\tau}, \frac{\rho_1 C_{O_2,1}/8}{\Delta\tau}\right) = -\frac{\rho_1}{\Delta\tau} \min(C_{H_2,1}, C_{O_2,1}/8). \quad (6.102)$$

The first term in the brackets is taken for either lean or stoichiometric mixtures, and the second for rich mixtures.

### 6.8.2 Ignition temperature and ignition concentration limits

In accordance with *Isserlin* (1987), if the mole-fraction of  $H_2$  in the air is between 4.1 and 74.2%, or if the mole-fraction of  $H_2$  in the hydrogen-oxygen mixtures is between 4 and 94%, and if the gas temperature increases above  $T_{1,ign} = 783.15 - 863.15K$ , ignition is possible. *Bröckerhoff* et al. (2002) reviewed 10 sources and summarized that if the mole-fraction of  $H_2$  in the air is between 4.0 and 75% the burning gas temperature is 803.15 to 953.15 K with most of the authors reporting temperatures between 833.15 and 857.15 K. The ignition temperature is a function of the local species concentrations and of the pressure of the mixture. *Belles* reported in 1958 a method for prediction of the ignition temperature based on equalizing the rate of the chain-branching reaction to the half of the rate of the chain-breaking reaction resulting in a transcendental equation

$$\frac{3.11T_{1,ign} \exp(-8550/T_{1,ign})}{Y_X p_{in\_atm}} = 1, \quad (6.103)$$

where the effective mole fraction of the third bodies for the formation of  $HO_2$  radicals is

$$Y_X = Y_{H_2} + 0.35Y_{O_2} + 0.43Y_{N_2} + 0.20Y_{Ar} + 1.47Y_{CO_2}. \quad (6.104)$$

In a similar way *Maas* and *Wanatz* proposed in 1988 a method for explicit computation of the ignition temperature defined as a temperature dividing the slow reaction from the rapid reaction region. *Moser* (1997) p. 38 approximated in 1997 this method by the following equation

$$T_{1,ign} = 1.03015 \times 10^3 + 1.08375 \times 10^2 \ln \left\{ 8.47444 \times 10^{-2} + p_{bar} \zeta \left[ 1 + p_{bar} \zeta \left( 7.40849 \times 10^{-2} + p_{bar} \zeta 1.45577 \times 10^{-3} \right) \right] \right\}, \quad (6.105)$$

where

$$\zeta = Y_H + 0.3Y_{O_2} + 0.5Y_{N_2} + 6.5Y_{H_2O}, \quad (6.106)$$

is the stokes efficiency and  $p_{bar} = p / 10^5$ . The approximation is reported to be valid in the region of  $p_{bar} = 1$  to 40 bar,  $T_1 = 700$  to 2500 K for a fuel equivalence ratio

$$\Phi = 0.2 \text{ to } 0.4, \quad (6.107)$$

with a stoichiometric mass concentration

$$C_{st,H_2} = 0.02818 \quad (6.108)$$

in air.

As already mentioned the ignition may start at a much lower temperature but the velocity of the reaction is very low. That is why the term ignition temperature is associated with a single-step reaction because the steps absorbing energy

necessary to create mutually interacting radicals are absorbed in a single reaction which simply releases the energy. The term ignition temperature is not necessary if one uses appropriate multi-step reactions and in addition resolves the spatial variation of the controlling variables properly. In this case, as in nature, if the local energy dissipation is higher than the locally released energy the reaction cannot propagate.

### 6.8.3 Detonability concentration limits

For the history of the detonation analysis see *Laffitte* (1938), *Wendlandt* (1924), *Chapman* (1899), *Jouguet* (1905), *Crussard* (1907) and *Zeldovich* (1940). *Belles* (1958) analyzed how strong a shock wave must be in order to obtain parameters satisfying Eq. (6.86). Then he computed the enthalpy of the gas behind the shock for different hydrogen concentrations and compared it with the available enthalpy increase due to combustion.

**Table 6.1.** Predicted and observed limits of detonability of hydrogen mixtures at 300K and 1 atm

System	Lean limit, molar concentration $H_2$ in %	Rich limit, molar concentration $H_2$ in %
$H_2-O_2$	16.3 (15*)	92.3 (90*)
$H_2-Air$	15.8 (18.3*)	59.7 (58.9*)

\* Experimentally observed detonability limits by *Mallard* and *Le Chatelier* (1881).

As a result he obtained the detonability limits given in Table 6.1.

### 6.8.4 The heat release due to combustion

The *heat* release per 1 kg of the final reaction product (in this case steam) is usually called the *enthalpy* of formation of steam. One should pay attention to the exact definition of the measured values. *Dorsey* (1951) provided a review of all the measurements up to 1940 and selected the value

$$\Delta h_{H_2O} [T_{H_2O,ref} = (273.1 + 15) K, p_{ref} = 1.01325 \text{ bar}] = 13\,425\,000 \text{ J/kg.} \quad (6.109)$$

defined as follow:

Enthalpy of formation of steam = the heat produced after reaction of 1/9 kg hydrogen with 8/9 kg oxygen at the condition  $T_{H_2O,ref} = 288.1K$  and  $p_{H_2O,ref} = 1.101325 \text{ bar}$  and then removed from the produced 1kg water-steam in order to obtain the same initial temperature and pressure.

Note that the enthalpy of formation of water is about six times larger than the latent heat of evaporation. The enthalpy of formation of steam is almost not a func-

tion of the pressure but is a slight function of the temperature due to the differences of the specific capacities at constant pressure of the stoichiometric mixture of the initial products and of the resulting steam.

$$\begin{aligned}
 h_{H_2O} &= \Delta h_{H_2O,ref} - \int_{T_{H_2O,ref}}^T \left[ c_{p,H_2O} - c_{p,(H_2+O_2)_{stoichiometric}} \right] dT' \approx \\
 &\approx \left[ 240.227 + 3.93108\bar{T} + 6.9003\bar{T}^2 - 3.094968\bar{T}^3 \right] 10^6 / 18.0154 \\
 &= \left( 13.334\ 536\ 01 + 0.118206\bar{T} + 0.383022\bar{T}^2 - 0.1717957\bar{T}^3 \right) 10^6 \text{ J/kg}, \quad (6.110)
 \end{aligned}$$

where  $\bar{T} = T/1000$ , Dorsey (1951). The difference of the measurements of the enthalpy of formation and of the specific heat at constant pressure gives an agreement of the computed values within 10% deviation from different authors. In Zwerev and Smirnov (1987) p. 169 the value  $\Delta h_{H_2O} = 13\ 275\ 363 \text{ J/kg}$  is used, which gives an adiabatic temperature of the burned gases of about 2773 K. In Landau and Lifshitz (1987) p. 655 the value  $\Delta h_{H_2O} = 13\ 333 \text{ kJ/kg}$  is reported. Lewis and von Elbe (1987) p. 696 reported in 1987 the value of 13 281 kJ/kg. Cox recommended in 1989 the following value  $\Delta h_{H_2O}(298.15K, 1atm) = 13\ 431 \pm 0.017\% \text{ kJ/kg}$ . A summary is given in Table 6.2. Further information on this topic can be found in Zwerev and Smirnov (1987), Cox (1989) and Lewis and von Elbe (1987).

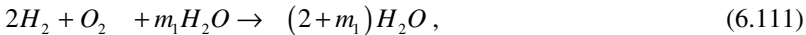
**Table 6.2.** Enthalpy of formation of steam

Author	$\Delta h_{H_2O}$ in MJ/kg	$T_{burned\ gases}$ in K
Dorsey (1951)	13.425	
Zwervev and Smirnov (1987) p. 169	13.275	2 773*
Landau and Lifshitz (1987) p. 655	13.333	
Lewis and von Elbe (1987) p. 696	13.281	
Cox (1989)	13.435 $\pm$ 0.017%	

\*computed

## 6.8.5 Equilibrium dissociation

**Identification of the chemical system.** Consider the idealized single-step reaction



followed by dissociation. Six compounds are usually identified in the chemical literature as important with their chemical identification symbols

$$Symb_{j=1,6} = (H_2O, H_2, O_2, HO, H, O) \quad (6.112)$$

consisting of two different chemical elements ( $O$ ,  $H$ ). Therefore  $6 - 2 = 4$  chemical reactions are required to describe the system. Water dissociates in two ways,



Hydrogen and oxygen dissociate as follows



The four expressions can be rewritten formally as

$$2 \times H_2O - 2 \times H_2 - 1 \times O_2 + 0 \times HO + 0 \times H + 0 \times O = 0 \quad (6.117)$$

$$2 \times H_2O - 1 \times H_2 + 0 \times O_2 - 2 \times HO + 0 \times H + 0 \times O = 0 \quad (6.118)$$

$$0 \times H_2O + 1 \times H_2 + 0 \times O_2 + 0 \times HO - 2 \times H + 0 \times O = 0 \quad (6.119)$$

$$0 \times H_2O + 0 \times H_2 + 1 \times O_2 + 0 \times HO + 0 \times H - 2 \times O = 0, \quad (6.120)$$

or

$$\sum_{j=1}^6 n_{ij} \text{Symb}_j = 0, \quad \text{for } i = 1, 4. \quad (6.121)$$

Here  $n_{ij}$  is the stoichiometric coefficient,  $< 0$  for reactants and  $> 0$  for products. The temperature  $T$  and the total pressure  $p$  of the mixture of compounds are given. We look for a solution of six molar concentrations

$$\mathbf{Y}^T = (Y_{H_2O}, Y_{H_2}, Y_{O_2}, Y_{OH}, Y_H, Y_O), \quad (6.122)$$

for which the system is in a chemical equilibrium. We know from *Dalton's* law that the system pressure is the sum of the partial pressures

$$p = \sum_{j=1}^6 p_j, \quad Y_i = \frac{p_j}{p}, \quad \sum_{j=1}^6 Y_j = 1 \quad (6.123)$$

and that the total number of hydrogen atoms is twice that of oxygen in non-dissociated as well in dissociated water

$$\frac{2Y_1 + 2Y_2 + Y_4 + Y_5}{Y_1 + 2Y_3 + Y_4 + Y_6} = 1. \quad (6.124)$$

We need four additional equations to close the system. The four equations are the enforced chemical equilibrium for the considered four reactions for the prescribed temperature  $T$  and pressure  $p$ . The chemical equilibrium condition is a complicated relation between partial pressure ratios and the system pressure and temperature. For a mixture of perfect fluids the resulting system consists of implicit transcendental equations. It is solvable by iteration. Now we consider this task in more detail.

**Chemical equilibrium conditions.** The so called *pressure equilibrium factors* are defined as follows:

$$K_p = \prod_{i=1}^{i_{\max}} p_i^{n_i} = p^{\sum_{i=1}^{i_{\max}} n_i} \prod_{i=1}^{i_{\max}} Y_i^{n_i}. \quad (6.125)$$

For a reference pressure equal to  $1\text{bar}$  and for pressures in dimensions of  $\text{bar}$  we have

$$K_{p1} = \frac{1}{p} \frac{Y_{H_2O}^2}{Y_{H_2}^2 Y_{O_2}} \quad (6.126)$$

$$K_{p2} = \frac{1}{p} \frac{Y_{H_2O}^2}{Y_{H_2} Y_{HO}^2}, \quad (6.127)$$

$$K_{p3} = \frac{1}{p} \frac{Y_{H_2}}{Y_H^2}, \quad (6.128)$$

$$K_{p4} = \frac{1}{p} \frac{Y_{O_2}}{Y_O^2}. \quad (6.129)$$

Each of the pressure equilibrium factors can be computed for the reference pressure and temperature  $p_0 = 1\text{bar}$  and  $T_0 = 0\text{K}$  using the chemical equilibrium condition

$$\ln K_{pi} = -\frac{\Delta H_{ch,0,i}}{TR} + \sum_{j=1}^{j_{\max}} n_{ij} \frac{s_{j0} M_j}{R} = -\frac{\Delta H_{ch,0,i}}{TR} + \sum_{j=1}^{j_{\max}} n_{ij} s_{0j} / R_j. \quad (3.130)$$

Here the  $R$  is the universal gas constant,  $R_j$  is the gas constant of the  $j$ -th component,

$$\Delta H_{ch,0,i} = \sum_{j=1}^{j_{\max}} h_{0j} n_{ij} M_j \quad (6.131)$$

and

$$\bar{s}_{0j} = s_{0j} / R_j \quad (6.132)$$

are the dimensionless reference entropies of the species  $j$ . Equation (6.130) is applied for each of the four selected chemical reactions resulting in

$$\ln K_{p1} = -\left(\Delta H_{ch,0,1}/R\right)/T - 2\bar{s}_{01} + 2\bar{s}_{02} + \bar{s}_{03}, \quad (6.133)$$

$$\ln K_{p2} = -\left(\Delta H_{ch,0,2}/R\right)/T - 2\bar{s}_{01} + \bar{s}_{02} + 2\bar{s}_{04}, \quad (6.134)$$

$$\ln K_{p3} = -\left(\Delta H_{ch,0,3}/R\right)/T - \bar{s}_{02} + 2\bar{s}_{05}, \quad (6.135)$$

$$\ln K_{p4} = -\left(\Delta H_{ch,0,4}/R\right)/T - \bar{s}_{03} + 2\bar{s}_{06}. \quad (6.136)$$

**Thermo-physical data required.** Thermo-physical data are available in the chemical literature e.g. *Chase* (1998) and *Robert, Rupley and Miller* (1987). Examples for molar masses and the enthalpies of formation for a reference pressure of 1bar for each of the compounds are given in Table 6.3.

**Table 6.3.** The molar mass and the enthalpy of formation for components of water vapor dissociation at pressure  $p = 0.1\text{MPa}$ , by *Chase et al.* (1998)

$j$	Symbol	$M_j$	$\bar{\Delta}h_j,$	$\Delta h_j$	$\bar{\Delta}h_j,$	$\Delta h_j$
		kg/kg-mole	$T = 0K$	$T = 0K$	$T = 298.15K$	$T = 298.15K$
			MJ/kg-mole	MJ/kg	MJ/kg-mole	MJ/kg
1	H <sub>2</sub> O	18.01528	-238.921	13.262	-241.826	13.424
2	H <sub>2</sub>	2.01588	0	0	0	0
3	O <sub>2</sub>	31.9988	0	0	0	0
4	HO	17.00734	38.39	2.257	38.987	2.292
5	H	1.00794	216.035	214.333	217.999	216.28
6	O	15.9994	246.79	15.425	249.173	15.574

Using the data in the table we obtain

$$\Delta H_{ch,0,1}/R = 57.471546K, \quad (6.137)$$

$$\Delta H_{ch,0,2}/R = 66.706116K, \quad (6.138)$$

$$\Delta H_{ch,0,3}/R = 51.966405K, \quad (6.139)$$

$$\Delta H_{ch,0,4}/R = 59.364404K. \quad (6.140)$$

*Ihara* (1977, 1979) provided the following approximation for the dimensionless entropies at the reference pressure and temperature for each species.



$$\begin{aligned}\bar{s}_{01} = \bar{s}_{0,H_2O} = & -\ln[1 - \exp(-5262/T)] - \ln[1 - \exp(-2294/T)] \\ & - \ln[1 - \exp(-5404/T)] + 4\ln(T) - 4.1164 - \ln(p/10^5),\end{aligned}\quad (6.141)$$

$$\begin{aligned}\bar{s}_{02} = \bar{s}_{0,H_2} = & \ln \left[ \frac{T}{682.6} + \frac{1}{24} + \frac{0.711}{T} + \frac{104}{T^2} + \left( \frac{T}{227.53} + 0.875 \right) \exp\left(-\frac{171}{T}\right) \right. \\ & \left. + \ln \left\{ \left[ 1 - \exp\left(-\frac{6338}{T}\right) \right]^{-1} + \frac{360.65}{T} \exp\left(-\frac{6338}{T}\right) \left[ 1 - \exp\left(-\frac{6338}{T}\right) \right]^{-3} \right\} \right] \\ & + 2.5\ln(T) - 2.6133 - \ln(p/10^5),\end{aligned}\quad (6.142)$$

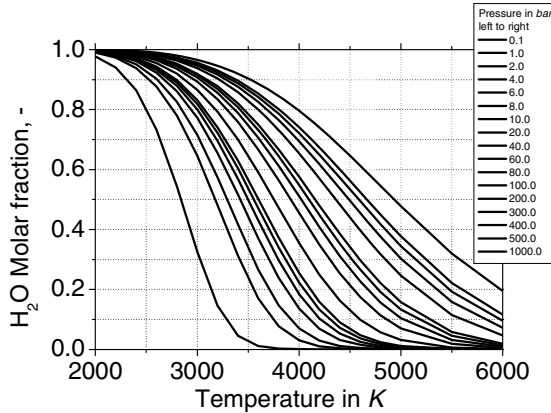
$$\begin{aligned}\bar{s}_{03} = \bar{s}_{0,O_2} = & -\ln[1 - \exp(-2239/T)] \\ & + \ln[3 + 2\exp(-11340/T) + \exp(-18878/T)] \\ & + 3.5\ln(T) + 0.114 - \ln(p/10^5),\end{aligned}\quad (6.143)$$

$$\begin{aligned}\bar{s}_{04} = \bar{s}_{0,OH} = & \ln(T/26.638 + 1/3 + 1.776/T) - \ln[1 - \exp(-5136/T)] \\ & + \ln[1 + \exp(-201/T)] + 2.5\ln(T) + 1.2787 - \ln(p/10^5),\end{aligned}\quad (6.144)$$

$$\bar{s}_{0,H} = 2.5\ln(T) - 2.9604 - \ln(p/10^5),\quad (6.145)$$

$$\begin{aligned}\bar{s}_{0,O} = & \ln[5 + 3\exp(-228/T) + \exp(-326/T) + 5\exp(-22830/T)] \\ & + 2.5\ln(T) + 0.4939 - \ln(p/10^5).\end{aligned}\quad (6.146)$$

Figure 6.3 demonstrates the result of the solution of the system of equations (6.123), (6.124), (6.126) through (6.129) for different temperatures and pressures with the algorithm developed by *Vasic* (1993). For comparison see the method reported by *Kesselman* et al. (1968) and the tables provided by *Vargaftik* (1983). We immediately realize how important it is to take into account the dissociation physics in analyzing combustion processes leading to temperatures higher than 1600K.



**Fig. 6.3** Molar fraction of dissociated steam as a function of temperature with pressure as a parameter

Thus initially we have a gas consisting of

$$C_{1,H_2} + C_{1,O_2} + C_{1,H_2O} = 1, \quad (6.147)$$

and finally a mixture of gases and radicals

$$C_{2,H_2} + C_{2,O_2} + C_{2,H_2O} + C_{2,H} + C_{2,O} + C_{2,OH} = 1. \quad (6.148)$$

Note that

$$C_{1,H_2} + C_{1,O_2} - C_{2,H_2} - C_{2,O_2} - C_{2,H} - C_{2,O} - C_{2,OH} = C_{2,H_2O} - C_{1,H_2O} \quad (6.149)$$

reflects in fact the net generation of stable steam and therefore

$$h_{formation} = (\Delta h_{ref} - c_{p2} T_{ref}) (C_{2,H_2O} - C_{1,H_2O}). \quad (6.150)$$

### 6.8.6 Source terms of the energy conservation of the gas phase

Bearing in mind the information already introduced the expression in the brackets in Eq. (6.84) can then be rearranged as follows:

$$\begin{aligned} \dot{q}''' + \sum_{i=1}^{i_{max}} \mu_i (h_i^* - h_i) + T \sum_{i=2}^{i_{max}} \mu_i (R_i - R_1) \\ = \dot{q}'''_{ref} + \mu_{H_2O} (h_{H_2O,ref} - h_{H_2O}) + T \left[ \mu_{H_2} (R_{H_2} - R_{H_2O}) + \mu_{O_2} (R_{O_2} - R_{H_2O}) \right] \\ = \dot{q}'''_{ref} + \mu_{H_2O} (h_{H_2O,ref} - h_{H_2O}) + T (\mu_{H_2} R_{H_2} + \mu_{O_2} R_{O_2} + \mu_{H_2O} R_{H_2O}). \end{aligned} \quad (6.151)$$

Note that for perfect gases  $\mu RT = \mu pv$  which results in a very interesting interpretation of the term

$$T(\mu_{H_2} R_{H_2} + \mu_{O_2} R_{O_2} + \mu_{H_2O} R_{H_2O}) \equiv \mu pv, \quad (6.152)$$

which is in fact the  $\mu pv$  work associated with the disappearance and appearance of components with respect to zero reference pressure.

Expressing on the right hand side all mass sources as a function of the hydrogen mass source only results in

$$\begin{aligned} &= \dot{q}_{ref}''' + \mu_{H_2} \left[ - \left( 1 + \frac{1}{2} \frac{M_{O_2}}{M_{H_2}} \right) (h_{H_2O,ref} - h_{H_2O}) \right. \\ &\quad \left. + T \left( R_{H_2} + \frac{1}{2} \frac{M_{O_2}}{M_{H_2}} R_{O_2} - \left( 1 + \frac{1}{2} \frac{M_{O_2}}{M_{H_2}} \right) R_{H_2O} \right) \right] \\ &= \dot{q}_{ref}''' - \mu_{H_2} \left[ 9(h_{H_2O,ref} - h_{H_2O}) - T(R_{H_2} + 8R_{O_2} - 9R_{H_2O}) \right]. \end{aligned} \quad (6.152)$$

Keeping in mind that

$$h_{H_2O,ref} - h_{H_2O} = c_{p,H_2O} (T_{H_2O,ref} - T), \quad (6.153)$$

and

$$\dot{q}_{ref}''' = \mu_{H_2O} \Delta h_{H_2O} = - (9\mu_{H_2}) \Delta h_{H_2O}, \quad (6.154)$$

we finally obtain

$$\begin{aligned} &= \dot{q}_{ref}''' - \mu_{H_2} \left[ 9c_{p,H_2O} (T_{H_2O,ref} - T) - T(R_{H_2} + 8R_{O_2} - 9R_{H_2O}) \right] \\ &= \dot{q}_{ref}''' + \mu_{H_2} \left[ T(R_{H_2} + 8R_{O_2} - 9R_{H_2O} + 9c_{p,H_2O}) - 9c_{p,H_2O} T_{H_2O,ref} \right] \\ &= -\mu_{H_2} \left[ 9(\Delta h_{H_2O} + c_{p,H_2O} T_{H_2O,ref}) - T(R_{H_2} + 8R_{O_2} - 9R_{H_2O} + 9c_{p,H_2O}) \right] \\ &= -\mu_{H_2} (1.19478267 \times 10^8 - 10515.3 \times T). \end{aligned} \quad (6.155)$$

Consequently the term  $\mu_{H_2} 10515.3 \times T < 0$  obviously has a cooling effect. At  $T = 3000K$  the cooling effect amounts to about 26% of the origination enthalpy.

### 6.8.7 Temperature and pressure changes in a closed control volume; adiabatic temperature of the burned gases

Thus, the temperature and pressure changes of the burning mixture in a closed volume are governed by the equations

$$\frac{1}{aT - cT^2} \frac{dT}{d\tau} = -\mu_{H_2} (\kappa - 1) \frac{1}{p}, \quad (6.156)$$

$$\frac{dp}{d\tau} = -\mu_{H_2} (\kappa - 1)(a - bT), \quad (6.157)$$

where  $a = 1.2 \times 10^8$ ,  $c = 10515.3$ ,  $b = 8466 + \frac{\kappa}{\kappa - 1} 2049.3$ . The hydrogen and oxygen concentrations change obeying the mass conservation

$$\frac{dC_{H_2}}{d\tau} = -\mu_{H_2} / \rho, \quad (6.158)$$

$$\frac{dC_{O_2}}{d\tau} = -8\mu_{H_2} / \rho, \quad (6.159)$$

where

$$\rho = \rho_0 = \text{const}. \quad (6.160)$$

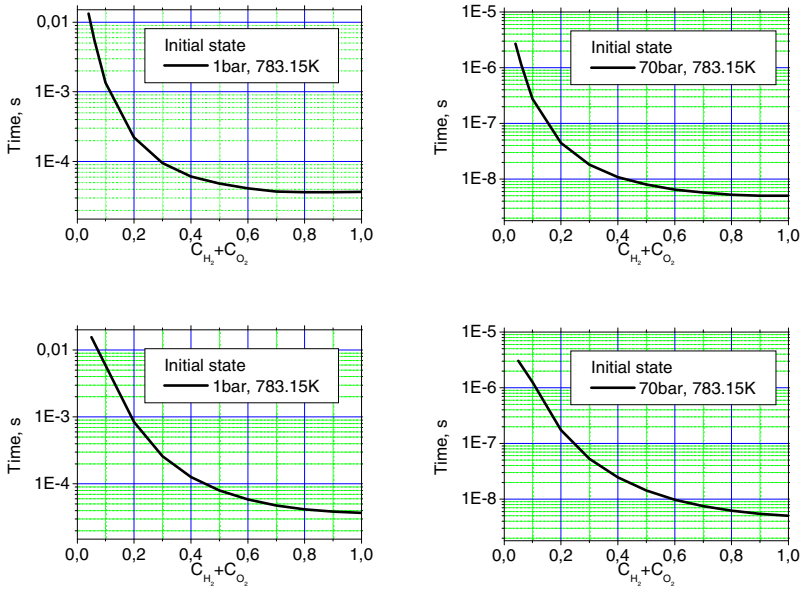
Therefore

$$\int_0^{\Delta\tau} \mu_{H_2} d\tau = -\rho_0 \Delta C_{H_2}. \quad (6.161)$$

Given the initial conditions, the system can be integrated numerically until the concentrations of the fuel and the oxidizer are below given limits for which no more oxidation is possible. The duration of the process gives the inherent time scale of the burning process as presented in Fig. 6.4. The higher the initial pressure the faster is the burning. This information is important for selecting of an appropriate integration time step for large scale computational analysis.

The hydrogen mass source term can be eliminated from Eqs. (6.156) and (6.157) and the equation obtained written in a compact form

$$\frac{a - bT}{aT - cT^2} dT = d \ln p \quad \text{or} \quad d \ln \frac{Tb^{1/c}}{(a - cT)^{\frac{c-1}{c}}} = d \ln p \quad (6.162)$$



**Fig. 6.4** Time elapsed from ignition to complete burning as a function of the initial mass concentration of the stoichiometric mixture fuel + oxidizer. Initial temperature 783.15K. Inert component is nitrogen. a) 1bar; b) 70bar; Inert component is steam: c) 1bar; d) 70bar

can be integrated analytically assuming that  $\kappa \approx const$ . The result

$$\frac{T}{T_0} \frac{(a - cT_0)^{\frac{c-1}{c}}}{(a - cT)^{\frac{c-1}{c}}} = \frac{p}{p_0} \quad (6.163)$$

can be further simplified by noting that  $c \gg 1$

$$\frac{T}{T_0} = \frac{p}{p_0} \left/ \left( 1 + \frac{c}{a} T_0 \frac{p - p_0}{p_0} \right) \right. \quad (6.164)$$

Bearing in mind that  $1 \gg \frac{c}{a} T_0 \frac{p - p_0}{p_0}$  the solution converges to those of heat input

only in a closed volume obtained previously. Using Eq. (6.163) and substituting for the pressure into Eq. (6.156) we obtain

$$\frac{1}{(a - cT)^m} \frac{dT}{d\tau} = -\mu_{H_2} \frac{T_0 (\kappa - 1)}{p_0 (a - cT_0)^{\frac{c-1}{c}}}, \quad (6.165)$$

where  $m = 2 - 1/c \approx 2$ . After integrating with respect to time and solving with respect to the temperature we obtain

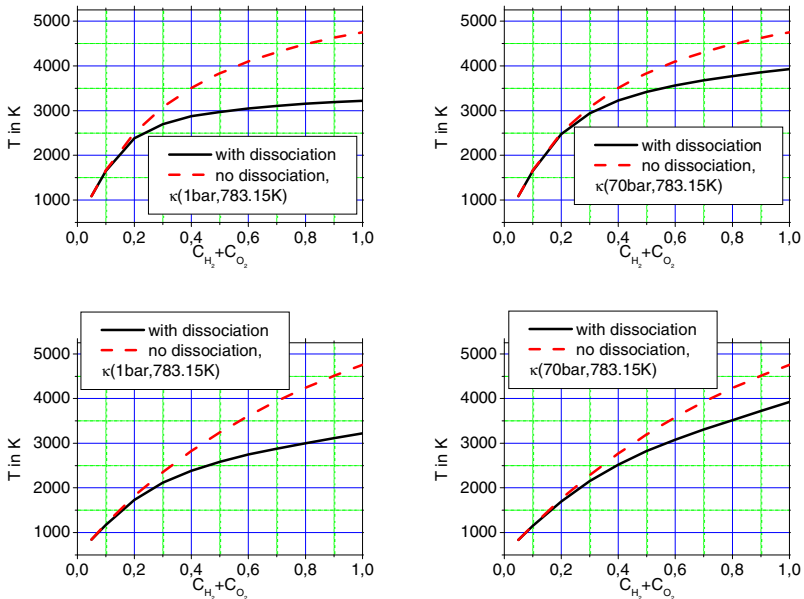
$$T = \frac{1}{c} \left\{ a - \left[ \frac{1}{(a - cT_0)^{m-1}} - \frac{(\kappa - 1)c(m-1)}{(a - cT_0)^{\frac{c-1}{c}}} \frac{T_0}{p_0} \int_0^{\Delta\tau} \mu_{H_2} d\tau \right]^{\frac{1}{m-1}} \right\}$$

$$\approx \frac{1}{c} \left\{ a - \frac{a - cT_0}{1 - (\kappa - 1)c \frac{T_0}{p_0} \int_0^{\Delta\tau} \mu_{H_2} d\tau} \right\}. \quad (6.166)$$

For the limiting case of complete consumption of the fuel or of the oxidizer over the time step  $\Delta\tau$  we have

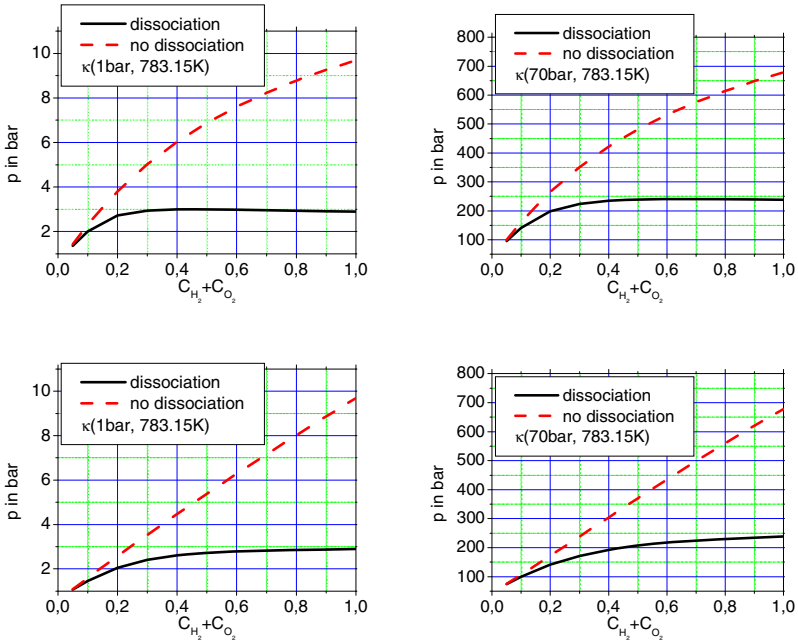
$$T = \frac{1}{c} \left\{ a - \left[ \frac{1}{(a - cT_0)^{m-1}} + \frac{(\kappa - 1)c(m-1)}{(a - cT_0)^{\frac{c-1}{c}}} \frac{T_0 \rho_0}{p_0} \min(\Delta C_{H_2}, \Delta C_{O_2}/8) \right]^{\frac{1}{m-1}} \right\}$$

$$\approx a/c - \frac{a/c - T_0}{1 + (\kappa - 1)c \frac{T_0 \rho_0}{p_0} \min(\Delta C_{H_2}, \Delta C_{O_2}/8)}. \quad (6.167)$$



**Fig. 6.5** Temperature after burning as a function of the initial mass concentration of the stoichiometric mixture fuel + oxidizer. Initial temperature 783.15 K. Initial pressure: inert component is nitrogen. a) 1 bar; b) 70 bar; inert component is steam: c) 1 bar; d) 70 bar

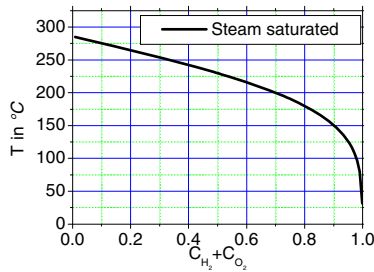
This result is valid for the case of no steam dissociation, i.e. for a burned products (steam) temperature below 1600 K. For higher temperatures  $\kappa$  and  $R$  are no longer constant and the integration has to be performed numerically. Note that for temperature higher than 1600 K the specific capacity at constant pressure changes dramatically due to the thermal dissociation and the “effective isentropic exponent” tends to unity which clearly reduces the right hand side of Eqs. (6.156) and (6.157). Figures 6.5 and 6.6 demonstrate the error in the final temperature and pressure made if the dissociation is not taken into account. For low pressure the difference in the final temperatures may be higher than 1500 K, and for high initial pressure the difference in the final pressures may be greater than 450 bar.



**Fig. 6.6** Pressure after burning as a function of the initial mass concentration of the stoichiometric mixture fuel + oxidizer. Initial temperature 783.15 K. Initial pressure: inert component is nitrogen: a) 1 bar; b) 70 bar; inert component is steam: c) 1 bar; d) 70 bar

An easy method of indication of combustible gases in pipelines intended to work with saturated steam is to measure their temperature. The accumulation of combustible gases is manifested in a temperature reduction. An example is given in Fig. 6.7 for steam lines at 70 bar nominal pressure.

In a coordinate system moving with the burning front velocity  $w_{1,b}$ , the temperature before the front is governed by heat conduction mainly,  $-w_{1,b} dT_1/dz - \lambda_1 d^2T_1/dz^2 = 0$ , with the solution  $T_1 = T_{10} \exp(-w_{1,b}z / \lambda_1)$ , where  $T_{10}$  is the flame temperature at the burning front, *Landau and Lifschitz* (1987) p. 645.



**Fig. 6.7** Initial temperature – saturated temperature under the given partial steam pressure as a function of the mass concentration of a stoichiometric hydrogen-oxygen mixture with initial presence of steam. Initial total pressure  $70\text{bar}$

Here are some velocities characterizing the interaction process for stoichiometric mixture: The propagation of a laminar combustion front is reported to be about  $1$  to  $30\text{ m/s}$  see *Landau and Lifschitz* (1987) p. 639 and *Schmidt* (1945), respectively. The sound velocity in stagnant mixture at  $783.15\text{K}$  is about  $664\text{m/s}$ . The detonation velocity of  $2819\text{ m/s}$  was measured by *Lewis and Friauf* in 1930 for initial mixture temperature of  $291\text{K}$  (see also *Oswatitsch* (1952) p.59). The authors computed a temperature of the burning products of about  $3583\text{K}$ . From these data we see that the propagation velocity of a combustion front is very small compared to the sound velocity and to the detonation velocity.

## 6.9 Constituents of sodium vapor

Water vapor at boiling point consists 100% of water molecules. In contrast, at the boiling point of sodium 11% of the matter is not in a mono-atomic state – the normal state at high temperatures (*Cordfunke and Konings* 1990). The sodium vapor is, therefore, a mixture of mono-atomic (monomer), diatomic (dimer), and probably four-atomic (tetramer) components with ion and electron components as well. The general way to compute the constituency of sodium vapor is associated with the classical approach of chemical thermodynamics for computing the equilibrium mole concentrations as shown in Chapter 4.2.4.3. There are differences in the literature in the implementation of this method by considering different number of components and the corresponding reactions. As an example I give here the approach clearly stated by *Golden and Tokar* (1967).

Given the temperature  $T$  and the total pressure  $p$  of a mixture of  $j_{\max} = 3$  compounds that may react in a number  $i_{\max} = 2$  of chemical reactions,

$$\sum_{j=1}^{j_{\max}} n_{ij} \text{Symb}_j = 0, \quad \text{for } i = 1, i_{\max}, \quad (6.168)$$



or in particular

$$-2Na + Na_2 = 0, \text{ reaction Nr. 1,} \quad (6.169)$$

$$-4Na + Na_4 = 0, \text{ reaction Nr. 2.} \quad (6.170)$$

Here  $n_{ij}$  is the stoichiometric coefficient,  $< 0$  for reactants and  $> 0$  for products. We look for a solution of three molar concentrations

$$\mathbf{Y}^T = (Y_1, Y_2, Y_4), \quad (6.171)$$

for which the system is in chemical equilibrium. We know from *Dalton's* law that the system pressure is the sum of the partial pressures

$$p = p_1 + p_2 + p_4, \quad (6.172)$$

$$Y_i = \frac{p_i}{p}, \quad (6.173)$$

$$Y_1 + Y_2 + Y_4 = 1. \quad (6.174)$$

For each chemical reaction  $i$  we have the condition enforcing chemical equilibrium, see Eq. (3.174),

$$\begin{aligned} K_{p_2} &= \exp \left[ -\frac{1}{TR} \left[ -2(h_{10} - Ts_{10})M_1 + (h_{20} - Ts_{20})M_2 \right] \right] \\ &= \exp \left[ -2(s_{10} - s_{20})/R_1 + \frac{2(h_{10} - h_{20})/R_1}{T} \right] = \exp(e_{a1} + e_{a2}/T), \end{aligned} \quad (6.175)$$

$$\begin{aligned} K_{p_4} &= \exp \left[ -\frac{1}{TR} \left[ -4(h_{10} - Ts_{10})M_1 + (h_{40} - Ts_{40})M_4 \right] \right] \\ &= \exp \left[ -4(s_{10} - s_{40})/R_1 + \frac{4(h_{10} - h_{40})/R_1}{T} \right] = \exp(e_{b1} + e_{b2}/T), \end{aligned} \quad (6.176)$$

with the chemical equilibrium factors defined as follows:

$$K_{p2} = \frac{p_0 Y_2}{p Y_1^2}, \quad (6.177)$$

$$K_{p4} = \frac{p_0^3 Y_4}{p^3 Y_1^4}. \quad (6.178)$$

Using the molar enthalpy of demerization and tetramerization given above recomputed per kg

$$h_{20} - h_{10} = \Delta h_{Na_2}^{2Na} / (2M_1) = -1.666315 \times 10^6 \text{ J/kg}, \quad (6.179)$$

$$h_{40} - h_{20} = \Delta h_{Na_4}^{4Na} / (4M_1) = -1.887987 \times 10^6 \text{ J/kg}, \quad (6.180)$$

results in

$$e_{a2} = 2(h_{10} - h_{20})/R_1 = 9117.50383 \text{ K}, \quad (6.181)$$

$$e_{b2} = 4(h_{10} - h_{40})/R_1 = 20660.83388 \text{ K}, \quad (6.182)$$

which are the values used here. *Stone et al. (1965)* used values for the enthalpy and entropy changes of the mixture due to each chemical reaction derived from their  $p$ - $v$ - $T$  data. They come to  $e_{a2} = 9217.72 \text{ K}$  and  $e_{b2} = 20772.05 \text{ K}$ , which are slightly different. For the entropy terms these authors come to

$$e_{a1} = -2(s_{10} - s_{20})/R_1 = -9.95845, \quad (6.183)$$

$$e_{b1} = -4(s_{10} - s_{40})/R_1 = -24.59115, \quad (6.184)$$

related to reference pressure  $p_0 = 1 \text{ atm}$ , which means  $s_{20} - s_{10} = -1820.006322 \text{ J/(kgK)}$ , and  $s_{40} - s_{10} = -2247.139287 \text{ J/(kgK)}$ . These values are used by *Hame (1986)*. Recomputed to  $p_0 = 1 \text{ Pa}$  results in  $e_{a1} = -21.4845$ ,  $e_{b1} = -59.1694$  which means  $s_{20} - s_{10} = -3926.50722 \text{ J/(kgK)}$ ,  $s_{40} - s_{10} = -5406.899772 \text{ J/(kgK)}$ . Therefore, for given pressure and temperature the algebraic system of Eqs. (6.174), (6.177), and (6.178) completely defines the mole concentrations. Replacing in Eq. (6.174) the mole concentrations

$$Y_2 = K_{p2} p Y_1^2, \quad (6.185)$$

$$Y_4 = K_{p4} p^3 Y_1^4 \quad (6.186)$$

results in the fourth-order equation

$$K_{p4} \frac{p^3}{p_0^3} Y_1^4 + K_{p2} \frac{p}{p_0} Y_1^2 + Y_1 - 1 = 0, \quad (6.187)$$

or

$$a_4 Y_1^4 + a_2 Y_1^2 + Y_1 - 1 = 0, \quad (6.188)$$

where

$$a_2 = K_{p2} p = p \exp(e_{a1} + e_{a2}/T), \quad (6.189)$$

$$a_4 = K_{p4} p^3 = p^3 \exp(e_{b1} + e_{b2}/T). \quad (6.190)$$

The solution of the fourth-order equation is performed by iteration using *Newton's* method starting with  $Y_1 = 0.8$  and iterating

$$f = Y_1 + a_2 Y_1^2 + a_4 Y_1^4 - 1, \quad (6.191)$$

$$df/dY_1 = 1 + 2a_2 Y_1 + 4a_4 Y_1^3, \quad (6.192)$$

$$\Delta Y_1 = f/(df/dY_1), \quad (6.193)$$

$$Y_1 = Y_1 - \Delta Y_1, \quad (6.194)$$

until  $|\Delta Y_1| < \varepsilon Y_1$ , where  $\varepsilon = 10^{-6}$ . So at each given temperature and pressure the mole concentrations are uniquely defined. The mixture molar mass is usually computed by assuming instantaneous adjustment of chemical equilibrium for each pressure and temperature. Therefore,

$$M(p, T) = Y_1 M_1 + Y_2 M_2 + Y_4 M_4 = M_1 (Y_1 + 2Y_2 + 4Y_4) = M_1 (4 - 3Y_1 - 2a_2 Y_1^2). \quad (6.195)$$

Knowing the mole concentrations  $Y_i$ , the molar masses  $M_i$ , and the mixture molar mass  $M$  the component mass concentrations  $C_i$  are then

$$C_i = Y_i M_i / M. \quad (6.196)$$

The method presented here for computation of the components of the gas mixture is sometimes called the *quasi-chemical method* in the literature. Thus, knowing the concentration of the constituents at each pressure and temperature, the calorific properties of the vapor can be computed.

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# 7 Exergy of multi-phase multi-component systems

## 7.1 Introduction

Fluids at pressures and temperatures higher than the environment pressures and temperatures may perform technical work at the costs of their internal energy. Experience shows that not all available internal energy of fluids may be consumed for performing technical work but only part of it. The industrial revolution initiated with the invention of the steam machine started also the discussion on how much of the internal fluid energy may be transferred under given circumstances in technical work. The result of this discussion is well presented in the references and text books *Baer* (1996), *Elsner* (1974), *Gouy* (1889), *Rant* (1956-1964), *Reynolds* and *Perkins* (1977), *Stephan and Mayinger* (1998) and *Zwicker* (1976). We will shortly demonstrate the main ideas on a single-phase multi-component open system in which the spatial intermixing at any time is assumed to be perfect.

We will emphasize the different definitions of the exergy used in the Anglo-Saxon thermodynamic literature and in the German literature. Thereafter, we will discuss different limiting cases. On the example of the heat pump we will demonstrate a practical application of the exergy.

Finally, we consider the exergy of multi-fluid mixtures for which each of the fluids consists of many chemical components.

## 7.2 The pseudo-exergy equation for single-fluid systems

The energy conservation equation written in a specific enthalpy form is

$$\rho \frac{dh}{d\tau} - \frac{dp}{d\tau} = DT^N, \quad (7.1)$$

where

$$DT^N = \rho(P_k + \varepsilon) + \dot{q}''' + \sum_{i=1}^{i_{\max}} \mu_{i,in} (h_{i,in} - h_i) - \sum_{i=1}^{i_{\max}} \mu_{i,out} (h_{i,out} - h_i) + \frac{1}{2} \left[ \mu_{in} (\mathbf{V}_{in} - \mathbf{V})^2 - \mu_{out} (\mathbf{V}_{out} - \mathbf{V})^2 \right] \quad (7.2)$$

are different components of the energy input into the flow.  $P_k$  and  $\varepsilon$  are irreversible components of the energy dissipation per unit volume of the flow for whatever reason. The thermal energy (heat) introduced per unit time into unit volume of the flow is  $\dot{q}'''$ . The specific energy of the system is influenced by the difference between the in-flowing and the system enthalpy of the chemical components  $\sum_{i=1}^{i_{\max}} \mu_{i,in} (h_{i,in} - h_i)$  as well as by the difference between the out-

flowing and the system enthalpy of the chemical components  $\sum_{i=1}^{i_{\max}} \mu_{i,out} (h_{i,out} - h_i)$ .

Sources and sinks of fluid mass in the system with a velocity different from the system fluid velocity also give rise to the internal energy change,  $\frac{1}{2} \left[ \mu_{in} (\mathbf{V}_{in} - \mathbf{V})^2 - \mu_{out} (\mathbf{V}_{out} - \mathbf{V})^2 \right]$ .  $\mu_{in} \geq 0$  and  $\mu_{out} \geq 0$  are the inflow and outflow mass per unit volume and unit time,  $\mu_i = \mu_{in} - \mu_{out}$ . The complete equivalent to this equation in its entropy form is

$$\rho \frac{ds}{d\tau} = \frac{1}{T} DT^N + \sum_{i=1}^{i_{\max}} \mu_i (s_i - s). \quad (7.3)$$

For the derivation see Chapter 5 or *Kolev (1995, 1997, 1998)*.

In order to come to the exergy definition *Elsner (1974)*, *Reynolds and Perkins (1977)* used the energy conservation equation for an infinite time step *without* irreversible entropy sources and combined it with the entropy conservation equation with irreversible entropy sources. Then they simply eliminate the thermal energy from both equations. In this way the exergy of the thermal energy is not defined, which is a disadvantage of their approach, the reason why we will not follow this approach but only use their exergy definition. Note that in our quantitative statement we take into account in the energy conservation equation the same irreversible terms as in the entropy equations. We emphasize that the two equations are mathematically equivalent (*Legendre* transformed by using the *Gibbs* definition of specific entropy).

Now we multiply the entropy equation by the environment temperature, subtract it from Eq. (7.1), multiply the resulting equation by the time differential and divide by the fluid density. The result is called here the *pseudo-exergy equation*

$$dh - v dp - T_{\infty} ds = v \left[ \left( 1 - \frac{T_{\infty}}{T} \right) DT^N + T_{\infty} \sum_{i=1}^{i_{\max}} \mu_i (s_i - s) \right] d\tau. \quad (7.4)$$

Note the appearance in the right-hand site of the so called efficiency coefficient  $1 - T_\infty / T$  called *Carnot* coefficient – compare with *Carnot* in 1824.

As already mentioned, there are different definitions of the exergy used in the Anglo-Saxon thermodynamic literature and in the German literature. Next we discuss the differences.

## 7.3 The fundamental exergy equation

### 7.3.1 The exergy definition in accordance with *Reynolds* and *Perkins*

Assuming constant environmental conditions,  $p_\infty = \text{const}$  and  $T_\infty = \text{const}$ , *Reynolds* and *Perkins* (1977) defined the specific quantity

$$e_x^{pdv} := e + p_\infty v - T_\infty s, \quad (7.5)$$

as a specific internal *pdv* exergy. *pdv* here is not an exponent but a superscript, whose meaning will be clear in a moment. Note that the exergy is a combination of state variables. Consequently the exergy itself is a state variable. The exergy is a remarkable quantity. As noted by the authors, it is obviously a function of the environmental temperature and pressure besides the two other state variables selected as thermodynamically independent,

$$e_x^{pdv} = e_x^{pdv}(e, v, T_\infty, p_\infty). \quad (7.6)$$

Note the similarity and the difference to the *Gibbs* function  $e + pv - Ts$ , which does not depend on the environmental parameter. The differential form of the exergy is

$$de_x^{pdv} = \left( \frac{\partial e_x^{pdv}}{\partial e} \right)_v de + \left( \frac{\partial e_x^{pdv}}{\partial v} \right)_e dv = \left( 1 - \frac{T_\infty}{T} \right) de + \left( p_\infty - \frac{T_\infty}{T} p \right) dv. \quad (7.7)$$

The last form is obtained having in mind the thermodynamic definition of temperature and pressure, see Eqs. (9) and (20) in *Reynolds* and *Perkins* (1977),

$$\left( \frac{\partial e_x^{pdv}}{\partial e} \right)_v = 1 - T_\infty \left( \frac{\partial s}{\partial e} \right)_v = 1 - \frac{T_\infty}{T}, \quad (7.8)$$

and

$$\left( \frac{\partial e_x^{pdv}}{\partial v} \right)_e = p_\infty - T_\infty \left( \frac{\partial s}{\partial v} \right)_e = p_\infty - \frac{T_\infty}{T} p, \quad (7.9)$$

see in *Reynolds* and *Perkins* (1977). Note that the partial derivative of the exergy with respect to the specific internal energy at constant specific volume, Eq. (7.8) gives exactly the *Carnot* coefficient. The minimum of this function is defined by simultaneously equating the partial derivatives to zero,



$$\left(\frac{\partial e_x^{pdv}}{\partial e}\right)_v = 1 - \frac{T_\infty}{T} = 0, \quad (7.10)$$

and

$$\left(\frac{\partial e_x^{pdv}}{\partial v}\right)_e = p_\infty - \frac{T_\infty}{T} p = 0. \quad (7.11)$$

This leads to the conclusion that the  $pdv$  exergy has an extremum and this extremum is at  $T = T_\infty$  and  $p = p_\infty$ , compare with *Reynolds* and *Perkins* (1977). Because the exergy is decreasing function with decreasing pressure and temperature we conclude that this extreme is a minimum.

### 7.3.2 The exergy definition in accordance with *Gouy* (l'énergie utilisable, 1889)

Assuming constant environmental conditions,  $p_\infty = const$  and  $T_\infty = const$ , *Gouy* in 1889 Eq. 2, p.506 defined the specific  $vdp$  exergy as

$$e_x^{vdp} := h - T_\infty s = e + pv - T_\infty s = e + p_\infty v - T_\infty s + (p - p_\infty)v = e_x^{pdv} + (p - p_\infty)v, \quad (7.12)$$

which differs from the *Reynolds* and *Perkins* (1977) definition

$$e_x^{pdv} := e + p_\infty v - T_\infty s, \quad (7.13)$$

by the residual

$$e_x^{vdp} - e_x^{pdv} = (p - p_\infty)v. \quad (7.14)$$

Note again that  $vdp$  here is not an exponent but a superscript, whose meaning will be clear in a moment. The name  $\epsilon\chi\epsilon\rho\gamma\upsilon\nu$  = available work, was given by *Rant* (1956-1964). Note again the similarity and the difference to the *Gibbs* function  $e + pv - Ts$ , which does not depend on the environmental parameter. Of course for relaxed fluid at environmental conditions  $p = p_\infty$ ,  $e_x^{vdp} = e_x^{pdv}$ , and both definitions possess a minimum. The differential form of the exergy definition is then

$$e_x^{vdp} = e_x^{pdv} + d[(p - p_\infty)v] = \left[1 - \frac{T_\infty}{T} + v\left(\frac{\partial p}{\partial e}\right)_v\right] de + \left[1 - \frac{T_\infty}{T} + \frac{v}{p}\left(\frac{\partial p}{\partial v}\right)_e\right] pdv. \quad (7.15)$$

The difference

$$e - e_x^{vdp} = p_\infty v - T_\infty s, \quad (7.16)$$

is called *anergy* - *Ruppel*. The anergy can not be transferred into mechanical work.

Using the two above introduced definitions of the exergy the pseudo-exergy equation (7.4) can be transferred into two alternative forms of the exergy equation.

### 7.3.3 The exergy definition appropriate for estimation of the volume change work

Using the differential form of the definition equation of the specific enthalpy

$$h := e + pv, \quad (7.17)$$

and replacing the specific enthalpy differential we obtain from Eq. (7.4)

$$de + pdv - T_\infty ds = v \left[ \left( 1 - \frac{T_\infty}{T} \right) DT^N + T_\infty \sum_{i=1}^{i_{\max}} \mu_i (s_i - s) \right] d\tau. \quad (7.18)$$

Replacing with

$$p = p + p_\infty - p_\infty, \quad (7.19)$$

and rearranging we obtain

$$de + p_\infty dv - T_\infty ds + (p - p_\infty) dv = v \left[ \left( 1 - \frac{T_\infty}{T} \right) DT^N + T_\infty \sum_{i=1}^{i_{\max}} \mu_i (s_i - s) \right] d\tau. \quad (7.20)$$

The assumption  $p_\infty = \text{const}$  and  $T_\infty = \text{const}$  allows us to write

$$de + p_\infty dv - T_\infty ds = d(e + p_\infty v - T_\infty s) = de_x \quad (7.21)$$

where the specific quantity  $e_x^{pdv}$ , defined by Eq. (7.5), arises. With this definition of the specific exergy we obtain an abbreviated notation of the integral form of Eq. (7.20)

$$\int_1^2 (p - p_\infty) dv = e_{x,1}^{pdv} - e_{x,2}^{pdv} + \int_{\tau_1}^{\tau_2} v \left[ \left( 1 - \frac{T_\infty}{T} \right) DT^N + T_\infty \sum_{i=1}^{i_{\max}} \mu_i (s_i - s) \right] d\tau. \quad (7.22)$$

The expression

$$-w_i^{pdv} = \int_1^2 (p - p_\infty) dv > 0, \quad (7.23)$$

defines the specific work which is required to change the volume of the fluid and is called *useful work* (Nutzarbeit), see Baer (1996), p. 50. In the literature this work is called also *expansion* or *compression work* depending on the sign. For  $p > p_\infty$  and  $dv > 0$  the expression defines the *specific expansion work per unit mass of the fluid*. Technical expansion work taken away from the fluid is

considered negative per definition. The understanding that Eq. (7.21) is a total differential was clearly expressed first by *Gouy* in 1889.

Thus the available technical work is defined exactly by the *fundamental exergy equation* in integral form

$$-w_{t,12}^{pdv} = e_{x,1}^{pdv} - e_{x,2}^{pdv} + \int_{\tau_1}^{\tau_2} v \left[ \left( 1 - \frac{T_\infty}{T} \right) DT^N + T_\infty \sum_{i=1}^{i_{\max}} \mu_i (s_i - s) \right] d\tau. \quad (7.24)$$

### 7.3.4 The exergy definition appropriate for estimation of the technical work

In the sense of the *Gouy* exergy definition, Eq. (7.4) reads

$$-vdp = -de_x^{vdp} + v \left[ \left( 1 - \frac{T_\infty}{T} \right) DT^N + T_\infty \sum_{i=1}^{i_{\max}} \mu_i (s_i - s) \right] d\tau, \quad (7.25)$$

or in the integral form

$$-w_{t,12}^{vdp} = e_{x,1}^{vdp} - e_{x,2}^{vdp} + \int_1^2 v \left[ \left( 1 - \frac{T_\infty}{T} \right) DT^N + T_\infty \sum_{i=1}^{i_{\max}} \mu_i (s_i - s) \right] d\tau, \quad (7.26)$$

with

$$w_{t,12}^{vdp} = \int_1^2 vdp, \quad (7.27)$$

interpreted as a *specific technical work*. Now the meaning of the superscript *vdp* is clear.

## 7.4 Some interesting consequences of the fundamental exergy equation

Several limiting cases can be derived from Eq. (7.26). Some of them are given as examples.

A very interesting consequence of the two different forms (7.24) and (7.26) of the fundamental exergy equation in integral form is obtained for closed cycle processes in which the end and the initial fluid parameters are equal and therefore the specific exergy difference between the states 1 and 2 is zero. For this case we obtain

$$\int_1^2 (p - p_\infty) dv = \int_1^2 vdp. \quad (7.28)$$

This conclusion is manifested in the classical example for computing the technical work of a piston air compressor; see e.g. *Stephan and Mayinger (1998)* p. 107.

Let us discuss some important cases for the engineering applications:

**Case 1:** For a *closed* and *adiabatic* system without internal irreversible energy dissipation, without external thermal energy supply or removal, the available technical work is exactly equal to the differences of the initial and the final exergy of the fluid

$$-w_{i,12}^{pdv} = e_{x,1}^{pdv} - e_{x,2}^{pdv}. \quad (7.29a)$$

$$-w_{i,12}^{vdp} = e_{x,1}^{vdp} - e_{x,2}^{vdp}. \quad (7.29b)$$

The above expression is valid also for closed systems with irreversible energy dissipation and heat exchange with the environment if it happens at environmental temperature.

**Case 2:** If the system is closed, and there is no irreversible energy dissipation, but there is thermal energy exchange with the environment under temperatures different from the environmental temperatures, the available technical work is

$$-w_{i,12}^{pdv} = e_{x,1}^{pdv} - e_{x,2}^{pdv} + \int_{\tau_1}^{\tau_2} \left(1 - \frac{T_\infty}{T}\right) v \dot{q}'''' d\tau. \quad (7.30a)$$

$$-w_{i,12}^{pdv} = e_{x,1}^{pdv} - e_{x,2}^{pdv} + \int_{\tau_1}^{\tau_2} \left(1 - \frac{T_\infty}{T}\right) v \dot{q}'''' d\tau. \quad (7.30b)$$

**Case 3:** If the real system is closed, and there is irreversible energy dissipation, and there is thermal energy exchange with the environment under temperatures different from the environmental temperatures, the available volume change work and the available technical work are

$$-w_{i,12,irev}^{pdv} = e_{x,1}^{pdv} - e_{x,2}^{pdv} + \int_{\tau_1}^{\tau_2} \left(1 - \frac{T_\infty}{T}\right) v \dot{q}'''' d\tau + \int_{\tau_1}^{\tau_2} \left(1 - \frac{T_\infty}{T}\right) (P_k + \varepsilon) d\tau, \quad (7.31a)$$

$$-w_{i,12,irev}^{vdp} = e_{x,1}^{vdp} - e_{x,2}^{vdp} + \int_{\tau_1}^{\tau_2} \left(1 - \frac{T_\infty}{T}\right) v \dot{q}'''' d\tau + \int_{\tau_1}^{\tau_2} \left(1 - \frac{T_\infty}{T}\right) (P_k + \varepsilon) d\tau, \quad (7.31b)$$

which leads to

$$w_{i,12,rev}^{pdv} > w_{i,12}^{pdv}, \quad (7.32a)$$

$$w_{i,12,rev}^{vdp} > w_{i,12}^{vdp}. \quad (7.32b)$$

**Case 4:** If there is a cyclic process with the final state equal to the initial state and thermal energy supply only, only part of the thermal energy can be transferred in the technical work

$$-W_{t,12,rev}^{pdv} = -W_{t,12,rev}^{vdp} = \int_{\tau_1}^{\tau_2} \left(1 - \frac{T_\infty}{T}\right) v \dot{q}''' d\tau, \quad (7.33)$$

compare with *Carnot* 1824.

**Case 5:** If there is a process resulting in irreversible energy dissipation only, the available technical work is only a part of this dissipation

$$-W_{t,12,irev}^{pdv} = -W_{t,12,irev}^{vdp} = \int_{\tau_1}^{\tau_2} \left(1 - \frac{T_\infty}{T}\right) v P_k d\tau. \quad (7.34)$$

## 7.5 Judging the efficiency of a heat pump as an example of application of the exergy

Let us consider a heat pump, *Thomson* (1852), working with a cyclic steady state process for which at the same spatial point in the system at any time  $e_{x,1}^{vdp} = e_{x,2}^{vdp}$ . The circuit of the system is closed and therefore there are no mass sinks or sources. Equation (7.26) therefore reads

$$-W_{t,12}^{vdp} = \int_1^2 \left(1 - \frac{T_\infty}{T}\right) v \dot{q}''' d\tau + \int_1^2 \left(1 - \frac{T_\infty}{T_{dissipation}}\right) v \dot{q}'''_{dissipation} d\tau. \quad (7.35)$$

Thermal power  $\dot{q}'''_{in}$  is introduced in the system at constant temperature  $T_{in}$ . Thermal power  $\dot{q}'''_{out}$  is removed by the system at  $T_{out}$ . Somewhere under way from the lower to the higher level there are also unwished heat losses,  $\dot{q}'''_{dissipation}$  happening at some  $T_{dissipation}$ . In this case the exergy equation reads

$$-W_{t,12}^{vdp} = \left(1 - \frac{T_\infty}{T_{in}}\right) \int_1^2 v \dot{q}'''_{in} d\tau - \left(1 - \frac{T_\infty}{T_{out}}\right) \int_1^2 v \dot{q}'''_{out} d\tau - \left(1 - \frac{T_\infty}{T_{dissipation}}\right) \int_1^2 v \dot{q}'''_{dissipation} d\tau. \quad (7.36)$$

We immediately realize that the mechanic work, which has to be performed on the fluid by the compressor, depends on the temperature levels at which energy is supplied from the environment and provided to the house. In the idealized case of introducing heat at environmental temperature and negligible energy dissipation

the obtained heat from unit fluid mass  $\int_1^2 v \dot{q}'''_{out} d\tau$  is proportional to the compressor work  $W_{t,12}^{vdp}$  but not equal to it. The ratio

$$cop := \frac{\int_1^2 v \dot{q}_{out}''' d\tau}{w_{t,12}^{vdp}} = \frac{T_{out}}{T_{out} - T_{\infty}} > 1 \quad (7.37)$$

is called *coefficient of performance*. The smaller the temperature difference which has to be overcome the larger is the amount of the transferred thermal energy per unit technical work. Usually the heat is introduced into the heat pump condenser at a temperature higher than the environmental temperature for instance if heat from another process is available. In this case the exergy of the heat introduced into the fluid,

$$\left(1 - \frac{T_{\infty}}{T_{in}}\right) \int_1^2 v \dot{q}_{in}''' d\tau,$$

is a measure of the technical work which can be gained by an ideal *Carnot* cycle, so not utilizing this exergy is considered as a loss. This leads some authors to introduce the so called exergetic coefficient of performance

$$ex\_cop := \frac{\left(1 - \frac{T_{\infty}}{T_{out}}\right) \int_1^2 v \dot{q}_{out}''' d\tau}{w_{t,12}^{vdp} + \left(1 - \frac{T_{\infty}}{T_{in}}\right) \int_1^2 v \dot{q}_{in}''' d\tau} = 1 - \frac{\left(1 - \frac{T_{\infty}}{T_{dissipation}}\right) \int_1^2 v \dot{q}_{dissipation}''' d\tau}{w_{t,12}^{vdp} + \left(1 - \frac{T_{\infty}}{T_{in}}\right) \int_1^2 v \dot{q}_{in}''' d\tau} < 1. \quad (7.38)$$

This coefficient reflects the self-understanding that the heat  $\int_1^2 v \dot{q}_{out}''' d\tau$  has an

exergetic equivalence (value) of  $\left(1 - \frac{T_{\infty}}{T_{out}}\right) \int_1^2 v \dot{q}_{out}''' d\tau$ , which can be used as useful technical work. Thus, the exergetic coefficient of performance is a ratio of exergies. Only if there are no heat losses in the system is the exergetic coefficient of performance equal to one. If the introduction of the thermal energy happens under the environmental temperature we have

$$ex\_cop := \left(1 - \frac{T_{\infty}}{T_{out}}\right) cop. \quad (7.39)$$

## 7.6 Three-fluid multi-component systems

Now we consider a system consisting of three different fluids designated with subscript  $l$ . Each of the fluids takes only  $\alpha_l \geq 0$  part of the total system volume. Each of the fluids consists of several chemical components, designated with

subscript  $i$ . Evaporation and condensation are allowed as well as mechanical transfer between the fluids having the same aggregate state as a constitutive fluid component. Details of the definitions of such a system are given in Chapter 1 or in *Kolev* (1994). Again as in the energy conservation equation for each fluid written in a specific enthalpy form we have

$$\alpha_i \left( \rho_i \frac{dh_i}{d\tau} - \frac{dp_i}{d\tau} \right) = DT_i^N, \quad (7.40)$$

where

$$\begin{aligned} DT_i^N = & \alpha_i \rho_i \left( \delta_i P_{kl} + \varepsilon_i \right) \\ & + \dot{q}_i''' + \sum_{i=1}^{i_{\max}} \mu_{il, \text{in}} (h_{il, \text{in}} - h_{il}) + \sum_{i=1}^{i_{\max}} \mu_{il, \text{out}} (h_{il, \text{out}} - h_{il}) \\ & + \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}} \left[ \mu_{Mml} (h_{Ml}^\sigma - h_{Ml}) + \sum_{n=1}^{n_{\max}} \mu_{nml} (h_{nm} - h_{nl}) \right] \\ & + \frac{1}{2} \left[ \mu_{l, \text{in}} (\mathbf{V}_{l, \text{in}} - \mathbf{V}_l)^2 - \mu_{l, \text{out}} (\mathbf{V}_{l, \text{out}} - \mathbf{V}_l)^2 + \sum_{m=1}^3 \mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)^2 \right]. \end{aligned} \quad (7.41)$$

are different components of the energy input into the flow.  $P_{kl}$  and  $\varepsilon_l$  are irreversible components of the energy dissipation per unit volume of the fluid mixture for whatever reason. The thermal energy introduced into the fluid per unit time into unit volume of the fluid mixture is  $\dot{q}_i'''$ . The specific energy of the fluid is influenced by the difference between the in-flowing and the fluid enthalpy of the chemical components  $\sum_{i=1}^{i_{\max}} \mu_{il, \text{in}} (h_{il, \text{in}} - h_{il})$  as well as by the difference between the out-flowing and the system enthalpy of the chemical components  $\sum_{i=1}^{i_{\max}} \mu_{il, \text{out}} (h_{il, \text{out}} - h_{il})$ . Injecting and removing of mass of the system with a velocity different to the system velocity also gives rise to the internal energy change,  $\frac{1}{2} \left[ \mu_{l, \text{in}} (\mathbf{V}_{l, \text{in}} - \mathbf{V}_l)^2 - \mu_{l, \text{out}} (\mathbf{V}_{l, \text{out}} - \mathbf{V}_l)^2 + \sum_{m=1}^3 \mu_{ml} (\mathbf{V}_m - \mathbf{V}_l)^2 \right]$ .  $\mu_{l, \text{in}}$  and  $\mu_{l, \text{out}}$  are the inflow and outflow mass per unit volume of the fluid mixture and unit time.  $\mu_{ml}$  is the mass transferred from fluid  $m$  to fluid  $l$  per unit volume of the mixture and unit time. The complete equivalent to this equation in its entropy form is

$$\rho_i \alpha_i \frac{ds_i}{d\tau} = \frac{1}{T_i} DT_i^N + \sum_{m=1}^{l_{\max}} \sum_{\substack{i=1 \\ m \neq l}}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - \mu_l s_l. \quad (7.42)$$

For the derivation see Chapter 5 or *Kolev* (1995, 1997, 1998).

Now we multiply the entropy equation by the environment temperature, subtract it from the energy conservation equation, multiply the resulting equation by the time differential and divide by the fluid density and fluid volume fraction. The result for  $\alpha_l \geq 0$  is

$$de_l + p_l dv_l - T_\infty ds_l = \frac{1}{\alpha_l \rho_l} \left[ \left( 1 - \frac{T_\infty}{T_l} \right) DT_l^N + T_\infty \left[ \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}^w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - \mu_l s_l \right] \right] d\tau. \quad (7.43)$$

Defining the fluid exergy as

$$e_{x,l}^{pdv} := e_l + p_{l\infty} v_l - T_\infty s_l, \quad (7.44)$$

and the useful volume change work

$$-w_{t,l,12}^{pdv} = \int_1^2 (p_l - p_\infty) dv_l, \quad (7.45)$$

and after integration we obtain

$$-w_{t,l,12}^{pdv} = e_{x,l,1}^{pdv} - e_{x,l,2}^{pdv} + \int_{\tau_1}^{\tau_2} \frac{1}{\alpha_l \rho_l} \left[ \left( 1 - \frac{T_\infty}{T_l} \right) DT_l^N + T_\infty \left[ \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}^w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - \mu_l s_l \right] \right] d\tau, \quad (7.46)$$

the fundamental exergy equation for each fluid inside the fluid mixture. The essential difference to the single phase formulation is that mass and thermal energy transport inside the mixture is also taken into account. This equation can be reduced to all simple cases discussed before by setting the corresponding simplifying assumptions.

In practical analysis not the available technical work from the each fluid but that from the mixture is of interests. To compute it we sum all the exergy equations and introduce the instantaneous local mass fraction of each fluid

$$x_l := \frac{\alpha_l \rho_l}{\rho}, \quad (7.47)$$

where

$$\rho = \frac{1}{v} := \sum_{l=1}^{l_{\max}} \alpha_l \rho_l, \quad (7.48)$$



is the mixture density. The result obtained after the integration

$$-w_{t,12}^{pdv} = -\int_1^2 \sum_{l=1}^{l_{\max}} x_l de_{x,l}^{pdv} + \int_{\tau_1}^{\tau_2} v \sum_{l=1}^{l_{\max}} \left\{ \begin{array}{l} \left( 1 - \frac{T_\infty}{T_l} \right) DT_l^N \\ + T_\infty \left[ \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}^w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - \mu_l s_l \right] \end{array} \right\} d\tau, \quad (7.49)$$

is the specific volume change work of the mixture.

$$-w_{t,12}^{pdv} = \int_1^2 \sum_{l=1}^{l_{\max}} x_l (p_l - p_\infty) dv_l, \quad (7.50)$$

or for the case of assumed equal fluid pressures (single pressure model) we have

$$-w_{t,12}^{pdv} = \int_1^2 (p - p_\infty) \sum_{l=1}^{l_{\max}} x_l dv_l. \quad (7.51)$$

In a similar way the alternative exergy equation is

$$-w_{t,12}^{vdp} = -\int_1^2 \sum_{l=1}^{l_{\max}} x_l de_{x,l}^{vdp} + \int_{\tau_1}^{\tau_2} v \sum_{l=1}^{l_{\max}} \left\{ \begin{array}{l} \left( 1 - \frac{T_\infty}{T_l} \right) DT_l^N \\ + T_\infty \left[ \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}^w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} - \mu_l s_l \right] \end{array} \right\} d\tau, \quad (7.52)$$

where the specific technical work of the mixture is

$$-w_{t,12}^{vdp} = \int_1^2 \sum_{l=1}^{l_{\max}} x_l v_l dp_l. \quad (7.53)$$

For the case of assumed equal fluid pressures (single pressure model) we have

$$-w_{t,12}^{vdp} = \int_1^2 \left( \sum_{l=1}^{l_{\max}} x_l v_l \right) dp = \int_1^2 v dp. \quad (7.54)$$

Note that due to evaporation, condensation, or other mass transfer processes the mass concentration between the two selected states may change, which make the estimation of the integrals more complicated. For constant mass concentrations the exergy equations simplify.

## 7.7 Practical relevance

The general exergy equations for mixtures provide a simple way for computing the available volume change work or the technical work for a variety of processes. The multi-fluid exergy equations allow us to estimate the changes of the available technical work of the system due to processes like evaporation or condensation in non-equilibrium mixtures, which seems to be not known up to now. It is a generalization of the exergy principle to multi-phase multi-component systems. For the limiting case of a single fluid the equations reduce to the known relationships of classical thermodynamics. To the author's knowledge, Eqs. (7.46), (7.49) and (7.52) are derived for the first time in *Kolev* (2001).

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# 8 One-dimensional three-fluid flows

We call “one dimensional” the flow in pipes or in a pipe networks. We understand here a flow with cross section averaged flow characteristics with special wall boundary layer treatment, like pressure loss modeling, heat transfer modeling etc. The flow axis is of course arbitrarily oriented in space. Therefore this class of flows is one dimensional along a curvilinear pipe axis. A *network* consists of pipes and knots. An example is illustrated in Fig. 8.1.

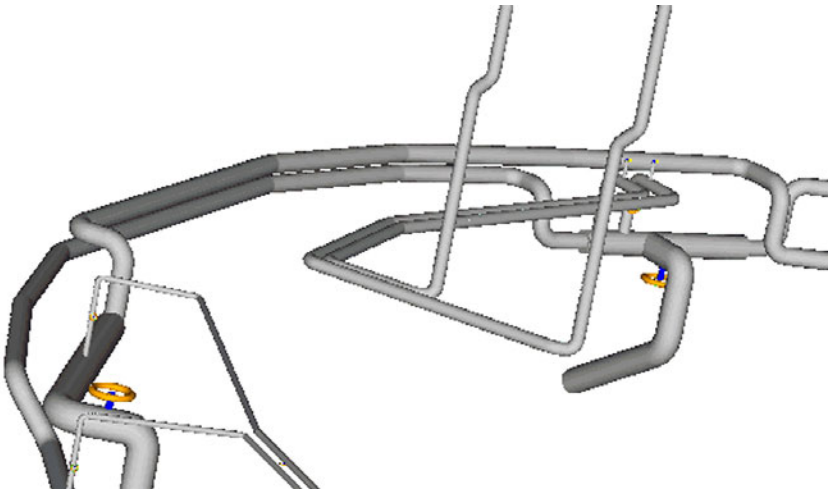


Fig. 8.1 Pipe network

## 8.1 Summary of the local volume- and time-averaged conservation equations

In order to facilitate the practical application of the conservation equations derived in the previous chapters we give here a summary of the equations simplified for the case of flow in the axial direction only. The mass conservation equation (1.62) derived in Chapter 1 simplifies as follows

$$\frac{\partial}{\partial \tau}(\alpha_i \rho_l \gamma_v) + \frac{\partial}{\partial z}(\alpha_i \rho_l w_l \gamma_z) = \gamma_v \sum_{m=1}^{l_{\max}, w} (\mu_{ml} - \mu_{lm}). \quad (8.1)$$

The mass conservation equation for each species inside the velocity field (1.90) derived in Chapter 1 simplifies as follows

$$\frac{\partial}{\partial \tau}(\alpha_i \rho_l C_{il} \gamma_v) + \frac{\partial}{\partial z} \left[ \alpha_i \rho_l \left( w_l C_{il} - D_{il}^* \frac{\partial C_{il}}{\partial z} \right) \gamma_z \right] = \gamma_v \sum_{m=1}^{l_{\max}, w} (\mu_{iml} - \mu_{ilm})$$

for  $\alpha_i \geq 0$ ,  $i = 1, i_{\max}$  (8.2)

or alternatively Eq. (1.96)

$$\alpha_i \rho_l \left( \gamma_v \frac{\partial C_{il}}{\partial \tau} + w_l \gamma_z \frac{\partial C_{il}}{\partial z} \right) - \frac{\partial}{\partial z} \left[ \alpha_i \rho_l \left( D_{il}^* \frac{\partial C_{il}}{\partial z} \right) \gamma_z \right] = \gamma_v (\mu_{il} - C_{il} \mu_l). \quad (8.3)$$

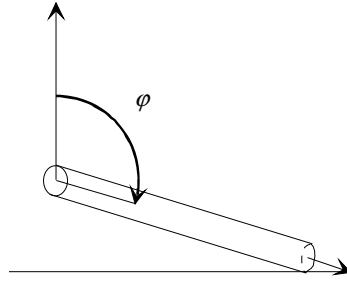
The particle number density equation (1.109) derived in Chapter 1 simplifies as follows

$$\frac{\partial}{\partial \tau} (n_l \gamma_v) + \frac{\partial}{\partial z} \left[ \left( w_l n_l - \frac{v_l'}{S c_l'} \frac{\partial n_l}{\partial z} \right) \gamma_z \right] = \gamma_v (\dot{n}_{l,kin} - \dot{n}_{l,coal} + \dot{n}_{l,sp}) \quad \text{for } \alpha_l \geq 0. \quad (8.4)$$

The momentum equation (2.233) derived in Chapter 2 simplifies as follows

$$\begin{aligned} & \alpha_l \rho_l \gamma_v \frac{\partial w_l}{\partial \tau} + \alpha_l^e \rho_l \gamma_z \frac{1}{2} \frac{\partial w_l^2}{\partial z} - \frac{\partial}{\partial z} \left( \alpha_l^e \rho_l v_l^* \gamma_z \frac{\partial w_l}{\partial z} \right) \\ & + \alpha_l^e \gamma_z \frac{\partial p_l}{\partial z} - \Delta p_m^{w\sigma^*} \frac{\partial \gamma_z}{\partial z} + f_{l\alpha} + \gamma_v \alpha_l \rho_l g \cos \varphi \\ & - \gamma_v \left\{ \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \bar{c}_{ml}^{vm} \left( \frac{\partial \Delta w_{ml}}{\partial \tau} + w_l \frac{\partial \Delta w_{ml}}{\partial z} \right) \right] - \rho_l c_{lw}^{vm} \frac{\partial w_l}{\partial \tau} \right\} \\ & = \gamma_v \left\{ \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \bar{c}_{ml}^d |\Delta w_{ml}| \cdot \Delta w_{ml} + \mu_{ml} (w_m - w_l) \right] + \bar{c}_{lw}^d |w_l| w_l \right. \\ & \quad \left. + \mu_{wl} (w_{wl} - w_l) - \mu_{lw} (w_{lw} - w_l) \right\}. \quad (8.5) \end{aligned}$$

Here  $\varphi$  is the polar angle between the flow direction and the upwards directed vertical as shown in Fig. 8.2



**Fig. 8.2** Definition of the polar angle

The entropy equation (5.125) derived in Chapter 5 simplifies as follows

$$\begin{aligned}
 & \rho_l \left( \alpha_l \gamma_v \frac{\partial s_l}{\partial \tau} + \alpha_l^e w_l \gamma_z \frac{\partial s_l}{\partial z} \right) - \frac{1}{T_l} \frac{\partial}{\partial z} \left[ \left( \alpha_l^e \lambda_l^* \frac{\partial T_l}{\partial z} \right) \gamma_z \right] \\
 & - \frac{\partial}{\partial z} \left\{ \alpha_l^e \rho_l \left[ \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \frac{\partial C_{il}}{\partial z} \right] \gamma_z \right\} \\
 & = \gamma_v \left[ \frac{1}{T_l} D T_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right] \text{ for } \alpha_l \geq 0, \quad (8.6)
 \end{aligned}$$

Alternatively, the temperature equation (5.176) derived in Chapter 5 can be used instead of the entropy equation for some applications

$$\begin{aligned}
 & \rho_l c_{pl} \left( \alpha_l \gamma_v \frac{\partial T_l}{\partial \tau} + \alpha_l^e w_l \gamma_z \frac{\partial T_l}{\partial z} \right) - \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } c's} \right] \left( \alpha_l \gamma_v \frac{\partial p}{\partial \tau} + \alpha_l^e w_l \gamma_z \frac{\partial p}{\partial z} \right) \\
 & - \frac{\partial}{\partial z} \left[ \left( \alpha_l^e \lambda_l^* \frac{\partial T_l}{\partial z} \right) \gamma_z \right] + T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \frac{\partial}{\partial z} \left[ \left( \alpha_l^e \rho_l D_{il}^* \frac{\partial C_{il}}{\partial z} \right) \gamma_z \right] \\
 & = \gamma_v \left[ D T_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right]. \quad (8.7)
 \end{aligned}$$

The volume conservation equation (5.188) derived in Chapter 5 simplifies for one-dimensional flow as follows

$$\frac{\gamma_v}{\rho a^2} \frac{\partial p}{\partial \tau} + \gamma_z \left( \sum_{l=1}^{l_{\max}} \frac{\alpha_l w_l}{\rho_l a_l^2} \right) \frac{\partial p}{\partial z} + \frac{\partial}{\partial z} \left( \sum_{l=1}^{l_{\max}} \alpha_l w_l \gamma_z \right) = \sum_{l=1}^{l_{\max}} D \alpha_l - \frac{\partial \gamma_v}{\partial \tau}, \quad (8.8)$$

where

$$D \alpha_l = \frac{1}{\rho_l} \left\{ \gamma_v \mu_l - \frac{1}{\rho_l} \left[ \frac{\partial \rho_l}{\partial s_l} \overline{D s_l^N} + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} \overline{D C_{il}^N} \right] \right\} \quad (8.9)$$

and

$$\overline{DC_{il}^N} = \frac{\partial}{\partial z} \left[ \alpha_l^e \rho_l \left( D_{il}^* \frac{\partial C_{il}}{\partial z} \right) \gamma_z \right] + \gamma_v (\mu_{il} - C_{il} \mu_l), \quad (8.10)$$

$$\begin{aligned} \overline{Ds_l^N} = & \frac{1}{T_l} \frac{\partial}{\partial z} \left[ \left( \alpha_l^e \lambda_l^* \frac{\partial T_l}{\partial z} \right) \gamma_z \right] + \frac{\partial}{\partial z} \left\{ \alpha_l^e \rho_l \left[ \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \frac{\partial C_{il}}{\partial z} \right] \gamma_z \right\} \\ & + \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right]. \end{aligned} \quad (8.11)$$

Remember that  $a$  is the *sonic velocity* in a homogeneous multi-phase mixture defined as

$$\frac{1}{\rho a^2} = \sum_{i=1}^3 \frac{\alpha_i}{\rho_i a_i^2} = \frac{1}{p} \sum_{i=1}^3 \frac{\alpha_i}{\kappa_i} = \frac{1}{\kappa p}, \quad (8.12)$$

and

$$\rho = \sum_{i=1}^3 \alpha_i \rho_i \quad (8.13)$$

is the mixture density.

## 8.2 Treatment of the field pressure gradient forces

### 8.2.1 Dispersed flows

The pressure terms in the momentum equation are now written for a continuous and a disperse phase

$$f_{d\alpha} = (p_d - p_c - \delta_d \sigma_{dc} \kappa_d - \Delta p_c^{d\sigma*}) \frac{\partial}{\partial z} (\alpha_d^e \gamma_z), \quad (8.14)$$

and

$$f_{c\alpha} = \Delta p_c^{d\sigma*} \frac{\partial}{\partial z} (\alpha_d^e \gamma_z). \quad (8.15)$$

where

$$\Delta p_c^{d\sigma*} = -0.37 c_{cd}^d \rho_c (w_c - w_d)^2 \quad (8.16)$$

is Eq. (2.178) in Chapter 2. An approximation is frequently used in the literature in the following form

$$\dots + \alpha_c \gamma_z \frac{\partial p}{\partial z} - \Delta p_c^{w\sigma} \frac{\partial \gamma_z}{\partial z} + \gamma_v \alpha_c \rho_c g \cos \varphi \dots, \quad (8.17)$$

$$\dots + \alpha_d \gamma_z \frac{\partial p}{\partial z} - 0.37 c_{cd}^d \rho_c |w_c - w_d| (w_c - w_d) \frac{\partial}{\partial z} (\alpha_d \gamma_z) + \gamma_v \alpha_d \rho_d g \cos \varphi \dots \quad (8.18)$$

### 8.2.2 Stratified flow

In stratified flow the gravitation is the dominant force. It is manifested in different pressure distribution over the cross section which gives rise to differences in the averaged system pressure and the averaged field pressures. In Chapter 2 the Eqs. (2.103) and (2.104)

$$\dots + \gamma_z \left[ \alpha_1 \frac{\partial p}{\partial z} - \alpha_1 \alpha_2 (\rho_1 - \rho_2) g H \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_1^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (8.19)$$

$$\dots + \gamma_z \left[ \alpha_2 \frac{\partial p}{\partial z} + \alpha_1 \alpha_2 (\rho_1 - \rho_2) g H \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_2^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (8.20)$$

are derived for rectangular channels and Eqs. (2.148) and (2.149)

$$\dots + \gamma_z \left[ \alpha_1 \frac{\partial p}{\partial z} + \alpha_1 \alpha_2 (\rho_2 - \rho_1) g \frac{\pi D_h}{4 \sin \theta} \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_1^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (8.21)$$

$$\dots + \gamma_z \left[ \alpha_2 \frac{\partial p}{\partial z} - \alpha_1 \alpha_2 (\rho_2 - \rho_1) g \frac{\pi D_h}{4 \sin \theta} \frac{\partial \alpha_1}{\partial z} \right] - \Delta p_2^{w\sigma} \frac{\partial \gamma_z}{\partial z} \dots \quad (8.22)$$

for pipes. Note that the term  $\dots \gamma_v \alpha_i \rho_i g \cos \varphi \dots$  does not appear any more in these equations.

### 8.3 Pipe deformation due to temporal pressure change in the flow

Pipes with elastic walls change the propagation velocity of pressure pulses because of the energy dissipation for mechanical deformation. The effect can be taken into account into the pressure equation (8.8) as follows

$$\begin{aligned} \gamma_v \left( \frac{1}{\rho a^2} + \frac{1}{\gamma_v} \frac{d\gamma_v}{dp} \right) \frac{\partial p}{\partial \tau} + \gamma_z \sum_{l=1}^3 \alpha_l w_l \left( \frac{1}{\rho_l a_l^2} + \frac{1}{\gamma_z} \frac{d\gamma_z}{dp} \right) \frac{\partial p}{\partial z} \\ + \gamma_z \frac{\partial}{\partial z} \left( \sum_{l=1}^3 \alpha_l w_l \right) = \sum_{l=1}^{l_{\max}} D \alpha_l . \end{aligned} \quad (8.23)$$

With the terms  $d\gamma_v/dp$  and  $d\gamma_z/dp$  we take into account the elasticity of the pipe. The method for computation of the two terms is given below.

For closed pipe with uniform pressure the deformation caused by the pressure can be computed using text book solution

$$dz / \Delta z = (\sigma_z - \mu^* \sigma_\theta - \mu^* \sigma_r) / E , \quad (8.24)$$

$$dr_i / r_i = (\sigma_\theta - \mu^* \sigma_z - \mu^* \sigma_r) / E , \quad (8.25)$$

$$d\delta / \delta = (\sigma_r - \mu^* \sigma_\theta - \mu^* \sigma_z) / E , \quad (8.26)$$

where the subscripts  $i$  and  $a$  stand for inner and outer wall radius, respectively. Here  $E$  is the *elasticity modulus* e.g. for steel  $E=1 \times 10^{11} \text{ Nm}^2$ ,  $\mu^*$  is the contraction number, e.g. 0.3. The axial, the radial and the angular stresses are functions of the pressure change of the fluid

$$\sigma_r = dp / \left[ (r_a / r_i)^2 - 1 \right] , \quad (8.27)$$

$$\sigma_r = -dp , \quad (8.28)$$

$$\sigma_\theta = dp (r_a^2 + r_i^2) / (r_a^2 - r_i^2) . \quad (8.29)$$

Here  $r_i$  and  $r_a$  are the inner and the outer radius.

$$\delta = r_a - r_i \quad (8.30)$$

is the wall thickness. The change of the flow volume due to elastic pipe deformation is therefore

$$\frac{1}{V_{\text{pipe}}} \frac{dV_{\text{pipe}}}{dp} = \frac{1}{\gamma_v} \frac{d\gamma_v}{dp} = \frac{1}{\pi r_i^2 \Delta z} \frac{d}{dp} (\pi r_i^2 \Delta z) = \frac{2}{r_i} \frac{dr_i}{dp} + \frac{1}{\Delta z} \frac{d\Delta z}{dp} = \frac{\psi}{E} , \quad (8.31)$$

where

$$\psi = \left[ 2(1 + \mu^*) r_a^2 + (3 - 6\mu^*) r_i^2 \right] / (r_a^2 - r_i^2) \quad (8.32)$$

is frequently called *geometry factor*. For long pipes the local distribution of pressure makes it difficult to compute the axial component. Often, only the radial component is used



$$\psi = 2 \left\{ \left[ r_a^2 + (1 - \mu^*) r_i^2 \right] / \left( r_a^2 - r_i^2 \right) + \mu^* \right\}. \quad (8.33)$$

From the geometry factor we see that the smaller the pipe wall thickness the stronger the effect of the pipe elasticity on the pressure wave propagation.

The expression containing the pipe deformation can be rewritten as follows

$$\frac{1}{\rho a^2} + \frac{1}{\gamma_v} \frac{d\gamma_v}{dp} = \frac{1}{\rho a^2} + \frac{\psi}{E} = \frac{1}{\rho a_{\text{eff}}^2},$$

where the effective velocity of sound is obviously a function of the elasticity of the pipe

$$\frac{1}{a_{\text{eff}}^2} = \frac{1}{a^2} + \frac{1}{a_{\text{elastic pipe}}^2}.$$

Here

$$a_{\text{elastic pipe}} = \sqrt{E/(\rho\psi)}.$$

Note that if  $a_{\text{elastic pipe}} \gg a$ ,  $a_{\text{eff}} \approx a$  and if  $a_{\text{elastic pipe}} \ll a$ ,  $a_{\text{eff}} \approx a_{\text{elastic pipe}}$ . For elastic ducts like human vanes, the influence of the wall elasticity is substantial. Note, that this expression was obtained for the first time by *Joukowsky* in 1898. *Joukowsky* extend the *Rieman*'s analysis to water flow in *elastic* pipe and explained the physical basics behind the reduction of the velocity of sound in elastic pipes. He also proved additional prove for this expression based on mechanical energy balance. *Joukowsky* computed numerically the influence of the axial extension and found that it is negligible for cases of metallic pipes used in urban environment.

More complicated is the case if plastic deformations are possible. A simple method proposed by *Pollak* (1976) for taking this effect in pipes is

$$\rho_{\text{wall}} d^2 r_i / d\tau^2 = (p / \delta_{\text{wall}} - \sigma / r_i) r_i / r_a,$$

where the stress-extension function  $\sigma = \sigma(\varepsilon)$  is material properties of the wall.

## 8.4 Some simple cases

*Concentration and specific entropy propagation in incompressible flow with prescribed velocity:* In many practical applications where the convection is the governing phenomenon the diffusion terms in the concentration and the entropy equations can be neglected. The resulting equations

$$\alpha_l \rho_l \left( \gamma_v \frac{\partial C_{il}}{\partial \tau} + w_l \gamma_z \frac{\partial C_{il}}{\partial z} \right) = \gamma_v (\mu_{il} - \mu_l C_{il}), \quad (8.34)$$

$$\alpha_l \rho_l \left( \gamma_v \frac{\partial s_l}{\partial \tau} + w_l \gamma_z \frac{\partial s_l}{\partial z} \right) = \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right], \quad (8.35)$$

can be easily analytically integrated. The eigenvalues of the characteristic matrix of the above system are  $\lambda_l = w_l \gamma_z / \gamma_v$ , where  $l=1,2,3$  for each velocity field and equation, respectively. This means that changes of the concentrations and of the specific entropies travel with the corresponding field velocity multiplied by  $\gamma_z / \gamma_v$ . For pipes with constant cross section the multiplier is unity. If one changes the coordinate system for each equation with an rectangular system having one of the axes tangential to the curve defined with inclination  $dz / d\tau = \lambda_l$  in the time - space plane, called characteristic curves, the equations take a very simple form

$$\frac{dC_{il}}{d\tau} = (\mu_{il} - \mu_l C_{il}) / (\alpha_l \rho_l), \quad (8.36)$$

$$\frac{ds_l}{d\tau} = \left( \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} s_{il} - \mu_l s_l \right) / (\alpha_l \rho_l), \quad (8.37)$$

called characteristic form (*Rieman*). In case of no sources, that is the right-hand side of the equations is equal to zero,

$$\frac{dC_{il}}{d\tau} = 0, \quad (8.38)$$

$$\frac{ds_l}{d\tau} = 0, \quad (8.39)$$

the concentrations and the specific entropies are constant

$$C_{il} = \text{const}, \quad (8.40)$$

$$s_l = \text{const}, \quad (8.41)$$

along the characteristic line. The asymptotic solutions are

$$C_{il\infty} = \mu_{il} / \mu_l, \quad (8.42)$$

$$s_{l\infty} = \frac{1}{\mu_l} \left( \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} s_{il} \right). \quad (8.43)$$

Rewriting the system in the following form

$$\frac{dC_{il}}{d\tau} = (C_{il\infty} - C_{il})\mu_l / (\alpha_l \rho_l), \quad (8.44)$$

$$\frac{ds_l}{d\tau} = (s_{l\infty} - s_l)\mu_l / (\alpha_l \rho_l), \quad (8.45)$$

and assuming that during a time step  $\Delta\tau$  the expressions  $C_{il\infty} = const$ ,  $s_{l\infty} = const$ , the equations can be analytically integrated with the result

$$C_{il,\tau+\Delta\tau} = C_{il\infty} - (C_{il\infty} - C_{il}) / \exp \left[ \int_{\tau}^{\tau+\Delta\tau} \mu_l / (\alpha_l \rho_l) d\tau \right], \quad (8.46)$$

$$s_{l,\tau+\Delta\tau} = s_{l\infty} - (s_{l\infty} - s_l) / \exp \left[ \int_{\tau}^{\tau+\Delta\tau} \mu_l / (\alpha_l \rho_l) d\tau \right]. \quad (8.47)$$

Note once more that the so obtained analytical solutions are valid along the characteristic curve.

*Steady state flow:* In steady state the time derivatives are equal to zero. I split the resulting system in two groups of equations. For the first group

$$\frac{dC_{il}}{dz} = \gamma_v (C_{il\infty} - C_{il})\mu_l / (\alpha_l \rho_l w_l \gamma_z), \quad (8.48)$$

$$\frac{ds_l}{dz} = \gamma_v (s_{l\infty} - s_l)\mu_l / (\alpha_l \rho_l w_l \gamma_z), \quad (8.49)$$

assuming that along  $\Delta z$  the expressions  $C_{il\infty} = const$ ,  $s_{l\infty} = const$  hold, the equations can be analytically integrated with the result

$$C_{il,z+\Delta z} = C_{il\infty} - (C_{il\infty} - C_{il}) / \exp \left[ \int_z^{z+\Delta z} \gamma_v \mu_l / (\alpha_l \rho_l w_l \gamma_z) dz \right], \quad (8.50)$$

$$s_{l,z+\Delta z} = s_{l\infty} - (s_{l\infty} - s_l) / \exp \left[ \int_z^{z+\Delta z} \gamma_v \mu_l / (\alpha_l \rho_l w_l \gamma_z) dz \right]. \quad (8.51)$$

The second group of equations is discussed below. The mass conservation equation

$$\frac{d}{dz} (\alpha_l \rho_l w_l \gamma_z) = \gamma_v \mu_l, \quad (8.52)$$

is expanded and the densities are replaced by their equals from the series expansion of the state equations. The result is

$$\frac{d\alpha_l}{dz} + \frac{\alpha_l}{w_l} \frac{dw_l}{dz} + \frac{\alpha_l}{\rho_l a_l^2} \frac{dp}{dz} + \alpha_l \frac{1}{\gamma_z} \frac{d\gamma_z}{dz} = D\alpha_l. \quad (8.53)$$

Using the concentration and the entropy equations we obtain

$$\begin{aligned}
 D\alpha_l &= \frac{1}{\rho_l w_l} \frac{\gamma_v}{\gamma_z} \mu_l - \frac{\alpha_l}{\rho_l} \left[ \frac{\partial \rho_l}{\partial s_l} \frac{ds_l}{dz} + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} \frac{dC_{il}}{dz} \right] \\
 &= \frac{1}{\rho_l w_l} \frac{\gamma_v}{\gamma_z} \left\{ \mu_l - \frac{1}{\rho_l} \left[ \frac{\partial \rho_l}{\partial s_l} (s_{l\infty} - s_l) \mu_l + \mu_l \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} (C_{i\infty} - C_{il}) \right] \right\}. \quad (8.54)
 \end{aligned}$$

Having in mind that

$$\sum_{l=1}^3 \frac{d\alpha_l}{dz} = 0, \quad (8.55)$$

the sum of the volume fraction equations become the so called pressure equation

$$\sum_{l=1}^3 \frac{\alpha_l}{w_l} \frac{dw_l}{dz} + \frac{1}{\rho a^2} \frac{dp}{dz} = \sum_{l=1}^3 D\alpha_l - \frac{1}{\gamma_z} \frac{d\gamma_z}{dz}. \quad (8.56)$$

Next we look for expression for the velocity gradients obtained from the momentum equations with the purpose to substitute them into the pressure equation and solve the so obtained equation with respect to the pressure gradient. The momentum equation is

$$\left( \alpha_l^e \rho_l + \frac{\gamma_v}{\gamma_z} \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm} \right) \frac{dw_l}{dz} - \frac{\gamma_v}{\gamma_z} \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm} \frac{dw_m}{dz} + \frac{\alpha_l}{w_l} \frac{dp}{dz} = \frac{Z_l}{w_l} \quad (8.57)$$

where

$$Z_l = \frac{\gamma_v}{\gamma_z} \left\{ \begin{aligned} & -\alpha_l \rho_l g + \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \bar{c}_{ml}^d |\Delta w_{ml}| \Delta w_{ml} + \mu_{ml} (w_m - w_l) \right] - \bar{c}_{lw}^d |w_l| w_l \\ & + \mu_{wl} (w_{wl} - w_l) - \mu_{lw} (w_{lw} - w_l) \end{aligned} \right\}, \quad (8.58)$$

contains no differential terms. The system of algebraic equations containing the velocity gradients is therefore

$$\mathbf{A} \left[ \frac{dw}{dz} \right] = \mathbf{B} - \mathbf{C} \frac{dp}{dz}. \quad (8.59)$$

The components of the algebraic vectors  $\mathbf{C}$  and  $\mathbf{B}$  are  $c_l = \alpha_l / w_l$  and  $b_l = Z_l / w_l$ , respectively. The diagonal and the non-diagonal elements of the matrix  $\mathbf{A}$  are

$$a_{ll} = \alpha_l^e \rho_l + \frac{\gamma_v}{\gamma_z} \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm}, \quad (8.60)$$

$$a_{lm} = a_{ml} = -\frac{\gamma_v}{\gamma_z} \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm}, \quad (8.61)$$

respectively. Having in mind that  $\bar{c}_{ml}^{vm} = \bar{c}_{lm}^{vm}$ , we realize that the coefficient matrix  $\mathbf{A}$  of the velocity derivatives is symmetric. If the virtual mass coefficients are set to zero there is only one velocity gradient in each momentum equation

$$\frac{1}{2} \frac{dw_l^2}{dz} + \frac{1}{\rho_l} \frac{dp}{dz} = \frac{Z_l}{\alpha_l \rho_l}. \quad (8.62)$$

The algebraic system (8.59) can be solved with respect to the velocity gradients to provide

$$\det |\mathbf{A}| = a_{11}a_{22}a_{33} + 2a_{12}a_{23}a_{13} - a_{13}^2a_{22} - a_{23}^2a_{11} - a_{12}^2a_{33} \neq 0. \quad (8.63)$$

Note that if one field does not exist the coupling coefficients with the other fields are by definition zeros. In such case the rank of the matrix is reduced by one. Solving with respect to the velocity gradients gives

$$\frac{dw_l}{dz} = w_l^* - R_l \frac{dp}{dz} \quad (8.64)$$

where

$$w_l^* = \left[ \sum_{m=1}^3 (Z_l / w_l) \bar{a}_{lm} \right] / \det |\mathbf{A}|, \quad (8.65)$$

$$R_l = \left[ \sum_{m=1}^3 (\alpha_l / w_l) \bar{a}_{lm} \right] / \det |\mathbf{A}|, \quad (8.66)$$

$$\bar{a}_{11} = a_{22}a_{33} - a_{23}^2, \quad \bar{a}_{22} = a_{11}a_{33} - a_{13}^2, \quad \bar{a}_{33} = a_{11}a_{22} - a_{12}^2, \quad (8.67-8.69)$$

$$\bar{a}_{12} = \bar{a}_{21} = a_{32}a_{13} - a_{12}a_{33}, \quad \bar{a}_{13} = \bar{a}_{31} = a_{12}a_{23} - a_{22}a_{13}, \quad \bar{a}_{23} = \bar{a}_{32} = a_{21}a_{13} - a_{23}a_{11}. \quad (8.70-8.72)$$

Replacing the so obtained velocity gradients into the pressure equation we obtain

$$\frac{dp}{dz} = - \left\{ \sum_{l=1}^3 D\alpha_l - \frac{1}{\gamma_z} \frac{d\gamma_z}{dz} - \sum_{l=1}^3 \frac{\alpha_l}{w_l} w_l^* \right\} / (1 - Ma^2) \sum_{l=1}^3 \frac{\alpha_l}{w_l} R_l. \quad (8.73)$$

Here

$$Ma^2 = \frac{1}{\rho a^2} / \sum_{l=1}^3 \frac{\alpha_l}{w_l} R_l \quad (8.74)$$

is the definition of a dimensionless number corresponding to the *Mach* number for single-phase flows. Obviously if this number tends to unity the pressure gradient tends to minus infinity. If this happens at some position in the channel no pressure disturbances coming from flow downwards can influence the flow. This state is called *critical flow*. It plays an important role in the technology. The critical flow is expected to happen in the smallest cross section of the pipe where the velocities have local maximums and the pressure local minimum, or at the end of pipes with constant cross section.

Let us summarize the resulting system in a form very convenient for numerical integration.

$$\frac{dp}{dz} = - \left\{ \sum_{l=1}^3 D\alpha_l - \frac{1}{\gamma_z} \frac{d\gamma_z}{dz} - \sum_{l=1}^3 \frac{\alpha_l}{w_l} w_l^* \right\} / (1 - Ma^2) \sum_{l=1}^3 \frac{\alpha_l}{w_l} R_l, \quad (8.75)$$

$$\frac{dw_l}{dz} = w_l^* - R_l \frac{dp}{dz}, \quad (8.76)$$

or alternatively

$$\frac{1}{2} \frac{dw_l^2}{dz} = \frac{1}{\det|\mathbf{A}|} \left[ \sum_{m=1}^3 \left( Z_l - \alpha_l \frac{dp}{dz} \right) \bar{a}_{lm} \right] = \frac{1}{\det|\mathbf{A}|} \left[ \sum_{m=1}^3 Z_l \bar{a}_{lm} - \left( \sum_{m=1}^3 \alpha_l \bar{a}_{lm} \right) \frac{dp}{dz} \right] \quad (8.77)$$

and

$$\frac{d\alpha_l}{dz} = D\alpha_l - \frac{\alpha_l}{w_l} \frac{dw_l}{dz} - \frac{\alpha_l}{\rho_l a_l^2} \frac{dp}{dz} - \alpha_l \frac{1}{\gamma_z} \frac{d\gamma_z}{dz}, \quad (8.78)$$

$$\frac{dC_{il}}{dz} = \gamma_v (C_{i\infty} - C_{il}) \mu_l / (\alpha_l \rho_l w_l \gamma_z), \quad (8.79)$$

$$\frac{ds_l}{dz} = \gamma_v (s_{l\infty} - s_l) \mu_l / (\alpha_l \rho_l w_l \gamma_z), \quad (8.80)$$

$$\frac{dn_l}{dz} = \frac{\gamma_v}{w_l \gamma_z} (\dot{n}_{l,kin} - \dot{n}_{l,coal} + \dot{n}_{l,sp}) - \frac{n_l}{w_l} \frac{dw_l}{dz} - \frac{n_l}{\gamma_z} \frac{d\gamma_z}{dz}. \quad (8.81)$$

*Nozzles frozen flow:* For steady state with neglected virtual mass forces the momentum equation takes the simple form (8.64). Replacing the density in the gas momentum equation with the expression for the isentropic state of change we obtain

$$\frac{1}{2} \frac{dw_l^2}{dz} + \frac{p_0}{\rho_{l0}} \frac{\kappa_l}{\kappa_l - 1} \frac{d}{dz} \varepsilon^{\frac{\kappa_l - 1}{\kappa_l}} = \frac{Z_l}{\alpha_l \rho_l}, \quad (8.82)$$

where

$$\varepsilon = \frac{P}{P_0} \quad (8.83)$$

is the pressure ratio. Here we designate with 0 some reference state, e.g. the state at the pipe inlet. For the second and the third velocity field Eq. (8.62) remains unchanged. Integrating between two points we obtain

$$\frac{1}{2} (w_1^2 - w_{10}^2) = \int_{z_0}^z \frac{Z_1}{\alpha_1 \rho_1} dz - \frac{P_0}{\rho_{10}} \frac{\kappa_1}{\kappa_1 - 1} \left( \varepsilon^{\frac{\kappa_1 - 1}{\kappa_1}} - 1 \right), \quad (8.84)$$

$$\frac{1}{2} (w_l^2 - w_{l0}^2) = \int_{z_0}^z \frac{Z_l}{\alpha_l \rho_l} dz - \frac{P_0}{\rho_{l0}} (\varepsilon - 1), \quad \text{for } l = 2, 3. \quad (8.85)$$

For short pipes, nozzles and orifices with negligible interfacial drag we obtain

$$w_1^2 = w_{10}^2 + 2 \frac{P_0}{\rho_{10}} \frac{\kappa_1}{\kappa_1 - 1} \left( 1 - \varepsilon^{\frac{\kappa_1 - 1}{\kappa_1}} \right), \quad (8.86)$$

$$w_l^2 = w_{l0}^2 + 2 \frac{P_0}{\rho_{l0}} (1 - \varepsilon), \quad \text{for } l = 2, 3. \quad (8.87)$$

The last two equations were obtained in 1949 by *Tangren* et al. For pure gas flow the equations results to the result obtained by *de Saint Venant* and *Wantzel* in 1838. In the literature concerning critical two-phase flow the velocity ratio  $S_1 = w_1 / w_2$  is frequently called *slip ratio* or *slip*. For two velocity fields the slip ratio at the outlet of the nozzle

$$S_1 = \left\{ \frac{w_{10}^2 + 2 \frac{P_0}{\rho_{10}} \frac{\kappa_1}{\kappa_1 - 1} \left( 1 - \varepsilon^{\frac{\kappa_1 - 1}{\kappa_1}} \right)}{w_{20}^2 + 2 \frac{P_0}{\rho_{20}} (1 - \varepsilon)} \right\}^{1/2} \quad (8.88)$$

depends also on the inlet history. For discharge from a vessel with stagnant mixture or flushing inlet flow we have

$$S_1 = \left\{ \frac{\rho_{20}}{\rho_{10}} \frac{\kappa_1}{\kappa_1 - 1} \left( 1 - \varepsilon^{\frac{\kappa_1 - 1}{\kappa_1}} \right) / (1 - \varepsilon) \right\}^{1/2}. \quad (8.89)$$

In the reality the interfacial drag will reduce considerable this value especially in the limiting case of diminishing velocity field.

The sum of Eqs. (8.86) and (8.87) multiplied by the corresponding densities and volume fractions at the nozzle outlet, gives the interesting expression

$$\frac{1}{2} \sum_{l=1}^3 \alpha_l \rho_l w_l^2 = \frac{1}{2} \sum_{l=1}^3 \alpha_l \rho_l w_{l0}^2 + p_0 \left[ \alpha_1 \varepsilon^{1/\kappa_1} \frac{\kappa_1}{\kappa_1 - 1} \left( 1 - \varepsilon^{\frac{\kappa_1 - 1}{\kappa_1}} \right) + (1 - \alpha_1)(1 - \varepsilon) \right]. \quad (8.90)$$

Equation (8.90) allows us to compute the mass flow rate by known slip and pressure ratios. Note that in accordance with the slip definition from the next section Eq. (8.90) can be rewritten as follows

$$\frac{1}{2} v_s f_0 G^2 = \frac{1}{2} v_{s0} f_{00} G_0^2 + p_0 \left[ \alpha_1 \varepsilon^{1/\kappa_1} \frac{\kappa_1}{\kappa_1 - 1} \left( 1 - \varepsilon^{\frac{\kappa_1 - 1}{\kappa_1}} \right) + (1 - \alpha_1)(1 - \varepsilon) \right].$$

*Nozzle flow with instantaneous heat exchange without mass exchange:* For this case

$$\frac{1}{2} \sum_{l=1}^3 \alpha_l \rho_l w_l^2 = \frac{1}{2} \sum_{l=1}^3 \alpha_l \rho_l w_{l0}^2 + p_0 \left[ \alpha_1 \varepsilon^{1/n} \frac{n}{n - 1} \left( 1 - \varepsilon^{\frac{n - 1}{n}} \right) + (1 - \alpha_1)(1 - \varepsilon) \right] \quad (8.91)$$

we have the polytropic state of change with the polytropic exponent

$$n = \frac{c_p}{c_p - R}, \quad (8.92)$$

and mixture specific heat defined as

$$c_p = \sum_{l=1}^3 \frac{\alpha_l \rho_l}{\rho} c_{pl}, \quad (8.93)$$

and the effective gas constant

$$R = \sum_{l=1}^3 \frac{\alpha_l \rho_l}{\rho} \bar{R}_l. \quad (8.94)$$

Note that for the second and the third field the pseudo gas constant is negligibly small.

## 8.5 Slip model – transient flow

Historically for the mathematical description of the mechanical interfacial interaction the so called slip models are used among others. In this technique field velocity ratios are modeled with empirical correlation replacing the complete description of the mechanical interaction by means of separated momentum



equations. Only the mixture momentum equation is used instead. Thus only the mechanical behavior of the mixture as a whole is considered properly. The adjustment of the slip ratio as a function of the local flow parameter is assumed to be instantaneous that is – inertialess. The disadvantages of the slip model are obvious. Nevertheless, the slip model provides for some application a reasonable simplicity and therefore will be described here. We described one of the families of the slip models in which the *slip velocity ratio* is defined as the field velocity divided by the center of mass mixture velocity.

$$S_l = w_l / w, \quad (8.95)$$

where the center of mass velocity

$$w = G / \rho, \quad (8.96)$$

is defined as the mixture mass flow rate

$$G = \sum_{l=1}^3 \alpha_l \rho_l w_l, \quad (8.97)$$

divided by the mixture density

$$\rho = \sum_{l=1}^3 \alpha_l \rho_l. \quad (8.98)$$

Instead of using local volume fractions the local mass flow concentrations defined as follows

$$X_l = \alpha_l \rho_l w_l / G, \quad (8.99)$$

will be used as very convenient. Obviously

$$\sum_{l=1}^3 X_l = 1, \quad (8.100)$$

per definition. Some useful consequences of the introduction of the slip ratios and the mass flow concentrations are

$$f_0 = \sum_{l=1}^3 \frac{X_l}{S_l}, \quad (8.101)$$

$$f_1 = \rho v_s = \sum_{l=1}^3 X_l S_l, \quad (8.102)$$

$$\rho w \equiv G, \quad (8.103)$$

$$v_s = \sum_{l=1}^3 X_l S_l v_l, \quad (8.104)$$

$$v_l = \left( \sum_{l=1}^3 \alpha_l \rho_l w_l^2 \right) / G^2 = v_s f_0, \quad (8.105)$$

$$\alpha_l = \frac{X_l v_l S_l}{v_s}, \quad (8.106)$$

$$w_l = v_s G / S_l. \quad (8.107)$$

With these definitions the conservation equations for field mass, mixture mass and mixture momentum

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l \gamma_v) + \frac{\partial}{\partial z} (\alpha_l \rho_l w_l \gamma_z) = \gamma_v \mu_l, \quad (8.108)$$

$$\frac{\partial}{\partial \tau} \left( \gamma_v \sum_{l=1}^3 \alpha_l \rho_l \right) + \frac{\partial}{\partial z} \left( \gamma_z \sum_{l=1}^3 \alpha_l \rho_l w_l \right) = \gamma_v \mu_l, \quad (8.109)$$

$$\begin{aligned} & \frac{\partial}{\partial \tau} \left( \gamma_v \sum_{l=1}^3 \alpha_l \rho_l w_l \right) + \frac{\partial}{\partial z} \left( \gamma_z \sum_{l=1}^3 \alpha_l \rho_l w_l^2 \right) + \gamma_z \frac{\partial p}{\partial z} + \gamma_v g \cos \varphi \sum_{l=1}^3 \alpha_l \rho_l \\ & = \gamma_v \sum_{l=1}^3 \left[ \bar{c}_{lw}^d |w_l| w_l + \mu_{wl} (w_{wl} - w_l) - \mu_{lw} (w_{lw} - w_l) \right], \end{aligned} \quad (8.110)$$

can be rewritten as follows

$$\frac{\partial}{\partial \tau} \left( \frac{X_l S_l}{v_s} \gamma_v \right) + \frac{\partial}{\partial z} (X_l G \gamma_z) = \gamma_v \mu_l, \quad (8.111)$$

$$\frac{\partial}{\partial \tau} \left( \frac{f_1}{v_s} \gamma_v \right) + \frac{\partial}{\partial z} (G \gamma_z) = \gamma_v \mu, \quad (8.112)$$

$$\frac{\partial}{\partial \tau} (G \gamma_v) + \frac{\partial}{\partial z} (v_l G^2 \gamma_z) + \gamma_z \frac{\partial p}{\partial z} = \gamma_v (f_{fr} + f_\mu - \rho g \cos \varphi). \quad (8.113)$$

Here  $f_{fr}$  is the frictional pressure drop gradient of the mixture depending on the local parameter.  $f_\mu$  is the force due to injection of suction. We assign the subscript  $m$  to those particular fields whose mass flow concentration is dependent and defined by the other ones through Eq.(8.100). The dependent variable vector is

$$\mathbf{U}^T = [G, p, X_{l,l \neq m}, s_l, C_{il}]. \quad (8.114)$$

The conservative equations can be rewritten in the non-conservative form by using the chain rule. At this place the assumption of quasi-constant slip ratio is introduced. With this assumption we obtain

$$\begin{aligned}
 dv_l = & -\frac{dp}{G^{*2}} - f_0 \sum_{l=1}^3 \frac{X_l S_l}{\rho_l^2} \left( \frac{\partial \rho_l}{\partial s_l} ds_l + \sum_{\substack{l=1 \\ l \neq m}}^3 \frac{\partial \rho_l}{\partial C_{il}} dC_{il} \right) \\
 & + \sum_{\substack{l=1 \\ l \neq m}}^3 \left[ f_0 (v_l S_l - v_m S_m) + v_s \left( \frac{1}{S_l} - \frac{1}{S_m} \right) \right] dX_l, \quad (8.115)
 \end{aligned}$$

$$d\rho = \frac{\rho^2}{f_1 f_0 G^{*2}} \left[ dp + f_0 G^{*2} \sum_{l=1}^3 \frac{X_l S_l}{\rho_l^2} \left( \frac{\partial \rho_l}{\partial s_l} ds_l + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} dC_{il} \right) + f_0 G^{*2} \sum_{\substack{l=1 \\ l \neq m}}^3 \left( \frac{S_l - S_m}{\rho} - v_l S_l + v_m S_m \right) dX_l \right], \quad (8.116)$$

where the local critical mass flow rate is

$$\frac{1}{G^{*2}} = f_0 \sum_{l=1}^3 \frac{X_l S_l}{G_l^{*2}} = f_0 \left( \frac{X_1 S_1}{\kappa_1 \rho_1 p} + \frac{X_2 S_2}{G_2^{*2}} + \frac{X_3 S_3}{G_3^{*2}} \right). \quad (8.117)$$

Replacing  $dv_l$  in the mixture mass equation we obtain Eq. (8.118). Equation (8.119) is obtained after differentiating the field mass conservation equations, adding them, and taking into account Eq. (8.116). For the limiting case of three velocity fields this equations reduces to the equation obtained by *Kolev* in (1985). The third group of  $l_{\max} - 1$  equations (8.120) is obtained from the mixture mass conservation equation by using the above derived differential relationships, and after some rearrangements. For two velocity fields these equations reduce to the equation obtained in *Kolev* (1986a, b). The last two equations are obtained easily from the entropy and momentum equations neglecting the diffusion terms. The equations below are obtained assuming also that  $\gamma_v = \gamma_z \neq f(\tau)$ , and setting

$$\gamma^* = \frac{1}{\gamma} \frac{\partial \gamma}{\partial z}.$$

$$\begin{aligned}
 & \frac{\partial G}{\partial \tau} + 2Gv_l \frac{\partial G}{\partial z} + \left( 1 - \frac{G^2}{G^{*2}} \right) \frac{\partial p}{\partial z} + G^2 \sum_{\substack{l=1 \\ l \neq m}}^3 \left[ f_0 (v_l S_l - v_m S_m) + v_s \left( \frac{1}{S_l} - \frac{1}{S_m} \right) \right] \frac{\partial X_l}{\partial z} \\
 & = \gamma_v (f_{fr} + f_{\mu} - \rho g \cos \varphi) - G^2 \left[ v_l \gamma^* - f_0 \sum_{l=1}^3 \frac{X_l S_l}{\rho_l^2} \left( \frac{\partial \rho_l}{\partial s_l} \frac{\partial s_l}{\partial z} + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} \frac{\partial C_{il}}{\partial z} \right) \right], \quad (8.118)
 \end{aligned}$$

$$\begin{aligned} & \frac{\partial p}{\partial \tau} + f_0 G^{*2} \left[ \frac{f_1}{\rho^2} \frac{\partial G}{\partial z} + \sum_{\substack{l=1 \\ l \neq m}}^3 \left( \frac{S_l - S_m}{\rho} - v_l S_l + v_m S_m \right) \frac{\partial X_l}{\partial \tau} \right] \\ & = -f_0 G^{*2} \left[ \frac{f_1}{\rho^2} (G\gamma^* - \mu) + \sum_{l=1}^3 \frac{X_l S_l}{\rho_l^2} \left( \frac{\partial \rho_l}{\partial s_l} \frac{\partial s_l}{\partial \tau} + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} \frac{\partial C_{il}}{\partial \tau} \right) \right], \end{aligned} \quad (8.119)$$

$$\frac{\partial}{\partial \tau} \left( \frac{X_l S_l}{v_s} \gamma_v \right) + \frac{\partial}{\partial z} (X_l G \gamma_z) = \gamma_v \mu_l, \quad (8.120)$$

$$\begin{aligned} & \frac{1}{v_s} \left( \sum_{k \neq l} X_k S_k \right) \frac{\partial}{\partial \tau} (X_l S_l) - X_l S_l \frac{1}{v_s} \sum_{k \neq l} \frac{\partial}{\partial \tau} (X_k S_k) \\ & + f_1 G \frac{\partial X_l}{\partial z} + (f_1 - S_l) X_l \frac{1}{\gamma} \frac{\partial}{\partial z} (G\gamma) = f_1 \mu_l - X_l S_l \mu, \end{aligned} \quad (8.121)$$

$l \neq m$

$$\frac{\partial s_l}{\partial \tau} + \frac{G v_s}{S_l} \frac{\partial s_l}{\partial z} = \frac{v_s}{X_l S_l} \left[ \frac{1}{T_l} D T_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right], \quad l = 1, l_{\max}, \quad (8.122)$$

$$\frac{\partial C_{il}}{\partial \tau} + \frac{G v_s}{S_l} \frac{\partial C_{il}}{\partial z} = \frac{v_s}{X_l S_l} (\mu_{il} - \mu_l C_{il}), \quad l = 1, l_{\max}, i = 1, i_{\max}. \quad (8.123)$$

This non-conservative form

$$\mathbf{A} \frac{\partial \mathbf{U}}{\partial \tau} + \mathbf{B} \frac{\partial \mathbf{U}}{\partial z} = \mathbf{C} \quad (8.124)$$

of the system is very suitable for numerical integration. The entropy and concentration equations can be integrated in each time step separately. With the thus-obtained values for the new time plane, the terms in the first two equations, containing entropies and concentrations can be calculated explicitly, and the resulting system can be integrated with one of the known integration methods (see *Kolev* (1986) p.157), as it was made for a three velocity fields by *Kolev* (1985).

## 8.6 Slip model – steady state. Critical mass flow rate

The local critical mass flow rate plays an important role for the modeling of the three-phase flow in technological equipment, because of the fact that the local mass flow rate in a confined geometry can not exceed the local critical mass flow rate, which is in many cases the limitation of the productivity of the particular apparatus. The local critical mass flow rate can be used as a boundary condition for the analysis of mixture discharge from three-dimensional pressurized vessels.

The purpose of this section is to derive from the steady state part of the system (8.118-8.123) an expression defining the local critical mass flow rate.

We obtain the definition of the local critical mass flow rate from the steady state part of the system, solved as an algebraic equation with respect to the space derivatives. The resulting form of the equivalent system of ordinary differential equations is Eqs. (8.125) to (8.129). As we see the critical condition is reached if the mass flow rate tends to the local critical mass flow rate. This leads to an infinite negative pressure gradient either in the smallest cross section of the channel, or at the end of the duct with constant cross section. In case of three velocity fields the expression defining the local critical mass flow rate reduces to the one obtained by *Kolev* (1985). For three velocity fields without inert components it is reduced to the one derived by *Kolev* (1977). In case of three velocity fields without inert components, incompressible liquid, and a perfect gas the general expression reduces to one obtained by *Nigmatulin* and *Ivandejev* (1977). In case of homogeneous two-phase flow the general expression reduces to the one obtained by *Wood* (1937).

$$\frac{dp}{dz} = - \left\{ -f_{fr} - f_{\mu} + 2Gv_l\mu + \rho\mathbf{g} + G \sum_{\substack{l=1 \\ l \neq m}}^3 \left[ \begin{aligned} & f_0 (v_l S_l - v_m S_m) \\ & + v_s \left( \frac{1}{S_l} - \frac{1}{S_m} \right) \end{aligned} \right] (\mu_l - X_l \mu) \right. \\ \left. - G^2 v_l \gamma^* - f_0 G \sum_{l=1}^3 \frac{S_l}{\rho_l^2} \left( \frac{\partial \rho_l}{\partial s_l} \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right] \right) \right. \\ \left. + \sum_{i=1}^{i_{\max}-1} \frac{\partial \rho_l}{\partial C_{il}} (\mu_{il} - \mu_l C_{il}) \right\} / \left( 1 - \frac{G^2}{G^{*2}} \right) \quad (8.125)$$

$$\frac{dG}{dz} = \mu - G\gamma^*, \quad (8.126)$$

$$\frac{dX_l}{dz} = (\mu_l - X_l \mu) / G, \quad l \neq m, \quad (8.127)$$

$$\frac{ds_l}{dz} = \frac{1}{\gamma X_l G} DS_l^N = \frac{1}{X_l G} \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) \right], \quad l = 1, l_{\max}, \quad (8.128)$$

$$\frac{dC_{il}}{dz} = \frac{1}{\gamma X_l G} DC_{il}^N = \frac{1}{X_l G} (\mu_{il} - \mu_l C_{il}), \quad l = 1, l_{\max}, \quad i = 1, i_{\max}. \quad (8.129)$$

*Single component flashing nozzle flow without external sources:* For a nozzle flow we have no friction forces. The gravitation is negligible. The system simplifies to the following form:

$$\frac{dp}{dz} = -\frac{G}{1-\frac{G^2}{G^{*2}}} \left\{ -f_0 \sum_{l=1}^3 \frac{S_l}{\rho_l^2} \frac{\partial \rho_l}{\partial s_l} \frac{1}{T_l} DT_l^N + \sum_{\substack{l=1 \\ l \neq m}}^3 \left[ f_0 (v_l S_l - v_m S_m) + v_s \left( \frac{1}{S_l} - \frac{1}{S_m} \right) \right] \mu_l \right\} \quad (8.130)$$

$$G = \text{const}, \quad (8.131)$$

$$\frac{dX_l}{dz} = \frac{\mu_l}{G}, \quad l \neq m, \quad (8.132)$$

$$\frac{ds_l}{dz} = \frac{1}{X_l G T_l} DT_l^N, \quad l = 1, l_{\max}. \quad (8.133)$$

The knowledge of the particular interfacial heat, mass and momentum transfer is the prerequisite for integrating the above system. In the 1960s several authors proposed different assumptions replacing the mechanistic description of the interfacial transfer mechanism, like homogeneity for the interfacial momentum description, thermodynamic equilibrium or limited non-equilibrium for the description of the heat and mass transfer etc. Probably the most successful approach was those proposed by *Henry and Fauske* in 1969. We will implement the main ideas of their model applying them consequently to three-field flow. We first compute the heat fluxes required to keep the thermal equilibrium between the fields. Then we will assume that only a part of this heat transfer happens in the real nozzle flow. Then we will compute the evaporation required to keep the liquid fields saturated and assume that only part of this evaporation happens in the real discharge after the nozzle. We will define under this assumption the critical mass flow rate in the throat of the nozzle and then use the simplified integrated *Euler* equations (8.90) to compute the critical pressure ratio. Then knowing the pressure ratio we compute the critical mass flow rate.

*Modeling of the interfacial heat transfer:* The transferred heat between the velocity fields will be a  $b_l$ -part of those required to equalize instantaneously their temperatures defined by Eqs. (6.13) and (6.15):

$$\dot{q}_{21}''' + \dot{q}_{31}''' = -b_1 c_{p1} \left( \frac{1}{n} - \frac{1}{\kappa_1} \right) \frac{X_1 G T_1}{p} \frac{dp}{dz}, \quad (8.134)$$

$$\dot{q}_{21}''' = -b_1 X_2 G T_2 c_{p2} \frac{n-1}{n} \frac{1}{p} \frac{dp}{dz}, \quad (8.135)$$

$$\dot{q}_{31}''' = -b_1 X_3 G T_3 c_{p3} \frac{n-1}{n} \frac{1}{p} \frac{dp}{dz}. \quad (8.136)$$

The heat removed by the liquids due to the evaporation is connected through the interfacial energy jump condition with the evaporated masses per unit time and unit mixture volume as follows

$$\dot{q}_2^{\prime\prime\sigma 1} = -\mu_{21}(h'' - h'), \quad (8.137)$$

$$\dot{q}_3^{\prime\prime\sigma 1} = -\mu_{31}(h'' - h'). \quad (8.138)$$

Thus, the sources in the entropy equations (8.133) are

$$DT_1^N = \dot{q}_{21}^{\prime\prime\prime} + \dot{q}_{21}^{\prime\prime\prime} + (\mu_{21} + \mu_{31})(h'' - h_1), \quad (8.139)$$

$$DT_2^N = -\dot{q}_{21}^{\prime\prime\prime} + \dot{q}_2^{\prime\prime\prime\sigma 1} - \mu_{21}(h' - h_2) = -\dot{q}_{21}^{\prime\prime\prime} - \mu_{21}(h'' - h_2), \quad (8.140)$$

$$DT_3^N = -\dot{q}_{31}^{\prime\prime\prime} + \dot{q}_3^{\prime\prime\prime\sigma 1} - \mu_{31}(h' - h_2) = -\dot{q}_{31}^{\prime\prime\prime} - \mu_{31}(h'' - h_3). \quad (8.141)$$

*Modeling of the interfacial mass transfer:* The evaporation required to keep the fields 2 and 3 saturated is computed from the entropy equations

$$\mu_{21} = -X_2 G \frac{T_2}{h'' - h_2} \left( \frac{ds'}{dp} - b_1 \frac{c_{p2}}{p} \frac{n-1}{n} \right) \frac{dp}{dz}, \quad (8.142)$$

$$\mu_{31} = -X_3 G \frac{T_3}{h'' - h_3} \left( \frac{ds'}{dp} - b_1 \frac{c_{p3}}{p} \frac{n-1}{n} \right) \frac{dp}{dz}. \quad (8.143)$$

The real evaporation may be a  $b_2$ -part of this amount

$$\mu_{21} = -b_2 X_2 G \frac{T_2}{h'' - h_2} \left( \frac{ds'}{dp} - b_1 \frac{c_{p2}}{p} \frac{n-1}{n} \right) \frac{dp}{dz}, \quad (8.144)$$

$$\mu_{31} = -b_2 X_3 G \frac{T_3}{h'' - h_3} \left( \frac{ds'}{dp} - b_1 \frac{c_{p3}}{p} \frac{n-1}{n} \right) \frac{dp}{dz}, \quad (8.145)$$

and therefore

$$\mu_{21} + \mu_{31} = -b_2 G \left[ \frac{X_2 T_2}{h'' - h_2} \left( \frac{ds'}{dp} - b_1 \frac{c_{p2}}{p} \frac{n-1}{n} \right) + \frac{X_3 T_3}{h'' - h_3} \left( \frac{ds'}{dp} - b_1 \frac{c_{p3}}{p} \frac{n-1}{n} \right) \right] \frac{dp}{dz}. \quad (8.146)$$

The change of the gas mass concentration is then

$$dX_1 = -b_2 \left[ \left( \frac{X_2 T_2}{h'' - h_2} + \frac{X_3 T_3}{h'' - h_3} \right) \frac{ds'}{dp} - b_1 \left( \frac{X_2 T_2 c_{p2}}{h'' - h_2} + \frac{X_3 T_3 c_{p3}}{h'' - h_3} \right) \frac{1}{p} \frac{n-1}{n} \right] dp. \quad (8.146)$$

*The critical mass flow rate:* The momentum equation is then rewritten as follows

$$\frac{dp}{dz} = -\frac{G}{1-\frac{G^2}{G^{*2}}} \left\{ \begin{aligned} & -f_0 \left( \frac{S_1}{\rho_1^2} \frac{\partial \rho_1}{\partial s_1} \frac{1}{T_1} DT_1^N + \frac{S_2}{\rho_2^2} \frac{\partial \rho_2}{\partial s_2} \frac{1}{T_2} DT_2^N + \frac{S_3}{\rho_3^2} \frac{\partial \rho_3}{\partial s_3} \frac{1}{T_3} DT_3^N \right) \\ & + \left[ f_0 (v_1 S_1 - v_2 S_2) + v_s \left( \frac{1}{S_1} - \frac{1}{S_2} \right) \right] (\mu_{21} + \mu_{31}) \\ & - \left[ f_0 (v_3 S_3 - v_2 S_2) + v_s \left( \frac{1}{S_3} - \frac{1}{S_2} \right) \right] \mu_{31} \end{aligned} \right\} \quad (8.147)$$

Substituting the sources we obtain

$$\begin{aligned} \frac{1}{G^2} &= \frac{1}{G^{*2}} + f_0 \frac{b_1}{p} \left[ \begin{aligned} & -\frac{S_1}{\rho_1^2} \frac{\partial \rho_1}{\partial s_1} X_1 c_{\rho 1} \left( \frac{1}{n} - \frac{1}{\kappa_1} \right) \\ & + \left( \frac{S_2}{\rho_2^2} \frac{\partial \rho_2}{\partial s_2} X_2 c_{\rho 2} + \frac{S_3}{\rho_3^2} \frac{\partial \rho_3}{\partial s_3} X_3 c_{\rho 3} \right) \frac{n-1}{n} \end{aligned} \right] \\ &+ b_2 \left[ \begin{aligned} & f_0 (v_1 S_1 - v_2 S_2) + v_s \left( \frac{1}{S_1} - \frac{1}{S_2} \right) \left[ \frac{X_2 T_2}{h'' - h_2} \left( \frac{ds'}{dp} - b_1 \frac{c_{\rho 2}}{p} \frac{n-1}{n} \right) \right. \\ & \left. - f_0 (h'' - h_1) \frac{S_1}{\rho_1^2} \frac{\partial \rho_1}{\partial s_1} \frac{1}{T_1} \left[ \frac{X_3 T_3}{h'' - h_3} \left( \frac{ds'}{dp} - b_1 \frac{c_{\rho 3}}{p} \frac{n-1}{n} \right) \right] \right] \\ & + b_2 X_2 \frac{S_2}{\rho_2^2} \frac{\partial \rho_2}{\partial s_2} f_0 (h'' - h_2) \left( \frac{ds'}{dp} - b_1 \frac{c_{\rho 2}}{p} \frac{n-1}{n} \right) \\ & - b_2 X_3 \left[ \begin{aligned} & f_0 (v_3 S_3 - v_2 S_2) + v_s \left( \frac{1}{S_3} - \frac{1}{S_2} \right) \right] \frac{T_3}{h'' - h_3} \left( \frac{ds'}{dp} - b_1 \frac{c_{\rho 3}}{p} \frac{n-1}{n} \right) \\ & + \frac{S_3}{\rho_3^2} \frac{\partial \rho_3}{\partial s_3} \frac{1}{T_3} f_0 (h'' - h_3) \end{aligned} \right] \end{aligned} \right]. \quad (8.148) \end{aligned}$$

It is a very good approximation to neglect the change of the liquid density with the change of the liquid entropy. In addition the relative deviation of the gas temperature with respect to the saturation is very small. Taking into account this simplification the final result is therefore



$$\frac{1}{f_0 G^2} = \frac{X_1 S_1}{\rho_1 p} \left( \frac{1}{\kappa_1} + \frac{b_1}{n} \right) + \frac{X_2 S_2}{G_2^{*2}} + \frac{X_3 S_3}{G_3^{*2}} + b_2 \left\{ \begin{aligned} & \left[ v_1 S_1 - v_2 S_2 + \frac{v_S}{f_0} \left( \frac{1}{S_1} - \frac{1}{S_2} \right) \right] \frac{X_2 T_2}{h'' - h_2} \left( \frac{ds'}{dp} - b_1 \frac{c_{p2}}{p} \frac{n-1}{n} \right) \\ & + \left[ v_1 S_1 - v_3 S_3 + \frac{v_S}{f_0} \left( \frac{1}{S_1} - \frac{1}{S_3} \right) \right] \frac{X_3 T_3}{h'' - h_3} \left( \frac{ds'}{dp} - b_1 \frac{c_{p3}}{p} \frac{n-1}{n} \right) \end{aligned} \right\}. \quad (8.149)$$

Thus at the critical cross section the critical mass flow rate is defined by the above equation. If the  $b_2$  coefficient is set to zero, the model results in the so called frozen model. If the  $b_2$  coefficient is set to unity the model is close to the so called thermal equilibrium model. For gas mass concentrations different from zero the second and the third terms are much smaller than the first one. For homogeneous flow we have

$$\frac{1}{G^2} = \frac{X_1}{\rho_1 p} \left( \frac{1}{\kappa_1} + \frac{b_1}{n} \right) + b_2 \left[ (v_1 - v_2) \frac{X_2 T_2}{h'' - h_2} \left( \frac{ds'}{dp} - b_1 \frac{c_{p2}}{p} \frac{n-1}{n} \right) + (v_1 - v_3) \frac{X_3 T_3}{h'' - h_3} \left( \frac{ds'}{dp} - b_1 \frac{c_{p3}}{p} \frac{n-1}{n} \right) \right]. \quad (8.150)$$

Comparing this result with the derivation by *Henry and Fauske* (1969)

$$\frac{1}{G^2} = \frac{X_1}{n \rho_1 p} + (v_1 - v_2) \left[ \frac{(1 - X_1) b_2}{s'' - s'} \frac{ds'}{dp} - X_1 \frac{c_{p1}}{p (s_1 - s_2)} \left( \frac{1}{n} - \frac{1}{\kappa_1} \right) \right], \quad (8.151)$$

we see a slight difference.

The gas density changes rapidly in the process in accordance with  $\rho_1 = \rho_{10} \varepsilon^{\frac{1}{n}}$ . Replacing in Eq. (8.149) and introducing the pressure ratios we obtain

$$\frac{1}{f_0 G^2} = \frac{X_1 S_1}{\rho_{10} p_0 \varepsilon^{\frac{n+1}{n}}} \left( \frac{1}{\kappa_1} + \frac{b_1}{n} \right) + \frac{X_2 S_2}{G_2^{*2}} + \frac{X_3 S_3}{G_3^{*2}} + b_2 \left\{ \begin{aligned} & \left[ \frac{v_{10} S_1}{\varepsilon^{\frac{1}{n}}} - v_2 S_2 + \frac{v_S}{f_0} \left( \frac{1}{S_1} - \frac{1}{S_2} \right) \right] \frac{X_2 T_2}{h'' - h_2} \left( \frac{ds'}{dp} - b_1 \frac{c_{p2}}{p_0 \varepsilon} \frac{n-1}{n} \right) \\ & + \left[ \frac{v_{10} S_1}{\varepsilon^{\frac{1}{n}}} - v_3 S_3 + \frac{v_S}{f_0} \left( \frac{1}{S_1} - \frac{1}{S_3} \right) \right] \frac{X_3 T_3}{h'' - h_3} \left( \frac{ds'}{dp} - b_1 \frac{c_{p3}}{p_0 \varepsilon} \frac{n-1}{n} \right) \end{aligned} \right\}, \quad (8.152)$$

where

$$v_s = X_1 S_1 \frac{v_{10}}{\varepsilon^n} + X_2 S_2 v_2 + X_3 S_3 v_3, \quad (8.153)$$

$$X_1 = X_{10} - b_2 \left[ \left( \frac{X_2 T_2}{h'' - h_2} + \frac{X_3 T_3}{h'' - h_3} \right) \frac{ds'}{dp} - b_1 \left( \frac{X_2 T_2 c_{p2}}{h'' - h_2} + \frac{X_3 T_3 c_{p3}}{h'' - h_3} \right) \frac{1}{p} \frac{n-1}{n} \right] p_0 (\varepsilon - 1), \quad (8.154)$$

$$X_2 = X_{20} + b_2 X_2 G \frac{T_2}{h'' - h_2} \left( \frac{ds'}{dp} - b_1 \frac{c_{p2}}{p} \frac{n-1}{n} \right) p_0 (\varepsilon - 1), \quad (8.155)$$

and

$$X_3 = 1 - X_1 - X_2. \quad (8.156)$$

Equation (8.152) together with Eq. (8.90)

$$\frac{1}{2} v_s f_0 G^2 = p_0 \left[ \alpha_1 \varepsilon^{1/\kappa_1} \frac{\kappa_1}{\kappa_1 - 1} \left( 1 - \varepsilon^{\frac{\kappa_1 - 1}{\kappa_1}} \right) + (1 - \alpha_1)(1 - \varepsilon) \right] \quad (8.157)$$

define the critical pressure ratio. Both equations have to be solved by iterations with respect to  $G$  and  $\varepsilon$ . For estimation of the slip ratios, a reasonable assumption is  $S_2 = S_3 = 1$  and  $S_1$  computed using

$$S_1 = 1 + \frac{X_1}{X_{1,\max}} \left( \frac{1 - X_1}{1 - X_{1,\max}} \right)^{\frac{1 - X_1}{X_{1,\max}}} (S_{1,\max} - 1), \quad (8.158)$$

$$S_{1,\max} = \left( \frac{\rho_2}{\rho_1} \right)^{0.1}, \quad (8.159)$$

$$X_{1,\max} = 0.1. \quad (8.160)$$

*Henry and Fauske* proposed to use

$$b_2 = \frac{X_{10,e}}{0.14} \quad \text{for } X_{10,e} < 0.14 \quad (8.161)$$

and  $b_2 = 1$  otherwise, where

$$X_{1,e} = \frac{\sum_{l=1}^3 X_{l0} s_{l0} - s'_0}{s''_0 - s'_0}. \quad (8.162)$$

As already mentioned the pressure gradient at the critical flow location is minus infinite. Such gradients are very difficult to be resolved numerically and special

attention has to be paid to the discretization. Non-equidistant discretization with sizes getting smaller in the proximity of the critical cross section is the right answer of this challenge.

As already mentioned the critical condition is reached if the mass flow rate tends to the local critical mass flow rate manifested by infinite negative pressure gradient either in the smallest cross section of the channel, or at the end of the duct with constant cross section. The numerical resolution of such gradients is difficult and needs special attention.

**Grid structure:** It is recommendable to use fine discretization around the critical cross section and not so fine elsewhere. I recommended in *Kolev* (1986) p. 177 to use a grid sizes building a geometrical sequence

$$\Delta z_k = \Delta z_1 c^{k-1}, \quad c < 1,$$

so that the sum of the grid lengths give the length of the pipe

$$\sum_{k=1}^{k_{\max}} \Delta z_k = \Delta z_1 (1 - c^{k_{\max}+1}) / (1 - c) = L$$

and therefore

$$\Delta z_1 = L(1 - c) / (1 - c^{k_{\max}+1}).$$

The constant  $c$  is controlling the grid sizes. Renormalization of the grid sizes is necessary in the numerical discretization in order to have a strict conservation of the length. Later on other strategies are reported e.g. *Wendt* (1992). Counting the  $z$ -coordinate from the entrance of the pipe the coordinates of the cell boundaries are defined as

$$\frac{z_k}{L} = \frac{\zeta_{crit}}{a} \left\{ \sinh [(\zeta_k - \zeta')c] + a \right\},$$

where

$$\zeta_k = (k-1)/k_{\max}, \quad k = 1, 2, \dots, k_{\max}, \quad a = \sinh(c\zeta'),$$

and

$$\zeta' = \frac{1}{2c} \ln \left[ \frac{1 + (e^c - 1)\zeta_{rit}}{1 + (e^{-c} - 1)\zeta_{rit}} \right].$$

Here the dimensionless position of the critical cross section is

$$\zeta_{crit} = z_{crit} / L.$$

The cell sizes are controlled by the choice of the constant  $c \approx 2 \dots 3$ .

For *Laval* nozzles I recommended in *Kolev* (1986) p. 178 the following discretization for the converging part decreasing cell sizes sequence

$$\Delta z_m = \Delta z_{conv,1} c_{conv}^{m-1}, c_{conv} < 1, m = 1, m_{max},$$

$$\Delta z_{conv,1} = z_{crit} (1 - c_{conv}) / (1 - c_{conv}^{m_{max}+1}),$$

and for the diverging part increasing cell sizes sequence

$$\Delta z_n = \Delta z_{div,1} c_{div}^{n-1}, c_{div} > 1, n = 1, n_{max},$$

$$\Delta z_{div,1} = (L - z_{crit}) (1 - c_{div}) / (1 - c_{div}^{n_{max}+1}),$$

where  $m_{max} + n_{max} = k_{max}$ .

**Iteration strategy:** The mathematical formalism describing the critical flow do not allow mass flow rates larger then the critical. Therefore the critical mass flow rate  $G^*$  can be approached only using increasing trial values of  $G_1, G_2, \dots, G^*$ . The value of the denominator of the pressure gradient  $N_1, N_2, \dots, \mathcal{E}$  is the controlled variable. The target is the small number  $\mathcal{E} \rightarrow 0$ . Simple way to reach the target is if the target is becoming negative to half the step  $\Delta G$  and to repeat the computation as long the  $\Delta G$  is becoming small enough. Optimization strategies are also possible. I recommended in *Kolev* (1986) to construct a polynomial based on the couples  $(G_i, N_i)$  then to set  $N = 0$  and to obtain the next guess for  $G^*$ .

I give in Appendix 1 additional 110 chronologically ordered references on the subject of critical two-phase flows for those of you who are interested on the history of this subject.

## 8.7 Forces acting on the pipes due to the flow – theoretical basics

In many applications the analysis of thermo-hydraulic processes in a networks or 3D facilities is done with the intention to estimate the thrusts acting on the pipes and facilities. These thrusts are used to design the mechanical supports of the constructions.

The purpose of this chapter is to describe the algorithm needed for computation of the pipe thrusts in the network operating with multi-phase flows.

In the following we assume a) that the effect of gravitational forces is negligible, and postulate that b) the positive flow force direction is the direction opposite to the positive flow velocity direction.

The solid structure experiences forces from the continuum flow wetting its internal surface and from the ambient fluid. There are normal and shear forces caused by the flow and acting on the internal wall surface and normal forces due the action of the ambient pressure on the external wall surface.

We designate with  $\mathbf{f}_{wp}$  the normal pressure force acting on the structure. This force consists of two components due to internal and ambient pressure, respectively

$$\mathbf{f}_{wp} = \iint_{A_{\text{wall-ambient}}} p_{\infty} \mathbf{n}_w da + \iint_{A_{\text{wall-flow}}} p \mathbf{n}_f da = \iint_{A_{\text{wall-ambient}}} p_{\infty} \mathbf{n}_w da - \iint_{A_{\text{wall-flow}}} p \mathbf{n}_w da . \quad (8.163)$$

Here the subscript  $w$  stands for wall,  $f$  for flow,  $p$  is the fluid pressure inside the pipe acting on the infinitesimal surface  $da$ ,  $p_{\infty}$  is the ambient pressure acting outside of the pipe on the same surface.  $\mathbf{n}_w$  is the unit vector normal to the internal wall surface pointing into the flow - see Fig. 8.3. The flow friction shear force,  $\mathbf{f}_{f\eta}$ , resists the flow and is positive per definition. The wall experiences a force,  $\mathbf{f}_{w\eta}$ , with magnitude equal to the magnitude of the flow friction force but with the opposite direction,

$$\mathbf{f}_{f\eta} + \mathbf{f}_{w\eta} = 0 . \quad (8.164)$$

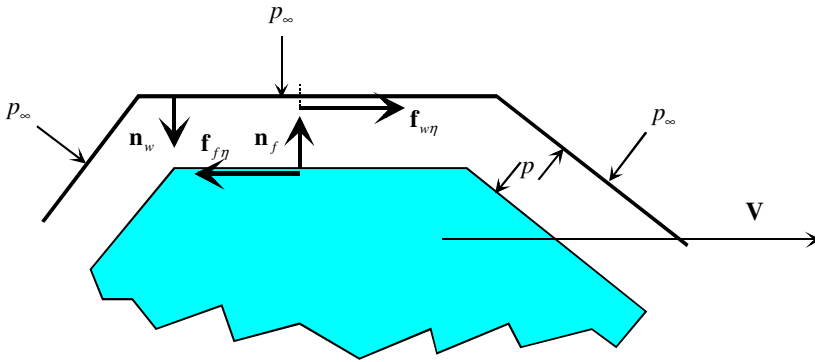


Fig. 8.3 Definitions of the force directions signs

The total force acting on the wall is therefore

$$\mathbf{f}_{\text{wall}} = \mathbf{f}_{wp} + \mathbf{f}_{w\eta} = \mathbf{f}_{wp} - \mathbf{f}_{f\eta} . \quad (8.165)$$

*Direct computation of the fluid friction force:* The fluid friction force can be computed directly by integrating the friction pressure loss through the total flow volume

$$\mathbf{f}_{f\eta} = - \iiint_{\text{flow volume}} (\nabla p)_{\text{friction}} dVol \approx - \iiint_{\text{flow volume}} \left( \frac{dp}{dz} \right)_{\text{friction}} dVol . \quad (8.166)$$

*Indirect computation of the fluid friction force:* The fluid friction force can be computed indirectly, using the momentum balance on the fluid in the control volume,

$$\begin{aligned} & \frac{\partial}{\partial \tau} \iiint_{\text{flow volume}} \sum_{l=1}^{l_{\max}} \alpha_l \rho_l \mathbf{V}_l \gamma_l dVol - \iint_{A_{\text{flow}}} p \mathbf{n}_f da_f \\ & + \iint_{A_{\text{open}}} \sum_{l=1}^{l_{\max}} (\alpha_l \rho_l \mathbf{V}_l \cdot \mathbf{n}_f) \mathbf{V}_l \gamma_l da + \mathbf{f}_{\eta} = 0 . \end{aligned} \quad (8.167)$$

i.e.

$$\begin{aligned}
 -\mathbf{f}_\eta &= \frac{\partial}{\partial \tau} \iiint_{\text{flow volume}} \sum_{l=1}^{l_{\max}} \alpha_l \rho_l \mathbf{V}_l \gamma_v dVol \\
 &- \iint_{A_{\text{flow}}} p \mathbf{n}_f da + \iint_{A_{\text{open}}} \sum_{l=1}^{l_{\max}} (\alpha_l \rho_l \mathbf{V}_l \cdot \mathbf{n}_f) \mathbf{V}_l \gamma_n da .
 \end{aligned} \tag{8.168}$$

Unlike the momentum flux

$$\iint_{A_{\text{open}}} \sum_{l=1}^{l_{\max}} (\alpha_l \rho_l \mathbf{V}_l \cdot \mathbf{n}_f) \mathbf{V}_l \gamma_n da$$

that acts only at the open flow areas and vanishes on the wet surface the friction tension acts at the wet walls and vanishes at the open areas. We split the area surrounding our flow control volume into a part contacting the wall and a open part crossed by the flow

$$A_{\text{flow}} = A_{\text{open}} + A_{\text{wall flow}} , \tag{8.169}$$

and therefore

$$\iint_{A_{\text{flow}}} p \mathbf{n}_f da = \iint_{A_{\text{open}}} p \mathbf{n}_f da + \iint_{A_{\text{wall flow}}} p \mathbf{n}_f da . \tag{8.170}$$

The total wall force is therefore

$$\begin{aligned}
 \mathbf{f}_{\text{wall}} = \mathbf{f}_{\text{wp}} - \mathbf{f}_{f\eta} &= \iint_{A_{\text{wall ambient}}} p_\infty \mathbf{n}_w da - \iint_{A_{\text{wall flow}}} p \mathbf{n}_w da - \iint_{A_{\text{open}}} p \mathbf{n}_f da - \iint_{A_{\text{wall flow}}} p \mathbf{n}_f da \\
 &+ \frac{\partial}{\partial \tau} \iiint_{\text{flow volume}} \sum_{l=1}^{l_{\max}} \alpha_l \rho_l \mathbf{V}_l \gamma_v dVol + \iint_{A_{\text{open}}} \sum_{l=1}^{l_{\max}} (\alpha_l \rho_l \mathbf{V}_l \cdot \mathbf{n}_f) \mathbf{V}_l \gamma_n da .
 \end{aligned} \tag{8.171}$$

Having in mind that the unit vectors of the wall surface and the flow surface are opposite at the wet wall

$$\mathbf{n}_w = -\mathbf{n}_f \tag{8.172}$$

we recognize that in all surfaces where the pressure acts simultaneously on the wall and at the flow  $A_{\text{wall flow}}$  the pressure force cancels and therefore

$$\begin{aligned}
 \mathbf{f}_{\text{wall}} &= - \iint_{A_{\text{wall ambient}}} p_\infty \mathbf{n}_f da - \iint_{A_{\text{open}}} p \mathbf{n}_f da \\
 &+ \frac{\partial}{\partial \tau} \iiint_{\text{flow volume}} \sum_{l=1}^{l_{\max}} \alpha_l \rho_l \mathbf{V}_l \gamma_v dVol + \iint_{A_{\text{open}}} \sum_{l=1}^{l_{\max}} (\alpha_l \rho_l \mathbf{V}_l \cdot \mathbf{n}_f) \mathbf{V}_l \gamma_n da .
 \end{aligned} \tag{8.173}$$

Therefore the wall force consists of three components

(i) pressure force

$$- \iint_{A_{\text{wall ambient}}} p_{\infty} \mathbf{n}_f da - \iint_{A_{\text{open}}} p \mathbf{n}_f da , \tag{8.174}$$

(ii) “wave” force

$$\frac{\partial}{\partial \tau} \iiint_{\text{flow volume}} \sum_{l=1}^{l_{\text{max}}} \alpha_l \rho_l \mathbf{V}_l \gamma_v dVol , \tag{8.175}$$

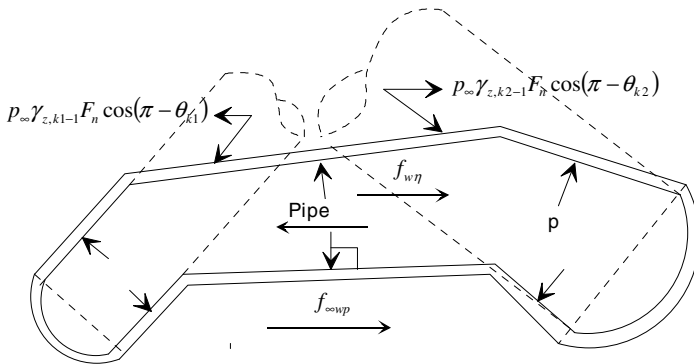
and

(iii) reaction thrust

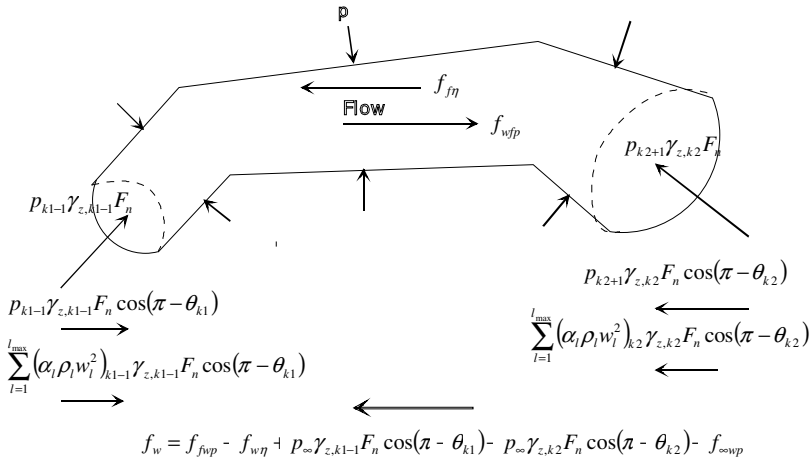
$$\iint_{A_{\text{open}}} \sum_{l=1}^{l_{\text{max}}} \alpha_l \rho_l \mathbf{V}_l^2 \gamma_n \mathbf{n}_f da . \tag{8.176}$$

The estimation of wall forces from the primary Eqs. (8.163-8.166) appears to be much simpler than the final expression Eq. (8.173). Often, however, pressure distribution on the wet wall is not known well enough to obtain reasonable results. That is why Eq. (8.173) is successfully used in practice, see e.g. *Abedin, Takeuchi and Zoung* (1986), *Lahey and Moody* (1977). As illustration of the application of the final equation (8.173) we consider a few practical cases.

*Forces on a pipe segment with two bends:* The pipe segment as given in Fig. 8.4 consists of one straight part and two elbows. The respective angles of the elbows are  $\theta_{k1}$ , and  $\theta_{k2}$ . The velocity in the segment is directed from left to the right. Compute the projection of all wall forces on the main axis of the straight part.



**Fig. 8.4** Forces acting at the flow control volume



**Fig. 8.5** Pressure forces acting on the flow control volume

The pipe segment may be divided into several control volumes starting with  $k1$  and ending with  $k2$ . Each control volume has a length  $\Delta z_k$  and volume  $\gamma_v \Delta z_k F_n$ , where  $\gamma_v$  is the part of the control volume  $\Delta z_k F_n$  occupied by flow. Here  $F_n$  is some constant cross section. The left surface passed by the flow at the left-hand side control volume  $k1$ , is  $\gamma_{z,k1-1} F_n$ , and right hand side surface passed by the flow of the right-hand side control volume  $k2$  is  $\gamma_{z,k2} F_n$ . The normal velocity at the entrance flow cross section of each velocity field is  $w_{l,k1-1}$ , and at the outlet velocity at the outlet flow cross section of each velocity field is  $w_{l,k2}$ , respectively. The flow pressure before the cell  $k1$  inside the pipe at the left is  $p_{k1-1}$ , and the pressure after the cell  $k2$  at the right is  $p_{k2+1}$ . The ambient pressure is  $p_{\infty}$ . The flow entering the control volume accelerates the flow with a force component in the  $z$  direction

$$-\cos(\pi - \theta_{k1}) \left( \sum_{l=1}^{l_{max}} \dot{m}_l w_l \right)_1 = -\gamma_{z,k1-1} F_n \cos(\pi - \theta_{k1}) \left( \sum_{l=1}^{l_{max}} \alpha_l \rho_l w_l^2 \right)_{k1-1}, \quad (8.177)$$

which is directed from the left to the right and is parallel to the flow axis, see Fig. 8.4 and 8.5. The flow leaving the control volume resists the control flow volume with the force component in the  $z$  direction

$$\cos(\pi - \theta_{k2}) \left( \sum_{l=1}^{l_{max}} \dot{m}_l w_l \right)_2 = \gamma_{z,k2} F_n \cos(\pi - \theta_{k2}) \left( \sum_{l=1}^{l_{max}} \alpha_l \rho_l w_l^2 \right)_{k2} \quad (8.178)$$

which is directed from the right to the left and is parallel to the flow axis. The pressure forces acting on the flow control volumes, see Fig. 8.5, are computed as follows. The projection of the pressure force acting at the inlet cross section on the axis,



$$-p_{k1-1}\gamma_{z,k1-1}F_n \cos(\pi - \theta_{k1}), \tag{8.179}$$

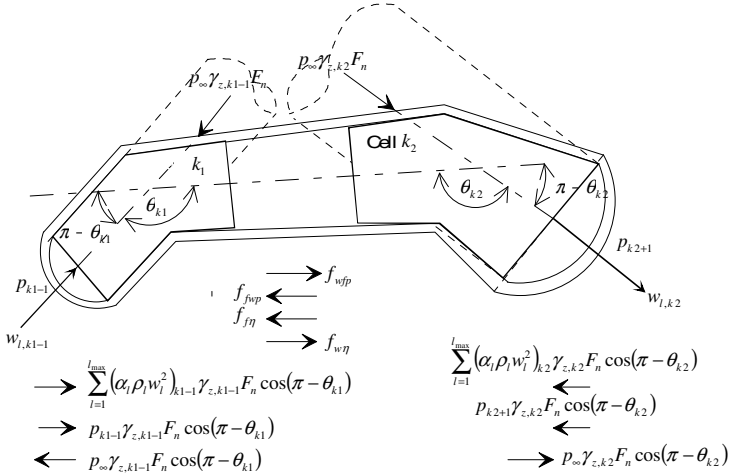
is directed from left to the right. The projection of the pressure force acting at the left side of the flow control volume on the axis,

$$p_{k2+1}\gamma_{z,k2}F_n \cos(\pi - \theta_{k2}), \tag{8.180}$$

is directed from right to the left. In addition at the projection of the pressure force acting on the flow volume from the side of the wet wall surface on the flow axis is  $f_{wfp}$ , and is directed from the left to the right. The opposite force  $f_{fwp}$  acts on the structure at the wet side. Both forces cancel after summation. Thus, the friction force resisting the flow is

$$\begin{aligned} -f_\eta &= F_n \frac{\partial}{\partial \tau} \sum_{k=k1}^{k2} \left( \gamma_v \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l \Delta z \right)_k \\ &- \left( p_{k1-1} + \left( \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l^2 \right)_{k1-1} \right) \gamma_{z,k1-1} F_n \cos(\pi - \theta_{k1}) \\ &+ \left( p_{k2+1} + \left( \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l^2 \right)_{k2} \right) \gamma_{z,k2} F_n \cos(\pi - \theta_{k2}). \end{aligned} \tag{8.181}$$

Now we consider as a control volume the structure only, see Fig. 8.6.



**Fig. 8.6** Total force acting on the pipe in axial direction

The projection of the ambient pressure forces acting at the left cross section  $\gamma_{z,k1-1}F_n$  on the axis is

$$p_{\infty}\gamma_{z,k1-1}F_n \cos(\pi - \theta_{k1}), \tag{8.182}$$

and is directed from the right to the left. The projection of the ambient pressure forces acting at the right cross section  $\gamma_{z,k2}F_n$  on the axis is

$$-p_\infty \gamma_{z,k2} F_n \cos(\pi - \theta_{k2}), \quad (8.183)$$

and is directed from the left to the right. In addition there is a pressure force acting from the fluid side at the wet wall. Its projection on the axis is  $\mathbf{f}_{fwp}$ . The resulting pressure force acting on the structure is

$$\mathbf{f}_{wp} = p_\infty \gamma_{z,k1-1} F_n \cos(\pi - \theta_{k1}) - p_\infty \gamma_{z,k2} F_n \cos(\pi - \theta_{k2}) + \mathbf{f}_{fwp}. \quad (8.184)$$

Now we have all components we need to compute the resulting force acting on the structure

$$\begin{aligned} \mathbf{f}_w &= \mathbf{f}_{wp} + \mathbf{f}_{w\eta} = \mathbf{f}_{wp} - \mathbf{f}_{f\eta} = F_n \frac{\partial}{\partial \tau} \sum_{k=k1}^{k2} \left( \gamma_v \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l \Delta z \right)_k \\ &- \left[ p_{k1-1} - p_\infty + \left( \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l^2 \right)_{k1-1} \right] \gamma_{z,k1-1} F_n \cos(\pi - \theta_{k1}) \\ &+ \left[ p_{k2+1} - p_\infty + \left( \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l^2 \right)_{k2} \right] \gamma_{z,k2} F_n \cos(\pi - \theta_{k2}). \end{aligned} \quad (8.185)$$

The wave force during a time step is computed using the old time level values, designated with  $a$ , and the new time level values - without indices - as follows

$$F_n \frac{\partial}{\partial \tau} \sum_{k=k1}^{k2} \left( \gamma_v \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l \Delta z \right)_k = \frac{1}{\Delta \tau} F_n \frac{\partial}{\partial \tau} \sum_{k=k1}^{k2} \left( \gamma_v \sum_{l=1}^{l_{\max}} [\alpha_l \rho_l w_l - (\alpha_l \rho_l w_l)_a] \Delta z \right)_k. \quad (8.186)$$

For the simple case of a discharging flow from a pipe with dead end,  $\gamma_{z,k1-1} = 0$ , and constant cross section  $F_n$ ,  $\gamma_{v,k} = 1$ , we obtain

$$\begin{aligned} \mathbf{f}_w &= \mathbf{f}_{wp} - \mathbf{f}_{f\eta} = \frac{1}{\Delta \tau} F_n \frac{\partial}{\partial \tau} \sum_{k=k1}^{k2} \left\{ \gamma_v \sum_{l=1}^{l_{\max}} [\alpha_l \rho_l w_l - (\alpha_l \rho_l w_l)_a] \Delta z \right\}_k \\ &+ \left[ p_{k2+1} - p_\infty + \left( \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l^2 \right)_{k2} \right] \gamma_{z,k2} F_n. \end{aligned} \quad (8.187)$$

For the same case as that previously discussed but with critical discharge the critical pressure,  $p_c$ , should replace  $p_{k2+1}$  in the above equation - see the discussion by Yano et al. (1982).

$$\mathbf{f}_w = \mathbf{f}_{wp} - \mathbf{f}_{f\eta} = \frac{1}{\Delta \tau} F_n \frac{\partial}{\partial \tau} \sum_{k=k1}^{k2} \left\{ \gamma_v \sum_{l=1}^{l_{\max}} [\alpha_l \rho_l w_l - (\alpha_l \rho_l w_l)_a] \Delta z \right\}_k$$

$$+ \left[ p_c - p_\infty + \left( \sum_{l=1}^{l_{\max}} \alpha_l \rho_l w_l^2 \right) \right]_{k2} \gamma_{z,k2} F_n. \quad (8.188)$$

## 8.8 Relief valves

### 8.8.1 Introduction

Relief valves are designed to keep the system pressure under a prescribed value, and therefore are important safety components of industrial networks. Operating valves excites considerable forces in pipe networks. If the design of the support structures is based on the *steady state analysis* only the structures may be destroyed by *transient forces*. That is why proper design of dynamic valve behavior is a necessary step towards modeling industrial pipe networks. Usually the second order ordinary differential equation describing the piston motion is solved by finite difference methods with time step limited by the non-dumped oscillation time constant of the valve. In this Section we demonstrate a piecewise analytical solution which removes this limitation and increases the stability of the numerical solution for the flow itself.

### 8.8.2 Valve characteristics, model formulation

The commonly used components of a relief valve are: valve housing, inlet, outlet, piston, rod assembly, spring, bellows, and valve adjusting ring assembly, see Fig. 8.7. *Back pressing valves* belong to the same category and can also be described by the models considered in this section. The main difference between relief and back pressing valves is their mode of operation. Relief valves are usually closed during regular operation and open if the inlet pressure exceeds a prescribed value, whereas back pressing valves are open under regular conditions and closed in case of a sudden pressure loss. Also the damping mechanisms may be different.

The dynamic behavior of a valve is uniquely defined at any moment  $\tau$  if we know the piston position  $z$  and the dependence of the smallest flow cross section as a function of the piston position  $A^*(z)$  - see the Nomenclature at the end of Section 8. The inlet-outlet flow paths are usually absorbed as a part of the pipe model. Knowing  $A^*(z)$  the dimensionless surface permeability at the contraction cross section used in the network analysis,

$$\gamma_z = A^*(z) / A_{norm} \quad (8.189)$$

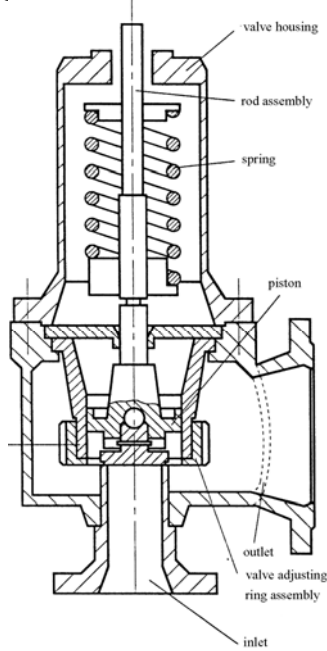
can be computed, where  $A_{norm}$  is some normalizing cross section in  $m^2$ , e.g. the maximum cross section in the pipe where the valve is installed. The quantitative characteristics needed to describe the dynamic behavior of the valve given in Fig. 8.8 are given in the Nomenclature. Valves without adjusting ring assembly,

$A_{out} = 0$ , do not develop additional reaction force. Valves provided with adjusting ring assembly with lower end being below the closing surface of the valve in closed position,

$$z_r \leq 0, \quad (8.190)$$

$$\varphi = 0, \quad (8.191)$$

develop the largest additional flow reaction force at the valve piston and consequently have the lowest closing pressure for the particulate valve geometry. If the position of the lower end of the valve-adjusting-ring assembly is above the closing surface of the valve during the valve operation



**Fig. 8.7** Valve components

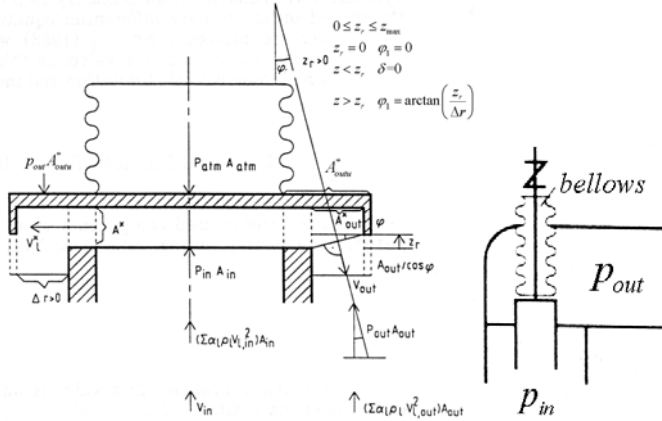


Fig. 8.8 Valve characteristics

$$z_r > 0, \tag{8.192}$$

there are two possible regimes for the flow to develop additional reaction force at the piston. The first regime is defined if the piston position is below the lower end of the valve adjusting ring assembly

$$z \leq z_r. \tag{8.193}$$

In this case the adjusting ring assembly has no influence on the flow behavior. We describe this by setting a multiplier

$$\varphi = \pi / 2, \tag{8.194}$$

for the additional flow reaction force. In the second regime the piston position is above the lower end of the valve adjusting ring assembly

$$z > z_r. \tag{8.195}$$

In this case the angle between the velocity vectors of the inlet and outlet flows is

$$\varphi = \arctan(z_r / \Delta r). \tag{8.196}$$

In this case the additional flow reaction force at the piston is between zero and the maximum possible value.

Before writing down the momentum equation for the moving assembly of the valve let us summarize the forces acting on the moving valve mechanism:

$$\text{Inertia force } m_v \frac{d^2 z}{d\tau^2}; \tag{8.197}$$

$$\text{Friction force } C_{fr} \frac{dz}{d\tau}; \tag{8.198}$$

$$\text{Spring force } C_{sp} (z + z_0); \quad (8.199)$$

$$\text{Gravitation force } m_v g; \quad (8.200)$$

Pressure difference forces for open valve

$$p_{in}(\tau)A_{in} + p_{out}(\tau)A_{out} - p_{outu}(\tau)A_{outu} - p_{atm}(\tau)A_{atm}, \quad (8.201)$$

or for

$$p_{out}(\tau) \approx p_{outu}(\tau), \quad (8.202)$$

$$p_{in}(\tau)A_{in} + p_{out}(\tau)(A_{out} - A_{outu}) - p_{atm}(\tau)A_{atm}. \quad (8.203)$$

The fluid momentum flux force for open valve is

$$\begin{aligned} f_{fl} &= A_{in} \sum_{l=1}^{l_{\max}} \alpha_l \rho_l V_{l,in}^2 + A_{out} \sum_{l=1}^{l_{\max}} \alpha_l \rho_l V_{l,out}^2 \approx \left( \sum_{l=1}^{l_{\max}} \alpha_l^* \rho_l^* V_l^* \right) A^{*2} \left( \frac{1}{A_{in}} + \frac{1}{A_{out}} \cos^2 \varphi \right) \\ &\approx \left( \sum_{l=1}^{l_{\max}} \alpha_l^* \rho_l^* V_l^* \right) A^{*2} \left( \frac{1}{A_{in}} + \frac{1}{A_{out}} \frac{\Delta r^2}{z_r^2 + \Delta r^2} \right) \end{aligned} \quad (8.204)$$

for  $A_{out} > 0$ , and

$$f_{fl} \approx \left( \sum_{l=1}^{l_{\max}} \alpha_l^* \rho_l^* V_l^* \right) A^{*2} / A_{in} \quad (8.205)$$

for  $A_{out} = 0$ . Here  $A_{out} \sum_{l=1}^{l_{\max}} \alpha_l \rho_l V_{l,out}^2$  is the additional flow reaction force at the piston. The field velocities  $V_l^*$  at the *vena contracta*,  $A^*$ , for open valve are computed with the fluid dynamic model of the flow in the network. Assuming

$$(\alpha_l \rho_l)_{in} \approx (\alpha_l \rho_l)_{out} = (\alpha_l \rho_l)^* \quad (8.206)$$

the inlet and outlet velocities can be approximately estimated as follows

$$V_{l,in} = V_l^* A^* / A_{in}, \quad (8.207)$$

$$V_{l,out} = V_l^* (A^* / A_{out}) \cos \varphi \quad \text{for } A_{out} > 0. \quad (8.208)$$

The momentum principle applied to the valve moving mechanism gives

$$\begin{aligned} m_v \frac{d^2 z}{d\tau^2} + C_{fr} \frac{dz}{d\tau} + C_{sp} (z + z_0) + m_v g - \\ \left[ p_{in}(\tau)A_{in} + p_{out}(\tau)(A_{out} - A_{outu}) - p_{atm}(\tau)A_{atm} \right] + f_{fl} = 0. \end{aligned} \quad (8.209)$$

Having in mind that for steady state the following force equilibrium is valid at the moment the valve is just starting to open,

$$m_v g + C_{sp} z_0 = p_{on}(\tau) A_{in} + p_{out}(\tau)(A_{out} - A_{outu}) - p_{atm}(\tau) A_{atm} \quad (8.210)$$

we replace

$$m_v g + C_{sp} z_0 - p_{out}(\tau)(A_{out} - A_{outu}) + p_{atm}(\tau) A_{atm} = p_{on}(\tau) A_{in} \quad (8.211)$$

and obtain a simpler dynamic force balance

$$m_v \frac{d^2 z}{d\tau^2} + C_{fr} \frac{dz}{d\tau} + C_{sp} z = [p_{in}(\tau) - p_{on}(\tau)] A_{in} + f_{fl} = 0. \quad (8.212)$$

In case  $A_{out} \approx A_{outu}$ ,  $p_{on}(\tau) \approx p_{on}$  is no longer a function of time. In this case instead of the input information  $z_0$ ,  $A_{out}$ ,  $A_{outu}$ ,  $A_{atm}$  only  $p_{on}$  is necessary for the description of the valve behavior.

Next we look for a piecewise analytical solution of this equation assuming that within short time interval

$$\Delta \tau, \quad (8.213)$$

the input pressure is constant

$$p_{in}(\tau) \approx const. \quad (8.214)$$

### 8.8.3 Analytical solution

As already mentioned the differential equation describing the piston motion is solved usually by finite difference methods – see *Ransom et al.* (1988). This is associated with a time step controlled by the criterion

$$\Delta \tau \approx 0.01 \Delta \tau_v, \quad (8.215)$$

where  $\Delta \tau_v$  is the oscillation time constant. In order to avoid this limitation a piecewise quasi analytical method is applied in this work. Next we describe the theoretical basics for this solution – for more information see *Greiner* (1984), *Sass, Bouché and Leitner* (1969), *Magnus* (1986). First we use the text book solution of the homogeneous equation

$$m_v \frac{d^2 z}{d\tau^2} + C_{fr} \frac{dz}{d\tau} + C_{sp} z = 0. \quad (8.216)$$

The eigenvalues  $\lambda_{1,2}$  are the roots of the equation

$$m_v \lambda^2 + C_{fr} \lambda + C_{sp} = 0, \quad (8.217)$$

$$\lambda_{1,2} = -\frac{1}{2} \frac{C_{fr}}{m_v} \pm \sqrt{\left(\frac{1}{2} \frac{C_{fr}}{m_v}\right)^2 - \frac{C_{sp}}{m_v}} = -\frac{1}{\Delta\tau_d} \pm \sqrt{-1} \frac{2\pi}{\Delta\tau_v}. \quad (8.218)$$

a) For the *harmonic oscillations* case,

$$\frac{C_{sp}}{m_v} > \left(\frac{1}{2} \frac{C_{fr}}{m_v}\right)^2, \quad (8.219)$$

the eigenvalues are complex:

$$\lambda_{1,2} = -\frac{1}{\Delta\tau_d} \pm i \frac{2\pi}{\Delta\tau_v}. \quad (8.220)$$

Here the time constants of the process are

(i) the *damping time constant*

$$\Delta\tau_d = 2m_v / C_{fr}, \quad (8.221)$$

and

(ii) the *oscillation period*

$$\Delta\tau_v = \frac{2\pi}{\sqrt{\frac{C_{sp}}{m_v} - \frac{1}{\Delta\tau_d^2}}}. \quad (8.222)$$

The oscillation frequency is then

$$\omega = 2\pi / \Delta\tau_v. \quad (8.223)$$

The text book parametric solution of the homogeneous equation is therefore

$$z(\tau) = e^{-\tau/\Delta\tau_d} [C_1 \cos(\omega\tau) + C_2 \sin(\omega\tau)] \quad (8.224)$$

and the general parametric solution of the non-homogeneous equation

$$z(\tau) = e^{-\tau/\Delta\tau_d} [C_1 \cos(\omega\tau) + C_2 \sin(\omega\tau)] + z_{\max}^{**} / 2, \quad (8.225)$$

where

$$z_{\max}^{**} = \frac{2}{C_{sp}} \{ [p_{in}(\tau) - p_{on}(\tau)] A_m + f_{fl} \}. \quad (8.226)$$

b) For the *aperiodic* case of

$$\frac{C_{sp}}{m_v} < \left(\frac{1}{2} \frac{C_{fr}}{m_v}\right)^2 \quad (8.227)$$



$$\Delta\tau_v = \frac{2\pi}{\sqrt{\frac{1}{\Delta\tau_d^2} - \frac{C_{sp}}{m_v}}} . \quad (8.228)$$

the eigenvalues are real

$$\lambda_{1,2} = -\frac{1}{\Delta\tau_d} \pm \frac{2\pi}{\Delta\tau_v} \quad (8.229)$$

and

$$\lambda_1 \neq \lambda_2 . \quad (8.230)$$

In this case the text book solution is

$$z(\tau) = C_1 e^{\lambda_1 \tau} + C_2 e^{\lambda_2 \tau} + z_{\max}^{**} / 2 . \quad (8.231)$$

c) For the *asymptotic* case

$$\frac{C_{sp}}{m_v} = \left( \frac{1}{2} \frac{C_{fr}}{m_v} \right)^2 \quad (8.232)$$

the eigenvalues are equal to each other

$$\lambda_{1,2} = \lambda = -\frac{1}{\Delta\tau_d} . \quad (8.233)$$

In this case the text book solution is

$$z(\tau) = \frac{z_{\max}^{**}}{2} + C_1 e^{\lambda \tau} + C_2 \tau e^{\lambda \tau} . \quad (8.234)$$

### 8.8.4 Fitting the piecewise solution on two known position – time points

a) *Harmonic oscillations case:*

Knowing two arbitrary  $(\tau, z)$  points

$$\tau = \tau_1, \quad z = z_1, \quad (8.235)$$

and

$$\tau = \tau_2, \quad z = z_2, \quad (8.236)$$

we compute the constants  $C_1$  and  $C_2$ . Replacing the denominator of the so obtained solution with its equivalent

$$\cos(\omega\tau_1)\sin(\omega\tau_2) - \cos(\omega\tau_2)\sin(\omega\tau_1) = \sin[\omega(\tau_2 - \tau_1)], \quad (8.237)$$

and substituting

$$z_1^* = (z_1 - z_{\max}^{**} / 2) e^{\tau_1 / \Delta\tau_d}, \quad (8.238)$$

and

$$z_2^* = (z_2 - z_{\max}^{**} / 2) e^{\tau_2 / \Delta\tau_d}, \quad (8.239)$$

we obtain finally

$$C_1 = \frac{z_1^* \sin(\omega\tau_2) - z_2^* \sin(\omega\tau_1)}{\sin[\omega(\tau_2 - \tau_1)]}, \quad (8.240)$$

$$C_2 = \frac{z_2^* \cos(\omega\tau_1) - z_1^* \cos(\omega\tau_2)}{\sin[\omega(\tau_2 - \tau_1)]}, \quad (8.241)$$

or after inserting into Eq. (8.225) and some rearrangements we obtain

$$z(\tau) = \frac{z_{\max}^{**}}{2} + \frac{z_1^{**} \sin[\omega(\tau_2 - \tau)] - z_2^{**} \sin[\omega(\tau_1 - \tau)]}{\sin[\omega(\tau_2 - \tau_1)]}, \quad (8.242)$$

where

$$z_1^{**} = (z_1 - z_{\max}^{**} / 2) e^{(\tau_1 - \tau) / \Delta\tau_d}, \quad (8.243)$$

and

$$z_2^{**} = (z_2 - z_{\max}^{**} / 2) e^{(\tau_2 - \tau) / \Delta\tau_d}. \quad (8.244)$$

b) *Aperiodic case*

Again knowing the solutions in two previous points we can estimate the constants. Inserting into Eq. (8.231) and rearranging we obtain finally

$$z(\tau) = \frac{z_{\max}^{**}}{2} + \frac{z_1^{**} \sinh[\omega(\tau_2 - \tau)] - z_2^{**} \sinh[\omega(\tau_1 - \tau)]}{\sinh[\omega(\tau_2 - \tau_1)]}. \quad (8.245)$$

Note the formal similarity to Eq. (8.242). The sine functions are here replaced by hyperbolic sinus.

c) *Asymptotic case*

Again knowing the solutions in two previous points we can estimate the constants. Inserting into Eq. (8.234) and rearranging we obtain finally

$$z(\tau) = \frac{z_{\max}^{**}}{2} + \frac{z_1^{**}(\tau_2 - \tau) - z_2^{**}(\tau_1 - \tau)}{\tau_2 - \tau_1}. \quad (8.246)$$

Note that Eqs. (8.242) and (8.245) reduces to the above equation for  $\Delta\tau_v \rightarrow \infty$ .

### 8.8.5 Fitting the piecewise solution on known velocity and position for a given time

For a description of reflection of the piston after the impact with the upper or lower lift limitation structures it is more appropriate to use a solution fitted to a single time point position and velocity  $\tau = \tau_1$ ,  $z = z_1$ ,  $\frac{dz}{d\tau} = w_{v,1}$ .

a) For the *harmonic oscillations* case we have for the integration constants

$$z(\tau) = e^{-\tau/\Delta\tau_d} \left[ C_1 \cos(\omega\tau) + C_2 \sin(\omega\tau) \right] + z_{\max}^{**}/2, \quad (8.247)$$

$$w_v = -\frac{1}{\Delta\tau_d} (z - z_{\max}^{**}/2) + \omega e^{-\tau/\Delta\tau_d} \left[ -C_1 \sin(\omega\tau) + C_2 \cos(\omega\tau) \right], \quad (8.248)$$

$$C_1 = e^{\tau_1/\Delta\tau_d} \left\{ (z_1 - z_{\max}^{**}/2) \left[ \cos(\omega\tau_1) - \frac{\sin(\omega\tau_1)}{\omega\Delta\tau_d} \right] - \frac{w_{v,1}}{\omega} \sin(\omega\tau_1) \right\}, \quad (8.249)$$

$$C_2 = e^{\tau_1/\Delta\tau_d} \left\{ (z_1 - z_{\max}^{**}/2) \left[ \sin(\omega\tau_1) + \frac{\cos(\omega\tau_1)}{\omega\Delta\tau_d} \right] + \frac{w_{v,1}}{\omega} \cos(\omega\tau_1) \right\}. \quad (8.250)$$

b) For the *aperiodic* case

$$z = C_1 e^{\lambda_1 \tau} + C_2 e^{\lambda_2 \tau} + z_{\max}^{**}/2, \quad (8.251)$$

$$w_v = C_1 \lambda_1 e^{\lambda_1 \tau} + C_2 \lambda_2 e^{\lambda_2 \tau}, \quad (8.252)$$

$$C_1 = \frac{-(z_1 - z_{\max}^{**}/2) \lambda_2 e^{\lambda_2 \tau_1} + w_{v,1} e^{\lambda_2 \tau_1}}{\lambda_1 e^{(\lambda_1 + \lambda_2) \tau_1} - \lambda_2 e^{(\lambda_1 + \lambda_2) \tau_1}}, \quad (8.253)$$

$$C_2 = \frac{(z_1 - z_{\max}^{**}/2) \lambda_1 e^{\lambda_1 \tau_1} - w_{v,1} e^{\lambda_1 \tau_1}}{\lambda_1 e^{(\lambda_1 + \lambda_2) \tau_1} - \lambda_2 e^{(\lambda_1 + \lambda_2) \tau_1}}. \quad (8.254)$$

c) For the *asymptotic* case

$$z(\tau) = \frac{z_{\max}^{**}}{2} + C_1 e^{\lambda\tau} + C_2 \tau e^{\lambda\tau}, \quad (8.255)$$

$$w_v = C_1 \lambda e^{\lambda\tau} + C_2 (1 + \lambda\tau) e^{\lambda\tau}, \quad (8.256)$$

$$C_1 = \left[ (1 + \lambda\tau_1) (z_1 - z_{\max}^{**}/2) - w_v \tau_1 \right] e^{-\lambda\tau_1}, \quad (8.257)$$

$$C_2 = \left[ w_v - (z_1 - z_{\max}^{**}/2) \lambda \right] e^{-\lambda\tau_1}. \quad (8.258)$$

### 8.8.6 Idealized valve characteristics

In order to understand some characterizing features of the valve dynamic we consider next some interesting simple cases.

a) No friction force

If there is no friction force,

$$C_{fr} = 0 \quad (8.259)$$

$$\Delta\tau_d \rightarrow \infty, \quad (8.260)$$

the solution is

$$z(\tau) = \frac{z_{\max}^{**}}{2} + C_1 \cos(\omega\tau) + C_2 \sin(\omega\tau), \quad (8.261)$$

which means that the oscillations are not damped.

Some important features of the dynamic behavior of the valve can be studied on the non-damped solution.

b) Opening just starts, idealized opening time:

Let us consider the opening process at the very beginning. For this case we have

$$\tau = 0, \quad (8.262)$$

$$z = 0, \quad (8.263)$$

and

$$dz/d\tau = 0, \quad (8.264)$$

and consequently

$$C_{fr} dz/d\tau = 0, \quad (8.265)$$

the friction force is small due to the small averaged velocity gradient at the moment of the opening. Further it can be assumed that

$$f_{\beta} = 0, \quad (8.266)$$

i. e. there is no flow reaction force.

$$z(\tau) = \frac{z_{\max}^*}{2} [1 - \cos(\omega\tau)] \geq 0, \quad (8.267)$$

where

$$z_{\max}^* = \frac{2A_m}{C_{sp}} [p_{in}(\tau) - p_{on}(\tau)] \quad (8.268)$$

is the maximum of the piston position which can be reached if there where no hardware limitation of the piston motion. Obviously  $z_{\max}^*$  depends an  $p_{in}(\tau)$ . If

$$z_{\max} < z_{\max}^* \quad (8.269)$$

we have stable opening process. The time needed for opening the valve is

$$\Delta\tau_{op} = \omega \arccos\left(1 - 2\frac{z_{\max}}{z_{\max}^*}\right), \quad (8.270)$$

where

$$-1 < 1 - 2\frac{z_{\max}}{z_{\max}^*} \leq 1 \quad \text{or} \quad 0 < z_{\max} < z_{\max}^*. \quad (8.271)$$

If

$$z_{\max} \geq z_{\max}^* \quad (8.272)$$

the non-damping harmonic oscillation of the piston position within

$$0 < z < z_{\max}^* \quad (8.273)$$

is expected. This operation regime of the valve is called *piston fluttering*.

c) Closing just starts, idealized closing time:

Now let as consider the case where the valve was completely open and starts to close, i.e.

$$\tau = 0, \quad (8.274)$$

$$z = z_{\max}, \quad (8.275)$$

$$dz/d\tau = 0, \quad (8.276)$$

and consequently

$$C_{fr} dz/d\tau = 0, \quad (8.277)$$

the solution is

$$z(\tau) = \frac{z_{\max}^*}{2} [1 - \cos(\omega\tau)] + z_{\max} \cos(\omega\tau) < z_{\max} . \quad (8.278)$$

The condition  $z < z_{\max}$  leads to

$$z_{\max}^{**} < 2z_{\max} , \quad (8.279)$$

which is the necessary condition that the open valve starts to close. Obviously

$$\left| z_{\max}^{**} \right| > \left| z_{\max}^* \right| \quad \text{for the some} \quad \left| p_{in}(\tau) - p_{on} \right| \quad (8.280)$$

due to the existence of the flow reaction thrust

$$f_{fl} > 0. \quad (8.281)$$

Consequently the closing time

$$\Delta\tau_{cl} = \omega \arccos \left( 1 - 2 \frac{z_{\max}}{z_{\max}^{**}} \right)^{-1} \quad (8.282)$$

is larger than the closing time for the same value of  $\left| p_{in}(\tau) - p_{on} \right|$ . For  $z_{\max}^{**} = 0$ ,  $\Delta\tau_{cl} = \Delta\tau_v / 4$ . For  $z_{\max}^{**} > 0$ ,  $\Delta\tau_{cl} > \Delta\tau_v / 4$  and for  $z_{\max}^{**} < 0$ ,  $\Delta\tau_{cl} < \Delta\tau_v / 4$ . We see that in the case of open valve even for pressure difference  $\left| p_{in}(\tau) - p_{on} \right| \approx 0$  the valve remain open at least  $\Delta\tau_v / 4$  due to the action of the flow reaction force.

### 8.8.7 Recommendations for the application of the model in system computer codes

Finally the following algorithm can be recommended for modeling the dynamic valve behavior. The algorithm consists of logical conditions needed to identify the valve regime and the appropriate solution of the valve dynamic equation.

1. The valve is closed and remains closed if

$$(z = 0 \text{ and } p_{in} \leq p_{on}). \quad (8.283)$$

2. The valve starts to open if

$$(z = 0 \text{ and } p_{in} > p_{on}). \quad (8.284)$$

For

$$\tau = \tau + \Delta\tau , \quad (8.285)$$

the piston is lifted to

$$z(\tau) = \frac{z_{\max}^*}{2} [1 - \cos(\omega \Delta \tau)]. \quad (8.286)$$

In this case if  $z < \varepsilon$  set  $z = \varepsilon$ , where

$$\varepsilon \rightarrow 0, \quad (8.287)$$

is the “computer zero”. This is very important for the practical application of the method for all possible time steps.

3. The valve was completely open and remains completely open if

$$z = z_{\max} \quad \text{and} \quad z_{\max}^{**} \geq 2z_{\max}. \quad (8.288)$$

4. The valve was completely open and starts to close

$$z = z_{\max} \quad \text{and} \quad z_{\max}^{**} < 2z_{\max}. \quad (8.289)$$

In this case after  $\Delta \tau$  the plug position is removed to the position

$$z(\tau) = \frac{z_{\max}^*}{2} [1 - \cos(\omega \Delta \tau)] + z_{\max} \cos(\omega \Delta \tau) < z_{\max} \quad (8.290)$$

In this case if  $z > z_{\max} - \varepsilon$  set  $z = z_{\max} - \varepsilon$ . This is very important for the practical application of the method for all possible time steps.

5. The valve is in operation if

$$\varepsilon < z = z_{\max} - \varepsilon. \quad (8.291)$$

In this case the preceding information for two points is used

$$\tau = \tau_1, \quad z = z_1, \quad (8.292)$$

and

$$\tau = \tau_2, \quad z = z_2, \quad (8.293)$$

to compute the constants of the piecewise analytical solution  $C_1$  and  $C_2$ . The plug

position at  $\tau$  for  $\frac{C_{sp}}{m_v} > \left(\frac{1}{2} \frac{C_{fr}}{m_v}\right)^2$  is

$$z(\tau) = \frac{z_{\max}^{**}}{2} + \frac{z_1^{**} \sin[\omega(\tau_2 - \tau)] - z_2^{**} \sin[\omega(\tau_1 - \tau)]}{\sin[\omega(\tau_2 - \tau_1)]} \quad (8.294)$$

for  $\frac{C_{sp}}{m_v} < \left(\frac{1}{2} \frac{C_{fr}}{m_v}\right)^2$ ,

$$z(\tau) = \frac{z_{\max}^{**}}{2} + \frac{z_1^{**} \sinh[\omega(\tau_2 - \tau)] - z_2^{**} \sinh[\omega(\tau_1 - \tau)]}{\sinh[\omega(\tau_2 - \tau_1)]} \quad (8.295)$$

and for  $\frac{C_{sp}}{m_v} = \left(\frac{1}{2} \frac{C_{fr}}{m_v}\right)^2$ ,

$$z(\tau) = \frac{z_{\max}^{**}}{2} + \frac{z_1^{**} (\tau_2 - \tau) - z_2^{**} (\tau_1 - \tau)}{\tau_2 - \tau_1}, \quad (8.296)$$

where if

$$z < 0, \quad (8.297)$$

$z$  is set to zero

$$z = 0, \quad (8.298)$$

and the valve is considered as completely closed, and if

$$z > z_{\max}, \quad (8.299)$$

$z$  is set to  $z_{\max}$ ,

$$z = z_{\max}, \quad (8.300)$$

and the valve is considered completely open. Note that if  $\varphi < 0.00001$   $\sin \varphi \approx \varphi$  and  $\sinh \varphi \approx \varphi$ .

Thus the mathematical description of the plug motion is completed. Having the prescribed dependence of the flow surface at the vena contracta as a function of the plug position

$$A^* = A^*(z) \quad (8.301)$$

we can estimate in any moment the surface available for the flow and consequently the flow reaction force  $f_{fr}$ .

### 8.8.8 Some illustrations of the valve performance model

This chapter is devoted to an example of the analytical work required to be done before implementing a single-component model into system computer codes. The system computer codes are very complex. The behavior of any single pipe network component has to be examined very carefully outside the code before coupling the model with the system code. The valve model is a typical example. It introduces such a non-linear interaction with the flows in pipe systems that without knowing exactly the valve model behavior it is very difficult to make error diagnostics during the development. Here we present some examples documented in *Kolev* (1993) and *Roloff-Bock* and *Kolev* (1998). The tests have been performed by *Roloff-Bock* and *Kolev* (1998).



We will have a look on the valve response to various forms of the driving force function  $F(\vartheta)$ . The fluid momentum flux force  $F_f$  is set to zero, therefore the driving force is directly proportional to the inlet pressure. For this purpose simply the inlet pressure  $p_m$  is provided as a time function, which will be either a harmonic oscillation or a step function.

1) Figure 8.9 shows the *response to a slowly varying driving function* with several values of the friction coefficient  $C_{fr}$ . The valve was closed at the beginning. When the inlet pressure exceeds the onset pressure  $p_{on}$ , the valve starts to open as mentioned in the previous chapter.

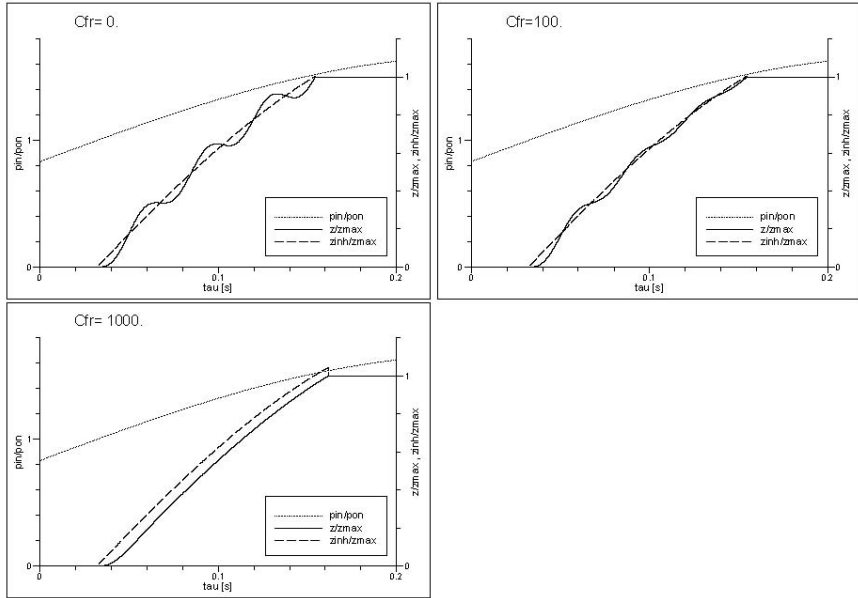
If there is no friction, i. e.  $C_{fr} = 0$ , the valve piston oscillates around the equilibrium position  $z_{inh}(\vartheta)$ , which follows the driving function, until the upper boundary at  $z_{max}$  is reached. No reflection is taken into account here, so the valve stays open.

If the friction coefficient is small with respect to the effect of the spring, the oscillations around the equilibrium position are damped.

If the friction coefficient is large, the valve piston follows the equilibrium position without oscillating, but with a delay.

2) Figure 8.10 shows the *response to a harmonic oscillation function*. The frequency of the driving function is of the same order of magnitude as the eigenfrequency of the free oscillating valve, which is in this case  $f_v = 28 \text{ Hz}$ . The valve is closed at the beginning. It starts to open, when the onset pressure  $p_{on}$  is exceeded by the inlet pressure  $p_m$ . Reflection is not taken into account here. Therefore the valve stays open, until the inlet pressure drops according to case 4 discussed in the previous Section.

When the lower boundary at  $z = 0$  is reached, the valve remains closed, until the inlet pressure again exceeds the onset pressure. Also in this case a delay due to the friction can be clearly seen.



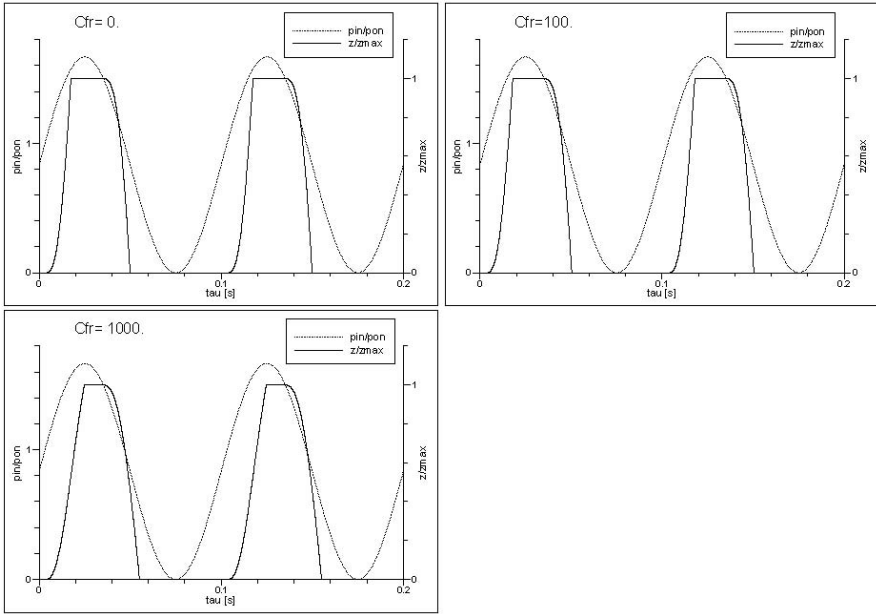
**Fig. 8.9** Valve response to a slowly varying driving function with frequency  $f_v = 1 \text{ Hz}$

3) Figure 8.11 shows the dependence of the valve response on the frequency of the driving function. No friction is taken into account this time. Reflection is also not considered here.

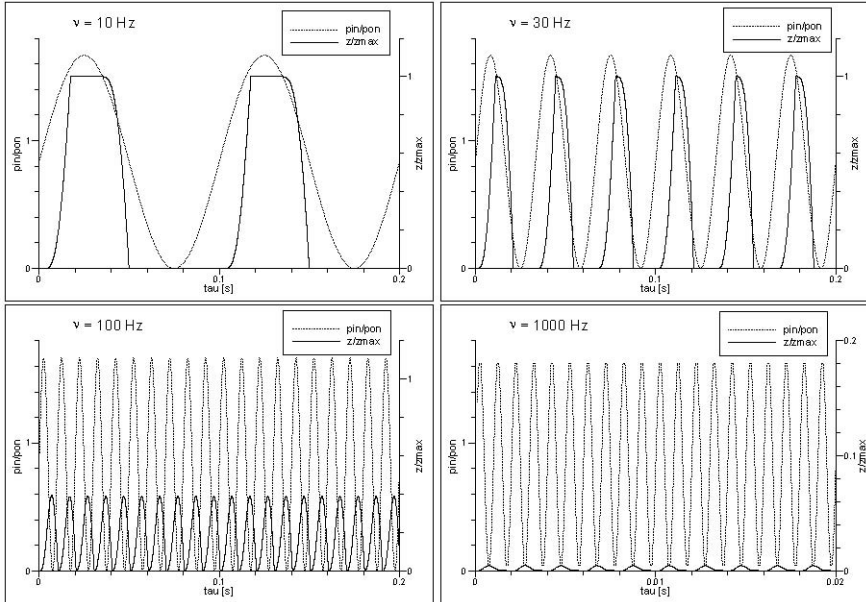
If the frequency  $\omega$  of the driving function is much larger than the eigen-frequency of the free oscillating valve, which in this example is  $f_v = 28 \text{ Hz}$ , then the amplitude of the valve oscillation drops according to the ratio of the two frequencies.

Figure 8.12 shows the valve response to a step function. The height of the step is varied. The previously closed valve starts oscillating around the new equilibrium position, which corresponds to the height of the step, as long as the value of the inlet pressure does not satisfy the conditions of case 3 discussed in the previous Section to hold up the valve in a completely open stage.

4) Figure 8.13 and 8.14 show the effect of the friction coefficient  $C_{fr}$  and the transition from the oscillation case to the aperiodic solution case and the asymptotic solution case. If the friction coefficient is small, the valve does some damped oscillations around the new equilibrium position. When the friction coefficient takes a value that counteracts the spring effect, the new equilibrium position is reached within one period and the oscillations are stopped. As the friction coefficient increases, the valve needs more and more time to reach the new equilibrium position.



**Fig. 8.10** Valve response to a harmonic oscillation with frequency  $\nu = 10 \text{ Hz}$



**Fig. 8.11** Dependence on the frequency  $\nu$  of the driving function

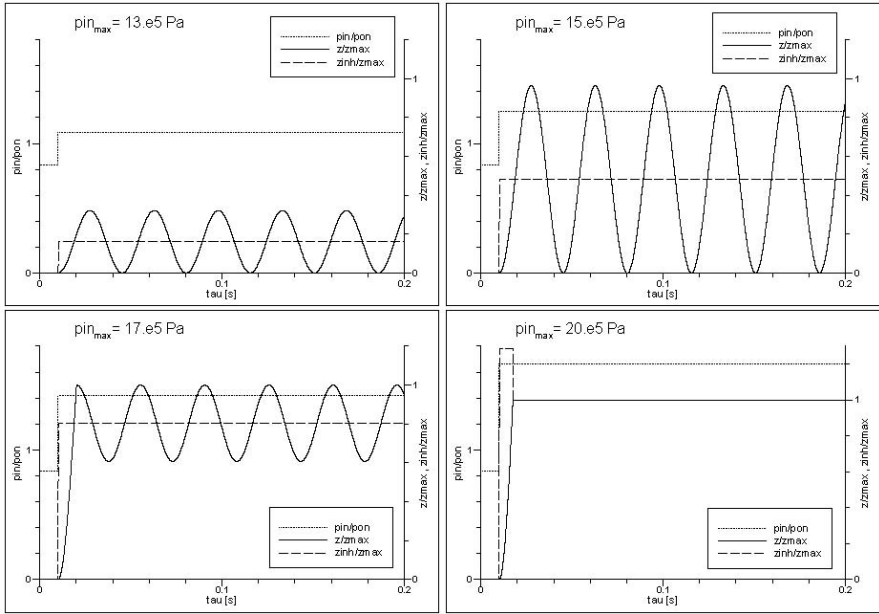


Fig. 8.12 Valve response to a step function

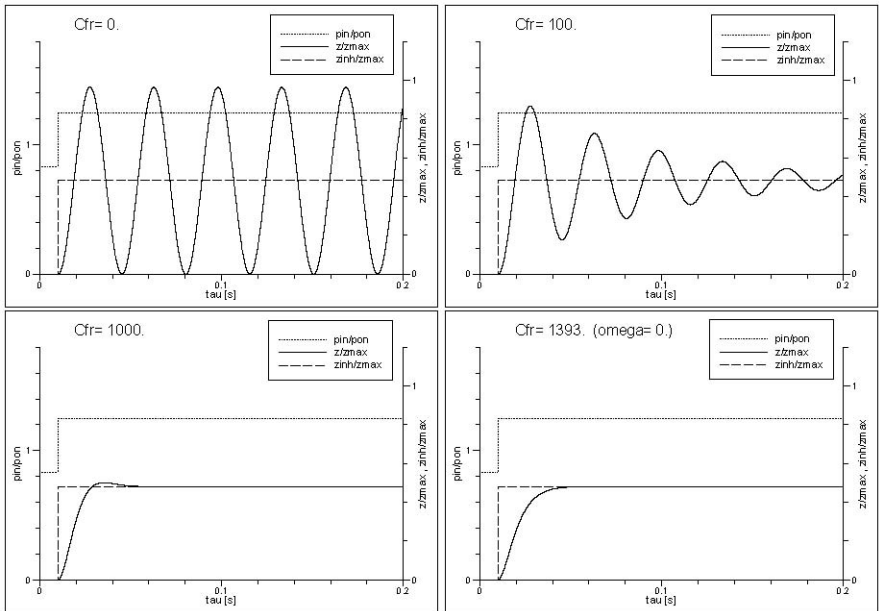


Fig. 8.13 Influence of the friction coefficient, oscillating and aperiodic case

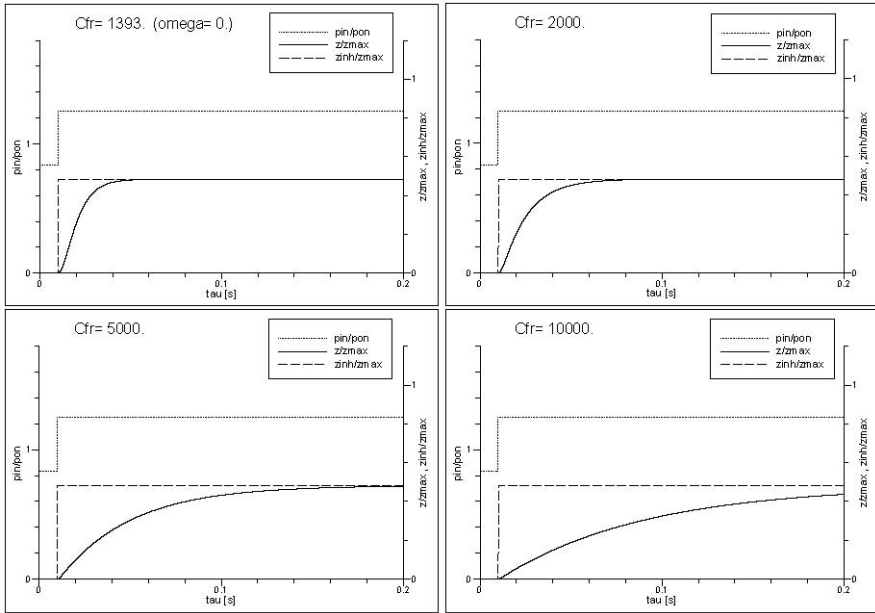


Fig. 8.14 Influence of the friction coefficient, asymptotic case

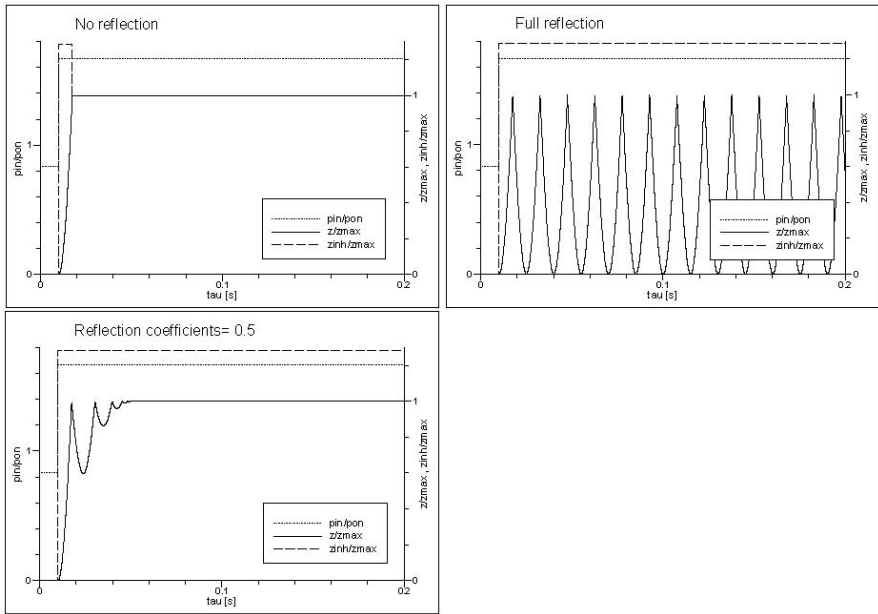
5) Figure 8.15 shows the effect of the reflection coefficients. The reflection coefficients for both boundaries are varied simultaneously. As an example the response to a large step satisfying the conditions of case 3 discussed in the previous chapter to hold up the valve in a completely open stage is considered.

With no reflection at the upper boundary, i. e.  $r_u = 0$ , the piston sticks to the boundary once it is reached.

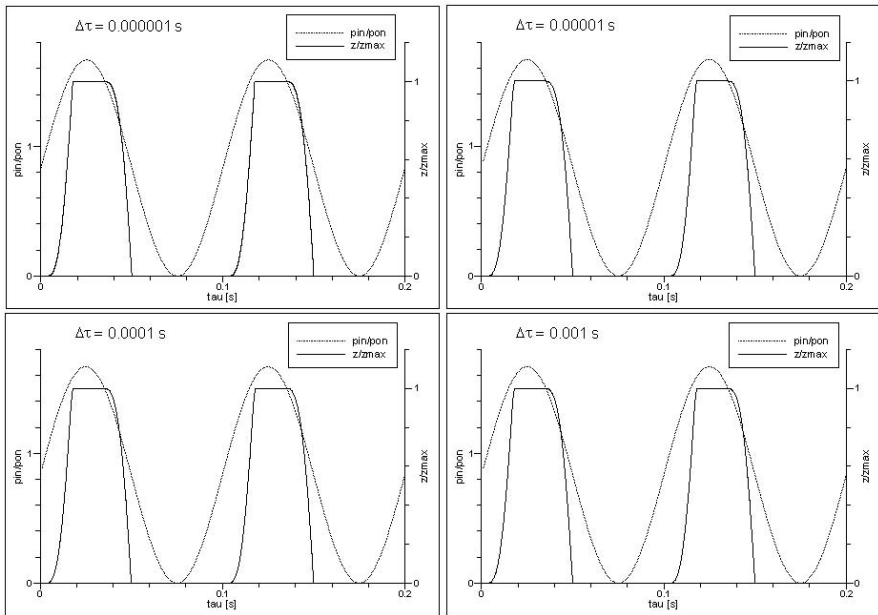
A reflection coefficient of  $r_u = 1$  means full reflection. In this case the valve piston retains all its velocity, but the direction of the movement is inverted. Therefore the piston goes back exactly in reverse manner.

With a reflection coefficient value of  $r_u = 0.5$ , half of the velocity amount is preserved. That means that every reversion goes only half the way back.

6) Figure 8.16 shows the dependence of the solution on the time step  $\Delta t$ . As can be seen clearly, the solution method gives good results also for very large time steps compared to a finite difference method approach.



**Fig. 8.15** Influence of the reflection coefficients



**Fig. 8.16** Dependence of the valve equation solution on the time step  $\Delta\tau$

### 8.8.9 Nomenclature for Section 8.8

#### Latin

$A^*(z)$	flow cross section or critical flow cross section as a function of the piston position, $m^2$ . This function is defined by the particular valve geometry.
$A^*$	vena contracta area i.e. the valve ring area $m^2$
$A_{norm}$	some normalizing cross section, $m^2$
$A_m$	$> 0$ , valve piston face area exposed to the inlet flow stream, $m^2$
$A_{out}$	$\geq 0$ , valve piston face ring area exposed to the outlet pressure by open and closed valve, $m^2$
$A_{outu}$	back valve piston area outside the bellow, $m^2$
$A_{amu}$	back valve piston area inside the bellow, $m^2$
$f_{fl}$	fluid force acting on the piston for open valve, $N$
$f_v$	eigen frequency of the piston, $1/s$
$C_{sp}$	spring constant, $N/m$
$C_{ft}$	valve damping coefficient, $Ns/m$
$m_v$	mass of the valve mechanism that is in motion (i.g. the valve piston, rod assembly combined with the spring and bellows), $kg$
$p_{in}(\tau)$	valve inlet pressure, $Pa$
$p_{out}(\tau)$	valve outlet pressure, $Pa$
$p_{on}$	pressure at which previously closed valve starts to open, $Pa$
$p_{am}$	atmospheric back pressure inside the bellow, $Pa$
$p_{outu}$	valve back pressure outside the bellow, $Pa$
$r_u$	reflection coefficient, <i>dimensionless</i>
$w_v$	velocity of the valve piston, $m/s$
$z$	piston position (i.e., $z$ coordinate, $0 \leq z \leq z_{max}$ ), $m$ $= 0$ , piston position for closed valve, $m$ $= z_{max}$ , piston position for completely open valve, valve lift, $m$
$z_0$	spring pressing distance for the normal valve operation, $m$
$z_r$	position of the lower end of the valve adjusting ring assembly, $0 \leq z_r \leq z_{max}$ , $m$

#### Greek

$\alpha_l$	volume fraction of the field $l$ , <i>dimensionless</i>
$\gamma_z$	$= A^*(z)/A_{norm}$ , surface permeability at the contraction cross section used in the network analysis, <i>dimensionless</i>

$\Delta r$	width of the expansion ring area formed by the inner surface of the valve adjusting ring assembly and the outer surface of the inlet, $m$
$\Delta \tau_{cl}$	closing time, $s$
$\Delta \tau_v$	oscillation period, $s$
$\Delta \tau_d$	damping time constant, $s$
$\lambda_{1,2}$	eigenvalues, $1/s$
$\rho_l$	density of field $l$ , $kg/m^3$
$\varphi$	$\leq \pi/2$ , angle of the flow path trajectory between valve inlet and valve outlet, $rad$
$\tau$	time, $s$
$\omega$	oscillation frequency, $2\pi/\Delta \tau$

**Subscripts**

1	field 1, gas
2	field 2, liquid
3	field 3, droplets
$l$	field $l$
$in$	valve inlet
$out$	valve outlet
$atm$	at atmospheric pressure
$norm$	scaling
$outu$	outside the bellow
$sp$	spring
$fr$	friction
$on$	previously closed valve starts to open
$r$	ring

**Superscripts**

\* at the smallest flow cross section called in Latin vena contracta

**8.9 Pump model**

In the previous section we considered the relief valves as an important component of pipe networks. Another important component is the pump. In this Section we describe the dynamic behavior of the *centrifugal pumps* by simple model. The model is appropriate for use in computer codes simulating system behavior of complex pipe networks e.g. *Kolev (1993)*.

The derived pump model is based on the following simplifying assumptions:

(a) Transient flow processes into the pump impeller can be represented by a sequence of steady state processes because the time needed by the flow particle to pass the pump is very short. The smaller the pumps dimension the more correct is this assumption and vice versa.



(b) The flow through the pump is incompressible.

### 8.9.1 Variables defining the pump behavior

The variables required to describe mathematically the pump behavior are defined below.

Geometry:

$D_{p1}$	inner pump impeller diameter, $m$
$D_{p2}$	outer pump impeller diameter, $m$
$\beta_1$	angle between the relative flow velocity $V_1^r$ and the impeller angular velocity $V_{\theta 1}$ , $rad$
$\beta_2$	angle between the relative flow velocity $V_2^r$ and the impeller angular velocity $V_{\theta 2}$ , $rad$
$I_{pr}$	moment of inertia of the pump rotor, $kgm^2$

Velocities:

$V_r$	radial (meridian) velocity, $m/s$
$V^r$	relative flow velocity for an observer rotating with the impeller, $m/s$
$V_1^r$	$= \dot{V} / A_1$ , relative flow velocity for an observer rotating with the impeller at the entrance cross section of the impeller, $m/s$
$V_2^r$	$= \dot{V} / A_2$ , relative flow velocity for an observer rotating with the impeller at the exit cross section of the impeller, $m/s$
$V_\theta$	angular impeller velocity, $m/s$
$V_{\theta i}$	$= \omega D_{pi} / 2 = \pi D_{pi} n$ , impeller angular velocity at the position defined with diameter $D_{pi}$ , $m/s$

Fluid characteristics:

$\rho$	fluid density, $kg/m^3$
$V_m$	$= \left( \sum_{l=1}^{l_{max}} \alpha_l \rho_l V_l \right) / \sum_{l=1}^{l_{max}} \alpha_l \rho_l$ , fixture velocity, $m/s$
$A$	flow cross section at which the mixture velocity, $m^2$

Pump characteristics:

$g$	$= 9.81$ , acceleration due to gravity, $m/s^2$
$\dot{V}$	volumetric through flow, $m^3/s$ . The volumetric flow is positive if it is in the same direction as the positive velocity in the control volume i.e. if it has the direction of the increasing cell indices.

$\dot{V}_R$	rated volumetric through flow characterizing the normal pump operation, $m^3/s$
$\dot{V}^*$	$= \dot{V} / \dot{V}_R$ , volumetric flow ratio, <i>dimensionless</i>
$\Delta p_{pump}$	pump pressure difference i.e. pressure at the pump outlet minus pressure at the pump inlet (used in the momentum balance), $Pa$ . The pressure difference is positive if it would accelerate the flow in the positive velocity direction.
$H$	$= \Delta p_{pump} / (\rho g)$ , total head rise of the pump (defined by empirical homologous pump performance model as a function of the volumetric through flow $\dot{V}$ ), $m$
$H_R$	rated total head rise of the pump characterizing the normal pump operation, $m$
$H^*$	$= H / H_R$ , head ratio, <i>dimensionless</i>
$M_{fl}$	torque acting at the pump rotor and caused by the flow in the pump (defined by empirical homologous pump performance model as a function of the volumetric through flow $\dot{V}$ ), $Nm$ . Negative, if it tends to decelerate the pump.
$M_{fl,R}$	rated pump torque characterizing the normal pump operation, $Nm$
$M^*$	$= M_{fl} / M_{fl,R}$ , torque ratio, <i>dimensionless</i>
$M_{fr}$	friction torque, $Nm$
$M_m$	motor torque, $Nm$
$\omega$	pump angular speed (rotational speed defined by a pump drive model). A pump operating in a normal pump regime has positive angular velocity, $rad/s$
$\omega_R$	rated pump angular speed characterizing the normal pump operation, $rad/s$
$\omega^*$	$\omega / \omega_R$ , angular velocity ratio, <i>dimensionless</i>
$n$	rotation per second, $1/s$
$n_R$	rated rotation per second characterizing the normal pump operation, $1/s$
$n^*$	$= n / n_R$ , rotation ratio, <i>dimensionless</i>
$P_{pump}$	$= M \omega = \eta_p \Delta p_{pump} A V_m = \eta_p \Delta p_{pump} \dot{V}$ , pump power introduced into the flow (used into the energy balance), $W$
$\eta_p$	efficiency, <i>dimensionless</i>

The rated values  $\dot{V}_R$ ,  $\omega_R$ ,  $H_R$  and  $M_{fl,R}$  are required input data for the pump model.

Dimensionless similarity criteria:

$$\varphi = \frac{\dot{V}}{V_\theta \pi D_{pi}^2 / 4} = \frac{8\dot{V}}{\pi \omega D_{pi}^3} = \frac{4\dot{V}}{\pi^2 n D_{pi}^3}, \text{ specific capacity, Keller 1934, } \textit{dimensionless}$$

$$\psi = \frac{\Delta p_{\text{pump}}}{\rho V_{\theta i}^2 / 2} = \frac{\rho g H}{\rho V_{\theta i}^2 / 2} = \frac{g H}{V_{\theta i}^2 / 2} = \frac{8 g H}{(\omega D_{pi})^2} = \frac{2 g H}{(\pi n D_{pi})^2}, \text{ specific head, } L. \text{ Prandtl 1912, dimensionless. } \Delta p_{\text{pump}} / \rho \text{ can be considered as a specific work done on the fluid and } V_{\theta i}^2 / 2 \text{ the specific kinetic energy at the outer impeller diameter.}$$

$$\lambda = \phi \psi / \eta_{\text{eff}}, \text{ power number, dimensionless}$$

$$\eta_{\text{eff}} = \text{ratio of the power inserted into the flow to the pump motor power, dimensionless}$$

$$M / (\rho \omega^2 D_{\rho 2}^5) = \text{specific torque, dimensionless}$$

In order to obtain bounded values in the homologous pump curve the following modified similarity numbers are used too:

$$g H / \dot{V}^2 = \text{modified specific head, dimensionless}$$

$$M / (\rho \dot{V}^2 D_{\rho 2}^3) = \text{modified specific torque, dimensionless}$$

$$n_s = \text{specific speed, } 1/s$$

The *specific speed* is the speed necessary for a geometrically similar model pump to provide a head of  $H = 1 \text{ m}$  and a volume flow rate of  $\dot{V} = 1 \text{ m}^3/s$ . Using the similarity relationships and the steady state operation point ( $H_R, \dot{V}_R$  and  $n_R$ ) the specific speed is computed with Eq. 8.5 in Kolve (1982) p. 251 as follows

$$n_s = n_R \left( \frac{\dot{V}_R}{1 \text{ m}^3/s} \right)^{1/2} / \left( \frac{H_R}{1 \text{ m}} \right)^{3/4}. \quad (8.302)$$

The dimension of the specific speed is the same as the dimension of the rated speed  $n_R$ . One should be careful with the definition of the specific speed in older sources because at least three metric systems were widely used in the past.

### 8.9.2 Theoretical basics

The text book theoretical basics, see Pohlentz (1975) p. 53, are briefly given here in order to understand the world wide accepted pump modeling technique. Consider a centrifugal pump with impeller rotating with  $n$  rotations per second that is with angular velocity

$$\omega = 2\pi n, \quad (8.303)$$

inner radius  $r_1$  and outer radius  $r_2$ . The angular velocity at the inner impeller radius is

$$V_{\theta 1} = \omega r_1 \quad (8.304)$$

and at the outer impeller radius

$$V_{\theta 2} = \omega r_2. \quad (8.305)$$

The relative flow velocity for an observer rotating with the impeller is  $V^r$ . The relative flow velocity is equal to the volumetric flow divided by the cross section normal to  $V^r$ . For incompressible flow we have

$$V_1^r = \dot{V} / A_1, \quad (8.306)$$

$$V_2^r = \dot{V} / A_2. \quad (8.307)$$

The vector sum of the relative flow velocity and the angular impeller velocity gives the absolute flow velocity  $\mathbf{V}^a$ .

$$\mathbf{V}^a = \mathbf{V}^r + \mathbf{V}_\theta^a. \quad (8.306)$$

We designate the angular component of this velocity magnitude with  $V_\theta^a$ .

The mass flow in  $kg/s$  entering the pump impeller is

$$\dot{m} = \rho \dot{V}. \quad (8.307)$$

The angle between the relative flow velocity  $V_1^r$  and the impeller angular velocity  $V_{\theta 1}$  is  $\beta_1$  and that the angle between the relative flow velocity  $V_2^r$  and the impeller angular velocity  $V_{\theta 2}$  is  $\beta_2$ . Therefore

$$V_{\theta 1}^a = V_{\theta 1} - V_1^r \cos \beta_1, \quad (8.308)$$

$$V_{\theta 2}^a = V_{\theta 2} - V_2^r \cos \beta_2. \quad (8.309)$$

The flow is stationary and incompressible. Thus the force acting on the flow in the angular direction at the inlet is

$$F_{\theta 1} = \dot{m} V_{\theta 1}^a \quad (8.310)$$

and at the impeller outlet

$$F_{\theta 2} = -\dot{m} V_{\theta 2}^a. \quad (8.311)$$

The corresponding torques are  $F_{\theta 1} r_1$  and  $F_{\theta 2} r_2$ , respectively. The resulting torque acting on the flow inside the impeller is

$$M_{\beta} = F_{\theta 2} r_2 + F_{\theta 1} r_1 = \dot{m} (V_{\theta 2}^a r_2 - V_{\theta 1}^a r_1) \quad (8.312)$$

The power inserted into the flow is therefore  $M_{\beta} \omega$  or

$$\Delta P_{\text{pump}} \dot{V} = M_{\beta} \omega = \omega \dot{m} (V_{\theta 2}^a r_2 - V_{\theta 1}^a r_1) = \rho \dot{V} (V_{\theta 2}^a \omega r_2 - V_{\theta 1}^a \omega r_1), \quad (8.313)$$

or after canceling the volumetric flux

$$\Delta p_{pump} = \rho (V_{\theta_2}^a \omega r_2 - V_{\theta_1}^a \omega r_1) = \rho (V_{\theta_2}^a V_{\theta_2} - V_{\theta_1}^a V_{\theta_1}). \quad (8.314)$$

This is the *main pump equation* describing the pressure rise as a function of the impeller speed and geometry. Replacing Eqs. (8.308) and (8.309) in the main pump equation we obtain

$$\begin{aligned} \Delta p_{pump} &= \rho \left[ (V_{\theta_2} - V_2^r \cos \beta_2) V_{\theta_2} - (V_{\theta_1} - V_1^r \cos \beta_1) V_{\theta_1} \right] \\ &= \rho \left( V_{\theta_2}^2 - V_{\theta_1}^2 + V_1^r V_{\theta_1} \cos \beta_1 - V_2^r V_{\theta_2} \cos \beta_2 \right). \end{aligned} \quad (8.315)$$

or after using the mass conservation equations (8.306) and (8.307) one obtains

$$\begin{aligned} \Delta p_{pump} &= \rho \left[ (r_2^2 - r_1^2) \omega^2 + \left( \frac{r_1}{A_1} \cos \beta_1 - \frac{r_2}{A_2} \cos \beta_2 \right) \omega \dot{V} \right] \\ &= \rho \left[ (r_2^2 - r_1^2) 4\pi^2 n^2 + \left( \frac{r_1}{A_1} \cos \beta_1 - \frac{r_2}{A_2} \cos \beta_2 \right) 2\pi n \dot{V} \right]. \end{aligned} \quad (8.316)$$

If we consider additionally that the flow exerts some pressure loss due to friction with the pump structures in the impeller region we finally obtain

$$\begin{aligned} \Delta p_{pump} &= \rho \left[ (r_2^2 - r_1^2) \omega^2 + \left( \frac{r_1}{A_1} \cos \beta_1 - \frac{r_2}{A_2} \cos \beta_2 \right) \omega \dot{V} \right] - \xi \frac{1}{2} \frac{\rho}{A_m} \dot{V}^2 \\ &= \rho \left[ (r_2^2 - r_1^2) 4\pi^2 n^2 + \left( \frac{r_1}{A_1} \cos \beta_1 - \frac{r_2}{A_2} \cos \beta_2 \right) 2\pi n \dot{V} \right] - \xi \frac{1}{2} \frac{\rho}{A_m} \dot{V}^2. \end{aligned} \quad (8.317)$$

where  $\xi$  is some friction loss coefficient and  $A_m$  some averaged flow cross section.

The pressure rise of the pump is usually expressed for liquids with the head  $H$

$$\Delta p_{pump} = \rho g H, \quad (8.318)$$

where  $g$  is the gravitational acceleration. Usually the main pump equation is written in the form

$$H = An^2 + Bn\dot{V} - C\dot{V}^2, \quad (8.319)$$

where

$$A = (r_2^2 - r_1^2) 4\pi^2 / g, \quad (8.320)$$

$$B = \left( \frac{r_1}{A_1} \cos \beta_1 - \frac{r_2}{A_2} \cos \beta_2 \right) 2\pi / g, \quad (8.321)$$

$$C = \xi \frac{1}{2} \frac{\rho}{A_m} / g. \quad (8.322)$$

The coefficients  $A$ ,  $B$  and  $C$  are not dependent on the fluid density and can be computed approximately from the particular pump geometry. The above equation contains the most important effect in the first quadrant i.e.  $\dot{V} > 0$  and  $H > 0$ . In fact such analysis does not take into account several effects. That is why in the practical applications the dependence  $H = H(n, \dot{V})$  is obtained experimentally for each particular pump family by the pump manufacturer. In the reality the ratio of the power inserted into the flow to the impeller power called efficiency is less than one

$$\eta_p = \rho g H \dot{V} / (M_{fl} \omega). \quad (8.323)$$

The power

$$W_{diss} = (1 - \eta_p) (M_{fl} \omega) \quad (8.324)$$

is transferred irreversibly from mechanical energy into internal energy of the flow per unit second due to internal friction, turbulization etc. Gao et al. (2011) reported efficiencies between 72 and 79% for axial pump usually used in nuclear power plants.

Obviously  $H$  is a function of  $n$  and  $\dot{V}$ . It is easy to show that introduction of a simple variable transformation reduces the three-dimensional relationship to a two-dimensional one. Dividing either by  $\dot{V}^2$  or by  $n^2$  one obtains

$$H / \dot{V}^2 = A (n / \dot{V})^2 + B n / \dot{V} - C = f(n / \dot{V}) \quad (8.325)$$

or

$$H / n^2 = A + B \dot{V} / n - C (\dot{V} / n)^2 = f(\dot{V} / n). \quad (8.326)$$

This is a remarkable quality of the main pump equation. In spite of the fact that this dependence is only an approximate one, the practical measurements

$$f(H, \dot{V}, n) = 0 \quad (8.327)$$

are surprisingly reduced to a single curve named the *similarity* or *homologous* curve

$$f_1(H / \dot{V}^2, n / \dot{V}) = 0 \quad (8.328)$$

or

$$f_2(H/n^2, \dot{V}/n) = 0. \quad (8.329)$$

Both above discussed forms are used in the literature in order to avoid division by zero in the four quadrant if  $n$  or  $\dot{V}$  tends to zero. This fact was recognized several decades ago. It was *Prandtl* in 1912, see *Prandtl* (1984), who introduced the so called specific head,

$$\psi = \frac{\Delta p_{pump}}{\rho V_{\theta 2}^2 / 2} = \frac{\rho g H}{\rho V_{\theta 2}^2 / 2} = \frac{g H}{V_{\theta 2}^2 / 2} = \frac{8 g H}{(\omega D_{p2})^2} = \frac{2 g H}{(\pi n D_{p2})^2}. \quad (8.330)$$

Here  $\Delta p_{pump} / \rho$  can be considered as a specific work done on the fluid and  $V_{\theta 2}^2 / 2$  the specific kinetic energy at the outer impeller diameter. It was *Keller* in 1934 who introduced the so called specific capacity of the pump defined as follows

$$\varphi = \frac{\dot{V}}{V_{\theta 2} \pi D_{p2}^2 / 4} = \frac{8 \dot{V}}{\pi \omega D_{p2}^3} = \frac{4 \dot{V}}{\pi^2 n D_{p2}^3}. \quad (8.331)$$

Like the measured dependence  $f(H, \dot{V}, n) = 0$  the presentation in the form

$$f_1(\varphi, \psi) = 0. \quad (8.332)$$

reduces with remarkable accuracy the data to a single curve. The very important practical significance of the above relationship is that keeping  $\varphi$  constant in two neighboring steady state points of operation by

$$\dot{V}_1 / n_1 = \dot{V}_2 / n_2 \quad (8.333)$$

$\psi$  is also constant, which means

$$H_1 / n_1^2 = H_2 / n_2^2. \quad (8.334)$$

Recently the main pump relationship was further modified by introducing dimensionless quantities namely quantities divided by the corresponding quantities characterizing the normal pump operation called some times rated quantities,  $H_R, n_R, \dot{V}_R$ ,

$$H^* = H / H_R, \quad (8.335)$$

$$n^* = n / n_R, \quad (8.336)$$

$$\dot{V}^* = \dot{V} / \dot{V}_R, \quad (8.337)$$

$$H^* / \dot{V}^{*2} = A^* (n^* / \dot{V}^*)^2 + B^* n^* / \dot{V}^* - C^* = f(n^* / \dot{V}^*), \quad (8.338)$$

or

$$H^*/n^{*2} = A^* + B^*\dot{V}^*/n^* - C^*(\dot{V}^*/n^*)^2 = f(\dot{V}^*/n^*), \tag{8.339}$$

where

$$A^* = An_R^2/H_R, \tag{8.340}$$

$$B^* = Bn_R\dot{V}_R/H_R \tag{8.341}$$

and

$$C^* = C\dot{V}_R^2/H_R. \tag{8.342}$$

Table 1 gives a possible way how the homologous curve can be stored for all four quadrants avoiding division with zero as already exercised in RELAP 5/Mod 2 computer code, see *Ransom et al. (1988) p. 27, 179*. Instead of the rotation ratio in Table 8.1 the angular velocity ratio is used. Note that the rotation ratio is equal to the angular velocity ratio

$$n^* = n/n_R = \omega/\omega_R = \omega^*. \tag{8.343}$$

Similarly the hydraulic torque acting on the pump rotor can be presented. One of the possibilities to provide the pump data for computer codes is to read the 16 curves as discussed above as input - see Figs. 8.17 and 8.18. This is getting popular through the using of the RELAP computer code.

But there is more efficient way to use the pump characteristics: It is recommended then to store the data in the form of the so called *Suter* diagram, which is much simpler to use in a computer code - see Fig. 8.19. The idea how to construct a *Suter* diagram is described in the next section.

**Table 8.1** Pump homologous curve definitions

Regime mode	Regime	$\omega^*$	$\dot{V}^*$	$\frac{\dot{V}^*}{\omega^*}$	Head	Key	Torque	Key	
1	Normal pump, or energy dissipating pump, or reverse turbine	$> 0$	$\geq 0$	$\leq 1$	$\frac{\dot{V}^*}{\omega^*}$	$\frac{H^*}{\omega^{*2}}$	HON	$\frac{M^*}{\omega^{*2}}$	TON
2		$> 0$	$\geq 0$	$> 1$	$\frac{\omega^*}{\dot{V}^*}$	$\frac{H^*}{\dot{V}^{*2}}$	HVN	$\frac{M^*}{\dot{V}^{*2}}$	TVN
3	Energy dissipation only	$> 0$	$< 0$	$\geq -1$	$\frac{\dot{V}^*}{\omega^*}$	$\frac{H^*}{\omega^{*2}}$	HOD	$\frac{M^*}{\omega^{*2}}$	TOD



4		> 0	< 0	< -1	$\frac{\omega^*}{\dot{V}^*}$	$\frac{\omega^{*2}}{H^*}$	HVD	$\frac{M^*}{\dot{V}^{*2}}$	TVD
5	Normal turbine, or energy dissipating turbine	$\leq 0$	$\leq 0$	$\leq 1$	$\frac{\dot{V}^*}{\omega^*}$	$\frac{H^*}{\omega^{*2}}$	HOT	$\frac{M^*}{\omega^{*2}}$	TOT
6		$\leq 0$	$\leq 0$	> 1	$\frac{\omega^*}{\dot{V}^*}$	$\frac{H^*}{\dot{V}^{*2}}$	HVT	$\frac{M^*}{\dot{V}^{*2}}$	TVT
7	Reverse pump or energy dissipating reverse pump	$\leq 0$	> 0	$\geq -1$	$\frac{\dot{V}^*}{\omega^*}$	$\frac{H^*}{\omega^{*2}}$	HOR	$\frac{M^*}{\omega^{*2}}$	TOR
8		$\leq 0$	> 0	< -1	$\frac{\omega^*}{\dot{V}^*}$	$\frac{H^*}{\dot{V}^{*2}}$	HVR	$\frac{M^*}{\dot{V}^{*2}}$	TVR

$\omega^*$  = rotational velocity ratio,  $\dot{V}^*$  = volumetric flow ratio,  $H^*$  = head ratio, and  $M^*$  = torque ratio. The key indicates which of the homologous parameters in each octant. The first sign is either H or T and means either head or torque, the second O or V indicates division either by  $\omega^{*2}$  or  $\dot{V}^{*2}$ , and the third N, D, T and R indicates the 1, 2, 3 and 4 quadrant, respectively.

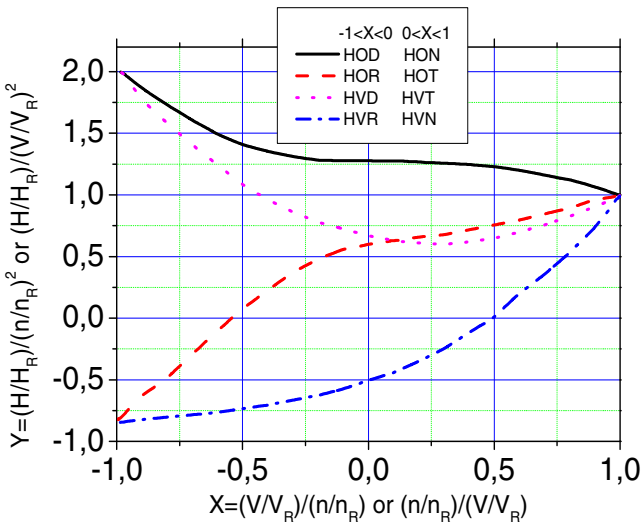


Fig. 8.17 Typical pump homologous head curves

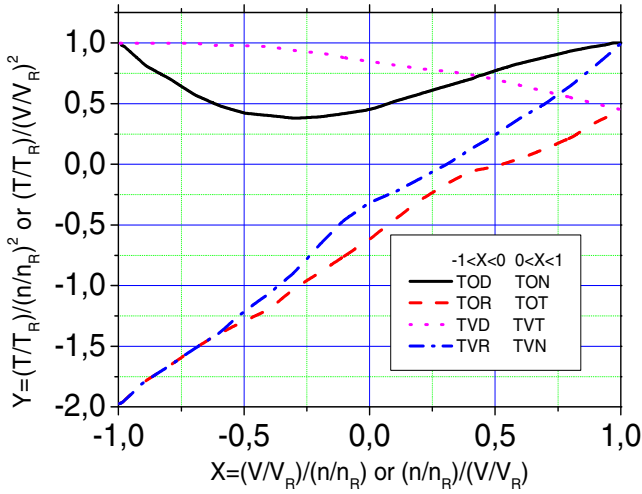


Fig. 8.18 Typical pump homologous torque curves

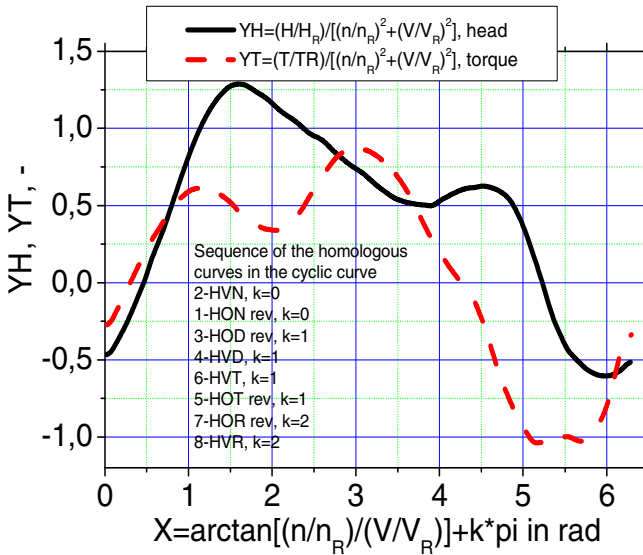


Fig. 8.19 Cyclic head and torque curves for single phase flow

### 8.9.3 Suter diagram

In order to avoid the division of the homologous curve into eight segments the proposal made by *Suter* in (1966) and modified by *Lang*, see in *Kastner, Riedle* and *Seeberger* (1983, *Kastner* and *Seeberger* (1983) is very useful. Instead of using either  $H^*/n^{*2}$  or  $H^*/\dot{V}^{*2}$  it is convenient to use only one function quantity

$$H^*/(n^{*2} + \dot{V}^{*2}), \quad (8.344)$$

which can be easily computed from the known eight homologous curves as follows

$$H^*/(n^{*2} + \dot{V}^{*2}) = \frac{H^*}{n^{*2}} \sqrt{1 + (\dot{V}^*/n^*)^2} = \frac{H^*}{\dot{V}^{*2}} \sqrt{1 + (n^*/\dot{V}^*)^2}, \quad (8.345)$$

or directly deduced from the experimental observation. The region splitting for the argument can also be avoided if one uses as argument

$$X = \arctan(n^*/\dot{V}^*) + k\pi, \quad (8.346)$$

where for the first  $n - \dot{V}$  quadrant  $k = 0$ , for the second and third quadrants  $k = 1$ , and for the fourth quadrant  $k = 2$ . The eight head curves presented in Figs. 8.17 and 8.18 reduce to only one curve, as shown in Fig. 8.19. In similar way the eight torque curves can be reduced to a single curve. Thus finally we can use the *Suter* diagram in the form

$$H^*/(n^{*2} + \dot{V}^{*2}) = f[\arctan(n^*/\dot{V}^*) + k\pi], \quad (8.347)$$

and

$$M_{\beta}^*/(n^{*2} + \dot{V}^{*2}) = f[\arctan(n^*/\dot{V}^*) + k\pi]. \quad (8.348)$$

Examples of such pump characteristics were obtained by *Hollander* in 1953 - see *Donsky* (1961) for three types of pumps in wide range of the so called specific speed defined as follows

$$n_s = n_R \left( \frac{\dot{V}_R}{1m^3/s} \right)^{1/2} \sqrt{\left( \frac{H_R}{1m} \right)^{3/4}}, \quad (8.349)$$

$n_s = 34.8$  rotations per minute for a *centrifugal* pump, 147.17 for a *diagonal* pump and 261.42 for an *axial* pump. *Fox* gives in 1977 the numerical values needed to draw the *Suter* diagram in the first three quadrants of the *Hollander* characteristics (see pp. 135 - 142 in the Russian translation). We supplement these data with the data for the fourth quadrant obtained from the original diagrams as given in *Donsky* (1961) and use in the computer code series IVA the resulting curves as given in Table 8.2 and plotted in Figs. 8.20 to 8.22 as one of the possible options to generate *Suter* diagrams knowing only the pump data for the point having optimum pump efficiency.

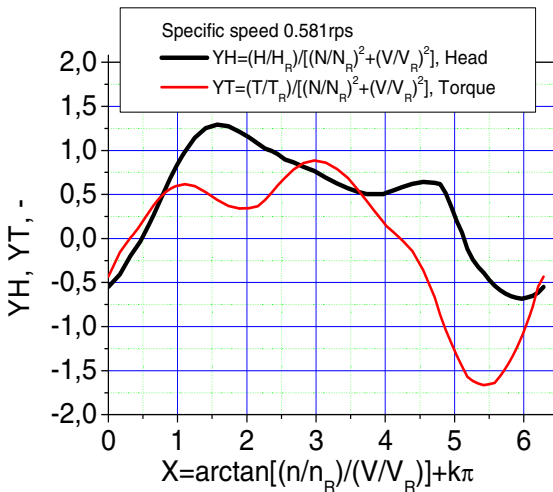
**Table 8.2** *Suter* diagram for three different specific speeds  $n_s$ 

X	$n_s = 0.581$		2.452		4.357	
	YH	YT	YH	YT	YH	YT
0.000	-0.549	-0.433	-1.644	-1.417	-0.952	-0.574
0.168	-0.408	-0.155	-1.098	-0.904	-0.874	-0.602
0.318	-0.198	0.009	-0.623	-0.424	-0.686	-0.542
0.464	-0.032	0.160	-0.350	-0.088	-0.399	-0.312
0.588	0.158	0.297	0.035	0.200	-0.076	0.021
0.695	0.332	0.415	0.308	0.397	0.219	0.303
0.785	0.500	0.500	0.500	0.500	0.500	0.500
0.876	0.650	0.555	0.626	0.579	0.803	0.619
0.983	0.817	0.596	0.776	0.651	1.088	0.741
1.107	0.984	0.616	0.968	0.728	1.409	0.904
1.249	1.143	0.594	1.197	0.882	1.817	1.214
1.406	1.254	0.526	1.479	1.147	2.268	1.626
1.571	1.290	0.440	1.960	1.481	2.729	1.960
1.736	1.274	0.370	2.103	1.538	3.183	2.310
1.893	1.214	0.342	2.187	1.548	3.474	2.647
2.034	1.145	0.345	2.265	1.623	3.576	2.934
2.159	1.080	0.367	2.359	1.711	3.508	3.031
2.266	1.020	0.437	2.474	1.907	3.451	2.945
2.356	0.994	0.520	2.637	2.079	3.272	2.756
2.447	0.958	0.604	2.802	2.356	2.853	2.547
2.554	0.897	0.691	2.900	2.541	2.484	2.182
2.678	0.865	0.783	2.976	2.722	2.161	1.801
2.820	0.812	0.857	2.890	2.749	1.823	1.442
2.976	0.767	0.884	2.624	2.496	1.440	1.030
3.142	0.691	0.859	2.170	2.103	1.082	0.669
3.307	0.623	0.787	1.555	1.525	0.787	0.417
3.463	0.569	0.686	0.992	1.036	0.704	0.415
3.605	0.529	0.552	0.616	0.664	0.616	0.504
3.730	0.504	0.428	0.415	0.387	0.462	0.372
3.836	0.503	0.319	0.279	0.183	0.260	0.106
3.972	0.506	0.187	0.150	-0.100	0.065	-0.075
4.018	0.520	0.141	0.112	-0.171	-0.166	-0.325
4.124	0.548	0.069	0.042	-0.318	-0.416	-0.582
4.249	0.584	-0.024	-0.096	-0.503	-0.687	-0.880
4.391	0.621	-0.144	-0.252	-0.711	-1.026	-1.171
4.547	0.642	-0.360	-0.448	-1.061	-1.508	-1.538
4.712	0.630	-0.671	-0.671	-1.501	-2.190	-2.329
4.791	0.616	-0.868	-0.786	-1.577	-2.427	-2.427
4.869	0.514	-1.035	-0.800	-1.681	-2.625	-2.625
4.948	0.367	-1.182	-0.859	-1.729	-2.785	-2.785
5.027	0.202	-1.321	-0.937	-1.745	-2.928	-2.928
5.105	0.068	-1.450	-1.008	-1.777	-3.071	-3.071

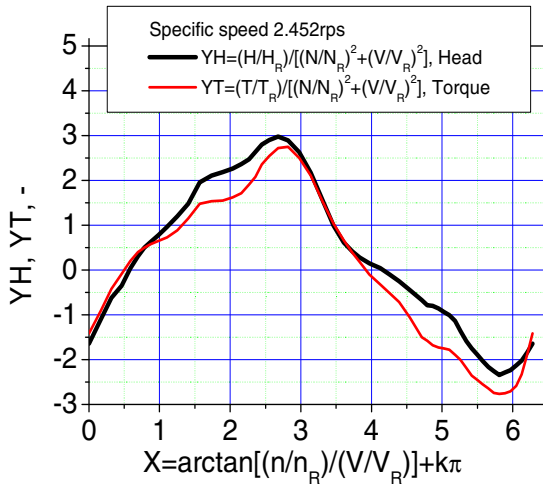
5.184	-0.120	-1.571	-1.137	-1.892	-3.192	-3.192
5.262	-0.239	-1.618	-1.367	-2.008	-3.113	-3.113
5.341	-0.325	-1.646	-1.582	-2.185	-3.034	-3.034
5.419	-0.389	-1.663	-1.751	-2.355	-2.955	-2.955
5.498	-0.470	-1.652	-1.888	-2.451	-2.876	-2.876
5.576	-0.534	-1.641	-2.026	-2.547	-2.716	-2.742
5.655	-0.586	-1.563	-2.163	-2.644	-2.436	-2.525
5.733	-0.625	-1.475	-2.258	-2.742	-2.155	-2.309
5.812	-0.653	-1.374	-2.341	-2.765	-1.875	-2.093
5.890	-0.672	-1.245	-2.286	-2.750	-1.595	-1.850
5.969	-0.682	-1.112	-2.232	-2.708	-1.519	-1.588
6.048	-0.671	-0.957	-2.130	-2.583	-1.448	-1.326
6.126	-0.652	-0.774	-2.021	-2.328	-1.262	-1.053
6.205	-0.616	-0.546	-1.840	-1.889	-1.092	-0.788
6.283	-0.549	-0.433	-1.644	-1.417	-0.952	-0.574

There are situations in the engineering practice where transients have to be analyzed for networks with old pumps for which limited data are available. One usually knows the optimum working point and sometimes the first quadrant characteristics. But the data for all four quadrants are required for transient analysis. In this case one can generate *approximate* four quadrant characteristics by interpolation between the *Hollander's* data. The way used to generate the *Suter* diagrams is the following: First the specific speed for the particular pump under consideration is computed. If

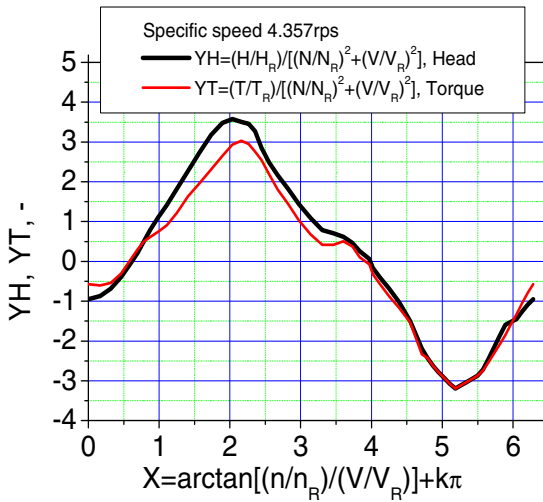
$$n_s < 0.581 \text{ rps} \quad (8.350)$$



**Fig. 8.20** Cyclic head and torque curves for single phase flow centrifugal pump. Measurements made by Hollander 1953



**Fig. 8.21** Cyclic head and torque curves for single phase flow diagonal pump. Measurements made by Hollander 1953



**Fig. 8.22** Cyclic head and torque curves for single phase flow axial pump. Measurements made by Hollander 1953

the *Suter* diagram for  $n_s = 0.581$  rps is used. If

$$n_s > 4.357 \text{ rps}$$

$$(8.351)$$

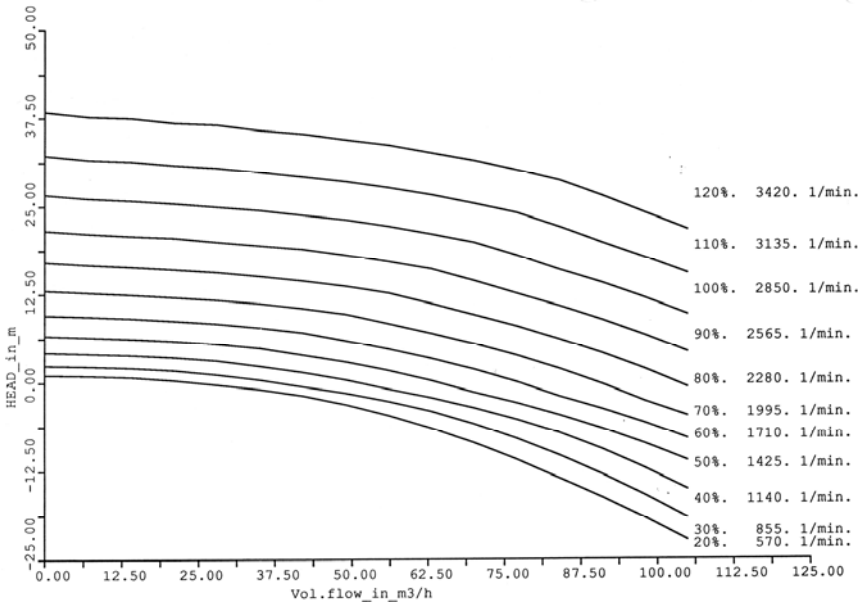
the *Suter* diagram for  $n_s = 4.357 \text{ rps}$  is used. Otherwise the interpolation between two neighboring characteristics is used for 0.581 and 2.452, or 2.452 and 4.357, respectively, depending on the particular specific velocity. The code prints comprehensive data for the first quadrant in order to check the so obtained pump characteristics - see Figs. 8.23 and 8.24.

As already mentioned, usually the first quadrant information is available. One should compare the so obtained characteristics with the known ones for the first quadrant and if the agreement is satisfactory the computation can continue. If the characteristics disagree the actual data should be obtained from the pump manufacturer. In the case that no data except first quadrant data are available the *Suter* diagram can be separately computed from the available first quadrant data and the residual part can be drawn parallel to some known characteristics for a pump with similar geometry.

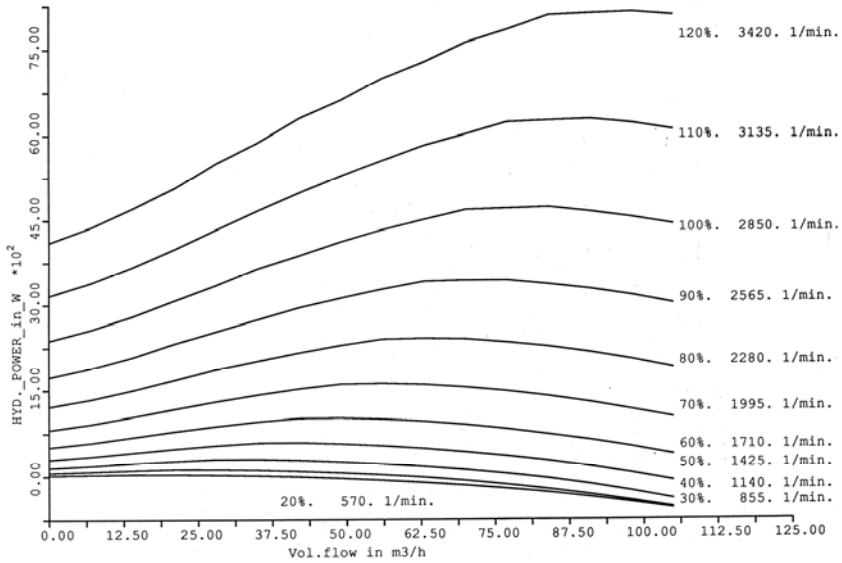
Characteristics of 1:4 and 1:5 models of large scale mixed flow and axial pumps are reported by *Kastner et al.* (1983a, b) and *Kennedy et al.* (1982) where the specific speed was  $n_s \approx 130 \text{ rpm}$ .

Large scale pump characteristics for two pumps are used by *Ransom* (1988).

Characteristics of centrifugal pumps with different constructions are available in *Pohlenz* (1977).



**Fig. 8.23** Head as a function of the volumetric flow. Parameter – rotational frequency



**Fig. 8.24** Hydraulic power as a function of the volumetric flow. Parameter – rotational frequency

Data in a form of the *Suter* diagram can be approximated by a quadratic polynomial of the main pump equation. In order to compute the coefficients  $A^*$ ,  $B^*$ , and  $C^*$  we need three characteristic points, e.g. the point with the maximum efficiency,

$$X = 1, YH = 1, \tag{8.352}$$

the point with the zero volumetric flux

$$X = 0, YH = YH_0, \tag{8.353}$$

and some third point

$$X = X_1, YH = YH_1, \tag{8.354}$$

Substituting into the main pump equation we obtain

$$1 = A^* + B^* - C^*, \tag{8.355}$$

$$C^* = -YH_0, \tag{8.356}$$

$$YH_1 = A^* X_1^2 + B^* X_1 + C^*. \tag{8.357}$$

Solving with respect to the unknown constants we obtain

$$A^* = [(1 - YH_0) X_1 - YH_1 + YH_0] / (X_1 - X_1^2), \tag{8.358}$$

$$B^* = [- (1 - YH_0) X_1^2 + YH_1 - YH_0] / (X_1 - X_1^2), \tag{8.359}$$



$$C^* = -YH_0 \quad (8.360)$$

An approximate model of the two-phase performance can be easily derived having in mind that the term  $C$  is in fact the frictional pressure loss coefficient for one-phase flow. The method of *Lockhart* and *Martinelli* can be applied here in order to compute also the frictional pressure drop exerted by the flow passing the pump. Thus a liquid only two-phase flow multiplier  $\Phi_{f0}^2$  should be computed for volumetric flow  $\dot{V}$  considered to consist only of liquid and with some averaged characteristic hydraulic diameter of the impeller. Thereafter the coefficient  $C^*$  is corrected as follows

$$C^* = -YH_0 \Phi_{f0}^2 \quad (8.361)$$

An immediate result of this approximation is that, if the pump is entered by a two-phase mixture instead of water only, the pressure head decreases due to the dramatic increase of the frictional pressure drop. Additional effects are the flow non-homogeneity and the disturbance of the flow velocity triangles in the impeller.

#### 8.9.4 Computational procedure

Input for each time  $\tau$  :

$\dot{V}_R$ ,  $\omega_R$ ,  $H_R$  and  $M_{fl,R}$  a priori defined.

$\dot{V} = \sum_{l=1}^{l_{\max}} \alpha_l w_l$  provided from the hydraulic modeling in the pipe network.

$\omega$  either prescribed or computed from the pump drive model.

Compute

$$\dot{V}^* = \dot{V} / \dot{V}_R,$$

$$\omega^* = \omega / \omega_R.$$

From the homologous curve compute by interpolation

$$H^* = H^*(V^*, \omega^*),$$

$$M^* = M^*(V^*, \omega^*),$$

Correct  $H^*$  in case of multi-phase flow.

Compute

$$\Delta p_{pump} = \rho g H^* H_R,$$

$$M_{fl} = M^* M_{fl,R}.$$

Use further

$$\Delta p_{pump}$$

in the momentum equations and the energy dissipation

$$\left| M_{fr} \omega - \Delta p_{pump} \dot{V} \right|$$

in the energy balance.

### 8.9.5 Centrifugal pump drive model

The angular momentum balance at the pump rotor reads

$$I_{pr} \frac{d\omega}{d\tau} + M_{fl} + M_{fr} - M_m = 0. \quad (8.362)$$

Here the motor torque is provided by the motor manufacturer as a function of the angular speed (e.g. for induction motors at constant voltage)

$$M_R = M_R(\omega). \quad (8.363)$$

$M_{fl}$  is the hydraulic torque computed from the homologous curve as a function of pump angular speed and flow rate. The frictional torque is a prescribed function of the angular velocity

$$M_{fr} = M_{fr}(\omega), \quad (8.364)$$

where for  $M_{fr} \geq 0 > 0$  for  $\omega \geq 0$  and  $M_{fr} \leq 0$  for  $\omega \leq 0$  (if reverse rotation is allowed at all). A cost down of pump is then defined if the motor torque is zero. Gao et al. (2011) reported interesting solution and data for cost down behavior of a pump in closed circuit filled with water.

### 8.9.6 Extension of the theory to multiphase flow

The theoretical basics for approximate description of the pump performance for multiphase flow are briefly given here. Consider a centrifugal pump with impeller rotating with  $n$  rotations per second or with the angular velocity

$$\omega = 2\pi n, \quad (8.365)$$

inner radius  $r_1$ , and outer radius  $r_2$ . The angular velocity at the inner impeller radius is

$$V_{\theta 1} = \omega r_1 \quad (8.366)$$

and at the outer impeller radius

$$V_{\theta 2} = \omega r_2. \quad (8.367)$$

The corresponding velocity field components at the inner impeller radius are

$$V_{\theta 1,l} = \omega r_1 \quad (8.368)$$

and at the outer impeller radius

$$V_{\theta 2,l} = \omega r_2. \quad (8.369)$$

The relative flow field velocity for an observer rotating with the impeller is  $V_l^r$ . The vector sum of the relative flow field velocity and the angular impeller velocity gives the absolute flow velocity  $V_l^a$ . We designate the angular component of this velocity with  $V_{\theta,l}^a$ . The mass flow in kg/s of each velocity field entering the pump impeller is  $\dot{m}_l$ . The flow is stationary and incompressible. Thus, the force acting on the flow in the angular direction at the inlet is

$$F_{\theta 1,l} = \dot{m}_l V_{\theta 1,l}^a \quad (8.370)$$

and at the impeller outlet

$$F_{\theta 2,l} = -\dot{m}_l V_{\theta 2,l}^a \quad (8.371)$$

The corresponding torques are  $F_{\theta 1,l} r_1$  and  $F_{\theta 2,l} r_2$ . The resulting torque acting on the flow inside the impeller is

$$M_{fl} = \sum_{l=1}^{l_{\max}} (F_{\theta 2,l} r_2 + F_{\theta 1,l} r_1) = \sum_{l=1}^{l_{\max}} \dot{m}_l (V_{\theta 2,l}^a r_2 - V_{\theta 1,l}^a r_1). \quad (8.372)$$

The power inserted into the flow is, therefore,  $M_{fl} \omega$  or

$$\Delta p_{pump} \dot{V} = M_{fl} \omega = \omega \sum_{l=1}^{l_{\max}} \dot{m}_l (V_{\theta 2,l}^a r_2 - V_{\theta 1,l}^a r_1). \quad (8.373)$$

Bearing in mind that

$$\dot{m} = \rho \dot{V}, \quad (8.374)$$

one obtains

$$\Delta p_{pump} = \omega \sum_{l=1}^{l_{\max}} \frac{\dot{m}_l}{\dot{V}} (V_{\theta 2,l}^a r_2 - V_{\theta 1,l}^a r_1) = \sum_{l=1}^{l_{\max}} \frac{\dot{m}_l}{\dot{V}} (V_{\theta 2,l}^a V_{\theta 2,l} - V_{\theta 1,l}^a V_{\theta 1,l}). \quad (8.375)$$

This is the *main pump equation* describing the pressure rise as a function of the impeller speed and geometry. Bearing in mind that the angle between the relative flow velocity  $V_{1,l}^r$  and the impeller angular velocity  $V_{\theta 1,l}$  is  $\beta_1$  and that the angle between the relative flow velocity  $V_{2,l}^r$  and the impeller angular velocity  $V_{\theta 2,l}$  is

$\beta_2$ , the projections of the absolute flow velocities on the angular impeller velocities are

$$V_{\theta 1,l}^a = V_{\theta 1,l} - V_{1,l}^r \cos \beta_1, \quad (8.376)$$

$$V_{\theta 2,l}^a = V_{\theta 2,l} - V_{2,l}^r \cos \beta_2. \quad (8.377)$$

Substituting into the main pump equation we obtain

$$\begin{aligned} \Delta p_{pump} &= \sum_{l=1}^{l_{\max}} \frac{\dot{m}_l}{\dot{V}} \left[ (V_{\theta 2,l} - V_{2,l}^r \cos \beta_2) V_{\theta 2,l} - (V_{\theta 1,l} - V_{1,l}^r \cos \beta_1) V_{\theta 1,l} \right] \\ &= \sum_{l=1}^{l_{\max}} \frac{\dot{m}_l}{\dot{V}} (V_{\theta 2,l}^2 - V_{\theta 1,l}^2 + V_{\theta 1,l} V_{1,l}^r \cos \beta_1 - V_{\theta 2,l} V_{2,l}^r \cos \beta_2). \end{aligned} \quad (8.378)$$

Bearing in mind that the relative flow velocity is equal to the volumetric flow divided by the cross section normal to  $V^r$ , i.e.,

$$V_{1,l}^r = \dot{V}_l / (\alpha_{1,l} A_1), \quad (8.379)$$

$$V_{2,l}^r = \dot{V}_l / (\alpha_{2,l} A_2), \quad (8.380)$$

one obtains

$$\begin{aligned} \Delta p_{pump} &= \sum_{l=1}^{l_{\max}} \frac{\dot{m}_l}{\dot{V}} \left[ (r_2^2 - r_1^2) \omega^2 + \left( \frac{r_1 \cos \beta_1}{\alpha_{1,l} A_1} - \frac{r_2 \cos \beta_2}{\alpha_{2,l} A_2} \right) \omega \dot{V}_l \right] \\ &= \sum_{l=1}^{l_{\max}} \frac{\dot{m}_l}{\dot{V}} \left[ (r_2^2 - r_1^2) 4\pi^2 n^2 + \left( \frac{r_1 \cos \beta_1}{\alpha_{1,l} A_1} - \frac{r_2 \cos \beta_2}{\alpha_{2,l} A_2} \right) 2\pi n \dot{V}_l \right]. \end{aligned} \quad (8.381)$$

If additionally we consider that the flow exerts some pressure loss due to friction with the pump structures in the impeller region, we finally obtain

$$\begin{aligned} \Delta p_{pump} &= \sum_{l=1}^{l_{\max}} \frac{\dot{m}_l}{\dot{V}} \left[ (r_2^2 - r_1^2) \omega^2 + \left( \frac{r_1 \cos \beta_1}{\alpha_{1,l} A_1} - \frac{r_2 \cos \beta_2}{\alpha_{2,l} A_2} \right) \omega \dot{V}_l \right] - \xi \frac{1}{2} \frac{\rho}{A_m} \dot{V}^2 \Phi_{20}^2 \\ &= \sum_{l=1}^{l_{\max}} \frac{\dot{m}_l}{\dot{V}} \left[ (r_2^2 - r_1^2) 4\pi^2 n^2 + \left( \frac{r_1 \cos \beta_1}{\alpha_{1,l} A_1} - \frac{r_2 \cos \beta_2}{\alpha_{2,l} A_2} \right) 2\pi n \dot{V}_l \right] - \xi \frac{1}{2} \frac{\rho}{A_m} \dot{V}^2 \Phi_{20}^2, \end{aligned} \quad (8.382)$$

where  $\xi$  is some friction loss coefficient computed as with the multiphase mass flow rate and liquid properties,  $\Phi_{20}^2$  is the *Lochart* and *Martinelli* two-phase multiplier, and  $A_m$  some averaged flow cross section. The pressure rise of the pump is usually expressed with the head  $H$

$$\Delta p_{pump} = \rho_h g H, \quad (8.383)$$

where  $g$  is the gravitational acceleration and

$$\rho_h = \sum_{l=1}^{l_{\max}} \alpha_l \rho_l \quad (8.384)$$

is the homogeneous mixture density. Usually, the main pump equation is written in the form

$$\begin{aligned} \rho_h g H = & (r_2^2 - r_1^2) 4\pi^2 \left[ \left( \sum_{l=1}^{l_{\max}} \dot{m}_l \right) / \dot{V} \right] n^2 \\ & + 2\pi n \dot{V} \sum_{l=1}^{l_{\max}} \left[ \frac{\dot{m}_l}{\dot{V}} \left( \frac{r_1 \cos \beta_1}{\alpha_{1,l} A_1} - \frac{r_2 \cos \beta_2}{\alpha_{2,l} A_2} \right) \frac{\dot{V}_l}{\dot{V}} \right] - \xi \frac{1}{2} \frac{\rho}{A_m} \dot{V}^2 \Phi_{20}^2, \end{aligned} \quad (8.385)$$

or

$$H^+ = A^+ n^2 + B^+ n \dot{V} - C^+ \dot{V}^2, \quad (8.386)$$

where

$$A^+ = (r_2^2 - r_1^2) \frac{4\pi^2}{\rho_h g} \left[ \left( \sum_{l=1}^{l_{\max}} \dot{m}_l \right) / \dot{V} \right] = A \left( \sum_{l=1}^{l_{\max}} \dot{m}_l \right) / (\rho_h \dot{V}) \quad (8.387)$$

$$\begin{aligned} B^+ &= \frac{2\pi}{\rho_h g} \sum_{l=1}^{l_{\max}} \left[ \frac{\dot{m}_l}{\dot{V}} \left( \frac{r_1 \cos \beta_1}{\alpha_{1,l} A_1} - \frac{r_2 \cos \beta_2}{\alpha_{2,l} A_2} \right) \frac{\dot{V}_l}{\dot{V}} \right] \\ &= B \frac{\sum_{l=1}^{l_{\max}} \left[ \frac{\dot{m}_l}{\dot{V}} \left( \frac{r_1 \cos \beta_1}{\alpha_{1,l} A_1} - \frac{r_2 \cos \beta_2}{\alpha_{2,l} A_2} \right) \frac{\dot{V}_l}{\dot{V}} \right]}{\rho_h \left( \frac{r_1 \cos \beta_1}{A_1} - \frac{r_2 \cos \beta_2}{A_2} \right)} \end{aligned} \quad (8.388)$$

$$C^+ = \xi \frac{1}{2} \frac{\rho}{\rho_h g A_m} \Phi_{20}^2 = C \frac{\rho}{\rho_h} \Phi_{20}^2 \approx C \Phi_{20}^2. \quad (8.389)$$

Obviously, for single-phase flow

$$\frac{1}{\rho_h \dot{V}} \sum_{l=1}^{l_{\max}} \dot{m}_l = 1 \quad (8.390)$$

$$\frac{\sum_{l=1}^{l_{\max}} \left[ \frac{\dot{m}_l}{\dot{V}} \left( \frac{r_1 \cos \beta_1}{\alpha_{1,l} A_1} - \frac{r_2 \cos \beta_2}{\alpha_{2,l} A_2} \right) \frac{\dot{V}_l}{\dot{V}} \right]}{\rho_h \left( \frac{r_1 \cos \beta_1}{A_1} - \frac{r_2 \cos \beta_2}{A_2} \right)} = 1 \quad (8.391)$$

$$\Phi_{20}^2 = 1 \quad (8.392)$$

and, therefore,

$$A^+ = A, \quad (8.393)$$

$$B^+ = B, \quad (8.394)$$

$$C^+ = C. \quad (8.395)$$

## Appendix 1: Chronological references to the subject critical two-phase flow

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# 9 Detonation waves caused by chemical reactions or by melt-coolant interactions

## 9.1 Introduction

Analyzing a fascinating physical phenomenon such as the melt-water detonation in this chapter we will give an interesting application of the theory of multi-phase flows – namely the analysis of the detonation wave propagation during the interaction of molten materials with liquids such as that of iron with water.

Melt-coolant interaction analysis plays an important role in risk estimation for facilities in which high-temperature molten material comes or may come into contact with low-temperature liquid coolant. Contacts of iron with water or uranium dioxide with water are some examples of this phenomenon. One of the intriguing discussions in this field is motivated by the observation that molten alumina produces *severe interactions* when dropped into water whereas molten uranium dioxide at a similar temperature produces only *moderate interactions*. Use of the solution of the *Fourier* equation for contact of melt and coolant during a prescribed time as presented in *Kolev (1999a)* does not yield any explanation for these differences. It was shown in *Kolev (1999a)* that the ratio of the discharged to the available thermal energy for  $UO_2$  at an initial temperature  $T_3 = 3000K$  in contact over  $\Delta\tau_{23} \approx 0.001s$  with water at temperature  $T_2 = 30^\circ C$  and atmospheric pressure is  $\eta_{th} = 0.000133/D_3$ . For the same conditions for  $Al_2O_3$  this ratio is  $\eta_{th} = 0.000129/D_3$ . Here  $D_3$  is the particle size. It is evident that the thermal properties for transient heat conduction by these fluid pairs do not differ to such a large extent that may explain the differences in the interaction. Thus the question still remains why these materials behave so differently. We will demonstrate in this chapter that the reason is *mainly the differences in their caloric equations of state* – this is the first target of this chapter.

The next interesting problem is associated with the question of how to judge the explosivity of melt-water mixtures in engineering premixing computations used for upper bound estimates of the possible detonation energy release. In *Kolev (1999b)* the criterion that mixtures of melt and water can only experience severe

detonations if one of the liquids continuously surrounds the other was introduced. The roots of this idea go back to the works by *Henry* and *Fauske* (1981a, b) and are discussed in detail by *Park* and *Corradini* (1991). In the same work the so called “stoichiometric” mixtures were introduced which are defined as mixtures in which there is an optimum amount of water available locally to permit the melt to transfer all its thermal energy by evaporation. The question whether dispersed systems such as melt droplets flowing together with water droplets may also be of some risk is still unresolved. We will answer in this work two questions associated with this issue. *First, is it possible to have detonation waves in such disperse-disperse systems and second, if so, what are the expected pressure wave magnitudes and can they be considered as dangerous?*

A further interesting aspect results from the detonation theory of melt-water interaction proposed by *Board* et al. (1975). Just as in detonation of reacting gases, *Board* et al. considered the fine fragmentation behind a shock front and the subsequent steam production and expansion as the driving force for transferring shock waves in self-sustaining detonation waves. The ideas propagated by *Board* have been discussed worldwide. In particular, his remark that not all the melt behind the front may necessarily participate in the energy transfer process was widely accepted and introduced in the detonation analysis – see for instance the work by *Shamoun* and *Corradini* (1995). *Board’s* remark that not all the coolant behind the front necessarily has to participate in the energy transfer process was controversially discussed. For instance, some authors thought that assuming that all the melt and coolant participate in the interaction is conservative, e.g. *Frost*, *Lee* and *Ciccarelli* (1991), and this may serve as an upper bound estimate. This was criticized by *Yuen* and *Theofanous* (1997). Based on their own experiments, these authors introduced the idea of not only a limited amount of melt but also of a limited amount of coolant participating in the interaction – the so-called micro-interaction concept. *Yuen* and *Theofanous* came to an important conclusion for practical applications: *mixtures judged as non-explosive assuming that all the water is participating in the debris quenching may be explosive if only part of the water is participating in the debris quenching*. Of course, this problem naturally disappears if numerical computations are used that feature a very fine discretization grid for resolving regions of near and far interactions numerically. In simulations with a characteristic grid size much larger than the near interaction length this problem is acute and a quasi-micro-interaction approach is thus inevitable. One of the targets in the analysis presented here is to discriminate for the three important materials pairs ( $UO_2$ -water,  $Al_2O_3$ -water and  $Fe$ -water) between coolant entrainment leading to detonations and entrainment leading to quenching, thereby preventing detonations.

The assumption of a homogeneous mixture was relaxed by *Berthoud* (1999) and *Scott* and *Berthoud* (1978) in a steady-state analysis and resulted in interesting estimates, one of them being the “thickness of the discontinuity front“ which depends on the initial particle size. Nevertheless, all applied computations

use a detonation theory based on homogeneity, inclusive of those of *Scott* and *Berthoud*.

The theory of weak pressure waves, called sound waves, and strong pressure waves, called shock waves, for single-phase flow is also reiterated briefly in Section 9.2 to assist the reader in the transition to the multi-phase theory. The model for detonation analysis for multi-phase flow that will be used in Section 9.4 including a brief discussion of the numerical method is presented in Section 9.3.

In Section 9.4 the application results will be discussed. Finally the answers to the questions posed above as obtained by this analysis will be summarized.

## 9.2 Single-phase theory

### 9.2.1 Continuum sound waves (*Laplace*)

Consider a small pressure pulse propagating through a continuum. The pressure disturbance causes velocity disturbance. We are interested in the velocity of propagation of this disturbance. Consider a control volume with infinitesimal thickness located around a plane that is perpendicular to the flow velocity. The control volume moves with the plane. The steady-state conservation equations for mass, momentum (neglecting body and viscous forces), energy (neglecting all dissipative effects such as work done by viscous stress and heat transfer due to thermal conduction and diffusion) are

$$d(\rho w) = 0, \quad (9.1)$$

$$d(\rho w^2) + dp = 0, \quad (9.2)$$

$$d\left[\rho w\left(h + \frac{1}{2}w^2\right)\right] = 0. \quad (9.3)$$

Now remember that all flow parameters are continuous in space. Using the mass conservation equation, the momentum and energy conservation equations reduce to

$$\rho d\left(\frac{1}{2}w^2\right) + dp = 0 \quad \text{or} \quad -w^2 d\rho + dp = 0 \quad (9.4)$$

$$d\left(h + \frac{1}{2}w^2\right) = 0 \quad (9.5)$$

If we exclude the differential of the kinetic energy, this yields  $\rho dh - dp = 0$  which after application of the *Gibbs* equation is simply

$$ds = 0. \quad (9.6)$$



Using the equation of state in the form

$$d\rho = \left( \frac{\partial \rho}{\partial s} \right)_p ds + \frac{d\rho}{a^2}, \quad (9.7)$$

where

$$\frac{1}{a^2} = \left( \frac{\partial \rho}{\partial p} \right)_s \quad (9.8)$$

the momentum equation becomes

$$\left[ 1 - \left( \frac{w}{a} \right)^2 \right] dp = 0. \quad (9.9)$$

For a pressure difference that is non-zero, this is satisfied only if the velocity of the flow with respect to the plane is equal to a value defined solely by the local fluid properties,

$$w = a,$$

and called the *sound velocity*. This remarkable result was first obtained by *Laplace* in 1816 and is generally accepted today. It simply says that:

- 1) Small pressure-velocity disturbances called *weak waves* or *acoustic waves* propagate with a velocity which is a function only of the local continuum parameters.
- 2) Across small pressure-velocity disturbances (neglecting dissipative effects) the continuity of the flow variables implies entropy conservation across any infinitesimal distance.

## 9.2.2 Discontinuum shock waves (*Rankine-Hugoniot*)

Now consider a plane normal to the flow velocity separated into two continuous regions featuring discontinuity at that plane. We designate the medium ahead of wave motion and the medium behind the wave with indices 1 and 2, respectively. Following *Landau* and *Lifshitz* (1953) we call the side of the plane facing the non-disturbed medium the down-flow side and the side facing the disturbed medium the up-flow side. Note that the major difference from the case with continuous flow parameters is that the *differentiation across the plane is not possible*. We will see that velocities before and after the plane are related through relationships that differ from the isentropic wave disturbance as discussed above. The conservation equations governing this process are similar to Eqs. (9.1), (9.2) and (9.3) but written as a primitive balance applied on the control volume:

$$\Delta(\rho w) = 0, \text{ or } \rho w = \text{const} \text{ or } w_1 = \rho w / \rho_1 \text{ or } w_2 = \rho w / \rho_2, \quad (9.10)$$

$$\Delta(\rho w^2) + \Delta p = 0, \text{ or } \rho_2 w_2^2 - \rho_1 w_1^2 + p_2 - p_1 = 0, \quad (9.11)$$

$$\Delta \left[ \rho w \left( h + \frac{1}{2} w^2 \right) \right] = 0, \text{ or } \rho_2 w_2 \left( h_2 + \frac{1}{2} w_2^2 \right) - \rho_1 w_1 \left( h_1 + \frac{1}{2} w_1^2 \right) = 0. \quad (9.12)$$

Thus, the parameters in front of and behind the discontinuity plane are related to each other by Eqs. (9.10) through (9.12). Using mass conservation, momentum and energy conservation can be rewritten as

$$(\rho w)^2 = \frac{p_2 - p_1}{\frac{1}{\rho_1} - \frac{1}{\rho_2}}, \quad (9.13)$$

$$h_2 - h_1 + \frac{1}{2} (\rho w)^2 \left( \frac{1}{\rho_2^2} - \frac{1}{\rho_1^2} \right) = 0. \quad (9.14)$$

By inserting the mass flow rate from the momentum equation into the energy equation we obtain

$$h_2 - h_1 - \frac{1}{2} (p_2 - p_1) \left( \frac{1}{\rho_2} + \frac{1}{\rho_1} \right) = 0. \quad (9.15)$$

Using the equation of state

$$h_2 = f(p_2, \rho_2) \quad (9.16)$$

(e.g.  $h_2 = c_{p2} T_2 = \frac{\kappa_2}{\kappa_2 - 1} \frac{p_2}{\rho_2}$  for perfect gases) the state behind the front is a

unique function of the state in the undisturbed region. Equations (9.15) and (9.16) have to be solved with respect to  $p_2$  by iterations. With the pressure difference known, the mass flow rate is then readily computed using Eq. (9.13). The velocities are then computed from the mass conservation equation. A very convenient formula for the velocity difference is

$$w_1 - w_2 = -\sqrt{(p_2 - p_1) \left( \frac{1}{\rho_1} - \frac{1}{\rho_2} \right)}, \quad (9.17)$$

which is the *shock wave velocity in a laboratory frame*. Equation (9.15) is frequently written in terms of specific internal energies

$$e_2 - e_1 + \frac{p_2 + p_1}{2} \left( \frac{1}{\rho_2} - \frac{1}{\rho_1} \right) = 0. \quad (9.18)$$

A similar procedure can be applied by using the equation of state in the form

$$e_2 = f(p_2, \rho_2). \quad (9.19)$$

This remarkable result was first obtained by *Rankine* in 1870 and independently by *Hugoniot* in 1887 and is widely accepted today. It says that *there is a velocity discontinuity in the shock plane*. The velocities in front of and behind the shock plane relative to the shock plane are a unique function of state of the undisturbed region 1.

The pressure and the density behind the shock front are higher than those in front of the plane. There is an entropy increase across the shock front. This conclusion was reached by *Rayleigh* (1910) and *Taylor* (1910). It is the only known case where flow processes in a non-viscous and non-dissipating medium take place irreversibly, that is with entropy increase, see the discussion by *Landau* and *Lifshitz* (1953) p. 443. The velocity in front of the shock with respect to the front is greater than the local sound velocity,  $w_1 > a_1$ , and the velocity behind the shock with respect to the front is less than the local sound velocity,  $w_2 < a_2$ . Whatever happens before the shock cannot influence the shock propagation. The events behind the shock influence the shock form and propagation.

In the laboratory frame of reference we have

$$w_1 = -w_{cs},$$

$$w_2 = -(w_{cs} - w_2^a).$$

Here  $w_{cs}$  is the velocity of our moving frame of reference and  $w_2^a$  is the absolute velocity behind the shock. Therefore

$$w_1 - w_2 = -w_2^a.$$

In the case of shock wave propagation in an explosive mixture, the compression may increase the temperature behind the shock to values that are higher than the ignition temperature. In the case of an exothermic chemical reaction, it is possible for the released energy behind the front to cause additional drive, resulting in a rise in the *velocity behind the wave with respect to the front* to the local velocity of sound,  $w_2 = a_2$ . This process is usually visualized in the literature as a piston moving with a velocity  $-(w_1 - a_2)$  pushing the reaction front ahead and leaving behind a refraction wave. In this case, neither the disturbance behind the shock nor the disturbance ahead of the shock can have any influence on the shock wave propagation. The propagation process becomes self-sustaining.

A shock wave fulfilling the condition

$$w_2 = a_2 \quad (9.20)$$

is called a *detonation wave*, *Chapman* (1899) and *Jouguet* (1905)

This condition is called the *Chapman-Jouguet* condition – see *Chapman* who found in 1899 that the velocity behind the front with respect to the front in reality take the minimum of the possible detonation states, and *Jouguet* who found in 1905 that the velocity behind the front with respect to the front is equal to the sound velocity of the burned products. *Crussard* showed in 1907 that both statements are identical.

It is important to note that the equations of state of the mixture ahead of the front and behind the front are different.

The energy conservation equation also has to take into account the reaction energy release. An additional term then occurs,  $\Delta h_{reaction}$ ,

$$h_2 - h_1 + \frac{1}{2} w_2^2 - \frac{1}{2} w_1^2 - \Delta h_{reaction} = 0$$

or

$$h_2 - h_1 - \Delta h_{reaction} - \frac{1}{2} (p_2 - p_1) \left( \frac{1}{\rho_2} + \frac{1}{\rho_1} \right) = 0. \quad (9.21)$$

The formation enthalpy  $\Delta h_{reaction}$  in *J* per *kg* of the final mixture at zero absolute temperature is related to the formation enthalpy  $\Delta h_{ref}$  at the reference conditions  $(p_{ref}, T_{ref})$  by the relation

$$\Delta h_{reaction} = (\Delta h_{ref} - c_{p2} T_{ref}) \Delta C_{st\_mix}. \quad (9.22)$$

Here the multiplier is first demonstrated for the case of an initially stoichiometric gas mixture containing hydrogen and oxygen without radicals in equilibrium

$$\Delta C_{st\_mix} = C_{1,H_2} + C_{1,O_2} - C_{2,H_2} - C_{2,O_2}.$$

The multiplier converts the formation enthalpy related to 1*kg* of hydrogen-oxygen mixture to that related to 1*kg* gas mixture. If for instance other gas components are in the mixture that do not participate into the reaction the heating effect of the reaction on the final products is smaller. If the resulting temperature after the detonation wave is so low that the dissociation can be neglected we have  $C_{2,H_2} = 0$ ,  $C_{2,O_2} = 0$ . The general form of the multiplier reads

$$\Delta C_{st\_mix} = \Delta C_{H_2} + \Delta C_{O_2} = (n_{H_2} M_{H_2} + n_{O_2} M_{O_2}) \min \left( \frac{\Delta C_{H_2}}{n_{H_2} M_{H_2}}, \frac{\Delta C_{O_2}}{n_{O_2} M_{O_2}} \right)$$

$$= 9 \min(\Delta C_{H_2}, \Delta C_{O_2} / 8),$$

where  $n_{H_2} = 2$ ,  $n_{O_2} = 1$  are the stoichiometric coefficients equal to the number of *kg-moles* of each species that participates in the reaction and  $M_{H_2} = 2 \text{ kg}$ ,  $M_{O_2} = 32 \text{ kg}$  are the *kg-mole masses*.

The *Rankine-Hugoniot* curve constructed on the basis of the equation of state before the shock is called the *shock adiabetic*. The *Rankine-Hugoniot* curve constructed on the basis of the equation of state after the shock is called the *detonation adiabetic*. The thermodynamic properties in front of and behind the detonation discontinuity therefore lie on different thermodynamic curves.

Detonation waves have a destructive potential. It is interesting to know how much energy can be released in this case and transferred into technical work. Using Eq. (7.26) we obtain

$$-w_{r,21}^{vdp} = h_2 - h_1 + T_\infty (s_1 - s_2). \quad (9.23)$$

### 9.2.3 The *Landau* and *Liftshitz* analytical solution for detonation in perfect gases

*Landau* and *Liftshitz* obtained in 1953 an analytical detonation solution for perfect gas mixtures. The main idea of the procedure is to express the state of the gas behind the detonation front as a function of the parameters before the detonation front. Starting with the definition equation for having detonation

$$w_2 = a_2 = \sqrt{\kappa_2 p_2 / \rho_2}, \quad (9.24)$$

and therefore  $(\rho w)^2 = \kappa_2 p_2 \rho_2$  or

$$\rho_2 = (\rho w)^2 / (\kappa_2 p_2), \quad (9.25)$$

the authors solved the system consisting of the last equation and the momentum Eq. (9.13) with respect to the pressure and density of the burned products. The result is

$$p_2 = \frac{p_1 + (\rho w)^2 \frac{1}{\rho_1}}{\kappa_2 + 1} = \frac{p_1 + \rho_1 w_1^2}{\kappa_2 + 1}. \quad (9.26)$$

$$\rho_2 = \frac{(\rho w)^2}{\kappa_2 p_2} = \frac{\kappa_2 + 1}{\kappa_2} \frac{(\rho w)^2}{p_1 + (\rho w)^2} \frac{1}{\rho_1} = \frac{\kappa_2 + 1}{\kappa_2} \frac{\rho_1^2 w_1^2}{p_1 + \rho_1 w_1^2}, \quad (9.27)$$

Using the burned gas density, the burned gas velocity relative to the moving front is

$$w_2 = \frac{\rho_1 w_1}{\rho_2} = \frac{\kappa_2}{\kappa_2 + 1} \frac{p_1 + \rho_1 w_1^2}{\rho_1 w_1}. \quad (9.28)$$

The enthalpies of the perfect gas mixtures are computed with the zero point selected at zero temperature as follows

$$h_1 = c_{p1} T_1 = \frac{\kappa_1}{\kappa_1 - 1} \frac{p_1}{\rho_1}, \quad (9.29)$$

$$h_2 = c_{p2} T_2 = \frac{\kappa_2}{\kappa_2 - 1} \frac{p_2}{\rho_2} = \frac{\kappa_2^2}{(\kappa_2^2 - 1)(\kappa_2 + 1)} \frac{(p_1 + \rho_1 w_1^2)^2}{\rho_1^2 w_1^2}. \quad (9.30)$$

Here  $T = \frac{p}{\rho R}$ ,  $\frac{c_p}{R} = \frac{\kappa}{\kappa - 1}$ , Eqs. (9.26) and (9.27) are used. Therefore

$$h_2 + \frac{1}{2} w_2^2 = \frac{1}{2} \frac{\kappa_2^2}{\kappa_2^2 - 1} \frac{(p_1 + \rho_1 w_1^2)^2}{\rho_1^2 w_1^2}. \quad (2.31)$$

Replacing all the terms in the energy conservation equation

$$h_2 - h_1 + \frac{1}{2} w_2^2 - \frac{1}{2} w_1^2 - \Delta h_{reaction} = 0, \quad (2.32)$$

one obtains

$$\frac{1}{2} \frac{\kappa_2^2}{\kappa_2^2 - 1} \frac{(p_1 + w_1^2 \rho_1)^2}{w_1^2 \rho_1^2} - \frac{\kappa_1}{\kappa_1 - 1} \frac{p_1}{\rho_1} - \frac{1}{2} w_1^2 - \Delta h_{reaction} = 0, \quad (2.33)$$

or

$$w_1^4 - 2 \left[ (\kappa_2^2 - 1) \Delta h_{reaction} + \frac{\kappa_2^2 - \kappa_1}{\kappa_1 - 1} \frac{p_1}{\rho_1} \right] w_1^2 + \left( \kappa_2 \frac{p_1}{\rho_1} \right)^2 = 0. \quad (2.34)$$

This is a bi-quadratic equation of the type

$$x^4 - 2px^2 + q = 0, \quad (2.35)$$

having the largest positive solution

$$x = \sqrt{p + \sqrt{p^2 - q}} = \sqrt{(p + \sqrt{q})/2} + \sqrt{(p - \sqrt{q})/2}. \quad (3.36)$$

For a strong detonation waves the equation simplifies to

$$w_1 = \sqrt{2(\kappa_2^2 - 1)\Delta h_{reaction}}. \quad (3.37)$$

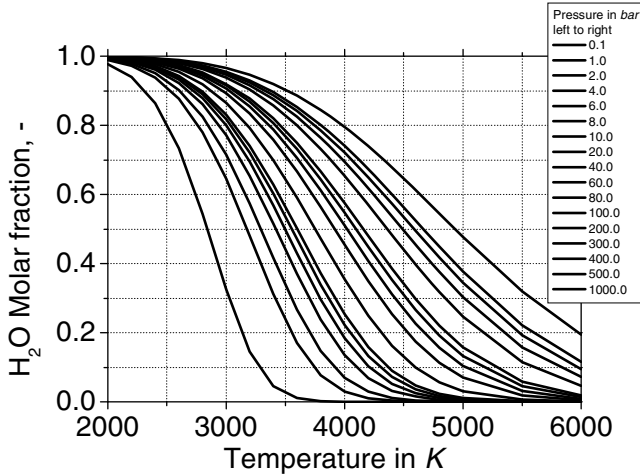
Knowing the velocity of the unburned products with respect to the detonation front, the pressure and the density behind the front are computed using Eqs. (2.26) and (2.27). The velocity of the burned products with respect to the front is then computed using Eq. (2.24). Note the important limitation of this theory

$$\kappa_1 = const, \quad (3.38)$$

$$\kappa_2 = const. \quad (3.39)$$

A systematic summary of data for the specific capacity was obtained in 1987 by Robert et al. (1987) for 165 gases and radicals. Some of these data are based on measurements up to 5000K and others are provided by quantum chemistry computations. These data help us to compute properly the specific capacity at constant pressure for mixtures of gases or radicals and therefore to compute properly the isentropic exponents of the burned products. This approach results in more realistic temperatures.

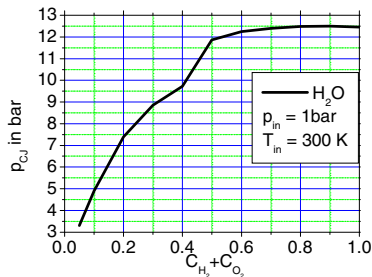
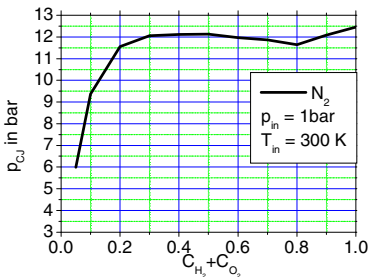
For large values of  $\Delta h_{reaction}$ , the computed temperatures, even with an appropriate isentropic exponent of the burned products, are unrealistically high. At high temperature the products obtained start to dissociate absorbing energy. This process limits the temperature increase itself.



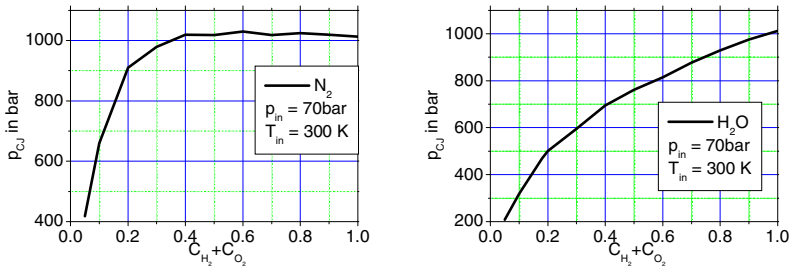
**Fig. 9.1** Molar fraction of dissociated steam as a function of temperature with pressure as a parameter

This was found first by *Zeldovich* in 1940 p. 544 who proposed using the specific heat as a function of temperature taking into account the dissociation. Figure 9.1 shows the molar fraction of the dissociated steam as a function of the local temperature for different pressures. We compute this concentration as discussed in Chapter 3. Other authors have introduced a coefficient of completion of the reaction less than unity depending on the temperature of the burned gases.

We give here two examples of detonation CJ pressures, CJ temperatures and CJ velocities in Figs. 9.2 through 9.4 (index 2 is replaced with CJ if the *Chapman-Jouguet* condition is fulfilled). We consider two cases one with nitrogen as an inert component and the second with steam. We present the results as a function of the concentration of the stoichiometric mixture of hydrogen and oxygen.

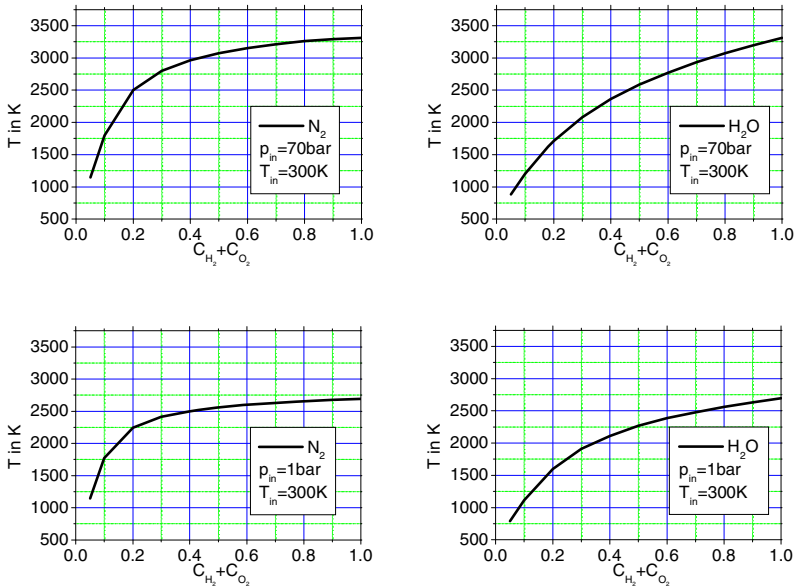






**Fig. 9.2** CJ pressure for stoichiometric hydrogen-oxygen mixtures with nitrogen or steam as an inert component

For the case of steam as an inert component we use the equation of state for dissociated steam. We see the effect of dissociation. It is manifested in the reduction of the effective isentropic exponent of the products close to unity. This makes the increases temperature and pressure with increasing energy release considerable smaller than if computed without consideration of the dissociation.

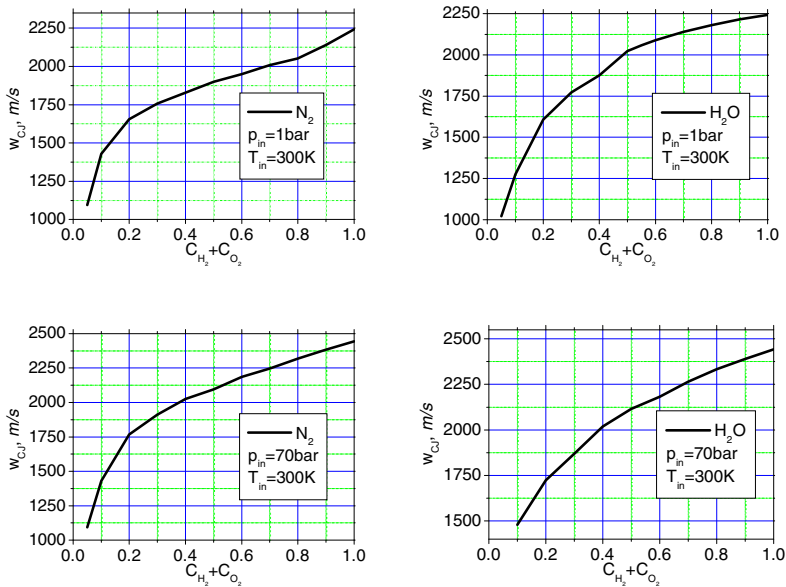


**Fig. 9.3** CJ temperature for stoichiometric hydrogen-oxygen mixtures with nitrogen or steam as an inert component

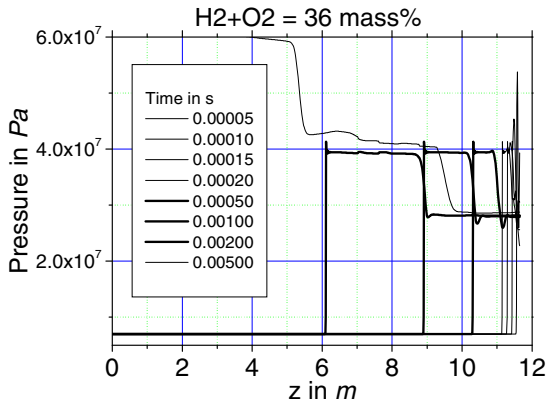
### 9.2.4 Numerical solution for detonation in closed pipes

Combustion in pipes with closed ends usually starts generating a shock wave which travels forwards and backwards between the front and the dead end. The reflected waves interact with the shock so that the picture of reacting detonation conditions is much more complex than that described by the idealized steady state theory. To demonstrate this we analyze a detonation process in a pipe  $0.1\text{m}$  in diameter and  $11.65\text{m}$  in length. The left half of the pipe is filled with steam, and the right hand side with an explosive mixture of stoichiometric hydrogen-oxygen and steam as an inert species. The initial pressure is selected to be  $70\text{ bar}$ . The initial temperature is  $400\text{ K}$ . The solution is obtained with single step combustion kinetic postulating an ignition temperature of  $783\text{ K}$ . We use a donor-cell discretization for the convective terms and second order discretization for the diffusion terms with IVA\_5M computer code. The axial discretization has 1165  $1\text{-cm}$  cells. The results are given below.

Figure 9.5 gives a typical development of the shock wave into a detonation wave after ignition at one end. The initial transient process, followed by the semi-steady-state formation of the detonation front, is clearly recognizable. The extension of the CJ pressure region followed by a refraction wave is also clearly visible. There



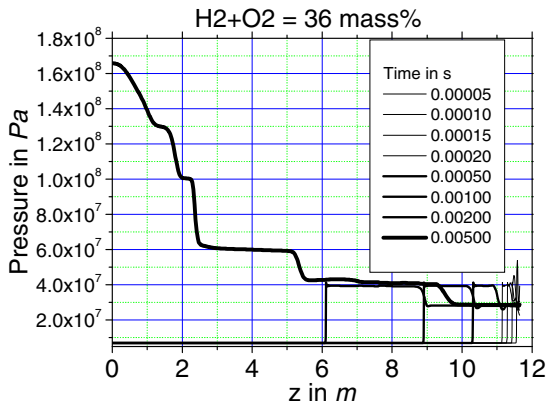
**Fig. 9.4** CJ velocity for stoichiometric hydrogen-oxygen mixtures with nitrogen or steam as an inert component



**Fig. 9.5** Pressure as a function of the distance for different times after the ignition

is the effect of reflection of the wave at the moment it reaches the separation point and penetrates the inert gases without fuel continuing to propagate and without the pushing effect of the combustion behind the shock front. The reflected shock wave is clearly visible from Fig. 9.6. The destructive potential of such waves for pipes not designed to sustain such pressures is obvious.

Increasing the concentration of the stoichiometric explosive mixture leads to an increase in the CJ pressure. The velocity of sound in the burning products is then higher. Therefore, the frequency of the wave reflections between the front and the dead end is much higher, as demonstrated in Fig. 9.7.



**Fig. 9.6** Pressure as a function of distance for different times after the ignition

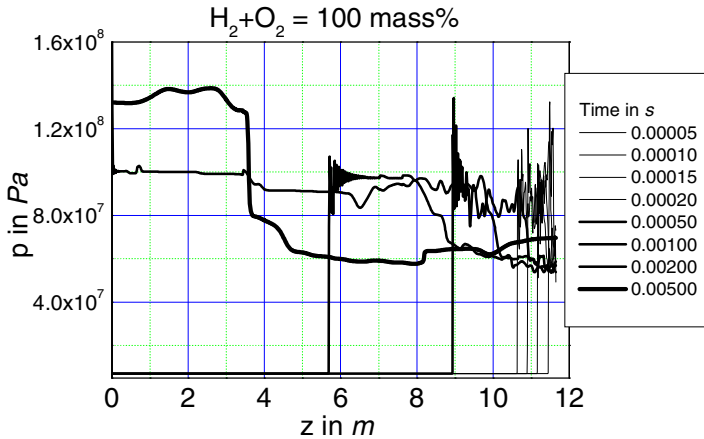


Fig. 9.7 Pressure as a function of distance for different times after the ignition

## 9.3 Multi-phase flow

### 9.3.1 Continuum sound waves

a) *No mechanical phase coupling (Wallis)*. Let us now examine the similar geometrical case of sound wave propagation in multi-phase flows. The local volume-averaged equations give:

$$d(\alpha_i \rho_i w_i) = 0, \quad (9.40)$$

$$d(\alpha_i \rho_i w_i^2) + \alpha_i dp = 0 \quad \text{or} \quad \rho_i d\left(\frac{1}{2} w_i^2\right) + dp = 0, \quad (9.41)$$

$$d\left[\alpha_i \rho_i w_i \left(h_i + \frac{1}{2} w_i^2\right)\right] = 0 \quad \text{or} \quad d\left(h_i + \frac{1}{2} w_i^2\right) = 0 \quad \text{or} \\ \rho_i dh_i - dp = 0 \quad \text{or} \quad ds_i = 0, \quad (9.42)$$

$$d(\alpha_i \rho_i w_i C_{ii}) = 0 \quad \text{or} \quad dC_{ii} = 0 \quad \text{for} \quad \alpha_i \geq 0, \quad (9.43)$$

$$d(w_i n_i) = 0 \quad \text{for} \quad \alpha_i \geq 0. \quad (9.44)$$

Equation (9.40) says that mass is conserved. Equation (9.41) says that force equals the time rate of change of momentum (neglecting body and viscous sources).

Equation (9.42) says that energy is conserved (neglecting all dissipative effects such as work done by viscous stress, and heat transfer due to thermal conduction and diffusion). Beyond *Wallis* (1969), we also take into account Eqs. (9.43) and (9.44). Equation (9.43) says that the species mass is conserved, neglecting diffusion inside the shock front. Equation (9.44) says that there is no fragmentation across the shock. Differentiating the mass conservation equations using the chain rule, dividing by  $\rho_l w_l$ , summing the resulting equations and allowing for the fact that the sum of the volume fraction is constant we then obtain

$$\sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l} d\rho_l + \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{w_l} dw_l = 0, \quad (9.45)$$

Using the equation of state

$$d\rho_l = \frac{dp}{a_l^2} + \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, \text{all } C\text{'s}} ds_l + \sum_{i=1, \dots, i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{il}} \right)_{s_l, p} dC_{il}, \quad (9.46)$$

where

$$\frac{1}{a_l^2} = \left( \frac{\partial \rho_l}{\partial p} \right)_{s_l, \text{all } C\text{'s}} \quad (9.47)$$

and taking into account the last form of Eqs. (9.42) and (9.43) we obtain

$$dp \left( \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} \right) + \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{w_l} dw_l = 0. \quad (9.48)$$

Excluding the velocity differential by using the momentum equation we obtain

$$\left( \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} \right) \left( 1 - \frac{\sum_{l=1}^{l_{\max}} \alpha_l}{\sum_{l=1}^{l_{\max}} \rho_l w_l^2} \right) dp = 0. \quad (9.49)$$

This equation says that for pressure perturbations that are not equal to zero the following condition must be satisfied:

$$\sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l w_l^2} = \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2}. \quad (9.50)$$

This result was obtained by *Wallis* (1969) p.142 in 1969 and is applicable to stratified flows *without strong coupling* between the phases and *no interfacial heat and mass transfer*. Equation (9.50) does not give any information on the magnitude of the phase velocities. If one assumes, that the velocity ratio for stratified flow is controlled by the two independently written momentum equations

without friction resulting in  $S = w_1/w_2 = \sqrt{\rho_2/\rho_1}$ , then the vapor velocity satisfying Eq. (9.50) is

$$w_1 = \sqrt{\left(\frac{\alpha_1}{\rho_1} + \frac{1-\alpha_1}{\rho_2/S^2}\right) / \left(\frac{\alpha_1 a_1^2}{\rho_1} + \frac{1-\alpha_1}{\rho_2 a_2^2}\right)}.$$

b) *Homogeneous multiphase flow (Wood)*. Note that the sound velocity in a multi-phase mixture is flow-pattern dependent and may be influenced by the amplitude and frequency. Let us recall the case for homogeneous multi-phase flow. In this case the sum of the momentum equations, the so called mixture momentum equation, is used instead of the separated momentum equations.

$$d(\rho_h w_h^2) + dp = 0 \quad \text{or} \quad dw_h = -\frac{1}{\rho_h w_h} dp, \quad (9.51)$$

where

$$\rho_h = \sum_{l=1}^{l_{\max}} \alpha_l \rho_l \quad (9.52)$$

is the mixture density. Equation (9.48) simplifies in this case to

$$dp \left( \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} \right) + \frac{1}{w_h} dw_h = 0. \quad (9.53)$$

By inserting the velocity differential from the mixture momentum equation, we obtain

$$\left[ \left( \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} \right) - \frac{1}{\rho_h w_h^2} \right] dp = 0. \quad (9.54)$$

For disturbances of small pressure amplitude we have the velocity of sound

$$w_h = \frac{1}{\sqrt{\rho \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2}}}. \quad (9.55)$$

This is the so called homogeneous mixture sound velocity. For two-phase flow this equation was first derived by *Wood* in 1930. It is widely accepted to be valid for bubbly flow, see the discussion by *Wallis* in 1969.

### 9.3.2 Discontinuous shock waves

We assume that the thickness of the shock front is small and that essentially no fragmentation, interfacial heat, or mass transfer take place. The control volume balance is, therefore,

$$\Delta(\alpha_l \rho_l w_l) = 0, \quad (9.56)$$

$$\Delta(\alpha_l \rho_l w_l^2) + \alpha_l \Delta p = 0, \quad (9.57)$$

$$\Delta \left[ \alpha_l \rho_l w_l \left( h_l + \frac{1}{2} w_l^2 \right) \right] = 0, \quad (9.58)$$

$$\Delta(\alpha_l \rho_l w_l C_{il}) = 0 \quad \text{for } \alpha_l \geq 0, \quad (9.59)$$

$$\Delta(w_l n_l) = 0 \quad \text{for } \alpha_l \geq 0. \quad (9.60)$$

*Homogeneous multiphase flow:* The problem is considerably simplified by the assumption that all the velocities are equal,

$$w_l = w_h. \quad (9.61)$$

The mass conservation for a single velocity field is then

$$\alpha_{l,2} \rho_{l,2} w_{h,2} - \alpha_{l,1} \rho_{l,1} w_{h,1} = 0 \quad \text{or} \quad \frac{\alpha_{l,2} \rho_{l,2}}{\rho_{h,2}} - \frac{\alpha_{l,1} \rho_{l,1}}{\rho_{h,1}} = 0 \quad \text{or} \\ x_{l,2} - x_{l,1} = 0, \quad (9.62)$$

which says that the field mass concentration remains constant across the shock front. This is a remarkable equation. It indicates that such calculations can be readily performed if one uses the following definition for the mass concentrations:

$$x_l = \frac{\alpha_l \rho_l}{\rho_h}. \quad (9.63)$$

The homogeneous density can thus be expressed as

$$\rho_h = \frac{1}{\sum_{l=1}^{l_{\max}} \frac{x_l}{\rho_l}}. \quad (9.64)$$

The mixture mass conservation results in

$$\rho_{h,1} w_{h,1} = \rho_{h,2} w_{h,2} = \rho_h w_h. \quad (9.65)$$

Using the mixture mass conservation, the mixture momentum equation can then be transformed into

$$\rho_{h,2} w_{h,2}^2 - \rho_{h,1} w_{h,1}^2 + p_2 - p_1 = 0 \quad \text{or} \quad (\rho_h w_h)^2 = \frac{p_2 - p_1}{\frac{1}{\rho_{h,1}} - \frac{1}{\rho_{h,2}}}. \quad (9.66)$$

The energy conservation for each velocity field is then

$$\alpha_{i,2} \rho_{i,2} w_{h,2} \left( h_{i,2} + \frac{1}{2} w_{h,2}^2 \right) - \alpha_{i,1} \rho_{i,1} w_{h,1} \left( h_{i,1} + \frac{1}{2} w_{h,1}^2 \right) = 0. \quad (9.67)$$

After applying the field-mass and the mixture-mass conservation equation and using the last form of the momentum equation, the single-field energy conservation that results is as follows:

$$h_{i,2} - h_{i,1} - \frac{1}{2} (p_2 - p_1) \left( \frac{1}{\rho_{h,1}} + \frac{1}{\rho_{h,2}} \right) = 0. \quad (9.68)$$

The assumption that multiphase flow consists of steam and water, which are in thermodynamic equilibrium, reduces Eq. (9.66) to the equations derived by *Fischer* in 1967.

From Eq. (7.52) we obtain the capability of the wave to perform technical work

$$-w_{i,2}^{vdp} = \sum_{l=1}^{l_{\max}} x_l \left[ h_{i,2} - h_{i,1} + T_{\infty} (s_{i,1} - s_{i,2}) \right]. \quad (9.69)$$

The velocity of discontinuum shock waves for homogeneous flow without mass exchange is then

$$w_{h,1} - w_{h,2} = -\sqrt{(p_2 - p_1) \left( \frac{1}{\rho_{h,1}} - \frac{1}{\rho_{h,2}} \right)},$$

For such flow  $X_1 = \text{const}$ . Assuming isentropic change of state of the gas and setting  $\rho_2 = \text{const}$  results for the mixture density in  $\frac{1}{\rho_{h,2}} = \frac{X_1}{\rho_{1,1} (p_2/p_1)^{1/\kappa}} + \frac{1-X_1}{\rho_2}$ .

Replacing and using the definitions  $X_1 = \alpha_{1,1} \rho_{1,1} / \rho_{h,1}$ ,  $\frac{1}{\rho_{h,1}} = \frac{X_1}{\rho_1} + \frac{1-X_1}{\rho_2}$  results in the simple expressions



$$\frac{\Delta w_h^2 \rho_{h,1}}{p_1 \alpha_{1,1}} = \left( \frac{p_2}{p_1} - 1 \right) \left( 1 - \frac{1}{(p_2/p_1)^{1/\kappa}} \right)$$

and

$$\frac{w_{h,1}^2}{p_1 / (\alpha_{1,1} \rho_{h,1})} = \frac{p_2/p_1 - 1}{1 - (p_1/p_2)^{1/\kappa}}.$$

The assumption  $\rho_{h,1} \approx \rho_2$ , valid for very small void fractions, reduces the above equation to Eq. (4) by *van Vijngaarden* (1971), which compares well with the experimental data for air–water shock wave propagation at initially subatmospheric pressure given in Table 9.8.

**Table 9.8** Shock wave velocity as a function of the pressure ratio (*van Vijngaarden* 1971), data for  $p_1 = 1$  bar

Exp. no.	$p_2/p_1$	$R_1$ mm	$\alpha_1$ %	Shock vel., front m/s
1	1.21	1.06	1.38	104.2
2	1.21	0.9	0.74	142.9
3	1.30	1.1	1.68	98
4	1.30	0.98	1.1	125
5	1.30	0.9	0.74	156.2
6	1.40	1.25	2.04	83.3
7	1.40	0.98	1.4	104.2
8	1.40	0.95	0.96	128.2
9	1.56	1.23	2.33	75.8
10	1.56	1	1.68	84.7
11	1.56	1	1.33	100
12	1.70	1.35	2.75	68.5
13	1.70	1.08	1.97	82
14	1.70	1	1.67	92.6
15	1.93	1.35	3.1	61
16	1.93	1.09	2.25	71.4
17	1.93	1	1.82	80.6
18	2.23	1.4	3.55	56.8
19	2.23	1.12	2.62	66.7
20	2.23	1.04	2.19	72.5
21	2.54	1.45	4.16	53.2
22	2.54	1.13	3.15	64.1
23	2.54	1	1.47	94.3
24	3.06	1.5	4.69	51

25	3.06	1.05	3.05	64.9
26	3.06	1	1.91	84.7
27	3.79	1.56	6.04	46.3
28	3.79	1.1	3.45	59.5
29	3.79	1	2.48	71.4
30	4.95	1.54	5.81	48.1
31	4.95	1.3	4.8	53.2
32	4.95	0.9	3.43	61.7

### 9.3.3 Melt-coolant interaction detonations

In the melt-coolant interaction, molten material is premixed with liquid that is in thermal nonequilibrium before the shock wave (Fig. 9.8).

The melt is in the state of film boiling. The incoming shock wave causes instabilities and fragments the  $f_3$  part of the melt mass. The microscopic particles generated release their thermal energy by direct contact with the  $f_2$  part of the surrounding coolant mass. This interaction is a short distance interaction.  $1 - f_2$  of the

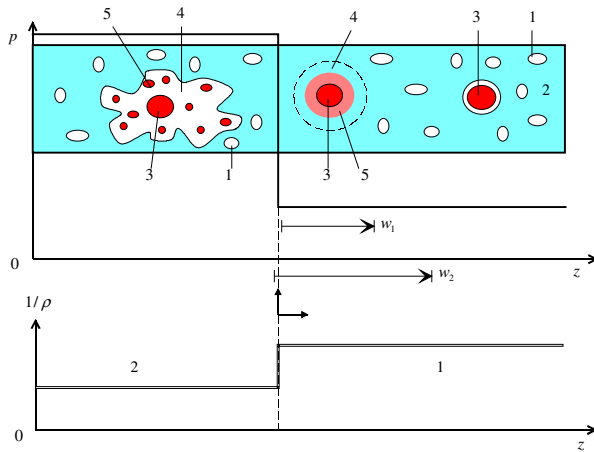


Fig. 9.9 Definitions of the velocity field in the five-field concept

coolant mass outside the short distance remains unaffected by this interaction. The physical problem can thus be specified using the following quantities:

$$p_1, T_{1,1}, T_{2,1}, T_{3,1}, x_1, x_3, \tag{9.50}$$

where the mass concentrations in the undisturbed mixture are

$$\text{vapor } x_1, \tag{9.71}$$

$$\text{liquid } x_2 = 1 - x_1 - x_3, \tag{9.72}$$

$$\text{melt} \quad x_3, \quad (9.73)$$

and the mass concentrations of the newly born fields behind the front are

$$\text{short distance liquid} \quad x_4 = f_2 x_2 = f_2 (1 - x_1 - x_3), \quad (9.74)$$

and

$$\text{microparticles} \quad x_5 = f_3 x_3. \quad (9.75)$$

Here  $f_2$  indicates the fraction of  $x_2$  that is within the short distance liquid.  $f_3$  is the fraction of the melt that is finely fragmented. These fractions are controlled by complex interactions that are not considered in detail in this study. We use the two fractions as parameters in our analysis.

One very useful simplification for such analyses was proposed by *Shamoun* and *Corradini* (1996). The authors proposed that the mixture of the entrained melt intermixed with the entrained liquid reaches thermal equilibrium so that both components have a common pressure  $p_2$  and a common temperature  $(T_m)_2$ . In accordance with this, we seek for the following unknown variables:

$$p_2, T_{1,2}, T_{2,2}, T_{3,2}, T_m, \quad (9.76)$$

which satisfy the *Chapman–Jouguet* condition behind the detonation front. *Shamoun* and *Corradini* (1996) used the assumption that the nonentrained melt and water do not change their initial temperatures and densities, and that the vapor phase experiences an isentropic change of state. These assumptions are in contradiction with the discontinuity of the properties across the shock. We relax these assumptions by using instead the corresponding conservation equations.

The densities and enthalpies required for this computation are functions of the pressure and temperature in front of and behind the shock. The following results are obtained for the vapor

$$\begin{aligned} T_{1,1}, T_{1,2}, (\rho_1)_1 &= \rho_1(p_1, T_{1,1}), (\rho_1)_2 = \rho_1(p_2, T_{1,2}), (h_1)_1 = h_1(p_1, T_{1,1}), \\ (h_1)_2 &= h_1(p_2, T_{1,2}), \end{aligned} \quad (9.77)$$

for the liquid, we have

$$\begin{aligned} T_{2,1}, T_{2,2}, (\rho_2)_1 &= \rho_2(p_1, T_{2,1}), (\rho_2)_2 = \rho_2(p_2, T_{2,2}), (h_2)_1 = h_2(p_1, T_{2,1}), \\ (h_2)_2 &= h_2(p_2, T_{2,2}), \end{aligned} \quad (9.78)$$

for the melt, we have

$$T_{3,1}, T_{3,2}, (\rho_3)_1 = \rho_3(p_1, T_{3,1}), (\rho_3)_2 = \rho_3(p_2, T_{3,2}), (h_3)_1 = h_3(p_1, T_{3,1}),$$

$$(h_3)_2 = h_3(p_2, T_{3,2}), \tag{9.79}$$

for the short distance entrained liquid, we have

$$\begin{aligned} T_{2,1}, T_{1,2}, (\rho_4)_1 &= \rho_2(p_1, T_{2,1}), (\rho_4)_2 = \rho_1(p_2, T_m), (h_4)_1 = h_2(p_1, T_{1,2}), \\ (h_4)_2 &= h_1(p_2, T_m), \end{aligned} \tag{9.80}$$

and for the microparticles, we have

$$\begin{aligned} T_{3,1}, T_{1,2}, (\rho_5)_1 &= \rho_3(p_1, T_{3,1}), (\rho_5)_2 = \rho_3(p_2, T_m), (h_5)_1 = h_3(p_1, T_{3,1}), \\ (h_5)_2 &= h_3(p_2, T_m). \end{aligned} \tag{9.81}$$

We assume that the pressure wave first causes microfragmentation and acceleration of the fine particles into the short distance liquid. The shock discontinuity then occurs. In this case, the mass conservation equation (9.62) holds. The homogeneous mixture densities are then

$$\rho_{h,1} = \frac{1}{\frac{x_1}{\rho_1(p_1, T_{1,1})} + \frac{1-x_1-x_3}{\rho_2(p_1, T_{2,1})} + \frac{x_3}{\rho_3(p_1, T_{3,1})}}, \tag{9.82}$$

$$\rho_{h,2} = \left[ \frac{x_1}{\rho_1(p_2, T_{1,2})} + \frac{(1-f_2)(1-x_1-x_3)}{\rho_2(p_2, T_{2,2})} + \frac{(1-f_3)x_3}{\rho_3(p_2, T_{3,2})} + \frac{f_2(1-x_1-x_3)}{\rho_1(p_2, T_m)} + \frac{f_3x_3}{\rho_3(p_2, T_m)} \right]^{-1}. \tag{9.83}$$

Then the detonation adiabat for each phase is as follows:

1) primary steam

$$(h_1)_2 = h_1(p_1, T_{1,1}) + \Delta h, \tag{9.84}$$

where

$$\Delta h = \frac{1}{2}(p_2 - p_1) \left( \frac{1}{\rho_{h,1}} + \frac{1}{\rho_{h,2}} \right), \tag{9.85}$$

$$h_1(p_2, T_{1,2}) = (h_1)_2 \quad \text{or} \quad T_{1,2} = T_1[p_2, (h_1)_2], \tag{9.86}$$

$$(\rho_1)_2 = \rho_1(p_2, T_{1,2}). \tag{9.88}$$

## 2) nonentrained liquid

$$(h_2)_2 = h_2(p_1, T_{2,1}) + \Delta h, \quad (9.88)$$

$$h_2(p_2, T_{2,2}) = (h_2)_2, \quad (9.89)$$

$$(\rho_2)_2 = \rho_2(p_2, T_{2,2}). \quad (9.90)$$

## 3) nonentrained melt

$$(h_3)_2 = h_3(p_1, T_{3,1}) + \Delta h, \quad (9.91)$$

$$h_3(p_2, T_{3,2}) = (h_3)_2, \quad (9.92)$$

$$(\rho_3)_2 = \rho_3(p_2, T_{3,2}). \quad (9.93)$$

## 4) entrained liquid

$$(h_4)_2 = h_2(p_1, T_{2,1}) + \Delta h, \quad (9.94)$$

$$h_1(p_2, T_m) = (h_4)_2, \quad (9.95)$$

$$(\rho_4)_2 = \rho_1(p_2, T_m). \quad (9.96)$$

## 5) entrained melt

$$(h_5)_2 = h_3(p_1, T_{3,1}) + \Delta h, \quad (9.97)$$

$$h_3(p_2, T_m) = (h_5)_2, \quad (9.98)$$

$$(\rho_5)_2 = \rho_3(p_2, T_m). \quad (9.99)$$

The definition of the homogeneous velocity of sound is rewritten in terms of the mass concentrations to give

$$\frac{1}{(\rho w_h)^2} = \sum_{l=1}^{l_{\max}} \frac{x_l}{(\rho_l a_l)^2}. \quad (9.100)$$

The *Chapman–Jouguet* condition is thus defined as

$$\frac{1}{(\rho_h w_h)_2^2} = \frac{1}{(\rho_h w_h)^2} = \frac{\frac{1}{\rho_{h,1}} - \frac{1}{\rho_{h,2}}}{p_2 - p_1} \quad (9.101)$$

or

$$\sum_{l=1}^{l_{\max}} \frac{x_l}{(\rho_l a_l)_2} = \frac{1}{p_2} - \frac{1}{p_1}, \quad (9.102)$$

or

$$\begin{aligned} & \frac{x_1}{[\rho_1(p_2, T_{1,2})a_1(p_2, T_{1,2})]^2} + \frac{(1-f_2)(1-x_1-x_3)}{[\rho_2(p_2, T_{2,2})a_2(p_2, T_{2,2})]^2} \\ & + \frac{(1-f_3)x_3}{[\rho_3(p_2, T_{3,2})a_3(p_2, T_{3,2})]^2} + \frac{f_2(1-x_1-x_3)}{[\rho_1(p_2, T_m)a_1(p_2, T_m)]^2} \\ & + \frac{f_3x_3}{[\rho_3(p_2, T_m)a_3(p_2, T_m)]^2} = \frac{1}{p_2} - \frac{1}{p_1}. \end{aligned} \quad (9.103)$$

The temperature of the mixture consisting of entrained melt and coolant must satisfy the sum of the detonation equations

$$\begin{aligned} & f_2(1-x_1-x_3)(h_4)_2 + f_3x_3(h_5)_2 \equiv f_2(1-x_1-x_3)h_1(p_2, T_m) + f_3x_3h_3(p_2, T_m) \\ & = \Delta h_{45,2}, \end{aligned} \quad (9.104)$$

where

$$\Delta h_{45,2} = f_2(1-x_1-x_3)h_2(p_1, T_{2,1}) + f_3x_3h_3(p_1, T_{3,1}) + [f_2(1-x_1-x_3) + f_3x_3]\Delta h. \quad (9.105)$$

Large coolant entrainment leads to quenching of the entrained melt and, therefore, makes detonation impossible. Theoretically, the coolant entrainment ratio should be smaller than a prescribed value dictated by the minimum microinteraction mixture temperature leading to the expansion:

$$f_2 < f_3 \frac{x_3}{1-x_1-x_3} \frac{h_3(p_1, T_{3,1}) + \Delta h - h_3(p_2, T_{m,\min})}{h_1(p_2, T_{m,\min}) - [h_2(p_1, T_{2,1}) + \Delta h]},$$

where  $T_{m,\min} = T'(p_2)$  for  $p_2 < p_c$  and  $T_{m,\min} = T_c$  for  $p_2 \geq p_c$ . The temperature inversion can be performed using *Newton's* iteration method

$$T_m^{n+1} = T_m^n - \frac{f_2(1-x_1-x_3)h_1(p_2, T_m^n) + f_3x_3h_3(p_2, T_m^n) - \Delta h_{45,2}}{f_2(1-x_1-x_3)c_{p,1}(p_2, T_m^n) + f_3x_3c_{p,3}(p_2, T_m^n)}. \quad (9.106)$$

The superscripts  $n$  and  $n+1$  denote the old and the new iteration values, respectively. Similarly for the vapor, nonentrained liquid, and nonentrained melt, we have

$$T_{1,2}^{n+1} = T_{1,2}^n - \frac{h_1(p_2, T_{1,2}^n) - h_1(p_1, T_{1,1}) - \Delta h}{c_{p,1}(p_1, T_{1,1})}, \quad (9.107)$$

$$T_{2,2}^{n+1} = T_{2,2}^n - \frac{h_2(p_2, T_{2,2}^n) - h_2(p_1, T_{2,1}) - \Delta h}{c_{p,2}(p_1, T_{2,1})}, \quad (9.108)$$

$$T_{3,2}^{n+1} = T_{3,2}^n - \frac{h_3(p_2, T_{3,2}^n) - h_3(p_1, T_{3,1}) - \Delta h}{c_{p,3}(p_1, T_{3,1})}. \quad (9.109)$$

### 9.3.4 Similarity to and differences from the *Yuen* and *Theofanous* formalism

The exact mathematical formalism used by *Yuen* and *Theofanous* is not reported in 1997. It seems that we use very much the same primitive conservation equations as these authors used. They then constructed *numerically* the *Rankine-Hugoniot* shock and detonation adiabat by using the primitive conservation equations and looked for the tangent to the detonation adiabat corresponding to the initial state, the so called *Chapman-Jouguet* point.

Instead, we use the formalism *analytically* transformed to the detonation adiabat for each of the five fields together with the condition that the speed of propagation behind the shock is equal to the homogeneous velocity of sound that corresponds to the local conditions behind the shock. The computer code written by *Huang Hu Lin* to solve this system of nonlinear transcendental equations allows rapid performance of a variety of computations. In addition, the initial vapor mass is compressed by the shock adiabat and the resulting vapor temperature behind the shock is different from that of the mixture of entrained melt and entrained liquid. Results of such computations will be shown in the next section.

### 9.3.5 Numerical solution method

In fact for the five unknowns,  $p_2, T_{1,2}, T_{2,2}, T_{3,2}, T_m$ , we have a system of five nonlinear transcendental equations (9.103) and (9.106-9.109). The system was coupled with the material properties library of the IVA6 computer code

*Kolev* (1999c). This library contains properties for water and water vapor in sub- and supercritical states, properties of different inert gases, properties for ten materials, which can be either solid or two-phase or liquid with iron, uranium dioxide and alumina among them. The equations are solved by iteration as follows:

- a) Assume initial values at the very beginning of the iterations as follows:

$$p_2 = p_1 + \Delta p, \text{ where}$$

$$\Delta p > 0, T_{1,2} = T_{1,1} \left( \frac{p_2}{p_1} \right)^{\frac{\kappa_1 - 1}{\kappa_1}}, T_{2,2} = T_{2,1}, T_{3,2} = T_{3,1}, T_m = T_{1,1}.$$

- b) Compute the densities behind the front by using the state equations.  
 c) Compute the pressure jump across the shock from the multi-phase *Chapman-Jouguet* condition.  
 d) Compute  $\Delta h$  using Eq. (9.85).  
 e) Compute the temperatures using Eqs. (9.106) to (9.109).  
 f) Repeat steps c) through d) until the change in the pressure jump from iteration to iteration becomes smaller than the prescribed.

**Conclusions.** The following conclusions can be drawn from Sections 9.2 and 9.3. The maximum of the pressure jump for multiphase detonations in melt-water systems can be described by using the assumption of homogeneous flows and relying on the *Rankine-Hugoniot* relations for multiple concentration fields having the same speed. The resulting transcendental system of algebraic equations has to be solved by iteration. Several assumptions with respect of the part of the fragmented particles and of the part of the entrained coolant in the interacting zone are required to close the system. Parametric study can provide the upper limit of the produced detonation shock waves which can be used in estimation of the explosive potential of mixtures of different material pairs.

Such a study will be presented in the next section. In Section 9.4 we present a parametric detonation study for the systems of  $UO_2$ -water,  $Al_2O_3$ -water and  $Fe$ -water. We will draw the conclusion that at the same initial temperatures, mass concentrations, entrainment factors and local volume fractions, the  $Al_2O_3$ -water system produces the highest detonation pressures, and that the dispersed systems consisting of melted particles and water droplets can create detonation pressures less than 200bar.



## 9.4 Detonation waves in water mixed with different molten materials

In Sections 9.2 and 9.3, see also *Kolev* (2000), we recalled the basics of the detonation theory for single- and multi-phase flows. We ended up with a model for detonation analysis for melt-coolant interaction systems. In this section, see also *Kolev* and *Hulin* (2001), we will apply the detonation theory of melt-water interaction to gain information required for an upper bound estimate of the energetic interactions for the nuclear industry. We will consider melt initially mixed with water that is in saturation at atmospheric pressure. In all examples the initial temperature of the melt is considered to be 3000 K. We will study the influence of the following parameters on the detonation pressure of self-sustaining waves: initial void fraction, initial mass concentration of the melt, melt entrainment ratio, water entrainment ratio, different melt material, etc.

In particular, we will answer the following questions:

- a) How can we judge the risk potential of dispersed-dispersed systems for in-vessel and ex-vessel analysis?
- b) Are there any differences between the detonation behavior of the following three pairs of materials:  $UO_2$  water,  $Al_2O_3$  water and  $Fe$  water that result from the detonation theory?
- c) Do detonation solutions exist for lean mixtures with a limited amount of water entrained?
- d) What is the maximum water entrainment ratio that allows detonation solutions for the above mentioned three material pairs and what are the associated maximum pressures for these solutions?

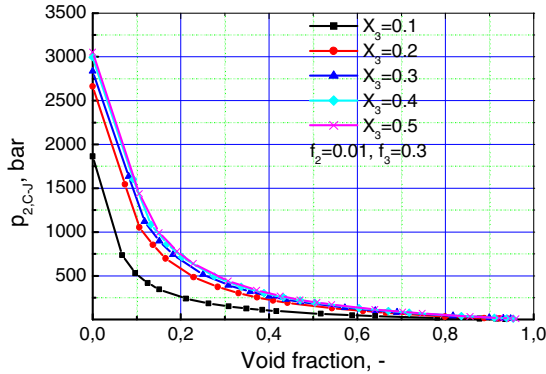
Note that postulating a given set of initial conditions that theoretically lead to detonations does not necessarily mean that such an event can really take place. The present analysis provides only an upper bound of solution sets.

### 9.4.1 $UO_2$ water system

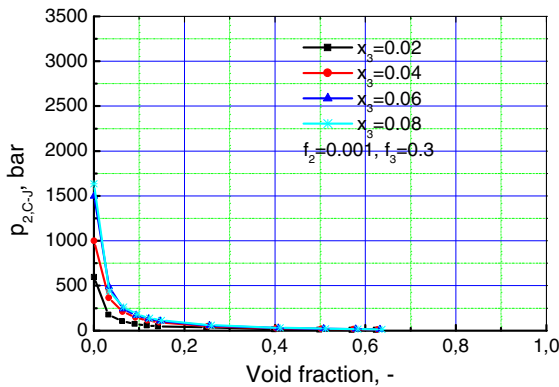
Consider a mixture consisting of uranium dioxide and water. We will represent the detonation pressure behind a self-sustaining shock as a function of the void volume fraction rather than as a function of the mixture specific volume as is usually done in the literature. This is due to our aim of deriving conclusions for systems for which the local void fraction controls the flow pattern. Figure 9.9 shows the pressure behind the self-sustained detonation wave as a function of vapor volume fraction for different mass concentrations of  $UO_2$ . The C-J pressures sharply increase with decreasing vapor volume fraction. Increasing the melt mass concentration  $x_3$  up to about 30% leads to an increase in the C-J pressure. Figure 9.10 demonstrates the behavior of the lean mixtures – mixtures having small amounts of melt. One sees

that even a very small amount of entrained water may cause considerable detonation pressures at low void fractions. This result confirms the warning expressed by *Yuen* and *Theofanous* to consider lean mixtures as potentially explosive.

We also see that lean systems are very sensitive to void fraction. For void fractions larger than 10% the lean mixture can in fact detonate but the detonation pressure has a low risk potential. Homogeneous low void fraction



**Fig. 9.10** The C-J pressure behind the detonation wave as a function of the volume fraction of vapor for large melt mass concentrations of  $UO_2$  (rich mixtures)

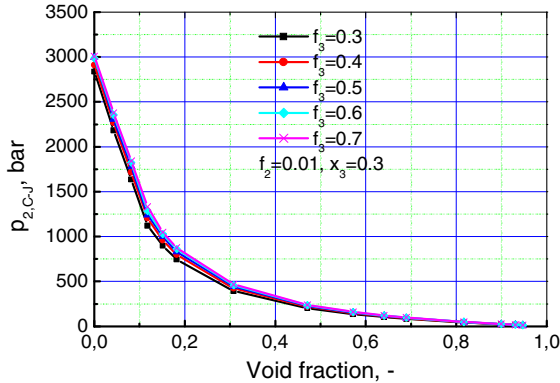


**Fig. 9.11** The C-J pressure behind the detonation wave as a function of the volume fraction of vapor for small melt mass concentrations of  $UO_2$  (lean mixtures)

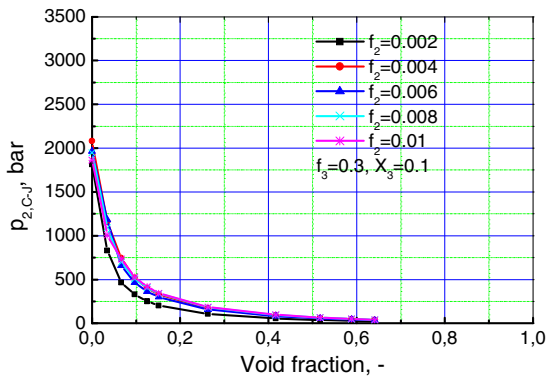
and low water fraction mixtures are in fact possible only in the case of fast transient water entrapment or water injection into the melt. Such systems possess very high detonation pressures. That is the reason why such systems, despite the fact that their existence is only locally possible, are real triggers in nature. Systems not having depletion paths may realize such states locally.

Large-scale existence of such mixtures is practically impossible in open systems having depletion paths for the participating materials. More discussion of this point is provided in *Henry and Fauske (1981a, b)*.

Figures 9.11 and 9.12 show the pressure as a function of the volume fraction for different entrained melt and liquid fractions, respectively. In Fig. 9.11 the



**Fig. 9.12** The C-J pressure behind the detonation wave as a function of the initial volume fraction of vapor. Parameter: melt entrained fraction  $f_3$  of  $UO_2$



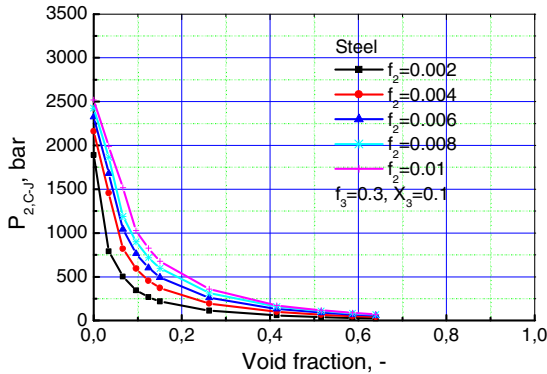
**Fig. 9.13** The C-J pressure behind the shock wave as a function of the initial volume fraction of vapor. Parameter: liquid entrained fraction  $f_2$  for  $UO_2$

entrained coolant controls the pressure and the amount of entrained melt does not influence the process for fixed coolant entrainment.

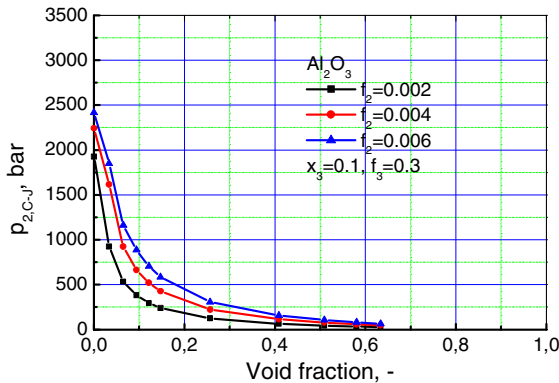
Figure 9.12 shows that at fixed melt entrainment an increase of the coolant entrainment up to a given maximum value may increase the detonation pressure.

Above this value we have predominant by fine debris quenching, and detonation solutions can no longer be obtained.

From this analysis we learn the following for *dispersed* systems: detonation solutions exist theoretically for mixtures consisting of *dispersed* melt and



**Fig. 9.14** The steel C-J pressure behind the shock wave as a function of the initial volume fraction of vapor. Parameter: liquid entrained fraction  $f_2$



**Fig. 9.15** The alumina C-J pressure behind the shock wave as a function of the initial volume fraction of vapor. Parameter: liquid entrained fraction  $f_2$

*dispersed* water. There is, however, no mechanism that would explain the degree of melt fragmentation and water entrainment required for these solutions. As a result, such mixtures cannot be considered to have explosive risk potential. Even assuming a hypothetical, as yet unknown, mechanism that may lead to the

required fragmentation, the resulting detonation pressures are of no concern for the so called in-vessel analysis in nuclear reactor safety.

*Comparison of the detonation behavior of different material pairs:* Here, we keep the parameter the same as for Fig. 9.5 but change only the melt material. The computational results for steel are presented in Fig. 9.13 and for alumina in Fig. 9.14. Comparing Figs. 9.5, 9.13 and 9.14 we come to a very interesting conclusion: at the same initial conditions the steel and the alumina possess the potential to create stronger detonations than the uranium dioxide. This is a fact known from experiments e.g. KROTOS, *Huhtniemi, Magalon and Hohmann (1977)* and the discussion by *Kolev (1999b)*. This means that it is solely the differences in the calorific thermal state properties (an example is given in Appendix 9.1 where the specific capacities at constant pressure for urania and zirconia as functions of the temperature are plotted) for the different materials that make any difference to the achievable detonation pressure under similar conditions.

### 9.4.2 Efficiencies

We will now compare for the three different pairs the fraction of the energy contained in the melt that is necessary to transport the mixture across the boundary of the control volume moving with the frame of reference,

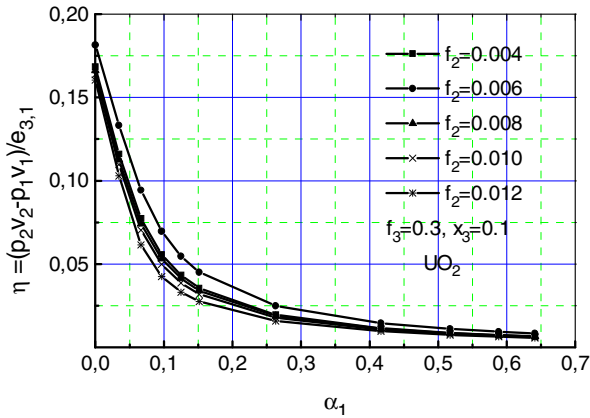
$$\eta_{th} = \frac{P_2 v_2 - P_1 v_1}{e_{3,1}},$$

and the fraction of the thermal energy discharge that is transformed into this work

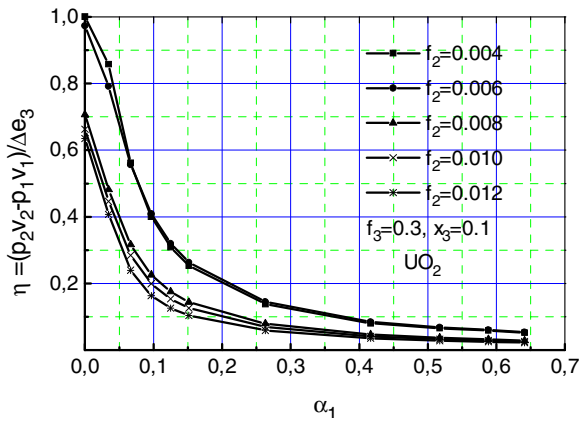
$$\eta = \frac{P_2 v_2 - P_1 v_1}{e_{3,1} - f_3 e_5 - (1 - f_3) e_{3,2}}.$$

From Figs. 9.15a, 9.16a and 9.17a we see that for all of the material pairs considered here lean mixtures with a melt mass fraction of 10%, and 30% melt entrainment cannot transform more than 20% of their thermal energy into mechanical flow transport energy. Increasing the void in the system reduces this figure dramatically.

From Figs. 9.15b, 9.16b and 9.17b we learn that increasing the entrained coolant may cause quenching and reduce the transformation of all the discharged thermal energy into mechanical flow transport energy. Surprisingly, this effect is not so manifested for the alumina-water system.

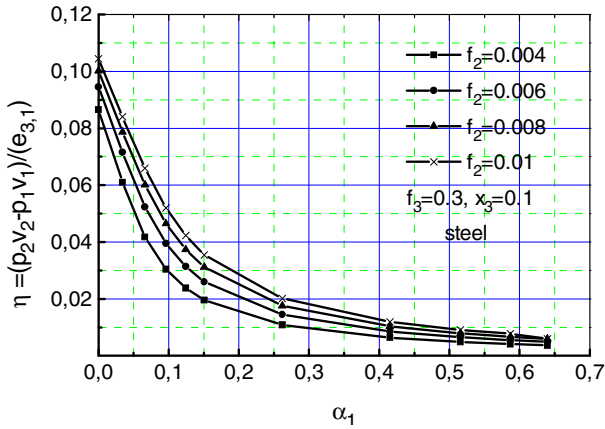


$$a) \eta_{th} = \frac{p_2 v_2 - p_1 v_1}{e_{3,1}}$$

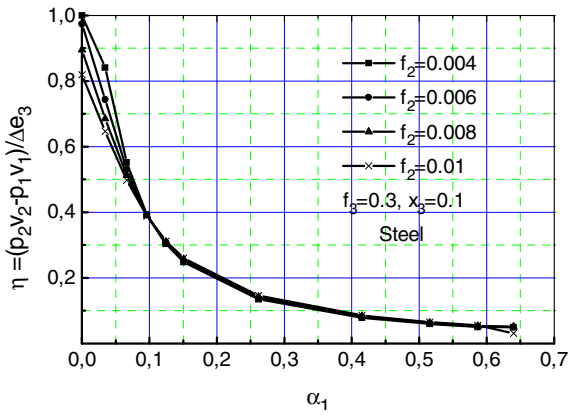


$$b) \eta = \frac{p_2 v_2 - p_1 v_1}{e_{3,1} - f_3 e_5 - (1 - f_3) e_{3,2}}$$

**Fig. 9.16** The two nominal efficiencies of uranium dioxide behind the shock wave as a function of the initial volume fraction of vapor. Parameter: liquid entrained fraction  $f_2$

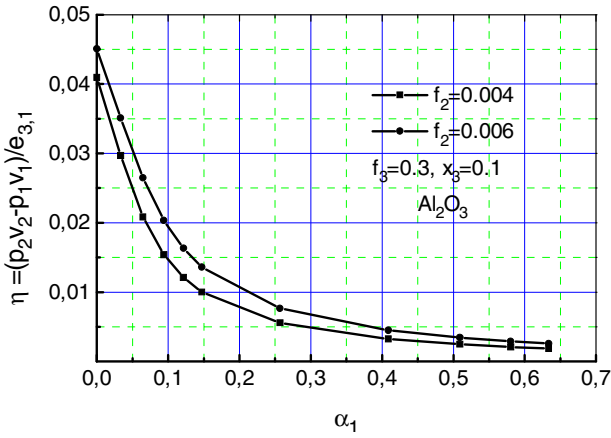


$$\eta_{th} = \frac{p_2 v_2 - p_1 v_1}{e_{3,1}}$$

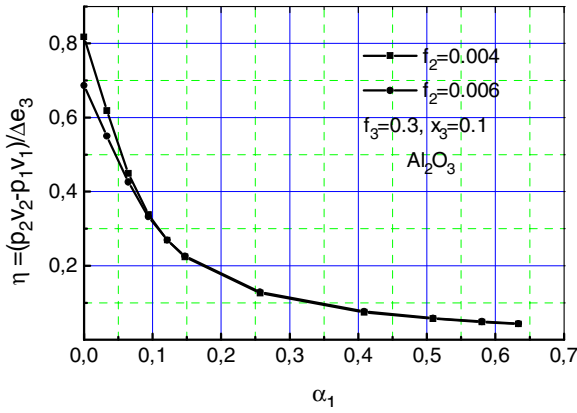


$$b) \eta = \frac{p_2 v_2 - p_1 v_1}{e_{3,1} - f_3 e_5 - (1 - f_3) e_{3,2}}$$

**Fig. 9.17** The two nominal efficiencies of steel behind the shock wave as a function of the initial volume fraction of vapor. Parameter: liquid entrained fraction  $f_2$



$$a) \eta_{th} = \frac{p_2 v_2 - p_1 v_1}{e_{3,1}}$$



$$b) \eta = \frac{p_2 v_2 - p_1 v_1}{e_{3,1} - f_3 e_5 - (1 - f_3) e_{3,2}}$$

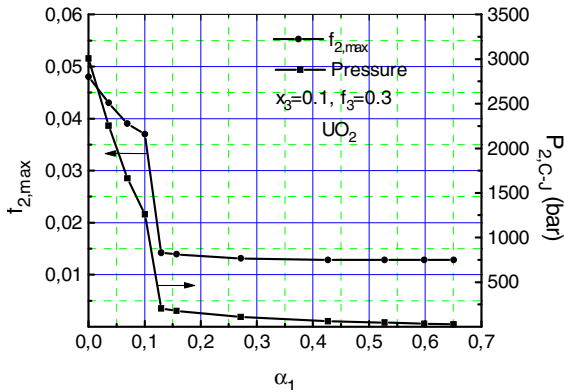
**Fig. 9.18** The two nominal efficiencies of steel behind the shock wave as a function of the initial volume fraction of vapor. Parameter: liquid entrained fraction  $f_2$



### 9.4.3 The maximum coolant entrainment ratio

As already mentioned, even theoretically large coolant entrainment ratios do not result in detonation solutions. In this chapter we compare the three pairs by presenting the maximum of the achievable coolant entrainment ratio that allows self-sustaining detonation waves as a function of the void fractions and the corresponding detonation pressures – Figs. 9.18, 9.19 and 9.20. The jumps in the entrainment ratios in Figs. 9.18 and 9.19 are associated with the latent heat of solidification. The larger water entrainment ratios give higher detonation pressures because the volume of the produced vapor is larger.

Surprisingly, the alumina-water system once again allows the highest water entrainment rates and produces higher detonation pressures at very low void fractions. This is due to the fact that the specific heat of uranium dioxide is less than that of alumina.



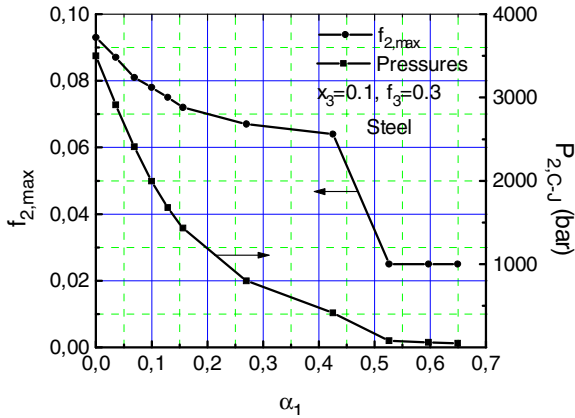
**Fig. 9.19** The maximum water mass entrained fraction and C-J pressure behind the shock wave for uranium dioxide as a function of the initial volume fraction of vapor

Assuming small coolant entrainment leads to large superheating of the evaporated entrained coolant surrounding the melt and therefore a temperature difference between this coolant and the melt limits locally the energy transferred from the melt to the coolant. This means there is also in this sense a natural limit of the energy transfer.

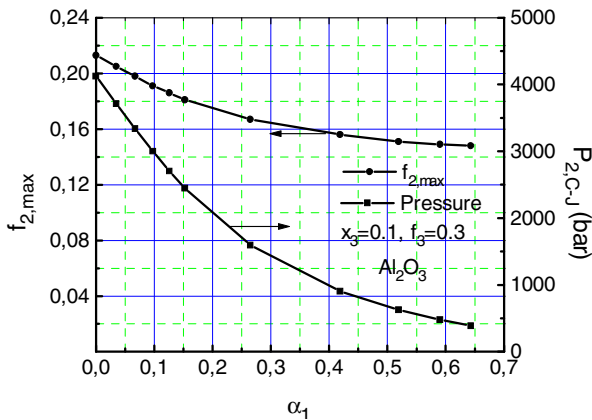
## 9.5 Conclusions

In this chapter we applied the detonation theory in order to learn features of melt-water interaction required for an upper bound estimate of the energetic interactions for the nuclear industry. We found the following results:

- 1) Mixtures consisting of dispersed melt and dispersed water have detonation solutions but there is no mechanism to explain the degree of melt fragmentation and of water entrainment into brought contact with the melt debris. Therefore, such mixtures cannot be considered to have explosive risk potential. Even assuming a hypothetical mechanism that may lead to the required fragmentation, the resulting detonation pressures are of no concern for in-vessel melt-water interaction risk analysis.



**Fig. 9.20** The maximum water mass entrained fraction and C-J pressure behind the shock wave for steel as a function of the initial volume fraction of vapor



**Fig. 9.21** The maximum water mass entrained fraction and C-J pressure behind the shock wave for alumina as a function of the initial volume fraction of vapor

- 2) Even a very small amount of entrained water may cause considerable detonation pressures at low void fractions. This result confirms the warning expressed by *Yuen and Theofanous* to consider lean mixtures as potentially explosive. We also see that lean systems are very sensitive to void fraction.

For void fractions higher than 10% the lean mixture can in fact detonate but the detonation pressure has low risk potential.

- 3) Homogeneous low void fraction and low water fraction mixtures are in fact possible only in the case of fast transient water entrapment or water injection into the melt. Such systems possess very high detonation pressures. That is the reason why such systems, despite the fact that their existence is only locally possible, are real triggers in nature. Large-scale existence of such mixtures is practically impossible in open systems having depletion paths for the participating materials. Systems not having depletion paths may realize such states locally.
- 4) For each entrained melt ratio there is a maximum of entrained coolant ratio that controls the maximum of the detonation pressure.
- 5) For the same initial conditions the detonation pressures for  $Al_2O_3$ -water and  $Fe$ -water systems are definitively larger than that for  $UO_2$ -water systems. The capability for heat discharge during postulated melt-water contact for these systems does not explain the differences in their detonation behavior. As a result, only the differences in the caloric equation of state explain the differences in the detonation behavior. In addition, at the same initial condition,  $UO_2$ -water systems can entrain a lower mass fraction of liquid than the other analyzed material pairs. This system reaches the thermal equilibrium state at a lower mixture temperature than the other analyzed material pairs.
- 6) The efficiency of transformation of thermal energy into mechanical flow transport energy depends strongly a) on the vapor fraction and b) on the entrained water. The higher the void fraction the lower the efficiency of transformation of thermal into mechanical flow transport energy. At levels below some characteristic water entrainment rates for low void fraction mixtures lower efficiency results due to the local non-availability of an adequate amount of water to evaporate. The theoretical maximum of the highest total efficiency of transformation of the all-internal melt energy into mechanical flow transport energy is lower than 20%. This very high figure is achievable in a small-scale laboratory experiment with very low initial void fraction, mainly in cases of entrapment or very fast melt injection into water. That is why we consider this value as a local maximum for trigger efficiencies which cannot be transferred to large scale systems with a variety of depletion paths such as those of in-vessel melt-water interaction.

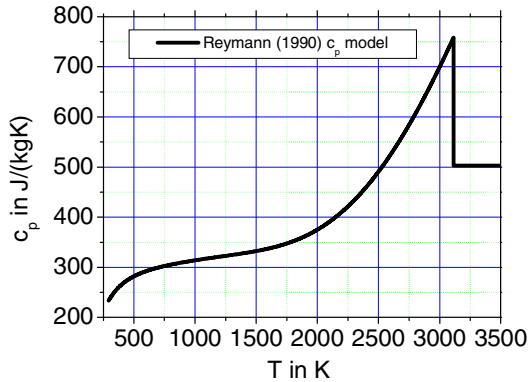
## 9.6 Practical significance

Judging the explosivity of melt-water and vapor mixtures is important for nuclear safety. There is a consensus among researchers that mixtures in which one of the liquids is continuous are explosive. But up to this study *Kolev (2000)*, *Kolev and Hulin (2001)* it was not clear how to judge the risk of mixtures consisting of melt

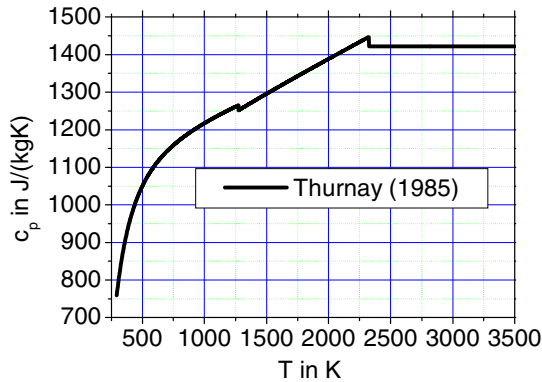
droplets, coolant droplets and gas. The practical outcome of this study is the demonstration that detonation pressure jumps in disperse systems are lower than 200 *bar* without considering the question whether they can be realized in nature or not. This knowledge classified such mixtures as not dangerous at all for in-vessel melt-water interactions.

Another unexpected outcome of this study is the different theoretical behavior of different melt-water pairs for the same initial temperatures, gas volume fractions and mass concentrations of the melt and water, resulting only from the difference of their caloric equations of state.

## Appendix 9.1 Specific heat capacity at constant pressure for urania and alumina



**Fig. A9-1** Specific capacity at constant pressure as a function of the temperature of solid and liquid  $\text{UO}_2$ , Reymann (1990)



**Fig. A9-2** Specific capacity at constant pressure as a function of the temperature of solid and liquid alumina, Thurnay (1985)

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# 10 Conservation equations in general curvilinear coordinate systems

## 10.1 Introduction

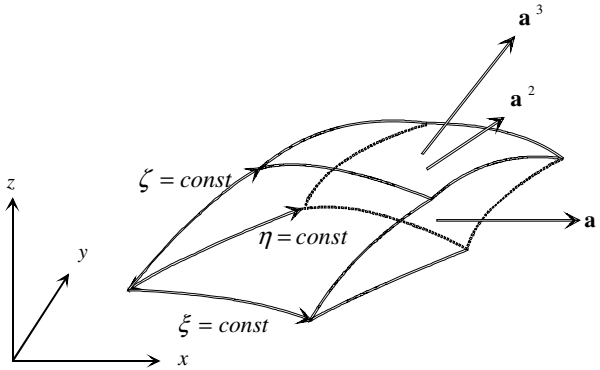
In 1974 *Vivand* and *Vinokur* published their remarkable works on conservation equations of gas dynamics in curvilinear coordinate systems. Since that time there have been many publications on different aspects of this topic. Several providers of computational fluid dynamics tools use the developed strategy for single-phase flows in attempting to extend the algorithm for multi-phase flows. The usually used approach is to write all partial differential equation in convection-diffusion form and to use the already existing transformations and numerical algorithms for single-phase flow. What remains outside the convection-diffusion terms is pooled as a source into the right hand side. The problems with this approach are two: (a) the remaining terms can contain substantial physics represented in differential terms and (b) the realized coupling between the equations is weak. The latter is manifested if one tries to use such codes for processes with strong feedback of the interfacial heat and mass transfer processes on the pressure, e.g. steam explosion, spontaneous flashing etc.

In this Chapter, the conservation equations rigorously derived in Chapters 1, 2, 5 for multi-phase flows based on local time and volume average *Kolev* (1994a, b, 1995, 1997, 1998, 1999) are rewritten in general curvilinear coordinate in order to facilitate their numerical integration for arbitrary geometry. In addition the so called conservation of total volume equation is also transformed. The latter is extremely useful for creating strongly coupled numerical solution algorithms. Note that the derivation in Chapter 1, 2, 5, *Kolev* (1994a, b, 1995, 1997, 1998, 1999), is performed for flows in heterogeneous porous structures. Each of the velocity fields was considered to consist of several chemical components. The concept of dynamic fragmentation and coalescence is used. The derivation given here was first published in *Kolev* (2001).

For the consideration presented below we assume that the flow happens in the physical space described by the Cartesian, left oriented coordinate systems  $(x,y,z)$  (Euclidean space) – Fig. 10.1. In this space the new curvilinear coordinate system is introduced having the coordinates  $\xi, \eta, \zeta$ . Another notation simultaneously used is  $x_i$  ( $i=1,2,3$ ):  $x_1, x_2, x_3$  and  $\xi^i$  ( $i=1,2,3$ ):



$\xi^1, \xi^2, \xi^3$ . The new coordinate system is called the transformed coordinate system. The curves defining the transformed system are smooth (at least one times differentiable). The transformation is invertible. The Jacobian, the metrics and inverted metrics tensors, the covariant and the contravariant vectors exist. Figure 10.1 shows the physical coordinate system, the transformed coordinate system and a control volume in the transformed coordinates system build by coordinate surfaces. The figure shows also the definition of the contravariant vectors perpendicular to the coordinate surfaces. In addition, the curvilinear coordinates system moves with velocity  $\mathbf{V}_{cs}$ .



**Fig. 10.1** The unit normal vectors to the coordinate surfaces form the contravariant base vectors

For better understanding of the material in this Chapter we present in Appendix 1 a brief introduction to vector analysis and in Appendix 2 some basics of the coordinate transformation theory. I strongly recommend the reader to go over these Appendixes before continuing reading this Chapter.

## 10.2 Field mass conservation equations

The local volume- and time-averaged mass conservation equation derived in Chapter 1, Eq. (1.62) *Kolev (1994a)* is

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l \gamma_v) + \nabla \cdot (\alpha_l \rho_l \mathbf{V}_l \gamma) = \gamma_v \mu_l \quad (10.1)$$

First we transform the time derivative having in mind that

$$\left( \frac{\partial \varphi}{\partial \tau} \right)_{x,y,z} = \left( \frac{\partial \varphi}{\partial \tau} \right)_{\xi,\eta,\zeta} - \mathbf{V}_{cs} \cdot \nabla \varphi \quad (10.2)$$

The result is

$$\frac{\partial}{\partial \tau}(\alpha_i \rho_i \gamma_v) + \nabla \cdot (\alpha_i \rho_i \mathbf{V}_i \gamma) - \mathbf{V}_{cs} \cdot \nabla (\alpha_i \rho_i \gamma) = \gamma_v \mu_i. \quad (10.3)$$

Here the time derivative  $(\partial / \partial \tau)_{\xi, \eta, \zeta}$  is understood to be at a fixed point in the transformed region and the subscripts are omitted for simplicity of the notation. Then having in mind that the divergence of a vector and the gradient of a scalar in the transformed region are

$$\nabla \cdot \mathbf{F} = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}) \right], \quad (10.4)$$

$$\nabla \varphi = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \varphi) \right], \quad (10.5)$$

we obtain finally

$$\begin{aligned} & \frac{\partial}{\partial \tau}(\alpha_i \rho_i \gamma_v) \\ & + \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\alpha_i \rho_i \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{V}_i) + \frac{\partial}{\partial \eta} (\alpha_i \rho_i \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{V}_i) + \frac{\partial}{\partial \zeta} (\alpha_i \rho_i \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{V}_i) \right] \\ & - \mathbf{V}_{cs} \cdot \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\alpha_i \rho_i \gamma_\xi \sqrt{g} \mathbf{a}^1) + \frac{\partial}{\partial \eta} (\alpha_i \rho_i \gamma_\eta \sqrt{g} \mathbf{a}^2) + \frac{\partial}{\partial \zeta} (\alpha_i \rho_i \gamma_\zeta \sqrt{g} \mathbf{a}^3) \right] = \gamma_v \mu_i. \end{aligned} \quad (10.6)$$

The term containing the velocity of the transformed coordinate system can be rewritten as follows

$$\begin{aligned} & \mathbf{V}_{cs} \cdot \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\alpha_i \rho_i \gamma_\xi \sqrt{g} \mathbf{a}^1) + \frac{\partial}{\partial \eta} (\alpha_i \rho_i \gamma_\eta \sqrt{g} \mathbf{a}^2) + \frac{\partial}{\partial \zeta} (\alpha_i \rho_i \gamma_\zeta \sqrt{g} \mathbf{a}^3) \right] \\ & = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\alpha_i \rho_i \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{V}_{cs}) + \frac{\partial}{\partial \eta} (\alpha_i \rho_i \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{V}_{cs}) \right. \\ & \quad \left. + \frac{\partial}{\partial \zeta} (\alpha_i \rho_i \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{V}_{cs}) \right] \\ & = -\alpha_i \rho_i \gamma_v \left( \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \xi} + \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \eta} + \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \zeta} \right). \end{aligned} \quad (10.7)$$

Having in mind that

$$\frac{1}{\sqrt{g}} \frac{d\sqrt{g}}{dt} = \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \xi} + \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \eta} + \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \zeta},$$

we obtain after multiplying by  $\sqrt{g}$  and using the reverse chain rule,

$$\sqrt{g} \frac{\partial}{\partial \tau} (\alpha_i \rho_i \gamma_v) + \alpha_i \rho_i \gamma_v \frac{\partial \sqrt{g}}{\partial \tau} = \frac{\partial}{\partial \tau} (\alpha_i \rho_i \gamma_v \sqrt{g}), \quad (10.8)$$

the following final conservative form of the mass conservation equation for multi-phase flows

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_i \rho_i \gamma_v \sqrt{g}) + \frac{\partial}{\partial \xi} [\alpha_i \rho_i \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs})] + \frac{\partial}{\partial \eta} [\alpha_i \rho_i \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs})] \\ & + \frac{\partial}{\partial \zeta} [\alpha_i \rho_i \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs})] = \gamma_v \sqrt{g} \mu_i. \end{aligned} \quad (10.9)$$

We see that the scalar

$$\alpha_i \rho_i \sqrt{g}$$

with the flux

$$\alpha_i \rho_i \sqrt{g} \mathbf{a}^i \cdot (\mathbf{V}_l - \mathbf{V}_{cs})$$

is subject to conservation in the transformed coordinate system. The velocity vector components

$$\bar{V}_l^i = \mathbf{a}^i \cdot (\mathbf{V}_l - \mathbf{V}_{cs})$$

will be called contravariant field relative velocity vector components. It is perpendicular to the coordinate surface defined by  $\xi^i = const$ . Thus the mass conservation equation is simply

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_i \rho_i \gamma_v \sqrt{g}) + \frac{\partial}{\partial \xi} (\alpha_i \rho_i \gamma_\xi \sqrt{g} \bar{V}_l^1) + \frac{\partial}{\partial \eta} (\alpha_i \rho_i \gamma_\eta \sqrt{g} \bar{V}_l^2) + \frac{\partial}{\partial \zeta} (\alpha_i \rho_i \gamma_\zeta \sqrt{g} \bar{V}_l^3) \\ & = \gamma_v \sqrt{g} \mu_i. \end{aligned} \quad (10.10)$$

Remember again that the time derivative is understood to be at a fixed point in the transformed region. Note that setting  $\alpha_i \rho_i \gamma_v = \alpha_i \rho_i \gamma = const$ ,  $\mathbf{V}_l = const$  and  $\mu_i = 0$ , and using the *first fundamental metric identity*, Peyret (1996)

$$\frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3) = 0, \quad (10.11)$$

results in

$$\frac{\partial \sqrt{g}}{\partial \tau} - \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \cdot \mathbf{V}_{cs}) - \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \cdot \mathbf{V}_{cs}) - \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \cdot \mathbf{V}_{cs}) = 0, \quad (10.12)$$

which is the *second fundamental metric identity*, *Thompson et al. (1985)*, p.159. As pointed out by *Thompson et al. (1985)* this identity should be used to numerically determine updated values of the *Jacobian*,  $\sqrt{g}$ , instead of updating it directly from the new values of the Cartesian coordinates. In the later case spurious source terms will appear, *Thompson et al. (1985)*.

### 10.3 Mass conservation equations for components inside the field – conservative form

The conservative form of the local volume- and time-averaged field mass conservation equations for each component inside a velocity field as derived in *Kolev (1994a)* is

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l C_{il} \gamma_v) + \nabla \cdot [\alpha_l \rho_l \gamma (\mathbf{V}_l C_{il} - D_{il}^* \nabla C_{il})] = \gamma_v \mu_{il}. \quad (10.13)$$

The time derivative and the convection term are transformed as in the previous section. For the transformation of the diffusion term we use the *Diffusion Laplacian* in the non-conservative form

$$\begin{aligned} & \sqrt{g} \nabla \cdot (\lambda \nabla \varphi) \\ &= \frac{\partial}{\partial \xi} \left[ \sqrt{g} \mathbf{a}^1 \cdot \lambda \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right] \\ &+ \frac{\partial}{\partial \eta} \left[ \sqrt{g} \mathbf{a}^2 \cdot \lambda \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right] \\ &+ \frac{\partial}{\partial \zeta} \left[ \sqrt{g} \mathbf{a}^3 \cdot \lambda \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right], \end{aligned} \quad (10.14)$$

where  $\varphi = \alpha_l \rho_l \gamma D_{il}^*$  is a scalar valued function of the local flow parameters. With this we obtain the final form which is additionally multiplied by the *Jacobian*

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (\alpha_l \rho_l C_{il} \sqrt{g} \gamma_v) \\
& + \frac{\partial}{\partial \xi} \left\{ \alpha_l \rho_l \sqrt{g} \gamma_\xi \left[ C_{il} \bar{V}_l^1 - D_{il}^* \left( \mathbf{a}^1 \cdot \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^1 \cdot \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^1 \cdot \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \\
& + \frac{\partial}{\partial \eta} \left\{ \alpha_l \rho_l \sqrt{g} \gamma_\eta \left[ C_{il} \bar{V}_l^2 - D_{il}^* \left( \mathbf{a}^2 \cdot \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \cdot \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^2 \cdot \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \\
& + \frac{\partial}{\partial \zeta} \left\{ \alpha_l \rho_l \sqrt{g} \gamma_\zeta \left[ C_{il} \bar{V}_l^3 - D_{il}^* \left( \mathbf{a}^3 \cdot \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^3 \cdot \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \cdot \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \\
& = \gamma_v \sqrt{g} \mu_{il}. \tag{10.15}
\end{aligned}$$

We see that the scalar

$$\alpha_l \rho_l C_{il} \sqrt{g}$$

with the convective flux

$$\alpha_l \rho_l C_{il} \sqrt{g} \mathbf{a}^i \cdot (\mathbf{V}_l - \mathbf{V}_{cs})$$

and a diffusion flux is subject to conservation in the transformed coordinate system. Using the elements of the inverted metric tensor which is symmetric per definition the notation simplifies to

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (\alpha_l \rho_l C_{il} \sqrt{g} \gamma_v) \\
& + \frac{\partial}{\partial \xi} \left\{ \alpha_l \rho_l \sqrt{g} \gamma_\xi \left[ C_{il} \bar{V}_l^1 - D_{il}^* \left( g^{11} \frac{\partial C_{il}}{\partial \xi} + g^{12} \frac{\partial C_{il}}{\partial \eta} + g^{13} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \\
& + \frac{\partial}{\partial \eta} \left\{ \alpha_l \rho_l \sqrt{g} \gamma_\eta \left[ C_{il} \bar{V}_l^2 - D_{il}^* \left( g^{21} \frac{\partial C_{il}}{\partial \xi} + g^{22} \frac{\partial C_{il}}{\partial \eta} + g^{23} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \\
& + \frac{\partial}{\partial \zeta} \left\{ \alpha_l \rho_l \sqrt{g} \gamma_\zeta \left[ C_{il} \bar{V}_l^3 - D_{il}^* \left( g^{31} \frac{\partial C_{il}}{\partial \xi} + g^{32} \frac{\partial C_{il}}{\partial \eta} + g^{33} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} = \gamma_v \sqrt{g} \mu_{il}. \tag{10.16}
\end{aligned}$$

Thus the *isotropic* convection-diffusion problem in the physical space turns out to be a *anisotropic* in the transformed space. One immediately recognizes the advantage of the orthogonal coordinate systems for which only the diagonal elements of the inverse matrices are different from zero

$$\begin{aligned}
& \frac{\partial}{\partial \tau} (\alpha_i \rho_i C_{il} \sqrt{g} \gamma_v) \\
& + \frac{\partial}{\partial \xi} \left[ \alpha_i \rho_i \sqrt{g} \gamma_\xi \left( C_{il} \bar{V}_l^1 - g^{11} D_{il}^* \frac{\partial C_{il}}{\partial \xi} \right) \right] \\
& + \frac{\partial}{\partial \eta} \left[ \alpha_i \rho_i \sqrt{g} \gamma_\eta \left( C_{il} \bar{V}_l^2 - g^{22} D_{il}^* \frac{\partial C_{il}}{\partial \eta} \right) \right] \\
& + \frac{\partial}{\partial \zeta} \left[ \alpha_i \rho_i \sqrt{g} \gamma_\zeta \left( C_{il} \bar{V}_l^3 - g^{33} D_{il}^* \frac{\partial C_{il}}{\partial \zeta} \right) \right] = \gamma_v \sqrt{g} \mu_{il}. \tag{10.17}
\end{aligned}$$

#### 10.4 Field mass conservation equations for components inside the field – non-conservative form

The non-conservative form is obtained by differentiating the time derivative and the convection term and comparing with the mass conservation equation. The result is

$$\begin{aligned}
& \alpha_i \rho_i \sqrt{g} \left( \gamma_v \frac{\partial C_{il}}{\partial \tau} + \gamma_\xi \bar{V}_l^1 \frac{\partial C_{il}}{\partial \xi} + \gamma_\eta \bar{V}_l^2 \frac{\partial C_{il}}{\partial \eta} + \gamma_\zeta \bar{V}_l^3 \frac{\partial C_{il}}{\partial \zeta} \right) \\
& - \frac{\partial}{\partial \xi} \left[ \alpha_i \rho_i \sqrt{g} \gamma_\xi D_{il}^* \left( g^{11} \frac{\partial C_{il}}{\partial \xi} + g^{12} \frac{\partial C_{il}}{\partial \eta} + g^{13} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \\
& - \frac{\partial}{\partial \eta} \left[ \alpha_i \rho_i \sqrt{g} \gamma_\eta D_{il}^* \left( g^{21} \frac{\partial C_{il}}{\partial \xi} + g^{22} \frac{\partial C_{il}}{\partial \eta} + g^{23} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \\
& - \frac{\partial}{\partial \zeta} \left[ \alpha_i \rho_i \sqrt{g} \gamma_\zeta D_{il}^* \left( g^{31} \frac{\partial C_{il}}{\partial \xi} + g^{32} \frac{\partial C_{il}}{\partial \eta} + g^{33} \frac{\partial C_{il}}{\partial \zeta} \right) \right] = \gamma_v \sqrt{g} (\mu_{il} - \mu_l C_{il}). \tag{10.18}
\end{aligned}$$

#### 10.5. Particles number conservation equations for each velocity field

The local volume- and time-averaged particle number density conservation equation for each velocity field derived in *Kolev (1994a)* is

$$\frac{\partial}{\partial \tau} (n_l \gamma_v) + \nabla \cdot \left[ \left( \mathbf{V}_l n_l - \frac{V_l^t}{Sc^t} \nabla n_l \right) \gamma \right] = \gamma_v (\dot{n}_{l,kin} - \dot{n}_{l,coal} + \dot{n}_{l,sp}). \quad (10.19)$$

By analogy to the component mass conservation equation derived in the previous section we have

$$\begin{aligned} & \frac{\partial}{\partial \tau} (n_l \sqrt{g} \gamma_v) \\ & + \frac{\partial}{\partial \xi} \left\{ \sqrt{g} \gamma_\xi \left[ n_l \bar{V}_l^1 - \frac{V_l^t}{Sc^t} \left( g^{11} \frac{\partial n_l}{\partial \xi} + g^{12} \frac{\partial n_l}{\partial \eta} + g^{13} \frac{\partial n_l}{\partial \zeta} \right) \right] \right\} \\ & + \frac{\partial}{\partial \eta} \left\{ \sqrt{g} \gamma_\eta \left[ n_l \bar{V}_l^2 - \frac{V_l^t}{Sc^t} \left( g^{21} \frac{\partial n_l}{\partial \xi} + g^{22} \frac{\partial n_l}{\partial \eta} + g^{23} \frac{\partial n_l}{\partial \zeta} \right) \right] \right\} \\ & + \frac{\partial}{\partial \zeta} \left\{ \sqrt{g} \gamma_\zeta \left[ n_l \bar{V}_l^3 - \frac{V_l^t}{Sc^t} \left( g^{31} \frac{\partial n_l}{\partial \xi} + g^{32} \frac{\partial n_l}{\partial \eta} + g^{33} \frac{\partial n_l}{\partial \zeta} \right) \right] \right\} \\ & = \gamma_v \sqrt{g} (\dot{n}_{l,kin} - \dot{n}_{l,coal} + \dot{n}_{l,sp}). \end{aligned} \quad (10.20)$$

## 10.6 Field entropy conservation equations – conservative form

The conservative form of the local volume- and time-averaged field entropy conservation equation Eq. (5.101b) as derived in Chapter 5, *Kolev* (1994a), is

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l s_l \gamma_v) + \nabla \cdot \left[ \alpha_l^e \rho_l \gamma \left( s_l \mathbf{V}_l - \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \nabla C_{il} \right) \right] - \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \gamma \nabla T_l) \\ & = \gamma_v \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) + \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} \right] \end{aligned} \quad (10.21)$$

for  $\alpha_l \geq 0$ . The entropy conservation equation in the transformed space is therefore

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l s_l \sqrt{g} \gamma_v) \\ & + \frac{\partial}{\partial \xi} \left\{ \alpha_l \rho_l \sqrt{g} \gamma_\xi \left\{ s_l \bar{V}_l^1 - \sum_{i=2}^{i_{\max}} (s_{il} - s_{il}) D_{il}^* \left( g^{11} \frac{\partial C_{il}}{\partial \xi} + g^{12} \frac{\partial C_{il}}{\partial \eta} + g^{13} \frac{\partial C_{il}}{\partial \zeta} \right) \right\} \right\} \end{aligned}$$

$$\begin{aligned}
 & + \frac{\partial}{\partial \eta} \left\{ \alpha_i \rho_i \sqrt{g} \gamma_\eta \left\{ s_i \bar{V}_i^2 - \sum_{i=2}^{i_{\max}} \left[ (s_{il} - s_{l1}) D_{il}^* \left( g^{21} \frac{\partial C_{il}}{\partial \xi} + g^{22} \frac{\partial C_{il}}{\partial \eta} + g^{23} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \right\} \\
 & + \frac{\partial}{\partial \zeta} \left\{ \alpha_i \rho_i \sqrt{g} \gamma_\zeta \left\{ s_i \bar{V}_i^3 - \sum_{i=2}^{i_{\max}} \left[ (s_{il} - s_{l1}) D_{il}^* \left( g^{31} \frac{\partial C_{il}}{\partial \xi} + g^{32} \frac{\partial C_{il}}{\partial \eta} + g^{33} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \right\} \\
 & - \frac{1}{T_l} \frac{\partial}{\partial \xi} \left[ \alpha_i^e \lambda_i^* \sqrt{g} \gamma_\xi \left( g^{11} \frac{\partial T_l}{\partial \xi} + g^{12} \frac{\partial T_l}{\partial \eta} + g^{13} \frac{\partial T_l}{\partial \zeta} \right) \right] \\
 & - \frac{1}{T_l} \frac{\partial}{\partial \eta} \left[ \alpha_i^e \lambda_i^* \sqrt{g} \gamma_\eta \left( g^{21} \frac{\partial T_l}{\partial \xi} + g^{22} \frac{\partial T_l}{\partial \eta} + g^{23} \frac{\partial T_l}{\partial \zeta} \right) \right] \\
 & - \frac{1}{T_l} \frac{\partial}{\partial \zeta} \left[ \alpha_i^e \lambda_i^* \sqrt{g} \gamma_\zeta \left( g^{31} \frac{\partial T_l}{\partial \xi} + g^{32} \frac{\partial T_l}{\partial \eta} + g^{33} \frac{\partial T_l}{\partial \zeta} \right) \right] \\
 & = \gamma_v \sqrt{g} \left[ \frac{1}{T_l} D T_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_l) + \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \sum_{i=1}^{i_{\max}} (\mu_{iml} - \mu_{ilm}) s_{il} \right]. \quad (10.22)
 \end{aligned}$$

## 10.7 Field entropy conservation equations – non-conservative form

The non-conservative form is obtained by differentiating the time derivative and the convection term and comparing with the mass conservation equation. The result is

$$\begin{aligned}
 & \alpha_i \rho_i \sqrt{g} \left( \gamma_v \frac{\partial s_l}{\partial \tau} + \gamma_\xi \bar{V}_i^1 \frac{\partial s_l}{\partial \xi} + \gamma_\eta \bar{V}_i^2 \frac{\partial s_l}{\partial \eta} + \gamma_\zeta \bar{V}_i^3 \frac{\partial s_l}{\partial \zeta} \right) \\
 & - \frac{\partial}{\partial \xi} \left\{ \alpha_i \rho_i \sqrt{g} \gamma_\xi \sum_{i=2}^{i_{\max}} \left[ (s_{il} - s_{l1}) D_{il}^* \left( g^{11} \frac{\partial C_{il}}{\partial \xi} + g^{12} \frac{\partial C_{il}}{\partial \eta} + g^{13} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \\
 & - \frac{\partial}{\partial \eta} \left\{ \alpha_i \rho_i \sqrt{g} \gamma_\eta \sum_{i=2}^{i_{\max}} \left[ (s_{il} - s_{l1}) D_{il}^* \left( g^{21} \frac{\partial C_{il}}{\partial \xi} + g^{22} \frac{\partial C_{il}}{\partial \eta} + g^{23} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \\
 & - \frac{\partial}{\partial \zeta} \left\{ \alpha_i \rho_i \sqrt{g} \gamma_\zeta \sum_{i=2}^{i_{\max}} \left[ (s_{il} - s_{l1}) D_{il}^* \left( g^{31} \frac{\partial C_{il}}{\partial \xi} + g^{32} \frac{\partial C_{il}}{\partial \eta} + g^{33} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\}
 \end{aligned}$$



$$\begin{aligned}
& -\frac{1}{T_l} \frac{\partial}{\partial \xi} \left[ \alpha_l^e \lambda_l^* \sqrt{g} \gamma_\xi \left( g^{11} \frac{\partial T_l}{\partial \xi} + g^{12} \frac{\partial T_l}{\partial \eta} + g^{13} \frac{\partial T_l}{\partial \zeta} \right) \right] \\
& -\frac{1}{T_l} \frac{\partial}{\partial \eta} \left[ \alpha_l^e \lambda_l^* \sqrt{g} \gamma_\eta \left( g^{21} \frac{\partial T_l}{\partial \xi} + g^{22} \frac{\partial T_l}{\partial \eta} + g^{23} \frac{\partial T_l}{\partial \zeta} \right) \right] \\
& -\frac{1}{T_l} \frac{\partial}{\partial \zeta} \left[ \alpha_l^e \lambda_l^* \sqrt{g} \gamma_\zeta \left( g^{31} \frac{\partial T_l}{\partial \xi} + g^{32} \frac{\partial T_l}{\partial \eta} + g^{33} \frac{\partial T_l}{\partial \zeta} \right) \right] \\
& = \gamma_v \sqrt{g} \left[ \frac{1}{T_l} DT_l^N + \sum_{i=1}^{i_{\max}} \mu_{il} (s_{ii} - s_l) \right]. \tag{10.23}
\end{aligned}$$

## 10.8 Irreversible power dissipation caused by the viscous forces

The irreversible power dissipation caused by the viscous forces due to deformation as a result of the mean velocities in the space is an important term in the energy conservation. Substituting for the stress tensor components using the Helmholtz and Stokes hypothesis, the following expression was obtained in Chapter 5 Eq. (5.154), Kolev (1994a),

$$\begin{aligned}
\gamma_v \frac{\alpha_l}{\alpha_l^e} \frac{P_{kl}}{v_l} &= 2 \left[ \gamma_x \left( \frac{\partial u_l}{\partial x} \right)^2 + \gamma_y \left( \frac{\partial v_l}{\partial y} \right)^2 + \gamma_z \left( \frac{\partial w_l}{\partial z} \right)^2 \right] \\
&+ \left( \frac{\partial v_l}{\partial x} + \frac{\partial u_l}{\partial y} \right) \left( \gamma_x \frac{\partial v_l}{\partial x} + \gamma_y \frac{\partial u_l}{\partial y} \right) + \left( \frac{\partial w_l}{\partial x} + \frac{\partial u_l}{\partial z} \right) \left( \gamma_x \frac{\partial w_l}{\partial x} + \gamma_z \frac{\partial u_l}{\partial z} \right) \\
&+ \left( \frac{\partial w_l}{\partial y} + \frac{\partial v_l}{\partial z} \right) \left( \gamma_y \frac{\partial w_l}{\partial y} + \gamma_z \frac{\partial v_l}{\partial z} \right) \\
&- \frac{2}{3} \left( \frac{\partial u_l}{\partial x} + \frac{\partial v_l}{\partial y} + \frac{\partial w_l}{\partial z} \right) \left( \gamma_x \frac{\partial u_l}{\partial r} + \gamma_y \frac{\partial v_l}{\partial y} + \gamma_z \frac{\partial w_l}{\partial z} \right) \tag{10.24}
\end{aligned}$$

The transformed form is then

$$\begin{aligned}
\gamma_v \frac{\alpha_l P_{kl}}{\alpha_l^e v_l} = & 2 \left[ \begin{aligned} & \gamma_\xi \left( a^{11} \frac{\partial u_l}{\partial \xi} + a^{21} \frac{\partial u_l}{\partial \eta} + a^{31} \frac{\partial u_l}{\partial \zeta} \right)^2 \\ & + \gamma_\eta \left( a^{12} \frac{\partial v_l}{\partial \xi} + a^{22} \frac{\partial v_l}{\partial \eta} + a^{32} \frac{\partial v_l}{\partial \zeta} \right)^2 \\ & + \gamma_\zeta \left( a^{13} \frac{\partial w_l}{\partial \xi} + a^{23} \frac{\partial w_l}{\partial \eta} + a^{33} \frac{\partial w_l}{\partial \zeta} \right)^2 \end{aligned} \right] \\
& + \left( \begin{aligned} & a^{11} \frac{\partial v_l}{\partial \xi} + a^{21} \frac{\partial v_l}{\partial \eta} + a^{31} \frac{\partial v_l}{\partial \zeta} \\ & + a^{12} \frac{\partial u_l}{\partial \xi} + a^{22} \frac{\partial u_l}{\partial \eta} + a^{32} \frac{\partial u_l}{\partial \zeta} \end{aligned} \right) \left( \begin{aligned} & \gamma_\xi \left( a^{11} \frac{\partial v_l}{\partial \xi} + a^{21} \frac{\partial v_l}{\partial \eta} + a^{31} \frac{\partial v_l}{\partial \zeta} \right) \\ & + \gamma_\eta \left( a^{12} \frac{\partial u_l}{\partial \xi} + a^{22} \frac{\partial u_l}{\partial \eta} + a^{32} \frac{\partial u_l}{\partial \zeta} \right) \end{aligned} \right) \\
& + \left( \begin{aligned} & a^{11} \frac{\partial w_l}{\partial \xi} + a^{21} \frac{\partial w_l}{\partial \eta} + a^{31} \frac{\partial w_l}{\partial \zeta} \\ & + a^{13} \frac{\partial u_l}{\partial \xi} + a^{23} \frac{\partial u_l}{\partial \eta} + a^{33} \frac{\partial u_l}{\partial \zeta} \end{aligned} \right) \left( \begin{aligned} & \gamma_\xi \left( a^{11} \frac{\partial w_l}{\partial \xi} + a^{21} \frac{\partial w_l}{\partial \eta} + a^{31} \frac{\partial w_l}{\partial \zeta} \right) \\ & + \gamma_\zeta \left( a^{13} \frac{\partial u_l}{\partial \xi} + a^{23} \frac{\partial u_l}{\partial \eta} + a^{33} \frac{\partial u_l}{\partial \zeta} \right) \end{aligned} \right) \\
& + \left( \begin{aligned} & a^{12} \frac{\partial w_l}{\partial \xi} + a^{22} \frac{\partial w_l}{\partial \eta} + a^{32} \frac{\partial w_l}{\partial \zeta} \\ & + a^{13} \frac{\partial v_l}{\partial \xi} + a^{23} \frac{\partial v_l}{\partial \eta} + a^{33} \frac{\partial v_l}{\partial \zeta} \end{aligned} \right) \left( \begin{aligned} & \gamma_\eta \left( a^{12} \frac{\partial w_l}{\partial \xi} + a^{22} \frac{\partial w_l}{\partial \eta} + a^{32} \frac{\partial w_l}{\partial \zeta} \right) \\ & + \gamma_\zeta \left( a^{13} \frac{\partial v_l}{\partial \xi} + a^{23} \frac{\partial v_l}{\partial \eta} + a^{33} \frac{\partial v_l}{\partial \zeta} \right) \end{aligned} \right) \\
& - \frac{2}{3} \left( \begin{aligned} & a^{11} \frac{\partial u_l}{\partial \xi} + a^{21} \frac{\partial u_l}{\partial \eta} + a^{31} \frac{\partial u_l}{\partial \zeta} \\ & + a^{12} \frac{\partial v_l}{\partial \xi} + a^{22} \frac{\partial v_l}{\partial \eta} + a^{32} \frac{\partial v_l}{\partial \zeta} \\ & + a^{13} \frac{\partial w_l}{\partial \xi} + a^{23} \frac{\partial w_l}{\partial \eta} + a^{33} \frac{\partial w_l}{\partial \zeta} \end{aligned} \right) \left( \begin{aligned} & \gamma_\xi \left( a^{11} \frac{\partial u_l}{\partial \xi} + a^{21} \frac{\partial u_l}{\partial \eta} + a^{31} \frac{\partial u_l}{\partial \zeta} \right) \\ & + \gamma_\eta \left( a^{12} \frac{\partial v_l}{\partial \xi} + a^{22} \frac{\partial v_l}{\partial \eta} + a^{32} \frac{\partial v_l}{\partial \zeta} \right) \\ & + \gamma_\zeta \left( a^{13} \frac{\partial w_l}{\partial \xi} + a^{23} \frac{\partial w_l}{\partial \eta} + a^{33} \frac{\partial w_l}{\partial \zeta} \right) \end{aligned} \right) \quad (10.25)
\end{aligned}$$

## 10.9 The non-conservative entropy equation in terms of temperature and pressure

The so called non-conservative entropy equation in terms of temperature and pressure was already derived in Chapter 5, Eq. (5.176) *Kolev* (1997):

$$\begin{aligned} & \rho_l c_{pl} \left[ \alpha_l \gamma_v \frac{\partial T_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) T_l \right] - \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } C_s} \right] \left[ \alpha_l \gamma_v \frac{\partial p}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) p \right] \\ & - \nabla \cdot (\alpha_l^e \lambda_l^* \nabla T) + T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \nabla (\alpha_l^e \rho_l D_{il}^* \gamma \nabla C_{il}) = \gamma_v \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right]. \end{aligned} \quad (10.26)$$

For the derivation of its counterpart in transformed coordinates we follow the same procedure as described in Chapter 5 for the derivation of Eq. (5.176) or in *Kolev* (1997). In Chapter 3 also in *Kolev* (1991) the differential relationship, Eq. (3.106), between the field temperature,  $T_l$ , and the field properties ( $s_l, C_{il}, p$ ) is found to be

$$c_{pl} \frac{dT_l}{T_l} - \bar{R}_l \frac{dp_l}{p_l} = ds_l - \sum_{i=2}^{i_{\max}} \left( \frac{\partial s_l}{\partial C_{il}} \right)_{p, T_l, \text{all } C_s \text{ except } C_{il}} dC_{il}, \quad (10.27)$$

where

$$\left( \frac{\partial s}{\partial C_i} \right)_{p, T, \text{all } C_s \text{ except } C_i} = s_{il} - s_l + \Delta s_{il}^{np}, \quad (10.28)$$

$$\frac{\rho_l \bar{R}_l T_l}{p} = \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } C_s} \right]. \quad (10.29)$$

One of the mass concentrations, arbitrarily numbered with subscript 1,  $C_{1l}$ , depends on all others and is computed as all others are known,

$$C_{1l} = 1 - \sum_{i=2}^{i_{\max}} C_{il}. \quad (10.30)$$

Equation (10.28) consists of two parts. For the case of a mixture consisting of ideal fluids the second part is equal to zero,

$$\Delta s_{il}^{np} = 0, \quad (10.31)$$

this also demonstrating the meaning of the subscript  $np$ , which stands for non-perfect fluid. The non-conservative form of the entropy equation in terms of

temperature and pressure is obtained by multiplying the  $i_{\max} - 1$  mass conservation equations (10.18) by  $s_{il} - s_{l1} + \Delta s_{il}^{np}$  and subtracting them from Eq. (10.23). The result is simplified by using

$$\sum_{i=1}^{i_{\max}} \mu_{il} (s_{il} - s_{l1}) - \sum_{i=2}^{i_{\max}} (s_{il} - s_{l1} + \Delta s_{il}^{np}) (\mu_{il} - C_{il} \mu_l) = - \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l), \quad (10.32)$$

neglecting the second order terms, and making the same assumption about the difference in the diffusion coefficients as made in *Kolev* (1991),

$$\begin{aligned} & \sqrt{g} \alpha_l \rho_l c_{pl} \left( \gamma_v \frac{\partial T_l}{\partial \tau} + \gamma_\xi \bar{V}_l \frac{\partial T_l}{\partial \xi} + \gamma_\eta \bar{V}_l^2 \frac{\partial T_l}{\partial \eta} + \gamma_\zeta \bar{V}_l^3 \frac{\partial T_l}{\partial \zeta} \right) \\ & - \alpha_l \sqrt{g} \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } C\text{'s}} \right] \left( \gamma_v \frac{\partial p_l}{\partial \tau} + \gamma_\xi \bar{V}_l \frac{\partial p_l}{\partial \xi} + \gamma_\eta \bar{V}_l^2 \frac{\partial p_l}{\partial \eta} + \gamma_\zeta \bar{V}_l^3 \frac{\partial p_l}{\partial \zeta} \right) \\ & + T_l \sum_{i=2}^{i_{\max}} \left\{ \Delta s_{il}^{np} \left[ \begin{aligned} & \frac{\partial}{\partial \xi} \left[ \alpha_l \rho_l \sqrt{g} \gamma_\xi D_{il}^* \left( g^{11} \frac{\partial C_{il}}{\partial \xi} + g^{12} \frac{\partial C_{il}}{\partial \eta} + g^{13} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right. \\ & \left. + \frac{\partial}{\partial \eta} \left[ \alpha_l \rho_l \sqrt{g} \gamma_\eta D_{il}^* \left( g^{21} \frac{\partial C_{il}}{\partial \xi} + g^{22} \frac{\partial C_{il}}{\partial \eta} + g^{23} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right. \\ & \left. + \frac{\partial}{\partial \zeta} \left[ \alpha_l \rho_l \sqrt{g} \gamma_\zeta D_{il}^* \left( g^{31} \frac{\partial C_{il}}{\partial \xi} + g^{32} \frac{\partial C_{il}}{\partial \eta} + g^{33} \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right] \right\} \\ & - \frac{\partial}{\partial \xi} \left[ \alpha_l \lambda_l^* \sqrt{g} \gamma_\xi \left( g^{11} \frac{\partial T_l}{\partial \xi} + g^{12} \frac{\partial T_l}{\partial \eta} + g^{13} \frac{\partial T_l}{\partial \zeta} \right) \right] \\ & - \frac{\partial}{\partial \eta} \left[ \alpha_l \lambda_l^* \sqrt{g} \gamma_\eta \left( g^{21} \frac{\partial T_l}{\partial \xi} + g^{22} \frac{\partial T_l}{\partial \eta} + g^{23} \frac{\partial T_l}{\partial \zeta} \right) \right] \\ & - \frac{\partial}{\partial \zeta} \left[ \alpha_l \lambda_l^* \sqrt{g} \gamma_\zeta \left( g^{31} \frac{\partial T_l}{\partial \xi} + g^{32} \frac{\partial T_l}{\partial \eta} + g^{33} \frac{\partial T_l}{\partial \zeta} \right) \right] \\ & = \gamma_v \sqrt{g} \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right]. \end{aligned} \quad (10.33)$$

Note that the prescribing of the velocity at the boundary as a boundary condition is associated with pressure decoupling across the interface.

## 10.10 The volume conservation equation

The so called volume conservation equation (5.188) was already derived by *Kolev* (1986) and published by *Kolev* in 1987:

$$\frac{\gamma_v}{\rho a^2} \frac{\partial p}{\partial \tau} + \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} (\mathbf{V}_l \gamma \cdot \nabla) p + \nabla \cdot \sum_{l=1}^{l_{\max}} (\alpha_l \mathbf{V}_l \gamma) = \sum_{l=1}^{l_{\max}} D\alpha_l - \frac{\partial \gamma_v}{\partial \tau},$$

where

$$D\alpha_l = \frac{1}{\rho_l} \left\{ \gamma_v \mu_l - \frac{1}{\rho_l} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, \text{all } C_{il}'} Ds_l^N + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{li}} \right)_{p, s, \text{all } C_{il}'} DC_{il}^N \right] \right\}, \quad (10.34)$$

$a$  is the *sonic velocity* in a “homogeneous” multi-phase mixture

$$\frac{1}{\rho a^2} = \sum_{l=1}^3 \frac{\alpha_l}{\rho_l a_l^2} = \frac{1}{p} \sum_{l=1}^3 \frac{\alpha_l}{\kappa_l} = \frac{1}{\kappa p}, \quad (10.35)$$

and

$$\rho = \sum_{l=1}^3 \alpha_l \rho_l \quad (10.36)$$

is the mixture density. It is very useful for designing numerical solution methods. We derive this equation for the transformed system following the same procedure as in Chapter 5, *Kolev* (1986, 1997). This means we start with the mass conservation for each velocity field, use the chain rule, divide by the field density and add the resulting field equations. The result is

$$\begin{aligned} & \sqrt{g} \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l} \left( \gamma_v \frac{\partial \rho_l}{\partial \tau} + \gamma_\xi \bar{V}_l^1 \frac{\partial \rho_l}{\partial \xi} + \gamma_\eta \bar{V}_l^2 \frac{\partial \rho_l}{\partial \eta} + \gamma_\zeta \bar{V}_l^3 \frac{\partial \rho_l}{\partial \zeta} \right) \\ & + \sum_{l=1}^{l_{\max}} \left[ \frac{\partial}{\partial \xi} (\alpha_l \gamma_\xi \sqrt{g} \bar{V}_l^1) + \frac{\partial}{\partial \eta} (\alpha_l \gamma_\eta \sqrt{g} \bar{V}_l^2) + \frac{\partial}{\partial \zeta} (\alpha_l \gamma_\zeta \sqrt{g} \bar{V}_l^3) \right] \\ & = \gamma_v \sqrt{g} \sum_{l=1}^{l_{\max}} \frac{\mu_l}{\rho_l} - \frac{\partial}{\partial \tau} (\gamma_v \sqrt{g}). \end{aligned} \quad (10.37)$$

The density derivatives were substituted using the differential form of the equation of state for each velocity field, Eq. (3.157) *Kolev* (1991)

$$d\rho = \frac{dp}{a^2} + \left( \frac{\partial \rho}{\partial s} \right)_{p, \text{all } C_s} ds + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho}{\partial C_i} \right)_{p, s, \text{all } C_s \text{ except } C_i} dC_i.$$

Using the following substitutions

$$\alpha_l \rho_l \left( \gamma_\xi \bar{V}_l^1 \frac{\partial C_{il}}{\partial \xi} + \gamma_\eta \bar{V}_l^2 \frac{\partial C_{il}}{\partial \eta} + \gamma_\zeta \bar{V}_l^3 \frac{\partial C_{il}}{\partial \zeta} \right) = DC_{il}^N \quad (10.38)$$

$$\alpha_l \rho_l \left( \gamma_\xi \bar{V}_l^1 \frac{\partial s_l}{\partial \xi} + \gamma_\eta \bar{V}_l^2 \frac{\partial s_l}{\partial \eta} + \gamma_\zeta \bar{V}_l^3 \frac{\partial s_l}{\partial \zeta} \right) = Ds_l^N \quad (10.39)$$

we finally obtain

$$\begin{aligned} & \gamma_v \sqrt{g} \frac{1}{\rho a^2} \frac{\partial p}{\partial \tau} + \sqrt{g} \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} \left( \gamma_\xi \bar{V}_l^1 \frac{\partial p}{\partial \xi} + \gamma_\eta \bar{V}_l^2 \frac{\partial p}{\partial \eta} + \gamma_\zeta \bar{V}_l^3 \frac{\partial p}{\partial \zeta} \right) \\ & + \sum_{l=1}^{l_{\max}} \left[ \frac{\partial}{\partial \xi} (\alpha_l \gamma_\xi \sqrt{g} \bar{V}_l^1) + \frac{\partial}{\partial \eta} (\alpha_l \gamma_\eta \sqrt{g} \bar{V}_l^2) + \frac{\partial}{\partial \zeta} (\alpha_l \gamma_\zeta \sqrt{g} \bar{V}_l^3) \right] \\ & = \sqrt{g} \sum_{l=1}^{l_{\max}} \frac{1}{\rho_l} \left[ \gamma_v \mu_l - \frac{1}{\rho_l} \left( \frac{\partial \rho_l}{\partial s_l} Ds_l^N + \sum_{i=2}^{l_{\max}} \frac{\partial \rho_l}{\partial C_{il}} DC_{il}^N \right) \right] - \frac{\partial}{\partial \tau} (\gamma_v \sqrt{g}). \quad (10.40) \end{aligned}$$

Note that this equation is not more complicated than its counterpart in Cartesian coordinates and can be used instead of one of the mass conservation equations. The volume conservation equation can be directly discretized and incorporated into the numerical scheme. Another possibility is to follow the same scheme as for deriving it analytically but starting with already discretized mass conservation equations. The coupling finally obtained is then *strictly consistent* with the discretized form of the mass conservation equations.

## 10.11 The momentum equations

The conservative form of the local volume- and time-averaged field momentum conservation equations (2.232) as derived in Chapter 2 *Kolev* (1994a) is

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l \mathbf{V}_l \gamma_v) + \nabla \cdot \left( \alpha_l^e \rho_l \gamma \left\{ \mathbf{V}_l \mathbf{V}_l - v_l^* \left[ \nabla \mathbf{V}_l + (\nabla \mathbf{V}_l)^T - \frac{2}{3} (\nabla \cdot \mathbf{V}_l) \mathbf{I} \right] \right\} \right) \\ & + \alpha_l^e \bar{\gamma} \nabla p + \alpha_l \rho_l \mathbf{g} \gamma_v \\ & - \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left\{ \bar{c}_{ml}^d |\Delta \mathbf{V}_m| \cdot \Delta \mathbf{V}_m + \bar{c}_{ml}^{vm} \left[ \frac{\partial}{\partial \tau} \Delta \mathbf{V}_m + (\mathbf{V}_l \cdot \nabla) \Delta \mathbf{V}_m \right] + \bar{c}_{ml}^L (\mathbf{V}_l - \mathbf{V}_m) \times (\nabla \times \mathbf{V}_m) \right\} \end{aligned}$$

$$= \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} \mathbf{V}_m - \mu_{lm} \mathbf{V}_l). \quad (10.41)$$

For the transformation it is more convenient to write the components in each Cartesian direction:

*x* – direction:

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l u_l \gamma_v) + \nabla \cdot \left( \alpha_l^e \rho_l \gamma \left\{ \mathbf{V}_l u_l - v_l^* \left[ \nabla u_l + \frac{\partial \mathbf{V}_l}{\partial x} - \frac{2}{3} (\nabla \cdot \mathbf{V}_l) \mathbf{i} \right] \right\} \right) \\ & + \alpha_l^e \gamma_x \frac{\partial p}{\partial x} + \alpha_l \rho_l g_x \gamma_v \\ & - \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left\{ \begin{aligned} & \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta u_{ml} + \bar{c}_{ml}^{vm} \left[ \frac{\partial \Delta u_{ml}}{\partial \tau} + (\mathbf{V}_l \cdot \nabla) \Delta u_{ml} \right] \\ & + \bar{c}_{ml}^L \left[ (v_l - v_m) \left( \frac{\partial v_m}{\partial x} - \frac{\partial u_m}{\partial y} \right) + (w_l - w_m) \left( \frac{\partial w_m}{\partial x} - \frac{\partial u_m}{\partial z} \right) \right] \end{aligned} \right\} \\ & = \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} u_m - \mu_{lm} u_l). \end{aligned} \quad (10.42)$$

*y* – direction:

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l v_l \gamma_v) + \nabla \cdot \left( \alpha_l^e \rho_l \gamma \left\{ \mathbf{V}_l v_l - v_l^* \left[ \nabla v_l + \frac{\partial \mathbf{V}_l}{\partial y} - \frac{2}{3} (\nabla \cdot \mathbf{V}_l) \mathbf{j} \right] \right\} \right) \\ & + \alpha_l^e \gamma_y \frac{\partial p}{\partial y} + \alpha_l \rho_l g_y \gamma_v \\ & - \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}, w} \left\{ \begin{aligned} & \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta v_{ml} + \bar{c}_{ml}^{vm} \left[ \frac{\partial \Delta v_{ml}}{\partial \tau} + (\mathbf{V}_l \cdot \nabla) \Delta v_{ml} \right] \\ & - \bar{c}_{ml}^L \left[ (u_l - u_m) \left( \frac{\partial v_m}{\partial x} - \frac{\partial u_m}{\partial y} \right) - (w_l - w_m) \left( \frac{\partial w_m}{\partial y} - \frac{\partial v_m}{\partial z} \right) \right] \end{aligned} \right\} \\ & = \gamma_v \sum_{m=1}^{3,w} (\mu_{ml} v_m - \mu_{lm} v_l). \end{aligned} \quad (10.43)$$

$z$  – direction:

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_i \rho_i w_i \gamma_v) + \nabla \cdot \left( \alpha_i^e \rho_i \gamma \left\{ \mathbf{V}_i w_i - v_i^* \left[ \nabla w_i + \frac{\partial \mathbf{V}_i}{\partial z} - \frac{2}{3} (\nabla \cdot \mathbf{V}_i) \mathbf{k} \right] \right\} \right) \\
 & + \alpha_i^e \gamma_z \frac{\partial p}{\partial z} + \alpha_i \rho_i g_z \gamma_v \\
 & - \gamma_v \sum_{\substack{l=1 \\ m \neq l}}^{l_{\max}, w} \left\{ \begin{aligned} & \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta w_{ml} + \bar{c}_{ml}^{vm} \left[ \frac{\partial \Delta w_{ml}}{\partial \tau} + (\mathbf{V}_l \cdot \nabla) \Delta w_{ml} \right] \\ & - \bar{c}_{ml}^L \left[ (u_l - u_m) \left( \frac{\partial w_m}{\partial x} - \frac{\partial u_m}{\partial z} \right) + (v_l - v_m) \left( \frac{\partial w_m}{\partial y} - \frac{\partial v_m}{\partial z} \right) \right] \end{aligned} \right\} \\
 & = \gamma_v \sum_{m=1}^{3, w} (\mu_{ml} w_m - \mu_{lm} w_l). \tag{10.44}
 \end{aligned}$$

We see that the velocity components can be treated as scalars, which are subject to advection and diffusion as any other scalar variable describing the flow.

First we transfer the bulk viscosity term

$$\nabla \cdot \left( \alpha_i^e \rho_i \gamma_i^* \frac{2}{3} (\nabla \cdot \mathbf{V}_i) \mathbf{i} \right),$$

because it appears in similar form in all the momentum equations. Using Eqs. (10.4) and (10.11) we find that the divergence of the field velocity is the number

$$\nabla \cdot \mathbf{V}_i = \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_i}{\partial \xi} + \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}_i}{\partial \eta} + \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}_i}{\partial \zeta}. \tag{10.45}$$

Using Eq. (10.4) we then find that the divergence of the vector is

$$\begin{aligned}
 & \sqrt{g} \nabla \cdot \left[ \alpha_i^e \rho_i \gamma_i^* \frac{2}{3} (\nabla \cdot \mathbf{V}_i) \mathbf{i} \right] \\
 & = \frac{\partial}{\partial \xi} \left[ \sqrt{g} \mathbf{a}^1 \cdot \mathbf{i} \gamma_\xi \alpha_i^e \rho_i \gamma_i^* \frac{2}{3} (\nabla \cdot \mathbf{V}_i) \right] + \frac{\partial}{\partial \eta} \left[ \sqrt{g} \mathbf{a}^1 \cdot \mathbf{i} \gamma_\eta \alpha_i^e \rho_i \gamma_i^* \frac{2}{3} (\nabla \cdot \mathbf{V}_i) \right] \\
 & + \frac{\partial}{\partial \zeta} \left[ \sqrt{g} \mathbf{a}^1 \cdot \mathbf{i} \gamma_\zeta \alpha_i^e \rho_i \gamma_i^* \frac{2}{3} (\nabla \cdot \mathbf{V}_i) \right] \tag{10.46}
 \end{aligned}$$

The term

$$\nabla \cdot \left( \alpha_i^e \rho_i \gamma_i^* \frac{\partial \mathbf{V}_i}{\partial x} \right)$$



presents the divergence of a vector. Again using Eq. (10.4) we obtain

$$\begin{aligned} & \sqrt{g} \nabla \cdot \left( \alpha_i^e \rho_i \gamma_i^* \frac{\partial \mathbf{V}_l}{\partial x} \right) \\ &= \frac{\partial}{\partial \xi} \left( \sqrt{g} \alpha_i^e \rho_i \gamma_i^* v_i^* \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_l}{\partial x} \right) + \frac{\partial}{\partial \eta} \left( \sqrt{g} \alpha_i^e \rho_i \gamma_i^* v_i^* a^2 \cdot \frac{\partial \mathbf{V}_l}{\partial x} \right) + \frac{\partial}{\partial \zeta} \left( \sqrt{g} \alpha_i^e \rho_i \gamma_i^* v_i^* a^3 \cdot \frac{\partial \mathbf{V}_l}{\partial x} \right). \end{aligned} \quad (10.47)$$

The component notation is

$$\begin{aligned} \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}}{\partial x} &= a^{11} \left( \frac{\partial u}{\partial \xi} a^{11} + \frac{\partial u}{\partial \eta} a^{21} + \frac{\partial u}{\partial \zeta} a^{31} \right) + a^{12} \left( \frac{\partial v}{\partial \xi} a^{11} + \frac{\partial v}{\partial \eta} a^{21} + \frac{\partial v}{\partial \zeta} a^{31} \right) \\ &+ a^{13} \left( \frac{\partial w}{\partial \xi} a^{11} + \frac{\partial w}{\partial \eta} a^{21} + \frac{\partial w}{\partial \zeta} a^{31} \right) \end{aligned} \quad (10.48)$$

$$\begin{aligned} \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}}{\partial y} &= a^{21} \left( \frac{\partial u}{\partial \xi} a^{12} + \frac{\partial u}{\partial \eta} a^{22} + \frac{\partial u}{\partial \zeta} a^{32} \right) + a^{22} \left( \frac{\partial v}{\partial \xi} a^{12} + \frac{\partial v}{\partial \eta} a^{22} + \frac{\partial v}{\partial \zeta} a^{32} \right) \\ &+ a^{23} \left( \frac{\partial w}{\partial \xi} a^{12} + \frac{\partial w}{\partial \eta} a^{22} + \frac{\partial w}{\partial \zeta} a^{32} \right) \end{aligned} \quad (10.49)$$

$$\begin{aligned} \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}}{\partial z} &= a^{31} \left( \frac{\partial u}{\partial \xi} a^{13} + \frac{\partial u}{\partial \eta} a^{23} + \frac{\partial u}{\partial \zeta} a^{33} \right) + a^{32} \left( \frac{\partial v}{\partial \xi} a^{13} + \frac{\partial v}{\partial \eta} a^{23} + \frac{\partial v}{\partial \zeta} a^{33} \right) \\ &+ a^{33} \left( \frac{\partial w}{\partial \xi} a^{13} + \frac{\partial w}{\partial \eta} a^{23} + \frac{\partial w}{\partial \zeta} a^{33} \right) \end{aligned} \quad (10.50)$$

with the second superscript indicating the Cartesian component of the contravariant vectors.

The interfacial virtual mass term is transformed as follows

$$\left( \frac{\partial \varphi}{\partial \tau} \right)_{x,y,z} + \mathbf{V}_l \cdot \nabla \varphi = \left( \frac{\partial \varphi}{\partial \tau} \right)_{\xi,\eta,\zeta} + V_l^1 \frac{\partial \varphi}{\partial \xi} + V_l^2 \frac{\partial \varphi}{\partial \eta} + V_l^3 \frac{\partial \varphi}{\partial \zeta}, \quad (10.51)$$

after using the Eqs. (10.2), (10.5) and (10.11). In a similar way the virtual mass term for the wall-field force is transformed. Note that in the moving coordinate system we have

$$\Delta \mathbf{V}_{wl} = \mathbf{V}_{cs} - \mathbf{V}_l. \quad (10.52)$$

The conservative form of the transformed equations is given below. Note that the equations are still Cartesian components of the vector momentum equation. The component equations are multiplied by  $\sqrt{g}$ .

$x$  - direction:

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_l \rho_l u_l \sqrt{g} \gamma_v) \\
 & + \frac{\partial}{\partial \xi} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\xi \left[ u_l \bar{V}_l^1 - v_l^* \left( g^{11} \frac{\partial u_l}{\partial \xi} + g^{12} \frac{\partial u_l}{\partial \eta} + g^{13} \frac{\partial u_l}{\partial \zeta} \right) + \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_l}{\partial x} - \frac{2}{3} a^{11} (\nabla \cdot \mathbf{V}_l) \right] \right\} \\
 & + \frac{\partial}{\partial \eta} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\eta \left[ u_l \bar{V}_l^2 - v_l^* \left( g^{21} \frac{\partial u_l}{\partial \xi} + g^{22} \frac{\partial u_l}{\partial \eta} + g^{23} \frac{\partial u_l}{\partial \zeta} \right) + \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}_l}{\partial x} - \frac{2}{3} a^{21} (\nabla \cdot \mathbf{V}_l) \right] \right\} \\
 & + \frac{\partial}{\partial \zeta} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\zeta \left[ u_l \bar{V}_l^3 - v_l^* \left( g^{31} \frac{\partial u_l}{\partial \xi} + g^{32} \frac{\partial u_l}{\partial \eta} + g^{33} \frac{\partial u_l}{\partial \zeta} \right) + \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}_l}{\partial x} - \frac{2}{3} a^{31} (\nabla \cdot \mathbf{V}_l) \right] \right\} \\
 & + \sqrt{g} \alpha_l^e \left( a^{11} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{21} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{31} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \\
 & - \gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left\{ \bar{c}_{ml}^{3,w} \left( \frac{\partial \Delta u_{ml}}{\partial \tau} + V_l^1 \frac{\partial \Delta u_{ml}}{\partial \xi} + V_l^2 \frac{\partial u_{ml}}{\partial \eta} + V_l^3 \frac{\partial \Delta u_{ml}}{\partial \zeta} \right) + \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| |\Delta u_{ml}| \right\} \\
 & - \gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left\{ \bar{c}_{ml}^L \left[ (v_l - v_m) \left( a^{11} \frac{\partial v_m}{\partial \xi} + a^{21} \frac{\partial v_m}{\partial \eta} + a^{31} \frac{\partial v_m}{\partial \zeta} \right) - a^{12} \frac{\partial u_m}{\partial \xi} - a^{22} \frac{\partial u_m}{\partial \eta} - a^{32} \frac{\partial u_m}{\partial \zeta} \right] \right. \\
 & \quad \left. + (w_l - w_m) \left( a^{11} \frac{\partial w_m}{\partial \xi} + a^{21} \frac{\partial w_m}{\partial \eta} + a^{31} \frac{\partial w_m}{\partial \zeta} \right) - a^{13} \frac{\partial u_m}{\partial \xi} - a^{23} \frac{\partial u_m}{\partial \eta} - a^{33} \frac{\partial u_m}{\partial \zeta} \right] \right\}
 \end{aligned}$$

$$= \gamma_v \sqrt{g} \left[ -\alpha_l \rho_l g_x + \sum_{m=1}^{3,w} (\mu_{ml} u_m - \mu_{lm} u_l) \right], \quad (10.53)$$

*y* – direction:

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l v_l \sqrt{g} \gamma_v) \\ & + \frac{\partial}{\partial \xi} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\xi \left[ v_l \bar{V}_l^1 - v_l^* \left( g^{11} \frac{\partial v_l}{\partial \xi} + g^{12} \frac{\partial v_l}{\partial \eta} + g^{13} \frac{\partial v_l}{\partial \zeta} \right) + \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_l}{\partial y} - \frac{2}{3} a^{12} (\nabla \cdot \mathbf{V}_l) \right] \right\} \\ & + \frac{\partial}{\partial \eta} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\eta \left[ v_l \bar{V}_l^2 - v_l^* \left( g^{21} \frac{\partial v_l}{\partial \xi} + g^{22} \frac{\partial v_l}{\partial \eta} + g^{23} \frac{\partial v_l}{\partial \zeta} \right) + \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}_l}{\partial y} - \frac{2}{3} a^{22} (\nabla \cdot \mathbf{V}_l) \right] \right\} \\ & + \frac{\partial}{\partial \zeta} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\zeta \left[ v_l \bar{V}_l^3 - v_l^* \left( g^{31} \frac{\partial v_l}{\partial \xi} + g^{32} \frac{\partial v_l}{\partial \eta} + g^{33} \frac{\partial v_l}{\partial \zeta} \right) + \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}_l}{\partial y} - \frac{2}{3} a^{32} (\nabla \cdot \mathbf{V}_l) \right] \right\} \\ & + \sqrt{g} \alpha_l^e \left( a^{12} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{22} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{32} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \\ & - \gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left\{ \bar{c}_{ml}^{vm} \left( \frac{\partial \Delta v_{ml}}{\partial \tau} + V_l^1 \frac{\partial \Delta v_{ml}}{\partial \xi} + V_l^2 \frac{\partial \Delta v_{ml}}{\partial \eta} + V_l^3 \frac{\partial \Delta v_{ml}}{\partial \zeta} \right) + \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| |\Delta v_{ml}| \right\} \end{aligned}$$

$$\begin{aligned}
 & -\gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left\{ \bar{c}_{ml}^L \left[ \begin{aligned} & -(u_l - u_m) \begin{pmatrix} a^{11} \frac{\partial v_m}{\partial \xi} + a^{21} \frac{\partial v_m}{\partial \eta} + a^{31} \frac{\partial v_m}{\partial \zeta} \\ -a^{12} \frac{\partial u_m}{\partial \xi} - a^{22} \frac{\partial u_m}{\partial \eta} - a^{32} \frac{\partial u_m}{\partial \zeta} \end{pmatrix} \\ & + (w_l - w_m) \begin{pmatrix} a^{12} \frac{\partial w_m}{\partial \xi} + a^{22} \frac{\partial w_m}{\partial \eta} + a^{32} \frac{\partial w_m}{\partial \zeta} \\ -a^{13} \frac{\partial v_m}{\partial \xi} - a^{23} \frac{\partial v_m}{\partial \eta} - a^{33} \frac{\partial v_m}{\partial \zeta} \end{pmatrix} \end{aligned} \right] \right\} \\
 & = \gamma_v \sqrt{g} \left[ -\alpha_l \rho_l g_y + \sum_{m=1}^{3,w} (\mu_{ml} v_m - \mu_{lm} v_l) \right], \tag{10.54}
 \end{aligned}$$

$z$  - direction:

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_l \rho_l w_l \sqrt{g} \gamma_v) \\
 & + \frac{\partial}{\partial \xi} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\xi \left[ w_l \bar{V}_l^1 - v_l^* \begin{pmatrix} g^{11} \frac{\partial w_l}{\partial \xi} + g^{12} \frac{\partial w_l}{\partial \eta} + g^{13} \frac{\partial w_l}{\partial \zeta} \\ + \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_l}{\partial z} - \frac{2}{3} a^{13} (\nabla \cdot \mathbf{V}_l) \end{pmatrix} \right] \right\} \\
 & + \frac{\partial}{\partial \eta} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\eta \left[ w_l \bar{V}_l^2 - v_l^* \begin{pmatrix} g^{21} \frac{\partial w_l}{\partial \xi} + g^{22} \frac{\partial w_l}{\partial \eta} + g^{23} \frac{\partial w_l}{\partial \zeta} \\ + \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}_l}{\partial z} - \frac{2}{3} a^{23} (\nabla \cdot \mathbf{V}_l) \end{pmatrix} \right] \right\} \\
 & + \frac{\partial}{\partial \zeta} \left\{ \alpha_l^e \rho_l \sqrt{g} \gamma_\zeta \left[ w_l \bar{V}_l^3 - v_l^* \begin{pmatrix} g^{31} \frac{\partial w_l}{\partial \xi} + g^{32} \frac{\partial w_l}{\partial \eta} + g^{33} \frac{\partial w_l}{\partial \zeta} \\ + \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}_l}{\partial y} - \frac{2}{3} a^{33} (\nabla \cdot \mathbf{V}_l) \end{pmatrix} \right] \right\}
 \end{aligned}$$

$$\begin{aligned}
 & +\sqrt{g}\alpha_i^e \left( a^{13}\gamma_\xi \frac{\partial p}{\partial \xi} + a^{23}\gamma_\eta \frac{\partial p}{\partial \eta} + a^{33}\gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \\
 & -\gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left\{ \bar{c}_{ml}^{vm} \left( \frac{\partial \Delta w_{ml}}{\partial \tau} + V_l^1 \frac{\partial \Delta w_{ml}}{\partial \xi} + V_l^2 \frac{\partial \Delta w_{ml}}{\partial \eta} + V_l^3 \frac{\partial \Delta w_{ml}}{\partial \zeta} \right) + \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta w_{ml} \right\} \\
 & -\gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^{3,w} \left\{ \bar{c}_{ml}^L \left[ \begin{array}{l} -\left( u_l - u_m \right) \begin{pmatrix} a^{11} \frac{\partial w_m}{\partial \xi} + a^{21} \frac{\partial w_m}{\partial \eta} + a^{31} \frac{\partial w_m}{\partial \zeta} \\ -a^{13} \frac{\partial u_m}{\partial \xi} - a^{23} \frac{\partial u_m}{\partial \eta} - a^{33} \frac{\partial u_m}{\partial \zeta} \end{pmatrix} \\ -\left( v_l - v_m \right) \begin{pmatrix} a^{12} \frac{\partial w_m}{\partial \xi} + a^{22} \frac{\partial w_m}{\partial \eta} + a^{32} \frac{\partial w_m}{\partial \zeta} \\ -a^{13} \frac{\partial v_m}{\partial \xi} - a^{23} \frac{\partial v_m}{\partial \eta} - a^{33} \frac{\partial v_m}{\partial \zeta} \end{pmatrix} \end{array} \right] \right\} \\
 & = \gamma_v \sqrt{g} \left[ -\alpha_l \rho_l g_z + \sum_{m=1}^{3,w} (\mu_{ml} w_m - \mu_{lm} w_l) \right]. \tag{10.55}
 \end{aligned}$$

## 10.12 The flux concept, conservative and semi-conservative forms

The purpose of this section is to introduce the so called *flux concept*. Within the flux concept the integration over a control volume gives simple balance expressions which are very convenient for constructing of numerical algorithms.

### 10.12.1 Mass conservation equation

The conservative form of the mass conservation equation for each species  $i$  inside the velocity field  $l$  is

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (\alpha_l \rho_l C_{il} \sqrt{g} \gamma_v) + \frac{\partial}{\partial \xi} (\gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{G}_{il}^c) + \frac{\partial}{\partial \eta} (\gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{G}_{il}^c) \\
 & + \frac{\partial}{\partial \zeta} (\gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{G}_{il}^c) = \gamma_v \sqrt{g} \mu_{il}, \tag{10.56}
 \end{aligned}$$

where the species mass flow rate vector defined as follows

$$\mathbf{G}_{il}^C = \alpha_l \rho_l (\mathbf{V}_l - \mathbf{V}_{cs}) C_{il} - \alpha_l \rho_l D_{il}^* \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) = \mathbf{G}_l + \mathbf{F}_{il}^C, \quad (10.57)$$

consists of a convective

$$\mathbf{G}_l = \alpha_l \rho_l (\mathbf{V}_l - \mathbf{V}_{cs}) \quad (10.58)$$

and of a diffusion

$$\mathbf{F}_{il}^C = -\alpha_l \rho_l D_{il}^* \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \quad (10.59)$$

component. The minus sign reflects the observation that the positive diffusion mass flow rate happens towards the decreasing concentrations. Note that for  $C_{il} = 1$  we have  $\mathbf{G}_{il}^C = \mathbf{G}_l$ . The corresponding mass conservation equation for each velocity field is

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l \sqrt{g} \gamma_v) + \frac{\partial}{\partial \xi} (\gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{G}_l) + \frac{\partial}{\partial \eta} (\gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{G}_l) \\ & + \frac{\partial}{\partial \zeta} (\gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{G}_l) = \gamma_v \sqrt{g} \mu_{il}. \end{aligned} \quad (10.60)$$

We multiply Eq. (10.60) by the concentration and subtract the resulting equation from Eq. (10.56). Then the field mass source term is split in two non-negative parts  $\mu_l = \mu_l^+ + \mu_l^-$ . The result is the so called semi-conservative form of the species mass conservation equation

$$\begin{aligned} & \sqrt{g} \left[ \alpha_l \rho_l \gamma_v \frac{\partial C_{il}}{\partial \tau} + \left( \gamma_\xi \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \gamma_\eta \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \gamma_\zeta \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \cdot \mathbf{G}_l \right] \\ & + \frac{\partial}{\partial \xi} (\gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}_{il}^C) + \frac{\partial}{\partial \eta} (\gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}_{il}^C) + \frac{\partial}{\partial \zeta} (\gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}_{il}^C) \\ & + \gamma_v \sqrt{g} \mu_l^+ C_{il} = \gamma_v \sqrt{g} (\mu_{il} - C_{il} \mu_l^-). \end{aligned} \quad (10.61)$$

We will keep in mind the procedure used to derive this equation which is simpler than the initial equation (10.56) because in the convection part it does not contain the derivatives of the contravariant vectors. By designing numerical methods for Eq. (10.61) we will follow the same procedure but applied to the discretized couple of equations (10.56) and (10.60).

### 10.12.2 Entropy equation

The flux notation of the entropy equation is

$$\begin{aligned}
 & \sqrt{g} \left[ \alpha_l \rho_l \gamma_v \frac{\partial s_l}{\partial \tau} + \left( \gamma_\xi \mathbf{a}^1 \frac{\partial s_l}{\partial \xi} + \gamma_\eta \mathbf{a}^2 \frac{\partial s_l}{\partial \eta} + \gamma_\zeta \mathbf{a}^3 \frac{\partial s_l}{\partial \zeta} \right) \cdot \mathbf{G}_l \right] \\
 & + \frac{1}{T} \left[ \frac{\partial}{\partial \xi} \left( \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}_l^T \right) + \frac{\partial}{\partial \eta} \left( \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}_l^T \right) + \frac{\partial}{\partial \zeta} \left( \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}_l^T \right) \right] \\
 & + \frac{\partial}{\partial \xi} \left( \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}_l^{\Delta s C} \right) + \frac{\partial}{\partial \eta} \left( \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}_l^{\Delta s C} \right) + \frac{\partial}{\partial \zeta} \left( \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}_l^{\Delta s C} \right) \\
 & + \gamma_v \sqrt{g} \mu_l^+ s_l = \gamma_v \sqrt{g} D s_l. \tag{10.62}
 \end{aligned}$$

We see that two diffusion fluxes additionally appear, the heat flux

$$\mathbf{F}_l^T = -\alpha_l \lambda_l^* \left( \mathbf{a}^1 \frac{\partial T_l}{\partial \xi} + \mathbf{a}^2 \frac{\partial T_l}{\partial \eta} + \mathbf{a}^3 \frac{\partial T_l}{\partial \zeta} \right), \tag{10.63}$$

and the entropy flux due to diffusion of species with different thermal properties

$$\mathbf{F}_l^{\Delta s C} = \sum_{i=2}^{i_{\max}} (s_{il} - s_{iI}) \mathbf{F}_{il}^C. \tag{10.64}$$

### 10.12.3 Temperature equation

The flux notation of the temperature equation is

$$\begin{aligned}
 & \sqrt{g} c_{pl} \left[ \alpha_l \rho_l \gamma_v \frac{\partial T_l}{\partial \tau} + \left( \gamma_\xi \mathbf{a}^1 \frac{\partial T_l}{\partial \xi} + \gamma_\eta \mathbf{a}^2 \frac{\partial T_l}{\partial \eta} + \gamma_\zeta \mathbf{a}^3 \frac{\partial T_l}{\partial \zeta} \right) \cdot \mathbf{G}_l \right] \\
 & - \alpha_l \sqrt{g} \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } C_s} \right] \left[ \gamma_v \frac{\partial p_l}{\partial \tau} + \left( \gamma_\xi \mathbf{a}^1 \frac{\partial p_l}{\partial \xi} + \gamma_\eta \mathbf{a}^2 \frac{\partial p_l}{\partial \eta} + \gamma_\zeta \mathbf{a}^3 \frac{\partial p_l}{\partial \zeta} \right) \cdot \mathbf{G}_l \frac{1}{\rho_l} \right] \\
 & + \frac{\partial}{\partial \xi} \left( \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}_l^T \right) + \frac{\partial}{\partial \eta} \left( \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}_l^T \right) + \frac{\partial}{\partial \zeta} \left( \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}_l^T \right) \\
 & - T_l \sum_{i=2}^{i_{\max}} \left\{ \Delta s_{il}^{np} \left( \frac{\partial}{\partial \xi} \left( \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}_{il}^C \right) + \frac{\partial}{\partial \eta} \left( \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}_{il}^C \right) + \frac{\partial}{\partial \zeta} \left( \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}_{il}^C \right) \right) \right\} \\
 & = \gamma_v \sqrt{g} \left[ D T_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right]. \tag{10.65}
 \end{aligned}$$

The flux notation of the particles number density equation is

$$\begin{aligned}
 & \frac{\partial}{\partial \tau} (n_l \sqrt{g} \gamma_v) + \frac{\partial}{\partial \xi} \left[ \sqrt{g} \gamma_\xi n_l \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \right] + \frac{\partial}{\partial \eta} \left[ \sqrt{g} \gamma_\eta n_l \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \right] \\
 & + \frac{\partial}{\partial \zeta} \left[ \sqrt{g} \gamma_\zeta n_l \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \right] \\
 & + \frac{\partial}{\partial \eta} (\sqrt{g} \gamma_\eta \mathbf{F}_l^n) + \frac{\partial}{\partial \xi} (\sqrt{g} \gamma_\xi \mathbf{F}_l^n) + \frac{\partial}{\partial \zeta} (\sqrt{g} \gamma_\zeta \mathbf{F}_l^n) \\
 & = \gamma_v \sqrt{g} (\dot{n}_{l,kin} - \dot{n}_{l,coal} + \dot{n}_{l,sp}), \tag{10.66}
 \end{aligned}$$

where the turbulent diffusion flux of particles is defined as follows

$$\mathbf{F}_l^n = -\frac{\nu_l'}{Sc^l} \left( \mathbf{a}^1 \frac{\partial n_l}{\partial \xi} + \mathbf{a}^2 \frac{\partial n_l}{\partial \eta} + \mathbf{a}^3 \frac{\partial n_l}{\partial \zeta} \right). \tag{10.67}$$

#### 10.12.4 Momentum conservation in the x-direction

The flux notation of the  $x$ -component of the momentum equation is

$$\begin{aligned}
 & \sqrt{g} \left[ \alpha_l \rho_l \gamma_v \frac{\partial u_l}{\partial \tau} + \left( \gamma_\xi \mathbf{a}^1 \frac{\partial u_l}{\partial \xi} + \gamma_\eta \mathbf{a}^2 \frac{\partial u_l}{\partial \eta} + \gamma_\zeta \mathbf{a}^3 \frac{\partial u_l}{\partial \zeta} \right) \cdot \mathbf{G}_l \right] \\
 & + \frac{\partial}{\partial \xi} (\gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}_l^u) + \frac{\partial}{\partial \eta} (\gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}_l^u) + \frac{\partial}{\partial \zeta} (\gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}_l^u) \\
 & + \sqrt{g} \alpha_l \left( a^{11} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{21} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{31} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \\
 & - \sqrt{g} \gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta u_{ml} - \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| \Delta u_{wl} \right] \\
 & - \gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^{l_{\max}} \bar{c}_{ml}^L \left[ (\mathbf{V}_l - \mathbf{V}_m) \times \left( \mathbf{a}^1 \times \frac{\partial \mathbf{V}_m}{\partial \xi} + \mathbf{a}^2 \times \frac{\partial \mathbf{V}_m}{\partial \eta} + \mathbf{a}^3 \times \frac{\partial \mathbf{V}_m}{\partial \zeta} \right) \right]_x \\
 & - \gamma_v \sqrt{g} \bar{c}_{wl}^L \left[ (\mathbf{V}_l - \mathbf{V}_{cs}) \times \left( \mathbf{a}^1 \times \frac{\partial \mathbf{V}_{cs}}{\partial \xi} + \mathbf{a}^2 \times \frac{\partial \mathbf{V}_{cs}}{\partial \eta} + \mathbf{a}^3 \times \frac{\partial \mathbf{V}_{cs}}{\partial \zeta} \right) \right]_x
 \end{aligned}$$



$$\begin{aligned}
 & -\gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm} \left[ \frac{\partial \Delta u_{ml}}{\partial \tau} + \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta u_{ml}}{\partial \xi} + \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta u_{ml}}{\partial \eta} \right] \\
 & + \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta u_{ml}}{\partial \zeta} \\
 & -\gamma_v \sqrt{g} \bar{c}_{wl}^{vm} \left[ \frac{\partial \Delta u_{csl}}{\partial \tau} + \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta u_{csl}}{\partial \xi} + \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta u_{csl}}{\partial \eta} \right] \\
 & + \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta u_{csl}}{\partial \zeta} \\
 & = \gamma_v \sqrt{g} \left[ -\alpha_l \rho_l g_x + \sum_{m=1}^{3,w} [\mu_{ml} (u_m - u_l)] + \mu_{lw} (u_{lw} - u_l) \right], \tag{10.68}
 \end{aligned}$$

where

$$\mathbf{F}_l^u = \mathbf{F}_l^{uv} + \mathbf{F}_l^{uvb} + \mathbf{F}_l^{uvT}, \tag{10.69}$$

is the diffusion momentum flux in the  $x$ -direction with components

$$\mathbf{F}_l^{uv} = -\alpha_l \rho_l \nu_l^* \left( \mathbf{a}^1 \frac{\partial u_l}{\partial \xi} + \mathbf{a}^2 \frac{\partial u_l}{\partial \eta} + \mathbf{a}^3 \frac{\partial u_l}{\partial \zeta} \right), \tag{10.70}$$

$$\mathbf{F}_l^{uvb} = \alpha_l^e \rho_l \nu_l^* \frac{2}{3} (\nabla \cdot \mathbf{V}_l) \mathbf{i}, \tag{10.71}$$

$$\mathbf{F}_l^{uvT} = -\alpha_l^e \rho_l \nu_l^* \frac{\partial \mathbf{V}_l}{\partial x}. \tag{10.72}$$

### 10.12.5 Momentum conservation in the $y$ -direction

The flux notation of the  $y$ -component of the momentum equation is

$$\begin{aligned}
 & \sqrt{g} \left[ \alpha_l \rho_l \gamma_v \frac{\partial v_l}{\partial \tau} + \left( \gamma_\xi \mathbf{a}^1 \frac{\partial v_l}{\partial \xi} + \gamma_\eta \mathbf{a}^2 \frac{\partial v_l}{\partial \eta} + \gamma_\zeta \mathbf{a}^3 \frac{\partial v_l}{\partial \zeta} \right) \cdot \mathbf{G}_l \right] \\
 & + \frac{\partial}{\partial \xi} \left( \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}_l^v \right) + \frac{\partial}{\partial \eta} \left( \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}_l^v \right) + \frac{\partial}{\partial \zeta} \left( \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}_l^v \right) \\
 & + \sqrt{g} \alpha_l \left( a^{12} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{22} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{32} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right)
 \end{aligned}$$

$$\begin{aligned}
 & -\sqrt{g}\gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta v_{ml} - \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| \Delta v_{wl} \right] \\
 & -\gamma_v \sqrt{g} \sum_{\substack{l=\max \\ m \neq l}}^L \bar{c}_{ml}^L \left[ (\mathbf{V}_l - \mathbf{V}_m) \times \left( \mathbf{a}^1 \times \frac{\partial \mathbf{V}_m}{\partial \xi} + \mathbf{a}^2 \times \frac{\partial \mathbf{V}_m}{\partial \eta} + \mathbf{a}^3 \times \frac{\partial \mathbf{V}_m}{\partial \zeta} \right) \right]_y \\
 & -\gamma_v \sqrt{g} \bar{c}_{wl}^L \left[ (\mathbf{V}_l - \mathbf{V}_{cs}) \times \left( \mathbf{a}^1 \times \frac{\partial \mathbf{V}_{cs}}{\partial \xi} + \mathbf{a}^2 \times \frac{\partial \mathbf{V}_{cs}}{\partial \eta} + \mathbf{a}^3 \times \frac{\partial \mathbf{V}_{cs}}{\partial \zeta} \right) \right]_y \\
 & -\gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm} \left[ \begin{aligned} & \frac{\partial \Delta v_{ml}}{\partial \tau} + \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta v_{ml}}{\partial \xi} + \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta v_{ml}}{\partial \eta} \\ & + \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta v_{ml}}{\partial \zeta} \end{aligned} \right] \\
 & -\gamma_v \sqrt{g} \bar{c}_{wl}^{vm} \left[ \begin{aligned} & \frac{\partial \Delta v_{csl}}{\partial \tau} + \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta v_{csl}}{\partial \xi} + \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta v_{csl}}{\partial \eta} \\ & + \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta v_{csl}}{\partial \zeta} \end{aligned} \right] \\
 & = \gamma_v \sqrt{g} \left[ -\alpha_l \rho_l g_y + \sum_{m=1}^{3,w} [\mu_{ml} (v_m - v_l)] + \mu_{lw} (v_{lw} - v_l) \right], \tag{10.73}
 \end{aligned}$$

where

$$\mathbf{F}_l^v = \mathbf{F}_l^{vv} + \mathbf{F}_l^{vvb} + \mathbf{F}_l^{vvt} \tag{10.74}$$

is the diffusion momentum flux in the  $y$ -direction with components

$$\mathbf{F}_l^{vv} = -\alpha_l \rho_l v_l^* \left( \mathbf{a}^1 \frac{\partial v_l}{\partial \xi} + \mathbf{a}^2 \frac{\partial v_l}{\partial \eta} + \mathbf{a}^3 \frac{\partial v_l}{\partial \zeta} \right), \tag{10.75}$$

$$\mathbf{F}_l^{vvb} = \alpha_l^e \rho_l v_l^* \frac{2}{3} (\nabla \cdot \mathbf{V}_l) \mathbf{j}, \tag{10.76}$$

$$\mathbf{F}_l^{vvt} = -\alpha_l^e \rho_l v_l^* \frac{\partial \mathbf{V}_l}{\partial y}.$$

### 10.12.6 Momentum conservation in the z-direction

The flux notation of the z-component of the momentum equation is

$$\begin{aligned}
& \sqrt{g} \left[ \alpha_l \rho_l \gamma_v \frac{\partial w_l}{\partial \tau} + \left( \gamma_\xi \mathbf{a}^1 \frac{\partial w_l}{\partial \xi} + \gamma_\eta \mathbf{a}^2 \frac{\partial w_l}{\partial \eta} + \gamma_\zeta \mathbf{a}^3 \frac{\partial w_l}{\partial \zeta} \right) \cdot \mathbf{G}_l \right] \\
& + \frac{\partial}{\partial \xi} \left( \gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}_l^w \right) + \frac{\partial}{\partial \eta} \left( \gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}_l^w \right) + \frac{\partial}{\partial \zeta} \left( \gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}_l^w \right) \\
& + \sqrt{g} \alpha_l \left( a^{13} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{23} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{33} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \\
& - \sqrt{g} \gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| \Delta w_{ml} - \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| \Delta w_{wl} \right] \\
& - \gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^{\max} \bar{c}_{ml}^L \left[ (\mathbf{V}_l - \mathbf{V}_m) \times \left( \mathbf{a}^1 \times \frac{\partial \mathbf{V}_m}{\partial \xi} + \mathbf{a}^2 \times \frac{\partial \mathbf{V}_m}{\partial \eta} + \mathbf{a}^3 \times \frac{\partial \mathbf{V}_m}{\partial \zeta} \right) \right]_z \\
& - \gamma_v \sqrt{g} \bar{c}_{wl}^L \left[ (\mathbf{V}_l - \mathbf{V}_{cs}) \times \left( \mathbf{a}^1 \times \frac{\partial \mathbf{V}_{cs}}{\partial \xi} + \mathbf{a}^2 \times \frac{\partial \mathbf{V}_{cs}}{\partial \eta} + \mathbf{a}^3 \times \frac{\partial \mathbf{V}_{cs}}{\partial \zeta} \right) \right]_z \\
& - \gamma_v \sqrt{g} \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm} \left[ \frac{\partial \Delta w_{ml}}{\partial \tau} + \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta w_{ml}}{\partial \xi} + \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta w_{ml}}{\partial \eta} \right. \\
& \quad \left. + \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta w_{ml}}{\partial \zeta} \right] \\
& - \gamma_v \sqrt{g} \bar{c}_{wl}^{vm} \left[ \frac{\partial \Delta w_{csl}}{\partial \tau} + \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta w_{csl}}{\partial \xi} + \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta w_{csl}}{\partial \eta} \right. \\
& \quad \left. + \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \Delta w_{csl}}{\partial \zeta} \right] \\
& = \gamma_v \sqrt{g} \left[ -\alpha_l \rho_l g_z + \sum_{m=1}^{3,w} [\mu_{ml} (w_m - w_l)] + \mu_{lw} (w_{lw} - w_l) \right], \tag{10.77}
\end{aligned}$$

where

$$\mathbf{F}_l^w = \mathbf{F}_l^{wv} + \mathbf{F}_l^{wvb} + \mathbf{F}_l^{wvT} \tag{10.78}$$

is the diffusion momentum flux in the  $z$ -direction with components

$$\mathbf{F}_l^{wv} = -\alpha_l \rho_l v_l^* \left( \mathbf{a}^1 \frac{\partial w_l}{\partial \xi} + \mathbf{a}^2 \frac{\partial w_l}{\partial \eta} + \mathbf{a}^3 \frac{\partial w_l}{\partial \zeta} \right), \quad (10.79)$$

$$\mathbf{F}_l^{wvb} = \alpha_l^e \rho_l v_l^* \frac{2}{3} (\nabla \cdot \mathbf{V}_l) \mathbf{k}, \quad (10.80)$$

$$\mathbf{F}_l^{w\theta T} = -\alpha_l^e \rho_l v_l^* \frac{\partial \mathbf{V}_l}{\partial z}. \quad (10.81)$$

### 10.13 Concluding remarks

The equations derived in this Chapter may be of interest not only for scientists and engineers developing computational models but also for those using intensively computational models. One may compare the set of equation presented in this Chapter with the sets solved by different commercial providers and reveal the physical phenomena which are still not taken into account by them. This is helpful to learn the limitations of the existing products.

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# 11 Type of the system of PDEs

*Understanding the type of the partial differential equation systems is an important prerequisite for building successful numerical methods. This chapter gives the definition equations of eigenvalues and eigenvectors of systems of PDEs with constant coefficients. In addition the way to transform the initial system into canonical form is given after determining the type of the system. Then the following questions are answered: What relation exists between the eigenvalues and (a) the propagation velocity of perturbations of the flow parameters, (b) the propagation velocity of harmonic oscillations of the flow parameters, and (c) the critical flows?*

This section is an English translation of the slightly modified Chap. 4 of *Kolev* (1986). The interested reader will find more information about the method of the characteristics in *Kolev* (1986).

## 11.1 Eigenvalues, eigenvectors, canonical form

The general form of the conservation laws presents a *semi-linear non-homogeneous system of partial differential equations*

$$\frac{\partial \mathbf{X}}{\partial \tau} + \frac{\partial \mathbf{Y}}{\partial z} = \mathbf{D}. \quad (11.1)$$

This form is frequently called the *primary* or *conservative* form. Let us assume a flow that can be completely described by the following vector of dependent variables:  $\mathbf{U} = \mathbf{U}(\tau, z)$ . Using the differential forms of the state variables it is possible to rewrite Eq. (11.1) in other forms, for instance in the so called *semi-conservative* forms

$$\mathbf{J}_1 \frac{\partial \mathbf{U}}{\partial \tau} + \frac{\partial \mathbf{Y}}{\partial z} = \mathbf{D} \quad (11.2)$$

or

$$\frac{\partial \mathbf{U}}{\partial \tau} + \mathbf{F} \frac{\partial \mathbf{Y}}{\partial z} = \mathbf{C} \quad (11.3)$$

where  $\mathbf{J}_1 = \partial \mathbf{X} / \partial \mathbf{U}$ ,  $\mathbf{F} = \mathbf{J}_1^{-1}$ ,  $\mathbf{C} = \mathbf{J}_1^{-1} \mathbf{D}$ , or in the *non-conservative* forms

$$\mathbf{J}_1 \frac{\partial \mathbf{U}}{\partial \tau} + \mathbf{J}_2 \frac{\partial \mathbf{U}}{\partial z} = \mathbf{D} \text{ or } \frac{\partial \mathbf{U}}{\partial \tau} + \mathbf{A} \frac{\partial \mathbf{U}}{\partial z} = \mathbf{C}, \quad (11.4, 5)$$

where  $\mathbf{J}_2 = \partial \mathbf{Y} / \partial \mathbf{U}$ ,  $\mathbf{A} = \mathbf{J}_1^{-1} \mathbf{J}_2$ . Writing Eq. (11.5) in component form yields two equations

$$L_1 = \frac{\partial u_1}{\partial \tau} + a_{11} \frac{\partial u_1}{\partial z} + a_{12} \frac{\partial u_2}{\partial z} - c_1 = 0, \quad (11.6)$$

$$L_2 = \frac{\partial u_2}{\partial \tau} + a_{21} \frac{\partial u_1}{\partial z} + a_{22} \frac{\partial u_2}{\partial z} - c_2 = 0. \quad (11.7)$$

Using a vector  $\mathbf{h}_i$  we build a linear combination of the two equations

$$L_i = h_{i1} \left( \frac{\partial u_1}{\partial \tau} + \underbrace{\frac{a_{11}h_{i1} + a_{21}h_{i2}}{h_{i1}}}_{\lambda_i} \frac{\partial u_1}{\partial z} \right) + h_{i2} \left( \frac{\partial u_2}{\partial \tau} + \underbrace{\frac{a_{12}h_{i1} + a_{22}h_{i2}}{h_{i2}}}_{\lambda_i} \frac{\partial u_2}{\partial z} \right) - h_{i1}c_1 - h_{i2}c_2 = 0. \quad (11.8)$$

If the components of the vector  $\mathbf{h}_i$  are computed so that the following conditions are fulfilled

$$a_{11}h_{i1} + a_{21}h_{i2} = \lambda_i h_{i1}, \quad (11.9)$$

$$a_{12}h_{i1} + a_{22}h_{i2} = \lambda_i h_{i2}, \quad (11.10)$$

or in matrix notation

$$(\mathbf{A}^T - \lambda_i \mathbf{I}) \mathbf{h}_i = 0, \quad (11.11)$$

the two Eqs. (11.8) for each  $\mathbf{h}_i$  can be written in a remarkably simple form as we will see later. Equation (11.11) is the definition equation for *eigenvectors* of the matrix  $\mathbf{A}^T$ . Because Eq. (11.11) is homogeneous, the condition to have linear independent solutions for  $\mathbf{h}_i$  is that the determinant of the coefficient matrix is equal to zero

$$|\mathbf{A}^T - \lambda_i \mathbf{I}| = 0. \quad (11.12)$$

Actually, this is the definition equation for the *eigenvalues* of the matrix  $\mathbf{A}^T$  called the *characteristic equation*. With respect to the unknown eigenvalues it is a polynomial of the  $n$ -th degree, where  $n$  is the rank of the matrix  $\mathbf{A}^T$ . Let us assume we have already obtained  $n$  real solutions for the eigenvalues such that at least two of them differ from each other. With them we can obtain at least  $n$  independent solutions for the eigenvectors. In our particular case of Eq. (11.8) we have

$$h_{11} \left( \frac{\partial u_1}{\partial \tau} + \lambda_1 \frac{\partial u_1}{\partial z} \right) + h_{12} \left( \frac{\partial u_2}{\partial \tau} + \lambda_1 \frac{\partial u_2}{\partial z} \right) = h_{11}c_1 + h_{12}c_2, \quad (11.13)$$

$$h_{21} \left( \frac{\partial u_1}{\partial \tau} + \lambda_2 \frac{\partial u_1}{\partial z} \right) + h_{22} \left( \frac{\partial u_2}{\partial \tau} + \lambda_2 \frac{\partial u_2}{\partial z} \right) = h_{21}c_1 + h_{22}c_2, \quad (11.14)$$

or rewritten as a scalar product

$$\mathbf{h}_i \left( \frac{\partial \mathbf{U}}{\partial \tau} + \lambda_i \frac{\partial \mathbf{U}}{\partial z} \right) = \mathbf{h}_i \mathbf{C}, \quad i = 1, n, \quad (11.15)$$

or in matrix notation

$$\mathbf{H} \frac{\partial \mathbf{U}}{\partial \tau} + \mathbf{A} \mathbf{H} \frac{\partial \mathbf{U}}{\partial z} = \mathbf{H} \mathbf{C}. \quad (11.16)$$

Here  $\mathbf{H}$  is a matrix whose rows are the eigenvectors  $\mathbf{h}_i$ .  $\mathbf{I}$  is the unit matrix. Equations (11.15) or the system (11.16) is called the *pseudo-canonical* form of the original system.

The total time derivative of  $\mathbf{U}$  along the curve defined by

$$dz/d\tau = \lambda_i \quad (11.17)$$

is

$$\frac{d\mathbf{U}}{d\tau} = \frac{\partial \mathbf{U}}{\partial \tau} + \lambda_i \frac{\partial \mathbf{U}}{\partial z}. \quad (11.18)$$

With this notation we reach the final, *canonical form* of the original system,

$$\mathbf{h}_i \frac{d\mathbf{U}}{d\tau} = \mathbf{h}_i \mathbf{C}, \quad i = 1, n. \quad (11.19)$$

The curves whose inclination in the space-time domain is defined by  $dz/d\tau = \lambda_i$  are called *characteristic curves*. Note that each equation in Eq. (11.19) is valid only along the corresponding characteristic curve. The eigenvalues and eigenvectors determine the type of the system – see *Courant and Hilbert* (1962):

- a) If all the solutions of the characteristic equation are *imaginer or complex* then the system is *elliptic* and from the type of the *potential equation*.
- b) If all the solutions are *real and equal*, then the system is *parabolic* and from the type of the *heat conduction equation*.
- c) If the characteristic equation *n different real solutions* or it has *n-real and at least two different solutions*, and the system  $\mathbf{A}^T$  has *n-linear independent eigenvectors*, then the system is *hyperbolic* and from the type of the *wave equation*.



## 11.2 Physical interpretation

The eigenvalues contain very important physical information. This has to be well understood for several reasons. The first one the need to distinguish between different types of systems of partial differential equations with different properties. So, the proper defined multi-phase flow equations are from the type of the wave equation. If due to erroneous formulation the type of the resulting system turns to be other than hyperbolic the model is a priori not acceptable. Second, the formulation of different numerical methods has to rely on the basic mathematical properties of the system in order not to lose them during the numerical integration.

### 11.2.1 Eigenvalues and propagation velocity of perturbations

First of all observe that the eigenvalue  $\lambda_i$  has dimensions of velocity. Attaching a coordinate system with one axis tangential to the characteristic line and moving it with the *characteristic velocity*  $\lambda_i$  allows the initial system of partial differential equations to be transforming to the simpler system of ordinary differential equations. This simply means that (a) in the corresponding canonical equation the change of one variable is connected with the change of the other variables appearing with their derivatives, and (b) the propagation velocity of a perturbation of the participating variables is equal to the characteristic velocity  $\lambda_i$ . This is remarkable. The absence of real eigenvalues in a wrong formulated model means that there is no real propagation velocity of the variable or variables describing the particular physical phenomenon. If the experimental observations show a real and finite propagation velocity such models are a priori in conflict with reality and therefore not acceptable. Using the language of the mathematicians:

Non-hyperbolic models cannot adequately describe flow phenomena.

### 11.2.2 Eigenvalues and propagation velocity of harmonic oscillations

Consider a small perturbation  $\Delta\mathbf{U}$  of the vector of the time-averaged dependent variables  $\bar{\mathbf{U}}$ ,

$$\mathbf{U} = \bar{\mathbf{U}} + \Delta\mathbf{U} \quad (11.20)$$

in a flow described by

$$\frac{\partial\mathbf{U}}{\partial\tau} + \mathbf{A} \frac{\partial\mathbf{U}}{\partial z} = \mathbf{C}, \quad (11.21)$$

resulting in

$$\frac{\partial \Delta \mathbf{U}}{\partial \tau} + \mathbf{A} \frac{\partial \Delta \mathbf{U}}{\partial z} + \frac{\partial \bar{\mathbf{U}}}{\partial \tau} + \mathbf{A} \frac{\partial \bar{\mathbf{U}}}{\partial z} = \mathbf{C}. \quad (11.22)$$

Assuming that

$$\frac{\partial \bar{\mathbf{U}}}{\partial \tau} + \mathbf{A} \frac{\partial \bar{\mathbf{U}}}{\partial z} \approx \mathbf{C} \quad (11.23)$$

results in

$$\frac{\partial \Delta \mathbf{U}}{\partial \tau} + \mathbf{A} \frac{\partial \Delta \mathbf{U}}{\partial z} = 0. \quad (11.25)$$

Consider harmonic oscillations of magnitude  $\Delta \bar{\mathbf{U}}$ , the frequency of the not damped oscillation  $f$  and the wave number  $k$  defined by

$$\mathbf{U} = \Delta \bar{\mathbf{U}} e^{i(f\tau - kz)}. \quad (11.26)$$

Substituting in the perturbation equation we obtain

$$-ik(\mathbf{A} - \mathbf{I}f/k)\Delta \bar{\mathbf{U}} e^{i(f\tau - kz)} = 0 \quad (11.27)$$

or

$$-ik(\mathbf{A} - \mathbf{I}f/k) = 0. \quad (11.28)$$

We see that  $f/k = \lambda_i$  defined by

$$|\mathbf{A} - \mathbf{I}f/k| = 0 \quad (11.29)$$

is the propagation velocity of the harmonic oscillations.

Therefore, if the eigenvalues exist and are real and finite they are equal to the propagation velocity of the harmonic oscillations inside the flow.

### 11.2.3 Eigenvalues and critical flow

What happens if the propagation velocity of a disturbance is equal to zero,  $\lambda_i = 0$ ? It follows from Eq. (11.29) that for this particular case

$$|\mathbf{A}| = 0. \quad (11.30)$$

If for instance pressure and velocity are coupled in one equation of the canonical form, the condition  $\lambda_i = 0$  means that a disturbance of the pressure and the

corresponding velocity disturbance cannot propagate in space. Let us take another look at this situation. Consider a steady state flow described by

$$\mathbf{A} \frac{d\mathbf{U}}{dz} = \mathbf{C}, \quad (11.31)$$

where  $\mathbf{U}$  contains also the pressure. Solving with respect to the derivatives we obtain

$$\frac{dp}{dz} = - \frac{\dots}{|\mathbf{A}|}. \quad (11.32)$$

Condition (11.30) means therefore

$$\frac{dp}{dz} = -\infty. \quad (11.33)$$

But this is the well known condition for critical flow. So we win another physical interpretation of the eigenvalues.

If one of the eigenvalues is equal to zero the flow is stagnant or critical. For the case of non-zero velocities, the changes of any parameters downwards the critical cross section does not influence the steady state mass flow rate. The critical steady state mass flow rate is only a function of the flow parameters upwards the flow.

Critical flow is a phenomenon that is desirable if it is wished to limit the mass flow rate in a facility and undesirable if for instance due to the critical mass flow rate the coolability of a system is endangered or the productivity of a facility is reduced.

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# 12 Numerical solution methods for multi-phase flow problems

*A class of implicit numerical solution methods for multi-phase flow problems based on the entropy concept is presented in this work. First order donor-cell finite difference approximation is used for the time derivatives and the convection terms. The diffusion terms are discretized using a second order finite difference approximation. A staggered grid is used for discretization of the momentum equations. The entropy concept permits analytical reduction of the algebraic problem to give pressure or pressure correction equations for each computational cell. Analytical backward substitution closes the iteration cycle. Three different methods applied in IVA computer codes are presented and the experience gained with these methods is discussed. The class of methods presented has been applied with success to a variety of practical problems in the field of nuclear engineering. High order upwinding and the constrained interpolation profile method are also presented.*

This Chapter is an extended version of the theoretical part of the publication *Kolev* (1996).

## 12.1 Introduction

Numerical methods for transient single-phase flow analysis are available and in widespread use for practical applications. Methods for cost-effective solution of multi-phase flow problems are still in their infancy. There is a significant experience base available for two-phase flows, e.g. gas/liquid and dispersed solid particles/gas flow for the special case of small concentrations in the dispersed phase. To my knowledge, there is no universal method for integrating systems of partial differential equations describing multi-phase flows. The purpose of this work is to present the methods derived for the computer codes IVA2 to IVA6 and to give a short description of the experience gained with these methods.

## 12.2 Formulation of the mathematical problem

Consider the following mathematical problem: A multi-phase flow is described by the following vector of dependent variables

$$\mathbf{U}^T = (\alpha_m, T_l, s_2, s_3, C_{il}, n_l, p, u_l, v_l, w_l), \text{ where } l = 1, 2, 3, \quad (12.1)$$

which is a function of the three space coordinates  $(r, \theta, z)$ , and of the time  $\tau$ ,

$$\mathbf{U} = \mathbf{U}(r, \theta, z, \tau). \quad (12.2)$$

The relationship  $\mathbf{U} = \mathbf{U}(r, \theta, z, \tau)$  is defined by the volume-averaged and successively time-averaged mass, momentum and energy conservation equations as well as by initial conditions, boundary conditions, and geometry. As shown in Chapters 1, 2 and 5 *Kolev* (1994a, b, 1995a), the conservation principles lead to the following system of 21 non-linear, non-homogeneous partial differential equations with variable coefficients

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l \gamma_v) + \nabla \cdot (\alpha_l \rho_l \mathbf{V}_l \gamma) = \gamma_v \mu_l \quad (12.3)$$

$$\alpha_l \rho_l \left[ \gamma_v \frac{\partial C_{il}}{\partial \tau} + (\mathbf{V}_l \mathcal{N}) C_{il} \right] - \nabla \cdot (\alpha_l \rho_l D_{il}^* \mathcal{N} C_{il}) + \gamma_v \mu_l^+ C_{il} = \gamma_v D C_{il} \quad (12.4)$$

$$\frac{\partial}{\partial \tau} (n_l \gamma_v) + \nabla \cdot (n_l \mathbf{V}_l \gamma) = \gamma_v (\dot{n}_{kin} - \dot{n}_{col} + \dot{n}_{isp}) \quad \text{for } \alpha_l > 0, \quad (12.5)$$

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_l \rho_l \mathbf{V}_l \gamma_v) + \nabla \cdot [(\alpha_l \rho_l \mathbf{V}_l \mathbf{V}_l + \bar{\tau}_l) \gamma] + \alpha_l \mathcal{N} p + (\alpha_l \rho_l \mathbf{g} + \mathbf{f}_l) \gamma_v \\ & = \gamma_v \left[ \mu_{wl} \mathbf{V}_{wl} - \mu_{lw} \mathbf{V}_{lw} + \sum_{m=1}^{l_{max}} (\mu_{ml} \mathbf{V}_m - \mu_{lm} \mathbf{V}_l) \right], \end{aligned} \quad (12.6)$$

$$\begin{aligned} & \rho_l c_{pl} \left[ \alpha_l \gamma_v \frac{\partial T_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) T_l \right] - \nabla \cdot (\alpha_l^e \lambda_l^* \mathcal{N} T) \\ & - \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, all\_c's} \right] \left[ \alpha_l \gamma_v \frac{\partial p}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) p \right] + T_l \sum_{i=2}^{l_{max}} \Delta s_{il}^{np} \nabla \cdot (\alpha_l^e \rho_l D_{il}^* \mathcal{N} C_{il}) \\ & = \gamma_v \left[ D T_l^N - T_l \sum_{i=2}^{l_{max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right], \quad \text{for } l = 1, \end{aligned} \quad (12.7a)$$

and

$$\alpha_l \rho_l \left[ \gamma_v \frac{\partial s_l}{\partial \tau} + (\mathbf{V}_l \mathcal{N}) s_l \right] - \frac{1}{T_l} \nabla \cdot (\alpha_l^e \lambda_l^* \mathcal{N} T) + \gamma_v \mu_l^+ s_l = \gamma_v D s_l \quad \text{for } l = 2, 3 \quad (12.7b)$$

This system is defined in the three-dimensional domain  $\mathbf{R}$ . The initial conditions of  $\mathbf{U}(\tau=0) = \mathbf{U}_a$  in  $\mathbf{R}$  and the boundary conditions acting at the interface separating the integration space from its environment are given. The solution required is for conditions after the time interval  $\Delta\tau$  has elapsed. The previous time variables are assigned the index  $a$ . The time variables not denoted with  $a$  are either in the new time plane, or are the best available guesses for the new time plane.

In order to enable modeling of flows with arbitrary obstacles and inclusions in the integration space as is usually expected for technical applications, surface permeabilities are defined

$$(\gamma_r, \gamma_\theta, \gamma_z) = \text{functions of } (r, \theta, z, \tau), \quad (12.8)$$

at the virtual surfaces that separate each computational cell from its environment. By definition, the surface permeabilities have values between one and zero,

$$0 \leq \text{each of all } (\gamma_r, \gamma_\theta, \gamma_z) \text{'s} \leq 1. \quad (12.9)$$

A volumetric porosity

$$\gamma_v = \gamma_v(r, \theta, z, \tau) \quad (12.10)$$

is assigned to each computational cell, with

$$0 < \gamma_v \leq 1. \quad (12.11)$$

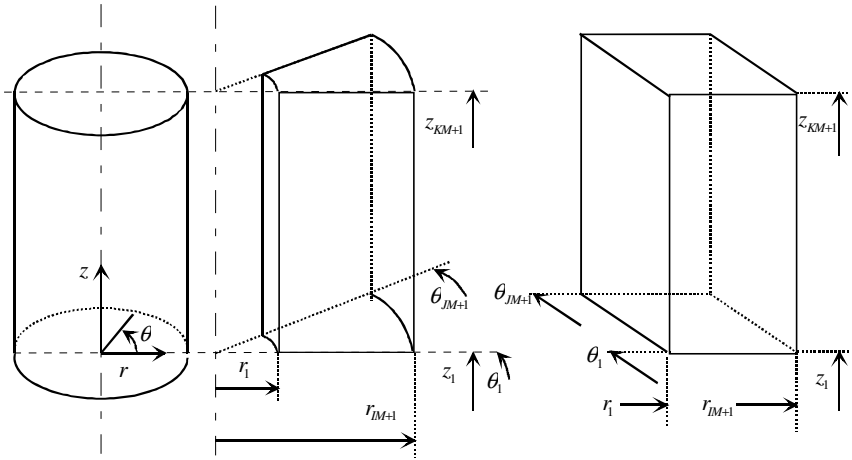
The surface permeabilities and the volume porosities are not expected to be smooth functions of the space coordinates in the region  $\mathbf{R}$  and of time. For this reason, one constructs a frame of geometrical flow obstacles which are functions of space and time. This permits a large number of extremely interesting technical applications to be done with this type of approach.

In order to construct useful numerical solutions it is essential that an appropriate set of constitutive relations be available: state equations, thermodynamic derivatives given in Chapter 3, equations for estimation of the transport properties, correlations modeling the heat, mass and momentum transport across the surfaces dividing the separate velocity fields given in Volume II, etc. These relationships together are called closure equations – see e.g. in *Kolev* (1990), *Kolev, Tomiyama and Sakaguchi* (1991), *Kolev* (1993, 1994a, 1995b, c, d).

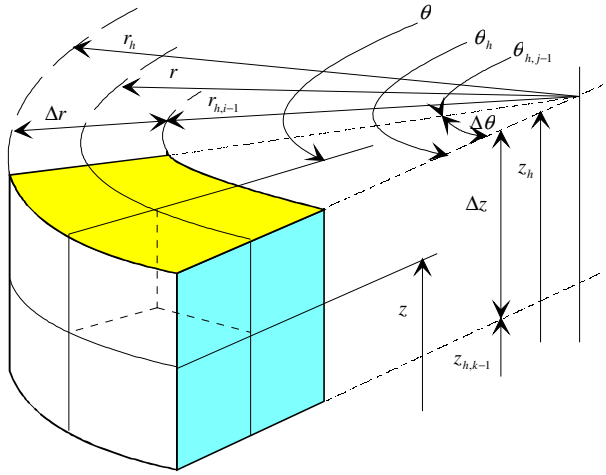
## 12.3 Space discretization and location of the discrete variables

The flow is defined in a domain to rectangular coordinate system; either Cartesian or cylindrical (see Fig. 12.1).

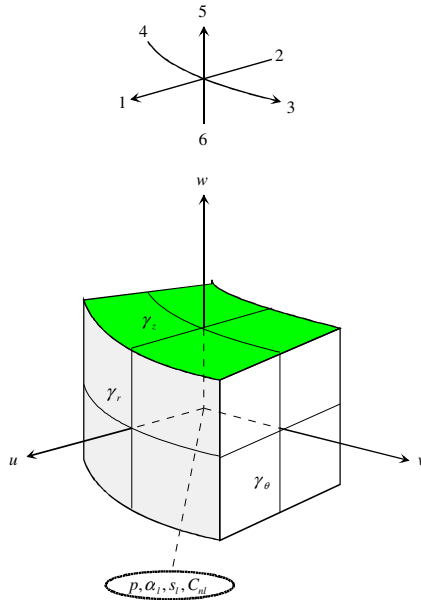
The domain is divided into computational cells as shown in Fig. 12.2, these having a center with the volume coordinates  $(r, \theta, z)$ . The sizes of each cell  $(\Delta r, \Delta \theta, \Delta z)$  are a result of non-uniform spatial discretization in each direction. The cell boundaries are defined by  $(r_h, \theta_h, z_h)$ . The distances between two adjacent cell centers are  $(\Delta r_h, \Delta \theta_h, \Delta z_h)$ , respectively. The dependent variables  $(\alpha_m, s_l, C_{il}, n_l, p)$ , the state and the transport properties  $(T_l, \rho_l, v_l, \lambda_l, etc)$ , and the volumetric porosity are located in the cell centers as shown in Fig. 12.3. The surface permeabilities  $(\gamma_r, \gamma_\theta, \gamma_z)$ , and the velocity components  $(u_l, v_l, w_l)$  are located in the cell interfaces as shown in Fig. 12.3. Thus, all dependent field variables are located at the cell center except the flow variables which are located at the cell surfaces. Fig. 12.3 also shows the control volume for integration of the field and field-component mass conservation equations, of the particle density conservation equations, and of the entropy conservation equations. Figures 12.4, 12.5 and 12.6 show the control volumes for integration of the momentum conservation equations in each separate direction. The momentum cells have the mass cell centers at their surfaces. This forms the so-called staggered grid system - see *Harlow and Amsden (1971)*. Such systems are employed extensively in fluid mechanics to avoid non-physical oscillations by allowing the use of first order discretization methods - see *Issa (1983)*, for example. The staggered grid system is not necessarily needed if high order discretization methods are used.



**Fig. 12.1** Flow domain definition

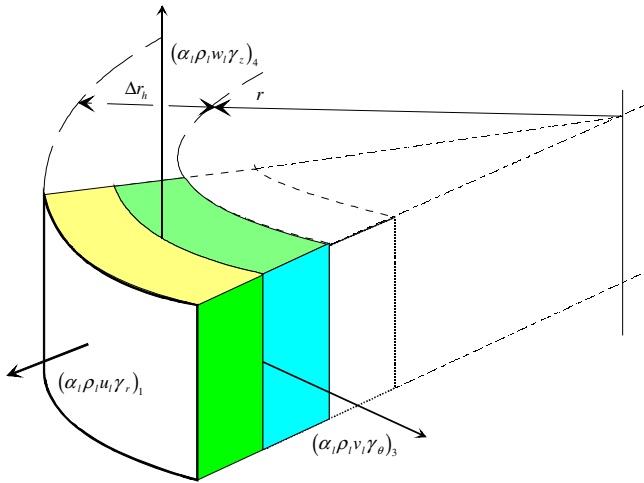


**Fig. 12.2** Geometrical sizes of the mass and energy conservation computational cell referred to as a non-staggered cell

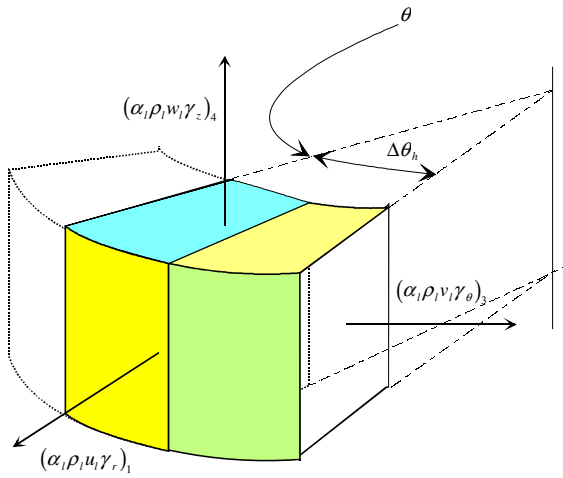


**Fig. 12.3** Location of the dependent variables and of the surface permeabilities

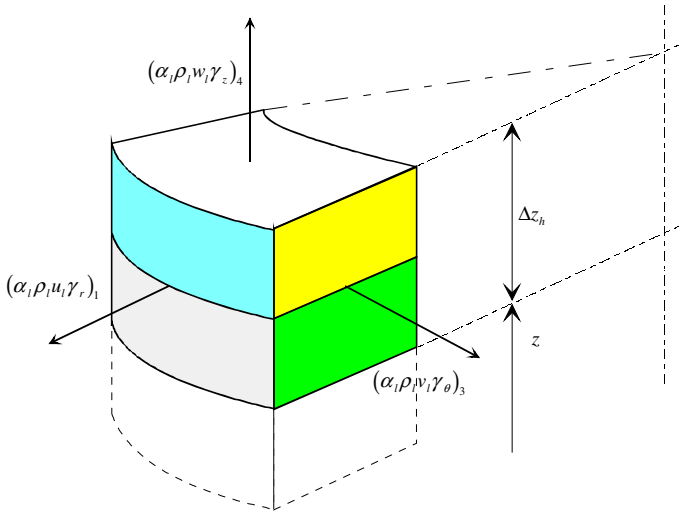




**Fig. 12.4** Staggered control volume for integration of the radial momentum conservation equations



**Fig. 12.5** Staggered control volume for integration of the angular momentum conservation equations



**Fig. 12.6** Staggered control volume for integration of the axial momentum conservation equations

Each mass and energy cell is identified by the integer indices  $(i, j, k)$  corresponding to the three spatial directions (see Fig. 12.7). The integer indices take values  $i$  from 1 to  $IM + 2$ ,  $j$  from 1 to  $JM + 2$ , and  $k$  from 1 to  $KM + 2$ . To achieve a code architecture that allows treatment of boundary cells in the same manner as non-boundary cells, I use a layer of fictitious cells surrounding the real cells as shown in Fig. 12.7. These cells have common indices  $i = 1$  or  $i = IM + 2$ ,  $j = 1$ , or  $j = JM + 2$ ,  $k = 1$ , or  $k = KM + 2$ , respectively. The boundary conditions in the form of prescribed time functions are likewise applied within these auxiliary cells.

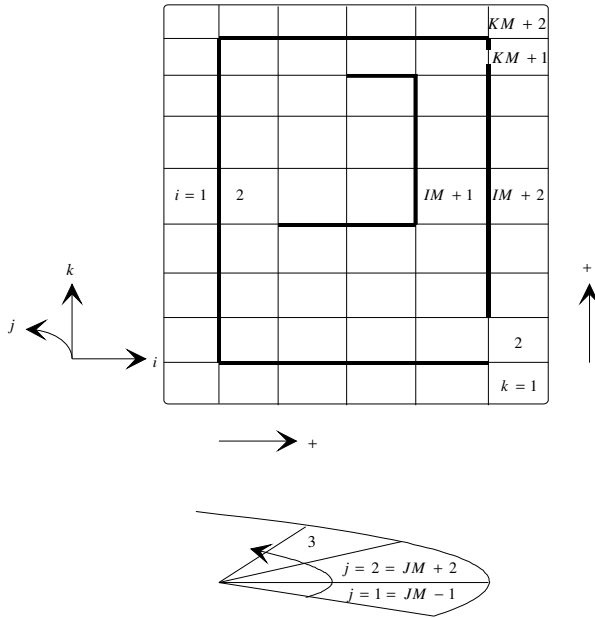
The flow field variables for the front walls of the cell, just like the velocities, are assigned the same indices as the cell itself. Thus, the back walls of the cells have the cell indices minus one for each particular direction.

It is convenient to denote with  $m = 1$  through 6 the cells  $i+1, i-1, j+1, j-1, k+1$  and  $k-1$ , respectively that surround the mass and energy conservation cell  $(i, j, k)$  (see Fig. 12.3). Similarly the corresponding walls attached to the  $(i, j, k)$  cell, and the flow properties defined at this surfaces are assigned designations  $m = 1$  through 6. For notation convenience I define the outwards directed normal velocities to the surfaces of each computational cell,  $\mathbf{V}_{lm}^n$ , as follows

$$\left(\mathbf{V}_{lm}^n\right)^T = \left(u_l, -u_{l,i-1}, v_l, -v_{l,j-1}, w_l, -w_{l,k-1}\right) \tag{12.12}$$

The donor-cell concept is used. This concept is based on the intuitive assumption that the material leaving the cell has the same properties as the cell. This concept stems from *Courant et al. (1952)*.

All field variables have three indices,  $i, j$  and  $k$ . For simplicity, I omit the indices except where they are not  $i, j, k$ . For example  $p_{i,j,k}$  is replaced by  $p$ .



**Fig. 12.7** Three-dimensional mesh construction, boundary layer cells, cell numbering

## 12.4 Discretization of the mass conservation equations

The first order donor cell discretized mass conservation equation (12.3) for each velocity field is given in *Appendix 12.1*. Introducing a number of convenient abbreviations also given in *Appendix 12.1* the following is obtained

$$(\alpha_i \rho_i \gamma_v - \alpha_{la} \rho_{la} \gamma_{va}) / \Delta \tau + \sum_{m=1}^6 (b_{lm+} \alpha_l \rho_l - b_{lm-} \alpha_{lm} \rho_{lm}) - \gamma_v \mu_l = 0. \quad (12.13)$$

Introducing the velocity normal to each surface of the discretization volume, as defined by Eq. (12.12), the  $b$  coefficients from *Appendix 12.1* can be conveniently written as

$$b_{lm+} = \beta_m \xi_{lm+} \mathbf{V}_{lm}^n \geq 0, \quad (12.14)$$

$$b_{lm-} = -\beta_m \xi_{lm-} \mathbf{V}_{lm}^n \geq 0, \quad (12.15)$$

$$\xi_{lm+} = \frac{1}{2} [1 + \text{sign}(\mathbf{V}_{lm}^n)], \quad (12.16)$$

$$\xi_{lm-} = 1 - \xi_{lm+}. \quad (12.17)$$

It is advisable to compute the geometry coefficients,

$$\beta_1 = \frac{r_h^k \gamma_r}{r^k \Delta r}, \quad \beta_2 = \frac{(r_h^k \gamma_r)_{i-1}}{r^k \Delta r}, \quad \beta_3 = \frac{\gamma_\theta}{r^k \Delta \theta}, \quad \beta_4 = \frac{\gamma_{\theta, j-1}}{r^k \Delta \theta}, \quad \beta_5 = \frac{\gamma_z}{\Delta z}, \quad \beta_6 = \frac{\gamma_{z, k-1}}{\Delta z},$$

once at the beginning of the integration and perform corrections only for those computational cells where there is a change of the geometry during the time considered. Normally there are only a few such cells relative to the total number of cells. Note that for the first order finite difference approximation

$$\nabla \gamma \approx \sum_{m=1}^6 \beta_m. \quad (12.18)$$

At this point the method used for computation of the field volumetric fractions in the computer codes IVA2 Kolev (1986, 1987, 1990, 1993), and IVA3 Kolev (1991) will be described. The method exploits the point *Gauss-Seidel* iteration assuming known velocity fields and thermal properties.

In the donor cell concept the term

$$B_{lm-} = b_{lm-} \alpha_{l,m} \rho_{l,m} \geq 0$$

plays an important role.  $B_{lm-}$  is in fact the mass flow entering the cell from the face  $m$  divided by the volume of the cell – it is in fact a non-zero volumetric mass source. Once computed for the mass conservation equation it is stored and used subsequently in all other conservation equations.

Consider the field variables  $\alpha_l \rho_l$  in the convective terms associated with the output flow in the new time plane, and  $\alpha_{lm} \rho_{lm}$  in the neighboring cells  $m$  as the best available guesses for the new time plane. Solving Eq. (12.13) with respect to  $\alpha_l \rho_l$  gives

$$\bar{\alpha}_l \bar{\rho}_l = \left[ \gamma_{va} \left( \mu_l + \frac{\alpha_{la} \rho_{la}}{\Delta \tau} \right) + \sum_{m=1}^6 B_{lm-} \right] / \left( \frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm+} \right). \quad (12.19)$$

Here  $\frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm+} > 0$  is ensured besides by Eq. (12.15) by Eq. (12.11) that does not allow  $\gamma_v$  to be zero. For just originating field we have

$$\bar{\alpha}_l = \frac{\Delta \tau}{\bar{\rho}_l} \frac{\gamma_{va} \mu_l + \sum_{m=1}^6 B_{lm-}}{\gamma_v + \Delta \tau \sum_{m=1}^6 b_{lm+}}.$$

Obviously the field can originate due to convection,  $\sum_{m=1}^6 B_{lm} > 0$ , or due to in-cell mass source,  $\mu_l > 0$ , or due to simultaneous appearance of the both phenomena. In case of origination caused by in-cell mass source terms it is important to define the initial density,  $\bar{\rho}_l$ , in order to compute  $\bar{\alpha}_l = \Delta\tau \frac{\mu_l}{\bar{\rho}_l}$ .

The best mass conservation in such procedures is ensured if the following sequence is used for computation of the volume fractions:

$$\alpha_2 = \bar{\alpha}_2 \bar{\rho}_2 / \rho_2, \alpha_3 = \bar{\alpha}_3 \bar{\rho}_3 / \rho_3, \alpha_1 = 1 - \alpha_2 - \alpha_3.$$

### 12.5 First order donor-cell finite difference approximations

Each velocity field is characterized by properties such as specific entropies, concentrations of the inert components, etc. denoted by  $\Phi$ . These properties can be transported (a) by convection, driven by the mass flow  $F$  and (b) by diffusion, controlled by the diffusion constant  $\Gamma$ . Differentiating the convective terms in the conservation equation and comparing these with the corresponding field mass conservation equation multiplied by  $\Phi$  leads to a considerably simpler semi-conservative form that contains the following connective terms

$$\dots F \frac{\partial \Phi}{\partial z} - \frac{\partial}{\partial z} \left( \Gamma \frac{\partial \Phi}{\partial z} \right) \dots \tag{12.20}$$

To allow use of the values for the mass flows at the boundary of the computational cell where these mass flows are originally defined, I discretize the following equivalent form

$$\dots \frac{\partial}{\partial z} \left( F \Phi - \Gamma \frac{\partial \Phi}{\partial z} \right) - \Phi \frac{\partial F}{\partial z} \dots \tag{12.21}$$

There are a number of methods for discretizing this type of equation for single-phase flows, see *Chow et al. (1978)*, *Patel et al. (1985)*, *Patel et al. (1986)*. Some of them take the form of an analytical solution of the simple convection-diffusion equation as an approximation formula for the profile  $\Phi = \Phi(z)$ . For multi-phase flows I use the simplest first order donor cell method for discretization of the convective terms and a second order central difference method for discretization of the diffusion term. The result

$$F \frac{\partial \Phi}{\partial z} - \frac{\partial}{\partial z} \left( \Gamma \frac{\partial \Phi}{\partial z} \right) = \frac{\partial}{\partial z} \left( F \Phi - \Gamma \frac{\partial \Phi}{\partial z} \right) - \Phi \frac{\partial F}{\partial z}$$

$$\begin{aligned}
 &\approx \frac{1}{\Delta z} \left[ F \left\{ \frac{1}{2} [1 + \text{sign}(F)] \Phi + \frac{1}{2} [1 - \text{sign}(F)] \Phi_{k+1} \right\} - \frac{\Gamma}{\Delta z_h} (\Phi_{k+1} - \Phi) \right] \\
 &- \frac{1}{\Delta z} \left[ F_{k-1} \left\{ \frac{1}{2} [1 + \text{sign}(F_{k-1})] \Phi_{k-1} + \frac{1}{2} [1 - \text{sign}(F_{k-1})] \Phi \right\} - \frac{\Gamma_{k-1}}{\Delta z_{h,k-1}} (\Phi - \Phi_{k-1}) \right] \\
 &- \frac{1}{\Delta z} \Phi (F - F_{k-1}), \tag{12.22}
 \end{aligned}$$

can be further simplified taking into account

$$\frac{1}{2} [1 + \text{sign}(F)] - 1 = -\frac{1}{2} [1 - \text{sign}(F)], \tag{12.23}$$

$$\frac{1}{2} [1 - \text{sign}(F_{k-1})] - 1 = -\frac{1}{2} [1 + \text{sign}(F_{k-1})], \tag{12.24}$$

as follows

$$\begin{aligned}
 &- \frac{1}{\Delta z} \left\{ -F \frac{1}{2} [1 - \text{sign}(F)] + \frac{\Gamma}{\Delta z_h} \right\} (\Phi_{k+1} - \Phi) \\
 &- \frac{1}{\Delta z} \left\{ F_{k-1} \frac{1}{2} [1 + \text{sign}(F_{k-1})] + \frac{\Gamma_{k-1}}{\Delta z_{h,k-1}} \right\} (\Phi_{k-1} - \Phi) \\
 &= -\frac{1}{\Delta z} \left( -F \xi_{15-} + \frac{\Gamma}{\Delta z_h} \right) (\Phi_{k+1} - \Phi) - \frac{1}{\Delta z} \left( F_{k-1} \xi_{16-} + \frac{\Gamma_{k-1}}{\Delta z_{h,k-1}} \right) (\Phi_{k-1} - \Phi) \\
 &= (b_{k+1} + b_{k-1}) \Phi - b_{k-1} \Phi_{k-1} - b_{k+1} \Phi_{k+1}. \tag{12.25}
 \end{aligned}$$

An important property of the linearized coefficients  $b_{k+1}$ ,  $b_{k-1}$  should be noted, namely that they are not negative  $b_{k+1} \geq 0$ ,  $b_{k-1} \geq 0$ . Two consequences of this property are:

- 1) An increase in  $\Phi_{k\pm 1}$  in the neighboring locations to  $k$  leads to an increase in  $\Phi$  and vice versa;
- 2) If the coefficients  $b_{k\pm 1}$  are equal to zero (e.g. due to the fact that  $\gamma_{z,k\pm 1} = 0$ ), the difference  $\Phi_{k\pm 1} - \Phi$  cannot influence the value of  $\Phi$ , so that there is real decoupling.
- 3) If the diffusion is neglected, the flow leaving the cell due to convection does not influence the specific properties  $\Phi$  of the velocity field in the cell. Inlet flows can influence the specific properties of the field in the cell only if they have specific properties that differ from those of the field in the cell considered.

The coefficient  $b$  contains information about the propagation speed of a disturbance  $\Phi$ , namely  $\Delta z / \Delta \tau \approx \textit{convection} + \textit{diffusion velocity}$ . The limitation of the time step is associated with the material velocity of the quality  $\Phi$ . This method is very appropriate for predominant convection

$$Pe = F \Delta z_h / \Gamma < 2, \tag{12.26}$$

see *Patankar* (1980). The local grid *Peclet* number  $Pe$  is the ratio of the amounts of the property  $\Phi$  transported by convection and diffusion, respectively. Large values of  $Pe$ , e.g.  $|Pe| > 10$ , mean that convection is predominant and small values that diffusion is predominant. For one-dimensional processes without sources of  $\Phi$  and with predominant convection, linearization of the profile leads to overestimation of the diffusion component of the flow. This is characteristic for coarse meshes, leading to  $|Pe| > 10$ . This consideration causes some investigators to look for a more realistic profile of the function  $\Phi = \Phi(z)$  as a basis for construction of discretization schemes without a strong upper limitation on mesh size.

## 12.6 Discretization of the concentration equations

Equation (12.4) is discretized following the procedure already described in Section 12.5. The result is given in *Appendix 12.2*. The abbreviated notation is

$$\alpha_{ia} \rho_{ia} \gamma_{va} \frac{C_{il} - C_{ila}}{\Delta \tau} - \sum_{m=1}^6 b_{lm} (C_{ilm} - C_{il}) + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ C_{il} = \frac{1}{2} (\gamma_v + \gamma_{va}) DC_{il} \tag{12.27}$$

where

$$b_{lm} = B_{lm-} + \beta_m D_{ilm}^* / \Delta L_{lm}. \tag{12.28}$$

Solving with respect to the unknown concentration we obtain

$$C_{il} = \frac{\alpha_{ia} \rho_{ia} \gamma_{va} C_{ila} + \Delta \tau \left[ \frac{1}{2} (\gamma_v + \gamma_{va}) DC_{il} + \sum_{m=1}^6 b_{lm} C_{ilm} \right]}{\alpha_{ia} \rho_{ia} \gamma_{va} + \Delta \tau \left[ \sum_{m=1}^6 b_{lm} + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ \right]}.$$

For the case of just an originating velocity field,  $\alpha_{ia} = 0$  and

$$\sum_{m=1}^6 b_{lm} + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ > 0,$$

we have

$$C_{il} = \frac{\frac{1}{2}(\gamma_v + \gamma_{va})DC_{il} + \sum_{m=1}^6 b_{lm} C_{ilm}}{\frac{1}{2}(\gamma_v + \gamma_{va})\mu_l^+ + \sum_{m=1}^6 b_{lm}}.$$

The diffusion coefficients  $D_{ilm}^*$  are defined at the surfaces  $m$  of the mass cell such that the diffusion flows through these surfaces are continuous. Because the diffusion coefficients  $D_{il}^*$  are located by definition in the center of the cell, with the coefficients acting at the cell surfaces required, it is necessary to compute these by some kind of averaging procedure. Harmonic averaging, described in *Appendix 12.3*, is the natural choice for diffusion processes. *Patankar (1978)* shows that this averaging procedure gives better results than simple arithmetic averaging for steady-state heat conduction problems in multi-composite materials and for diffusion in single-phase flows. This type of averaging procedure has the following interesting properties:

- (a) If one of the diffusion coefficients  $D_{il}^*$  or  $D_{il,i+1}^*$  is equal to zero,  $D_{il}^*$  is likewise equal to zero, meaning that no diffusion can take place because one of the two elementary cells hinders it, this matching behavior in reality.
- (b) The above effect also takes place if in one of the two neighboring cells the velocity field  $l$  does not exist.

The superficial mass flow rate  $G_{lm}$  of the velocity field  $l$  through the surfaces  $m=1$  through 6, computed by means of the donor-cell concept is also given in *Appendix 12.2*. In this case the absolute value of the *Peclet* number is

$$|Pe| = |G_{lm}| \Delta L_{hm} / D_{ilm}^*, \quad D_{ilm}^* > 0 \quad (12.29)$$

$$|Pe| \rightarrow +\infty \quad D_{ilm}^* = 0, \quad (12.30)$$

where  $\Delta L_{hm}$  is the distance between the centers of the elementary cells belonging to the surface  $m$ . The mass flows  $G_{lm}$  are used simultaneously in several different conservation equations. It is advisable to compute these once at the beginning of the cycle. Because different dependent variables have different diffusion coefficients, the *diffusion* part is specific to each equation.

## 12.7 Discretization of the entropy equation

Equation (12.7) is discretized following the procedure already described in Section 12.5. The result is



$$\alpha_{la} \rho_{la} \gamma_{va} \frac{s_l - s_{la}}{\Delta \tau} - \sum_{m=1}^6 B_{lm-} (s_{lm} - s_l) + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ s_l = \frac{1}{2} (\gamma_v + \gamma_{va}) Ds_l^* \quad (12.31)$$

where

$$Ds_l^* = Ds_l + \frac{1}{\gamma_v T_l} \sum_{m=1}^6 \beta_m \frac{D_{lm}^T}{\Delta L_{lm}} (T_{lm} - T_l). \quad (12.32)$$

The computation of the harmonic averaged thermal conductivity coefficients is given in *Appendix 12.3*. Solving with respect to the unknown specific entropy we obtain

$$s_l = \frac{\alpha_{la} \rho_{la} \gamma_{va} s_{la} + \Delta \tau \left[ \frac{1}{2} (\gamma_v + \gamma_{va}) Ds_l^* + \sum_{m=1}^6 B_{lm-} s_{lm} \right]}{\alpha_{la} \rho_{la} \gamma_{va} + \Delta \tau \left[ \sum_{m=1}^6 B_{lm-} + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ \right]}.$$

For the case of just an originating velocity field,  $\alpha_{la} = 0$  and

$$\sum_{m=1}^6 B_{lm-} + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ > 0,$$

we have

$$s_l = \frac{\frac{1}{2} (\gamma_v + \gamma_{va}) Ds_l^* + \sum_{m=1}^6 B_{lm-} s_{lm}}{\frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ + \sum_{m=1}^6 B_{lm-}}.$$

## 12.8 Discretization of the temperature equation

In place of the entropy equation, the temperature equation can also be used. The temperature equation

$$\begin{aligned} & \rho_l c_{pl} \left[ \alpha_l \gamma_v \frac{\partial T_l}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) T_l \right] - \nabla \cdot (\alpha_l^e \lambda_l^* \nabla T) \\ & - \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } C's} \right] \left[ \alpha_l \gamma_v \frac{\partial p}{\partial \tau} + (\alpha_l^e \mathbf{V}_l \gamma \cdot \nabla) p \right] \\ & = \gamma_v \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right] \end{aligned} \quad (12.33)$$

is discretized following the procedure already described in Section 12.5. The result is

$$\begin{aligned}
& \alpha_{la} \rho_{la} c_{pa} \gamma_{va} \frac{T_l - T_{la}}{\Delta \tau} - \sum_{m=1}^6 \left( B_{lm-c_{pm}} + \beta_m \frac{\alpha_{lm}^* \lambda_{lm}}{\Delta L_{lm}} \right) (T_{lm} - T_l) \\
&= \frac{1}{2} (\gamma_v + \gamma_{va}) \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right] \\
&+ \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } c's} \right]_a \left[ \alpha_{la} \gamma_{va} \frac{p - p_a}{\Delta \tau} - \sum_{m=1}^6 \frac{B_{lm-c_{pm}}}{\rho_{lm}} (p_m - p) \right]. \quad (12.34)
\end{aligned}$$

Computation of the harmonic averaged thermal conductivity coefficients is shown in *Appendix 12.3*. Bearing in mind that the linearized source terms can be rewritten as linear functions of the temperatures

$$DT_l^N = c_l^T - \sum_{k=1}^3 a_{lk}^{*T} T_k, \quad (12.35)$$

we obtain the following form for the above equation

$$\begin{aligned}
& \left[ \frac{\alpha_{la} \rho_{la} c_{pa}}{\Delta \tau} + a_{ll}^{*T} + \frac{1}{\gamma_{va}} \sum_{m=1}^6 \left( B_{lm-c_{pm}} + \beta_m \frac{\alpha_{lm}^* \lambda_{lm}}{\Delta L_{lm}} \right) \right. \\
& \left. + \frac{1}{2} \left( 1 + \frac{\gamma_v}{\gamma_{va}} \right) \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right] T_l + \frac{1}{2} \left( 1 + \frac{\gamma_v}{\gamma_{va}} \right) \sum_{k \neq l} a_{lk}^{*T} T_k \\
&= \frac{\alpha_{la} \rho_{la} c_{pa}}{\Delta \tau} T_{la} + \frac{1}{2} \left( 1 + \frac{\gamma_v}{\gamma_{va}} \right) c_l^T + \frac{1}{\gamma_{va}} \sum_{m=1}^6 \left( B_{lm-c_{pm}} + \beta_m \frac{\alpha_{lm}^* \lambda_{lm}}{\Delta L_{lm}} \right) T_{lm} \\
&+ \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, \text{all } c's} \right]_a \left[ \alpha_{la} \frac{p - p_a}{\Delta \tau} - \frac{1}{\gamma_{va}} \sum_{m=1}^6 \frac{B_{lm-c_{pm}}}{\rho_{lm}} (p_m - p) \right], \quad (12.36)
\end{aligned}$$

or in a short notation

$$a_{ll}^T T_l + \sum_{k \neq l} a_{lk}^T T_k = b_l, \quad (12.37)$$

or in a matrix notation for all velocity fields

$$\mathbf{AT} = \mathbf{B} \quad (12.38)$$

where

$$Dp_l = \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, all\_C's} \right]_a \left[ \alpha_{la} \frac{p - p_a}{\Delta \tau} - \frac{1}{\gamma_{va}} \sum_{m=1}^6 \frac{B_{lm-}}{\rho_{lm}} (p_m - p) \right], \quad (12.39)$$

$$b_{lm-}^* = B_{lm-} c_{pm} + \beta_m \frac{\alpha_{lm}^* \lambda_{lm}}{\Delta L_{lm}}, \quad (12.40)$$

$$a_{ll}^T = \frac{\alpha_{la} \rho_{la} c_{pa}}{\Delta \tau} + \frac{1}{2} \left( 1 + \frac{\gamma_v}{\gamma_{va}} \right) \left[ a_{ll}^{*T} + \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right] + \frac{1}{\gamma_{va}} \sum_{m=1}^6 b_{lm-}^*, \quad (12.41)$$

$$a_{lm}^T = \frac{1}{2} \left( 1 + \frac{\gamma_v}{\gamma_{va}} \right) a_{lm}^{*T} \quad \text{for } l \neq m, \quad (12.42)$$

$$b_l = \frac{\alpha_{la} \rho_{la} c_{pa}}{\Delta \tau} T_{la} + \frac{1}{2} \left( 1 + \frac{\gamma_v}{\gamma_{va}} \right) c_l^T + \frac{1}{\gamma_{va}} \sum_{m=1}^6 b_{lm-}^* T_{lm} + Dp_l. \quad (12.43)$$

This algebraic system can be solved for the temperatures for

$$\det \mathbf{A} = a_{11} a_{22} a_{33} + a_{12} a_{23} a_{31} + a_{21} a_{32} a_{13} - a_{31} a_{22} a_{13} - a_{32} a_{23} a_{11} - a_{12} a_{21} a_{33} \neq 0. \quad (12.44)$$

The result is

$$\mathbf{T} = \mathbf{A}^{-1} \mathbf{B} \quad (12.45)$$

with components

$$T_l = \left( \sum_{m=1}^3 b_m \bar{a}_{lm} \right) / \det \mathbf{A}, \quad (12.46)$$

The  $\bar{a}$  values are given in *Appendix 12.5*.

An important property of the diagonal elements of the matrix  $\mathbf{A}$

$$a_{ll}^T = \frac{\alpha_{la} \rho_{la} c_{pa}}{\Delta \tau} + \frac{1}{2} \left( 1 + \frac{\gamma_v}{\gamma_{va}} \right) \left[ a_{ll}^{*T} + \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right] + \frac{1}{\gamma_{va}} \sum_{m=1}^6 b_{lm-}^* \quad (12.47)$$

should be noted: The  $m$ -th diagonal element equal to zero indicates that at that time the velocity field does not exist, and will not originate in the next time step. For this reason, the rank of the matrix is reduced by one. Even if the field does not exist but will originate in the next time step either by convection or by mass transfer from the neighboring field, or from other mass sources, or by an arbitrary combination of these three processes, the diagonal element is not zero and the initial temperature is induced properly. It is obvious that, if one neglects all convective terms and mass sources, no initial value for a non-existent field can be defined. For numerical computations I recommend normalization of the diagonal

elements  $a_{il,norm}^T = \left| a_{il}^T \Delta \tau / (\rho_{la} c_{pla}) \right|$  and their comparison with e.g.  $\varepsilon = 0.001$ . If  $a_{il,norm}^T < \varepsilon$ , there is no velocity field  $l$ , and the velocity field  $l$  will not originate in the next time step.

The three equations obtained in this manner are used to construct the pressure-temperature coupling. I use the term partial decoupling of the temperature equations from each other (PDTE) to describe this decoupling procedure. This step is extremely important for creation of a stable numerical algorithm even when using first order donor cell discretization. The coupling coefficients between the velocity fields correspond to each of the flow patterns modeled. For a number of these the coupling is strong e.g. bubble-liquid, for others not so strong, e.g. large diameter droplets-gas. Coupling is non-linear in all cases and must be resolved by iteration.

### 12.9. Physical significance of the necessary convergence condition

Writing for each cell  $(i,j,k)$  a single algebraic equation (12.27), I obtain a system of algebraic equations for all  $C_{il}$  values. This system has a specific seven diagonal structure. The coefficient matrix is defined positive. Suppose that all other variables except the  $C_{il}$  values are known. A possible method for the solution of this system for all  $C_{il}$  values is the *Gauss-Seidel* iterative method in one of its several variants. This method is not the most effective but is frequently used because of its simplicity. The necessary condition for convergence of this method is the predominance of the elements on the main diagonal compared to the other elements

$$|b_i| \geq \sum_{m=1}^6 |b_{im}| \quad \text{for all equations,} \tag{12.48}$$

$$|b_i| > \sum_{m=1}^6 |b_{im}| \quad \text{at least for one equation.} \tag{12.49}$$

This is the well-known *Scarborough* criterion, see *Scarborough* (1958). Because this is only a necessary condition, convergence is possible even if this criterion is violated. But satisfaction of this condition gives the confidence that the algebraic system can be solved at least with one iteration method, *Patankar* (1980). Bearing in mind that all of the elements of the sum

$$b = \gamma_v (\alpha_{la} \rho_{la} / \Delta \tau + \mu_l^+) + \sum_{m=1}^6 b_{im} \tag{12.50}$$

are non-negative, the *Scarborough* criterion reduces to

$$\gamma_v(\alpha_{la}\rho_{la} / \Delta\tau + \mu_l^+) \geq 0 \quad \text{for all equations} \quad (12.51)$$

$$\gamma_v(\alpha_{la}\rho_{la} / \Delta\tau + \mu_l^+) > 0 \quad \text{at least for one equation.} \quad (12.52)$$

If  $\alpha_{la} > 0$  at least for one elementary cell, the above conditions are always satisfied. If the velocity field  $l$  is missing from the integration domain,  $\alpha_{la} = 0$  for all cells, satisfaction of  $\mu_l^+ > 0$  is necessary for at least one single point.

The *Scarborough* criterion will now be considered from a different point of view. Assuming that the properties associated with the flow leaving the cell are known at the new point in time and that the properties associated with the flow entering the cell are the best guesses for the new point in time, and solving Eq. (12.27) with respect to  $C_{il}$ , the following is obtained

$$C_{il} = \frac{\alpha_{la}\rho_{la}\gamma_{va}C_{ila} + \Delta\tau \left[ \frac{1}{2}(\gamma_v + \gamma_{va})DC_{il} + \sum_{m=1}^6 b_{lm}C_{ilm} \right]}{\alpha_{la}\rho_{la}\gamma_{va} + \Delta\tau \left[ \sum_{m=1}^6 b_{lm} + \frac{1}{2}(\gamma_v + \gamma_{va})\mu_l^+ \right]}. \quad (12.53)$$

This resembles the use of the point *Jacobi* method for the solution of the above equation. The method consists in a successive visiting of all cells in the definition domain as many times as necessary until the improvement in the solution from iteration to iteration falls below a defined small value. Where space velocity distribution is known, this method works without any problems. Even though this method has the lowest convergence rate compared to other existing methods, this illustrates an important feature, namely the computation of the initial values for field properties when the field is in the process of origination within the current time step,  $\alpha_{la} = 0$ , i.e.

$$C_{il} = \frac{\frac{1}{2}(\gamma_v + \gamma_{va})DC_{il} + \sum_{m=1}^6 b_{lm}C_{ilm}}{\frac{1}{2}(\gamma_v + \gamma_{va})\mu_l^+ + \sum_{m=1}^6 b_{lm}}. \quad (12.54)$$

Note that initial value for  $C_{ila}$  is not required in this case. Only the following is required

$$\frac{1}{2}(\gamma_v + \gamma_{va})\mu_l^+ + \sum_{m=1}^6 b_{lm} > 0. \quad (12.55)$$

The velocity field can originate in several possible ways. Two of these are as follows:

- a) no convection and diffusion takes place but the source terms differ from zero

$$C_{il} = DC_{il} / \mu_l^+; \quad (12.56)$$

b) source terms in the cell are zero but convection or diffusion takes place

$$C_{il} = \left( \sum_{m=1}^6 b_{lm} C_{ilm} \right) / \sum_{m=1}^6 b_{lm} . \quad (12.57)$$

The *Scarborough* criterion is not satisfied if

$$\alpha_{ia} \rho_{ia} \gamma_{va} + \Delta \tau \left[ \sum_{m=1}^6 b_{lm} + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_i^+ \right] = 0 \quad (12.58)$$

for all cells. In this case  $C_{il}$  is, however, not defined in accordance with Eq. (12.53).

In other words, this criterion simply states that if the field (a) does not exist in the entire integration domain and (b) will not originate at the next moment, its specific properties, such as concentrations, entropies, etc. are not defined.

Note: In creation of the numerical method, it is necessary to set the *initial values* for the field properties once before starting simulation so as to avoid multiplication with undefined numbers  $\alpha_{ia} \rho_{ia} C_{ila} / \Delta \tau$ . If the field disappears during the transient, the field will retain its specific properties from the last time step. On origination of a field, its specific properties are obtained by averaging the specific properties of the flows entering it with weighting coefficients equal to the corresponding mass flow divided by the net mass flow into the cell, as obtained say using Eq. (12.54). In this case the previous value for the specific properties, such as  $C_{ila}$ , etc. do not influence the result.

## 12.10. Implicit discretization of momentum equations

Figure 12.4 shows the control volume for discretization of the momentum equation in the  $r$  direction. Compared to the mass conservation cell at the centre, for which pressure is defined, the volume  $r$  is displaced by  $\Delta r / 2$  in the  $r$  direction. Indices 1 through 6 are used to denote the front and back surfaces of the elementary cell in  $r, \theta$ , and  $z$  directions, respectively, similar to the manner shown in Fig. 12.3.

The pressure difference  $p_{i+1} - p$  is the driving force for the velocity change in the  $r$  direction. As already mentioned, this type of construction for a staggered mesh of elementary cells is really only necessary in the case of schemes with the first order of accuracy for the space discretization of the convective terms in the momentum equations. The reasons behind this are as follows: If the mesh is not staggered for the case

$$p_{i+1} = p_{i-1} \quad (12.59)$$

and

$$p \neq p_{i\pm 1} , \quad (12.60)$$

the pressure gradient

$$(p_{i+1} - p_{i-1}) / (\Delta r_h + \Delta r_{h,i-1}) \quad (12.61)$$

is equal to zero which means that in spite of the non-uniform pressure field the velocity component  $u$  is then not influenced in any way by the pressure force, which does not correspond to reality. Such schemes are numerically unstable.

By definition, the permeabilities, the porosities, mass flow rates, densities, and volumetric fractions are not defined at those points where they are needed for discretization of the momentum equation. As a result, surface properties for the staggered cells are derived by surface weighting of the surface properties for the adjacent non-staggered cells. Volumetric properties of the staggered cells are computed by volumetric weighting of the corresponding volumetric properties for the adjacent non-staggered cells. For velocities transporting material between two adjacent non-staggered cells the donor cell principle is used to compute the convected property.

This calculation entails additional computational effort. Despite this, this method has been in widespread use for single-phase flows for the last three decades. A number of examples of the way in which cell mixing properties are computed will now be given.

Instead of direct discretization of the momentum equations, I derive the discretized working form through the following steps:

1. Implicit discretization of the momentum equation.
2. Implicit discretization of the mass conservation equation for the same velocity field.
3. Subtraction of the mass conservation equation obtained in this way multiplication by  $u_i$  from the discretized momentum equation.

The result is given in *Appendix 12.4*. Note that the centrifugal force gives the effective force in the  $r$ -direction which results from fluid motion in the  $\theta$  direction. For computation of the centrifugal force in the relatively large elementary cell both components for surfaces 3 and 4 are added, as was performed during the derivation of the momentum equation itself. For a very small  $\Delta\theta$  this yields the following

$$\lim_{\Delta\theta \rightarrow 0} \frac{\sin(\Delta\theta/2)}{\Delta\theta} = \frac{1}{2} \quad (12.62)$$

which reduces the expression for the centrifugal force to just this expression in the momentum equation. The viscous components that counteract the centrifugal force are computed in an analogous manner.

The momentum equation is now rewritten in the following compact form

$$\left[ \frac{\alpha_{lua} \rho_{lua}}{\Delta \tau} - \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} + c_{lu} |u_l| + a_{l,conv} + \mu_{wl} - \mu_{lw} \right] u_l + \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} u_m$$

$$\begin{aligned}
 &= b_{l,conv} + \alpha_{lua} \rho_{lua} (u_{la} / \Delta \tau - g_u) - \sum_{\substack{m=1 \\ m \neq l}}^3 b_{lm} + \mu_{wl} u_{wl} - \mu_{lw} u_{lw} \\
 &- \alpha_{lua} \frac{1}{\Delta r_h} \frac{\gamma_{rua}}{\gamma_{vu}} (p_{i+1} - p), \tag{12.63}
 \end{aligned}$$

where

$$a_{l,conv} = -\frac{1}{\gamma_{vu}} \sum_{m=1}^6 b u_{lm}, \tag{12.64}$$

$$b_{l,conv} = -\frac{1}{\gamma_{vu}} \left( \bar{R}_{ul} + \sum_{m=1}^6 b u_{lm} u_{lm} \right), \tag{12.65}$$

represents the discretized convective term.

$$\begin{aligned}
 \bar{R}_{ul} &= -\frac{\sin(\Delta \theta / 2)}{(r + \Delta r_h / 2)^{\kappa} \Delta \theta} \left\{ (\alpha_l \rho_l v_l \gamma_{\theta})_3 v_{l3} + (\alpha_l \rho_l v_l \gamma_{\theta})_4 v_{l4} \right. \\
 &- \left. \frac{2}{(r + \Delta r_h / 2)^{\kappa}} \left\{ (\alpha \rho v)_{l3} \left[ \left( \frac{\partial v_l}{\partial \theta} \right)_3 + u_{l3} \right] \gamma_3 + (\alpha \rho v)_{l4} \left[ \left( \frac{\partial v_l}{\partial \theta} \right)_4 + u_{l4} \right] \gamma_4 \right\} \right\}. \tag{12.66}
 \end{aligned}$$

All velocities  $u_l$  are taken as being in the new time plane, a so called implicit formulation. This process is repeated for the chosen number of the velocity fields. The result is a system of algebraic equations with respect to the velocities

$$\mathbf{A} \mathbf{u} = \mathbf{B} - \mathbf{a} (p_{i+1} - p), \tag{12.67}$$

where the elements of the matrix  $\mathbf{A}$  are

$$a_{lm} = a_{lm}^* - \mu_{ml}, \tag{12.68}$$

$$a_{ml}^* = a_{ml}^* = -\bar{c}_{ml}^d - \bar{c}_{lm}^d - (\bar{c}_{ml}^{vm} + \bar{c}_{lm}^{vm}) / \Delta \tau \quad \text{for } m \neq l, \tag{12.69}$$

$$a_{ll} = \frac{\alpha_{lua} \rho_{lua}}{\Delta \tau} - \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} + c_{lu} |u_l| + a_{l,conv} + \mu_{wl} - \mu_{lw}, \tag{12.70}$$

the elements of the algebraic vector  $\mathbf{B}$  are

$$b_l = b_{l,conv} + \alpha_{lua} \rho_{lua} (u_{la} / \Delta \tau - g_u) - \sum_{\substack{m=1 \\ m \neq l}}^3 b_{lm} + \mu_{wl} u_{wl} - \mu_{lw} u_{lw}, \tag{12.71}$$



where

$$b_{lm} = -b_{ml} = (\bar{c}_{ml}^{vm} + \bar{c}_{lm}^{vm})(u_{ma} - u_{la}) / \Delta \tau. \quad (12.72)$$

Note that by definition, if one velocity field does not exist,  $\alpha_l = 0$ , so that the coefficients describing its coupling with the other fields are then equal to zero,  $\bar{c}_{ml}^{vm} = 0$ ,  $\bar{c}_{lm}^{vm} = 0$ . This algebraic system can be solved to derive the  $l$  velocities provided that

$$\det \mathbf{A} = a_{11}a_{22}a_{33} + a_{12}a_{23}a_{31} + a_{21}a_{32}a_{13} - a_{31}a_{22}a_{13} - a_{32}a_{23}a_{11} - a_{12}a_{21}a_{33} \neq 0. \quad (12.73)$$

The result is

$$\mathbf{u} = \mathbf{du} - \mathbf{RU}(p_{i+1} - p), \quad (12.74)$$

where

$$\mathbf{du} = \mathbf{A}^{-1} \mathbf{B} \quad (12.75)$$

with components

$$du_l = \left( \sum_{m=1}^3 b_m \bar{a}_{lm} \right) / \det \mathbf{A}, \quad (12.76)$$

and

$$\mathbf{RU} = \mathbf{A}^{-1} \mathbf{a}, \quad (12.77)$$

with components

$$RU_l = \frac{1}{\Delta r_h} \frac{\gamma_{lua}}{\gamma_{vu}} \left( \sum_{m=1}^3 \alpha_{lua} \bar{a}_{lm} \right) / \det \mathbf{A}, \quad (12.78)$$

and the  $\bar{a}$  values given in *Appendix 12.5*. Equation (12.74) shows that applying a spatial pressure difference to a multi-field mixture with different field densities results in relative motion between the fields. This relative motion between two adjacent fields results in forces that act at the interfaces.

An important property of the diagonal elements

$$\frac{\alpha_{lua} \rho_{lua}}{\Delta \tau} - \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} + c_{lu} |u_l| + a_{l,conv} + \mu_{wl} - \mu_{lw} \quad (12.79)$$

of the matrix  $\mathbf{A}$  should be noted: The  $m$ -th diagonal element equal to zero indicates that at that time the velocity field does not exist, and will not originate in the next time step. For this reason, the rank of the matrix is reduced by one. Even

if the field does not exist, but will originate in the next time step either by convection, or by mass transfer from the neighboring field, or from other mass sources, or by an arbitrary combination of these three processes, the diagonal element is not zero and the initial velocity is induced properly.

It is obvious that, if one neglects all convective terms and mass sources, no initial value for a non-existent field can be defined. For numerical computations we recommend normalization of the diagonal elements  $a_{ll, norm} = |a_{ll} \Delta \tau / \rho_l|$  and their comparison with e.g.  $\varepsilon = 0.001$ . If  $a_{ll, norm} < \varepsilon$ , there is no velocity field  $l$ , and the velocity field  $l$  will not originate in the next time step.

If the convection, diffusion and mass source terms are disregarded, the matrix  $\mathbf{A}$  is symmetric and the expressions for the relative velocities become very simple.

The three equations obtained in this manner are used to construct the pressure-velocity coupling. I use the term partial decoupling of the momentum equations from each other (PDME) to describe this decoupling procedure. Examples of the usefulness of this method are given in *Kolev, Tomiyama and Sakaguchi* (1991). Note the difference between this procedure and the decoupling procedure used in COBRA-TF, *Kelly and Kohrt* (1983), where decoupling is performed with a lower degree of implicitness by solving for the directional mass flow rates instead of for the velocities. This step is extremely important for creation of a stable numerical algorithm even when using first order donor cell discretization. The coupling coefficients between the velocity fields correspond to each of the flow patterns modeled. For a number of these the coupling is strong, e.g. bubble-liquid, for others not so strong, e.g. large diameter droplets-gas. Coupling is non-linear in all cases and must be resolved by iteration.

I derive the discretized working form of the momentum equations in the other two directions analogously to Eq. (12.63). The corresponding control volumes are shown in Figs. 12.5 and 12.6. The result is given in *Appendixes 12.6 and 12.7*. The abbreviated notations are

$$\begin{aligned} & \left[ \frac{\alpha_{lva} \rho_{lva}}{\Delta \tau} - \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} + c_{lv} |v_l| + a_{l, conv} + \mu_{wl} - \mu_{lw} \right] v_l + \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} v_m \\ & = b_{l, conv} + \alpha_{lva} \rho_{lva} (v_{la} / \Delta \tau - g_v) - \sum_{\substack{m=1 \\ m \neq l}}^3 b_{lm} + \mu_{wl} v_{wl} - \mu_{lw} v_{lw} \\ & - \alpha_{lva} \frac{1}{r^k \Delta \theta_h} \frac{\gamma_{\theta va}}{\gamma_{vv}} (p_{j+1} - p), \end{aligned} \quad (12.80)$$

where

$$a_{l, conv} = - \frac{1}{\gamma_{vv}} \sum_{m=1}^6 b v_{lm}, \quad (12.81)$$

$$b_{l,conv} = -\frac{1}{\gamma_{vw}} \left( \bar{R}_{vl} + \sum_{m=1}^6 b v_{lm} v_{lm} \right), \quad (12.82)$$

or

$$\mathbf{v} = \mathbf{d}\mathbf{v} - \mathbf{R}\mathbf{V} \left( p_{j+1} - p \right), \quad (12.83)$$

and

$$\begin{aligned} & \left[ \frac{\alpha_{lwa} \rho_{lwa}}{\Delta \tau} - \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} + c_{lw} |w_l| + a_{l,conv} + \mu_{wl} - \mu_{lw} \right] w_l + \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} w_m \\ & = b_{l,conv} + \alpha_{lwa} \rho_{lwa} (w_{la} / \Delta \tau - g_w) - \sum_{\substack{m=1 \\ m \neq l}}^3 b_{lm} + \mu_{wl} w_{wl} - \mu_{lw} w_{lw} \\ & - \alpha_{lwa} \frac{1}{\Delta z_h} \frac{\gamma_{zwa}}{\gamma_{vw}} (p_{k+1} - p), \end{aligned} \quad (12.84)$$

where

$$a_{l,conv} = -\frac{1}{\gamma_{vw}} \sum_{m=1}^6 b w_{lm}, \quad (12.85)$$

$$b_{l,conv} = -\frac{1}{\gamma_{vw}} \left( \bar{R}_{wl} + \sum_{m=1}^6 b w_{lm} w_{lm} \right), \quad (12.86)$$

or

$$\mathbf{w} = \mathbf{d}\mathbf{w} - \mathbf{R}\mathbf{W} \left( p_{k+1} - p \right). \quad (12.87)$$

The  $a_l$  and  $b_l$  terms reflect the actions of drag and of added mass forces. The  $a_{l,conv}$  and  $b_{l,conv}$  terms reflect the actions of the spatial inertia and viscous forces. This general structure of Eqs. (12.63), (12.80), and (12.84) proved its worth during testing of the code, permitting effects to be introduced step by step.

Note that a non-slip boundary condition at the wall is easily introduced by computing the wall friction force resisting the flow as follows

Wall friction force =

$$\alpha_i^* \rho_l \nu \left[ \frac{1}{(r + \Delta r / 2)^K r_h^K \Delta \theta} \left( \frac{1 - \gamma_3}{\Delta \theta_h} + \frac{1 - \gamma_4}{\Delta \theta_{h,j-1}} \right) + \frac{1}{\Delta z} \left( \frac{1 - \gamma_5}{\Delta z_h} + \frac{1 - \gamma_6}{\Delta z_{h,k-1}} \right) \right] 2u_l$$

$$- \alpha_i^* \rho_l \nu \frac{2 \sin(\Delta \theta / 2)}{(r + \Delta r / 2)^{2K} \Delta \theta} \left( \frac{1 - \gamma_3}{\Delta \theta_h} + \frac{1 - \gamma_4}{\Delta \theta_{h,j-1}} \right) 2v_l \quad (12.88)$$

where  $\alpha_i^* = 1$  for  $l = \text{continuum}$  and  $l = \text{disperse}$ .

The forms of the momentum equations obtained in this way are

$$V_{lm}^n = dV_{lm}^n - RVel_{lm}(p_m - p), \quad (12.89)$$

$$dV_l^{n=1,6} = (du_l, -du_{l,i-1}, dv_l, -dv_{l,j-1}, dw_l, -dw_{l,k-1}), \quad (12.90)$$

$$RVel_l^{n=1,6} = (RU_l, RU_{l,i-1}, RV_l, RV_{l,j-1}, RW_l, RW_{l,k-1}). \quad (12.91)$$

Remember that the normal velocities, Eq. (12.12), are defined as positive if directed from the control volume to the environment on each of the six surfaces  $m$  of the computational cell.

## 12.11 Pressure equations for IVA2 and IVA3 computer codes

The mixture volume conservation equation (5.188) derived in Chapter 5 and *Kolev* (1995a) is

$$\frac{\gamma_v}{\rho a^2} \frac{\partial p}{\partial \tau} + \sum_{i=1}^3 \frac{\alpha_i}{\rho_i a_i^2} \mathbf{V}_i \gamma \cdot \nabla p + \nabla \cdot \left( \sum_{i=1}^3 \alpha_i \mathbf{V}_i \gamma \right) = \sum_{i=1}^3 D\alpha_i - \frac{\partial \gamma_v}{\partial \tau}. \quad (12.92)$$

Here  $a$  is the sonic velocity in a homogeneous multi-phase mixture defined as follows

$$\frac{1}{\rho a^2} = \sum_{i=1}^3 \frac{\alpha_i}{\rho_i a_i^2}, \quad (12.93)$$

and

$$\rho = \sum_{i=1}^3 \alpha_i \rho_i \quad (12.94)$$

is the mixture density. The mixture volume conservation equation is discretized directly using the donor cell concept in the IVA2 computer code. The result is Eq. (38) in *Kolev* (1987),

$$\gamma_{va} \sum_{l=1}^3 \frac{\alpha_{al}}{\rho_{la} a_{la}^2} \frac{p - p_a}{\Delta \tau} + \sum_{m=1}^6 \beta_m \sum_{l=1}^3 \alpha_{lm}^* V_{lm}^n = \sum_{l=1}^3 D\alpha_l - \frac{\gamma_v - \gamma_{va}}{\Delta \tau} \quad (12.95)$$

where

$$D\alpha_l = \frac{1}{2}(\gamma_v + \gamma_{va}) \frac{1}{\rho_{la}} \left\{ \mu_l - \frac{1}{\rho_{la}} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_a Ds_l^N + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{il}} \right)_a DC_{il}^N \right] \right\}$$

$$\approx \frac{1}{2}(\gamma_v + \gamma_{va}) \frac{1}{\rho_{la}} \left\{ \mu_l - \frac{1}{\rho_{la}} \left[ \left( \frac{\partial \rho_l}{\partial T_l} \right)_{p, \text{all\_C's}} \frac{1}{c_{pl}} \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right] \right. \right. \\ \left. \left. + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{il}} \right)_{p, T_l, \text{all\_C's\_except\_} C_{il}} (\mu_{il} - C_{il} \mu_l) \right] \right\}, \quad (12.96)$$

$$\alpha_{lm}^* = \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \left[ 1 + \frac{1}{\rho_{ma} a_{ma}^2} (p_m - p) \right], \quad (12.97)$$

$$Ds_l^N = Ds_l - \mu_l^+ s_l, \quad (12.98)$$

$$DC_{il}^N = DC_{il} - \mu_l^+ s_{il}. \quad (12.99)$$

Substituting the momentum equation (12.89) for the normal velocities

$$V_{lm}^n = dV_{lm}^n - RVel_{lm} (p_m - p),$$

one finally obtains the pressure equation used in IVA2:

$$c p + \sum_{m=1}^6 c_m p_m = d, \quad (12.100)$$

where

$$c = \frac{\gamma_{va}}{\Delta \tau} \sum_{l=1}^3 \frac{\alpha_{al}}{\rho_{la} a_{la}^2} - \sum_{m=1}^6 c_m, \quad (12.101)$$

$$c_m = -\beta_m \sum_{l=1}^3 \alpha_{lm}^* RVel_{lm}, \quad (12.102)$$

$$d = \sum_{l=1}^3 D\alpha_l - \frac{\gamma_v - \gamma_{va}}{\Delta \tau} + p_a \frac{\gamma_{va}}{\Delta \tau} \sum_{l=1}^3 \frac{\alpha_{al}}{\rho_{la} a_{la}^2} - \sum_{m=1}^6 \beta_m \sum_{l=1}^3 \alpha_{lm}^* V_{lm}^n. \quad (12.103)$$

An improvement of the performance of the method for immersing of one phase inside the others for what ever reason it happens is obtained by using

$$\frac{1}{2} \sum_{l=1}^3 \frac{\alpha_{al} + \alpha_l}{\rho_{la} a_{la}^2} \text{ instead } \sum_{l=1}^3 \frac{\alpha_{al}}{\rho_{la} a_{la}^2}.$$

The mixture volume conservation equation used in IVA3 was derived in the same way but starting with mass conservation equations that had already been discretized instead of starting with the analytical one. This ensures full compatibility of the pressure equation obtained in this way with the discretized mass conservation equations.

$$\begin{aligned} & \gamma_v \sum_{l=1}^3 \frac{\alpha_{la}}{\rho_{la}} (\rho_l - \rho_{la}) / \Delta \tau + \sum_{l=1}^3 \frac{1}{\rho_{la}} \left\{ \sum_{m=1}^6 \beta_m [\xi_{lm+} \alpha_l \rho_l + \xi_{lm-} (\alpha_l \rho_l)_m] V_{lm}^n \right\} \\ & = \frac{\gamma_v}{\rho_{la}} \sum_{l=1}^3 \mu_l - (\gamma_v - \gamma_{va}) / \Delta \tau . \end{aligned} \quad (12.104)$$

Replacing the linearized state equation (3.137)

$$\begin{aligned} \rho_l - \rho_{la} & = \frac{1}{a_{la}^2} (p - p_a) + \left( \frac{\partial \rho_l}{\partial s_l} \right)_{p, all\_C's} (s_l - s_{la}) \\ & + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{i,l}} \right)_{p, s_l, all\_C's\_except\_C_{i,l}} (C_{i,l} - C_{i,la}) \end{aligned}$$

from Kolev (1991) I obtain

$$\begin{aligned} & (p - p_a) \frac{\gamma_v}{\Delta \tau} \sum_{l=1}^3 \frac{\alpha_{la}}{\rho_{la} a_{la}^2} + \sum_{l=1}^3 \frac{1}{\rho_{la}} \left\{ \sum_{m=1}^6 \beta_m [\xi_{lm+} \alpha_l \rho_l + \xi_{lm-} (\alpha_l \rho_l)_m] V_{lm}^n \right\} \\ & = \sum_{l=1}^3 D \alpha_l - \frac{\gamma_v - \gamma_{va}}{\Delta \tau} , \end{aligned} \quad (12.105)$$

where

$$\begin{aligned} D \alpha_l & = \frac{\gamma_v}{\rho_{la}} \left\{ \mu_l - \frac{\alpha_{la}}{\Delta \tau} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_a (s_l - s_{la}) + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{il}} \right)_a (C_{il} - C_{ila}) \right] \right\} \\ & \approx \frac{\gamma_v}{\rho_{la}} \left\{ \mu_l - \alpha_{la} \sum_{i=2}^{i_{\max}} \left[ \left( \frac{\partial \rho_l}{\partial C_{il}} \right)_{p, T_l, all\_C's\_except\_C_{il}} - \left( \frac{\partial \rho_l}{\partial T_l} \right)_{p, all\_C's} \frac{T_l}{c_{pl}} \Delta s_{il}^{np} \right] \frac{C_{il} - C_{ila}}{\Delta \tau} \right\} . \end{aligned} \quad (12.106)$$

Substituting the momentum equation (12.89) for the normal velocities  $V_{lm}^n$ , one finally obtains the pressure equation used in IVA3:

$$c p + \sum_{m=1}^6 c_m p_m = d , \quad (12.107)$$

where

$$c = \frac{\gamma_v}{\Delta \tau} \left( \sum_{l=1}^3 \frac{\alpha_{la}}{\rho_{la} \alpha_{la}^2} \right) - \sum_{m=1}^6 c_m, \quad (12.108)$$

$$c_m = -\beta_m \sum_{l=1}^3 \frac{1}{\rho_{la}} \left[ \xi_{lm+} \alpha_l \rho_l + \xi_{lm-} (\alpha_l \rho_l)_m \right] RVel_{lm}, \quad (12.109)$$

$$d = \frac{\gamma_v}{\Delta \tau} \left( \sum_{l=1}^3 \frac{\alpha_{la}}{\rho_{la} \alpha_{la}^2} \right) P_a - \frac{\gamma_v - \gamma_{va}}{\Delta \tau} + \sum_{l=1}^3 \left\{ D\alpha_l - \frac{1}{\rho_{la}} \sum_{m=1}^6 \beta_m \left[ \xi_{lm+} \alpha_l \rho_l + \xi_{lm-} (\alpha_l \rho_l)_m \right] dV_{lm}^n \right\}. \quad (12.110)$$

Writing this equation for each particular cell I obtain a system of  $IM \times JM \times KM$  algebraic equations with respect to the pressures in the new time plane. The coefficient matrix has the expected 7 diagonal symmetrical structure with the guaranteed diagonal dominance, see Eq. (12.108). The system coefficients are continuous non-linear functions of the solutions of the system. Therefore the system is non-linear. It has to be solved by iterations. There are two iteration cycles one called outer and another called inner. Inside the inner cycle the coefficients of the system are considered constant. Even in this case for large size of the problems an iterative procedure of solving the system of algebraic equations is necessary – see for brief introduction Appendix 12.9. I solve this system using one of the 4 successive relaxation methods built into the IVA3 computer code. The relaxation coefficient used here is unity. The four methods in IVA3 differ in the computational effort associated with direct inversion during the iterations. The first three methods solve directly the pressure equation plane by plane, for rectangle, cylinder, circle. The fourth method implements strong coupling between pressure and velocity along one line. This method is called the line-by-line solution method. The user can select the appropriate one of the four methods for the geometry of the problem which has to be simulated.

## 12.12 A Newton-type iteration method for multi-phase flows

The methods used in IVA2 and IVA3 are found to be converging and numerically stable. The reason for introduction of a new method is based on the following observations:

- (a) In spite of the fact that the mixture volume conservation residuals are reduced to the values of the computational zeros, there are limitations on the reduction of the mass residuals for all mass conservation equations;
- (b) Although strict convergence has been demonstrated in hundreds of numerical experiments, it has never been proven analytically.

The method presented below resolves the above two dilemmas: It (i) simultaneously leads to reduction of all residuals to strict computer zeros, and (ii) proof of convergence is derived from the principle of the Newton-type iteration method. Numerical experiments show that this method is 20% faster than the previous two methods, in spite of the fact that a single iteration step takes more time. This method does, however, require a preconditioning step. For the first iteration step I normally use the IVA3 method. All preceding methods are of course retained in IVA4.

A variety of iteration methods for single- and two-phase flows were proposed by Patankar and Spalding (1967, 1972), Patankar (1975, 1978, 1980, 1981), Patankar, Rafiinejad and Spalding (1975), Patankar, Basn and Alpay (1977), Caretto, Gosman, Patankar and Spalding (1973), Spalding (1976, 1979, 1980a, b, 1981a, b, c), Chow and Tien (1978), Connell and Stow (1986), van Doormaal and Raithby (1984), Haaland (1984), Amsden (1985), Köller (1980), Latimaer and Pollard (1985), Neuberger (1984), Prakash (1984), Vanka (1985), Patel and Markatos (1986), Roscoe (1976) having widespread use in a variety of applications, e.g., TRAC development, Addessio et al. (1984, 1985), Andersen and Schang (1984), Dearing (1985), Knight (1984), Liles and Reed (1978), Liles et al. (1981), Liles and Mahaffy (1984), Mahaffy and Liles (1979), Mahaffy (1979), Prior (1979), Rohatgi (1985), Sargis and Chan (1984), Williams and Liles (1984), the COBRA-TF development, Kelly and Kohrt (1983), and the COBRA/TRAC development, Thurgood, Cuta, Koontz and Kelly (1983).

The IVA4 method described here can be considered as a generalization of this family of methods for multi-phase flow that use the highly efficient entropy concept and allowing analytical derivation of the pressure correction equations in contrast with all preceding *Newtonian* iteration methods. The new method will be described in this Section.

Consider the following system of algebraic non-linear equations

$$f_l^\alpha \equiv \left[ \alpha_l \rho_l \gamma_v - (\alpha_l \rho_l \gamma_v)_a \right] / \Delta \tau + \sum_{m=1}^6 [b_{lm+} \alpha_l \rho_l - b_{lm-} (\alpha_l \rho_l)_m] - \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l = 0 \quad (12.111)$$

$$f_l^s \equiv \alpha_a \rho_{la} \gamma_{va} \frac{s_l - s_{la}}{\Delta \tau} - \sum_{m=1}^6 [b_{lm} \alpha_{lm} \rho_{lm} (s_{lm} - s_l)] + \frac{1}{2} (\gamma_v + \gamma_{va}) (\mu_l^+ s_l - Ds_l^*) = 0 \quad (12.112)$$

$$f_{il}^c \equiv \alpha_{la} \rho_{la} \gamma_{va} \frac{C_{il} - C_{ila}}{\Delta \tau} - \sum_{m=1}^6 \left\{ b_{lm-} (\alpha_l \rho_l)_m + \beta_m \frac{D_{ilm}^*}{\Delta L_{h,m}} A(|Pe_m|) \right\} (C_{il,m} - C_{il}) \left. \right\} + \frac{1}{2} (\gamma_v + \gamma_{va}) (\mu_l^+ C_{il} - DC_{il}) = 0 \quad (12.113)$$

$$f_l^n \equiv \left[ n_l \gamma_v - (n_l \gamma_v)_a \right] / \Delta \tau + \sum_{m=1}^6 (b_{lm+} n_l - b_{lm-} n_{lm}) - \frac{1}{2} (\gamma_v + \gamma_{va}) Dn_l = 0 \quad (12.114)$$

$$V_{lm}^n = dV_{lm}^n - RVel_{lm} (p_m - p), \quad (12.115)$$



resulting from the first order donor-cell discretization of the 21 defining equations describing the multi-phase, multi-component flow and the condition

$$\sum_{l=1}^{l_{\max}} \alpha_l - 1 = 0. \quad (12.116)$$

For simplicity, diffusion terms in the concentration and entropy equations are included within the  $DS_l$  and  $DC_{il}$  terms. The system can be written in the following compact form

$$\mathbf{F}(\mathbf{U}) = 0, \quad (12.117)$$

where

$$\mathbf{U}^T = (\alpha_l, s_l, C_{il}, n_l, p, p_m). \quad (12.118)$$

In order to solve the system of Eqs. (12.111-12.115) I proceeded as follows. The sum of the terms on the left hand side should be equal to zero if the current values of all variables satisfy the equations. The mass, entropy, concentration and particle density equations will generally not be satisfied when the new velocities computed from the momentum equations are used to compute the convective terms in Eqs. (12.111) through (12.114). There will be some residual error in each equation as a result of the new velocities given by

$$f_l^\alpha = \frac{\partial f_l^\alpha}{\partial \alpha_l} (\alpha_l - \alpha_{l0}) \rightarrow 0 \quad (12.119)$$

$$f_l^s = \frac{\partial f_l^s}{\partial s_l} (s_l - s_{l0}) \rightarrow 0 \quad (12.120)$$

$$f_{il}^C = \frac{\partial f_{il}^C}{\partial C_{il}} (C_{il} - C_{i0}) \rightarrow 0 \quad (12.121)$$

$$f_l^n = \frac{\partial f_l^n}{\partial n_l} (n_l - n_{l0}) \rightarrow 0 \quad (12.122)$$

$$f_\Sigma = \sum_{l=1}^{l_{\max}} \alpha_l - 1 \rightarrow 0 \quad (12.123)$$

where

$$\alpha_{l0} = \left[ \gamma_{va} \alpha_{la} \rho_{la} / \Delta \tau + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l + \sum_{m=1}^6 b_{lm} \alpha_{lm} \rho_{lm} \right] / \frac{\partial f_l^\alpha}{\partial \alpha_l}, \quad (12.124)$$

$$s_{l0} = \left[ \gamma_{va} \alpha_{la} \rho_{la} s_{la} / \Delta \tau + \frac{1}{2} (\gamma_v + \gamma_{va}) Ds_l^* + \sum_{m=1}^6 b_{lm} \alpha_{lm} \rho_{lm} s_{lm} \right] / \frac{\partial f_l^s}{\partial s_l}, \quad (12.125)$$

$$C_{i|0} = \left[ \begin{array}{l} \gamma_{va} \alpha_{la} \rho_{la} C_{ila} / \Delta \tau + \frac{1}{2} (\gamma_v + \gamma_{va}) DC_{il} \\ + \sum_{m=1}^6 \left[ b_{lm-} (\alpha_l \rho_l)_m + \beta_m \frac{D_{ilm}^*}{\Delta L_{h,m}} A(|Pe_m|) \right] C_{ilm} \end{array} \right] / \frac{\partial f_{il}^C}{\partial C_{il}}, \quad (12.126)$$

$$n_{i0} = \left[ \gamma_{va} n_{la} / \Delta \tau + \frac{1}{2} (\gamma_v + \gamma_{va}) Dn_i + \sum_{m=1}^6 b_{lm-} n_{lm} \right] / \frac{\partial f_l^n}{\partial n_i}. \quad (12.127)$$

All terms are computed using currently known values for each of the variables not marked with  $a$ . The equations are simultaneously satisfied when  $f_l^\alpha, f_l^s, f_{il}^C, f_l^n$  and  $f_\Sigma$  simultaneously approach zero for all cells in the mesh. The variation of each of the dependent variables required to bring the residual errors to zero can be obtained by using the block *Newton-Raphson* method for two-phase flow see for example *Liles and Reed (1978), Thurgood et al. (1983)*. This is implemented by linearizing the equations with respect to the dependent variables  $\alpha_l, s_l, C_{il}, n_i, p$  and  $p_m$  to obtain the following equations for each cell.

$$\frac{\partial f_l^\alpha}{\partial \alpha_l} \delta \alpha_l + \frac{\partial f_l^\alpha}{\partial s_l} \delta s_l + \sum_{i=2}^{i_{\max}} \frac{\partial f_l^\alpha}{\partial C_{il}} \delta C_{il} + \frac{\partial f_l^\alpha}{\partial p} \delta p + \sum_{m=1}^6 \frac{\partial f_l^\alpha}{\partial p_m} \delta p_m = -f_l^\alpha, \quad (12.128)$$

$$\frac{\partial f_l^s}{\partial s_l} \delta s_l + \frac{\partial f_l^s}{\partial p} \delta p + \sum_{m=1}^6 \frac{\partial f_l^s}{\partial p_m} \delta p_m = -f_l^s, \quad (12.129)$$

$$\frac{\partial f_{il}^C}{\partial C_{il}} \delta C_{il} + \frac{\partial f_{il}^C}{\partial p} \delta p + \sum_{m=1}^6 \frac{\partial f_{il}^C}{\partial p_m} \delta p_m = -f_{il}^C, \quad (12.130)$$

$$\frac{\partial f_l^n}{\partial n_i} \delta n_i + \frac{\partial f_l^n}{\partial p} \delta p + \sum_{m=1}^6 \frac{\partial f_l^n}{\partial p_m} \delta p_m = -f_l^n, \quad (12.131)$$

$$\sum_{l=1}^{l_{\max}} \delta \alpha_l = -f_\Sigma. \quad (12.132)$$

This equation has the form

$$\frac{\partial \mathbf{f}}{\partial \mathbf{U}} \delta \mathbf{U} = -\mathbf{f}. \quad (12.133)$$

$Det(\partial \mathbf{f} / \partial \mathbf{U})$  is the *Jacobian* of the system of equations evaluated for the set of dependent variables given by the vector  $\mathbf{U}$ .  $\delta \mathbf{U}$  is the solution vector containing the linear variation of the dependent variables, and  $-\mathbf{f}$  is a vector containing the

negative of the residual errors required to bring the error for each equation to zero. The matrix  $\partial \mathbf{f} / \partial \mathbf{U}$  is composed of analytical derivatives of each of the increments of the residuals with respect to the dependent variables. The derivatives needed to construct the *Jacobian* are easily obtained and summarized in *Appendix 12.8*. The velocities are assumed to exhibit linear dependence on the pressures, this allowing one to obtain the derivatives of the velocities with respect to pressure directly from the momentum equations (12.89). The result is given also in *Appendix 12.8*. The linear variation of the velocity with respect to pressure is given by

$$\delta V_{lm}^n = -RVel_{lm}(\delta p_m - \delta p). \quad (12.134)$$

Note that this form of the velocity increment equation is applicable only after an explicit pass of the explicit parts of the momentum equations.

After all of the derivatives for the above equations have been calculated, the system of equations (12.133) is solved analytically by elimination for the unknown increments. In order to eliminate the volume fraction increments we divide the mass conservation equation by  $\partial f_i^\alpha / \partial \alpha_i$  and sum the resulting equations.

Replacing  $\sum_{l=1}^{l_{\max}} \delta \alpha_l$  by  $-f_\Sigma$  I eliminate the  $\delta \alpha_l$  values

$$\begin{aligned} & \sum_{l=1}^{l_{\max}} \left( \frac{\partial f_l^\alpha / \partial s_l}{\partial f_l^\alpha / \partial \alpha_l} \delta s_l + \sum_{i=2}^{i_{\max}} \frac{\partial f_l^\alpha / \partial C_{il}}{\partial f_l^\alpha / \partial \alpha_l} \delta C_{il} + \frac{\partial f_l^\alpha / \partial p}{\partial f_l^\alpha / \partial \alpha_l} \delta p + \sum_{m=1}^6 \frac{\partial f_l^\alpha / \partial p_m}{\partial f_l^\alpha / \partial \alpha_l} \delta p_m \right) \\ & = f_\Sigma - \sum_{l=1}^{l_{\max}} \frac{f_l^\alpha}{\partial f_l^\alpha / \partial \alpha_l} = f_\Sigma - \sum_{l=1}^{l_{\max}} (\alpha_l - \alpha_{l0}) = \sum_{l=1}^{l_{\max}} \alpha_{l0} - 1. \end{aligned} \quad (12.135)$$

The next step is to solve the entropy, concentration and particle density equations for the entropy, concentration and particle density increments. Thus for  $\partial f_i^s / \partial s_i > 0$ ,

$$\delta s_i = -\frac{f_i^s}{\partial f_i^s / \partial s_i} - \frac{\partial f_i^s / \partial p}{\partial f_i^s / \partial s_i} \delta p - \sum_{m=1}^6 \frac{\partial f_i^s / \partial p_m}{\partial f_i^s / \partial s_i} \delta p_m \quad (12.136)$$

$$\delta C_{il} = -\frac{f_{il}^C}{\partial f_{il}^C / \partial C_{il}} - \frac{\partial f_{il}^C / \partial p}{\partial f_{il}^C / \partial C_{il}} \delta p - \sum_{m=1}^6 \frac{\partial f_{il}^C / \partial p_m}{\partial f_{il}^C / \partial C_{il}} \delta p_m \quad (12.137)$$

For  $\partial f_i^s / \partial s_i = 0$ ,

$$\delta s_i = 0, \quad (12.138)$$

$$\delta C_{il} = 0. \quad (12.139)$$

Similarly for  $\partial f_l^n / \partial n_l > 0$

$$\delta n_l = -\frac{f_l^n}{\partial f_l^n / \partial n_l} - \frac{\partial f_l^n / \partial p}{\partial f_l^n / \partial n_l} \delta p - \sum_{m=1}^6 \frac{\partial f_l^n / \partial p_m}{\partial f_l^n / \partial n_l} \delta p_m, \quad (12.140)$$

and for  $\partial f_l^n / \partial n_l = 0$ ,

$$\delta n_l = 0. \quad (12.141)$$

Note that computation of the increments is only permissible if  $\partial f_l^s / \partial s_l > 0$ . If  $\partial f_l^s / \partial s_l = 0$ , the velocity field  $l$  did not exist, and will not originate in the time step considered. In this case, the properties of the velocity field  $l$  are not defined and not required for the computation of the pressure field as  $\partial f_l^\alpha / \partial s_l = 0$ ,  $\partial f_l^\alpha / \partial C_{il} = 0$ .

If the velocity field  $l$  did not exist at the old time level but has just originated, this means that

$$\alpha_{la} \leq 0 + \varepsilon \text{ and } \partial f_l^s / \partial s_l > 0, \quad (12.142)$$

I then obtain from Eqs. (12.125) through (12.127) the exact initial values of the entropies, of the concentrations, and of the particle number densities

$$s_l = s_{l0} > 0, \quad (12.143)$$

$$C_{il} = C_{il0}, \quad 0 \leq C_{il} \leq 1, \quad (12.144)$$

$$n_l = n_{l0} \geq 0, \quad (12.145)$$

which obviously do not depend on the non-defined old values  $s_{la}$ ,  $C_{ila}$  and  $n_{la}$ , respectively. The necessary temperature inversion is performed by the iterations using Eq. (3.129)

$$T_l = T_l^- \exp \left\{ \left[ s_l - s_l^- \left( T_l^-, p, C_{il, i=1, i_{\max}} \right) \right] / c_{pl} \right\}. \quad (12.146)$$

Here the superscript “-” denotes the previous iteration value. The initial density is computed using the state equation of the new velocity field

$$\rho_l = \rho_l \left( T_l, p, C_{il, i=2, i_{\max}} \right) \quad (12.147)$$

in this case only. In all other cases, the temperature and the density are calculated using the linearized equation of state for each velocity field (3.106) and (3.137):

$$T_l = T_{la} + \frac{\partial T_l}{\partial p} \delta p + \frac{\partial T_l}{\partial s_l} \delta s_l + \sum_{i=2}^{i_{\max}} \frac{\partial T_l}{\partial C_{il}} \delta C_{il}, \quad (12.148)$$

$$\rho_l = \rho_{la} + \frac{\partial \rho_l}{\partial p} \delta p + \frac{\partial \rho_l}{\partial s_l} \delta s_l + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} \delta C_{il}. \quad (12.149)$$

The above demonstrates the remarkable property of the semi-conservative form of the discretized entropy and concentration equations that they do not depend on the field volume fractions in the actual computational cell. We therefore replace the entropy and the concentration increments in Eq. (12.135), substituting for  $\partial f_l^s / \partial s_l > 0$

$$\overline{\alpha s_l} = \frac{\partial f_l^\alpha / \partial s_l}{\partial f_l^\alpha / \partial \alpha_l} \frac{\partial f_l^s}{\partial s_l} = \frac{\alpha_l}{\rho_l} \frac{\partial \rho_l}{\partial s_l} \frac{\partial f_l^s}{\partial s_l}, \quad (12.150)$$

$$\overline{\alpha C_{il}} = \frac{\partial f_l^\alpha / \partial C_{il}}{\partial f_l^\alpha / \partial \alpha_l} \frac{\partial f_{il}^c}{\partial C_{il}} = \frac{\alpha_l}{\rho_l} \frac{\partial \rho_l}{\partial C_{il}} \frac{\partial f_{il}^c}{\partial C_{il}}, \quad (12.151)$$

and for  $\partial f_l^s / \partial s_l = 0$

$$\overline{\alpha s_l} = 0, \quad (12.152)$$

$$\overline{\alpha C_{il}} = 0, \quad (12.153)$$

and finally obtain the so called pressure correction equation which is the discrete analog of the *Poisson*-type equation for multi-phase flows

$$c \delta p + \sum_{m=1}^6 c_m \delta p_m = D^*, \quad (12.154)$$

where

$$\begin{aligned} c &= \sum_{l=1}^{l_{\max}} \left( \frac{\partial f_l^\alpha / \partial p}{\partial f_l^\alpha / \partial \alpha_l} - \overline{\alpha s_l} \frac{\partial f_l^s}{\partial p} - \sum_{i=2}^{i_{\max}} \overline{\alpha C_{il}} \frac{\partial f_{il}^c}{\partial p} \right) \\ &= \sum_{l=1}^{l_{\max}} \left[ \frac{\partial f_l^\alpha / \partial p}{\partial f_l^\alpha / \partial \alpha_l} - \frac{\alpha_l}{\rho_l} \frac{1}{\partial f_l^s / \partial s_l} \left( \frac{\partial \rho_l}{\partial s_l} \frac{\partial f_l^s}{\partial p} + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} \frac{\partial f_{il}^c}{\partial p} \right) \right] \\ &= \sum_{l=1}^{l_{\max}} \frac{1}{\rho_l} \frac{\partial f_l^\alpha}{\partial n_l} \left[ \frac{\partial f_l^\alpha}{\partial p} - \left( \frac{\partial f_l^s}{\partial p} - \sum_{i=2}^{i_{\max}} \frac{\partial s_l}{\partial C_{il}} \frac{\partial f_{il}^c}{\partial p} + \sum_{i=2}^{i_{\max}} c_{pl} \Delta \bar{R}_{il} \frac{\partial f_{il}^c}{\partial p} \right) \right] \\ &= \frac{1}{\rho a^2} + \sum_{m=1}^6 c_m^*, \end{aligned} \quad (12.155)$$

and

$$c_m^* = \beta_m \sum_{l=1}^{l_{\max}} \frac{RVel_{lm}}{\rho_l \frac{\partial f_l^n}{\partial n_l}} \left[ \xi_{lm+} \alpha_l \rho_l + \xi_{lm-} (\alpha_l \rho_l)_m \left( \begin{array}{c} 1 - (s_{lm} - s_l) + \sum_{i=2}^{i_{\max}} \frac{\partial s_l}{\partial C_{il}} (C_{ilm} - C_{il}) \\ - \sum_{i=2}^{i_{\max}} c_{pl} \Delta \bar{R}_{il} (C_{ilm} - C_{il}) \end{array} \right) \right], \quad (12.156)$$

$$\begin{aligned} c_m &= \sum_{l=1}^{l_{\max}} \left( \frac{\partial f_l^\alpha / \partial p_m}{\partial f_l^\alpha / \partial \alpha_l} - \overline{\alpha s_l} \frac{\partial f_l^s}{\partial p_m} - \sum_{i=2}^{i_{\max}} \overline{\alpha C_{il}} \frac{\partial f_{il}^C}{\partial p_m} \right) \\ &= \sum_{l=1}^{l_{\max}} \left[ \frac{\partial f_l^\alpha / \partial p_m}{\partial f_l^\alpha / \partial \alpha_l} - \frac{\alpha_l}{\rho_l} \frac{1}{\partial f_l^s / \partial s_l} \left( \frac{\partial \rho_l}{\partial s_l} \frac{\partial f_l^s}{\partial p_m} + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} \frac{\partial f_{il}^C}{\partial p_m} \right) \right] \\ &= \sum_{l=1}^{l_{\max}} \frac{1}{\rho_l \frac{\partial f_l^n}{\partial n_l}} \left[ \partial f_l^\alpha / \partial p_m - \left( \frac{\partial f_l^s}{\partial p_m} - \sum_{i=2}^{i_{\max}} \frac{\partial s_l}{\partial C_{il}} \frac{\partial f_{il}^C}{\partial p_m} + \sum_{i=2}^{i_{\max}} c_{pl} \Delta \bar{R}_{il} \frac{\partial f_{il}^C}{\partial p_m} \right) \right] \\ &= -\beta_m \sum_{l=1}^{l_{\max}} \frac{1}{\rho_l \frac{\partial f_l^n}{\partial n_l}} \left[ \xi_{lm+} \alpha_l \rho_l RVel_{lm} + \xi_{lm-} (\alpha_l \rho_l)_m \left( RVel_{lm} - \frac{V_{lm}^n}{\rho_{lm} a_{lm}^2} \right) \right. \\ &\quad \left. \left( 1 - (s_{lm} - s_l) + \sum_{i=2}^{i_{\max}} \frac{\partial s_l}{\partial C_{il}} (C_{ilm} - C_{il}) - \sum_{i=2}^{i_{\max}} c_{pl} \Delta \bar{R}_{il} (C_{ilm} - C_{il}) \right) \right] \\ &= \beta_m \sum_{l=1}^{l_{\max}} \frac{1}{\rho_l \frac{\partial f_l^n}{\partial n_l}} \xi_{lm-} V_{lm}^n \frac{\alpha_{lm}}{a_{lm}^2} - c_m^*, \end{aligned} \quad (12.157)$$

$$\begin{aligned} D^* &= f_\Sigma - \sum_{l=1}^{l_{\max}} \left( \frac{f_l^\alpha}{\partial f_l^\alpha / \partial \alpha_l} - \overline{\alpha s_l} f_l^s - \sum_{i=2}^{i_{\max}} \overline{\alpha C_{il}} f_{il}^C \right) \\ &= f_\Sigma - \sum_{l=1}^{l_{\max}} \left[ \frac{f_l^\alpha}{\partial f_l^\alpha / \partial \alpha_l} - \frac{\alpha_l}{\rho_l} \frac{1}{\partial f_l^s / \partial s_l} \left( \frac{\partial \rho_l}{\partial s_l} f_l^s + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} f_{il}^C \right) \right] \\ &= f_\Sigma - \sum_{l=1}^{l_{\max}} \frac{1}{\rho_l \frac{\partial f_l^n}{\partial n_l}} \left[ f_l^\alpha - \frac{1}{\partial \rho_l}{\partial s_l} \left( \frac{\partial \rho_l}{\partial s_l} f_l^s + \sum_{i=2}^{i_{\max}} \frac{\partial \rho_l}{\partial C_{il}} f_{il}^C \right) \right] \end{aligned}$$

$$= f_{\Sigma} - \sum_{l=1}^{i_{\max}} \frac{1}{\rho_l} \frac{\partial f_l^n}{\partial n_l} \left[ f_l^\alpha - \left( f_l^s - \sum_{i=2}^{i_{\max}} \frac{\partial s_l}{\partial C_{il}} f_{il}^C + \sum_{i=2}^{i_{\max}} c_{pl} \Delta \bar{R}_{il} f_{il}^C \right) \right], \quad (12.158)$$

for  $\partial f_l^s / \partial s_l > 0$ . Using Eqs. (3.135) and (3.136), see also in *Kolev* (1990), we obtain for the ratio

$$\begin{aligned} \frac{\partial \rho_l / \partial C_{il}}{\partial \rho_l / \partial s_l} &= \frac{c_{pl}}{T_l} \frac{\left( \frac{\partial \rho_l}{\partial C_{il}} \right)_{p, T_l, \text{all } C's \text{ except } C_{il}}}{\left( \frac{\partial \rho_l}{\partial T_l} \right)_{p, \text{all } C's}} - \left( \frac{\partial s_l}{\partial C_{il}} \right)_{p, T_l, \text{all } C's \text{ except } C_{il}} \\ &= c_{pl} \Delta \bar{R}_{il} - \frac{\partial s_l}{\partial C_{il}}, \end{aligned} \quad (12.159)$$

bearing in mind Eqs. (3.57) and (3.58), see also in *Kolev* (1990),

$$\Delta \bar{R}_{il} = \frac{\left( \frac{\partial \rho_l}{\partial C_{il}} \right)_{p, T_l, \text{all } C's \text{ except } C_{il}}}{T_l \left( \frac{\partial \rho_l}{\partial T_l} \right)_{p, \text{all } C's}} = \frac{\frac{1}{T_l} \sum_{i=2}^{i_{\max}} \left[ \frac{1}{(\partial \rho_{il} / \partial p_{il})_{T_l}} - \frac{1}{(\partial \rho_{1l} / \partial p_{1l})_{T_l}} \right]}{-\frac{1}{\rho_l} \sum_{i=1}^{i_{\max}} \left( \frac{\partial \rho_{il}}{\partial T_l} \right)_{p_{il}} / \left( \frac{\partial \rho_{il}}{\partial p_{il}} \right)_{T_l}} \quad (12.160)$$

the term  $\Delta \bar{R}_{il}$  can be interpreted as the relative deviation of the pseudo gas constant for the component  $l$  from the pseudo gas constant of component 1. For ideal gas mixtures this term is simply  $\Delta \bar{R}_{il} = (R_{il} - R_{1l}) / R_{1l}$ .

Obviously if  $\Delta \tau \rightarrow 0$ ,  $c_m \rightarrow 0$ , and  $c \rightarrow \frac{1}{\rho a^2}$  because

$$\frac{\partial f_l^n}{\partial n_l} = \frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 \beta_m \xi_{lm+} \mathbf{V}_{lm}^n > 0.$$

Note that the coefficients in this equation are analytically defined. I emphasize this point as in the current state of the art for modeling of multi-dimensional, multi-phase flows this step is performed numerically e.g. *Liles and Reed* (1978), *Thurgood et al.* (1983), *Kolev* (1986), *Bohl et al.* (1988), this being a much more time consuming approach.

The linear variation of the pressure in cell  $(i, j, k)$  as a function of the surrounding cell pressures is given by Eq. (12.154). A similar equation can be derived for each cell in the mesh. This set of equations for the pressure variation in each mesh cell must be simultaneously satisfied. The solution to this equation set can be obtained by direct inversion for problems involving only a few mesh cells,

or using the *Gauss-Seidel* iteration technique for problems involving a large number of mesh cells. In IVA4 the pressure corrector method exploits the algebraic solvers already discussed in Section 10. For brief introduction see Appendix 12.9.

The computer time required to solve Eq. (12.133) can be greatly reduced if the non-linear coefficients  $dV_{lm}^n$  and  $RVel_{lm}$  are assumed to remain constant during a time step  $\Delta\tau$ , with the solution then obtained only for the linearized system (12.133). Checks are made on the value of each of the new time variables to assure that the variations of the new time variables from the old variables lie within reasonable limits. If the new time variables have non-physical values. e.g., void fractions less than zero or greater than one, or if the variation of the new time variable from the old variable is implausibly large, then the solution is run back to the beginning of the time step, the variables are set to their old time value, the time step is reduced and the computation repeated. This is implemented to ensure that the linearized equations are sufficiently representative of the non-linear equations to provide an acceptable level of calculation accuracy. The time step size is controlled as a function of the rate of change of the independent variables for the same reason.

## 12.13 Integration procedure: implicit method

The integration procedure is a logical sequence of the steps needed to obtain a set of dependent variables for each computational cell, these variables satisfying the conservation equations for each time step under the simplifying assumptions introduced and the working hypothesis for any given class of initial and boundary conditions.

The following procedure was found to lead to unconditionally stable solutions.

1. Read the information defining the problem and the required integration accuracy:

- Logical control information;
- Geometry;
- Initial conditions;
- Boundary conditions;
- External sources;
- Variable permeabilities;
- Heat structure definitions.

Perform as many time steps as required to reach the prescribed process time. A single time cycle consists of the following steps:

2. Perform computations before starting the outer iterations:

Numerics:

- Impose actual geometry;
- Store old time level information;
- Impose boundary conditions;



- Impose structure heat sources;

Cell by cell constitutive relations:

- Estimate the equations of state for each component of the velocity field (thermo-physical and transport properties for simple constituents);
- Estimate the equations of state for the mixtures of which each field consists;
- Identify the flow pattern;
- Estimate constitutive relationships dependent on flow pattern;
- Compute energy and mass source terms for each particular cell for the flow pattern identified;
- Compute interfacial drag and virtual mass coefficients;
- Compute wall-fluid interaction drag coefficients;

Numerics:

- Compute coefficients of the discretized momentum equations;
- Compute the linearized coefficients for solving the local momentum equation with respect to the local field velocities for each direction;
- Estimate velocities based on the old time level pressures;
- Impose cyclic boundary conditions in the case of  $\theta$  closed cylindrical geometry.

3. Perform outer iterations:

- Estimate new  $s_l, \approx C_{nl}, \approx \alpha_l$  ;
- Repeat this step as many times as necessary to satisfy with prescribed accuracy all entropy and concentration equations and the appropriate two of the three mass conservation equations;
- Compute the coefficients of the pressure equation;
- Solve the pressure equation for  $\tau + \Delta \tau$  ;
- Perform convergence, accuracy and time step control;
- Compute velocities for  $\tau + \Delta \tau$  ;
- Impose cyclic boundary conditions if required;
- Compute  $s_l, C_{nl}, \alpha_l$  for  $\tau + \Delta \tau$  ;
- Control convergence;
- Check against general accuracy requirements, if not fulfilled perform the next outer iteration; if no convergence is achieved reduce time step, recover the old time level situation and repeat the outer iterations until convergence is achieved;
- Perform the next outer iterations until all general accuracy requirements are satisfied;

4. Perform computations after successful time step:

- Perform temperature inversion;
- Optimize time step for the next integration step;

5. Write restart information for prescribed step frequency before specified CPU time has elapsed, and at the end of the simulation.
6. Write information for post processing of the results.

## 12.14 Time step and accuracy control

The time step limitation dictated by the linear stability analysis for implicit donor-cell methods is

$$\Delta\tau_{CFL} < \max\left(\frac{\Delta r_h}{u_l}, \frac{r_h \Delta\theta_h}{v_l}, \frac{\Delta z_h}{w_l}\right) \quad (12.161a)$$

for all cells. This is the so called material *Courant, Friedrichs and Levi* (MCFL) criterion. Numerous numerical experiments have shown that this method can work properly in many cases with larger time steps. The MCFL criterion is nevertheless retained to ensure convergence in all cases. In addition to this limitation, there are two reasons leading to further time step limits: (a) linearization limits, (b) definition limits for the dependent variables:

- (a) Linearization limits: The change of the dependent variables within a time step in each computational cell should not exceed a prescribed value. This condition is associated with the linearization of the strongly non-linear system of 21 PDEs and the state equations for each time step, which is not considered in the classical *von Neumann* linear stability analysis of 1D numerical scheme for differential equations with constant coefficients.
- (b) Definition limits for the dependent variables: We illustrate this problem by the following example. The velocity field mass is non-negative

$$\Delta\tau_{\max}^{\alpha} < (\alpha_l^* \rho_l \gamma_v - \alpha_{la} \rho_{la} \gamma_{va}) / \left\{ \gamma_v \mu_l - \sum_m [b_{lm+} \alpha_l \rho_l - b_{lm-} (\alpha_l \rho_l)_m] \right\}, \quad (12.161b)$$

where  $\alpha_l^* = 0$  for a decreasing volumetric fraction, that is for  $\bar{\alpha}_l < \alpha_{la}$ . For an increasing volumetric fraction,  $\bar{\alpha}_l > \alpha_{la}$ , the volumetric fraction of the velocity field cannot exceed the value of one by definition,  $\alpha_l^* = 1$ .

The outer iterations are considered as successfully completed if pressure and velocity increments from iteration to iteration and the relative mass conservation error reach values smaller than those prescribed. Additional time step optimization is imposed in order to keep the time step such as to have only a prescribed number of outer iterations, e.g. 6.

High order methods for discretization of the time derivatives require storing information for past time steps. This is the limitation that forces most of the authors to use in the multiphase flows first order explicit or implicit Euler methods. Such schemes can be easily extended to second order by using the trapezoid rule for integration over the time resulting in the popular *Crank-Nicolson* method

$$\Phi - \Phi_a = \frac{1}{2} \Delta \tau [f(\tau, \Phi_a) + f(\tau + \Delta \tau, \Phi)],$$

where  $f$  contains the remaining part of the discretized equation. Usually this method is using as a two step predictor corrector method

$$\Phi^* = \Phi_a + \frac{1}{2} \Delta \tau f(\tau, \Phi_a),$$

$$\Phi = \Phi^* + \frac{1}{2} \Delta \tau f(\tau + \Delta \tau, \Phi^*).$$

The information required is stored in any case for the old and for the new time level. The corrector step can be repeated until the solution does not change any more within prescribed limits. Alternatively the method of *Adams-Bashforth*, see in *Ferziger and Paric* (2002) p.139, can also be used

$$\Phi - \Phi_a = \frac{1}{2} \Delta \tau [3f(\tau, \Phi_a) - f(\tau + \Delta \tau, \Phi)].$$

## 12.15 High order discretization schemes for convection-diffusion terms

### 12.15.1 Space exponential scheme

*Patankar and Spalding* (1972) propose the  $\Phi$  profile between the locations  $(k, k+1)$  and  $(k-1, k)$  to be approximated by means of the following functions

$$\frac{\Phi^* - \Phi}{\Phi_{k+1} - \Phi} = \frac{\exp\left(Pe \frac{z^* - z}{\Delta z_h}\right) - 1}{\exp(Pe) - 1}, \quad Pe = \frac{F \Delta z_h}{\Gamma}, \quad z \leq z^* \leq z_{k+1} \quad (12.162)$$

$$\frac{\Phi^* - \Phi_{k-1}}{\Phi - \Phi_{k-1}} = \frac{\exp\left(Pe_{k-1} \frac{z^* - z_{k-1}}{\Delta z_{h,k-1}}\right) - 1}{\exp(Pe_{k-1}) - 1}, \quad Pe_{k-1} = \frac{F_{k-1} \Delta z_{h,k-1}}{\Gamma_{k-1}}, \quad z_{k-1} \leq z^* \leq z. \quad (12.163)$$

For more details see *Patankar* (1980). Some reason for the choice even of this functional form is the solution of the equation

$$\frac{d}{dz} \left( F \Phi - \Gamma \frac{d\Phi}{dz} \right) = 0, \quad (12.164)$$

by the following boundary conditions

$$z^* = z, \quad \Phi^* = \Phi, \quad (12.165)$$

$$z^* = z_{k+1}, \quad \Phi^* = \Phi_{k+1} \quad (12.166)$$

for the first equation and

$$z^* = z_{k-1}, \quad \Phi^* = \Phi_{k-1}, \quad (12.167)$$

$$z^* = z, \quad \Phi^* = \Phi \quad (12.168)$$

for the second. *Tomiyama* (1990) found that the analytical solution of a transient convection-diffusion equation without source terms is also fully consistent with the functions (12.162) and (12.163). These functions have interesting properties, namely, the total convection-diffusion flow on the boundaries 5 and 6

$$J_{k-1} = \left( F\Phi^* - \Gamma \frac{d\Phi^*}{dz} \right)_{k-1} = F_{k-1} \left[ \Phi_{k-1} + \frac{\Phi_{k-1} - \Phi}{\exp(Pe_{k-1}) - 1} \right] \quad (12.169)$$

$$J = \left( F\Phi^* - \Gamma \frac{d\Phi^*}{dz} \right) = F \left[ \Phi + \frac{\Phi - \Phi_{k+1}}{\exp(Pe) - 1} \right] \quad (12.170)$$

depends on  $z$  only by the averaging method for the computation of  $\Gamma$  and  $\Gamma_{k-1}$ , but *not on  $z$  alone*. Using the so obtained flows on the boundaries of the computational cell we obtain the finite difference analog of

$$\frac{d}{dz} \left( F\Phi - \Gamma \frac{d\Phi}{dz} \right) - \Phi \frac{dF}{dz} = 0 \quad (12.171)$$

namely

$$\begin{aligned} & - \underbrace{\frac{F}{\Delta z} \frac{\Gamma}{\Delta z_h} \frac{1}{\exp(Pe) - 1}}_{b_{k+1}} (\Phi_{k+1} - \Phi) - \underbrace{\frac{F_{k-1}}{\Delta z} \frac{\Gamma_{k-1}}{\Delta z_{h,k-1}} \frac{Pe_{k-1} \exp(Pe_{k-1})}{\exp(Pe_{k-1}) - 1}}_{b_{k-1}} (\Phi_{k-1} - \Phi) \\ & = (b_{k+1} + b_{k-1})\Phi - b_{k+1}\Phi_{k+1} - b_{k-1}\Phi_{k-1}, \end{aligned} \quad (12.172)$$

where

$$b_{k+1} = \frac{F}{\Delta z} \frac{\Gamma}{\Delta z_h} \frac{1}{\exp(Pe) - 1}, \quad (12.173)$$

$$b_{k-1} = \frac{F_{k-1}}{\Delta z} \frac{\Gamma_{k-1}}{\Delta z_{h,k-1}} \frac{Pe_{k-1} \exp(Pe_{k-1})}{\exp(Pe_{k-1}) - 1}. \quad (12.174)$$

The computation of exponents is in general *time consuming*. *Patankar* (1981) proposed a very accurate *approximation* of the coefficients

$$b_{k+1} = \frac{1}{\Delta z} \left\{ -F \frac{1}{2} [1 - \text{sign}(F)] + \frac{\Gamma}{\Delta z_h} A(|Pe|) \right\} \quad (12.175)$$

$$b_{k-1} = \frac{1}{\Delta z} \left\{ -F_{k-1} \frac{1}{2} [1 + \text{sign}(F_{k-1})] + \frac{\Gamma_{k-1}}{\Delta z_{h,k-1}} A(|Pe_{k-1}|) \right\}, \quad (12.176)$$

where

$$A(|Pe|) = \max \left[ 0, (1 - 0.1|Pe|)^5 \right]. \quad (12.177)$$

Comparing these coefficients with the coefficients obtained with the linear profile, we see that for values of the local grid *Peclet* numbers  $|Pe|$  greater than about 6 the diffusion flow is reduced to zero and the both schemes are nearly equivalent. By construction of a numerical algorithm two methods can be easily used simultaneously, programming the second one and setting optionally  $A = 1$  if the use of the first method is desired.

Using such profile approximations for the discretization poses some *unanswered questions*:

- The profile of  $\Phi$  in the case of no *source terms* (of  $\Phi$ ) has to be not the same as in the case of source terms different from zero;
- The *convection in the other two directions* in the three-dimensional flow will influence the profile in the first direction and vice versa, which is not taken into account into the above discretization schemes;
- During *strong transients* it is difficult to guess an appropriate profile of  $\Phi$  ;
- For most cases of practical interest, especially in three dimensions where very fine meshes are out of the question, the actual component grid *Peclet* numbers are likely to be orders of magnitude larger than 2 or 6 throughout the bulk of the flow domain. This means that the donor-cell upwind and the space exponential schemes are operating as first order upwind almost everywhere in the flow field except for a very small fraction of grid points near boundaries and stagnation regions where the convecting velocities are small. Thus instead of solving a high-convection problem, these methods simulate a low convection problem in which the effective local component grid *Peclet* number can never be greater than 2 *Leonard and Mokhtari (1990)*.

There are attempts to answer the first two questions. *Chow and Tien (1978)* p. 91, successfully take into account some influence of the source term in their profile approximation. The resulting scheme is more accurate than the above described two schemes and possesses unconditional stability. *Prakash (1984)* takes into account the influence of the flow in the other two directions in his profile approximation. *Patel et al. (1985)*, comparing the qualities of 7 discretization schemes with analytical one- and two-dimensional solutions, emphasize the advantages of the above discussed local analytical discretization scheme, and of the *Leonard super upwind scheme (1978)*, *Leonard and Mokhtari (1990)*, *Leonard (1990)*. In a later publication, *Patel and Markatos (1986)* compare 8 new schemes

and emphasize the advantages of the quadratic-upstream-extended-revised scheme proposed by *Pollard* and *Ali* (1982).

*Günther* comparing the qualities of 7 discretization schemes, emphasizes the advantages of the local-analytical-method of second order of accuracy in space (donor-cell) proposed in *Günther* (1988).

The general tendency in the development of such schemes is to use further the donor-cell concept and increase the accuracy of the space approximations using schemes with an order of accuracy higher than one.

### 12.15.2 High order upwinding

The main idea behind the high order differencing methods is the approximation of the  $\Phi$  function by using information from more than two adjacent cells. Usually the *Lagrange* polynomial parabolic approximation  $m$ -th order is used

$$\Phi(z_h) = \sum_m \frac{\sum_{l \neq m} (z_k - z_l)}{\sum_{l \neq m} (z_m - z_l)} \Phi(z_m) \tag{12.178}$$

where, e.g.

$$m = k - 2, k - 1, k, k + 1, k + 2.$$

Numerical schemes for convection problems are usually constructed non-symmetrically, which means using more information for the approximation of the actual  $\Phi$  from the upwind direction than from the downwind direction. Compared with symmetrical schemes the non-symmetrical schemes are more stable. Let us illustrate the main idea by the example of the successful third-order upwind scheme QUICK - see *Leonard* and *Mokharti* (1990):

$$\Phi_3 = \text{Eq. (12.178)} \tag{12.179}$$

where

$$m = k - 1, k, k + 1 \quad \text{for } w \geq 0,$$

$$m = k, k + 1, k + 2 \quad \text{for } w < 0,$$

and

$$\Phi_6 = \text{Eq. (12.178)} \tag{12.180}$$

where

$$m = k - 2, k - 1, k \quad \text{for } w \geq 0,$$

$$m = k - 1, k, k + 1 \quad \text{for } w_{k-1} < 0.$$

For approximation of the diffusion terms QUICK uses the usual second order symmetric approximation. The disadvantage of the high order upwind schemes is the so called undershoot and overshoot if one simulates sharp discontinuities in the

integration region. What is helpful to resolve this problem is a simple limiting strategy which maintains the good resolution of higher order upwinding while eliminating undershoots and overshoots without introducing artificial diffusion or destroying conservation. *Leonard and Mokhtari (1990)* propose an extremely simple technique which can be extended to arbitrary high order accuracy. Next we explain the main idea of this proposal. In the case of positive direction of the velocity, the character of the functional form of  $\Phi_m$ ,  $m = k - 1, k, k + 1$  is checked. One distinguishes between locally monotonic and locally non-monotonic behavior:

1) Locally monotonic behavior is defined if

$$\Phi_{k+1} > \Phi > \Phi_{k-1} \quad (12.181)$$

or

$$\Phi_{k-1} < \Phi < \Phi_{k+1}. \quad (12.182)$$

In this case the convected face-value,  $\Phi_5$ , computed by the high order differencing should lie between adjacent node values

$$\Phi < \Phi_5 < \Phi_{k+1}, \quad (12.183)$$

or

$$\Phi > \Phi_5 > \Phi_{k-1}, \quad (12.184)$$

respectively. Otherwise  $\Phi_5$  is limited by one of the adjacent node values. If

$$\Phi_{k-1} = \Phi \quad (12.185)$$

*Leonard and Mokhtari* use simply the upwind value

$$\Phi_5 = \Phi \quad (\equiv \Phi_{k-1}). \quad (12.186)$$

2) Locally non-monotonic behavior is defined if

$$\Phi > \Phi_{k+1} \quad \text{and} \quad \Phi > \Phi_{k-1}, \quad (12.187)$$

or

$$\Phi < \Phi_{k+1} \quad \text{and} \quad \Phi < \Phi_{k-1}. \quad (12.188)$$

In this case one can use the linear extrapolation of the values of the two adjacent upwind points, namely

$$\Phi_5 = \Phi + \frac{z_n - z}{z - z_{k-1}} (\Phi - \Phi_{k-1}), \quad (12.189)$$

which is simply a reduction of order of the approximation to second order upwind.

One should be careful in applying the high order differencing in a region divided by permeabilities which can be zeros. If e.g. in the above discussed cases  $\gamma_{z,k-1} = 0$  we should use simply the first order upwind value

$$\Phi_s = \Phi_k . \tag{12.190}$$

The existing codes using the classical donor-cell upwind or local analytical schemes can be easily upgraded to incorporate cost-effective high order non-oscillatory methods in order to use them in cases of coarse grids where the local grid *Peclet* number is greater than two or six respectively.

### 12.15.3 Constrained interpolation profile (CIP) method

*Yabe, Xiao and Utsumi* (2001) published a family of methods with the remarkable quality of resolving discontinuities with much smaller number of cells than any other known methods – see in *Tanaka, Nakamura and Yabe* (2000) p. 588, Fig. 19. The basic idea of this group of methods is to use the primitive dependent variables and their derivatives as a set of dependent variables. For the first group the conservation equations are used and for the second group the corresponding derivatives of these equations are used. Instead of reviewing here all the methods of this group, we give few examples of the so called exactly conservative scheme for the CIP method. The remarkable quality of this scheme is the enforced conservation of given properties even by using a non-conservative form of the conservation equations.

#### 12.15.3.1 Exactly conservative scheme for transport equations in non-conservative form

*Tanaka, Nakamura and Yabe* (2000) proposed to approximate the  $f$  profile between the locations  $z_{h,k-1} \leq z^* \leq z_h$  by means of the *fourth order* polynomial spline function

$$f^* - f_{k-1} = a(z^* - z_{h,k-1})^4 + b(z^* - z_{h,k-1})^3 + c(z^* - z_{h,k-1})^2 + g_{k-1}(z^* - z_{h,k-1}), \tag{12.191}$$

which already satisfies the continuity and the smoothness conditions at the left boundary  $z^* = z_{h,k-1}$ ,

$$f^* = f_{k-1}, \quad \frac{df}{dz^*} = g_{k-1}, \tag{12.192}$$

and has to satisfy the continuity and the smoothness conditions also at the right boundary  $z^* = z_h$

$$f^* = f, \quad \frac{df}{dz^*} = g. \tag{12.193}$$



This leads to the additional two equations for computing the unknown coefficients  $a$ ,  $b$ , and  $c$

$$\Delta z_{k-1}^4 a + \Delta z_{k-1}^3 b + \Delta z_{k-1}^2 c = f - f_{k-1} - g_{k-1} \Delta z_{k-1}, \tag{12.194}$$

$$4\Delta z_{k-1}^3 a + 3\Delta z_{k-1}^2 b + 2\Delta z_{k-1} c = g - g_{k-1}. \tag{12.195}$$

In addition the authors enforce the so called continuity constraint. We use here the integral as proposed by the authors in *Tanaka, Nakamura and Yabe (2000)* divided by the distance between two points  $\Delta z_{k-1} = z_h - z_{h,k-1}$  which is the weighted averaged over this distance,

$$f_{av} = \frac{1}{\Delta z_{k-1}} \int_{z_{k-1}}^z f^* dz^*, \tag{12.196}$$

resulting in the third equation

$$\frac{1}{5} \Delta z_{k-1}^4 a + \frac{1}{4} \Delta z_{k-1}^3 b + \frac{1}{3} \Delta z_{k-1}^2 c = f_{av} - g_{k-1} \frac{1}{2} \Delta z_{k-1} - f_{k-1}, \tag{12.197}$$

required to compute the three unknown constants. Solving with respect the coefficients one obtains

$$a = \frac{30}{\Delta z_{k-1}^4} \left[ f_{av} - \frac{1}{2} (f + f_{k-1}) + (g - g_{k-1}) \frac{1}{12} \Delta z_{k-1} \right], \tag{12.198}$$

$$b = \frac{4}{\Delta z_{k-1}^3} \left[ -15 f_{av} + 7f + 8f_{k-1} - \left( g - \frac{3}{2} g_{k-1} \right) \Delta z_{k-1} \right], \tag{12.199}$$

$$c_{k-1} = \frac{3}{\Delta z_{k-1}^2} \left[ 10 f_{av} - 4f - 6f_{k-1} + (g - 3g_{k-1}) \frac{1}{2} \Delta z_{k-1} \right]. \tag{12.200}$$

The CIP method is illustrated on integrating the following simple equation

$$\frac{\partial f}{\partial \tau} + w \frac{\partial f}{\partial z} = 0, \tag{12.201}$$

written in canonical form  $df/d\tau = 0$  along the characteristic line defined by  $dz/d\tau = w$ . The solution for the point  $k - 1$  is

$$f_{k-1}^{n+1} (\tau + \Delta \tau, z_{h,k-1}) = f_{k-1}^* (\tau, z_{h,k-1} - w \Delta \tau), \tag{12.202}$$

or in terms of the already find approximations for  $w > 0$

$$f_{k-1}^{n+1} = a_{k-1} \xi^4 + b_{k-1} \xi^3 + c_{k-1} \xi^2 + g_{k-2} \xi + f_{k-2}, \tag{12.203}$$

where  $\xi = z_h - z_{h,k-1} - w\Delta\tau = \Delta z_{k-1}(1 - w\Delta\tau/\Delta z_{k-1})$ . We will need in a moment also the derivative

$$g_{k-1}^{n+1} = 4a_{k-1}\xi^3 + 3b_{k-1}\xi^2 + 2c_{k-1}\xi + g_{k-2}, \tag{12.204}$$

and the flux

$$F = \int_0^{\Delta\tau} f w d\tau = \int_{\xi}^{\Delta z_{k-1}} f_{k-1}^*(z) dz = \frac{1}{5} a_{k-1} (\Delta z_{k-1}^5 - \xi^5) + \frac{1}{4} b_{k-1} (\Delta z_{k-1}^4 - \xi^4) + \frac{1}{3} c_{k-1} (\Delta z_{k-1}^3 - \xi^3) + \frac{1}{2} g_{k-2} (\Delta z_{k-1}^2 - \xi^2) + f_{k-2} (\Delta z_{k-1} - \xi). \tag{12.205}$$

For  $w < 0$  we have

$$f_{k-1}^{n+1} = a\xi^4 + b\xi^3 + c\xi^2 + g_{k-1}\xi + f_{k-1}. \tag{12.206}$$

where  $\xi = w\Delta\tau$ . Again the derivative is

$$g_{k-1}^{n+1} = 4a\xi^3 + 3b\xi^2 + 3c\xi^2 + g_{k-1}, \tag{12.207}$$

and the flux is

$$F = \int_0^{\Delta\tau} f w d\tau = \int_{\xi}^0 f(z) dz = -\left( \frac{1}{5} a\xi^5 + \frac{1}{4} b\xi^4 + \frac{1}{3} c\xi^3 + \frac{1}{2} g_{k-1}\xi^2 + f_{k-1}\xi \right). \tag{12.208}$$

For solving the non homogeneous convection equation

$$\frac{\partial f}{\partial \tau} + \frac{\partial}{\partial z}(wf) = 0 \tag{12.209}$$

let us apply the splitting procedure as applied by *Xiao, Ikebata and Hasegawa* (2004). The solution of the homogeneous convection  $\frac{\partial f}{\partial \tau} + w\frac{\partial f}{\partial z} = 0$  is done by the pseudo-characteristic method as already discussed. It gives  $f^{*n+1}$ . Then the second step of the splitting provides  $f^{*n+1}$  as follows:

$$f^{n+1} = f^{*n+1} - f^{*n+1} \int_{\tau}^{\tau+\Delta\tau} \frac{\partial w}{\partial z} d\tau'. \tag{12.210}$$

*Xiao, Ikebata and Hasegawa* (2004) used the donor-cell concept for computing the integral term as follows

$$\int_{\tau}^{\tau+\Delta\tau} \frac{\partial w}{\partial z} d\tau' = \frac{\Delta\tau}{2} \frac{1}{\Delta z} (w - w_{k-1} + w^{*n+1} - w_{k-1}^{*n+1}) \quad \text{for } w > 0, \tag{12.211}$$

$$\int_{\tau}^{\tau+\Delta\tau} \frac{\partial w}{\partial z} d\tau' = \frac{\Delta\tau}{\Delta z_{k+1}} \frac{1}{2} (w_{k+1} - w + w_{k+1}^{*n+1} - w^{*n+1}) \quad \text{for } w \leq 0. \quad (12.212)$$

**12.15.3.2 Computing the weighted averages in the new time plane**

The non homogeneous convection defined by

$$\frac{\partial f}{\partial \tau} + \frac{\partial}{\partial z}(wf) = 0 \quad (12.213)$$

can be rewritten in integral form for  $w > 0$

$$\int_0^{\Delta z} f^{*n+1} dz - \int_0^{\Delta z} f^* dz = \frac{1}{\Delta z} \left( \int_{\tau}^{\tau+\Delta\tau} f_{k-1}^* w_{k-1} d\tau' - \int_{\tau}^{\tau+\Delta\tau} f^* w d\tau' \right). \quad (12.214)$$

Substituting  $w d\tau = dz$  in the left hand site integrals and setting as a lower and upper boundaries the departure and the end point of the corresponding characteristic lines at the both ends of the cell we obtain

$$\int_0^{\Delta z} f^{*n+1} dz - \int_0^{\Delta z} f^* dz = \frac{1}{\Delta z} \left( \int_{\Delta z_{k-1} - w_{k-1} \Delta \tau}^{\Delta z_{k-1}} f_{k-1}^* w_{k-1} d\tau' - \int_{\Delta z - w \Delta \tau}^{\Delta z} f^* w d\tau' \right). \quad (12.215a)$$

The general notation of this result is

$$f_{av}^{n+1} = \frac{1}{\Delta z} (F_{k-1} + f_{av} \Delta z - F) = f_{av} + \frac{1}{\Delta z} (F_{k-1} - F). \quad (12.215b)$$

where the  $F$ 's are the fluxes already defined by Eqs. (12.205) and (12.208): Remember that these fluxes are depending on the velocity direction. With other words, the profile in the new time plane between  $z_{h,k-1}$  and  $z_h$  is the transported profile from the old time plane within the boundaries  $z_{h,k-1} - w_{k-1} \Delta \tau$  and  $z_h - w \Delta \tau$ .

**12.15.3.3 Choice of the gradients**

In the ordinary cubic-spline interpolation, continuity of the value, the first derivative, and the second derivative is required to generate the piecewise cubic polynomials from the data given at some discrete points. *Yabe and Takewaki* (1986) showed that this procedure is not suitable for the present problem because the profile generated by the classical spline method is not consistent with the physical nature of the governing equations.

The CIP method relaxes the construction of the spline based on gradients computed with the values at the discrete points. Instead, it requires gra-

dients physically based on the governing equations which is the main source of success of this method.

To estimate the gradient the authors differentiate the model equation with respect to the spatial coordinate,

$$\frac{\partial g}{\partial \tau} + w \frac{\partial g}{\partial z} = \frac{dg}{d\tau} = 0. \quad (12.216)$$

Transforming the equation in the orthogonal coordinate system attached with the one coordinate to the characteristic line defined by  $dz/d\tau = w$ , gives

$$\frac{dg}{d\tau} = 0. \quad (12.217)$$

For the solution of this particular case we have Eqs. (12.204) and (12.207).

#### 12.15.3.4 Phase discontinuity treated with CIP

For propagation of a volumetric fraction described by the equation

$$\frac{\partial \alpha}{\partial \tau} + w \frac{\partial \alpha}{\partial z} = 0, \quad (12.218)$$

which may be monotonic or may possess sharp discontinuities, and at the same time has always to have values between 0 and 1, *Yabe and Xiao (1993)*, *Yabe, Xiao and Mochizuki (1995)*, proposed to use the following transformation

$$\Phi = \tan \left[ (1-\varepsilon)\pi(\alpha - 0.5) \right], \quad (12.219)$$

$$\alpha = 0.5 + \frac{\arctan \Phi}{(1-\varepsilon)\pi}, \quad (12.220)$$

where  $\varepsilon$  is a small positive constant. The factor  $(1-\varepsilon)$  enables one to get for  $\Phi \rightarrow -\infty$   $\alpha \rightarrow 0$  and  $\Phi \rightarrow \infty$   $\alpha \rightarrow 1$ , and to tune the sharpness of the layer to be resolved. Differentiating and replacing in the original equation results in

$$\frac{1}{(1-\varepsilon)\pi(1+\Phi^2)} \left( \frac{\partial \Phi}{\partial \tau} + w \frac{\partial \Phi}{\partial z} \right) = 0, \quad (12.221)$$

or

$$\frac{\partial \Phi}{\partial \tau} + w \frac{\partial \Phi}{\partial z} = 0, \quad (12.222)$$

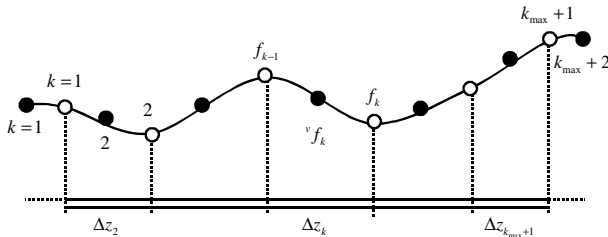
which again is solved by the CIP method. The merit of this transformation is that, although  $\Phi$  may be slightly diffusive and may have undulation when the discretized equation is solved, the inversely transformed value of  $\Phi$  is always limited to

a range between 0 and 1. The method is appropriate even for the description of the boundary of a rigid body moving in space. The method is not recommended in combination with a low order solution method like upwind etc. For more discussion see *Tanaka, Nakamura and Yabe (2000) p. 567.*

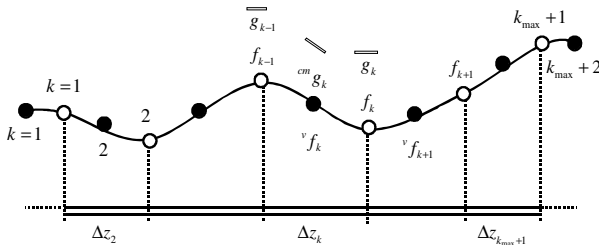
## 12.16 Problem solution examples to the basics of the CIP method

### 12.16.1 Discretization concept

Before starting with exercises on the above introduced group of CIP methods I will define the double grid discretization concept on a single pipe as shown in Fig. 12.8. The pipe is discretized on  $k_{\max}$  peaces with a length  $\Delta z_k$ ,  $k=1, k_{\max}$ . There are  $f_k$ ,  $k=1, k_{\max}+1$  function values defined at the cell boundaries and  ${}^v f_k$ ,  $k=1, k_{\max}+2$  cell averages. Note that the cell averages are not equivalent with the cell value at the cell middle. The cell averages  ${}^v f_1$  and  ${}^v f_{k_{\max}+2}$  are auxiliary values introduced to organize coupling with other pipes.



**Fig. 12.8** One dimensional pipe discretization concept on double grid. The empty symbols are boundary values. The filled symbols are spatially averaged values within the cell. The filled symbols  $k=1$  and  $k_{\max}+2$  are auxiliary averaged values defined to couple this pipe with other regions



**Fig. 12.9** One dimensional pipe discretization concept for double grid. Position of the spatial derivatives:  $g_k$  at the cell boundaries and  ${}^v g_k$  at the cell middle

In addition we define the inclinations  $g_k$  of the function  $f$  at the cell boundaries as given in Fig. 12.9 for the cell boundaries and  ${}^{cm}g_k$  for the middle of the cell. All of these values or parts of them are available from the integration process depending on its accuracy. For convenience we will notate only that subscript that differs from  $k$ .

### 12.16.2 Second order constrained interpolation profiles

**Problem 1.** Given are the fluid properties  $f$  and  $f_{k+1}$  at two different points being at distance  $\Delta z_{k+1}$  from each other. Position auxiliary coordinate system at the boundary point  $k$  so that

$$z' = 0, \Phi = f, \tag{12.223}$$

$$z' = \Delta z_{k+1}, \Phi = f_{k+1}. \tag{12.224}$$

Find a polynomial of second degree that satisfy the condition

$$\frac{1}{\Delta z_{k+1}} \int_0^{\Delta z_{k+1}} \Phi(z') dz' = {}^v f_{k+1}. \tag{12.225}$$

**Solution to problem 1:** The function, its spatial derivative, and its undetermined integral we are looking for have the form

$$\Phi = cz'^2 + gz' + f, \tag{12.226}$$

$$\frac{d\Phi}{dz} = 2cz' + g, \tag{12.227}$$

$$\int \Phi(z) dz = \frac{1}{3} cz'^3 + \frac{1}{2} gz'^2 + fz'. \tag{12.228}$$

Equations (12.226) satisfy already the conditions (12.223). These equations together with condition (12.224) and condition (12.225) applied to Eq. (12.225) give the following system of algebraic equations

$$\Delta z_{k+1}^2 c + \Delta z_{k+1} g = f_{k+1} - f, \tag{12.229}$$

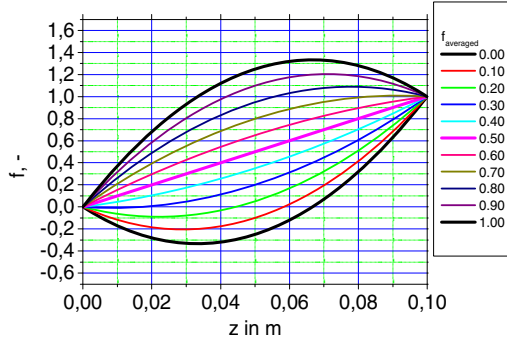
$$\frac{1}{3} \Delta z_{k+1}^2 c + g \frac{1}{2} \Delta z_{k+1} + f = {}^v f_{k+1}. \tag{12.230}$$

that is solved with respect to  $c$  and  $g$ . The result is

$$g^R = \frac{2}{\Delta z_{k+1}} (3 {}^v f_{k+1} - 2f - f_{k+1}), \tag{12.231}$$

$$c^R = \frac{6}{\Delta z_{k+1}^2} \left[ \frac{1}{2}(f + f_{k+1}) - {}^v f_{k+1} \right]. \tag{12.232}$$

Note that in our absolute coordinate system the derivative at the point  $z = z_k$  is  $g = g^R$  because in both coordinate systems the positive axis point to the same direction.



**Fig. 12.10** Quadratic constrained interpolation for  $z' = 0$ ,  $f = 0$ ,  $z' = 0.1$ ,  $f_{k+1} = 1$ ,  ${}^v f_{k+1} = 0, 0.1, 0.2, 0.3, 0.4, 0.5, 0.6, 0.7, 0.8, 0.9, 1$

**Problem 2:** Given are the values and the derivatives of a function  $f(z)$  at two points having a distance  $\Delta z = 0.1m$  from each other:

$$\text{BC 1: } z = 0, \Phi = 1; \tag{12.233}$$

$$\text{BC 2: } z = \Delta z_h, \Phi = 0. \tag{12.234}$$

Compute the profile between these points by using the results of problem 1 for

$$f_{av} = 0, 0.1, 0.2, 0.3, 0.4, 0.5, 0.6, 0.7, 0.8, 0.9, 1, \tag{12.235}$$

and plot them.

**Solution to problem 2:** The computed values are plotted in Fig. 12.10.

### 12.16.3 Third order constrained interpolation profiles

**Problem 3.** Given are the fluid properties  $f$  and  $f_{k+1}$  at two different points being at distance  $\Delta z_{k+1}$  from each other. Position auxiliary coordinate system at the boundary point  $k$  so that

$$z' = 0, \Phi = f, \tag{12.236}$$

$$z' = \Delta z_{k+1}, \quad \Phi = f_{k+1}. \tag{12.237}$$

Find a polynomial of third degree that satisfy the conditions

$$\frac{1}{\Delta z_{k+1}} \int_0^{\Delta z_{k+1}} \Phi(z') dz' = {}^v f_{k+1}, \tag{12.238}$$

$$z' = \frac{1}{2} \Delta z_{k+1}, \quad \frac{d\Phi}{dz} = {}^{cm} g_{k+1}. \tag{12.239}$$

**Solution to problem 3:** The function, its spatial derivative, and its undetermined integral we are looking for have the form

$$\Phi = bz'^3 + cz'^2 + gz' + f, \tag{12.240}$$

$$\frac{d\Phi}{dz} = 3bz'^2 + 2cz' + g, \tag{12.241}$$

$$\int \Phi(z') dz'' = \frac{1}{4}bz'^4 + \frac{1}{3}cz'^3 + \frac{1}{2}gz'^2 + fz'. \tag{12.242}$$

Equation (12.240) already satisfies the conditions (12.236). These equations together with conditions (12.237) and (12.238) give the following system of algebraic equations

$$\Delta z_{k+1}^3 b + \Delta z_{k+1}^2 c + \Delta z_{k+1} g = f_{k+1} - f, \tag{12.243}$$

$$\frac{3}{4} \Delta z_{k+1}^2 b + \Delta z_{k+1} c + g = {}^{cm} g_{k+1}, \tag{12.244}$$

$$\frac{1}{4} \Delta z_{k+1}^3 b + \frac{1}{3} \Delta z_{k+1}^2 c + \frac{1}{2} \Delta z_{k+1} g = {}^v f_{k+1} - f. \tag{12.245}$$

Solving with respect the unknown coefficients we have

$$b = \frac{4}{\Delta z_{k+1}^3} (f_{k+1} - f - {}^{cm} g_{k+1} \Delta z), \tag{12.246}$$

$$c = \frac{3}{\Delta z_{k+1}^2} (-2 {}^v f_{k+1} - f_{k+1} + 5f + 2 {}^{cm} g_{k+1} \Delta z_{k+1}), \tag{12.247}$$

$$g = \frac{2}{\Delta z_{k+1}} [-\Delta z_{k+1} {}^{cm} g_{k+1} + 3({}^v f_{k+1} - f)]. \tag{12.248}$$



### 12.16.4 Fourth order constrained interpolation profiles

**Problem 4.** Given are the fluid properties  $f_{da}$  and  $f$  and their spatial derivatives  $g_{da}$  and  $g$  at two different points being at distance  $\Delta z$  from each other:

$$z = 0, \quad \Phi = f_{da}, \quad \frac{d\Phi}{dz} = g_{da}, \quad (12.249)$$

$$z = \Delta z, \quad \Phi = f, \quad \frac{d\Phi}{dz} = g. \quad (12.250)$$

Find a polynomial of fourth degree that satisfy the condition

$$\frac{1}{\Delta z} \int_0^{\Delta z} \Phi(z) dz = f_{av}. \quad (12.251)$$

**Solution to problem 4:** The function, its spatial derivative, and its undetermined integral that we are looking for have the form

$$\Phi = az^4 + bz^3 + cz^2 + g_{da}z + f_{da}, \quad (12.252)$$

$$\frac{d\Phi}{dz} = 4az^3 + 3bz^2 + 2cz + g_{da}, \quad (12.253)$$

$$\int \Phi(z) dz = \frac{1}{5}az^5 + \frac{1}{4}bz^4 + \frac{1}{3}cz^3 + \frac{1}{2}g_{da}z^2 + f_{da}z. \quad (12.254)$$

Equations (12.252) and (12.253) already satisfy the conditions (12.249). These equations together with condition (12.251) applied to Eq. (12.254) give a system of algebraic equations

$$\Delta z_{h,up}^4 a + \Delta z_{h,up}^3 b + \Delta z_{h,up}^2 c = f_{up} - f - g_{-} \Delta z_{h,up}, \quad (12.255)$$

$$4\Delta z_{h,up}^3 a + 3\Delta z_{h,up}^2 b + 2\Delta z_{h,up} c = g_{up} - g, \quad (12.256)$$

$$\frac{1}{5}\Delta z_{h,up}^4 a + \frac{1}{4}\Delta z_{h,up}^3 b + \frac{1}{3}\Delta z_{h,up}^2 c = f_{av,up} - g \frac{1}{2}\Delta z_h - f, \quad (12.257)$$

from which the constants

$$a = \frac{30}{\Delta z^4} \left[ f_{av} - \frac{1}{2}(f + f_{da}) + \frac{1}{12}(g - g_{da})\Delta z \right], \quad (12.258)$$

$$b = \frac{4}{\Delta z^3} \left[ -15f_{av} + 7f + 8f_{da} - \left( g - \frac{3}{2}g_{da} \right) \Delta z \right], \quad (12.259)$$

$$c = \frac{3}{\Delta z^2} \left[ 10f_{av} - 4f - 6f_{da} + (g - 3g_{da}) \frac{1}{2} \Delta z \right], \quad (12.260)$$

are uniquely defined. The function (12.252) is called *monotonic*, if its values inside  $\Delta z$  satisfy the condition

$$f_{da} \leq \Phi \leq f, \quad (12.261)$$

which means that the zero-first derivative

$$\frac{d\Phi}{dz} = 4az^3 + 3bz^2 + 2cz + g_{da} = 0 \quad (12.262)$$

has its roots outside the interval

$$0 \leq z_{1,2,3} \leq \Delta z, \quad (12.263)$$

or do not have real roots. If the roots are inside the interval one have to prove whether the maximum and minimum are inside the definition region of  $\Phi$ . In this case the approximation is acceptable, else appropriate limiters of the approximation have to be introduces.

**Problem 5:** Given are the values and the derivatives of a function  $f(z)$  at two points having a distance  $\Delta z = 0.1m$  from each other:

$$\text{BC 1: } z = 0, \Phi = 1, \frac{d\Phi}{dz} = 0; \quad (12.264)$$

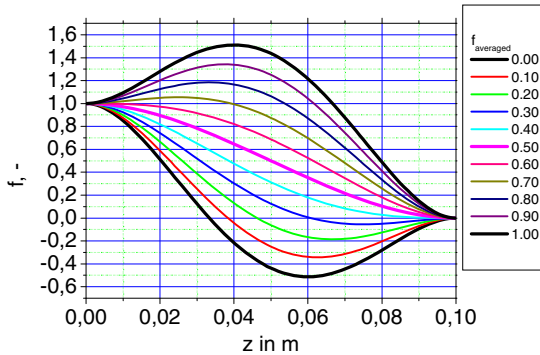
$$\text{BC 2: } z = \Delta z_h, \Phi = 0, \frac{d\Phi}{dz} = 0. \quad (12.265)$$

Compute the profile between these points by using the results of problem 1 for

$$f_{av} = 0, 0.1, 0.2, 0.3, 0.4, 0.5, 0.6, 0.7, 0.8, 0.9, 1. \quad (12.266)$$

Analyze the obtained solutions whether they are monotone or not.

**Solution to 5:** Using the set of the polynomial coefficients defined by Eqs. (12.258 to 12.260) we compute the family of curves presented in Fig. 12.11 which is the solution of this problem.



**Fig. 12.11** Constrained fourth order profile for different averaged values of the function

We see that our approximation function does not guarantee monotonic change between the minimum and the maximum values for all  $f_{av}$ .

**Problem 6:** Given are the values and the derivatives of a function  $f(z)$  at two points having a distance  $\Delta z = 0.1m$  from each other:

$$\text{BC 1: } z = 0, \Phi = 0, \frac{d\Phi}{dz} = 0; \tag{12.267}$$

$$\text{BC 2: } z = \Delta z_h, \Phi = 1, \frac{d\Phi}{dz} = 0. \tag{12.268}$$

Compute the profile between these points by using the results of problem 4 for

$$f_{av} = 0, 0.1, 0.2, 0.3, 0.4, 0.5, 0.6, 0.7, 0.8, 0.9, 1. \tag{12.269}$$

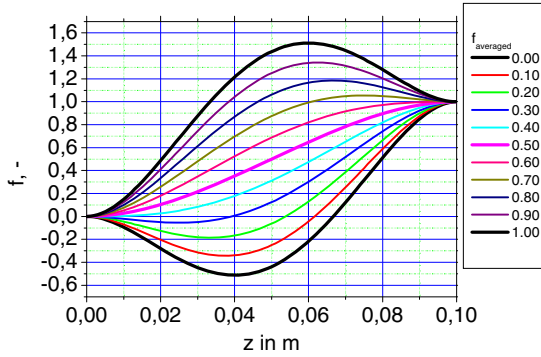
**Solution to 6:** The solution is presented at Fig. 12.12. We see again that only two of the curves are monotonic functions.

**Problem 7:** Given are the conditions for Problem 5 and 6. Construct the profile between these points by using the results of problem 4 for

$$f_{av} = 0, 0.1, 0.2, 0.3, 0.4, 0.5, 0.6, 0.7, 0.8, 0.9, 1 \tag{12.270}$$

in such a way that the function values monotonically changes between the boundary values. As already mentioned, monotonically change of the function means that all values within the definition region take values between the minimum and maximum values,

$$f_{\min} \leq \Phi \leq f_{\max}. \tag{12.271}$$



**Fig. 12.12** Constrained fourth order profile for different averaged values of the function

**Solution to 7:** We consider two cases: For case 1 the larger value is at the left boundary and for case 2 the larger value is at the right boundary.

**Case 1.** For

$$f_{da} > f \tag{12.272}$$

we have

$$f_{\max} = f_{da}, \tag{12.273}$$

$$f_{\min} = f. \tag{12.274}$$

For

$$\frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}} < \frac{1}{2}, \tag{12.275}$$

we construct a piecewise linear function that satisfy condition (12.251)

$$z < \Delta z', \quad \Phi(z) = f_{\max} + (f_{\min} - f_{\max}) \frac{z}{\Delta z'}, \tag{12.276}$$

$$z \geq \Delta z', \quad \Phi(z) = f_{\min}, \tag{12.277}$$

where

$$\Delta z' = 2\Delta z \frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}}. \tag{12.278}$$

For

$$\frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}} \geq \frac{1}{2}, \tag{12.279}$$

we construct a piecewise linear function that satisfy condition (12.251)

$$z < \Delta z', \quad \Phi(z) = f_{\max}, \tag{12.280}$$

$$z \geq \Delta z', \quad \Phi(z) = f_{\max} + (f_{\min} - f_{\max}) \frac{z - \Delta z'}{\Delta z - \Delta z'}, \tag{12.281}$$

where

$$\Delta z' = \Delta z \left( 2 \frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}} - 1 \right). \tag{12.282}$$

Now we combine this function and apply it only to those curves in Fig. 12.12 that are non monotonic. The result is presented in Fig. 12.13.

**Case 2:** For the case

$$f_{da} < f \tag{2.284}$$

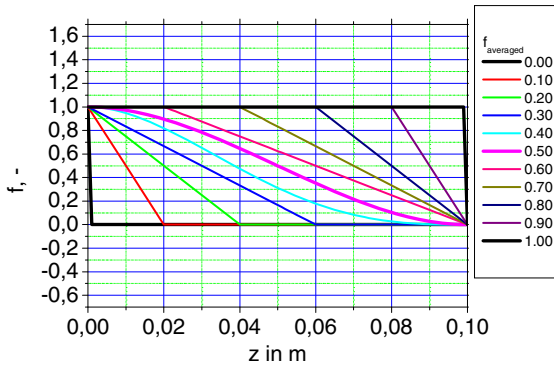
we have

$$f_{\max} = f, \tag{2.285}$$

$$f_{\min} = f_{da}. \tag{2.286}$$

For

$$\frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}} < \frac{1}{2}, \tag{12.286}$$



**Fig. 12.13** Constrained fourth order profile for different averaged values of the function with imposed step linear limiters guaranteeing the monotonic change and the integral condition (12.251)

we construct a piecewise linear function that satisfy condition (12.251)

$$z < \Delta z', \quad \Phi(z) = f_{\min}, \tag{12.287}$$

$$z \geq \Delta z', \quad \Phi(z) = f_{\min} + (f_{\max} - f_{\min}) \frac{z - \Delta z'}{\Delta z - \Delta z'}, \tag{12.288}$$

where

$$\Delta z' = \left( 1 - 2 \frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}} \right) \Delta z. \tag{12.289}$$

For the case

$$\frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}} \geq \frac{1}{2}, \tag{12.290}$$

we construct a piecewise linear function that satisfy condition (12.251)

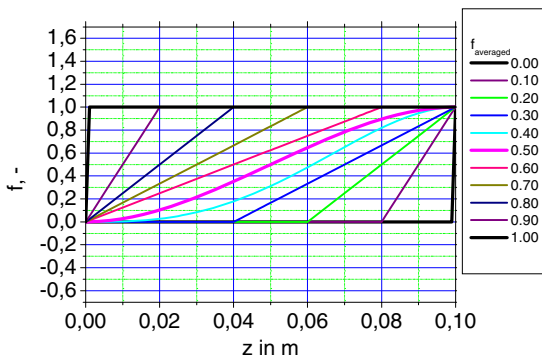
$$z < \Delta z', \quad \Phi(z) = f_{\min} + (f_{\max} - f_{\min}) \frac{z}{\Delta z}, \tag{12.291}$$

$$z \geq \Delta z', \quad \Phi(z) = f_{\max}, \tag{12.292}$$

where

$$\Delta z' = \left( 1 - \frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}} \right) 2\Delta z. \tag{12.293}$$

Now we combine this function and apply it only to those curves in Fig. 12.12 that are non monotonic. The result is presented in Fig. 12.14.



**Fig. 12.14** Constrained fourth order profile for different averaged values of the function with imposed step linear limiters guaranteeing the monotonic change and the integral condition (12.251)

The discussed limiters are summarized in Table 12.1.

**Table 12.1** Summary limiters, *Kolev (2005)*

Case	$\bar{f} = \frac{f_{av} - f_{\min}}{f_{\max} - f_{\min}}$	$\Delta z'$	$z$	$g(z)$	$f(z)$
$f_{da} > f$ $f_{\max} = f_{da}$ $f_{\min} = f$	$< \frac{1}{2}$	$2\Delta z' \bar{f}$	$< \Delta z'$	$\frac{f_{\min} - f_{\max}}{\Delta z'}$	$f_{\max} + gz$
			$\geq \Delta z'$	0	$f_{\min}$
	$\geq \frac{1}{2}$	$\Delta z'(2\bar{f} - 1)$	$< \Delta z'$	0	$f_{\max}$
			$\geq \Delta z'$	$\frac{f_{\min} - f_{\max}}{\Delta z' - \Delta z'}$	$f_{\max} + g(z - \Delta z')$
$f_{da} < f$ $f_{\max} = f$ $f_{\min} = f_{da}$	$< \frac{1}{2}$	$\Delta z'(1 - 2\bar{f})$	$< \Delta z'$	0	$f_{\min}$
			$\geq \Delta z'$	$\frac{f_{\max} - f_{\min}}{\Delta z' - \Delta z'}$	$f_{\min} + g(z - \Delta z')$
	$\geq \frac{1}{2}$	$2\Delta z'(1 - \bar{f})$	$< \Delta z'$	$\frac{f_{\max} - f_{\min}}{\Delta z'}$	$f_{\min} + gz$
			$\geq \Delta z'$	0	$f_{\max}$

**Problem 8:** Given the following partial differential equation

$$\frac{\partial f}{\partial \tau} + w \frac{\partial f}{\partial z} = 0, \tag{12.294}$$

The initial

$$f(\tau = 0, 0 \leq z \leq z_{\max}) \tag{12.295}$$

and the left boundary values

$$f(\tau \geq 0, z = 0) \tag{12.296}$$

are known. The convection velocity is a positive constant

$$w = const > 0. \tag{12.297}$$

Design *Tanaka et al. (2000)* CIP-CSL4 symbolic algorithm for solution of Eq. (12.294) over the length  $0 \leq z \leq z_{\max}$  for  $\tau = \Delta \tau$ .

**Solution to problem 5:** We divide the total length on  $k_{\max} - 1$  peaces. For simplicity but without losing the generality we will use an equidistant discretization  $\Delta z = z_{\max} / (k_{\max} - 1)$ . At each point we compute the coefficients using Eqs. (12.258 to 12.260). For the positive velocity we visit all the points  $k = 2, k_{\max}$  and compute the values and the derivatives the new time plane. In addition we compute also the fluxes.

$$w \geq 0, \quad \zeta = \Delta z_{k-1} - w\Delta\tau = \Delta z_{k-1} \left( 1 - \frac{w\Delta\tau}{\Delta z_{k-1}} \right), \quad k_{da} = k - 1, \quad (12.298)$$

$$\Phi^{n+1} = \Phi_{kda}(\zeta) = a_{kda}\zeta^4 + b_{kda}\zeta^3 + c_{kda}\zeta^2 + g_{kda}\zeta + f_{kda}, \quad (12.299)$$

$$g^{n+1} = 4a_{kda}\zeta^3 + 3b_{kda}\zeta^2 + 2c_{kda}\zeta + g_{kda}, \quad (12.300)$$

$$F = \int_0^{\Delta\tau} \Phi w d\tau = \int_{\zeta}^{\Delta z_{h,kda}} \Phi_{kda}(z) dz = \frac{1}{5} a_{kda} (\Delta z_{kda}^5 - \zeta^5) + \frac{1}{4} b_{kda} (\Delta z_{kda}^4 - \zeta^4) + \frac{1}{3} c_{kda} (\Delta z_{kda}^3 - \zeta^3) + \frac{1}{2} g_{kda} (\Delta z_{kda}^2 - \zeta^2) + f_{kda} (\Delta z_{kda} - \zeta). \quad (12.301)$$

The second set of equations is valid for negative velocity:

$$w < 0, \quad \zeta = -w\Delta\tau, \quad k_{da} = k, \quad (12.302)$$

$$\Phi^{n+1} = \Phi_{kda}(\zeta) = a_{kda}\zeta^4 + b_{kda}\zeta^3 + c_{kda}\zeta^2 + g_{kda}\zeta + f_{kda}, \quad (12.303)$$

$$g^{n+1} = 4a_{kda}\zeta^3 + 3b_{kda}\zeta^2 + 2c_{kda}\zeta + g_{kda}, \quad (12.304)$$

$$F = \int_0^{\Delta\tau} \Phi w d\tau = \int_{\zeta}^0 \Phi(z) dz = - \left( \frac{1}{5} a_{kda} \zeta^5 + \frac{1}{4} b_{kda} \zeta^4 + \frac{1}{3} c_{kda} \zeta^3 + \frac{1}{2} g_{kda} \zeta^2 + f_{kda} \zeta \right). \quad (12.305)$$

Note that the time step is so selected that the origin of the characteristic line is always within  $\Delta z$ . This is the so called *Courant-Friedrihs-Levi* condition

$$Cu = \frac{|w|\Delta\tau}{\Delta z} < 1. \quad (12.306)$$

Of course the pseudo-characteristic method allows violating this criterion for constant velocity. On have simply to contain appropriate procedure for searching the cell where the characteristic line originates.

For constant velocity the non homogeneous convection defined by Eq. (12.294) can be rewritten as

$$\frac{\partial f}{\partial \tau} + \frac{\partial}{\partial z}(wf) = 0. \quad (12.307)$$



We rewrite this equation in integral form for  $w > 0$ . The result is

$$\int_0^{\Delta z} f^{*n+1} dz - \int_0^{\Delta z} f^* dz = \frac{1}{\Delta z} \left( \int_{\tau}^{\tau+\Delta\tau} f_{k-1}^* w_{k-1} d\tau' - \int_{\tau}^{\tau+\Delta\tau} f^* w d\tau' \right). \tag{12.308}$$

Substituting  $w d\tau = dz$  in the left hand site integrals and setting as a lower and upper boundaries the departure and the end point of the corresponding characteristic lines at the both ends of the cell we obtain

$$\int_0^{\Delta z} f^{*n+1} dz - \int_0^{\Delta z} f^* dz = \frac{1}{\Delta z} \left( \int_{\Delta z_{k-1} - w_{k-1} \Delta\tau}^{\Delta z_{k-1}} f_{k-1}^* w_{k-1} d\tau' - \int_{\Delta z - w \Delta\tau}^{\Delta z} f^* w d\tau' \right). \tag{12.309}$$

The general notation of this result is

$$f_{av}^{n+1} = \frac{1}{\Delta z} (F_{k-1} + f_{av} \Delta z - F) = f_{av} + \frac{1}{\Delta z} (F_{k-1} - F). \tag{12.310}$$

where the  $F$ 's are the fluxes already defined with Eqs. (12.301) and (12.305) depending on the velocity direction, respectively. With other words, the profile in the new time plane between  $z_{h,k-1}$  and  $z_h$  is the transported profile from the old time plane within the boundaries  $z_{h,k-1} - w_{k-1} \Delta\tau$  and  $z_h - w \Delta\tau$ . Having the new values  $f$  and  $f_{av}$  we can repeat the time step as much times as necessary.

**Problem 9:** Given the initial values for  $f$  and  $f_{av}$  defining a quadratic profile as follows

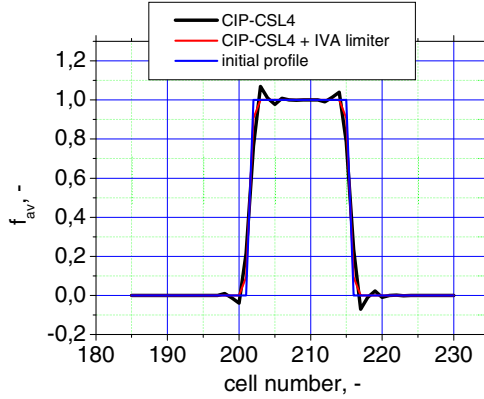
$$f = \begin{cases} 1 & \text{for } k = 2 : 16 \\ 0 & \text{elsewhere} \end{cases}, \tag{12.311}$$

$$f_{av} = \begin{cases} 1 & \text{for } k = 2 : 15 \\ 0 & \text{elsewhere} \end{cases}, \tag{12.312}$$

transported by the velocity

$$w = 1 \tag{12.313}$$

in accordance with equation (12.294). The special step is  $\Delta z = 0.01$ . The time step is  $\Delta\tau = 0.002$ . Find the profile after 1000 time steps using the *Tanaka's* et al. method. Check the conservation of the quantity  $\sum_{\text{all } k's} f_{av,k}^0 = 14$ . Compare the computed position of the quadratic wave with its expected analytical position. The analytical position is  $z = 2.01$  corresponding to cell number  $k = 201$  for  $f$  and  $k = 202$  for  $f_{av}$ .



**Fig. 12.15** Propagation of rectangular profile after 1000 time steps. Velocity  $w = 1$ ,  $\Delta z = 0.01$ ,  $\Delta \tau = 0.002$

**Solution to 9:** The solution to 9 for the averaged values designated with CIP-CSL4

is presented in Fig. 12.15. The quantity  $\sum_{all\ k's} f_{av,k}^{2000} = 14$  is perfectly conserved. The arrival position of the wave after 2s is also very well predicted. The sharp discontinuity is very well preserved. We observe around the sharp discontinuities the typical *oscillations characteristic for all high order methods applied without special measures called limiters*.

**Problem 10:** Solve the problem 9 by using the limiters from problem 7.

**Solution to 10:** The solution for  $f$  is presented in Fig. 12.16 designated with CIP method and the IVA computer code limiters derived for problem 7. We see an excellent agreement. The corresponding values of  $f_{av}$  are presented in Fig. 12.15 for comparison. This result is again better than CIP-CSL4.

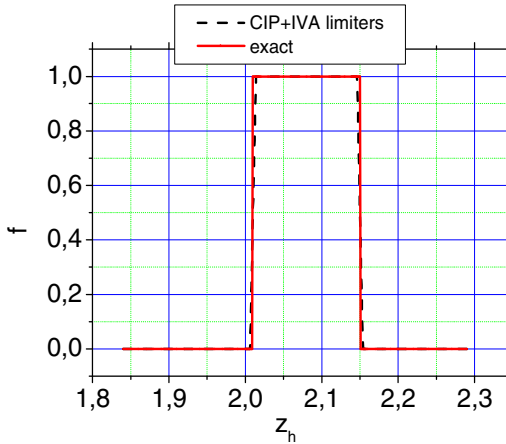
**Problem 11:** Given the model equation

$$\frac{\partial f}{\partial \tau} + \frac{\partial}{\partial z}(wf) = \mu . \tag{12.314}$$

Design a CIP-CSL4 method for solving it numerically by using time splitting and the method proposed by Xiao et al. (2004).

**Solution to 11:** We first rewrite the model equation in non conservative form,

$$\frac{\partial f}{\partial \tau} + w \frac{\partial f}{\partial z} = \mu - f \frac{\partial w}{\partial z} , \tag{12.315}$$



**Fig. 12.16** Propagation of rectangular profile after 1000 time steps. Velocity  $w = 1$ ,  $\Delta z = 0.01$ ,  $\Delta \tau = 0.002$ , Kolev (2005). Conservation error equal to zero.

and then split it into two parts: the *homogeneous advection part*

$$\frac{\partial f}{\partial \tau} + w \frac{\partial f}{\partial z} = 0, \tag{12.316}$$

and the *source part*

$$\frac{\partial f}{\partial \tau} = \mu - f \frac{\partial w}{\partial z}, \tag{12.317}$$

in such a way that their sum results in the original equation. The approximate solution for  $w > 0$  is then

$$f_{k-1}^{n+1} = f_{k-1}^{*n+1} + \int_0^{\Delta \tau} \mu_{k-1} d\tau - \frac{1}{2} (f_{k-1}^{*n+1} + f_{k-1}^n) \int_0^{\Delta \tau} \frac{\partial w}{\partial z} d\tau. \tag{12.318}$$

For  $w < 0$  we have

$$f_{k-1}^{n+1} = f_{k-1}^{*n+1} + \int_0^{\Delta \tau} \mu d\tau - \frac{1}{2} (f_{k-1}^{*n+1} + f_{k-1}^n) \int_0^{\Delta \tau} \frac{\partial w}{\partial z} d\tau, \tag{12.319}$$

The intermediate solution designated with \* are obtained for the homogeneous part as already discussed. This includes also the derivatives and the fluxes. The integrals are computed by Xiao et al. (2004), using the donor-cell concept.

$$\int_{\tau}^{\tau+\Delta \tau} \frac{\partial w}{\partial z} d\tau' = \frac{\Delta \tau}{\Delta z} \frac{1}{2} (w - w_{k-1} + w^{*n+1} - w_{k-1}^{*n+1}) \quad \text{for } w > 0, \tag{12.320}$$

$$\int_{\tau}^{\tau+\Delta\tau} \frac{\partial w}{\partial z} d\tau' = \frac{\Delta\tau}{\Delta z_{k+1}} \frac{1}{2} (w_{k+1} - w + w_{k+1}^{*n+1} - w^{*n+1}) \quad \text{for } w < 0. \quad (12.321)$$

Then the integration of original model equation is used to obtain the solutions for the averaged values

$$f_{av}^{n+1} = f_{av} + \frac{1}{\Delta z} (F_{k-1} - F) + \int_0^{\Delta\tau} \mu d\tau. \quad (12.322)$$

To obtain the conservation equation controlling the change of the derivatives we differentiate Eq. (12.314) with respect to the spatial coordinate. The result written in non conservative form is then

$$\frac{\partial g}{\partial \tau} + w \frac{\partial g}{\partial z} = \frac{\partial \mu}{\partial z} - f \frac{\partial}{\partial z} \frac{\partial w}{\partial z} - 2g \frac{\partial w}{\partial z}. \quad (12.323)$$

Again using time splitting the solution is combination of the advection part and the non advection part

$$g_{k-1}^{*n+1} = 4a_{k-1} \xi^3 + 3b_{k-1} \xi^2 + 2c_{k-1} \xi + g_{k-2}, \quad (12.324)$$

$$g_{k-1}^{n+1} = g_{k-1}^{*n+1} + \int_0^{\Delta\tau} \frac{\partial \mu_{k-1}}{\partial z} d\tau - \frac{1}{2} (f_{k-1}^{*n+1} + f_{k-1}^n) \int_0^{\Delta\tau} \frac{\partial}{\partial z} \left( \frac{\partial w}{\partial z} \right) d\tau - (g_{k-1}^{*n+1} + g_{k-1}^n) \int_0^{\Delta\tau} \frac{\partial w}{\partial z} d\tau, \quad (12.325)$$

$$\int_0^{\Delta\tau} \frac{\partial}{\partial z} \left( \frac{\partial w}{\partial z} \right) d\tau = \frac{\Delta\tau}{\Delta z + \Delta z_{k+1}} \left( \frac{w - w_{k-1}}{\Delta z} - \frac{w_{k-1} - w_{k-2}}{\Delta z_{k-1}} + \frac{w^* - w_{k-1}^*}{\Delta z} - \frac{w_{k-1}^* - w_{k-2}^*}{\Delta z_{k-1}} \right), \quad (12.326)$$

$$\int_0^{\Delta\tau} \frac{\partial \mu_{k-1}}{\partial z} d\tau = 2 \frac{\Delta\tau}{\Delta z + \Delta z_{k+1}} (\mu_{k+1} + \mu), \quad (12.327)$$

for  $w > 0$ , and

$$g_{k-1}^{*n+1} = 4a\xi^3 + 3b\xi^2 + 3c\xi^2 + g_{k-1} \quad (12.328)$$

$$g_{k-1}^{n+1} = g_{k-1}^{*n+1} + \int_0^{\Delta\tau} \frac{\partial \mu}{\partial z} d\tau - \frac{1}{2} (f_{k-1}^{*n+1} + f_{k-1}^n) \int_0^{\Delta\tau} \frac{\partial}{\partial z} \left( \frac{\partial w}{\partial z} \right) d\tau - (g_{k-1}^{*n+1} + g_{k-1}^n) \int_0^{\Delta\tau} \frac{\partial w}{\partial z} d\tau, \quad (12.329)$$

$$\int_0^{\Delta\tau} \frac{\partial \mu}{\partial z} d\tau = 2 \frac{\Delta\tau}{\Delta z + \Delta z_{k+1}} (\mu_{k+1} + \mu), \quad (12.330)$$

for  $w < 0$ . For computation of the second derivatives at the boundaries the internal neighbors have to be used as an approximation.

**Problem 12.** Given the mass conservation equation for multi-phase flow

$$\frac{\partial}{\partial \tau}(\alpha_l \rho_l \gamma_v) + \nabla \cdot (\alpha_l \rho_l \mathbf{V}_l \gamma) = \gamma_v \mu_l . \quad (12.331)$$

Derive an expression for the convective velocity of the volume fraction for a semi-conservative form. Write the corresponding conservative form also.

**Solution to 12:** Using the chain rule we differentiate the two products of the left hand site. Then we divide by the density and the volume porosity. Taking into account that

$$\frac{\partial \rho_l}{\rho_l} = \partial \ln \rho_l , \quad (12.332)$$

$$\frac{\partial \gamma_v}{\gamma_v} = \partial \ln \gamma_v , \quad (12.333)$$

substituting

$$\bar{\mathbf{V}}_l = \frac{\gamma \mathbf{V}_l}{\gamma_v} , \quad (12.334)$$

and after some rearrangements we finally obtain

$$\frac{\partial \alpha_l}{\partial \tau} + \nabla \cdot (\bar{\mathbf{V}}_l \alpha_l) = \frac{\mu_l}{\rho_l} - \alpha_l \left( \frac{\partial \ln \gamma_v \rho_l}{\partial \tau} + \bar{\mathbf{V}}_l \cdot \nabla \ln \gamma_v \rho_l \right) , \quad (12.335)$$

or

$$\frac{\partial \alpha_l}{\partial \tau} + \bar{\mathbf{V}}_l \cdot \nabla \alpha_l = \frac{\mu_l}{\rho_l} - \alpha_l \left( \frac{\partial \ln \gamma_v \rho_l}{\partial \tau} + \bar{\mathbf{V}}_l \cdot \nabla \ln \gamma_v \rho_l + \nabla \cdot \bar{\mathbf{V}}_l \right) . \quad (12.336)$$

or

$$\frac{d \alpha_l}{d \tau} = \frac{\mu_l}{\rho_l} - \alpha_l \left( \frac{d \ln \gamma_v \rho_l}{d \tau} + \nabla \cdot \bar{\mathbf{V}}_l \right) , \quad (12.337)$$

with the characteristic velocity  $\bar{\mathbf{V}}_l$ , which is the real convection velocity of the volume fraction. Designing the semi-*Lagrangian* splitting for the cell boundary volume fractions we obtain

$${}^z \alpha_l^* = {}^z \alpha_{l, \Delta z - w_l \Delta \tau}^n + \int_0^{\Delta \tau} \frac{{}^z \mu_l}{{}^z \rho_l} d\tau' \quad \text{for } {}^z w_l > 0 , \quad (12.338)$$

$${}^z \alpha_l^* = {}^z \alpha_{l, -w_l \Delta \tau}^n + \int_0^{\Delta \tau} \frac{{}^z \mu_l}{{}^z \rho_l} d\tau' \quad \text{for } {}^z w_l < 0 , \quad (12.339)$$

$${}^z\alpha_l^{n+1} = {}^z\alpha_l^* - \frac{1}{2}({}^z\alpha_l + {}^z\alpha_l^*) \int_0^{\Delta\tau} \left( \frac{d \ln {}^z\gamma_v {}^z\rho_l}{d\tau} + \nabla \cdot {}^z\bar{\mathbf{V}}_l \right) d\tau, \quad (12.340)$$

$$\begin{aligned} \alpha_l^{n+1} &= \alpha_l + \frac{1}{\Delta z} (F_{\alpha,k-1} - F_\alpha) + \int_0^{\Delta\tau} \frac{\mu_l}{\rho_l} d\tau' \\ &- \frac{1}{2}(\alpha_l^{n+1} + \alpha_l) \int_0^{\Delta\tau} \left( \frac{\partial \ln \gamma_v \rho_l}{\partial \tau} + \bar{\mathbf{V}}_l \cdot \nabla \ln \gamma_v \rho_l \right) d\tau \end{aligned} \quad (12.341)$$

The same procedure can be applied to obtain conservative and non conservative equations with respect to  $\alpha_l \rho_l$ . The result is

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l) + \nabla \cdot (\alpha_l \rho_l \bar{\mathbf{V}}_l) = \mu_l - \alpha_l \rho_l \left[ \frac{\partial}{\partial \tau} (\ln \gamma_v) + \bar{\mathbf{V}}_l \cdot \nabla \cdot (\ln \gamma_v) \right], \quad (12.342)$$

$$\frac{d}{d\tau} (\alpha_l \rho_l) = \mu_l - \alpha_l \rho_l \left[ \frac{d}{d\tau} (\ln \gamma_v) + \nabla \cdot \bar{\mathbf{V}}_l \right], \quad (12.343)$$

where the total derivatives of the last equation are along the curve defined by  $dz/d\tau = \bar{\mathbf{V}}_l$ .

**Problem 13:** Design a concept for pipe flow for definition of the surface permeabilities and the volume porosities.

**Solution to 13:** One of the possible solutions is the following. The axis of the pipe is discretized on  $k_m$  non equidistant segments with length  $\Delta z_k$ . At the centre of each segment we define a volumetric porosity  $\gamma_{cv,k}$ . At the beginning and the end of each segment  $k$  we define surface permeabilities  $\gamma_{z,k}$  and  $\gamma_{z,k+1}$ . For convenience we will omit the subscript notation except it is different from  $k$ . We require also an additional set of volumetric porosities positioned at the cell boundaries. One thinkable approach is

$$\gamma_v = \frac{\gamma_{vc,k-1} \Delta z_{k-1} + \gamma_{vc} \Delta z}{\Delta z_{k-1} + \Delta z}, \quad (12.344)$$

which is exactly the volume porosity of a staggered cell formed between two centres. There are solution methods that require the surface permeability at the centre of the cell. In this case a good approximation may be

$$\gamma_{zc} = \gamma_{vc}. \quad (12.345)$$

For the boundaries we have  $\gamma_{v,1} = \gamma_{vc,1}$ ,  $\gamma_{v,k_m+1} = \gamma_{vc,k_m}$ .

**Problem 14.** Given the momentum conservation equation for multiphase flow

$$\frac{\partial}{\partial \tau}(\alpha_i \rho_i w_i \gamma_v) + \frac{\partial}{\partial z}(\alpha_i \rho_i w_i w_i \gamma_z) + \alpha_i \gamma_z \frac{\partial p}{\partial z} = \gamma_v f_z. \quad (12.346)$$

Write the form appropriate to use semi-*Lagrangian* methods with respect to the volume flux.

**Solution to 14:** We introduce the velocity  $\bar{V}$  and rearrange to obtain the conservative form with respect to the field specific momentum

$$\frac{\partial}{\partial \tau}(\alpha_i \rho_i w_i) + \frac{\partial}{\partial z}(\alpha_i \rho_i w_i \bar{V}_i) + \alpha_i \frac{\gamma_z}{\gamma_v} \frac{\partial p}{\partial z} = f_z - \alpha_i \rho_i w_i \left( \frac{\partial \ln \gamma_v}{\partial \tau} + \bar{V}_i \frac{\partial}{\partial z} \ln \gamma_v \right). \quad (12.347)$$

The corresponding non conservative form is

$$\begin{aligned} & \frac{\partial}{\partial \tau}(\alpha_i \rho_i w_i) + \bar{V}_i \frac{\partial}{\partial z}(\alpha_i \rho_i w_i) + \alpha_i \frac{\gamma_z}{\gamma_v} \frac{\partial p}{\partial z} \\ &= f_z - \alpha_i \rho_i w_i \left( \frac{\partial \ln \gamma_v}{\partial \tau} + \bar{V}_i \frac{\partial}{\partial z} \ln \gamma_v + \frac{\partial \bar{V}_i}{\partial z} \right). \end{aligned} \quad (12.348)$$

The pseudo-canonical form along the characteristic line defined by  $dz/d\tau = \bar{V}$  is then

$$\frac{d}{d\tau}(\alpha_i \rho_i w_i) + \alpha_i \frac{\gamma_z}{\gamma_v} \frac{\partial p}{\partial z} = f_z - \alpha_i \rho_i w_i \left( \frac{d \ln \gamma_v}{d\tau} + \frac{\partial \bar{V}_i}{\partial z} \right). \quad (12.349)$$

The pseudo-canonical form with respect to the volume flux is

$$\frac{d}{d\tau}(\alpha_i w_i) + \frac{\alpha_i \gamma_z}{\rho_i \gamma_v} \frac{\partial p}{\partial z} = \frac{f_z}{\rho_i} - \alpha_i w_i \left( \frac{d \ln \gamma_v \rho_i}{d\tau} + \frac{\partial \bar{V}_i}{\partial z} \right). \quad (12.350)$$

The last two differential expressions can be replaced using Eq. (12.337) resulting in

$$\frac{d}{d\tau}(\alpha_i w_i) + \frac{\alpha_i \gamma_z}{\rho_i \gamma_v} \frac{\partial p}{\partial z} = \frac{f_z}{\rho_i} + \left( \frac{d\alpha_i}{d\tau} - \frac{\mu_i}{\rho_i} \right) w_i. \quad (12.351)$$

Using semi-*Lagrangian* operator splitting method we can compute the intermediate cell boundary values as follows

$$(\alpha_i w_i)^* = (\alpha_i w_i)_{\Delta z_{i-1} - V_i \Delta \tau} + \int_0^{\Delta \tau} \left[ \frac{f_z}{\rho_i} + \left( \frac{d\alpha_i}{d\tau} - \frac{\mu_i}{\rho_i} \right) w_i \right] d\tau', \text{ for } w_i > 0, \quad (12.352)$$

and

$$(\alpha_l w_l)^* = (\alpha_l w_l)_{-V_l \Delta \tau} + \int_0^{\Delta \tau} \left[ \frac{f_z}{\rho_l} + \left( \frac{d\alpha_l}{d\tau} - \frac{\mu_l}{\rho_l} \right) w_l \right] d\tau', \text{ for } w_l < 0. \quad (12.353)$$

In a second step we have for both boundaries

$$(\alpha_l w_l) = (\alpha_l w_l)^* - \frac{\alpha_l \gamma_z}{\rho_l \gamma_v} \frac{\partial p}{\partial z} \Delta \tau, \quad (12.354)$$

$$(\alpha_l w_l)_{k-1} = (\alpha_l w_l)_{k-1}^* - \left( \frac{\alpha_l}{\rho_l} \right)_{k-1} \left( \frac{\gamma_z}{\gamma_v} \right)_{k-1} \left( \frac{\partial p}{\partial z} \right)_{k-1} \Delta \tau. \quad (12.355)$$

**Problem 15.** Given are  $l_{\max}$  mass conservation equations for multiphase flow

$$\frac{\partial}{\partial \tau} (\alpha_l \rho_l \gamma_v) + \nabla \cdot (\alpha_l \rho_l \mathbf{V}_l \gamma) = \gamma_v \mu_l. \quad (12.356)$$

Combine the equations in such a way to obtain one equation in which the time derivative of the volume fraction is eliminated. Use the differential form of the state equation  $\rho_l = \rho_l(p_l, s_l, C_{il})$  to obtain *Poison*-like equation for multiphase flow.

**Solution to 15:** We first differentiate using the chain rule to obtain

$$\gamma_v \frac{\partial \alpha_l}{\partial \tau} + \alpha_l \frac{\partial \gamma_v}{\partial \tau} + \frac{1}{\rho_l} \left[ \alpha_l \gamma_v \frac{\partial \rho_l}{\partial \tau} + (\alpha_l \mathbf{V}_l \gamma \cdot \nabla) \rho_l \right] + \nabla \cdot (\alpha_l \mathbf{V}_l \gamma) = \gamma_v \frac{\mu_l}{\rho_l}. \quad (12.357)$$

Then using the differential form of the state equation we replace all density derivatives with their equals. The terms not containing pressure derivatives are putted in the right hand site. The result is

$$\gamma_v \frac{\partial \alpha_l}{\partial \tau} + \alpha_l \frac{\partial \gamma_v}{\partial \tau} + \frac{\alpha_l}{\rho_l a_l^2} \left[ \gamma_v \frac{\partial p}{\partial \tau} + (\mathbf{V}_l \gamma \cdot \nabla) p \right] + \nabla \cdot (\alpha_l \mathbf{V}_l \gamma) = D\alpha_l. \quad (12.358)$$

Dividing by the volumetric porosity, adding all the mass conservation equations, and having in mind that  $\sum_{l=1}^{l_{\max}} \alpha_l = 1$ , and therefore  $\sum_{l=1}^{l_{\max}} d\alpha_l = 0$  we obtain the so called volume conservation equation

$$\frac{\gamma_v}{\rho a^2} \frac{\partial p}{\partial \tau} + \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} (\mathbf{V}_l \gamma \cdot \nabla) p + \nabla \cdot \sum_{l=1}^{l_{\max}} (\alpha_l \mathbf{V}_l \gamma) = \sum_{l=1}^{l_{\max}} D\alpha_l - \frac{\partial \gamma_v}{\partial \tau}. \quad (12.359)$$

Multiplying both sites by  $\rho a^2 / \gamma_v$  we finally obtain

$$\frac{\partial p}{\partial \tau} + \mathbf{V}_p \cdot \nabla p + \rho a^2 \nabla \cdot \left( \sum_{l=1}^{l_{\max}} \alpha_l \bar{\mathbf{V}}_l \right) = \rho a^2 \left( \frac{1}{\gamma_v} \sum_{l=1}^{l_{\max}} D\alpha_l - \frac{\partial \ln \gamma_v}{\partial \tau} \right). \quad (12.360)$$



We may call

$$\mathbf{V}_p = \rho a^2 \left( \sum_{l=1}^{l_{\max}} \frac{\alpha_l}{\rho_l a_l^2} \frac{\mathbf{V}_l \gamma}{\gamma_v} \right), \quad (12.361)$$

pressure pseudo-convection velocity.

**Problem 16:** Given the volume conservation equation (12.361). Find a way to couple its discretized form with the discretized form of the momentum equations.

**Solution to 16:** If we define pressure  ${}^v p$  at the centre of each cell, we may consider as possible discretization of the volume flux

$$\frac{1}{\rho a^2} \left( \frac{\partial {}^v p}{\partial \tau} + \mathbf{V}_p \cdot \nabla {}^v p \right) + \frac{1}{\Delta z} \sum_{l=1}^{l_{\max}} \left( \alpha_l w_l \frac{\gamma_z}{\gamma_v} - \left( \alpha_l w_l \frac{\gamma_z}{\gamma_v} \right)_{k-1} \right) = \frac{1}{\gamma_v} \sum_{l=1}^{l_{\max}} D \alpha_l - \frac{\partial \ln \gamma_v}{\partial \tau}. \quad (12.362)$$

Using Eqs. (12.354) and (12.355) we express the volume fluxes at the new time plane with their equals. The result is

$$\begin{aligned} & \frac{1}{\rho a^2} \left( \frac{\partial {}^v p}{\partial \tau} + \mathbf{V}_p \cdot \nabla {}^v p \right) - \frac{\Delta \tau}{\Delta z} \left( \frac{\gamma_z}{\gamma_v} \right)^2 \frac{\partial {}^z p}{\partial z} \sum_{l=1}^{l_{\max}} \left( \frac{\alpha_l}{\rho_l} \right) \\ & + \frac{\Delta \tau}{\Delta z} \left( \frac{\gamma_z}{\gamma_v} \right)_{k-1}^2 \left( \frac{\partial {}^z p}{\partial z} \right)_{k-1} \sum_{l=1}^{l_{\max}} \left( \frac{\alpha_l}{\rho_l} \right)_{k-1} \\ & = \frac{1}{\gamma_v} \sum_{l=1}^{l_{\max}} D \alpha_l - \frac{\partial \ln \gamma_v}{\partial \tau} - \frac{1}{\Delta z} \sum_{l=1}^{l_{\max}} \left[ \left( \alpha_l w_l \right)^* \frac{\gamma_z}{\gamma_v} - \left( \alpha_l w_l \right)^*_{k-1} \left( \frac{\gamma_z}{\gamma_v} \right)_{k-1} \right]. \quad (12.363) \end{aligned}$$

For the discretization of the spatial derivatives of the pressures we may use explicit, implicit or semi-implicit methods. Let us start with an explicit discretization. The result is

$$\begin{aligned} & \frac{1}{\rho a^2} \left( \frac{{}^v p^{n+1} - {}^v p}{\Delta \tau} + \mathbf{V}_p \cdot \nabla {}^v p \right) \\ & - \frac{\Delta \tau}{\Delta z} \left( \frac{\gamma_z}{\gamma_v} \right)^2 2 \frac{{}^v p_{k+1} - {}^v p}{\Delta z + \Delta z_{k+1}} \sum_{l=1}^{l_{\max}} \left( \frac{\alpha_l}{\rho_l} \right) \\ & + \frac{\Delta \tau}{\Delta z} \left( \frac{\gamma_z}{\gamma_v} \right)_{k-1}^2 2 \frac{{}^v p - {}^v p_{k-1}}{\Delta z + \Delta z_{k-1}} \sum_{l=1}^{l_{\max}} \left( \frac{\alpha_l}{\rho_l} \right)_{k-1} \end{aligned}$$

$$= \frac{1}{\gamma_v} \sum_{l=1}^{l_{\max}} D\alpha_l - \frac{1}{\Delta\tau} \ln \frac{\gamma_v^{n+1}}{\gamma_v} - \frac{1}{\Delta z} \sum_{l=1}^{l_{\max}} \left[ (\alpha_l w_l)^* \frac{\gamma_z}{\gamma_v} - (\alpha_l w_l)_{k-1}^* \left( \frac{\gamma_z}{\gamma_v} \right)_{k-1} \right]. \quad (12.364)$$

Using the IVA-donor cell formula for first order explicit discretization of the term

$$w_p \frac{\partial^v p}{\partial z},$$

$$\begin{aligned} w_p \frac{\partial^v p}{\partial z} &= \frac{\partial}{\partial z} (w_p {}^v p) - {}^v p \frac{\partial w_p}{\partial z} \\ &= -\frac{1}{\Delta z} (-w_p \xi_{p5-}) ({}^v p_{k+1} - {}^v p) - \frac{1}{\Delta z} (w_{p_{k-1}} \xi_{p6-}) ({}^v p_{k-1} - {}^v p), \end{aligned} \quad (12.365)$$

where  $\frac{1}{2} [1 - \text{sign}(w_p)] = \xi_{p5-}$  and  $\frac{1}{2} [1 + \text{sign}(w_{p_{k-1}})] = \xi_{p6-}$  we obtain

$$\begin{aligned} &\frac{1}{\rho a^2} {}^v p^{n+1} - \frac{\Delta\tau}{\Delta z} \left[ \frac{1}{\rho a^2} (-w_p \xi_{p5-}) + \frac{2\Delta\tau}{(\Delta z + \Delta z_{k+1})} \left( \frac{\gamma_z}{\gamma_v} \right)^2 \sum_{l=1}^{l_{\max}} \left( \frac{\alpha_l}{\rho_l} \right) \right] ({}^v p_{k+1} - {}^v p) \\ &- \frac{\Delta\tau}{\Delta z} \left[ \frac{1}{\rho a^2} (w_{p_{k-1}} \xi_{p6-}) + \frac{2\Delta\tau}{(\Delta z + \Delta z_{k-1})} \left( \frac{\gamma_z}{\gamma_v} \right)_{k-1}^2 \sum_{l=1}^{l_{\max}} \left( \frac{\alpha_l}{\rho_l} \right)_{k-1} \right] ({}^v p_{k-1} - {}^v p) \\ &= \frac{{}^v p}{\rho a^2} + \Delta\tau \frac{1}{\gamma_v} \sum_{l=1}^{l_{\max}} D\alpha_l - \ln \frac{\gamma_v^{n+1}}{\gamma_v} - \frac{\Delta\tau}{\Delta z} \sum_{l=1}^{l_{\max}} \left[ (\alpha_l w_l)^* \frac{\gamma_z}{\gamma_v} - (\alpha_l w_l)_{k-1}^* \left( \frac{\gamma_z}{\gamma_v} \right)_{k-1} \right]. \end{aligned} \quad (12.366)$$

Substituting in the above equation with

$$c_{k+1} = - \left[ \frac{\Delta\tau}{\Delta z} (-w_p \xi_{p5-}) + \frac{2\rho a^2}{\Delta z (\Delta z + \Delta z_{k+1})} \left( \Delta\tau \frac{\gamma_z}{\gamma_v} \right)^2 \sum_{l=1}^{l_{\max}} \left( \frac{\alpha_l}{\rho_l} \right) \right], \quad (12.367)$$

$$c_{k-1} = - \left[ \frac{\Delta\tau}{\Delta z} (w_{p_{k-1}} \xi_{p6-}) + \frac{2\rho a^2}{\Delta z (\Delta z + \Delta z_{k-1})} \left( \Delta\tau \frac{\gamma_z}{\gamma_v} \right)_{k-1}^2 \sum_{l=1}^{l_{\max}} \left( \frac{\alpha_l}{\rho_l} \right)_{k-1} \right], \quad (12.368)$$

$$D = {}^v p + \rho a^2 \left\{ \Delta\tau \frac{1}{\gamma_v} \sum_{l=1}^{l_{\max}} D\alpha_l - \ln \frac{\gamma_v^{n+1}}{\gamma_v} - \frac{\Delta\tau}{\Delta z} \sum_{l=1}^{l_{\max}} \left[ (\alpha_l w_l)^* \frac{\gamma_z}{\gamma_v} - (\alpha_l w_l)_{k-1}^* \left( \frac{\gamma_z}{\gamma_v} \right)_{k-1} \right] \right\}, \quad (12.369)$$

we obtain

$${}^v p^{n+1} + c_{k+1} ({}^v p_{k+1} - {}^v p) + c_{k-1} ({}^v p_{k-1} - {}^v p) = D. \quad (12.370)$$

For implicit discretization of the pressure gradients we have

$$c_{k-1} {}^v p_{k-1}^{n+1} + (1 - c_{k+1} - c_{k-1}) {}^v p^{n+1} + c_{k+1} {}^v p_{k+1}^{n+1} = D. \quad (12.371)$$

This is the pressure equation similar to the discretized *Poison* equation for single phase flow.

**Problem 17:** Given the system of algebraic equations

$$c_{k-} p_{k-1} + c_k p_k + c_{k+} p_{k+1} = d_k \quad (12.372)$$

for a pipe segment defined by  $k = k_{in} + 1, k_{out} - 1$  internal cells. Find for all cells the dependence of a type

$$c_{k-}^* p_{k_{in}} + c_k^* p_k + c_{k+}^* p_{k+1} = d_k^*. \quad (12.373)$$

This means that each cell pressure depend on the inlet pressure  $p_{k_{in}}$  and the right neighbor pressure  $p_{k+1}$ .

**Solution to 17:** For the first real cell we have

$$c_{k_{in}+1,-} p_{k_{in}} + c_{k_{in}+1} p_{k_{in}+1} + c_{k_{in}+1,+} p_{k_{in}+2} = d_{k_{in}+1}, \quad (12.374)$$

and therefore  $c_{k_{in}+1,-}^* = c_{k_{in}+1,-}$ ,  $c_{k_{in}+1}^* = c_{k_{in}+1}$ ,  $c_{k_{in}+1,+}^* = c_{k_{in}+1,+}$  and  $d_{k_{in}+1}^* = d_{k_{in}+1}$ . The task then is reduced to eliminating  $p_{k-1}$  from the following system

$$c_{k-1,-}^* p_{k_{in}} + c_{k-1}^* p_{k-1} + c_{k-1,+}^* p_k = d_{k-1}^* \quad (12.375)$$

$$c_{k-} p_{k-1} + c_k p_k + c_{k+} p_{k+1} = d_k \quad (12.376)$$

resulting in

$$-c_{k-} c_{k-1,-}^* p_{k_{in}} + (c_{k-1}^* c_k - c_{k-} c_{k-1,+}^*) p_k + c_{k-1}^* c_{k+} p_{k+1} = c_{k-1}^* d_k - c_{k-} d_{k-1}^*$$

with the recursion

$$c_{k-}^* = -c_{k-} c_{k-1,-}^*, \quad c_k^* = c_{k-1}^* c_k - c_{k-} c_{k-1,+}^*, \quad c_{k+}^* = c_{k-1}^* c_{k+}, \quad d_k^* = c_{k-1}^* d_k - c_{k-} d_{k-1}^*. \quad (12.377)$$

For the last real cells we have

$$c_{k_{out}-1,-}^* p_{k_{in}} + c_{k_{out}-1}^* p_{k_{out}-1} + c_{k_{out}-1,+}^* p_{k_{out}} = d_{k_{out}-1}^*. \quad (12.378)$$

Therefore for given outlet pressure next left pressure is analytically computed. This is the so called forwards elimination pass of the three diagonal solution algorithms.

**Problem 18:** Given the system of algebraic equations as in the previous case. Find for all cells the dependence of a type

$$c_{k-}^{**} p_{k-1} + c_k^{**} p_k + c_{k+}^{**} p_{k_{out}} = d_k^{**}. \quad (12.379)$$

This means that each cell pressure depend on the outlet pressure  $p_{k_{out}}$  and the left neighbor pressure  $p_{k-1}$ .

**Solution to 18:** The pressure equation for the real cell having its neighbor  $k_{out}$  gives the initial values for the recursive coefficients  $c_{k_{out}-1,-}^{**} = c_{k_{out}-1,-}^{**}$ ,  $c_{k_{out}-1}^{**} = c_{k_{out}-1}^{**}$ ,  $c_{k_{out}-1,+}^{**} = c_{k_{out}-1,+}^{**}$ ,  $d_{k_{out}-1}^{**} = d_{k_{out}-1}^{**}$ . Then eliminating  $p_{k+1}$  from the system

$$c_{k+1,-}^{**} p_k + c_{k+1}^{**} p_{k+1} + c_{k+1,+}^{**} p_{k_{out}} = d_{k+1}^{**}, \quad (12.380)$$

$$c_{k-} p_{k-1} + c_k p_k + c_{k+} p_{k+1} = d_k, \quad (12.381)$$

results in

$$c_{k+1}^{**} c_{k-} p_{k-1} + (c_{k+1}^{**} c_k - c_k c_{k+1,-}^{**}) p_k - c_k c_{k+1,+}^{**} p_{k_{out}} = c_{k+1}^{**} d_k - c_k d_{k+1}^{**}, \quad (12.382)$$

and therefore

$$c_{k-}^{**} = c_{k+1}^{**}, \quad c_k^{**} = c_{k+1}^{**} c_k - c_k c_{k+1,-}^{**}, \quad c_{k+}^{**} = -c_k c_{k+1,+}^{**}, \quad d_k^{**} = c_{k+1}^{**} d_k - c_k d_{k+1}^{**}. \quad (12.383)$$

Therefore for given outlet pressure next left pressure is analytically computed.

**Problem 19:** Given the system of algebraic equations as in the previous two cases. Find the dependence of the pressure in any cells from the inlet and outlet pressure.

**Solution to 19:** Writing Eqs. (12.373) and (12.380)

$$c_{k-1,-}^* p_{k_{in}} + c_{k-1}^* p_{k-1} + c_{k-1,+}^* p_k = d_{k-1}^*,$$

$$c_{k-}^{**} p_{k-1} + c_k^{**} p_k + c_{k+}^{**} p_{k_{out}} = d_k^{**},$$

and eliminating the  $p_{k-1}$  results in

$$p_k = \frac{c_{k-}^{**} d_{k-1}^* - c_{k-1}^* d_k^{**} - c_{k-}^{**} c_{k-1,-}^* p_{k_{in}} + c_{k-1}^* c_{k+}^{**} p_{k_{out}}}{c_{k-1,+}^* c_{k-}^{**} - c_{k-1}^* c_k^{**}}. \quad (12.384)$$

The solutions to the problems 17, 18 and 19 can be used to design variety of methods for solving algebraic systems for flows in pipe networks in which  $k$  may be the real cell selected as a knot. In any case, by iterative solutions of the algebraic

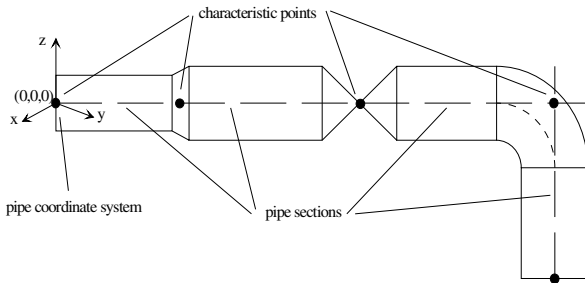
systems for many pipes, it is recommendable to rank the pipes depending of the number of knots they possess and to solve the equations for each pipe following this order. So the new information propagates faster in the iteration process.

## 12.17 Pipe networks: some basic definitions

The methods for integration of the systems of partial differential equations governing the multi phase flows in 3D presented in the previous Sections can be used also for describing the flow in pipe networks. The system takes a very simple form. Usually the  $z$ -direction components of all the equations are necessary in differential and in finite difference form with some small modifications for the change of the axis angle. In addition coupling through the so called knots has to be mathematically defined. We call the pipe network flow 1.5-dimensional flow rather than one dimensional, because of the cross connections among the pipes. It is very important before attempting to design a new computational model for flows in pipe networks to formally describe the pipe network in quantitative characteristics. This is the subject of this section.

### 12.17.1 Pipes

A *pipe* is a one-dimensional flow channel, of which the axis runs arbitrarily through three-dimensional space – see Fig. 12.17.



**Fig. 12.17** Pipe definition: coordinate system, characteristic points, pipe sections

At specified *characteristic points* the pipe may contain

- a) elbows,
- b) reductions and expansions,
- c) components such as valves, pumps etc.

or may experience sudden changes in

- a) the pipe inner diameter defining the flow cross section, or in
- b) the pipe hydraulic diameter,

- c) the pipe material, or in
- d) the wall thickness or in
- e) the roughness,

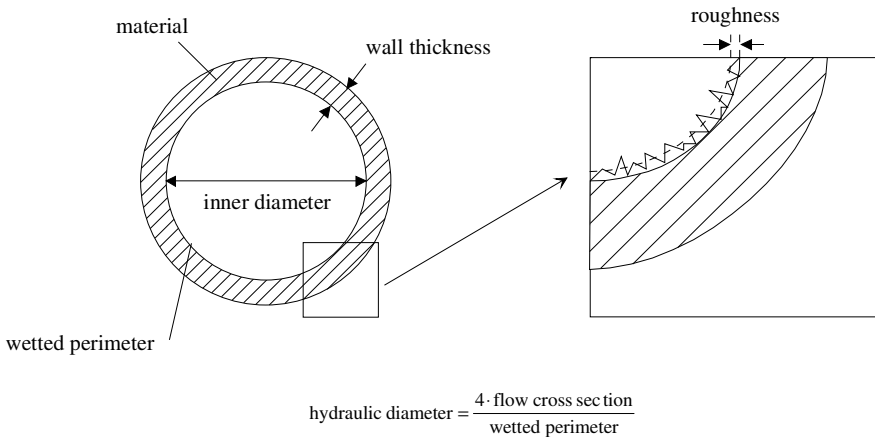
or

- d) may be interconnected with the beginning or the end of other pipes and form a knot. The coordinates of all these *characteristic points* define the axis of the pipe in the space.

The characteristic points divide the pipe into straight *sections*, where some specific pipe attributes like

- a) inner diameter,
- b) hydraulic diameter,
- c) wall thickness,
- d) the roughness and
- e) the material

of the pipe are constant – see Fig. 12.18.



**Fig. 12.18** Definitions of the pipe attributes: inner diameter defining the flow cross section, hydraulic diameter, material, wall thickness, roughness

Pipes are identified by integer numbers ranging from 1 to the total number of pipes inside the network and for convenience by a text identifier - *name*. This name has to be unique throughout the network. Sometimes it may happen that pipes have exactly the same geometry and run parallel in space, i.e. they are of the same type. It is not necessary to model each of them separately. The resulting flow is instead computed by multiplying the single pipe flow by the number of parallel pipes.

In addition, *each* pipe has its own *normalizing diameter* defining the normalizing flow area. All pipe flow cross section areas are divided by the normalizing cross section for the pipe. The results are the so called flow permeabilities defining the flow cross section. Describing the portion of the cross

section available for the flow is the technique used in the IVA code series and is applied in the three-dimensional analysis as well as in the one-dimensional network analysis.

### 12.17.2 Axis in the space

Before a pipe can be inserted into the network, its geometry has to be defined. The pipe is usually defined by an identifier name and the coordinates of its characteristic points.

$$(x_p, y_p, z_p)$$

The pipe is defined in its own rectangular left oriented coordinate system (Fig. 12.17). The coordinate system is usually attached to the beginning of the pipe. That means every pipe starts at the point  $(0,0,0)$  in its own coordinate system. The coordinates of every point of interest called characteristic points, such as elbows, components, area changes or changes in material, are defined with respect to this point (*relative coordinates*). This gives an opportunity to create libraries with standardized pipes. Defining pipes in *absolute coordinates* is also possible. In this case the coordinate system is not necessarily connected to the pipe start point.

The positive orientation of the pipe axis is defined through the order of the characteristic points from the start point to the pipe end. This direction corresponds to the increasing cell indices of the discrete control volumes created after the pipe definition for computational analysis.

Two characteristic angles are specified (see Fig. 12.19) for use internally in computer codes. The first one,

$$\theta = \arccos \left[ \frac{(\mathbf{r}_{k-1,k} \cdot \mathbf{r}_{k+1,k})}{(|\mathbf{r}_{k-1,k}| |\mathbf{r}_{k+1,k}|)} \right],$$

is called the *deviation angle*. Here  $k-1$ ,  $k$  and  $k+1$  define three sequential characteristic points and  $\mathbf{r}$ 's are the vectors between them. This is the angle of a pipe section defined as the deviation from the positive oriented axis of the previous pipe section. The second one,

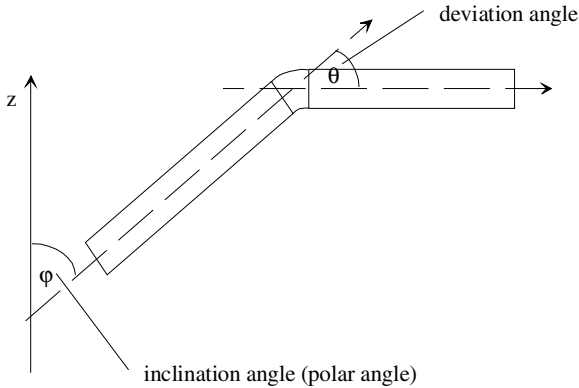
$$\varphi = \arccos \left[ \frac{(z_{k+1} - z_k)}{|\mathbf{r}_{k+1,k}|} \right],$$

is called *inclination angle (polar angle)*.

Here  $k$  and  $k+1$  are the two end points of the segment. The inclination angle is defined as the deviation of the positive oriented section axis with respect to the upwards oriented vertical direction (the negative gravity direction).

The length of each pipe section,

$$|\mathbf{r}_{k+1,k}| = \sqrt{(x_{k+1} - x_k)^2 + (y_{k+1} - y_k)^2 + (z_{k+1} - z_k)^2},$$



**Fig. 12.19** Definition of deviation and inclination angles

and its characteristic angles can then be computed automatically with the already specified information. Both angles and the section length are basic geometrical inputs for the definition of the mathematical flow description problem.

### 12.17.3 Diameters of pipe sections

A *pipe section* is a part of the pipe being between two neighboring characteristic points, see Fig. 12.17. The pipe section is per definition a *straight* piece of a pipe. For every pipe section

- the pipe inner diameter defining the flow cross section,
- the pipe hydraulic diameter,
- the pipe material,
- the wall thickness and
- the roughness.

have to be defined, see Fig. 12.18.

The default for the inner diameter is the normalizing diameter of the pipe.

The hydraulic diameter of a flow channel is defined as 4 times the flow cross section divided by the wetted perimeter corresponding to this cross section. By default the hydraulic diameter is equal to the inner diameter. That means the pipes are assumed by default to be circular tubes.

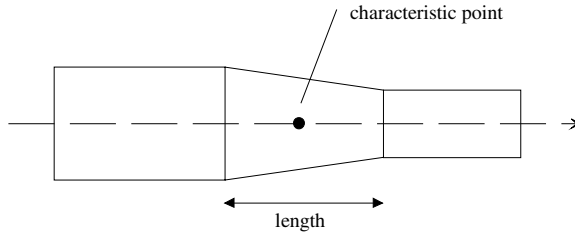
Examples for default definitions are: the material is stainless steel, the wall thickness is 0.1 times the inner diameter, and the roughness is 0.00004 *m*. These values can then be changed by per input.

### 12.17.4 Reductions

A smooth change in the pipe inner diameter or hydraulic diameter is expressed through a form piece called *reduction* – see Fig. 12.20. The reduction has a



specified length and is centered on the corresponding characteristic point. At its ends it has the diameters of the adjacent pipe sections and the diameter varies linearly between the two ends. On the contrary, an *abrupt area change* is expressed exactly through the characteristic point itself.



**Fig. 12.20** Definition of reductions

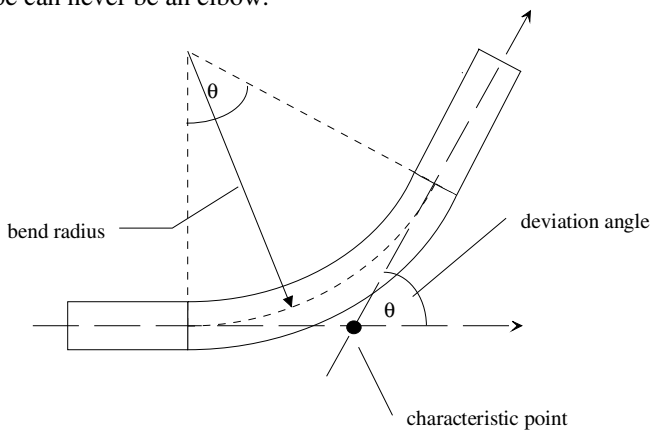
### 12.17.5 Elbows

*Elbows* are associated with points, where the pipe segment axis changes its direction in space – see Fig. 12.21. The flow axis coincides before and after the bend with the pipe segment axis. Note that the cross point of these axes is the characteristic point with which an elbow is associated.

The flow axis of an elbow possesses a bend radius which is in fact the curvature radius of the flow axis. Example for default value for the bend radius is 1.5 times the inner diameter of the preceding section. This value can be then changed by input. For example characteristic points which are not associated with an elbow, may receive a very large default value of the bend radius – 100 *m*.

Therefore a pipe section possesses an elbow on one of its ends, if the deviation angle is greater than zero and the bend radius of the corresponding characteristic point is less than the specified default.

Note that in accordance with the above definition the start point or the end point of a pipe can never be an elbow.



**Fig. 12.21** Definition of elbows

### 12.17.6 Creating a library of pipes

After the pipe data have been specified correctly, the pipe definition has to be automatically saved in a file. A *pipe library* is a file system containing an arbitrary number of pipe defining files. Creating a library of files “pipes” allows one later to simply interconnect them and use some of them repeatedly for different problems.

### 12.17.7 Sub system network

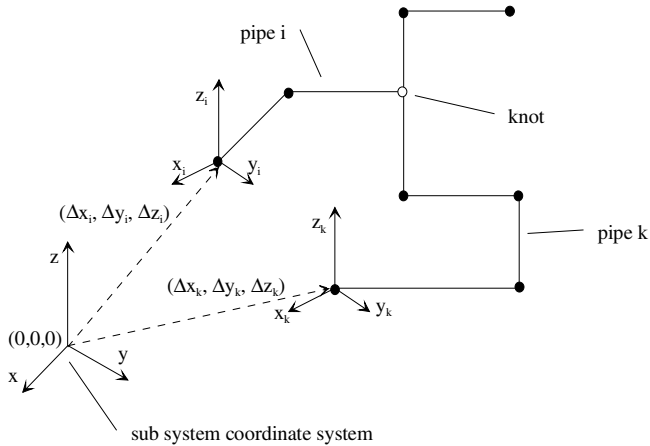
A *sub system network* consists of a number of *pipes* already defined in the pipe library which are linked together through *knots* (Fig. 12.22). The sub system network is defined in its own coordinate system having default coordinates

$$(0,0,0).$$

The sub system network definition contains also

- a) a list of all involved pipes and
- b) the shift of the particular pipe attached coordinate system in the new sub system coordinate system.

Note: a sub system network can also consist of only one pipe.



**Fig. 12.22** Definition of sub system network

Data of any pipe in the sub system network list should be allowed to be changed if necessary. Earlier defined pipes can be loaded into the list. Pipes can also be removed from the list.

Once the list of pipes necessary to form the network is complete, the pipes can be linked together by editing coordinates of the starting points.

$$(x_0, y_0, z_0)$$

They act as an offset to the particular pipe attached coordinate system. The pipe internal coordinate system itself is not affected.

### 12.17.8 Discretization of pipes

For the numerical computation of the system of differential equations, we use usually a finite volume technique to discretize the system. Therefore each pipe is divided into finite control volumes, at whose centers the flow properties like pressure, temperature, mass etc. are defined. We call these the real cells of the pipe. They are numbered with increasing cell indices e.g.  $i$ . The increasing cell indices define explicitly the positive flow direction as illustrated in Fig. 12.23. The staggered grid method implies a second set of cells, called the momentum cells, e.g.  $k$ . These are used for the discretization of the momentum equations and are located at the upper boundary of the corresponding real cells. The velocity is defined for these momentum cells. In addition, two auxiliary cells with zero length are introduced: one representing the pipe inlet and the other representing the pipe outlet. These two are needed to set proper boundary conditions at the pipe edges. The relation between the real cell numbering and the computational cell numbering is simply  $k = i + 2(n - 1) + 1$ , where  $n$  is the consecutive number of pipes. It is wise to store the computational cell numbers for the entrance and exit cells of each pipe. Then the cycles can be organized by visiting in each pipe the cells from the entrance to the exit.

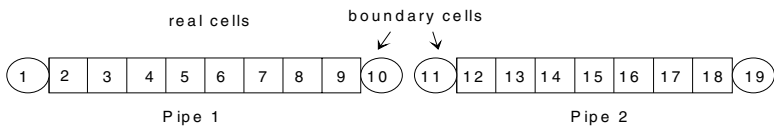


Fig. 12.23 Pipes

Equidistant discretization is recommended but not always possible in technical systems for which the exact positioning of bents, valves etc. is important. A remedy is to try to discretize each segment equidistantly.

### 12.17.9 Knots

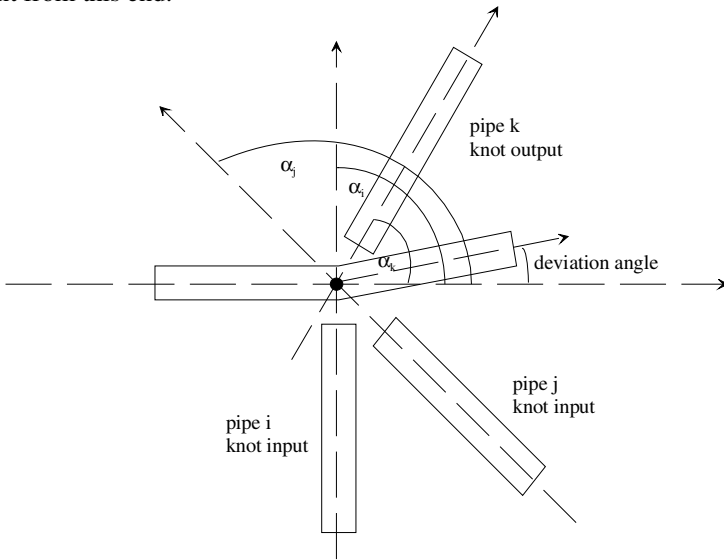
Several pipes can be linked together through knots to design a pipe network. The knots are arbitrarily numbered with indices ranging from 1 to the total number of knots inside the network. The flow passes through the auxiliary cells of the pipes starting or ending at the knot. The pipes starting out of the knot are called knot outputs; the pipes ending into the knot are called knot inputs. The auxiliary cells represent the openings of the knot into the pipe. They are assumed identical with the knot cell and therefore have the same cell properties. A good knot model

provides a momentum transfer from one pipe into another. Therefore the angles between the pipes have to be specified. The inclination angles of the pipes connected through the knot are defined as the deviation from the positive direction of the knot cell, regardless of whether the corresponding pipe is a knot input or a knot output (see also Fig. 12.24).

How to recognize *knots*?

A knot must be a characteristic point belonging to one of the pipes included in the sub system satisfying the following conditions:

If one of the ends of at least one other pipe coincides with a characteristic point of the pipe, this characteristic point is identified also as a knot. The same is valid, if one of the ends of the pipe itself coincides with a characteristic point of the pipe different from this end.



**Fig. 12.24** Definition of knots

Thus, inspecting all of the characteristic points of one pipe and checking whether the first or the last points of other pipes - or of the same pipe, if the characteristic point is different from this end - coincide with it, we identify whether knots are present in this pipe or not. If during this search pipe ends of other pipes are identified to belong to a knot, they are marked as “*already belongs to knot*”.

The procedure is repeated for all pipes successively in order of their appearance in the network list. The pipe ends already having the mark “*already belongs to knot*” are excluded from the further checks. Thus, all participating pipes are visited. Finally the total number of knots is identified. The knots are numbered from one to their total number in the order in which they are determined.

Additional information is derived from the above specified knot data for the computational analysis by answering the following questions: How many pipes enter the knot? Which pipes are these? How many pipes exit the knot? Which pipes are these? What are the inclination angles of the entering pipes with respect

to the positive oriented axis pointing to the knot of the pipe which the knot belongs to? What are the inclination angles of the exiting pipes with respect to the positive oriented axis pointing to the knot of the pipe which the knot belongs to? The pipe to which the knot belongs is excluded from this questionnaire.

The inclination angles of the entering and exiting pipes with respect to the positive oriented axis pointing to the knot of the pipe which the knot belongs to, are computed automatically. If the characteristic point identified as a knot belongs to a change in direction, the angle is computed with respect to the section before the characteristic point. In this case the direction change is not considered as an elbow, but as a sharp kink (see Fig. 12.23).

So, the network definition can be then saved. A prior defined network definition file can be also loaded for modification or check if required. The strategy described in Section 12.16 was programmed by *Iris Roloff-Bock* in the graphical preprocessing system NETGEN and later transformed into SONIA graphical preprocessing module by *Tony Chen*. With both systems a convenient way is provided for defining a network flow problems for the IVA computer code system.

### Appendix 12.1 Definitions applicable to discretization of the mass conservation equations

Following Section 12.5 the first order donor-cell discretized mass conservation equation (12.3) for each velocity field is

$$\begin{aligned}
 & (\alpha_l \rho_l \gamma_v - \alpha_{la} \rho_{la} \gamma_{va}) / \Delta \tau \\
 & + \frac{r_h^k \gamma_r}{r^k \Delta r} u_l \left\{ \frac{1}{2} [1 + \text{sign}(u_l)] \alpha_l \rho_l + \frac{1}{2} [1 - \text{sign}(u_l)] (\alpha_l \rho_l)_{i+1} \right\} \\
 & - \frac{(r_h^k \gamma_r)_{i-1}}{r^k \Delta r} u_{l,i-1} \left\{ \frac{1}{2} [1 + \text{sign}(u_{l,i-1})] (\alpha_l \rho_l)_{i-1} + \frac{1}{2} [1 - \text{sign}(u_{l,i-1})] \alpha_l \rho_l \right\} \\
 & + \frac{\gamma_\theta}{r^k \Delta \theta} v_l \left\{ \frac{1}{2} [1 + \text{sign}(v_l)] \alpha_l \rho_l + \frac{1}{2} [1 - \text{sign}(v_l)] (\alpha_l \rho_l)_{j+1} \right\} \\
 & - \frac{\gamma_{\theta,j-1}}{r^k \Delta \theta} v_{l,j-1} \left\{ \frac{1}{2} [1 + \text{sign}(v_{l,j-1})] (\alpha_l \rho_l)_{j-1} + \frac{1}{2} [1 - \text{sign}(v_{l,j-1})] \alpha_l \rho_l \right\} \\
 & + \frac{\gamma_z}{\Delta z} w_l \left\{ \frac{1}{2} [1 + \text{sign}(w_l)] \alpha_l \rho_l + \frac{1}{2} [1 - \text{sign}(w_l)] (\alpha_l \rho_l)_{k+1} \right\} \\
 & - \frac{\gamma_{z,k-1}}{\Delta z} w_{l,k-1} \left\{ \frac{1}{2} [1 + \text{sign}(w_{l,k-1})] (\alpha_l \rho_l)_{k-1} + \frac{1}{2} [1 - \text{sign}(w_{l,k-1})] \alpha_l \rho_l \right\}
 \end{aligned}$$

$$-\frac{1}{2}(\gamma_v + \gamma_{va})\mu_l = 0$$

Introducing the velocity normal to each surface of the discretization volume, Eq. (12.12), the  $b$  coefficients can be conveniently written as follows

$$b_{l1+} = +\beta_1 \frac{1}{2} [1 + \text{sign}(u_l)] u_l = +\beta_1 \xi_{l1+} u_l,$$

$$b_{l2+} = -\beta_2 \frac{1}{2} [1 - \text{sign}(u_{l,i-1})] u_{l,i-1} = -\beta_2 \xi_{l2+} u_{l,i-1},$$

$$b_{l3+} = +\beta_3 \frac{1}{2} [1 + \text{sign}(v_l)] v_l = +\beta_3 \xi_{l3+} v_l,$$

$$b_{l4+} = -\beta_4 \frac{1}{2} [1 - \text{sign}(v_{l,j-1})] v_{l,j-1} = -\beta_4 \xi_{l4+} v_{l,j-1},$$

$$b_{l5+} = +\beta_5 \frac{1}{2} [1 + \text{sign}(w_l)] w_l = +\beta_5 \xi_{l5+} w_l,$$

$$b_{l6+} = -\beta_6 \frac{1}{2} [1 - \text{sign}(w_{l,k-1})] w_{l,k-1} = -\beta_6 \xi_{l6+} w_{l,k-1},$$

or compactly written

$$b_{lm+} = \beta_m \xi_{lm+} V_{lm}^n,$$

and

$$b_{l1-} = -\beta_1 \frac{1}{2} [1 - \text{sign}(u_l)] u_l = -\beta_1 \xi_{l1-} u_l,$$

$$b_{l2-} = +\beta_2 \frac{1}{2} [1 + \text{sign}(u_{l,i-1})] u_{l,i-1} = +\beta_2 \xi_{l2-} u_{l,i-1},$$

$$b_{l3-} = -\beta_3 \frac{1}{2} [1 - \text{sign}(v_l)] v_l = -\beta_3 \xi_{l3-} v_l,$$

$$b_{l4-} = +\beta_4 \frac{1}{2} [1 + \text{sign}(v_{l,j-1})] v_{l,j-1} = +\beta_4 \xi_{l4-} v_{l,j-1},$$

$$b_{l5-} = -\beta_5 \frac{1}{2} [1 - \text{sign}(w_l)] w_l = -\beta_5 \xi_{l5-} w_l,$$

$$b_{l6-} = +\beta_6 \frac{1}{2} [1 + \text{sign}(w_{l,k-1})] w_{l,k-1} = +\beta_6 \xi_{l6-} w_{l,k-1},$$

or compactly written

$$b_{lm-} = -\beta_m \xi_{lm-} V_{lm}^n,$$

where the signed integers are

$$\xi_{l1+} = \frac{1}{2} [1 + \text{sign}(u_l)],$$

$$\xi_{l2+} = \frac{1}{2} [1 - \text{sign}(u_{l,i-1})],$$

$$\xi_{l3+} = \frac{1}{2} [1 + \text{sign}(v_l)],$$

$$\xi_{l4+} = \frac{1}{2} [1 - \text{sign}(v_{l,j-1})],$$

$$\xi_{l5+} = \frac{1}{2} [1 + \text{sign}(w_l)],$$

$$\xi_{l6+} = \frac{1}{2} [1 - \text{sign}(w_{l,k-1})],$$

and

$$\xi_{l1-} = \frac{1}{2} [1 - \text{sign}(u_l)],$$

$$\xi_{l2-} = \frac{1}{2} [1 + \text{sign}(u_{l,i-1})],$$

$$\xi_{l3-} = \frac{1}{2} [1 - \text{sign}(v_l)],$$

$$\xi_{l4-} = \frac{1}{2} [1 + \text{sign}(v_{l,j-1})],$$

$$\xi_{l5-} = \frac{1}{2} [1 - \text{sign}(w_l)],$$

$$\xi_{l6-} = \frac{1}{2} [1 + \text{sign}(w_{l,k-1})].$$

Obviously

$$\xi_{lm-} = 1 - \xi_{lm+}.$$

Note that the  $b_{lm+}$  coefficients are volume flows leaving the cell through the face  $m$  divided by the cell volume.  $b_{lm-}$  coefficients are volume flows entering the cell through the face  $m$  divided by the cell volume. Therefore the  $b$  coefficients have

the physical meaning of volumetric mass sources and sinks in the cell due to convection.

## Appendix 12.2 Discretization of the concentration equations

The concentration Eq. (12.4) will now be discretized following the procedure already described. The result is

$$\begin{aligned}
 & \alpha_{la} \rho_{la} \gamma_v (C_{il} - C_{ila}) / \Delta \tau \\
 & - \frac{r_h^k \gamma_r}{r^k \Delta r} \left\{ -u_l \frac{1}{2} [1 - \text{sign}(u_l)] (\alpha_l \rho_l)_{i+1} + \frac{D_{il1}^*}{\Delta r_h} A(|Pe_1|) \right\} (C_{il,i+1} - C_{il}) \\
 & - \frac{(r_h^k \gamma_r)_{i-1}}{r^k \Delta r} \left\{ u_{l,i-1} \frac{1}{2} [1 + \text{sign}(u_{l,i-1})] (\alpha_l \rho_l)_{i-1} + \frac{D_{il2}^*}{\Delta r_{h,i-1}} A(|Pe_2|) \right\} (C_{il,i-1} - C_{il}) \\
 & - \frac{\gamma_\theta}{r^k \Delta \theta} \left\{ -v_l \frac{1}{2} [1 - \text{sign}(v_l)] (\alpha_l \rho_l)_{j+1} + \frac{D_{il3}^*}{r^k \Delta \theta_h} A(|Pe_3|) \right\} (C_{il,j+1} - C_{il}) \\
 & - \frac{\gamma_{\theta,j-1}}{r^k \Delta \theta} \left\{ v_{l,j-1} \frac{1}{2} [1 + \text{sign}(v_{l,j-1})] (\alpha_l \rho_l)_{j-1} + \frac{D_{il4}^*}{r^k \Delta \theta_{h,j-1}} A(|Pe_4|) \right\} (C_{il,j-1} - C_{il}) \\
 & - \frac{\gamma_z}{\Delta z} \left\{ -w_l \frac{1}{2} [1 - \text{sign}(w_l)] (\alpha_l \rho_l)_{k+1} + \frac{D_{il5}^*}{\Delta z_h} A(|Pe_5|) \right\} (C_{il,k+1} - C_{il}) \\
 & - \frac{\gamma_{z,k-1}}{\Delta z} \left\{ w_{l,k-1} \frac{1}{2} [1 + \text{sign}(w_{l,k-1})] (\alpha_l \rho_l)_{k-1} + \frac{D_{il6}^*}{\Delta z_{h,k-1}} A(|Pe_6|) \right\} (C_{il,k-1} - C_{il}) \\
 & + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ C_{il} = \frac{1}{2} (\gamma_v + \gamma_{va}) DC_{il},
 \end{aligned}$$

where

$$A(|Pe_m|) = 1,$$

or

$$\begin{aligned}
 & \alpha_{la} \rho_{la} \gamma_{va} (C_{il} - C_{ila}) / \Delta \tau \\
 & - \left\{ b_{l-} (\alpha_l \rho_l)_{i+1} + \beta_l \frac{D_{il1}^*}{\Delta r_h} A(|Pe_1|) \right\} (C_{il,i+1} - C_{il})
 \end{aligned}$$



$$\begin{aligned}
& - \left\{ b_{l2-} (\alpha_l \rho_l)_{i-1} + \beta_2 \frac{D_{il2}^*}{\Delta r_{h,i-1}} A(|Pe_2|) \right\} (C_{il,i-1} - C_{il}) \\
& - \left\{ b_{l3-} (\alpha_l \rho_l)_{j+1} + \beta_3 \frac{D_{il3}^*}{r^x \Delta \theta_h} A(|Pe_3|) \right\} (C_{il,j+1} - C_{il}) \\
& - \left\{ b_{l4-} (\alpha_l \rho_l)_{j-1} + \beta_4 \frac{D_{il4}^*}{r^x \Delta \theta_{h,j-1}} A(|Pe_4|) \right\} (C_{il,j-1} - C_{il}) \\
& - \left\{ b_{l5-} (\alpha_l \rho_l)_{k+1} + \beta_5 \frac{D_{il5}^*}{\Delta z_h} A(|Pe_5|) \right\} (C_{il,k+1} - C_{il}) \\
& - \left\{ b_{l6-} (\alpha_l \rho_l)_{k-1} + \beta_6 \frac{D_{il6}^*}{\Delta z_{h,k-1}} A(|Pe_6|) \right\} (C_{il,k-1} - C_{il}) \\
& + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ C_{il} = \frac{1}{2} (\gamma_v + \gamma_{va}) DC_{il},
\end{aligned}$$

or compactly written

$$\begin{aligned}
& \alpha_{la} \rho_{la} \gamma_{va} (C_{il} - C_{ila}) / \Delta \tau - \sum_{m=1}^6 \left[ \underbrace{B_{lm-} + \beta_m \frac{D_{ilm}^*}{\Delta L_{h,m}} A(|Pe_m|)}_{b_m} \right] (C_{il,m} - C_{il}) \\
& + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ C_{il} = \frac{1}{2} (\gamma_v + \gamma_{va}) DC_{il}.
\end{aligned}$$

The harmonic averaged diffusion coefficients are given in Appendix 12.3. The  $l$ -field mass flow rate across the  $m$ -face is computed by means of the donor-cell concept as follows

$$\begin{aligned}
G_{l1} &= (\alpha_l \rho_l u_l)_1 = \left[ \xi_{l1+} \alpha_l \rho_l + \xi_{l1-} (\alpha_l \rho_l)_{i+1} \right] u_l, \\
G_{l2} &= (\alpha_l \rho_l u_l)_2 = \left[ \xi_{l2-} (\alpha_l \rho_l)_{i-1} + \xi_{l2+} \alpha_l \rho_l \right] u_{l,i-1}, \\
G_{l3} &= (\alpha_l \rho_l v_l)_3 = \left[ \xi_{l3+} \alpha_l \rho_l + \xi_{l3-} (\alpha_l \rho_l)_{j+1} \right] v_l, \\
G_{l4} &= (\alpha_l \rho_l v_l)_4 = \left[ \xi_{l4-} (\alpha_l \rho_l)_{j-1} + \xi_{l4+} \alpha_l \rho_l \right] v_{l,j-1}, \\
G_{l5} &= (\alpha_l \rho_l w_l)_5 = \left[ \xi_{l5+} \alpha_l \rho_l + \xi_{l5-} (\alpha_l \rho_l)_{k+1} \right] w_l,
\end{aligned}$$

$$G_{l_6} = (\alpha_l \rho_l w_l)_6 = \left[ \xi_{l_6^-} (\alpha_l \rho_l)_{k-1} + \xi_{l_6^+} \alpha_l \rho_l \right] w_{l,k-1}.$$

Note that the donor-cell concept takes into account the velocity directions. Computation of the harmonic averaged thermal conductivity coefficients is shown also in *Appendix 12.3*.

### Appendix 12.3 Harmonic averaged diffusion coefficients

A natural averaging of the coefficients describing diffusion across the face  $m$ , having surface cross section  $S_m$  is then the harmonic averaging

$$\frac{D_{l,m}^\Phi}{\Delta L_{h,m}} = \left( \frac{\Phi_l}{\Delta V} \right)_m S_m = S_m \frac{2(\Phi_l)(\Phi_l)_m}{\Delta V_m (\Phi_l) + \Delta V (\Phi_l)_m}$$

where in the right hand side  $m = 1, 2, 3, 4, 5, 6$  is equivalent to  $i + 1, i - 1, j + 1, j - 1, k + 1, k - 1$ , respectively regarding the properties inside a control volumes.  $\Delta V$  is the non-staggered cell volume, and  $\Delta V_m$  is the volume of the cell at the other side of face  $m$ . It guaranties that if the field in one of the neighboring cell is missing the diffusion coefficient is zero. This property is derived from the solution of the steady state one-dimensional diffusion equations.

For computation of

$$\frac{D_{il,m}^*}{\Delta L_{h,m}} = \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_m S_m = S_m \frac{2(\alpha_l \rho_l D_{il}^*)(\alpha_l \rho_l D_{il}^*)_m}{\Delta V_m (\alpha_l \rho_l D_{il}^*) + \Delta V (\alpha_l \rho_l D_{il}^*)_m}$$

we simply set  $\Phi_l = \alpha_l \rho_l D_{il}^*$ .

For computation of  $\frac{D_{l,m}^T}{\Delta L_{h,m}}$  we simply set  $\Phi_l = \alpha_l \lambda_l$ .

For computation of  $\frac{D_{il,m}^{sC}}{\Delta L_{h,m}}$  we simply set  $\Phi_l = \alpha_l \rho_l D_{il}^* (s_{il} - s_{il})$ . Note that

$$\frac{D_{il,m}^C}{\Delta L_{h,m}} = 0$$

for  $s_{il} = s_{il}$  or  $s_{il,m} = s_{il,m}$ .

For computation of the turbulent particle diffusion coefficient  $\frac{D_{l,m}^n}{\Delta L_{h,m}}$  we

simply set  $\Phi_l = \frac{v_l^t}{Sc^t}$ .

For computation of  $\frac{D_{l,m}^v}{\Delta L_{h,m}}$  we simply set  $\Phi_l = \alpha_l \rho_l v_l^*$ .

In case of cylindrical or Cartesian coordinate systems we have zero off-diagonal diffusion terms and

$$\frac{D_{l1}^\Phi}{\Delta r_h} = \frac{2(\Phi_l)(\Phi_l)_{i+1}}{\Delta r(\Phi_l)_{i+1} + \Delta r_{i+1}\Phi_l},$$

$$\frac{D_{l2}^\Phi}{\Delta r_{h,i-1}} = \frac{2(\Phi_l)(\Phi_l)_{i-1}}{\Delta r(\Phi_l)_{i-1} + \Delta r_{i-1}\Phi_l},$$

$$\frac{D_{l3}^\Phi}{r^\kappa \Delta \theta_h} = \frac{2(\Phi_l)(\Phi_l)_{j+1}}{r^\kappa [\Delta \theta(\Phi_l)_{j+1} + \Delta \theta_{j+1}\Phi_l]},$$

$$\frac{D_{l4}^\Phi}{r^\kappa \Delta \theta_{h,j-1}} = \frac{2(\Phi_l)(\Phi_l)_{j-1}}{r^\kappa [\Delta \theta(\Phi_l)_{j-1} + \Delta \theta_{j-1}\Phi_l]},$$

$$\frac{D_{l5}^\Phi}{\Delta z_h} = \frac{2(\Phi_l)(\Phi_l)_{k+1}}{\Delta z(\Phi_l)_{k+1} + \Delta z_{k+1}\Phi_l},$$

$$\frac{D_{l6}^\Phi}{\Delta z_{h,k-1}} = \frac{2(\Phi_l)(\Phi_l)_{k-1}}{\Delta z(\Phi_l)_{k-1} + \Delta z_{k-1}\Phi_l}.$$

## Appendix 12.4. Discretized radial momentum equation

The  $r$ -momentum equation discretized by using the donor-cell concept is

$$\begin{aligned} & \gamma_{va} \alpha_{lua} \rho_{lua} (u_l - u_{la}) / \Delta \tau \\ & - \frac{r_{i+1}^\kappa}{(r + \Delta r_h / 2)^\kappa \Delta r_h} \left\{ -\min [0, (\alpha_l \rho_l u_l \gamma_r)_1] + (\alpha_l \rho_l v_l \gamma_r)_1 \frac{1}{\Delta r_{i+1}} \right\} (u_{li+1} - u_l) \\ & - \frac{r^\kappa}{(r + \Delta r_h / 2)^\kappa \Delta r_h} \left\{ \max [0, (\alpha_l \rho_l u_l \gamma_r)_2] + (\alpha_l \rho_l v_l \gamma_r)_2 \frac{1}{\Delta r} \right\} (u_{li-1} - u_l) \\ & - \frac{1}{(r + \Delta r_h / 2)^\kappa \Delta \theta} \left\{ -\min [0, (\alpha_l \rho_l v_l \gamma_\theta)_3] + (\alpha_l \rho_l v_l \gamma_\theta)_3 \frac{1}{r_h^\kappa \Delta \theta_h} \right\} (u_{lj+1} - u_l) \end{aligned}$$

$$\begin{aligned}
 & -\frac{1}{(r+\Delta r_h/2)^k \Delta \theta} \left\{ \max \left[ 0, (\alpha_i \rho_i v_i \gamma_\theta)_4 \right] + (\alpha_i \rho_i v_i \gamma_\theta)_4 \frac{1}{r_h^k \Delta \theta_{h,j-1}} \right\} (u_{ij-1} - u_i) \\
 & -\frac{1}{\Delta z} \left\{ -\min \left[ 0, (\alpha_i \rho_i w_i \gamma_z)_5 \right] + (\alpha_i \rho_i v_i \gamma_z)_5 \frac{1}{\Delta z_h} \right\} (u_{ik+1} - u_i) \\
 & -\frac{1}{\Delta z} \left\{ \max \left[ 0, (\alpha_i \rho_i w_i \gamma_z)_6 \right] + (\alpha_i \rho_i v_i \gamma_z)_6 \frac{1}{\Delta z_{h,k-1}} \right\} (u_{ik-1} - u_i) \\
 & -\frac{\sin(\Delta \theta / 2)}{(r+\Delta r_h/2)^k \Delta \theta} \left\{ \frac{2}{(r+\Delta r_h/2)^k} \left\{ \begin{aligned} & (\alpha_i \rho_i v_i \gamma_\theta)_3 v_{i3} + (\alpha_i \rho_i v_i \gamma_\theta)_4 v_{i4} \\ & (\alpha \rho v)_{i3} \left[ \left( \frac{\partial v_i}{\partial \theta} \right)_3 + u_{i3} \right] \gamma_3 \\ & + (\alpha \rho v)_{i4} \left[ \left( \frac{\partial v_i}{\partial \theta} \right)_4 + u_{i4} \right] \gamma_4 \end{aligned} \right\} \right\} \\
 & + \alpha_{lua} \frac{\gamma_{rua}}{\Delta r_h} (p_{i+1} - p) \\
 & + \gamma_{vu} \left[ \alpha_{lua} \rho_{lua} g_r + c_{wlu} |u_l| u_l - \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \bar{c}_{ml}^d (u_m - u_l) + \bar{c}_{ml}^{vm} \frac{\partial}{\partial \tau} (u_m - u_l) \right] \right] \\
 & = \gamma_{vu} \left[ \mu_{wlu} (u_{wl} - u_l) - \mu_{lwu} (u_{lw} - u_l) + \sum_{m \neq l} \mu_{mlu} (u_m - u_l) \right] + f_{vlu}.
 \end{aligned}$$

Now we introduce the volume weighting coefficient

$$C^* = \frac{V_{i+1}}{V + V_{i+1}}.$$

This coefficient allows us easily to compute the mass flows entering the staggered cell through each face divided by the staggered cell volume by using the  $B_{lm-}$  flux densities already computed for the non-staggered cells

$$bu_{l,m} = -C^* \left( B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \right) - (1 - C^*) \left( B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \right)_{i+1}.$$

With this abbreviation we have

$$\gamma_{vu} \alpha_{lua} \rho_{lua} (u_l - u_{la}) / \Delta \tau + \sum_{m=1}^6 b u_{l,m} (u_{l,m} - u_l) \dots$$

Comparison with simple test problems for pure radial flow, pure rotational flow and superposition of both gives rise to improvements that will be described in a moment.

The first improvement is to use the following expression for the centrifugal force and its viscous counterpart

$$\begin{aligned} &-\frac{1}{2} \frac{1}{(r + \Delta r_h / 2)^k} \left\{ (\alpha_l \rho_l v_l \gamma_\theta)_3 v_{l3} + (\alpha_l \rho_l v_l \gamma_\theta)_4 v_{l4} \right. \\ &\left. - \frac{2}{(r + \Delta r_h / 2)^k} \left\{ (\alpha \rho v)_{l3} \left[ \left( \frac{\partial v_l}{\partial \theta} \right)_3 + u_{l3} \right] \gamma_3 + (\alpha \rho v)_{l4} \left[ \left( \frac{\partial v_l}{\partial \theta} \right)_4 + u_{l4} \right] \gamma_4 \right\} \right\} \end{aligned}$$

instead of the original form. The method of computing the centrifugal force can be understood as first computing the average flow rotation frequency and then the corresponding force. The analogy to the single-phase rigid body steady rotation, Appendix 2.3,

$$p_{i+1} - p = \frac{1}{2} \rho \omega^2 (r_{i+1}^2 - r^2)$$

or

$$\frac{p_{i+1} - p}{\Delta r_h} = \rho \omega^2 (r + \Delta r_h / 2),$$

is obvious.

The second improvement regards the convective term in the  $r$  direction. The exact solution of the single-phase steady state mass and momentum equation for radial flow gives the well-known *Bernulli* equation

$$\frac{1}{\Delta r_h} (p_{i+1} - p) = -\frac{1}{2} \rho \frac{1}{\Delta r_h} (u_{i+1}^{*2} - u^{*2}),$$

see Appendix 2.3. In this equation the velocities are defined at the same places where the pressures are defined. This is not the case with our finite difference equation. In order to make the finite difference form equivalent to *Bernulli* form we

use the single-phase case to derive a corrector multiplier as follows. We still use the donor-cell idea to compute the displaced velocities

$$u > 0 \qquad \qquad \qquad u \leq 0$$

$$u^*_{i+1} = u \frac{r_h}{r_{i+1}} \quad u^*_{i+1} = u_{i+1} \frac{r_{h,i+1}}{r_{i+1}}$$

$$u^* = u_{i-1} \frac{r_{h,i-1}}{r} \quad u^* = u \frac{r_h}{r}$$

Replacing in the *Bernulli* equation results in

$$\rho \frac{1}{\Delta r_h} \frac{1}{2} \left( u \frac{r_h}{r_{i+1}} + u_{i-1} \frac{r_{h,i-1}}{r} \right) \left( u \frac{r_h}{r_{i+1}} - u_{i-1} \frac{r_{h,i-1}}{r} \right) + \frac{1}{\Delta r_h} (p_{i+1} - p) = 0 \quad \text{for } u > 0$$

and

$$\rho \frac{1}{\Delta r_h} \frac{1}{2} \left( u_{i+1} \frac{r_{h,i+1}}{r_{i+1}} + u \frac{r_h}{r} \right) \left( u_{i+1} \frac{r_{h,i+1}}{r_{i+1}} - u \frac{r_h}{r} \right) + \frac{1}{\Delta r_h} (p_{i+1} - p) = 0 \quad \text{for } u \leq 0,$$

respectively. We multiply and divide the convective term by  $(u + u_{i-1})(u - u_{i-1})$  and use outside this product

$$u_{i-1} = u \frac{r_h}{r_{h,i-1}}.$$

The result is

$$\rho \frac{1}{\Delta r_h} \frac{\frac{1}{r_{i+1}^2} - \frac{1}{r^2}}{\frac{1}{r_h^2} - \frac{1}{r_{h,i-1}^2}} \frac{1}{2} (u + u_{i-1})(u - u_{i-1}) + \frac{1}{\Delta r_h} (p_{i+1} - p) = 0, \quad u > 0.$$

For the singularity  $r_{h,i-1} = 0$  we have

$$\rho \frac{1}{\Delta r_h} \left( \frac{r_h}{r_{i+1}} \right)^2 \frac{1}{2} (u + u_{i-1})(u - u_{i-1}) + \frac{1}{\Delta r_h} (p_{i+1} - p) = 0.$$

Similarly for negative velocity between the two pressures we have

$$u_{i+1} = u \frac{r_h}{r_{h,i+1}}$$

and therefore

$$\rho \frac{1}{\Delta r_h} \frac{\frac{1}{r_{i+1}^2} - \frac{1}{r^2}}{\frac{1}{r_{h,i+1}^2} - \frac{1}{r_h^2}} \frac{1}{2} (u_{i+1} + u)(u_{i+1} - u) + \frac{1}{\Delta r_h} (p_{i+1} - p) = 0, \quad u \leq 0.$$

Thus the final form of the multiphase convective term is written as

$$-\frac{1}{\Delta r_h} \frac{\frac{1}{r_{i+1}^2} - \frac{1}{r^2}}{\frac{1}{r_h^2} - \frac{1}{r_{h,i-1}^2}} \left\{ -\min [0, (\alpha_l \rho_l u_l \gamma_r)_1] + (\alpha_l \rho_l v_l \gamma_r)_1 \frac{1}{\Delta r_{i+1}} \right\} (u_{i+1} - u_l)$$

$$-\frac{1}{\Delta r_h} \frac{\frac{1}{r_{i+1}^2} - \frac{1}{r^2}}{\frac{1}{r_{h,i+1}^2} - \frac{1}{r_h^2}} \left\{ \max [0, (\alpha_l \rho_l u_l \gamma_r)_2] + (\alpha_l \rho_l v_l \gamma_r)_2 \frac{1}{\Delta r} \right\} (u_{i-1} - u_l).$$

It is convenient to compute the geometry coefficients at the start

$$\beta_{r1} = \frac{1}{\Delta r_h}, \beta_{r2} = \beta_{r1}, \beta_{r3} = \frac{1}{(r + \Delta r_h / 2)^\kappa \Delta \theta}, \beta_{r4} = \beta_{r3}, \beta_{r5} = \frac{1}{\Delta z}, \beta_{r6} = \beta_{r5}.$$

The  $bu$  coefficients are defined as follows:

$$bu_{l1} = -\beta_{r1} \frac{\frac{1}{r_{i+1}^2} - \frac{1}{r^2}}{\frac{1}{r_h^2} - \frac{1}{r_{h,i-1}^2}} \left\{ -\min [0, (\alpha_l \rho_l u_l \gamma_r)_1] + (\alpha_l \rho_l v_l \gamma_r)_1 \frac{1}{\Delta r_{i+1}} \right\},$$

$$bu_{l2} = -\beta_{r2} \frac{\frac{1}{r_{i+1}^2} - \frac{1}{r^2}}{\frac{1}{r_{h,i+1}^2} - \frac{1}{r_h^2}} \left\{ \max [0, (\alpha_l \rho_l u_l \gamma_r)_2] + (\alpha_l \rho_l v_l \gamma_r)_2 \frac{1}{\Delta r} \right\},$$

$$bu_{l3} = -\beta_{r3} \left\{ -\min [0, (\alpha_l \rho_l v_l \gamma_\theta)_3] + (\alpha_l \rho_l v_l \gamma_\theta)_3 \frac{1}{r_h^\kappa \Delta \theta_h} \right\},$$

$$bu_{l4} = -\beta_{r4} \left\{ \max [0, (\alpha_l \rho_l v_l \gamma_\theta)_4] + (\alpha_l \rho_l v_l \gamma_\theta)_4 \frac{1}{r_h^\kappa \Delta \theta_{h,j-1}} \right\},$$

$$bu_{i5} = -\beta_{r5} \left\{ -\min \left[ 0, (\alpha_i \rho_l w_l \gamma_z)_5 \right] + (\alpha_i \rho_l v_l \gamma_z)_5 \frac{1}{\Delta z_h} \right\},$$

$$bu_{i6} = -\beta_{r6} \left\{ \max \left[ 0, (\alpha_i \rho_l w_l \gamma_z)_6 \right] + (\alpha_i \rho_l v_l \gamma_z)_6 \frac{1}{\Delta z_{h,k-1}} \right\}.$$

Now we introduce the volume weighting coefficient

$$C^* = \frac{V_{i+1}}{V + V_{i+1}}.$$

This coefficient allows us easily to compute the mass flows entering the staggered cell through each face divided by the staggered cell volume by using the  $B_{lm-}$  flux densities already computed for the non-staggered cells

$$bu_{l,m} = -C^* \left( B_{lm-} + \beta_m \frac{D_{l,m}''}{\Delta L_{n,m}} \right) - (1 - C^*) \left( B_{lm-} + \beta_m \frac{D_{l,m}''}{\Delta L_{n,m}} \right)_{i+1}.$$

With this abbreviation we have

$$\gamma_{va} \alpha_{lva} \rho_{lva} (u_l - u_{la}) / \Delta \tau + \sum_{m=1}^6 bu_{l,m} (u_{l,m} - u_l) \dots$$

**Appendix 12.5 The  $\bar{a}$  coefficients for Eq. (12.46)**

$$\begin{aligned} \bar{a}_{11} &= a_{22}a_{33} - a_{32}a_{23} & \bar{a}_{12} &= a_{32}a_{13} - a_{12}a_{33} & \bar{a}_{13} &= a_{12}a_{23} - a_{22}a_{13} \\ \bar{a}_{21} &= a_{23}a_{31} - a_{21}a_{33} & \bar{a}_{22} &= a_{11}a_{33} - a_{31}a_{13} & \bar{a}_{23} &= a_{21}a_{13} - a_{23}a_{11} \\ \bar{a}_{31} &= a_{21}a_{32} - a_{31}a_{22} & \bar{a}_{32} &= a_{12}a_{31} - a_{32}a_{11} & \bar{a}_{33} &= a_{11}a_{22} - a_{21}a_{12} \end{aligned}$$

**Appendix 12.6 Discretization of the angular momentum equation**

The  $\theta$  momentum equation discretized by using the donor cell concept is

$$\begin{aligned} &\gamma_{vv} \alpha_{lva} \rho_{lva} (v_l - v_{la}) / \Delta \tau \\ &- \frac{r_h^{\kappa}}{r^{\kappa} \Delta r} \left\{ -\min \left[ 0, (\alpha_i \rho_l u_l \gamma_r)_1 \right] + (\alpha_i \rho_l v_l \gamma_r)_1 \frac{1}{\Delta r_h} \right\} (v_{i+1} - v_l) \end{aligned}$$



$$\begin{aligned}
 & -\frac{r_{h,i-1}^\kappa}{r^\kappa \Delta r} \left\{ \max \left[ 0, (\alpha_i \rho_l u_l \gamma_r)_2 \right] + (\alpha_i \rho_l v_l \gamma_r)_2 \frac{1}{\Delta r_{h,i-1}} \right\} (v_{i-1} - v_l) \\
 & -\frac{1}{r^\kappa \Delta \theta_h} \left\{ -\min \left[ 0, (\alpha_i \rho_l v_l \gamma_\theta)_3 \right] + (\alpha_i \rho_l v_l \gamma_\theta)_3 \frac{1}{r^\kappa \Delta \theta_{j+1}} \right\} (v_{ij+1} - v_l) \\
 & -\frac{1}{r^\kappa \Delta \theta_h} \left\{ \max \left[ 0, (\alpha_i \rho_l v_l \gamma_\theta)_4 \right] + (\alpha_i \rho_l v_l \gamma_\theta)_4 \frac{1}{r^\kappa \Delta \theta} \right\} (v_{ij-1} - v_l) \\
 & -\frac{1}{\Delta z} \left\{ -\min \left[ 0, (\alpha_i \rho_l w_l \gamma_z)_5 \right] + (\alpha_i \rho_l v_l \gamma_z)_5 \frac{1}{\Delta z_h} \right\} (v_{lk+1} - v_l) \\
 & -\frac{1}{\Delta z} \left\{ \max \left[ 0, (\alpha_i \rho_l w_l \gamma_z)_6 \right] + (\alpha_i \rho_l v_l \gamma_z)_6 \frac{1}{\Delta z_{h,k-1}} \right\} (v_{lk-1} - v_l) \\
 & + \frac{\sin(\Delta\theta/2)}{r^\kappa \Delta \theta_h} \\
 & \times \left[ \begin{aligned} & \left( (\alpha_i \rho_l v_l \gamma_\theta)_3 u_{l3} + (\alpha_i \rho_l v_l \gamma_\theta)_4 u_{l4} - \left\{ \begin{aligned} & (\alpha \rho v)_{ij+1} \gamma_{\theta 3} \left[ r \frac{\partial}{\partial r} \left( \frac{v_l}{r} \right) + \frac{1}{r} \frac{\partial u_l}{\partial \theta} \right]_3 \\ & + (\alpha \rho v)_l \gamma_{\theta 4} \left[ r \frac{\partial}{\partial r} \left( \frac{v_l}{r} \right) + \frac{1}{r} \frac{\partial u_l}{\partial \theta} \right]_4 \end{aligned} \right\} \right] \end{aligned} \right. \\
 & \left. + \alpha_{lva} \frac{\gamma_\theta}{r^\kappa \Delta \theta_h} (p_{j+1} - p) \right. \\
 & \left. + \gamma_{vv} \left[ \alpha_{lva} \rho_{lva} g_\theta + c_{wlv} |v_l| v_l - \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \bar{c}_{ml}^d (v_m - v_l) + \bar{c}_{ml}^{vm} \frac{\partial}{\partial \tau} (v_m - v_l) \right] \right] \right. \\
 & \left. = \gamma_{vv} \left[ \mu_{wlv} (v_{wl} - v_l) - \mu_{lvv} (v_{lv} - v_l) + \sum_{\substack{m=1 \\ m \neq l}}^3 \mu_{mlv} (v_m - v_l) \right] + f_{vlv} \right.
 \end{aligned}$$

Introducing the geometry coefficients

$$\beta_{\theta 1} = \frac{r_h^\kappa}{r^\kappa \Delta r}, \beta_{\theta 2} = \frac{r_{h,i-1}^\kappa}{r^\kappa \Delta r}, \beta_{\theta 3} = \frac{1}{r^\kappa \Delta \theta_h}, \beta_{\theta 4} = \beta_{\theta 3}, \beta_{\theta 5} = \frac{1}{\Delta z}, \beta_{\theta 6} = \beta_{\theta 5},$$

the  $bv$  coefficients are defined as follows

$$\begin{aligned}
 bv_{i1} &= -\beta_{\theta 1} \left\{ -\min \left[ 0, (\alpha_i \rho_l u_l \gamma_r)_1 \right] + (\alpha_i \rho_l v_l \gamma_r)_1 \frac{1}{\Delta r_h} \right\}, \\
 bv_{i2} &= -\beta_{\theta 2} \left\{ \max \left[ 0, (\alpha_i \rho_l u_l \gamma_r)_2 \right] + (\alpha_i \rho_l v_l \gamma_r)_2 \frac{1}{\Delta r_{h,i-1}} \right\}, \\
 bv_{i3} &= -\beta_{\theta 3} \left\{ -\min \left[ 0, (\alpha_i \rho_l v_l \gamma_\theta)_3 \right] + (\alpha_i \rho_l v_l \gamma_\theta)_3 \frac{1}{r^k \Delta \theta_{j+1}} \right\}, \\
 bv_{i4} &= -\beta_{\theta 4} \left\{ \max \left[ 0, (\alpha_i \rho_l v_l \gamma_\theta)_4 \right] + (\alpha_i \rho_l v_l \gamma_\theta)_4 \frac{1}{r^k \Delta \theta} \right\}, \\
 bv_{i5} &= -\beta_{\theta 5} \left\{ -\min \left[ 0, (\alpha_i \rho_l w_l \gamma_z)_5 \right] + (\alpha_i \rho_l v_l \gamma_z)_5 \frac{1}{\Delta z_h} \right\}, \\
 bv_{i6} &= -\beta_{\theta 6} \left\{ \max \left[ 0, (\alpha_i \rho_l w_l \gamma_z)_6 \right] + (\alpha_i \rho_l v_l \gamma_z)_6 \frac{1}{\Delta z_{h,k-1}} \right\}.
 \end{aligned}$$

Now we introduce the volume weighting coefficient

$$C^* = \frac{V_{j+1}}{V + V_{j+1}}.$$

This coefficient allows us easily to compute the mass flows entering the staggered cell through each face divided by the staggered cell volume by using the  $B_{m-}$  flux densities already computed for the non-staggered cells

$$bv_{l,m} = -C^* \left( B_{m-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \right) - (1 - C^*) \left( B_{m-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \right)_{j+1}.$$

With this abbreviation we have

$$\gamma_{vw} \alpha_{lva} \rho_{lva} (v_l - v_{la}) / \Delta \tau + \sum_{m=1}^6 bv_{l,m} (v_{l,m} - v_l) \dots$$

## Appendix 12.7 Discretization of the axial momentum equation

The  $z$  momentum equation discretized by using the donor cell concept is

$$\gamma_{vw} \alpha_{lva} \rho_{lva} (w_l - w_{la}) / \Delta \tau$$

$$\begin{aligned}
 & -\frac{r_h^\kappa}{r^\kappa \Delta r} \left\{ -\min \left[ 0, (\alpha_l \rho_l u_l \gamma_r)_1 \right] + (\alpha_l \rho_l v_l \gamma_r)_1 \frac{1}{\Delta r_h} \right\} (w_{li+1} - w_l) \\
 & -\frac{r_{h,i-1}^\kappa}{r^\kappa \Delta r} \left\{ \max \left[ 0, (\alpha_l \rho_l u_l \gamma_r)_2 \right] + (\alpha_l \rho_l v_l \gamma_r)_2 \frac{1}{\Delta r_{h,i-1}} \right\} (w_{li-1} - w_l) \\
 & -\frac{1}{r^\kappa \Delta \theta} \left\{ -\min \left[ 0, (\alpha_l \rho_l v_l \gamma_\theta)_3 \right] + (\alpha_l \rho_l v_l \gamma_\theta)_3 \frac{1}{r^\kappa \Delta \theta_h} \right\} (w_{lj+1} - w_l) \\
 & -\frac{1}{r^\kappa \Delta \theta} \left\{ \max \left[ 0, (\alpha_l \rho_l v_l \gamma_\theta)_4 \right] + (\alpha_l \rho_l v_l \gamma_\theta)_4 \frac{1}{r^\kappa \Delta \theta_{h,j-1}} \right\} (w_{lj-1} - w_l) \\
 & -\frac{1}{\Delta z_h} \left\{ -\min \left[ 0, (\alpha_l \rho_l w_l \gamma_z)_5 \right] + (\alpha_l \rho_l v_l \gamma_z)_5 \frac{1}{\Delta z_{k+1}} \right\} (w_{lk+1} - w_l) \\
 & -\frac{1}{\Delta z_h} \left\{ \max \left[ 0, (\alpha_l \rho_l w_l \gamma_z)_6 \right] + (\alpha_l \rho_l v_l \gamma_z)_6 \frac{1}{\Delta z} \right\} (w_{lk-1} - w_l) + \alpha_{lwa} \frac{\gamma_z}{\Delta z} (p_{k+1} - p) \\
 & + \gamma_{vw} \left[ \alpha_{lwa} \rho_{lwa} g_z + c_{wlw} |w_l| w_l - \sum_{\substack{m=1 \\ m \neq l}}^3 \left[ \bar{c}_{ml}^d (w_m - w_l) + \bar{c}_{ml}^{vm} \frac{\partial}{\partial \tau} (w_m - w_l) \right] \right] \\
 & = \gamma_{vw} \left[ \mu_{wlw} (w_{wl} - w_l) - \mu_{lww} (w_{lw} - w_l) + \sum_{\substack{m=1 \\ m \neq l}}^3 \mu_{mlw} (w_m - w_l) \right] + f_{vbw}
 \end{aligned}$$

Introducing the geometry coefficients

$$\beta_{z1} = \frac{r_h^\kappa}{r^\kappa \Delta r}, \beta_{z2} = \frac{r_{h,i-1}^\kappa}{r^\kappa \Delta r}, \beta_{z3} = \frac{1}{r^\kappa \Delta \theta}, \beta_{z4} = \beta_{z3}, \beta_{z5} = \frac{1}{\Delta z_h}, \beta_{z6} = \beta_{z5}$$

the  $bw$  coefficients are defined as follows

$$\begin{aligned}
 bw_{l1} &= -\beta_{z1} \left\{ -\min \left[ 0, (\alpha_l \rho_l u_l \gamma_r)_1 \right] + (\alpha_l \rho_l v_l \gamma_r)_1 \frac{1}{\Delta r_h} \right\}, \\
 bw_{l2} &= -\beta_{z2} \left\{ \max \left[ 0, (\alpha_l \rho_l u_l \gamma_r)_2 \right] + (\alpha_l \rho_l v_l \gamma_r)_2 \frac{1}{\Delta r_{h,i-1}} \right\}, \\
 bw_{l3} &= -\beta_{z3} \left\{ -\min \left[ 0, (\alpha_l \rho_l v_l \gamma_\theta)_3 \right] + (\alpha_l \rho_l v_l \gamma_\theta)_3 \frac{1}{r^\kappa \Delta \theta_h} \right\},
 \end{aligned}$$

$$bw_{l4} = -\beta_{z4} \left\{ \max \left[ 0, (\alpha_l \rho_l v_l \gamma_\theta)_4 \right] + (\alpha_l \rho_l v_l \gamma_\theta)_4 \frac{1}{r^{\kappa} \Delta \theta_{h,j-1}} \right\},$$

$$bw_{l5} = -\beta_{z5} \left\{ -\min \left[ 0, (\alpha_l \rho_l w_l \gamma_z)_5 \right] + (\alpha_l \rho_l v_l \gamma_z)_5 \frac{1}{\Delta z_{k+1}} \right\},$$

$$bw_{l6} = -\beta_{z6} \left\{ \max \left[ 0, (\alpha_l \rho_l w_l \gamma_z)_6 \right] + (\alpha_l \rho_l v_l \gamma_z)_6 \frac{1}{\Delta z} \right\}.$$

Now we introduce the volume weighting coefficient

$$C^* = \frac{V_{k+1}}{V + V_{k+1}}.$$

This coefficient allows us easily to compute the mass flows entering the staggered cell through each face divided by the staggered cell volume by using the  $B_{lm-}$  flux densities already computed for the non-staggered cells

$$bw_{l,m} = -C^* \left( B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{n,m}} \right) - (1 - C^*) \left( B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{n,m}} \right)_{k+1}.$$

With this abbreviation we have

$$\gamma_{vw} \alpha_{lwa} \rho_{lwa} (w_l - w_{la}) / \Delta \tau + \sum_{m=1}^6 bw_{l,m} (w_{l,m} - w_l) \dots$$

## Appendix 12.8 Analytical derivatives for the residual error of each equation with respect to the dependent variables

The velocities are assumed to be linearly dependent on the pressures, with the result that derivatives for the velocities with respect to pressure can be obtained directly from the momentum equations

$$V_{lm}^n = dV_{lm}^n - RVel_{lm} (p_m - p).$$

The result is

$$\frac{\partial V_{lm}^n}{\partial p} = RVel_{lm}, \quad \frac{\partial V_{lm}^n}{\partial p_m} = -RVel_{lm}, \quad \frac{\partial V_{lm}^n}{\partial p_x} = 0 \quad \text{for } x \neq m,$$

and therefore

$$\frac{\partial b_{lm+}}{\partial p} = \beta_m \xi_{lm+} \frac{\partial V_{lm}^n}{\partial p} = \beta_m \xi_{lm+} RVel_{lm},$$

$$\frac{\partial b_{lm+}}{\partial p_m} = \beta_m \xi_{lm+} \frac{\partial V_{lm}^n}{\partial p_m} = -\beta_m \xi_{lm+} R V e l_{lm},$$

$$\frac{\partial b_{lm+}}{\partial p_x} = 0 \quad \text{for } x \neq m,$$

$$\frac{\partial b_{lm-}}{\partial p} = -\beta_m \xi_{lm-} \frac{\partial V_{lm}^n}{\partial p} = -\beta_m \xi_{lm-} R V e l_{lm}, \quad \frac{\partial b_{lm}}{\partial p} = \frac{\partial b_{lm-}}{\partial p},$$

$$\frac{\partial b_{lm-}}{\partial p_m} = -\beta_m \xi_{lm-} \frac{\partial V_{lm}^n}{\partial p_m} = \beta_m \xi_{lm-} R V e l_{lm}, \quad \frac{\partial b_{lm}}{\partial p_m} = \frac{\partial b_{lm-}}{\partial p_m},$$

$$\frac{\partial b_{lm-}}{\partial p_x} = 0 \quad \text{for } x \neq m, \quad \frac{\partial b_{lm}}{\partial p_x} = 0 \quad \text{for } x \neq m.$$

*Particle number density derivatives:* Taking the derivatives of the LHS of the discretized particle number density conservation equation

$$\begin{aligned} f_l^n &\equiv [n_l \gamma_v - (n_l \gamma_v)_a] / \Delta \tau + \sum_{m=1}^6 (b_{lm+} n_l - b_{lm-} n_{lm}) - \frac{1}{2} (\gamma_v + \gamma_{va}) D n_l \\ &= \frac{\partial f_l^n}{\partial n_l} (n_l - n_{l0}) = 0 \end{aligned}$$

we obtain:

$$\frac{\partial f_l^n}{\partial n_l} = \frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm+} > 0,$$

$$\frac{\partial f_l^n}{\partial p_m} = \sum_{x=1}^6 \left( n_l \frac{\partial b_{lm+}}{\partial p_x} - n_{lm} \frac{\partial b_{lm-}}{\partial p_x} \right) = -\beta_m (\xi_{lm+} n_l + \xi_{lm-} n_{lm}) R V e l_{lm},$$

for  $m = 1, 6$ ,

$$\frac{\partial f_l^n}{\partial p} = \sum_{m=1}^6 \left( n_l \frac{\partial b_{lm+}}{\partial p} - n_{lm} \frac{\partial b_{lm-}}{\partial p} \right) = \sum_{m=1}^6 [\beta_m (\xi_{lm+} n_l + \xi_{lm-} n_{lm}) R V e l_{lm}] = -\sum_{m=1}^6 \frac{\partial f_l^n}{\partial p_m}.$$

*Mass conservation derivatives:* Taking the derivatives of the LHS of the discretized mass conservation equation

$$\begin{aligned} f_l^\alpha &\equiv [\alpha_l \rho_l \gamma_v - (\alpha_l \rho_l \gamma_v)_a] / \Delta \tau + \sum_{m=1}^6 [b_{lm+} \alpha_l \rho_l - b_{lm-} (\alpha_l \rho_l)_m] - \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l \\ &= \frac{\partial f_l^\alpha}{\partial \alpha_l} (\alpha_l - \alpha_{l0}) = 0 \end{aligned}$$

we obtain

$$\begin{aligned}
 \frac{\partial f_l^\alpha}{\partial \alpha_l} &= \rho_l \left( \frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm^+} \right) \equiv \rho_l \frac{\partial f_l^n}{\partial n_l} > 0, \\
 \frac{\partial f_l^\alpha}{\partial s_l} &= \alpha_l \frac{\partial \rho_l}{\partial s_l} \left( \frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm^+} \right) \equiv \frac{\alpha_l}{\rho_l} \frac{\partial \rho_l}{\partial s_l} \frac{\partial f_l^\alpha}{\partial \alpha_l} \equiv \alpha_l \frac{\partial \rho_l}{\partial s_l} \frac{\partial f_l^n}{\partial n_l}, \\
 \frac{\partial f_l^\alpha}{\partial C_{il}} &= \alpha_l \frac{\partial \rho_l}{\partial C_{il}} \left( \frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm^+} \right) \equiv \frac{\alpha_l}{\rho_l} \frac{\partial \rho_l}{\partial C_{il}} \frac{\partial f_l^\alpha}{\partial \alpha_l} \equiv \alpha_l \frac{\partial \rho_l}{\partial C_{il}} \frac{\partial f_l^n}{\partial n_l}, \\
 \frac{\partial f_l^\alpha}{\partial p} &= \alpha_l \frac{\partial \rho_l}{\partial p} \left( \frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm^+} \right) + \sum_{m=1}^6 \left[ \alpha_l \rho_l \frac{\partial b_{lm^+}}{\partial p} - (\alpha_l \rho_l)_m \frac{\partial b_{lm^-}}{\partial p} \right] \\
 &= \frac{\alpha_l}{\rho_l \alpha_l^2} \frac{\partial f_l^\alpha}{\partial \alpha_l} + \sum_{m=1}^6 \left\{ \beta_m \left[ \xi_{lm^+} \alpha_l \rho_l + \xi_{lm^-} (\alpha_l \rho_l)_m \right] RVel_{lm} \right\}, \\
 \frac{\partial f_l^\alpha}{\partial p_m} &= \alpha_l \rho_l \sum_{x=1}^6 \frac{\partial b_{lx^+}}{\partial p_m} - \sum_{x=1}^6 \left[ (\alpha_l \rho_l)_x \frac{\partial b_{lx^-}}{\partial p_m} \right] - \alpha_{lm} b_{lm^-} \frac{\partial \rho_{lm}}{\partial p_m} \\
 &= -\beta_m \left[ \xi_{lm^+} \alpha_l \rho_l + \xi_{lm^-} (\alpha_l \rho_l)_m \right] RVel_{lm} + \beta_m \alpha_{lm} \xi_{lm^-} V_{lm}^n \frac{\partial \rho_{lm}}{\partial p_m} \\
 &= -\beta_m \left[ \xi_{lm^+} \alpha_l \rho_l RVel_{lm} + \xi_{lm^-} (\alpha_l \rho_l)_m \left( RVel_{lm} - \frac{V_{lm}^n}{\rho_{lm}} \frac{\partial \rho_{lm}}{\partial p_m} \right) \right] \quad \text{for } m = 1, 6.
 \end{aligned}$$

*Specific entropy derivatives:* Taking the derivatives of the LHS of the discretized entropy conservation equation

$$\begin{aligned}
 f_l^s &\equiv \alpha_a \rho_{la} \gamma_{va} \frac{s_l - s_{la}}{\Delta \tau} - \sum_{m=1}^6 [b_{lm^-} \alpha_{lm} \rho_{lm} (s_{lm} - s_l)] + \frac{1}{2} (\gamma_v + \gamma_{va}) (\mu_l^+ s_l - Ds_l^*) \\
 &= \frac{\partial f_l^s}{\partial s_l} (s_l - s_{l0}) = 0
 \end{aligned}$$

we obtain:

$$\begin{aligned}
 \frac{\partial f_l^s}{\partial s_l} &= \frac{\alpha_{la} \rho_{la} \gamma_{va}}{\Delta \tau} + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ + \sum_{m=1}^6 b_{lm^-} (\alpha_l \rho_l)_m \geq 0, \\
 \frac{\partial f_l^s}{\partial p} &= -\sum_{m=1}^6 (\alpha_l \rho_l)_m (s_{lm} - s_l) \frac{\partial b_{lm^-}}{\partial p} = \sum_{m=1}^6 (\alpha_l \rho_l)_m (s_{lm} - s_l) \beta_m \xi_{lm^-} RVel_{lm},
 \end{aligned}$$

$$\begin{aligned} \frac{\partial f_l^s}{\partial p_m} &= -\sum_{x=1}^6 (\alpha_l \rho_l)_x (s_{lx} - s_l) \frac{\partial b_{lx}}{\partial p_m} - b_{lm} \alpha_{lm} (s_{lm} - s_l) \frac{\partial \rho_{lm}}{\partial p_m} \\ &= -\beta_m \xi_{lm} \alpha_{lm} (s_{lm} - s_l) \left( \rho_{lm} R V e_{lm} - V_{lm}^n \frac{\partial \rho_{lm}}{\partial p_m} \right) \text{ for } m = 1, 6. \end{aligned}$$

*Inert mass conservation equation derivatives:* Taking the derivatives of the LHS of the discretized inert mass conservation equation

$$\begin{aligned} f_{il}^c &\equiv \alpha_{la} \rho_{la} \gamma_{va} \frac{C_{il} - C_{ila}}{\Delta \tau} - \sum_{m=1}^6 \left\{ \left[ b_{lm} (\alpha_l \rho_l)_m + \beta_m \frac{D_{ilm}^*}{\Delta L_{h,m}} A(|Pe_m|) \right] (C_{il,m} - C_{il}) \right\} \\ &+ \frac{1}{2} (\gamma_v + \gamma_{va}) (\mu_l^+ C_{il} - DC_{il}) = \frac{\partial f_{il}^c}{\partial C_{il}} (C_{il} - C_{il0}) = 0 \end{aligned}$$

we obtain

$$\begin{aligned} \frac{\partial f_{il}^c}{\partial C_{il}} &= \frac{\gamma_{va} \alpha_{la} \rho_{la}}{\Delta \tau} + \frac{1}{2} (\gamma_v + \gamma_{va}) \mu_l^+ + \sum_{m=1}^6 \left[ b_{lm} (\alpha_l \rho_l)_m + \beta_m \frac{D_{ilm}^*}{\Delta L_{h,m}} A(|Pe_m|) \right] \geq 0, \\ \frac{\partial f_{il}^c}{\partial p} &= -\sum_{m=1}^6 (\alpha_l \rho_l)_m (C_{ilm} - C_{il}) \frac{\partial b_{lm}}{\partial p} = \sum_{m=1}^6 (\alpha_l \rho_l)_m (C_{ilm} - C_{il}) \beta_m \xi_{lm} R V e_{lm}, \\ \frac{\partial f_{il}^c}{\partial p_m} &= -\sum_{x=1}^6 (\alpha_l \rho_l)_x (C_{ilx} - C_{il}) \frac{\partial b_{lx}}{\partial p_m} - b_{lm} \alpha_{lm} (C_{ilm} - C_{il}) \frac{\partial \rho_{lm}}{\partial p_m} \\ &= -\alpha_{lm} (C_{ilm} - C_{il}) \left( \rho_{lm} \frac{\partial b_{lm}}{\partial p_m} + b_{lm} \frac{\partial \rho_{lm}}{\partial p_m} \right) \\ &= -\beta_m \xi_{lm} \alpha_{lm} (C_{ilm} - C_{il}) \left( \rho_{lm} R V e_{lm} - V_{lm}^n \frac{\partial \rho_{lm}}{\partial p_m} \right) \text{ for } m=1,6. \end{aligned}$$

### Appendix 12.9 Simple introduction to iterative methods for solution of algebraic systems

Most of the iterative methods for solving the linear system of equations

$$\mathbf{Ax} = \mathbf{b}$$

are based on the regular splitting of the coefficient matrix

$$\mathbf{A} = \mathbf{B} - \mathbf{R}.$$

All three matrices above are of the same range. The iteration method is then

$$\mathbf{B}\mathbf{x}^{n+1} = \mathbf{R}\mathbf{x}^n + \mathbf{b}$$

or after eliminating  $\mathbf{R}$

$$\mathbf{B}\mathbf{x}^{n+1} = \mathbf{B}\mathbf{x}^n + \mathbf{b} - \mathbf{A}\mathbf{x}^n .$$

Introducing the residual error vector at the previous solution

$$\mathbf{r}^n = \mathbf{b} - \mathbf{A}\mathbf{x}^n ,$$

we have a method for computing the next approximation

$$\mathbf{x}^{n+1} = \mathbf{x}^n + \mathbf{B}^{-1}\mathbf{r}^n ,$$

by adding to the previous estimate for the solution a correction vector  $\Delta\mathbf{x}^n$  ,

$$\mathbf{x}^{n+1} = \mathbf{x}^n + \Delta\mathbf{x}^n .$$

$\Delta\mathbf{x}^n$  is in fact the solution of the algebraic system

$$\mathbf{B}\Delta\mathbf{x}^n = \mathbf{r}^n .$$

The matrix  $\mathbf{B}$  has to be selected so as to enable non-expensive solution of the above system in terms of computer time and storage. Selecting

$$\mathbf{B} = \text{diag}(\mathbf{A})$$

results in the *Jacobi* method. Selection of  $\mathbf{B}$  to be the lower off diagonal part including the diagonal leads to the *Gauss-Seidel* method. Selecting  $\mathbf{B}$  to consist of non-expensively invertible blocks of  $\mathbf{A}$  leads to the so called *Block-Gauss-Seidel* method. The secret of creating a powerful method is in appropriate selection of  $\mathbf{B}$ . The reader will find valuable information on this subject in *Saad* (1996).

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# 13 Numerical methods for multi-phase flow in curvilinear coordinate systems

*This chapter presents a numerical solution method for multi-phase flow analysis based on local volume and time averaged conservation equations. The emphasis of this development was to create a computer code architecture that absorb all the constitutive physics and functionality from the past 25years development of the three fluid multi-component IVA-entropy concept for multi-phase flows into a boundary fitted orthogonal coordinate framework. Collocated discretization for the momentum equations is used followed by weighted averaging for the staggered grids resulting in analytical expressions for the normal velocities. Using the entropy concept analytical reduction to a pressure-velocity coupling is found. The performance of the method is demonstrated by comparison of two cases for which experimental results and numerical solution with the previous method are available. The agreement demonstrates the success of this development.*

## 13.1 Introduction

We extend now the method described in the previous chapter to more arbitrary geometry. Instead of considering Cartesian or cylindrical geometry only, we will consider an integration space called a *block* in which the computational finite volumes fit inside the block so that the outermost faces of the external layer of the finite volumes create the face of the block. Similarly, bodies immersed into this space have external faces identical with the faces of the environmental computational cells. Such blocks can be inter-connected. With this technology multi-phase flows in arbitrary interfaces can be conveniently handled.

For understanding the material presented in this section I strongly recommend going over Appendixes 1 and 2 before continuing reading.

Before starting with the description of the new method let us summarize briefly the state of the art in this field.

In the last ten years the numerical modeling of single-phase flow in boundary fitted coordinates is becoming standard in the industry. This is not the case with the numerical modeling of multiphase flows. There are some providers of single-phase-computer codes claiming that their codes can simulate multi-phase flows. Taking close looks of the solution methods of these codes reveals that existing single phase solvers are used and a provision is given to the user to add an other

velocity field and define explicit the interfacial interaction physics. This strategy does not account for the feed back of the strong interfacial interactions on the mathematical solution methods - see the discussion in *Miettinen* and *Schmidt* (2002). Multi-phase flow simulations require specific solution methods accounting for this specific physics - see for instance the discussion by *Antal* et al. (2000).

There are groups of methods that are solving single phase conservation equations with surface tracking, see the state of the art part of *Tryggvason* et al. (2001) work. This is in fact a direct numerical simulation that is outside of the scope of this chapter. To mention few of them: In Japan a powerful family of cubic-interpolation methods (CIP) is developed based on the pseudo-characteristic method of lines *Takewaki*, *Nishiguchi* and *Yabe* (1985), *Takewaki* and *Yabe* (1987), *Nakamura*, *Tanaka*, *Yabe* and *Takizawa* (2001), *Yabe*, *Xiao* and *Utsumi* (2001), *Yabe*, *Tanaka*, *Nakamura* and *Xiao* (2001), *Yabe*, *Xiao* and *Utsumi* (2001), *Yabe* and *Takei* (1988), *Xiao* and *Yabe* (2001), *Xiao*, *Yabe*, *Peng* and *Kobayashi* (2002), *Xiao* (2003), *Xiao* and *Ikebata* (2003), *Yabe* and *Wang* (1991), *Yabe* and *Aoki* (1991), *Yabe* et al. (1991). In USA particle tracking and level-set surface tracking methods are very popular; see for instance *Sussman*, *Smereka* and *Osher* (1994), *Osher* and *Fredkiw* (2003), *Swthian* (1996), *Tryggvason* et al. (2001). The third group of DNS method with surface tracking is the lattice-*Boltzman* family, see *Hou* et al. (1995), *Nourgaliev* et al. (2002) and the references given there. To the family of emerging methods the so called diffuse interface methods based on high order thermodynamics can be mentioned, see *Verschuieren* (1999), *Jamet* et al. (2001). Let us emphasize once again, that unlike those methods, our work is concentrated on methods solving the local volume and time averaged multiphase flow equations which are much different then the single phase equations.

In Europe two developments for solving two-fluid conservation equations in unstructured grids are known to me. *Staedke* et al. (1998) developed a solution method based on the method of characteristics using unstructured grid in a single domain. The authors added artificial terms to enforce hyperbolicity in the initially incomplete system of partial differential equations that contain derivatives which do not have any physical meaning. *Toumi* et al. (2000) started again from the incomplete system for two-fluid two phase flows without interaction terms and included them later for a specific class of processes; see *Kumbaro* et al. (2002). These authors extended the approximate *Riemann* solver originally developed for single phase flows by *Roe* to two-fluid flows. One application example of the method is demonstrated in a single space domain in *Toumi* et al. (2000), *Kumbaro* et al. (2002). No industrial applications of these two methods have been reported so far. One should note that it is well know that if proper local volume averaging is applied the originating interfacial interaction terms provide naturally hyperbolicity of the system of PDS and there is no need for artificial terms without any physical meaning. An example for the resolution of this problem is given by *van Wijngaarden* in 1976 among many others.

In USA *Lahey* and *Drew* demonstrated clearly in 1999 how by careful elaboration of the constitutive relationships starting from first principles variety of steady state processes including *frequency dependent acoustics* can be successfully simulated. Actually, the idea by *Lahey* and *Drew* (1999) is a further development

of the proposal made by *Harlow* and *Amsden* in 1975 where liquid (1) in vapor (2) and vapor (3) in liquid (4) are grouped in two velocity fields 1 + 2 and 3 + 4. The treatment of *Lahey* and *Drew* (1999) is based on four velocity fields. *Antal* et al. (2000) started developing the NPHASE multi-domain multi-phase flows code based on the single phase *Rhie* and *Chow* numerical method extended to multiphase flows. Application example is given for T-pipe bubble flows with 10 groups of bubble diameters. Two works that can be considered as a subset of this approach are reported in *Tomiyama* et al. (2000) *Gregor*, *Petelin* and *Tiselj* (2000). Another direction of development in USA that can be observed is the use of the volume of fluid method with computing the surface tension force as a volumetric force *Hirt* (1993), *Kothe* et al. (1996), *Brackbill* et al. (1992), *Rider* and *Kothe* (1998).

## 13.2 Nodes, grids, meshes, topology - some basic definitions

**Database concept:** Let us consider the data base concept. The data volume is made up of points, or *nodes*, which themselves define in their neighborhood a volume element. We use hexahedrons (Fig. 4 in Appendix 1). A hexahedron is a 3D volume element with six sides and eight vertices. The vertices are connected in an order that mimics the way nodes are numbered in the data volume. The nodes in the data volume are numbered by beginning with 1 at the data volume's origin. Node numbering increases, with  $x$  changing fastest. This means node numbering increases along the  $x$  axis first, the  $y$  axis second, and the  $z$  axis last, until all nodes are numbered. The numbering of the vertices of the volume elements follows the same rule. It starts with the vertex being closest to the data volume's origin, moves along  $x$ , then  $y$ , and then  $z$ .

**Grid:** A *grid* is a set of locations in a 3D data volume defined with  $x$ ,  $y$ , and  $z$  coordinates. The locations are called *nodes*, which are connected in a specific order to create the topology of the grid. A grid can be regular or irregular depending on how its nodes are represented as points.

A *regular grid*'s nodes are evenly spaced in  $x$ ,  $y$ , and  $z$  directions, respectively. A *regular grid*'s nodes are specified with  $x$ ,  $y$ , and  $z$  offsets from the data volume's origin. A regular grid may have equidistant or non-equidistant spacing. If equidistant spacing is used all areas of the data volumes have the same resolution. This suits data with regular sample intervals.

An *irregular grid*'s nodes need not be evenly spaced or in a rectangular configuration. This suits data with a specific area of interest that require finer sampling. Because nodes may not be evenly spaced, each node's  $xyz$  coordinates must be explicitly listed.

**Topology:** A *topology* defines an array of elements by specifying the connectivity of the element's vertices or nodes. It builds a volume from separate elements by specifying how they are connected together. The elements can be



3D volume elements or 2D surface elements. A topology is either regular or irregular, depending upon what types of elements it defines and how they are structured.

A *regular topology* defines a data volume's node connectivity. We assume a hexahedron volume element type. A regular topology can be used for regular or irregular grids.

An *irregular topology* defines the node connectivity of either a data volume or geometric surface elements. An irregular topology data volume can be composed of either hexahedron or tetrahedron volume elements. Geometry objects are composed of points, lines, or polygons. An element list has to explicitly specify how the nodes connect to form these elements.

**Volume elements:** The volume elements are the smallest building blocks of a data volume topology.

**Mesh:** A *mesh* is a grid combined with specific topology for the volume of data. We distinguish the following mesh types: regular meshes, irregular or structured meshes, unstructured meshes, and geometry meshes.

A *regular mesh* consists of grid having regular spacing and regular topology that consists of simple, rectangular array of volume elements.

An *irregular mesh* explicitly specifies the  $xyz$  coordinates of each node in a node list. As in the regular mesh, the topology is regular, although individual elements are formed by explicit  $xyz$  node locations. The grid may be irregular or rectilinear.

An unstructured mesh explicitly defines the topology. Each topology element is explicitly defined by its node connectivity in an element list. The grid may be regular or irregular.

In this sense we are dealing in this Chapter with *multi-blocks* each of them consisting of

- irregular grid's nodes, irregular meshes,
- regular topology with hexahedron volume element type.

The integration space is built by a specified number of interconnected blocks.

### 13.3 Formulation of the mathematical problem

Consider the following mathematical problem: A multi-phase flow is described by the following vector of dependent variables

$$\mathbf{U}^T = (\alpha_m, T_1, s_2, s_3, C_{il}, n_l, p, u_l, v_l, w_l),$$

where

$$l = 1, 2, 3, \quad i1 = 1 \dots n1, \quad i2 = 1 \dots n2, \quad i3 = 1 \dots n3$$

which is a function of the three space coordinates  $(x, y, z)$ , and of the time  $\tau$ ,

$$\mathbf{U} = \mathbf{U}(x, y, z, \tau).$$

The relationship  $\mathbf{U} = \mathbf{U}(x, y, z, \tau)$  is defined by the volume-averaged and successively time-averaged mass, momentum and energy conservation equations derived in Chapters 1, 2, 5 *Kolev* (1994a, b, 1995, 1997, 1998) as well as by initial conditions, boundary conditions, and geometry. The conservation equations are transformed in a curvilinear coordinate system  $\xi, \eta, \zeta$  as shown in Chapter 11, *Kolev* (2001). The flux form of these equations is given in Chapter 11. As shown in Chapter 11, *Kolev* (2001), the conservation principles lead to a system of  $19+n1+n2+n3$  non-linear, non-homogeneous partial differential equations with variable coefficients. This system is defined in the three-dimensional domain  $\mathbf{R}$ . The initial conditions of  $\mathbf{U}(\tau = 0) = \mathbf{U}_a$  in  $\mathbf{R}$  and the boundary conditions acting at the interface separating the integration space from its environment are given. The solution required is for conditions after the time interval  $\Delta\tau$  has elapsed. The previous time variables are assigned the index  $a$ . The time variables not denoted with  $a$  are either in the new time plane, or are the best available guesses for the new time plane.

In order to enable modeling of flows with arbitrary obstacles and inclusions in the integration space as is usually expected for technical applications, surface permeabilities are defined

$$(\gamma_\xi, \gamma_\eta, \gamma_\zeta) = \text{functions of } (\xi, \eta, \zeta, \tau),$$

at the virtual surfaces that separate each computational cell from its environment. By definition, the surface permeabilities have values between one and zero,

$$0 \leq \text{each of all } (\gamma_\xi, \gamma_\eta, \gamma_\zeta) \text{ 's} \leq 1.$$

A volumetric porosity

$$\gamma_v = \gamma_v(\xi, \eta, \zeta, \tau)$$

is assigned to each computational cell, with

$$0 < \gamma_v \leq 1.$$

The surface permeabilities and the volume porosities are not expected to be smooth functions of the space coordinates in the region  $\mathbf{R}$  and of time. For this reason, one constructs a frame of geometrical flow obstacles which are functions of space and time. This permits a large number of extremely interesting technical applications of this type of approach.

In order to construct useful numerical solutions it is essential that an appropriate set of constitutive relations be available: state equations, thermodynamic derivatives, equations for estimation of the transport properties, correlations modeling the heat, mass and momentum transport across the surfaces dividing the separate velocity fields, etc. These relationships together are called closure equations. This very complex problem will not be discussed in Volume II. Only the numerics will be addressed here.

## 13.4 Discretization of the mass conservation equations

### 13.4.1 Integration over a finite time step and finite control volume

We start with the conservation equation (10.56) for the species  $i$  inside the velocity field  $l$  in the curvilinear coordinate system

$$\begin{aligned} & \frac{\partial}{\partial \tau} (\alpha_i \rho_l C_{il} \sqrt{g} \gamma_v) + \frac{\partial}{\partial \xi} (\gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{G}_{il}) + \frac{\partial}{\partial \eta} (\gamma_\eta \sqrt{g} \mathbf{a}^2 \cdot \mathbf{G}_{il}) \\ & + \frac{\partial}{\partial \zeta} (\gamma_\zeta \sqrt{g} \mathbf{a}^3 \cdot \mathbf{G}_{il}) = \gamma_v \sqrt{g} \mu_{il}, \end{aligned} \quad (13.1)$$

where the species mass flow rate vector is defined as follows

$$\mathbf{G}_{il} = \alpha_i \rho_l \left[ C_{il} (\mathbf{V}_l - \mathbf{V}_{cs}) - D_{il} \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]. \quad (13.2)$$

Note that for  $C_{il} = 1$  we have the mass flow vector of the velocity field

$$\mathbf{G}_l = \alpha_l \rho_l (\mathbf{V}_l - \mathbf{V}_{cs}). \quad (13.3)$$

Next we will use the following basic relationships from Appendix 2 between the surface vectors and the contravariant vectors, and between the *Jacobian* determinant and the infinitesimal spatial and volume increments

$$\sqrt{g} \mathbf{a}^1 = \frac{\mathbf{S}^1}{\partial \eta \partial \zeta}, \quad \sqrt{g} \mathbf{a}^2 = \frac{\mathbf{S}^2}{\partial \xi \partial \zeta}, \quad \sqrt{g} \mathbf{a}^3 = \frac{\mathbf{S}^3}{\partial \xi \partial \eta}, \quad \sqrt{g} = \frac{dV}{d\xi d\eta d\zeta}. \quad (13.4-7)$$

We will integrate both sides of the equation over the time, and spatial intervals  $\partial \tau$ ,  $\partial \xi$ ,  $\partial \eta$ , and  $\partial \zeta$  respectively. We start with the first term

$$\begin{aligned} & \iiint_{\Delta V} \left[ \int_0^{\Delta \tau} \frac{\partial}{\partial \tau} (\alpha_i \rho_l C_{il} \sqrt{g} \gamma_v) \partial \tau \right] \partial \xi \partial \eta \partial \zeta = \left[ \int_0^{\Delta \tau} \frac{\partial}{\partial \tau} (\alpha_i \rho_l C_{il} \gamma_v) \partial \tau \right] \iiint_{\Delta V} dV \\ & = \left[ (\alpha_i \rho_l C_{il} \gamma_v) - (\alpha_i \rho_l C_{il} \gamma_v)_a \right] \Delta V. \end{aligned} \quad (13.8)$$

This result is obtained under the assumption that there is no spatial variation of the properties  $(\alpha_i \rho_l C_{il} \gamma_v)$  inside the cell. The integration of the other terms gives the following results:

$$\begin{aligned} & \int_0^{\Delta \tau} \iiint_{\Delta V} \frac{\partial}{\partial \xi} (\gamma_\xi \sqrt{g} \mathbf{a}^1 \cdot \mathbf{G}_{il}) \partial \xi \partial \eta \partial \zeta \partial \tau = \Delta \tau \int_0^{\Delta \xi} \frac{\partial}{\partial \xi} (\gamma_\xi \mathbf{S}^1 \cdot \mathbf{G}_{il}) \partial \xi \\ & = \Delta \tau \left[ (\gamma_\xi \mathbf{S}^1 \cdot \mathbf{G}_{il}) - (\gamma_\xi \mathbf{S}^1 \cdot \mathbf{G}_{il})_{i-1} \right] \end{aligned}$$

$$= \Delta \tau \Delta V \left[ \gamma_{\xi} \frac{S_1}{\Delta V} (\mathbf{e}^1 \cdot \mathbf{G}_{il}) - \gamma_{\xi, j-1} \frac{S_2}{\Delta V} (\mathbf{e}^1 \cdot \mathbf{G}_{il})_{i-1} \right], \quad (13.9)$$

$$\begin{aligned} & \int_0^{\Delta \tau} \iiint_{\Delta V} \frac{\partial}{\partial \eta} (\gamma_{\eta} \sqrt{g} \mathbf{a}^2 \cdot \mathbf{G}_{il}) \partial \xi \partial \eta \partial \zeta \partial \tau = \Delta \tau \int_0^{\Delta \eta} \frac{\partial}{\partial \eta} (\gamma_{\eta} \mathbf{S}^2 \cdot \mathbf{G}_{il}) \partial \eta \\ & = \Delta \tau \left[ (\gamma_{\eta} \mathbf{S}^2 \cdot \mathbf{G}_{il}) - (\gamma_{\eta} \mathbf{S}^2 \cdot \mathbf{G}_{il})_{j-1} \right] \\ & = \Delta \tau \Delta V \left[ \gamma_{\eta} \frac{S_3}{\Delta V} (\mathbf{e}^2 \cdot \mathbf{G}_{il}) - \gamma_{\eta, j-1} \frac{S_4}{\Delta V} (\mathbf{e}^2 \cdot \mathbf{G}_{il})_{j-1} \right], \end{aligned} \quad (13.10)$$

$$\begin{aligned} & \int_0^{\Delta \tau} \iiint_{\Delta V} \frac{\partial}{\partial \zeta} (\gamma_{\zeta} \sqrt{g} \mathbf{a}^3 \cdot \mathbf{G}_{il}) \partial \xi \partial \eta \partial \zeta \partial \tau = \Delta \tau \int_0^{\Delta \zeta} \frac{\partial}{\partial \zeta} (\gamma_{\zeta} \mathbf{S}^3 \cdot \mathbf{G}_{il}) \partial \zeta \\ & = \Delta \tau \left[ (\gamma_{\zeta} \mathbf{S}^3 \cdot \mathbf{G}_{il}) - (\gamma_{\zeta} \mathbf{S}^3 \cdot \mathbf{G}_{il})_{k-1} \right] \\ & = \Delta \tau \Delta V \left[ \gamma_{\zeta} \frac{S_5}{\Delta V} (\mathbf{e}^3 \cdot \mathbf{G}_{il}) - \gamma_{\zeta, k-1} \frac{S_6}{\Delta V} (\mathbf{e}^3 \cdot \mathbf{G}_{il})_{k-1} \right], \end{aligned} \quad (13.11)$$

$$\int_0^{\Delta \tau} \iiint_{\Delta V} \gamma_v \sqrt{g} \mu_{il} \partial \xi \partial \eta \partial \zeta \partial \tau = \Delta \tau \iiint_{\Delta V} \gamma_v \mu_{il} dV = \Delta \tau \gamma_v \mu_{il} \Delta V. \quad (13.12)$$

It is convenient to introduce the numbering at the surfaces of the control volumes 1 to 6 corresponding to high- $i$ , low- $i$ , high- $j$ , low- $j$ , high- $k$  and low- $k$ , respectively. We first define the unit surface vector  $(\mathbf{e})^m$  at each surface  $m$  as outwards directed:

$$(\mathbf{e})^1 = \mathbf{e}^1, (\mathbf{e})^2 = -\mathbf{e}_{i-1}^1, (\mathbf{e})^3 = \mathbf{e}^2, (\mathbf{e})^4 = -\mathbf{e}_{j-1}^2, (\mathbf{e})^5 = \mathbf{e}^3, (\mathbf{e})^6 = -\mathbf{e}_{k-1}^3. \quad (13.13-18)$$

With this we have a short notation of the corresponding discretized concentration conservation equation

$$\alpha_i \rho_i C_{il} \gamma_v - \alpha_{ia} \rho_{ia} C_{ila} \gamma_{va} + \Delta \tau \sum_{m=1}^6 \beta_m (\mathbf{e})^m \cdot \mathbf{G}_{il, m} = \Delta \tau \gamma_v \mu_{il}. \quad (13.19)$$

We immediately recognize that it is effective to compute once the geometry coefficients

$$\begin{aligned} \beta_1 &= \gamma_{\xi} \frac{S_1}{\Delta V}, \beta_2 = \gamma_{\xi, j-1} \frac{S_2}{\Delta V}, \beta_3 = \gamma_{\eta} \frac{S_3}{\Delta V}, \beta_4 = \gamma_{\eta, j-1} \frac{S_4}{\Delta V}, \\ \beta_5 &= \gamma_{\zeta} \frac{S_5}{\Delta V}, \beta_6 = \gamma_{\zeta, k-1} \frac{S_6}{\Delta V}, \end{aligned} \quad (13.20-25)$$

before the process simulation, to store them, and to update only those that change during the computation. Secondly, we see that these coefficients contain exact physical geometry information. Note that for cylindrical coordinate systems we have

$$\beta_1 = \frac{r_h^\kappa \gamma_r}{r^\kappa \Delta r}, \beta_2 = \frac{(r_h^\kappa \gamma_r)_{i-1}}{r^\kappa \Delta r}, \beta_3 = \frac{\gamma_\theta}{r^\kappa \Delta \theta}, \beta_4 = \frac{\gamma_{\theta, j-1}}{r^\kappa \Delta \theta},$$

$$\beta_5 = \frac{\gamma_z}{\Delta z}, \beta_6 = \frac{\gamma_{z, k-1}}{\Delta z}, \quad (13.26-31)$$

and for Cartesian setting  $\kappa = 0$  and  $r = x$ ,  $\theta = y$ ,

$$\beta_1 = \frac{\gamma_x}{\Delta x}, \beta_2 = \frac{\gamma_{x, i-1}}{\Delta x}, \beta_3 = \frac{\gamma_y}{\Delta y}, \beta_4 = \frac{\gamma_{y, j-1}}{\Delta y}, \beta_5 = \frac{\gamma_z}{\Delta z}, \beta_6 = \frac{\gamma_{z, k-1}}{\Delta z}, \quad (13.32-39)$$

compare with Section 11.4. Setting  $C_{il} = 1$  in Eq. (13.19) we obtain the discretized mass conservation equation of each velocity field

$$\alpha_l \rho_l \gamma_v - \alpha_{la} \rho_{la} \gamma_{va} + \Delta \tau \sum_{m=1}^6 \beta_m (\mathbf{e})^m \cdot \mathbf{G}_{l,m} = \Delta \tau \gamma_v \mu_l. \quad (13.40)$$

Next we derive the useful non-conservative form of the concentration equations. We multiply Eq. (13.40) by the concentration at the new time plane and subtract the resulting equation from Eq. (13.19). Then the field mass source term is split in two non-negative parts  $\mu_l = \mu_l^+ + \mu_l^-$ . The result is

$$\alpha_{la} \rho_{la} (C_{il} - C_{ila}) \gamma_{va} + \Delta \tau \sum_{m=1}^6 \beta_m (\mathbf{e})^m \cdot (\mathbf{G}_{il,m} - C_{il} \mathbf{G}_{l,m}) + \Delta \tau \gamma_v \mu_l^+ C_{il} = \Delta \tau \gamma_v DC_{il}. \quad (13.41)$$

where  $DC_{il} = \mu_{il} + \mu_l^- C_{il}$ . Note that

$$\mathbf{G}_{il,m} - C_{il} \mathbf{G}_{l,m}$$

$$= \left[ \alpha_l \rho_l (\mathbf{V}_l - \mathbf{V}_{cs}) \right]_m (C_{il,m} - C_{il}) - \left[ \alpha_l \rho_l D_{il}^* \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_m. \quad (13.42)$$

Up to this point of the derivation we did not made any assumption about the computation of the properties at the surfaces of the control volume.

### 13.4.2 The donor-cell concept

The concept of the so called *donor-cell* for the convective terms is now introduced. Flow of given scalars takes the values of the scalars at the cell where the flow is coming from. Mathematically it is expressed as follows. First we define velocity normal to the cell surfaces and outwards directed

$$\mathbf{V}_{l,m}^n = (\mathbf{e})^m \cdot (\mathbf{V}_l - \mathbf{V}_{cs})_m, \quad (13.43)$$

then the switch functions (to store them use *signet integers* in computer codes, it saves memory)

$$\xi_{lm+} = \frac{1}{2} \left[ 1 + \text{sign} \left( V_{lm}^n \right) \right], \quad (13.44)$$

$$\xi_{lm-} = 1 - \xi_{lm+}, \quad (13.45)$$

and then the  $b$  coefficients as follows

$$b_{lm+} = \beta_m \xi_{lm+} V_{lm}^n \geq 0, \quad (13.46)$$

$$b_{lm-} = -\beta_m \xi_{lm-} V_{lm}^n \geq 0. \quad (13.47)$$

If the normal outwards directed velocity is positive the  $+b$  coefficients are unity and the  $-b$  coefficients are zero and vice versa. In this case the normal mass flow rate at the surfaces is

$$(\mathbf{e})^m \cdot \mathbf{G}_{l,m} = \left[ \alpha_l \rho_l \mathbf{e} \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \right]_m = (\xi_{lm+} \alpha_l \rho_l + \xi_{lm-} \alpha_{l,m} \rho_{l,m}) V_{lm}^n. \quad (13.48)$$

Using the above result the mass conservation equations for each field result is

$$\alpha_l \rho_l \gamma_v - \alpha_{la} \rho_{la} \gamma_{va} + \Delta \tau \sum_{m=1}^6 (b_{lm+} \alpha_l \rho_l - B_{lm-}) - \Delta \tau \gamma_v \mu_l = 0. \quad (13.49)$$

In the donor-cell concept the term

$$B_{lm-} = b_{lm-} \alpha_{l,m} \rho_{l,m} \quad (13.50)$$

plays an important role.  $B_{lm-}$  is in fact the mass flow entering the cell from the face  $m$  divided by the volume of the cell. Once computed for the mass conservation equation it is stored and used subsequently in all other conservation equations.

At this point the method used for computation of the field volumetric fractions by iteration using the point *Gauss-Seidel* method for known velocity vectors and thermal properties will be described.

Consider the field variables  $\alpha_l \rho_l$  in the convective terms associated with the output flow in the new time plane, and  $\alpha_{lm} \rho_{lm}$  in the neighboring cells  $m$  as the best available guesses for the new time plane. Solving Eq. (13.49) with respect to  $\alpha_l \rho_l$  gives

$$\bar{\alpha}_l \bar{\rho}_l = \left[ \gamma_{va} \left( \mu_l + \frac{\alpha_{la} \rho_{la}}{\Delta \tau} \right) + \sum_{m=1}^6 B_{lm-} \right] / \left( \frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm+} \right). \quad (13.51)$$

Here

$$\frac{\gamma_v}{\Delta \tau} + \sum_{m=1}^6 b_{lm+} > 0 \quad (13.52)$$

is ensured because  $\gamma_v$  is not allowed to be zero. For a field that is just originating we have

$$\bar{\alpha}_l = \frac{\Delta\tau}{\bar{\rho}_l} \frac{\gamma_{va}\mu_l + \sum_{m=1}^6 B_{lm-}}{\gamma_v + \Delta\tau \sum_{m=1}^6 b_{lm+}}. \tag{13.53}$$

Obviously the field can originate due to convection,  $\sum_{m=1}^6 B_{lm-} > 0$ , or due to an in-cell mass source,  $\mu_l > 0$ , or due to the simultaneous appearance of both phenomena. In case of origination caused by in-cell mass source terms it is important to define the initial density,  $\bar{\rho}_l$ , in order to compute  $\bar{\alpha}_l = \Delta\tau\mu_l / \bar{\rho}_l$ .

The best mass conservation in such procedures is ensured if the following sequence is used for computation of the volume fractions:

$$\alpha_2 = \bar{\alpha}_2 \bar{\rho}_2 / \rho_2, \alpha_3 = \bar{\alpha}_3 \bar{\rho}_3 / \rho_3, \alpha_1 = 1 - \alpha_2 - \alpha_3. \tag{13.54-56}$$

For designing the pressure-velocity coupling the form of the discretized mass conservation is required that explicitly contains the normal velocities,

$$(\alpha_l \rho_l \gamma_v - \alpha_{la} \rho_{la} \gamma_{va}) / \Delta\tau + \sum_{m=1}^6 \beta_m (\xi_{lm+} \alpha_l \rho_l + \xi_{lm-} \alpha_{lm} \rho_{lm}) V_{lm}^n - \gamma_v \mu_l = 0. \tag{13.57}$$

The mass flow rate of the species  $i$  inside the field  $l$  at the cell surface  $m$  is then

$$\begin{aligned} (\mathbf{e})^m \cdot \mathbf{G}_{il,m} &= \left[ \alpha_l \rho_l C_{il} \mathbf{e} \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \right]_m - \left[ \alpha_l \rho_l D_{il}^* \mathbf{e} \cdot \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_m \\ &= (\xi_{lm+} \alpha_l \rho_l C_{il} + \xi_{lm-} \alpha_{l,m} \rho_{l,m} C_{il,m}) V_{lm}^n - \left[ \alpha_l \rho_l D_{il}^* \mathbf{e} \cdot \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_m, \end{aligned} \tag{13.58}$$

and consequently

$$\begin{aligned} &(\mathbf{e})^m \cdot (\mathbf{G}_{il,m} - C_{il} \mathbf{G}_{l,m}) \\ &= \xi_{lm-} \alpha_{l,m} \rho_{l,m} V_{lm}^n (C_{il,m} - C_{il}) - \left[ \alpha_l \rho_l D_{il}^* \mathbf{e} \cdot \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_m. \end{aligned} \tag{13.59}$$

Thus Eq. (13.41) takes the intermediate form

$$\begin{aligned}
 & \alpha_{la} \rho_{la} (C_{il} - C_{ila}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 \left\{ \begin{aligned} & B_{lm-} (C_{il,m} - C_{il}) \\ & + \beta_m \left[ \alpha_l \rho_l D_{il}^* \mathbf{e} \cdot \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right] \right\} \\ & = \Delta \tau \gamma_v (\mu_{il} - C_{il} \mu_l). \end{aligned} \right. \quad (13.60)
 \end{aligned}$$

### 13.4.3 Two methods for computing the finite difference approximations of the contravariant vectors at the cell center

The contravariant vectors for each particular surface can be expressed by

$$\mathbf{a}^1 = \frac{\mathbf{S}^1}{\Delta V} \partial \xi, \quad \mathbf{a}^2 = \frac{\mathbf{S}^2}{\Delta V} \partial \eta, \quad \mathbf{a}^3 = \frac{\mathbf{S}^3}{\Delta V} \partial \zeta. \quad (13.61-63)$$

Note that the contravariant vectors normal to each control volume surface are conveniently computed for **equidistant discretization in the computational space** as follows

$$(\mathbf{a}^1)_1 = \frac{\mathbf{S}^1}{\Delta V_1} = \frac{(\mathbf{S})_1}{\Delta V_1} = \frac{S_1}{\Delta V_1} (\mathbf{e})^1, \quad (\mathbf{a}^1)_2 = \frac{\mathbf{S}^1}{\Delta V_2} = -\frac{(\mathbf{S})_2}{\Delta V_2} = -\frac{S_2}{\Delta V_2} (\mathbf{e})^2, \quad (13.64-65)$$

$$(\mathbf{a}^2)_3 = \frac{\mathbf{S}^2}{\Delta V_3} = \frac{(\mathbf{S})_3}{\Delta V_3} = \frac{S_3}{\Delta V_3} (\mathbf{e})^3, \quad (\mathbf{a}^2)_4 = \frac{\mathbf{S}^2}{\Delta V_4} = -\frac{(\mathbf{S})_4}{\Delta V_4} = -\frac{S_4}{\Delta V_4} (\mathbf{e})^4, \quad (13.66-67)$$

$$(\mathbf{a}^3)_5 = \frac{\mathbf{S}^3}{\Delta V_5} = \frac{(\mathbf{S})_5}{\Delta V_5} = \frac{S_5}{\Delta V_5} (\mathbf{e})^5, \quad (\mathbf{a}^3)_6 = \frac{\mathbf{S}^3}{\Delta V_6} = -\frac{(\mathbf{S})_6}{\Delta V_6} = -\frac{S_6}{\Delta V_6} (\mathbf{e})^6, \quad (13.68-69)$$

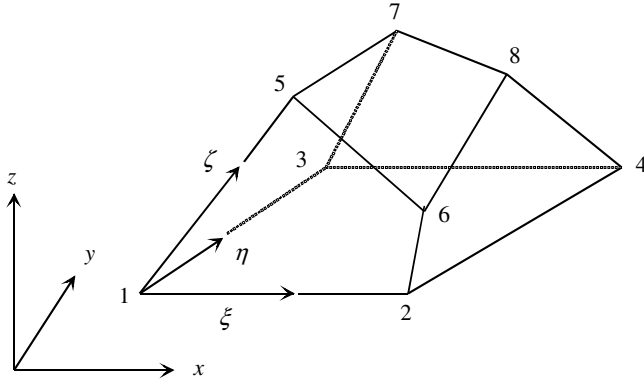
where the volume associated with these vectors is

$$\overline{\Delta V}_m = \frac{1}{2} (\Delta V + \Delta V_m). \quad (13.70)$$

*The finite volume method:* There are two practicable methods for approximation of the contravariant vectors at the cell center. The first one makes use of the already computed normal interface vectors in the following way:

$$\mathbf{a}_c^1 = \frac{1}{2} [(\mathbf{a}^1)_1 + (\mathbf{a}^1)_2], \quad \mathbf{a}_c^2 = \frac{1}{2} [(\mathbf{a}^2)_3 + (\mathbf{a}^2)_4], \quad \mathbf{a}_c^3 = \frac{1}{2} [(\mathbf{a}^3)_5 + (\mathbf{a}^3)_6] \quad (13.71-73)$$





**Fig. 13.1** Numbering of the vertices

*The finite difference method:* The second method uses the coordinates of the vertices of the control volume directly, Fig. 13.1. First we define the position at the cell surfaces that will be used to compute the transformation metrics as follows:

$$\mathbf{r}_{s1} = \frac{1}{4}(\mathbf{r}_2 + \mathbf{r}_4 + \mathbf{r}_8 + \mathbf{r}_6), \quad \mathbf{r}_{s2} = \frac{1}{4}(\mathbf{r}_1 + \mathbf{r}_3 + \mathbf{r}_7 + \mathbf{r}_5), \quad (13.74-75)$$

$$\mathbf{r}_{s3} = \frac{1}{4}(\mathbf{r}_3 + \mathbf{r}_4 + \mathbf{r}_8 + \mathbf{r}_7), \quad \mathbf{r}_{s4} = \frac{1}{4}(\mathbf{r}_1 + \mathbf{r}_2 + \mathbf{r}_6 + \mathbf{r}_5), \quad (13.76-77)$$

$$\mathbf{r}_{s5} = \frac{1}{4}(\mathbf{r}_5 + \mathbf{r}_6 + \mathbf{r}_8 + \mathbf{r}_7), \quad \mathbf{r}_{s6} = \frac{1}{4}(\mathbf{r}_1 + \mathbf{r}_2 + \mathbf{r}_4 + \mathbf{r}_3). \quad (13.78-79)$$

Then we compute the *inverse metrics* of the coordinate transformation for **equidistant discretization in the transformed space**

$$\begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix} = \begin{pmatrix} x_{s1} - x_{s2} & x_{s3} - x_{s4} & x_{s5} - x_{s6} \\ y_{s1} - y_{s2} & y_{s3} - y_{s4} & y_{s5} - y_{s6} \\ z_{s1} - z_{s2} & z_{s3} - z_{s4} & z_{s5} - z_{s6} \end{pmatrix}. \quad (13.80)$$

Then we compute the *Jacobian determinant* and the *metrics* of the coordinate transformation for equidistant discretization in the transformed space.

As already mentioned all this information belongs to the center of the cell. However, the off-diagonal geometry information is required at the cell surfaces. For both cases we use the two corresponding neighbor vectors to compute the contravariant vectors at the cell surfaces as follows

$$\begin{aligned}
(\mathbf{a}^2)_1 &= \frac{1}{2}(\mathbf{a}_c^2 + \mathbf{a}_{c,i+1}^2), \quad (\mathbf{a}^3)_1 = \frac{1}{2}(\mathbf{a}_c^3 + \mathbf{a}_{c,i+1}^3), \\
(\mathbf{a}^2)_2 &= \frac{1}{2}(\mathbf{a}_c^2 + \mathbf{a}_{c,i-1}^2), \quad (\mathbf{a}^3)_2 = \frac{1}{2}(\mathbf{a}_c^3 + \mathbf{a}_{c,i-1}^3), \\
(\mathbf{a}^1)_3 &= \frac{1}{2}(\mathbf{a}_c^1 + \mathbf{a}_{c,j+1}^1), \quad (\mathbf{a}^3)_3 = \frac{1}{2}(\mathbf{a}_c^3 + \mathbf{a}_{c,j+1}^3), \\
(\mathbf{a}^1)_4 &= \frac{1}{2}(\mathbf{a}_c^1 + \mathbf{a}_{c,j-1}^1), \quad (\mathbf{a}^3)_4 = \frac{1}{2}(\mathbf{a}_c^3 + \mathbf{a}_{c,j-1}^3), \\
(\mathbf{a}^1)_5 &= \frac{1}{2}(\mathbf{a}_c^1 + \mathbf{a}_{c,k+1}^1), \quad (\mathbf{a}^2)_5 = \frac{1}{2}(\mathbf{a}_c^2 + \mathbf{a}_{c,k+1}^2), \\
(\mathbf{a}^1)_6 &= \frac{1}{2}(\mathbf{a}_c^1 + \mathbf{a}_{c,k-1}^1), \quad (\mathbf{a}^2)_6 = \frac{1}{2}(\mathbf{a}_c^2 + \mathbf{a}_{c,k-1}^2).
\end{aligned} \tag{13.81-94}$$

### 13.4.4 Discretization of the diffusion terms

#### 13.4.4.1 General

Our next task is to find appropriate finite difference approximation for the six diffusion terms

$$\left[ \alpha_l \rho_l D_{il}^* \mathbf{e} \cdot \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_m,$$

The geometric properties computed by using the control volume approach in the previous section are used to transform the diagonal diffusion terms as direct finite differences

$$\begin{aligned}
& \sum_{m=1}^6 \beta_m \left[ \alpha_l \rho_l D_{il}^* \mathbf{e} \cdot \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_m \\
&= \beta_1 \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_1 \left[ C_{il,i+1} - C_{il} + \frac{\overline{\Delta V}_1}{S_1} \left( \mathbf{e} \cdot \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{e} \cdot \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_1 \\
&+ \beta_2 \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_2 \left[ C_{il,i-1} - C_{il} + \frac{\overline{\Delta V}_2}{S_2} \left( \mathbf{e} \cdot \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{e} \cdot \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_2 \\
&+ \beta_3 \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_3 \left[ C_{il,j+1} - C_{il} + \frac{\overline{\Delta V}_3}{S_3} \left( \mathbf{e} \cdot \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{e} \cdot \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_3 \\
&+ \beta_4 \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_4 \left[ C_{il,j-1} - C_{il} + \frac{\overline{\Delta V}_4}{S_4} \left( \mathbf{e} \cdot \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{e} \cdot \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_4
\end{aligned}$$

$$\begin{aligned}
& +\beta_5 \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_5 \left[ C_{il,k+1} - C_{il} + \frac{\overline{\Delta V}_5}{S_5} \left( \mathbf{e} \cdot \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{e} \cdot \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} \right)_5 \right] \\
& +\beta_6 \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_6 \left[ C_{il,k-1} - C_{il} + \frac{\overline{\Delta V}_6}{S_6} \left( \mathbf{e} \cdot \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{e} \cdot \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} \right)_6 \right]. \quad (13.95)
\end{aligned}$$

A natural averaging of the diffusion coefficients is then the harmonic averaging as given in Appendix 12.1

$$\frac{D_{il,m}^C}{\Delta L_{h,m}} = \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_m S_m = S_m \frac{2(\alpha_l \rho_l D_{il}^*)(\alpha_l \rho_l D_{il}^*)_m}{\Delta V_m (\alpha_l \rho_l D_{il}^*) + \Delta V (\alpha_l \rho_l D_{il}^*)_m}, \quad (13.96)$$

where in the right hand side  $m = 1, 2, 3, 4, 5, 6$  is equivalent to  $i + 1, i - 1, j + 1, j - 1, k + 1, k - 1$ , respectively regarding the properties inside a control volumes. It guaranties that if the field in one of the neighboring cells is missing the diffusion coefficient is zero.

#### 13.4.4.2 Orthogonal coordinate systems

In the case of orthogonal coordinate systems we see that:

- the off-diagonal diffusion terms are equal to zero,
- the finite volume approximations of the diagonal terms are obtained without the need to know anything about the contravariant vectors.

This illustrates the advantage of using orthogonal coordinate systems. This is valid for any diffusion terms in the conservation equations, e.g. the thermal heat diffusion terms in the energy conservation equations, the viscous diffusion terms in the momentum equations etc.

#### 13.4.4.3 Off-diagonal diffusion terms in the general case

The geometric coefficients of the off-diagonal diffusion terms can then be computed as follows

$$d_{12} = (\mathbf{e})^1 \cdot (\mathbf{a}^2)_1 = (e)^{11} (a^{21})_1 + (e)^{12} (a^{22})_1 + (e)^{13} (a^{23})_1, \quad (13.97)$$

$$d_{13} = (\mathbf{e})^1 \cdot (\mathbf{a}^3)_1 = (e)^{11} (a^{31})_1 + (e)^{12} (a^{32})_1 + (e)^{13} (a^{33})_1, \quad (13.98)$$

$$d_{22} = (\mathbf{e})^2 \cdot (\mathbf{a}^2)_2 = (e)^{21} (a^{21})_2 + (e)^{22} (a^{22})_2 + (e)^{23} (a^{23})_2 = -(d_{12})_{i-1}, \quad (13.99)$$

$$d_{23} = (\mathbf{e})^2 \cdot (\mathbf{a}^3)_2 = (e)^{21} (a^{31})_2 + (e)^{22} (a^{32})_2 + (e)^{23} (a^{33})_2 = -(d_{13})_{i-1}, \quad (13.100)$$

$$d_{31} = (\mathbf{e})^3 \cdot (\mathbf{a}^1)_3 = (e)^{31} (a^{11})_3 + (e)^{32} (a^{12})_3 + (e)^{33} (a^{13})_3, \quad (13.101)$$

$$d_{33} = (\mathbf{e})^3 \cdot (\mathbf{a}^3)_3 = (e)^{31} (a^{31})_3 + (e)^{32} (a^{32})_3 + (e)^{33} (a^{33})_3, \quad (13.102)$$

$$d_{41} = (\mathbf{e})^4 \cdot (\mathbf{a}^1)_4 = (e)^{41} (a^{11})_4 + (e)^{42} (a^{12})_4 + (e)^{43} (a^{13})_4 = -(d_{31})_{j-1}, \quad (13.103)$$

$$d_{43} = (\mathbf{e})^4 \cdot (\mathbf{a}^3)_4 = (e)^{41} (a^{31})_4 + (e)^{42} (a^{32})_4 + (e)^{43} (a^{33})_4 = -(d_{33})_{j-1}, \quad (13.104)$$

$$d_{51} = (\mathbf{e})^5 \cdot (\mathbf{a}^1)_5 = (e)^{51} (a^{11})_5 + (e)^{52} (a^{12})_5 + (e)^{53} (a^{13})_5, \quad (13.105)$$

$$d_{52} = (\mathbf{e})^5 \cdot (\mathbf{a}^2)_5 = (e)^{51} (a^{21})_5 + (e)^{52} (a^{22})_5 + (e)^{53} (a^{23})_5, \quad (13.106)$$

$$d_{61} = (\mathbf{e})^6 \cdot (\mathbf{a}^1)_6 = (e)^{61} (a^{11})_6 + (e)^{62} (a^{12})_6 + (e)^{63} (a^{13})_6 = -(d_{51})_{k-1}, \quad (13.107)$$

$$d_{62} = (\mathbf{e})^6 \cdot (\mathbf{a}^2)_6 = (e)^{61} (a^{21})_6 + (e)^{62} (a^{22})_6 + (e)^{63} (a^{23})_6 = -(d_{52})_{k-1}. \quad (13.108)$$

With this notation the diffusion term takes the form

$$\begin{aligned} & \sum_{m=1}^6 \beta_m \left[ \alpha_l \rho_l D_{il}^* \mathbf{e} \cdot \left( \mathbf{a}^1 \frac{\partial C_{il}}{\partial \xi} + \mathbf{a}^2 \frac{\partial C_{il}}{\partial \eta} + \mathbf{a}^3 \frac{\partial C_{il}}{\partial \zeta} \right) \right]_{\perp m} \\ &= \sum_{m=1}^6 \beta_m \frac{D_{il,m}^C}{\Delta L_{h,m}} \left( C_{il,m} - C_{il} + \frac{\overline{\Delta V}_m}{S_m} DI - C_{il,m} \right) \end{aligned} \quad (13.109)$$

where

$$DI - C_{il,1} = d_{12} \left. \frac{\partial C_{il}}{\partial \eta} \right|_1 + d_{13} \left. \frac{\partial C_{il}}{\partial \zeta} \right|_1, \quad DI - C_{il,2} = d_{22} \left. \frac{\partial C_{il}}{\partial \eta} \right|_2 + d_{23} \left. \frac{\partial C_{il}}{\partial \zeta} \right|_2, \quad (13.110-111)$$

$$DI - C_{il,3} = d_{31} \left. \frac{\partial C_{il}}{\partial \xi} \right|_3 + d_{33} \left. \frac{\partial C_{il}}{\partial \zeta} \right|_3, \quad DI - C_{il,4} = d_{41} \left. \frac{\partial C_{il}}{\partial \xi} \right|_4 + d_{43} \left. \frac{\partial C_{il}}{\partial \zeta} \right|_4, \quad (13.112-113)$$

$$DI - C_{il,5} = d_{51} \left. \frac{\partial C_{il}}{\partial \xi} \right|_5 + d_{52} \left. \frac{\partial C_{il}}{\partial \eta} \right|_5, \quad DI - C_{il,6} = d_{61} \left. \frac{\partial C_{il}}{\partial \xi} \right|_6 + d_{62} \left. \frac{\partial C_{il}}{\partial \eta} \right|_6. \quad (13.114-115)$$

The twelve concentration derivatives are computed as follows

$$\left. \frac{\partial C_{il}}{\partial \eta} \right|_1 = \frac{1}{4} (C_{il,j+1} + C_{il,i+1,j+1} - C_{il,j-1} - C_{il,i+1,j-1}), \quad (13.116)$$

$$\left. \frac{\partial C_{il}}{\partial \zeta} \right|_1 = \frac{1}{4} (C_{il,k+1} + C_{il,i+1,k+1} - C_{il,k-1} - C_{il,i+1,k-1}), \quad (13.117)$$

$$\left. \frac{\partial C_{il}}{\partial \eta} \right|_2 = \frac{1}{4} (C_{il,i-1,j+1} + C_{il,j+1} - C_{il,i-1,j-1} - C_{il,j-1}), \quad (13.118)$$

$$\left. \frac{\partial C_{il}}{\partial \zeta} \right|_2 = \frac{1}{4} (C_{il,k+1} + C_{il,i-1,k+1} - C_{il,k-1} - C_{il,i-1,k-1}), \quad (13.119)$$

$$\left. \frac{\partial C_{il}}{\partial \xi} \right|_3 = \frac{1}{4} (C_{il,i+1} + C_{il,i+1,j+1} - C_{il,i-1} - C_{il,i-1,j+1}), \quad (13.120)$$

$$\left. \frac{\partial C_{il}}{\partial \zeta} \right|_3 = \frac{1}{4} (C_{il,k+1} + C_{il,j+1,k+1} - C_{il,k-1} - C_{il,j+1,k-1}), \quad (13.121)$$

$$\left. \frac{\partial C_{il}}{\partial \xi} \right|_4 = \frac{1}{4} (C_{il,i+1} + C_{il,i+1,j-1} - C_{il,i-1} - C_{il,i-1,j-1}), \quad (13.122)$$

$$\left. \frac{\partial C_{il}}{\partial \zeta} \right|_4 = \frac{1}{4} (C_{il,k+1} + C_{il,j-1,k+1} - C_{il,k-1} - C_{il,j-1,k-1}), \quad (13.123)$$

$$\left. \frac{\partial C_{il}}{\partial \xi} \right|_5 = \frac{1}{4} (C_{il,i+1,k+1} + C_{il,i+1} - C_{il,i-1,k+1} - C_{il,i-1}), \quad (13.124)$$

$$\left. \frac{\partial C_{il}}{\partial \eta} \right|_5 = \frac{1}{4} (C_{il,j+1} + C_{il,j+1,k+1} - C_{il,j-1} - C_{il,j-1,k+1}), \quad (13.125)$$

$$\left. \frac{\partial C_{il}}{\partial \xi} \right|_6 = \frac{1}{4} (C_{il,i+1} + C_{il,i+1,k-1} - C_{il,i-1} - C_{il,i-1,k-1}), \quad (13.126)$$

$$\left. \frac{\partial C_{il}}{\partial \eta} \right|_6 = \frac{1}{4} (C_{il,j+1} + C_{il,j+1,k-1} - C_{il,j-1} - C_{il,j-1,k-1}). \quad (13.127)$$

#### 13.4.4.4 Final form of the finite volume concentration equation

Thus the final form of the discretized concentration equation (13.1) is

$$\begin{aligned} & \alpha_{ia} \rho_{ia} (C_{il} - C_{ila}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 \left( B_{m-} + \beta_m \frac{D_{il,m}^C}{\Delta L_{h,m}} \right) (C_{il,m} - C_{il}) + \Delta \tau \gamma_v \mu_i^+ C_{il} \\ & = \Delta \tau \gamma_v D C_{il} + \Delta \tau \sum_{m=1}^6 \beta_m \frac{D_{il,m}^C}{\Delta L_{h,m}} \frac{\overline{\Delta V}_m}{S_m} D I_{-} C_{il,m}, \end{aligned} \quad (13.128)$$

Solving with respect to the unknown concentration we obtain

$$C_{il} = \frac{\alpha_{la} \rho_{la} \gamma_{va} C_{ila} + \Delta \tau \left\{ \gamma_v DC_{il} + \sum_{m=1}^6 \left[ B_{lm^-} C_{il,m} + \beta_m \frac{D_{il,m}^C}{\Delta L_{h,m}} \left( C_{il,m} + \frac{\overline{\Delta V}_m}{S_m} DI - C_{il,m} \right) \right] \right\}}{\alpha_{la} \gamma_{va} \rho_{la} + \Delta \tau \left[ \gamma_v \mu_l^+ + \sum_{m=1}^6 \left( B_{lm^-} + \beta_m \frac{D_{il,m}^C}{\Delta L_{h,m}} \right) \right]} \quad (13.129)$$

For the case of a just originating velocity field,  $\alpha_{la} = 0$  and

$$\gamma_v \mu_l^+ + \sum_{m=1}^6 \left( b_{lm^-} \alpha_{l,m} \rho_{l,m} + \beta_m \frac{D_{il,m}^C}{\Delta L_{h,m}} \right) > 0, \quad (13.130)$$

we have

$$C_{il} = \frac{\gamma_v DC_{il} + \sum_{m=1}^6 \gamma_v DC_{il} + \sum_{m=1}^6 \left[ B_{lm^-} C_{il,m} + \beta_m \frac{D_{il,m}^C}{\Delta L_{h,m}} \left( C_{il,m} + \frac{\overline{\Delta V}_m}{S_m} DI - C_{il,m} \right) \right]}{\gamma_v \mu_l^+ + \sum_{m=1}^6 \left( B_{lm^-} + \beta_m \frac{D_{il,m}^C}{\Delta L_{h,m}} \right)} \quad (13.131)$$

## 13.5 Discretization of the entropy equation

The entropy equation (10.62) is discretized following the procedure already described in the previous section. The result is

$$\alpha_{la} \rho_{la} (s_l - s_{la}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 B_{lm^-} (s_{l,m} - s_l) + \Delta \tau \gamma_v \mu_l^+ s_l = \Delta \tau \gamma_v Ds_l^*, \quad (13.132)$$

where

$$Ds_l^* = Ds_l + \frac{1}{\gamma_v} \sum_{m=1}^6 \beta_m \left[ \frac{1}{T_l} \frac{D_{l,m}^T}{\Delta L_{h,m}} \left( T_{l,m} - T_l + \frac{\overline{\Delta V}_m}{S_m} DI - T_{l,m} \right) + \frac{D_{l,m}^{sC}}{\Delta L_{h,m}} \sum_{i=2}^{i_{\max}} \left( C_{il,m} - C_{il} + \frac{\overline{\Delta V}_m}{S_m} DI - C_{il} \right) \right] \quad (13.133)$$

The term  $DI - T_{l,m}$  is computed by replacing the concentrations in Eqs. (11.110-127) with the corresponding temperatures. The computation of the harmonic averaged thermal conductivity coefficients is given in Appendix 12.1. Solving with respect to the unknown specific entropy we obtain

$$s_l = \frac{\alpha_{la} \rho_{la} \gamma_{va} s_{la} + \Delta \tau \left( \gamma_v D s_l^* + \sum_{m=1}^6 B_{lm-} s_{lm} \right)}{\alpha_{la} \rho_{la} \gamma_{va} + \Delta \tau \left( \sum_{m=1}^6 B_{lm-} + \gamma_v \mu_l^+ \right)}. \quad (13.134)$$

For the case of a just originating velocity field,  $\alpha_{la} = 0$  and  $\sum_{m=1}^6 B_{lm-} + \gamma_v \mu_l^+ > 0$ , we have

$$s_l = \frac{\gamma_v D s_l^* + \sum_{m=1}^6 B_{lm-} s_{lm}}{\sum_{m=1}^6 B_{lm-} + \gamma_v \mu_l^+}. \quad (13.135)$$

### 13.6 Discretization of the temperature equation

The temperature equation (10.65) is discretized following the procedure already described in the previous section. The result is

$$\begin{aligned} & \alpha_{la} \rho_{la} c_{pla} (T_l - T_{la}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 \left( B_{lm-} c_{pl,m} + \beta_m \frac{D_{l,m}^T}{\Delta L_{h,m}} \right) (T_{l,m} - T_l) \\ & \left[ 1 - \rho_l \left( \frac{\partial h_l}{\partial p} \right)_{T_l, all\_C's} \right] \left[ \alpha_{la} (p - p_a) \gamma_{va} - \Delta \tau \sum_{m=1}^6 \frac{B_{lm-}}{\rho_{lm}} (p_m - p) \right] \\ & = \Delta \tau \gamma_v \left[ DT_l^N - T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} (\mu_{il} - C_{il} \mu_l) \right] \\ & - \Delta \tau \sum_{m=1}^6 \beta_m \left[ \frac{D_{l,m}^C}{\Delta L_{h,m}} T_l \sum_{i=2}^{i_{\max}} \Delta s_{il}^{np} \left( C_{il,m} - C_{il} + \frac{\overline{\Delta V}_m}{S_m} DI - C_{il} \right) - \frac{D_{l,m}^T}{\Delta L_{h,m}} \frac{\overline{\Delta V}_m}{S_m} DI - T_l \right]. \end{aligned} \quad (13.136)$$

## 13.7 Discretization of the particle number density equation

The particle number density equation (10.66) is discretized following the procedure already described in the previous section. The result is

$$\begin{aligned} n_l \gamma_v - n_{la} \gamma_{va} + \Delta \tau \sum_{m=1}^6 \beta_m \left[ (\mathbf{e})^m \cdot (\mathbf{V}_{l,m} - \mathbf{V}_{cs}) n_{l,m} - \frac{D_{l,m}^n}{\Delta L_{h,m}} (n_{l,m} - n_l) \right] \\ = \Delta \tau \gamma_v (\dot{n}_{l,kin} - \dot{n}_{l,coal} + \dot{n}_{l,sp}) + \Delta \tau \sum_{m=1}^6 \beta_m \frac{D_{l,m}^n}{\Delta L_{h,m}} \frac{\overline{\Delta V}_m}{S_m} DI_- n_{l,m}. \end{aligned} \quad (13.137)$$

The turbulent diffusion coefficient is again a result of harmonic volume averaging – Appendix 12.1.

$$\frac{D_{l,m}^n}{\Delta L_{h,m}} = \left( \frac{\frac{V_l'}{Sc^t}}{\Delta V} \right)_m S_m = S_m \frac{2 \left( \frac{V_l'}{Sc^t} \right) \left( \frac{V_l'}{Sc^t} \right)_m}{\Delta V_m \left( \frac{V_l'}{Sc^t} \right) + \Delta V \left( \frac{V_l'}{Sc^t} \right)_m}. \quad (13.138)$$

## 13.8 Discretization of the x momentum equation

The  $x$  momentum equation (10.68) is discretized as already discussed in the previous Section. The result is

$$\begin{aligned} \alpha_{la} \rho_{la} (u_l - u_{la}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 \left\{ B_{lm} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m1} (e)^{m1} \right] \right\} (u_{l,m} - u_l) \\ + \Delta \tau \alpha_l \left( a^{11} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{21} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{31} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \\ - \Delta \tau \gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| (u_m - u_l) + \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| (u_{cs} - u_l) \right] \\ - \Delta \tau \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{l_{max}} \bar{c}_{ml}^L [(v_l - v_m) b_{m,3} - (w_l - w_m) b_{m,2}] \\ - \Delta \tau \gamma_v \bar{c}_{wl}^L [(v_l - v_{cs}) b_{w,3} - (w_l - w_{cs}) b_{w,2}] \\ - \Delta \tau \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm} \left( \frac{\partial \Delta u_{ml}}{\partial \tau} + \bar{v}^1 \frac{\partial \Delta u_{ml}}{\partial \xi} + \bar{v}^2 \frac{\partial \Delta u_{ml}}{\partial \eta} + \bar{v}^3 \frac{\partial \Delta u_{ml}}{\partial \zeta} \right) \end{aligned}$$



$$\begin{aligned}
& -\gamma_v \Delta \tau \bar{c}_{wl}^{vm} \left( \frac{\partial \Delta u_{csl}}{\partial \tau} + \bar{V}^1 \frac{\partial \Delta u_{csl}}{\partial \xi} + \bar{V}^2 \frac{\partial \Delta u_{csl}}{\partial \eta} + \bar{V}^3 \frac{\partial \Delta u_{csl}}{\partial \zeta} \right) \\
& = \Delta \tau \gamma_v \left[ -\alpha_i \rho_i g_x + \sum_{m=1}^{3,w} [\mu_{ml} (u_m - u_l)] + \mu_{lw} (u_{lw} - u_l) \right] \\
& + \Delta \tau \sum_{m=1}^6 \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left( \begin{aligned} & DI_{-} u_{l,m} - \frac{2}{3} (e)^{m1} DI_{-} u_{l,m}^b + DI_{-} vis_{m}^{uT} \\ & + \frac{1}{3} (e)^{m1} [(e)^{m2} (v_{l,m} - v_l) + (e)^{m3} (w_{l,m} - w_l)] \end{aligned} \right). \quad (13.139)
\end{aligned}$$

A natural averaging of the diffusion coefficients is the harmonic averaging – Appendix 12.1

$$\frac{D_{il,m}^v}{\Delta L_{h,m}} = \left( \frac{\alpha_i^e \rho_i v_l^*}{\Delta V} \right)_m S_m = S_m \frac{2(\alpha_i^e \rho_i v_l^*)(\alpha_i^e \rho_i v_l^*)_m}{\Delta V_m (\alpha_i^e \rho_i v_l^*) + \Delta V (\alpha_i^e \rho_i v_l^*)_m}. \quad (13.140)$$

It is valid for all momentum equations. Note how we arrive to the integral form of the pressure term:

$$\begin{aligned}
& \frac{1}{\Delta V} \int_0^{\Delta \tau} \iiint_{\Delta V} \sqrt{g} \alpha_i \left( a^{11} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{21} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{31} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \partial \xi \partial \eta \partial \zeta \partial \tau \\
& = \Delta \tau \alpha_i \left( a^{11} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{21} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{31} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right). \quad (13.141)
\end{aligned}$$

The  $b$  coefficients of in the lift force expressions result from the Cartesian component decomposition:

$$b_{m,1} = a^{12} \frac{\partial w_m}{\partial \xi} - a^{13} \frac{\partial v_m}{\partial \xi} + a^{22} \frac{\partial w_m}{\partial \eta} - a^{23} \frac{\partial v_m}{\partial \eta} + a^{32} \frac{\partial w_m}{\partial \zeta} - a^{33} \frac{\partial v_m}{\partial \zeta}, \quad (13.142)$$

$$b_{m,2} = a^{13} \frac{\partial u_m}{\partial \xi} - a^{11} \frac{\partial w_m}{\partial \xi} + a^{23} \frac{\partial u_m}{\partial \eta} - a^{21} \frac{\partial w_m}{\partial \eta} + a^{33} \frac{\partial u_m}{\partial \zeta} - a^{31} \frac{\partial w_m}{\partial \zeta}, \quad (13.143)$$

$$b_{m,3} = a^{11} \frac{\partial v_m}{\partial \xi} - a^{12} \frac{\partial u_m}{\partial \xi} + a^{21} \frac{\partial v_m}{\partial \eta} - a^{22} \frac{\partial u_m}{\partial \eta} + a^{31} \frac{\partial v_m}{\partial \zeta} - a^{32} \frac{\partial u_m}{\partial \zeta}, \quad (13.144)$$

and

$$b_{w,1} = a^{12} \frac{\partial w_{cs}}{\partial \xi} - a^{13} \frac{\partial v_{cs}}{\partial \xi} + a^{22} \frac{\partial w_{cs}}{\partial \eta} - a^{23} \frac{\partial v_{cs}}{\partial \eta} + a^{32} \frac{\partial w_{cs}}{\partial \zeta} - a^{33} \frac{\partial v_{cs}}{\partial \zeta}, \quad (13.145)$$

$$b_{w,2} = a^{13} \frac{\partial u_{cs}}{\partial \xi} - a^{11} \frac{\partial w_{cs}}{\partial \xi} + a^{23} \frac{\partial u_{cs}}{\partial \eta} - a^{21} \frac{\partial w_{cs}}{\partial \eta} + a^{33} \frac{\partial u_{cs}}{\partial \zeta} - a^{31} \frac{\partial w_{cs}}{\partial \zeta}, \quad (13.146)$$

$$b_{w,3} = a^{11} \frac{\partial v_{cs}}{\partial \xi} - a^{12} \frac{\partial u_{cs}}{\partial \xi} + a^{21} \frac{\partial v_{cs}}{\partial \eta} - a^{22} \frac{\partial u_{cs}}{\partial \eta} + a^{31} \frac{\partial v_{cs}}{\partial \zeta} - a^{32} \frac{\partial u_{cs}}{\partial \zeta}. \quad (13.147)$$

We proceed in a similar way for the other momentum equations.

### 13.9 Discretization of the y momentum equation

The result of the discretization of the y momentum equation (10.73) is

$$\begin{aligned} & \alpha_{la} \rho_{la} (v_l - v_{la}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 \left\{ B_{lm} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m^2} (e)^{m^2} \right] \right\} (v_{l,m} - v_l) \\ & + \Delta \tau \alpha_l \left( a^{12} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{22} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{32} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \\ & - \Delta \tau \gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| (v_m - v_l) + \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| (v_{cs} - v_l) \right] \\ & - \Delta \tau \gamma_v \sum_{\substack{l=1 \\ m \neq l}}^{l_{\max}} \bar{c}_{ml}^L [(w_l - w_m) b_{m,1} - (u_l - u_m) b_{m,3}] \\ & - \Delta \tau \gamma_v \bar{c}_{wl}^L [(w_l - w_{cs}) b_{w,1} - \Delta u_{lw} (u_l - u_{cs}) b_{w,3}] \\ & - \Delta \tau \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm} \left( \frac{\partial \Delta v_{ml}}{\partial \tau} + \bar{V}^1 \frac{\partial \Delta v_{ml}}{\partial \xi} + \bar{V}^2 \frac{\partial \Delta v_{ml}}{\partial \eta} + \bar{V}^3 \frac{\partial \Delta v_{ml}}{\partial \zeta} \right) \\ & - \gamma_v \Delta \tau \bar{c}_{wl}^{vml} \left( \frac{\partial \Delta v_{csl}}{\partial \tau} + \bar{V}^1 \frac{\partial \Delta v_{csl}}{\partial \xi} + \bar{V}^2 \frac{\partial \Delta v_{csl}}{\partial \eta} + \bar{V}^3 \frac{\partial \Delta v_{csl}}{\partial \zeta} \right) \\ & = \Delta \tau \gamma_v \left[ -\alpha_l \rho_l g_y + \sum_{m=1}^{3,w} [\mu_{ml} (v_m - v_l)] + \mu_{lw} (v_{lw} - v_l) \right] \\ & + \Delta \tau \sum_{m=1}^6 \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left( \begin{aligned} & DI_{-v_{l,m}} - \frac{2}{3} (e)^{m^2} DI_{-u_{l,m}^b} + DI_{-vis_{lm}^{vT}} \\ & + \frac{1}{3} (e)^{m^2} [(e)^{m^1} (u_{l,m} - u_l) + (e)^{m^3} (w_{l,m} - w_l)] \end{aligned} \right). \quad (13.148) \end{aligned}$$

### 13.10 Discretization of the $z$ momentum equation

The result of the discretization of the  $z$  momentum equation (10.77) is

$$\begin{aligned}
& \alpha_{ia} \rho_{ia} (w_l - w_{la}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 \left\{ B_{lm^-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m3} (e)^{m3} \right] \right\} (w_{l,m} - w_l) \\
& + \Delta \tau \alpha_l \left( a^{13} \gamma_\xi \frac{\partial p}{\partial \xi} + a^{23} \gamma_\eta \frac{\partial p}{\partial \eta} + a^{33} \gamma_\zeta \frac{\partial p}{\partial \zeta} \right) \\
& - \Delta \tau \gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| (w_m - w_l) + \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| (w_{cs} - w_l) \right] \\
& - \Delta \tau \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{\max} \bar{c}_{ml}^L [(u_l - u_m) b_{m,2} - (v_l - v_m) b_{m,1}] \\
& - \Delta \tau \gamma_v \bar{c}_{wl}^L [(u_l - u_{cs}) b_{w,2} - (v_l - v_{cs}) b_{w,1}] \\
& - \Delta \tau \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^{vm} \left( \frac{\partial \Delta w_{ml}}{\partial \tau} + \bar{v}^1 \frac{\partial \Delta w_{ml}}{\partial \xi} + \bar{v}^2 \frac{\partial \Delta w_{ml}}{\partial \eta} + \bar{v}^3 \frac{\partial \Delta w_{ml}}{\partial \zeta} \right) \\
& - \gamma_v \Delta \tau \bar{c}_{wl}^{vm} \left( \frac{\partial \Delta w_{csl}}{\partial \tau} + \bar{v}^1 \frac{\partial \Delta w_{csl}}{\partial \xi} + \bar{v}^2 \frac{\partial \Delta w_{csl}}{\partial \eta} + \bar{v}^3 \frac{\partial \Delta w_{csl}}{\partial \zeta} \right) \\
& = \Delta \tau \gamma_v \left[ -\alpha_l \rho_l g_z + \sum_{m=1}^{3,w} [\mu_{ml} (w_m - w_l)] + \mu_{lw} (w_{lw} - w_l) \right] \\
& + \Delta \tau \sum_{m=1}^6 \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left( \begin{aligned} & DI - w_{l,m} - \frac{2}{3} (e)^{m3} DI - u_{l,m}^b + DI - vis_{lm}^{wT} \\ & + \frac{1}{3} (e)^{m3} [(e)^{m1} (u_{l,m} - u_l) + (e)^{m2} (v_{l,m} - v_l)] \end{aligned} \right). \quad (13.149)
\end{aligned}$$

### 13.11 Pressure-velocity coupling

*The IVA3 method:* An important target of the numerical methods is to guarantee a strict mass conservation in the sense of the overall mass balance as for the single cell as well for the sum of the cells inside the physical domain of interest. We use the discretized mass conservation equations of each field in a special way to construct the so called pressure-velocity coupling keeping in mind the above requirement. First we note that the difference resulting from the time derivative divided by the new time level density can be rearranged as follows

$$\frac{1}{\rho_l}(\alpha_l \rho_l \gamma_v - \alpha_{la} \rho_{la} \gamma_{va}) = (\alpha_l - \alpha_{la}) \gamma_v + \alpha_{la} (\gamma_v - \gamma_{va}) + \frac{\alpha_{la}}{\rho_l} (\rho_l - \rho_{la}) \gamma_{va} . \quad (13.150)$$

Then, we divide each of the discretized field mass conservation equations by the corresponding new time level density. Having in mind Eq. (13.150) we obtain

$$\begin{aligned} & (\alpha_l - \alpha_{la}) \gamma_v + \frac{\alpha_{la}}{\rho_l} (\rho_l - \rho_{la}) \gamma_{va} + \Delta \tau \frac{1}{\rho_l} \sum_{m=1}^6 \beta_m (\xi_{lm+} \alpha_l \rho_l + \xi_{lm-} \alpha_{lm} \rho_{lm}) V_{lm}^n \\ & = \frac{1}{\rho_l} \Delta \tau \gamma_v \mu_l - \alpha_{la} (\gamma_v - \gamma_{va}) . \end{aligned} \quad (13.151)$$

We sum all of the  $l_{max}$  mass conservation equations. The first term disappears because the sum of all volume fractions is equal to unity. In the resulting equation the temporal density difference is replaced by the linearized form of the equation of state, Eq. (3.173),

$$\rho_l - \rho_{la} = \frac{1}{a_{la}^2} (p - p_a) + \left( \frac{\partial \rho_l}{\partial s_l} \right)_a (s_l - s_{la}) + \sum_{i=2}^{i_{max}} \left( \frac{\partial \rho_l}{\partial C_{i,l}} \right)_a (C_{i,l} - C_{i,la}) . \quad (13.152)$$

The result is

$$(p - p_a) \gamma_{va} \sum_{l=1}^{l_{max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} + \Delta \tau \sum_{l=1}^{l_{max}} \frac{1}{\rho_l} \sum_{m=1}^6 \beta_m (\xi_{lm+} \alpha_l \rho_l + \xi_{lm-} \alpha_{lm} \rho_{lm}) V_{lm}^n = \Delta \tau \sum_{l=1}^{l_{max}} D \alpha_l , \quad (13.153)$$

where

$$\begin{aligned} \Delta \tau D \alpha_l & = \frac{1}{\rho_l} \Delta \tau \gamma_v \mu_l - \alpha_{la} (\gamma_v - \gamma_{va}) \\ & - \gamma_{va} \frac{\alpha_{la}}{\rho_l} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_a (s_l - s_{la}) + \sum_{i=2}^{i_{max}} \left( \frac{\partial \rho_l}{\partial C_{i,l}} \right)_a (C_{i,l} - C_{i,la}) \right] . \end{aligned} \quad (13.154)$$

This equation is equivalent exactly to the sum of the discretized mass conservation equations divided by the corresponding densities. It takes into account the influence of the variation of the density with the time on the pressure change. The spatial variation of the density in the second term is still not resolved. With the next step we will derive a approximated approach to change also the influence of the spatial variation of the density on the pressure change.

Writing the discretized momentum equation in the linearized form

$$V_{lm}^n = dV_{lm}^n - (\mathbf{e})^m \cdot \mathbf{V}_{cs,m} - R V e_{lm} (p_m - p) , \quad (13.155)$$

and replacing we finally obtain the so called pressure equation

$$\begin{aligned}
 & p\gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} - \Delta\tau \sum_{m=1}^6 \beta_m \left[ \sum_{l=1}^{l_{\max}} \left( \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \frac{\rho_{lm}}{\rho_l} \right) R V e l_{lm} \right] (p_m - p) \\
 & = p_a \gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} - \Delta\tau \sum_{l=1}^{l_{\max}} \frac{1}{\rho_l} \sum_{m=1}^6 \beta_m (\xi_{lm+} \alpha_l \rho_l + \xi_{lm-} \alpha_{lm} \rho_{lm}) \left[ dV_{lm}^n - (\mathbf{e})^m \cdot \mathbf{V}_{cs,m} \right] \\
 & + \Delta\tau \sum_{l=1}^{l_{\max}} D\alpha_l . \tag{13.156}
 \end{aligned}$$

Defining the coefficients

$$c_m = -\Delta\tau \beta_m \sum_{l=1}^{l_{\max}} \left( \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \frac{\rho_{lm}}{\rho_l} \right) R V e l_{lm} , \tag{13.157}$$

$$c = p\gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} - \sum_{m=1}^6 c_m , \tag{13.158}$$

$$\begin{aligned}
 d & = p_a \gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} + \Delta\tau \sum_{l=1}^{l_{\max}} D\alpha_l \\
 & - \Delta\tau \sum_{m=1}^6 \beta_m \sum_{l=1}^{l_{\max}} \left( \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \frac{\rho_{lm}}{\rho_l} \right) \left[ dV_{lm}^n - (\mathbf{e})^m \cdot \mathbf{V}_{cs,m} \right] , \tag{13.159}
 \end{aligned}$$

we obtain the pressure equation

$$cp + \sum_{m=1}^6 c_m p_m = d , \tag{13.160}$$

connecting each cell pressure with the pressure of the surrounding cells. We see that the system of algebraic equations has *positive diagonal elements*, is *symmetric* and *strictly diagonally dominant* because

$$|c| > \sum_{m=1}^6 |c_m| . \tag{13.161}$$

These are very important properties.

*The IVA2 method:* The spatial deviation of the density of the surrounding cells from the density of the cell considered can be introduced into Eq. (13.153) as follows

$$\beta_m (\xi_{lm+} \alpha_l \rho_l + \xi_{lm-} \alpha_{lm} \rho_{lm}) = \beta_m (\xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm}) \rho_l + \beta_m \xi_{lm-} \alpha_{lm} (\rho_{lm} - \rho_l) . \tag{13.161}$$

The result is

$$\begin{aligned}
 & (p - p_a) \gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} + \Delta \tau \sum_{l=1}^{l_{\max}} \sum_{m=1}^6 \beta_m (\xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm}) V_{lm}^n \\
 & + \Delta \tau \sum_{l=1}^{l_{\max}} \frac{1}{\rho_l} \sum_{m=1}^6 \beta_m \xi_{lm-} \alpha_{lm} (\rho_{lm} - \rho_l) V_{lm}^n \\
 & = \sum_{l=1}^{l_{\max}} \left[ \begin{array}{l} \frac{1}{\rho_l} \Delta \tau \gamma_v \mu_l - \alpha_{la} (\gamma_v - \gamma_{va}) \\ -\gamma_{va} \frac{\alpha_{la}}{\rho_l} \left[ \left( \frac{\partial \rho_l}{\partial s_l} \right)_a (s_l - s_{la}) + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{i,l}} \right)_a (C_{i,l} - C_{i,la}) \right] \end{array} \right]. \quad (13.163)
 \end{aligned}$$

The spatial density variation can again be expressed as follows, Eq. (3.173),

$$\rho_{lm} - \rho_l = \frac{1}{a_{la}^2} (p_m - p) + \left( \frac{\partial \rho_l}{\partial s_l} \right)_a (s_{lm} - s_l) + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{i,l}} \right)_a (C_{i,lm} - C_{i,l}). \quad (13.164)$$

With this we obtain

$$\begin{aligned}
 & (p - p_a) \gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} + \Delta \tau \sum_{l=1}^{l_{\max}} \sum_{m=1}^6 \beta_m \left[ \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \left( 1 + \frac{p_m - p}{\rho_l a_{la}^2} \right) \right] V_{lm}^n \\
 & = \Delta \tau \sum_{l=1}^{l_{\max}} D \alpha_l, \quad (13.165)
 \end{aligned}$$

where

$$\begin{aligned}
 & \rho_l \Delta \tau D \alpha_l = \Delta \tau \gamma_v \mu_l - \alpha_{la} \rho_l (\gamma_v - \gamma_{va}) - \left( \frac{\partial \rho_l}{\partial s_l} \right)_a \left[ \begin{array}{l} \gamma_{va} \alpha_{la} (s_l - s_{la}) \\ -\Delta \tau \sum_{m=1}^6 b_{lm-} \alpha_{lm} (s_{lm} - s_l) \end{array} \right] \\
 & - \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{i,l}} \right)_a \left[ \gamma_{va} \alpha_{la} (C_{i,l} - C_{i,la}) - \Delta \tau \sum_{m=1}^6 b_{lm-} \alpha_{lm} (C_{i,lm} - C_{i,l}) \right], \quad (13.166)
 \end{aligned}$$

Equation (13.165) is equivalent exactly to the sum of the discretized mass conservation equations. The discretized concentration equation divided by the old time level density is

$$\gamma_{va} \alpha_{la} (C_{il} - C_{ila}) - \Delta \tau \sum_{m=1}^6 b_{lm-} \alpha_{l,m} \frac{\rho_{l,m}}{\rho_{la}} (C_{il,m} - C_{il}) = \frac{\Delta \tau}{\rho_{la}} DC_{il}^N, \quad (13.167)$$

where

$$DC_{il}^N = \gamma_v (DC_{il} - \mu_l^+ C_{il}) + \sum_{m=1}^6 \beta_m \frac{D_{il,m}^*}{\Delta L_{h,m}} (C_{il,m} - C_{il} + DI - C_{il,m}), \quad (13.168)$$

does not contain time derivatives and convection terms. Even these terms are the most strongly varying in transient processes during a single time step. In the case of negligible diffusion  $DC_{il}^N$  contains only source terms.

We realize that the expression on the right hand side is very similar to the left hand side of the concentration equation divided by the old time level density. A very useful approximation is then

$$\begin{aligned} & \gamma_{va} \alpha_{la} (C_{i,l} - C_{i,la}) - \Delta \tau \sum_{m=1}^6 b_{lm-} \alpha_{lm} (C_{i,lm} - C_{i,l}) \\ & \approx \alpha_{la} (C_{il} - C_{ila}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 b_{lm-} \alpha_{l,m} \frac{\rho_{l,m}}{\rho_{la}} (C_{il,m} - C_{il}) = \frac{\Delta \tau}{\rho_{la}} DC_{il}^N \end{aligned} \quad (13.169)$$

and

$$\begin{aligned} & \gamma_{va} \alpha_{la} (s_l - s_{la}) - \Delta \tau \sum_{m=1}^6 b_{lm-} \alpha_{lm} (s_{lm} - s_l) \\ & \approx \alpha_{la} (s_l - s_{la}) \gamma_{va} - \Delta \tau \sum_{m=1}^6 b_{lm-} \alpha_{l,m} \frac{\rho_{l,m}}{\rho_{la}} (s_{l,m} - s_l) = \frac{\Delta \tau}{\rho_{la}} Ds_l^N, \end{aligned} \quad (13.170)$$

Thus  $D\alpha_l$  can be approximated as follows

$$D\alpha_l = \gamma_v \frac{\mu_l}{\rho_l} - \alpha_{la} \left( \frac{\gamma_v - \gamma_{va}}{\Delta \tau} \right) - \frac{1}{\rho_l \rho_{la}} \left\{ \left( \frac{\partial \rho_l}{\partial s_l} \right)_a Ds_l^N + \sum_{i=2}^{i_{\max}} \left( \frac{\partial \rho_l}{\partial C_{i,l}} \right)_a DC_{il}^N \right\}. \quad (13.171)$$

Replacing with the normal velocities computed from the discretized momentum equation in linearized form we finally obtain the so called pressure equation

$$\begin{aligned} & p \gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} - \Delta \tau \sum_{m=1}^6 \beta_m \sum_{l=1}^{l_{\max}} \left[ \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \left( 1 + \frac{p_m - p}{\rho_l a_{la}^2} \right) \right] RVel_{lm} (p_m - p) \\ & = p_a \gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} + \Delta \tau \sum_{l=1}^{l_{\max}} D\alpha_l \\ & - \Delta \tau \sum_{m=1}^6 \beta_m \sum_{l=1}^{l_{\max}} \left[ \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \left( 1 + \frac{p_m - p}{\rho_l a_{la}^2} \right) \right] \left[ dV_{lm}^n - (\mathbf{e})^m \cdot \mathbf{V}_{cs,m} \right] \end{aligned} \quad (13.172)$$

or

$$cp + \sum_{m=1}^6 c_m p_m = d, \quad (13.173)$$

where

$$c_m = -\Delta\tau\beta_m \sum_{l=1}^{l_{\max}} \left[ \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \left( 1 + \frac{p_m - p}{\rho_l a_{la}^2} \right) \right] R V e l_{lm}, \quad (13.174)$$

$$c = p\gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} - \sum_{m=1}^6 c_m, \quad (13.175)$$

$$d = p_a \gamma_{va} \sum_{l=1}^{l_{\max}} \frac{\alpha_{la}}{\rho_l a_{la}^2} + \Delta\tau \sum_{l=1}^{l_{\max}} D\alpha_l - \Delta\tau \sum_{m=1}^6 \beta_m \sum_{l=1}^{l_{\max}} \left[ \xi_{lm+} \alpha_l + \xi_{lm-} \alpha_{lm} \left( 1 + \frac{p_m - p}{\rho_l a_{la}^2} \right) \right] \left[ dV_{lm}^n - (\mathbf{e})^m \cdot \mathbf{V}_{cs,m} \right]. \quad (13.176)$$

The advantage of Eq. (13.172) for the very first outer iteration step is that it takes the influence of all sources on the pressure change which is not the case in Eq. (13.156). The advantage of Eq. (13.156) for all subsequent outer iterations is that it reduces the residuals to very low value which is not the case with Eq. (13.172) because of the approximations (13.169) and (13.170). An additional source of numerical error is that the new density is usually computed within the outer iteration by using Eq. (13.159) and not Eq. (13.164). Combined, both equations result in a useful algorithm. As a predictor step use Eq. (13.172) and for all other iterations use Eq. (13.156).

## 13.12 Staggered x momentum equation

Two families of methods are known in the literature for solving partial differential equations with low order methods, the so called co-located and staggered grid methods. In the co-located methods all dependent variables are defined at the center of the mass control volume. In these methods unless the staggered grid method is used, discretization of order higher than the first order is required to create a stable numerical method. In the staggered grid method all dependent variables are defined in the center of the mass control volume except the velocities which are defined at the faces of the volume. In both cases the velocities are required for the center as well for the faces, so that the one group of velocities is usually computed by interpolation from the known other group. The control volume for the staggered grid methods consists of the half of the volumes belonging to each face. Strictly speaking the required geometrical information that has to be stored for these methods is four times those for the co-located methods. A compromise between low order methods using low storage and stability is to derive the discre-



tized form of the momentum equation in the staggered cell from already discretized momentum equations in the two neighboring cells. This is possible for the following reason. Momentum equations are force balances per unit mixture volume and therefore they can be volumetrically averaged over the staggered grids. In this section we will use this idea. As already mentioned the staggered computational cell in the  $\xi$  direction consists of the half of the mass control volumes belonging to the both sites of the  $\xi$  face. We will discretize the three components of the momentum equation in this staggered cell. Then we will use the dot product of the so discretized vector momentum equation with the unit face vector to obtain the normal velocity at the cell face. In doing this, we will try to keep the computational effort small by finding common coefficients for all three equations. This approach leads to a pressure gradient component normal to the face instead of a pressure gradient to each of the Cartesian directions, which is simply numerically treated. This is the key for designing implicit or semi-implicit methods.

**Time derivatives:** We start with the term  $\alpha_{la}\rho_{la}(u_l - u_{la})\gamma_{va}$ , perform volume averaging

$$\begin{aligned} & \overline{\alpha_{la}\rho_{la}(u_l - u_{la})\gamma_{va}} \\ &= \alpha_{la}\rho_{la}(u_l - u_{la})\frac{\gamma_{va}\Delta V}{\Delta V + \Delta V_{i+1}} + \alpha_{la,i+1}\rho_{la,i+1}(u_{l,i+1} - u_{la,i+1})\frac{\gamma_{va,i+1}\Delta V_{i+1}}{\Delta V + \Delta V_{i+1}} \end{aligned} \quad (13.177)$$

and approximate the average with

$$\overline{\alpha_{la}\rho_{la}(u_l - u_{la})\gamma_{va}} \approx (\alpha_{la}\rho_{la}\gamma_{va})_u (u_l^u - u_{la}^u), \quad (13.178)$$

where

$$(\alpha_{la}\rho_{la}\gamma_{va})_u = \alpha_{la}\rho_{la}\frac{\gamma_{va}\Delta V}{\Delta V + \Delta V_{i+1}} + \alpha_{la,i+1}\rho_{la,i+1}\frac{\gamma_{va,i+1}\Delta V_{i+1}}{\Delta V + \Delta V_{i+1}}. \quad (13.179)$$

Note that this procedure of averaging does not give

$$u_l^u = \frac{1}{2}(u_l + u_{l,i+1}) \quad (13.180)$$

in the general case. Similarly we have for the other directions momentum equations

$$\overline{\alpha_{la}\rho_{la}(v_l - v_{la})\gamma_{va}} \approx (\alpha_{la}\rho_{la}\gamma_{va})_v (v_l^u - v_{la}^u), \quad (13.181)$$

$$\overline{\alpha_{la}\rho_{la}(w_l - w_{la})\gamma_{va}} \approx (\alpha_{la}\rho_{la}\gamma_{va})_w (w_l^u - w_{la}^u). \quad (13.182)$$

We realize that in this way of approximation the component velocity differences for all three Cartesian directions possess a common coefficient.

**Convective terms:** The following approximation for the convective terms is proposed

$$\begin{aligned} & \overline{\sum_{m=1}^6 b_{lm-} \alpha_{l,m} \rho_{l,m} (u_{l,m} - u_l)} \\ &= \frac{1}{\Delta V + \Delta V_{i+1}} \left\{ \Delta V \sum_{m=1}^6 B_{lm-} (u_{l,m} - u_l) + \Delta V_{i+1} \sum_{m=1}^6 [B_{lm-} (u_{l,m} - u_l)]_{i+1} \right\} \\ &\approx \sum_{m=1}^6 b_{lm-}^u \alpha_{l,m}^u \rho_{l,m}^u (u_{l,m}^u - u_l^u), \end{aligned} \tag{13.183}$$

where

$$b_{lm-}^u \alpha_{l,m}^u \rho_{l,m}^u = C^* B_{lm-} + (1 - C^*) (B_{lm-})_{i+1}, \tag{13.184}$$

is the  $m$ -th face mass flow into the staggered cell divided by its volume and

$$C^* = \frac{\Delta V}{\Delta V + \Delta V_{i+1}}. \tag{13.185}$$

As in the case of the time derivatives we realize that in this way of approximation the component velocity differences for all three Cartesian directions possess a common coefficient.

**Diagonal diffusion terms:** We apply a similar procedure to the diagonal diffusion terms

$$\begin{aligned} & \overline{\sum_{m=1}^6 \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m1} (e)^{m1} \right] (u_{l,m} - u_l)} = \\ &= C^* \sum_{m=1}^6 \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m1} (e)^{m1} \right] (u_{l,m} - u_l) \\ &+ (1 - C^*) \sum_{m=1}^6 \left\{ \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m1} (e)^{m1} \right] (u_{l,m} - u_l) \right\}_{i+1} \\ &\approx \sum_{m=1}^6 \left( \begin{aligned} & C^* \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m1} (e)^{m1} \right] \\ & + \left\{ (1 - C^*) \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m1} (e)^{m1} \right] \right\}_{i+1} \end{aligned} \right) (u_{l,m}^u - u_l^u). \end{aligned} \tag{13.186}$$

Thus the combined convection-diffusion terms are finally approximated as follows

$$\begin{aligned} & -\sum_{m=1}^6 \left\{ B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m1} (e)^{m1} \right] \right\} (u_{l,m} - u_l) \\ & \approx \sum_{m=1}^6 a_{lm,cd} (u_{l,m}^u - u_l^u) + \sum_{m=1}^6 a_{lm,u\_dif} (u_{l,m}^u - u_l^u), \end{aligned} \quad (13.187)$$

where

$$a_{lm,cd} = -C^* \left( B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \right) - (1 - C^*) \left( B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \right)_{i+1}, \quad (13.188)$$

$$a_{lm,u\_dif} = -\frac{1}{3} \left\{ C^* \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} (e)^{m1} (e)^{m1} + (1 - C^*) \left[ \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} (e)^{m1} (e)^{m1} \right]_{i+1} \right\}. \quad (13.189)$$

Similarly we have

$$\begin{aligned} & -\sum_{m=1}^6 \left\{ B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m2} (e)^{m2} \right] \right\} (v_{l,m} - v_l) \\ & \approx \sum_{m=1}^6 a_{lm,cd} (v_{l,m}^u - v_l^u) + \sum_{m=1}^6 a_{lm,v\_dif} (v_{l,m}^u - v_l^u), \end{aligned} \quad (13.190)$$

$$a_{lm,v\_dif} = -\frac{1}{3} \left\{ C^* \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} (e)^{m2} (e)^{m2} + (1 - C^*) \left[ \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} (e)^{m2} (e)^{m2} \right]_{i+1} \right\}. \quad (13.191)$$

$$\begin{aligned} & -\sum_{m=1}^6 \left\{ B_{lm-} + \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left[ 1 + \frac{1}{3} (e)^{m3} (e)^{m3} \right] \right\} (w_{l,m} - w_l) \\ & \approx \sum_{m=1}^6 a_{lm,cd} (w_{l,m}^u - w_l^u) + \sum_{m=1}^6 a_{lm,w\_dif} (w_{l,m}^u - w_l^u), \end{aligned} \quad (13.192)$$

$$a_{lm,w\_dif} = -\frac{1}{3} \left\{ C^* \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} (e)^{m3} (e)^{m3} + (1 - C^*) \left[ \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} (e)^{m3} (e)^{m3} \right]_{i+1} \right\}. \quad (13.193)$$

We realize again that the coefficients  $a_{lm,cd}$  are common for all the momentum equations in the staggered cell.

**Drag force terms:** The following approximation contains in fact computation of the volume averages of the linearized drag coefficients.

$$\begin{aligned} & \overline{\gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| (u_m - u_l) + \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| (u_{cs} - u_l) \right]} \\ & \approx \sum_{\substack{m=1 \\ m \neq l}}^3 (\gamma_v c_{ml}^d)_u (u_m^u - u_l^u) + (\gamma_v c_{wl}^d)_u (u_{cs}^u - u_l^u), \end{aligned} \quad (13.194)$$

where

$$(\gamma_v c_{ml}^d)_u = \frac{1}{\Delta V + \Delta V_{i+1}} (\gamma_v \Delta V \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| + \gamma_{v,i+1} \Delta V_{i+1} \bar{c}_{ml,i+1}^d |\Delta \mathbf{V}_{ml}|_{i+1}), \quad (13.195)$$

$$(\gamma_v c_{wl}^d)_u = \frac{1}{\Delta V + \Delta V_{i+1}} (\gamma_v \Delta V \bar{c}_{wl}^d |\Delta \mathbf{V}_{csl}| + \gamma_{v,i+1} \Delta V_{i+1} \bar{c}_{wl,i+1}^d |\Delta \mathbf{V}_{csl}|_{i+1}). \quad (13.196)$$

Similarly we have

$$\begin{aligned} & \overline{\gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| (v_m - v_l) + \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| (v_{cs} - v_l) \right]} \\ & \approx \sum_{\substack{m=1 \\ m \neq l}}^3 (\gamma_v c_{ml}^d)_u (v_m^u - v_l^u) + (\gamma_v c_{wl}^d)_u (v_{cs}^u - v_l^u), \end{aligned} \quad (13.197)$$

$$\begin{aligned} & \overline{\gamma_v \left[ \sum_{\substack{m=1 \\ m \neq l}}^3 \bar{c}_{ml}^d |\Delta \mathbf{V}_{ml}| (w_m - w_l) + \bar{c}_{wl}^d |\Delta \mathbf{V}_{wl}| (w_{cs} - w_l) \right]} \\ & \approx \sum_{\substack{m=1 \\ m \neq l}}^3 (\gamma_v c_{ml}^d)_u (w_m^u - w_l^u) + (\gamma_v c_{wl}^d)_u (w_{cs}^u - w_l^u). \end{aligned} \quad (13.198)$$

Again the drag term coefficients for the momentum equations in the staggered cell are common.

**Gravitational force:** The volume averaging for the gravitational force gives

$$\overline{\alpha_l \rho_l g_x \gamma_v} = g_x (\alpha_{la} \rho_{la} \gamma_{va})_u, \quad (13.199)$$

$$\overline{\alpha_l \rho_l g_y \gamma_v} = g_y (\alpha_{la} \rho_{la} \gamma_{va})_u, \quad (13.200)$$

$$\overline{\alpha_l \rho_l g_z \gamma_v} = g_z (\alpha_{la} \rho_{la} \gamma_{va})_u. \quad (13.201)$$

**Interfacial momentum transfer due to mass transfer:** The interfacial momentum transfer is again approximated by first volume averaging the volume mass source terms due to interfacial mass transfer. The source terms due to external injection or suction are computed exactly because it is easy to prescribe the velocities corresponding to the sources at the mass cell center.

$$-\overline{\left[ \sum_{m=1}^{3,w} [\mu_{ml} \gamma_v (u_m - u_l)] + \mu_{lw} \gamma_v (u_{lw} - u_l) \right]} \\ \approx \sum_{m=1}^3 (\gamma_v \mu_{ml})_u (u_m^u - u_l^u) + (\gamma_v \mu_{wl})_u (u_{wl}^u - u_l^u) - (\gamma_v \mu_{lw})_u (u_{lw}^u - u_l^u) \quad (13.202)$$

$$(\gamma_v \mu_{ml})_u = \frac{1}{\Delta V + \Delta V_{i+1}} (\gamma_v \Delta V \mu_{ml} + \gamma_{v,i+1} \Delta V_{i+1} \mu_{ml,i+1}), \quad (13.203)$$

$$(\gamma_v \mu_{wl})_u = \frac{1}{\Delta V + \Delta V_{i+1}} (\gamma_v \Delta V \mu_{wl} + \gamma_{v,i+1} \Delta V_{i+1} \mu_{wl,i+1}), \quad (13.204)$$

$$(\gamma_v \mu_{lw})_u = \frac{1}{\Delta V + \Delta V_{i+1}} (\gamma_v \Delta V \mu_{lw} + \gamma_{v,i+1} \Delta V_{i+1} \mu_{lw,i+1}), \quad (13.205)$$

$$(\gamma_v \mu_{wl})_u u_{wl} = \frac{1}{\Delta V + \Delta V_{i+1}} (\gamma_v \Delta V \mu_{wl} u_{wl} + \gamma_{v,i+1} \Delta V_{i+1} \mu_{wl,i+1} u_{wl,i+1}), \quad (13.206)$$

$$(\gamma_v \mu_{lw})_u u_{lw} = \frac{1}{\Delta V + \Delta V_{i+1}} (\gamma_v \Delta V \mu_{lw} u_{lw} + \gamma_{v,i+1} \Delta V_{i+1} \mu_{lw,i+1} u_{lw,i+1}), \quad (13.207)$$

Similarly we have

$$-\overline{\left[ \sum_{m=1}^{3,w} [\mu_{ml} \gamma_v (v_m - v_l)] + \mu_{lw} \gamma_v (v_{lw} - v_l) \right]} \\ \approx \sum_{m=1}^3 (\gamma_v \mu_{ml})_u (v_m^u - v_l^u) + (\gamma_v \mu_{wl})_u (v_{wl}^u - v_l^u) - (\gamma_v \mu_{lw})_u (v_{lw}^u - v_l^u), \quad (13.208)$$

$$-\overline{\left[ \sum_{m=1}^{3,w} [\mu_{ml} \gamma_v (w_m - w_l)] + \mu_{lw} \gamma_v (w_{lw} - w_l) \right]} \\ \approx \sum_{m=1}^3 (\gamma_v \mu_{ml})_u (w_m^u - w_l^u) + (\gamma_v \mu_{wl})_u (w_{wl}^u - w_l^u) - (\gamma_v \mu_{lw})_u (w_{lw}^u - w_l^u). \quad (13.209)$$

**Lift force, off-diagonal viscous forces:** The lift force and the off-diagonal viscous forces are explicitly computed by strict volume averaging.

Pressure gradient:

$$\overline{\alpha_l \gamma_\xi (\nabla p)} \cdot \mathbf{i} = \alpha_{lu} \gamma_\xi (\nabla p) \cdot \mathbf{i}, \quad (13.210)$$

where

$$\alpha_{lu} = \frac{\gamma_v \Delta V \alpha_l + \gamma_{v,i+1} \Delta V_{i+1} \alpha_{l,i+1}}{\gamma_v \Delta V + \gamma_{v,i+1} \Delta V_{i+1}}. \quad (13.211)$$

Similarly we have

$$\overline{\alpha_l \gamma_\xi (\nabla p)} \cdot \mathbf{j} = \alpha_{lu} \gamma_\xi (\nabla p) \cdot \mathbf{j}, \quad (13.212)$$

$$\overline{\alpha_l \gamma_\xi (\nabla p)} \cdot \mathbf{k} = \alpha_{lu} \gamma_\xi (\nabla p) \cdot \mathbf{k}. \quad (13.213)$$

Virtual mass force:

$$\begin{aligned} & \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{3,cs} \overline{c_{ml}^{vm}} \left( \frac{\partial \Delta u_{ml}}{\partial \tau} + \overline{V}^1 \frac{\partial \Delta u_{ml}}{\partial \xi} + \overline{V}^2 \frac{\partial \Delta u_{ml}}{\partial \eta} + \overline{V}^3 \frac{\partial \Delta u_{ml}}{\partial \zeta} \right) \\ &= \sum_{\substack{m=1 \\ m \neq l}}^{3,cs} \left( \gamma_v \overline{c_{ml}^{vm}} \right)_u \left[ \begin{aligned} & \left( \frac{u_m^u - u_l^u}{\Delta \tau} - \frac{u_{ma}^u - u_{la}^u}{\Delta \tau} + (\overline{V}^1)_1 (u_{m,i+1} - u_{l,i+1} - u_m + u_l) \right) \\ & + (\overline{V}^2)_1 \left( \frac{\partial u_m}{\partial \eta} \Big|_1 - \frac{\partial u_l}{\partial \eta} \Big|_1 \right) + (\overline{V}^3)_1 \left( \frac{\partial u_m}{\partial \zeta} \Big|_1 - \frac{\partial u_l}{\partial \zeta} \Big|_1 \right) \end{aligned} \right]. \quad (13.214) \end{aligned}$$

Similarly we have

$$\begin{aligned} & \gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{3,cs} \overline{c_{ml}^{vm}} \left( \frac{\partial \Delta v_{ml}}{\partial \tau} + \overline{V}^1 \frac{\partial \Delta v_{ml}}{\partial \xi} + \overline{V}^2 \frac{\partial \Delta v_{ml}}{\partial \eta} + \overline{V}^3 \frac{\partial \Delta v_{ml}}{\partial \zeta} \right) \\ &= \sum_{\substack{m=1 \\ m \neq l}}^{3,cs} \left( \gamma_v \overline{c_{ml}^{vm}} \right)_u \left[ \begin{aligned} & \left( \frac{v_m^u - v_l^u}{\Delta \tau} - \frac{v_{ma}^u - v_{la}^u}{\Delta \tau} + (\overline{V}^1)_1 (v_{m,i+1} - v_{l,i+1} - v_m + v_l) \right) \\ & + (\overline{V}^2)_1 \left( \frac{\partial v_m}{\partial \eta} \Big|_1 - \frac{\partial v_l}{\partial \eta} \Big|_1 \right) + (\overline{V}^3)_1 \left( \frac{\partial v_m}{\partial \zeta} \Big|_1 - \frac{\partial v_l}{\partial \zeta} \Big|_1 \right) \end{aligned} \right], \quad (13.215) \end{aligned}$$

$$\gamma_v \sum_{\substack{m=1 \\ m \neq l}}^{3,cs} \overline{c_{ml}^{vm}} \left( \frac{\partial \Delta w_{ml}}{\partial \tau} + \overline{V}^1 \frac{\partial \Delta w_{ml}}{\partial \xi} + \overline{V}^2 \frac{\partial \Delta w_{ml}}{\partial \eta} + \overline{V}^3 \frac{\partial \Delta w_{ml}}{\partial \zeta} \right)$$

$$= \sum_{\substack{m=1 \\ m \neq l}}^{3,cs} (\gamma_v \bar{c}_{ml}^{vm})_u \left[ \begin{aligned} & \left( \frac{w_m^u - w_l^u}{\Delta \tau} - \frac{w_{ma}^u - w_{la}^u}{\Delta \tau} + (\bar{V}^1)_1 (w_{m,i+1} - w_{l,i+1} - w_m + w_l) \right) \\ & + (\bar{V}^2)_1 \left( \frac{\partial w_m}{\partial \eta} \Big|_{\cdot 1} - \frac{\partial w_l}{\partial \eta} \Big|_{\cdot 1} \right) + (\bar{V}^3)_1 \left( \frac{\partial w_m}{\partial \zeta} \Big|_{\cdot 1} - \frac{\partial w_l}{\partial \zeta} \Big|_{\cdot 1} \right) \end{aligned} \right]. \tag{13.216}$$

Implicit treatment of the interfacial interaction:

$$\begin{aligned} & (\alpha_{la} \rho_{la} \gamma_{va})_u \frac{u_l^u - u_{la}^u}{\Delta \tau} + \sum_{m=1}^6 a_{lm,cd} (u_{l,m}^u - u_l^u) + \alpha_{l,u} \gamma_\xi (\nabla p) \cdot \mathbf{i} \\ & - \sum_{\substack{m=1 \\ m \neq l}}^{3,cs} (\gamma_v \bar{c}_{ml}^{vm})_u \left[ \begin{aligned} & \left( \frac{u_m^u - u_l^u}{\Delta \tau} - \frac{u_{ma}^u - u_{la}^u}{\Delta \tau} + (\bar{V}^1)_1 (u_{m,i+1} - u_{l,i+1} - u_m + u_l) \right) \\ & + (\bar{V}^2)_1 \left( \frac{\partial u_m}{\partial \eta} \Big|_{\cdot 1} - \frac{\partial u_l}{\partial \eta} \Big|_{\cdot 1} \right) + (\bar{V}^3)_1 \left( \frac{\partial u_m}{\partial \zeta} \Big|_{\cdot 1} - \frac{\partial u_l}{\partial \zeta} \Big|_{\cdot 1} \right) \end{aligned} \right] - (\gamma_v f_l^L)_u \\ & = - \sum_{m=1}^6 a_{lm,u\_dif} (u_{l,m}^u - u_l^u) - g_x (\alpha_{la} \rho_{la} \gamma_{va})_u \\ & + \sum_{m=1}^3 (\gamma_v \mu_{ml})_u (u_m^u - u_l^u) + (\gamma_v \mu_{wl})_u (u_{wl}^u - u_l^u) - (\gamma_v \mu_{lw})_u (u_{lw}^u - u_l^u) \\ & + \sum_{\substack{m=1 \\ m \neq l}}^3 (\gamma_v c_{ml}^d)_u (u_m^u - u_l^u) + (\gamma_v c_{vl}^d)_u (u_{cs}^u - u_l^u) + Vi s_l^u. \end{aligned} \tag{13.217}$$

For all three velocity fields we have a system of algebraic equations with respect to the corresponding field velocity components in the  $x$  direction

$$\begin{aligned} & \left[ (\alpha_{la} \rho_{la} \gamma_{va})_u \frac{1}{\Delta \tau} - \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} + a_{l,cd} + (\gamma_v \bar{c}_{vl}^{vm})_u \frac{1}{\Delta \tau} + (\gamma_v c_{vl}^d)_u + (\gamma_v \mu_{wl})_u - (\gamma_v \mu_{lw})_u \right] u_l^u \\ & + \sum_{\substack{m=1 \\ m \neq l}}^3 a_{lm} u_m^u = b_l - \alpha_{l,u} \gamma_\xi (\nabla p) \cdot \mathbf{i} \end{aligned} \tag{13.218}$$

or

$$\mathbf{A}^u \mathbf{u}^u = \mathbf{b} u^u - \mathbf{a}^u \gamma_\xi (\nabla p) \cdot \mathbf{i}. \tag{13.219}$$

The non-diagonal and the diagonal elements of the  $\mathbf{A}^u$  matrix are

$$a_{ml} = - \left[ \left( \gamma_v \bar{c}_{ml}^{vm} \right)_u \frac{1}{\Delta \tau} + \left( \gamma_v \mu_{ml} \right)_u + \left( \gamma_v c_{ml}^d \right)_u \right], \quad (13.220)$$

and

$$a_{ll} = \left( \alpha_{la} \rho_{la} \gamma_{va} \right)_u \frac{1}{\Delta \tau} - \sum_{m=1}^3 a_{lm} + a_{l,cd} + \left( \gamma_v \bar{c}_{wl}^{vm} \right)_u \frac{1}{\Delta \tau} + \left( \gamma_v c_{wl}^d \right)_u + \left( \gamma_v \mu_{wl} \right)_u - \left( \gamma_v \mu_{lw} \right)_u, \quad (13.221)$$

respectively, where

$$a_{l,cd} = - \sum_{m=1}^6 a_{lm,cd}. \quad (13.222)$$

The elements of the algebraic vector  $\mathbf{bu}^u$  are

$$\begin{aligned} bu_l &= bu_{l,conv} + Vis_l^u + \left( \alpha_{la} \rho_{la} \gamma_{va} \right)_u \left( \frac{u_{la}^u}{\Delta \tau} - g_x \right) - \sum_{\substack{m=1 \\ m \neq l}}^{3,cs} bu_{lm} + \left( \gamma_v f_l^L \right)_u \\ &+ \left[ \left( \gamma_v \bar{c}_{csl}^{vm} \right)_u \frac{1}{\Delta \tau} + \left( \gamma_v c_{wl}^d \right)_u \right] u_{cs}^u + \left[ \left( \gamma_v \mu_{wl} \right)_u - \left( \gamma_v \mu_{lw} \right)_u \right] u_{lw}^u \end{aligned} \quad (13.223)$$

where

$$bu_{ml} = -bu_{lm} = \left( \gamma_v \bar{c}_{ml}^{vm} \right)_u \left[ - \frac{u_{ma}^u - u_{la}^u}{\Delta \tau} + \left( \bar{V}^1 \right)_1 (u_{m,i+1} - u_{l,i+1} - u_m + u_l) + \left( \bar{V}^2 \right)_1 \left( \frac{\partial u_m}{\partial \eta} \Big|_1 - \frac{\partial u_l}{\partial \eta} \Big|_1 \right) + \left( \bar{V}^3 \right)_1 \left( \frac{\partial u_m}{\partial \zeta} \Big|_1 - \frac{\partial u_l}{\partial \zeta} \Big|_1 \right) \right] \quad (13.224)$$

and

$$bu_{l,conv} = - \sum_{m=1}^6 \left[ \left( a_{lm,cd} + a_{lm,u\_dif} \right) u_{l,m}^u - a_{lm,u\_dif} u_l^u \right]. \quad (13.225)$$

The elements of the algebraic vector  $\mathbf{a}^u$  are  $\alpha_{l,u}$ , where  $l = 1, 2, 3$ . Note that by definition, if one velocity field does not exist,  $\alpha_l = 0$ , the coefficients describing its coupling with the other fields are then equal to zero. Similarly we can discretize the momentum equations in the same staggered control volume for the other Cartesian components. The result in component form is then



$$\mathbf{A}^u \mathbf{u}^u = \mathbf{b} \mathbf{u}^u - \mathbf{a}^u \gamma_\xi (\nabla p) \cdot \mathbf{i}, \quad (13.226)$$

$$\mathbf{A}^v \mathbf{v}^u = \mathbf{b} \mathbf{v}^u - \mathbf{a}^u \gamma_\xi (\nabla p) \cdot \mathbf{j}, \quad (13.227)$$

$$\mathbf{A}^w \mathbf{w}^u = \mathbf{b} \mathbf{w}^u - \mathbf{a}^u \gamma_\xi (\nabla p) \cdot \mathbf{k}. \quad (13.228)$$

It is remarkable that the  $\mathbf{A}$  matrix and the coefficients of the pressure gradient are common for all the three systems of equations. If we take the dot product of each  $u$ - $v$ - $w$  equation with the unit normal vector at the control volume face we then obtain

$$\begin{aligned} & \mathbf{A}^u \left[ (e^{11})_1 \mathbf{u}^u + (e^{12})_1 \mathbf{v}^u + (e^{13})_1 \mathbf{w}^u \right] \\ &= (e^{11})_1 \mathbf{b} \mathbf{u}^u + (e^{12})_1 \mathbf{b} \mathbf{v}^u + (e^{13})_1 \mathbf{b} \mathbf{w}^u - \mathbf{a}^u \gamma_\xi \left[ (\mathbf{e}^1)_1 \cdot (\nabla p) \right]. \end{aligned} \quad (13.229)$$

Having in mind that the outwards pointing normal face velocity is

$$\bar{\mathbf{V}}^n = (\mathbf{e}^1)_1 \cdot \mathbf{V}^u, \quad (13.230)$$

and

$$\frac{\partial p}{\partial \xi} = (\mathbf{e}^1)_1 \cdot (\nabla p), \quad (13.231)$$

we obtain finally

$$\mathbf{A}^u \bar{\mathbf{V}}^n = \mathbf{b}^u - \mathbf{a}^u \gamma_\xi \frac{\partial p}{\partial \xi}, \quad (13.232)$$

where

$$\mathbf{b}^u = (e^{11})_1 \mathbf{b} \mathbf{u}^u + (e^{12})_1 \mathbf{b} \mathbf{v}^u + (e^{13})_1 \mathbf{b} \mathbf{w}^u. \quad (13.233)$$

This algebraic system can be solved with respect to each field velocity provided that

$$\det \mathbf{A}^u = a_{11} a_{22} a_{33} + a_{12} a_{23} a_{31} + a_{21} a_{32} a_{13} - a_{31} a_{22} a_{13} - a_{32} a_{23} a_{11} - a_{12} a_{21} a_{33} \neq 0. \quad (13.234)$$

The result is

$$\bar{\mathbf{V}}^n = \mathbf{d} \mathbf{V}_\xi - \mathbf{R} \mathbf{V}_\xi \gamma_\xi (p_{i+1} - p), \quad (13.235)$$

where

$$\mathbf{dV}_\xi = (\mathbf{A}^u)^{-1} \mathbf{b}^u \quad (13.236)$$

with components

$$d\bar{V}_{\xi,l} = \left( \sum_{m=1}^3 b_m \bar{a}_{lm} \right) / \det \mathbf{A}^u, \quad (13.237)$$

and

$$\mathbf{RV}_\xi = (\mathbf{A}^u)^{-1} \mathbf{a}_u, \quad (13.238)$$

with components

$$RU_l = \left( \sum_{m=1}^3 \alpha_{ma,u} \bar{a}_{lm} \right) / \det \mathbf{A}_u, \quad (13.239)$$

and the  $\bar{a}$  values are

$$\begin{aligned} \bar{a}_{11} &= a_{22}a_{33} - a_{32}a_{23}, & \bar{a}_{12} &= a_{32}a_{13} - a_{12}a_{33}, & \bar{a}_{13} &= a_{12}a_{23} - a_{22}a_{13}, \\ \bar{a}_{21} &= a_{23}a_{31} - a_{21}a_{33}, & \bar{a}_{22} &= a_{11}a_{33} - a_{31}a_{13}, & \bar{a}_{23} &= a_{21}a_{13} - a_{23}a_{11}, \\ \bar{a}_{31} &= a_{21}a_{32} - a_{31}a_{22}, & \bar{a}_{32} &= a_{12}a_{31} - a_{32}a_{11}, & \bar{a}_{33} &= a_{11}a_{22} - a_{21}a_{12}. \end{aligned} \quad (13.240-248)$$

Actually, not the absolute but the relative normal face velocity is required to construct the pressure equation which is readily obtained

$$\mathbf{V}^n = \bar{\mathbf{V}}^n - (\mathbf{e})^1 \cdot \mathbf{V}_{cs}. \quad (13.249)$$

## Appendix 13.1 Harmonic averaged diffusion coefficients

A natural averaging of the coefficients describing diffusion across the face  $m$ , having surface cross section  $S_m$  is then the harmonic averaging

$$\frac{D_{l,m}^\Phi}{\Delta L_{h,m}} = \left( \frac{\Phi_l}{\Delta V} \right) S_m = S_m \frac{2(\Phi_l)(\Phi_l)_m}{\Delta V_m (\Phi_l) + \Delta V (\Phi_l)_m}$$

where on the right hand side  $m = 1, 2, 3, 4, 5, 6$  is equivalent to  $i + 1, i - 1, j + 1, j - 1, k + 1, k - 1$ , respectively regarding the properties inside a control volumes.  $\Delta V$  is the non-staggered cell volume, and  $\Delta V_m$  is the volume of the cell at the other side of face  $m$ . It guaranties that if the field in one of the neighboring cells is missing the diffusion coefficient is zero. This property is derived from the solution of the steady state one-dimensional diffusion equations.

For computation of

$$\frac{D_{il,m}^*}{\Delta L_{h,m}} = \left( \frac{\alpha_l \rho_l D_{il}^*}{\Delta V} \right)_m S_m = S_m \frac{2(\alpha_l \rho_l D_{il}^*)(\alpha_l \rho_l D_{il}^*)_m}{\Delta V_m (\alpha_l \rho_l D_{il}^*) + \Delta V (\alpha_l \rho_l D_{il}^*)_m}$$

we simply set  $\Phi_l = \alpha_l \rho_l D_{il}^*$ .

For computation of  $\frac{D_{l,m}^T}{\Delta L_{h,m}}$  we simply set  $\Phi_l = \alpha_l \lambda_l$ .

For computation of  $\frac{D_{il,m}^{sC}}{\Delta L_{h,m}}$  we simply set  $\Phi_l = \alpha_l \rho_l D_{il}^* (s_{il} - s_{l_l})$ . Note that

$$\frac{D_{il,m}^C}{\Delta L_{h,m}} = 0$$

for  $s_{il} = s_{l_l}$  or  $s_{il,m} = s_{l_l,m}$ .

For computation of the turbulent particle diffusion coefficient  $\frac{D_{l,m}^n}{\Delta L_{h,m}}$  we simply

set  $\Phi_l = \frac{V_l^t}{Sc^t}$ .

For computation of  $\frac{D_{l,m}^v}{\Delta L_{h,m}}$  we simply set  $\Phi_l = \alpha_l \rho_l V_l^*$ .

In the case of cylindrical or Cartesian coordinate systems we have zero off-diagonal diffusion terms and

$$\frac{D_{l1}^\Phi}{\Delta r_h} = \frac{2(\Phi_l)(\Phi_l)_{i+1}}{\Delta r(\Phi_l)_{i+1} + \Delta r_{i+1}\Phi_l},$$

$$\frac{D_{l2}^\Phi}{\Delta r_{h,i-1}} = \frac{2(\Phi_l)(\Phi_l)_{i-1}}{\Delta r(\Phi_l)_{i-1} + \Delta r_{i-1}\Phi_l},$$

$$\frac{D_{l3}^\Phi}{r^\kappa \Delta \theta_h} = \frac{2(\Phi_l)(\Phi_l)_{j+1}}{r^\kappa [\Delta \theta(\Phi_l)_{j+1} + \Delta \theta_{j+1}\Phi_l]},$$

$$\frac{D_{l4}^\Phi}{r^\kappa \Delta \theta_{h,j-1}} = \frac{2(\Phi_l)(\Phi_l)_{j-1}}{r^\kappa [\Delta \theta(\Phi_l)_{j-1} + \Delta \theta_{j-1}\Phi_l]},$$

$$\frac{D_{l5}^\Phi}{\Delta z_h} = \frac{2(\Phi_l)(\Phi_l)_{k+1}}{\Delta z(\Phi_l)_{k+1} + \Delta z_{k+1}\Phi_l},$$

$$\frac{D_{l6}^\Phi}{\Delta z_{h,k-1}} = \frac{2(\Phi_l)(\Phi_l)_{k-1}}{\Delta z(\Phi_l)_{k-1} + \Delta z_{k-1}\Phi_l}.$$

## Appendix 13.2 Off-diagonal viscous diffusion terms of the x momentum equation

The off-diagonal viscous diffusion terms in the x momentum equation are

$$\begin{aligned} & \sum_{m=1}^6 \beta_m \frac{D_{l,m}^v}{\Delta L_{h,m}} \left( DI_{-} u_{l,m} - \frac{2}{3} (e)^{m1} DI_{-} u_{l,m}^b + DI_{-} vis_{lm}^{uT} \right) \\ &= \beta_1^* \left\{ q_{x1,12} \left. \frac{\partial u_l}{\partial \eta} \right|_1 + q_{x1,13} \left. \frac{\partial u_l}{\partial \zeta} \right|_1 + q_{x1,22} \left. \frac{\partial v_l}{\partial \eta} \right|_1 + q_{x1,23} \left. \frac{\partial v_l}{\partial \zeta} \right|_1 + q_{x1,32} \left. \frac{\partial w_l}{\partial \eta} \right|_1 + q_{x1,33} \left. \frac{\partial w_l}{\partial \zeta} \right|_1 \right\} \\ &+ \beta_2^* \left\{ q_{x2,12} \left. \frac{\partial u_l}{\partial \eta} \right|_2 + q_{x2,13} \left. \frac{u_l}{\partial \zeta} \right|_2 + q_{x2,22} \left. \frac{\partial v_l}{\partial \eta} \right|_2 + q_{x2,23} \left. \frac{\partial v_l}{\partial \zeta} \right|_2 + q_{x2,32} \left. \frac{\partial w_l}{\partial \eta} \right|_2 + q_{x2,33} \left. \frac{\partial w_l}{\partial \zeta} \right|_2 \right\} \\ &+ \beta_3^* \left\{ q_{x3,11} \left. \frac{\partial u_l}{\partial \xi} \right|_3 + q_{x3,13} \left. \frac{\partial u_l}{\partial \zeta} \right|_3 + q_{x3,21} \left. \frac{\partial v_l}{\partial \xi} \right|_3 + q_{x3,23} \left. \frac{\partial v_l}{\partial \zeta} \right|_3 + q_{x3,31} \left. \frac{\partial w_l}{\partial \xi} \right|_3 + q_{x3,33} \left. \frac{\partial w_l}{\partial \zeta} \right|_3 \right\} \\ &+ \beta_4^* \left\{ q_{x4,11} \left. \frac{\partial u_l}{\partial \xi} \right|_4 + q_{x4,13} \left. \frac{\partial u_l}{\partial \zeta} \right|_4 + q_{x4,21} \left. \frac{\partial v_l}{\partial \xi} \right|_4 + q_{x4,23} \left. \frac{\partial v_l}{\partial \zeta} \right|_4 + q_{x4,31} \left. \frac{\partial w_l}{\partial \xi} \right|_4 + q_{x4,33} \left. \frac{\partial w_l}{\partial \zeta} \right|_4 \right\} \\ &+ \beta_5^* \left\{ q_{x5,11} \left. \frac{\partial u_l}{\partial \xi} \right|_5 + q_{x5,12} \left. \frac{\partial u_l}{\partial \eta} \right|_5 + q_{x5,21} \left. \frac{\partial v_l}{\partial \xi} \right|_5 + q_{x5,22} \left. \frac{\partial v_l}{\partial \eta} \right|_5 + q_{x5,31} \left. \frac{\partial w_l}{\partial \xi} \right|_5 + q_{x5,32} \left. \frac{\partial w_l}{\partial \eta} \right|_5 \right\} \\ &+ \beta_6^* \left\{ q_{x6,11} \left. \frac{\partial u_l}{\partial \xi} \right|_6 + q_{x6,12} \left. \frac{\partial u_l}{\partial \eta} \right|_6 + q_{x6,21} \left. \frac{\partial v_l}{\partial \xi} \right|_6 + q_{x6,22} \left. \frac{\partial v_l}{\partial \eta} \right|_6 + q_{x6,31} \left. \frac{\partial w_l}{\partial \xi} \right|_6 + q_{x6,32} \left. \frac{\partial w_l}{\partial \eta} \right|_6 \right\}. \end{aligned}$$

Here the coefficients

$$\beta_m^* = \beta_m = \frac{D_{l,m}^v}{\Delta L_{h,m}} \frac{\overline{\Delta V}_m}{S_m}$$

are used also in the other momentum equations. The following 36 coefficients are functions of the geometry only.

$$\begin{aligned} q_{x1,12} &= \frac{4}{3} (e)^{11} (a^{21})_1 + (e)^{12} (a^{22})_1 + (e)^{13} (a^{23})_1 = d_{12} + \frac{1}{3} (e)^{11} (a^{21})_1, \\ q_{x1,13} &= \frac{4}{3} (e)^{11} (a^{31})_1 + (e)^{12} (a^{32})_1 + (e)^{13} (a^{33})_1 = d_{13} + \frac{1}{3} (e)^{11} (a^{31})_1, \end{aligned}$$

$$q_{x1,22} = (e)^{12} (a^{21})_1 - \frac{2}{3} (e)^{11} (a^{22})_1,$$

$$q_{x1,23} = (e)^{12} (a^{31})_1 - \frac{2}{3} (e)^{11} (a^{32})_1,$$

$$q_{x1,32} = (e)^{13} (a^{21})_1 - \frac{2}{3} (e)^{11} (a^{23})_1,$$

$$q_{x1,33} = (e)^{13} (a^{31})_1 - \frac{2}{3} (e)^{11} (a^{33})_1,$$

$$\begin{aligned} q_{x2,12} &= \frac{4}{3} (e)^{21} (a^{21})_2 + (e)^{22} (a^{22})_2 + (e)^{23} (a^{23})_2 = d_{22} + \frac{1}{3} (e)^{21} (a^{21})_2 \\ &= -(q_{x1,12})_{i-1}, \end{aligned}$$

$$\begin{aligned} q_{x2,13} &= \frac{4}{3} (e)^{21} (a^{31})_2 + (e)^{22} (a^{32})_2 + (e)^{23} (a^{33})_2 = d_{23} + \frac{1}{3} (e)^{21} (a^{31})_2 \\ &= -(q_{x1,13})_{i-1}, \end{aligned}$$

$$q_{x2,22} = (e)^{22} (a^{21})_2 - \frac{2}{3} (e)^{21} (a^{22})_2 = -(q_{x1,22})_{i-1},$$

$$q_{x2,23} = (e)^{22} (a^{31})_2 - \frac{2}{3} (e)^{21} (a^{32})_2 = -(q_{x1,23})_{i-1},$$

$$q_{x2,32} = (e)^{23} (a^{21})_2 - \frac{2}{3} (e)^{21} (a^{23})_2 = -(q_{x1,32})_{i-1},$$

$$q_{x2,33} = (e)^{23} (a^{31})_2 - \frac{2}{3} (e)^{21} (a^{33})_2 = -(q_{x1,33})_{i-1},$$

$$q_{x3,11} = \frac{4}{3} (e)^{31} (a^{11})_3 + (e)^{32} (a^{12})_3 + (e)^{33} (a^{13})_3 = d_{31} + \frac{1}{3} (e)^{31} (a^{11})_3,$$

$$q_{x3,13} = \frac{4}{3} (e)^{31} (a^{31})_3 + (e)^{32} (a^{32})_3 + (e)^{33} (a^{33})_3 = d_{33} + \frac{1}{3} (e)^{31} (a^{31})_3,$$

$$q_{x3,21} = (e)^{32} (a^{11})_3 - \frac{2}{3} (e)^{31} (a^{12})_3,$$

$$q_{x3,23} = (e)^{32} (a^{31})_3 - \frac{2}{3} (e)^{31} (a^{32})_3,$$

$$q_{x3,31} = (e)^{33} (a^{11})_3 - \frac{2}{3} (e)^{31} (a^{13})_3,$$

$$q_{x3,33} = (e)^{33} (a^{31})_3 - \frac{2}{3} (e)^{31} (a^{33})_3,$$

$$\begin{aligned} q_{x4,11} &= \frac{4}{3} (e)^{41} (a^{11})_4 + (e)^{42} (a^{12})_4 + (e)^{43} (a^{13})_4 = d_{41} + \frac{1}{3} (e)^{41} (a^{11})_4 \\ &= -(q_{x3,11})_{j-1}, \end{aligned}$$

$$\begin{aligned} q_{x4,13} &= \frac{4}{3} (e)^{41} (a^{31})_4 + (e)^{42} (a^{32})_4 + (e)^{43} (a^{33})_4 = d_{43} + \frac{1}{3} (e)^{41} (a^{31})_4 \\ &= -(q_{x3,13})_{j-1}, \end{aligned}$$

$$q_{x4,21} = (e)^{42} (a^{11})_4 - \frac{2}{3} (e)^{41} (a^{12})_4 = -(q_{x3,21})_{j-1},$$

$$q_{x4,23} = (e)^{42} (a^{31})_4 - \frac{2}{3} (e)^{41} (a^{32})_4 = -(q_{x3,23})_{j-1},$$

$$q_{x4,31} = (e)^{43} (a^{11})_4 - \frac{2}{3} (e)^{41} (a^{13})_4 = -(q_{x3,31})_{j-1},$$

$$q_{x4,33} = (e)^{43} (a^{31})_4 - \frac{2}{3} (e)^{41} (a^{33})_4 = -(q_{x3,33})_{j-1},$$

$$q_{x5,11} = \frac{4}{3} (e)^{51} (a^{11})_5 + (e)^{52} (a^{12})_5 + (e)^{53} (a^{13})_5 = d_{51} + \frac{1}{3} (e)^{51} (a^{11})_5,$$

$$q_{x5,12} = \frac{4}{3} (e)^{51} (a^{21})_5 + (e)^{52} (a^{22})_5 + (e)^{53} (a^{23})_5 = d_{52} + \frac{1}{3} (e)^{51} (a^{21})_5,$$

$$q_{x5,21} = (e)^{52} (a^{11})_5 - \frac{2}{3} (e)^{51} (a^{12})_5,$$

$$q_{x5,22} = (e)^{52} (a^{21})_5 - \frac{2}{3} (e)^{51} (a^{22})_5,$$

$$q_{x5,31} = (e)^{53} (a^{11})_5 - \frac{2}{3} (e)^{51} (a^{13})_5,$$

$$q_{x5,32} = (e)^{53} (a^{21})_5 - \frac{2}{3} (e)^{51} (a^{23})_5,$$

$$\begin{aligned} q_{x6,11} &= \frac{4}{3} (e)^{61} (a^{11})_6 + (e)^{62} (a^{12})_6 + (e)^{63} (a^{13})_6 = d_{61} + \frac{1}{3} (e)^{61} (a^{11})_6 \\ &= -(q_{x5,11})_{k-1}, \end{aligned}$$

$$\begin{aligned}
q_{x6,12} &= \frac{4}{3}(e)^{61} (a^{21})_6 + (e)^{62} (a^{22})_6 + (e)^{63} (a^{23})_6 = d_{62} + \frac{1}{3}(e)^{61} (a^{21})_6 \\
&= -(q_{x5,12})_{k-1}, \\
q_{x6,21} &= (e)^{62} (a^{11})_6 - \frac{2}{3}(e)^{61} (a^{12})_6 = -(q_{x5,21})_{k-1}, \\
q_{x6,22} &= (e)^{62} (a^{21})_6 - \frac{2}{3}(e)^{61} (a^{22})_6 = -(q_{x5,22})_{k-1}, \\
q_{x6,31} &= (e)^{63} (a^{11})_6 - \frac{2}{3}(e)^{61} (a^{13})_6 = -(q_{x5,31})_{k-1}, \\
q_{x6,32} &= (e)^{63} (a^{21})_6 - \frac{2}{3}(e)^{61} (a^{23})_6 = -(q_{x5,32})_{k-1}.
\end{aligned}$$

### Appendix 13.3 Off-diagonal viscous diffusion terms of the $y$ momentum equation

The off-diagonal viscous diffusion terms in the  $y$  momentum equation

$$\sum_{m=1}^6 \beta_m \frac{D_{l,m}^y}{\Delta L_{h,m}} \left( DI_{-} v_{l,m} - \frac{2}{3}(e)^{m2} DI_{-} u_{l,m}^b + DI_{-} vis_{lm}^{yT} \right)$$

are computed using the same procedure as those for  $x$  equation replacing simply the subscript  $x$  with  $y$  and using the following geometry coefficients.

$$\begin{aligned}
q_{y1,12} &= (e)^{11} (a^{22})_1 - \frac{2}{3}(e)^{12} (a^{21})_1, \\
q_{y1,13} &= (e)^{11} (a^{32})_1 - \frac{2}{3}(e)^{12} (a^{31})_1, \\
q_{y1,22} &= (e)^{11} (a^{21})_1 + \frac{4}{3}(e)^{12} (a^{22})_1 + (e)^{13} (a^{23})_1, \\
q_{y1,23} &= (e)^{11} (a^{31})_1 + \frac{4}{3}(e)^{12} (a^{32})_1 + (e)^{13} (a^{33})_1, \\
q_{y1,32} &= (e)^{13} (a^{22})_1 - \frac{2}{3}(e)^{12} (a^{23})_1, \\
q_{y1,33} &= (e)^{13} (a^{32})_1 - \frac{2}{3}(e)^{12} (a^{33})_1,
\end{aligned}$$

$$q_{y2,12} = (e)^{21} (a^{22})_2 - \frac{2}{3} (e)^{22} (a^{21})_2 = -(q_{y1,12})_{i-1},$$

$$q_{y2,13} = (e)^{21} (a^{32})_2 - \frac{2}{3} (e)^{22} (a^{31})_2 = -(q_{y1,13})_{i-1},$$

$$q_{y2,22} = (e)^{21} (a^{21})_2 + \frac{4}{3} (e)^{22} (a^{22})_2 + (e)^{23} (a^{23})_2 = -(q_{y1,22})_{i-1},$$

$$q_{y2,23} = (e)^{21} (a^{31})_2 + \frac{4}{3} (e)^{22} (a^{32})_2 + (e)^{23} (a^{33})_2 = -(q_{y1,23})_{i-1},$$

$$q_{y2,32} = (e)^{23} (a^{22})_2 - \frac{2}{3} (e)^{22} (a^{23})_2 = -(q_{y1,32})_{i-1},$$

$$q_{y2,33} = (e)^{23} (a^{32})_2 - \frac{2}{3} (e)^{22} (a^{33})_2 = -(q_{y1,33})_{i-1},$$

$$q_{y3,11} = (e)^{31} (a^{12})_3 - \frac{2}{3} (e)^{32} (a^{11})_3,$$

$$q_{y3,13} = (e)^{31} (a^{32})_3 - \frac{2}{3} (e)^{32} (a^{31})_3,$$

$$q_{y3,21} = (e)^{31} (a^{11})_3 + \frac{4}{3} (e)^{32} (a^{12})_3 + (e)^{33} (a^{13})_3,$$

$$q_{y3,23} = (e)^{31} (a^{31})_3 + \frac{4}{3} (e)^{32} (a^{32})_3 + (e)^{33} (a^{33})_3,$$

$$q_{y3,31} = (e)^{33} (a^{12})_3 - \frac{2}{3} (e)^{32} (a^{13})_3,$$

$$q_{y3,33} = (e)^{33} (a^{32})_3 - \frac{2}{3} (e)^{32} (a^{33})_3,$$

$$q_{y4,11} = (e)^{41} (a^{12})_4 - \frac{2}{3} (e)^{42} (a^{11})_4 = -(q_{y3,11})_{j-1},$$

$$q_{y4,13} = (e)^{41} (a^{32})_4 - \frac{2}{3} (e)^{42} (a^{31})_4 = -(q_{y3,13})_{j-1},$$

$$q_{y4,21} = (e)^{41} (a^{11})_4 + \frac{4}{3} (e)^{42} (a^{12})_4 + (e)^{43} (a^{13})_4 = -(q_{y3,21})_{j-1},$$

$$q_{y4,23} = (e)^{41} (a^{31})_4 + \frac{4}{3} (e)^{42} (a^{32})_4 + (e)^{43} (a^{33})_4 = -(q_{y3,23})_{j-1},$$



$$q_{y4,31} = (e)^{43} (a^{12})_4 - \frac{2}{3} (e)^{42} (a^{13})_4 = -(q_{y3,31})_{j-1},$$

$$q_{y4,33} = (e)^{43} (a^{32})_4 - \frac{2}{3} (e)^{42} (a^{33})_4 = -(q_{y3,33})_{j-1},$$

$$q_{y5,11} = (e)^{51} (a^{12})_5 - \frac{2}{3} (e)^{52} (a^{11})_5,$$

$$q_{y5,12} = (e)^{51} (a^{22})_5 - \frac{2}{3} (e)^{52} (a^{21})_5,$$

$$q_{y5,21} = (e)^{51} (a^{11})_5 + \frac{4}{3} (e)^{52} (a^{12})_5 + (e)^{53} (a^{13})_5,$$

$$q_{y5,22} = (e)^{51} (a^{21})_5 + \frac{4}{3} (e)^{52} (a^{22})_5 + (e)^{53} (a^{23})_5,$$

$$q_{y5,31} = (e)^{53} (a^{12})_5 - \frac{2}{3} (e)^{52} (a^{13})_5,$$

$$q_{y5,32} = (e)^{53} (a^{22})_5 - \frac{2}{3} (e)^{52} (a^{23})_5,$$

$$q_{y6,11} = (e)^{61} (a^{12})_6 - \frac{2}{3} (e)^{62} (a^{11})_6 = -(q_{y5,11})_{k-1},$$

$$q_{y6,12} = (e)^{61} (a^{22})_6 - \frac{2}{3} (e)^{62} (a^{21})_6 = -(q_{y5,12})_{k-1},$$

$$q_{y6,21} = (e)^{61} (a^{11})_6 + \frac{4}{3} (e)^{62} (a^{12})_6 + (e)^{63} (a^{13})_6 = -(q_{y5,21})_{k-1},$$

$$q_{y6,22} = (e)^{61} (a^{21})_6 + \frac{4}{3} (e)^{62} (a^{22})_6 + (e)^{63} (a^{23})_6 = -(q_{y5,22})_{k-1},$$

$$q_{y6,31} = (e)^{63} (a^{12})_6 - \frac{2}{3} (e)^{62} (a^{13})_6 = -(q_{y5,31})_{k-1},$$

$$q_{y6,32} = (e)^{63} (a^{22})_6 - \frac{2}{3} (e)^{62} (a^{23})_6 = -(q_{y5,32})_{k-1}.$$

## Appendix 13.4 Off-diagonal viscous diffusion terms of the $z$ momentum equation

The off-diagonal viscous diffusion terms in the  $z$  momentum equation

$$\sum_{m=1}^6 \beta_m \frac{D_{l,m}^V}{\Delta L_{h,m}} \left( DI_{-w_{l,m}} - \frac{2}{3} (e)^{m3} DI_{-u_{l,m}^b} + DI_{-vis_m^{wT}} \right)$$

are computed using the same procedure as those for  $x$  equation replacing simply the subscript  $x$  with  $z$  and using the following geometry coefficients.

$$q_{z1,12} = (e)^{11} (a^{23})_1 - \frac{2}{3} (e)^{13} (a^{21})_1,$$

$$q_{z1,13} = (e)^{11} (a^{33})_1 - \frac{2}{3} (e)^{13} (a^{31})_1,$$

$$q_{z1,22} = (e)^{12} (a^{23})_1 - \frac{2}{3} (e)^{13} (a^{22})_1,$$

$$q_{z1,23} = (e)^{12} (a^{33})_1 - \frac{2}{3} (e)^{13} (a^{32})_1,$$

$$q_{z1,32} = (e)^{11} (a^{21})_1 + (e)^{12} (a^{22})_1 + \frac{4}{3} (e)^{13} (a^{23})_1,$$

$$q_{z1,33} = (e)^{11} (a^{31})_1 + (e)^{12} (a^{32})_1 + \frac{4}{3} (e)^{13} (a^{33})_1,$$

$$q_{z2,12} = (e)^{21} (a^{23})_2 - \frac{2}{3} (e)^{23} (a^{21})_2 = -(q_{z1,12})_{i-1},$$

$$q_{z2,13} = (e)^{21} (a^{33})_2 - \frac{2}{3} (e)^{23} (a^{31})_2 = -(q_{z1,13})_{i-1},$$

$$q_{z2,22} = (e)^{22} (a^{23})_2 - \frac{2}{3} (e)^{23} (a^{22})_2 = -(q_{z1,22})_{i-1},$$

$$q_{z2,23} = (e)^{22} (a^{33})_2 - \frac{2}{3} (e)^{23} (a^{32})_2 = -(q_{z1,23})_{i-1},$$

$$q_{z2,32} = (e)^{21} (a^{21})_2 + (e)^{22} (a^{22})_2 + \frac{4}{3} (e)^{23} (a^{23})_2 = -(q_{z1,32})_{i-1},$$

$$q_{z2,33} = (e)^{21} (a^{31})_2 + (e)^{22} (a^{32})_2 + \frac{4}{3} (e)^{23} (a^{33})_2 = -(q_{z1,33})_{i-1},$$

$$q_{z3,11} = (e)^{31} (a^{13})_3 - \frac{2}{3} (e)^{33} (a^{11})_3,$$

$$q_{z3,13} = (e)^{31} (a^{33})_3 - \frac{2}{3} (e)^{33} (a^{31})_3,$$

$$q_{z3,21} = (e)^{32} (a^{13})_3 - \frac{2}{3} (e)^{33} (a^{12})_3,$$

$$q_{z3,23} = (e)^{32} (a^{33})_3 - \frac{2}{3} (e)^{33} (a^{32})_3,$$

$$q_{z3,31} = (e)^{31} (a^{11})_3 + (e)^{32} (a^{12})_3 + \frac{4}{3} (e)^{33} (a^{13})_3,$$

$$q_{z3,33} = (e)^{31} (a^{31})_3 + (e)^{32} (a^{32})_3 + \frac{4}{3} (e)^{33} (a^{33})_3,$$

$$q_{z4,11} = (e)^{41} (a^{13})_4 - \frac{2}{3} (e)^{43} (a^{11})_4 = -(q_{z3,11})_{j-1},$$

$$q_{z4,13} = (e)^{41} (a^{33})_4 - \frac{2}{3} (e)^{43} (a^{31})_4 = -(q_{z3,13})_{j-1},$$

$$q_{z4,21} = (e)^{42} (a^{13})_4 - \frac{2}{3} (e)^{43} (a^{12})_4 = -(q_{z3,21})_{j-1},$$

$$q_{z4,23} = (e)^{42} (a^{33})_4 - \frac{2}{3} (e)^{43} (a^{32})_4 = -(q_{z3,23})_{j-1},$$

$$q_{z4,31} = (e)^{41} (a^{11})_4 + (e)^{42} (a^{12})_4 + \frac{4}{3} (e)^{43} (a^{13})_4 = -(q_{z3,31})_{j-1},$$

$$q_{z4,33} = (e)^{41} (a^{31})_4 + (e)^{42} (a^{32})_4 + \frac{4}{3} (e)^{43} (a^{33})_4 = -(q_{z3,33})_{j-1},$$

$$q_{z5,11} = (e)^{51} (a^{13})_5 - \frac{2}{3} (e)^{53} (a^{11})_5,$$

$$q_{z5,12} = (e)^{51} (a^{23})_5 - \frac{2}{3} (e)^{53} (a^{21})_5,$$

$$q_{z5,21} = (e)^{52} (a^{13})_5 - \frac{2}{3} (e)^{53} (a^{12})_5,$$

$$q_{z5,22} = (e)^{52} (a^{23})_5 - \frac{2}{3} (e)^{53} (a^{22})_5,$$

$$q_{z5,31} = (e)^{51} (a^{11})_5 + (e)^{52} (a^{12})_5 + \frac{4}{3} (e)^{53} (a^{13})_5,$$

$$q_{z5,32} = (e)^{51} (a^{21})_5 + (e)^{52} (a^{22})_5 + \frac{4}{3} (e)^{53} (a^{23})_5,$$

$$q_{z6,11} = (e)^{61} (a^{13})_6 - \frac{2}{3} (e)^{63} (a^{11})_6 = -(q_{z5,11})_{k-1},$$

$$q_{z6,12} = (e)^{61} (a^{23})_6 - \frac{2}{3} (e)^{63} (a^{21})_6 = -(q_{z5,12})_{k-1},$$

$$q_{z6,21} = (e)^{62} (a^{13})_6 - \frac{2}{3} (e)^{63} (a^{12})_6 = -(q_{z5,21})_{k-1},$$

$$q_{z6,22} = (e)^{62} (a^{23})_6 - \frac{2}{3} (e)^{63} (a^{22})_6 = -(q_{z5,22})_{k-1},$$

$$q_{z6,31} = (e)^{61} (a^{11})_6 + (e)^{62} (a^{12})_6 + \frac{4}{3} (e)^{63} (a^{13})_6 = -(q_{z5,31})_{k-1},$$

$$q_{z6,32} = (e)^{61} (a^{21})_6 + (e)^{62} (a^{22})_6 + \frac{4}{3} (e)^{63} (a^{23})_6 = -(q_{z5,32})_{k-1}.$$

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## 14 Visual demonstration of the method

The method presented in this monograph is implemented in the computer code IVA, Kolev (1996, 1999a, 1999b). The verification of the method is presented in the last Chapter of Volume II. The purpose of this Chapter is to demonstrate the capability of the method in simulating very complex flows – consisting of molten material, gases and water in all their possible combinations. The cases considered here are already documented and discussed in separate publications and the interested reader should go to the original sources for details, especially if he or she is interested in discussions on the physical phenomena and the conclusions for practical applications. What were not given in these publications are the videos corresponding to the discussed cases.

We will proceed here with a short description of the case considered, provide the video, and discuss briefly the methods, which are operating within the mathematical simulations. The visualization tool used here is SONJA, Kolev, Chen, Kollmann and Schlicht (1998).

### 14.1 Melt-water interactions

Next generation nuclear reactors are designed to sustain any catastrophic melt of the reactor core and to keep the accident localized inside a specially designed reactor building called containment.

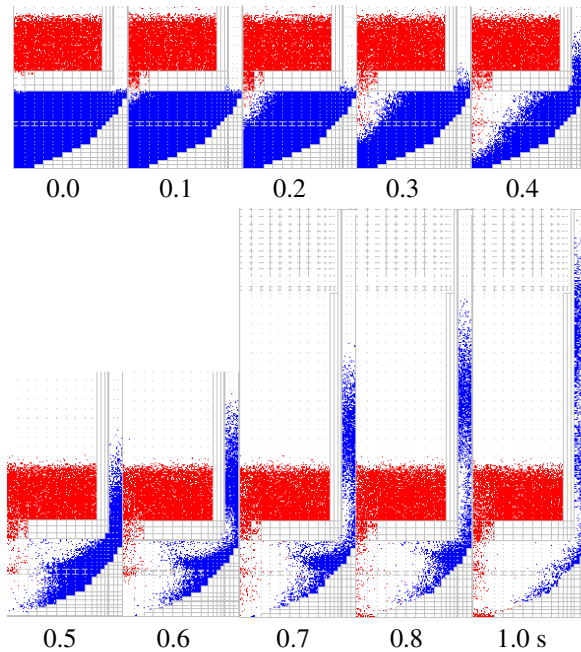
#### 14.1.1 Cases 1 to 4

In Kolev (1998a) a postulated in-vessel melt-water interaction is analyzed caused by failure of the crust above the core support plate and below the molten pool. The initial conditions are given below: The lower head is assumed to be filled with residual water, 15 t, at saturation conditions up to the lower support plate. The molten pool consisting of 101 t Corium with initial temperature 3100 K - see Fig. 14.1. The release cross sections for the first three cases are effectively 0.3267, 0.5809 and 0.9076 m<sup>2</sup>, and the release time constants are 16.43, 9.24 and 5.91 s respectively. The release cross sections are defined with an axial permeability of 0.31 multiplied by the total failed cross sections. Making use of the symmetry we simulate the process as being two dimensional by using 1424 computational cells. Note that the distribution plate within the lower head is modeled. The 15% reduction of the venting cross section caused by the 8 core barrel supports is taken into account. Case 4 is

similar to the previous case. The only difference is that the total lower support plate is assumed to be permeable with 0.31-permeability. Because cases 1 to 3 did not lead to any melt elevation at all we decided to consider the very hypothetical case 4 to see whether even under these hypothetical conditions the feared slug formation is possible. The most important results of the simulation are described below.

### 14.1.1.1 Case 1

Figure 14.1 presents the material relocation as a function of space and time. We see that after 0.1 s the melt reaches the water. The fragmentation produces larger melt interfacial area density and steam production driving the water depletion. After 1 s there is almost no water left in the lower head.

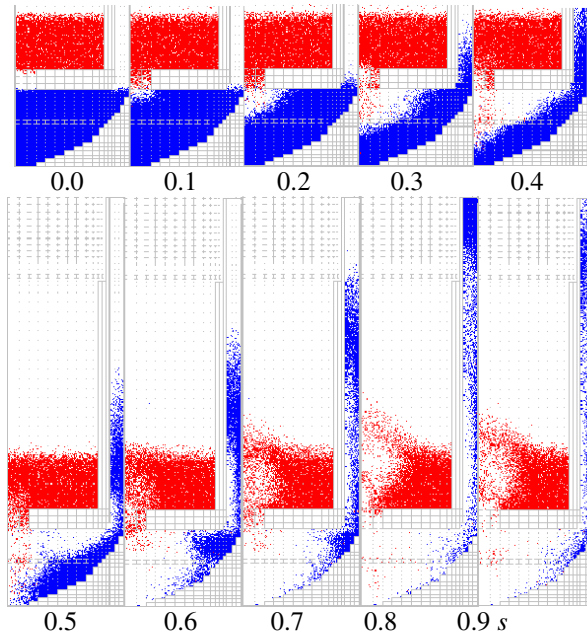


**Fig. 14.1** IVA5 EPR in-vessel melt water interaction analysis: Case 1, failure cross section  $0.3267 \text{ m}^2$ . Material relocation as a function of space and time

The pressures versus time functions at four different positions along the symmetry axis, given in Kolev (1998a), compared with the material relocation maps reveal the cyclic nature of this kind of interaction. The initial melt-water interaction is not strong enough to completely deplete the water. This occurs after the third interaction. In this case the lower head pressure was not large enough to invert the melt relocation in the cross section of the failure. The main conclusion drawn in Kolev (1998a) for this particular case is that, during the process considered, the total melt mass integrated over all cells with water



volume fraction above 0.52 (within the water - gas mixture) was at maximum several kilograms which presents no danger for RPV upper head integrity.



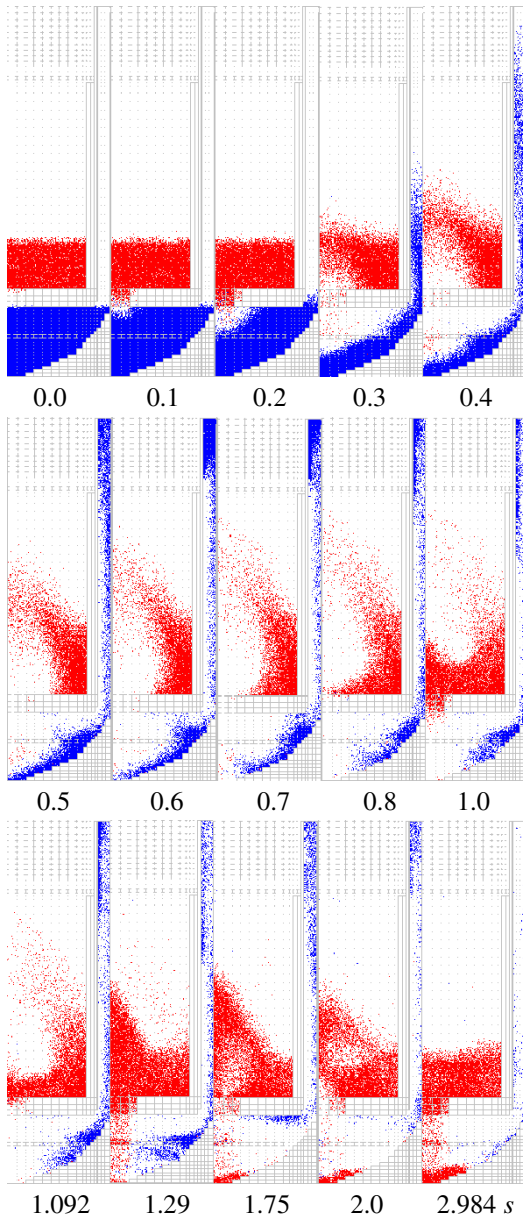
**Fig. 14.2** IVA5 EPR in-vessel melt water interaction analysis: Case 2, failure cross section  $0.5809 \text{ m}^2$ . Material relocation as a function of space and time

### 14.1.1.2 Case 2

This case is similar to case 1 with the difference that the water depletion from the lower head happens faster, as demonstrated in Fig. 14.2, where the material relocation as a function of time and space is presented.

### 14.1.1.3 Case 3

The material relocation during the transient is presented on Fig. 14.3. In Kolev (1998a) the pressure as a function of time at the symmetry axis for four different positions: bottom, below and above the flow distribution plate, and below the lower core support plate is also presented. Again, as in cases 1 and 2, the first interaction was not strong enough to invert the melt relocation. The second interaction is started if the melt reaches the flow distribution plate. The flow distribution plate obviously serves as a trigger. This is evident from the fact that the pressure spikes occur here first and then propagate in all directions. The second interaction succeeded in halting the melt relocation.



**Fig. 14.3** IVA5 EPR in-vessel melt water interaction analysis: Case 3, failure cross section  $0.9076 \text{ m}^2$ . a) Material relocation as a function of space and time; b) Melt being below the lower core support plate as a function of time. Without interaction the mass after 3 s must be about  $59 t$

The expanding steam creates RT instabilities and disintegrates the melt above the pressure source, relaxing the pressure source. Losing kinetic energy due to the action of gravity, the dispersed melt then settles down. The sideward sloshing of the remaining continuous melt is reflected from the walls. Thus, at about 1.0 *s* the melt starts to flow into the lower head again. The water was depleted from the lower head after the second interaction. Part of the water entered the primary circuit but part was reflected from the upper closure of the down-comer. At about 1.5 *s* both streams of melt and water penetrate each other and the third, most violent interaction is observed. This produces enough pressure to blow the water completely up from the lower head, but not enough to completely invert the melt stream - it stops it for about 0.3 *s*. The melt relocation then starts again and continues undisturbed.

The main conclusion drawn in Kolev (1998a) for this particular case is that the possible energy releases in this case are below the “no deformation limit” and therefore present no danger for the integrity of the containment.

We would like to emphasize the cyclic nature of this kind of in-vessel interaction. The first interaction is not always the strongest one. The next ones may be much stronger in so far as water is still available in the lower head. The water is depleted by cyclic explosions, down to the complete removal from the lower head and the down-comer. The available large venting area makes this possible.

#### **14.1.1.4 Case 4**

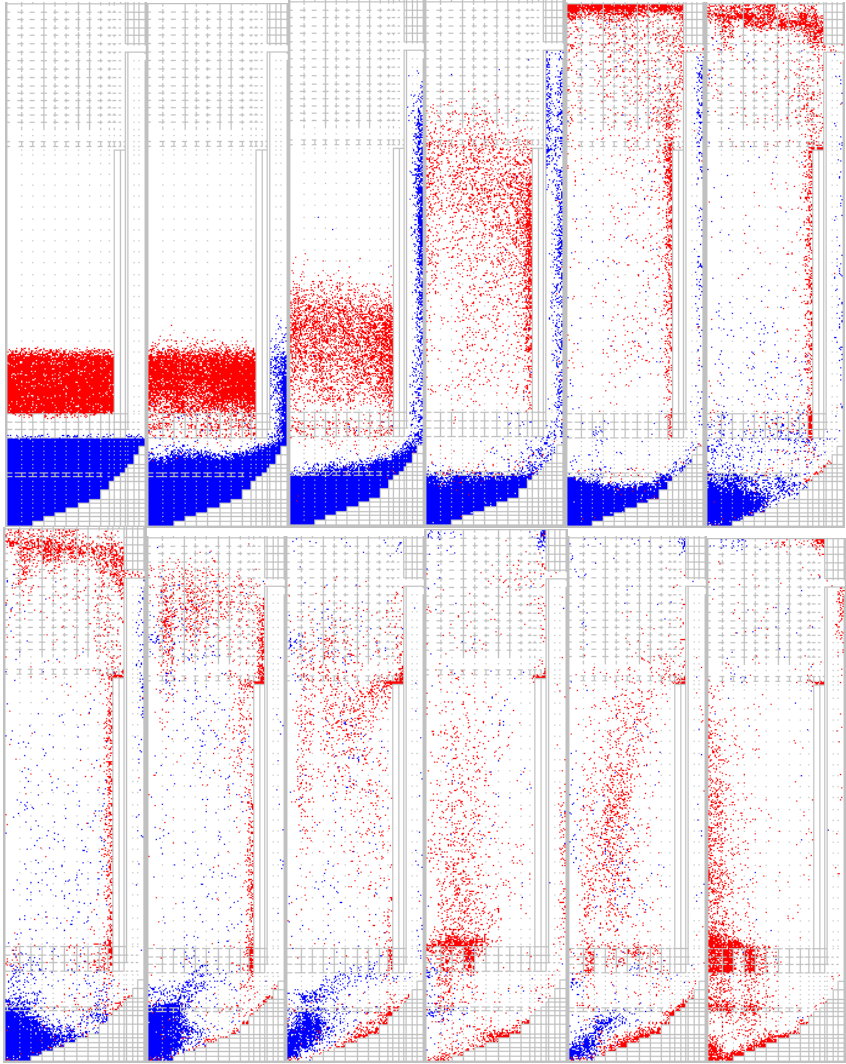
As expected, case 4 gives the most violent interaction compared to the previous three cases. The first explosion after 0.2 *s* creates enough pressure to accelerate the melt upwards. The melt reaches the top and starts to drop down. After 0.7 *s* the second explosion takes place as a result of entrapment. The second interaction produces pressure spikes of up to 25 *bar* in the region of most intensive interaction.

The reason for the violent explosion is the dramatic increase in the interfacial melt area. The result is violent acceleration of the melt upwards again. There is a third cycle of the interaction after 2.2 *s*. The melt falling down accumulates in the external region of the lower head and intermixes again with the remaining water. The result is a third pressure excursion.

Again as in the previous cases we observe cyclic interactions. Almost the total amount of water is depleted outside the reactor pressure vessel (RPV) after the third second. The melt mass inside the RPV in three-phase bubbly flow is not larger than 120 *kg*. The melt inside the RPV that is in “stoichiometric” three-phase flow is after 0.25 *s* in its major part of the in dispersed droplet-droplet state.

#### **14.1.1.5 Model elements addressed in cases 1 to 4**

1. All conservation equations for field 1 and field 3 including particle number conservation equations;
2. Geometry description using direction-permeabilities;



**Fig. 14.4** IVA5 EPR in-vessel melt water interaction analysis: Case 4, failure cross section-lower support plate with permeability of 0.31. Material relocation as a function of space. Parameter - time: 0, 0.182, 0.256, 0.502, 0.893, 1.360, 1.515, 1.903, 2.206, 2.279, 2.500, and 2.993

3. Numerical method resolving short pressure spikes and complex material relocations as presented in Kolev (1998a);
4. Constitutive relations for mass, momentum, and heat transfer for bubbles, droplets and molten particles;
5. Constitutive relations for radiation interaction between very hot melt and surrounding water-steam-air mixture;
6. Constitutive relations controlling the dynamics fragmentation and coalescence.

#### **14.1.1.6 Available file on CD**

*fig\_14.04.gif, fig\_14.04.html*

#### **14.1.2 Cases 5, 6 and 7**

In Kolev (1998b) an ex-vessel melt-water interaction caused by symmetric lower head unzipping within the reactor pit was analyzed. The melt-water interaction outside the reactor pressure vessel (RPV) in a typical pressurized water reactor (PWR) during a severe accident is associated with the particular geometry of each reactor system and the applied accident management strategy. If melt is relocated in the lower head (LH) the nature of the buoyancy driven convection causes the largest heat flux at the circumferential level of the melt surface. The heat flux decreases with decreasing arc distance measured from the lowest point of the vessel. In this situation pressure inside the vessel or the stress caused by the weight of the melt itself causes circumferential failure Kolev (1998b). It is interesting to know what will happen if water is present at the time of lower head failure caused by melt attack for typical PWR. For this purpose we selected the following two hypothetical cases:

*Case 5 Lower head unzipping and relocation in cavity partially filled with water. No water is available in the RPV.*

*Case 6 Lower head unzipping and relocation in dry cavity. Water is available in the RPV.*

In case 5 the melt is relocated into the lower head, the region around the reactor is free of water but there is some residual water in the pit below the reactor. Without external cooling ablation takes place in this plane within 10 to 20 *min*. The result is unzipping of the lower head. The lower head relocation due to gravity is limited by the bottom of the pit. The water in the pit is displaced upwards. Part of the water comes into contact with the melt exercising large scale relocation and surface instabilities caused by the impact.

In case 6 injection of water takes place after the melt relocation into the lower head. The injection does not stop the ablation process and the unzipping occurs just as in case 5. The lower head falls into the dry pit. The accelerated liquids are the melt and the water, the later being in a film boiling above the melt. The large scale melt relocation (sloshing) process starts and stratified melt-water interaction takes place.

### **14.1.2.1 Initial conditions**

We specify the following initial conditions for the two cases:

**Case 5** In this case the lower head with the melt inside it is accelerated due to gravitation only up to 5.3 *m/s* velocity before the lower head reaches the pit bottom. The lower head contains 104 *t* oxide melt. The mass of the water in the reactor pit is 53 *t*. The temperature of the water is equal to the saturation temperature at pressure 1 *bar*, which is assumed as initial system pressure.

**Case 6** In this case the lower head with the melt inside it is accelerated due to gravitation only up to 5.3 *m/s* velocity before the lower head reaches the pit bottom. The lower head contains 104 *t* melt. The mass of the water inside the RPV is 27 *t*. The temperature of the water is equal to the saturation temperature at pressure 1 *bar*, which is assumed as initial system pressure. The volumetric fraction of the steam inside the water is 0.3.

### **14.1.2.2 Driving instability for Case 5 and 6**

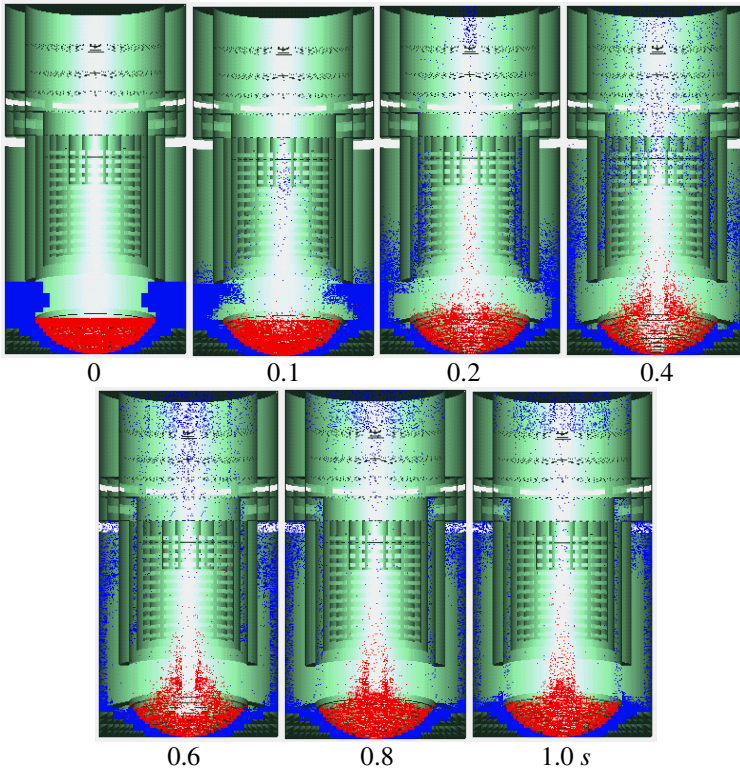
At the moment at which the accelerated lower head hits the bottom of the pit the stagnation at the walls create pressures which are reflected to the melt surface. The surface acts as reflector again and so pressure waves interfere, resulting in a complicated pattern in the space-time domain. The manifestation of the physics is clearly seen from Fig. 1 in Kolev (1998b) in which the pressure at the lowest point is presented. Pressure spikes up to 880 *bar* are computed. The frequency of 750 *Hz* is characteristic for the particular lower head geometry. The pressure waves are the reason for small- and large-scale melt motion. The large-scale motion causes surface instability, which is partially resolved by the code numeric. The small-scale instabilities are partially resolved by the fragmentation models. Just after the interpenetrating of some fragmented melt and water occurs, the triggering happens. The local triggers are inherently predicted by the code. No artificial trigger is applied. The system possesses its internal capabilities for triggering explosions.

### **14.1.2.3 Material relocation**

#### **Case 5**

In this case the melt-water contact starts at the corner of the relocated LH - see Fig. 14.5.

Then the toroidal pressure wave presses the melt in the corner downwards. Due to mass conservation, the melt erupts in the symmetry axis. The toroidal pressure wave accelerates the water upward in the pit. Thus we observe that the path of the melt outside the vessel is practically blocked and melt relocation into the containment due to this event is not possible. The upward melt relocation does not have enough momentum to destroy the upper RPV head. Losing its momentum the melt drops down and starts to gather in the lower head. The large amount of



**Fig. 14.5** An IVA5 simulation of PWR ex-vessel melt-water interaction. Case 1: Lower head unzipping and relocation in cavity partially filled with water. No water is available in the RPV. Material relocation after melt-water interaction

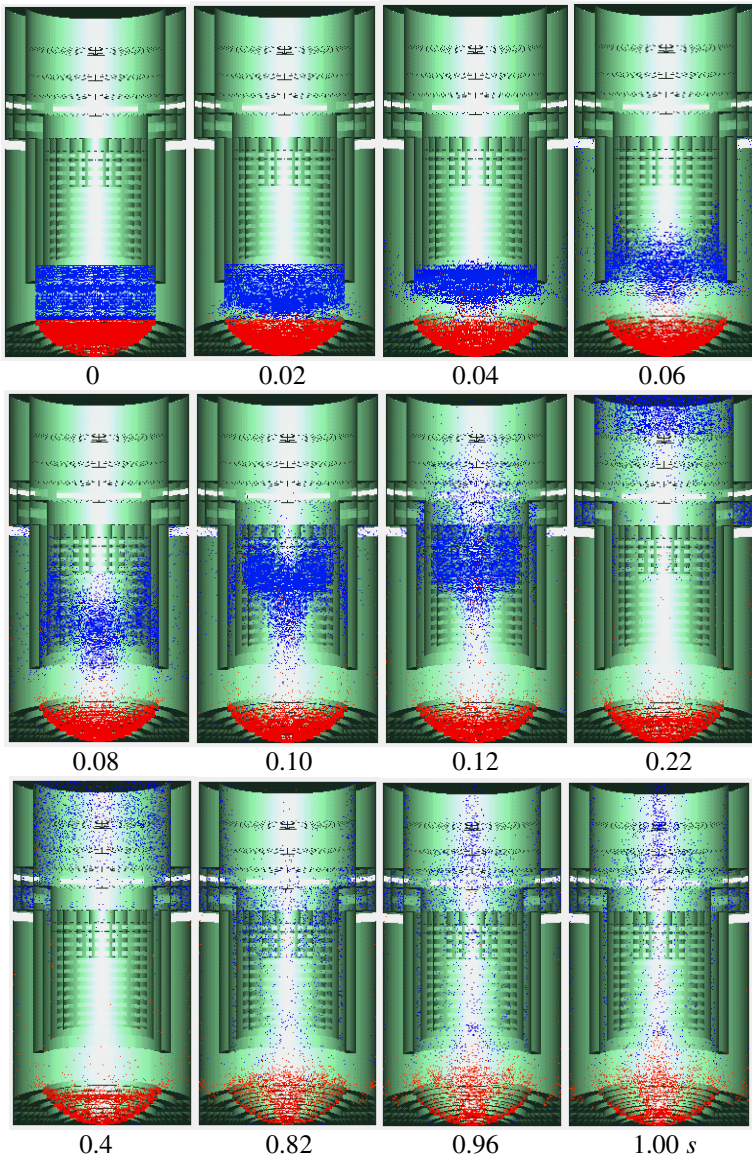
the water initially above the melt level is depleted. The remaining water will cyclically receive portions of melt due to melt sloshing and be depleted after one or two additional cycles.

The most jeopardized region is the pit wall around the region between the LH and the remaining RPV. The peripheral melt-water interaction in this case reflects the melt into the internal region of the reactor, thus limiting the melt depletion in the containment.

### Case 6

This case is characterized by more violent interaction than in case 5. The inertia of the water holds the water for some time in an intensive unstable contact with the melt. The water above the melt serves as a constraint for pressure build before the expansion phase. The contact surface is much larger than in case 5. That is why the thermal energy transferred into evaporation is much larger than in case 5. The melt sloshing due to the instability, as previously described, provides additional penetration of melt into the water due to surface instabilities considerably accelerated by the local thermal interactions. It seems that the most intensive interaction starts at one





**Fig. 14.6** An IVA5 simulation of PWR ex-vessel melt-water interaction. Case 2: Lower head unzipping and relocation in dry cavity. Water is available in RPV. Material relocation after melt-water interaction

half of the LH melt surface radius. This is evident from the pressure-time histories plotted in Kolev (1998b) for three different points over the radius at the level of the initial melt surface. The pressure wave rises first in the corner and then propagates towards the axis of the RPV being strongly amplified at a half of the radius. We observe two interactions, the first within the first 0.015 s being much shorter than the



second. The interaction is so violent that within about 0.2 s all the water is 10 m above its initial state - see Fig. 14.6. The water fragments during the flight and reaches the upper head in a droplet form. The impact with the upper head leads to the pressure increase. This impact does not lead to upper head failure and missile formation.

The explosion disperses not only the water but the melt too. After about 0.8 s the water falls down and interacts with the melt again. The kind of interaction is very interesting: It resembles two clouds of different materials penetrating into each other. Intensive evaporation causes local pressure to again increase, which again accelerates the remaining water upwards.

In general this mode of material relocation possesses some potential for depleting some melt into the containment.

Again as in case 5 the most jeopardized region is the pit wall around the region between the LH and the remaining RPV. In Kolev (1998b) the pressures in this region as a function of time are shown. The maximum pressure at the vertical wall inside the pit is about 17.5 bar (triangle, rise time 0.03 s, decrease time 0.15 s) decreasing as one moves upwards to 12.5 bar.

**Case 7** is the same as case 6 except the initial pressure which was 20 bar. In the attached movie to case 7 the comparison between large scale relocation of the lower head and small relocation of the lower head is demonstrated. Small scale relocation of the lower head leads to considerable low melt dispersion outside the vessel.

One of the important conclusions drawn in Kolev (1998b) from studying cases 5 and 6 is that no significant loads result in this case for the RPV upper head. Missile formation jeopardizing the containment integrity is not possible.

#### **14.1.2.4 Model elements addressed in cases 5, 6 and 7**

1. The same model elements as in cases 1 to 4;
2. In addition, the driving instability is not only the gravitational acceleration as in cases 1 to 4 but the deceleration due to the impact of the lower head with the bottom.

This case is extraordinarily complicated for the numerics and is recommended as a test for numerical models that will be developed in the future.

#### **14.1.2.5 Available files on CD**

*figs\_14.05\_14.06.gif, figs\_14.05\_14.06.html*  
*case\_14.07.gif, case\_14.07.html*

#### **14.1.3 Cases 8 to 10**

In Kolev (2000a) three-dimensional cases of melt-water interactions are presented.

We postulate two cases 8 and 9 in which a molten sea is formed within the destroyed core of a modern PWR. In both cases the process is modeled as three dimensional. We make use of the existing symmetry and consider only 45° of the

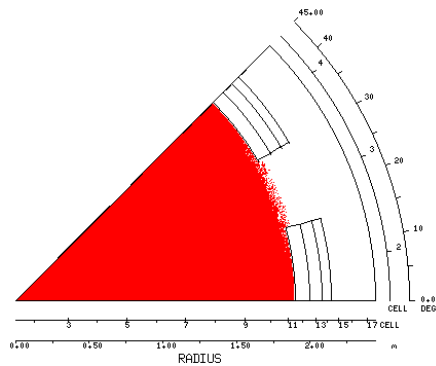
horizontal cross section. Both cases differ from each other in the elevation of the molten sea and in the side of the release slide as follows: *Case 8*: 15° slide, about 100 t of melt pool, bottom position. We call this case *inertialess* release. *Case 8*: 5° slide, 153 t of melt pool, elevated position. We call this case *inertial* release.

While in case 8 the melt cannot reach considerable penetration velocity because of the short distance to the water pool, in case 9 the melt is considerably accelerated. This causes considerable differences in the physical behavior of the system.

Note that in the case 7 the release is considered to happen simultaneously from 8 symmetric positions in the RPV and in case 8 from 4.

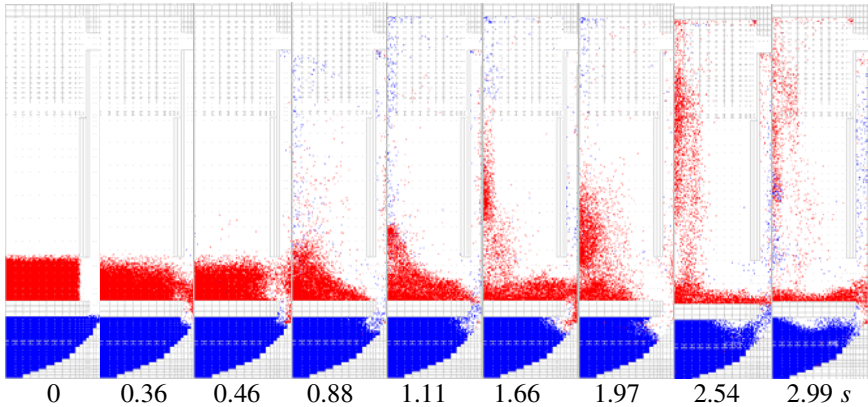
### 14.1.3.1 Case 8

Figure 14.7 shows the initial conditions in a horizontal plane crossing the molten pool. We see also the radial discretization used in the computational analysis.



**Fig. 14.7** IVA5 PWR in-vessel melt water interaction analysis: Case 8, side failure of the heavy reflector, molten pool at the lowest position. Initial conditions and geometry in a plane crossing the molten pool

In Fig. 14.7 the material relocation as function of time and space is presented. The discretization in the vertical plane used for the computational analysis is also visible. Reading Fig. 14.8 together with the pressure histories presented in Kolev (2000a) we realize the following behavior: Again as in the lower support failure cases 1 to 4 the processes are cyclic starting with small pressure increase leading to small expansion, collapse and steam explosion due to entrapment, characterized by the second sharp pressure peak. The explosion causes material relocation. The descending melt gathers again in the pool and causes a third violent pressure excursion producing the largest pressure inside the RPV. Whilst in cases 1 to 4 the water is completely depleted after 3 s in case 7 from the initial amount of about 15 t, 11 t are steel in the RPV after the third second. We observe that the inertialess side penetration of melt into the water in the lower head is very difficult and causes very low water depletion. The reverse injection of the water-steam mixture into the molten pool is possible due to the assumed large release cross section of the slice. This is the reason again for large-scale melt dispersion and depletion. After the third second from the initial mass of about 100 t, 50 t are depleted into the primary circuit.

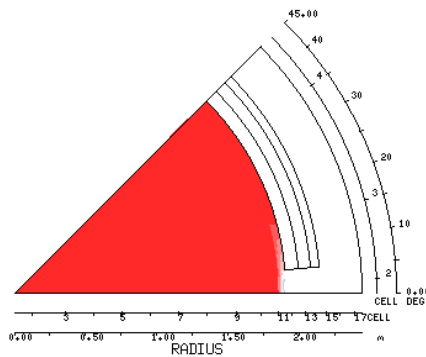


**Fig. 14.8** Material relocation as a function of space and time

Note that there are three different geometrical regions: lower head, down comer and the empty space inside the core region. The melt is being dispersed through all of these regions. A coherent explosion influencing all regions in a sense that effective chain fragmentation happens simultaneously is impossible. The so called *premixing mass* that is the melt mass being in bubble three-phase flow is of particular interest. Only in bubble three-phase flow we have the most effective heat transfer mechanism from melt to liquid. During the all transient maximum 120 kg of melt are occasionally being in bubble three-phase flow. Thus, again we conclude that there is not enough available energy as a pressure source which can cause failure of the upper head endangering the containment integrity.

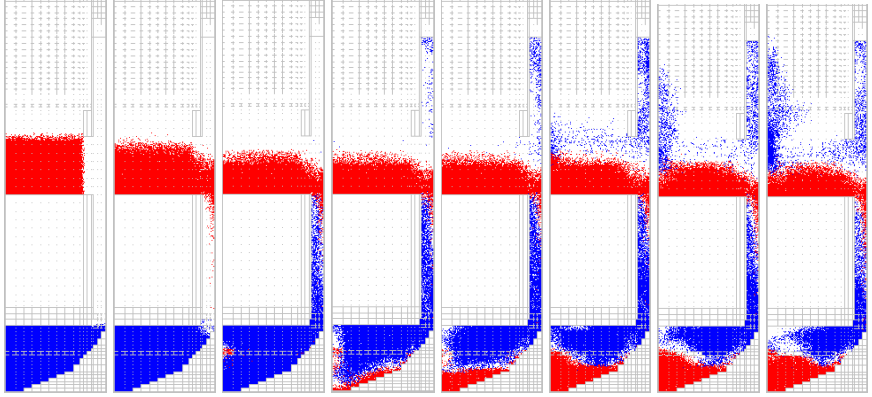
### 14.1.3.2 Case 9

Figure 14.9 shows the initial conditions in a plane crossing the molten pool.

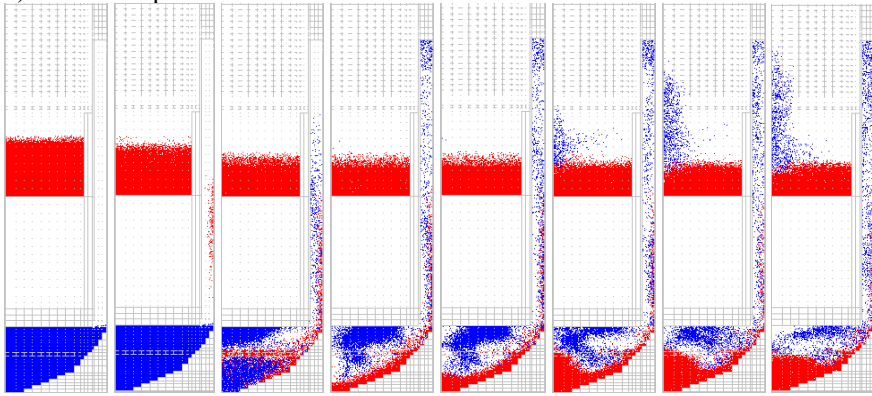


**Fig. 14.9** IVA5 PWR in-vessel melt-water interaction analysis: Case 9, side failure of the heavy reflector, molten pool at the elevated position. Initial conditions and geometry in a plane crossing the molten pool

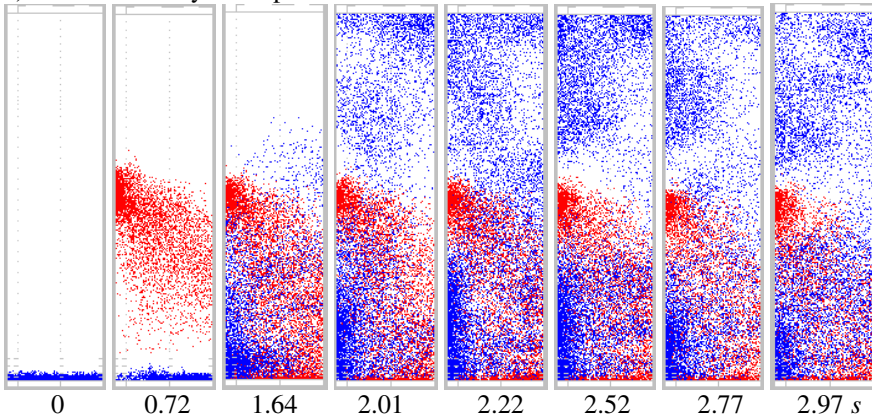
a) Failure plane



b) Outermost plane



c) Down-corer cylinder plane



0 0.72 1.64 2.01 2.22 2.52 2.77 2.97 s

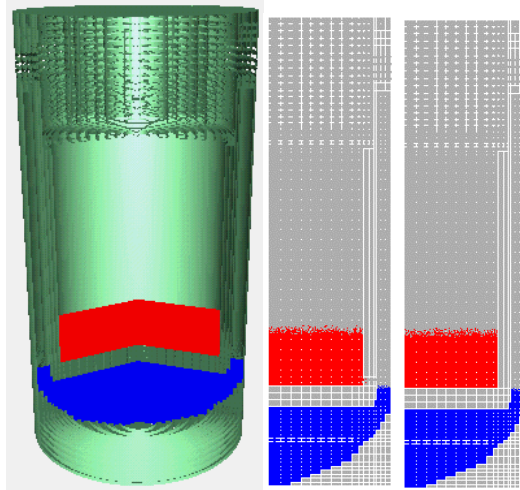
Fig. 14.10 Material relocation as a function of space and time

We see also the radial discretization and the three angular sectors used in the computational analysis. Figure 14.10 shows the discretization in the vertical plane used for the computational analysis. In Fig. 14.10 the material relocation as function of time and space is shown. The material relocation is presented for three different planes. The first one is the vertical plane of the failure slice. The second one is the outermost vertical plane. The third plane is in fact parallel to the cylindrical wall of the down-comer - the outermost layer.

We observe here very complex three-dimensional flow. The melt jet is disintegrated shortly after impacting the RPV wall. After the first melt reaches the lower head, the sloshing of the water in the failure plane is so strong that the dispersed melt penetrates the water very easily in the non-failure planes - see Fig. 14.10. As seen from Fig. 14.10 a large space of the down-comer is occupied by dispersed melt-water-steam flow. If one observes the movies produced with this material, one recognizes concentration water droplet waves cyclically passing the dispersed melt. Because of the acceleration of the falling melt the water expulsion and later interactions do not force elevation of the melt. Thus, the depletion of the melt into the primary circuit is limited in this case. The pressures curves in Kolev (2000a) show very noisy interaction in the lower head. The lower head is filled with melt. Again as in case 8, there are three different geometrical regions where melt and water coexist: lower head, down comer and the empty space inside the core region. This melt mass is dispersed through all of these regions. A coherent explosion influencing all regions in a sense that effective chain fragmentation happens simultaneously is impossible. There is a maximum of about 1.5 t premixed melt mass. We see that from all analyzed situations 1 to 7 this is the one with the largest damage potential. The large amount of potentially explosive mixture is in fact found in the lower head. The mixture in the down-comer is rather dispersed.

### **14.1.3.3 Case 10**

In this case we consider - *45° slice, about 120 t of melt pool, bottom position*. In cases 8 and 9 we have considered melt release from the side of the heavy reflector having very pessimistic initial conditions of simultaneous release from 4 or 8 positions. In this section we consider release from one spot located at the bottom of the molten pool as presented in Fig. 14.11. As can be seen from Fig. 14.11 and Fig. 14.12, the process is asymmetric and has to be considered as a complete three-dimensional problem.



**Fig. 14.11** a, b, c IVA5 model of case 9. Discretization scheme for case 10 in two vertical planes: a)  $j = 2$ , b)  $j = 6$

**Geometry:** The calculation has been performed in a three-dimensional cylindrical geometry. The IVA computational model consists of  $16 \times 8 \times 69 = 8\,832$  computational cells. These computational cells are denoted with indices  $(i, j, k)$ , where  $i$  stands for the cell index in radial,  $j$  for azimuthal and  $k$  for axial direction. Schematic plots of the discretization are given in Figs. 14.11 and 14.12. Note that the lowest cell indices in all three directions have the number 2.

**Initial conditions:** The initial conditions are specified as follows: The initial system pressure is assumed to be  $3\text{ bar}$ , the system temperature is equal to the corresponding saturation temperature. We assume about  $120\text{ t}$  of oxide melt inside the heavy reflector and about  $17\text{ t}$  of saturated water inside the lower head (see Fig. 14.13). The remaining part of the computational domain is filled with air.

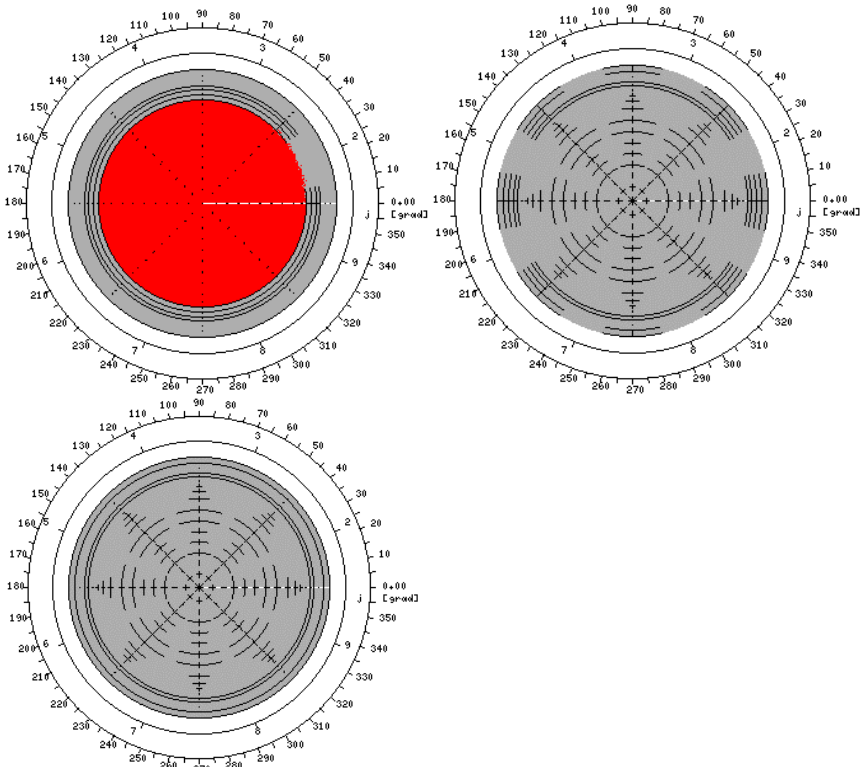


Fig. 14.12 Cross sections at different heights: a)  $k = 26$ , b)  $k = 62$ , c)  $k = 70$

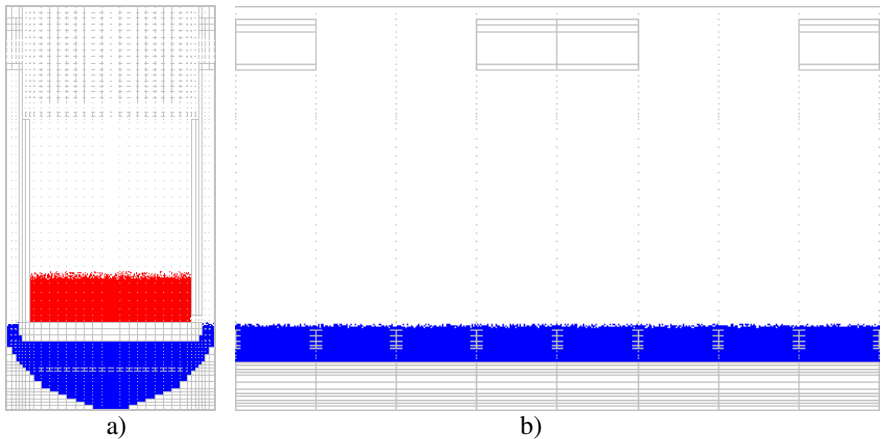
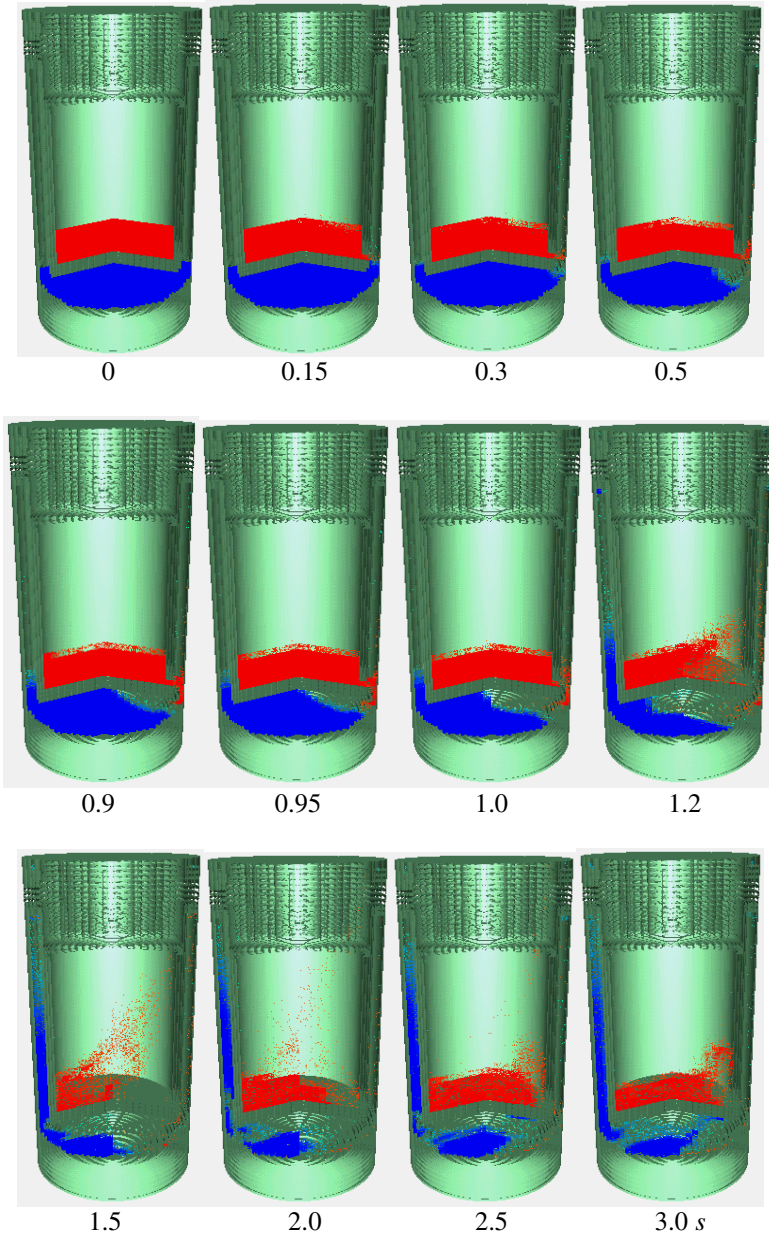


Fig. 14.13 Initial conditions for location of melt and water. a) Cross section in  $r$ - $z$  plane. b) Surface close to the internal wall of the vessel



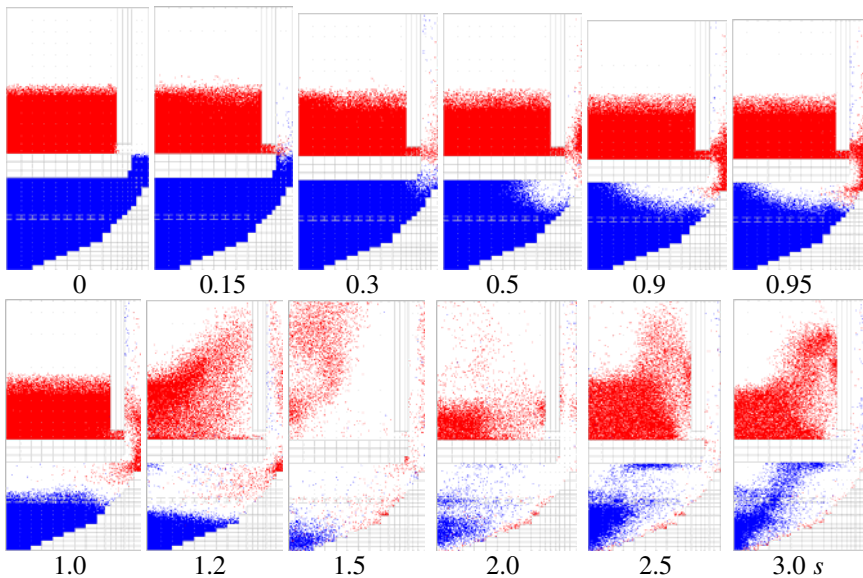


**Fig. 14.14** Material relocation at different times (planes  $j = 2$  and  $j = 6$ )

**Results:** The computational results are presented as time functions for pressures at different positions in Kolev (2000a), as maps of the material relocation, Figs. 14.14 to 14.17, and as time functions of the total mass being in bubbly flow in Ko-

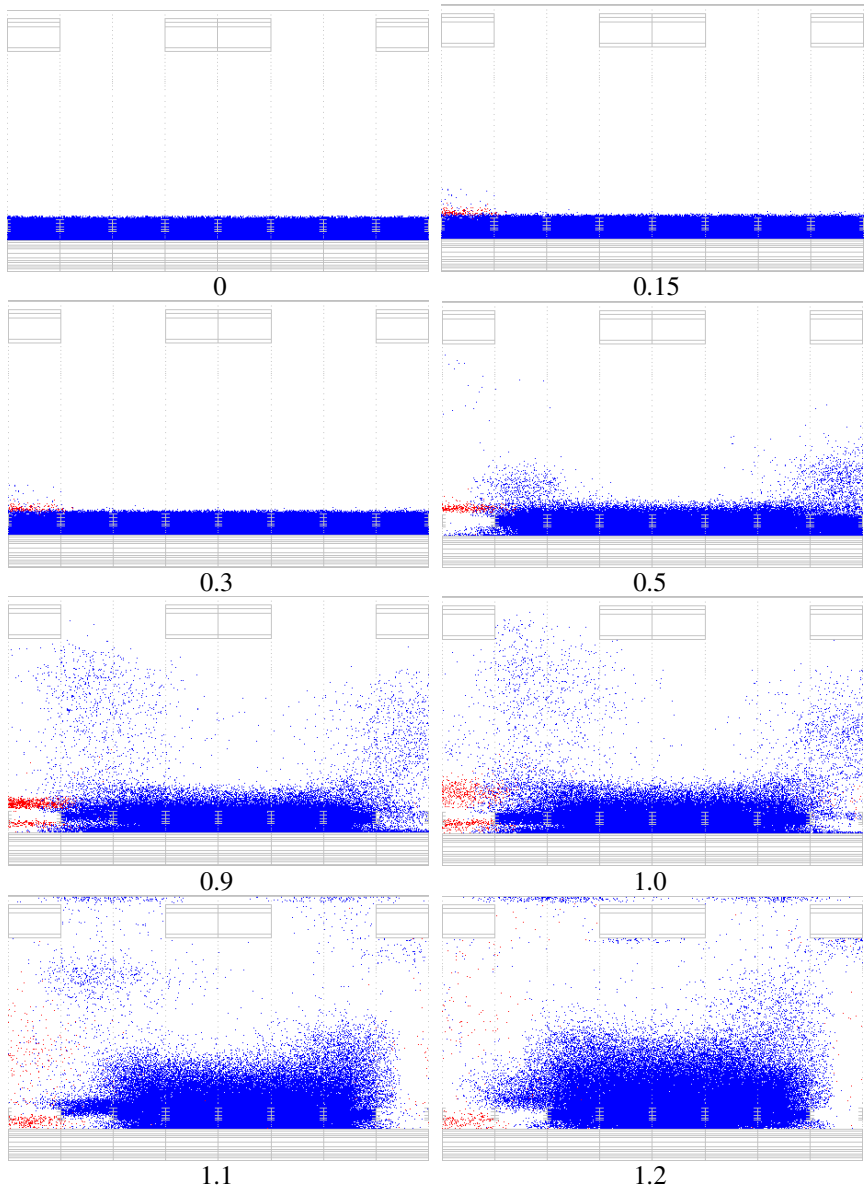


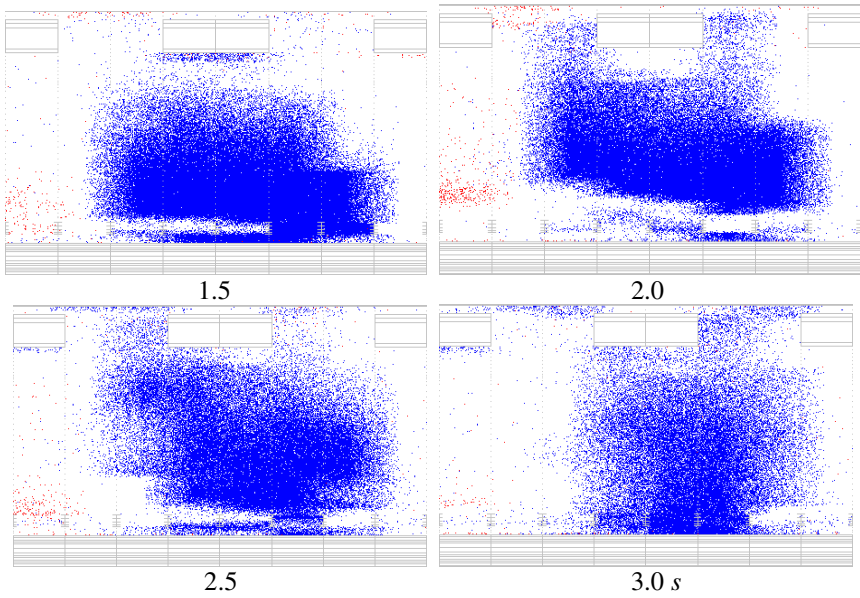
lev (2000a). We observe from the pressure time histories that the first contact of the melt with water happens after about 0.15 s. It produces a pressure peak of about 5.2 bar. Thereafter melt intermixing with water takes place for up to about 1 s. At that time explosive events lead to strong local pressure changes with time, followed cyclically by two other weaker interactions. Pressure peaks of up to 12.4 bar are observed. The comparison of the pressure time histories with the material relocation pictures reveals that, after the first melt-water contact happens, the interactions locally produce such a pressure that causes melt accumulation immediately after the release. Just after the pressure release the accumulated melt drops down into the water and creates an impact intermixing which leads to explosive interactions. The result is acceleration of the water into the depletion cross sections. The local pressure increase causes also a stop of the melt release. Moreover, it blows multi-phase mixture into the melt through the opening cross section in the heavy reflector. The melt contains a certain amount of water. This is the reason for internal interaction inside the molten pool leading to local dispersion. Due to the limited amount of water content in the mixture the interaction is not strong enough to accelerate the melt up to the top. After some time the melt settles down and starts again to flow continuously through the failure cross section (see Fig. 14.14).



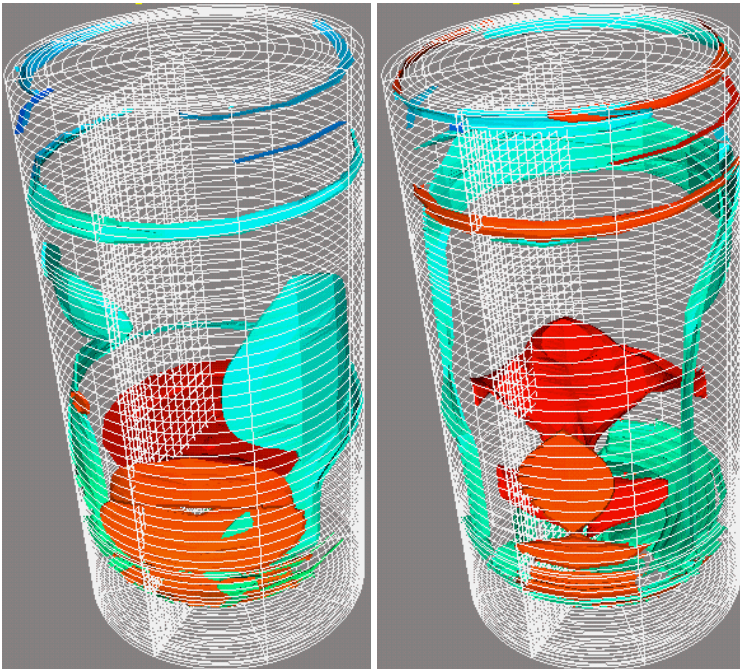
**Fig. 14.15** Material relocation in the lower head at different times (plane  $j = 2$ )

The asymmetric behavior of melt and water can be clearly seen in Fig. 14.16, where two planes are shown, one through the sector containing the leak and the other through the diametrically opposed sector.





**Fig. 14.16** Material relocation in the peripheral direction at different times (plane  $i = 15$ )



**Fig. 14.17** Material relocation – melt and water iso-surfaces, 1.07, 1.5 s

How the water behaves, is illustrated by comparing Figs. 14.15, 14.16 and 14.17. On the side of the melt ingression the melt-water interaction pushes the water away. The intermixed masses are only about 300 kg, so that the cyclically released mechanical energies are not sufficient for complete water depletion from the reactor pressure vessel. As seen from Fig. 14.16 a massive bulk of water is subject to azimuthal concentration waves. The water entrainment happens mainly due to fragmentation and mechanical steam-water interaction. The asymmetric geometry, as is clearly seen from Fig. 14.17, causes asymmetric behavior in the down comer. No more than 1.3 t of melt is able to release its energy.

Finally, the pressurization events are not amplified but are essentially damped while propagating up to the upper head - as can be seen in the pressure time histories presented in Kolev (2000a).

#### **14.1.3.4 Model elements addressed in cases 8 to 10**

1. The same model elements as in cases 1 to 7;
2. In addition, the numerical method is challenged for all three dimensions.

Again these cases are extraordinarily complicated for the numeric and are recommended as a test for numerical models that will be developed in the future.

#### **14.1.3.5 Available files on CD**

*fig\_14.08.gif, fig\_14.08.html*  
*fig\_14.09.gif, fig\_14.09.html*  
*fig\_14.10.gif, fig\_14.10.html*  
*fig\_14.10.gif, fig\_14.10.html*  
*fig\_14.10.gif, fig\_14.10.html*

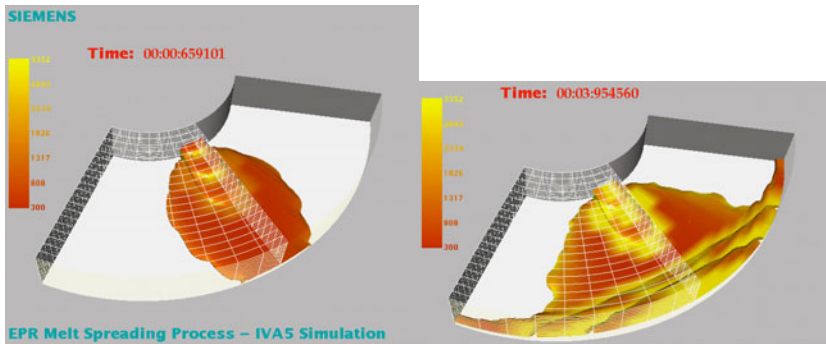
#### **14.1.4 Cases 11 and 12**

In Kolev (2000a) two cases are presented for ex-vessel melt spreading into a compartment with a flat horizontal surface, one without a water layer and one with a water layer.

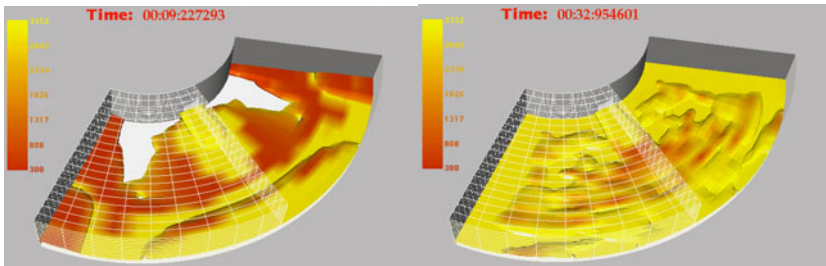
##### **14.1.4.1 Case 11**

We simulated gravitational sloshing of 300 t of non-modified melt at 3000 K in a dry compartment comprising a 110° sector with horizontal spreading surface. Note that for this melt the dynamic viscosity of 0.0044 kg/(ms) has no influence on the spreading process, as in this particular case the spreading is inertia dominated. The melt enters the computational region at the axis of the spreading compartment. As the problem is symmetric, only half of the geometry is simulated. Figs. 14.18 and 14.19 shows the discretization net and the surface dividing volumetric concentration of the melt being less and larger than 0.005. The result shows that the spreading is turbulent and heterogeneous. Two symmetric vortices are formed within 20

s. In the center of the vortices there are regions which are not reached by the melt in this initial period. The melt covers the total area in about 30 s. Gravitational surface waves travel for some time after this.



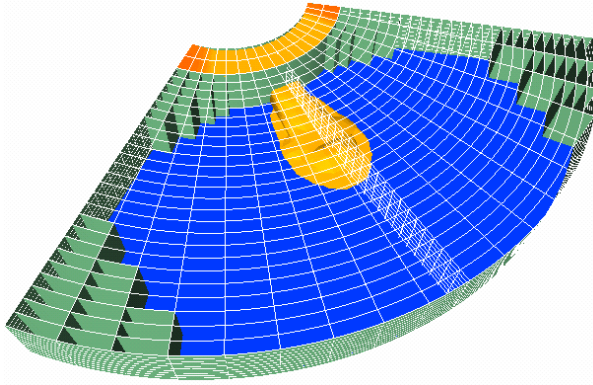
**Fig. 14.18** An IVA5 prediction of the dry melt spreading dynamics in a EPR spreading compartment. Iso-surface of the melt volume fraction with overlaid melt temperature at the surface



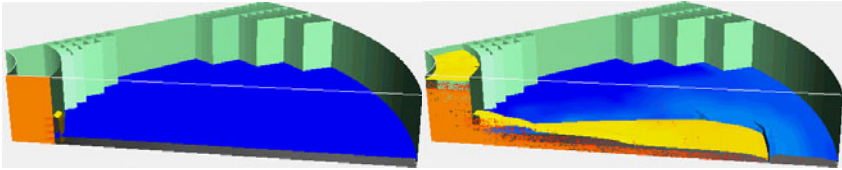
**Fig. 14.19** (Cont.)

#### 14.1.4.2 Case 12

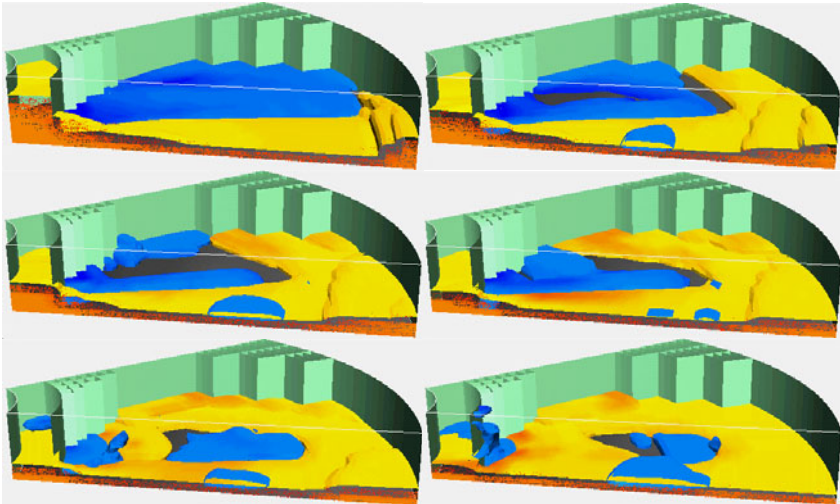
In this case we analyze the spreading using similar initial and boundary conditions as in the previous case with the only difference that the floor is covered with 1 cm water and the geometry is slightly different.



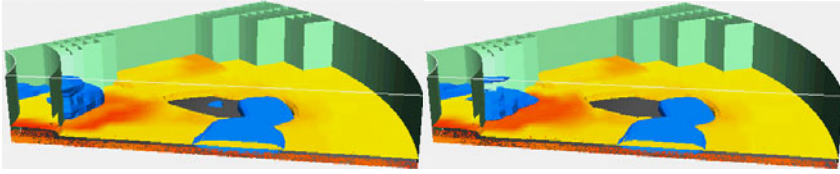
**Fig. 14.20** IVA5 discretization model for simulation of melt spreading in presence of water



**Fig. 14.21** An IVA5 prediction of the wet melt spreading dynamics in a EPR spreading compartment. Iso-surface of the melt volume fraction with overlaid melt temperature at the surface. Iso-surface of the water volume fraction with overlaid water temperature at the surface - blue. Spreading of melt in presence of thin water layer (1cm). Parameter: Time 0, 2, 3.8, 6.2, 8, 10, 14.4, 16.8, 17.8, 18 s







**Fig. 14.22** (Cont.)

Figures 14.20 through 14.22 present the melt-water interaction in this case. The melt and the water are presented with iso-surfaces dividing volume fractions less than and larger than 0.5%. In Kolev (2000a) the pressures at two different places in the spreading compartment are presented. The pressure spikes are characteristic for inertial water entrapment. The discretization used in this case is not fine enough to resolve such events with large confidence.

#### **14.1.4.3 Model elements addressed in cases 11 and 12**

1. The same model elements as in cases 1 to 10;
2. In addition, the numerical method is challenged for modeling of stratification of multiple materials.
3. The iso-surfaces possess as an attribute the temperature of the constituent. The thermal radiation and temperature reduction of the surface are clearly visible.

#### **14.1.4.4 Available files on CD**

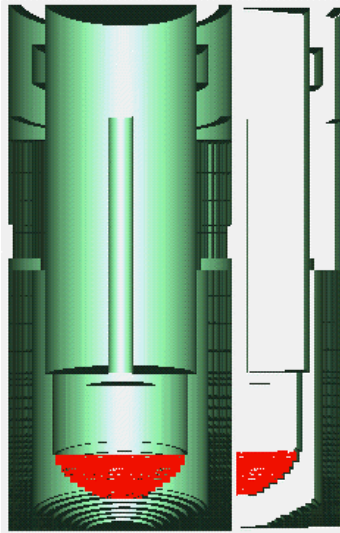
*fig\_14.18.gif, fig\_14.18.html*  
*fig\_14.20.gif, fig\_14.20.html*

### **14.1.5 Case 13**

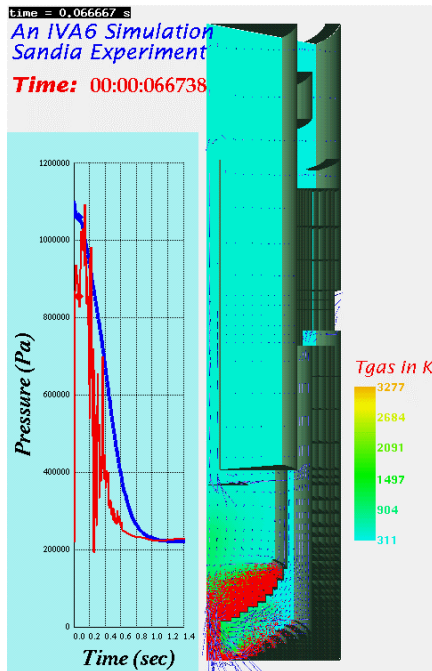
Figures 14.21 and 14.22 demonstrate a vessel filled with inert gas, hydrogen and molten material that is discharged in the container air atmosphere through a defined opening at the lower heat. The pressure in the vessel and in the stagnation point of the jet is presented also in the Fig. 14.22.

#### **14.1.5.1 Model elements addressed in case 13**

1. Fragmentation and the associated thermal interaction between the inert gas and the molten material;
2. Burning of hydrogen during the process.



**Fig. 14.23** A high pressure discharge of melt into a cavity, geometry (Sandia National Laboratory), initial conditions



**Fig. 14.24** A high pressure discharge of melt into a cavity.

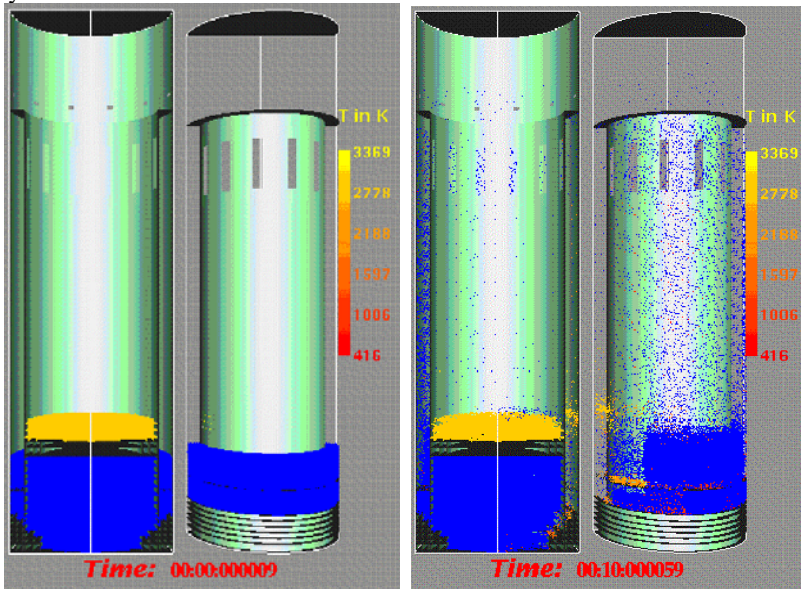


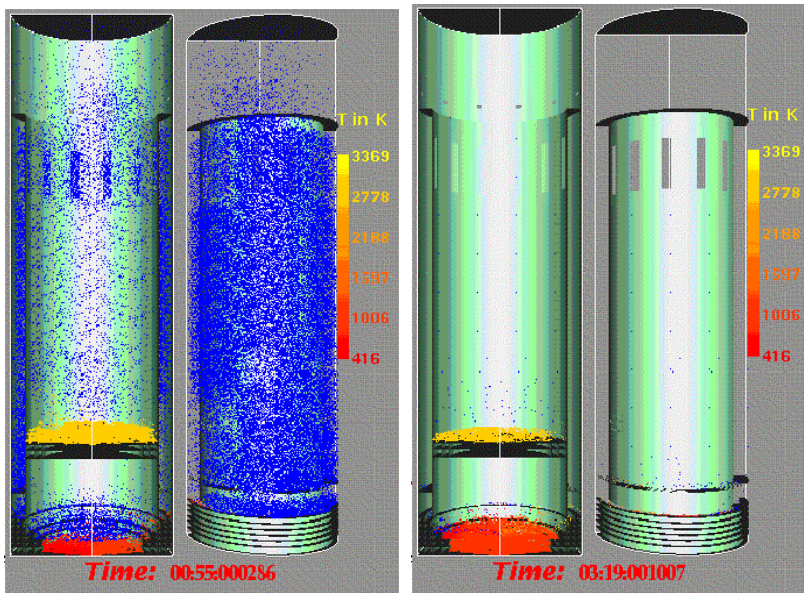
### 14.1.5.2 Available files on CD

*fig\_14.22.gif, fig\_14.22.html*

### 14.1.6 Case 14

In Kolev (2001a) severe accident control for a large-scale boiling water reactor through in-vessel melt retention by external reactor pressure vessel cooling is considered. Side release of melt from a postulated molten pool in the core region is analyzed.





**Fig. 14.25** Material relocation in the reactor pressure vessel as a function of time and space.  $0.18 \text{ m}^2$  melt release cross section. Legend: blue – water, red – melt, transparent – gas

We assume a vertical opening of the core shroud wall with the height of  $0.9 \text{ m}$  and horizontal size of  $20 \text{ cm}$  which gives  $0.18 \text{ m}^2$  cross section. The discretization contains  $(8 \times 6 \times 29) = 1392$  cells for the half of the reactor vessel. We assume  $1.6 \text{ t}$  steam within the reactor vessel with temperature  $419.15 \text{ K}$ ,  $124 \text{ t}$  water with temperature  $417.75 \text{ K}$ ,  $186 \text{ t}$  melt with temperature  $3000 \text{ K}$  – see Fig. 14.23. The initial pressure is  $4.1 \text{ bar}$ . We assume pressure  $4.1 \text{ bar}$  at the 6 safety valve lines that are open and remain open.

#### 14.1.6.1 Results of the computational analysis

We present in Fig. 14.23 the material relocation as a function of the time and space. In Kolev (2001a) the pressures as functions of time in the down-comer periphery below the initial water level, the pressures as functions of time in the lower head at different positions, and the pressures as functions of time in the upper head at different positions are presented. Comparing the material relocation with the pressure history we reveal several interesting phenomena. The process requires  $200 \text{ s}$  to come to the natural end at which there is no more water to interact with the melt any more. This is much longer than the characteristic time of  $1$  to  $3 \text{ s}$  for melt-water interactions in PWR's – see in all previously discussed cases. During this time the melt releases a considerable part of its internal energy as seen from the color of the melt representing the melt temperature. The melt reaches actually the lower head in a form of solid particles. For about  $70 \text{ s}$  the level of the pressure is around  $12 \text{ bar}$ . During this time there is pulsating water exchange between the lower head and the down-comer. Melt-water interactions in the down-comer cause

considerable water dispersion waves. Water cyclically comes in contact with the melt and is then being repulsed. Short living pressure pulses up to 25 *bar* are also shown in Kolev (2001a).

Limitation of the present day computer power is the main reason not to attack this complicated multi-phase problem with the same resolution as is now standard for single-phase flow. Grid dependence is investigated increasing the number of the cells up to  $(12 \times 12 \times 29) = 4176$  for half of the reactor vessel. We found that the pressures satisfy the findings that the uncertainty in the pressure prediction is within 50% of the pressure increase. The main conclusion of studding this particular case made in Kolev (2001a) was that the melt-water interaction causes pressure spikes below 30 *bar* maximum pressure which again having the uncertainties in mind with which such computations are associated Kolev (2001a) is of no concern for the integrity of the pressure vessel. Therefore there is no danger that the external cooling strategy will not start properly.

#### **14.1.6.2 Model elements addressed in case 14**

1. All elements as in the all previous cases;
2. The main problem associated with this analysis is the long duration of the melt-water interaction compared to all previous cases and the need to resolve pressure spikes as in the cases of short time interactions.

#### **14.1.6.3 Available files on CD**

*fig\_14.23.gif, fig\_14.23.html*

## **14.2 Pipe networks**

### **14.2.1 Case 15**

This case is in contrast with all previous cases. It is a simulation of processes inside a complicated network consisting of pipes, valves etc. – see Fig. 14.24 Kolev (2000b). The heat exchanger has as primary medium water at 320°C at about 160 *bar* pressure and as secondary medium water with 50°C at about 6 *bar*. A break inside the heat exchanger is simulated. Flashing water enters in the secondary side and creates non-stable pressure increase. The first 0.2 *s* are characterized by strong condensation oscillation in the secondary side. Complicated interactions between the valves and the break are extremely challenging to the solution method.

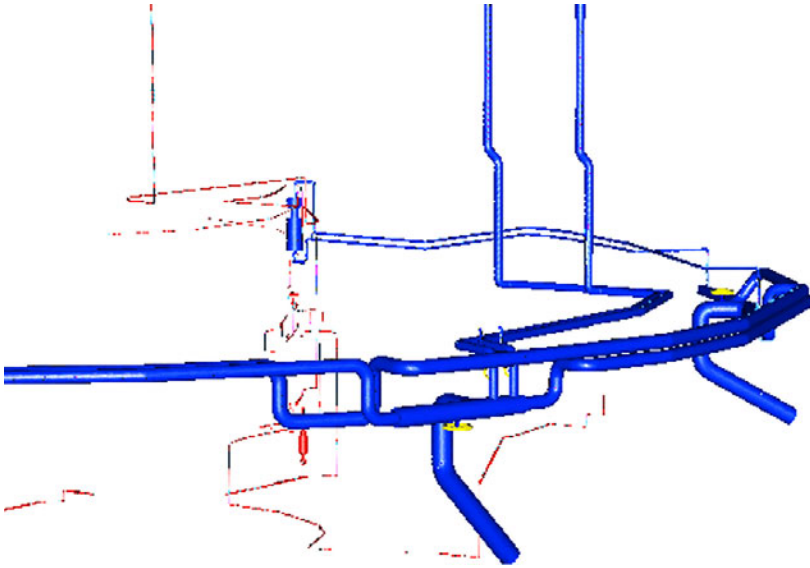


Fig. 14.26 Pipe break in a high-pressure heat exchanger. Pressure as a function of time

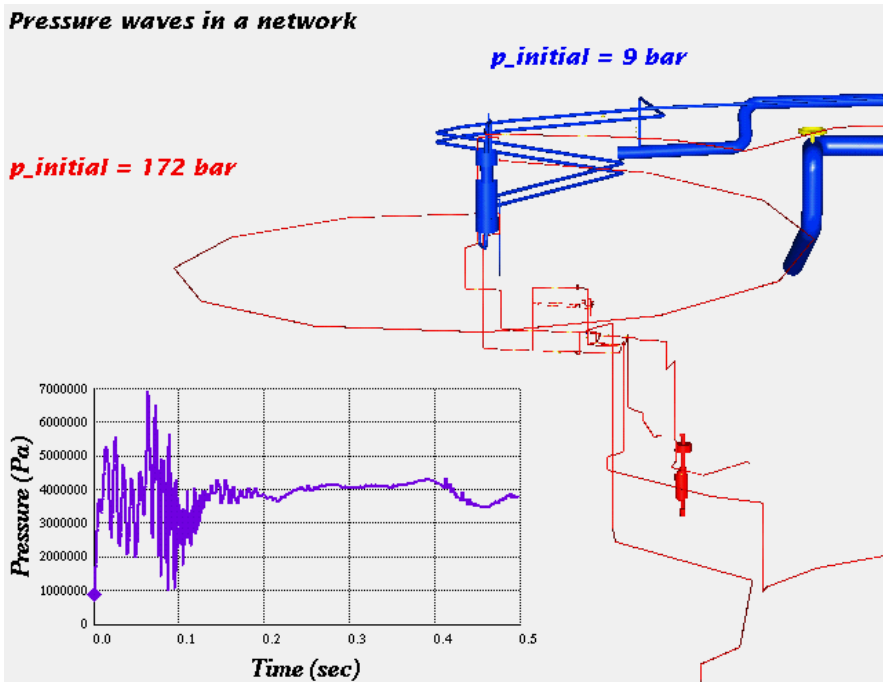


Fig. 14.27 (Cont.)

**14.2.1.1 Model elements addressed in case 15**

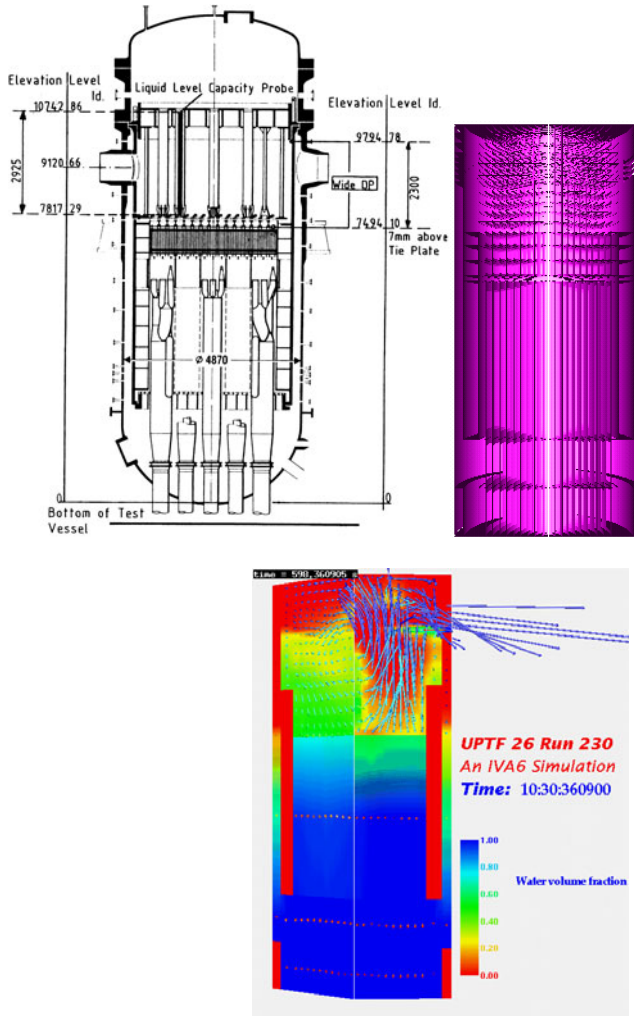
1. Two-phase flow in complex pipe networks;
2. Networks containing valves interacting with the flow;
3. Completely opposite physical phenomena like flushing and condensation shocks happen in the same network;
4. Constitutive models for flushing and condensation shocks.

**14.2.1.2 Available files on CD**

*fig\_14.24.gif, fig\_14.24.html*

### 14.3 3D steam-water interactions

#### 14.3.1 Case 16



**Fig. 14.28** a) 3D test section simulating full scale PWR for hot leg water injection in steaming core. b) IVA6 geometry model for UPTF-Test 26 Run 230 simulation. c) Water volume concentration and steam velocity as a function of space at different times. Velocities at the entrance of the main circulation pipe for the corresponding times

Figure 14.25 a) and b) presents a technical facility designed for studying of mechanical steam-water interaction in complicated geometry – see Kolev, Seitz and

Roloff (2001b). A water jet is injected from the side of the upper part, and steam is injected inside the vessel on a specified horizontal level through a part of the cross section. Figure 14.25 c) presents an IVA6 simulation of this process.

#### **14.3.1.1 Model elements addressed in case 16**

1. Geometry representation using the concept of the heterogeneous permeabilities;
2. Fragmentation and coalescence dynamics, represented here with conservation equations for a given class of particle number densities, etc.

#### **14.3.1.2 Available files on CD**

*fig\_14.25.gif, fig\_14.25.html*

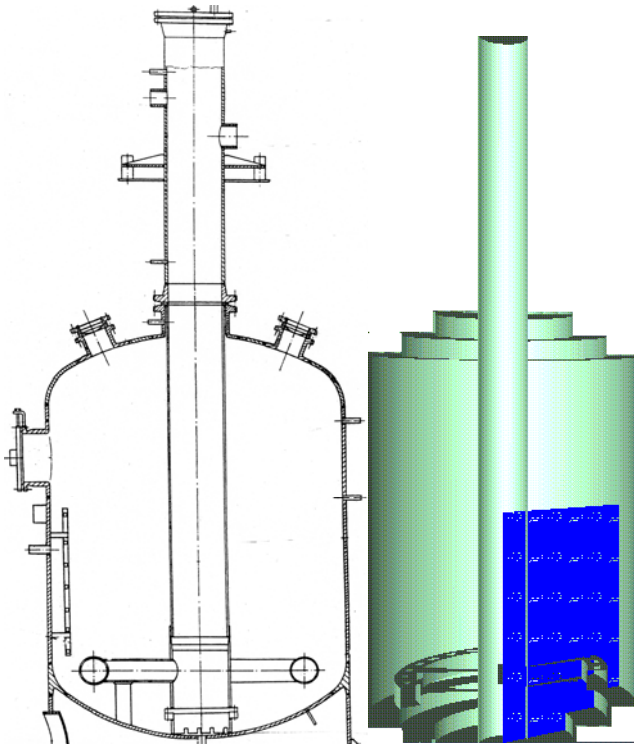
## **14.4 Three-dimensional steam-water interaction in presence of non-condensable gases**

### **14.4.1 Case 17**

Controlling the pressure of pressurized system is one of the most important safety tasks of the design engineer. Usually safety and relieve valves helps to release working material if for some reason the pressure increases. In many cases the working material is not released to the atmosphere but to an especially designed for this purpose vessel. For systems working with steam this may happen in water vessels in order to reduce dramatically the volume of the steam. The vessels are usually filled partially with water and partially with inert gas. Knowing the phenomena in such a release process allows designing properly the system. Let consider one such case which geometry and initial conditions are presented in Fig. 14.29.

The biggest challenge here is the very sensitive condensation process at bubble and droplet surfaces that changes dramatically. We open the valve over about 22 seconds. Thereafter the valve is closed. Several interesting processes can be recognized from the attached movie. Initially the pressure in the pressure dome increases up to 10 bar in order to overcome the inertia of the water and to remove it from the dome. In front of the water we observe how the initial volume of the inert nitrogen was dramatically reduced. The water first penetrates in the release vessel followed by the inert gases. This causes increasing the water level. Due to the intensive input of the mechanical energy the system is so much turbulized that within a few seconds all the vessel is occupied by rotating two phases, two component mixture. The condensation of the steam happens in the immediate neighborhood of the distribution nozzles. Large bubble oscillation in the vessel causes increasing and decreasing resistance to the steam resulting in some macroscopic pressure oscillations. After closing the valve, the pressure in the dome decreases exponential-

ly. After reaching a values lower then the vessel pressure some waters penetrates back into the distribution system. Because there is only pure steam in the dome the condensation creates a vacuum of 0.2 bar. The inverse pressure difference causes change of the flow direction in the nozzles. The stratification process at that moment in the vessel has just started and there is enough rotation energy keeping the mixture none stratified. The result is the two-phase penetration back into the dome. But the gas inside the vessel was nitrogen. Therefore nitrogen is transferred in this way back into the dome reducing the non wished vacuum production in the dome.



**Fig. 14.29** Release of saturated steam ( $p = 175 \text{ bar}$ ) into vessel partially filled with  $50^\circ\text{C}$  water through  $40 \text{ cm}^2$  cross section valve.

#### **14.4.1.1 Model elements addressed in case 17**

1. Geometry representation using the concept of the heterogeneous permeabilities, permeabilities being a time functions;
2. Fragmentation and coalescence dynamics, represented here with conservation equations for a given class of particle number densities, etc.;
3. Steam condensation (bubbles, droplets) in subcooled liquid;



4. Steam condensation (bubbles, droplets) in subcooled liquid in presence of non-condensable gases;
5. Dynamic flow regime transitions.

#### **14.4.1.2 Available files on CD**

*fig\_14.26.gif, fig\_14.26.html*

## **14.5 Three dimensional steam production in boiling water reactor**

### **14.5.1 Case 18**

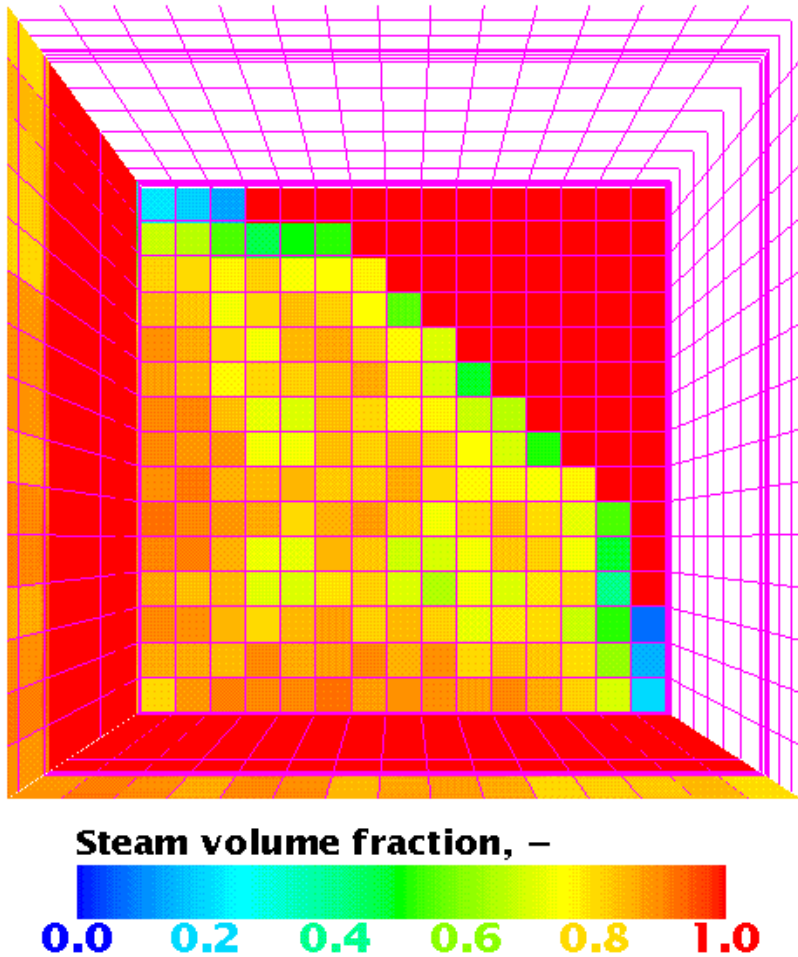
Modern design of boiling water reactors requires analysis of the steam quality produced by the nuclear reactor. We give here an example from a design case study in which 1/4<sup>th</sup> of the modern reactor vessel was presented by about 10000 cells. The thermal heat release in the reactor core is 3D-non uniform. The computation is performing starting with an arbitrary but meaningful initial state until reaching the steady state. The results are presented in Fig.14.27 for the horizontal plane being at the outlet of the core. Such studies are valuable tools for optimizing design geometry.

#### **14.5.1.1 Model elements addressed in case 18**

1. Geometry representation using the concept of the heterogeneous permeabilities;
2. Fragmentation and coalescence dynamics, represented here with conservation equations for a given class of particle number densities, etc.;
3. Dynamic flow regime transitions.
4. Thermal structure-flow interactions.
5. All forced convection heat transfer regimes: forced convection, subcooled boiling, saturated forced convection boiling, departure of nucleate boiling etc.
6. All kind of interaction of three velocity field with each other.
7. Appearance and disappearance of velocity fields.

#### **14.5.1.2 Available files on CD**

*fig\_14.27.gif, fig\_14.27.html*



**Fig. 14.30** Volume fraction of steam in a boiling water reactor - top view into 1/4<sup>th</sup> of the reactor vessel

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# Appendix 1 Brief introduction to vector analysis

Before starting to study the theory of multi-phase flows it is advisable to refresh your knowledge on vector analysis. My favorite choice is the book “Calculus and Analytic Geometry” by *Thomas et al.* (1998). Of course you may use any text book to this topic also. Here only a brief summary is given in order to assist in the understanding of the vector notations in this book.

**Right handed Cartesian coordinate system:** To locate points in space, we use three mutually perpendicular coordinate axes  $(x, y, z)$  as proposed by *Descartes*. When you hold your right hand so that the fingers curl from the positive  $x$ -axis toward the positive  $y$ -axis and your thumb points along the positive  $z$ -axis the coordinate systems is *right handed*. We use right handed coordinate systems.

**Vector (The Gibbs - Heaviside concept from 1870):** A *vector* in a space is a directed line segment. Two vectors are *equal* or *the same* if they have the same length and direction.

**The vector between two points:** The vector from point  $P_1(x_1, y_1, z_1)$  to point  $P_2(x_2, y_2, z_2)$  is

$$\vec{P_1P_2} = (x_2 - x_1)\mathbf{i} + (y_2 - y_1)\mathbf{j} + (z_2 - z_1)\mathbf{k}. \quad (1)$$

**Midpoints:** The position vector of the midpoint  $M$  of the line segment joining points  $P_1(x_1, y_1, z_1)$  and  $P_2(x_2, y_2, z_2)$  is

$$\mathbf{r}_M = \frac{x_2 + x_1}{2}\mathbf{i} + \frac{y_2 + y_1}{2}\mathbf{j} + \frac{z_2 + z_1}{2}\mathbf{k}. \quad (2)$$

**Centroid of a triangle:** The position vector of the center of mass of a triangle defined by the points  $P_0(x_0, y_0, z_0)$ ,  $P_1(x_1, y_1, z_1)$  and  $P_2(x_2, y_2, z_2)$  is

$$\mathbf{r}_{cm} = \frac{x_2 + x_1 + x_0}{3}\mathbf{i} + \frac{y_2 + y_1 + y_0}{3}\mathbf{j} + \frac{z_2 + z_1 + z_0}{3}\mathbf{k}. \quad (3)$$

**Centroid of the union of non-overlapping plane regions:** The Alexandrian Greek *Pappus* knew in the third century that a centroid of the union of two non-overlapping plane regions lies on the line segment joining their individual

centroids. More specifically, suppose  $m_1$  and  $m_2$  are the masses of thin plates  $P_1$  and  $P_2$  that cover non overlapping regions in the  $xy$  plane. Let  $\mathbf{c}_1$  and  $\mathbf{c}_2$  be the vectors from the origin to the respective centers of mass of  $P_1$  and  $P_2$ . Then the center of mass of the union  $P_1 \cup P_2$  of the two plates is determined by the vector

$$\mathbf{r}_{cm} = \frac{m_1\mathbf{c}_1 + m_2\mathbf{c}_2}{m_1 + m_2}. \quad (4)$$

This equation is known as the *Pappus's* formula. For more than two non-overlapping plates, as long as their number is finite, the formula generalizes to

$$\mathbf{r}_{cm} = \frac{\sum_m m_m \mathbf{c}_m}{\sum_m m_m}. \quad (5)$$

This formula is especially useful for finding the centroid of a plate of irregular shape that is made up of pieces of constant density whose centroids we know from geometry. We find the centroid of each piece and apply the above equation to find the centroid of the plane.

**Magnitude:** The *magnitude* (length) of the vector  $\mathbf{A} = a_1\mathbf{i} + a_2\mathbf{j} + a_3\mathbf{k}$  is

$$|\mathbf{A}| = |a_1\mathbf{i} + a_2\mathbf{j} + a_3\mathbf{k}| = \sqrt{a_1^2 + a_2^2 + a_3^2}. \quad (6)$$

Vectors with magnitude equal to one are called *unit vectors*. Unit vectors are built from direction cosines. Vectors with magnitude zero are called *zero vectors* and are denoted with  $\mathbf{0}$ .

**Lines and line segments in space:** The vector equation for the line through  $P_0(x_0, y_0, z_0)$  parallel to a vector  $\mathbf{v} = A\mathbf{i} + B\mathbf{j} + C\mathbf{k}$  is

$$\vec{P_0P} = t\mathbf{v}, \quad -\infty < t < \infty \quad (7)$$

or

$$(x - x_0)\mathbf{i} + (y - y_0)\mathbf{j} + (z - z_0)\mathbf{k} = t(A\mathbf{i} + B\mathbf{j} + C\mathbf{k}), \quad -\infty < t < \infty, \quad (8)$$

where all points  $P$  lie on the line and  $t$  is a parameter. For line segments  $t$  is bounded, e.g.  $a < t < b$ .

**Standard parametrization of the line through  $P_0(x_0, y_0, z_0)$  parallel to  $\mathbf{v} = A\mathbf{i} + B\mathbf{j} + C\mathbf{k}$ :** The standard parametrization of the line through  $P_0(x_0, y_0, z_0)$  parallel to the vector  $\mathbf{v} = A\mathbf{i} + B\mathbf{j} + C\mathbf{k}$  is

$$x = x_0 + tA, \quad y = y_0 + tB, \quad z = z_0 + tC, \quad -\infty < t < \infty. \quad (9)$$

In other words the points  $(x, y, z) = (x_0 + tA, y_0 + tB, z_0 + tC)$  in the interval  $-\infty < t < \infty$  make up the line.

**Position vector defining a line:** The vector

$$\mathbf{r}(x, y, z) = \mathbf{r}(t) = (x_0 + tA)\mathbf{i} + (y_0 + tB)\mathbf{j} + (z_0 + tC)\mathbf{k}, \quad -\infty < t < \infty \quad (10)$$

from the origin to the point  $P(x, y, z)$  is the *position vector*. The position vector defining a line crossing the points  $P_1(x_1, y_1, z_1)$  and  $P_2(x_1, y_1, z_1)$  is

$$\mathbf{r}(x, y, z) = \mathbf{r}(t) = [x_1 + t(x_2 - x_1)]\mathbf{i} + [y_1 + t(y_2 - y_1)]\mathbf{j} + [z_1 + t(z_2 - z_1)]\mathbf{k}, \quad 0 < t < 1. \quad (11)$$

**Derivative of a position vector defining a line:** At any point  $t$  the derivative of the position vector is

$$\frac{d\mathbf{r}}{dt} = \lim_{\Delta t \rightarrow 0} \frac{\mathbf{r}(t + \Delta t) - \mathbf{r}(t)}{\Delta t} = A\mathbf{i} + B\mathbf{j} + C\mathbf{k}, \quad (12)$$

which is parallel to the line.

**Position vector defining a curve:** The vector

$$\mathbf{r}(x, y, z) = \mathbf{r}(t) = f(t)\mathbf{i} + g(t)\mathbf{j} + h(t)\mathbf{k}, \quad -\infty < t < \infty \quad (13)$$

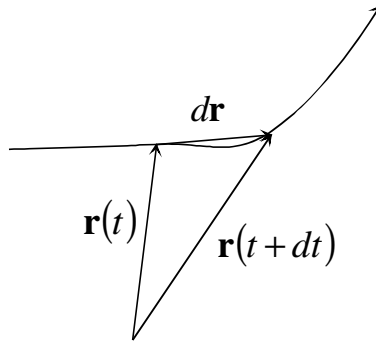
from the origin to the point  $P(x, y, z)$  that belongs to the curve is the *position vector* of the curve.

**Derivative of a position vector defining a curve:** At any point  $t$  the derivative of the position vector is the vector

$$\frac{d\mathbf{r}}{dt} = \lim_{\Delta t \rightarrow 0} \frac{\mathbf{r}(t + \Delta t) - \mathbf{r}(t)}{\Delta t} = \frac{\partial f}{\partial t}\mathbf{i} + \frac{\partial g}{\partial t}\mathbf{j} + \frac{\partial h}{\partial t}\mathbf{k}, \quad (14)$$

- see Fig. A1.1.

The curve traced by  $\mathbf{r}$  is smooth if  $d\mathbf{r}/dt$  is continuous and never  $\mathbf{0}$ , i.e. if  $f$ ,  $g$ , and  $h$  have continuous first derivatives that are not simultaneously 0. The vector  $d\mathbf{r}/dt$ , when different from  $\mathbf{0}$ , is also a vector *tangent* to the curve. The tangent line through the point  $[f(t_0), g(t_0), h(t_0)]$  is defined to be a line through the point



**Fig. A1.1** Infinitesimal change of the position vector

parallel to  $d\mathbf{r}/dt$  at  $t = t_0$ . The magnitude of the derivative of the position vector defining the curve is

$$\left| \frac{d\mathbf{r}}{dt} \right| = \sqrt{\left( \frac{df}{dt} \right)^2 + \left( \frac{dg}{dt} \right)^2 + \left( \frac{dh}{dt} \right)^2} = \sqrt{\left( \frac{dx}{dt} \right)^2 + \left( \frac{dy}{dt} \right)^2 + \left( \frac{dz}{dt} \right)^2}. \quad (15)$$

For the case where the position vector defines a line between two points  $P_1(x_1, y_1, z_1)$  and  $P_2(x_2, y_2, z_2)$  the magnitude of the derivative of the position vector is

$$\left| \frac{d\mathbf{r}}{dt} \right| = \sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2} \quad (16)$$

the distance between the two points.

**Arc length:** The length of a smooth curve  $\mathbf{r}(t) = f(t)\mathbf{i} + g(t)\mathbf{j} + h(t)\mathbf{k}$ ,  $a < t < b$ , that is traced exactly once as  $t$  increases from  $t = a$  to  $t = b$  is

$$L = \int_a^b \sqrt{\left( \frac{df}{dt} \right)^2 + \left( \frac{dg}{dt} \right)^2 + \left( \frac{dh}{dt} \right)^2} dt = \int_a^b \sqrt{\left( \frac{dx}{dt} \right)^2 + \left( \frac{dy}{dt} \right)^2 + \left( \frac{dz}{dt} \right)^2} dt = \int_a^b \left| \frac{d\mathbf{r}}{dt} \right| dt. \quad (17)$$

For the case where the position vector defines a line between two points  $P_1(x_1, y_1, z_1)$  and  $P_2(x_2, y_2, z_2)$  the length of the line segment between the two points is

$$L = \int_0^1 \sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2} dt = \sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2}. \quad (18)$$

**Arc length parameter base point**  $P_0(x_0, y_0, z_0)$  : If we choose  $P_0(t_0)$  on a smooth curve  $C$  parametrized by  $t$ , each value of  $t$  determines a point  $P[x(t), y(t), z(t)]$  on  $C$  and a “directed distance”

$$s(t) = \int_{t_0}^t \left| \frac{d\mathbf{r}}{dt} \right| d\tau, \quad (19)$$

measured along  $C$  from the base point. If  $t > t_0$ ,  $s(t)$  is the distance from  $P_0(t_0)$  to  $P(t)$ . If  $t < t_0$ ,  $s(t)$  is the negative of the distance. Each value of  $s$  determines a point on  $C$  and parametrizes  $C$  with respect to  $s$ . We call  $s$  an *arc length parameter* for the curve. The parameter’s value increases in direction of increasing  $t$ . For a smooth curve the derivatives beneath the radical are continuous and the fundamental theorem of calculus tells us that  $s$  is a differentiable function of  $t$  with derivative

$$\frac{ds}{dt} = \left| \frac{d\mathbf{r}}{dt} \right|. \quad (20)$$

Note that, while the base point  $P_0(t_0)$  plays a role in defining  $s$ , it plays no role in defining  $ds/dt$ . Note also that  $ds/dt > 0$  since the curve is smooth and  $s$  increases with increasing  $t$ .

**Time derivatives, velocity, speed, acceleration, direction of motion** If

$$\mathbf{r}(x, y, z) = \mathbf{r}(\tau) = f(\tau)\mathbf{i} + g(\tau)\mathbf{j} + h(\tau)\mathbf{k} \quad (21)$$

is a position vector of a particle moving along a smooth curve in space, and the components are smooth functions of time  $\tau$ , then

$$\mathbf{v}(t) = \frac{d\mathbf{r}}{dt} \quad (22)$$

is the particle’s *velocity vector*, tangent to a curve. At any time  $\tau$ , the direction of  $\mathbf{v}$  is the *direction of motion*. The magnitude of  $\mathbf{v}$  is the particle’s speed.

$$\text{Speed} = |\mathbf{v}(\tau)| = \left| \frac{d\mathbf{r}}{d\tau} \right|. \quad (23)$$

The derivative

$$\mathbf{a} = \frac{d\mathbf{v}}{d\tau} = \frac{d^2\mathbf{r}}{d\tau^2}, \quad (24)$$

when it exists, is the particle’s *acceleration vector*.



**Differentiation rules for time derivatives:** Because the derivatives of vector functions may be computed component by component, the rules for differentiating vector functions have the same form as the rules for differentiating scalar functions.

$$\text{Constant Function Rule: } \frac{d\mathbf{C}}{d\tau} = \mathbf{0} \quad (\text{any constant vector } \mathbf{C}). \quad (25)$$

$$\text{Scalar Multiple Rules: } \frac{d}{d\tau}(c\mathbf{u}) = c \frac{d\mathbf{u}}{d\tau} \quad (\text{any number } c) \quad (26)$$

$$\frac{d}{d\tau}(f\mathbf{u}) = f \frac{d\mathbf{u}}{d\tau} + \mathbf{u} \frac{df}{d\tau} \quad (\text{any differentiable scalar function}). \quad (27)$$

If  $\mathbf{u}$  and  $\mathbf{v}$  are differentiable vector functions of  $t$ , then

$$\text{Sum Rule: } \frac{d}{d\tau}(\mathbf{u} + \mathbf{v}) = \frac{d\mathbf{u}}{d\tau} + \frac{d\mathbf{v}}{d\tau}. \quad (28)$$

$$\text{Difference Rule: } \frac{d}{d\tau}(\mathbf{u} - \mathbf{v}) = \frac{d\mathbf{u}}{d\tau} - \frac{d\mathbf{v}}{d\tau}. \quad (29)$$

$$\text{Dot Product Rule: } \frac{d}{d\tau}(\mathbf{u} \cdot \mathbf{v}) = \frac{d\mathbf{u}}{d\tau} \cdot \frac{d\mathbf{v}}{d\tau}. \quad (30)$$

$$\text{Cross Product Rule: } \frac{d}{d\tau}(\mathbf{u} \times \mathbf{v}) = \frac{d\mathbf{u}}{d\tau} \times \mathbf{v} + \mathbf{u} \times \frac{d\mathbf{v}}{d\tau} \quad (\text{in that order}). \quad (31)$$

**Chain Rule (Short Form):** If  $\mathbf{r}$  is a differentiable function of  $\tau$  and  $\tau$  is a differentiable function of  $s$ , then

$$\frac{d\mathbf{r}}{ds} = \frac{d\mathbf{r}}{d\tau} \frac{d\tau}{ds}. \quad (32)$$

**Unit tangent vector  $\mathbf{T}$ :** The *unit tangent vector* of a differentiable curve  $\mathbf{r}(t)$  is

$$\mathbf{T} = \frac{d\mathbf{r}}{ds} = \frac{d\mathbf{r}/dt}{ds/dt} = \frac{d\mathbf{r}/dt}{|d\mathbf{r}/dt|}, \quad (33)$$

or

$$\mathbf{T} = \frac{d\mathbf{r}}{ds} = \left( \frac{\partial x}{\partial s} \mathbf{i} + \frac{\partial y}{\partial s} \mathbf{j} + \frac{\partial z}{\partial s} \mathbf{k} \right). \quad (34)$$

For the case where the position vector defines a line between two points  $P_1(x_1, y_1, z_1)$  and  $P_2(x_2, y_2, z_2)$  the unit tangent vector is

$$\mathbf{T} = \frac{d\mathbf{r}}{ds} = \frac{d\mathbf{r}/dt}{ds/dt} = \frac{d\mathbf{r}/dt}{|d\mathbf{r}/dt|} = \frac{(x_2 - x_1)\mathbf{i} + (y_2 - y_1)\mathbf{j} + (z_2 - z_1)\mathbf{k}}{\sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2}}. \quad (35)$$

The unit tangent vector plays an important role in the theory of curvilinear coordinate transformation.

**Curvature of space curve:** As a particle moves along a smooth curve in the space,  $\mathbf{T} = d\mathbf{r}/ds$  turns as the curve bends. Since  $\mathbf{T}$  is a unit vector, its length remains constant and only its direction changes as the particle moves along the curve. The rate at which  $\mathbf{T}$  turns per unit of the length along the curve is called *curvature*. The traditional symbol for the curvature is the Greek letter kappa,  $\kappa$ . If  $\mathbf{T}$  is the tangent vector of a smooth curve, the curvature function of the curve is

$$\kappa = \left| \frac{d\mathbf{T}}{ds} \right|. \quad (36)$$

**The principal unit normal vector for space curve:** Since  $\mathbf{T}$  has a constant length, the vector  $d\mathbf{T}/ds$  is orthogonal to  $\mathbf{T}$ . Therefore, if we divide  $d\mathbf{T}/ds$  by its magnitude  $\kappa$ , we obtain a unit vector orthogonal to  $\mathbf{T}$ . At a point where  $\kappa \neq 0$ , the principal unit normal vector for a curve in the space is

$$\mathbf{N} = \frac{d\mathbf{T}/ds}{|d\mathbf{T}/ds|} = \frac{1}{\kappa} \frac{d\mathbf{T}}{ds}. \quad (37)$$

The vector  $d\mathbf{T}/ds$  points in the direction in which  $\mathbf{T}$  turns as the curve bends. Therefore, if we face the direction of increasing length, the vector  $d\mathbf{T}/ds$  points toward the right if  $\mathbf{T}$  turns clockwise and towards the left if  $\mathbf{T}$  turns counterclockwise. Because the arc length parameter for a smooth curve is defined with  $ds/dt$  positive,  $ds/dt = |ds/dt|$ , the chain rule gives

$$\mathbf{N} = \frac{d\mathbf{T}/ds}{|d\mathbf{T}/ds|} = \frac{(d\mathbf{T}/dt)(dt/ds)}{|d\mathbf{T}/dt||dt/ds|} = \frac{d\mathbf{T}/dt}{|d\mathbf{T}/dt|}. \quad (38)$$

**Scalar (dot) product:** The *scalar product (dot product)*  $\mathbf{A} \cdot \mathbf{B}$  (“**A dot B**”) of vectors

$$\mathbf{A} = a_1\mathbf{i} + a_2\mathbf{j} + a_3\mathbf{k}, \quad (39)$$

and

$$\mathbf{B} = b_1\mathbf{i} + b_2\mathbf{j} + b_3\mathbf{k} \quad (40)$$

is the number

$$\mathbf{A} \cdot \mathbf{B} = |\mathbf{A}||\mathbf{B}|\cos\theta \quad (41)$$

where  $\theta$  is the angle between  $\mathbf{A}$  and  $\mathbf{B}$ . In words, it is the length of  $\mathbf{A}$  times the length  $\mathbf{B}$  times the cosine of the angle between  $\mathbf{A}$  and  $\mathbf{B}$ . The law of cosines for the triangle whose sides represent  $\mathbf{A}$ ,  $\mathbf{B}$  and  $\mathbf{C}$  is  $|\mathbf{C}|^2 = |\mathbf{A}|^2 + |\mathbf{B}|^2 - 2|\mathbf{A}||\mathbf{B}|\cos\theta$  is used to obtain

$$\mathbf{A} \cdot \mathbf{B} = a_1 b_1 + a_2 b_2 + a_3 b_3. \quad (42)$$

To find the scalar product of two vectors we multiply their corresponding components and add the results.

$$\mathbf{A} \cdot \mathbf{A} = a_1^2 + a_2^2 + a_3^2 = |\mathbf{A}|^2 \quad (43)$$

is called the *Euclidian norm*.

**Splitting a vector into components parallel and orthogonal to another vector:** The vector  $\mathbf{B}$  can be split into a component which is parallel to  $\mathbf{A}$

$$\text{proj}_{\mathbf{A}} \mathbf{B} = \left( \frac{\mathbf{B} \cdot \mathbf{A}}{\mathbf{A} \cdot \mathbf{A}} \right) \mathbf{A} \quad (44)$$

and a component orthogonal to  $\mathbf{A}$ ,  $\mathbf{B} - \text{proj}_{\mathbf{A}} \mathbf{B}$ ,

$$\mathbf{B} = \left( \frac{\mathbf{B} \cdot \mathbf{A}}{\mathbf{A} \cdot \mathbf{A}} \right) \mathbf{A} + \left[ \mathbf{B} - \left( \frac{\mathbf{B} \cdot \mathbf{A}}{\mathbf{A} \cdot \mathbf{A}} \right) \mathbf{A} \right]. \quad (45)$$

**Equations for a plane in space:** Suppose a plane  $M$  passes through a point  $P_0(x_0, y_0, z_0)$  and is normal (perpendicular) to the non-zero vector

$$\mathbf{n} = A\mathbf{i} + B\mathbf{j} + C\mathbf{k}. \quad (46)$$

Then  $M$  is the set of all points  $P(x, y, z)$  for which  $\vec{P_0P}$  is orthogonal to  $\mathbf{n}$ . That is,  $P$  lies on  $M$  if and only if

$$\mathbf{n} \cdot \vec{P_0P} = 0. \quad (47)$$

This equation is equivalent to

$$(A\mathbf{i} + B\mathbf{j} + C\mathbf{k}) \cdot [(x - x_0)\mathbf{i} + (y - y_0)\mathbf{j} + (z - z_0)\mathbf{k}] = 0 \quad (48)$$

or

$$A(x - x_0) + B(y - y_0) + C(z - z_0) = 0. \quad (49)$$

Note that the angle between two intersecting planes is defined to be the (acute) angle determined by their normal vectors.

**A plane determined by three points:** Consider a plane determined by the three points  $P_0(x_0, y_0, z_0)$ ,  $P_1(x_1, y_1, z_1)$  and  $P_2(x_2, y_2, z_2)$ . The vector connecting point  $P_0$  with point  $P_1$  is then

$$\vec{P_0P_1} = (x_1 - x_0)\mathbf{i} + (y_1 - y_0)\mathbf{j} + (z_1 - z_0)\mathbf{k}. \quad (50)$$

The vector connecting point  $P_0$  with point  $P_2$  is then

$$\vec{P_0P_2} = (x_2 - x_0)\mathbf{i} + (y_2 - y_0)\mathbf{j} + (z_2 - z_0)\mathbf{k}. \quad (51)$$

The vector normal to the plane is defined by

$$\vec{P_0P_1} \times \vec{P_0P_2} = \begin{vmatrix} \mathbf{i} & \mathbf{j} & \mathbf{k} \\ x_1 - x_0 & y_1 - y_0 & z_1 - z_0 \\ x_2 - x_0 & y_2 - y_0 & z_2 - z_0 \end{vmatrix}. \quad (52)$$

The plane is then defined by

$$\begin{vmatrix} \mathbf{i} & \mathbf{j} & \mathbf{k} \\ x_1 - x_0 & y_1 - y_0 & z_1 - z_0 \\ x_2 - x_0 & y_2 - y_0 & z_2 - z_0 \end{vmatrix} \cdot [(x - x_0)\mathbf{i} + (y - y_0)\mathbf{j} + (z - z_0)\mathbf{k}] = 0, \quad (53)$$

or

$$\begin{vmatrix} x_0 - x & y_0 - y & z_0 - z \\ x_1 - x & y_1 - y & z_1 - z \\ x_2 - x & y_2 - y & z_2 - z \end{vmatrix} = 0. \quad (54)$$

**Laws of the dot product:** The dot product is commutative

$$\mathbf{A} \cdot \mathbf{B} = \mathbf{B} \cdot \mathbf{A}. \quad (55)$$

If  $c$  is any number (or scalar), then

$$(c\mathbf{A}) \cdot \mathbf{B} = \mathbf{A} \cdot (c\mathbf{B}) = c(\mathbf{A} \cdot \mathbf{B}). \quad (56)$$

If  $\mathbf{C} = c_1\mathbf{i} + c_2\mathbf{j} + c_3\mathbf{k}$  is any third vector, then

$$\mathbf{A} \cdot (\mathbf{B} + \mathbf{C}) = \mathbf{A} \cdot \mathbf{B} + \mathbf{A} \cdot \mathbf{C}, \quad (57)$$

that is the dot product obeys the *distributive law*. Combined with the commutative law it is also evident that

$$(\mathbf{A} + \mathbf{B}) \cdot \mathbf{C} = \mathbf{A} \cdot \mathbf{C} + \mathbf{B} \cdot \mathbf{C}. \quad (58)$$

The last two equations permit us to multiply sums of vectors by the familiar laws of algebra. For example

$$(\mathbf{A} + \mathbf{B}) \cdot (\mathbf{C} + \mathbf{D}) = \mathbf{A} \cdot \mathbf{C} + \mathbf{A} \cdot \mathbf{D} + \mathbf{B} \cdot \mathbf{C} + \mathbf{B} \cdot \mathbf{D}. \quad (59)$$

**Angle between two non-zero vectors:** The angle between two non-zero vectors  $\mathbf{A}$  and  $\mathbf{B}$  is

$$\theta = \arccos \left( \frac{\mathbf{A} \cdot \mathbf{B}}{|\mathbf{A}| |\mathbf{B}|} \right). \quad (60)$$

**Perpendicular (orthogonal) vectors:** Non-zero vectors  $\mathbf{A}$  and  $\mathbf{B}$  are *perpendicular (orthogonal)* if and only if

$$\mathbf{A} \cdot \mathbf{B} = 0. \quad (61)$$

**The gradient vector (gradient):** The *gradient vector (gradient)* of the differentiable function  $f(x, y, z)$  at point  $P_0(x_0, y_0, z_0)$  is the vector

$$\nabla f = \frac{\partial f}{\partial x} \mathbf{i} + \frac{\partial f}{\partial y} \mathbf{j} + \frac{\partial f}{\partial z} \mathbf{k} \quad (62)$$

obtained by evaluating the partial derivatives of  $f$  at  $P_0(x_0, y_0, z_0)$ . The notation  $\nabla f$  is read “grad  $f$ ” as well as “gradient of  $f$ ” and “del  $f$ ”. The symbol  $\nabla$  by itself is read “del”. Another notation for the gradient is  $\text{grad } f$ , read the way it is written.

**Equations for tangent planes and normal lines:** If

$$\mathbf{r}(x, y, z) = \mathbf{r}(t) = g(t)\mathbf{i} + h(t)\mathbf{j} + k(t)\mathbf{k}$$

is a smooth curve on the level surface  $f(x, y, z) = c$  of a differentiable function  $f$ , then  $f[g(t), h(t), k(t)] = c$ . Differentiating both sides of this equation with respect to  $t$  leads to

$$\frac{d}{dt} f[g(t), h(t), k(t)] = \frac{dc}{dt} \quad (63)$$

$$\begin{aligned} \frac{df}{dt} &= \frac{\partial f}{\partial x} \frac{dg}{dt} + \frac{\partial f}{\partial y} \frac{dh}{dt} + \frac{\partial f}{\partial z} \frac{dk}{dt} = \left( \frac{\partial f}{\partial x} \mathbf{i} + \frac{\partial f}{\partial y} \mathbf{j} + \frac{\partial f}{\partial z} \mathbf{k} \right) \cdot \left( \frac{dg}{dt} \mathbf{i} + \frac{dh}{dt} \mathbf{j} + \frac{dk}{dt} \mathbf{k} \right) \\ &= \nabla f \cdot \frac{d\mathbf{r}}{dt} = 0. \end{aligned} \quad (64)$$

At every point along the curve,  $\nabla f$  is orthogonal to the curve's velocity vector  $\frac{d\mathbf{r}}{dt}$ , and therefore orthogonal to the surface  $f(x, y, z) = c$ . The unit normal vector to this surface is

$$\mathbf{n} = \frac{\nabla f}{|\nabla f|}. \quad (65)$$

If the curve passes through the point  $P_0(x_0, y_0, z_0)$ :

The *tangent plane* at the point  $P_0$  on the level surface  $f(x, y, z) = c$  is the plane through  $P_0$  normal to  $\nabla f|_{P_0}$ ,

$$\left( \frac{\partial f}{\partial x} \mathbf{i} + \frac{\partial f}{\partial y} \mathbf{j} + \frac{\partial f}{\partial z} \mathbf{k} \right) \cdot [(x - x_0) \mathbf{i} + (y - y_0) \mathbf{j} + (z - z_0) \mathbf{k}] = 0; \quad (66)$$

The *normal line* of the surface at  $P_0$  is the line through  $P_0$  parallel to  $\nabla f|_{P_0}$ .

**Gradient of a plane defined by three points:** Consider a plane determined by the three points  $P_0(x_0, y_0, z_0)$ ,  $P_1(x_1, y_1, z_1)$  and  $P_2(x_2, y_2, z_2)$

$$\begin{aligned} f(x, y, z) &= \begin{vmatrix} x_0 - x & y_0 - y & z_0 - z \\ x_1 - x & y_1 - y & z_1 - z \\ x_2 - x & y_2 - y & z_2 - z \end{vmatrix} \\ &= (x_0 - x)(y_1 - y)(z_2 - z) + (x_2 - x)(y_0 - y)(z_1 - z) + (x_1 - x)(y_2 - y)(z_0 - z) \\ &\quad - (x_2 - x)(y_1 - y)(z_0 - z) - (x_1 - x)(y_0 - y)(z_2 - z) - (x_0 - x)(y_2 - y)(z_1 - z) = 0 \end{aligned} \quad (67)$$

The *gradient vector (gradient)* of the differentiable function  $f(x, y, z)$  at point  $P(x, y, z)$  is the vector

$$\nabla f = \frac{\partial f}{\partial x} \mathbf{i} + \frac{\partial f}{\partial y} \mathbf{j} + \frac{\partial f}{\partial z} \mathbf{k}, \quad (68)$$

with components

$$\frac{\partial f}{\partial x} = (y_1 - y_2)(z_0 - z) + (y_2 - y_0)(z_1 - z) + (y_0 - y_1)(z_2 - z), \quad (69)$$

$$\frac{\partial f}{\partial y} = (z_1 - z_2)(x_0 - x) + (z_2 - z_0)(x_1 - x) + (z_0 - z_1)(x_2 - x), \quad (70)$$

$$\frac{\partial f}{\partial z} = (x_1 - x_2)(y_0 - y) + (x_2 - x_0)(y_1 - y) + (x_0 - x_1)(y_2 - y). \quad (71)$$

Note that the vectors  $\vec{P_0P_1}$ ,  $\vec{P_0P_2}$ , and the vector normal to the plane forms a left oriented vector system. The gradient vector is therefore parallel to the surface normal vector.

**Curvature of a surface:** Consider the smooth curves

$$\mathbf{r}_1(x, y, z) = \mathbf{r}_1(t) = g_1(t)\mathbf{i} + h_1(t)\mathbf{j} + k_1(t)\mathbf{k}, \quad (72)$$

$$\mathbf{r}_2(x, y, z) = \mathbf{r}_2(t) = g_2(t)\mathbf{i} + h_2(t)\mathbf{j} + k_2(t)\mathbf{k}, \quad (73)$$

on the level surface  $f(x, y, z) = c$  of a differentiable function  $f$ . In this case  $f[g_1(t), h_1(t), k_1(t)] = c$  and  $f[g_2(t), h_2(t), k_2(t)] = c$  along the curves. The curves cross the point  $P_0(x_0, y_0, z_0)$ . At this point they are orthogonal. The tangents at  $P_0$  of these curves are  $\mathbf{T}_1$  and  $\mathbf{T}_2$ .  $s_1$  and  $s_2$  are the arc distances counted from  $P_0$ . As a particle moves along the first curve in the space,  $\mathbf{T}_1 = d\mathbf{r}_1/ds_1$  turns as the curve bends. Since  $\mathbf{T}_1$  is a unit vector, its length remains constant and only its direction changes as the particle moves along the curve. The rate at which  $\mathbf{T}_1$  turns per unit of the length along the curve is called the *curvature*  $\kappa_1$ . As  $\mathbf{T}_1$  is the tangent vector of a smooth curve, the curvature function of the curve is

$$\kappa_1 = \left| \frac{d\mathbf{T}_1}{ds_1} \right| = |(\mathbf{T}_1 \cdot \nabla) \mathbf{T}_1|. \quad (74)$$

Similarly, the curvature function of the second curve is

$$\kappa_2 = \left| \frac{d\mathbf{T}_2}{ds_2} \right| = |(\mathbf{T}_2 \cdot \nabla) \mathbf{T}_2|. \quad (75)$$

Both curvature functions are called *principle curvatures*. Their reciprocals are called *principal radii*. The *curvature function of the level surface*  $f(x, y, z) = c$  is defined as

$$\kappa = \kappa_1 + \kappa_2. \quad (76)$$

**The gradient along a direction normal and tangential to a surface:**

Consider the level surface  $f(x, y, z) = c$  of a differentiable function  $f$  with a normal unit vector

$$\mathbf{n} = \frac{\nabla f}{|\nabla f|}. \quad (77)$$

The gradient of any differentiable scalar function in space can then be split into a component parallel to the normal vector and a component tangential to the surface:

$$\nabla = \left( \frac{\nabla \cdot \mathbf{n}}{\mathbf{n} \cdot \mathbf{n}} \right) \mathbf{n} + \left[ \nabla - \left( \frac{\nabla \cdot \mathbf{n}}{\mathbf{n} \cdot \mathbf{n}} \right) \mathbf{n} \right] = \nabla_n + \nabla_t, \quad (78)$$

where

$$\nabla_n = (\nabla \cdot \mathbf{n}) \mathbf{n} \quad (79)$$

and

$$\nabla_t = \nabla - \nabla_n. \quad (80)$$

**Curvature of a surface defined by the gradient of its unit normal vector:** Consider a smooth surface in space defined by its normal vector

$$\mathbf{n}(x, y, z) = g(x, y, z)\mathbf{i} + h(x, y, z)\mathbf{j} + k(x, y, z)\mathbf{k}. \quad (81)$$

In this case there is a convenient method reported by *Brackbill et al. (1992)* for computation of the curvature of this surface

$$\kappa = -(\nabla \cdot \mathbf{n}). \quad (82)$$

Let us estimate as an example the curvature of a liquid layer in stratified flow between two horizontal planes in the plane  $y = \text{const}$ . The interface is described by the curve  $z = \delta_2(x)$  where  $\delta_2(x)$  is the film thickness. Thus the level surface function is  $f(x, y, z) \equiv \delta_2(x) - z = 0$ . The gradient of the level surface function is

$$\nabla f(x, y = \text{const}, z) = \frac{\partial f(x, z)}{\partial x} \mathbf{i} + \frac{\partial f(x, z)}{\partial z} \mathbf{k} = \frac{\partial \delta_2(x)}{\partial x} \mathbf{i} - \mathbf{k}, \quad (83)$$

and the magnitude of the gradient

$$|\nabla f| = \left\{ 1 + \left[ \frac{\partial \delta_2(x)}{\partial x} \right]^2 \right\}^{1/2}. \quad (84)$$

The normal vector is



$$\mathbf{n}_2 = -\frac{\nabla f}{|\nabla f|} = -\left( \frac{\frac{\partial \delta_2(x)}{\partial x}}{\left\{1 + \left[\frac{\partial \delta_2(x)}{\partial x}\right]^2\right\}^{1/2}} \mathbf{i} - \frac{1}{\left\{1 + \left[\frac{\partial \delta_2(x)}{\partial x}\right]^2\right\}^{1/2}} \mathbf{k} \right). \quad (85)$$

Note that for decreasing film thickness with  $x$  the normal vector points outside the liquid. This is the result of selecting  $f(x, y, z) \equiv \delta_2(x) - z = 0$  instead of  $f(x, y, z) \equiv z - \delta_2(x) = 0$  which is also possible. The curvature is then

$$\begin{aligned} \kappa_1 &= -(\nabla \cdot \mathbf{n}_1) = \nabla \cdot \left[ \frac{\nabla f}{|\nabla f|} \right] = \nabla \cdot \left( \frac{\frac{\partial \delta_2(x)}{\partial x}}{\left\{1 + \left[\frac{\partial \delta_2(x)}{\partial x}\right]^2\right\}^{1/2}} \mathbf{i} - \frac{1}{\left\{1 + \left[\frac{\partial \delta_2(x)}{\partial x}\right]^2\right\}^{1/2}} \mathbf{k} \right) \\ &= \frac{\partial}{\partial x} \frac{\frac{\partial \delta_2(x)}{\partial x}}{\left\{1 + \left[\frac{\partial \delta_2(x)}{\partial x}\right]^2\right\}^{1/2}} - \frac{\partial}{\partial z} \frac{1}{\left\{1 + \left[\frac{\partial \delta_2(x)}{\partial x}\right]^2\right\}^{1/2}}. \end{aligned} \quad (86)$$

Having in mind that

$$\frac{\partial}{\partial z} \frac{1}{\left\{1 + \left[\frac{\partial \delta_2(x)}{\partial x}\right]^2\right\}^{1/2}} = 0 \quad (87)$$

one finally obtain the well-known expression

$$\kappa_2 = \frac{\partial^2 \delta_2(x)}{\partial x^2} \bigg/ \left\{1 + \left[\frac{\partial \delta_2(x)}{\partial x}\right]^2\right\}^{3/2}. \quad (88)$$

**Speed of displacement of geometrical surface (Truesdell and Toupin 1960):** Consider a set of geometrical surfaces defined by the equation

$$\mathbf{r} = \mathbf{r}(t_1, t_2, \tau), \quad (89)$$

where  $t_1$  and  $t_2$  are the coordinates of a point on this surface and  $\tau$  is the time. The velocity of the surface point  $(t_1, t_2)$  is defined by

$$\mathbf{V}_\sigma = \left( \frac{\partial \mathbf{r}}{\partial \tau} \right)_{t_1, t_2}. \quad (90)$$

Expressing the surface equation with

$$f(x, y, z, \tau) = 0 \quad (91)$$

we have

$$\frac{\partial f}{\partial \tau} + \mathbf{V}_\sigma \cdot \nabla f = 0. \quad (92)$$

Knowing that the outwards directed unit surface vector is

$$\mathbf{n} = \frac{\nabla f}{|\nabla f|} \quad (93)$$

the surface velocity can be expressed as

$$\mathbf{V}_\sigma \cdot \mathbf{n} = -\frac{\partial f / \partial \tau}{|\nabla f|}. \quad (94)$$

**Potential function:** If the vector  $\mathbf{F}$  is defined on  $D$  and

$$\mathbf{F} = \nabla f \quad (95)$$

for some scalar function on  $D$  then  $f$  is called a *potential function* for  $\mathbf{F}$ . The important property of the potential functions is derived from the following consideration. Suppose  $A$  and  $B$  are two points in  $D$  and that the curve

$$\mathbf{r}(x, y, z) = \mathbf{r}(t) = f(t)\mathbf{i} + g(t)\mathbf{j} + h(t)\mathbf{k}, \quad a \leq t \leq b, \quad (96)$$

is smooth on  $D$  joining  $A$  and  $B$ . Along the curve,  $f$  is a differentiable function of  $t$  and

$$\begin{aligned} \frac{df}{dt} &= \frac{\partial f}{\partial x} \frac{dx}{dt} + \frac{\partial f}{\partial y} \frac{dy}{dt} + \frac{\partial f}{\partial z} \frac{dz}{dt} = \left( \frac{\partial f}{\partial x} \mathbf{i} + \frac{\partial f}{\partial y} \mathbf{j} + \frac{\partial f}{\partial z} \mathbf{k} \right) \cdot \left( \frac{\partial x}{\partial t} \mathbf{i} + \frac{\partial y}{\partial t} \mathbf{j} + \frac{\partial z}{\partial t} \mathbf{k} \right) \\ &= \nabla f \cdot \frac{d\mathbf{r}}{dt} = \mathbf{F} \cdot \frac{d\mathbf{r}}{dt}. \end{aligned} \quad (97)$$

Therefore,

$$\int_c \mathbf{F} \cdot d\mathbf{r} = \int_a^b \mathbf{F} \cdot \frac{d\mathbf{r}}{dt} dt = \int_a^b \frac{df}{dt} dt = f(A) - f(B), \quad (98)$$

depends only of the values of  $f$  at  $A$  and  $B$  and not on the path between. In other words if

$$\mathbf{F} = M(x, y, z)\mathbf{i} + N(x, y, z)\mathbf{j} + P(x, y, z)\mathbf{k} \quad (99)$$

is a field whose component functions have continuous first partial derivatives and there is a solution of the equation

$$M(x, y, z)\mathbf{i} + N(x, y, z)\mathbf{j} + P(x, y, z)\mathbf{k} = \frac{\partial f}{\partial x}\mathbf{i} + \frac{\partial f}{\partial y}\mathbf{j} + \frac{\partial f}{\partial z}\mathbf{k}, \quad (100)$$

which means solution of the system

$$\frac{\partial f}{\partial x} = M(x, y, z), \quad (101)$$

$$\frac{\partial f}{\partial y} = N(x, y, z), \quad (102)$$

$$\frac{\partial f}{\partial z} = P(x, y, z), \quad (103)$$

the vector  $\mathbf{F}$  possesses a potential function  $f$  and is called *conservative*.

**Differential form, exact differential form:** The form

$$M(x, y, z)dx + N(x, y, z)dy + P(x, y, z)dz \quad (104)$$

is called the *differential form*. A differential form is *exact* on a domain  $D$  in space if

$$M(x, y, z)dx + N(x, y, z)dy + P(x, y, z)dz = \frac{\partial f}{\partial x}dx + \frac{\partial f}{\partial y}dy + \frac{\partial f}{\partial z}dz \quad (105)$$

for some scalar function  $f$  through  $D$ . In this case  $f$  is a smooth function of  $x$ ,  $y$ , and  $z$ . Obviously the test for the differential form being exact is the same as the test for  $\mathbf{F}$ 's being conservative that is the test for the existence of  $f$ . Another way for testing whether the differential form is exact is *Euler's relations* saying that if the differential form is exact the mutual cross derivatives are equal, that is

$$\frac{\partial}{\partial y}\left(\frac{\partial f}{\partial x}\right) = \frac{\partial}{\partial x}\left(\frac{\partial f}{\partial y}\right), \quad \frac{\partial}{\partial z}\left(\frac{\partial f}{\partial x}\right) = \frac{\partial}{\partial x}\left(\frac{\partial f}{\partial z}\right), \quad \text{and} \quad \frac{\partial}{\partial y}\left(\frac{\partial f}{\partial z}\right) = \frac{\partial}{\partial z}\left(\frac{\partial f}{\partial y}\right). \quad (106)$$

**Directional derivatives:** Suppose that the function  $f(x, y, z)$  is defined through the region  $R$  in the  $xyz$  space, that  $P_0(x_0, y_0, z_0)$  is a point in  $R$ , and that  $\mathbf{u} = u_1\mathbf{i} + u_2\mathbf{j} + u_3\mathbf{k}$  is a unit vector. Then the equations

$$x = x_0 + su_1, \quad y = y_0 + su_2, \quad z = z_0 + su_3, \quad -\infty < s < \infty, \quad (107)$$

parametrize the line through  $P_0(x_0, y_0, z_0)$  parallel to  $\mathbf{u}$ . The parameter  $s$  measures the arc length from  $P_0$  in the direction  $\mathbf{u}$ . We find the rate of change of  $f$  at  $P_0$  in the direction  $\mathbf{u}$  by calculating  $df/ds$  at  $P_0$ . The derivative of  $f$  at  $P_0$  in the direction of the unit vector  $\mathbf{u} = u_1\mathbf{i} + u_2\mathbf{j} + u_3\mathbf{k}$  is the number

$$\left(\frac{df}{ds}\right)_{\mathbf{u}, P_0} = \lim_{s \rightarrow 0} \frac{f(x_0 + su_1, y_0 + su_2, z_0 + su_3) - f(x_0, y_0, z_0)}{s} \quad (108)$$

provided the limit exists. Useful relations for practical calculations are

$$\begin{aligned} \left(\frac{df}{ds}\right)_{\mathbf{u}, P_0} &= \left[ \left(\frac{df}{dx}\right)_{P_0} \mathbf{i} + \left(\frac{df}{dy}\right)_{P_0} \mathbf{j} + \left(\frac{df}{dz}\right)_{P_0} \mathbf{k} \right] \cdot (u_1\mathbf{i} + u_2\mathbf{j} + u_3\mathbf{k}) = (\nabla f)_{P_0} \cdot \mathbf{u} \\ &= |\nabla f| \cdot |\mathbf{u}| \cos \theta. \end{aligned} \quad (109)$$

At any given point,  $f$  increases most rapidly in the direction of  $\nabla f$  and decreases most rapidly in the direction of  $-\nabla f$ . In any direction orthogonal to  $\nabla f$ , the derivative is zero.

**Derivatives of a function along the unit tangent vector:** Suppose that the function  $f(x, y, z)$  is defined through the region  $R$  in the  $xyz$  space, and is differentiable, and that

$$\mathbf{T} = \frac{d\mathbf{r}}{ds} = \left( \frac{\partial x}{\partial s} \mathbf{i} + \frac{\partial y}{\partial s} \mathbf{j} + \frac{\partial z}{\partial s} \mathbf{k} \right) \quad (110)$$

is a unit tangent vector to the differentiable curve  $\mathbf{r}(t)$  through the region  $R$ .  $P_0(x_0, y_0, z_0)$  is a point at  $\mathbf{r}(t)$  in  $R$ . Then the derivative of  $f$  at  $P_0$  along the unit tangent vector is the number

$$\begin{aligned} \frac{\partial f}{\partial s} &= \frac{\partial f}{\partial x} \frac{\partial x}{\partial s} + \frac{\partial f}{\partial y} \frac{\partial y}{\partial s} + \frac{\partial f}{\partial z} \frac{\partial z}{\partial s} = \left( \frac{\partial f}{\partial x} \mathbf{i} + \frac{\partial f}{\partial y} \mathbf{j} + \frac{\partial f}{\partial z} \mathbf{k} \right) \cdot \left( \frac{\partial x}{\partial s} \mathbf{i} + \frac{\partial y}{\partial s} \mathbf{j} + \frac{\partial z}{\partial s} \mathbf{k} \right) \\ &= (\nabla f) \cdot \frac{\partial \mathbf{r}}{\partial s} = (\nabla f) \cdot \mathbf{T}. \end{aligned} \quad (111)$$

**The vector (cross) products:** If  $\mathbf{A}$  and  $\mathbf{B}$  are two non-zero vectors in space which are not parallel, they determine a plane. We select a unit vector  $\mathbf{n}$  perpendicular to the plane by the *right-handed rule*. This means we choose  $\mathbf{n}$  to be unit (normal) vector that points the way your right thumb points when your fingers curl through the angle  $\theta$  from  $\mathbf{A}$  to  $\mathbf{B}$ . We then define the *vector product*  $\mathbf{A} \times \mathbf{B}$  (“ $\mathbf{A}$  cross  $\mathbf{B}$ ”) to be the vector

$$\mathbf{A} \times \mathbf{B} = (|\mathbf{A}||\mathbf{B}|\sin \theta) \mathbf{n}. \quad (112)$$

The vector  $\mathbf{A} \times \mathbf{B}$  is orthogonal to both  $\mathbf{A}$  and  $\mathbf{B}$  because it is a scalar multiple of  $\mathbf{n}$ . The vector product of  $\mathbf{A}$  and  $\mathbf{B}$  is often called the cross product of  $\mathbf{A}$  and  $\mathbf{B}$  because of the cross in the notation  $\mathbf{A} \times \mathbf{B}$ .

Since the sines of 0 and  $\pi$  are both zero in the definition equation, it makes sense to define the cross product of two parallel non-zero vectors to be  $\mathbf{0}$ .

If one or both of the vectors  $\mathbf{A}$  and  $\mathbf{B}$  are zero, we also define  $\mathbf{A} \times \mathbf{B}$  to be zero. This way, the cross product of two vectors  $\mathbf{A}$  and  $\mathbf{B}$  is zero if and only if  $\mathbf{A}$  and  $\mathbf{B}$  are parallel or one or both of them are zero.

Note that

$$\mathbf{A} \times \mathbf{B} = -(\mathbf{B} \times \mathbf{A}). \quad (113)$$

**The determination formula for  $\mathbf{A} \times \mathbf{B}$ :** If  $\mathbf{A} = a_1\mathbf{i} + a_2\mathbf{j} + a_3\mathbf{k}$ , and  $\mathbf{B} = b_1\mathbf{i} + b_2\mathbf{j} + b_3\mathbf{k}$ , then

$$\mathbf{A} \times \mathbf{B} = \begin{vmatrix} \mathbf{i} & \mathbf{j} & \mathbf{k} \\ a_1 & a_2 & a_3 \\ b_1 & b_2 & b_3 \end{vmatrix}. \quad (114)$$

As an example consider the expression appearing in the multi-phase fluid mechanics for computation of the lift forces  $(\mathbf{V}_l - \mathbf{V}_m) \times (\nabla \times \mathbf{V}_m)$  as a function of the velocity vectors. The result is given in two steps

$$(\nabla \times \mathbf{V}_m) = \begin{vmatrix} \mathbf{i} & \mathbf{j} & \mathbf{k} \\ \frac{\partial}{\partial x} & \frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ u_m & v_m & w_m \end{vmatrix} = \left( \frac{\partial w_m}{\partial y} - \frac{\partial v_m}{\partial z} \right) \mathbf{i} - \left( \frac{\partial w_m}{\partial x} - \frac{\partial u_m}{\partial z} \right) \mathbf{j} + \left( \frac{\partial v_m}{\partial x} - \frac{\partial u_m}{\partial y} \right) \mathbf{k}, \quad (115)$$

$$\begin{aligned} (\mathbf{V}_l - \mathbf{V}_m) \times (\nabla \times \mathbf{V}_m) &= \\ &= \left[ (v_l - v_m) \left( \frac{\partial v_m}{\partial x} - \frac{\partial u_m}{\partial y} \right) + (w_l - w_m) \left( \frac{\partial w_m}{\partial x} - \frac{\partial u_m}{\partial z} \right) \right] \mathbf{i} \\ &- \left[ (u_l - u_m) \left( \frac{\partial v_m}{\partial x} - \frac{\partial u_m}{\partial y} \right) - (w_l - w_m) \left( \frac{\partial w_m}{\partial y} - \frac{\partial v_m}{\partial z} \right) \right] \mathbf{j} \\ &+ \left[ -(u_l - u_m) \left( \frac{\partial w_m}{\partial x} - \frac{\partial u_m}{\partial z} \right) - (v_l - v_m) \left( \frac{\partial w_m}{\partial y} - \frac{\partial v_m}{\partial z} \right) \right] \mathbf{k}. \end{aligned} \quad (116)$$

**Parallel vectors:** Non-zero vectors  $\mathbf{A}$  and  $\mathbf{B}$  are parallel if and only if

$$\mathbf{A} \times \mathbf{B} = \mathbf{0}. \quad (117)$$

**The associative and distributive laws for the cross product:** The scalar distributive law is

$$(r\mathbf{A}) \times (s\mathbf{B}) = (rs)(\mathbf{A} \times \mathbf{B}), \quad (118)$$

with the special case

$$(-\mathbf{A}) \times \mathbf{B} = \mathbf{A} \times (-\mathbf{B}) = -(\mathbf{A} \times \mathbf{B}). \quad (119)$$

The vector distributive laws are

$$\mathbf{A} \times (\mathbf{B} + \mathbf{C}) = \mathbf{A} \times \mathbf{B} + \mathbf{A} \times \mathbf{C}, \quad (120)$$

$$(\mathbf{B} + \mathbf{C}) \times \mathbf{A} = \mathbf{B} \times \mathbf{A} + \mathbf{C} \times \mathbf{A}. \quad (121)$$

Note the interesting vector identities, *Thompson et al. (1985)*, p. 100 and 101,

$$(\mathbf{A} \times \mathbf{B}) \cdot (\mathbf{C} \times \mathbf{D}) = (\mathbf{A} \cdot \mathbf{C})(\mathbf{B} \cdot \mathbf{D}) - (\mathbf{A} \cdot \mathbf{D})(\mathbf{B} \cdot \mathbf{C}), \quad (122)$$

and

$$\mathbf{A} \times (\mathbf{B} \times \mathbf{C}) = (\mathbf{A} \cdot \mathbf{C})\mathbf{B} - (\mathbf{A} \cdot \mathbf{B})\mathbf{C}. \quad (123)$$

**The area of a parallelogram:** Because  $\mathbf{n}$  is a unit vector, the magnitude of  $\mathbf{A} \times \mathbf{B}$  is

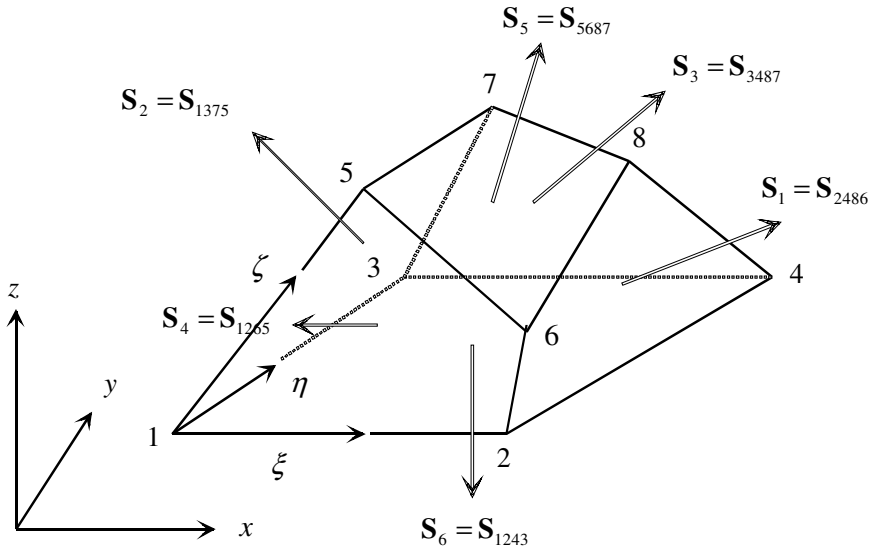
$$|\mathbf{A} \times \mathbf{B}| = |\mathbf{A}||\mathbf{B}||\sin \theta| |\mathbf{n}| = |\mathbf{A}||\mathbf{B}||\sin \theta|. \quad (124)$$

This is the area of the parallelogram determined by  $\mathbf{A}$  and  $\mathbf{B}$ ,  $|\mathbf{A}|$  being the base of the parallelogram and  $|\mathbf{B}||\sin \theta|$  the height.

**Area of union of two triangle plane regions:** Consider four vertices points defined by the position vector

$$\mathbf{r}_i = x_i \mathbf{i} + y_i \mathbf{j} + z_i \mathbf{k}, \quad (125)$$

where  $i = 5, 6, 8, 7$ , see Fig. A1.2.



**Fig. A1.2** Tetrahedron defined by joint triangles

In general the four points must not belong to the same plane. The surface vector defined by these points is computed as follows. First we compute the surface vectors of the triangles 587 and 568

$$\mathbf{S}_{587} = \frac{1}{2}(\mathbf{r}_5 - \mathbf{r}_7) \times (\mathbf{r}_8 - \mathbf{r}_7), \quad (126)$$

$$\mathbf{S}_{568} = \frac{1}{2}(\mathbf{r}_6 - \mathbf{r}_8) \times (\mathbf{r}_5 - \mathbf{r}_6). \quad (127)$$

The vector sum of the above two vectors yields an expression for the vector surface

$$\mathbf{S}_{5687} = \frac{1}{2}(\mathbf{r}_6 - \mathbf{r}_5) \times (\mathbf{r}_8 - \mathbf{r}_5), \quad (128)$$

*Kordulla and Vinokur (1983)*. In fact this is the surface defined by the vector product of the vectors joining the couple of the opposite vertices, respectively. One should note that the first, the second and the resulting vector form right handed coordinate systems in order to specify the outward direction of the surface if it is part of the control volume surface. Thus, for the hexadron presented in Fig. A1.2, we have

$$\mathbf{S}_1 = \mathbf{S}_{2486} = \frac{1}{2}(\mathbf{r}_4 - \mathbf{r}_6) \times (\mathbf{r}_8 - \mathbf{r}_2), \quad (129)$$

$$\mathbf{S}_2 = \mathbf{S}_{1375} = \frac{1}{2}(\mathbf{r}_3 - \mathbf{r}_5) \times (\mathbf{r}_1 - \mathbf{r}_7), \quad (130)$$

$$\mathbf{S}_3 = \mathbf{S}_{3487} = \frac{1}{2}(\mathbf{r}_4 - \mathbf{r}_7) \times (\mathbf{r}_3 - \mathbf{r}_8), \quad (131)$$

$$\mathbf{S}_4 = \mathbf{S}_{1265} = \frac{1}{2}(\mathbf{r}_6 - \mathbf{r}_1) \times (\mathbf{r}_5 - \mathbf{r}_2), \quad (132)$$

$$\mathbf{S}_5 = \mathbf{S}_{5687} = \frac{1}{2}(\mathbf{r}_8 - \mathbf{r}_5) \times (\mathbf{r}_7 - \mathbf{r}_6), \quad (133)$$

$$\mathbf{S}_6 = \mathbf{S}_{1243} = \frac{1}{2}(\mathbf{r}_4 - \mathbf{r}_1) \times (\mathbf{r}_2 - \mathbf{r}_3). \quad (134)$$

**The triple scalar or box product:** The product  $(\mathbf{A} \times \mathbf{B}) \cdot \mathbf{C}$  is called the triple scalar product of  $\mathbf{A}$ ,  $\mathbf{B}$ , and  $\mathbf{C}$  (in that order). As you can see from the formula

$$|(\mathbf{A} \times \mathbf{B}) \cdot \mathbf{C}| = |\mathbf{A} \times \mathbf{B}| |\mathbf{C}| \cos \theta, \quad (135)$$

the absolute value of the product is the volume of the parallelepiped (parallelogram-sided box) determined by  $\mathbf{A}$ ,  $\mathbf{B}$ , and  $\mathbf{C}$ . The number  $|\mathbf{A} \times \mathbf{B}|$  is the area of the parallelogram. The number  $|\mathbf{C}| \cos \theta$  is the parallelogram's height. Because of the geometry  $(\mathbf{A} \times \mathbf{B}) \cdot \mathbf{C}$  is called the *box product* of  $\mathbf{A}$ ,  $\mathbf{B}$ , and  $\mathbf{C}$ .

By treating the planes of  $\mathbf{B}$  and  $\mathbf{C}$  and of  $\mathbf{C}$  and  $\mathbf{A}$  as the base planes of the parallelepiped determined by  $\mathbf{A}$ ,  $\mathbf{B}$ , and  $\mathbf{C}$ , we see that

$$(\mathbf{A} \times \mathbf{B}) \cdot \mathbf{C} = (\mathbf{B} \times \mathbf{C}) \cdot \mathbf{A} = (\mathbf{C} \times \mathbf{A}) \cdot \mathbf{B}. \quad (136)$$

Since the dot product is commutative, the above equation gives

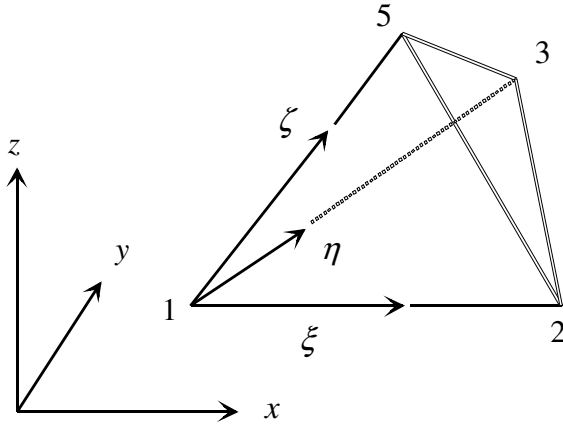
$$(\mathbf{A} \times \mathbf{B}) \cdot \mathbf{C} = \mathbf{A} \cdot (\mathbf{B} \times \mathbf{C}). \quad (137)$$

The triple scalar product can be evaluated as a determinant

$$(\mathbf{A} \times \mathbf{B}) \cdot \mathbf{C} = \begin{vmatrix} a_1 & a_2 & a_3 \\ b_1 & b_2 & b_3 \\ c_1 & c_2 & c_3 \end{vmatrix}. \quad (138)$$

**Volume of a tetrahedron:** Consider a tetrahedron defined by the four vertices numbered as shown in Fig. A1.3.



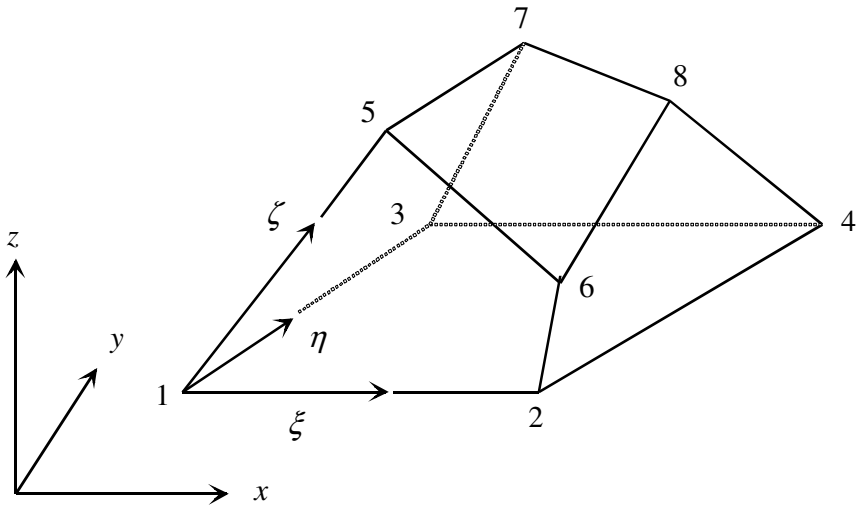


**Fig. A1.3** Definition of the vertices numbering of the tetrahedron

The volume of the tetrahedron is then

$$V = \frac{1}{6}(\mathbf{r}_5 - \mathbf{r}_1) \cdot [(\mathbf{r}_2 - \mathbf{r}_1) \times (\mathbf{r}_3 - \mathbf{r}_1)]. \quad (139)$$

**Volume of a hexahedron:** Consider a hexahedron defined by the eight vertices numbered starting with base in the counterclockwise direction as shown in Fig. A1.4.



**Fig. A1.4** Definition of the vertices numbering of the hexahedron

An alternative approach for computation of the volume of the hexahedron is to compute the volumes of the tetrahedra making up the hexahedron. *Kordulla* and *Vinokur* (1983) used symmetric partitioning of the faces decomposing the hexahedron into four tetrahedra and obtained

$$V_{12435687} = -\frac{1}{3}(\mathbf{r}_8 - \mathbf{r}_1) \cdot (\mathbf{S}_{1243} + \mathbf{S}_{1265} + \mathbf{S}_{1375}) = -\frac{1}{3}(\mathbf{r}_8 - \mathbf{r}_1) \cdot (\mathbf{S}_2 + \mathbf{S}_4 + \mathbf{S}_6). \quad (140)$$

Here the surfaces  $\mathbf{S}_{1243}$ ,  $\mathbf{S}_{1265}$ , and  $\mathbf{S}_{1375}$  belongs to the same vertices and  $\mathbf{r}_8 - \mathbf{r}_1$  is the diagonal joining these vertices with the opposite ones.

**Dyadic product of two vectors:** Consider the vectors  $\mathbf{A} = a_1\mathbf{i} + a_2\mathbf{j} + a_3\mathbf{k}$ ,  $\mathbf{B} = b_1\mathbf{i} + b_2\mathbf{j} + b_3\mathbf{k}$ . The dyadic product of the two vectors is written as  $\mathbf{AB}$  without any sign between them, and is defined as the following second order tensor

$$\mathbf{AB} = \begin{pmatrix} a_1b_1 & a_1b_2 & a_1b_3 \\ a_2b_1 & a_2b_2 & a_2b_3 \\ a_3b_1 & a_3b_2 & a_3b_3 \end{pmatrix}. \quad (141)$$

Note that the dyadic product of  $\mathbf{AB}$  is not equivalent to  $\mathbf{BA}$ . As examples we give some dyadic products between the vectors  $\nabla = \frac{\partial}{\partial x}\mathbf{i} + \frac{\partial}{\partial y}\mathbf{j} + \frac{\partial}{\partial z}\mathbf{k}$  and  $\mathbf{V} = u\mathbf{i} + v\mathbf{j} + w\mathbf{k}$ , which are used in fluid mechanics for definition of the friction forces in flows:

$$\nabla\mathbf{V} = \begin{pmatrix} \frac{\partial u}{\partial x} & \frac{\partial v}{\partial x} & \frac{\partial w}{\partial x} \\ \frac{\partial u}{\partial y} & \frac{\partial v}{\partial y} & \frac{\partial w}{\partial y} \\ \frac{\partial u}{\partial z} & \frac{\partial v}{\partial z} & \frac{\partial w}{\partial z} \end{pmatrix}, \quad (142)$$

$$(\nabla\mathbf{V})^T = \begin{pmatrix} \frac{\partial u}{\partial x} & \frac{\partial u}{\partial y} & \frac{\partial u}{\partial z} \\ \frac{\partial v}{\partial x} & \frac{\partial v}{\partial y} & \frac{\partial v}{\partial z} \\ \frac{\partial w}{\partial x} & \frac{\partial w}{\partial y} & \frac{\partial w}{\partial z} \end{pmatrix}. \quad (144)$$

Another dyadic product of the velocity vector

$$\mathbf{V}\mathbf{V} = \begin{pmatrix} uu & uv & uw \\ vu & vv & vw \\ wu & wv & ww \end{pmatrix}, \quad (145)$$

is used to describe the convective momentum transport.

**Eigenvalues:** The eigenvalues  $\lambda_i$  of a matrix  $\mathbf{A}$  are the solution of the characteristic polynomial

$$|\mathbf{A} - \lambda\mathbf{I}| = \det(\mathbf{A} - \lambda\mathbf{I}) = 0, \quad (146)$$

where  $\mathbf{I}$  is the identity matrix.

**Eigenvectors:** The right eigenvector of a matrix  $\mathbf{A}$  corresponding to an eigenvalue  $\lambda_i$  of  $\mathbf{A}$  is a vector  $\mathbf{K}^{(i)} = [k_1^{(i)}, k_2^{(i)}, \dots, k_m^{(i)}]^T$  satisfying

$$\mathbf{A}\mathbf{K}^{(i)} = \lambda_i\mathbf{K}^{(i)}. \quad (147)$$

Similarly, a left eigenvector of a matrix  $\mathbf{A}$  corresponding to an eigenvalue  $\lambda_i$  of  $\mathbf{A}$  is a vector  $\mathbf{L}^{(i)} = [k_1^{(i)}, k_2^{(i)}, \dots, k_m^{(i)}]^T$  satisfying

$$\mathbf{A}\mathbf{L}^{(i)} = \lambda_i\mathbf{L}^{(i)}. \quad (148)$$

**Diagonalizable matrix:** A matrix  $\mathbf{A}$  is said to be diagonalizable if  $\mathbf{A}$  can be expressed as

$$\mathbf{A} = \mathbf{K}\mathbf{H}\mathbf{K}^{-1}, \quad (149)$$

in terms of a diagonal matrix  $\mathbf{H}$  and a matrix  $\mathbf{K}$ . The diagonal element of  $\mathbf{H}$  are the eigenvalues  $\lambda_i$  of  $\mathbf{A}$  and the columns of  $\mathbf{K}$  are the right eigenvectors of  $\mathbf{A}$  corresponding to the eigenvalues  $\lambda_i$ , that is

$$\mathbf{H} = \begin{bmatrix} \lambda_1 & 0 & \cdot & 0 \\ 0 & \lambda_2 & \cdot & 0 \\ \cdot & \cdot & \cdot & \cdot \\ 0 & 0 & \cdot & \lambda_m \end{bmatrix}, \quad \mathbf{K} = [\mathbf{K}^{(1)}, \mathbf{K}^{(2)}, \dots, \mathbf{K}^{(m)}]^T, \quad \mathbf{A}\mathbf{K}^{(i)} = \lambda_i\mathbf{K}^{(i)}. \quad (150)$$

**Rotation around axes:** Given a point  $(x_0, y_0, z_0)$ . A new position of this point  $(x, y, z)$  obtained after rotation by angle  $\varphi$  with respect to the  $x$ ,  $y$  or  $z$  axes respectively is

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = \mathbf{R}_i \begin{pmatrix} x_0 \\ y_0 \\ z_0 \end{pmatrix}, \quad (151)$$

where

$$\mathbf{R}_1 = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \varphi & -\sin \varphi \\ 0 & \sin \varphi & \cos \varphi \end{pmatrix}, \mathbf{R}_2 = \begin{pmatrix} \cos \varphi & 0 & \sin \varphi \\ 0 & 1 & 0 \\ -\sin \varphi & 0 & \cos \varphi \end{pmatrix}, \mathbf{R}_3 = \begin{pmatrix} \cos \varphi & -\sin \varphi & 0 \\ \sin \varphi & \cos \varphi & 0 \\ 0 & 0 & 1 \end{pmatrix}. \quad (152,153,154)$$

**Scaling:** Given a point  $(x_0, y_0, z_0)$ . A new position of this point  $(x, y, z)$  obtained after scaling by

$$\mathbf{S}_1 = \begin{pmatrix} s_x & 0 & 0 \\ 0 & s_y & 0 \\ 0 & 0 & s_z \end{pmatrix}, \quad (155)$$

is

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = \mathbf{S}_i \begin{pmatrix} x_0 \\ y_0 \\ z_0 \end{pmatrix}. \quad (156)$$

**Translation:** Graphical systems make use of the general notation of transformation operation defining by multiplying the coordinates with a 4 by 4 matrix

$$\begin{pmatrix} x \\ y \\ z \\ w \end{pmatrix} = \mathbf{T} \begin{pmatrix} x_0 \\ y_0 \\ z_0 \\ 1 \end{pmatrix}, \quad (157)$$

called translation matrix

$$\mathbf{T} = \begin{pmatrix} 1 & 0 & 0 & \Delta x \\ 0 & 1 & 0 & \Delta y \\ 0 & 0 & 1 & \Delta z \\ 0 & 0 & 0 & 1 \end{pmatrix}. \quad (158)$$

$w$  is used for hardware scaling in graphical representations. Replacing the upper 3 by 3 elements with the matrices defining rotation or scaling, results in general notation of transformation. Then multiple transformations on objects are simply sequences of matrices multiplications. Note that the round off errors may damage

the objects coordinates after many multiplications. Therefore one should design systems by starting the transformations always from a initial state.

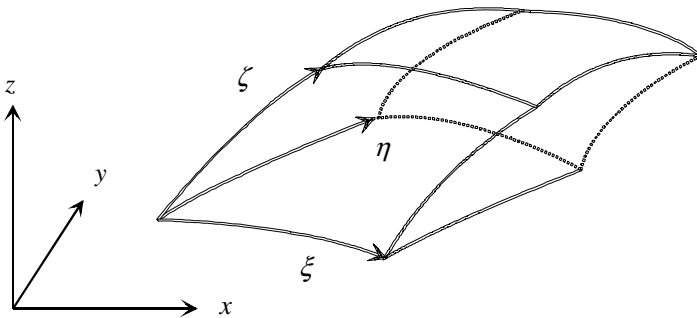
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## Appendix 2 Basics of the coordinate transformation theory

Cartesian coordinate systems are usually the first choice of the scientist and the engineer. Unfortunately in nature and technology the boundaries of a flow may have a very complicated form. Prescribing physically adequate boundary condition at such surfaces is frequently not a solvable problem. The next choice of course is one of the well-known curvilinear orthogonal coordinate systems, polar, or bipolar, or cylindrical.

The purpose of this Appendix is to collect from the literature the most important basics on which the general coordinate transformation theory relies. The material is ordered in a generic way. Each statement follows from the already defined statement or statements. For those who would like to perfect his knowledge in this field there are two important books that are my favorite choices: *Thomas et al. (1998)* - about the calculus and analytic geometry, and *Thompson et al. (1985)* - about numerical grid generation. For a complete collection of basic definitions, rules and useful formula see *Miki and Takagi (1984)*.



**Fig. A2.1** Cartesian and curvilinear coordinate systems

Consider three non identical and non-parallel curves in the space  $\xi, \eta, \zeta$  having only one common point designated as an origin and presented in Fig. A2.1. The curves are smooth (at least one times differentiable). We call these curves *coordinate lines* or *curvilinear coordinates*. The curvilinear coordinate lines of a three-dimensional system can be considered also as space curves formed by the intersection of surfaces on which one of the coordinates is hold constant. One

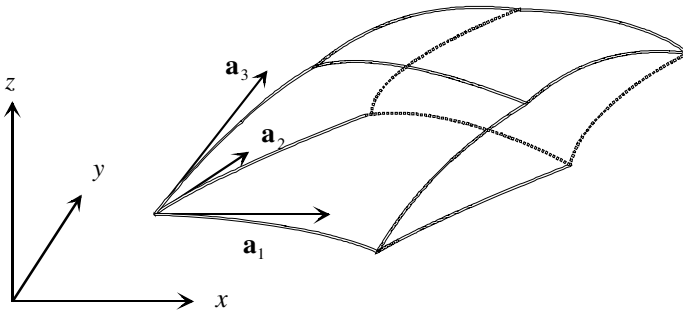
coordinate varies along a coordinate line, while the other two are constant thereon. Thus, we have the coordinate transformation defined by

$$x = f(\xi, \eta, \zeta), \quad (1)$$

$$y = g(\xi, \eta, \zeta), \quad (2)$$

$$z = h(\xi, \eta, \zeta). \quad (3)$$

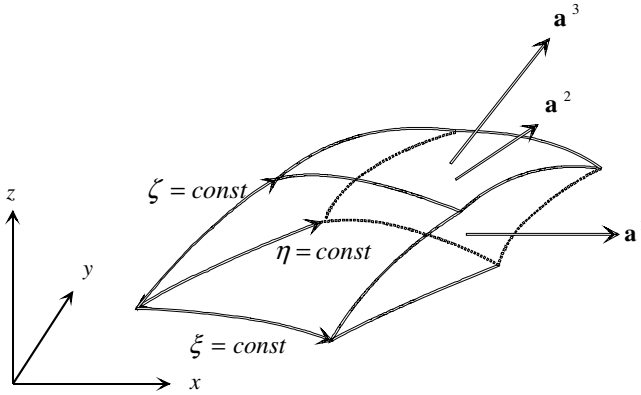
The surfaces defined by  $\xi = \text{const}$ ,  $\eta = \text{const}$ ,  $\zeta = \text{const}$ , are called coordinate surfaces. We talk then of physical and transformed (computational) space. In the *physical space* the coordinates  $(x, y, z)$  become independent variables. In the *transformed space* the transformed computational coordinates  $(\xi, \eta, \zeta)$  become independent variables.



**Fig. A2.2** The covariant tangent vectors to the coordinate lines form the so called covariant base vectors of the curvilinear coordinate system

The tangent vectors to the coordinate lines form the so called *base vectors of the coordinate system*. These base vectors are called *covariant base vectors*, Fig. A2.2. The normal vectors to the coordinate surfaces form the so called *contravariant base vectors*. The two types of the base vectors are illustrated in Figs. A2.2 and A2.3, showing an element of volume with six sides, each of which lies on some coordinate surface.

Another frequently used notation is  $x_i$  ( $i=1,2,3$ ):  $x_1, x_2, x_3$  and  $\xi^i$  ( $i=1,2,3$ ):  $\xi^1, \xi^2, \xi^3$ . In the latter case the superscripts serve only as labels and not as powers.



**Fig. A2.3** The contravariant normal vectors to the coordinate surfaces form the so called contravariant base vectors of the curvilinear coordinate system

**Metrics and inverse metrics:** The differential increments in the transformed coordinate system as a function of the differential increments in the physical coordinate system can be computed as follows

$$\begin{pmatrix} d\xi \\ d\eta \\ d\zeta \end{pmatrix} = \begin{pmatrix} \frac{\partial \xi}{\partial x} & \frac{\partial \xi}{\partial y} & \frac{\partial \xi}{\partial z} \\ \frac{\partial \eta}{\partial x} & \frac{\partial \eta}{\partial y} & \frac{\partial \eta}{\partial z} \\ \frac{\partial \zeta}{\partial x} & \frac{\partial \zeta}{\partial y} & \frac{\partial \zeta}{\partial z} \end{pmatrix} \begin{pmatrix} dx \\ dy \\ dz \end{pmatrix}. \quad (4)$$

The elements of the matrix in Eq. (4) are called *metrics* of the coordinate transformation, *Anderson* (1995) p.178. An alternative notation of the metrics is  $a^{ij}$ . The differential increments in the physical coordinate system are a function of the differential increments in the transformed coordinate system

$$\begin{pmatrix} dx \\ dy \\ dz \end{pmatrix} = \begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix} \begin{pmatrix} d\xi \\ d\eta \\ d\zeta \end{pmatrix} \quad \text{or} \quad \begin{pmatrix} dx \\ dy \\ dz \end{pmatrix} = \mathbf{J} \begin{pmatrix} d\xi \\ d\eta \\ d\zeta \end{pmatrix}. \quad (5)$$

where



$$\mathbf{J}(\xi, \eta, \zeta) = \begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix} \quad (6)$$

is the so called *Jacobian matrix* of the coordinate transformation  $x = f(\xi, \eta, \zeta)$ ,  $y = g(\xi, \eta, \zeta)$ ,  $z = h(\xi, \eta, \zeta)$ . The elements of the *Jacobian matrix* are called *inverse metrics* of the coordinate transformation Anderson (1995). An alternative notation of the inverse metrics is  $a_{ij}^T$ . The relation between the so called *metrics* and the inverse metrics is obvious by noting that the transformation is invertible

$$\begin{pmatrix} d\xi \\ d\eta \\ d\zeta \end{pmatrix} = \begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix}^{-1} \begin{pmatrix} dx \\ dy \\ dz \end{pmatrix} = \begin{pmatrix} \frac{\partial \xi}{\partial x} & \frac{\partial \xi}{\partial y} & \frac{\partial \xi}{\partial z} \\ \frac{\partial \eta}{\partial x} & \frac{\partial \eta}{\partial y} & \frac{\partial \eta}{\partial z} \\ \frac{\partial \zeta}{\partial x} & \frac{\partial \zeta}{\partial y} & \frac{\partial \zeta}{\partial z} \end{pmatrix} \begin{pmatrix} dx \\ dy \\ dz \end{pmatrix}, \quad (7)$$

and therefore

$$\begin{aligned} \mathbf{J}^{-1} &= \begin{pmatrix} \frac{\partial \xi}{\partial x} & \frac{\partial \xi}{\partial y} & \frac{\partial \xi}{\partial z} \\ \frac{\partial \eta}{\partial x} & \frac{\partial \eta}{\partial y} & \frac{\partial \eta}{\partial z} \\ \frac{\partial \zeta}{\partial x} & \frac{\partial \zeta}{\partial y} & \frac{\partial \zeta}{\partial z} \end{pmatrix} = \begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix}^{-1} \\ &= \frac{1}{\sqrt{g}} \begin{pmatrix} \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} & \frac{\partial x}{\partial \zeta} \frac{\partial z}{\partial \eta} - \frac{\partial x}{\partial \eta} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \zeta} - \frac{\partial x}{\partial \zeta} \frac{\partial y}{\partial \eta} \\ \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \xi} - \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \zeta} \frac{\partial z}{\partial \xi} - \frac{\partial x}{\partial \xi} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \zeta} \frac{\partial y}{\partial \xi} - \frac{\partial x}{\partial \xi} \frac{\partial y}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \eta} - \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \xi} & \frac{\partial x}{\partial \eta} \frac{\partial z}{\partial \xi} - \frac{\partial x}{\partial \xi} \frac{\partial z}{\partial \eta} & \frac{\partial x}{\partial \xi} \frac{\partial y}{\partial \eta} - \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \xi} \end{pmatrix}, \quad (8) \end{aligned}$$

where

$$\sqrt{g} = |\mathbf{J}(\xi, \eta, \zeta)|$$

$$= \frac{\partial x}{\partial \xi} \left( \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} \right) - \frac{\partial x}{\partial \eta} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \xi} \right) + \frac{\partial x}{\partial \zeta} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \eta} - \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \xi} \right), \quad (9)$$

is called the *Jacobian determinant* or *Jacobian* of the coordinate transformation. An alternative notation of the metrics is  $a^{ij}$ . We see that

$$\mathbf{J} = a_{ij}^T = (a^{ij})^{-1}, \quad (10)$$

$$a^{ij} = \mathbf{J}^{-1} = (a_{ij}^T)^{-1}, \quad (11)$$

$$a_{ij} = \left[ (a^{ij})^{-1} \right]^T = \left[ (a^{ij})^T \right]^{-1}. \quad (12)$$

**Covariant base vectors:** The *tangent vectors*  $(\mathbf{a}_1, \mathbf{a}_2, \mathbf{a}_3)$  to the three curvilinear coordinate lines represented by  $(\xi, \eta, \zeta)$  are called the three *covariant* base vectors of the curvilinear coordinate system, and are designated with

$$\mathbf{a}_1 : \frac{\partial \mathbf{r}}{\partial \xi} = \frac{\partial x}{\partial \xi} \mathbf{i} + \frac{\partial y}{\partial \xi} \mathbf{j} + \frac{\partial z}{\partial \xi} \mathbf{k}, \quad (13)$$

$$\mathbf{a}_2 : \frac{\partial \mathbf{r}}{\partial \eta} = \frac{\partial x}{\partial \eta} \mathbf{i} + \frac{\partial y}{\partial \eta} \mathbf{j} + \frac{\partial z}{\partial \eta} \mathbf{k}, \quad (14)$$

$$\mathbf{a}_3 : \frac{\partial \mathbf{r}}{\partial \zeta} = \frac{\partial x}{\partial \zeta} \mathbf{i} + \frac{\partial y}{\partial \zeta} \mathbf{j} + \frac{\partial z}{\partial \zeta} \mathbf{k}. \quad (15)$$

The components of the three *covariant* base vectors of the curvilinear coordinate system are the columns of the matrix of the inverse metrics (Jacobian matrix – Eq. 6). They are *not* unit vectors. The corresponding unit vectors are

$$\mathbf{e}_1 = \frac{\mathbf{a}_1}{|\mathbf{a}_1|}, \quad \mathbf{e}_2 = \frac{\mathbf{a}_2}{|\mathbf{a}_2|}, \quad \mathbf{e}_3 = \frac{\mathbf{a}_3}{|\mathbf{a}_3|}. \quad (16,17,18)$$

**Contravariant base vectors:** The *normal vectors* to a coordinate surface on which the coordinates  $\xi$ ,  $\eta$  and  $\zeta$  are constant, respectively are given by

$$\mathbf{a}^1 : \nabla \xi = \frac{\partial \xi}{\partial x} \mathbf{i} + \frac{\partial \xi}{\partial y} \mathbf{j} + \frac{\partial \xi}{\partial z} \mathbf{k}, \quad (19)$$

$$\mathbf{a}^2 : \nabla \eta = \frac{\partial \eta}{\partial x} \mathbf{i} + \frac{\partial \eta}{\partial y} \mathbf{j} + \frac{\partial \eta}{\partial z} \mathbf{k}, \quad (20)$$

$$\mathbf{a}^3 : \nabla \zeta = \frac{\partial \zeta}{\partial x} \mathbf{i} + \frac{\partial \zeta}{\partial y} \mathbf{j} + \frac{\partial \zeta}{\partial z} \mathbf{k} . \quad (21)$$

The components of the three *contravariant* base vectors of the curvilinear coordinate system are the rows of the matrix of the metrics – Eq. (4). They are *not* unit vectors. The corresponding unit vectors are

$$\xi = \text{const} , \quad \mathbf{e}^1 = \frac{\mathbf{a}^1}{|\mathbf{a}^1|} , \quad (22)$$

$$\eta = \text{const} , \quad \mathbf{e}^2 = \frac{\mathbf{a}^2}{|\mathbf{a}^2|} , \quad (23)$$

$$\zeta = \text{const} , \quad \mathbf{e}^3 = \frac{\mathbf{a}^3}{|\mathbf{a}^3|} . \quad (24)$$

At any given point on the surface defined with  $\xi = \text{const}$  ,  $\xi$  increases most rapidly in the direction of  $\nabla \xi$  and decreases most rapidly in the direction of  $-\nabla \xi$  . In any direction orthogonal to  $\nabla \xi$  , the derivative is zero. The corresponding statements are valid for the remaining surfaces  $\eta = \text{const}$  and  $\zeta = \text{const}$  .

These normal vectors to the three coordinate surfaces are called the three *contravariant* base vectors of the curvilinear coordinate system.

**Cartesian vector in co- and contravariant coordinate systems:** Any Cartesian vector  $\mathbf{V} = u\mathbf{i} + v\mathbf{j} + w\mathbf{k}$  has two sets of distinct components with respect to the frames of the covariant and contravariant base vectors

$$\mathbf{V} = V^1 \mathbf{a}_1 + V^2 \mathbf{a}_2 + V^3 \mathbf{a}_3 , \quad (25)$$

$$\mathbf{V} = V'_1 \mathbf{a}^1 + V'_2 \mathbf{a}^2 + V'_3 \mathbf{a}^3 . \quad (26)$$

The components  $V^i$  and  $V'_i$  are called *contravariant* and *covariant components* of  $\mathbf{V}$  respectively. The components are easily computed by equalizing the Cartesian components as follows

$$\begin{aligned} u\mathbf{i} + v\mathbf{j} + w\mathbf{k} &= V^1 \mathbf{a}_1 + V^2 \mathbf{a}_2 + V^3 \mathbf{a}_3 \\ &= (V^1 a_{11} + V^2 a_{21} + V^3 a_{31}) \mathbf{i} + (V^1 a_{12} + V^2 a_{22} + V^3 a_{32}) \mathbf{j} \\ &+ (V^1 a_{13} + V^2 a_{23} + V^3 a_{33}) \mathbf{k} , \end{aligned} \quad (27)$$

or

$$\begin{pmatrix} a_{11} & a_{21} & a_{31} \\ a_{12} & a_{22} & a_{32} \\ a_{13} & a_{23} & a_{33} \end{pmatrix} \begin{pmatrix} V'^1 \\ V'^2 \\ V'^3 \end{pmatrix} = \begin{pmatrix} u \\ v \\ w \end{pmatrix} \quad \text{or} \quad \mathbf{J} \begin{pmatrix} V'^1 \\ V'^2 \\ V'^3 \end{pmatrix} = \begin{pmatrix} u \\ v \\ w \end{pmatrix} \quad (28)$$

or after comparing with Eq. (9) we have

$$\begin{pmatrix} V'^1 \\ V'^2 \\ V'^3 \end{pmatrix} = \begin{pmatrix} a_{11} & a_{21} & a_{31} \\ a_{12} & a_{22} & a_{32} \\ a_{13} & a_{23} & a_{33} \end{pmatrix}^{-1} \begin{pmatrix} u \\ v \\ w \end{pmatrix} = \mathbf{J}^{-1} \cdot \mathbf{V}. \quad (29)$$

Therefore

$$V'^1 = a^{11}u + a^{12}v + a^{13}w = \mathbf{a}^1 \cdot \mathbf{V}, \quad (30)$$

$$V'^2 = a^{21}u + a^{22}v + a^{23}w = \mathbf{a}^2 \cdot \mathbf{V}, \quad (31)$$

$$V'^3 = a^{31}u + a^{32}v + a^{33}w = \mathbf{a}^3 \cdot \mathbf{V}, \quad (32)$$

and Eq. (25) can be rewritten as follows

$$\mathbf{V} = (\mathbf{a}^1 \cdot \mathbf{V})\mathbf{a}_1 + (\mathbf{a}^2 \cdot \mathbf{V})\mathbf{a}_2 + (\mathbf{a}^3 \cdot \mathbf{V})\mathbf{a}_3, \quad (33)$$

compare with Eq. (35) in *Thompson, Warsi and Mastin* (1985) p. 109. Similarly we have

$$\begin{pmatrix} V'_1 \\ V'_2 \\ V'_3 \end{pmatrix} = \begin{pmatrix} a^{11} & a^{21} & a^{31} \\ a^{12} & a^{22} & a^{32} \\ a^{13} & a^{23} & a^{33} \end{pmatrix}^{-1} \begin{pmatrix} u \\ v \\ w \end{pmatrix} = \left[ (a^{ij})^T \right]^{-1} = \left[ (a^{ij})^{-1} \right]^T = a_{ij}^T = \mathbf{J} \cdot \mathbf{V}, \quad (34)$$

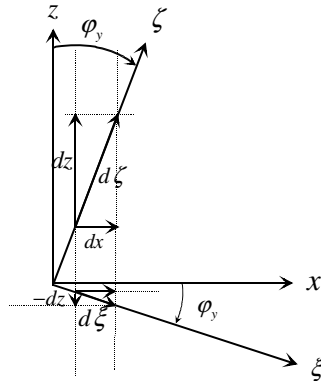
and

$$\mathbf{V} = (\mathbf{a}_1 \cdot \mathbf{V})\mathbf{a}^1 + (\mathbf{a}_2 \cdot \mathbf{V})\mathbf{a}^2 + (\mathbf{a}_3 \cdot \mathbf{V})\mathbf{a}^3, \quad (35)$$

compare with Eq. (36) in *Thompson, Warsi and Mastin* (1985) p. 109. Remember again that the components  $\mathbf{a}^i \cdot \mathbf{V}$  and  $\mathbf{a}_i \cdot \mathbf{V}$  are called contravariant and covariant components of  $\mathbf{V}$  respectively.

**Gravitational acceleration vector in inclined Cartesian coordinate system:** Given the Cartesian coordinate system  $(x, y, z)$  in which the gravitational acceleration is  $\mathbf{g} = g_x \mathbf{i} + g_y \mathbf{j} + g_z \mathbf{k}$  with  $g_x = 0$ ,  $g_y = 0$  and  $g_z = -|\mathbf{g}|$ . Ro-

tate this coordinate system clockwise around  $y$  on angle  $\varphi_y$  as shown in Fig. 1. The new coordinate system  $(\xi, \eta, \zeta)$  is also Cartesian. Compute the Jacobian matrix of the coordinate transformation and the inversed Jacobian matrix of the new system. Compute the components of the gravitational vector in the new system. Solve the same task for rotation around the other axis.



**Fig. 4.** Rotation of Cartesian coordinate system in  $y = \text{const}$  plane

From geometrical consideration the solution the components of the Jacobian matrix are

$$\mathbf{J}_{y=\eta} = \begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix} = \begin{pmatrix} \cos \varphi_y & 0 & \sin \varphi_y \\ 0 & 1 & 0 \\ -\sin \varphi_y & 0 & \cos \varphi_y \end{pmatrix}.$$

Obviously the co-variant vectors are unit vectors. The inversed Jacobian matrix of the system is then

$$a_{y=\eta}^{ij} = \mathbf{J}_{y=\eta}^{-1} = \begin{pmatrix} \cos \varphi_y & 0 & \sin \varphi_y \\ 0 & 1 & 0 \\ -\sin \varphi_y & 0 & \cos \varphi_y \end{pmatrix}^{-1} = \begin{pmatrix} \cos \varphi_y & 0 & -\sin \varphi_y \\ 0 & 1 & 0 \\ \sin \varphi_y & 0 & \cos \varphi_y \end{pmatrix}.$$

The components of the gravitational vector

$$\mathbf{g} = g^1 \mathbf{a}_1 + g^2 \mathbf{a}_2 + g^3 \mathbf{a}_3$$

in the new system are

$$\begin{aligned} g_\xi &= g'^1 = a^{11}g_x + a^{12}g_y + a^{13}g_z = |\mathbf{g}| \sin \varphi_y, \\ g_\eta &= g'^2 = a^{21}g_x + a^{22}g_y + a^{23}g_z = 0, \\ g_\zeta &= g'^3 = a^{31}g_x + a^{32}g_y + a^{33}g_z = -|\mathbf{g}| \cos \varphi_y, \end{aligned}$$

which in this case is easily directly verified by direct geometrical considerations.

For rotation the Cartesian system around the  $x$  axis on angle  $\varphi_x$  the result is:

$$\mathbf{J}_{x=\xi} = \begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \varphi_x & \sin \varphi_x \\ 0 & -\sin \varphi_x & \cos \varphi_x \end{pmatrix},$$

$$a_{x=\xi}^{ij} = \mathbf{J}_{x=\xi}^{-1} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \varphi_x & -\sin \varphi_x \\ 0 & \sin \varphi_x & \cos \varphi_x \end{pmatrix},$$

and therefore

$$\begin{aligned} g_\xi &= 0, \\ g_\eta &= |\mathbf{g}| \sin \varphi_x, \\ g_\zeta &= -|\mathbf{g}| \cos \varphi_x. \end{aligned}$$

For rotation the Cartesian system around the  $z$  axis on angle  $\varphi_z$  the result is:

$$\mathbf{J}_{z=\zeta} = \begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix} = \begin{pmatrix} \cos \varphi_z & \sin \varphi_z & 0 \\ -\sin \varphi_z & \cos \varphi_z & 0 \\ 0 & 0 & 1 \end{pmatrix},$$

$$a_{z=\zeta}^{ij} = \mathbf{J}_{z=\zeta}^{-1} = \begin{pmatrix} \cos \varphi_z & -\sin \varphi_z & 0 \\ \sin \varphi_z & \cos \varphi_z & 0 \\ 0 & 0 & 1 \end{pmatrix},$$

and therefore

$$\begin{aligned} g_\xi &= 0, \\ g_\eta &= 0, \\ g_\zeta &= -|\mathbf{g}|, \end{aligned}$$

which is the expected result because in this case the z-axis remains parallel to the gravitational direction.

**Gravitational acceleration vector in inclined cylindrical coordinate system:** Given the task as described in the previous problem. Consider the case with the rotation around the y-axis. In addition for the transformed coordinate system use instead of Cartesian coordinate system  $(\xi, \eta, \zeta)$  a orthogonal cylindrical one  $(r, \theta, \zeta)$  for which at  $\theta=0$   $r$  coincides with  $\xi$ . Compute the gravitational components along the two new axes. Because  $\zeta$  remains the same the component

$$g_\zeta = -|\mathbf{g}| \cos \varphi_y,$$

along it remain as it is. In the plane  $\zeta = \text{const}$  the component  $|\mathbf{g}| \sin \varphi$  is a constant along  $\xi$  with components

$$g_r = |\mathbf{g}| \sin \varphi_y \cos \theta,$$

$$g_\theta = -|\mathbf{g}| \sin \varphi_y \sin \theta.$$

Observe that this components do not depend on the radius  $r$  and on the axial coordinate  $\zeta$ . They only depend on the two angles  $\varphi_y$  and  $\theta$  which is intuitively expected. The same result can be obtained using the general algorithm. As it will be shown later for transferring the Cartesian coordinate system  $(\xi, \eta, \zeta)$  into cylindrical one  $(r, \theta, \zeta)$  the *metrics* of the coordinate transformation are

$$\mathbf{J}_{\xi=r, \eta \rightarrow \theta, \zeta=\zeta}^{-1} = \begin{pmatrix} \cos \theta & \sin \theta & 0 \\ -\frac{1}{r} \sin \theta & \frac{1}{r} \cos \theta & 0 \\ 0 & 0 & 1 \end{pmatrix}.$$

The first and the third contravariant vectors are unit vectors, the unit vector of the second has components  $(-\sin \theta \quad \cos \theta \quad 0)$ . With this the gravitational vector in the final coordinate system is represented by

$$\begin{aligned} g_r &= g'^1 = a^{11} g_\xi + a^{12} g_\eta + a^{13} g_\zeta = |\mathbf{g}| \sin \varphi_y \cos \theta, \\ g_\theta &= g'^2 = a^{21} g_\xi + a^{22} g_\eta + a^{23} g_\zeta = -|\mathbf{g}| \sin \varphi_y \sin \theta, \\ g_\zeta &= -|\mathbf{g}| \cos \varphi_y. \end{aligned}$$

This is a demonstration that several successive transformations of a vector in new coordinate systems can be obtained by

$$\mathbf{g}_{final} = \mathbf{J}_n^{-1} \left( \dots \left( \mathbf{J}_2^{-1} \left( \mathbf{J}_1^{-1} \mathbf{g}_{initial} \right) \right) \right).$$

**Increment of the position vector:** The infinitesimal increment of the position vector is

$$d\mathbf{r} = \mathbf{a}_1 d\xi + \mathbf{a}_2 d\eta + \mathbf{a}_3 d\zeta. \quad (36)$$

**Increment of the arc length:** The infinitesimal increment of the arc length along a general space curve can be approximated with the magnitude of the infinitesimal increment of the position vector

$$ds = |d\mathbf{r}| = \sqrt{d\mathbf{r} \cdot d\mathbf{r}} \quad (37)$$

or

$$\begin{aligned} |ds|^2 &= d\mathbf{r} \cdot d\mathbf{r} = (\mathbf{a}_1 d\xi + \mathbf{a}_2 d\eta + \mathbf{a}_3 d\zeta) \cdot (\mathbf{a}_1 d\xi + \mathbf{a}_2 d\eta + \mathbf{a}_3 d\zeta) \\ &= \left( \begin{aligned} &(\mathbf{a}_1 \cdot \mathbf{a}_1) d\xi d\xi + (\mathbf{a}_2 \cdot \mathbf{a}_1) d\xi d\eta + (\mathbf{a}_3 \cdot \mathbf{a}_1) d\xi d\zeta \\ &+ (\mathbf{a}_1 \cdot \mathbf{a}_2) d\eta d\xi + (\mathbf{a}_2 \cdot \mathbf{a}_2) d\eta d\eta + (\mathbf{a}_3 \cdot \mathbf{a}_2) d\eta d\zeta \\ &+ (\mathbf{a}_1 \cdot \mathbf{a}_3) d\zeta d\xi + (\mathbf{a}_2 \cdot \mathbf{a}_3) d\zeta d\eta + (\mathbf{a}_3 \cdot \mathbf{a}_3) d\zeta d\zeta \end{aligned} \right). \end{aligned} \quad (38)$$

The tensor built by the dot products is obviously symmetric, and is called the *covariant tensor*.

**Covariant metric tensor:** The *covariant metric tensor* is defined as



$$g_{ij} = \begin{pmatrix} \mathbf{a}_1 \cdot \mathbf{a}_1 & \mathbf{a}_1 \cdot \mathbf{a}_2 & \mathbf{a}_1 \cdot \mathbf{a}_3 \\ \mathbf{a}_2 \cdot \mathbf{a}_1 & \mathbf{a}_2 \cdot \mathbf{a}_2 & \mathbf{a}_2 \cdot \mathbf{a}_3 \\ \mathbf{a}_3 \cdot \mathbf{a}_1 & \mathbf{a}_3 \cdot \mathbf{a}_2 & \mathbf{a}_3 \cdot \mathbf{a}_3 \end{pmatrix}, \quad (39)$$

which is a symmetric tensor.

**Angles between the unit covariant vectors:** The angles between the covariant unit vectors can be expressed in terms of the elements of the covariant metric tensor as follows

$$\theta_{12} = \arccos \left( \frac{\mathbf{a}_1 \cdot \mathbf{a}_2}{|\mathbf{a}_1| |\mathbf{a}_2|} \right) = \arccos \left( \frac{g_{12}}{\sqrt{g_{11} g_{22}}} \right), \quad (40)$$

$$\theta_{13} = \arccos \left( \frac{\mathbf{a}_1 \cdot \mathbf{a}_3}{|\mathbf{a}_1| |\mathbf{a}_3|} \right) = \arccos \left( \frac{g_{13}}{\sqrt{g_{11} g_{33}}} \right), \quad (41)$$

$$\theta_{23} = \arccos \left( \frac{\mathbf{a}_2 \cdot \mathbf{a}_3}{|\mathbf{a}_2| |\mathbf{a}_3|} \right) = \arccos \left( \frac{g_{23}}{\sqrt{g_{22} g_{33}}} \right). \quad (42)$$

It is obvious that only in a specially designed curvilinear coordinate systems can the covariant vectors be mutually perpendicular. Such systems are called *orthogonal*. In this case the transformed region is *rectangular*.

General curvilinear coordinate systems are *not orthogonal* in many applications. In such systems the transformed region where the curvilinear coordinates are independent variables *can be thought of as being rectangular*, and can be treated as such from a coding standpoint by formation of the finite difference equations and in the solution thereof. The problem is thus much simpler in the transformed field, *since the boundaries here are all thought of as rectangular*.

**Orthogonality:** In an orthogonal coordinate system: a) the two types of base vectors are parallel and b) three base vectors on each type are mutually perpendicular. The consequence is that the off diagonal terms of covariant metric tensor are zeros.

**Surface area increment:** Consider a parallelepiped with finite sizes  $\mathbf{a}_1 d\xi$ ,  $\mathbf{a}_2 d\eta$ , and  $\mathbf{a}_3 d\zeta$ . The areas of the surfaces defined with constant curvilinear coordinates are

$$\xi = \text{const}, \quad dS^1 = (\mathbf{a}_2 \times \mathbf{a}_3) d\eta d\zeta, \quad (43)$$

$$\eta = \text{const}, \quad dS^2 = (\mathbf{a}_3 \times \mathbf{a}_1) d\xi d\zeta, \quad (44)$$

$$\zeta = \text{const}, \quad dS^3 = (\mathbf{a}_1 \times \mathbf{a}_2) d\xi d\eta, \quad (45)$$

respectively. Using the vector identity

$$(\mathbf{A} \times \mathbf{B}) \cdot (\mathbf{C} \times \mathbf{D}) = (\mathbf{A} \cdot \mathbf{C})(\mathbf{B} \cdot \mathbf{D}) - (\mathbf{A} \cdot \mathbf{D})(\mathbf{B} \cdot \mathbf{C}) \quad (46)$$

for  $\mathbf{C} = \mathbf{A}$  and  $\mathbf{D} = \mathbf{B}$

$$(\mathbf{A} \times \mathbf{B}) \cdot (\mathbf{A} \times \mathbf{B}) = (\mathbf{A} \cdot \mathbf{A})(\mathbf{B} \cdot \mathbf{B}) - (\mathbf{A} \cdot \mathbf{B})^2, \quad (47)$$

or

$$|\mathbf{A} \times \mathbf{B}| = \sqrt{(\mathbf{A} \cdot \mathbf{A})(\mathbf{B} \cdot \mathbf{B}) - (\mathbf{A} \cdot \mathbf{B})^2} \quad (48)$$

the magnitude of the increment of the surface area is

$$\xi = \text{const},$$

$$dS^1 = |\mathbf{a}_2 \times \mathbf{a}_3| d\eta d\zeta = \sqrt{(\mathbf{a}_2 \cdot \mathbf{a}_2)(\mathbf{a}_3 \cdot \mathbf{a}_3) - (\mathbf{a}_2 \cdot \mathbf{a}_3)^2} d\eta d\zeta = \sqrt{g_{22}g_{33} - g_{23}^2} d\eta d\zeta, \quad (49)$$

$$\eta = \text{const},$$

$$dS^2 = |\mathbf{a}_1 \times \mathbf{a}_3| d\xi d\zeta = \sqrt{(\mathbf{a}_1 \cdot \mathbf{a}_1)(\mathbf{a}_3 \cdot \mathbf{a}_3) - (\mathbf{a}_3 \cdot \mathbf{a}_1)^2} d\xi d\zeta = \sqrt{g_{11}g_{33} - g_{31}^2} d\xi d\zeta, \quad (50)$$

$$\zeta = \text{const},$$

$$dS^3 = |\mathbf{a}_1 \times \mathbf{a}_2| d\xi d\eta = \sqrt{(\mathbf{a}_1 \cdot \mathbf{a}_1)(\mathbf{a}_2 \cdot \mathbf{a}_2) - (\mathbf{a}_1 \cdot \mathbf{a}_2)^2} d\xi d\eta = \sqrt{g_{11}g_{22} - g_{12}^2} d\xi d\eta. \quad (51)$$

**Infinite volume in curvilinear coordinate systems:** An infinitesimal parallelepiped formed by the vectors  $\mathbf{a}_1 d\xi$ ,  $\mathbf{a}_2 d\eta$ , and  $\mathbf{a}_3 d\zeta$  has a infinitesimal volume computed using the box product rule

$$\begin{aligned} dV &= (\mathbf{a}_1 d\xi \times \mathbf{a}_2 d\eta) \cdot \mathbf{a}_3 d\zeta = (\mathbf{a}_1 \times \mathbf{a}_2) \cdot \mathbf{a}_3 d\xi d\eta d\zeta \\ &= \begin{vmatrix} a_{11} & a_{12} & a_{13} \\ a_{21} & a_{22} & a_{23} \\ a_{31} & a_{32} & a_{33} \end{vmatrix} d\xi d\eta d\zeta = \begin{vmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{vmatrix} d\xi d\eta d\zeta, \end{aligned} \quad (52)$$

or using the *Jacobian determinant* or *Jacobian* of the coordinate transformation  $x = f(\xi, \eta, \zeta)$ ,  $y = g(\xi, \eta, \zeta)$ ,  $z = h(\xi, \eta, \zeta)$  we have

$$dV = |\mathbf{J}(\xi, \eta, \zeta)| d\xi d\eta d\zeta. \quad (53)$$

Another notation of the *Jacobian* is

$$\sqrt{g} = |\mathbf{J}(\xi, \eta, \zeta)| = (\mathbf{a}_1 \times \mathbf{a}_2) \cdot \mathbf{a}_3 = (\mathbf{a}_2 \times \mathbf{a}_3) \cdot \mathbf{a}_1 = (\mathbf{a}_3 \times \mathbf{a}_1) \cdot \mathbf{a}_2, \quad (54)$$

so that the volume increment can also be written as

$$dV = \sqrt{g} d\xi d\eta d\zeta. \quad (55)$$

Note the cyclic permutation of the subscripts inside the brackets in Eq. (54). In finite difference form we have

$$\Delta V = \sqrt{g} \Delta\xi \Delta\eta \Delta\zeta. \quad (56)$$

Selecting strict equidistant discretization into the transformed space with  $\Delta\xi = 1$ ,  $\Delta\eta = 1$ ,  $\Delta\zeta = 1$  we have a visualization of the meaning of the *Jacobian*, namely

$$\sqrt{g} = \Delta V. \quad (57)$$

Some authors are using this approach to compute the *Jacobian* from the volume of the computational cell provided the discretization of the computational domain is strictly equidistant with steps of unity in all directions. The real volume of the cell is computed in this case from the Cartesian coordinates of the vertices of the real cell as will be shown later.

***The Jacobian of the coordinate transformation in terms of dot products of the covariant vectors:*** The square of the Jacobian

$$g = |\mathbf{J}(\xi, \eta, \zeta)|^2 = [\mathbf{a}_1 \cdot (\mathbf{a}_2 \times \mathbf{a}_3)]^2 \quad (58)$$

can be rewritten as a function of the dot products of the components of the *Jacobian* determinant as follows. Using the vector identities

$$(\mathbf{A} \times \mathbf{B}) \cdot (\mathbf{C} \times \mathbf{D}) = (\mathbf{A} \cdot \mathbf{C})(\mathbf{B} \cdot \mathbf{D}) - (\mathbf{A} \cdot \mathbf{D})(\mathbf{B} \cdot \mathbf{C}) \quad (59)$$

for  $\mathbf{C} = \mathbf{A}$  and  $\mathbf{D} = \mathbf{B}$  results in

$$(\mathbf{A} \times \mathbf{B}) \cdot (\mathbf{A} \times \mathbf{B}) = (\mathbf{A} \cdot \mathbf{A})(\mathbf{B} \cdot \mathbf{B}) - (\mathbf{A} \cdot \mathbf{B})^2 \quad (60)$$

or

$$(\mathbf{A} \cdot \mathbf{B})^2 = (\mathbf{A} \cdot \mathbf{A})(\mathbf{B} \cdot \mathbf{B}) - (\mathbf{A} \times \mathbf{B}) \cdot (\mathbf{A} \times \mathbf{B}). \quad (61)$$

With this result we have

$$[\mathbf{a}_1 \cdot (\mathbf{a}_2 \times \mathbf{a}_3)]^2 = (\mathbf{a}_1 \cdot \mathbf{a}_1)(\mathbf{a}_2 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3) - [\mathbf{a}_1 \times (\mathbf{a}_2 \times \mathbf{a}_3)] \cdot [\mathbf{a}_1 \times (\mathbf{a}_2 \times \mathbf{a}_3)]. \quad (62)$$

Using the same identity Eq. (59)

$$(\mathbf{a}_2 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3) = (\mathbf{a}_2 \cdot \mathbf{a}_2)(\mathbf{a}_3 \cdot \mathbf{a}_3) - (\mathbf{a}_2 \cdot \mathbf{a}_3)^2 \quad (63)$$

and by the vector identity

$$\mathbf{A} \times (\mathbf{B} \times \mathbf{C}) = (\mathbf{A} \cdot \mathbf{C})\mathbf{B} - (\mathbf{A} \cdot \mathbf{B})\mathbf{C} \quad (64)$$

we have

$$\mathbf{a}_1 \times (\mathbf{a}_2 \times \mathbf{a}_3) = (\mathbf{a}_1 \cdot \mathbf{a}_3)\mathbf{a}_2 - (\mathbf{a}_1 \cdot \mathbf{a}_2)\mathbf{a}_3. \quad (65)$$

Finally we obtain

$$\begin{aligned} & \left[ \mathbf{a}_1 \cdot (\mathbf{a}_2 \times \mathbf{a}_3) \right]^2 = (\mathbf{a}_1 \cdot \mathbf{a}_1) \left[ (\mathbf{a}_2 \cdot \mathbf{a}_2)(\mathbf{a}_3 \cdot \mathbf{a}_3) - (\mathbf{a}_2 \cdot \mathbf{a}_3)^2 \right] - \left| (\mathbf{a}_1 \cdot \mathbf{a}_3)\mathbf{a}_2 - (\mathbf{a}_1 \cdot \mathbf{a}_2)\mathbf{a}_3 \right|^2 \\ & = (\mathbf{a}_1 \cdot \mathbf{a}_1)(\mathbf{a}_2 \cdot \mathbf{a}_2)(\mathbf{a}_3 \cdot \mathbf{a}_3) - (\mathbf{a}_1 \cdot \mathbf{a}_1)(\mathbf{a}_2 \cdot \mathbf{a}_3)^2 - (\mathbf{a}_2 \cdot \mathbf{a}_2)(\mathbf{a}_1 \cdot \mathbf{a}_3)^2 - (\mathbf{a}_3 \cdot \mathbf{a}_3)(\mathbf{a}_1 \cdot \mathbf{a}_2)^2 \\ & \quad + 2(\mathbf{a}_1 \cdot \mathbf{a}_2)(\mathbf{a}_3 \cdot \mathbf{a}_2)(\mathbf{a}_1 \cdot \mathbf{a}_3) \\ & = \begin{vmatrix} \mathbf{a}_1 \cdot \mathbf{a}_1 & \mathbf{a}_1 \cdot \mathbf{a}_2 & \mathbf{a}_1 \cdot \mathbf{a}_3 \\ \mathbf{a}_2 \cdot \mathbf{a}_1 & \mathbf{a}_2 \cdot \mathbf{a}_2 & \mathbf{a}_2 \cdot \mathbf{a}_3 \\ \mathbf{a}_3 \cdot \mathbf{a}_1 & \mathbf{a}_3 \cdot \mathbf{a}_2 & \mathbf{a}_3 \cdot \mathbf{a}_3 \end{vmatrix} = |g_{ij}| \\ & = g_{11}(g_{22}g_{33} - g_{23}^2) - g_{12}(g_{21}g_{33} - g_{23}g_{31}) + g_{13}(g_{21}g_{32} - g_{22}g_{31}), \end{aligned} \quad (66)$$

which is the determinant of the covariant symmetric tensor.

**Relation between the partial derivatives with respect to Cartesian and curvilinear coordinates:** Partial derivatives with respect to *Cartesian* coordinates are related to partial derivatives with respect to the curvilinear coordinates by the chain rule. If  $\varphi(x, y, z)$  is a scalar-valued function of the main variables  $x, y, z$ , through the three intermediate variables  $\xi, \eta, \zeta$ , then

$$\frac{\partial \varphi}{\partial x} = \frac{\partial \varphi}{\partial \xi} \frac{\partial \xi}{\partial x} + \frac{\partial \varphi}{\partial \eta} \frac{\partial \eta}{\partial x} + \frac{\partial \varphi}{\partial \zeta} \frac{\partial \zeta}{\partial x} = \frac{\partial \varphi}{\partial \xi} a^{11} + \frac{\partial \varphi}{\partial \eta} a^{21} + \frac{\partial \varphi}{\partial \zeta} a^{31}, \quad (67)$$

$$\frac{\partial \varphi}{\partial y} = \frac{\partial \varphi}{\partial \xi} \frac{\partial \xi}{\partial y} + \frac{\partial \varphi}{\partial \eta} \frac{\partial \eta}{\partial y} + \frac{\partial \varphi}{\partial \zeta} \frac{\partial \zeta}{\partial y} = \frac{\partial \varphi}{\partial \xi} a^{12} + \frac{\partial \varphi}{\partial \eta} a^{22} + \frac{\partial \varphi}{\partial \zeta} a^{32}, \quad (68)$$

$$\frac{\partial \varphi}{\partial z} = \frac{\partial \varphi}{\partial \xi} \frac{\partial \xi}{\partial z} + \frac{\partial \varphi}{\partial \eta} \frac{\partial \eta}{\partial z} + \frac{\partial \varphi}{\partial \zeta} \frac{\partial \zeta}{\partial z} = \frac{\partial \varphi}{\partial \xi} a^{13} + \frac{\partial \varphi}{\partial \eta} a^{23} + \frac{\partial \varphi}{\partial \zeta} a^{33}, \quad (69)$$

with the second superscript indicating the Cartesian component of the contravariant vectors. If

$$\mathbf{V}(x, y, z) = u(x, y, z)\mathbf{i} + v(x, y, z)\mathbf{j} + w(x, y, z)\mathbf{k}$$

is a vector-valued function of the main variables  $x, y, z$ , through the three intermediate variables  $\xi, \eta, \zeta$ , then

$$\begin{aligned} \frac{\partial \mathbf{V}}{\partial x} &= \left( \frac{\partial u}{\partial \xi} a^{11} + \frac{\partial u}{\partial \eta} a^{21} + \frac{\partial u}{\partial \zeta} a^{31} \right) \mathbf{i} + \left( \frac{\partial v}{\partial \xi} a^{11} + \frac{\partial v}{\partial \eta} a^{21} + \frac{\partial v}{\partial \zeta} a^{31} \right) \mathbf{j} \\ &+ \left( \frac{\partial w}{\partial \xi} a^{11} + \frac{\partial w}{\partial \eta} a^{21} + \frac{\partial w}{\partial \zeta} a^{31} \right) \mathbf{k}, \end{aligned} \quad (70)$$

$$\begin{aligned} \frac{\partial \mathbf{V}}{\partial y} &= \left( \frac{\partial u}{\partial \xi} a^{12} + \frac{\partial u}{\partial \eta} a^{22} + \frac{\partial u}{\partial \zeta} a^{32} \right) \mathbf{i} + \left( \frac{\partial v}{\partial \xi} a^{12} + \frac{\partial v}{\partial \eta} a^{22} + \frac{\partial v}{\partial \zeta} a^{32} \right) \mathbf{j} \\ &+ \left( \frac{\partial w}{\partial \xi} a^{12} + \frac{\partial w}{\partial \eta} a^{22} + \frac{\partial w}{\partial \zeta} a^{32} \right) \mathbf{k}, \end{aligned} \quad (71)$$

$$\begin{aligned} \frac{\partial \mathbf{V}}{\partial z} &= \left( \frac{\partial u}{\partial \xi} a^{13} + \frac{\partial u}{\partial \eta} a^{23} + \frac{\partial u}{\partial \zeta} a^{33} \right) \mathbf{i} + \left( \frac{\partial v}{\partial \xi} a^{13} + \frac{\partial v}{\partial \eta} a^{23} + \frac{\partial v}{\partial \zeta} a^{33} \right) \mathbf{j} \\ &+ \left( \frac{\partial w}{\partial \xi} a^{13} + \frac{\partial w}{\partial \eta} a^{23} + \frac{\partial w}{\partial \zeta} a^{33} \right) \mathbf{k}. \end{aligned} \quad (72)$$

**The divergence theorem (Gauss-Ostrogradskii):** The divergence theorem says that under suitable conditions the outward flux of a vector field across a closed surface (oriented outward) equals the triple integral of the divergence of the field over the region enclosed by the surface. The flux of vector  $\mathbf{F} = M\mathbf{i} + N\mathbf{j} + P\mathbf{k}$  across a closed oriented surface  $S$  in the direction of the surface's outward unit normal field  $\mathbf{n}$  equals the integral of  $\nabla \cdot \mathbf{F}$  over the region  $D$  enclosed by the surface:

$$\iint_S \mathbf{F} \cdot \mathbf{n} d\sigma = \iiint_D \nabla \cdot \mathbf{F} dV. \quad (73)$$

**Volume of space enveloped by a closed surface:** Consider in a space closed surface  $S$  described by the position vector  $\mathbf{r}(x, y, z)$ . The volume inside the surface can be expressed in terms of the area integral over its boundary using the divergence theorem

$$\begin{aligned} \iint_S \mathbf{r} \cdot \mathbf{n} d\sigma &= \iiint_D \nabla \cdot \mathbf{r} dV = \iiint_D \left( \frac{\partial}{\partial x} \mathbf{i} + \frac{\partial}{\partial y} \mathbf{j} + \frac{\partial}{\partial z} \mathbf{k} \right) \cdot (x\mathbf{i} + y\mathbf{j} + z\mathbf{k}) dV \\ &= \iiint_D \left( \frac{\partial x}{\partial x} + \frac{\partial y}{\partial y} + \frac{\partial z}{\partial z} \right) dV = 3 \iiint_D dV = 3D, \end{aligned} \quad (74)$$

or

$$D = \frac{1}{3} \iint_S \mathbf{r} \cdot \mathbf{n} d\sigma. \quad (75)$$

A practical application of this relation is the computation of the volume enclosed in plane elements with area  $S_m$ , outwards unit normal vectors  $\mathbf{n}_m$ , and position vector of the centroids of the faces  $\mathbf{r}_m$ ,

$$D = \frac{1}{3} \sum_m S_m \mathbf{n}_m \cdot \mathbf{r}_m. \quad (76)$$

**The divergence theorem for a curvilinear coordinate system:** Consider a differential element of volume  $D$  bounded by six faces lying on coordinate surfaces, as shown in Fig. A2.3.

*Divergence:* Applying the *Divergence Theorem* we obtain

$$\begin{aligned} & \iiint_D (\nabla \cdot \mathbf{F}) \sqrt{g} \, d\xi d\eta d\zeta \\ &= \iint_{S_1} \mathbf{F} \cdot (\mathbf{a}_2 \times \mathbf{a}_3) d\eta d\zeta - \iint_{S_2} \mathbf{F} \cdot (\mathbf{a}_2 \times \mathbf{a}_3) d\eta d\zeta \\ &+ \iint_{S_3} \mathbf{F} \cdot (\mathbf{a}_3 \times \mathbf{a}_1) d\xi d\zeta - \iint_{S_4} \mathbf{F} \cdot (\mathbf{a}_3 \times \mathbf{a}_1) d\xi d\zeta \\ &+ \iint_{S_5} \mathbf{F} \cdot (\mathbf{a}_1 \times \mathbf{a}_2) d\xi d\eta - \iint_{S_6} \mathbf{F} \cdot (\mathbf{a}_1 \times \mathbf{a}_2) d\xi d\eta. \end{aligned} \quad (77)$$

Dividing by  $D$  and letting  $D$  approach zero we obtain the so called *conservative* expression for the divergence

$$\nabla \cdot \mathbf{F} = \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} [(\mathbf{a}_2 \times \mathbf{a}_3) \cdot \mathbf{F}] + \frac{\partial}{\partial \eta} [(\mathbf{a}_3 \times \mathbf{a}_1) \cdot \mathbf{F}] + \frac{\partial}{\partial \zeta} [(\mathbf{a}_1 \times \mathbf{a}_2) \cdot \mathbf{F}] \right\}. \quad (78)$$

Having in mind that  $\mathbf{a}_2$  and  $\mathbf{a}_3$  are independent of  $\xi$ , and  $\mathbf{a}_1$  and  $\mathbf{a}_3$  are independent of  $\eta$ , and  $\mathbf{a}_1$  and  $\mathbf{a}_2$  are independent of  $\zeta$  results in the so called *first fundamental metric identity*, Peyret (1996),

$$\frac{\partial}{\partial \xi} (\mathbf{a}_2 \times \mathbf{a}_3) + \frac{\partial}{\partial \eta} (\mathbf{a}_3 \times \mathbf{a}_1) + \frac{\partial}{\partial \zeta} (\mathbf{a}_1 \times \mathbf{a}_2) = 0. \quad (79)$$

The so called *non-conservative* expression for the divergence is then

$$\nabla \cdot \mathbf{F} = \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \cdot \frac{\partial \mathbf{F}}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \cdot \frac{\partial \mathbf{F}}{\partial \eta} + (\mathbf{a}_1 \times \mathbf{a}_2) \cdot \frac{\partial \mathbf{F}}{\partial \zeta} \right]. \quad (80)$$

*Curl:* The Divergence Theorem is valid if the dot product is replaced by a cross product. The conservative form of the *curl* is

$$\nabla \times \mathbf{F} = \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} [(\mathbf{a}_2 \times \mathbf{a}_3) \times \mathbf{F}] + \frac{\partial}{\partial \eta} [(\mathbf{a}_3 \times \mathbf{a}_1) \times \mathbf{F}] + \frac{\partial}{\partial \zeta} [(\mathbf{a}_1 \times \mathbf{a}_2) \times \mathbf{F}] \right\}, \quad (81)$$

and the non-conservative form

$$\nabla \times \mathbf{F} = \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \times \frac{\partial \mathbf{F}}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \times \frac{\partial \mathbf{F}}{\partial \eta} + (\mathbf{a}_1 \times \mathbf{a}_2) \times \frac{\partial \mathbf{F}}{\partial \zeta} \right]. \quad (82)$$

*Gradient:* The Divergence Theorem is valid if the tensor or the vector  $\mathbf{F}$  is replaced by a scalar  $\varphi$ . The conservative form of the *gradient vector (gradient)* of the differentiable function  $\varphi$  is

$$\nabla \varphi = \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} [(\mathbf{a}_2 \times \mathbf{a}_3) \varphi] + \frac{\partial}{\partial \eta} [(\mathbf{a}_3 \times \mathbf{a}_1) \varphi] + \frac{\partial}{\partial \zeta} [(\mathbf{a}_1 \times \mathbf{a}_2) \varphi] \right\}, \quad (83)$$

and the non-conservative form

$$\nabla \varphi = \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \frac{\partial \varphi}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \frac{\partial \varphi}{\partial \eta} + (\mathbf{a}_1 \times \mathbf{a}_2) \frac{\partial \varphi}{\partial \zeta} \right]. \quad (84)$$

*Laplacian:* The expression for the *Laplacian* follows from the expression for the divergence and gradient as follows: The conservative form is

$$\begin{aligned} \nabla^2 \varphi &= \nabla \cdot (\nabla \varphi) \\ &= \frac{1}{\sqrt{g}} \frac{\partial}{\partial \xi} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \cdot \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} [(\mathbf{a}_2 \times \mathbf{a}_3) \varphi] + \frac{\partial}{\partial \eta} [(\mathbf{a}_3 \times \mathbf{a}_1) \varphi] \right. \right. \\ &\quad \left. \left. + \frac{\partial}{\partial \zeta} [(\mathbf{a}_1 \times \mathbf{a}_2) \varphi] \right\} \right] \\ &\quad + \frac{1}{\sqrt{g}} \frac{\partial}{\partial \eta} \left[ (\mathbf{a}_1 \times \mathbf{a}_3) \cdot \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} [(\mathbf{a}_2 \times \mathbf{a}_3) \varphi] + \frac{\partial}{\partial \eta} [(\mathbf{a}_3 \times \mathbf{a}_1) \varphi] \right. \right. \\ &\quad \left. \left. + \frac{\partial}{\partial \zeta} [(\mathbf{a}_1 \times \mathbf{a}_2) \varphi] \right\} \right] \\ &\quad + \frac{1}{\sqrt{g}} \frac{\partial}{\partial \zeta} \left[ (\mathbf{a}_1 \times \mathbf{a}_2) \cdot \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} [(\mathbf{a}_2 \times \mathbf{a}_3) \varphi] + \frac{\partial}{\partial \eta} [(\mathbf{a}_3 \times \mathbf{a}_1) \varphi] \right. \right. \\ &\quad \left. \left. + \frac{\partial}{\partial \zeta} [(\mathbf{a}_1 \times \mathbf{a}_2) \varphi] \right\} \right]. \quad (85) \end{aligned}$$

The non-conservative form is

$$\begin{aligned} \nabla^2 \varphi &= \nabla \cdot (\nabla \varphi) \\ &= \frac{1}{\sqrt{g}} \left\{ \begin{aligned} &\frac{\partial}{\partial \xi} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \cdot \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \frac{\partial \varphi}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \frac{\partial \varphi}{\partial \eta} + (\mathbf{a}_1 \times \mathbf{a}_2) \frac{\partial \varphi}{\partial \zeta} \right] \right] \\ &+ \frac{\partial}{\partial \eta} \left[ (\mathbf{a}_1 \times \mathbf{a}_3) \cdot \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \frac{\partial \varphi}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \frac{\partial \varphi}{\partial \xi} + (\mathbf{a}_1 \times \mathbf{a}_2) \frac{\partial \varphi}{\partial \zeta} \right] \right] \\ &+ \frac{\partial}{\partial \zeta} \left[ (\mathbf{a}_1 \times \mathbf{a}_2) \cdot \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \frac{\partial \varphi}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \frac{\partial \varphi}{\partial \xi} + (\mathbf{a}_1 \times \mathbf{a}_2) \frac{\partial \varphi}{\partial \zeta} \right] \right] \end{aligned} \right\}. \end{aligned} \tag{86}$$

The above relations were used for the first time to derive conservative transformed equations in the gas dynamics by *Vivand* and *Vinokur* in 1974.

**Contravariant metric tensor:** It is very interesting to note that in the *Laplacian* the nine scalar products form a symmetric tensor

$$g^{ij} = \begin{pmatrix} \mathbf{a}^1 \cdot \mathbf{a}^1 & \mathbf{a}^1 \cdot \mathbf{a}^2 & \mathbf{a}^1 \cdot \mathbf{a}^3 \\ \mathbf{a}^2 \cdot \mathbf{a}^1 & \mathbf{a}^2 \cdot \mathbf{a}^2 & \mathbf{a}^2 \cdot \mathbf{a}^3 \\ \mathbf{a}^3 \cdot \mathbf{a}^1 & \mathbf{a}^3 \cdot \mathbf{a}^2 & \mathbf{a}^3 \cdot \mathbf{a}^3 \end{pmatrix}, \tag{87}$$

called the *contravariant metric tensor*. Using the identity

$$(\mathbf{A} \times \mathbf{B}) \cdot (\mathbf{C} \times \mathbf{D}) = (\mathbf{A} \cdot \mathbf{C})(\mathbf{B} \cdot \mathbf{D}) - (\mathbf{A} \cdot \mathbf{D})(\mathbf{B} \cdot \mathbf{C}) \tag{88}$$

we compute the six contravariant metric numbers as follows

$$\begin{aligned} \mathbf{a}^1 \cdot \mathbf{a}^1 &= \frac{1}{g} (\mathbf{a}_2 \times \mathbf{a}_3) \cdot (\mathbf{a}_2 \times \mathbf{a}_3) = \frac{1}{g} \left[ (\mathbf{a}_2 \cdot \mathbf{a}_2)(\mathbf{a}_3 \cdot \mathbf{a}_3) - (\mathbf{a}_2 \cdot \mathbf{a}_3)^2 \right] \\ &= \frac{1}{g} (g_{22}g_{33} - g_{23}^2), \end{aligned} \tag{89}$$

$$\begin{aligned} \mathbf{a}^2 \cdot \mathbf{a}^2 &= \frac{1}{g} (\mathbf{a}_3 \times \mathbf{a}_1) \cdot (\mathbf{a}_3 \times \mathbf{a}_1) = \frac{1}{g} \left[ (\mathbf{a}_3 \cdot \mathbf{a}_3)(\mathbf{a}_1 \cdot \mathbf{a}_1) - (\mathbf{a}_3 \cdot \mathbf{a}_1)^2 \right] \\ &= \frac{1}{g} (g_{33}g_{11} - g_{31}^2), \end{aligned} \tag{90}$$

$$\mathbf{a}^3 \cdot \mathbf{a}^3 = \frac{1}{g} (\mathbf{a}_1 \times \mathbf{a}_2) \cdot (\mathbf{a}_1 \times \mathbf{a}_2) = \frac{1}{g} \left[ (\mathbf{a}_1 \cdot \mathbf{a}_1)(\mathbf{a}_2 \cdot \mathbf{a}_2) - (\mathbf{a}_1 \cdot \mathbf{a}_2)^2 \right]$$



$$= \frac{1}{g} (g_{11}g_{22} - g_{12}^2), \quad (91)$$

$$\begin{aligned} \mathbf{a}^1 \cdot \mathbf{a}^2 &= \frac{1}{g} (\mathbf{a}_2 \times \mathbf{a}_3) \cdot (\mathbf{a}_3 \times \mathbf{a}_1) = \frac{1}{g} [(\mathbf{a}_2 \cdot \mathbf{a}_3)(\mathbf{a}_3 \cdot \mathbf{a}_1) - (\mathbf{a}_2 \cdot \mathbf{a}_1)(\mathbf{a}_3 \cdot \mathbf{a}_3)] \\ &= \frac{1}{g} (g_{23}g_{31} - g_{21}g_{33}), \end{aligned} \quad (92)$$

$$\begin{aligned} \mathbf{a}^1 \cdot \mathbf{a}^3 &= \frac{1}{g} (\mathbf{a}_2 \times \mathbf{a}_3) \cdot (\mathbf{a}_1 \times \mathbf{a}_2) = \frac{1}{g} [(\mathbf{a}_2 \cdot \mathbf{a}_1)(\mathbf{a}_3 \cdot \mathbf{a}_2) - (\mathbf{a}_2 \cdot \mathbf{a}_2)(\mathbf{a}_3 \cdot \mathbf{a}_1)] \\ &= \frac{1}{g} (g_{21}g_{32} - g_{22}g_{31}), \end{aligned} \quad (93)$$

$$\begin{aligned} \mathbf{a}^2 \cdot \mathbf{a}^3 &= \frac{1}{g} (\mathbf{a}_3 \times \mathbf{a}_1) \cdot (\mathbf{a}_1 \times \mathbf{a}_2) = \frac{1}{g} [(\mathbf{a}_3 \cdot \mathbf{a}_1)(\mathbf{a}_1 \cdot \mathbf{a}_2) - (\mathbf{a}_3 \cdot \mathbf{a}_2)(\mathbf{a}_1 \cdot \mathbf{a}_1)] \\ &= \frac{1}{g} (g_{31}g_{12} - g_{32}g_{11}). \end{aligned} \quad (94)$$

Computing the determinant of the contravariant metric tensor we realize that

$$\begin{vmatrix} \mathbf{a}^1 \cdot \mathbf{a}^1 & \mathbf{a}^1 \cdot \mathbf{a}^2 & \mathbf{a}^1 \cdot \mathbf{a}^3 \\ \mathbf{a}^2 \cdot \mathbf{a}^1 & \mathbf{a}^2 \cdot \mathbf{a}^2 & \mathbf{a}^2 \cdot \mathbf{a}^3 \\ \mathbf{a}^3 \cdot \mathbf{a}^1 & \mathbf{a}^3 \cdot \mathbf{a}^2 & \mathbf{a}^3 \cdot \mathbf{a}^3 \end{vmatrix} = \frac{1}{g}. \quad (95)$$

**Relation between the contravariant and the covariant base vectors, dual vectors:** The non-conservative form of the *gradient vector* of the differentiable function  $\varphi$

$$\nabla \varphi = \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \frac{\partial \varphi}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \frac{\partial \varphi}{\partial \eta} + (\mathbf{a}_1 \times \mathbf{a}_2) \frac{\partial \varphi}{\partial \zeta} \right] \quad (96)$$

is applied for all of the coordinates  $\xi, \eta, \zeta$ . Since the three coordinates are independent of each other their derivatives with respect to each other are zero. With this in mind we obtain

$$\nabla \xi = \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \frac{\partial \xi}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \frac{\partial \xi}{\partial \eta} + (\mathbf{a}_1 \times \mathbf{a}_2) \frac{\partial \xi}{\partial \zeta} \right] = \frac{1}{\sqrt{g}} (\mathbf{a}_2 \times \mathbf{a}_3), \quad (97)$$

$$\nabla \eta = \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \frac{\partial \eta}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \frac{\partial \eta}{\partial \eta} + (\mathbf{a}_1 \times \mathbf{a}_2) \frac{\partial \eta}{\partial \zeta} \right] = \frac{1}{\sqrt{g}} (\mathbf{a}_3 \times \mathbf{a}_1), \quad (98)$$

$$\nabla \zeta = \frac{1}{\sqrt{g}} \left[ (\mathbf{a}_2 \times \mathbf{a}_3) \frac{\partial \zeta}{\partial \xi} + (\mathbf{a}_3 \times \mathbf{a}_1) \frac{\partial \zeta}{\partial \eta} + (\mathbf{a}_1 \times \mathbf{a}_2) \frac{\partial \zeta}{\partial \zeta} \right] = \frac{1}{\sqrt{g}} (\mathbf{a}_1 \times \mathbf{a}_2), \quad (99)$$

or simply

$$\mathbf{a}^1 = \frac{1}{\sqrt{g}} (\mathbf{a}_2 \times \mathbf{a}_3), \quad (100)$$

$$\mathbf{a}^2 = \frac{1}{\sqrt{g}} (\mathbf{a}_3 \times \mathbf{a}_1), \quad (101)$$

$$\mathbf{a}^3 = \frac{1}{\sqrt{g}} (\mathbf{a}_1 \times \mathbf{a}_2). \quad (102)$$

The contravariant vectors are not unit vectors. This result allows us to simplify the expressions for the gradient, divergence, curl, *Laplacian*, etc. Note that with these results the *first fundamental metric identity*,

$$\frac{\partial}{\partial \eta} (\mathbf{a}_2 \times \mathbf{a}_3) + \frac{\partial}{\partial \xi} (\mathbf{a}_3 \times \mathbf{a}_1) + \frac{\partial}{\partial \zeta} (\mathbf{a}_1 \times \mathbf{a}_2) = 0, \quad (103)$$

can be written as

$$\frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3) = 0. \quad (104)$$

It is very important that the grid generated for numerical integration fulfills the above identity strictly. Otherwise spurious numerical errors are introduced in the computational results.

Note the useful relations between the covariant and the contravariant vectors

$$\mathbf{a}_1 \cdot \mathbf{a}^1 = \frac{1}{\sqrt{g}} \mathbf{a}_1 \cdot (\mathbf{a}_2 \times \mathbf{a}_3) = 1, \quad \mathbf{a}_2 \cdot \mathbf{a}^1 = 0, \quad \mathbf{a}_3 \cdot \mathbf{a}^1 = 0, \quad (105-107)$$

$$\mathbf{a}_1 \cdot \mathbf{a}^2 = 0, \quad \mathbf{a}_2 \cdot \mathbf{a}^2 = \frac{1}{\sqrt{g}} \mathbf{a}_2 \cdot (\mathbf{a}_3 \times \mathbf{a}_1) = 1, \quad \mathbf{a}_3 \cdot \mathbf{a}^2 = 0, \quad (108-110)$$

$$\mathbf{a}_1 \cdot \mathbf{a}^3 = 0, \quad \mathbf{a}_2 \cdot \mathbf{a}^3 = 0, \quad \mathbf{a}_3 \cdot \mathbf{a}^3 = \frac{1}{\sqrt{g}} \mathbf{a}_3 \cdot (\mathbf{a}_1 \times \mathbf{a}_2) = 1. \quad (111-113)$$

These properties distinguish the couple of vectors  $\mathbf{a}_i$  and  $\mathbf{a}^i$ . Such vectors are called *dual*. Using these properties Eqs. (33) and (35) can be easily proved.

**Approximating the contravariant base vectors:** Consider a parallelepiped with finite sizes  $\mathbf{a}_1\Delta\xi$ ,  $\mathbf{a}_2\Delta\eta$ , and  $\mathbf{a}_3\Delta\zeta$ . The areas of the surfaces defined with constant curvilinear coordinates are

$$\xi = \text{const}, \quad \mathbf{S}^1 = (\mathbf{a}_2 \times \mathbf{a}_3) \Delta\eta\Delta\zeta = \sqrt{g} \mathbf{a}^1 \Delta\eta\Delta\zeta, \quad (114)$$

$$\eta = \text{const}, \quad \mathbf{S}^2 = (\mathbf{a}_3 \times \mathbf{a}_1) \Delta\xi\Delta\zeta = \sqrt{g} \mathbf{a}^2 \Delta\xi\Delta\zeta, \quad (115)$$

$$\zeta = \text{const}, \quad \mathbf{S}^3 = (\mathbf{a}_1 \times \mathbf{a}_2) \Delta\xi\Delta\eta = \sqrt{g} \mathbf{a}^3 \Delta\xi\Delta\eta, \quad (116)$$

respectively. Solving with respect to the contravariant vectors gives

$$\mathbf{a}^1 = \frac{\mathbf{S}^1}{\sqrt{g} \Delta\eta\Delta\zeta}, \quad (117)$$

$$\mathbf{a}^2 = \frac{\mathbf{S}^2}{\sqrt{g} \Delta\xi\Delta\zeta}, \quad (118)$$

$$\mathbf{a}^3 = \frac{\mathbf{S}^3}{\sqrt{g} \Delta\xi\Delta\eta}. \quad (119)$$

If one selects  $\Delta\xi = 1$ ,  $\Delta\eta = 1$ , and  $\Delta\zeta = 1$ , the contravariant vectors are then

$$\mathbf{a}^1 = \frac{\mathbf{S}^1}{\sqrt{g}}, \quad (120)$$

$$\mathbf{a}^2 = \frac{\mathbf{S}^2}{\sqrt{g}}, \quad (121)$$

$$\mathbf{a}^3 = \frac{\mathbf{S}^3}{\sqrt{g}}. \quad (122)$$

The contravariant metric tensor is in this case

$$g^{ij} = \mathbf{a}^i \cdot \mathbf{a}^j = \frac{\mathbf{S}^i \cdot \mathbf{S}^j}{g}. \quad (123)$$

In a sense of the numerical construction of the computational grid the above formulas can be used for computation of the contravariant vectors for equidistant grids with steps of unity in all directions in the transformed space.

**Relations between the inverse metrics and the metrics of the coordinate transformation:** Usually in practical problems the inverse metrics are

computed analytically or numerically and then the metrics are computed using the following procedure

$$\mathbf{a}^1 := \frac{\partial \xi}{\partial x} \mathbf{i} + \frac{\partial \xi}{\partial y} \mathbf{j} + \frac{\partial \xi}{\partial z} \mathbf{k} = \frac{1}{\sqrt{g}} (\mathbf{a}_2 \times \mathbf{a}_3) = \frac{1}{\sqrt{g}} \begin{vmatrix} \mathbf{i} & \mathbf{j} & \mathbf{k} \\ \frac{\partial x}{\partial \eta} & \frac{\partial y}{\partial \eta} & \frac{\partial z}{\partial \eta} \\ \frac{\partial x}{\partial \zeta} & \frac{\partial y}{\partial \zeta} & \frac{\partial z}{\partial \zeta} \end{vmatrix}, \quad (124)$$

where the Cartesian components are

$$\frac{\partial \xi}{\partial x} = \frac{1}{\sqrt{g}} \left( \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial z}{\partial \eta} \frac{\partial y}{\partial \zeta} \right), \quad (125)$$

$$\frac{\partial \xi}{\partial y} = -\frac{1}{\sqrt{g}} \left( \frac{\partial x}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial z}{\partial \eta} \frac{\partial x}{\partial \zeta} \right), \quad (126)$$

$$\frac{\partial \xi}{\partial z} = \frac{1}{\sqrt{g}} \left( \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \zeta} - \frac{\partial y}{\partial \eta} \frac{\partial x}{\partial \zeta} \right). \quad (127)$$

Similarly we have for the other vectors

$$\mathbf{a}^2 := \frac{\partial \eta}{\partial x} \mathbf{i} + \frac{\partial \eta}{\partial y} \mathbf{j} + \frac{\partial \eta}{\partial z} \mathbf{k} = \frac{1}{\sqrt{g}} (\mathbf{a}_3 \times \mathbf{a}_1) = \frac{1}{\sqrt{g}} \begin{vmatrix} \mathbf{i} & \mathbf{j} & \mathbf{k} \\ \frac{\partial x}{\partial \zeta} & \frac{\partial y}{\partial \zeta} & \frac{\partial z}{\partial \zeta} \\ \frac{\partial x}{\partial \xi} & \frac{\partial y}{\partial \xi} & \frac{\partial z}{\partial \xi} \end{vmatrix} \quad (128)$$

resulting in

$$\frac{\partial \eta}{\partial x} = \frac{1}{\sqrt{g}} \left( \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \xi} - \frac{\partial z}{\partial \zeta} \frac{\partial y}{\partial \xi} \right), \quad (129)$$

$$\frac{\partial \eta}{\partial y} = -\frac{1}{\sqrt{g}} \left( \frac{\partial x}{\partial \zeta} \frac{\partial z}{\partial \xi} - \frac{\partial z}{\partial \zeta} \frac{\partial x}{\partial \xi} \right), \quad (130)$$

$$\frac{\partial \eta}{\partial z} = \frac{1}{\sqrt{g}} \left( \frac{\partial x}{\partial \zeta} \frac{\partial y}{\partial \xi} - \frac{\partial y}{\partial \zeta} \frac{\partial x}{\partial \xi} \right), \quad (131)$$

and

$$\mathbf{a}^3 := \frac{\partial \zeta}{\partial x} \mathbf{i} + \frac{\partial \zeta}{\partial y} \mathbf{j} + \frac{\partial \zeta}{\partial z} \mathbf{k} = \frac{1}{\sqrt{g}} (\mathbf{a}_1 \times \mathbf{a}_2) = \frac{1}{\sqrt{g}} \begin{vmatrix} \mathbf{i} & \mathbf{j} & \mathbf{k} \\ \frac{\partial x}{\partial \xi} & \frac{\partial y}{\partial \xi} & \frac{\partial z}{\partial \xi} \\ \frac{\partial x}{\partial \eta} & \frac{\partial y}{\partial \eta} & \frac{\partial z}{\partial \eta} \end{vmatrix} \quad (132)$$

resulting in

$$\frac{\partial \zeta}{\partial x} = \frac{1}{\sqrt{g}} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \eta} - \frac{\partial z}{\partial \xi} \frac{\partial y}{\partial \eta} \right), \quad (133)$$

$$\frac{\partial \zeta}{\partial y} = -\frac{1}{\sqrt{g}} \left( \frac{\partial x}{\partial \xi} \frac{\partial z}{\partial \eta} - \frac{\partial z}{\partial \xi} \frac{\partial x}{\partial \eta} \right), \quad (134)$$

$$\frac{\partial \zeta}{\partial z} = \frac{1}{\sqrt{g}} \left( \frac{\partial x}{\partial \xi} \frac{\partial y}{\partial \eta} - \frac{\partial y}{\partial \xi} \frac{\partial x}{\partial \eta} \right). \quad (135)$$

A plausibility proof of this computation is obtained by comparing the partial derivatives with the components of Eq. (8).

### **Restatement of the conservative derivative operations:**

In what follows the *covariant* differential operators in conservative form are given.

*Divergence:*

$$\nabla \cdot \mathbf{F} = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}) \right]. \quad (136)$$

*Curl:*

$$\nabla \times \mathbf{F} = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \times \mathbf{F}) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \times \mathbf{F}) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \times \mathbf{F}) \right]. \quad (137)$$

*Gradient:*

$$\nabla \varphi = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \varphi) \right]. \quad (138)$$

From this expression the conservative expressions for the first derivatives result

$$\frac{\partial \varphi}{\partial x} = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} a^{11} \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} a^{12} \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} a^{13} \varphi) \right], \quad (139)$$

$$\frac{\partial \varphi}{\partial y} = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} a^{21} \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} a^{22} \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} a^{23} \varphi) \right], \quad (140)$$

$$\frac{\partial \varphi}{\partial z} = \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} a^{31} \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} a^{32} \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} a^{33} \varphi) \right]. \quad (141)$$

Using the first fundamental metric identity the above expressions reduce to the simple non-conservative form, which are also immediately obtained by the chain rule. If conservative integration schemes are designed the use of the conservative form is strongly recommended.

*Laplacian:*

$$\begin{aligned} \nabla^2 \varphi &= \nabla \cdot (\nabla \varphi) \\ &= \frac{1}{\sqrt{g}} \left\{ \begin{aligned} &\frac{\partial}{\partial \xi} \left\{ \mathbf{a}^1 \cdot \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \varphi) \right] \right\} \\ &+ \frac{\partial}{\partial \eta} \left\{ \mathbf{a}^2 \cdot \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \varphi) \right] \right\} \\ &+ \frac{\partial}{\partial \zeta} \left\{ \mathbf{a}^3 \cdot \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \varphi) \right] \right\} \end{aligned} \right\}. \end{aligned} \quad (142)$$

*Diffusion Laplacian:* Usually the Laplacian appears in the diffusion terms of the conservation equations in the form

$$\begin{aligned} \nabla \cdot (\lambda \nabla \varphi) &= \frac{1}{\sqrt{g}} \left\{ \begin{aligned} &\frac{\partial}{\partial \xi} \left\{ \mathbf{a}^1 \cdot \lambda \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \varphi) \right] \right\} \\ &+ \frac{\partial}{\partial \eta} \left\{ \mathbf{a}^2 \cdot \lambda \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \varphi) \right] \right\} \\ &+ \frac{\partial}{\partial \zeta} \left\{ \mathbf{a}^3 \cdot \lambda \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \varphi) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \varphi) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \varphi) \right] \right\} \end{aligned} \right\}, \end{aligned}$$

(143)

where  $\lambda$  is a scalar-valued function of the local flow parameters.

**Restatement of the non-conservative derivative operations:**

In what follows the *covariant* differential operators in non-conservative form are given.

$$\text{Divergence: } \nabla \cdot \mathbf{F} = \mathbf{a}^1 \cdot \frac{\partial \mathbf{F}}{\partial \xi} + \mathbf{a}^2 \cdot \frac{\partial \mathbf{F}}{\partial \eta} + \mathbf{a}^3 \cdot \frac{\partial \mathbf{F}}{\partial \zeta}. \quad (144)$$

$$\text{Curl: } \nabla \times \mathbf{F} = \mathbf{a}^1 \times \frac{\partial \mathbf{F}}{\partial \xi} + \mathbf{a}^2 \times \frac{\partial \mathbf{F}}{\partial \eta} + \mathbf{a}^3 \times \frac{\partial \mathbf{F}}{\partial \zeta}. \quad (145)$$

$$\text{Gradient: } \nabla \varphi = \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta}. \quad (146)$$

From this expression the non-conservative expressions for the first derivatives result

$$\frac{\partial \varphi}{\partial x} = a^{11} \frac{\partial \varphi}{\partial \xi} + a^{21} \frac{\partial \varphi}{\partial \eta} + a^{31} \frac{\partial \varphi}{\partial \zeta}, \quad (147)$$

$$\frac{\partial \varphi}{\partial y} = a^{12} \frac{\partial \varphi}{\partial \xi} + a^{22} \frac{\partial \varphi}{\partial \eta} + a^{32} \frac{\partial \varphi}{\partial \zeta}, \quad (148)$$

$$\frac{\partial \varphi}{\partial z} = a^{13} \frac{\partial \varphi}{\partial \xi} + a^{23} \frac{\partial \varphi}{\partial \eta} + a^{33} \frac{\partial \varphi}{\partial \zeta}, \quad (149)$$

with the second superscript indicating the Cartesian components of the contravariant vectors.

*Laplacian:*

$$\begin{aligned}
 \nabla^2 \varphi = \nabla \cdot (\nabla \varphi) &= \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} \left[ \sqrt{g} \mathbf{a}^1 \cdot \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right] \right. \\
 &\quad \left. + \frac{\partial}{\partial \eta} \left[ \sqrt{g} \mathbf{a}^2 \cdot \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right] \right. \\
 &\quad \left. + \frac{\partial}{\partial \zeta} \left[ \sqrt{g} \mathbf{a}^3 \cdot \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right] \right\} \\
 &= \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} \left[ \sqrt{g} \left( g^{11} \frac{\partial \varphi}{\partial \xi} + g^{12} \frac{\partial \varphi}{\partial \eta} + g^{13} \frac{\partial \varphi}{\partial \zeta} \right) \right] \right. \\
 &\quad \left. + \frac{\partial}{\partial \eta} \left[ \sqrt{g} \left( g^{21} \frac{\partial \varphi}{\partial \xi} + g^{22} \frac{\partial \varphi}{\partial \eta} + g^{23} \frac{\partial \varphi}{\partial \zeta} \right) \right] \right. \\
 &\quad \left. + \frac{\partial}{\partial \zeta} \left[ \sqrt{g} \left( g^{31} \frac{\partial \varphi}{\partial \xi} + g^{32} \frac{\partial \varphi}{\partial \eta} + g^{33} \frac{\partial \varphi}{\partial \zeta} \right) \right] \right\}. \tag{150}
 \end{aligned}$$

*Diffusion Laplacian:*

$$\begin{aligned}
 \nabla \cdot (\lambda \nabla \varphi) &= \frac{1}{\sqrt{g}} \left\{ \frac{\partial}{\partial \xi} \left[ \sqrt{g} \mathbf{a}^1 \cdot \lambda \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right] \right. \\
 &\quad \left. + \frac{\partial}{\partial \eta} \left[ \sqrt{g} \mathbf{a}^2 \cdot \lambda \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right] \right. \\
 &\quad \left. + \frac{\partial}{\partial \zeta} \left[ \sqrt{g} \mathbf{a}^3 \cdot \lambda \left( \mathbf{a}^1 \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \frac{\partial \varphi}{\partial \zeta} \right) \right] \right\}
 \end{aligned}$$



$$= \frac{1}{\sqrt{g}} \left\{ \begin{aligned} & \frac{\partial}{\partial \xi} \left[ \sqrt{g} \lambda \left( g^{11} \frac{\partial \varphi}{\partial \xi} + g^{12} \frac{\partial \varphi}{\partial \eta} + g^{13} \frac{\partial \varphi}{\partial \zeta} \right) \right] \\ & + \frac{\partial}{\partial \eta} \left[ \sqrt{g} \lambda \left( g^{21} \frac{\partial \varphi}{\partial \xi} + g^{22} \frac{\partial \varphi}{\partial \eta} + g^{23} \frac{\partial \varphi}{\partial \zeta} \right) \right] \\ & + \frac{\partial}{\partial \zeta} \left[ \sqrt{g} \lambda \left( g^{31} \frac{\partial \varphi}{\partial \xi} + g^{32} \frac{\partial \varphi}{\partial \eta} + g^{33} \frac{\partial \varphi}{\partial \zeta} \right) \right] \end{aligned} \right\}. \quad (151)$$

*Divergence of a tensor:* Consider the dyadic product of the vector  $\mathbf{V}$ ,  $\mathbf{V}\mathbf{V}$ , which is a second order tensor. The divergence of this tensor is then

$$\nabla \cdot (\mathbf{V}\mathbf{V}) = [\nabla \cdot (\mathbf{V}u)]\mathbf{i} + [\nabla \cdot (\mathbf{V}v)]\mathbf{j} + [\nabla \cdot (\mathbf{V}w)]\mathbf{k}. \quad (152)$$

In this case the already derived expressions for divergence of vectors can be used. Another example of the divergence of a tensor is the expression

$$\nabla \cdot (\eta \nabla \mathbf{V})^T = \left[ \nabla \cdot \left( \eta \frac{\partial \mathbf{V}}{\partial x} \right) \right] \mathbf{i} + \left[ \nabla \cdot \left( \eta \frac{\partial \mathbf{V}}{\partial y} \right) \right] \mathbf{j} + \left[ \nabla \cdot \left( \eta \frac{\partial \mathbf{V}}{\partial z} \right) \right] \mathbf{k}. \quad (153)$$

Also in this case the already derived expressions for divergence of vectors can be used.

*Diffusion Laplacian of a tensor:* Usually the diffusion Laplacian of a tensor appears in the diffusion terms of the momentum conservation equations in the form

$$\nabla \cdot (\eta \nabla \mathbf{V})$$

which is equivalent to

$$\nabla \cdot (\eta \nabla \mathbf{V}) = [\nabla \cdot (\eta \nabla u)]\mathbf{i} + [\nabla \cdot (\eta \nabla v)]\mathbf{j} + [\nabla \cdot (\eta \nabla w)]\mathbf{k}. \quad (154)$$

In this case the already derived expressions for each component are used.

***Derivatives along the normal to the curvilinear coordinate surface:*** If  $\varphi(\xi, \eta, \zeta)$  is a scalar-valued function of the main variables  $\xi, \eta, \zeta$ , through the three intermediate variables  $x, y, z$ , then the derivatives along the normal to the curvilinear coordinate surface are

$$\xi = \text{const}, \quad \frac{\partial \varphi}{\partial \xi^2} = \frac{\mathbf{a}^1}{|\mathbf{a}^1|} \cdot \nabla \varphi = \mathbf{e}^1 \cdot \nabla \varphi, \quad (155)$$

$$\eta = \text{const}, \quad \frac{\partial \varphi}{\partial \eta^*} = \frac{\mathbf{a}^2}{|\mathbf{a}^2|} \cdot \nabla \varphi = \mathbf{e}^2 \cdot \nabla \varphi, \quad (156)$$

$$\zeta = \text{const}, \quad \frac{\partial \varphi}{\partial \zeta^*} = \frac{\mathbf{a}^3}{|\mathbf{a}^3|} \cdot \nabla \varphi = \mathbf{e}^3 \cdot \nabla \varphi. \quad (157)$$

**Derivatives along the tangent of the curvilinear coordinate lines:** If  $\varphi(\xi, \eta, \zeta)$  is a scalar-valued function of the main variables  $\xi, \eta, \zeta$ , through the three intermediate variables  $x, y, z$ , then

$$\begin{aligned} \frac{\partial \varphi}{\partial \xi} &= \frac{\partial \varphi}{\partial x} \frac{\partial x}{\partial \xi} + \frac{\partial \varphi}{\partial y} \frac{\partial y}{\partial \xi} + \frac{\partial \varphi}{\partial z} \frac{\partial z}{\partial \xi} = \left( \frac{\partial \varphi}{\partial x} \mathbf{i} + \frac{\partial \varphi}{\partial y} \mathbf{j} + \frac{\partial \varphi}{\partial z} \mathbf{k} \right) \cdot \left( \frac{\partial x}{\partial \xi} \mathbf{i} + \frac{\partial y}{\partial \xi} \mathbf{j} + \frac{\partial z}{\partial \xi} \mathbf{k} \right) \\ &= (\nabla \varphi) \cdot \frac{\partial \mathbf{r}}{\partial \xi} = \mathbf{a}_1 \cdot \nabla \varphi, \end{aligned} \quad (158)$$

$$\begin{aligned} \frac{\partial \varphi}{\partial \eta} &= \frac{\partial \varphi}{\partial x} \frac{\partial x}{\partial \eta} + \frac{\partial \varphi}{\partial y} \frac{\partial y}{\partial \eta} + \frac{\partial \varphi}{\partial z} \frac{\partial z}{\partial \eta} = \left( \frac{\partial \varphi}{\partial x} \mathbf{i} + \frac{\partial \varphi}{\partial y} \mathbf{j} + \frac{\partial \varphi}{\partial z} \mathbf{k} \right) \cdot \left( \frac{\partial x}{\partial \eta} \mathbf{i} + \frac{\partial y}{\partial \eta} \mathbf{j} + \frac{\partial z}{\partial \eta} \mathbf{k} \right) \\ &= (\nabla \varphi) \cdot \frac{\partial \mathbf{r}}{\partial \eta} = \mathbf{a}_2 \cdot \nabla \varphi, \end{aligned} \quad (159)$$

$$\begin{aligned} \frac{\partial \varphi}{\partial \zeta} &= \frac{\partial \varphi}{\partial x} \frac{\partial x}{\partial \zeta} + \frac{\partial \varphi}{\partial y} \frac{\partial y}{\partial \zeta} + \frac{\partial \varphi}{\partial z} \frac{\partial z}{\partial \zeta} = \left( \frac{\partial \varphi}{\partial x} \mathbf{i} + \frac{\partial \varphi}{\partial y} \mathbf{j} + \frac{\partial \varphi}{\partial z} \mathbf{k} \right) \cdot \left( \frac{\partial x}{\partial \zeta} \mathbf{i} + \frac{\partial y}{\partial \zeta} \mathbf{j} + \frac{\partial z}{\partial \zeta} \mathbf{k} \right) \\ &= (\nabla \varphi) \cdot \frac{\partial \mathbf{r}}{\partial \zeta} = \mathbf{a}_3 \cdot \nabla \varphi. \end{aligned} \quad (160)$$

These are the three directional derivatives of  $\varphi(x, y, z)$  along  $\xi, \eta, \zeta$ , respectively. Another quick check is obtained by replacing the non-conservative form of the gradient in the transformed coordinate systems and using the dual properties of the co- and contravariant vectors.

**Derivatives along the normal to the coordinate lines and tangent to the curvilinear coordinate surface:** The vector  $\mathbf{a}^i \times \mathbf{a}_i$  is normal to the coordinate line on which  $\xi^i$  varies and is also tangent to the coordinate surface on which  $\xi^i$  is constant. Using the relations between the covariant and the contravariant vectors and the vector identity

$$(\mathbf{B} \times \mathbf{C}) \times \mathbf{A} = -[(\mathbf{A} \cdot \mathbf{C})\mathbf{B} - (\mathbf{A} \cdot \mathbf{B})\mathbf{C}] \quad (161)$$

we have

$$\begin{aligned}\mathbf{a}^1 \times \mathbf{a}_1 &= \frac{1}{\sqrt{g}} (\mathbf{a}_2 \times \mathbf{a}_3) \times \mathbf{a}_1 = -\frac{1}{\sqrt{g}} [(\mathbf{a}_1 \cdot \mathbf{a}_3) \mathbf{a}_2 - (\mathbf{a}_1 \cdot \mathbf{a}_2) \mathbf{a}_3] \\ &= -\frac{1}{\sqrt{g}} (g_{13} \mathbf{a}_2 - g_{12} \mathbf{a}_3),\end{aligned}\quad (162)$$

$$\begin{aligned}\mathbf{a}^2 \times \mathbf{a}_2 &= \frac{1}{\sqrt{g}} (\mathbf{a}_3 \times \mathbf{a}_1) \times \mathbf{a}_2 = -\frac{1}{\sqrt{g}} [(\mathbf{a}_2 \cdot \mathbf{a}_1) \mathbf{a}_3 - (\mathbf{a}_2 \cdot \mathbf{a}_3) \mathbf{a}_1] \\ &= -\frac{1}{\sqrt{g}} (g_{21} \mathbf{a}_3 - g_{23} \mathbf{a}_1),\end{aligned}\quad (163)$$

$$\begin{aligned}\mathbf{a}^3 \times \mathbf{a}_3 &= \frac{1}{\sqrt{g}} (\mathbf{a}_1 \times \mathbf{a}_2) \times \mathbf{a}_3 = -\frac{1}{\sqrt{g}} [(\mathbf{a}_3 \cdot \mathbf{a}_2) \mathbf{a}_1 - (\mathbf{a}_3 \cdot \mathbf{a}_1) \mathbf{a}_2] \\ &= -\frac{1}{\sqrt{g}} (g_{32} \mathbf{a}_1 - g_{31} \mathbf{a}_2).\end{aligned}\quad (164)$$

The magnitude is then

$$\begin{aligned}|\mathbf{a}^1 \times \mathbf{a}_1|^2 &= \frac{1}{g} (g_{13}^2 g_{22} - 2g_{12} g_{13} g_{32} + g_{12}^2 g_{33}) \\ &= \frac{1}{g} [g_{13} (g_{13} g_{22} - g_{12} g_{32}) - g_{12} (g_{13} g_{32} - g_{12} g_{33})],\end{aligned}\quad (165)$$

$$\begin{aligned}|\mathbf{a}^2 \times \mathbf{a}_2|^2 &= \frac{1}{g} (g_{21}^2 g_{33} - 2g_{21} g_{23} g_{31} + g_{23}^2 g_{11}) \\ &= \frac{1}{g} [g_{21} (g_{21} g_{33} - g_{23} g_{31}) - g_{23} (g_{21} g_{31} - g_{23} g_{11})],\end{aligned}\quad (166)$$

$$\begin{aligned}|\mathbf{a}^3 \times \mathbf{a}_3|^2 &= \frac{1}{g} (g_{32}^2 g_{11} - 2g_{31} g_{21} g_{32} + g_{31}^2 g_{22}) \\ &= \frac{1}{g} [g_{32} (g_{32} g_{11} - g_{31} g_{21}) - g_{31} (g_{21} g_{32} - g_{31} g_{22})].\end{aligned}\quad (167)$$

Comparing with

$$g = g_{11} (g_{22} g_{33} - g_{23}^2) - g_{12} (g_{21} g_{33} - g_{23} g_{31}) + g_{13} (g_{21} g_{32} - g_{22} g_{31}) \quad (168)$$

results in computationally cheaper expressions

$$|\mathbf{a}^1 \times \mathbf{a}_1|^2 = \frac{1}{g} g_{11} (g_{22} g_{33} - g_{23}^2) - 1, \quad (169)$$

$$|\mathbf{a}^2 \times \mathbf{a}_2|^2 = \frac{1}{g} g_{22} (g_{11} g_{33} - g_{31}^2) - 1, \quad (170)$$

$$|\mathbf{a}^3 \times \mathbf{a}_3|^2 = \frac{1}{g} g_{33} (g_{11} g_{22} - g_{21}^2) - 1. \quad (171)$$

With this result the derivatives normal to the coordinate line on which  $\xi^i$  varies and are also tangent to the coordinate surface on which  $\xi^i$  is constant are

$$\frac{\mathbf{a}^1 \times \mathbf{a}_1}{|\mathbf{a}^1 \times \mathbf{a}_1|} \cdot \nabla \varphi = -\frac{1}{\sqrt{g_{11} (g_{22} g_{33} - g_{23}^2) - g}} (g_{13} \mathbf{a}_2 - g_{12} \mathbf{a}_3) \cdot \nabla \varphi, \quad (172)$$

$$\frac{\mathbf{a}^2 \times \mathbf{a}_2}{|\mathbf{a}^2 \times \mathbf{a}_2|} \cdot \nabla \varphi = -\frac{1}{\sqrt{g_{22} (g_{11} g_{33} - g_{31}^2) - g}} (g_{21} \mathbf{a}_3 - g_{23} \mathbf{a}_1) \cdot \nabla \varphi, \quad (173)$$

$$\frac{\mathbf{a}^3 \times \mathbf{a}_3}{|\mathbf{a}^3 \times \mathbf{a}_3|} \cdot \nabla \varphi = -\frac{1}{\sqrt{g_{33} (g_{11} g_{22} - g_{21}^2) - g}} (g_{32} \mathbf{a}_1 - g_{31} \mathbf{a}_2) \cdot \nabla \varphi. \quad (174)$$

Replacing the non-conservative form of the gradient in the transformed coordinate systems and using the dual properties of the co- and contravariant vectors results in

$$\frac{\mathbf{a}^1 \times \mathbf{a}_1}{|\mathbf{a}^1 \times \mathbf{a}_1|} \cdot \nabla \varphi = -\frac{g_{13} \frac{\partial \varphi}{\partial \eta} - g_{12} \frac{\partial \varphi}{\partial \zeta}}{\sqrt{g_{11} (g_{22} g_{33} - g_{23}^2) - g}}, \quad (175)$$

$$\frac{\mathbf{a}^2 \times \mathbf{a}_2}{|\mathbf{a}^2 \times \mathbf{a}_2|} \cdot \nabla \varphi = -\frac{g_{23} \frac{\partial \varphi}{\partial \xi} - g_{21} \frac{\partial \varphi}{\partial \zeta}}{\sqrt{g_{22} (g_{11} g_{33} - g_{31}^2) - g}}, \quad (176)$$

$$\frac{\mathbf{a}^3 \times \mathbf{a}_3}{|\mathbf{a}^3 \times \mathbf{a}_3|} \cdot \nabla \varphi = -\frac{g_{32} \frac{\partial \varphi}{\partial \xi} - g_{31} \frac{\partial \varphi}{\partial \eta}}{\sqrt{g_{33} (g_{11} g_{22} - g_{21}^2) - g}}. \quad (177)$$

**Moving coordinate system:** Consider a coordinate transformation defined by

$$x = f(\xi, \eta, \zeta), \quad (178)$$

$$y = g(\xi, \eta, \zeta), \quad (179)$$

$$z = h(\xi, \eta, \zeta), \quad (180)$$

which moves with time in space. The velocity

$$\mathbf{V}_{cs} = \left( \frac{\partial x}{\partial \tau} \right)_{\eta, \xi, \zeta} \mathbf{i} + \left( \frac{\partial y}{\partial \tau} \right)_{\eta, \xi, \zeta} \mathbf{j} + \left( \frac{\partial z}{\partial \tau} \right)_{\eta, \xi, \zeta} \mathbf{k}, \quad (181)$$

is the *grid point velocity*. Note an interesting property of the time derivative of the covariant unit vectors

$$\frac{\partial \mathbf{a}_1}{\partial \tau} = \frac{\partial}{\partial \tau} \left( \frac{\partial \mathbf{r}}{\partial \xi} \right) = \frac{\partial}{\partial \xi} \left( \frac{\partial \mathbf{r}}{\partial \tau} \right) = \frac{\partial \mathbf{V}_{cs}}{\partial \xi}, \quad (182)$$

$$\frac{\partial \mathbf{a}_2}{\partial \tau} = \frac{\partial}{\partial \tau} \left( \frac{\partial \mathbf{r}}{\partial \eta} \right) = \frac{\partial}{\partial \eta} \left( \frac{\partial \mathbf{r}}{\partial \tau} \right) = \frac{\partial \mathbf{V}_{cs}}{\partial \eta}, \quad (183)$$

$$\frac{\partial \mathbf{a}_3}{\partial \tau} = \frac{\partial}{\partial \tau} \left( \frac{\partial \mathbf{r}}{\partial \zeta} \right) = \frac{\partial}{\partial \zeta} \left( \frac{\partial \mathbf{r}}{\partial \tau} \right) = \frac{\partial \mathbf{V}_{cs}}{\partial \zeta}. \quad (184)$$

If  $f = f(\tau, x, y, z)$  is a function of time and space using the chain rule we obtain

$$\begin{aligned} \frac{df}{d\tau} &= \left( \frac{\partial f}{\partial \tau} \right)_{x,y,z} \\ &+ \left[ \left( \frac{\partial f}{\partial x} \right)_{t,y,z} \mathbf{i} + \left( \frac{\partial f}{\partial y} \right)_{t,x,z} \mathbf{j} + \left( \frac{\partial f}{\partial z} \right)_{t,x,y} \mathbf{k} \right] \cdot \left[ \left( \frac{dx}{d\tau} \right)_{\xi,\eta,\zeta} \mathbf{i} + \left( \frac{dy}{d\tau} \right)_{\xi,\eta,\zeta} \mathbf{j} + \left( \frac{dz}{d\tau} \right)_{\xi,\eta,\zeta} \mathbf{k} \right]. \end{aligned} \quad (185)$$

Comparing the second part of the right hand side of the above equation with the definition of the gradient and of the grid point velocity we see that

$$\left( \frac{df}{d\tau} \right)_{\xi,\eta,\zeta} = \left( \frac{\partial f}{\partial \tau} \right)_{x,y,z} + (\nabla f) \cdot \mathbf{V}_{cs}, \quad (186)$$

the absolute change with the time of a function  $f$  is the velocity with the respect to the stationary coordinate system  $xyz$  plus a component resulting from the movement of the coordinate system  $\xi, \eta, \zeta$ . Therefore

$$\left( \frac{\partial f}{\partial \tau} \right)_{x,y,z} = \left( \frac{\partial f}{\partial \tau} \right)_{\xi,\eta,\zeta} - \mathbf{V}_{cs} \cdot \nabla f. \quad (187)$$

Here the time derivative  $(df/dt)_{\xi,\eta,\zeta}$  is understood to be at a fixed point in the transformed region. In a moving coordinate system the orientation of the unit covariant vectors will change. That is the elementary parallelepiped volume built by these vectors, the Jacobian, changes also. The derivative of the Jacobian of the coordinate transformation with time can be visualized by dotting the change of the normal component to the covariant vector with the corresponding elementary coordinate surfaces formed by the other two vectors

$$\left( \frac{d\sqrt{g}}{d\tau} \right)_{\xi, \eta, \zeta} = \frac{d\mathbf{a}_1}{d\tau} \cdot (\mathbf{a}_2 \times \mathbf{a}_3) + \frac{d\mathbf{a}_2}{d\tau} \cdot (\mathbf{a}_3 \times \mathbf{a}_1) + \frac{d\mathbf{a}_3}{d\tau} \cdot (\mathbf{a}_1 \times \mathbf{a}_2) \quad (188)$$

or

$$\frac{d\sqrt{g}}{d\tau} = \sqrt{g} \left( \frac{d\mathbf{a}_1}{d\tau} \cdot \mathbf{a}^1 + \frac{d\mathbf{a}_2}{d\tau} \cdot \mathbf{a}^2 + \frac{d\mathbf{a}_3}{d\tau} \cdot \mathbf{a}^3 \right), \quad (189)$$

see *Thompson, Warsi and Mastin* (1985) p.130. Having in mind that

$$\frac{\partial \mathbf{a}_1}{\partial \tau} = \frac{\partial \mathbf{V}_{cs}}{\partial \xi}, \quad (190)$$

$$\frac{\partial \mathbf{a}_2}{\partial \tau} = \frac{\partial \mathbf{V}_{cs}}{\partial \eta}, \quad (191)$$

$$\frac{\partial \mathbf{a}_3}{\partial \tau} = \frac{\partial \mathbf{V}_{cs}}{\partial \zeta}, \quad (192)$$

results in

$$\frac{d\sqrt{g}}{d\tau} = \sqrt{g} \left( \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \xi} + \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \eta} + \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \zeta} \right). \quad (193)$$

Note some very interesting properties of the following expression

$$\begin{aligned} & \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 f \cdot \mathbf{V}_{cs}) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 f \cdot \mathbf{V}_{cs}) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 f \cdot \mathbf{V}_{cs}) \\ &= f \sqrt{g} \left( \mathbf{a}^1 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \xi} + \mathbf{a}^2 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \eta} + \mathbf{a}^3 \cdot \frac{\partial \mathbf{V}_{cs}}{\partial \zeta} \right) \\ &+ \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 f) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 f) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 f) \right] \cdot \mathbf{V}_{cs} \\ &= f \frac{d\sqrt{g}}{d\tau} + \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 f) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 f) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 f) \right] \cdot \mathbf{V}_{cs} \end{aligned} \quad (194)$$

**Transformed fluid properties conservation equation:** Consider  $f = f(\tau, x, y, z)$  being a function of time and space which is controlled by the following conservation equation

$$\left( \frac{\partial f}{\partial \tau} \right)_{x,y,z} + \nabla \cdot (f\mathbf{V}) - \nabla \cdot (\lambda \nabla f) = \mu. \quad (195)$$

The convection-diffusion problem described by this equation is called *isotropic*. Using the expression for the time derivative, the conservative form of the divergence of a vector and gradient of a scalar in the transformed space, we obtain

$$\begin{aligned}
& \frac{\partial}{\partial \tau}(f\sqrt{g}) \\
& + \frac{\partial}{\partial \xi} \left( f\sqrt{g} \mathbf{a}^1 \cdot (\mathbf{V} - \mathbf{V}_{cs}) - \mathbf{a}^1 \cdot \lambda \left[ \frac{\partial}{\partial \xi}(\sqrt{g} \mathbf{a}^1 f) + \frac{\partial}{\partial \eta}(\sqrt{g} \mathbf{a}^2 f) + \frac{\partial}{\partial \zeta}(\sqrt{g} \mathbf{a}^3 f) \right] \right) \\
& + \frac{\partial}{\partial \eta} \left( f\sqrt{g} \mathbf{a}^2 \cdot (\mathbf{V} - \mathbf{V}_{cs}) - \mathbf{a}^2 \cdot \lambda \left[ \frac{\partial}{\partial \xi}(\sqrt{g} \mathbf{a}^1 f) + \frac{\partial}{\partial \eta}(\sqrt{g} \mathbf{a}^2 f) + \frac{\partial}{\partial \zeta}(\sqrt{g} \mathbf{a}^3 f) \right] \right) \\
& + \frac{\partial}{\partial \zeta} \left( f\sqrt{g} \mathbf{a}^3 \cdot (\mathbf{V} - \mathbf{V}_{cs}) - \mathbf{a}^3 \cdot \lambda \left[ \frac{\partial}{\partial \xi}(\sqrt{g} \mathbf{a}^1 f) + \frac{\partial}{\partial \eta}(\sqrt{g} \mathbf{a}^2 f) + \frac{\partial}{\partial \zeta}(\sqrt{g} \mathbf{a}^3 f) \right] \right) \\
& = \sqrt{g} \mu. \tag{196}
\end{aligned}$$

Using the non-conservative form of the Diffusion *Laplacian* results in a simpler form which is still conservative with respect to the spatial derivatives in the transformed coordinate system

$$\begin{aligned}
& \frac{\partial}{\partial \tau}(f\sqrt{g}) \\
& + \frac{\partial}{\partial \xi} \left\{ \sqrt{g} \left[ f\mathbf{a}^1 \cdot (\mathbf{V} - \mathbf{V}_{cs}) - \lambda \left( \mathbf{a}^1 \cdot \mathbf{a}^1 \frac{\partial f}{\partial \xi} + \mathbf{a}^1 \cdot \mathbf{a}^2 \frac{\partial f}{\partial \eta} + \mathbf{a}^1 \cdot \mathbf{a}^3 \frac{\partial f}{\partial \zeta} \right) \right] \right\} \\
& + \frac{\partial}{\partial \eta} \left\{ \sqrt{g} \left[ f\mathbf{a}^2 \cdot (\mathbf{V} - \mathbf{V}_{cs}) - \lambda \left( \mathbf{a}^2 \cdot \mathbf{a}^1 \frac{\partial f}{\partial \xi} + \mathbf{a}^2 \cdot \mathbf{a}^2 \frac{\partial f}{\partial \eta} + \mathbf{a}^2 \cdot \mathbf{a}^3 \frac{\partial f}{\partial \zeta} \right) \right] \right\} \\
& + \frac{\partial}{\partial \zeta} \left\{ \sqrt{g} \left[ f\mathbf{a}^3 \cdot (\mathbf{V} - \mathbf{V}_{cs}) - \lambda \left( \mathbf{a}^3 \cdot \mathbf{a}^1 \frac{\partial f}{\partial \xi} + \mathbf{a}^3 \cdot \mathbf{a}^2 \frac{\partial f}{\partial \eta} + \mathbf{a}^3 \cdot \mathbf{a}^3 \frac{\partial f}{\partial \zeta} \right) \right] \right\} = \sqrt{g} \mu. \tag{197}
\end{aligned}$$

The equation can be rewritten in more compact form using the relative contravariant velocity components in the transformed space defined as follows

$$\bar{\mathbf{V}} = \begin{pmatrix} \mathbf{a}^1 \cdot (\mathbf{V} - \mathbf{V}_{cs}) \\ \mathbf{a}^2 \cdot (\mathbf{V} - \mathbf{V}_{cs}) \\ \mathbf{a}^3 \cdot (\mathbf{V} - \mathbf{V}_{cs}) \end{pmatrix} = \begin{pmatrix} \bar{V}^1 \\ \bar{V}^2 \\ \bar{V}^3 \end{pmatrix} \quad (198)$$

and the terms of the inverse metrics matrices

$$\begin{aligned} & \frac{\partial}{\partial \tau} (f \sqrt{g}) \\ & + \frac{\partial}{\partial \xi} \left\{ \sqrt{g} \left[ f \bar{V}^1 - \lambda \left( g^{11} \frac{\partial f}{\partial \xi} + g^{12} \frac{\partial f}{\partial \eta} + g^{13} \frac{\partial f}{\partial \zeta} \right) \right] \right\} \\ & + \frac{\partial}{\partial \eta} \left\{ \sqrt{g} \left[ f \bar{V}^2 - \lambda \left( g^{21} \frac{\partial f}{\partial \xi} + g^{22} \frac{\partial f}{\partial \eta} + g^{23} \frac{\partial f}{\partial \zeta} \right) \right] \right\} \\ & + \frac{\partial}{\partial \zeta} \left\{ \sqrt{g} \left[ f \bar{V}^3 - \lambda \left( g^{31} \frac{\partial f}{\partial \xi} + g^{32} \frac{\partial f}{\partial \eta} + g^{33} \frac{\partial f}{\partial \zeta} \right) \right] \right\} = \sqrt{g} \mu. \end{aligned} \quad (199)$$

We see a remarkable property of this equation. In the transformed coordinate system the property  $f \sqrt{g}$  having convection-diffusion flux  $i$ -components

$$\hat{i}^i = \sqrt{g} \left[ f \mathbf{a}^i \cdot (\mathbf{V} - \mathbf{V}_{cs}) - \lambda \left( g^{i1} \frac{\partial f}{\partial \xi} + g^{i2} \frac{\partial f}{\partial \eta} + g^{i3} \frac{\partial f}{\partial \zeta} \right) \right] \quad (200)$$

is conserved. The convective component can be visualized as

$$\sqrt{g} f \mathbf{a}^i \cdot (\mathbf{V} - \mathbf{V}_{cs}) = \sqrt{g} f |\mathbf{a}^i| \mathbf{e}^i \cdot (\mathbf{V} - \mathbf{V}_{cs}) = \sqrt{g} f |\mathbf{a}^i| V^{n,i}. \quad (201)$$

Here

$$V^{n,i} = \mathbf{e}^i \cdot (\mathbf{V} - \mathbf{V}_{cs}) \quad (202)$$

is the relative velocity component normal to the surface defined by  $\xi^i = const$ . Remember Eqs. (117-119) giving

$$\sqrt{g} \mathbf{a}^i = \frac{\mathbf{S}^i}{d\xi^j d\xi^k}, \quad (203)$$

or



$$\sqrt{g} |\mathbf{a}^i| = \frac{|\mathbf{S}^i|}{d\xi^j d\xi^k}. \quad (204)$$

For equidistant discretization in the computational space we have

$$\sqrt{g} |\mathbf{a}^i| = |\mathbf{S}^i| = S^i. \quad (205)$$

Therefore the convective flux can be presented in several ways

$$\sqrt{g} f \mathbf{a}^i \cdot (\mathbf{V} - \mathbf{V}_{cs}) = \sqrt{g} f |\mathbf{a}^i| \mathbf{e}^i \cdot (\mathbf{V} - \mathbf{V}_{cs}) = S^i V^{n,i} f = \sqrt{g} \bar{V}^i f = \mathbf{S}^i \cdot (\mathbf{V} - \mathbf{V}_{cs}) f. \quad (206)$$

With this the conservation equation can be rewritten in terms of the normal velocity components

$$\begin{aligned} & \frac{\partial}{\partial \tau} (f \sqrt{g}) \\ & + \frac{\partial}{\partial \xi} \left[ S^1 V^{n,1} f - \lambda \sqrt{g} \left( g^{11} \frac{\partial f}{\partial \xi} + g^{12} \frac{\partial f}{\partial \eta} + g^{13} \frac{\partial f}{\partial \zeta} \right) \right] \\ & + \frac{\partial}{\partial \eta} \left[ S^2 V^{n,2} f - \lambda \sqrt{g} \left( g^{21} \frac{\partial f}{\partial \xi} + g^{22} \frac{\partial f}{\partial \eta} + g^{23} \frac{\partial f}{\partial \zeta} \right) \right] \\ & + \frac{\partial}{\partial \zeta} \left[ S^3 V^{n,3} f - \lambda \sqrt{g} \left( g^{31} \frac{\partial f}{\partial \xi} + g^{32} \frac{\partial f}{\partial \eta} + g^{33} \frac{\partial f}{\partial \zeta} \right) \right] = \sqrt{g} \mu. \end{aligned} \quad (207)$$

The convection-diffusion problem in the transformed space is called *anisotropic*. Thus, the isotropic convection-diffusion problem in the physical space turns out to be anisotropic in the transformed space.

Note that it is computationally more effective to store  $g^{ij} \sqrt{g}$  instead of  $g^{ij}$  if  $\sqrt{g}$  does not change with the time.

For an orthogonal transformed system the equation simplifies much more

$$\begin{aligned} & \frac{\partial}{\partial \tau} (f \sqrt{g}) + \frac{\partial}{\partial \xi} \left[ \sqrt{g} \left( f \bar{V}^1 - \lambda g^{11} \frac{\partial f}{\partial \xi} \right) \right] + \frac{\partial}{\partial \eta} \left[ \sqrt{g} \left( f \bar{V}^2 - \lambda g^{22} \frac{\partial f}{\partial \eta} \right) \right] \\ & + \frac{\partial}{\partial \zeta} \left[ \sqrt{g} \left( f \bar{V}^3 - \lambda g^{33} \frac{\partial f}{\partial \zeta} \right) \right] = \sqrt{g} \mu. \end{aligned} \quad (208)$$

The convection-diffusion problem in an orthogonal transformed system is called *orthotropic*.

**Time dependent grid metric identity (second fundamental metric identity):**

Consider a conservation equation in the transformed space for  $f = \text{const}$ , and  $\mathbf{V} = \text{const}$  without diffusion and mass sources

$$\begin{aligned} & \frac{\partial \sqrt{g}}{\partial \tau} - \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \cdot \mathbf{V}_{cs}) - \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \cdot \mathbf{V}_{cs}) - \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \cdot \mathbf{V}_{cs}) \\ &= -\mathbf{V} \cdot \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1) - \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2) - \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3) \right]. \end{aligned} \quad (209)$$

Using the first fundamental metric identity *Thompson et al. (1985)*, p.159, obtained

$$\frac{\partial \sqrt{g}}{\partial \tau} - \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \cdot \mathbf{V}_{cs}) - \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \cdot \mathbf{V}_{cs}) - \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \cdot \mathbf{V}_{cs}) = 0. \quad (210)$$

This is an important identity for the numerical analysis called sometimes the *second fundamental metric identity*. As pointed out by *Thompson et al. (1985)* this identity should be used to numerically determine updated values of the *Jacobian*,  $\sqrt{g}$ , instead of updating it directly from the new values of the Cartesian coordinates. In the later case spurious source terms will appear *Thompson et al. (1985)*.

**Transformed total time derivative of a scalar:** The total derivative of a scalar  $\varphi$  traveling in the Cartesian coordinate system with velocity  $\mathbf{V}$  is defined with

$$\frac{d\varphi}{d\tau} = \frac{\partial \varphi}{\partial \tau} + \mathbf{V} \cdot \nabla \varphi. \quad (211)$$

Using the already derived relations we obtain

$$\begin{aligned} \frac{d\varphi}{d\tau} &= \left( \frac{\partial \varphi}{\partial \tau} \right)_{x,y,z} + \mathbf{V}_l \cdot \nabla \varphi \\ &= \left( \frac{\partial \varphi}{\partial \tau} \right)_{\xi,\eta,\zeta} + \mathbf{a}^1 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \varphi}{\partial \xi} + \mathbf{a}^2 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \varphi}{\partial \eta} + \mathbf{a}^3 \cdot (\mathbf{V}_l - \mathbf{V}_{cs}) \frac{\partial \varphi}{\partial \zeta} \\ &= \left( \frac{\partial \varphi}{\partial \tau} \right)_{\xi,\eta,\zeta} + \bar{V}_l^1 \frac{\partial \varphi}{\partial \xi} + \bar{V}_l^2 \frac{\partial \varphi}{\partial \eta} + \bar{V}_l^3 \frac{\partial \varphi}{\partial \zeta}. \end{aligned} \quad (212)$$

### Examples of some frequently used curvilinear coordinate transformations:

Next we apply the derivations from this Appendix to two well-known coordinate systems – the cylindrical and the spherical. If the derived formulas are applied correctly the well-known expressions will finally arise for the differential operators. It is recommended to go through these examples in order to understand better the transformation theory.

*Cylindrical coordinates:* The transformation is defined analytically by

$$x = \xi \cos \eta, \quad \xi = +\sqrt{x^2 + y^2}, \quad (213,214)$$

$$y = \xi \sin \eta, \quad \eta = \arctan(y/x), \quad (215,216)$$

$$z = \zeta, \quad \zeta = z. \quad (217,218)$$

Here  $\xi$  is the radial coordinate,  $\eta$  is the azimuthal coordinate and the axial coordinate  $\zeta$  is equivalent with  $z$ . An equidistant transformation in the computational space is defined by

$$x(\xi_i, \eta_j) = \left[ r_{\min} + (r_{\max} - r_{\min}) \frac{\xi_i}{i_{\max}} \right] \cos \left( \frac{\theta_{\max} - \theta_{\min}}{j_{\max}} \eta_j \right),$$

$$y(\xi_i, \eta_j) = \left[ r_{\min} + (r_{\max} - r_{\min}) \frac{\xi_i}{i_{\max}} \right] \sin \left( \frac{\theta_{\max} - \theta_{\min}}{j_{\max}} \eta_j \right),$$

$$z(\zeta_k) = \frac{z_{\max} - z_{\min}}{k_{\max}} \zeta_k,$$

where

$$\xi_i = 0, 1, 2, \dots, i_{\max}, \quad \eta_j = 0, 1, 2, \dots, j_{\max}, \quad \zeta_k = 0, 1, 2, \dots, k_{\max}.$$

The *inverse metrics* of the coordinate transformation are then

$$\begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix} = \begin{pmatrix} \cos \eta & -\xi \sin \eta & 0 \\ \sin \eta & \xi \cos \eta & 0 \\ 0 & 0 & 1 \end{pmatrix}. \quad (219)$$

The *Jacobian* is then

$$\begin{aligned} \sqrt{g} &= \frac{\partial x}{\partial \xi} \left( \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} \right) - \frac{\partial x}{\partial \eta} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \xi} \right) + \frac{\partial x}{\partial \zeta} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \eta} - \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \xi} \right) \\ &= \xi. \end{aligned} \quad (220)$$

The *metrics* of the coordinate transformation are

$$\begin{aligned} \begin{pmatrix} \frac{\partial \xi}{\partial x} & \frac{\partial \xi}{\partial y} & \frac{\partial \xi}{\partial z} \\ \frac{\partial \eta}{\partial x} & \frac{\partial \eta}{\partial y} & \frac{\partial \eta}{\partial z} \\ \frac{\partial \zeta}{\partial x} & \frac{\partial \zeta}{\partial y} & \frac{\partial \zeta}{\partial z} \end{pmatrix} &= \frac{1}{\sqrt{g}} \begin{pmatrix} \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} & \frac{\partial x}{\partial \zeta} \frac{\partial z}{\partial \eta} - \frac{\partial x}{\partial \eta} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \zeta} - \frac{\partial x}{\partial \zeta} \frac{\partial y}{\partial \eta} \\ \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} & \frac{\partial x}{\partial \zeta} \frac{\partial z}{\partial \eta} - \frac{\partial x}{\partial \eta} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \zeta} - \frac{\partial x}{\partial \zeta} \frac{\partial y}{\partial \eta} \\ \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} & \frac{\partial x}{\partial \zeta} \frac{\partial z}{\partial \eta} - \frac{\partial x}{\partial \eta} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \zeta} - \frac{\partial x}{\partial \zeta} \frac{\partial y}{\partial \eta} \end{pmatrix} \\ &= \begin{pmatrix} \cos \eta & \sin \eta & 0 \\ -\frac{1}{\xi} \sin \eta & \frac{1}{\xi} \cos \eta & 0 \\ 0 & 0 & 1 \end{pmatrix}. \end{aligned} \quad (221)$$

The components of the three *covariant* base vectors of the curvilinear coordinate system are the columns of the matrix of the inverse metrics. They are given below together with the corresponding covariant unit vectors.

$$\begin{aligned} \mathbf{a}_1 &= \frac{\partial x}{\partial \xi} \mathbf{i} + \frac{\partial y}{\partial \xi} \mathbf{j} + \frac{\partial z}{\partial \xi} \mathbf{k} = \cos \eta \mathbf{i} + \sin \eta \mathbf{j} & |\mathbf{a}_1| &= 1 & \mathbf{e}_1 &= \cos \eta \mathbf{i} + \sin \eta \mathbf{j}, \\ \mathbf{a}_2 &= \frac{\partial x}{\partial \eta} \mathbf{i} + \frac{\partial y}{\partial \eta} \mathbf{j} + \frac{\partial z}{\partial \eta} \mathbf{k} = -\xi \sin \eta \mathbf{i} + \xi \cos \eta \mathbf{j} & |\mathbf{a}_2| &= \xi & \mathbf{e}_2 &= -\sin \eta \mathbf{i} + \cos \eta \mathbf{j} \\ \mathbf{a}_3 &= \frac{\partial x}{\partial \zeta} \mathbf{i} + \frac{\partial y}{\partial \zeta} \mathbf{j} + \frac{\partial z}{\partial \zeta} \mathbf{k} = \mathbf{k} & |\mathbf{a}_3| &= 1 & \mathbf{e}_3 &= \mathbf{k} \end{aligned} \quad (222-230)$$

The components of the three *contravariant* base vectors of the curvilinear coordinate system are the rows of the matrix of the metrics. They are given below together with the corresponding contravariant unit vectors.

$$\begin{aligned} \mathbf{a}^1 &= \frac{\partial \xi}{\partial x} \mathbf{i} + \frac{\partial \xi}{\partial y} \mathbf{j} + \frac{\partial \xi}{\partial z} \mathbf{k} = \cos \eta \mathbf{i} + \sin \eta \mathbf{j} & |\mathbf{a}^1| &= 1 & \mathbf{e}^1 &= \cos \eta \mathbf{i} + \sin \eta \mathbf{j} \\ \mathbf{a}^2 &= \frac{\partial \eta}{\partial x} \mathbf{i} + \frac{\partial \eta}{\partial y} \mathbf{j} + \frac{\partial \eta}{\partial z} \mathbf{k} = -\frac{1}{\xi} \sin \eta \mathbf{i} + \frac{1}{\xi} \cos \eta \mathbf{j} & |\mathbf{a}^2| &= \frac{1}{\xi} & \mathbf{e}^2 &= -\sin \eta \mathbf{i} + \cos \eta \mathbf{j} \\ \mathbf{a}^3 &= \frac{\partial \zeta}{\partial x} \mathbf{i} + \frac{\partial \zeta}{\partial y} \mathbf{j} + \frac{\partial \zeta}{\partial z} \mathbf{k} = \mathbf{k} & |\mathbf{a}^3| &= 1 & \mathbf{e}^3 &= \mathbf{k} \end{aligned}$$

(231-239)

We see that a) the two types of base vectors are parallel and that b) three base vectors on each type are mutually perpendicular which means that the cylindrical coordinate system is orthogonal.

The gradient of the vector is then

$$\begin{aligned}\nabla \cdot \mathbf{F} &= \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} \mathbf{a}^1 \cdot \mathbf{F}) + \frac{\partial}{\partial \eta} (\sqrt{g} \mathbf{a}^2 \cdot \mathbf{F}) + \frac{\partial}{\partial \zeta} (\sqrt{g} \mathbf{a}^3 \cdot \mathbf{F}) \right], \\ &= \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} |\mathbf{a}^1| \mathbf{e}^1 \cdot \mathbf{F}) + \frac{\partial}{\partial \eta} (\sqrt{g} |\mathbf{a}^2| \mathbf{e}^2 \cdot \mathbf{F}) + \frac{\partial}{\partial \zeta} (\sqrt{g} |\mathbf{a}^3| \mathbf{e}^3 \cdot \mathbf{F}) \right], \\ &= \frac{1}{\xi} \frac{\partial}{\partial \xi} (\xi F_\xi) + \frac{1}{\xi} \frac{\partial}{\partial \eta} (F_\eta) + \frac{\partial}{\partial \zeta} (F_\zeta),\end{aligned}\quad (240)$$

where  $F_\xi = \mathbf{e}^1 \cdot \mathbf{F}$ ,  $F_\eta = \mathbf{e}^2 \cdot \mathbf{F}$ ,  $F_\zeta = \mathbf{e}^3 \cdot \mathbf{F}$  are the components of  $\mathbf{F}$  normal to the coordinate surfaces defined by  $\xi = const$ ,  $\eta = const$ ,  $\zeta = const$ , respectively. This is the expected form for cylindrical coordinates – compare for instance *Bird, Stewart and Lightfoot* (1960), p. 739B.

The covariant metric tensor is

$$g_{ij} = \begin{pmatrix} \mathbf{a}_1 \cdot \mathbf{a}_1 & \mathbf{a}_1 \cdot \mathbf{a}_2 & \mathbf{a}_1 \cdot \mathbf{a}_3 \\ \mathbf{a}_2 \cdot \mathbf{a}_1 & \mathbf{a}_2 \cdot \mathbf{a}_2 & \mathbf{a}_2 \cdot \mathbf{a}_3 \\ \mathbf{a}_3 \cdot \mathbf{a}_1 & \mathbf{a}_3 \cdot \mathbf{a}_2 & \mathbf{a}_3 \cdot \mathbf{a}_3 \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \xi^2 & 0 \\ 0 & 0 & 1 \end{pmatrix}, \quad (241)$$

which is a symmetric tensor. The angles between the covariant unit vectors are

$$\theta_{12} = \arccos \left( \frac{g_{12}}{\sqrt{g_{11}g_{22}}} \right) = \frac{\pi}{2}, \quad (242)$$

$$\theta_{13} = \arccos \left( \frac{g_{13}}{\sqrt{g_{11}g_{33}}} \right) = \frac{\pi}{2}, \quad (243)$$

$$\theta_{23} = \arccos \left( \frac{g_{23}}{\sqrt{g_{22}g_{33}}} \right) = \frac{\pi}{2}. \quad (244)$$

The magnitude of the increment of the surface area is

$$\xi = const, \quad dS^1 = \sqrt{g_{22}g_{33} - g_{23}^2} d\eta d\zeta = \xi d\eta d\zeta, \quad (245)$$

$$\eta = const, \quad dS^2 = \sqrt{g_{11}g_{33} - g_{13}^2} d\xi d\zeta = d\xi d\zeta, \quad (246)$$

$$\zeta = const, \quad dS^3 = \sqrt{g_{11}g_{22} - g_{12}^2} d\xi d\eta = \xi d\xi d\eta. \quad (247)$$

The volume increment is then

$$dV = \sqrt{g} d\xi d\eta d\zeta = \xi d\xi d\eta d\zeta. \quad (248)$$

The elements of the contravariant metric tensor are

$$g^{ij} = \begin{pmatrix} \mathbf{a}^1 \cdot \mathbf{a}^1 & \mathbf{a}^1 \cdot \mathbf{a}^2 & \mathbf{a}^1 \cdot \mathbf{a}^3 \\ \mathbf{a}^2 \cdot \mathbf{a}^1 & \mathbf{a}^2 \cdot \mathbf{a}^2 & \mathbf{a}^2 \cdot \mathbf{a}^3 \\ \mathbf{a}^3 \cdot \mathbf{a}^1 & \mathbf{a}^3 \cdot \mathbf{a}^2 & \mathbf{a}^3 \cdot \mathbf{a}^3 \end{pmatrix}$$

$$= \frac{1}{g} \begin{pmatrix} g_{22}g_{33} - g_{23}^2 & g_{21}g_{33} - g_{23}g_{31} & g_{21}g_{32} - g_{22}g_{31} \\ g_{21}g_{33} - g_{23}g_{31} & g_{11}g_{33} - g_{13}^2 & g_{11}g_{32} - g_{12}g_{31} \\ g_{21}g_{32} - g_{22}g_{31} & g_{11}g_{32} - g_{12}g_{31} & g_{11}g_{22} - g_{12}^2 \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \frac{1}{\xi^2} & 0 \\ 0 & 0 & 1 \end{pmatrix}. \quad (249)$$

With this result we can compute the Diffusion Laplacian of a scalar for the orthogonal system

$$\begin{aligned} & \nabla \cdot (\lambda \nabla \varphi) \\ &= \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} \left( \sqrt{g} \lambda g^{11} \frac{\partial \varphi}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( \sqrt{g} \lambda g^{22} \frac{\partial \varphi}{\partial \eta} \right) + \frac{\partial}{\partial \zeta} \left( \sqrt{g} \lambda g^{33} \frac{\partial \varphi}{\partial \zeta} \right) \right] \\ &= \frac{1}{\xi} \frac{\partial}{\partial \xi} \left( \xi \lambda \frac{\partial \varphi}{\partial \xi} \right) + \frac{1}{\xi} \frac{\partial}{\partial \eta} \left( \lambda \frac{1}{\xi} \frac{\partial \varphi}{\partial \eta} \right) + \frac{\partial}{\partial \zeta} \left( \lambda \frac{\partial \varphi}{\partial \zeta} \right). \end{aligned} \quad (250)$$

This is the well-known form in cylindrical coordinates – compare for instance with *Bird, Stewart and Lightfoot* (1960), p. 739B

Spherical coordinates:

$$x = \xi \sin \eta \cos \zeta, \quad \xi = +\sqrt{x^2 + y^2 + z^2}, \quad (251,252)$$

$$y = \xi \sin \eta \sin \zeta, \quad x = \arctan \frac{\sqrt{x^2 + y^2}}{z}, \quad (253,254)$$

$$z = \xi \cos \eta, \quad \zeta = \arctan(y/x). \quad (255,256)$$

Here  $\xi$  is the radial coordinate,  $\eta$  is the azimuthal coordinate and  $\zeta$  is the polar coordinate. An equidistant transformation in the computational space is defined by

$$x(\xi_i, \eta_j) = \left[ r_{\min} + (r_{\max} - r_{\min}) \frac{\xi_i}{i_{\max}} \right] \sin \left( \frac{\theta_{\max} - \theta_{\min}}{j_{\max}} \eta_j \right) \cos \left( \frac{\varphi_{\max} - \varphi_{\min}}{k_{\max}} \zeta_j \right), \quad (257)$$

$$y(\xi_i, \eta_j) = \left[ r_{\min} + (r_{\max} - r_{\min}) \frac{\xi_i}{i_{\max}} \right] \sin \left( \frac{\theta_{\max} - \theta_{\min}}{j_{\max}} \eta_j \right) \sin \left( \frac{\varphi_{\max} - \varphi_{\min}}{k_{\max}} \zeta_j \right), \quad (258)$$

$$z(\xi_i, \eta_j) = \left[ r_{\min} + (r_{\max} - r_{\min}) \frac{\xi_i}{i_{\max}} \right] \cos \left( \frac{\theta_{\max} - \theta_{\min}}{j_{\max}} \eta_j \right), \quad (259)$$

where

$$\xi_i = 0, 1, 2, \dots, i_{\max}, \quad \eta_j = 0, 1, 2, \dots, j_{\max}, \quad \zeta_k = 0, 1, 2, \dots, k_{\max}.$$

The *inverse metrics* of the coordinate transformation are then

$$\begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix} = \begin{pmatrix} \sin \eta \cos \zeta & \xi \cos \eta \cos \zeta & -\xi \sin \eta \sin \zeta \\ \sin \eta \sin \zeta & \xi \cos \eta \sin \zeta & \xi \sin \eta \cos \zeta \\ \cos \eta & -\xi \sin \eta & 0 \end{pmatrix}. \quad (260)$$

The *Jacobian* is then

$$\begin{aligned} \sqrt{g} &= \frac{\partial x}{\partial \xi} \left( \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} \right) - \frac{\partial x}{\partial \eta} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \xi} \right) + \frac{\partial x}{\partial \zeta} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \eta} - \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \xi} \right) \\ &= \xi^2 \sin \eta. \end{aligned} \quad (261)$$

The *metrics* of the coordinate transformation are

$$\begin{pmatrix} \frac{\partial \xi}{\partial x} & \frac{\partial \xi}{\partial y} & \frac{\partial \xi}{\partial z} \\ \frac{\partial \eta}{\partial x} & \frac{\partial \eta}{\partial y} & \frac{\partial \eta}{\partial z} \\ \frac{\partial \zeta}{\partial x} & \frac{\partial \zeta}{\partial y} & \frac{\partial \zeta}{\partial z} \end{pmatrix} = \frac{1}{\sqrt{g}} \begin{pmatrix} \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} & \frac{\partial x}{\partial \zeta} \frac{\partial z}{\partial \eta} - \frac{\partial x}{\partial \eta} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \zeta} - \frac{\partial x}{\partial \zeta} \frac{\partial y}{\partial \eta} \\ \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \xi} - \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \xi} \frac{\partial z}{\partial \zeta} - \frac{\partial x}{\partial \zeta} \frac{\partial z}{\partial \xi} & \frac{\partial x}{\partial \zeta} \frac{\partial y}{\partial \xi} - \frac{\partial x}{\partial \xi} \frac{\partial y}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \eta} - \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \xi} & \frac{\partial x}{\partial \eta} \frac{\partial z}{\partial \xi} - \frac{\partial x}{\partial \xi} \frac{\partial z}{\partial \eta} & \frac{\partial x}{\partial \xi} \frac{\partial y}{\partial \eta} - \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \xi} \end{pmatrix}$$

$$= \begin{pmatrix} \sin \eta \cos \zeta & \sin \eta \sin \zeta & \cos \eta \\ \frac{1}{\xi} \cos \zeta \cos \eta & \frac{1}{\xi} \sin \zeta \cos \eta & -\frac{1}{\xi} \sin \eta \\ -\frac{1}{\xi} \frac{\sin \zeta}{\sin \eta} & \frac{1}{\xi} \frac{\cos \zeta}{\sin \eta} & 0 \end{pmatrix}. \quad (262)$$

The components of the three *covariant* base vectors of the curvilinear coordinate system are the columns of the matrix of the inverse metrics. They are given below together with the corresponding covariant unit vectors.

$$\begin{aligned} \mathbf{a}_1 &= \frac{\partial x}{\partial \xi} \mathbf{i} + \frac{\partial y}{\partial \xi} \mathbf{j} + \frac{\partial z}{\partial \xi} \mathbf{k} = \sin \eta \cos \zeta \mathbf{i} + \sin \eta \sin \zeta \mathbf{j} + \cos \eta \mathbf{k}, & |\mathbf{a}_1| &= 1, \\ \mathbf{e}_1 &= \sin \eta \cos \zeta \mathbf{i} + \sin \eta \sin \zeta \mathbf{j} + \cos \eta \mathbf{k}, \end{aligned} \quad (263-265)$$

$$\begin{aligned} \mathbf{a}_2 &= \frac{\partial x}{\partial \eta} \mathbf{i} + \frac{\partial y}{\partial \eta} \mathbf{j} + \frac{\partial z}{\partial \eta} \mathbf{k} = \xi \cos \eta \cos \zeta \mathbf{i} + \xi \cos \eta \sin \zeta \mathbf{j} - \xi \sin \eta \mathbf{k}, & |\mathbf{a}_2| &= \xi, \\ \mathbf{e}_2 &= \cos \eta \cos \zeta \mathbf{i} + \cos \eta \sin \zeta \mathbf{j} - \sin \eta \mathbf{k}, \end{aligned} \quad (266-268)$$

$$\begin{aligned} \mathbf{a}_3 &= \frac{\partial x}{\partial \zeta} \mathbf{i} + \frac{\partial y}{\partial \zeta} \mathbf{j} + \frac{\partial z}{\partial \zeta} \mathbf{k} = -\xi \sin \eta \sin \zeta \mathbf{i} + \xi \sin \eta \cos \zeta \mathbf{j}, & |\mathbf{a}_3| &= \xi \sin \eta, \\ \mathbf{e}_3 &= -\sin \zeta \mathbf{i} + \cos \zeta \mathbf{j}. \end{aligned} \quad (269-271)$$

The components of the three *contravariant* base vectors of the curvilinear coordinate system are the rows of the matrix of the metrics. They are given below together with the corresponding contravariant unit vectors.

$$\begin{aligned} \mathbf{a}^1 &= \frac{\partial \xi}{\partial x} \mathbf{i} + \frac{\partial \xi}{\partial y} \mathbf{j} + \frac{\partial \xi}{\partial z} \mathbf{k} = \sin \eta \cos \zeta \mathbf{i} + \sin \eta \sin \zeta \mathbf{j} + \cos \eta \mathbf{k}, & |\mathbf{a}^1| &= 1, \\ \mathbf{e}^1 &= \sin \eta \cos \zeta \mathbf{i} + \sin \eta \sin \zeta \mathbf{j} + \cos \eta \mathbf{k}, \end{aligned} \quad (272-274)$$

$$\begin{aligned} \mathbf{a}^2 &= \frac{\partial \eta}{\partial x} \mathbf{i} + \frac{\partial \eta}{\partial y} \mathbf{j} + \frac{\partial \eta}{\partial z} \mathbf{k} = \frac{1}{\xi} \cos \zeta \cos \eta \mathbf{i} + \frac{1}{\xi} \sin \zeta \cos \eta \mathbf{j} - \frac{1}{\xi} \sin \eta \mathbf{k}, & |\mathbf{a}^2| &= \frac{1}{\xi}, \\ \mathbf{e}^2 &= \cos \zeta \cos \eta \mathbf{i} + \sin \zeta \cos \eta \mathbf{j} - \sin \eta \mathbf{k}, \end{aligned} \quad (275-277)$$

$$\begin{aligned} \mathbf{a}^3 &= \frac{\partial \zeta}{\partial x} \mathbf{i} + \frac{\partial \zeta}{\partial y} \mathbf{j} + \frac{\partial \zeta}{\partial z} \mathbf{k} = -\frac{1}{\xi} \frac{\sin \zeta}{\sin \eta} \mathbf{i} + \frac{1}{\xi} \frac{\cos \zeta}{\sin \eta} \mathbf{j}, & |\mathbf{a}^3| &= \frac{1}{\xi \sin \eta}, \\ \mathbf{e}^3 &= -\sin \zeta \mathbf{i} + \cos \zeta \mathbf{j}. \end{aligned} \quad (278-280)$$

We see that a) the two types of base vectors are parallel and that b) three base vectors of each type are mutually perpendicular which means that the cylindrical coordinate system is *orthogonal*.



$$\begin{aligned}\nabla \cdot \mathbf{F} &= \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} (\sqrt{g} |\mathbf{a}^1| \mathbf{e}^1 \cdot \mathbf{F}) + \frac{\partial}{\partial \eta} (\sqrt{g} |\mathbf{a}^2| \mathbf{e}^2 \cdot \mathbf{F}) + \frac{\partial}{\partial \zeta} (\sqrt{g} |\mathbf{a}^3| \mathbf{e}^3 \cdot \mathbf{F}) \right] \\ &= \frac{1}{\xi^2 \sin \eta} \left[ \frac{\partial}{\partial \xi} (\xi^2 \sin \eta F_\xi) + \frac{\partial}{\partial \eta} (\xi \sin \eta F_\eta) + \frac{\partial}{\partial \zeta} (\xi F_\zeta) \right].\end{aligned}\quad (281)$$

This is the expected form for spherical coordinates – compare for instance with *Bird, Stewart and Lightfoot (1960), p. 739C*.

The *covariant metric tensor* is

$$g_{ij} = \begin{pmatrix} \mathbf{a}_1 \cdot \mathbf{a}_1 & \mathbf{a}_1 \cdot \mathbf{a}_2 & \mathbf{a}_1 \cdot \mathbf{a}_3 \\ \mathbf{a}_2 \cdot \mathbf{a}_1 & \mathbf{a}_2 \cdot \mathbf{a}_2 & \mathbf{a}_2 \cdot \mathbf{a}_3 \\ \mathbf{a}_3 \cdot \mathbf{a}_1 & \mathbf{a}_3 \cdot \mathbf{a}_2 & \mathbf{a}_3 \cdot \mathbf{a}_3 \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \xi^2 & 0 \\ 0 & 0 & \xi^2 \sin^2 \eta \end{pmatrix}, \quad (282)$$

which is a symmetric tensor. The angles between the covariant unit vectors are

$$\theta_{12} = \arccos \left( \frac{g_{12}}{\sqrt{g_{11}g_{22}}} \right) = \frac{\pi}{2}, \quad (283)$$

$$\theta_{13} = \arccos \left( \frac{g_{13}}{\sqrt{g_{11}g_{33}}} \right) = \frac{\pi}{2}, \quad (284)$$

$$\theta_{23} = \arccos \left( \frac{g_{23}}{\sqrt{g_{22}g_{33}}} \right) = \frac{\pi}{2}. \quad (285)$$

The magnitude of the increment of the surface area is

$$\xi = \text{const}, \quad dS^1 = \sqrt{g_{22}g_{33} - g_{23}^2} d\eta d\zeta = \xi^2 \sin \eta d\eta d\zeta, \quad (286)$$

$$\eta = \text{const}, \quad dS^2 = \sqrt{g_{11}g_{33} - g_{13}^2} d\xi d\zeta = \xi \sin \eta d\xi d\zeta, \quad (287)$$

$$\zeta = \text{const}, \quad dS^3 = \sqrt{g_{11}g_{22} - g_{12}^2} d\xi d\eta = \xi d\xi d\eta. \quad (288)$$

The volume increment is then

$$dV = \sqrt{g} d\xi d\eta d\zeta = \xi^2 \sin \eta d\xi d\eta d\zeta. \quad (289)$$

The elements of the contravariant metric tensor are

$$g^{ij} = \begin{pmatrix} \mathbf{a}^1 \cdot \mathbf{a}^1 & \mathbf{a}^1 \cdot \mathbf{a}^2 & \mathbf{a}^1 \cdot \mathbf{a}^3 \\ \mathbf{a}^2 \cdot \mathbf{a}^1 & \mathbf{a}^2 \cdot \mathbf{a}^2 & \mathbf{a}^2 \cdot \mathbf{a}^3 \\ \mathbf{a}^3 \cdot \mathbf{a}^1 & \mathbf{a}^3 \cdot \mathbf{a}^2 & \mathbf{a}^3 \cdot \mathbf{a}^3 \end{pmatrix}$$

$$= \frac{1}{g} \begin{pmatrix} g_{22}g_{33} - g_{23}^2 & g_{21}g_{33} - g_{23}g_{31} & g_{21}g_{32} - g_{22}g_{31} \\ g_{21}g_{33} - g_{23}g_{31} & g_{11}g_{33} - g_{13}^2 & g_{11}g_{32} - g_{12}g_{31} \\ g_{21}g_{32} - g_{22}g_{31} & g_{11}g_{32} - g_{12}g_{31} & g_{11}g_{22} - g_{12}^2 \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \frac{1}{\xi^2} & 0 \\ 0 & 0 & \frac{1}{\xi^2 \sin^2 \eta} \end{pmatrix}. \tag{290}$$

With this result we can compute the Diffusion Laplacian of a scalar for the orthogonal system

$$\begin{aligned} & \nabla \cdot (\lambda \nabla \varphi) \\ &= \frac{1}{\sqrt{g}} \left[ \frac{\partial}{\partial \xi} \left( \sqrt{g} \lambda g^{11} \frac{\partial \varphi}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( \sqrt{g} \lambda g^{22} \frac{\partial \varphi}{\partial \eta} \right) + \frac{\partial}{\partial \zeta} \left( \sqrt{g} \lambda g^{33} \frac{\partial \varphi}{\partial \zeta} \right) \right] \\ &= \frac{1}{\xi^2} \frac{\partial}{\partial \xi} \left( \xi^2 \lambda \frac{\partial \varphi}{\partial \xi} \right) + \frac{1}{\xi^2 \sin \eta} \frac{\partial}{\partial \eta} \left( \sin \eta \lambda \frac{\partial \varphi}{\partial \eta} \right) + \frac{1}{\xi^2 \sin^2 \eta} \frac{\partial}{\partial \zeta} \left( \lambda \frac{\partial \varphi}{\partial \zeta} \right). \end{aligned} \tag{291}$$

This is the well-known form in spherical coordinates – compare for instance with *Bird, Stewart and Lightfoot (1960), p. 739C*.

**Elliptic grid generation systems:** The most popular grid generating system makes use of the *Laplace* equation. The solution of the *Laplace* equations (harmonic functions) obeys minimum and maximum values only at the boundary (maximum principle) and is smooth (derivatives of all orders exists). One dimensional stretching along the three axes correspondingly, delivers also a solution which has the similar properties as the solution of the *Laplace* equation. The later is proven to be equivalent to solving the *Poison*-like equations, *Thompson, Warsi and Mastin (1985)*,

$$\xi_{xx} + \xi_{yy} + \xi_{zz} = P(\xi, \eta, \zeta), \tag{292}$$

$$\eta_{xx} + \eta_{yy} + \eta_{zz} = Q(\xi, \eta, \zeta), \tag{293}$$

$$\zeta_{xx} + \zeta_{yy} + \zeta_{zz} = R(\xi, \eta, \zeta), \tag{294}$$

with boundary conditions prescribed at the surface containing the domain (*Dirichlet* boundary value problem). The control functions *P*, *Q* and *R* define the concentration of the grid lines to a prescribed lines or points. Interchanging the dependent  $(\xi, \eta, \zeta)$  and the independent  $(x, y, z)$  variables results in

$$g^{11} \frac{\partial^2 \mathbf{r}}{\partial \xi^2} + g^{22} \frac{\partial^2 \mathbf{r}}{\partial \eta^2} + g^{33} \frac{\partial^2 \mathbf{r}}{\partial \zeta^2} + 2 \left( g^{12} \frac{\partial^2 \mathbf{r}}{\partial \xi \partial \eta} + g^{13} \frac{\partial^2 \mathbf{r}}{\partial \xi \partial \zeta} + g^{23} \frac{\partial^2 \mathbf{r}}{\partial \eta \partial \zeta} \right)$$

$$+P \frac{\partial \mathbf{r}}{\partial \xi} + Q \frac{\partial \mathbf{r}}{\partial \eta} + R \frac{\partial \mathbf{r}}{\partial \zeta} = 0,$$

or rewritten in components form

$$g^{11} \frac{\partial^2 x}{\partial \xi^2} + g^{22} \frac{\partial^2 x}{\partial \eta^2} + g^{33} \frac{\partial^2 x}{\partial \zeta^2} + 2 \left( g^{12} \frac{\partial^2 x}{\partial \xi \partial \eta} + g^{13} \frac{\partial^2 x}{\partial \xi \partial \zeta} + g^{23} \frac{\partial^2 x}{\partial \eta \partial \zeta} \right) + P \frac{\partial x}{\partial \xi} + Q \frac{\partial x}{\partial \eta} + R \frac{\partial x}{\partial \zeta} = 0, \quad (295)$$

$$g^{11} \frac{\partial^2 y}{\partial \xi^2} + g^{22} \frac{\partial^2 y}{\partial \eta^2} + g^{33} \frac{\partial^2 y}{\partial \zeta^2} + 2 \left( g^{12} \frac{\partial^2 y}{\partial \xi \partial \eta} + g^{13} \frac{\partial^2 y}{\partial \xi \partial \zeta} + g^{23} \frac{\partial^2 y}{\partial \eta \partial \zeta} \right) + P \frac{\partial y}{\partial \xi} + Q \frac{\partial y}{\partial \eta} + R \frac{\partial y}{\partial \zeta} = 0, \quad (296)$$

$$g^{11} \frac{\partial^2 z}{\partial \xi^2} + g^{22} \frac{\partial^2 z}{\partial \eta^2} + g^{33} \frac{\partial^2 z}{\partial \zeta^2} + 2 \left( g^{12} \frac{\partial^2 z}{\partial \xi \partial \eta} + g^{13} \frac{\partial^2 z}{\partial \xi \partial \zeta} + g^{23} \frac{\partial^2 z}{\partial \eta \partial \zeta} \right) + P \frac{\partial z}{\partial \xi} + Q \frac{\partial z}{\partial \eta} + R \frac{\partial z}{\partial \zeta} = 0, \quad (297)$$

where

$$g^{ij} = \frac{1}{g} \begin{pmatrix} g_{22}g_{33} - g_{23}^2 & g_{21}g_{33} - g_{23}g_{31} & g_{21}g_{32} - g_{22}g_{31} \\ g_{21}g_{33} - g_{23}g_{31} & g_{11}g_{33} - g_{13}^2 & g_{11}g_{32} - g_{12}g_{31} \\ g_{21}g_{32} - g_{22}g_{31} & g_{11}g_{32} - g_{12}g_{31} & g_{11}g_{22} - g_{12}^2 \end{pmatrix} \quad (298)$$

are the elements of the contravariant metric tensor which is symmetric, *Thompson, Warsi and Mastin* (1985). The elements of the contravariant metric tensor are computed as a function of the elements of the covariant metric tensor. The covariant metric tensor is defined as

$$g_{ij} = \begin{pmatrix} \mathbf{a}_1 \cdot \mathbf{a}_1 & \mathbf{a}_1 \cdot \mathbf{a}_2 & \mathbf{a}_1 \cdot \mathbf{a}_3 \\ \mathbf{a}_2 \cdot \mathbf{a}_1 & \mathbf{a}_2 \cdot \mathbf{a}_2 & \mathbf{a}_2 \cdot \mathbf{a}_3 \\ \mathbf{a}_3 \cdot \mathbf{a}_1 & \mathbf{a}_3 \cdot \mathbf{a}_2 & \mathbf{a}_3 \cdot \mathbf{a}_3 \end{pmatrix}$$

$$\begin{aligned}
 & \left( \begin{array}{ccc} \left(\frac{\partial x}{\partial \xi}\right)^2 + \left(\frac{\partial y}{\partial \xi}\right)^2 + \left(\frac{\partial z}{\partial \xi}\right)^2 & \frac{\partial x}{\partial \xi} \frac{\partial x}{\partial \eta} + \frac{\partial y}{\partial \xi} \frac{\partial y}{\partial \eta} + \frac{\partial z}{\partial \xi} \frac{\partial z}{\partial \eta} & \frac{\partial x}{\partial \xi} \frac{\partial x}{\partial \zeta} + \frac{\partial y}{\partial \xi} \frac{\partial y}{\partial \zeta} + \frac{\partial z}{\partial \xi} \frac{\partial z}{\partial \zeta} \\ \frac{\partial x}{\partial \xi} \frac{\partial x}{\partial \eta} + \frac{\partial y}{\partial \xi} \frac{\partial y}{\partial \eta} + \frac{\partial z}{\partial \xi} \frac{\partial z}{\partial \eta} & \left(\frac{\partial x}{\partial \eta}\right)^2 + \left(\frac{\partial y}{\partial \eta}\right)^2 + \left(\frac{\partial z}{\partial \eta}\right)^2 & \frac{\partial x}{\partial \eta} \frac{\partial x}{\partial \zeta} + \frac{\partial y}{\partial \eta} \frac{\partial y}{\partial \zeta} + \frac{\partial z}{\partial \eta} \frac{\partial z}{\partial \zeta} \\ \frac{\partial x}{\partial \xi} \frac{\partial x}{\partial \zeta} + \frac{\partial y}{\partial \xi} \frac{\partial y}{\partial \zeta} + \frac{\partial z}{\partial \xi} \frac{\partial z}{\partial \zeta} & \frac{\partial x}{\partial \eta} \frac{\partial x}{\partial \zeta} + \frac{\partial y}{\partial \eta} \frac{\partial y}{\partial \zeta} + \frac{\partial z}{\partial \eta} \frac{\partial z}{\partial \zeta} & \left(\frac{\partial x}{\partial \zeta}\right)^2 + \left(\frac{\partial y}{\partial \zeta}\right)^2 + \left(\frac{\partial z}{\partial \zeta}\right)^2 \end{array} \right), \\
 & \hspace{15em} (299)
 \end{aligned}$$

where

$$\mathbf{a}_1 = \frac{\partial x}{\partial \xi} \mathbf{i} + \frac{\partial y}{\partial \xi} \mathbf{j} + \frac{\partial z}{\partial \xi} \mathbf{k}, \quad (300)$$

$$\mathbf{a}_2 = \frac{\partial x}{\partial \eta} \mathbf{i} + \frac{\partial y}{\partial \eta} \mathbf{j} + \frac{\partial z}{\partial \eta} \mathbf{k}, \quad (301)$$

$$\mathbf{a}_3 = \frac{\partial x}{\partial \zeta} \mathbf{i} + \frac{\partial y}{\partial \zeta} \mathbf{j} + \frac{\partial z}{\partial \zeta} \mathbf{k}. \quad (302)$$

In fact all coefficients are functions of the elements in the *Jacobian matrix* of the coordinate transformation  $x = f(\xi, \eta, \zeta)$ ,  $y = g(\xi, \eta, \zeta)$ ,  $z = h(\xi, \eta, \zeta)$ ,

$$\mathbf{J}(\xi, \eta, \zeta) = \begin{pmatrix} \frac{\partial x}{\partial \xi} & \frac{\partial x}{\partial \eta} & \frac{\partial x}{\partial \zeta} \\ \frac{\partial y}{\partial \xi} & \frac{\partial y}{\partial \eta} & \frac{\partial y}{\partial \zeta} \\ \frac{\partial z}{\partial \xi} & \frac{\partial z}{\partial \eta} & \frac{\partial z}{\partial \zeta} \end{pmatrix}. \quad (303)$$

The elements of the Jacobian matrix are called *inverse metrics* of the coordinate transformation. The *Jacobian determinant* or *Jacobian* of the coordinate transformation is

$$\begin{aligned}
 \sqrt{g} &= |\mathbf{J}(\xi, \eta, \zeta)| \\
 &= \frac{\partial x}{\partial \xi} \left( \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \eta} \right) - \frac{\partial x}{\partial \eta} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \zeta} - \frac{\partial y}{\partial \zeta} \frac{\partial z}{\partial \xi} \right) + \frac{\partial x}{\partial \zeta} \left( \frac{\partial y}{\partial \xi} \frac{\partial z}{\partial \eta} - \frac{\partial y}{\partial \eta} \frac{\partial z}{\partial \xi} \right). \\
 & \hspace{15em} (304)
 \end{aligned}$$

For two dimensional case for instance for not existing  $j$ -direction set  $\partial x_j / \xi_j = 1$ ,  $\partial x_i / \xi_{j \neq i} = 0$ . A simple prescription for computation of  $g g^{ij}$  is

$$g g^{ij} = \alpha_{ij} = \sum_{m=1}^3 A_{mi} A_{mj} , \tag{298a}$$

Miki and Takagi (1984), where  $A_{mi}$  is the  $mi^{\text{th}}$  cofactor of the *Jacobian* matrix of the coordinate transformation, Eq. (303).  $A_{mi} = (\text{Subdeterminant } mi)(-1)^{m+i}$ .

*Boundary conditions:* The surface containing the domain of interest is divided into 6 non-overlapping patches,  $\Gamma_i$  ( $i = 1,2,3,4,5,6$ ), building 3 pairs perpendicular to each transformed directions. The boundary conditions, that is, the coordinates at the 6 patches are prescribed in advance as follows

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} f_i(\xi, \eta, \zeta) \\ g_i(\xi, \eta, \zeta) \\ h_i(\xi, \eta, \zeta) \end{pmatrix}, \quad (\xi, \eta, \zeta) \in \Gamma_i, \quad (i = 1, 2), \tag{305}$$

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} f_i(\xi, \eta_i, \zeta) \\ g_i(\xi, \eta_i, \zeta) \\ h_i(\xi, \eta_i, \zeta) \end{pmatrix}, \quad (\xi, \eta_i, \zeta) \in \Gamma_i, \quad (i = 3, 4), \tag{306}$$

$$\begin{pmatrix} x \\ y \\ z \end{pmatrix} = \begin{pmatrix} f_i(\xi, \eta, \zeta_i) \\ g_i(\xi, \eta, \zeta_i) \\ h_i(\xi, \eta, \zeta_i) \end{pmatrix}, \quad (\xi, \eta, \zeta_i) \in \Gamma_i, \quad (i = 5, 6). \tag{307}$$

*Line spacing control functions:* The line spacing control functions

$$P(\xi, \eta, \zeta) = -\sum_{l=1}^n a_{l1} \text{sign}(\xi - \xi_l) \exp(-b_{l1} T_l), \tag{308}$$

$$Q(\xi, \eta, \zeta) = -\sum_{l=1}^n a_{l2} \text{sign}(\eta - \eta_l) \exp(-b_{l2} T_l), \tag{309}$$

$$R(\xi, \eta, \zeta) = -\sum_{l=1}^n a_{l3} \text{sign}(\zeta - \zeta_l) \exp(-b_{l3} T_l), \tag{310}$$

where

$$T = \sqrt{c_{11}(\xi - \xi_l)^2 + c_{12}(\eta - \eta_l)^2 + c_{13}(\zeta - \zeta_l)^2}, \tag{311}$$

are recommended by Miki and Takagi (1984). The sets of the coefficients  $a_{lm}$  and  $c_{lm}$  attracts the  $\xi$ ,  $\eta$ , and/or  $\zeta = \text{constant}$  planes to the specified plane, coordinate line, or grid point. The range of the attraction effect is determined by the decay factor  $b_{lm}$ . The effect of the coefficient is demonstrated below:

Attracted planes	Planes, line, or point attracting neighboring planes
------------------	--

$a_{11} \neq 0$ for $\xi = \text{const}$ plane	$a_{11} \neq 0, a_{12} = 0, a_{13} = 0$ for $\xi = \xi_l$ plane $c_{11} \neq 0, c_{12} = 0, c_{13} = 0$
$a_{12} \neq 0$ for $\eta = \text{const}$ plane	$a_{11} = 0, a_{12} \neq 0, a_{13} = 0$ for $\eta = \eta_l$ plane $c_{11} = 0, c_{12} \neq 0, c_{13} = 0$
$a_{13} \neq 0$ for $\zeta = \text{const}$ plane	$a_{11} = 0, a_{12} = 0, a_{13} \neq 0$ for $\zeta = \zeta_l$ plane $c_{11} = 0, c_{12} = 0, c_{13} \neq 0$

For  $(\eta, \zeta) = (\eta_l, \zeta_l)$  line:

$a_{11} = 0, a_{12} \neq 0, a_{13} \neq 0$   
 $c_{11} = 0, c_{12} \neq 0, c_{13} \neq 0$

For  $(\xi, \zeta) = (\xi_l, \zeta_l)$  line:

$a_{11} = 0, a_{12} = 0, a_{13} \neq 0$   
 $c_{11} = 0, c_{12} = 0, c_{13} \neq 0$

For  $(\xi, \eta) = (\xi_l, \eta_l)$  line:

$a_{11} \neq 0, a_{12} \neq 0, a_{13} = 0$   
 $c_{11} \neq 0, c_{12} \neq 0, c_{13} = 0$

For  $(\xi, \eta, \zeta) = (\xi_l, \eta_l, \zeta_l)$  point:

$a_{11} \neq 0, a_{12} \neq 0, a_{13} \neq 0$   
 $c_{11} \neq 0, c_{12} \neq 0, c_{13} \neq 0$

*Simplification for 2D plane:* For grid generation in 2D plane the grid generating system  $(\xi, \eta)$  simplifies to

$$g^{11} \frac{\partial^2 x}{\partial \xi^2} + g^{22} \frac{\partial^2 x}{\partial \eta^2} + 2g^{12} \frac{\partial^2 x}{\partial \xi \partial \eta} + P \frac{\partial x}{\partial \xi} + Q \frac{\partial x}{\partial \eta} = 0, \tag{312}$$

$$g^{11} \frac{\partial^2 y}{\partial \xi^2} + g^{22} \frac{\partial^2 y}{\partial \eta^2} + 2g^{12} \frac{\partial^2 y}{\partial \xi \partial \eta} + P \frac{\partial y}{\partial \xi} + Q \frac{\partial y}{\partial \eta} = 0, \tag{313}$$

where

$$g^{ij} = \frac{1}{g} \begin{pmatrix} g_{22} & g_{21} \\ g_{21} & g_{11} \end{pmatrix} = \frac{1}{g} \begin{pmatrix} \left( \frac{\partial x}{\partial \eta} \right)^2 + \left( \frac{\partial y}{\partial \eta} \right)^2 & \frac{\partial x}{\partial \xi} \frac{\partial x}{\partial \eta} + \frac{\partial y}{\partial \xi} \frac{\partial y}{\partial \eta} \\ \frac{\partial x}{\partial \xi} \frac{\partial x}{\partial \eta} + \frac{\partial y}{\partial \xi} \frac{\partial y}{\partial \eta} & \left( \frac{\partial x}{\partial \xi} \right)^2 + \left( \frac{\partial y}{\partial \xi} \right)^2 \end{pmatrix}, \tag{314}$$

and

$$\sqrt{g} = \frac{\partial x}{\partial \xi} \frac{\partial y}{\partial \eta} - \frac{\partial x}{\partial \eta} \frac{\partial y}{\partial \xi}. \quad (315)$$

*Numerical solution:* Central difference numerical scheme is applied for computation of the partial derivatives assuming  $\Delta \xi = \Delta \eta = \Delta \zeta = 1$  :

$$\frac{\partial^2 x}{\partial \xi^2} = \frac{1}{4}(x_{i+1} - 2x + x_{i-1}), \quad \frac{\partial x}{\partial \xi} = \frac{1}{2}(x_{i+1} - x_{i-1}), \quad (316, 317)$$

$$\frac{\partial^2 x}{\partial \eta^2} = \frac{1}{4}(x_{j+1} - 2x + x_{j-1}), \quad \frac{\partial x}{\partial \eta} = \frac{1}{2}(x_{j+1} - x_{j-1}), \quad (318, 319)$$

$$\frac{\partial^2 x}{\partial \zeta^2} = \frac{1}{4}(x_{k+1} - 2x + x_{k-1}), \quad \frac{\partial x}{\partial \zeta} = \frac{1}{2}(x_{k+1} - x_{k-1}), \quad (320, 321)$$

$$\frac{\partial^2 y}{\partial \xi^2} = \frac{1}{4}(y_{i+1} - 2y + y_{i-1}), \quad \frac{\partial y}{\partial \xi} = \frac{1}{2}(y_{i+1} - y_{i-1}), \quad (322, 323)$$

$$\frac{\partial^2 y}{\partial \eta^2} = \frac{1}{4}(y_{j+1} - 2y + y_{j-1}), \quad \frac{\partial y}{\partial \eta} = \frac{1}{2}(y_{j+1} - y_{j-1}), \quad (324, 325)$$

$$\frac{\partial^2 y}{\partial \zeta^2} = \frac{1}{4}(y_{k+1} - 2y + y_{k-1}), \quad \frac{\partial y}{\partial \zeta} = \frac{1}{2}(y_{k+1} - y_{k-1}), \quad (326, 327)$$

$$\frac{\partial^2 z}{\partial \xi^2} = \frac{1}{4}(z_{i+1} - 2z + z_{i-1}), \quad \frac{\partial z}{\partial \xi} = \frac{1}{2}(z_{i+1} - z_{i-1}), \quad (326, 327)$$

$$\frac{\partial^2 z}{\partial \eta^2} = \frac{1}{4}(z_{j+1} - 2z + z_{j-1}), \quad \frac{\partial z}{\partial \eta} = \frac{1}{2}(z_{j+1} - z_{j-1}), \quad (328, 329)$$

$$\frac{\partial^2 z}{\partial \zeta^2} = \frac{1}{4}(z_{k+1} - 2z + z_{k-1}), \quad \frac{\partial z}{\partial \zeta} = \frac{1}{2}(z_{k+1} - z_{k-1}), \quad (330, 331)$$

$$\frac{\partial^2 x}{\partial \xi \partial \eta} = \frac{\partial}{\partial \xi} \frac{\partial x}{\partial \eta} = \frac{1}{4}(x_{i+1,j+1} - x_{i+1,j-1} - x_{i-1,j+1} + x_{i-1,j-1}), \quad (332, 333)$$

$$\frac{\partial^2 x}{\partial \xi \partial \zeta} = \frac{\partial}{\partial \xi} \frac{\partial x}{\partial \zeta} = \frac{1}{4}(x_{i+1,k+1} - x_{i+1,k-1} - x_{i-1,k+1} + x_{i-1,k-1}), \quad (334, 335)$$

$$\frac{\partial^2 x}{\partial \eta \partial \zeta} = \frac{\partial}{\partial \eta} \frac{\partial x}{\partial \zeta} = \frac{1}{4}(x_{j+1,k+1} - x_{j+1,k-1} - x_{j-1,k+1} + x_{j-1,k-1}), \quad (336, 337)$$

$$\frac{\partial^2 y}{\partial \xi \partial \eta} = \frac{\partial}{\partial \xi} \frac{\partial y}{\partial \eta} = \frac{1}{4} (y_{i+1,j+1} - y_{i+1,j-1} - y_{i-1,j+1} + y_{i-1,j-1}), \quad (339, 340)$$

$$\frac{\partial^2 y}{\partial \xi \partial \zeta} = \frac{\partial}{\partial \xi} \frac{\partial y}{\partial \zeta} = \frac{1}{4} (y_{i+1,k+1} - y_{i+1,k-1} - y_{i-1,k+1} + y_{i-1,k-1}), \quad (341, 342)$$

$$\frac{\partial^2 y}{\partial \eta \partial \zeta} = \frac{\partial}{\partial \eta} \frac{\partial y}{\partial \zeta} = \frac{1}{4} (y_{j+1,k+1} - y_{j+1,k-1} - y_{j-1,k+1} + y_{j-1,k-1}), \quad (343, 344)$$

$$\frac{\partial^2 z}{\partial \xi \partial \eta} = \frac{\partial}{\partial \xi} \frac{\partial z}{\partial \eta} = \frac{1}{4} (z_{i+1,j+1} - z_{i+1,j-1} - z_{i-1,j+1} + z_{i-1,j-1}), \quad (345, 346)$$

$$\frac{\partial^2 z}{\partial \xi \partial \zeta} = \frac{\partial}{\partial \xi} \frac{\partial z}{\partial \zeta} = \frac{1}{4} (z_{i+1,k+1} - z_{i+1,k-1} - z_{i-1,k+1} + z_{i-1,k-1}), \quad (347, 348)$$

$$\frac{\partial^2 z}{\partial \eta \partial \zeta} = \frac{\partial}{\partial \eta} \frac{\partial z}{\partial \zeta} = \frac{1}{4} (z_{j+1,k+1} - z_{j+1,k-1} - z_{j-1,k+1} + z_{j-1,k-1}). \quad (349, 350)$$

Replacing in Eq. (312) results in

$$c_{k-1}x_{k-1} + c_{j-1}x_{j-1} + c_{i-1}x_{i-1} + cx + c_{i+1}x_{i+1} + c_{j+1}x_{j+1} + c_{k+1}x_{k+1} = d_1,$$

$$c_{k-1}y_{k-1} + c_{j-1}y_{j-1} + c_{i-1}y_{i-1} + cy + c_{i+1}y_{i+1} + c_{j+1}y_{j+1} + c_{k+1}y_{k+1} = d_2,$$

$$c_{k-1}z_{k-1} + c_{j-1}z_{j-1} + c_{i-1}z_{i-1} + cz + c_{i+1}z_{i+1} + c_{j+1}z_{j+1} + c_{k+1}z_{k+1} = d_3,$$

where

$$c_{k+1} = g^{33}/2 + R, \quad c_{j+1} = g^{22}/2 + Q, \quad c_{i+1} = g^{11}/2 + P,$$

$$c_{k-1} = g^{33}/2 - R, \quad c_{j-1} = g^{22}/2 - Q, \quad c_{i-1} = g^{11}/2 - P,$$

$$c = -(g^{11} + g^{22} + g^{33}) = -(c_{k-1} + c_{j-1} + c_{i-1} + c_{i+1} + c_{j+1} + c_{k+1}),$$

$$d_1 = -g^{12} (x_{i+1,j+1} - x_{i+1,j-1} - x_{i-1,j+1} + x_{i-1,j-1})$$

$$-g^{13} (x_{i+1,k+1} - x_{i+1,k-1} - x_{i-1,k+1} + x_{i-1,k-1}) - g^{22} (x_{j+1,k+1} - x_{j+1,k-1} - x_{j-1,k+1} + x_{j-1,k-1}),$$

$$d_2 = -g^{12} (y_{i+1,j+1} - y_{i+1,j-1} - y_{i-1,j+1} + y_{i-1,j-1})$$

$$-g^{13} (y_{i+1,k+1} - y_{i+1,k-1} - y_{i-1,k+1} + y_{i-1,k-1}) - g^{22} (y_{j+1,k+1} - y_{j+1,k-1} - y_{j-1,k+1} + y_{j-1,k-1}),$$

$$d_3 = -g^{12} (z_{i+1,j+1} - z_{i+1,j-1} - z_{i-1,j+1} + z_{i-1,j-1})$$



$$-g^{13} (z_{i+1,k+1} - z_{i+1,k-1} - z_{i-1,k+1} + z_{i-1,k-1}) - g^{22} (z_{j+1,k+1} - z_{j+1,k-1} - z_{j-1,k+1} + z_{j-1,k-1}),$$

The resulting system is solved using SOR methods. Note that the coefficient matrix is the same for all directions. Therefore once the matrix is inverted it can be used for the three right hand sites which saves computer time. Do not forget to treat properly the boundary conditions by putting into the right hand sites the known boundary terms.

*Orthogonal systems:* Numerical orthogonality is reached by setting the off-diagonal elements of the contravariant metric tensor equal to zero. This result in

$$\frac{\partial}{\partial \xi} \left( \frac{\partial x}{\partial \xi} \sqrt{\frac{g_{22} g_{33}}{g_{11}}} \right) + \frac{\partial}{\partial \eta} \left( \frac{\partial x}{\partial \eta} \sqrt{\frac{g_{11} g_{33}}{g_{22}}} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\partial x}{\partial \zeta} \sqrt{\frac{g_{11} g_{22}}{g_{33}}} \right) = 0, \quad (351)$$

$$\frac{\partial}{\partial \xi} \left( \frac{\partial y}{\partial \xi} \sqrt{\frac{g_{22} g_{33}}{g_{11}}} \right) + \frac{\partial}{\partial \eta} \left( \frac{\partial y}{\partial \eta} \sqrt{\frac{g_{11} g_{33}}{g_{22}}} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\partial y}{\partial \zeta} \sqrt{\frac{g_{11} g_{22}}{g_{33}}} \right) = 0, \quad (352)$$

$$\frac{\partial}{\partial \xi} \left( \frac{\partial z}{\partial \xi} \sqrt{\frac{g_{22} g_{33}}{g_{11}}} \right) + \frac{\partial}{\partial \eta} \left( \frac{\partial z}{\partial \eta} \sqrt{\frac{g_{11} g_{33}}{g_{22}}} \right) + \frac{\partial}{\partial \zeta} \left( \frac{\partial z}{\partial \zeta} \sqrt{\frac{g_{11} g_{22}}{g_{33}}} \right) = 0. \quad (353)$$

see *Thompson and Warsi (1982)*. If, in addition to orthogonality, the condition  $g_{11} = g_{22} = g_{33} = const$  is imposed, then the above three equations reduce to *Laplace* equations.

For 2D systems the equations set (351-353) reduces to

$$\frac{\partial}{\partial \xi} \left( f_r \frac{\partial x}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( \frac{1}{f_r} \frac{\partial x}{\partial \eta} \right) = 0, \quad (354)$$

$$\frac{\partial}{\partial \xi} \left( f_r \frac{\partial y}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( \frac{1}{f_r} \frac{\partial y}{\partial \eta} \right) = 0, \quad (355)$$

where

$$f_r = \sqrt{g_{22}/g_{11}} > 0 \quad (356)$$

is called “distortion function”. It specifies the ratio of the sides of small rectangle in the  $x, y$  plane which is mapped onto a rectangle in the  $\xi, \eta$  plane. A reasonable upper limit for  $f_r$  is 10. The condition of orthogonality also leads to

$$\frac{\partial x}{\partial \eta} = -f_r \frac{\partial y}{\partial \xi}, \quad \frac{\partial y}{\partial \eta} = f_r \frac{\partial x}{\partial \xi}. \quad (357, 358)$$

This conditions have to be satisfied also at the boundary. For

$$f_r(\xi, \eta) = 1, \tag{359}$$

Eqs. (357, 358) become the *Cauchy-Riemann* equations which are used in conformal mapping. *Tamamidis* and *Assanis* (1991) recognized that non uniform  $f_r$  over the domain of interest can be used for controlling the spacing of the nodes even generating orthogonal grids. The authors successfully examined the properties of

$$\frac{\partial^2 f_r}{\partial \xi^2} + \frac{\partial^2 f_r}{\partial \eta^2} = G(\xi, \eta) \tag{360}$$

for generating a smooth  $f_r$ 's over the domain of interests.  $G(\xi, \eta)$  typically involve some combination of sinusoidal, cosinusoidal and exponential functions, i.e.,

$$G(\xi, \eta) = \text{const } f_1(\xi) f_2(\eta). \tag{361}$$

An initial approximation can be obtained by setting  $f_r = 1$ . The authors generated coefficients of the algebraic systems finite difference to Eqs. (354) and (355) having the same sign by using for the first order differences: forward finite differences for  $x_\xi$  when  $f_\xi$  is positive, and backward differences when  $f_\xi$  is negative; backward finite differences for  $x_\eta$  when  $f_\eta$  is positive, and forward differences when  $f_\eta$  is negative.

Setting

$$f = (g_{11}g_{22}g_{33})^{-1/2} = g^{-1/2} \tag{362}$$

the equations set (351-353) can be rewritten as

$$\frac{\partial}{\partial \xi} \left( fg_{22}g_{33} \frac{\partial x}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( fg_{11}g_{33} \frac{\partial x}{\partial \eta} \right) + \frac{\partial}{\partial \zeta} \left( fg_{11}g_{22} \frac{\partial x}{\partial \zeta} \right) = 0, \tag{363}$$

$$\frac{\partial}{\partial \xi} \left( fg_{22}g_{33} \frac{\partial y}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( fg_{11}g_{33} \frac{\partial y}{\partial \eta} \right) + \frac{\partial}{\partial \zeta} \left( fg_{11}g_{22} \frac{\partial y}{\partial \zeta} \right) = 0, \tag{364}$$

$$\frac{\partial}{\partial \xi} \left( fg_{22}g_{33} \frac{\partial z}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( fg_{11}g_{33} \frac{\partial z}{\partial \eta} \right) + \frac{\partial}{\partial \zeta} \left( fg_{11}g_{22} \frac{\partial z}{\partial \zeta} \right) = 0. \tag{365}$$

*Christov* (1982) proposed to use  $f = f(x, y, z)$  for controlling the spacing of the grid by controlling actually the local value of the Jacobian. For 2D,

$$\frac{\partial}{\partial \xi} \left( fg_{22} \frac{\partial x}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( fg_{11} \frac{\partial x}{\partial \eta} \right) = 0, \tag{366}$$

$$\frac{\partial}{\partial \xi} \left( fg_{22} \frac{\partial y}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( fg_{11} \frac{\partial y}{\partial \eta} \right) = 0, \quad (367)$$

*Christov* proposed to impose a control function of the type

$$f = \sqrt{1 + z_x^2 + z_y^2}, \quad (368)$$

that is simply the slope of a predefined surface  $z = z(x, y)$ . The stronger the slope of the chosen surface at a given point the more dense the grid around this point.

*The Sorenson method for enforcing orthogonality at the boundaries: Sorenson* (1989) oriented his development towards creating a multiple interconnected blocks. He simplified the  $P$ ,  $Q$ , and  $R$  functions to facilitate attractions to the corresponding outer boundary only by specifying attenuation factors so as not to influence the opposite boundary surface e.g.

$$\begin{aligned} P(\xi, \eta, \zeta) = & p_1(\eta, \zeta) \exp[-a(\xi - 1)] + p_2(\eta, \zeta) \exp[-a(\xi_{\max} - \xi)] \\ & + p_3(\xi, \zeta) \exp[-a(\eta - 1)] + p_4(\xi, \zeta) \exp[-a(\eta_{\max} - \eta)] \\ & + p_5(\xi, \eta) \exp[-a(\zeta - 1)] + p_6(\xi, \eta) \exp[-a(\zeta_{\max} - \zeta)]. \end{aligned} \quad (369)$$

By appropriate selection of the  $p$ ,  $q$ ,  $r$  coefficients he enforced the orthogonality at the boundary. As example consider face 3. At this face the derivatives  $\mathbf{r}_\xi$ ,  $\mathbf{r}_\zeta$ ,  $\mathbf{r}_{\xi\xi}$ ,  $\mathbf{r}_{\xi\xi}$ , and  $\mathbf{r}_{\xi\xi}$  are computed from the point values known at the surface once for the generation process. The derivatives  $\mathbf{r}_\eta$  are found from the desired clustering and orthogonality on face 3. The desired orthogonality and clustering is specified by the following three relations  $\mathbf{r}_\xi \cdot \mathbf{r}_\eta = 0$ ,  $\mathbf{r}_\zeta \cdot \mathbf{r}_\eta = 0$ ,  $\mathbf{r}_\eta \cdot \mathbf{r}_\eta = S^2$ , where  $S$  is the height to be imposed on the cell on the boundary or in expanded form

$$x_\xi x_\eta + y_\xi y_\eta + z_\xi z_\eta = 0, \quad (370)$$

$$x_\zeta x_\eta + y_\zeta y_\eta + z_\zeta z_\eta = 0, \quad (371)$$

$$x_\eta^2 + y_\eta^2 + z_\eta^2 = S^2. \quad (372)$$

Solving the first two equations with respect to  $x_\eta$  and  $y_\eta$  results in

$$x_\eta = z_\eta \left( -z_\xi y_\zeta + z_\zeta y_\xi \right) / \left( x_\xi y_\zeta - x_\zeta y_\xi \right), \quad (373)$$

$$y_\eta = z_\eta \left( -x_\xi z_\zeta + x_\zeta z_\xi \right) / \left( x_\xi y_\zeta - x_\zeta y_\xi \right), \quad (374)$$

or comparing with the cofactors of the *Jacobian* matrix of the coordinate transformation

$$x_\eta = z_\eta (-A_{12}) / (-A_{32}), \quad (375)$$

$$y_\eta = z_\eta (-A_{22}) / (-A_{32}). \quad (376)$$

Substituting in the third equation and solving with respect to  $z_\eta$  we obtain

$$z_\eta = \frac{SA_{32}}{\pm\sqrt{A_{12}^2 + A_{22}^2 + A_{32}^2}}, \quad (377)$$

and consequently

$$x_\eta = \frac{SA_{12}}{\pm\sqrt{A_{12}^2 + A_{22}^2 + A_{32}^2}}, \quad (378)$$

$$y_\eta = \frac{SA_{22}}{\pm\sqrt{A_{12}^2 + A_{22}^2 + A_{32}^2}}. \quad (379)$$

The positive sign is used for right-handed and the negative for left-handed coordinate systems. Now we have the derivatives  $\mathbf{r}_\eta$ . These derivatives can be differentiated to obtain the mixed derivatives  $\mathbf{r}_{\eta\xi}$ ,  $\mathbf{r}_{\eta\zeta}$ . All these derivatives are obtained as a function of the information stored at the surface and have to be computed once over the generation process. The only derivatives lacking is

$$(\mathbf{r}_{\eta\eta})_{i,1,k} = \frac{-7\mathbf{r}_{i,1,k} + 8\mathbf{r}_{i,2,k} - \mathbf{r}_{i,3,k}}{2(\Delta\eta)^2} - \frac{3\mathbf{r}_{\eta i,1,k}}{\Delta\eta}, \quad (380)$$

which results from the *Taylor* series. Now the governing three equations written for the point  $(i,1,k)$

$$\begin{aligned} & \frac{\partial \mathbf{r}}{\partial \xi} p_{i,1,k} + \frac{\partial \mathbf{r}}{\partial \eta} q_{i,1,k} + \frac{\partial \mathbf{r}}{\partial \zeta} r_{i,1,k} \\ &= -\frac{1}{g} \left[ \alpha_{11} \frac{\partial^2 \mathbf{r}}{\partial \xi^2} + \alpha_{33} \frac{\partial^2 \mathbf{r}}{\partial \zeta^2} + 2 \left( \alpha_{12} \frac{\partial^2 \mathbf{r}}{\partial \xi \partial \eta} + \alpha_{13} \frac{\partial^2 \mathbf{r}}{\partial \xi \partial \zeta} + \alpha_{23} \frac{\partial^2 \mathbf{r}}{\partial \eta \partial \zeta} \right) \right] - \frac{\alpha_{22}}{g} \frac{\partial^2 \mathbf{r}}{\partial \eta^2} \end{aligned}$$

can be solved with respect to the unknown coefficients  $p$ ,  $q$  and  $r$ . In each successive iteration the value of the derivatives  $(\mathbf{r}_{\eta\eta})_{i,1,k}$  and consequently the values of  $p$ ,  $q$  and  $r$  are updated until convergence is reached.

*Practical recommendation:* Specify first the 6 non-overlapping patches. Divide the peripheral boundary on 4 patches. Generate the grid on each boundary line. Use simple 2D generating system and generate approximate 2D grid at each patch. Improve the 2D grid by using elliptic 2D generating system. Use simple algebraic system to generate initial guess for the 3D system. Generate 3D grids by using 3D

elliptic generating system. If other domain has to be attached to the already generated grid specify the patch at which the new domain will be attached. Then use the same procedure to generate the 3D grid by using the already generated grid on the common patch. Plot the grid and control the result.

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