

Photophysics of Hybrid and Inorganic Lead Halide Perovskites



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1 Introduction

This chapter focuses on the optoelectronic properties of hybrid and all-inorganic lead halide perovskites. Such materials adopt the chemical stoichiometry ABX_3 and are strong contenders for applications in solar photovoltaics. Among leading candidates are systems where the A cation is organic (e.g., methylammonium, $CH_3NH_3^+$, MA; formamidinium, $CH(NH_2)_2^+$, FA) or inorganic (e.g., Cs^+), the B cation is Pb^{2+} , and the X anion is I^- , Br^- , or Cl^- . Important systems come from the methylammonium lead halide ($CH_3NH_3PbX_3$, or $MAPbX_3$), formamidinium lead halide [$CH(NH_2)_2PbX_3$, or $FAPbX_3$], and cesium lead halide ($CsPbX_3$) families. Alloys are also possible and $APbX_3$ materials can be produced as mixed cation, mixed anion, or even mixed cation/mixed anion alloys. Common mixed cation systems include $A = FA/Cs$ or $A = FA/MA$ alloys, e.g., $FA_{1-y}Cs_yPbI_3$ or $FA_{1-y}MA_yPbI_3$. Mixed anion materials are often mixtures of iodine and bromine such as $MAPb(I_{1-x}Br_x)_3$ and $FAPb(I_{1-x}Br_x)_3$ while mixed cation/anion systems include $FA_xMA_yCs_{1-x-y}Pb(I_{1-z}Br_z)_3$ (FAMACs).

Given prior discussion about the unique structural properties of ABX_3 materials, we simply recall here that $APbX_3$ perovskites adopt cubic/quasi-cubic structures at room temperature with corner sharing $[PbI_6]^{4-}$ octahedra and with A^+ cations (MA, FA, Cs) filling octahedral voids. Such structures satisfy the Goldschmidt tolerance factors required of ideal cubic structures ($0.9 \leq t \leq 1.0$) and for structures having tilted octahedra ($0.7 < t < 0.9$) [1]. The compositional diversity of mixed cation and mixed anion systems is limited by the existence of non-perovskite (δ_{ortho} - and δ_{hex} -phases) phases that appear when A^+ ionic radii are insufficient to stabilize interstitial voids in the structure [2, 3]. Although $APbX_3$ perovskites can adopt other (e.g.,

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orthorhombic) crystal structures at different temperatures [4], we focus on the photophysical properties of photovoltaically relevant cubic/pseudo-cubic structures in what follows.

The primary motivation for investigating and ultimately understanding the optical response of APbX_3 perovskites stems from their successful implementation in high-efficiency photovoltaics. Today, perovskite solar cell power conversion efficiencies (PCEs) routinely exceed 20%. A maximum PCE of 25.7% has been reported in NREL's benchmark efficiency chart [5] and will undoubtedly be supplanted shortly. These values collectively represent a remarkable rise of perovskite solar cell efficiencies given their modest starting value of 3.8% in 2009. In short, APbX_3 perovskite solar cells are, from a PCE perspective, on par with crystalline silicon.

Responsible for this success are extraordinary and fortuitous perovskite properties. This entails facile solution processability, crystallinity despite low temperature processing, chemical and stoichiometric diversity, and large absorption efficiencies, all simultaneously coupled to low exciton binding energies, large carrier mobilities, and favorable energetics to engender defect tolerance. However, despite extensive research into improving perovskite solar cell PCEs, performance bottlenecks still remain. This prevents them from reaching their full Shockley-Queisser efficiency of ~31% for single-junction devices. A need therefore exists to fully understand the fundamental optical and electrical properties of APbX_3 systems to realize their ultimate performance potentials.

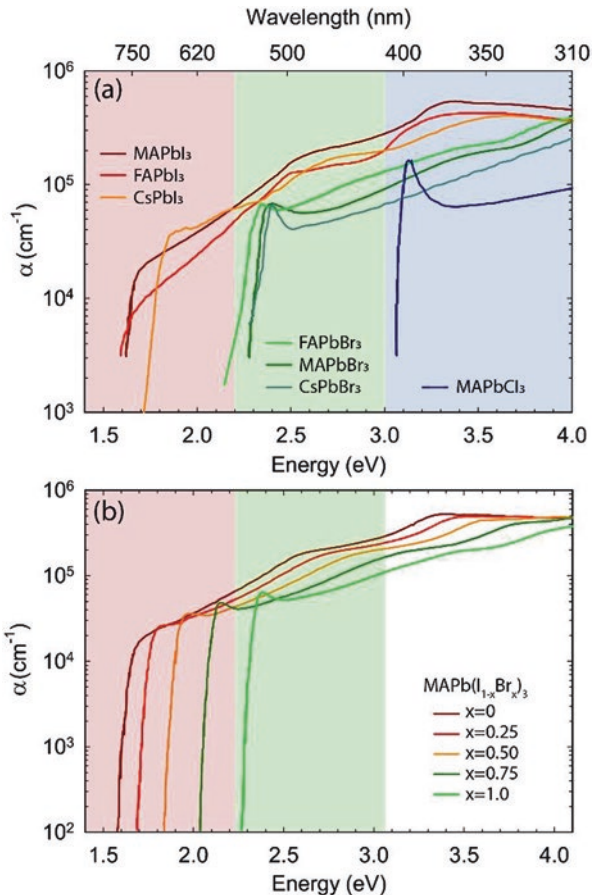
2 Absorption

A key feature of lead halide perovskites is their favorable absorption properties. This includes sizable absorption coefficients (α), band edges close to the desired Shockley-Queisser value of 1.55 eV, and tunable absorption edges in mixed halide alloys. Figure 1a highlights these features by showing reported absorption spectra for common APbX_3 systems.

In the red, at approximately 1.6 eV lie MAPbI_3 , FAPbI_3 , and CsPbI_3 . Near 2.2 eV are MAPbBr_3 , FAPbBr_3 , and CsPbBr_3 . Further to the blue at ~3.1 eV is MAPbCl_3 . The figure makes apparent that perovskite band gaps are sensitive to the choice of halide anion, whether I^- , Br^- , or Cl^- . This has previously suggested that the A-site cation plays a lesser role in determining the optical response of these materials. Instead, optical transitions are thought to be primarily established by perovskite's inorganic $[\text{PbI}_6]^{4-}$ framework [10]. This is supported by electronic structure calculations, which suggest A-site cation-related states being energetically removed from corresponding band edges. Cation-influenced octahedral tilting and lattice contraction [11] do, however, influence band edge energies, as evidenced by measurements on mixed cation perovskites such as $\text{MA}_{1-x}\text{FA}_x\text{PbI}_3$ or $\text{FA}_{1-x}\text{Cs}_x\text{PbI}_3$ where band gaps can be altered using cation stoichiometry [12–15].

Figure 1b further illustrates the sensitivity of perovskite band gaps to halide stoichiometry by showing how increasing the Br fraction (x) in a $\text{MAPb}(\text{I}_{1-x}\text{Br}_x)_3$ alloy

Fig. 1 Summary of the absorption spectra of various APbX_3 perovskites. Data extracted from References [6–9]. Spectral band edges highlighted using colored regions



causes its E_g to progressively shift towards the MAPbBr_3 limit. Analogous trends are observed with $\text{MAPb}(\text{Cl}_{1-x}\text{Br}_x)_3$ [16, 17] as well as with $\text{FAPb}(\text{I}_{1-x}\text{Br}_x)_3$ [18]. The formation of continuous $\text{MAPb}(\text{I}_{1-x}\text{Cl}_x)_3$ alloys is prevented by large differences in I^- and Cl^- ionic radii such that little if any Cl incorporation is achieved. Consequently, such systems are denoted $\text{MAPbI}_3(\text{Cl})$ in what follows [19]. This ability to compositionally tune band gaps makes mixed halide alloys of potential use in tandem (perovskite/silicon) solar cells.

Figure 1 summarizes the absorption coefficients of these materials. Evident are sizable band edge values, which lie between 10^4 and 10^5 cm^{-1} . These α -values correspond favorably to those of other photovoltaically relevant semiconductors. To illustrate, GaAs has an absorption coefficient of $\alpha \sim 10^4 \text{ cm}^{-1}$ at its absorption edge. References [6, 20–22] highlight this favorable comparison by visually illustrating perovskite α -values relative to those of other semiconductors across a range of energies.

Table 1 summarizes compiled E_g and α -values for the various APbX_3 perovskites being discussed. Apart from the general trends noted above, there is a sizable variability in reported values. MAPbI_3 band gaps, for instance, range from 1.5 to

Table 1 Absorption parameters of various lead halide perovskites

System	E_g (eV)	E_b (meV)	$\alpha_{2.0\text{ eV}}$ ($\times 10^5\text{ cm}^{-1}$)	$\alpha_{2.3\text{ eV}}$ ($\times 10^5\text{ cm}^{-1}$)	$\alpha_{3.1\text{ eV}}$ ($\times 10^5\text{ cm}^{-1}$)
MAPbI ₃	~1.5 [23]	1.7–2.1 [31]	0.23 [24, 32]	0.47 [32]	1.09 [29]
	1.55 [24, 25]	6 [41]	0.24 [21]	0.51 [29]	2.04 [25]
	1.553 [26]	7.4 [42]	0.26 [29]	0.53 [24]	2.34 [21]
	1.557 [27]	9 [43]	0.38 [31, 55]	0.59 [21]	2.43 [58]
	1.56 [28, 29]	12.3 [40]	0.39 [36]	0.64 [31, 33]	2.56 [24, 35]
	1.574 [30]	13.1 [44]	0.44 [56]	0.86 [36]	2.76 [31]
	1.59 [21, 31]	14–18 [45]	0.45 [26]	0.91 [58]	2.80 [32]
	1.6 [32–35]	16–22 [46]	0.46 [23]	0.93 [25]	2.86 [26]
	1.61 [36–38]	25 [47]	0.47 [33, 35]	0.94 [55]	2.89 [23]
	1.63 [39]	29 [48]	0.50 [57]	0.95 [35]	2.94 [33]
	1.646 [40]	32 [49]	0.58 [58]	1.01 [23]	2.99 [28]
		37 [50]	0.60 [20]	1.04 [20, 26]	3.02 [59]
		40 [51]	0.66 [30, 38]	1.19 [56]	3.15 [36]
		45 [52]	0.67 [20]	1.22 [38]	3.24 [37]
		50 [53]	0.68 [25]	1.34 [59]	3.31 [38]
		55 [54]	0.73 [39]	1.39 [46]	3.33 [56]
		0.75 [59]	1.44 [39]	3.38 [55]	
		0.78 [46]	1.47 [30]	3.51 [46]	
		0.79 [37]	1.50 [37]	3.68 [39]	
		0.92 [28]	1.55 [28]	4.19 [34]	
	1.70 [34]	2.88 [34]	4.51 [30]		
FAPbI ₃	1.45 [22]	8.4 [40]	0.25 [55]	0.71 [55]	3.59 [55]
	1.55 [55]	10 [60]	0.43 [22]	1.62 [22]	4.52 [22]
	1.566 [40]	31.6 [61]			
CsPbI ₃	1.738 [8]	20 [62]	–	0.71 [8]	2.34 [8]
FAPbBr ₃	2.351 [40]	24 [60]	–	0.74 [63]	1.86 [63]
		31.8 [40]			
MAPbBr ₃	2.24 [24]	14 [65]	–	0.08 [55]	0.73 [64]
	~2.3 [56]	30.9–36.3		0.12 [24]	1.04 [24]
	2.309 [64]	[40]		0.13 [56]	1.11 [55]
	2.384–2.392	40.3 [44]		0.43 [64]	1.24 [56]
	[40]	41 [66]		0.91 [63]	1.77 [63]
	76 [53]				
	150 [67]				
CsPbBr ₃	2.3 [9]	40 [62]	–	0.02 [68]	0.78 [9]
	2.37 [68]			0.13 [9]	0.89 [69]
	2.39 [69]			0.23 [63]	1.02 [68]
	2.4 [70]			0.31 [69]	1.36 [63]
			0.57 [70]	1.45 [70]	
MAPbCl ₃	2.97 [24]	41 [71]	–	–	0.35 [24]
		50 [72]			1.0 [55]

1.646 eV. This is also true of MAPbBr₃ where E_g -values range from 2.24 to 2.392 eV. In either case, E_g spreads are of order 150 meV.

Associated absorption coefficients are also highly variable, as evident from tabulated α -values compiled at three different energies (2.0 eV, 2.3 eV, and 3.1 eV). In particular, Table 1 shows that MAPbI₃ α -values at 2.0 eV range from 0.23 to

$1.70 \times 10^5 \text{ cm}^{-1}$. At 2.3 eV, α -values range from 0.47 to $2.88 \times 10^5 \text{ cm}^{-1}$. Analogous variations exist with other perovskites. This variability and lack of accord are summarized visually in References [6, 36, 40, and 58].

Many reasons exist for apparent differences in reported optical parameters. Much has to do with variations in sample quality that stem from the numerous approaches used to prepare perovskite thin films. They include solution deposition (doctor blading, spray coating, slot-die coating, inkjet printing, etc.), solution deposition with solvent recrystallization (two-step spin-coating or antisolvent treatment), hot casting, chemical vapor deposition, and low-pressure vapor-assisted solution processing [73, 74]. Sample quality variability is especially highlighted when thin films are compared to APbX_3 single crystals, which possess larger grains, reduced morphological disorder, and correspondingly reduced surface roughness [75].

Consequently, what results are thin film/single crystal specimens that possess varying degrees of crystallinity, thicknesses, apparent grain sizes, surface roughness, etc. All lead to measurement variations. Fujiwara [6], for example, attributes reported E_g and α -value variations to surface roughness that introduces significant light scattering to spectroscopic ellipsometry measurements. This degrades subsequent model extraction of perovskite optical constants (i.e., frequency-dependent refractive indices and permittivities) and leads to an underestimation of perovskite band gaps. When such surface roughness variations are explicitly accounted for, a closer agreement between researcher-reported α and E_g values is realized.

Beyond band gaps and absorption coefficients, Fig. 1 reveals other intriguing features of APbX_3 materials. For MAPbI_3 , FAPbI_3 and CsPbI_3 , band edges resemble those of bulk, direct gap semiconductors. No apparent excitonic features are seen. The absence of an excitonic response is corroborated by numerous estimates of their exciton binding energies (E_b), as summarized in Table 1. These estimates arise from magnetoabsorption and temperature-dependent emission and absorption measurements as well as from modeling experimental APbX_3 band edge absorption spectra with Elliott's model [76].

Table 1 shows a spread of reported E_b values. As with E_g and α , large variations can be seen where for MAPbI_3 E_b ranges from 1.7 to 55 meV. Despite this, reported binding energies are of order kT and suggest that the optical response of MAPbI_3 and FAPbI_3 can be described in terms of free carriers. This conclusion is corroborated by time-resolved emission, transient differential absorption, and THz conductivity studies [76]. Note that this is not necessarily true of Br- and Cl-based APbX_3 materials such as CsPbBr_3 or MAPbCl_3 where prominent band edge excitonic features are seen in the linear absorption. The suggestion is supported by their generally larger E_b -values in Table 1.

The specific origin of the optical transitions in lead halide perovskites has been the subject of numerous studies. Most entail density functional theory calculations to varying degrees of approximation [24, 36, 77]. Without delving into specifics, consensus exists that valence to conduction band transitions occurs at the R symmetry point and involves valence bands that originate from the antibonding combination of halide p and Pb(6s) orbitals. Corresponding conduction bands largely arise from Pb(6p) orbitals [77]. These bands are also responsible for above gap

transitions and produce apparent structure at higher energies. For example, a feature in the absorption spectrum of MAPbI₃ close to 2.5 eV (Fig. 1) is attributed to a valence/conduction band transition at the M symmetry point [24, 76, 77]. The anti-bonding nature of the APbX₃ valence band is supported by apparent increases in perovskite band gap with increasing temperature. This contrasts itself to the response of traditional, tetrahedrally coordinated semiconductors where band gaps decrease (increase) with increasing (decreasing) temperature.

3 Carrier Dynamics

Having briefly summarized the general absorptive properties of APbX₃ perovskites, we now provide insight into their carrier recombination processes, following photoexcitation. This is important since the fate of photogenerated carriers is fundamental to device operation and ultimately to their efficiencies. A kinetic model is therefore developed that qualitatively and quantitatively rationalizes the intrinsic photophysics of APbX₃ systems [47, 78]. This includes experimental observations of photoluminescence, time-correlated emission decays, and transient differential absorption dynamics. In addition to numerical simulations, analytical approximations are provided to better visualize qualitative trends in both emission intensities and time-correlated decays. Although the model does not explicitly consider device operation, interested readers may refer to Reference [79] for an extension that includes charge extracting interfaces. Such a model has been used to establish the performance bottleneck(s) of a high-efficiency FAMACs solar cell.

In general, the primary recombination processes considered are (a) bimolecular (radiative) electron-hole recombination, (b) carrier trapping, and (c) nonradiative Auger recombination. The latter is nominally only important at high carrier densities, far beyond 1 sun conditions. These processes are summarized in Fig. 2 with the

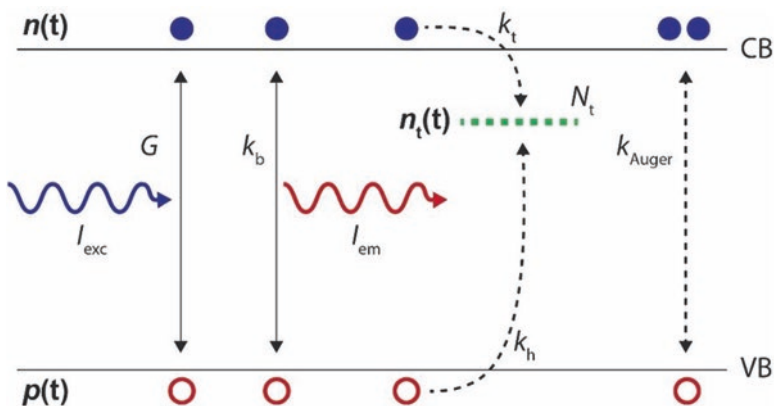


Fig. 2 Schematic illustration of photophysical processes occurring in lead halide perovskites, following photoexcitation

illustration showing photoexcitation creating transient electron and hole populations [$n(t)$ and $p(t)$] in the perovskite conduction and valence bands. Carriers subsequently recombine via the three processes outlined above. Although the identity of APbX_3 trap states remains debated, there appears to be some agreement that such states are electron traps. This is assumed in what follows. Note that other versions of this model exist, which account for exciton dissociation, unintentional doping, and carrier diffusion. The interested reader is therefore referred to References [80], [81], and [49] for details.

Kinetic expressions that summarize the model are

$$\begin{aligned}\frac{dn}{dt} &= G - k_t (N_t - n_t) n - k_b np - \frac{1}{2} k_{\text{Auger}} [n^2 p + np^2] \\ \frac{dp}{dt} &= G - k_h n_t p - k_b np - \frac{1}{2} k_{\text{Auger}} [n^2 p + np^2] \\ \frac{dn_t}{dt} &= k_t (N_t - n_t) n - k_h n_t p\end{aligned}\quad (1)$$

where G is the initial charge generation rate ($\text{cm}^{-3} \text{s}^{-1}$), linked to the excitation intensity (I_{exc} , W cm^{-2}), i.e., $G = \frac{I_{\text{exc}} \alpha}{h\nu}$ with α the absorption coefficient (cm^{-1}) and $h\nu$ the photon energy (J), n (p) is the electron (hole) concentration (cm^{-3}) in the perovskite conduction (valence) band, k_t is an electron trapping rate constant ($\text{cm}^3 \text{s}^{-1}$), N_t is a corresponding trap density (cm^{-3}), n_t is the trap population (cm^{-3}), k_b is a bimolecular radiative recombination rate constant ($\text{cm}^3 \text{s}^{-1}$, referred to as k_2 in the literature), and k_{Auger} is the Auger, three-carrier rate constant ($\text{cm}^6 \text{s}^{-1}$, referred to as k_3 in the literature).

Numerous studies now provided estimates for the various rate constants in Eq. (1) and Fig. 2. These literature estimates are summarized in Table 2 across various APbX_3 systems. An inspection shows that most work has focused on MAPbI_3 and $\text{MAPbI}_3(\text{Cl})$ thin films with relatively less work carried out on corresponding FA-based materials.

Table 2 also makes apparent that while variations in rate constants exist across systems and even within a given material, there is general consistency in their values. Bimolecular radiative rate constants are of order $10^{-10} \text{ cm}^3 \text{ s}^{-1}$, while Auger rate constants are of order $10^{-28} \text{ cm}^6 \text{ s}^{-1}$ [85]. Measured pseudo-first-order rate constants for electron trapping are of order $k_t \sim 10^7 \text{ s}^{-1}$ from where corresponding k_t values are of order $k_t \sim 10^{-10} \text{ cm}^3 \text{ s}^{-1}$, provided estimated trap densities of order $N_t \sim 10^{-16} \text{ cm}^{-3}$.

The general photogenerated carrier dynamics, predicted by Eq. (1) at different excitation intensities, are now illustrated. Implicit to the discussion is continuous wave (CW) excitation. An identical analysis can be conducted assuming pulsed excitation. However, this is not pursued here since common applications of perovskite materials generally entail CW excitation conditions. Interested readers may refer to References [97] and [98] for details of a pulsed excitation analysis.

Table 2 Literature rate constants for various lead halide perovskites

	MAPbI ₃	MAPbI ₃ (Cl)	MAPbBr ₃	FAPbI ₃	FAPb(I _{1-x} Br _x) ₃	FAPbBr ₃	FAMACs
$k_1 \approx k_e N_i$ ($\times 10^7 \text{ s}^{-1}$)	0.091 (TOPO) [82] 0.66 [82] 0.72–1.3 [83] 1.4–1.5 [84] 1.5 [85, 86] 1.8 [87]	0.49 [84] 1.2 [85] 1.2 [88]	–	0.7 [85]	–	2.1 [85]	0.4 [89]
N_i ($\times 10^{16} \text{ cm}^{-3}$)	10^{-4} – 10^{-3} (emissive) [90] 0.90–86 [78] 5.0–16 [59] 6.0 [80]	<0.05 [80] <0.1–4.0 [91] 1.0 [92] 2.5 [81]	–	–	–	–	0.54–60 [79]
k_6 ($\times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$)	0.1–1.0 [86] 0.35 [82] 0.4 (TOPO) [82] 0.60 [85] 1.5 [93] 1.7 [87] 3.5 [80] 6.0 [94] 6.2–37 [83] 9.2–9.4 [84] 13 [49] 23 [95]	0.23–0.78 [91] 0.87 [84] 1.1 [85, 88] 4.9 [80] 11 [96]	4.9 [93]	1.1 [85]	1.0–30 [18] ($x = 0$ – 1)	11 [85]	0.30 [89]
k_{Auger} ($\times 10^{-28} \text{ cm}^6 \text{ s}^{-1}$)	0.37–1.3 [84] ~ 1 [86] 1.6 [85] 3.4 [93] ~ 10 [94]	0.20 [85] 0.23 [88] 0.99 [84] 3.8 [96]	13.5 [93]	0.20 [85]	0.20–1.5 [18] ($x = 0$ – 1)	1.5 [85]	0.01 [89]

4 Low Excitation Intensities

At low excitation intensities, carrier trapping dominates radiative recombination. Low photogenerated carrier densities further mean that nonradiative Auger pathways can be ignored. Consequently, under the assumption that the material is intrinsic, Eq. (1) reduces to

$$\begin{aligned}\frac{dn}{dt} &\approx G - k_t N_t n \\ \frac{dp}{dt} &\approx G - k_h p^2\end{aligned}$$

The equations make apparent that at steady state $n = \frac{G}{k_t N_t} \propto G$ and $p = \sqrt{\frac{G}{k_h}} \propto \sqrt{G}$. Since the emission rate and corresponding intensity, I_{em} , are proportional to the product of n and p , I_{em} grows with increasing I_{exc} (G) in a power law fashion. Its growth exponent is $\sim 3/2$, i.e., $I_{em} \propto G^{3/2}$.

Figure 3a illustrates this for the case where $N_t = 10^{16} \text{ cm}^{-3}$. Employed rate constants for the numerical simulation of Eq. (1) are $k_b = 10^{-10} \text{ cm}^3 \text{ s}^{-1}$, $k_t = 10^{-9} \text{ cm}^3 \text{ s}^{-1}$, $k_h = 10^{-11} \text{ cm}^3 \text{ s}^{-1}$, and $k_{Auger} = 10^{-28} \text{ cm}^6 \text{ s}^{-1}$. The model therefore reveals that I_{em} grows as $G^{1.6}$ at low G where recombination is primarily trap-mediated (shaded red region). Figure 3b shows identical behavior for simulations where N_t has been varied between $N_t = 10^{15}$ and 10^{18} cm^{-3} . In all cases, fit-extracted growth exponents range from $m = 1.5$ to 1.6. Of note is the increasing range of G -values where $m \sim 1.5$. This is rationalized by delayed trap saturation due to larger N_t -values.

Figure 3c shows experimental data for a MAPbI₃ single crystal and thin film [78] that corroborate these model predictions. Acquired over ~ 3 orders of magnitude in G , the data reveal $m \sim 1.5$ power law growth exponents for either material, as established by fits to low G I_{em} -values (dashed lines with fit-extracted m -values shown). Differences in the range of I_{em} -values over which $m \sim 1.5$ qualitatively agree with Fig. 3b and suggest underlying N_t -value variations between MAPbI₃ single crystals and thin films. Other data acquired on MAPbI₃ and MAPbI₃(Cl) over 8 orders of magnitude in G reveal identical $m \sim 1.5$ power law growth exponents at low G [47]. This further corroborates the analysis and conclusions drawn here. Note that under (low irradiance) *pulsed* excitation, analogous power law growth of the integrated emission intensity is predicted with an ideal model growth exponent of $m = 2.0$ [97, 98]. Such quadratic emission growth has previously been observed for MAPbI₃ and MAPbI₃(Cl) thin films [47].

Next, by assuming above-simplified kinetic expression for n and p , corresponding (normalized) photoluminescence decays take the form

$$PL(t) = \frac{I_{em}(t)}{I_{em,max}} \approx \frac{e^{-k_t N_t t}}{1 + k_h G t}. \quad (2)$$

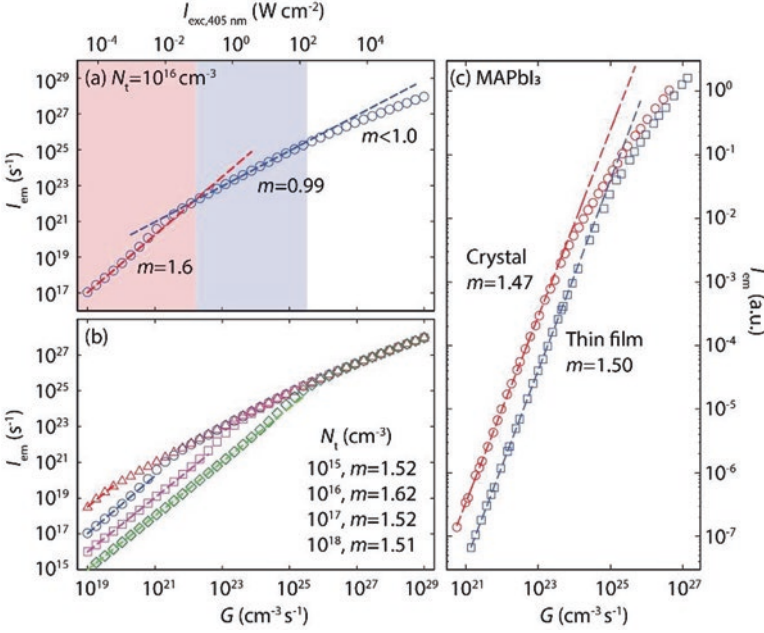


Fig. 3 (a) I_{em} versus G for $N_t = 10^{16} \text{ cm}^{-3}$. Dashed lines are linear fits to the data in a given growth regime (shaded regions). Fit-extracted growth exponents shown. G -values are linked to associated I_{exc} , assuming 405 nm excitation with $\alpha = 10^5 \text{ cm}^{-1}$, for reference purposes. (b) I_{em} versus G for N_t varying between $N_t = 10^{15} - 10^{18} \text{ cm}^{-3}$. Dashed lines are linear fits in the trap-mediated recombination regime. Fit-extracted growth exponents shown. (c) I_{em} versus G data for a MAPbI₃ single crystal and thin film. Data from Reference [78]. Dashed lines are fits to the data at smaller G with fit-extracted growth exponents shown

At small G or short times, the numerator in Eq. (2) dominates. Decays are therefore near exponential with an associated pseudo-first-order decay constant of $k_t N_t \sim 10^7 \text{ s}^{-1}$ (Table 2). Figure 4a illustrates this for the model predicted decay when $G = 10^{15} \text{ cm}^{-3}$. An accompanying (superimposed) dotted line is Eq. (2). This qualitative prediction is supported by numerous studies, which report near exponential decays at low I_{exc} for various perovskite systems [80, 81, 84, 87, 99].

5 Intermediate Excitation Intensities

As I_{exc} increases, progressive trap filling ($n_t \rightarrow N_t$) reduces the electron trapping rate such that $n \rightarrow p$. This simplifies Eq. (1) and leads to the following effective rate expressions:

$$\frac{dn}{dt} \approx G - k_b n^2$$

$$\frac{dp}{dt} \approx G - k_h N_t p - k_b p^2.$$

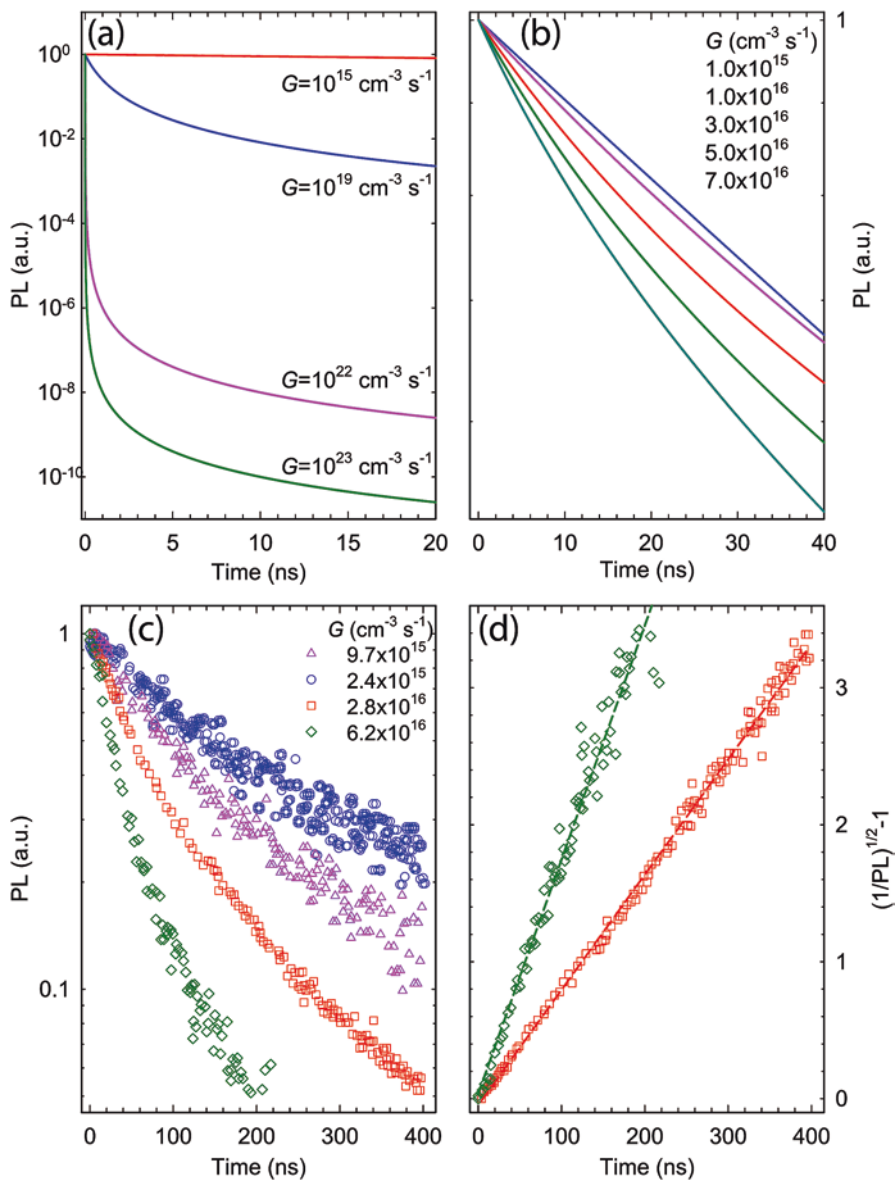


Fig. 4 (a) Predicted PL decays from Eq. (1) for variable G between $G = 10^{15}$ and 10^{23} cm^{-3} . Superimposed over the data are analytical predictions from Eqs. (2)–(4) (dotted lines). (b) Predicted decays for $G = 10^{15}$ – 7×10^{16} cm^{-3} . (c) MAPbI₃ PL decays from Reference [87]. (d) Large G MAPbI₃ PL decays from Reference [87] replotted. Dashed lines are linear fits to the data

Invoking steady state conditions then means that $n = \sqrt{\frac{G}{k_b}} \propto \sqrt{G}$ with $p = \frac{-k_t N_t + \sqrt{(k_t N_t)^2 + 4G}}{2} \propto \sqrt{G}$. The emission intensity therefore transitions to power law growth with a growth exponent of $m \sim 1$. Figure 3a illustrates this transition by showing a fit to the model data, highlighted in the shaded blue region. A fit-extracted power law growth exponent of $m = 0.99$ is found. Analogous behavior is observed in Fig. 3b when N_t is varied between 10^{15} and 10^{18} cm^{-3} . In whole, the model data makes evident that increasing N_t extends the region of trap-mediated recombination ($m \sim 1.5$) to larger G -values before bimolecular (radiative) recombination ($m \sim 1.0$) dominates carrier dynamics following photoexcitation.

Using the above simplified rate expressions, associated (normalized) photoluminescence decays adopt the following bimolecular form:

$$\text{PL}(t) = \frac{I_{\text{em}}(t)}{I_{\text{em,max}}} = \frac{1}{(1 + k_b G t) [1 + (k_t N_t + k_b G t)]} \quad (3)$$

where Fig. 4a shows model-predicted decays for $G = 10^{19}$ and $G = 10^{22} \text{ cm}^{-3}$ using the same parameters employed earlier at low I_{exc} . Analytical results from Eq. (3) are superimposed atop the model decays and are in excellent agreement. Model decays also highlight the transition in kinetic response in this I_{exc} regime. Figure 4b illustrates this for G -values between $G = 10^{16}$ and 10^{17} cm^{-3} where for smaller G -values, near exponential behavior is seen. With increasing G , an apparent near exponential to bimolecular transition occurs. Such $\text{PL}(t)$ transitions have previously been reported in the literature [80, 81, 87, 99] and an example from a MAPbI₃ thin film is shown in Fig. 4c.

6 High Excitation Intensities

At sufficiently high excitation intensities, trap saturation causes bimolecular (radiative) recombination to dominate. In this limit, Eq. (1) effectively becomes

$$\begin{aligned} \frac{dn}{dt} &\approx G - k_b n^2 \\ \frac{dp}{dt} &\approx G - k_b p^2 \end{aligned}$$

so that $n = p = \sqrt{\frac{G}{k_b}} \propto \sqrt{G}$. I_{em} thus continues to grow in a power law fashion with a growth exponent of $m \sim 1$. Figure 3 illustrates this for $N_t = 10^{16} \text{ cm}^{-3}$ (Fig. 3a) and across N_t -values between $N_t = 10^{15}$ and 10^{18} cm^{-3} (Fig. 3b). It should be mentioned that a growth exponent of $m \sim 1$ is common to this intensity regime under pulsed excitation [47, 97, 98].

An associated (normalized) photoluminescence decay takes the bimolecular form

$$\text{PL}(t) = \frac{I_{\text{em}}(t)}{I_{\text{em,max}}} \approx \frac{1}{(1 + k_b G t)^2}, \quad (4)$$

which is near identical to the expression derived earlier for intermediate excitation intensities (Eq. 3). Figure 4a plots a model-predicted decay for $G = 10^{23} \text{ cm}^{-3}$ with Eq. (4) superimposed. Again, there is excellent agreement with the analytical approximation.

Beyond bimolecular fits, predicted bimolecular decays can be confirmed visually by plotting $\sqrt{\frac{I_{\text{em,max}}}{I_{\text{em}}(t)} - 1}$ versus time. In this case, linear behavior is expected [100]. Figure 4d illustrates this for the large G experimental data in Fig. 4c. Evident linear behavior is observed, as highlighted by dashed, linear fits.

Above this excitation regime, the onset of Auger-mediated nonradiative recombination causes emission efficiencies to decrease. This stems from competitive, nonradiative recombination of carriers. What results then is sublinear growth of I_{em} with an associated power law growth exponent $m < 1$. Figure 3a explicitly illustrates the onset of Auger recombination for $N_t = 10^{16} \text{ cm}^{-3}$ at high G where the simulated data shows $m < 1$. It can also be shown that in this regime, plotting $\left[\frac{I_{\text{em,max}}}{I_{\text{em}}(t)} - 1 \right]$ versus time yields linear behavior [100].

7 Emission Quantum Yields

Equation (1) simultaneously allows internal emission quantum yields (QYs) to be estimated through the ratio of the bimolecular radiative recombination rate to the initial carrier generation rate, i.e.

$$\text{QY} = \frac{k_b n p}{G}. \quad (5)$$

The importance of this metric is that high emission efficiencies are requisite for optimizing APbX₃ photovoltaic performance. More specifically, it is the associated external quantum efficiency (EQE, $\text{EQE} = \eta_e \text{QY}$ where η_e is a photon extraction efficiency) that is fundamentally linked to the maximum open circuit voltage (and PCE) achievable in a solar cell [101]. The seemingly contradictory conclusion that arises then is that a good solar cell must also be a good light emitter [102].

Figure 5 shows model-predicted (internal) QYs plotted as functions of N_t when N_t is varied between 10^{15} cm^{-3} and 10^{18} cm^{-3} . As with Figs. 3 and 4, employed rate constants have been kept constant at their nominal literature values of $k_b \sim 10^{-10} \text{ cm}^3 \text{ s}^{-1}$, $k_t \sim 10^{-9} \text{ cm}^3 \text{ s}^{-1}$, and $k_n \sim 10^{-11} \text{ cm}^3 \text{ s}^{-1}$. Fig. 5 also provides model-predicted EQEs via $\text{EQE} = \frac{\eta_e (k_b n p)}{G}$.

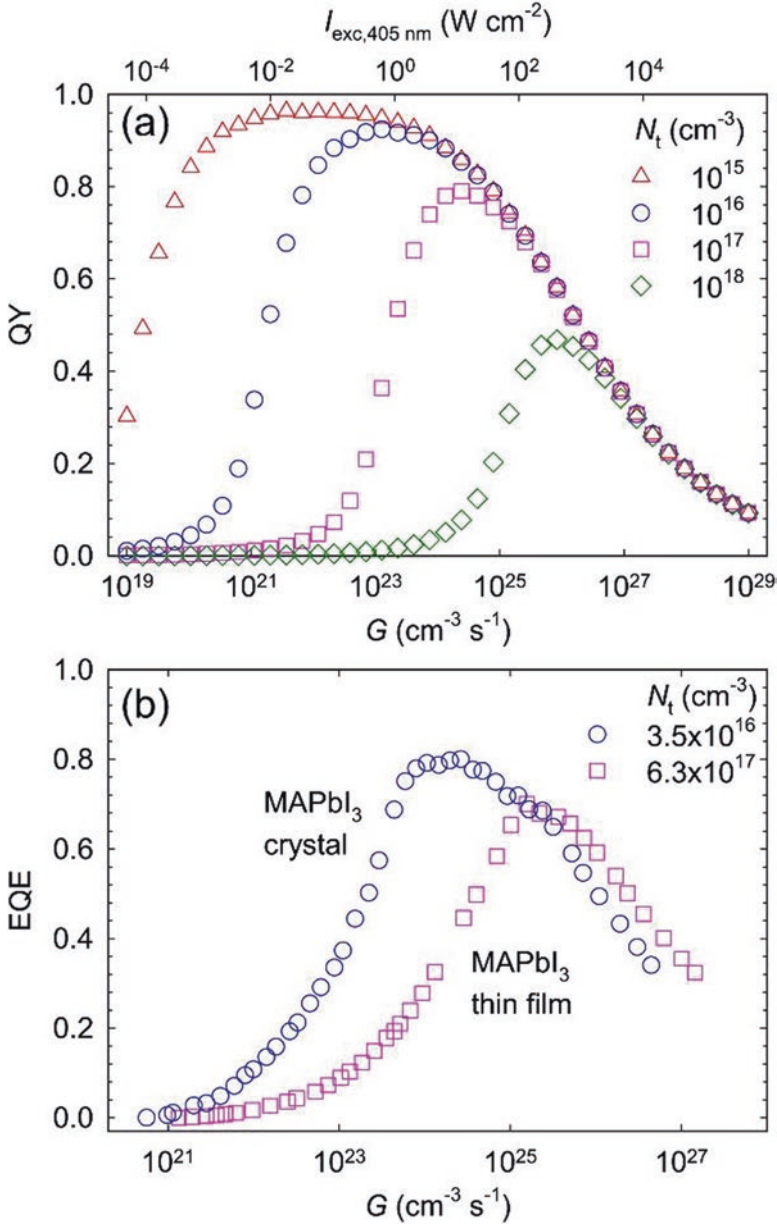


Fig. 5 (a) Predicted (internal) QYs for variable G between $G = 10^{19}$ – 10^{27} cm^{-3} and N_t between 10^{15} and 10^{18} cm^{-3} . For reference purposes, G -values are linked to associated I_{exc} assuming 405 nm excitation with $\alpha = 10^5 \text{ cm}^{-1}$. (b) Experimental EQE estimates for a MAPbI₃ single crystal and thin film. Data from Reference [78]. Fit extracted N_t -values shown

Table 3 Literature APbX₃ emission EQEs under effective 1 sun conditions. Additives shown in parenthesis: K = potassium, Rb = rubidium, TOPO = trioctylphosphine oxide, pyr = pyridine, TPPO = triphenylphosphine oxide, SPPO = 9,9-spirobifluoren-2-yl-diphenyl-phosphine oxide)

Material	EQE film (%)	EQE film treated (%)	EQE device (%)
MAPbI ₃	0.92 [103] 3 [82] 5 [104] 5.8 [78] 10 [105]	12–13 (TOPO) [106] 20.3 (TOPO) [103] 35 (TOPO) [82]	5 [106]
MAPbI ₃ (Cl)	20 [104] 50–60 [107]	24–25 (pyr) [108]	–
MAPbBr ₃	2–3 [109] 15 [104]	1.2 [110] 32 (TPPO) [111] 85 (SPPO) [111]	–
FAMACs	2.4 [112] 66 (K) [113]	3.6 (Rb) [112]	0.43–2.83 [79] 14.5 (K) [113]
FAMAPb(I _{1-x} Br _x) ₃	-	-	0.1–3.0 [114]

The resulting figure illustrates several things. First, as discussed earlier, at low I_{exc} , trapping dominates carrier recombination. This leads to low QYs. Attesting to this are experimental perovskite, 1 sun EQEs in Table 3. Values for thin films, (surface) treated thin films, and devices are shown. Inspection of the data reveals that reported 1 sun EQEs are generally on the order of several percent with some notable exceptions. This is consistent with the lower QYs seen at low G in Fig. 5. Of note is that this bulk perovskite data contrasts itself to analogous results, summarized in Reference [115] for perovskite nanocrystals (NCs). In these materials, EQEs regularly approach or attain unity values.

Next, Table 3 shows that treating lead halide perovskite thin films with Lewis bases such as trioctylphosphine oxide (TOPO) or pyridine improves their EQEs. However, they only increase values to numbers of order 10%. This indicates that while defect passivation is possible, significant trap densities still remain. This is again unlike the case of perovskite NCs where effective surface passivation does appear possible and which leads to unity EQEs [116]. Finally, Table 3 shows that device EQEs are all suppressed from corresponding thin film or treated thin film values due to the inevitable competition for carriers by electron and hole extraction interfaces present in working devices.

Figure 5 shows that maximal QYs are achieved upon trap saturation at high excitation intensities. In the case where $N_t = 10^{16} \text{ cm}^{-3}$, near unity internal QYs are realized close to 1 sun conditions. The figure further shows that increasing N_t simply leads to progressively larger I_{exc} -values required to achieve peak QYs, which themselves become progressively smaller. In all cases, maximum QYs persist until a critical G beyond which the onset of Auger-mediated carrier recombination causes them to fall as discussed earlier.

This QY behavior has previously been modeled by reducing the kinetics in Eq. (1) to [85].

$$\frac{dN}{dt} = G - AN - BN^2 - CN^3. \quad (6)$$

In Eq. (6), N is an effective electron-hole density, A is a generic first-order rate constant that describes nonradiative, trap-mediated recombination of photogenerated carriers, B is a second-order (radiative) rate constant, and C is a third-order rate constant that accounts for nonradiative Auger recombination. The rate constants A , B , and C are effectively k_1 , k_b , and k_{Auger} in Table 2. A corresponding internal QY is then

$$\text{QY} = \frac{BN^2}{AN + BN^2 + CN^3}. \quad (7)$$

Equation (7) can be recast in terms of EQE via $\text{EQE} = \frac{\eta_e BN^2}{AN + \eta_e BN^2 + CN^3}$. In either case, a peak QY can be found by taking the derivative of QY (EQE) with respect to N to find a critical point (i.e., $\frac{d\text{QY}}{dN} = \frac{d\text{EQE}}{dN} = 0$) [85]. A resulting optimal carrier density (N_{opt}) is then

$$N_{\text{opt}} = \sqrt{\frac{A}{C}} \sim \sqrt{\frac{k_1}{k_{\text{Auger}}}} \quad (8)$$

and is linked to a corresponding maximum internal QY of

$$\text{QY}_{\text{max}} = \frac{1}{1 + \frac{2\sqrt{AC}}{B}} \sim \frac{1}{1 + \frac{2\sqrt{k_1 k_{\text{Auger}}}}{k_2}} \quad (9)$$

with a corresponding maximum EQE of $\text{EQE}_{\text{max}} = \frac{1}{1 + \frac{2\sqrt{AC}}{\eta_e B}} \sim \frac{1}{1 + \frac{2\sqrt{k_1 k_{\text{Aug}}}}{\eta_e k_2}}$.

Equations (6) and (7) can be used to model experimental EQEs to extract relevant rate constants. They are, however, not predictive in that tabulated rate constants from Table 2 cannot be used to estimate QYs and EQE a priori. This is because Figs. 3–5 show that trap saturation occurs due to $n_t \rightarrow N_t$. Consequently, the pseudo-first-order trapping rate constant (here A) is I_{exc} -dependent. Immediate application of Eqs. (8) and (9) therefore leads to predicted N_{opt} and QY_{max} smaller than those revealed by full numerical simulations of Eq. (1) (Fig. 5).

Finally, beyond trap-mediated recombination, an important reason for small overall EQE values and for why EQEs are smaller than internal QYs is photon trapping within the perovskite. This originates from refractive index differences with the surrounding medium and is most prominent in APbX_3 films and crystals where physical sizes approach the wavelength of light. This leads to estimated photon extraction efficiencies of $\eta_e \sim 10\%$ [104, 105]. Note that such trapping is not significant for NCs as they effectively behave as dipole emitters. This rationalizes why unity/near unity EQEs are readily seen with these materials [115].

8 Ion Migration

Finally, despite all of the notable properties of APbX_3 materials, preventing their widespread commercialization is their well-known environmental sensitivities. Addressing this have been a number of studies [117, 118]. Less recognized but equally important are intrinsic instabilities linked to ion migration. Namely, both cations and anions in APbX_3 materials are mobile with ion mobilities stemming from vacancy-mediated ion migration under bias or under illumination.

For cations, evidence for bias-induced A^+ migration comes from observed device hysteresis thought to stem from cation accumulation at electrodes [119–121]. The phenomenon is better illustrated using more direct measurements such as time-of-flight secondary ion mass spectrometry and super-resolution infrared absorption measurements [122, 123], which explicitly reveal cation migration and accumulation near electrodes.

For anions, a well-known phenomenon is light-induced photosegregation whereby shining light on mixed I^-/Br^- systems [e.g., $\text{MAPb}(\text{I}_{1-x}\text{Br}_x)_3$] induces halide segregation. This results in the formation of I-rich inclusions within parent, mixed halide perovskites. Such photoinduced halide segregation is experimentally observed as progressive photoluminescence redshifts due to photogenerated carriers being funneled to I-rich domains. Such domains act as emissive recombination centers because of favorable (valence) band offsets relative to those of parent mixed halide materials. Observed redshifts/photosegregation are reversible when specimens are no longer illuminated with remixing being entropically driven. References [4, 124, and 125] provide more details on the phenomenon.

At this point, an inevitable question that arises is the connection between the earlier photocarrier dynamics and the bias-/light-induced ion migration described here. Since ion migration, whether cation or anion, is thought to be vacancy-mediated and since point defects are likely responsible for carrier trapping, a connection between the two is suggested. Evidence for this can already be found in the literature where References [126, 127, and 128] already suggest that decreasing defect densities mitigates ion migration. Studies linking the two topics, however, are limited. Consequently, more work is required to establish a comprehensive picture of ion-inclusive carrier dynamics that follow photoexcitation of APbX_3 systems. As such, linking the two sets the direction for future investigations of these unique materials.

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