

# Chapter 15

## Overview of Molecular Basics



- 1. (a)** For  $T = 273$  K and pressure 1 atmosphere, that is  $10^6$  dyne  $\text{cm}^{-2}$  (760 mm of Hg), find the density,  $n$ , of an ideal gas in  $\text{cm}^{-3}$ . Repeat for conditions in a molecular cloud, that is  $T = 10$  K, pressure  $10^{-12}$  mm of Hg.
- (b)** For both sets of conditions, find the mean free path,  $\lambda$ , which is defined as  $1/(\sigma n)$ , and the mean time between collisions,  $\tau$ , which is  $1/(\sigma n v)$ , where  $v$  is the average velocity. In both cases, take  $\sigma = 10^{-16}$   $\text{cm}^{-3}$ . For the laboratory, take the average velocity to be  $300$   $\text{m s}^{-1}$ ; for the molecular cloud, take the average velocity of  $\text{H}_2$  as  $0.2$   $\text{km s}^{-1}$ .
- (c)** Suppose that the population of the upper level of a molecule decays in  $10^5$  s. How many collisions in both cases occur before a decay?
- (d)** For extinction we define the penetration depth,  $\lambda_v$ , in analogy with the mean free path. When  $\lambda_v = 1$  the light from a background star is reduced by a factor 0.3678. For a density of atoms  $n$ ,  $\lambda_v$ , in cm, is  $2 \times 10^{21}/n$ . Calculate the value of  $\lambda_v$  for a molecular cloud and for standard laboratory conditions. The parameters for both are given in part **(a)** of this problem.

- 2. (a)** The result of problem 2(c) of chapter 13 gives  $T_k = 21.2 (m/m_H) (\Delta V_t)^2$  where  $\Delta V_t$  is the FWHP thermal width, i.e. there is no turbulence and the gas has a Maxwell–Boltzmann distribution. Apply this formula to the CO molecule (mass  $28 m_H$ ) for a gas of temperature  $T$ . What is  $\Delta V_t$  for  $T = 10$  K,  $T = 100$  K,  $T = 200$  K?
- (b)** The observed linewidth is  $3$   $\text{km s}^{-1}$  in a dark cloud for which  $T = 10$  K. What is the turbulent velocity width in such a cloud if the relation between the observed FWHP linewidth,  $\Delta V_{1/2}$ , the thermal linewidth,  $\Delta V_t$  and the turbulent linewidth  $\Delta V_{\text{turb}}$  is

$$\Delta V_{1/2}^2 = \Delta V_t^2 + \Delta V_{\text{turb}}^2 ?$$

3. The following expression is appropriate for the spontaneous decay between two rotational levels,  $(u, l)$  of a linear molecule:  $A_{ul} = 1.165 \times 10^{-11} \mu_0^2 \nu^3 (J+1)/(2J+3)$  where  $\nu$  is in GHz,  $\mu_0$  is in Debyes and  $J$  is the lower level in the transition from  $J+1 \rightarrow J$ . Use this to estimate the Einstein  $A$  coefficient for a system with a dipole moment of 0.1 Debye for a transition from the  $J=1$  level to the  $J=0$  level at 115.271 GHz.

4. To determine whether a given level is populated, one frequently makes use of the concept of the “critical density”,  $n^*$ , defined as:

$$A_{ul} = n^* \langle \sigma v \rangle .$$

where  $u$  is the quantum number of the upper rotational level, and  $l$  is that for the lower level. If we take  $\langle \sigma v \rangle$  to be  $10^{-10} \text{ cm}^3 \text{ s}^{-1}$ , determine  $n^*$  from the following  $A_{ul}$  coefficients

$$\text{CS} : A_{10} = 1.8 \times 10^{-6} \text{ s}^{-1}$$

$$\text{CS} : A_{21} = 2.2 \times 10^{-5} \text{ s}^{-1}$$

$$\text{CO} : A_{10} = 7.4 \times 10^{-8} \text{ s}^{-1} .$$

5. Suppose the effective radius  $r_e = 1.1 \times 10^{-8} \text{ cm}$  and the reduced mass,  $m_r$ , of a perfectly rigid molecule is 10 atomic mass units, AMU (an AMU is 1/16 of the mass of a 16-oxygen atom;  $1 \text{ AMU} = 1.660 \times 10^{-24} \text{ g}$ ), where  $\Theta = m_r r_e^2$ .

(a) Calculate the lowest four rotational frequencies and energies of the levels above the ground state. One needs a simplified version of Eq. 15.11 to 15.13 from ‘Tools’ for the rotational constant is

$$B_e = \frac{\hbar}{4\pi \Theta_e} \quad (15.1)$$

The energy of level  $J$  is:

$$E_{\text{rot}} = W(J) = \frac{\hbar^2}{2\Theta_e} J(J+1) - hD [J(J+1)]^2 . \quad (15.2)$$

and the frequency is the difference between the energy of level  $J+1$  and  $J$  divided by the Planck constant:

$$\nu(J) = \frac{1}{h} [W(J+1) - W(J)] = 2 B_e [(J+1) - 4D(J+1)^3] \quad (15.3)$$

(b) Repeat if the reduced mass is  $(2/3) \text{ AMU}$  with a separation of  $0.75 \times 10^{-8} \text{ cm}$ ; this is appropriate for the HD molecule. The HD molecule has a dipole moment  $\mu_0 = 10^{-4} \text{ Debye}$ , caused by the fact that the center of mass is not coincident with

**Table 15.1** Parameters of the more commonly observed carbon monoxide lines (problem 6)

Chemical <sup>a</sup> formula	Molecule name	Transition	$\nu/\text{GHz}^b$	$E_u/\text{K}^c$	$A_{ij}/\text{s}^{-1d}$
$\text{C}^{18}\text{O}$	Carbon monoxide	$J = 1 - 0$	109.782182	5.3	$6.5 \times 10^{-8}$
$^{13}\text{CO}$	Carbon monoxide	$J = 1 - 0$	110.201370	5.3	$6.5 \times 10^{-8}$
CO	Carbon monoxide	$J = 1 - 0$	115.271203	5.5	$7.4 \times 10^{-8}$
$\text{C}^{18}\text{O}$	Carbon monoxide	$J = 2 - 1$	219.560319	15.9	$6.2 \times 10^{-7}$
$^{13}\text{CO}$	Carbon monoxide	$J = 2 - 1$	220.398714	15.9	$6.2 \times 10^{-7}$
CO	Carbon monoxide	$J = 2 - 1$	230.538001	16.6	$7.1 \times 10^{-7}$

<sup>a</sup> If isotope not explicitly given, this is the most abundant variety, i.e.,  $^{12}\text{C}$  is C,  $^{16}\text{O}$  is O,  $^{14}\text{N}$  is N

<sup>b</sup> From Lovas (1992, J. Chem. Phys. Ref. Data 21, 18)

<sup>c</sup> Energy above the ground state in Kelvins

<sup>d</sup> Spontaneous transition rate, i.e., the Einstein A coefficient

the center of charge. Take the expression for  $A(ul)$  from Problem 3 and apply to the  $J = 1 - 0$  and  $J = 2 - 1$  transitions of the HD molecule.

(c) Find the “critical density”,  $n^* \approx 10^{10} A(ul)$ .

6. The  $^{12}\text{C}^{16}\text{O}$  molecule has  $B_e = 57.6360 \text{ GHz}$  and  $D_e = 0.185 \text{ MHz}$ . Calculate the energies for the  $J = 1, 2, 3, 4, 5$  levels and line frequencies for the  $J = 1 - 0, 2 - 1, 3 - 2, 4 - 3$  and  $5 - 4$  transitions. Use the expression energy  $E(J)/h \approx B_e J(J+1) - D_e J^2(J+1)^2$  for the energy calculation. Check the results against the relevant parts of Table 16.1 in ‘Tools’, given here as Table 15.1.

7. Apply for  $J = 0, 1$  the analysis in problem 6 to the linear molecule  $\text{HC}_{11}\text{N}$ , which has  $B_e = 169.06295 \text{ MHz}$  and  $D_e = 0.24 \text{ Hz}$ . Estimate  $J$  for a transition near 20 GHz. What is the error if one neglects the distortion term?

8. In the following, we neglect the distortion term  $D_e$  and assume that the population is in LTE. The population in a given  $J$  level for a linear molecule is given by Eq. (15.33):

$$n(J)/n(\text{total}) = (2J + 1)e^{B_0 J(J+1)/kT} / Z$$

where  $Z$ , the partition function, does not depend on  $J$ . Differentiate  $n(J)$  with respect to  $J$  to find the state which has the largest population for a fixed value of temperature,  $T$ . Calculate this for CO if  $T = 10 \text{ K}$  and  $T = 100 \text{ K}$ . Repeat for CS ( $B_0 = 24.584 \text{ GHz}$ ) and  $\text{HC}_{11}\text{N}$ , for  $T = 10 \text{ K}$ .

9. Extend Eq. (15.33 in ‘Tools’), which is:

$$N(J)/N(\text{total}) = \frac{(2J + 1)}{Z} \exp\left[-\frac{h B_e J(J + 1)}{kT}\right]$$

to include the optical depth relation Eq. (15.26), which is:

$$N_l = 1.95 \times 10^3 \frac{g_l v^2}{g_u A_{ul}} \int T_B dV$$

to obtain an estimate of which  $J$  level has the largest optical depth,  $\tau$ , in the case of emission for a linear molecule.

(a) Show that when the expression for the  $A$  coefficient for a linear molecule is inserted into Eq. (15.26 of 'Tools'), we have

$$N_1 = \frac{1.67 \times 10^{14}}{\mu_0^2 \nu[\text{GHz}]} \times \frac{2J+1}{J+1} T_{\text{ex}} \tau \Delta \nu,$$

where  $\mu$  is in Debyes and  $\nu$  is in  $\text{km s}^{-1}$ .

(b) Use the above expression to estimate whether the  $J$  for the maximum  $T_{\text{MB}} = T_{\text{ex}} \tau$  is larger or smaller than the  $J$  obtained in Problem 8.

**10.** Find the ratio of the intensities of the  $J = 2 - 1$  to  $J = 1 - 0$  transitions for a linear molecule if the excitation temperature of the system,  $T$ , is very large compared to the energy of the  $J = 2$  level above the ground state, and both lines are optically thin. What is the ratio if both are optically thick? Use the last equation in the statement of Problem 9 of this Chapter.

**11.** The ammonia molecule,  $\text{NH}_3$ , is an oblate symmetric top. For ammonia,  $B = 298$  GHz,  $C = 189$  GHz. If  $T \gg B, C$ , the value of  $Z$ , the partition function, with  $C$  and  $B$  in GHz, is  $Z = 168.7 \sqrt{(T^3)/(B^2 A)}$ .

(a) Evaluate  $Z$  for  $\text{NH}_3$  for  $T = 50$  K, 100 K, 200 K, 300 K. For this approximation to be valid, what is a lower limit to the value of  $T$ ?

(b) The (3,3) levels are 120 K above ground. Use the partition function and

$$n(J)/n(\text{total}) = (2J+1)e^{120/T}/Z$$

to calculate the ratio of the total population to that in the (3,3) levels.

(c) If only metastable ( $J = K$ ) levels are populated, use the definition of  $Z$  as a sum over all populated states, and

$$n(J)/n(\text{total}) = (2J+1)e^{(BJ(J+1)+K^2(C-B))/kT}/Z$$

and the  $B$  and  $C$  values for  $\text{NH}_3$  to obtain the ratio between the population of the (3,3) levels and all metastable levels.

**12.** The selection rules for dipole transitions of the doubly deuterated isotopomer  $\text{D}_2\text{CO}$  differ from that of  $\text{H}_2\text{CO}$  since  $\text{D}_2\text{CO}$  has two Bosons, so the symmetry of the total wavefunction must be symmetric. Determine these rules following the procedure in Sect. 15.6.2.